

Lecture 24 *Relawavity* Introduction Part 2

Thursday 4.07.2016

Relawavity: Relativistic wave mechanics II. 2nd-order effects

(Unit 3 4.05.16)

- Review of Doppler-shift and Rapidity ρ_{AB} calculation: *Galileo's Revenge Part I Lect. 23 p.64-75*
 - Relating rapidity ρ_{AB} and relativity velocity parameter $\beta_{AB}=u_{AB}/c$
- Review of $\frac{1}{2}$ -sum- $\frac{1}{2}$ -difference Phase and Group factors giving relativistic space-axes and time-axes
 - Colliding-CW space-time (x, ct) -graph vs Colliding PW space-time (R, L) -baseball diamond
- Review of $\frac{1}{2}$ -sum- $\frac{1}{2}$ -difference of phasor angular velocity: *Galileo's Revenge Part II* (Pirelli site)
 - Elementary models: 2-comb Moire' patterns and cosine-law constructions
- Bob, Alice, and Carla combine Doppler shifted $\frac{1}{2}$ -sum- $\frac{1}{2}$ -difference Phase and Group factors
 - Doppler shifted Phase vector \mathbf{P}' and Group vector \mathbf{G}' in per-space-time Minkowski coordinate grid in space-time
 - Animations that compare Doppler shifted colliding CW with colliding PW
- The 16 parameters of Doppler-shifted 2-CW Minkowski geometry
 - Doppler shifted Phase parameters
 - Doppler shifted Group parameters
 - Lorentz transformation matrix and Two Famous-Name Coefficients
- Thales Mean Geometry (*Thales of Miletus 624-543 BCE*) and its role in Relawavity
 - Detailed geometric construction of relawavity plot for 1-octave Doppler ($\beta_{AB}=u_{AB}/c=3/5$)
- Stellar aberration and the Epstein approach to SR

Review Doppler-shift and Rapidity calculation

ALICE'S
LASER
GAUNTLET



Alice: Hey, Bob and Carla! Read off your Doppler shift ratios $\langle B|A \rangle$ and $\langle C|A \rangle$ to my 600THz beam.

Also, rapidity ρ_{BA} and ρ_{CA} relative to me.

Now, Carla, what's your rapidity ρ_{CB} relative to Bob?



Doppler ratio:

$$\langle R|S \rangle = \frac{v_{RECEIVER}}{v_{SOURCE}}$$

rapidity:

$$\rho_{RS} = \log_e \langle R|S \rangle$$

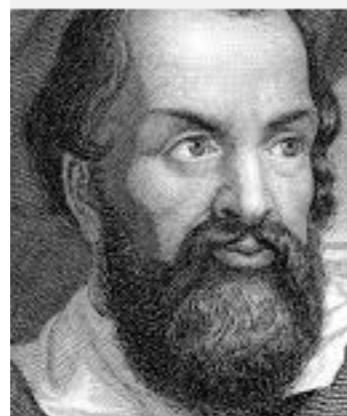
or:

$$\langle R|S \rangle = e^{\rho_{RS}}$$

Definition of Rapidity

$$\langle B|A \rangle = \frac{v_B}{v_A}$$

Happy now, Galileo?



is time-reversed

$$\langle A|B \rangle = \frac{v_A}{v_B}$$

I got $\langle C|B \rangle = \langle C|A \rangle \langle A|B \rangle = (2/3)(1/2) = 1/3$,
and $\rho_{CB} = \rho_{CA} + \rho_{AB} = -1.10$

We're in Splitsville!

Carla: I see Doppler Red shift to 400THz

I got $\langle C|A \rangle = 2/3$,
and $\rho_{CA} = \ln(2/3) = -0.41$

Bob: I see Doppler Blue shift to 1200THz

I got $\langle B|A \rangle = 2$,
and $\rho_{BA} = \ln(2) = +0.69$

Bob-Alice Doppler ratio:

$$\langle B|A \rangle = \frac{v_B}{v_A} = \frac{1200}{600} = \frac{2}{1}$$

Bob-Alice rapidity:

$$\rho_{BA} = \log_e \langle B|A \rangle = \log_e \frac{2}{1}$$

$$\rho_{BA} = 0.69 \quad (\text{so: } \rho_{AB} = -0.69)$$

Carla-Alice Doppler ratio:

$$\langle C|A \rangle = \frac{v_C}{v_A} = \frac{400}{600} = \frac{2}{3}$$

Carla-Alice rapidity:

$$\rho_{CA} = \log_e \langle C|A \rangle = \log_e \frac{2}{3}$$

$$\rho_{CA} = -0.41$$

Carla-Bob Doppler ratio:

$$\langle C|B \rangle = \frac{v_C}{v_B} = \frac{v_C}{v_A} \frac{v_A}{v_B} = \langle C|A \rangle \langle A|B \rangle$$

Carla-Bob rapidity:

$$e^{\rho_{CB}} = e^{\rho_{CA}} e^{\rho_{AB}} \text{ implies:}$$

Galileo's Revenge (part 1)

Rapidity adds just like
Galilean velocity

$$\rho_{CB} = \rho_{CA} + \rho_{AB} = -0.41 - 0.69 = -1.10$$

Review of Doppler-shift and Rapidity ρ_{AB} calculation: *Galileo's Revenge Part I Lect. 23 p.64-75*

→ Relating rapidity ρ_{AB} and relativity velocity parameter $\beta_{AB}=u_{AB}/c$



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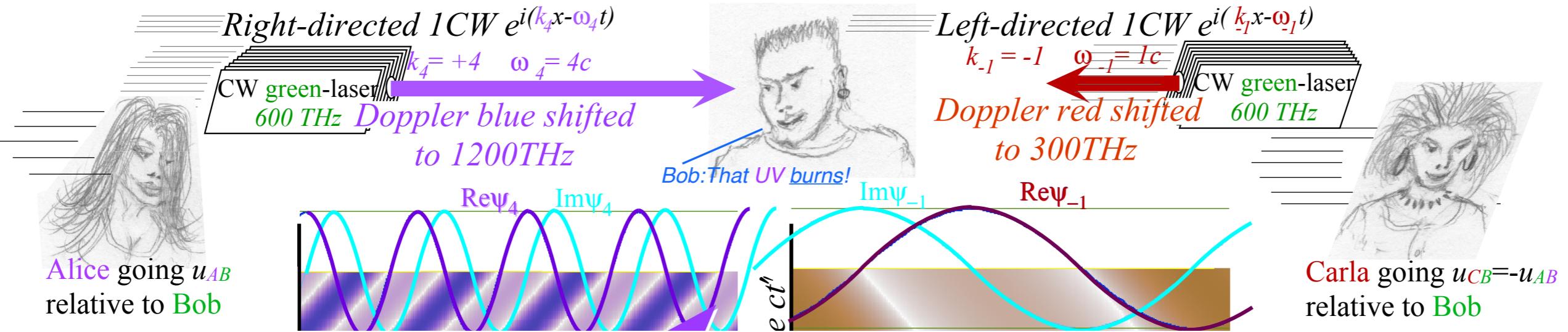
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More at Pirelli Challenge page: [*'Un Grande Affare' - Light Meets Light*](#)

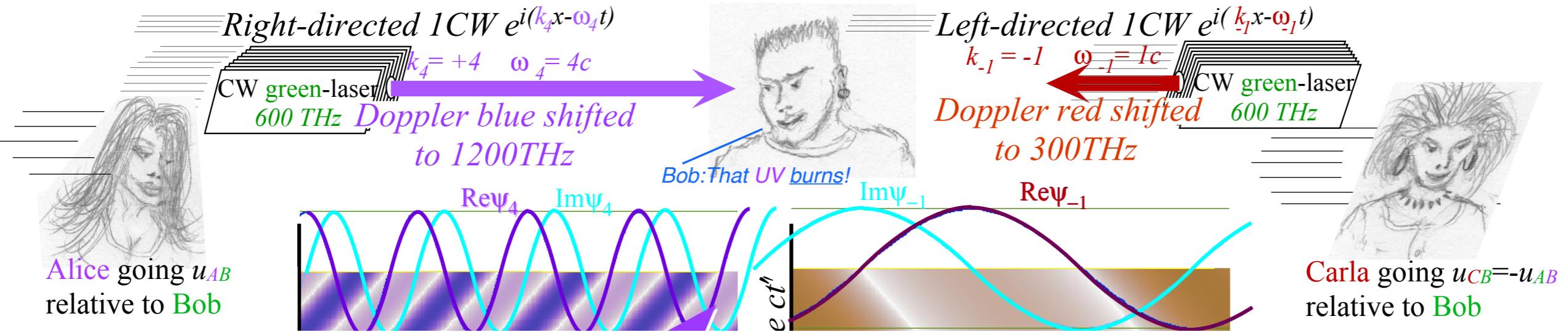
Relating rapidity ρ_{AB} and relativity velocity parameter $\beta_{AB}=u_{AB}/c$

Imagine Bob sees a pair of counter-propagating laser beams with wavevectors $k_R=+\omega_R/c$ and $k_L=-\omega_L/c$ $\omega_R=\omega_A$ going left-to-right (from Alice's 600 THz laser) and $\omega_L=\omega_C$ going right-to-left (from Carla's 600 THz laser).



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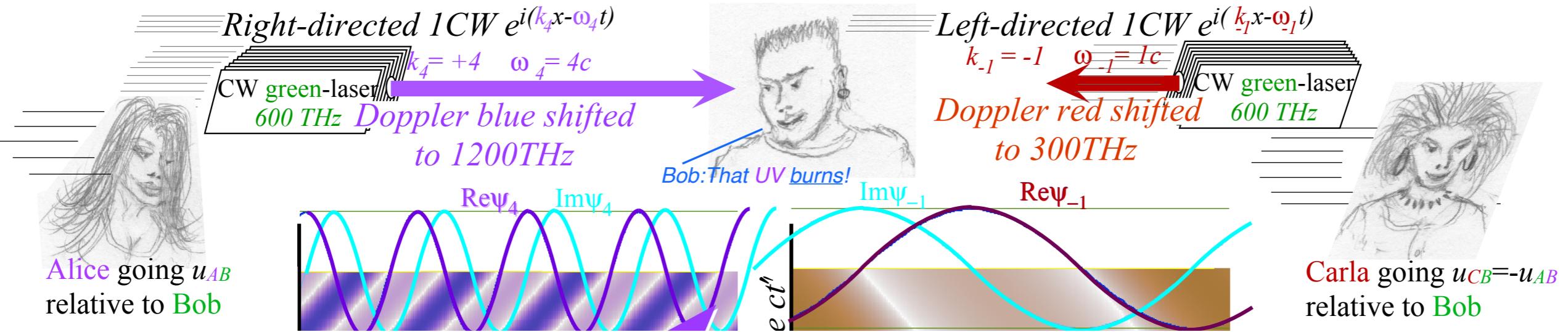


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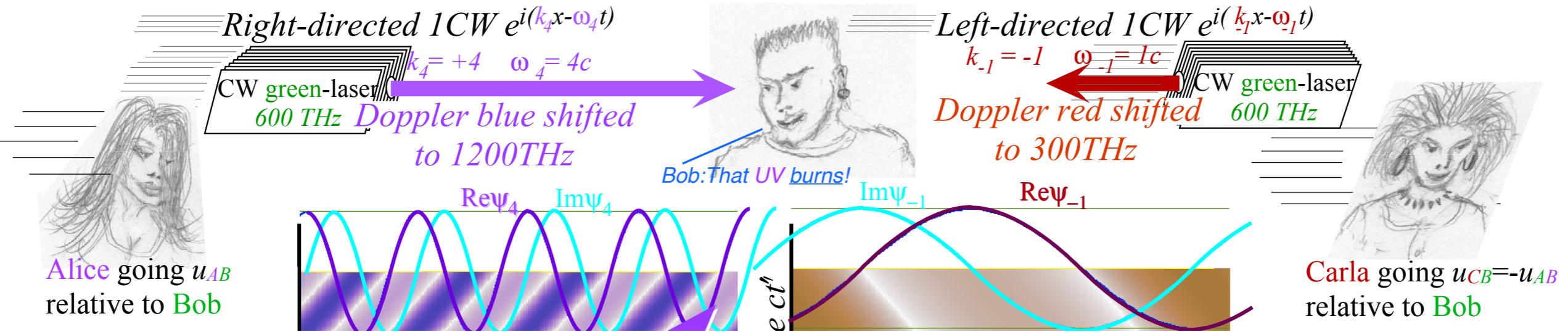
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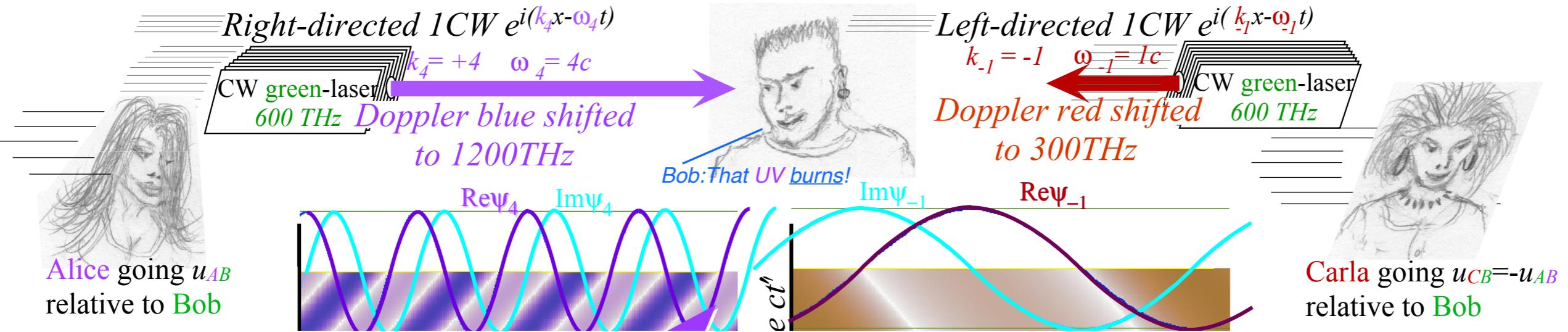
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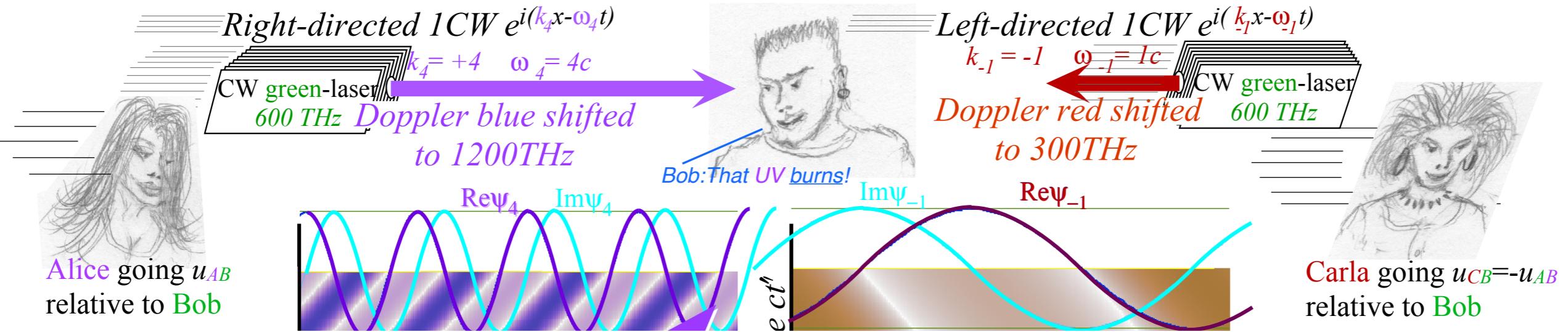
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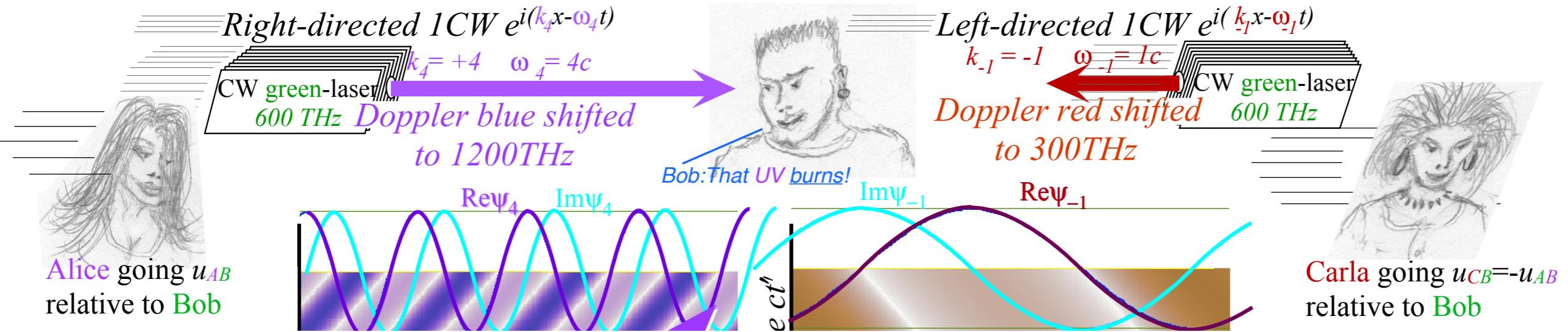
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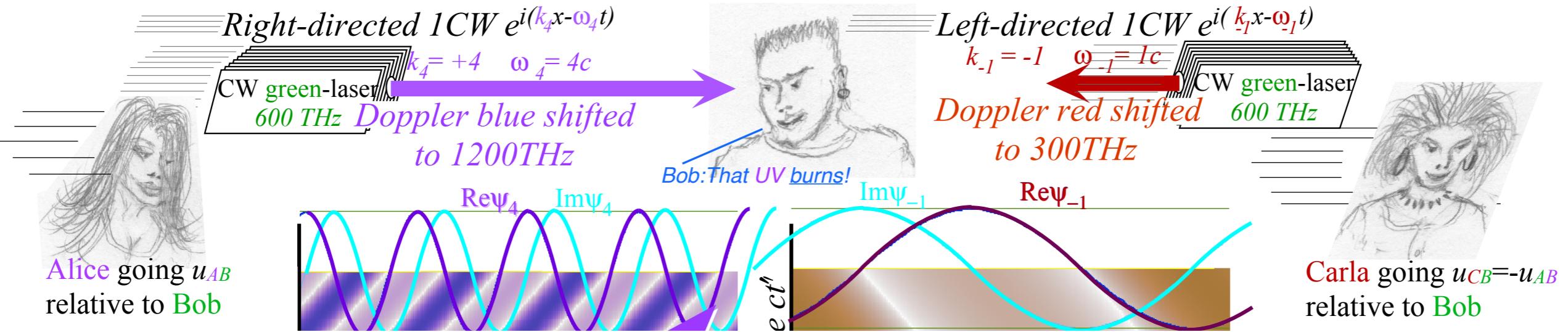
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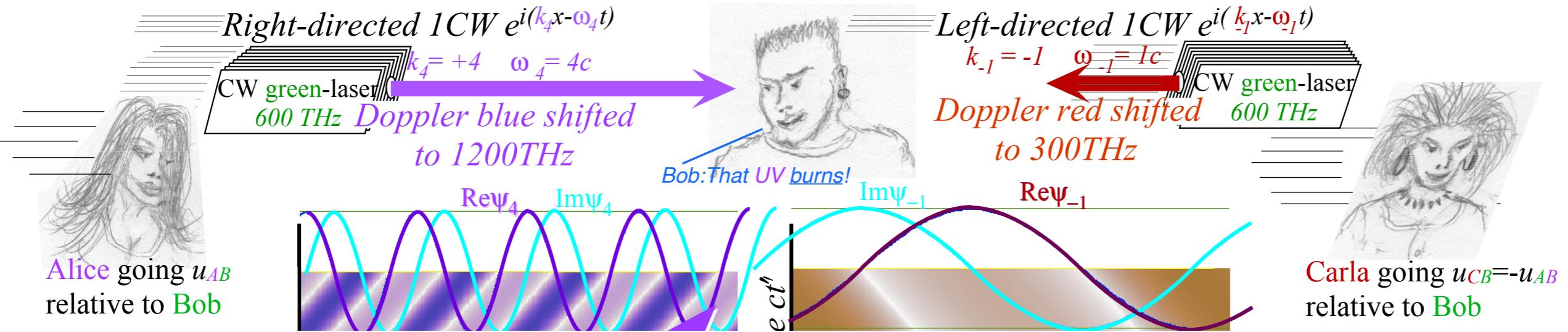
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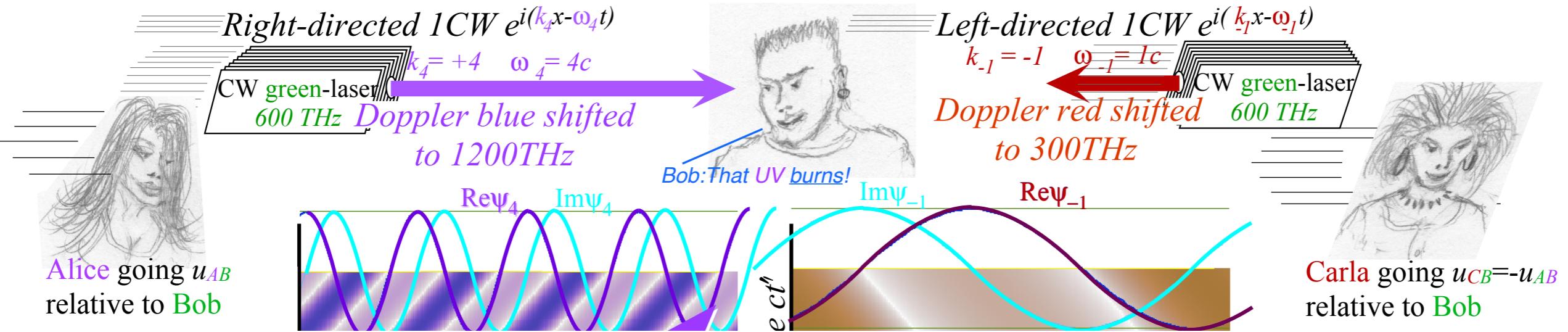
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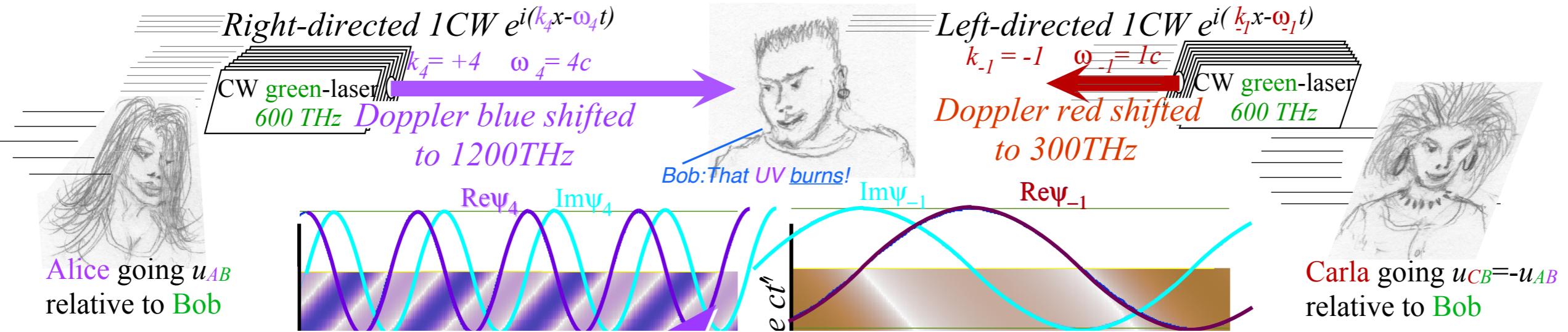
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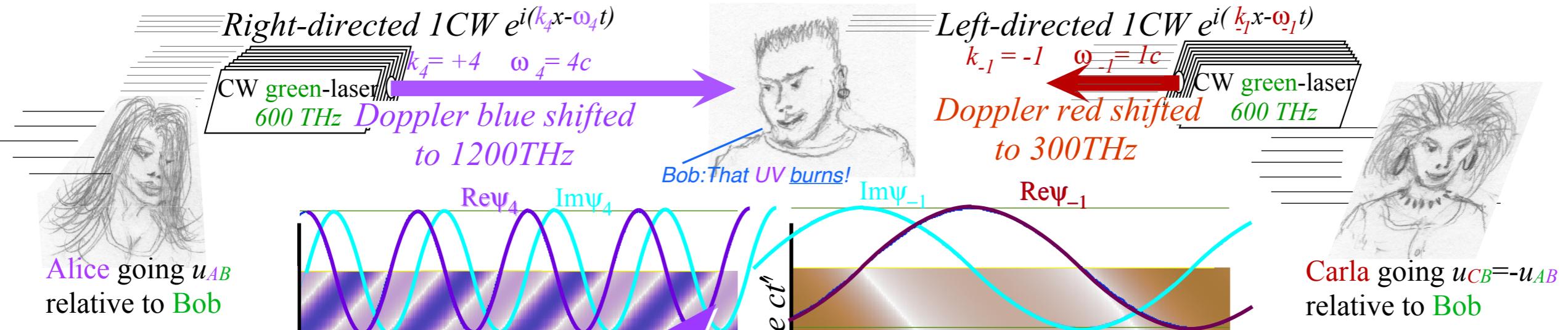
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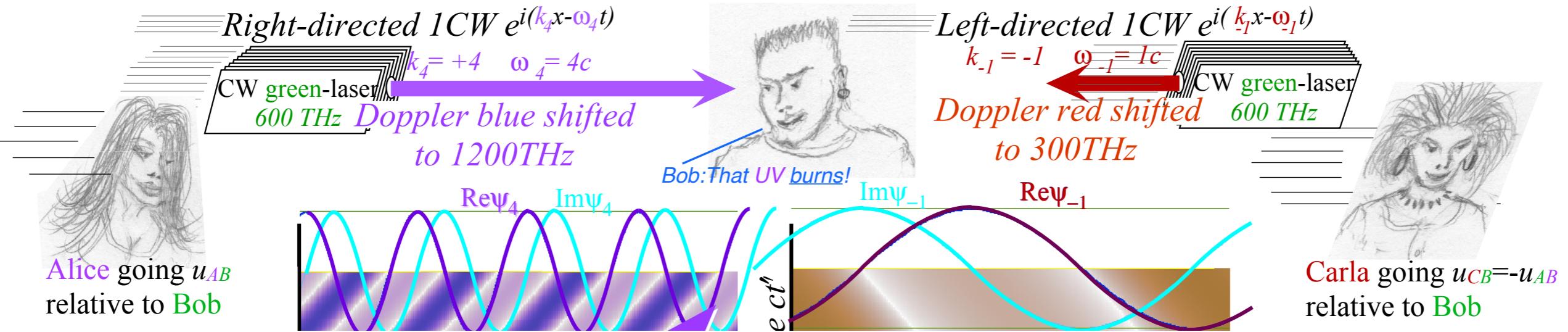
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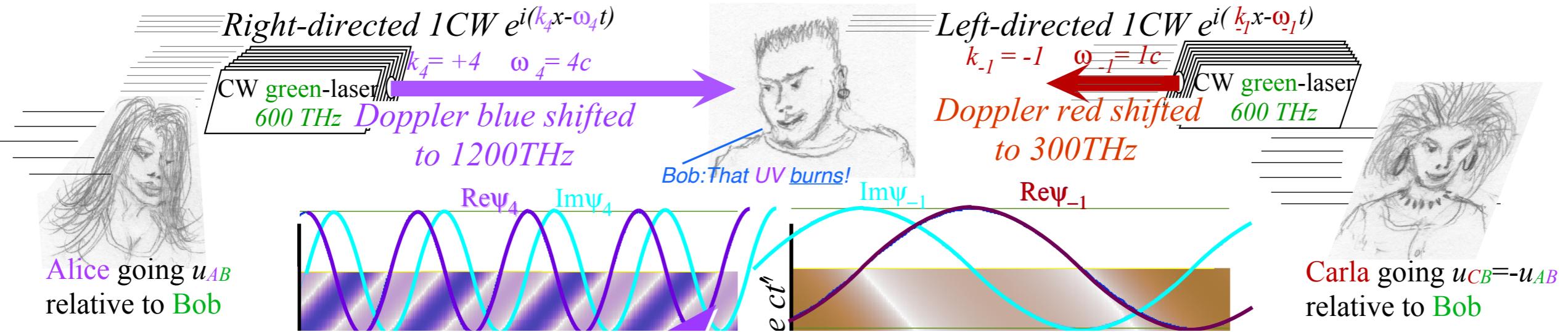
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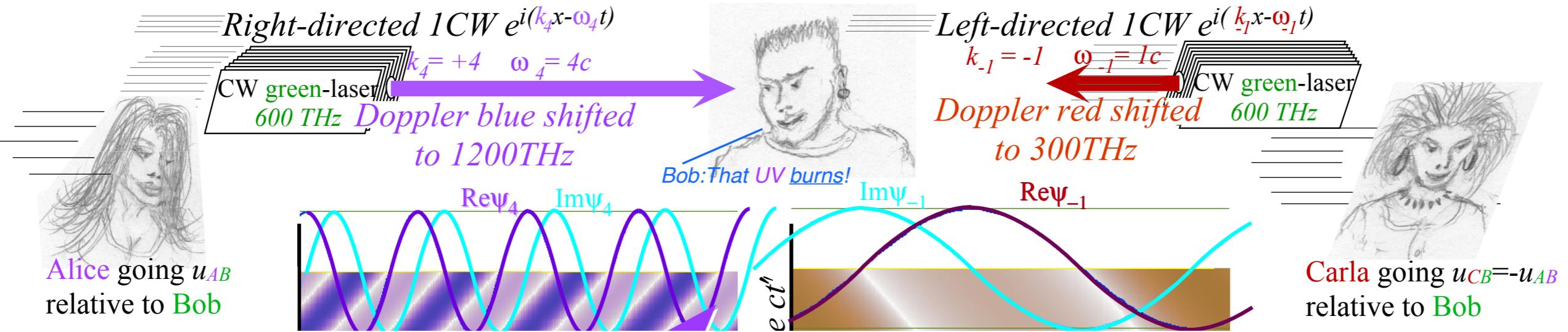
Given:

$$\omega_R = e^{\rho_{AB}} \omega_{600}$$

$$\text{and: } \omega_L = e^{\rho_{CB}} \omega_{600} = e^{-\rho_{AB}} \omega_{600}$$

Relating rapidity ρ_{AB} and relativity velocity parameter $\beta_{AB}=u_{AB}/c$

Imagine Bob sees a pair of counter-propagating laser beams with wavevectors $k_R=+\omega_R/c$ and $k_L=-\omega_L/c$ $\omega_R=\omega_A$ going left-to-right (from Alice's 600 THz laser) and $\omega_L=\omega_C$ going right-to-left (from Carla's 600 THz laser).



We ask two questions:

- (1.) To what velocity u_E must Bob accelerate so he sees beams with equal frequency ω_E ?
- (2.) What is that frequency ω_E ?

Reply to Query (1.) has a *Jeopardy*-style answer-by-question:

What is the beam group velocity?

$$\text{Given: } \omega_{group} = \frac{\omega_R - \omega_L}{2} \text{ and: } k_{group} = \frac{k_R - k_L}{2}$$

$$u_E = V_{group} = \frac{\omega_{group}}{k_{group}} = \frac{\omega_R - \omega_L}{k_R - k_L} = c \frac{\omega_R - \omega_L}{\omega_R + \omega_L} = c \frac{1200 - 300}{1200 + 300} = \frac{3}{5} c$$

with $k_R=+\omega_R/c$ and $k_L=-\omega_L/c$

$$\frac{u_E}{c} = \frac{u_{AB}}{c} = \frac{e^{\rho_{AB}} - e^{-\rho_{AB}}}{e^{\rho_{AB}} + e^{-\rho_{AB}}} = \frac{\sinh \rho_{AB}}{\cosh \rho_{AB}} = \tanh \rho_{AB} = \frac{3}{5}$$

Using Rapidity:
 $\rho_{AB} = \log_e \langle A | B \rangle$

Given: $\omega_R = e^{\rho_{AB}} \omega_{600}$
 and: $\omega_L = e^{\rho_{CB}} \omega_{600} = e^{-\rho_{AB}} \omega_{600}$

Reply to Query (2.) in similar style:

What ω_E is blue-shift $b\omega_L$ of ω_L and red-shift ω_R/b of ω_R ? Blue-shift $b = e^{\rho_{AB}}$ Red-shift $r = b^{-1} = e^{-\rho_{AB}}$

$$\omega_E = b\omega_L = \omega_R/b \Rightarrow b = \sqrt{\omega_R/\omega_L} \Rightarrow \omega_E = \sqrt{\omega_R \cdot \omega_L} = \sqrt{1200 \cdot 300} = 600 \text{ THz}$$

(Geometric Mean)

Review of Doppler-shift and Rapidity ρ_{AB} calculation: *Galileo's Revenge Part I Lect. 23 p.64-75*

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More at Pirelli Challenge page: [*'Un Grande Affare' - Light Meets Light*](#)



Alice: OK, Bob.
We're gonna' hit
you from both
sides, now!

Colliding 2CW laser beams



Bob: Yikes!



Carla:
Look out, Bob!

Right-moving wave $e^{i(kx-\omega t)}$

CW Dye-laser
600 THz

Alice's laser

$$k = +2$$

$$\omega = 2c$$



Bob: Yikes!

Left-moving wave $e^{i(-kx-\omega t)}$

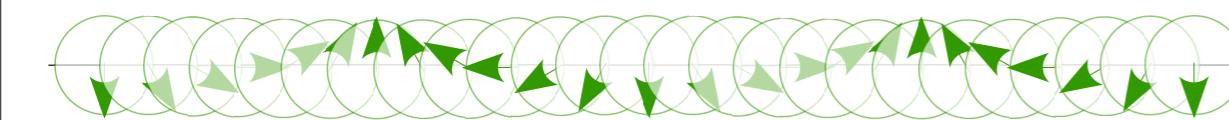
CW Dye-laser
600 THz

Carla's laser

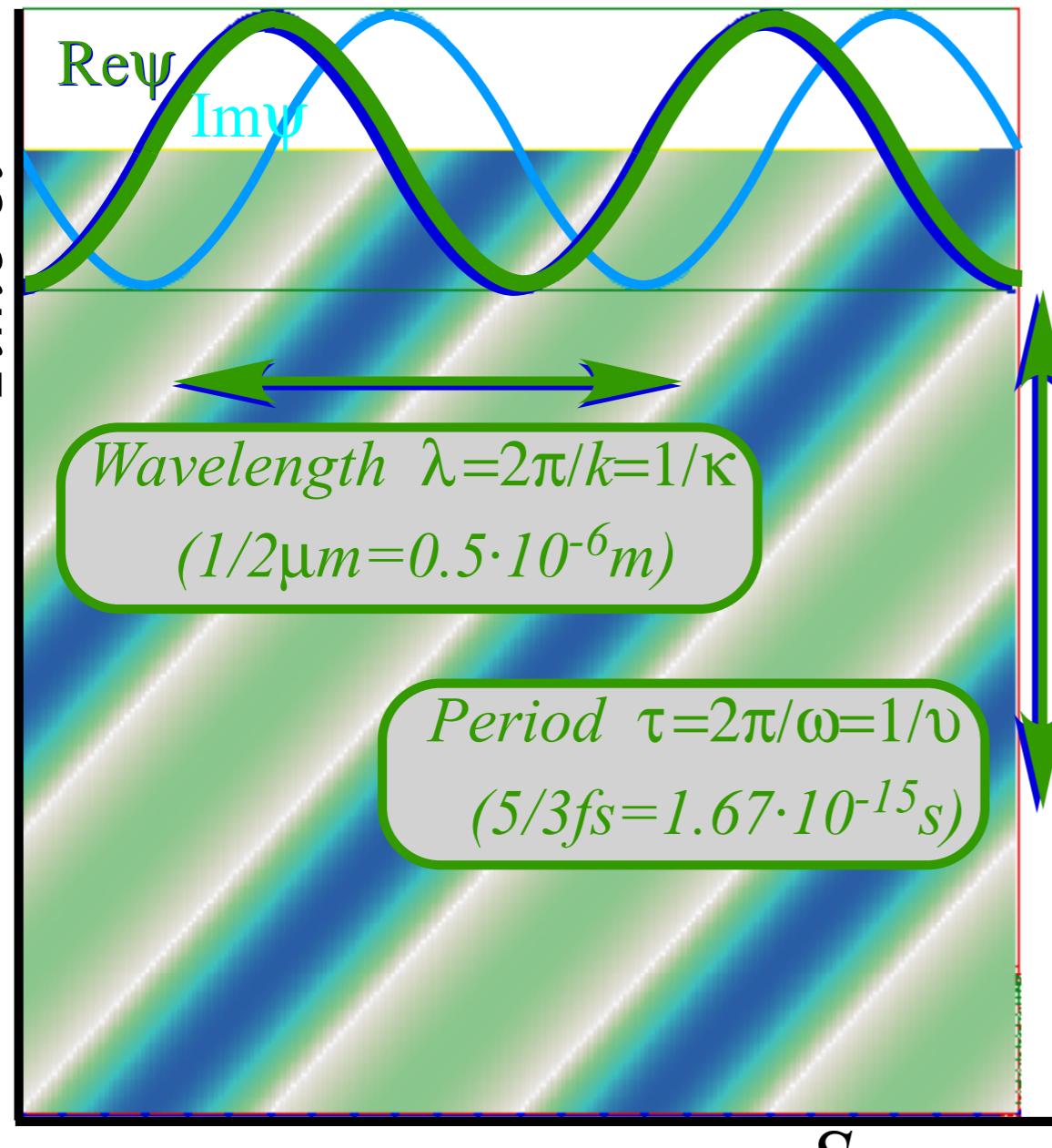
$$k = -2$$

$$\omega = 2c$$

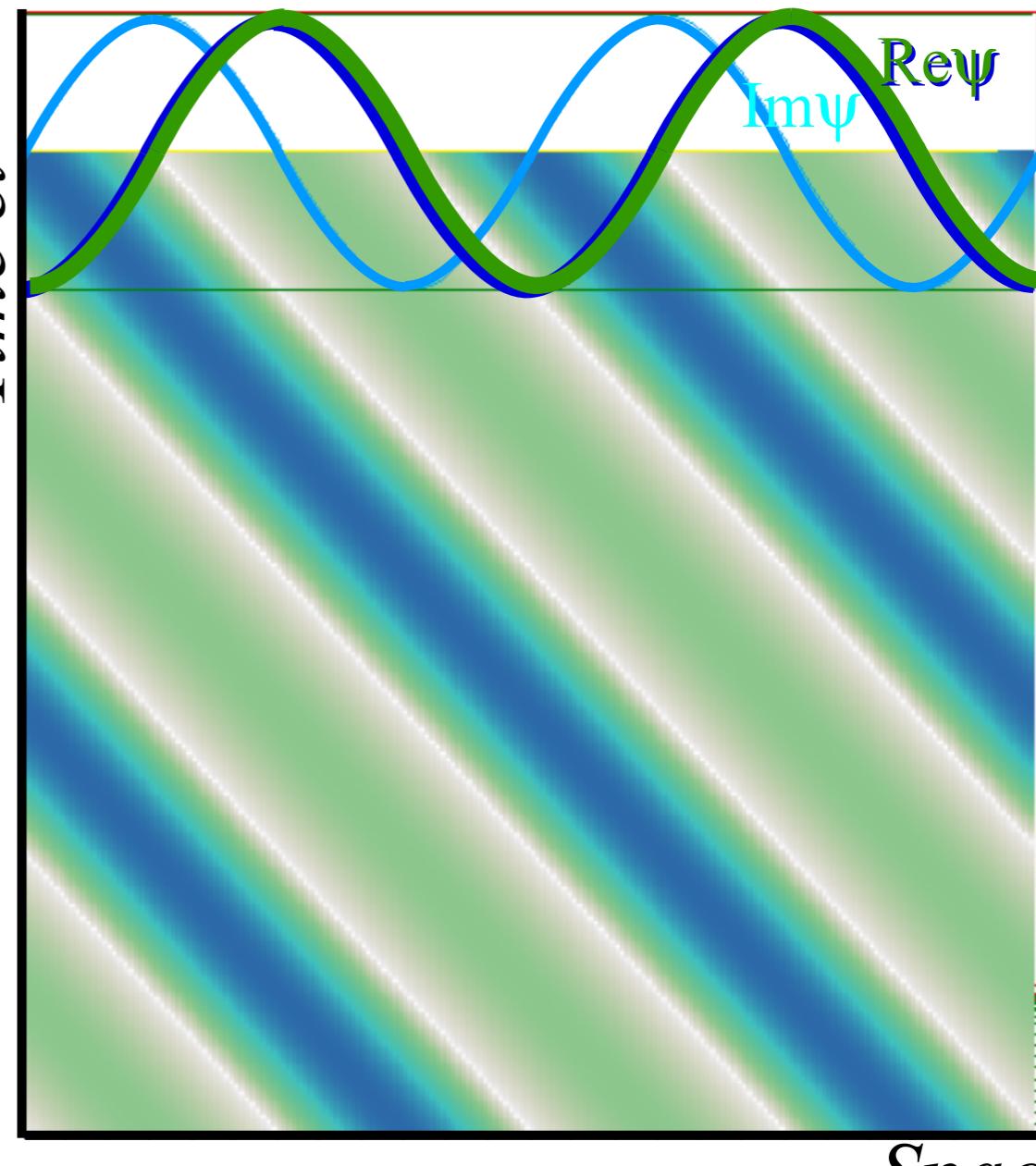
$\text{Re}\psi$
 $\text{Im}\psi$



Time ct

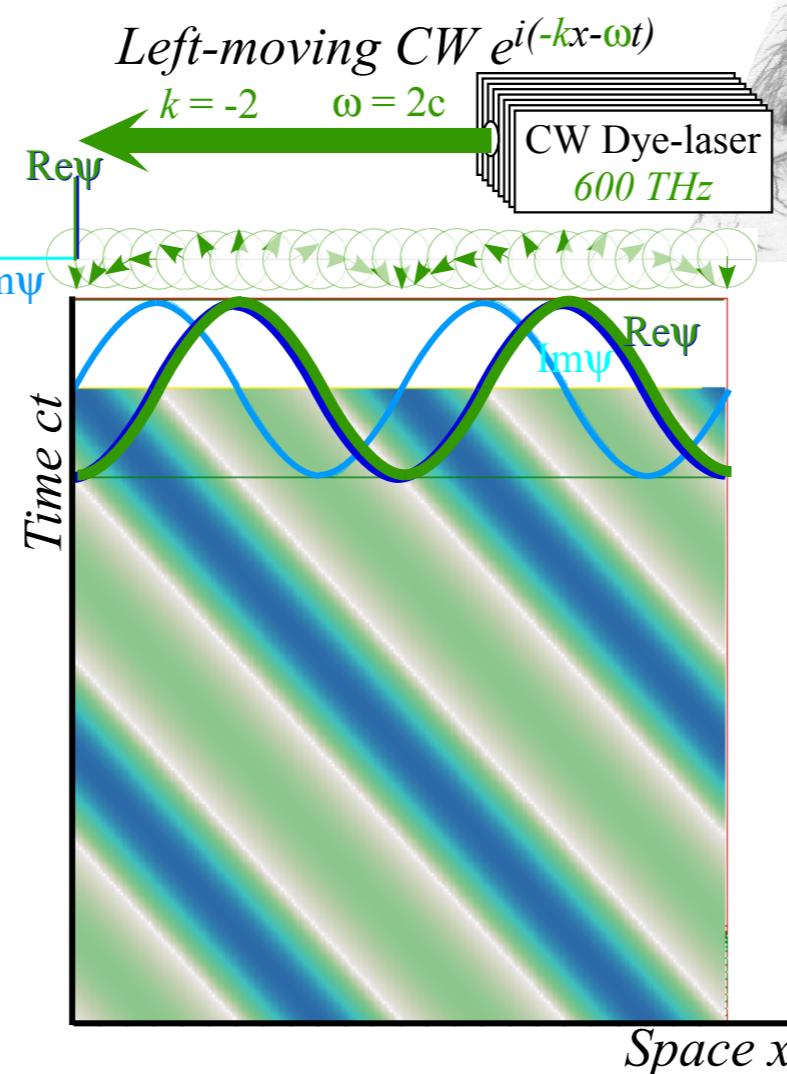
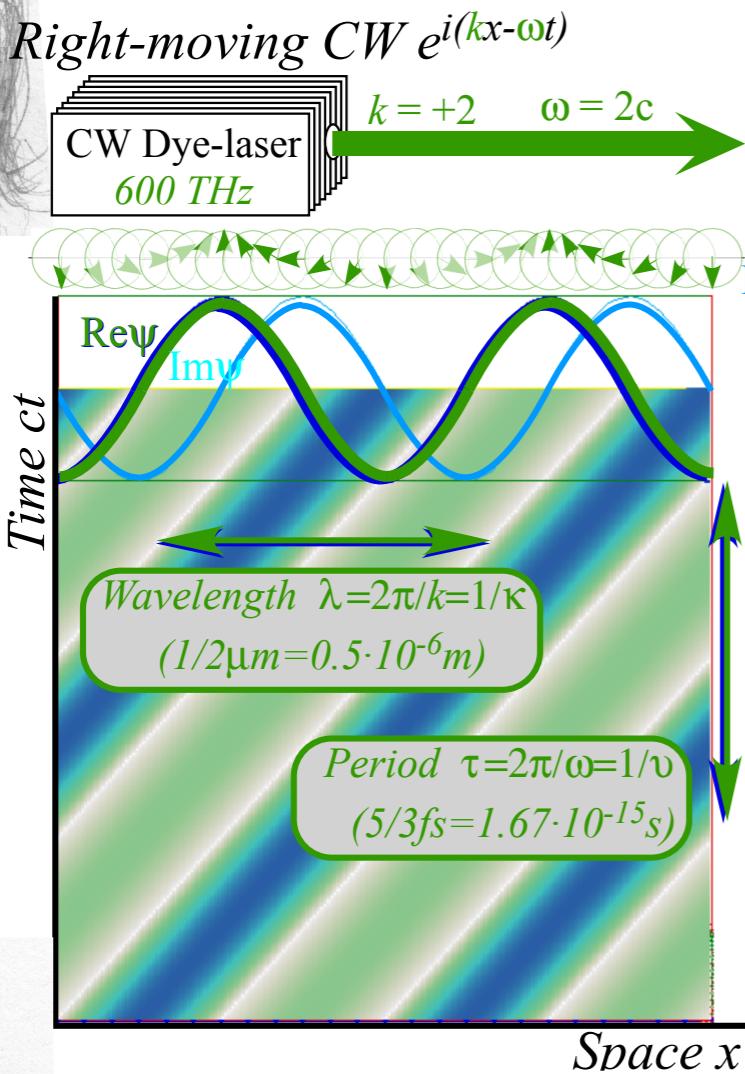


Time ct



Space x

Space x

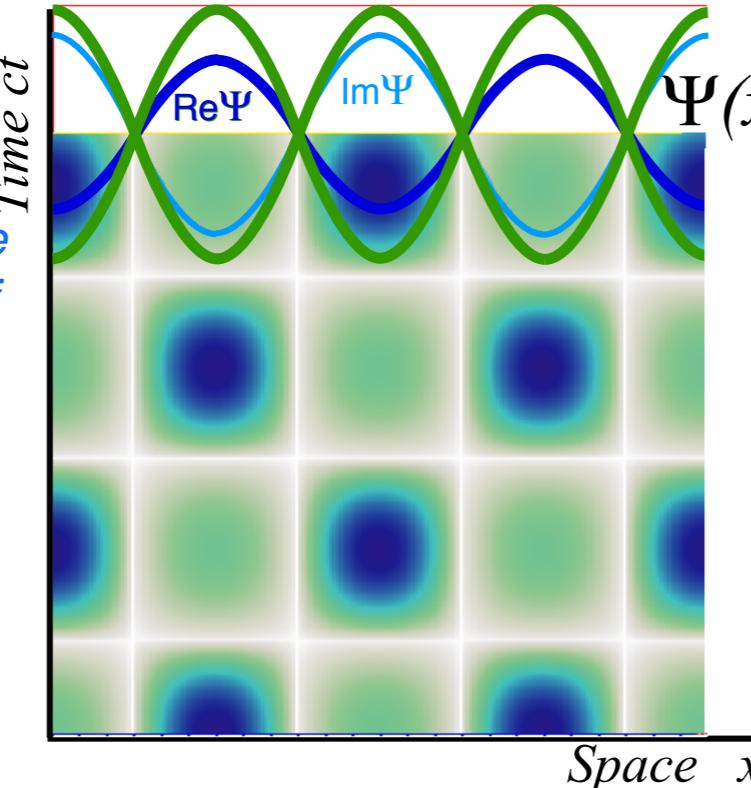


Carla:

Easy!

