

# Lecture 31

## Thur. 12.17.2015

Review of 16 relativity functions of  $\rho$  and related geometric approach to relativity  
Animation of  $e^{\rho=2}$  spacetime and per-spacetime plots

➔ *Rapidity*  $\rho$  related to *stellar aberration angle*  $\sigma$  and L. C. Epstein's approach to relativity

Longitudinal hyperbolic  $\rho$ -geometry connects to transverse circular  $\sigma$ -geometry

“Occams Sword” and summary of 16 parameter functions of  $\rho$  and  $\sigma$

Applications to optical waveguide, spherical waves, and accelerator radiation

*Learning about sin! and COS and...*

Derivation of relativistic quantum mechanics

What's the matter with mass? Shining some light on the Elephant in the room

Relativistic action and Lagrangian-Hamiltonian relations

Poincare' and Hamilton-Jacobi equations

Relativistic optical transitions and Compton recoil formulae

Feynman diagram geometry

Compton recoil related to rocket velocity formula

Comparing 2<sup>nd</sup>-quantization “photon” number  $N$  and 1<sup>st</sup>-quantization wavenumber  $\kappa$

*Relativity* in accelerated frames

Laser up-tuning by Alice and down-tuning by Carla makes  $g$ -acceleration grid

Analysis of constant- $g$  grid compared to zero- $g$  Minkowski grid

Animation of mechanics and metrology of constant- $g$  grid

# Lecture 31

## Thur. 12.10.2015

- ➔ Review of 16 *relawavity* functions of  $\rho$  and related geometric approach to relativity  
Animation of  $e^{\rho=2}$  spacetime and per-spacetime plots

*Rapidity  $\rho$  related to stellar aberration angle  $\sigma$  and L. C. Epstein's approach to relativity*

*Longitudinal hyperbolic  $\rho$ -geometry connects to transverse circular  $\sigma$ -geometry*

*“Occams Sword” and summary of 16 parameter functions of  $\rho$  and  $\sigma$*

*Applications to optical waveguide, spherical waves, and accelerator radiation*

*Learning about **sin** and **cos** and...*

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*Feynman diagram geometry*

*Compton recoil related to rocket velocity formula*

*Comparing 2<sup>nd</sup>-quantization “photon” number  $N$  and 1<sup>st</sup>-quantization wavenumber  $\kappa$*

*Relawavity in accelerated frames*

*Laser up-tuning by Alice and down-tuning by Carla makes  $g$ -acceleration grid*

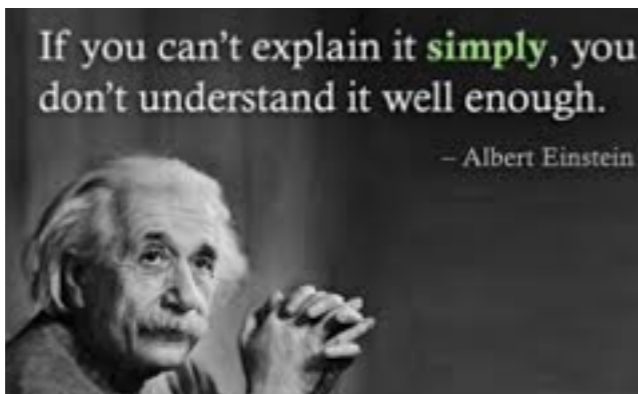
*Analysis of constant- $g$  grid compared to zero- $g$  Minkowski grid*

*Animation of mechanics and metrology of constant- $g$  grid*

# Two Famous-Name Coefficients

Review of Lect. 30 p.106

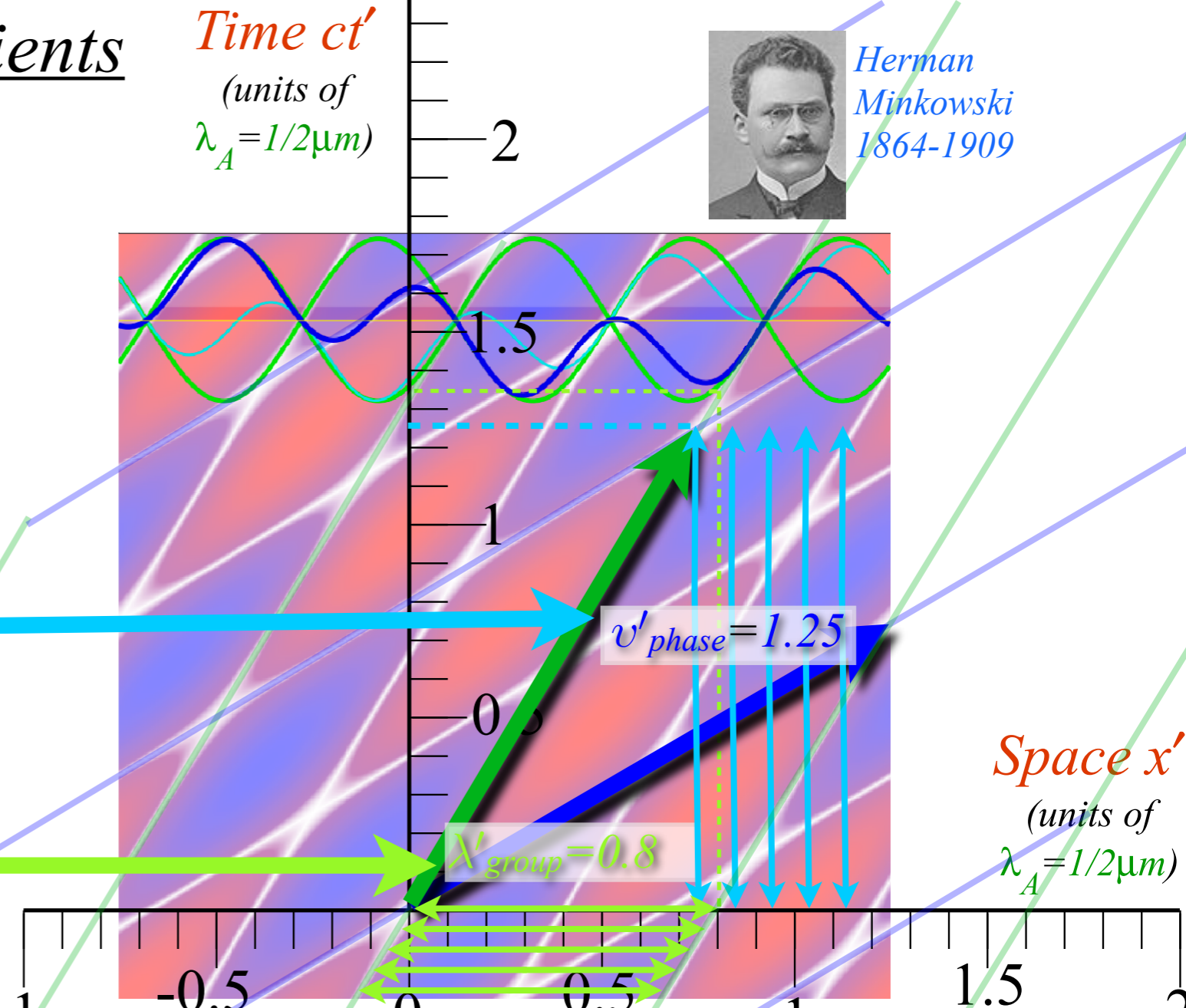
Albert Einstein  
1859-1955



Time  $ct'$   
(units of  $\lambda_A = 1/2\mu\text{m}$ )



Herman Minkowski  
1864-1909



This number is called an: **Einstein time-dilation** (dilated by 25% here)

This number is called a: **Lorentz length-contraction** (contracted by 20% here)



Hendrik A. Lorentz  
1853-1928

phase	$b_{RED}^{Doppler}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
group	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\text{coth } \rho$	$e^{+\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

## Old-Fashioned Notation

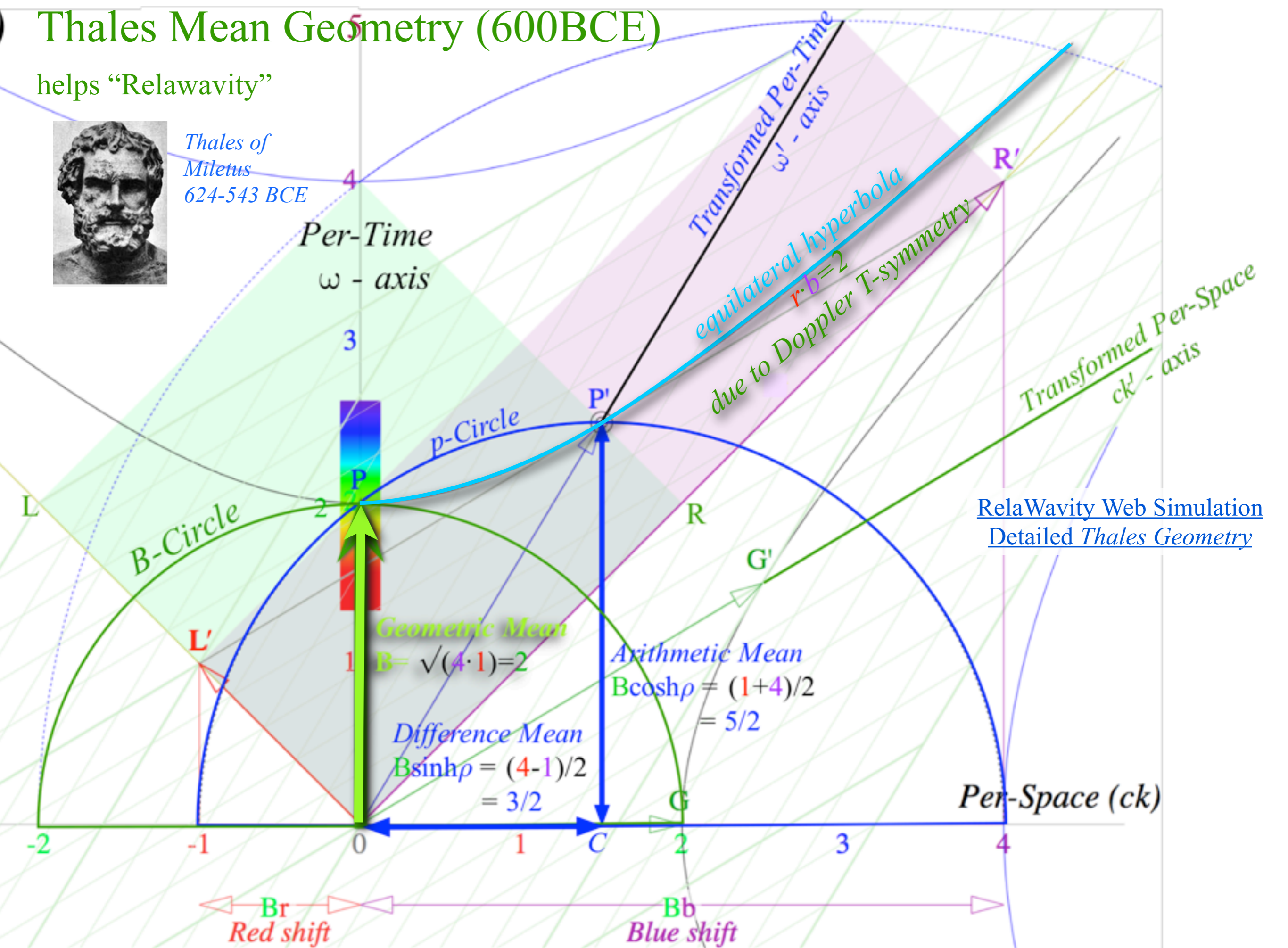
RelaWavity Web Simulation - Relativistic Terms  
(Expanded Table)

# Thales Mean Geometry (600BCE)

helps "Relativity"



Thales of Miletus  
624-543 BCE

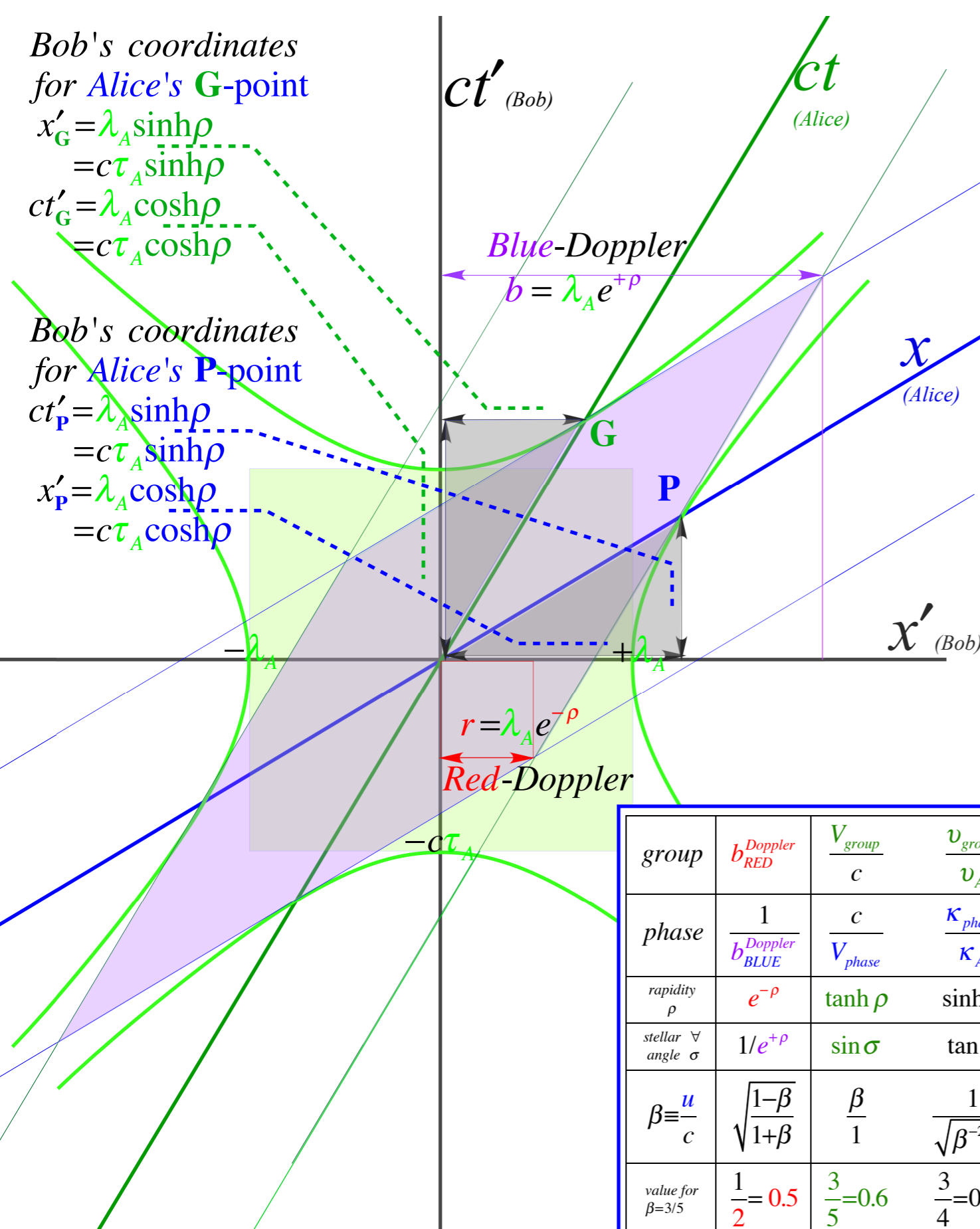


Bob's coordinates  
for Alice's **G**-point

$$\begin{aligned} x'_G &= \lambda_A \sinh \rho \\ &= c\tau_A \sinh \rho \\ ct'_G &= \lambda_A \cosh \rho \\ &= c\tau_A \cosh \rho \end{aligned}$$

Bob's coordinates  
for Alice's **P**-point

$$\begin{aligned} ct'_P &= \lambda_A \sinh \rho \\ &= c\tau_A \sinh \rho \\ x'_P &= \lambda_A \cosh \rho \\ &= c\tau_A \cosh \rho \end{aligned}$$



Space-time parameters

$$\begin{aligned} \lambda_{phase} &= \lambda_A \operatorname{csch} \rho \\ \lambda_{group} &= \lambda_A \operatorname{sech} \rho \\ c\tau_{phase} &= c\tau_A \operatorname{sech} \rho \\ c\tau_{group} &= c\tau_A \operatorname{csch} \rho \end{aligned}$$

Per-space-time parameters

$$\begin{aligned} cK_{phase} &= cK_A \sinh \rho \\ cK_{group} &= cK_A \cosh \rho \\ v_{phase} &= v_A \cosh \rho \\ v_{group} &= v_A \sinh \rho \end{aligned}$$

[RelaWavity Web Simulation](#)  
( $ct'$  vs  $x'$ ) with parameter table

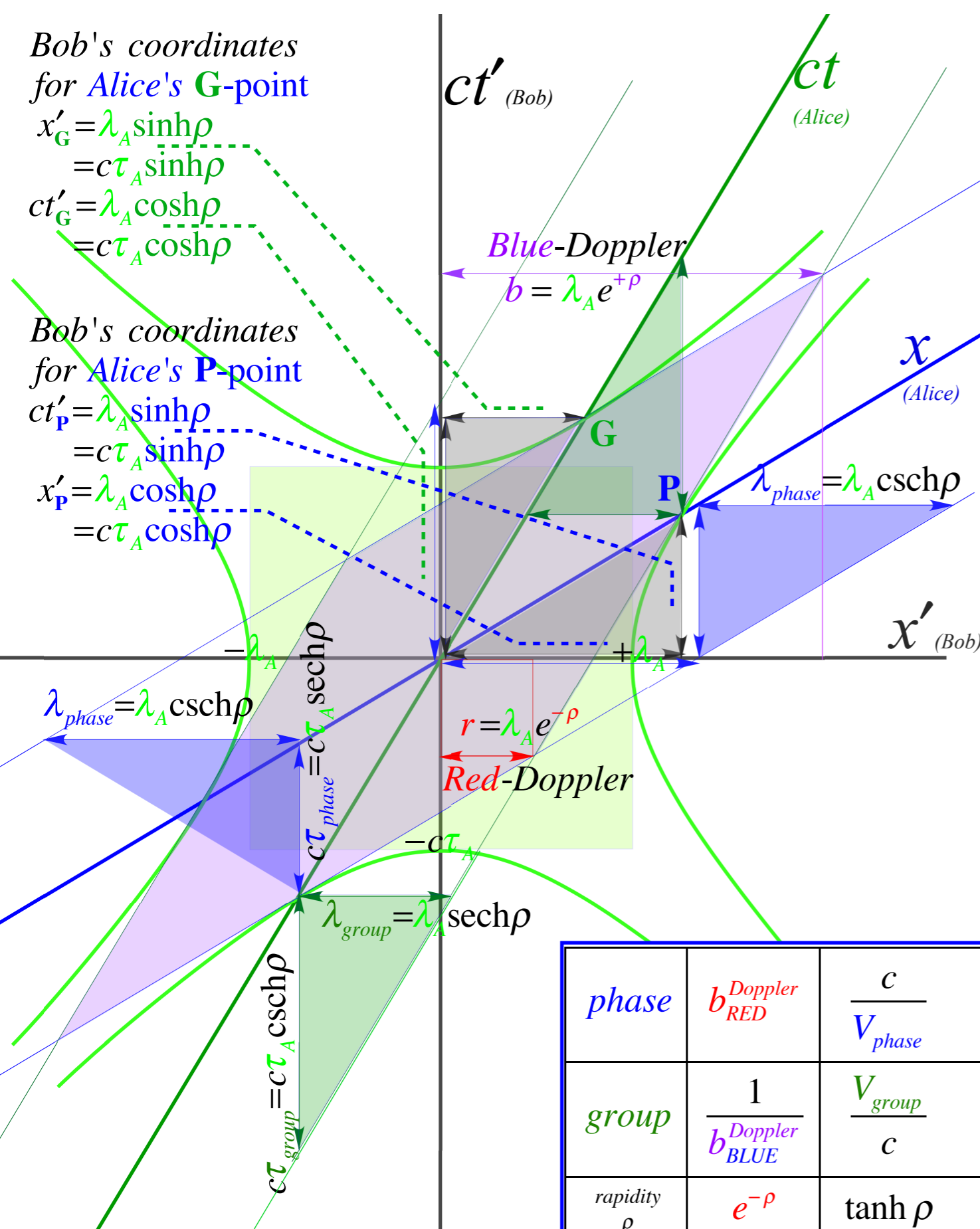
group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar $\nabla$ angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$
effects	$b_{RED}^{Doppler}$	$V_{group}$	past-future asymmetry (off-diagonal Lorentz-transform)	$x$ -contraction <sup>(Lorentz)</sup> $\tau_{phase}$ -contraction	$t$ -dilation <sup>(Einstein)</sup> $v_{phase}$ -dilation (on-diagonal Lorentz-transform)	inverse asymmetry	$V_{phase}$	$b_{BLUE}^{Doppler}$

Bob's coordinates  
for Alice's **G**-point

$$\begin{aligned} x'_G &= \lambda_A \sinh \rho \\ &= c\tau_A \sinh \rho \\ ct'_G &= \lambda_A \cosh \rho \\ &= c\tau_A \cosh \rho \end{aligned}$$

Bob's coordinates  
for Alice's **P**-point

$$\begin{aligned} ct'_P &= \lambda_A \sinh \rho \\ &= c\tau_A \sinh \rho \\ x'_P &= \lambda_A \cosh \rho \\ &= c\tau_A \cosh \rho \end{aligned}$$



Space-time parameters

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$$\begin{aligned} cK_{phase} &= cK_A \sinh \rho \\ cK_{group} &= cK_A \cosh \rho \\ v_{phase} &= v_A \cosh \rho \\ v_{group} &= v_A \sinh \rho \end{aligned}$$

$$\lambda_{phase} = \lambda_A \operatorname{csch} \rho$$

$$c\tau_{phase} = c\tau_A \operatorname{sech} \rho$$

$$\lambda_{group} = \lambda_A \operatorname{sech} \rho$$

$$c\tau_{group} = c\tau_A \operatorname{csch} \rho$$

$$r = \lambda_A e^{-\rho}$$

$$b = \lambda_A e^{+\rho}$$

<i>phase</i>	$b_{RED}^{Doppler}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
<i>group</i>	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
<i>rapidity</i> $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
<i>value for</i> $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

[RelaWavity Web Simulation](#)  
(*ct' vs x'*) with parameter table

Bob's coordinates for Alice's G-point

$$x'_G = \lambda_A \sinh \rho = c\tau_A \sinh \rho$$

$$ct'_G = \lambda_A \cosh \rho = c\tau_A \cosh \rho$$

Bob's coordinates for Alice's P-point

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$$x'_P = \lambda_A \cosh \rho = c\tau_A \cosh \rho$$

Space-time parameters

$$\lambda_{phase} = \lambda_A \operatorname{csch} \rho$$

$$\lambda_{group} = \lambda_A \operatorname{sech} \rho$$

$$c\tau_{phase} = c\tau_A \operatorname{sech} \rho$$

$$c\tau_{group} = c\tau_A \operatorname{csch} \rho$$

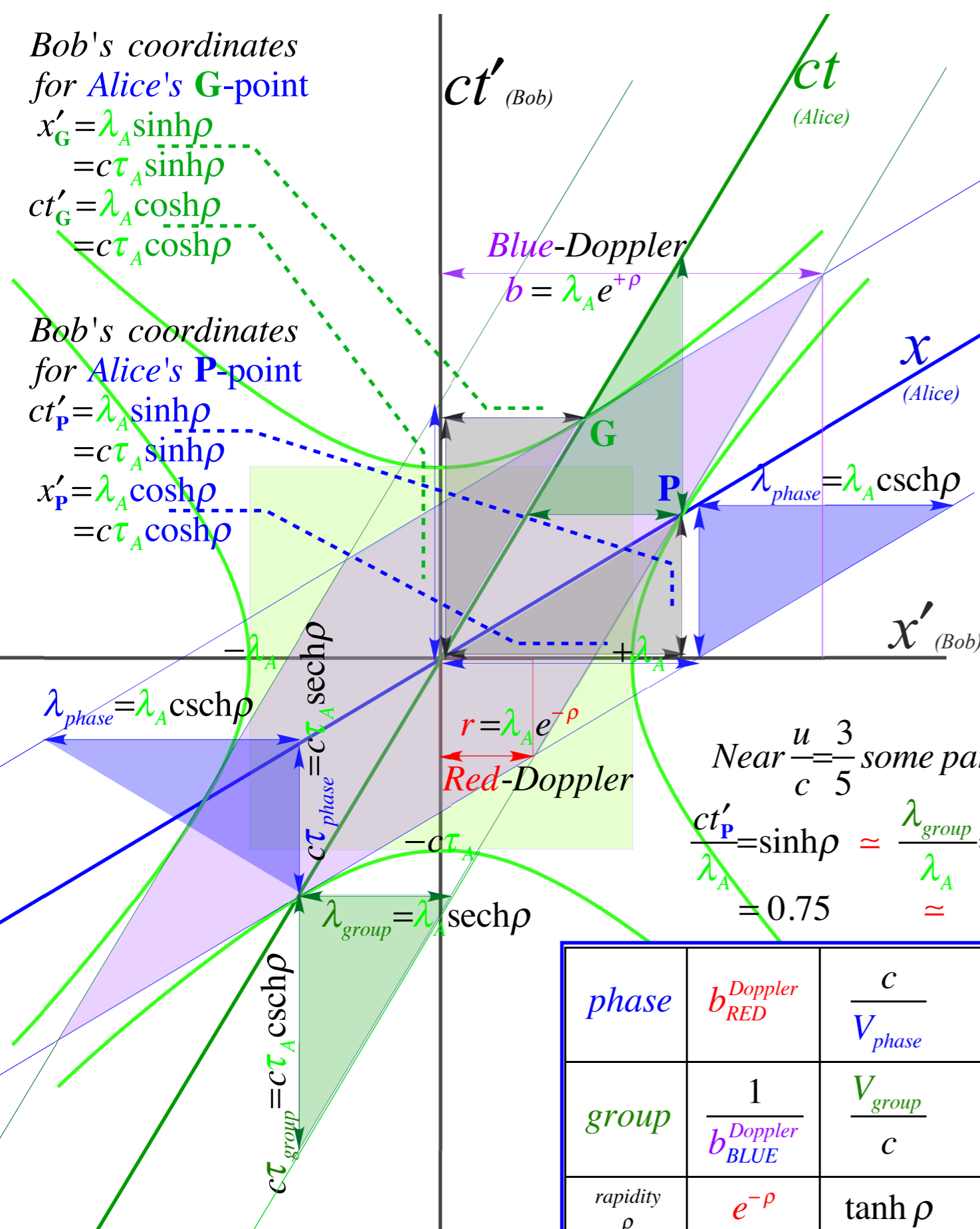
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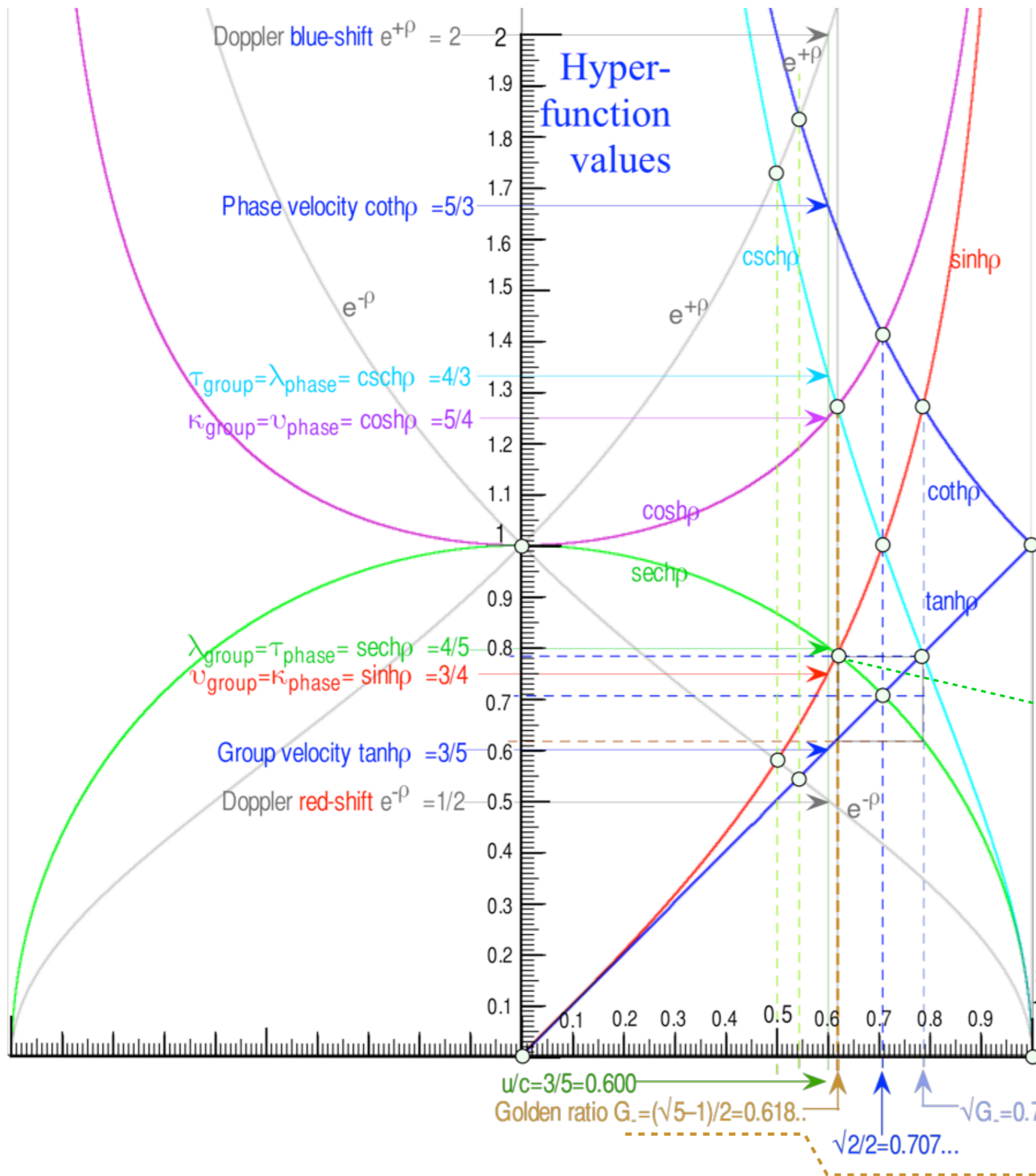
Near  $\frac{u}{c} = \frac{3}{5}$  some parameters are so close they *may be confused* on graphs :

$$\frac{ct'_P}{\lambda_A} = \sinh \rho \approx \frac{\lambda_{group}}{\lambda_A} = \frac{\tau_{phase}}{\tau_A} = \operatorname{sech} \rho \quad \text{and} \quad \frac{x'_P}{c\tau_A} = \cosh \rho \approx \frac{\lambda_{phase}}{\lambda_A} = \frac{\tau_{group}}{\tau_A} = \operatorname{csch} \rho$$

$$= 0.75 \qquad \qquad \qquad = 0.80 \qquad \qquad \qquad = 1.25 \qquad \qquad \qquad = 1.33$$

phase	$b_{RED}^{Doppler}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
group	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
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( $ct'$  vs  $x'$ ) with parameter table



If  $\frac{u}{c} = \tanh \rho = 0.618\dots$  (Golden-Mean  $G_-$ )

two parameters become *exactly equal* :

$$\frac{ct'_P}{c\tau_A} = \sinh \rho = \frac{\lambda_{group}}{\lambda_A} = \frac{\tau_{phase}}{\tau_A} = \operatorname{sech} \rho$$

$$= 0.786\dots = \sqrt{G_-} = 0.786\dots$$

and

$$\frac{x'_P}{\lambda_A} = \cosh \rho = \frac{\lambda_{phase}}{\lambda_A} = \frac{\tau_{group}}{\tau_A} = \operatorname{csch} \rho$$

$$= 1.272\dots = 1/\sqrt{G_-} = 1.272\dots$$

Solve :

$$\operatorname{sech} \rho = \sinh \rho$$

or:

$$\sinh \rho \cosh \rho = 1$$

or:

$$\sinh 2\rho = 2$$

$$\rho = \frac{1}{2} \sinh^{-1} 2 = 0.7218\dots$$

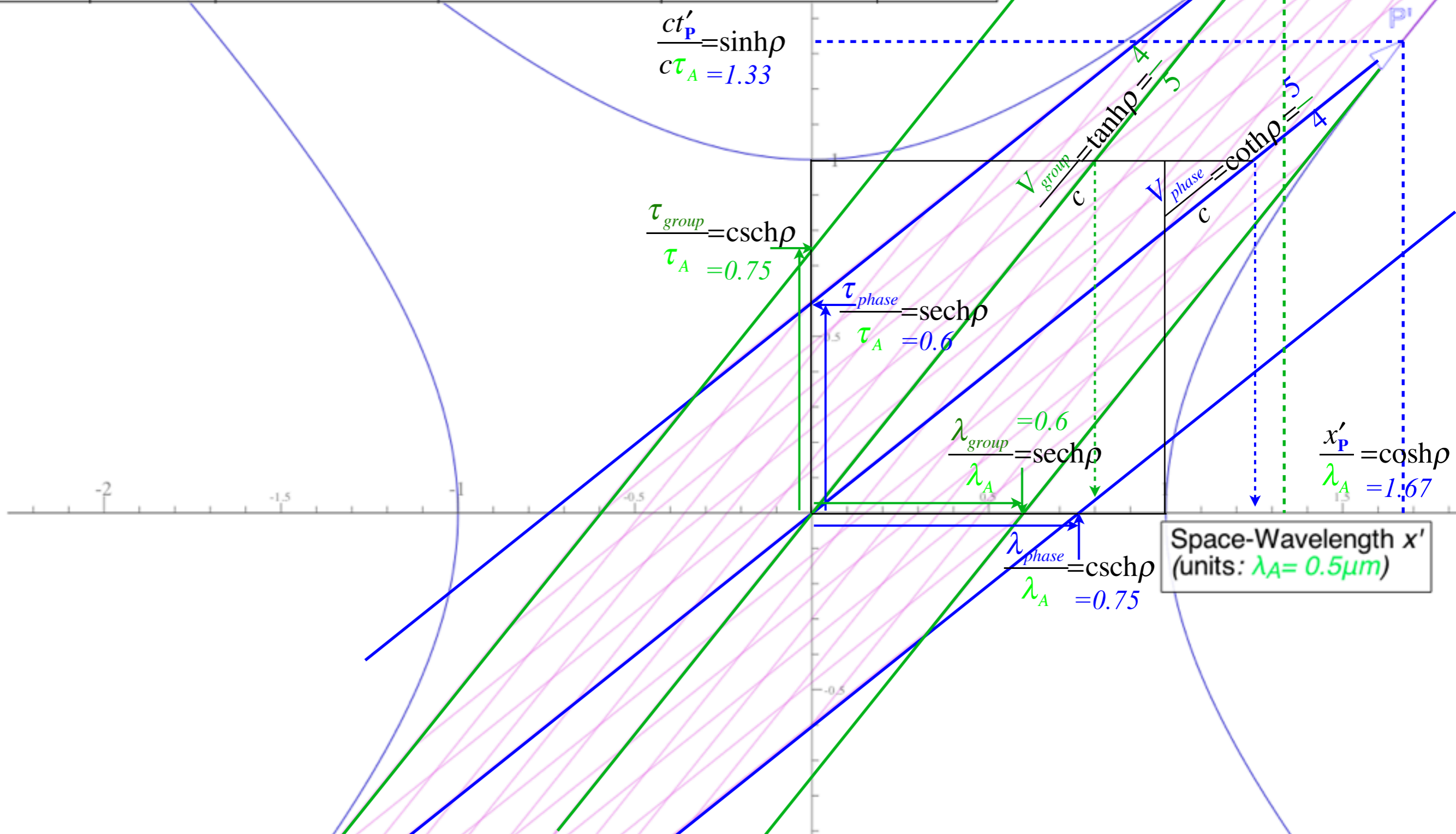
$$\tanh \rho = 0.618\dots = \frac{\sqrt{5}-1}{2}$$

Group velocity  
 $u/c = \sin(\sigma)$   
 $1 = \tanh(\rho)$

$u/c = 3/5 = 0.600$   
Golden ratio  $G_+ = (\sqrt{5}-1)/2 = 0.618\dots$   
 $\sqrt{2}/2 = 0.707\dots$   
 $\sqrt{G_-} = 0.7862\dots$



time	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
value for $\beta=0.80$	<b>0.33</b>	<b>0.80</b>	<b>1.34</b>	<b>0.60</b>	<b>1.67</b>	<b>0.75</b>	<b>1.25</b>	<b>3.01</b>



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➔ Animation of  $e^{\rho=2}$  spacetime and per-spacetime plots

*Rapidity  $\rho$  related to stellar aberration angle  $\sigma$  and L. C. Epstein's approach to relativity*

