

# Tunable Laser Spectra of the Infrared-Active Fundamentals of Cubane

A. S. Pine,\*† A. G. Maki,† A. G. Robiette,‡ B. J. Krohn,⊥ J. K. G. Watson,|| and Th. Urbanek§










Contribution from the Molecular Spectroscopy Division, National Bureau of Standards, Washington, DC 20234, the Department of Chemistry, University of Reading, Whiteknights, Reading, UK RG6 2AD, the University of California, Theoretical Division, Los Alamos National Laboratory, Los Alamos, New Mexico 87545, the Herzberg Institute of Astrophysics, National Research Council of Canada, Ottawa, Ontario, Canada K1A 0R6, and the Institut für Organische Chemie I, Universität Dusseldorf, Dusseldorf, FDR. Received September 19, 1983

**Abstract:** Confirmation of the octahedral symmetry and improved bond length measurements are obtained from high-resolution tunable laser spectra of the three infrared-active fundamental vibrations of cubane in the vapor phase. The C-H stretching band ( $\nu_{10}$ ) was recorded with a difference-frequency laser and is found to be severely perturbed by a second-order Coriolis resonance with another nearby (nominally inactive) C-H stretch. The C-H bend ( $\nu_{11}$ ) and the C-C stretch ( $\nu_{12}$ ), which were studied with diode lasers, are relatively unperturbed, revealing the symmetry from the nuclear spin weight intensities of the ro-vibrational clusters and the bond lengths from an analysis of the rotational fine structure. The data also provided the  $f_{1u}$ -block  $\zeta$  constants, which have been used, together with vibrational fundamentals from an earlier solid-state study of cubane and its isotopic derivatives, to determine a quadratic force field for the molecule which results in some minor reassignments of the modes.

To structural chemists, cubane is one of the most intriguing of all organic molecules because of its high symmetry and unique bonding. As its name implies, cubane has a skeleton of eight carbon atoms located at the vertices of a cube, each bonded to three others at  $90^\circ$  (instead of the usual  $109.5^\circ$  tetrahedral bond angle of a saturated hydrocarbon) and to a single outboard hydrogen located along the body diagonal. Cubane was first synthesized by Eaton and Cole<sup>1</sup> in 1964; the sample used here was prepared<sup>2</sup> by using the method given by Chapman et al.<sup>3</sup> for a brominated precursor. At room temperature the molecule occurs in the solid phase as millimeter-sized crystallites with rhombohedral facets; the vapor pressure is about 0.5 torr. An X-ray study of the crystal by Fleischer<sup>4</sup> yielded an average C-C bond distance of 1.551 (5) Å and a C-H bond length of 1.06 (7) Å.

In the isolated molecule there are 18 fundamental vibrational modes, listed in Table I, but the cubic symmetry ( $O_h$ ) restricts the infrared activity to but 3 of these and the Raman activity to 8. In the solid phase, the rhombohedral distortion of the molecule by its crystal neighbors activates the remaining modes and splits some degeneracies. This has enabled Della et al.<sup>5</sup> in a thorough infrared and Raman study of the vibrational spectra of solid cubane and four deuterium-substituted derivatives, to determine all of the vibrational mode wavenumbers listed in the first column of Table I along with their symmetries and dominant characters. More recently, Cole et al.<sup>6</sup> examined the three infrared-active  $f_{1u}$  modes in the vapor phase at moderate resolution ( $\sim 0.06$  cm<sup>-1</sup> unapodized) with a Fourier transform interferometer in order to obtain rotation-vibration constants. Though unable to resolve the octahedral fine structure expected for a spherical top molecule, they were able to distinguish the  $J$  manifolds, resulting in reasonable rotational constants and band centers. We have recorded all three infrared-active fundamentals at Doppler-limited resolution using tunable laser instrumentation. The two lower frequency modes—namely the C-C stretch,  $\nu_{12}$ , at 852 cm<sup>-1</sup> and the C-H bend,  $\nu_{11}$ , at 1235 cm<sup>-1</sup>—were studied with Pb-salt diode lasers, whereas the high-frequency C-H stretch,  $\nu_{10}$ , at 2990 cm<sup>-1</sup> was recorded with a difference-frequency spectrometer.<sup>7</sup> Octahedral fine structure is observed in all three bands, but assignments and complete analyses have been accomplished as yet only for the lower

Table I. Vibrational Fundamentals (cm<sup>-1</sup>) of Cubane (C<sub>8</sub>H<sub>8</sub>)

solid <sup>a</sup>	vapor <sup>b</sup>	mode	symmetry	activity <sup>c</sup>	approximate description
2995	2990	$\nu_1$	$a_{1g}$	R, p	C-H
2978		$\nu_{10}$	$f_{1u}$	IR	C-H
~2978		$\nu_3$	$a_{2u}$		C-H
2970		$\nu_{13}$	$f_{2g}$	R, dp	C-H
1230	1235	$\nu_{11}$	$f_{1u}$	IR	
1182		$\nu_{14}$	$f_{2g}$	R, dp	
1151		$\nu_7$	$e_u$		
1130		$\nu_9$	$f_{1g}$		
1083		$\nu_5$	$e_g$	R, dp	
1036		$\nu_{17}$	$f_{2u}$		
1002	852	$\nu_2$	$a_{1g}$	R, p	C-C
912		$\nu_6$	$e_g$	R, dp	C-C
853		$\nu_{12}$	$f_{1u}$	IR	C-C
839		$\nu_4$	$a_{2u}$		
829		$\nu_{18}$	$f_{2u}$		C-C
821		$\nu_{15}$	$f_{2g}$	R, dp	C-C
665		$\nu_{16}$	$f_{2g}$	R, dp	
617		$\nu_8$	$e_u$		

<sup>a</sup> Reference 5. <sup>b</sup> Reference 6. <sup>c</sup> IR = infrared, R = Raman, p = polarized, dp = depolarized.

two bands which are relatively unperturbed. Hence we concentrate on the diode laser results in this paper since much structural

\* National Bureau of Standards

† University of Reading.

‡ Los Alamos National Laboratory.

⊥ National Research Council of Canada.

|| Institut für Organische Chemie I.

(1) P. E. Eaton and T. W. Cole, Jr., *J. Am. Chem. Soc.*, **86**, 3157-3158 (1964).

(2) H.-D. Martin, P. Pfohler, T. Urbanek, and R. Walsh, *Chem. Ber.*, **116**, 1415-1421 (1983).

Table II. Nuclear Spin Statistical Weights in Cubane ( $C_8H_8$ )

species	weight <sup>a,b</sup>
Single Levels	
$A_1$	23
$A_2$	13
$E$	20
$F_1$	27
$F_2$	33
Clusters	
$A_1 + F_1 + F_2 + A_2$	96
$F_1 + E + F_2$	80
$F_2 + E + F_1$	80
$A_1 + F_1 + E$	70
$F_1 + F_2$	60
$E + F_2 + A_2$	66
$F_2 + F_1$	60

<sup>a</sup> Reference 9. <sup>b</sup> Reference 10.

information has been obtained from them.