You get zeros of any wave-sum $e^{ia} + e^{ib}$ by factoring it into phase and group parts.

Bob:
Cool!
You guys
made me
a space-time
graph out of
real zeros.

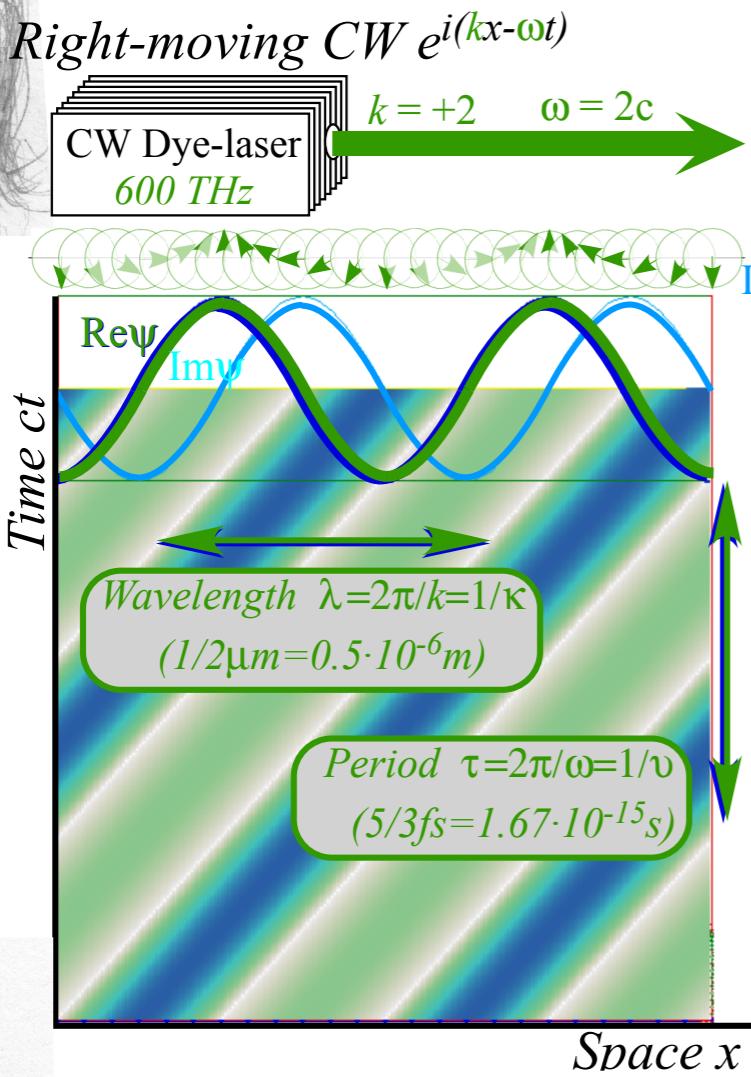


How'd it
do that?

$$\Psi(x, t) = e^{i a} + e^{i b}$$

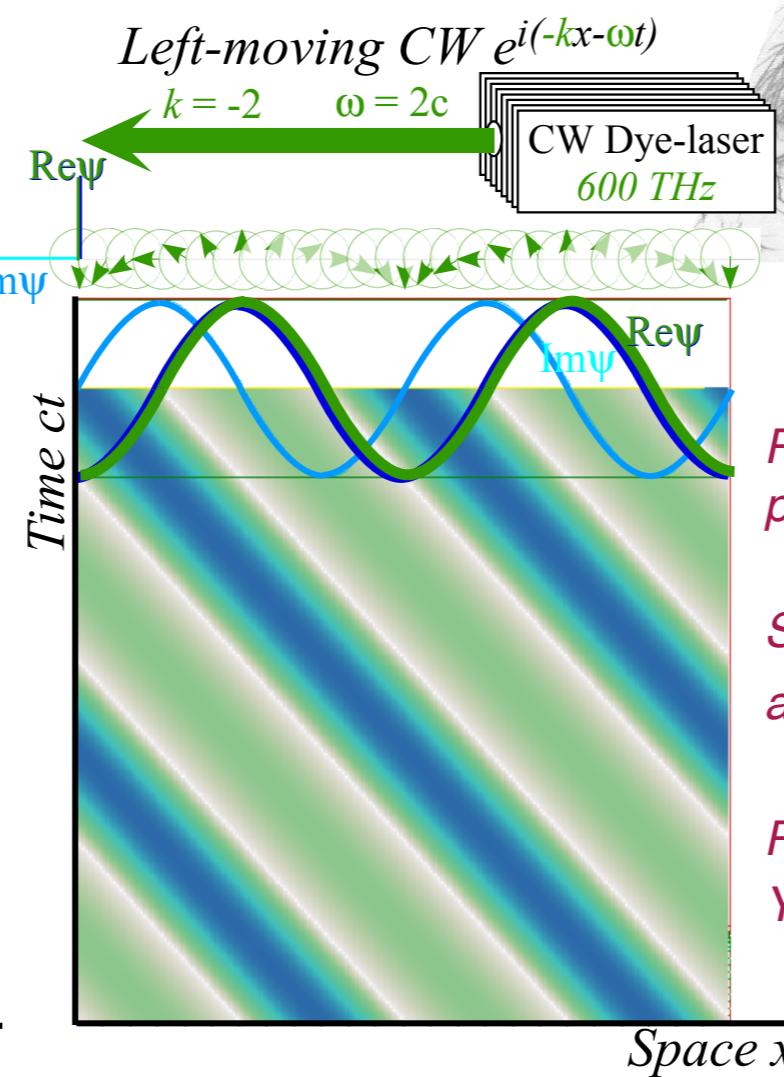
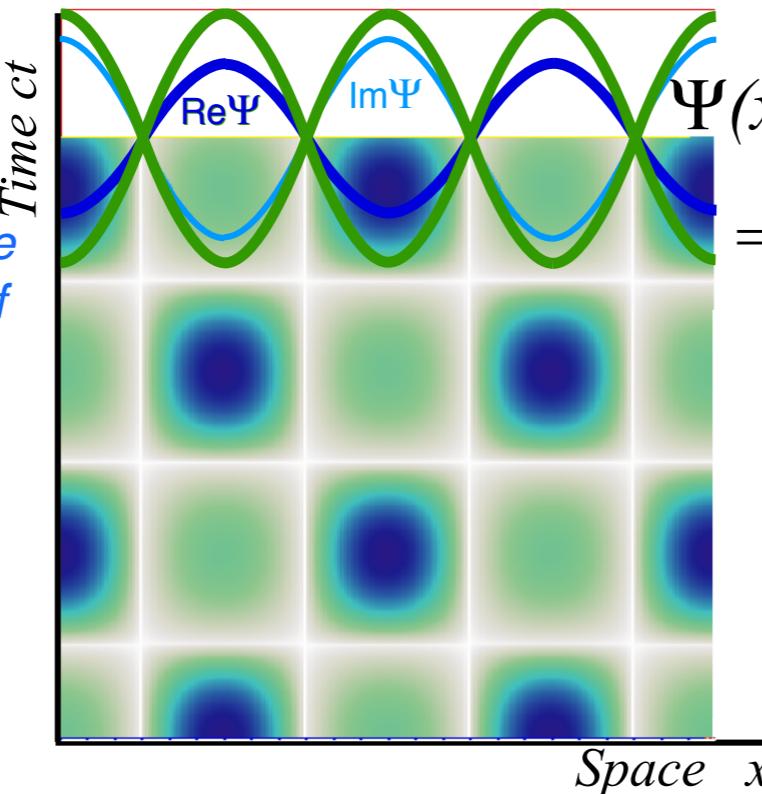
[BohrIt Web Simulation](#)
[1 CW ct vs x Plot \(\$ck = +1\$ \)](#)
[Single panel with Zero Tracers](#)

[BohrIt Web Simulation](#)
[2 CW ct vs x Plot \(\$ck = \pm 2\$ \)](#)
[Multi-panel with Zero Tracers](#)



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Easy!
You get zeros of any wave-sum $e^{ia} + e^{ib}$
by factoring it into phase and group parts.

Remember your algebra? Exponents of
products add.

So, half-sum $\frac{a+b}{2}$ plus half-diff $\frac{a-b}{2}$ gives a ,
and half-sum $\frac{a+b}{2}$ minus half-diff $\frac{a-b}{2}$ gives b .

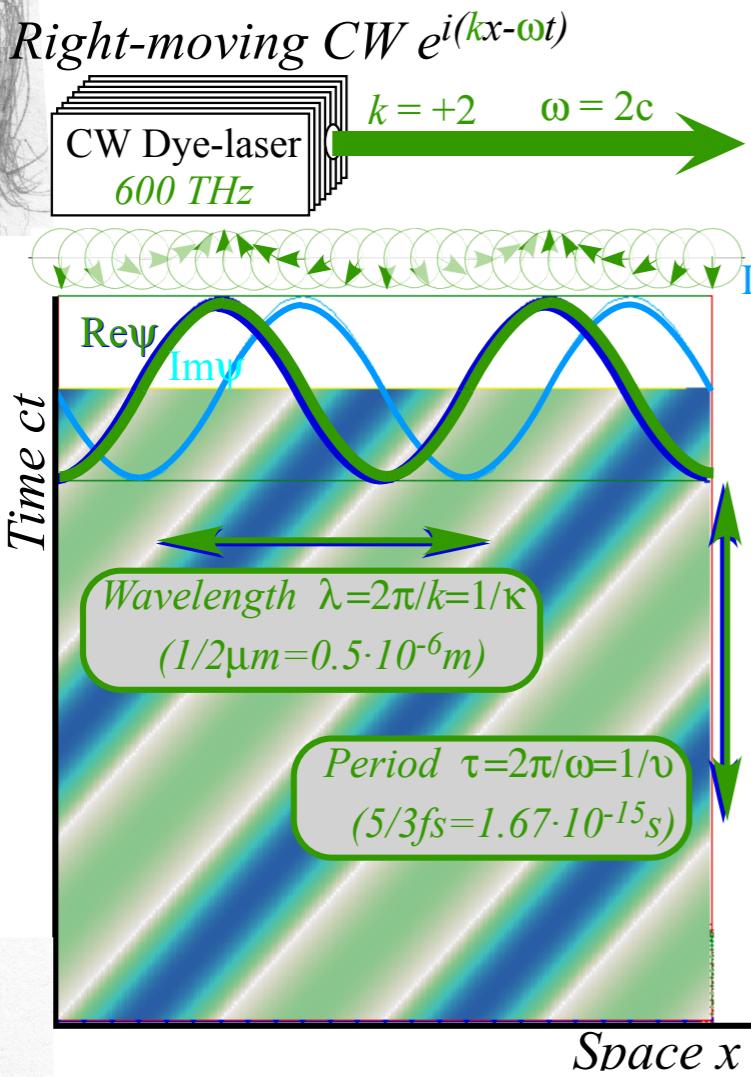
Presto!
You factor $e^{ia} + e^{ib}$ into $e^{i\frac{a+b}{2}} \left(e^{i\frac{a-b}{2}} + e^{-i\frac{a-b}{2}} \right)$

Alice 1CW phase: $a = kx - \omega t$

Carla 1CW phase: $b = -kx - \omega t$

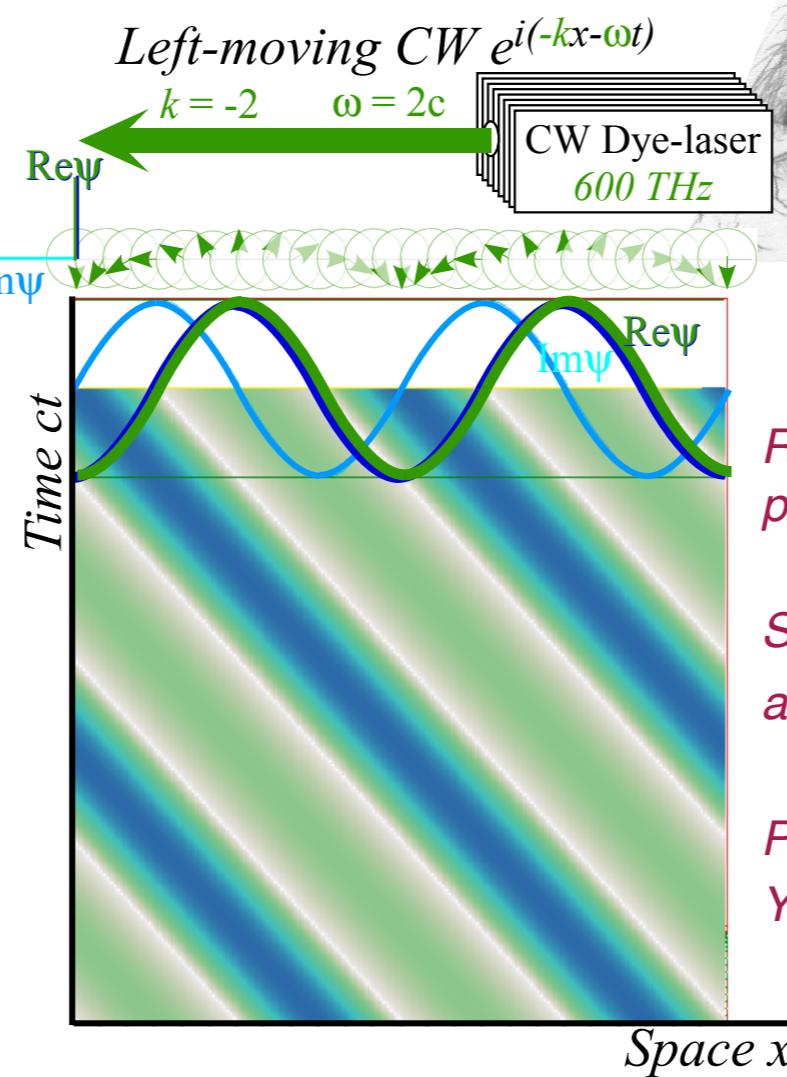
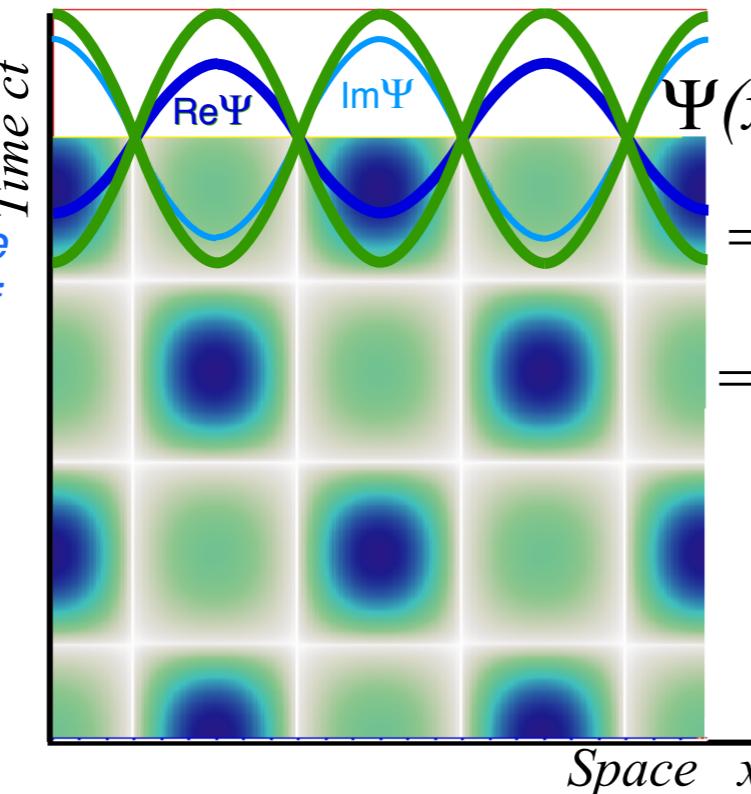
$$\Psi(x, t) = e^{ia} + e^{ib}$$

$$= e^{i\frac{a+b}{2}} \left(e^{i\frac{a-b}{2}} + e^{-i\frac{a-b}{2}} \right)$$



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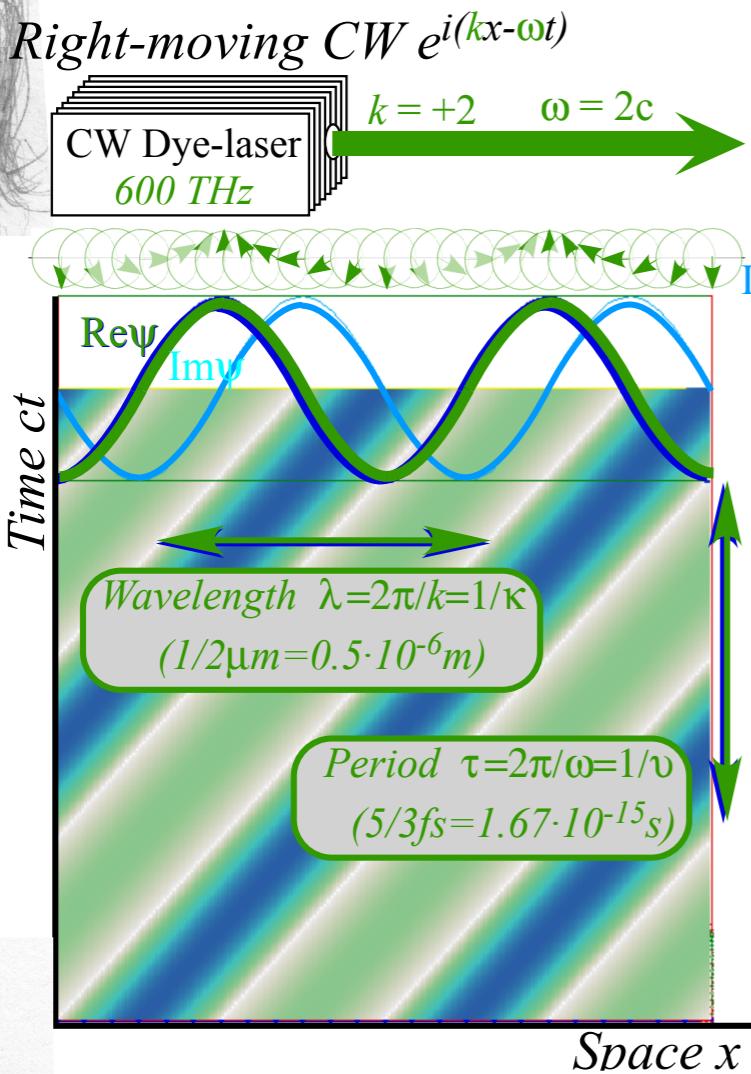
Carla 1CW phase: $b = -kx - \omega t$

Bob's 2CW Group-phase: $+k = \frac{a-b}{2}$

Group wave: $e^{-ikx} + e^{-ikx} = 2\cos kx$

is standing wave (does not vary with time t)

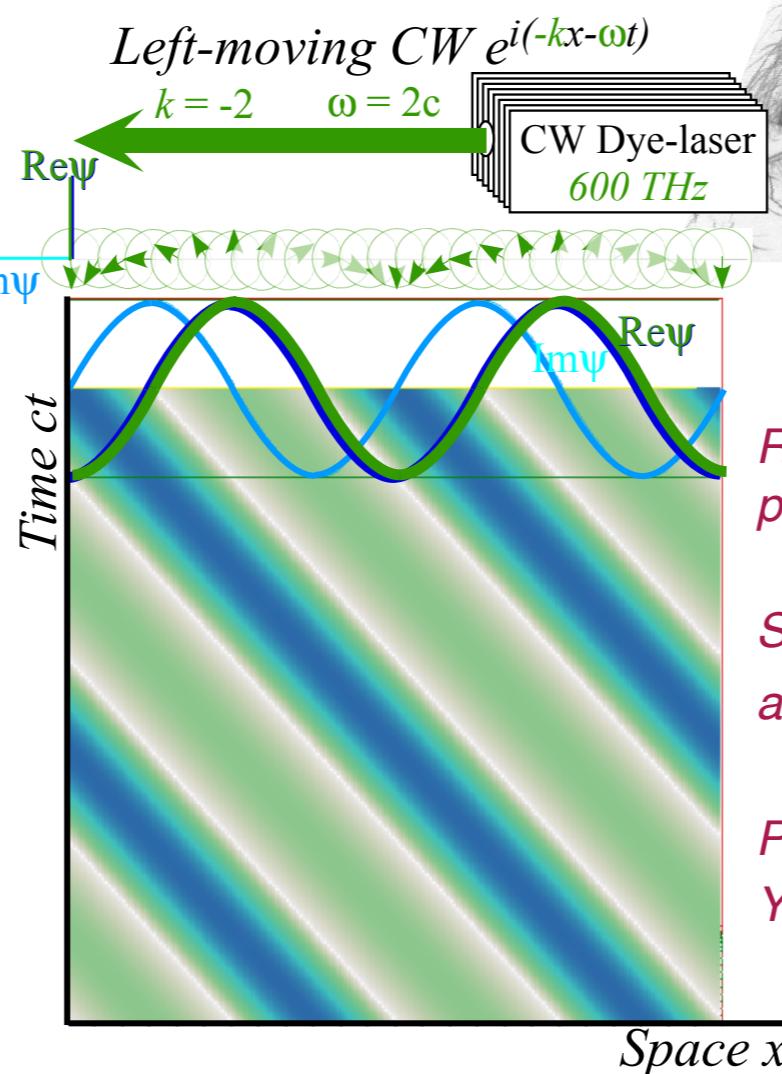
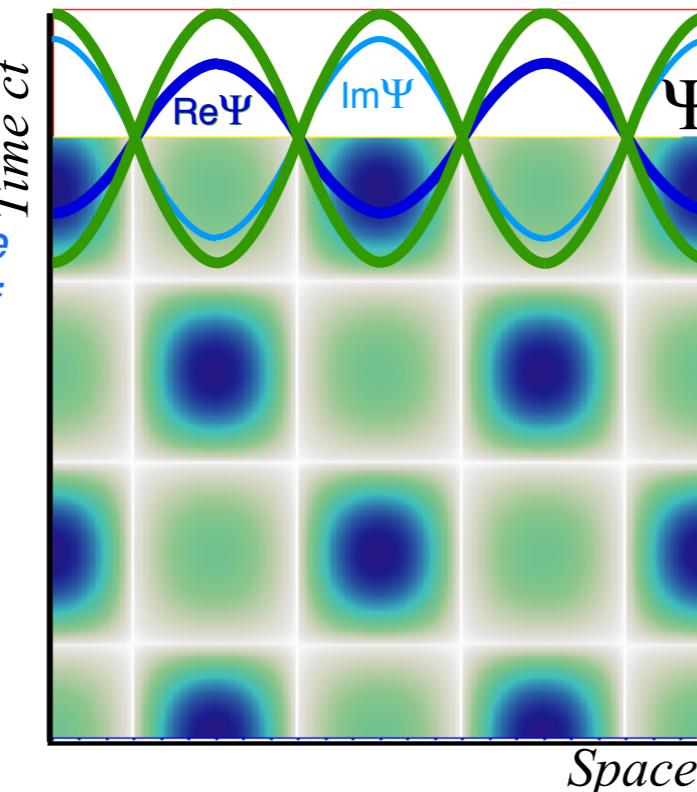
$$\begin{aligned} \Psi(x, t) &= e^{ia} + e^{ib} \\ &= e^{i\frac{a+b}{2}} \left(e^{i\frac{a-b}{2}} + e^{-i\frac{a-b}{2}} \right) \\ &= e^{-i\omega t} (e^{ikx} + e^{-ikx}) \end{aligned}$$



Bob: Let's plot this in per-spacetime?!

Cool!
You guys
made me
a space-time
graph out of
real zeros.

How'd it
do that?



$$\begin{aligned}\Psi(x,t) &= e^{ia} + e^{ib} \\ &= e^{i\frac{a+b}{2}} (e^{i\frac{a-b}{2}} + e^{-i\frac{a-b}{2}}) \\ &= e^{-i\omega t} (e^{ikx} + e^{-ikx}) \\ \Psi(x,t) &= e^{-i\omega t} 2 \cos(kx)\end{aligned}$$



Carla:

Easy!

You get zeros of any wave-sum $e^{ia} + e^{ib}$ by factoring it into phase and group parts.

Remember your algebra? Exponents of products add.

So, half-sum $\frac{a+b}{2}$ plus half-diff $\frac{a-b}{2}$ gives a , and half-sum $\frac{a+b}{2}$ minus half-diff $\frac{a-b}{2}$ gives b .

Presto!
You factor $e^{ia} + e^{ib}$ into $e^{i\frac{a+b}{2}} \left(e^{i\frac{a-b}{2}} + e^{-i\frac{a-b}{2}} \right)$

Alice 1CW phase: $a = kx - \omega t$

Carla 1CW phase: $b = -kx - \omega t$

Bob's 2CW Group-phase: $+k = \frac{a-b}{2}$
Wave

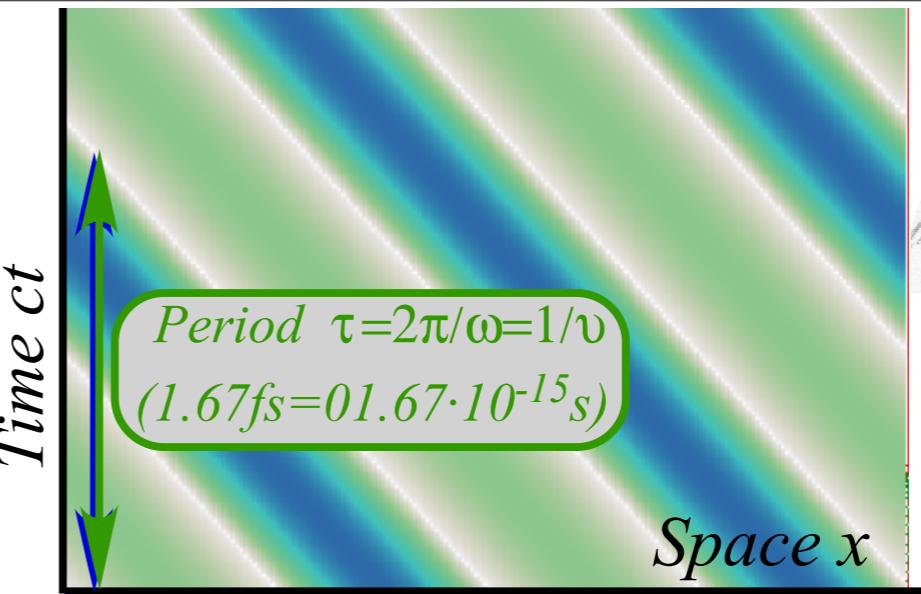
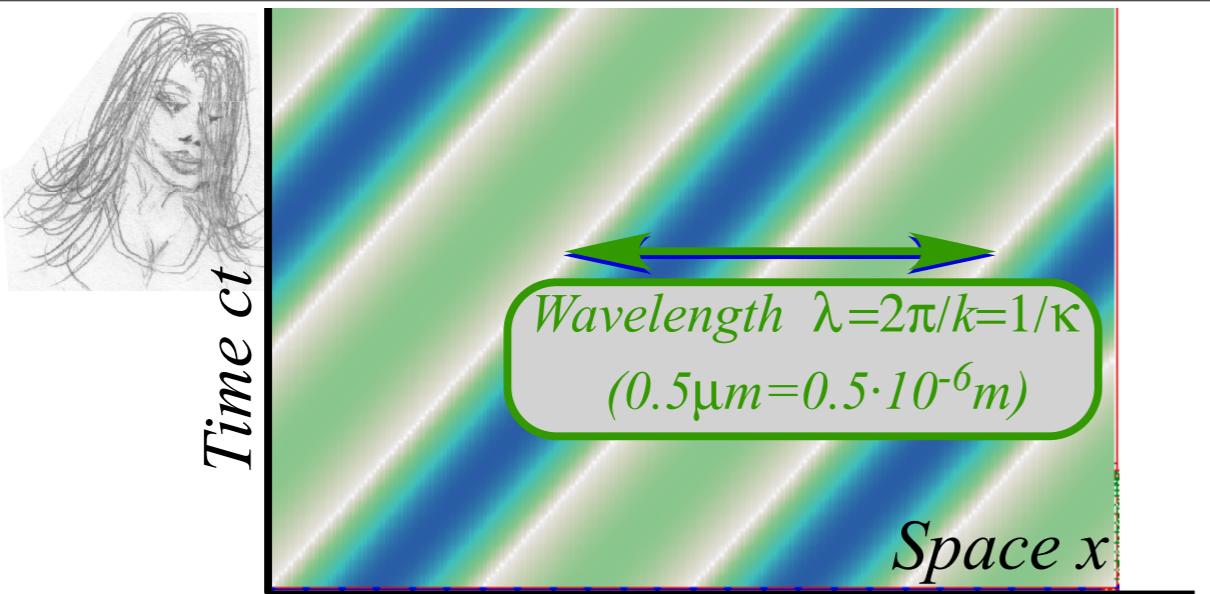
Group wave: $e^{-ikx} + e^{-ikx} = 2 \cos(kx)$

is standing wave (does not vary with time t)

Bob's 2CW Phase-phase: $-\omega = \frac{a+b}{2}$
Wave

Phase wave real part: $\text{Re}(e^{-i\omega t}) = \cos(\omega t)$

is "instanton" wave (does not vary in space x)

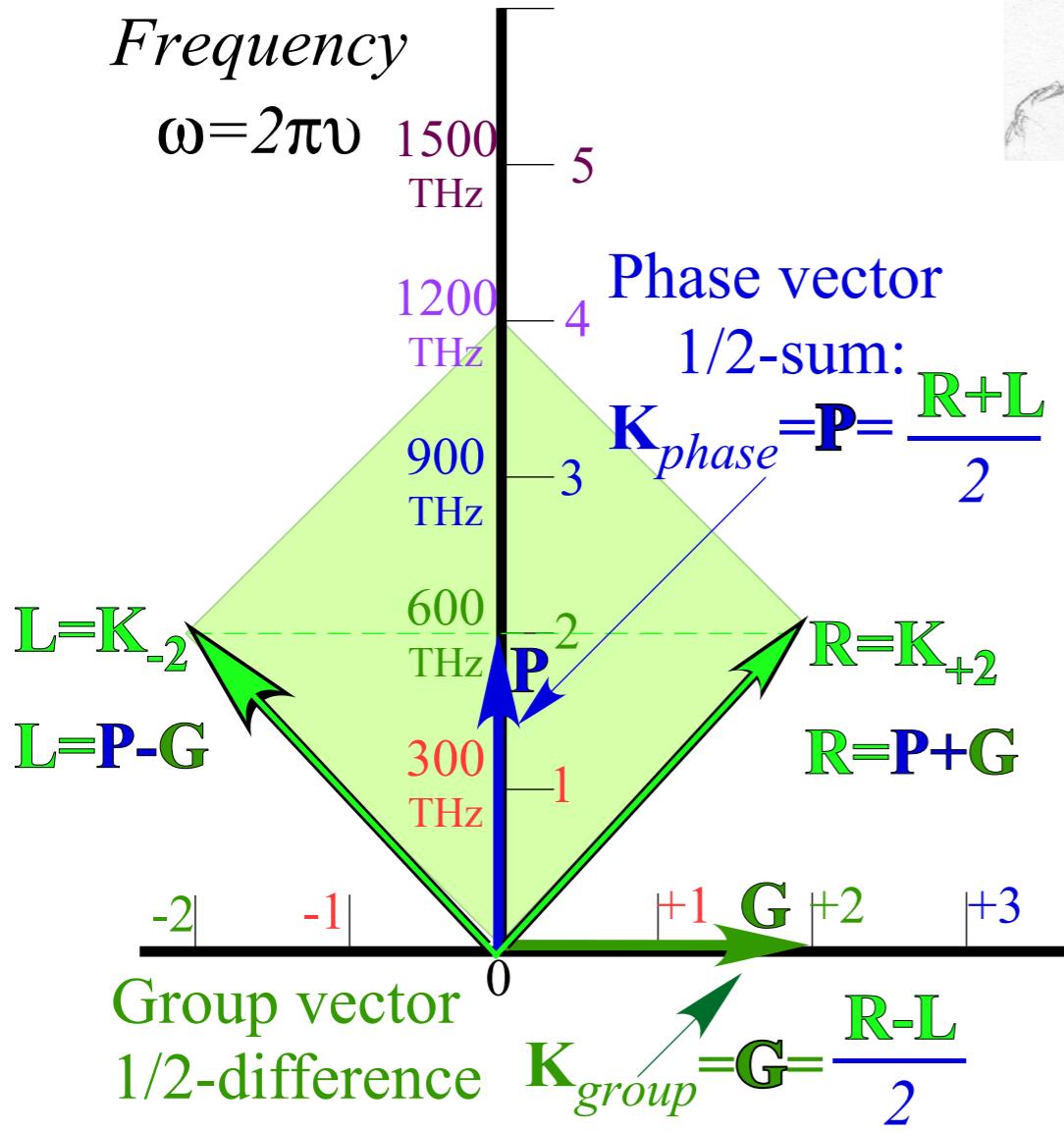


Carla:
OK, Bob!
It looks like a
baseball diamond
with
P at Pitcher's mound
and
G at the Grandstand*.
I'm on 1st base! (**R**)

*Thanks,
Woody!

$$\Psi(x,t) = (e^{-i\omega t})(2\cos kx) = e^{i(kx-\omega t)} + e^{i(-kx-\omega t)}$$

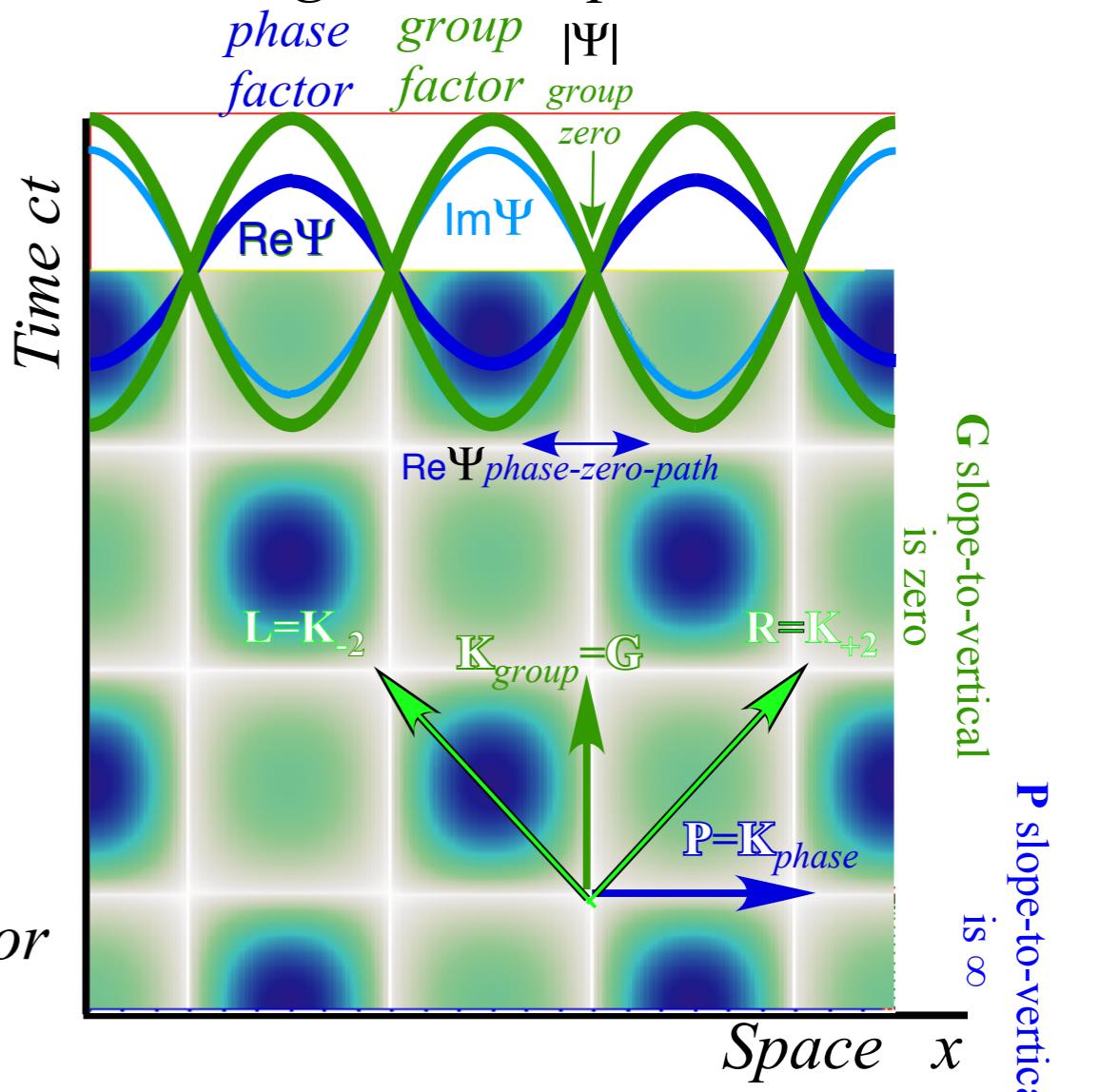
Standing 2CW in per-space-time



Bob:
The **P** and **G**
vectors are
scale models
of zero-grid
lattice vectors
(but **P** and **G**
switch places)

$$ck=2\pi\kappa c$$

Standing 2CW in space-time



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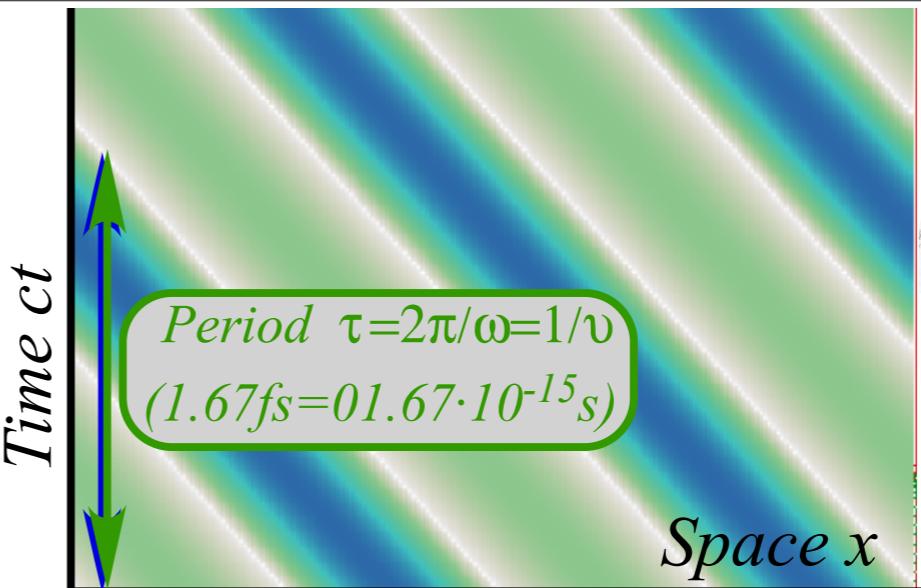
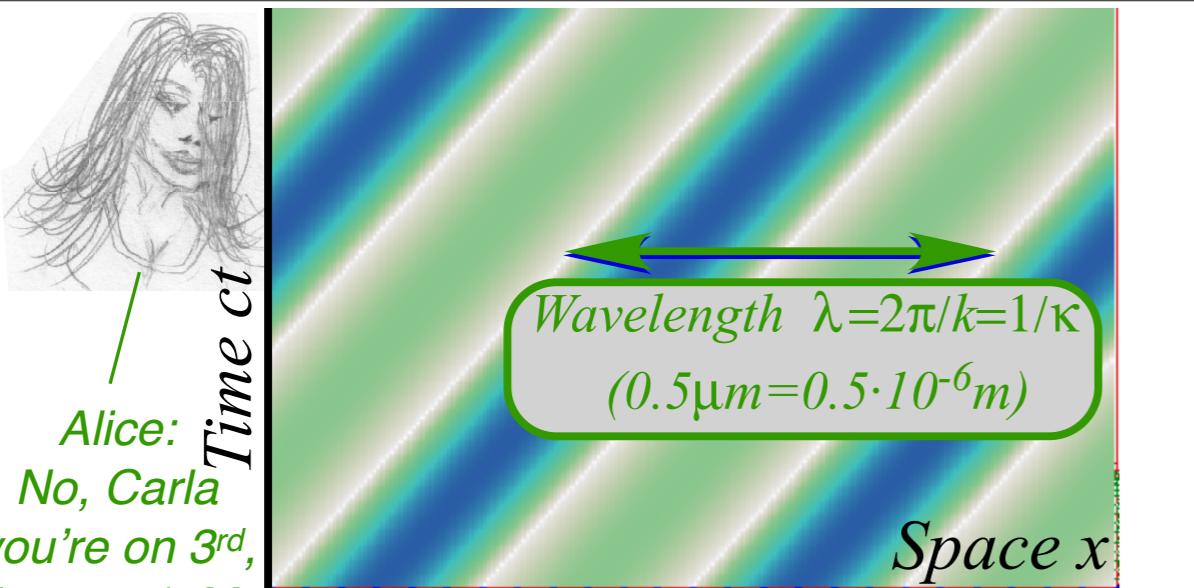
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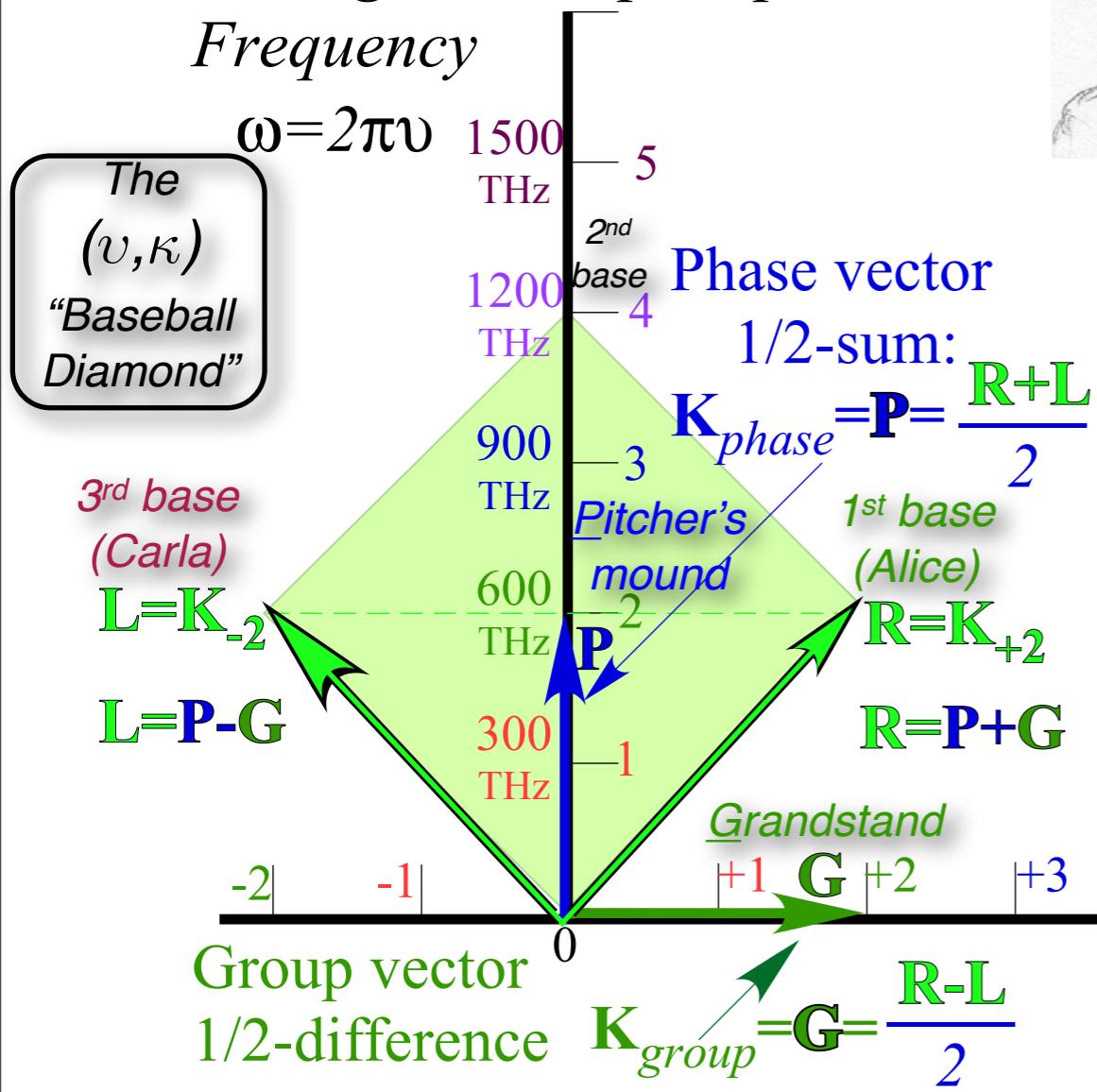
Stellar aberration and the Epstein approach to SR



