

Auxiliary slides for seminar at Rochester Institute of Optics June 19, 2018

Relawavity:

Helping to clarify Quantum Theory and Relativity
by effecting a wavy marriage
between this enigmatic pair

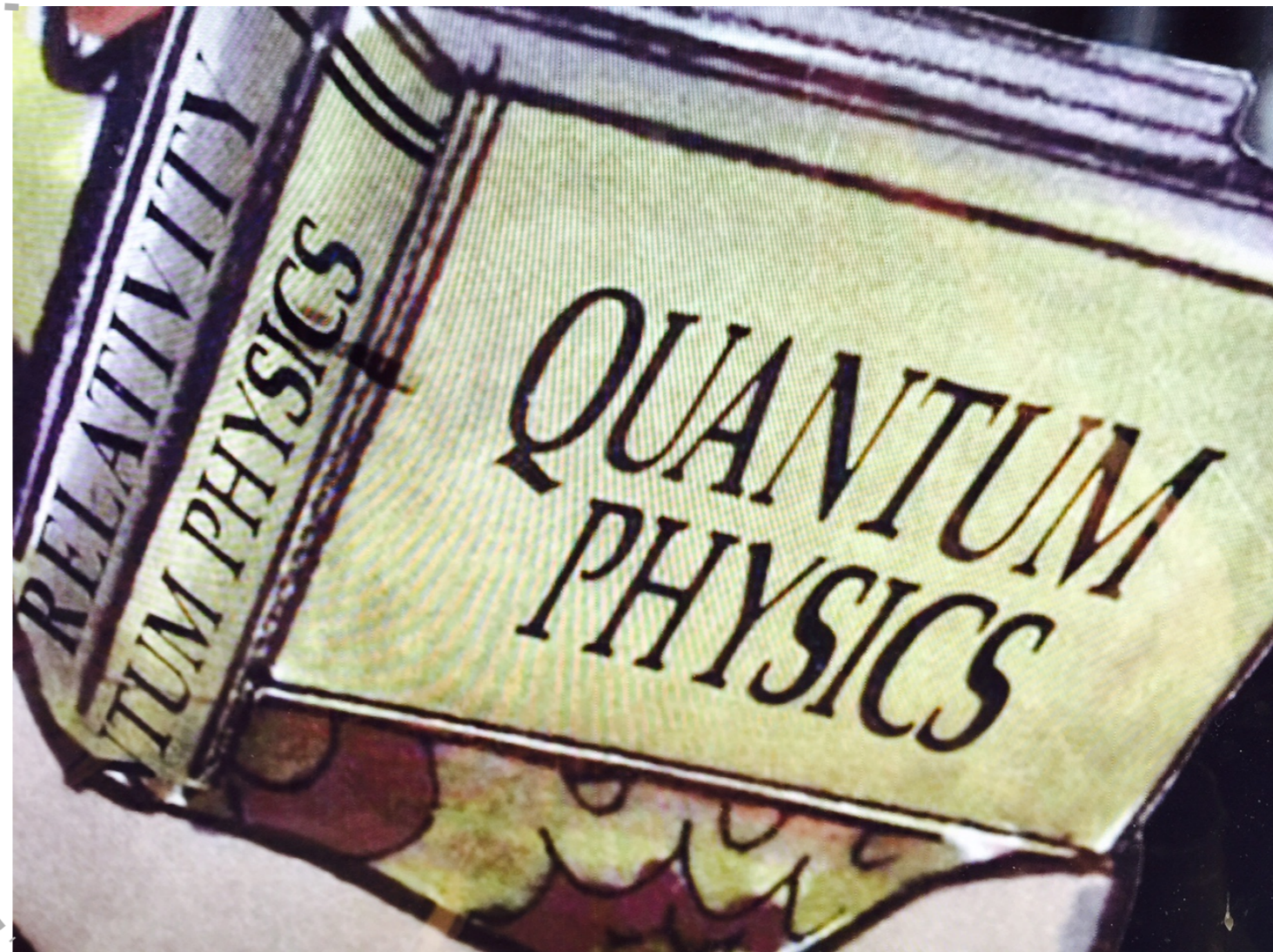


(Why a *Men In Black* candidate shot little Suzy)

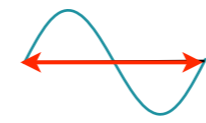
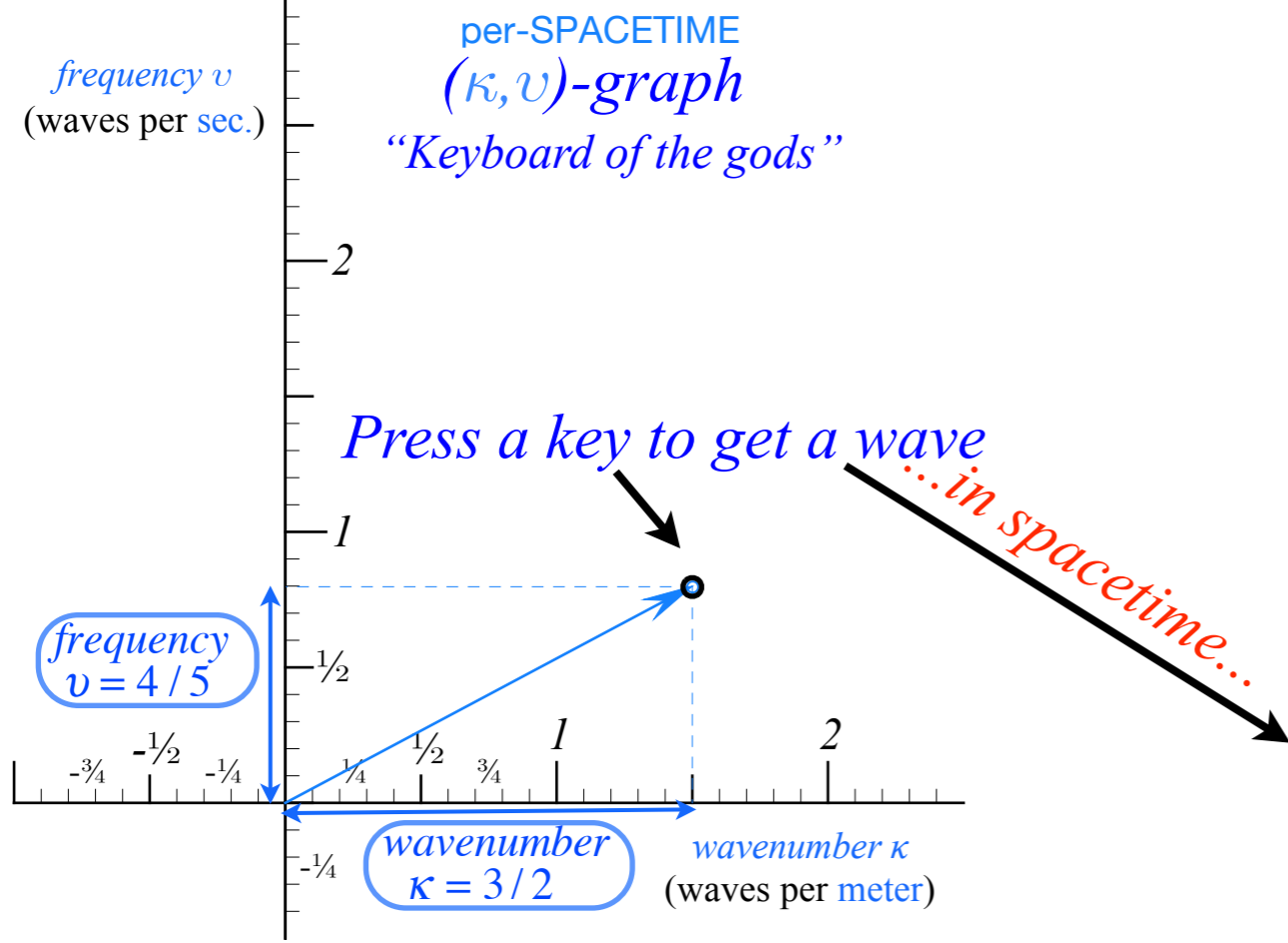
Bad Suzy!

Relativity and Quantum Theory
need to be unified in *one* book
half the size of those old tomes!

We call that a *Relawavity* book.
(It's a *lot* **lighter**!)



Analyzing wave velocity by **per-space-per-time** and **space-time** graphs



"Keyboard of the gods" is known as "Fourier-space"



Jean-Baptiste
Joseph Fourier
1768-1830

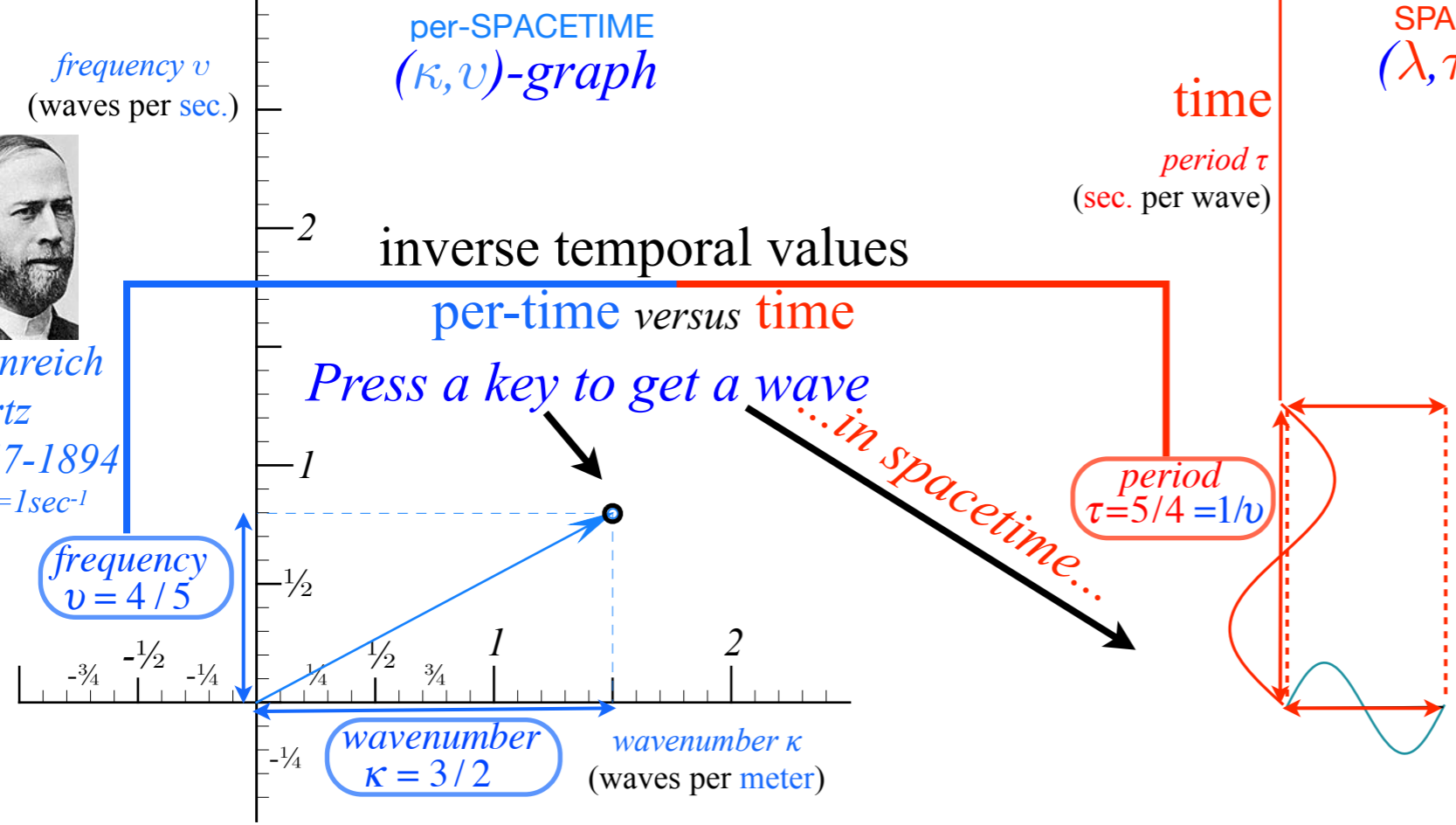
•How to understand waves
and
wave velocity V_{wave}

[RelaWavity Web Simulation](#)
[Keyboard of the Gods](#)
(per-Time vs per-Space)

Analyzing wave velocity by **per-space-per-time** and **space-time** graphs



Heinrich Hertz
1857-1894
1Hz=1sec⁻¹



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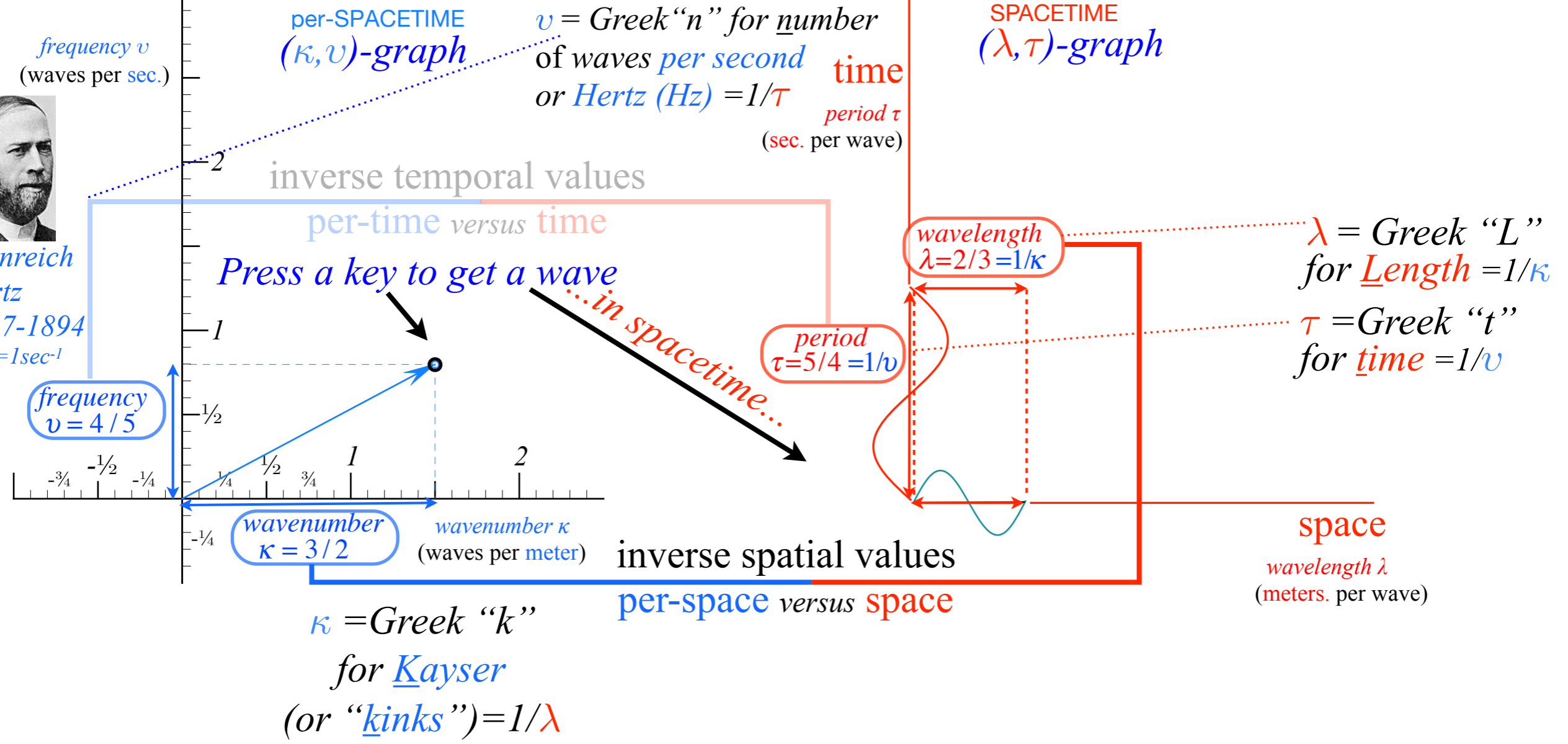
Jean-Baptiste Joseph Fourier
1768-1830

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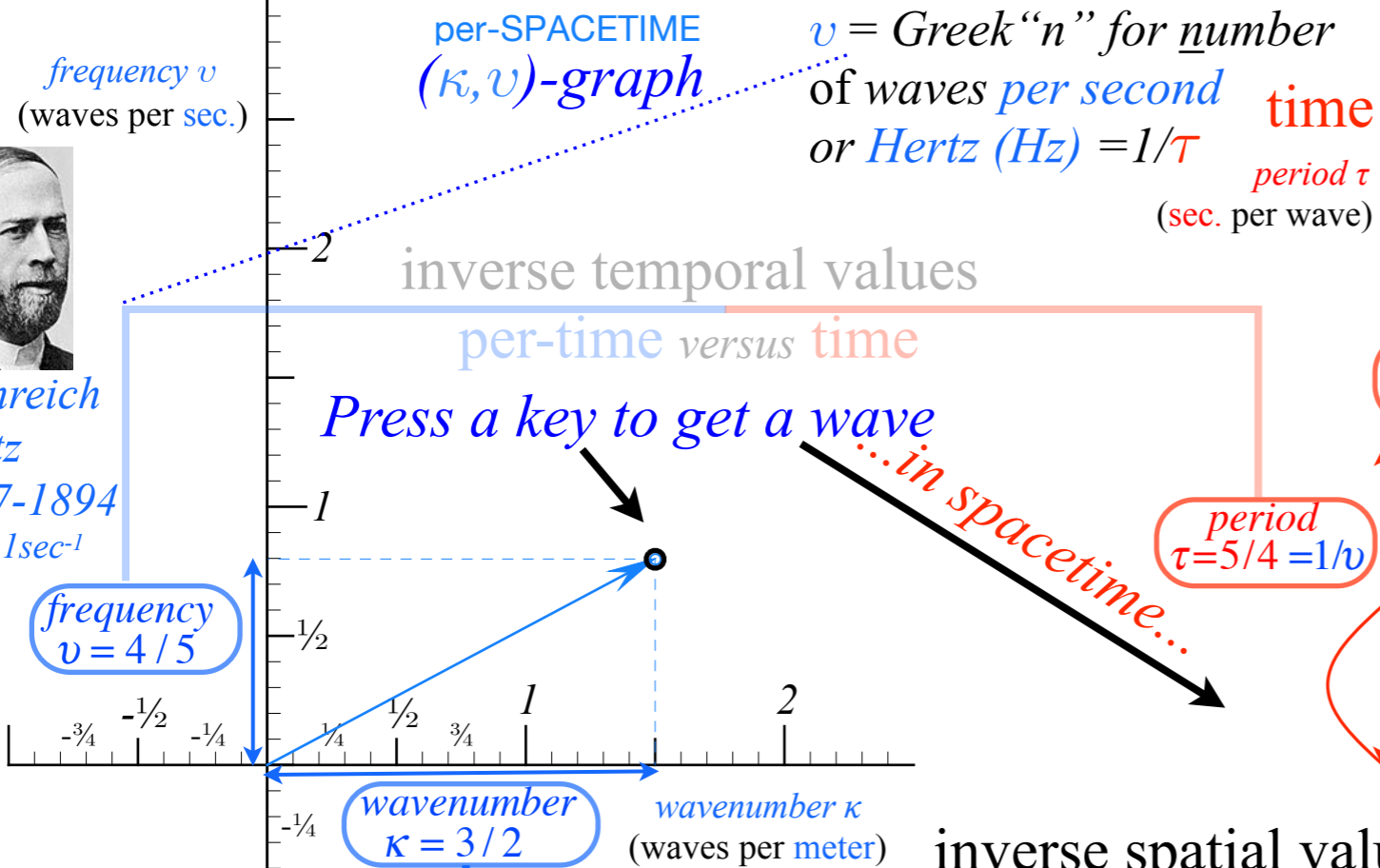
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1857-1894
1Hz=1sec⁻¹



per-SPACETIME
(κ, ν)-graph
 $\nu = \text{Greek "n" for number of waves per second or Hertz (Hz)} = 1/\tau$
time
period τ
(sec. per wave)

SPACETIME
(λ, τ)-graph

inverse temporal values

per-time versus time

Press a key to get a wave

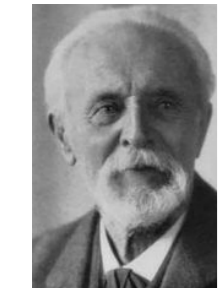
...in spacetime...

period $\tau = 5/4 = 1/\nu$

wavelength $\lambda = 2/3 = 1/\kappa$

$\lambda = \text{Greek "L" for Length} = 1/\kappa$

$\tau = \text{Greek "t" for time} = 1/\nu$



Heinrich Kayser
1853-1940
1Kayser=1cm⁻¹

inverse spatial values

$\kappa = \text{Greek "k" for Kayser (or "kinks")} = 1/\lambda$

per-space versus space

space
wavelength λ
(meters. per wave)

"Keyboard of the gods" is known as "Fourier-space"

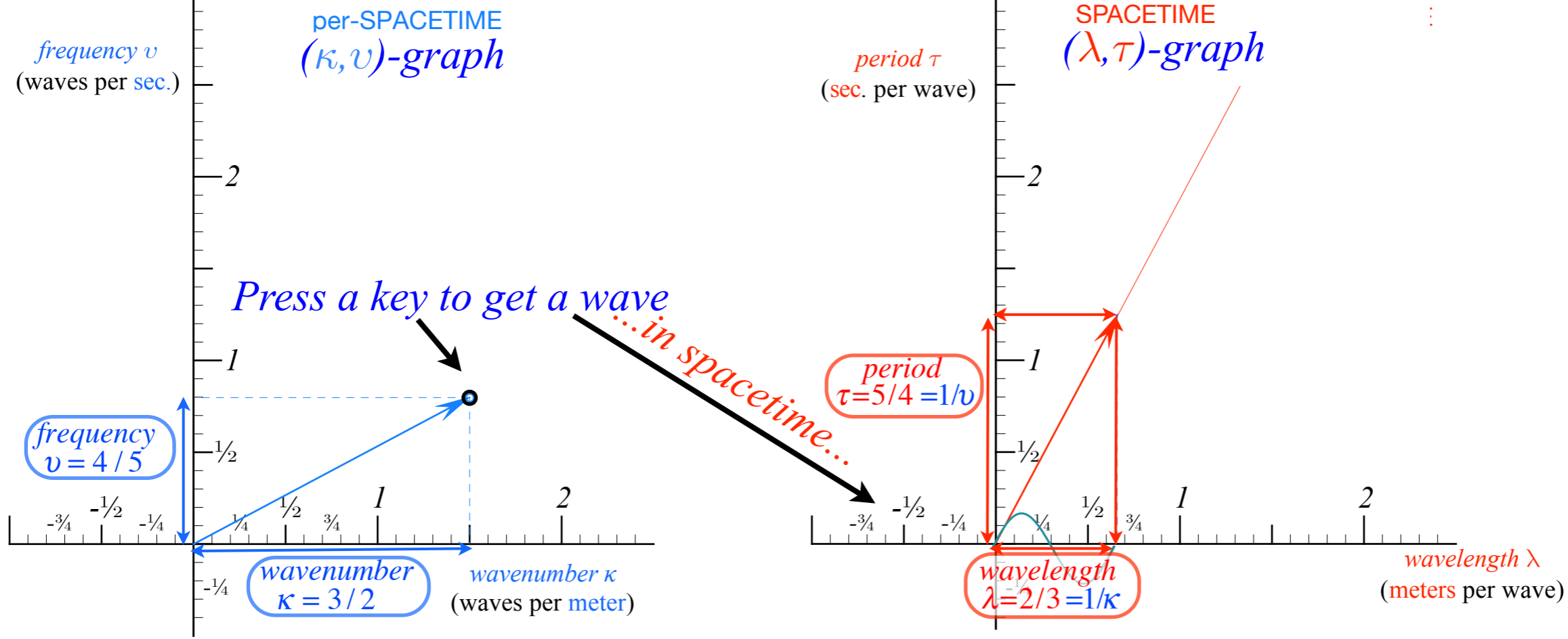


Jean-Baptiste Joseph Fourier
1768-1830

[RelaWavity Web Simulation](#)
[Keyboard of the Gods](#)
[\(Dual Plot\)](#)

•How to understand waves
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wave velocity V_{wave}

Analyzing wave velocity by **per-space-per-time** and **space-time** graphs



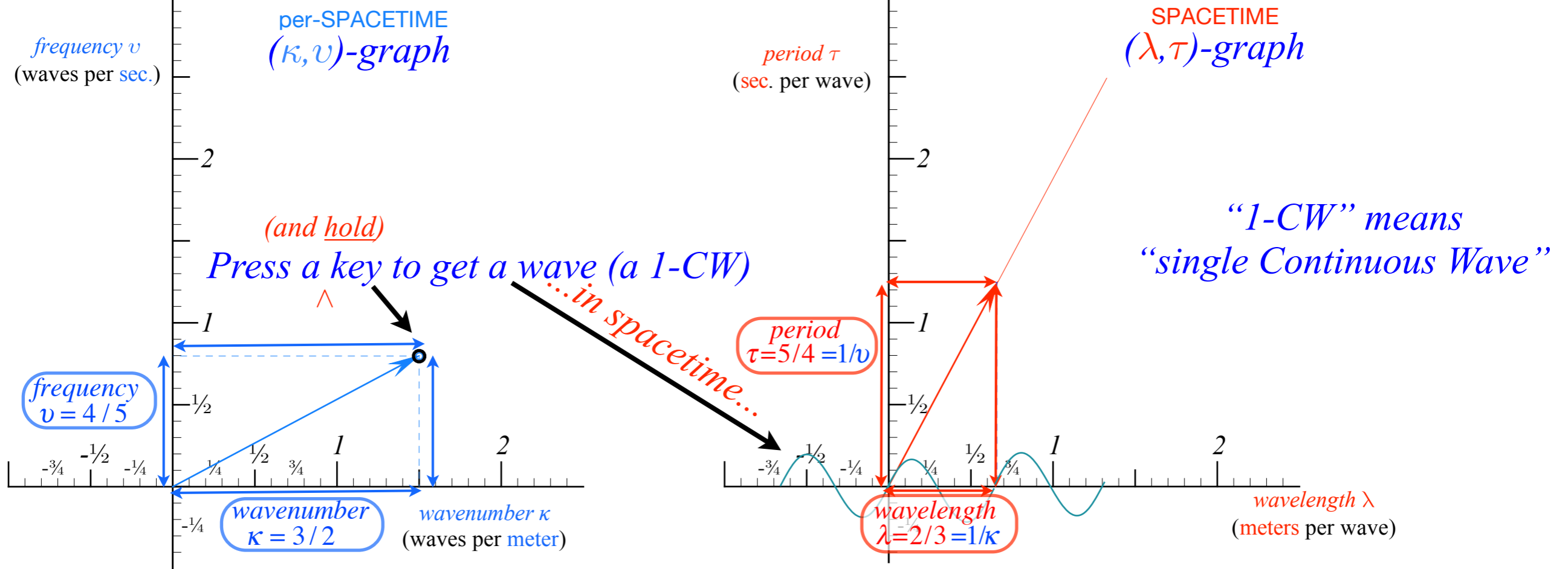
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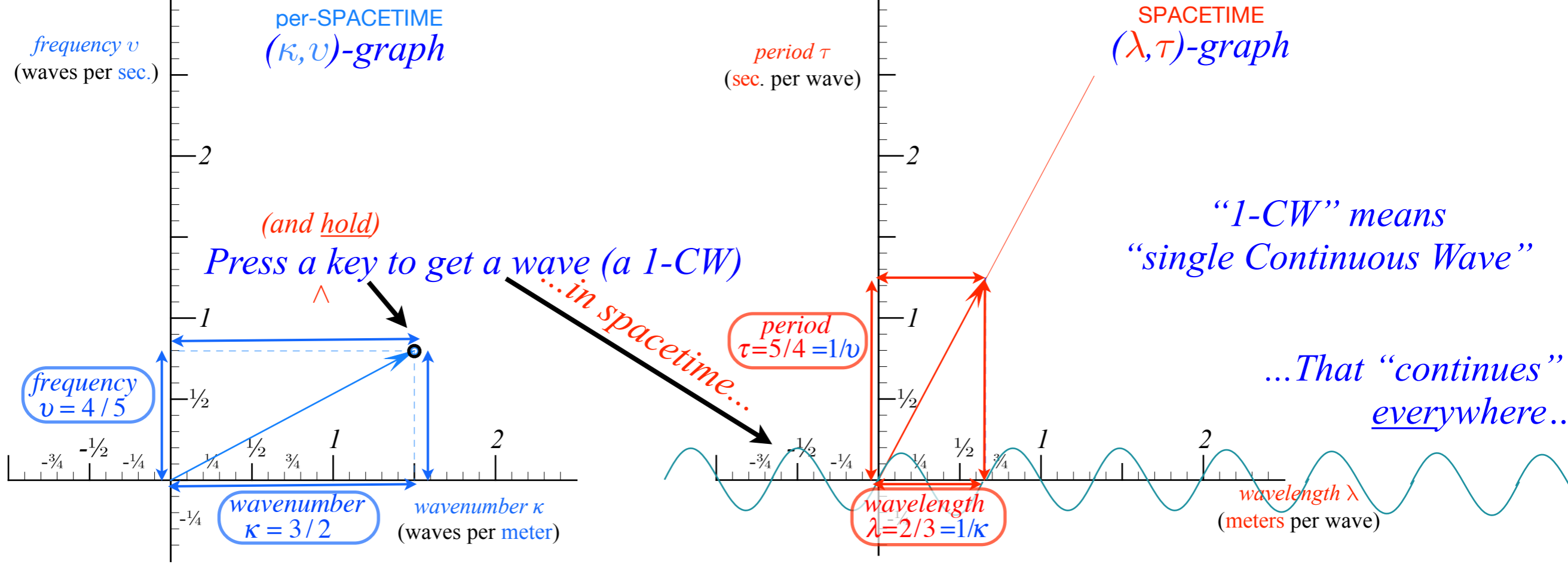


Jean-Baptiste
Joseph Fourier
1768-1830

[RelaWavity Web Simulation](#)
[Keyboard of the Gods](#)
(Dual Plot)

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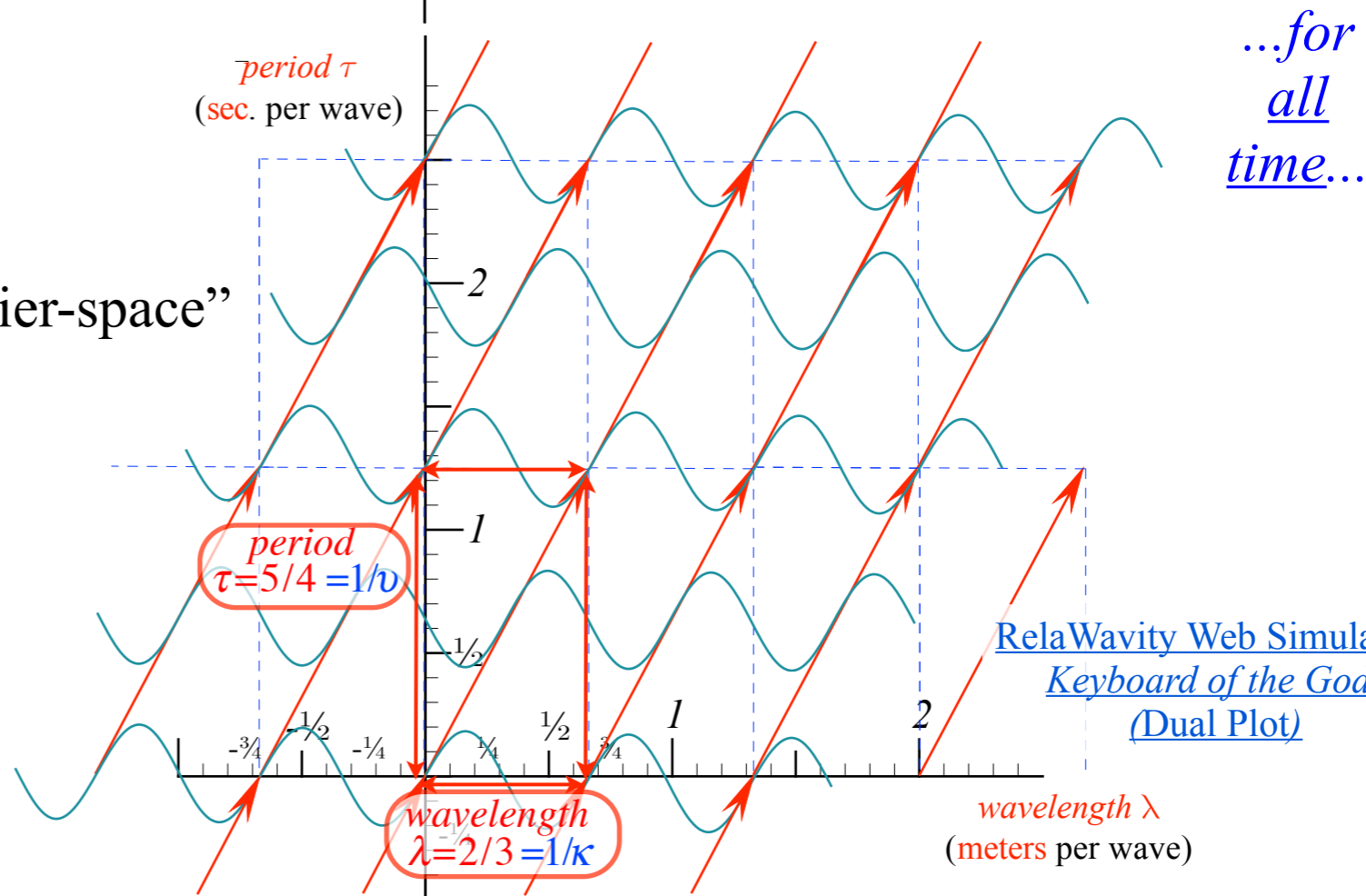
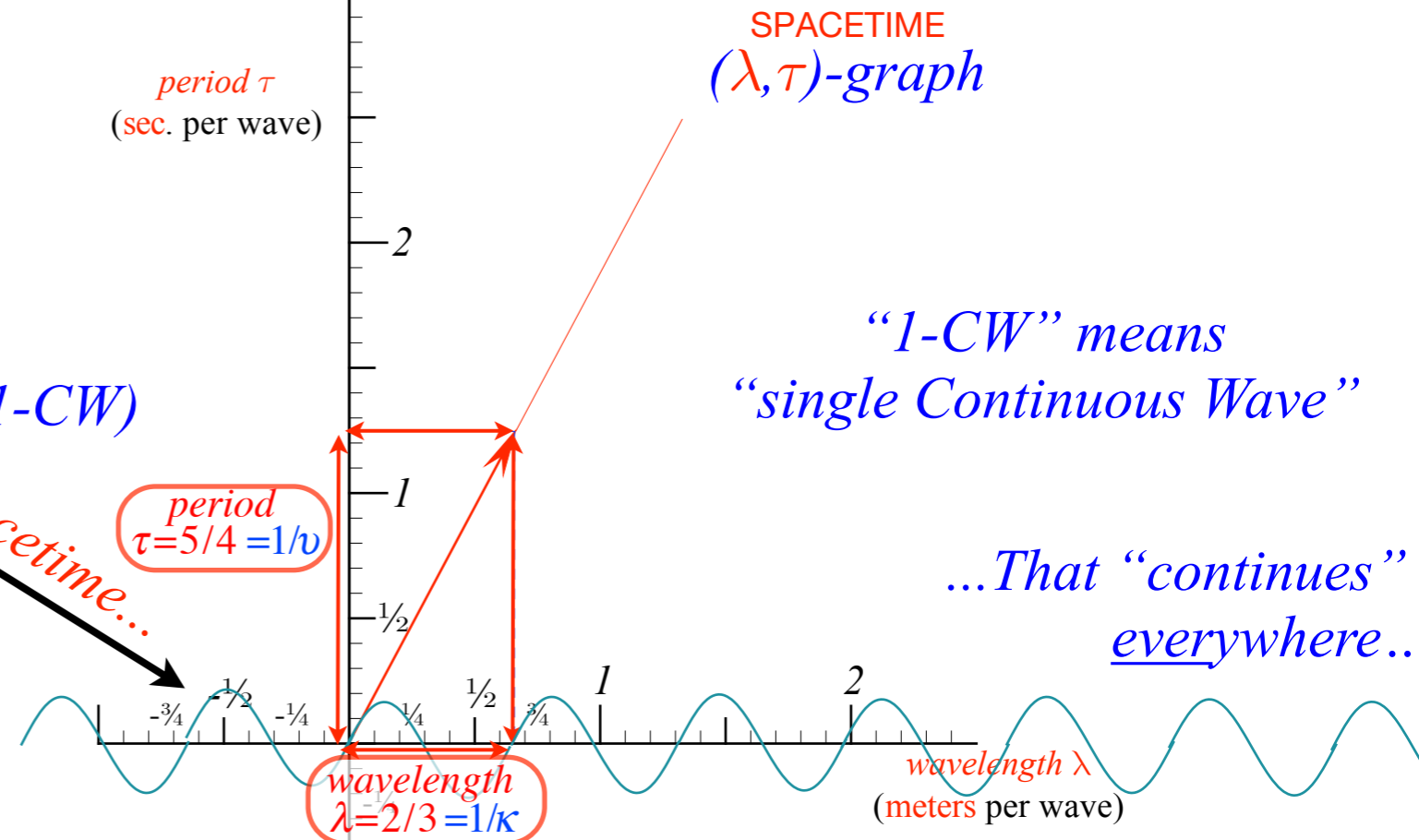
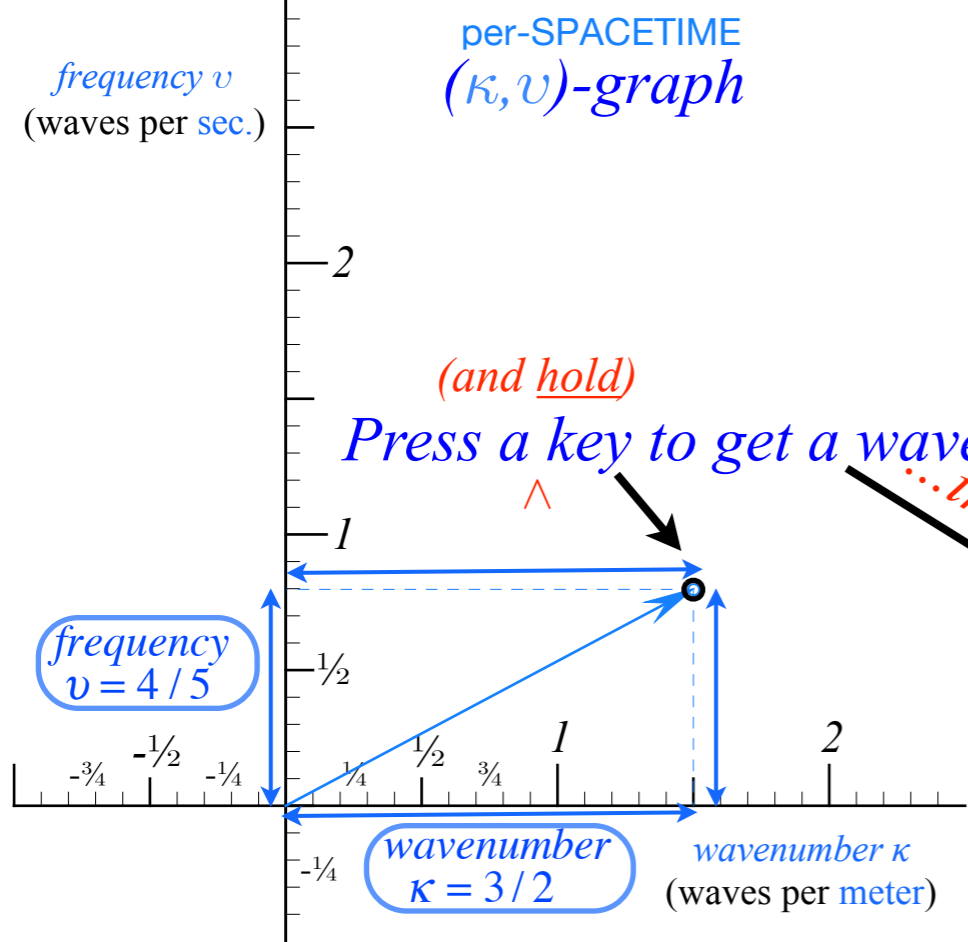