Because of the simplifying effects of its high symmetry, cubane may well be the largest molecule (in terms of numbers of atoms) for which the rotational structure will ever be resolvable at the Doppler limit. Benzene<sup>8</sup> was the former record holder, whereas most chain hydrocarbons with four or more carbons yield only a dense, unresolvable continuum with Doppler-limited or lower resolution. Also fortuitously, the very rigid carbon framework of cubane causes the lowest frequency carbon deformation modes to occur above 600  $cm^{-1}$ . This results in only a modest thermal population of about 6 and 7% for the lowest  $\nu_8$  and  $\nu_{16}$  modes, respectively, at room temperature and a ground-state occupancy of nearly 61%. Thus hot bands do not interfere significantly with the spectra of the fundamentals, although they are in evidence. Another minor complication arises from the  $^{13}C$  isotope in 1.1% natural abundance which leads to an 8.9% population for  $^{13}C^{12}C_7H_8$  relative to pure  $^{12}C_8H_8$ .

The prior X-ray data<sup>4</sup> and low-resolution spectra<sup>5,6</sup> were certainly indicative of and consistent with the cubic symmetry of the molecule, but were not conclusive. The high-resolution spectrum can, however, establish with certainty the cubic symmetry from the spin-statistical weights of the octahedral fine structure transitions due to the eight indistinguishable hydrogens and carbons. The spin weights have been worked out recently by Balasubramanian et al.<sup>9</sup> and independently by Weber<sup>10</sup> with the results shown in Table II. As is characteristic of heavy spherical top spectra, many individual ro-vibrational transition fall into nearly degenerate clusters corresponding to molecular rotations about high-symmetry threefold or fourfold axes.<sup>11</sup> These clusters exhibit distinctive intensity alternation patterns, also given in Table II, which help to establish the  $J$  assignments. The cluster intensity patterns for cubane show less dramatic alternation than those for smaller spherical tops, such as  $CH_4^{12}$  or  $SF_6$ ,<sup>13</sup> because of the larger number of possible spin wave function combinations. However, they are readily discernible in the diode laser spectra as illustrated in Figure 1 by the R(36) manifold of  $\nu_{12}$ , which conclusively

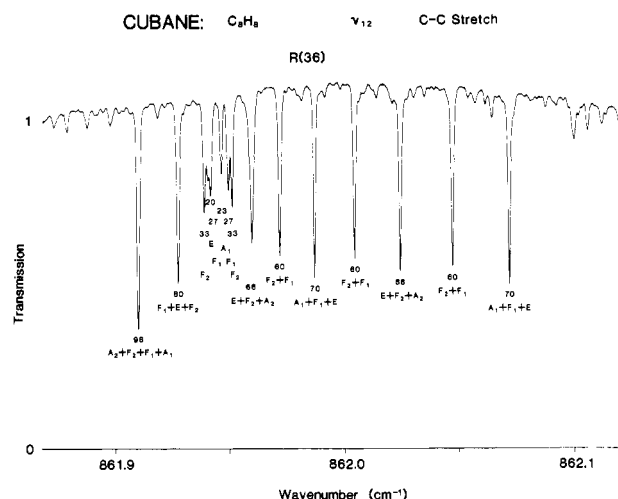


Figure 1. Diode laser transmission trace of the R(36) manifold of the C-C stretching band,  $\nu_{12}$ , of cubane illustrating alternating intensities due to nuclear spin statistical weights for indicated transitions and clusters.

confirms the cubic symmetry of cubane. This intensity alternation also leads to an unambiguous  $J$  assignment for the P- and R-branch manifolds.

### Experimental Details

The cubane sample of a few milligrams was stored in an evacuated, sealed-off glass vial and transferred by sublimation under vacuum at liquid nitrogen temperature to another vial with a Kovar-seal mounted stainless steel valve. Other than pumping for a few minutes on the sample held at 77 K, no further purification was attempted or required. The absorption cell was a glass tube 50 cm long by 2.7 cm in diameter with NaCl windows glued on (with Glyptal) at a tilt angle of  $\sim 10^\circ$  to avoid spectral channeling. The cell stopcock had a greaseless Teflon seat since Della et al.<sup>5</sup> cautioned against cubane dissolving in stopcock grease. The absorption cell was filled with 0.50 torr of cubane at room temperature, and no sample degradation was observed over a period of more than 1 month. Spectral measurements were performed at  $T = 296$  K in a single-pass geometry; absorption depths ranged from 10 to 60% for the measured allowed lines.

Spectra of the C-H stretching band were recorded with a scan-controlled difference-frequency laser<sup>7</sup> with digital data acquisition and processing. Complete coverage, base line normalization, Beer's law correction, and precision calibration against  $CH_4^{14}$  standards were achieved for this band. For the lower frequency bands, Pb-salt diode lasers were obtained commercially<sup>15</sup> and were used with a monochromator mode sorter and HgCdTe detectors. Two diodes were employed, one for each band, yielding 30–40% coverage, which was an adequate random sample in both cases to obtain a useful analysis. Calibration was accomplished by reference to  $OCS^{16}$  absorption lines in the 850- $cm^{-1}$  region and  $N_2O^{17}$  lines near 1235  $cm^{-1}$  with use of an  $\sim 7.5$  cm long temperature-controlled germanium etalon for interpolation. All scans were within 0.5  $cm^{-1}$  of a reference line, the scans being limited primarily by diode laser mode-hopping and thermal stability. Slow thermal drifts of the etalon were partially compensated by averaging two sequential scans in opposite directions. These drifts, plus subsidiary mode leakage and diode laser amplitude fluctuations (chiefly from mode competition), limited the wavenumber precision to  $\sim \pm 0.0003$   $cm^{-1}$  as determined from out least-squares fitting. Uncertainties in the absolute calibration contribute an additional uncertainty in the band center of  $\pm 0.0001$   $cm^{-1}$  for  $\nu_{12}$  and  $\pm 0.0003$   $cm^{-1}$  for  $\nu_{11}$ . Line centers were read from charts with an HP9825 computer/digitizer. No base line corrections were applied to the diode laser data shown here.

### Spectra

Figures 2 and 3 show several illustrative traces of R- and P-branch manifolds for the  $\nu_{11}$  and  $\nu_{12}$  bands of cubane. In Figure 2a we note the reversal of the R(36) pattern for  $\nu_{11}$  compared to that of  $\nu_{12}$  in Figure 1. This indicates that the sign of  $g$ , the

(3) N. B. Chapman, J. M. Key, and K. J. Toyne, *J. Org. Chem.*, **35**, 3860–3867 (1970).

(4) E. B. Fleischer, *J. Am. Chem. Soc.*, **86**, 3889–3890 (1964).

(5) E. W. Della, E. F. McCoy, H. K. Patney, G. S. Jones, and F. A. Miller, *J. Am. Chem. Soc.*, **101**, 7441–7457 (1979).

(6) T. W. Cole, Jr., J. Perkins, S. Putnam, P. W. Pakes, and H. S. Strauss, *J. Phys. Chem.*, **85**, 2185–2189 (1981).

(7) A. S. Pine and A. G. Robiette, to be published.

(8) J. Pliva and A. S. Pine, *J. Mol. Spectrosc.*, **93**, 209–236 (1982).

(9) K. Balasubramanian, K. S. Pitzer, and H. S. Strauss, *J. Mol. Spectrosc.*, **93**, 447–449 (1982).

(10) A. Weber, private communication.

(11) W. G. Harter and C. W. Patterson, *J. Chem. Phys.*, **66**, 4872–4885, 4886–4892 (1977).

