

Lecture 7

Thur. 9.17.2013

Geometry and Motion of Isotropic Harmonic Oscillators (Ch. 9 and Ch. 11 of Unit 1)

Geometry of idealized “Sophomore-physics Earth”

Coulomb field outside Isotropic Harmonic Oscillator (IHO) field inside

Contact-geometry of potential curve(s)

“Crushed-Earth” models: 3 key energy “steps” and 4 key energy “levels”

Earth matter vs nuclear matter:

*Introducing the “neutron starlet” and “**Black-Hole-Earth**”*

Isotropic harmonic oscillator dynamics in 1D, 2D, and 3D

Sinusoidal space-time dynamics derived by geometry

Isotropic harmonic oscillator orbits in 1D and 2D (You get 3D for free!)

Constructing 2D Isotropic harmonic oscillator orbits using phasor plots

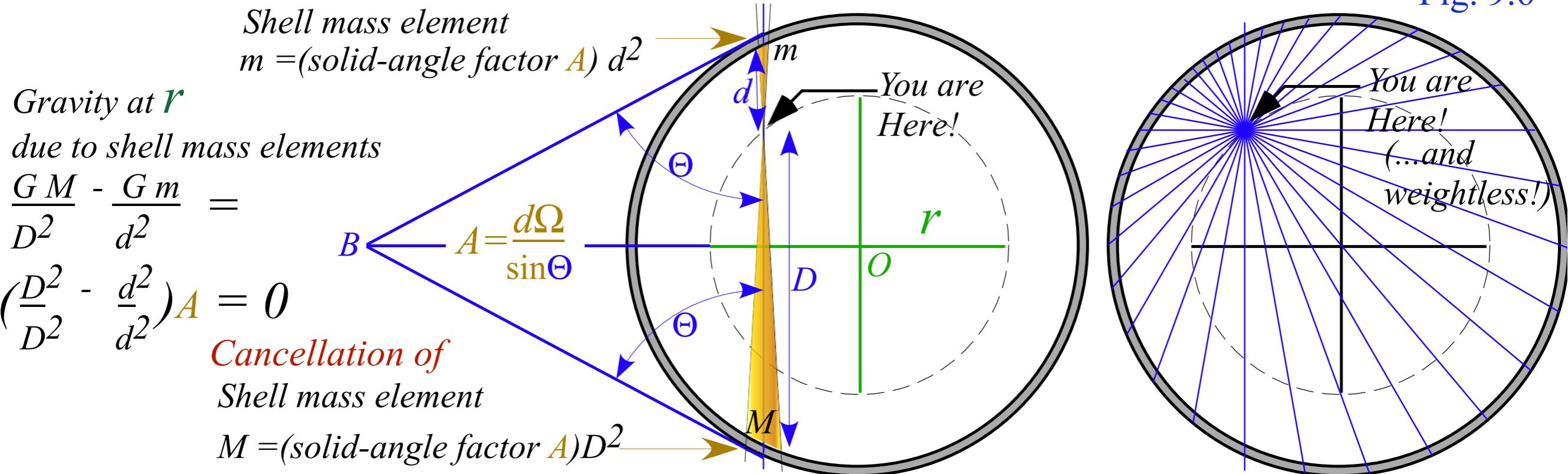
*Examples with x-y **phase lag** : $\alpha_{x-y} = \alpha_x - \alpha_y = 15^\circ, 30^\circ, \text{ and } \pm 75^\circ$*

Geometry of idealized “Sophomore-physics Earth”

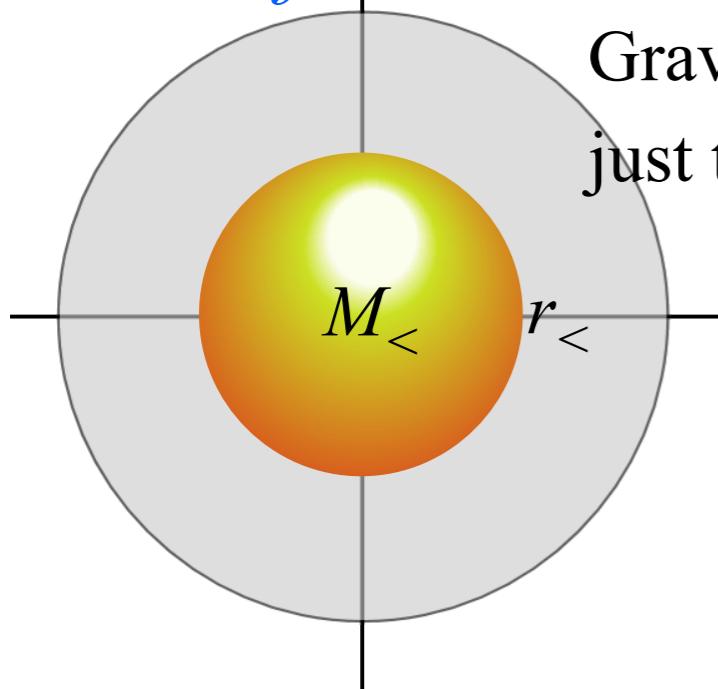
- *Coulomb field outside* *Isotropic Harmonic Oscillator (IHO) field inside*
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- Earth matter vs nuclear matter:*
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Coulomb force vanishes inside-spherical shell (Gauss-law)

Unit 1
Fig. 9.6



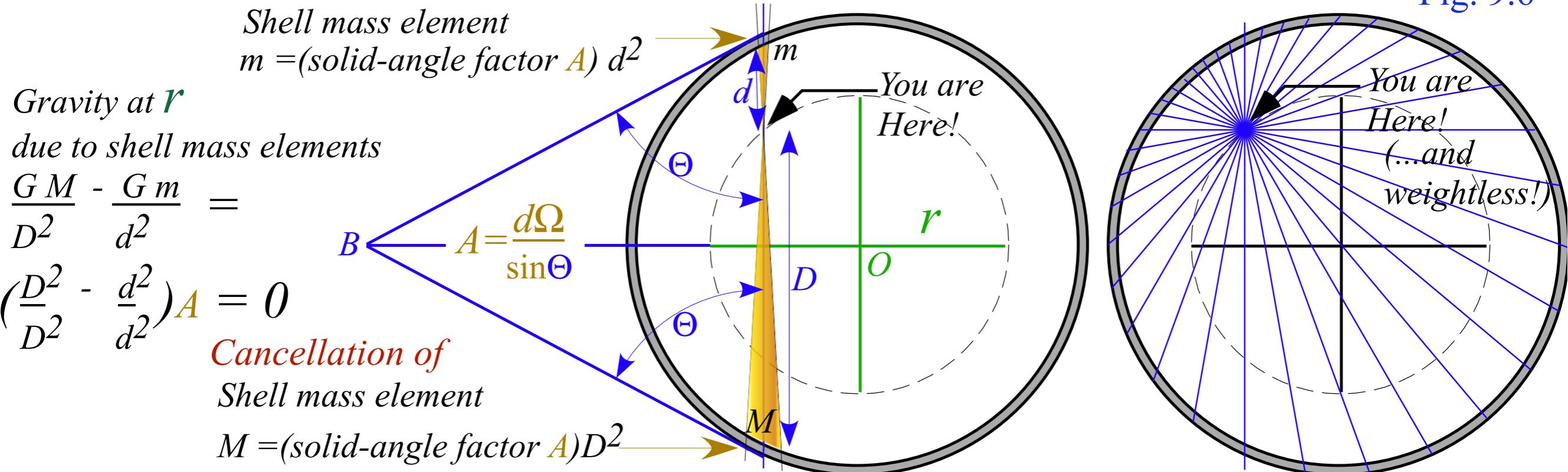
Coulomb force inside-spherical body due to stuff below you, only.



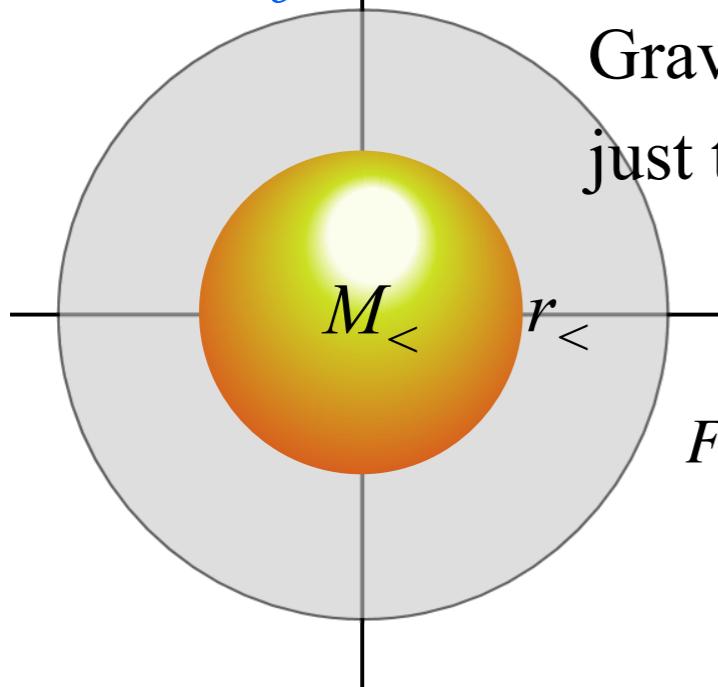
Gravitational force at $r_<$ is just that of planet $M_<$ below $r_<$

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Unit 1
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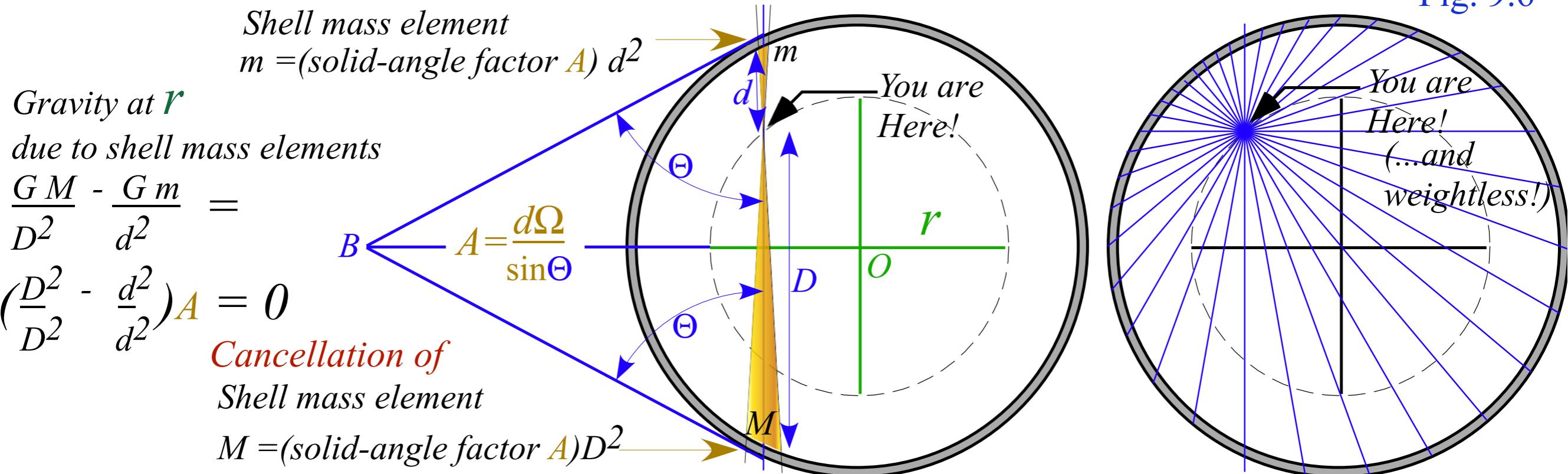


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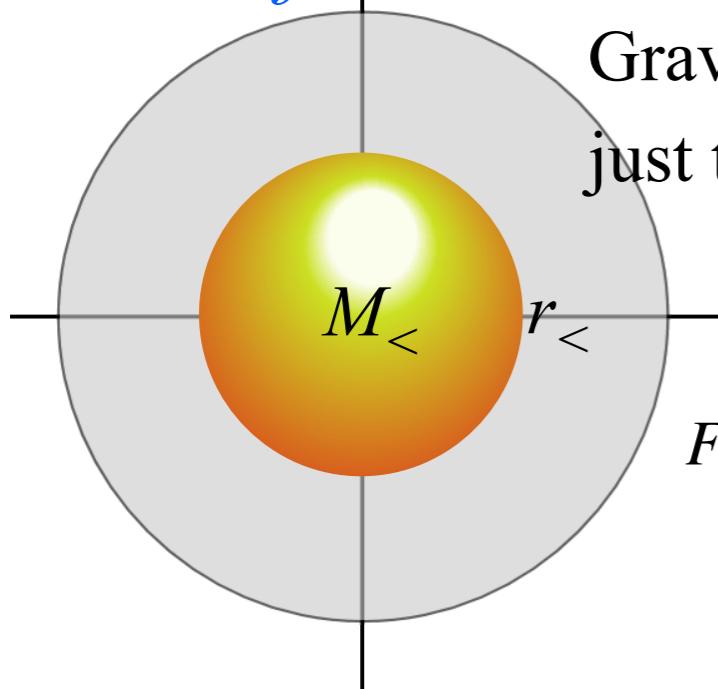
$$F^{inside}(r_<) = G \frac{m M_<}{r_<} = G m \frac{4\pi}{3} \frac{M_<}{\frac{4\pi}{3} r_<} r_<$$

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Unit 1
Fig. 9.6



Coulomb force inside-spherical body due to stuff below you, only.



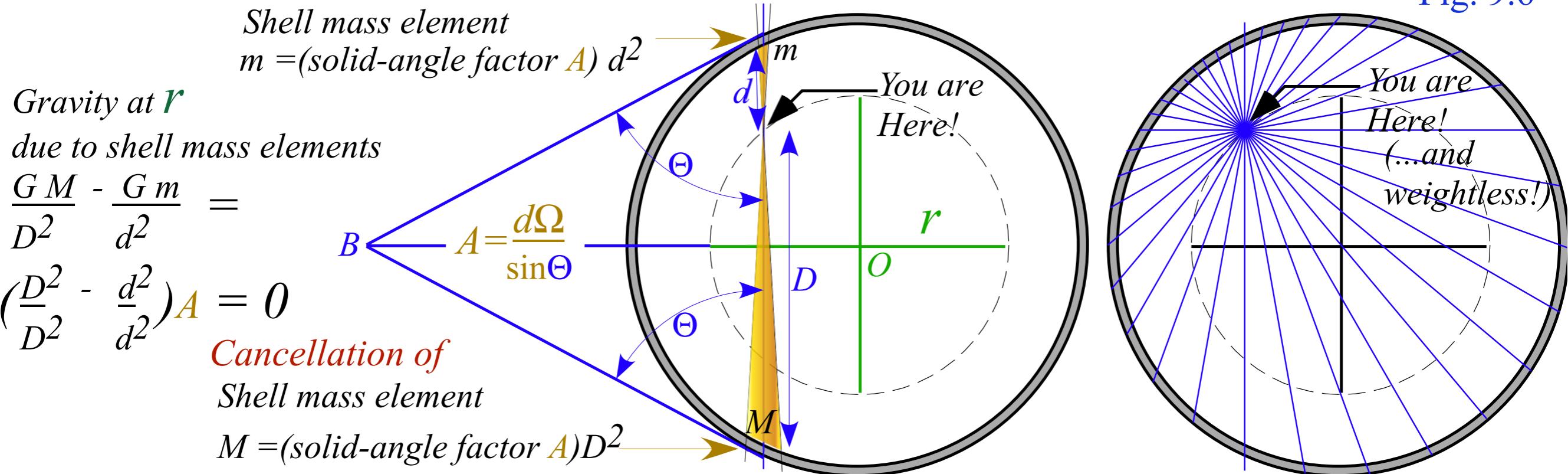
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$$F^{inside}(r_<) = G \frac{mM_<}{r_<} = Gm \frac{4\pi}{3} \frac{M_<}{\frac{4\pi}{3} r_<} r_< = Gm \frac{4\pi}{3} \rho_{\oplus} r_< = mg \frac{r_<}{R_{\oplus}} \equiv mg \cdot x$$