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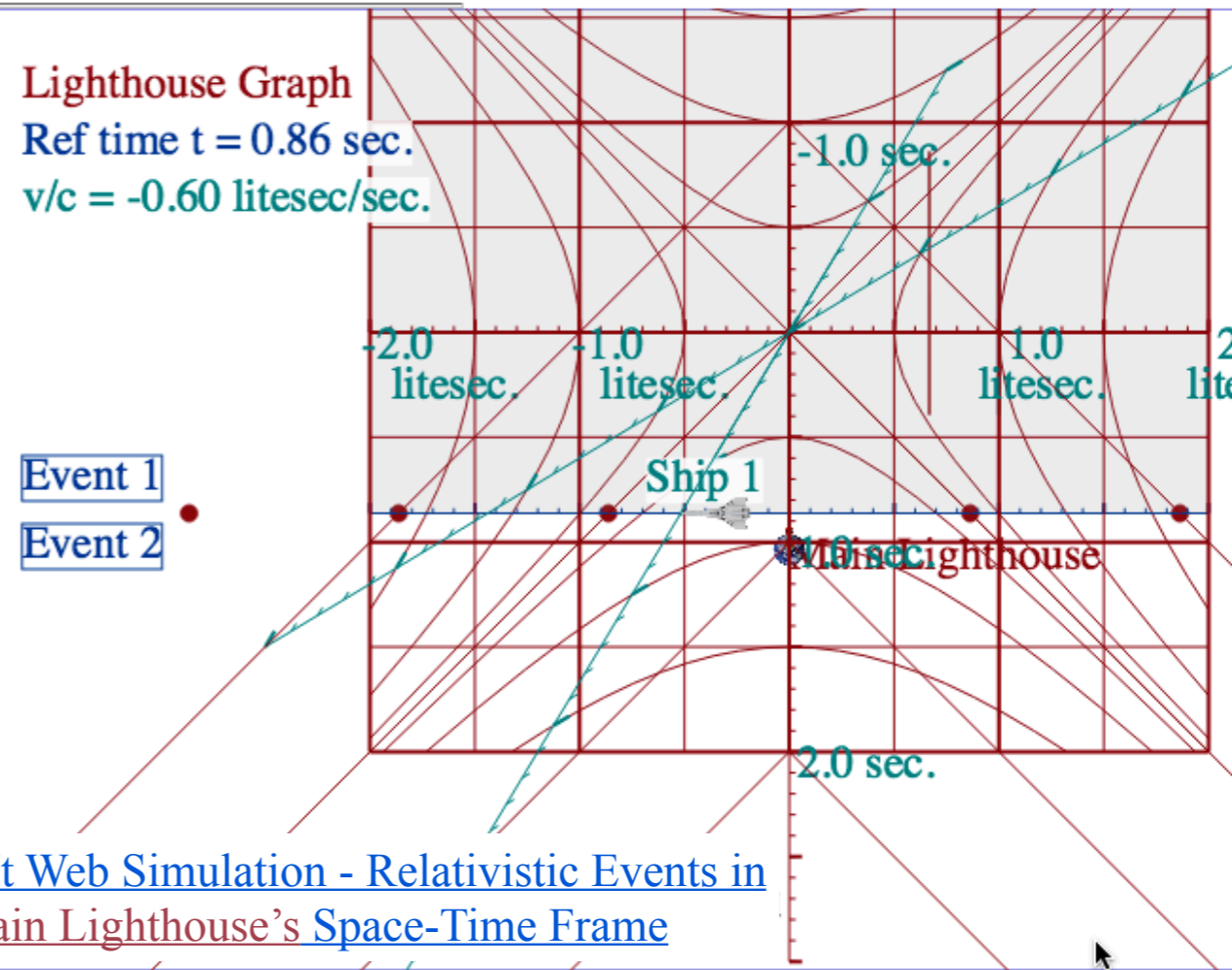
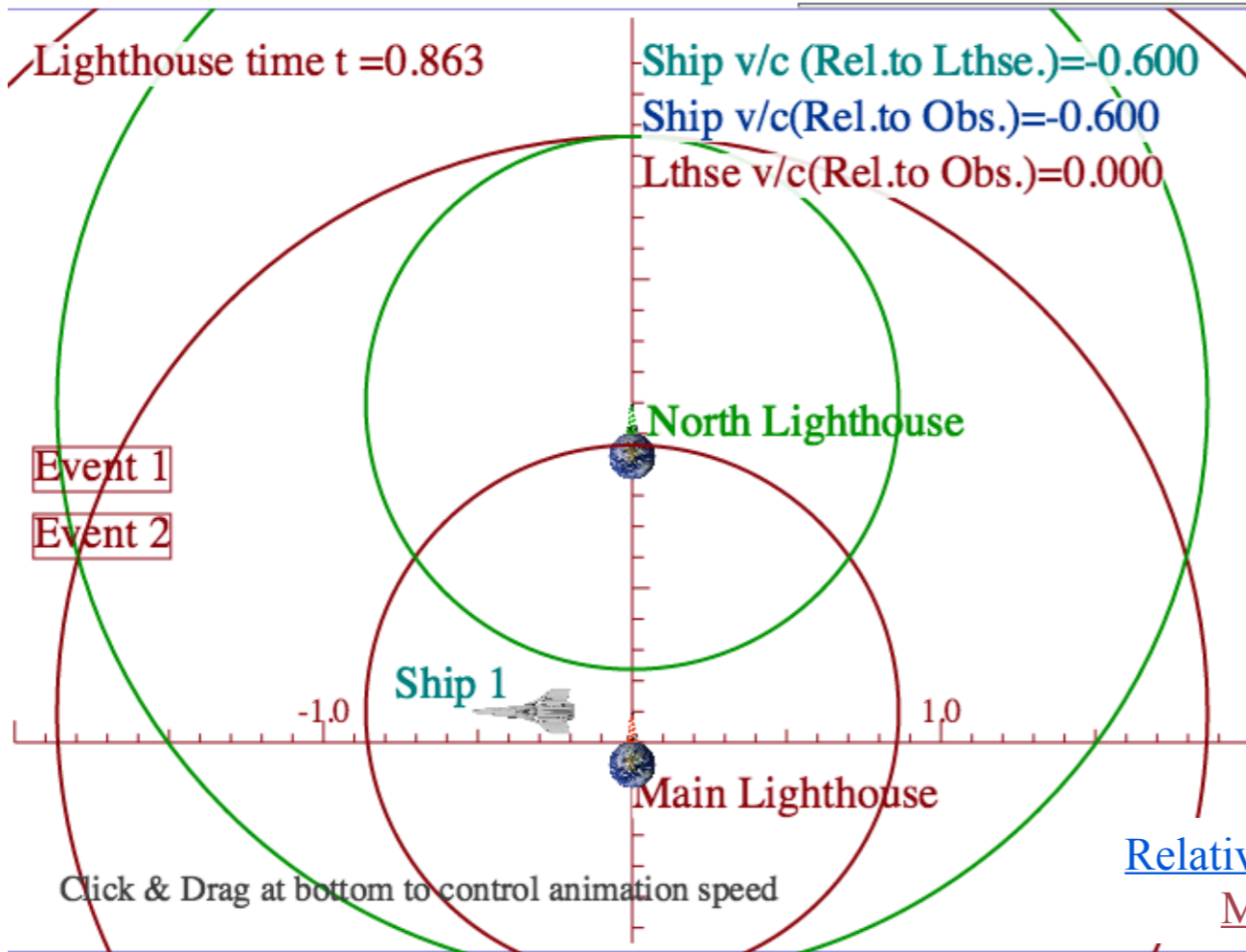
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*Relativity in accelerated frames*

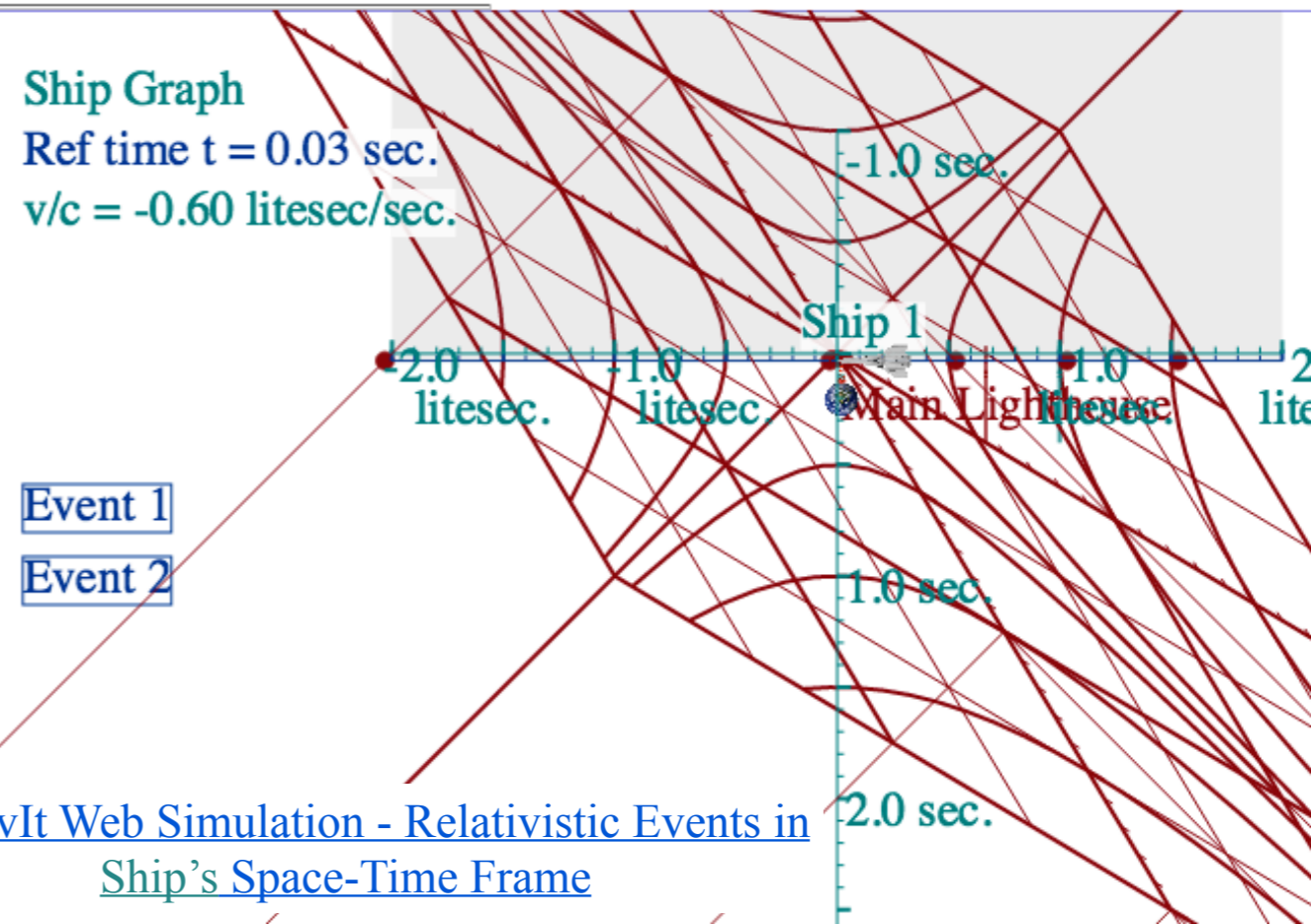
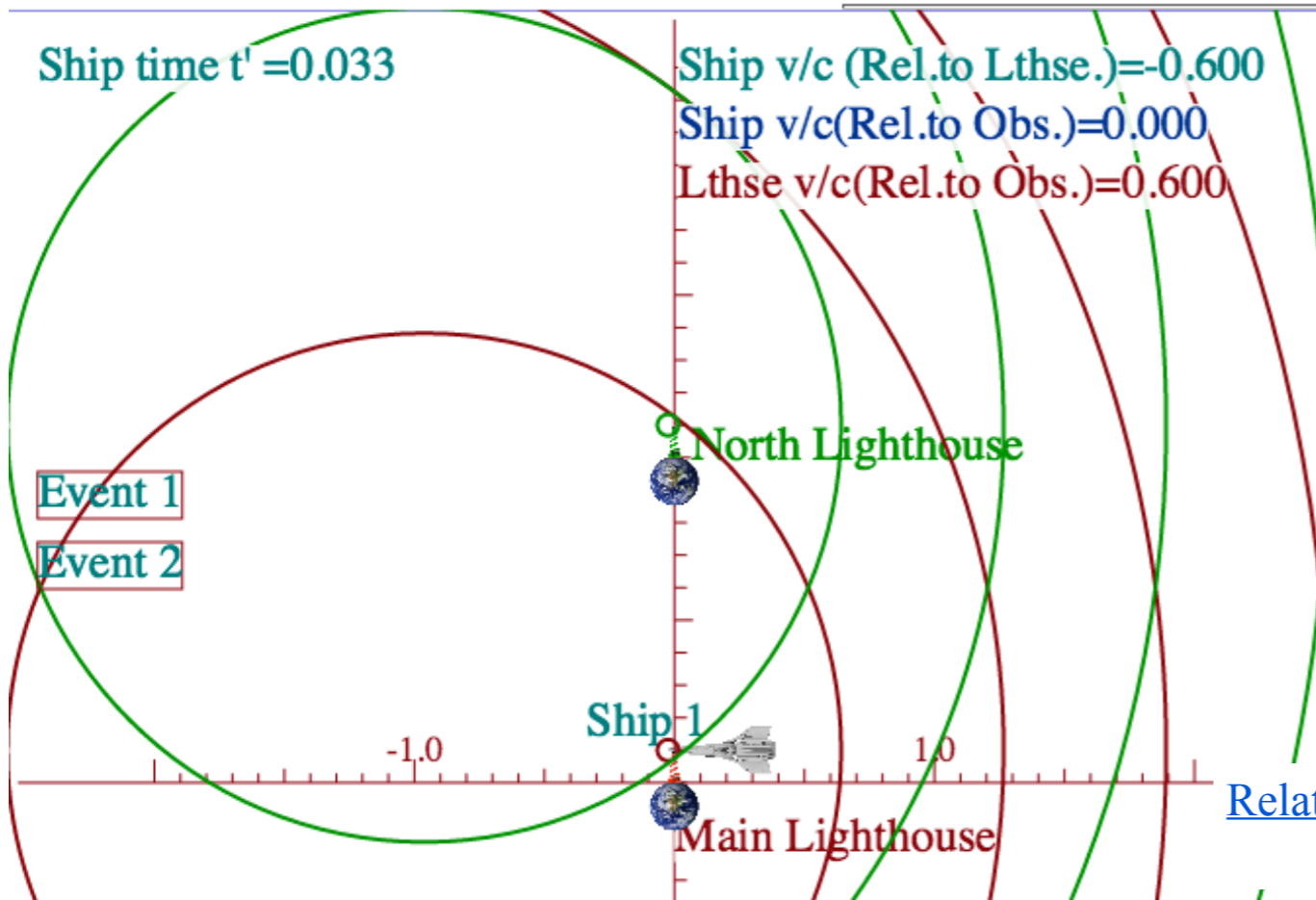
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RelativIt Web Simulation - Relativistic Events in Main Lighthouse's Space-Time Frame



RelativIt Web Simulation - Relativistic Events in Ship's Space-Time Frame

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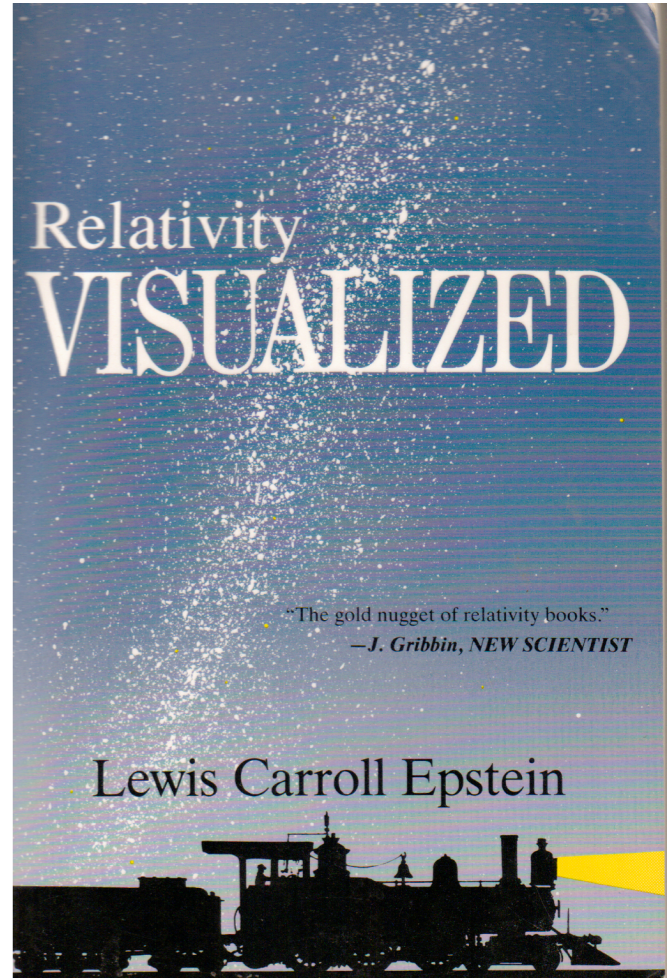
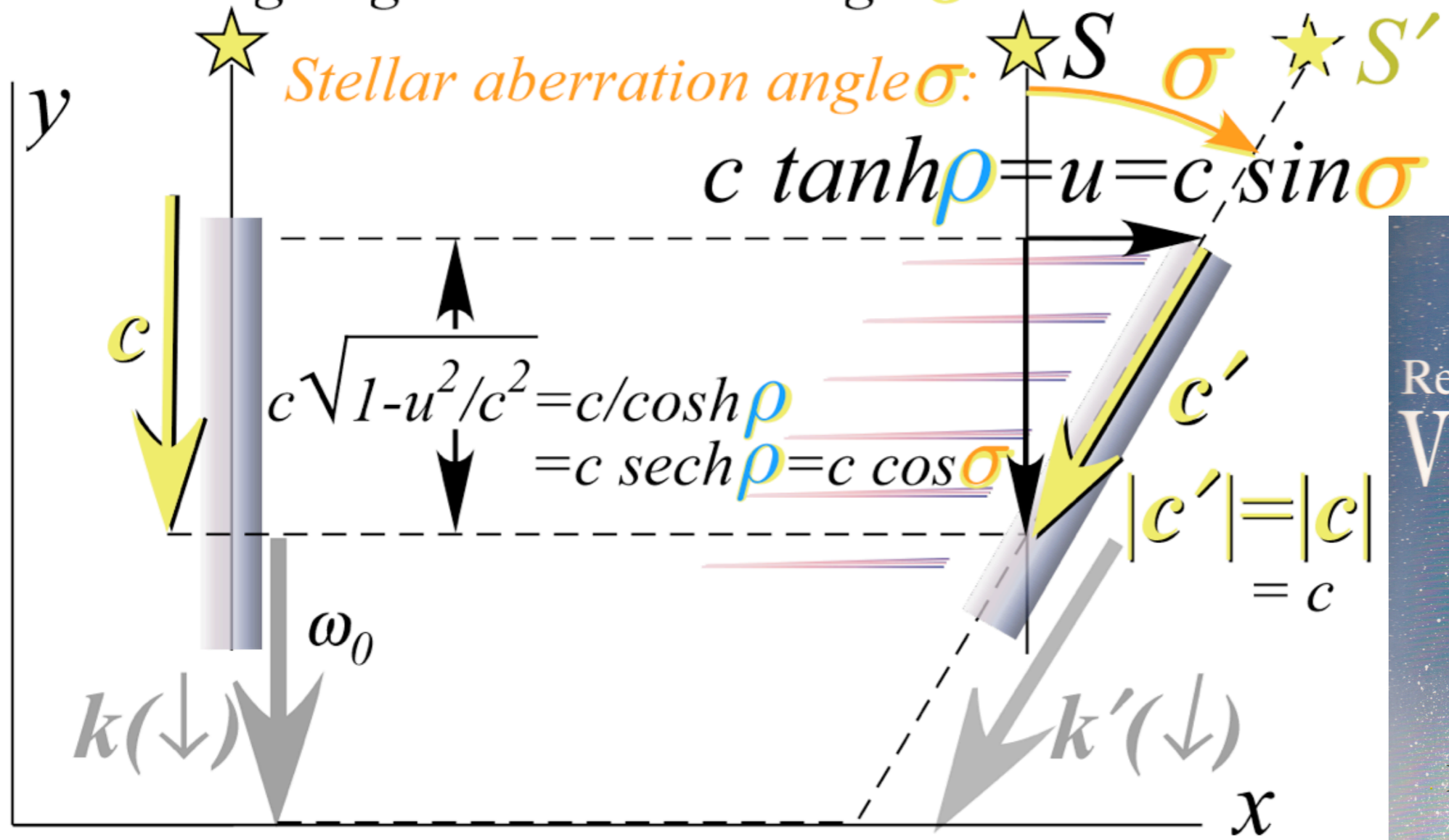
Comparing Longitudinal relativity parameter: Rapidity  $\rho = \log_e(\text{Doppler Shift})$

to a Transverse\*relativity parameter: Stellar aberration angle  $\sigma$

\*Lewis Carroll Epstein, *Relativitätstheorie*, Birkhäuser, (2004) Earlier English version (1985)-

Observer fixed below star sees it directly overhead.  
 Observer going  $u$  sees star at angle  $\sigma$  in  $u$  direction.

We used notion  $\sigma$  for stellar-ab-angle, (a “flipped-out”  $\rho$ ). Epstein not interested in  $\rho$  analysis or in relation of  $\sigma$  and  $\rho$ .



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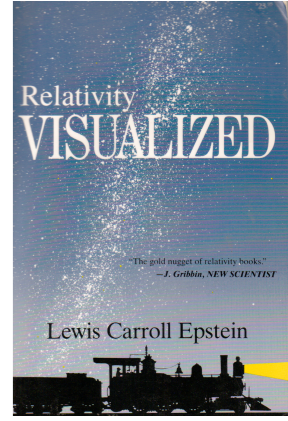
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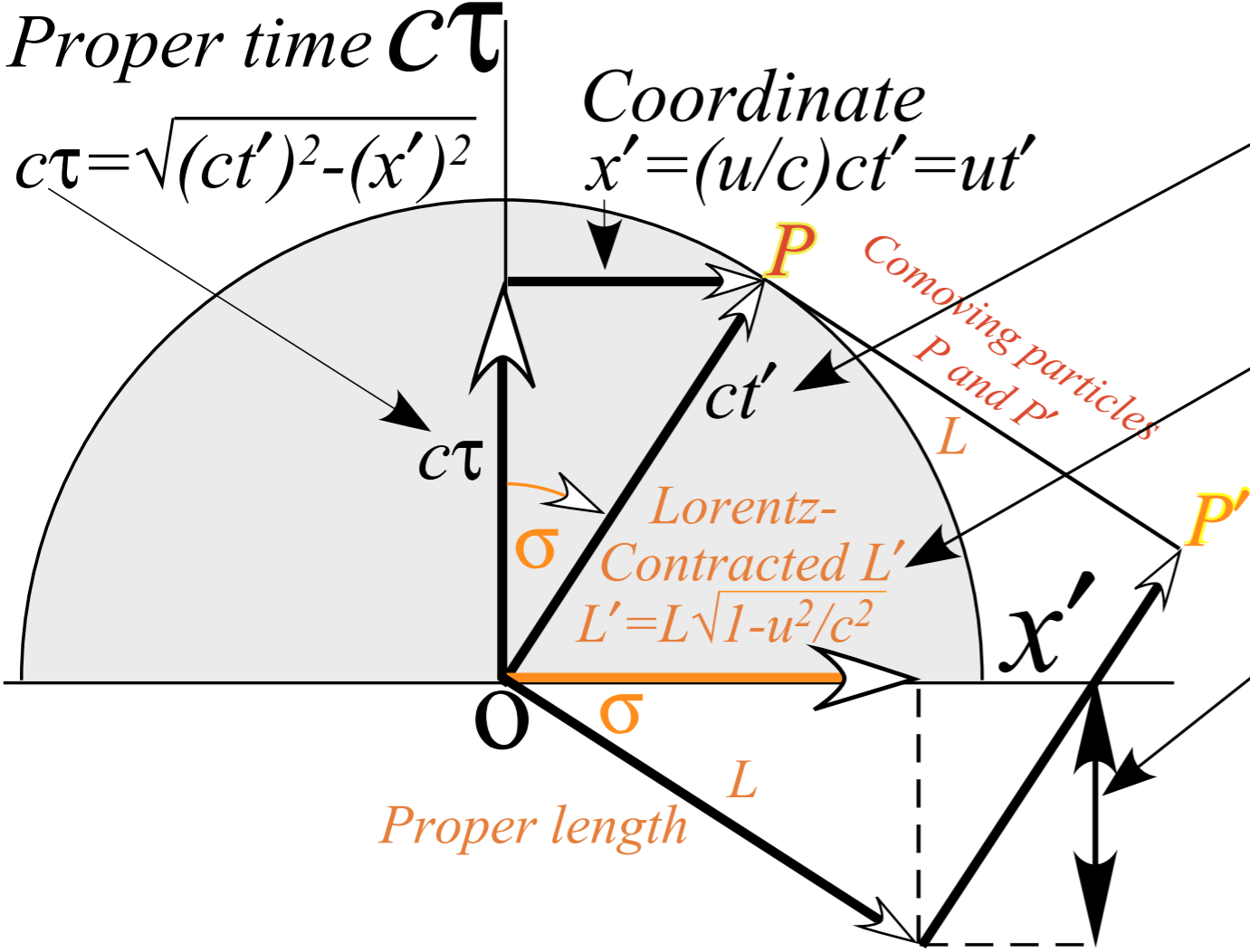
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Proper time  $c\tau$  vs. coordinate space  $x$  - (L. C. Epstein's "Cosmic Speedometer")

Particles  $P$  and  $P'$  have speed  $u$  in  $(x', ct')$  and speed  $c$  in  $(x, c\tau)$



Einstein time dilation:  
 $ct' = c\tau \sec\sigma = c\tau \cosh\rho = c\tau / \sqrt{1-u^2/c^2}$

Lorentz length contraction:  
 $L' = L \operatorname{sech}\rho = L \cos\sigma = L \cdot \sqrt{1-u^2/c^2}$

Proper Time asimultaneity:  
 $c \Delta\tau = L' \sinh\rho = L \cos\sigma \sinh\rho$   
 $= L \cos\sigma \tan\sigma$   
 $= L \sin\sigma = L / \sqrt{c^2/u^2 - 1} \sim L u/c$



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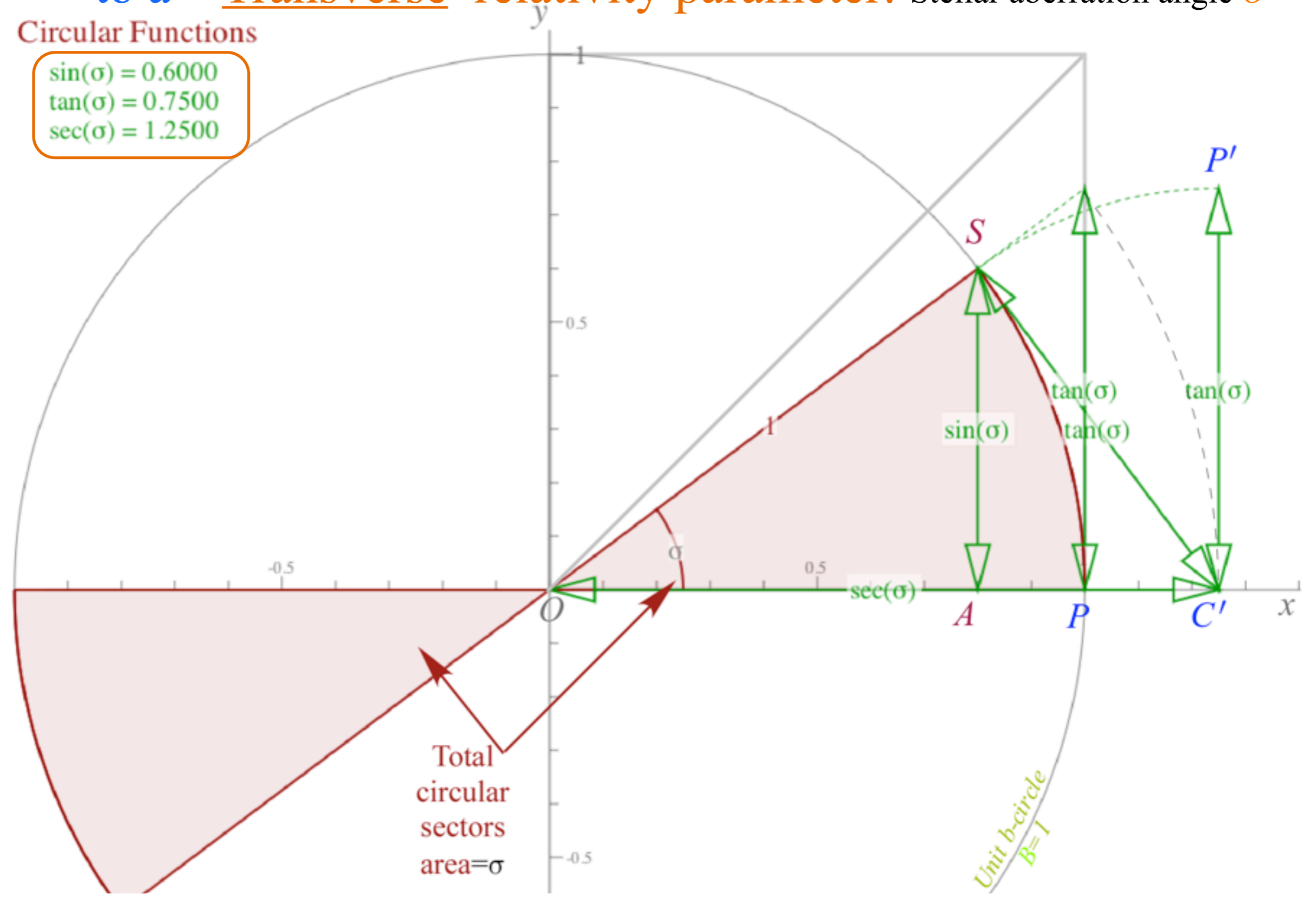


# Comparing Longitudinal relativity parameter: Rapidity $\rho = \log_e(\text{Doppler Shift})$

to a Transverse relativity parameter: Stellar aberration angle  $\sigma$

(a) Circular Functions

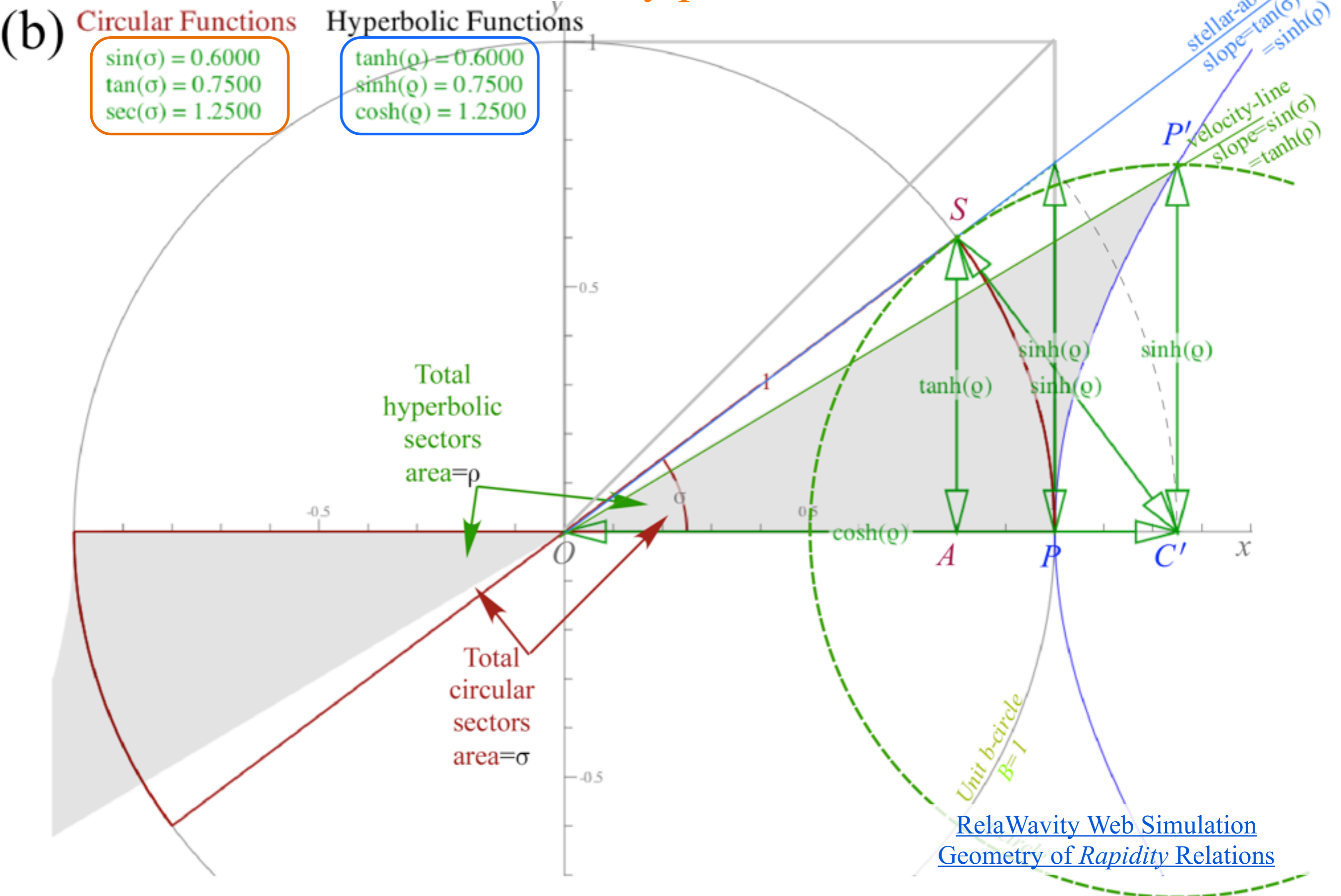
$\sin(\sigma) = 0.6000$   
 $\tan(\sigma) = 0.7500$   
 $\sec(\sigma) = 1.2500$



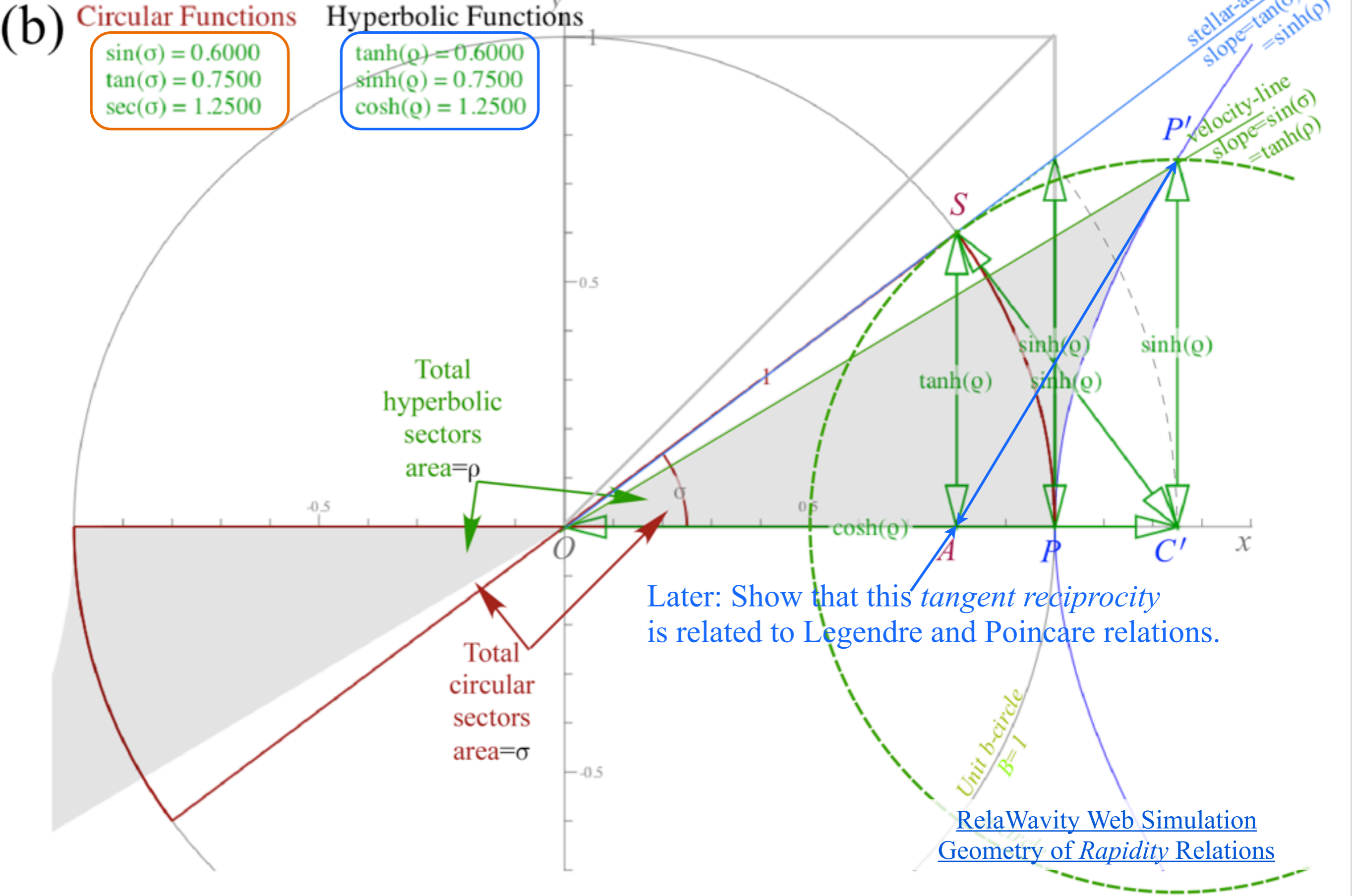
[RelaWavity Web Simulation](#)  
[Geometry of Stellar Aberration Angle](#)

# Comparing Longitudinal relativity parameter: Rapidity $\rho = \log_e(\text{Doppler Shift})$

to a Transverse relativity parameter: Stellar aberration angle  $\sigma$



# Comparing Longitudinal relativity parameter: Rapidity $\rho = \log_e(\text{Doppler Shift})$ to a Transverse relativity parameter: Stellar aberration angle $\sigma$



### Circular Functions

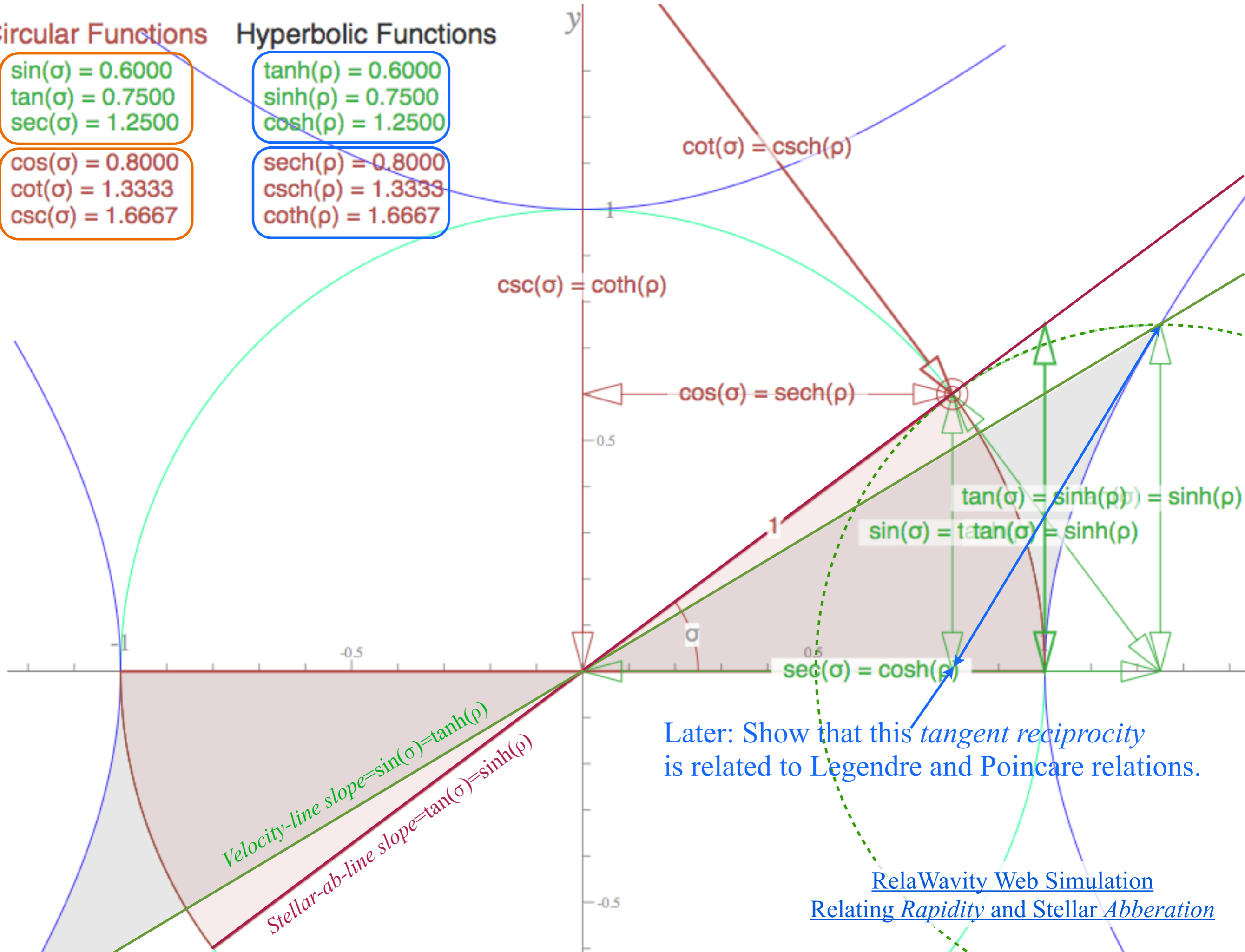
### Hyperbolic Functions

$$\begin{aligned} \sin(\sigma) &= 0.6000 \\ \tan(\sigma) &= 0.7500 \\ \sec(\sigma) &= 1.2500 \end{aligned}$$

$$\begin{aligned} \tanh(\rho) &= 0.6000 \\ \sinh(\rho) &= 0.7500 \\ \cosh(\rho) &= 1.2500 \end{aligned}$$

$$\begin{aligned} \cos(\sigma) &= 0.8000 \\ \cot(\sigma) &= 1.3333 \\ \csc(\sigma) &= 1.6667 \end{aligned}$$

$$\begin{aligned} \operatorname{sech}(\rho) &= 0.8000 \\ \operatorname{csch}(\rho) &= 1.3333 \\ \operatorname{coth}(\rho) &= 1.6667 \end{aligned}$$



Later: Show that this *tangent reciprocity* is related to Legendre and Poincare relations.