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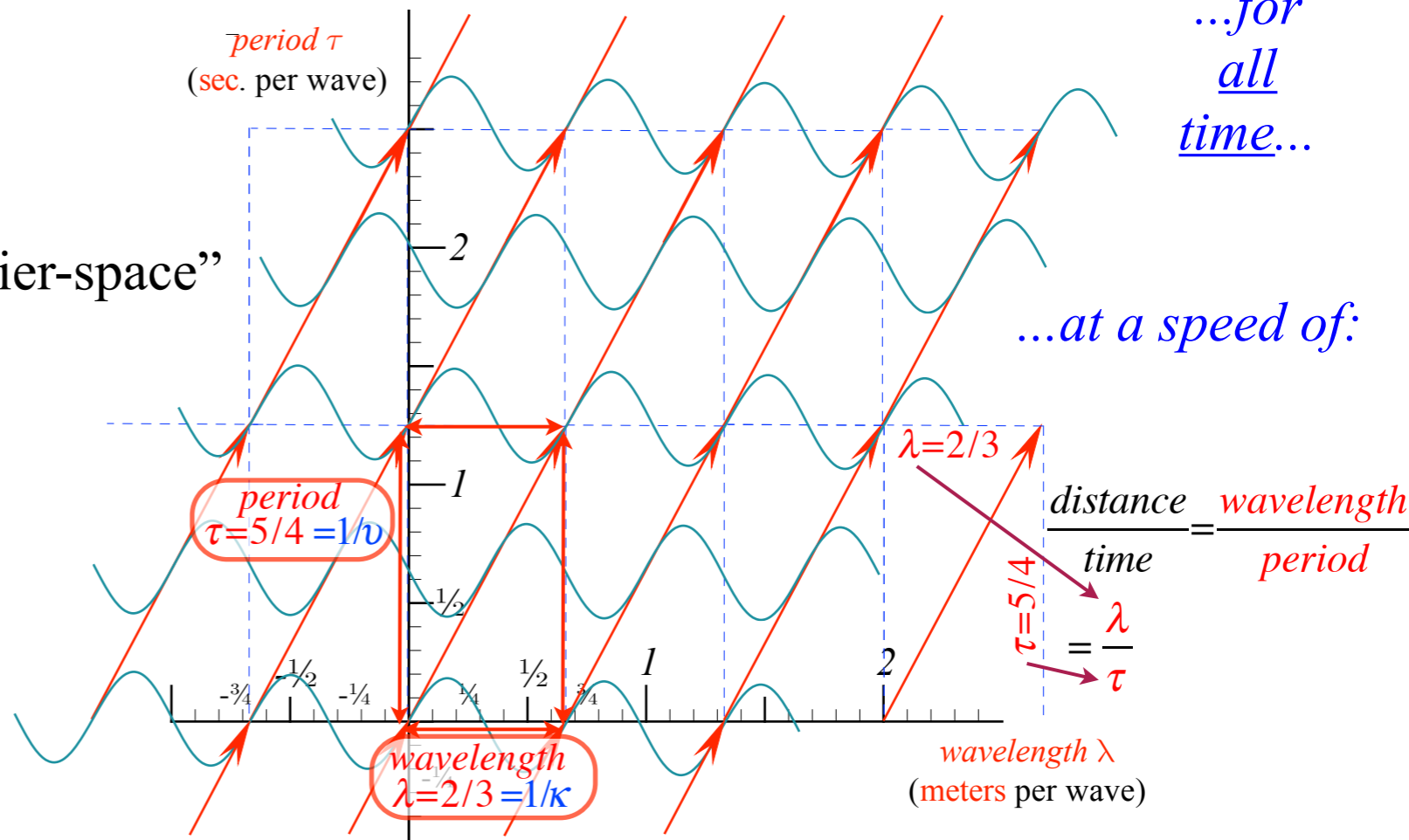
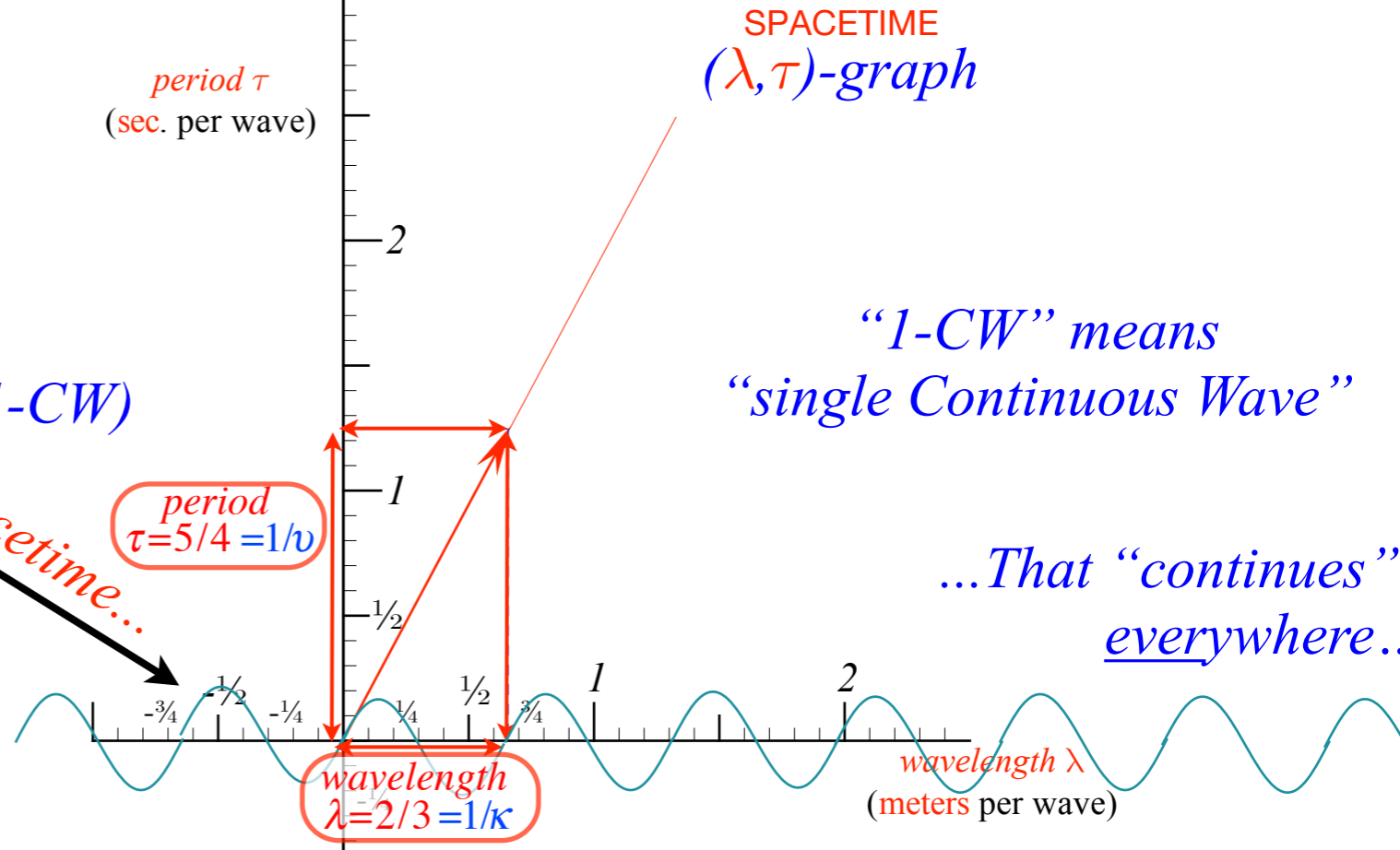
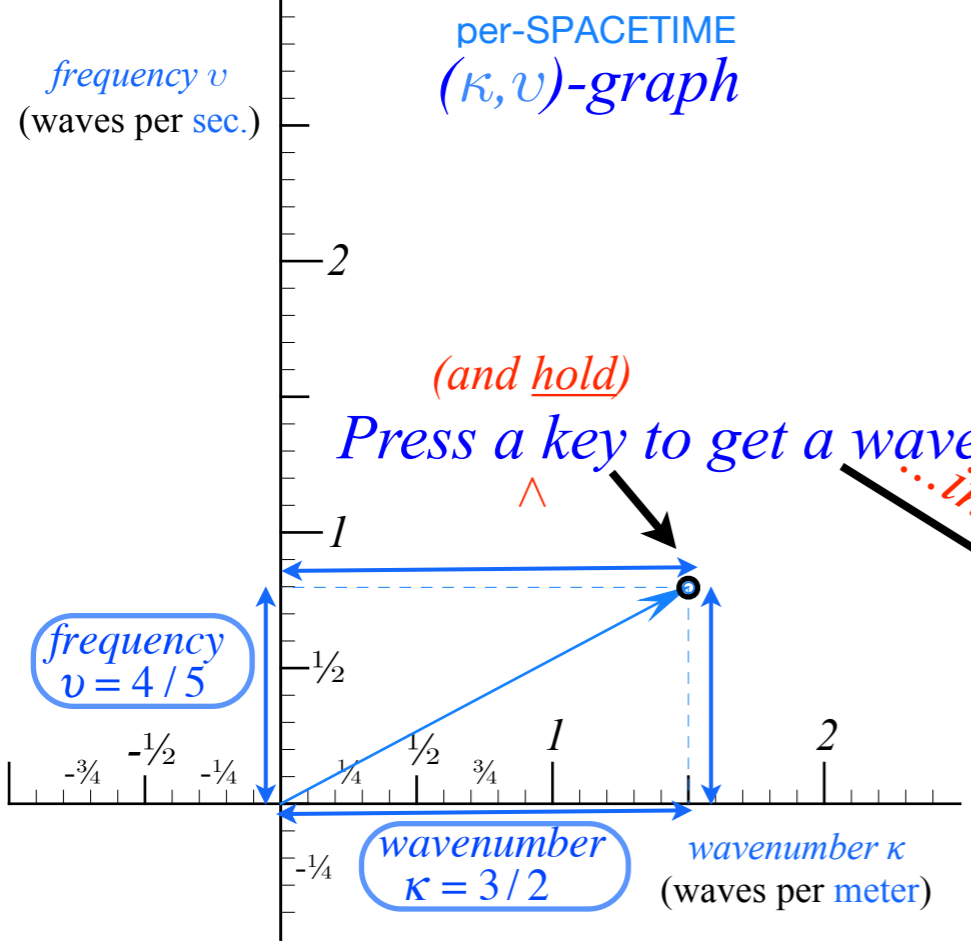


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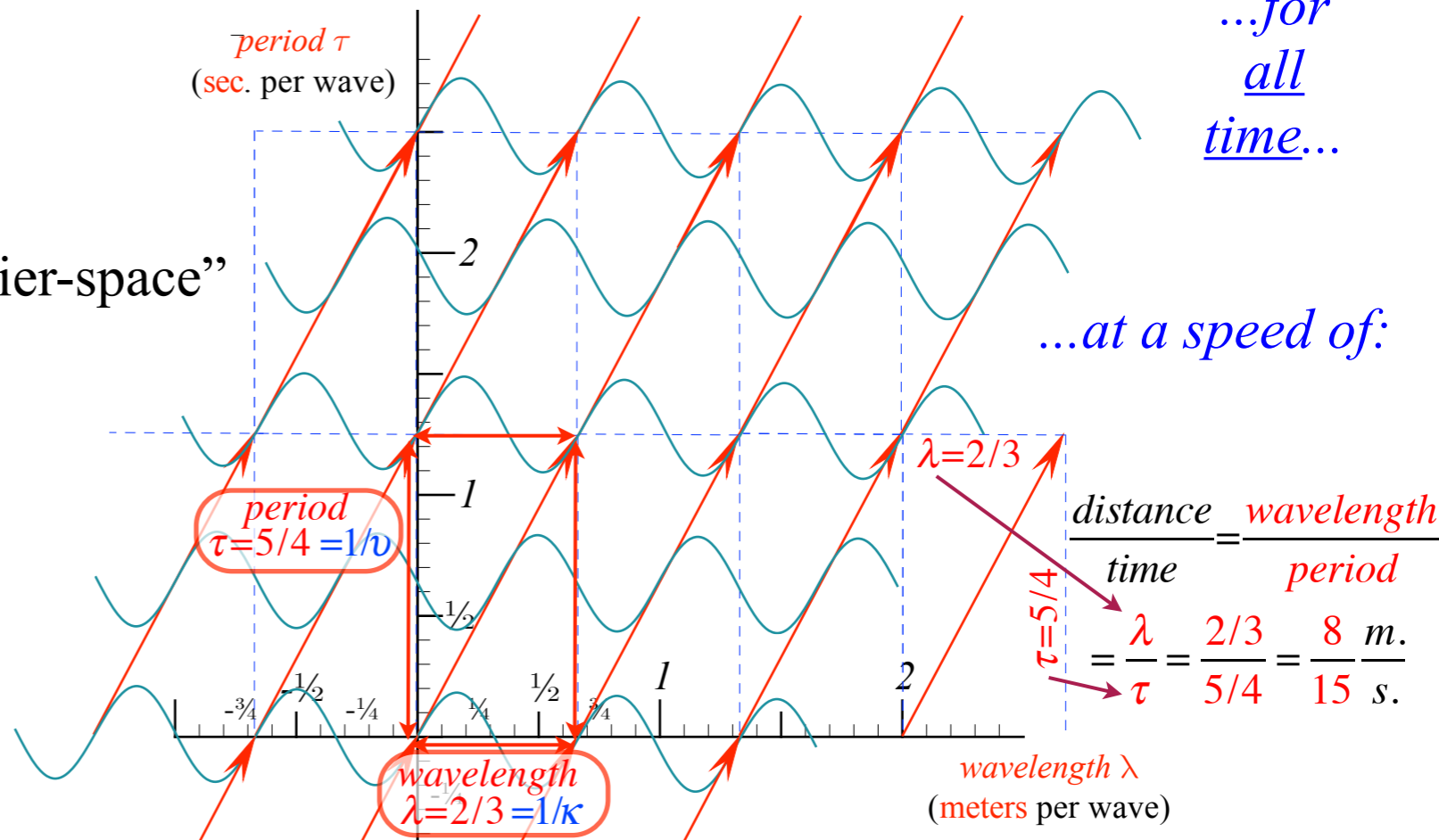
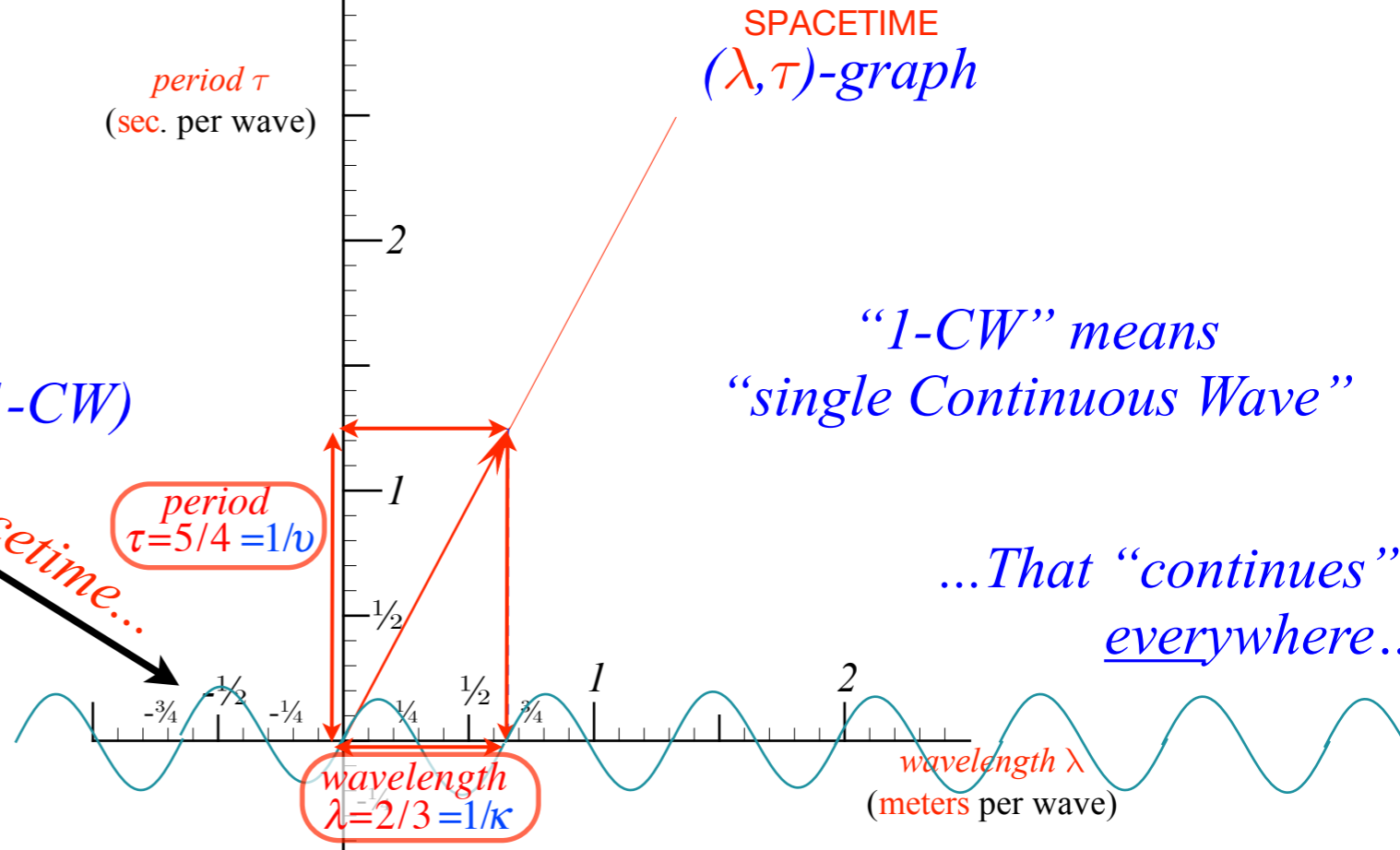
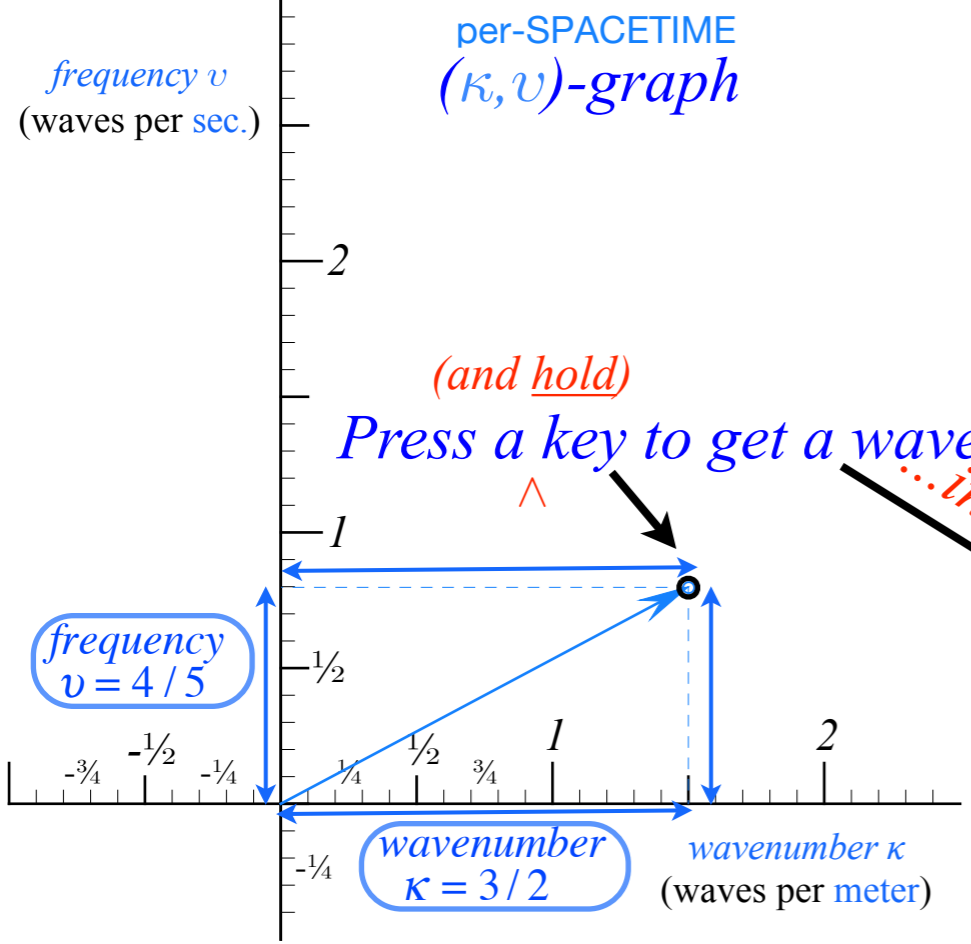
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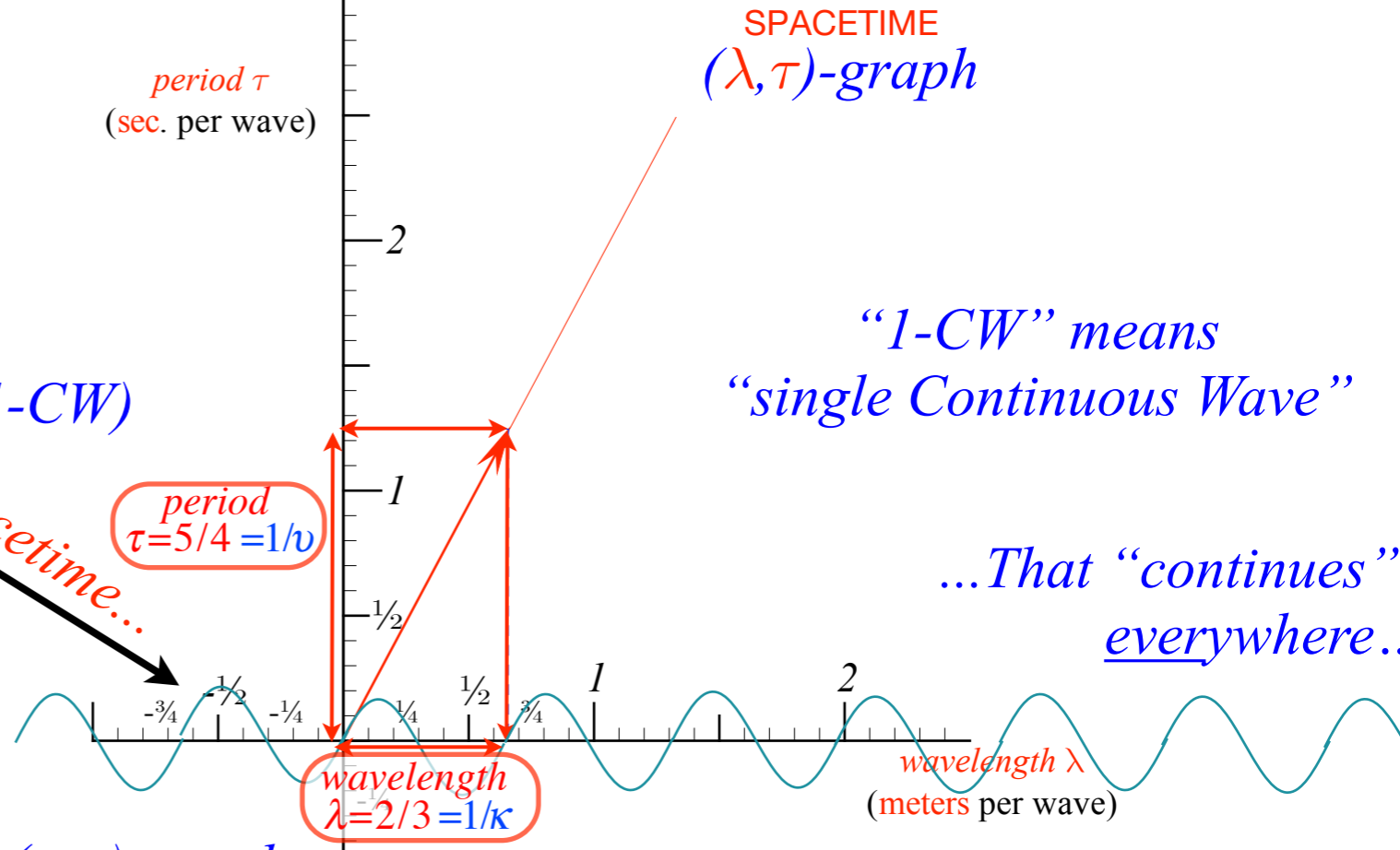
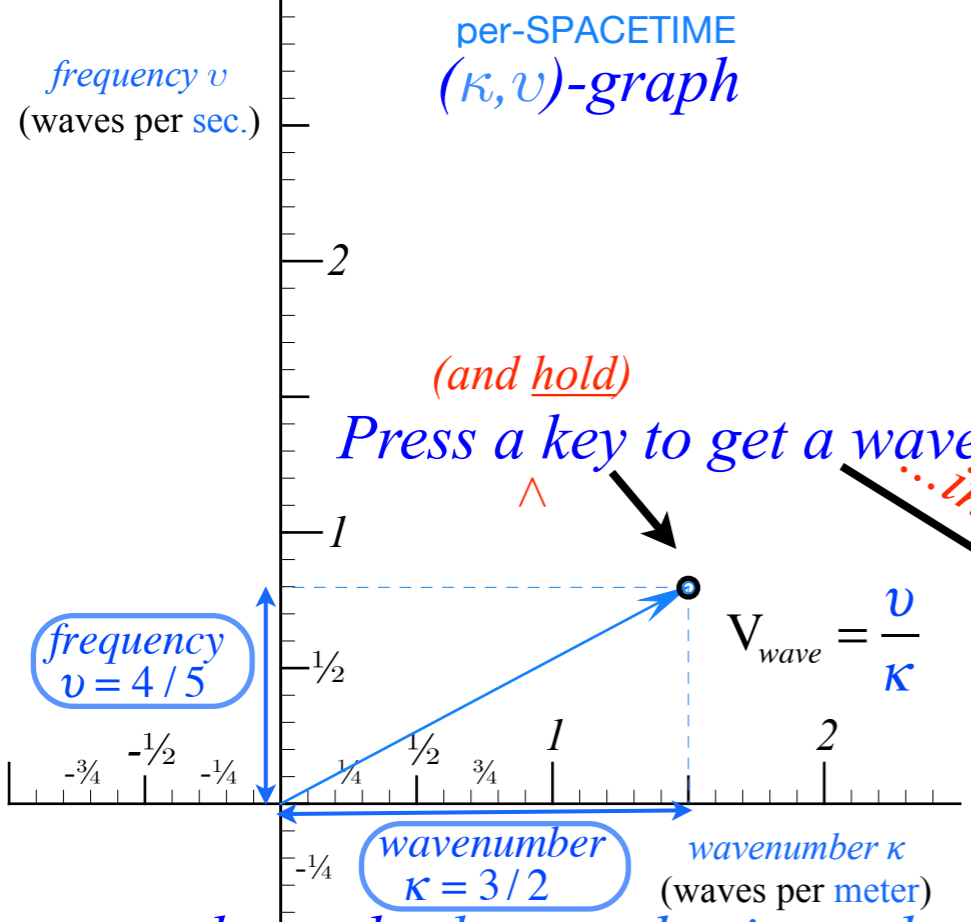


Jean-Baptiste Joseph Fourier
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•How to understand waves and wave velocity V_{wave}

wave-speed equals slope-to-vertical λ/τ in (λ, τ) -graph

Analyzing wave velocity by per-space-per-time and space-time graphs



(and hold)
Press a key to get a wave (a 1-CW)

...in spacetime...

"1-CW" means "single Continuous Wave"

...That "continues" everywhere..

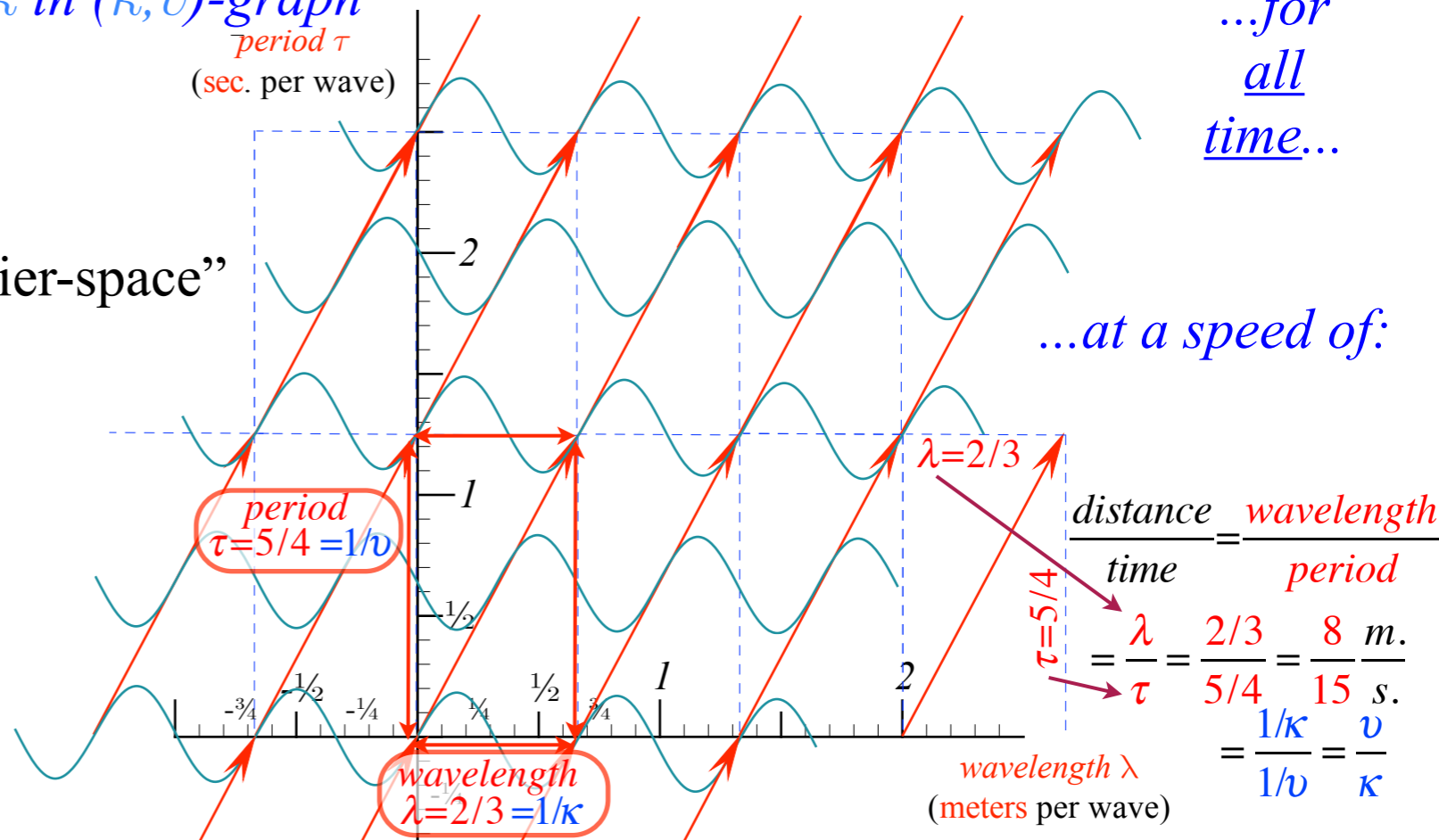
wave-speed equals slope-to-horizontal ν/κ in (κ, ν) -graph

...for all time...

"Keyboard of the gods" is known as "Fourier-space"



Jean-Baptiste Joseph Fourier
1768-1830

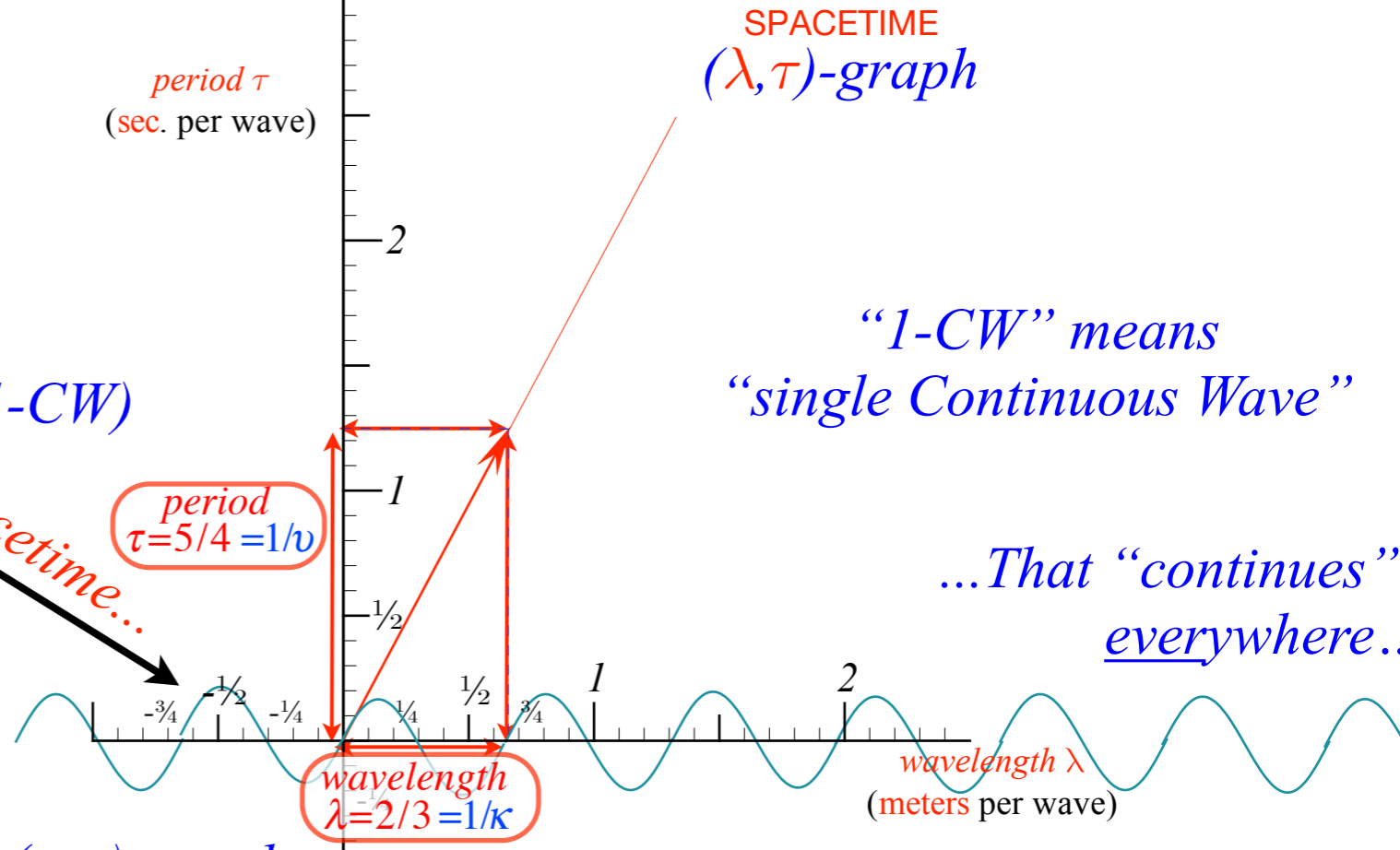
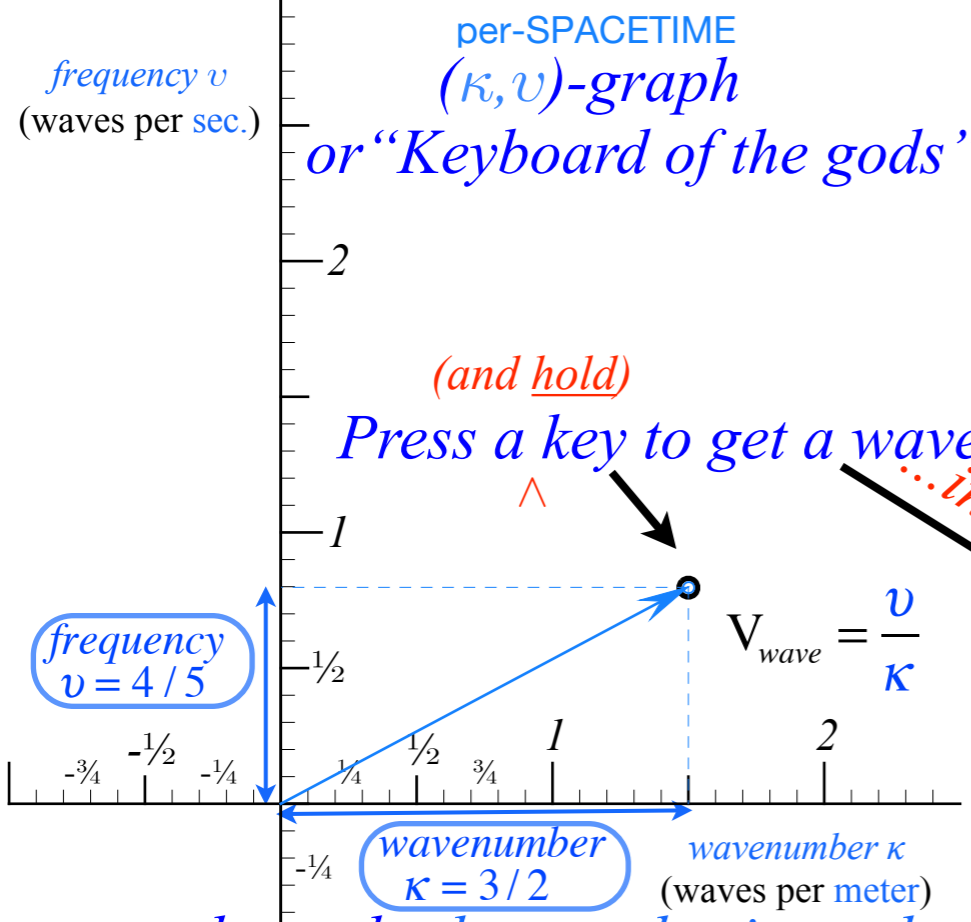


...at a speed of:

•How to understand waves and wave velocity V_{wave}

wave-speed equals slope-to-vertical λ/τ in (λ, τ) -graph

Analyzing wave velocity by per-space-per-time and space-time graphs



wave-speed equals slope-to-horizontal ν/κ in (κ, ν)-graph

...for all time..

wave-velocity formulas

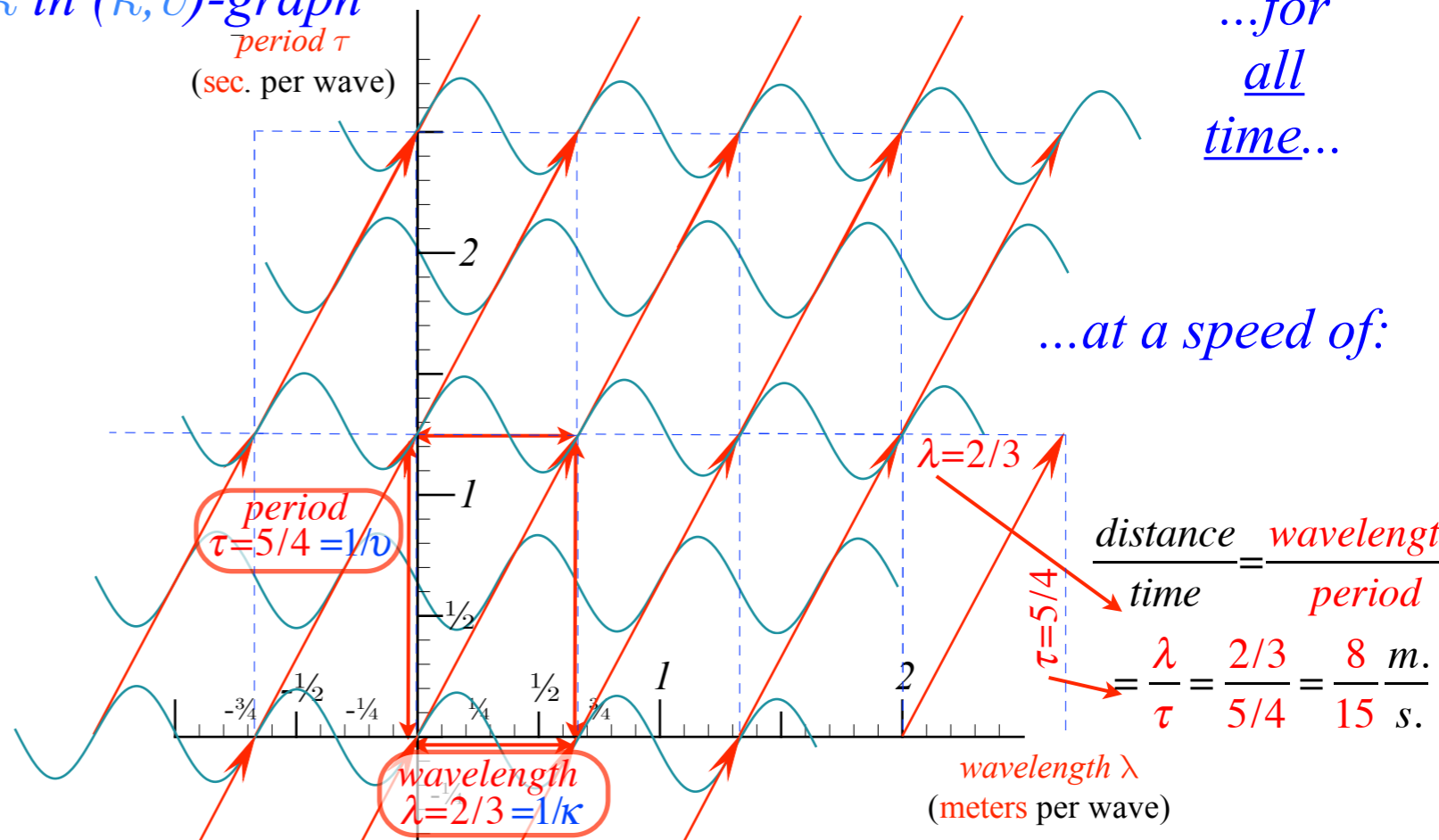
$$\frac{\text{distance}}{\text{time}} = \frac{\text{wavelength}}{\text{period}} = \frac{\text{frequency}}{\text{wavenumber}}$$

$$V_{\text{wave}} = \frac{\lambda}{\tau} = \frac{1/\kappa}{1/\nu} = \frac{\nu}{\kappa} = \frac{1/\tau}{1/\lambda}$$

$$= \frac{2/3}{5/4} = \frac{4/5}{3/2} = \frac{8 \text{ m.}}{15 \text{ s.}}$$

wave arithmetic is simpler to explain using fractions

•How to understand waves and "1st quantization"



wave-speed equals slope-to-vertical λ/τ in (λ, τ)-graph

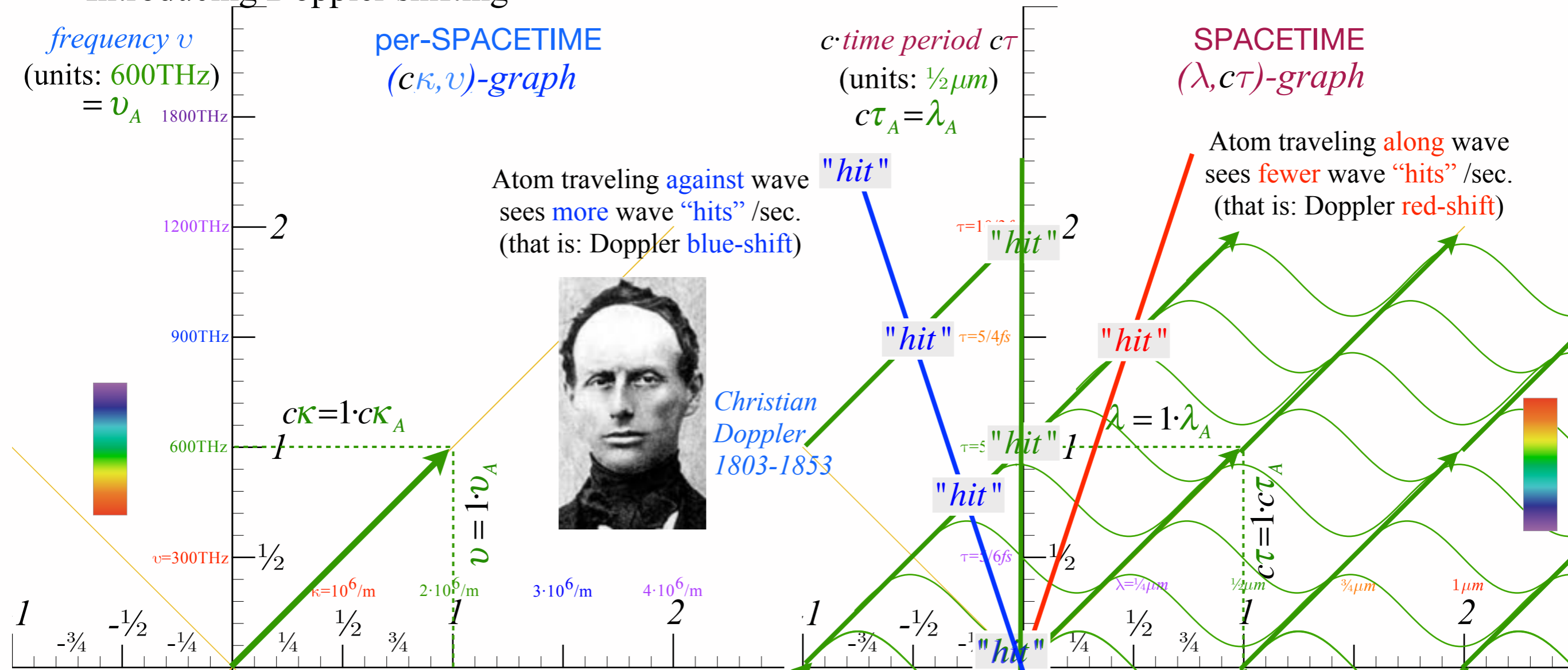
Introducing Doppler shifting

frequency ν
(units: 600THz)
 $= \nu_A$ 1800THz

per-SPACETIME
 $(c\kappa, \nu)$ -graph

$c \cdot$ time period $c\tau$
(units: $\frac{1}{2}\mu m$)
 $c\tau_A = \lambda_A$

SPACETIME
 $(\lambda, c\tau)$ -graph



Atom traveling **against** wave sees **more** wave "hits" /sec. (that is: Doppler **blue-shift**)

Atom traveling **along** wave sees **fewer** wave "hits" /sec. (that is: Doppler **red-shift**)



Christian Doppler
1803-1853

$$c = \frac{\lambda}{\tau} = \frac{\nu}{\kappa} = \frac{\omega}{k}$$

rescaled by c to:

$$1 = \frac{\lambda}{c\tau} = \frac{\nu}{c\kappa} = \frac{\omega}{ck}$$

Move fast enough this way then the "green" wave gets **redder** and **redder** until it dies

Move fast enough this way then the "green" wave gets **bluer** and **bluer** until YOU die

Frequency AND Amplitude decrease exponentially

Frequency AND Amplitude increase exponentially

Introducing Doppler shifting and why c is so constant (and so slow)

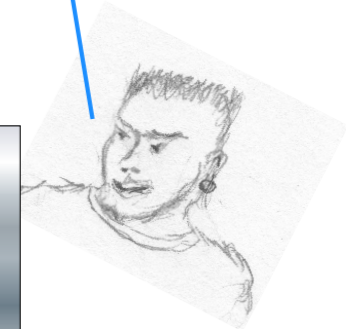
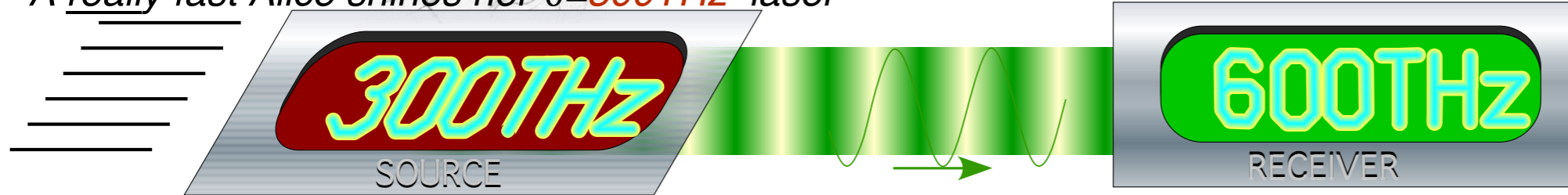
(a)



Bob: "Alice! My frequency meter reads $\nu=600\text{THz}$ for your laser beam."