*Thanks, Woody!

$$\Psi(x,t) = (e^{-i\omega t})(2\cos kx) = e^{i(kx-\omega t)} + e^{i(-kx-\omega t)}$$

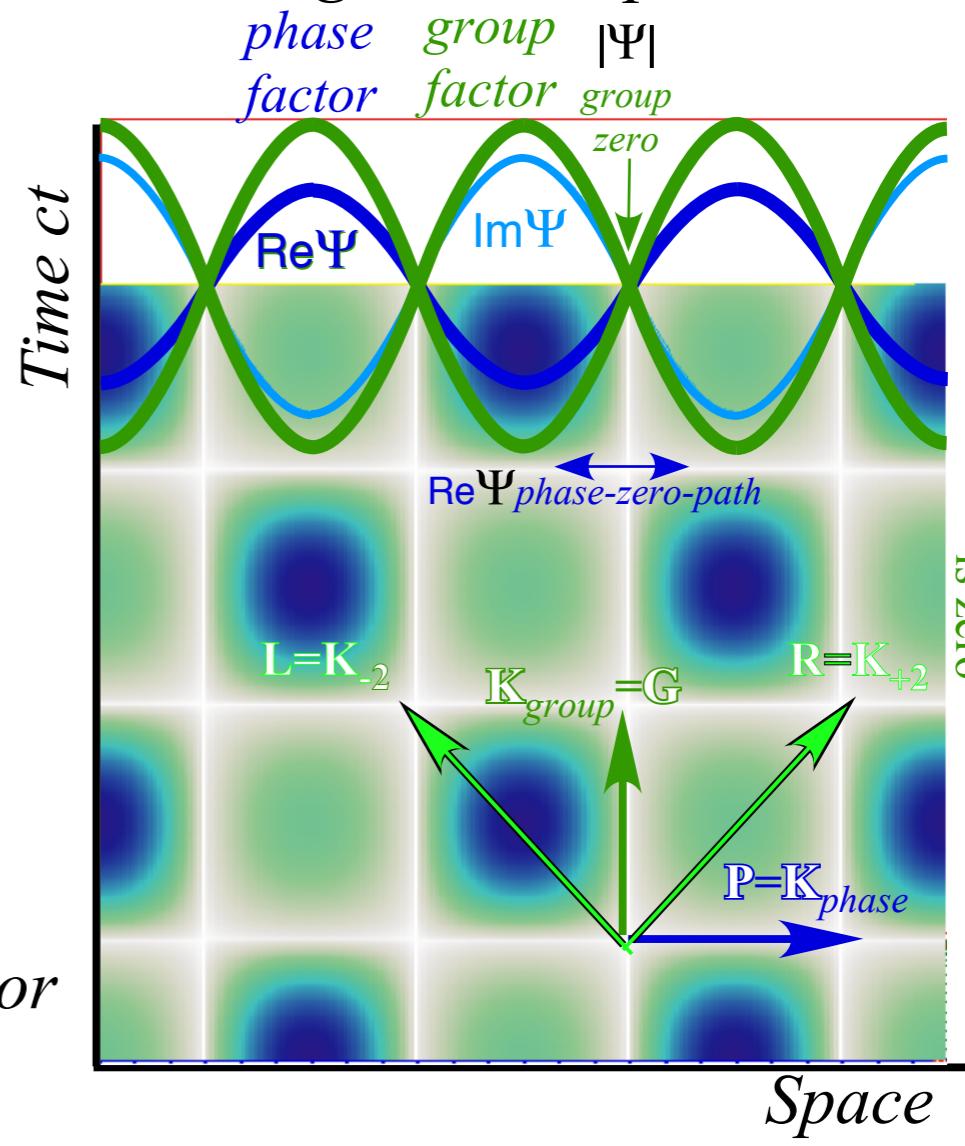
Standing 2CW in per-space-time



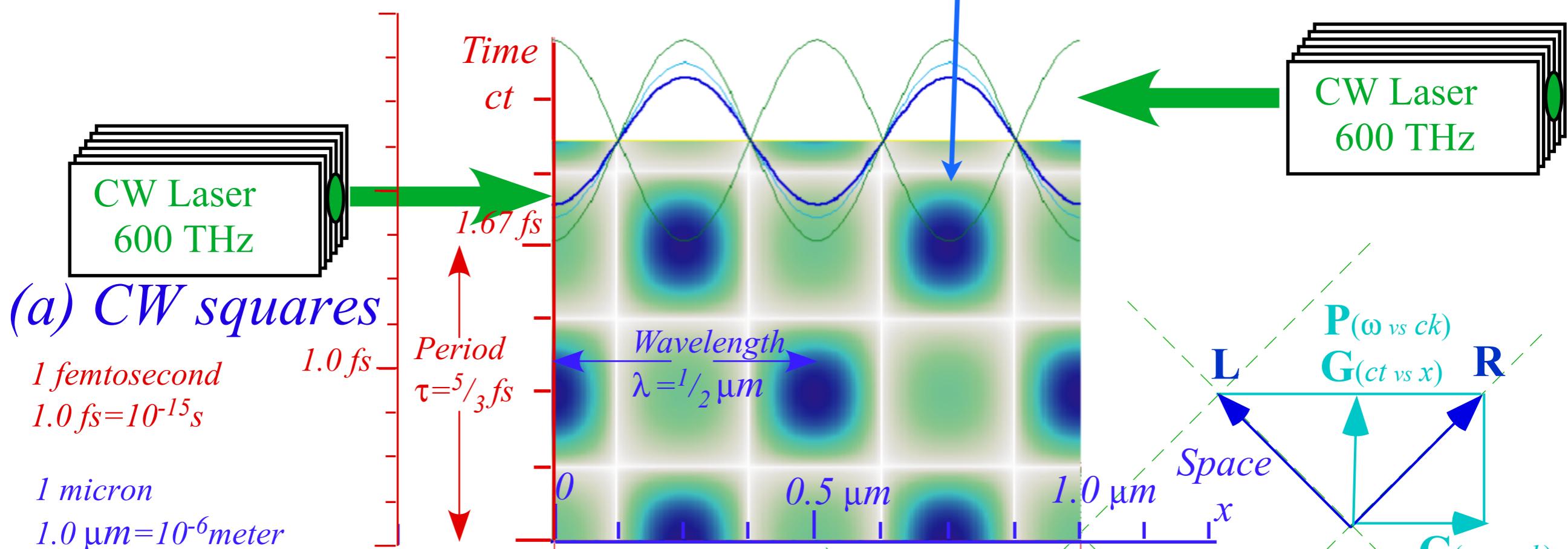
Bob: The P and G vectors are scale models of zero-grid lattice vectors (but P and G switch places)

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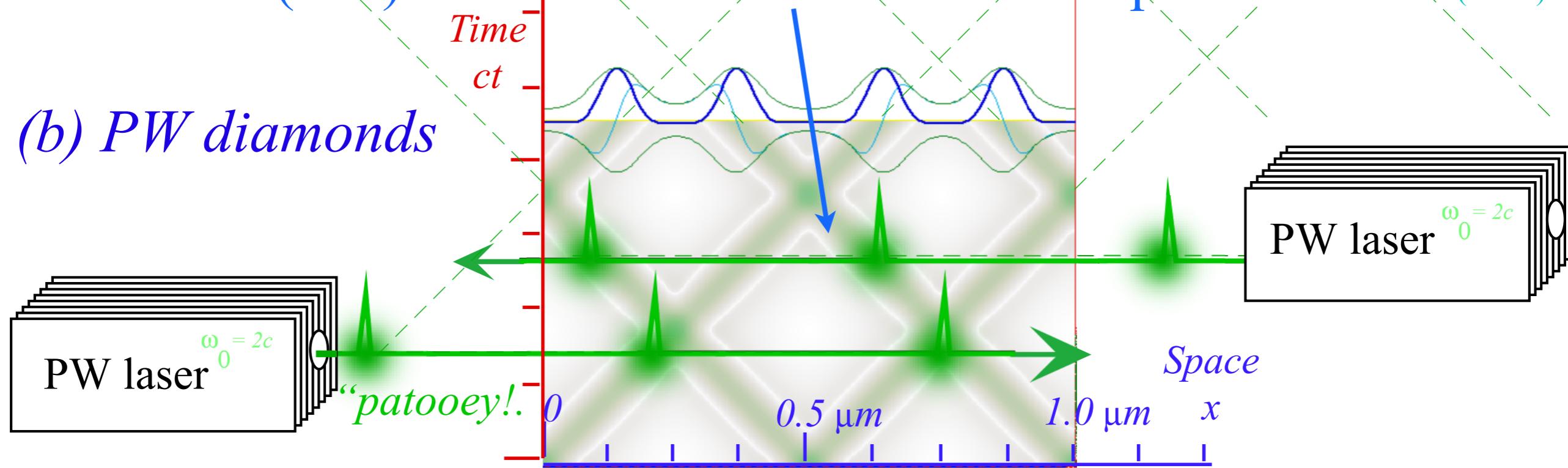
Standing 2CW in space-time

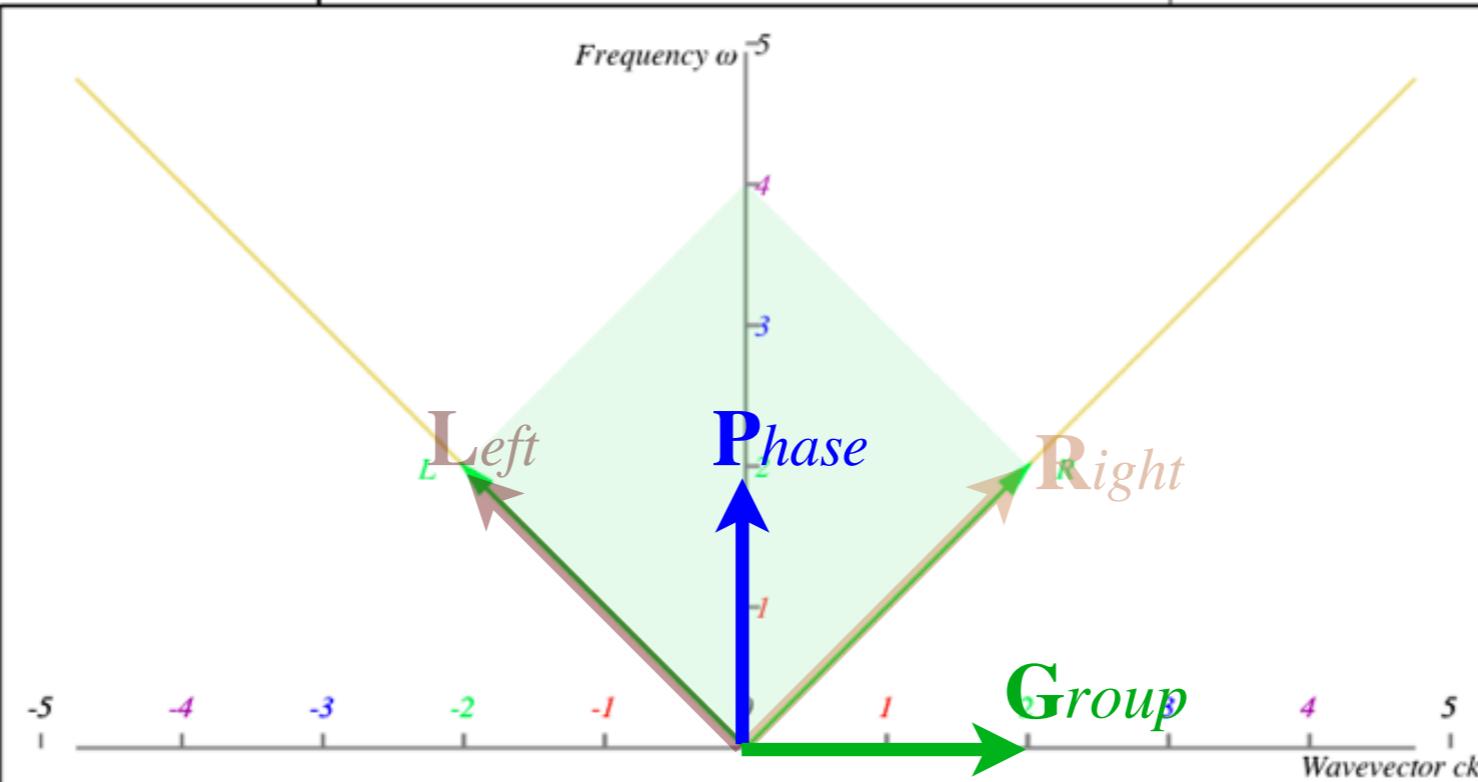
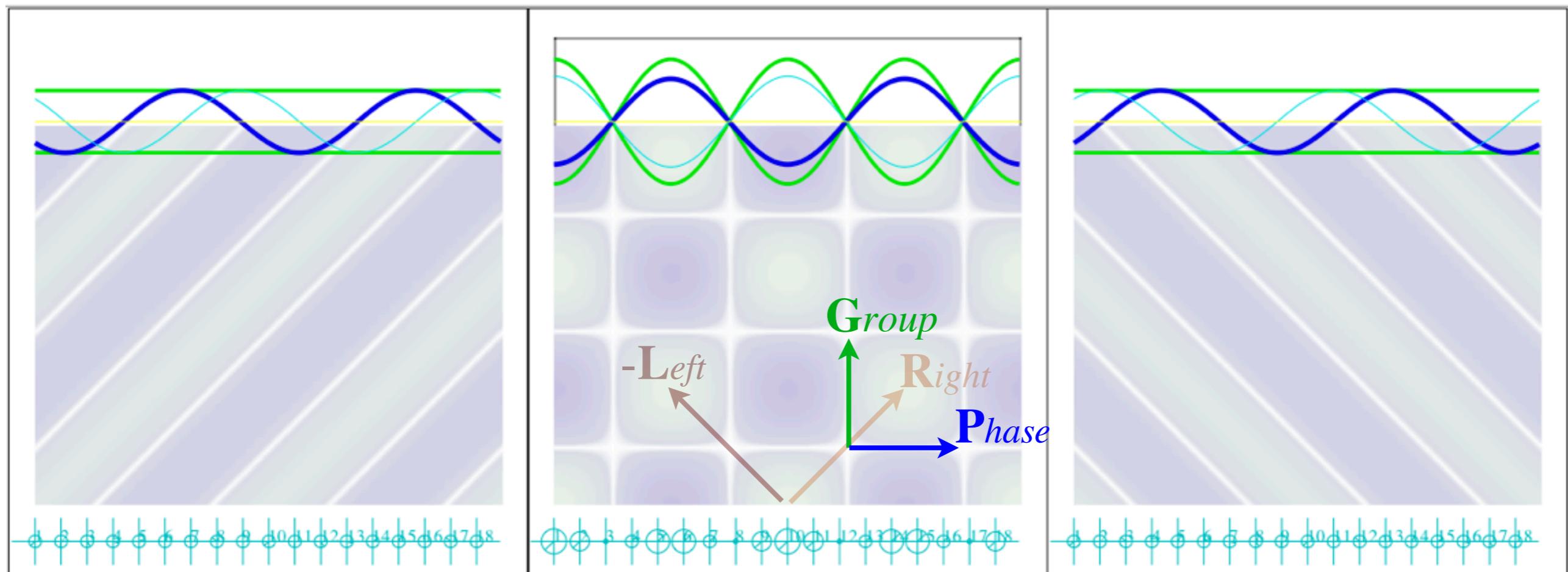


Continuous Waves (CW) trace “Cartesian squares” in space-time



Pulse Waves (PW) trace “baseball diamonds” in space-time

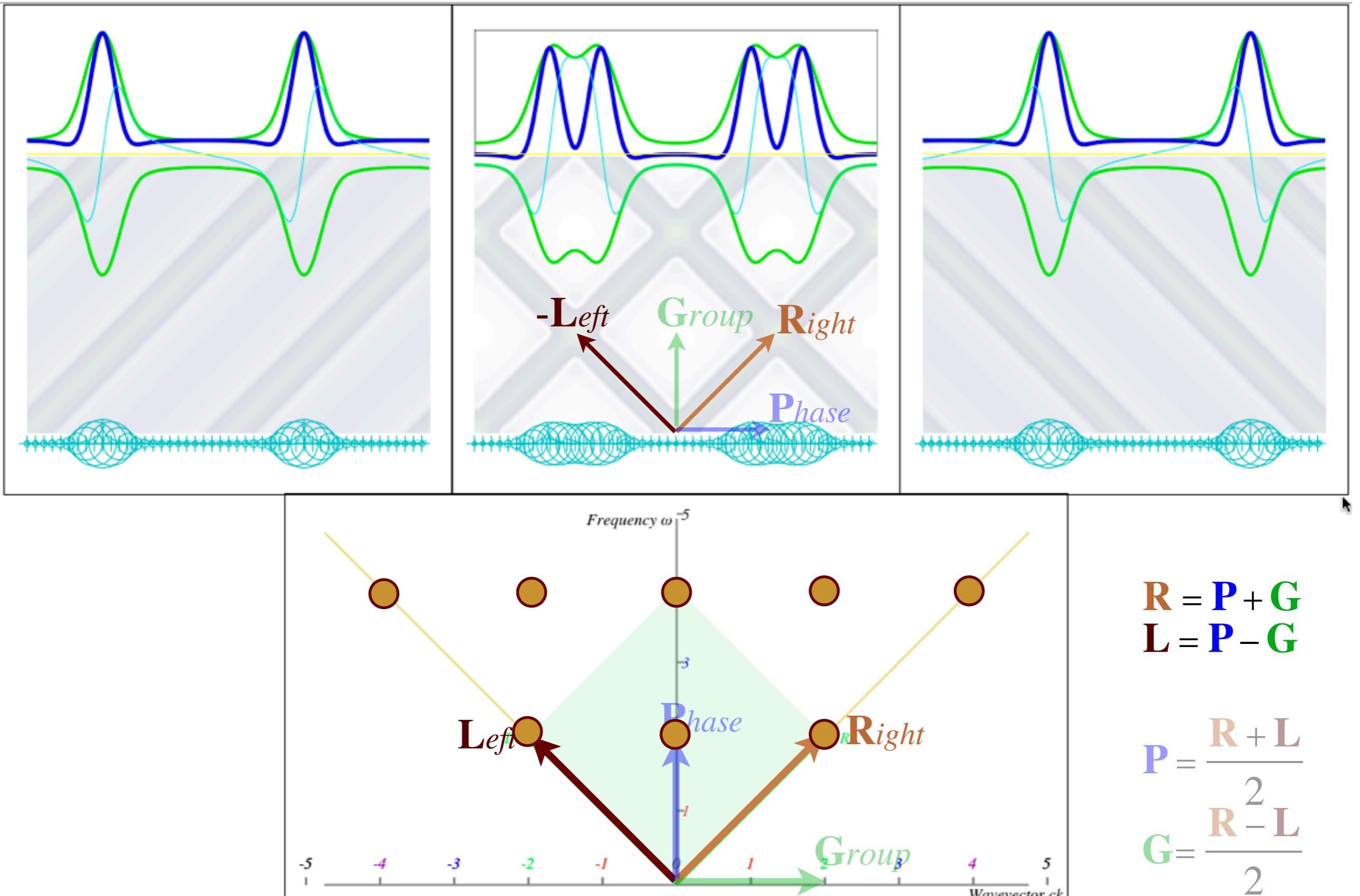




$$\begin{aligned} \mathbf{R} &= \mathbf{P} + \mathbf{G} \\ \mathbf{L} &= \mathbf{P} - \mathbf{G} \\ \mathbf{P} &= \frac{\mathbf{R} + \mathbf{L}}{2} \\ \mathbf{G} &= \frac{\mathbf{R} - \mathbf{L}}{2} \end{aligned}$$

[BohrIt Web Simulation](#)
[2 CW ct vs x Plot](#)
 $(ck = \pm 2)$

[RelaWavity Site](#)
[Phase and Group Vectors in per-Time vs per-Space](#)



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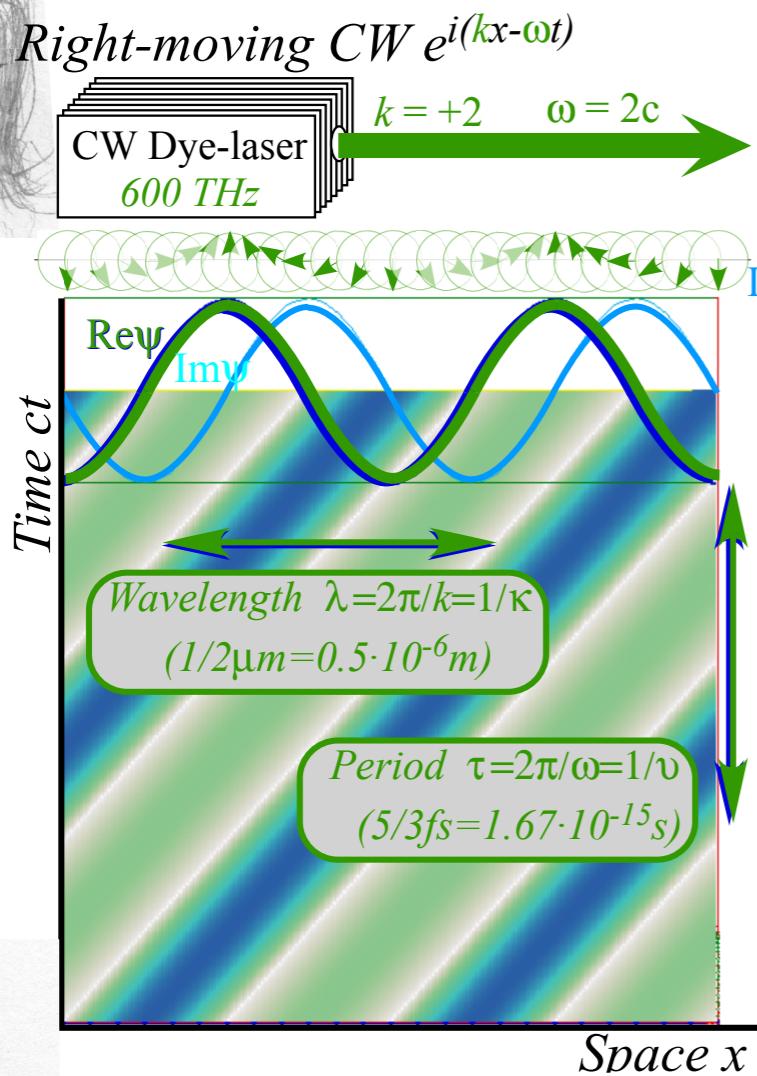
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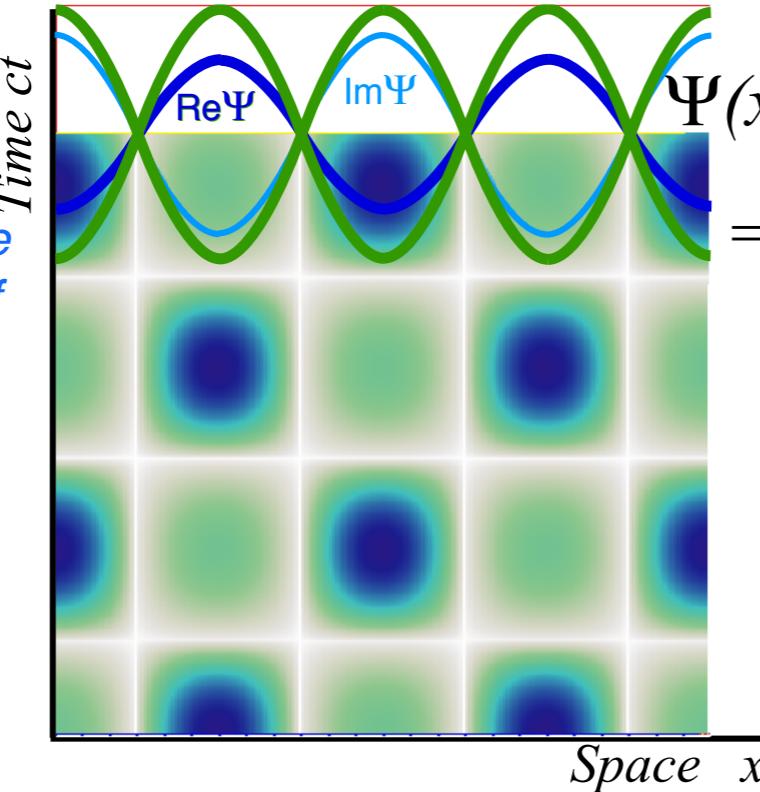
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Bob:
 Cool!
 You guys
 made me
 a space-time
 graph out of
 real zeros.

How'd it
do that?



$$\Psi(x,t) = e^{ia} + e^{ib}$$

$$= e^{i\frac{a+b}{2}} (e^{i\frac{a-b}{2}} + e^{-i\frac{a-b}{2}})$$



Carla:

Easy!

You get zeros of any wave-sum $e^{ia} + e^{ib}$ by factoring it into phase and group parts.

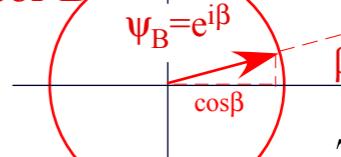
Remember your algebra? Exponents of products add.

So, half-sum $\frac{a+b}{2}$ plus half-diff $\frac{a-b}{2}$ gives a , and half-sum $\frac{a+b}{2}$ minus half-diff $\frac{a-b}{2}$ gives b .

Presto!

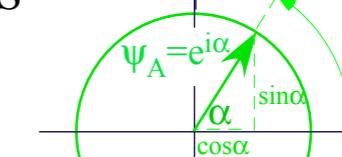
You factor $e^{ia} + e^{ib}$ into $e^{i\frac{a+b}{2}} \left(e^{i\frac{a-b}{2}} + e^{-i\frac{a-b}{2}} \right)$

Red phasor B



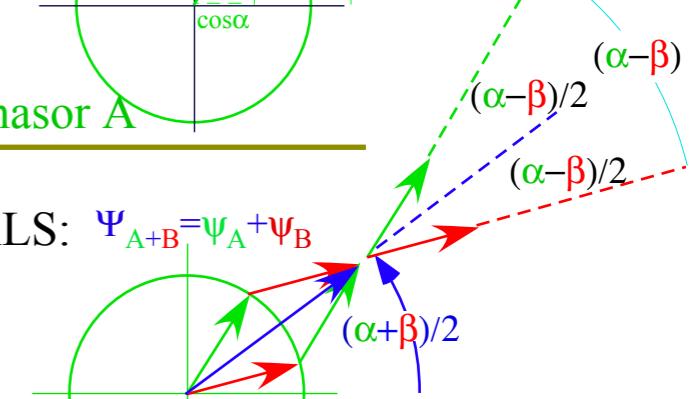
Typical
Phasor Sum:

PLUS



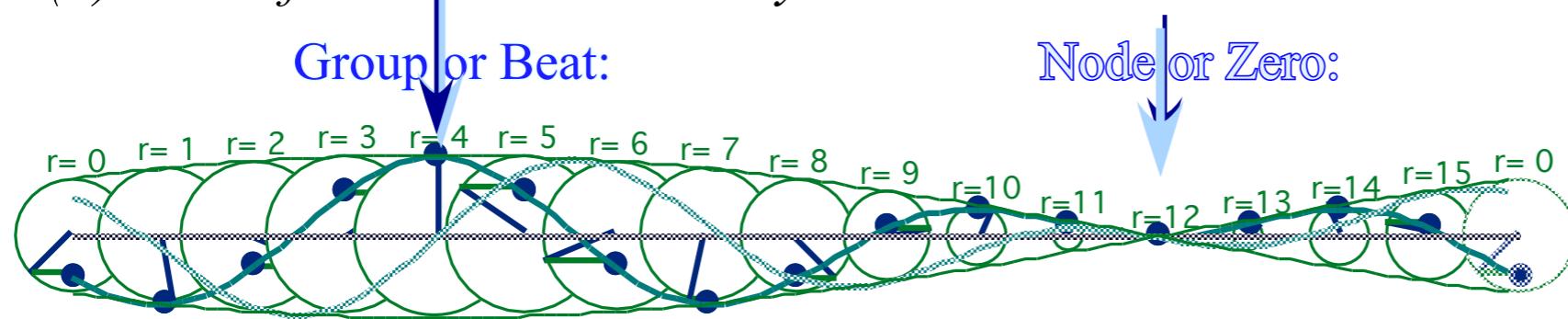
Green phasor A

EQUALS: $\Psi_{A+B} = \Psi_A + \Psi_B$

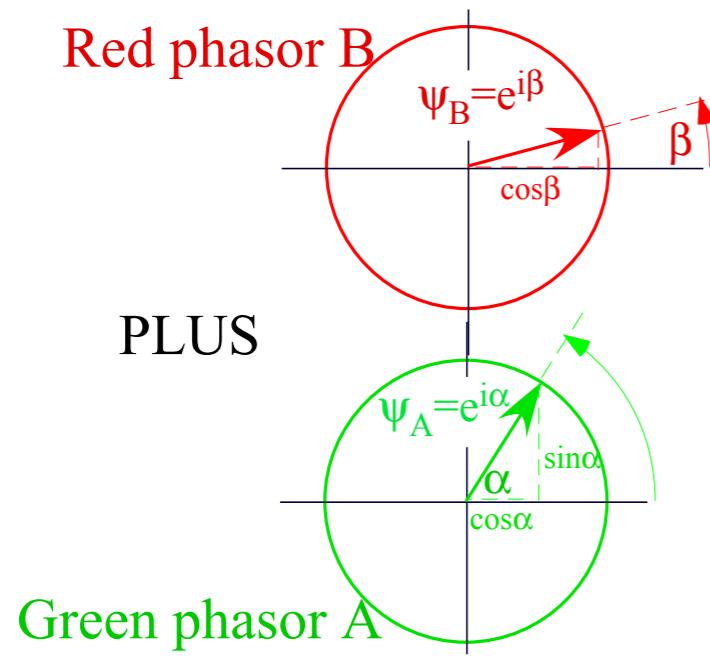


More at Pirelli Challenge page: 'Un Grande Affare' - Light Meets Light

(a) Sum of Wave Phasor Array



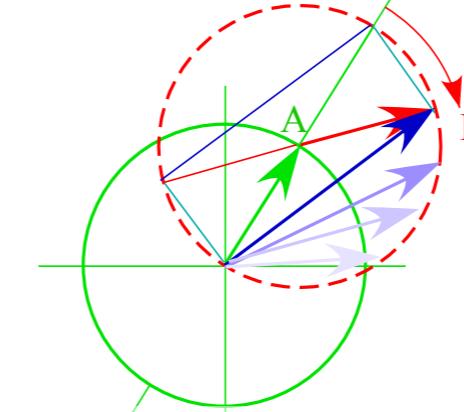
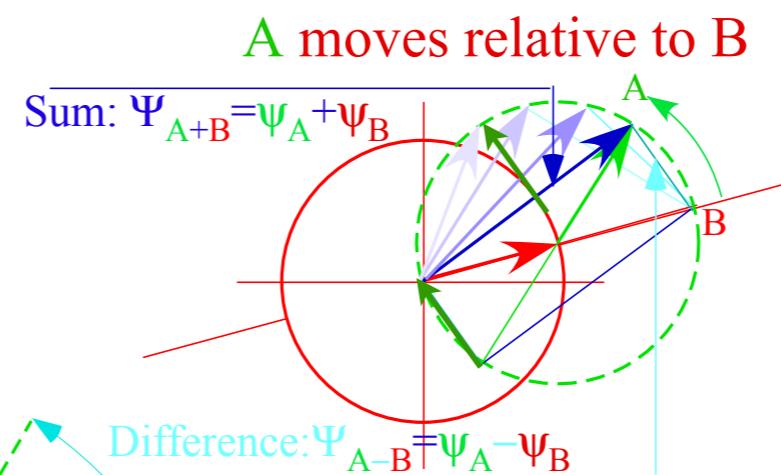
(b) Typical Phasor Sum:



$$\text{EQUALS: } \Psi_{A+B} = \Psi_A + \Psi_B$$

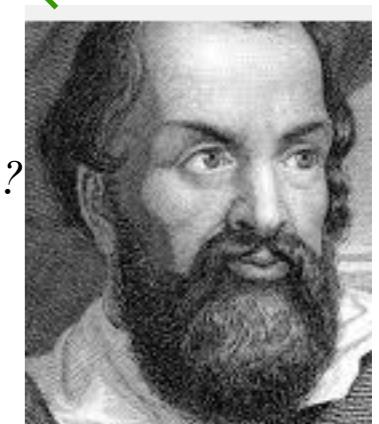
More at Pirelli Challenge page: '[Un Grande Affare](#)' - Light Meets Light

(c) Phasor-relative views



Geometry of the
Half-sum
Phase
and
Half-difference
Group

Happy now?



Galileo's Revenge (part 2)
Phasor angular velocity
adds just like
Galilean velocity

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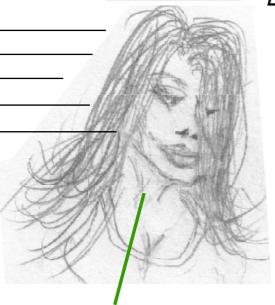
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More at Pirelli Challenge page: [*'Un Grande Affare' - Light Meets Light*](#)

Right-directed 1CW $e^{i(k_4 x - \omega_4 t)}$

$$k_4 = +4 \quad \omega_4 = 4c$$

CW green-laser
600 THz Doppler blue shifted
to 1200THz



Left-directed 1CW $e^{i(k_{-1} x - \omega_{-1} t)}$

$$k_{-1} = -1$$

$$\omega_{-1} = 1c$$

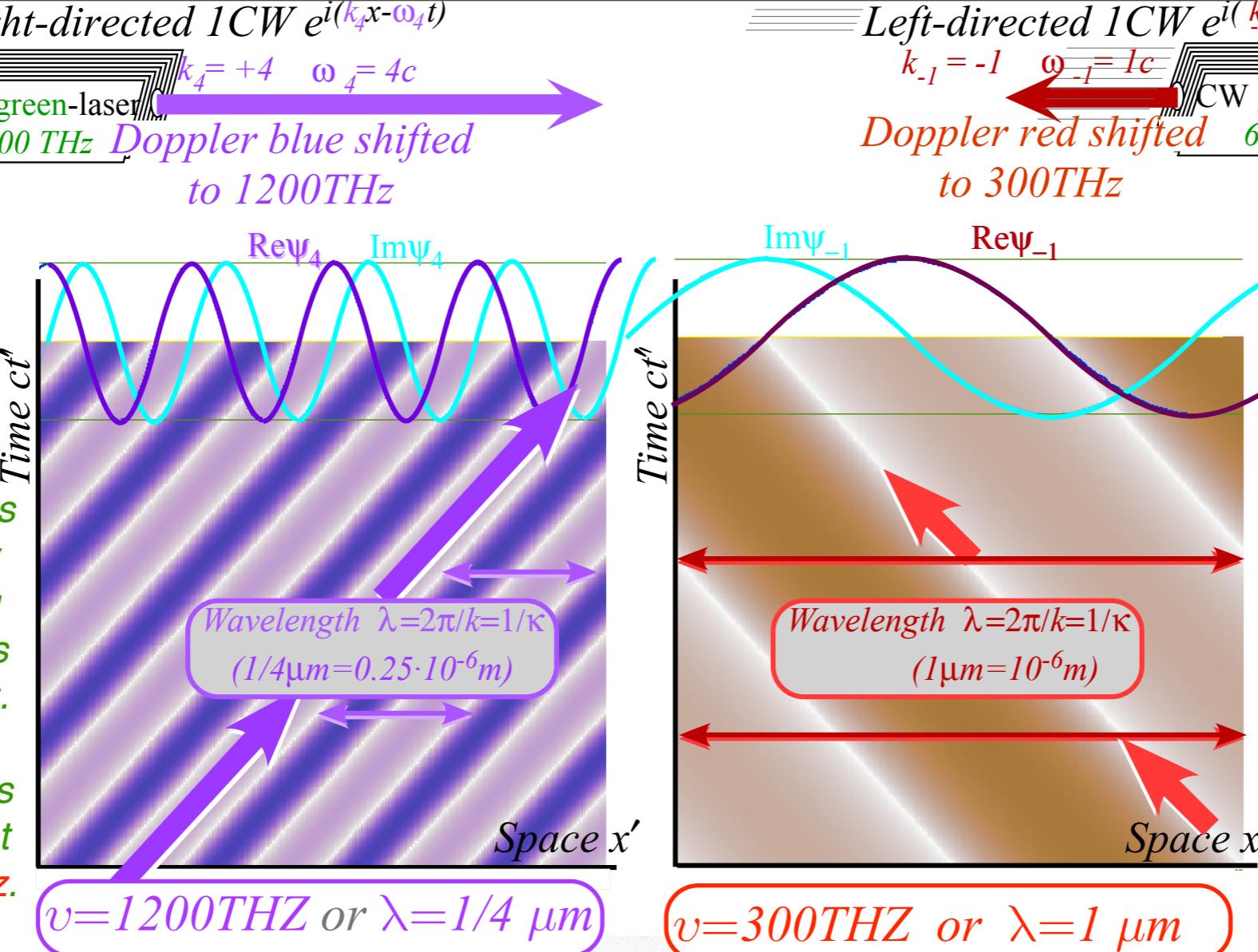
CW green-laser
600 THz Doppler red shifted
to 300THz



Alice:

Now our 600THz lasers move left-to-right. My 600THz laser is going so fast its beam blasts you with UV 1200THz.

Carla's 600THz laser is going away so you get a nice infrared 300THz.



[BohrIt Web Simulation](#)

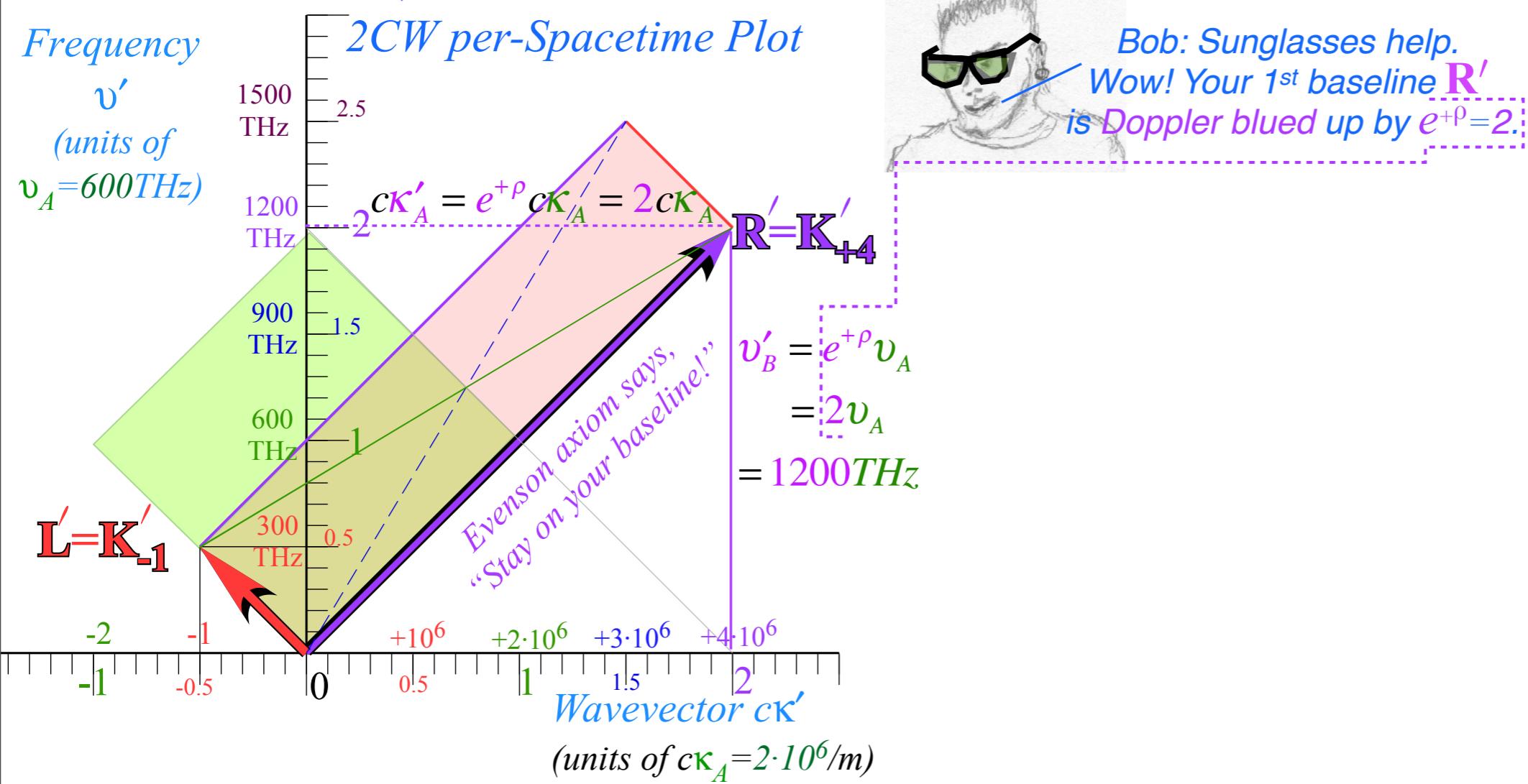
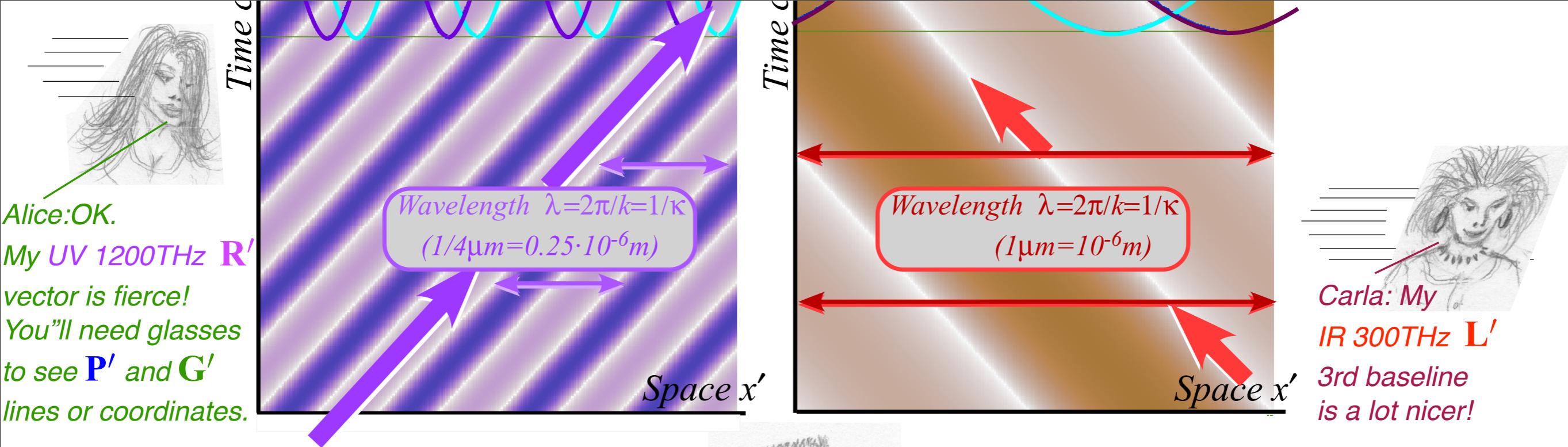
[1 CW ct vs x Plot](#)
($ck = +4$)

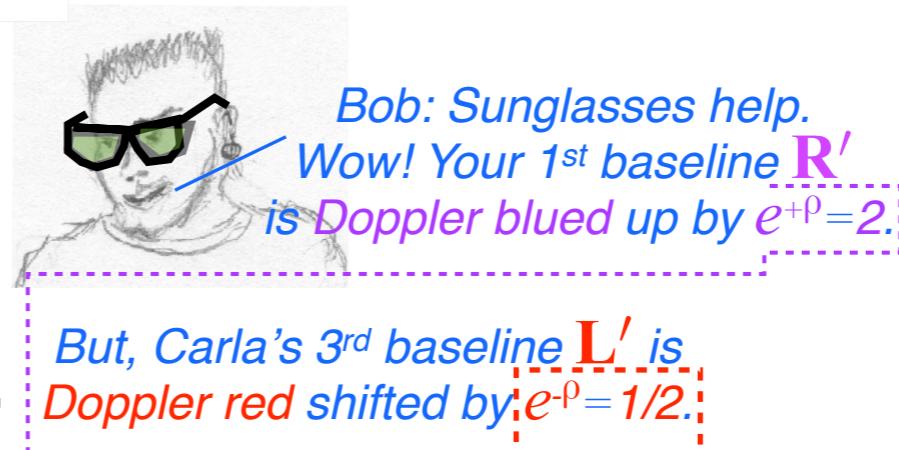
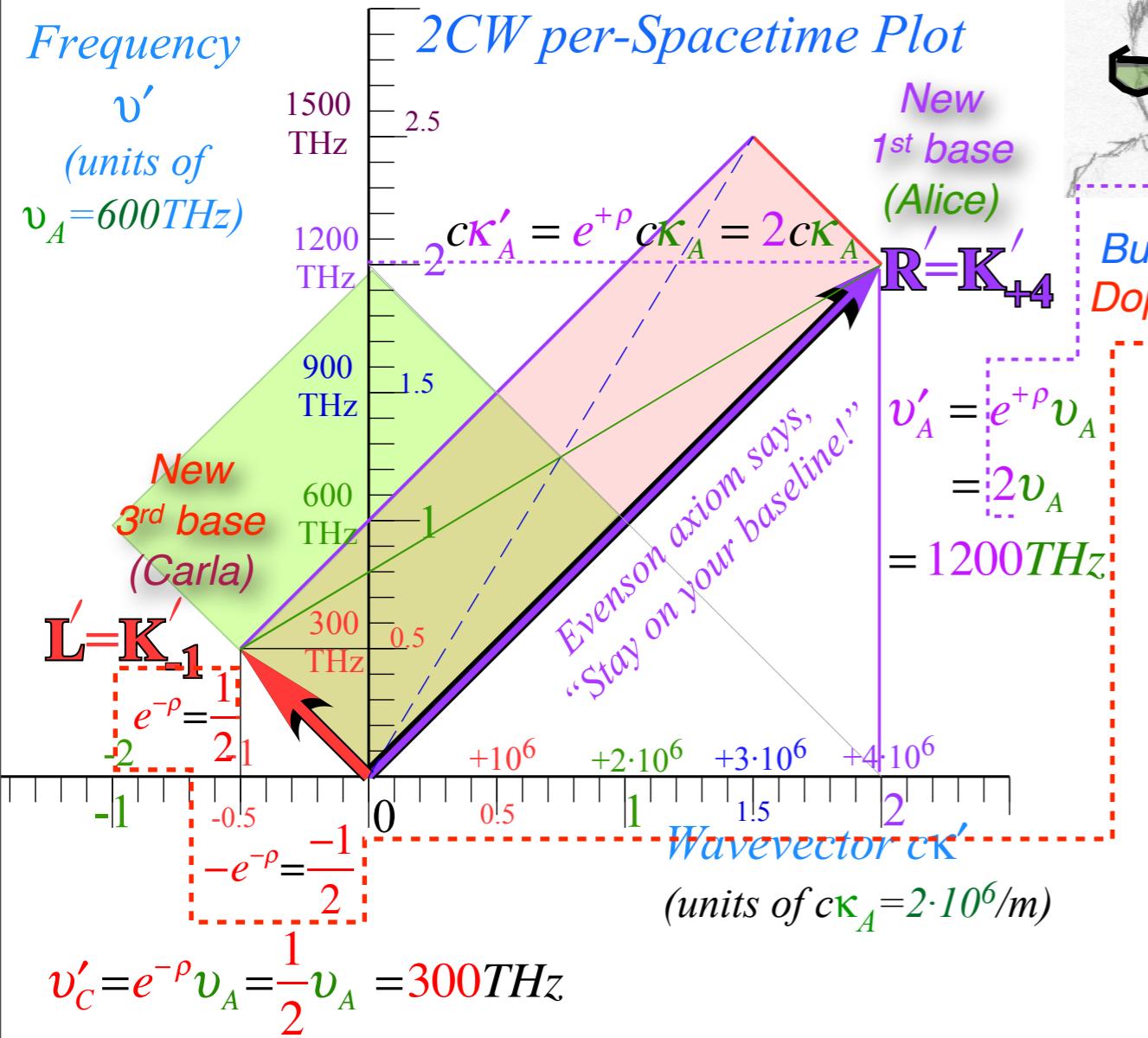
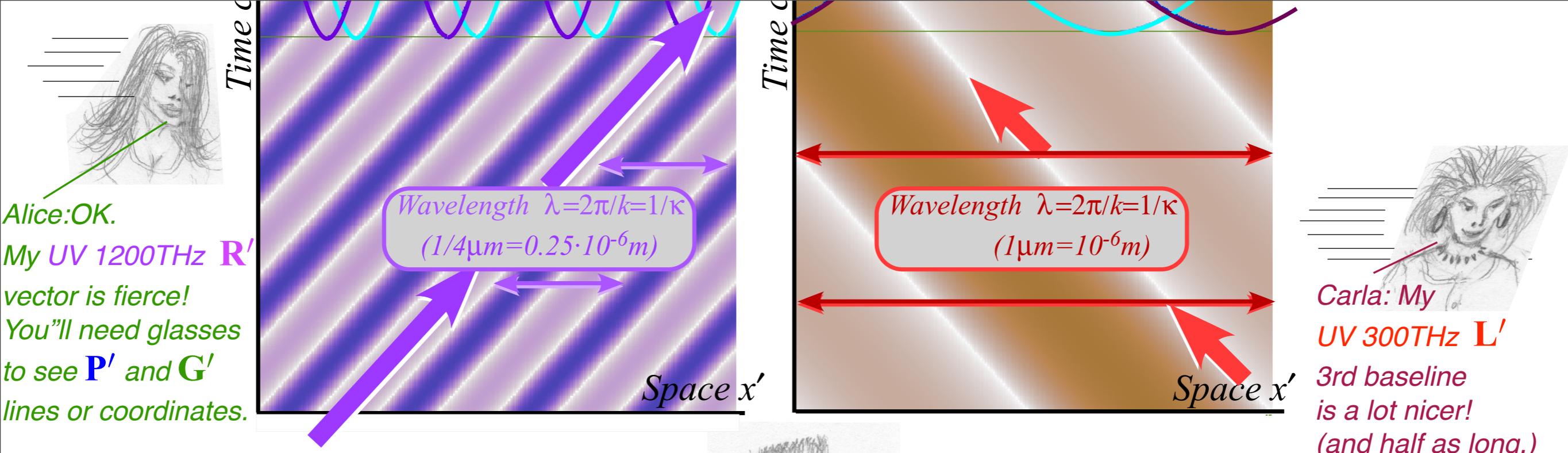


Bob: That UV burns!
I need to put on my sunglasses.

