

Lecture 6

Thur. 9.08.2016

Geometry of common power-law potentials

Geometric (Power) series

“Zig-Zag” exponential geometry

Projective or perspective geometry

Parabolic geometry of harmonic oscillator $kr^2/2$ potential and $-kr^1$ force fields

Coulomb geometry of $-1/r$ -potential and $-1/r^2$ -force fields

Compare mks units of Coulomb Electrostatic vs. Gravity

Geometry of idealized “Sophomore-physics Earth”

Coulomb field outside

Isotropic Harmonic Oscillator (IHO) field inside

Contact-geometry of potential curve(s)

“Crushed-Earth” models: 3 key energy “steps” and 4 key energy “levels”

Earth matter vs nuclear matter:

*Introducing the “neutron starlet” and “**Black-Hole-Earth**”*

Introducing 2D IHO orbits and phasor geometry

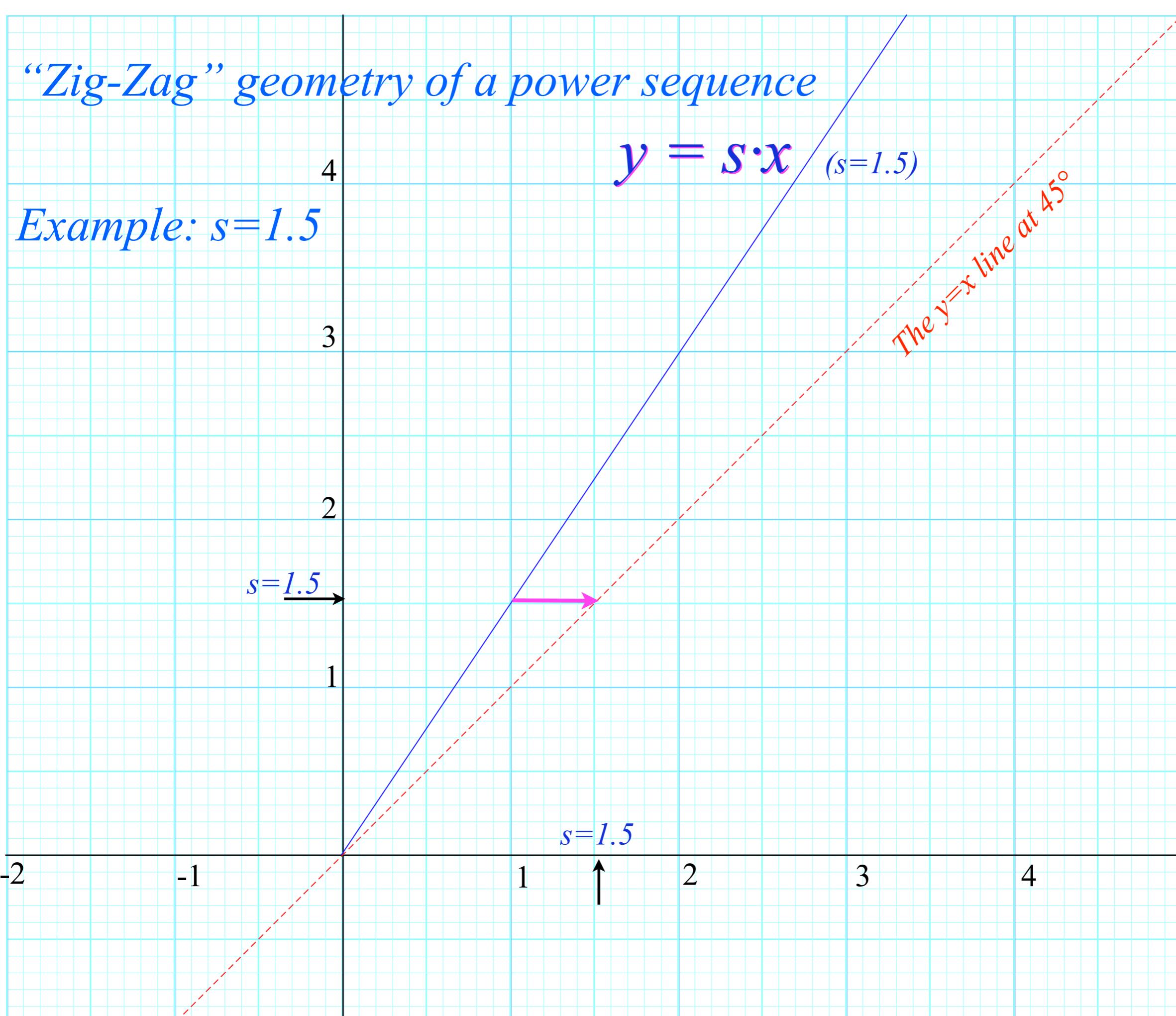
Phasor “clock” geometry

Geometry of common power-law potentials

- *Geometric (Power) series*
- “Zig-Zag” exponential geometry
- Projective or perspective geometry
- Parabolic geometry of harmonic oscillator $kr^2/2$ potential and $-kr^l$ force fields
- Coulomb geometry of $-1/r$ -potential and $-1/r^2$ -force fields
- Compare mks units of Coulomb Electrostatic vs. Gravity

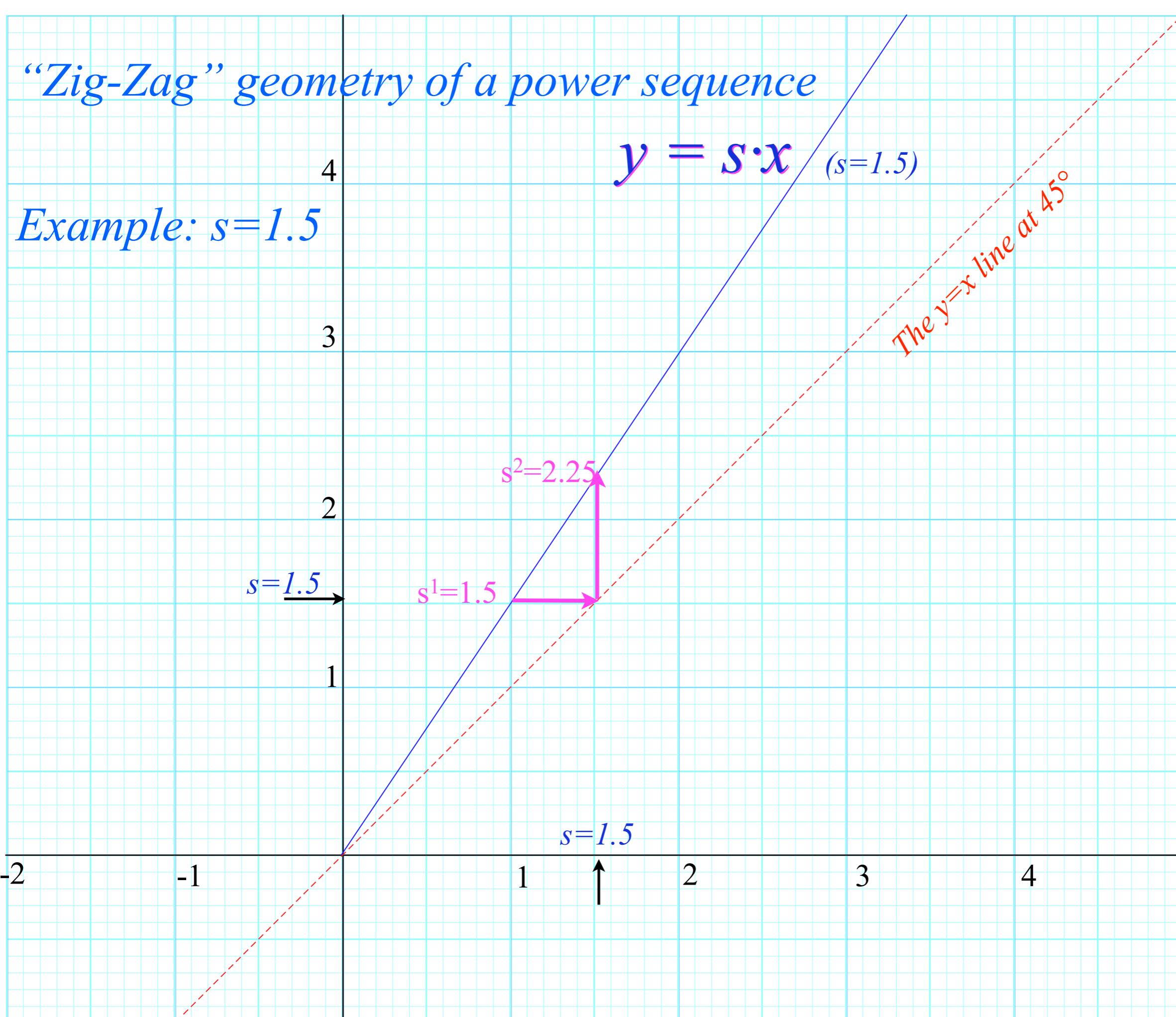
“Zig-Zag” geometry of a power sequence

Example: $s=1.5$



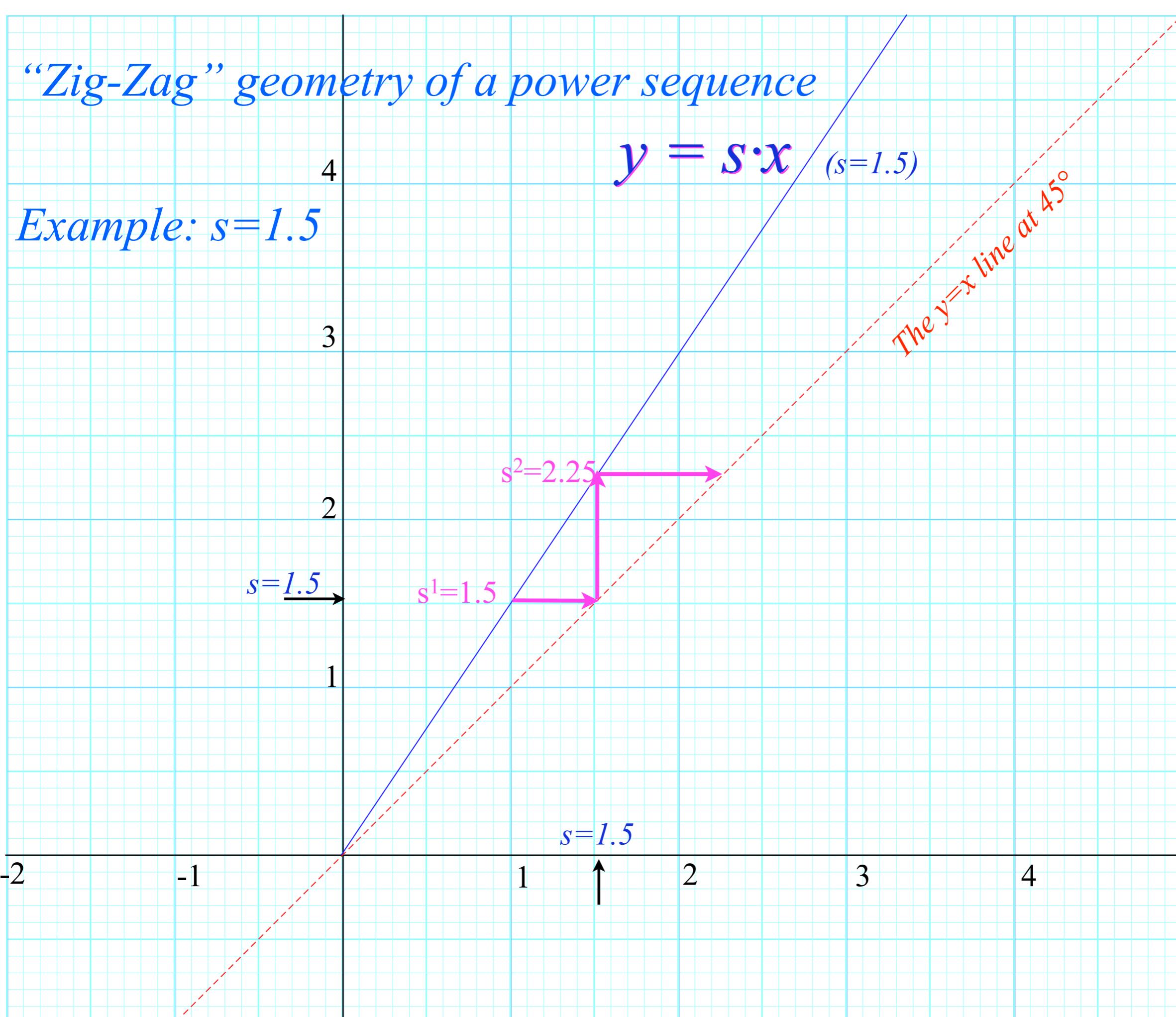
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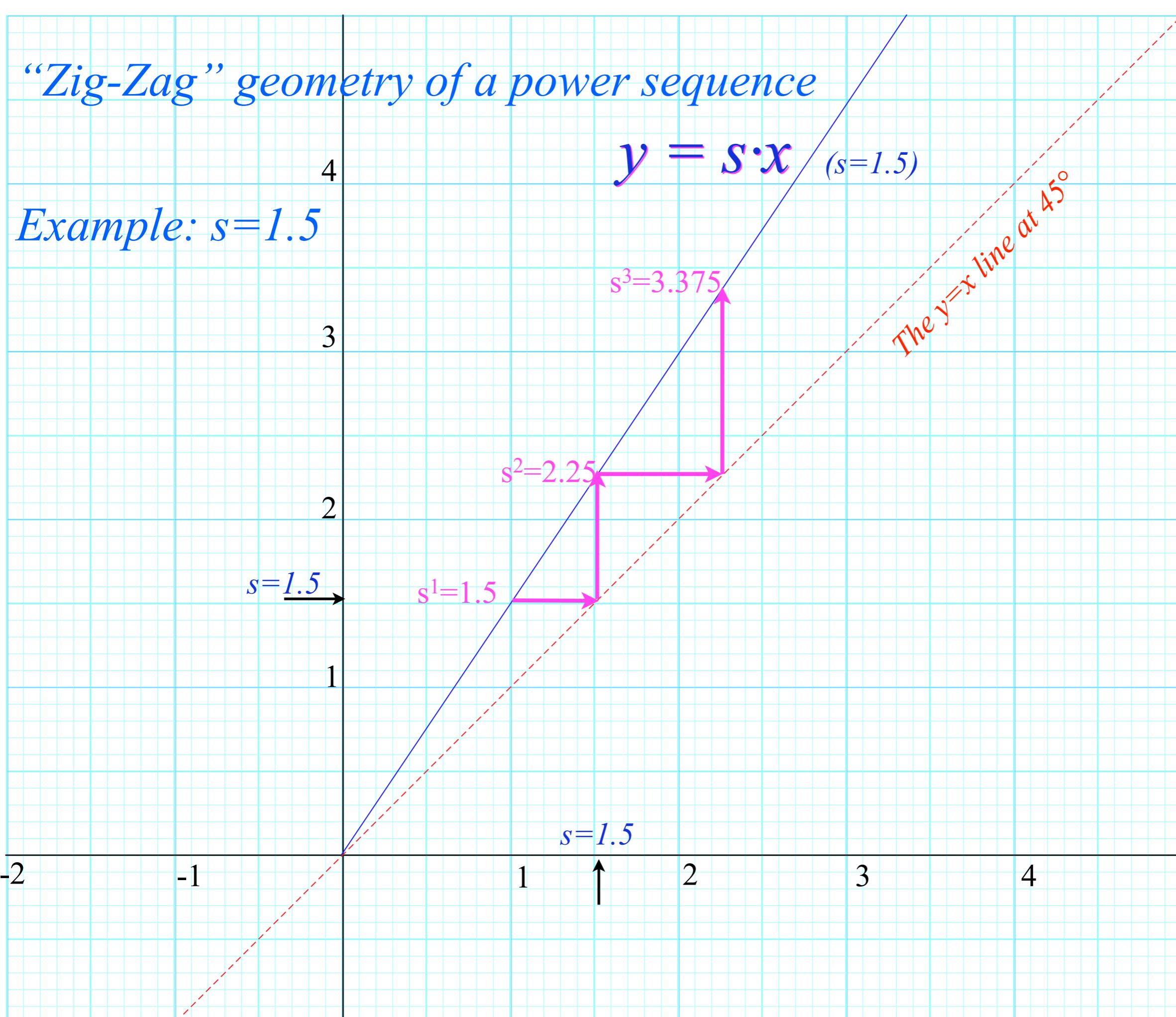
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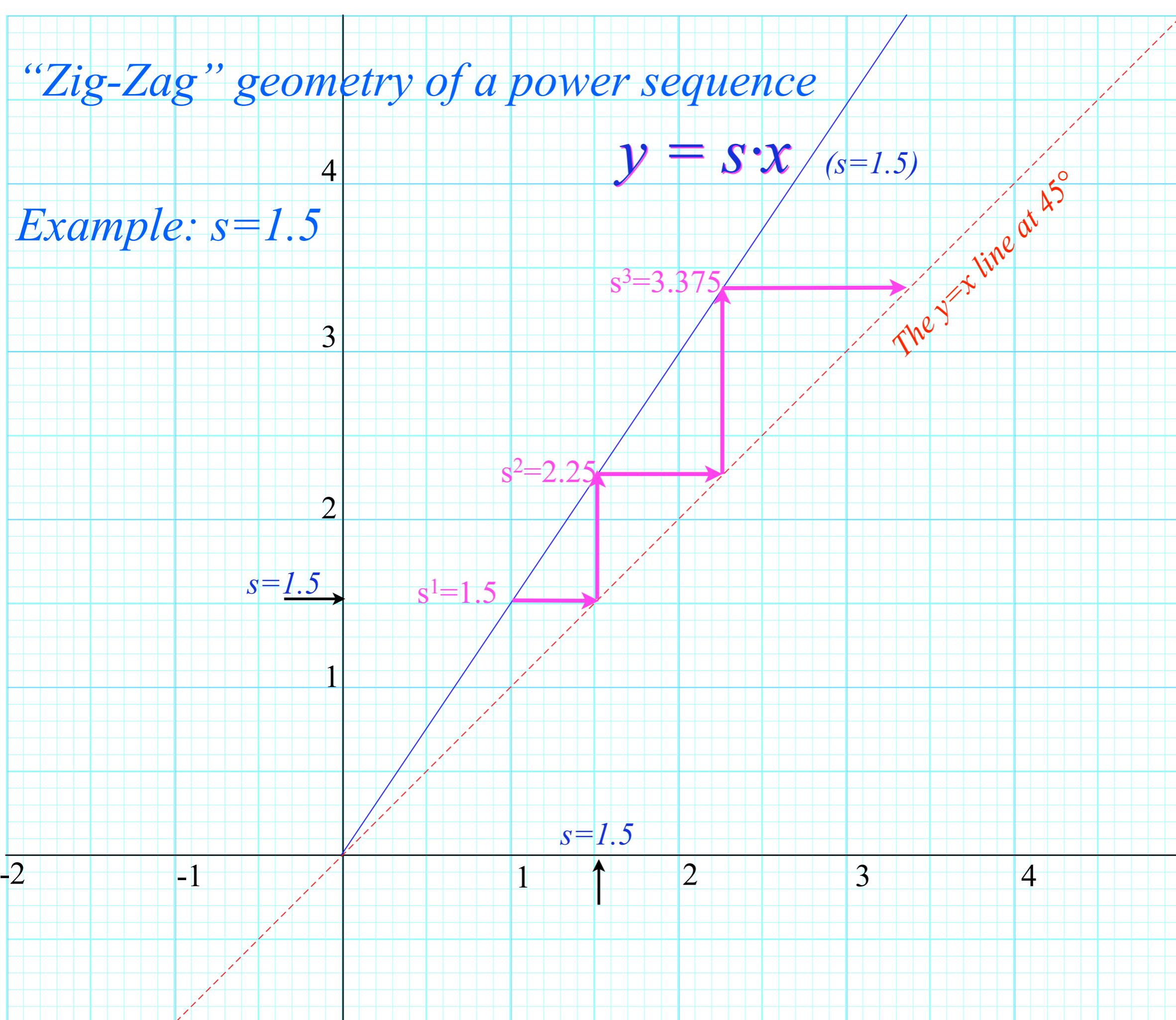
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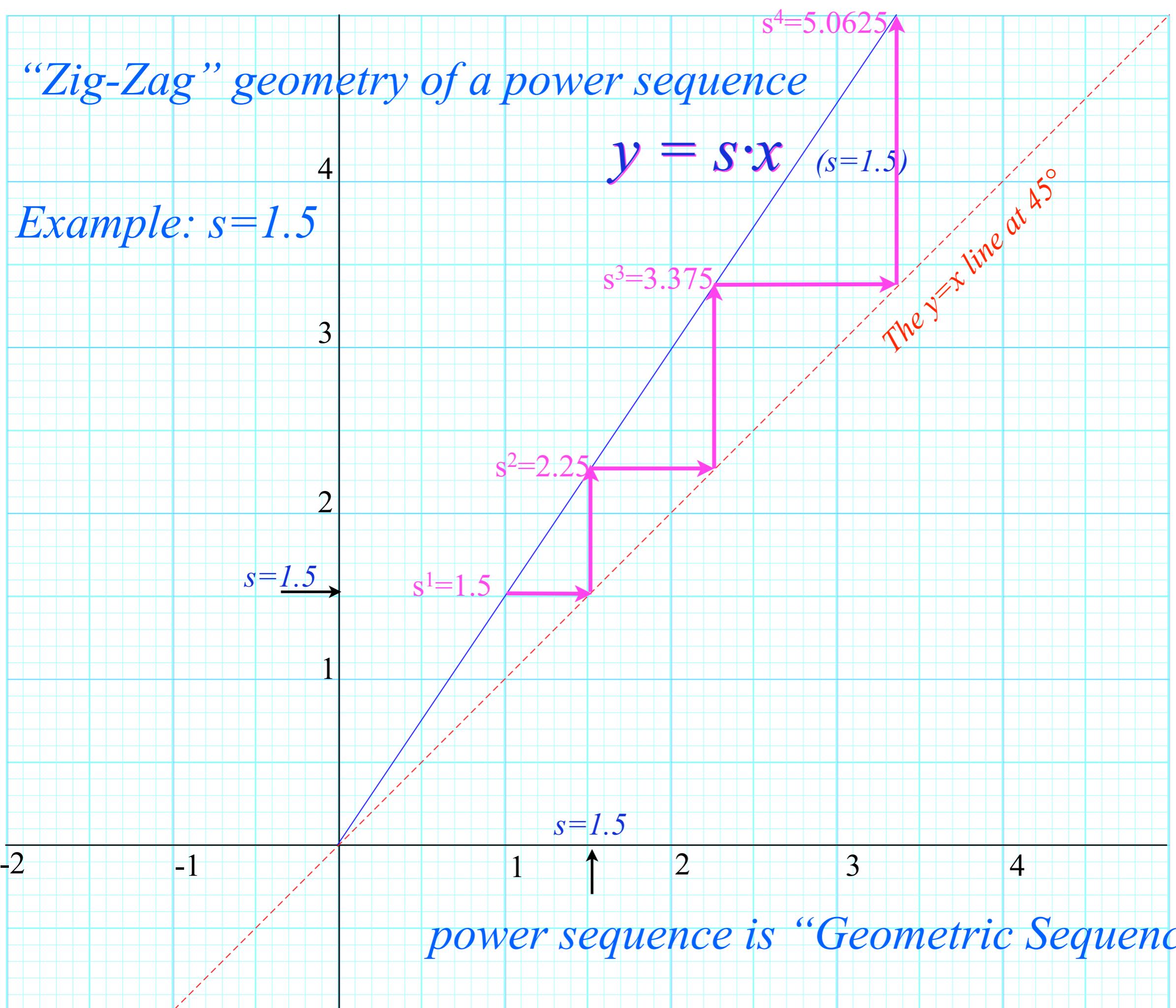
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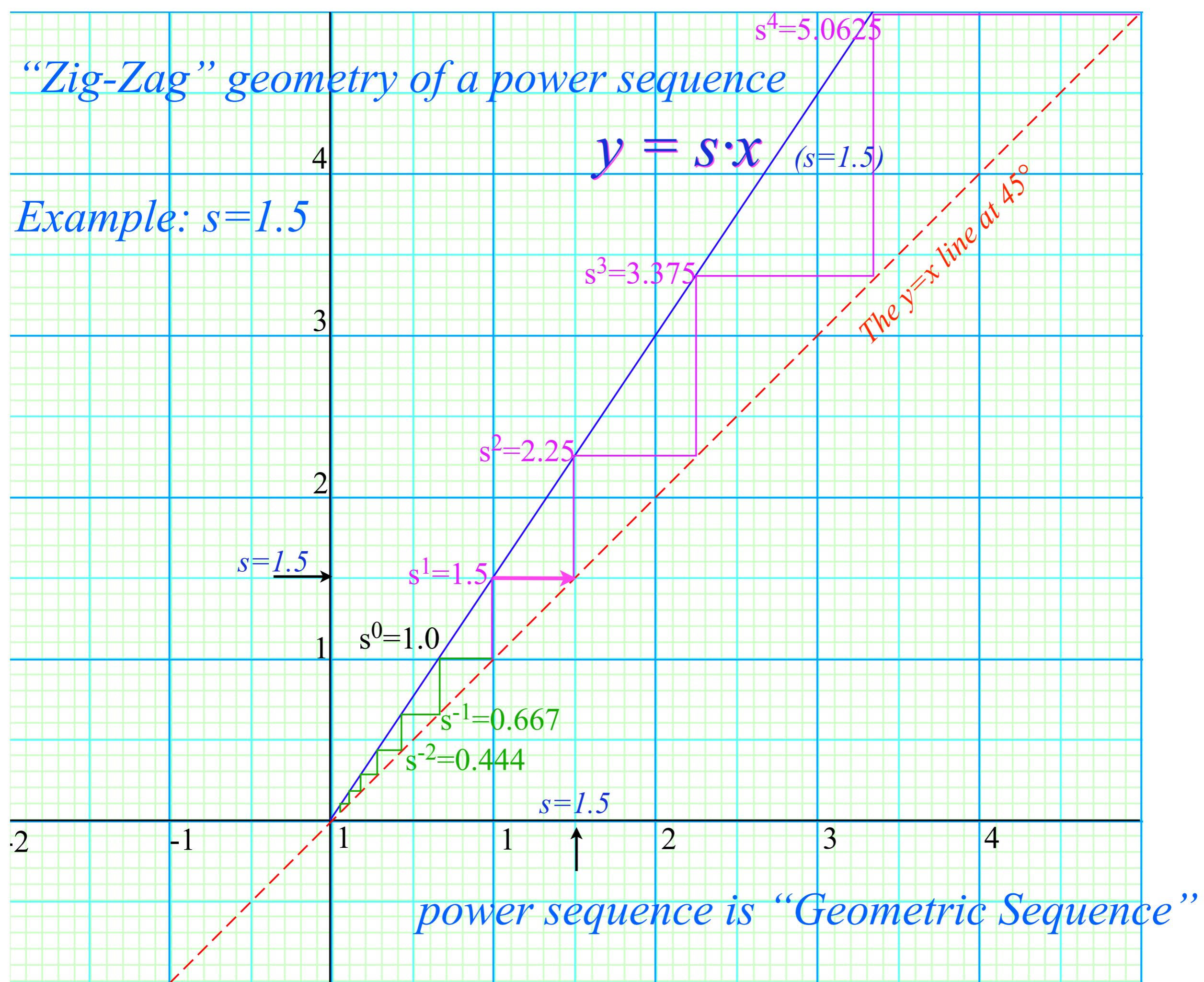
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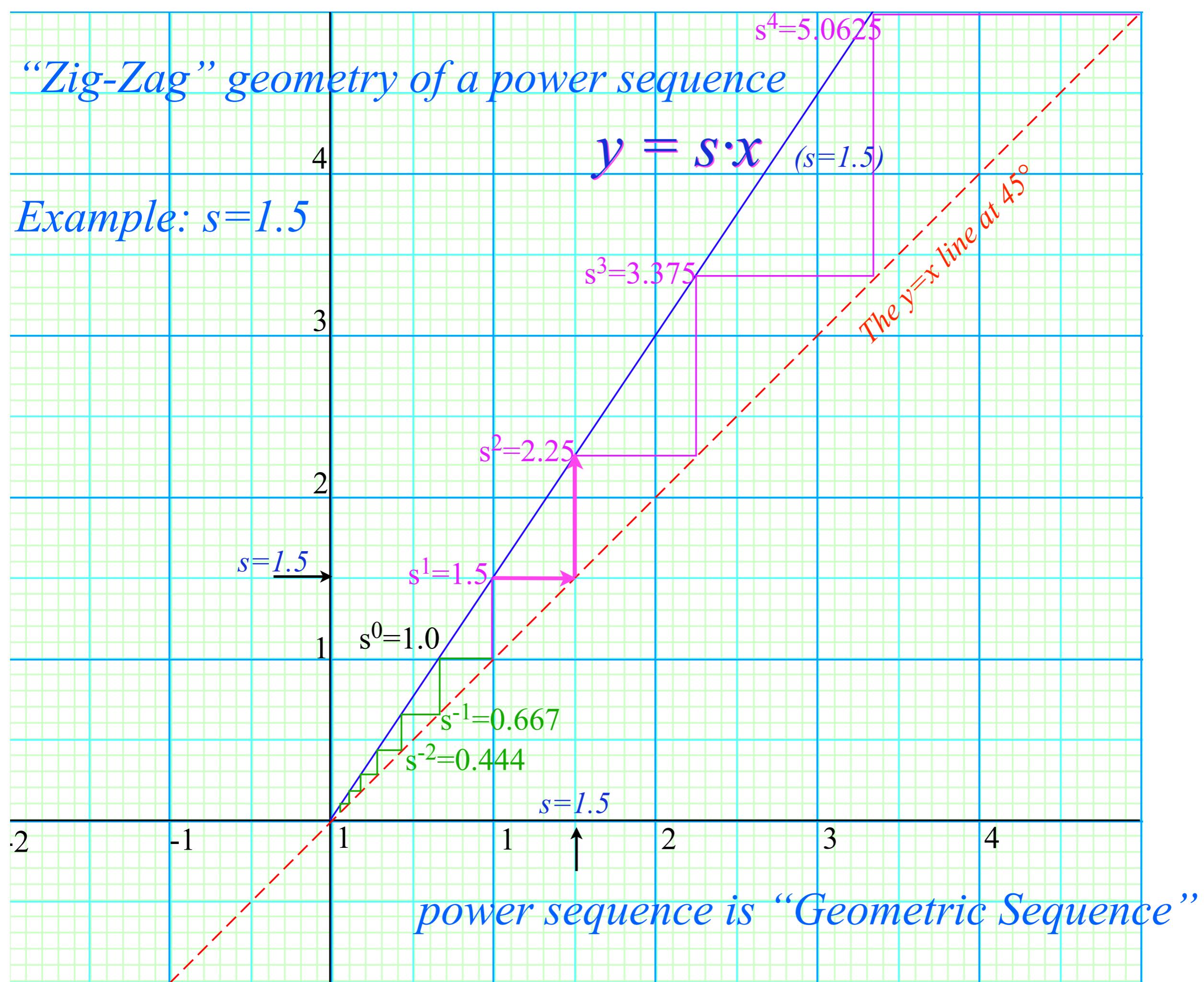
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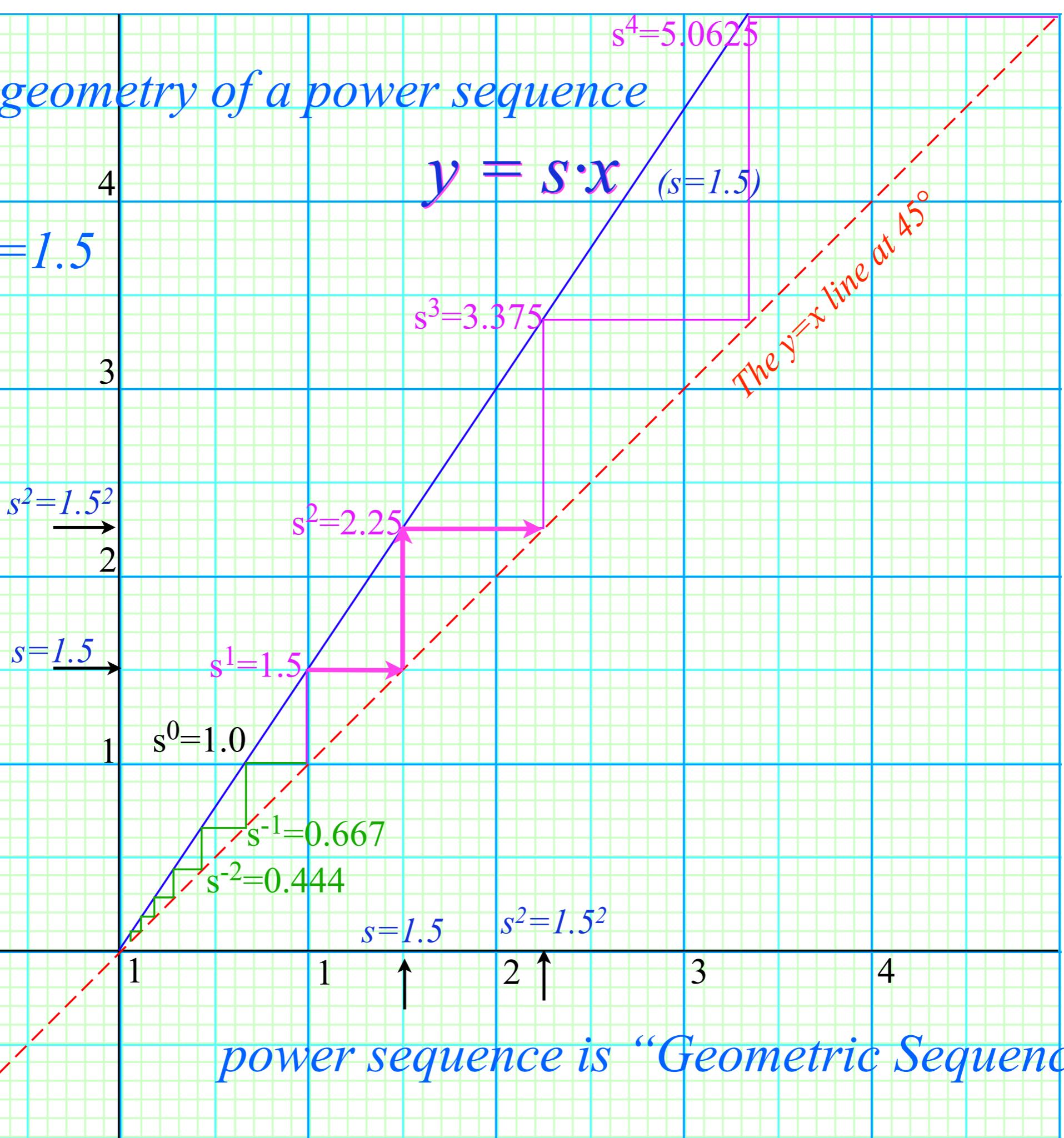
“Zig-Zag” geometry of a power sequence

Example: $s=1.5$



“Zig-Zag” geometry of a power sequence

Example: $s=1.5$



“Zig-Zag” geometry of a power sequence

Example: $s=1.5$

$$s^3=1.5^3$$

$$s^2=1.5^2$$

$$s=1.5$$

$$s^0=1.0$$

$$\begin{aligned}s^{-1} &= 0.667 \\ s^{-2} &= 0.444\end{aligned}$$

$$y = s \cdot x \quad (s=1.5)$$

$$s^4=5.0625$$

The $y=x$ line at 45°

$$s^3=3.375$$

$$s^2=2.25$$

$$s^1=1.5$$

$$s=1.5$$

$$s^2=1.5^2$$

power sequence is “Geometric Sequence”

“Zig-Zag” geometry of a power sequence

Example: $s=1.5$

$$s^3 = 1.5^3$$

$$s^2 = 1.5^2$$

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$$s^2 = 1.5^3$$

power sequence is “Geometric Sequence”

“Zig-Zag” geometry of a power sequence

$$y = s \cdot x \quad (s=1.5)$$

Example: $s=1.5$

...and
exponential
function...

$$s=1.5$$

$$s^4=5.0625$$

$$s^3=3.375$$

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$$s^0=1.0$$

$$s^{-1}=0.667$$

$$s^{-2}=0.444$$

$$s=1.5$$

The $y=x$ line at 45°

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The $y=x$ line at 45°

Example: $s=1.5$

...and
exponential
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Approximating

$$y = s^x$$

$$s=1.5$$

power sequence is “Geometric Sequence”

“Zig-Zag” geometry of a power sequence

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The $y=x$ line at 45°

Example: $s=1.5$

...and
exponential
function...

Approximating

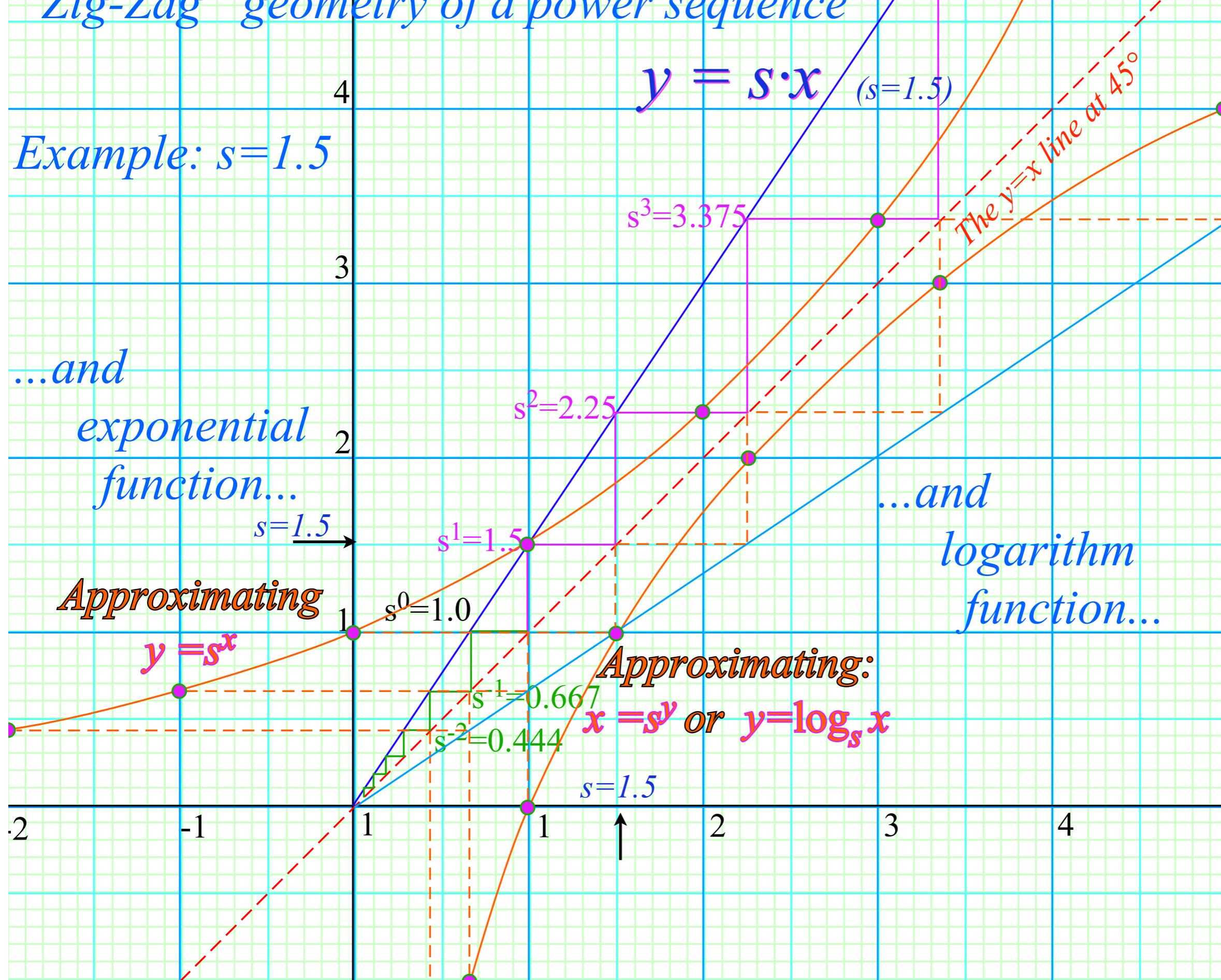
$$y = s^x$$

$$s=1.5$$

...and
logarithm
function...