Alice: "Well, what is its wavelength λ , Bob!"

A really fast Alice shines her $\nu=300\text{THz}$ laser



(b)

frequency $\nu=\omega/2\pi$
(Inverse period $\nu=1/\tau$)

$(\omega = ck)$

or

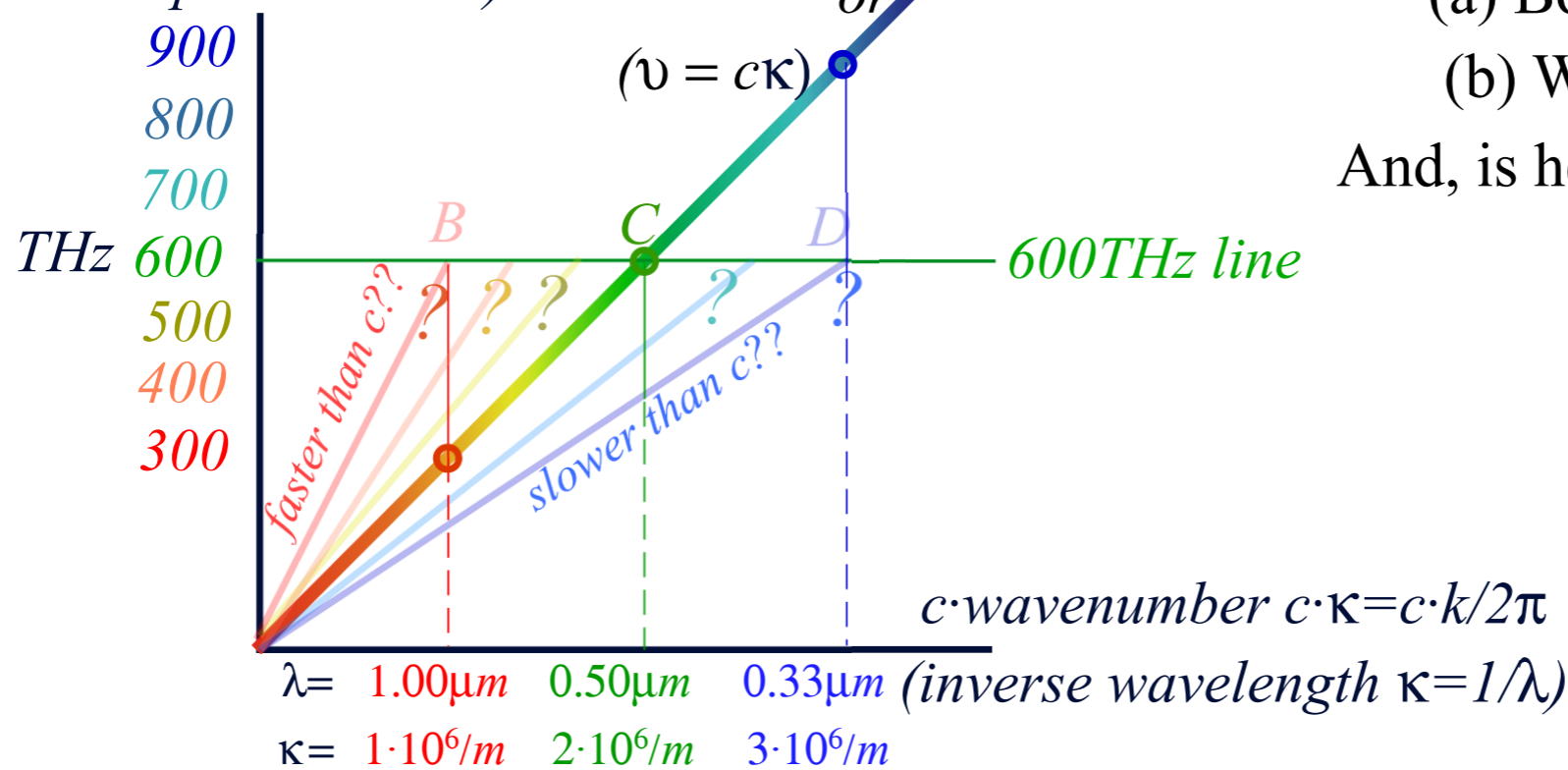
$(\nu = c\kappa)$

Alice's 300THz laser approaches Bob.

(a) Bob sees $\nu=600\text{THz}$.

(b) What $\lambda=1/\kappa$ does Bob measure?

And, is he seeing a 'phony' green?



Introducing Doppler shifting and why c is constant

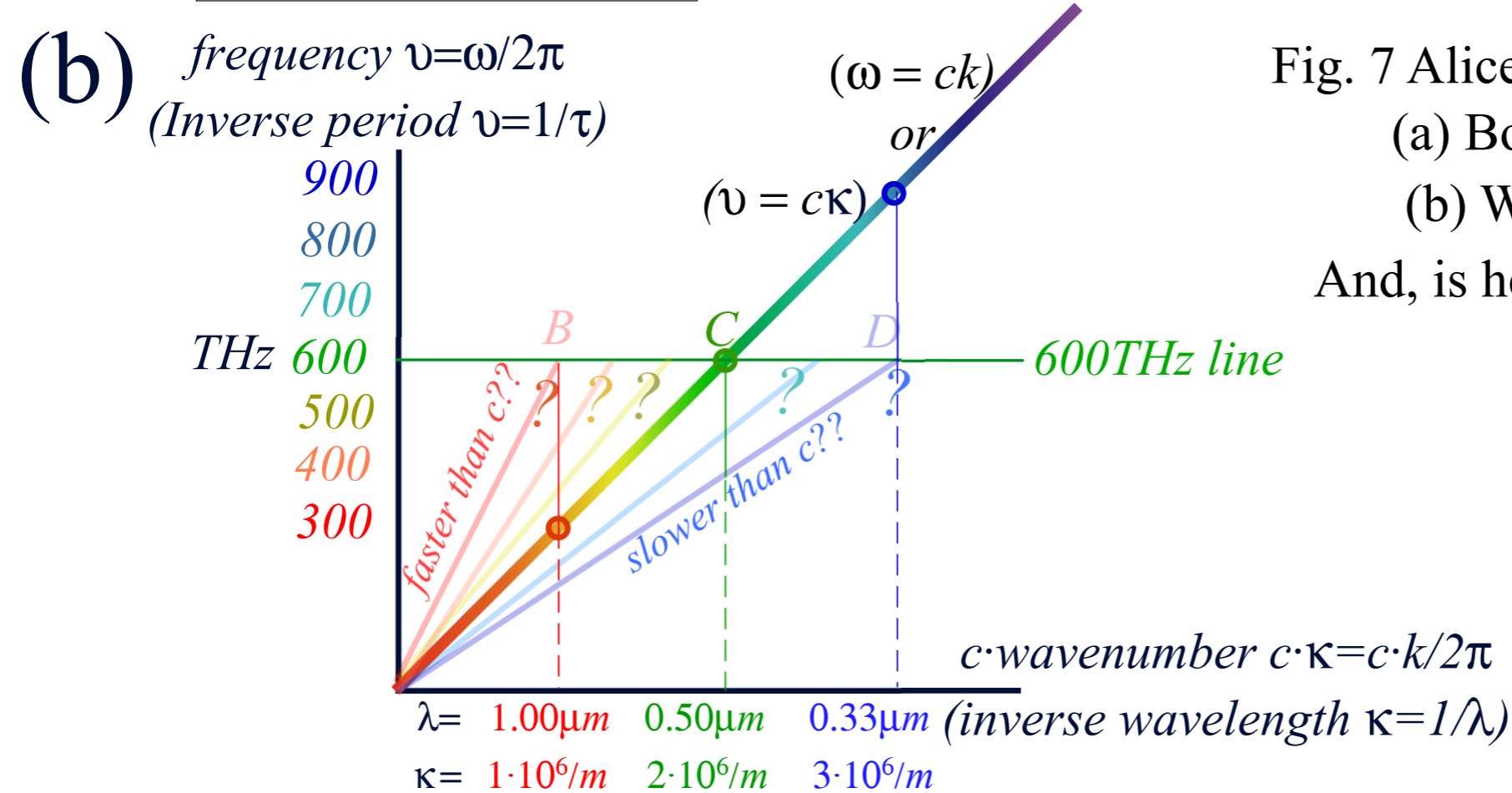
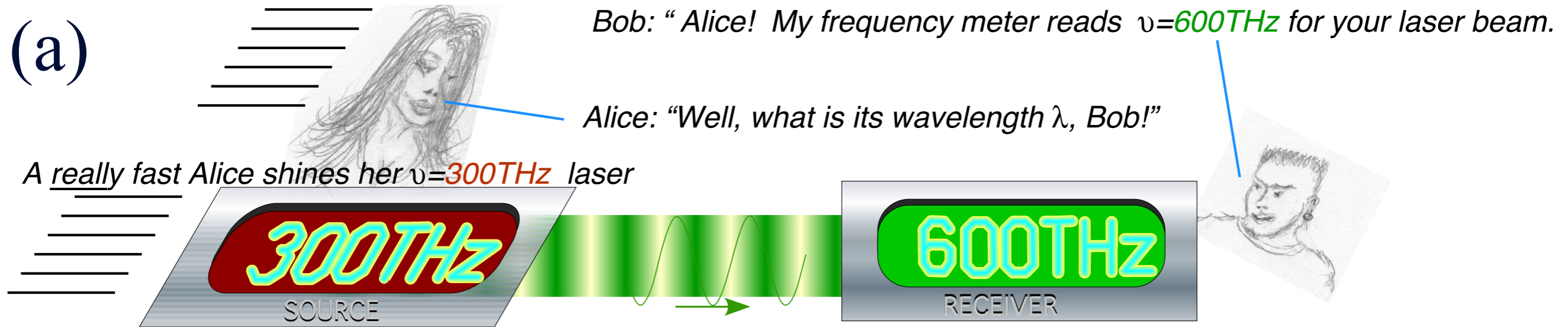


Fig. 7 Alice's 300THz laser approaches Bob.

(a) Bob sees $\nu=600\text{THz}$.

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Years of spectroscopy rule out 'phony' 600THz blue-green that do not have wavelength $\lambda=0.5\text{micron}$.

The only choice is C.

Introducing Doppler shifting and why c is constant

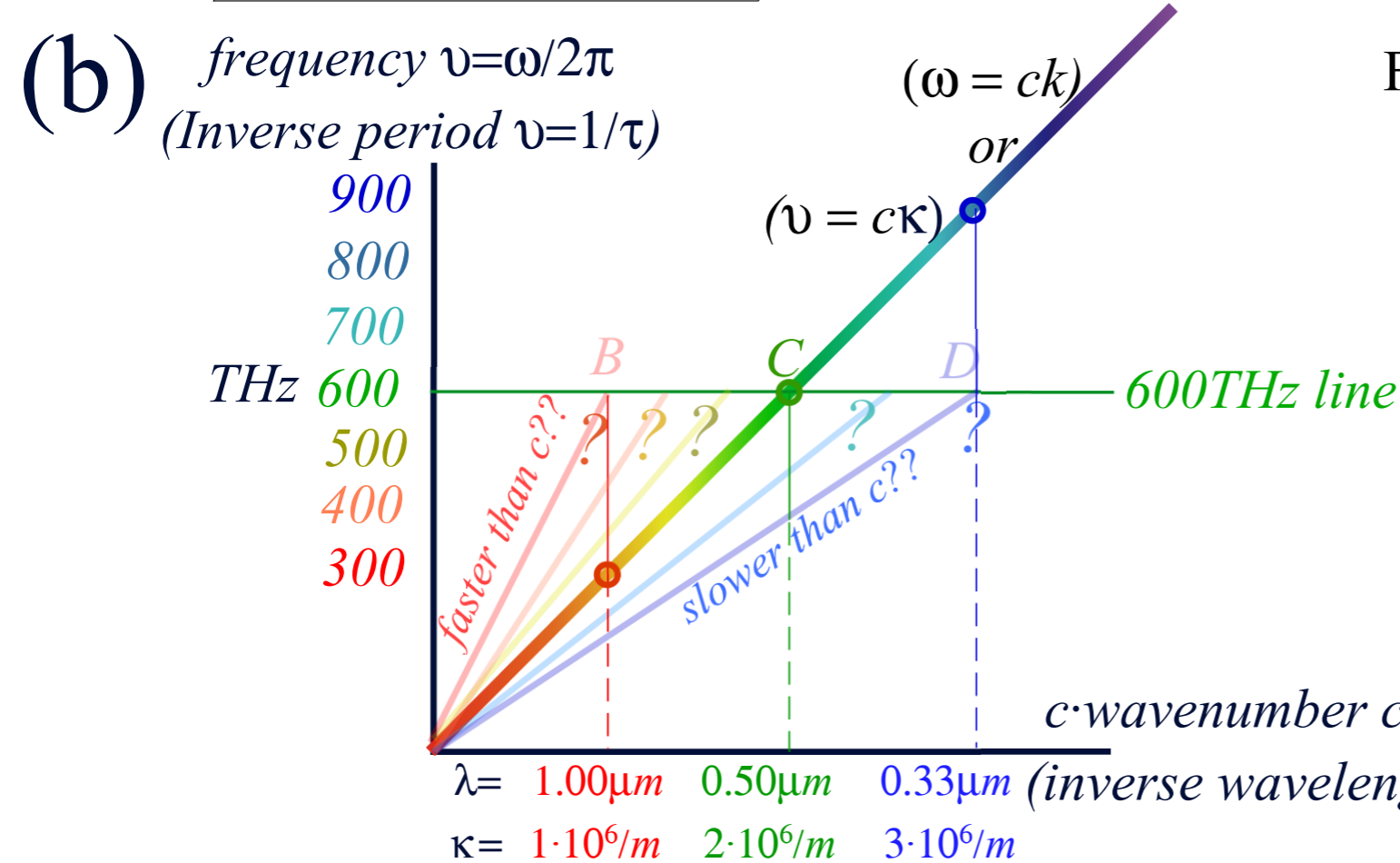
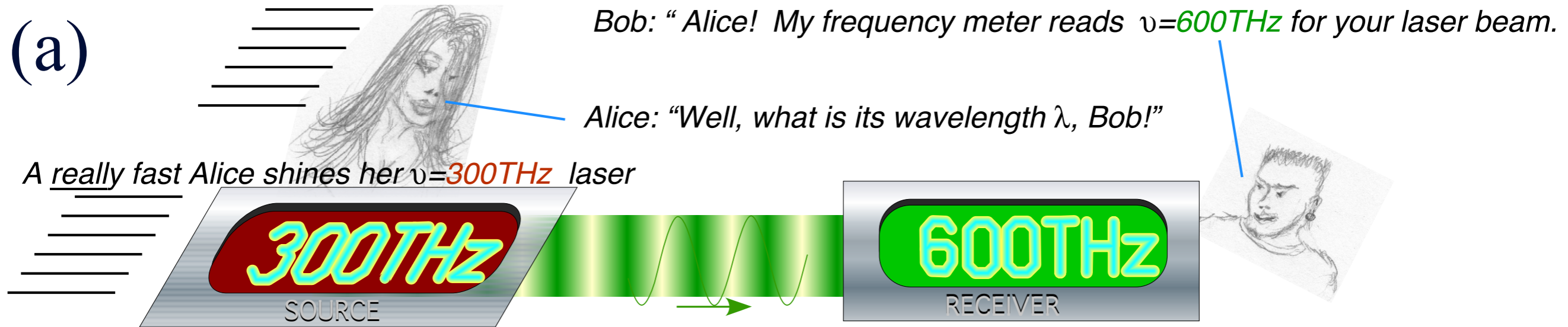


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The only choice is C. Also the only possible 600THz light speed is $c = \frac{\nu}{\kappa} = \frac{600 \cdot 10^{12}}{2 \cdot 10^6} = 3 \cdot 10^8 \text{ m} \cdot \text{s}^{-1}$

Introducing Doppler shifting and why c is constant

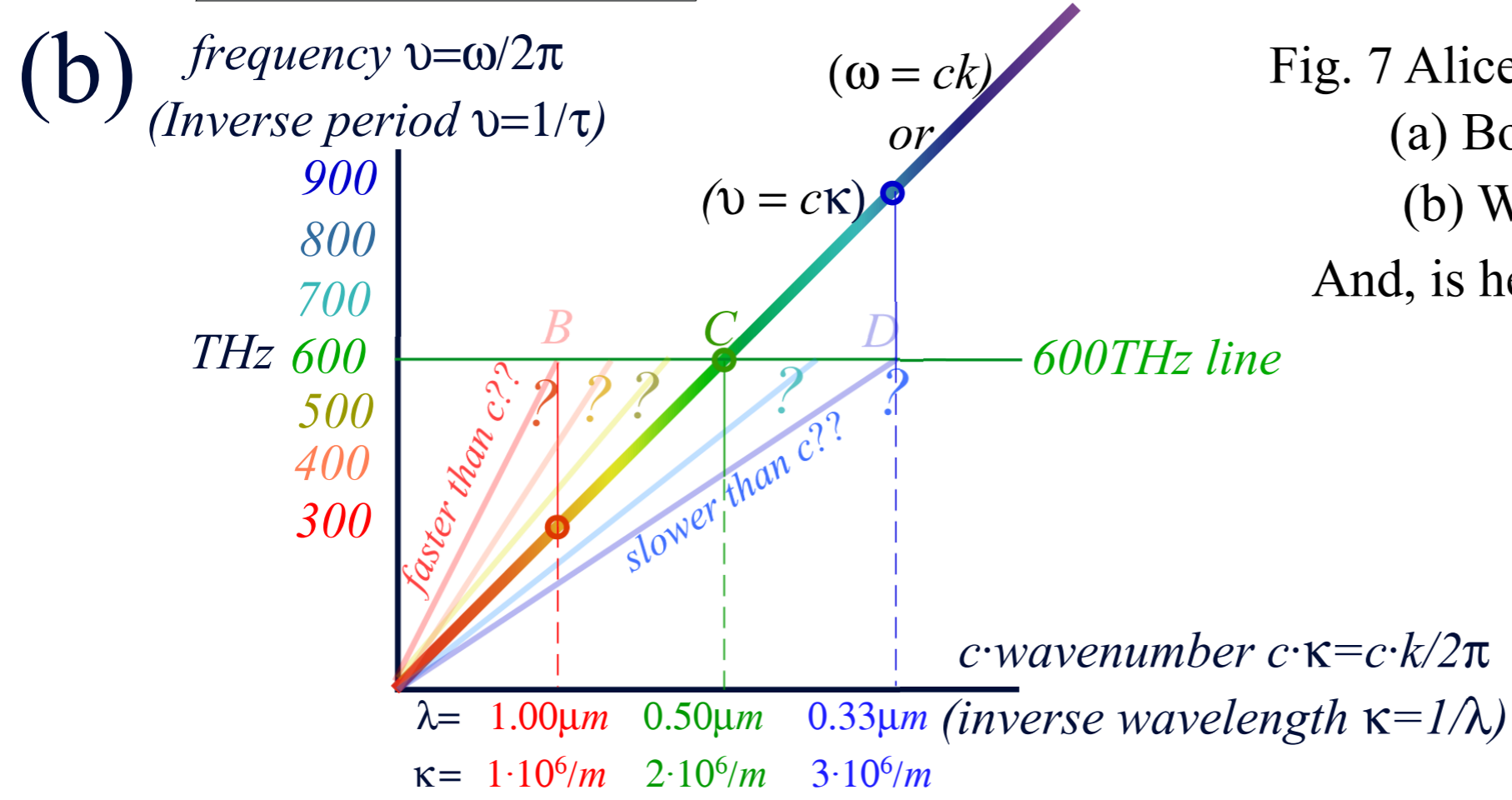
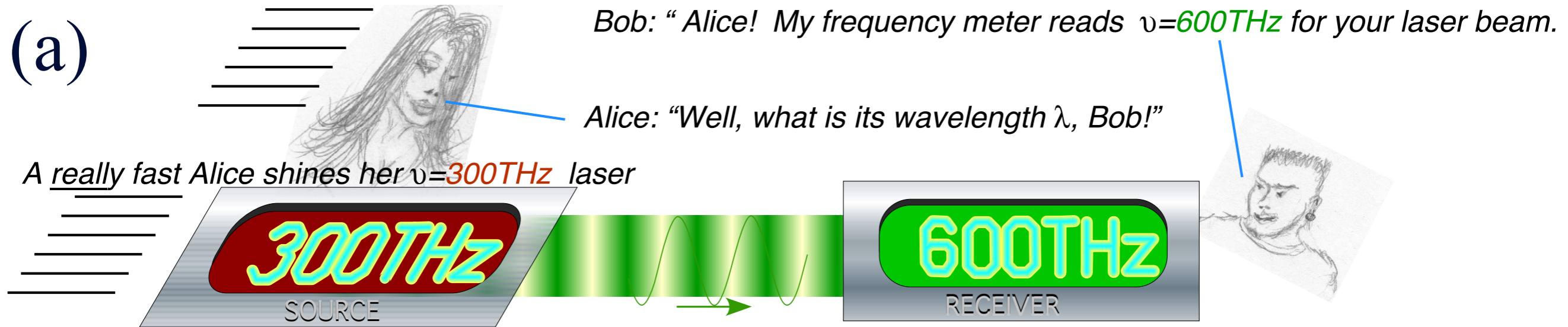


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Actually: $2.99792458 \cdot 10^8 \text{ m} \cdot \text{s}^{-1}$

Introducing Doppler shifting and why c is constant

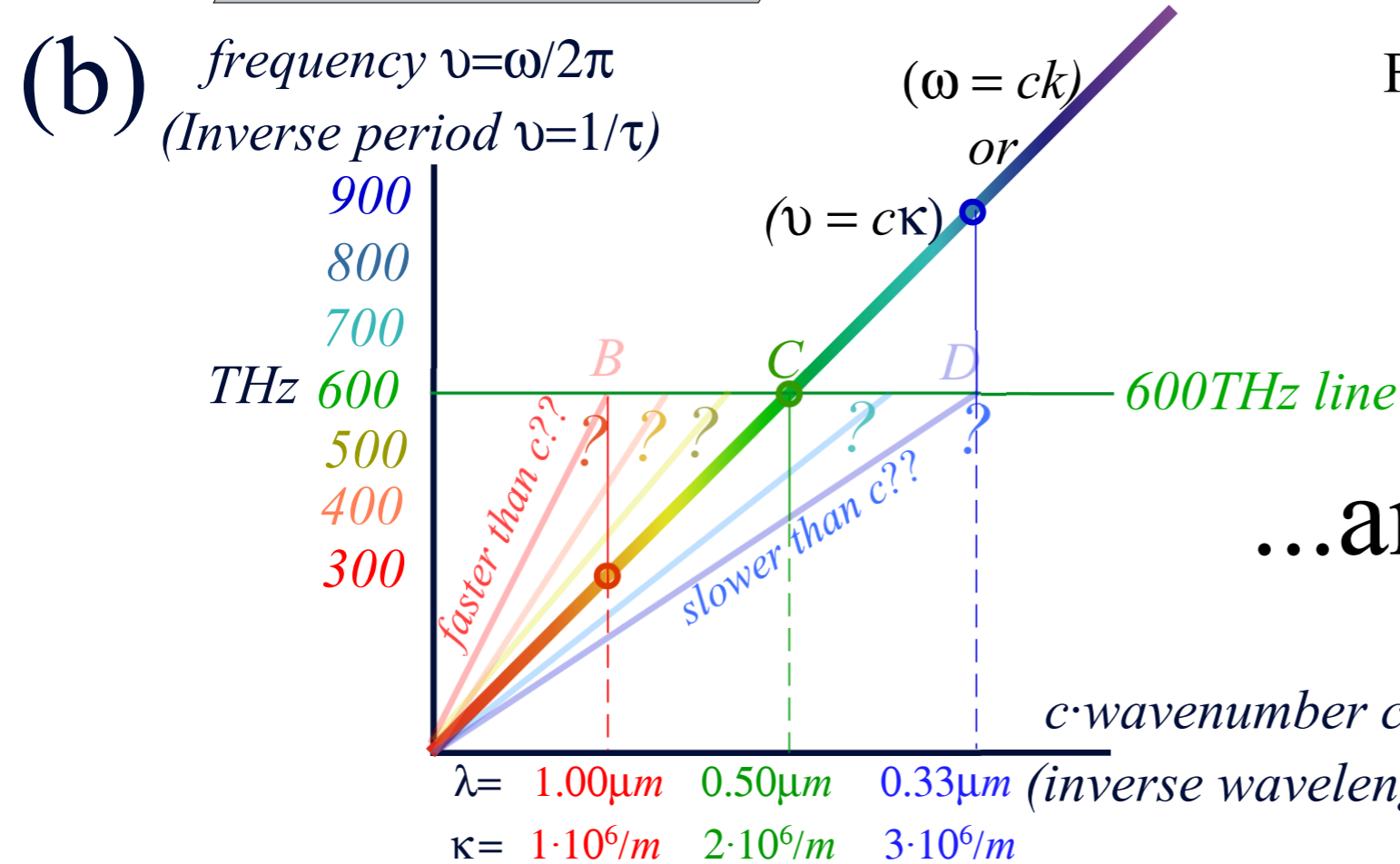
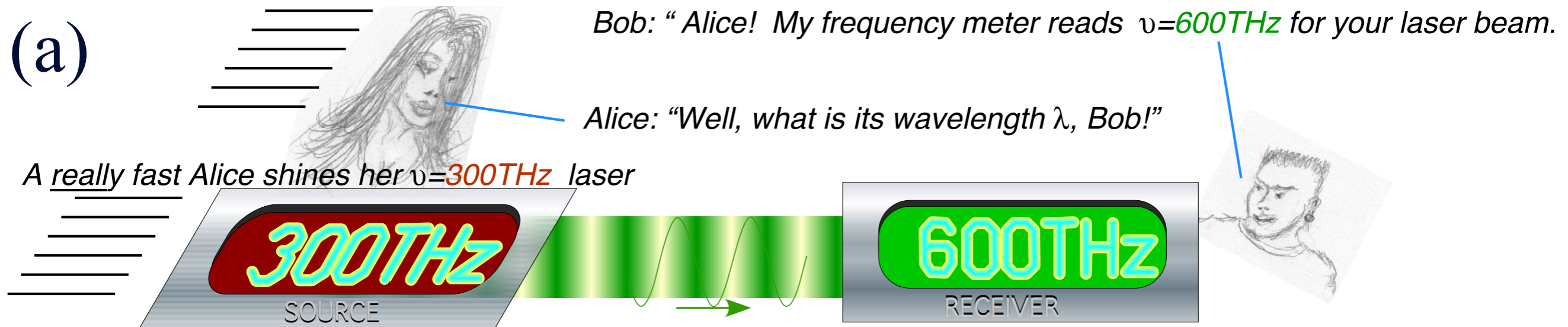


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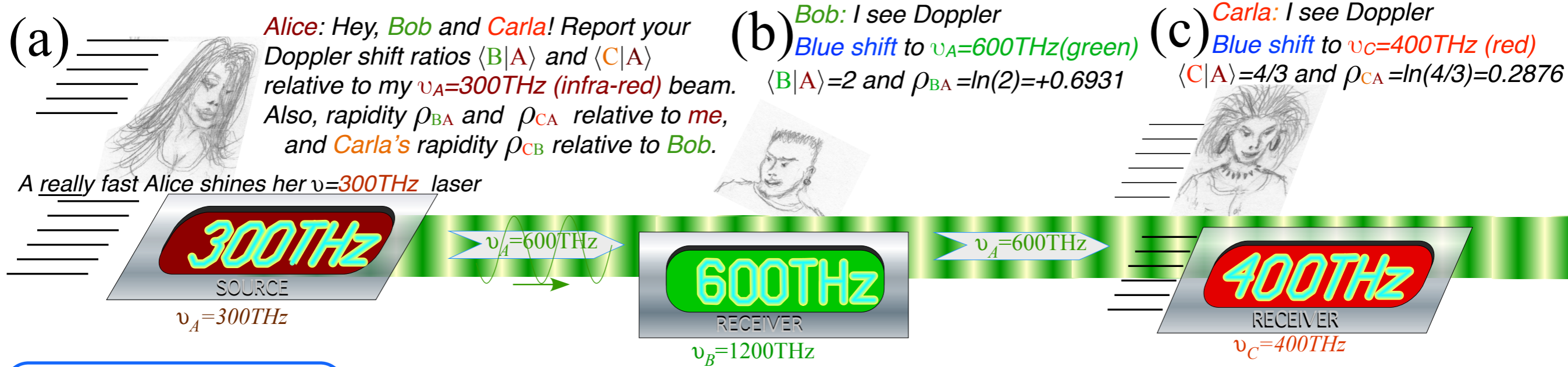
And, is he seeing a 'phony' green?

...and Dispersion-Free!

Years of spectroscopy rule out 'phony' 600THz blue-green that do not have wavelength $\lambda=0.5\text{micron}$.

The only choice is C. Also the only possible 600THz light speed is $c = \frac{\nu}{\kappa} = \frac{600 \cdot 10^{12}}{2 \cdot 10^6} = 3 \cdot 10^8 \text{ m} \cdot \text{s}^{-1}$

Actually: $2.99792458 \cdot 10^8 \text{ m} \cdot \text{s}^{-1}$



Doppler ratio:

$$\langle R|S \rangle = \frac{\nu_{RECEIVER}}{\nu_{SOURCE}}$$

$$\rho_{RS} = \ln \langle R|S \rangle$$

or:

$$\langle R|S \rangle = e^{\rho_{RS}} = e^{-\rho_{SR}}$$

Definition of Rapidity ρ_{RS}

Bob-Alice Doppler ratio:

$$\langle B|A \rangle = \frac{\nu_B}{\nu_A} = \frac{600}{300} = \frac{2}{1}$$

Bob-Alice rapidity:

$$\rho_{BA} = \ln \langle B|A \rangle = \ln \frac{2}{1} = 0.6931$$

$$\rho_{AB} = \ln \langle A|B \rangle = \ln \frac{1}{2} = -0.6931 = -\rho_{BA}$$

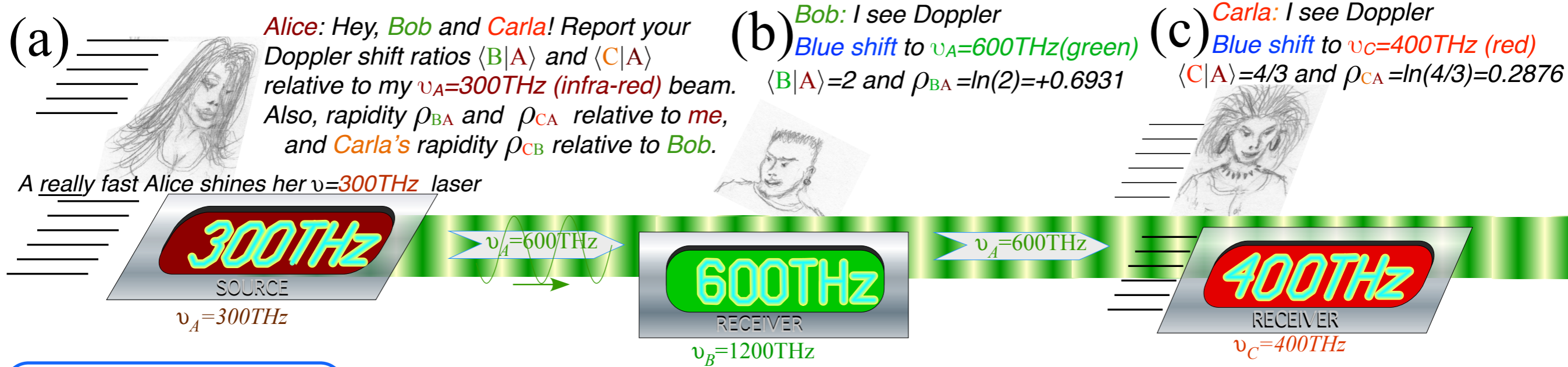
Carla-Alice Doppler ratio:

$$\langle C|A \rangle = \frac{\nu_C}{\nu_A} = \frac{400}{300} = \frac{4}{3}$$

Carla-Alice rapidity:

$$\rho_{CA} = \ln \langle C|A \rangle = \ln \frac{4}{3} = 0.2876$$

Introducing Doppler Arithmetic and rapidity ρ



Doppler ratio:

$$\langle R|S \rangle = \frac{\nu_{RECEIVER}}{\nu_{SOURCE}}$$

$$\rho_{RS} = \ln \langle R|S \rangle$$

or:

$$\langle R|S \rangle = e^{\rho_{RS}} = e^{-\rho_{SR}}$$

Definition of Rapidity ρ_{RS}

Bob-Alice Doppler ratio:

$$\langle B|A \rangle = \frac{\nu_B}{\nu_A} = \frac{600}{300} = 2$$

Bob-Alice rapidity:

$$\rho_{BA} = \ln \langle B|A \rangle = \ln \frac{2}{1} = 0.6931$$

$$\rho_{AB} = \ln \langle A|B \rangle = \ln \frac{1}{2} = -0.6931 = -\rho_{BA}$$

Carla-Alice Doppler ratio:

$$\langle C|A \rangle = \frac{\nu_C}{\nu_A} = \frac{400}{300} = \frac{4}{3}$$

Carla-Alice rapidity:

$$\rho_{CA} = \ln \langle C|A \rangle = \ln \frac{4}{3} = 0.2876$$

Carla-Bob Doppler ratio:

$$\langle C|B \rangle = \frac{\nu_C}{\nu_B} = \frac{\nu_C}{\nu_A} \frac{\nu_A}{\nu_B} = \langle C|A \rangle \langle A|B \rangle = \frac{4}{3} \frac{1}{2} = \frac{2}{3}$$

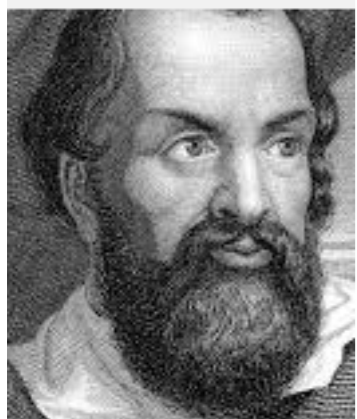
Carla-Bob rapidity:

$$e^{\rho_{CB}} = e^{\rho_{CA}} e^{\rho_{AB}} = e^{\rho_{CA} + \rho_{AB}}$$

$$\rho_{CB} = \rho_{CA} + \rho_{AB} = 0.2876 - 0.6931 = -0.4055$$

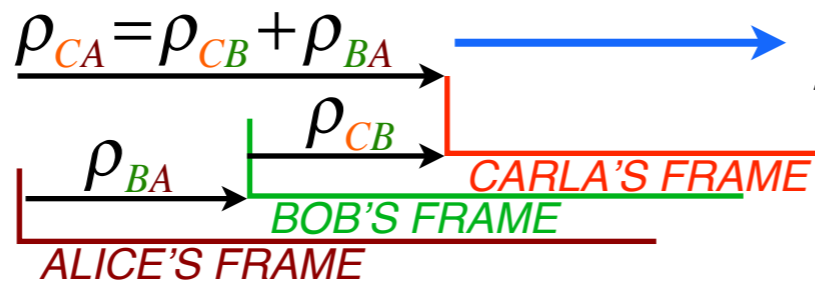
$$= \ln \frac{4}{3} + \ln \frac{1}{2} = \ln \frac{2}{3}$$

Galileo Galilei

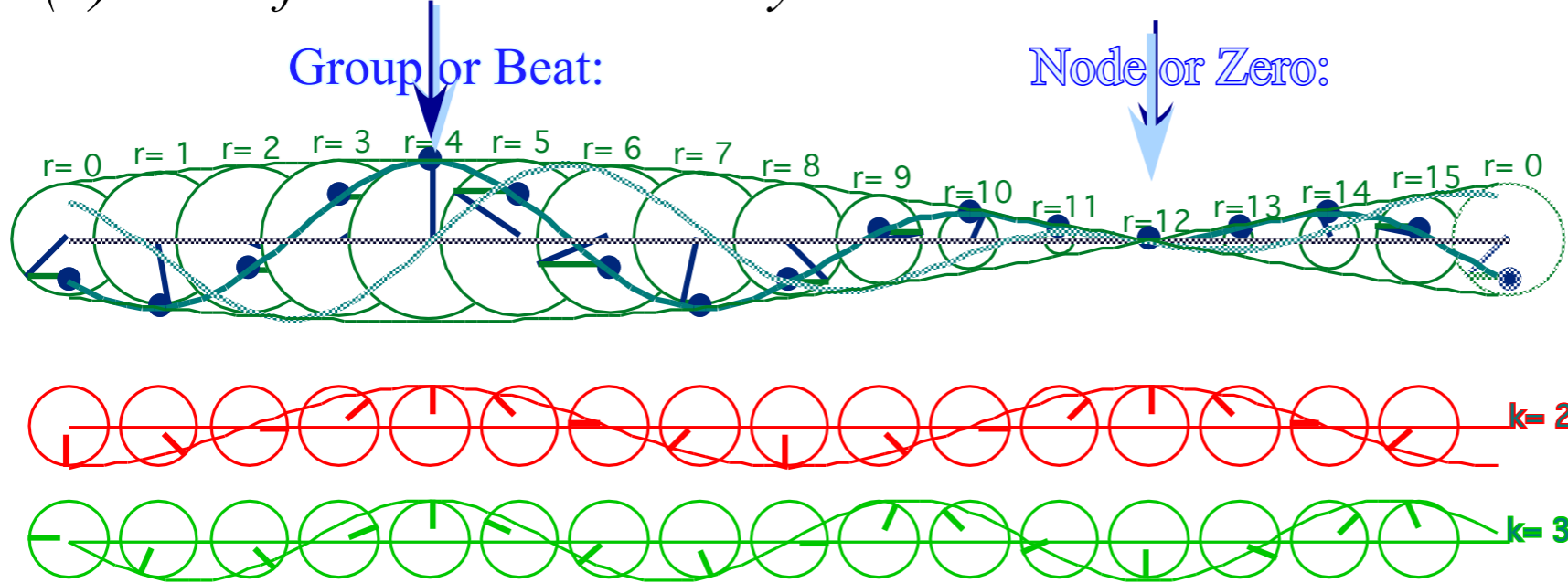


1564-1642

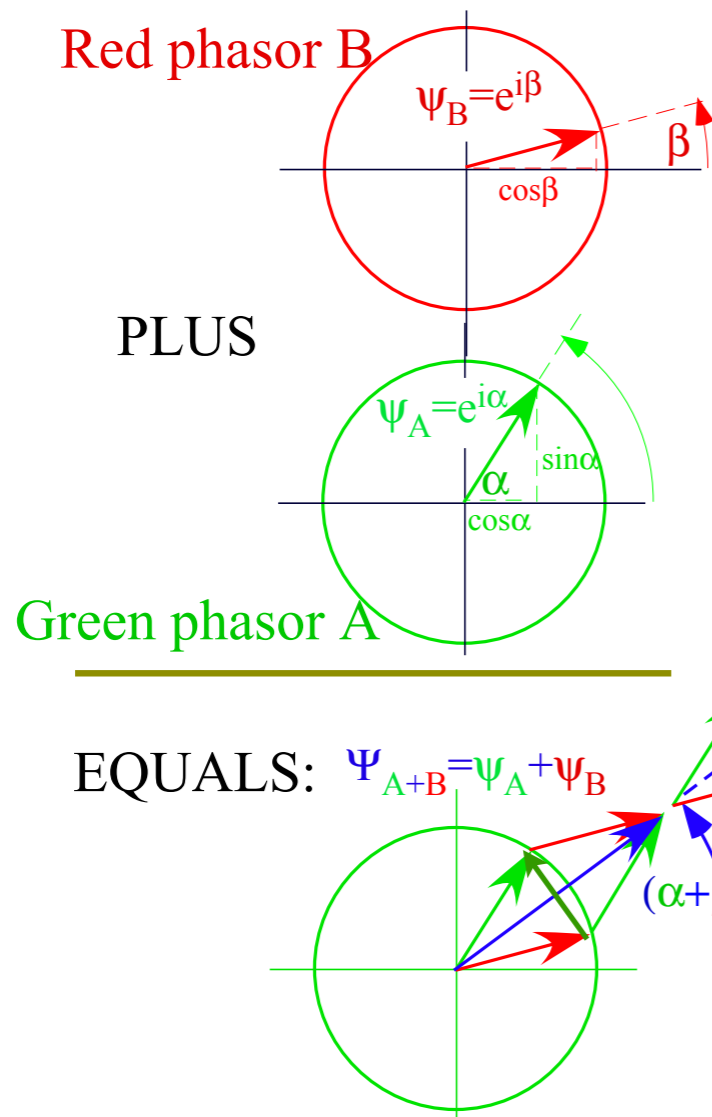
Galileo's Revenge (part 1)
Rapidity adds just like Galilean velocity



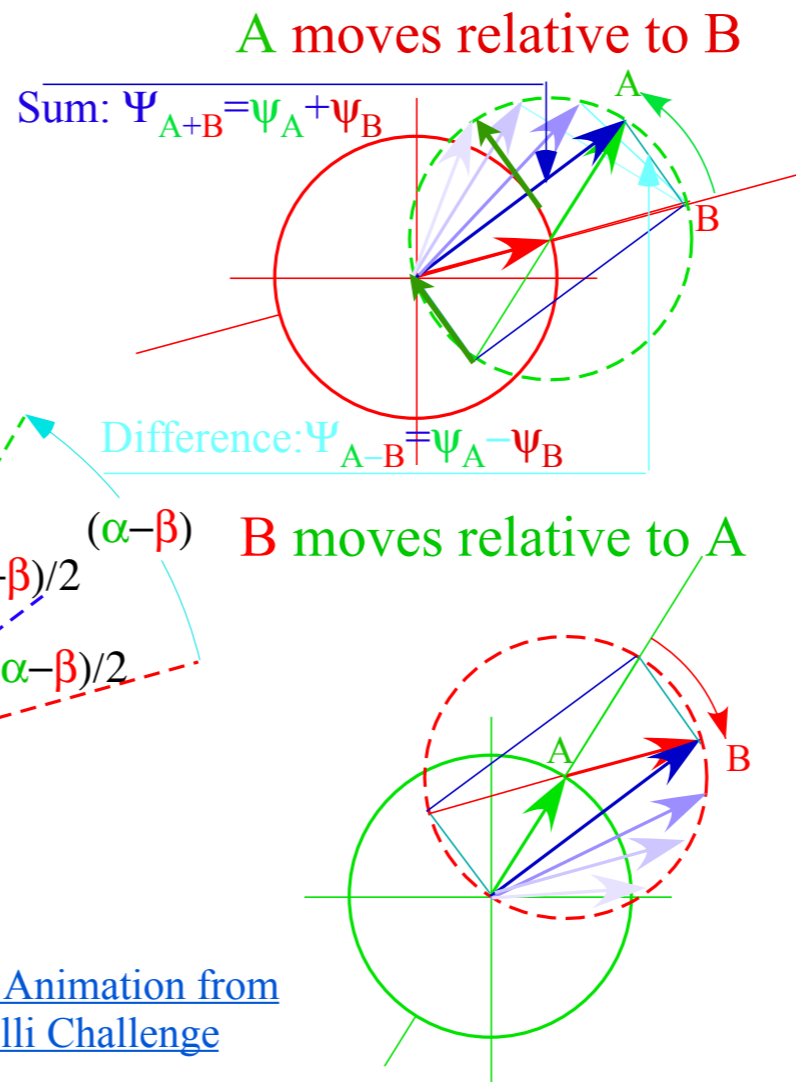
(a) Sum of Wave Phasor Array



(b) Typical Phasor Sum:

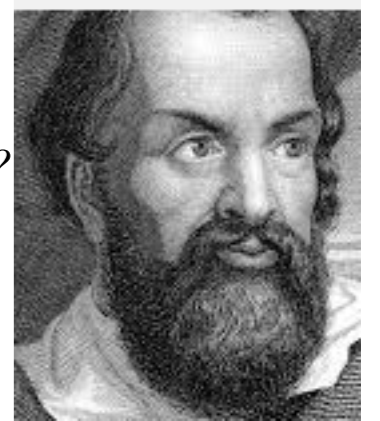


(c) Phasor-relative views



Geometry of the Half-sum Phase and Half-difference Group

Happy now?



