

Lecture 29

Tue. 12.14.2017

Relativity : a novel introduction to relativistic mechanics I.

(Unit 8 12.14.07)

Why Men in Black shot little Suzie...Learning about sin!, cos and...Trigonometric road maps

Hyper-Trigonometric *Relativity* geometry and Euler exponential algebra

1CW wavefunctions and phasors

Per-space-per-time vs Space-time

Wave velocity formulas

Introducing Doppler shifting

Why is c so constant?!

Introducing Doppler Arithmetic and *Rapidity* ρ

Optical interference “baseball-diamond” displays *phase* and *group* velocity

Details of 2CW wavefunctions in rest frame

Pulse waves (PW) versus Continuous Waves (CW)

Doppler shifted “baseball-diamond” displays Lorentz frame transformation

Analyzing wave velocity by *per-space-per-time* and *space-time* graphs

16 coefficients of relativistic 2CW interference

Two “famous-name” coefficients and the Lorentz transformation

Thales geometry of Lorentz transformation

Rapidity ρ related to *stellar aberration angle* σ and L. C. Epstein’s approach to relativity

Longitudinal hyperbolic ρ -geometry connects to transverse circular σ -geometry

“Occams Sword” and geometry of 16 parameter functions of ρ and σ

Application to TE-Waveguide modes and synchrotron beam relativity

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For an introductory, web based development of this and other concepts in special relativity see our entrant in the 2005 Pirelli Challenge:

*A **Colorful** Road to Relativity*

Using Occam's Razors

and

Evenson's Lasers

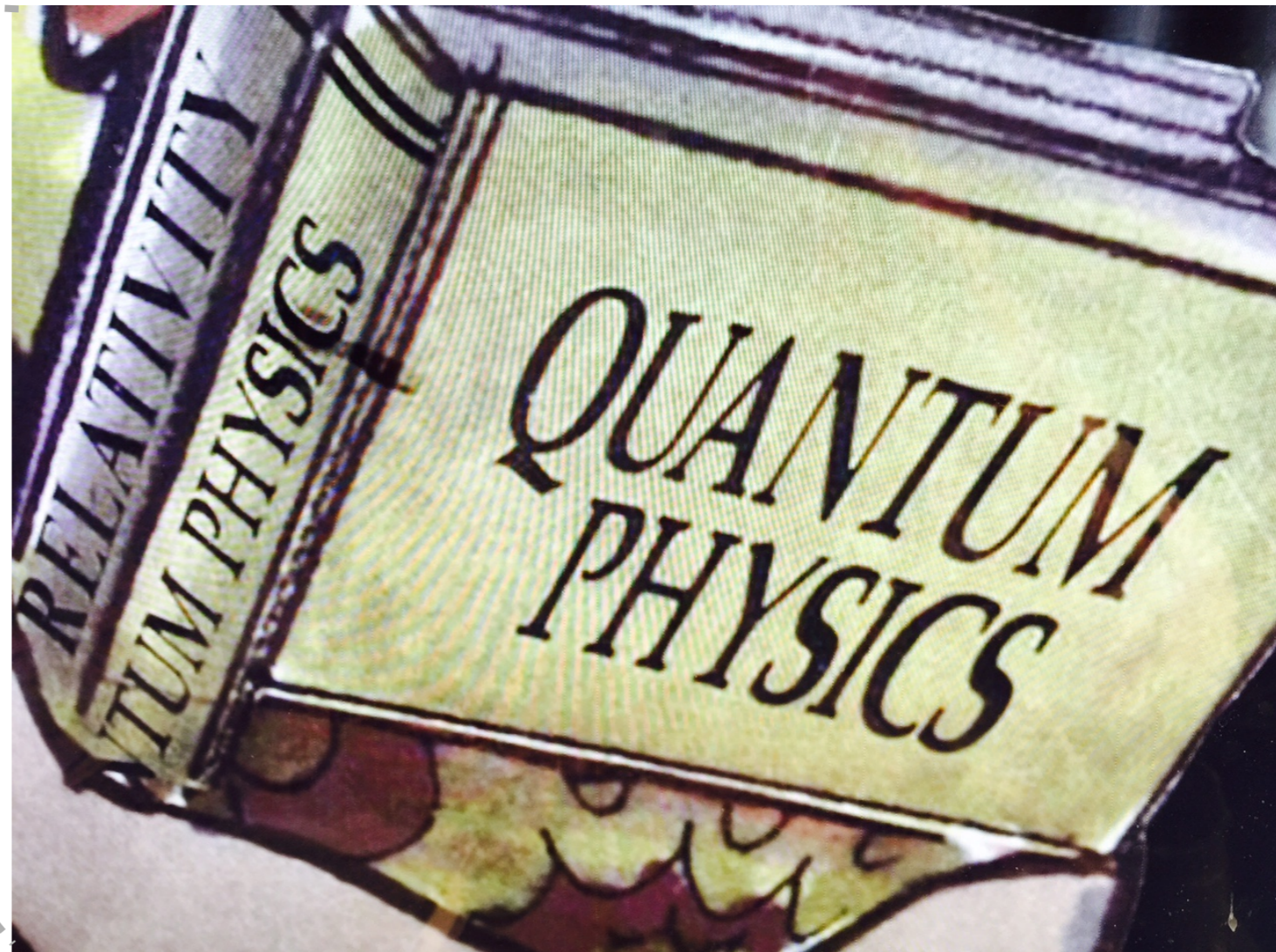


(Why a *Men In Black* candidate shot little Suzy)

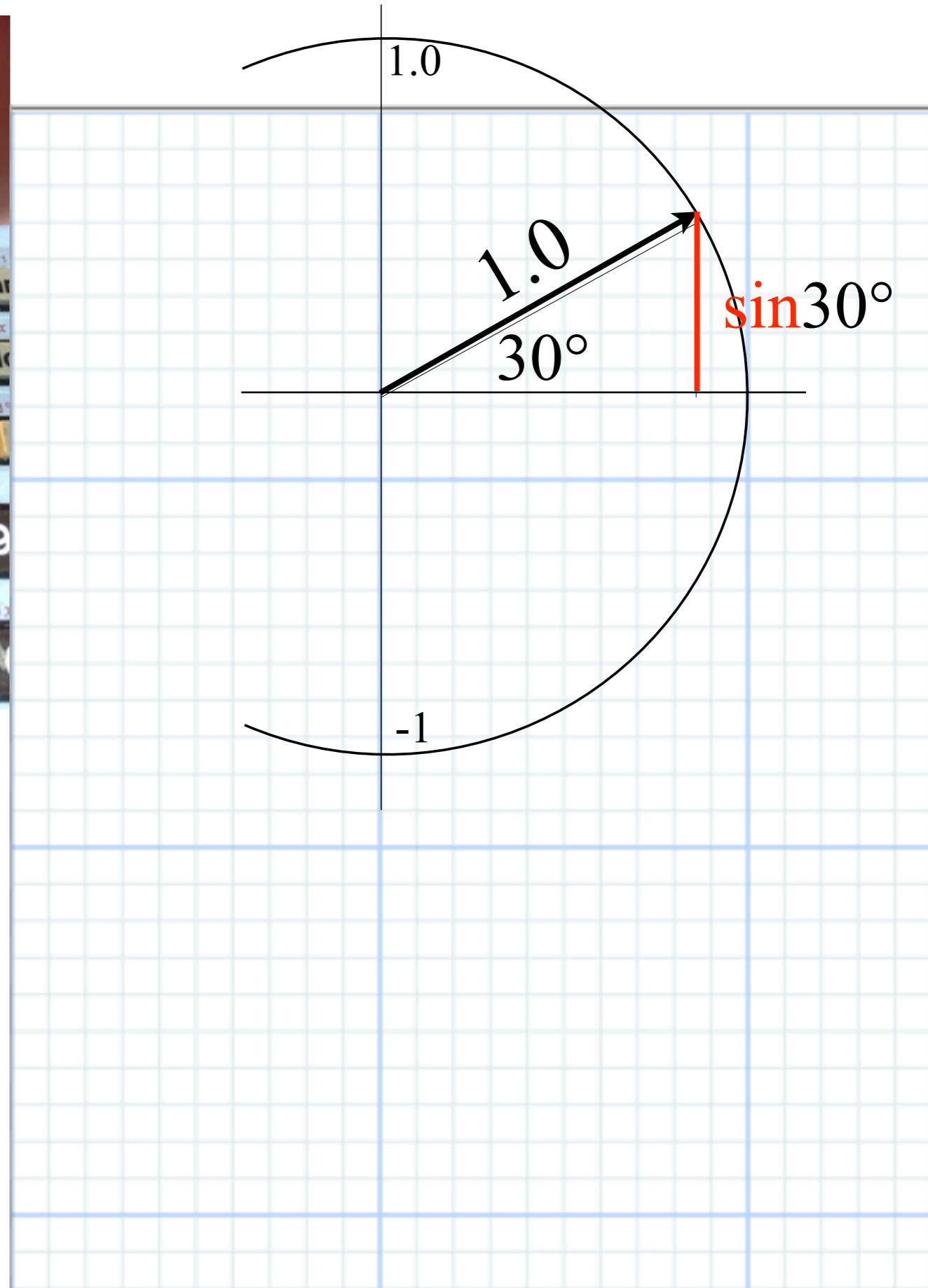
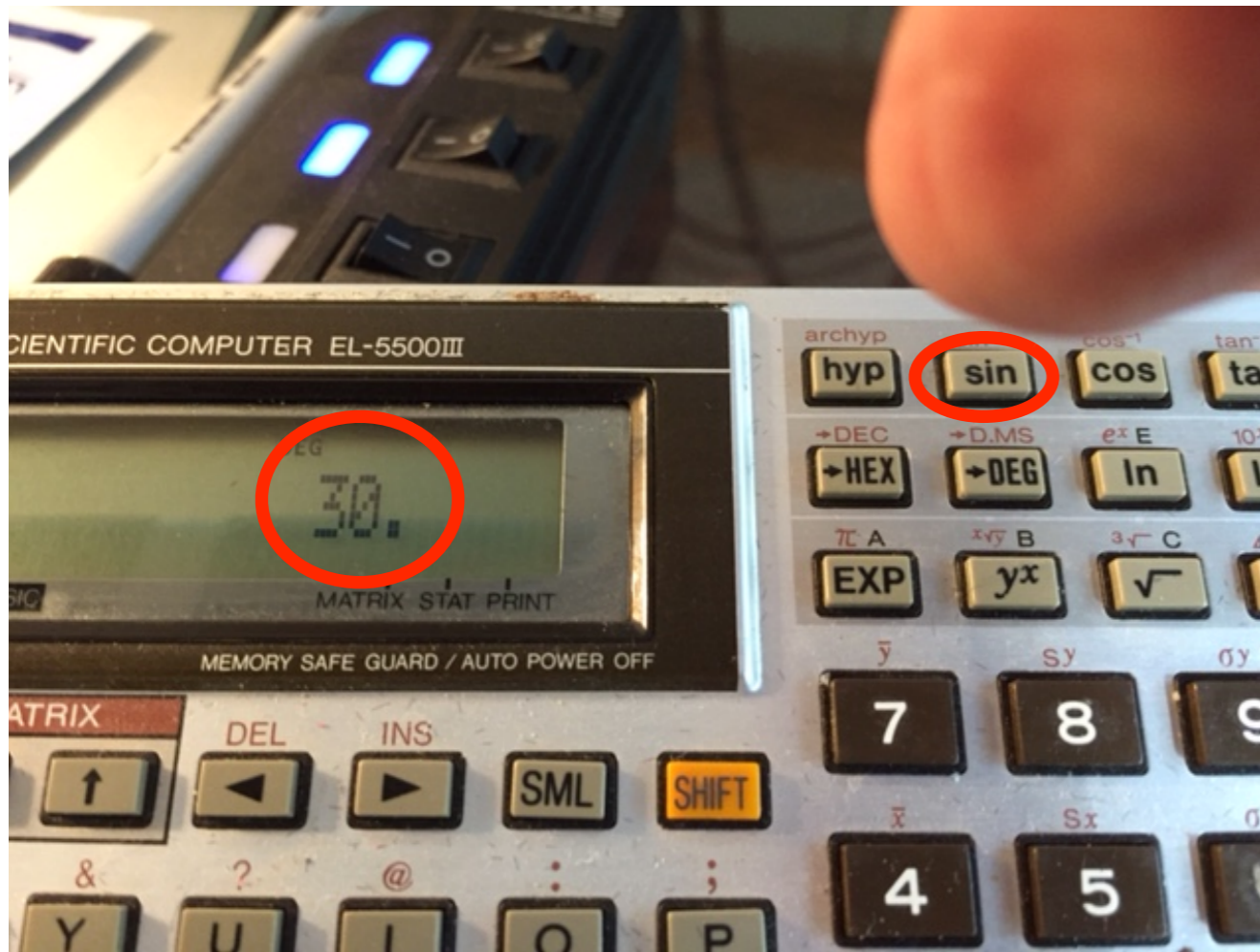
Bad Suzy!

Relativity and Quantum Theory
need to be unified in *one* book
half the size of those old tomes!

We call that a *Relawavity* book.
(It's a *lot* **lighter!**)

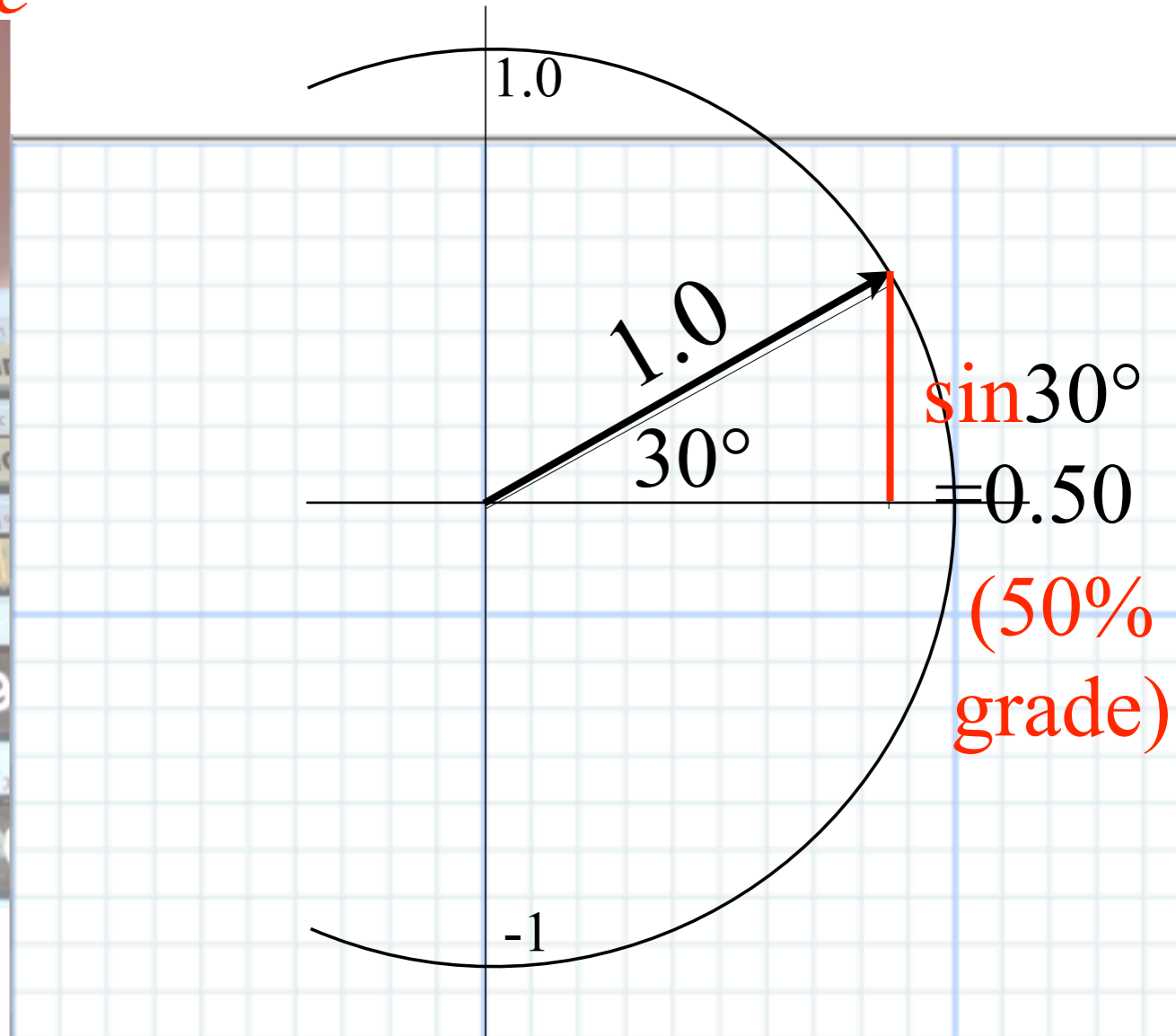
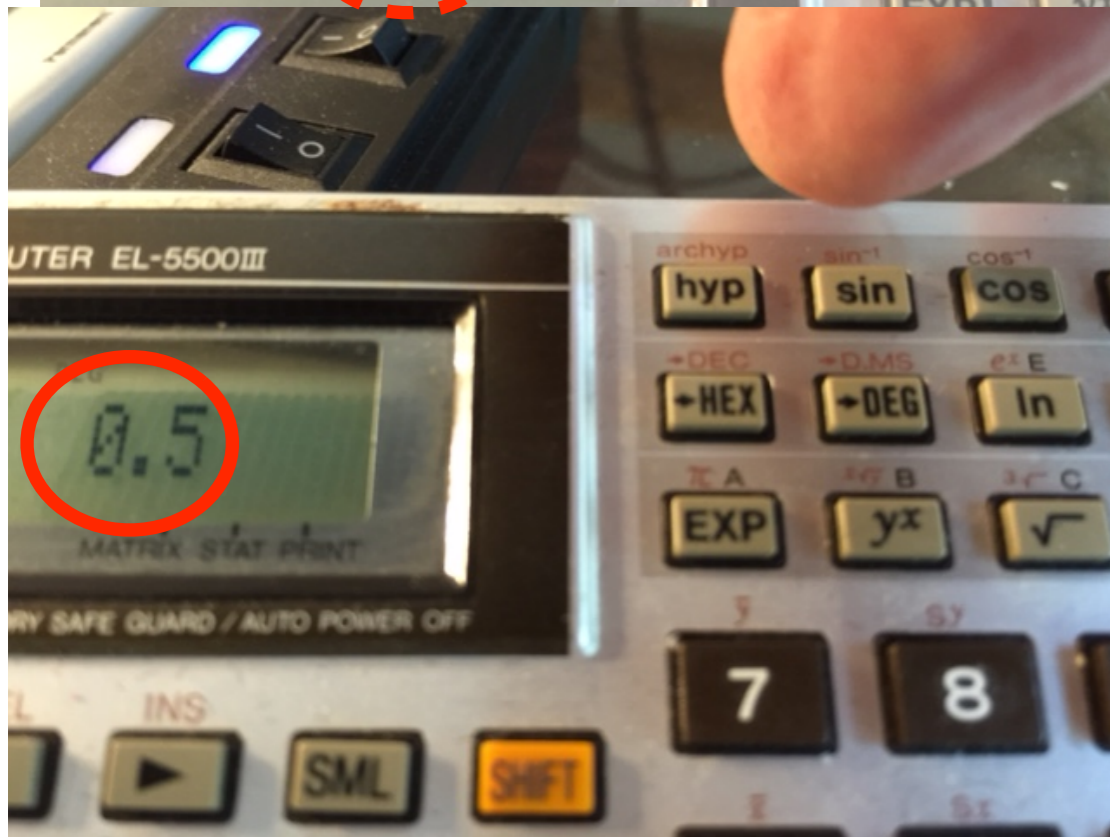
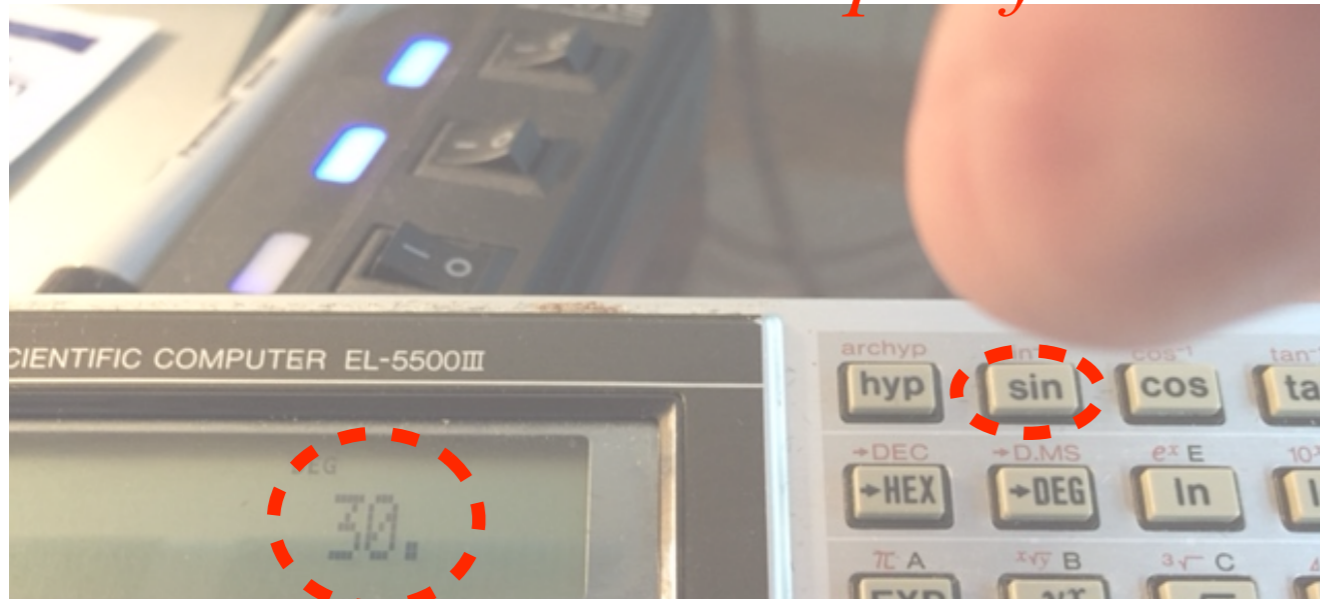


Learning about SIN



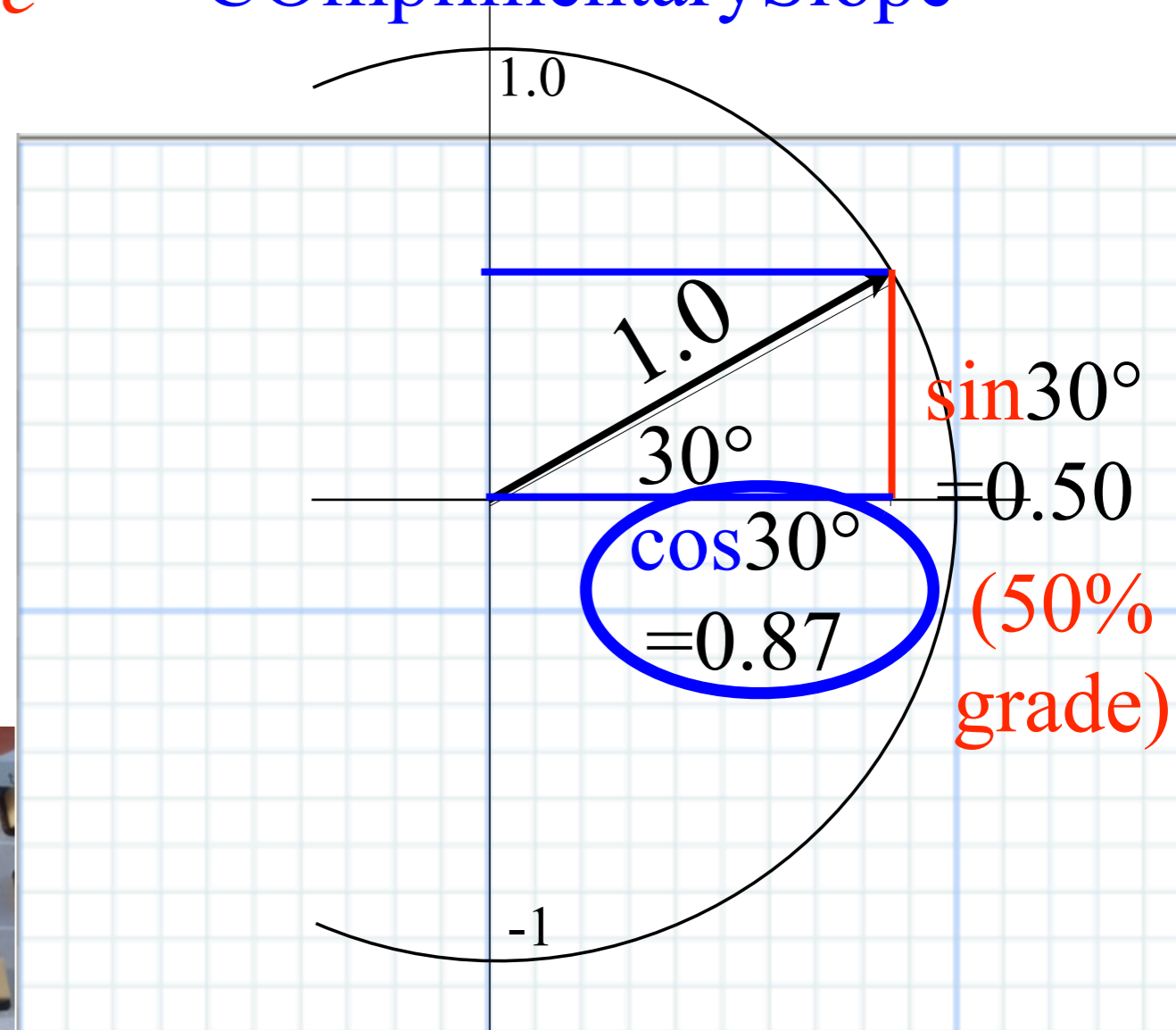
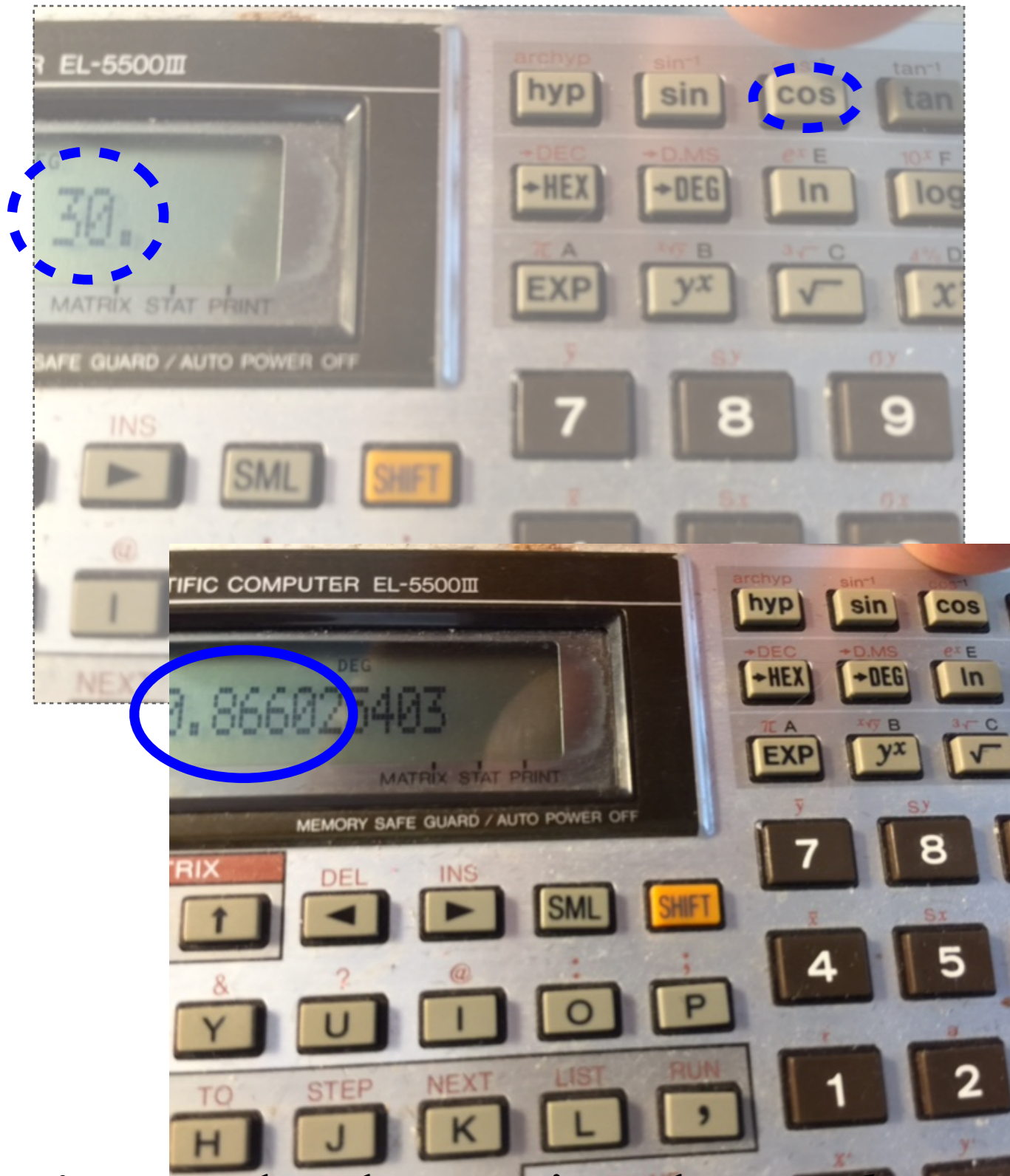
Learning about SIN

“Slope of INcline”



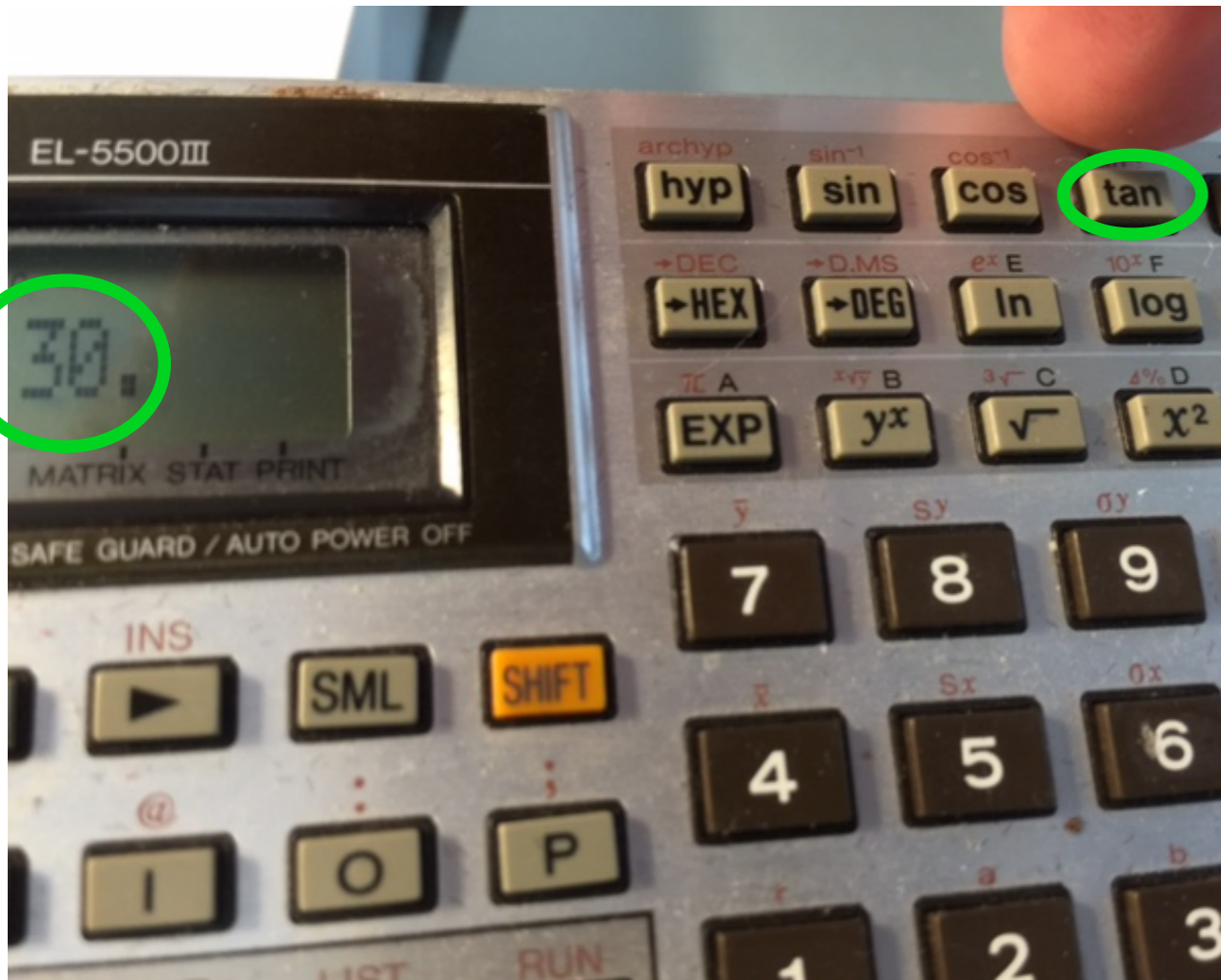
It's mostly about triangles *and sine-waves*

Learning about **SIN** and the **COS** in “*Slope of INcline*” “**C**OMplimentary**S**lope”



It's mostly about triangles *and sine-waves and cosine-waves*

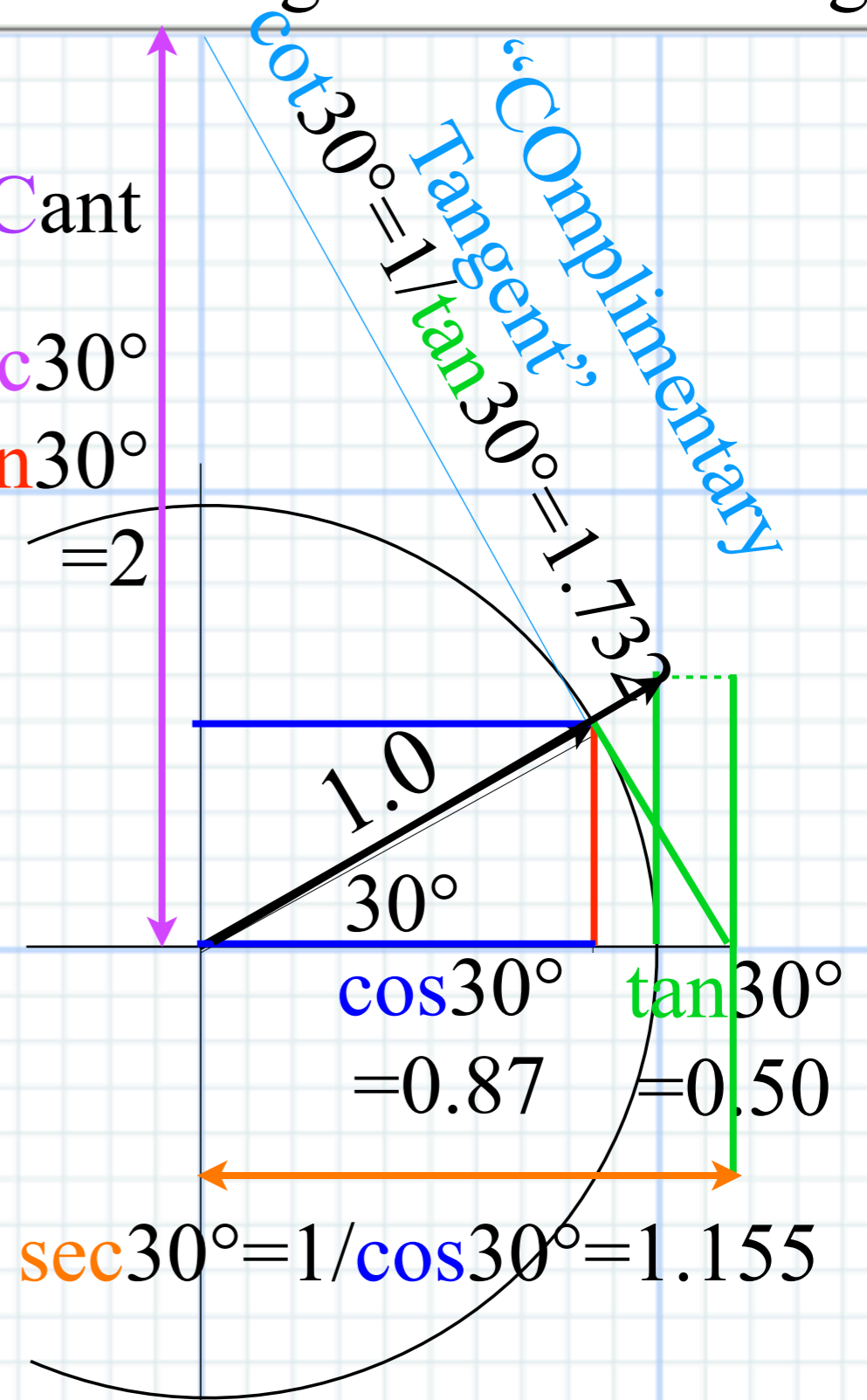
Learning about **SIN** and the **COS**in and **TAN**gent and **CO**Tangent *“Slope of INcline”*



...and
CoSeCant

$$\text{csc}30^\circ = 1/\text{sin}30^\circ$$

$$= 2$$



...and **SEC**ant

It's mostly about triangles *and sine-waves and cosine-waves*

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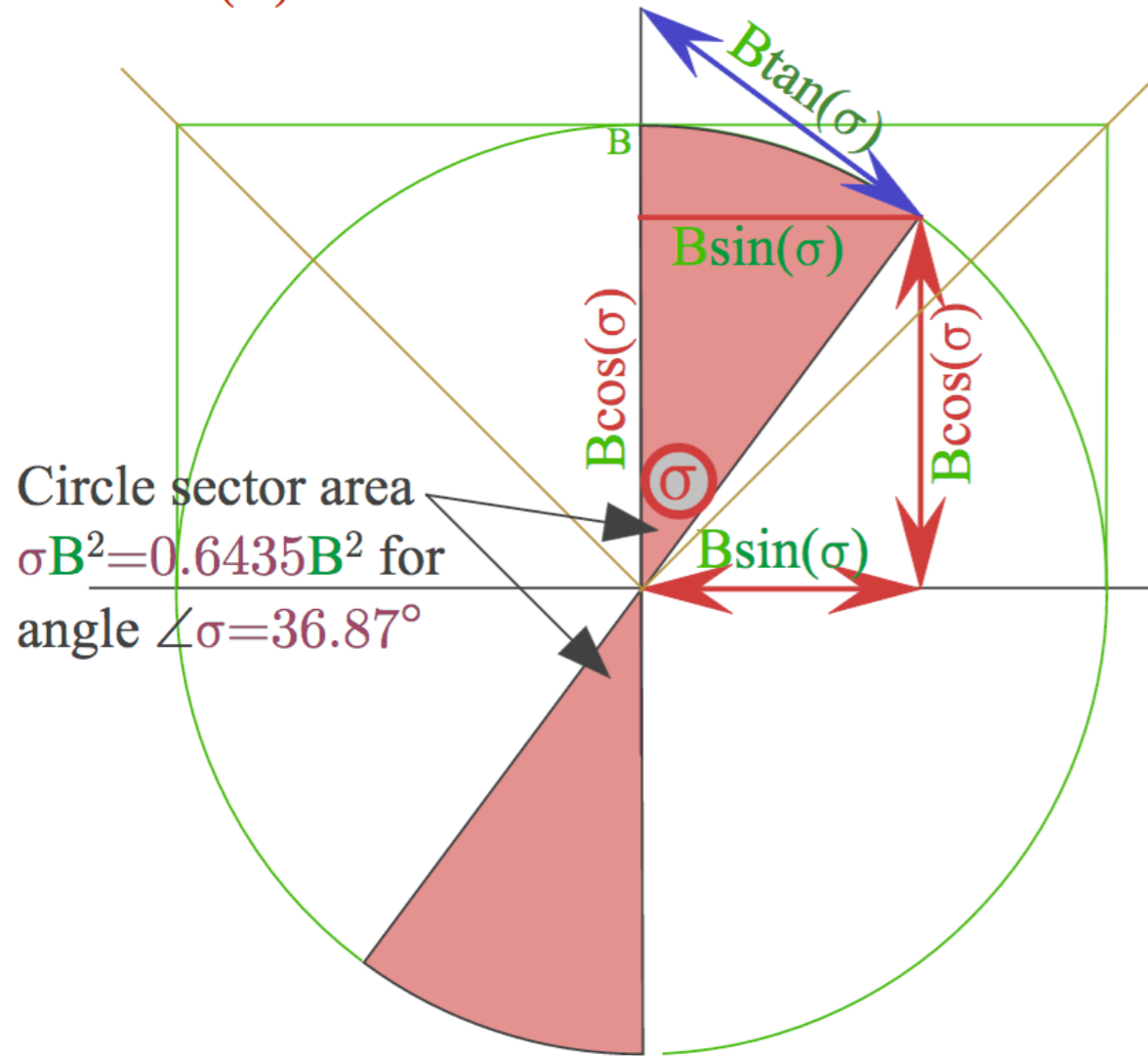
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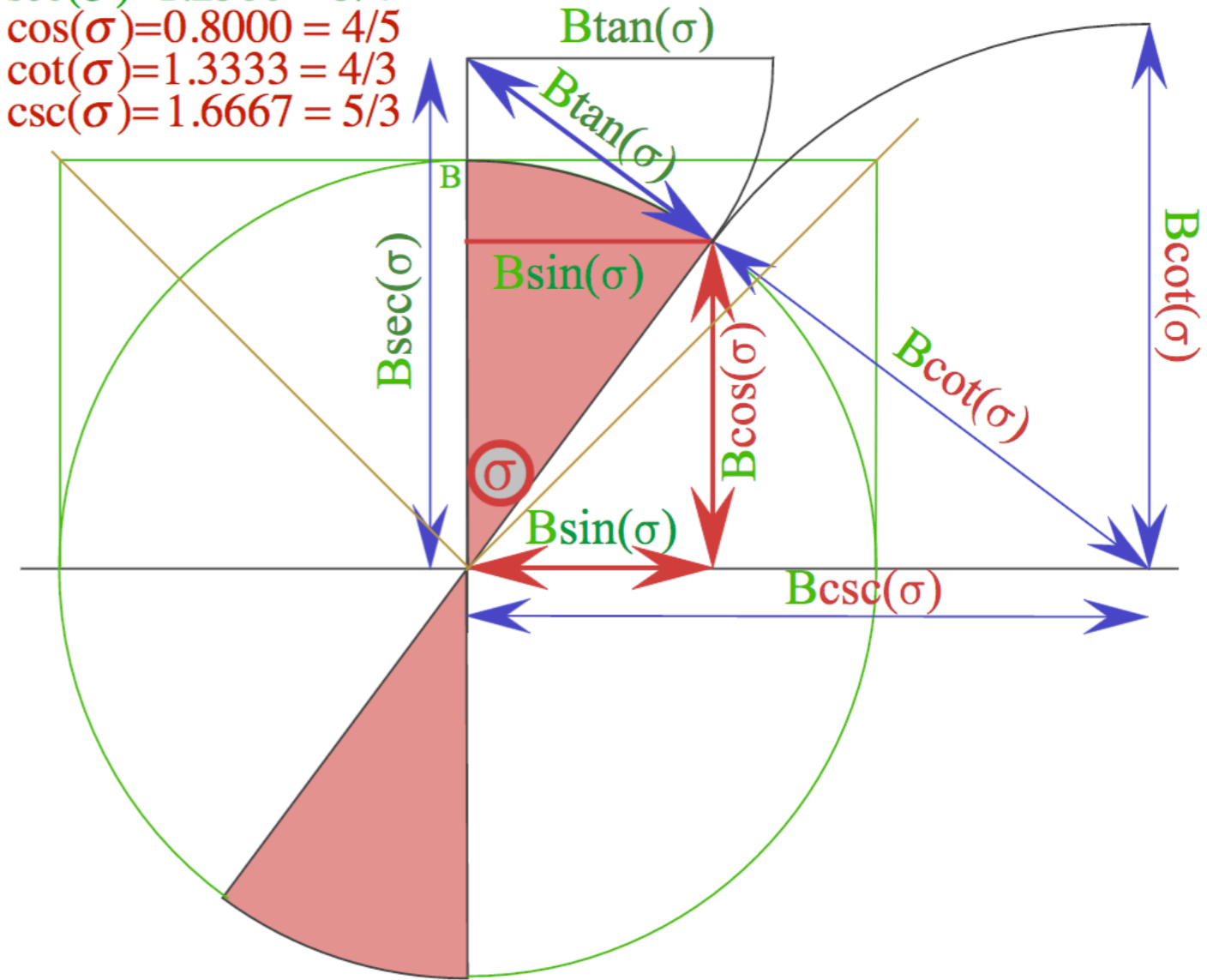
Application to TE-Waveguide modes and synchrotron beam relativity

Trigonometric road maps

(a) $\sin(\sigma) = 0.6000 = 3/5$
 $\tan(\sigma) = 0.7500 = 3/4$
 $\cos(\sigma) = 0.8000 = 4/5$



(b) $\sin(\sigma) = 0.6000 = 3/5$
 $\tan(\sigma) = 0.7500 = 3/4$
 $\sec(\sigma) = 1.2500 = 5/4$
 $\cos(\sigma) = 0.8000 = 4/5$
 $\cot(\sigma) = 1.3333 = 4/3$
 $\csc(\sigma) = 1.6667 = 5/3$



*All this physics of relativity
 is mostly simple trigonometry
 of optical wave interference!*

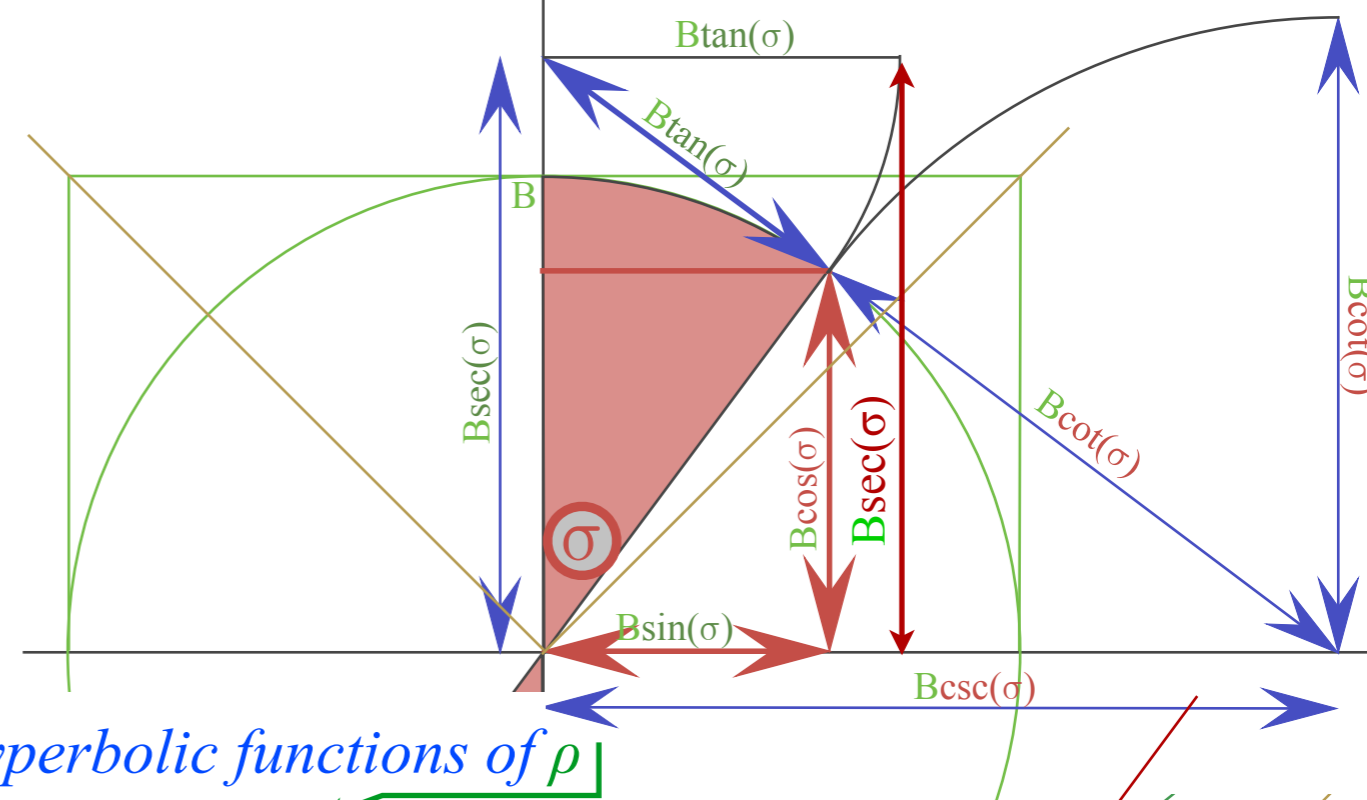
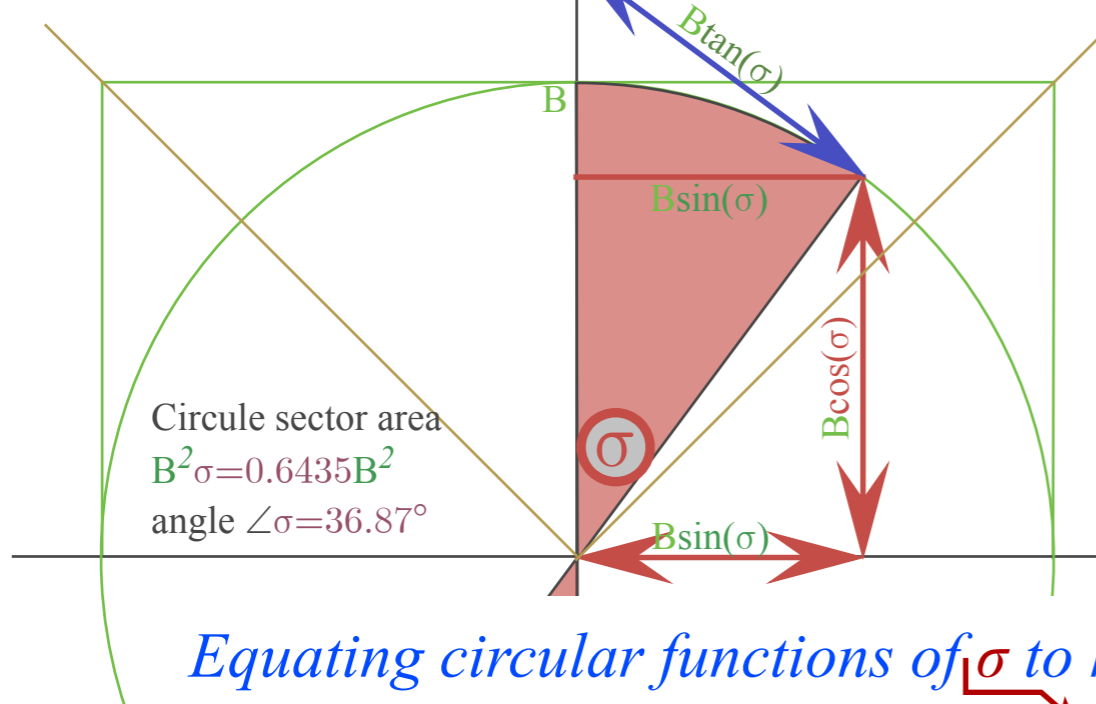
*And, it derives fundamentals
 of quantum theory, too!*

Trigonometric road maps become hyperbolic trig maps...

(a) $\sin(\sigma) = 0.6000 = 3/5$

(b)

$\cos(\sigma) = 0.8000 = 4/5$



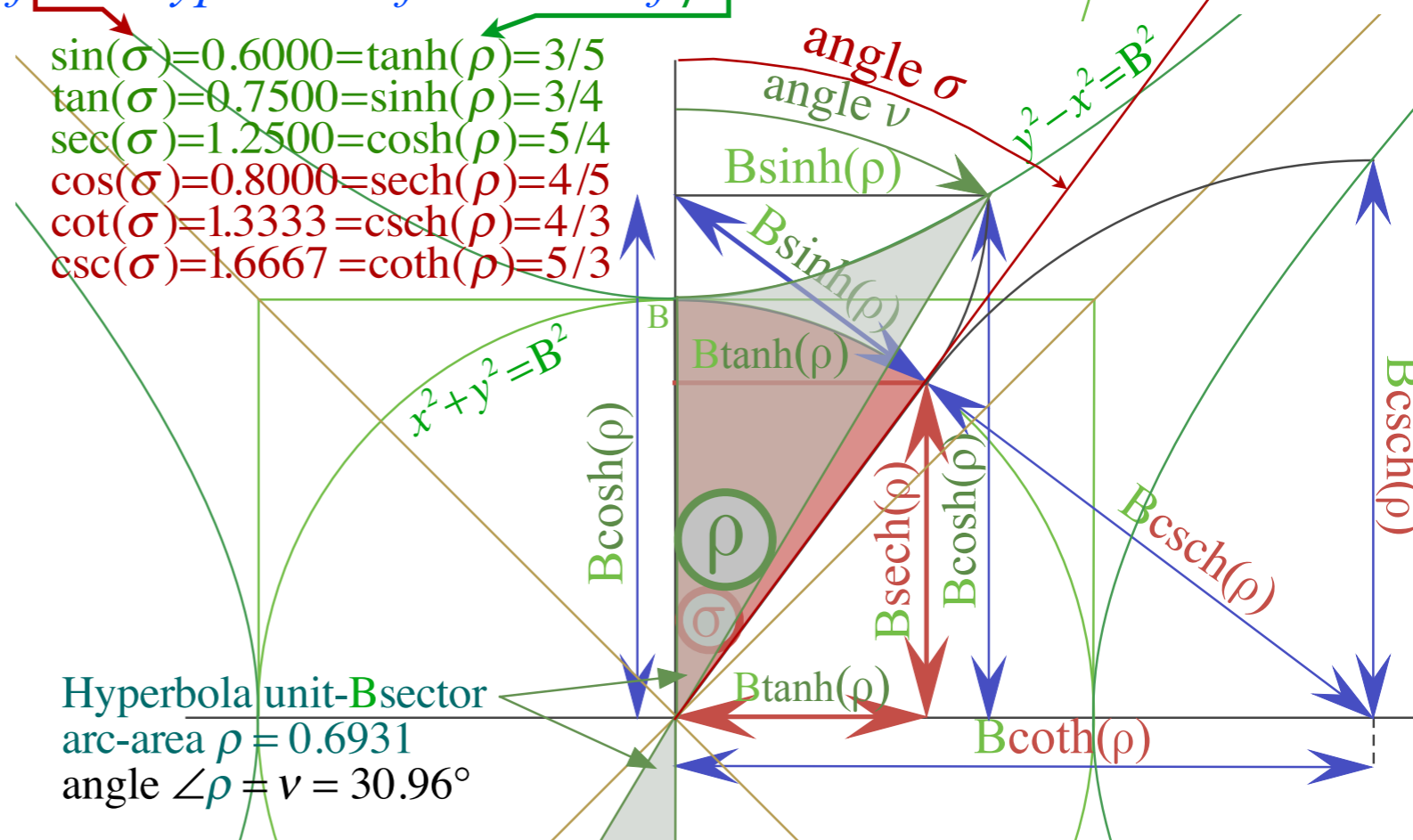
Equating circular functions of σ to hyperbolic functions of ρ

$\sin(\sigma) = 0.6000 = \tanh(\rho) = 3/5$
 $\tan(\sigma) = 0.7500 = \sinh(\rho) = 3/4$
 $\sec(\sigma) = 1.2500 = \cosh(\rho) = 5/4$
 $\cos(\sigma) = 0.8000 = \operatorname{sech}(\rho) = 4/5$
 $\cot(\sigma) = 1.3333 = \operatorname{csch}(\rho) = 4/3$
 $\operatorname{csc}(\sigma) = 1.6667 = \operatorname{coth}(\rho) = 5/3$

All this physics of relativity is mostly simple trigonometry of optical wave interference!

And, it derives fundamentals of quantum theory, too!