[RelaWavity Web Simulation](#)  
[Relating Rapidity and Stellar Abberation](#)

# Lecture 31

## Thur. 12.10.2015

Review of 16 relativity functions of  $\rho$  and related geometric approach to relativity  
Animation of  $e^{\rho=2}$  spacetime and per-spacetime plots

*Rapidity*  $\rho$  related to *stellar aberration angle*  $\sigma$  and L. C. Epstein's approach to relativity

Longitudinal hyperbolic  $\rho$ -geometry connects to transverse circular  $\sigma$ -geometry

➔ “Occams Sword” and summary of 16 parameter functions of  $\rho$  and  $\sigma$   
Applications to optical waveguide, spherical waves, and accelerator radiation

*Learning about sin! and COS and...*

Derivation of relativistic quantum mechanics

What's the matter with mass? Shining some light on the Elephant in the room

Relativistic action and Lagrangian-Hamiltonian relations

Poincare' and Hamilton-Jacobi equations

Relativistic optical transitions and Compton recoil formulae

Feynman diagram geometry

Compton recoil related to rocket velocity formula

Comparing 2<sup>nd</sup>-quantization “photon” number  $N$  and 1<sup>st</sup>-quantization wavenumber  $\kappa$

*Relativity* in accelerated frames

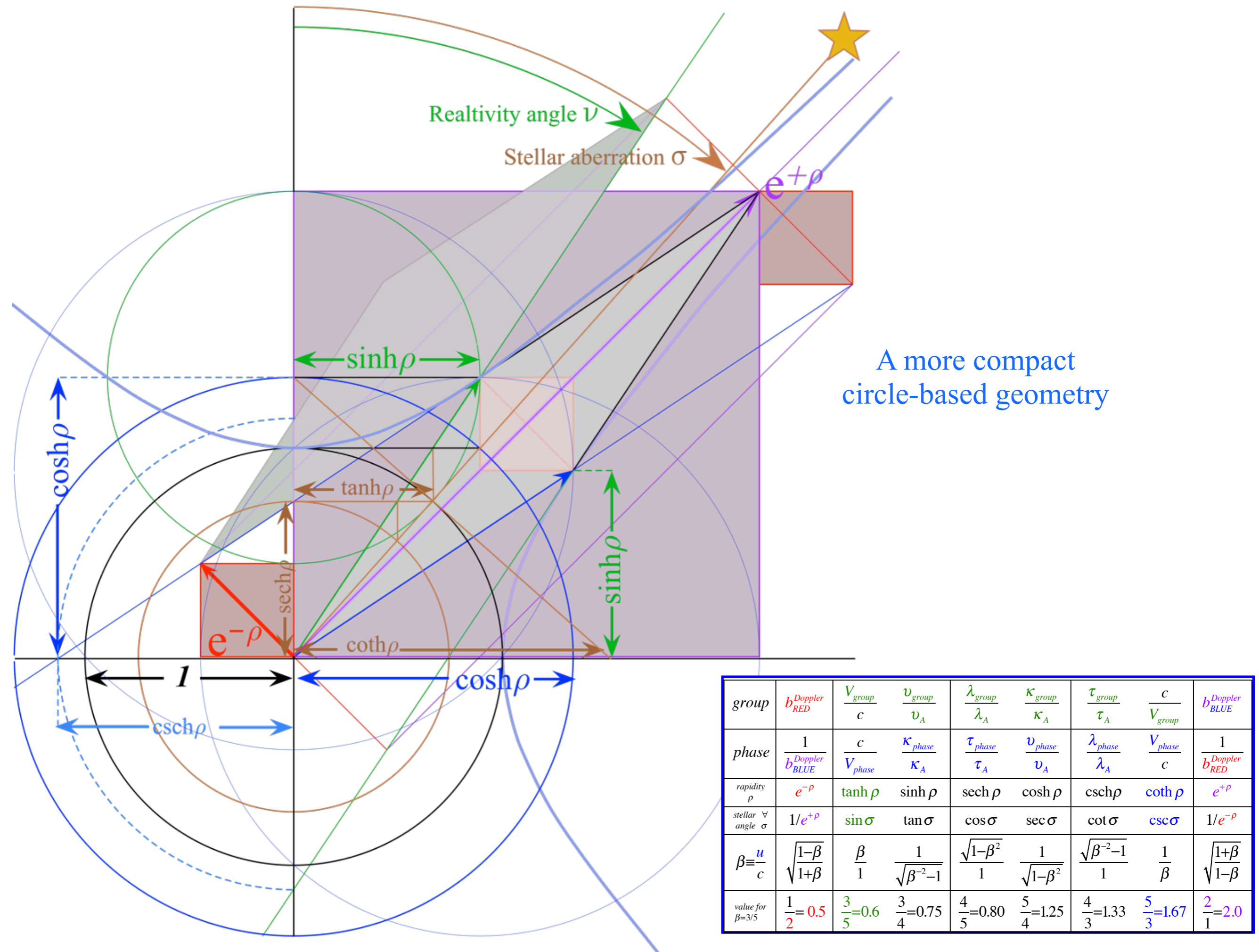
Laser up-tuning by Alice and down-tuning by Carla makes  $g$ -acceleration grid

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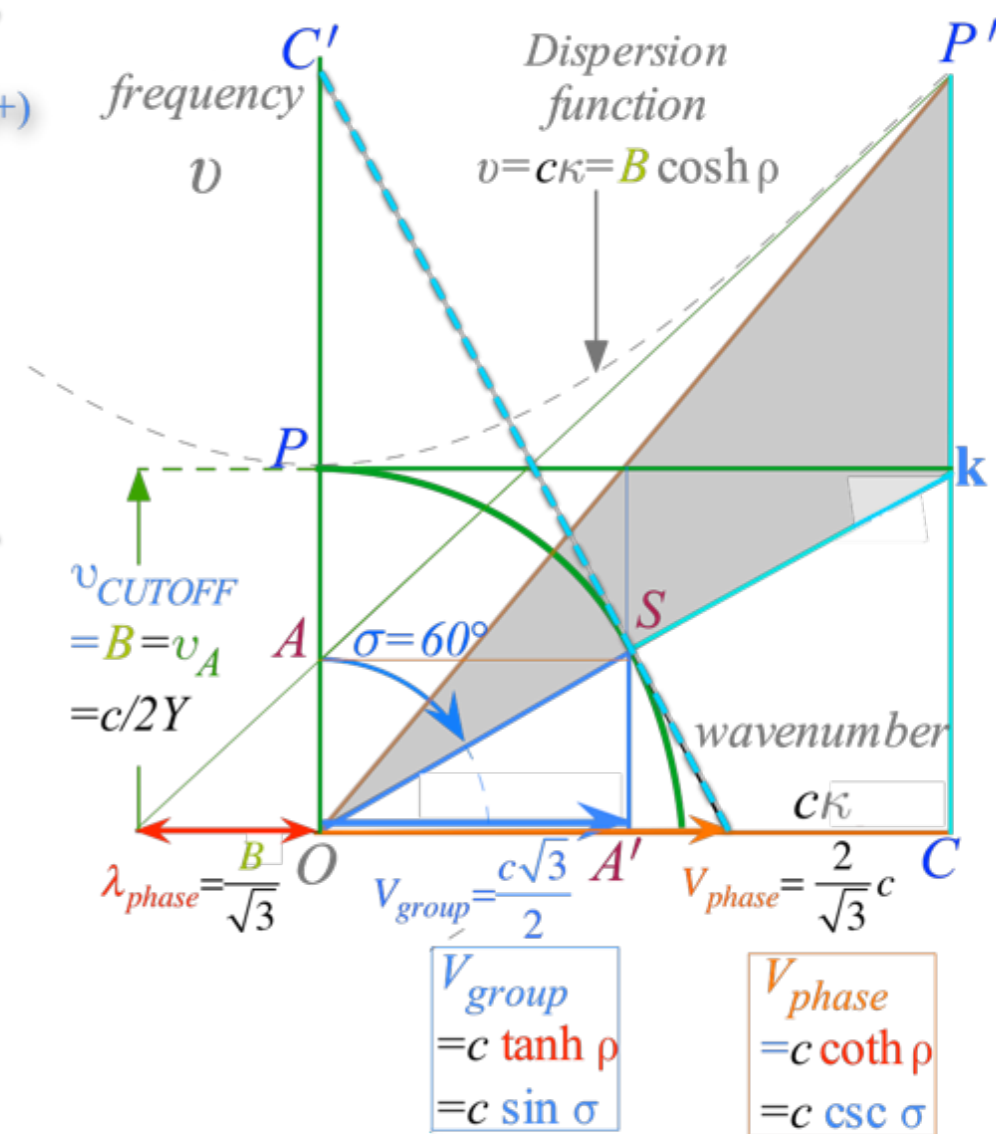
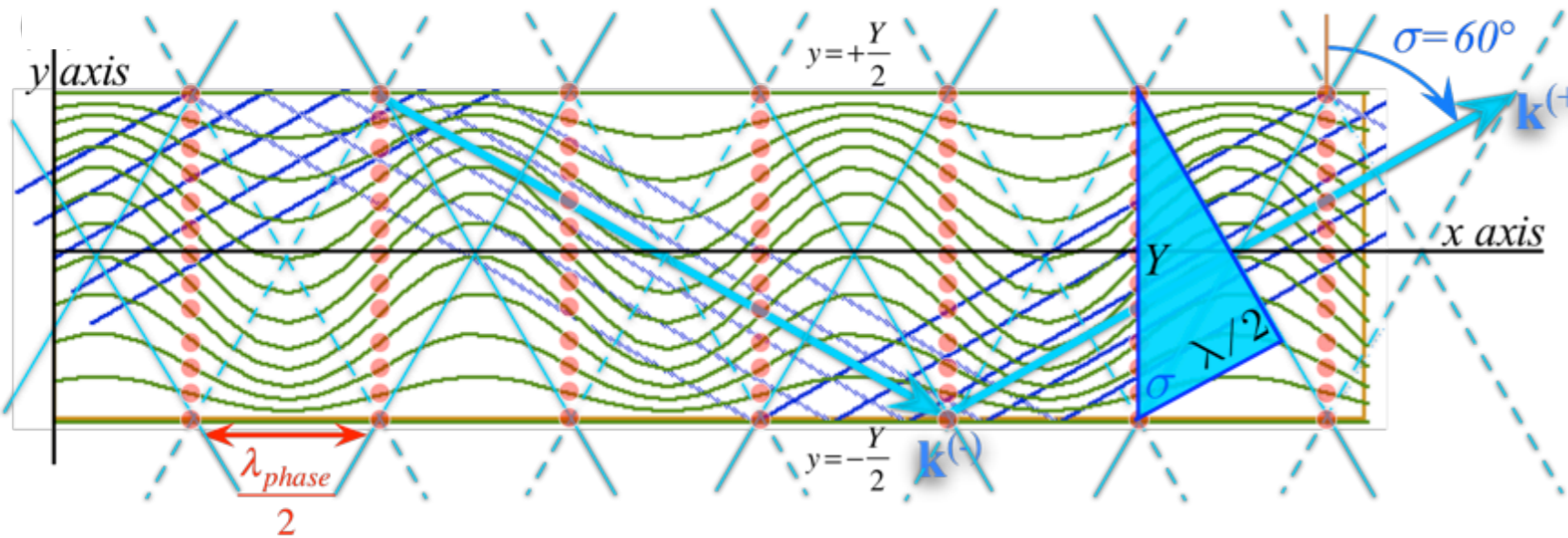
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# Optical wave guide relativistic geometry aided by Occam's Sword

geometry applies to  $(x,y)$  space-space  
 to  $(k_x, k_y)$  per-space-per-space  
 to  $(x, ct)$  space-time

Relativistic mode with near-c  $V_{group}=c/2$  and  $V_{phase}=2c$ . (Low dispersion.)

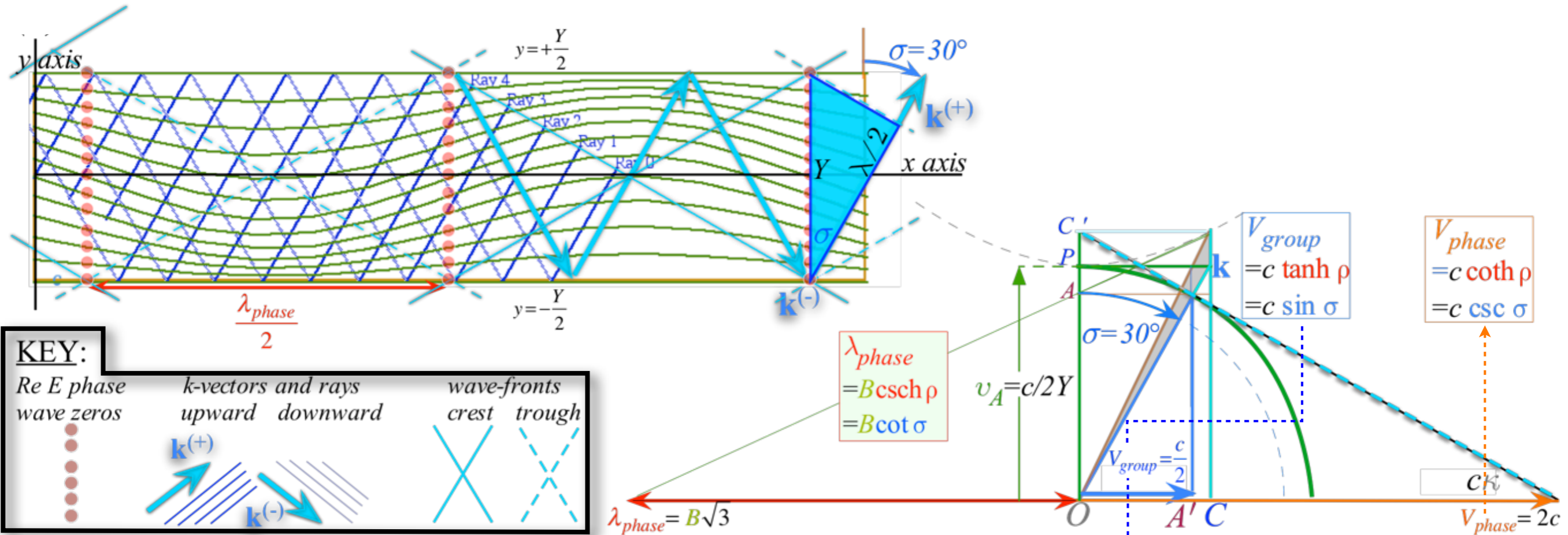


**KEY:**

Re E phase wave zeros 	k-vectors and rays upward downward 	wave-fronts crest trough 
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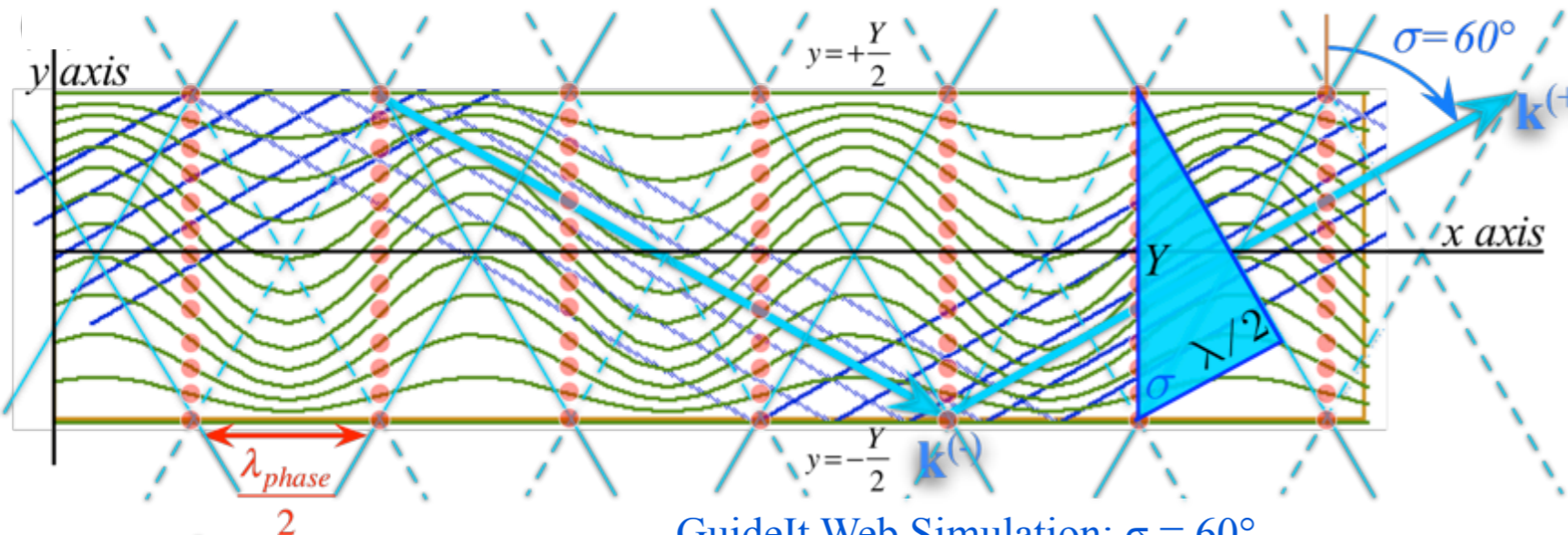


Example of near-cut-off mode with low  $V_{group} = c/2$  and high  $V_{phase} = 2c$ . (High dispersion.)

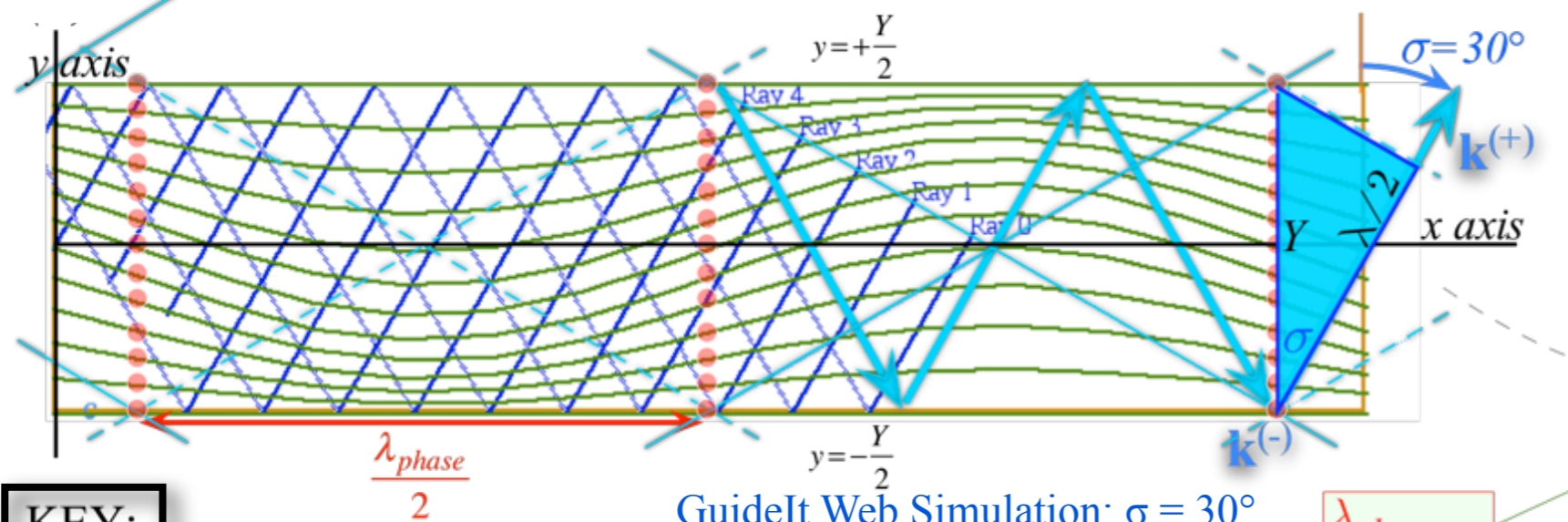
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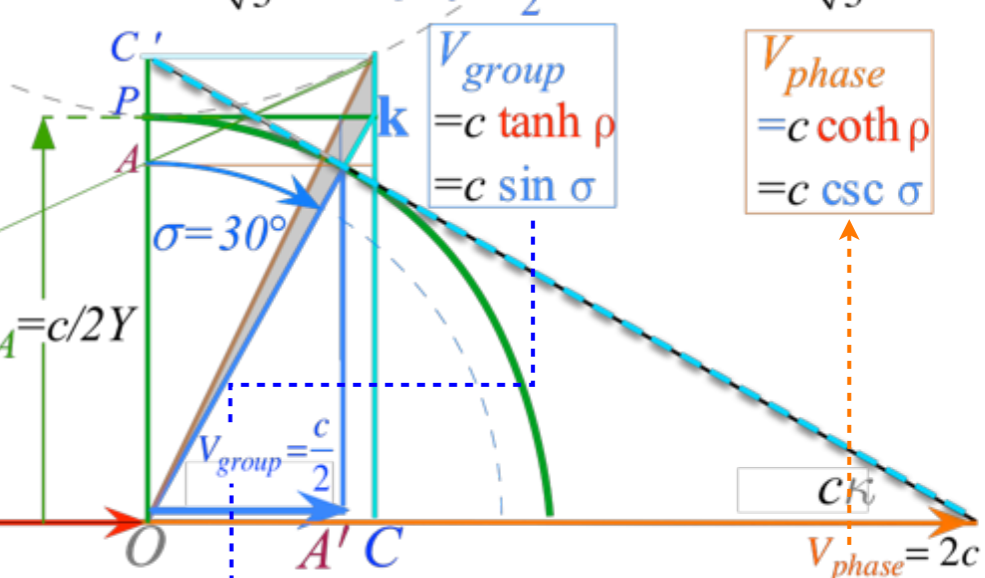
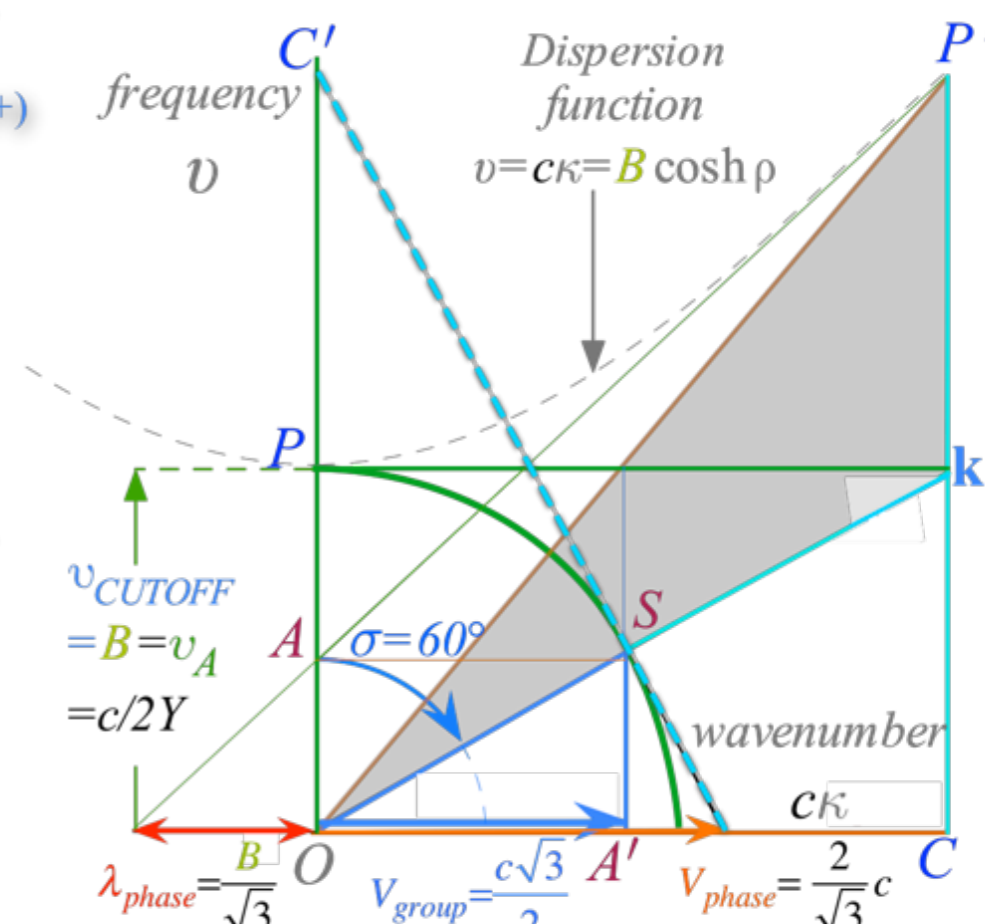
GuideIt Web Simulation:  $\sigma = 60^\circ$



GuideIt Web Simulation:  $\sigma = 30^\circ$

**KEY:**

Re E phase wave zeros	k-vectors and rays upward downward	wave-fronts crest trough



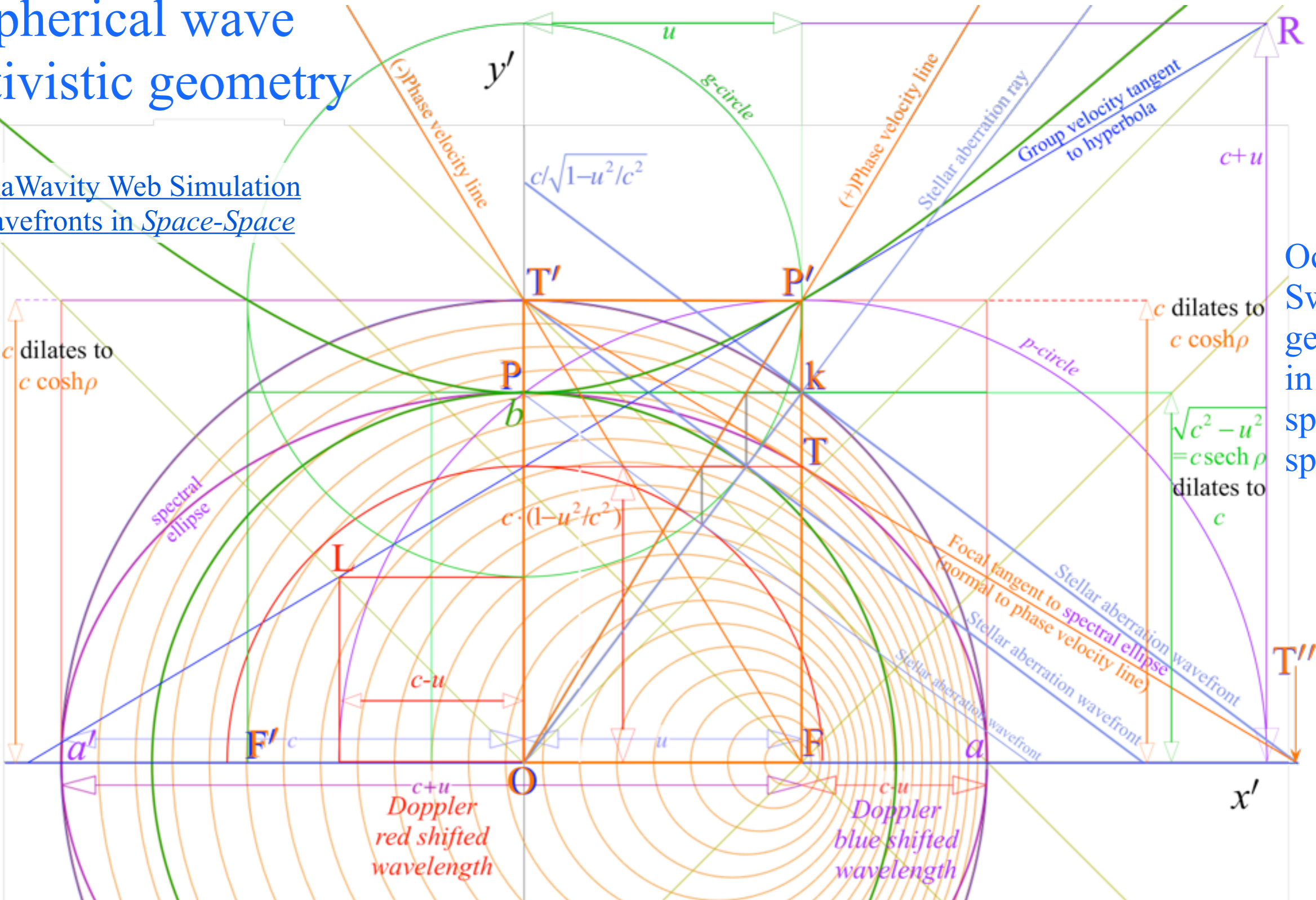
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# Spherical wave relativistic geometry

RelaWavity Web Simulation  
Wavefronts in Space-Space



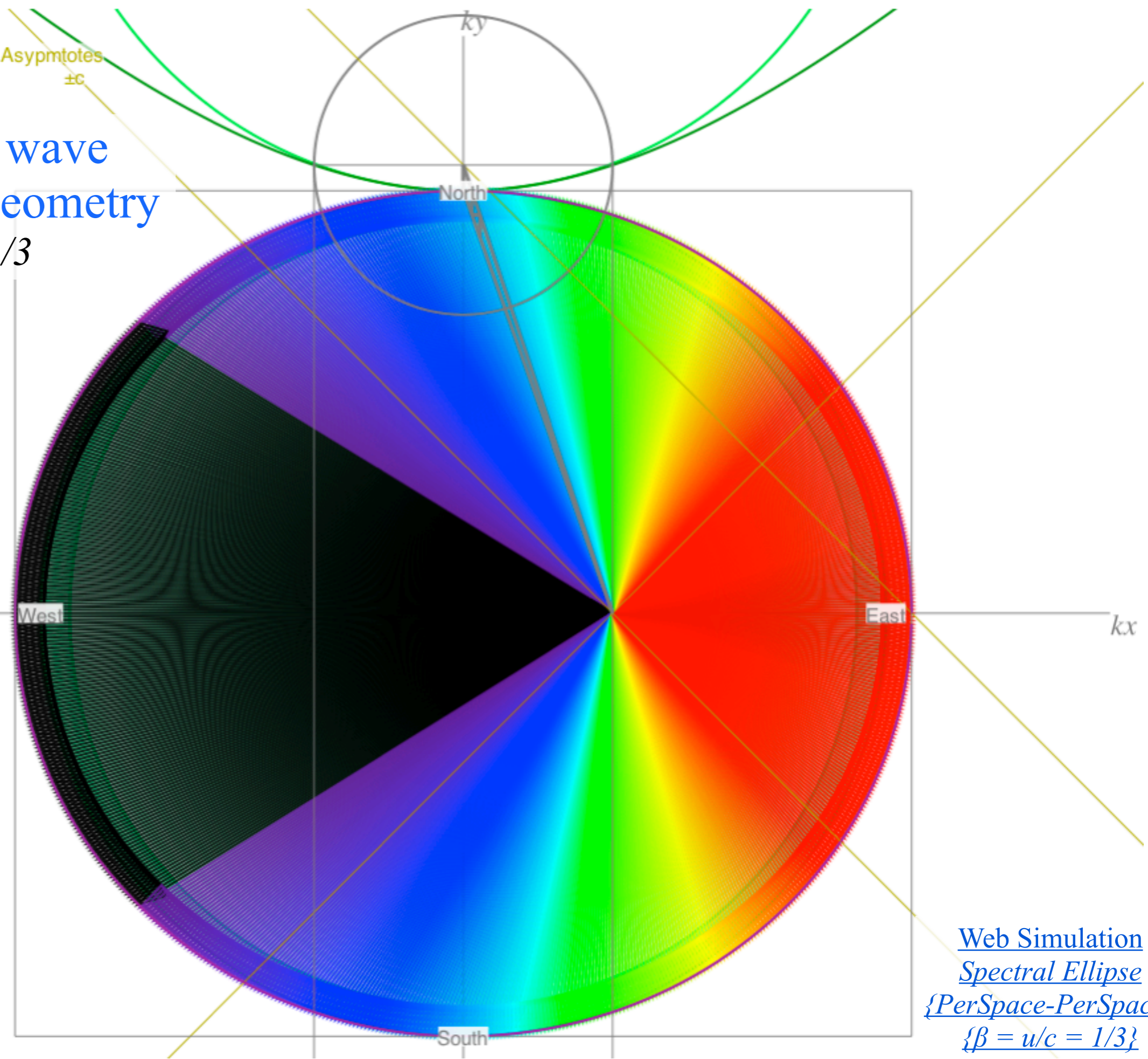
Occam  
Sword  
geometry  
in (x,y)  
space-  
space

<p>Doppler Red <math>\lambda=c+u</math> dilates to: <math>(c+u) \cosh \rho = c \sqrt{\frac{c+u}{c-u}} = ce^{+\rho}</math></p>	<p>ellipse focal length <math>FO = u = c \tanh \rho</math> dilates to: <math>u \cosh \rho = c \sinh \rho</math></p>	<p>Doppler Blue <math>\lambda=c-u</math> dilates to: <math>(c-u) \cosh \rho = c \sqrt{\frac{c-u}{c+u}} = ce^{-\rho}</math></p>
<p>ellipse major radius <math>a=OFa=c</math> dilates to: <math>c \cosh \rho = c/\sqrt{1-u^2/c^2}</math></p>	<p>ellipse latus radius <math>FT=c(1-u^2/c^2)</math> dilates to: <math>c(1-u^2/c^2) \cosh \rho = c\sqrt{1-u^2/c^2} = c \operatorname{sech} \rho</math></p>	<p>Base height <math>FTk = \sqrt{c^2 - u^2}</math> dilates to: <math>\sqrt{c^2 - u^2} \cosh \rho = c</math> (equal to ellipse minor radius <math>b</math>)</p>

Applications of Einstein dilation factor:  
 $\gamma = \cosh \rho = 1/\sqrt{1-u^2/c^2}$