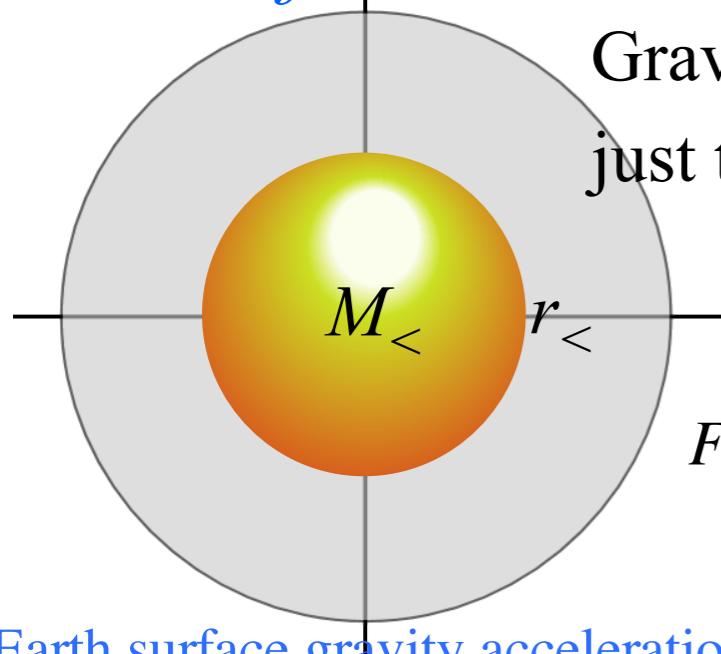
Note:
Hooke's (linear) force law
for $r_<$ inside uniform body

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Unit 1
Fig. 9.6



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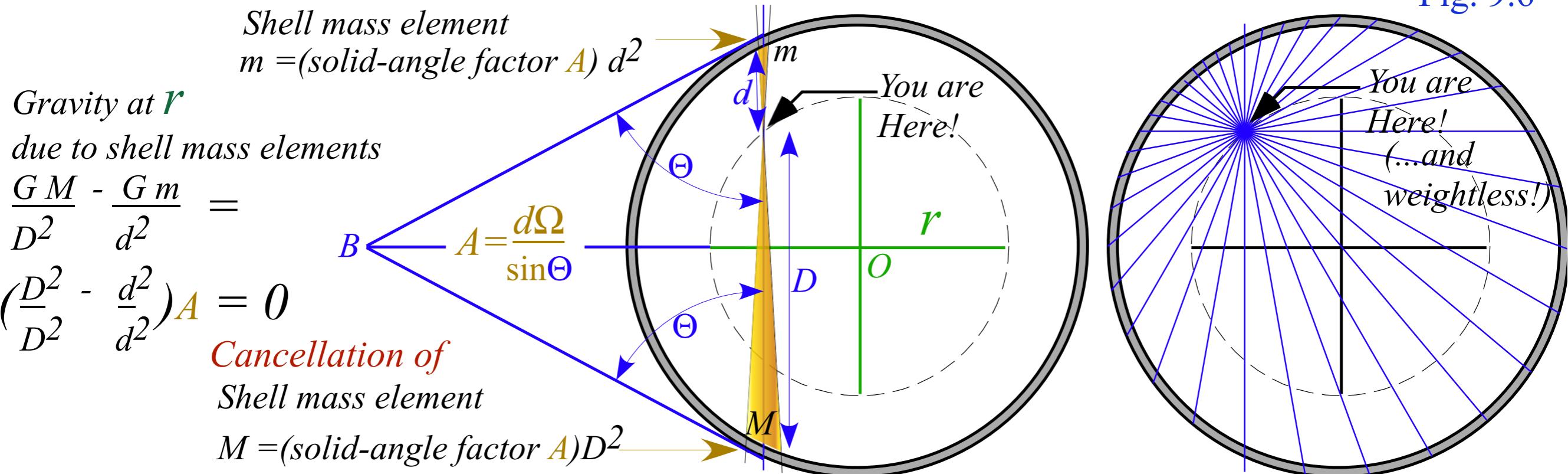
$$\text{Earth surface gravity acceleration: } g = G \frac{M_{\oplus}}{R_{\oplus}^2} = G \frac{M_{\oplus}}{R_{\oplus}^3} R_{\oplus} = G \frac{4\pi}{3} \frac{M_{\oplus}}{4\pi R_{\oplus}^3} R_{\oplus} = G \frac{4\pi}{3} \rho_{\oplus} R_{\oplus} = 9.8 \text{ m/s}^2$$

Note:
Hooke's (linear) force law for $r_<$ inside uniform body

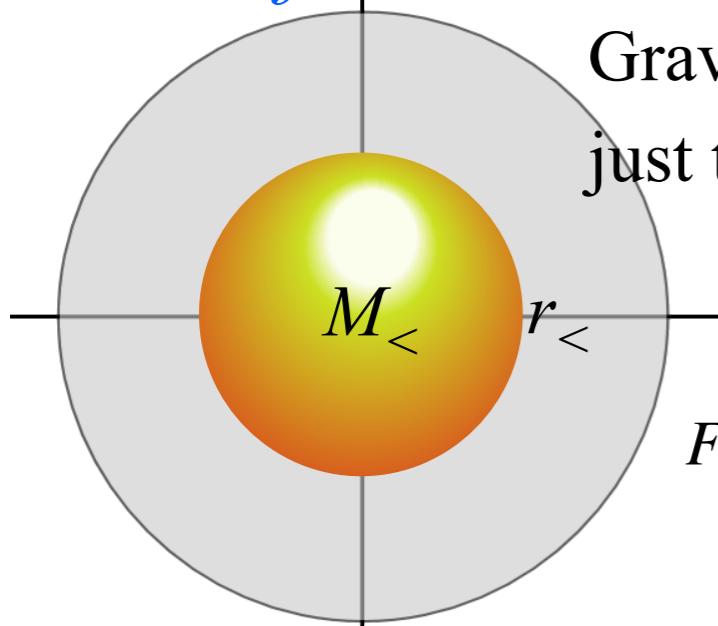


Coulomb force vanishes inside-spherical shell (Gauss-law)

Unit 1
Fig. 9.6



Coulomb force inside-spherical body due to stuff below you, only.



Gravitational force at $r_<$ is just that of planet $m_<$ below $r_<$

$$F_{\text{inside}}(r_<) = G \frac{m M_<}{r_<} = G m \frac{4\pi}{3} \frac{M_<}{\frac{4\pi}{3} r_<} r_< = G m \frac{4\pi}{3} \rho_< r_<^2 = m g \frac{r_<}{R_<} = m g \cdot x$$

Earth surface gravity acceleration: $g = G \frac{M_\oplus}{R_\oplus^2} = G \frac{M_\oplus}{R_\oplus^3} R_\oplus = G \frac{4\pi}{3} \frac{M_\oplus}{\frac{4\pi}{3} R_\oplus^3} R_\oplus = G \frac{4\pi}{3} \rho_\oplus R_\oplus = 9.8 \text{ m/s}^2$

$$G = 6.67384(80) \cdot 10^{-11} \text{ Nm}^2/\text{C}^2 \sim (2/3) 10^{-10}$$

$$\text{Earth radius: } R_\oplus = 6.371 \cdot 10^6 \text{ m} \approx 6.4 \cdot 10^6 \text{ m}$$

$$\text{Earth mass: } M_\oplus = 5.9722 \times 10^{24} \text{ kg.} \approx 6.0 \cdot 10^{24} \text{ kg.}$$

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Hooke's (linear) force law for $r_<$ inside uniform body

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$$\text{Solar radius: } R_\odot = 6.955 \times 10^8 \text{ m.} \approx 7.0 \cdot 10^8 \text{ m.}$$

$$\text{Solar mass: } M_\odot = 1.9889 \times 10^{30} \text{ kg.} \approx 2.0 \cdot 10^{30} \text{ kg.}$$

Geometry of idealized “Sophomore-physics Earth”

Coulomb field outside

Isotropic Harmonic Oscillator (IHO) field inside

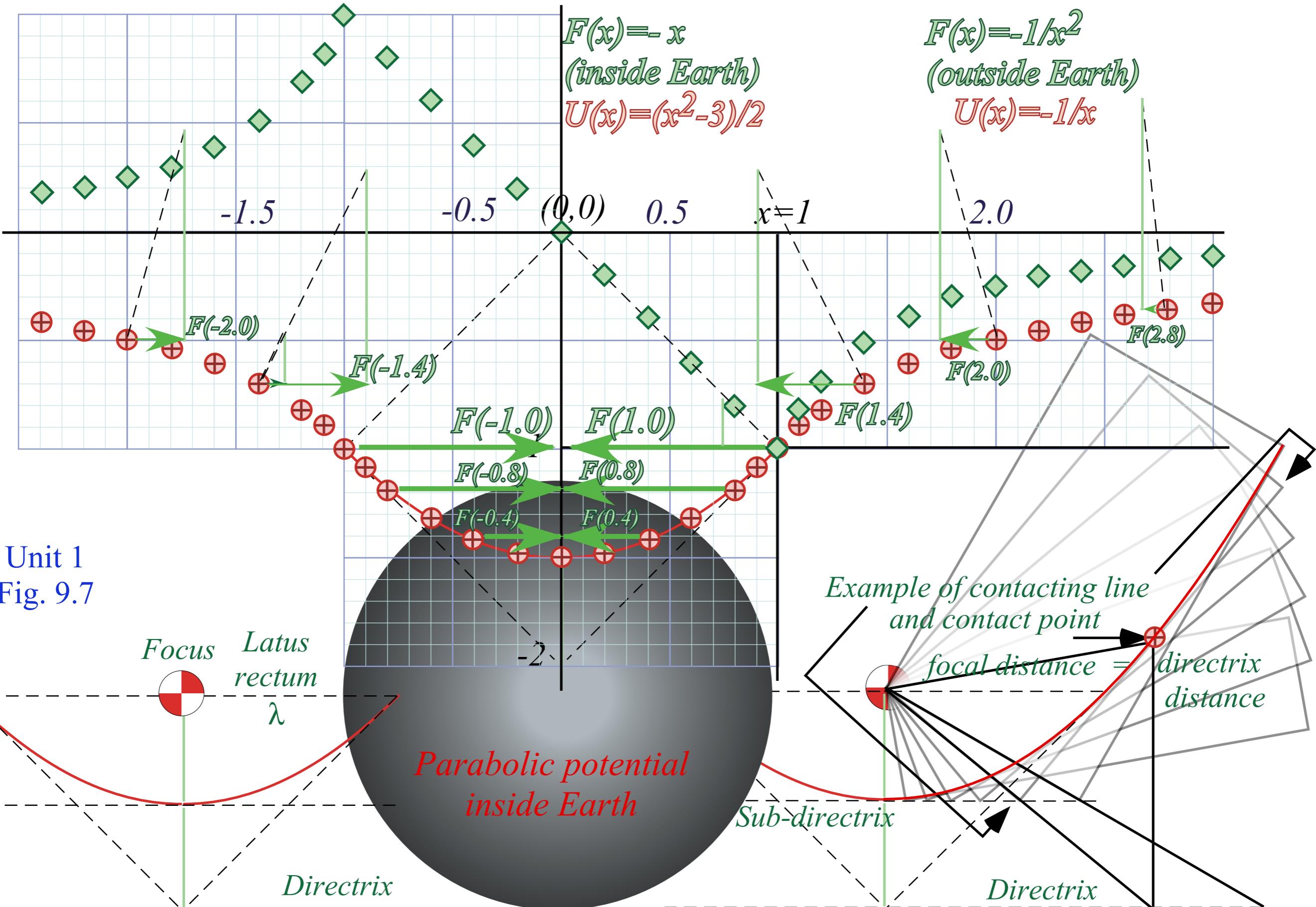
→ *Contact-geometry of potential curve(s)*

“Crushed-Earth” models: 3 key energy “steps” and 4 key energy “levels”

Earth matter vs nuclear matter:

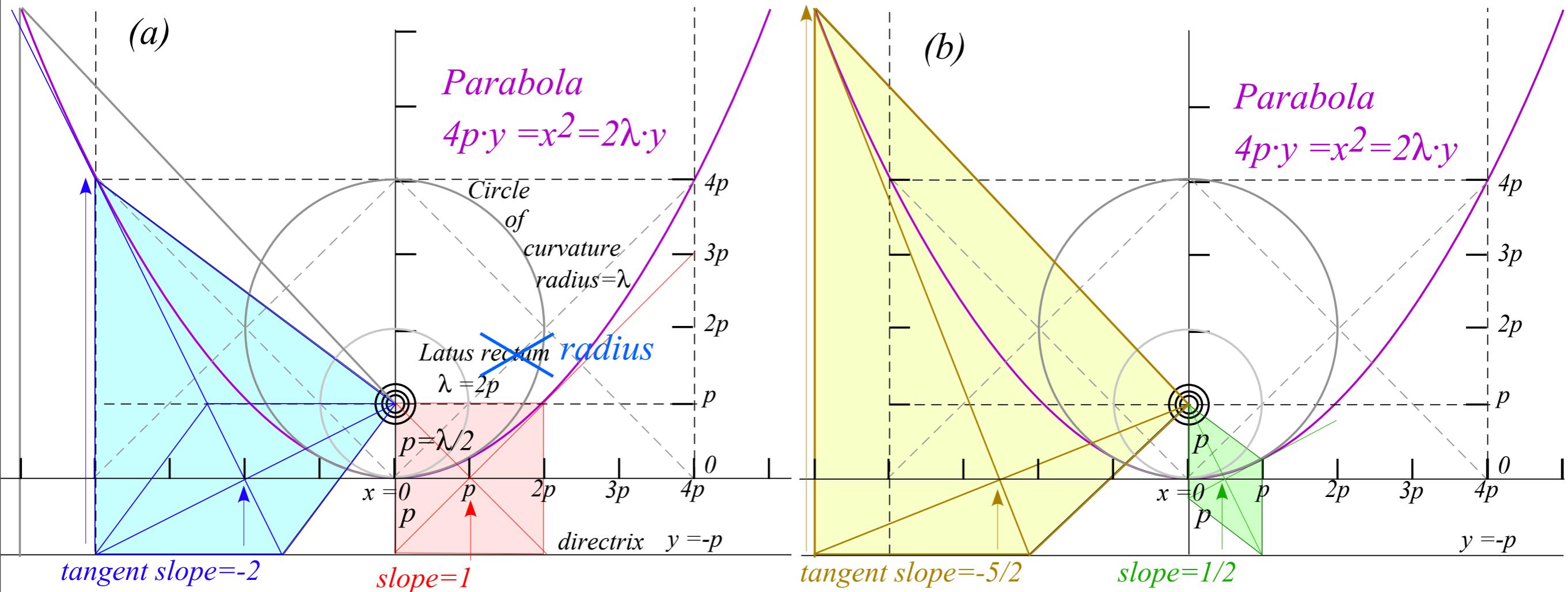
*Introducing the “neutron starlet” and “**Black-Hole-Earth**”*

The ideal “Sophomore-Physics-Earth” model of geo-gravity



...conventional parabolic geometry...carried to extremes...