Hyperbola unit-B sector
arc-area $\rho = 0.6931$
angle $\angle \rho = \nu = 30.96^\circ$



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Hyper-Trigonometric Relativity geometry and Euler exponential algebra ←

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Hyper-Trigonometric algebra easily derives Circular-Trigonometric-algebra

Exponential derived by infinite- n -compounding limit of the interest rate- r formula.

$$e^{rt} = \lim_{n \rightarrow \infty} \left(1 + \frac{rt}{n} \right)^n$$

Infinite- n limit of binomial series is an exponential power- p series of $(rt)^p$ with $1/p!$ coefficients.

$$e^{rt} = 1 + rt + \frac{(rt)^2}{2} + \frac{(rt)^3}{2 \cdot 3} + \frac{(rt)^4}{2 \cdot 3 \cdot 4} + \frac{(rt)^5}{2 \cdot 3 \cdot 4 \cdot 5} + \frac{(rt)^6}{2 \cdot 3 \cdot 4 \cdot 5 \cdot 6} + \dots$$
$$e^{-rt} = 1 - rt + \frac{(rt)^2}{2} - \frac{(rt)^3}{2 \cdot 3} + \frac{(rt)^4}{2 \cdot 3 \cdot 4} - \frac{(rt)^5}{2 \cdot 3 \cdot 4 \cdot 5} + \frac{(rt)^6}{2 \cdot 3 \cdot 4 \cdot 5 \cdot 6} - \dots$$

Half-sum and half difference of $e^{\pm rt}$ series define the hyperbolic cosine ($\cosh(rt)$) and sine ($\sinh(rt)$).

$$\frac{e^{+rt} + e^{-rt}}{2} = 1 + \frac{(rt)^2}{2} + \frac{(rt)^4}{2 \cdot 3 \cdot 4} + \frac{(rt)^6}{2 \cdot 3 \cdot 4 \cdot 5 \cdot 6} - \dots = \cosh(rt)$$
$$\frac{e^{+rt} - e^{-rt}}{2} = rt + \frac{(rt)^3}{2 \cdot 3} + \frac{(rt)^5}{2 \cdot 3 \cdot 4 \cdot 5} + \dots = \sinh(rt)$$

Hyper-Trig
 $\cosh \rho$ and $\sinh \rho$

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$$e^{-rt} = 1 - rt + \frac{(rt)^2}{2} - \frac{(rt)^3}{2 \cdot 3} + \frac{(rt)^4}{2 \cdot 3 \cdot 4} - \frac{(rt)^5}{2 \cdot 3 \cdot 4 \cdot 5} + \frac{(rt)^6}{2 \cdot 3 \cdot 4 \cdot 5 \cdot 6} - \dots = \cosh(rt) - \sinh(rt)$$

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$$\frac{e^{+rt} - e^{-rt}}{2} = rt + \frac{(rt)^3}{2 \cdot 3} + \frac{(rt)^5}{2 \cdot 3 \cdot 4 \cdot 5} + \dots = \sinh(rt)$$

Hyper-Trig
 $\cosh \rho$ and $\sinh \rho$

Replace rate r with imaginary rate ir and $i = \sqrt{-1}$ powers $i^0=1, i^1=i, i^2=-1, i^3=-i, i^4=1, i^5=i, i^6=-1, i^7=-i, \dots$

Then *hyper*-sine-cosine becomes the *circular*-sine-cosine.

$$\frac{e^{+i rt} + e^{-i rt}}{2} = 1 - \frac{(rt)^2}{2} + \frac{(rt)^4}{2 \cdot 3 \cdot 4} - \frac{(rt)^6}{2 \cdot 3 \cdot 4 \cdot 5 \cdot 6} - \dots = \cos rt$$

$$\frac{e^{+i rt} - e^{-i rt}}{2} = i rt - i \frac{(rt)^3}{2 \cdot 3} + i \frac{(rt)^5}{2 \cdot 3 \cdot 4 \cdot 5} - \dots = i \sin rt$$

Circular-Trig
 $\cos \sigma$ and $\sin \sigma$

Sum and difference of this pair gives the Euler-DeMoivre relations of exponentials vs trig-functions.

$$e^{+i\sigma} = \cos \sigma + i \sin \sigma ,$$

$$e^{+\rho} = \cosh \rho + \sinh \rho ,$$

$$e^{-i\sigma} = \cos \sigma - i \sin \sigma .$$

$$e^{-\sigma} = \cosh \rho - \sinh \rho .$$

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1CW Laser-phasor wave function

Dimensionless Light wave-velocity $c/c=1$

$$\frac{v_{\text{light}}}{c} = \frac{\lambda}{c\tau} = \frac{v}{c\kappa} = 1 = \frac{\omega \text{ angular}}{ck \text{ units}}$$

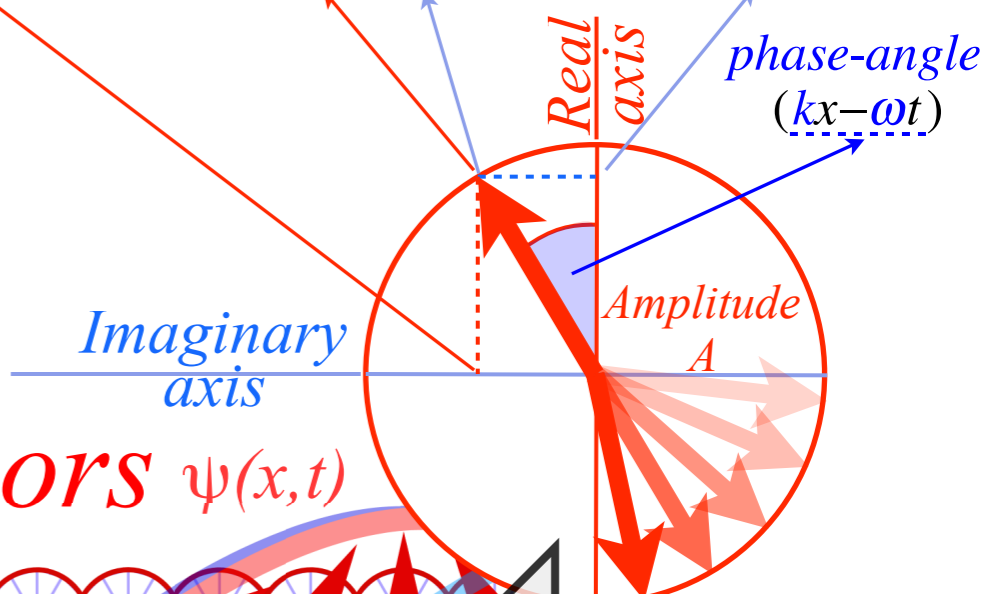
“winks”
“n”
“kinks”

angular frequency: $\omega = 2\pi\nu$
angular wave number: $k = 2\pi\kappa$

$k = \text{wavevector}$

$$\psi = A \cdot e^{i(kx - \omega t)} = A \cdot \cos(kx - \omega t) + iA \cdot \sin(kx - \omega t)$$

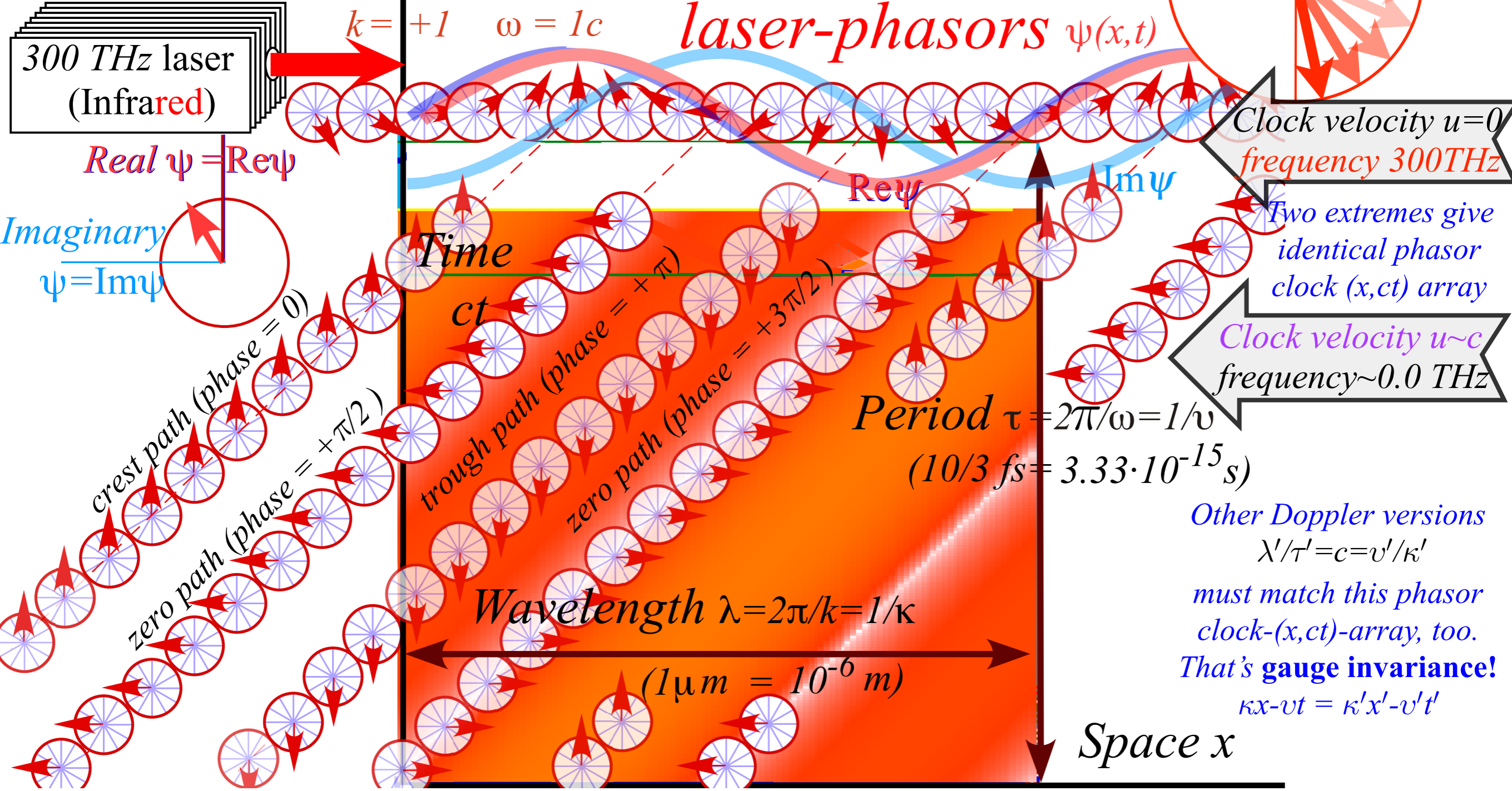
\uparrow Amplitude A \uparrow phase-angle
 \uparrow phase-angle $(kx - \omega t)$



laser-phasors $\psi(x,t)$

300 THz laser
(Infrared)

$k = +1$ $\omega = 1c$



Clock velocity $u=0$
frequency 300THz

Two extremes give
identical phasor
clock (x,ct) array

Clock velocity $u \sim c$
frequency ~ 0.0 THz

Period $\tau = 2\pi/\omega = 1/\nu$
(10/3 fs = $3.33 \cdot 10^{-15}$ s)

Other Doppler versions
 $\lambda'/\tau' = c = v'/\kappa'$
must match this phasor
clock- (x,ct) -array, too.
That's gauge invariance!
 $\kappa x - \nu t = \kappa' x' - \nu' t'$

Real $\psi = \text{Re}\psi$

Imaginary
 $\psi = \text{Im}\psi$

crest path (phase = 0)
zero path (phase = $+\pi/2$)
trough path (phase = $+\pi$)
zero path (phase = $+3\pi/2$)

Time

ct

Wavelength $\lambda = 2\pi/k = 1/\kappa$

($1 \mu m = 10^{-6} m$)

Space x

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The "Keyboard of the gods" : per-space-per-time plot versus space-time Minkowski plot



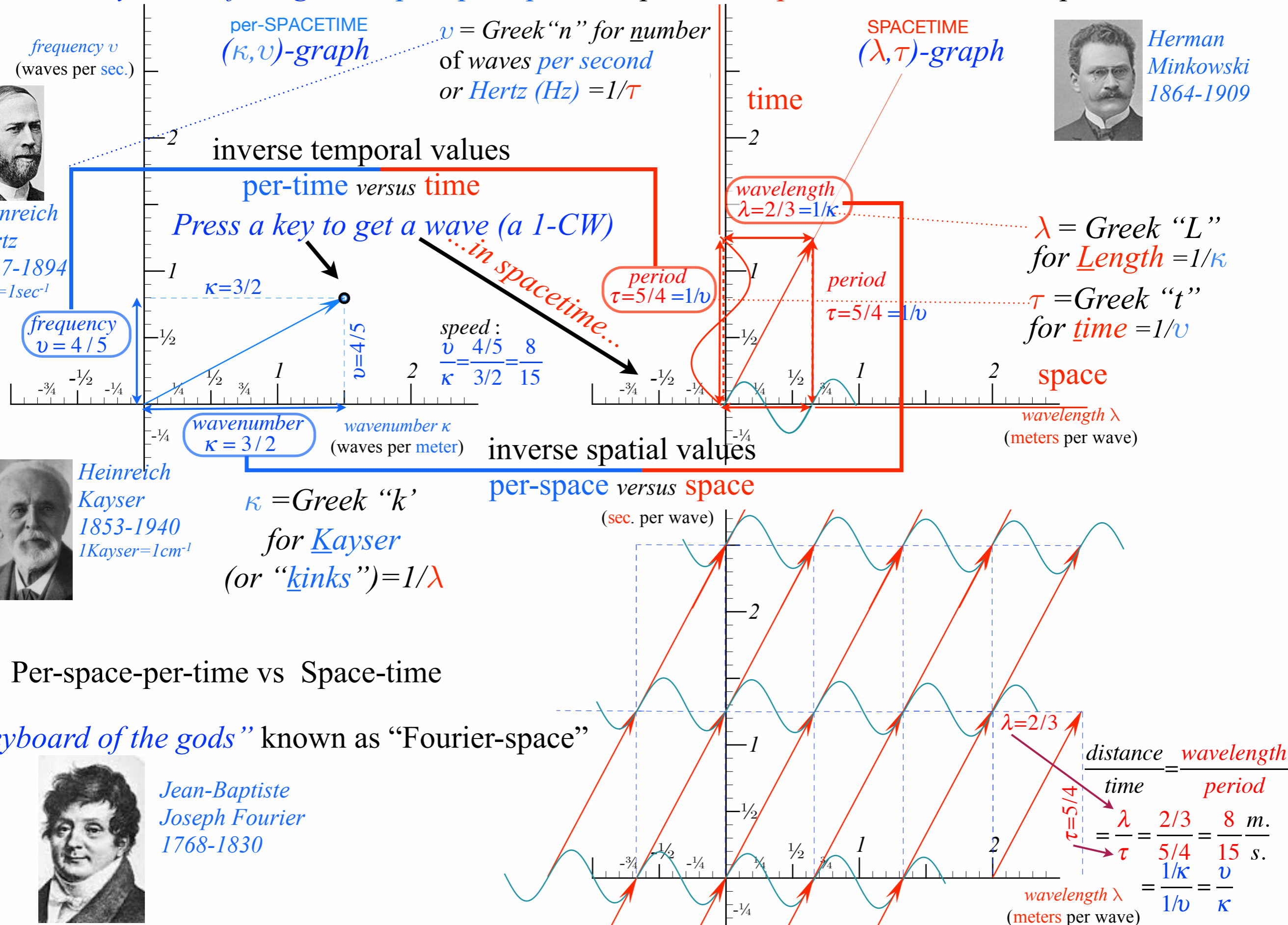
Heinrich Hertz
1857-1894
1Hz=1sec⁻¹



Heinrich Kayser
1853-1940
1Kayser=1cm⁻¹



Herman Minkowski
1864-1909



Per-space-per-time vs Space-time

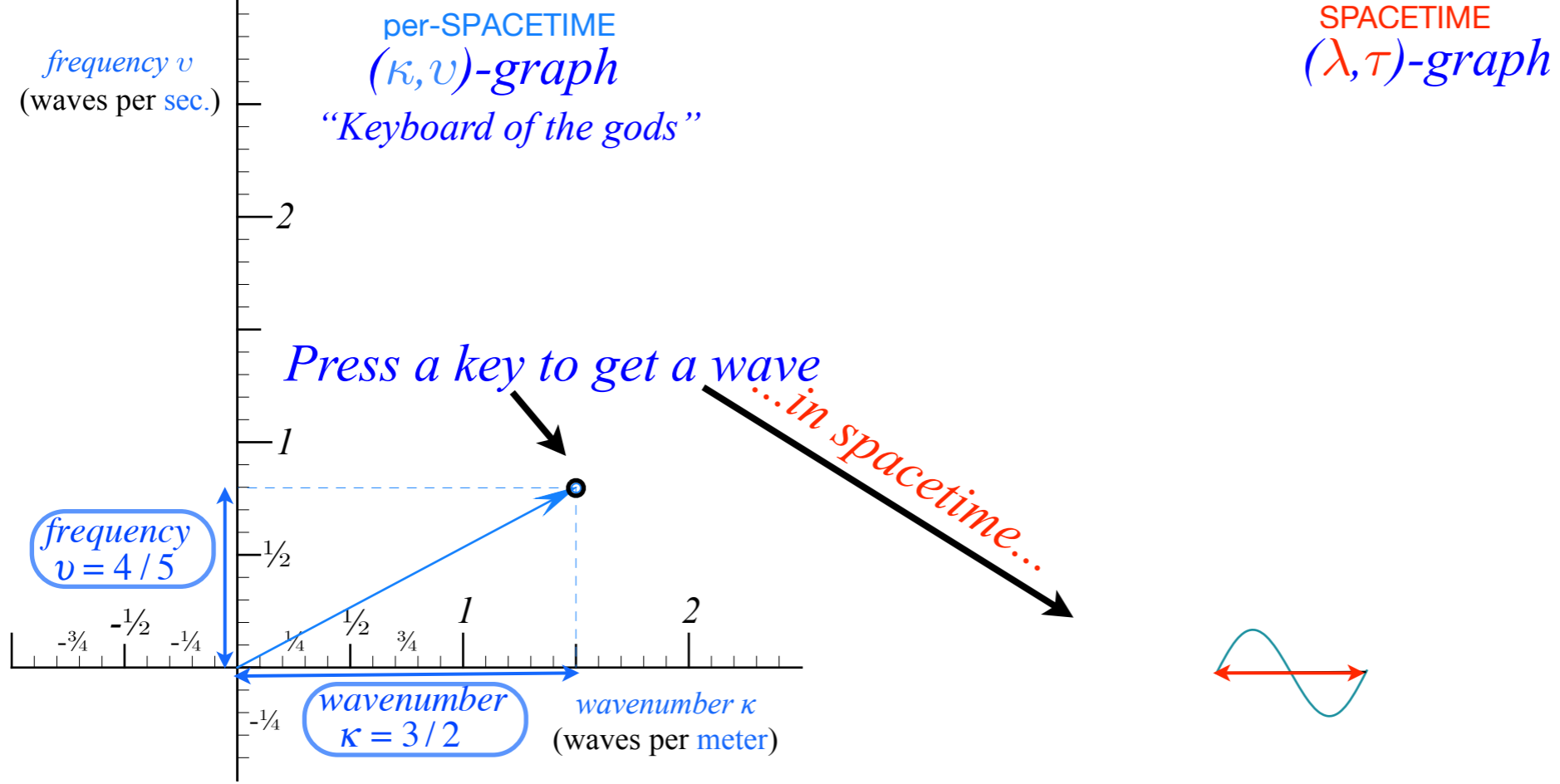
"Keyboard of the gods" known as "Fourier-space"



Jean-Baptiste Joseph Fourier
1768-1830

Fig. 5 Comparing a wave point in Kaiser-Hertz per-space-time to its Minkowski space-time view.

Analyzing wave velocity by **per-space-per-time** and **space-time** graphs



"Keyboard of the gods" is known as "Fourier-space"



Jean-Baptiste
Joseph Fourier
1768-1830

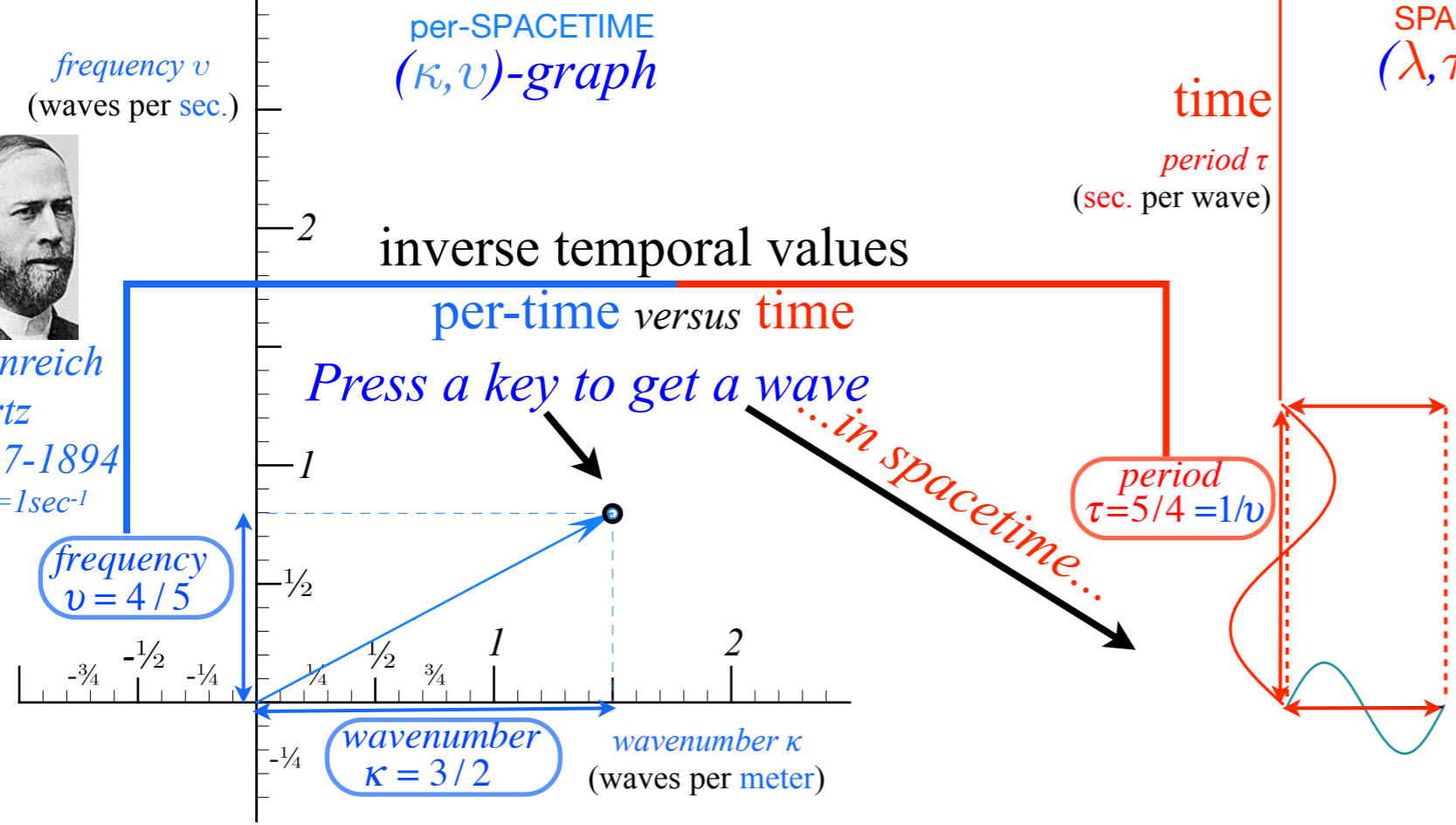
•How to understand waves
and
wave velocity V_{wave}

[RelaWavity Web Simulation](#)
[Keyboard of the Gods](#)
(per-Time vs per-Space)

Analyzing wave velocity by per-space-per-time and space-time graphs



Heinrich Hertz
1857-1894
1Hz=1sec⁻¹



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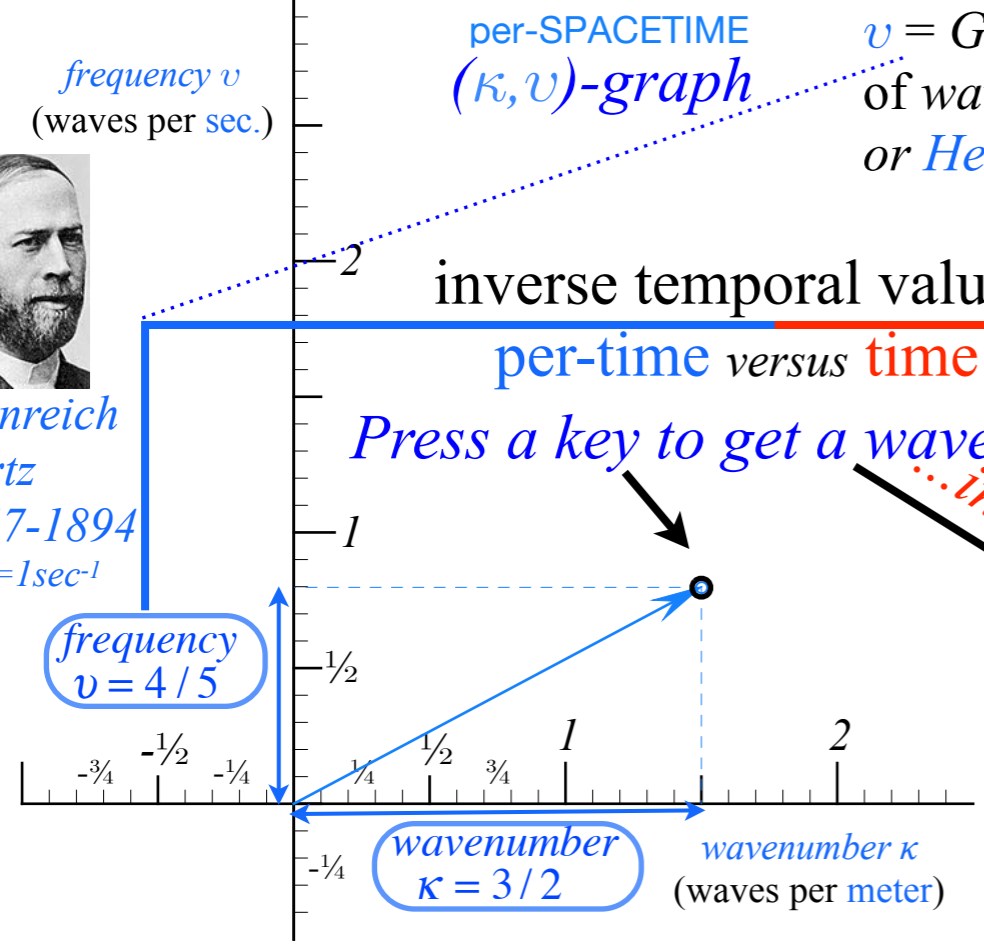
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Analyzing wave velocity by **per-space-per-time** and **space-time** graphs



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1857-1894
1Hz=1sec⁻¹



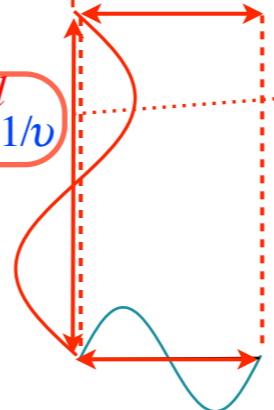
$v = \text{Greek "n" for number of waves per second or Hertz (Hz)} = 1/\tau$
 time period τ (sec. per wave)

SPACETIME (λ, τ) -graph

Press a key to get a wave
...in spacetime...

period $\tau = 5/4 = 1/v$

$\tau = \text{Greek "t" for time} = 1/v$



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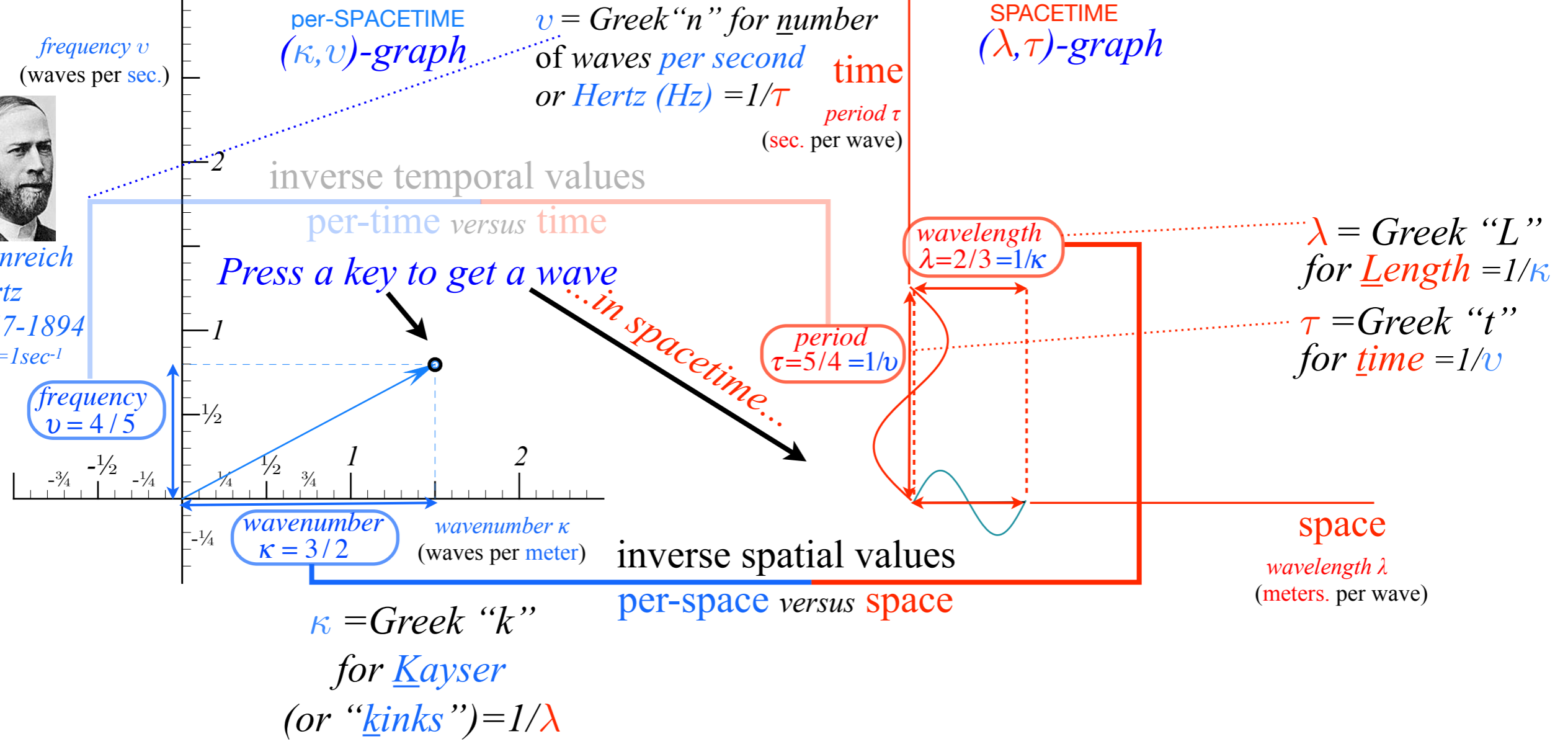
•How to understand waves and wave velocity V_{wave}

[RelaWavity Web Simulation](#)
[Keyboard of the Gods](#)
 (Dual Plot)

Analyzing wave velocity by per-space-per-time and space-time graphs



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1857-1894
1Hz=1sec⁻¹



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•How to understand waves
and
wave velocity V_{wave}

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Heinrich Hertz
1857-1894
1Hz=1sec⁻¹

frequency ν
(waves per sec.)
 $\nu = 4/5$

per-SPACETIME
 (κ, ν) -graph

ν = Greek "n" for number
of waves per second
or Hertz (Hz) = $1/\tau$
time
period τ
(sec. per wave)

SPACETIME
 (λ, τ) -graph

inverse temporal values

per-time versus time

Press a key to get a wave

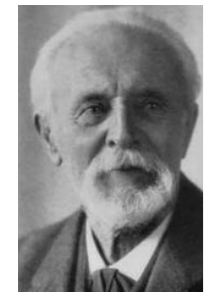
...in spacetime...

period
 $\tau = 5/4 = 1/\nu$

wavelength
 $\lambda = 2/3 = 1/\kappa$

λ = Greek "L" for Length = $1/\kappa$

τ = Greek "t" for time = $1/\nu$



Heinrich Kayser
1853-1940
1Kayser=1cm⁻¹

wavenumber
 $\kappa = 3/2$
wavenumber κ
(waves per meter)

inverse spatial values

per-space versus space

κ = Greek "k" for Kayser
(or "kinks") = $1/\lambda$

space
wavelength λ
(meters. per wave)

"Keyboard of the gods" is known as "Fourier-space"

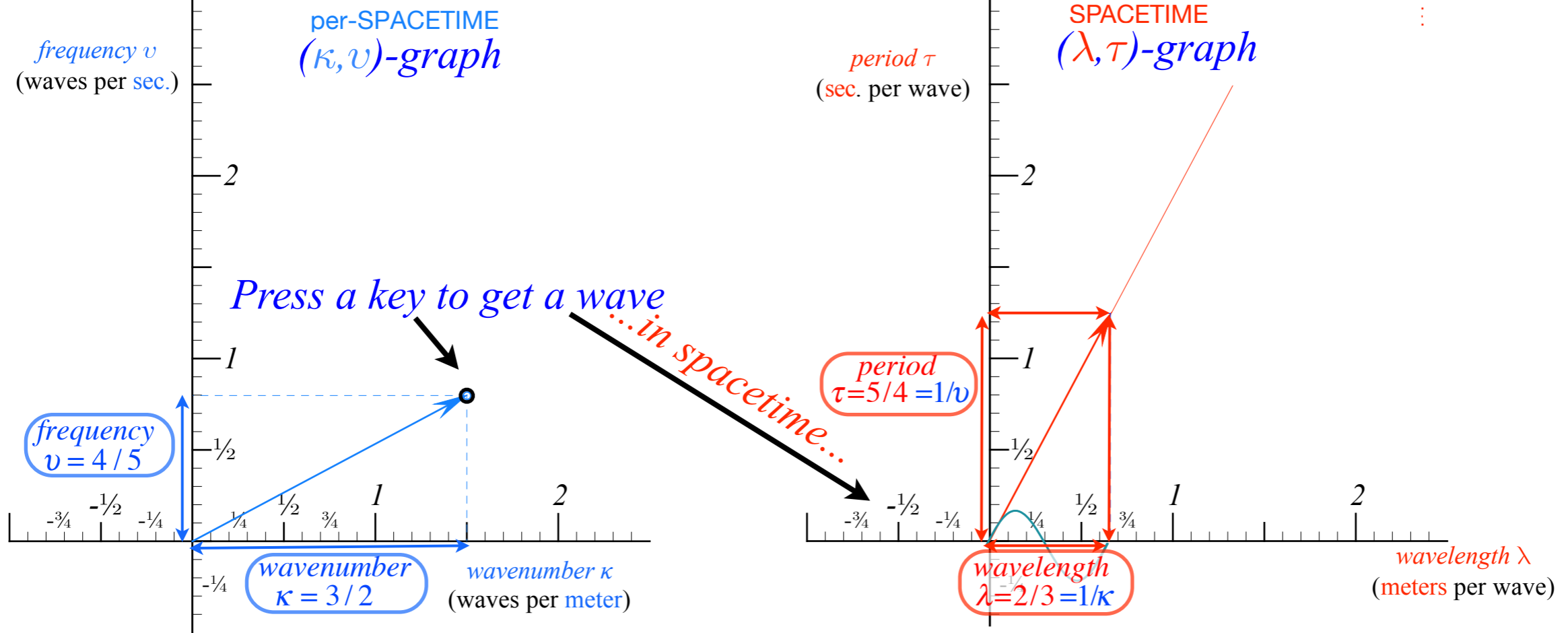


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[RelaWavity Web Simulation](#)
[Keyboard of the Gods](#)
(Dual Plot)

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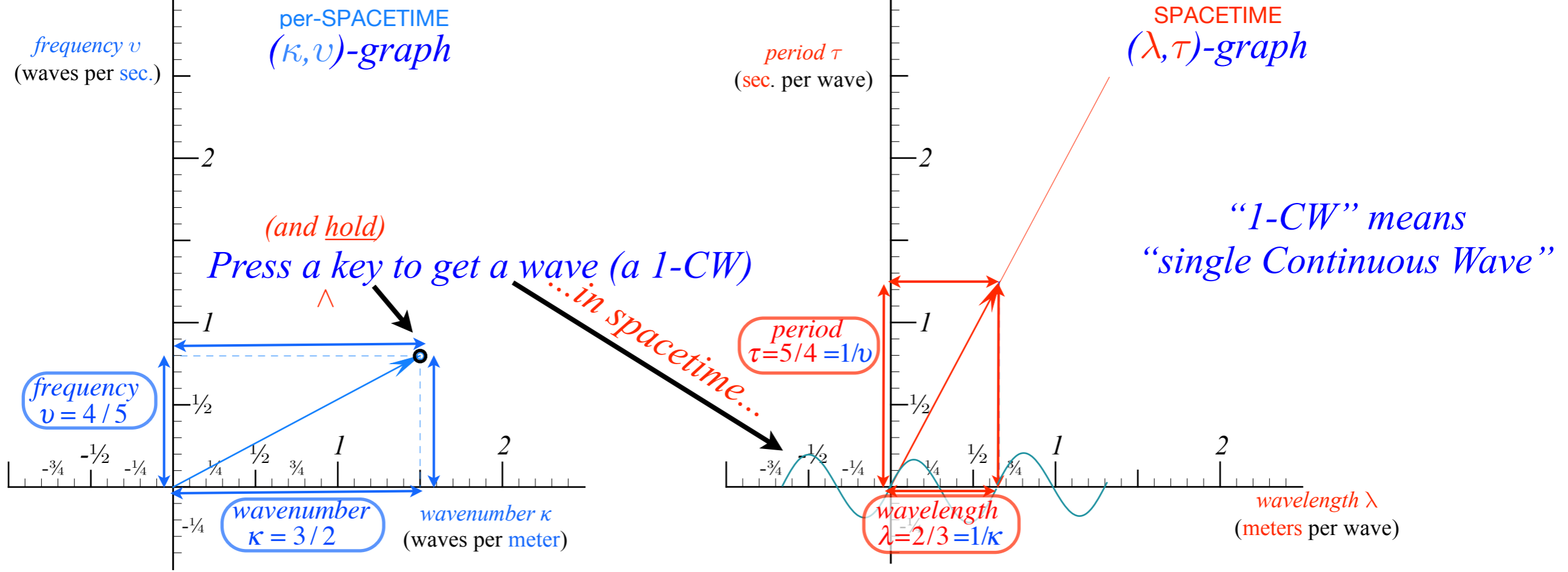
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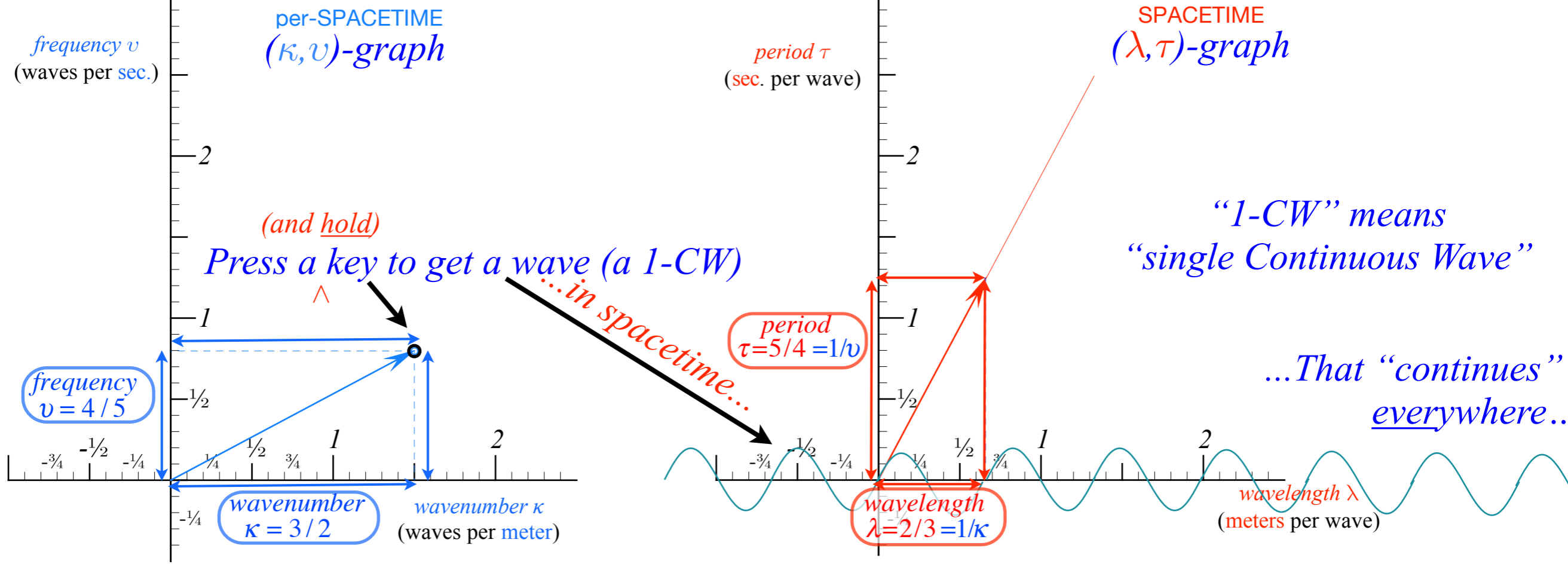


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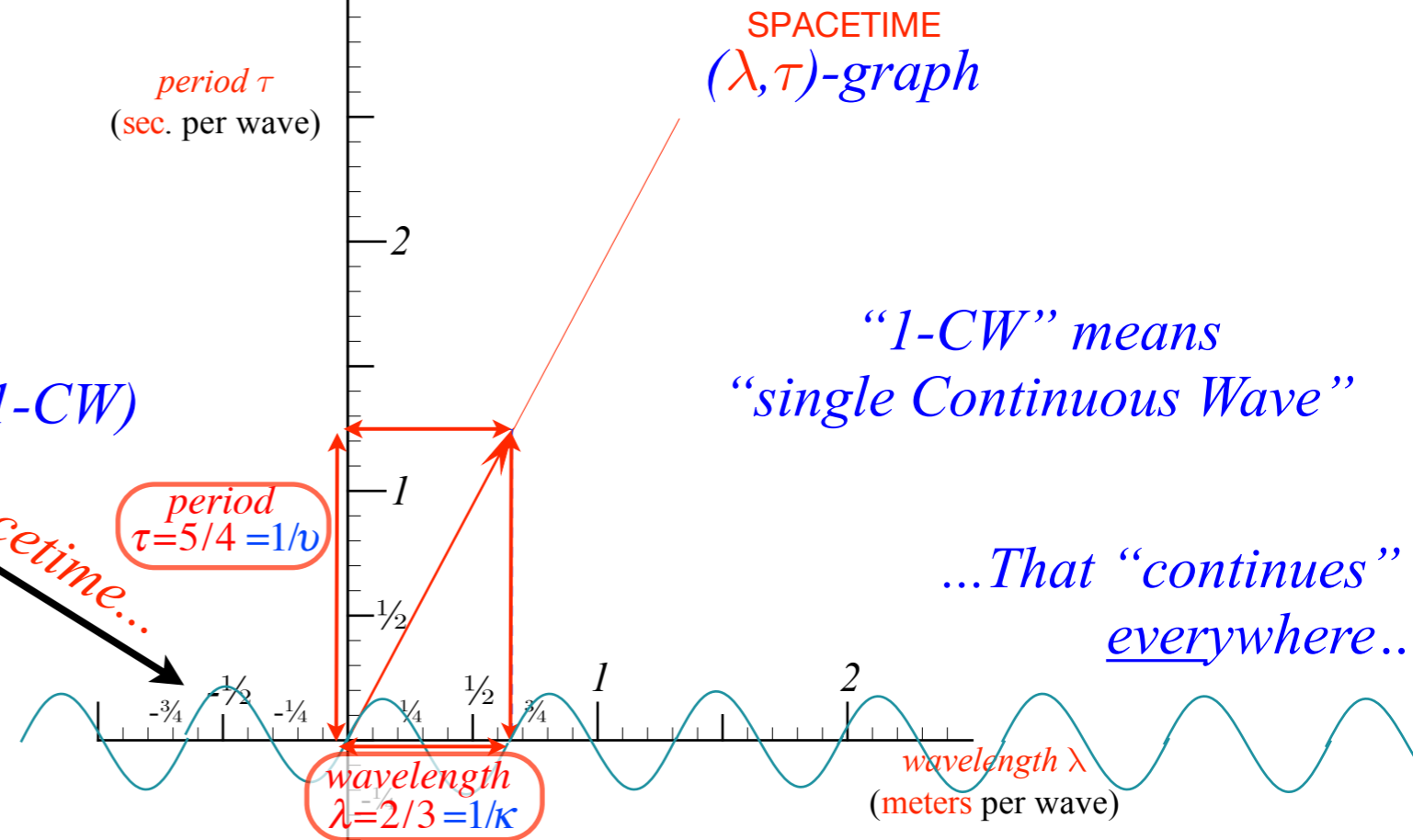
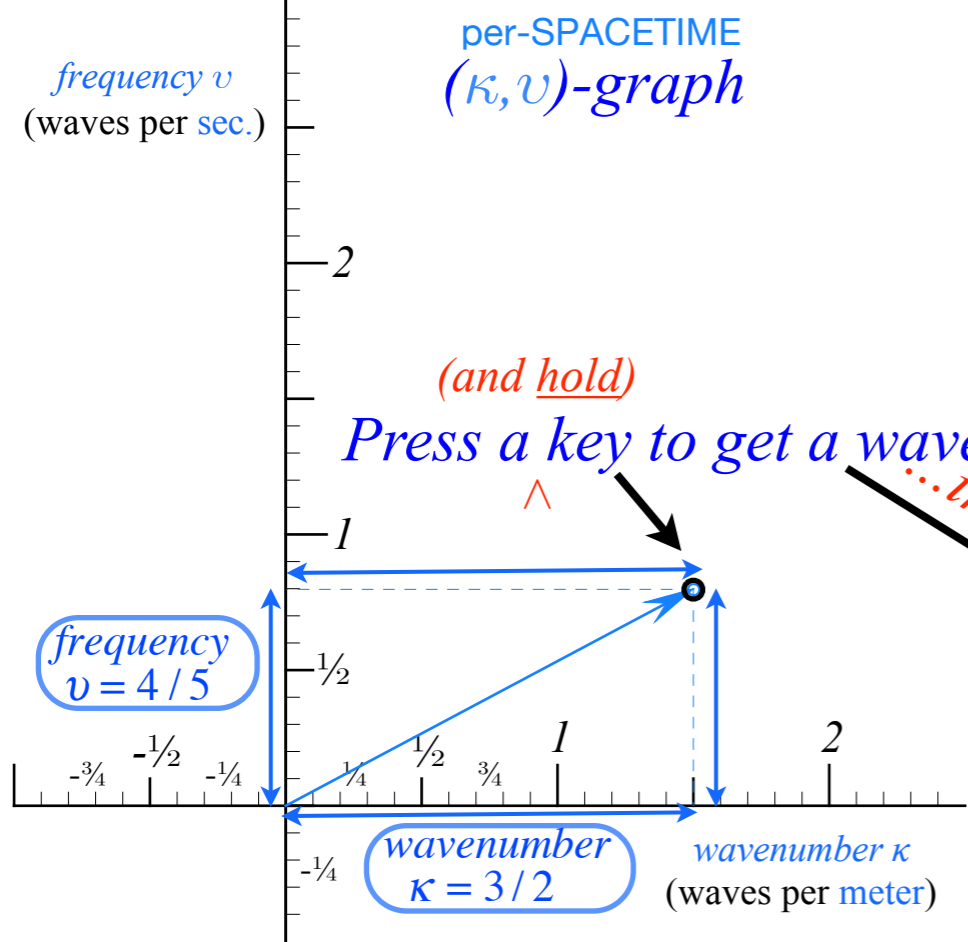


Jean-Baptiste
Joseph Fourier
1768-1830

•How to understand waves
and
wave velocity V_{wave}

[RelaWavity Web Simulation](#)
[Keyboard of the Gods](#)
(Dual Plot)

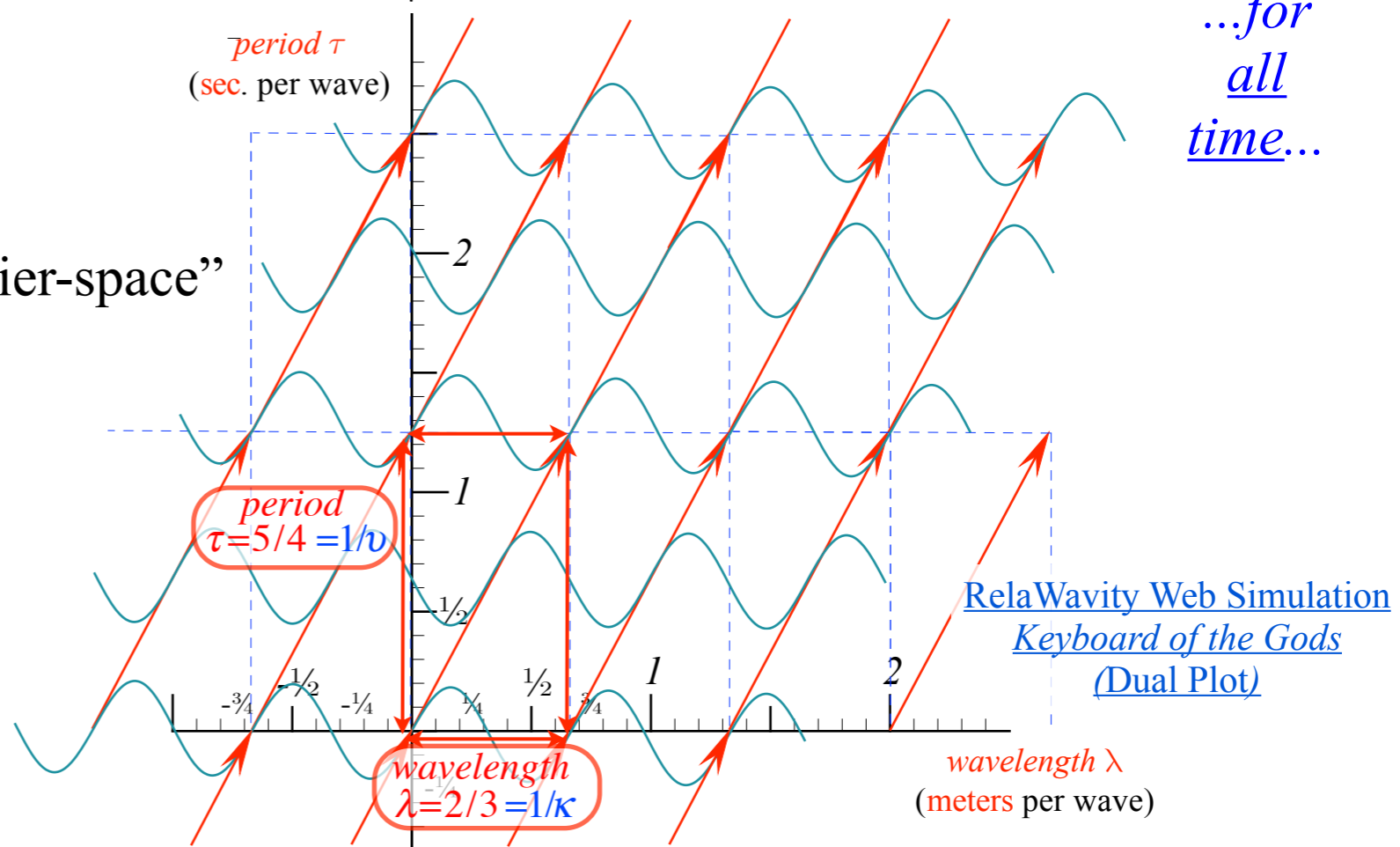
Analyzing wave velocity by **per-space-per-time** and **space-time** graphs