Spherical wave  
relativistic geometry

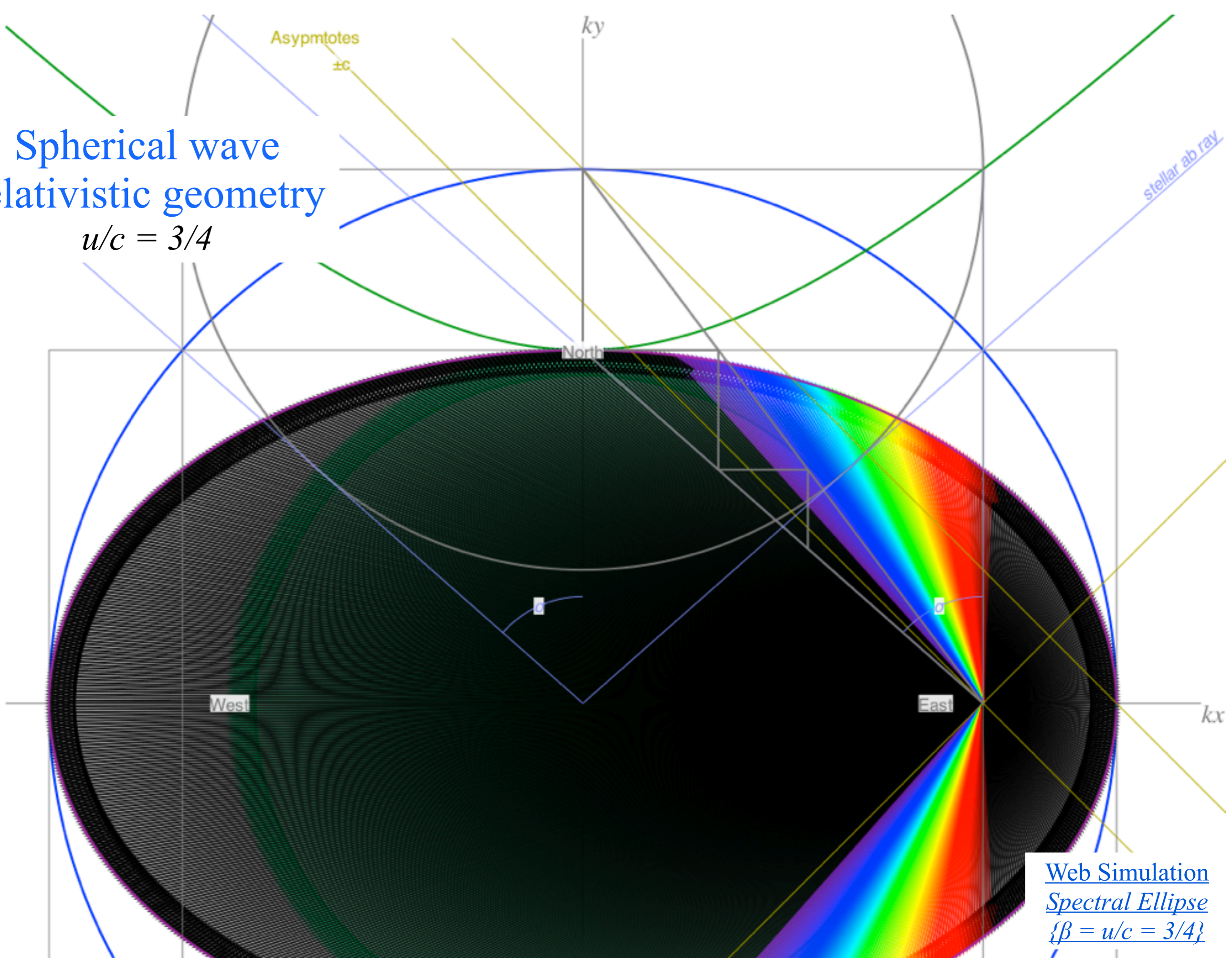
$u/c = 1/3$



[Web Simulation](#)  
[Spectral Ellipse](#)  
[{PerSpace-PerSpace}](#)  
[{ \$\beta = u/c = 1/3\$ }](#)



Spherical wave  
relativistic geometry  
 $u/c = 3/4$



Web Simulation  
Spectral Ellipse  
 $\{\beta = u/c = 3/4\}$

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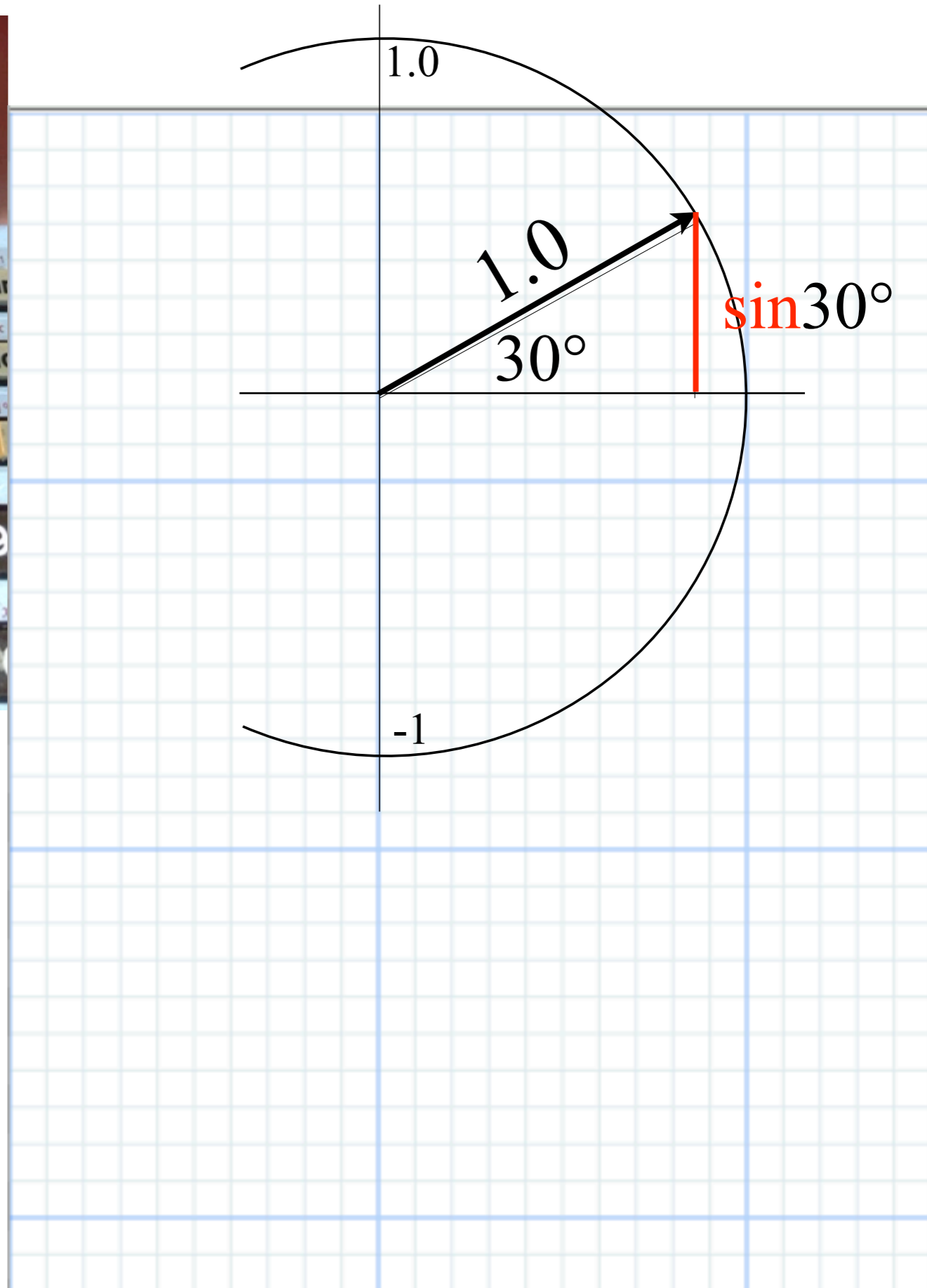
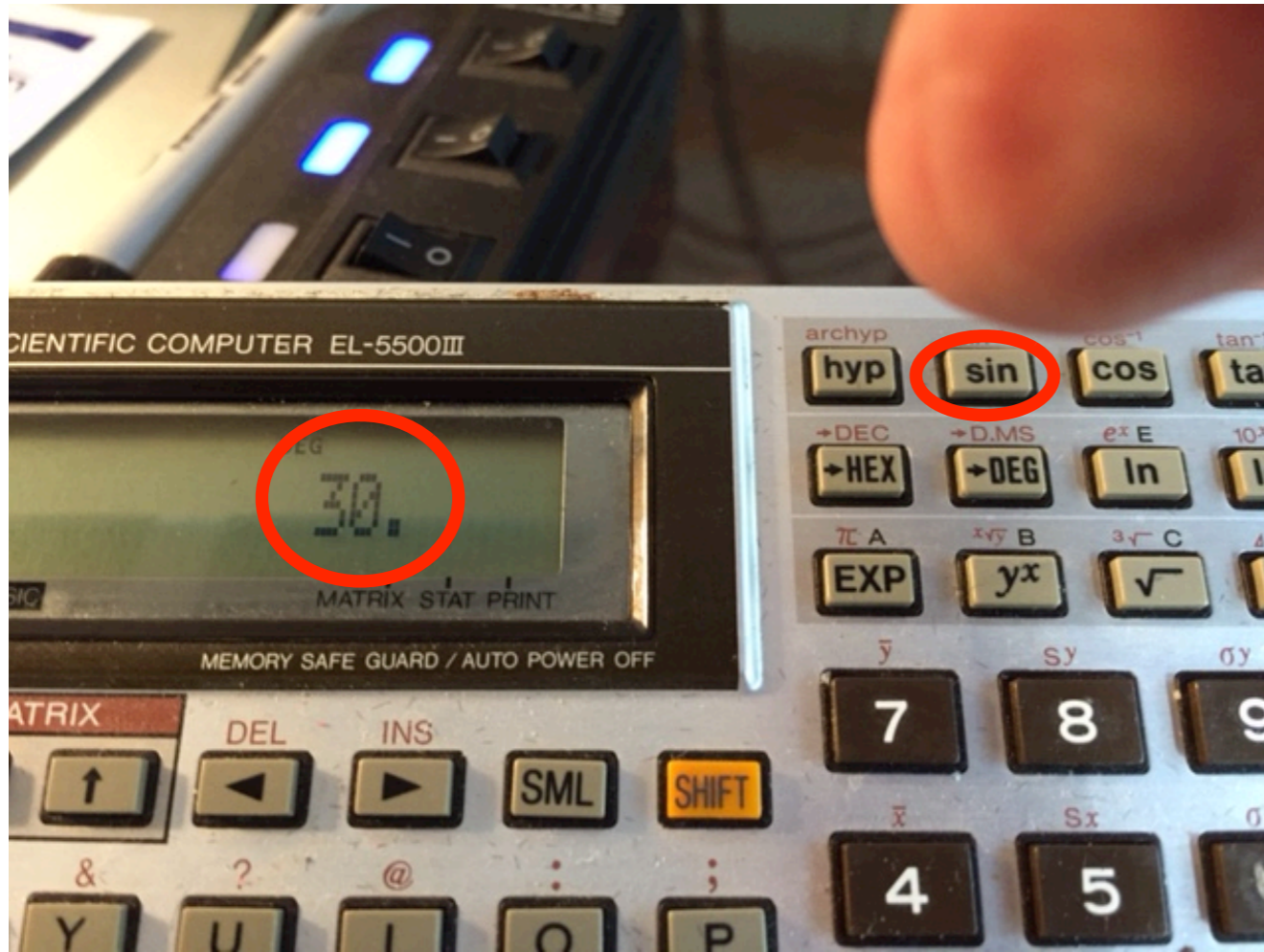
*Relativity* in accelerated frames

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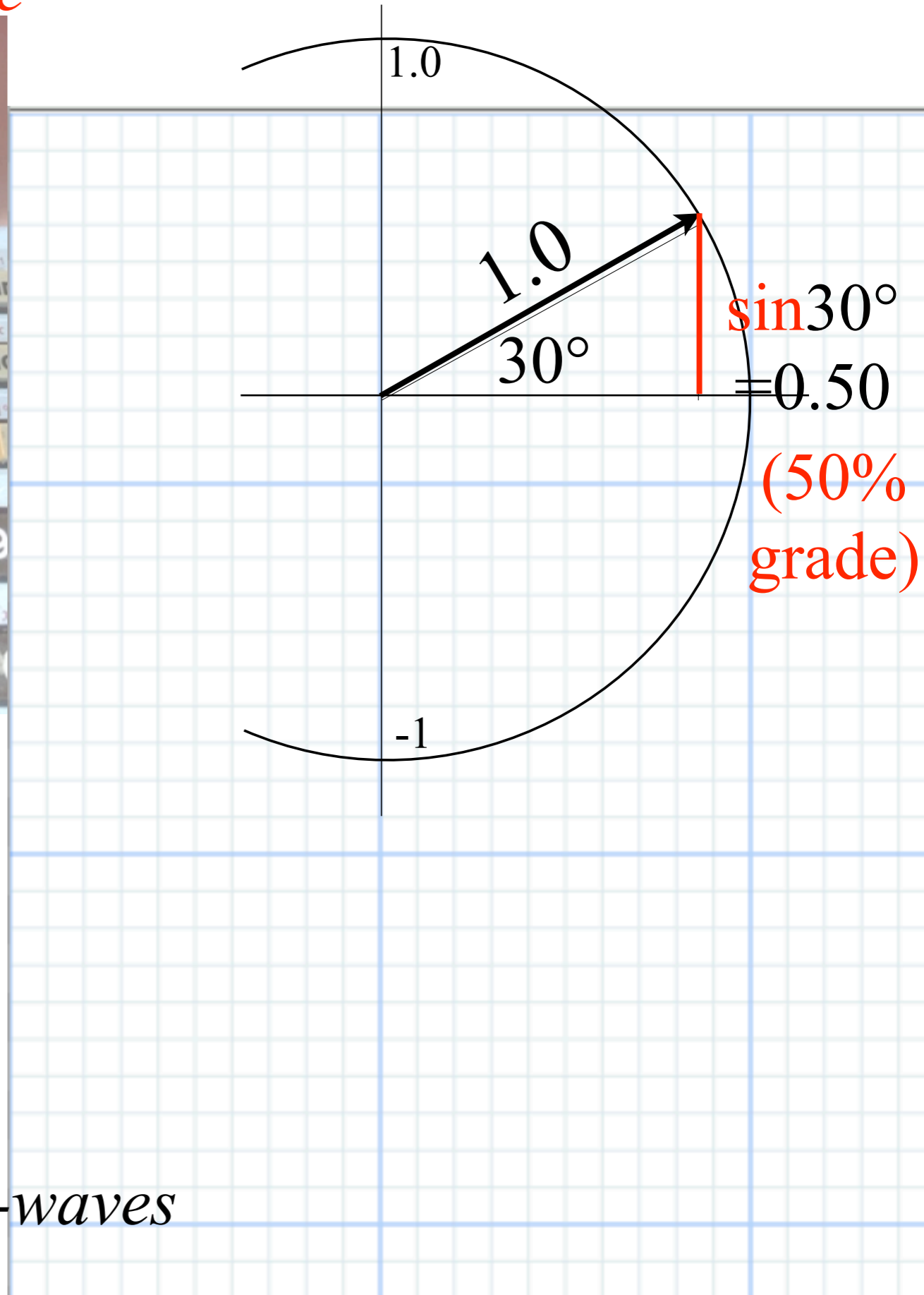
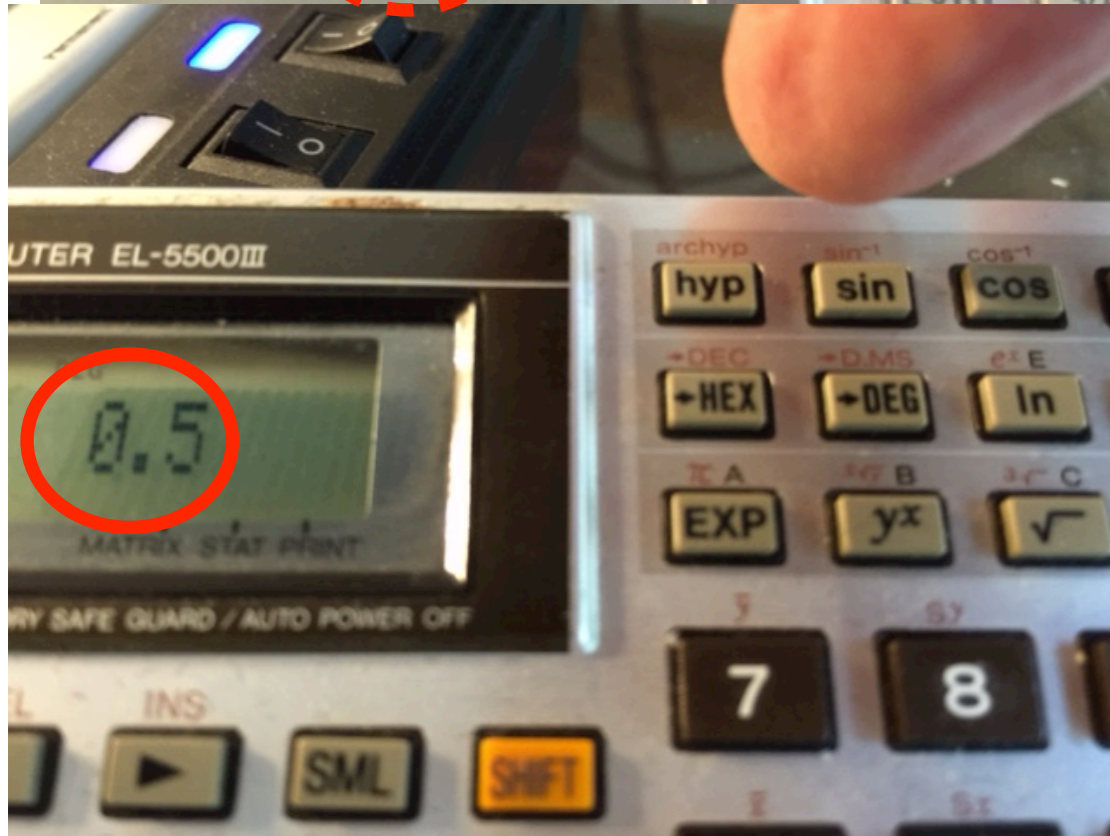
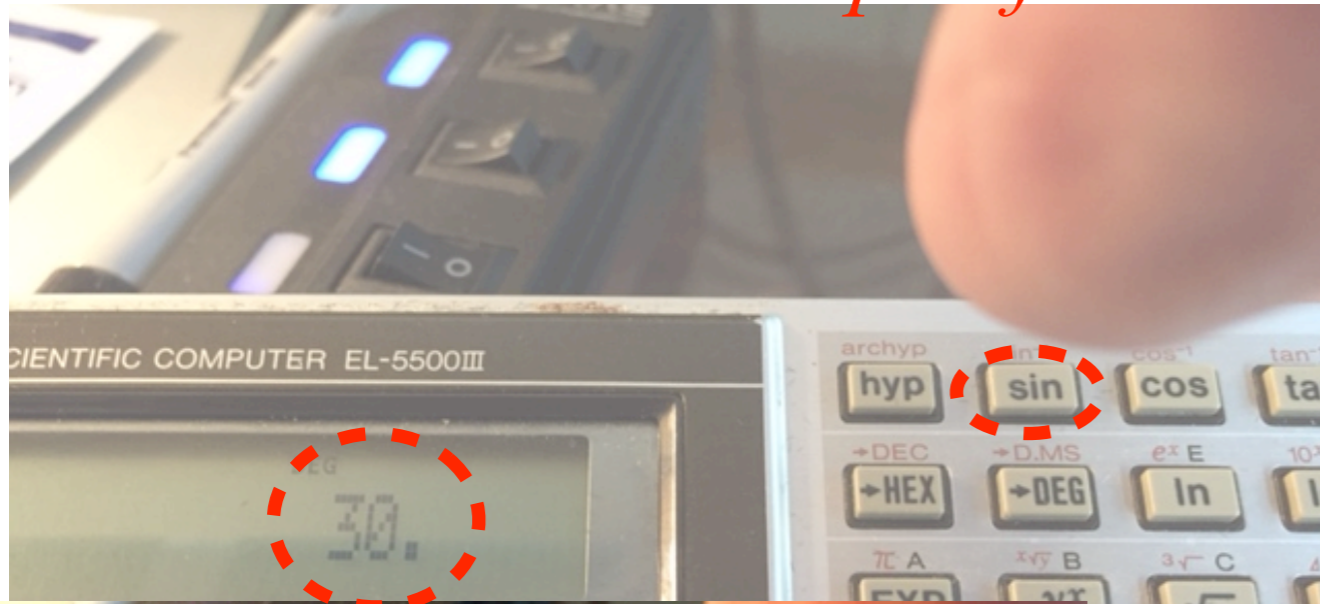
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# Learning about SIN

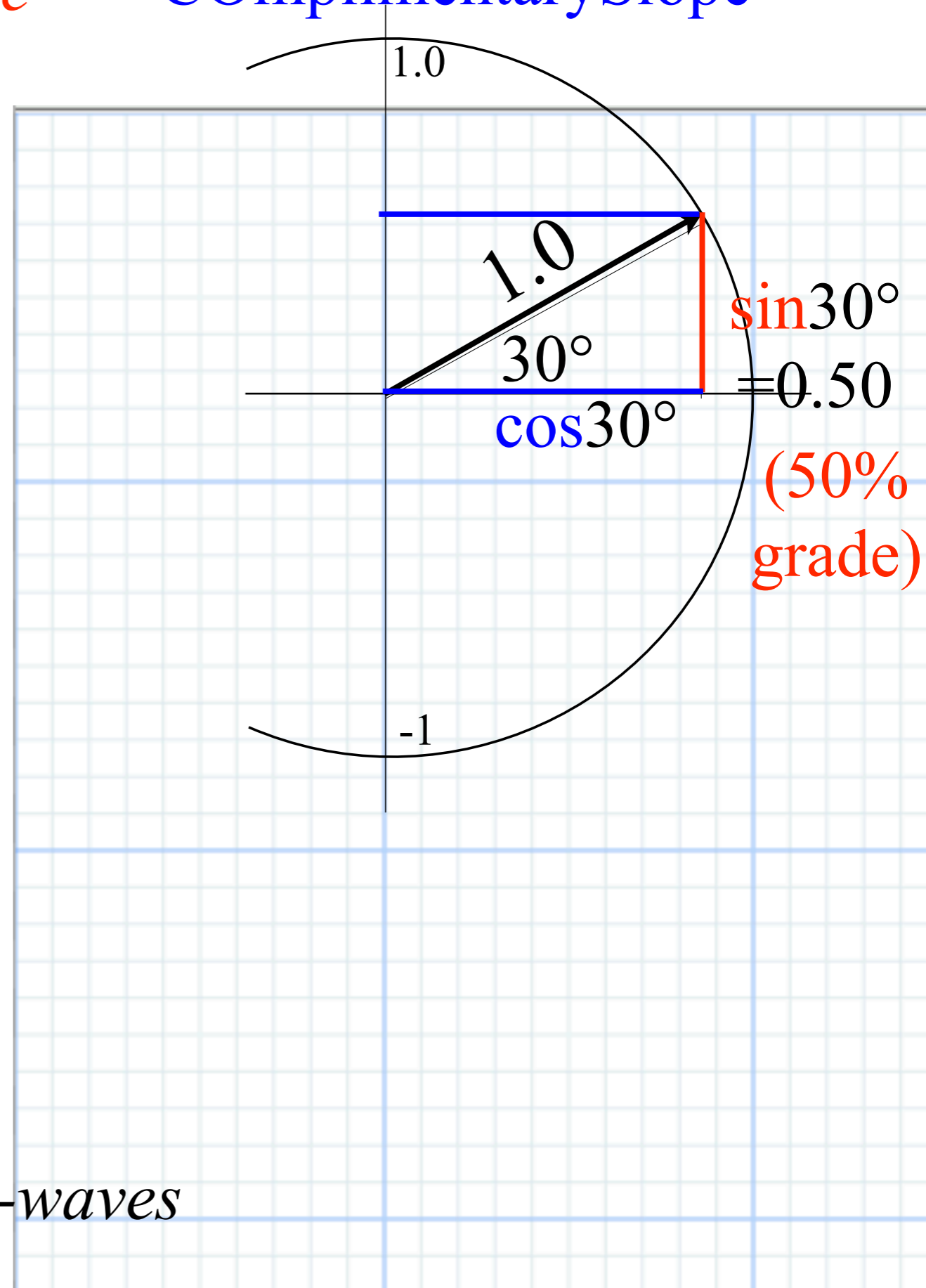
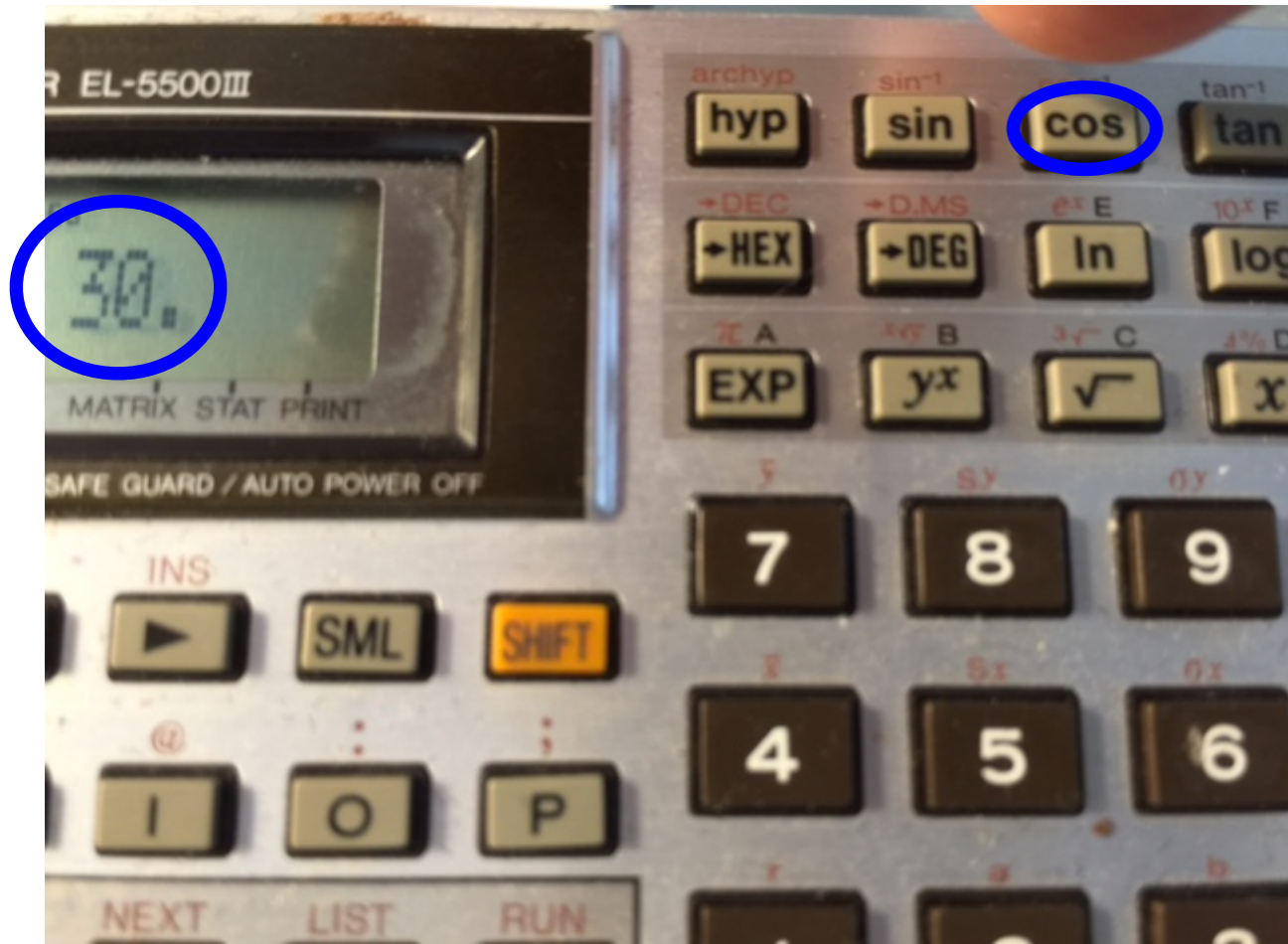


# Learning about SIN “Slope of INcline”



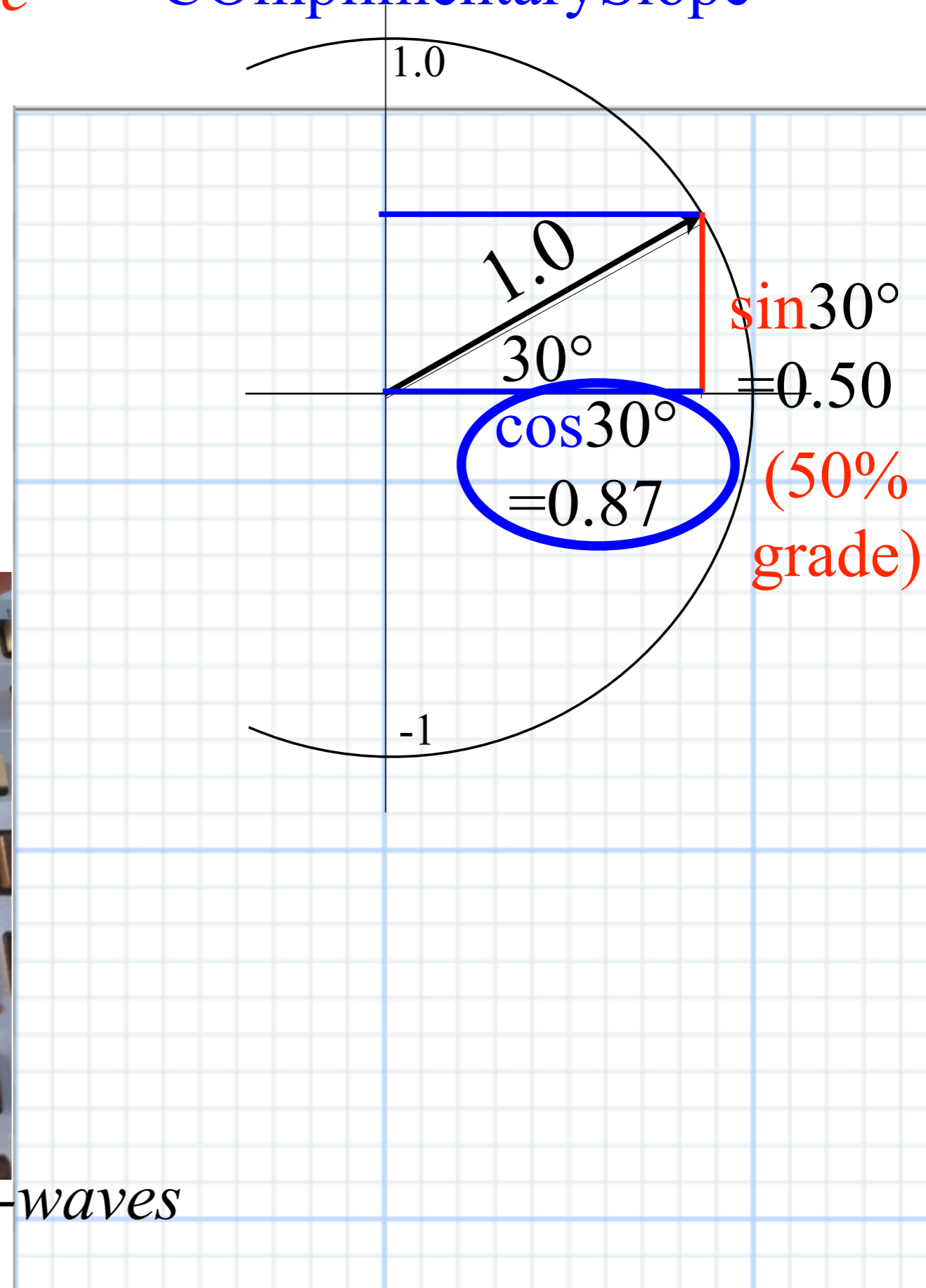
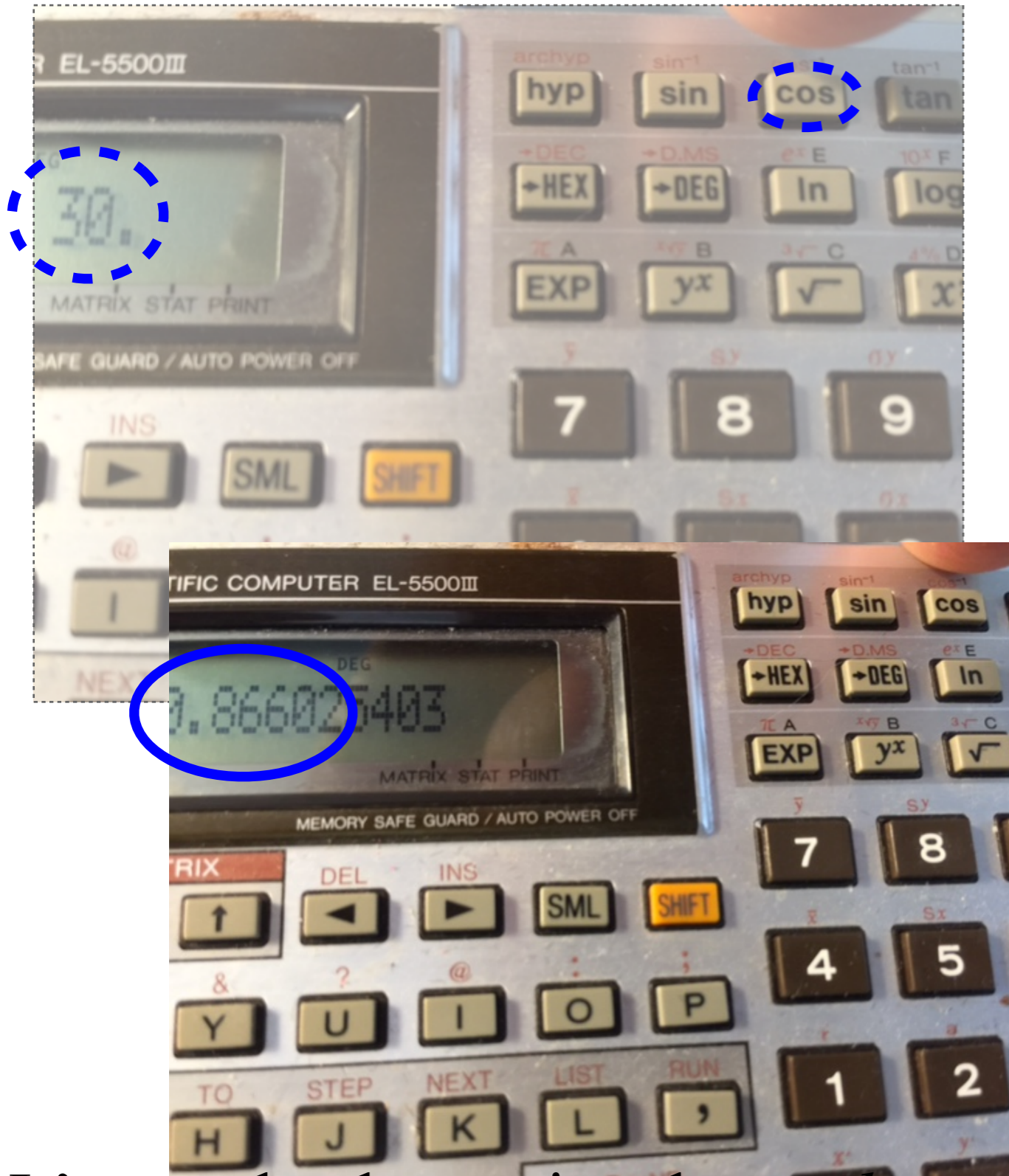
It's mostly about triangles *and sine-waves*