(Review of Lect. 6 p.29)



Unit 1
 Fig. 9.4

Geometry of idealized “Sophomore-physics Earth”

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Isotropic Harmonic Oscillator (IHO) field inside

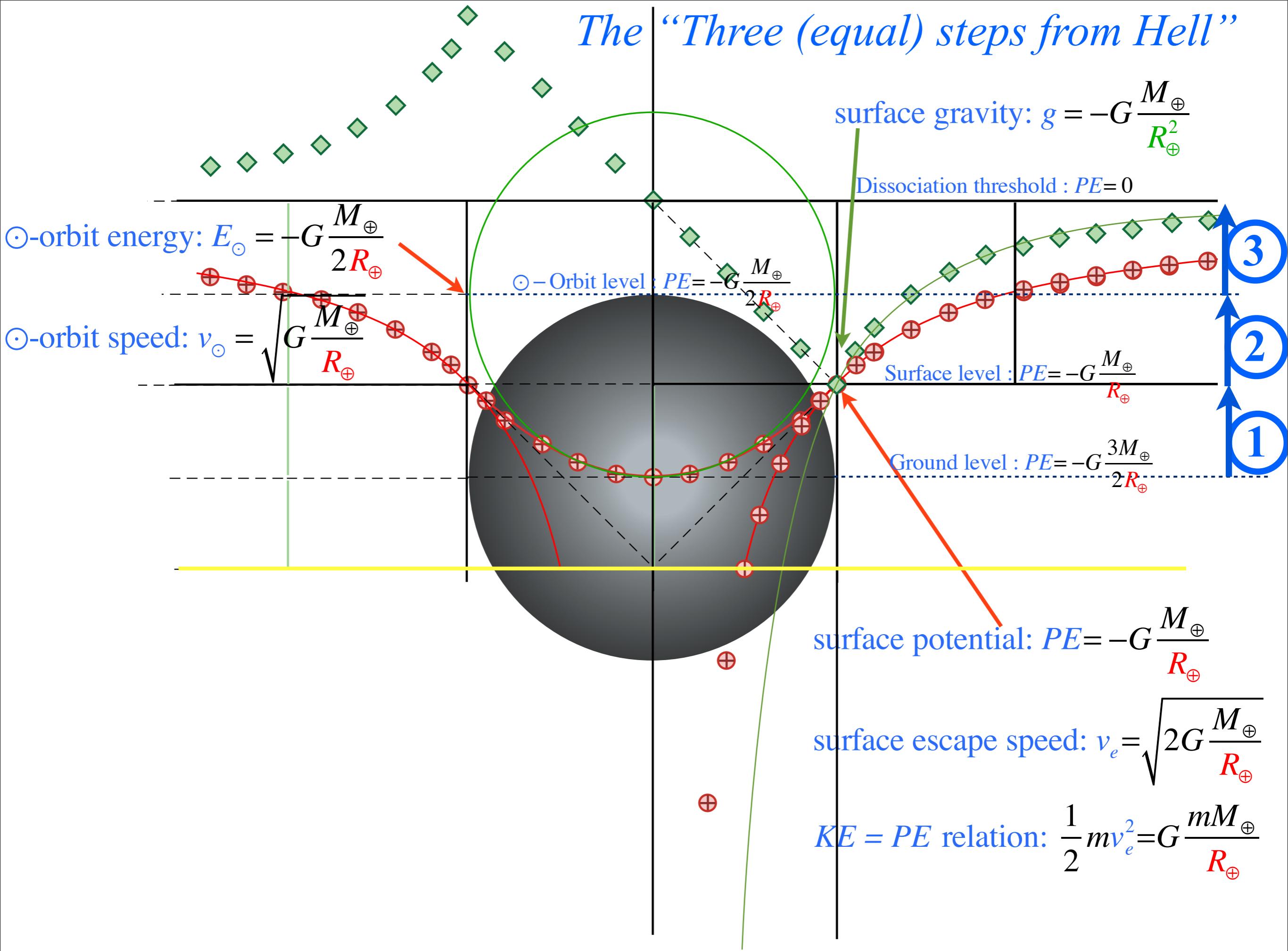
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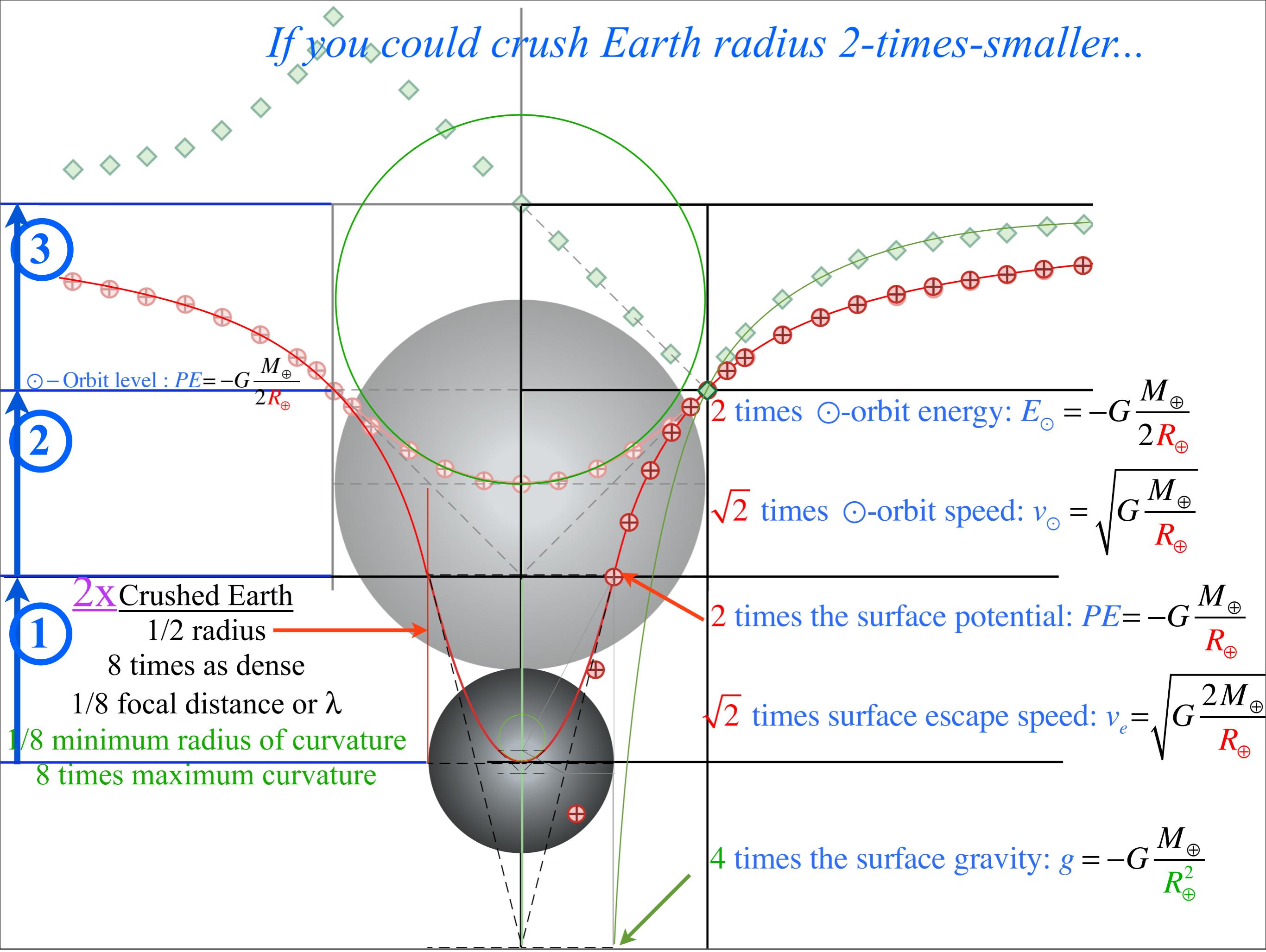
Earth matter vs nuclear matter:

*Introducing the “neutron starlet” and “**Black-Hole-Earth**”*

The “Three (equal) steps from Hell”



If you could crush Earth radius 2-times-smaller...



Geometry of idealized “Sophomore-physics Earth”

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Examples of “crushed” matter

Earth matter Earth mass : $M_{\oplus} = 5.9722 \times 10^{24} \text{ kg.} \approx 6.0 \cdot 10^{24} \text{ kg.}$ Density $\rho_{\oplus} ??$

Earth radius : $R_{\oplus} = 6.371 \cdot 10^6 \text{ m} \approx 6.4 \cdot 10^6 \text{ m}$ Earth volume : $(4\pi / 3)R_{\oplus}^3 \approx 4 \cdot 260 \cdot 10^{18} \approx 10^{21} \text{ m}^3$

$$(6.4)^3 \approx 262 \text{ and } (4\pi/3)260 = 1098 \approx 10^3$$

Examples of “crushed” matter

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Density of solid Fe= $7.9 \cdot 10^3 \text{ kg/m}^3$
Density of liquid Fe= $6.9 \cdot 10^3 \text{ kg/m}^3$

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Nuclear matter Nucleon mass = $1.67 \cdot 10^{-27} \text{ kg.} \sim 2 \cdot 10^{-27} \text{ kg.}$

Say a nucleus of atomic weight 50 has a radius of 3 fm , or 50 nucleons each with a mass $2 \cdot 10^{-27} \text{ kg.}$

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$$36\pi=113 \sim 10^2$$

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Nuclear density is $10^{-25+43} = 10^{18} \text{ kg/m}^3$ or a *trillion* (10^{12}) kilograms in the size of a fingertip (1cc).

Earth radius crushed by a factor of $0.5 \cdot 10^{-5}$ to $R_{\text{crush}\oplus} \approx 300 \text{ m}$ would approach neutron-star density.

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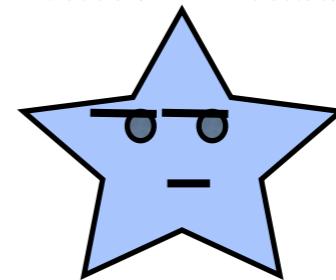
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Introducing the “Neutron starlet” 1 cm^3 of nuclear matter: mass = 10^{12} kg.



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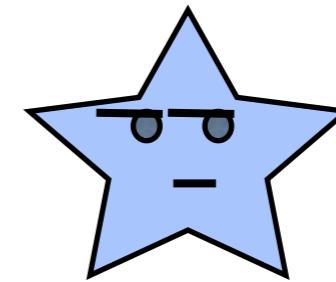
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Introducing the “Neutron starlet” 1 cm^3 of nuclear matter: mass = 10^{12} kg.



Introducing the “Black Hole Earth” Suppose Earth is crushed so that its

surface escape velocity is the speed of light $c \cong 3.0 \cdot 10^8 \text{ m/s.}$

$c \equiv 299,792,458 \text{ m/s (EXACTLY)}$

$$c = \sqrt{(2GM/R_{\odot})}$$

$$R_{\odot} = 2GM/c^2 = 8.9 \text{ mm} \sim 1 \text{ cm} \quad (\text{fingertip size!})$$

$$G = 6.67384(80) \cdot 10^{-11} \text{ Nm}^2/\text{C}^2 \sim (2/3)10^{-10}$$

Isotropic harmonic oscillator dynamics in 1D, 2D, and 3D

→ *Sinusoidal space-time dynamics derived by geometry*

Isotropic harmonic oscillator orbits in 1D and 2D (You get 3D for free!)

Constructing 2D Isotropic harmonic oscillator orbits using phasor plots

Examples with x-y phase lag : $\alpha_{x-y} = \alpha_x - \alpha_y = 15^\circ, 30^\circ, \text{ and } \pm 75^\circ$

[BoxIt simulation of U\(2\) orbits](#)

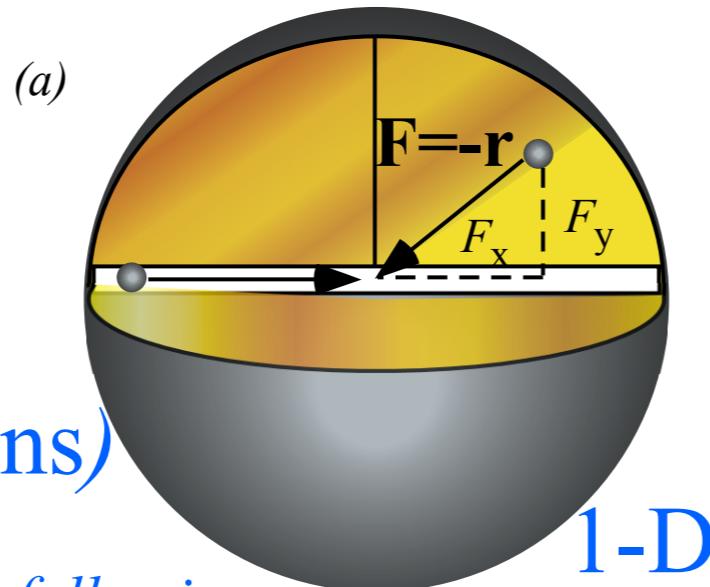
<http://www.uark.edu/ua/modphys/markup/BoxItWeb.html>

Isotropic Harmonic Oscillator phase dynamics in uniform-body

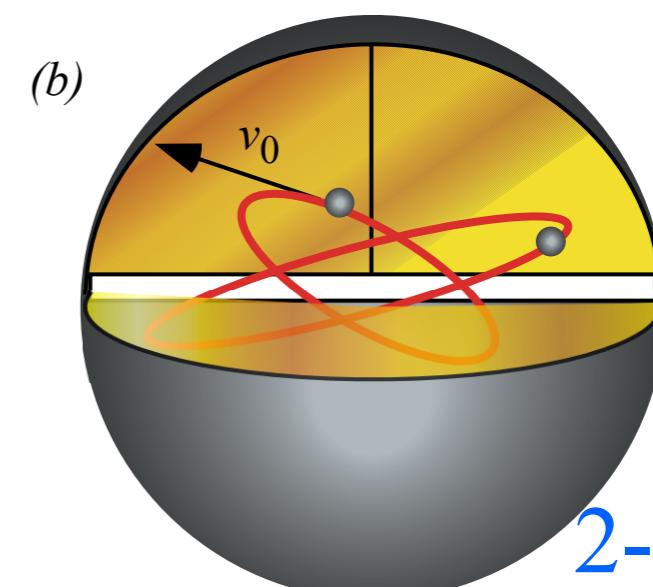
I.H.O. Force law

$$F = -x \quad (\text{1-Dimension})$$

$$\mathbf{F} = -\mathbf{r} \quad (\text{2 or 3-Dimensions})$$



1-D



2-D

Unit 1
Fig. 9.10

Each dimension x , y , or z obeys the following:

$$\text{Total } E = KE + PE = \frac{1}{2}mv^2 + U(x) = \frac{1}{2}mv^2 + \frac{1}{2}kx^2 = \text{const.}$$

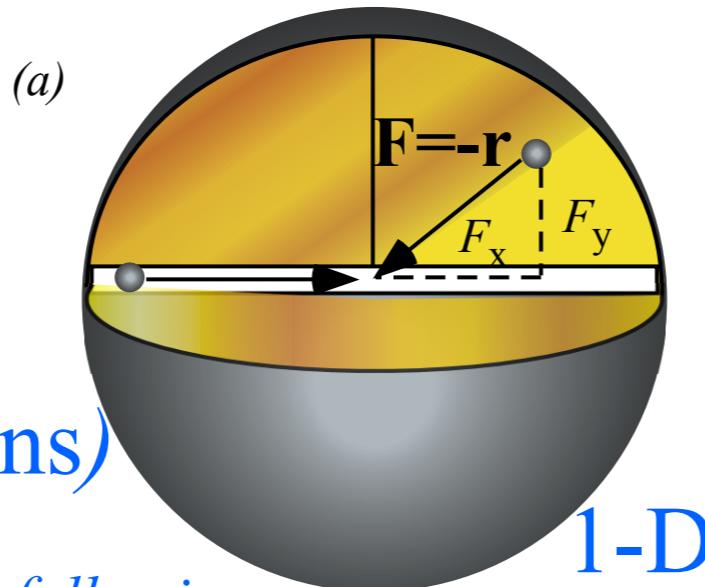
(Paths are always
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Isotropic Harmonic Oscillator phase dynamics in uniform-body

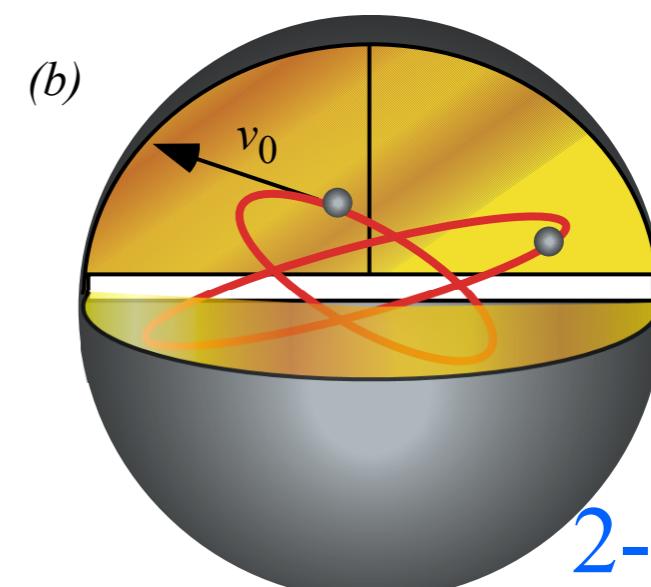
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$$F = -x \quad (\text{1-Dimension})$$

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1-D



Unit 1
Fig. 9.10

2-D or 3-D

(Paths are *always* 2-D ellipses if viewed right!)