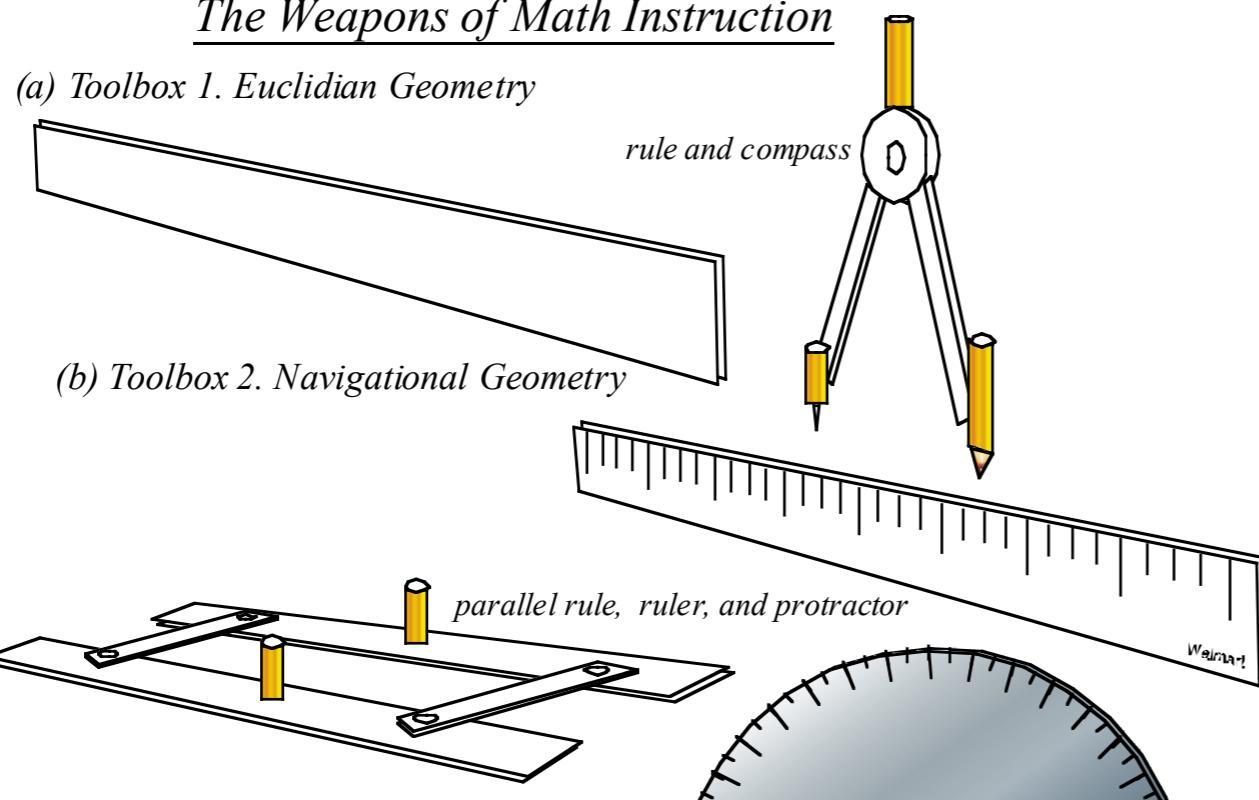
Approximating:

$$x = s^y \text{ or } y = \log_s x$$

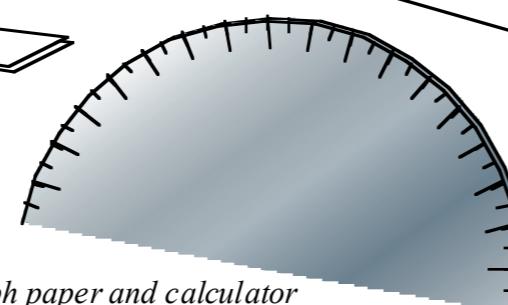


The Weapons of Math Instruction

(a) Toolbox 1. Euclidian Geometry

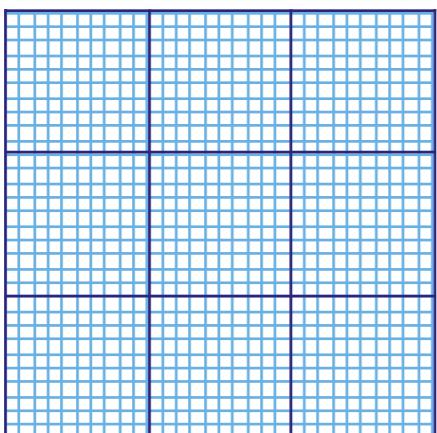


(b) Toolbox 2. Navigational Geometry



So far we mostly use
Toolbox (a-b)

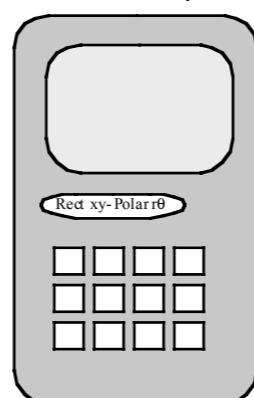
(c) Toolbox 3. Analytical geometry



Graph paper and calculator

Complex algebra and calculus
 $1/z = r^{-1} e^{-i\theta}$

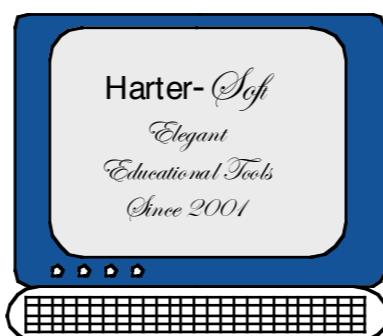
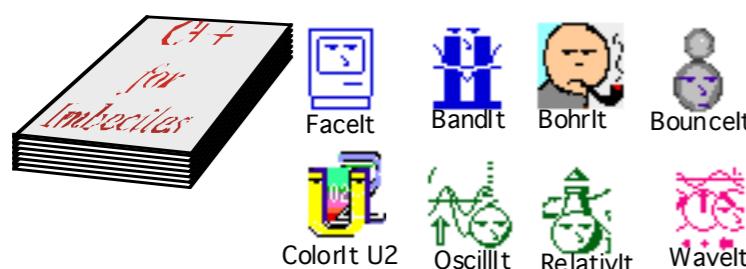
$$\int 1/z \, dz = \ln z$$

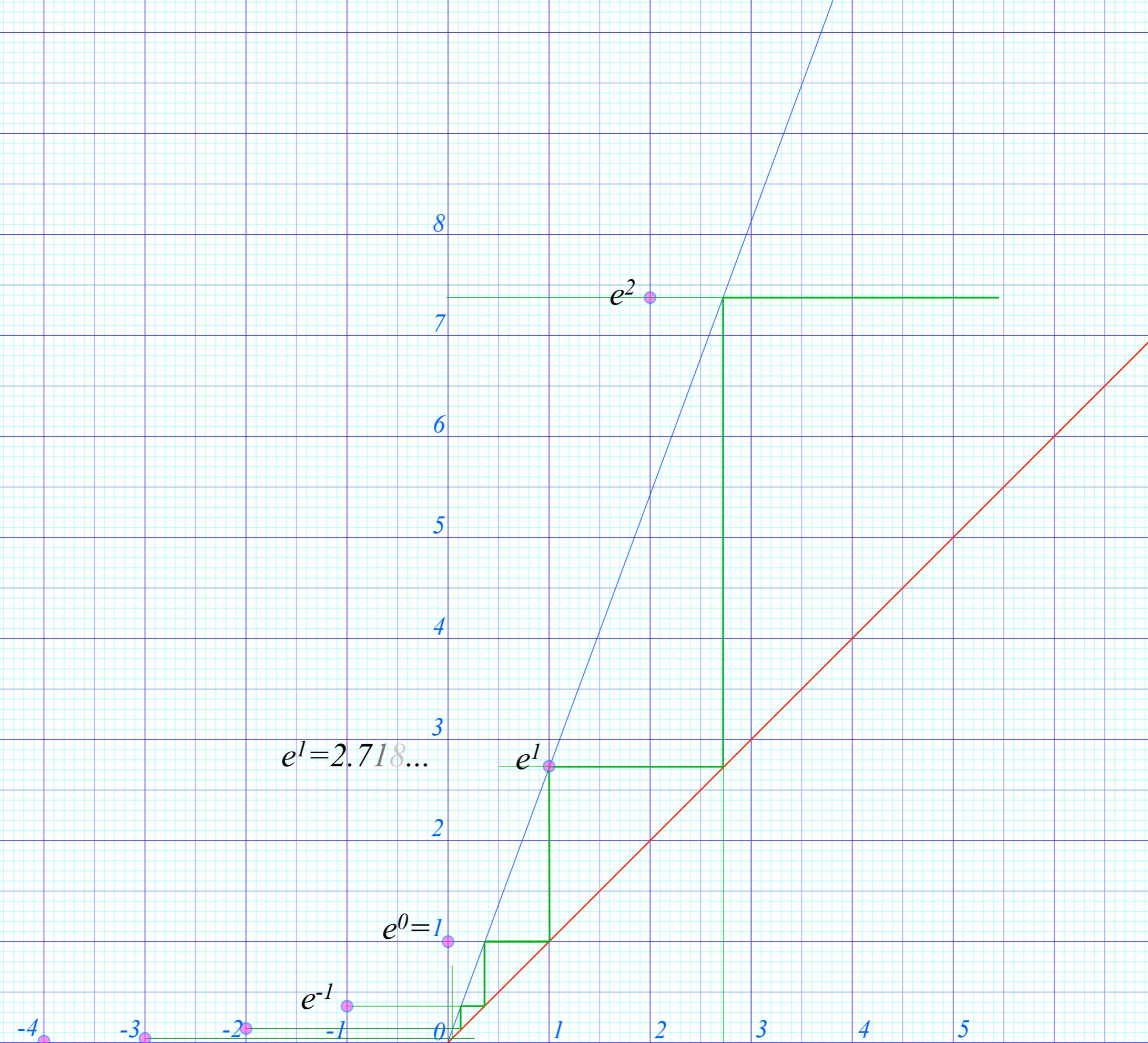


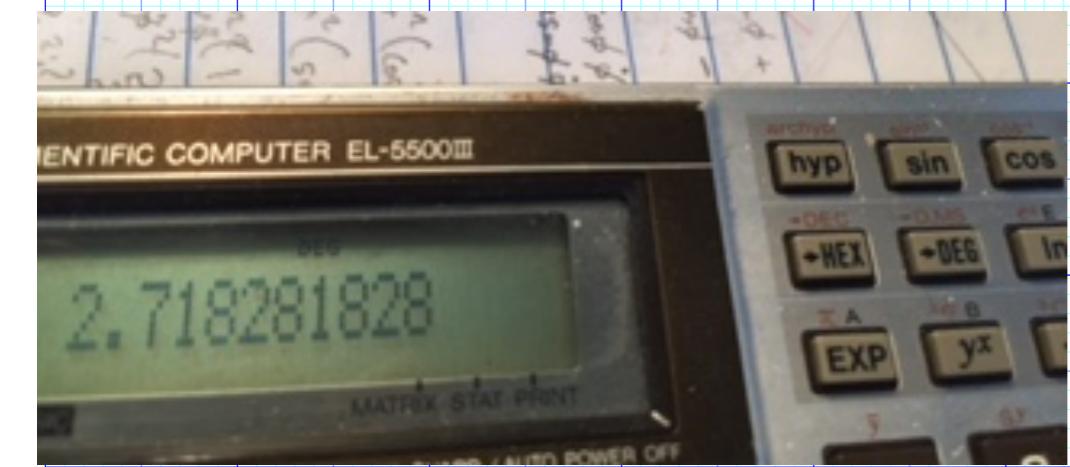
What follows uses
Toolbox (c) ...

...and Toolbox (d)

(d) Toolbox 4. Computer geometry..Anything goes!



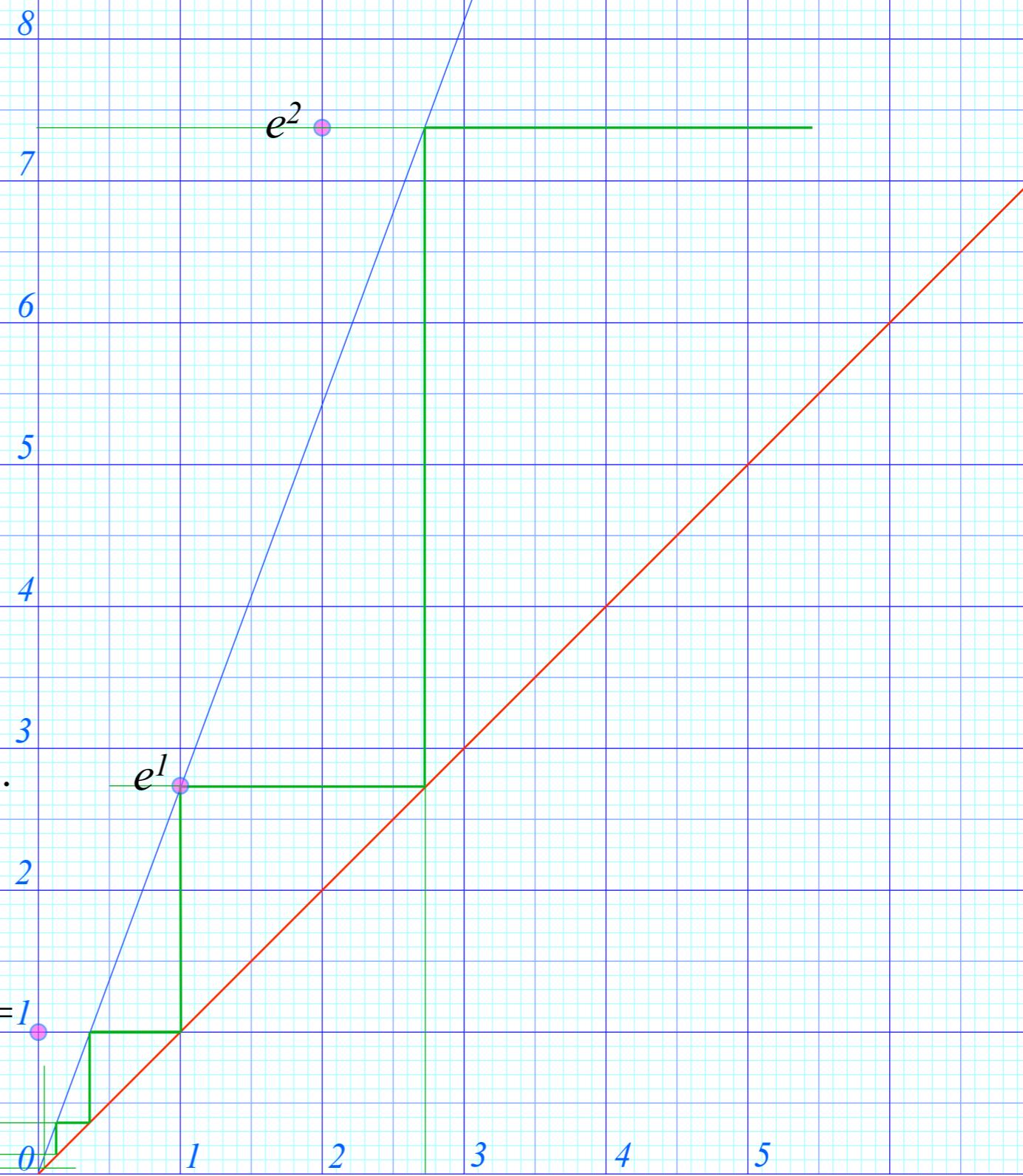




$$e^I = 2.718\dots$$

$$e^{-I}$$

$$e^0 = 1$$



Geometry of common power-law potentials

Geometric (Power) series

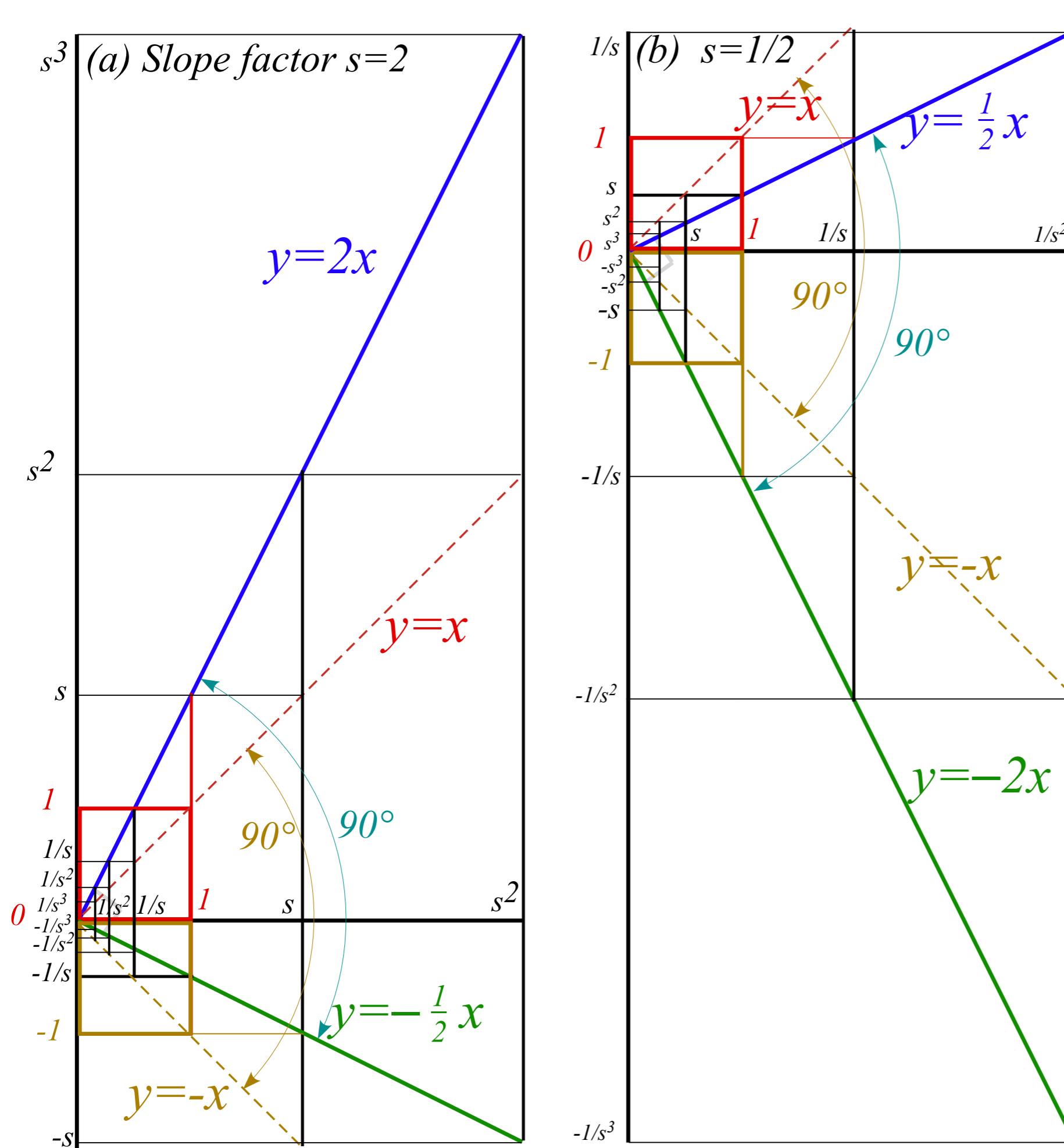
“Zig-Zag” exponential geometry

 *Projective or perspective geometry*

Parabolic geometry of harmonic oscillator $kr^2/2$ potential and $-kr^l$ force fields

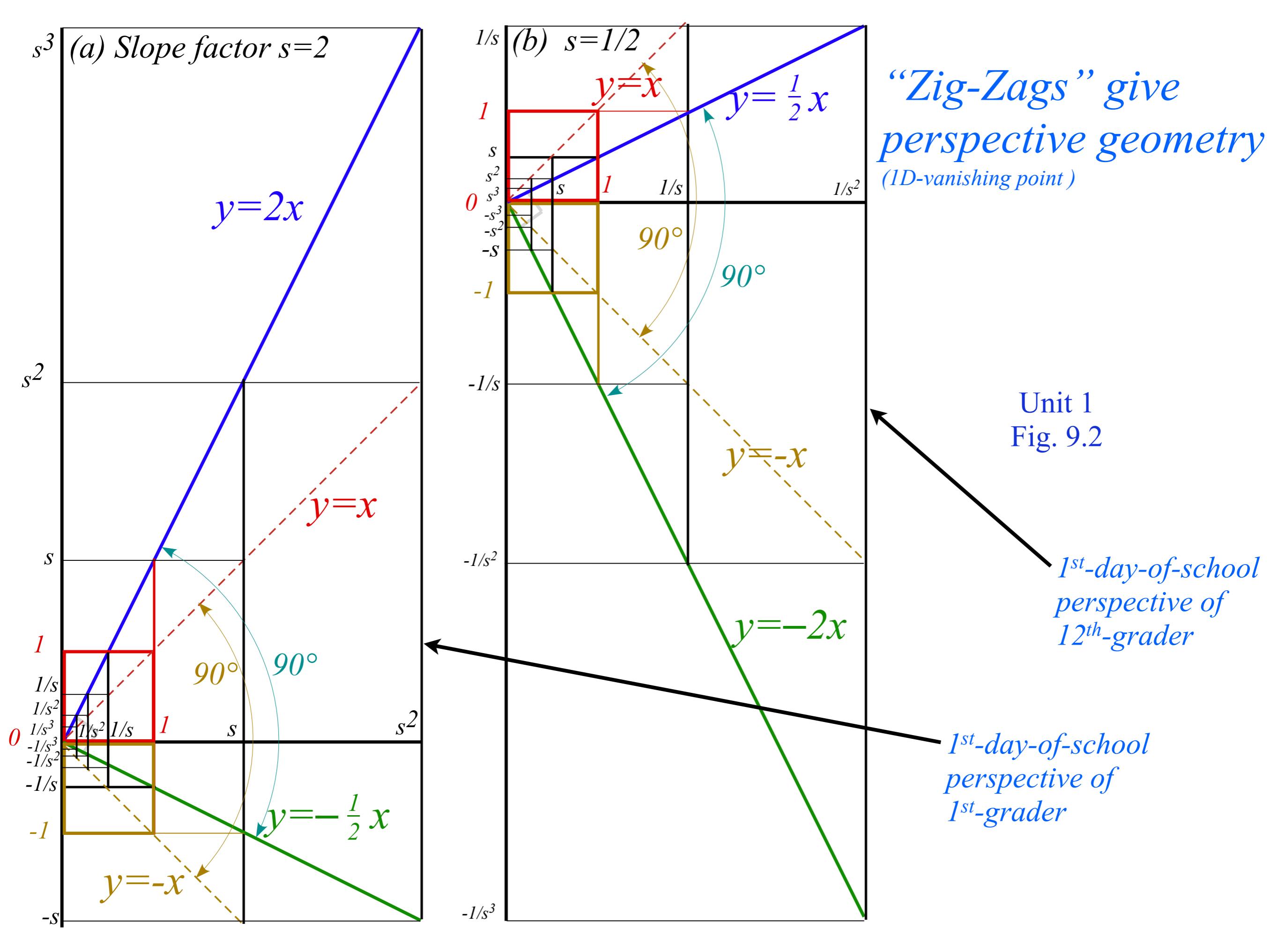
Coulomb geometry of $-1/r$ -potential and $-1/r^2$ -force fields

Compare mks units of Coulomb Electrostatic vs. Gravity



“Zig-Zags” give
perspective geometry
(1D-vanishing point)

Unit 1
Fig. 9.2



Geometry of common power-law potentials

Geometric (Power) series

“Zig-Zag” exponential geometry

Projective or perspective geometry

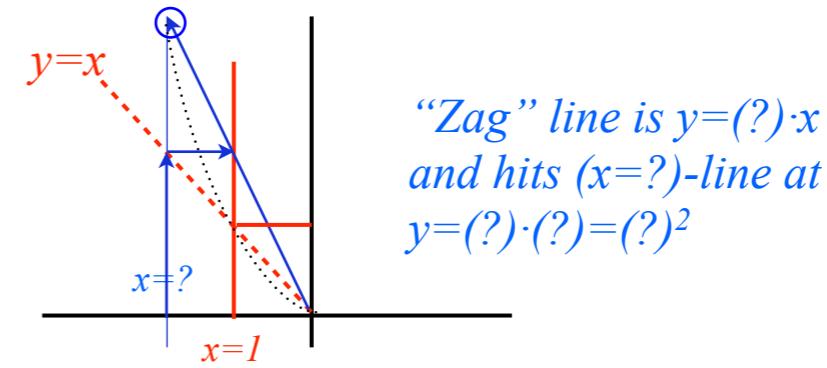
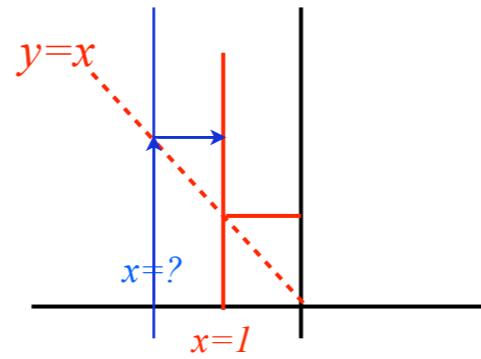
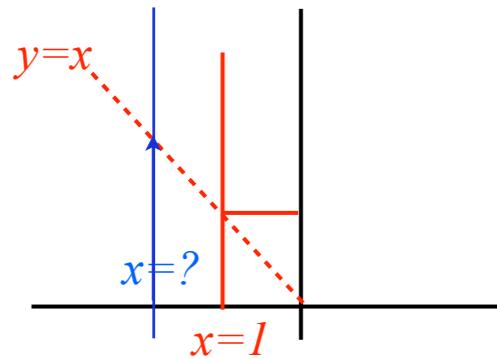
→ *Parabolic geometry of harmonic oscillator $kr^2/2$ potential and $-kr^l$ force fields*

Coulomb geometry of $-1/r$ -potential and $-1/r^2$ -force fields

Compare mks units of Coulomb Electrostatic vs. Gravity

Each $y=x^2$ parabola point found by just one “Zig-Zag”

1. Pick an $(x=?)$ -line
2. “Zig” from its $y=x$ intersection to $x=1$ line
3. “Zag” from origin back to $(x=?)$ -line

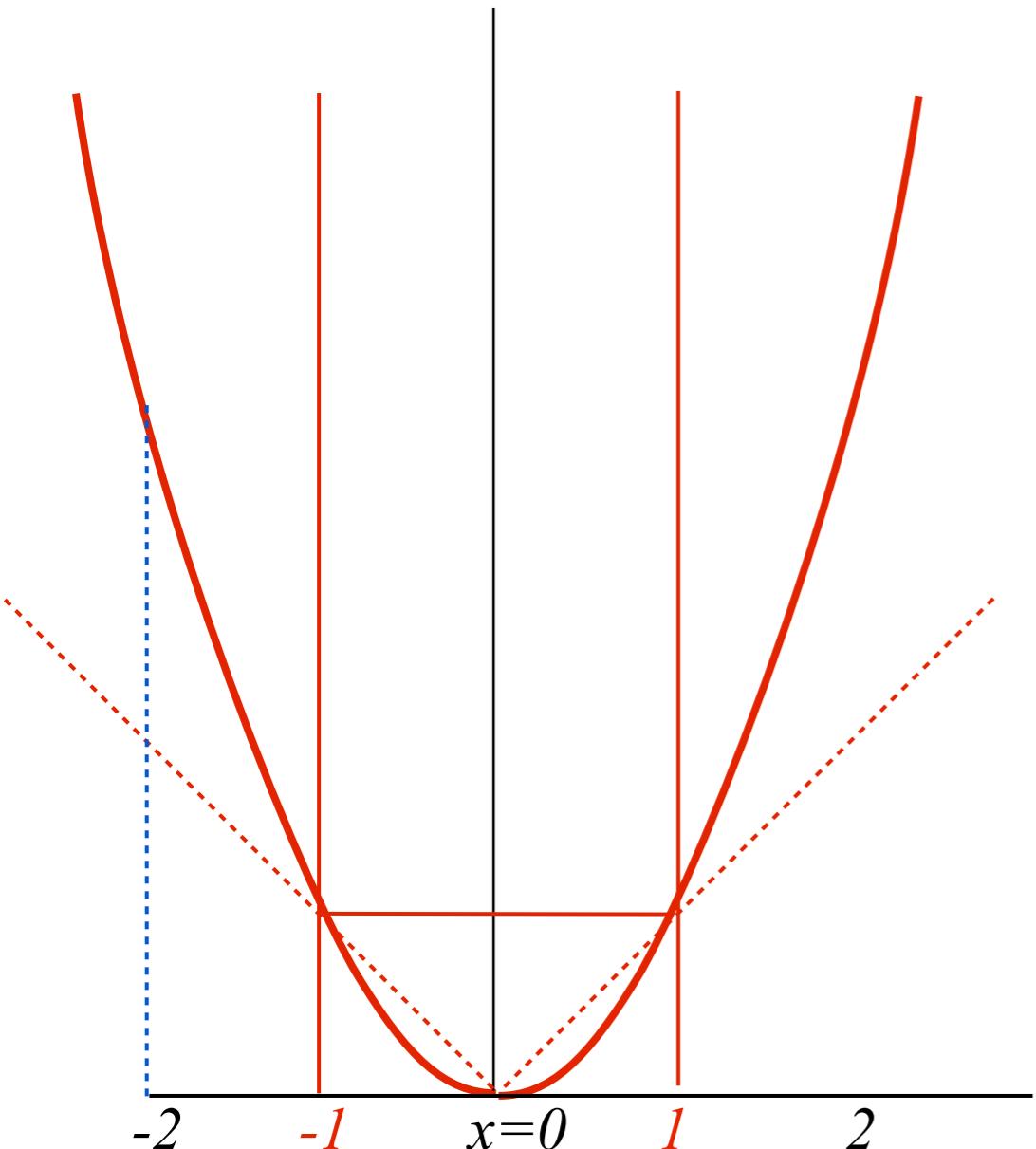
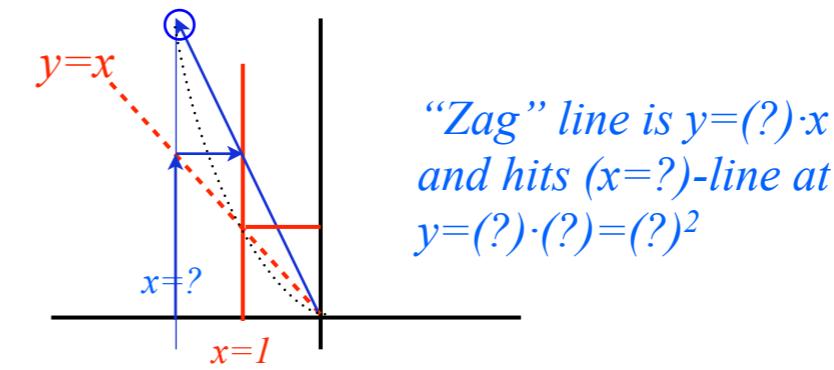
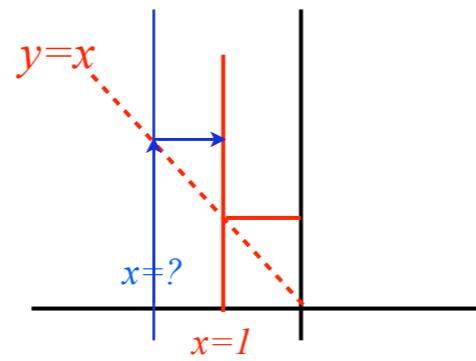
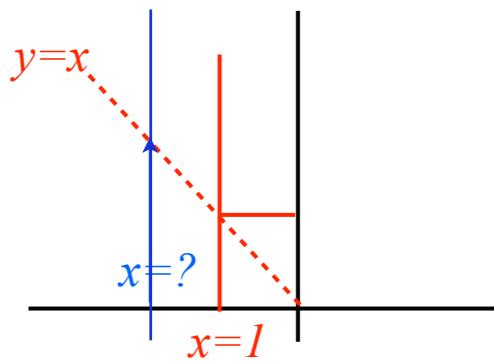


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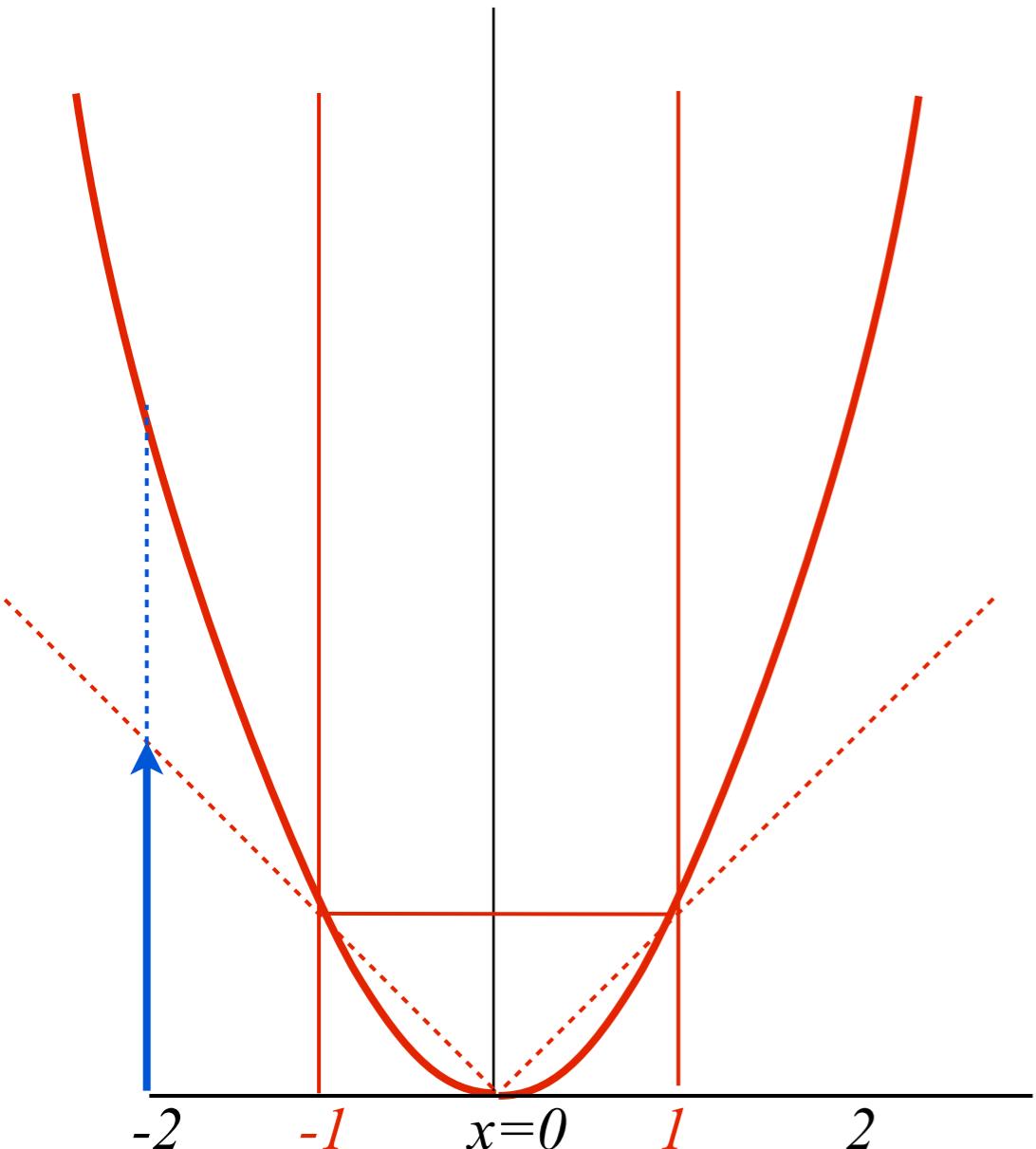
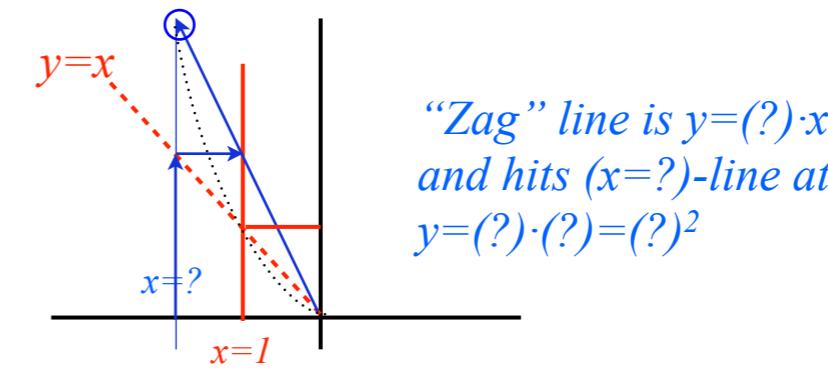
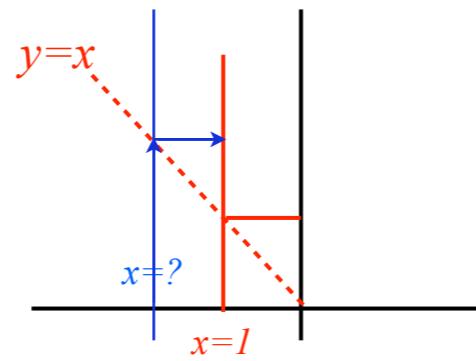
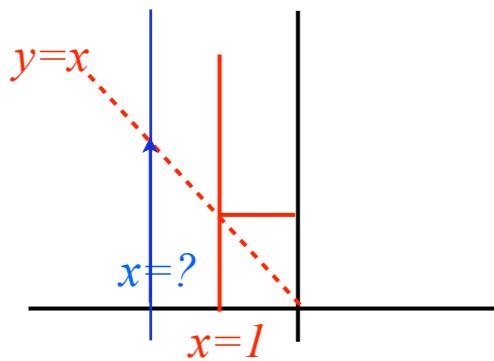
3. “Zag” from origin back to $(x=?)$ -line



Unit 1
Fig. 9.1

Each $y=x^2$ parabola point found by just one “Zig-Zag”

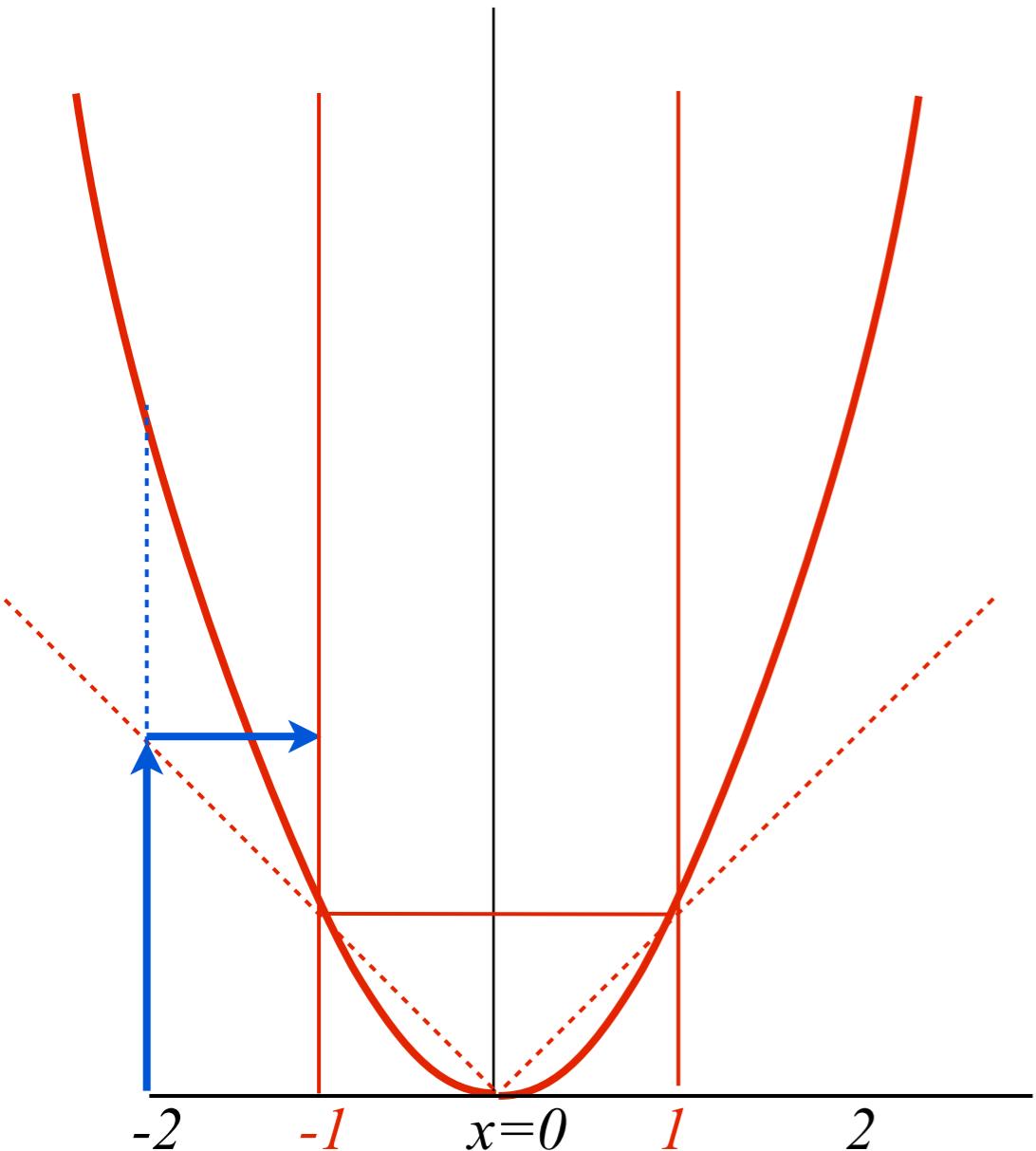
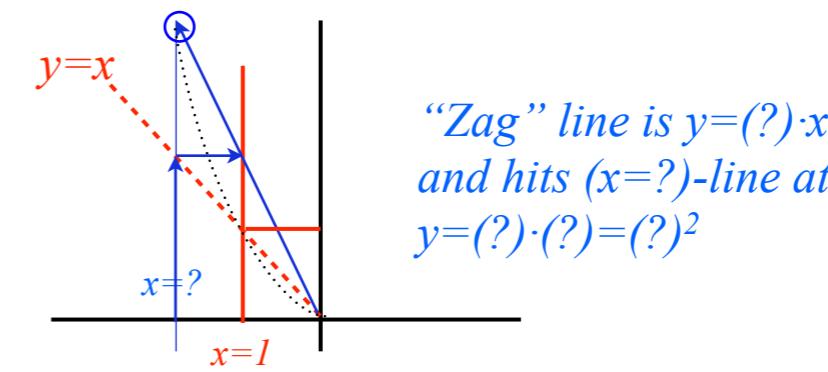
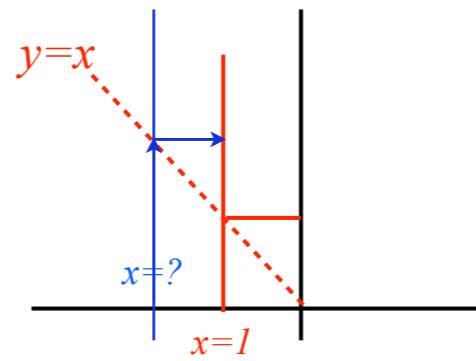
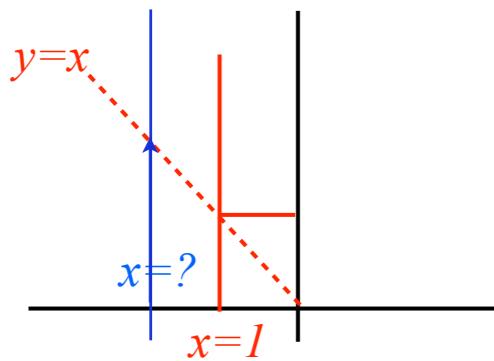
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Unit 1
Fig. 9.1