# Learning about **SIN** and the **COS**in “*Slope of INcline*” “**C**Omplimentary**S**lope”



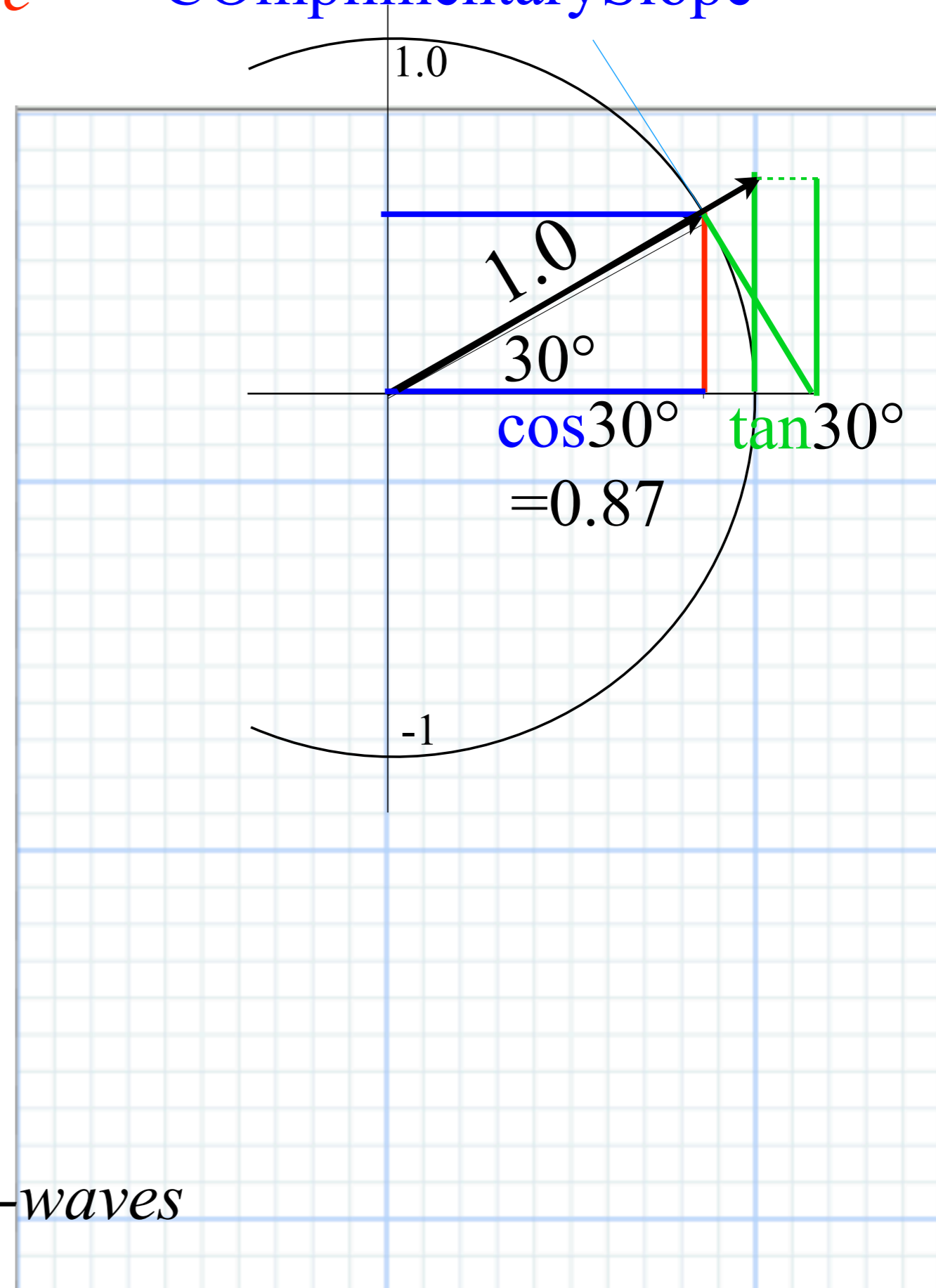
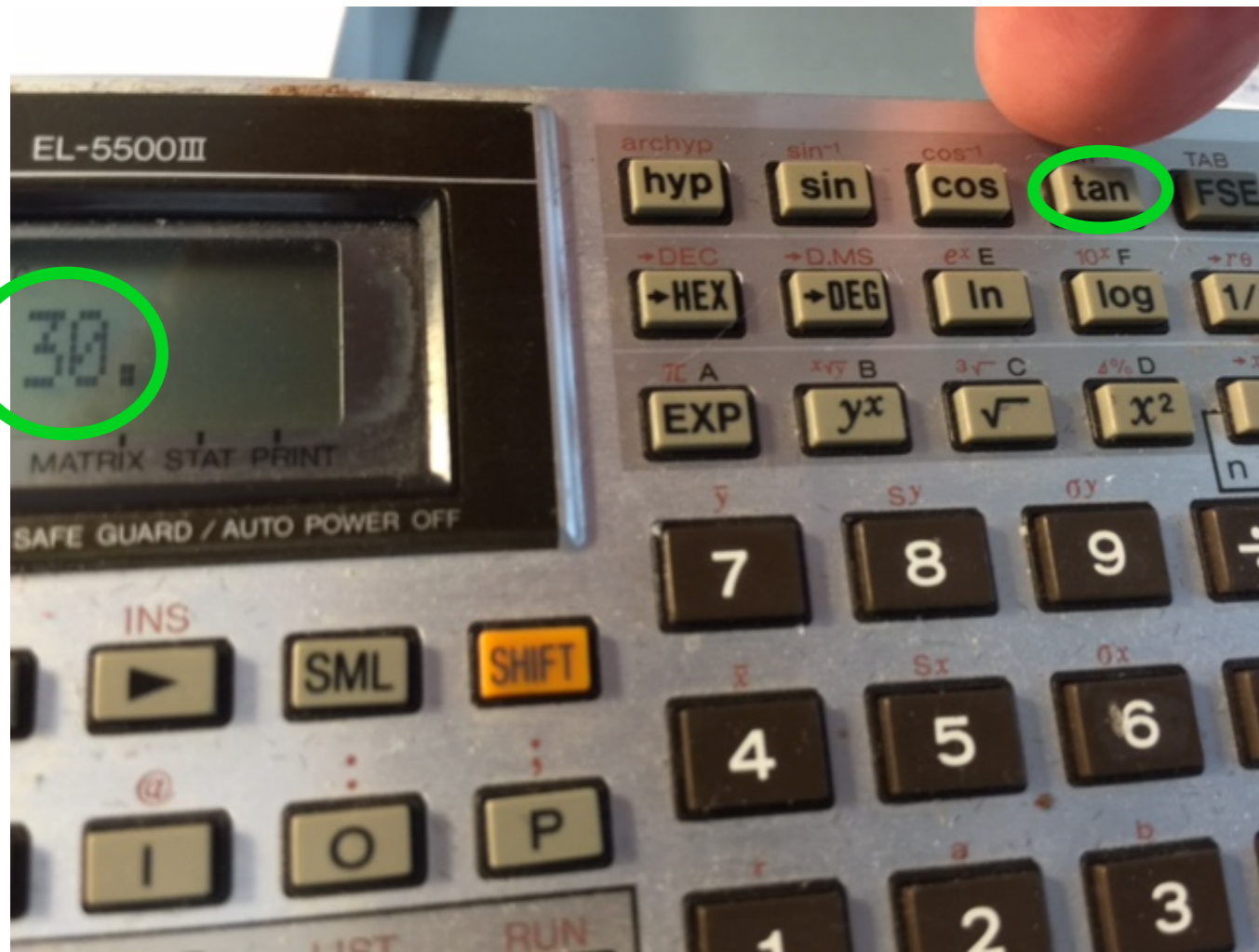
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# Learning about SIN and the COS in “Slope of INcline” “COmplimentary Slope”



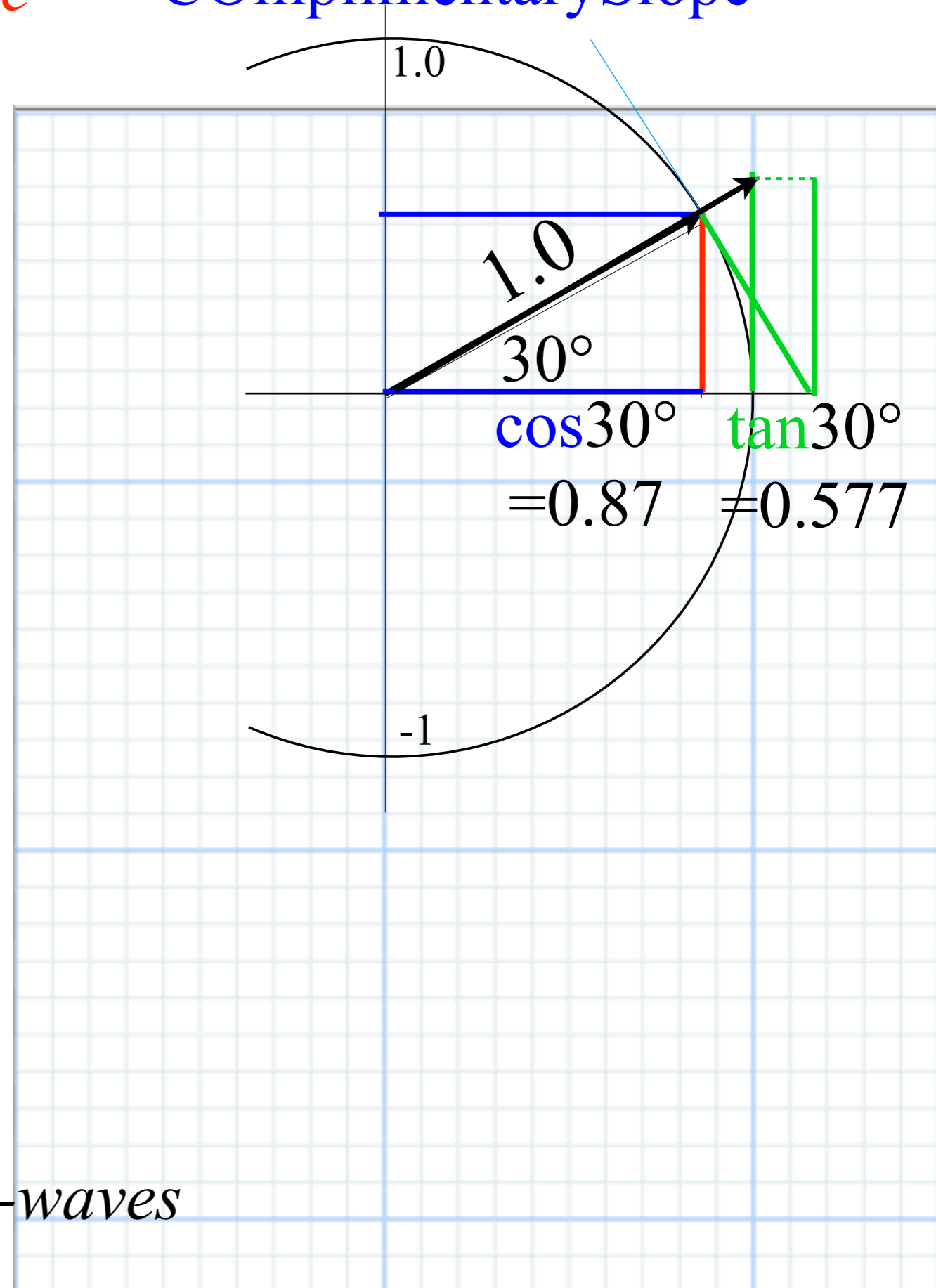
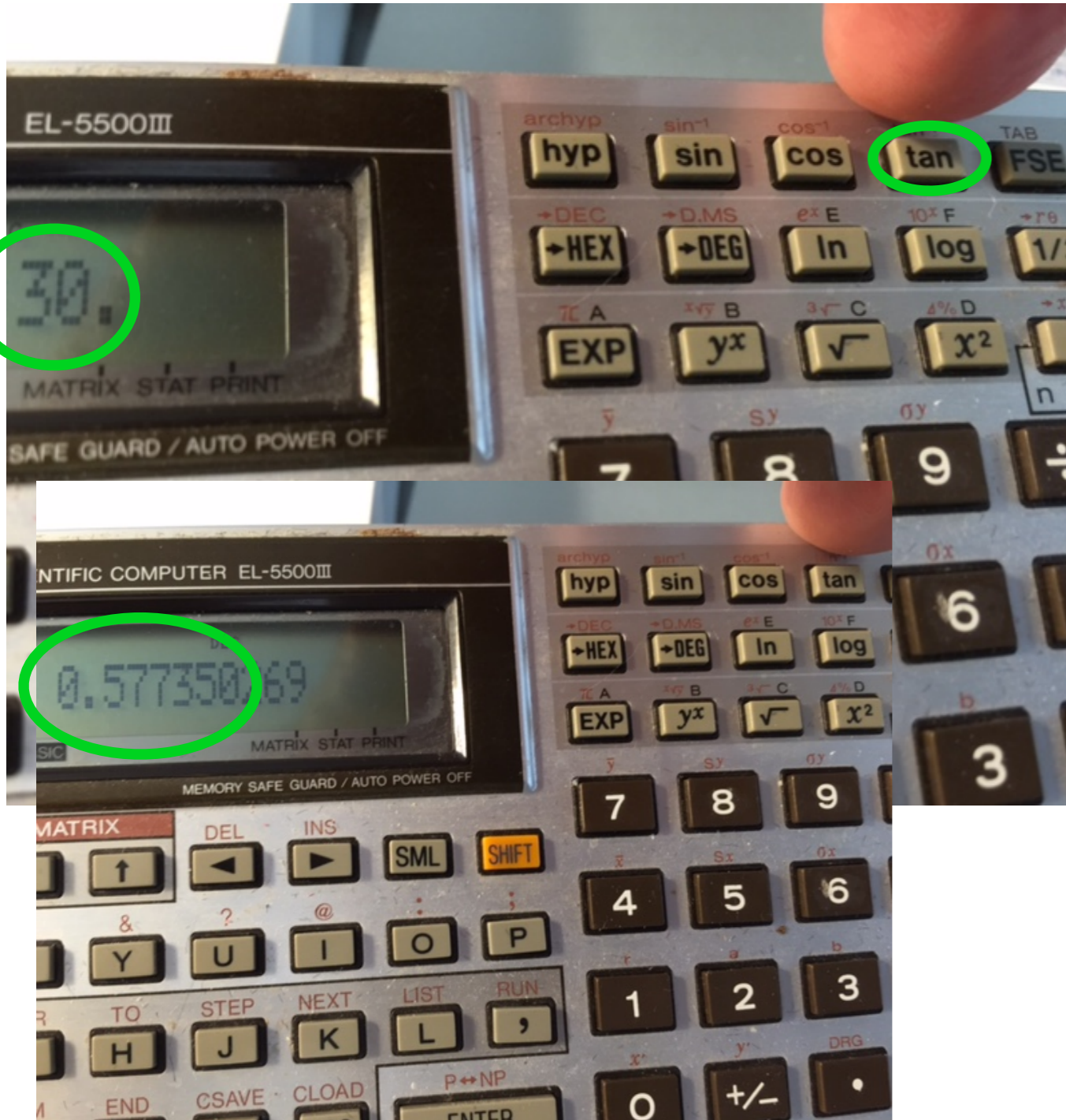
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# Learning about **SIN** and the **COS**in and **TAN**gent “*Slope of INcline*” “*COmplimentarySlope*”



It's mostly about triangles *and sine-waves*

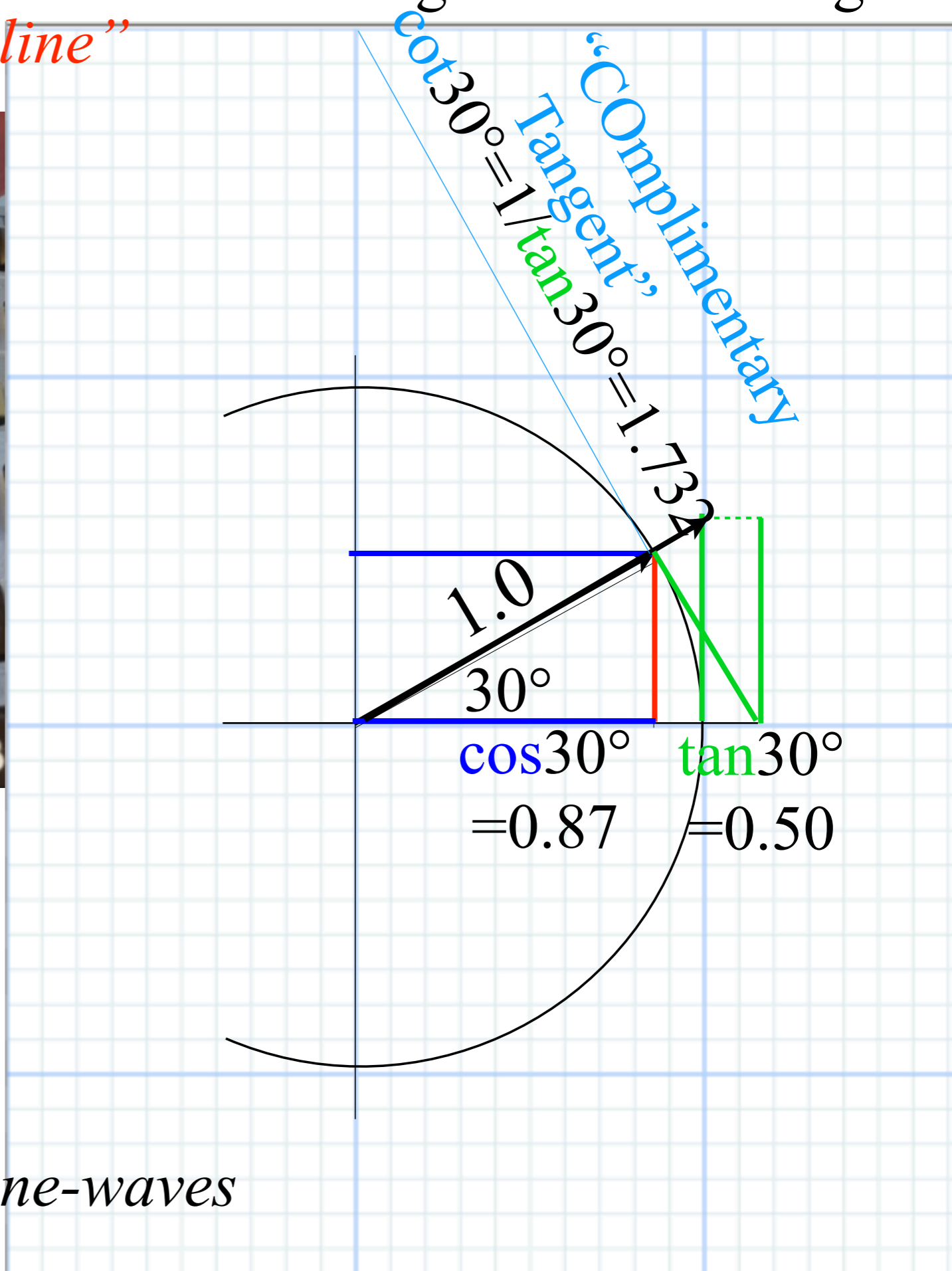
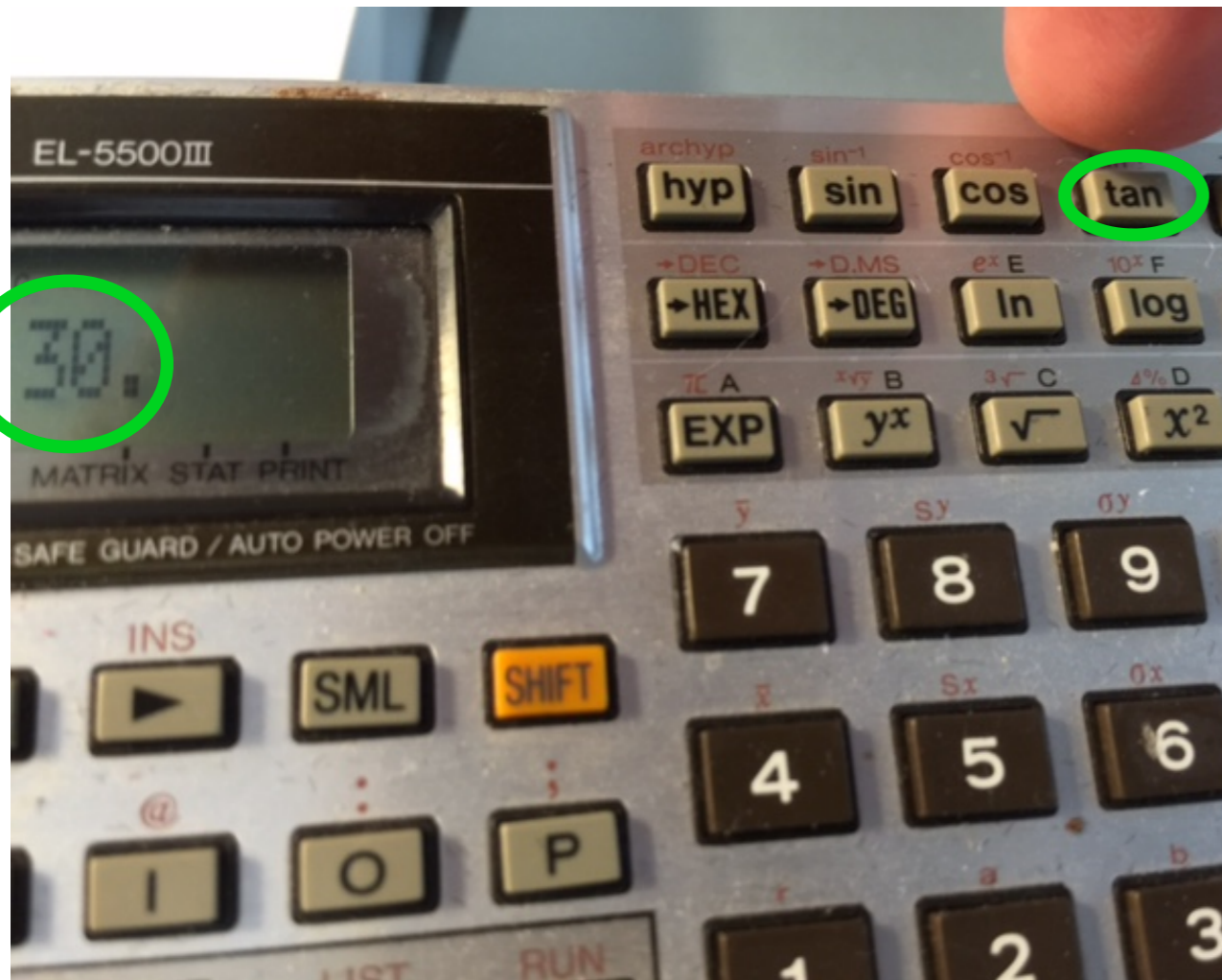
# Learning about **SIN** and the **COS**in and **TAN**gent “*Slope of INcline*” “*C*Omplimentary*S*lope”



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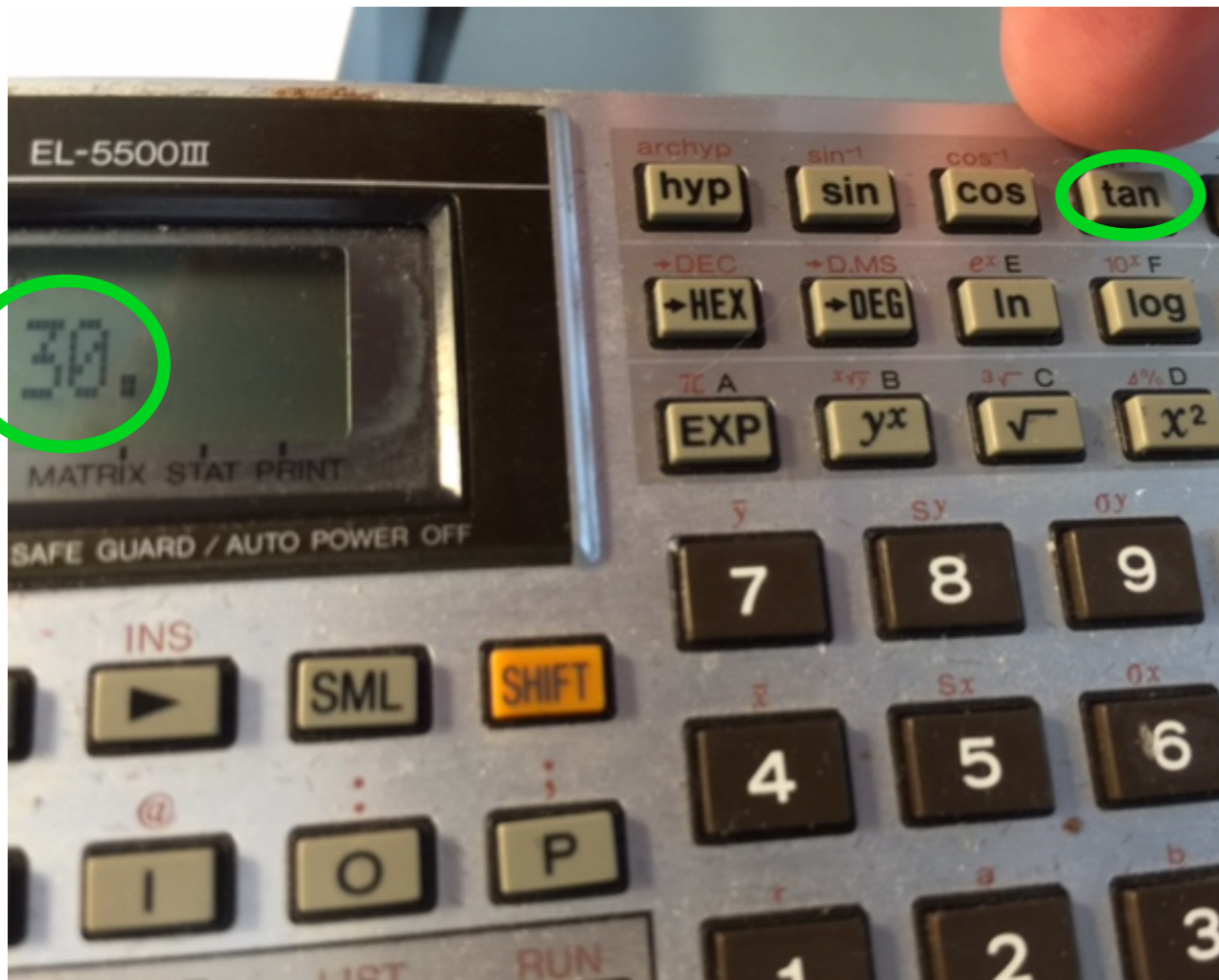


# Learning about **SIN** and the **COS**in and **TAN**gent and **CO**Tangent *“Slope of INcline”*



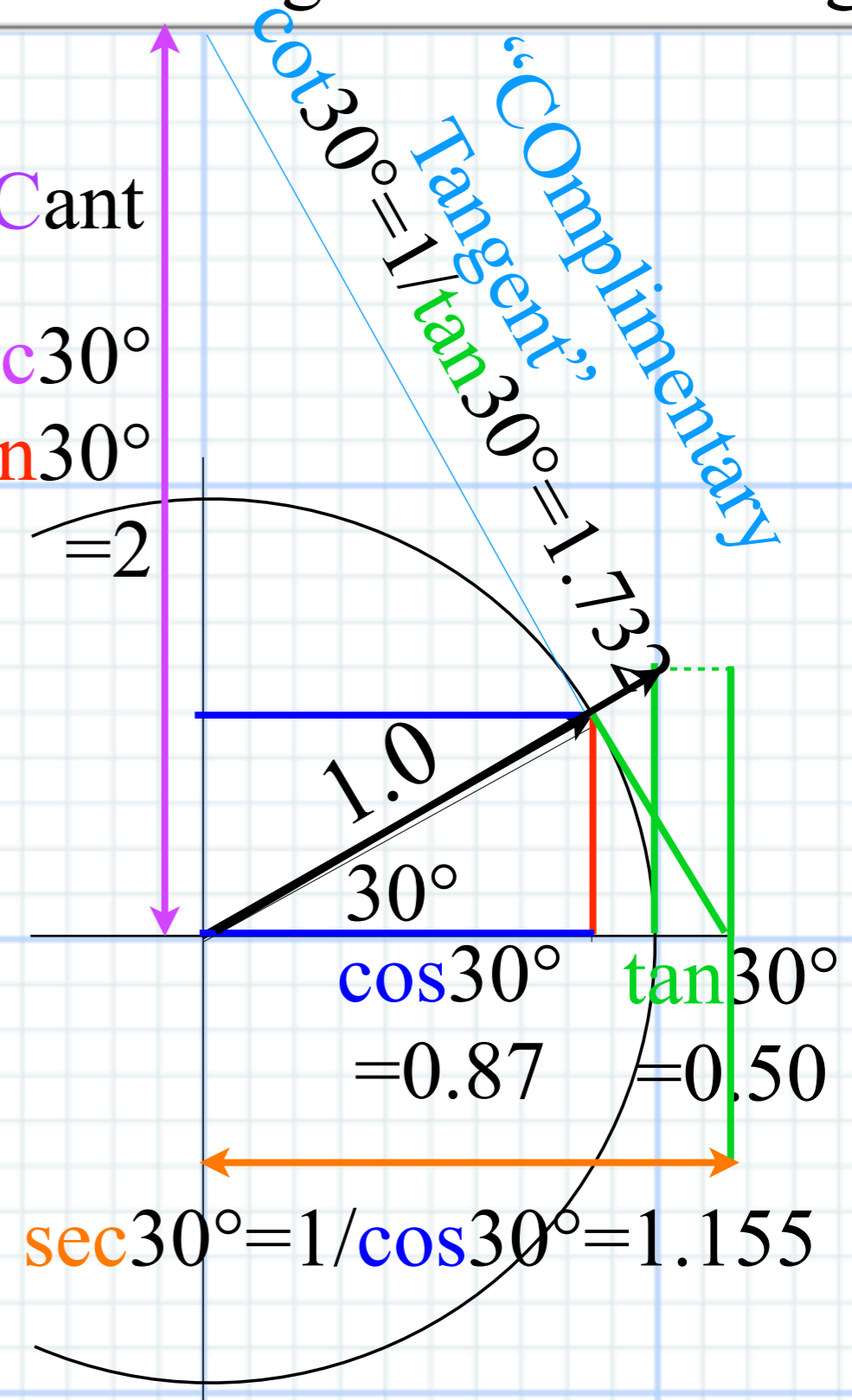
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# Learning about **SIN** and the **COS**in and **TAN**gent and **CO**Tangent *“Slope of INcline”*



...and  
**CoSeCant**

$$\text{csc}30^\circ = 1/\text{sin}30^\circ = 2$$



...and **SECant**

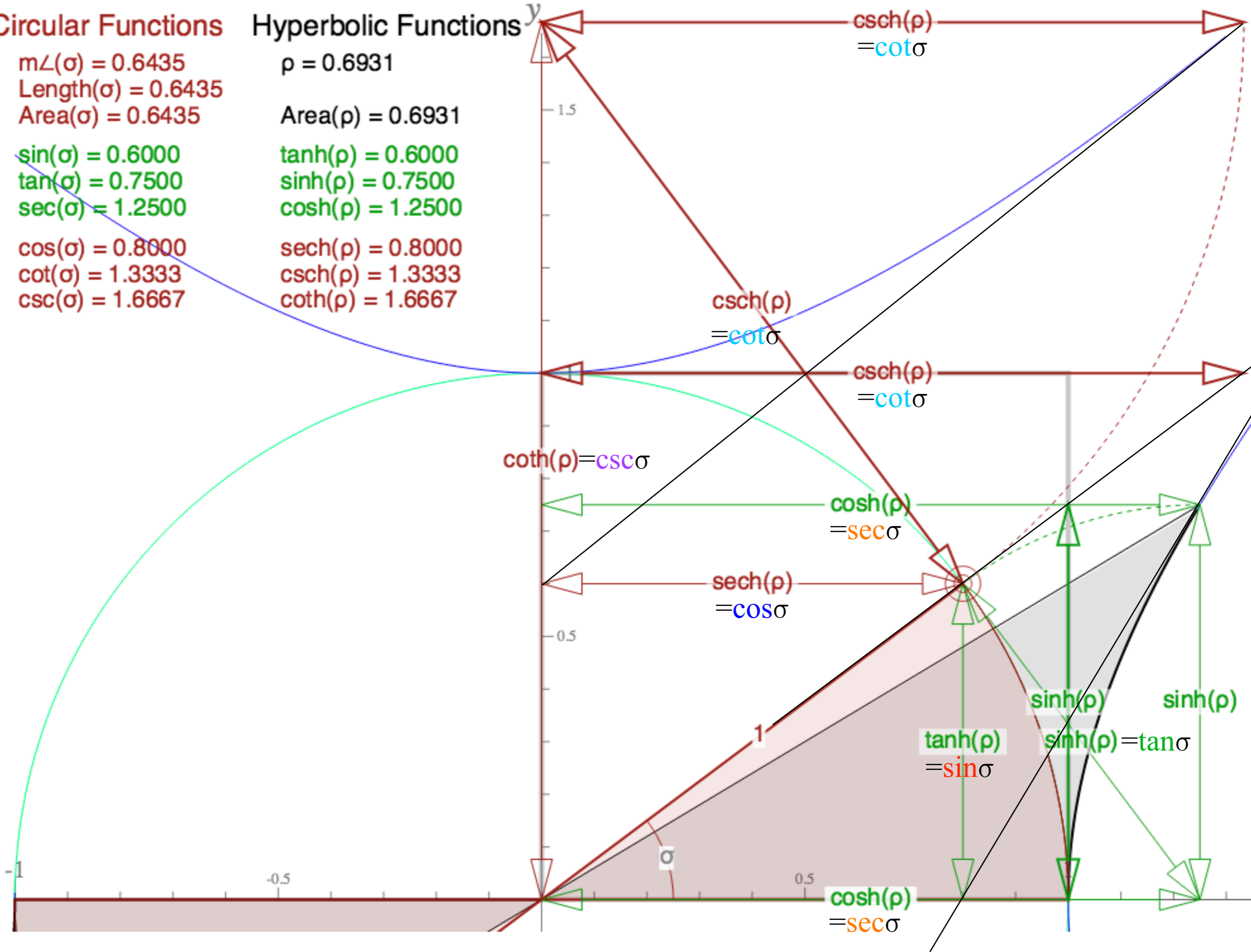
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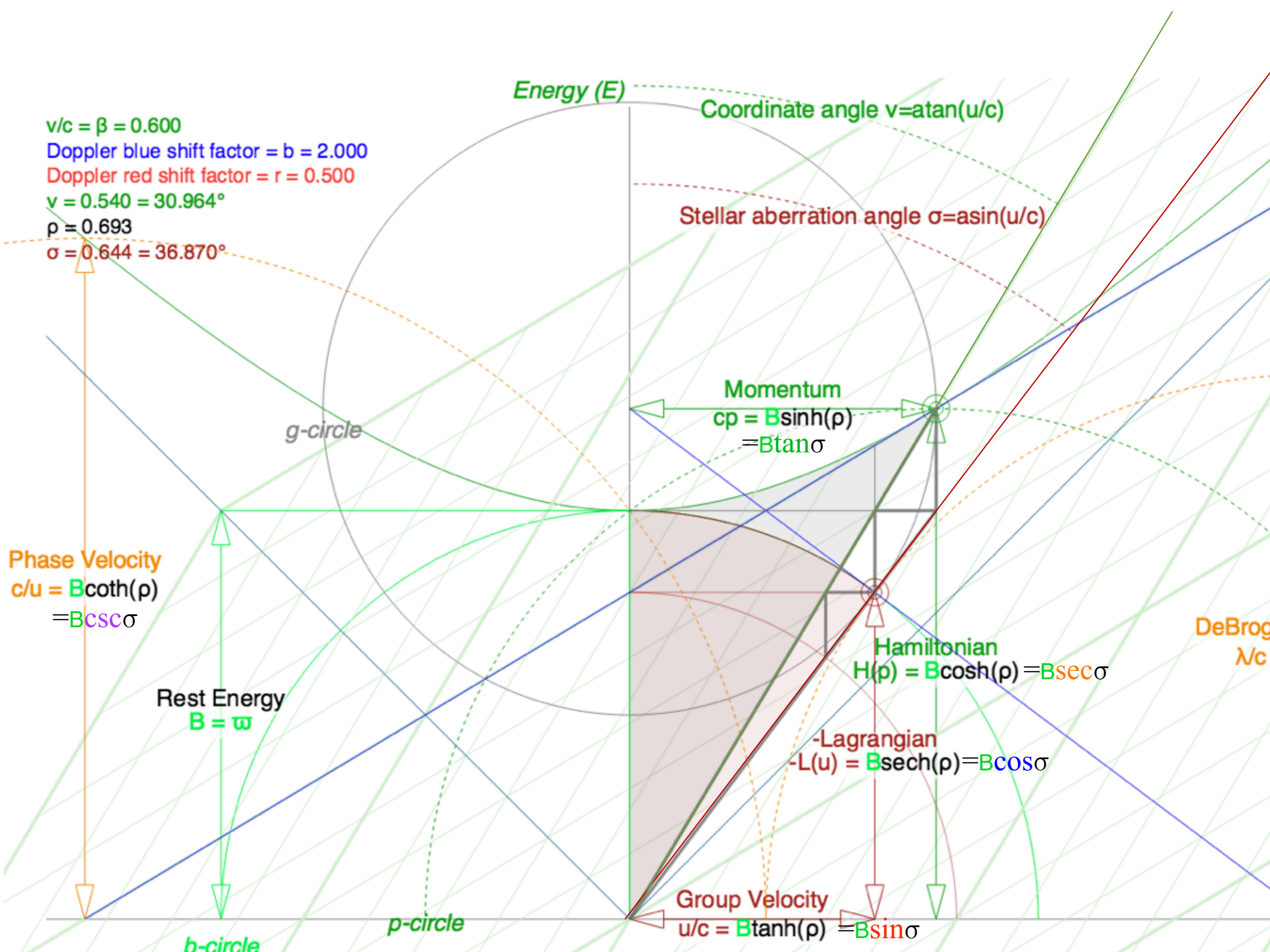
### Circular Functions

$m\angle(\sigma) = 0.6435$   
 $\text{Length}(\sigma) = 0.6435$   
 $\text{Area}(\sigma) = 0.6435$   
 $\sin(\sigma) = 0.6000$   
 $\tan(\sigma) = 0.7500$   
 $\sec(\sigma) = 1.2500$   
 $\cos(\sigma) = 0.8000$   
 $\cot(\sigma) = 1.3333$   
 $\csc(\sigma) = 1.6667$

### Hyperbolic Functions

$\rho = 0.6931$   
 $\text{Area}(\rho) = 0.6931$   
 $\tanh(\rho) = 0.6000$   
 $\sinh(\rho) = 0.7500$   
 $\cosh(\rho) = 1.2500$   
 $\text{sech}(\rho) = 0.8000$   
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# Using (some) wave parameters to develop relativistic quantum theory

$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \text{ (for } u \ll c \text{)}$$

$$cK_{phase} = B \sinh \rho \approx B \rho \text{ (for } u \ll c \text{)}$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2$$

$$\sinh \rho \approx \rho$$

$$B = v_A$$

$$B = v_A = cK_A$$

At low speeds:

<i>group</i>	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
<i>phase</i>	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
<i>rapidity</i> $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
<i>stellar</i> $\nabla$ <i>angle</i> $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
<i>value for</i> $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

[RelaWavity Web Simulation - Relativistic Terms](#)  
(Expanded Table)

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$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

At low speeds:

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

$$\sinh \rho \approx \rho \approx \frac{u}{c}$$

$$B = v_A$$

$$B = v_A = cK_A$$

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar $\nabla$ angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

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$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

$$\sinh \rho \approx \rho \approx \frac{u}{c}$$

$$B = v_A$$

$$B = v_A = cK_A$$

At low speeds:

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2 \quad \Leftarrow \text{for } (u \ll c)$$

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar $\nabla$ angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$



# Using (some) wave parameters to develop relativistic quantum theory

$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \quad (\text{for } u \ll c)$$

$$c\kappa_{phase} = B \sinh \rho \approx B \rho \quad (\text{for } u \ll c)$$

$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

$$\sinh \rho \approx \rho \approx \frac{u}{c}$$

$$B = v_A$$

$$B = v_A = c\kappa_A$$

At low speeds:

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2$$

for  $(u \ll c) \Rightarrow$

$$\kappa_{phase} \approx \frac{B}{c^2} u$$

time	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar $\nabla$ angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\operatorname{csc} \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

# Using (some) wave parameters to develop relativistic quantum theory

$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \quad (\text{for } u \ll c)$$

$$cK_{phase} = B \sinh \rho \approx B \rho \quad (\text{for } u \ll c)$$

$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

$$\sinh \rho \approx \rho \approx \frac{u}{c}$$

$$B = v_A$$

$$B = v_A = cK_A$$

At low speeds:

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2 \quad \Leftarrow \text{for } (u \ll c) \Rightarrow$$

$$K_{phase} \approx \frac{B}{c^2} u$$

$v_{phase}$  and  $K_{phase}$  resemble formulae for Newton's kinetic energy and momentum

Resembles:  $const. + \frac{1}{2} Mu^2$

Resembles:  $Mu$

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar $\nabla$ angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

# Using (some) wave parameters to develop relativistic quantum theory

$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \quad (\text{for } u \ll c)$$

$$cK_{phase} = B \sinh \rho \approx B \rho \quad (\text{for } u \ll c)$$

$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

$$\sinh \rho \approx \rho \approx \frac{u}{c}$$

$$B = v_A$$

$$B = v_A = cK_A$$

At low speeds:

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2 \quad \Leftarrow \text{for } (u \ll c) \Rightarrow$$

$$K_{phase} \approx \frac{B}{c^2} u$$

$v_{phase}$  and  $K_{phase}$  resemble formulae for Newton's kinetic energy  $\frac{1}{2} Mu^2$  and momentum  $Mu$ .