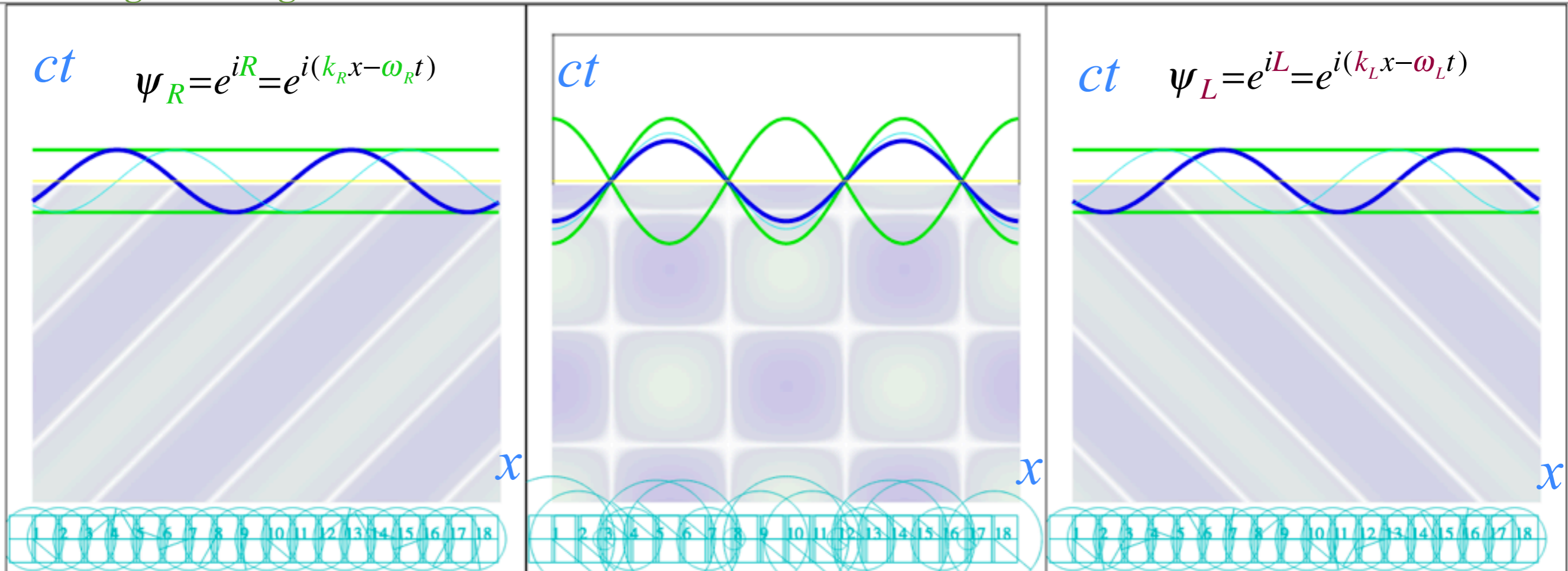
Galileo's Revenge (part 2)
Phasor angular velocity adds just like Galilean velocity

[Link to Animation from Pirelli Challenge](#)

right-moving CW laser

Colliding 2CW laser beams

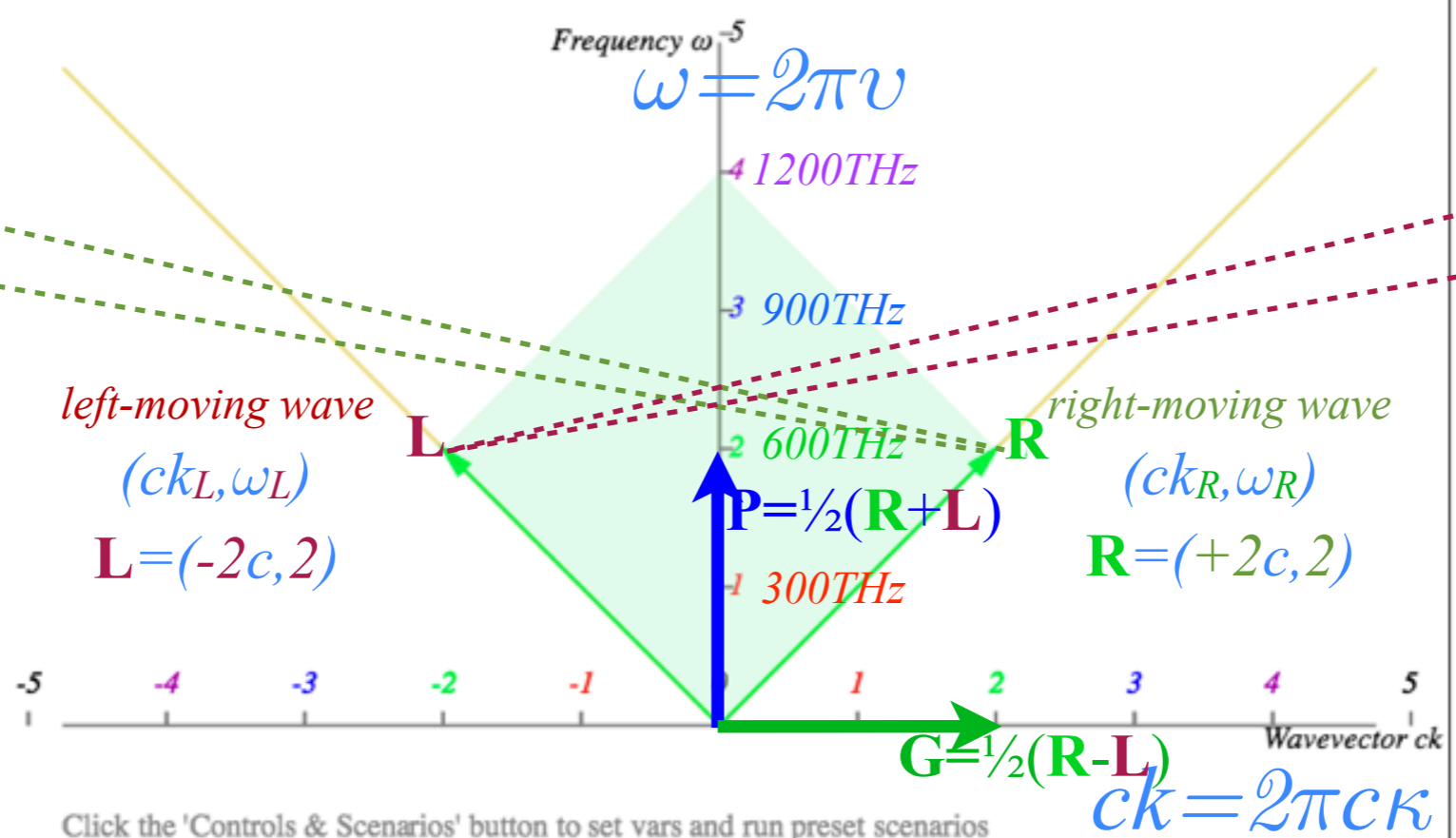
left-moving CW laser



right-moving wave
Spacetime (x, ct)

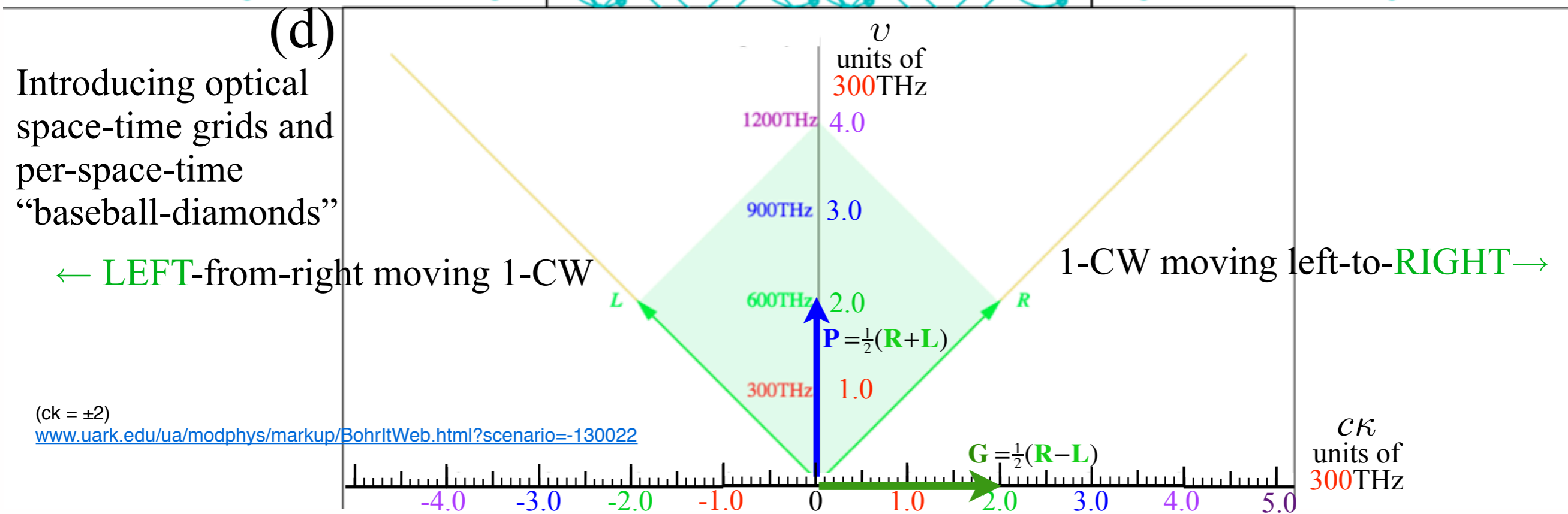
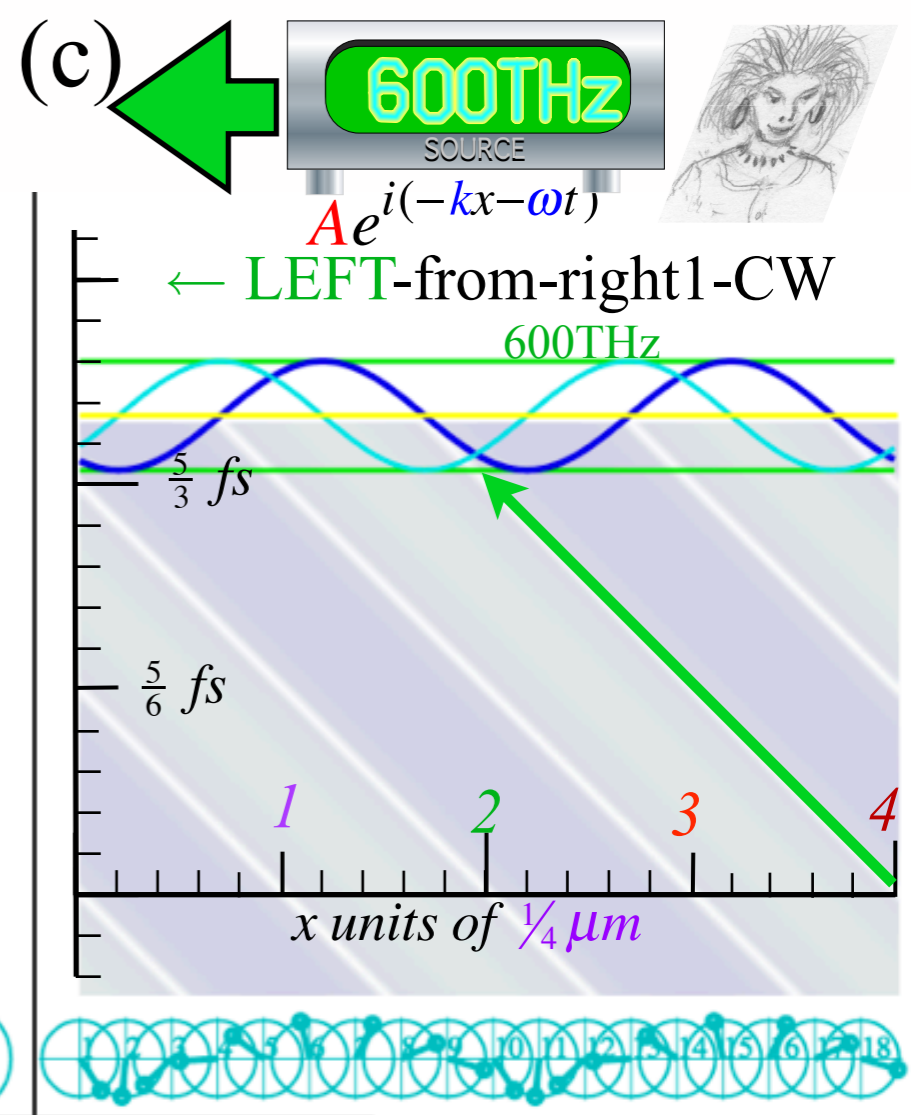
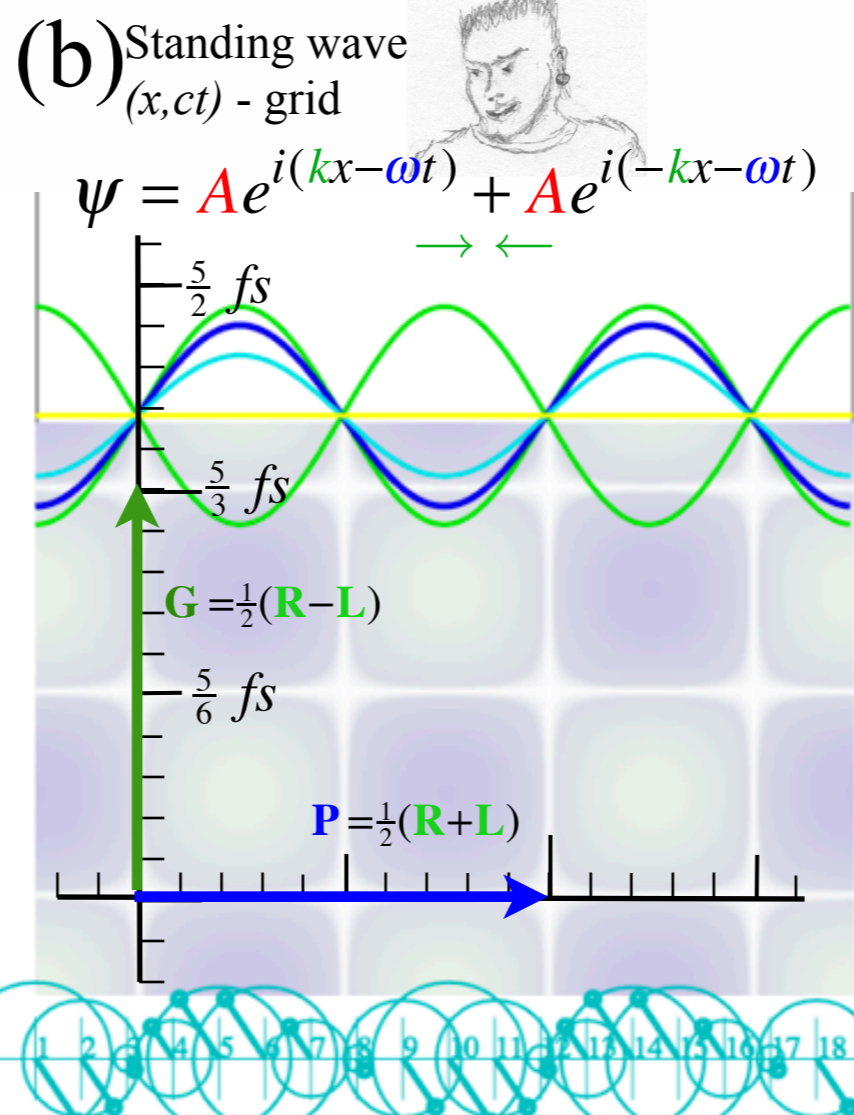
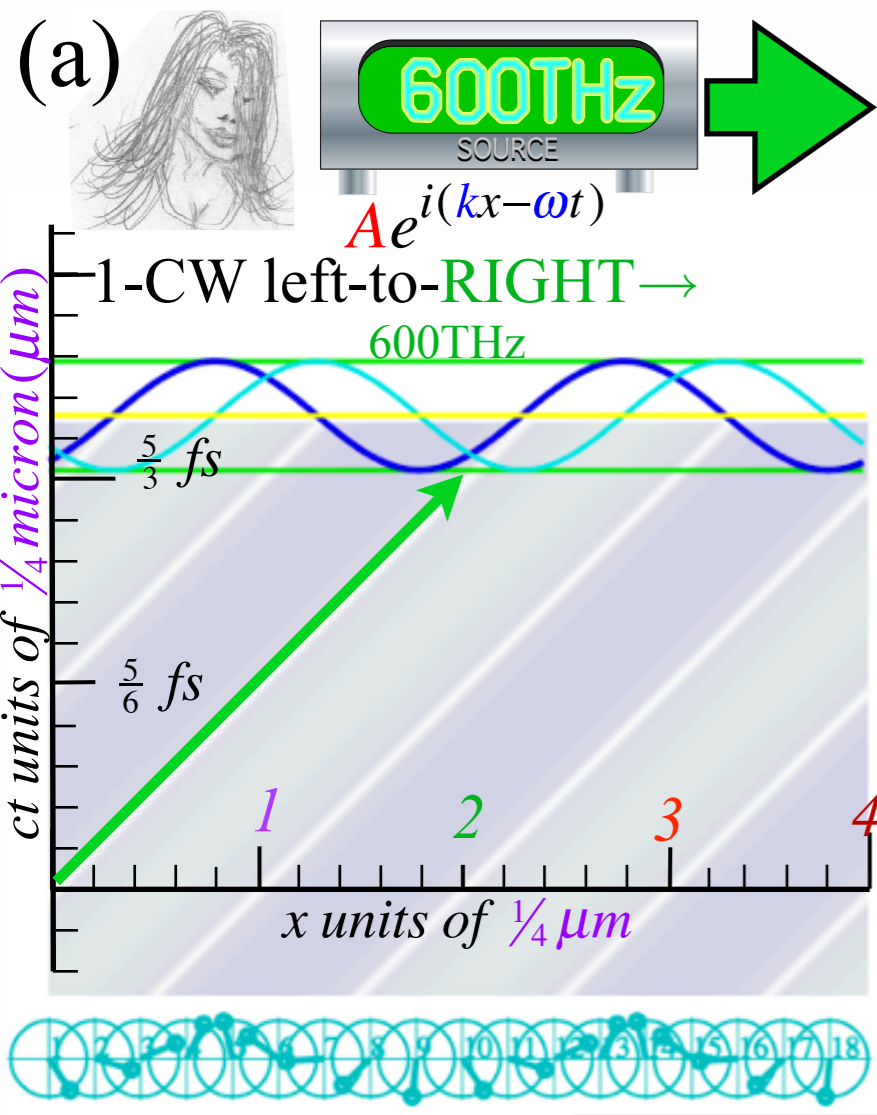
left-moving wave
Spacetime (x, ct)

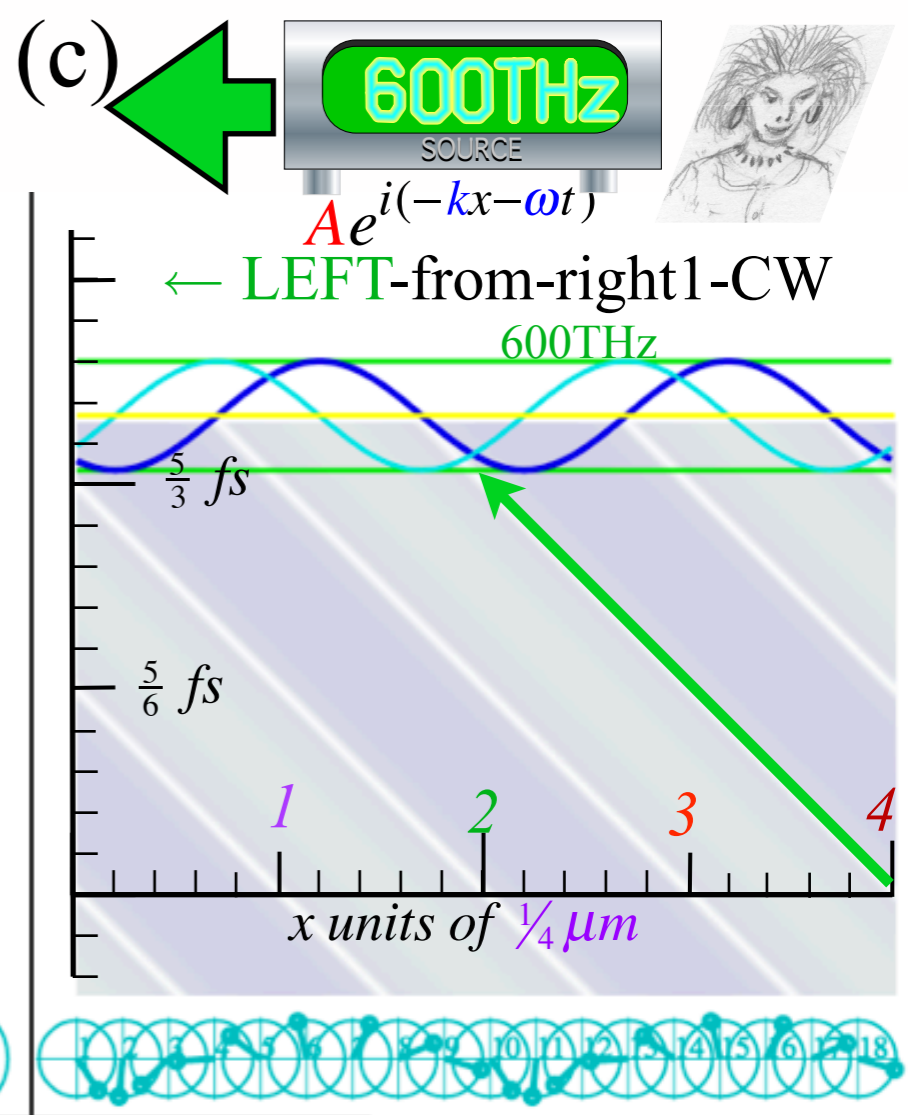
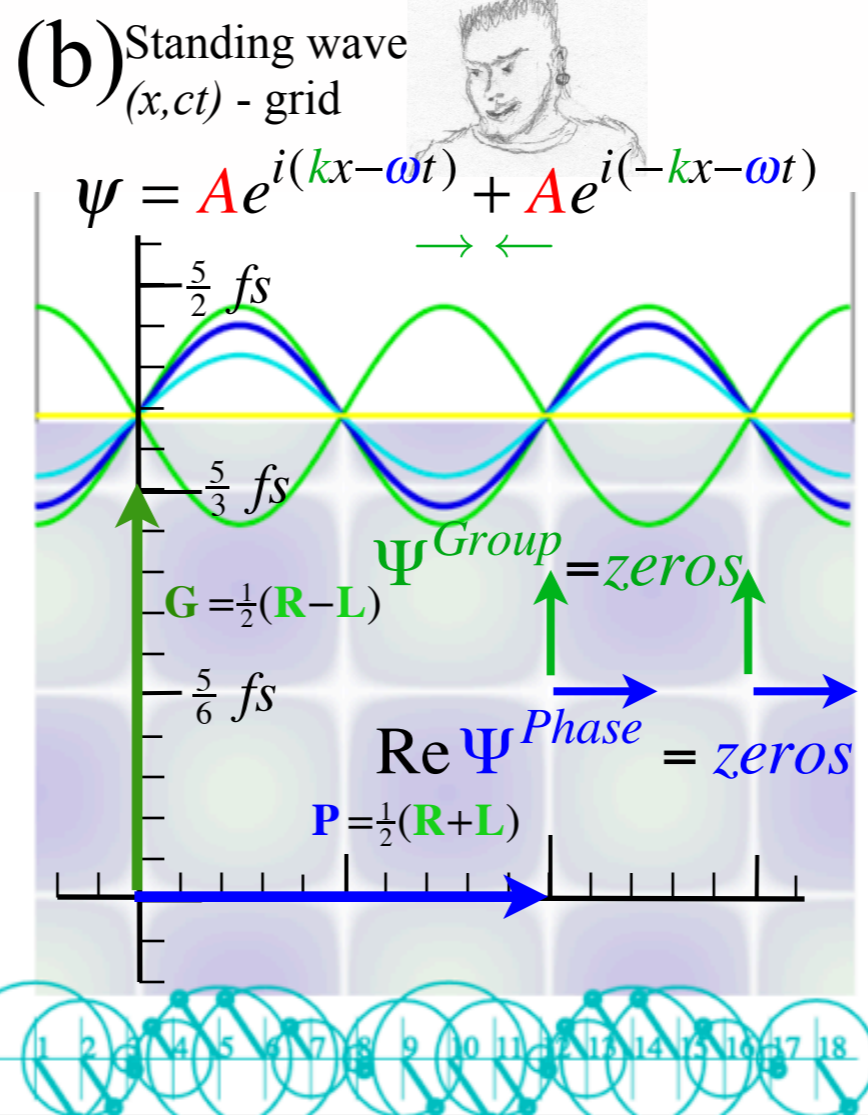
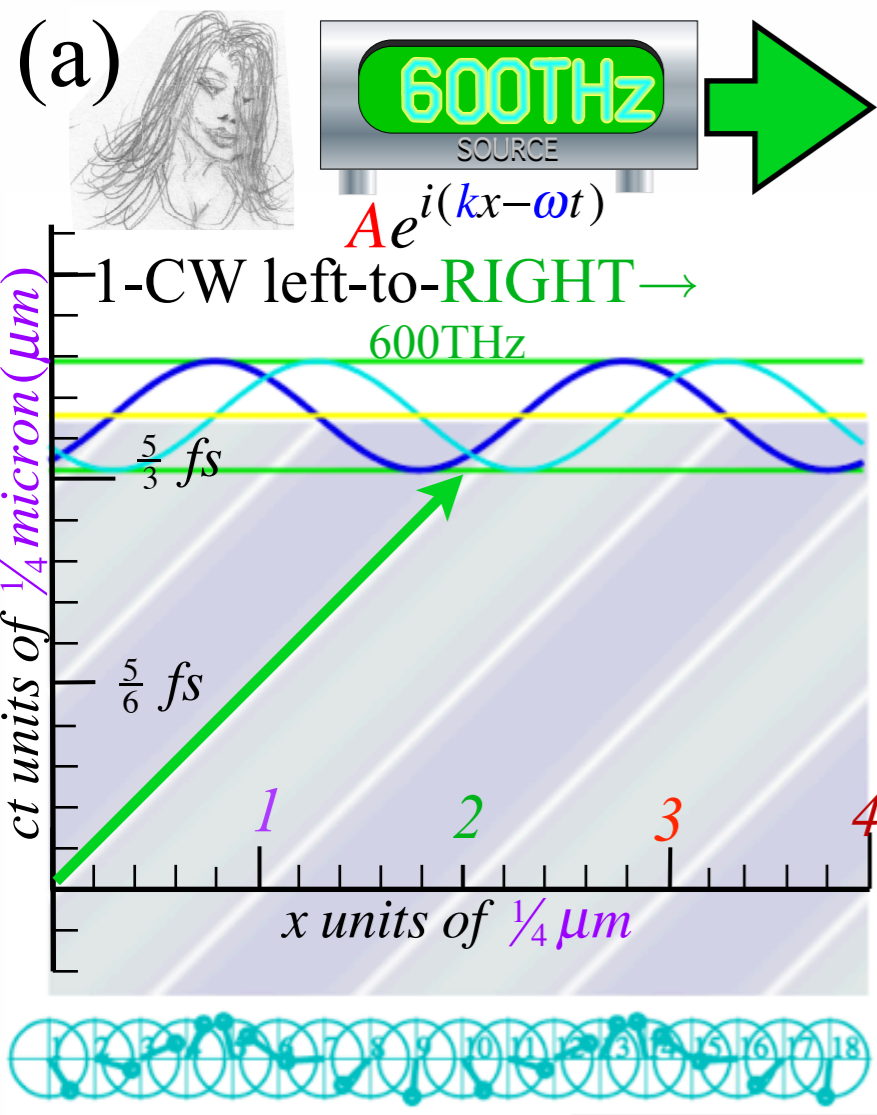
Per-Spacetime
(ck, ω)



BohrIt Web Simulation 2
CW ct vs x Plot (ck = ±2)

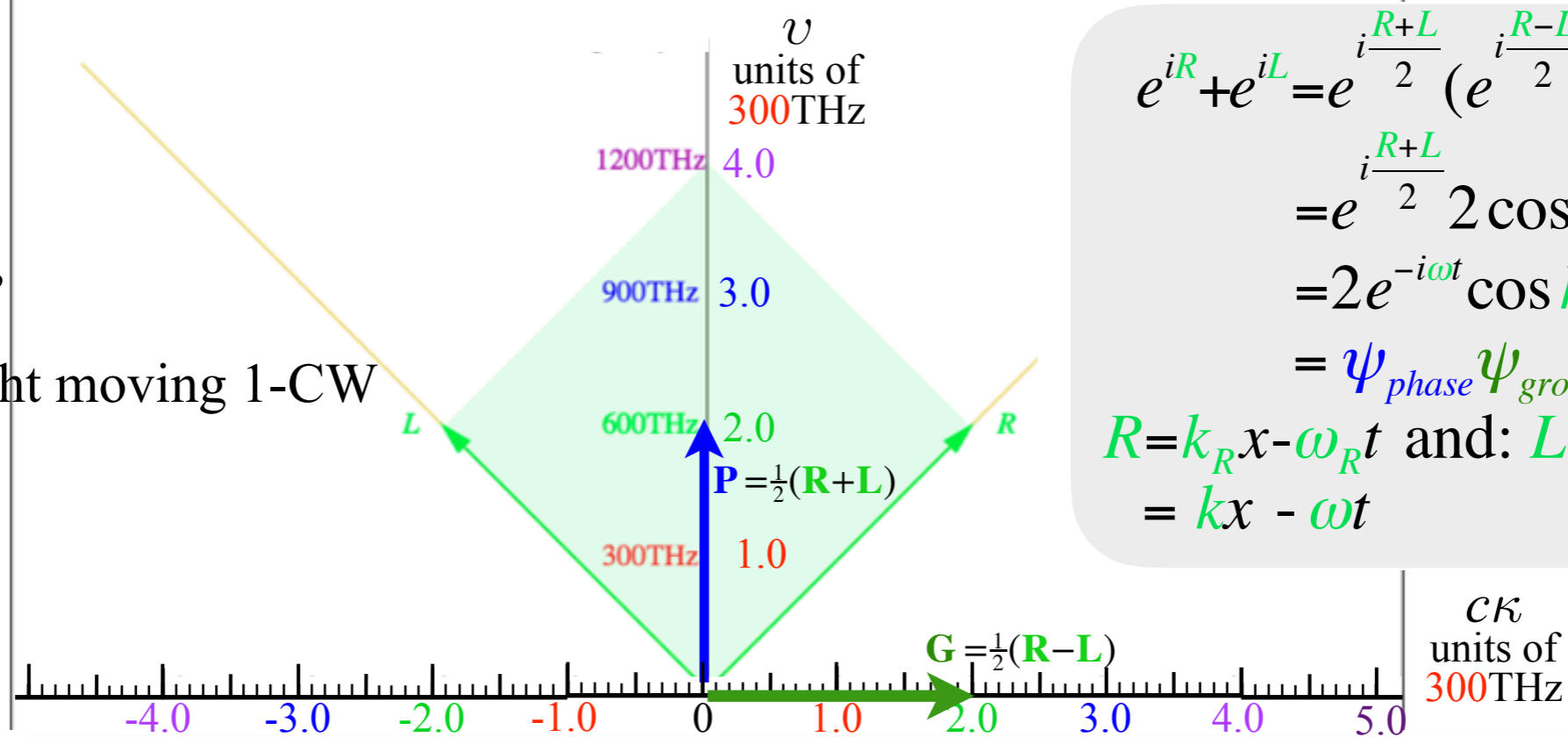
Click the 'Controls & Scenarios' button to set vars and run preset scenarios
Set the right & left-ward k values with clicks near the dispersion curve or ck axis.





(d) Introducing optical space-time grids and per-space-time “baseball-diamonds”

\leftarrow LEFT-from-right moving 1-CW



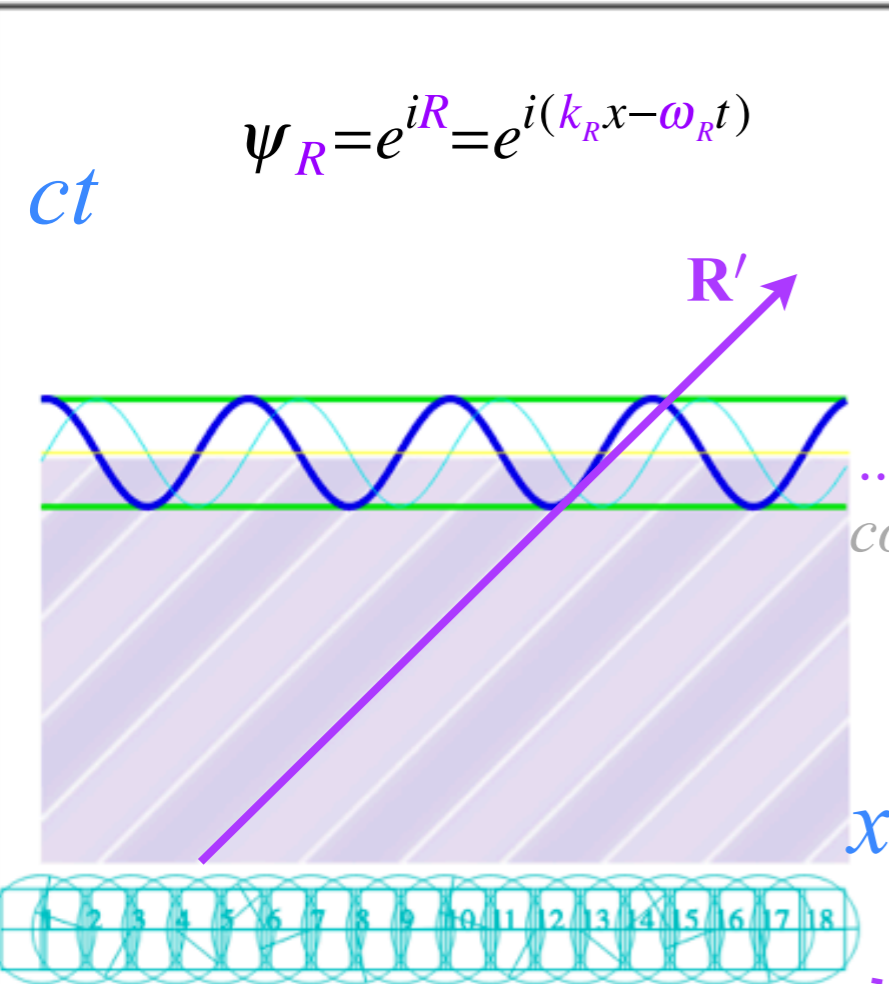
$$\begin{aligned}
 e^{iR} + e^{iL} &= e^{i\frac{R+L}{2}} (e^{i\frac{R-L}{2}} + e^{-i\frac{R-L}{2}}) \\
 &= e^{i\frac{R+L}{2}} 2 \cos \frac{R-L}{2} \\
 &= 2e^{-i\omega t} \cos kx \\
 &= \psi_{phase} \psi_{group} \\
 R &= k_R x - \omega_R t \text{ and: } L = -k_L x - \omega_L t \\
 &= kx - \omega t \qquad \qquad \qquad = -kx - \omega t
 \end{aligned}$$

right-moving Doppler blue shifted wave

left-moving Doppler red shifted wave

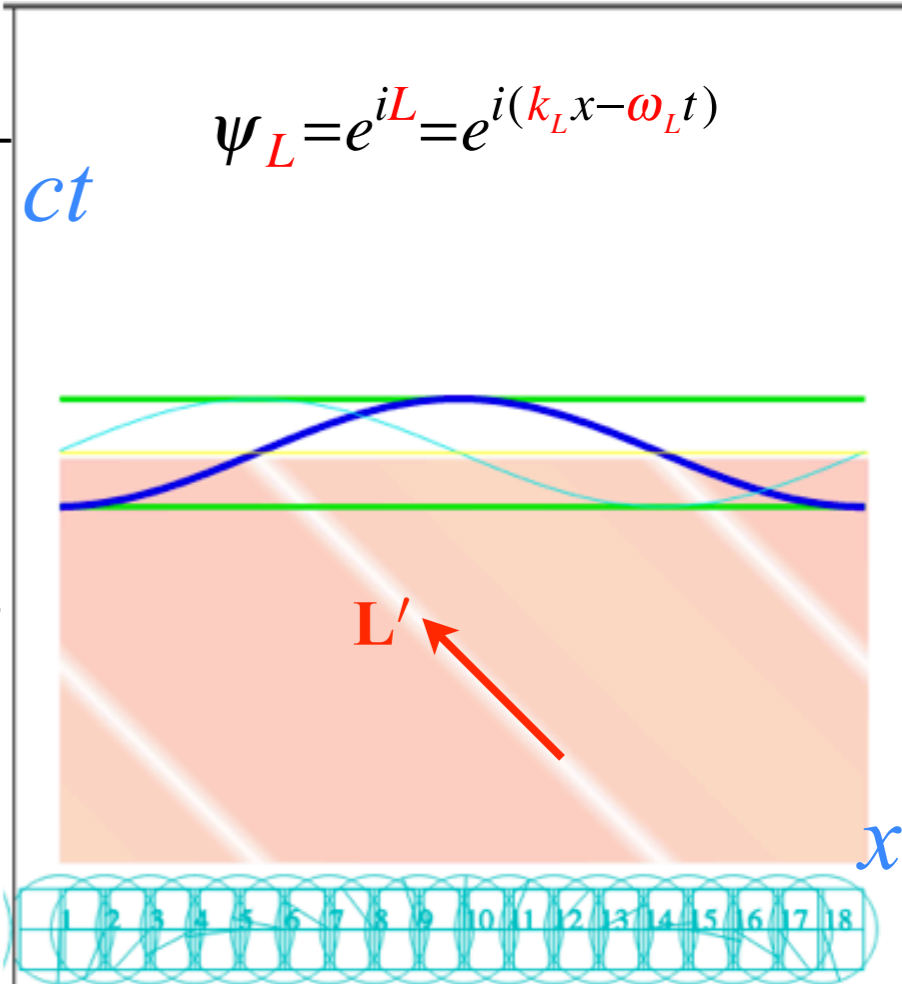


Rapidly moving Bob sees...



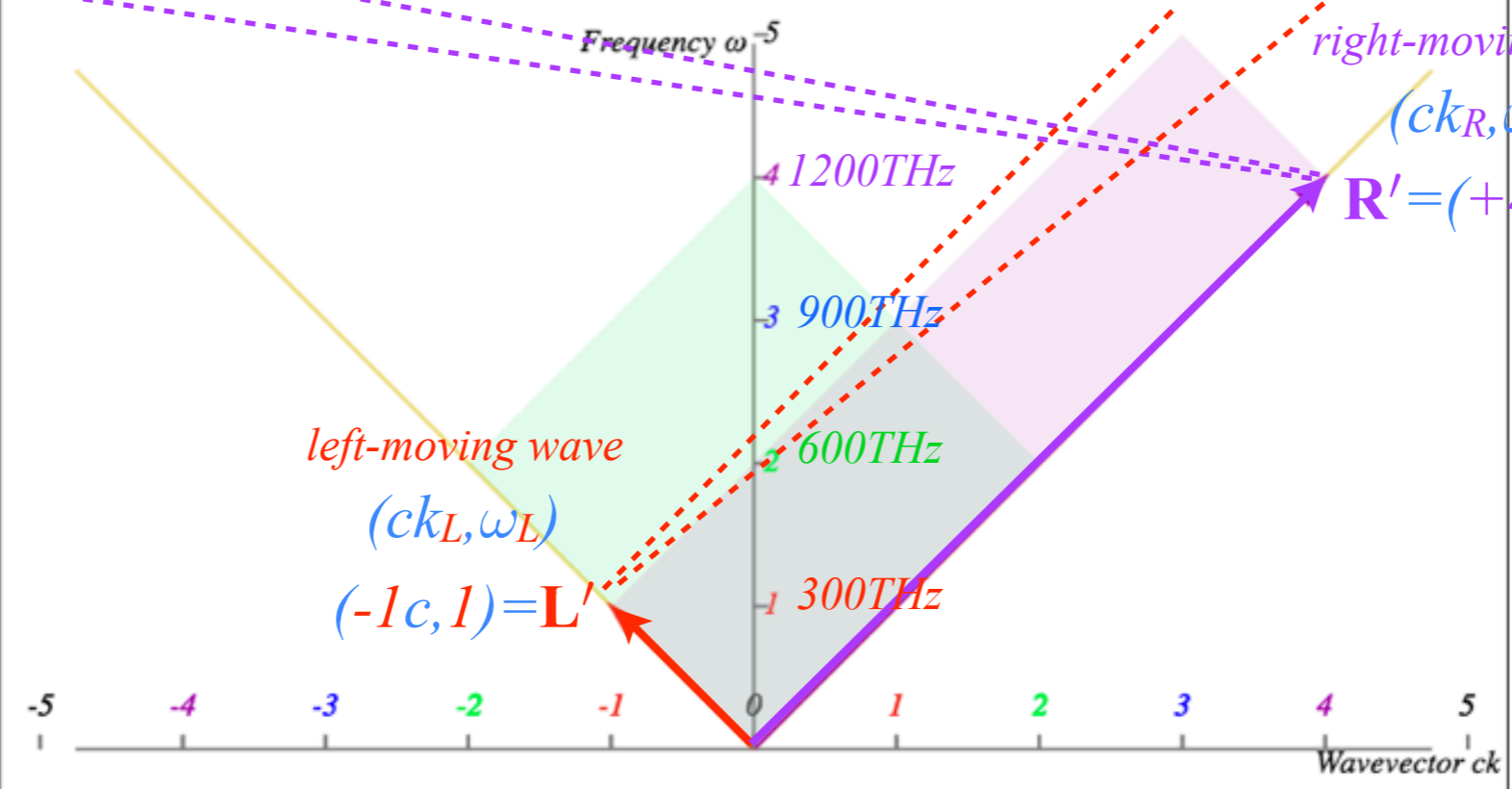
...Blue shifted wave coming at him and...

...Red shifted wave behind him.