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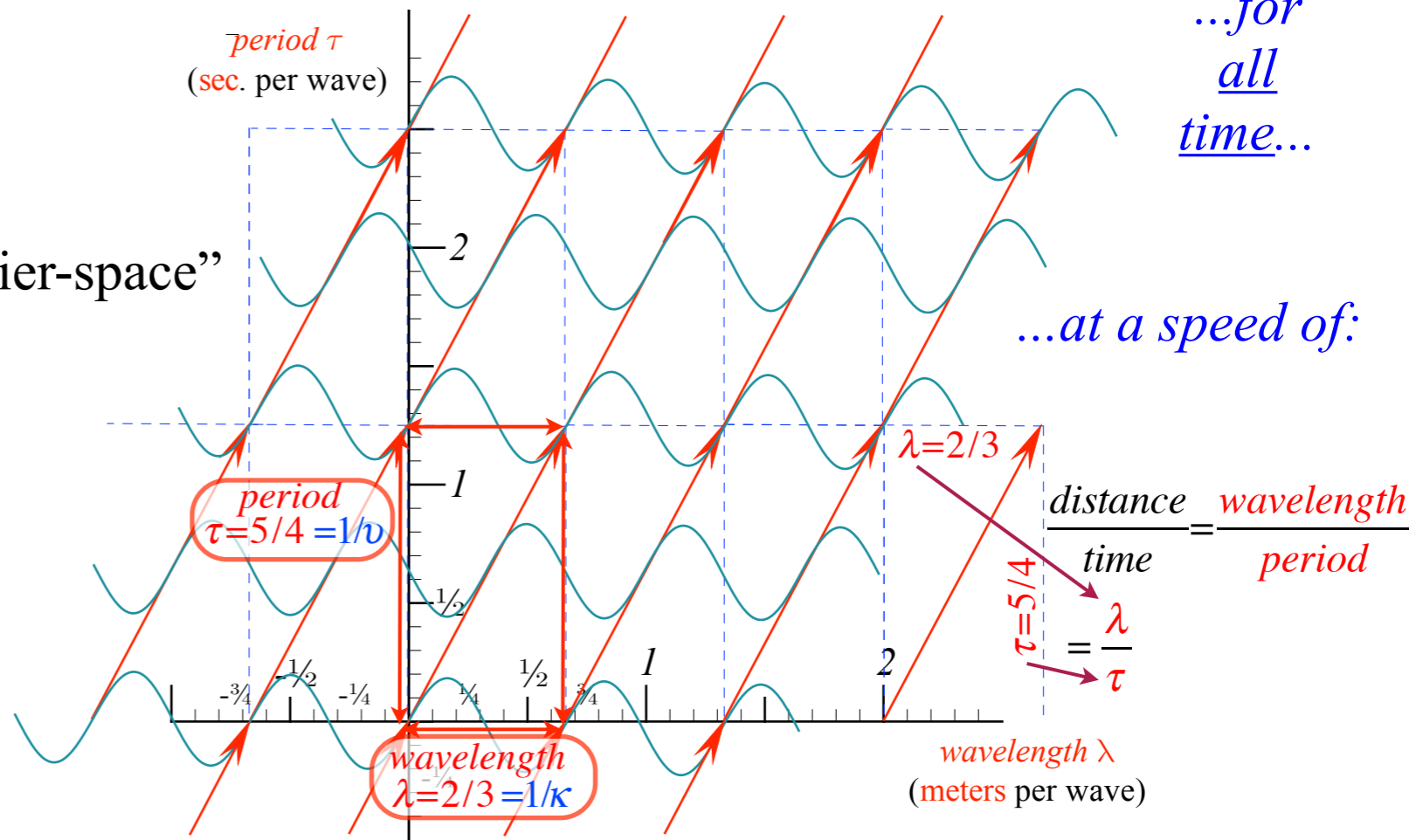
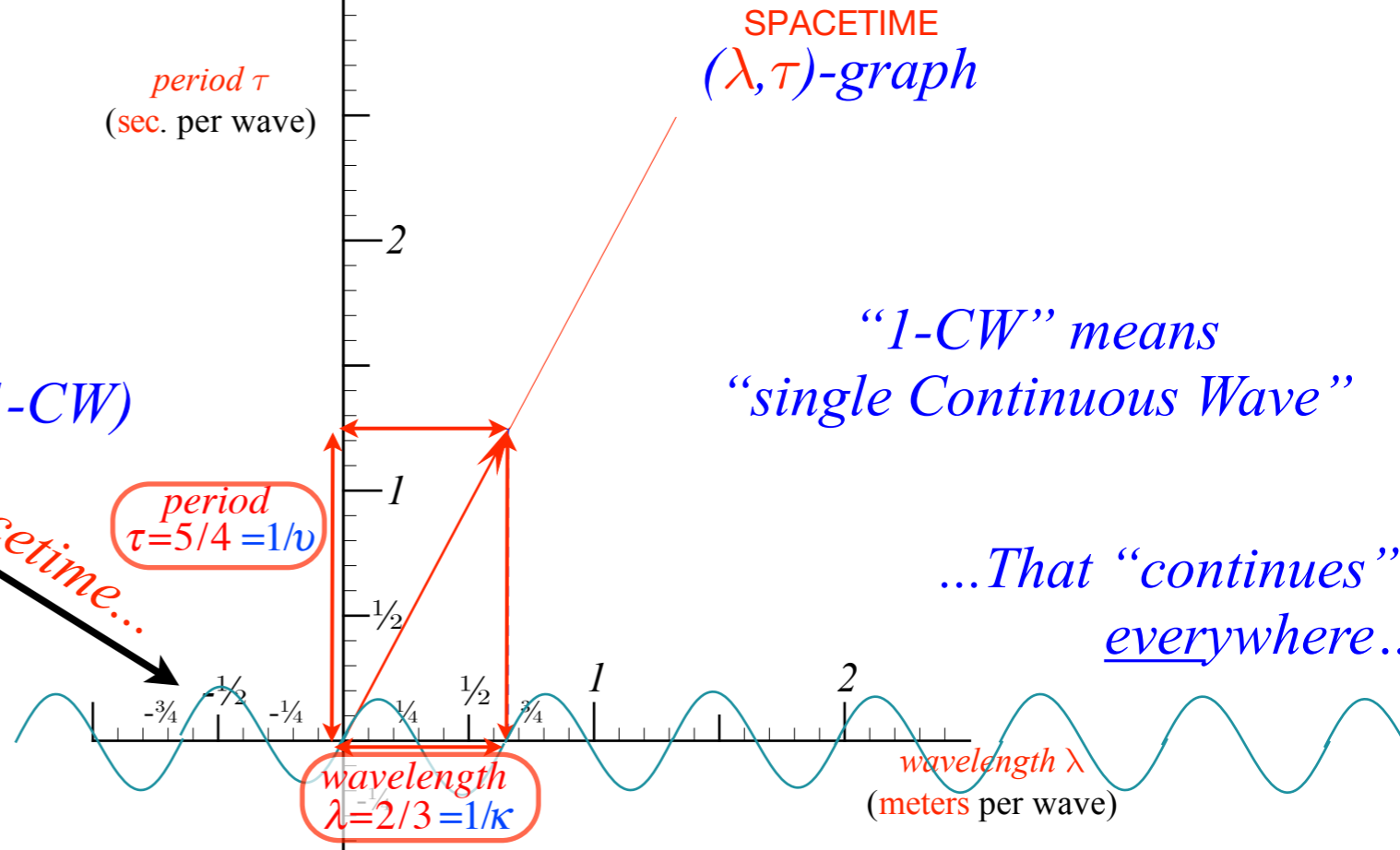
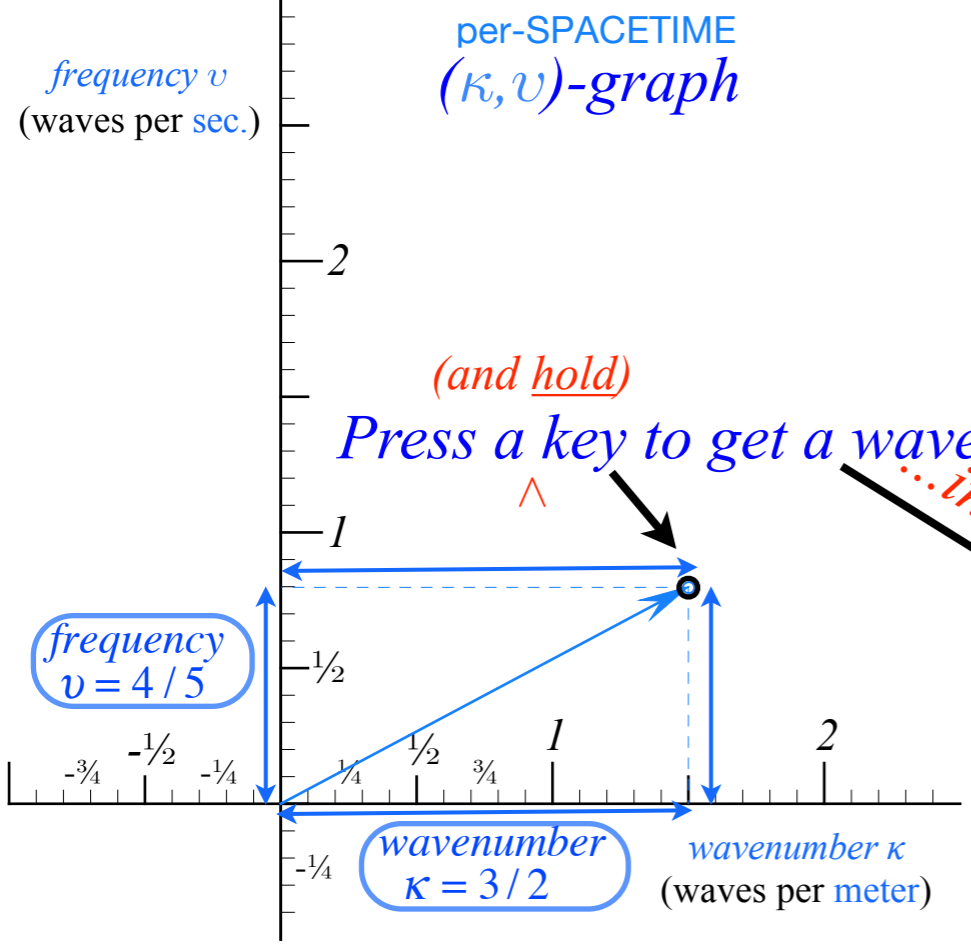
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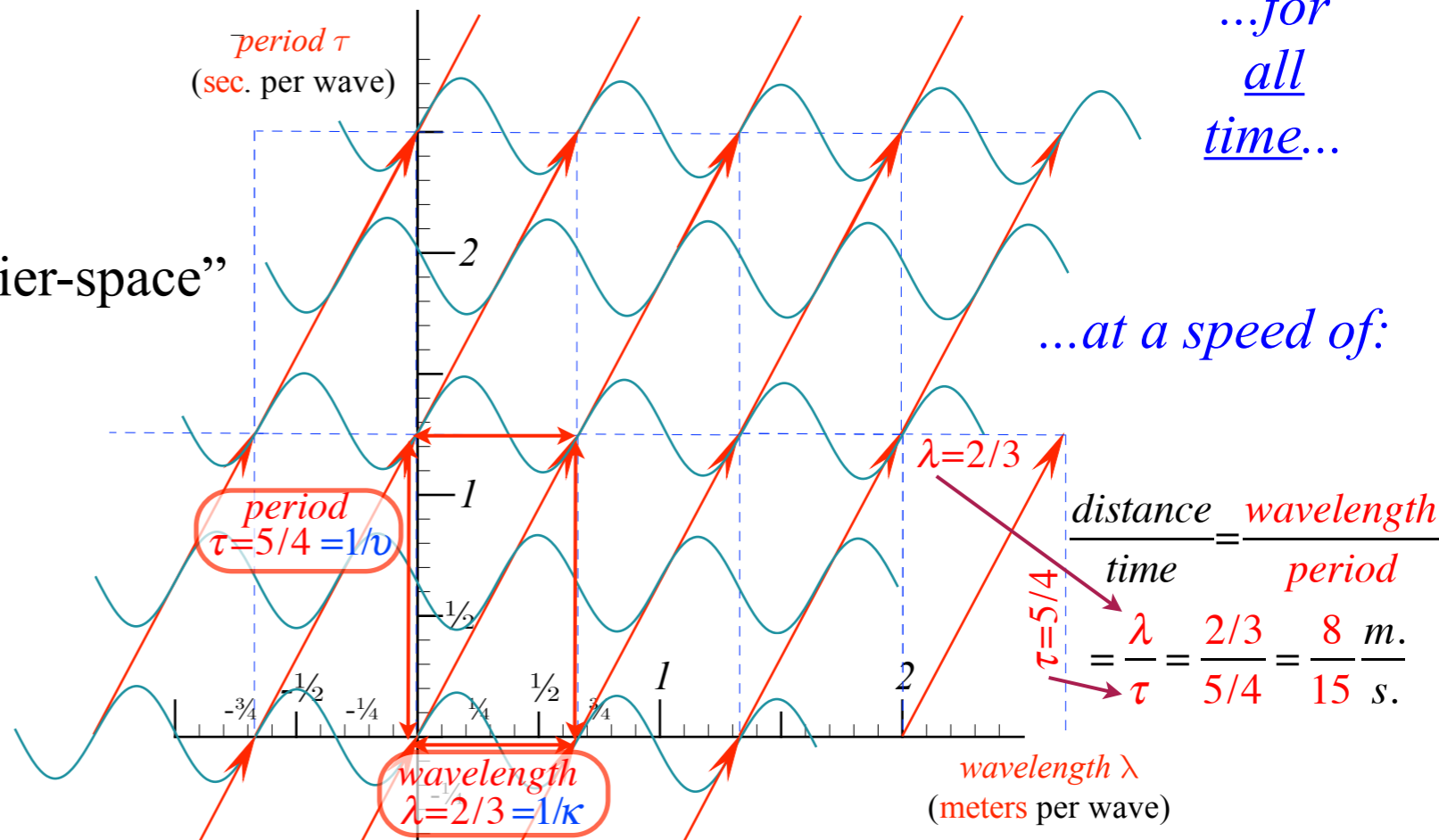
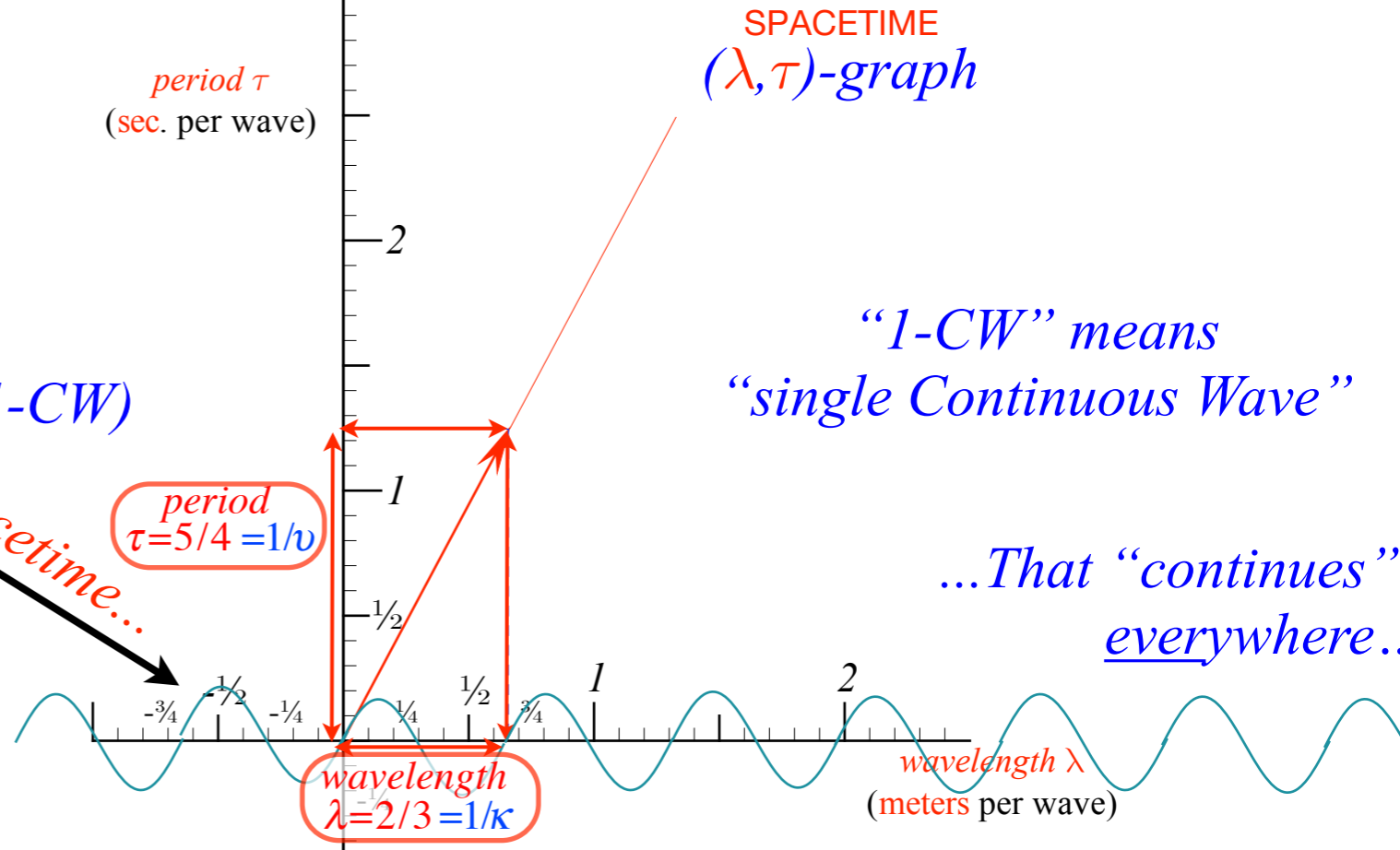
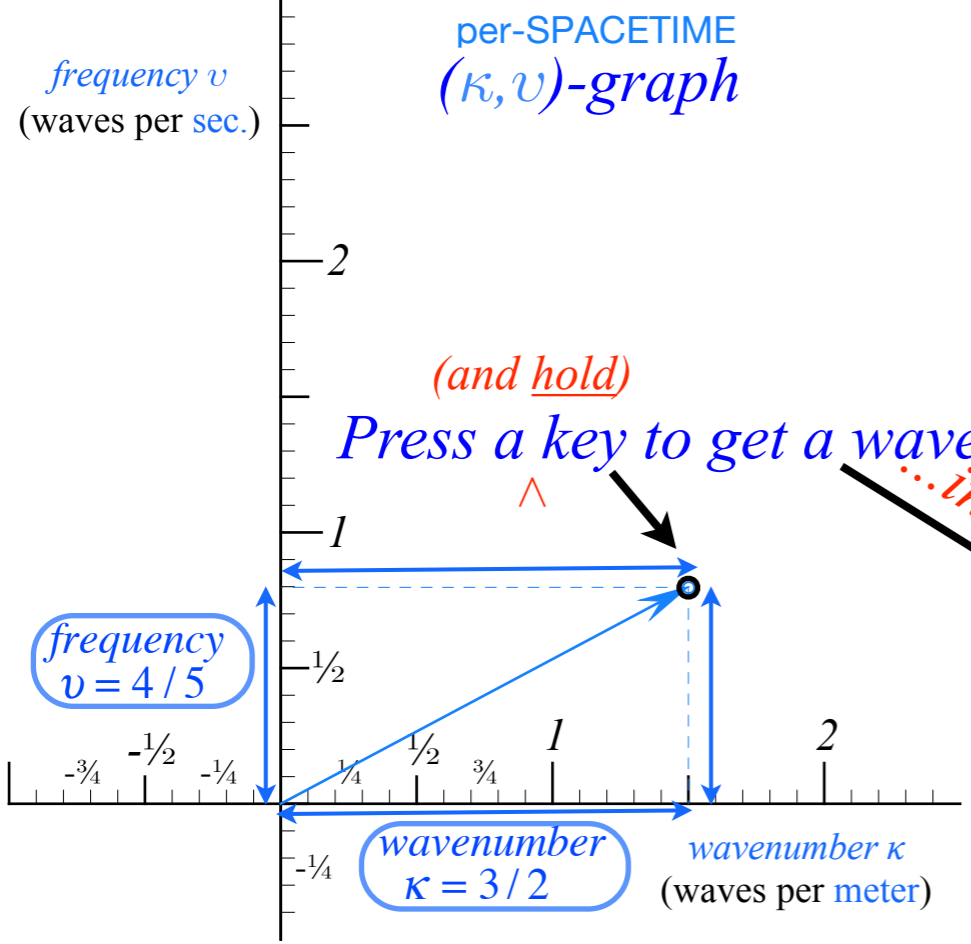
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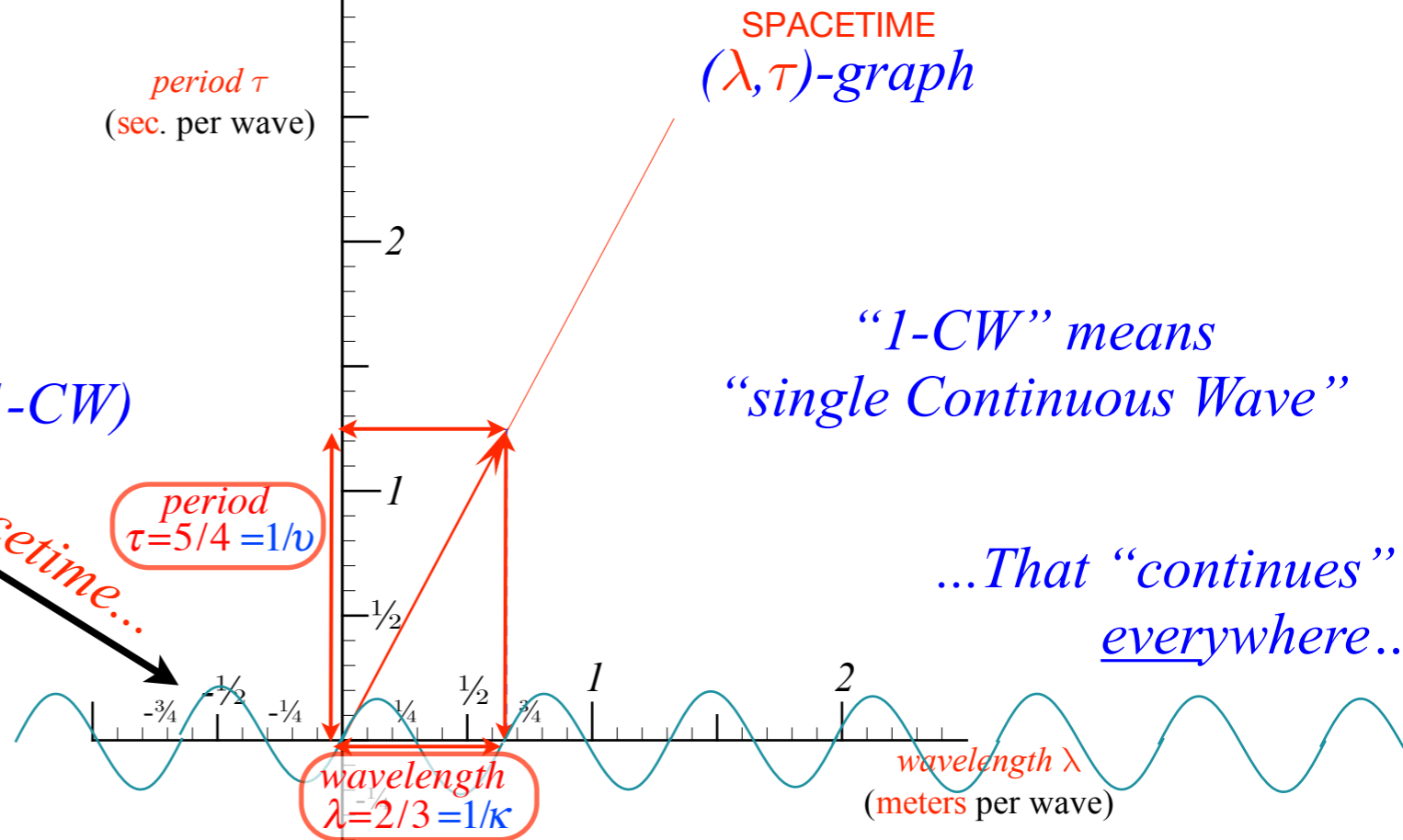
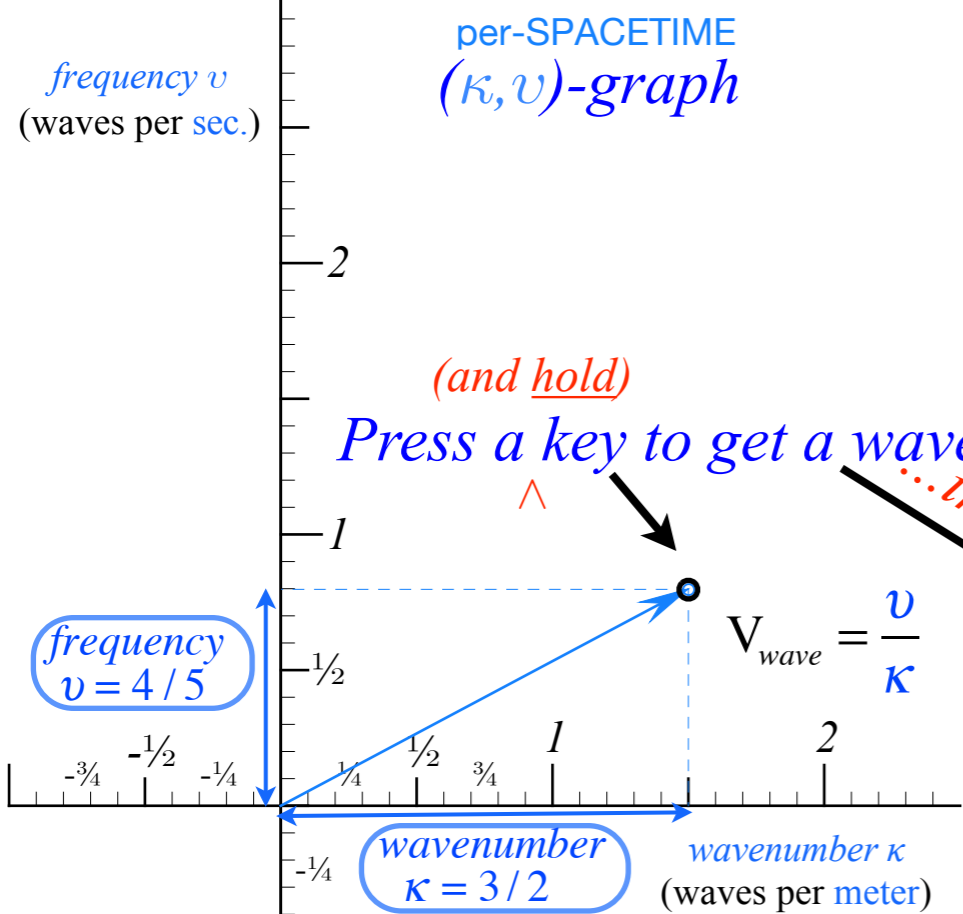


Jean-Baptiste Joseph Fourier
1768-1830

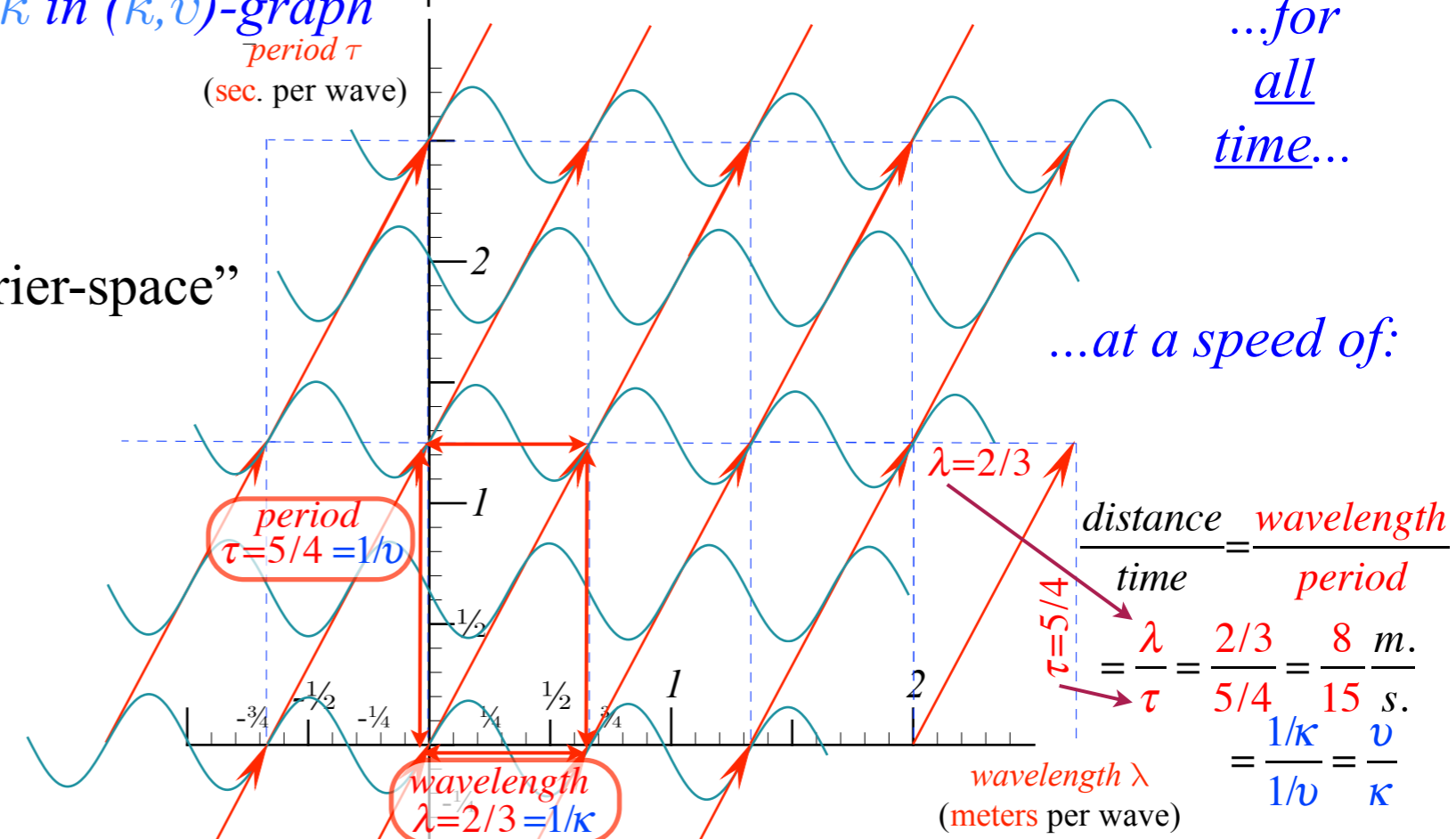
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wave-speed equals slope-to-vertical λ/τ in (λ, τ) -graph

Analyzing wave velocity by **per-space-per-time** and **space-time** graphs



wave-speed equals slope-to-horizontal ν/κ in (κ, ν) -graph



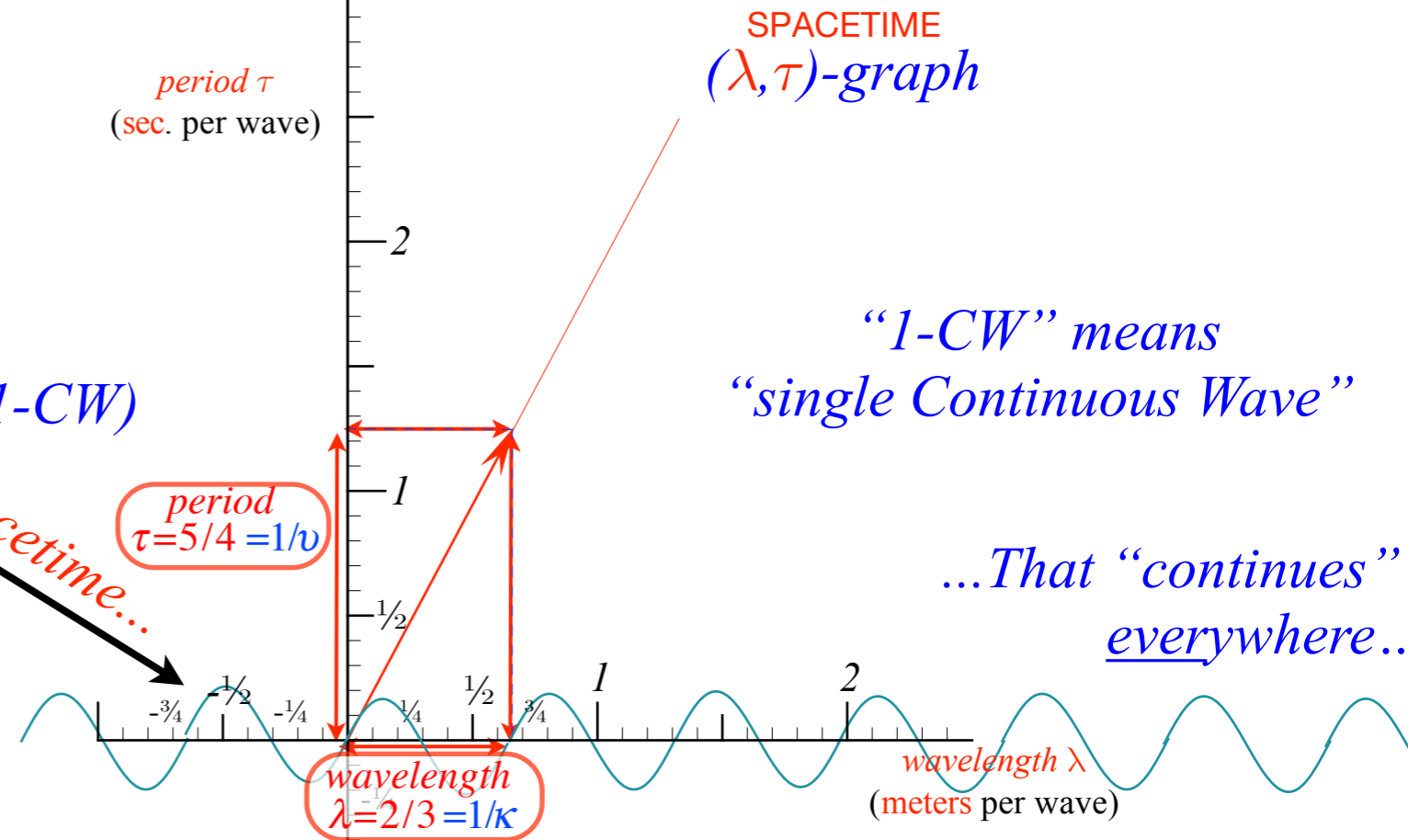
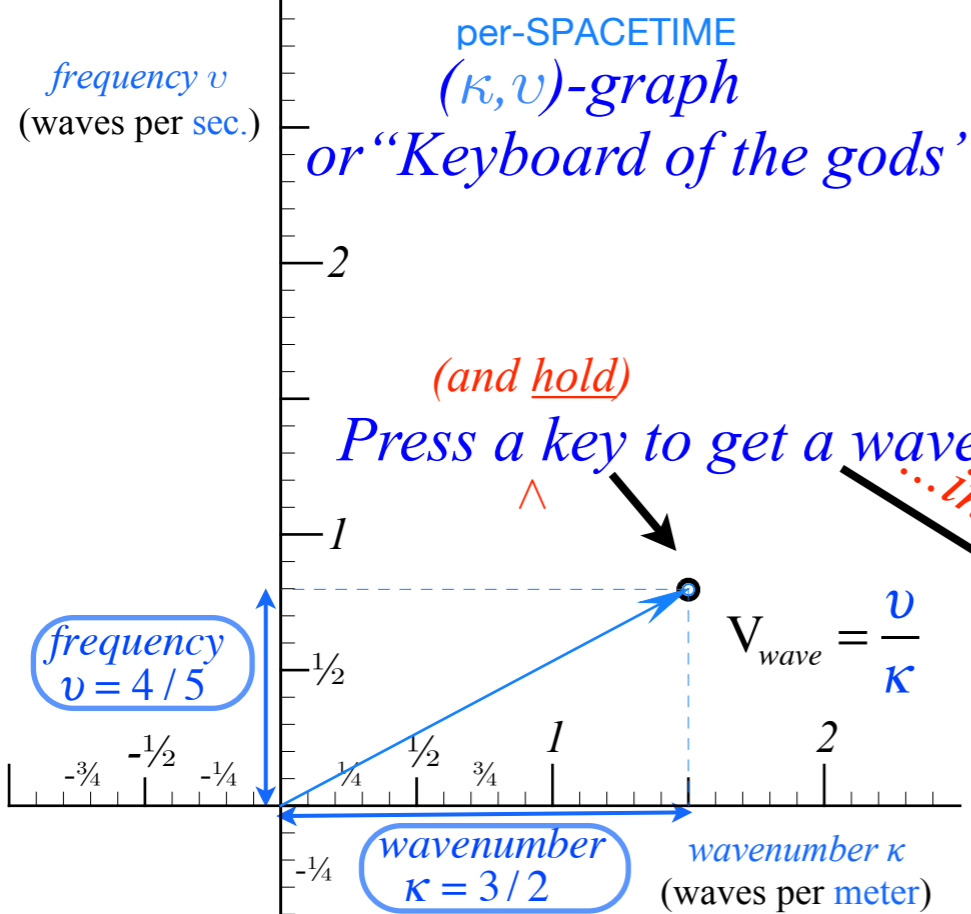
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wave-speed equals slope-to-horizontal ν/κ in (κ, ν) -graph

...for all time..

wave-velocity formulas

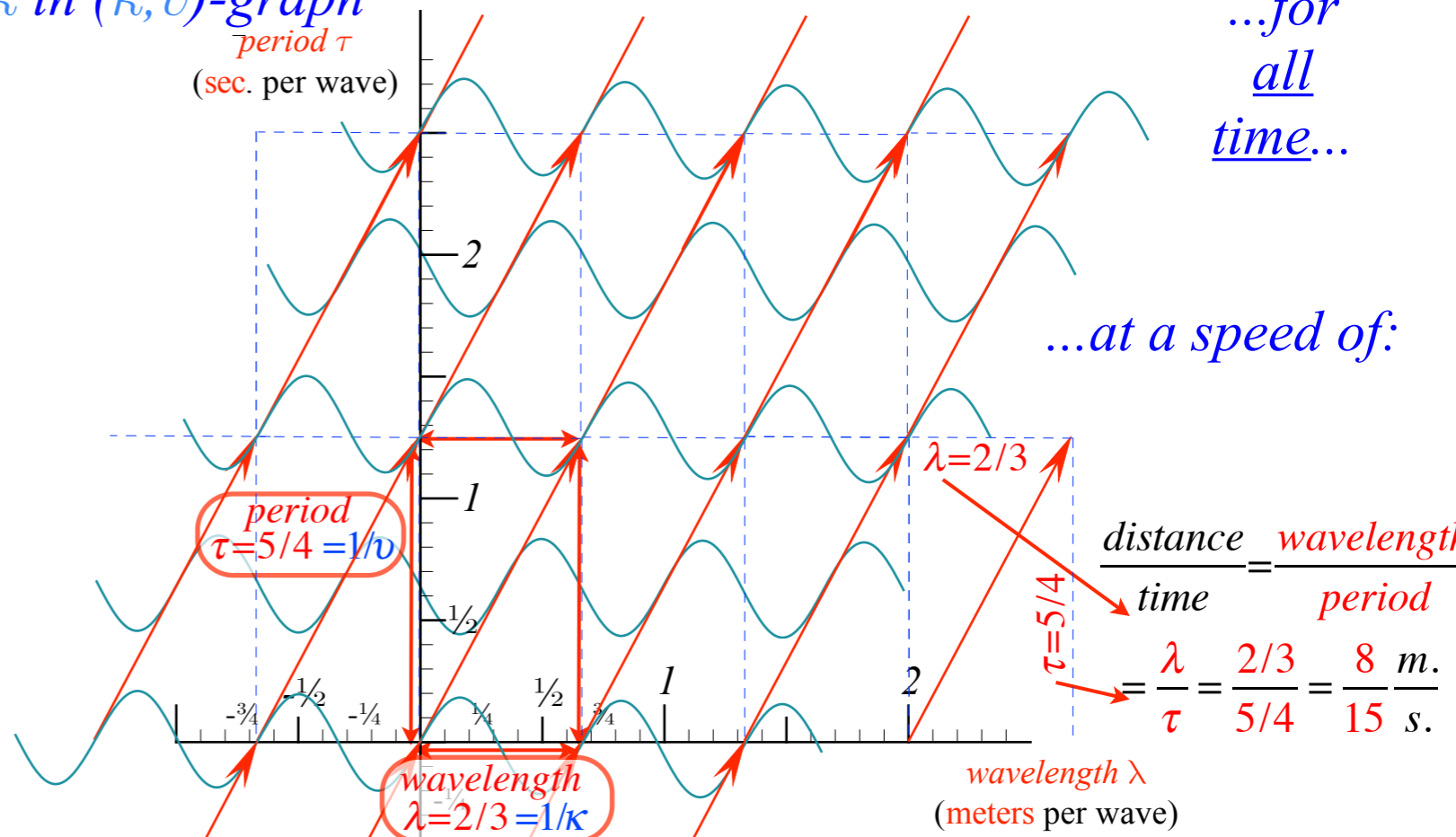
$$\frac{\text{distance}}{\text{time}} = \frac{\text{wavelength}}{\text{period}} = \frac{\text{frequency}}{\text{wavenumber}}$$

$$V_{\text{wave}} = \frac{\lambda}{\tau} = \frac{1/\kappa}{1/\nu} = \frac{\nu}{\kappa} = \frac{1/\tau}{1/\lambda}$$

$$= \frac{2/3}{5/4} = \frac{4/5}{3/2} = \frac{8 \text{ m.}}{15 \text{ s.}}$$

wave arithmetic is simpler to explain using fractions

•How to understand waves and "1st quantization"



wave-speed equals slope-to-vertical λ/τ in (λ, τ) -graph

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Introducing Doppler shifting and why c is so constant (and so slow)

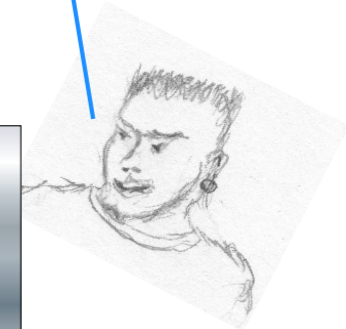
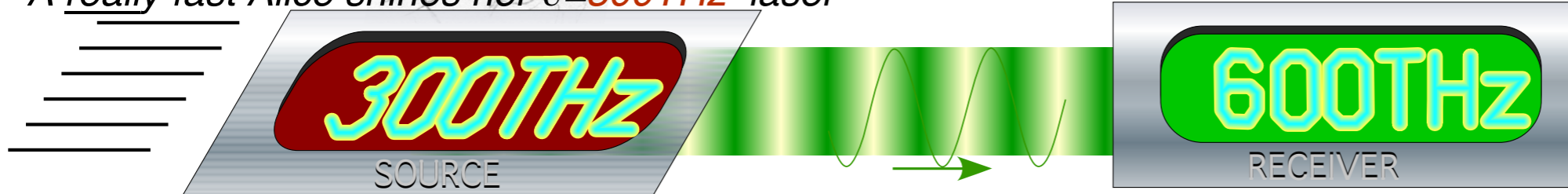
(a)



Bob: "Alice! My frequency meter reads $\nu=600\text{THz}$ for your laser beam."

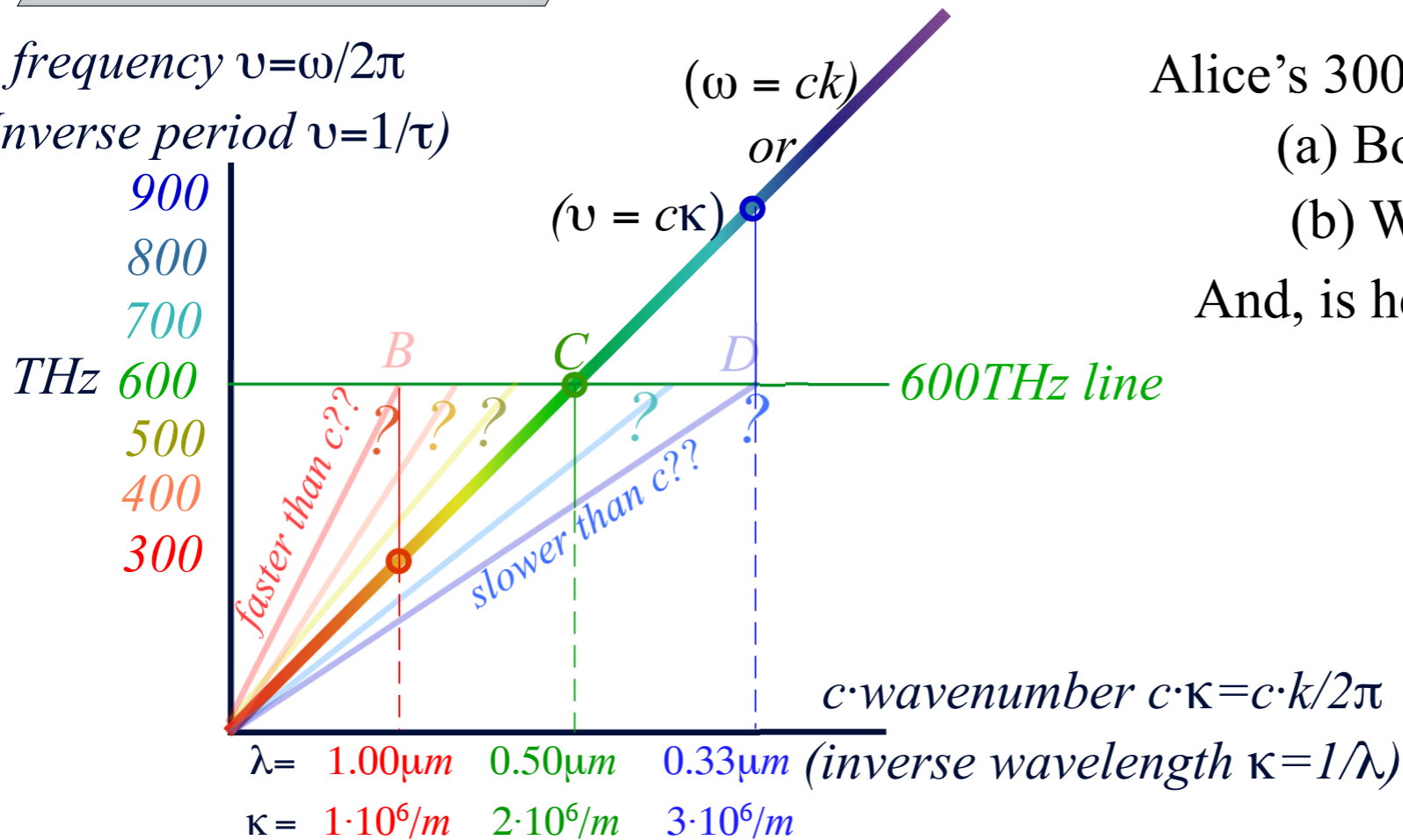
Alice: "Well, what is its wavelength λ , Bob!"

A really fast Alice shines her $\nu=300\text{THz}$ laser



(b)

frequency $\nu=\omega/2\pi$
(Inverse period $\nu=1/\tau$)



Alice's 300THz laser approaches Bob.

(a) Bob sees $\nu=600\text{THz}$.

(b) What $\lambda=1/\kappa$ does Bob measure?

And, is he seeing a 'phony' green?

Introducing Doppler shifting and why c is constant

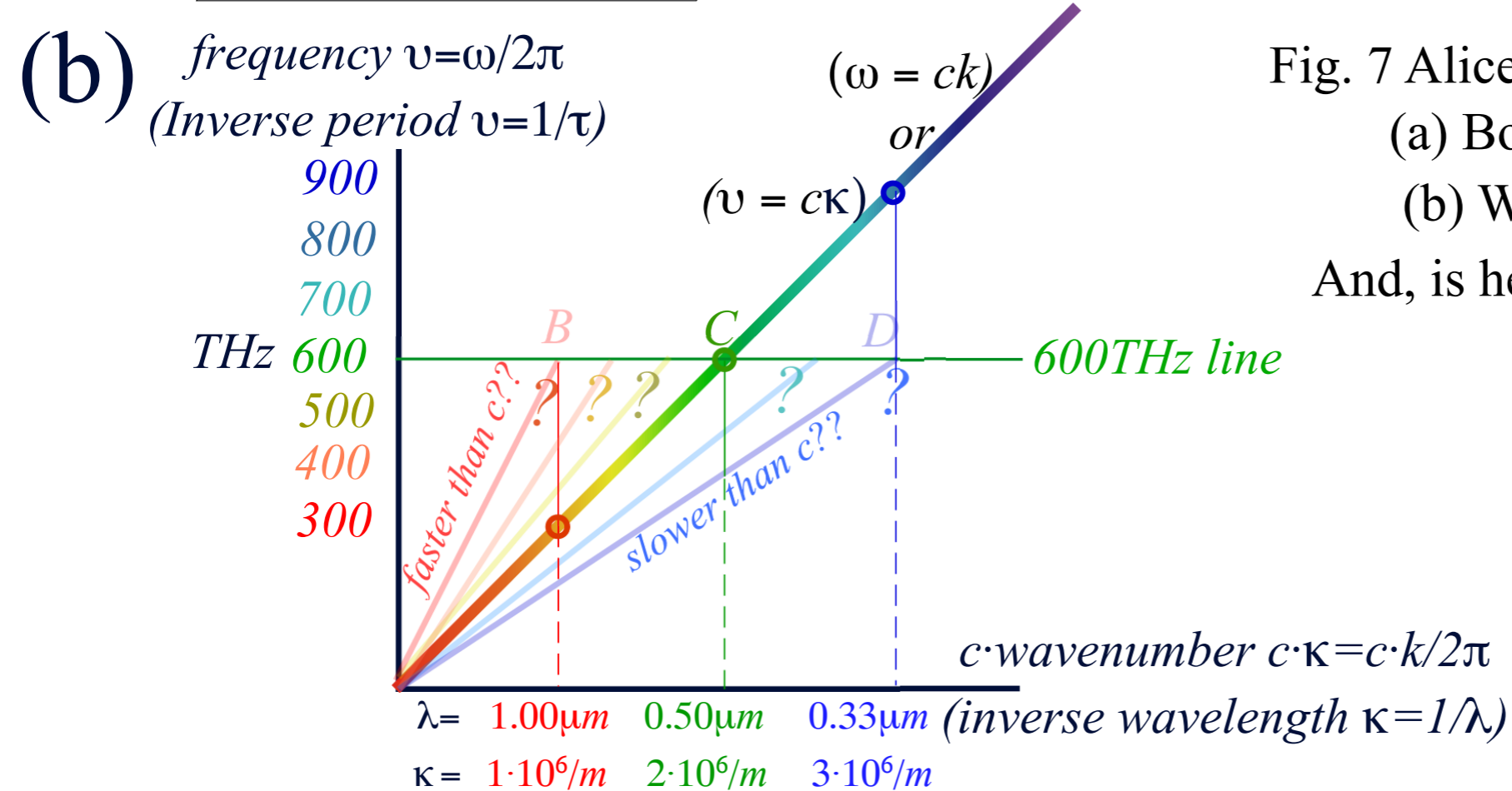
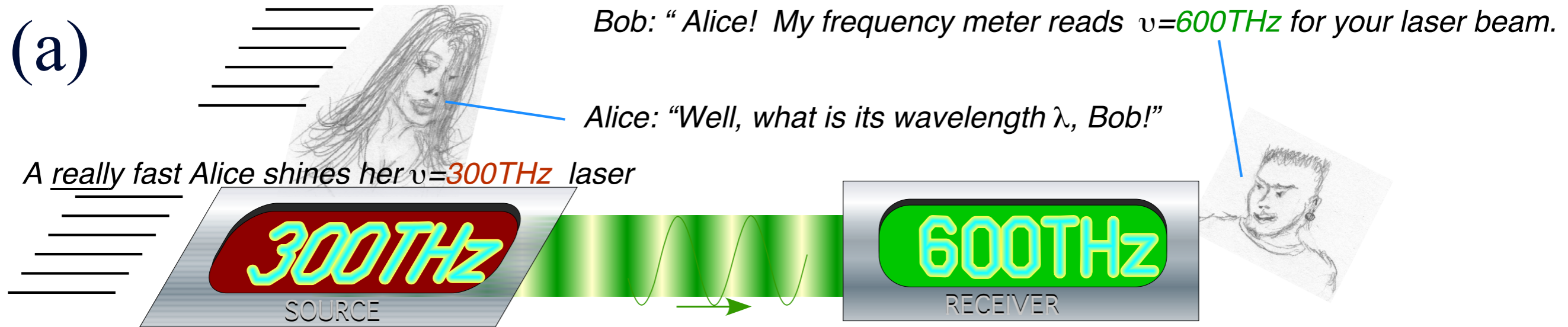


Fig. 7 Alice's 300THz laser approaches Bob.

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And, is he seeing a 'phony' green?

Years of spectroscopy rule out 'phony' 600THz blue-green that do not have wavelength $\lambda=0.5\text{micron}$.

The only choice is C.

Introducing Doppler shifting and why c is constant

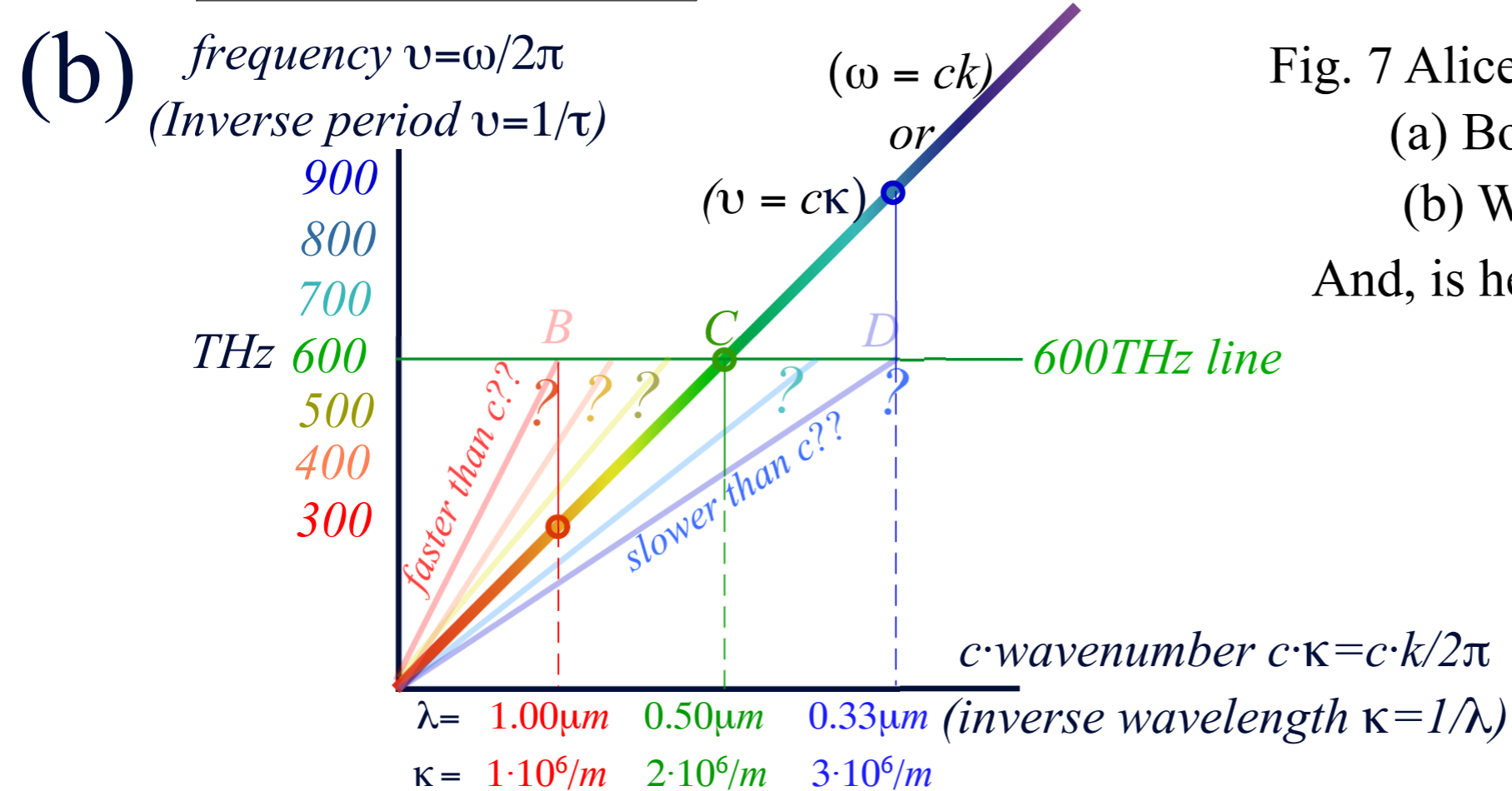
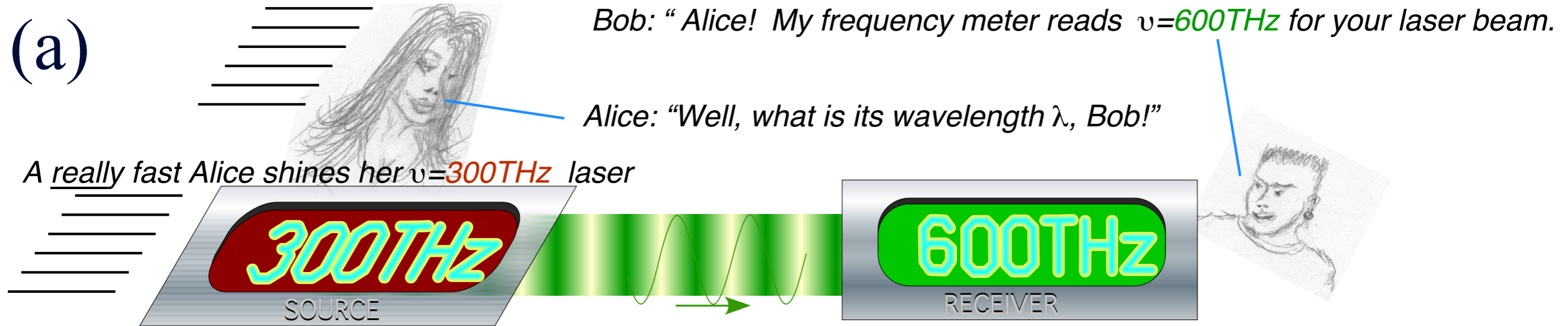


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Introducing Doppler shifting and why c is constant

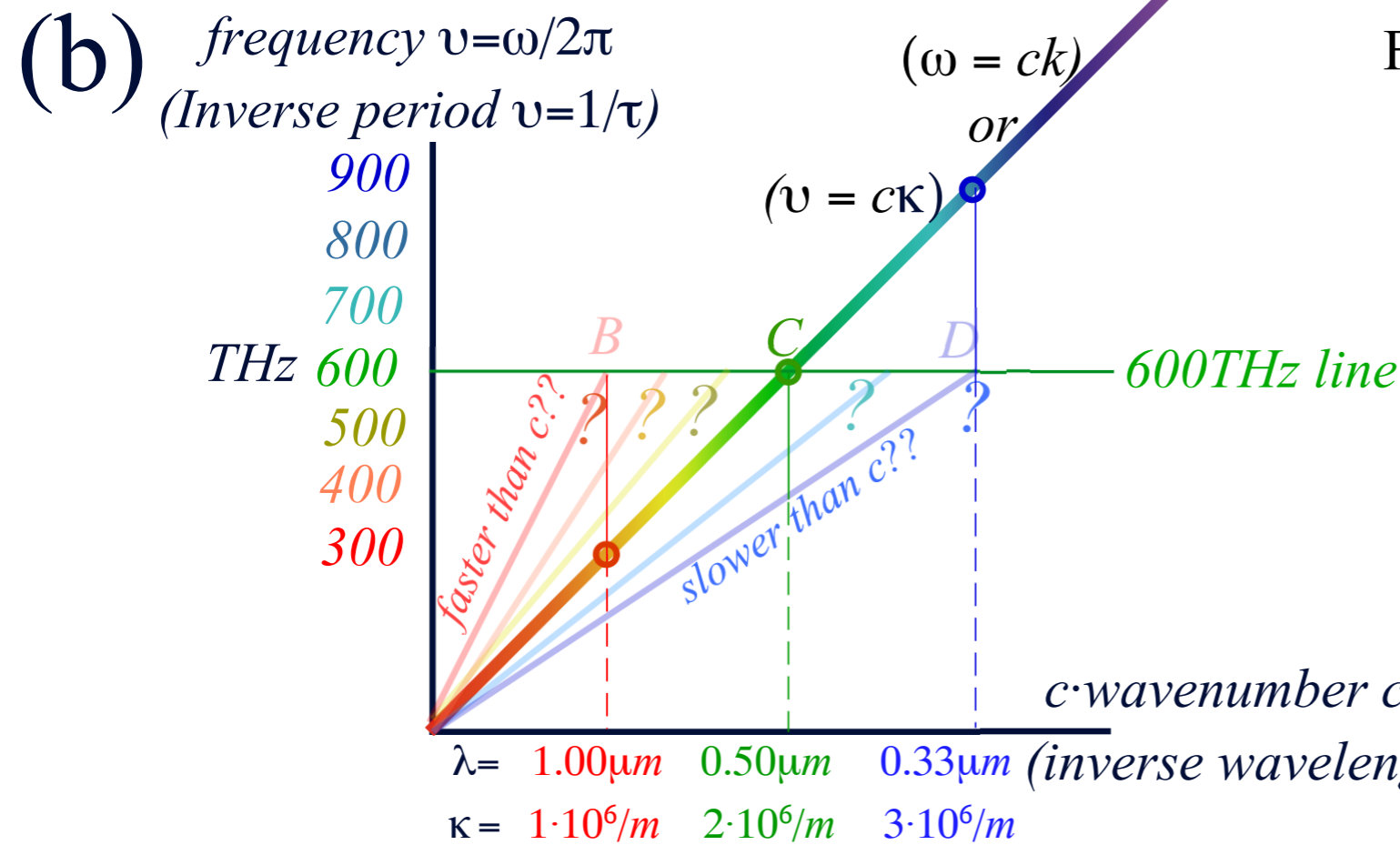
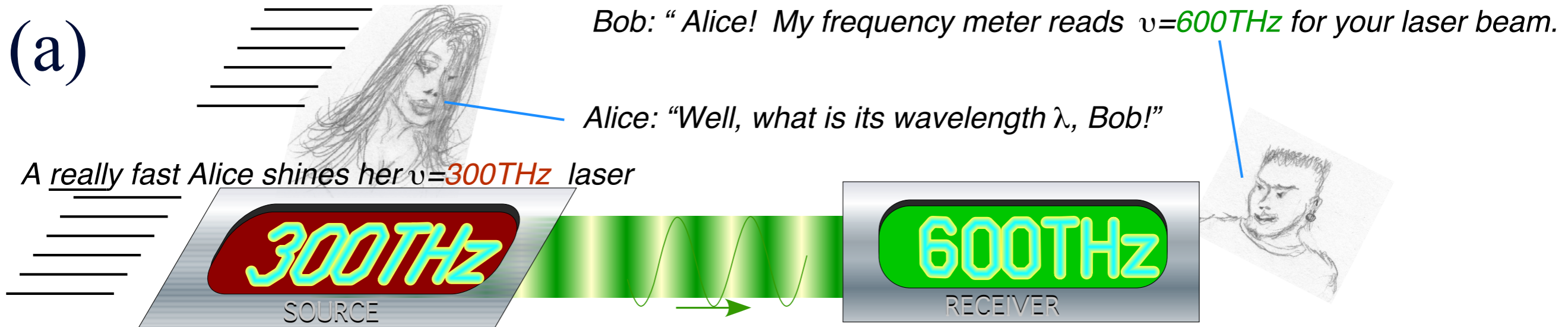


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Actually: $2.99792458 \cdot 10^8 \text{ m} \cdot \text{s}^{-1}$

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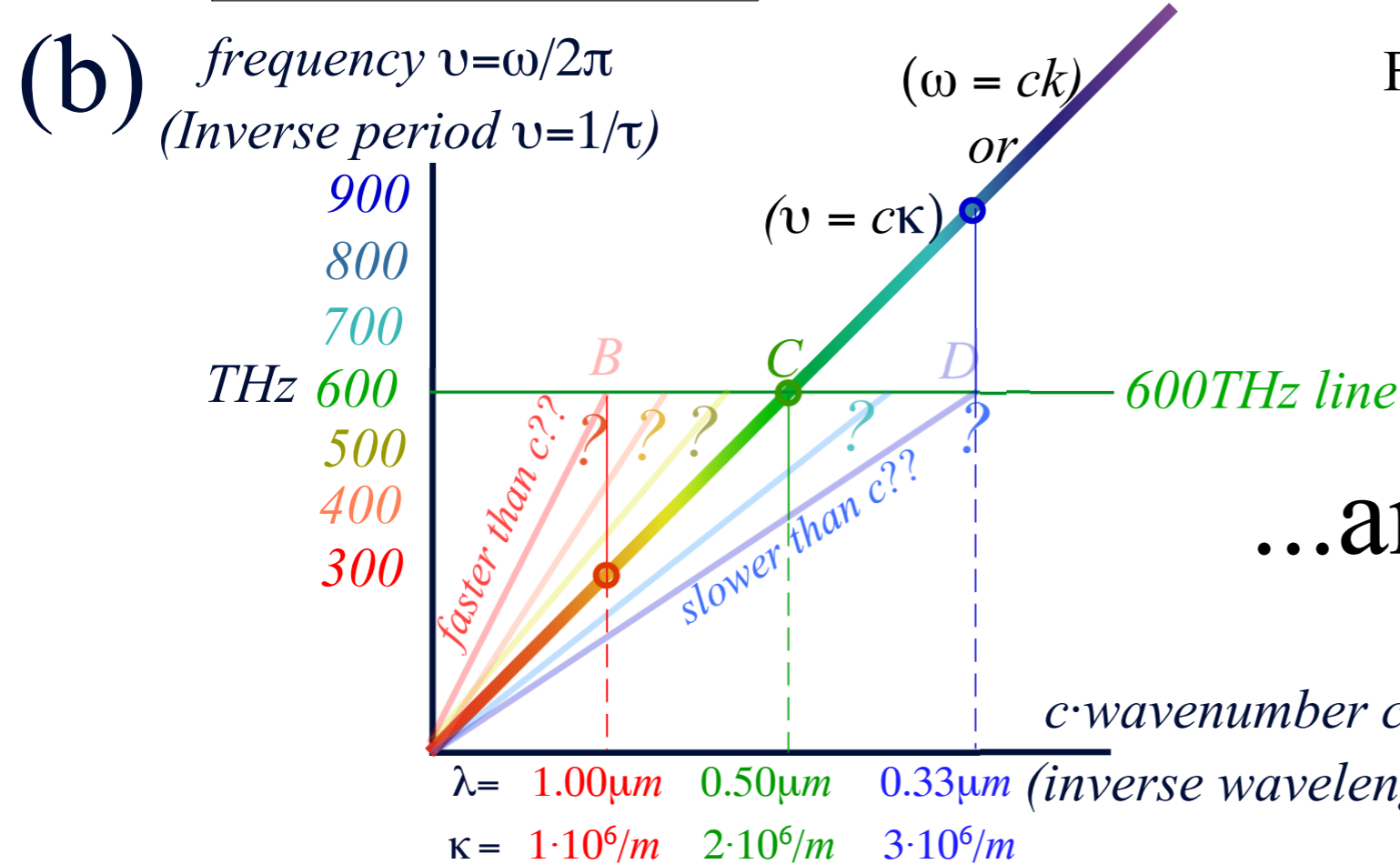
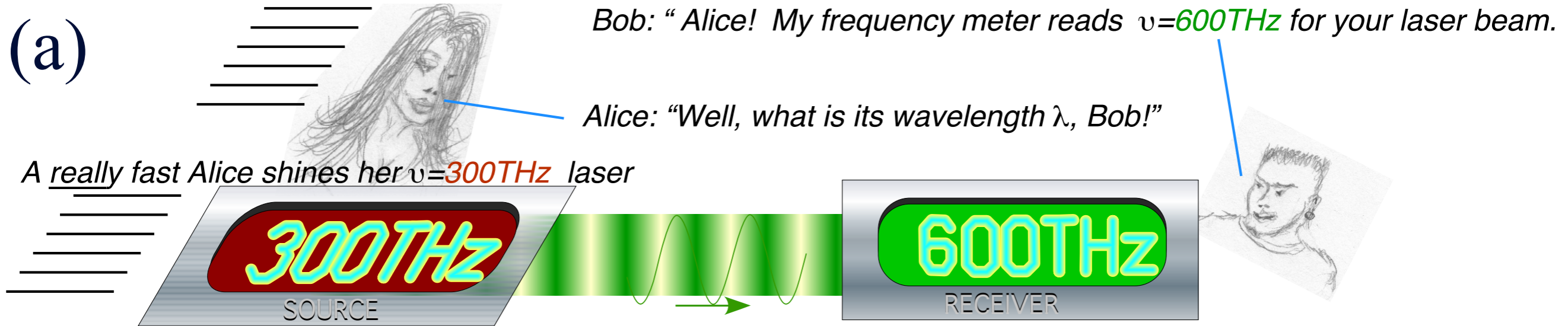


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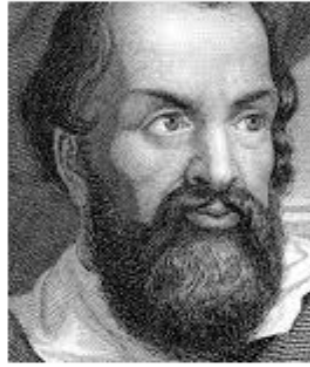
...and Dispersion-Free!