Rescale  $v_{phase}$  by  $h$  so:  $M = \frac{hB}{c^2}$

Resembles:  $const. + \frac{1}{2} Mu^2$

Resembles:  $Mu$

time	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar $\nabla$ angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

# Using (some) wave parameters to develop relativistic quantum theory

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$$cK_{phase} = B \sinh \rho \approx B \rho \quad (\text{for } u \ll c)$$

$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

$$\sinh \rho \approx \rho \approx \frac{u}{c}$$

$$B = v_A$$

$$B = v_A = cK_A$$

At low speeds:

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2 \quad \Leftarrow \text{for } (u \ll c) \Rightarrow$$

$$K_{phase} \approx \frac{B}{c^2} u$$

Rescale  $v_{phase}$  by  $h$  so:  $M = \frac{hB}{c^2}$

$v_{phase}$  and  $K_{phase}$  resemble formulae for Newton's kinetic energy  $\frac{1}{2} Mu^2$  and momentum  $Mu$ .

So attach scale factor  $h$  to match units.

Resembles:  $const. + \frac{1}{2} Mu^2$

Resembles:  $Mu$

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar $\nabla$ angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

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$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \quad (\text{for } u \ll c)$$

$$cK_{phase} = B \sinh \rho \approx B \rho \quad (\text{for } u \ll c)$$

$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

$$\sinh \rho \approx \rho \approx \frac{u}{c}$$

$$B = v_A$$

$$B = v_A = cK_A$$

At low speeds:

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2 \quad \Leftarrow \text{for } (u \ll c) \Rightarrow$$

$$K_{phase} \approx \frac{B}{c^2} u$$

$v_{phase}$  and  $K_{phase}$  resemble formulae for Newton's kinetic energy  $\frac{1}{2} Mu^2$  and momentum  $Mu$ .

Rescale  $v_{phase}$  by  $h$  so:  $M = \frac{hB}{c^2}$

$$hv_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2 \quad \Leftarrow \text{for } (u \ll c) \Rightarrow$$

$$hK_{phase} \approx \frac{hB}{c^2} u$$

So attach scale factor  $h$  to match units.

Resembles:  $const. + \frac{1}{2} Mu^2$

Resembles:  $Mu$

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

# Using (some) wave parameters to develop relativistic quantum theory

$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \quad (\text{for } u \ll c)$$

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$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

$$\sinh \rho \approx \rho \approx \frac{u}{c}$$

$$B = v_A$$

$$B = v_A = cK_A$$

At low speeds:

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2$$

for  $(u \ll c) \Rightarrow$

$$K_{phase} \approx \frac{B}{c^2} u$$

Rescale  $v_{phase}$  by  $h$  so:  $M = \frac{hB}{c^2}$

or:  $hB = Mc^2$  (The famous  $Mc^2$  shows up here!)

$v_{phase}$  and  $K_{phase}$  resemble formulae for Newton's kinetic energy  $\frac{1}{2} Mu^2$  and momentum  $Mu$ .

$$hv_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2$$

Resembles:  $const. + \frac{1}{2} Mu^2$

for  $(u \ll c) \Rightarrow$

$$hK_{phase} \approx \frac{hB}{c^2} u$$

Resembles:  $Mu$

So attach scale factor  $h$  to match units.

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

# Using (some) wave parameters to develop relativistic quantum theory

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$$cK_{phase} = B \sinh \rho \approx B \rho \quad (\text{for } u \ll c)$$

$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

$$\sinh \rho \approx \rho \approx \frac{u}{c}$$

$$B = v_A$$

$$B = v_A = cK_A$$

At low speeds:

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2$$

$\Leftarrow$  for  $(u \ll c) \Rightarrow$

$$K_{phase} \approx \frac{B}{c^2} u$$

$v_{phase}$  and  $K_{phase}$  resemble formulae for Newton's kinetic energy  $\frac{1}{2} Mu^2$  and momentum  $Mu$ .

Rescale  $v_{phase}$  by  $h$  so:  $M = \frac{hB}{c^2}$

or:  $hB = Mc^2$  (The famous  $Mc^2$  shows up here!)

$$hv_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2$$

$\Leftarrow$  for  $(u \ll c) \Rightarrow$

$$hK_{phase} \approx \frac{hB}{c^2} u$$

So attach scale factor  $h$  to match units.

$$hv_{phase} \approx Mc^2 + \frac{1}{2} Mu^2$$

$\Leftarrow$  for  $(u \ll c) \Rightarrow$

$$hK_{phase} \approx Mu$$

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar $\nabla$ angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

# Using (some) wave parameters to develop relativistic quantum theory

$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \quad (\text{for } u \ll c)$$

$$cK_{phase} = B \sinh \rho \approx B \rho \quad (\text{for } u \ll c)$$

$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

$$\sinh \rho \approx \rho \approx \frac{u}{c}$$

$$B = v_A$$

$$B = v_A = cK_A$$

At low speeds:

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2$$

$\Leftarrow$  for  $(u \ll c) \Rightarrow$

$$K_{phase} \approx \frac{B}{c^2} u$$

$v_{phase}$  and  $K_{phase}$  resemble formulae for Newton's kinetic energy  $\frac{1}{2} Mu^2$  and momentum  $Mu$ .

Rescale  $v_{phase}$  by  $h$  so:  $M = \frac{hB}{c^2}$  or:  $hB = Mc^2$

(The famous  $Mc^2$  shows up here!)

$$hv_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2$$

$\Leftarrow$  for  $(u \ll c) \Rightarrow$

$$hK_{phase} \approx \frac{hB}{c^2} u$$

So attach scale factor  $h$  to match units.

$$hv_{phase} \approx Mc^2 + \frac{1}{2} Mu^2$$

$\Leftarrow$  for  $(u \ll c) \Rightarrow$

$$hK_{phase} \approx Mu$$

Lucky coincidences?? Cheap trick??

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$



# Using (some) wave parameters to develop relativistic quantum theory

$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \quad (\text{for } u \ll c)$$

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$$B = v_A$$

$$B = v_A = cK_A$$

At low speeds:

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2 \quad \Leftarrow \text{for } (u \ll c) \Rightarrow$$

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$v_{phase}$  and  $K_{phase}$  resemble formulae for Newton's kinetic energy  $\frac{1}{2} Mu^2$  and momentum  $Mu$ .

Rescale  $v_{phase}$  by  $h$  so:  $M = \frac{hB}{c^2}$  or:  $hB = Mc^2$

(The famous  $Mc^2$  shows up here!)

$$hv_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2 \quad \Leftarrow \text{for } (u \ll c) \Rightarrow$$

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So attach scale factor  $h$  to match units.

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Lucky coincidences?? Cheap trick??  
...Try exact  $v_{phase}$  ...

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
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$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
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RelaWavity Web Simulation - Relativistic Terms  
(Expanded Table)

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phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
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(old-fashioned notation)

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Planck (1900)  $\uparrow$   
Einstein (1905)  $\uparrow$

(old-fashioned notation)

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phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\text{coth } \rho$	$e^{+\rho}$
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Max Planck  
1858-1947

# Using (some) wave parameters to develop relativistic quantum theory

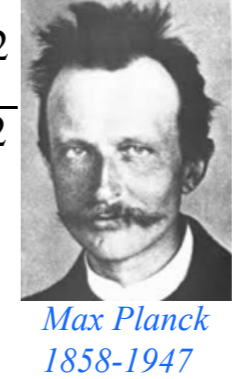
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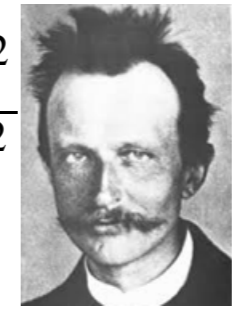
Planck (1900)

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phase	$b_{BLUE}^{Doppler}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$b_{RED}^{Doppler}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\text{coth } \rho$	$e^{+\rho}$
stellar $\nabla$ angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
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# Using (some) wave parameters to develop relativistic quantum theory



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1858-1947

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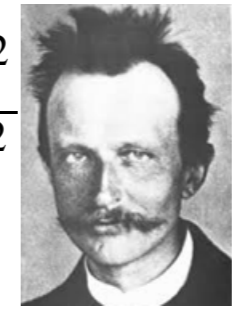
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This motivates the "particle" normalization  $\int \Psi^* \Psi dV = N$   $\Psi = \sqrt{\frac{\epsilon_0}{hv}} E$

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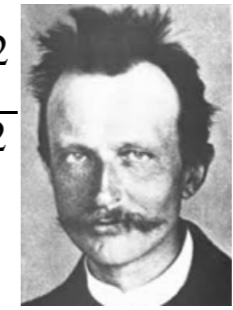
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This motivates the "particle" normalization  $\int \Psi^* \Psi dV = N \quad \Psi = \sqrt{\frac{\epsilon_0}{h\nu}} E$

Big worry: Is not oscillator energy quadratic in frequency  $\nu$ ?  
HO energy =  $\frac{1}{2} A^2 \nu^2$

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Rescale  $v_{phase}$  by  $h$  so:  $M = \frac{hB}{c^2}$  or:  $hB = Mc^2$  (The famous  $Mc^2$  shows up here!)

$$hv_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2 \quad \Leftarrow \text{for } (u \ll c) \Rightarrow$$

$$hK_{phase} \approx \frac{hB}{c^2} u$$

So attach scale factor  $h$  (or  $hN$ ) to match units.

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Lucky coincidences?? Cheap trick??  
...Try exact  $v_{phase}$  ...

Need to replace  $h$  with  $hN$  to match e.m. energy density  $\epsilon_0 \mathbf{E} \cdot \mathbf{E} = hNv_{phase}$

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Planck (1900)

$$= \text{Total Energy: } E = \frac{Mc^2}{\sqrt{1-u^2/c^2}}$$

Einstein (1905)

This motivates the "particle" normalization  $\int \Psi^* \Psi dV = N \quad \Psi = \sqrt{\frac{\epsilon_0}{h\nu}} E$

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$
phase	$b_{BLUE}^{Doppler}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$		
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$		
stellar angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$
	$\frac{2}{1} = 2.0$						

Big worry: Is not oscillator energy quadratic in frequency  $\nu$ ?  
HO energy =  $\frac{1}{2} A^2 \nu^2$   
Resolution and dirty secret:  $\mathbf{E}$ ,  $N$ , and  $v_{phase}$  are all frequencies!  
So  $\epsilon_0 \mathbf{E} \cdot \mathbf{E} = hNv_{phase}$  is quadratic in  $v_{phase}$

# Using (some) wave parameters to develop relativistic quantum theory

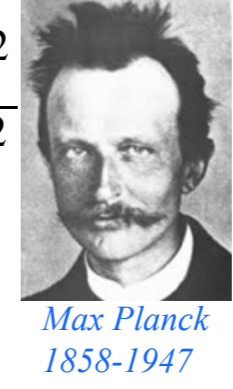
$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \quad (\text{for } u \ll c)$$

$$cK_{phase} = B \sinh \rho \approx B \rho \quad (\text{for } u \ll c)$$

$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

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At low speeds:

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Need to replace  $h$  with  $hN$  to match e.m. energy density  $\epsilon_0 \mathbf{E} \cdot \mathbf{E} = hN v_{phase}$

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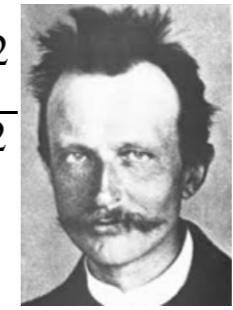
Einstein (1905)  $\uparrow$

$$hK_{phase} = hB \sinh \rho = Mc^2 \sinh \rho$$

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	
phase	$b_{BLUE}^{Doppler}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$b_{RED}^{Doppler}$
rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\text{coth } \rho$	$e^{+\rho}$
stellar angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
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# Using (some) wave parameters to develop relativistic quantum theory



Max Planck  
1858-1947

$$B = v_A$$

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$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

$$\sinh \rho \approx \rho \approx \frac{u}{c}$$

$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \quad (\text{for } u \ll c)$$

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$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

At low speeds:

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2 \quad \Leftarrow \text{for } (u \ll c) \Rightarrow$$

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$$hcK_{phase} = hB \sinh \rho = Mc^2 \sinh \rho$$

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$$cp = \frac{Mc u}{\sqrt{1-u^2/c^2}}$$

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rapidity $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\text{coth } \rho$
stellar $\nabla$ angle $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$
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phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$
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$$cp = \frac{Muc}{\sqrt{1-u^2/c^2}}$$

Momentum:  $hK_{phase} = p = \frac{Mu}{\sqrt{1-u^2/c^2}}$

DeBroglie (1921)

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⇐ for ( $u \ll c$ ) ⇒

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$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	
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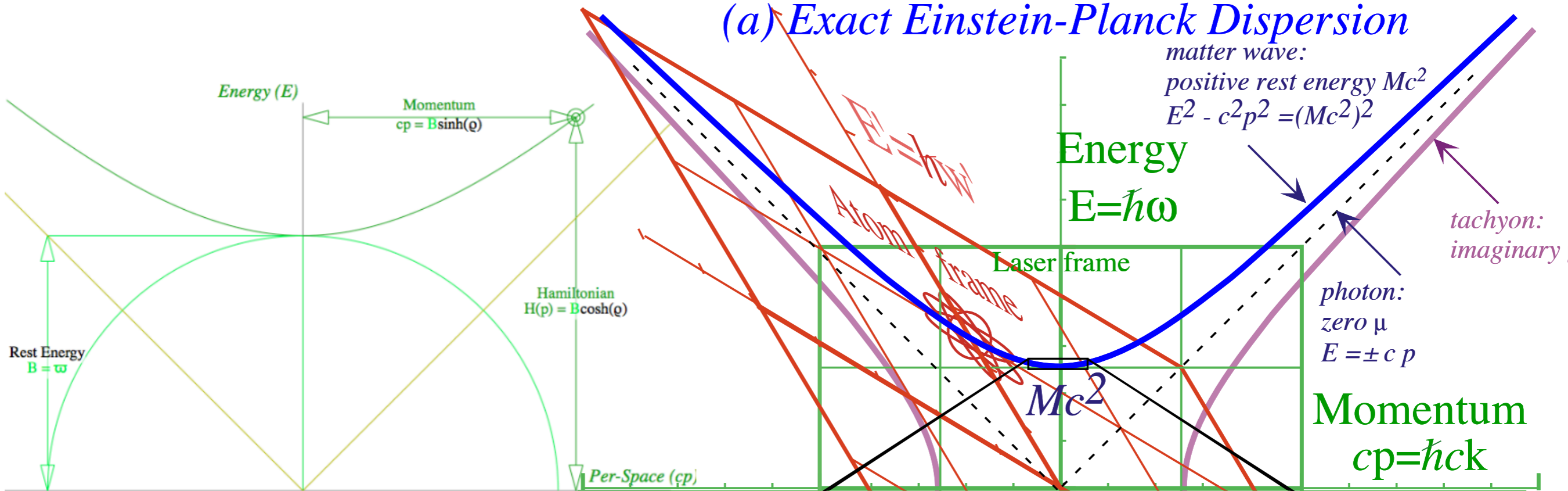
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Momentum:  $hK_{phase} = p = \frac{Mu}{\sqrt{1-u^2/c^2}}$

DeBroglie (1921)

# Using (some) wave coordinates for relativistic quantum theory



Mass (resting)

$$hB = h\nu_A = Mc^2 = hck_A$$

Energy

$$h\nu_{phase} = E = h\nu_A \cosh \rho$$

Momentum

$$hck_{phase} = cp = hck_A \sinh \rho = h\nu_A \sinh \rho$$

Energy versus Momentum

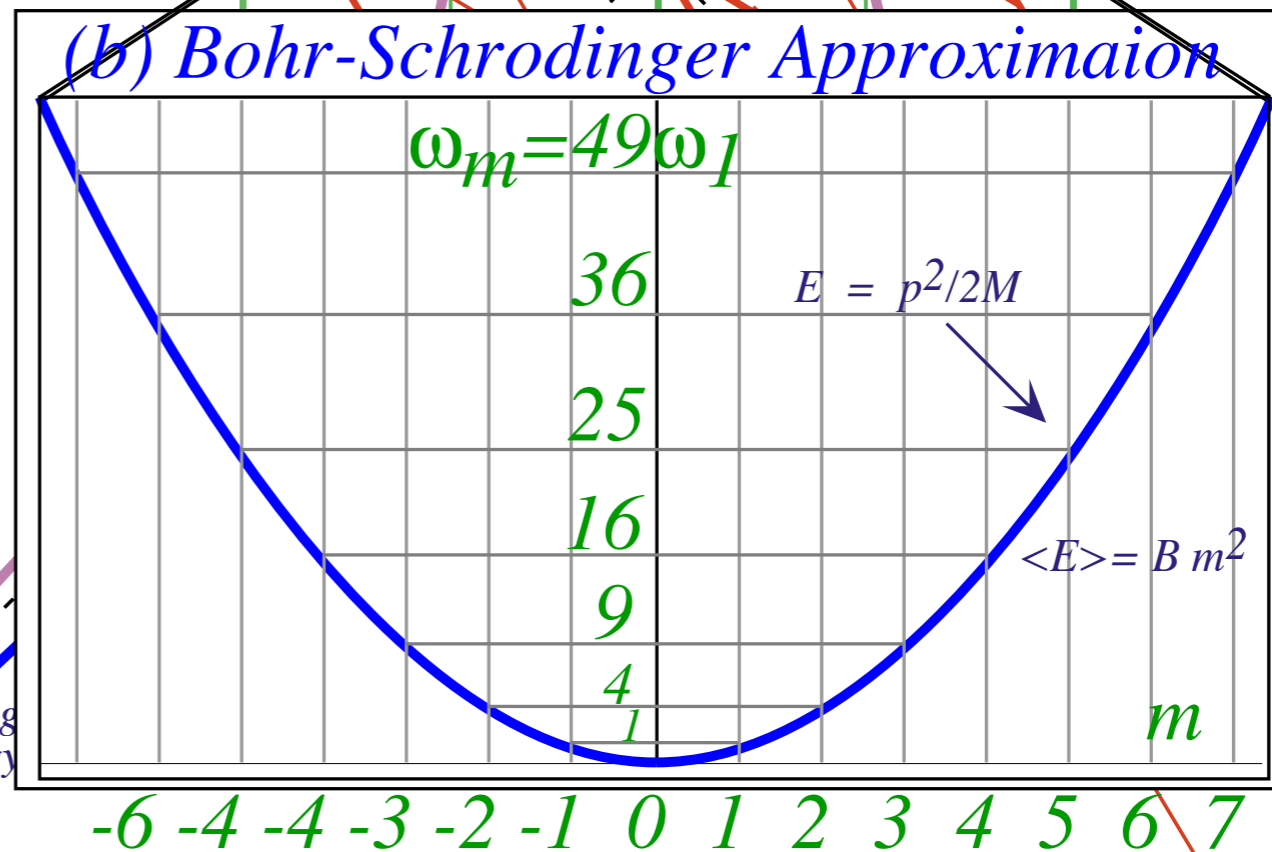
$$E^2 = (Mc^2)^2 \cosh^2 \rho$$

$$= (Mc^2)^2 (1 + \sinh^2 \rho) = (Mc^2)^2 + (cp)^2 \Rightarrow E = \pm \sqrt{(Mc^2)^2 + (cp)^2} \approx Mc^2 + \frac{p^2}{2M}$$



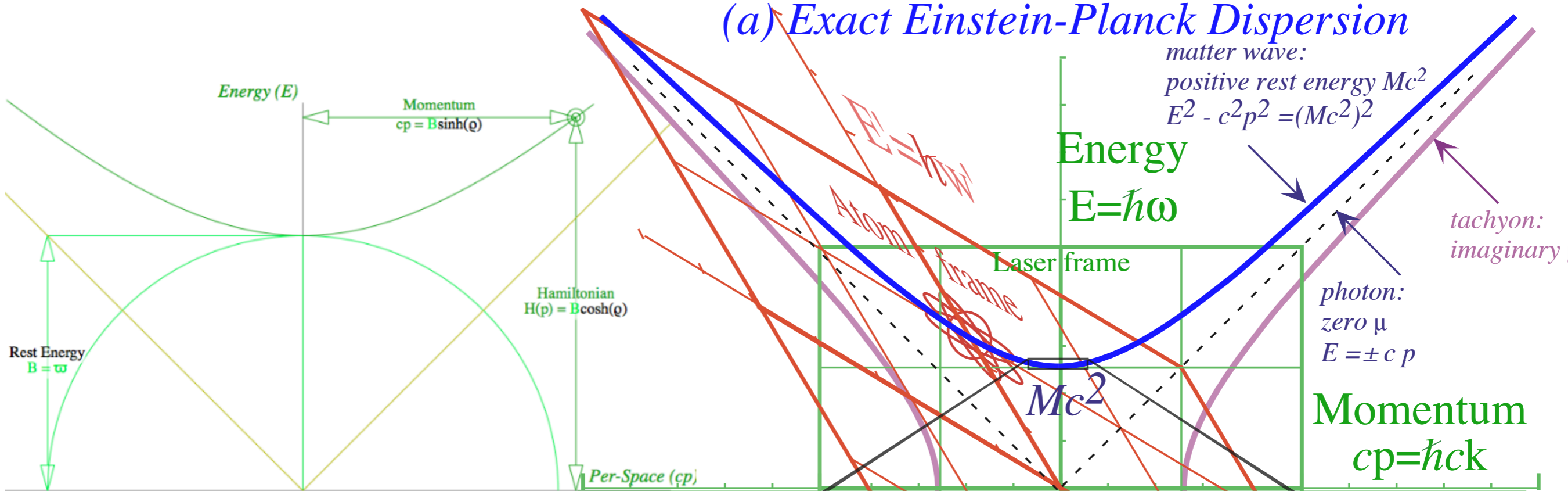
Niels Bohr  
1885-1962

(b) Bohr-Schrodinger Approximaion



Erwin Schrodinger  
1887-1961

# Using (some) wave coordinates for relativistic quantum theory



Mass (resting)

$$hB = h\nu_A = Mc^2 = hc\kappa_A$$

Energy

$$h\nu_{\text{phase}} = E = h\nu_A \cosh \rho$$

Momentum

$$hc\kappa_{\text{phase}} = cp = hc\kappa_A \sinh \rho = h\nu_A \sinh \rho$$

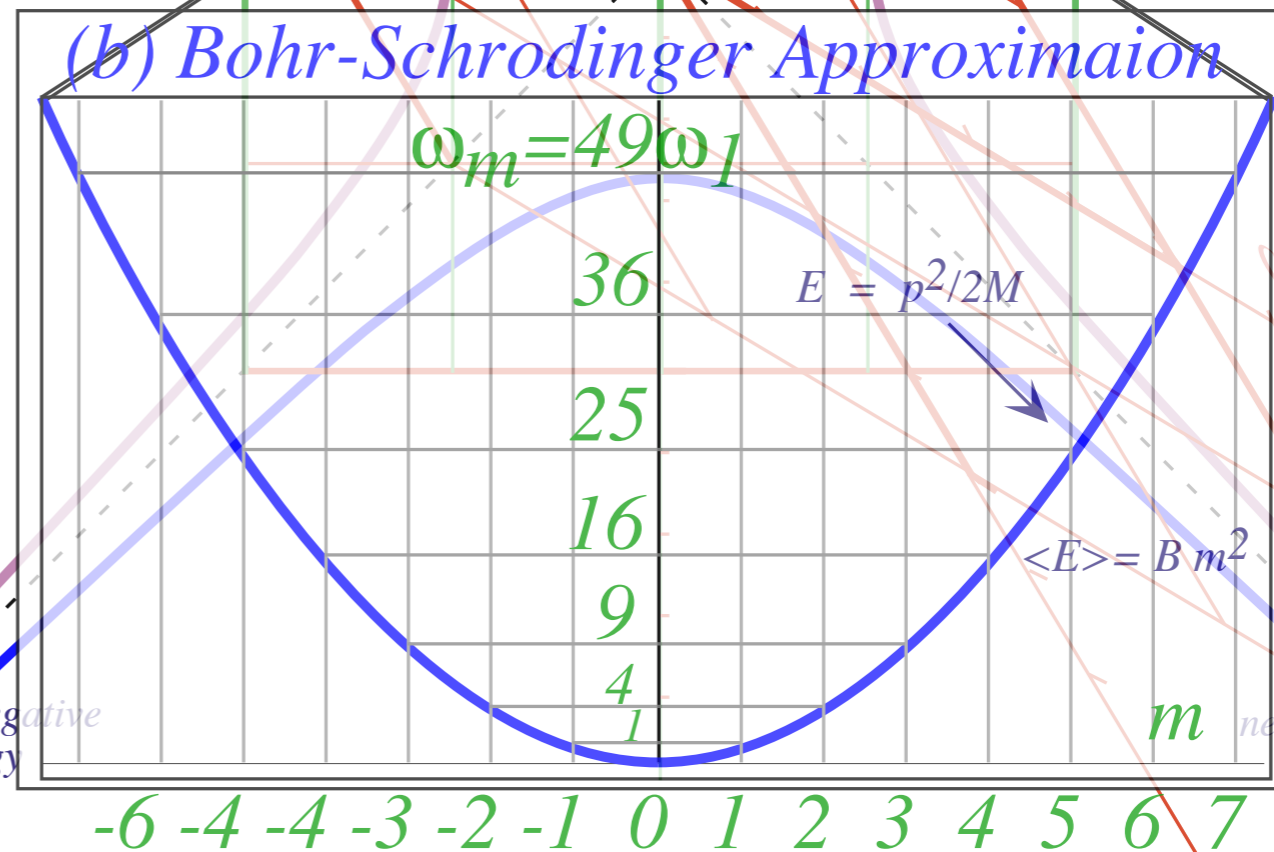
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low speed approximation

(b) Bohr-Schrodinger Approximaion



## Relativity variable tables

<i>group</i>	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
<i>phase</i>	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
<i>rapidity</i> $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
<i>stellar</i> $\nabla$ <i>angle</i> $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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<i>effects</i>	$b_{RED}^{Doppler}$	$V_{group}$	<i>past-future asymmetry</i> (off-diagonal Lorentz-transform)	$x$ -contraction <sup>(Lorentz)</sup> $\tau_{phase}$ -contraction	$t$ -dilation <sup>(Einstein)</sup> $v_{phase}$ -dilation (on-diagonal Lorentz-transform)	<i>inverse asymmetry</i>	$V_{phase}$	$b_{BLUE}^{Doppler}$

## Relativistic quantum mechanics variable tables

<i>group</i>	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
<i>phase</i>	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
<i>rapidity</i> $\rho$	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
<i>stellar</i> $\nabla$ <i>angle</i> $\sigma$	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{\beta}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{1-\beta^2}}{\beta}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
<i>value for</i> $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$
<i>functions</i>		$V_{group} = c \tanh \rho$	<i>momentum</i> $cp = Mc^2 \sinh \rho$	<i>-Lagrangian</i> $L = -Mc^2 \operatorname{sech} \rho$	<i>Hamiltonian</i> $H = Mc^2 \cosh \rho$	<i>DeBroglie</i> $\lambda = \alpha \operatorname{csch} \rho$	$V_{phase} = c \coth \rho$	

# Lecture 31

## Thur. 12.10.2015

Review of 16 relativity functions of  $\rho$  and related geometric approach to relativity  
Animation of  $e^{\rho=2}$  spacetime and per-spacetime plots

*Rapidity*  $\rho$  related to *stellar aberration angle*  $\sigma$  and L. C. Epstein's approach to relativity

Longitudinal hyperbolic  $\rho$ -geometry connects to transverse circular  $\sigma$ -geometry

“Occams Sword” and summary of 16 parameter functions of  $\rho$  and  $\sigma$

Applications to optical waveguide, spherical waves, and accelerator radiation

*Learning about sin! and COS and...*

Derivation of relativistic quantum mechanics

- ➔ What's the matter with mass? Shining some light on the Elephant in the room
- Relativistic action and Lagrangian-Hamiltonian relations
- Poincare' and Hamilton-Jacobi equations

Relativistic optical transitions and Compton recoil formulae

Feynman diagram geometry

Compton recoil related to rocket velocity formula

Comparing 2<sup>nd</sup>-quantization “photon” number  $N$  and 1<sup>st</sup>-quantization wavenumber  $\kappa$

*Relativity* in accelerated frames

Laser up-tuning by Alice and down-tuning by Carla makes  $g$ -acceleration grid

Analysis of constant- $g$  grid compared to zero- $g$  Minkowski grid

Animation of mechanics and metrology of constant- $g$  grid

# Definition(s) of mass for relativity/quantum

Given: Energy:  $E = Mc^2 \cosh \rho$

$$= h\nu_{\text{phase}}$$

momentum:  $cp = Mc^2 \sinh \rho$

$$= hc\kappa_{\text{phase}}$$

velocity:  $u = c \tanh \rho = \frac{d\nu}{d\kappa}$

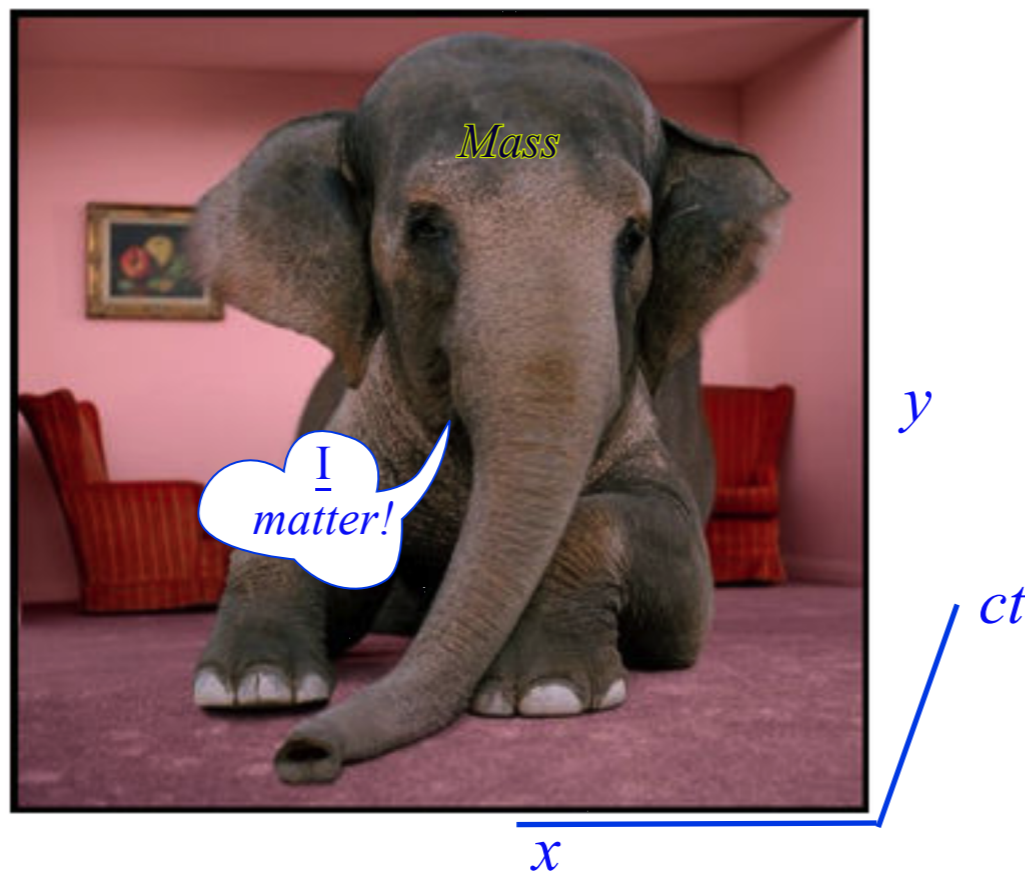
Rest Mass  $M_{\text{rest}}$  (Einstein's mass)

$$hB = h\nu_A = Mc^2 = hc\kappa_A$$

Defines invariant hyperbola(s)

$$E = \pm \sqrt{(Mc^2)^2 + (cp)^2}$$

- *What's the matter with Mass?*



*Shining some light on the elephant in the spacetime room*



# Definition(s) of mass for relativity/quantum

Given: Energy:  $E = Mc^2 \cosh \rho$

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$$\frac{h\nu_{phase}}{c^2} = M_{rest} = \frac{hc\kappa_{phase}}{c^2} \quad \begin{array}{l} \text{Rest} \\ \text{Mass} \end{array}$$

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Defines invariant hyperbola(s)

momentum:  $cp = Mc^2 \sinh \rho$

$$hB = h\nu_A = Mc^2 = hcK_A$$

$$E = \pm \sqrt{(Mc^2)^2 + (cp)^2}$$

$$= hcK_{phase}$$

$$\frac{h\nu_{phase}}{c^2} = M_{rest} = \frac{hcK_{phase}}{c^2} \quad \frac{\text{Rest}}{\text{Mass}}$$

velocity:  $u = c \tanh \rho = \frac{d\nu}{dK}$

Momentum Mass  $M_{mom}$  (Galileo's mass) Defined by ratio  $p/u$  of relativistic momentum to group velocity.

$$M_{mom} \equiv \frac{p}{u} = \frac{M_{rest} c \sinh \rho}{c \tanh \rho}$$

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$$= h\nu_{phase}$$

momentum:  $cp = Mc^2 \sinh \rho$

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velocity:  $u = c \tanh \rho = \frac{d\nu}{d\kappa}$

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Rest Mass  $M_{rest}$  (Einstein's mass)

$$h\nu_A = hc\kappa_A = Mc^2$$

$$\frac{h\nu_{phase}}{c^2} = M_{rest} = \frac{hc\kappa_{phase}}{c^2} \quad \frac{\text{Rest}}{\text{Mass}}$$

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Limiting cases:  $M_{mom} \xrightarrow{u \rightarrow c} M_{rest} e^{\rho} / 2$

$$M_{mom} \xrightarrow{u \ll c} M_{rest}$$

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Limiting cases:  $M_{eff} \xrightarrow{u \rightarrow c} M_{rest} e^{3\rho/2}$

$$M_{eff} \xrightarrow{u \ll c} M_{rest}$$



# Definition(s) of mass for relativity/quantum

Given: Energy:  $E = Mc^2 \cosh \rho$

$= h\nu_{phase}$

momentum:  $cp = Mc^2 \sinh \rho$

$= h c k_{phase}$

velocity:  $u = c \tanh \rho = \frac{d\nu}{dk}$

Defines invariant hyperbola(s)

$$E = \pm \sqrt{(Mc^2)^2 + (cp)^2}$$

Rest Mass  $M_{rest}$  (Einstein's mass)

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$M_{eff} \xrightarrow{u \ll c} M_{rest}$

More common derivation using group velocity:  $u \equiv V_{group} = \frac{d\omega}{dk} = \frac{d\nu}{dk}$

$$M_{eff} \equiv \frac{dp}{du} = \frac{\hbar dk}{dV_{group}} = \frac{\hbar}{\frac{d}{dk} \frac{d\omega}{dk}} = \frac{\hbar}{\frac{d^2 \omega}{dk^2}} = \frac{M_{rest}}{(1 - u^2/c^2)^{3/2}}$$

# Definition(s) of mass for relativity/quantum

Given: Energy:  $E = Mc^2 \cosh \rho$   
 $= h\nu_{phase}$

Rest Mass  $M_{rest}$  (Einstein's mass)

Defines invariant hyperbola(s)

momentum:  $cp = Mc^2 \sinh \rho$   
 $= hck_{phase}$

$$hB = h\nu_A = Mc^2 = hck_A$$

$$E = \pm \sqrt{(Mc^2)^2 + (cp)^2}$$

Group velocity:  $u = c \tanh \rho = \frac{d\nu}{dk}$

$$\frac{h\nu_{phase}}{c^2} = M_{rest} = \frac{hck_{phase}}{c^2} \quad \text{Rest Mass}$$

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$$M_{mom} \equiv \frac{p}{u} = \frac{M_{rest} c \sinh \rho}{c \tanh \rho}$$

Limiting cases:  $M_{mom} \xrightarrow{u \rightarrow c} M_{rest} e^{\rho/2}$

$$M_{mom} \xrightarrow{u \ll c} M_{rest}$$

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That is ratio of change  $dp = Mc \cosh \rho d\rho$  in momentum to change  $du = c \operatorname{sech}^2 \rho d\rho$  in velocity

$$M_{eff} \equiv \frac{dp}{du} = M_{rest} \frac{c \cosh \rho}{c \operatorname{sech}^2 \rho} = M_{rest} \cosh^3 \rho \quad \text{Effective Mass}$$

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general wave formula

to accompany

$$V_{group} = \frac{d\omega}{dk}$$

# Definition(s) of mass for relativity/quantum

## How much mass does a $\gamma$ -photon have?

Rest Mass (a)  $\gamma$ -rest mass:  $M_{rest}^{\gamma} = 0$ ,

Momentum Mass (b)  $\gamma$ -momentum mass:  $M_{mom}^{\gamma} = \frac{p}{c} = \frac{h\kappa}{c} = \frac{h\nu}{c^2}$ ,

Effective Mass (c)  $\gamma$ -effective mass:  $M_{eff}^{\gamma} = \infty$ .

Newton complained about his “corpuscles” of light having “fits” (going crazy).