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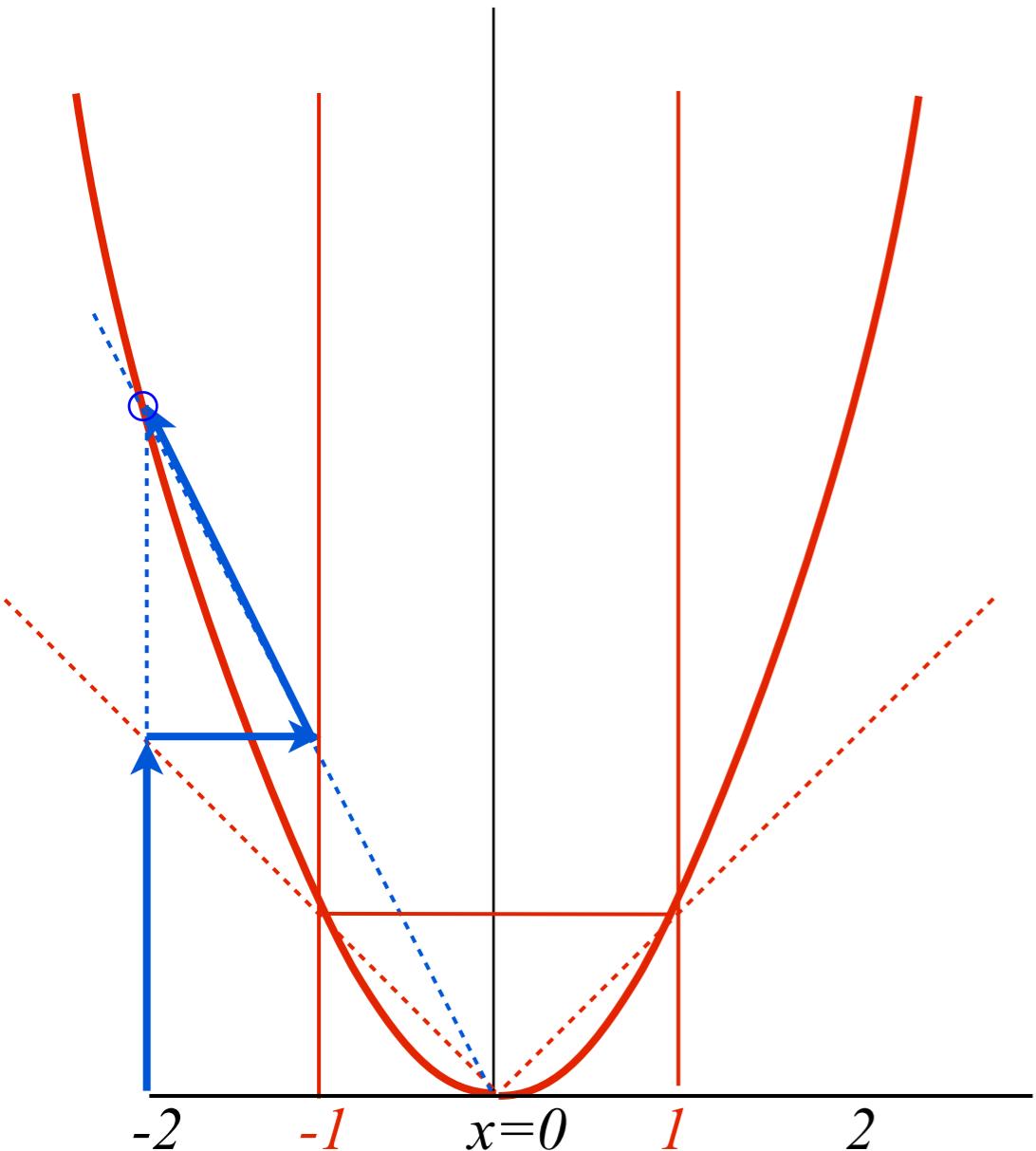
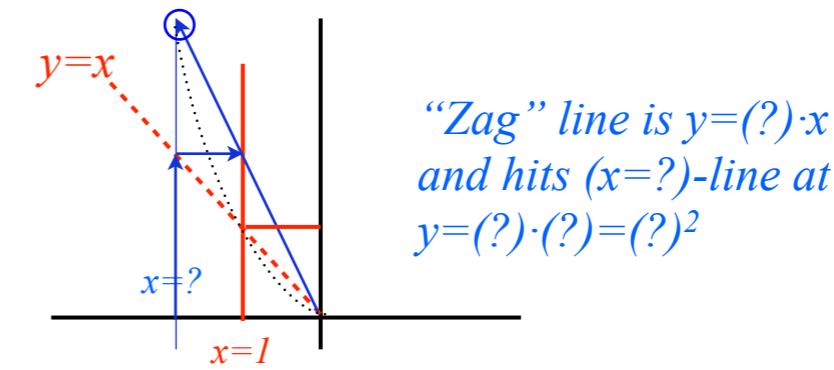
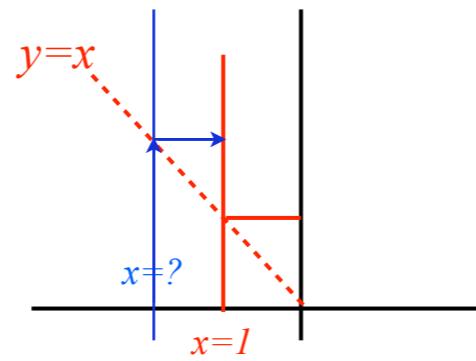
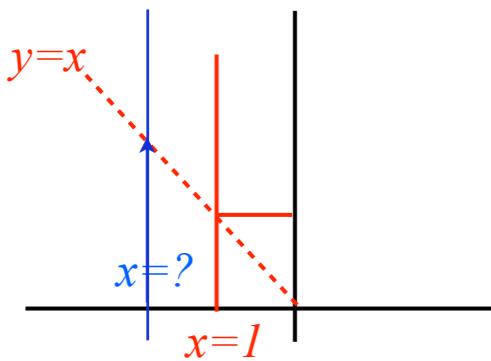
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Unit 1
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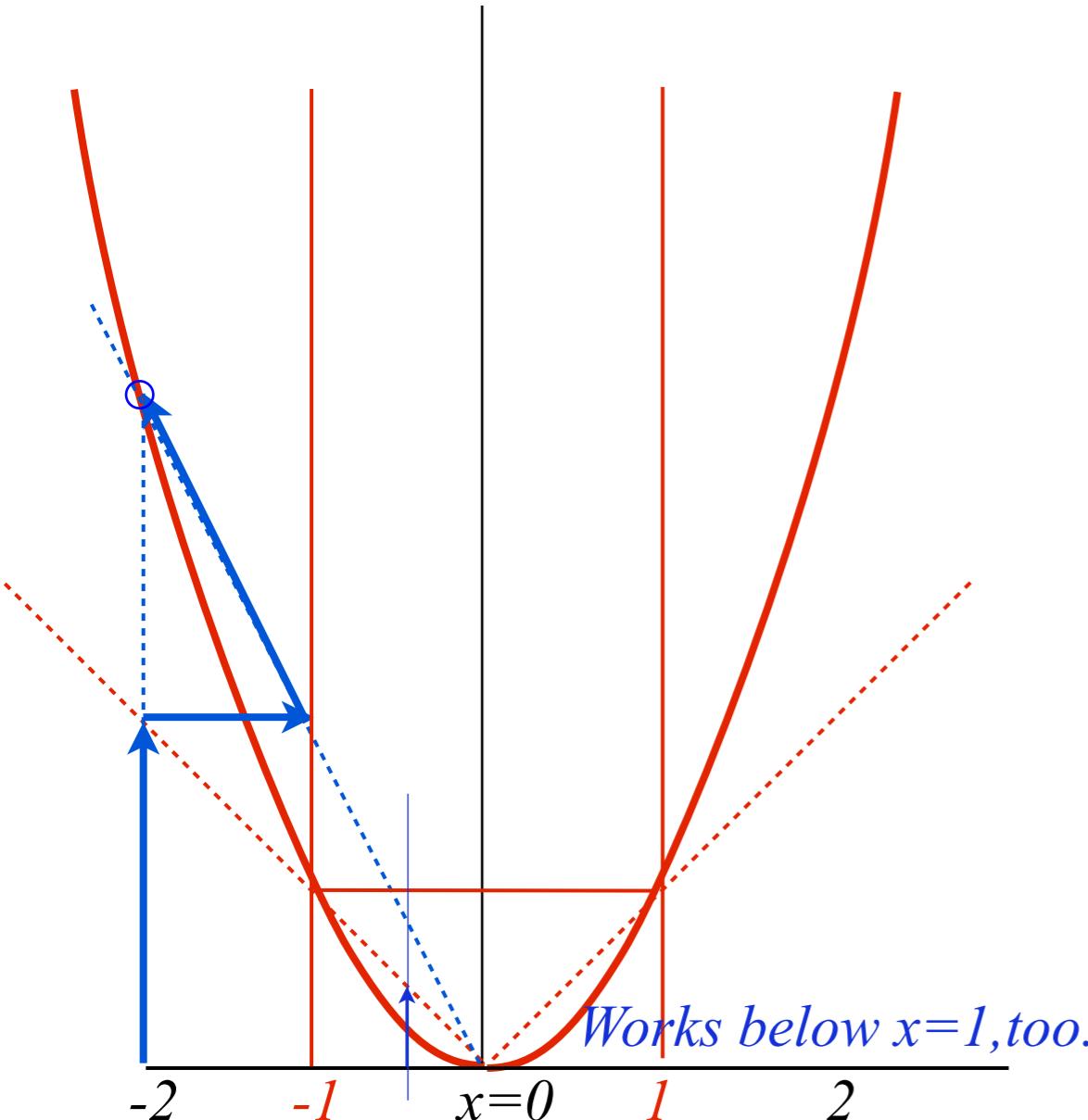
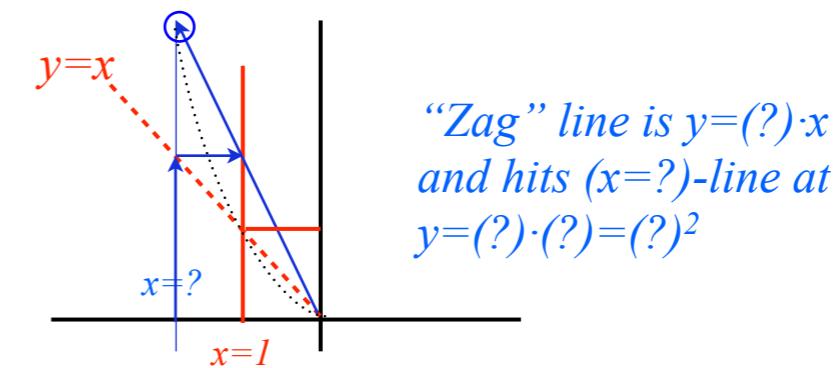
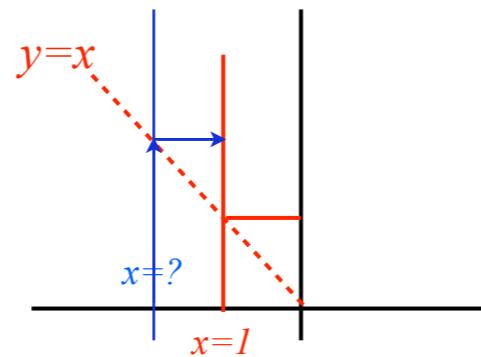
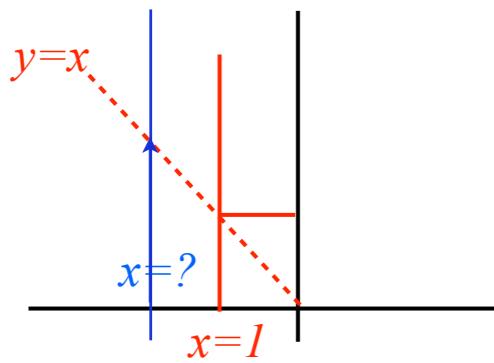
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Unit 1
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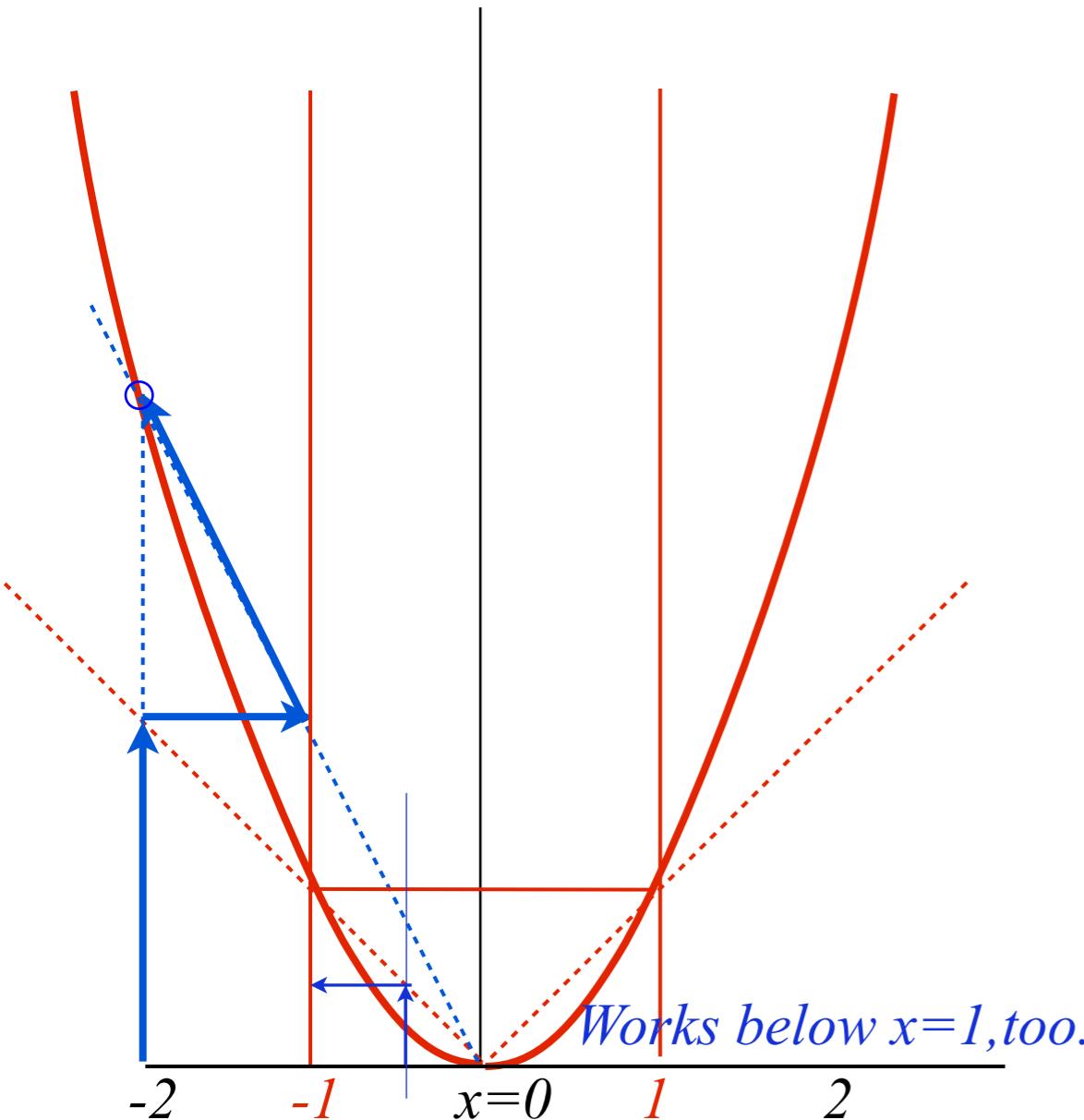
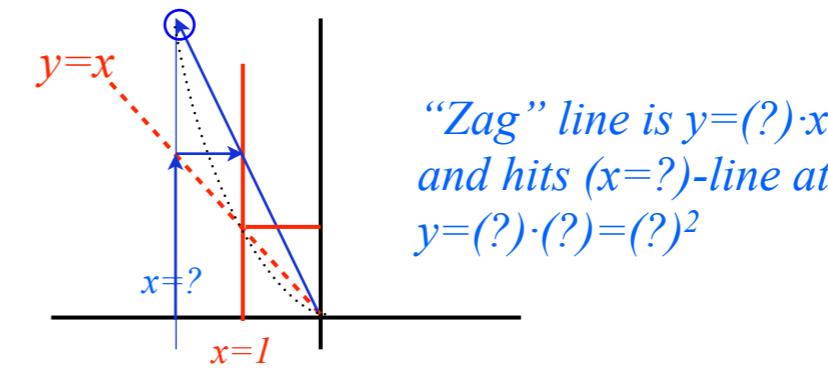
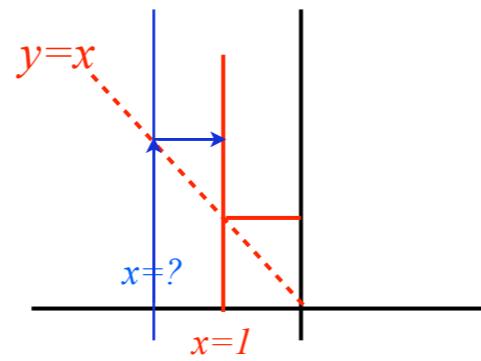
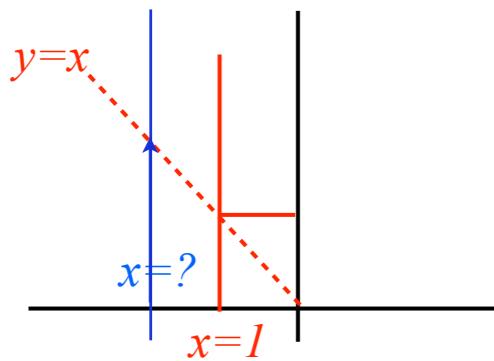
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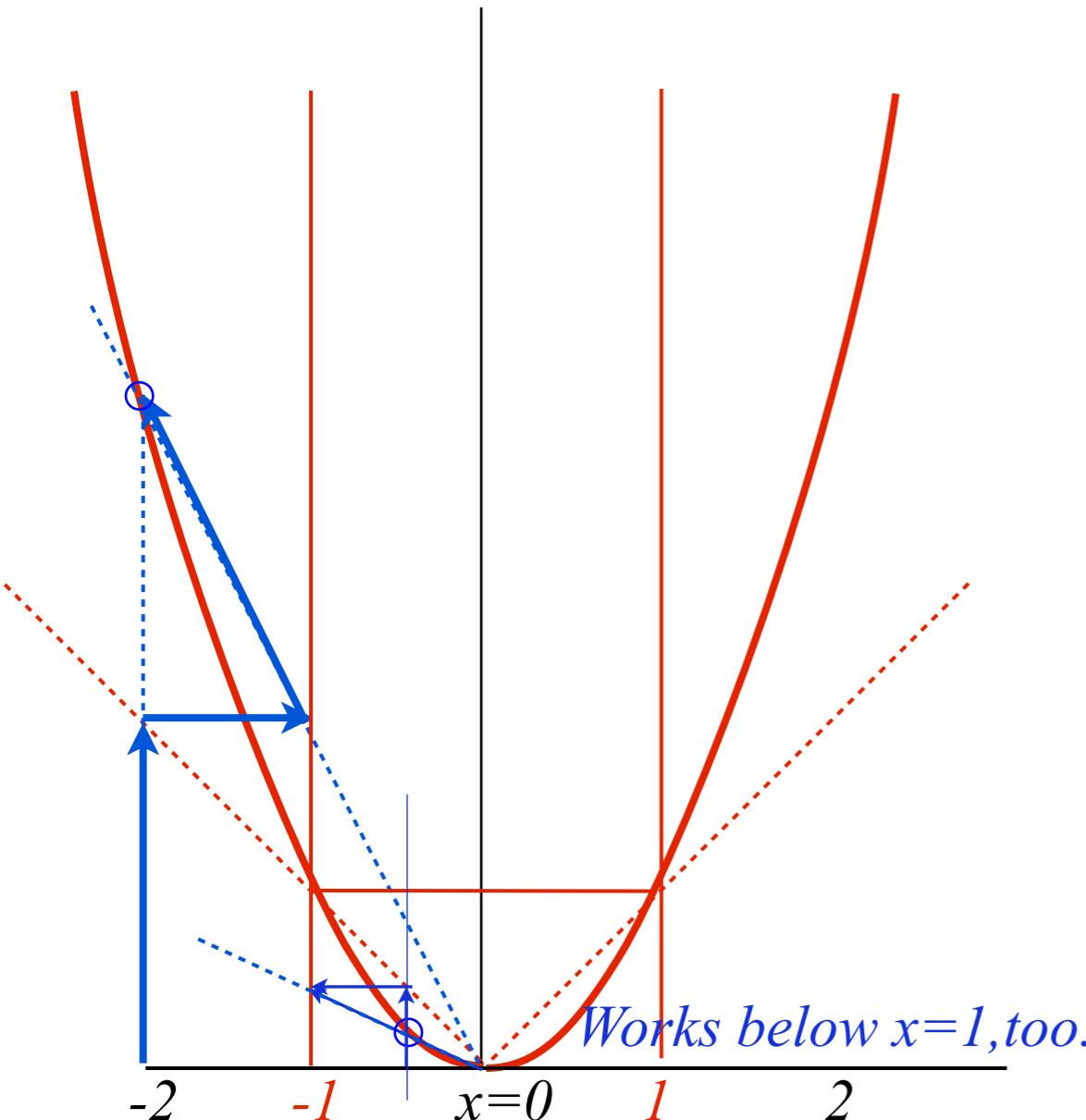
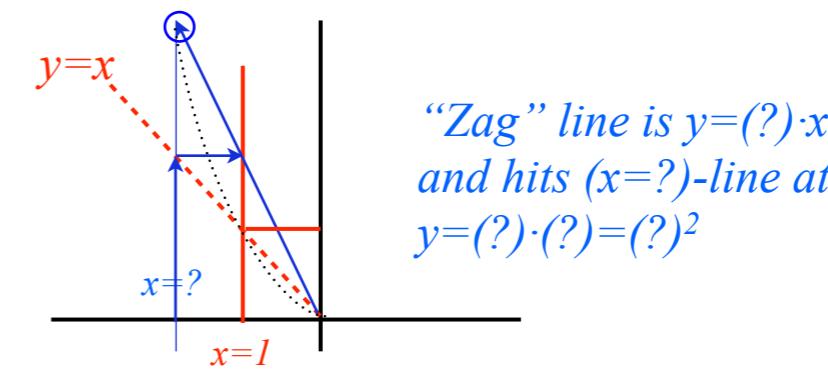
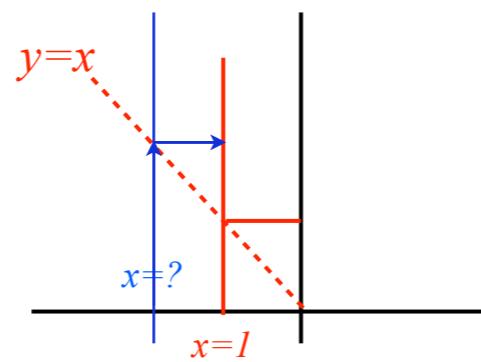
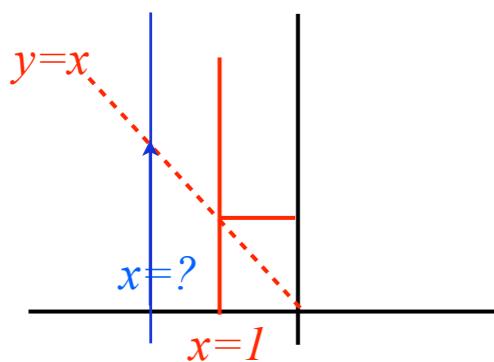
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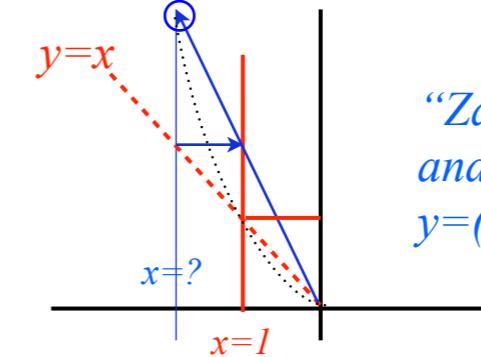
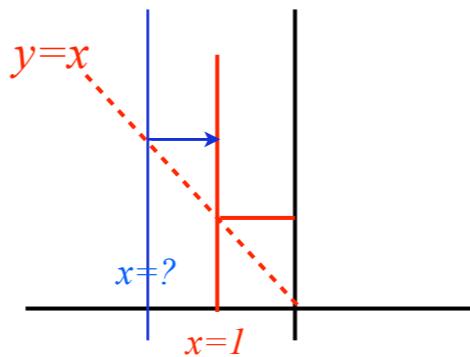
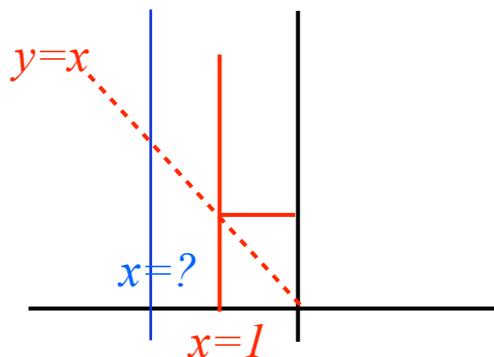
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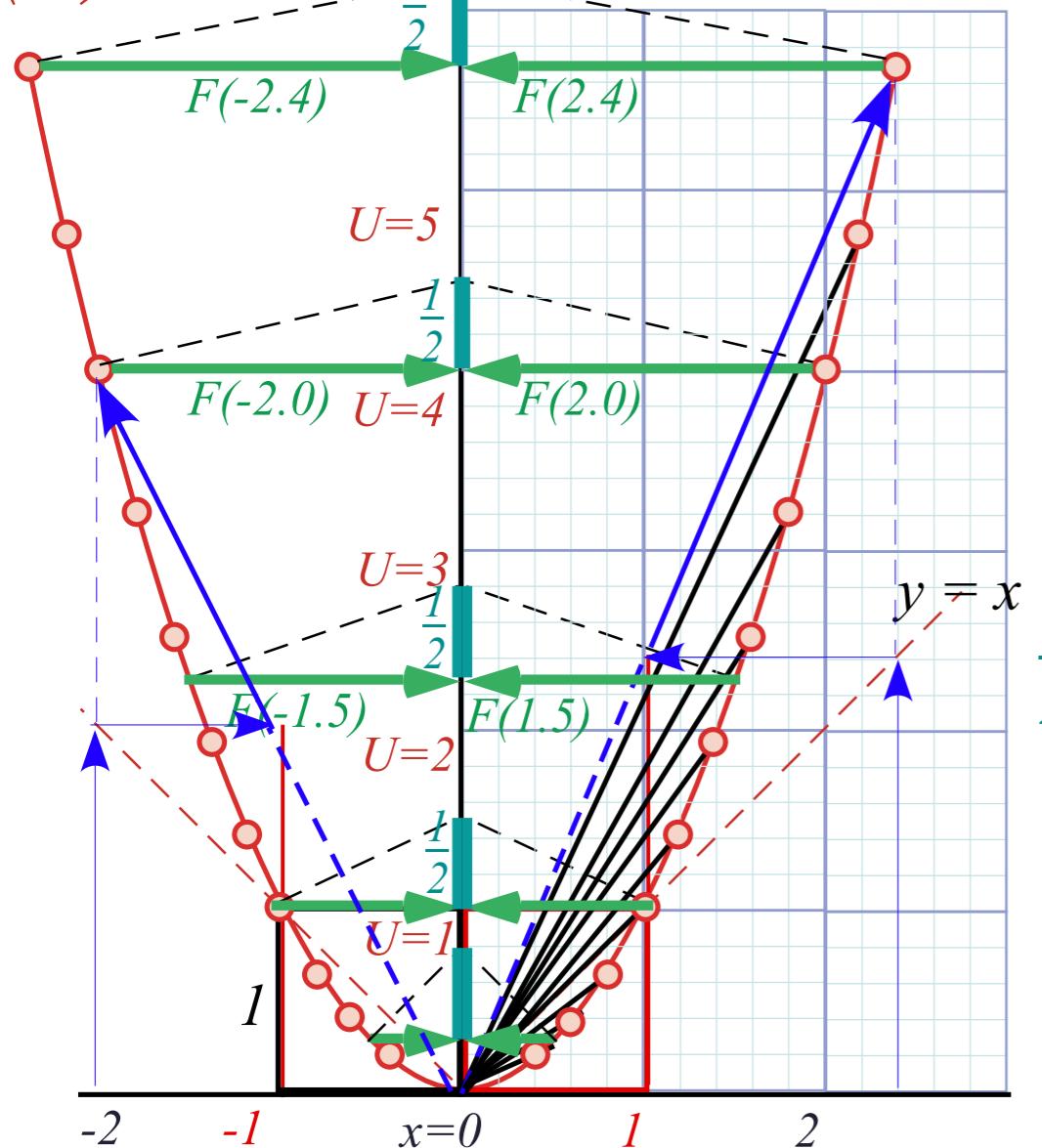
Unit 1
Fig. 9.1

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(a) Oscillator potential $U(x)=x^2$

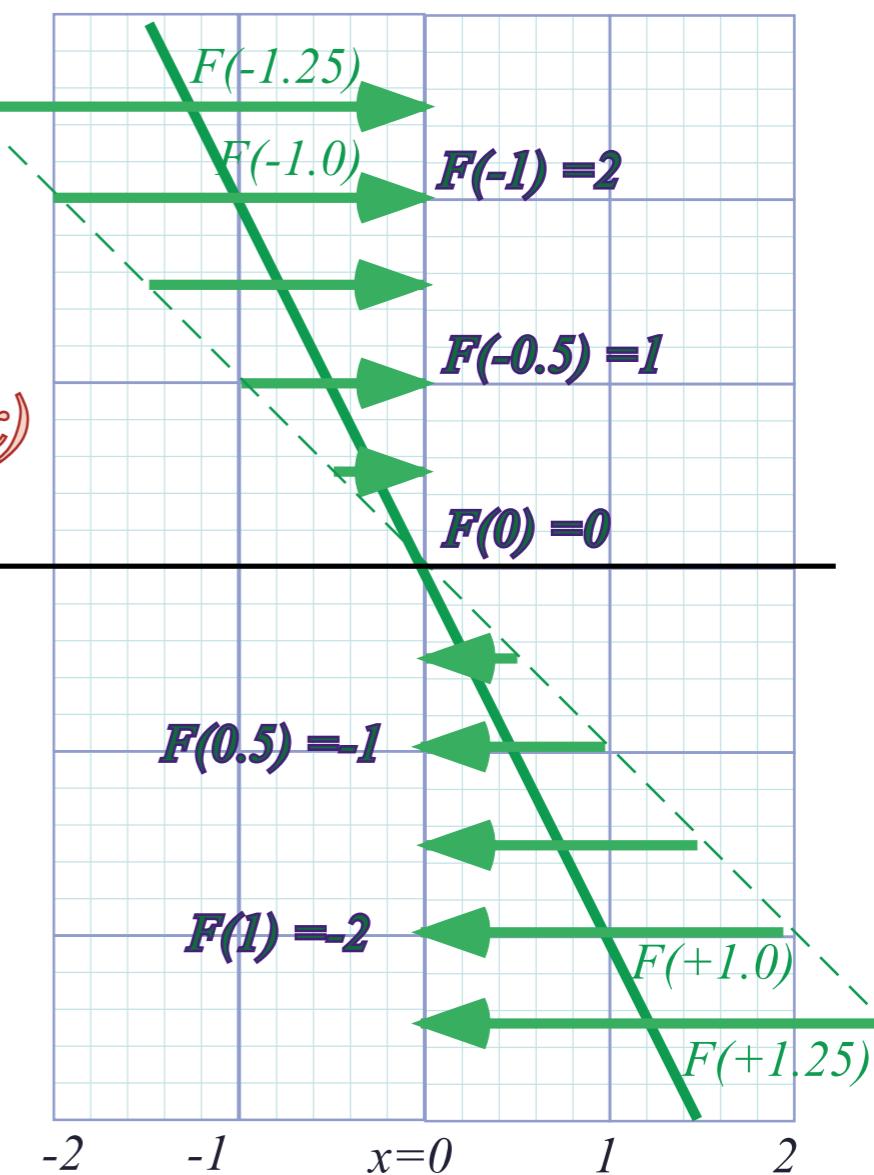


$$\frac{F(x)}{1} = -\frac{\Delta U}{\Delta x}$$

$\frac{1}{2} F(x)$ $U(x)$
 $\frac{1}{2} \Delta U$ Δx

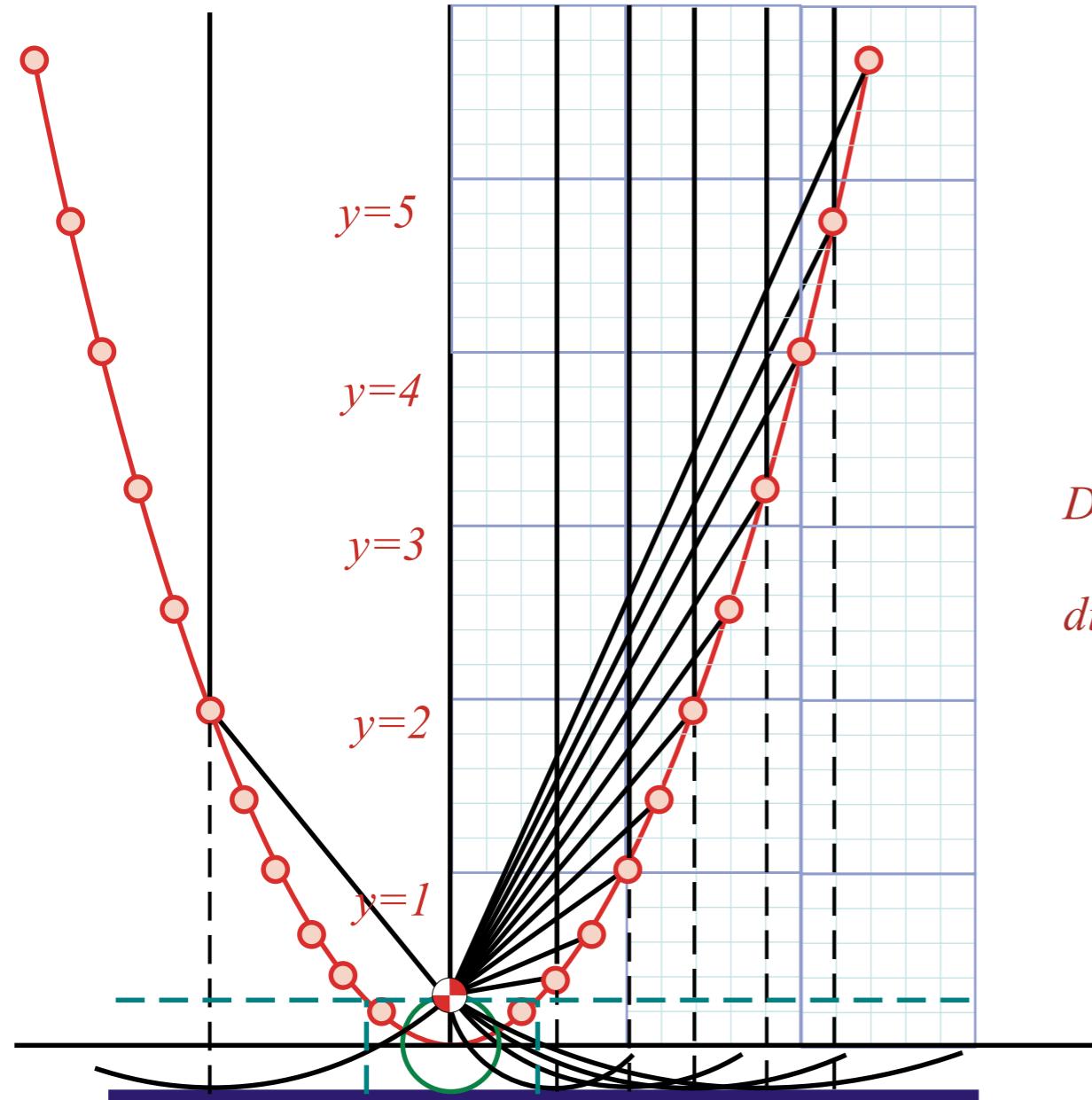
Unit 1
Fig. 9.1

(b) Hooke-Law Force $F(x) = -2x$

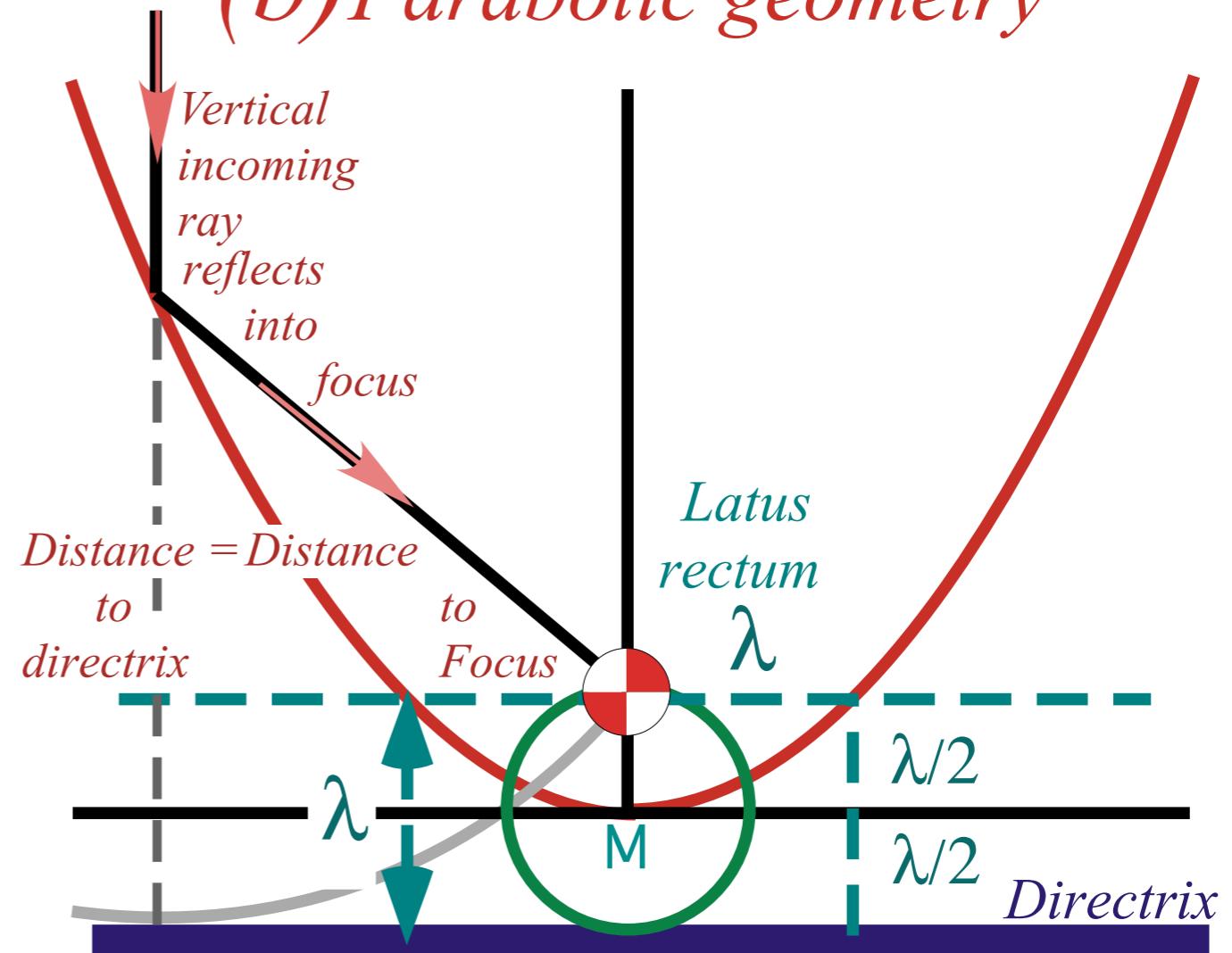


A more conventional parabolic geometry... (uses focal point)

(a) Parabolic Reflector $y=x^2$



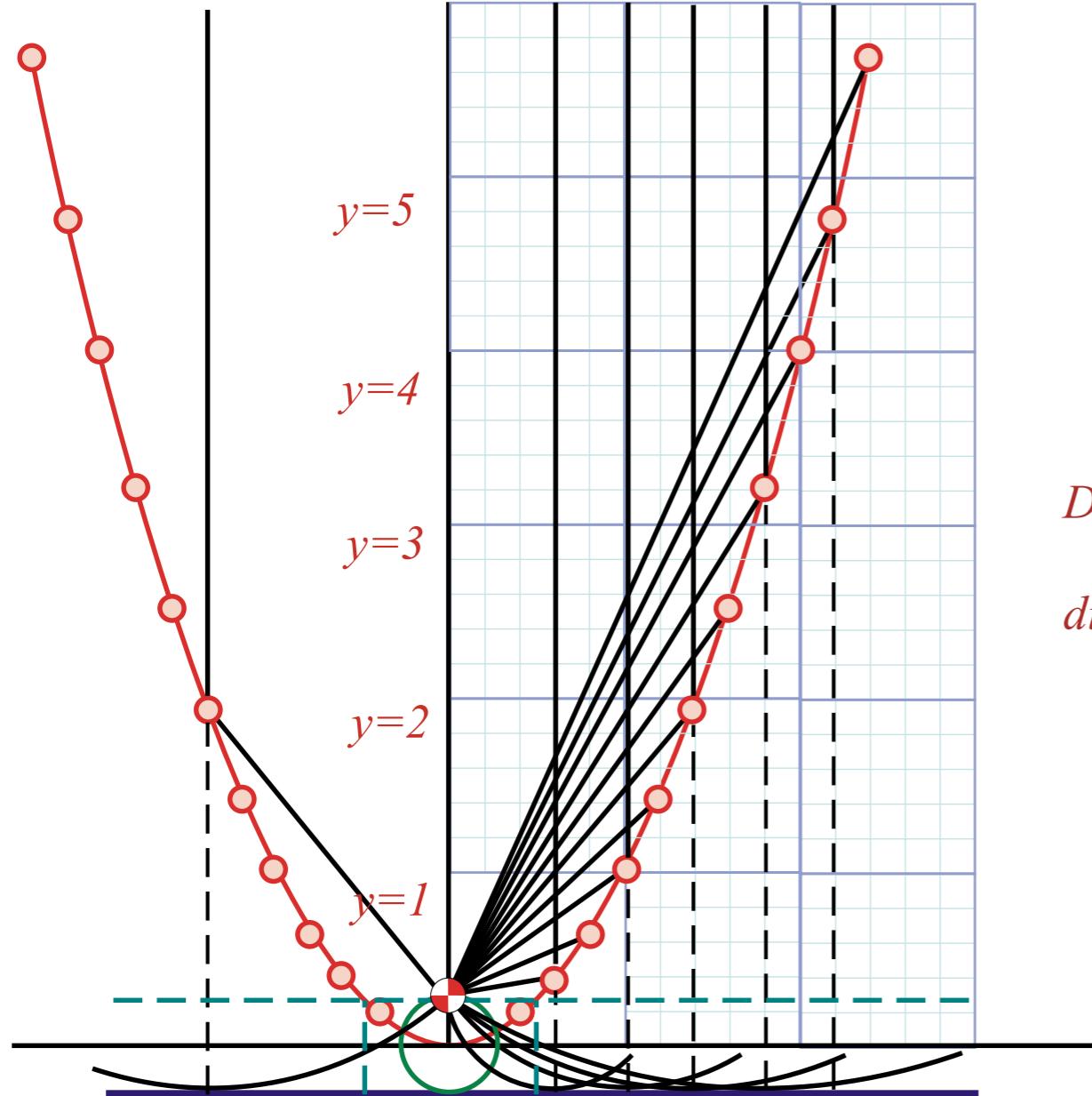
(b) Parabolic geometry



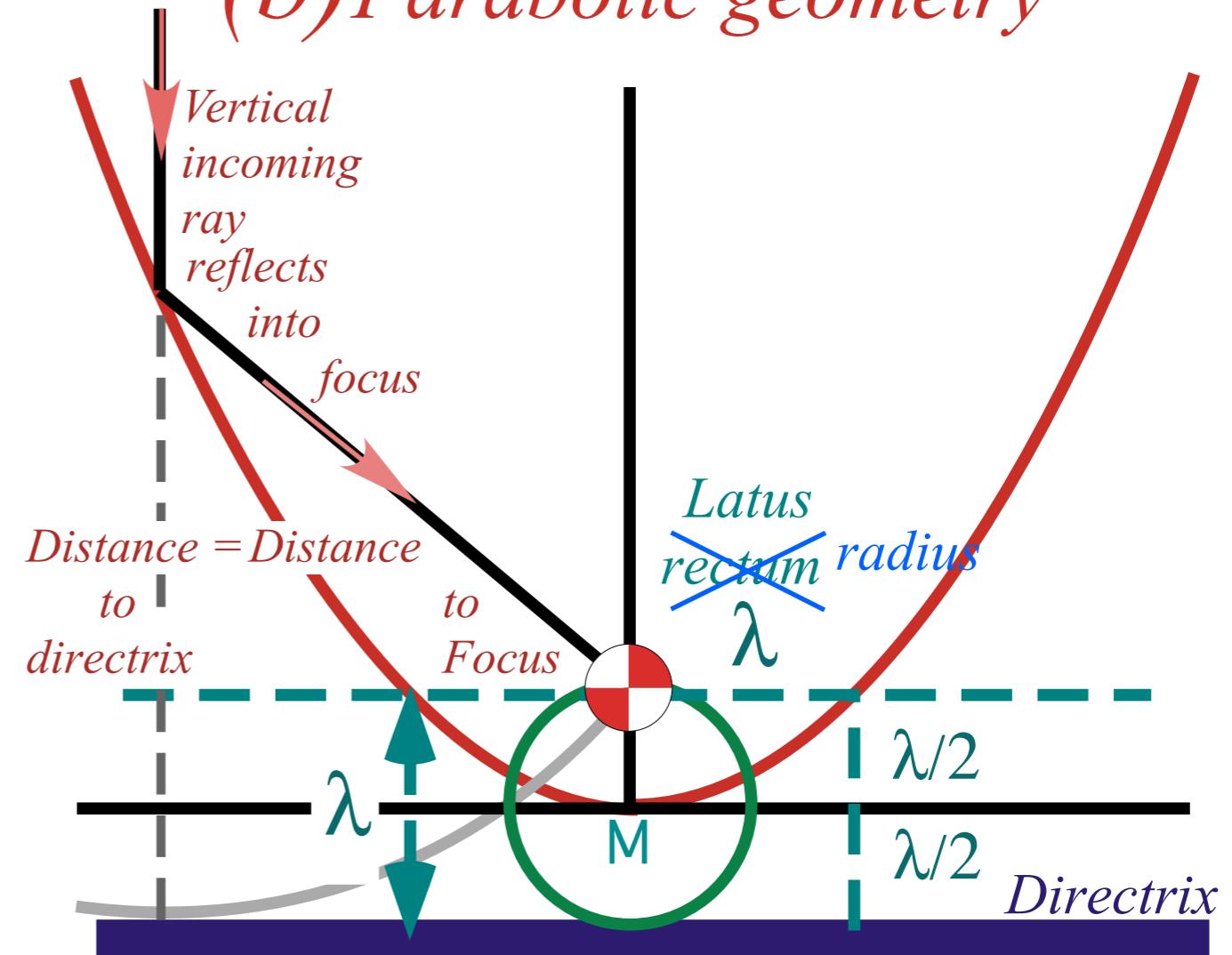
Unit 1
Fig. 9.3

A more conventional parabolic geometry...