[BohrIt Web Simulation](#)

[1 CW ct vs x Plot](#)
($ck = -1$)





*But, Carla's 3rd baseline \mathbf{L}' is
Doppler red shifted by $e^{-\rho} = 1/2$.*

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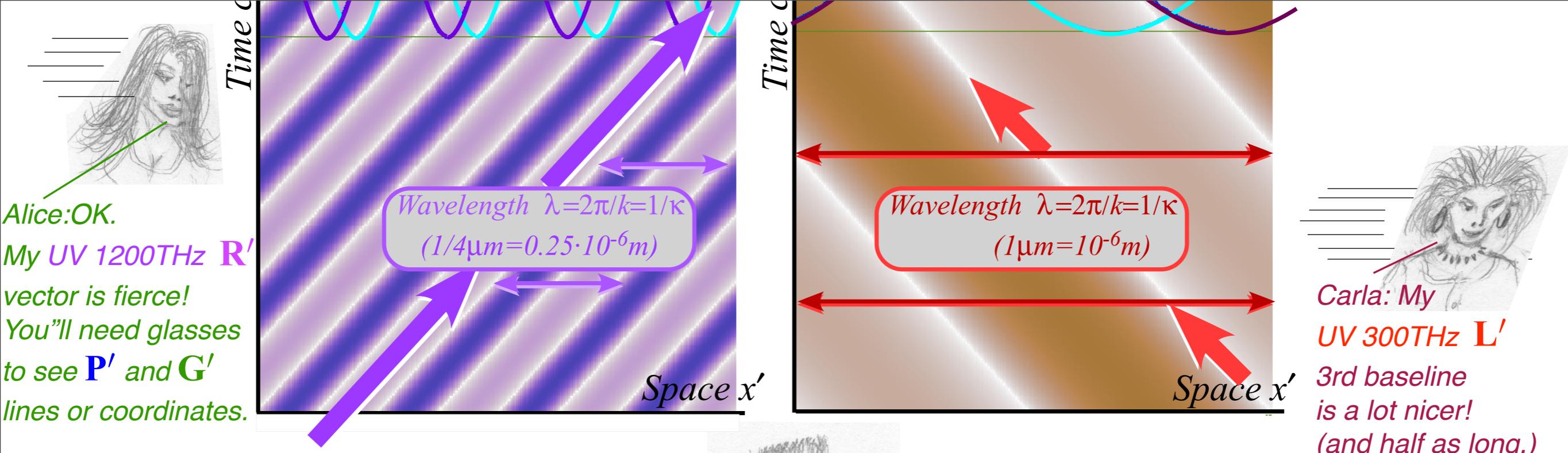
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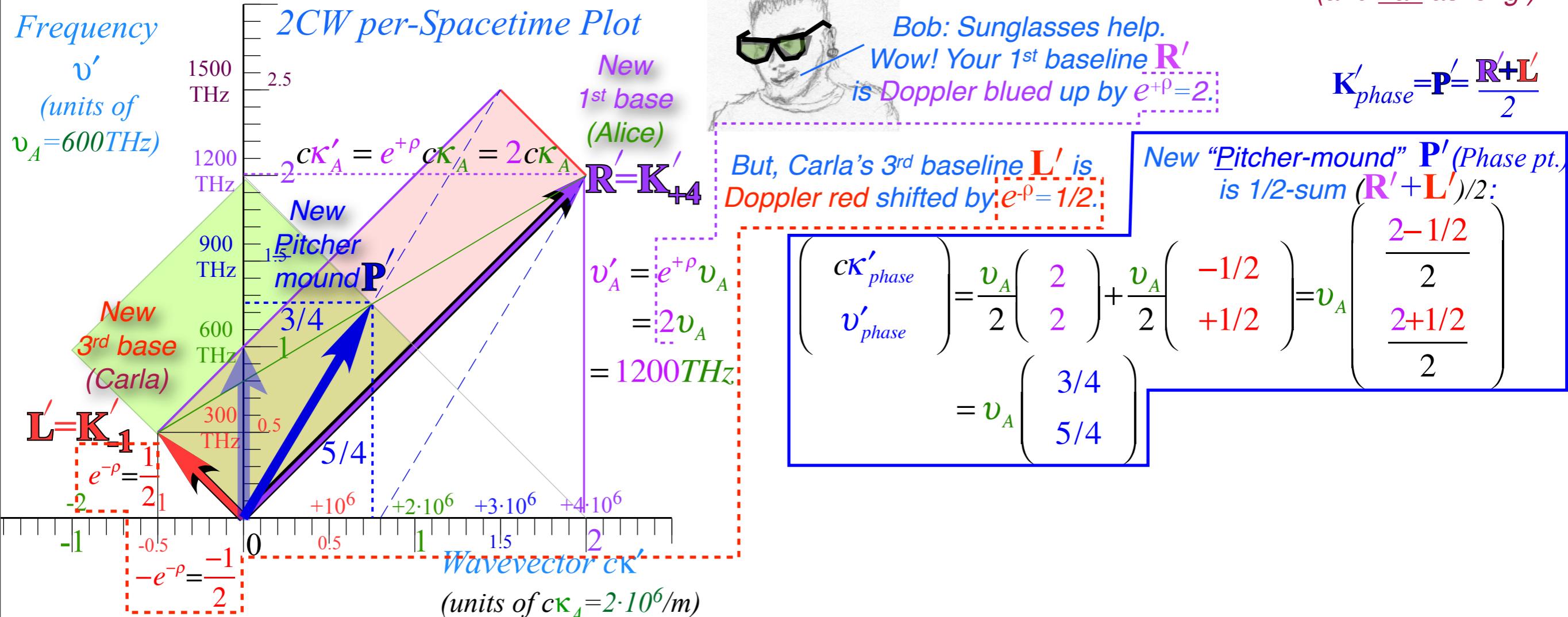
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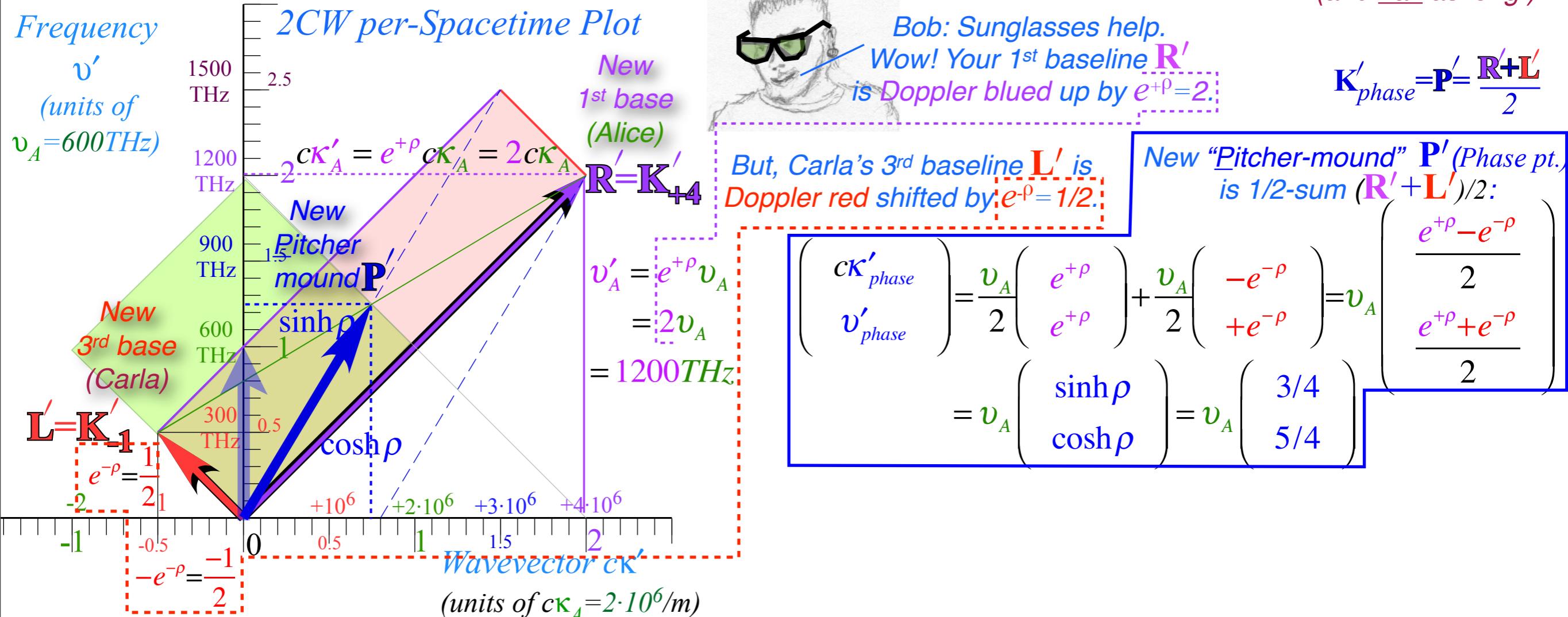
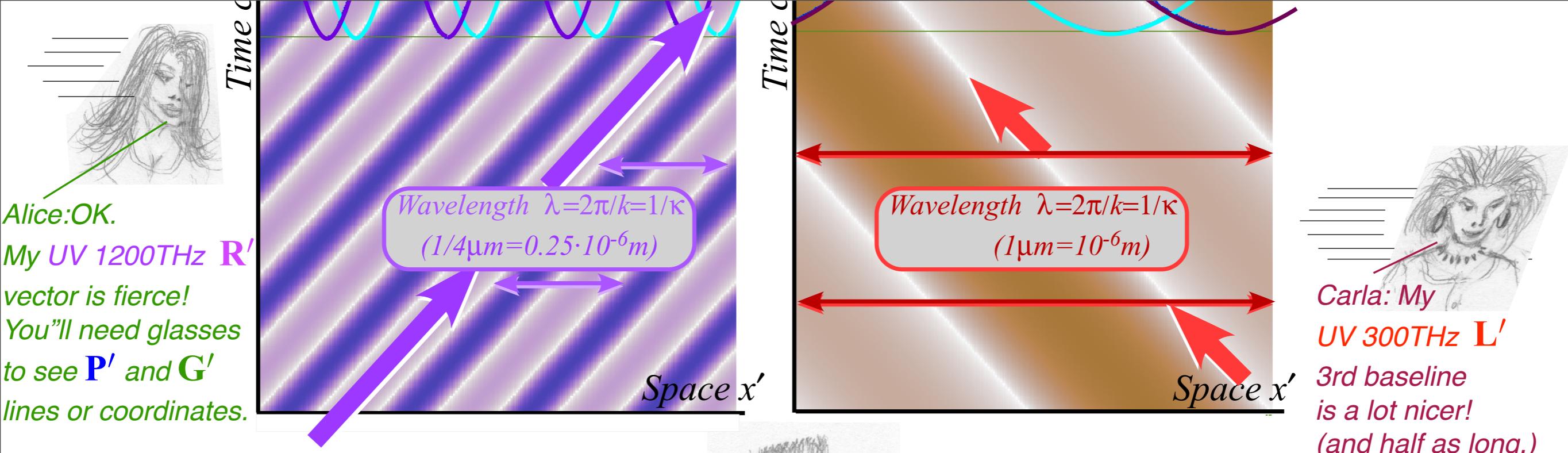
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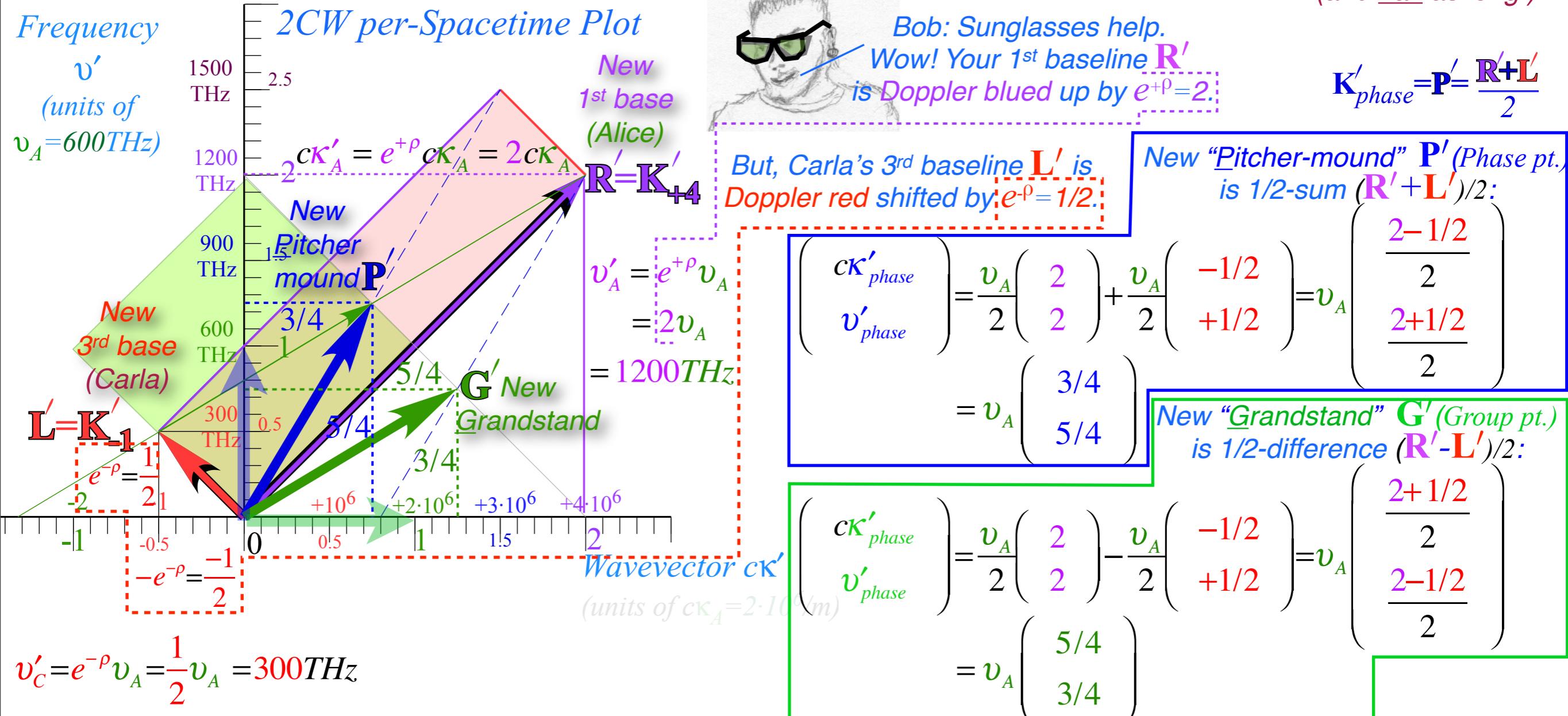
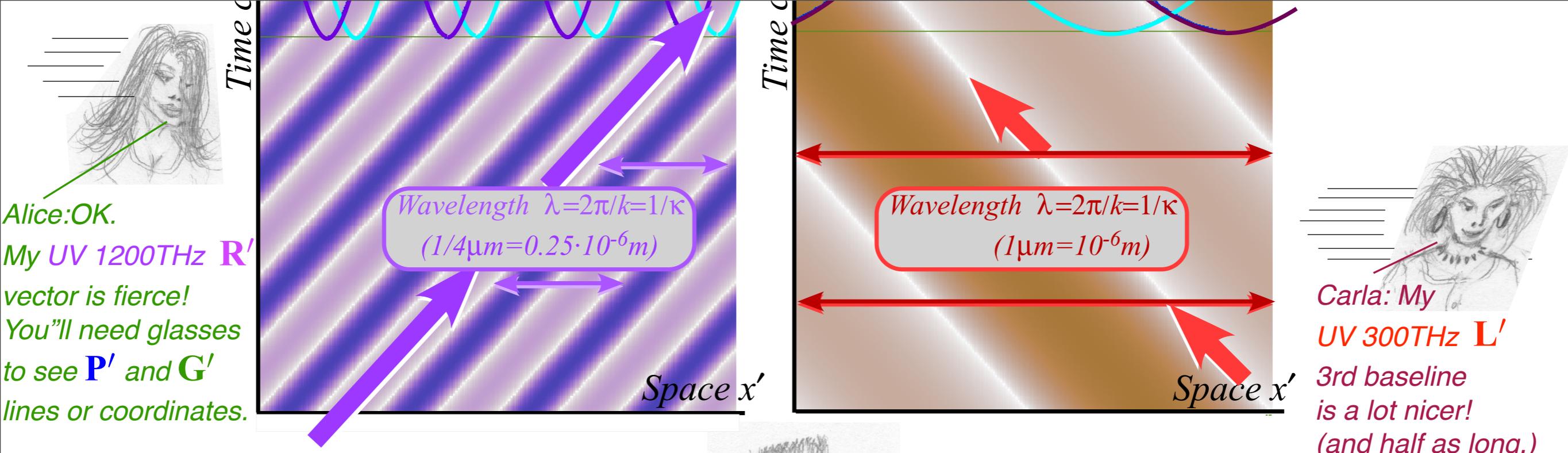
Carla: My
UV 300THz \mathbf{L}'
3rd baseline
is a lot nicer!
(and half as long.)

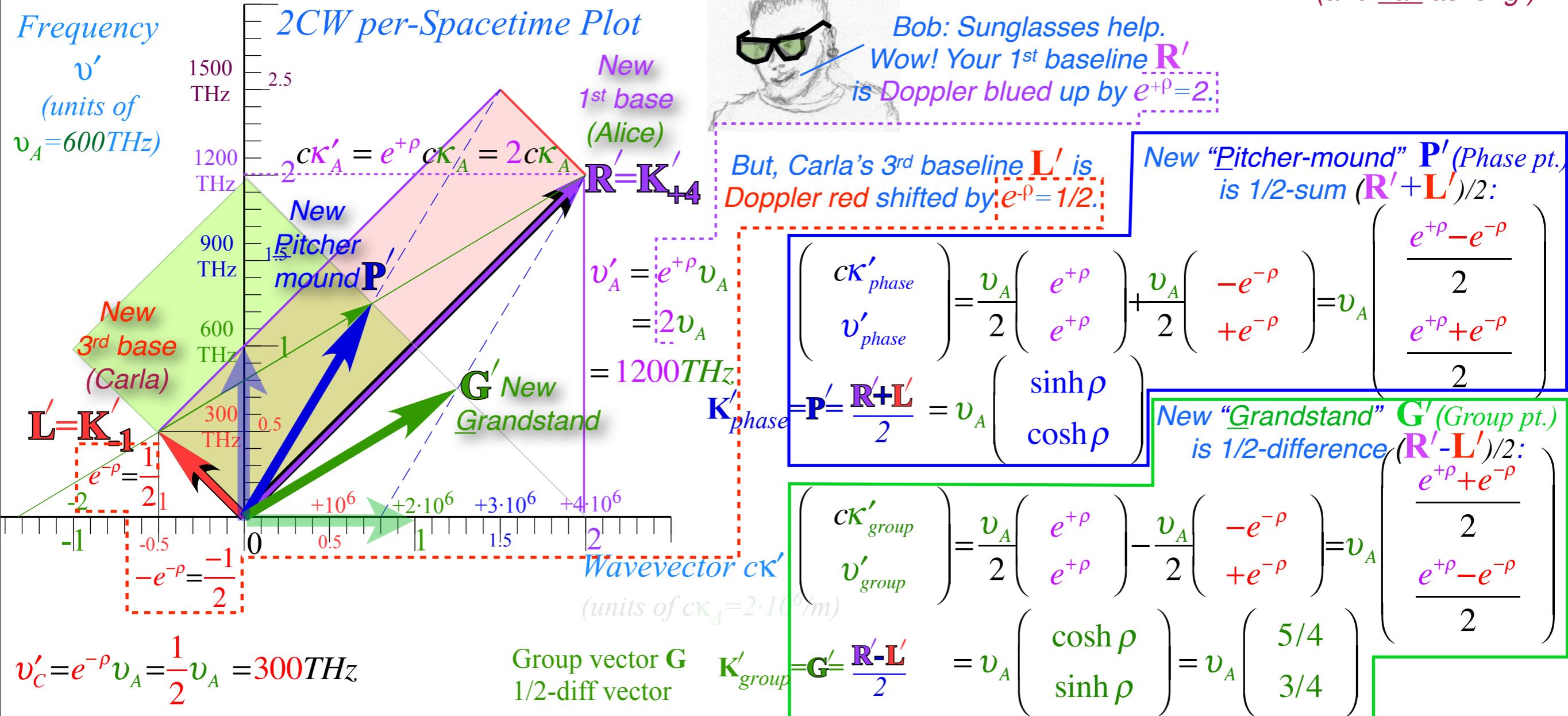
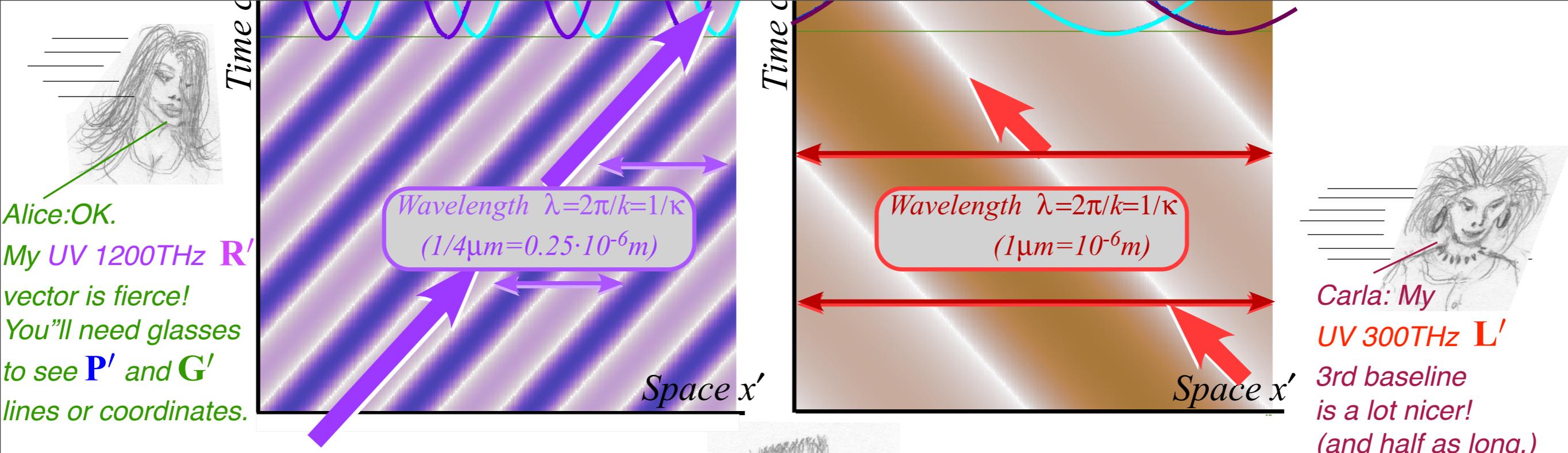




$$v'_C = e^{-\rho} v_A = \frac{1}{2} v_A = 300 \text{ THz}$$

RelaWavity Simulation
Shifted ($b=2$) Phase and Group Vectors in per-Time vs per-Space





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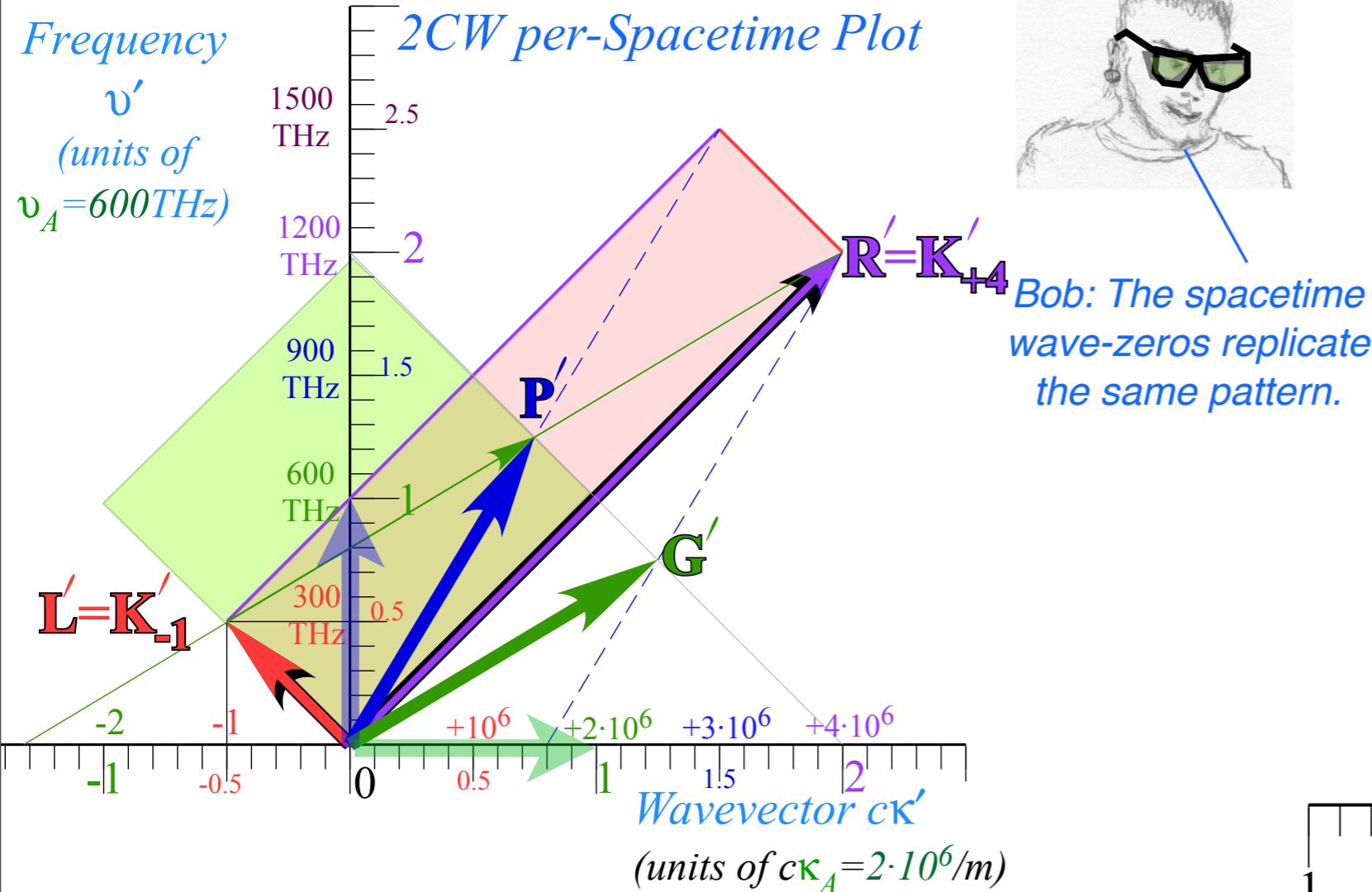
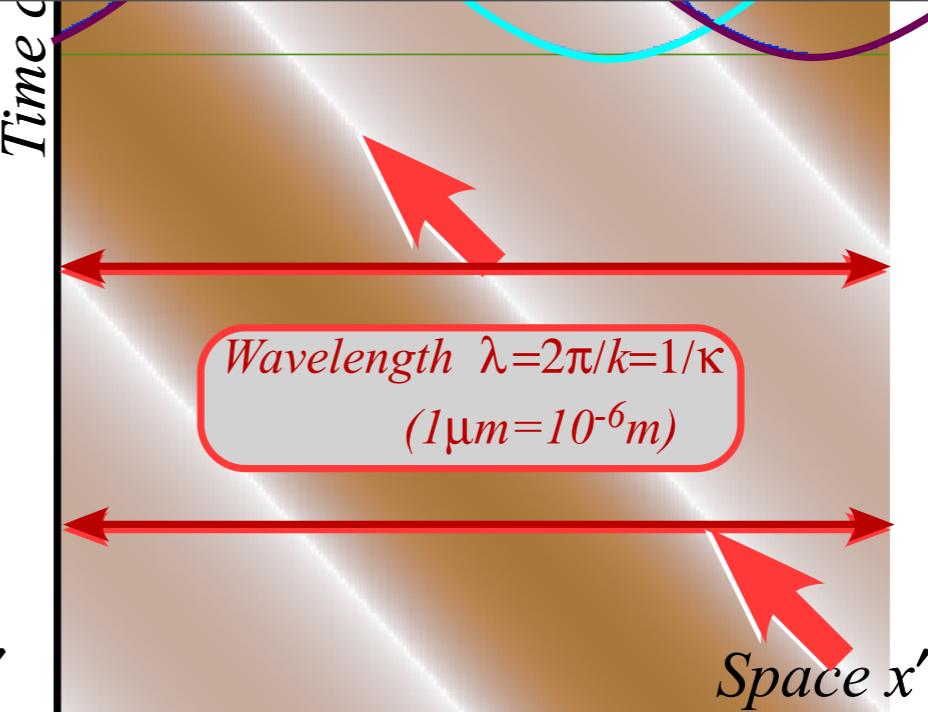
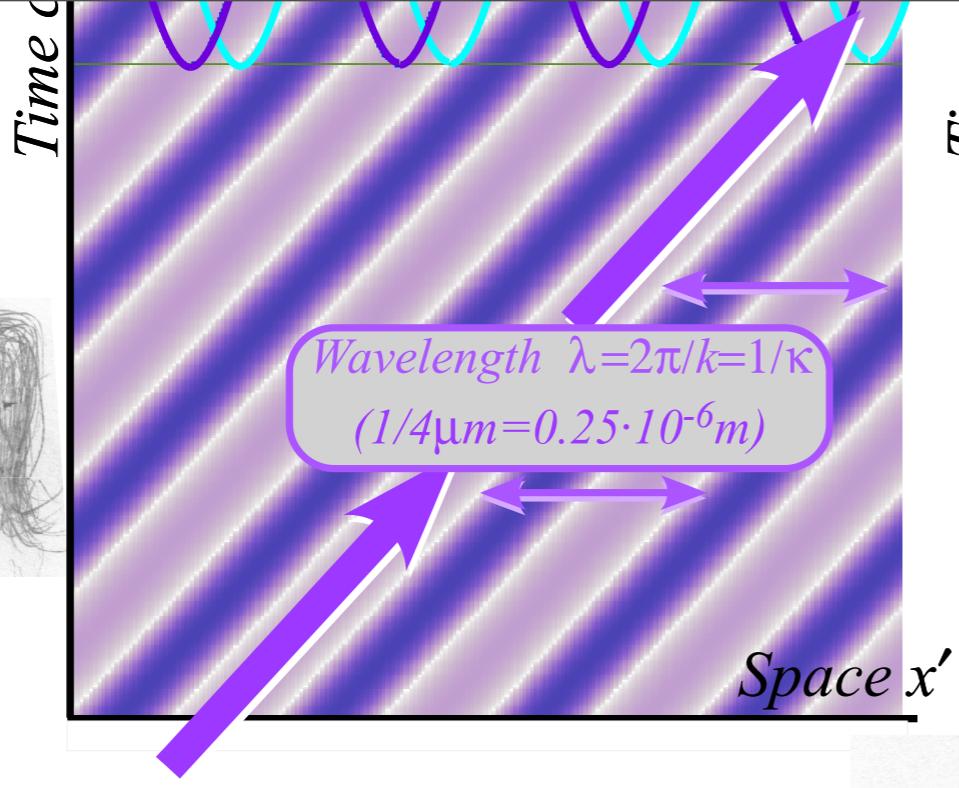
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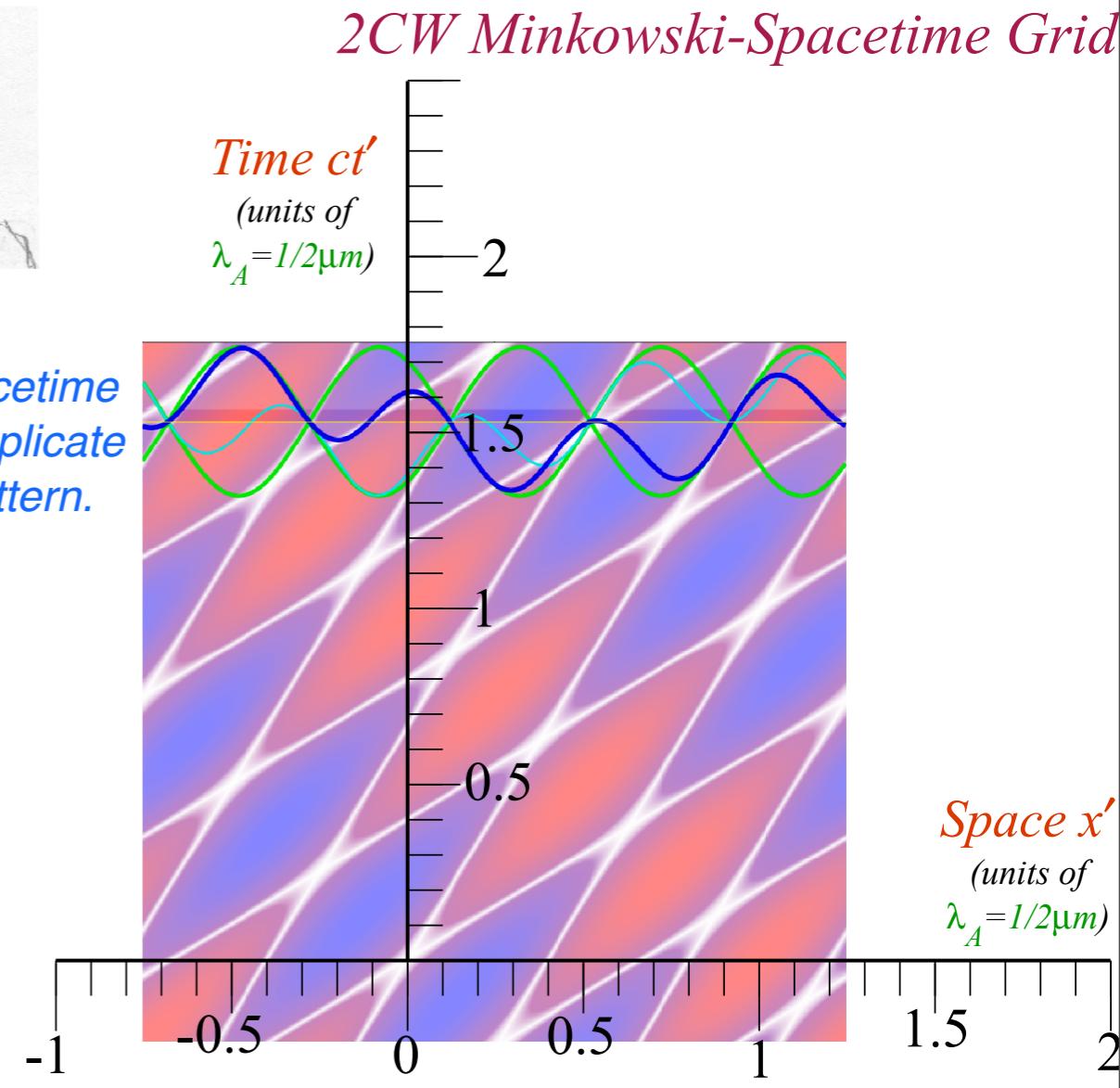