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Galileo Galilei



1564-1642

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Rapidity adds just like
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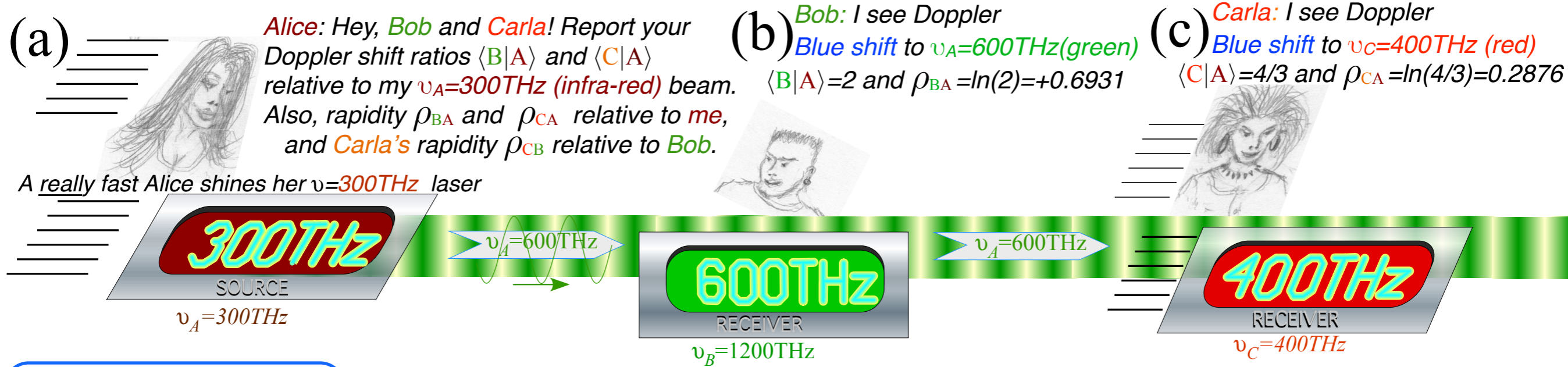
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Doppler ratio:

$$\langle R|S \rangle = \frac{\nu_{RECEIVER}}{\nu_{SOURCE}}$$

$$\rho_{RS} = \ln \langle R|S \rangle$$

or:

$$\langle R|S \rangle = e^{\rho_{RS}} = e^{-\rho_{SR}}$$

Definition of Rapidity ρ_{RS}

Bob-Alice Doppler ratio:

$$\langle B|A \rangle = \frac{\nu_B}{\nu_A} = \frac{600}{300} = \frac{2}{1}$$

Bob-Alice rapidity:

$$\rho_{BA} = \ln \langle B|A \rangle = \ln \frac{2}{1} = 0.6931$$

$$\rho_{AB} = \ln \langle A|B \rangle = \ln \frac{1}{2} = -0.6931 = -\rho_{BA}$$

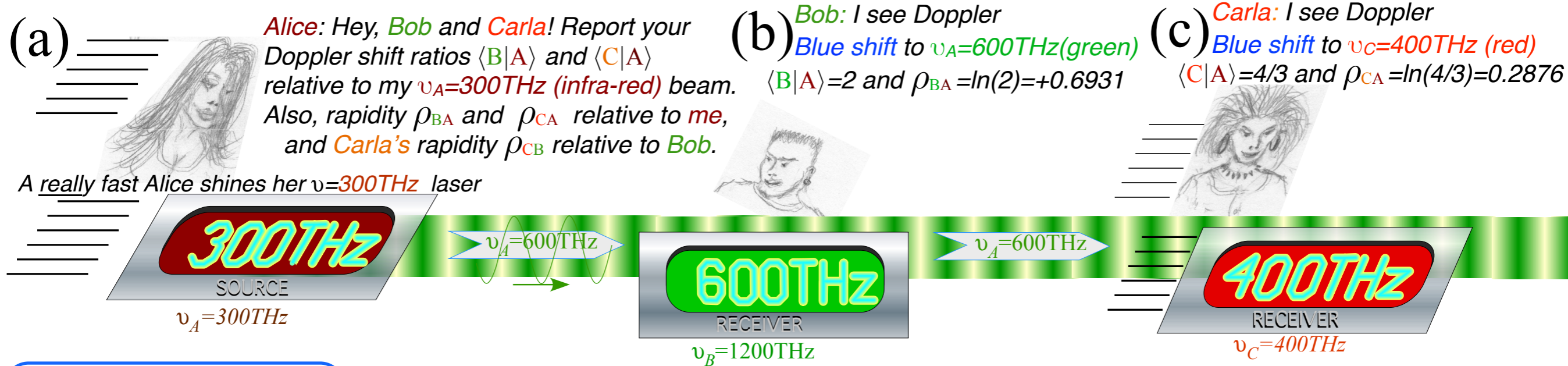
Carla-Alice Doppler ratio:

$$\langle C|A \rangle = \frac{\nu_C}{\nu_A} = \frac{400}{300} = \frac{4}{3}$$

Carla-Alice rapidity:

$$\rho_{CA} = \ln \langle C|A \rangle = \ln \frac{4}{3} = 0.2876$$

Introducing Doppler Arithmetic and rapidity ρ



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Carla-Bob Doppler ratio:

$$\langle C|B \rangle = \frac{\nu_C}{\nu_B} = \frac{\nu_C}{\nu_A} \frac{\nu_A}{\nu_B} = \langle C|A \rangle \langle A|B \rangle = \frac{4}{3} \frac{1}{2} = \frac{2}{3}$$

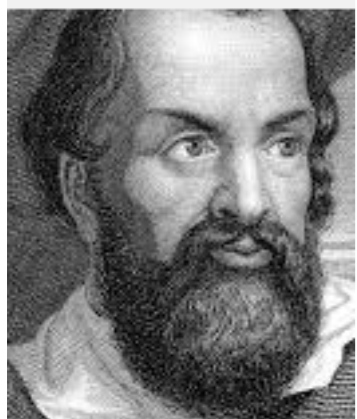
Carla-Bob rapidity:

$$e^{\rho_{CB}} = e^{\rho_{CA}} e^{\rho_{AB}} = e^{\rho_{CA} + \rho_{AB}}$$

$$\rho_{CB} = \rho_{CA} + \rho_{AB} = 0.2876 - 0.6931 = -0.4055$$

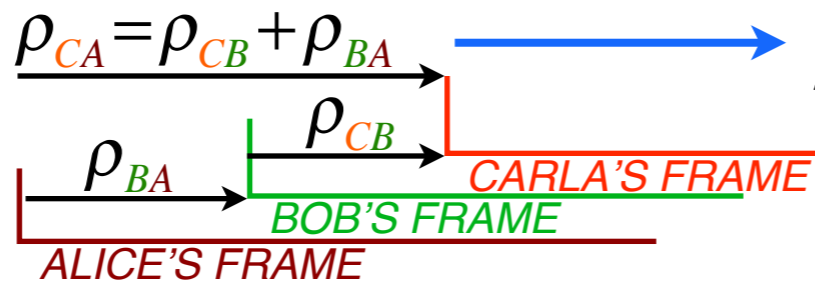
$$= \ln \frac{4}{3} + \ln \frac{1}{2} = \ln \frac{2}{3}$$

Galileo Galilei

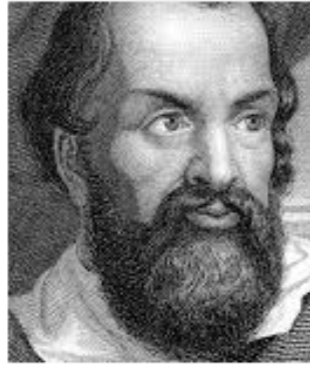


1564-1642

Galileo's Revenge (part 1)
Rapidity adds just like Galilean velocity



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1564-1642

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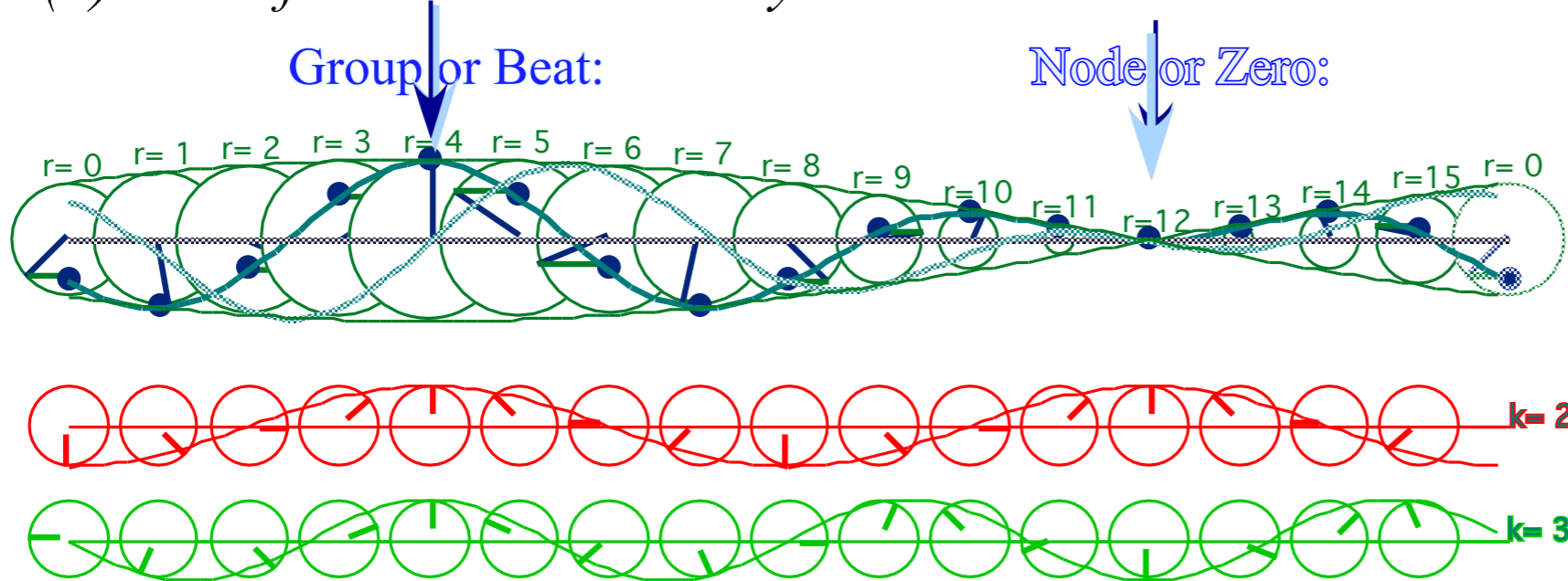
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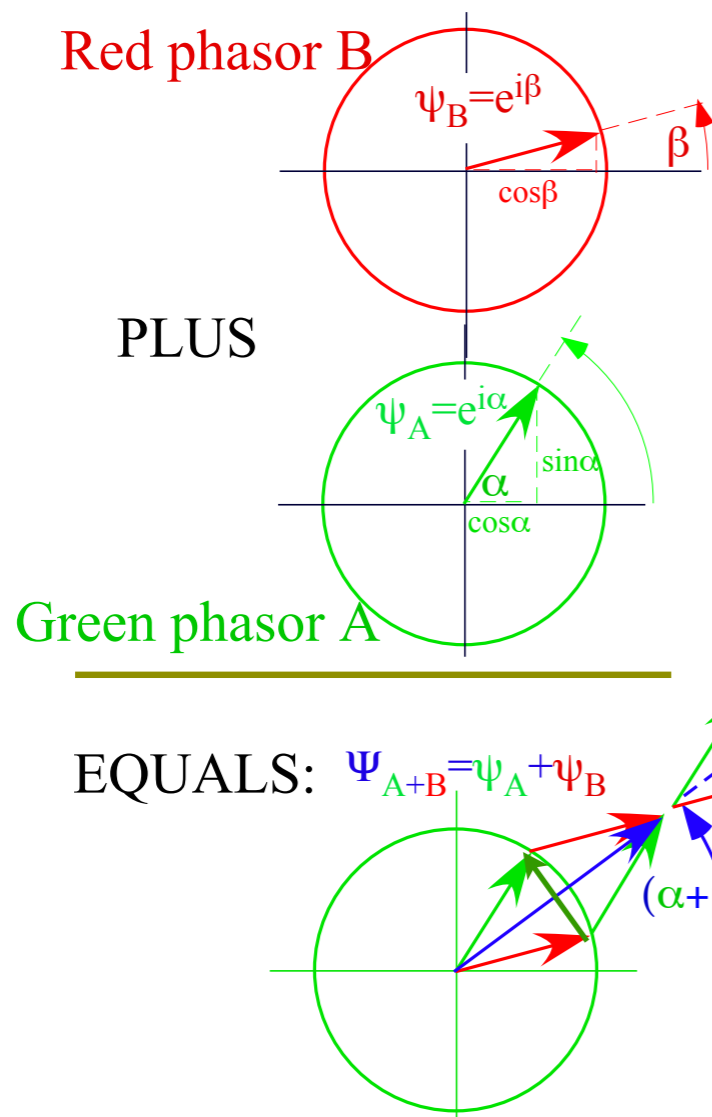
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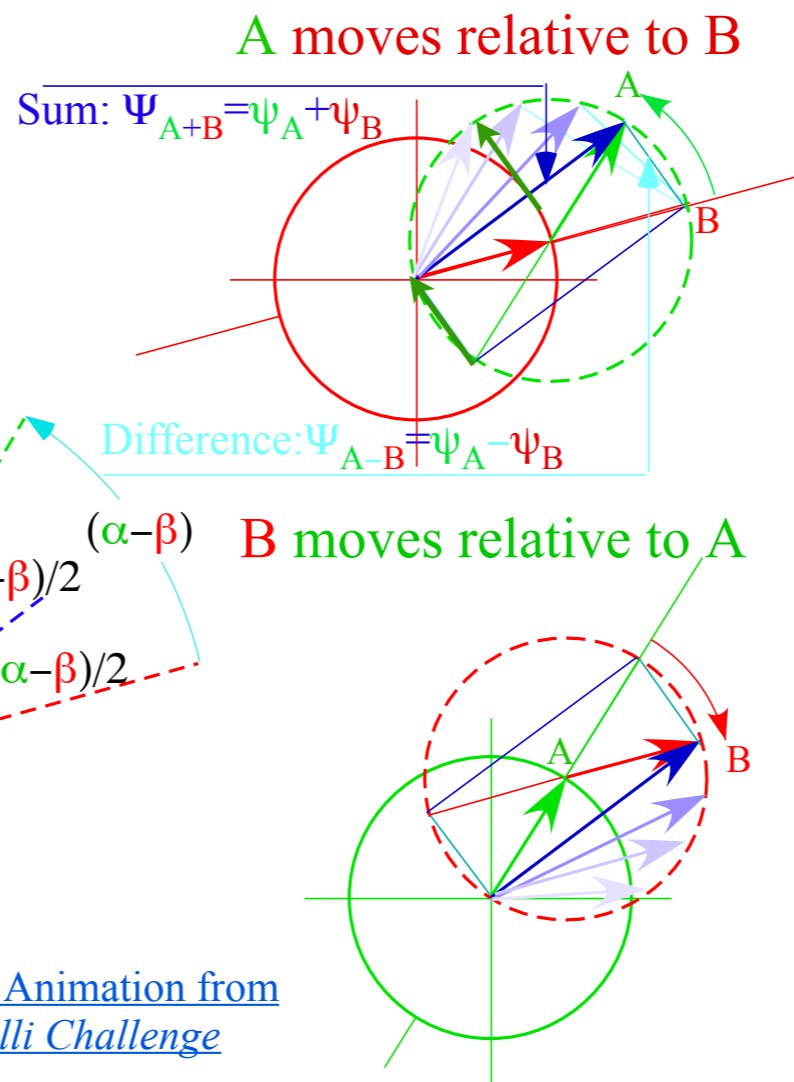
(a) Sum of Wave Phasor Array



(b) Typical Phasor Sum:

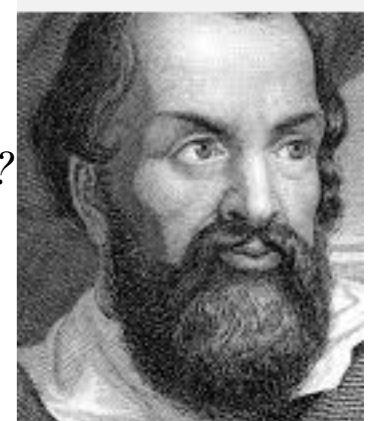


(c) Phasor-relative views



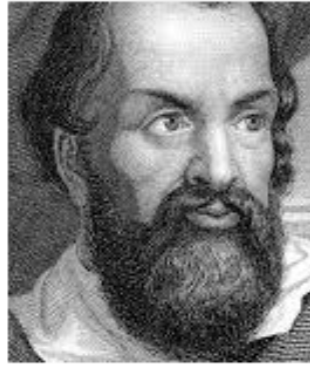
Geometry of the Half-sum Phase and Half-difference Group

Happy now?



Galileo's Revenge (part 2)
Phasor angular velocity adds just like Galilean velocity

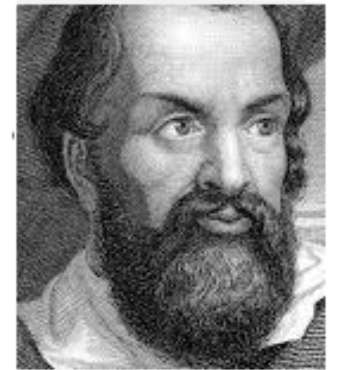
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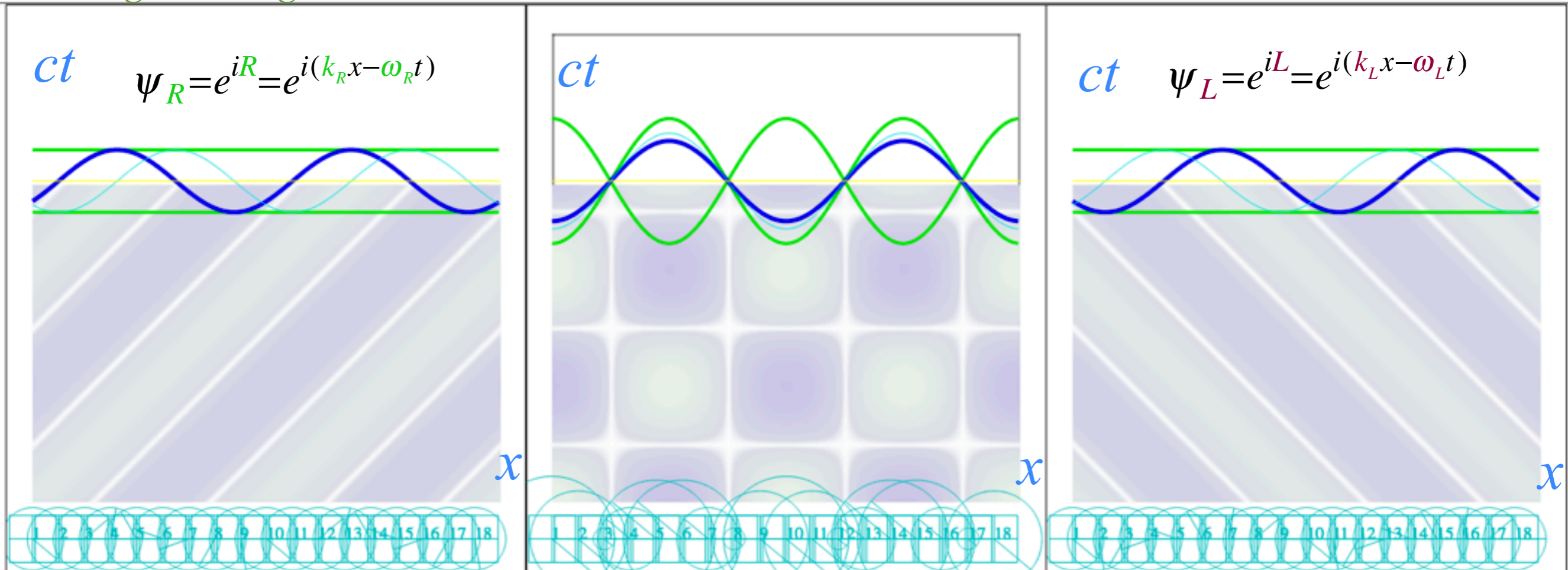
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right-moving CW laser

Colliding 2CW laser beams

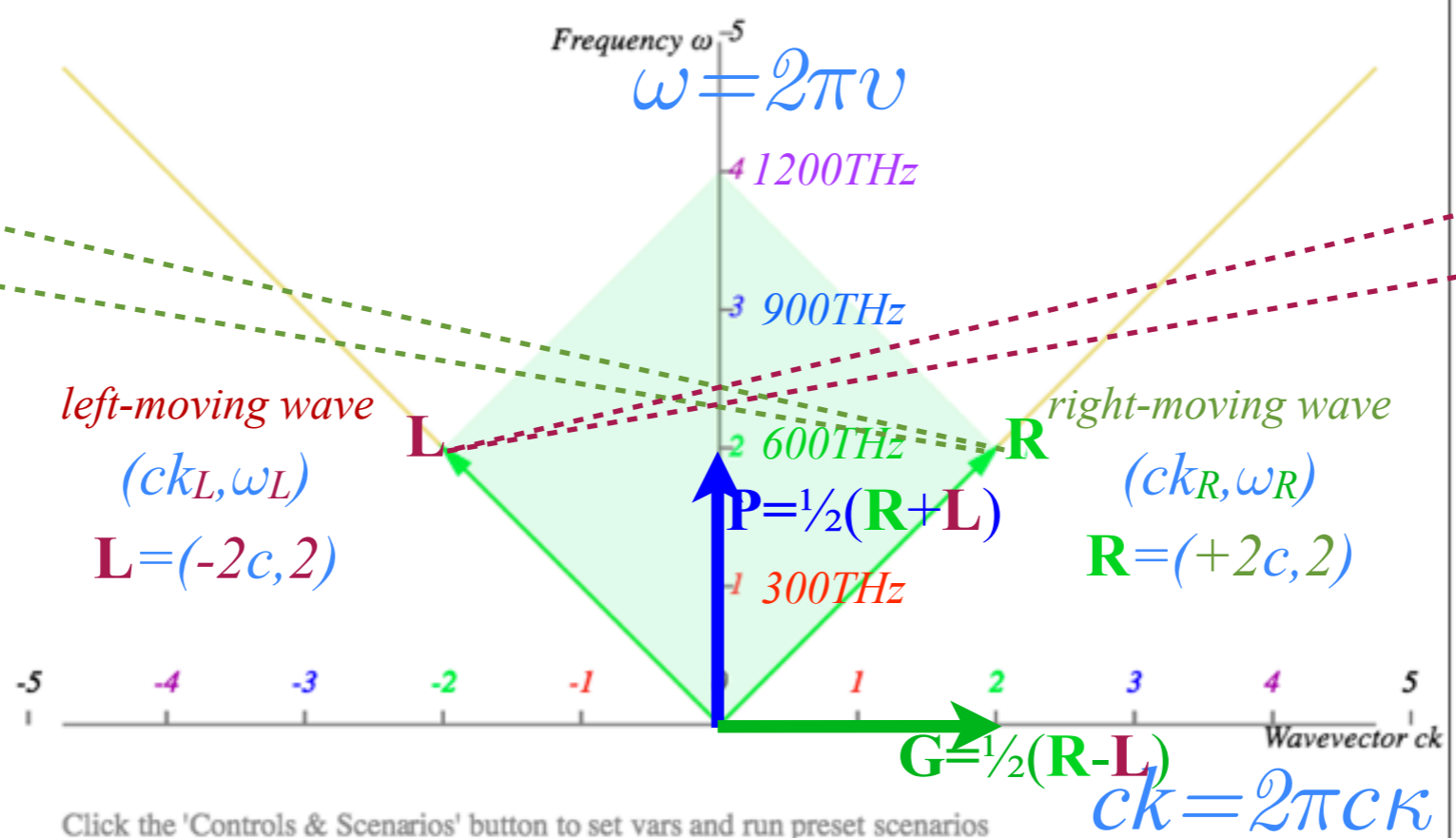
left-moving CW laser



right-moving wave
Spacetime (x, ct)

left-moving wave
Spacetime (x, ct)

Per-Spacetime
(ck, ω)



Click the 'Controls & Scenarios' button to set vars and run preset scenarios
Set the right & left-ward k values with clicks near the dispersion curve or ck axis.

BohrIt Web Simulation 2
CW ct vs x Plot (ck = ±2)

Parameters

BohrIt Panel 3x1: w/ k-Phasors

Configurations

Use Old ST Use Old Phasor Canvas

Canvas

Time Behavior Loop back to t=0

Retain Space-Time Plot

Align k-Phasors for Reset T=0

x-Phasor Locations Fixed at Bottom

Type of KE Photon: $\omega(k)-k$

Points per Well =

Space-Time Pixels per Phasor

Display

E Phasor Scale

X Phasor Scale

ψ Scale

Propagate Mouse Scale

Changes

$|\psi|$ Line Width

Re(ψ) Line Width

Im(ψ) Line Width

Phasor Line Width

Zero Tracer Line Width

Trace Group Zeros Trace Phase Zeros

Extra Coordinate Grid Axes w/ P & G Vectors

Background ST Plot Re(ψ)

Zero enhancement Threshold =

Crest-Trough distinction term =

Group & Phase Vectors Both

Right & Left **K** Vectors Both

Shaded Regions Show Both

Axis Titles - + Vertical: +

Axis Labels - Show Vertical: Show

Colors

Color Scheme Journal Color

Global alpha =

Space-Time background alpha =

Peak: Hue= Val=

Trough: Hue= Val=

Zero: Hue= Val=

Persistent Parameters

Default Space-Time Granularity

Best for tweaking the responsiveness, also vary/set persistent space-time granularity below

Scenarios

	Basic CW +1 >	
<	CW Light ± 1	+1..
-1..		>
<	PW Lite ± 1	+1..
-1..		>
	CW Light ± 2	
<	PW Lite ± 2	+2..
-2..		>
	CW Light -1 \diamond +4	

-4..		>
	CW Light -2 \diamond +8	
	PW Lit3 -2 \diamond +8	

Matter Wave: Bohr-Schrödinger Approximation

Bohr-Schrödinger {Quadratic dispersion}

CW
k=+1,+2
CW
k=+2,+3
CW k=-1,+2

RelaWavity Scenarios

Dispersion Plot (300 THz Scale)

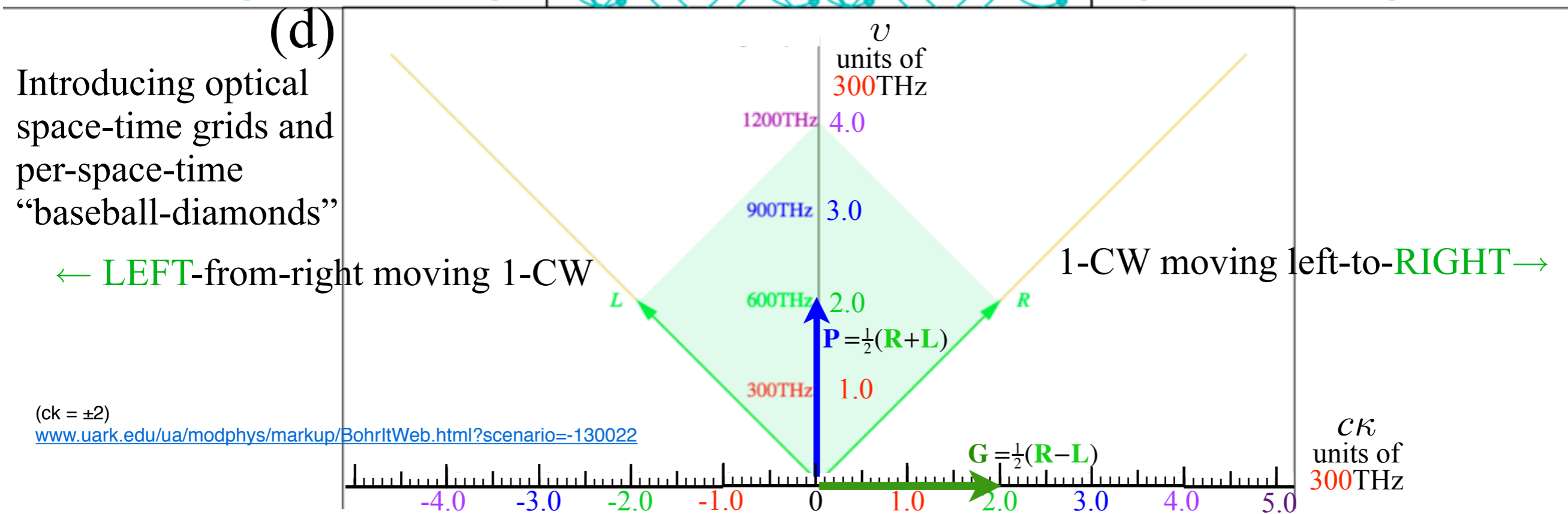
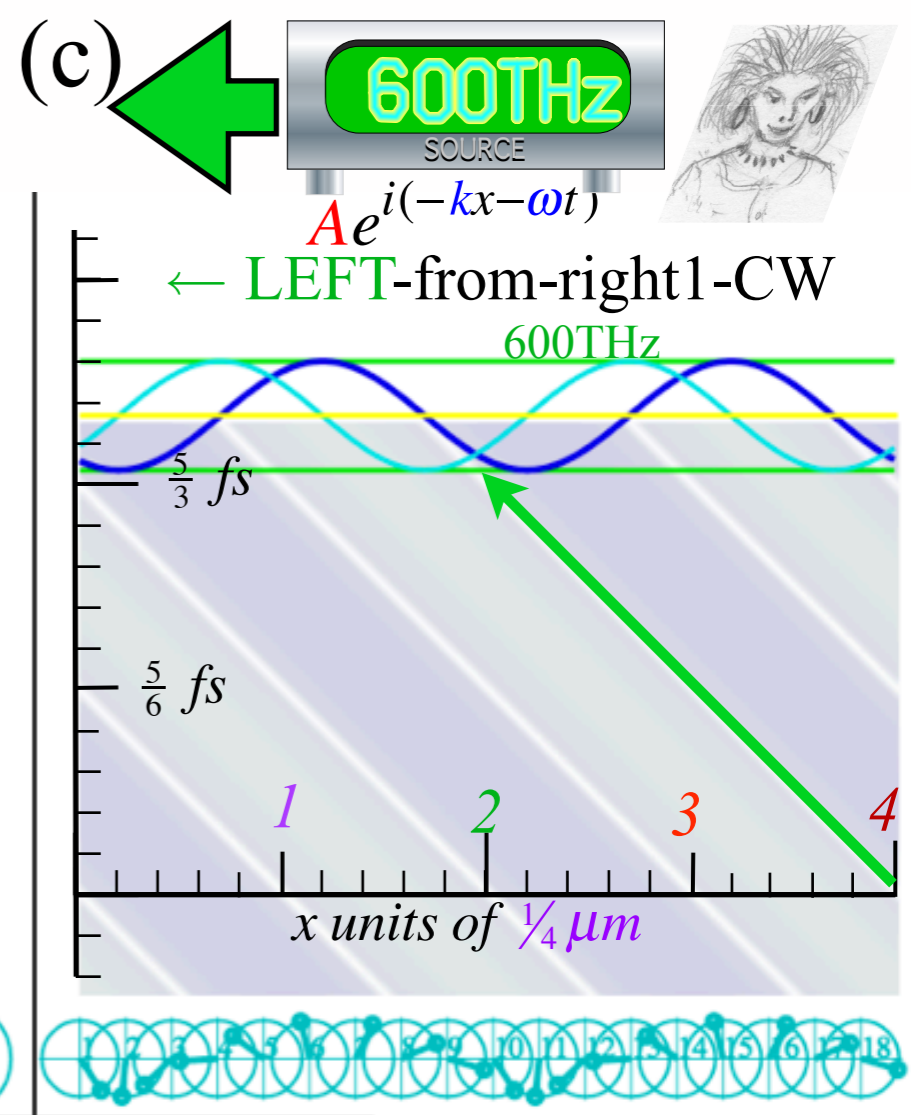
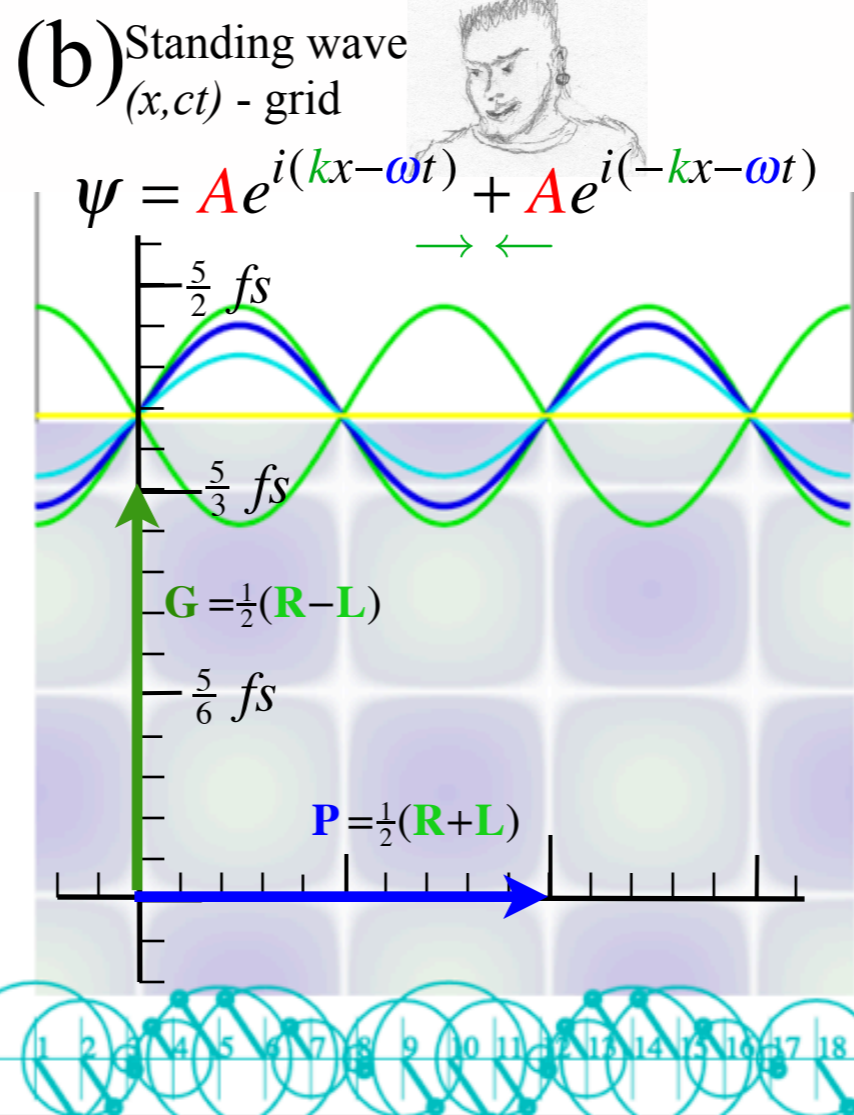
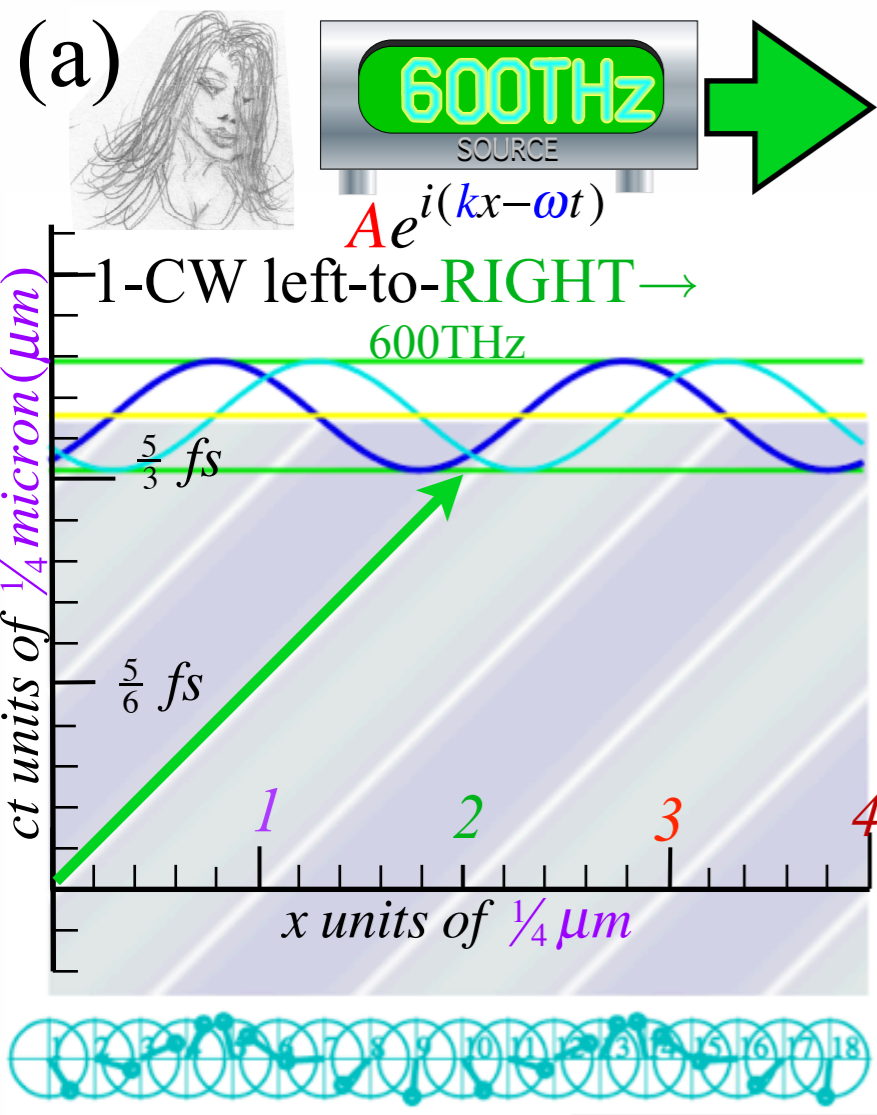
Make these two lower for finer/narrower zero lines. Note: they do vary with above settings

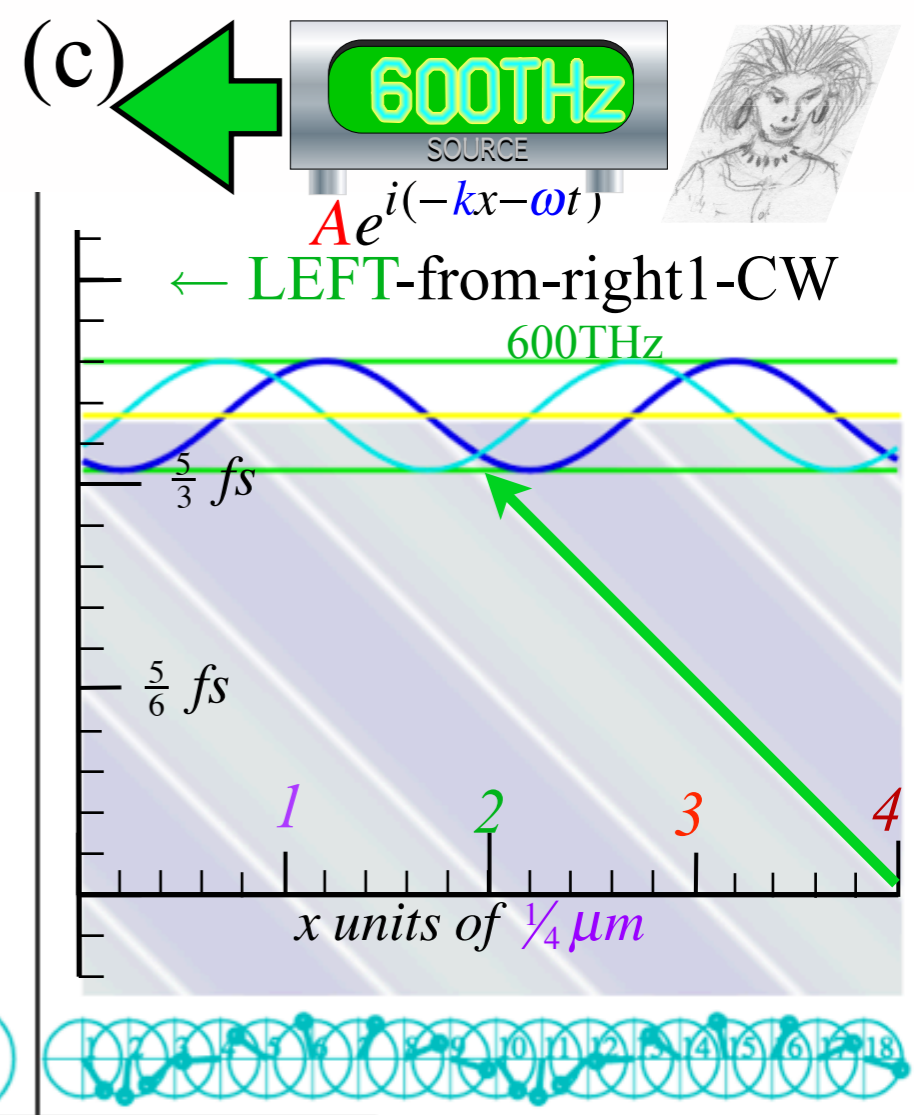
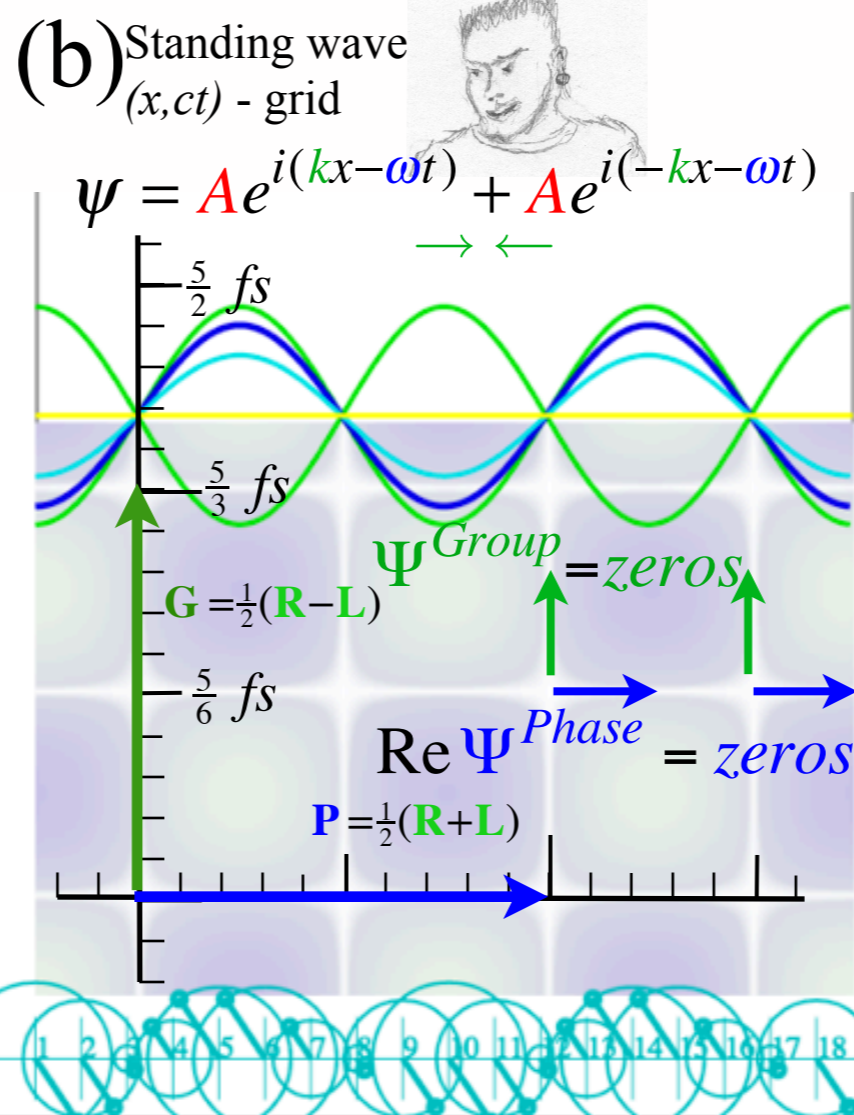
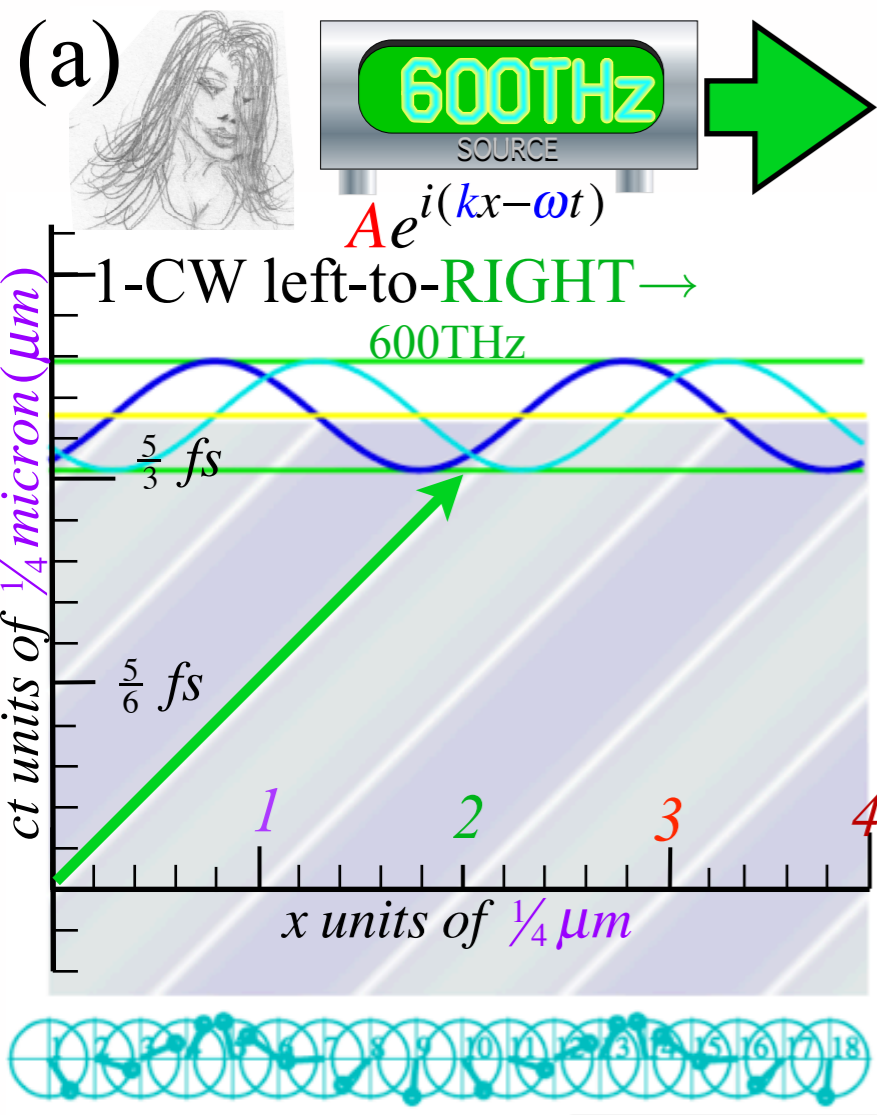
CW with ck = 2
CW with ck = 4
2 CW with ck = ± 2
RelaWavity $\beta = 0.0, v = 600$ THz
2 PW with ck = ± 2
RelaWavity $\beta = 0.0, v = 600$ THz
2 CW with ck = -1, 4
RelaWavity $\beta = 0.6, v = \sqrt{(300*1200)} = 600$ THz

k-Phasor Plot (100 THz Scale)

CW with ck = 3
CW with ck = -3
CW with ck = 6
CW with ck = 12
2 CW with ck = ± 6
RelaWavity $\beta = 0.0, v = 600$ THz
2 CW with ck = -3, 12
RelaWavity $\beta = 0.6, v = \sqrt{(300*1200)} = 600$ THz

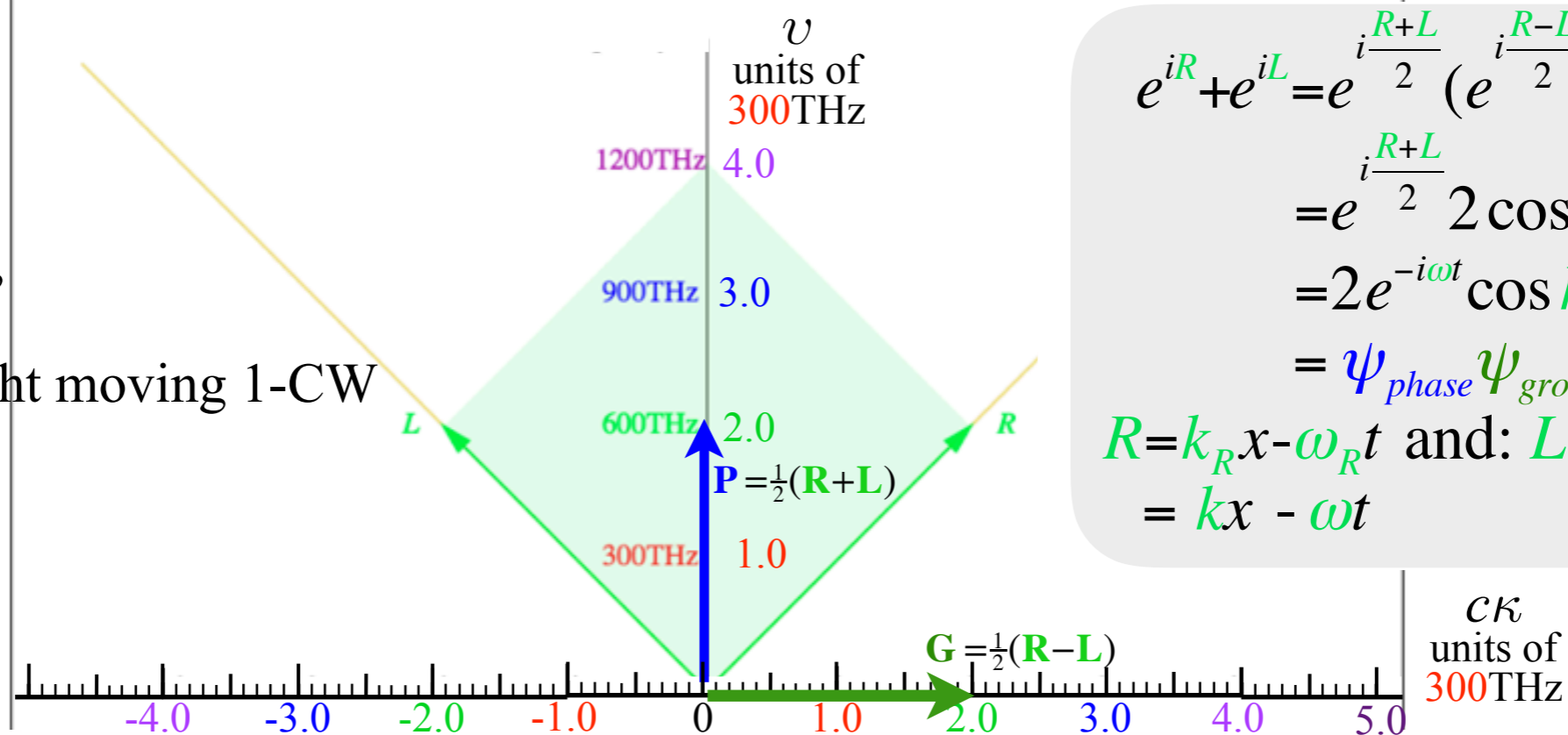
In the APP, right click on a scenario button to expose the actual scenario string





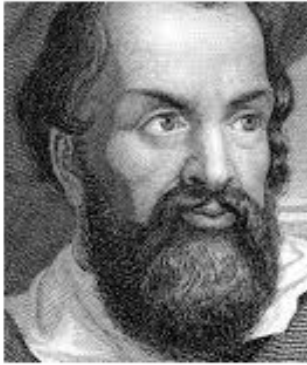
(d) Introducing optical space-time grids and per-space-time “baseball-diamonds”

\leftarrow LEFT-from-right moving 1-CW



$$\begin{aligned}
 e^{iR} + e^{iL} &= e^{i\frac{R+L}{2}} (e^{i\frac{R-L}{2}} + e^{-i\frac{R-L}{2}}) \\
 &= e^{i\frac{R+L}{2}} 2 \cos \frac{R-L}{2} \\
 &= 2e^{-i\omega t} \cos kx \\
 &= \psi_{phase} \psi_{group} \\
 R &= k_R x - \omega_R t \text{ and: } L = -k_L x - \omega_L t \\
 &= kx - \omega t \qquad \qquad \qquad = -kx - \omega t
 \end{aligned}$$

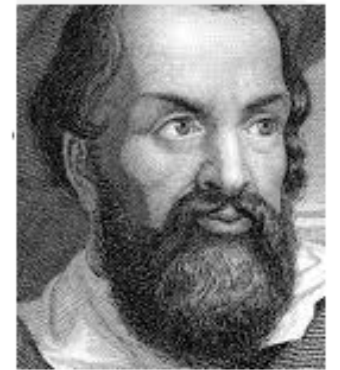
Galileo Galilei



1564-1642

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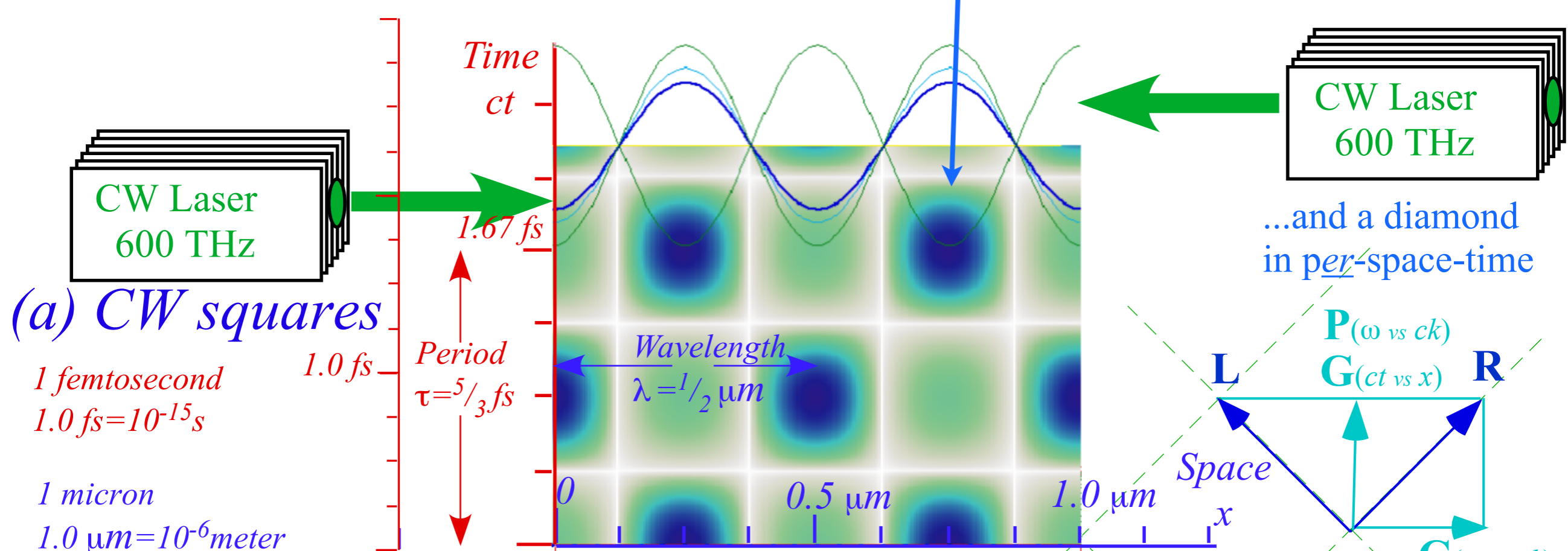
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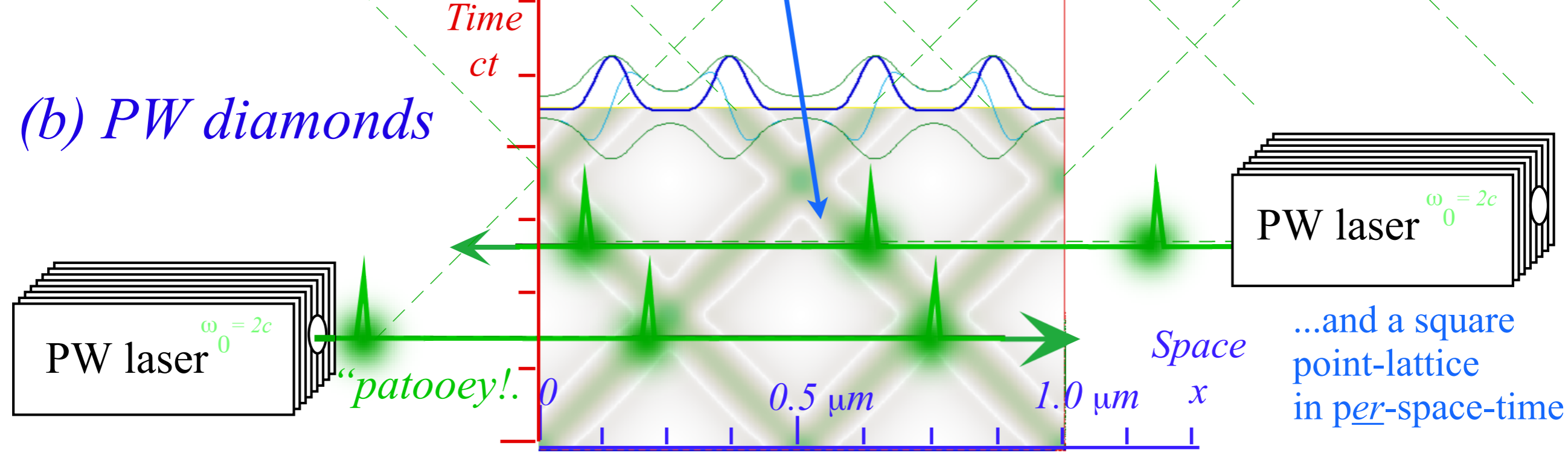
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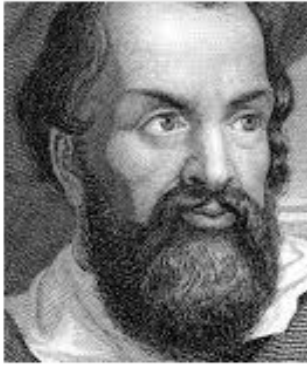
Continuous Waves (CW) trace “Cartesian squares” in space-time