(a) Parabolic Reflector $y=x^2$



(b) Parabolic geometry

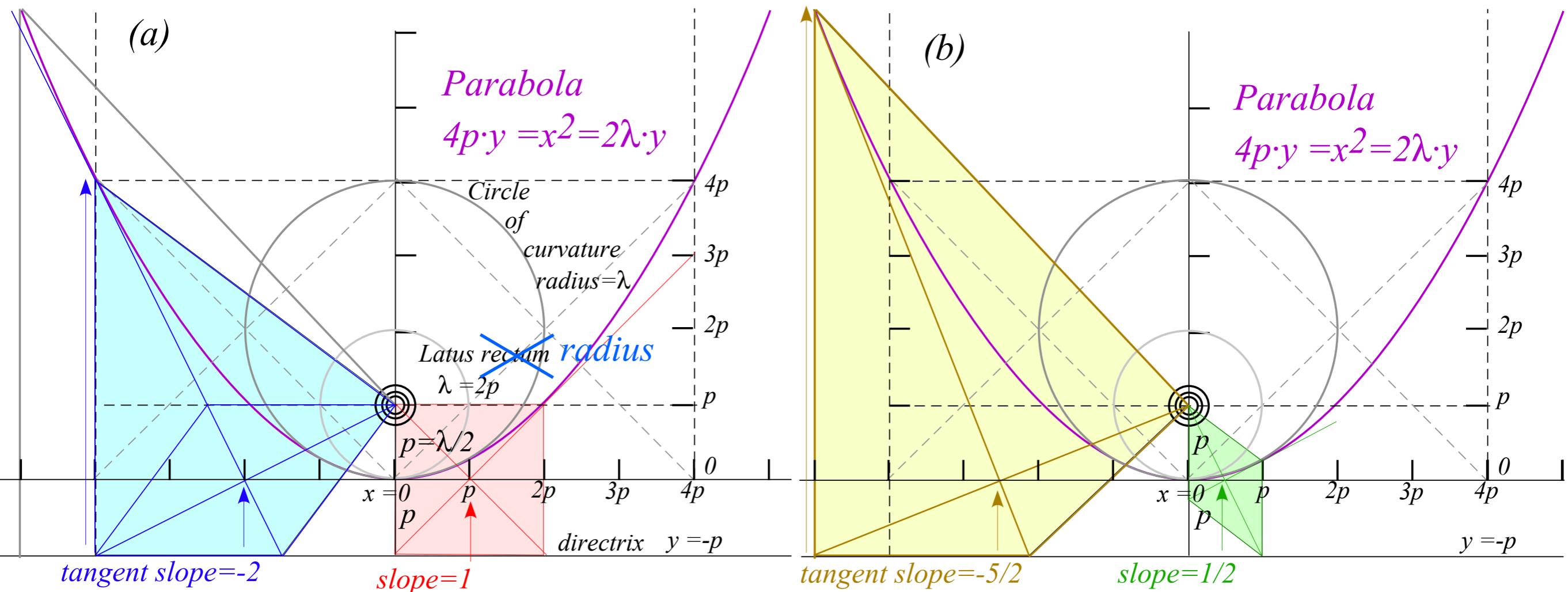


Better name[†] for λ : *latus radius*

[†] Old term *latus rectum* is exclusive copyright of
X-Treme Roidrage Gyms
Venice Beach, CA 90017

Unit 1
Fig. 9.3

...conventional parabolic geometry...carried to extremes...



Unit 1
Fig. 9.4

Geometry of common power-law potentials

Geometric (Power) series

“Zig-Zag” exponential geometry

Projective or perspective geometry

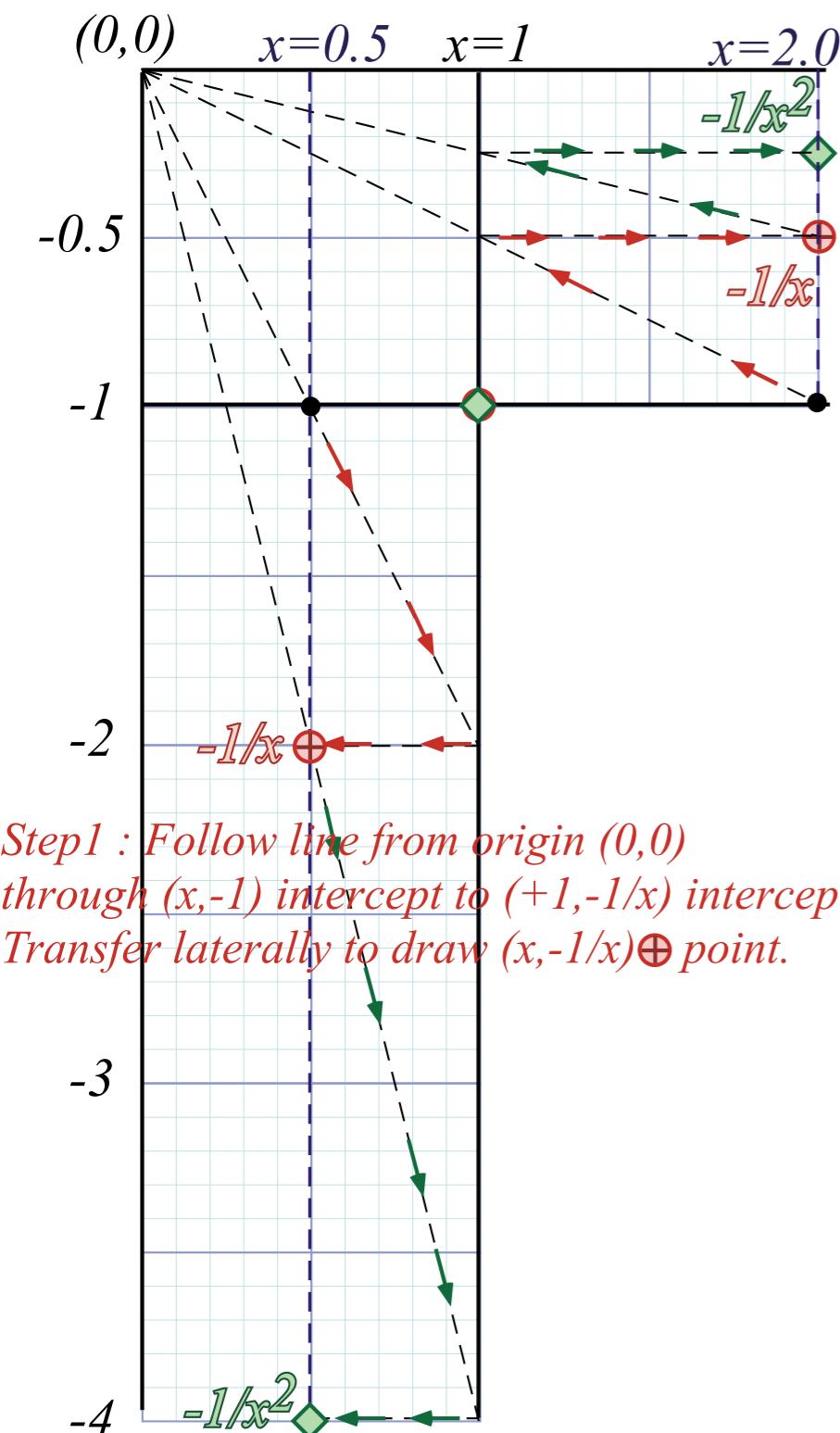
Parabolic geometry of harmonic oscillator $kr^2/2$ potential and $-kr^l$ force fields

→ *Coulomb geometry of $-1/r$ -potential and $-1/r^2$ -force fields*

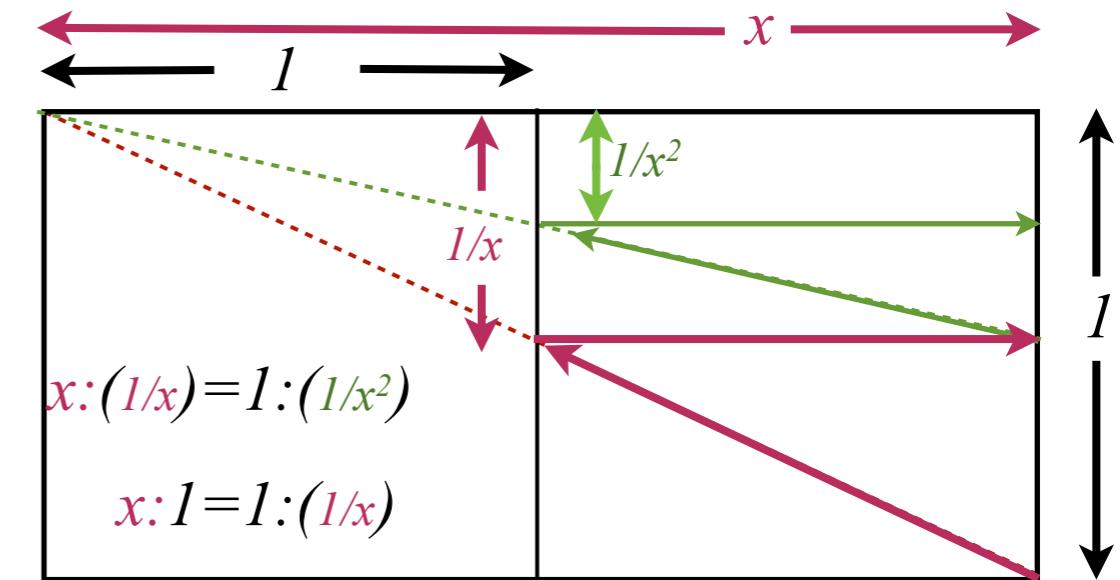
Compare mks units of Coulomb Electrostatic vs. Gravity

Unit 1
Fig. 9.4

Coulomb geometry
Force and Potential
 $F(x) = -1/r^2$ $U(x) = -1/r$

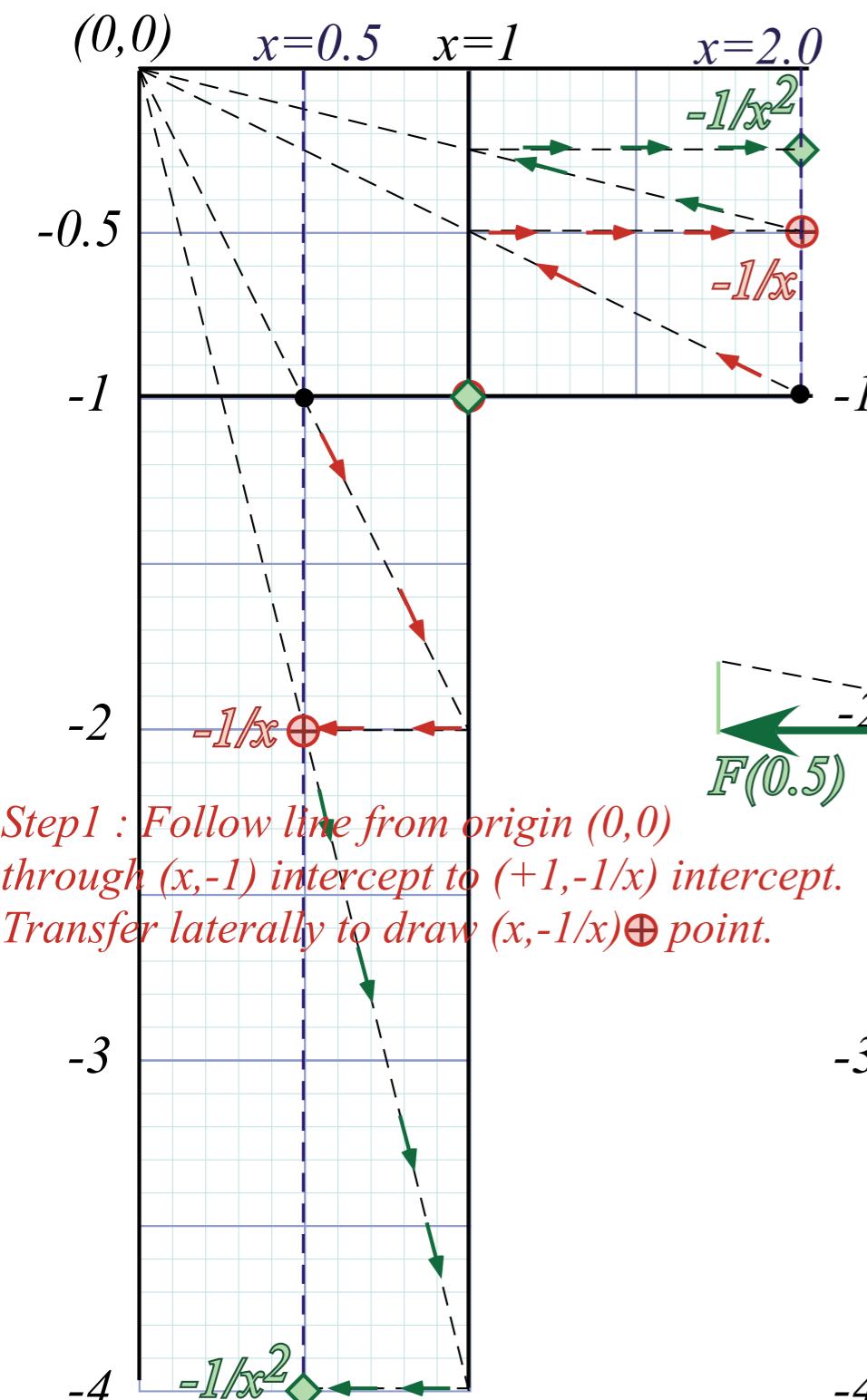


Step2 : Follow line from origin (0,0) through $(x,-1/x)$ point \oplus to $(+1,-1/x^2)$ intercept. Transfer laterally to draw $(x,-1/x^2)\diamond$ point.

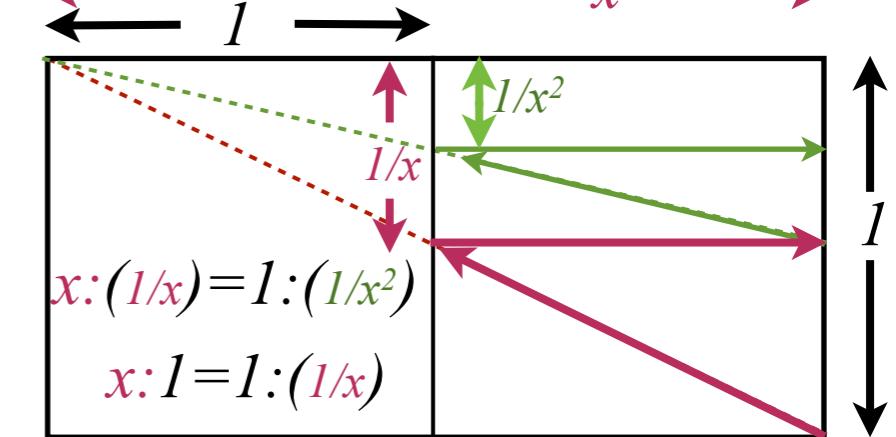
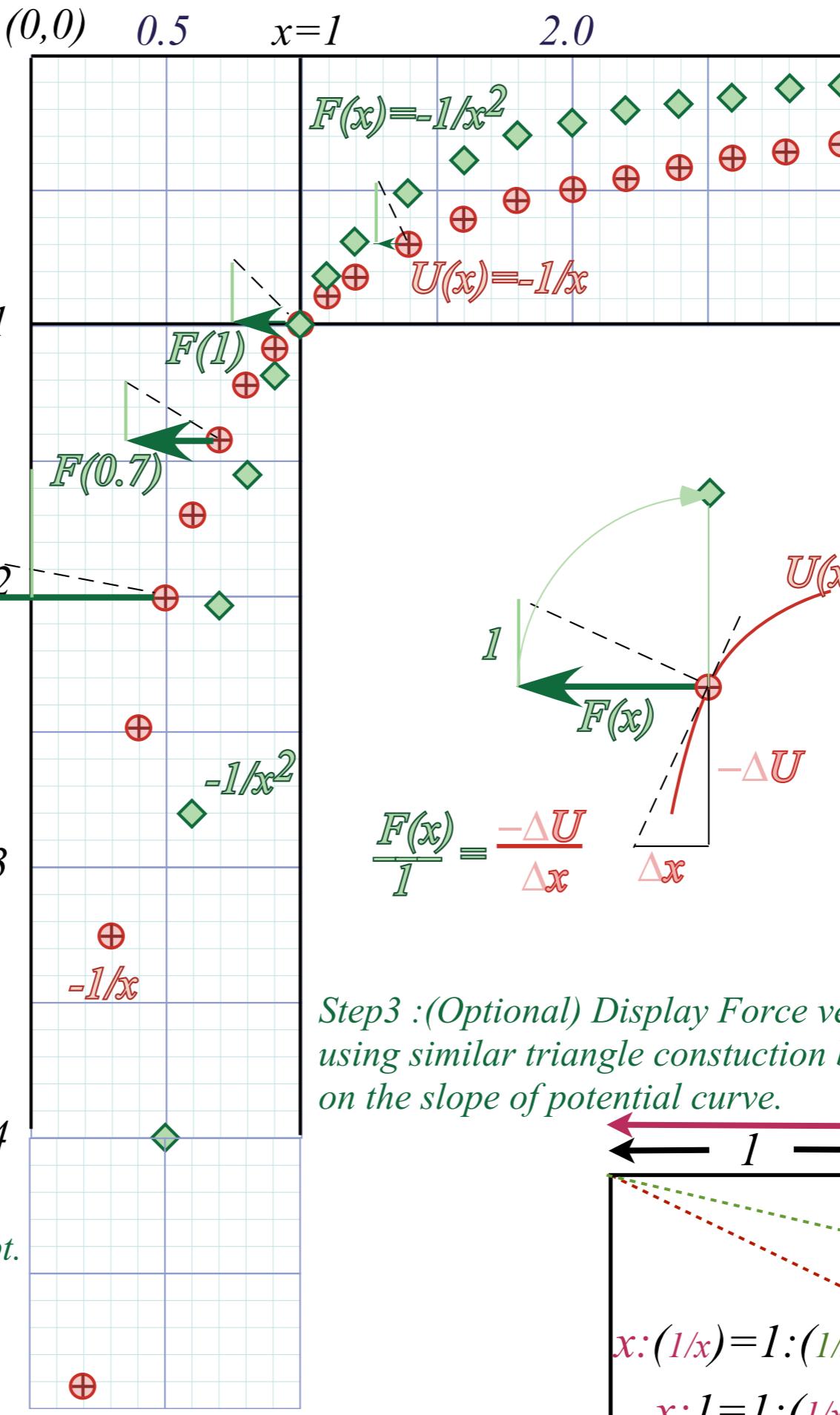


Unit 1
Fig. 9.4

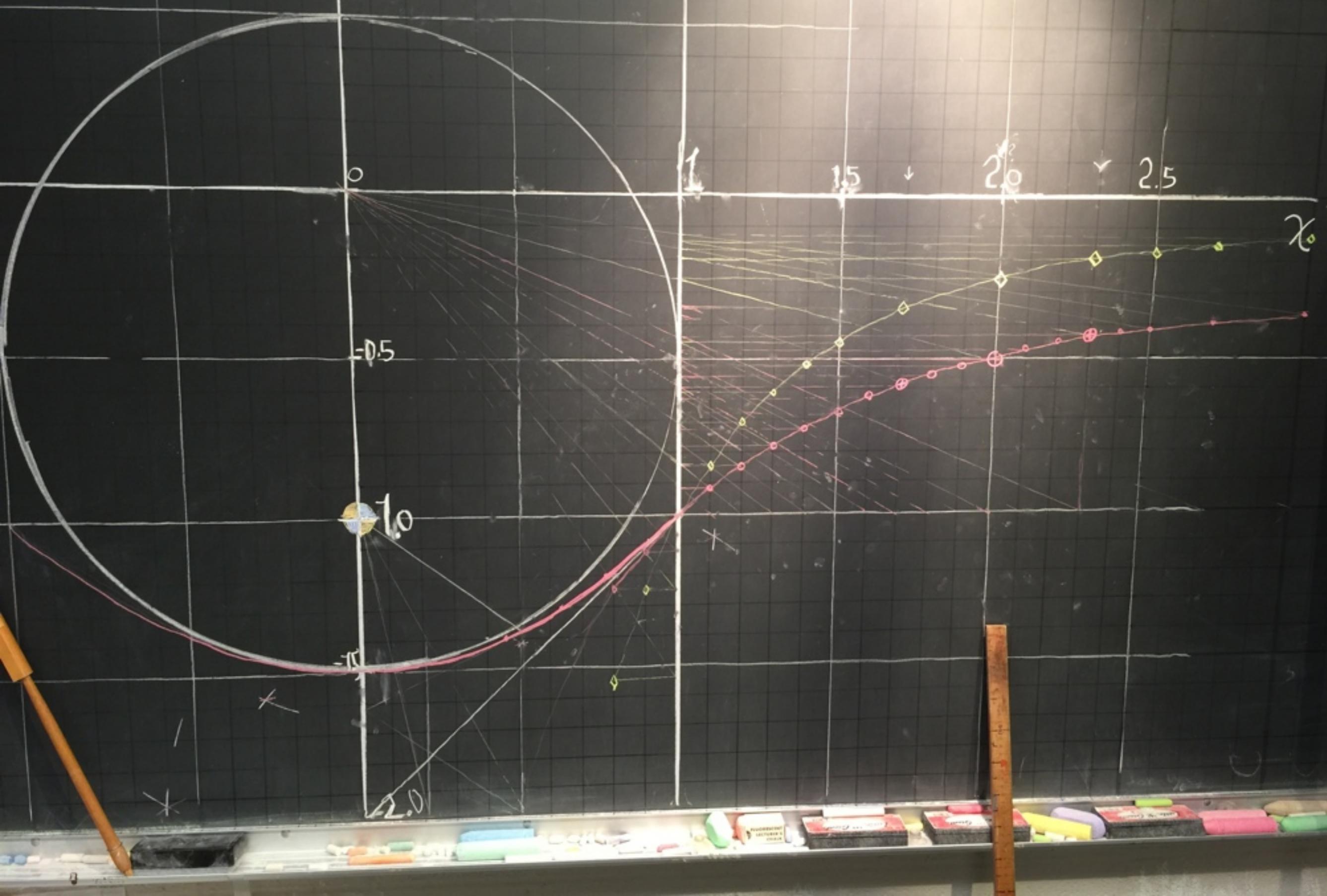
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Step2 : Follow line from origin $(0,0)$ through $(x,-1/x)$ point \oplus to $(+1,-1/x^2)$ intercept.
Transfer laterally to draw $(x,-1/x^2)\diamond$ point.



$$V(x) = \frac{1}{x}$$
$$F(x) = -\frac{1}{x^2}$$



Geometry of common power-law potentials

Geometric (Power) series

“Zig-Zag” exponential geometry

Projective or perspective geometry

Parabolic geometry of harmonic oscillator $kr^2/2$ potential and $-kr^l$ force fields

Coulomb geometry of $-1/r$ -potential and $-1/r^2$ -force fields

→ *Compare mks units of Coulomb Electrostatic vs. Gravity*

Compare mks units for Coulomb fields

1. Electrostatic force between q (Coulombs) and Q (C.)

$$F^{elec.}(r) = \pm \frac{1}{4\pi\epsilon_0} \frac{qQ}{r^2} \quad \text{where: } \frac{1}{4\pi\epsilon_0} \approx \boxed{?.? \cdot 10^?}$$

$\frac{\text{Newtons} \cdot \text{meter} \cdot \text{square}}{\text{per square Coulomb}}$

Compare mks units for Coulomb fields

1. Electrostatic force between q (Coulombs) and Q (C.)

!!!!

$$F^{elec.}(r) = \pm \frac{1}{4\pi\epsilon_0} \frac{qQ}{r^2} \quad \text{where: } \frac{1}{4\pi\epsilon_0} \stackrel{\sim 9E9 \sim 10^{10}}{\cong} 9,000,000,000 \frac{\text{Newtons} \cdot \text{meter} \cdot \text{square}}{\text{per square Coulomb}}$$

Compare mks units for Coulomb fields

1. Electrostatic force between q (Coulombs) and Q (C.) !!!!

$$F^{elec.}(r) = \pm \frac{1}{4\pi\epsilon_0} \frac{qQ}{r^2} \quad \text{where: } \frac{1}{4\pi\epsilon_0} \cong 9,000,000,000 \frac{\text{Newtons} \cdot \text{meter} \cdot \text{square}}{\text{per square Coulomb}}$$

More precise value for electrostatic constant : $1/4\pi\epsilon_0 = 8.987,551 \cdot 10^9 \text{ Nm}^2/\text{C}^2 \sim 9 \cdot 10^9 \sim 10^{10}$

quantum of charge: $|e| = 1.6022 \cdot 10^{-19} \text{ Coulomb}$



Repulsive (+)(+) or (-)(-)

Attractive (+)(-) or (-)(+)

...but 1 Ampere = 1 Coulomb/sec.

Compare mks units for Coulomb fields

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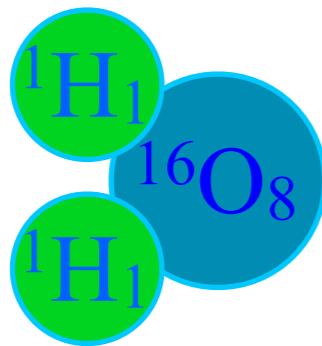
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"Fingertip Physics" of Ch. 8 notes that 1 $(\text{cm})^3$ = 1 gm of water (1/18 Mole) has (1/18) 6 \cdot 10^{23} molecules Avogadro's Number

$\sim 0.3 \cdot 10^{23}$

H_2O Molecular weight ~ 18



Compare mks units for Coulomb fields

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$$F^{elec.}(r) = \pm \frac{1}{4\pi\epsilon_0} \frac{qQ}{r^2} \quad \text{where: } \frac{1}{4\pi\epsilon_0} \cong 9,000,000,000 \frac{\text{Newtons} \cdot \text{meter} \cdot \text{square}}{\text{per square Coulomb}}$$

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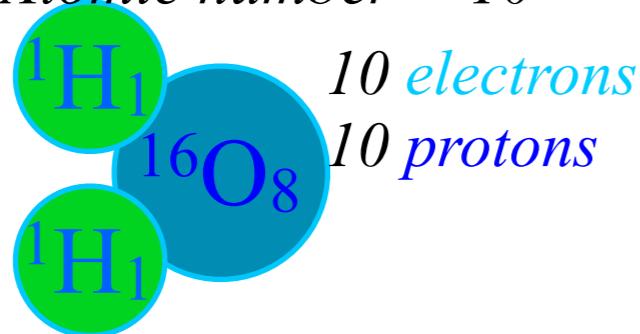
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"Fingertip Physics" of Ch. 9 notes that $1 \text{ (cm)}^3 = 1 \text{ gm}$ of water (1/18 Mole) has (1/18) $6 \cdot 10^{23}$ molecules or $\sim 3 \cdot 10^{23}$ electrons
 $\sim 0.3 \cdot 10^{23}$ and $\sim 3 \cdot 10^{23}$ protons.

H_2O Molecular weight ~ 18

Atomic number = 10



Compare mks units for Coulomb fields

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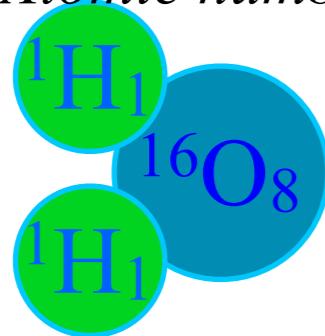
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 $\sim 0.3 \cdot 10^{23}$ and $\sim 3 \cdot 10^{23}$ protons.

H_2O Molecular weight ~ 18

Atomic number = 10



10 electrons That is $\sim -3 \cdot 10^{23} 1.6022 \cdot 10^{-19} \text{ Coulomb}$ or about $-0.5 \cdot 10^{+5} \text{ C}$ or $-50,000 \text{ Coulomb}$
 10 protons plus $\sim +3 \cdot 10^{23} 1.6022 \cdot 10^{-19} \text{ Coulomb}$ or about $+0.5 \cdot 10^{+5} \text{ C}$ or $+50,000 \text{ Coulomb}$

Equals zero total charge

Compare mks units for Coulomb fields

1. Electrostatic force between q (Coulombs) and Q (C.) !!!!

$$F^{elec.}(r) = \pm \frac{1}{4\pi\epsilon_0} \frac{qQ}{r^2} \quad \text{where: } \frac{1}{4\pi\epsilon_0} \approx 9,000,000,000 \frac{\text{Newtons} \cdot \text{meter} \cdot \text{square}}{\text{per square Coulomb}}$$

$\sim 9E9 \sim 10^{10}$

More precise value for electrostatic constant : $1/4\pi\epsilon_0 = 8.987,551 \cdot 10^9 \text{ Nm}^2/\text{C}^2 \sim 9 \cdot 10^9 \sim 10^{10}$



Repulsive (+)(+) or (-)(-)
Attractive (+)(-) or (-)(+)
vs
Always Attractive (so far)

↑COMPARE!↓

quantum of charge: $|e| = 1.6022 \cdot 10^{-19} \text{ Coulomb}$

BIG
vs
small



2. Gravitational force between m (kilograms) and M (kg.)

$$F^{grav.}(r) = -G \frac{mM}{r^2} \quad \text{where: } G = \boxed{?.? \cdot 10^?} \frac{\text{Newtons} \cdot \text{meter} \cdot \text{square}}{\text{per square Coulomb}}$$

Compare mks units for Coulomb fields

1. Electrostatic force between q (Coulombs) and Q (C.) !!!!

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$\sim 2/3 10^{-10} \sim 10^{-10}$

More precise value for gravitational constant : $G=6.67384(80) \cdot 10^{-11} \text{ Nm}^2/\text{C}^2 \sim (2/3)10^{-10}$

Compare mks units for Coulomb fields

1. Electrostatic force between q (Coulombs) and Q (C.)

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Repulsive (+)(+) or (-)(-)
Attractive (+)(-) or (-)(+)

quantum of charge: $|e|=1.6022 \cdot 10^{-19}$ Coulomb

Discussion of repulsive force and PE in Ch. 9...

1(a). Electrostatic potential energy between q (Coulombs) and Q (C.)

$$U(r) = \frac{1}{4\pi\epsilon_0} \frac{qQ}{r} \quad \text{where: } \frac{1}{4\pi\epsilon_0} \cong 9,000,000,000 \frac{\sim 9E9 \sim 10^{10} \text{ Joule}}{\text{per square Coulomb}} \quad !!!!$$

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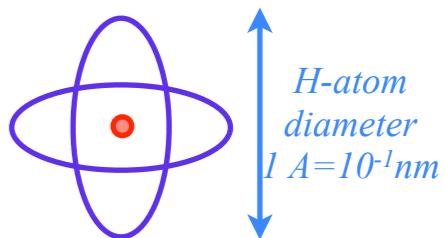
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Nuclear size $\sim 10^{-15} \text{ m} = 1 \text{ femtometer} = 1 \text{ fm}$


Atomic size $\sim 1 \text{ Angstrom} = 10^{-10} \text{ m}$



Compare mks units for Coulomb fields

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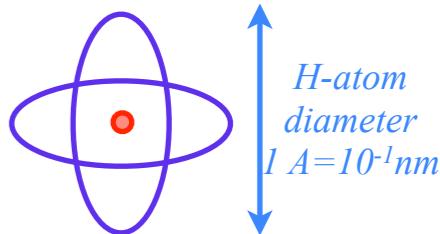
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Big molecule $\sim 10 \text{ Angstrom} = 10^{-9} \text{ m} = 1 \text{ nanometer} = 1 \text{ nm}$



Compare mks units for Coulomb fields

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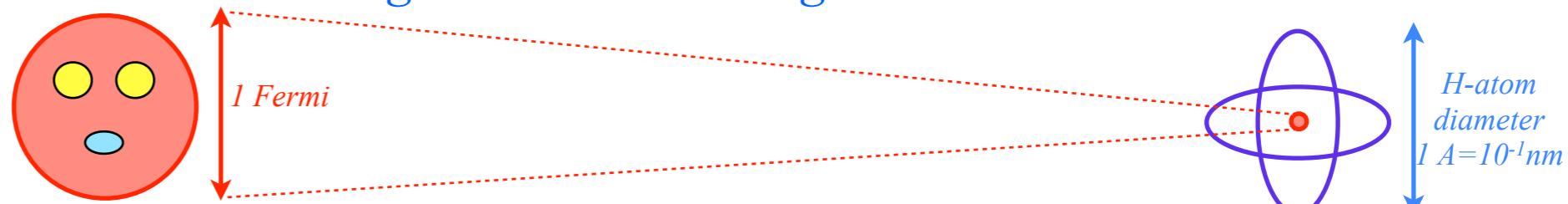
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Big molecule $\sim 10 \text{ Angstrom} = 10^{-9} \text{ m} = 1 \text{ nanometer} = 1nm$

also: $1fm = 10^{-13} \text{ cm} = 1 \text{ Fermi} = 1Fm$



Compare mks units for Coulomb fields

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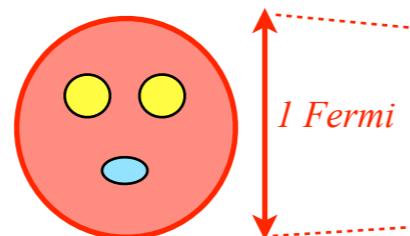
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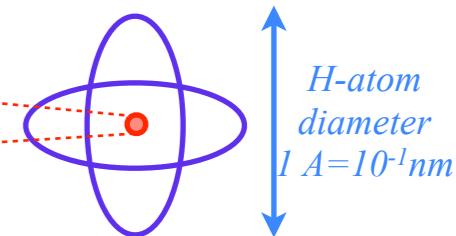
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Atomic size $\sim 1 \text{ Angstrom} = 10^{-10} \text{ m}$
Big molecule $\sim 10 \text{ Angstrom} = 10^{-9} \text{ m} = 1 \text{ nanometer} = 1nm$



nuclear radii are 100,000 to 1,000,000 times smaller than atomic/chemical radii

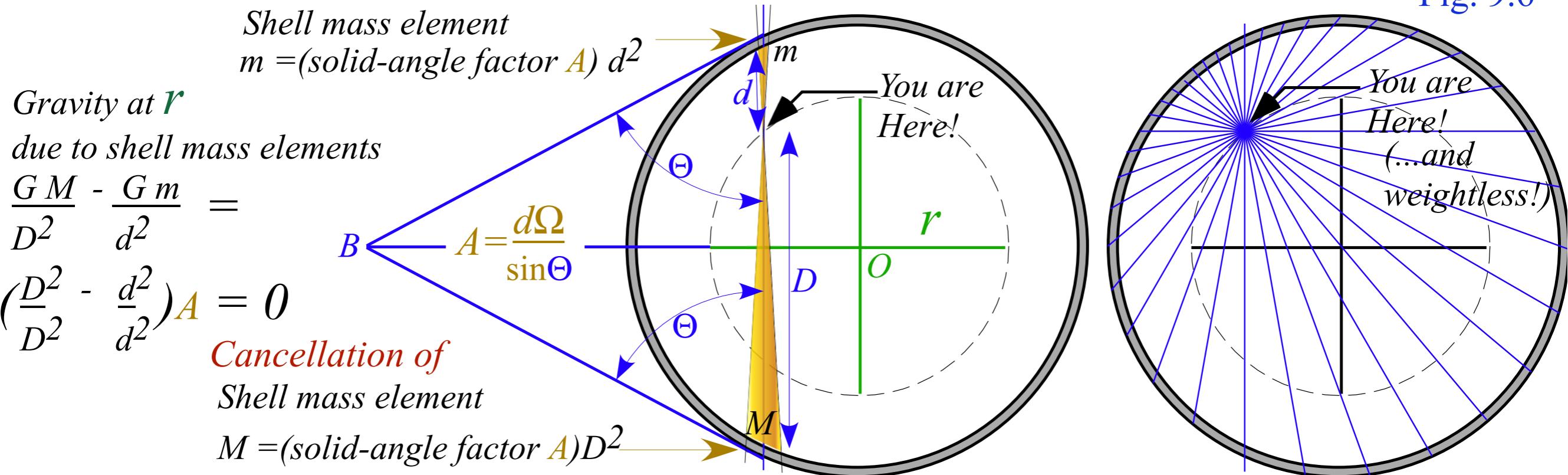
...so nuclear qQ/r energy 100,000 to 1,000,000 times **bigger** than of atomic/chemical...

Geometry of idealized “Sophomore-physics Earth”

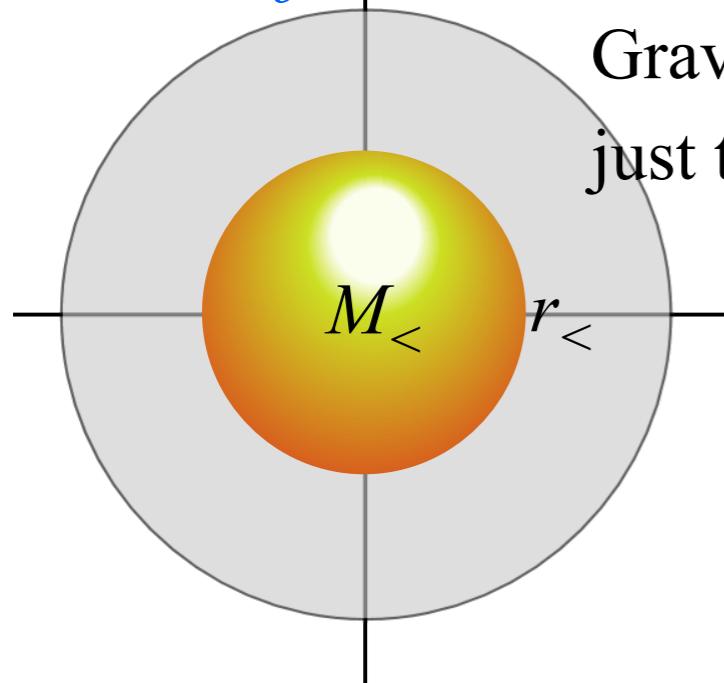
- *Coulomb field outside* *Isotropic Harmonic Oscillator (IHO) field inside*
- Contact-geometry of potential curve(s)*
- “Crushed-Earth” models: 3 key energy “steps” and 4 key energy “levels”*
- Earth matter vs nuclear matter:*
- Introducing the “neutron starlet” and “**Black-Hole-Earth**”*

Coulomb force vanishes inside-spherical shell (Gauss-law)

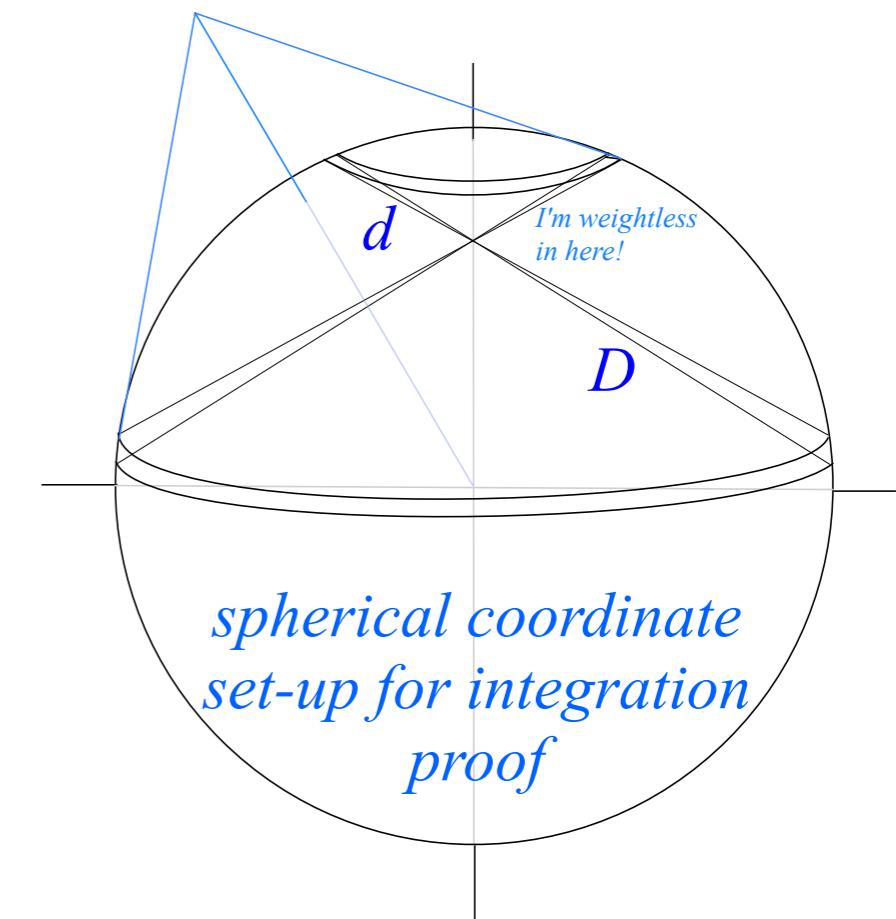
Unit 1
Fig. 9.6



Coulomb force inside-spherical body due to stuff below you, only.