(12) G. Herzberg, "Infrared and Raman Spectra of Polyatomic Molecules", D. Van Nostrand Co., Princeton, N. J., 1945, Chapter I, Section 3.

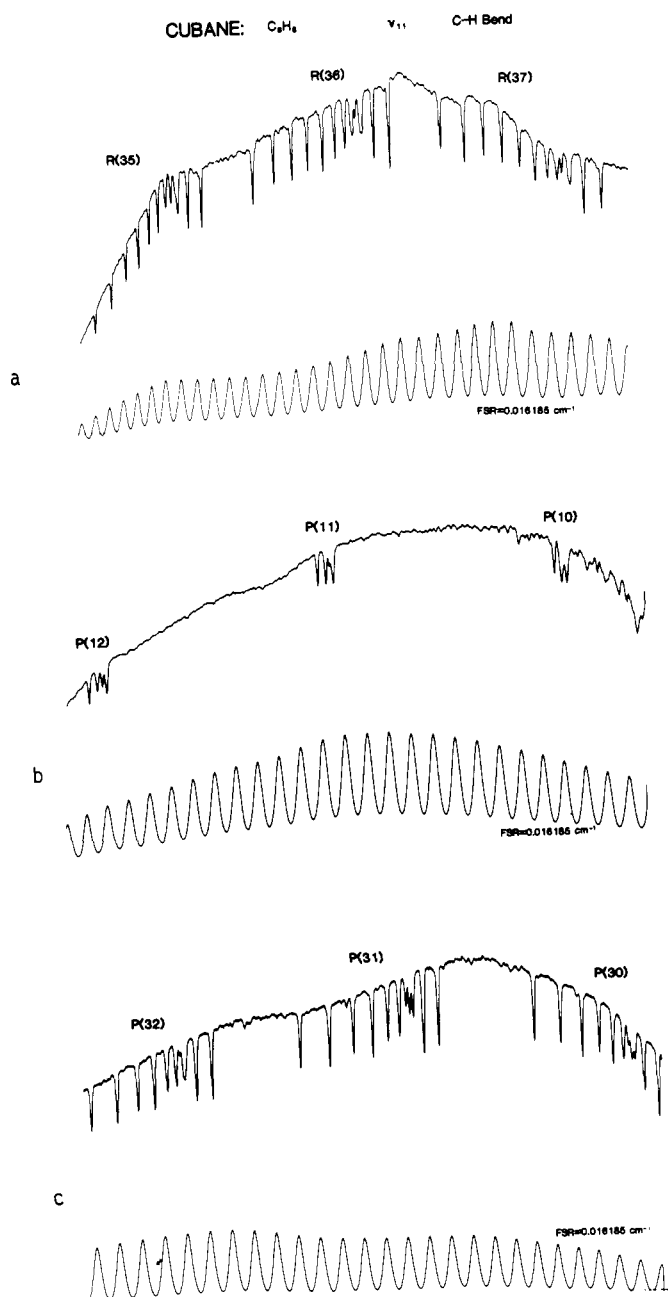
(13) C. D. Cantrell and H. W. Galbraith, *J. Mol. Spectrosc.*, **58**, 158–164 (1975).

(14) G. Tarrago, M. Dang-Nhu, G. Poussique, G. Guelachvili, and C. Amiot, *Mol. Phys.*, **57**, 246–263 (1975).

(15) Laser Analytics Division of Spectra-Physics, Bedford, MA.

(16) A. G. Maki, J. S. Wells, F. R. Peterson, W. B. Olson, A. Fayt, and J. P. Sattler, *J. Phys. Chem. Ref. Data*, in press.

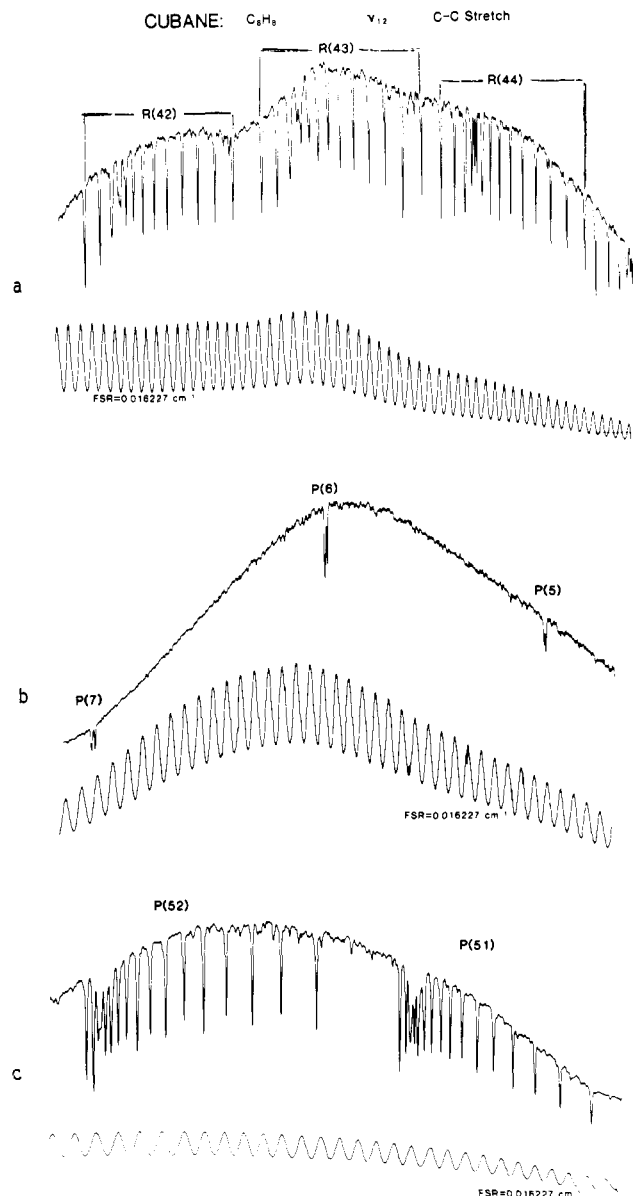
(17) G. Guelachvili, *Can. J. Phys.*, **60**, 1334–1347 (1982).



**Figure 2.** Diode laser transmission traces for selected  $J$  manifolds in the P and R branches of the C-H bending band,  $\nu_{11}$ , of cubane. Frequency increases to the right.

octahedral splitting coefficient, is opposite for the two bands and roughly equal in magnitude. The lower  $J$  P-branch manifolds in Figure 2b exhibit an extraneous feature to the high-frequency side of P(10), seen also in the lower resolution  $\nu_{11}$  spectrum of Cole et al.,<sup>6</sup> which probably arises from either a hot band or a  $^{13}\text{C}^{12}\text{C}_7\text{H}_8$  isotope band. As will be shown below, the  $\nu_{11}$  Q branch itself is quite narrow for transitions from the ground vibrational state, comprising only the single strongest peak in the Cole<sup>6</sup> spectrum, and cannot account for the structure near P(10). The alternating cluster intensity ratios are again evident for the higher  $J$  P-branch manifolds of  $\nu_{11}$  shown in Figure 2c.