Pulse Waves (PW) trace “baseball diamonds” in space-time



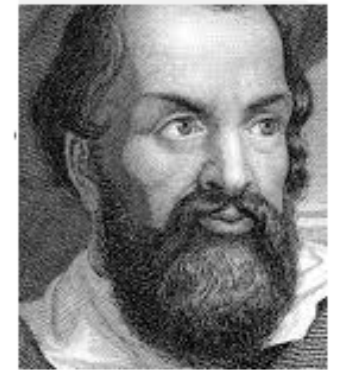
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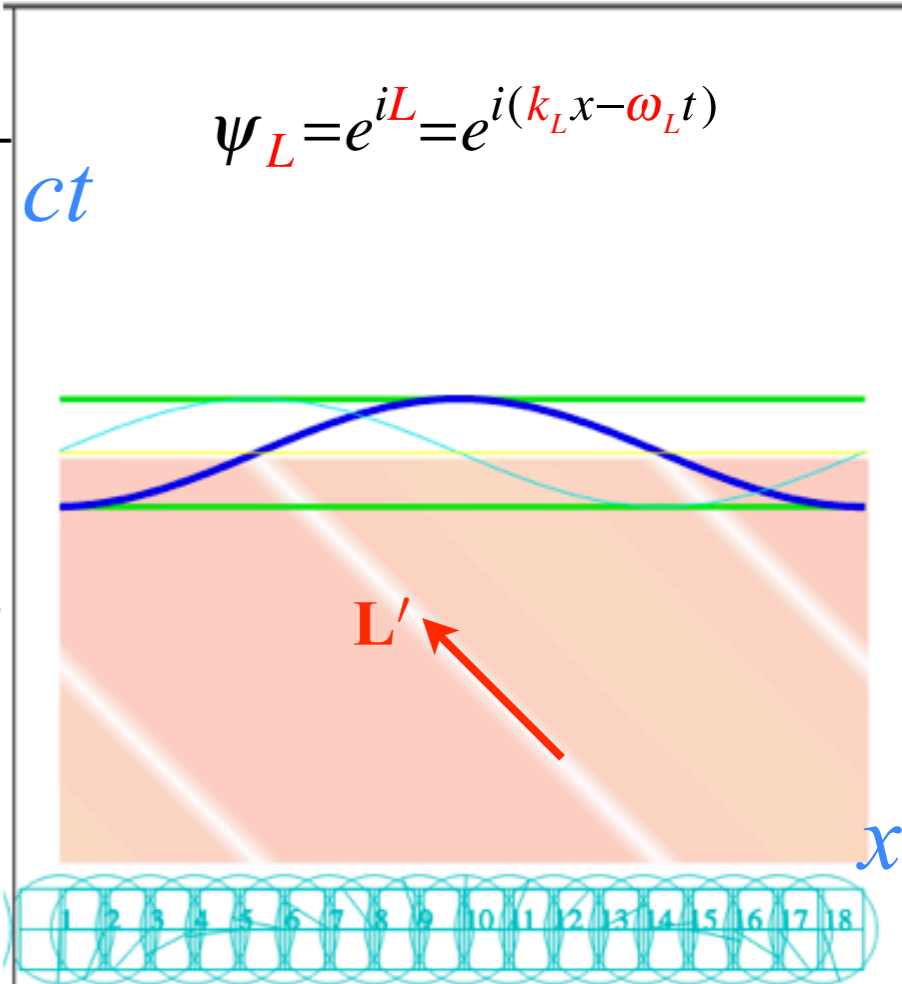
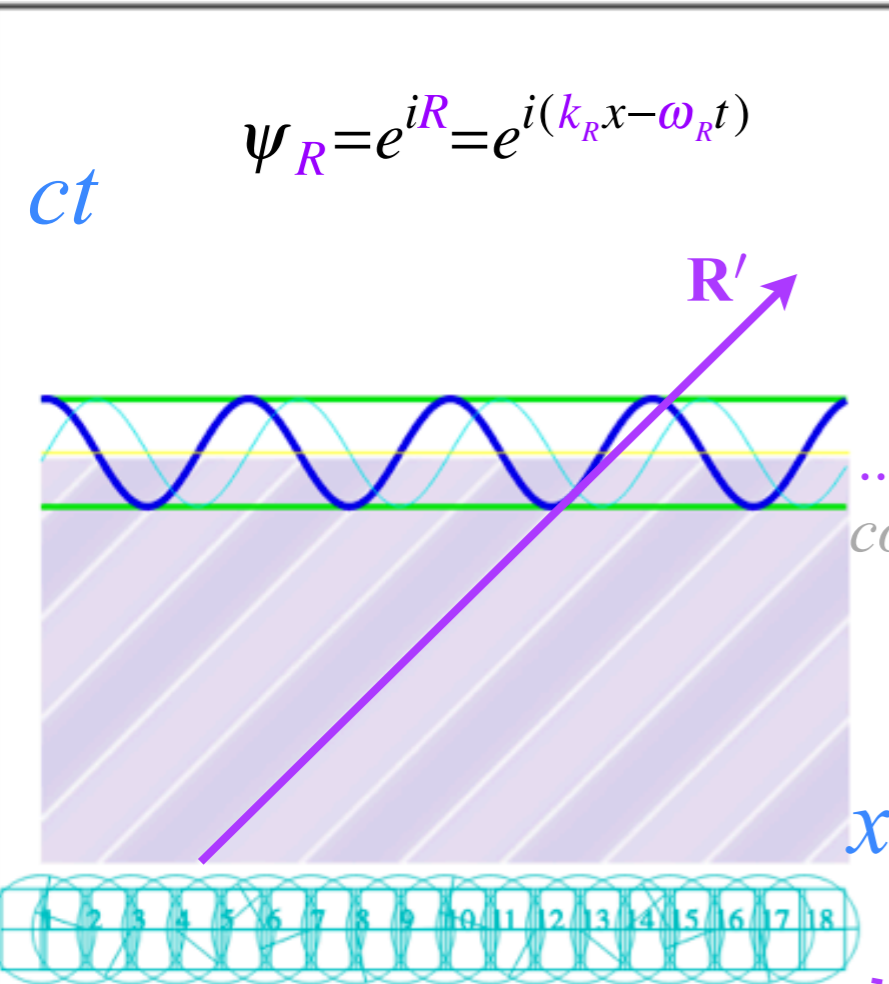
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right-moving Doppler blue shifted wave

left-moving Doppler red shifted wave

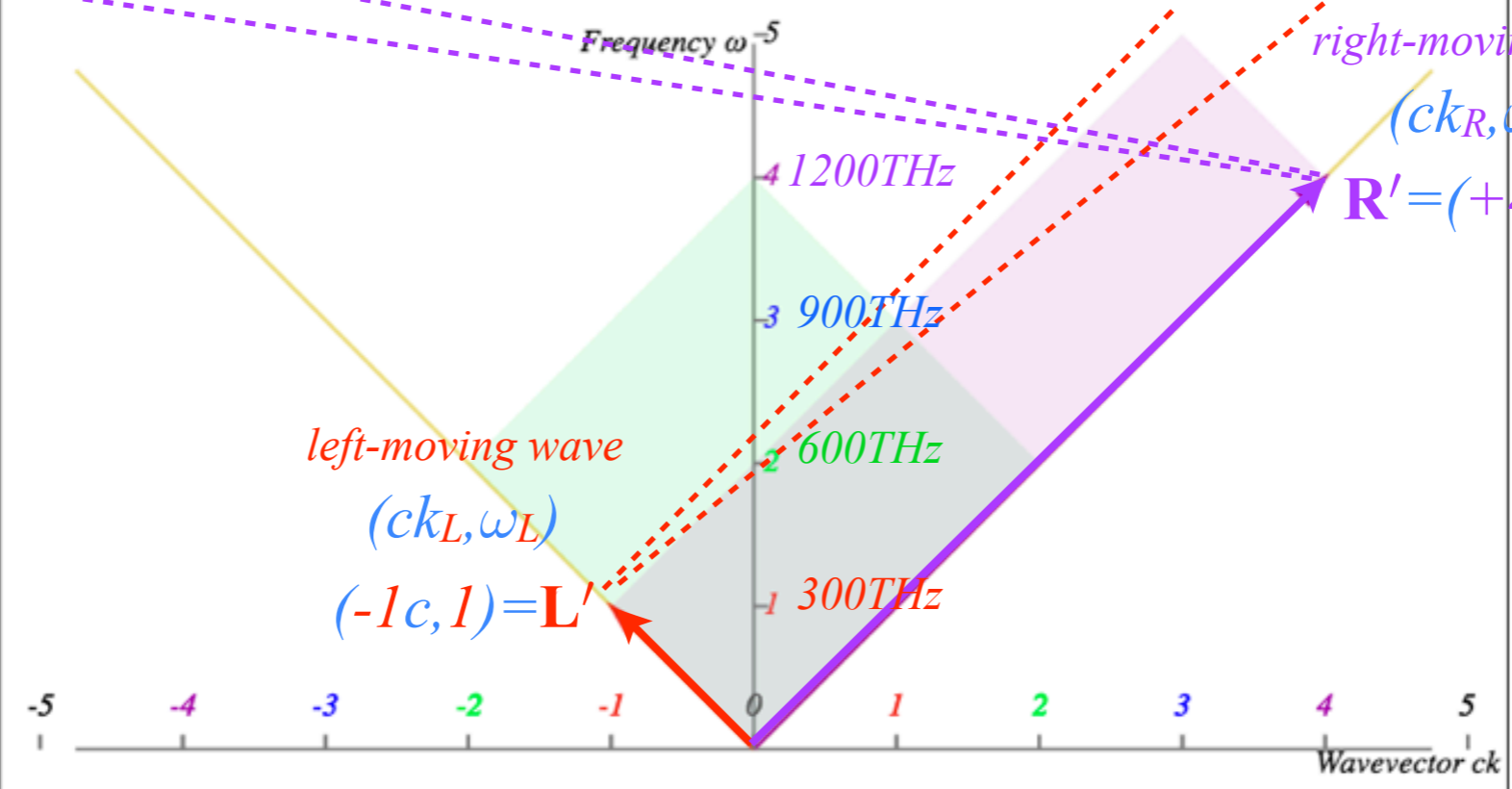


Rapidly moving Bob sees...



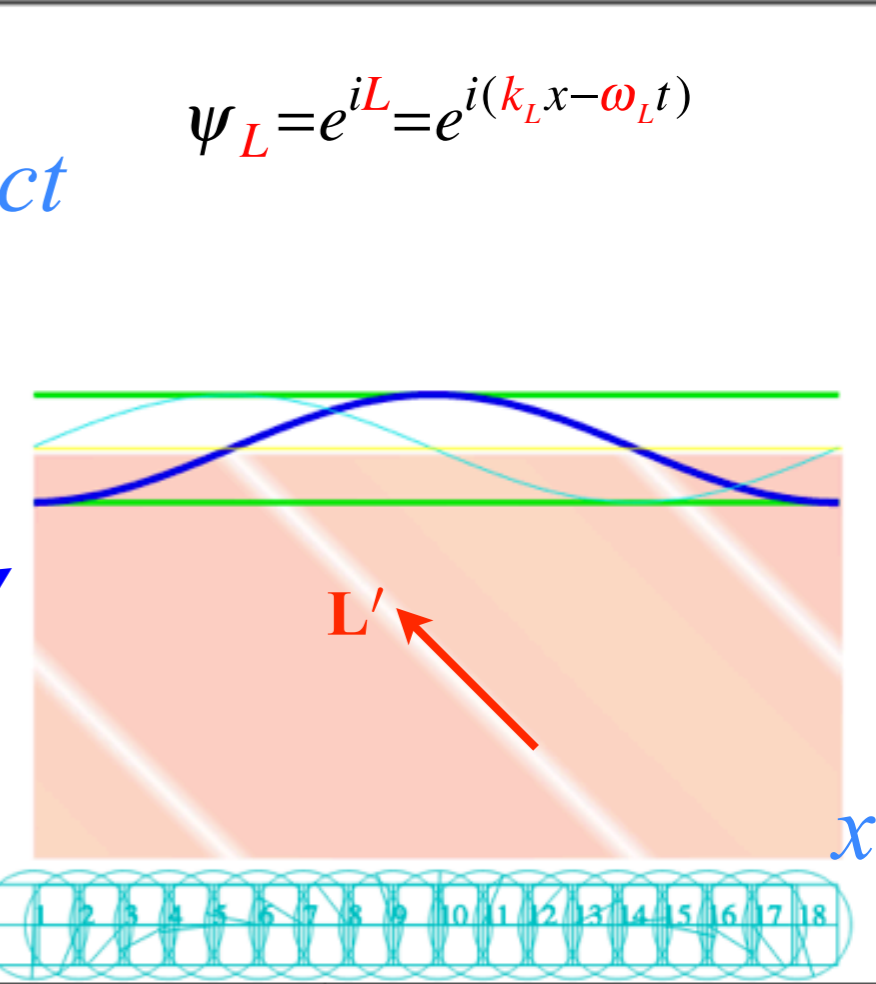
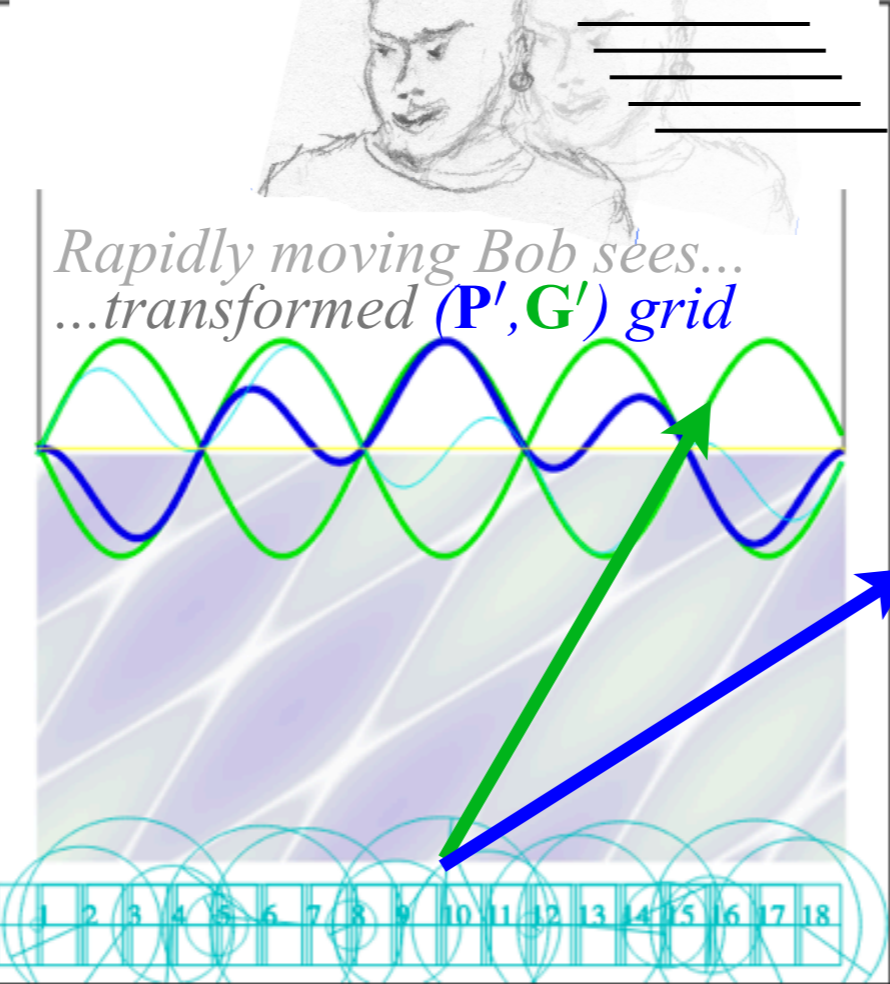
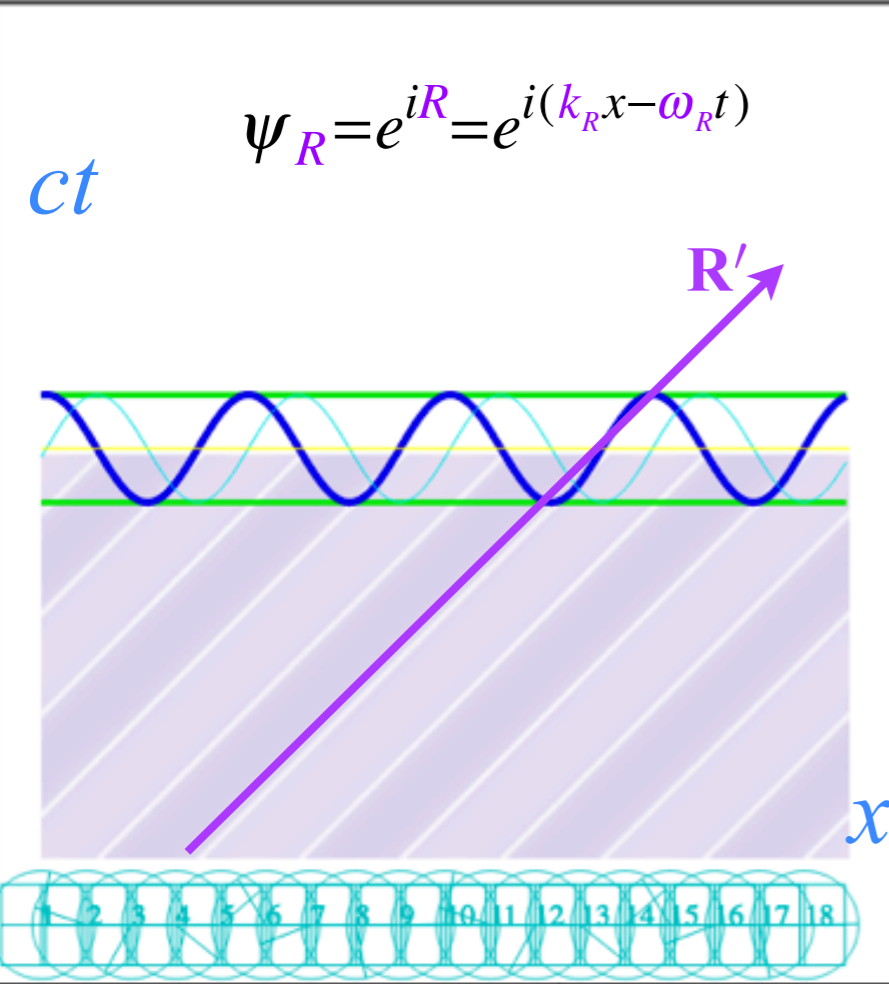
[Web Simulation 1 CW ct vs x Plot \(ck = +4\)](#)

[Web Simulation 1 CW ct vs x Plot \(ck = -1\)](#)



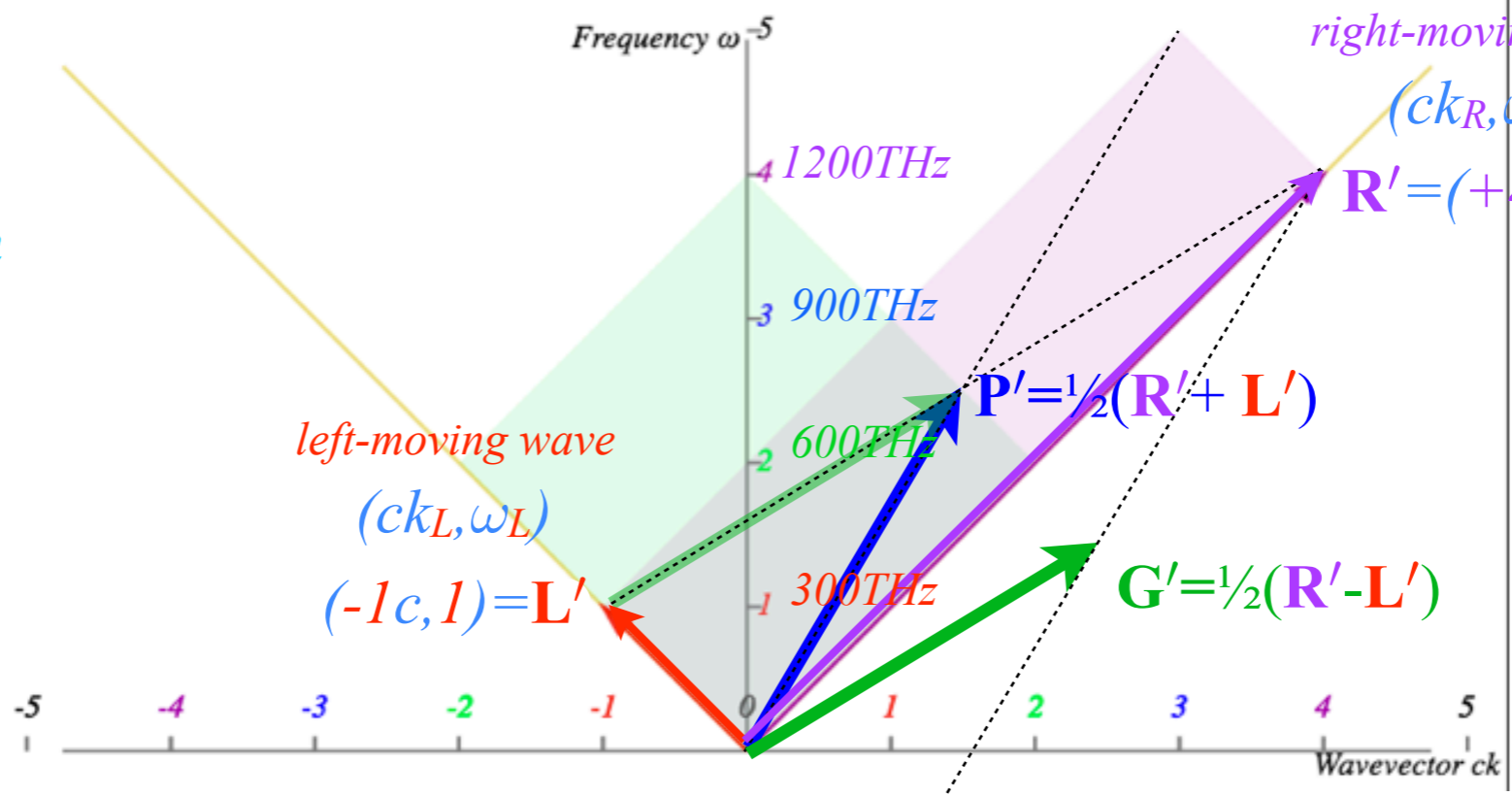
right-moving Doppler blue shifted wave

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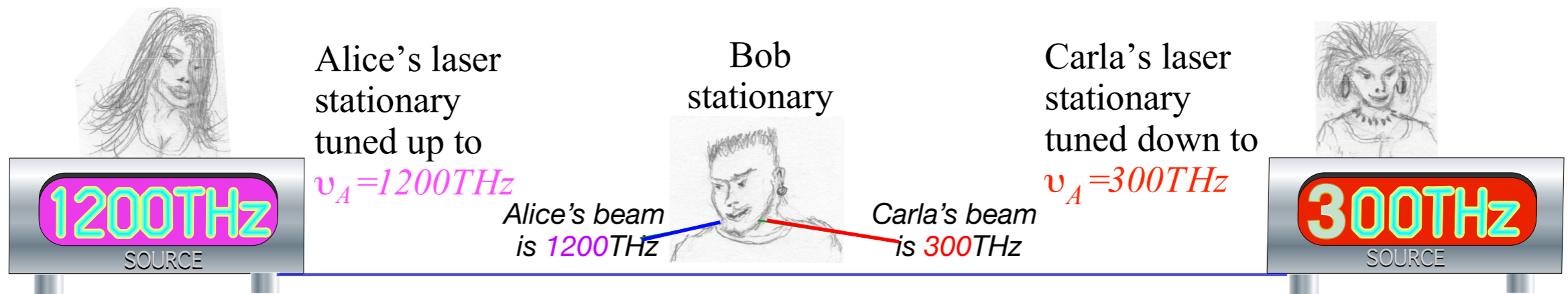
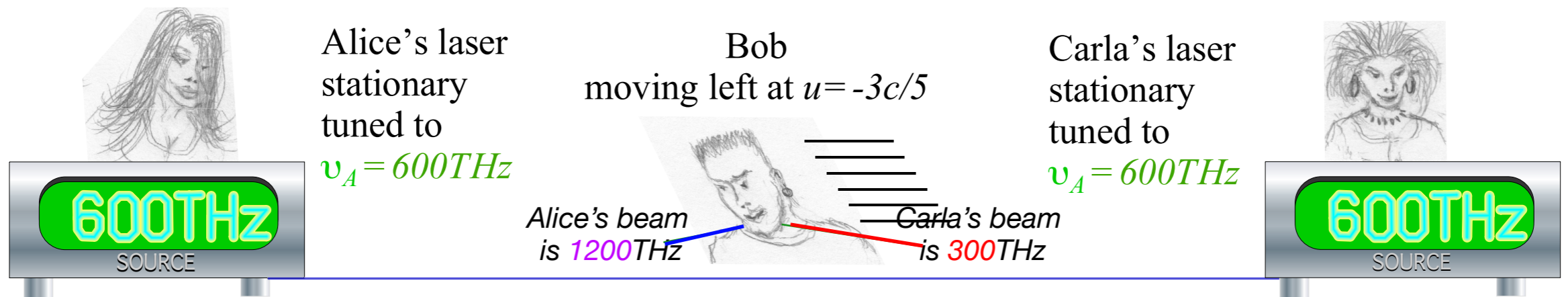
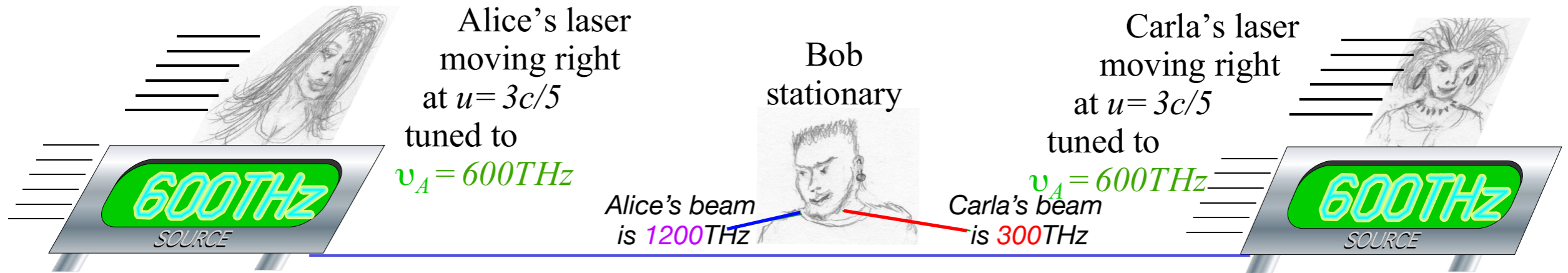


...Doppler shifts give Lorentz transformation of both these graphs

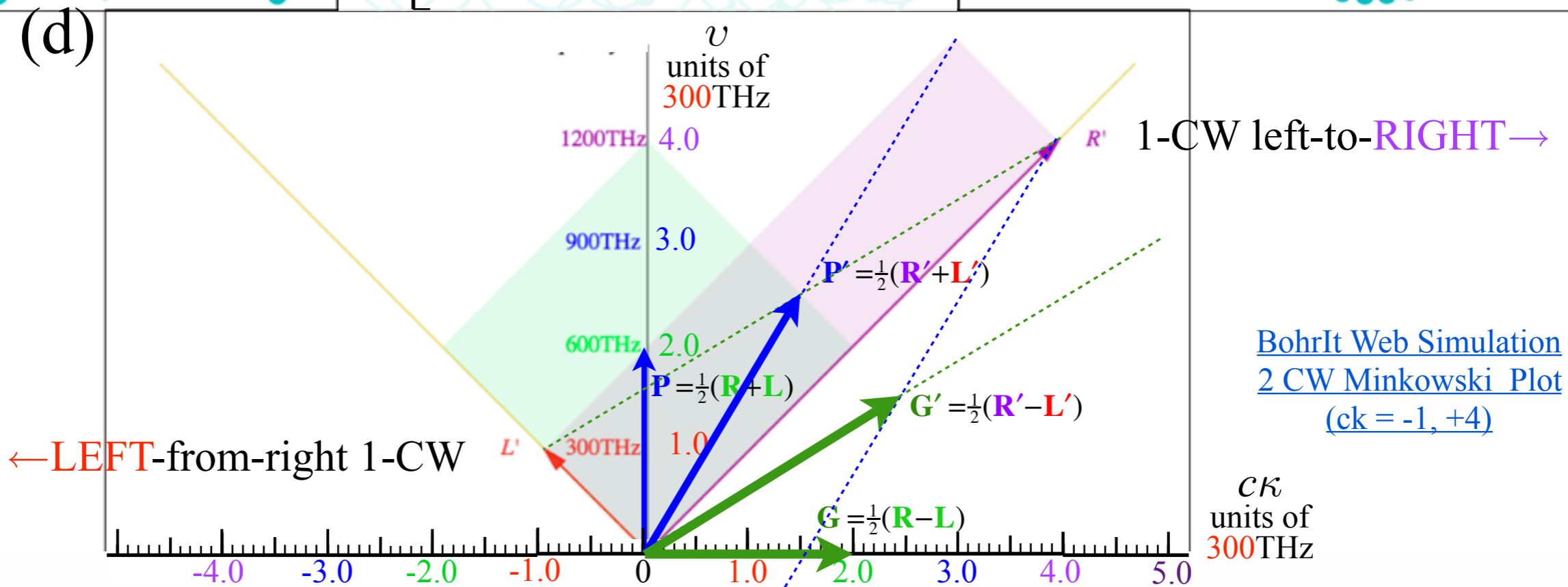
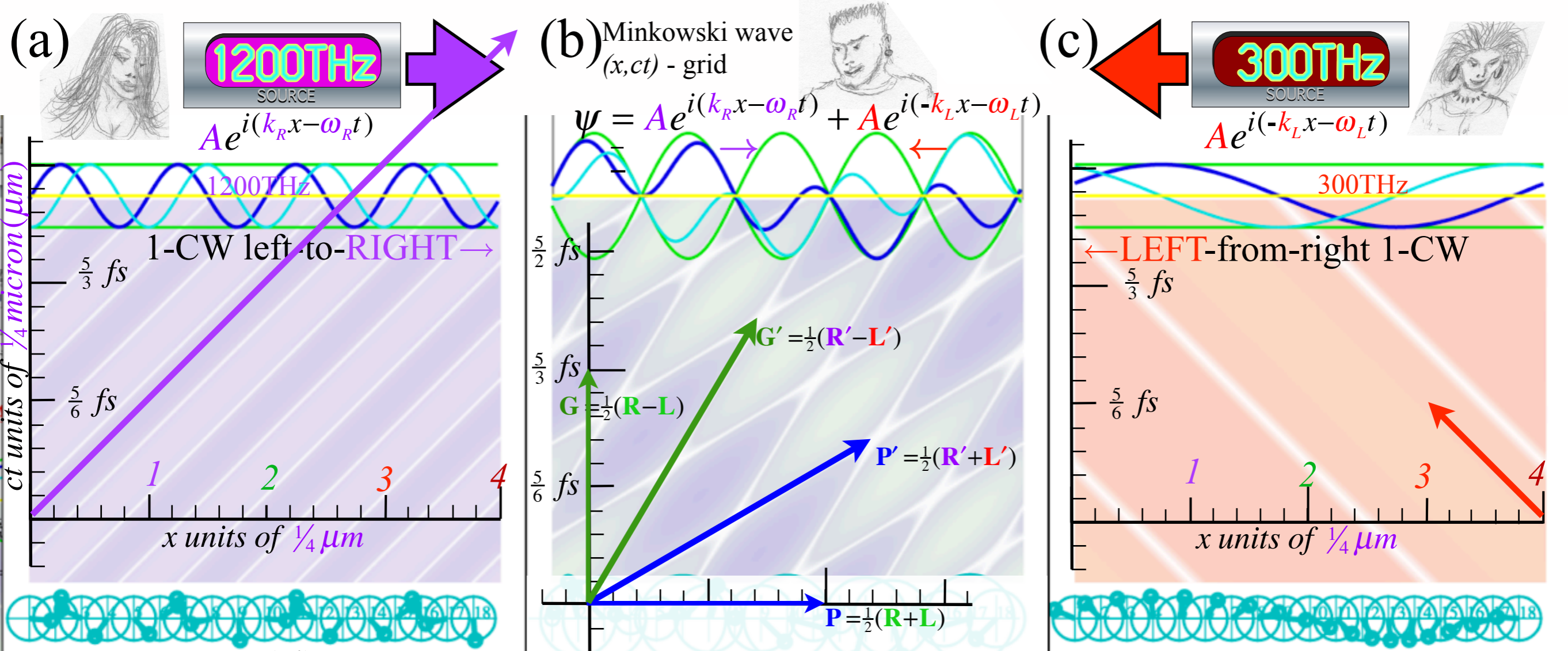
Per-Spacetime (ck, ω)

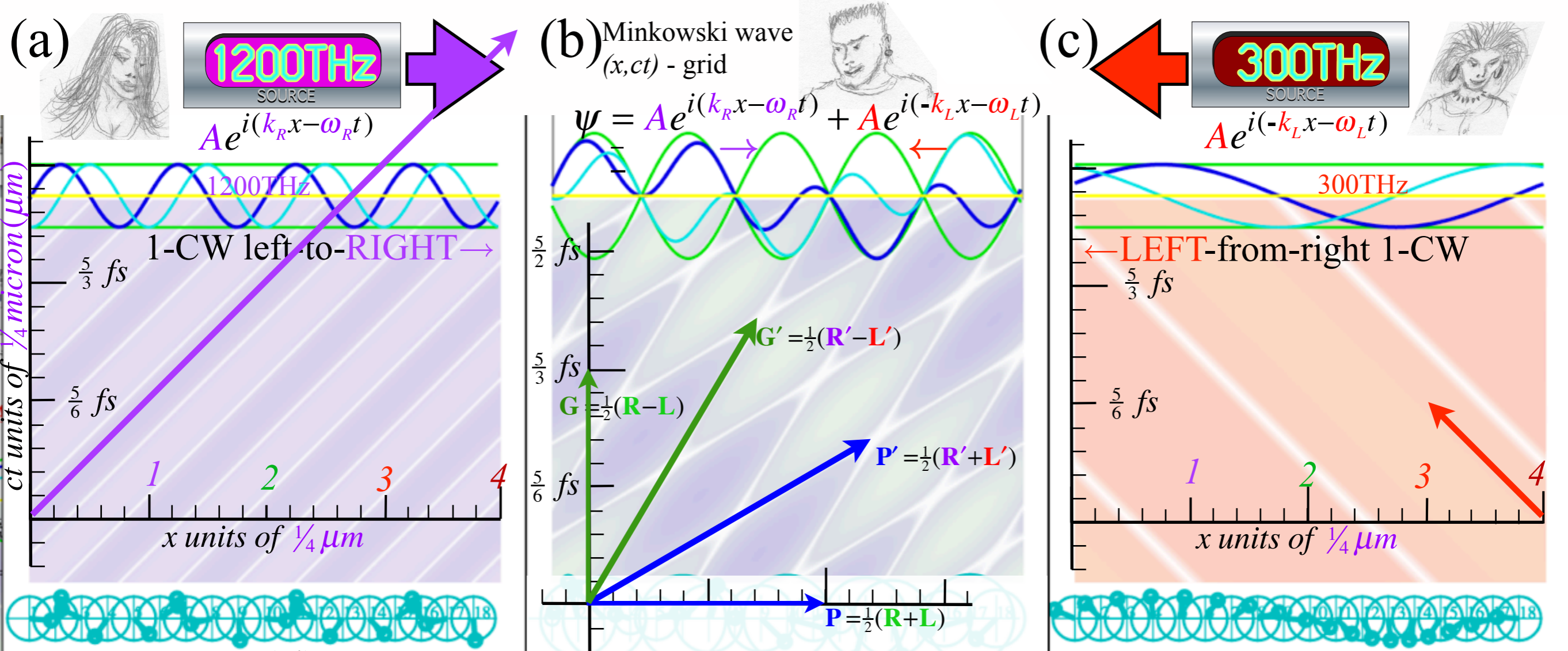


Three scenarios that look the same to Bob



Much cheaper (and safer) to do the 3rd scenario!\$!





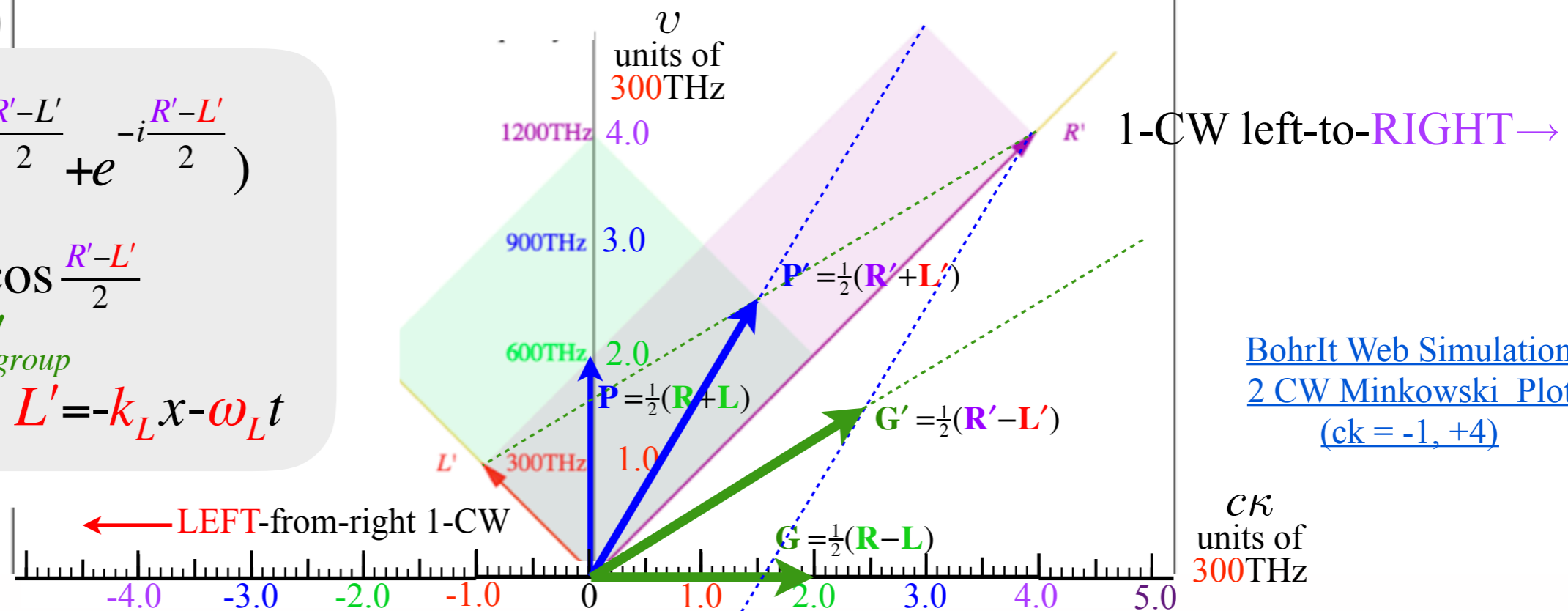
(d)

$$e^{iR'} + e^{iL'} = e^{i\frac{R'+L'}{2}} (e^{i\frac{R'-L'}{2}} + e^{-i\frac{R'-L'}{2}})$$

$$= e^{i\frac{R'+L'}{2}} 2 \cos \frac{R'-L'}{2}$$

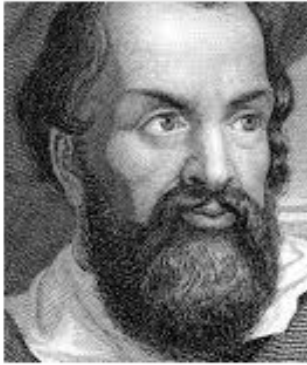
$$= \psi'_{phase} \psi'_{group}$$

$$R' = k_R x - \omega_R t \text{ and: } L' = -k_L x - \omega_L t$$



BohrIt Web Simulation
 2 CW Minkowski Plot
 (ck = -1, +4)

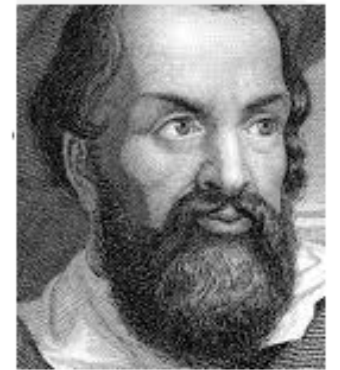
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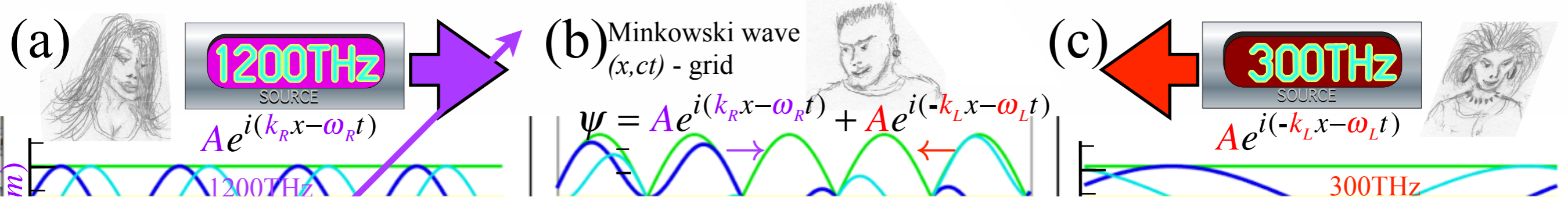
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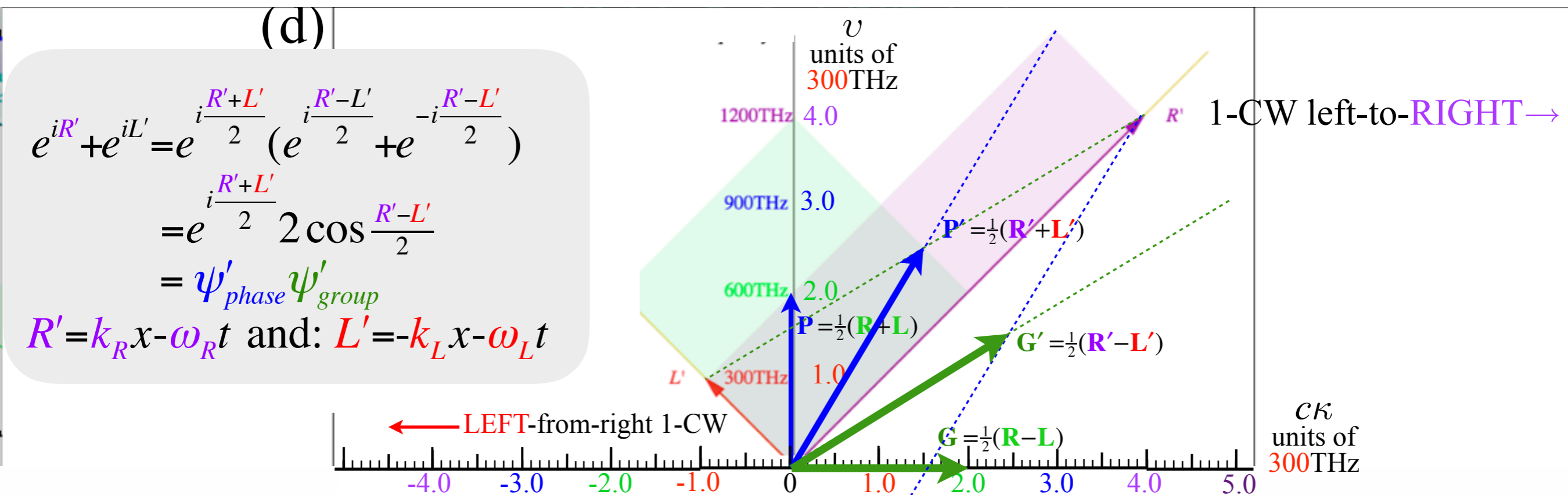
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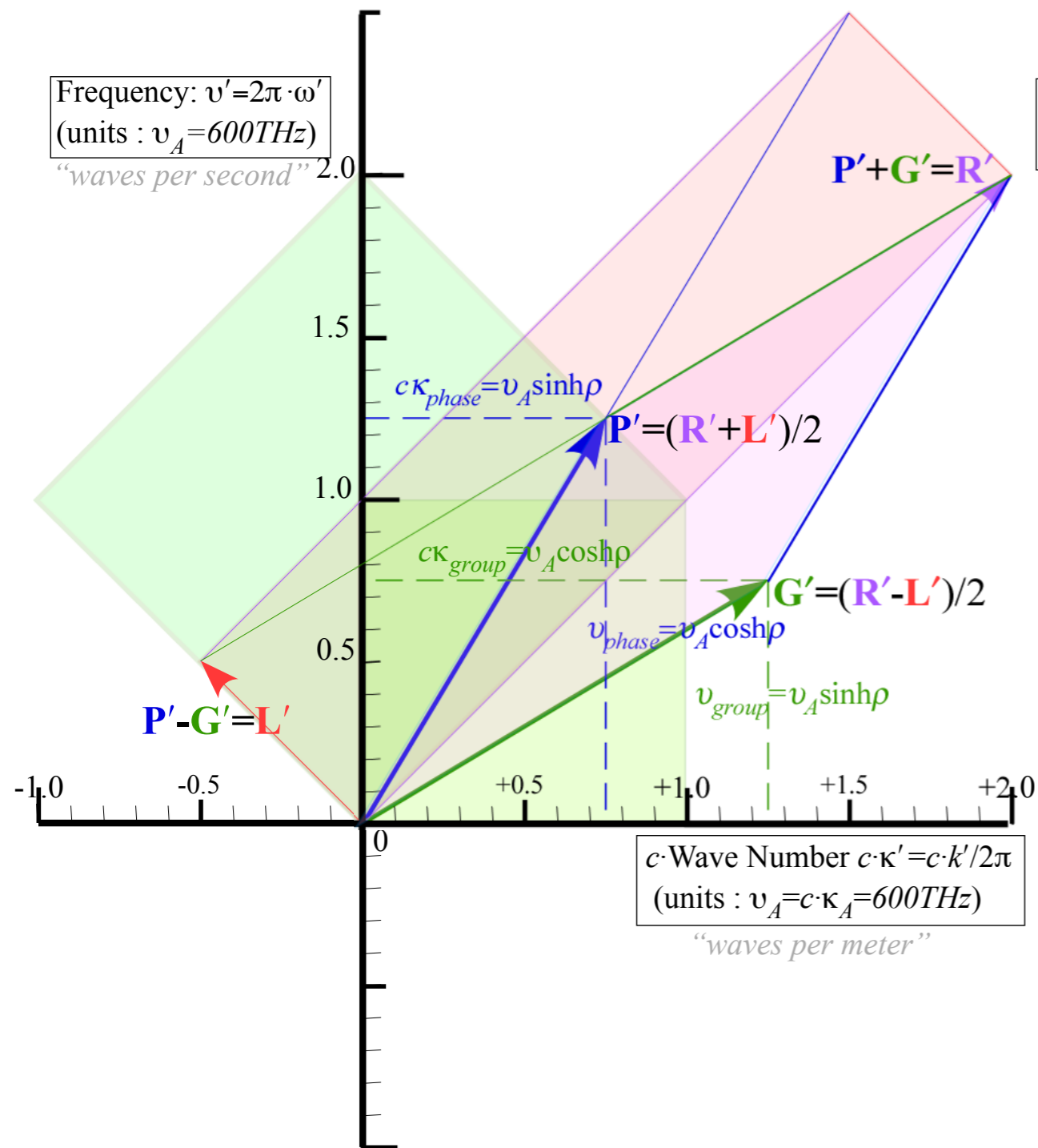


$$\mathbf{P}' = \begin{pmatrix} v'_{phase} \\ cK'_{phase} \end{pmatrix} = \frac{1}{2}(\mathbf{R}' + \mathbf{L}') = v_A \begin{pmatrix} \frac{1}{2}(e^\rho + e^{-\rho}) \\ \frac{1}{2}(e^\rho - e^{-\rho}) \end{pmatrix} = v_A \begin{pmatrix} \cosh \rho \\ \sinh \rho \end{pmatrix} = v_A \begin{pmatrix} \frac{5}{2} \\ \frac{3}{2} \end{pmatrix} \text{ Bob's View} \quad \text{or: } v_A \begin{pmatrix} 1 \\ 0 \end{pmatrix} \text{ Alice's View}$$

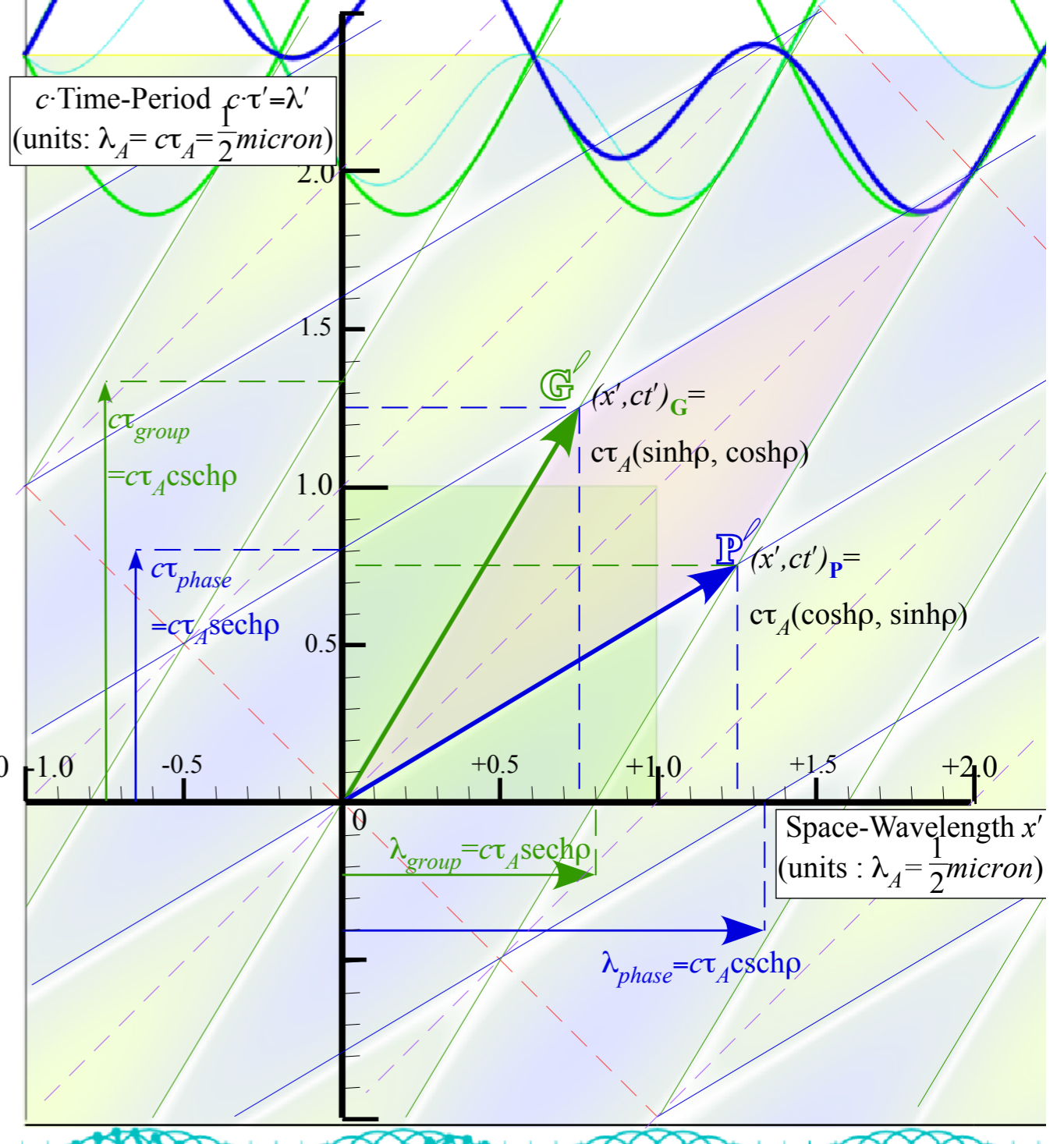
$$\mathbf{G}' = \begin{pmatrix} v'_{group} \\ cK'_{group} \end{pmatrix} = \frac{1}{2}(\mathbf{R}' - \mathbf{L}') = v_A \begin{pmatrix} \frac{1}{2}(e^\rho - e^{-\rho}) \\ \frac{1}{2}(e^\rho + e^{-\rho}) \end{pmatrix} = v_A \begin{pmatrix} \sinh \rho \\ \cosh \rho \end{pmatrix} = v_A \begin{pmatrix} \frac{3}{2} \\ \frac{5}{2} \end{pmatrix} \text{ Bob's View} \quad \text{or: } v_A \begin{pmatrix} 0 \\ 1 \end{pmatrix} \text{ Alice's View}$$