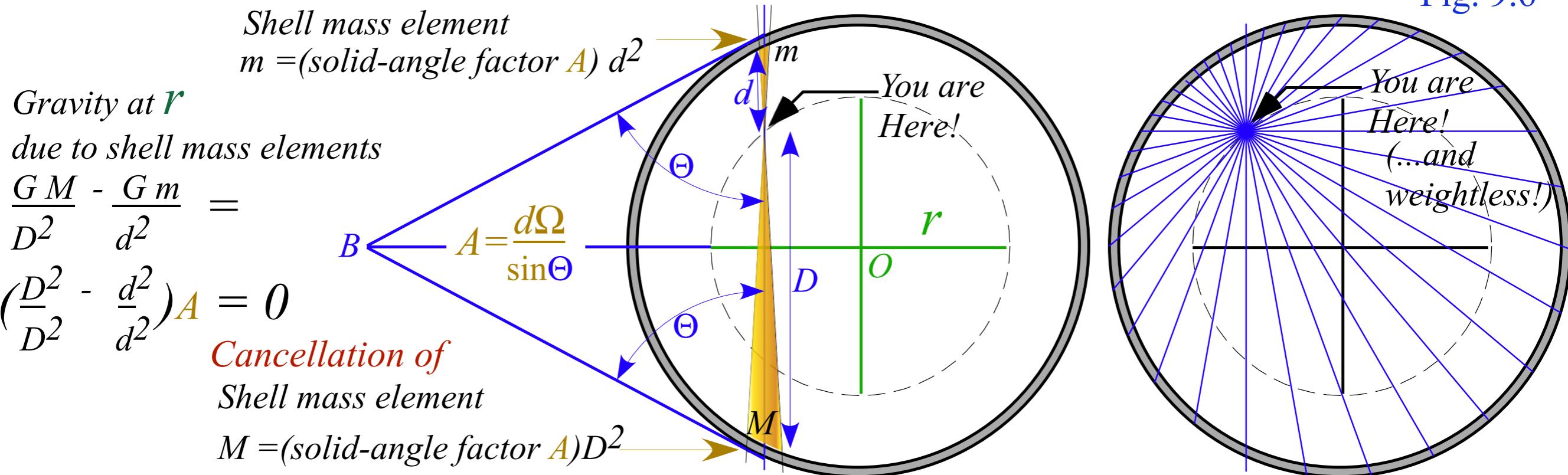


Gravitational force at $r_<$ is just that of planet $M_<$ below $r_<$

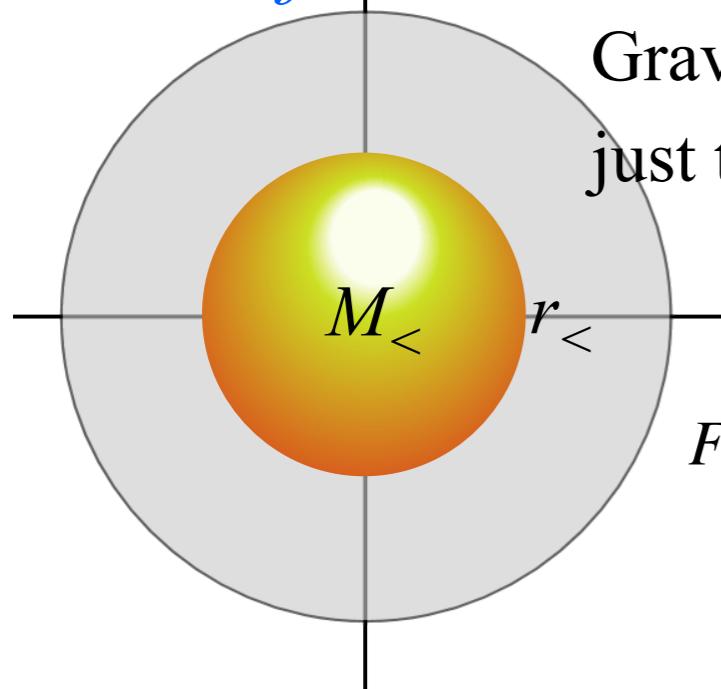


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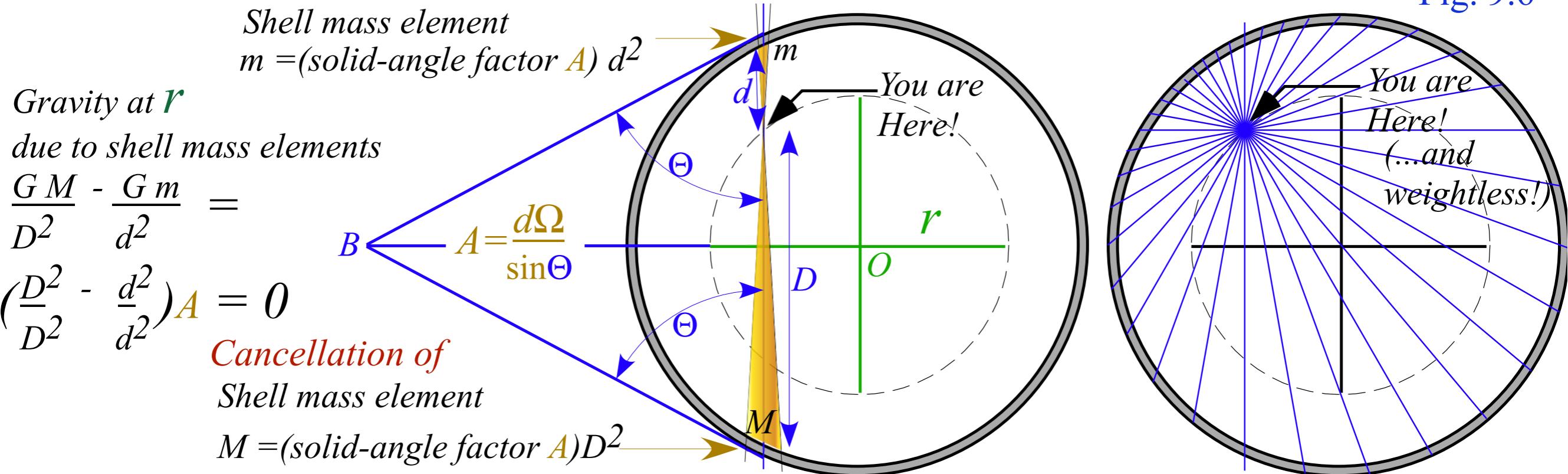
$$F^{inside}(r_<) = G \frac{mM_<}{r_<} = Gm \frac{4\pi}{3} \frac{M_<}{\frac{4\pi}{3} r_<} r_< = Gm \frac{4\pi}{3} \rho_{\oplus} r_<$$

Note:
Hooke's (linear) force law
for $r_<$ inside uniform body

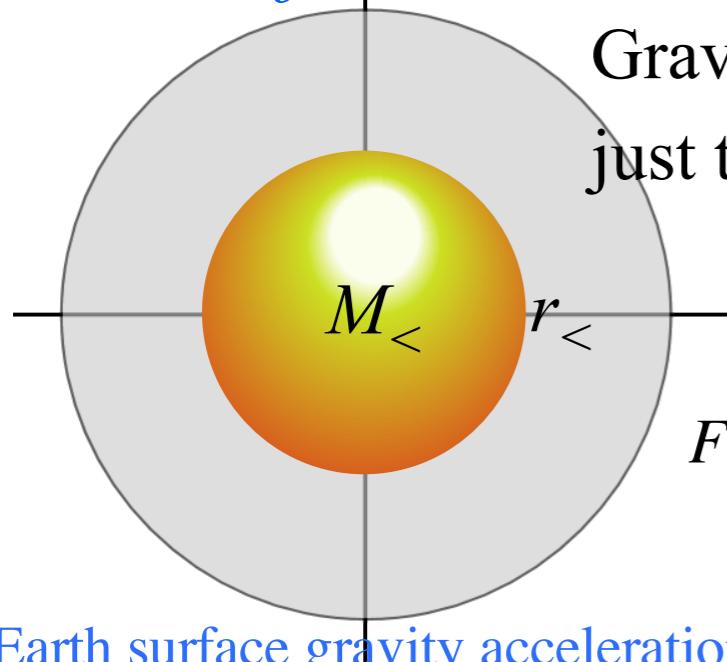
$$\downarrow \quad \downarrow \\ F^{inside}(r_<) = m g \frac{r_<}{R_{\oplus}} \equiv m g \cdot x$$

Coulomb force vanishes inside-spherical shell (Gauss-law)

Unit 1
Fig. 9.6



Coulomb force inside-spherical body due to stuff below you, only.



Gravitational force at $r_<$ is
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Earth surface gravity acceleration: $g = G \frac{M_{\oplus}}{R_{\oplus}^2} = G \frac{M_{\oplus}}{R_{\oplus}^3} R_{\oplus} = G \frac{4\pi}{3} \frac{M_{\oplus}}{4\pi R_{\oplus}^3} R_{\oplus} = G \frac{4\pi}{3} \rho_{\oplus} R_{\oplus} = 9.8 \text{ m/s}^2$

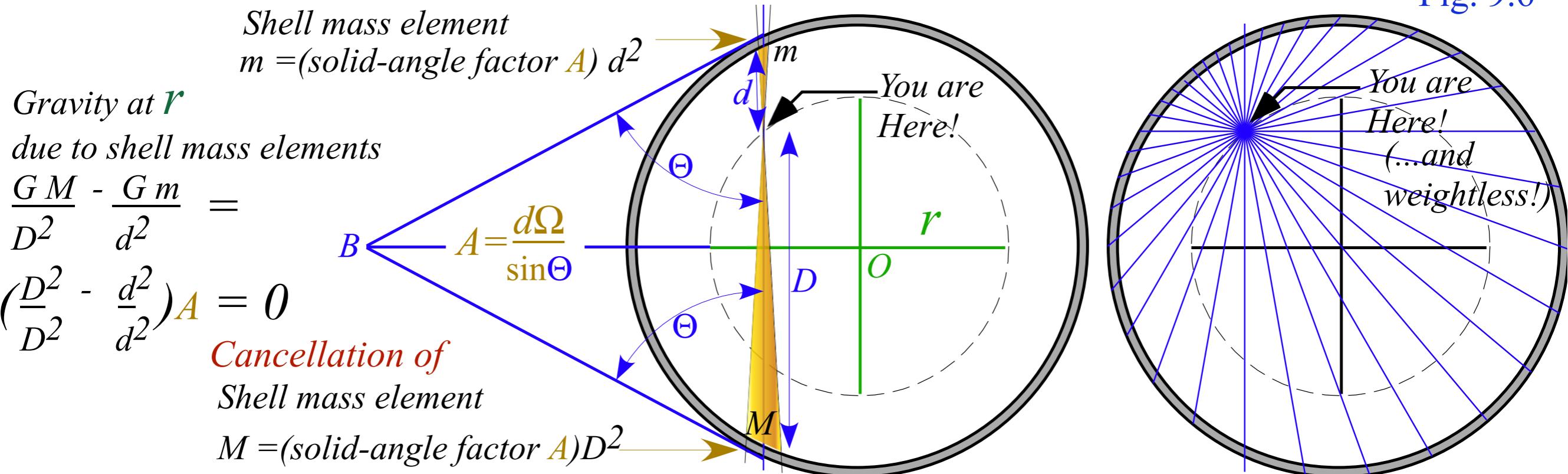
$G = 6.67384(80) \cdot 10^{-11} \text{ Nm}^2/\text{C}^2 \sim (2/3) 10^{-10}$

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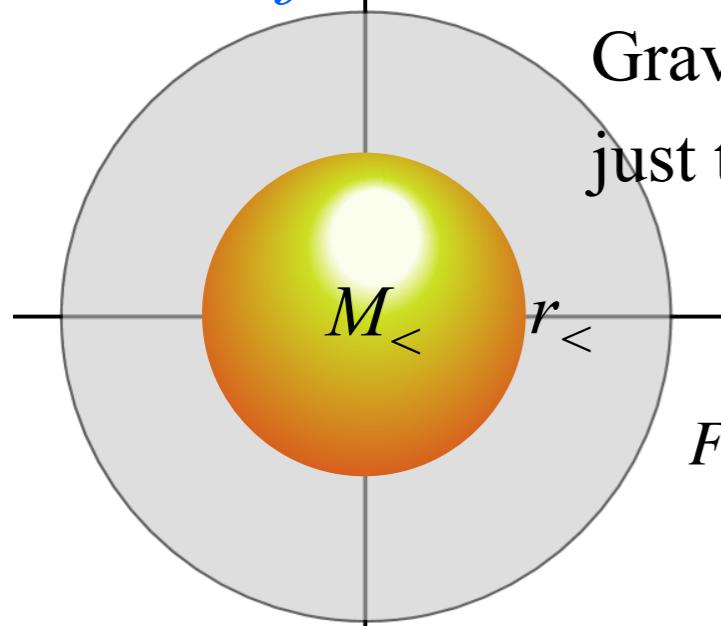


Coulomb force vanishes inside-spherical shell (Gauss-law)

Unit 1
Fig. 9.6



Coulomb force inside-spherical body due to stuff below you, only.



Gravitational force at $r_<$ is just that of planet $M_<$ below $r_<$

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$$\text{Earth surface gravity acceleration: } g = G \frac{M_+}{R_+^2} = G \frac{M_+}{R_+^3} R_+ = G \frac{4\pi}{3} \frac{M_+}{\frac{4\pi}{3}R_+^3} R_+ = G \frac{4\pi}{3} \rho_+ R_+ = 9.8 \text{ m/s}^2$$

$$G = 6.67384(80) \cdot 10^{-11} \text{ Nm}^2/\text{C}^2 \sim (2/3) 10^{-10}$$

$$\text{Earth radius: } R_+ = 6.371 \cdot 10^6 \text{ m} \approx 6.4 \cdot 10^6 \text{ m}$$

$$\text{Earth mass: } M_+ = 5.9722 \times 10^{24} \text{ kg.} \approx 6.0 \cdot 10^{24} \text{ kg.}$$

Note:
Hooke's (linear) force law for $r_<$ inside uniform body

$$\downarrow \quad \downarrow$$

$$F_{\text{inside}}(r_<) = G \frac{mM_<}{r_<} = Gm \frac{4\pi}{3} \frac{M_<}{\frac{4\pi}{3}r_<} r_< = Gm \frac{4\pi}{3} \rho_+ r_< = mg \frac{r_<}{R_+} \equiv mg \cdot x$$

$$\text{Solar radius: } R_\odot = 6.955 \times 10^8 \text{ m.} \approx 7.0 \cdot 10^8 \text{ m.}$$

$$\text{Solar mass: } M_\odot = 1.9889 \times 10^{30} \text{ kg.} \approx 2.0 \cdot 10^{30} \text{ kg.}$$

Geometry of idealized “Sophomore-physics Earth”

Coulomb field outside

Isotropic Harmonic Oscillator (IHO) field inside

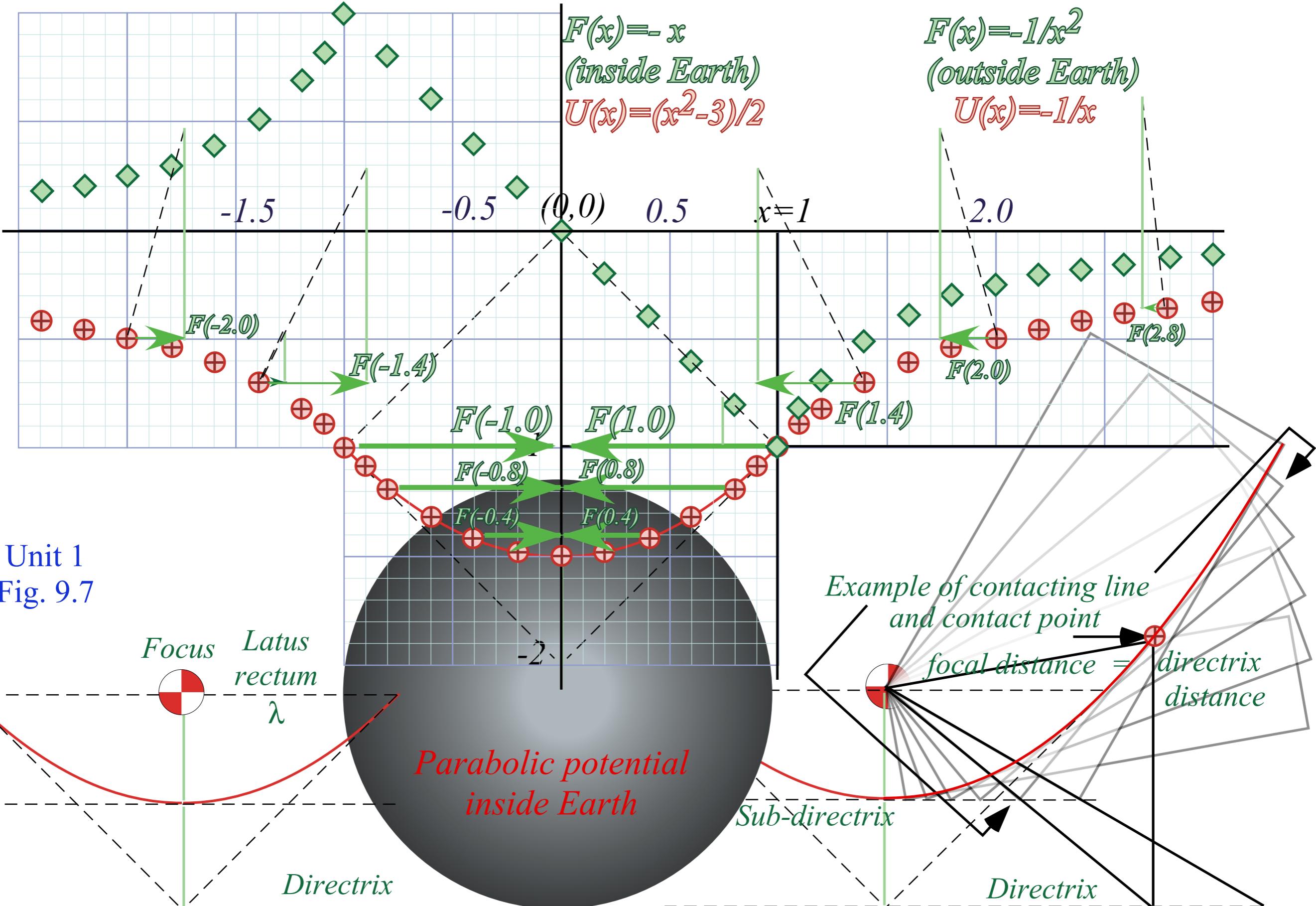
→ *Contact-geometry of potential curve(s)*

“Crushed-Earth” models: 3 key energy “steps” and 4 key energy “levels”

Earth matter vs nuclear matter:

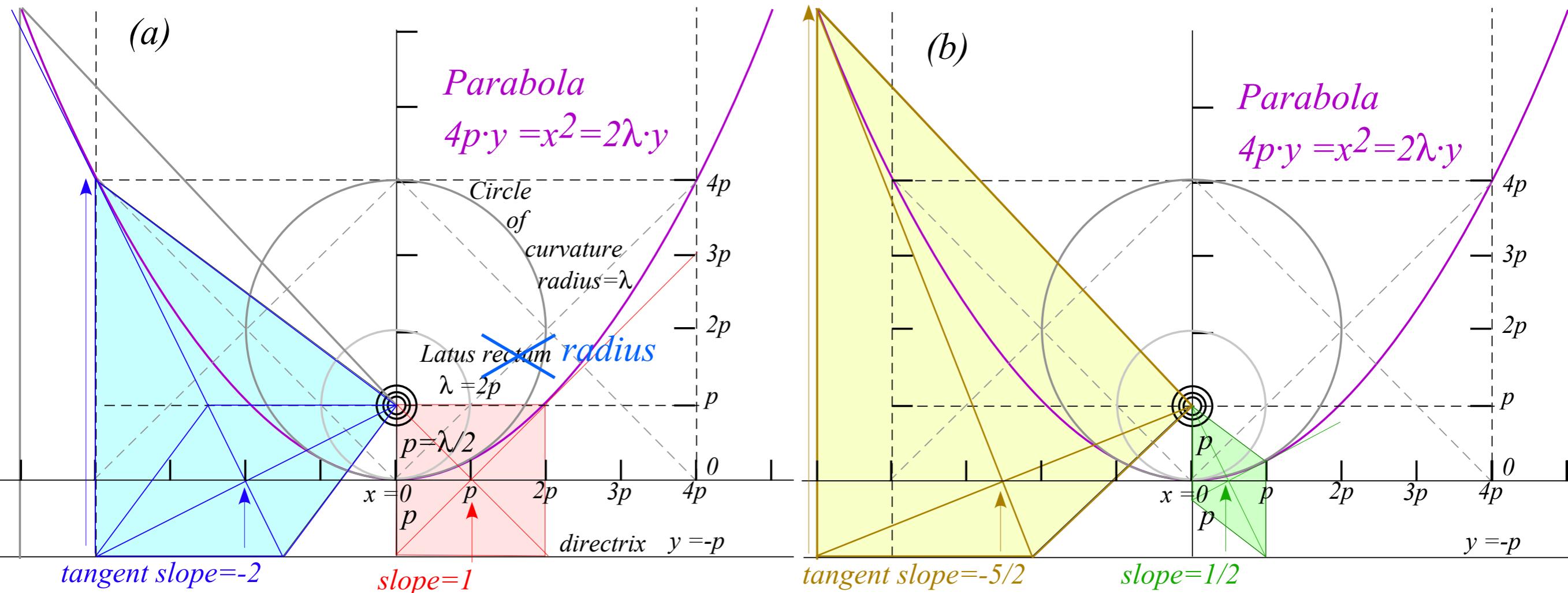
*Introducing the “neutron starlet” and “**Black-Hole-Earth**”*

The ideal “Sophomore-Physics-Earth” model of geo-gravity



...conventional parabolic geometry...carried to extremes...

(From p.18)



Unit 1
Fig. 9.4

Geometry of idealized “Sophomore-physics Earth”

Coulomb field outside

Isotropic Harmonic Oscillator (IHO) field inside

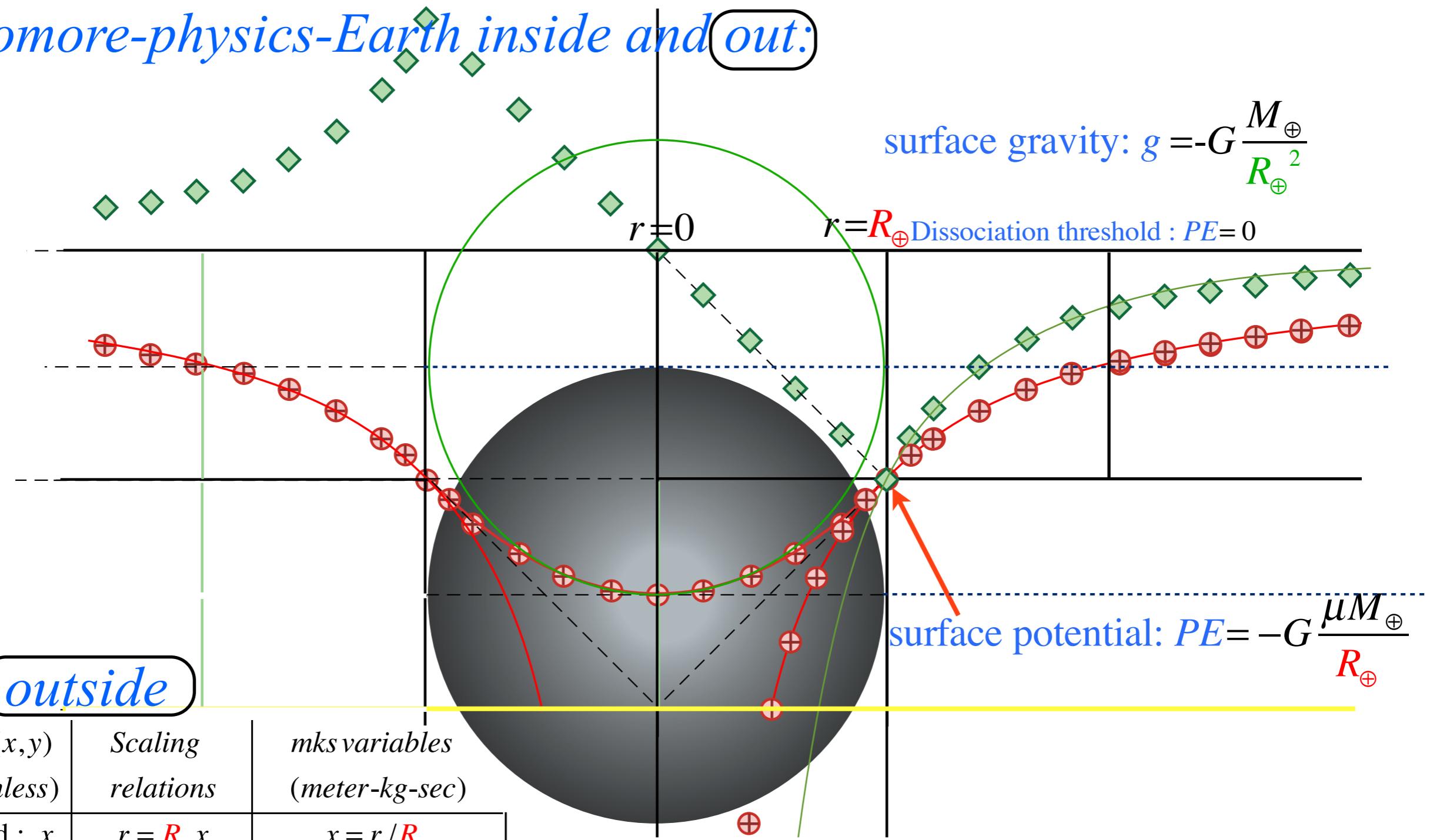
Contact-geometry of potential curve(s)

→ “Crushed-Earth” models: 3 key energy “steps” and 4 key energy “levels”

Earth matter vs nuclear matter:

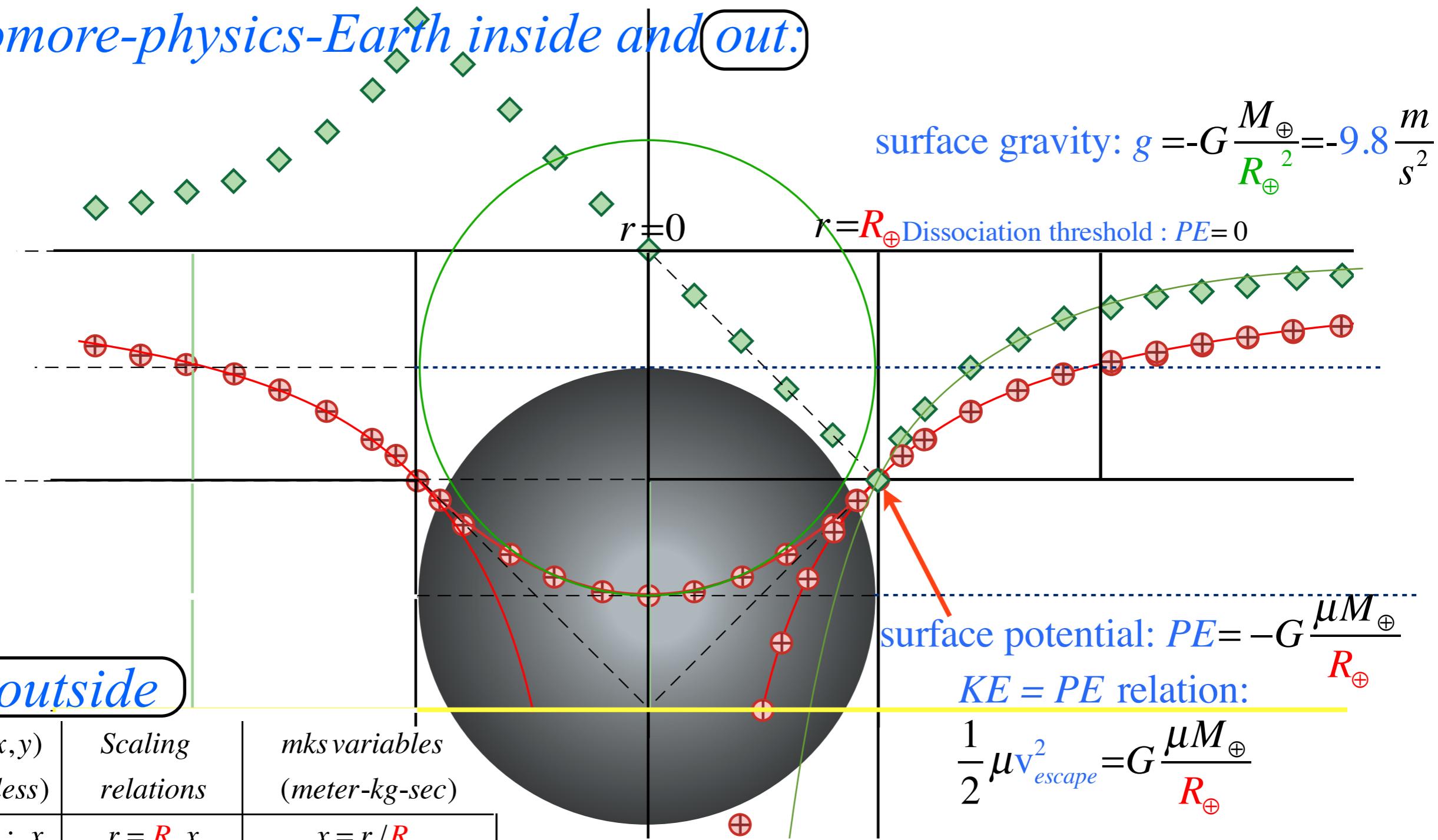
*Introducing the “neutron starlet” and “**Black-Hole-Earth**”*

Sophomore-physics-Earth inside and out:



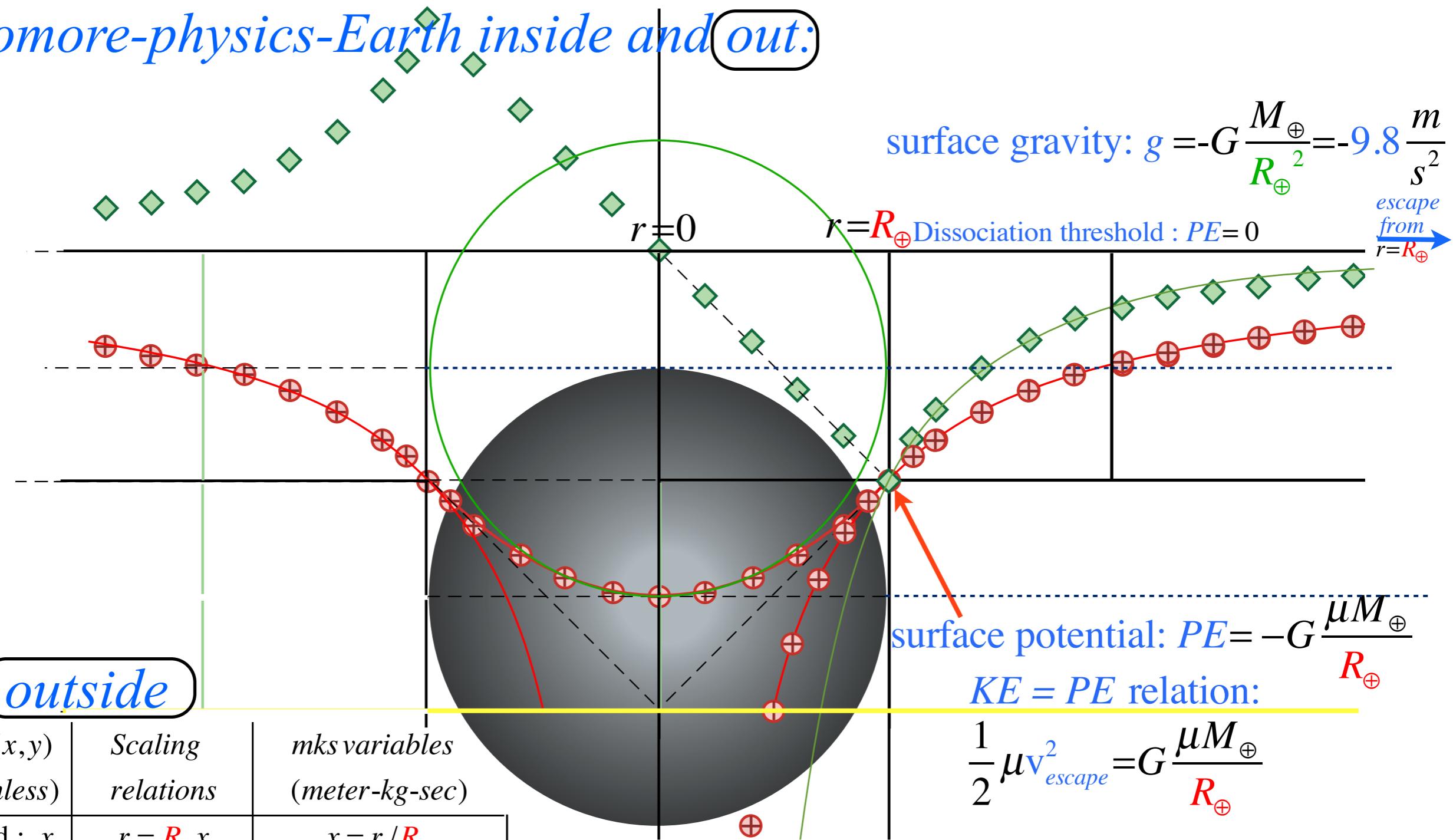
Geometric (x, y) (Dimensionless)	Scaling relations	mks variables (meter-kg-sec)
space coord.: x	$r = R_{\oplus}x$	$x = r / R_{\oplus}$
PE for $ x \geq 1$: $y^{PE} = \frac{-1}{x}$	$PE^{mks}(r)$ $= \frac{GM\mu}{R_{\oplus}} y^{PE}$	$PE^{mks}(r) = -\frac{GM\mu}{r}$ $= -\frac{GM\mu}{R_{\oplus}} \frac{1}{x}$

Sophomore-physics-Earth inside and out:



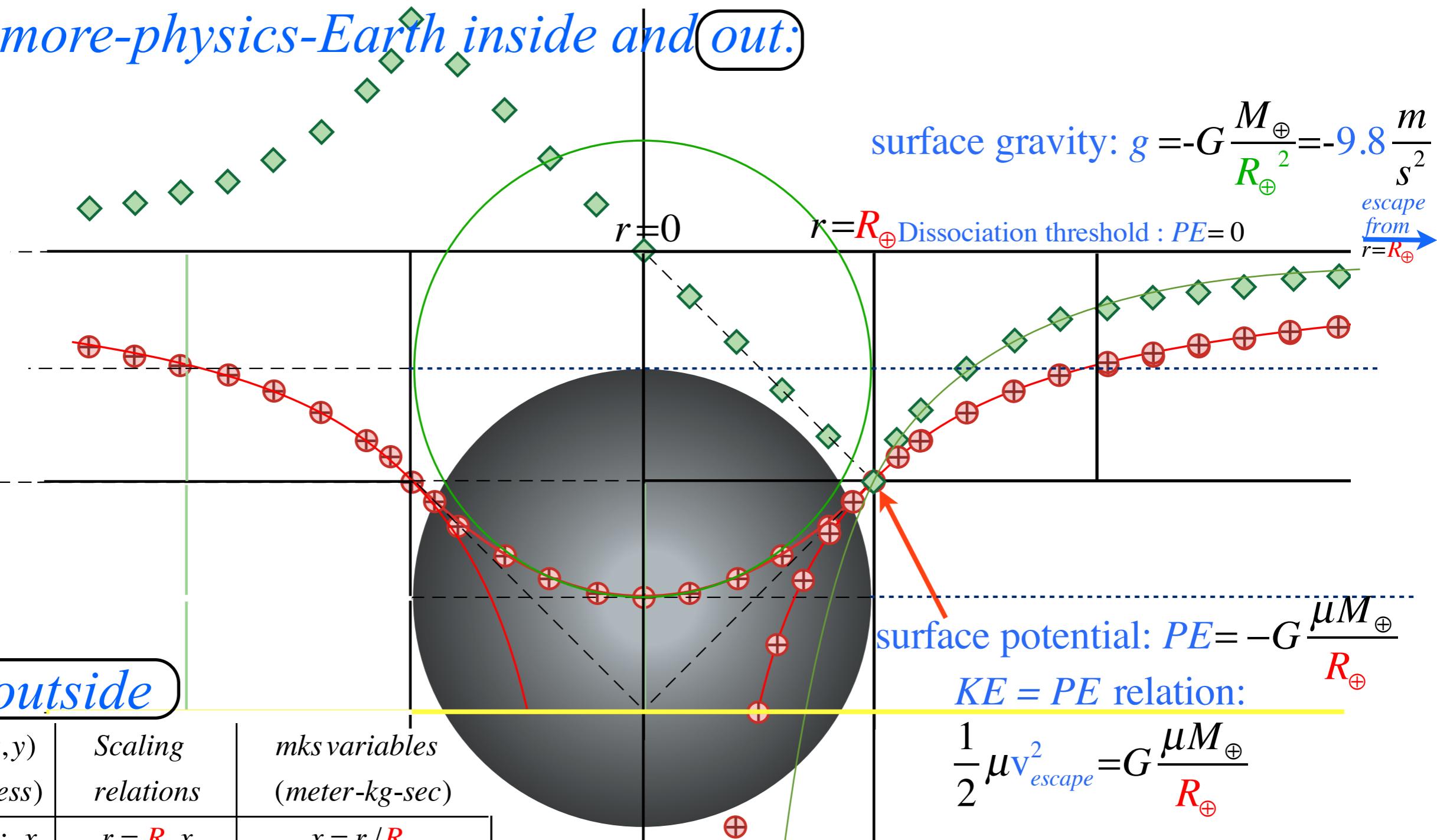
Geometric(x, y) (Dimensionless)	Scaling relations	mks variables (meter-kg-sec)
space coord.: x	$r = R_{\oplus}x$	$x = r / R_{\oplus}$
PE for $ x \geq 1$:	$PE^{\text{mks}}(r)$ $y^{PE} = \frac{-1}{x}$	$PE^{\text{mks}}(r) = -\frac{GM\mu}{r}$ $= -\frac{GM\mu}{R_{\oplus}} \frac{1}{x}$

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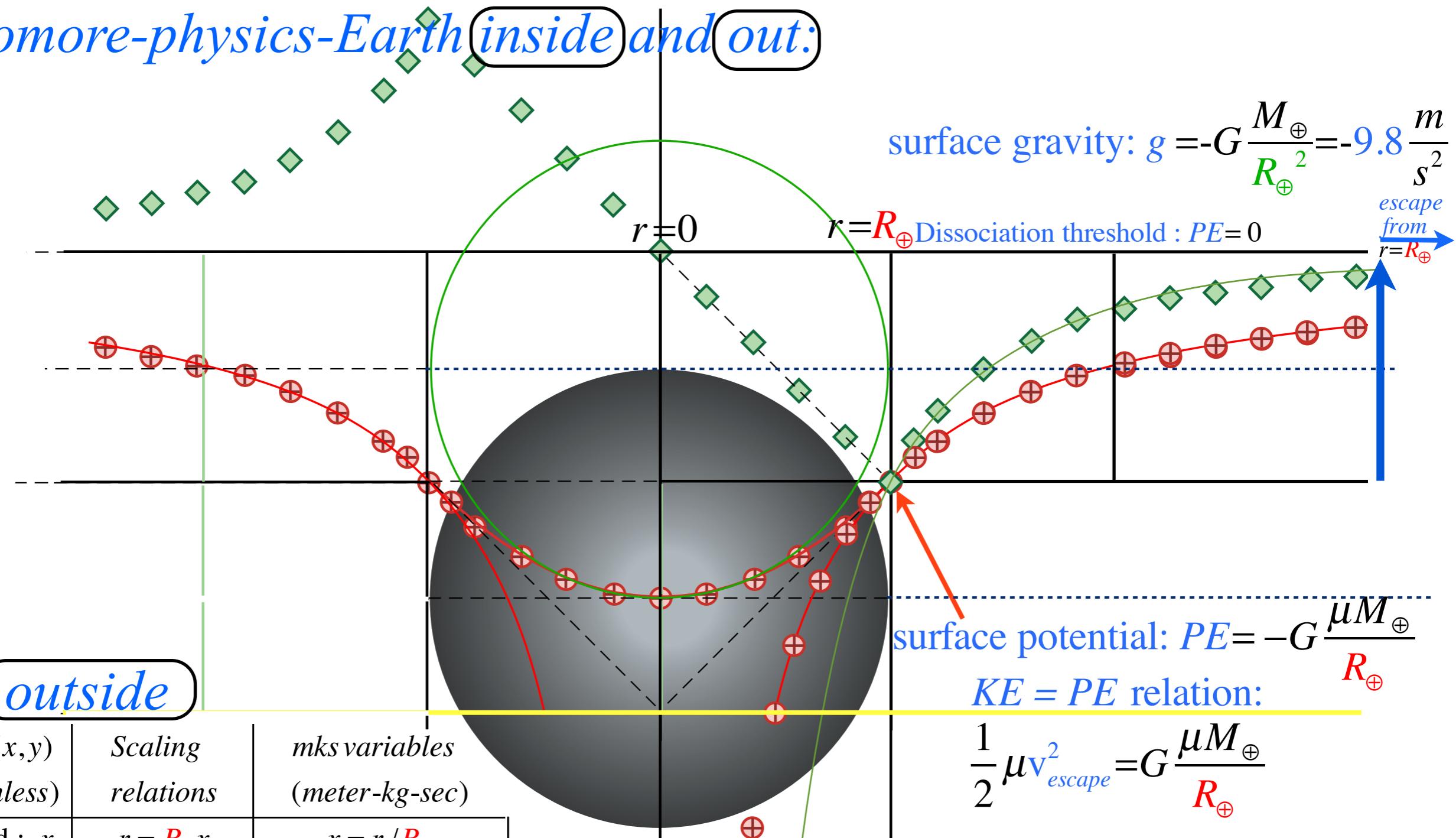