The  $\nu_{12}$ -band manifolds shown in Figure 3 are similar but exhibit some marked deviations from the regular spherical top  $F_4$  patterns usually encountered.<sup>18</sup> First we note that the high  $J$  manifolds near R(43) in Figure 3a are nearly overlapping whereas the P manifolds in Figure 3c up to  $J \sim 52$  are still well separated.



**Figure 3.** Diode laser transmission traces for selected  $J$  manifolds in the P and R branches of the C-C stretching band,  $\nu_{12}$ , of cubane. Frequency increases to the right.

Without centrifugal distortion and other perturbations, the P-branch manifolds would overlap at lower  $J$  than the R-branch manifolds. Also the  $F_4$  patterns of the P(51) and P(52) manifolds themselves are severely compressed on the threefold (left or lower wavenumber) side, indicative, of a perturbation requiring higher-order terms in the theoretical fit. The low  $J$  P-branch multiplets in Figure 3b illustrate that the Q-branch tail is not extensive and that the spectra are relatively free of interfering isotope and hot bands. Because of diode laser idiosyncrasies, we were not able to scan the Q branch of  $\nu_{12}$ . As will be seen later, however, we do obtain some Q-branch information because of the mixing between the Coriolis sublevels in this band.

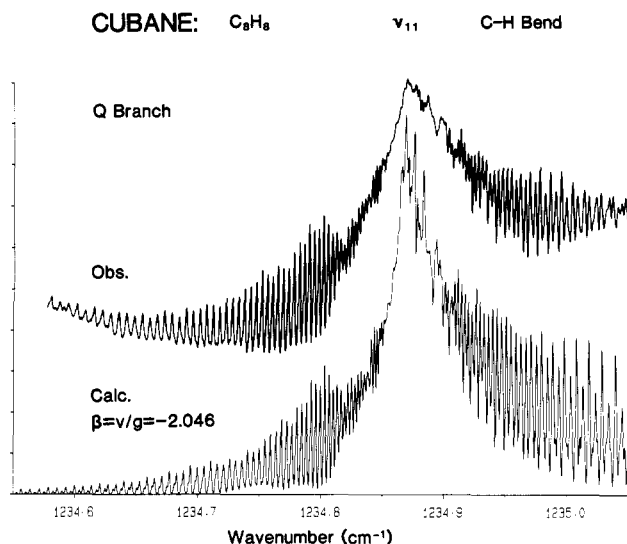
In Figure 4 we show the observed Q branch of  $\nu_{11}$ , the C-H bend, compared to a simulated spectrum based on the very simple model<sup>18,19</sup>

$$\nu_Q = m + \nu J(J+1) - 2gQ_4 \quad (1)$$

where the band center,  $m$ , and the octahedral splitting constant,  $g$ , are determined from the P and R branches discussed later.  $Q_4$  is a symmetry-adapted vector-coupling coefficient for the Q

(18) B. Bobin and K. Fox, *J. Phys., Lett. (Orsay, Fr.)*, **34**, 571-582 (1973); J. Moret-Bailly, *J. Mol. Spectrosc.*, **15**, 344-354 (1965).

(19) E. G. Brock, B. J. Krohn, R. S. McDowell, C. W. Patterson, and D. F. Smith, *J. Mol. Spectrosc.*, **76**, 301-321 (1979).



**Figure 4.** Comparison of the diode laser transmission trace (inverted) for the Q branch of the C-H bending mode,  $\nu_{11}$ , of cubane with the calculated intensity spectrum for  $\beta = 2.046$ .

branch<sup>18</sup> and  $v$ , a  $B' - B''$  type term, is adjusted for best agreement with experiment. The overall shape of the Q branch depends on the ratio<sup>19</sup>  $\beta = v/g = -2.046$  (8) and is seen to mimic the data faithfully out to very high  $J$  in the wings where centrifugal distortion effects, omitted from the model, become significant. The strong intensity peaks in the center of the Q branch in Figure 4 are somewhat saturated in the observed spectrum since base line and Beer's law corrections were not made. The calculated spectrum is presented with the  $y$  axis in absorbance units while the observed spectrum has a percent absorption scale with an absorption of more than 90% at the band center. Fortunately, the observed Q branch appears relatively uncluttered by extraneous isotope and hot-band structure over the range shown. The wavenumber scale in the observed trace is not exactly coincident with the simulation owing to slight nonlinearities in the tuning rate of the diode laser. However, the detailed band contour is very sensitive to the constants because of the extensive blending of clusters, which accounts for the relatively small uncertainty in  $\beta$  above.

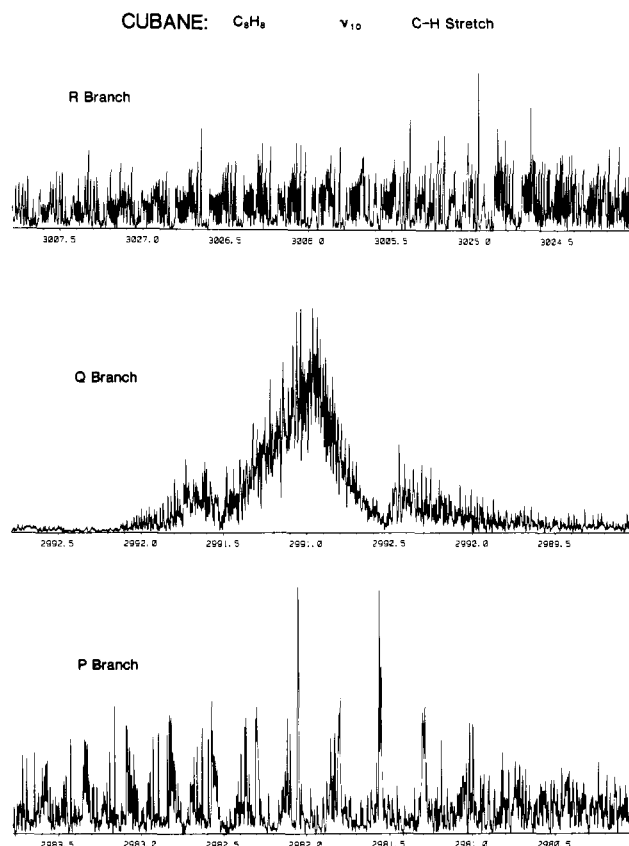
Portions of the  $\nu_{10}$ , C-H stretching, band are presented in Figure 5, illustrating the lack of recognizable  $F_4$  patterns in the P and R branches and the subsidiary  $Q^+$  and  $Q^-$  subbranches on either side of the stronger "allowed"  $Q^0$  subbranch.