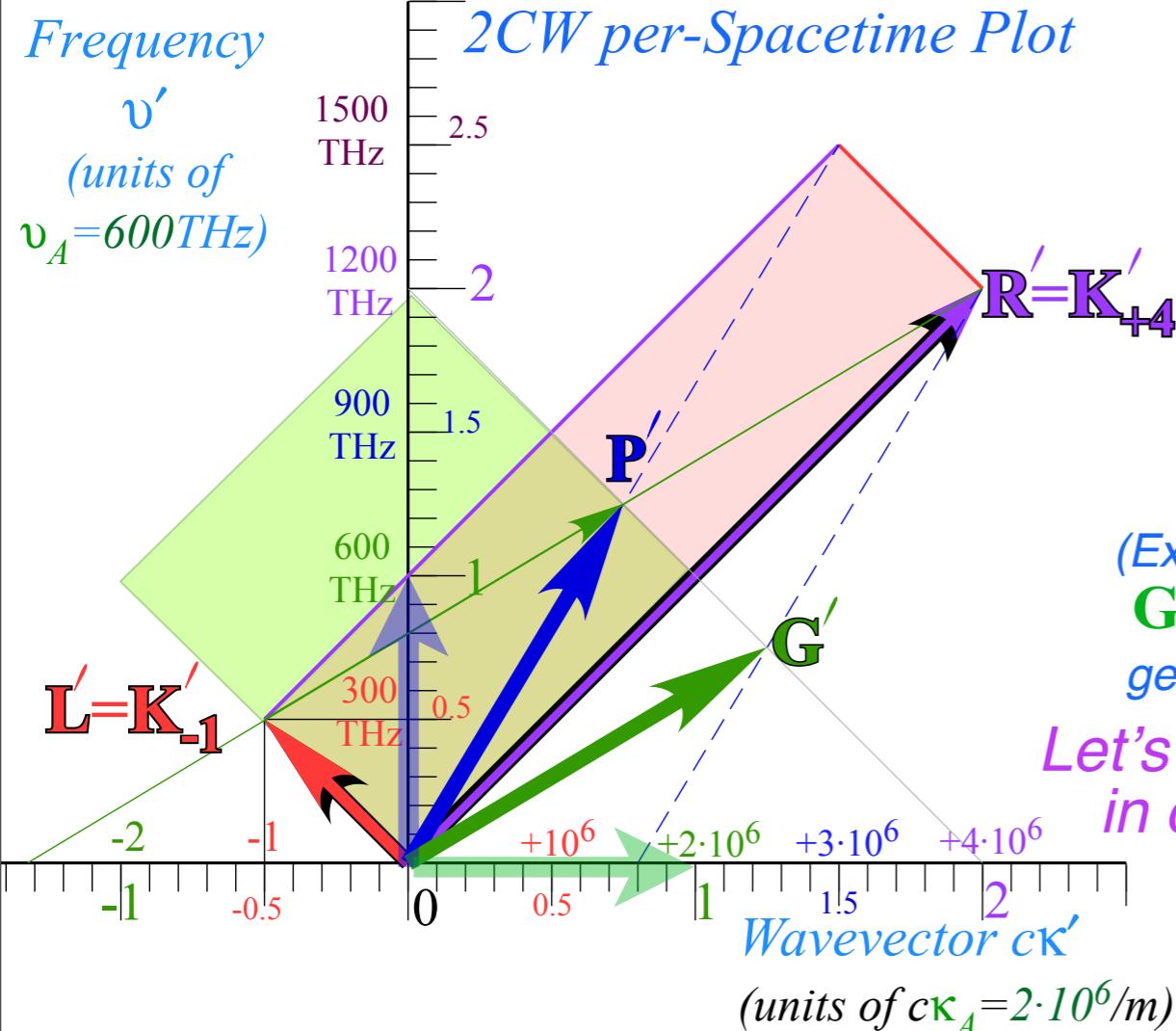
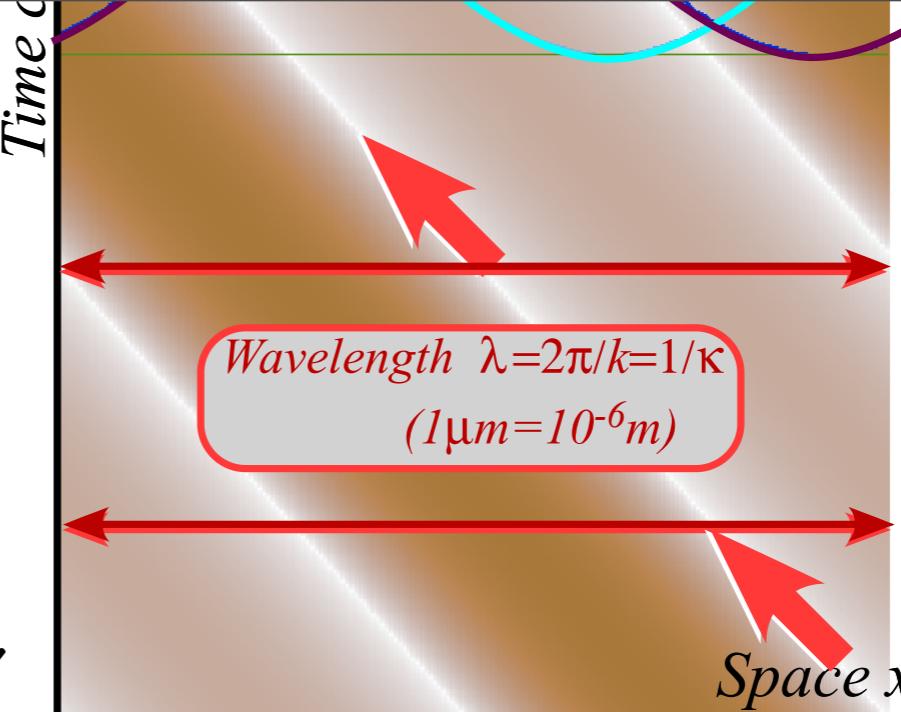
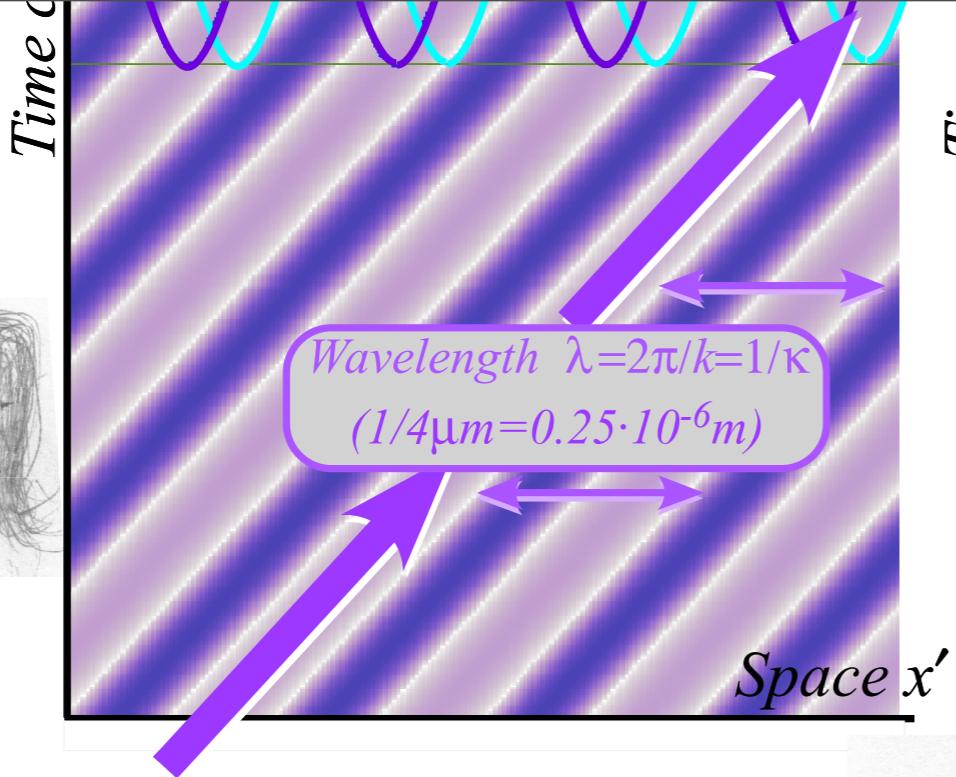
Phase vector P
1/2-sum vector

$$K'_{phase} = \frac{P' + L'}{2}$$

Group vector G
1/2-diff vector

$$K'_{group} = \frac{G' - L'}{2}$$



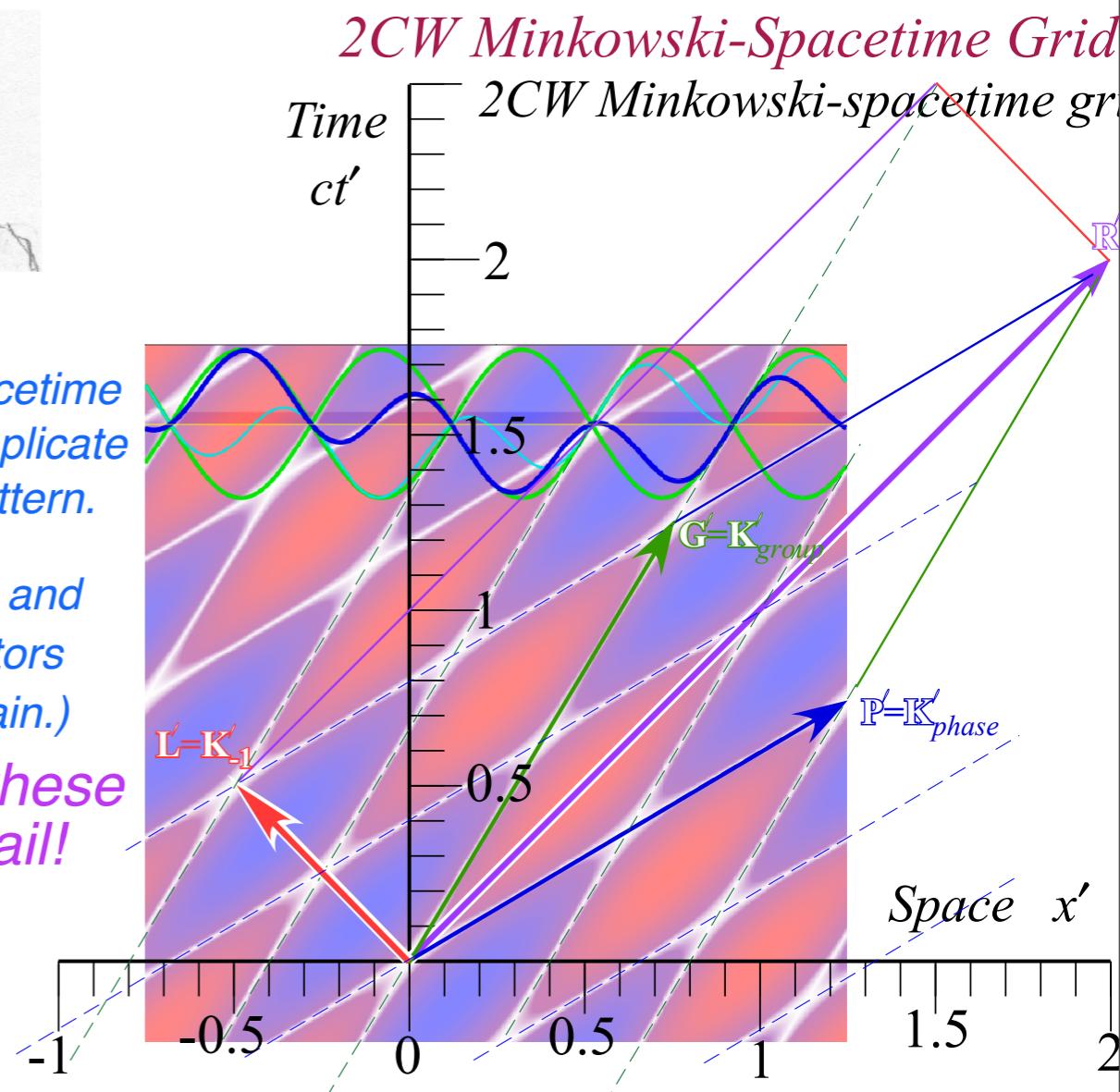


Phase vector P
1/2-sum vector

$$K'_{\text{phase}} = \frac{P' + L'}{2}$$

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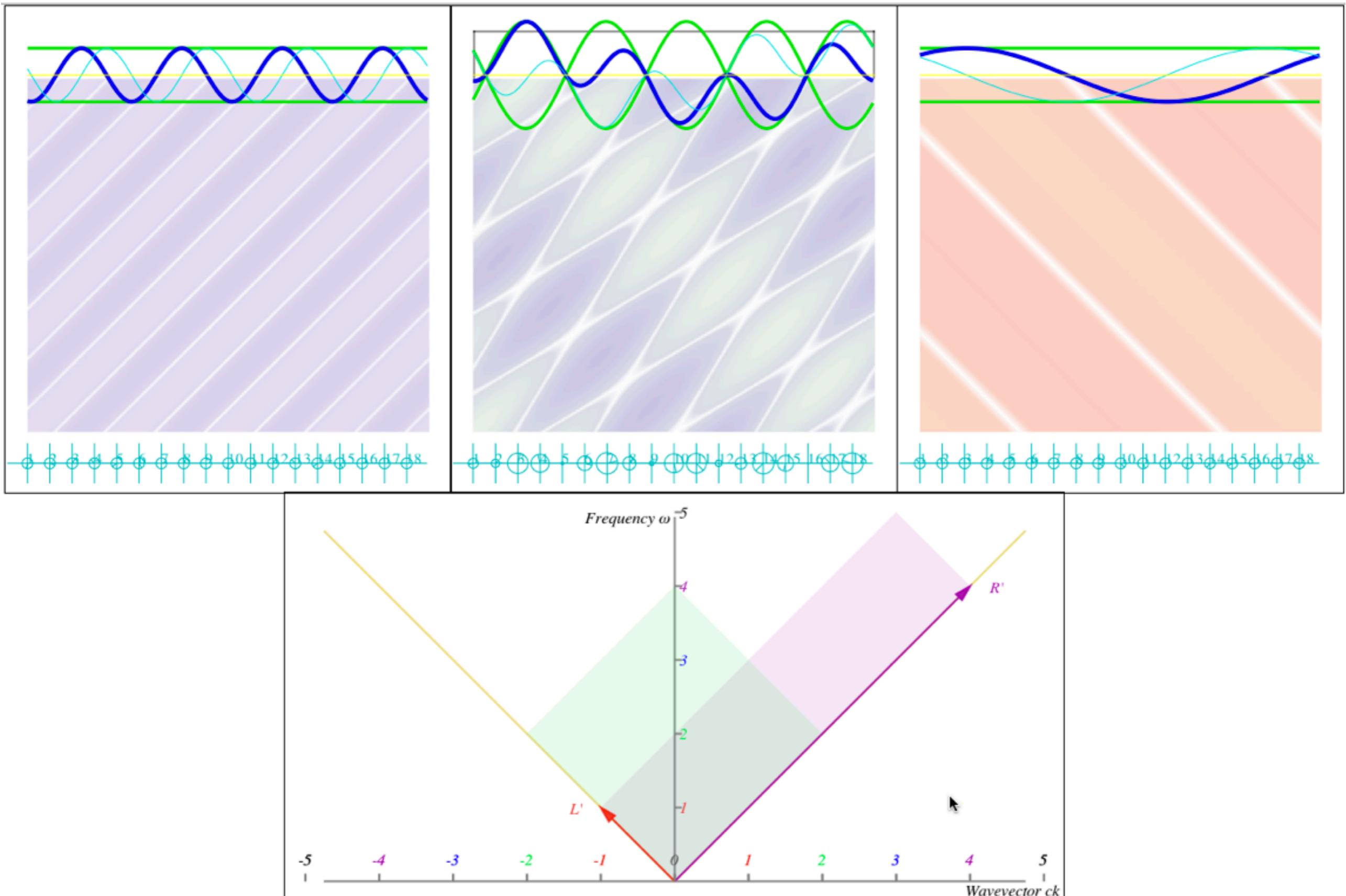
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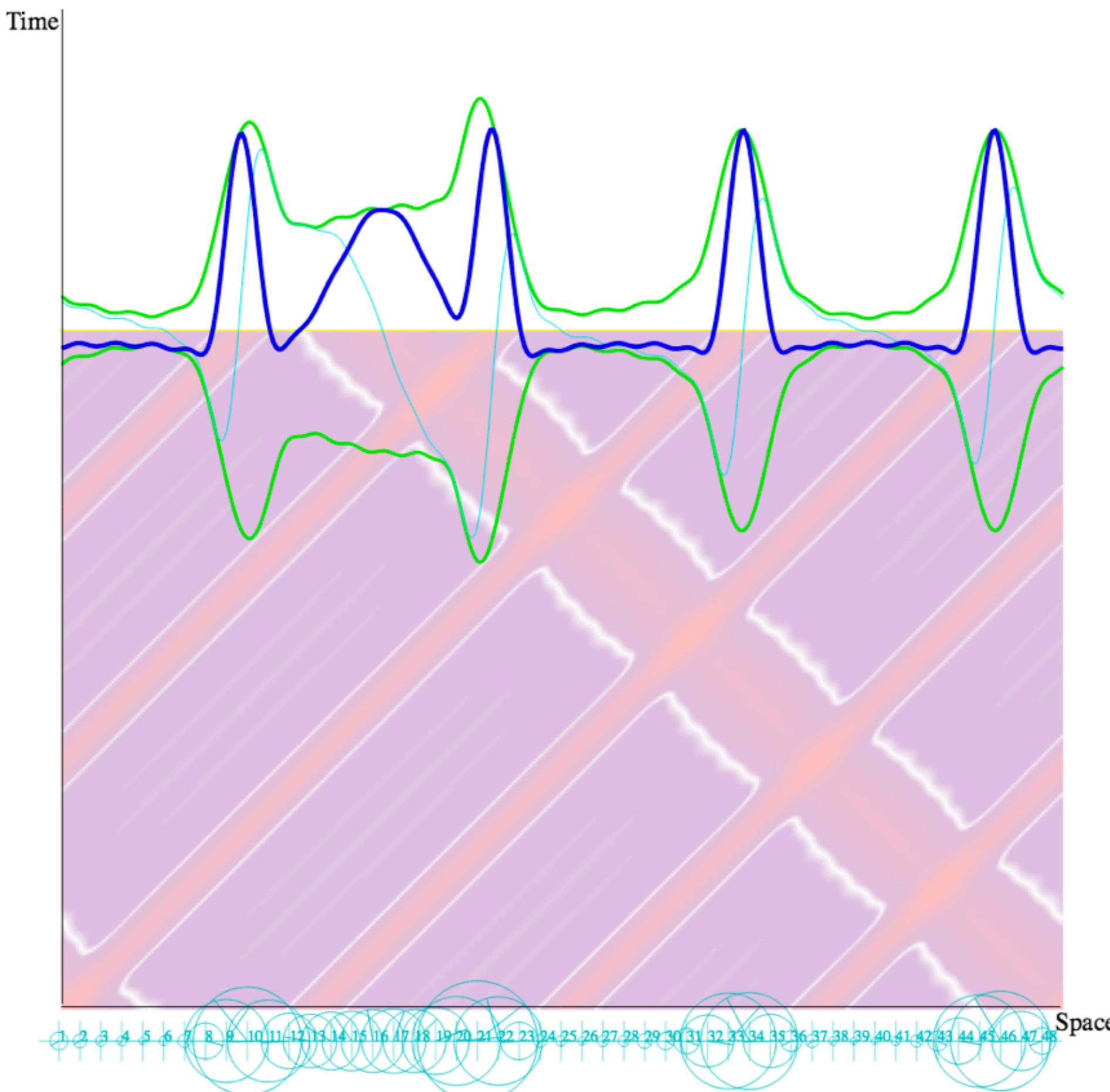
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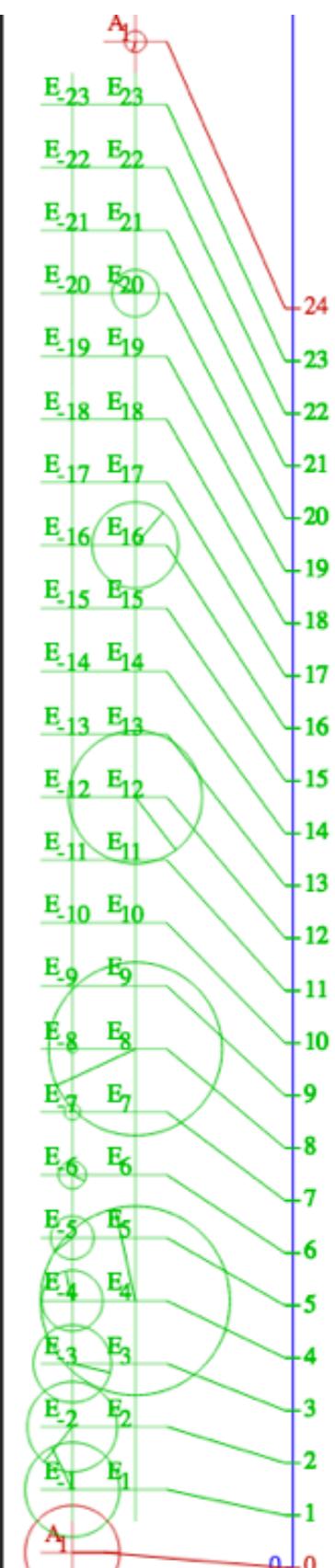
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[BohrIt Web Simulation](#)
[2 CW Minkowski Plot \(\$ck = -1, +4\$ \)](#)



BohrIt Web Simulation
2 PW ct vs x Plot ($\beta = u/c = 3/5$)



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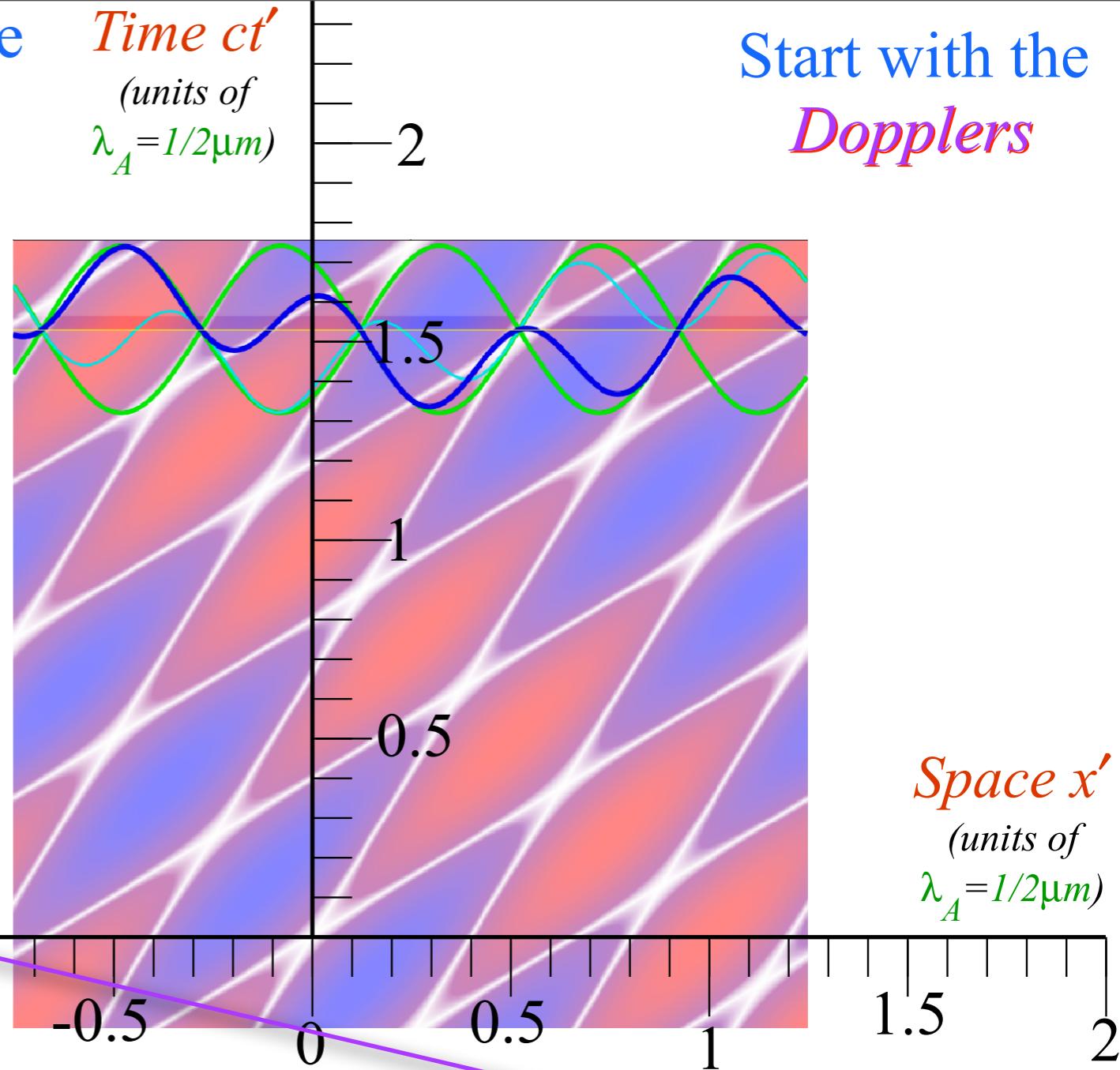
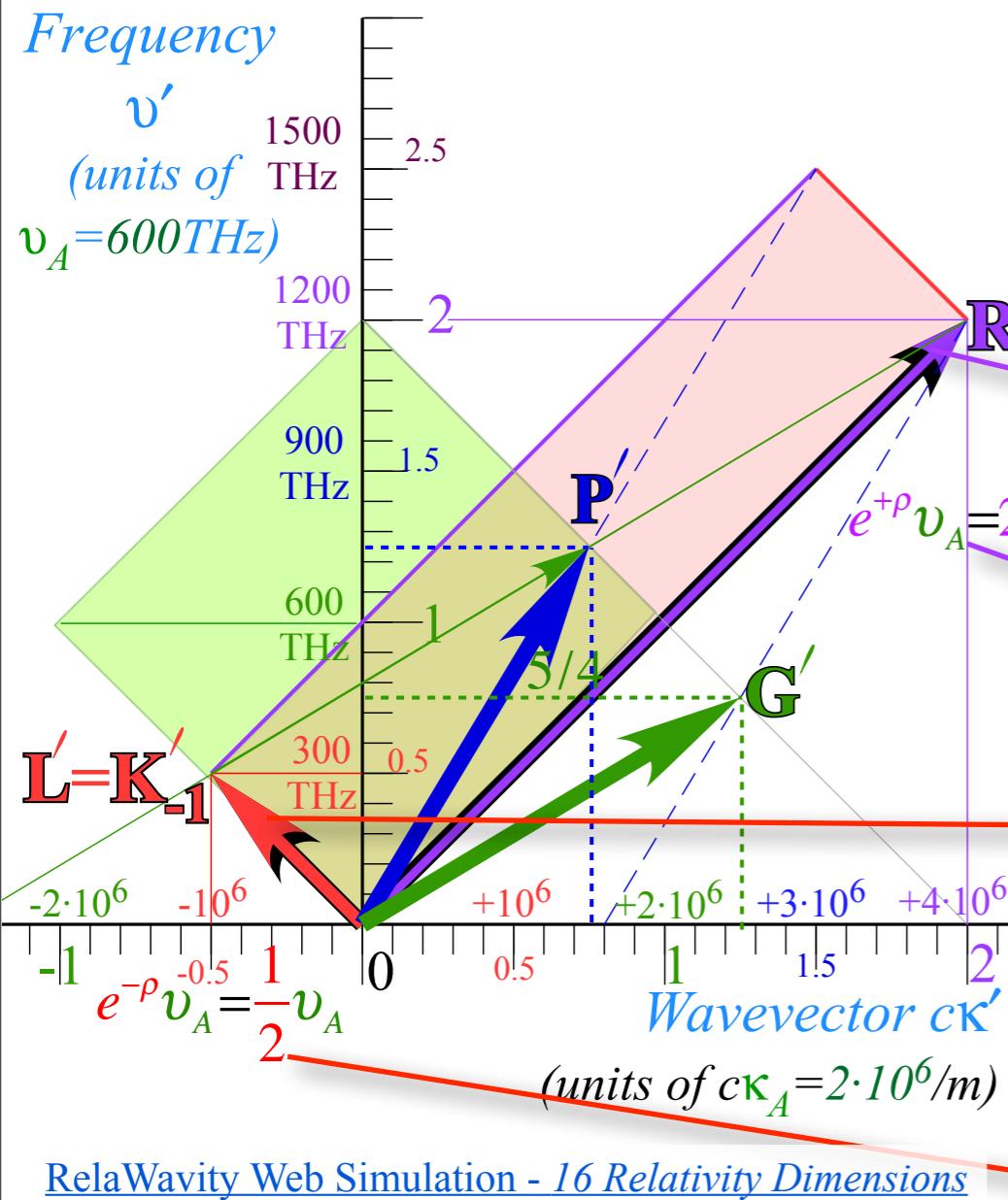
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The 16 parameters of 2CW interference



phase	$b_{\text{Doppler RED}}$	$\frac{c}{V_{\text{phase}}}$	$\frac{\kappa_{\text{phase}}}{K_A}$	$\frac{\tau_{\text{phase}}}{\tau_A}$	$\frac{v_{\text{phase}}}{v_A}$	$\frac{\lambda_{\text{phase}}}{\lambda_A}$	$\frac{V_{\text{phase}}}{c}$	$b_{\text{Doppler BLUE}}$
group	$b_{\text{Doppler BLUE}}$	$\frac{1}{c}$	$\frac{V_{\text{group}}}{c}$	$\frac{v_{\text{group}}}{v_A}$	$\frac{\lambda_{\text{group}}}{\lambda_A}$	$\frac{K_{\text{group}}}{K_A}$	$\frac{\tau_{\text{group}}}{\tau_A}$	$\frac{1}{b_{\text{Doppler RED}}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

Start with the
Dopplers

The 16 parameters of 2CW interference

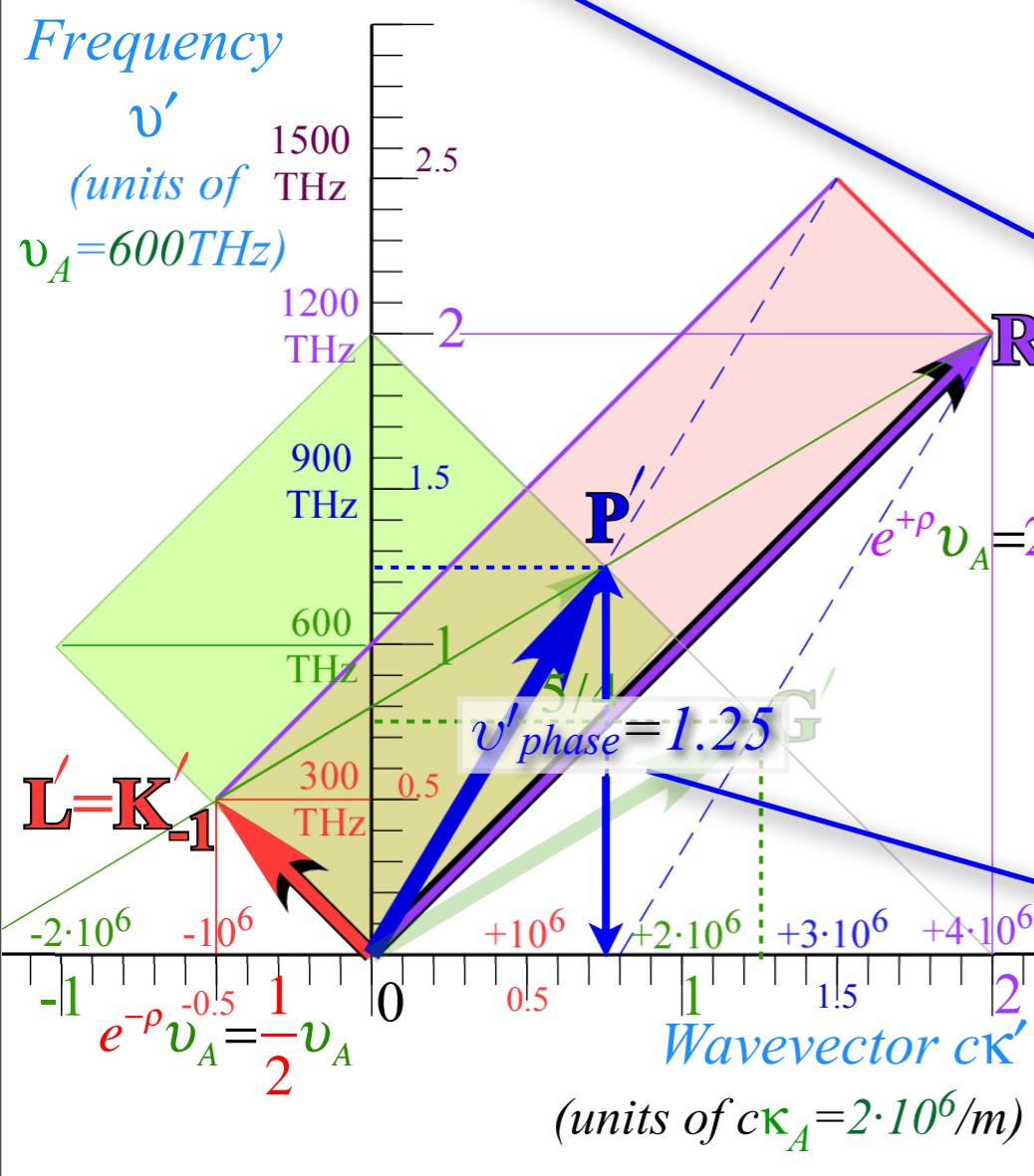
$$\mathbf{P}' = \begin{pmatrix} c\kappa'_{phase} \\ v'_{phase} \end{pmatrix} = v_A \begin{pmatrix} \sinh \rho \\ \cosh \rho \end{pmatrix} = v_A \begin{pmatrix} 3/4 \\ 5/4 \end{pmatrix}$$

Phase frequency
 $v'_{phase} = v_A \cosh \rho = 5/4 = 1.25$

flips to

Phase period $\tau = 1/v$
 $\tau'_{phase} = \tau_A \operatorname{sech} \rho = 4/5 = 0.8$

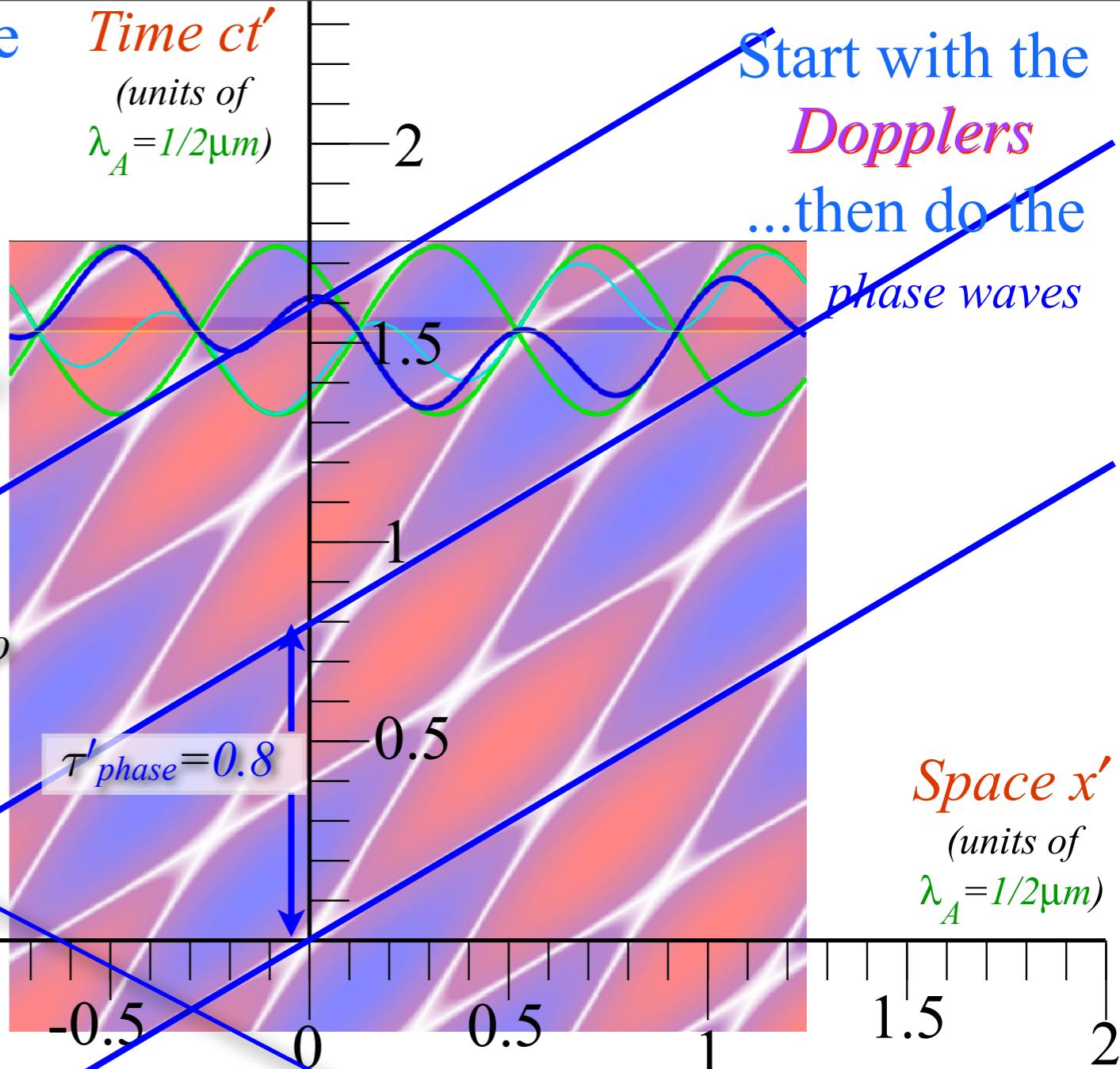
$$\tau'_{phase} = \tau_A \operatorname{sech} \rho = 4/5$$



RelaWavity Web Simulation - 16 Relativity Dimensions

Time ct'

(units of $\lambda_A = 1/2 \mu m$)



$\tau'_{phase} = 0.8$

phase	$b_{RED}^{Doppler}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
group	1	V_{group}	v_{group}	λ_{group}	κ_{group}	τ_{group}	c	$b_{RED}^{Doppler}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
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Start with the
Dopplers
...then do the
phase waves

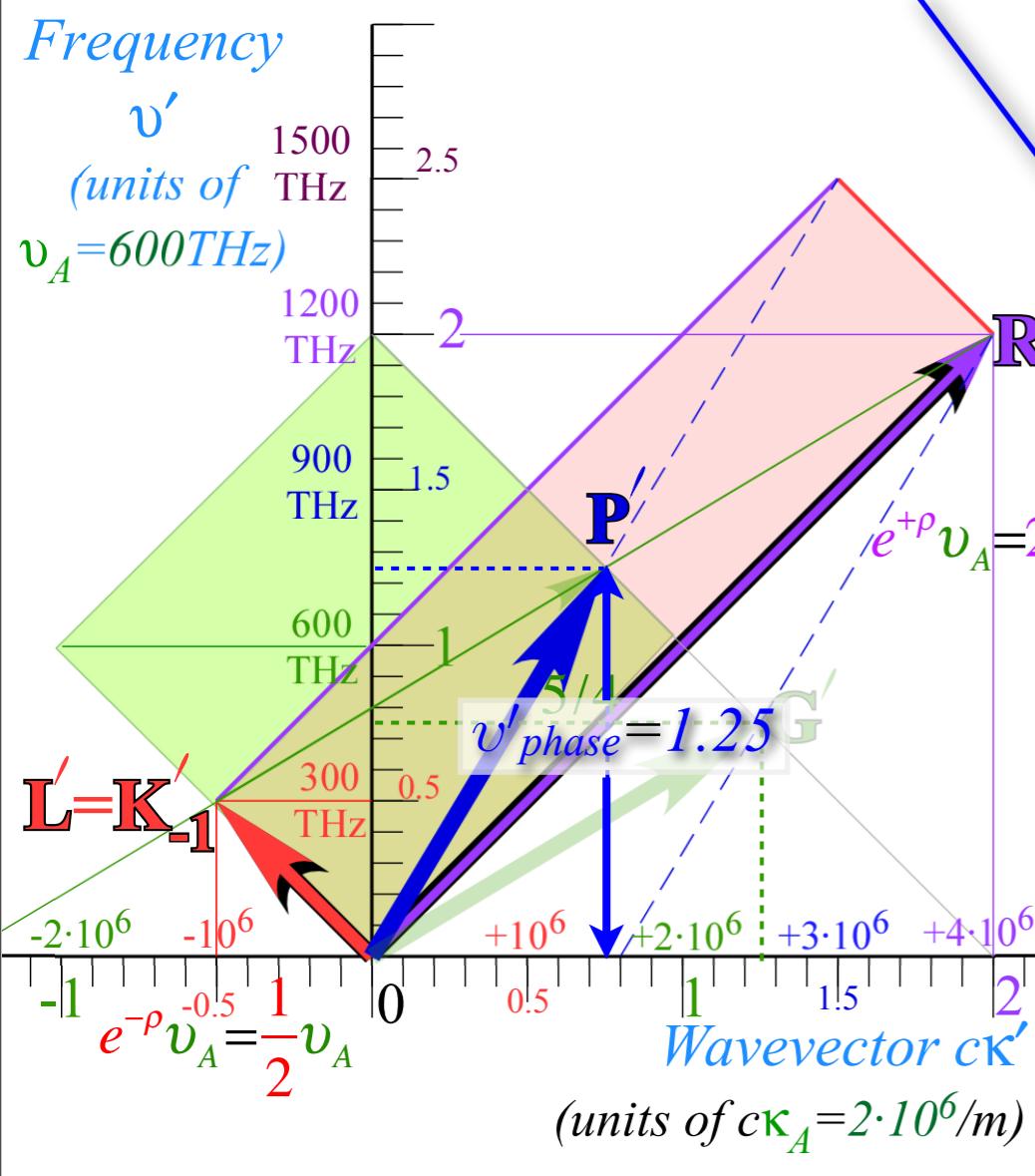
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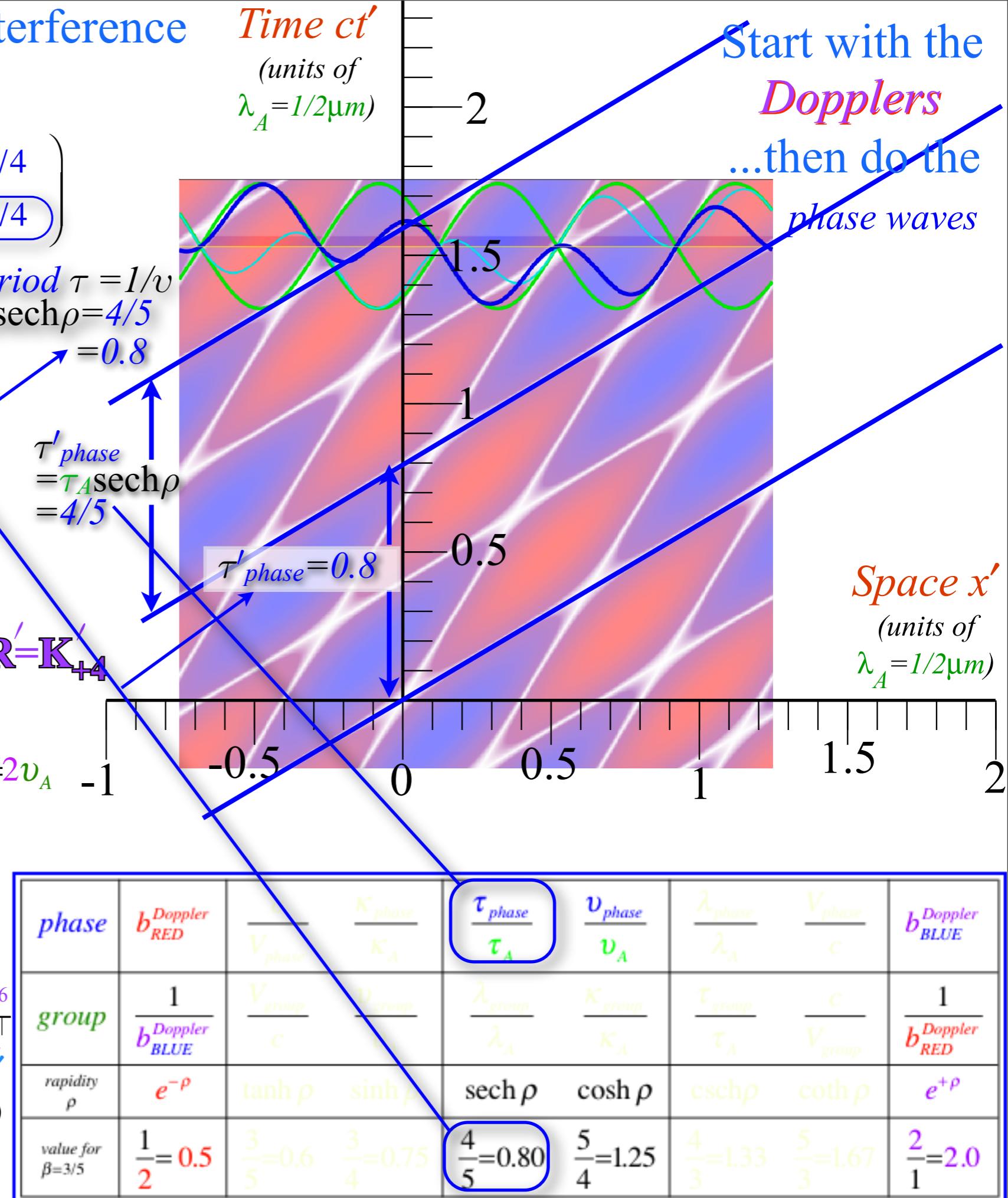
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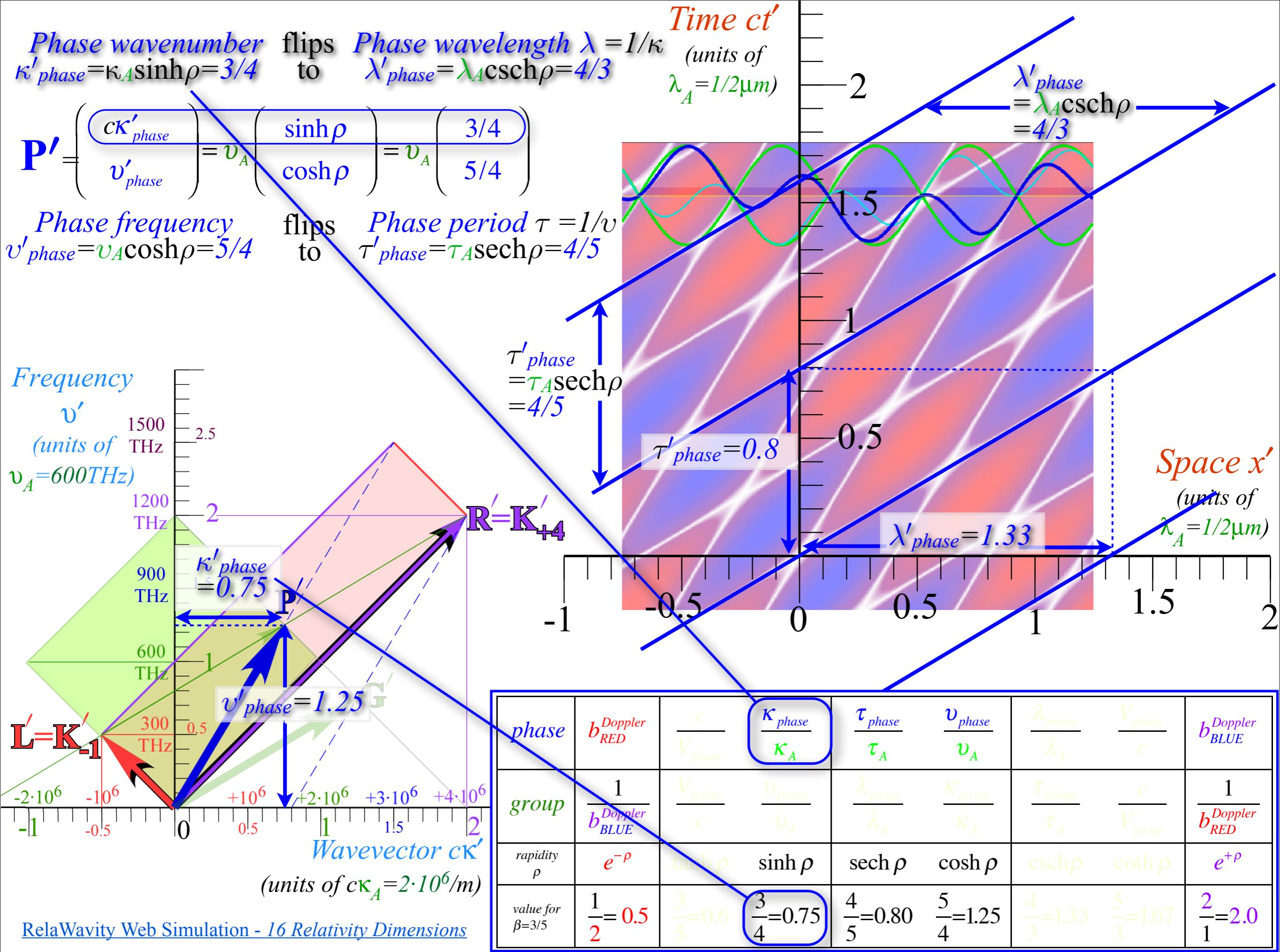
flips to