Each dimension x , y , or z obeys the following:

$$\text{Total } E = KE + PE = \frac{1}{2}mv^2 + U(x) = \frac{1}{2}mv^2 + \frac{1}{2}kx^2 = \text{const.}$$

Equations for x -motion

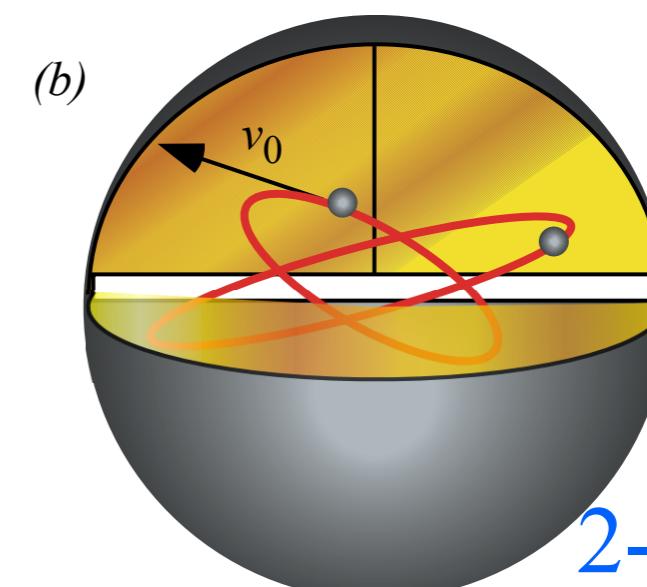
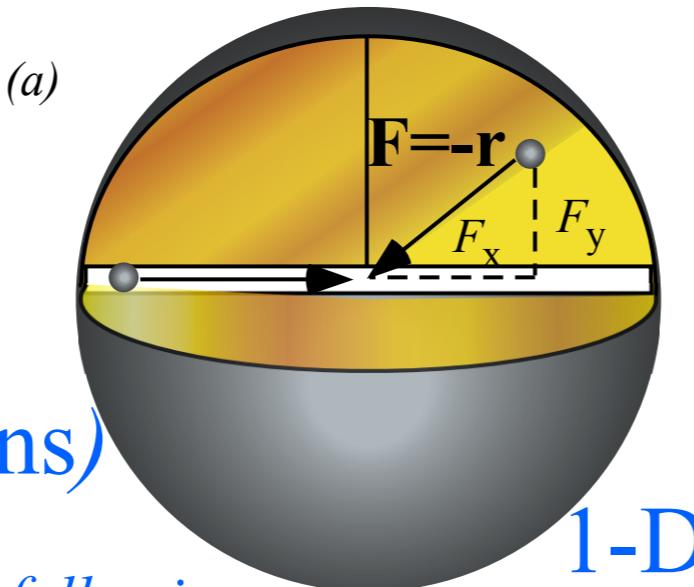
[$x(t)$ and $v_x = v(t)$] are given first. They apply as well to dimensions [$y(t)$ and $v_y = v(t)$] and [$z(t)$ and $v_z = v(t)$] in the ideal isotropic case.

Isotropic Harmonic Oscillator phase dynamics in uniform-body

I.H.O. Force law

$$F = -x \quad (\text{1-Dimension})$$

$$\mathbf{F} = -\mathbf{r} \quad (\text{2 or 3-Dimensions})$$



Unit 1
Fig. 9.10

Each dimension x, y , or z obeys the following:

$$\text{Total } E = KE + PE = \frac{1}{2}mv^2 + U(x) = \frac{1}{2}mv^2 + \frac{1}{2}kx^2 = \text{const.}$$

Equations for x -motion

[$x(t)$ and $v_x = v(t)$] are given first. They apply as well to dimensions [$y(t)$ and $v_y = v(t)$] and [$z(t)$ and $v_z = v(t)$] in the ideal isotropic case.

$$1 = \frac{mv^2}{2E} + \frac{kx^2}{2E} = \left(\frac{v}{\sqrt{2E/m}} \right)^2 + \left(\frac{x}{\sqrt{2E/k}} \right)^2$$

$$1 = \frac{mv^2}{2E} + \frac{kx^2}{2E} = (\cos\theta)^2 + (\sin\theta)^2$$

Another example of the old “scale-a-circle” trick...

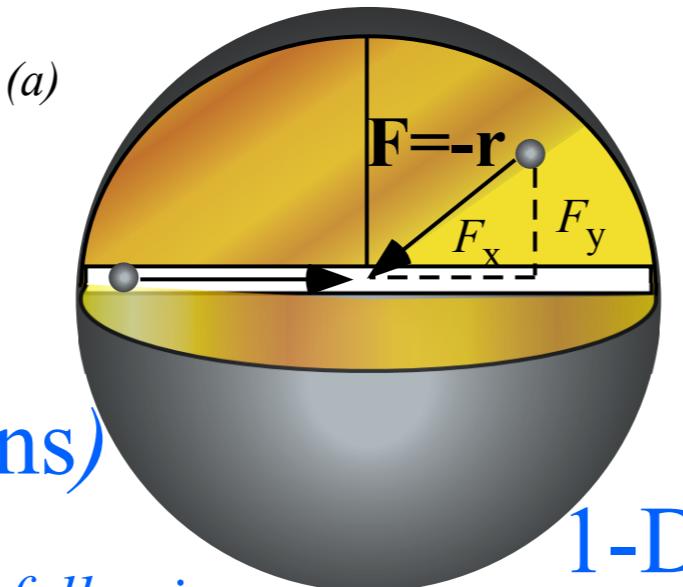
$$\text{Let : (1)} \quad v = \sqrt{2E/m} \cos\theta, \quad \text{and : (2)} \quad x = \sqrt{2E/k} \sin\theta$$

Isotropic Harmonic Oscillator phase dynamics in uniform-body

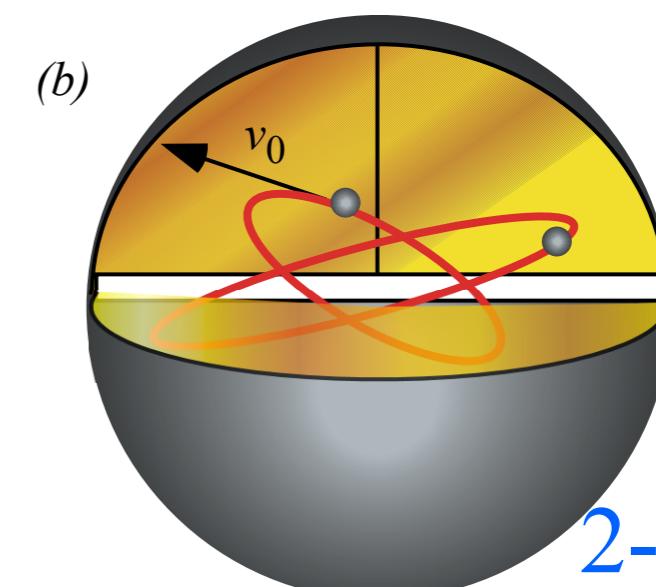
I.H.O. Force law

$$F = -x \quad (\text{1-Dimension})$$

$$\mathbf{F} = -\mathbf{r} \quad (\text{2 or 3-Dimensions})$$



1-D



Unit 1
Fig. 9.10

2-D or 3-D

(Paths are always 2-D ellipses if viewed right!)

Each dimension x, y , or z obeys the following:

$$\text{Total } E = KE + PE = \frac{1}{2}mv^2 + U(x) = \frac{1}{2}mv^2 + \frac{1}{2}kx^2 = \text{const.}$$

Equations for x -motion

[$x(t)$ and $v_x = v(t)$] are given first. They apply as well to dimensions [$y(t)$ and $v_y = v(t)$] and [$z(t)$ and $v_z = v(t)$] in the ideal isotropic case.

$$1 = \frac{mv^2}{2E} + \frac{kx^2}{2E} = \left(\frac{v}{\sqrt{2E/m}} \right)^2 + \left(\frac{x}{\sqrt{2E/k}} \right)^2$$

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Another example of the old “scale-a-circle” trick...

$$\text{Let : (1)} \quad v = \sqrt{2E/m} \cos\theta, \quad \text{and : (2)} \quad x = \sqrt{2E/k} \sin\theta$$

$$\text{def. (3)} \quad \omega = \frac{d\theta}{dt}$$

$$\sqrt{\frac{2E}{m}} \cos\theta = v = \frac{dx}{dt} = \frac{d\theta}{dt} \frac{dx}{d\theta} = \omega \frac{dx}{d\theta}$$

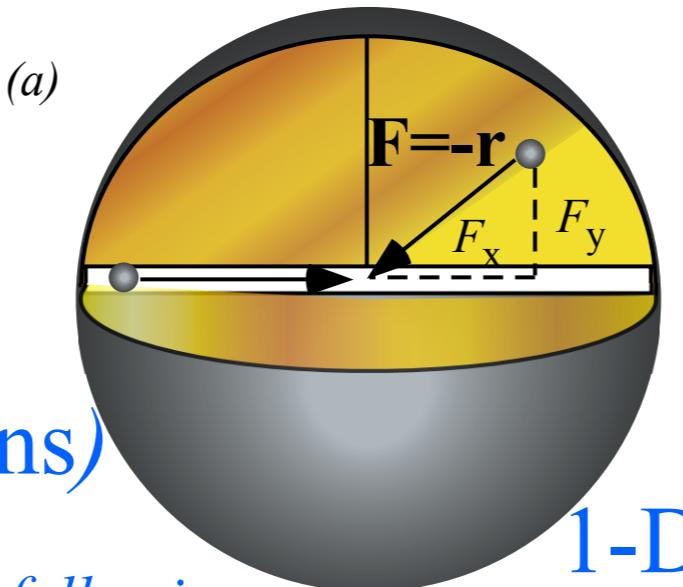
by (1) by def. (3)

Isotropic Harmonic Oscillator phase dynamics in uniform-body

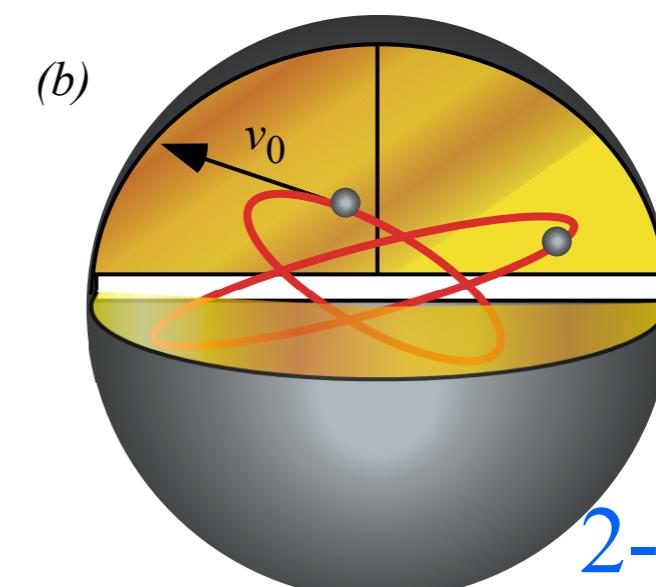
I.H.O. Force law

$$F = -x \quad (\text{1-Dimension})$$

$$\mathbf{F} = -\mathbf{r} \quad (\text{2 or 3-Dimensions})$$



1-D



Unit 1
Fig. 9.10

2-D or 3-D

(Paths are always 2-D ellipses if viewed right!)

Each dimension x, y , or z obeys the following:

$$\text{Total } E = KE + PE = \frac{1}{2}mv^2 + U(x) = \frac{1}{2}mv^2 + \frac{1}{2}kx^2 = \text{const.}$$

Equations for x -motion

[$x(t)$ and $v_x = v(t)$] are given first. They apply as well to dimensions [$y(t)$ and $v_y = v(t)$] and [$z(t)$ and $v_z = v(t)$] in the ideal isotropic case.

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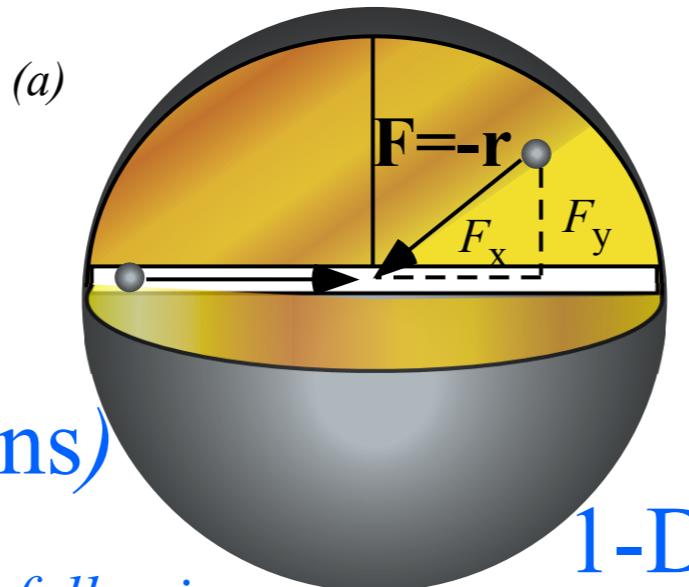
by (1) by def. (3) by (2)

Isotropic Harmonic Oscillator phase dynamics in uniform-body

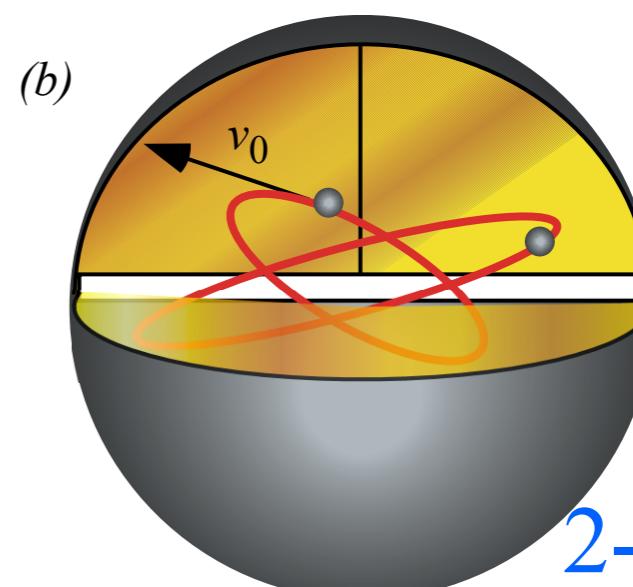
I.H.O. Force law

$F = -x$ (1-Dimension)

F = -r (2 or 3-Dimensions)



1- Γ



Unit 1

Fig. 9.10

2-D or 3-D

(Paths are *always* 2-D ellipses if viewed right!)

Equations for x-motion

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$$def.\ (3) \quad \omega = \frac{d\theta}{dt}$$

$$\sqrt{\frac{2E}{m}} \cos \theta = v = \frac{dx}{dt} = \frac{d\theta}{dt} \frac{dx}{d\theta} = \omega \frac{dx}{d\theta} = \omega \sqrt{\frac{2E}{k}} \cos \theta$$

by (1)
by def. (3)
by (2)

$$\omega = \frac{d\theta}{dt} = \sqrt{\frac{k}{m}}$$

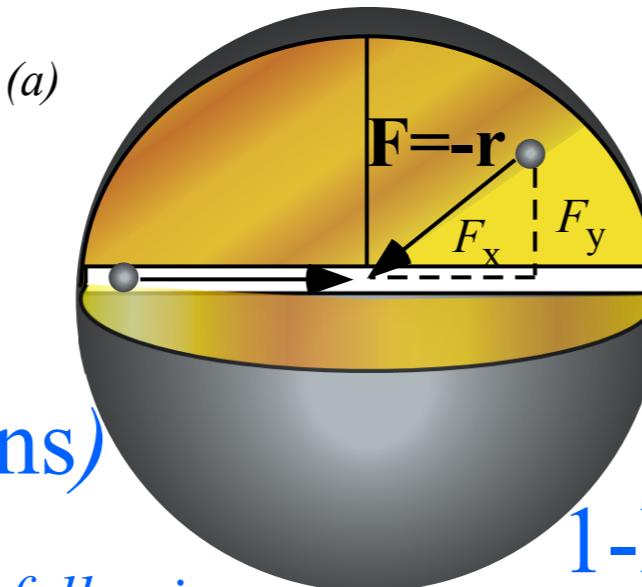
divide this by (1)

Isotropic Harmonic Oscillator phase dynamics in uniform-body

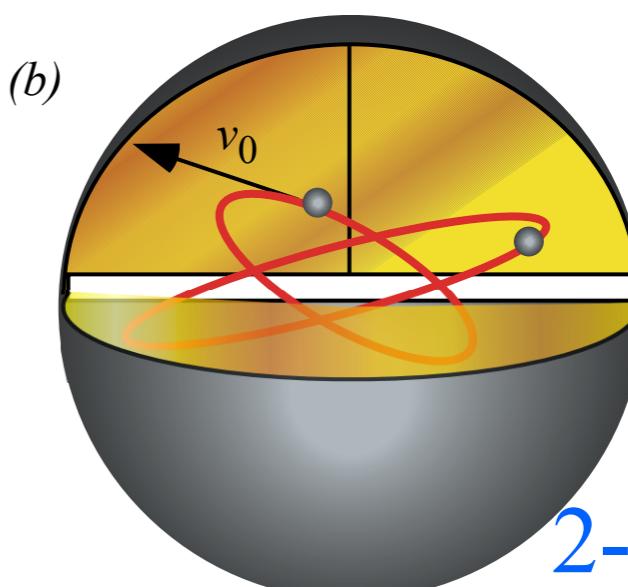
I.H.O. Force law

$F = -x$ (1-Dimension)

F = -r (2 or 3-Dimensions)



1-P



Unit 1

Fig. 9.10

2-D or 3-D

(Paths are *always* 2-D ellipses if viewed right!)