#### Analysis

Although the spectra of  $\nu_{11}$  and  $\nu_{12}$  exhibit easily recognizable spherical top cluster patterns, neither band is adequately fit by the usual dominant approximation (the diagonal  $F_4$  expansion<sup>18</sup>). We use then a third-order perturbation theory approach developed by Krohn and Watson<sup>20</sup> summarized by

$$\begin{aligned} \nu_{R,P} = & m + nM + pM^2 + qM^3 + \dots \\ & + (g + hM + kM^2 + \dots)F_4(R,P) \\ & + (z' + z''M + \dots)F_6(R,P) \\ & + f(B\zeta, g, h, (v-p), M, F_4, F_6) \end{aligned} \quad (2)$$

where  $M = J + 1$  for the R branch and  $M = -J$  for the P branch and  $J$  is the ground state angular momentum quantum number. The scalar terms on the first line and the diagonal  $F_4$  and  $F_6$  octahedral coefficients on the second and third lines are augmented by corrections due to off-diagonal matrix elements on the fourth line given in detail in Table II of Krohn and Watson.<sup>20</sup> These corrections involve the second-order expansion parameter  $g^2/B\zeta$  and the third-order parameters  $g^3/(B\zeta)^2$ ,  $g^2(v-p)/(B\zeta)^2$ , and  $gh/B\zeta$ . Thus a nonlinear least-squares fit of this model to the data yields the constants  $B\zeta$  and  $v$ , which are not ordinarily ob-



**Figure 5.** Excerpts of the R, Q, and P branches of the C-H stretching mode,  $\nu_{10}$ , recorded with a difference-frequency laser spectrometer. The intensity of the Q branch has been reduced by two compared to the P and R branches for convenience in plotting.

tained from the P and R branches of an unperturbed infrared band of a spherical top. The convergence of the fit depends on the smallness of the perturbation theory expansion parameters. For the  $\nu_{11}$  and  $\nu_{12}$  bands of cubane the convergence is rapid (two iterations) and the results are presented in Table III. Parenthetically we note here that for the C-H stretching band,  $\nu_{10}$ , where  $\zeta \sim 0$ , the above expansion is inappropriate and a full diagonalization will be required. Omitted from the  $\nu_{11}$  and  $\nu_{12}$  fits were certain sums over off-diagonal  $F_4$  coefficients<sup>20</sup> which have not yet been tabulated. These contributions, which can be described as the second-order effects of an off-diagonal second-order contribution within a near-degenerate set, have a fourth-order numerator and a first-order denominator.<sup>20</sup> These corrections for cubane have been estimated as  $<5 \times 10^{-6} \text{ cm}^{-1}$  for  $J < 55$ , and so are entirely negligible compared to the measurement precision.

The overall root mean square error for the fit to 722 transitions (300 measured lines or clusters) of the  $\nu_{12}$  band is  $3.3 \times 10^{-4} \text{ cm}^{-1}$  and that to 580 transitions (248 measured lines or clusters) of the  $\nu_{11}$  band is  $2.6 \times 10^{-4} \text{ cm}^{-1}$ . The least-squares weighting for unresolved clusters was proportional to the relative spin weights of the transition components. A few threefold clusters in the high  $J(>50)$  P-branch manifolds of  $\nu_{12}$  were given reduced or zero weights because of observed - calculated wavenumber residuals greater than three times the root mean square error, which indicates the presence of a local resonant perturbation. The slightly worse fit of the  $\nu_{12}$  band probably reflects this perturbation.

The most likely candidate for the perturbation of  $\nu_{12}$  is a first-order Coriolis resonance with  $\nu_{18}$  ( $f_{2u}$ ,  $829 \text{ cm}^{-1}$ ) lying just  $24 \text{ cm}^{-1}$  below. The force-field calculation described later yields a rather large Coriolis coupling constant,  $|f_{12,18}| = 0.435$ , between these two vibrations. A higher-order rotational perturbation of  $\nu_{12}$  by the even closer  $\nu_4$  ( $a_{2u}$ ,  $839 \text{ cm}^{-1}$ ) is also a possibility. On the other hand,  $\nu_{11}$  is relatively isolated, being about  $90 \text{ cm}^{-1}$  above the nearest u species fundamental,  $\nu_7$  ( $e_u$ ,  $1151 \text{ cm}^{-1}$ ), which could participate in a Coriolis interaction. Possible candidates for a

(20) B. J. Krohn and J. K. G. Watson, to be published.

Table III. Vibration-Rotation Constants ( $\text{cm}^{-1}$ ) for Cubane

	C-C stretch		C-H bend	
	$\nu_{12}^a$	$\nu_{12}^b$	$\nu_{11}^a$	$\nu_{11}^b$
$m$	851.82750 (5) <sup>c</sup>	851.279 (8)	1234.86714 (3)	1234.670 (2)
$n$	0.2698460 (17)	0.26968 (4)	0.1734839 (17)	0.17338 (4)
$p/10^{-4}$	1.30935 (46)	1.160 (5)	0.51243 (35)	0.750 (8)
$q/10^{-7}$	-3.848 (13)	-2.68 (4)	-0.481 (20)	-0.083 (4)
$g/10^{-5}$	2.3955 (38)		-2.2504 (14)	
$h/10^{-7}$	1.5220 (27)		0.318 (38)	
$k/10^{-10}$	-7.48 (17)			
$z'/10^{-10}$	1.199 (19)		0.096 (38)	
$z''/10^{-13}$	-4.77 (49)			
$v/10^{-4}$	1.564 (92)		0.461 (2) <sup>d</sup>	
$B\zeta/10^{-2}$	-2.247 (31)	-1.786 (?) <sup>e</sup>	2.511 (49)	2.993 (?) <sup>e</sup>

<sup>a</sup> Present work. <sup>b</sup> Reference 6. <sup>c</sup> Errors in parentheses are one standard deviation. <sup>d</sup> From Q-branch simulation. <sup>e</sup> Estimated in ref 6 by using the  $\zeta$  sum rule; no errors stated.

Fermi-type resonance are for the most part well above  $\nu_{11}$ , the nearest being  $\nu_8 + \nu_{16}$  ( $\sim 1282 \text{ cm}^{-1}$ ).

Also in Table III we compare our present results with the lower resolution Fourier transform spectrum results of Cole et al.<sup>6</sup> Since they did not resolve the octahedral splittings, they could obtain only the scalar constants. The rotational constants are in reasonable agreement, but their band centers are too low—probably reflecting a misassignment of the  $J$  values by 2 for  $\nu_{12}$  and 1 for  $\nu_{11}$ , rather than a miscalibration. In the present study all the parameters listed are statistically well determined except for  $z'$  in  $\nu_{11}$  which is marginally about three standard deviations. Also the parameter  $v$  could not be determined with statistical significance in the  $\nu_{11}$  fit, so it was fixed to the value obtained from the Q-branch simulation. Note that for  $\nu_{12}$ ,  $v$  is well determined (and relatively uncorrelated with its partner,  $p$ ) even though we have not measured the Q branch.

The relationships between the fitting parameters  $n$ ,  $p$ ,  $v$ , etc., and the more usual molecular constants are to a good approximation

$$n = B_v + B_0 - 2B\zeta_v \quad (3)$$

$$(2p + v)/3 = B_v - B_0 \quad (4)$$

The values of  $n$ ,  $p$ ,  $v$ , and  $B\zeta_v$  given in Table III can be used to derive the ground-state  $B$  value. We obtain  $B_0 = 0.11183$  (25)  $\text{cm}^{-1}$  from  $\nu_{11}$ , and  $B_0 = 0.11238$  (16)  $\text{cm}^{-1}$  from  $\nu_{12}$ . The error limits on  $B_0$  arise from the experimental uncertainties in  $B\zeta_v$ . A weighted mean value of  $B_0$  from the two bands is  $B_0 = 0.1122$  (2)  $\text{cm}^{-1}$ , which in turn leads to  $\zeta_{11} = 0.224$  (4) and  $\zeta_{12} = -0.200$  (3).