(All this would be evidence of *triple Schizophrenia*.)

$$M_{mom}^{\gamma} = \frac{h\nu}{c^2} = \nu(1.2 \cdot 10^{-51}) \text{kg} \cdot \text{s} = 4.5 \cdot 10^{-36} \text{kg} \quad (\text{for: } \nu=600\text{THz})$$

# Lecture 31

Thur. 12.10.2015

Review of 16 relativity functions of  $\rho$  and related geometric approach to relativity  
Animation of  $e^{\rho=2}$  spacetime and per-spacetime plots

*Rapidity*  $\rho$  related to *stellar aberration angle*  $\sigma$  and L. C. Epstein's approach to relativity

Longitudinal hyperbolic  $\rho$ -geometry connects to transverse circular  $\sigma$ -geometry

“Occams Sword” and summary of 16 parameter functions of  $\rho$  and  $\sigma$

Applications to optical waveguide, spherical waves, and accelerator radiation

*Learning about sin! and COS and...*

**Derivation of relativistic quantum mechanics**

What's the matter with mass? Shining some light on the Elephant in the room

➔ **Relativistic action and Lagrangian-Hamiltonian relations**

Poincare' and Hamilton-Jacobi equations

Relativistic optical transitions and Compton recoil formulae

Feynman diagram geometry

Compton recoil related to rocket velocity formula

Comparing 2<sup>nd</sup>-quantization “photon” number  $N$  and 1<sup>st</sup>-quantization wavenumber  $\kappa$

*Relativity* in accelerated frames

Laser up-tuning by Alice and down-tuning by Carla makes  $g$ -acceleration grid

Analysis of constant- $g$  grid compared to zero- $g$  Minkowski grid

Animation of mechanics and metrology of constant- $g$  grid

# Relativistic action $S$ and Lagrangian-Hamiltonian relations

Define *Lagrangian*  $L$  using invariant wave phase  $\Phi = kx - \omega t = k'x' - \omega't'$  for wave of  $k = k_{phase}$  and  $\omega = \omega_{phase}$ .

$$L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega$$

$$\hbar \equiv \frac{h}{2\pi}$$

$$\begin{aligned} h\nu_A &= Mc^2 = hc\kappa_A \\ h\nu_{phase} &= E = h\nu_A \cosh \rho \\ hc\kappa_{phase} &= cp = h\nu_A \sinh \rho \end{aligned}$$

Prior wave relations

← linear Hz  
format

angular phasor →  
format

$$\begin{aligned} \hbar\omega_A &= Mc^2 = \hbar ck_A \\ \hbar\omega_{phase} &= E = \hbar\omega_A \cosh \rho \\ \hbar ck_{phase} &= cp = \hbar\omega_A \sinh \rho \end{aligned}$$

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Use DeBroglie-momentum  $p = \hbar k$  relation and Planck-energy  $E = \hbar \omega$  relation

$$L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega$$

$$\hbar \equiv \frac{h}{2\pi}$$

$$p = \hbar k = Mc \sinh \rho$$

$$E = \hbar \omega = Mc^2 \cosh \rho$$

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$$p = \hbar k = Mc \sinh \rho$$

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Prior wave relations

← linear Hz format      angular phasor →  
format                      format

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Use DeBroglie-momentum  $p = \hbar k$  relation and Planck-energy  $E = \hbar \omega$  relation to define *Hamiltonian*  $H = E$

$$L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E \equiv p\dot{x} - E \equiv pu - H = L \quad \text{Legendre transformation}$$

$$p = \hbar k = Mc \sinh \rho$$

$$E = \hbar \omega = Mc^2 \cosh \rho = H$$

$$\begin{aligned} h\nu_A &= Mc^2 = hc\kappa_A \\ h\nu_{phase} &= E = h\nu_A \cosh \rho \\ hc\kappa_{phase} &= cp = h\nu_A \sinh \rho \end{aligned}$$

Prior wave relations

← linear Hz format      angular phasor →  
format                      format

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*Legendre transformation*

Use *Group velocity* :  $u = \frac{dx}{dt} = c \tanh \rho$

$$p = \hbar k = Mc \sinh \rho$$

$$E = \hbar \omega = Mc^2 \cosh \rho = H$$

$$\begin{aligned} h\nu_A &= Mc^2 = hc\kappa_A \\ h\nu_{phase} &= E = h\nu_A \cosh \rho \\ hc\kappa_{phase} &= cp = h\nu_A \sinh \rho \end{aligned}$$

Prior wave relations

← linear Hz format      angular phasor →  
format                      format

$$\begin{aligned} \hbar \omega_A &= Mc^2 = \hbar c k_A \\ \hbar \omega_{phase} &= E = \hbar \omega_A \cosh \rho \\ \hbar c k_{phase} &= cp = \hbar \omega_A \sinh \rho \end{aligned}$$

$$\hbar \equiv \frac{h}{2\pi}$$

# Relativistic **action S** and Lagrangian-Hamiltonian relations

Define *Lagrangian L* using invariant wave phase  $\Phi = kx - \omega t = k'x' - \omega't'$  for wave of  $k = k_{phase}$  and  $\omega = \omega_{phase}$ .

Use DeBroglie-momentum  $p = \hbar k$  relation and Planck-energy  $E = \hbar \omega$  relation to define *Hamiltonian H = E*

$$L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E \equiv p \dot{x} - E \equiv pu - H = L$$

Legendre transformation

Use Group velocity :  $u = \frac{dx}{dt} = c \tanh \rho$

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$$L \text{ is : } Mc^2 \frac{-1}{\cosh \rho} = -Mc^2 \operatorname{sech} \rho$$

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Note:  $Mc u = Mc^2 \tanh \rho$

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Compare *Lagrangian*  $L$

$$L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho$$

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Compare *Lagrangian*  $L$

$$L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho$$

with *Hamiltonian*  $H = E$

$$H = \hbar \omega = Mc^2 / \sqrt{1 - \frac{u^2}{c^2}} = Mc^2 \cosh \rho$$

$$\hbar \omega_A = Mc^2 = \hbar c k_A$$

Prior wave relations

$$\hbar \omega_{phase} = E = \hbar \omega_A \cosh \rho$$

← linear Hz format

angular phasor →

$$\hbar c k_{phase} = cp = \hbar \omega_A \sinh \rho$$

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Also:  $cp = Mc^2 \sinh \rho$

Compare Lagrangian  $L$

$$L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho$$

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$$H = \hbar\omega = Mc^2 / \sqrt{1 - \frac{u^2}{c^2}} = Mc^2 \cosh \rho$$

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Prior wave relations  $\hbar = h/2\pi$

$$\hbar\omega_{phase} = E = \hbar\omega_A \cosh \rho$$

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Note:  $Mc u = Mc^2 \tanh \rho = Mc^2 \sin \sigma$   
 Also:  $cp = Mc^2 \sinh \rho = \hbar ck = Mc^2 \tan \sigma$

Compare Lagrangian L

$$\dot{S} = L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho = -Mc^2 \cos \sigma$$

with Hamiltonian H=E

$$H = \hbar\omega = Mc^2 / \sqrt{1 - \frac{u^2}{c^2}} = Mc^2 \cosh \rho = Mc^2 \sec \sigma$$

$$= Mc^2 \sqrt{1 + \sinh^2 \rho} = Mc^2 \sqrt{1 + (cp)^2}$$

Including stellar angle  $\sigma$

Define Action  $S = \hbar\Phi$

$$\hbar\omega_A = Mc^2 = \hbar ck_A$$

$$\hbar\omega_{phase} = E = \hbar\omega_A \cosh \rho$$

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Prior wave relations

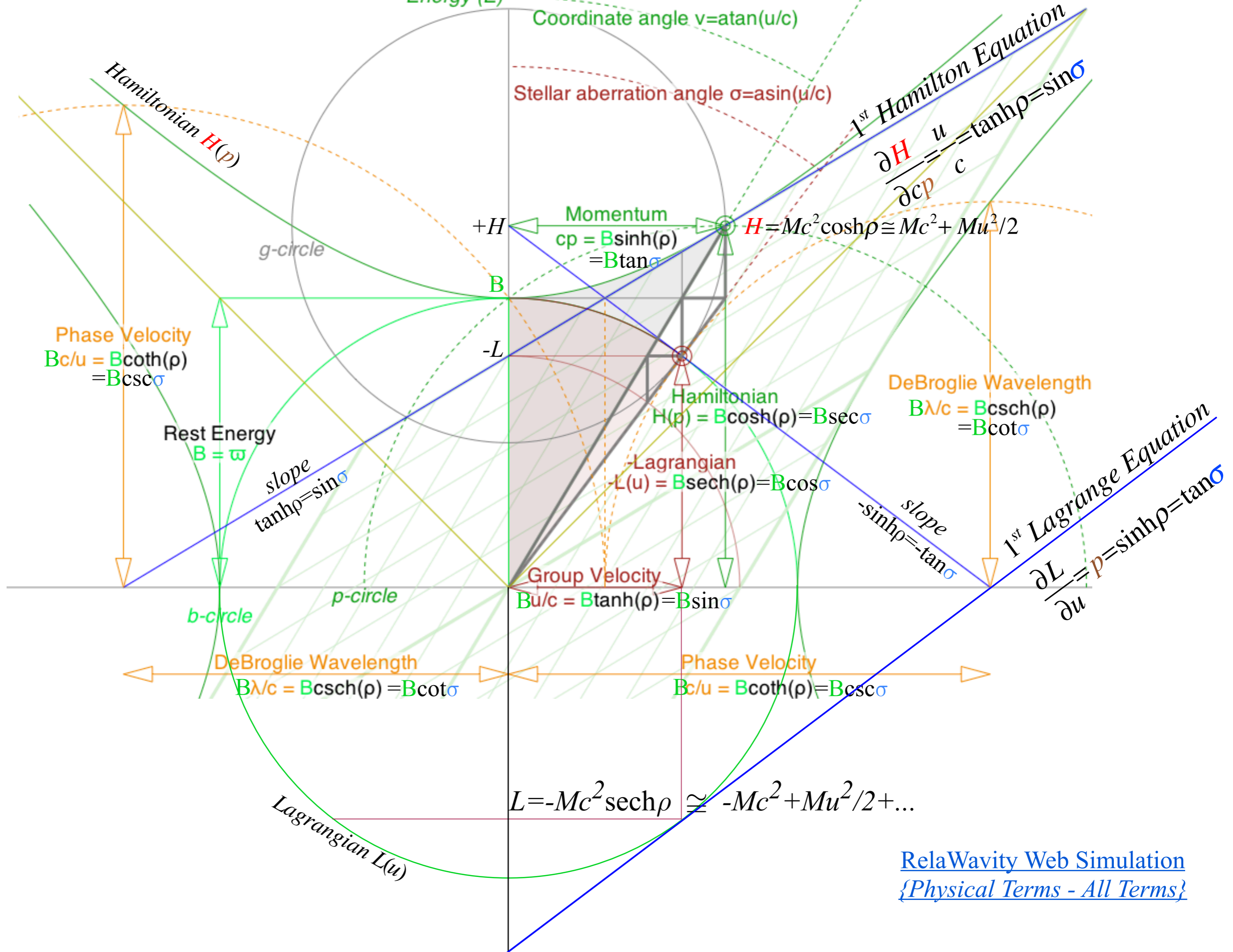
← linear Hz format      angular phasor format →

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Compare *Lagrangian L*

$$\dot{S} = L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho = -Mc^2 \cos \sigma$$

with *Hamiltonian H = E*

$$H = \hbar \omega = Mc^2 / \sqrt{1 - \frac{u^2}{c^2}} = Mc^2 \cosh \rho = Mc^2 \sec \sigma$$

$$= Mc^2 \sqrt{1 + \sinh^2 \rho} = Mc^2 \sqrt{1 + (cp)^2}$$

Define *Action S = \hbar \Phi*

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Prior wave relations

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$$dS \equiv Ldt \equiv \hbar d\Phi = \hbar k dx - \hbar \omega dt = p dx - H dt \quad \text{Poincare Invariant action differential}$$

Compare *Lagrangian L*

$$\dot{S} = L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho = -Mc^2 \cos \sigma$$

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$$\frac{\partial S}{\partial x} = p \quad \frac{\partial S}{\partial t} = -H$$

*Hamilton-Jacobi equations*

Compare *Lagrangian L*

$$\dot{S} = L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho = -Mc^2 \cos \sigma$$

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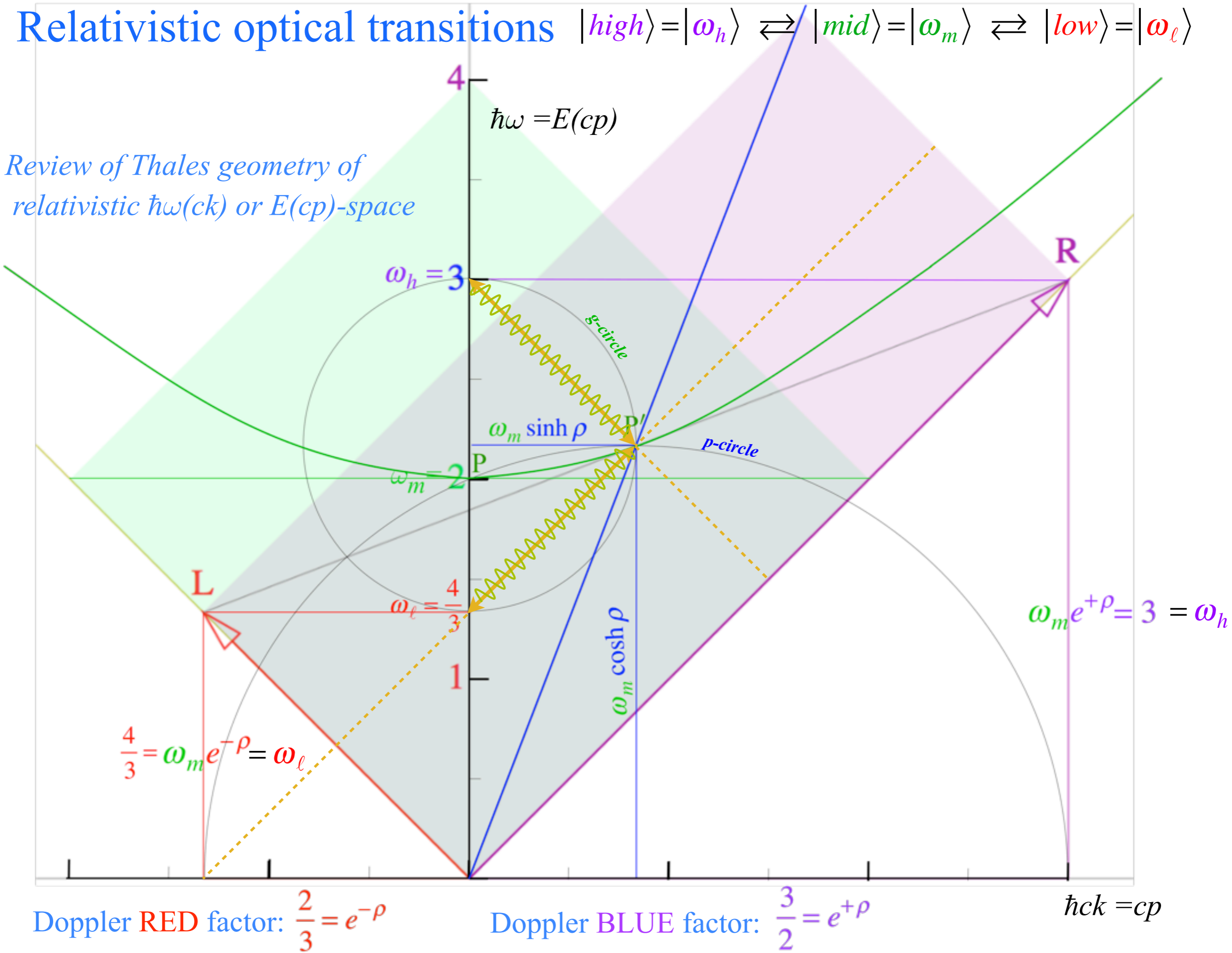
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# Relativistic optical transitions $|high\rangle = |\omega_h\rangle \rightleftharpoons |mid\rangle = |\omega_m\rangle \rightleftharpoons |low\rangle = |\omega_l\rangle$

*Review of Thales geometry of relativistic  $\hbar\omega(ck)$  or  $E(cp)$ -space*

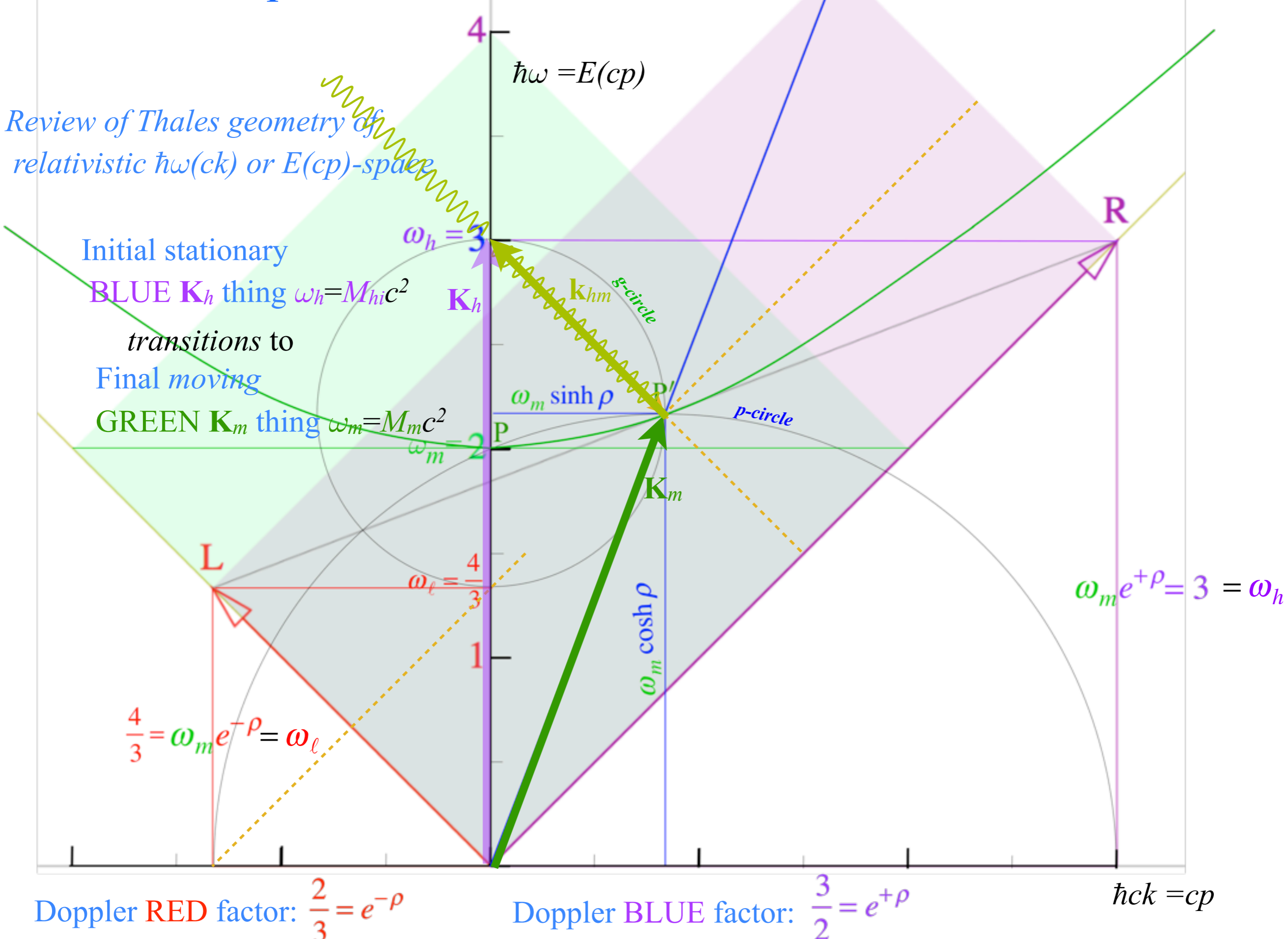








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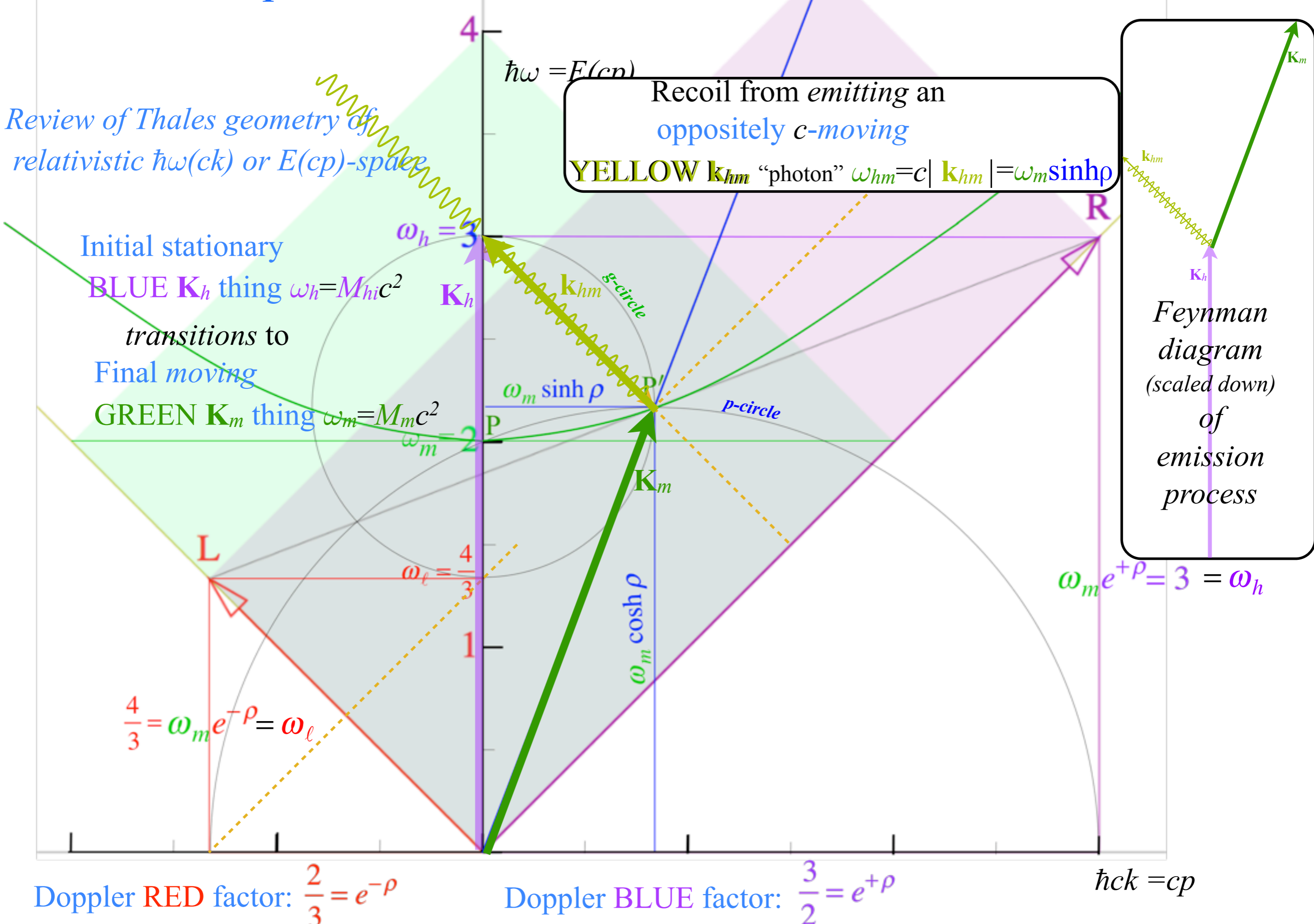
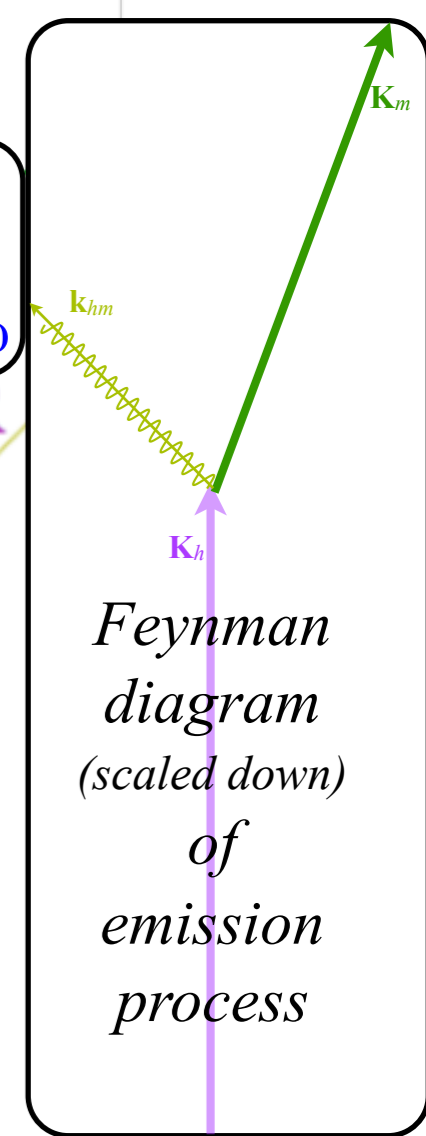
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Initial stationary BLUE  $\mathbf{K}_h$  thing  $\omega_h = M_h c^2$   
 transitions to Final moving GREEN  $\mathbf{K}_m$  thing  $\omega_m = M_m c^2$

Recoil from emitting an oppositely  $c$ -moving YELLOW  $\mathbf{k}_{hm}$  "photon"  $\omega_{hm} = c|\mathbf{k}_{hm}| = \omega_m \sinh \rho$



Doppler RED factor:  $\frac{2}{3} = e^{-\rho}$

Doppler BLUE factor:  $\frac{3}{2} = e^{+\rho}$

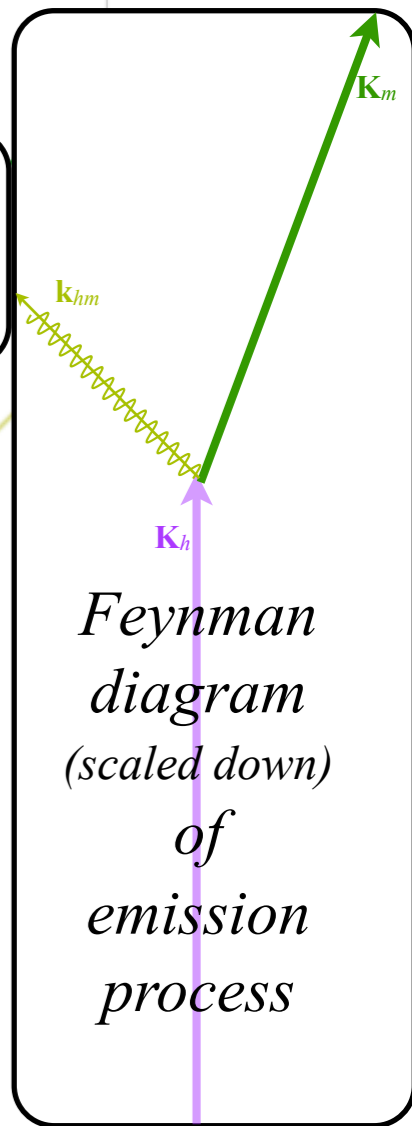
$\hbar ck = cp$

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Take-away point 0  
 Classical (and spectroscopic) Energy-momentum conservation is due to conservation in quantum-phase space-time "wiggle-count"

Doppler RED factor:  $\frac{2}{3} = e^{-\rho}$

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$\hbar ck = cp$

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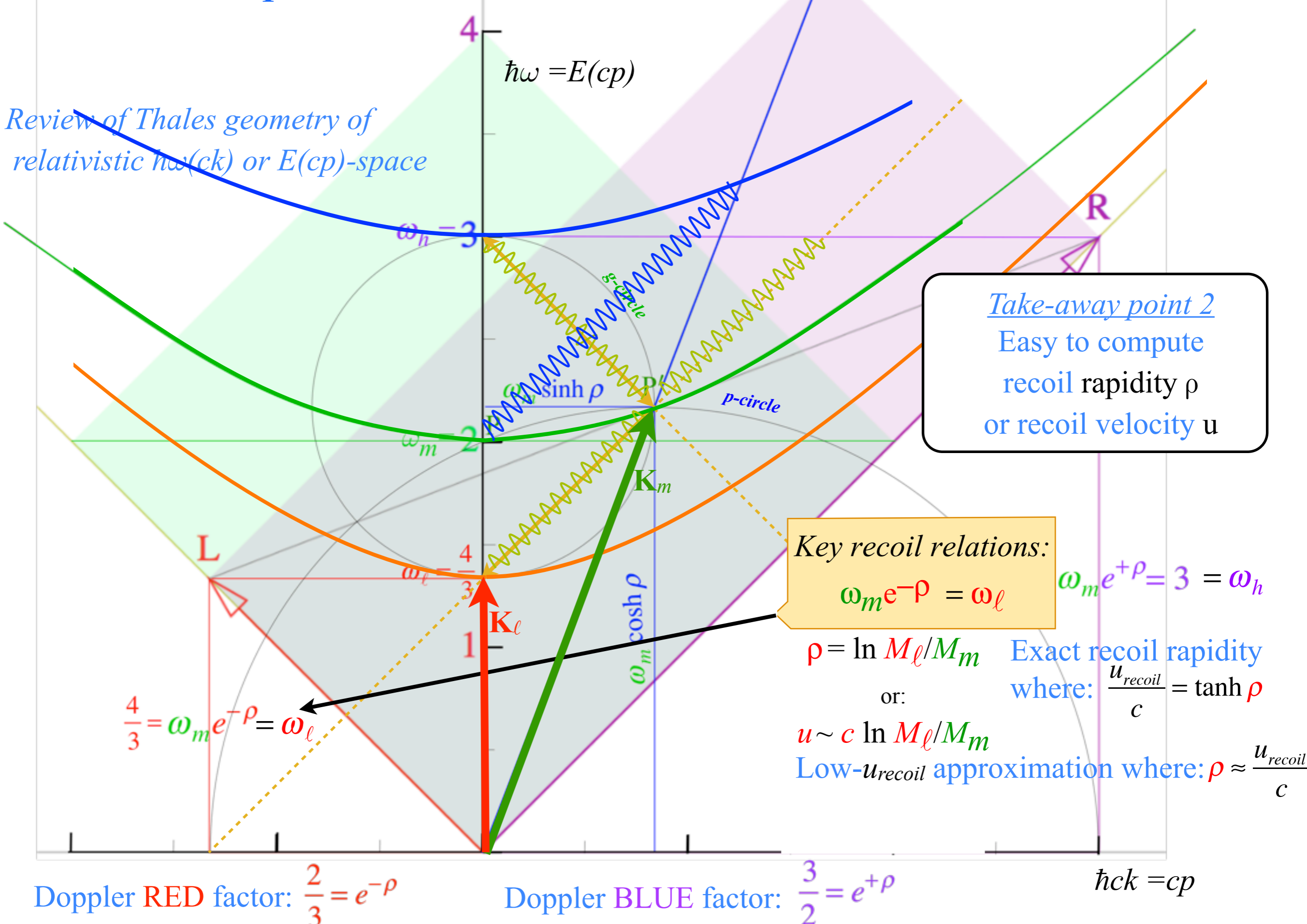
*Relativity* in accelerated frames

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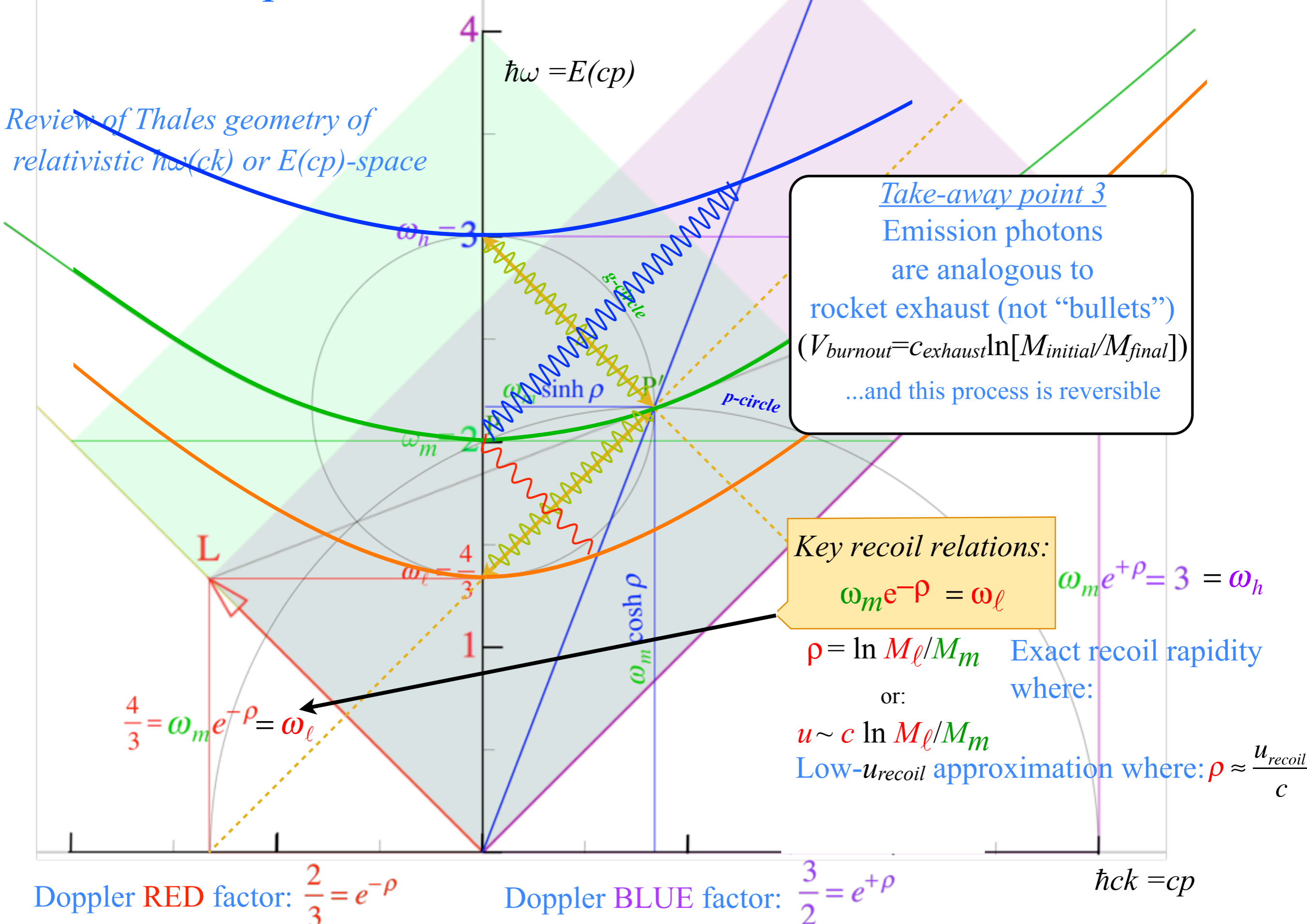
Animation of mechanics and metrology of constant- $g$  grid

Relativistic optical transitions  $|high\rangle = |\omega_h\rangle \rightleftharpoons |mid\rangle = |\omega_m\rangle \rightleftharpoons |low\rangle = |\omega_l\rangle$



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Review of Thales geometry of relativistic  $\hbar\omega(c k)$  or  $E(cp)$ -space



Doppler RED factor:  $\frac{2}{3} = e^{-\rho}$

Doppler BLUE factor:  $\frac{3}{2} = e^{+\rho}$

$\hbar ck = cp$





# Lecture 31

Thur. 12.10.2015

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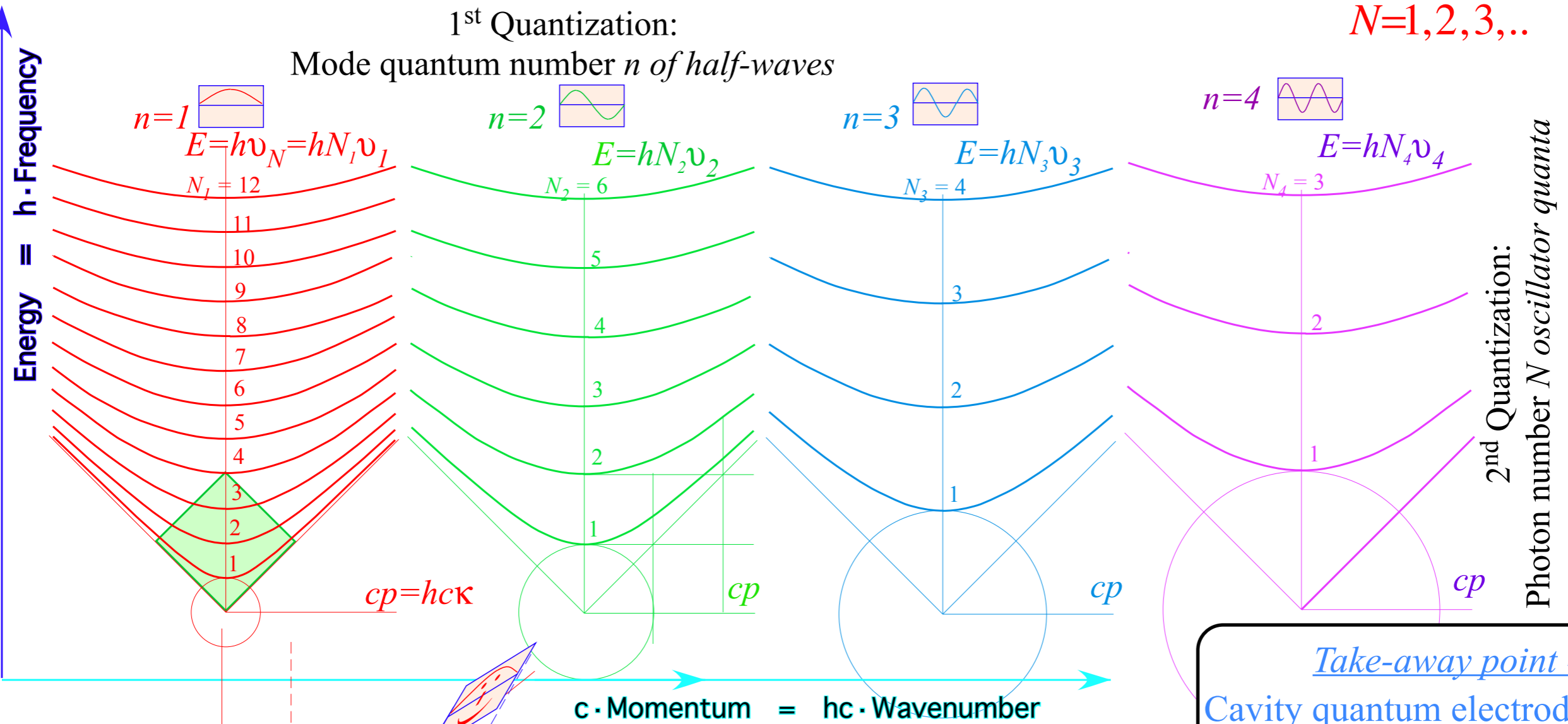
# 2<sup>nd</sup> Quantization: NEWS FLASH!!! $h\nu$ is actually $hN\nu$