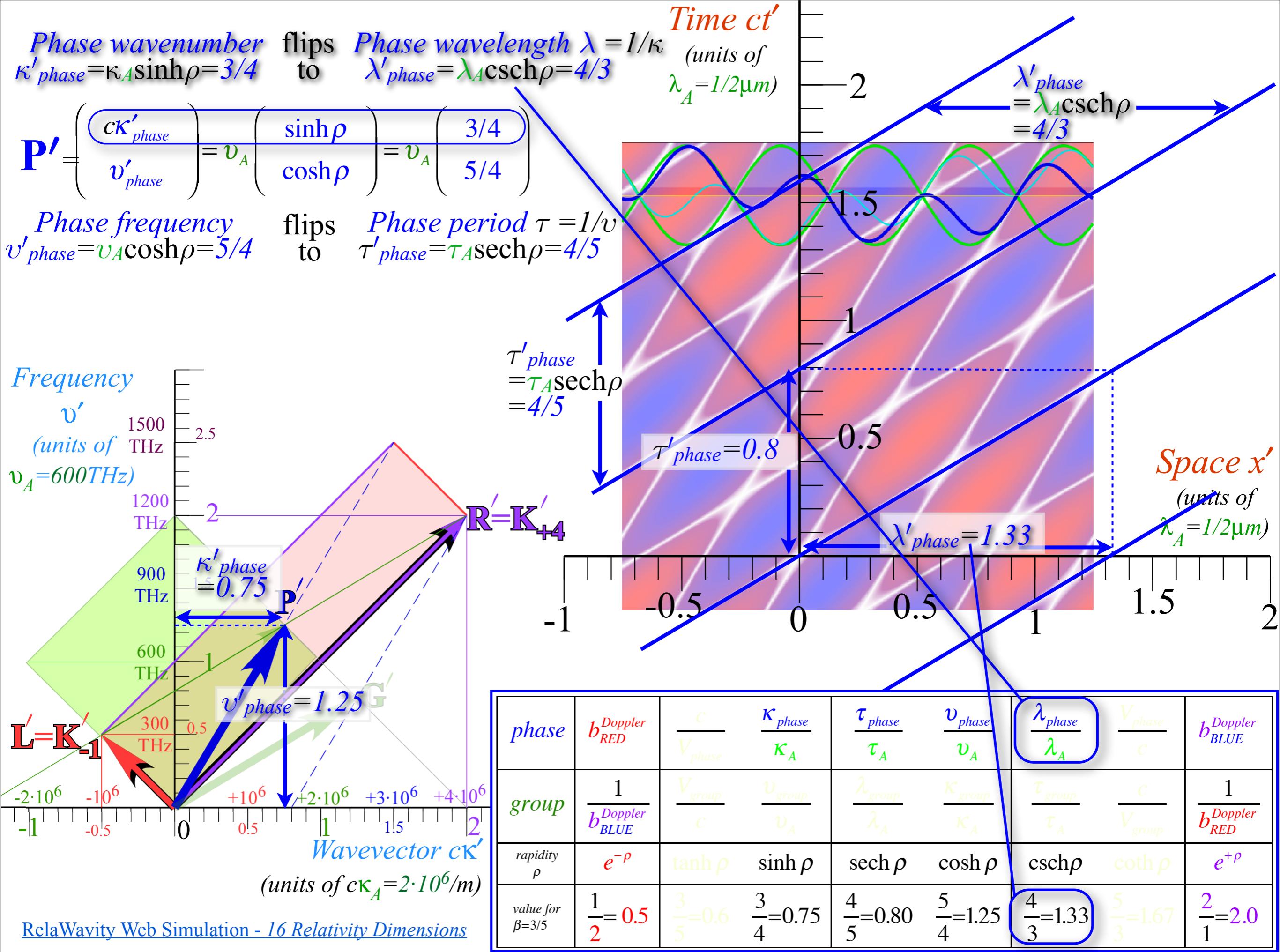
Phase period $\tau = 1/v$
 $\tau'_{phase} = \tau_A \operatorname{sech} \rho = 4/5$

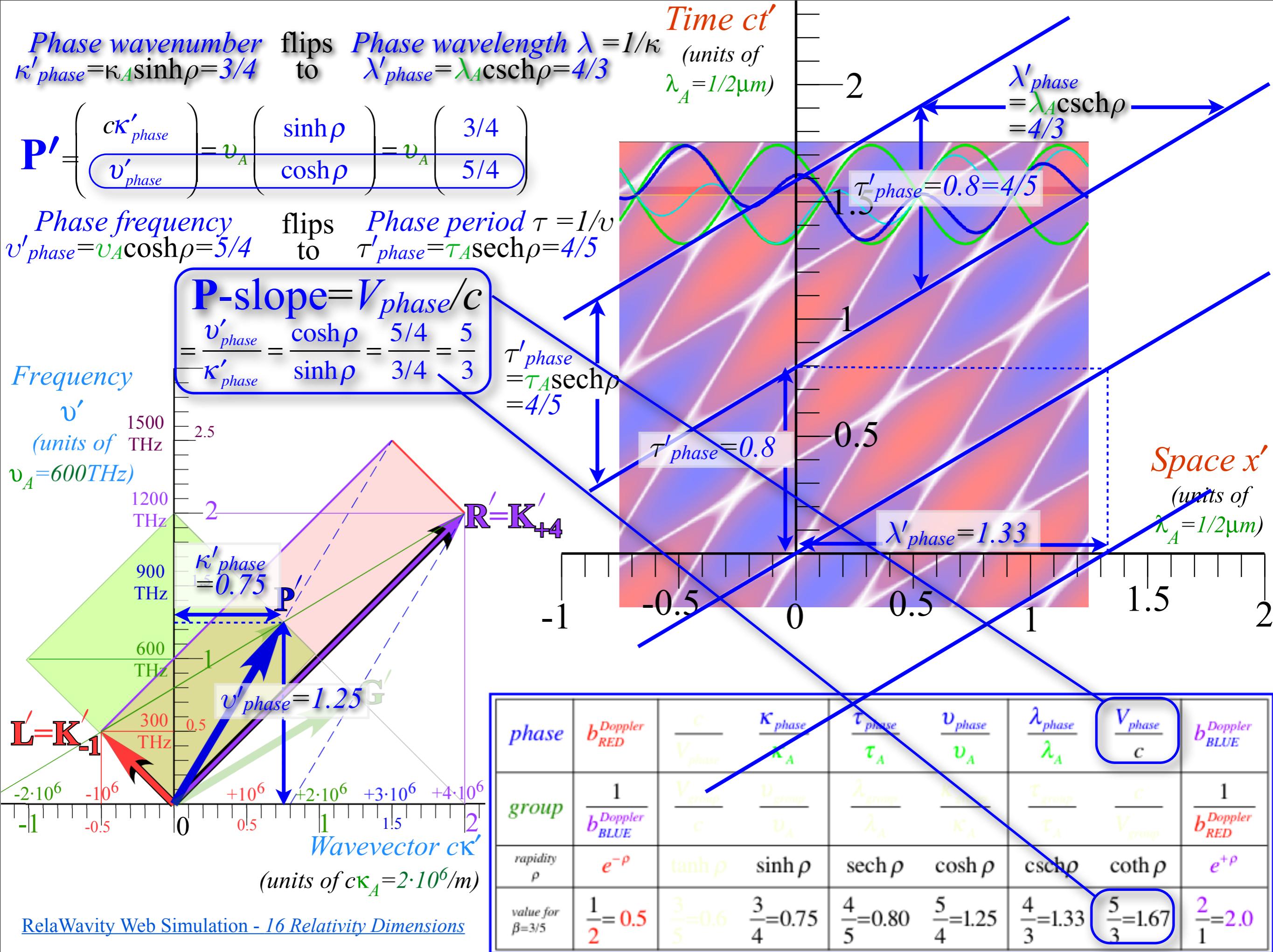


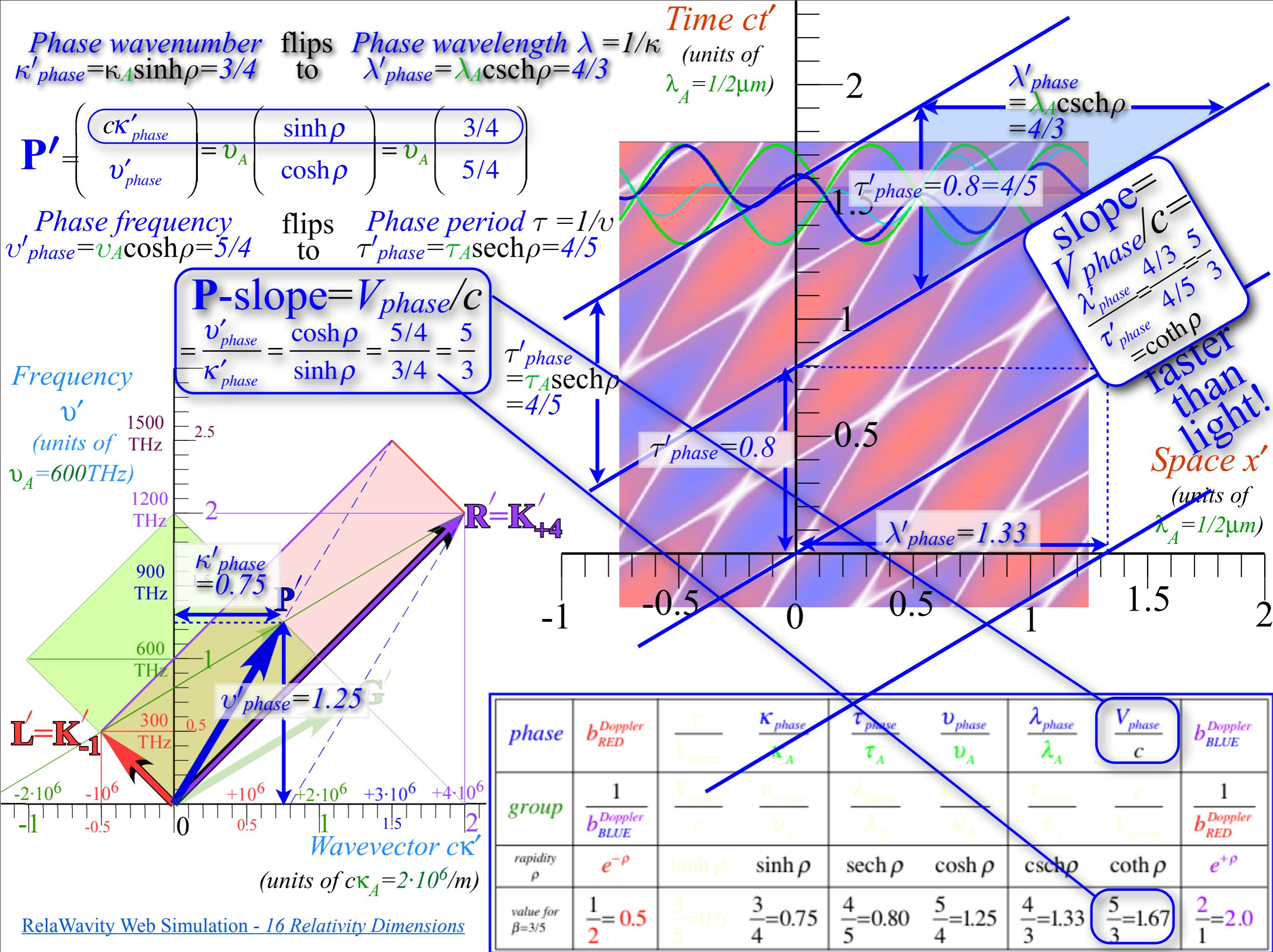
RelaWavity Web Simulation - 16 Relativity Dimensions











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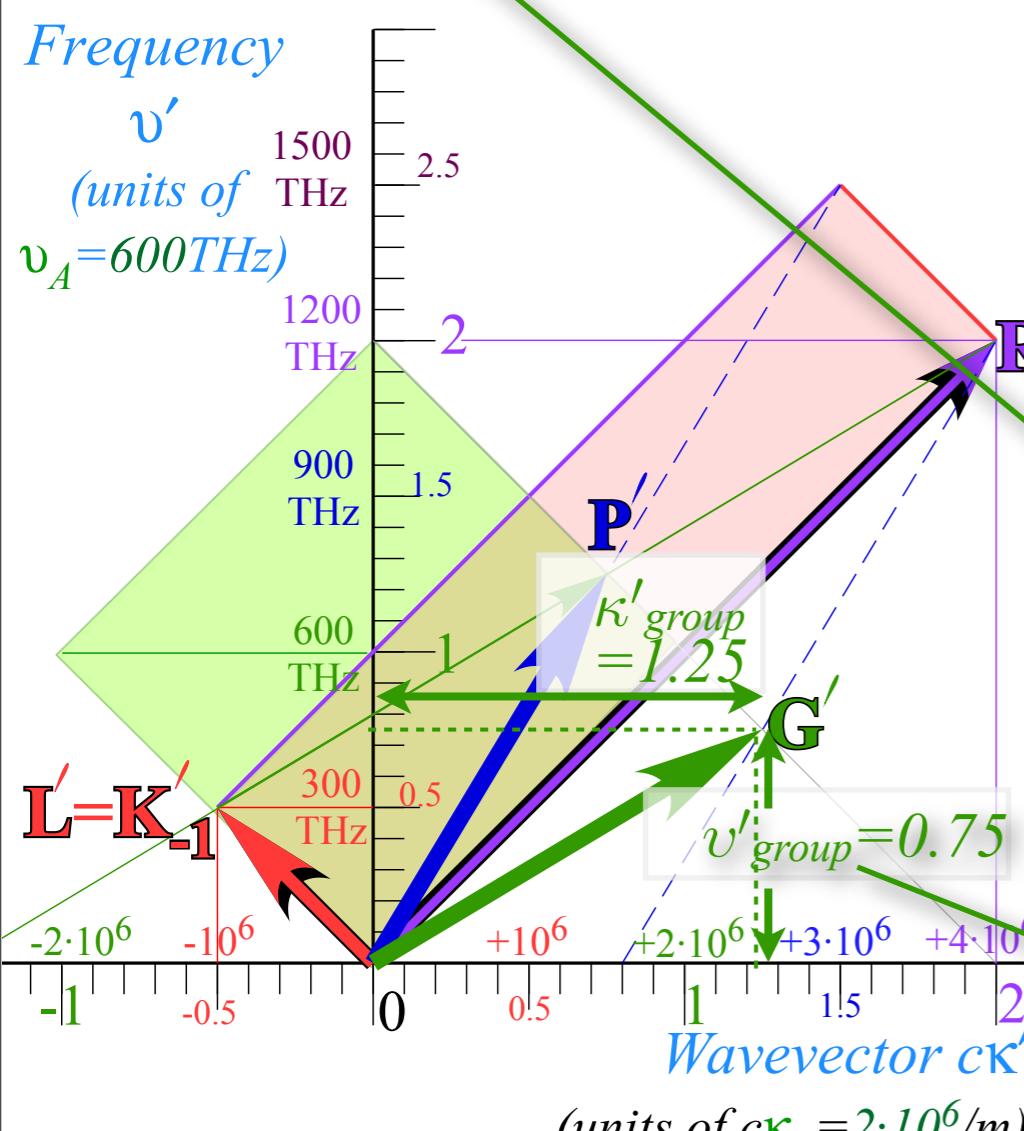
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Group frequency
 $v'_{group} = v_A \sinh \rho = 3/4 = 0.75$

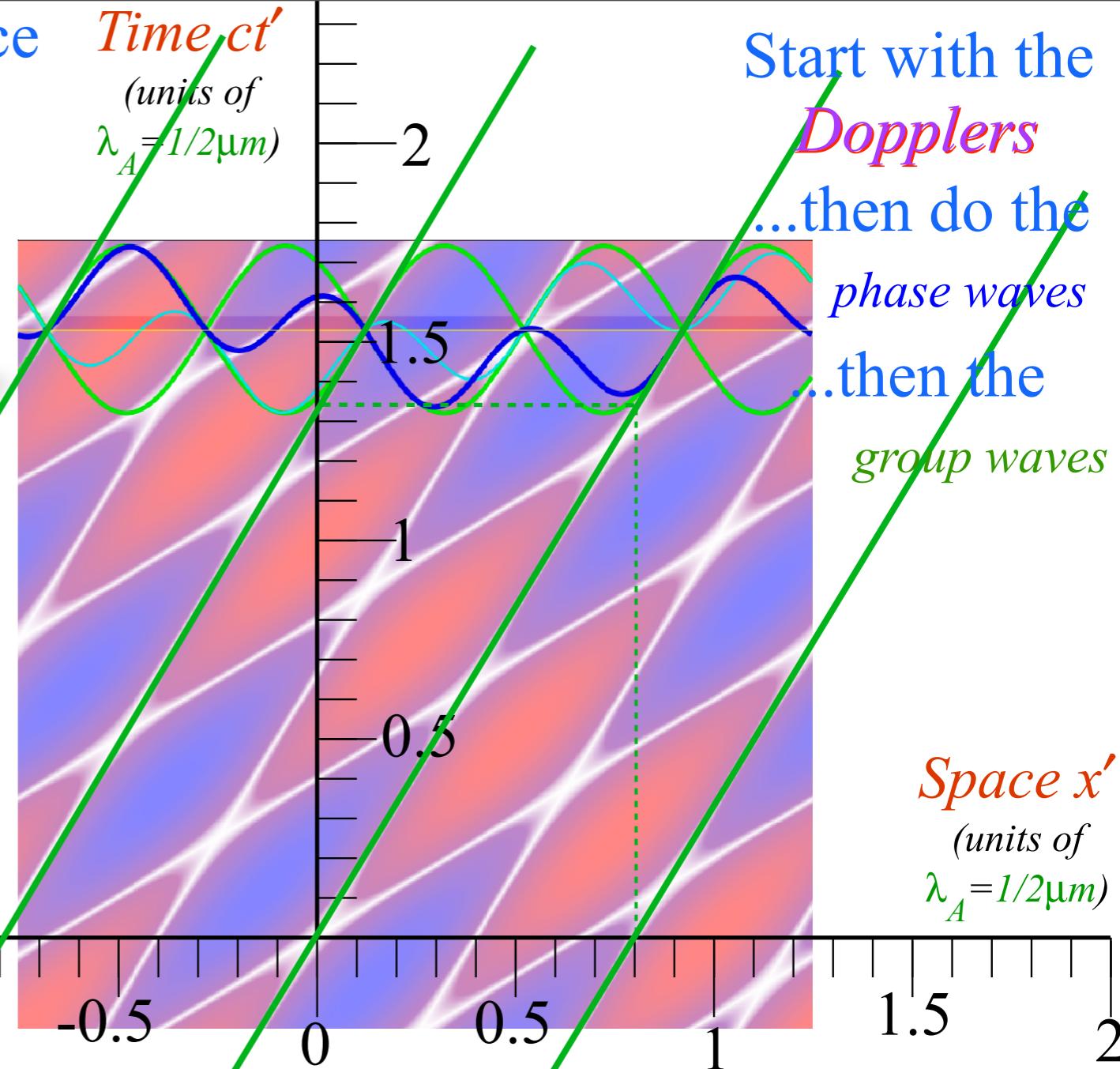
flips to

Group period $\tau = 1/v$
 $\tau'_{group} = \tau_A \text{csch} \rho = 4/3 = 1.33$



RelaWavity Web Simulation - 16 Relativity Dimensions

phase	$b_{Doppler}^{RED}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler}^{BLUE}$
group	$\frac{1}{b_{Doppler}^{Blue}}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler}^{Red}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

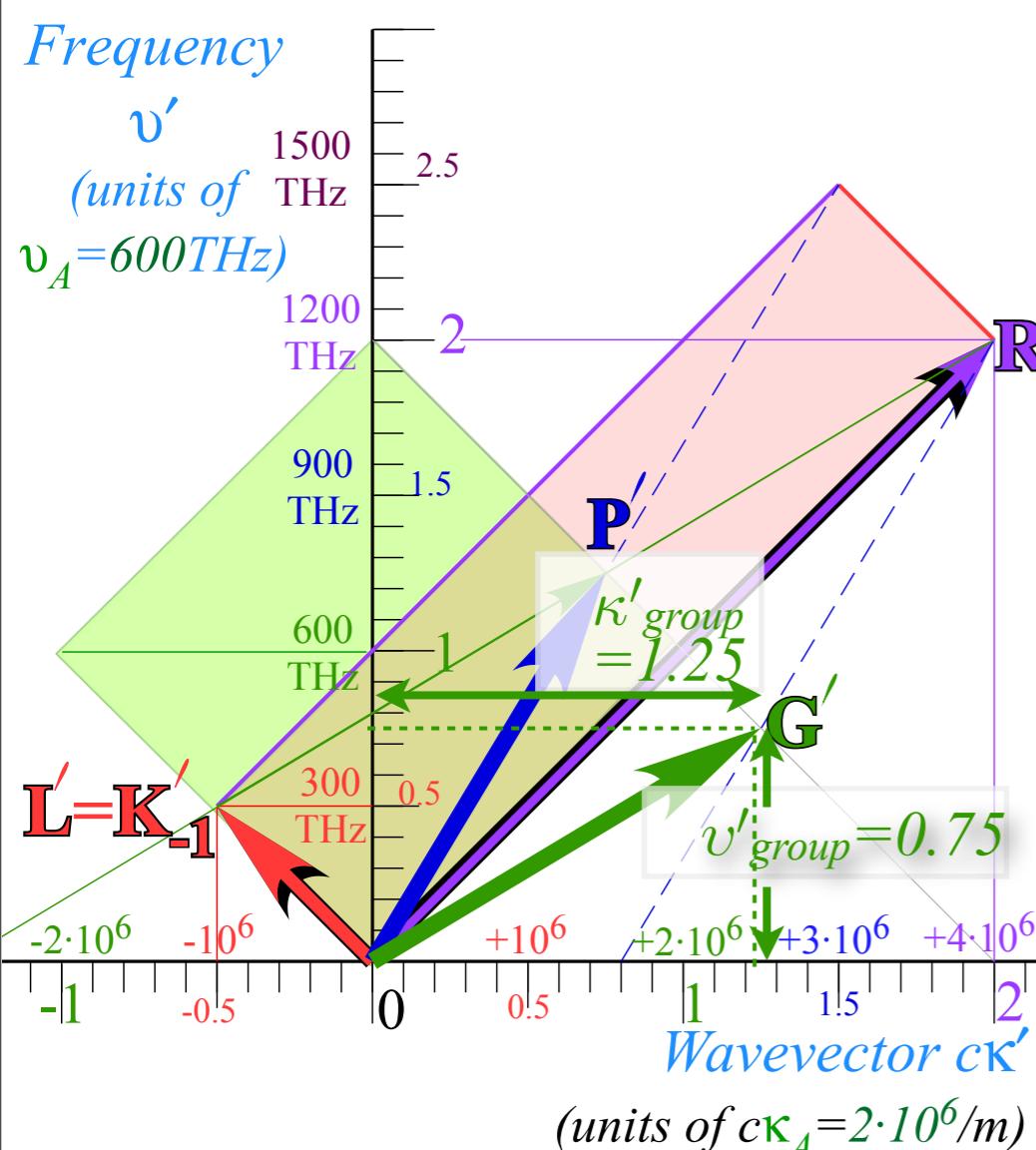


The 16 dimensions of 2CW interference

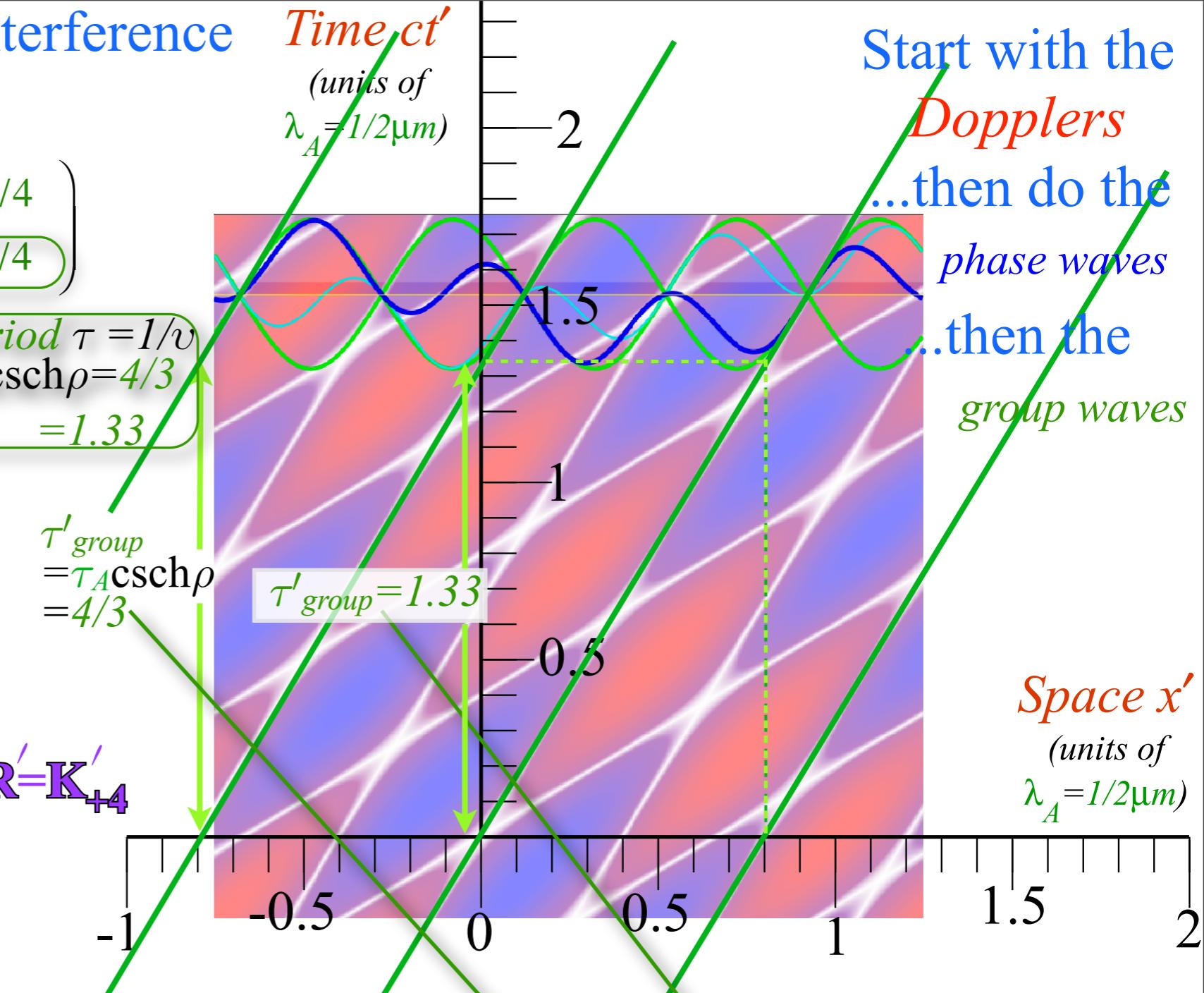
$$\mathbf{G}' = \begin{pmatrix} cK'_{group} \\ v'_{group} \end{pmatrix} = v_A \begin{pmatrix} \cosh \rho \\ \sinh \rho \end{pmatrix} = v_A \begin{pmatrix} 5/4 \\ 3/4 \end{pmatrix}$$

$$Group\ frequency \\ v'_{group} = v_A \sinh \rho = 3/4 \\ = 0.75$$

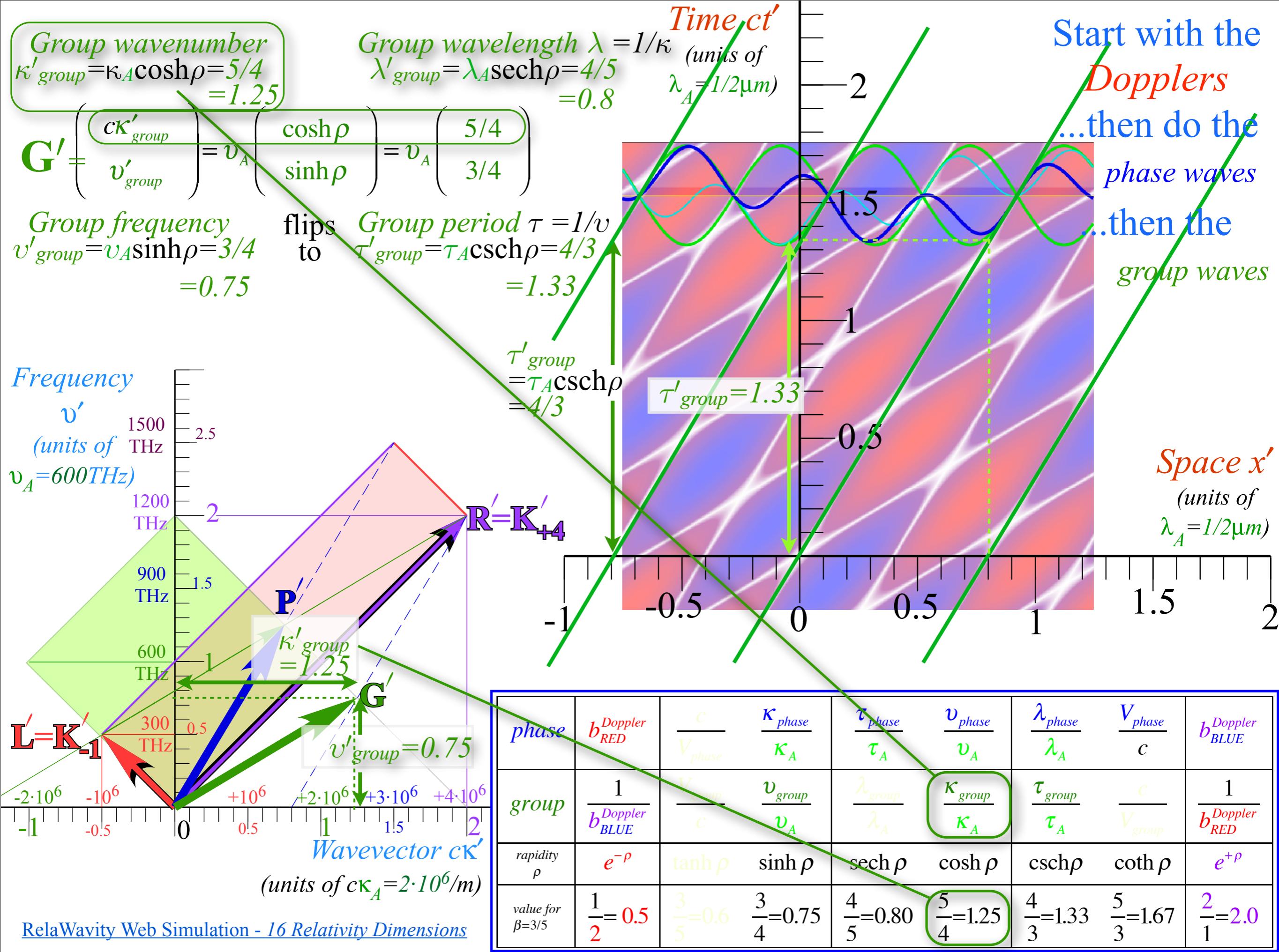
flips to $\tau'_{group} = \tau_A \text{csch} \rho = 4/3 = 1.33$

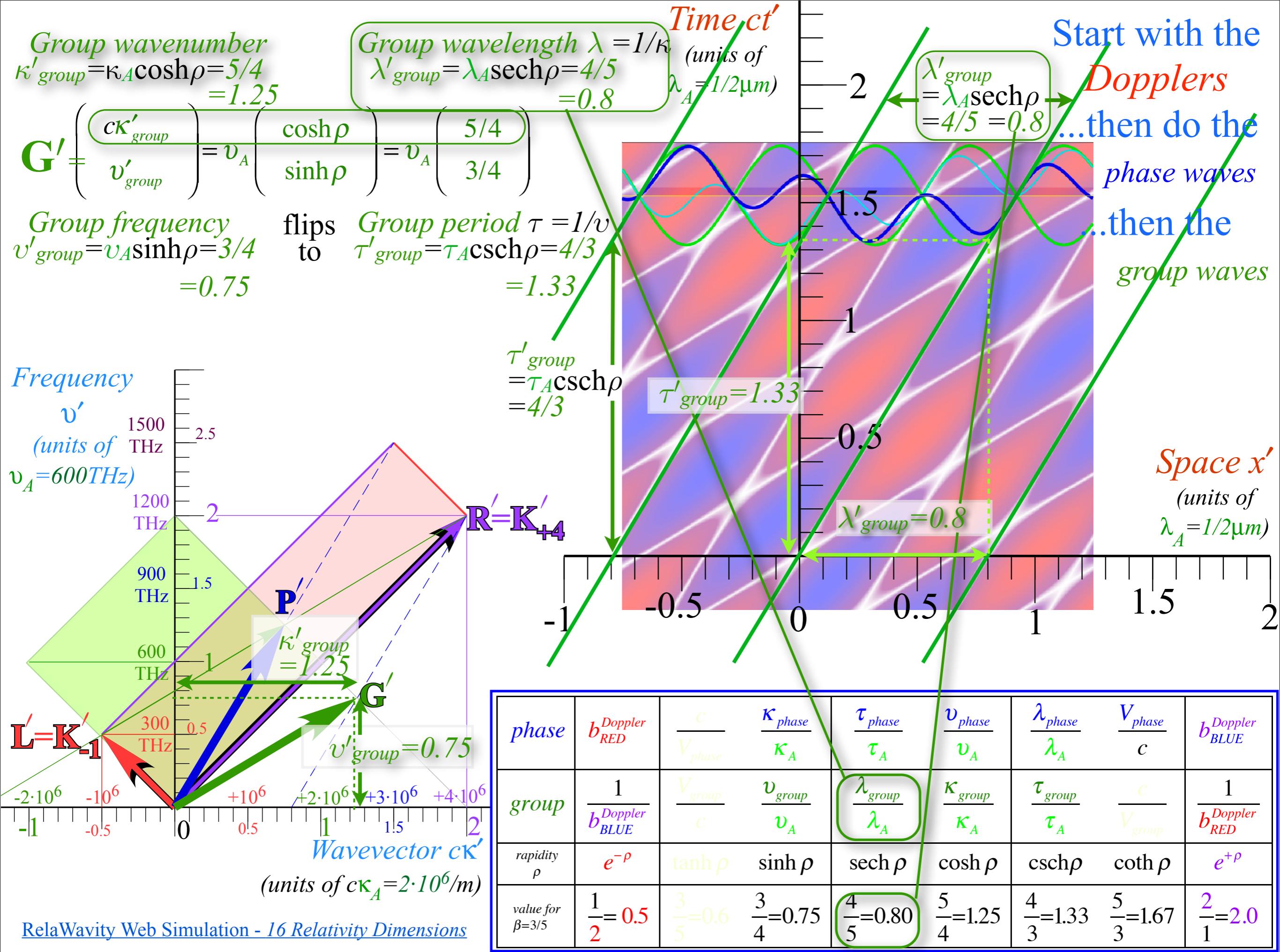


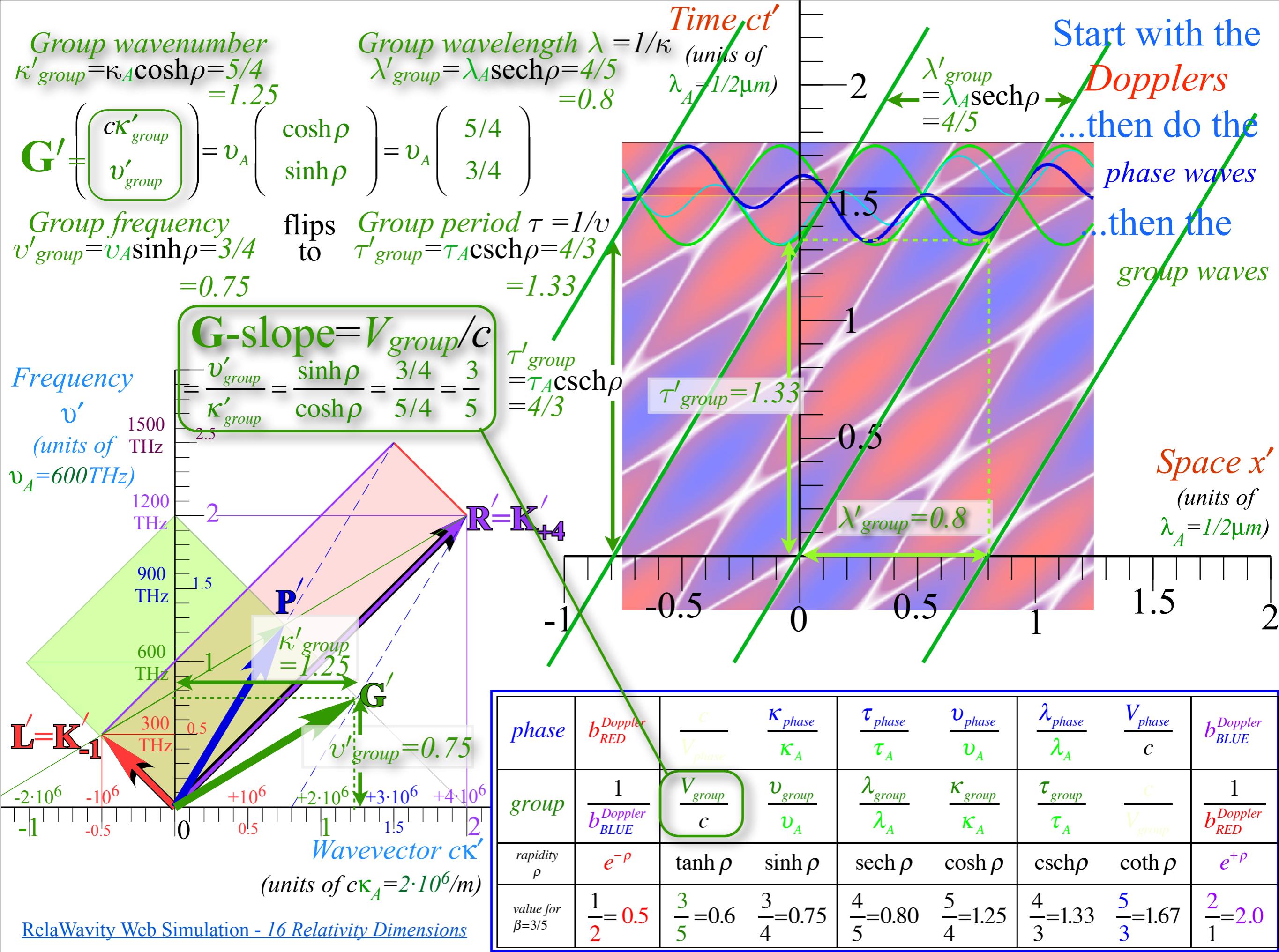
[RelaWavity Web Simulation - 16 Relativity Dimensions](#)

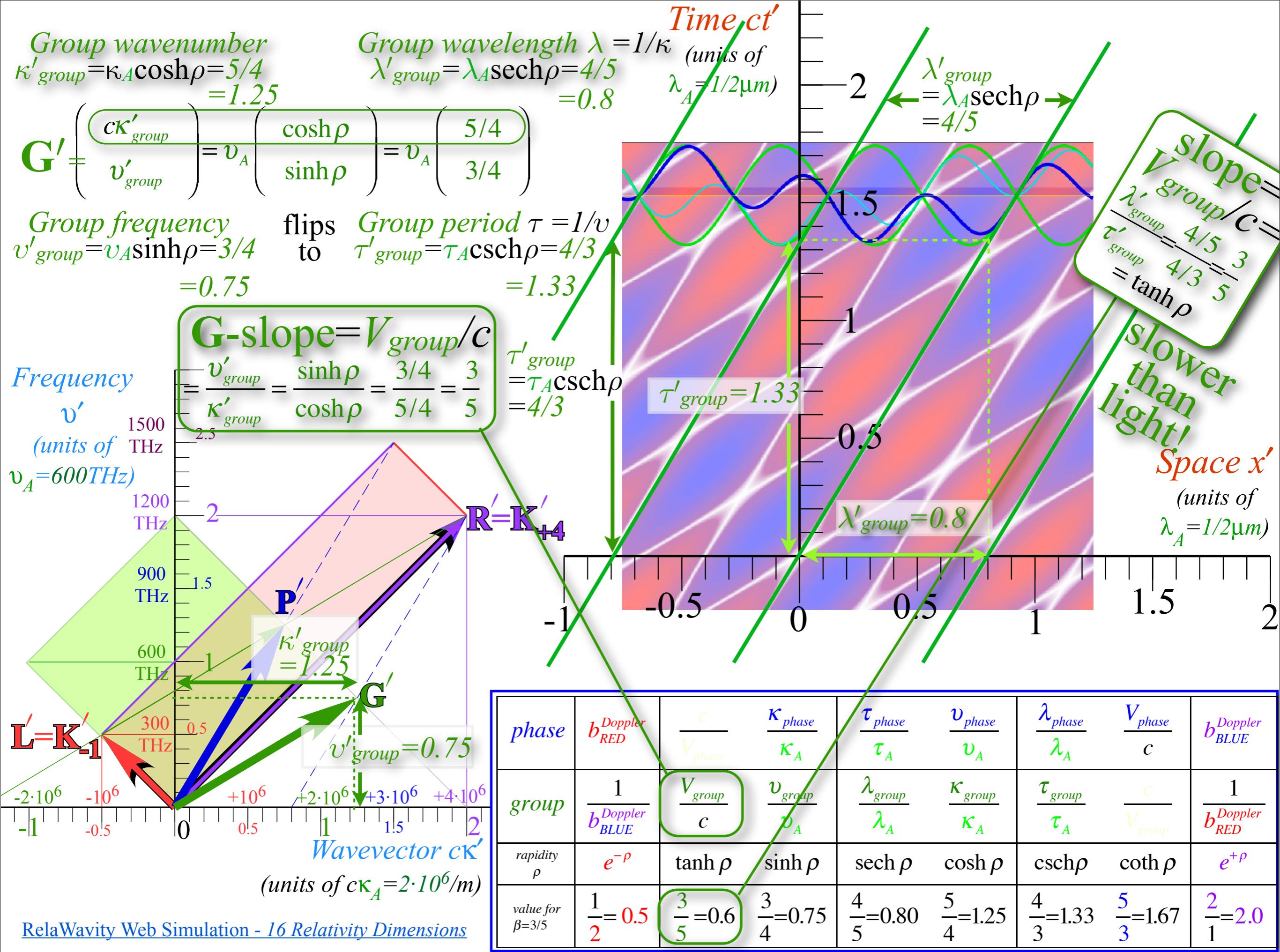


<i>phase</i>	$b_{RED}^{Doppler}$	$\frac{c}{V_{phase}}$	κ_{phase}	τ_{phase}	v_{phase}	λ_{phase}	V_{phase}	$b_{BLUE}^{Doppler}$
<i>group</i>	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{V_{group}}{c}$	v_{group}	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
<i>rapidity</i> ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
<i>value for</i> $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$









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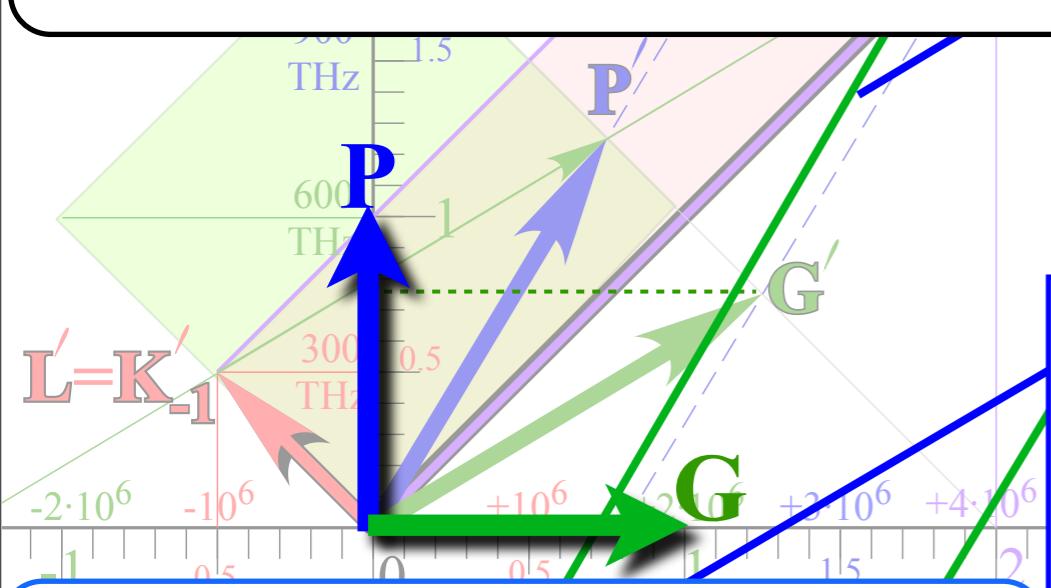
More at Pirelli Challenge page: [*'Un Grande Affare' - Light Meets Light*](#)

Lorentz transformations...