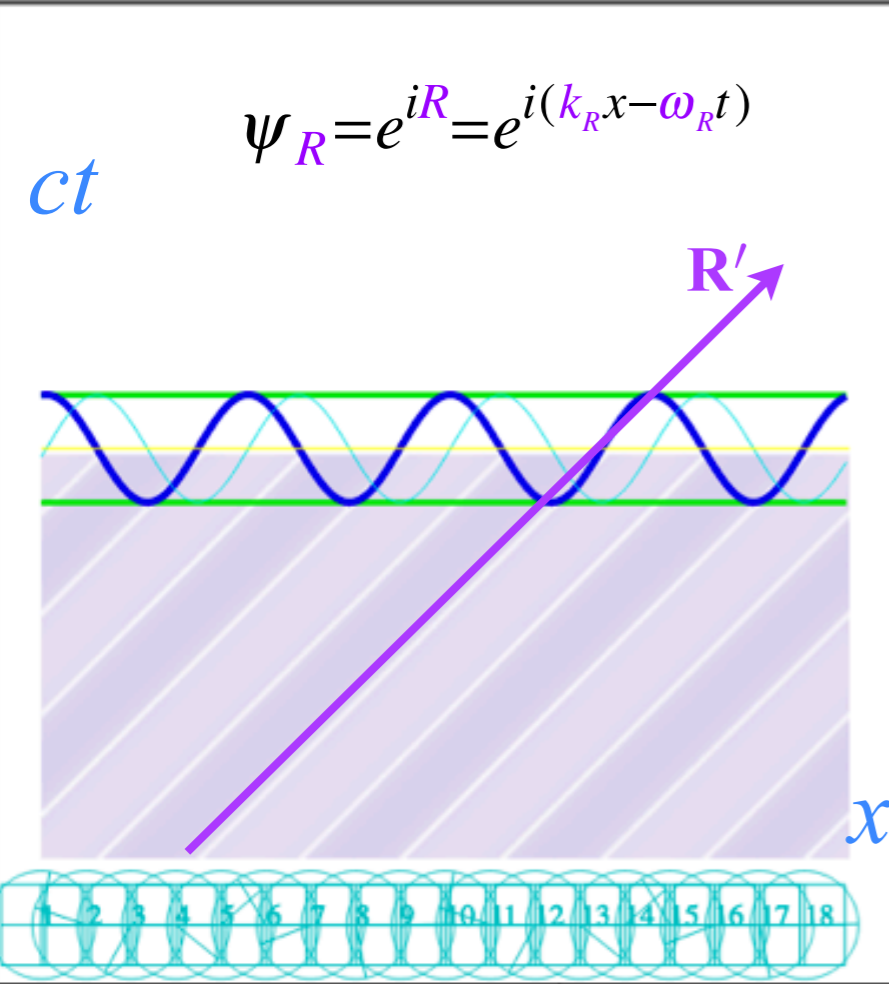
[Web Simulation](#)
1 CW ct vs x Plot
($ck = +4$)

[Web Simulation](#)
1 CW ct vs x Plot
($ck = -1$)

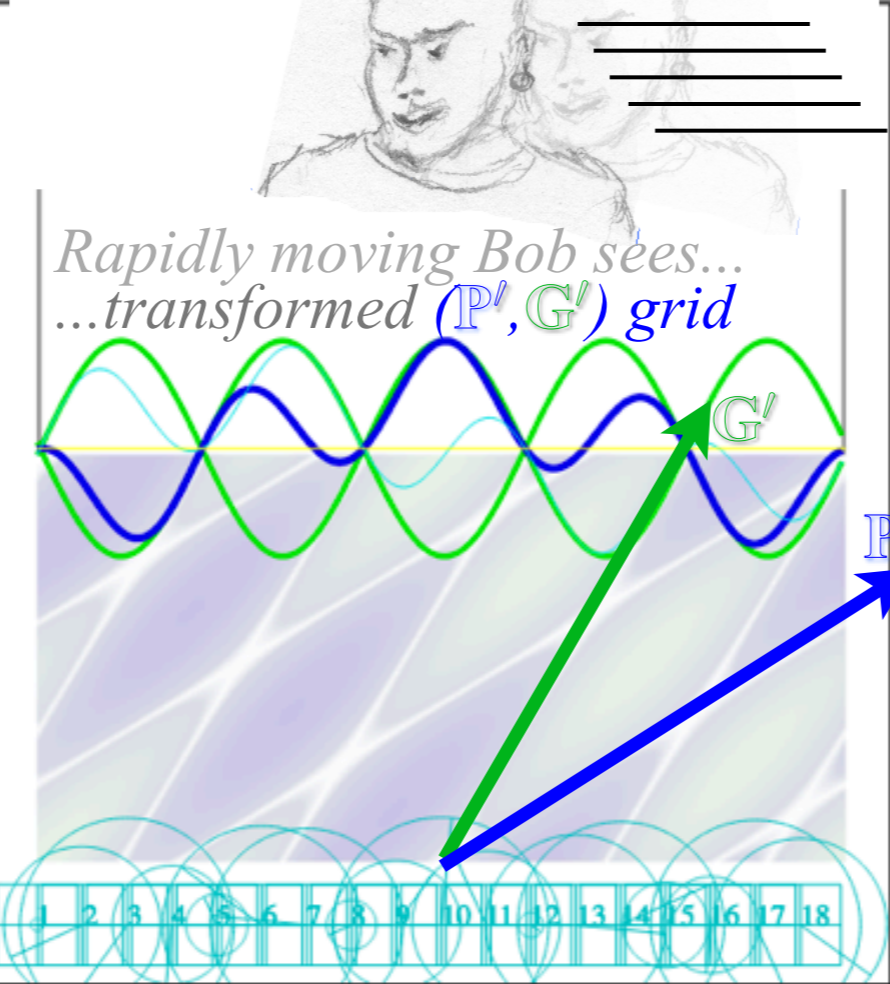


right-moving Doppler blue shifted wave

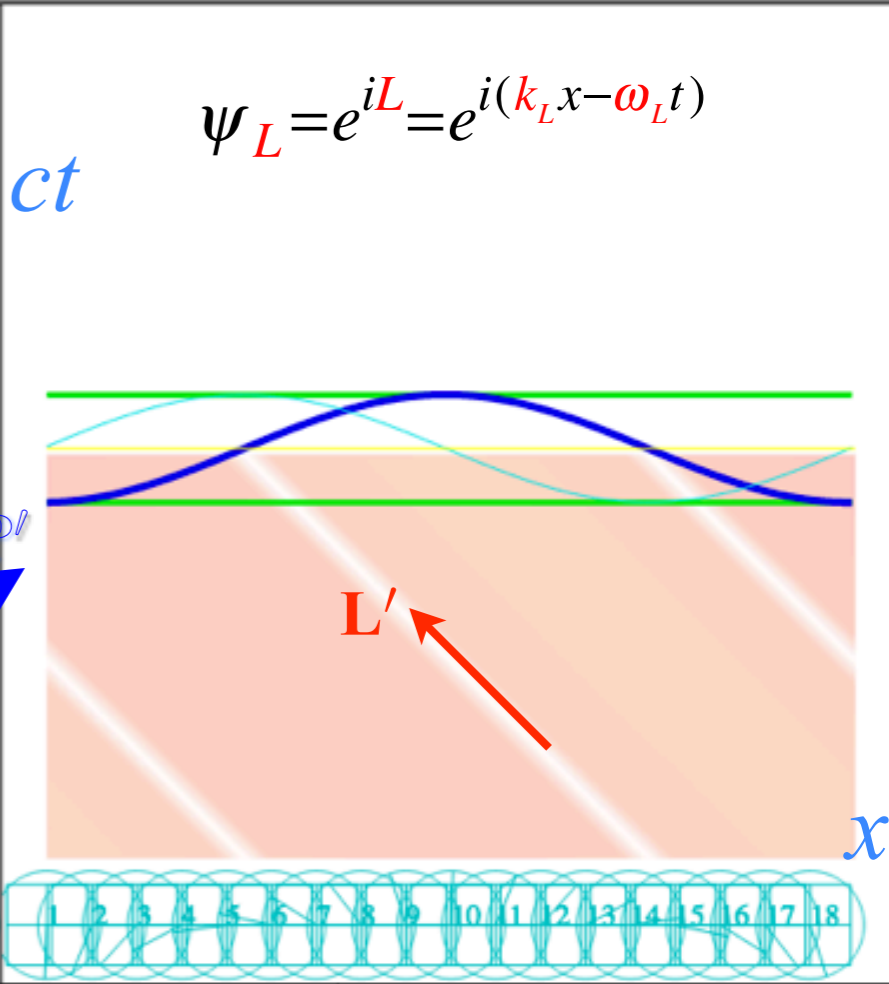
left-moving Doppler red shifted wave



$$\psi_R = e^{iR} = e^{i(k_R x - \omega_R t)}$$



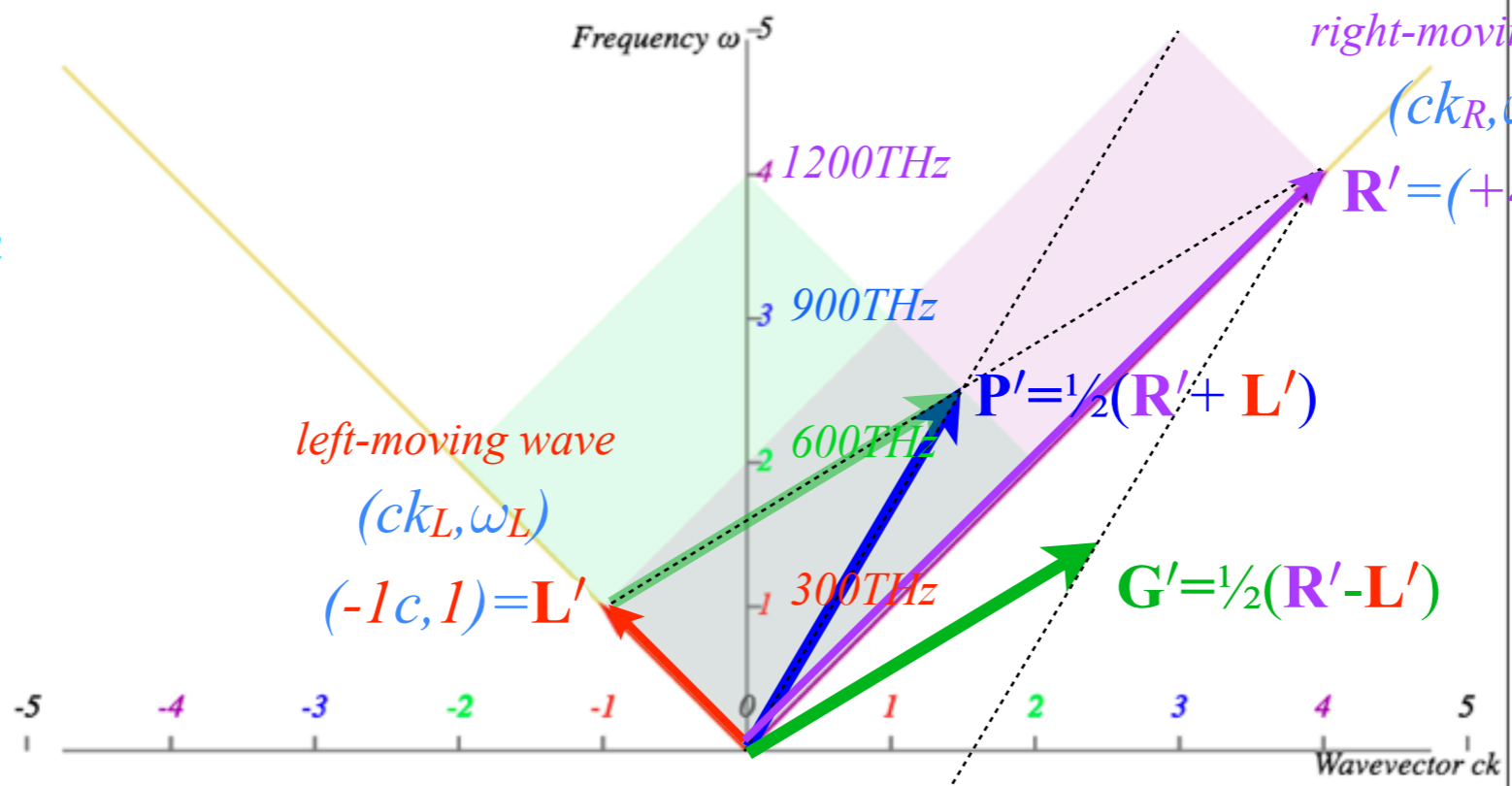
Rapidly moving Bob sees...
...transformed (P', G') grid



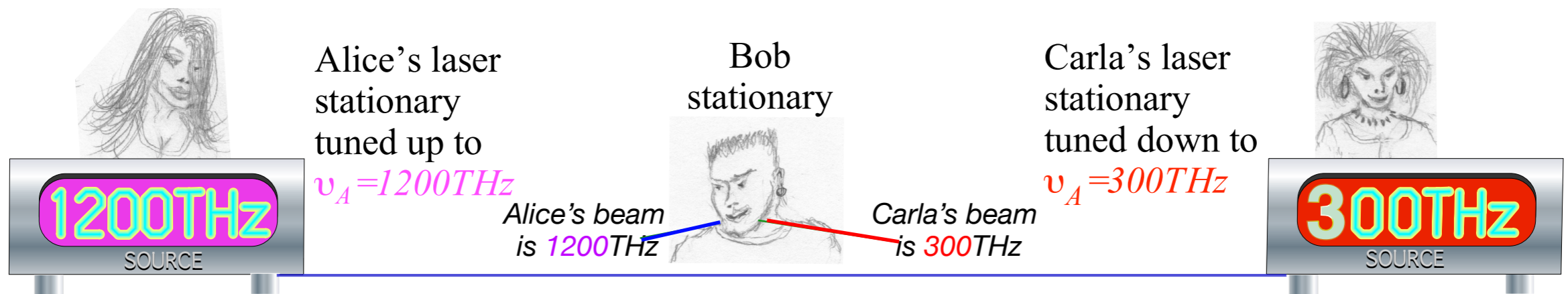
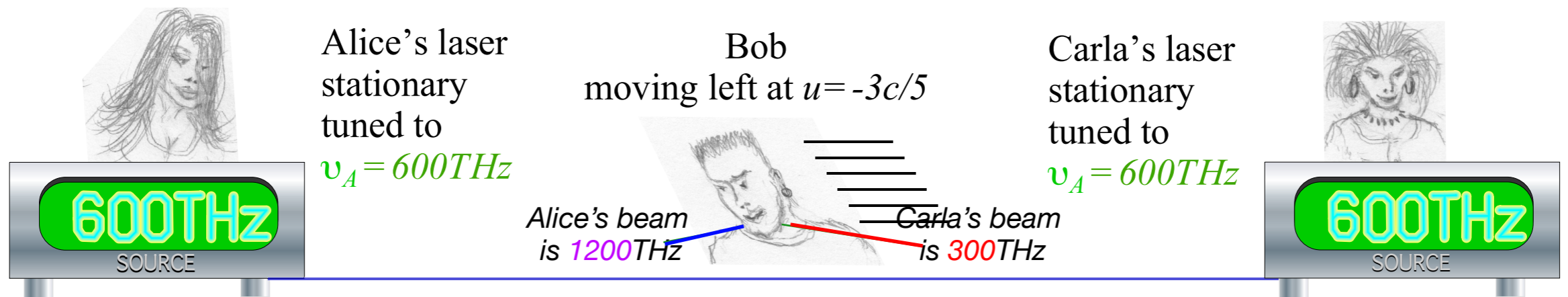
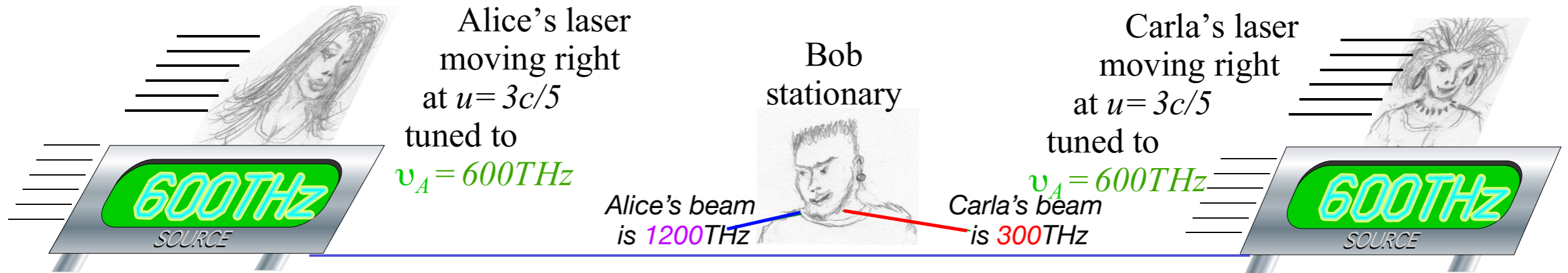
$$\psi_L = e^{iL} = e^{i(k_L x - \omega_L t)}$$

...Doppler shifts give
Lorentz transformation
of both these graphs

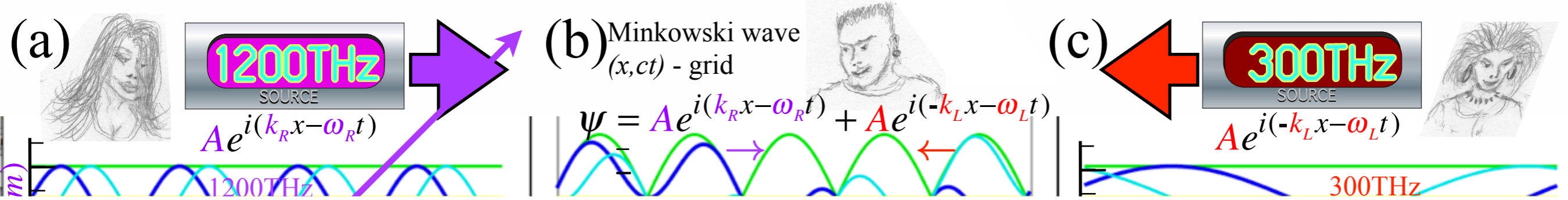
Per-Spacetime
(ck, ω)



Three scenarios that look the same to Bob



Much cheaper (and safer) to do the 3rd scenario!\$!



$$\mathbf{P}' = \begin{pmatrix} v'_{phase} \\ cK'_{phase} \end{pmatrix} = \frac{1}{2}(\mathbf{R}' + \mathbf{L}') = v_A \begin{pmatrix} \frac{1}{2}(e^\rho + e^{-\rho}) \\ \frac{1}{2}(e^\rho - e^{-\rho}) \end{pmatrix} = v_A \begin{pmatrix} \cosh \rho \\ \sinh \rho \end{pmatrix} = v_A \begin{pmatrix} \frac{5}{2} \\ \frac{3}{2} \end{pmatrix} \text{ Bob's View} \quad \text{or: } v_A \begin{pmatrix} 1 \\ 0 \end{pmatrix} \text{ Alice's View}$$

$$\mathbf{G}' = \begin{pmatrix} v'_{group} \\ cK'_{group} \end{pmatrix} = \frac{1}{2}(\mathbf{R}' - \mathbf{L}') = v_A \begin{pmatrix} \frac{1}{2}(e^\rho - e^{-\rho}) \\ \frac{1}{2}(e^\rho + e^{-\rho}) \end{pmatrix} = v_A \begin{pmatrix} \sinh \rho \\ \cosh \rho \end{pmatrix} = v_A \begin{pmatrix} \frac{3}{2} \\ \frac{5}{2} \end{pmatrix} \text{ Bob's View} \quad \text{or: } v_A \begin{pmatrix} 0 \\ 1 \end{pmatrix} \text{ Alice's View}$$

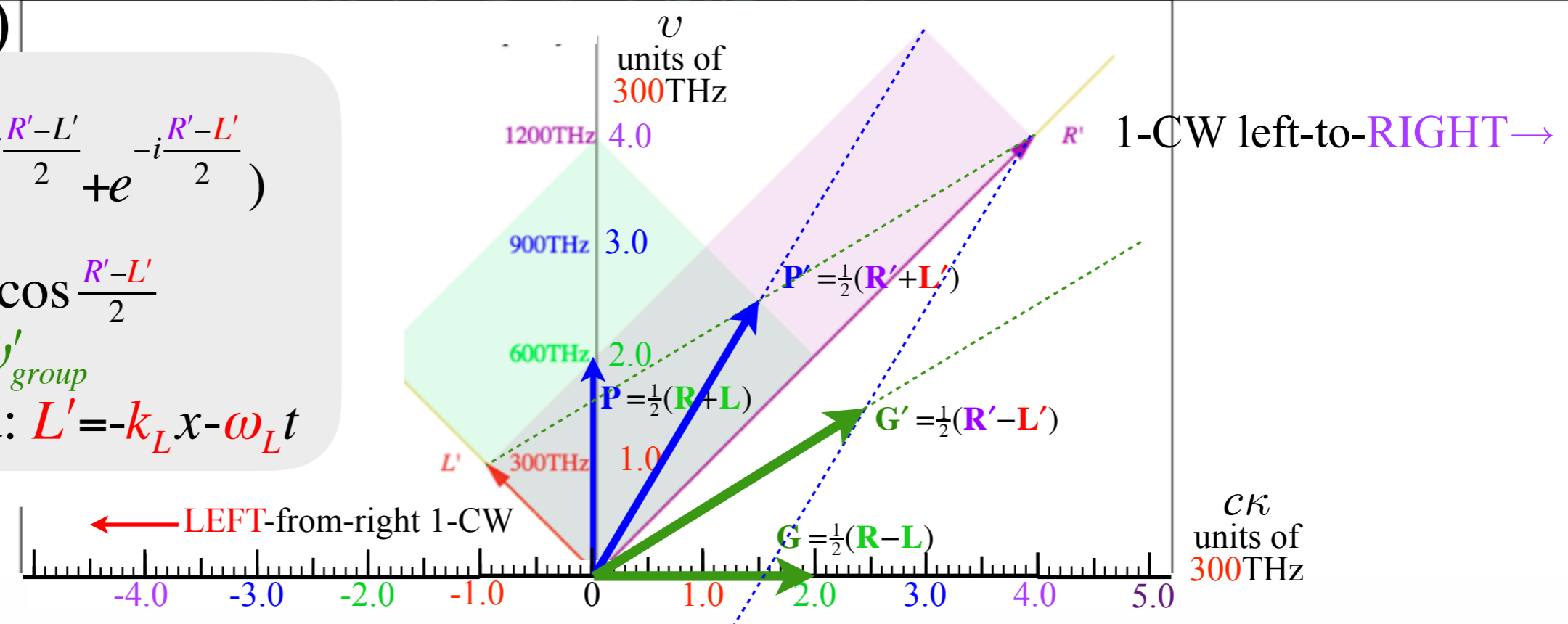
(d)

$$e^{iR'} + e^{iL'} = e^{i\frac{R'+L'}{2}} \left(e^{i\frac{R'-L'}{2}} + e^{-i\frac{R'-L'}{2}} \right)$$

$$= e^{i\frac{R'+L'}{2}} 2 \cos \frac{R'-L'}{2}$$

$$= \psi'_{phase} \psi'_{group}$$

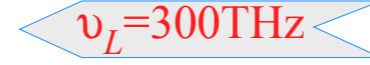
$R' = k_R x - \omega_R t$ and: $L' = -k_L x - \omega_L t$



Doppler Jeopardy


$$v_R = 600\text{THz}$$




$$v_L = 300\text{THz}$$

- (1.) To what velocity u_E must Bob accelerate so he sees beams with equal frequency ω_E ?
- (2.) What is that frequency ω_E ?

Doppler Jeopardy

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(2.) What is that frequency ω_E ?

Query (1.) has a Jeopardy-style answer-by-question: What is beam group velocity?

$$u_E = V_{group} = \frac{\omega_{group}}{k_{group}} = \frac{\omega_R - \omega_L}{k_R - k_L} = c \frac{\omega_R - \omega_L}{\omega_R + \omega_L}$$

Doppler Jeopardy

$$\nu_R = 600 \text{ THz}$$



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(1.) To what velocity u_E must Bob accelerate so he sees beams with equal frequency ω_E ?

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Query (1.) has a Jeopardy-style answer-by-question: What is beam group velocity?

$$\frac{300}{900}$$

$$u_E = V_{group} = \frac{\omega_{group}}{k_{group}} = \frac{\omega_R - \omega_L}{k_R - k_L} = c \frac{\omega_R - \omega_L}{\omega_R + \omega_L} \quad V_{group} = c \frac{\omega_R - \omega_L}{\omega_R + \omega_L} = c \frac{600 - 300}{600 + 300} = \frac{1}{3} c$$

Doppler Jeopardy

$$\nu_R = 600 \text{ THz}$$



$$\nu_L = 300 \text{ THz}$$

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Query (2.) similarly: What ω_E is blue-shift $b\omega_L$ of ω_L and red-shift ω_R/b of ω_R ?

$$\omega_E = b\omega_L = \omega_R/b \quad \Rightarrow \quad b = \sqrt{\omega_R / \omega_L} \quad \Rightarrow \quad \omega_E = \sqrt{\omega_R \cdot \omega_L}$$

Doppler Jeopardy

$$\nu_R = 600 \text{ THz}$$



$$\nu_L = 300 \text{ THz}$$

(1.) To what velocity u_E must Bob accelerate so he sees beams with equal frequency ω_E ?

(2.) What is that frequency ω_E ?

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Query (2.) similarly: What ω_E is blue-shift $b\omega_L$ of ω_L and red-shift ω_R/b of ω_R ?

$$\sqrt{6 \cdot 3} = 3\sqrt{2} = 4.24$$

$$\omega_E = b\omega_L = \omega_R/b \Rightarrow b = \sqrt{\omega_R / \omega_L} \Rightarrow \omega_E = \sqrt{\omega_R \cdot \omega_L}$$

$$\begin{aligned} \omega_E &= \sqrt{\omega_R \cdot \omega_L} \\ &= \sqrt{180000} \\ &= 424 \end{aligned}$$

Geometric mean

Doppler Jeopardy

$$\nu_R = 600 \text{ THz}$$



$$\nu_L = 300 \text{ THz}$$

(1.) To what velocity u_E must Bob accelerate so he sees beams with equal frequency ω_E ?

(2.) What is that frequency ω_E ?

Query (1.) has a Jeopardy-style answer-by-question: What is beam group velocity?

$$\frac{300}{900}$$

$$u_E = V_{group} = \frac{\omega_{group}}{k_{group}} = \frac{\omega_R - \omega_L}{k_R - k_L} = c \frac{\omega_R - \omega_L}{\omega_R + \omega_L} \quad V_{group} = c \frac{\omega_R - \omega_L}{\omega_R + \omega_L} = c \frac{600 - 300}{600 + 300} = \frac{1}{3} c$$

Query (2.) similarly: What ω_E is blue-shift $b\omega_L$ of ω_L and red-shift ω_R/b of ω_R ?

$$\sqrt{6 \cdot 3} = 3\sqrt{2} = 4.24$$

$$\omega_E = b\omega_L = \omega_R/b \Rightarrow b = \sqrt{\omega_R/\omega_L} \Rightarrow \omega_E = \sqrt{\omega_R \cdot \omega_L}$$

$$\omega_E = \sqrt{\omega_R \cdot \omega_L} = \sqrt{180000} = 424$$

V_{group}/c is ratio of difference mean $\omega_{group} = \frac{\omega_R - \omega_L}{2}$ to arithmetic mean $\omega_{phase} = \frac{\omega_R + \omega_L}{2}$. Frequency $\omega_E = B$ is the **geometric mean** $\sqrt{\omega_R \cdot \omega_L}$ of left and right-moving frequencies defining the geometry

Geometric mean

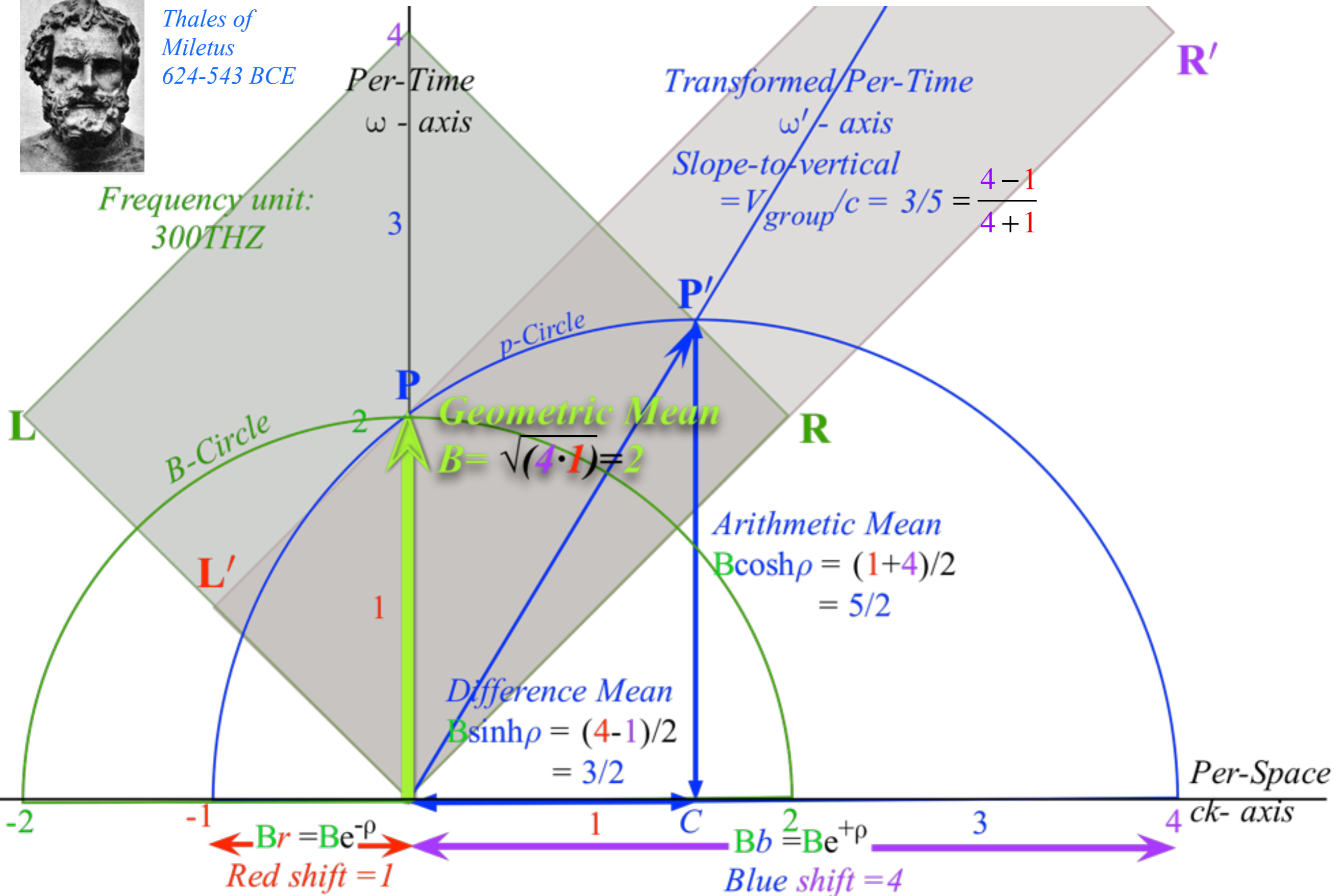
Thales Mean Geometry (600BCE)

helps "Relativity"



Thales of Miletus
624-543 BCE

Frequency unit:
300THZ



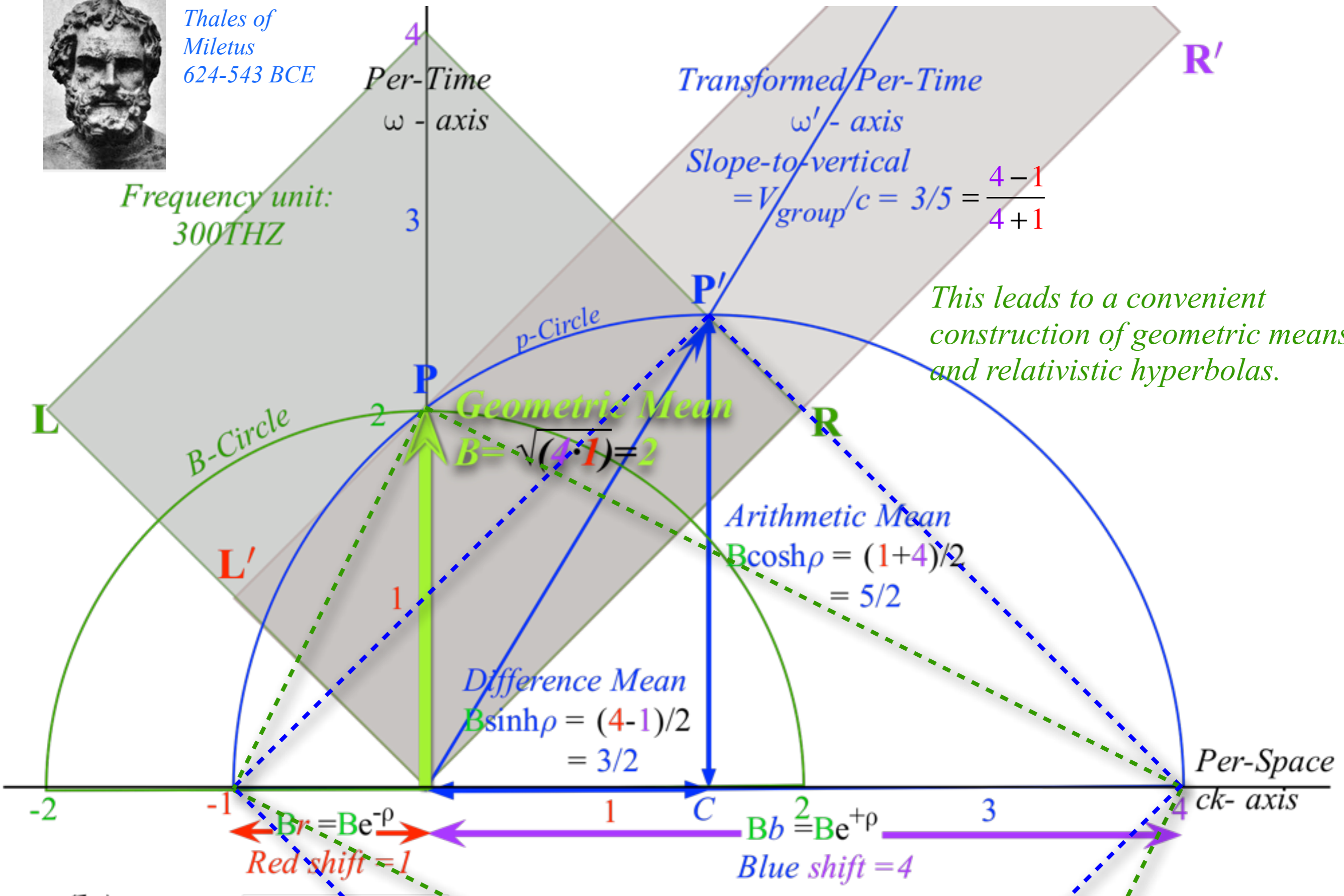
Thales Mean Geometry (600BCE)

helps “Relativity” *Thales showed a circle diameter subtends a right angle with any circle point P*



*Thales of Miletus
624-543 BCE*

*Frequency unit:
300THZ*



Thales Mean Geometry (600BCE)

helps “Relativity” *Thales showed a circle diameter subtends a right angle with any circle point P*



Thales of Miletus
624-543 BCE

Frequency unit:
300THZ

Per-Time
 ω - axis

Transformed/Per-Time
 ω' - axis

Slope-to-vertical

$$= V_{\text{group}}/c = 3/5 = \frac{4-1}{4+1}$$

equilateral hyperbola
 $r \cdot b = 2$

R'

This leads to a convenient construction of geometric means and relativistic hyperbolas.

L

B-Circle

Geometric Mean

$$B = \sqrt{(4 \cdot 1)} = 2$$

R

Arithmetic Mean

$$B \cosh \rho = (1+4)/2 = 5/2$$

Difference Mean

$$B \sinh \rho = (4-1)/2 = 3/2$$

Per-Space
ck- axis

-2

-1

1

2

3

4

$$Br = Be^{-\rho}$$

Red shift = 1

$$Bb = Be^{+\rho}$$

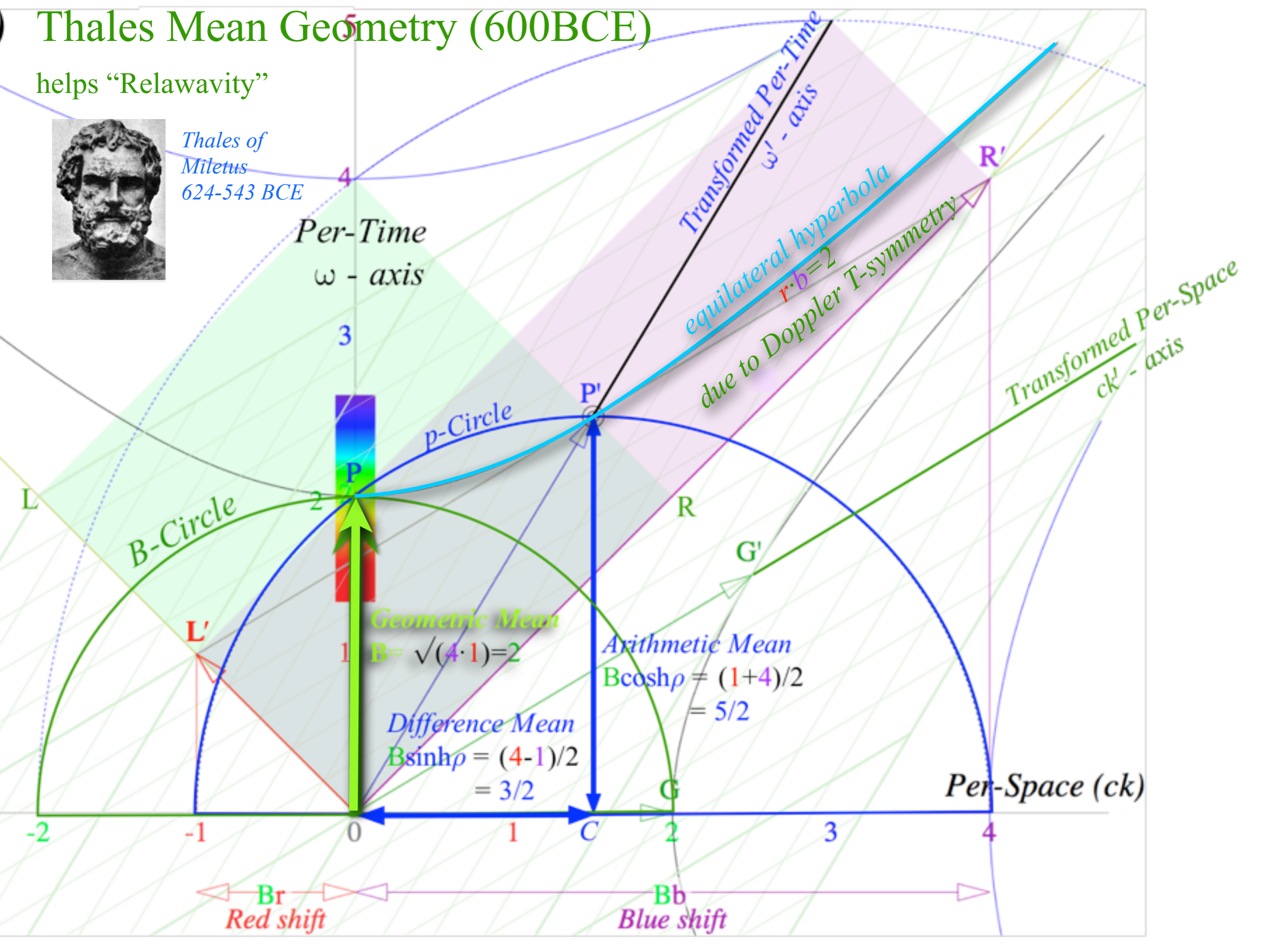
Blue shift = 4

Thales Mean Geometry (600BCE)

helps "Relativity"



Thales of Miletus
624-543 BCE



Per-Time (ω)

Acoustical base frequency = $B = 600\text{Hz}$

Hi freq = 1200.000

Lo freq = 300.000

Laser base frequency = $B = 600\text{THz}$

Doppler blue shift factor = $b = 2.000$

Doppler red shift factor = $r = 0.500$

$q = 0.693$

CW Light Axioms

All colors go c: $\omega/k = c$ or L&R on diagonals

Time Reversal ($r \leftrightarrow b$): $r = 1/b$

$G' = G \cosh(q) + P \sinh(q)$

$P' = G \sinh(q) + P \cosh(q)$

$G = G' \cosh(q) - P' \sinh(q)$

$P = -G' \sinh(q) + P' \cosh(q)$

H. sapiens Visual Best=600THz

600Hz = Auditory Base

Visual Min=400THz

20Hz = Auditory Min

-2.0 -1.5 -1.0 -0.5 0.0 0.5 1.0 1.5 2.0 Per-Space (ck)

[RelaWavity Web Simulation](#)

[Detailed Thales Geometry](#)

$B e^{-q}$
Red shift

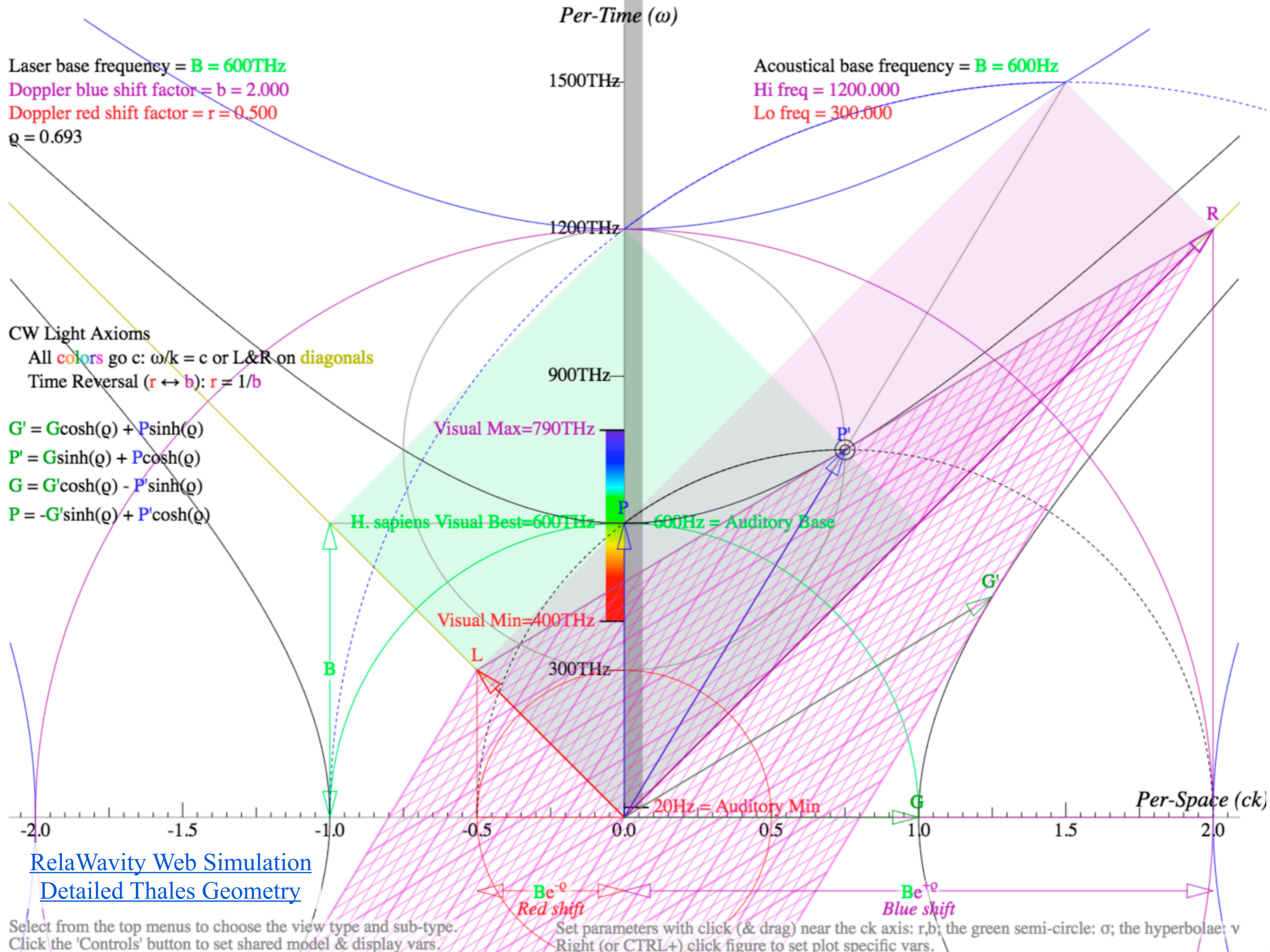
$B e^{+q}$
Blue shift

Select from the top menus to choose the view type and sub-type.

Click the 'Controls' button to set shared model & display vars.

Set parameters with click (& drag) near the ck axis: r,b; the green semi-circle: σ ; the hyperbolae: v

Right (or CTRL+) click figure to set plot specific vars.