Each dimension x, y, or z obeys the following:

$$Total\ E = KE + PE = \frac{1}{2}mv^2 + U(x) = \frac{1}{2}mv^2 + \frac{1}{2}kx^2 = const.$$

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the old “scale-a-circle”
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Let : (1) $v = \sqrt{2E/m} \cos\theta$, and : (2) $x = \sqrt{2E/k} \sin\theta$ def. (3) $\omega = \frac{d\theta}{dt}$

$$\sqrt{\frac{2E}{m}} \cos \theta = v = \frac{dx}{dt} = \frac{d\theta}{dt} \frac{dx}{d\theta} = \omega \frac{dx}{d\theta} = \omega \sqrt{\frac{2E}{k}} \cos \theta$$

by (1) *by def. (3)* *by (2)*

$$\omega = \frac{d\theta}{dt} = \sqrt{\frac{k}{m}}$$

divide this by (1)

by integration given constant ω :

Isotropic harmonic oscillator dynamics in 1D, 2D, and 3D

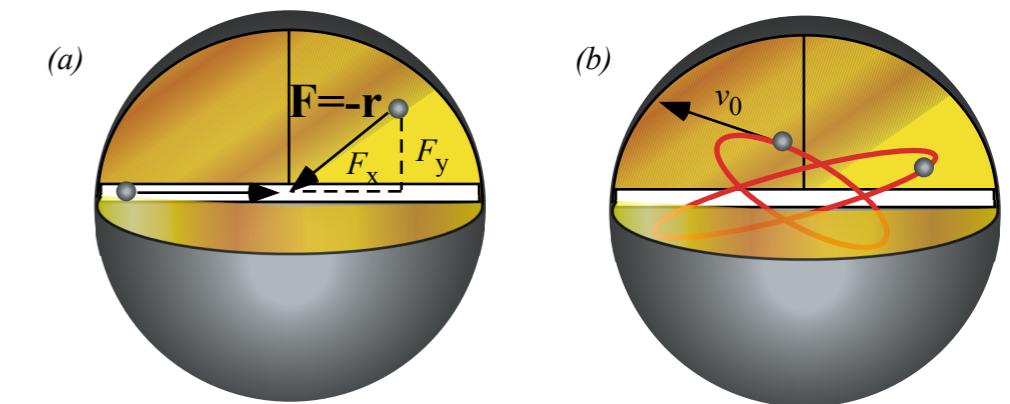
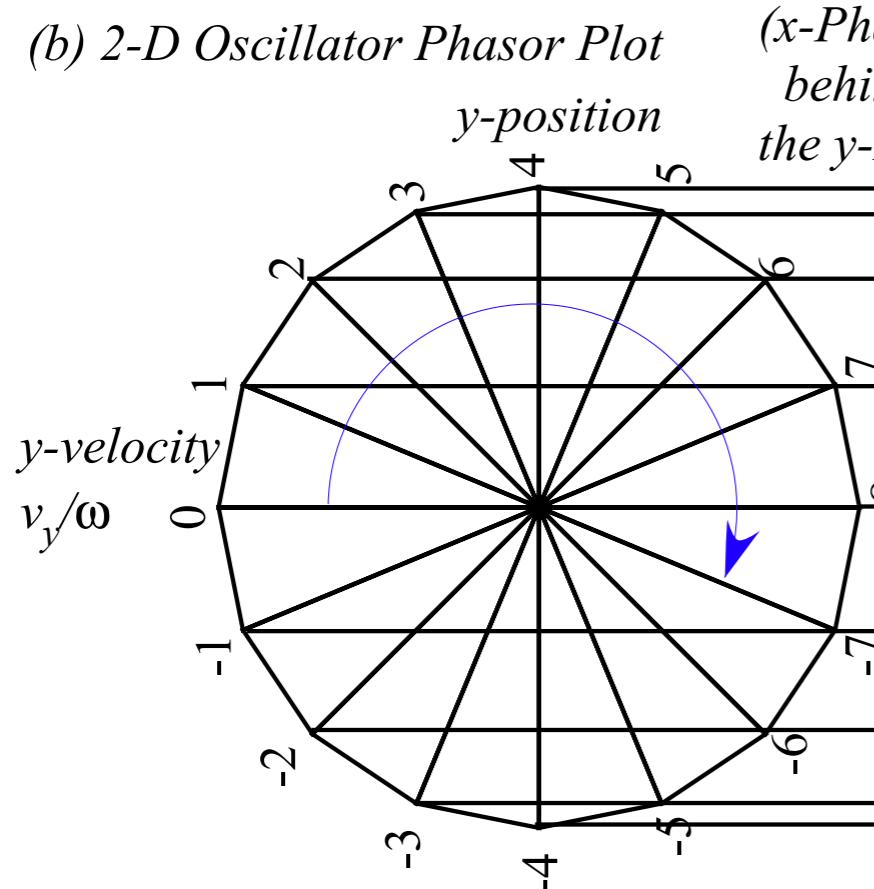
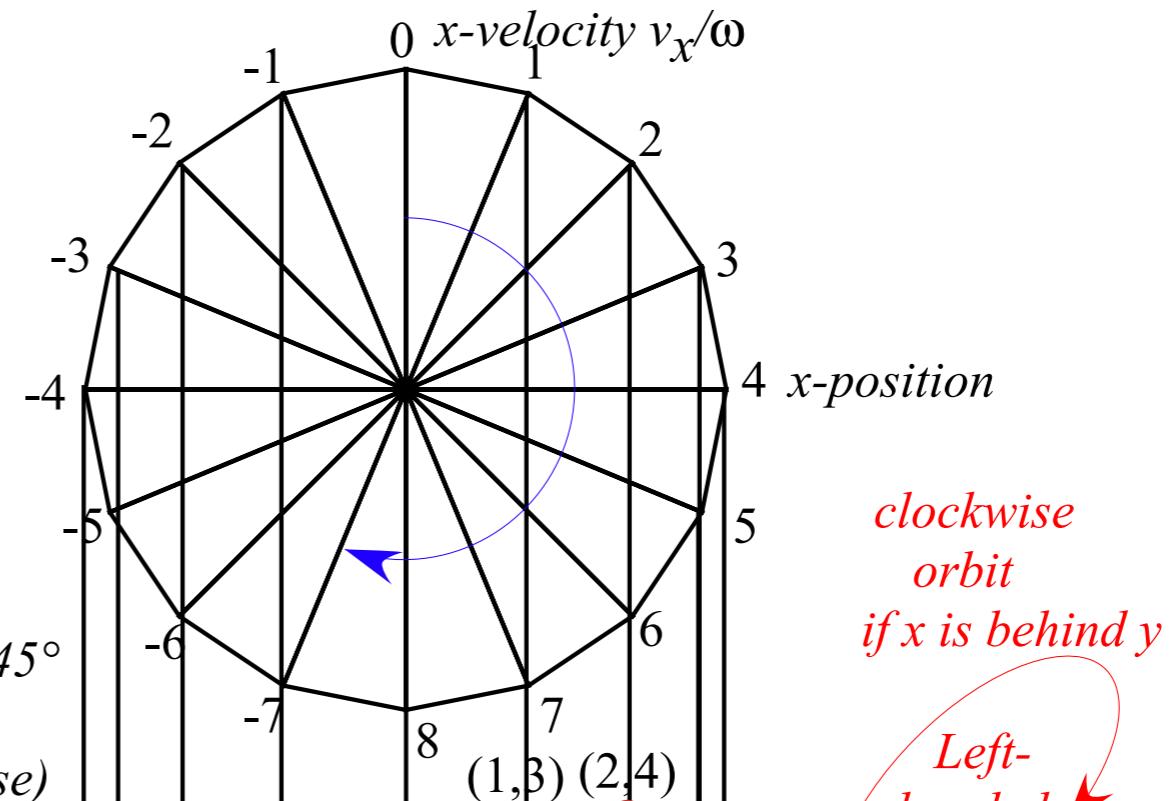
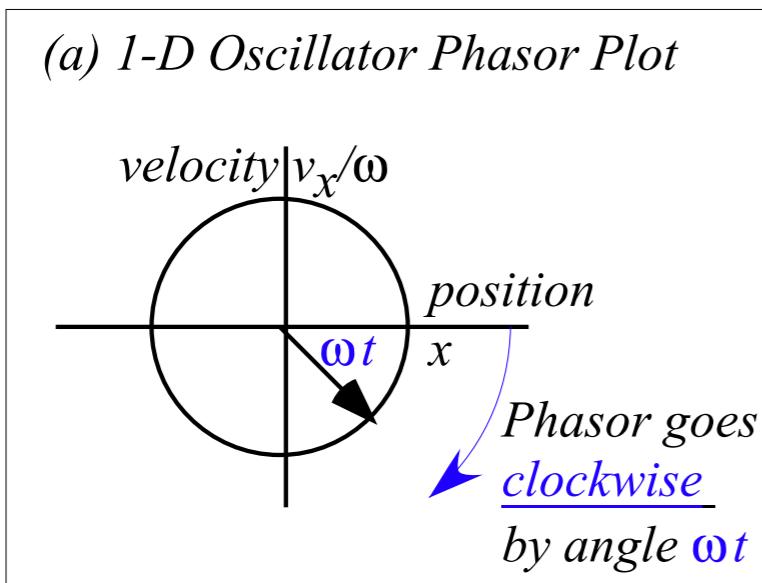
Sinusoidal space-time dynamics derived by geometry

Isotropic harmonic oscillator orbits in 1D and 2D (You get 3D for free!)

→ *Constructing 2D Isotropic harmonic oscillator orbits using phasor plots*

Examples with x-y phase lag : $\alpha_{x-y} = \alpha_x - \alpha_y = 15^\circ, 30^\circ, \text{ and } \pm 75^\circ$

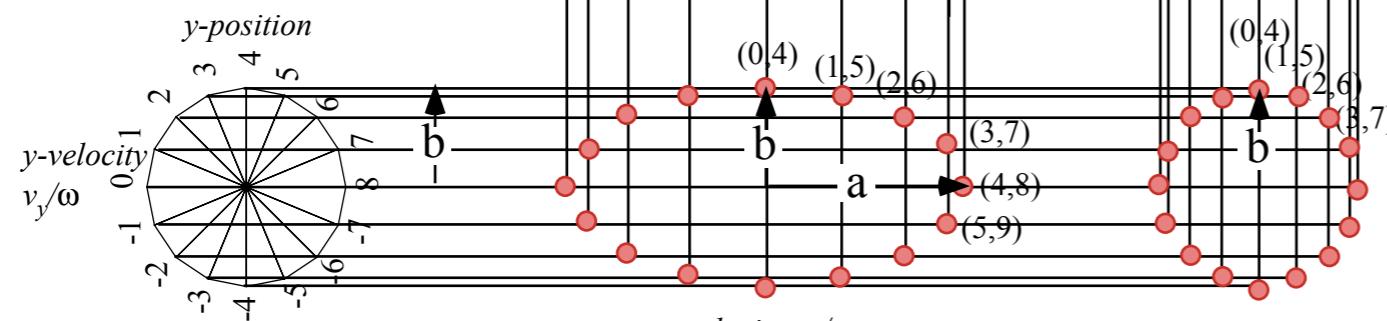
Isotropic Harmonic Oscillator phase dynamics in uniform-body



Unit 1
Fig. 9.10

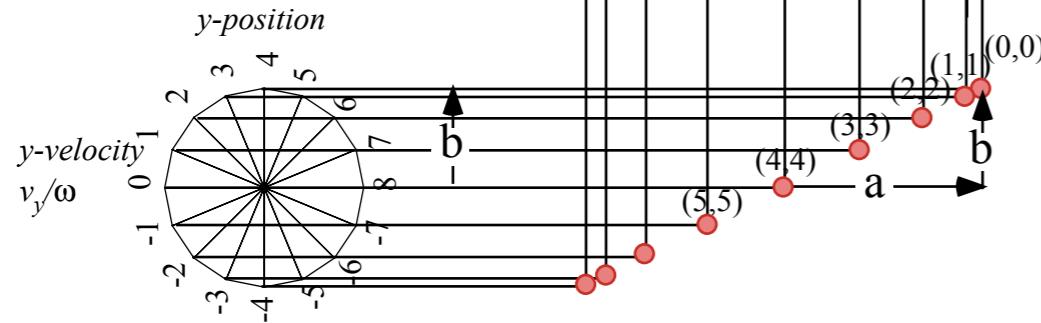
Unit 1
Fig. 9.12

(a) Phasor Plots
for
2-D Oscillator
or
Two 1D Oscillators
(x-Phase 90° behind
the y-Phase)



(b)
x-Phase 0° behind
the y-Phase

(In-phase case)



*These are more generic examples
with radius of x-phasor differing
from that of the y-phasor.*