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R_{\oplus} -escape-velocity

$$v_{\text{escape}} = \sqrt{2G \frac{M_{\oplus}}{R_{\oplus}}}$$

Sophomore-physics-Earth *inside and out:*



Geometric (x, y) (Dimensionless)	Scaling relations	mks variables (meter-kg-sec)
space coord.: x	$r = R_{\oplus} x$	$x = r / R_{\oplus}$
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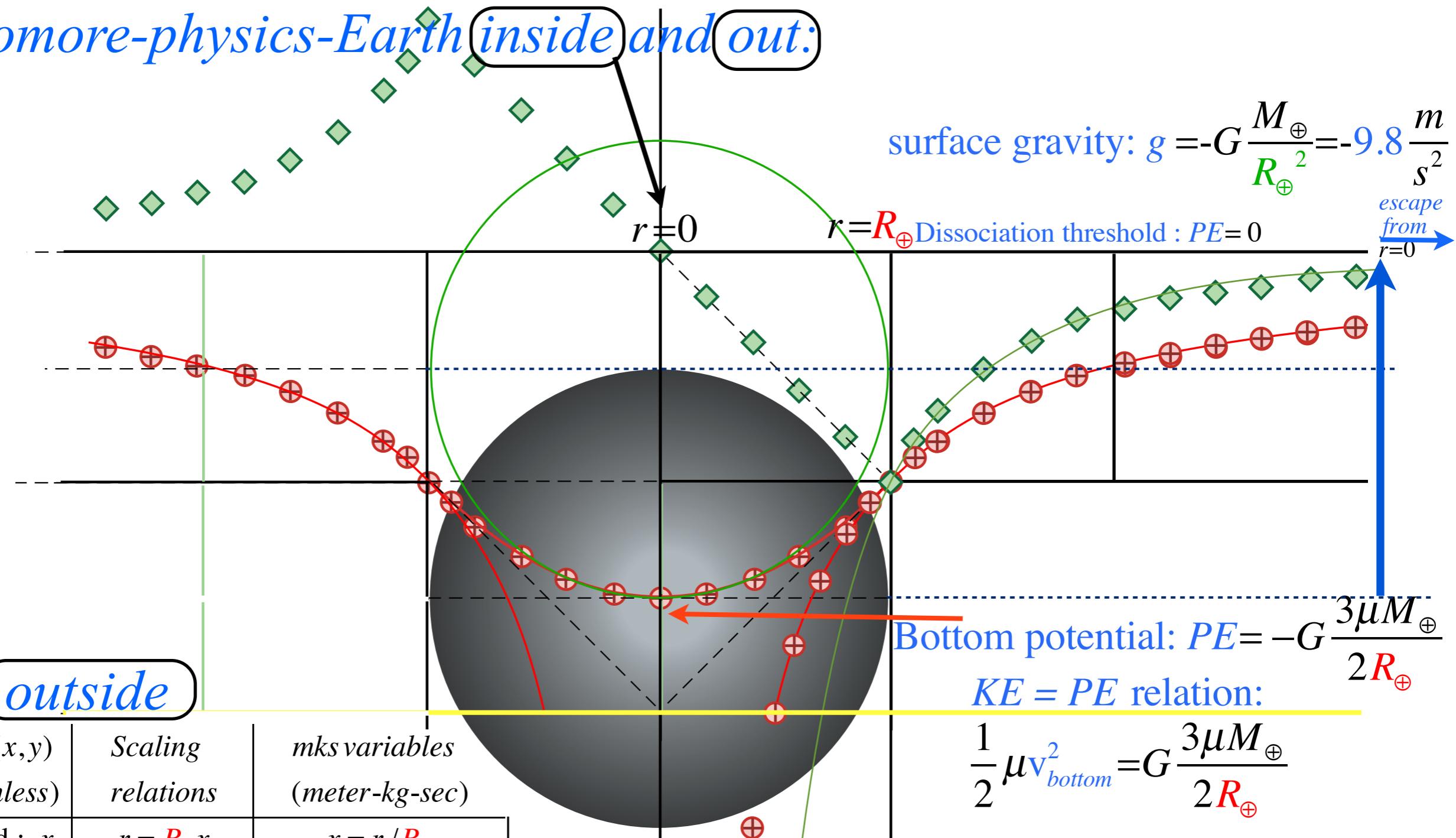
PE for $ x < 1$:	$y^{PE} = \frac{x^2}{2} - \frac{3}{2}$	$inside$
$Force$ for $ x < 1$:	$y^{Force} = -x$	$F^{mks}(r) = -\frac{GM\mu}{R_{\oplus}^3} r$

R_{\oplus} -escape-velocity

$$v_{\text{escape}} = \sqrt{2G \frac{M_{\oplus}}{R_{\oplus}}}$$

11.1 km/sec

Sophomore-physics-Earth *inside and out:*



Geometric (x, y) (Dimensionless)	Scaling relations	mks variables (meter-kg-sec)
space coord.: x	$r = R_\oplus x$	$x = r / R_\oplus$
PE for $ x \geq 1$: $y^{PE} = -\frac{1}{x}$	$PE^{mks}(r)$ $= \frac{GM\mu}{R_\oplus} y^{PE}$	$PE^{mks}(r) = -\frac{GM\mu}{r}$ $= -\frac{GM\mu}{R_\oplus} \frac{1}{x}$
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<i>inside</i>
PE for $ x < 1$: $y^{PE} = \frac{x^2}{2} - \frac{3}{2}$

$(r=0)$ -escape-velocity

$$v_{bottom} = \sqrt{3G \frac{M_\oplus}{R_\oplus}}$$

$$13.7 \text{ km/sec}$$

surface gravity: $g = -G \frac{M_\oplus}{R_\oplus^2} = 9.8 \frac{\text{m}}{\text{s}^2}$

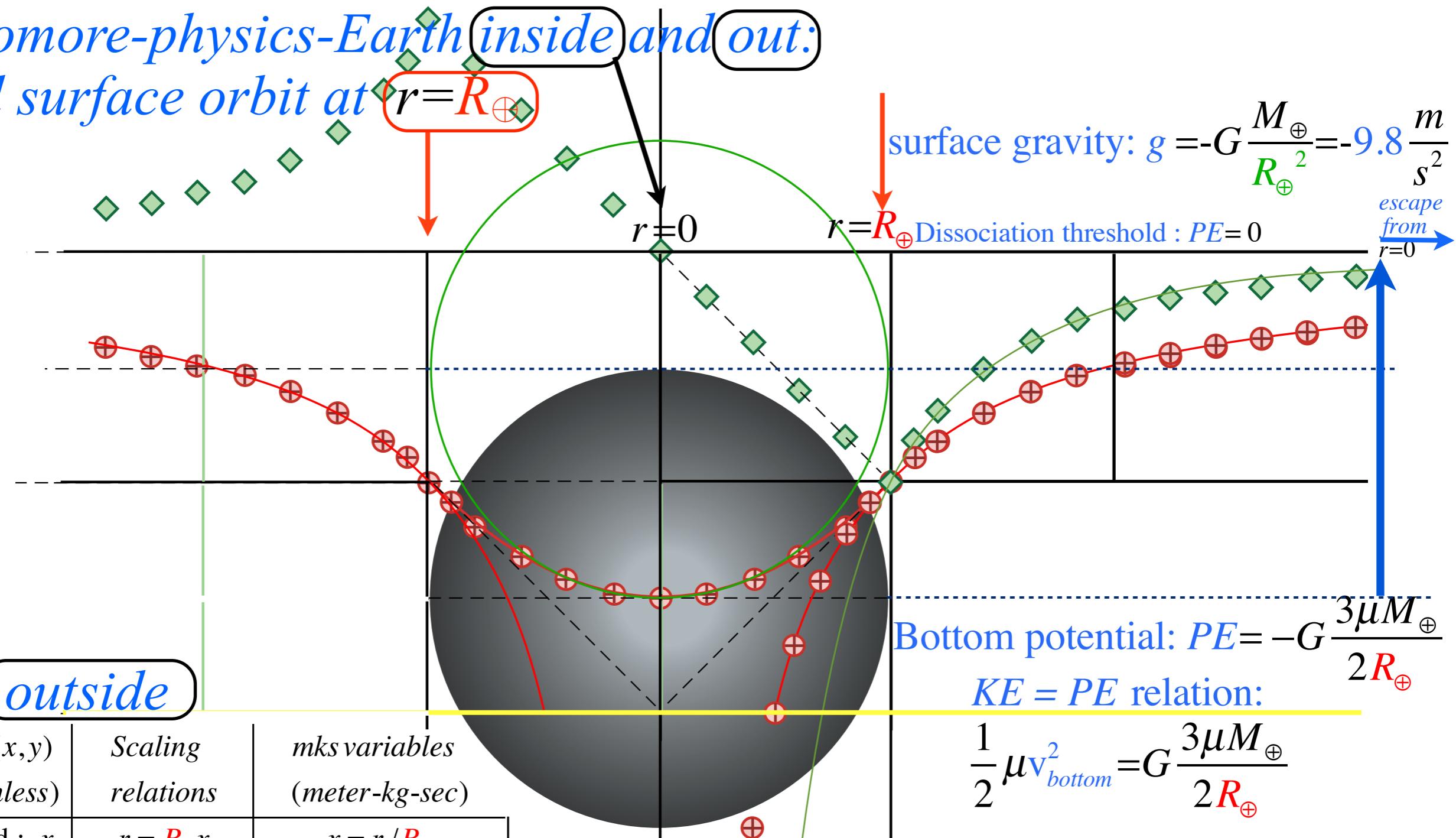
Dissociation threshold : $PE = 0$

Bottom potential: $PE = -G \frac{3\mu M_\oplus}{2R_\oplus}$

$KE = PE$ relation:

$$\frac{1}{2} \mu v_{bottom}^2 = G \frac{3\mu M_\oplus}{2R_\oplus}$$

Sophomore-physics-Earth inside and out: ...and surface orbit at $r=R_{\oplus}$



Geometric (x, y) (Dimensionless)	Scaling relations	mks variables (meter-kg-sec)
space coord.: x	$r = R_{\oplus}x$	$x = r / R_{\oplus}$
PE for $ x \geq 1$: $y^{PE} = \frac{-1}{x}$	$PE^{mks}(r) = -\frac{GM\mu}{r}$ $= \frac{GM\mu}{R_{\oplus}} y^{PE}$	$PE^{mks}(r) = -\frac{GM\mu}{r}$ $= -\frac{GM\mu}{R_{\oplus}} \frac{1}{x}$
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PE for $ x < 1$: $y^{PE} = \frac{x^2}{2} - \frac{3}{2}$	$Force$ for $ x < 1$: $y^{Force} = -x$	$inside$
$PE^{mks}(r) = \frac{GM\mu}{R_{\oplus}} \left(\frac{r^2}{2R_{\oplus}^2} - \frac{3}{2} \right)$	$F^{mks}(r) = -\frac{GM\mu}{R_{\oplus}^3} r$	

Bottom potential: $PE = -G \frac{3\mu M_{\oplus}}{2R_{\oplus}}$

$KE = PE$ relation:

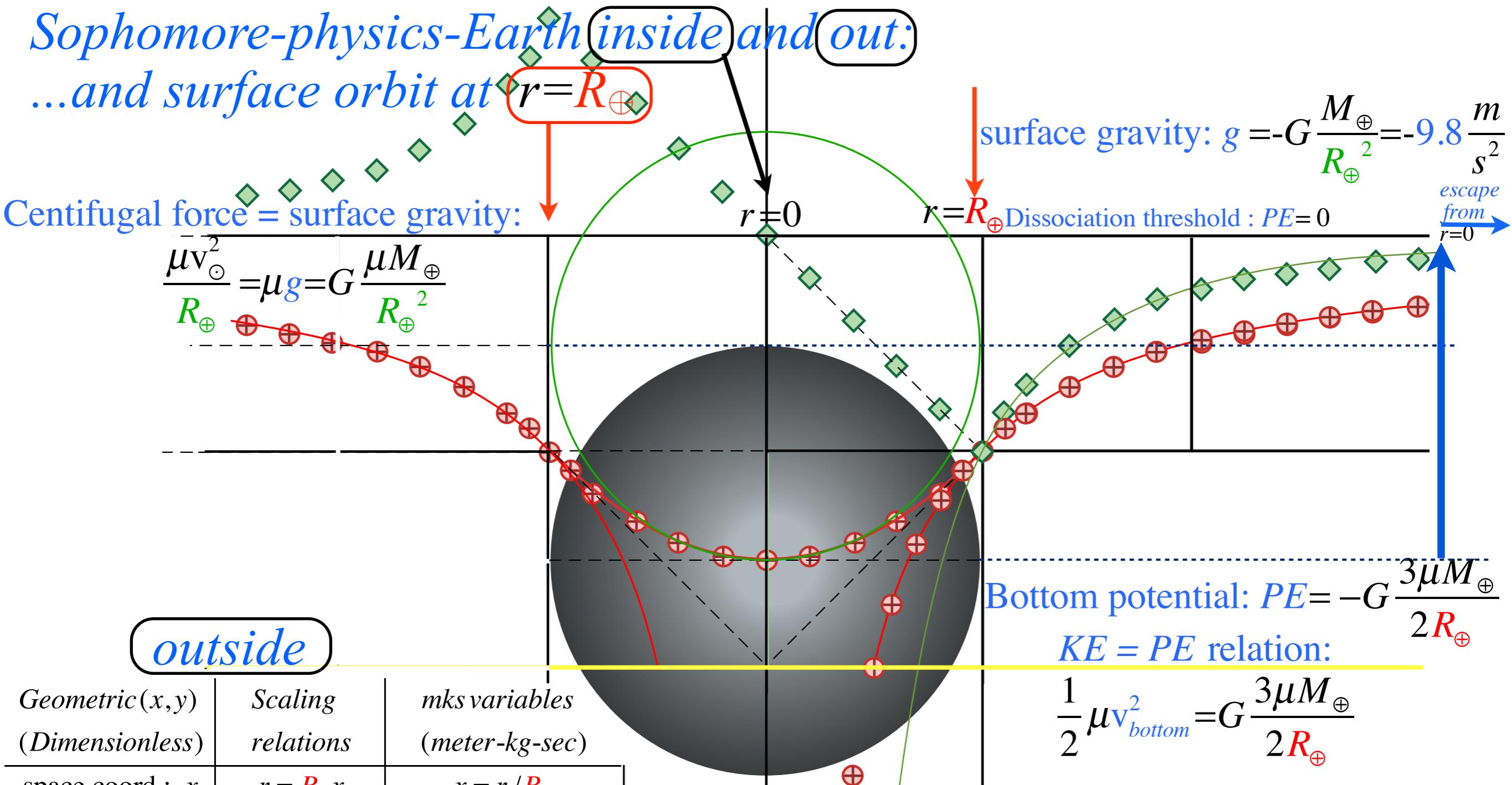
$$\frac{1}{2} \mu v_{bottom}^2 = G \frac{3\mu M_{\oplus}}{2R_{\oplus}}$$

$(r=0)$ -escape-velocity

$$v_{bottom} = \sqrt{3G \frac{M_{\oplus}}{R_{\oplus}}} = 13.7 \text{ km/sec}$$

Sophomore-physics-Earth inside and out:

...and surface orbit at



outside

Geometric(x, y) (Dimensionless)	Scaling relations	mks variables (meter-kg-sec)
space coord.: x	$r = R_\oplus x$	$x = r / R_\oplus$
PE for $ x \geq 1$:	$PE^{mks}(r) = -\frac{GM\mu}{r}$	$y^{PE} = \frac{GM\mu}{R_\oplus} x$
$Force$ for $ x \geq 1$:	$F^{mks}(r) = -\frac{GM\mu}{r^2}$	$y^{Force} = -\frac{GM\mu}{R_\oplus^2} x$

inside

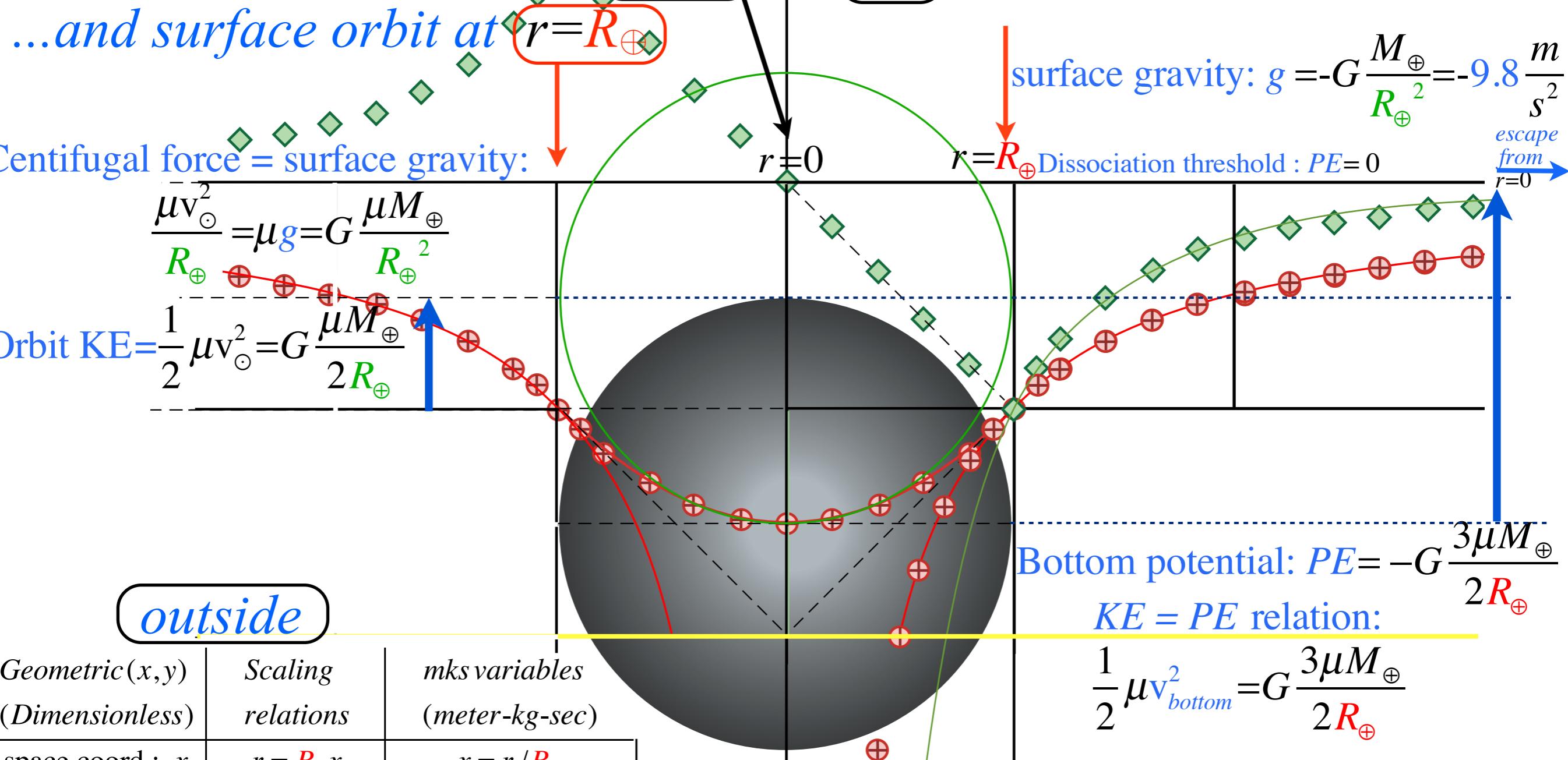
$$PE^{mks}(r) = \frac{GM\mu}{R_\oplus} \left(\frac{r^2}{2R_\oplus^2} - \frac{3}{2} \right)$$

$$F^{mks}(r) = -\frac{GM\mu}{R_\oplus^3} r$$

$(r=0)$ -escape-velocity

$$v_{bottom} = \sqrt{3G \frac{M_\oplus}{R_\oplus}}$$

Sophomore-physics-Earth inside and out:



Geometric (x, y) (Dimensionless)	Scaling relations	mks variables (meter-kg-sec)
space coord.: x	$r = R_\oplus x$	$x = r / R_\oplus$
PE for $ x \geq 1$:	$PE^{mks}(r) = -\frac{GM\mu}{r}$	$PE^{mks}(r) = -\frac{GM\mu}{r}$
	$y^{PE} = \frac{-1}{x}$	$y^{PE} = \frac{-1}{x}$
$Force$ for $ x \geq 1$:	$F^{mks}(r) = -\frac{GM\mu}{r^2}$	$F^{mks}(r) = -\frac{GM\mu}{r^2}$
	$y^{Force} = \frac{-1}{x^2}$	$y^{Force} = \frac{-1}{x^2}$

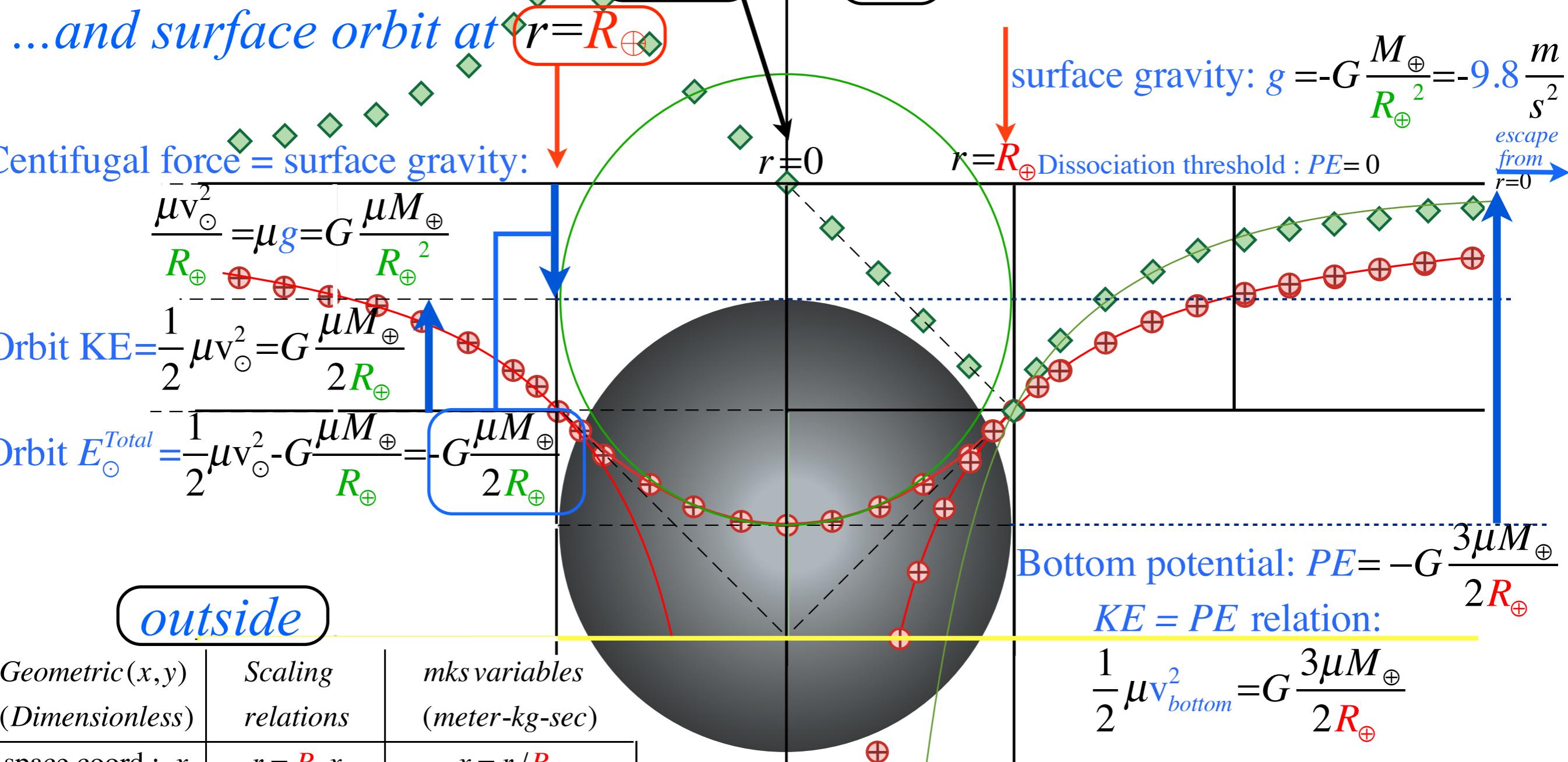
inside
PE for $ x < 1$:

$(r=0)$ -escape-velocity

$$v_{bottom} = \sqrt{3G \frac{M_\oplus}{R_\oplus}}$$

$$13.7 \text{ km/sec}$$

Sophomore-physics-Earth inside and out:



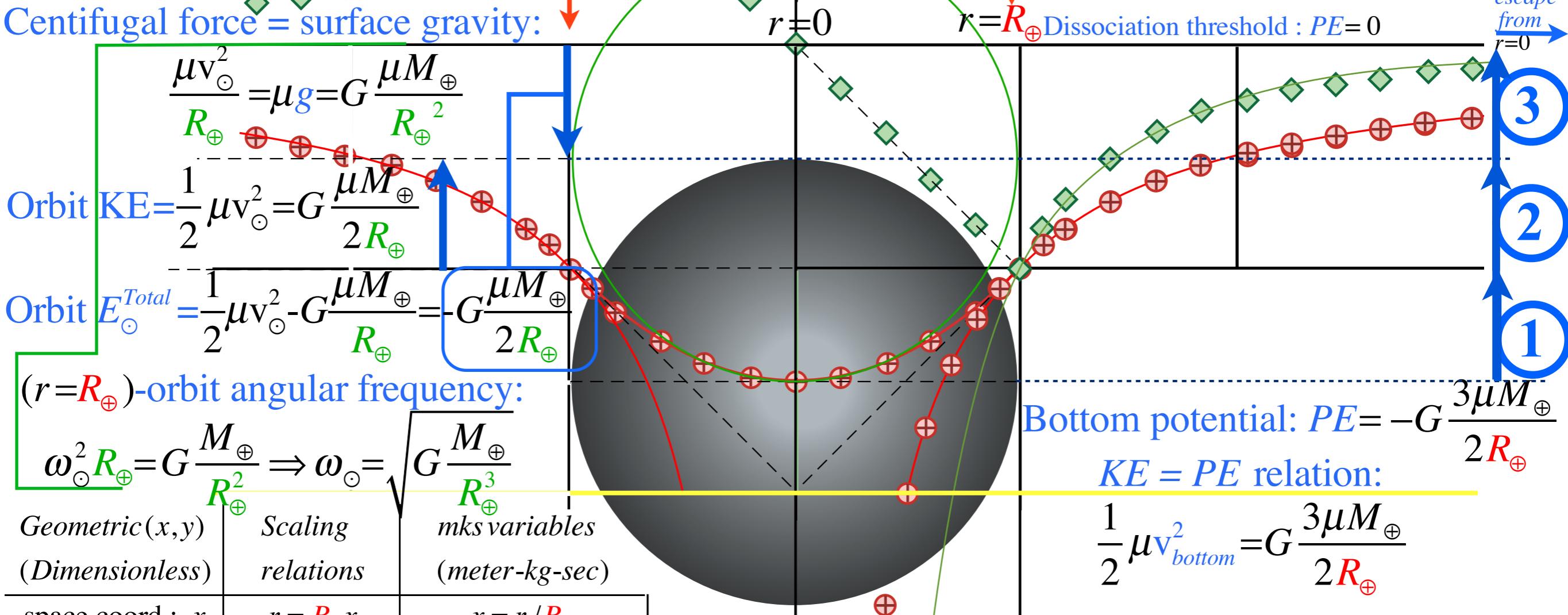
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			$F^{mks}(r) = -\frac{GM\mu}{R_{\oplus}^3} r$

$$v_{bottom} = \sqrt{3G \frac{M_{\oplus}}{R_{\oplus}}} = 13.7 \text{ km/sec}$$

Sophomore-physics-Earth inside and out: “3-steps out of (or into) Hell”

...and surface orbit at $r=R_\oplus$



space coord.: x	$r = R_\oplus x$	$x = r / R_\oplus$
PE for $ x \geq 1$:	$PE^{\text{mks}}(r) = -\frac{GM\mu}{r}$	$y^{PE} = \frac{GM\mu}{R_\oplus} x$
$Force$ for $ x \geq 1$:	$F^{\text{mks}}(r) = -\frac{GM\mu}{r^2}$	$y^{Force} = -\frac{GM\mu}{R_\oplus^2} \frac{1}{x^2}$

$inside$
PE for $ x < 1$: $y^{PE} = \frac{x^2}{2} - \frac{3}{2}$

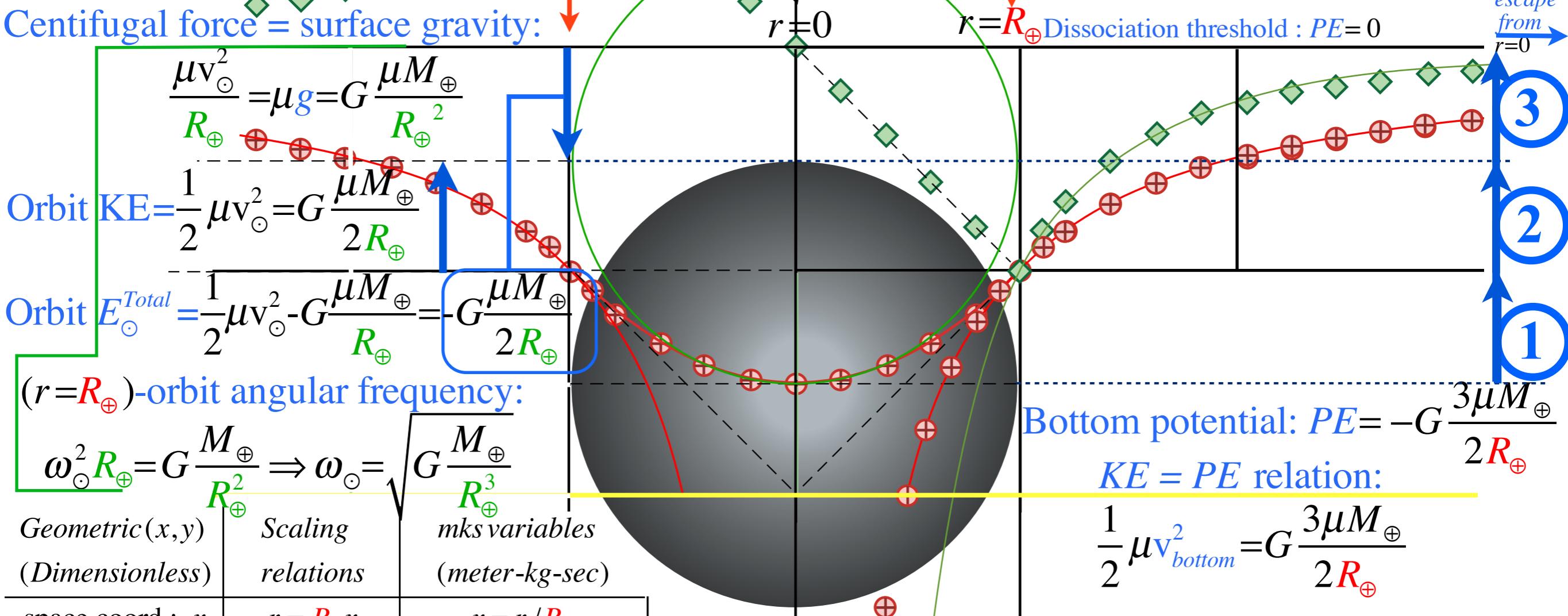
$(r=0)$ -escape-velocity

$$v_{\text{bottom}} = \sqrt{3G \frac{M_\oplus}{R_\oplus}}$$

13.7 km/sec

Sophomore-physics-Earth inside and out: “3-steps out of (or into) Hell”

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space coord.: x	$r = R_\oplus x$	$x = r / R_\oplus$
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$Force$ for $ x < 1$:	$F^{mks}(r) = -\frac{GM\mu}{R_\oplus^3} r$	

($r=0$)-escape-velocity

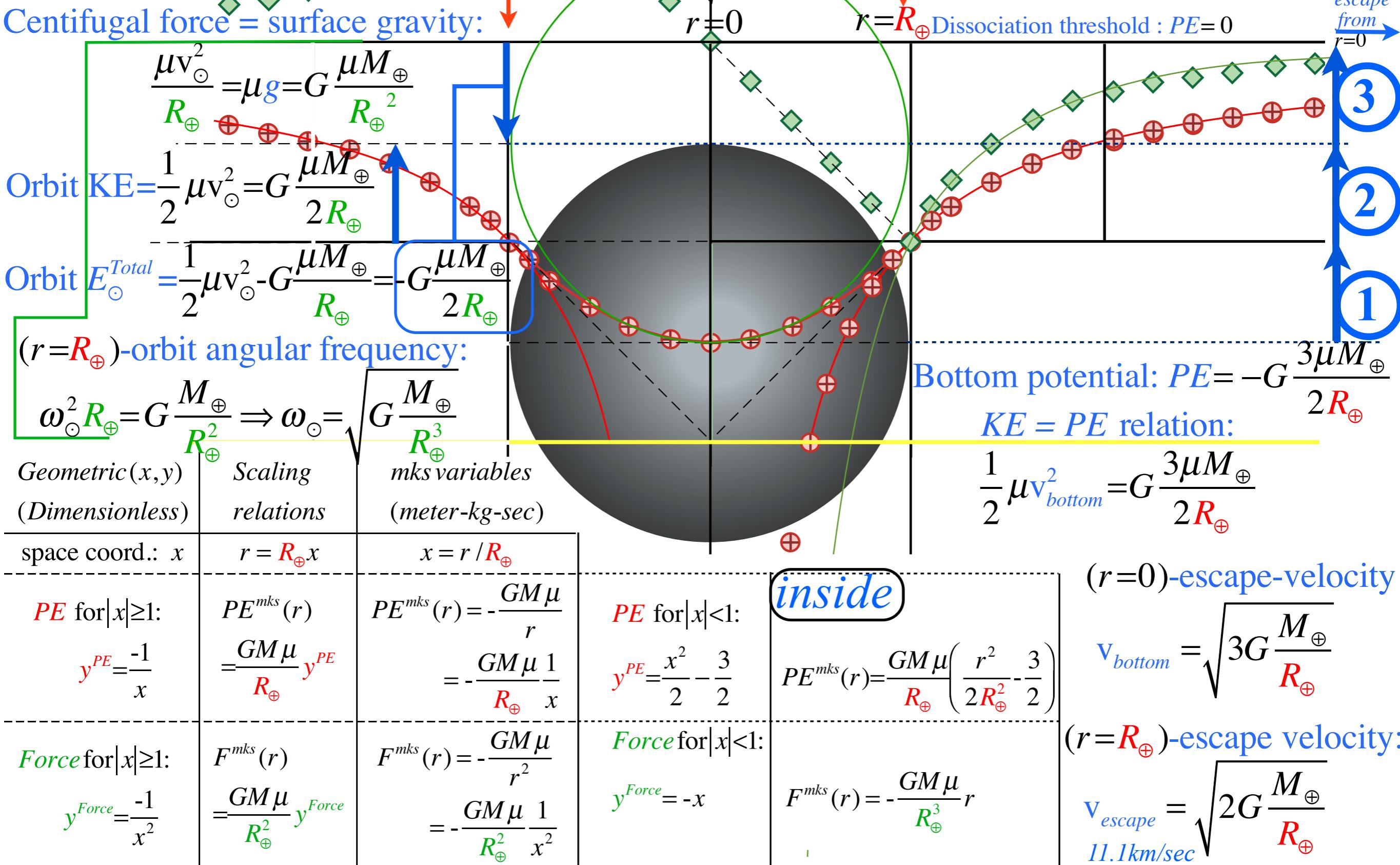
$$v_{bottom} = \sqrt{3G \frac{M_\oplus}{R_\oplus}} = 13.7 \text{ km/sec}$$

($r=R_\oplus$)-orbit speed:

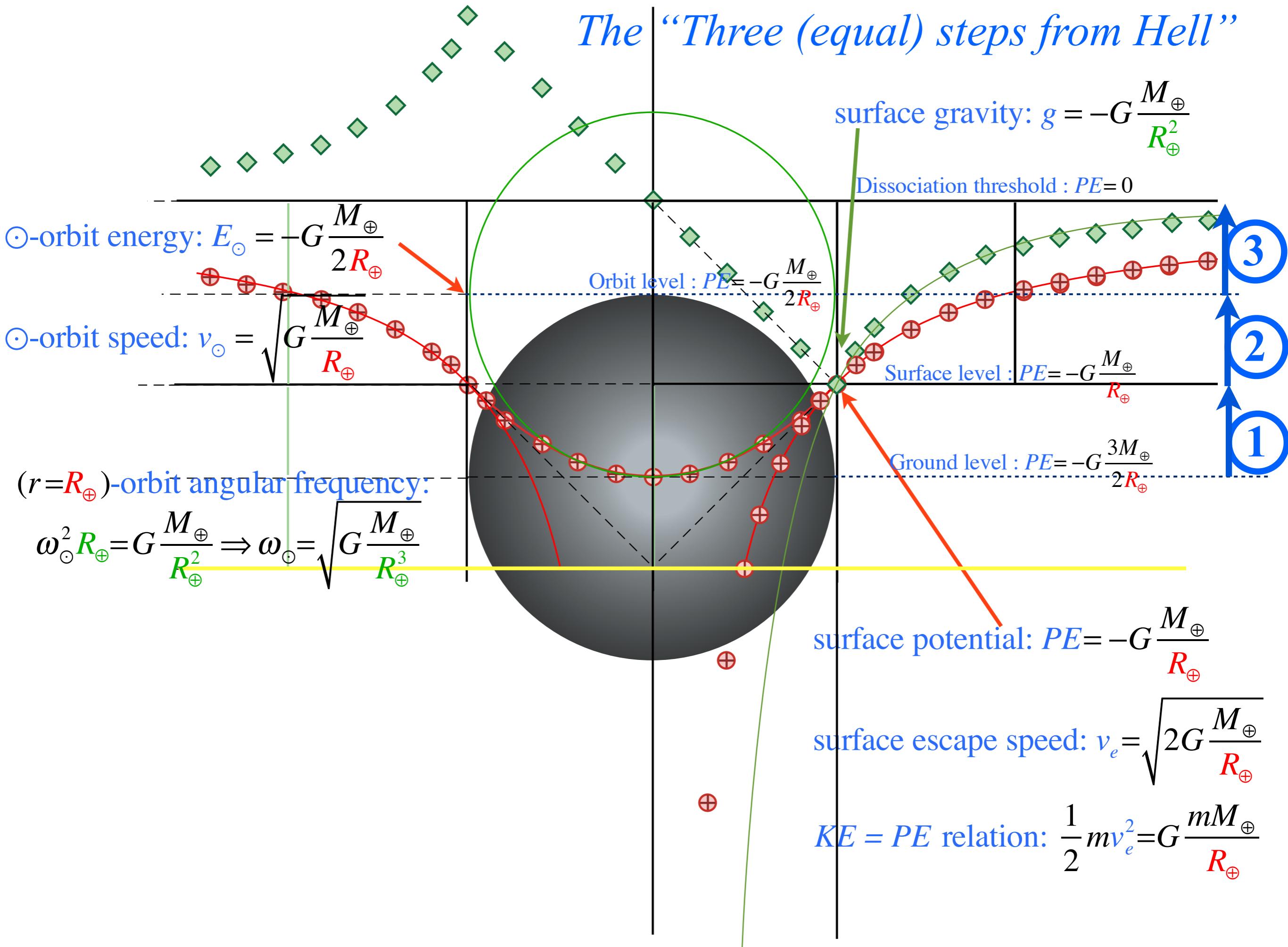
$$v_\oplus = \sqrt{G \frac{M_\oplus}{R_\oplus}} = \sqrt{g R_\oplus} = 7.9 \text{ km/sec}$$

Sophomore-physics-Earth inside and out: “3-steps out of (or into) Hell”