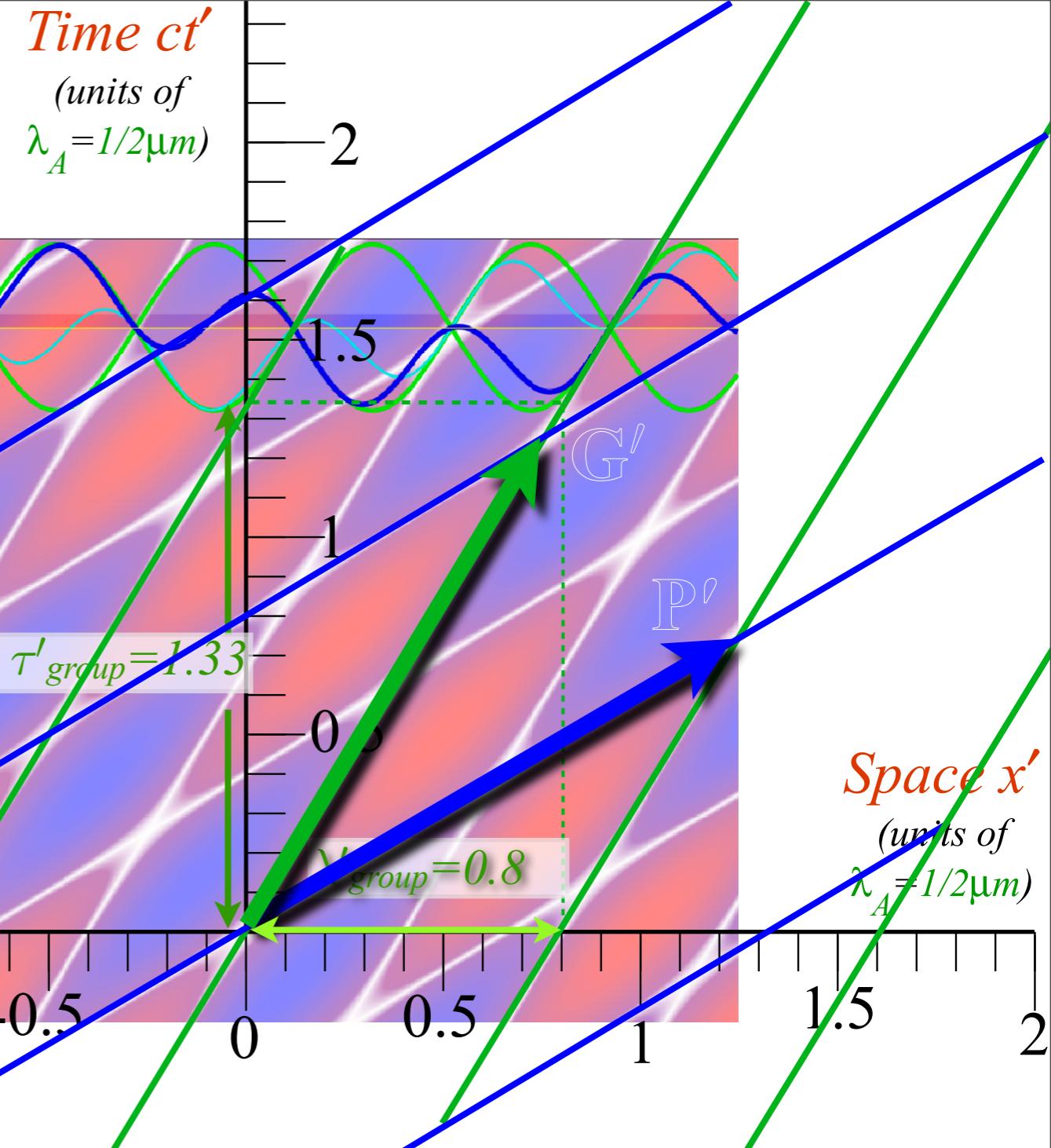
write \mathbf{G}' and \mathbf{P}' in terms of \mathbf{G} and \mathbf{P} using $\cosh\rho$ and $\sinh\rho$

$$\begin{aligned}\mathbf{G}' &= \begin{pmatrix} cK'_{group} \\ v'_{group} \end{pmatrix} = v_A \begin{pmatrix} \cosh\rho \\ \sinh\rho \end{pmatrix} = v_A \begin{pmatrix} 5/4 \\ 3/4 \end{pmatrix} \\ &= v_A \begin{pmatrix} 1 \\ 0 \end{pmatrix} \cosh\rho + v_A \begin{pmatrix} 0 \\ 1 \end{pmatrix} \sinh\rho \\ \mathbf{G}' &= \mathbf{G} \cosh\rho + \mathbf{P} \sinh\rho\end{aligned}$$

$$\begin{aligned}\mathbf{P}' &= \begin{pmatrix} cK'_{phase} \\ v'_{phase} \end{pmatrix} = v_A \begin{pmatrix} \sinh\rho \\ \cosh\rho \end{pmatrix} = v_A \begin{pmatrix} 3/4 \\ 5/4 \end{pmatrix} \\ &= v_A \begin{pmatrix} 1 \\ 0 \end{pmatrix} \sinh\rho + v_A \begin{pmatrix} 0 \\ 1 \end{pmatrix} \cosh\rho \\ \mathbf{P}' &= \mathbf{G} \sinh\rho + \mathbf{P} \cosh\rho\end{aligned}$$



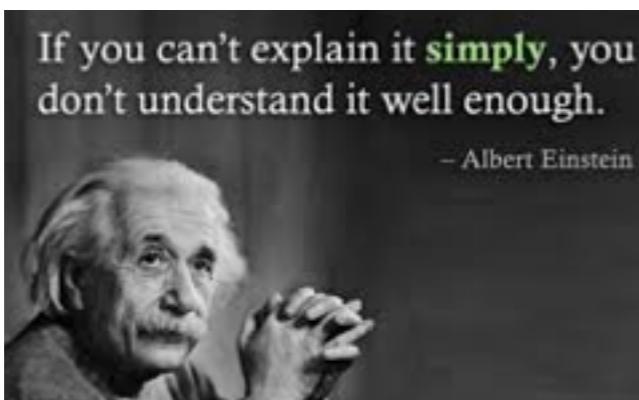
$$\begin{pmatrix} \cosh\rho & \sinh\rho \\ \sinh\rho & \cosh\rho \end{pmatrix} \text{ Lorentz transform matrix}$$



phase	$b_{RED}^{Doppler}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
group	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
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Two Famous-Name Coefficients

Albert Einstein
1859-1955

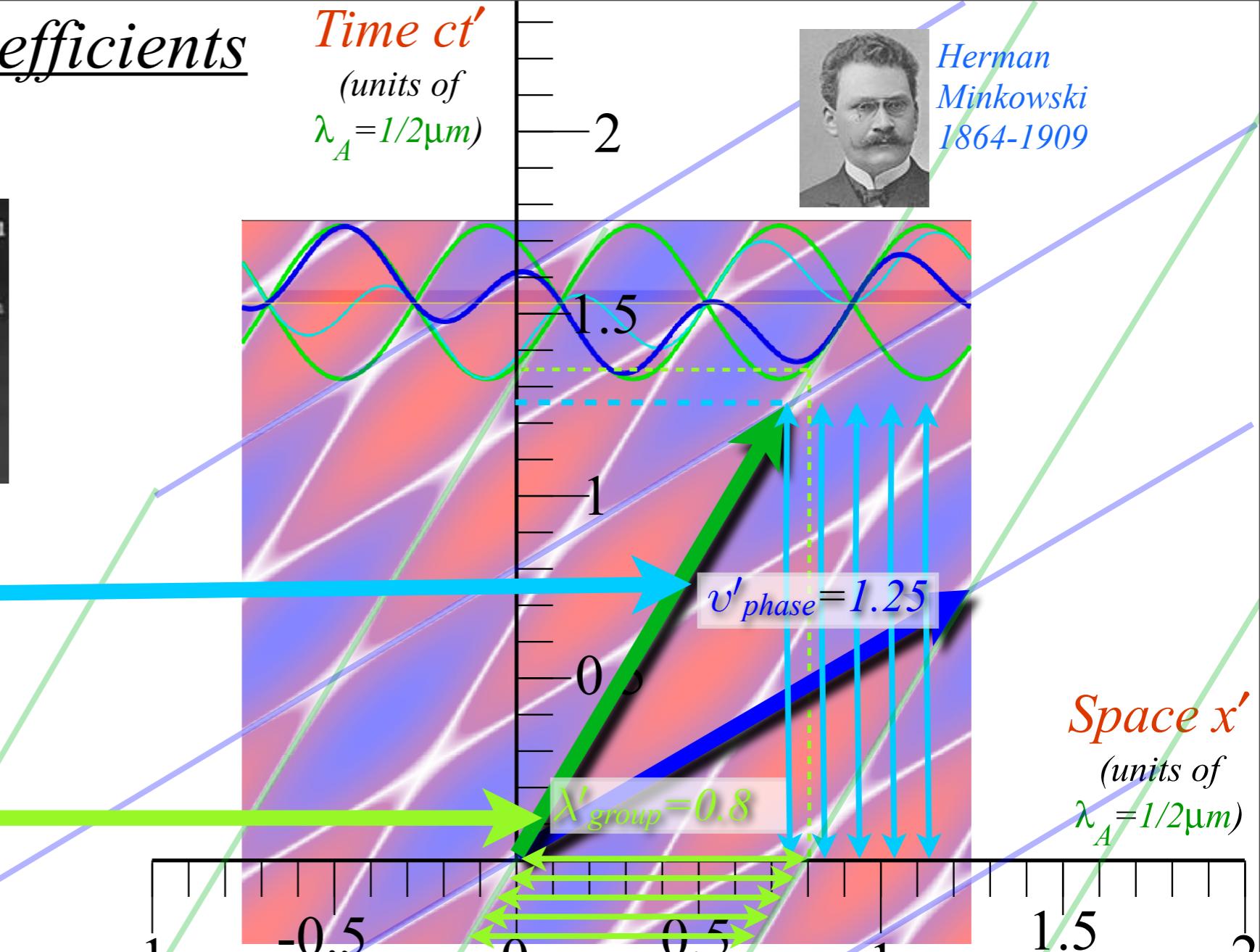


This number
is called an: **Einstein time-dilation**
(dilated by 25% here)

This number
is called a: **Lorentz length-contraction**
(contracted by 20% here)



Hendrik A.
Lorentz
1853-1928



phase	$b_{RED}^{Doppler}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
group	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

Old-Fashioned Notation
→ RelaWavity Web Simulation
Relativistic Terms (Dual plot w/expanded table)

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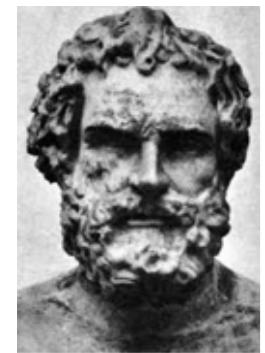
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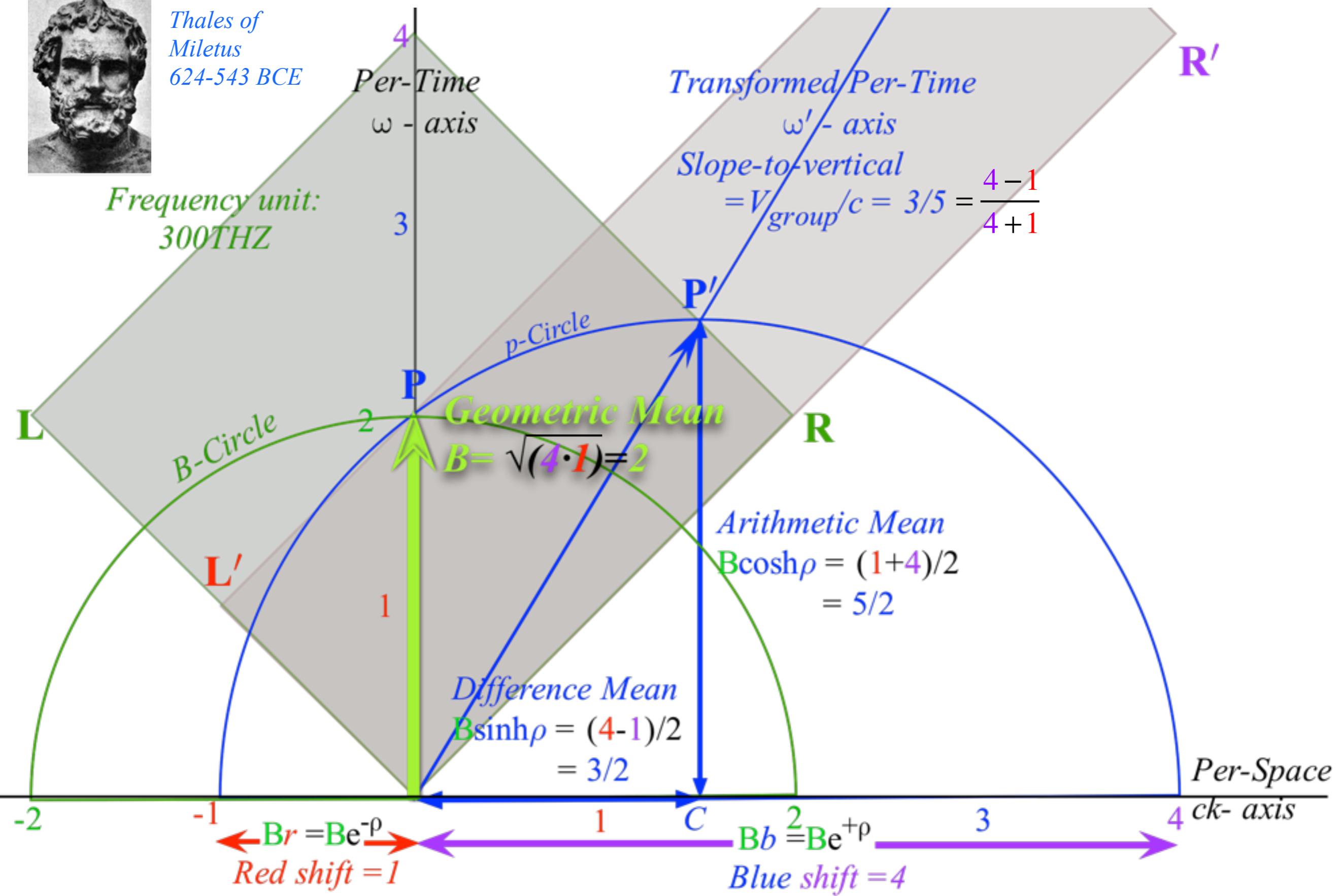
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Thales Mean Geometry (600BCE)

helps “Relativity”

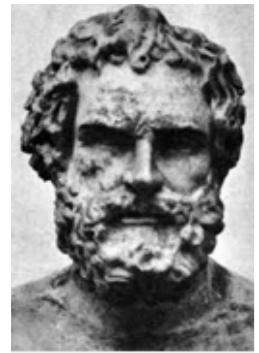


*Thales of
Miletus
624-543 BCE*

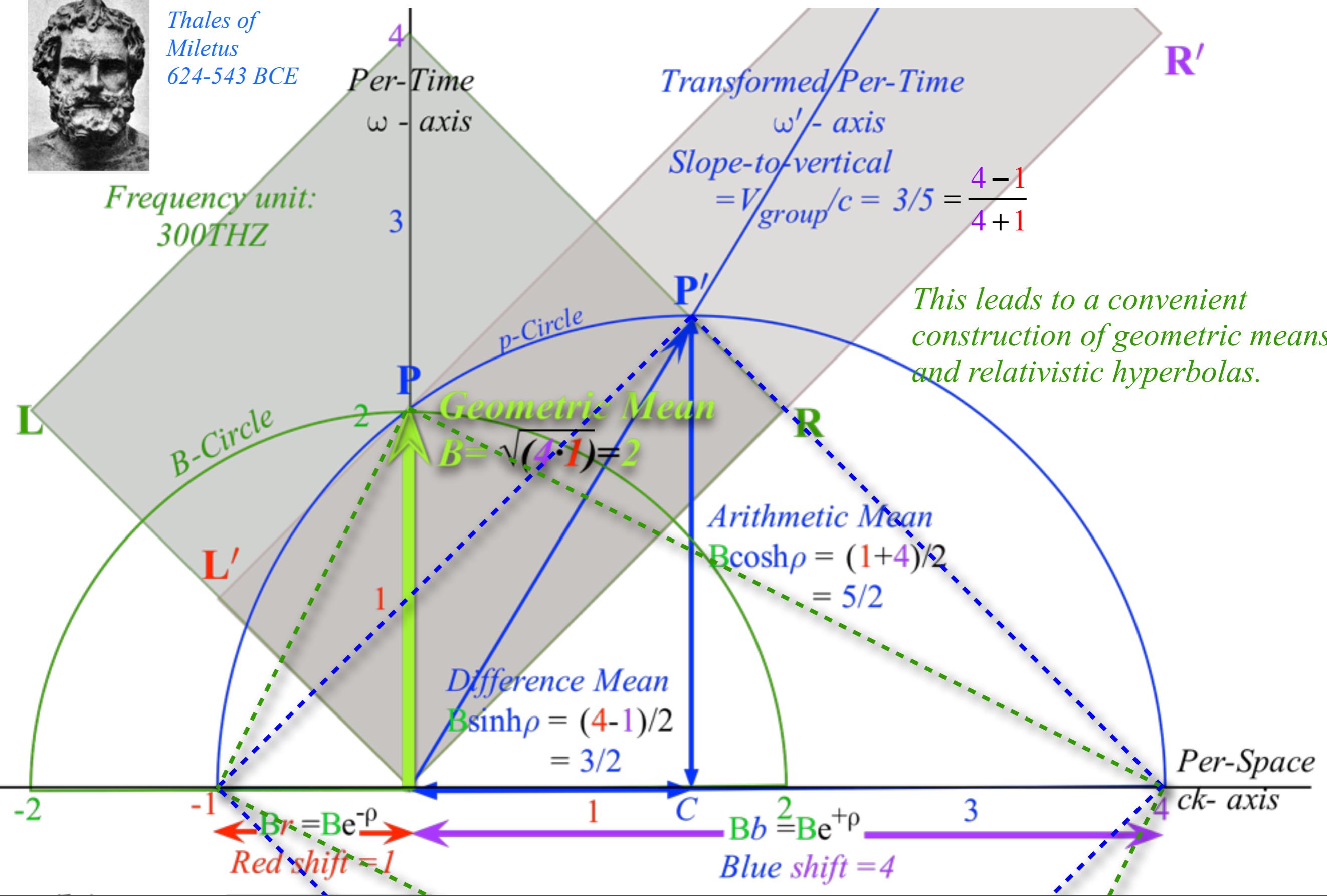


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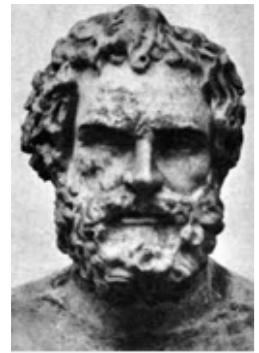


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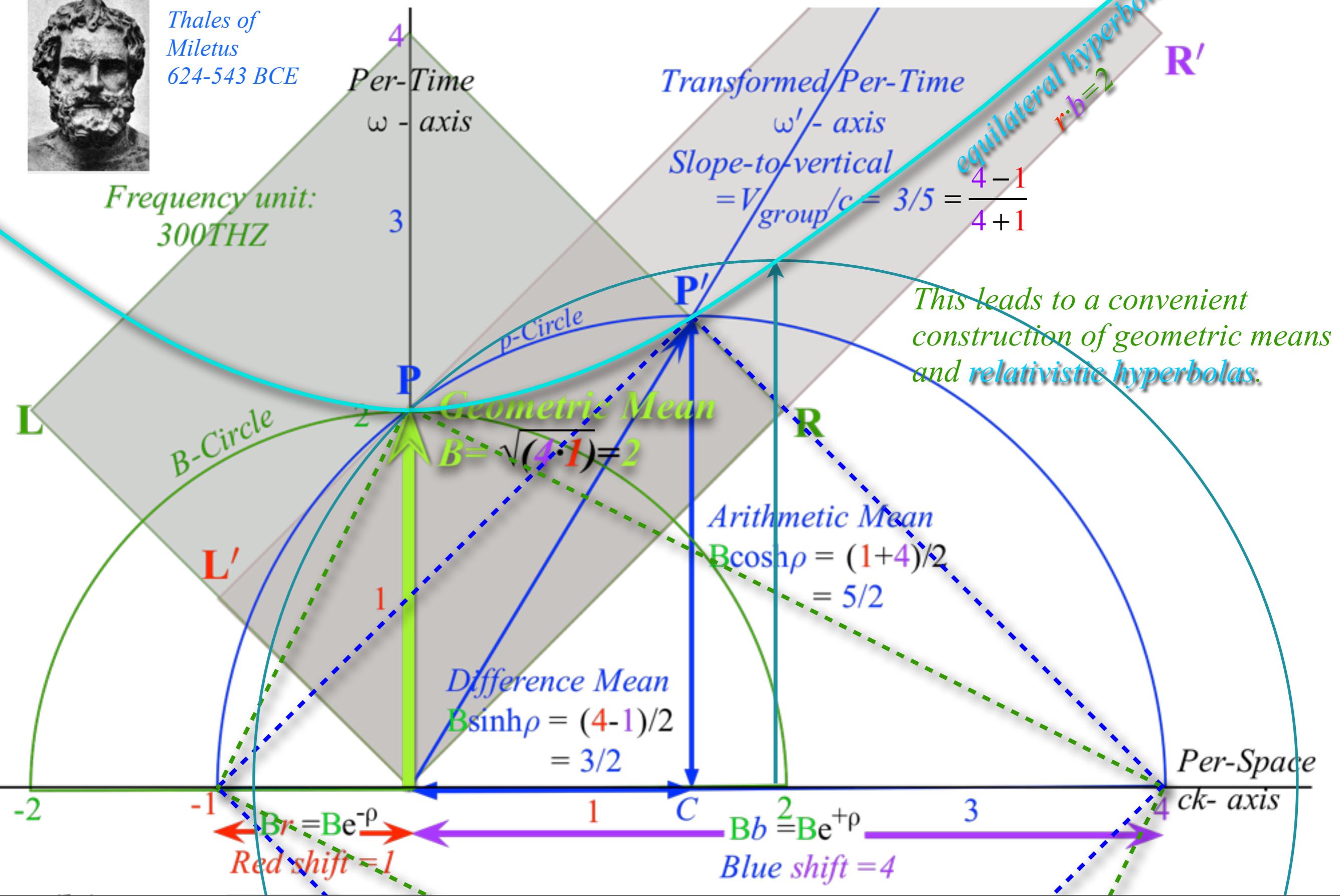


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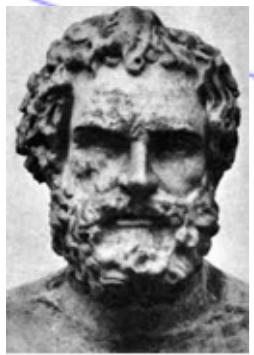


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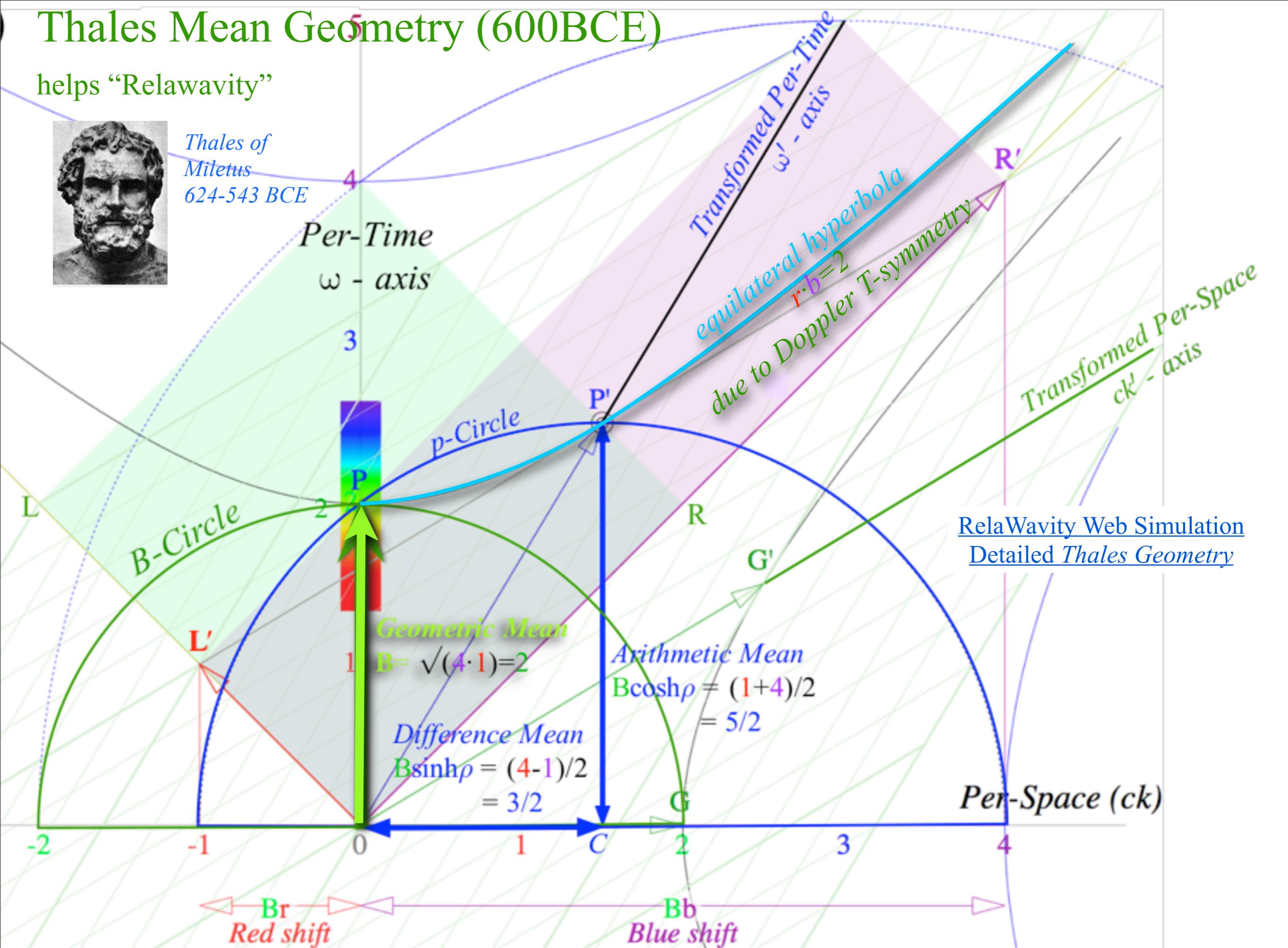


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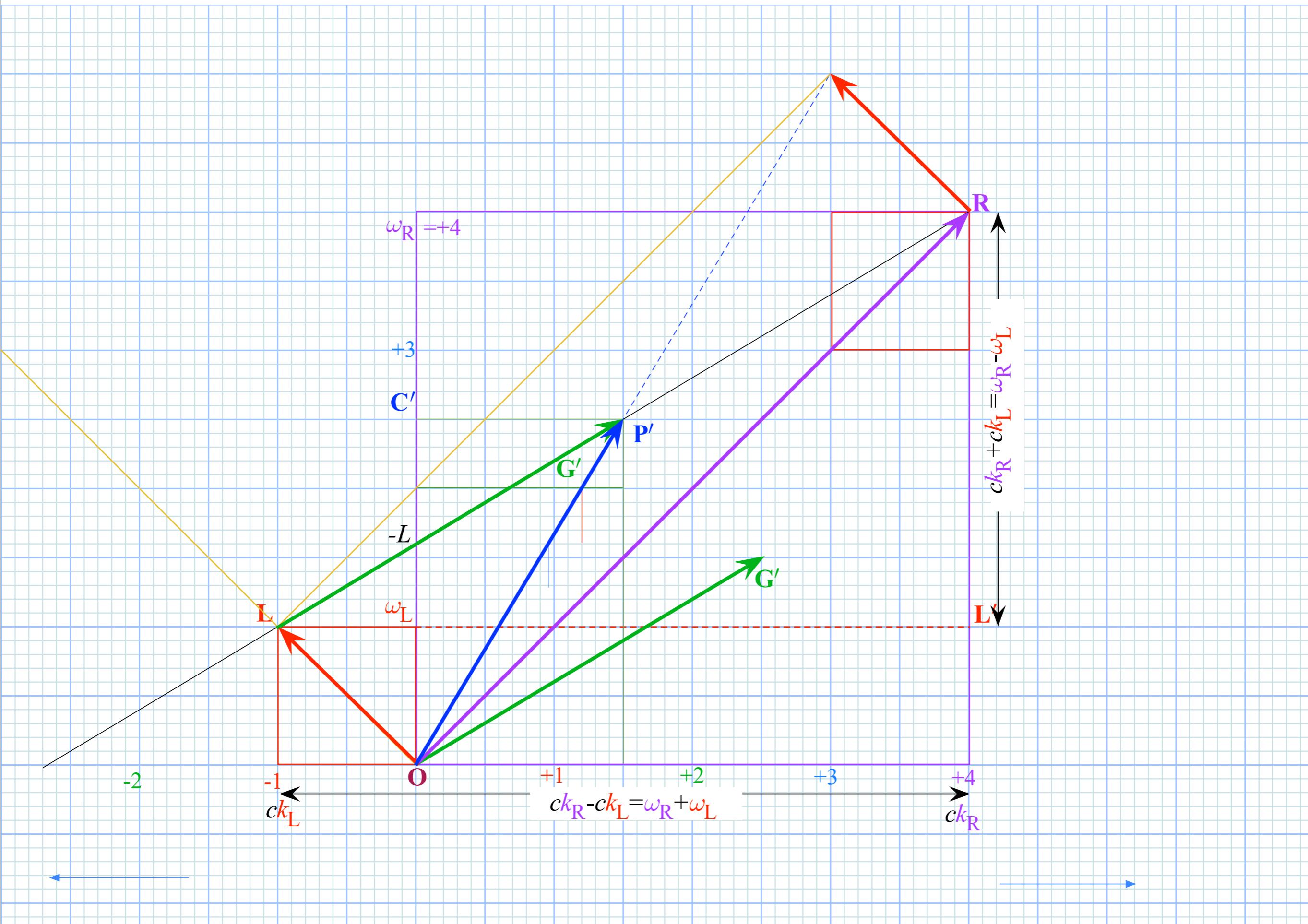
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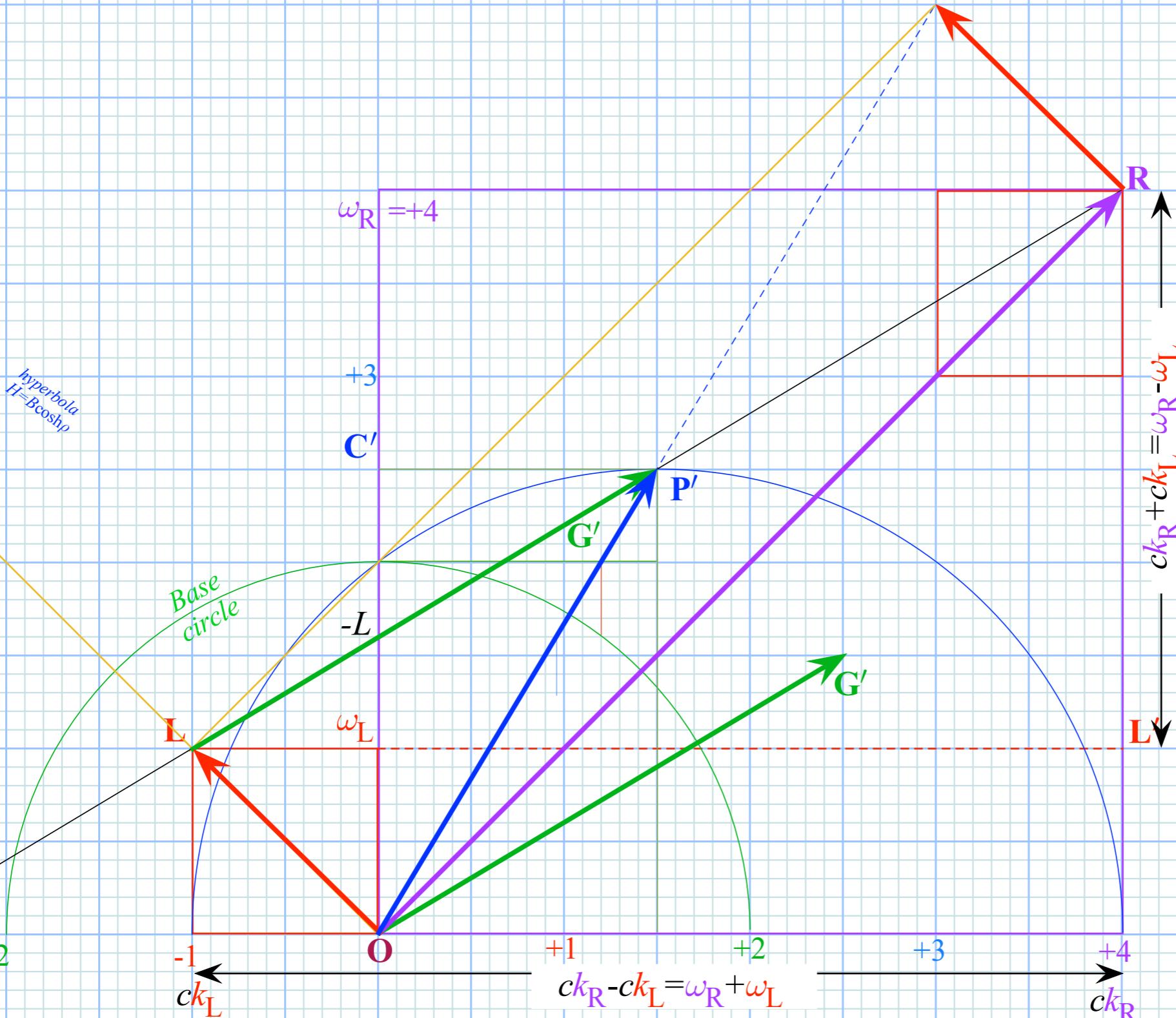
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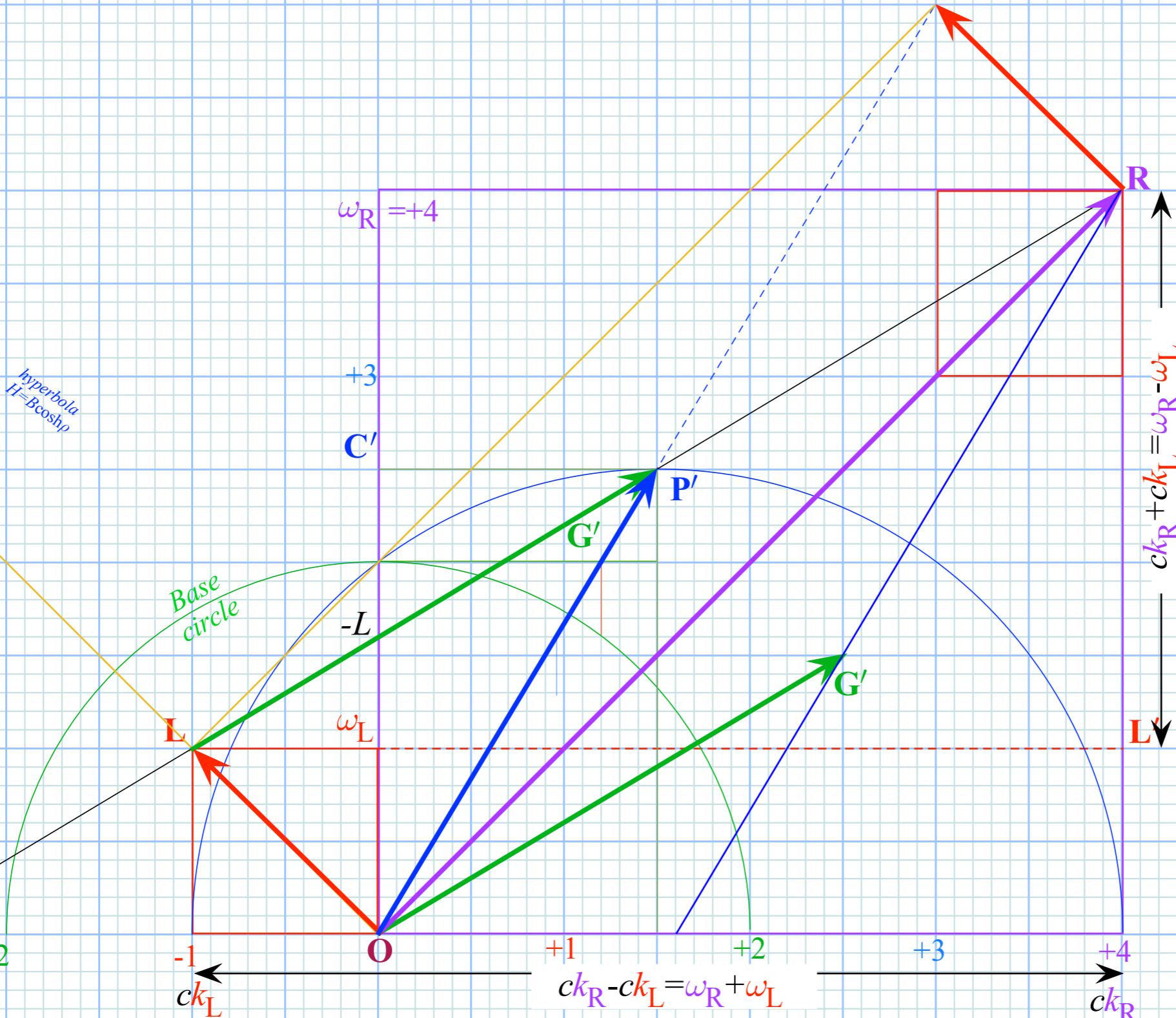
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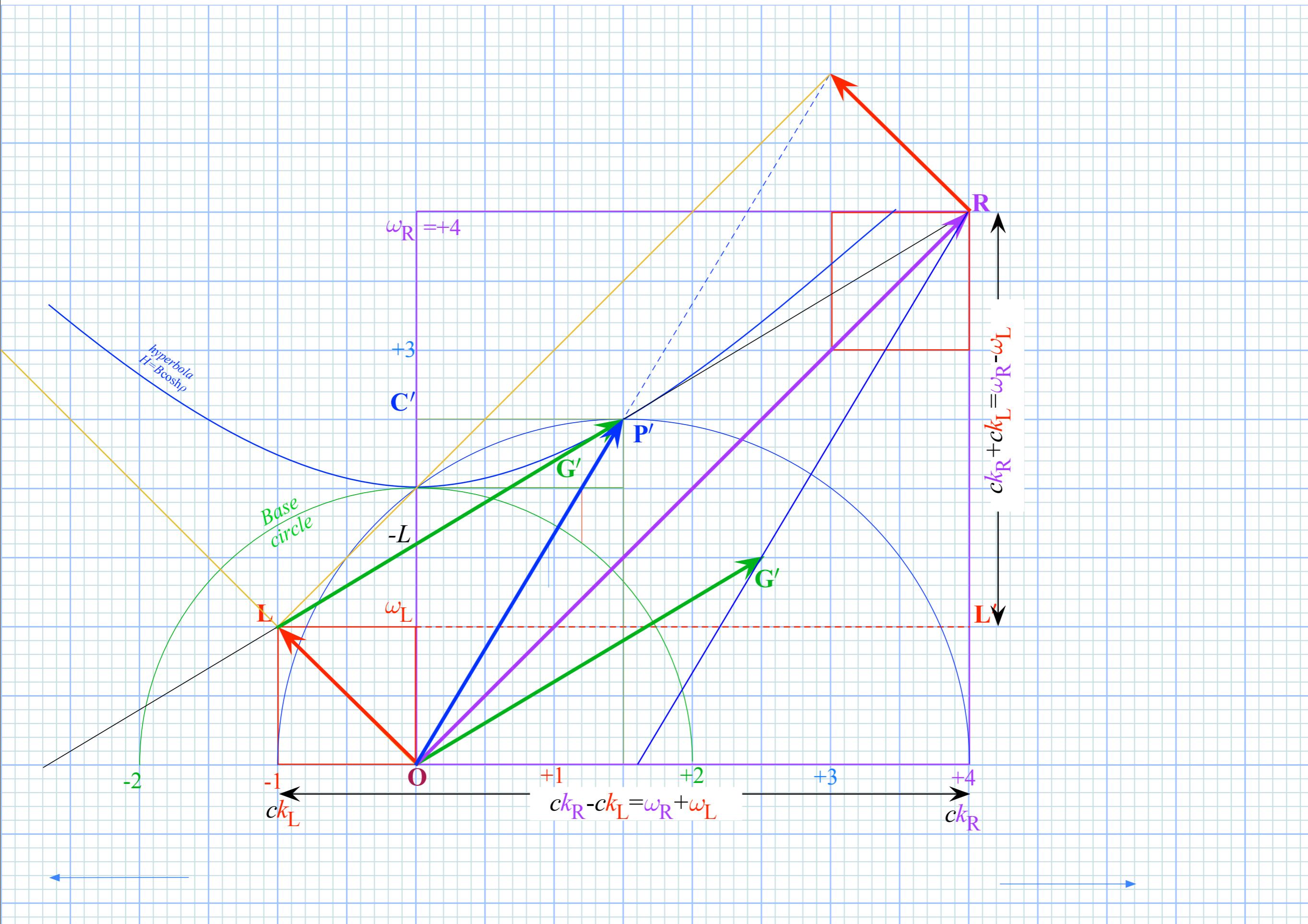
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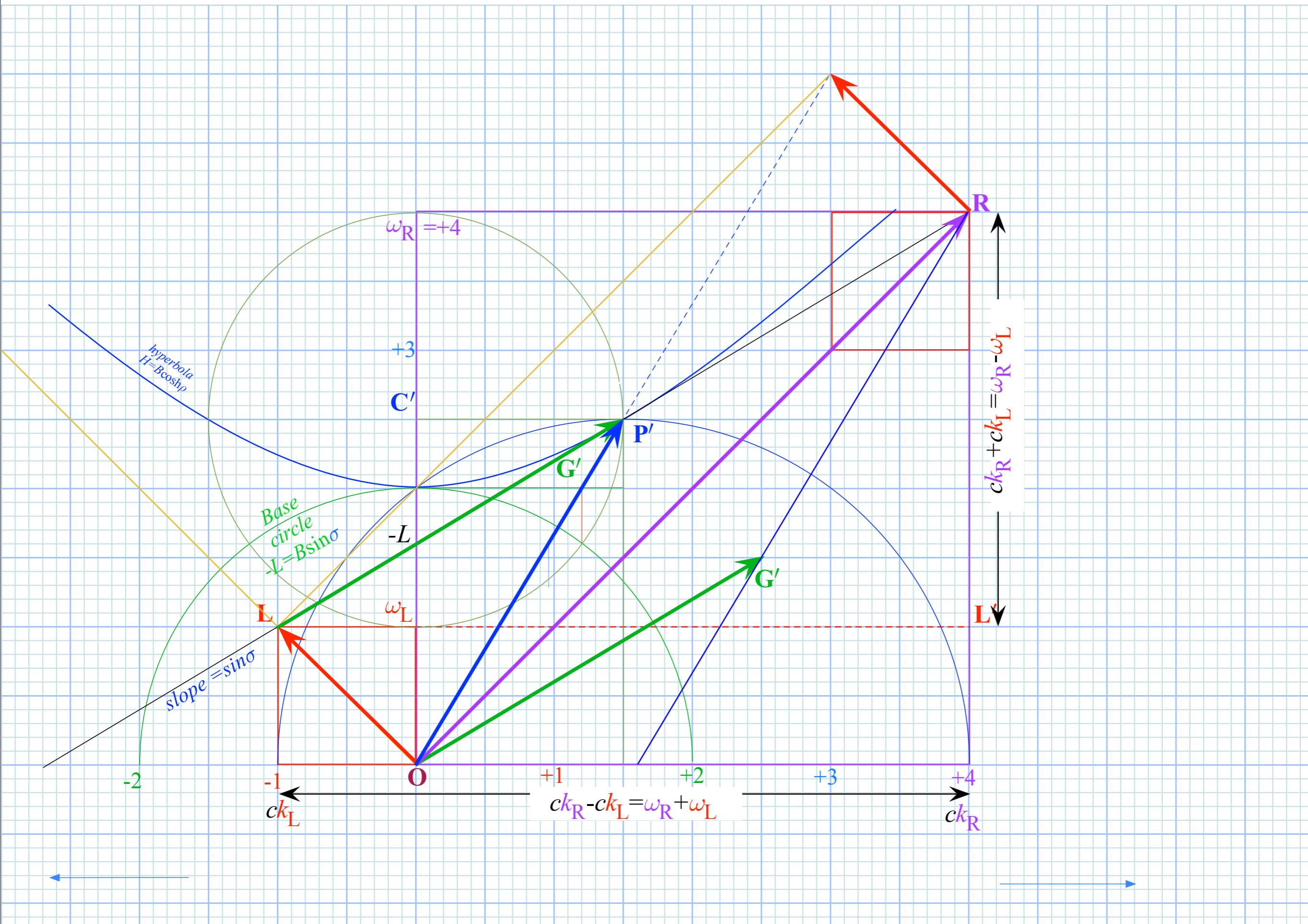
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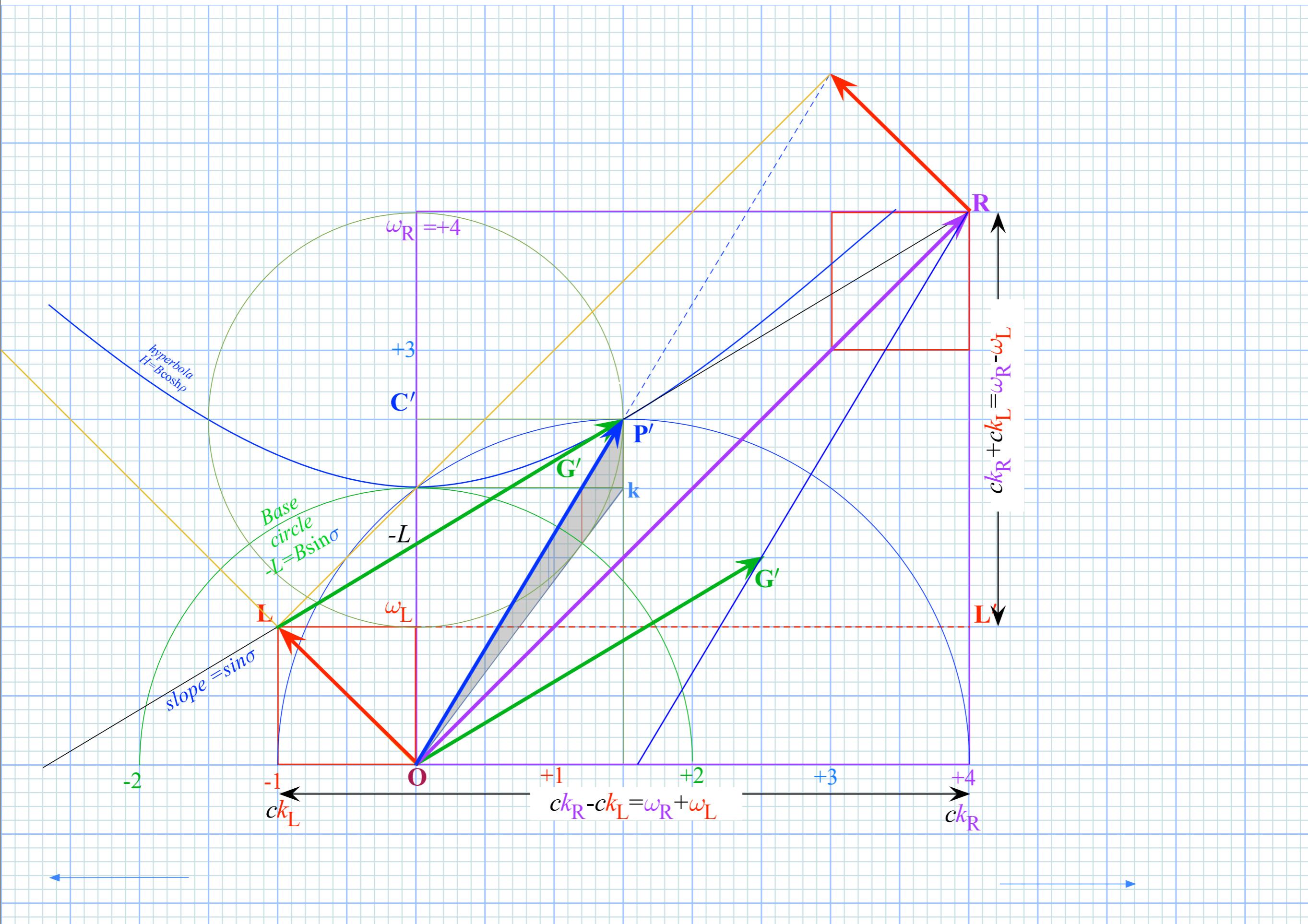