Isotropic harmonic oscillator dynamics in 1D, 2D, and 3D

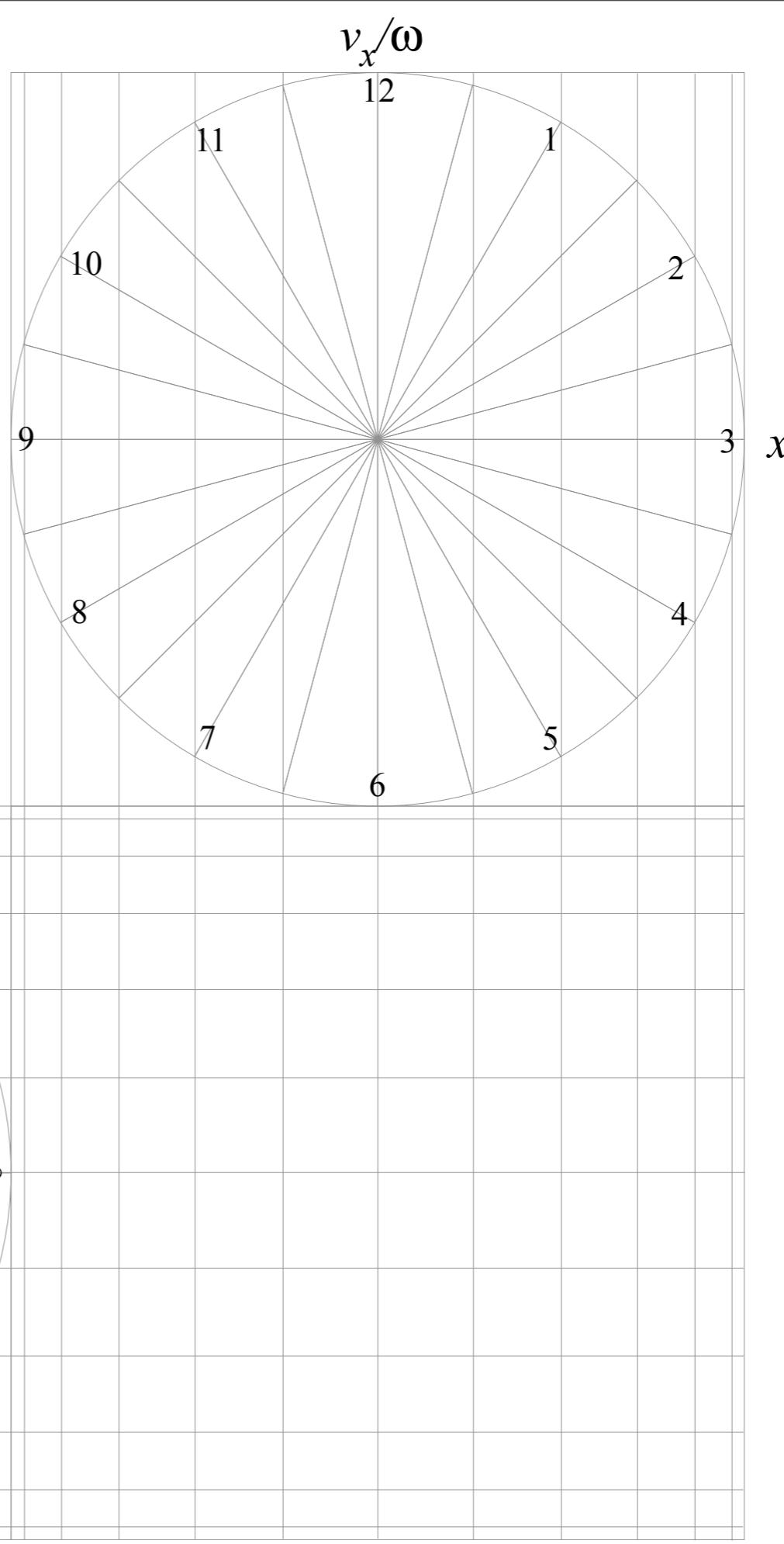
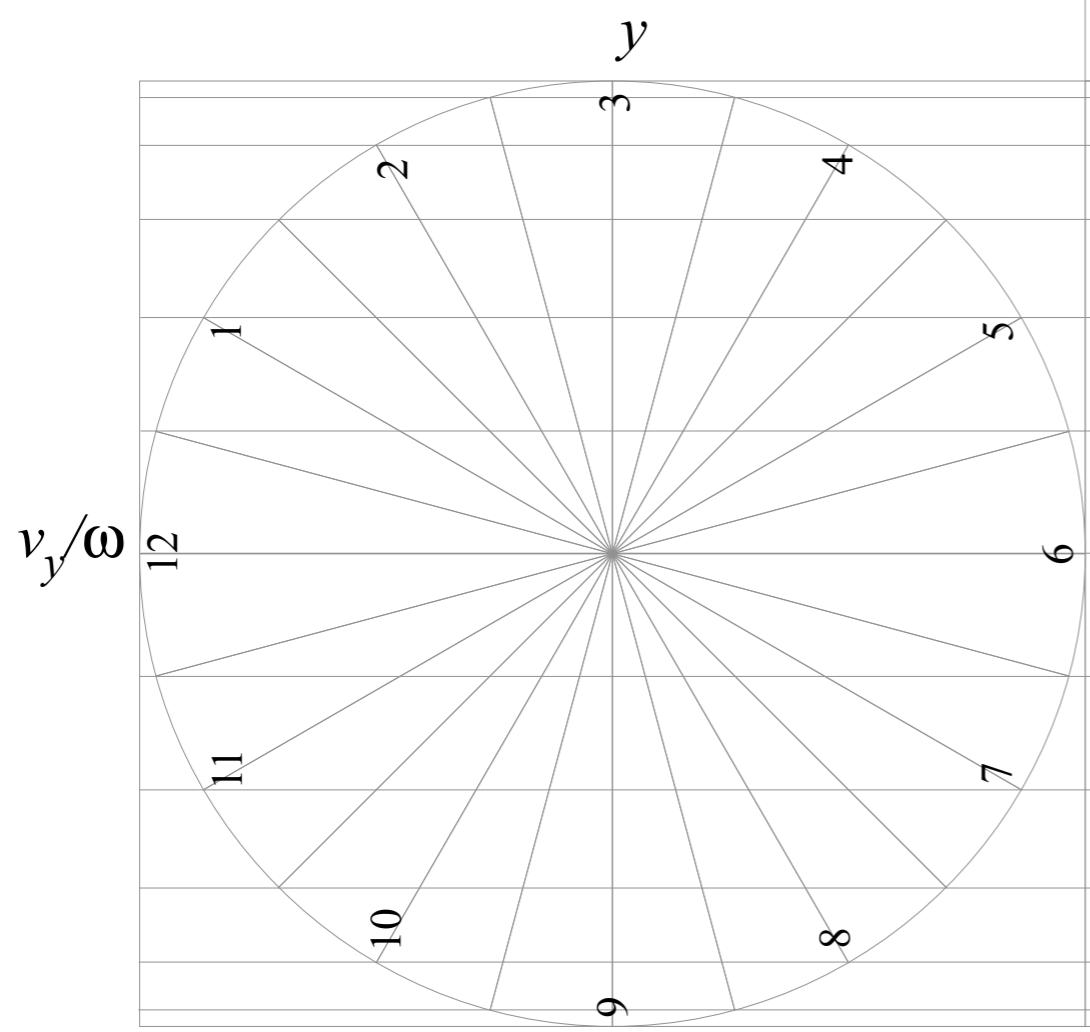
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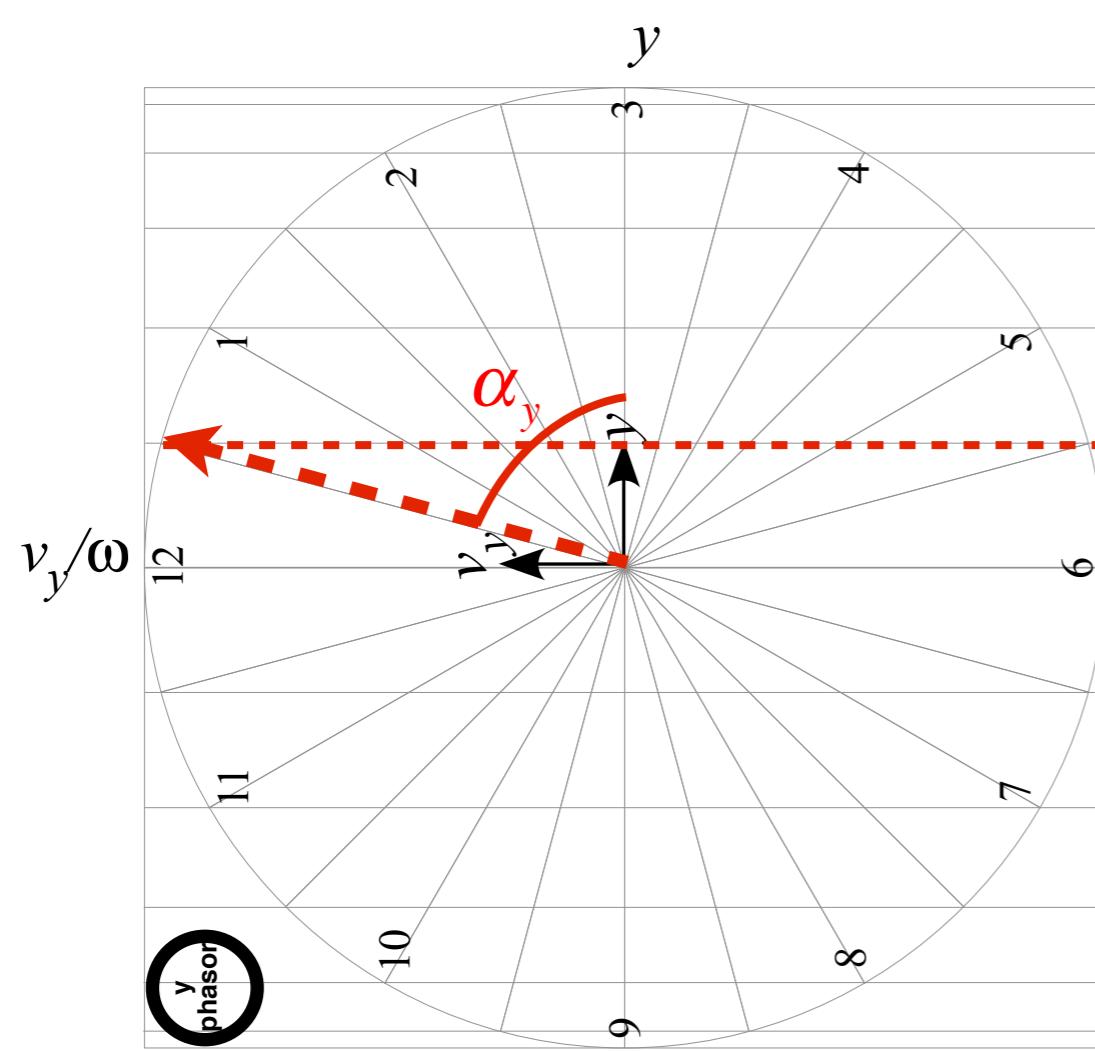
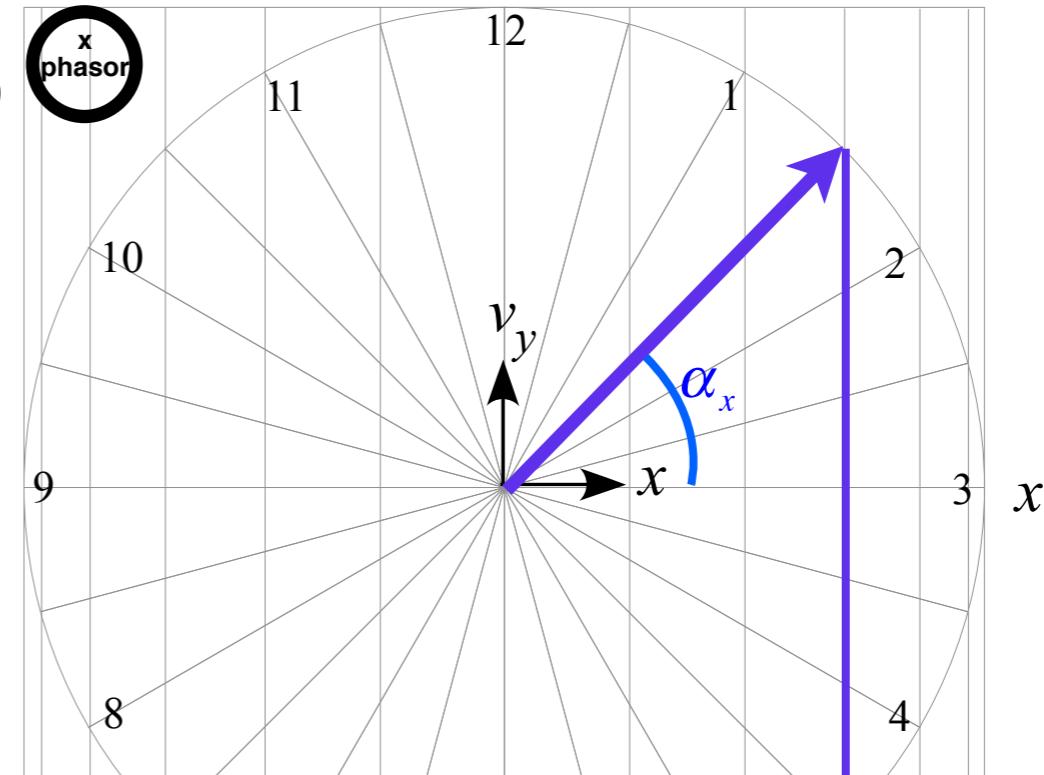
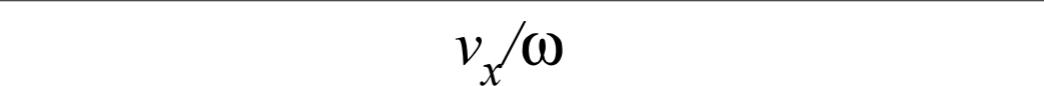


$$x(t) = A_x \cos(\omega \cdot t - \alpha_x) = A_x \cos(\alpha_x - \omega \cdot t)$$

$$\frac{v_x(t)}{\omega} = -A_x \sin(\omega \cdot t - \alpha_x) = A_x \sin(\alpha_x - \omega \cdot t)$$

$$y(t) = A_y \cos(\omega \cdot t - \alpha_y) = A_y \cos(\alpha_y - \omega \cdot t)$$

$$\frac{v_y(t)}{\omega} = -A_y \sin(\omega \cdot t - \alpha_y) = A_y \sin(\alpha_y - \omega \cdot t)$$



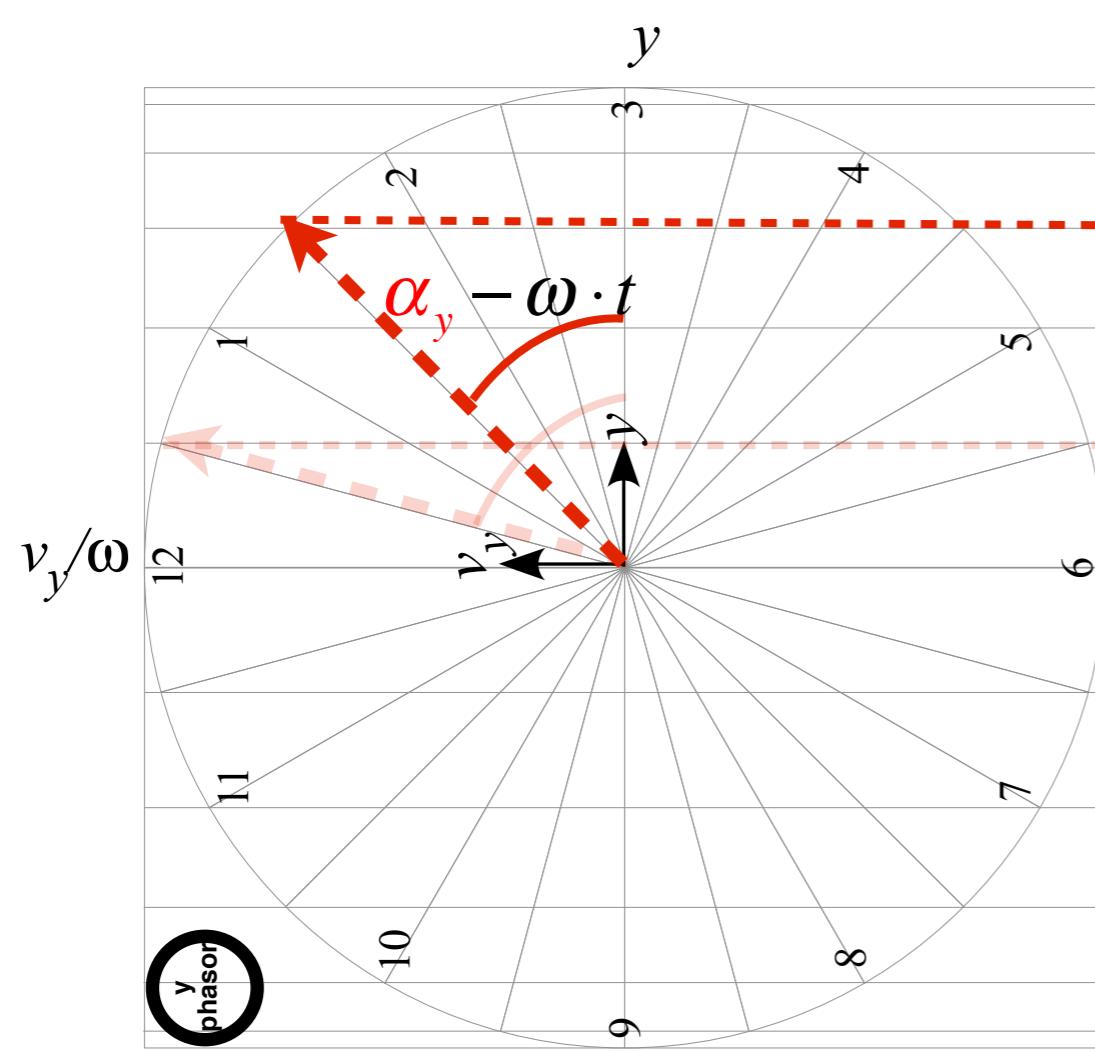
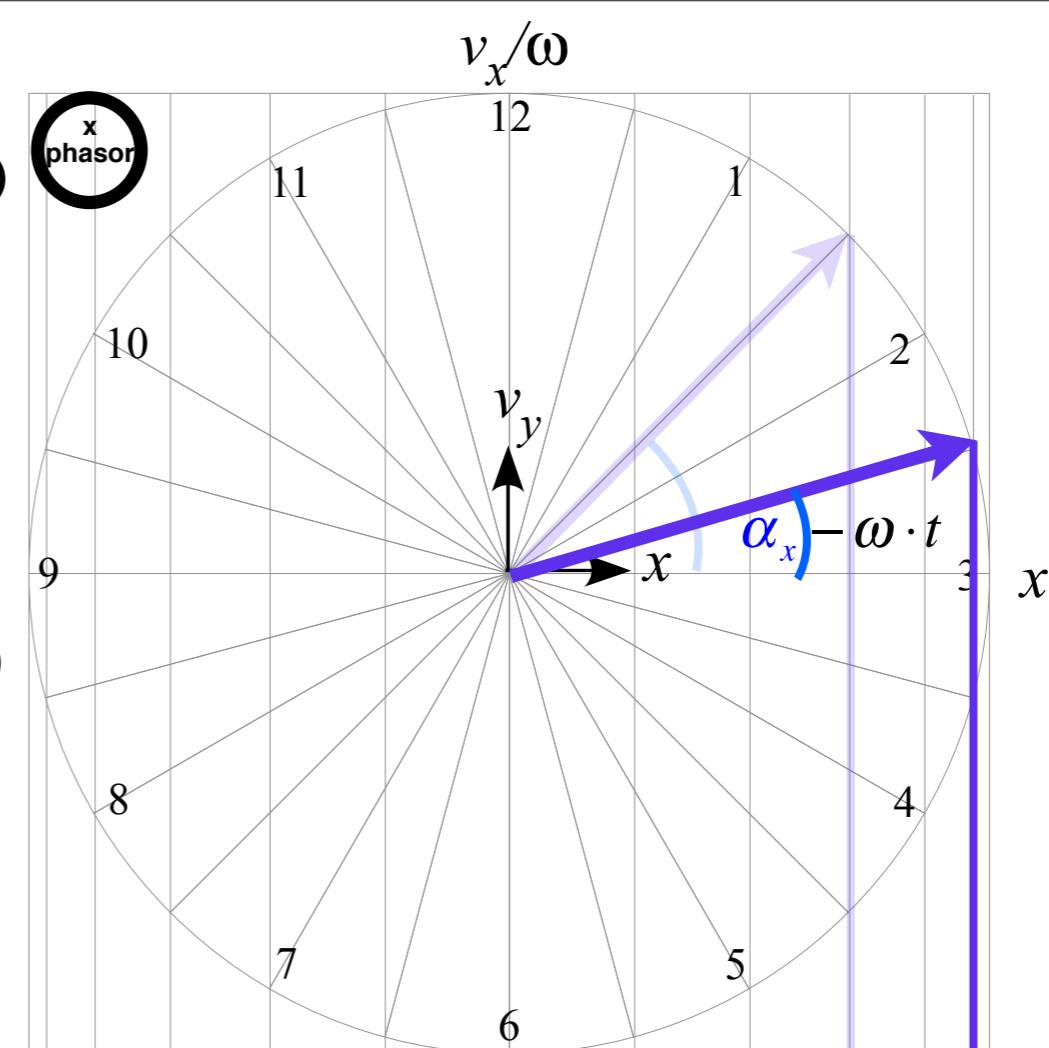
Initial phases at t=0

$$x(t) = A_x \cos(\omega \cdot t - \alpha_x) = A_x \cos(\alpha_x - \omega \cdot t)$$

$$\frac{v_x(t)}{\omega} = -A_x \sin(\omega \cdot t - \alpha_x) = A_x \sin(\alpha_x - \omega \cdot t)$$