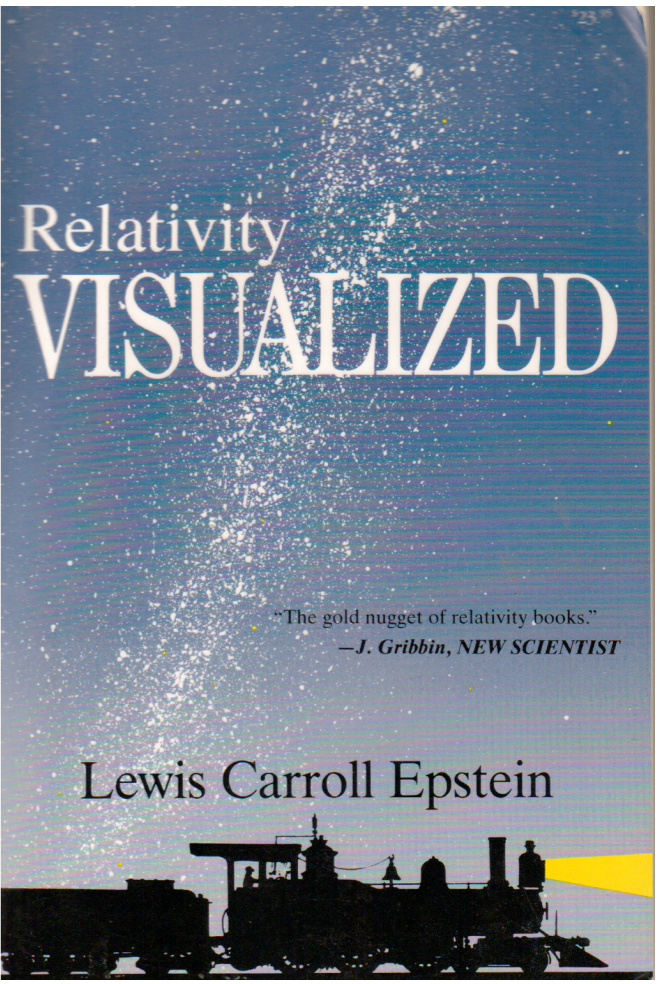
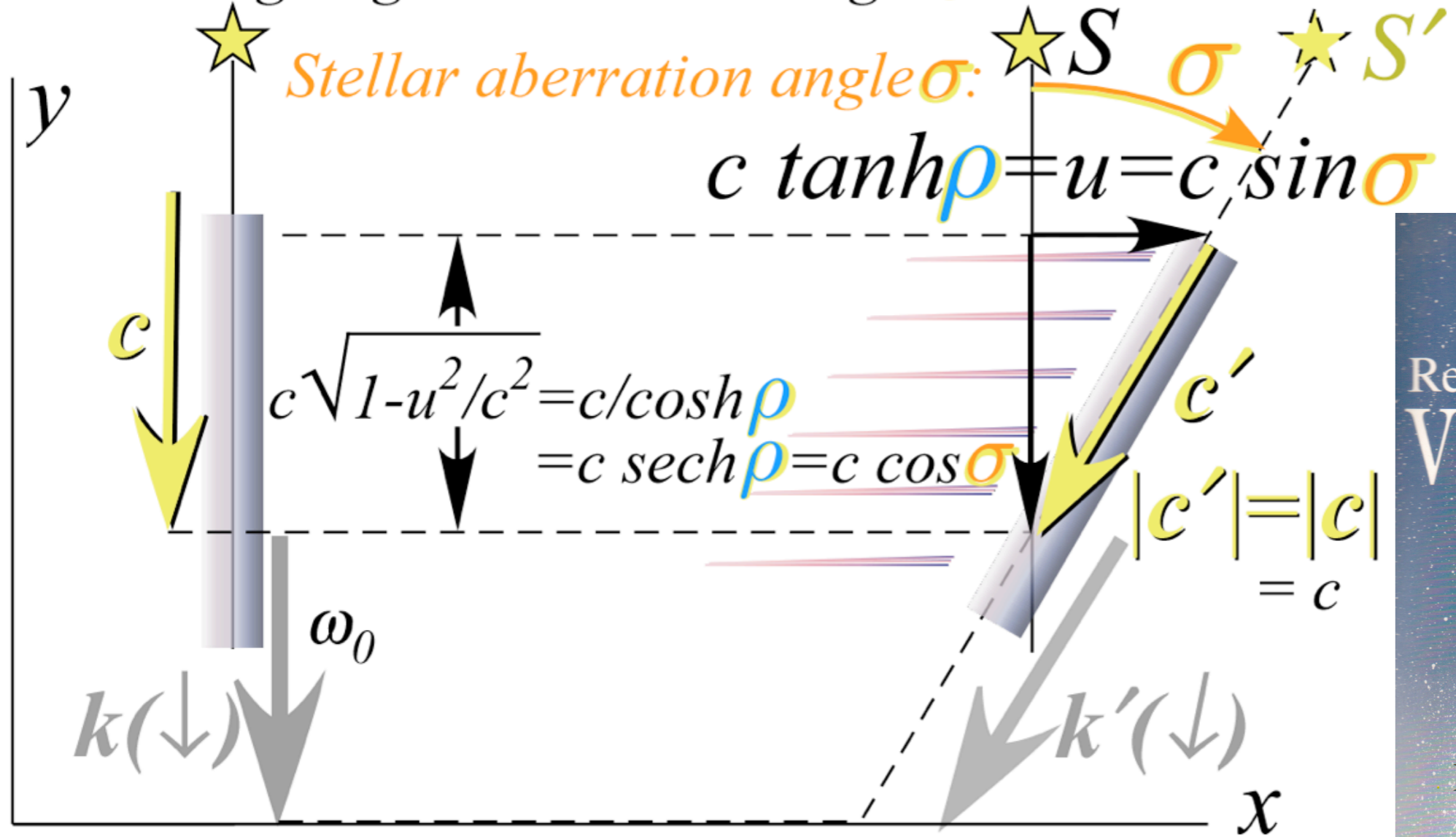
Comparing Longitudinal relativity parameter: Rapidity $\rho = \log_e(\text{Doppler Shift})$

to a Transverse*relativity parameter: Stellar aberration angle σ

*Lewis Carroll Epstein, *Relativitätstheorie*, Birkhäuser, (2004) Earlier English version (1985)-

Observer fixed below star sees it directly overhead.
 Observer going u sees star at angle σ in u direction.

We used notion σ for stellar-ab-angle, (a “flipped-out” ρ). Epstein not interested in ρ analysis or in relation of σ and ρ .



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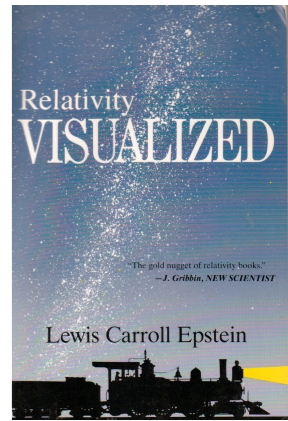
Purchase at:



Comparing Longitudinal relativity parameter: Rapidity $\rho = \log_e(\text{Doppler Shift})$

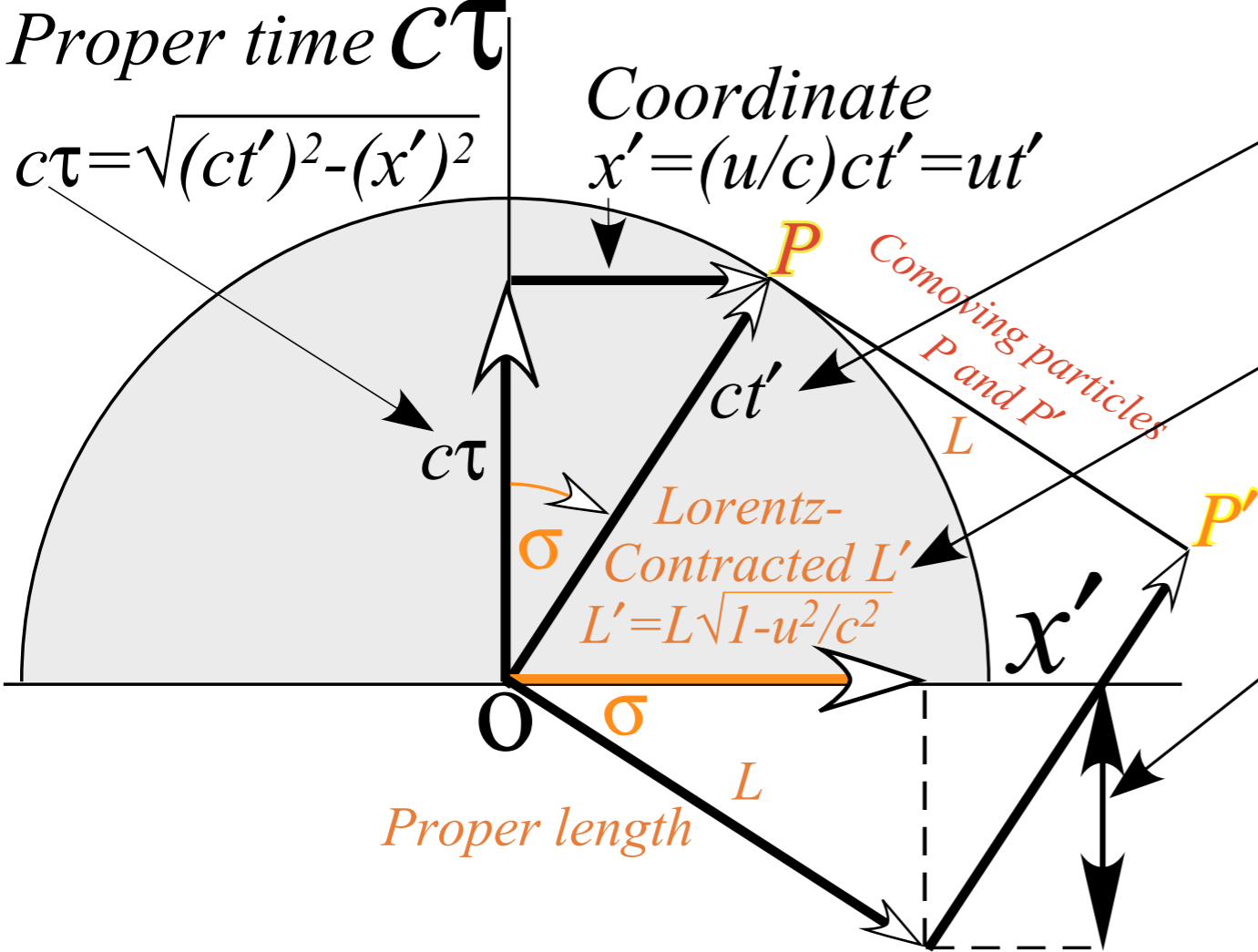
to a Transverse*relativity parameter: Stellar aberration angle σ

*Lewis Carroll Epstein, *Relativitätstheorie*, Birkhäuser, (2004) Earlier English version (1985)-



Proper time $c\tau$ vs. coordinate space x - (L. C. Epstein's "Cosmic Speedometer")

Particles P and P' have speed u in (x', ct') and speed c in $(x, c\tau)$



Einstein time dilation:
 $ct' = c\tau \sec\sigma = c\tau \cosh\rho = c\tau / \sqrt{1-u^2/c^2}$

Lorentz length contraction:
 $L' = L \operatorname{sech}\rho = L \cos\sigma = L \cdot \sqrt{1-u^2/c^2}$

Proper Time asimultaneity:
 $c \Delta\tau = L' \sinh\rho = L \cos\sigma \sinh\rho$
 $= L \cos\sigma \tan\sigma$
 $= L \sin\sigma = L / \sqrt{c^2/u^2 - 1} \sim L u/c$

This map has circle sector arc-area $\sigma = 0.6435$

set to angle $\angle\sigma = 36.87^\circ = 0.6435 \text{radian}$

$$\begin{aligned} \sin(\sigma) &= 0.6000 &= \tanh(\rho) &= 3/5 \\ \tan(\sigma) &= 0.7500 &= \sinh(\rho) &= 3/4 \\ \sec(\sigma) &= 1.2500 &= \cosh(\rho) &= 5/4 \\ \cos(\sigma) &= 0.8000 &= \operatorname{sech}(\rho) &= 4/5 \\ \cot(\sigma) &= 1.3333 &= \operatorname{csch}(\rho) &= 4/3 \\ \csc(\sigma) &= 1.6667 &= \operatorname{coth}(\rho) &= 5/3 \end{aligned}$$

$$\cosh(\rho) + \sinh(\rho) = \frac{5}{4} + \frac{3}{4} = 2.0 = e^{+\rho}$$

$$\cosh(\rho) - \sinh(\rho) = \frac{5}{4} - \frac{3}{4} = 1/2 = e^{-\rho}$$

$$\cosh(\rho) = \frac{e^{+\rho} + e^{-\rho}}{2} \quad \text{Half-Sum-}$$

Half-Difference

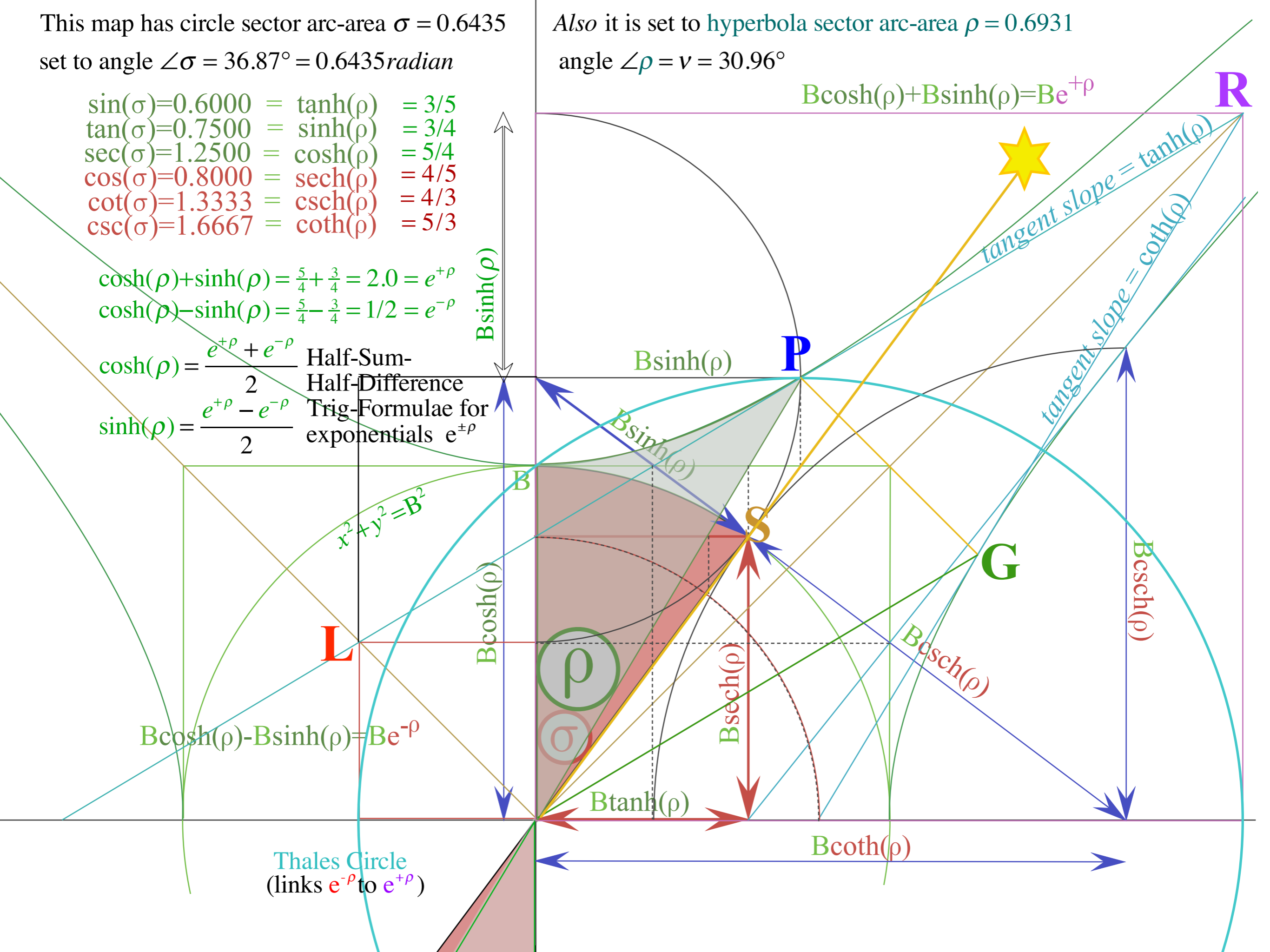
$$\sinh(\rho) = \frac{e^{+\rho} - e^{-\rho}}{2} \quad \text{Trig-Formulae for}$$

exponentials $e^{\pm\rho}$

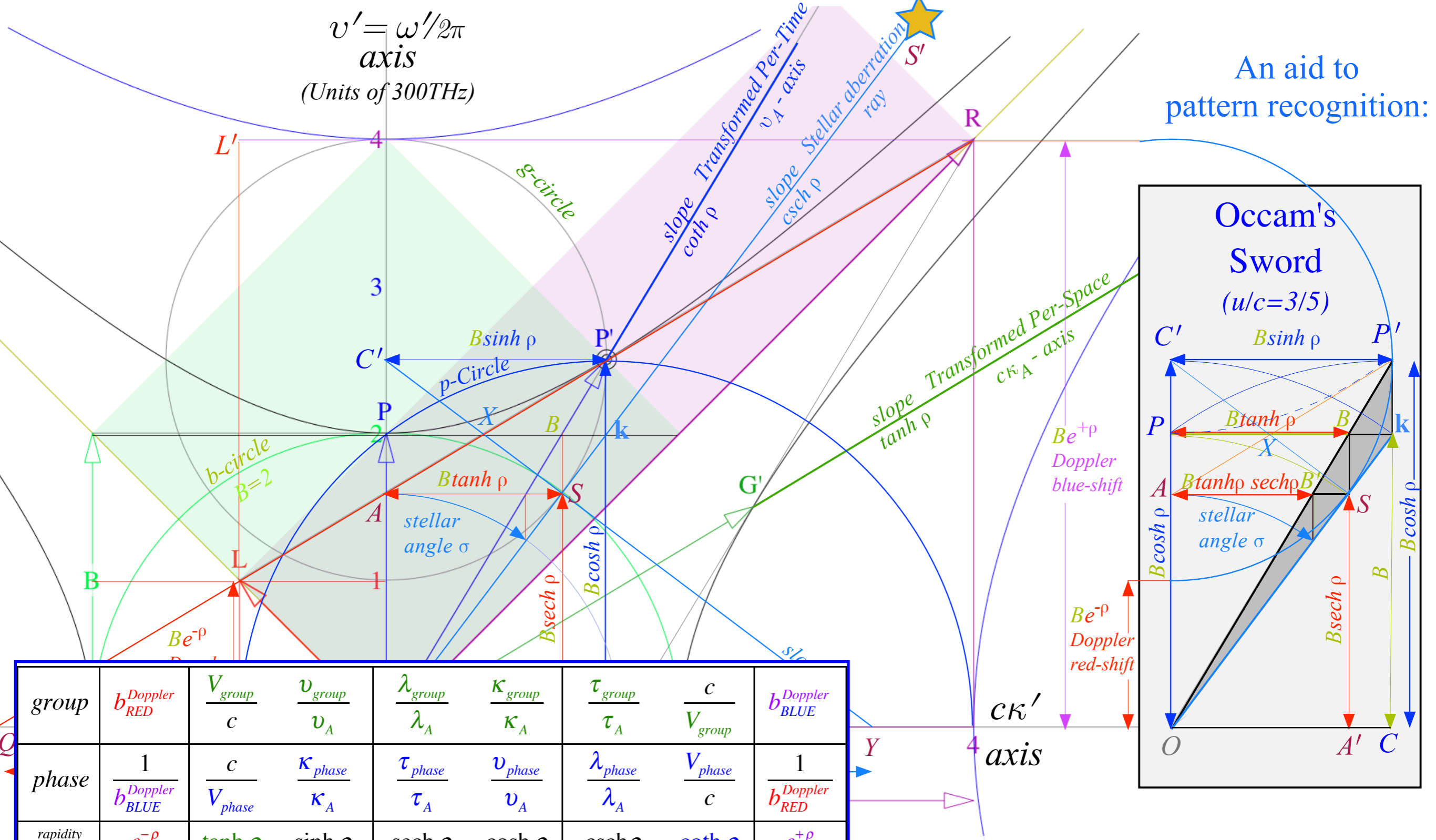
Also it is set to hyperbola sector arc-area $\rho = 0.6931$

angle $\angle\rho = \nu = 30.96^\circ$

$$B\cosh(\rho) + B\sinh(\rho) = B e^{+\rho}$$



Thales Circle
(links $e^{-\rho}$ to $e^{+\rho}$)



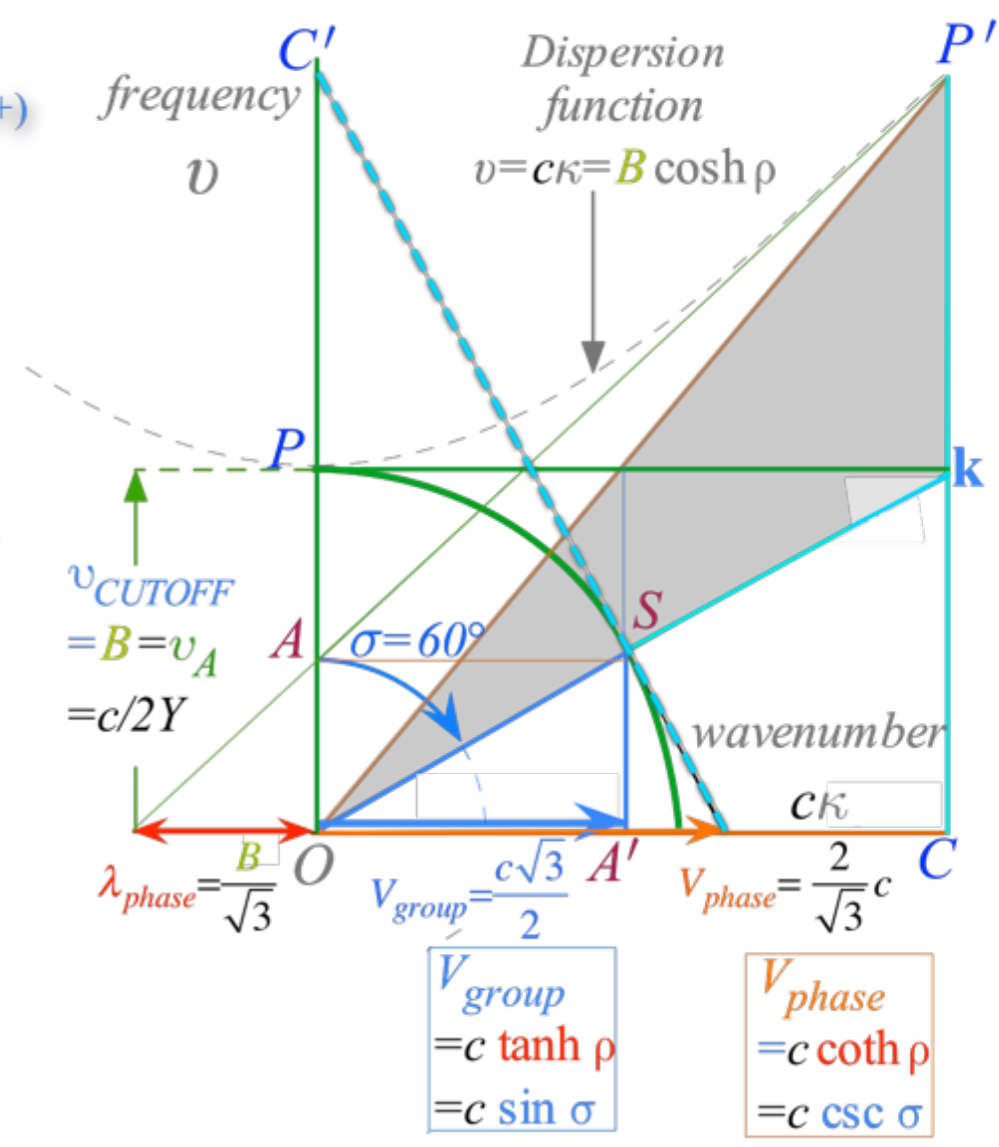
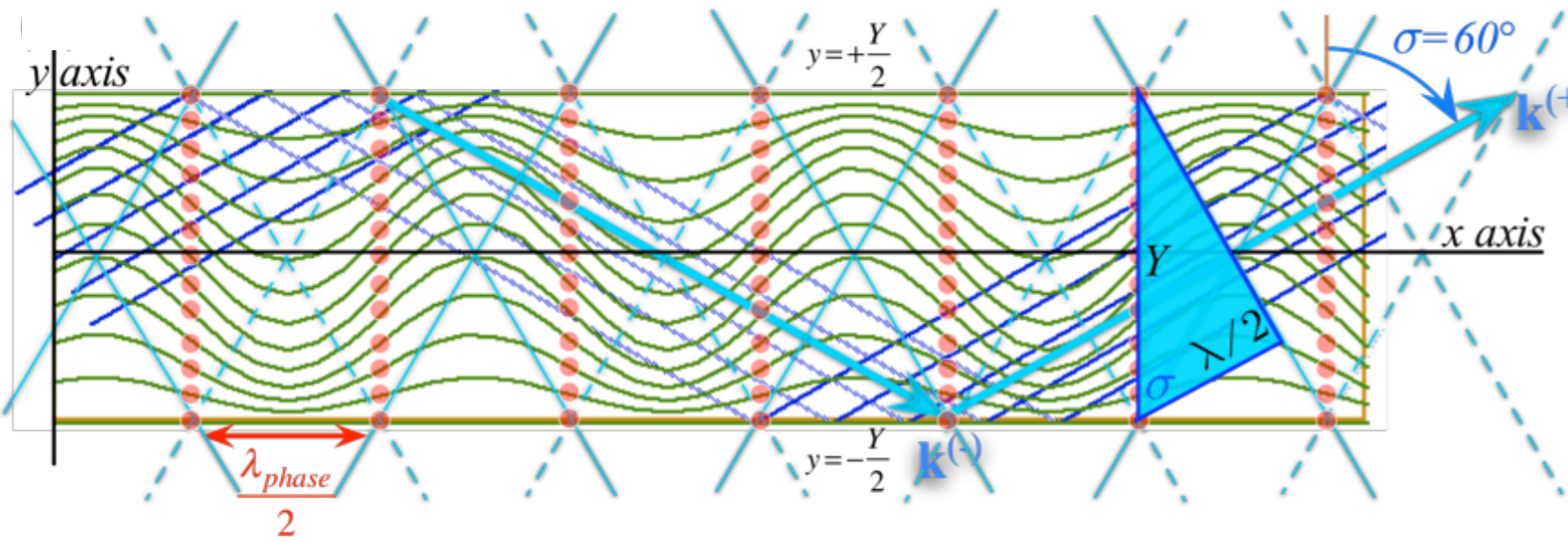
<i>group</i>	$b_{\text{Doppler RED}}$	$\frac{V_{\text{group}}}{c}$	$\frac{\nu_{\text{group}}}{\nu_A}$	$\frac{\lambda_{\text{group}}}{\lambda_A}$	$\frac{\kappa_{\text{group}}}{\kappa_A}$	$\frac{\tau_{\text{group}}}{\tau_A}$	$\frac{c}{V_{\text{group}}}$	$b_{\text{Doppler BLUE}}$
<i>phase</i>	$\frac{1}{b_{\text{Doppler BLUE}}}$	$\frac{c}{V_{\text{phase}}}$	$\frac{\kappa_{\text{phase}}}{\kappa_A}$	$\frac{\tau_{\text{phase}}}{\tau_A}$	$\frac{\nu_{\text{phase}}}{\nu_A}$	$\frac{\lambda_{\text{phase}}}{\lambda_A}$	$\frac{V_{\text{phase}}}{c}$	$\frac{1}{b_{\text{Doppler RED}}}$
<i>rapidity</i> ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\coth \rho$	$e^{+\rho}$
<i>stellar</i> ∇ <i>angle</i> σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
<i>value for</i> $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

Table of 12 wave parameters
(includes inverses) for relativity
...and values for $u/c=3/5$
[RelaWavity Web Simulation](#)
[Expanded Relativistic Relations](#)

Optical wave guide relativistic geometry aided by Occam's Sword

geometry applies to (x,y) space-space
 to (k_x, k_y) per-space-per-space
 to (x, ct) space-time

Relativistic mode with near-c $V_{group}=c/2$ and $V_{phase}=2c$. (Low dispersion.)

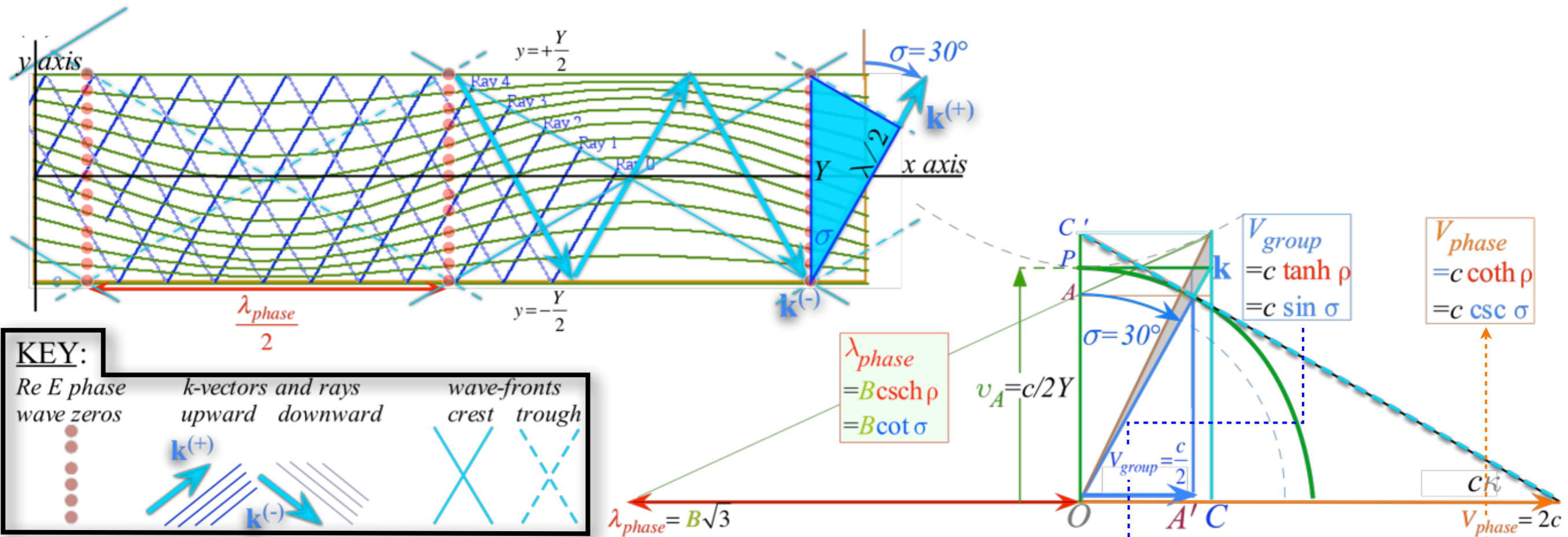


KEY:

<p><i>Re E phase</i></p> <p>wave zeros</p>	<p><i>k-vectors and rays</i></p> <p>upward downward</p>	<p><i>wave-fronts</i></p> <p>crest trough</p>
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Optical wave guide relativistic geometry aided by Occam's Sword

geometry applies to (x,y) space-space
 to (k_x, k_y) per-space-per-space
 to (x, ct) space-time

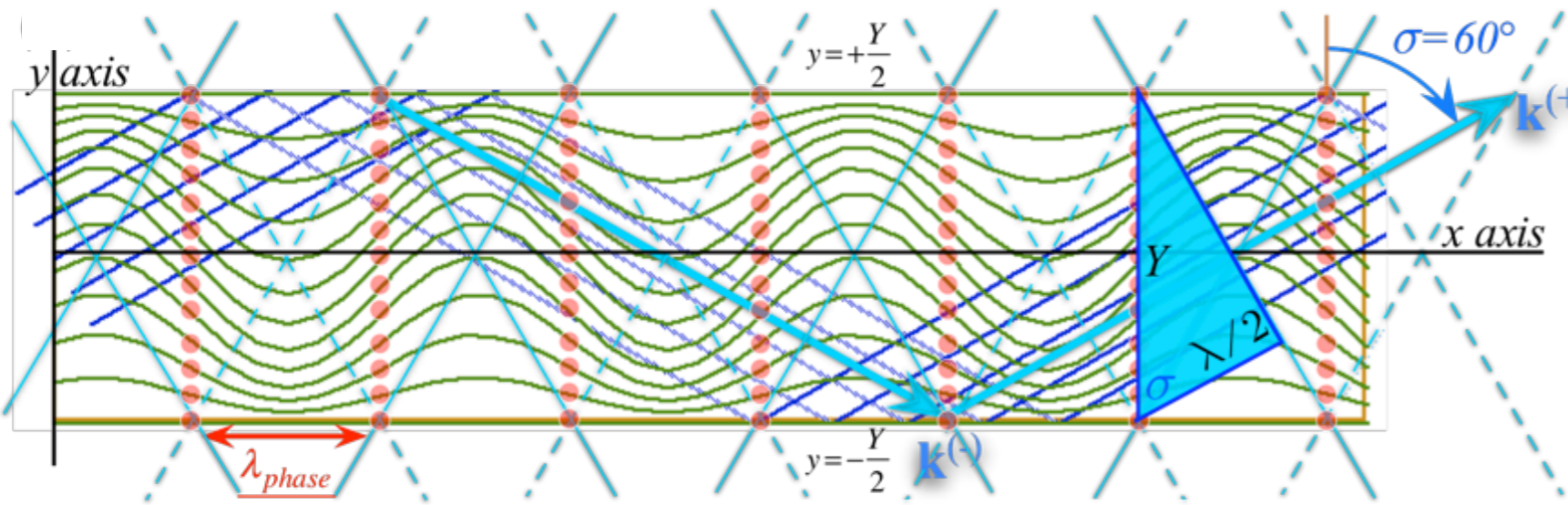


Example of near-cut-off mode with low $V_{group} = c/2$ and high $V_{phase} = 2c$. (High dispersion.)

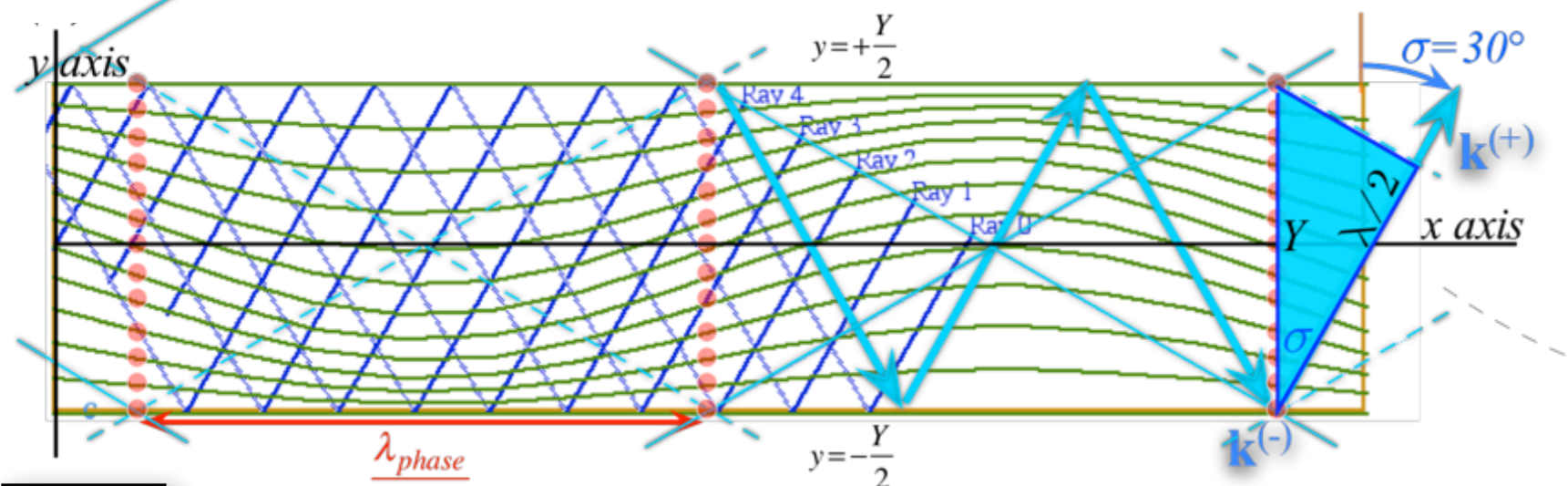
Optical wave guide relativistic geometry aided by Occam's Sword

geometry applies to (x,y) space-space
to (k_x, k_y) per-space-per-space
to (x, ct) space-time

Relativistic mode with near-c $V_{group}=c/2$ and $V_{phase}=2c$. (Low dispersion.)



GuidIt Web Simulation: $\sigma = 60^\circ$

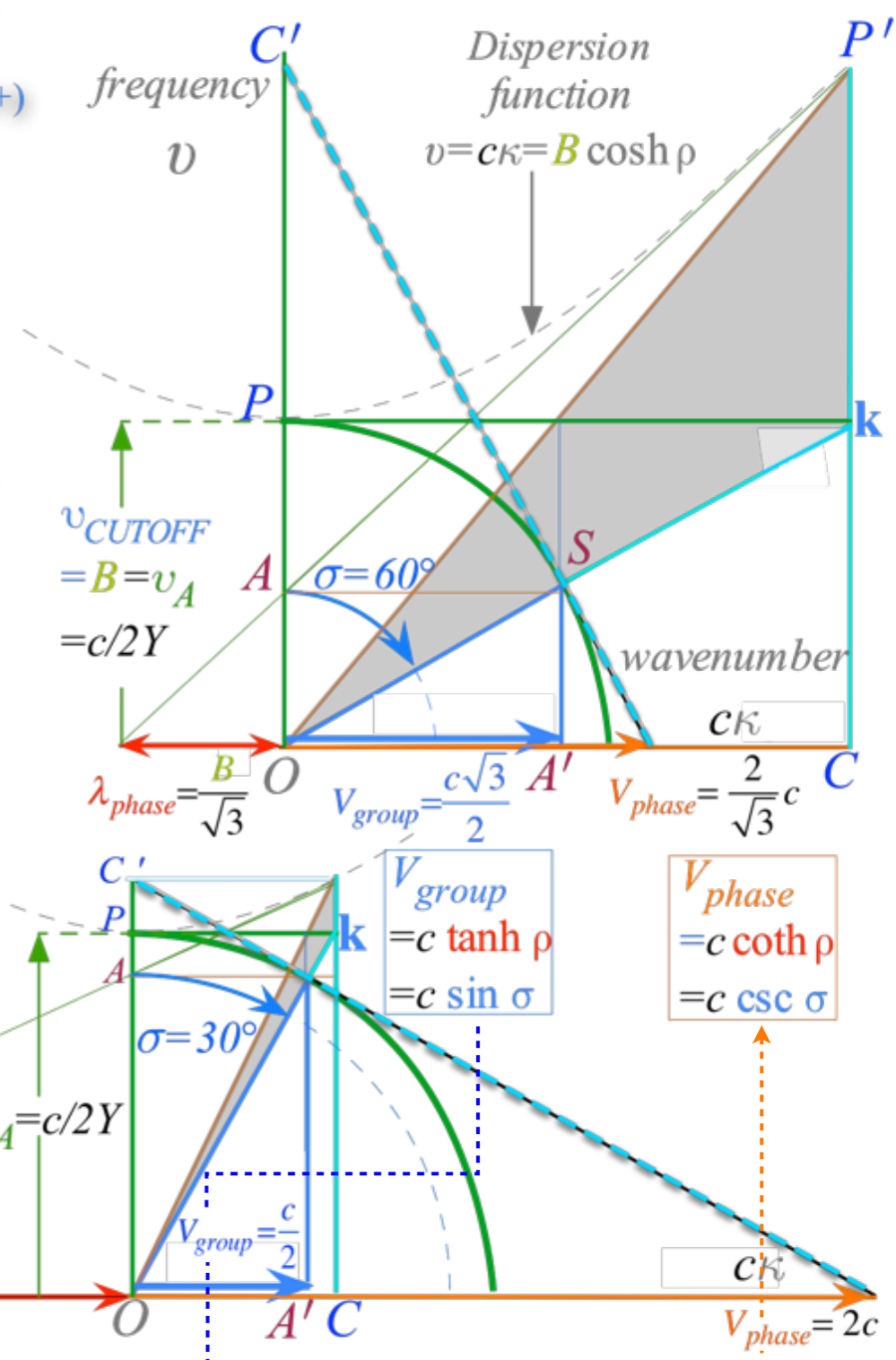


GuidIt Web Simulation: $\sigma = 30^\circ$

KEY:

Re E phase wave zeros	k-vectors and rays upward downward	wave-fronts crest trough

$k^{(+)}$
 $k^{(-)}$



$\lambda_{phase} = B\sqrt{3}$

$\lambda_{phase} = B \csc \rho$
 $= B \cot \sigma$

$v_A = c/2Y$

$V_{group} = c \tanh \rho$
 $= c \sin \sigma$

$V_{phase} = c \coth \rho$
 $= c \csc \sigma$

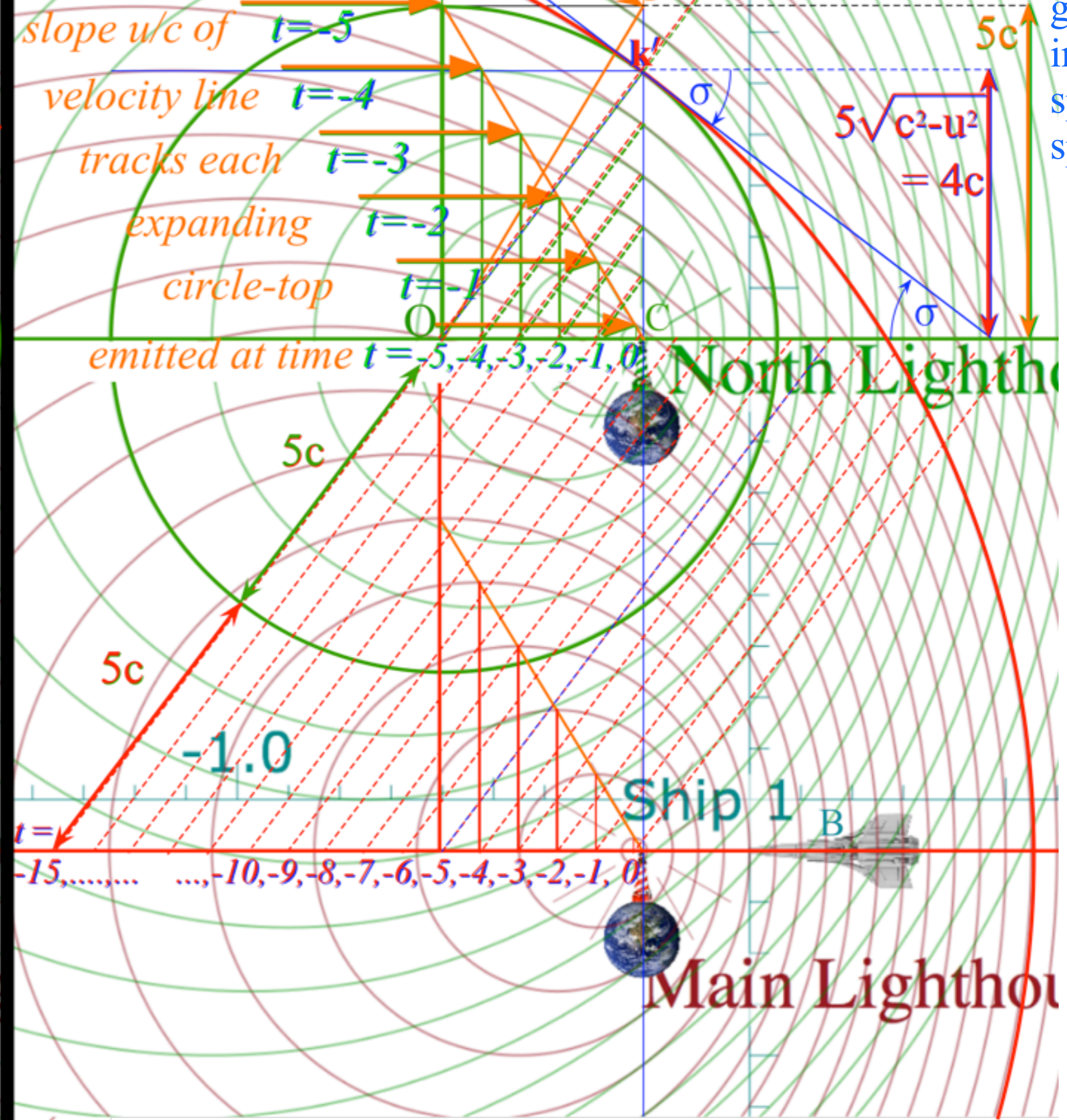
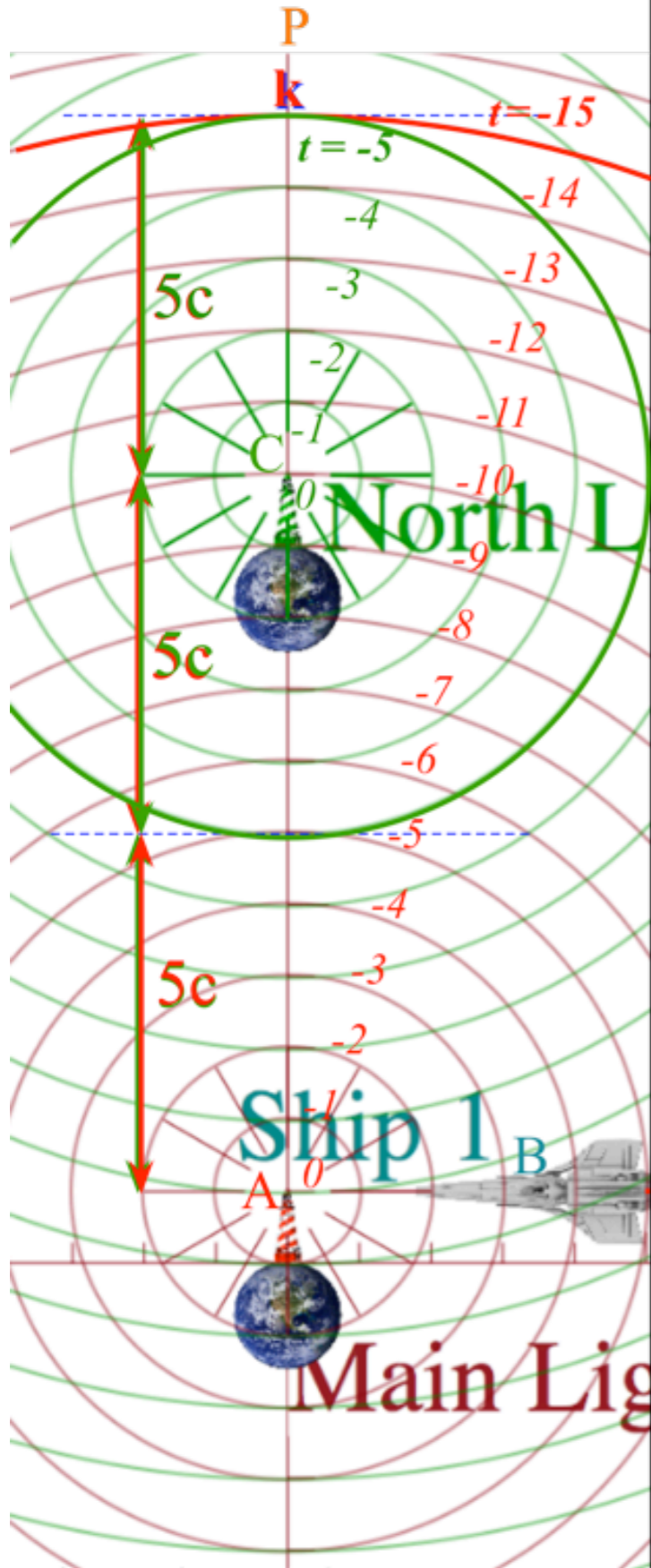
Example of near-cut-off mode with low $V_{group}=c/2$ and high $V_{phase}=2c$. (High dispersion.)