The  $\zeta$  sum rule,  $\zeta_{10} + \zeta_{11} + \zeta_{12} = 0$ , for the three  $f_{1u}$  modes<sup>6,21</sup> yields  $\zeta_{10} \sim -0.024$  for the C-H stretching mode,  $\nu_{10}$ . This small value of  $\zeta_{10}$ , along with the very close proximity of the  $a_{2u}$  symmetry C-H stretching mode,  $\nu_3$ , leads to the extremely perturbed spectrum of  $\nu_{10}$  seen in Figure 5. The coupling between the  $f_{1u}$  and  $a_{2u}$  modes is through a second-order Coriolis operator of the kind

$$q_3[q_{10x}(J_y J_z + J_z J_y)] + \text{cyclic permutations}$$

analogous to the  $\nu_1/\nu_3$  coupling observed for the "local" hydrogen stretching modes in silane and germane.<sup>22-26</sup> The other C-H stretching modes of  $g$  parity cannot perturb  $\nu_{10}$  in any way. Furthermore, since all remaining fundamentals are less than 1250  $\text{cm}^{-1}$ , strong resonances with binary combinations cannot occur.

### Force-Field Calculations

The  $\zeta$  constants provide valuable information on the quadratic

force field. The 18 normal modes of cubane span the representation<sup>5,6</sup>

$$2a_{1g} + 2e_g + f_{1g} + 4f_{2g} + 2a_{2u} + 2e_u + 3f_{1u} + 2f_{2u}$$

so that only the  $f_{1u}$  (infrared-active) and  $f_{2g}$  (Raman-active) blocks involve more than a  $2 \times 2$  G-F calculation. A complete harmonic force field calculation has been carried out, using as data the solid-state vibrational wavenumbers measured by Della et al.<sup>5</sup> for cubane- $d_0$ , cubane- $d_2$ , cubane- $d_6$ , and cubane- $d_8$ , together with the gas-phase  $\zeta_{11}$  and  $\zeta_{12}$  for cubane- $d_0$  determined here.

Full details of the force-field work, i.e., the coordinates used to represent the force field and the interpretation of the potential constants obtained, will be published elsewhere.<sup>27</sup> For the purposes of this paper we note only that all force constants are well determined by the data apart from those in the  $4 \times 4$   $f_{2g}$  block (where at present one constraint has been made): gas-phase  $\zeta$  constants for the  $f_{2g}$  modes are needed to remove the indeterminacy in this block. Illustrative results from the force-field refinement are given in Tables IV and V. Table IV shows the wavenumber fit achieved for the four isotopic species included. The experimental data on the vibrational modes of cubane- $d_1$ , and the more tentative and incomplete assignments for cubane- $d_4$ , cubane- $d_5$ , cubane- $d_7$ , and  $^{13}\text{C}$ -substituted species, are also well reproduced. A few of the assignments of Della et al.<sup>5</sup> have been revised, notably the  $f_{2u}$  modes of cubane- $d_8$  and two cubane- $d_6$  fundamentals which correlate with these. It is also possible that minor changes are still needed in some of the fundamentals lying between 800 and 850  $\text{cm}^{-1}$  in cubane- $d_0$ ,  $-d_1$ , and  $-d_2$ . On the whole, however, the force-field calculation confirms the earlier assignments quite satisfactorily.

Table V gives more details of the  $f_{1u}$  and  $f_{2g}$  blocks of cubane- $d_0$  and cubane- $d_8$ . The sums of the calculated  $\zeta$  constants are in agreement with those derived by McDowell,<sup>21</sup> and individual  $\zeta$ 's fall within the limits given by McDowell in a later paper.<sup>28</sup> For the case of cubane these are  $-1/2 \leq \zeta_v \leq +1$  in both the  $f_{1u}$  and  $f_{2g}$  blocks. McDowell's results also require that  $\zeta_v = +1/2$  for the single  $f_{1g}$  mode and both  $f_{2u}$  modes in any isotopic species of cubane with  $O_h$  symmetry, irrespective of the force field. This was found to be so at all stages of the calculations.

The force field also yields information about the character of the  $\nu_{11}$  and  $\nu_{12}$  modes which we have analyzed. Cole et al.<sup>6</sup> drew attention to the  $f_{1u}$   $\zeta$ 's expected for hypothetical "pure" normal coordinates, i.e., 0 for a C-H stretching mode, +0.5 for a C-H bending mode, and -0.5 for a C-C stretching mode. The observed  $\zeta_{11}$  and  $\zeta_{12}$  for  $\text{C}_8\text{H}_8$  (+0.22 and -0.20, respectively) show that there is appreciable mixing between the  $f_{1u}$  C-H bending and C-C stretching coordinates, but it is clear that  $\nu_{11}$  is best described as C-H bending and  $\nu_{12}$  as C-C stretching as previous authors have done. The mixed nature of the normal modes is also shown by the substantial isotope shift in both  $\nu_{11}$  and  $\nu_{12}$  on going from  $\text{C}_8\text{H}_8$  to  $\text{C}_8\text{D}_8$ . In fact the calculated  $\zeta$ 's for  $\text{C}_8\text{D}_8$  (Table V) suggest

(21) R. S. McDowell, *J. Chem. Phys.*, **43**, 319-323 (1965).

(22) A. Cabana, D. L. Gray, I. M. Mills, and A. G. Robiette, *J. Mol. Spectrosc.*, **66**, 174-176 (1977).

(23) A. Cabana, D. L. Gray, A. G. Robiette, and G. Pierre, *Mol. Phys.*, **36**, 1503-1516 (1978).

(24) A. Owyong, P. Esherick, A. G. Robiette, and R. S. McDowell, *J. Mol. Spectrosc.*, **86**, 209-215 (1981).

(25) P. Lepage, J. P. Champion, and A. G. Robiette, *J. Mol. Spectrosc.*, **89**, 440-448 (1981).

(26) S. Q. Mao, R. Saint-Loup, A. Aboumjad, P. Lepage, H. Berger, and A. G. Robiette, *J. Raman Spectrosc.*, **13**, 257-261 (1982).

(27) A. G. Robiette and I. M. Mills, unpublished work.

(28) R. S. McDowell, *J. Chem. Phys.*, **46**, 1535-1536 (1967).

**Table IV.** Calculated Vibrational Fundamentals ( $\text{cm}^{-1}$ ) of Cubane- $d_0$ , - $d_2$ , - $d_6$ , and - $d_8$ , Compared with Solid-State Experimental Data<sup>a-c</sup>