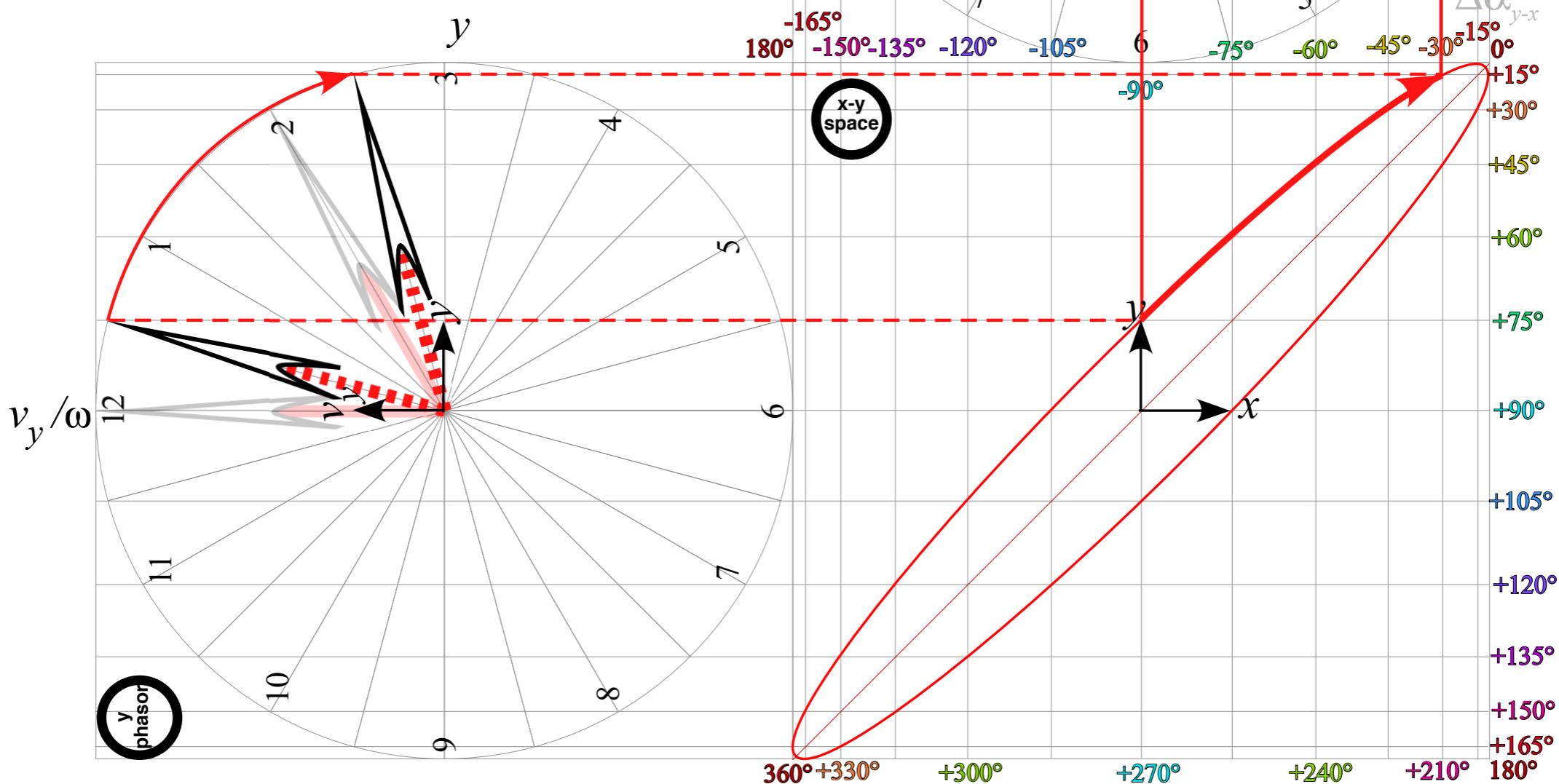
$$y(t) = A_y \cos(\omega \cdot t - \alpha_y) = A_y \cos(\alpha_y - \omega \cdot t)$$

$$\frac{v_y(t)}{\omega} = -A_y \sin(\omega \cdot t - \alpha_y) = A_y \sin(\alpha_y - \omega \cdot t)$$



Later phases at time t

Ellipsometry Contact Plots
vs.
Relative phase $\Delta\alpha = \alpha_y - \alpha_x$
 $\Delta\alpha = +15^\circ$
(Left-polarized clockwise case)



+30° case

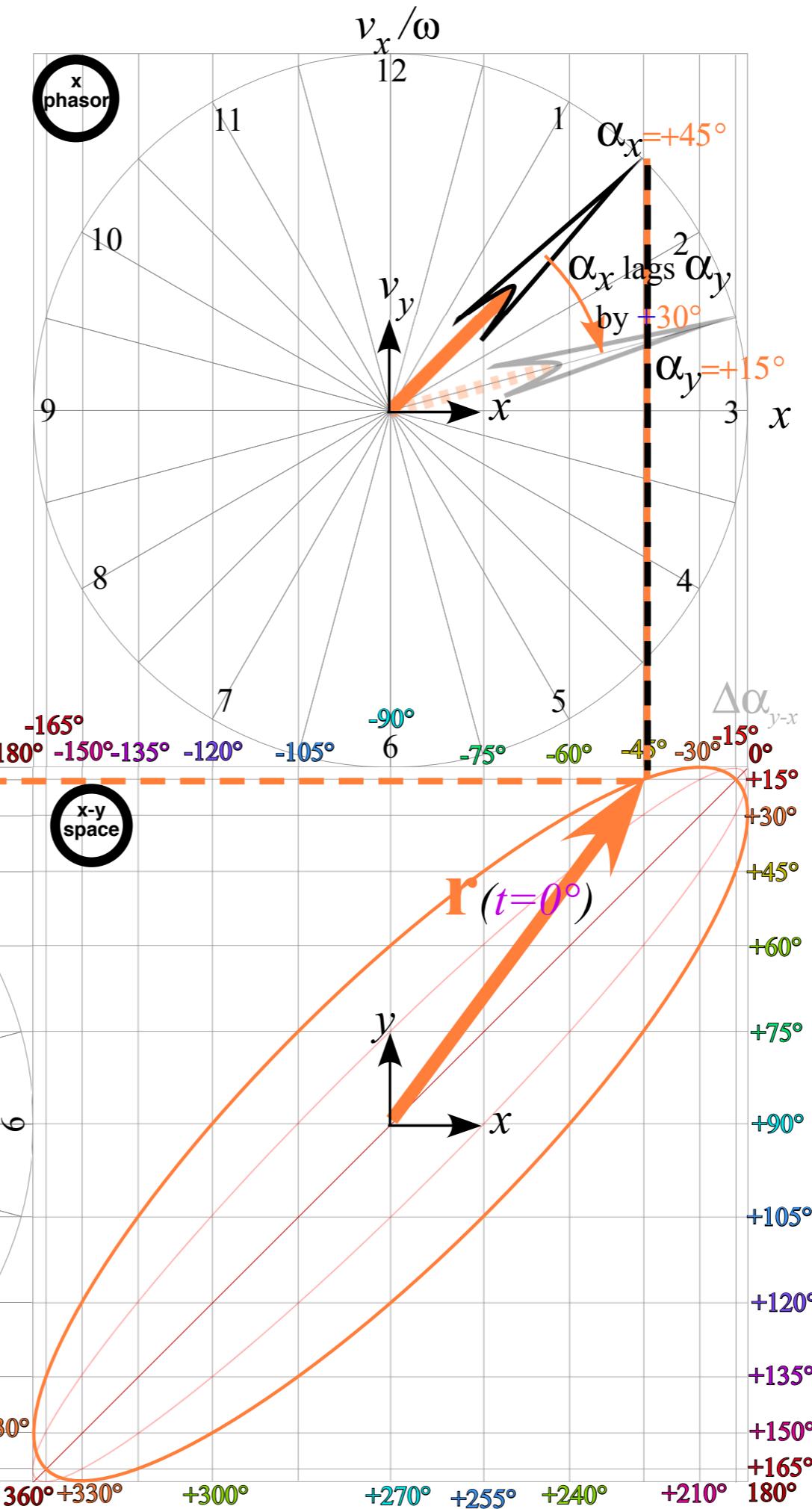
Ellipsometry Contact Plots vs.

Relative phase $\Delta\alpha = \alpha_y - \alpha_x$

$$\Delta\alpha = \alpha_x - \alpha_y = +30^\circ$$

Left-polarized (clockwise) orbit

Initial phase ($\alpha_x = 45^\circ, \alpha_y = +15^\circ$)



+30° case

Ellipsometry Contact Plots

vs.

Relative phase $\Delta\alpha = \alpha_y - \alpha_x$

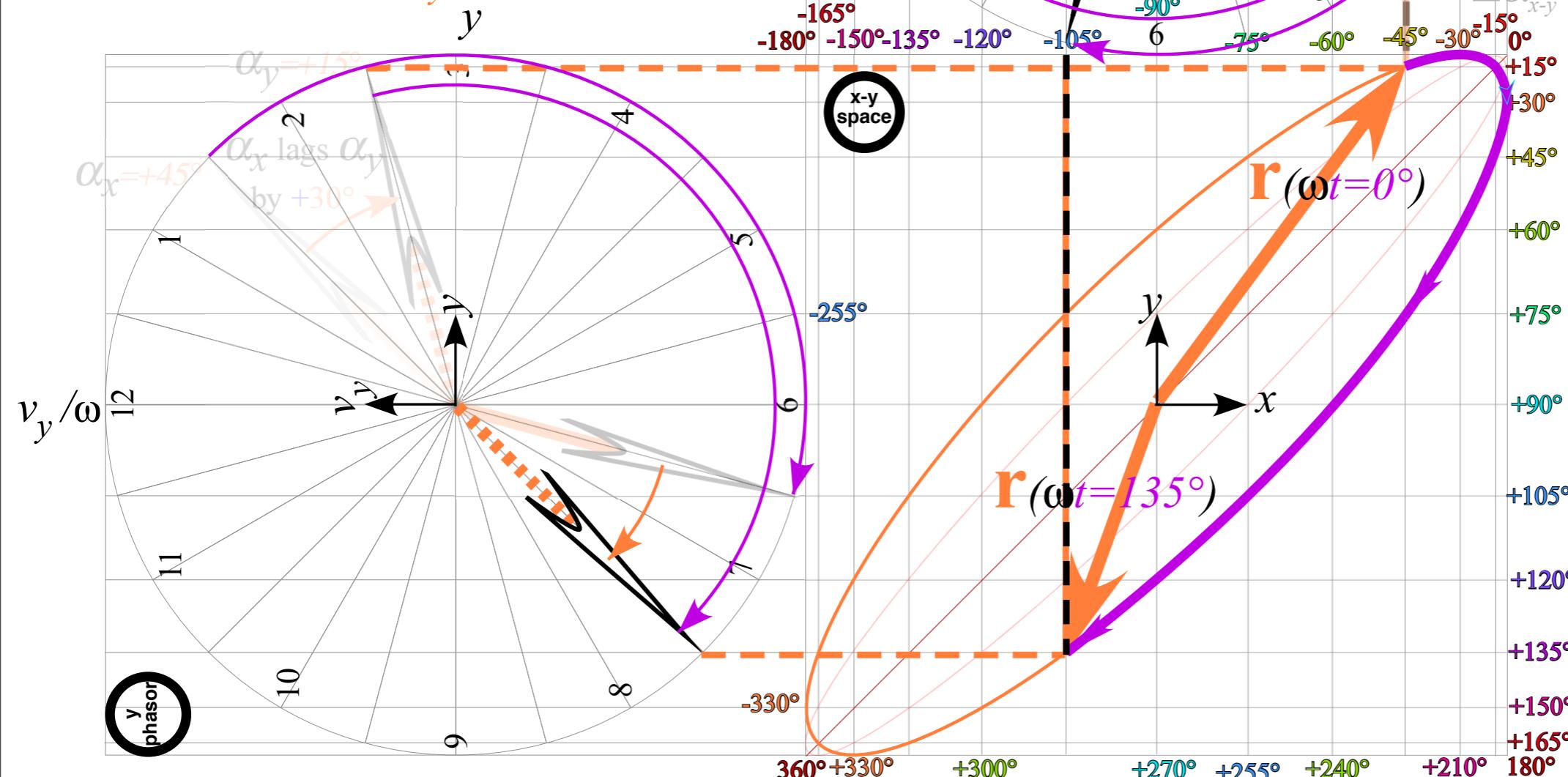
$$\Delta\alpha = \alpha_x - \alpha_y = +30^\circ$$

Left-polarized (clockwise) orbit

Initial phase ($\alpha_x = 45^\circ, \alpha_y = +15^\circ$)

...after time advances by 150°:

$$(\alpha_x = 45^\circ - 150^\circ = -105^\circ, \alpha_y = +15^\circ - 150^\circ = -135^\circ)$$