(a) Spherical wave pair
In Alice-Carla frame

stellar angle $\sigma = \sin^{-1}(u/c)$
velocity angle $v = \tan^{-1}(u/c)$
slope u/c of $t = -5$
velocity line $t = -4$
tracks each $t = -3$
expanding $t = -2$
circle-top $t = -1$
emitted at time $t = -5, -4, -3, -2, -1, 0$

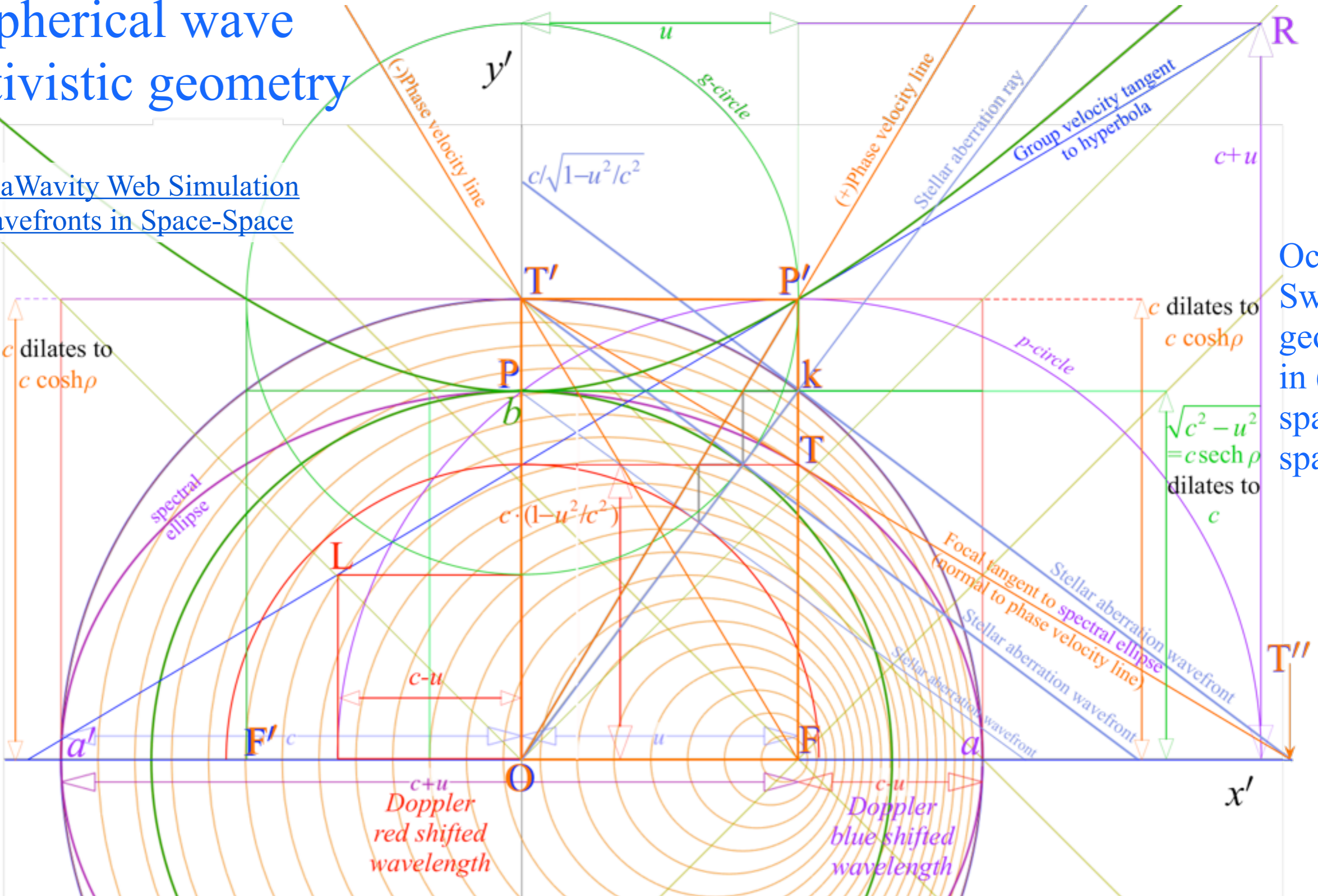
(b) Spherical wave pair
In Bob's frame: $u_x/c = -3/5$

Occam
Sword
geometry
in (x,y)
space-
space



Spherical wave relativistic geometry

RelaWavity Web Simulation
Wavefronts in Space-Space



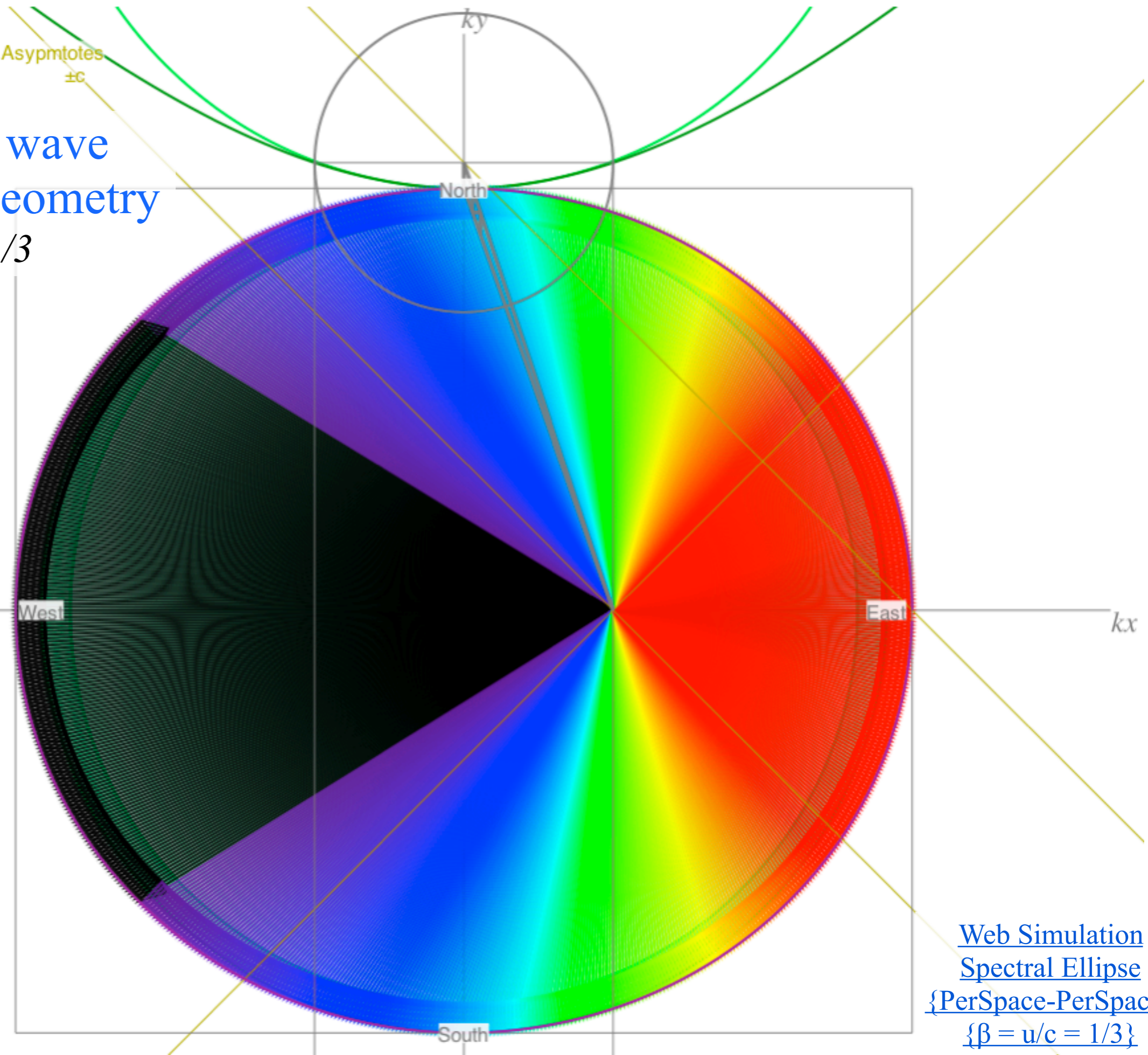
Occam
Sword
geometry
in (x,y)
space-space

<p>Doppler Red $\lambda=c+u$ dilates to: $(c+u) \cosh \rho = c \sqrt{\frac{c+u}{c-u}} = ce^{+\rho}$</p>	<p>ellipse focal length $FO = u = c \tanh \rho$ dilates to: $u \cosh \rho = c \sinh \rho$</p>	<p>Doppler Blue $\lambda=c-u$ dilates to: $(c-u) \cosh \rho = c \sqrt{\frac{c-u}{c+u}} = ce^{-\rho}$</p>
<p>ellipse major radius $a = OFa = c$ dilates to: $c \cosh \rho = c / \sqrt{1-u^2/c^2}$</p>	<p>ellipse latus radius $FT = c(1-u^2/c^2)$ dilates to: $c(1-u^2/c^2) \cosh \rho = c \sqrt{1-u^2/c^2} = c \operatorname{sech} \rho$</p>	<p>Base height $FTk = \sqrt{c^2 - u^2}$ dilates to: $\sqrt{c^2 - u^2} \cosh \rho = c$ (equal to ellipse minor radius b)</p>

Applications of Einstein dilation factor:
 $\gamma = \cosh \rho = 1 / \sqrt{1-u^2/c^2}$

Spherical wave
relativistic geometry

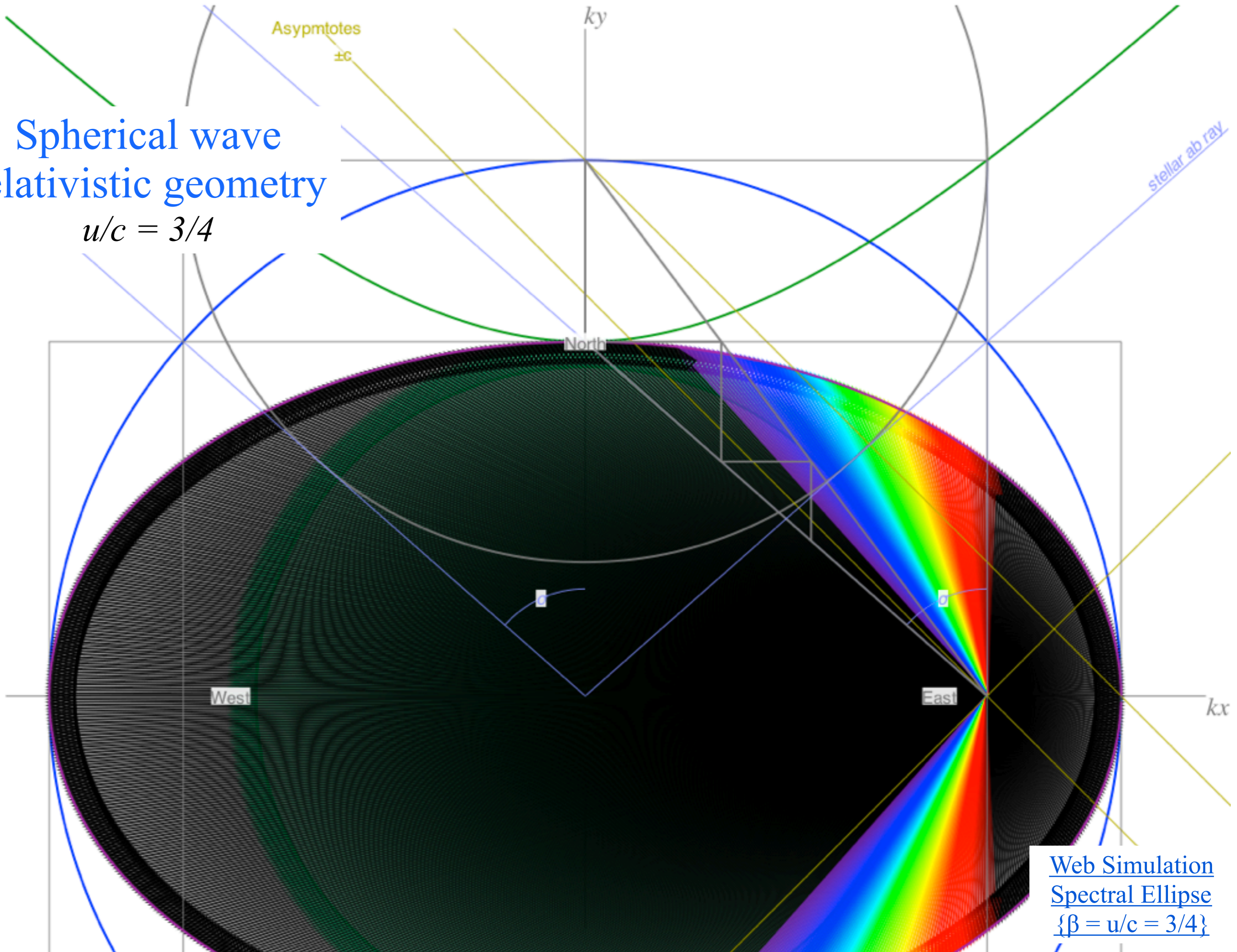
$$u/c = 1/3$$



[Web Simulation](#)
[Spectral Ellipse](#)
{PerSpace-PerSpace}
{ $\beta = u/c = 1/3$ }

Spherical wave
relativistic geometry

$$u/c = 3/4$$



Web Simulation
Spectral Ellipse
{ $\beta = u/c = 3/4$ }

Using (some) wave parameters to develop relativistic quantum theory ...and classical mechanics

$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \text{ (for } u \ll c)$$

$$cK_{phase} = B \sinh \rho \approx B \rho \text{ (for } u \ll c)$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2$$

$$\sinh \rho \approx \rho$$

$$B = v_A$$

$$B = v_A = cK_A$$

At low speeds: ...

<i>group</i>	$b_{Doppler RED}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler BLUE}$
<i>phase</i>	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
<i>rapidity</i> ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
<i>stellar angle</i> σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
<i>value for</i> $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

[RelaWavity Web Simulation - Relativistic Terms](#)
(Expanded Table)

Using (some) wave parameters to develop relativistic quantum theory ...and classical mechanics

$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \quad (\text{for } u \ll c)$$

$$c\kappa_{phase} = B \sinh \rho \approx B \rho \quad (\text{for } u \ll c)$$

$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

At low speeds:

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

$$\sinh \rho \approx \rho \approx \frac{u}{c}$$

$$B = v_A$$

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group	$b_{Doppler RED}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler BLUE}$
phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
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At low speeds: \Leftarrow for $(u \ll c)$

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2$$

group	$b_{Doppler RED}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler BLUE}$
phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
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At low speeds: \Leftarrow for $(u \ll c) \Rightarrow$

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2 \qquad K_{phase} \approx \frac{B}{c^2} u$$

time	$b_{Doppler RED}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler BLUE}$
space	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
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$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2$$

$$K_{phase} \approx \frac{B}{c^2} u$$

v_{phase} and K_{phase} resemble formulae for Newton's kinetic energy and momentum

Resembles: $const. + \frac{1}{2} Mu^2$

Resembles: Mu

group	$b_{Doppler RED}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler BLUE}$
phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
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At low speeds:

⇐ for ($u \ll c$) ⇒

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v_{phase} and K_{phase} resemble formulae for Newton's kinetic energy $\frac{1}{2} Mu^2$ and momentum Mu .

Rescale v_{phase} by h so: $M = \frac{hB}{c^2}$

Resembles: $const. + \frac{1}{2} Mu^2$

Resembles: Mu

time	$b_{Doppler RED}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler BLUE}$
space	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
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v_{phase} and K_{phase} resemble formulae for Newton's kinetic energy $\frac{1}{2} Mu^2$ and momentum Mu .

Rescale v_{phase} by h so: $M = \frac{hB}{c^2}$

So attach scale factor h to match units.

Resembles: $const. + \frac{1}{2} Mu^2$

Resembles: Mu

group	$b_{Doppler RED}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler BLUE}$
phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

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Rescale v_{phase} by h so: $M = \frac{hB}{c^2}$

$$hv_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2 \quad \Leftarrow \text{for } (u \ll c) \Rightarrow$$

$$hK_{phase} \approx \frac{hB}{c^2} u$$

So attach scale factor h to match units.

Resembles: $const. + \frac{1}{2} Mu^2$

Resembles: Mu

group	$b_{Doppler RED}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler BLUE}$
phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

Using (some) wave parameters to develop relativistic quantum theory ...and classical mechanics

$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \quad (\text{for } u \ll c)$$

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$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

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$$B = v_A$$

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\Leftarrow for $(u \ll c) \Rightarrow$

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Rescale v_{phase} by h so: $M = \frac{hB}{c^2}$

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So attach scale factor h to match units.

Resembles: $const. + \frac{1}{2} Mu^2$

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phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
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$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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[RelaWavity Web Simulation - Relativistic Terms](#)
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phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
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$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
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(old-fashioned notation)

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Planck (1900)

$$= \text{Total Energy: } E = \frac{Mc^2}{\sqrt{1-u^2/c^2}}$$

Einstein (1905)

(old-fashioned notation)

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phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\text{coth } \rho$	$e^{+\rho}$
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Max Planck
1858-1947

Using (some) wave parameters to develop relativistic quantum theory ...and classical mechanics

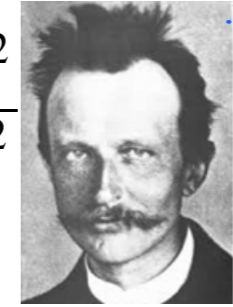
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So attach scale factor h (or hN) to match units.

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Lucky coincidences?? Cheap trick??
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Need to replace h with hN to match e.m. energy density
 $\epsilon_0 E^* \cdot E = hN v_{phase}$

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phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$
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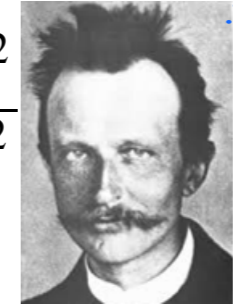
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Quantized amplitude

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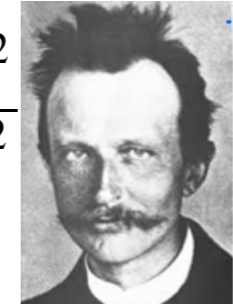
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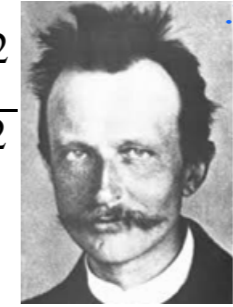
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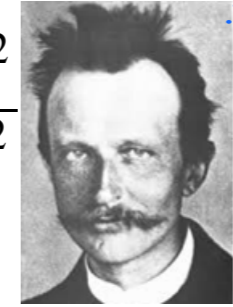
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Resolution and dirty secret: \mathbf{E} , N , and v_{phase} are all frequencies!

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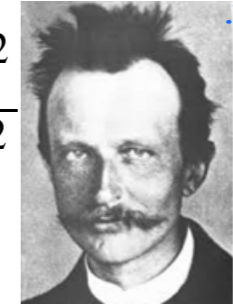
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Rescale v_{phase} by h so: $M = \frac{hB}{c^2}$ or: $hB = Mc^2$

(The famous Mc^2 shows up here!)

$$hv_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2$$

for $(u \ll c) \Rightarrow$

$$hK_{phase} \approx \frac{hB}{c^2} u$$

So attach scale factor h (or hN) to match units.

$$hv_{phase} \approx Mc^2 + \frac{1}{2} Mu^2$$

for $(u \ll c) \Rightarrow$

$$hK_{phase} \approx Mu$$

~~Natural wave conspiracy~~
~~Lucky coincidences??~~ ~~Expensive Cheap trick??~~
...Try exact v_{phase} and K_{phase} ...

Need to replace h with hN to match e.m. energy density $\epsilon_0 \mathbf{E} \cdot \mathbf{E} = hN v_{phase}$

$$hv_{phase} = hB \cosh \rho = Mc^2 \cosh \rho$$

Planck (1900)

$$= \text{Total Energy: } E = \frac{Mc^2}{\sqrt{1-u^2/c^2}}$$

Einstein (1905)

This motivates the "particle" normalization $\int \Psi^* \Psi dV = N$ $\Psi = \sqrt{\frac{\epsilon_0}{h\nu}} E$

$$hK_{phase} = hB \sinh \rho = Mc^2 \sinh \rho$$

group	$b_{\text{Doppler RED}}$	$\frac{V_{\text{group}}}{c}$	$\frac{v_{\text{group}}}{v_A}$	$\frac{\lambda_{\text{group}}}{\lambda_A}$	$\frac{K_{\text{group}}}{K_A}$	$\frac{\tau_{\text{group}}}{\tau_A}$	$\frac{V_{\text{phase}}}{c}$
phase	$b_{\text{Doppler BLUE}}$	$\frac{c}{V_{\text{phase}}}$	$\frac{K_{\text{phase}}}{K_A}$	$\frac{\tau_{\text{phase}}}{\tau_A}$	$\frac{v_{\text{phase}}}{v_A}$		
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$		
stellar angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$

$$\frac{1}{\sqrt{\beta^2-1}} = \frac{\frac{u}{c}}{\sqrt{1-\frac{u^2}{c^2}}} \quad (\text{old-fashioned notation})$$

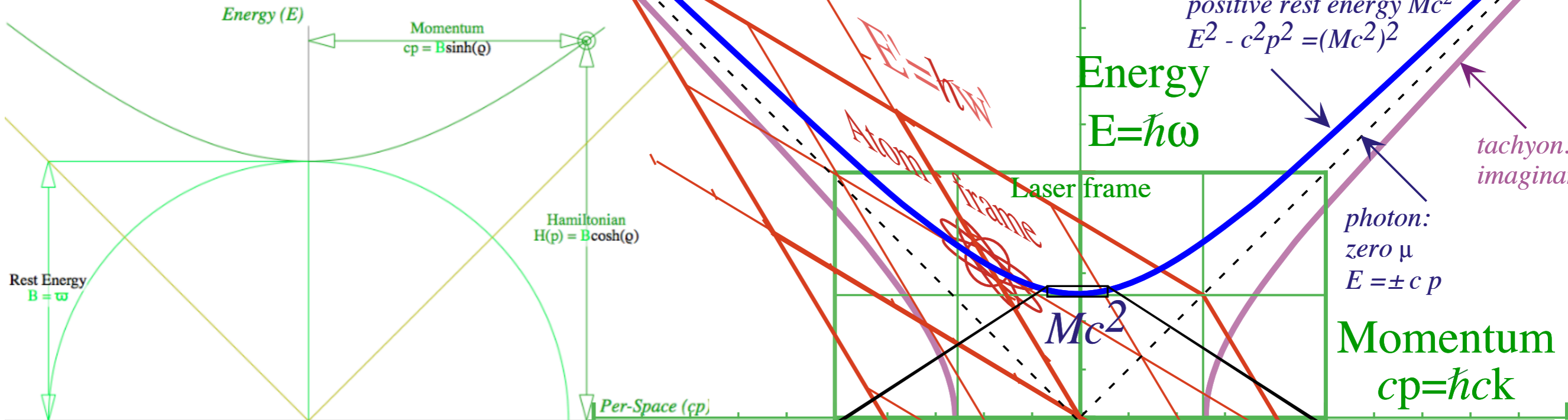
$$cp = \frac{Muc}{\sqrt{1-u^2/c^2}}$$

Momentum: $hK_{phase} = p = \frac{Mu}{\sqrt{1-u^2/c^2}}$

DeBroglie (1921)

Using (some) wave coordinates for relativistic quantum theory

(a) Exact Einstein-Planck Dispersion



Mass (resting)

$$hB = \hbar \omega_A = Mc^2 = \hbar c \kappa_A$$

Energy

$$\hbar \omega_{phase} = E = \hbar \omega_A \cosh \rho$$

Momentum

$$\hbar c \kappa_{phase} = cp = \hbar c \kappa_A \sinh \rho = \hbar \omega_A \sinh \rho$$

Energy versus Momentum

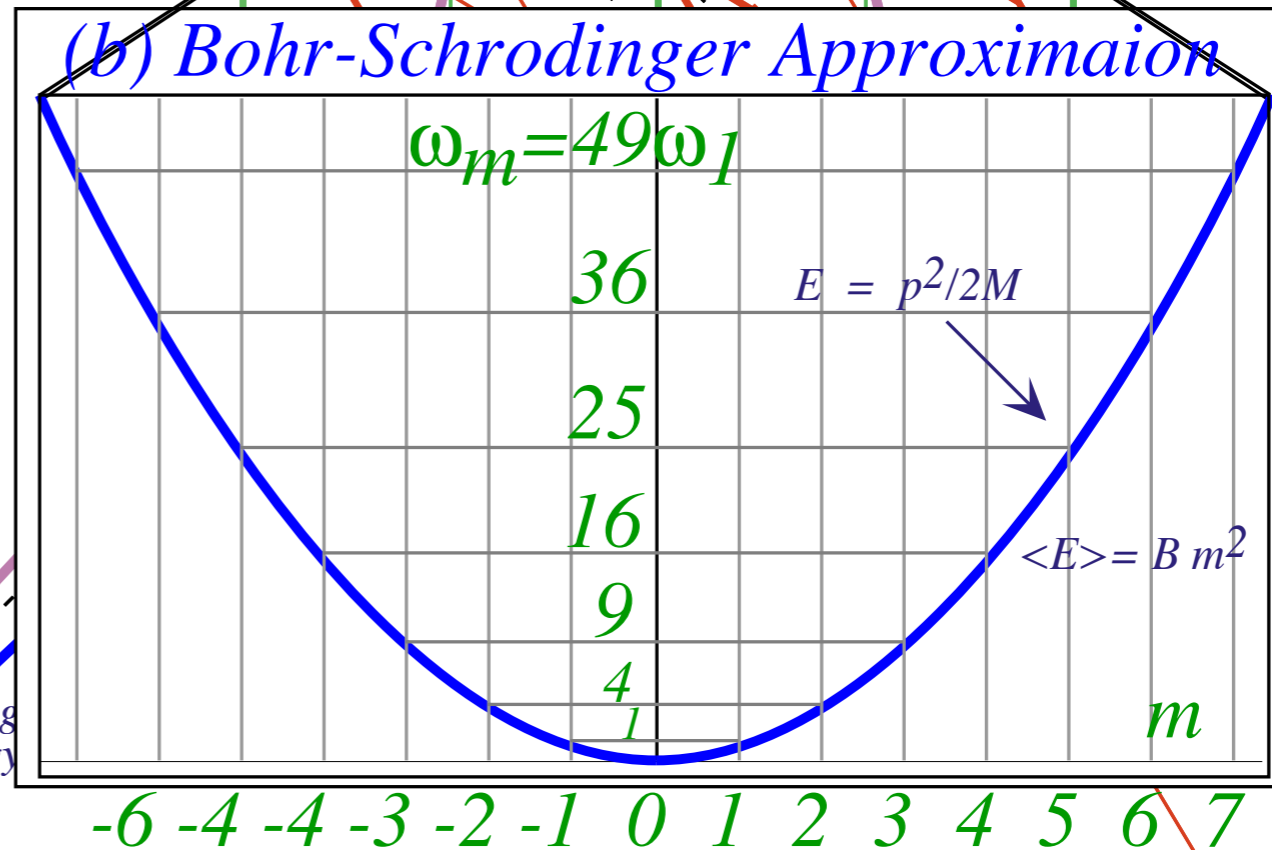
$$E^2 = (Mc^2)^2 \cosh^2 \rho$$

$$= (Mc^2)^2 (1 + \sinh^2 \rho) = (Mc^2)^2 + (cp)^2 \Rightarrow E = \pm \sqrt{(Mc^2)^2 + (cp)^2} \approx Mc^2 + \frac{p^2}{2M}$$

(b) Bohr-Schrodinger Approximation



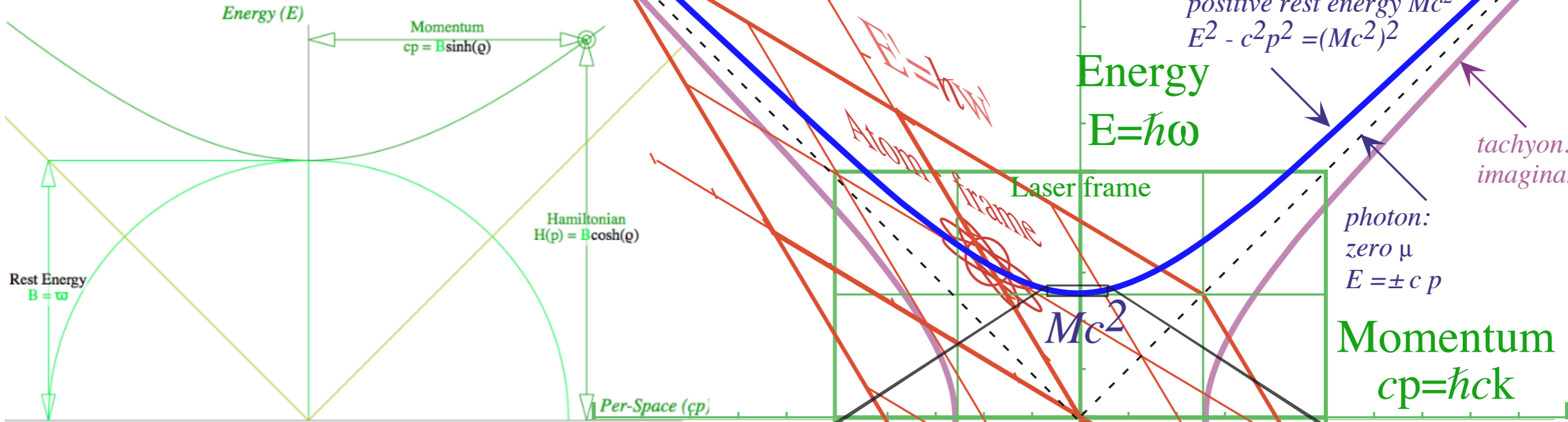
Niels Bohr
1885-1962



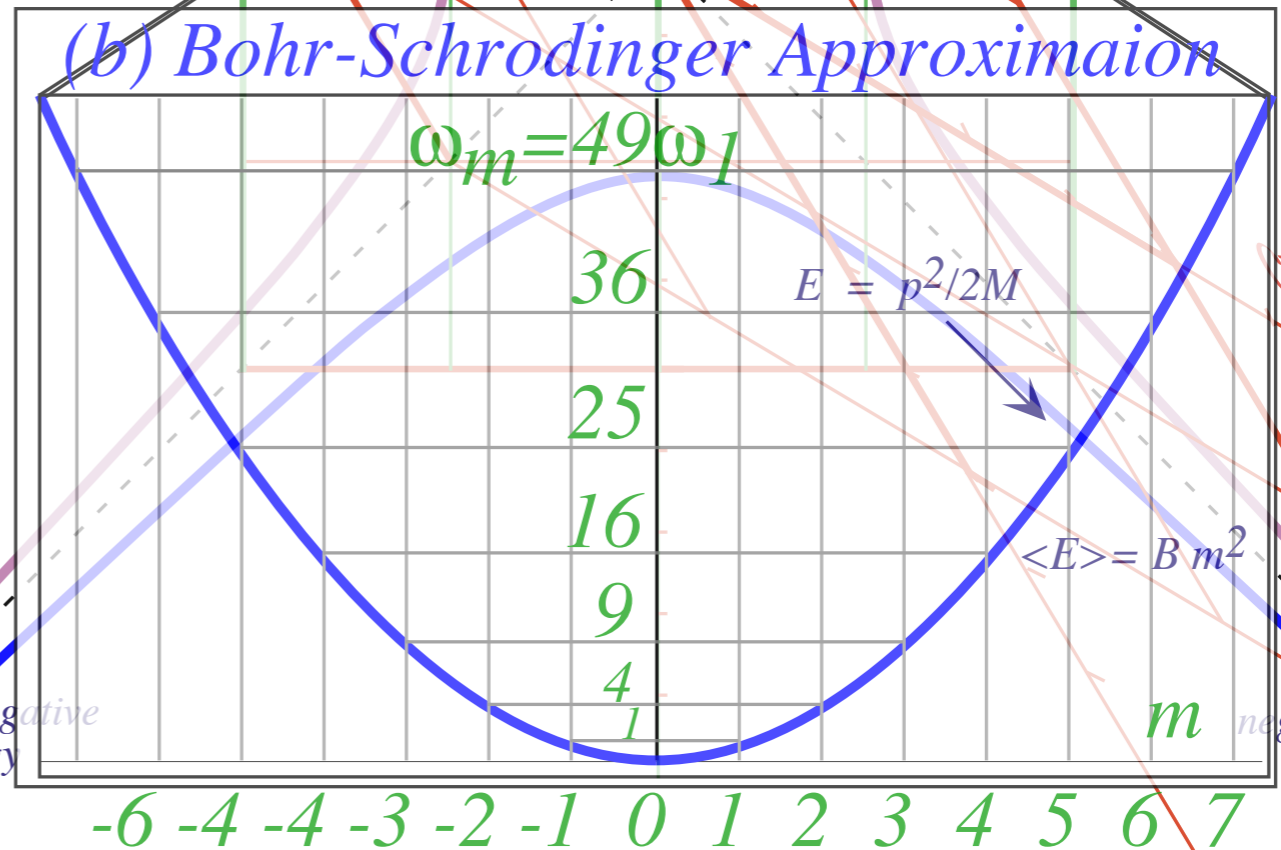
Erwin Schrodinger
1887-1961

Using (some) wave coordinates for relativistic quantum theory

(a) Exact Einstein-Planck Dispersion



(b) Bohr-Schrodinger Approximation



Mass (resting)

$$hB = h\nu_A = Mc^2 = hc\kappa_A$$

Energy

$$h\nu_{\text{phase}} = E = h\nu_A \cosh \rho$$

Momentum

$$hc\kappa_{\text{phase}} = cp = hc\kappa_A \sinh \rho = h\nu_A \sinh \rho$$

Energy versus Momentum

$$E^2 = (Mc^2)^2 \cosh^2 \rho$$

$$= (Mc^2)^2 (1 + \sinh^2 \rho) = (Mc^2)^2 + (cp)^2 \Rightarrow E = \pm \sqrt{(Mc^2)^2 + (cp)^2} \approx Mc^2 + \frac{p^2}{2M}$$

low speed approximation

Relativity variable tables

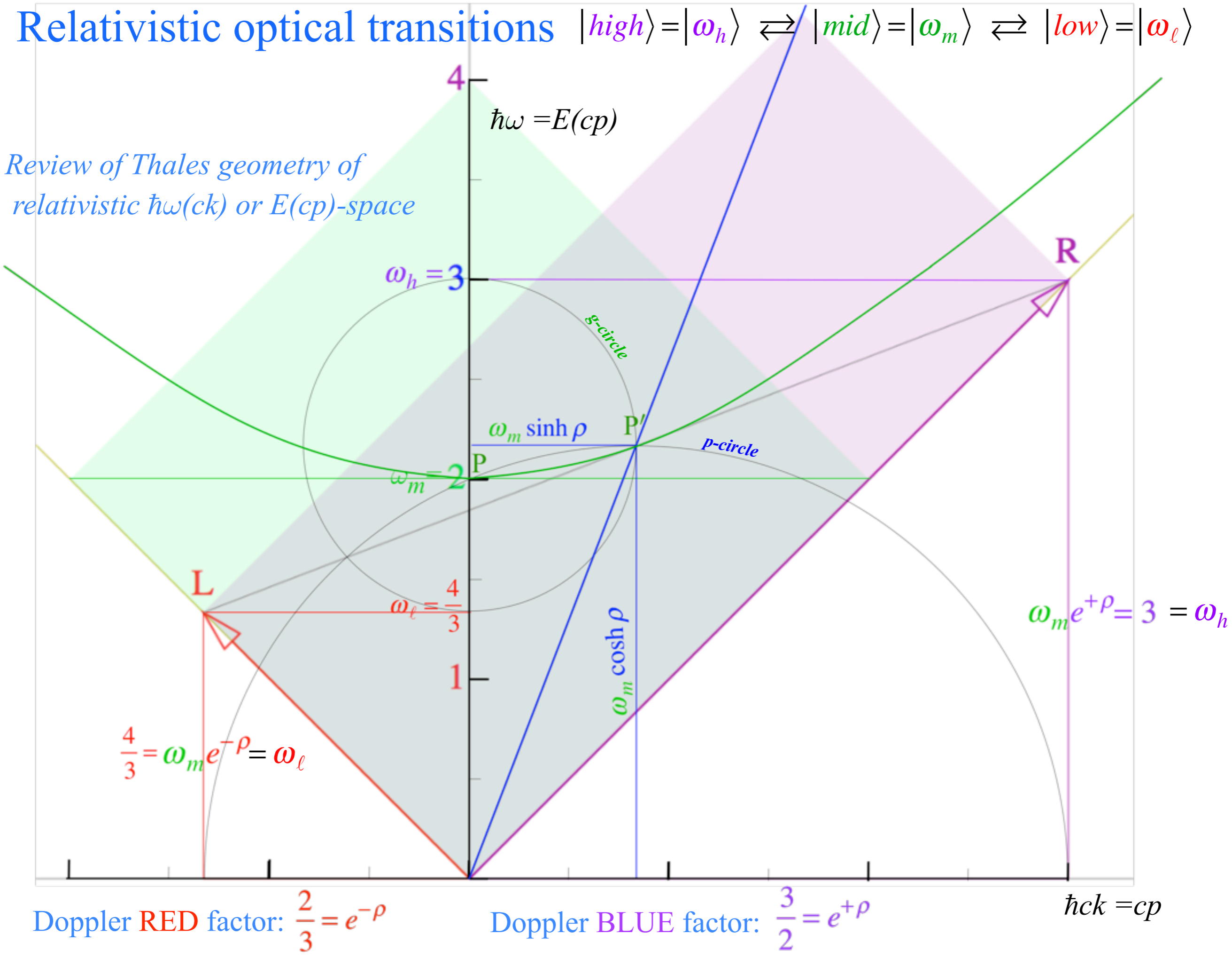
<i>group</i>	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
<i>phase</i>	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
<i>rapidity</i> ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
<i>stellar</i> ∇ <i>angle</i> σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
<i>value for</i> $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$
<i>effects</i>	$b_{RED}^{Doppler}$	V_{group}	<i>past-future asymmetry</i> (off-diagonal Lorentz-transform)	<i>x-contraction</i> (Lorentz) τ_{phase} -contraction	<i>t-dilation</i> (Einstein) v_{phase} -dilation (on-diagonal Lorentz-transform)	<i>inverse asymmetry</i>	V_{phase}	$b_{BLUE}^{Doppler}$

Relativistic quantum mechanics variable tables

<i>group</i>	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
<i>phase</i>	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
<i>rapidity</i> ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
<i>stellar</i> ∇ <i>angle</i> σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{\beta}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{1-\beta^2}}{\beta}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
<i>value for</i> $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$
<i>functions</i>		$V_{group} = c \tanh \rho$	<i>momentum</i> $cp = Mc^2 \sinh \rho$	<i>-Lagrangian</i> $L = -Mc^2 \operatorname{sech} \rho$	<i>Hamiltonian</i> $H = Mc^2 \cosh \rho$	<i>DeBroglie</i> $\lambda = \alpha \operatorname{csch} \rho$	$V_{phase} = c \operatorname{coth} \rho$	

Relativistic optical transitions $|high\rangle = |\omega_h\rangle \rightleftharpoons |mid\rangle = |\omega_m\rangle \rightleftharpoons |low\rangle = |\omega_l\rangle$