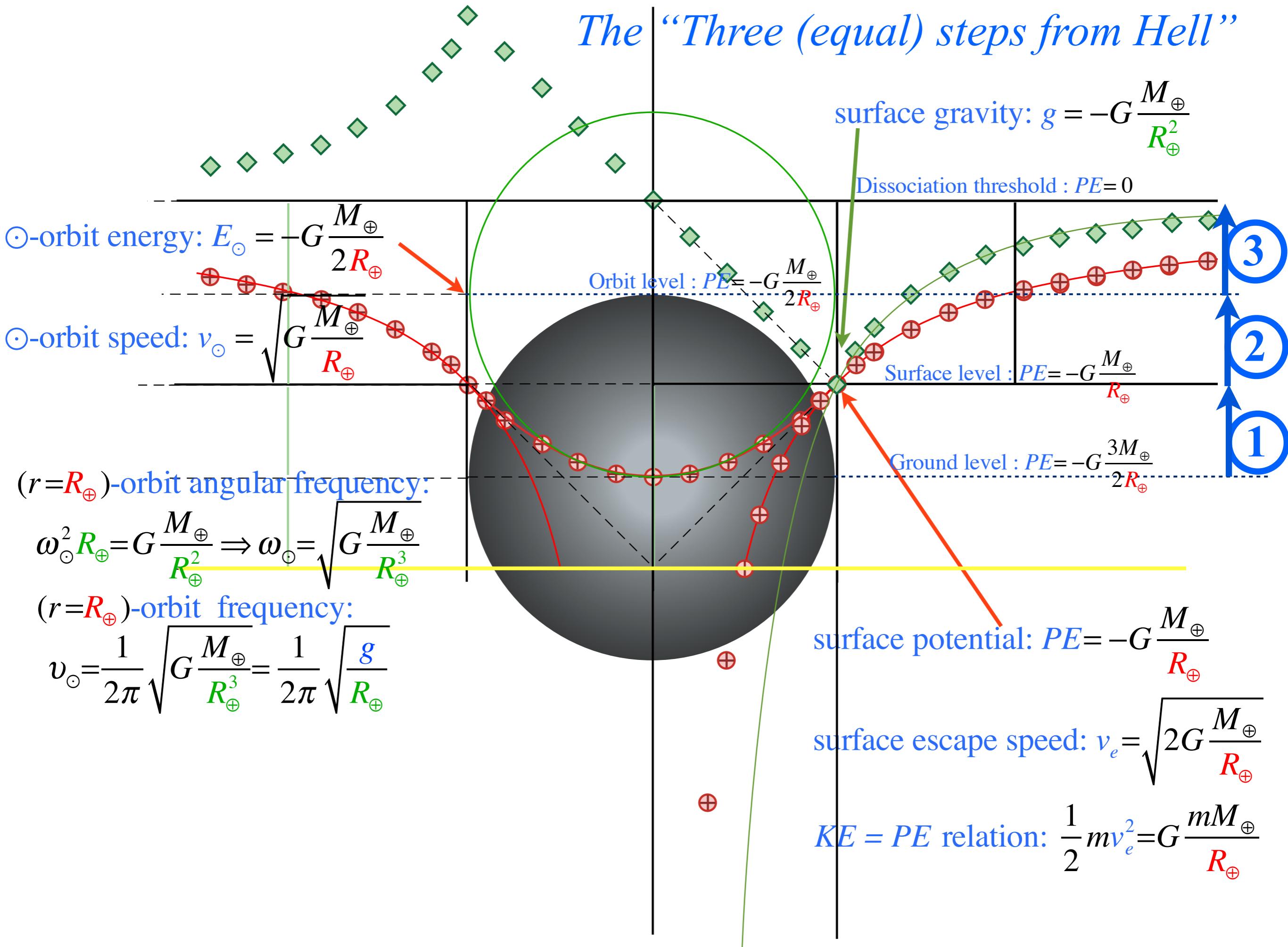
...and surface orbit at $r=R_\oplus$



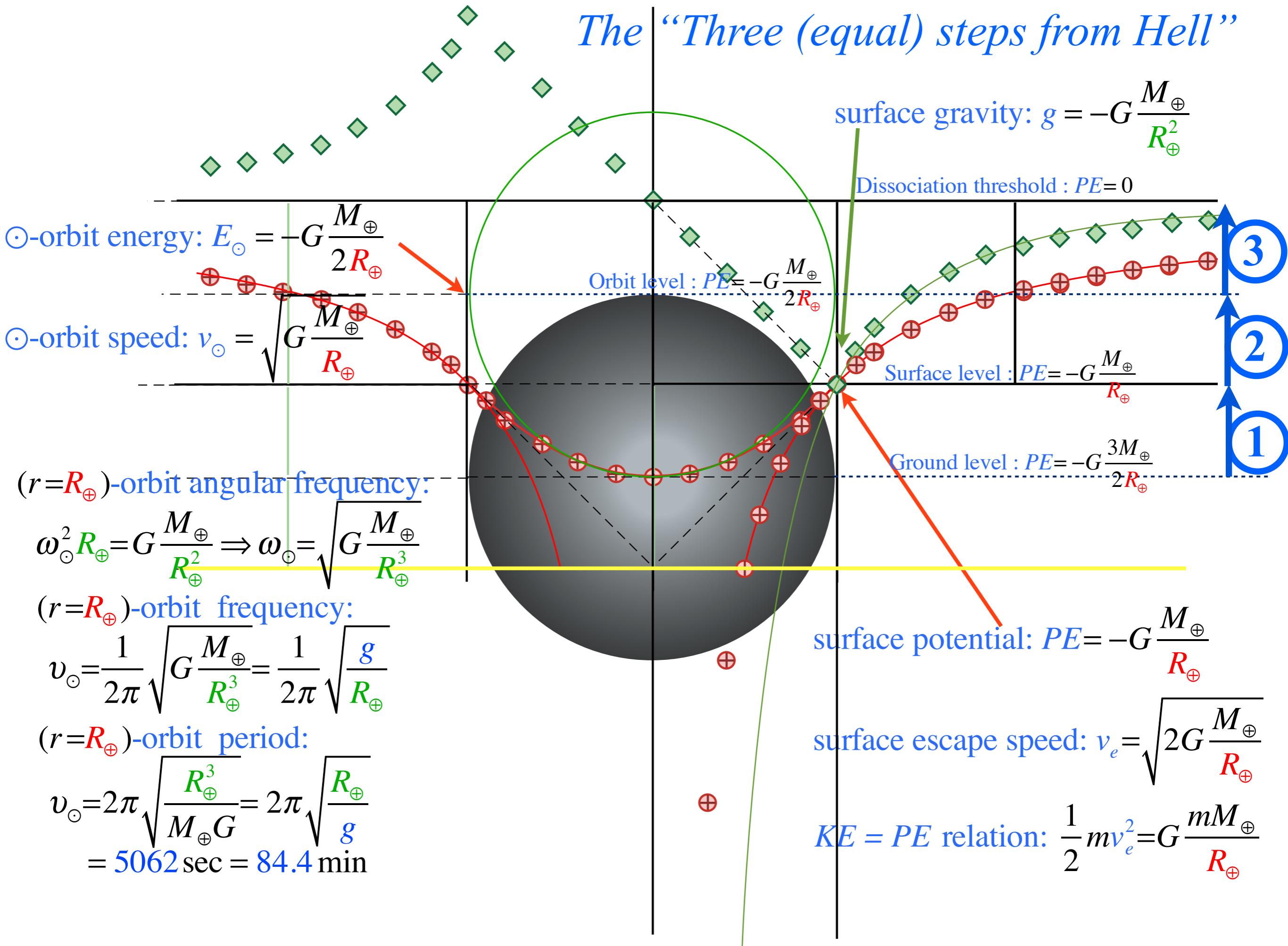
The “Three (equal) steps from Hell”



The “Three (equal) steps from Hell”

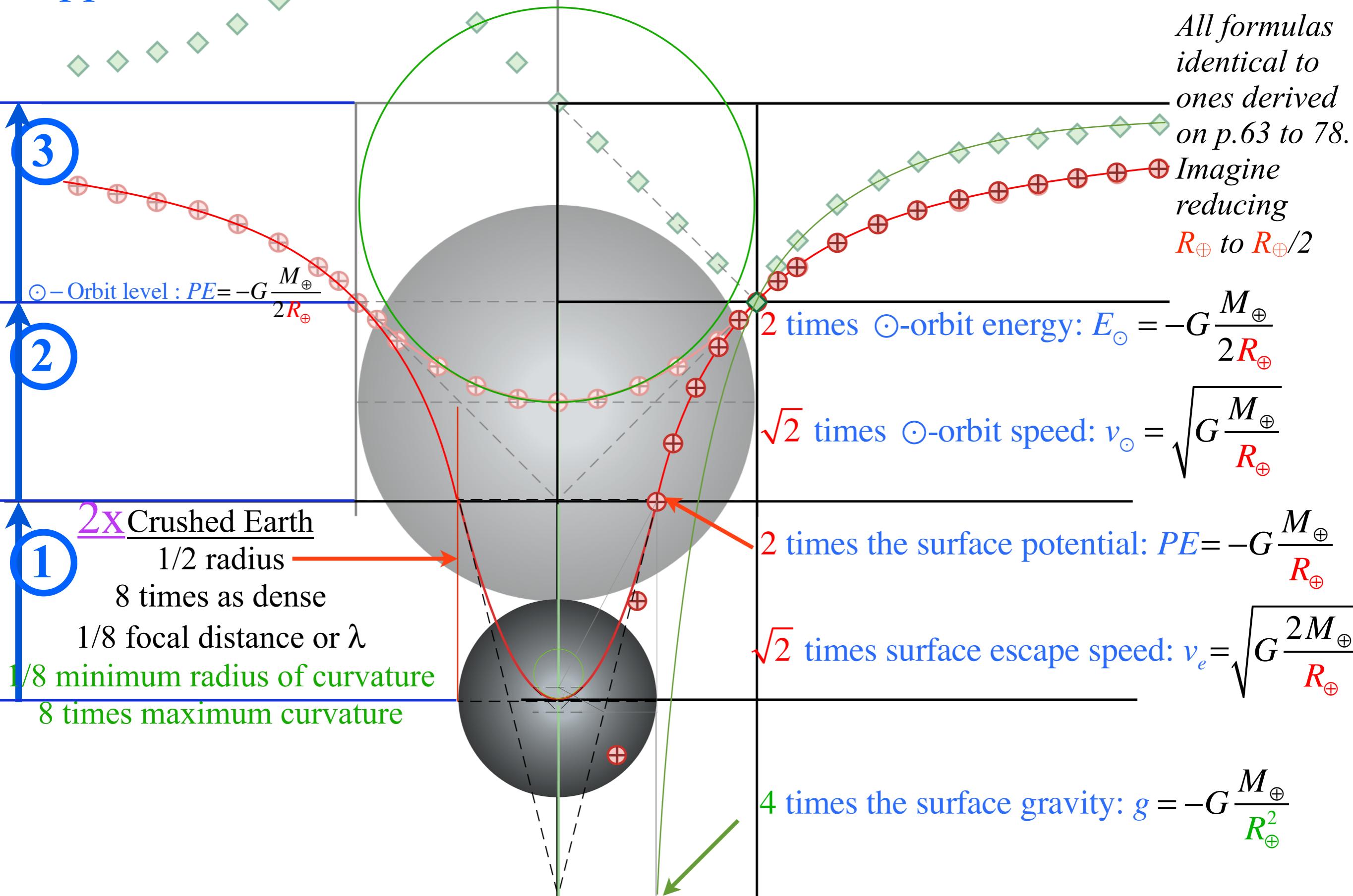


The “Three (equal) steps from Hell”



Suppose Earth radius crushed to 1/2: ($R_{\oplus}=6.4 \cdot 10^6 \text{ m}$ crushed to $R_{\oplus}/2=3.2 \cdot 10^6 \text{ m}$)

All formulas identical to ones derived on p.63 to 78.
Imagine reducing R_{\oplus} to $R_{\oplus}/2$



Geometry of idealized “Sophomore-physics Earth”

Coulomb field outside

Isotropic Harmonic Oscillator (IHO) field inside

Contact-geometry of potential curve(s)

“Crushed-Earth” models: 3 key energy “steps” and 4 key energy “levels”

Earth matter vs nuclear matter:

*Introducing the “neutron starlet” and “**Black-Hole-Earth**”*



Examples of “crushed” matter

Earth matter Earth mass : $M_{\oplus} = 5.9722 \times 10^{24} \text{ kg.} \approx 6.0 \cdot 10^{24} \text{ kg.}$ Density $\rho_{\oplus} = ??$

Earth radius : $R_{\oplus} = 6.371 \cdot 10^6 \text{ m} \approx 6.4 \cdot 10^6 \text{ m}$ Earth volume : $(4\pi / 3)R_{\oplus}^3 \approx 4 \cdot 260 \cdot 10^{18} \approx 10^{21} \text{ m}^3$

$$(6.4)^3 \approx 262 \text{ and } (4\pi/3)260 = 1098 \approx 10^3$$

Examples of “crushed” matter

$$\rho_{\oplus} = \frac{M_{\oplus}}{(4\pi/3)R_{\oplus}^3}$$

Earth matter Earth mass : $M_{\oplus} = 5.9722 \times 10^{24} \text{ kg.} \approx 6.0 \cdot 10^{24} \text{ kg.}$ Density $\rho_{\oplus} \sim 6.0 \cdot 10^{24-21} \sim 6 \cdot 10^3 \text{ kg/m}^3$

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$$(6.4)^3 \approx 262 \text{ and } (4\pi/3)260 = 1089 \sim 10^3$$

Density of solid Fe = $7.9 \cdot 10^3 \text{ kg/m}^3$
Density of liquid Fe = $6.9 \cdot 10^3 \text{ kg/m}^3$

Examples of “crushed” matter

$$\rho_{\oplus} = \frac{M_{\oplus}}{(4\pi/3)R_{\oplus}^3}$$

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Nuclear matter Nucleon mass = $1.67 \cdot 10^{-27} \text{ kg.} \sim 2 \cdot 10^{-27} \text{ kg.}$ (“fingertip physics”)

Say a nucleus of atomic weight 50 has a radius of 3 fm , or 50 nucleons each with a mass $2 \cdot 10^{-27} \text{ kg.}$

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Say a nucleus of atomic weight 50 has a radius of 3 fm , or 50 nucleons each with a mass $2 \cdot 10^{-27} \text{ kg.}$

That's $100 \cdot 10^{-27} = 10^{-25} \text{ kg}$ packed into a volume of $4\pi/3r^3 = 4\pi/3 (3 \cdot 10^{-15})^3 \text{ m}^3$ or about 10^{-43} m^3 .

$$4\pi/3 \cdot 3^3 = 36\pi = 113 \sim 10^2$$

Examples of “crushed” matter

$$\rho_{\oplus} = \frac{M_{\oplus}}{(4\pi/3)R_{\oplus}^3}$$

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Nuclear density is $10^{-25+43} = 10^{18} \text{ kg/m}^3$ or a trillion (10^{12}) kilograms in a **fingertip** (1 cm^3)³.

$$(1 \text{ cm})^3 = (10^{-2} \text{ m})^3 = 10^{-6} \text{ m}^3$$

Examples of “crushed” matter

$$\rho_{\oplus} = \frac{M_{\oplus}}{(4\pi/3)R_{\oplus}^3}$$

Earth matter Earth mass : $M_{\oplus} = 5.9722 \times 10^{24} \text{ kg.} \approx 6.0 \cdot 10^{24} \text{ kg.}$ Density $\rho_{\oplus} \sim 6.0 \cdot 10^{24-21} \sim 6 \cdot 10^3 \text{ kg/m}^3$

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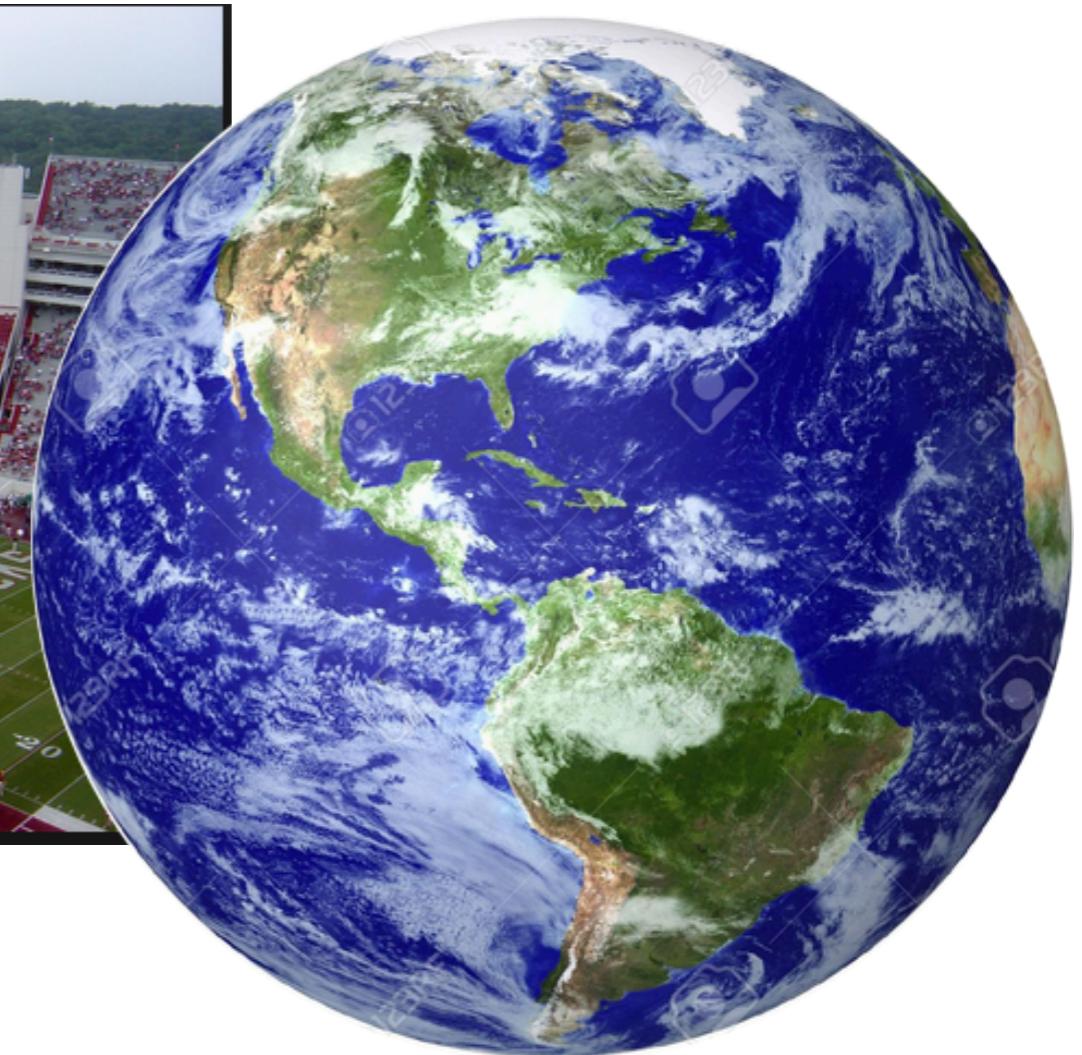
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Nuclear density is $10^{-25+43} = 10^{18} \text{ kg/m}^3$ or a trillion (10^{12}) kilograms in a **fingertip** (1 cm^3)³.

Earth radius crushed by a factor of $0.5 \cdot 10^{-5}$ to $R_{\text{crush}\oplus} \approx 300 \text{ m}$ would approach neutron-star density.



Geometry and algebra of idealized “Sophomore-physics Earth” fields

Coulomb field outside

Isotropic Harmonic Oscillator (IHO) field inside

Contact-geometry of potential curve(s) and “kite” geometry

“Ordinary-Earth” models: 3 key energy “steps” and 4 key energy “levels”

“Crushed-Earth” models: 3 key energy “steps” and 4 key energy “levels”

Earth matter vs nuclear matter:

→ *Introducing the “neutron starlet”*

*Fantasizing a “**Black-Hole-Earth**”*

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$$\rho_{\oplus} = \frac{M_{\oplus}}{(4\pi/3)R_{\oplus}^3}$$

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Earth radius : $R_{\oplus} = 6.371 \cdot 10^6 \text{ m} \approx 6.4 \cdot 10^6 \text{ m}$ Earth volume : $(4\pi/3)R_{\oplus}^3 \approx 4 \cdot 260 \cdot 10^{18} \sim 10^{21} \text{ m}^3$

Nuclear matter Nucleon mass = $1.67 \cdot 10^{-27} \text{ kg.} \sim 2 \cdot 10^{-27} \text{ kg.}$

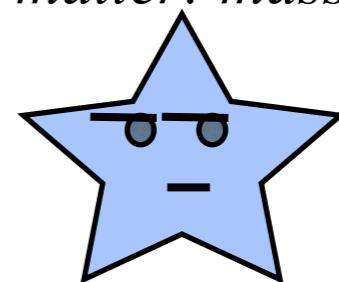
Say a nucleus of atomic weight 50 has a radius of 3 fm , or 50 nucleons each with a mass $2 \cdot 10^{-27} \text{ kg.}$

That's $100 \cdot 10^{-27} = 10^{-25} \text{ kg}$ packed into a volume of $4\pi/3r^3 = 4\pi/3 (3 \cdot 10^{-15})^3 \text{ m}^3$ or about 10^{-43} m^3 .

Nuclear density is $10^{-25+43} = 10^{18} \text{ kg/m}^3$ or a *trillion* (10^{12}) kilograms in the size of a fingertip.

Earth radius crushed by a factor of $0.5 \cdot 10^{-5}$ to $R_{\text{crush}\oplus} \approx 300 \text{ m}$ would approach neutron-star density.

Introducing the “Neutron starlet” 1 cm^3 of nuclear matter: mass = 10^{12} kg.



Geometry and algebra of idealized “Sophomore-physics Earth” fields

Coulomb field outside

Isotropic Harmonic Oscillator (IHO) field inside

Contact-geometry of potential curve(s) and “kite” geometry

“Ordinary-Earth” models: 3 key energy “steps” and 4 key energy “levels”

“Crushed-Earth” models: 3 key energy “steps” and 4 key energy “levels”

Earth matter vs nuclear matter:

Introducing the “neutron starlet”

→ *Fantasizing a “**Black-Hole-Earth**”*

Examples of “crushed” matter

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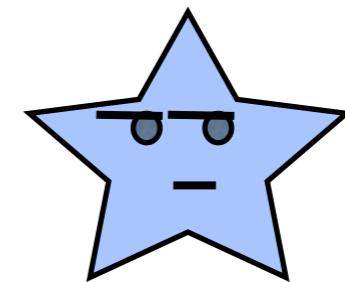
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surface escape velocity is the speed of light $c \cong 3.0 \cdot 10^8 \text{ m/s.}$

$c \equiv 299,792,458 \text{ m/s (EXACTLY)}$

$$V_{\text{escape}} = \sqrt{(2GM/R_{\oplus})}$$

(from p. 65, 66..., 75)

$$G = 6.67384(80) \cdot 10^{-11} \text{ Nm}^2/\text{C}^2 \sim (2/3)10^{-10}$$

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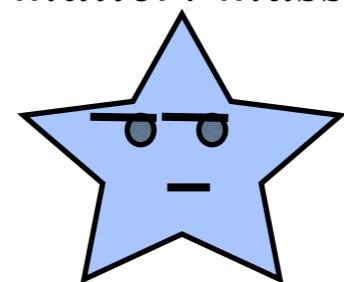
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$$c = \sqrt{(2GM/R_{\oplus})}$$

$$R_{\oplus} = 2GM/c^2 = 8.9 \text{ mm} \sim 1 \text{ cm}$$

(fingertip size!)

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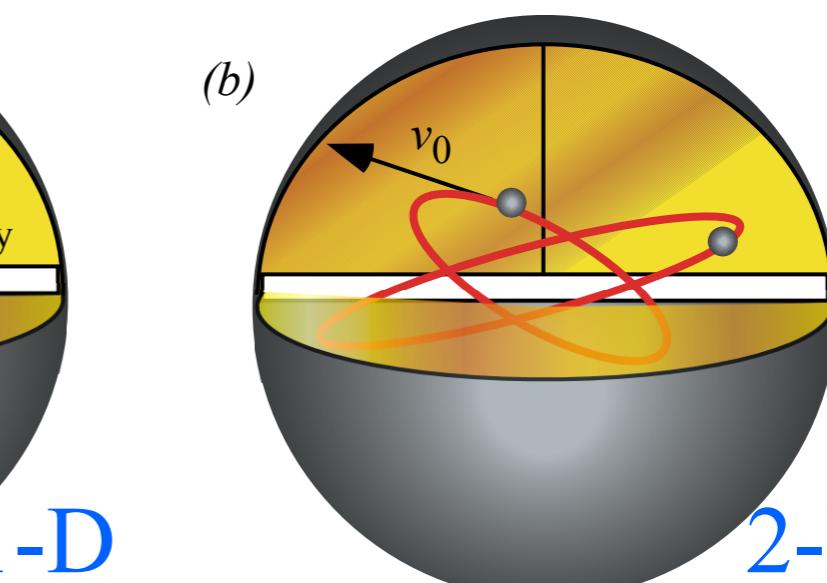
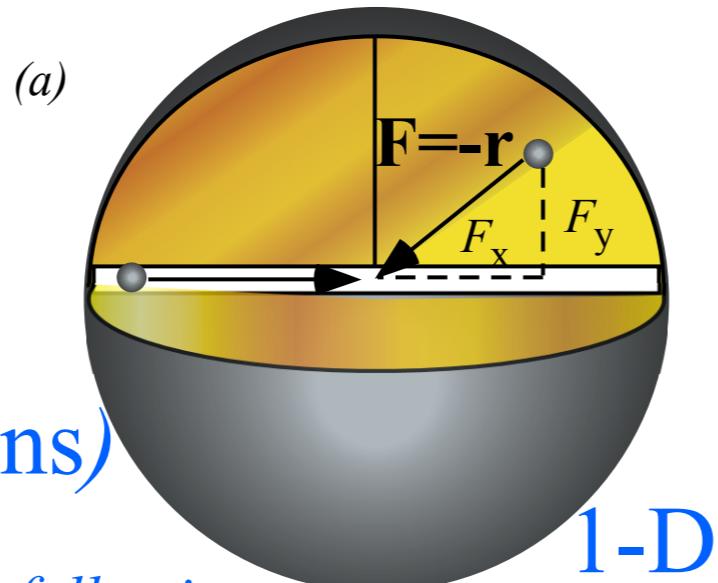
→ *Introducing 2D IHO orbits and phasor geometry*
Phasor “clock” geometry

Isotropic Harmonic Oscillator phase dynamics in uniform-body

I.H.O. Force law

$$F = -x \quad (\text{1-Dimension})$$

$$\mathbf{F} = -\mathbf{r} \quad (\text{2 or 3-Dimensions})$$



Unit 1
Fig. 9.10

1-D

2-D
(Paths are always
2-D ellipses if
viewed right!)

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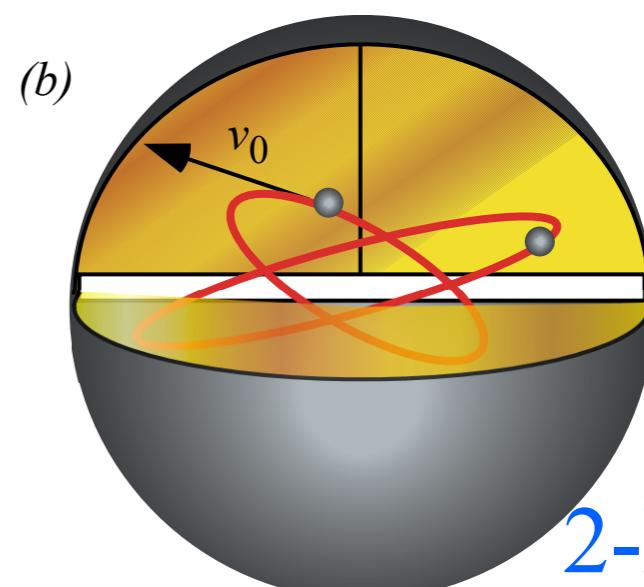
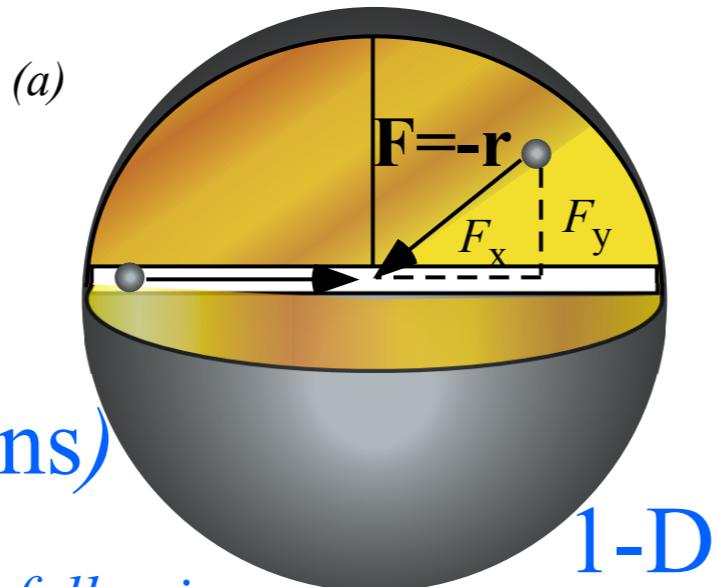
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2-D or 3-D

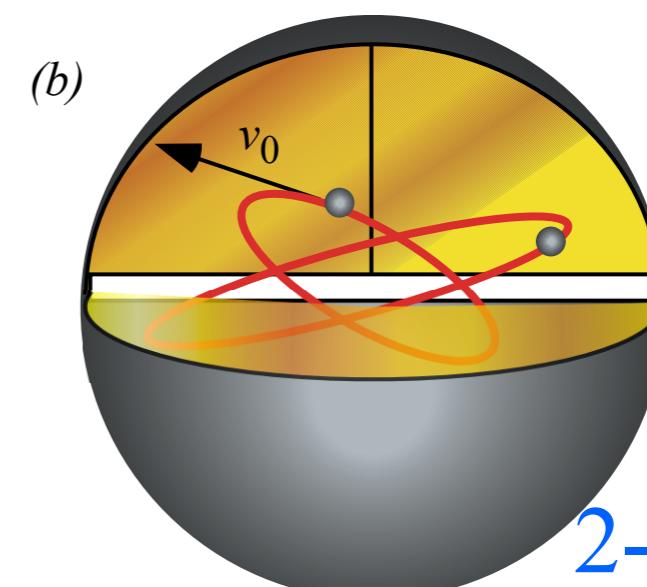
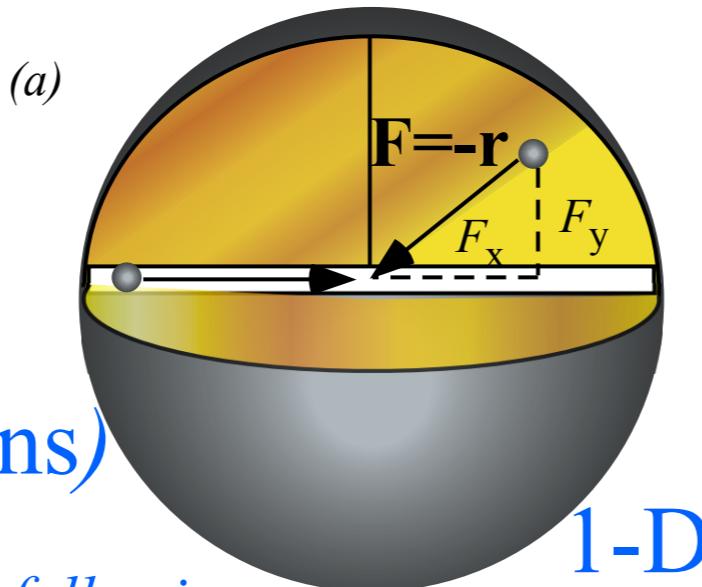
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velocity:

$$\text{Let : (1)} \quad v = \sqrt{2E/m} \cos\theta, \quad \text{and : (2)} \quad x = \sqrt{2E/k} \sin\theta$$

Another example of the old “scale-a-circle” trick...

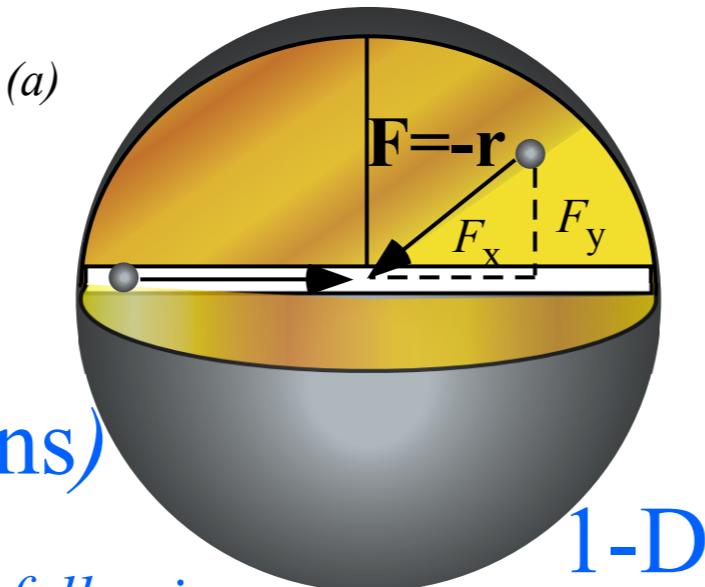
position:

Isotropic Harmonic Oscillator phase dynamics in uniform-body

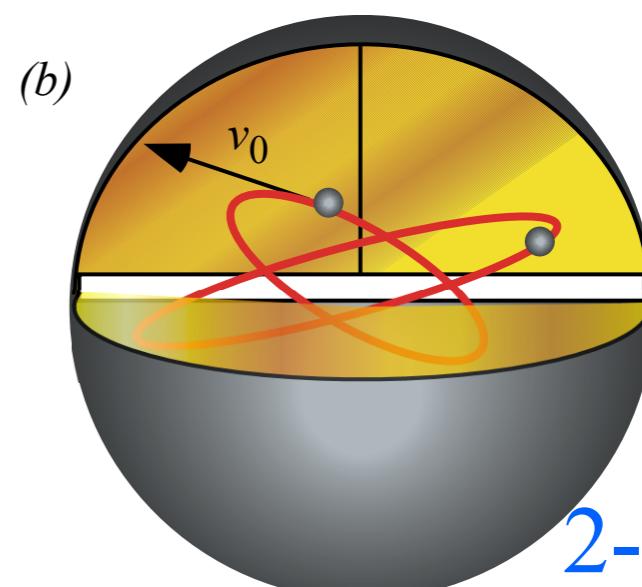
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1-D



Unit 1
Fig. 9.10

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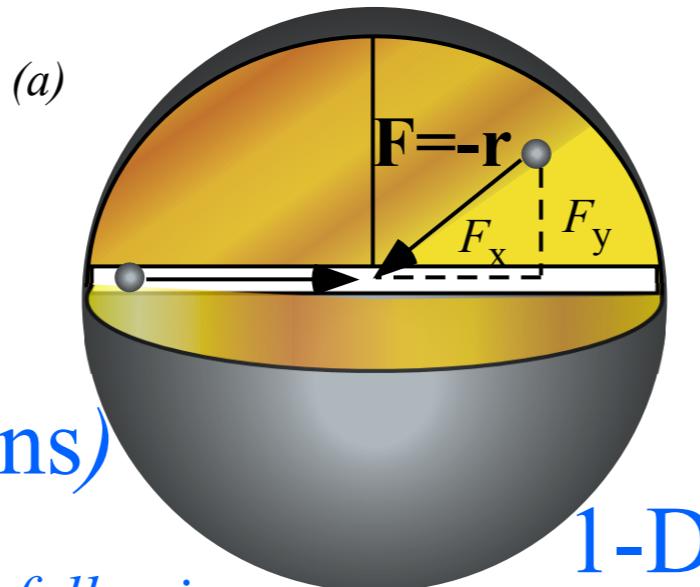
$$\text{angular velocity: } \omega = \frac{d\theta}{dt}$$

Isotropic Harmonic Oscillator phase dynamics in uniform-body

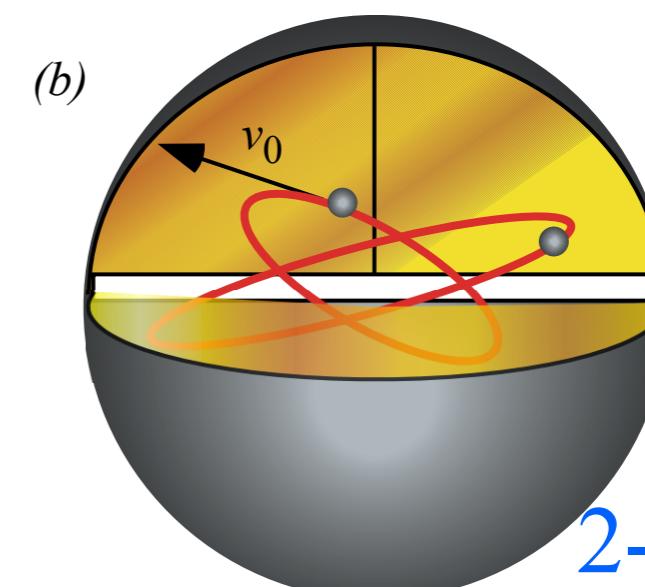
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1-D



Unit 1
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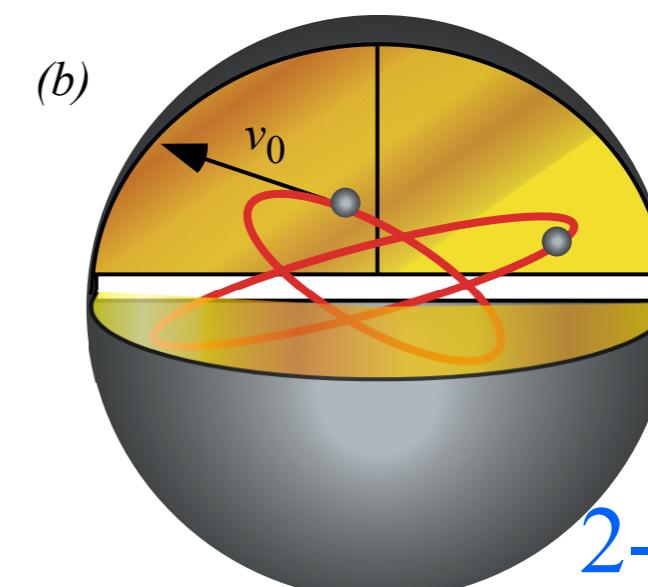
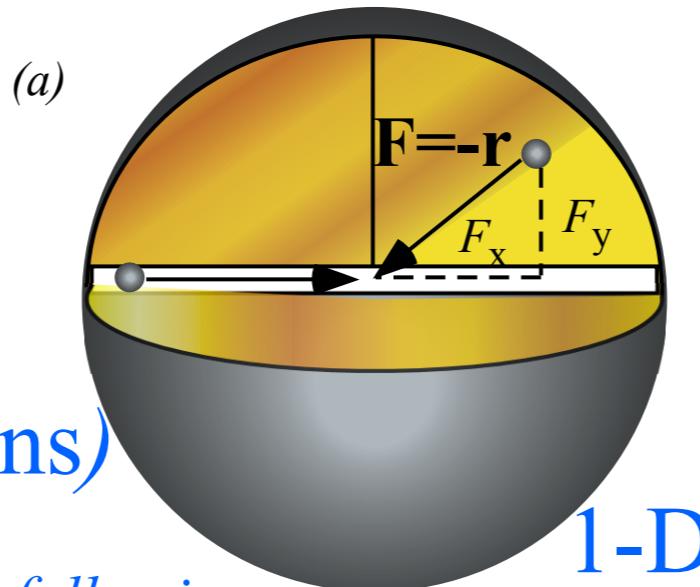
by (1)

Isotropic Harmonic Oscillator phase dynamics in uniform-body

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Unit 1
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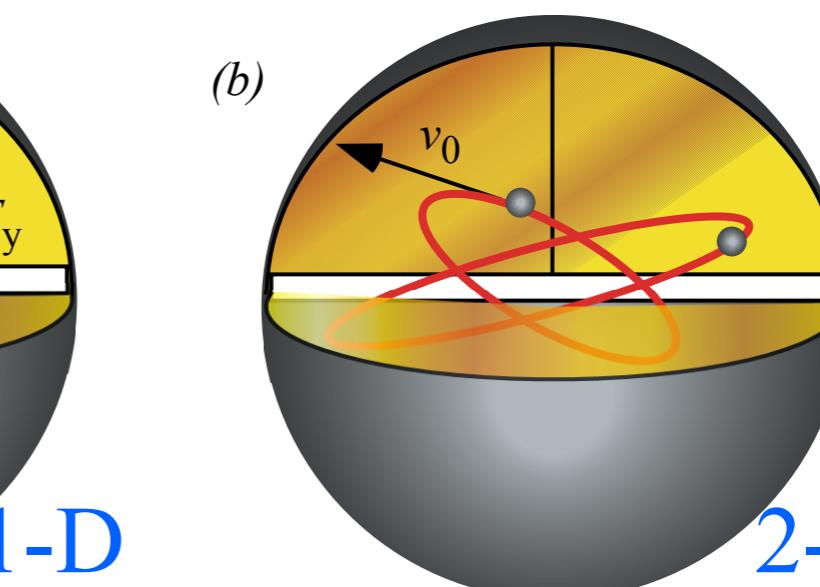
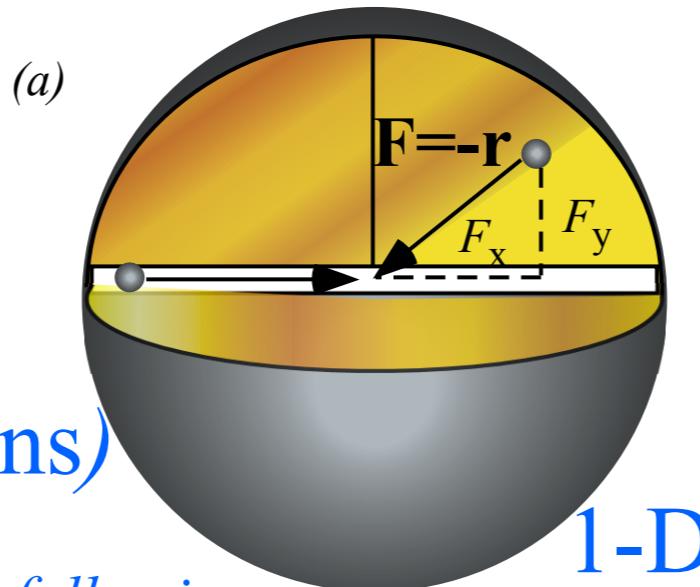
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Isotropic Harmonic Oscillator phase dynamics in uniform-body

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Unit 1
Fig. 9.10

1-D

2-D or 3-D

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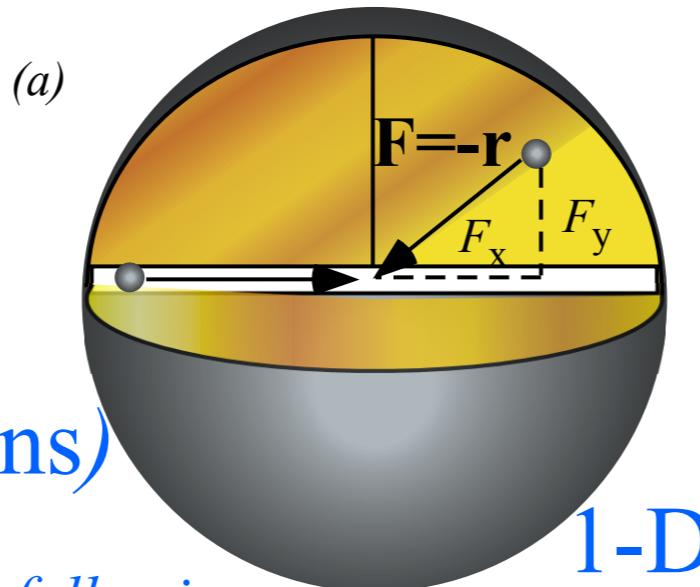
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Isotropic Harmonic Oscillator phase dynamics in uniform-body