$D_{3d}$ symmetry species	$d_0$ calcd	o-c	$d_2$ calcd	o-c	$d_6$ calcd	o-c	$d_8$ calcd	o-c
$A_{1g}$	3017 ( $A_{1g}$ )	-22	3011	-18	2996	-18	2240 ( $A_{1g}$ )	19
	2992 ( $F_{2g}$ )	-22	2220	17	2232	22	2208 ( $F_{2g}$ )	24
	1182 ( $F_{2g}$ )	0	1174	0	1081	2	1072 ( $F_{2g}$ )	0
	1003 ( $A_{1g}$ )	-1	991	0	967	0	955 ( $A_{1g}$ )	1
	826 ( $F_{2g}$ )	-5	819	2	725	0	711 ( $F_{2g}$ )	4
	665 ( $F_{2g}$ )	0	654	-3	581	-2	578 ( $F_{2g}$ )	8
$A_{2g}$	1131 ( $F_{1g}$ )	-1	1131	(0)	883	1	883 ( $F_{1g}$ )	1
$E_g$	2992 ( $F_{2g}$ )	-22	2991	-21	2210	19	2208 ( $F_{2g}$ )	24
	1182 ( $F_{2g}$ )	0	1169	-1	1146	-1	1072 ( $F_{2g}$ )	0
	1131 ( $F_{1g}$ )	-1	1099	1	1035	0	1027 ( $E_g$ )	0
	1085 ( $E_g$ )	-2	1014	2	983	2	883 ( $F_{1g}$ )	1
	913 ( $E_g$ )	-1	878	-3	755	3	711 ( $F_{2g}$ )	4
	826 ( $F_{2g}$ )	-5	740	-2	702	2	682 ( $E_g$ )	2
$A_{1u}$	665 ( $F_{2g}$ )	0	632	0	602	-4	578 ( $F_{2g}$ )	8
	1036 ( $F_{2u}$ )	0	1036	0	977	-2	977 ( $F_{2u}$ )	(0)
	829 ( $F_{2u}$ )	0	829	9	622	1	622 ( $F_{2u}$ )	1
	2998 ( $F_{1u}$ )	-20	2995	-27	2997	-23	2221 ( $A_{2u}$ )	(18)
	2994 ( $A_{2u}$ )	(-16)	2219	21	2220	22	2217 ( $F_{1u}$ )	23
	1230 ( $F_{1u}$ )	0	1220	2	1101	-1	1082 ( $F_{1u}$ )	1
$A_{2u}$	855 ( $F_{1u}$ )	-2	849	2	811	-4	802 ( $A_{2u}$ )	5
	841 ( $A_{2u}$ )	-2	824	3	688	2	683 ( $F_{1u}$ )	3
$E_u$	2998 ( $F_{1u}$ )	-20	2998	-21	2217	19	2217 ( $F_{1u}$ )	23
	1230 ( $F_{1u}$ )	0	1205	-4	1162	2	1082 ( $F_{1u}$ )	1
	1150 ( $E_u$ )	0	1111	-2	1014	0	977 ( $F_{2u}$ )	(0)
	1036 ( $F_{2u}$ )	0	993	2	972	3	958 ( $E_u$ )	2
	855 ( $F_{1u}$ )	-2	845	-1	791	-5	683 ( $F_{1u}$ )	3
	829 ( $F_{2u}$ )	0	712	-1	650	1	622 ( $F_{2u}$ )	1
	618 ( $E_u$ )	-1	574	0	537	1	525 ( $E_u$ )	2

<sup>a</sup> For cubane- $d_0$  and - $d_8$  the  $O_h$  symmetry species are indicated. The underlining of a calculated wavenumber implies a change in assignment from earlier work. <sup>b</sup> Columns headed o-c are observed - calculated differences, with observed wavenumbers taken from ref 5. A difference in parentheses indicates that the wavenumber was estimated from the product rule rather than observed directly.

**Table V.** Calculated Wavenumbers and  $\xi$  Constants of the  $f_{1u}$  and  $f_{2g}$  Blocks of Cubane- $d_0$  and Cubane- $d_8$ <sup>a</sup>

symmetry	$r$	$C_8H_8$		$C_8D_8$	
		$\nu_r$ ( $\text{cm}^{-1}$ )	$\xi_r^b$	$\nu_r$ ( $\text{cm}^{-1}$ )	$\xi_r$
$f_{1u}$	10	2998.3	-0.023	2217.2	-0.053
	11	1230.3	0.224	1082.3	-0.032
	12	855.2	-0.201	682.8	0.084
$f_{2g}$	13	2992.0	-0.088	2207.9	-0.074
	14	1182.2	0.857	1071.9	0.944
	15	826.2	0.412	711.5	-0.335
	16	665.4	-0.181	577.8	0.465

<sup>a</sup> The calculated  $\xi$  constants for the  $f_{1g}$  mode and both  $f_{2u}$  modes in  $C_8H_8$  and  $C_8D_8$  are all +0.5, in agreement with the sum rules and limiting values derived by McDowell (cf. ref 20 a and 21). <sup>b</sup> The calculated  $f_{1u}$   $\xi$  constants for  $C_8H_8$  should be compared with the observed values, i.e.,  $\xi_{11}(\text{obsd}) = +0.224$ ,  $\xi_{12}(\text{obsd}) = -0.200$  (see text); the sum rule would imply  $\xi_{10}(\text{obsd}) = -0.024$ .

that in the fully deuterated molecule the C-D bending and C-C stretching are mixed essentially 50/50, since both  $\xi_{11}$  and  $\xi_{12}$  are close 0.

### Molecular Structure

Since we have confirmed the cubic symmetry of cubane, only the C-C and C-H bond lengths are required to specify the structure of the molecule; these are designated here by  $R$  and  $r$  respectively for the ground vibrational state. The moment of inertia is then

$$I_0 = 4(M_C + M_H)R^2 + (16M_H/3^{1/2})[rR + (r^2/3^{1/2})] \quad (5)$$

which is related to the ground-state  $B_0$  value by

$$B_0 = h/8\pi^2 c I_0 \quad (6)$$

$B_0$  has been determined here for only one isotopic species, so we must make an assumption about one of the bond lengths in order to determine the other.

The molecule closest in structure to cubane studied earlier is probably *syn*-tricyclo[4.2.0.0<sup>2,5</sup>]octane, for which Andersen and Fernholt<sup>29</sup> found mean bond lengths of  $r_c(\text{CH}) = 1.111(7) \text{ \AA}$  and  $r_a(\text{CC}) = 1.566(3) \text{ \AA}$  by electron diffraction. If we take for cubane  $r = 1.11(2) \text{ \AA}$ , then from eq 5 and 6 we obtain  $R = 1.565(4) \text{ \AA}$  using our value for  $B_0 = 0.1122(2) \text{ cm}^{-1}$ , in excellent correspondence with  $r_a(\text{CC})$  above. A recent ab initio investigation of cubane by Almlof and Jonvik,<sup>30</sup> believed to be close to the Hartree-Fock limit, gave a C-C bond length of 1.570  $\text{\AA}$ . The latter authors commented that the bond lengths obtained by X-ray diffraction may be shorter than the true bond lengths, as neither libration nor bonding effects on the electron density distribution were explicitly considered in the X-ray study. Thus even though the X-ray value of 1.551(5)  $\text{\AA}$  for the C-C bond in the crystal<sup>4</sup> is considerably lower than our gas-phase value, we are reasonably confident in our value, particularly since it is within the range (1.555-1.566  $\text{\AA}$ ) determined for other four-membered-ring hydrocarbons.<sup>31</sup>

The influence on the C-C bond length by the assumption of the C-H bond length is relatively small since, by virtue of their lighter mass, the hydrogens have less leverage on  $I_0$ . An uncertainty in  $r$  leads to a reduced uncertainty in  $R$  by the ratio

$$dR/dr = -2y/3^{1/2}[y + (M_C/M_H)] \sim -0.142 \quad (7a)$$

(29) B. Andersen and L. Fernholt, *Acta Chem. Scand.*, **24**, 445-452 (1970).

(30) J. Almlof and T. Jonvik, *Chem. Phys. Lett.*, **92**, 267-270 (1982).