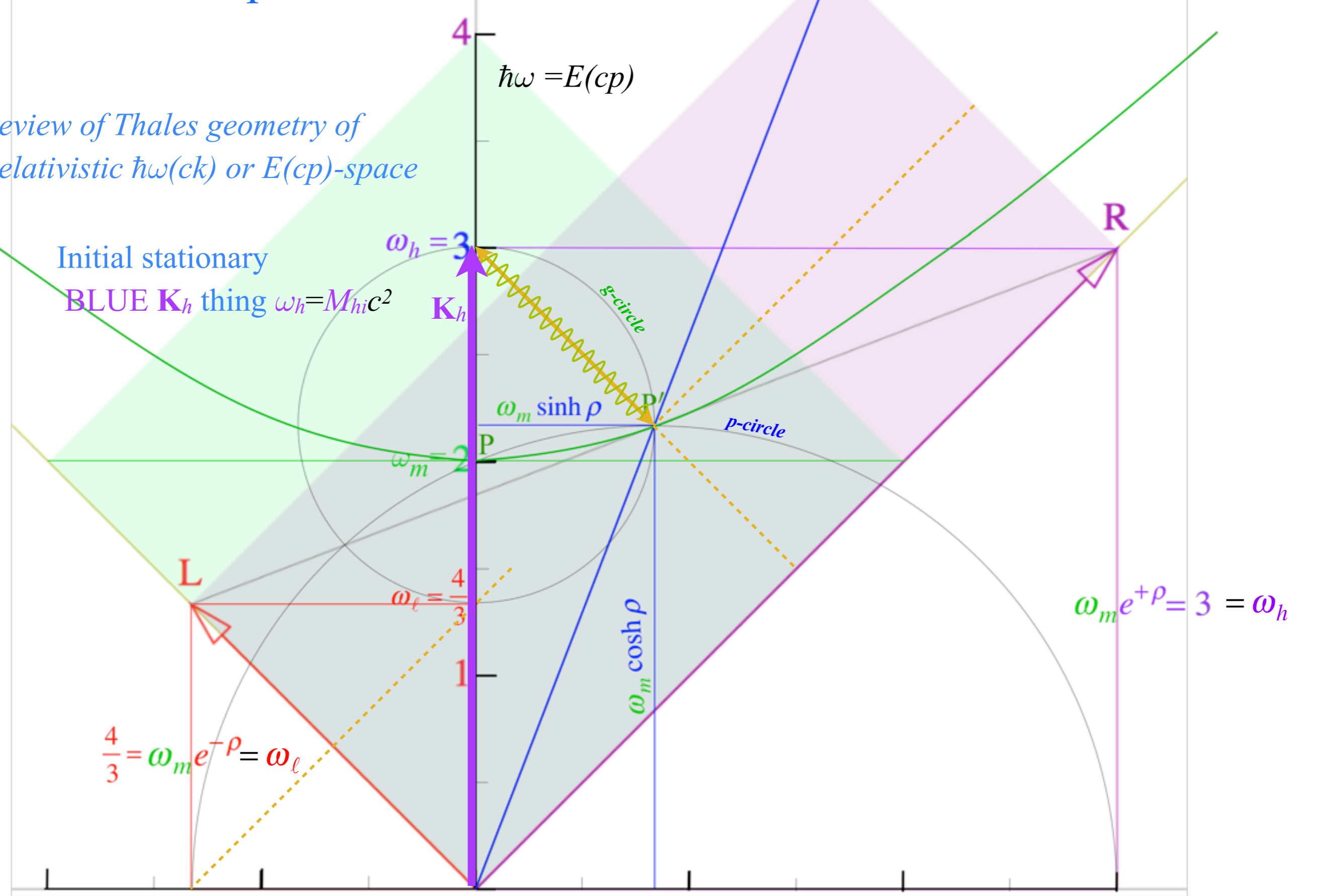
Review of Thales geometry of relativistic $\hbar\omega(ck)$ or $E(cp)$ -space



Relativistic optical transitions $|high\rangle = |\omega_h\rangle \rightleftharpoons |mid\rangle = |\omega_m\rangle \rightleftharpoons |low\rangle = |\omega_l\rangle$

Review of Thales geometry of relativistic $\hbar\omega(ck)$ or $E(cp)$ -space

Initial stationary BLUE K_h thing $\omega_h = M_{hi}c^2$



Doppler RED factor: $\frac{2}{3} = e^{-\rho}$

Doppler BLUE factor: $\frac{3}{2} = e^{+\rho}$

$\hbar ck = cp$

$\hbar\omega = E(cp)$

$\omega_h = 3$

$\omega_m = 2$

$\omega_l = \frac{4}{3}$

$\omega_m e^{+\rho} = 3 = \omega_h$

$\frac{4}{3} = \omega_m e^{-\rho} = \omega_l$

$\omega_m \cosh \rho$

$\omega_m \sinh \rho$

g-circle

p-circle

K_h

P'

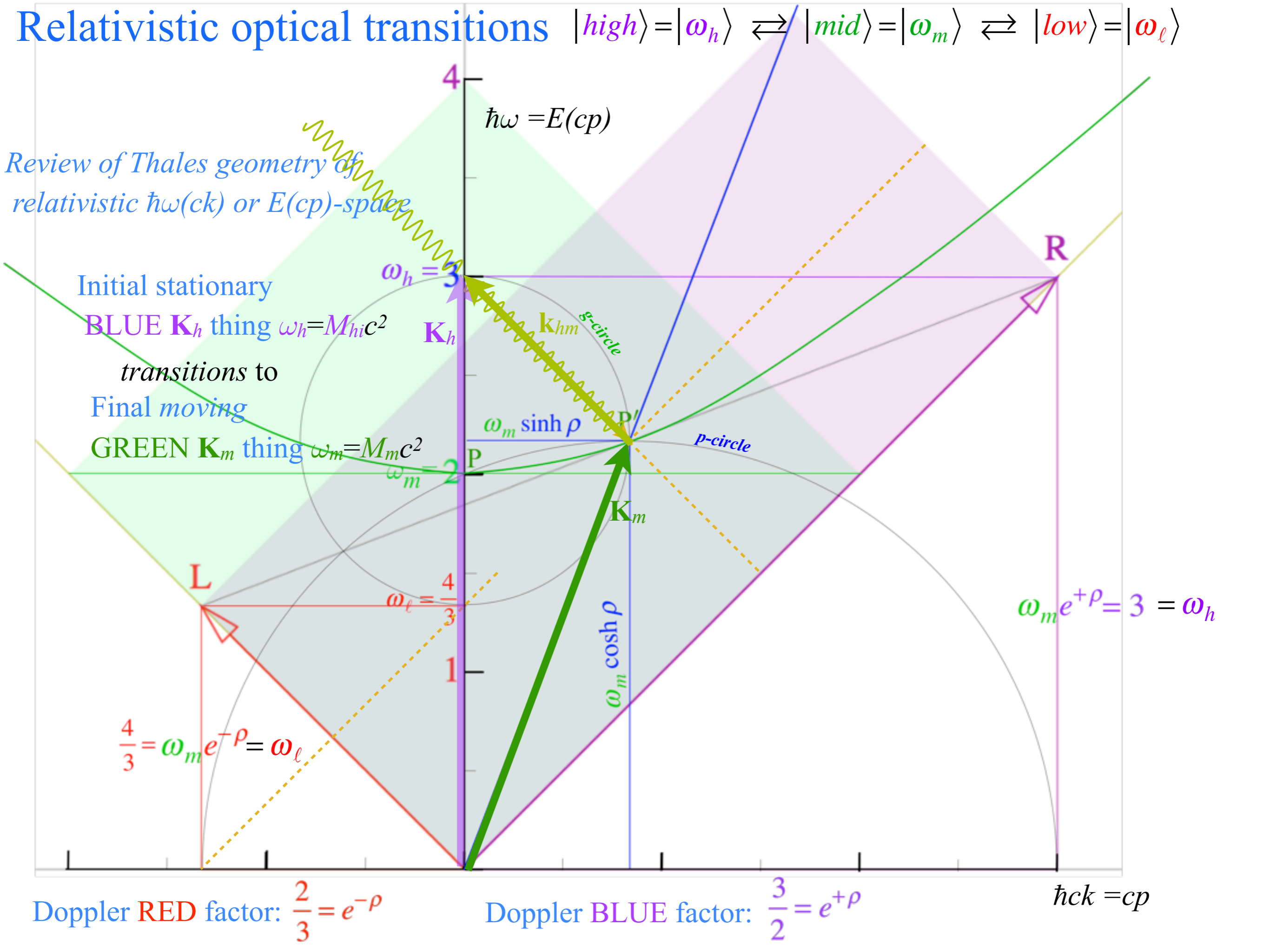
P

L

R

Relativistic optical transitions $|high\rangle = |\omega_h\rangle \rightleftharpoons |mid\rangle = |\omega_m\rangle \rightleftharpoons |low\rangle = |\omega_l\rangle$

Review of Thales geometry of relativistic $\hbar\omega(ck)$ or $E(cp)$ -space



Initial stationary
BLUE K_h thing $\omega_h = M_h c^2$

transitions to
Final moving
GREEN K_m thing $\omega_m = M_m c^2$

Doppler RED factor: $\frac{2}{3} = e^{-\rho}$

Doppler BLUE factor: $\frac{3}{2} = e^{+\rho}$

$\hbar ck = cp$

$\frac{4}{3} = \omega_m e^{-\rho} = \omega_l$

$\omega_m e^{+\rho} = 3 = \omega_h$

$\omega_m \sinh \rho$

$\omega_m \cosh \rho$

$\omega_h = 3$

$\omega_m = 2$

$\omega_l = \frac{4}{3}$

$\hbar\omega = E(cp)$

R

L

K_h

K_m

k_{hm}

g-circle

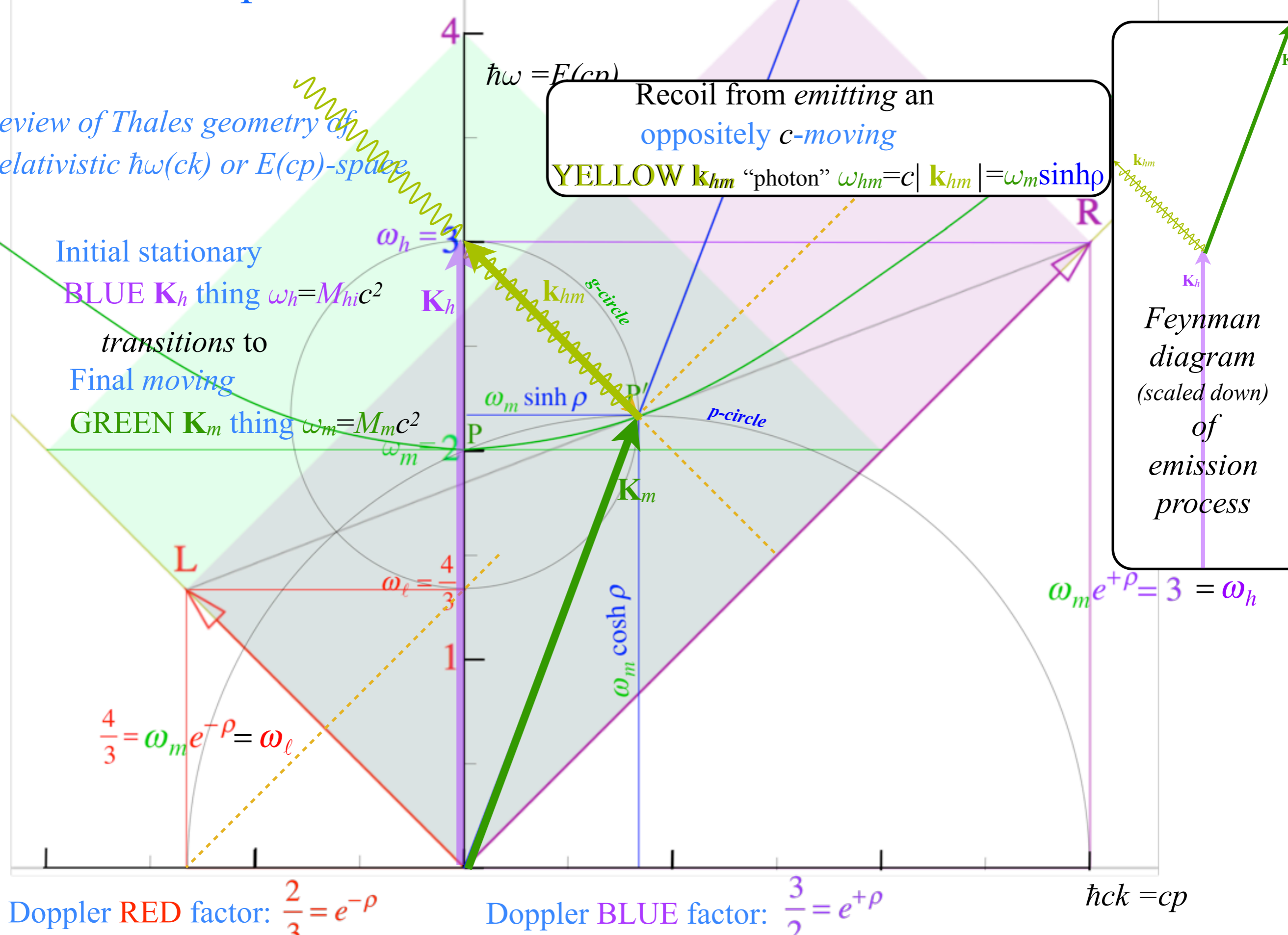
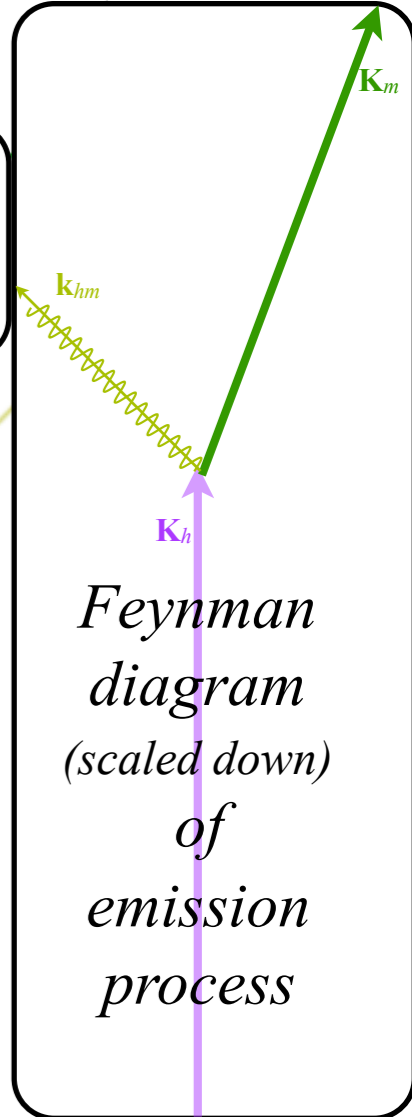
p-circle

Relativistic optical transitions $|high\rangle = |\omega_h\rangle \rightleftharpoons |mid\rangle = |\omega_m\rangle \rightleftharpoons |low\rangle = |\omega_l\rangle$

Review of Thales geometry of relativistic $\hbar\omega(cp)$ or $E(cp)$ -space

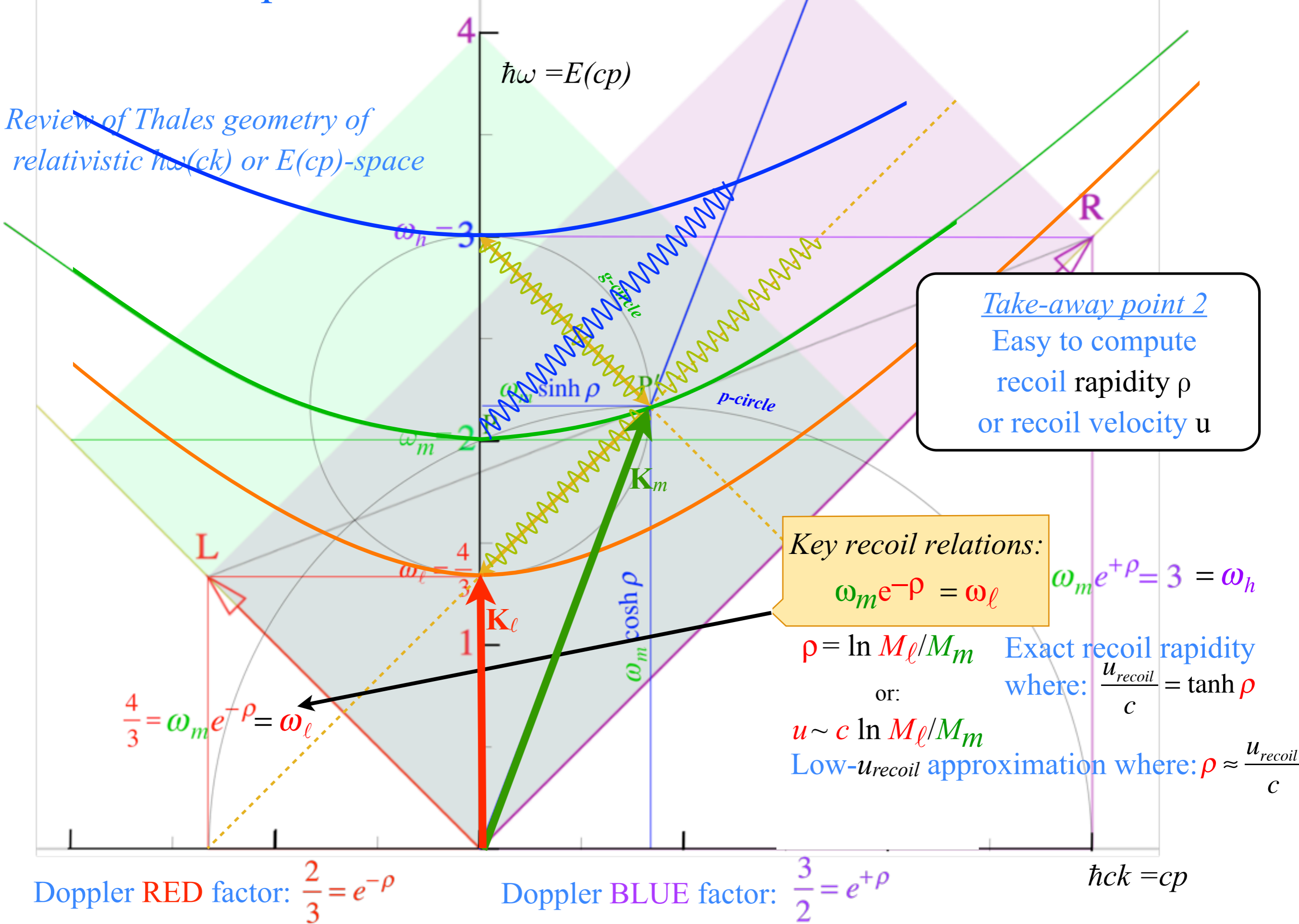
Initial stationary BLUE K_h thing $\omega_h = M_h c^2$
 transitions to Final moving GREEN K_m thing $\omega_m = M_m c^2$

Recoil from emitting an oppositely c -moving YELLOW k_{hm} "photon" $\omega_{hm} = c|k_{hm}| = \omega_m \sinh \rho$



Relativistic optical transitions $|high\rangle = |\omega_h\rangle \rightleftharpoons |mid\rangle = |\omega_m\rangle \rightleftharpoons |low\rangle = |\omega_l\rangle$

Review of Thales geometry of relativistic $\hbar\omega(ck)$ or $E(cp)$ -space



Take-away point 2
 Easy to compute
 recoil rapidity ρ
 or recoil velocity u

Key recoil relations:
 $\omega_m e^{-\rho} = \omega_l$ $\omega_m e^{+\rho} = 3 = \omega_h$

$\rho = \ln M_l / M_m$ Exact recoil rapidity
 where: $\frac{u_{recoil}}{c} = \tanh \rho$

or:
 $u \sim c \ln M_l / M_m$
 Low- u_{recoil} approximation where: $\rho \approx \frac{u_{recoil}}{c}$

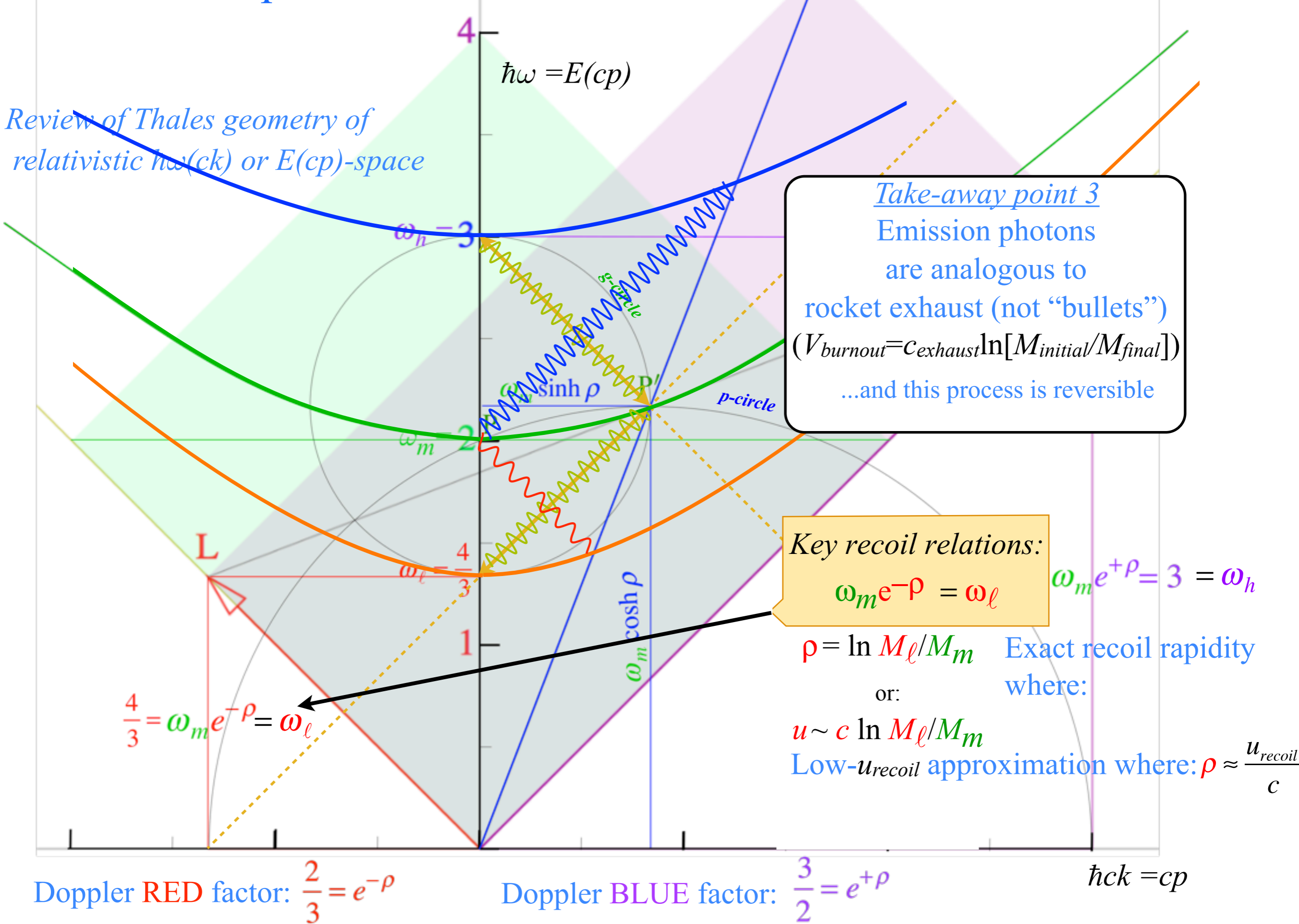
Doppler RED factor: $\frac{2}{3} = e^{-\rho}$

Doppler BLUE factor: $\frac{3}{2} = e^{+\rho}$

$\hbar ck = cp$

Relativistic optical transitions $|high\rangle = |\omega_h\rangle \rightleftharpoons |mid\rangle = |\omega_m\rangle \rightleftharpoons |low\rangle = |\omega_l\rangle$

Review of Thales geometry of relativistic $\hbar\omega(ck)$ or $E(cp)$ -space



Take-away point 3
 Emission photons are analogous to rocket exhaust (not “bullets”)
 ($V_{burnout} = c_{exhaust} \ln[M_{initial}/M_{final}]$)
 ...and this process is reversible

Key recoil relations:
 $\omega_m e^{-\rho} = \omega_l$

$\omega_m e^{+\rho} = 3 = \omega_h$

$\rho = \ln M_l/M_m$ Exact recoil rapidity where:

or:
 $u \sim c \ln M_l/M_m$

Low- u_{recoil} approximation where: $\rho \approx \frac{u_{recoil}}{c}$

Doppler RED factor: $\frac{2}{3} = e^{-\rho}$

Doppler BLUE factor: $\frac{3}{2} = e^{+\rho}$

$\hbar ck = cp$

(p, q) - coordinates

rest frequency: rapidity:

$$\omega_q = \omega_m e^{q\rho} \qquad \rho_p = p\rho$$

$$P_{p,q} = (ck_{p,q}, \omega_{p,q})$$

$$= \omega_m e^{q\rho} (\sinh p\rho, \cosh p\rho)$$

All-rational-fraction lattice defined by discrete sub-group of Lorentz Poincare Group
(Feynman path integrals defined by group transformations)

+2

+1

0

-1

-2

$L = \text{lefthand shift power}$
 $\omega_L = \omega_m e^{L\rho}$

-2

-1

0

+1

+2

$R = \text{righthand shift power}$
 $\omega_R = \omega_m e^{R\rho}$

$(p, q) - (R, L)$
coordinate

transformations:

$$p = \frac{R-L}{2}, \quad q = \frac{R+L}{2}$$

$$R = p+q, \quad L = q-p$$

Acceleration by chirping laser pairs

Varying acceleration (General case)

From Lect. 35
ModPhys (2012)

Only green-light is seen by observers on the green accelerated trajectory

Varying local acceleration $\rho = \rho(\tau)$

$$u = \frac{dx}{dt} = c \tanh(\tau)$$

$$\frac{dt}{d\tau} = \cosh \rho(\tau)$$

$$\frac{dx}{d\tau} = \frac{dx}{dt} \frac{dt}{d\tau} = c \tanh \rho(\tau) \cosh \rho(\tau) = c \sinh \rho(\tau)$$

$$ct = c \int \cosh \rho(\tau) d\tau$$

$$x = c \int \sinh \rho(\tau) d\tau$$

Constant local acceleration $\rho = \frac{g\tau}{c}$ "Einstein Elevator"

$$ct = c \int \cosh \frac{g\tau}{c} d\tau = \frac{c^2}{g} \sinh \frac{g\tau}{c}$$

$$x = c \int \sinh \frac{g\tau}{c} d\tau = \frac{c^2}{g} \cosh \frac{g\tau}{c}$$

Previous examples involved constant velocity

Constant velocity $\rho = \rho_0 = \text{const.}$ "Lorentz transformation"

$$ct = c \int \cosh \rho_0 d\tau = c\tau \cosh \rho_0$$

$$x = c \int \sinh \rho_0 d\tau = c\tau \sinh \rho_0$$

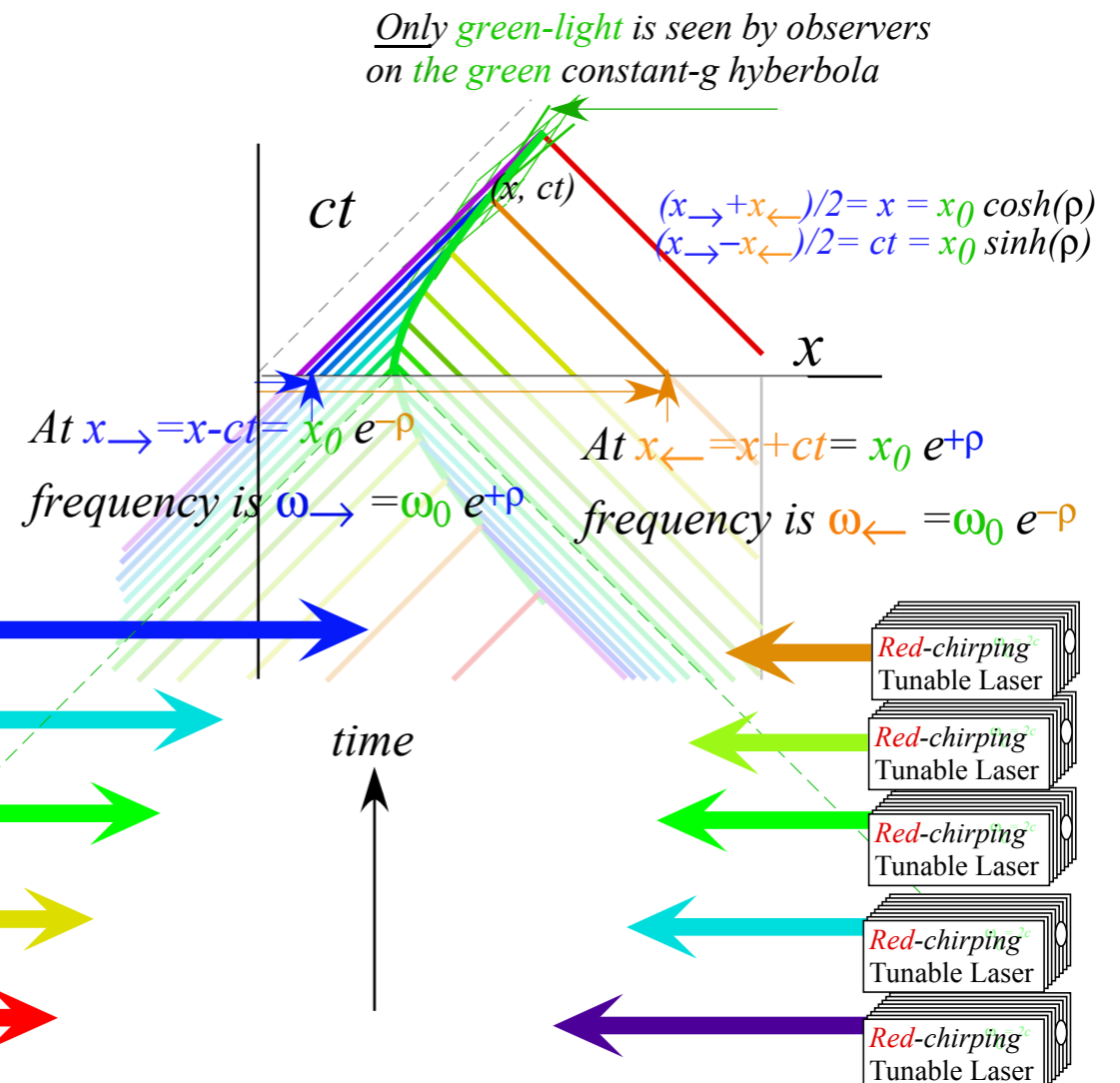
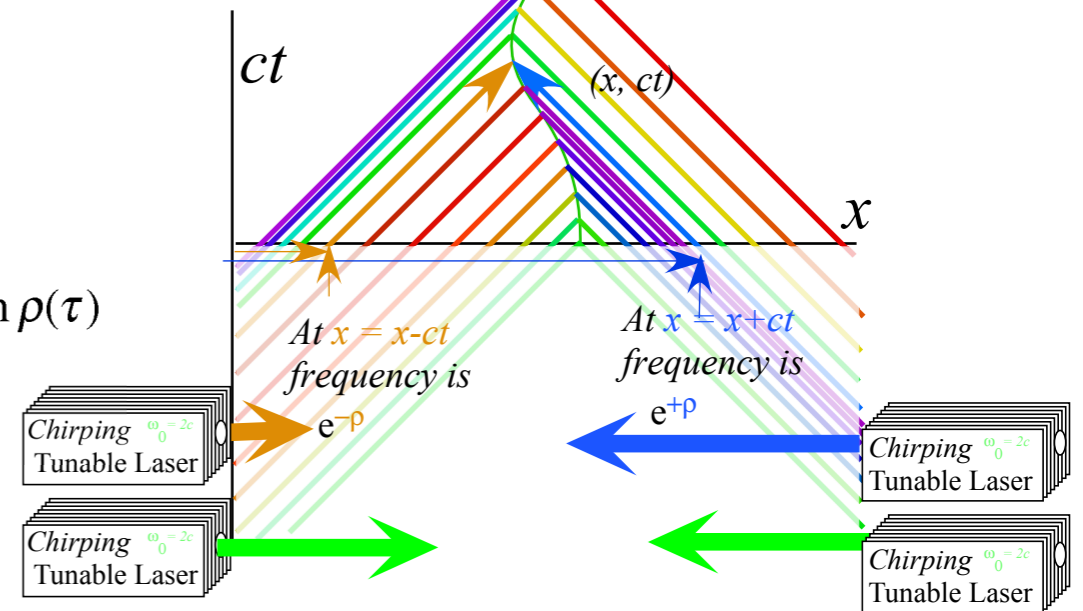
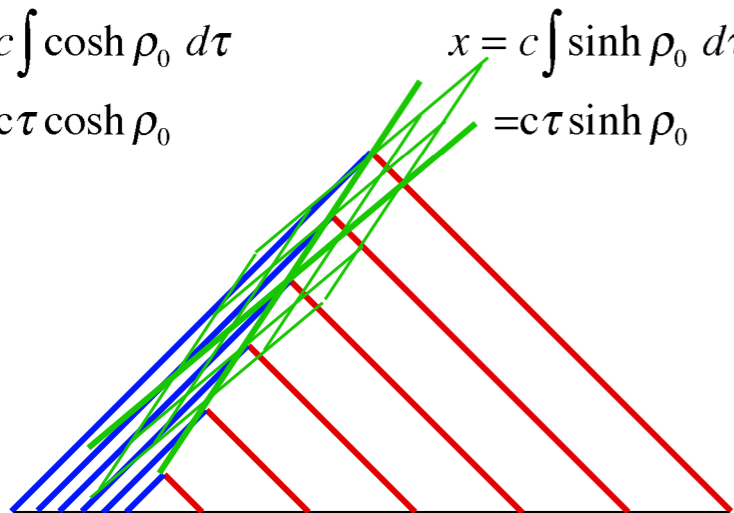
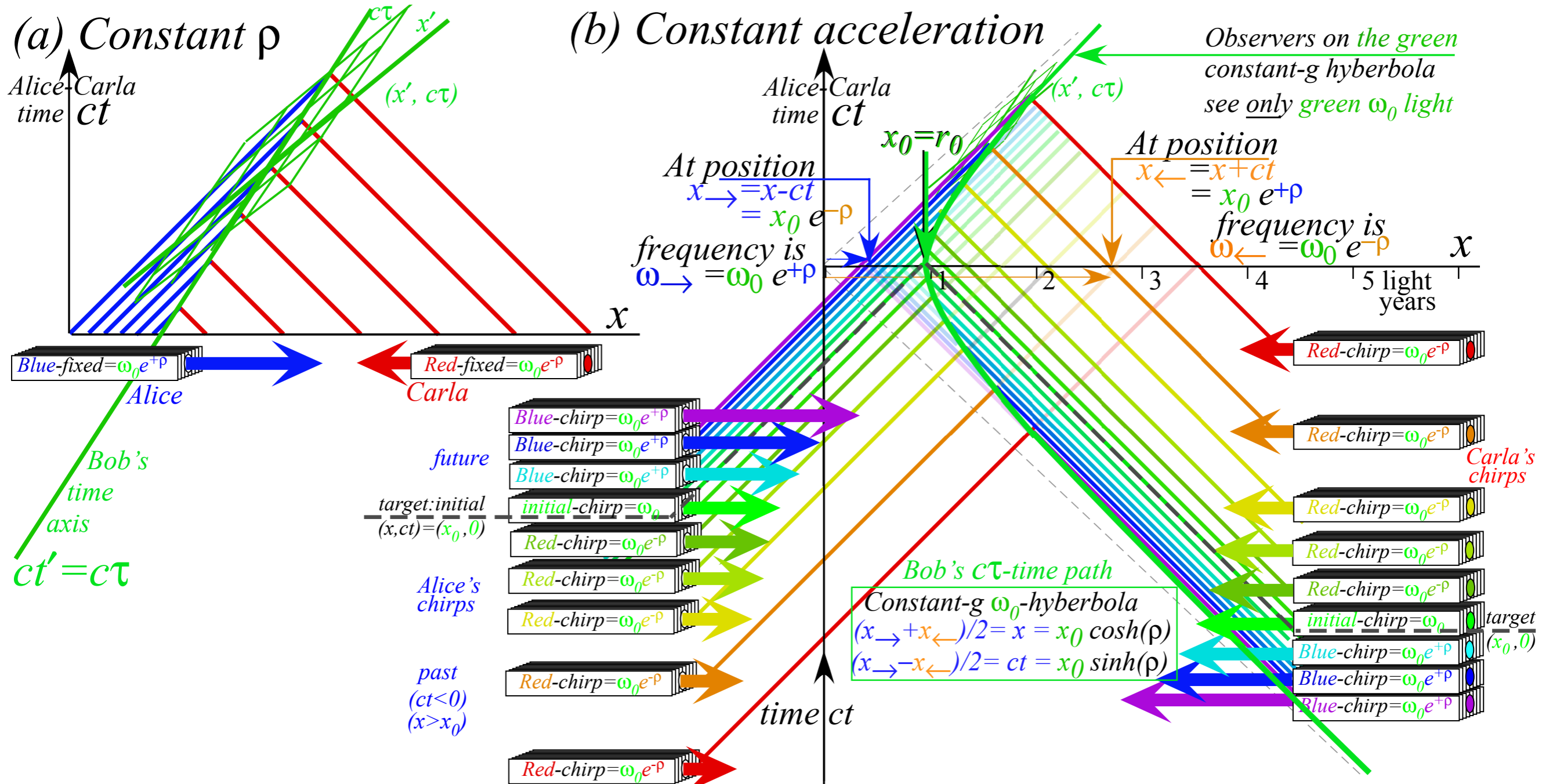
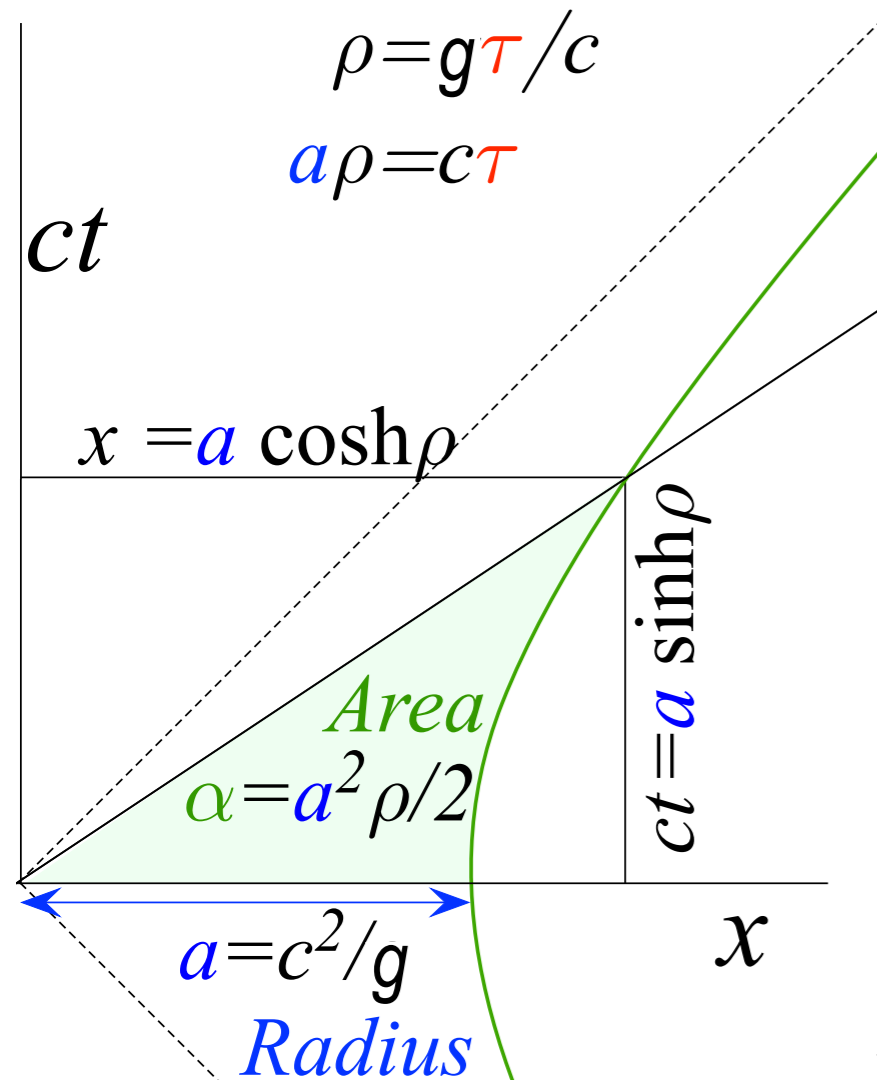


Fig. 8.1 Optical wave frames by red-and-blue-chirped lasers (a)Varying acceleration (b)Constant g



(a) Constant acceleration g
 Rapidity ρ vs proper time τ



(b) Traveler paths of acceleration g_q

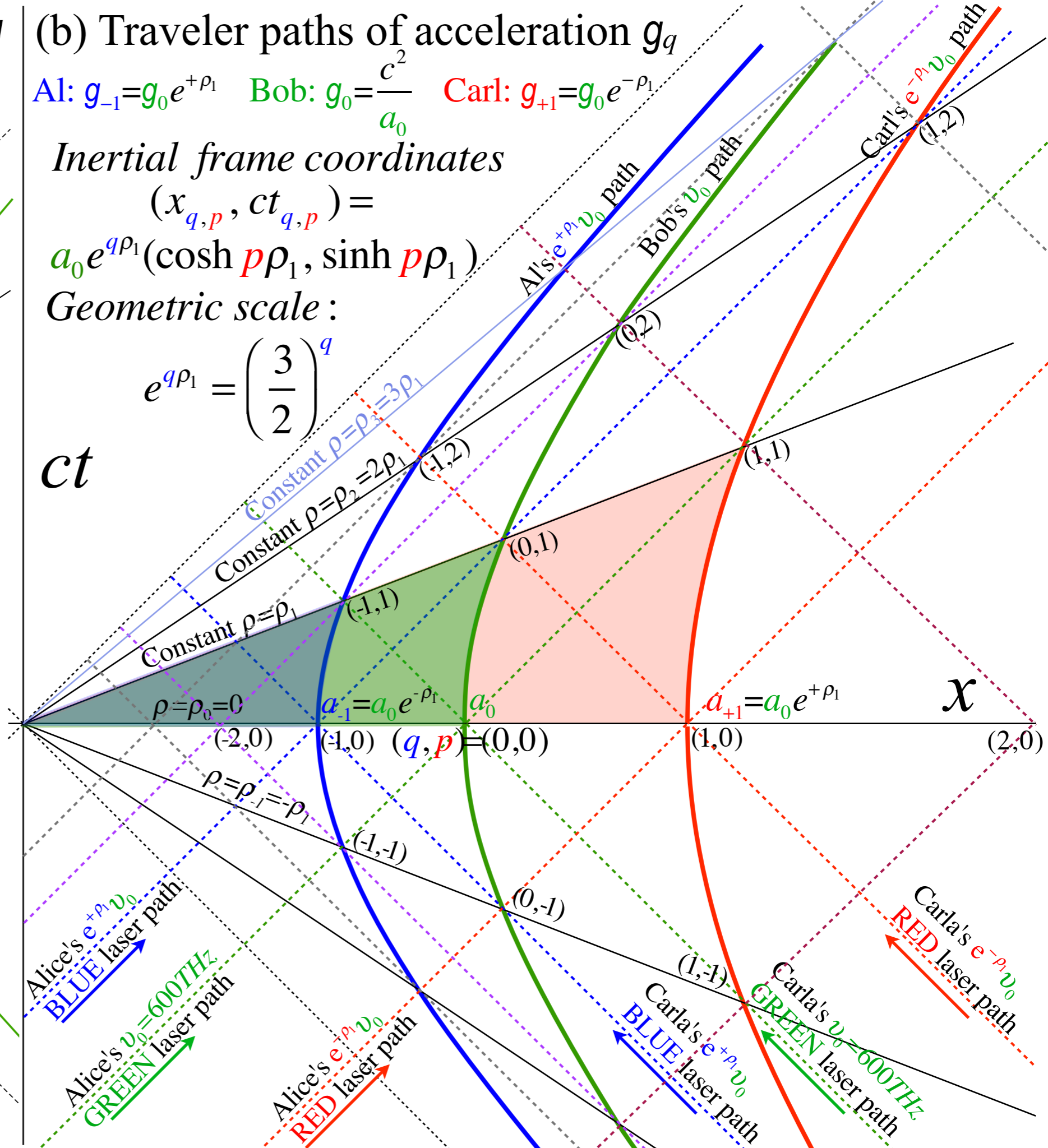
Al: $g_{-1} = g_0 e^{+\rho_1}$ Bob: $g_0 = \frac{c^2}{a_0}$ Carl: $g_{+1} = g_0 e^{-\rho_1}$

Inertial frame coordinates

$(x_{q,p}, ct_{q,p}) =$
 $a_0 e^{q\rho_1} (\cosh p\rho_1, \sinh p\rho_1)$

Geometric scale:

$e^{q\rho_1} = \left(\frac{3}{2}\right)^q$



RelativIt Web Simulation
{Accelerated proper-time
frame}