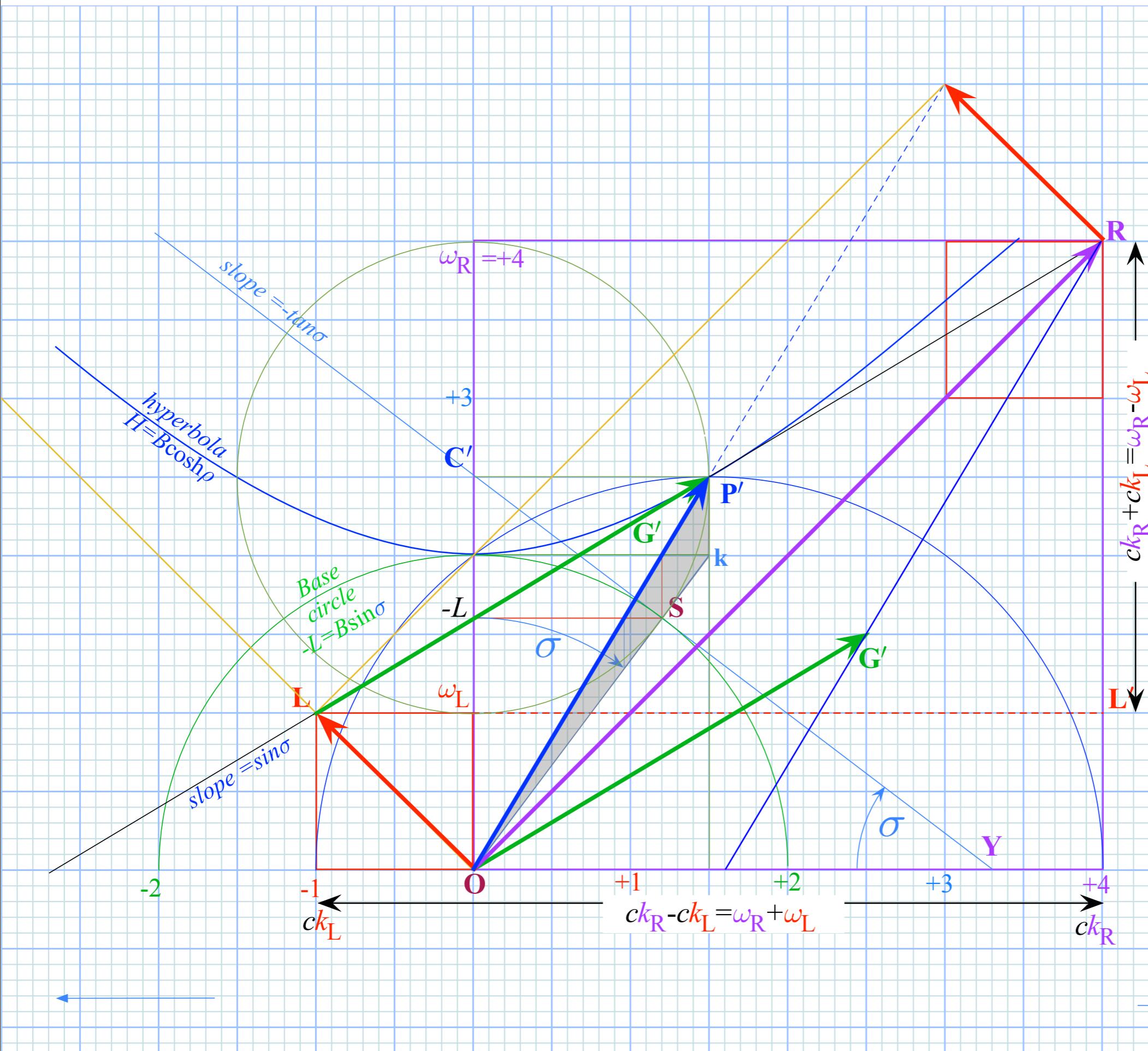




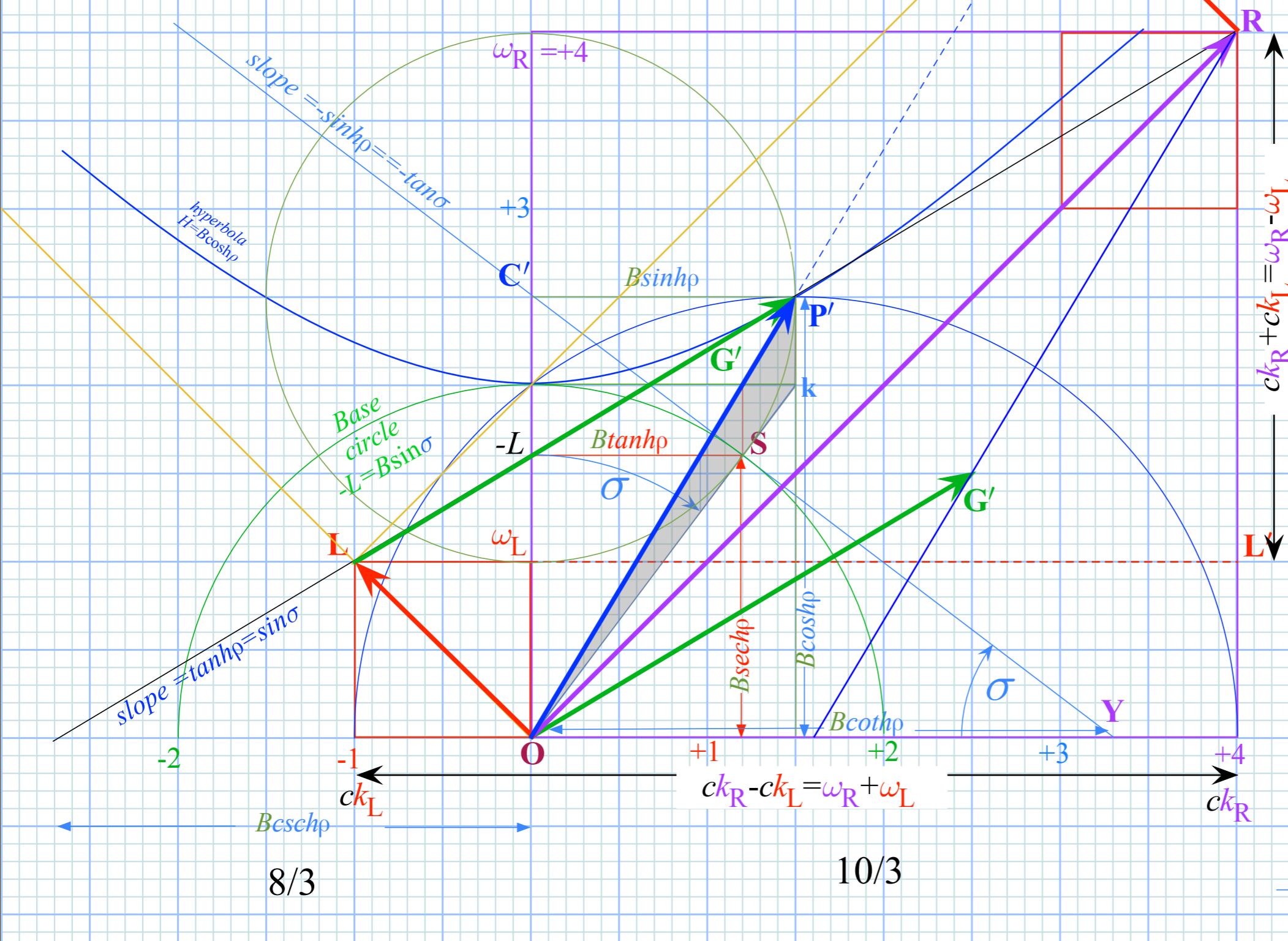








$$\rho = \log_e 2 = \text{Arctanh}(\frac{3}{5}) = \tanh^{-1}(\frac{3}{5}) = 0.6931$$



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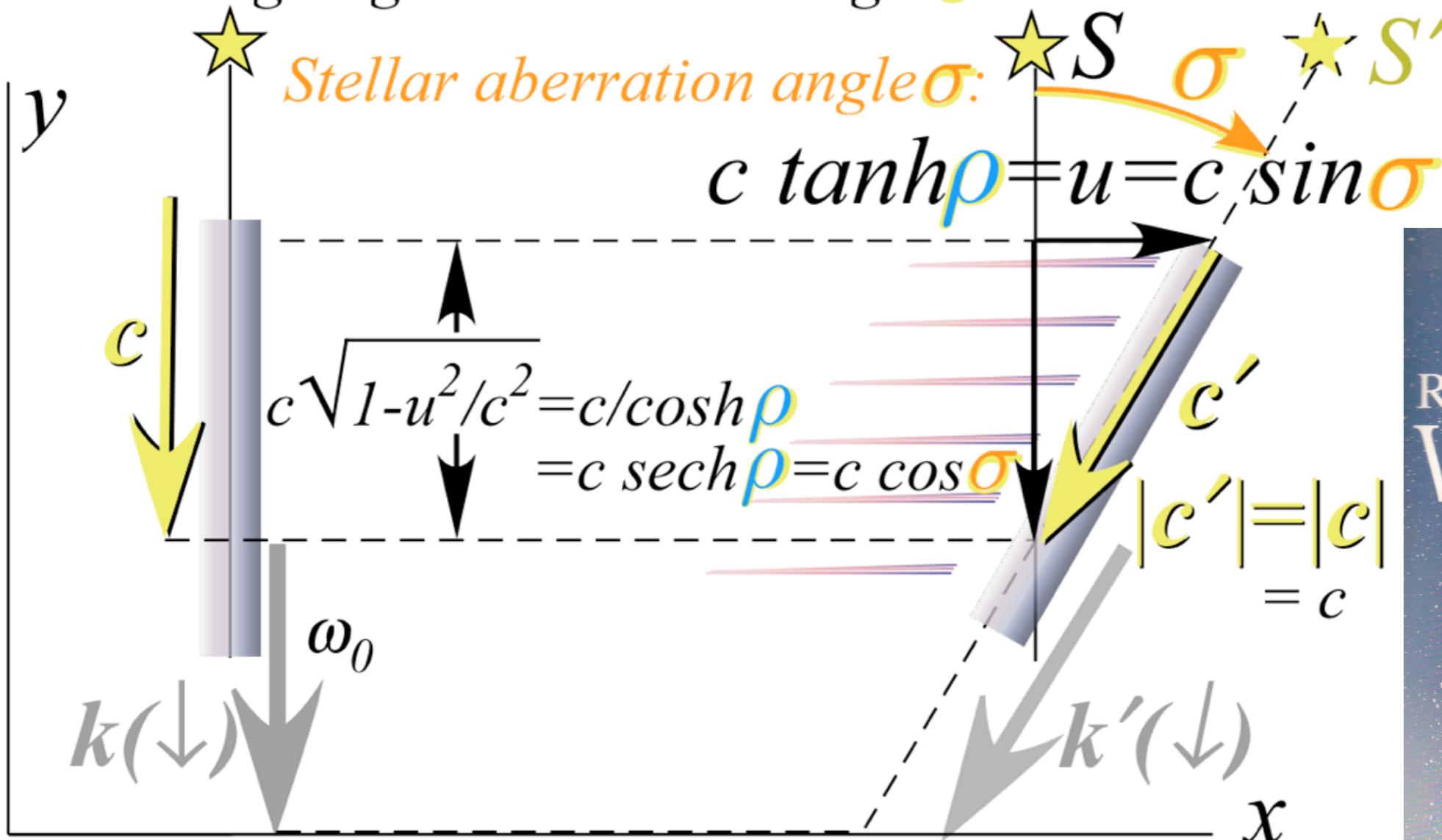


Comparing Longitudinal relativity parameter: Rapidity $\rho = \log_e(\text{Doppler Shift})$ to a Transverse* relativity parameter: Stellar aberration angle σ

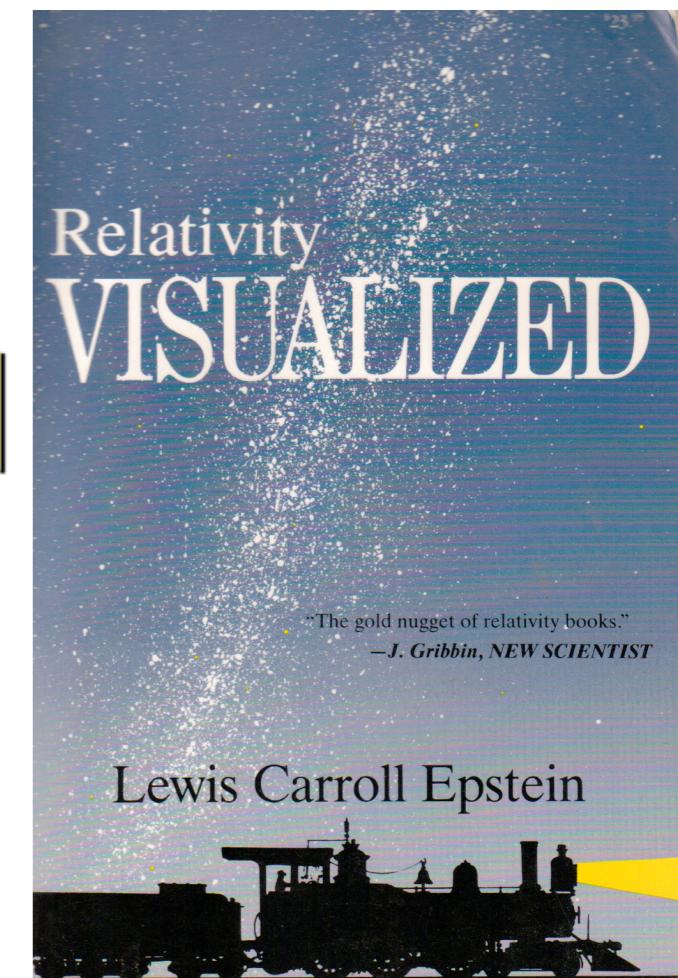
*Lewis Carroll Epstein, *Relativitätstheorie*, Birkhäuser, (2004) Earlier English version (1985)

Observer fixed below star sees it directly overhead.

Observer going u sees star at angle σ in u direction.

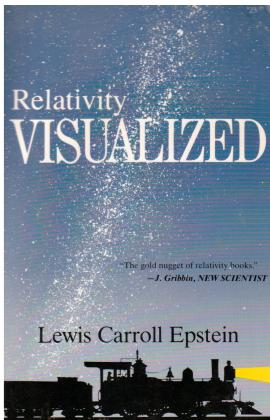


We used notion σ for stellar-ab-angle, (a “flipped-out” ρ). Epstein not interested in ρ analysis or in relation of σ and ρ .



Comparing Longitudinal relativity parameter: Rapidity $\rho = \log_e(\text{Doppler Shift})$ to a Transverse* relativity parameter: Stellar aberration angle σ

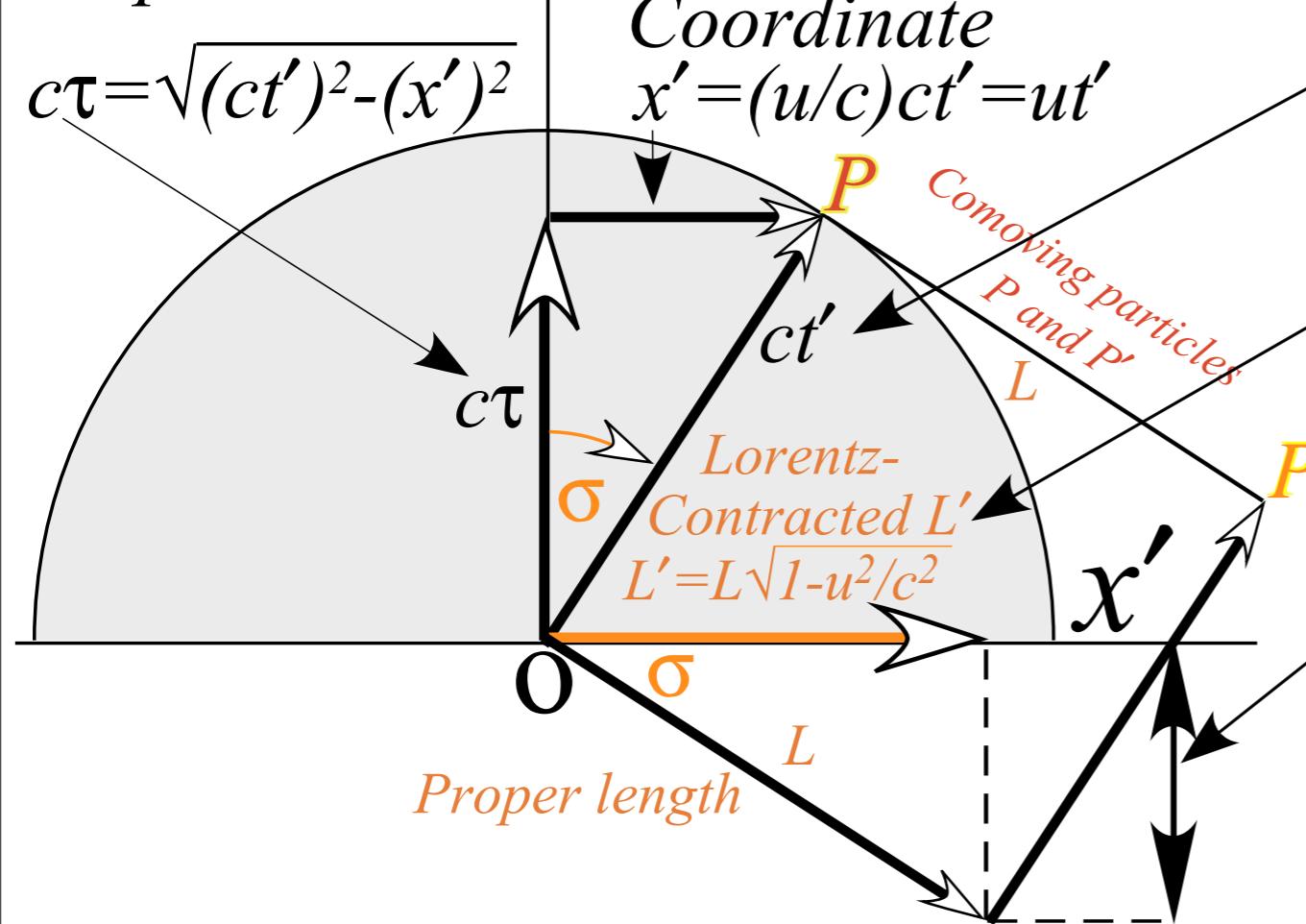
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Proper time $c\tau$ vs. coordinate space x - (L. C. Epstein's "Cosmic Speedometer")

Particles P and P' have speed u in (x', ct') and speed c in $(x, c\tau)$

Proper time $C\tau$



Einstein time dilation:

$$ct' = c\tau \sec \sigma = c\tau \cosh \rho = c\tau / \sqrt{1-u^2/c^2}$$

Lorentz length contraction:

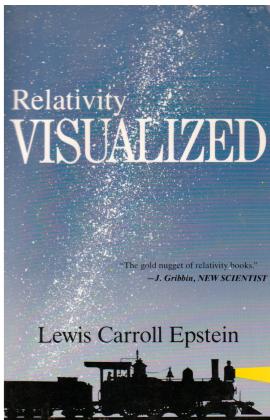
$$L' = L \operatorname{sech} \rho = L \cos \sigma = L \cdot \sqrt{1-u^2/c^2}$$

Proper Time asimultaneity:

$$\begin{aligned} c \Delta \tau &= L' \sinh \rho = L \cos \sigma \sinh \rho \\ &= L \cos \sigma \tan \sigma \\ &= L \sin \sigma = L / \sqrt{c^2/u^2 - 1} \sim L u/c \end{aligned}$$

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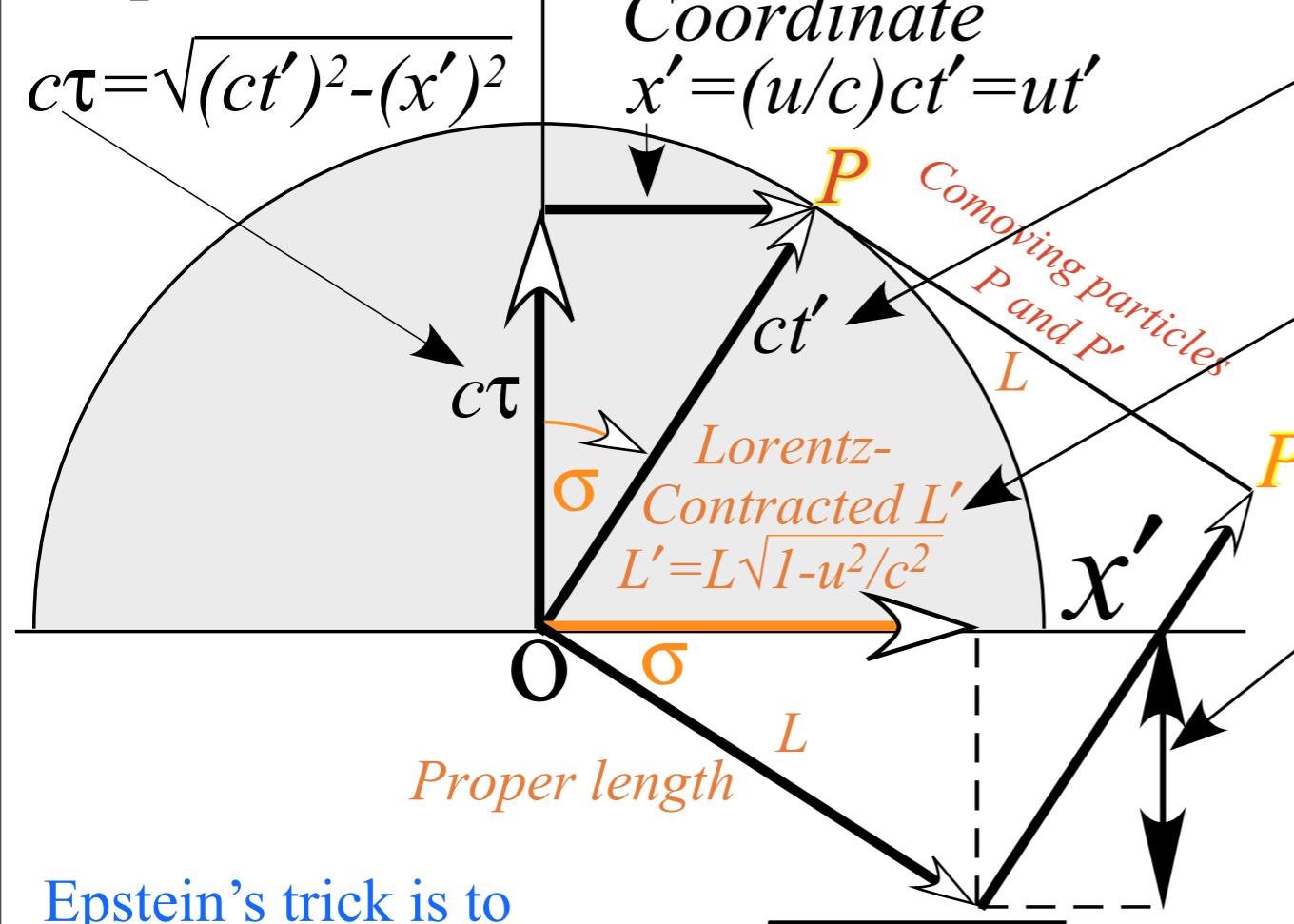
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Proper Time asimultaneity:

$$\begin{aligned} c \Delta \tau &= L' \sinh \rho = L \cos \sigma \sinh \rho \\ &= L \cos \sigma \tan \sigma \\ &= L \sin \sigma = L / \sqrt{c^2/u^2 - 1} \sim L u/c \end{aligned}$$

Epstein's trick is to

turn a hyperbolic form $c\tau = \sqrt{(ct')^2 - (x')^2}$

into a circular form:

$$\sqrt{(c\tau)^2 + (x')^2} = (ct')$$

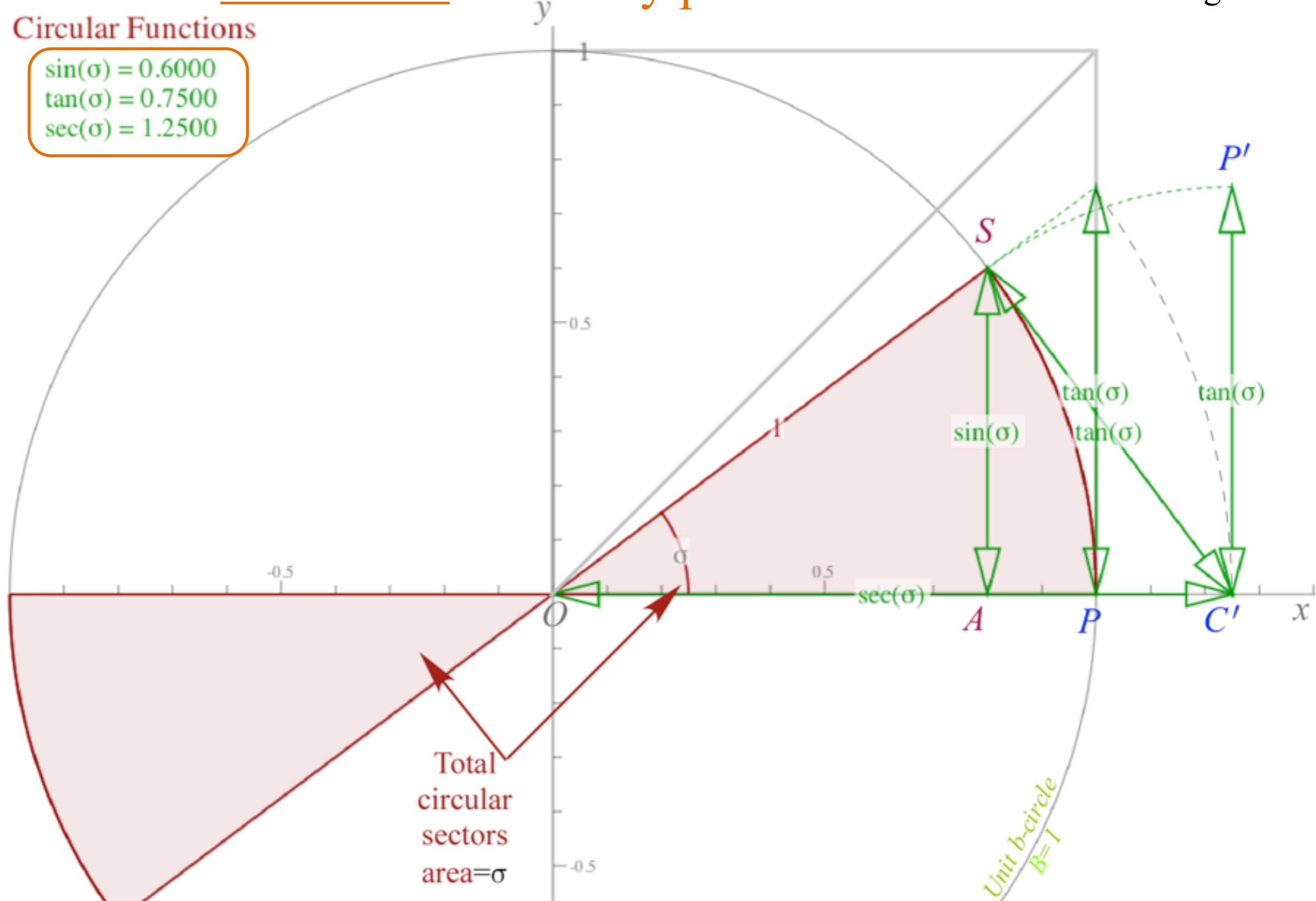
Then everything (and everybody) always goes speed c through $(x', c\tau)$ space!

Comparing Longitudinal relativity parameter: Rapidity $\rho = \log_e(\text{Doppler Shift})$ to a Transverse relativity parameter: Stellar aberration angle σ

(a)

Circular Functions

$$\begin{aligned}\sin(\sigma) &= 0.6000 \\ \tan(\sigma) &= 0.7500 \\ \sec(\sigma) &= 1.2500\end{aligned}$$



RelaWavity Web Simulation
Geometry of Stellar Aberration Angle

Comparing Longitudinal relativity parameter: Rapidity $\rho = \log_e(\text{Doppler Shift})$ to a Transverse relativity parameter: Stellar aberration angle σ

(b)

Circular Functions

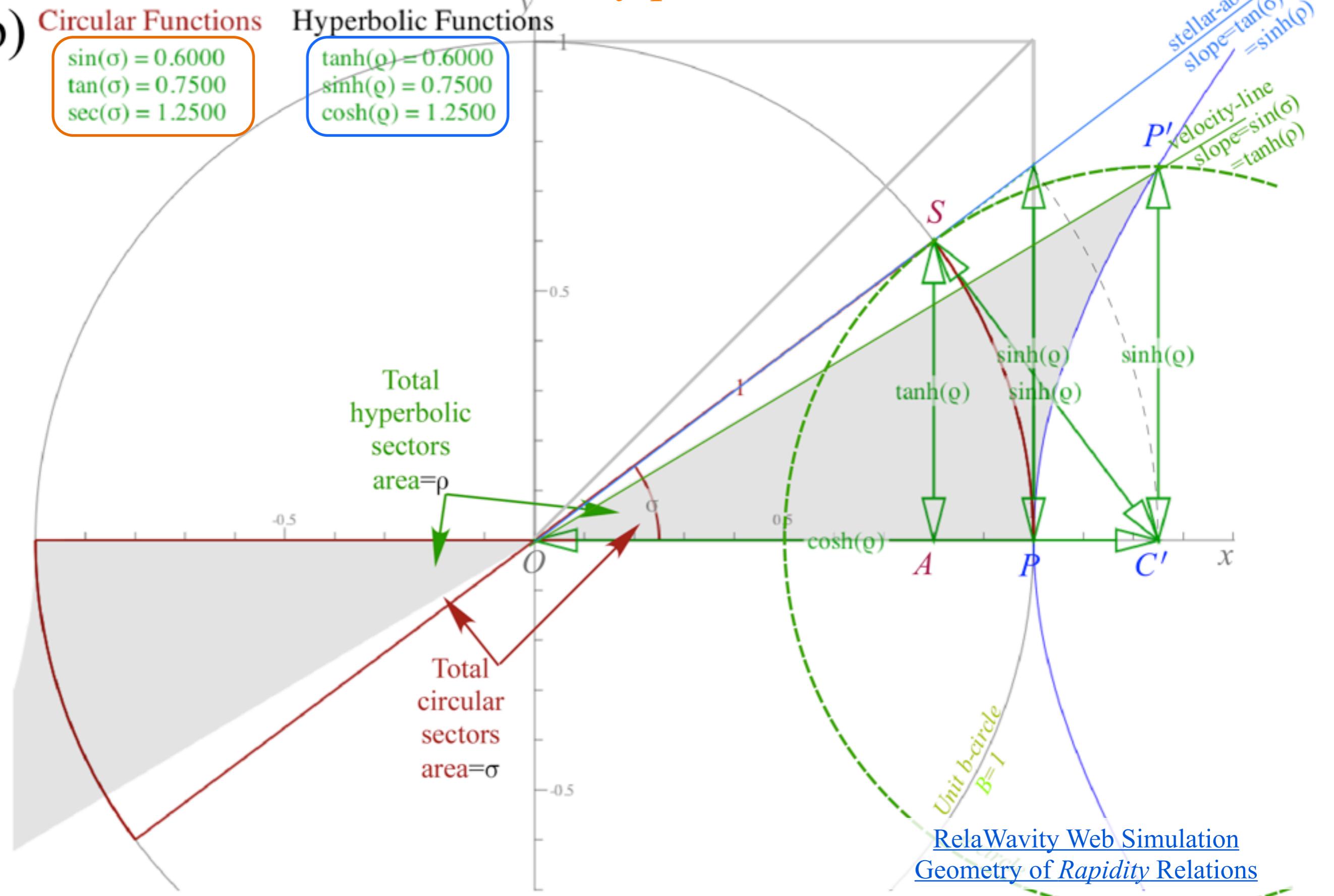
$$\begin{aligned}\sin(\sigma) &= 0.6000 \\ \tan(\sigma) &= 0.7500 \\ \sec(\sigma) &= 1.2500\end{aligned}$$

Hyperbolic Functions

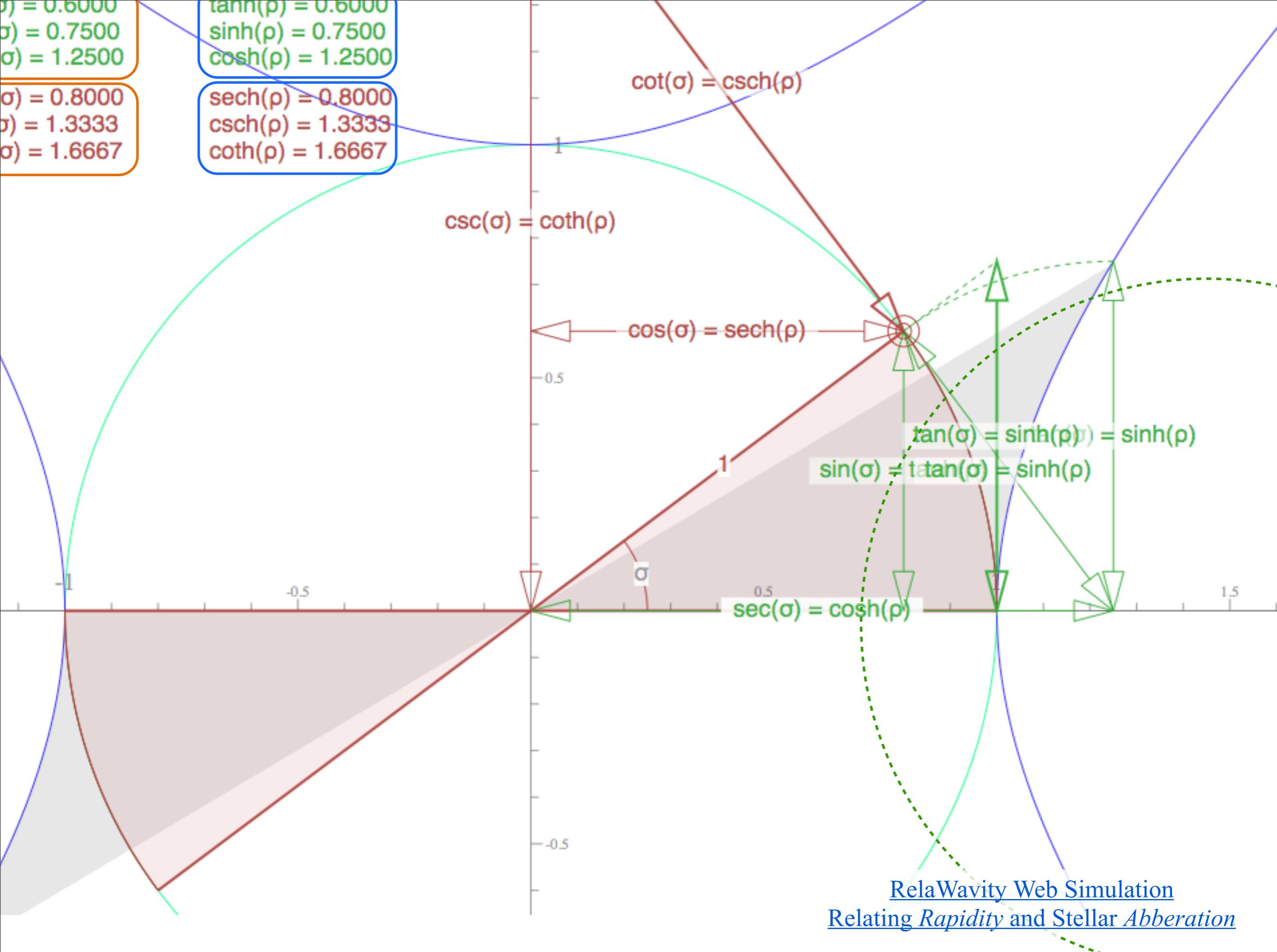
$$\begin{aligned}\tanh(\rho) &= 0.6000 \\ \sinh(\rho) &= 0.7500 \\ \cosh(\rho) &= 1.2500\end{aligned}$$

Total hyperbolic sectors area = ρ

Total circular sectors area = σ

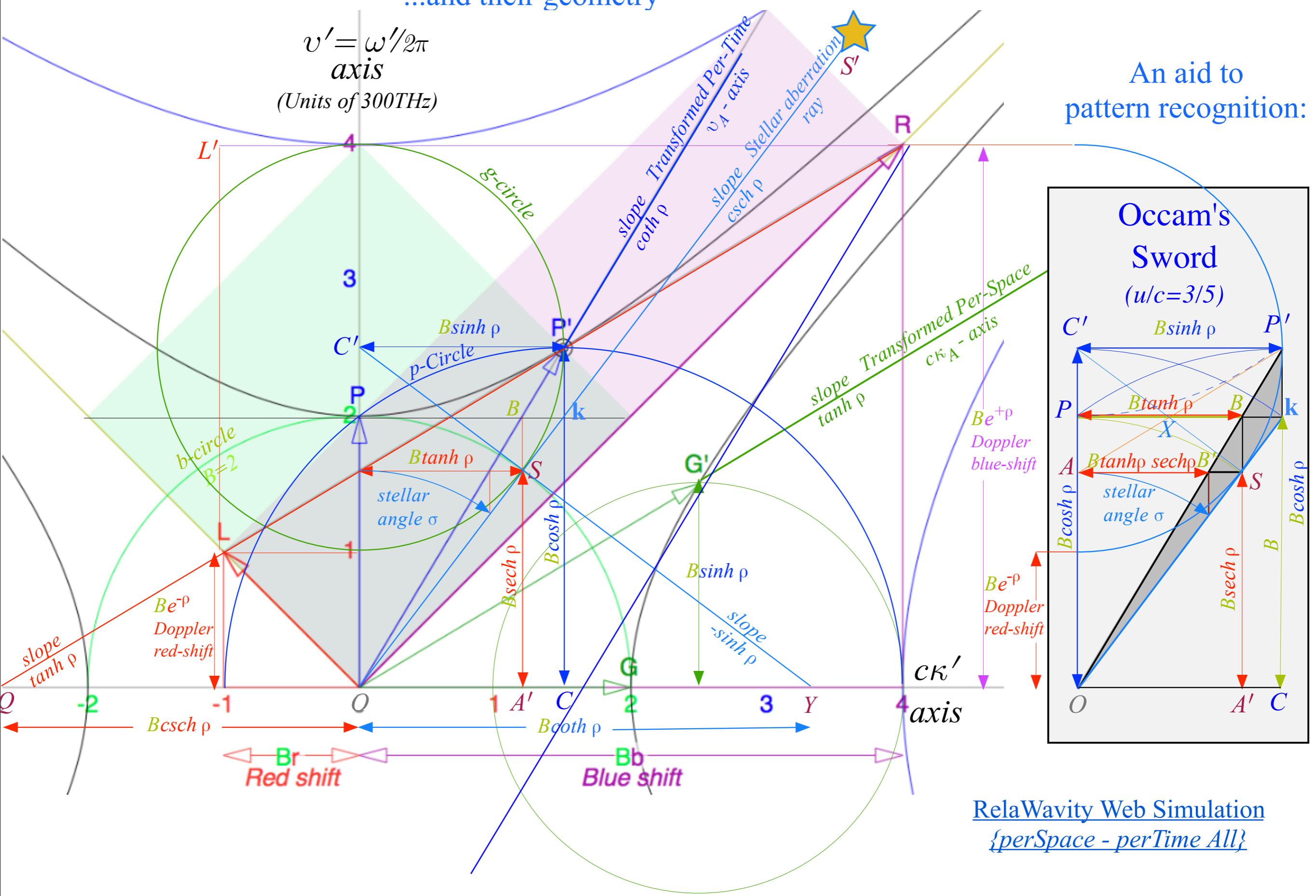


[RelaWavity Web Simulation](#)
[Geometry of Rapidity Relations](#)

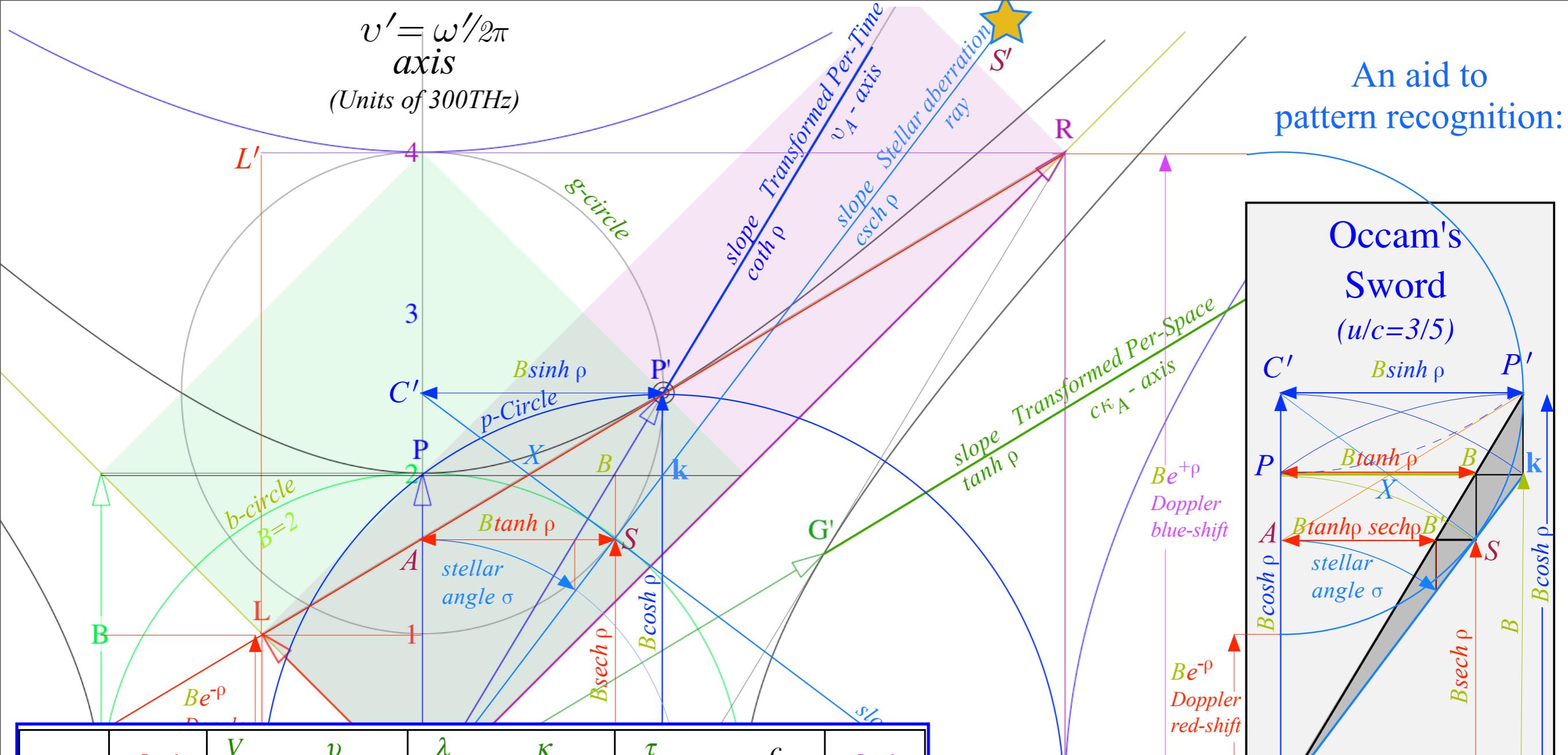


Summary of optical wave parameters for relativity and QM

...and their geometry



An aid to
pattern recognition:



group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

Table of 12 wave parameters
(includes inverses) for relativity
...and values for $u/c=3/5$

[RelaWavity Web Simulation](#)
[Relativistic Terms \(Dual plot w/expanded table\)](#)