(31) J. H. Callomon, E. Hirota, K. Kuchitsu, W. J. Lafferty, A. G. Maki, and C. S. Pote, Eds., "Structure Data of Free Polyatomic Molecules", Springer-Verlag, Berlin, 1976, Landolt-Bornstein New Series, Vol. II/7.

with

$$y = 1 + (2r/3^{1/2}R) \quad (7b)$$

as obtained from eq 5 by zeroing the variation in  $I_0$ . The assumption about  $r$  could be eliminated altogether in future work on isotopic variants of cubane.  $C_8D_8$  would be an obvious possibility, but a symmetric top species such as  $C_8H_7D$ ,  $C_8HD_7$ , or  $^{13}C^{12}C_7H_8$  might in practice prove easier. The  $^{13}C$  species could be examined in natural abundance if a suitable band could be found.

It should also be mentioned that our present uncertainty in  $B_0$  from the perturbation theory fit to  $\nu_{11}$  and  $\nu_{12}$  of  $C_8H_8$  might be reduced by an analysis of the present spectrum of the  $\nu_{10}$  band. This band is so extensively perturbed that the usual concepts of "allowed" and "forbidden" transitions become meaningless, and so assignment of this band should yield numerous ground-state combination differences leading to a higher precision  $B_0$  value as well as ground-state octahedral splitting and centrifugal distortion parameters.

Registry No. Cubane, 277-10-1.

## Rotational Spectroscopy of Molecular Complexes of $BF_3$ with $NCCN$ , $CO_2$ , and $N_2O$ <sup>†</sup>

K. R. Leopold, G. T. Fraser, and W. Klemperer\*

Contribution from the Department of Chemistry, Harvard University, Cambridge, Massachusetts 02138. Received July 27, 1983

**Abstract:** The microwave spectrum of the van der Waals complex  $NCCN \cdot BF_3$  has been obtained by molecular beam electric resonance spectroscopy. This molecule is shown to be a symmetric rotor. The rotational constants for the  $^{11}B$  and  $^{10}B$  species are  $B_0 = 672.2$  (2) and  $675.6$  (2) MHz, respectively. These values are consistent with a B-N bond length of  $2.647$  (3) Å. Radio-frequency and microwave transitions of the complex  $CO_2 \cdot BF_3$  have been observed and establish that this molecule is an asymmetric rotor. One radio-frequency transition observed for  $N_2O \cdot BF_3$  suggests an asymmetric structure for this complex as well.

### Introduction

Addition complexes of the strong Lewis acid- $BF_3$  are fundamental in classical donor-acceptor chemistry. Recently the van der Waals complexes of  $BF_3$  with Ar, CO, and  $N_2$  have been formed in an adiabatic expansion and structurally characterized by their microwave spectra.<sup>1</sup> The van der Waals bonds form along the  $C_3$  axis of the  $BF_3$  in a manner analogous to the binding in the "classical complexes". This behavior seems chemically reasonable and suggests that the van der Waals interaction may also be viewed in terms of a Lewis acid-base model. Such a picture has successfully provided an intuitive understanding of van der Waals interactions in a variety of weakly bound systems.

In view of the chemical character of weakly bound systems it seems reasonable to consider whether, in fact, there is a smooth transition between van der Waals and covalent bonding. Despite structural similarities between the covalent and van der Waals complexes of  $BF_3$ , for example, dramatic differences exist. The B-N bond length in  $BF_3$ -amines is  $1.6$  Å while that in  $N_2 \cdot BF_3$  is  $2.9$  Å. Such a large variation in bond length with nitrogen donor may permit, at least in principle, the observation of a smooth transition between "covalent" and "van der Waals" binding, provided the nature of the nitrogen donor were properly chosen. Two classes of nitrogen donors which may be used for this purpose are the amines and the cyanides. Microwave data for a number of the gas-phase amines are available, but the analogous data for the cyanides are lacking. With the simplest of these, HCN and  $CH_3CN$ ,  $BF_3$  forms solid compounds. Cyanogen, on the other hand, does not form a solid with  $BF_3$ , and so was chosen as an interesting system to study.

The structure of  $NCCN \cdot BF_3$  is also of interest for comparison with a number of other previously studied "van der Waals" molecules. Recent rotational spectroscopic studies of  $CO_2 \cdot HF^2$  and  $NCCN \cdot HF$ ,<sup>3</sup> as well as  $SCO \cdot HF^2$  and  $CO_2 \cdot HCl$ ,<sup>4</sup> have shown these complexes to have linear, hydrogen bonded structures. The linear structures of the  $CO_2 \cdot HX$  complexes are not readily un-

derstandable from consideration of oxygen lone pairs on  $CO_2$ , as might be expected if chemical reasoning were applicable. Moreover, the geometry of the closely related, isoelectronic species  $N_2O \cdot HF$  is quite different, having a bent, hydrogen bonded structure. In view of the similarities between  $CO_2$  and  $N_2O$  and the dissimilarities in their complexes with HF, a simple HOMO-LUMO approach does not appear applicable to these systems, since any argument based on lone pairs would have to predict the same structure for  $N_2O \cdot HF$  and  $CO_2 \cdot HF$ . In contrast, HF and  $BF_3$  appear quite similar in their weak interactions. For example, studies of the complexes of CO and  $N_2$  with HF and  $BF_3$  show that HF and  $BF_3$  behave similarly as simple Lewis acids. Specifically, viewed as electron-pair acceptors, both HF and  $BF_3$  accept electrons along their symmetry axes. Thus, given the observed behavior of  $CO_2$  with HF, the complex  $CO_2 \cdot BF_3$  is expected to be a symmetric top in which the  $CO_2$  axis is coincident with the symmetry axis of the  $BF_3$ . Likewise, the linear structure of  $NCCN \cdot HF$  suggests a similar geometry for  $NCCN \cdot BF_3$ .

We have studied the complexes  $NCCN \cdot BF_3$ ,  $CO_2 \cdot BF_3$ , and  $N_2O \cdot BF_3$  by rotational spectroscopy using the molecular beam electric resonance technique. The spectra show that the complex  $NCCN \cdot BF_3$  does indeed have the anticipated  $C_{3v}$  structure, while  $CO_2 \cdot BF_3$  and  $N_2O \cdot BF_3$  do not. This result further emphasizes the complexity of the binding in  $CO_2$  and  $N_2O$  systems.

### Experimental

A beam of  $NCCN \cdot BF_3$  was formed by expanding a mixture of 1%  $BF_3$  and 25%  $NCCN$  in Ar through a  $25\text{-}\mu\text{m}$  nozzle at room temperature. The stagnation pressure was typically 2.5 atm. Under these conditions,

(1) K. C. Janda, L. S. Bernstein, J. M. Steed, S. E. Novick, and W. Klemperer, *J. Am. Chem. Soc.*, **100**, 8074 (1978).

(2) F. A. Baiocchi, T. A. Dixon, C. H. Joyner, and W. Klemperer, *J. Chem. Phys.*, **74**, 6544 (1981).

(3) A. C. Legon, P. D. Soper, and W. H. Flygare, *J. Chem. Phys.*, **74**, 4936 (1981).

(4) R. S. Altman, T. A. Dixon, M. D. Marshall, and W. Klemperer, *J. Chem. Phys.*, **77**, 4344 (1982).

<sup>†</sup> This work was supported by the National Science Foundation.