(  $h\nu_{phase}=E=h\nu_A \cosh \rho$  ) is actually (  $hN\nu_{phase}=E_N=hN\nu_A \cosh \rho$  with quantum numbers )

$N=1,2,3,..$

1<sup>st</sup> Quantization:

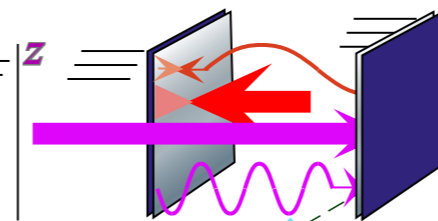
Mode quantum number  $n$  of half-waves



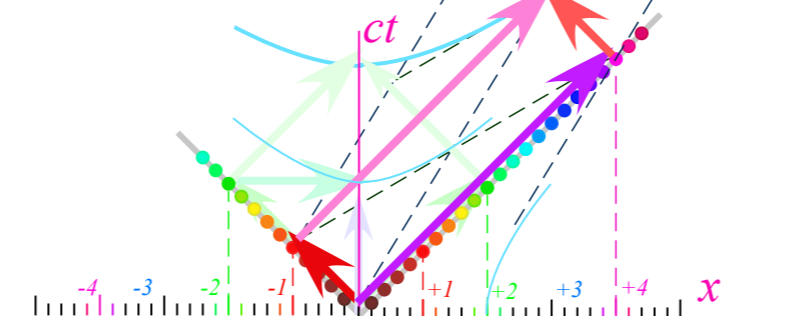
2<sup>nd</sup> Quantization:  
Photon number  $N$  oscillator quanta

Boosted wave mode

Boosted cavity wave has invariant mode number  $n$  photon number  $N_n$



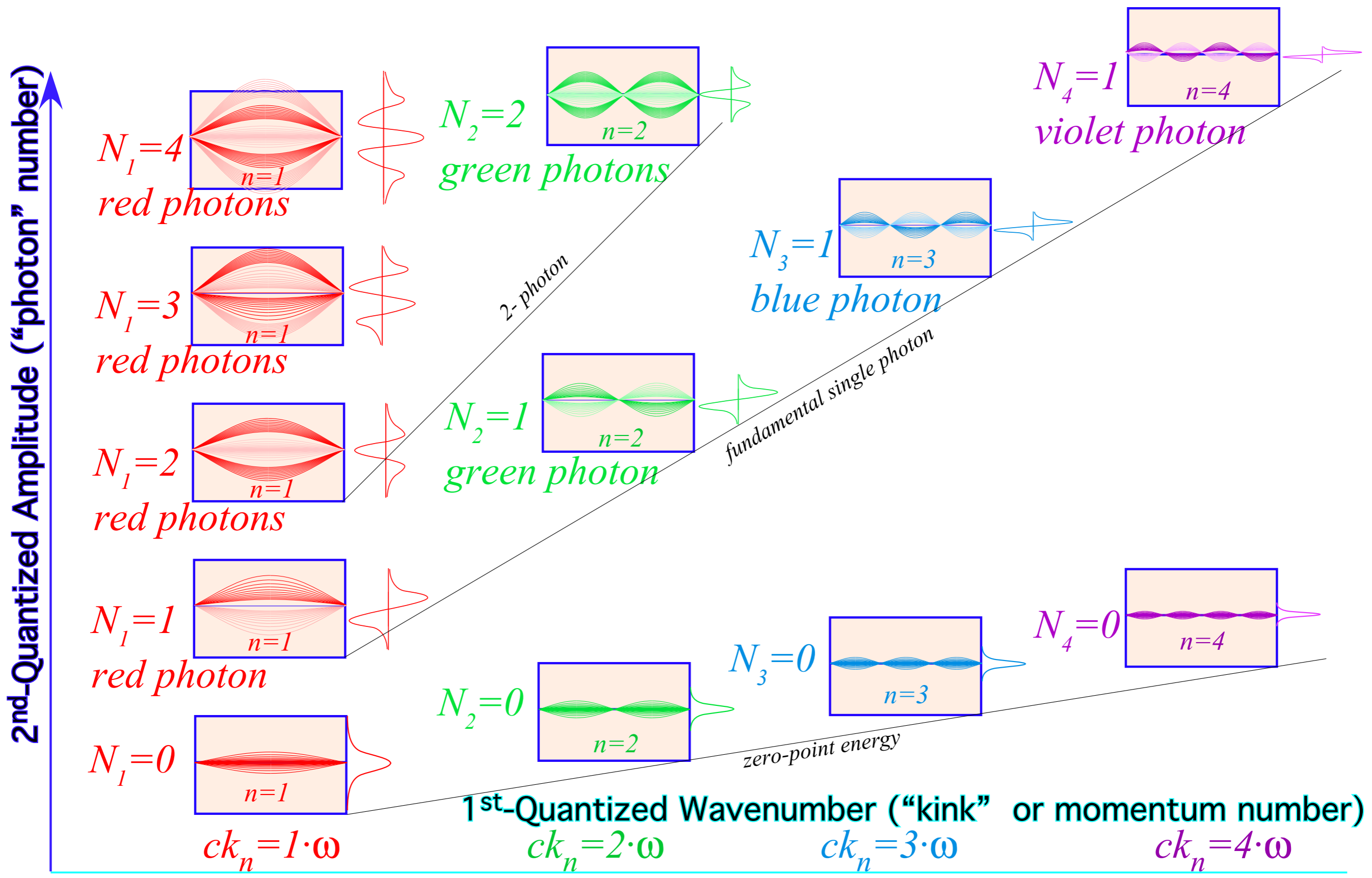
Lorentz contracted cavity length  $L=3.2$   
Proper length  $l=4.0$



Take-away point 4  
Cavity quantum electrodynamics (CQED) and spectra are analogous to molecular rovibronic dynamics with rotation-vibration algebra replaced by Lorentz-Poincare-Dirac algebra (and geometry!)

2<sup>nd</sup> Quantization: NEWS FLASH!!!  $h\nu$  is actually  $hN\nu$

$(h\nu_{phase}=E=h\nu_A \cosh \rho)$  is actually  $(hN\nu_{phase}=E_N=hN\nu_A \cosh \rho \quad (N=1,2,..) )$



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# Acceleration by chirping laser pairs

## Varying acceleration (General case)

From Lect. 35  
ModPhys (2012)

Only green-light is seen by observers on the green accelerated trajectory

Varying local acceleration  $\rho = \rho(\tau)$

$$u = \frac{dx}{dt} = c \tanh(\tau)$$

$$\frac{dt}{d\tau} = \cosh \rho(\tau)$$

$$\frac{dx}{d\tau} = \frac{dx}{dt} \frac{dt}{d\tau} = c \tanh \rho(\tau) \cosh \rho(\tau) = c \sinh \rho(\tau)$$

$$ct = c \int \cosh \rho(\tau) d\tau$$

$$x = c \int \sinh \rho(\tau) d\tau$$

Constant local acceleration  $\rho = \frac{g\tau}{c}$  "Einstein Elevator"

$$ct = c \int \cosh \frac{g\tau}{c} d\tau = \frac{c^2}{g} \sinh \frac{g\tau}{c}$$

$$x = c \int \sinh \frac{g\tau}{c} d\tau = \frac{c^2}{g} \cosh \frac{g\tau}{c}$$

Previous examples involved constant velocity

Constant velocity  $\rho = \rho_0 = \text{const.}$  "Lorentz transformation"

$$ct = c \int \cosh \rho_0 d\tau = c\tau \cosh \rho_0$$

$$x = c \int \sinh \rho_0 d\tau = c\tau \sinh \rho_0$$

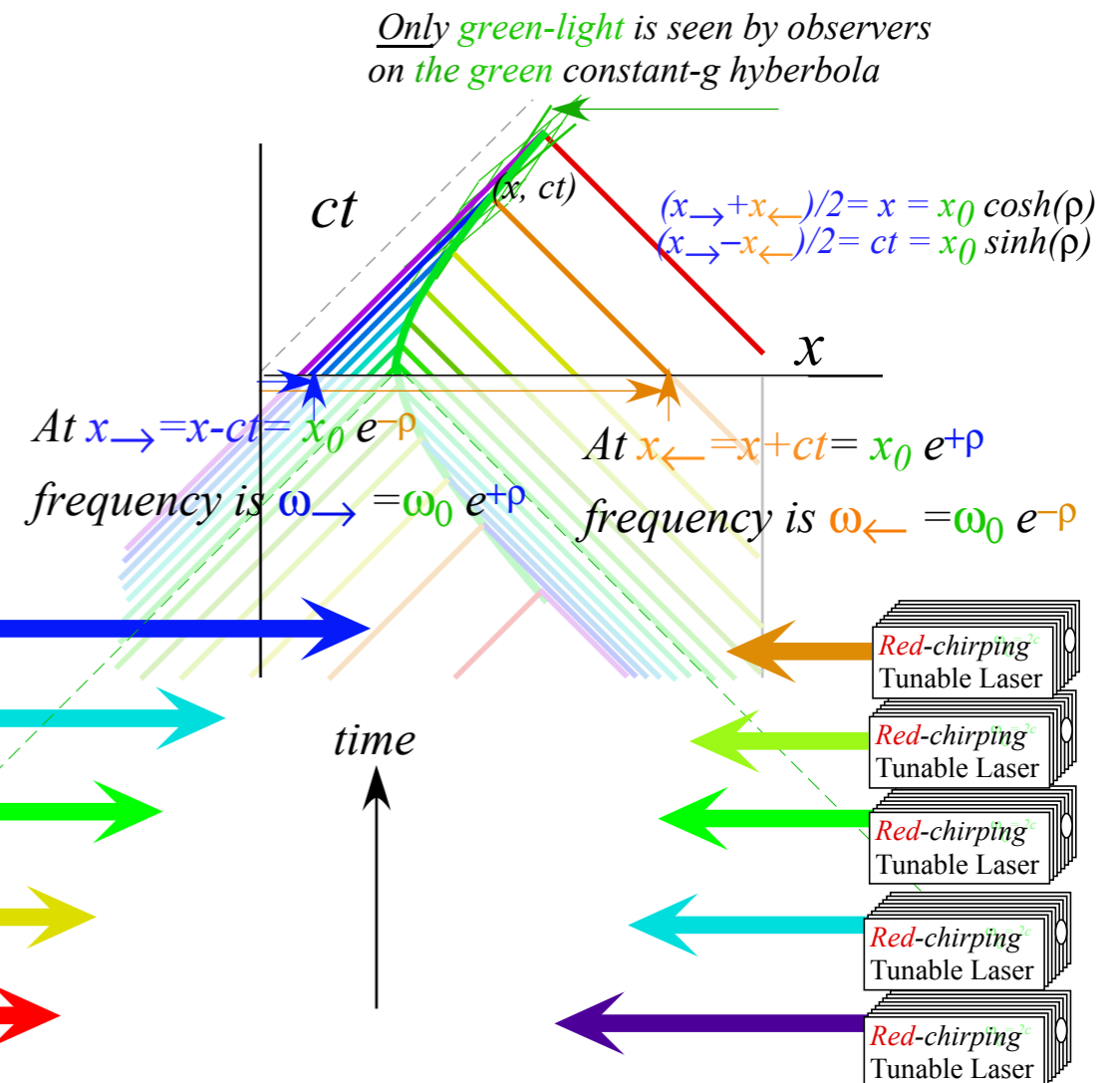
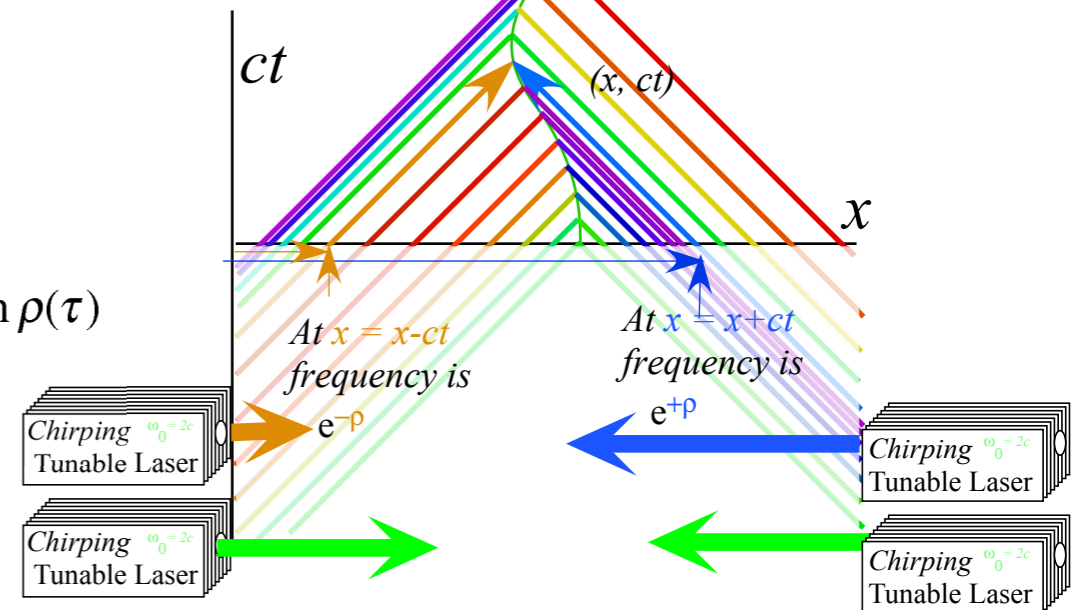
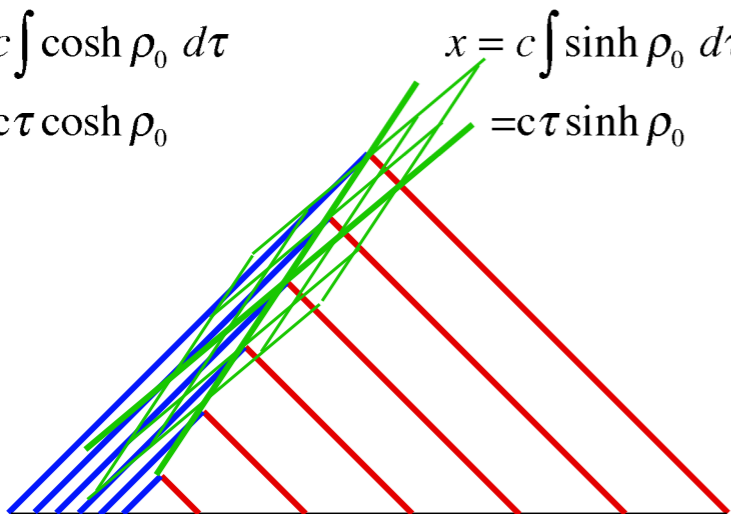
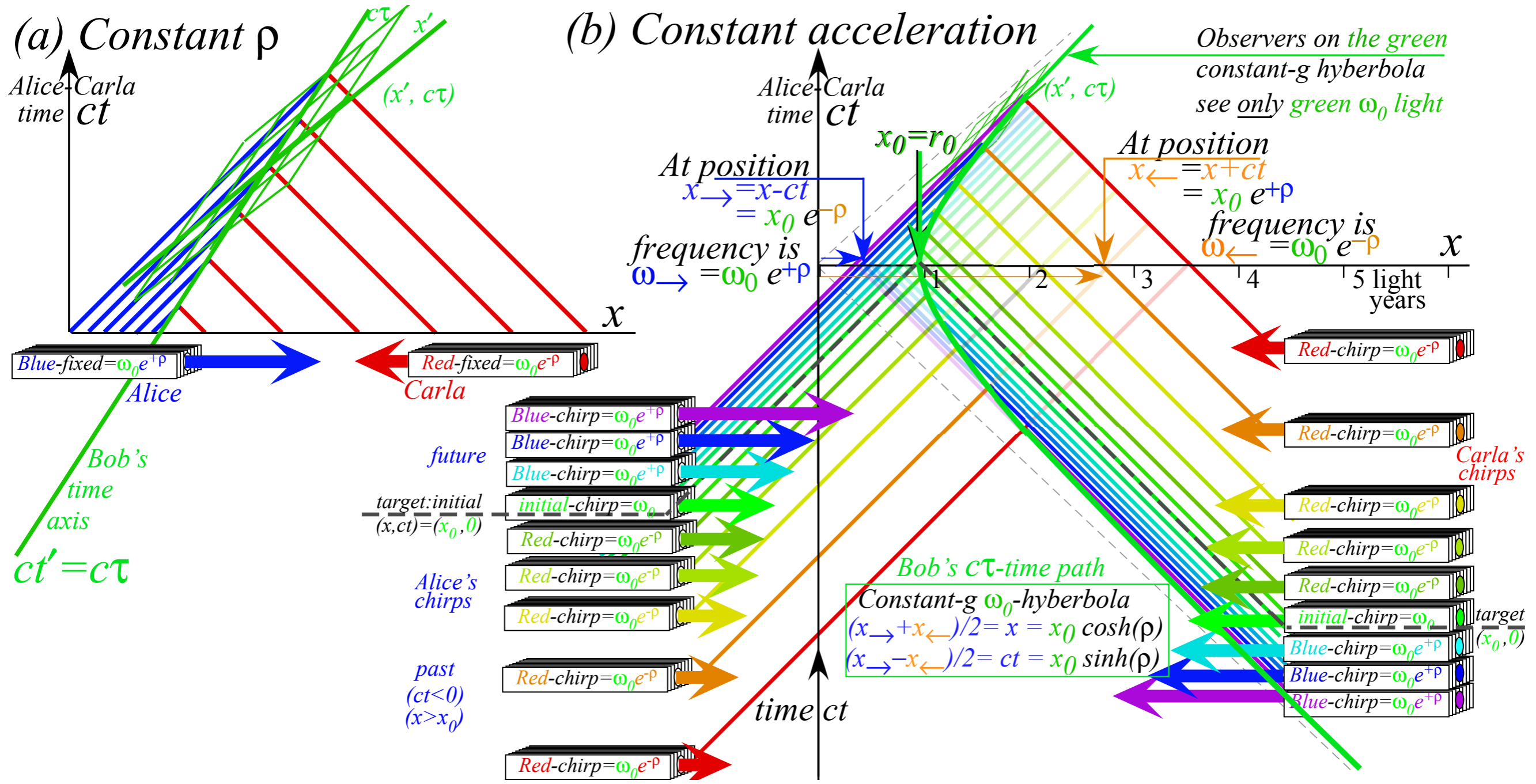
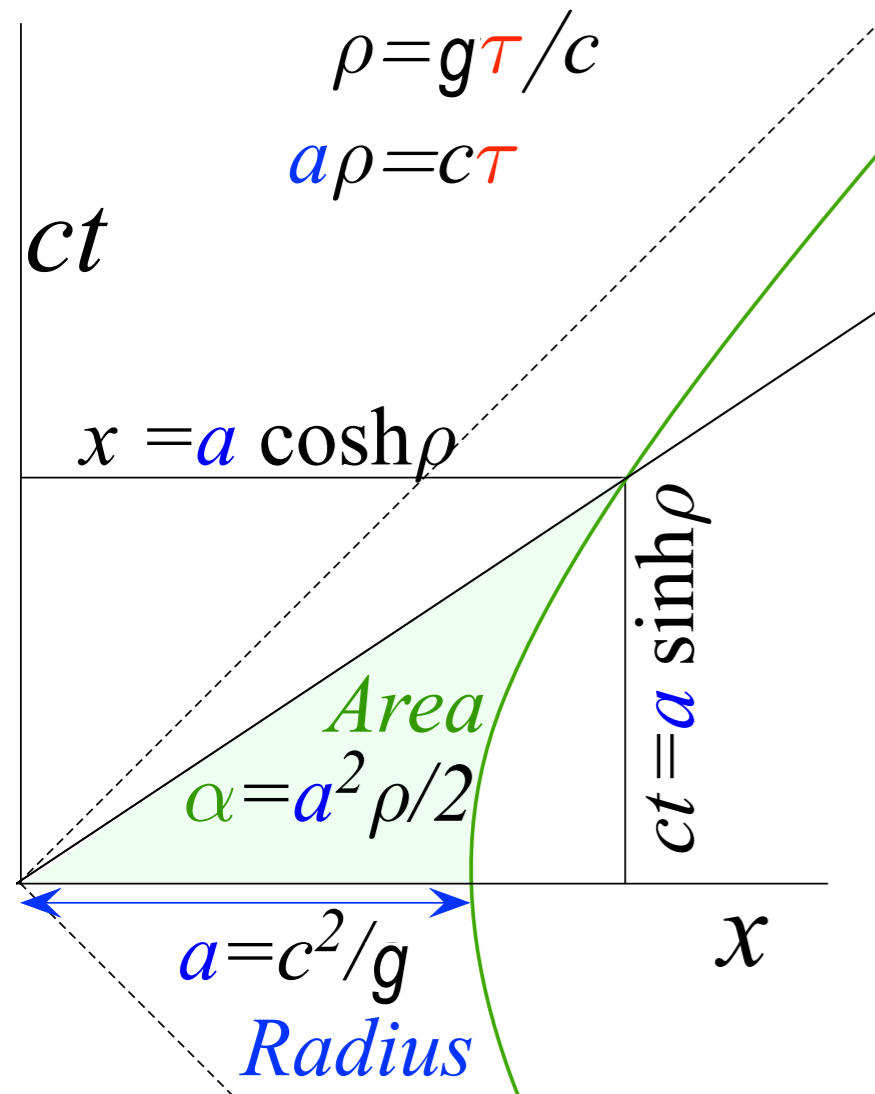


Fig. 8.1 Optical wave frames by red-and-blue-chirped lasers (a)Varying acceleration (b)Constant g



(a) Constant acceleration  $g$   
 Rapidity  $\rho$  vs proper time  $\tau$



(b) Traveler paths of acceleration  $g_q$

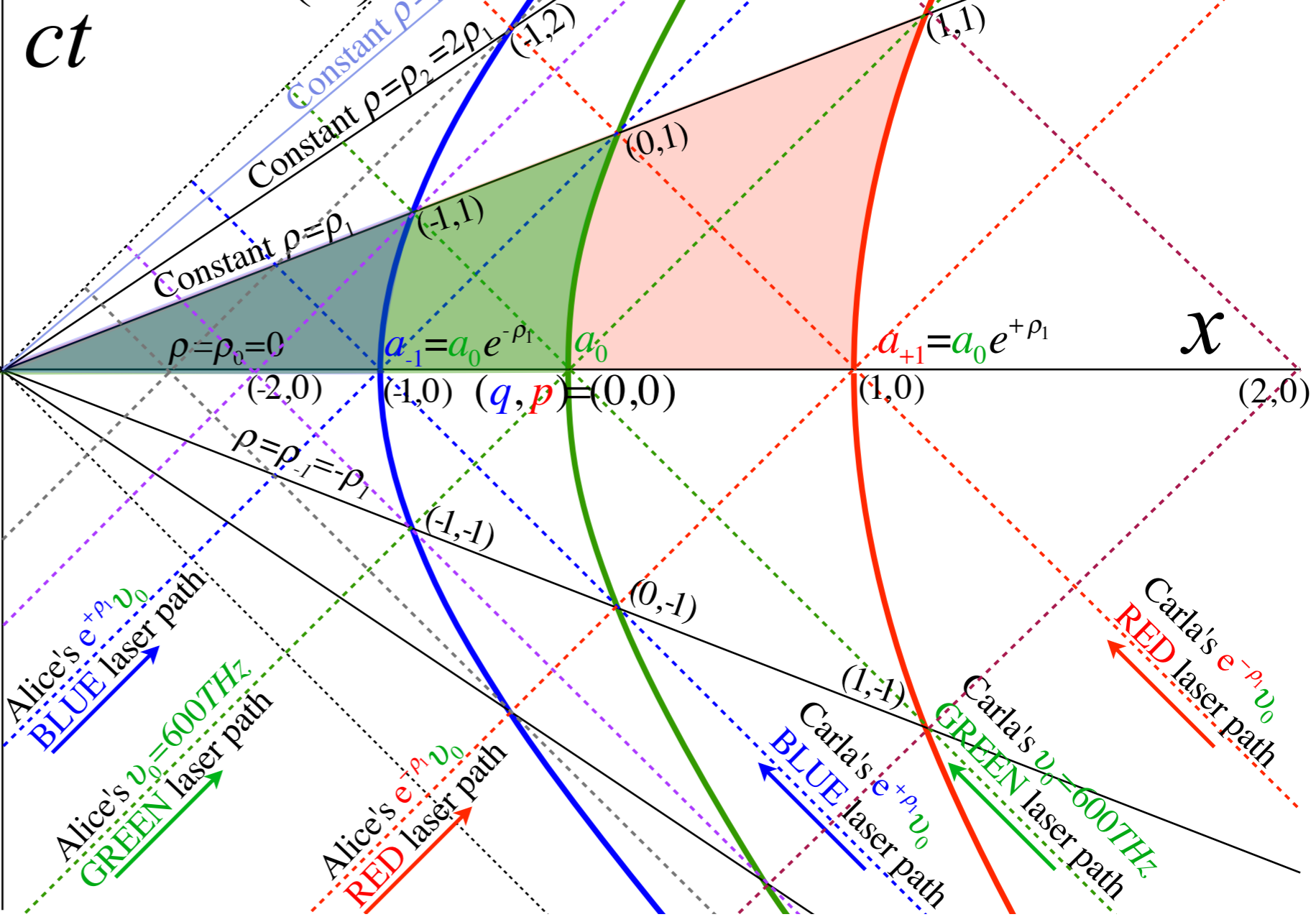
Al:  $g_{-1} = g_0 e^{+\rho_1}$     Bob:  $g_0 = \frac{c^2}{a_0}$     Carl:  $g_{+1} = g_0 e^{-\rho_1}$

Inertial frame coordinates

$(x_{q,p}, ct_{q,p}) =$   
 $a_0 e^{q\rho_1} (\cosh p\rho_1, \sinh p\rho_1)$

Geometric scale:

$e^{q\rho_1} = \left(\frac{3}{2}\right)^q$





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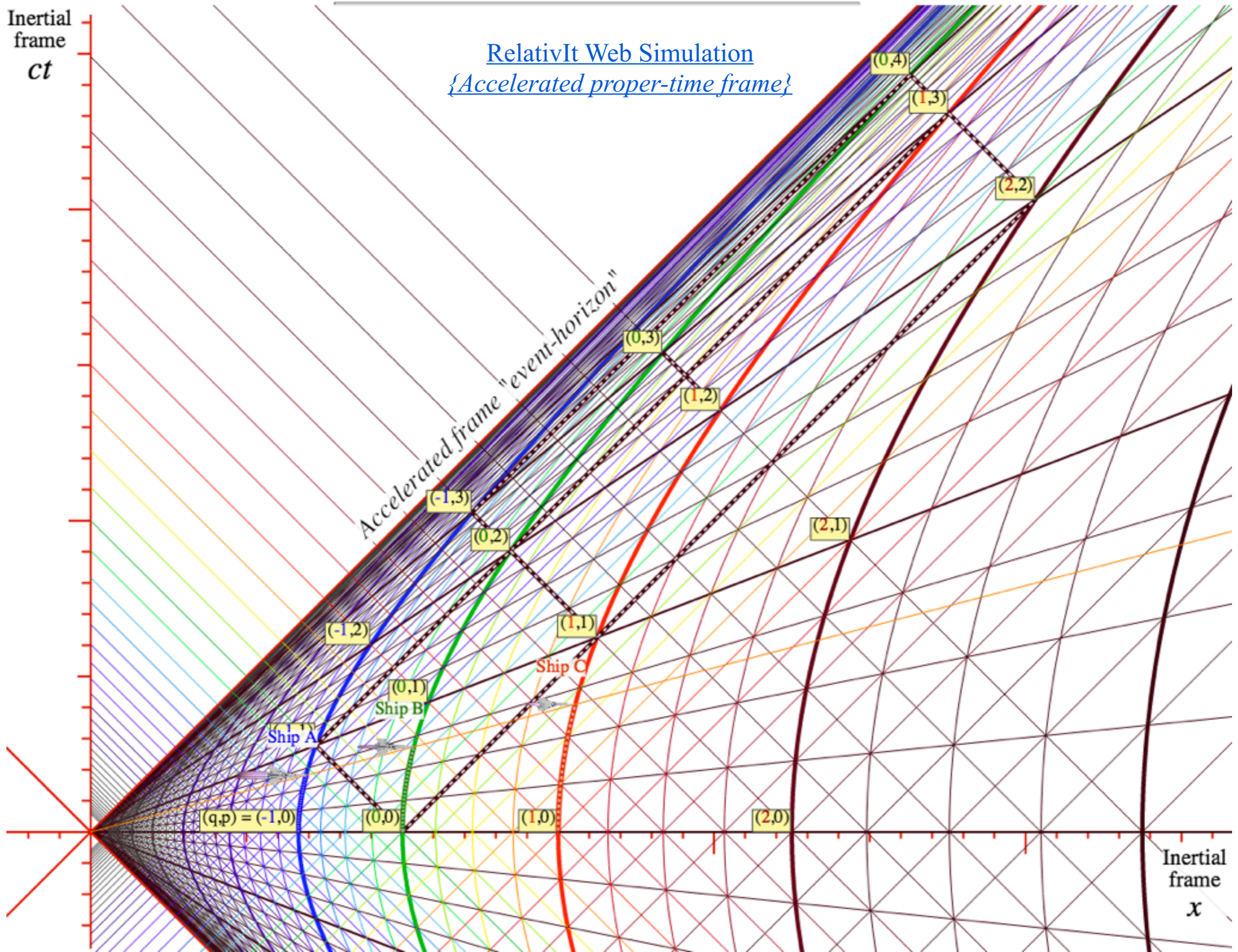
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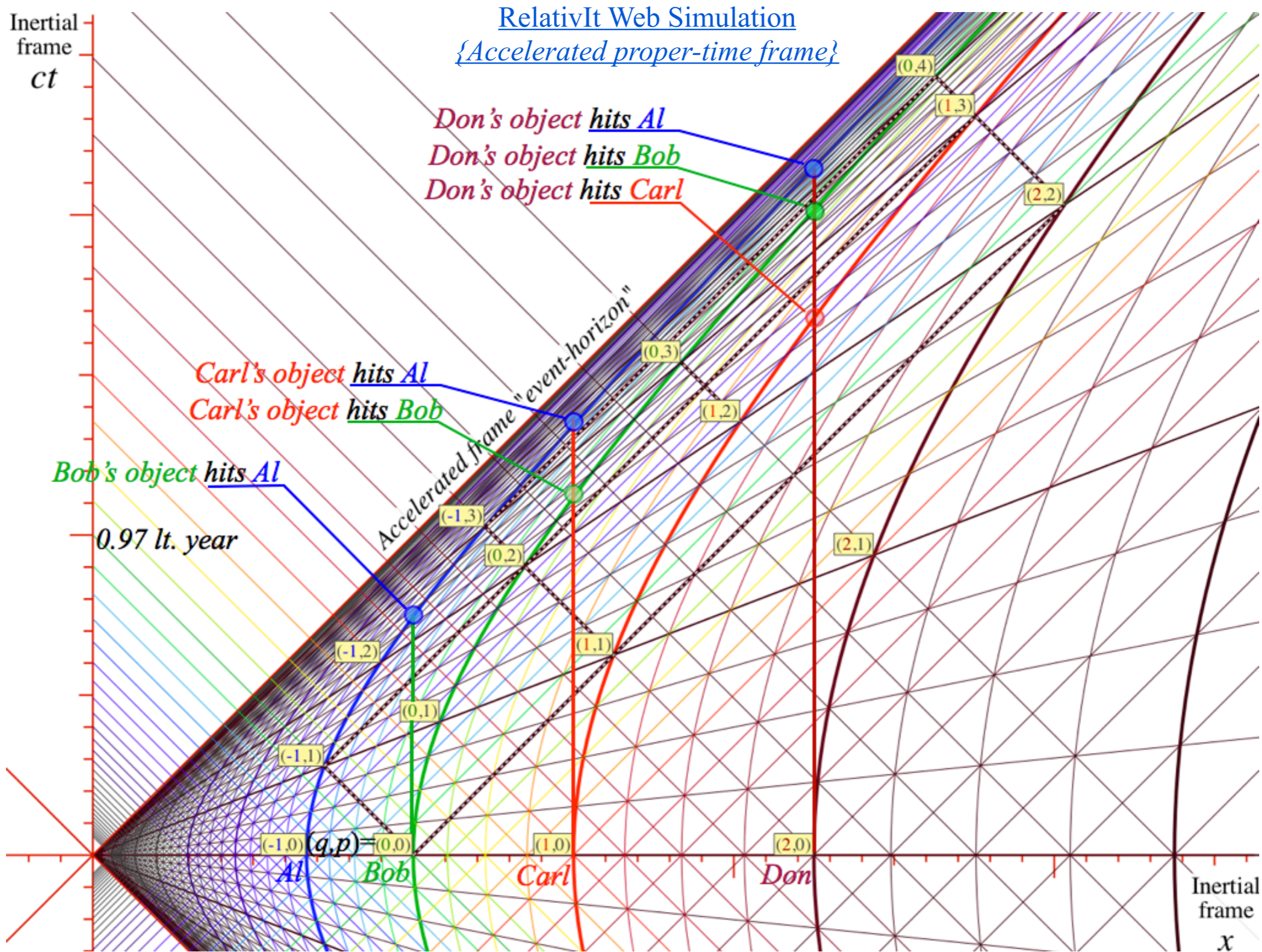
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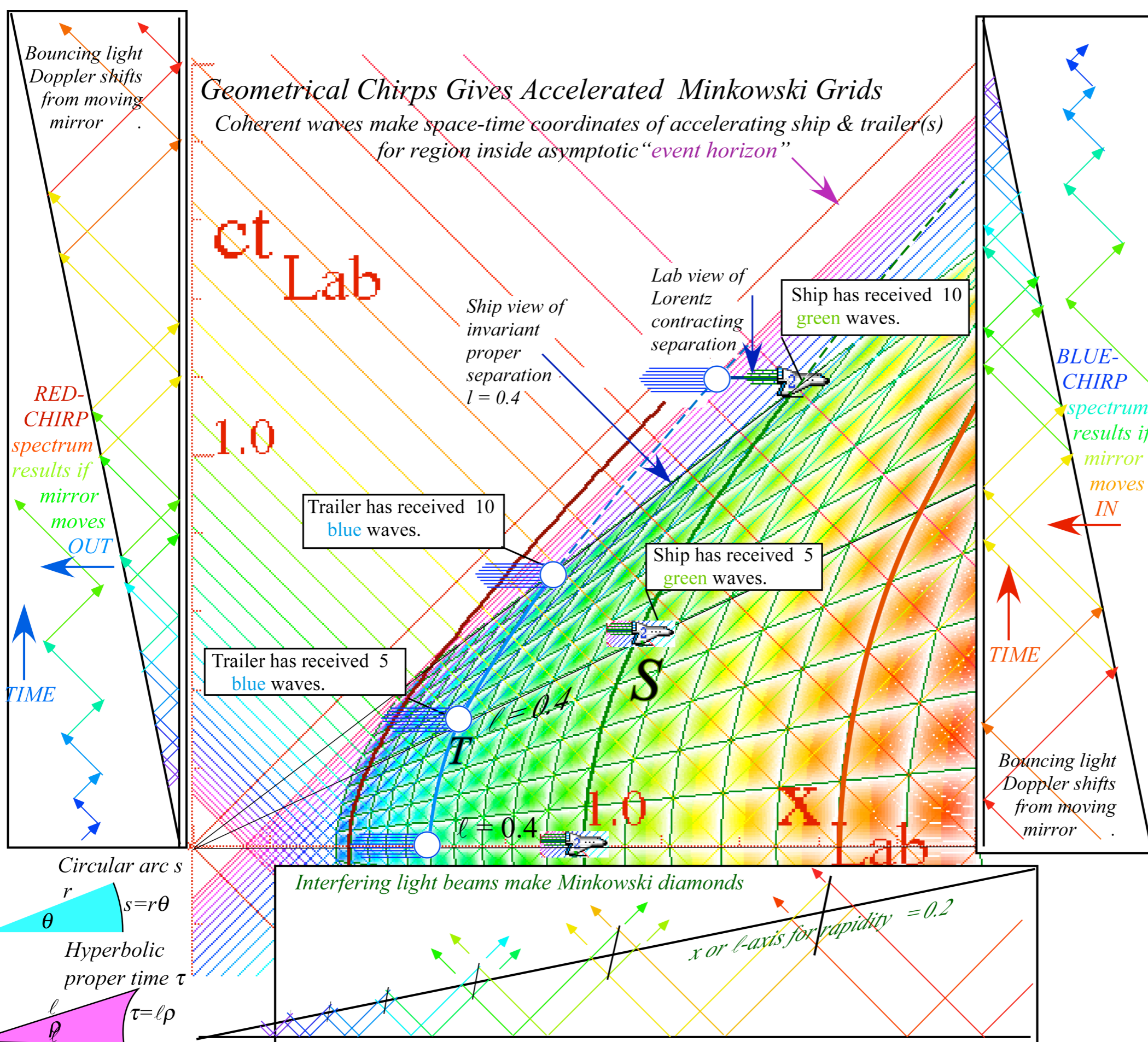


Fig. 8.2 Accelerated reference frames and their trajectories painted by chirped coherent light