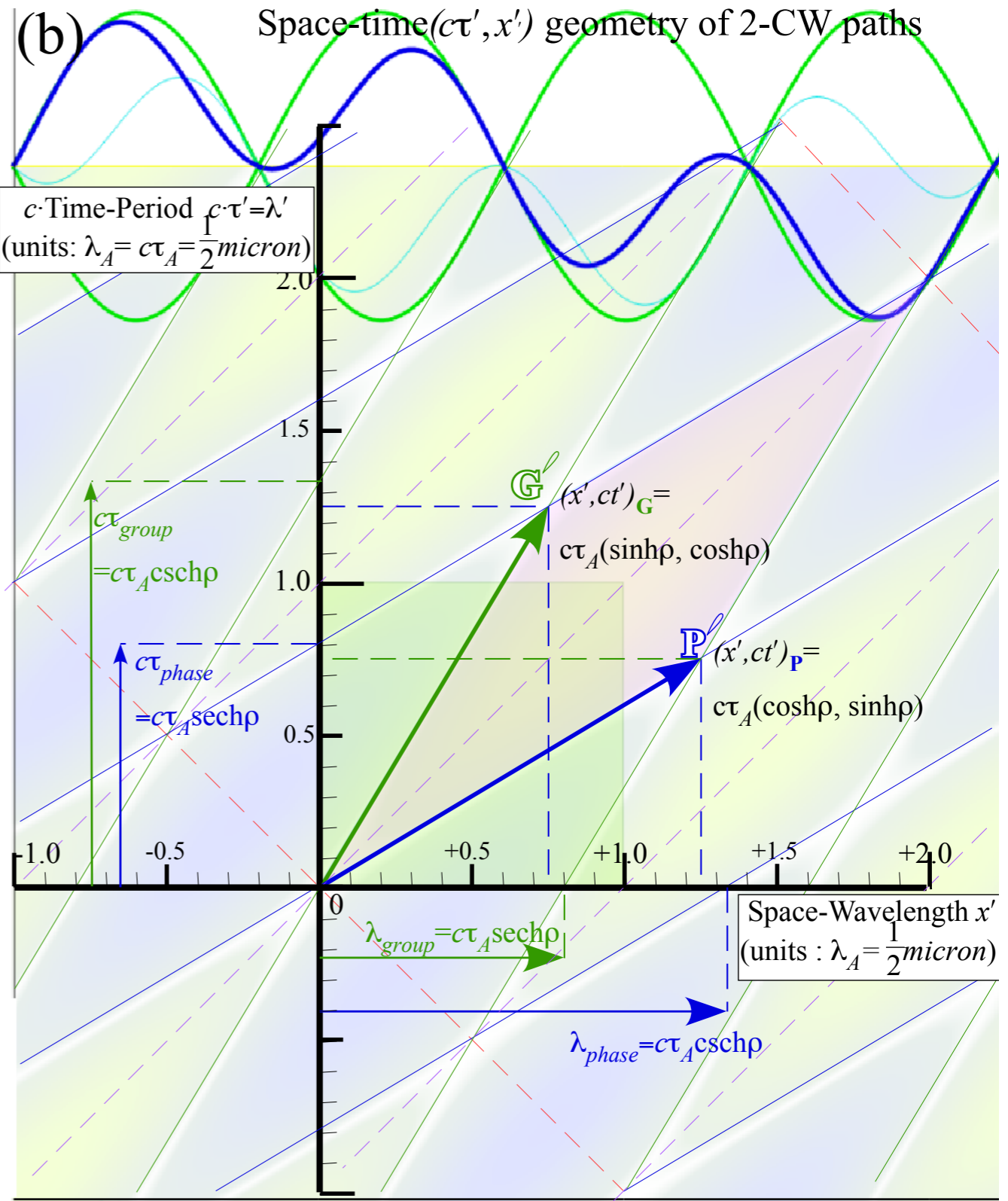
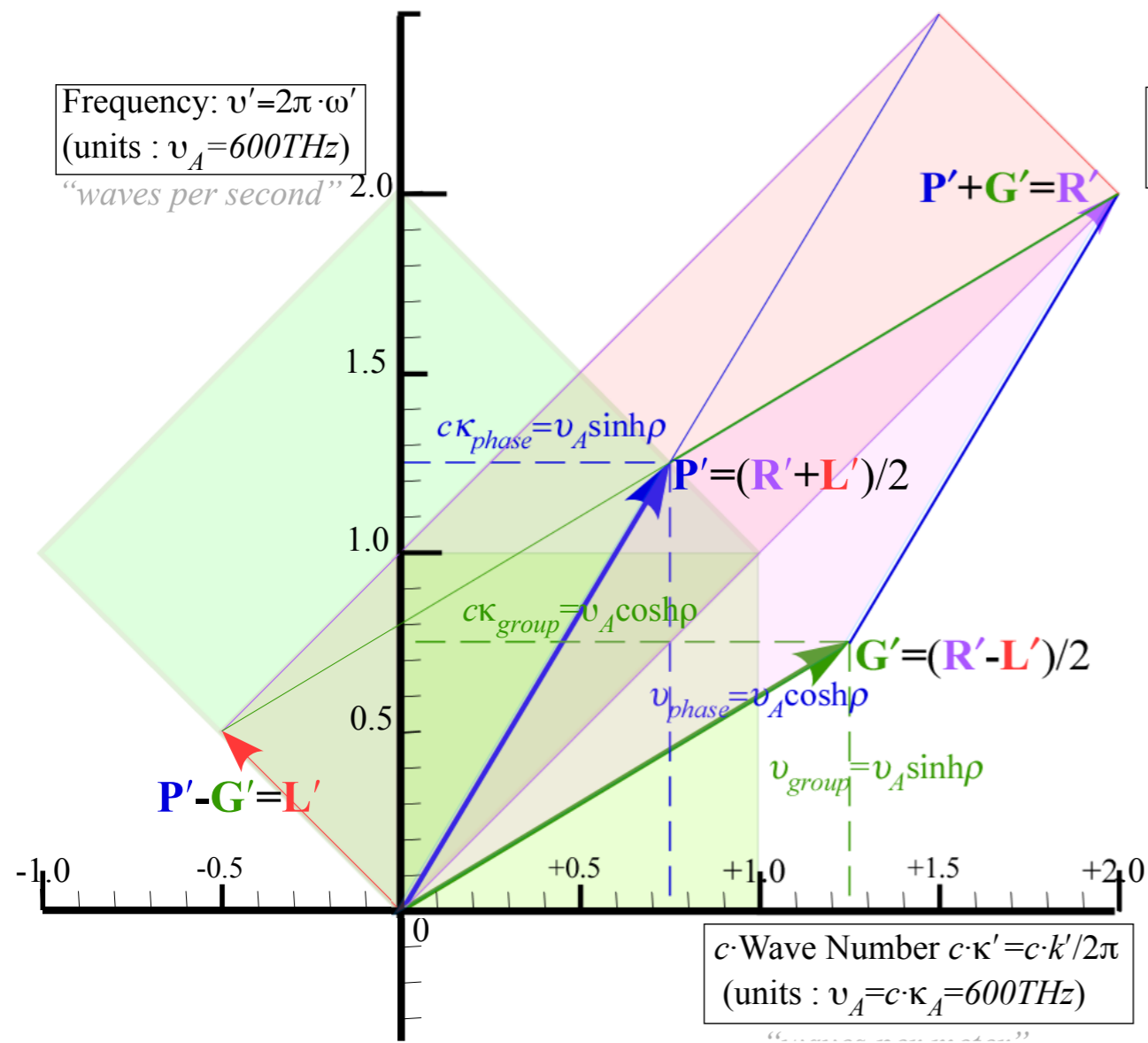
(a) Per-space-time $(\nu', c\kappa')$ geometry of 2-CW vectors



(b) Space-time $(c\tau', x')$ geometry of 2-CW paths



(a) Per-space-time $(v', c\kappa')$ geometry of 2-CW vectors

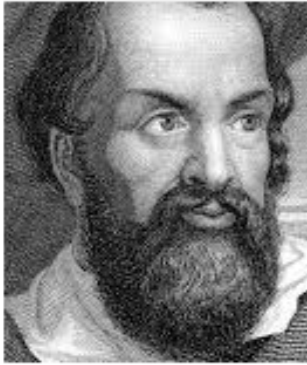


The slope of Bob's group vector \mathbf{G}' in $(c\kappa, v)$ -plot is actual group wave velocity in c -units.

$$\frac{V^{\text{group}}}{c} = \frac{v'_{\text{group}}}{c\kappa'_{\text{group}}} = \frac{\sinh \rho}{\cosh \rho} = \tanh \rho = \frac{\frac{3}{2}}{\frac{5}{2}} = \frac{3}{5} \equiv \frac{u}{c} \equiv \beta$$

Group vector \mathbf{G}' in (x, ct) -plot has 3/5 slope relative to time axis

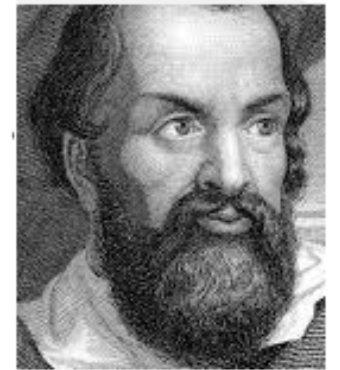
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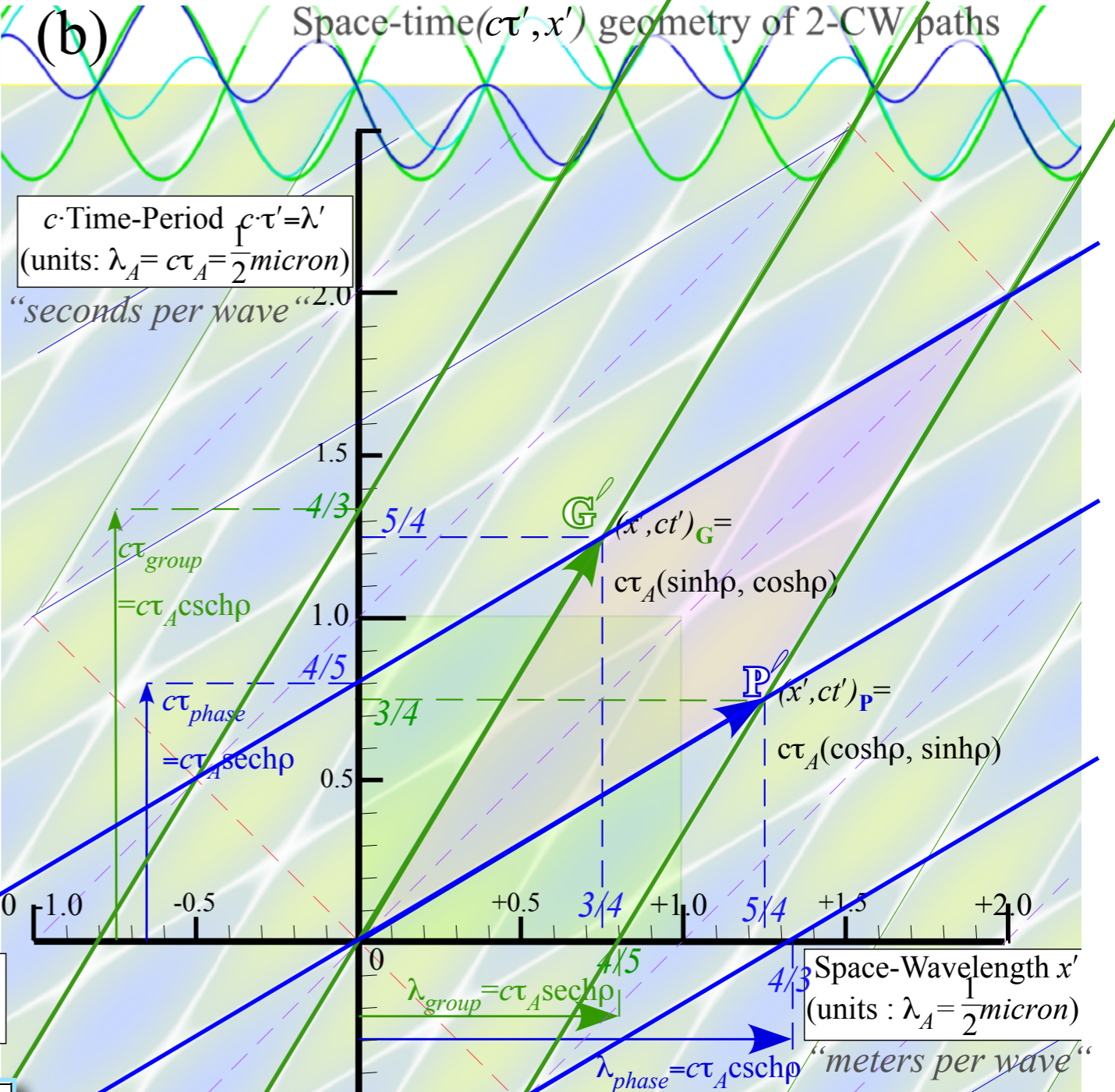
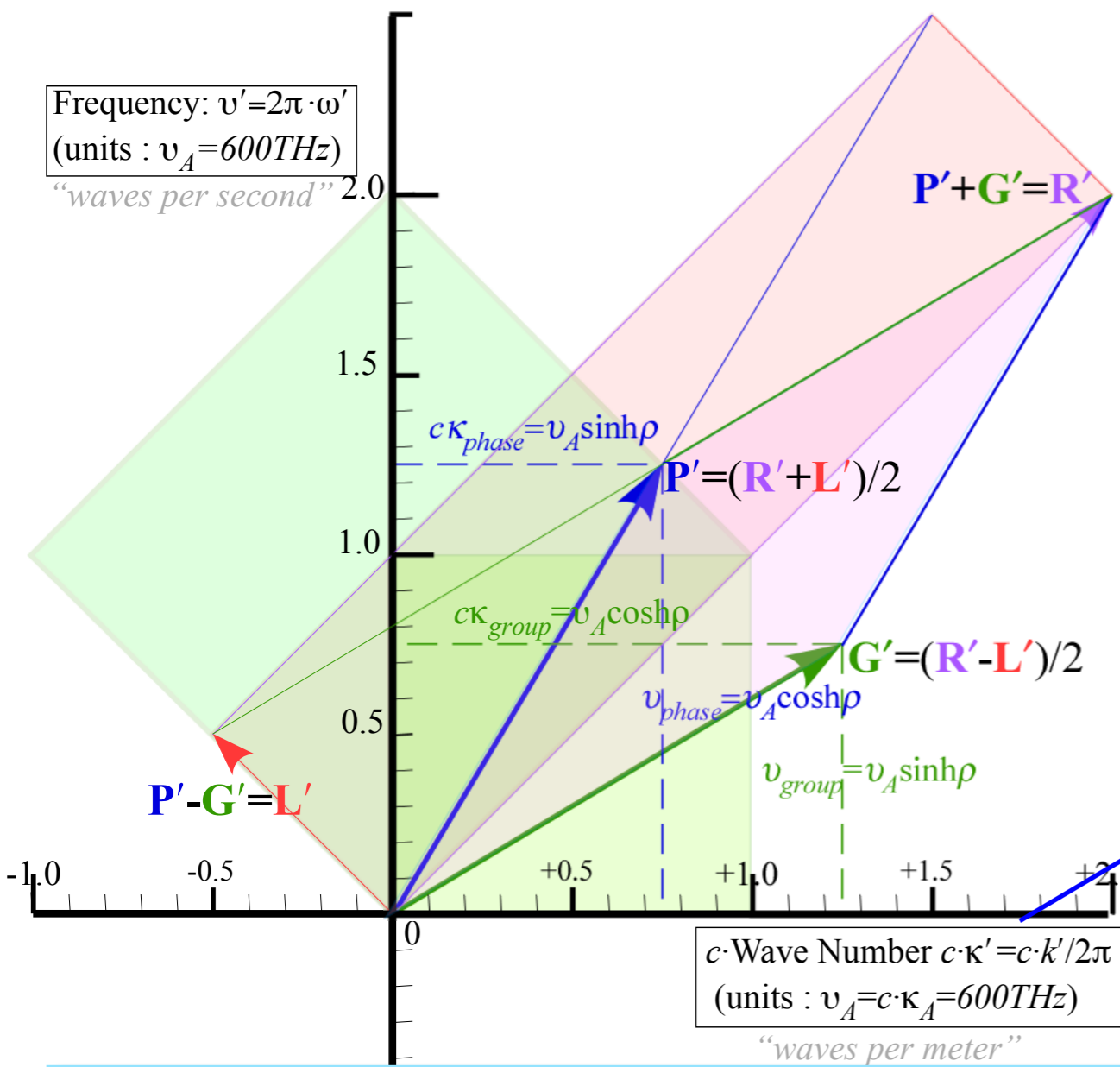
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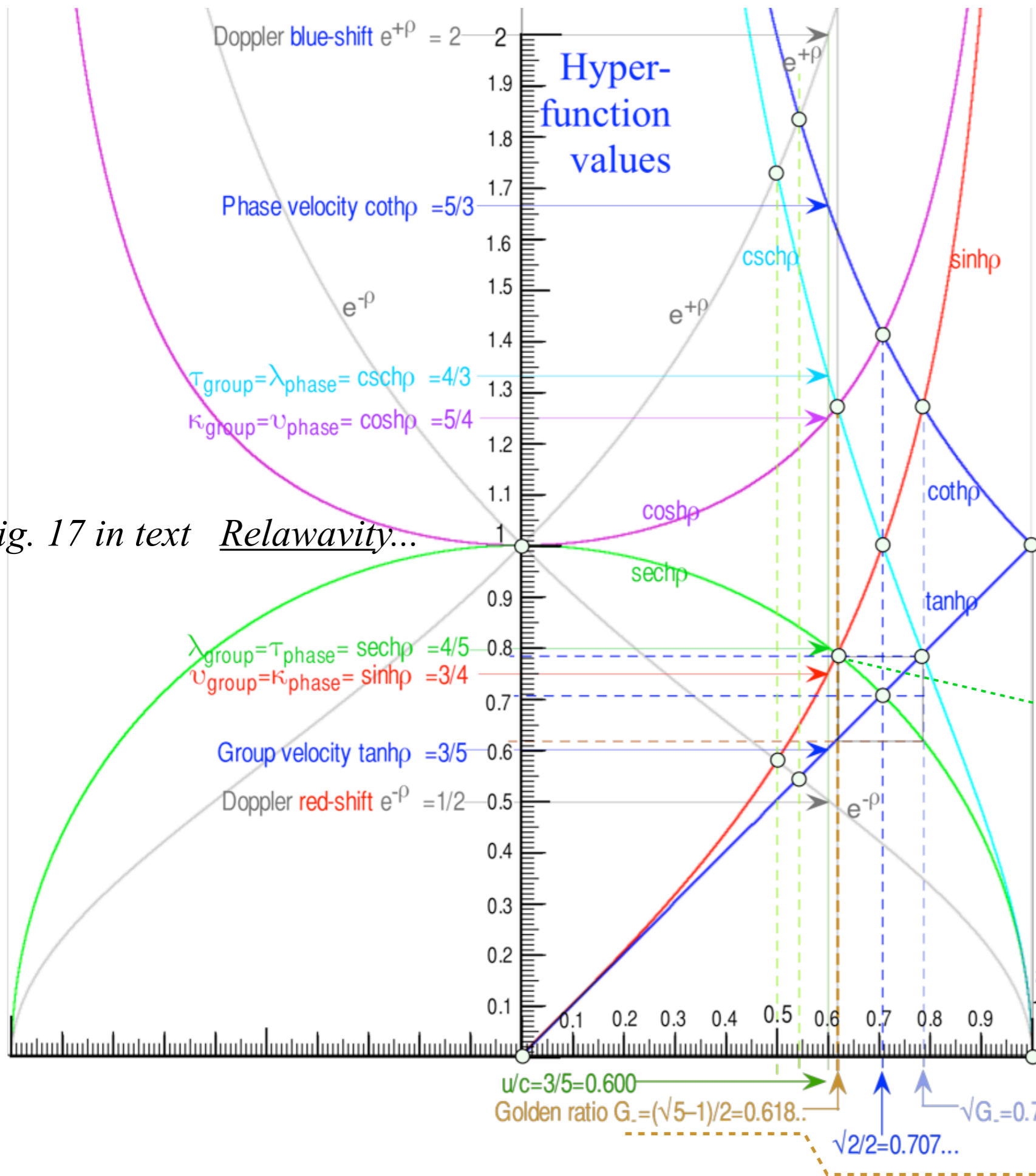
Application to TE-Waveguide modes and synchrotron beam relativity

(a) Per-space-time $(v', c\kappa')$ geometry of 2-CW vectors



group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

Fig. 11 in text *Relativity...*



If $\frac{u}{c} = \tanh \rho = 0.618\dots$ (Golden-Mean G_-)

two parameters become *exactly equal* :

$$\frac{ct'_P}{c\tau_A} = \sinh \rho = \frac{\lambda_{group}}{\lambda_A} = \frac{\tau_{phase}}{\tau_A} = \operatorname{sech} \rho = 0.786\dots = \sqrt{G_-}$$

and

$$\frac{x'_P}{\lambda_A} = \cosh \rho = \frac{\lambda_{phase}}{\lambda_A} = \frac{\tau_{group}}{\tau_A} = \operatorname{csch} \rho = 1.272\dots = 1/\sqrt{G_-} = 1.272\dots$$

Fig. 17 in text Relativity...

Solve :

$$\operatorname{sech} \rho = \sinh \rho$$

or:

$$\sinh \rho \cosh \rho = 1$$

or:

$$\sinh 2\rho = 2$$

$$\rho = \frac{1}{2} \sinh^{-1} 2 = 0.7218\dots$$

$$\tanh \rho = 0.618\dots = \frac{\sqrt{5}-1}{2}$$

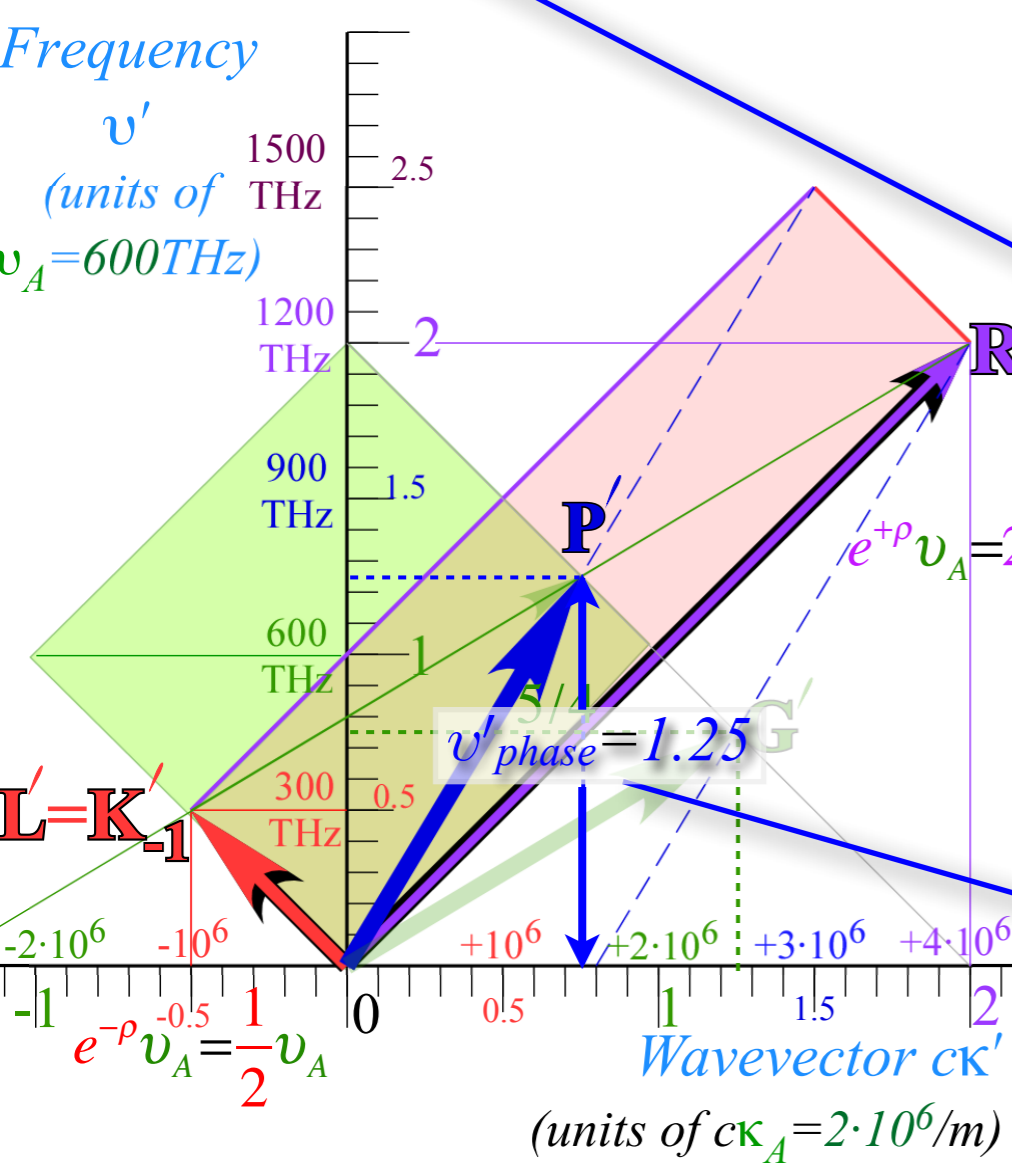
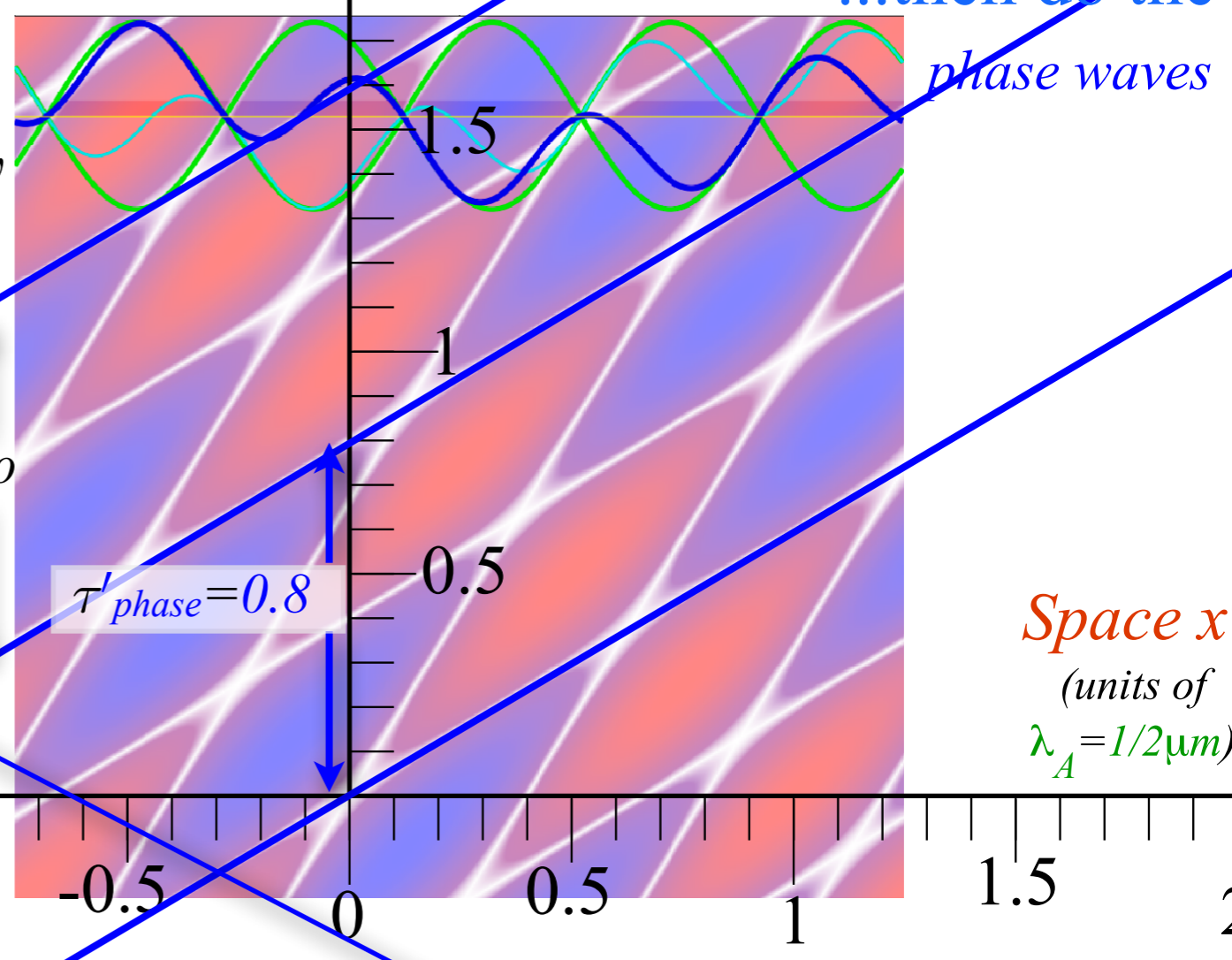
The 16 dimensions of 2CW interference

Time ct'
(units of $\lambda_A = 1/2 \mu m$)

Start with the **Dopplers**
...then do the **phase waves**

$$\mathbf{P}' = \begin{pmatrix} c\mathbf{K}'_{phase} \\ v'_{phase} \end{pmatrix} = v_A \begin{pmatrix} \sinh \rho \\ \cosh \rho \end{pmatrix} = v_A \begin{pmatrix} 3/4 \\ 5/4 \end{pmatrix}$$

Phase frequency $v'_{phase} = v_A \cosh \rho = 5/4 = 1.25$
 flips to Phase period $\tau'_{phase} = \tau_A \operatorname{sech} \rho = 4/5 = 0.8$

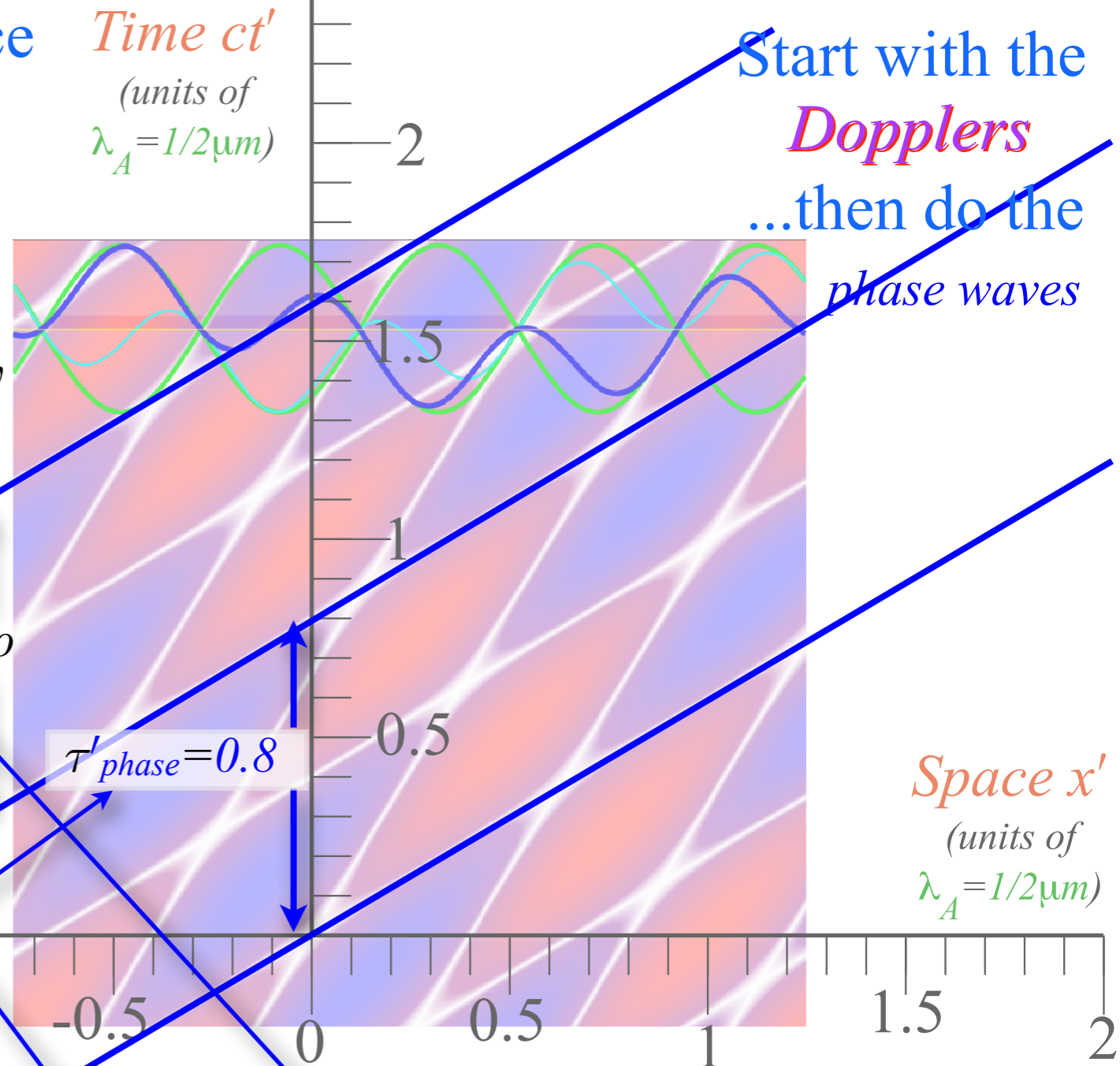
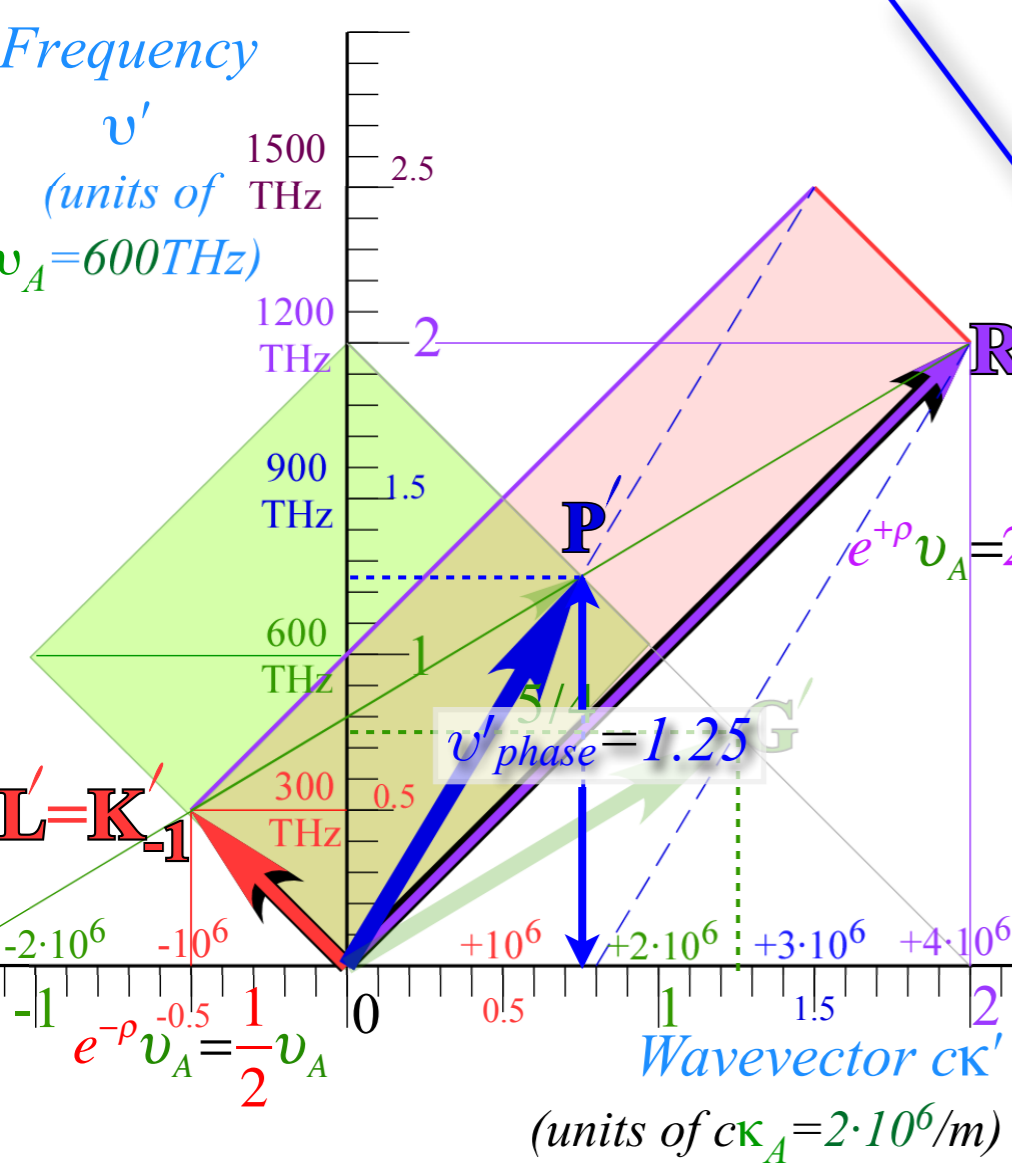


phase	$b_{Doppler RED}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler BLUE}$
group	1	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	1
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

The 16 dimensions of 2CW interference

$$\mathbf{P}' = \begin{pmatrix} c\mathbf{K}'_{phase} \\ v'_{phase} \end{pmatrix} = v_A \begin{pmatrix} \sinh \rho \\ \cosh \rho \end{pmatrix} = v_A \begin{pmatrix} 3/4 \\ 5/4 \end{pmatrix}$$

Phase frequency $v'_{phase} = v_A \cosh \rho = 5/4 = 1.25$ flips to Phase period $\tau'_{phase} = \tau_A \operatorname{sech} \rho = 4/5 = 0.8$



Start with the Dopplers ... then do the phase waves

phase	$b_{RED}^{Doppler}$	$\frac{v_{phase}}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
group	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

Group wavenumber
 $\kappa'_{group} = \kappa_A \cosh \rho = 5/4 = 1.25$

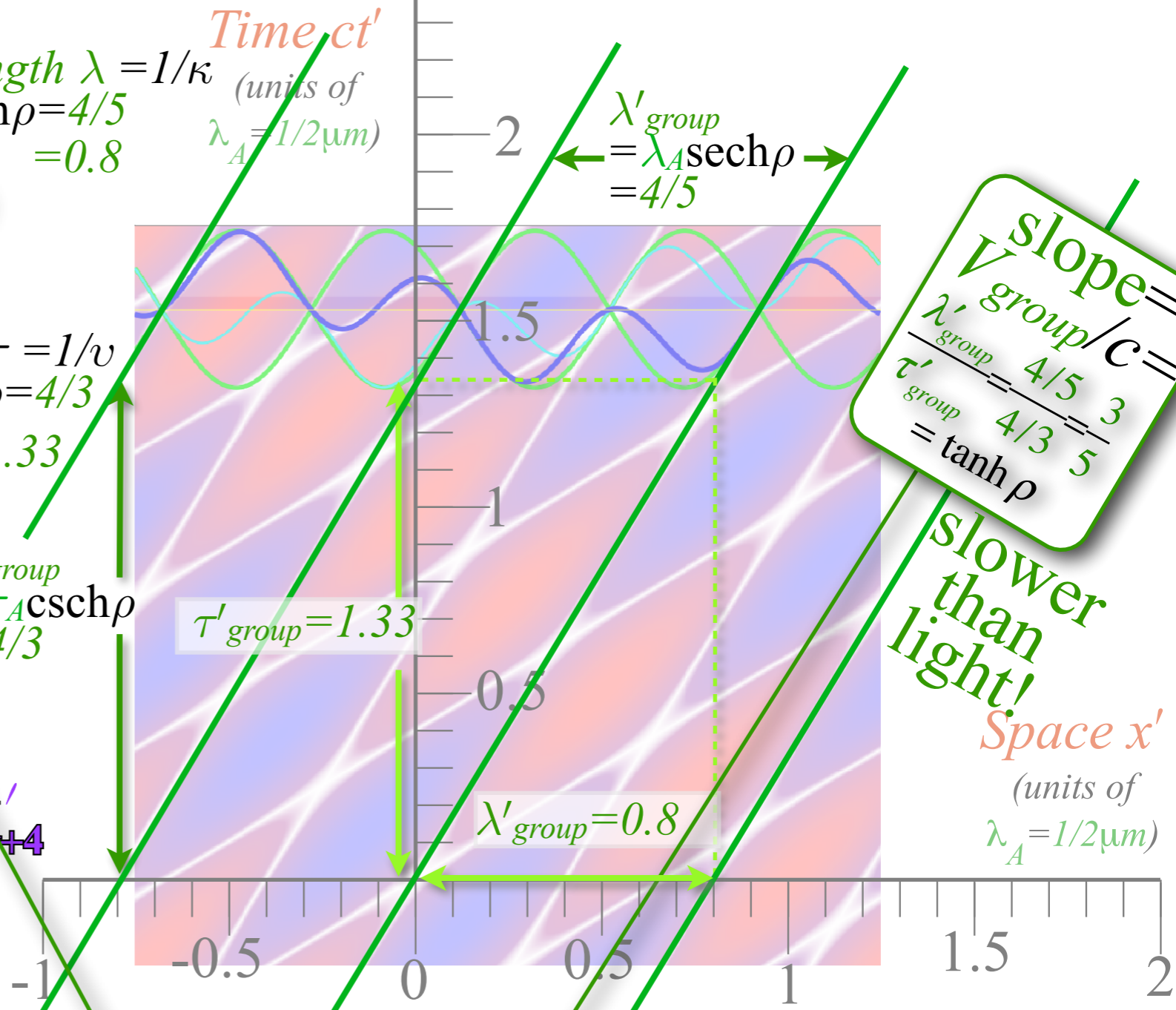
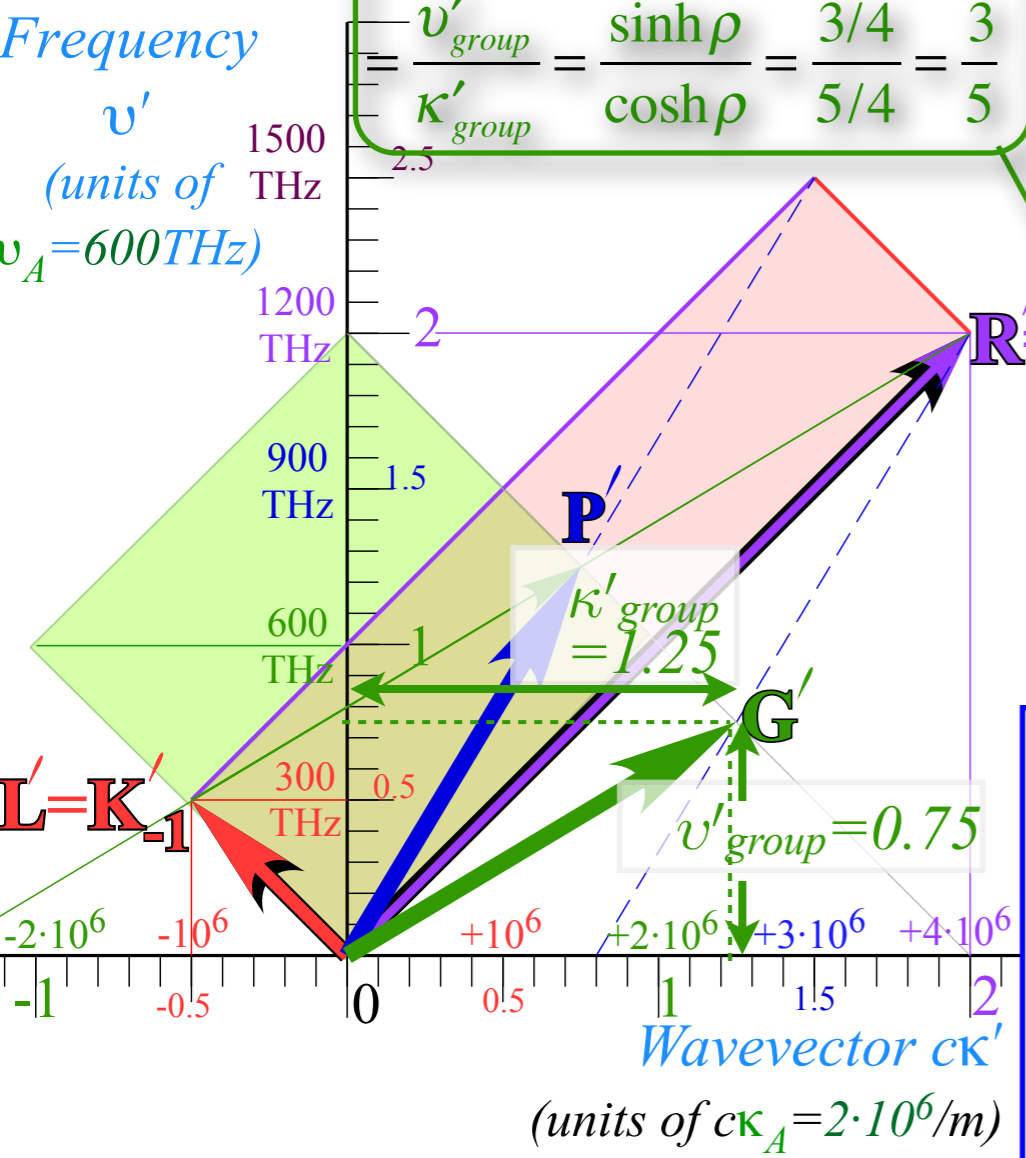
Group wavelength $\lambda = 1/\kappa$ (units of $\lambda_A = 1/2 \mu m$)
 $\lambda'_{group} = \lambda_A \operatorname{sech} \rho = 4/5 = 0.8$

$$\mathbf{G}' = \begin{pmatrix} c\kappa'_{group} \\ v'_{group} \end{pmatrix} = v_A \begin{pmatrix} \cosh \rho \\ \sinh \rho \end{pmatrix} = v_A \begin{pmatrix} 5/4 \\ 3/4 \end{pmatrix}$$

Group frequency
 $v'_{group} = v_A \sinh \rho = 3/4 = 0.75$

flips to Group period $\tau = 1/v$
 $\tau'_{group} = \tau_A \operatorname{csch} \rho = 4/3 = 1.33$

G-slope = V_{group}/c
 $\frac{v'_{group}}{\kappa'_{group}} = \frac{\sinh \rho}{\cosh \rho} = \frac{3/4}{5/4} = \frac{3}{5}$

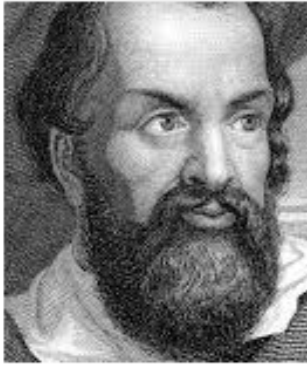


slope = $V_{group}/c = \frac{\lambda'_{group}}{\tau'_{group}} = \frac{4/5}{4/3} = \frac{3}{5} = \tanh \rho$

Slower than light!
 Space x' (units of $\lambda_A = 1/2 \mu m$)

phase	$b_{Doppler RED}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler BLUE}$
group	$\frac{1}{b_{Doppler BLUE}}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

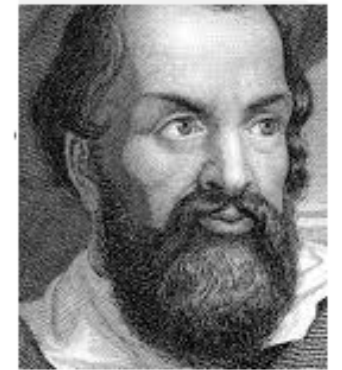
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Lorentz transformations...

write \mathbf{G}' and \mathbf{P}' in terms of \mathbf{G} and \mathbf{P} using $\cosh\rho$ and $\sinh\rho$

$$\mathbf{G}' = \begin{pmatrix} c\mathbf{K}'_{group} \\ v'_{group} \end{pmatrix} = v_A \begin{pmatrix} \cosh\rho \\ \sinh\rho \end{pmatrix} = v_A \begin{pmatrix} 5/4 \\ 3/4 \end{pmatrix}$$

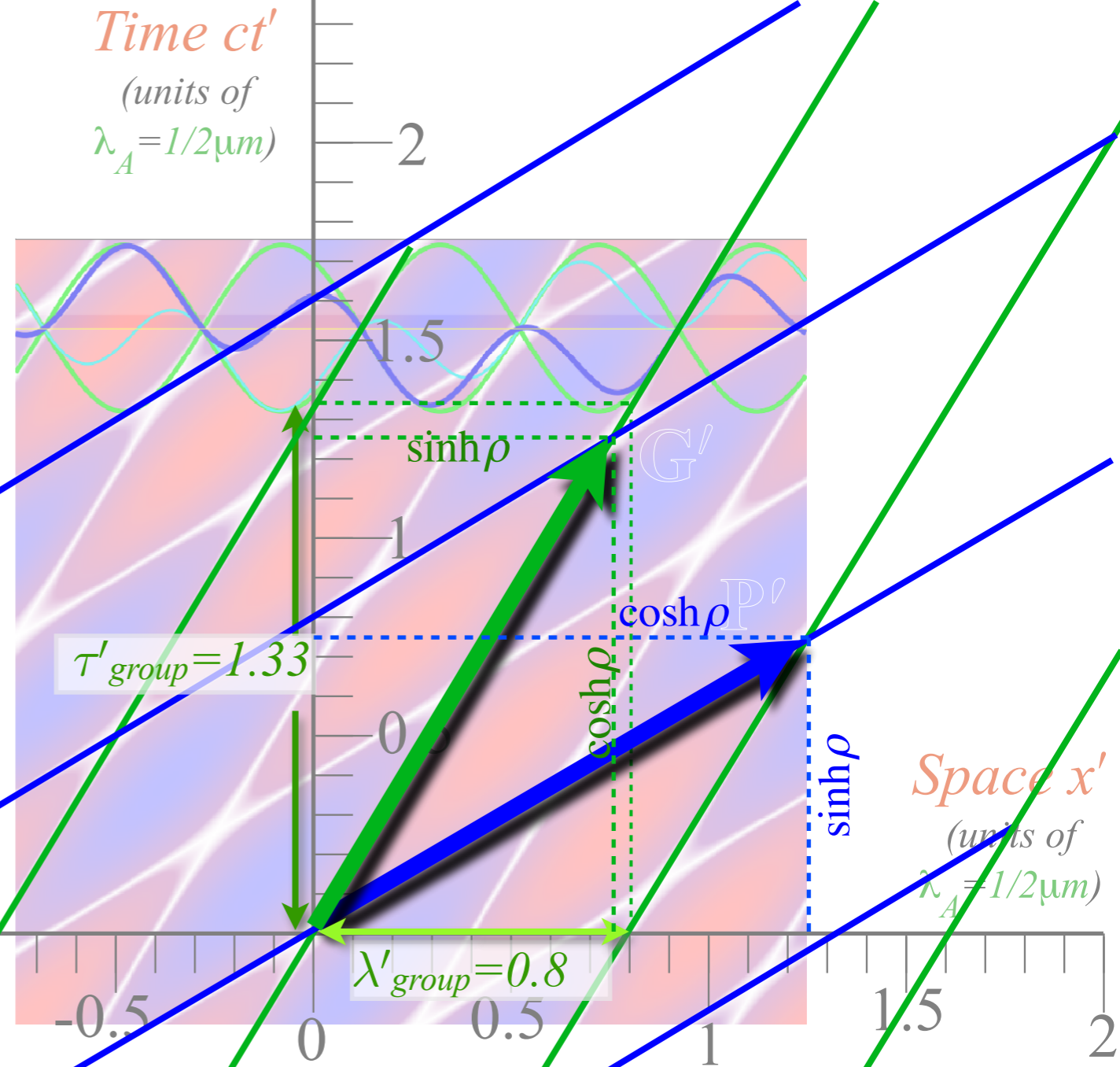
$$= v_A \begin{pmatrix} 1 \\ 0 \end{pmatrix} \cosh\rho + v_A \begin{pmatrix} 0 \\ 1 \end{pmatrix} \sinh\rho$$

$$\mathbf{G}' = \mathbf{G} \cosh\rho + \mathbf{P} \sinh\rho$$

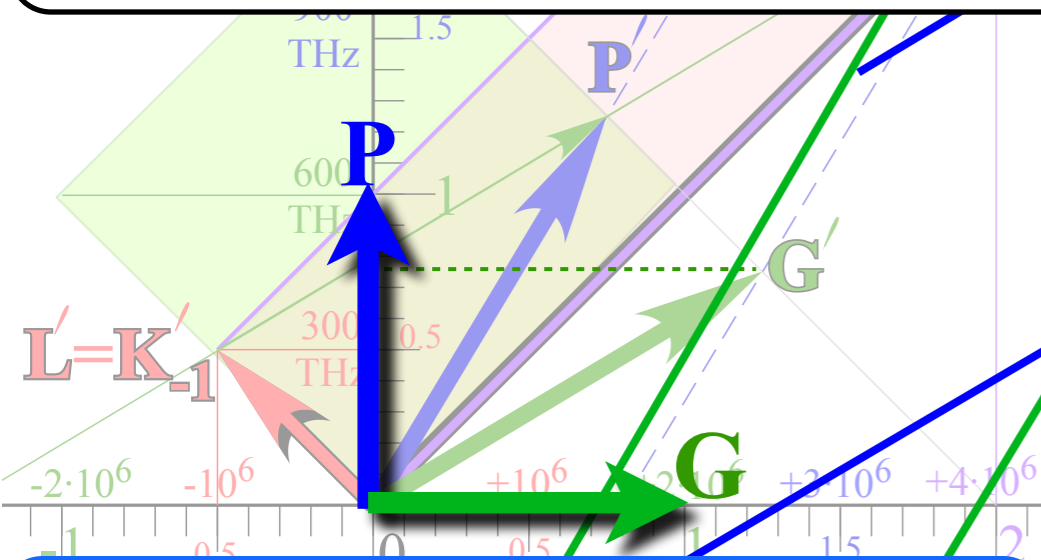
$$\mathbf{P}' = \begin{pmatrix} c\mathbf{K}'_{phase} \\ v'_{phase} \end{pmatrix} = v_A \begin{pmatrix} \sinh\rho \\ \cosh\rho \end{pmatrix} = v_A \begin{pmatrix} 3/4 \\ 5/4 \end{pmatrix}$$

$$= v_A \begin{pmatrix} 1 \\ 0 \end{pmatrix} \sinh\rho + v_A \begin{pmatrix} 0 \\ 1 \end{pmatrix} \cosh\rho$$

$$\mathbf{P}' = \mathbf{G} \sinh\rho + \mathbf{P} \cosh\rho$$



RelaWavity Web Simulation - 16 Relativity Dimensions



$$\begin{pmatrix} \cosh\rho & \sinh\rho \\ \sinh\rho & \cosh\rho \end{pmatrix} \text{ Lorentz transform matrix}$$

phase	$b^{\text{Doppler RED}}$	$\frac{c}{V_{phase}}$	$\frac{\mathbf{K}_{phase}}{\mathbf{K}_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b^{\text{Doppler BLUE}}$
group	$\frac{1}{b^{\text{Doppler BLUE}}}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\mathbf{K}_{group}}{\mathbf{K}_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b^{\text{Doppler RED}}}$
rapidity ρ	$e^{-\rho}$	$\tanh\rho$	$\sinh\rho$	$\text{sech}\rho$	$\cosh\rho$	$\text{csch}\rho$	$\text{coth}\rho$	$e^{+\rho}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

Two Famous-Name Coefficients

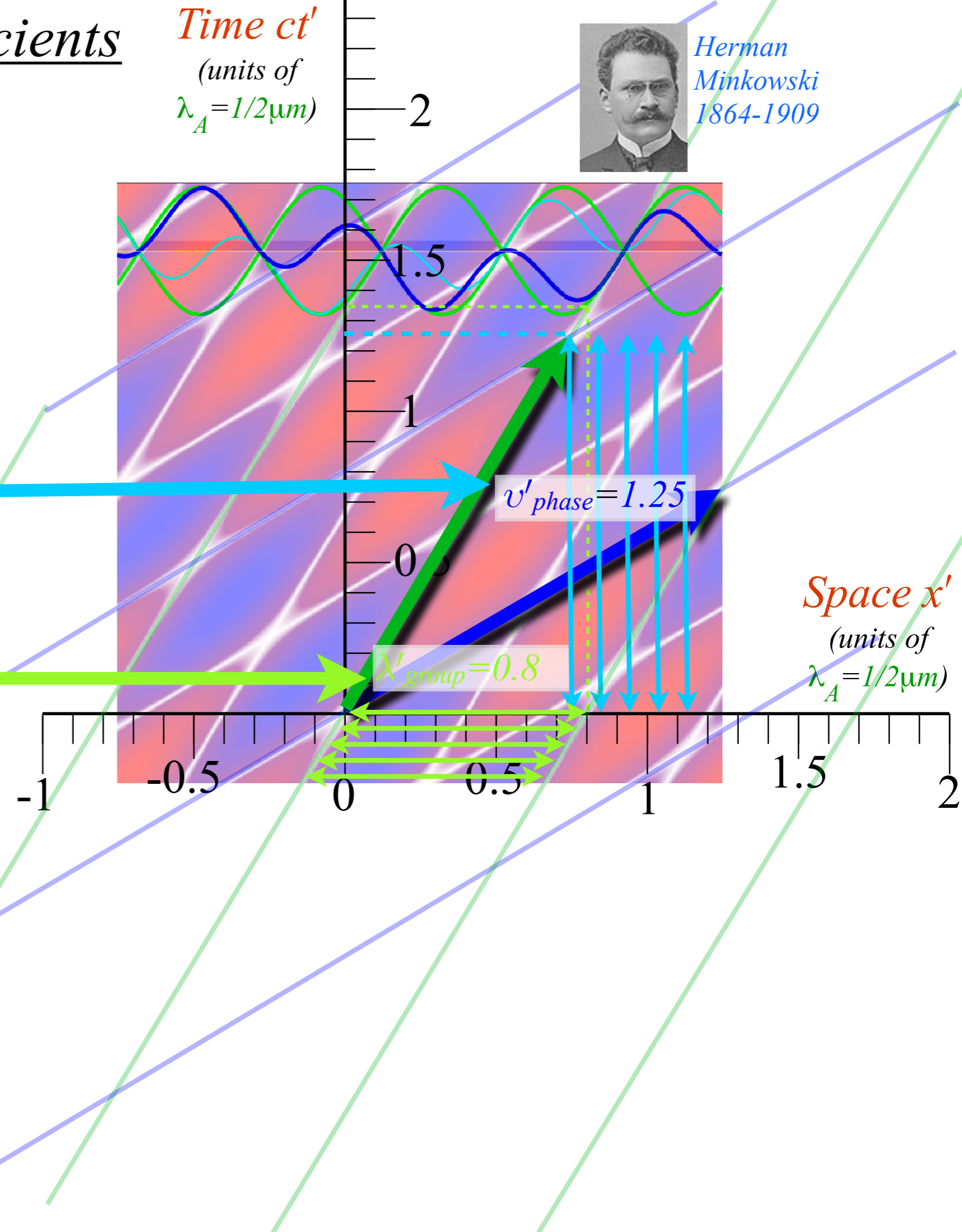
Time ct'
(units of $\lambda_A = 1/2\mu\text{m}$)



Herman Minkowski
1864-1909

This number is called an: **Einstein time-dilation**
(dilated by 25% here)

This number is called a: **Lorentz length-contraction**
(contracted by 20% here)



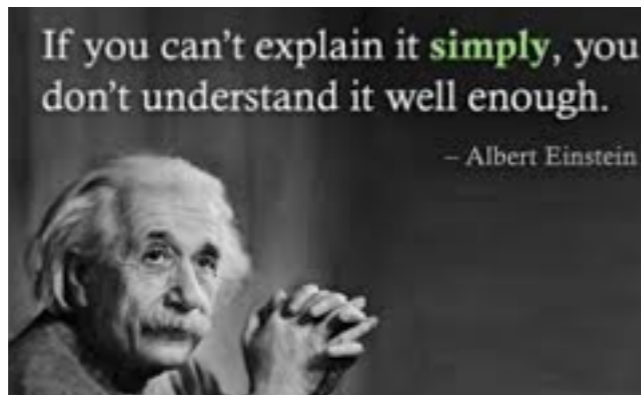
$v'_{\text{phase}} = 1.25$

$v'_{\text{group}} = 0.8$

Space x'
(units of $\lambda_A = 1/2\mu\text{m}$)