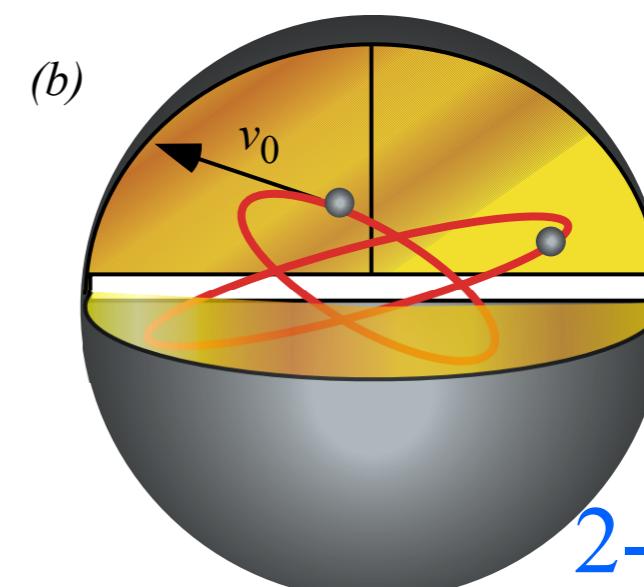
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1-D



Unit 1
Fig. 9.10

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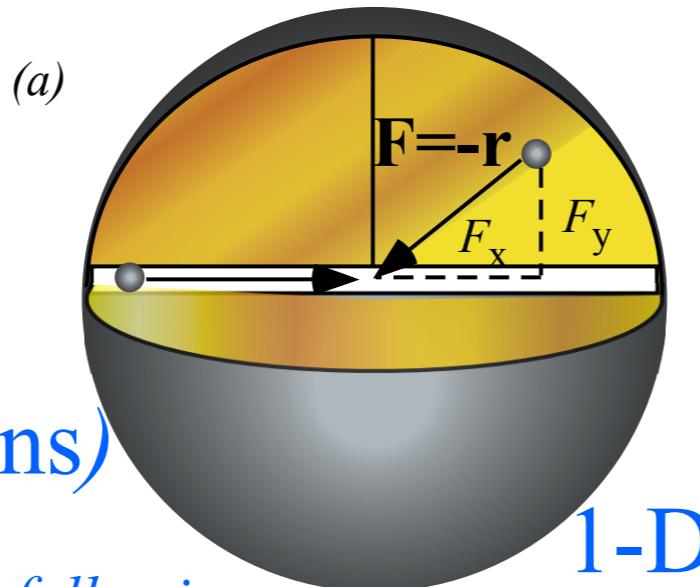
by (1)
 by def. (3)
 by (2)

Isotropic Harmonic Oscillator phase dynamics in uniform-body

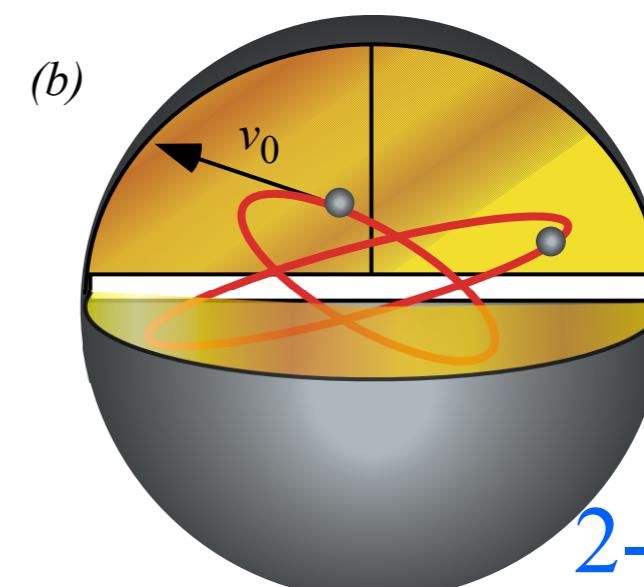
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1-D



Unit 1
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by (1) by def. (3) by (2)

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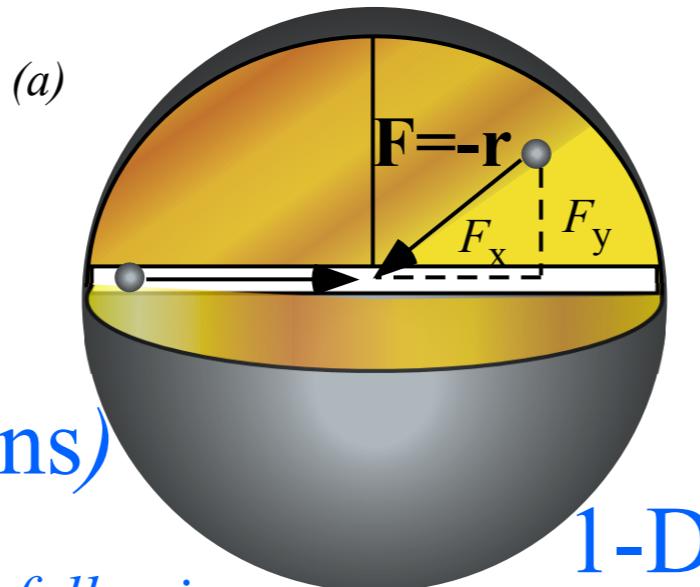
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Isotropic Harmonic Oscillator phase dynamics in uniform-body

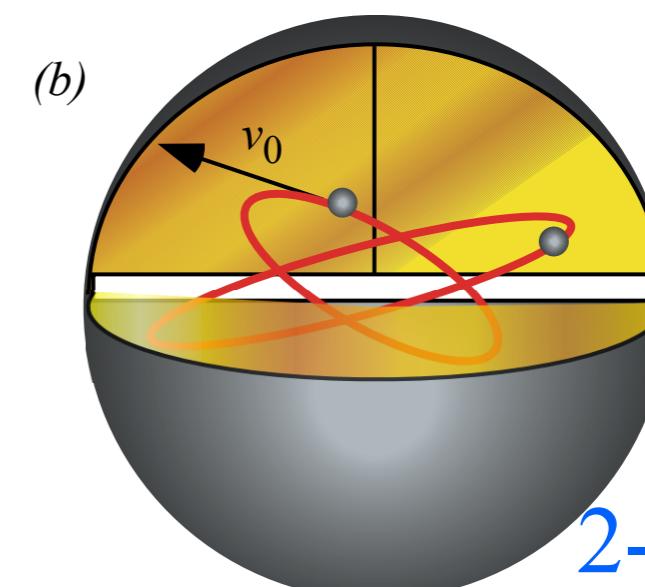
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1-D



Unit 1
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Another example of the old “scale-a-circle” trick...

$$\text{Let : (1)} \quad v = \sqrt{2E/m} \cos\theta, \quad \text{and : (2)} \quad x = \sqrt{2E/k} \sin\theta \quad \text{def. (3)} \quad \omega = \frac{d\theta}{dt}$$

$$\sqrt{\frac{2E}{m}} \cos\theta = v = \frac{dx}{dt} = \frac{d\theta}{dt} \frac{dx}{d\theta} = \omega \frac{dx}{d\theta} = \omega \sqrt{\frac{2E}{k}} \cos\theta$$

by (1) by def. (3) by (2)

$$\omega = \frac{d\theta}{dt} = \frac{\sqrt{\frac{2E}{m}} \cos\theta}{\sqrt{\frac{2E}{k}} \cos\theta}$$

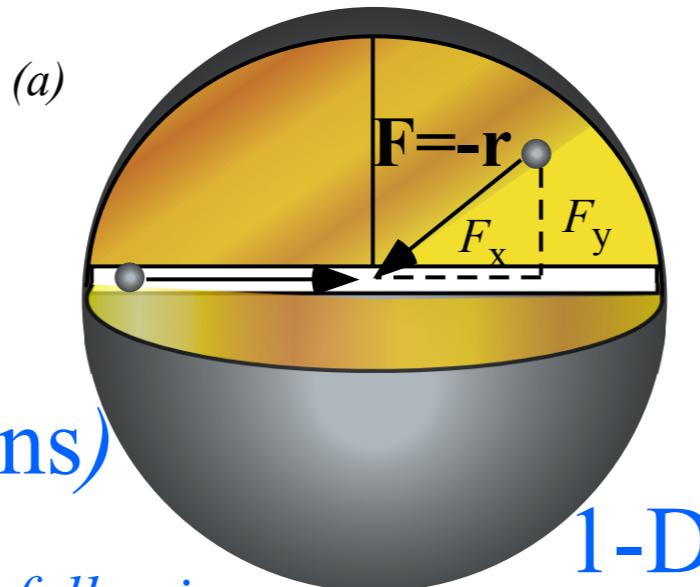
divide (1)
by (2) derivative

Isotropic Harmonic Oscillator phase dynamics in uniform-body

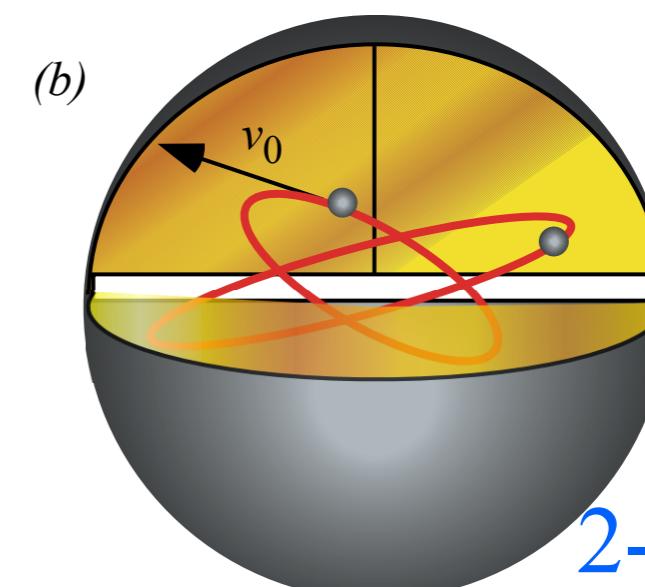
I.H.O. Force law

$$F = -x \quad (\text{1-Dimension})$$

$$\mathbf{F} = -\mathbf{r} \quad (\text{2 or 3-Dimensions})$$



1-D



Unit 1
Fig. 9.10

2-D or 3-D

(Paths are always 2-D ellipses if viewed right!)

Each dimension x, y , or z obeys the following:

$$\text{Total } E = KE + PE = \frac{1}{2}mv^2 + U(x) = \frac{1}{2}mv^2 + \frac{1}{2}kx^2 = \text{const.}$$

Equations for x -motion

[$x(t)$ and $v_x = v(t)$] are given first. They apply as well to dimensions [$y(t)$ and $v_y = v(t)$] and [$z(t)$ and $v_z = v(t)$] in the ideal isotropic case.

$$1 = \frac{mv^2}{2E} + \frac{kx^2}{2E} = \left(\frac{v}{\sqrt{2E/m}} \right)^2 + \left(\frac{x}{\sqrt{2E/k}} \right)^2$$

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by (1) by def. (3) by (2)

$$\omega = \frac{d\theta}{dt} = \frac{\sqrt{\frac{2E}{m}} \cos\theta}{\sqrt{\frac{2E}{k}} \cos\theta} = \sqrt{\frac{k}{m}}$$

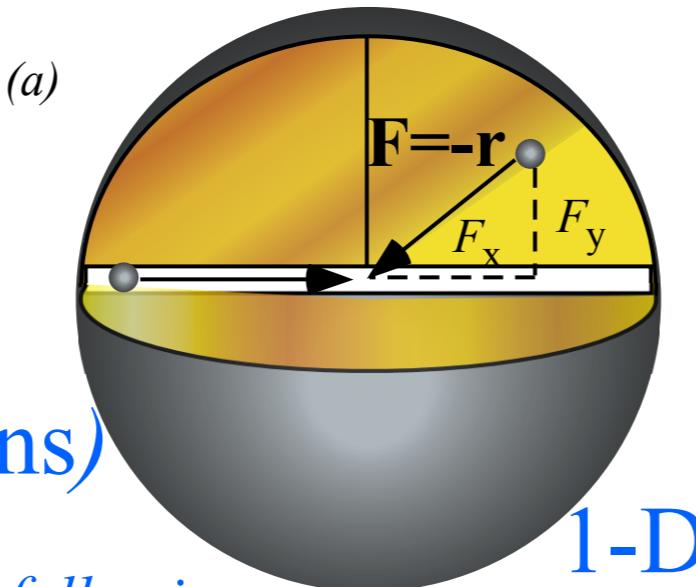
by def. (3) divide (1)
by (2) derivative

Isotropic Harmonic Oscillator phase dynamics in uniform-body

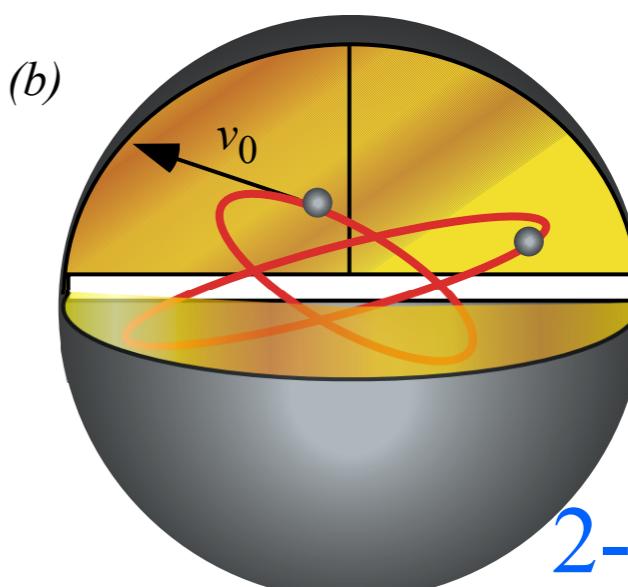
I.H.O. Force law

$F = -x$ (1-Dimension)

$\mathbf{F} = -\mathbf{r}$ (2 or 3-Dimensions)



1-D



Unit 1

Fig. 9.10

2-D or 3-D

(Paths are *always* 2-D ellipses if viewed right!)

Each dimension x, y, or z obeys the following:

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*Another example of
the old “scale-a-circle”
trick...*

Let : (1) $v = \sqrt{2E/m} \cos\theta$, and : (2) $x = \sqrt{2E/k} \sin\theta$ def. (3) $\omega = \frac{d\theta}{dt}$

$$\sqrt{\frac{2E}{m}} \cos \theta = v = \frac{dx}{dt} = \frac{d\theta}{dt} \frac{dx}{d\theta} = \omega \frac{dx}{d\theta} = \omega \sqrt{\frac{2E}{k}} \cos \theta$$

by (1) *by def. (3)* *by (2)*

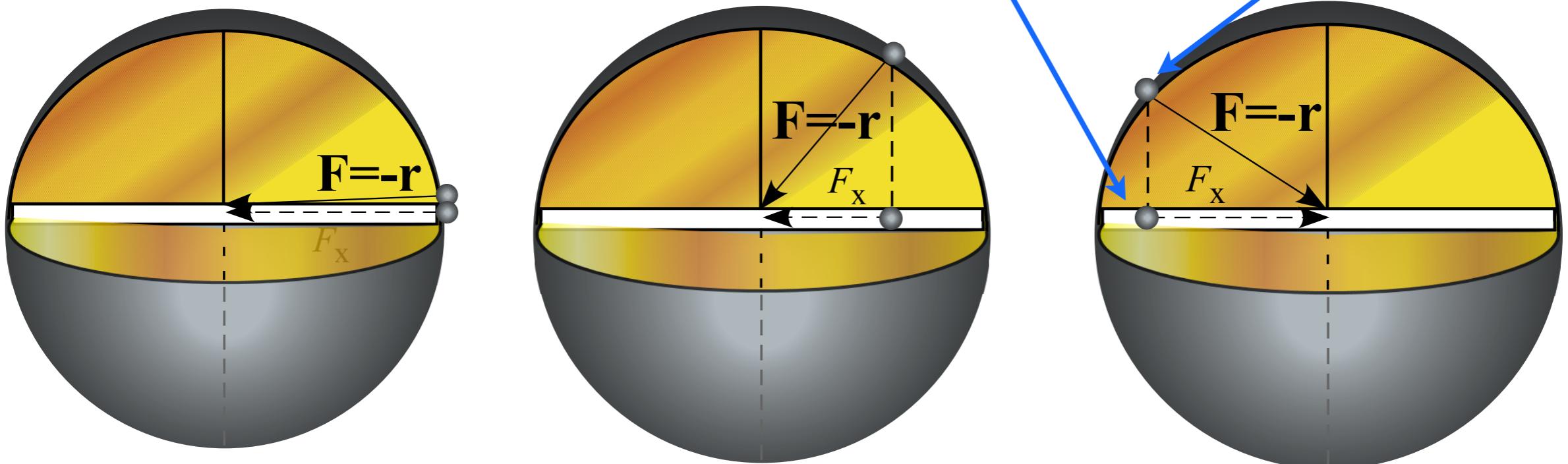
by def. (3)

by integration given constant ω :

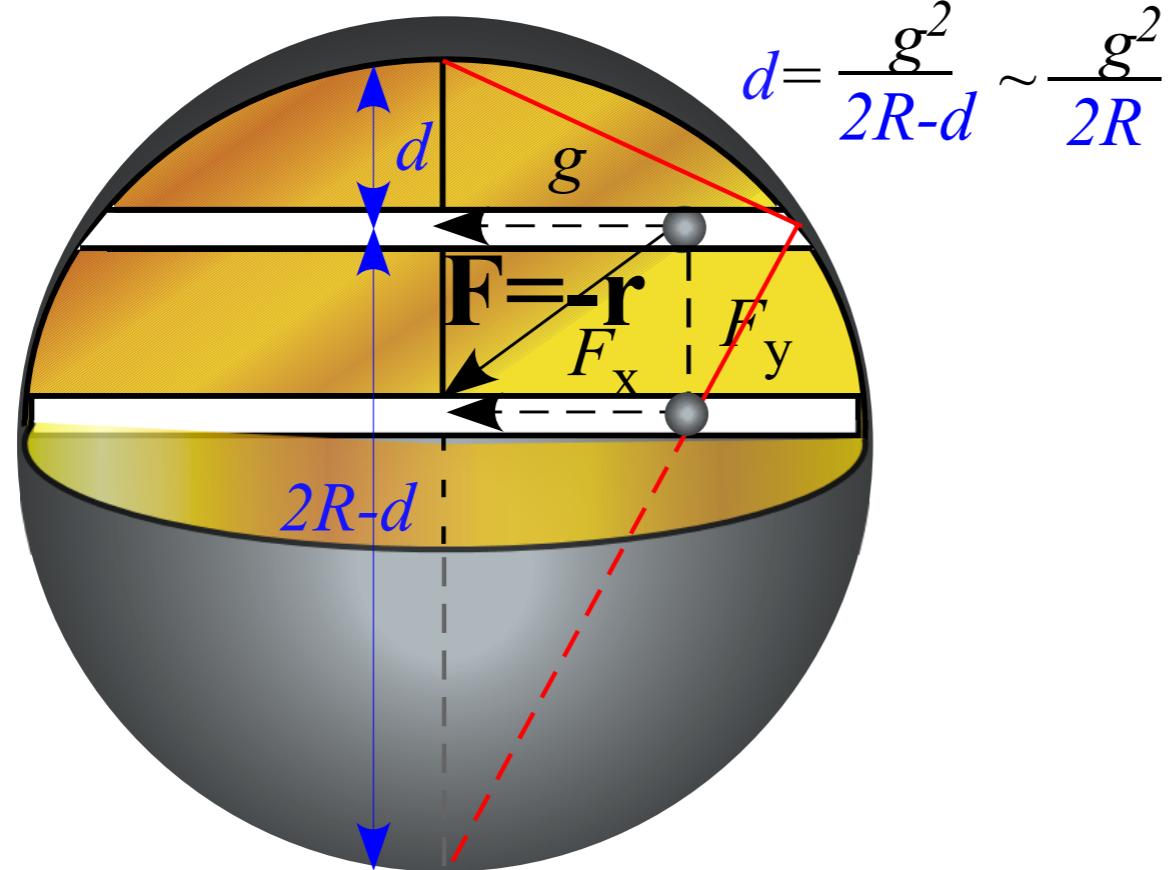


Introducing 2D IHO orbits and phasor geometry
Phasor “clock” geometry

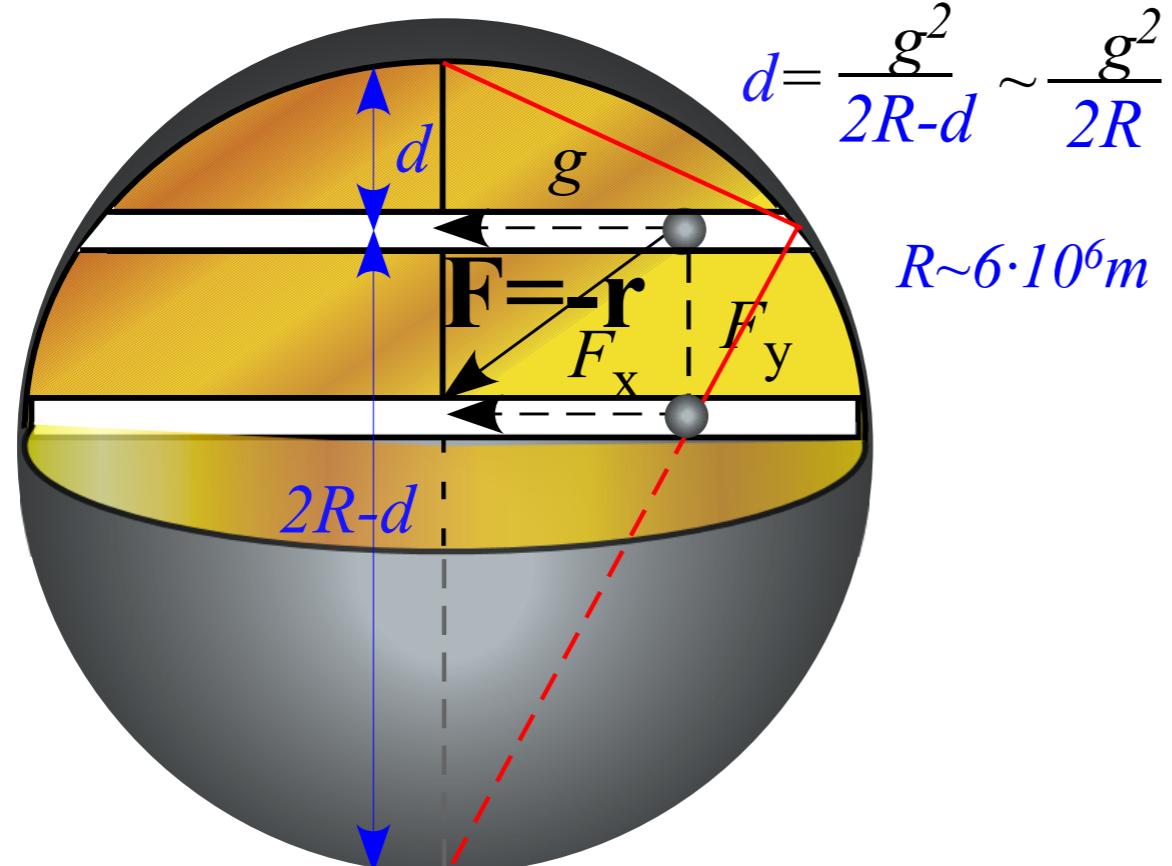
Isotropic Harmonic Oscillator makes tunneling ball track orbiting ball



*I*sotropic *H*armonic *O*scillator makes balls in parallel tunnel track each other



*I*sotropic *H*armonic *O*scillator makes balls in parallel tunnels track each other...



$$d = \frac{g^2}{2R-d} \sim \frac{g^2}{2R}$$

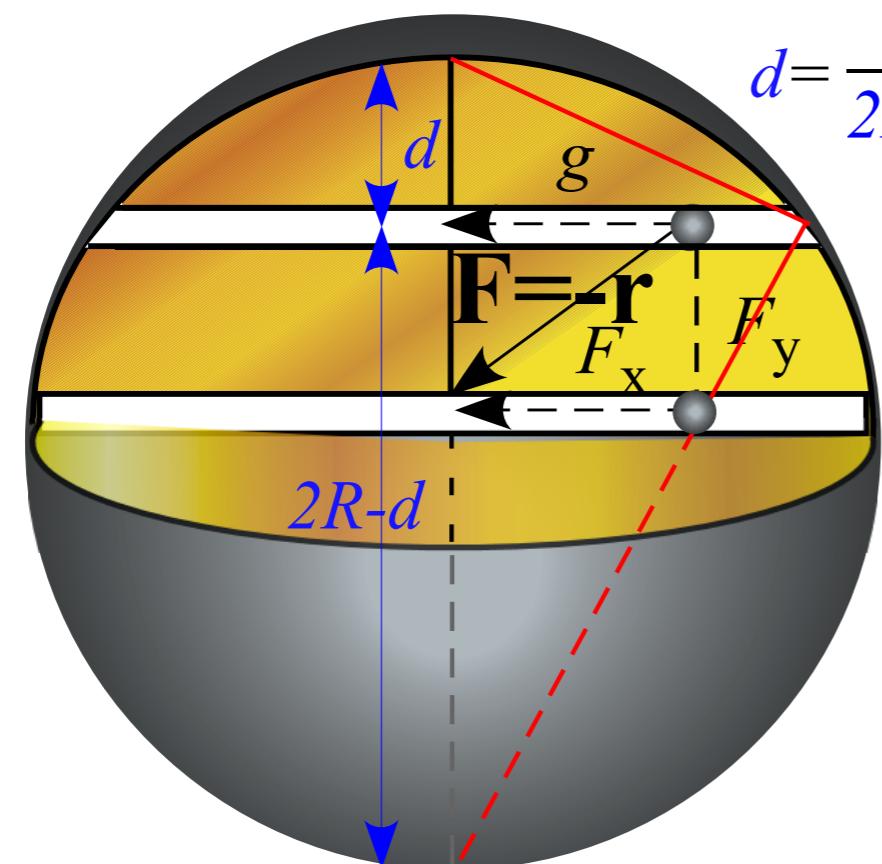
$$R \sim 6 \cdot 10^6 m$$

$$d \sim \frac{1}{2R}$$

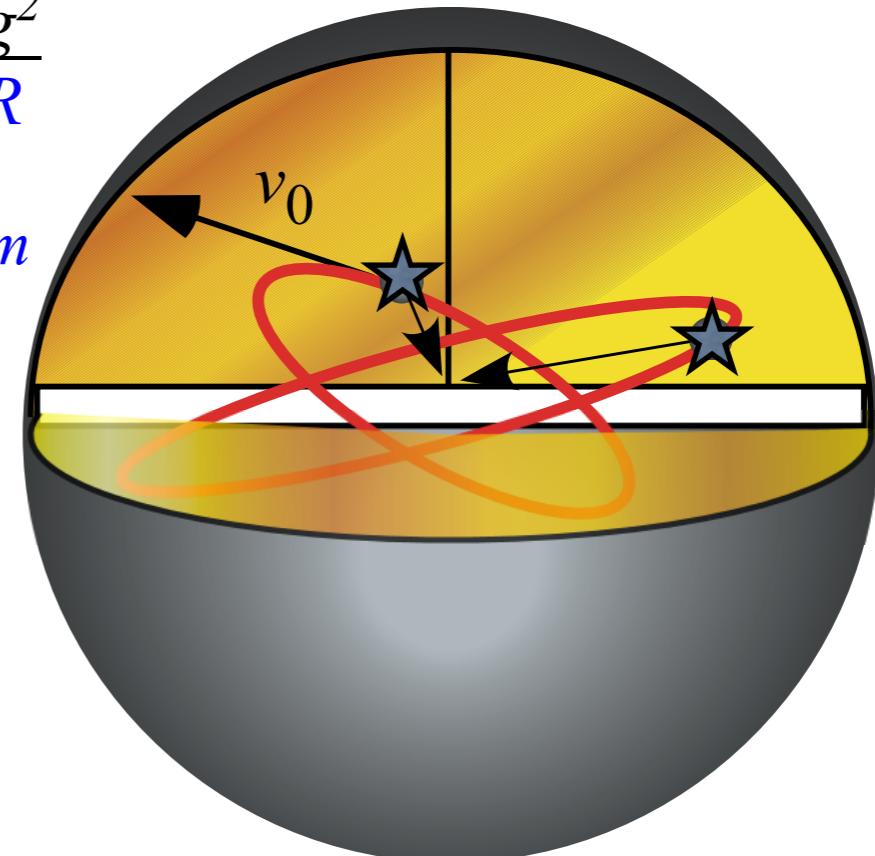
...even if track length is just $g = 1m$ so $d \sim (1/12)\text{micron}$

They all take about 84 minutes to go from right to left and back, again.

*I*sotropic *H*armonic *O*scillator makes balls in parallel tunnels track each other...



$$d = \frac{g^2}{2R-d} \sim \frac{g^2}{2R}$$
$$R \sim 6 \cdot 10^6 m$$

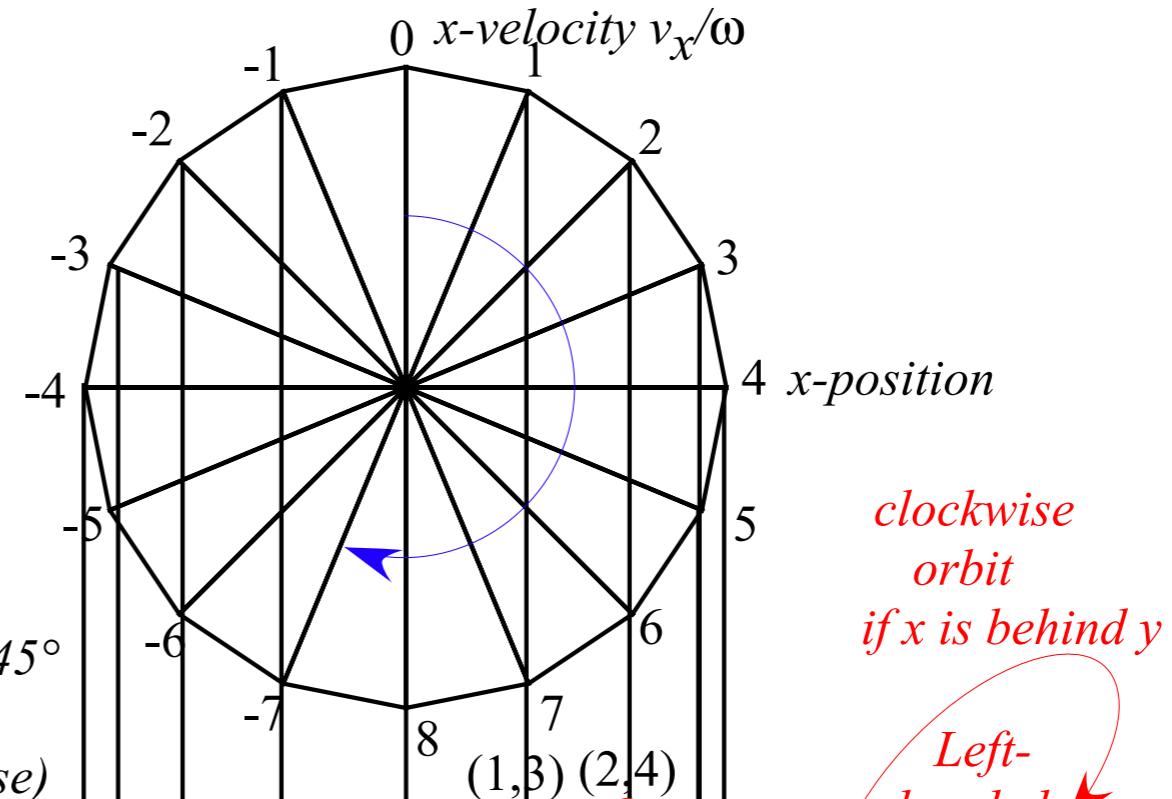
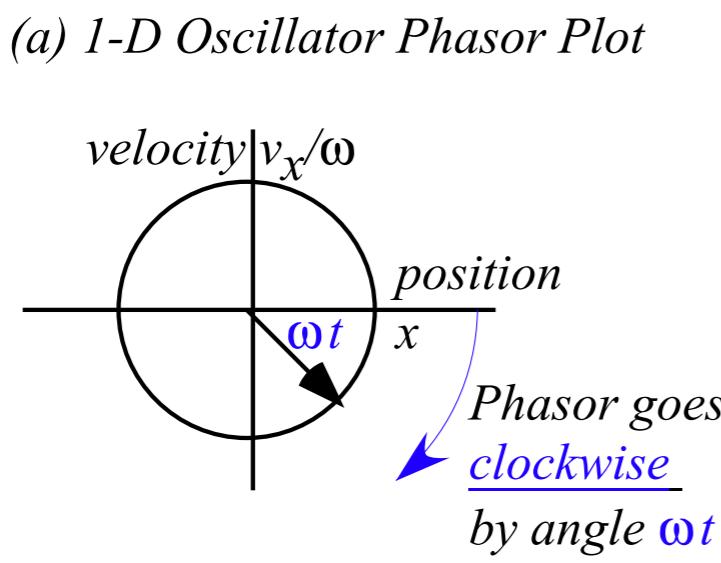


...even if track length is just $g = 1m$ so $d = (1/12)\text{micron}$

The all take about 84 minutes to go from right to left and back, again.

Most neutron starlet ()orbits are centered ellipses

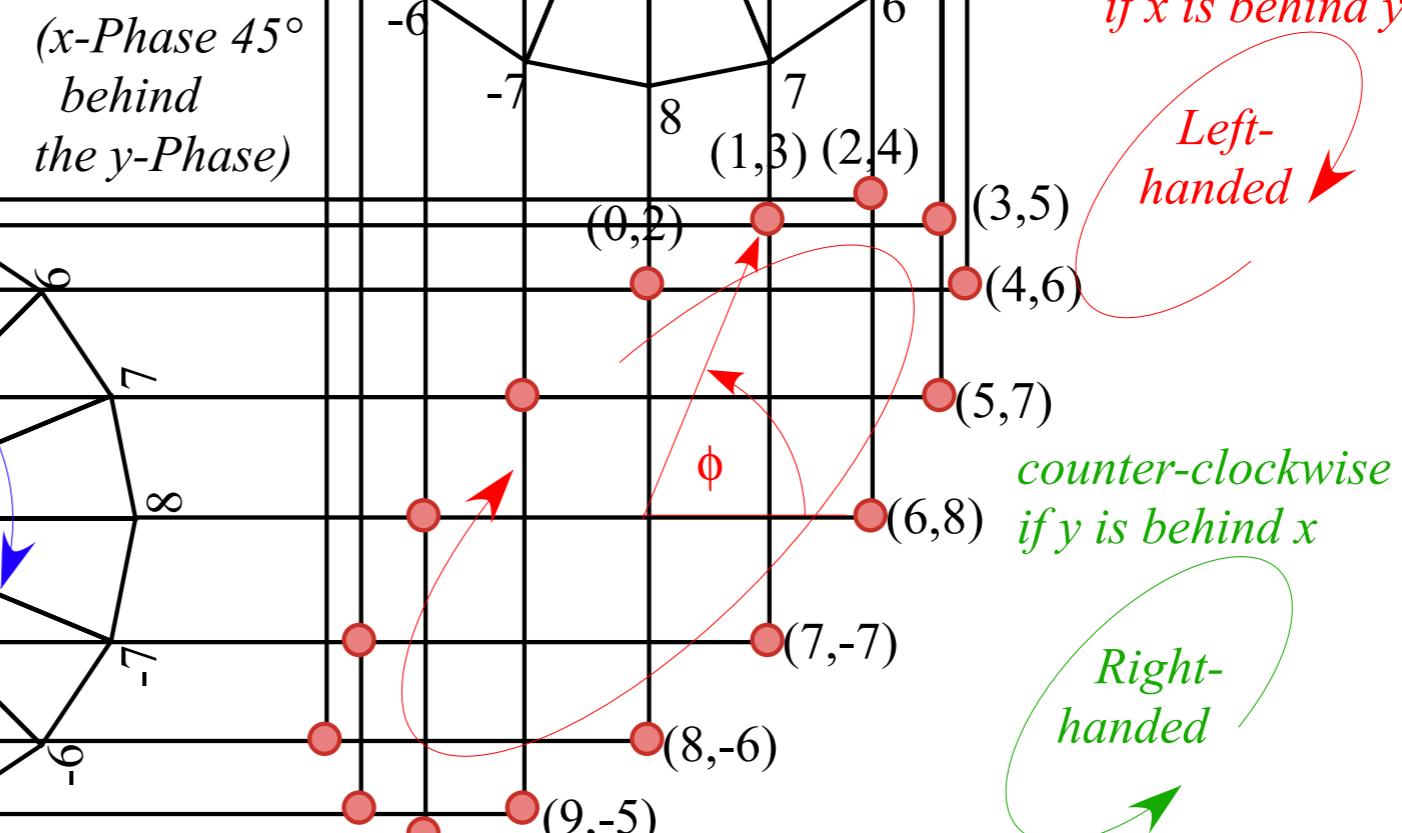
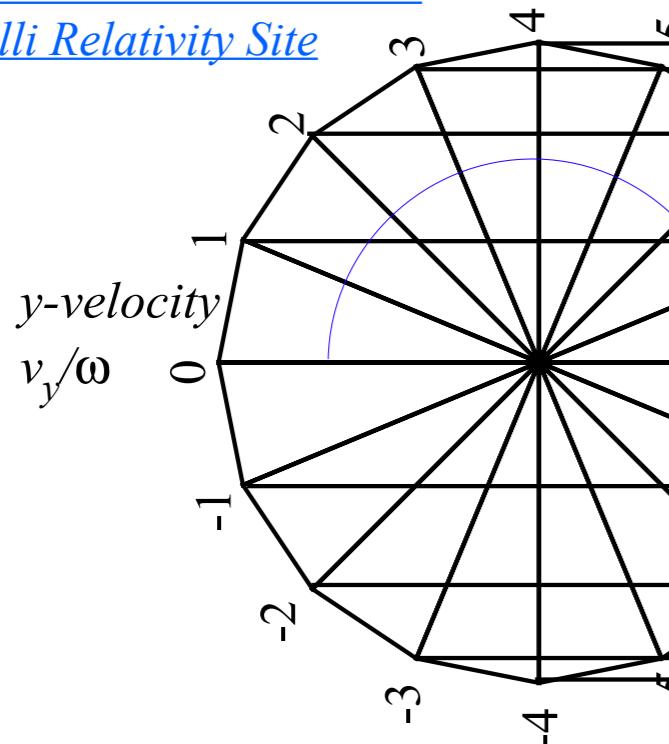
Isotropic Harmonic Oscillator phase dynamics in uniform-body



Unit 1
Fig. 9.10

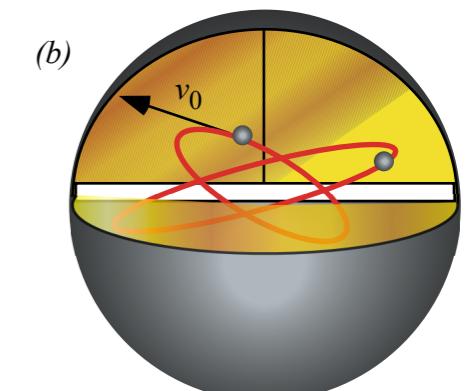
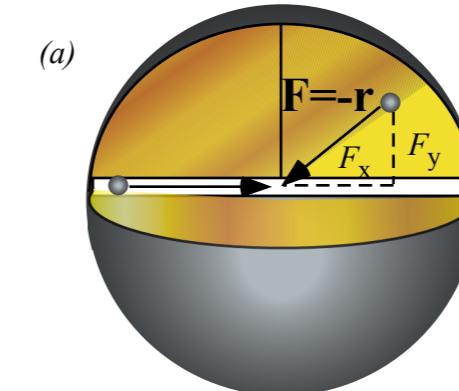
(b) 2-D Oscillator Phasor Plot

[Introduction to Phasors at our Pirelli Relativity Site](#)



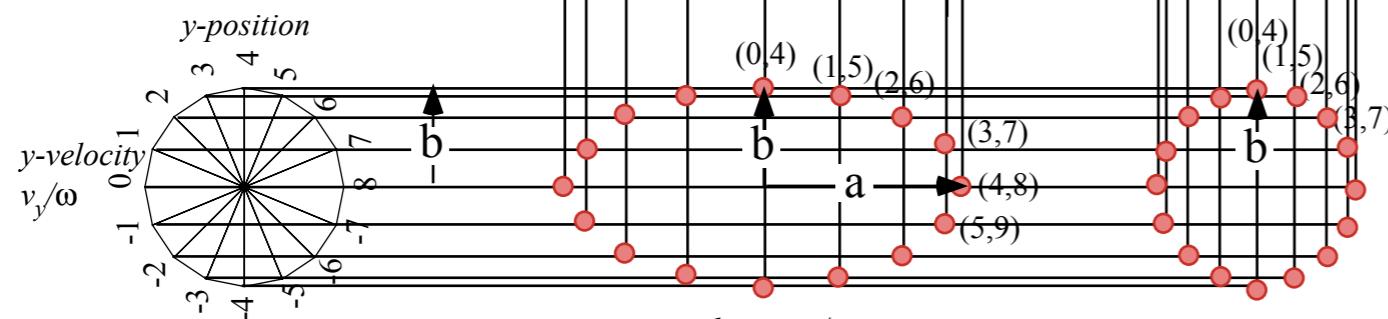
[BoxIt web simulation - With y-Phasor is on other side of xy plot](#)

[RelaWavity web simulation - Contact ellipsometry](#)



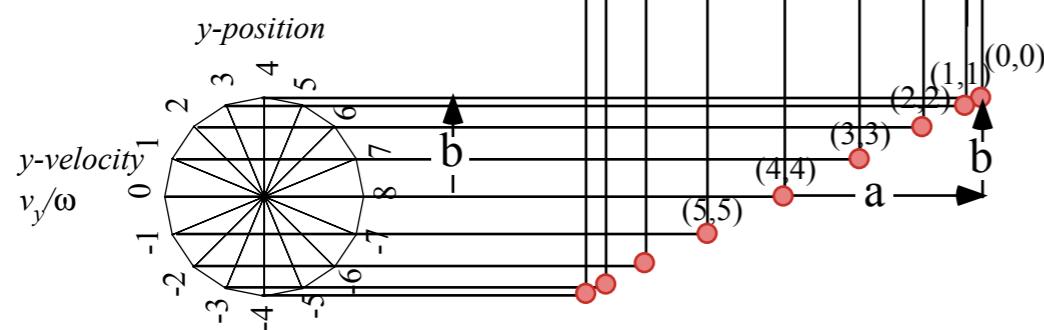
Unit 1
Fig. 9.12

(a) Phasor Plots
for
2-D Oscillator
or
Two 1D Oscillators
(x-Phase 90° behind
the y-Phase)



(b)
x-Phase 0° behind
the y-Phase

(In-phase case)



*These are more generic examples
with radius of x-phasor differing
from that of the y-phasor.*

[RelaWavity web simulation - Contact ellipsometry \(User Mouse Input allowed for setting phasor values\)](#)