Two Famous-Name Coefficients

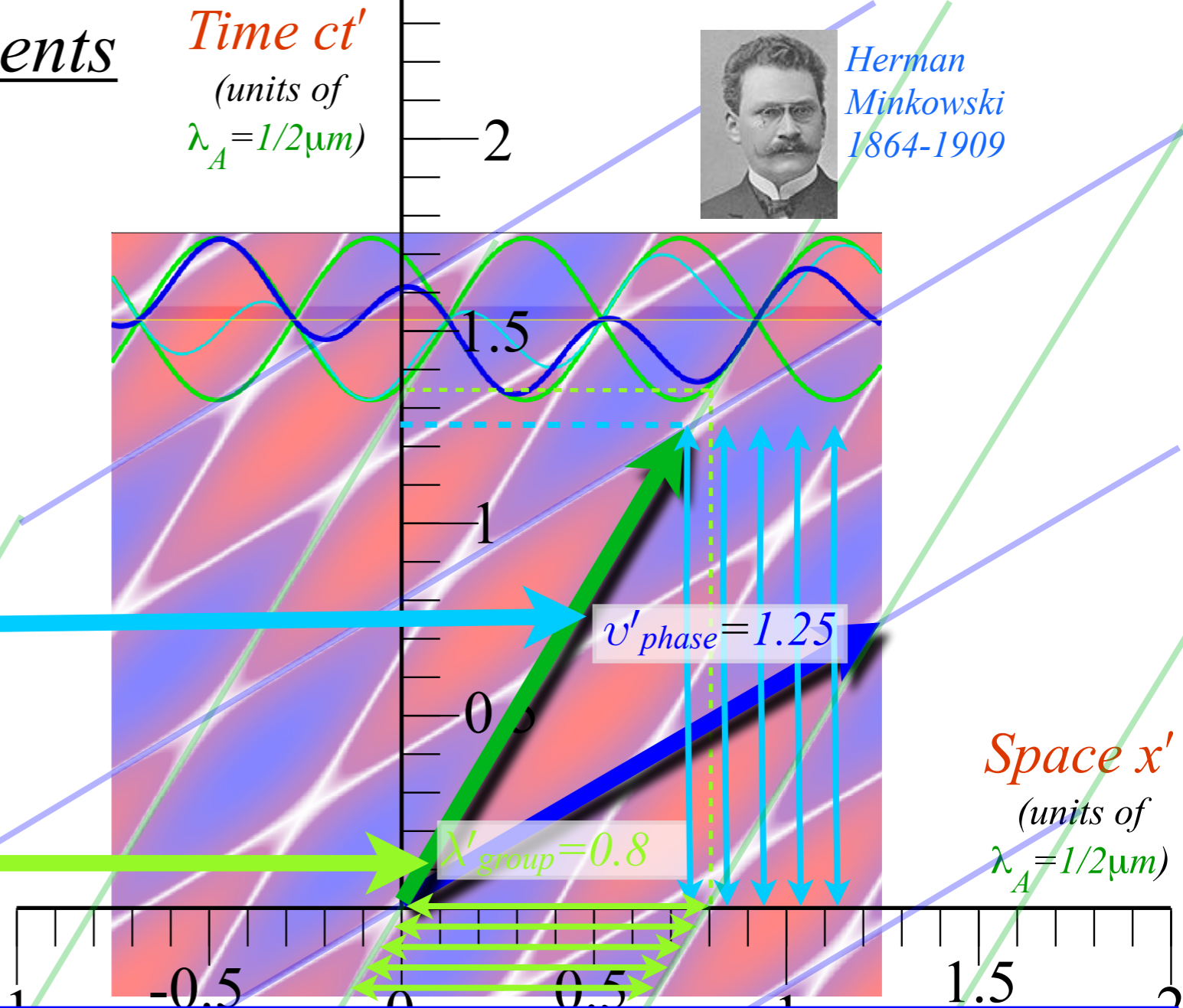
Albert Einstein
1859-1955



Time ct'
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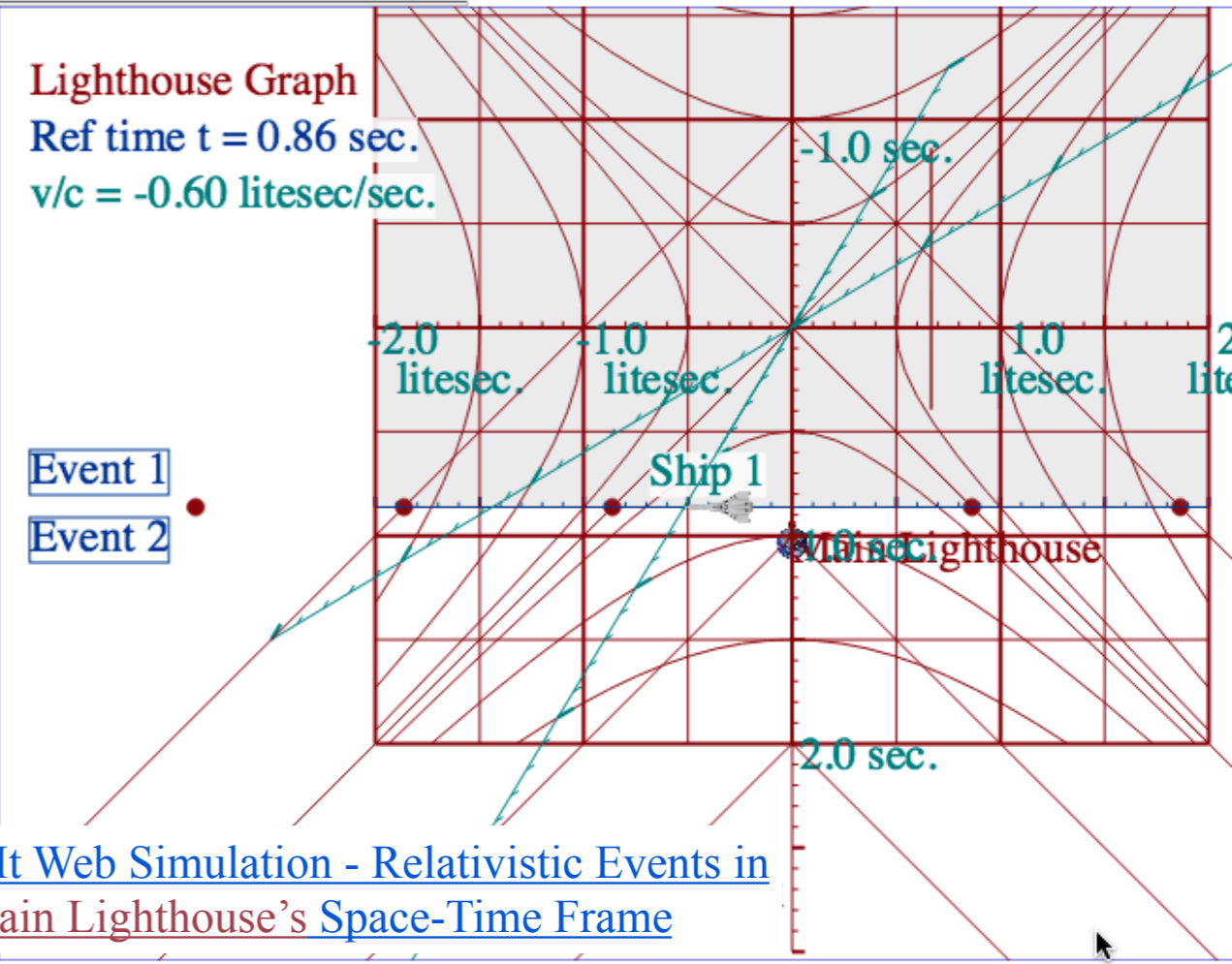
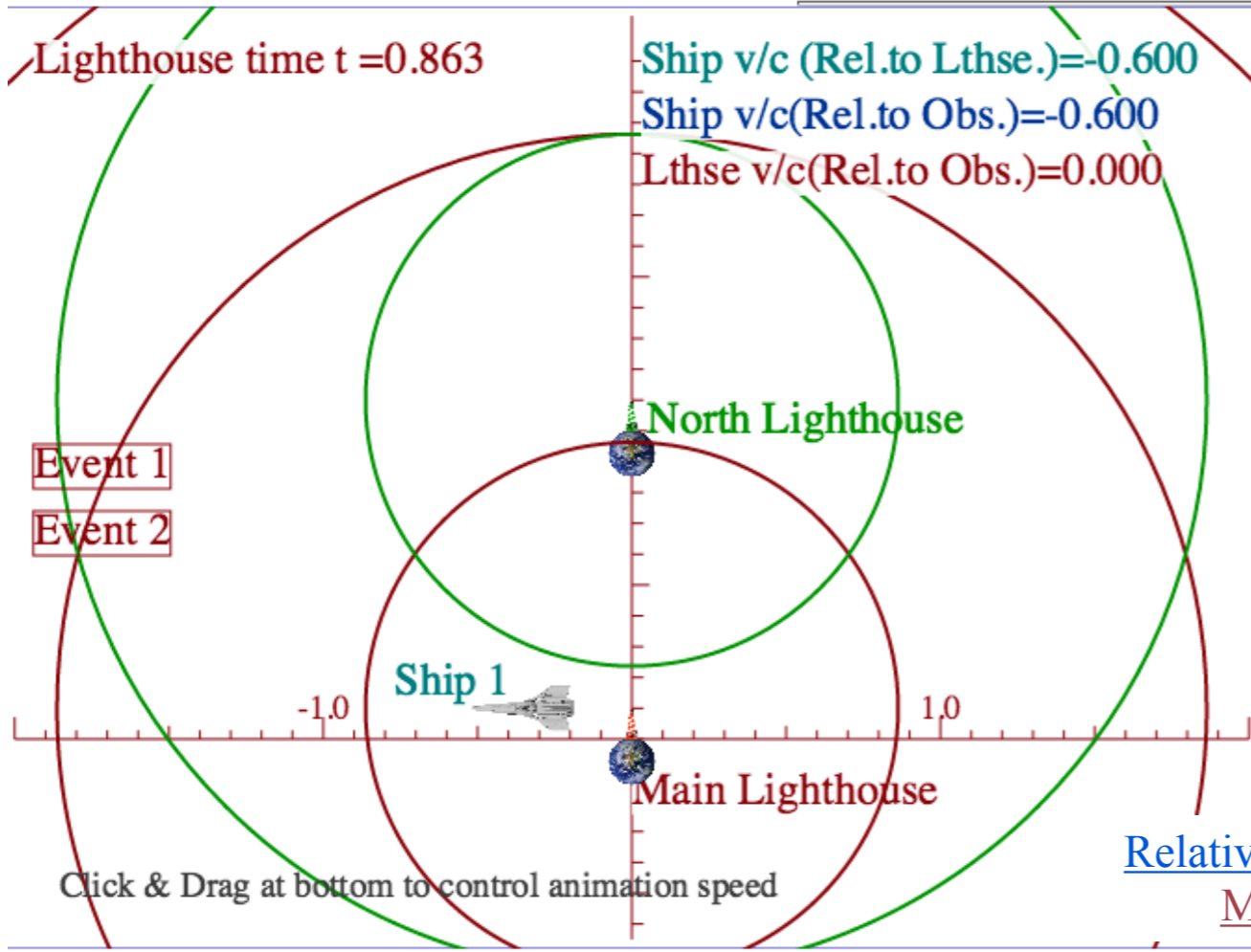
Hendrik A. Lorentz
1853-1928

phase	$b_{RED}^{Doppler}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
group	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\text{coth } \rho$	$e^{+\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

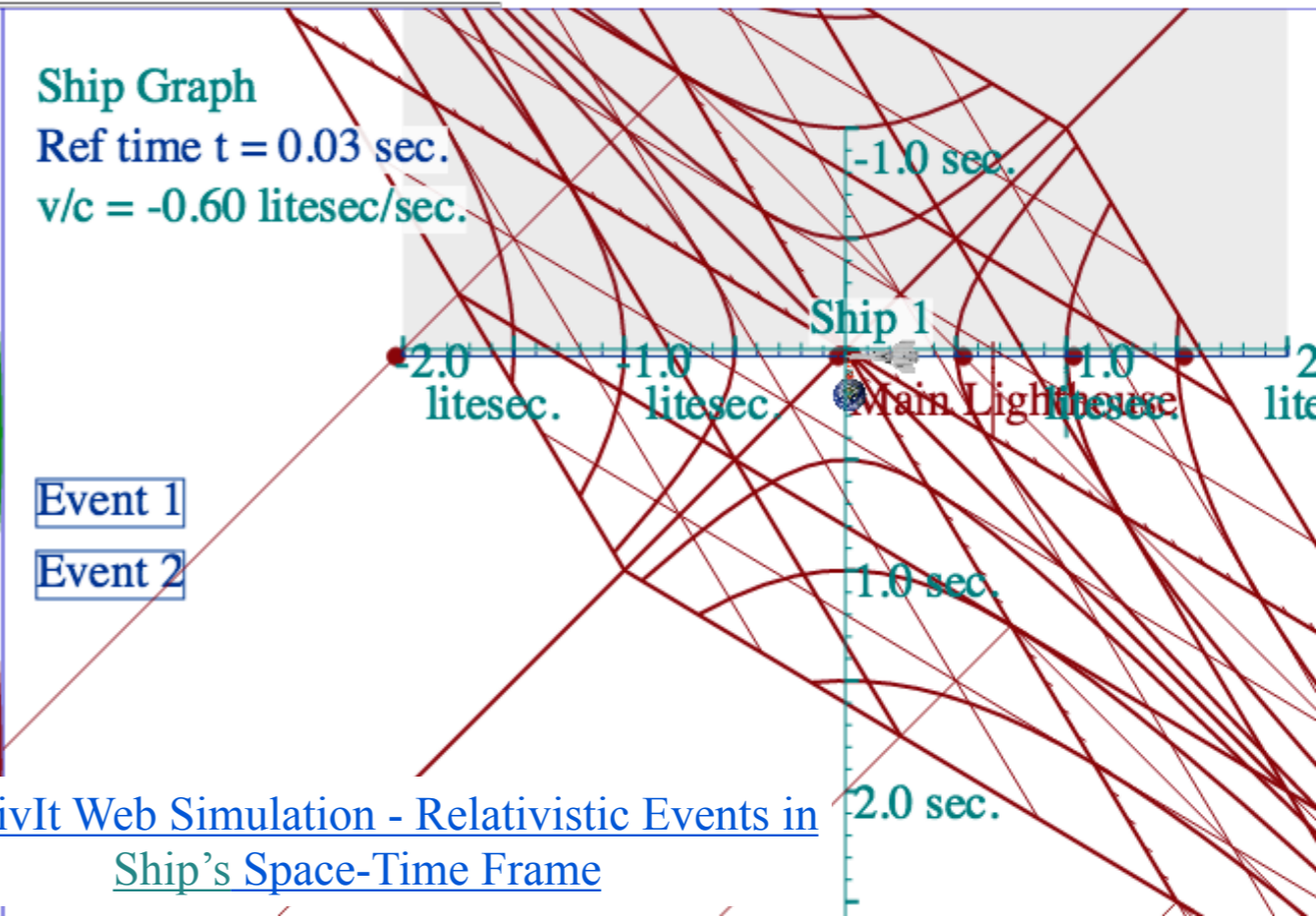
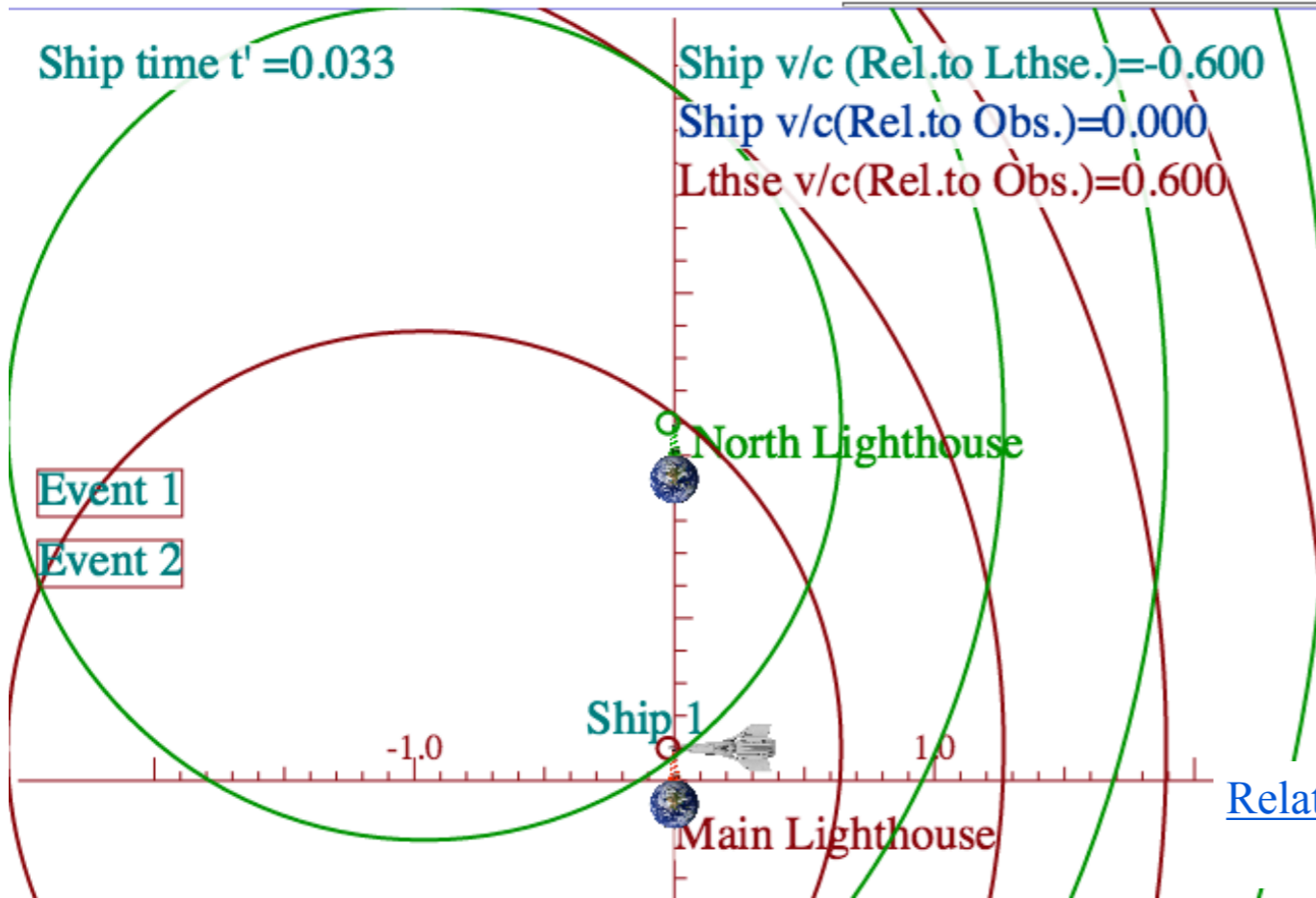
Old-Fashioned Notation

RelaWavity Web Simulation - Relativistic Terms
(Expanded Table)

<i>group</i>	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
<i>phase</i>	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
<i>rapidity</i> ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
<i>stellar</i> ∇ <i>angle</i> σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
<i>value for</i> $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$
<i>effects</i>	$b_{RED}^{Doppler}$	V_{group}	<i>past-future asymmetry</i> (off-diagonal Lorentz-transform)	<i>x-contraction</i> ^(Lorentz) τ_{phase} -contraction	<i>t-dilation</i> ^(Einstein) v_{phase} -dilation (on-diagonal Lorentz-transform)	<i>inverse asymmetry</i>	V_{phase}	$b_{BLUE}^{Doppler}$

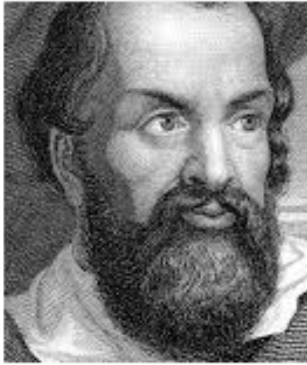


[RelativIt Web Simulation - Relativistic Events in Main Lighthouse's Space-Time Frame](#)



[RelativIt Web Simulation - Relativistic Events in Ship's Space-Time Frame](#)

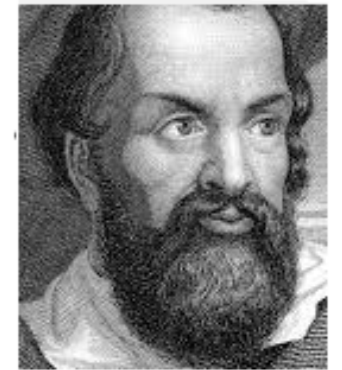
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Doppler Jeopardy

$\nu_R = 600\text{THz}$



$\nu_L = 300\text{THz}$

- (1.) To what velocity u_E must Bob accelerate so he sees beams with equal frequency ω_E ?
- (2.) What is that frequency ω_E ?

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(1.) To what velocity u_E must Bob accelerate so he sees beams with equal frequency ω_E ?

(2.) What is that frequency ω_E ?

Query (1.) has a Jeopardy-style answer-by-question: What is beam group velocity?

$$u_E = V_{group} = \frac{\omega_{group}}{k_{group}} = \frac{\omega_R - \omega_L}{k_R - k_L} = c \frac{\omega_R - \omega_L}{\omega_R + \omega_L}$$

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$$\frac{300}{900}$$

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Query (2.) similarly: What ω_E is blue-shift $b\omega_L$ of ω_L and red-shift ω_R/b of ω_R ?

$$\omega_E = b\omega_L = \omega_R/b \quad \Rightarrow \quad b = \sqrt{\omega_R / \omega_L} \quad \Rightarrow \quad \omega_E = \sqrt{\omega_R \cdot \omega_L}$$

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$$\sqrt{6 \cdot 3} = 3\sqrt{2} = 4.24$$

$$\omega_E = b\omega_L = \omega_R/b \Rightarrow b = \sqrt{\omega_R / \omega_L} \Rightarrow \omega_E = \sqrt{\omega_R \cdot \omega_L}$$

$$\begin{aligned} \omega_E &= \sqrt{\omega_R \cdot \omega_L} \\ &= \sqrt{180000} \\ &= 424 \end{aligned}$$

Geometric mean

Doppler Jeopardy

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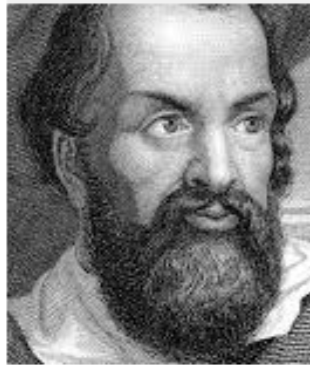
$$\omega_E = b\omega_L = \omega_R/b \Rightarrow b = \sqrt{\omega_R / \omega_L} \Rightarrow \omega_E = \sqrt{\omega_R \cdot \omega_L}$$

$$\omega_E = \sqrt{\omega_R \cdot \omega_L} = \sqrt{180000} = 424$$

V_{group}/c is ratio of difference mean $\omega_{group} = \frac{\omega_R - \omega_L}{2}$ to arithmetic mean $\omega_{phase} = \frac{\omega_R + \omega_L}{2}$. Frequency $\omega_E = B$ is the **geometric mean** $\sqrt{\omega_R \cdot \omega_L}$ of left and right-moving frequencies defining the geometry

Geometric mean

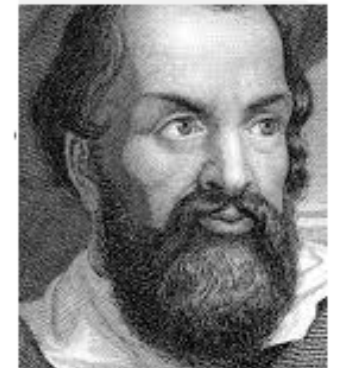
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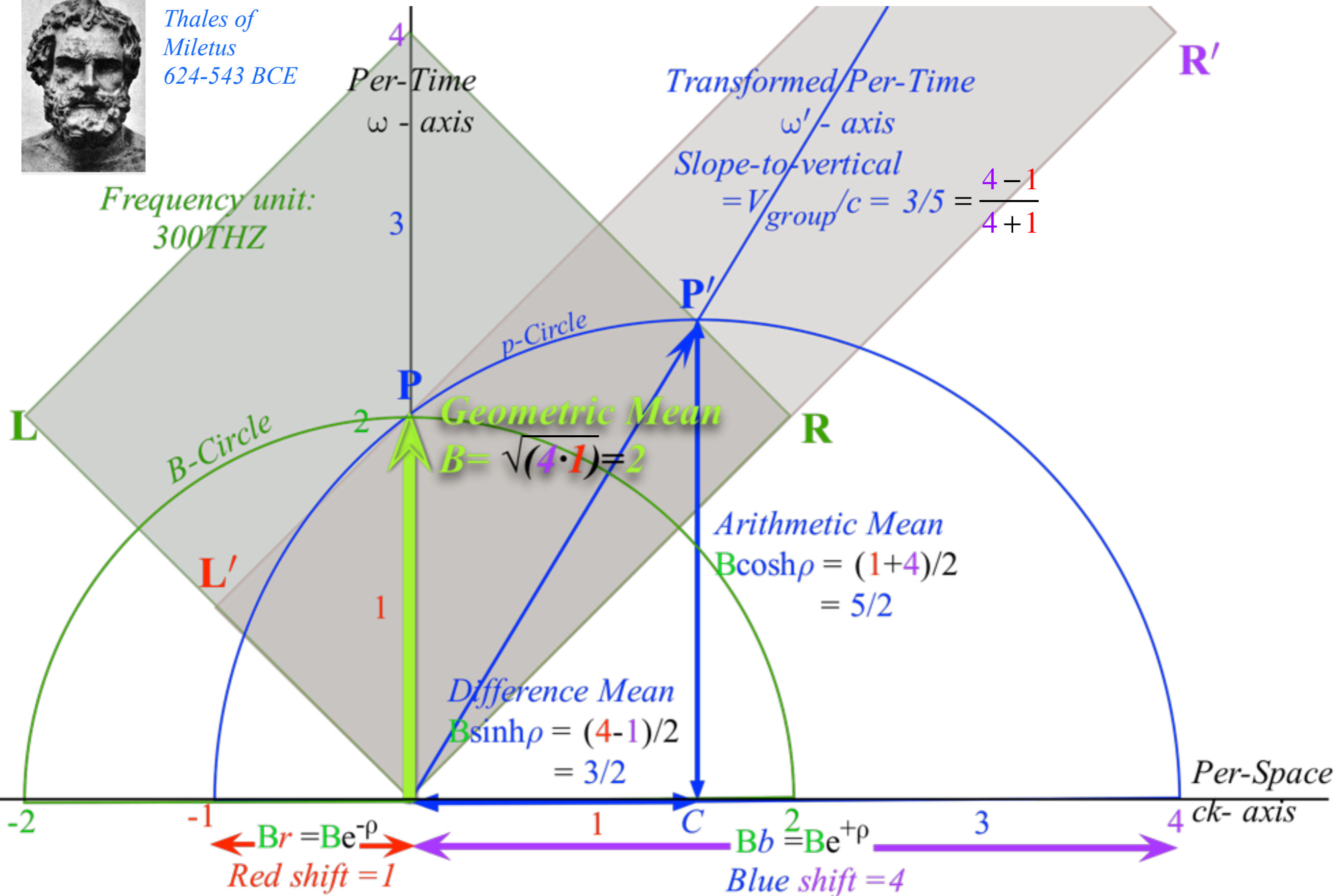
Thales Mean Geometry (600BCE)

helps "Relativity"



Thales of Miletus
624-543 BCE

Frequency unit:
300THZ



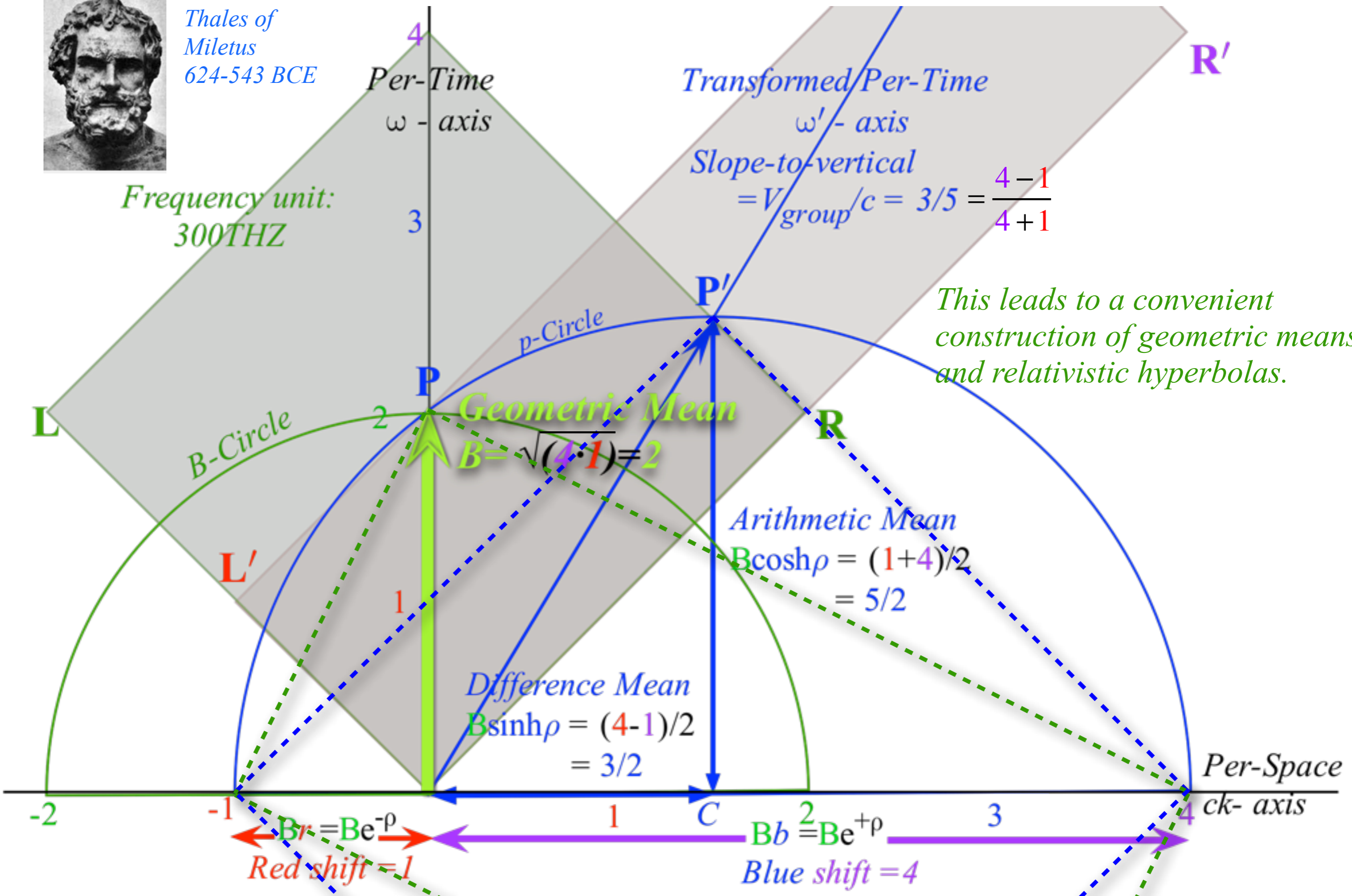
Thales Mean Geometry (600BCE)

helps “Relativity” *Thales showed a circle diameter subtends a right angle with any circle point P*

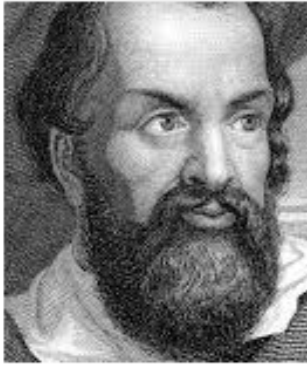


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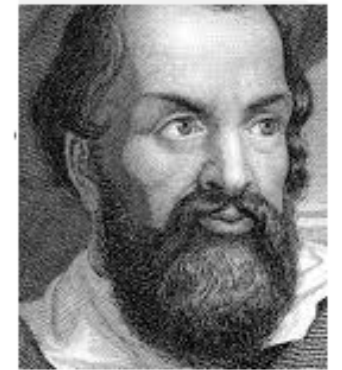
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Thales geometry of Lorentz transformation \rightarrow ...and invariant hyperbolas \leftarrow

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Thales of Miletus
624-543 BCE

Frequency unit:
300THZ

L

B-Circle

L'

-2

-1

$B r = B e^{-\rho}$
Red shift = 1

Per-Time
 ω - axis

4

3

2

1

P

Geometric Mean

$$B = \sqrt{(4 \cdot 1)} = 2$$

Difference Mean

$$B \sinh \rho = (4 - 1) / 2 = 3 / 2$$

1

Transformed/Per-Time

ω' - axis

Slope-to-vertical

$$= V_{\text{group}} / c = 3/5 = \frac{4 - 1}{4 + 1}$$

P'

R

Arithmetic Mean

$$B \cosh \rho = (1 + 4) / 2 = 5 / 2$$

C

$B b = B e^{+\rho}$

Blue shift = 4

equilateral hyperbola
 $r \cdot b = 2$

R'

This leads to a convenient construction of geometric means and relativistic hyperbolas.

Per-Space
ck- axis

3

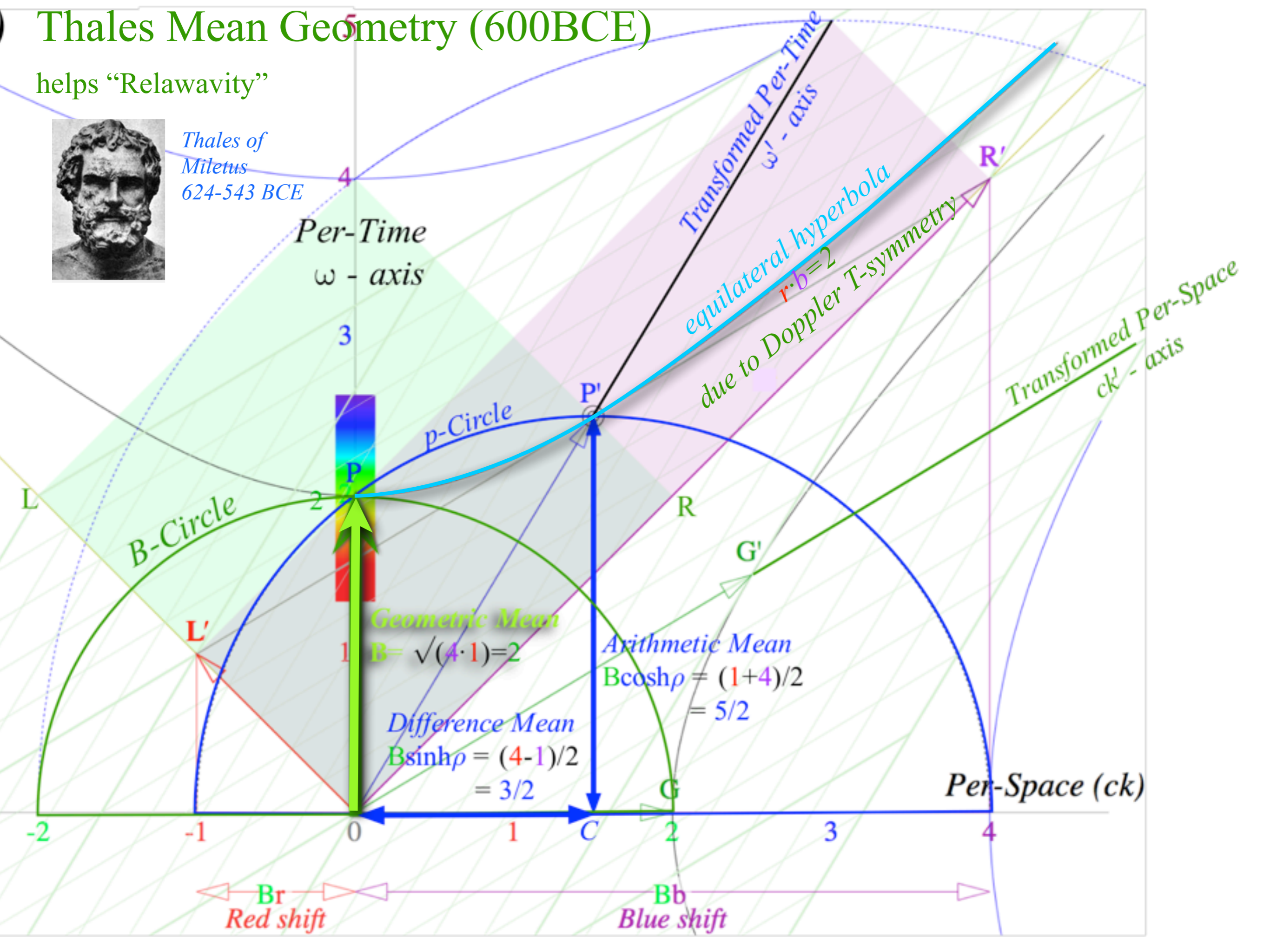
4

Thales Mean Geometry (600BCE)

helps "Relativity"



Thales of Miletus
624-543 BCE



Per-Time
 ω - axis

Transformed Per-Space
 ck' - axis

Transformed Per-Time
 ω' - axis

equilateral hyperbola
due to Doppler T-symmetry
 $r \cdot b = 2$

B-Circle

p-Circle

Geometric Mean
 $B = \sqrt{(4 \cdot 1)} = 2$

Arithmetic Mean
 $B \cosh \rho = (1+4)/2 = 5/2$

Difference Mean
 $B \sinh \rho = (4-1)/2 = 3/2$

B_r
Red shift

B_b
Blue shift

Per-Space (ck)

Per-Time (ω)

Acoustical base frequency = $B = 600\text{Hz}$

Hi freq = 1200.000

Lo freq = 300.000

Laser base frequency = $B = 600\text{THz}$

Doppler blue shift factor = $b = 2.000$

Doppler red shift factor = $r = 0.500$

$q = 0.693$

CW Light Axioms

All colors go c: $\omega/k = c$ or L&R on diagonals

Time Reversal ($r \leftrightarrow b$): $r = 1/b$

$$G' = G \cosh(q) + P \sinh(q)$$

$$P' = G \sinh(q) + P \cosh(q)$$

$$G = G' \cosh(q) - P' \sinh(q)$$

$$P = -G' \sinh(q) + P' \cosh(q)$$

H. sapiens Visual Best = 600THz

$600\text{Hz} = \text{Auditory Base}$

Visual Min = 400THz

$20\text{Hz} = \text{Auditory Min}$

-2.0 -1.5 -1.0 -0.5 0.0 0.5 1.0 1.5 2.0

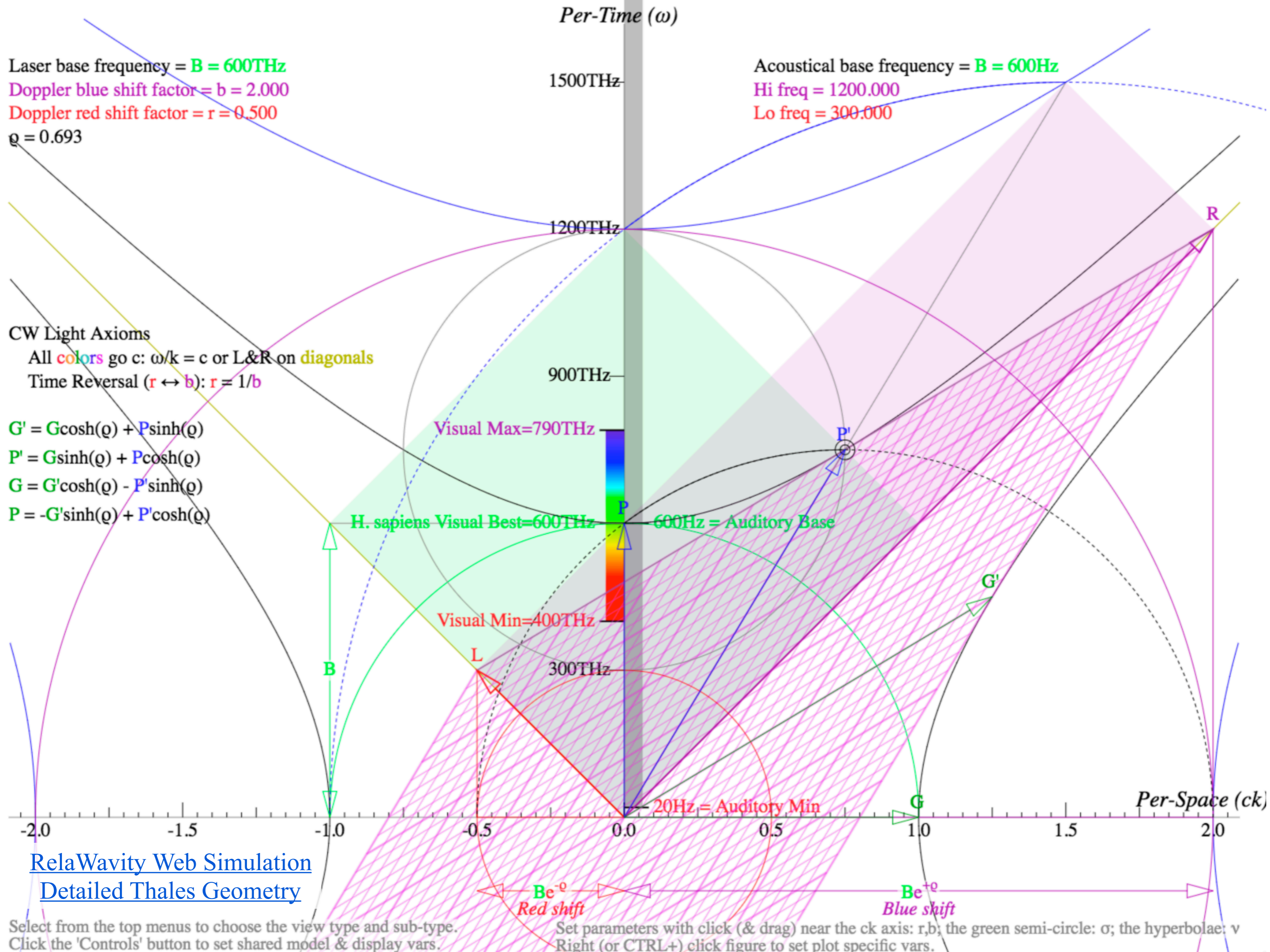
Per-Space (ck)

[RelaWavity Web Simulation](#)
[Detailed Thales Geometry](#)

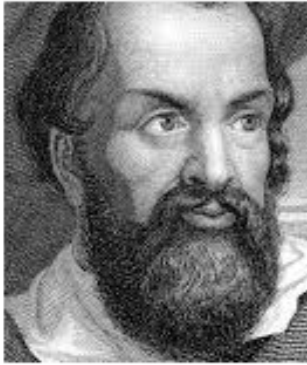
$B e^{-q}$
Red shift

$B e^{+q}$
Blue shift

Select from the top menus to choose the view type and sub-type.
Click the 'Controls' button to set shared model & display vars.
Set parameters with click (& drag) near the ck axis: r,b; the green semi-circle: σ ; the hyperbolae: v
Right (or CTRL+) click figure to set plot specific vars.



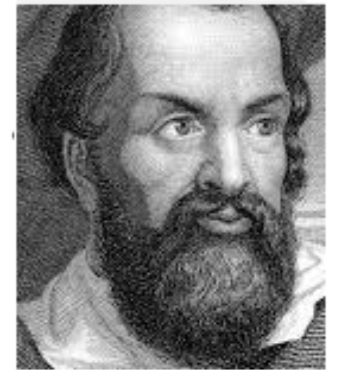
Galileo Galilei



1564-1642

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Galilean velocity*



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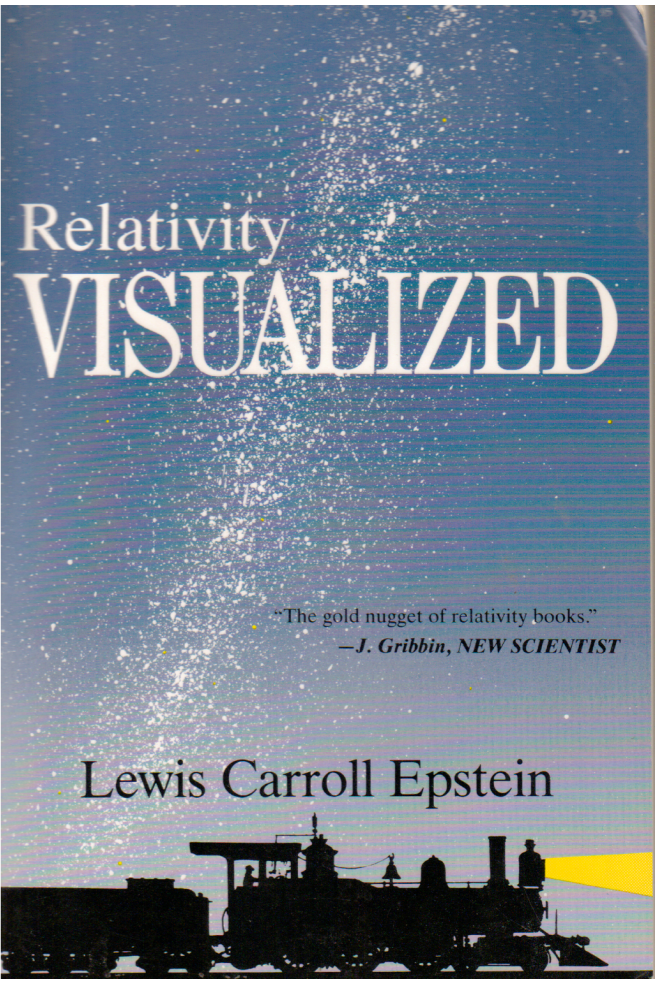
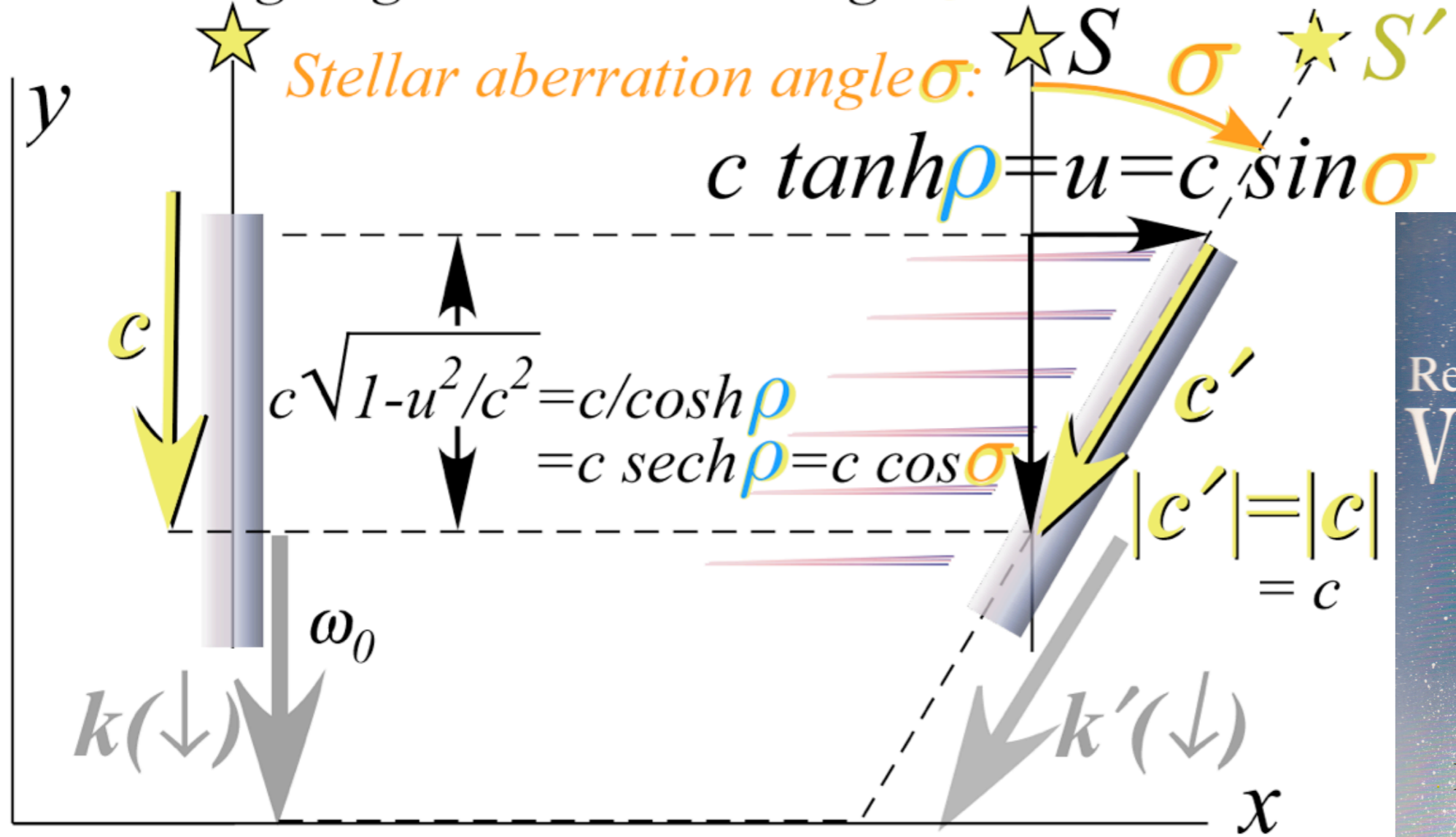
Comparing Longitudinal relativity parameter: Rapidity $\rho = \log_e(\text{Doppler Shift})$

to a Transverse*relativity parameter: Stellar aberration angle σ

*Lewis Carroll Epstein, *Relativitätstheorie*, Birkhäuser, (2004) Earlier English version (1985)-

Observer fixed below star sees it directly overhead.
 Observer going u sees star at angle σ in u direction.

We used notion σ for stellar-ab-angle, (a “flipped-out” ρ). Epstein not interested in ρ analysis or in relation of σ and ρ .



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 World's Greatest Place to Buy Books!

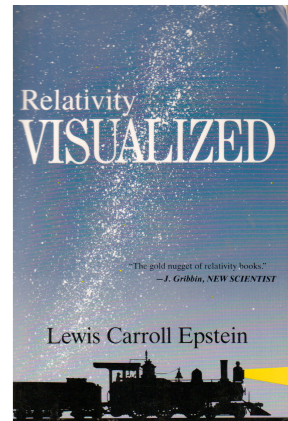
Purchase at:



Comparing Longitudinal relativity parameter: Rapidity $\rho = \log_e(\text{Doppler Shift})$

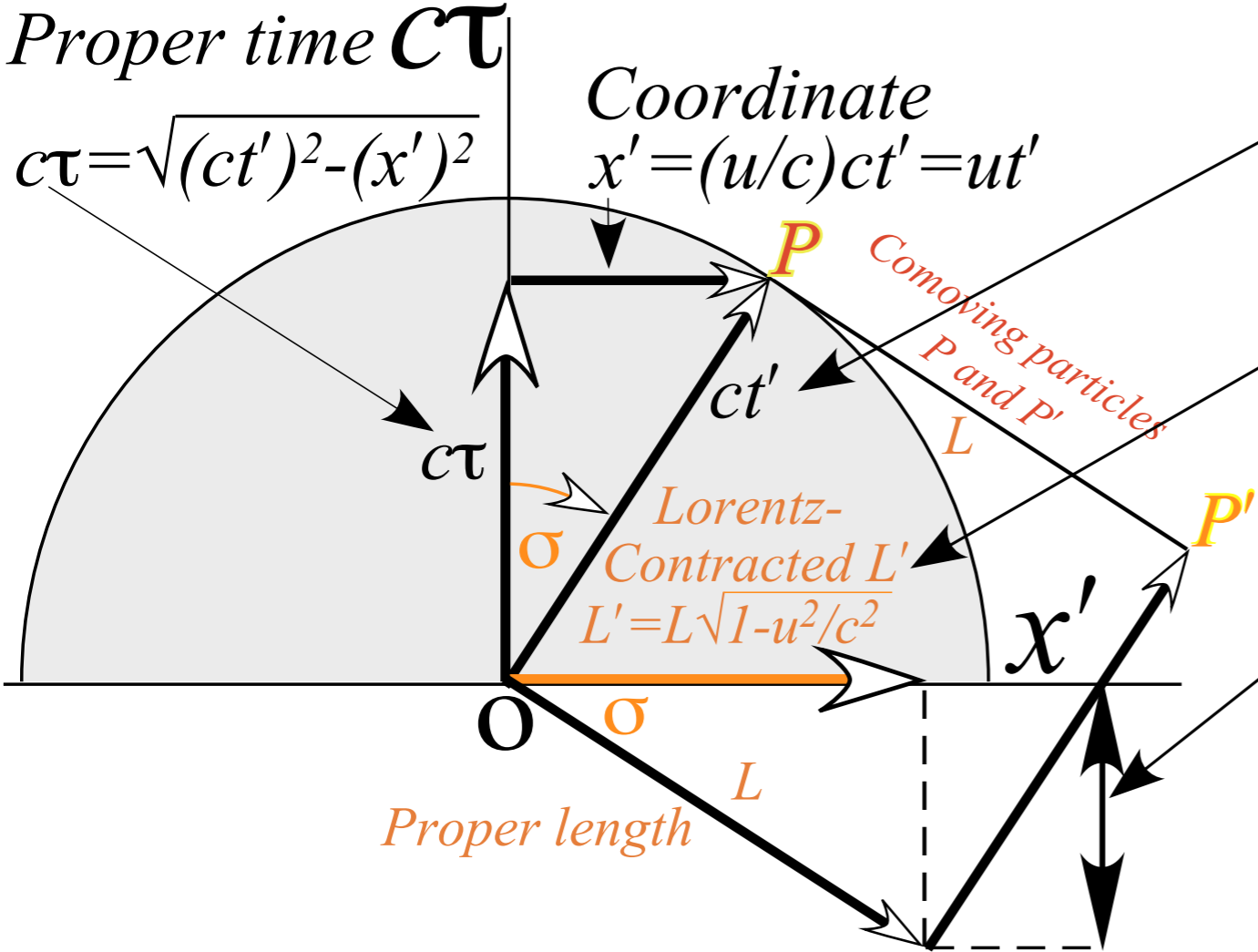
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Proper time $c\tau$ vs. coordinate space x - (L. C. Epstein's "Cosmic Speedometer")

Particles P and P' have speed u in (x', ct') and speed c in $(x, c\tau)$



Einstein time dilation:
 $ct' = c\tau \sec\sigma = c\tau \cosh\rho = c\tau / \sqrt{1-u^2/c^2}$

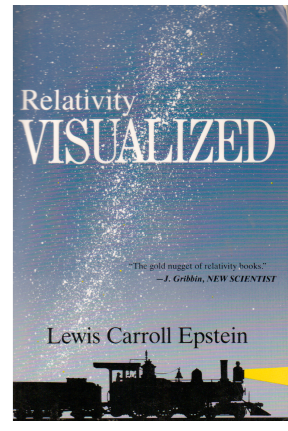
Lorentz length contraction:
 $L' = L \operatorname{sech}\rho = L \cos\sigma = L \cdot \sqrt{1-u^2/c^2}$

Proper Time asimultaneity:
 $c \Delta\tau = L' \sinh\rho = L \cos\sigma \sinh\rho$
 $= L \cos\sigma \tan\sigma$
 $= L \sin\sigma = L / \sqrt{c^2/u^2 - 1} \sim L u/c$

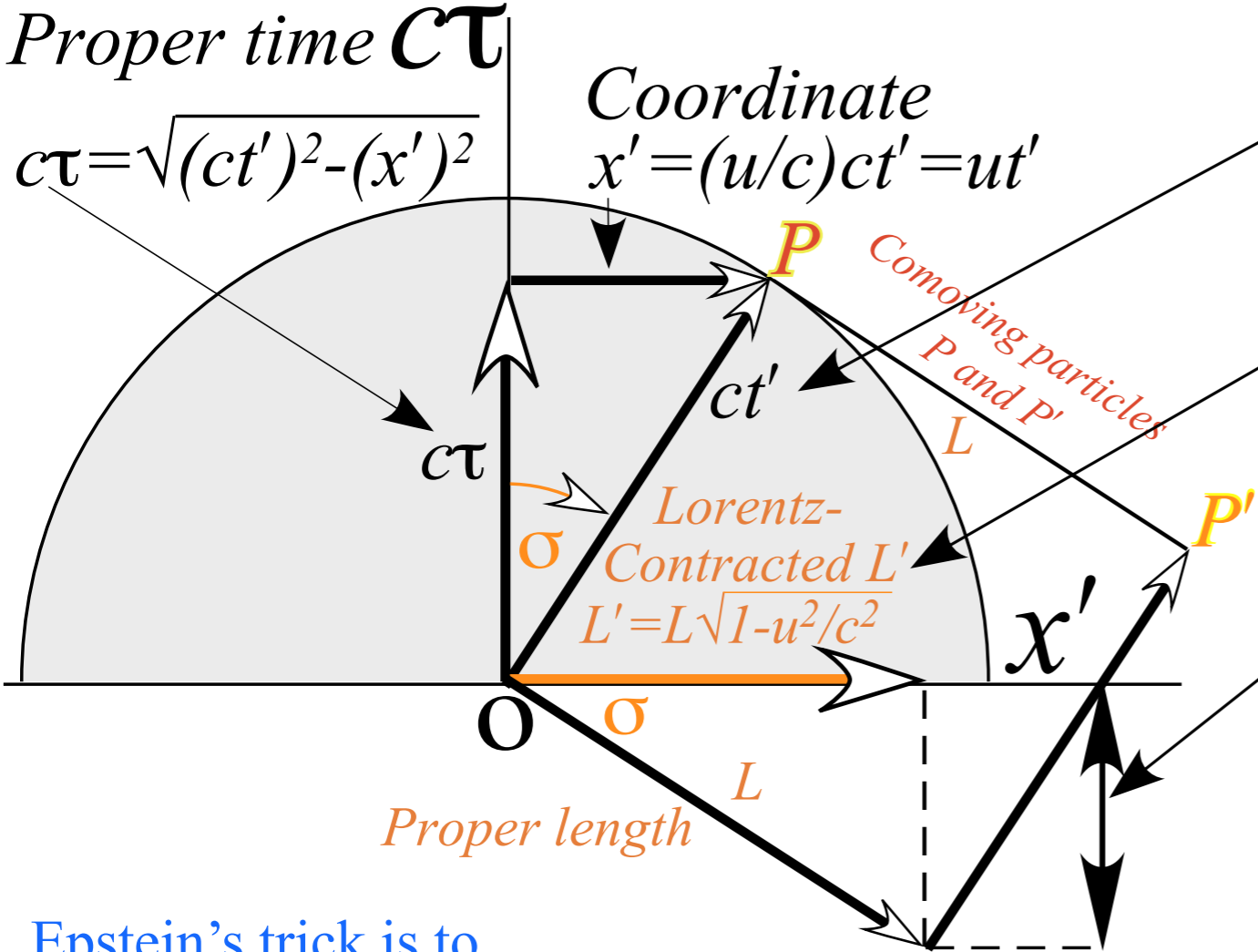
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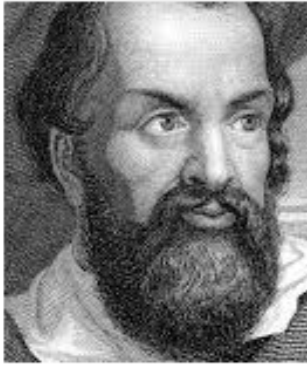
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 $= L \sin\sigma = L / \sqrt{c^2/u^2 - 1} \sim L u/c$

Epstein's trick is to turn a hyperbolic form $c\tau = \sqrt{(ct')^2 - (x')^2}$ into a circular form:

$\sqrt{(c\tau)^2 + (x')^2} = (ct')$

Then everything (and everybody) always goes speed c through $(x', c\tau)$ space!

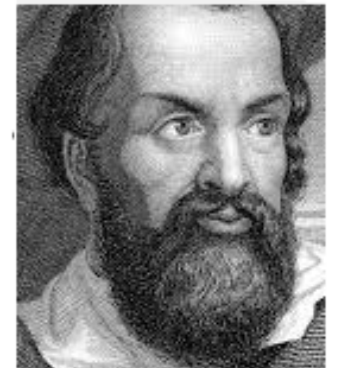
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This map has circle sector arc-area $\sigma = 0.6435$

set to angle $\angle\sigma = 36.87^\circ = 0.6435 \text{radian}$

$$\begin{aligned} \sin(\sigma) &= 0.6000 &= \tanh(\rho) &= 3/5 \\ \tan(\sigma) &= 0.7500 &= \sinh(\rho) &= 3/4 \\ \sec(\sigma) &= 1.2500 &= \cosh(\rho) &= 5/4 \\ \cos(\sigma) &= 0.8000 &= \operatorname{sech}(\rho) &= 4/5 \\ \cot(\sigma) &= 1.3333 &= \operatorname{csch}(\rho) &= 4/3 \\ \csc(\sigma) &= 1.6667 &= \operatorname{coth}(\rho) &= 5/3 \end{aligned}$$

$$\cosh(\rho) + \sinh(\rho) = \frac{5}{4} + \frac{3}{4} = 2.0 = e^{+\rho}$$

$$\cosh(\rho) - \sinh(\rho) = \frac{5}{4} - \frac{3}{4} = 1/2 = e^{-\rho}$$

$$\cosh(\rho) = \frac{e^{+\rho} + e^{-\rho}}{2} \quad \text{Half-Sum-}$$

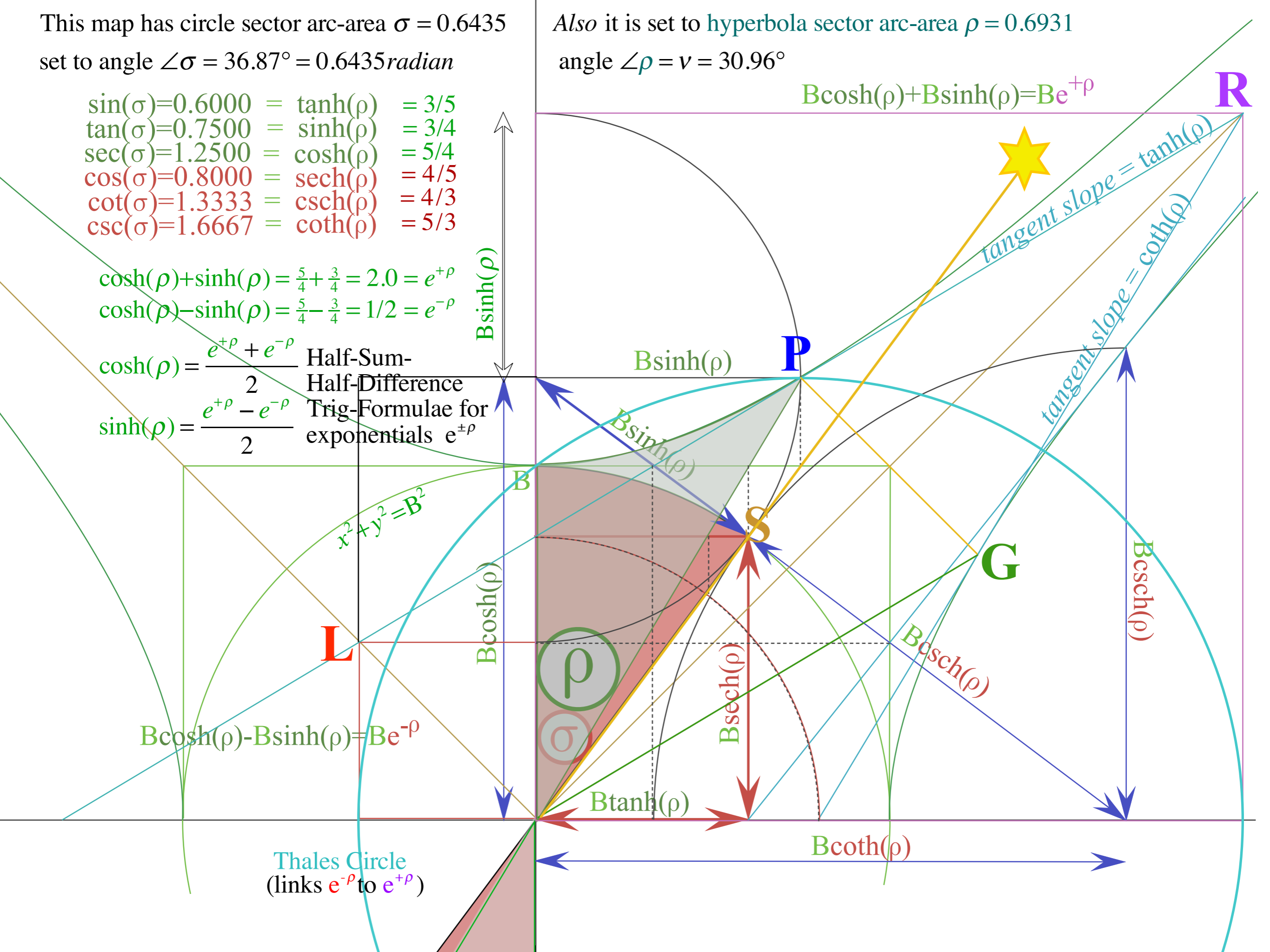
$$\sinh(\rho) = \frac{e^{+\rho} - e^{-\rho}}{2} \quad \text{Half-Difference}$$

Trig-Formulae for
exponentials $e^{\pm\rho}$

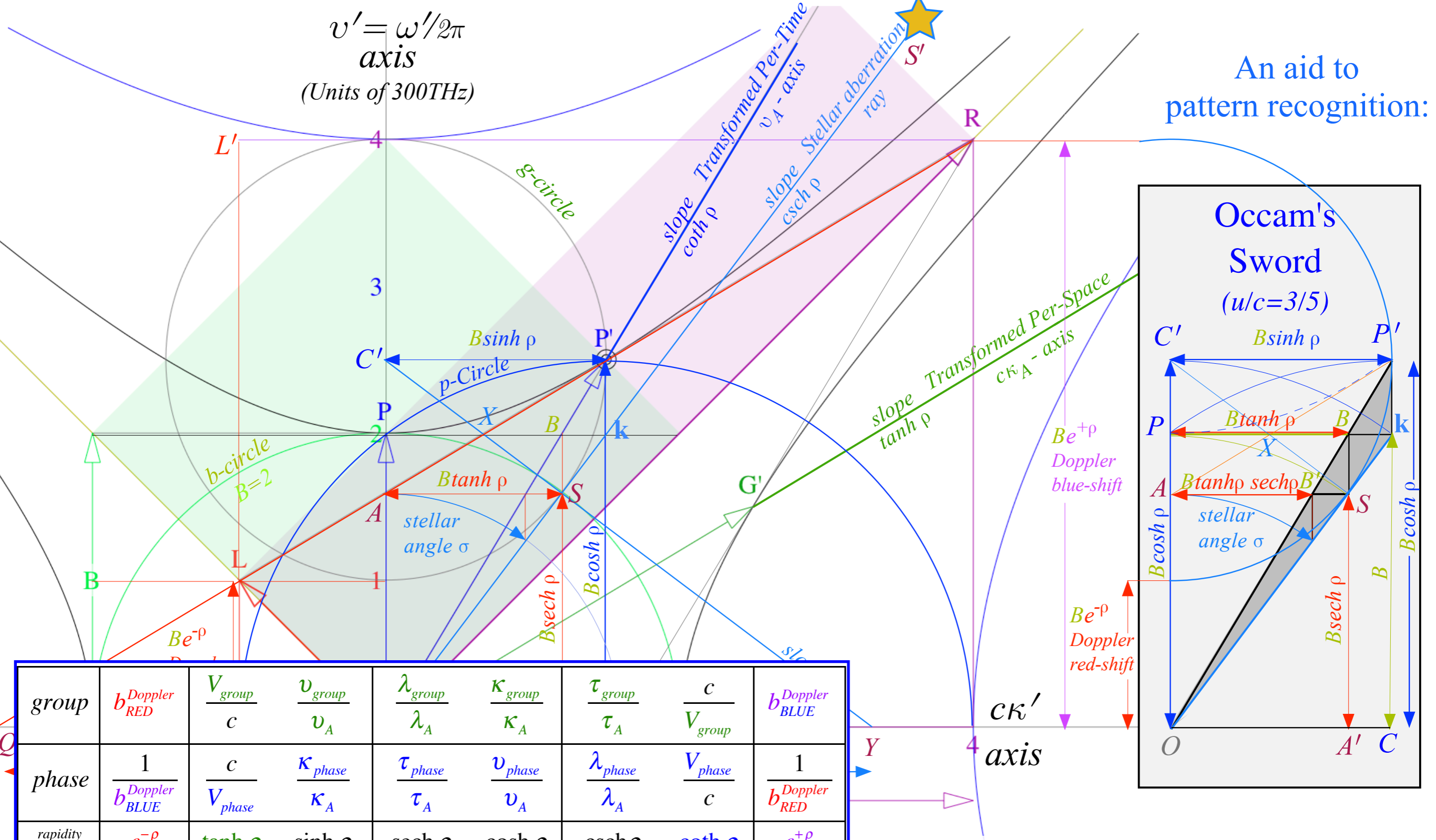
Also it is set to hyperbola sector arc-area $\rho = 0.6931$

angle $\angle\rho = \nu = 30.96^\circ$

$$B\cosh(\rho) + B\sinh(\rho) = B e^{+\rho}$$



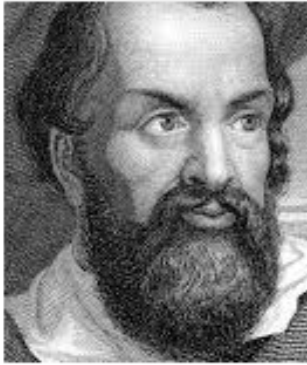
Thales Circle
(links $e^{-\rho}$ to $e^{+\rho}$)



<i>group</i>	$b_{Doppler RED}$	$\frac{V_{group}}{c}$	$\frac{\nu_{group}}{\nu_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$b_{Doppler BLUE}$
<i>phase</i>	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{\nu_{phase}}{\nu_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$\frac{1}{b_{Doppler RED}}$
<i>rapidity</i> ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
<i>stellar</i> ∇ <i>angle</i> σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
<i>value for</i> $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

Table of 12 wave parameters
(includes inverses) for relativity
...and values for $u/c=3/5$
[RelaWavity Web Simulation](#)
[Expanded Relativistic Relations](#)

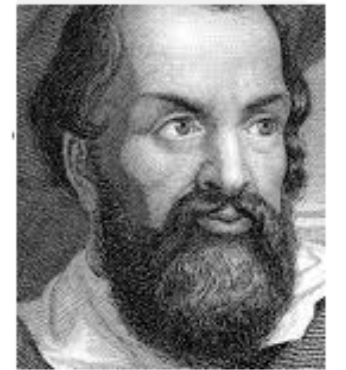
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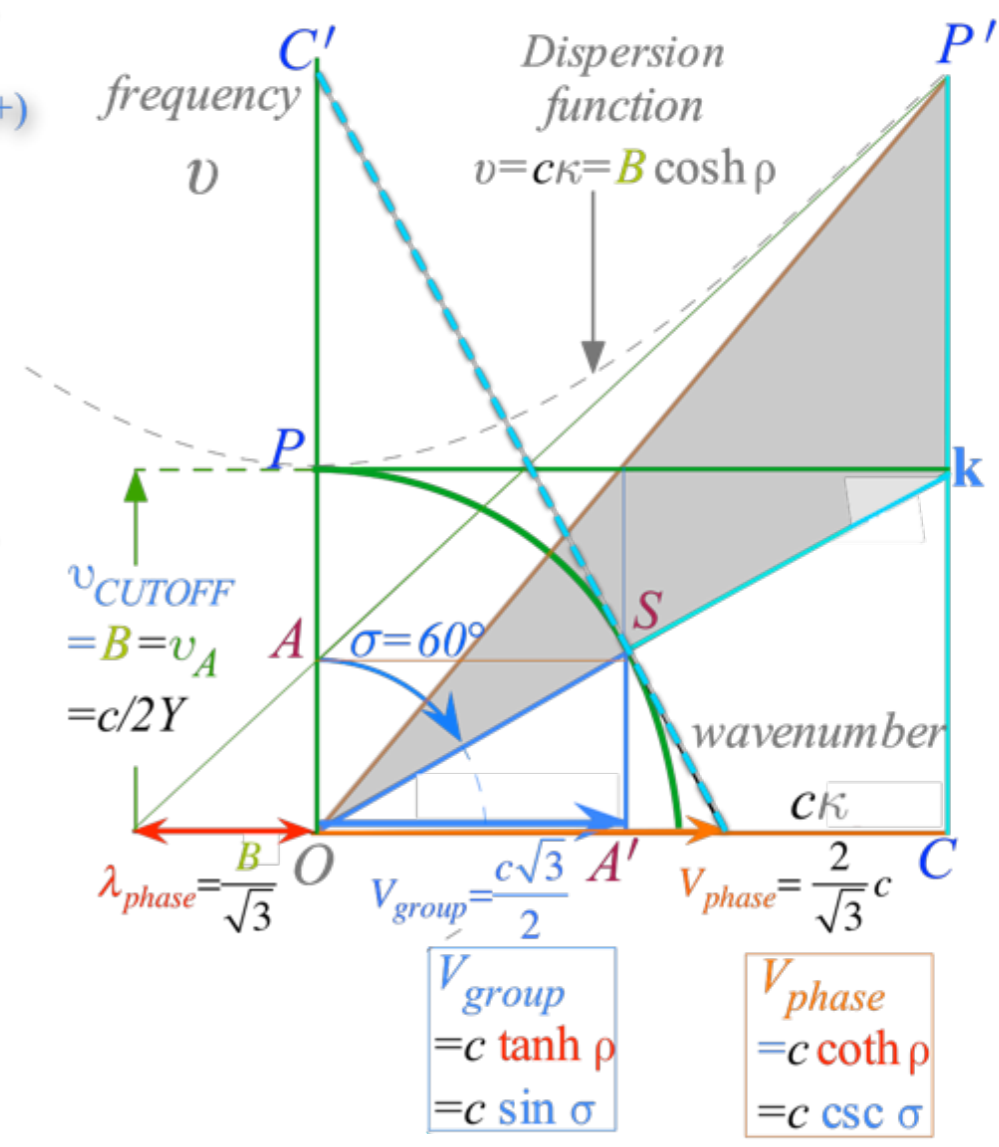
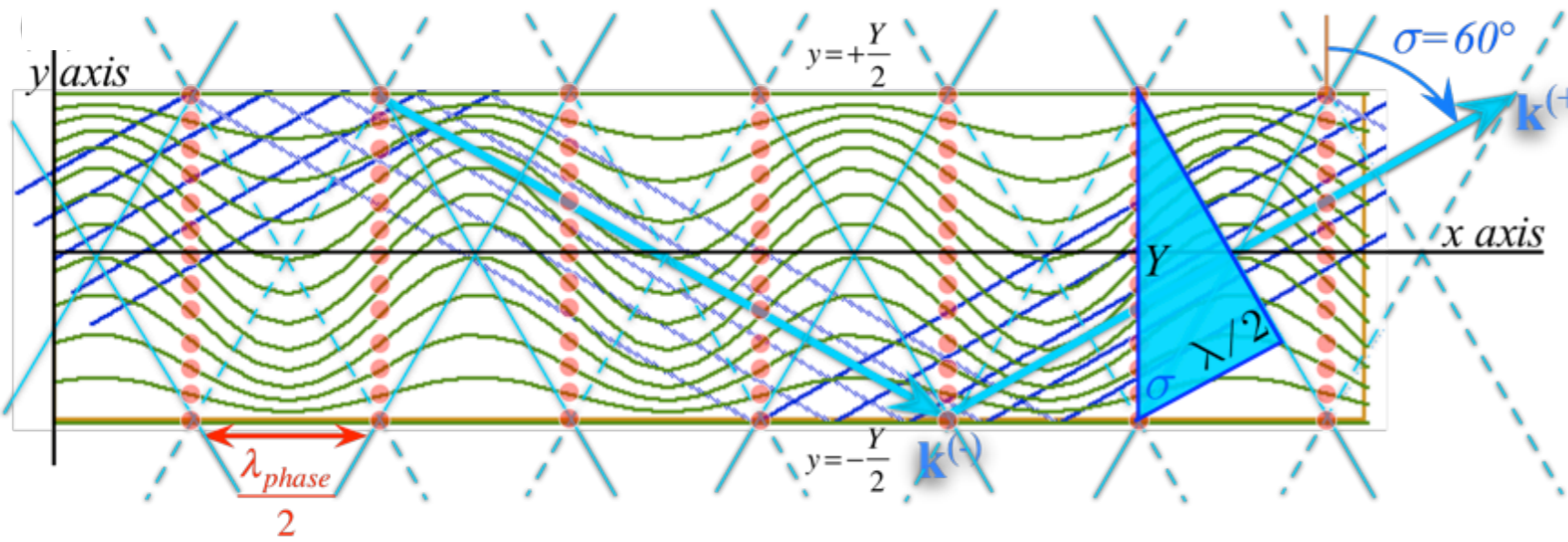
“Occams Sword” and geometry of 16 parameter functions of ρ and σ

➔ Application to TE-Waveguide modes and synchrotron beam relativity ←

Optical wave guide relativistic geometry aided by Occam's Sword

geometry applies to (x,y) space-space
 to (k_x, k_y) per-space-per-space
 to (x, ct) space-time

Relativistic mode with near- c $V_{group}=c/2$ and $V_{phase}=2c$. (Low dispersion.)



KEY:

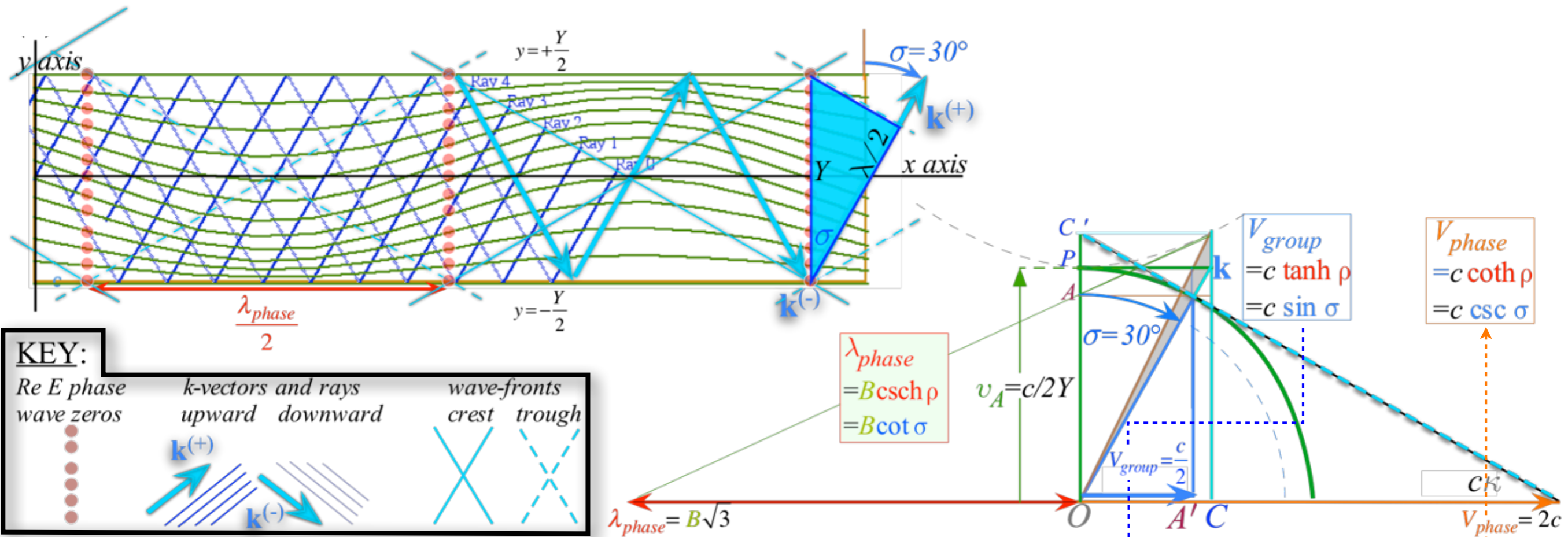
<i>Re E phase</i>	<i>k-vectors and rays</i>	<i>wave-fronts</i>
<i>wave zeros</i>	<i>upward downward</i>	<i>crest trough</i>

$$V_{group} = c \tanh \rho = c \sin \sigma$$

$$V_{phase} = c \coth \rho = c \csc \sigma$$

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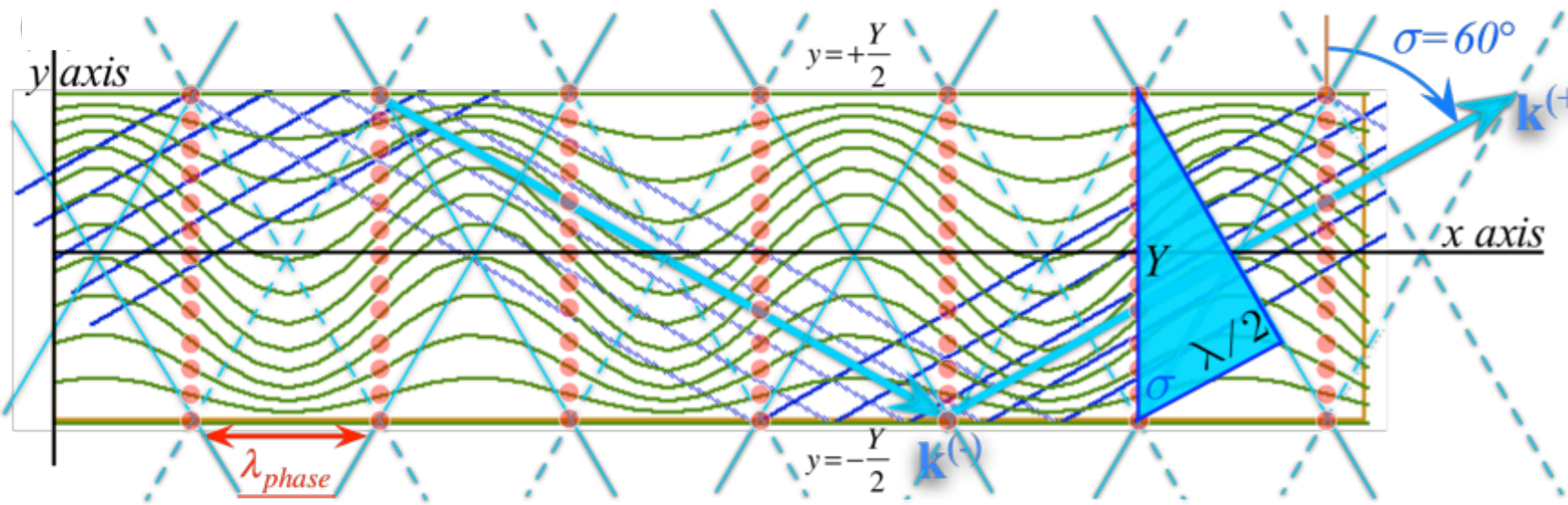


Example of near-cut-off mode with low $V_{group} = c/2$ and high $V_{phase} = 2c$. (High dispersion.)

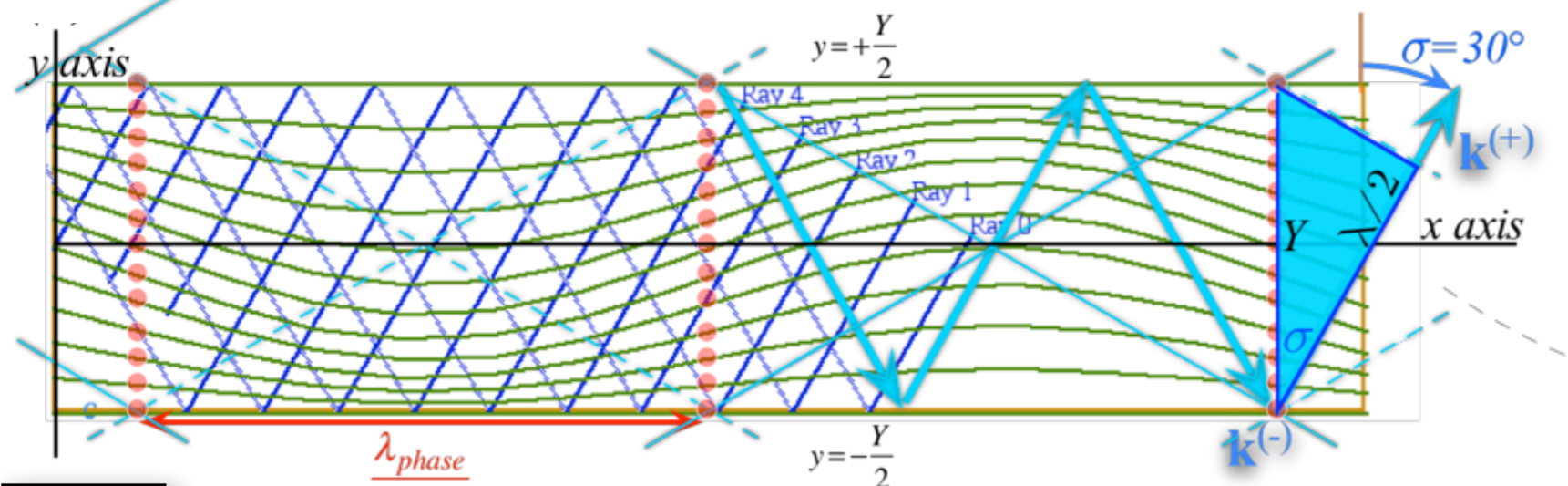
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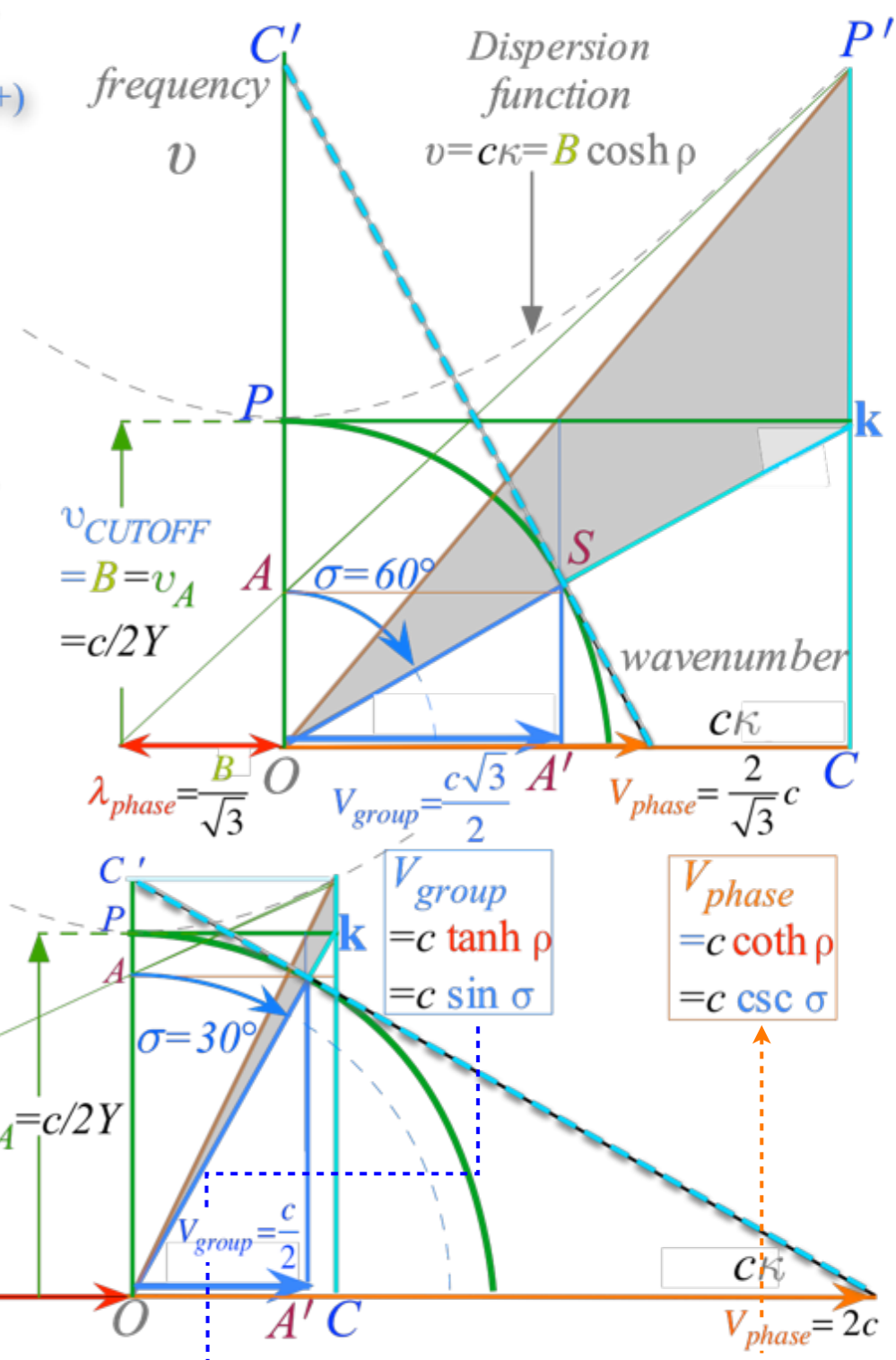
GuideIt Web Simulation: $\sigma = 60^\circ$



GuideIt Web Simulation: $\sigma = 30^\circ$

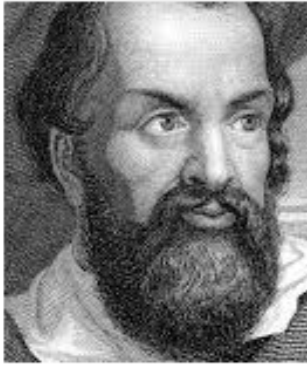
KEY:

Re E phase wave zeros	k -vectors and rays upward downward	wave-fronts crest trough



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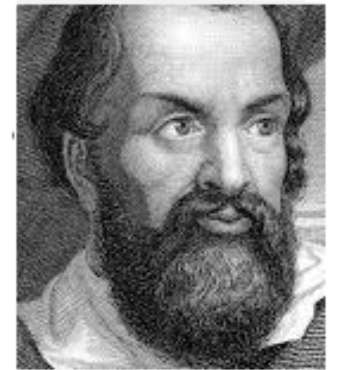
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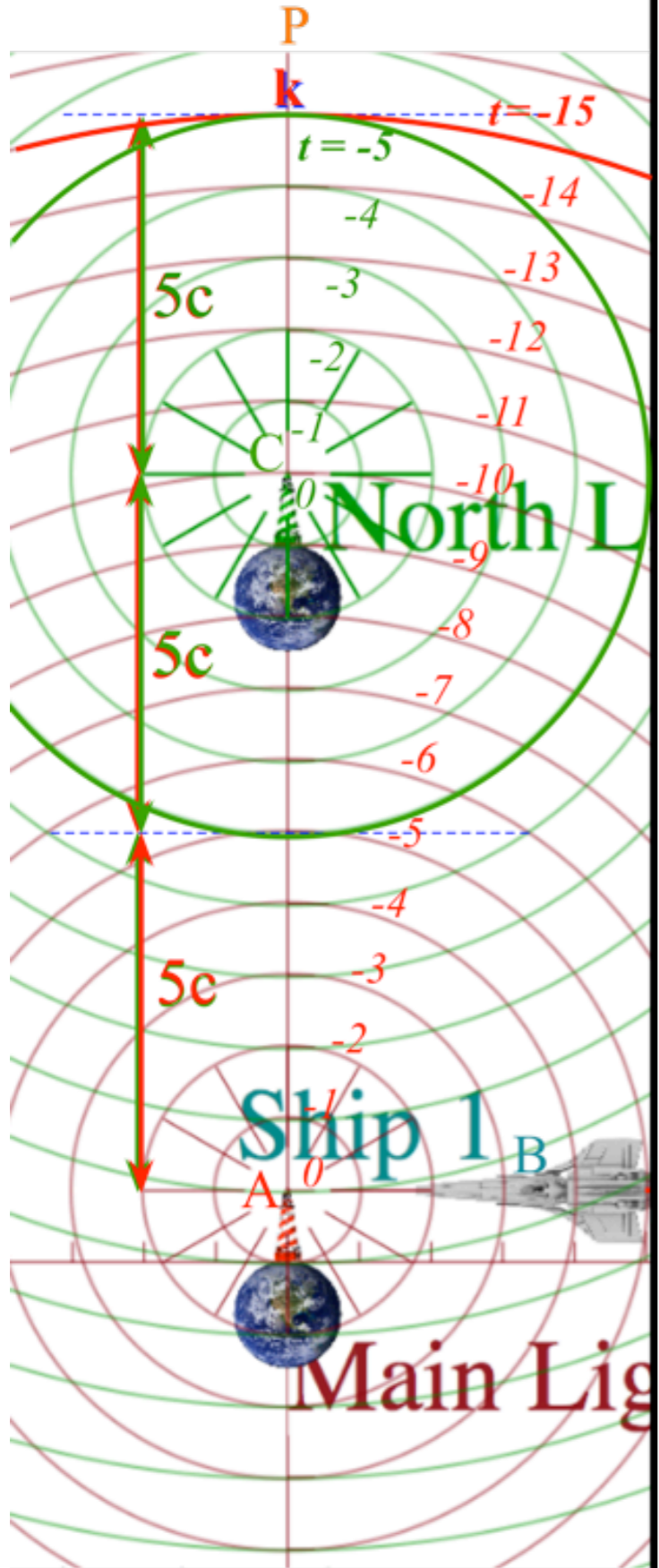
“Occams Sword” and geometry of 16 parameter functions of ρ and σ

Application to TE-Waveguide modes \rightarrow and synchrotron beam relativity \leftarrow

(a) Spherical wave pair
In Alice-Carla frame

Spherical wave relativistic geometry

Also, aided by Occam's Sword

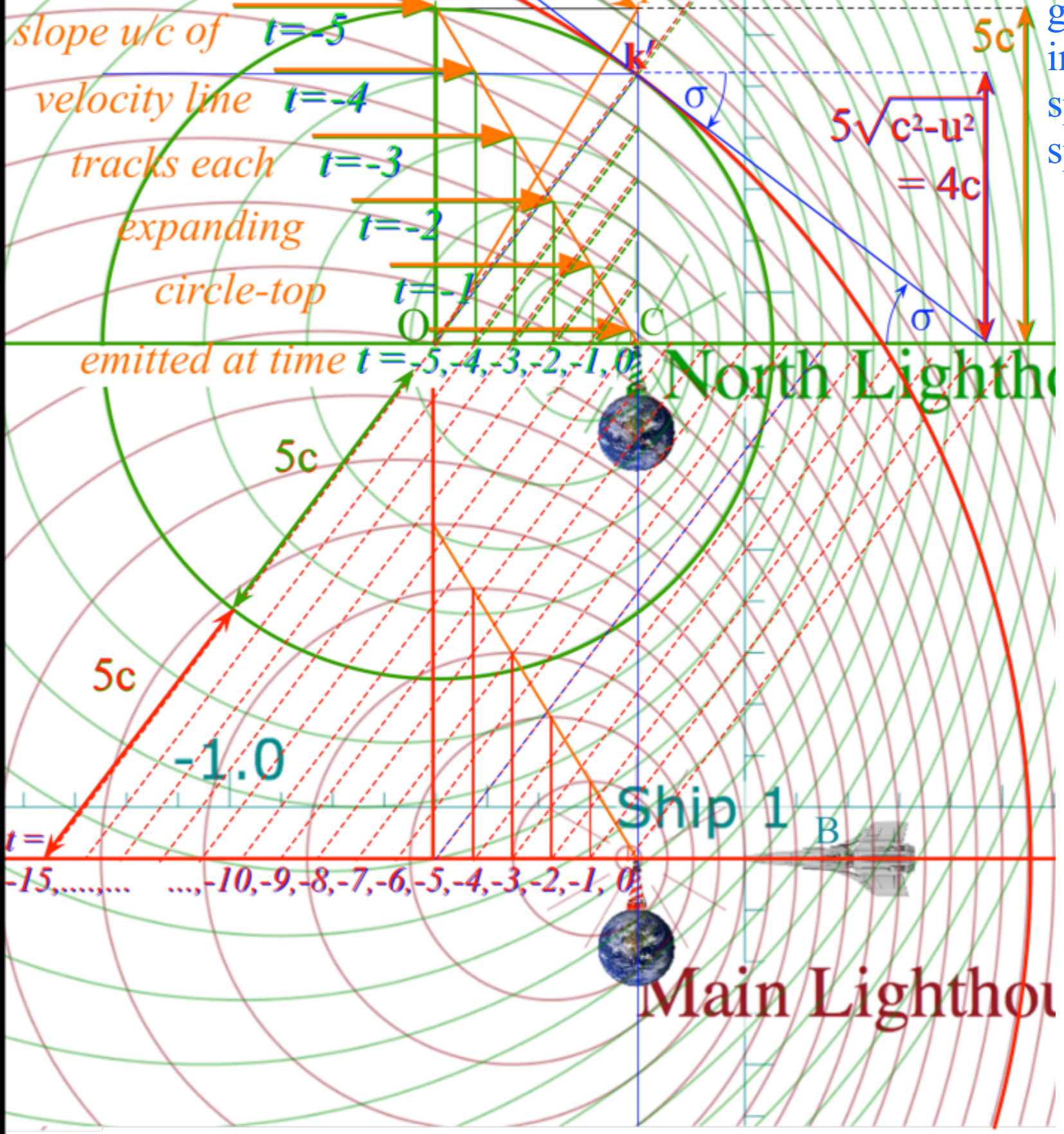
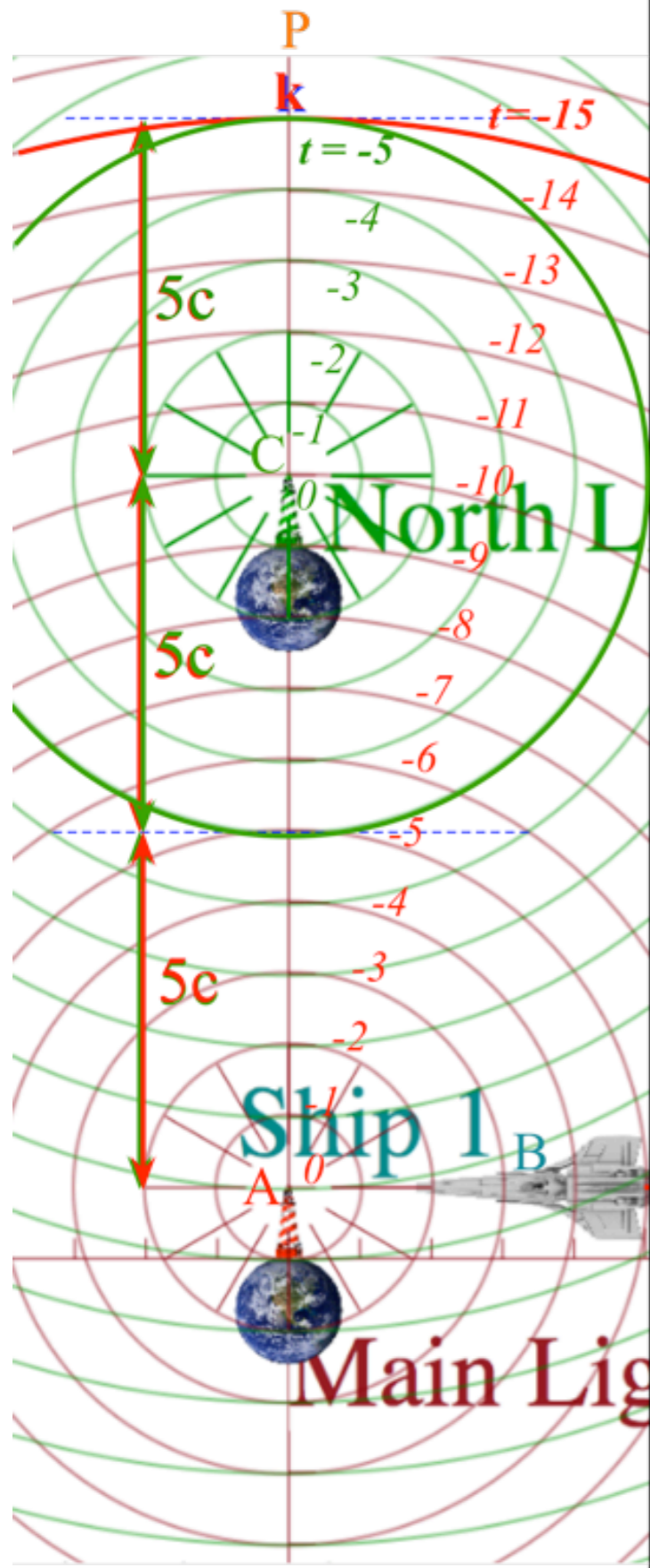


(a) Spherical wave pair
In Alice-Carla frame

$\text{stellar angle } \sigma = \sin^{-1}(u/c)$
 $\text{velocity angle } \nu = \tan^{-1}(u/c)$
 $\text{slope } u/c \text{ of } t = -5$
 $\text{velocity line } t = -4$
 $\text{tracks each expanding circle-top emitted at time } t = -5, -4, -3, -2, -1, 0$

(b) Spherical wave pair
In Bob's frame: $u_x/c = -3/5$

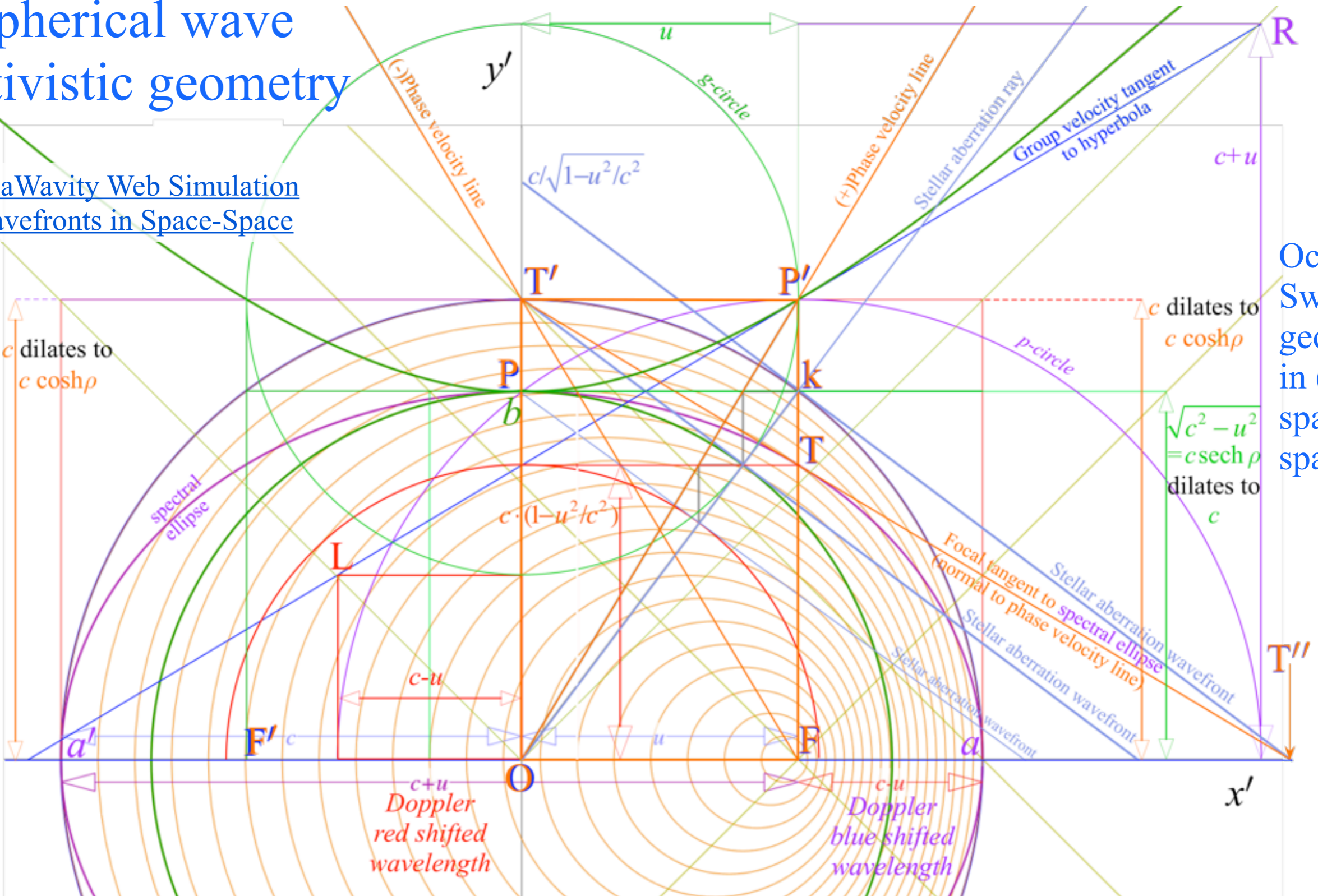
Occam
Sword
geometry
in (x,y)
space-
space



$t = -15, \dots, -10, -9, -8, -7, -6, -5, -4, -3, -2, -1, 0$

Spherical wave relativistic geometry

RelaWavity Web Simulation
Wavefronts in Space-Space

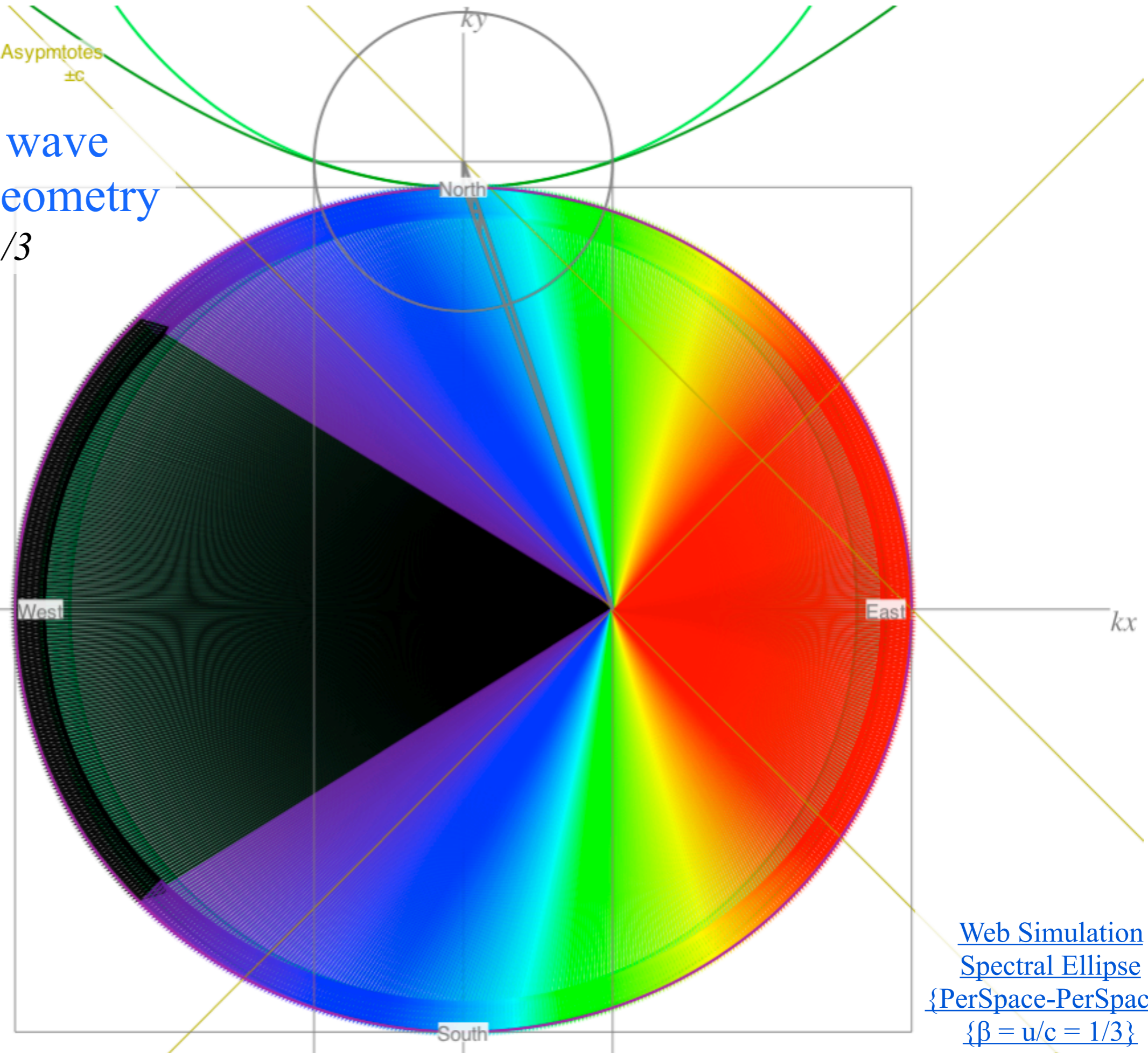


Occam
Sword
geometry
in (x,y)
space-space

<p>Doppler Red $\lambda=c+u$ dilates to: $(c+u) \cosh \rho = c \sqrt{\frac{c+u}{c-u}} = ce^{+\rho}$</p> <p>ellipse major radius $a=OFa=c$ dilates to: $c \cosh \rho = c/\sqrt{1-u^2/c^2}$</p>	<p>Applications of Einstein dilation factor: $\gamma = \cosh \rho = 1/\sqrt{1-u^2/c^2}$</p>	<p>ellipse focal length $FO = u = c \tanh \rho$ dilates to: $u \cosh \rho = c \sinh \rho$</p> <p>ellipse latus radius $FT = c(1-u^2/c^2)$ dilates to: $c(1-u^2/c^2) \cosh \rho = c\sqrt{1-u^2/c^2} = c \operatorname{sech} \rho$</p>	<p>Doppler Blue $\lambda=c-u$ dilates to: $(c-u) \cosh \rho = c \sqrt{\frac{c-u}{c+u}} = ce^{-\rho}$</p> <p>Base height $FTk = \sqrt{c^2 - u^2}$ dilates to: $\sqrt{c^2 - u^2} \cosh \rho = c$ (equal to ellipse minor radius b)</p>
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Spherical wave
relativistic geometry

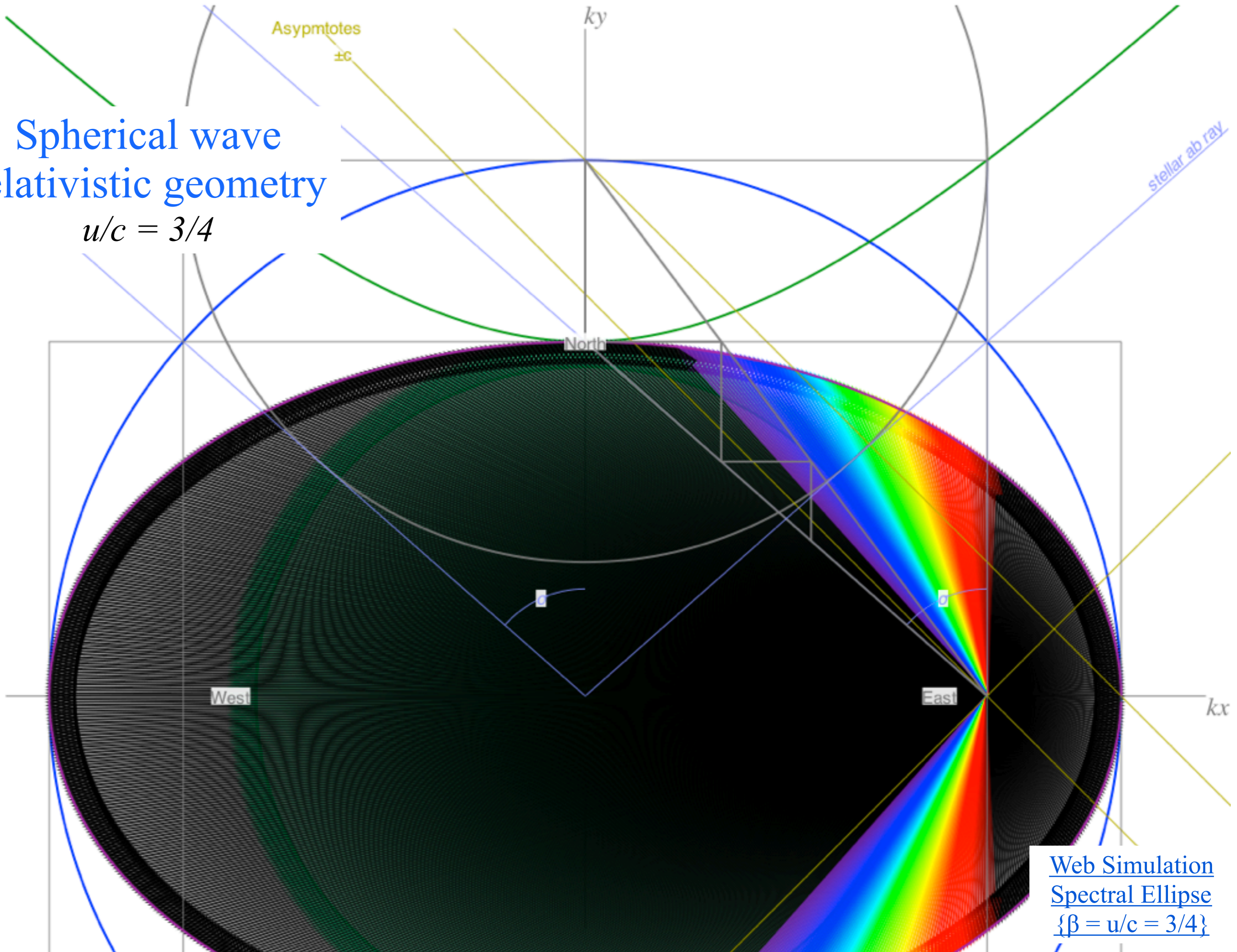
$$u/c = 1/3$$



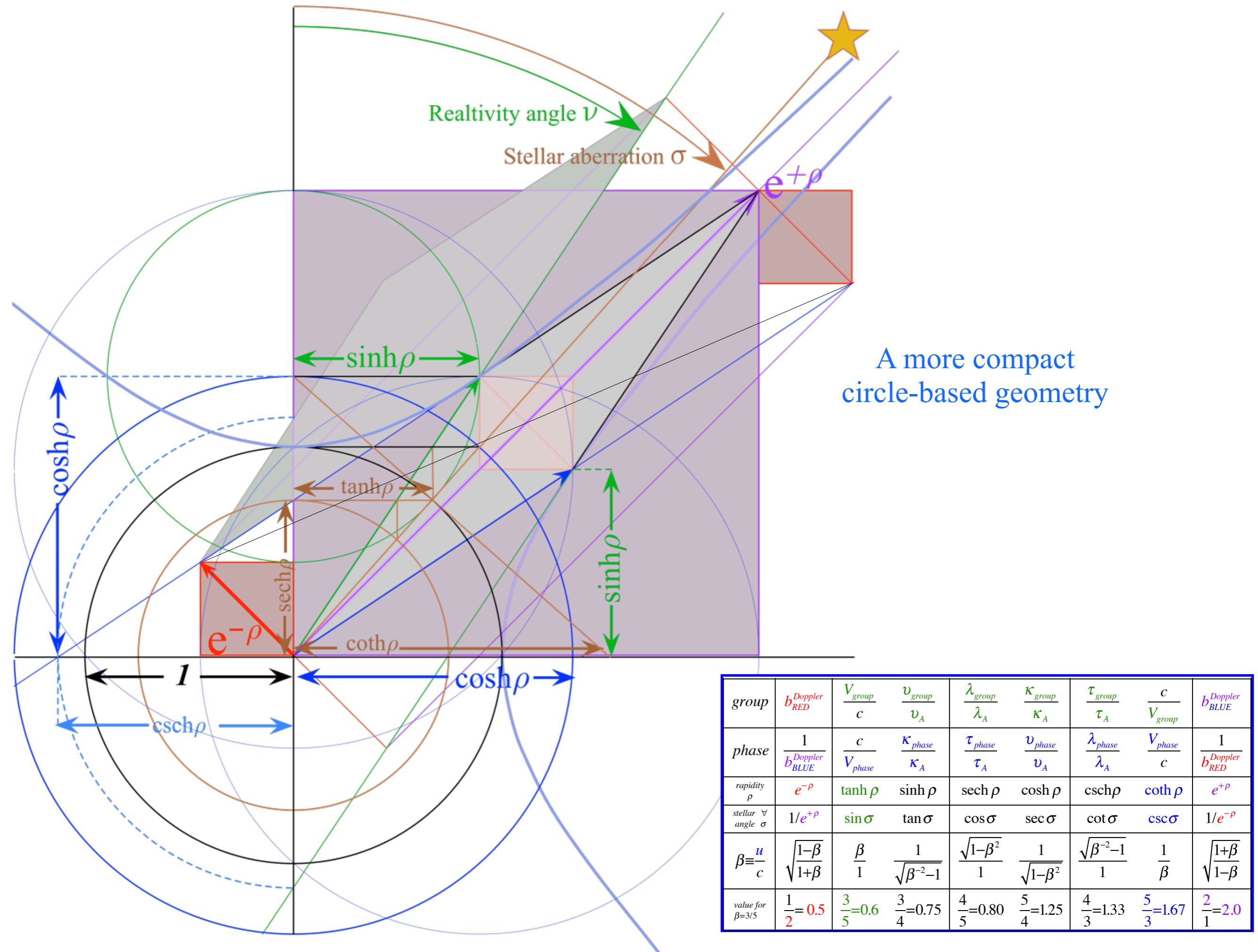
[Web Simulation](#)
[Spectral Ellipse](#)
{PerSpace-PerSpace}
{ $\beta = u/c = 1/3$ }

Spherical wave
relativistic geometry

$$u/c = 3/4$$



Web Simulation
Spectral Ellipse
{ $\beta = u/c = 3/4$ }



group	$b_{\text{RED}}^{\text{Doppler}}$	$\frac{V_{\text{group}}}{c}$	$\frac{v_{\text{group}}}{v_A}$	$\frac{\lambda_{\text{group}}}{\lambda_A}$	$\frac{\kappa_{\text{group}}}{\kappa_A}$	$\frac{\tau_{\text{group}}}{\tau_A}$	$\frac{c}{V_{\text{group}}}$	$b_{\text{BLUE}}^{\text{Doppler}}$
phase	$\frac{1}{b_{\text{BLUE}}^{\text{Doppler}}}$	$\frac{c}{V_{\text{phase}}}$	$\frac{\kappa_{\text{phase}}}{\kappa_A}$	$\frac{\tau_{\text{phase}}}{\tau_A}$	$\frac{v_{\text{phase}}}{v_A}$	$\frac{\lambda_{\text{phase}}}{\lambda_A}$	$\frac{V_{\text{phase}}}{c}$	$\frac{1}{b_{\text{RED}}^{\text{Doppler}}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$