-75° case

Ellipsometry Contact Plots

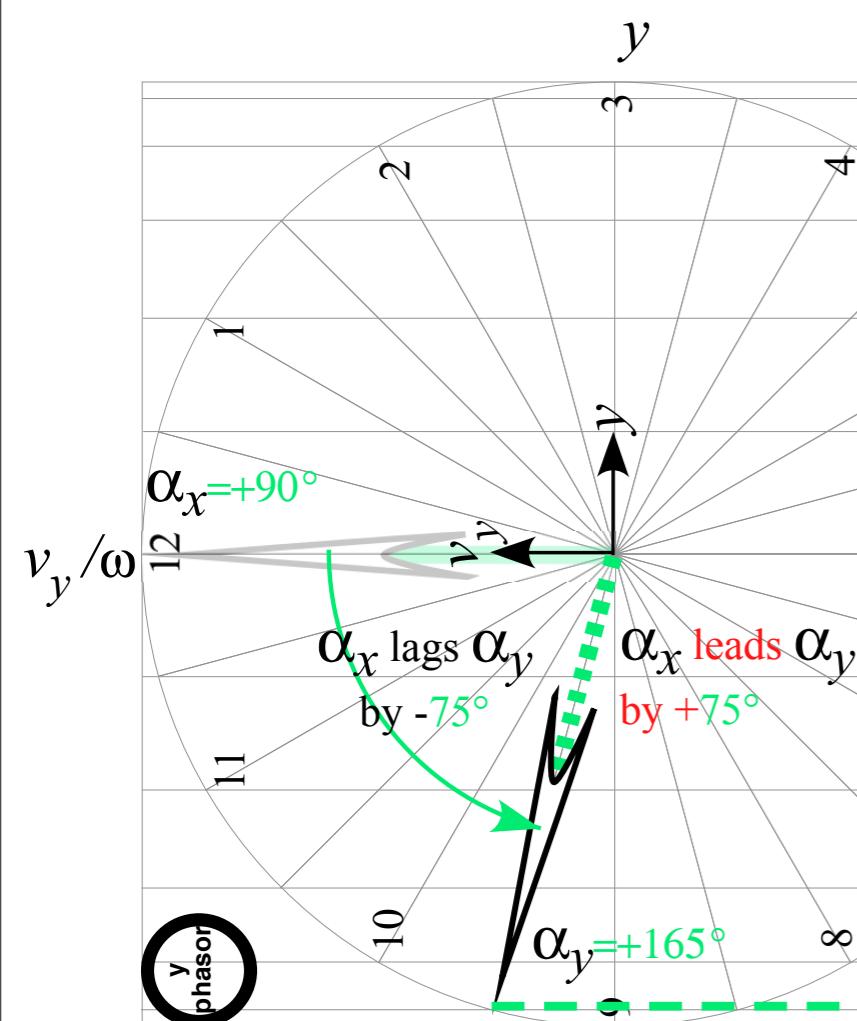
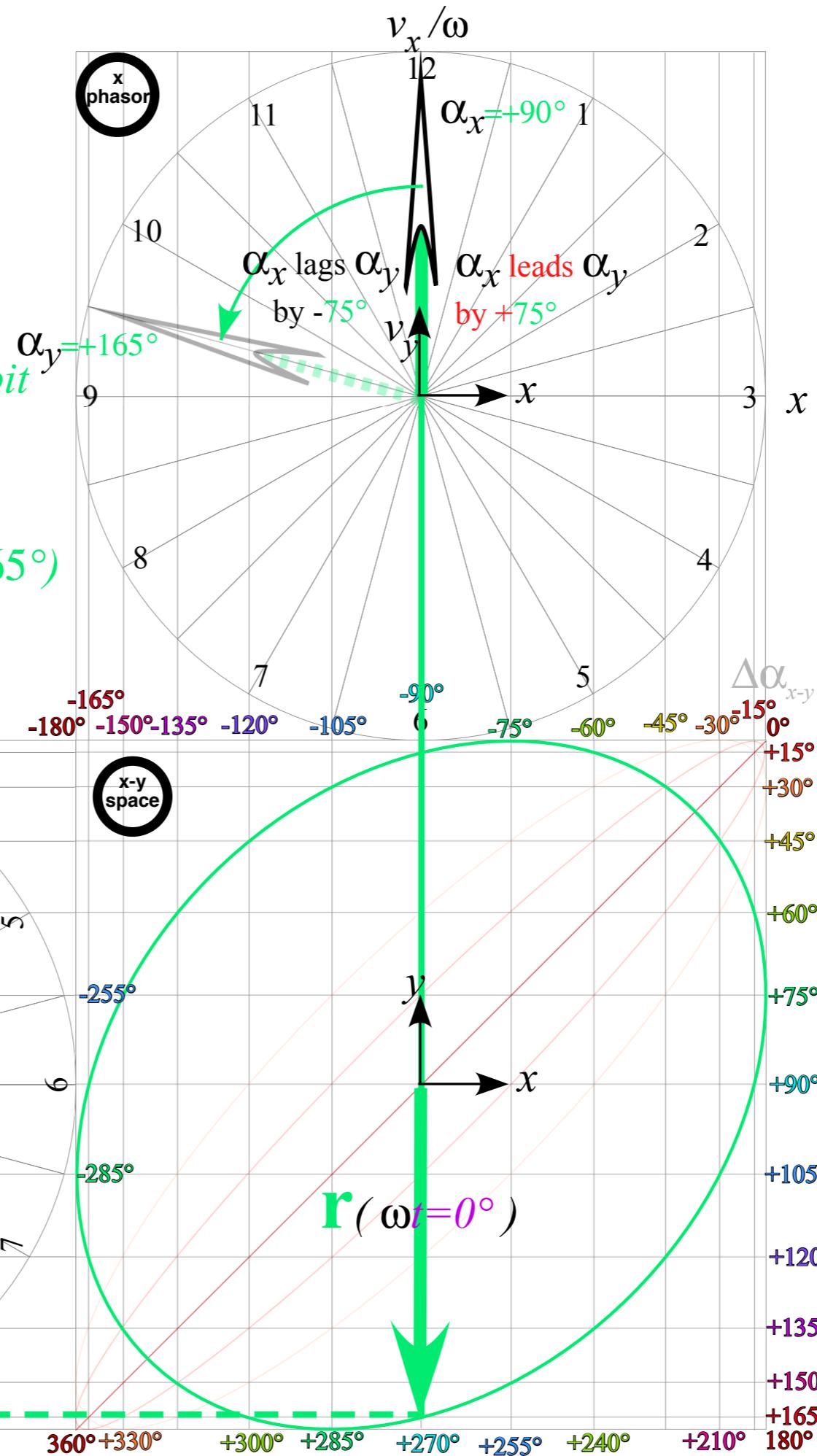
vs.

Relative phase $\Delta\alpha = \alpha_y - \alpha_x$

$$\Delta\alpha = \alpha_x - \alpha_y = -75^\circ$$

Right-polarized (anti-clockwise) orbit

Initial phase ($\alpha_x = 90^\circ, \alpha_y = +165^\circ$)



-75° case

Ellipsometry Contact Plots

vs.

Relative phase $\Delta\alpha = \alpha_y - \alpha_x$

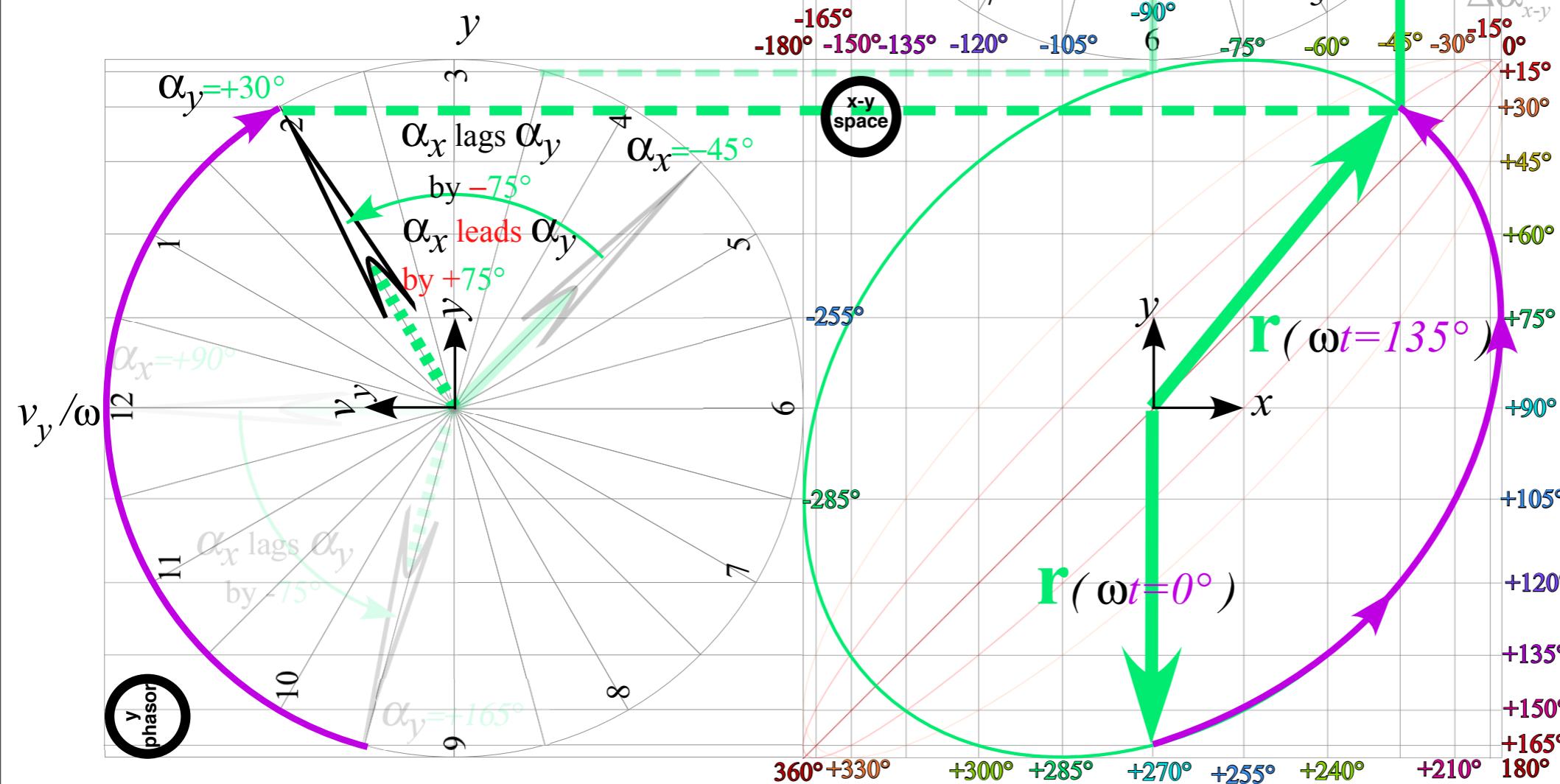
$$\Delta\alpha = \alpha_x - \alpha_y = -75^\circ$$

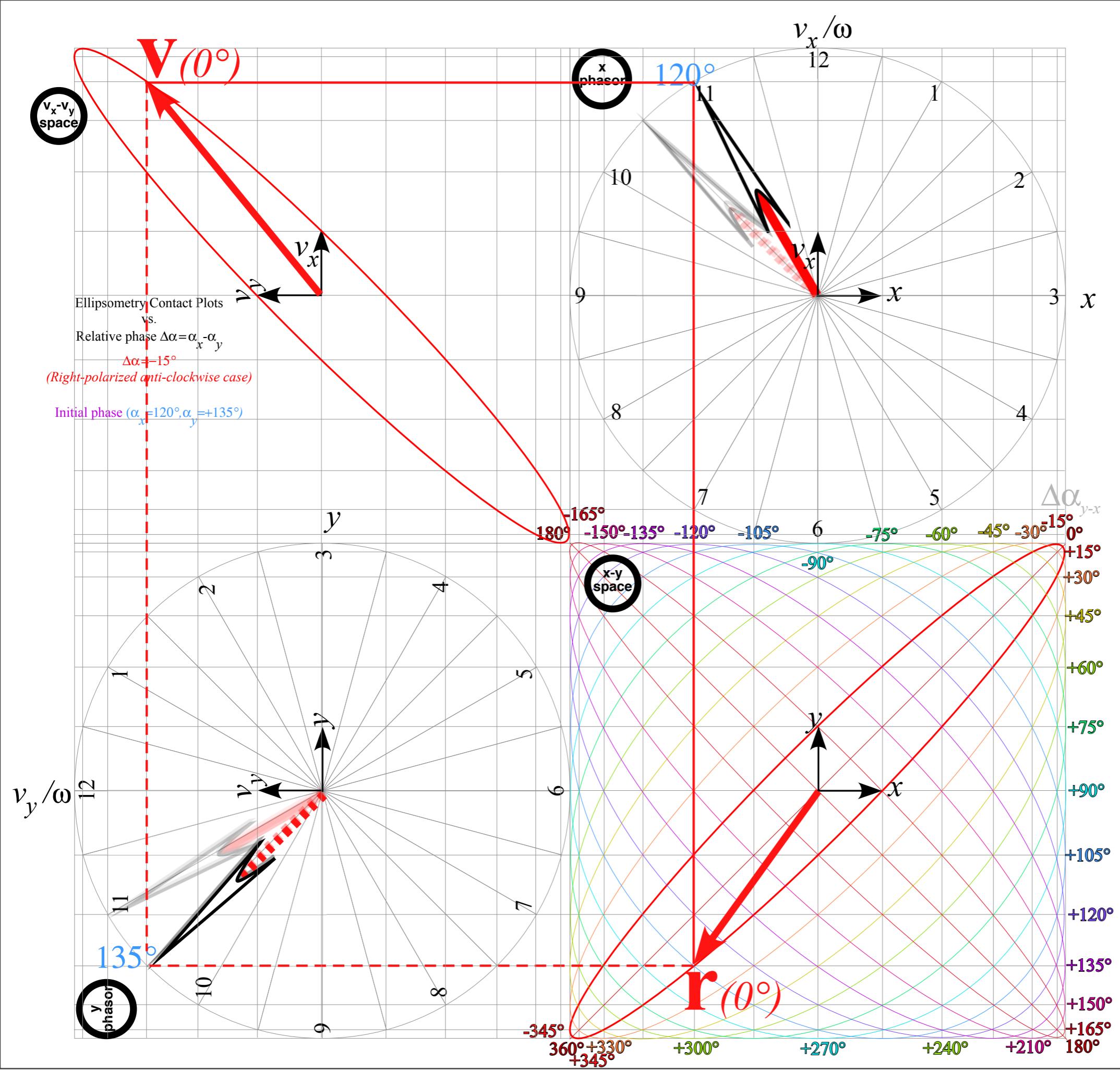
Right-polarized (anti-clockwise) orbit

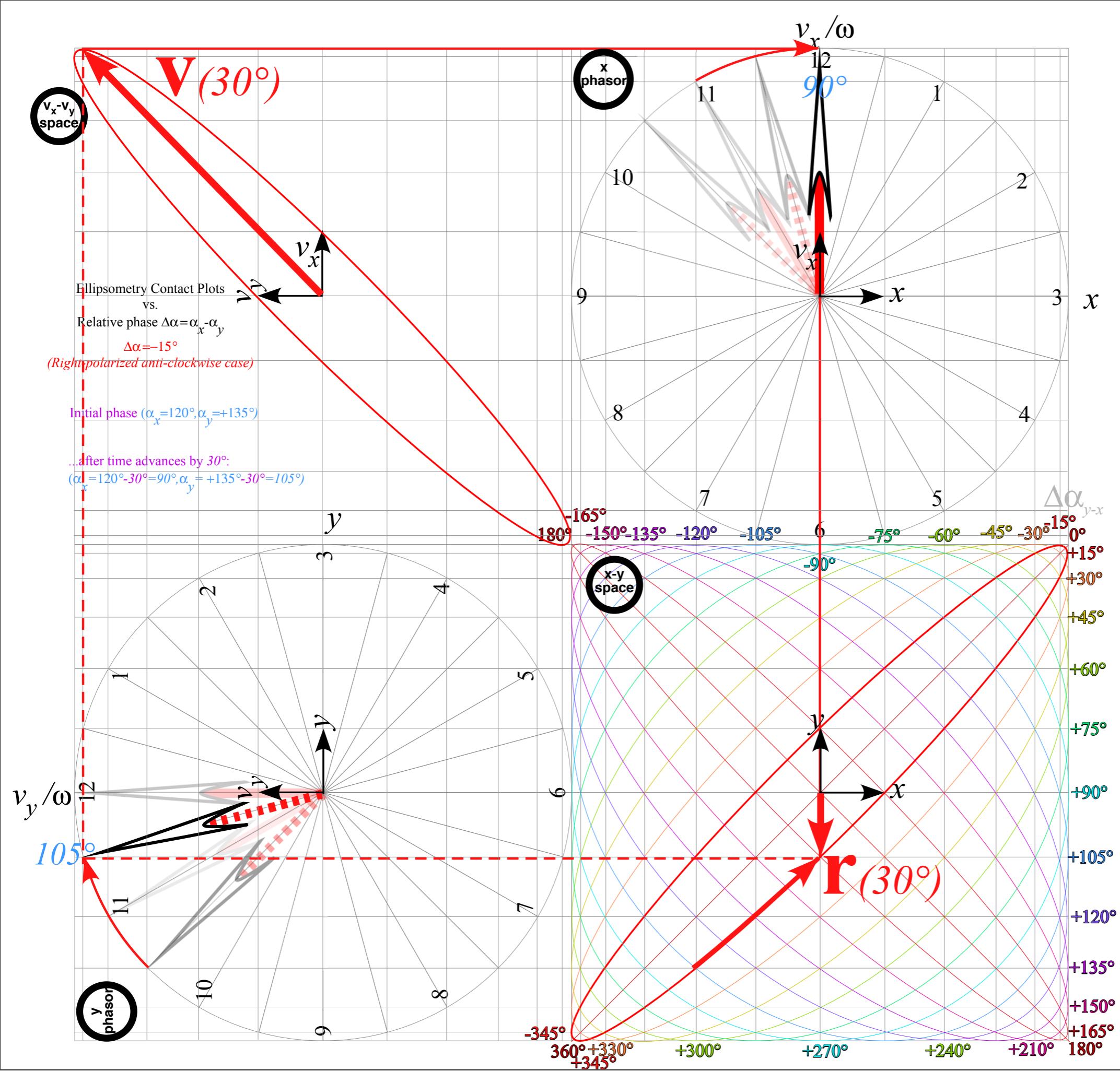
Initial phase ($\alpha_x = 90^\circ, \alpha_y = +165^\circ$)

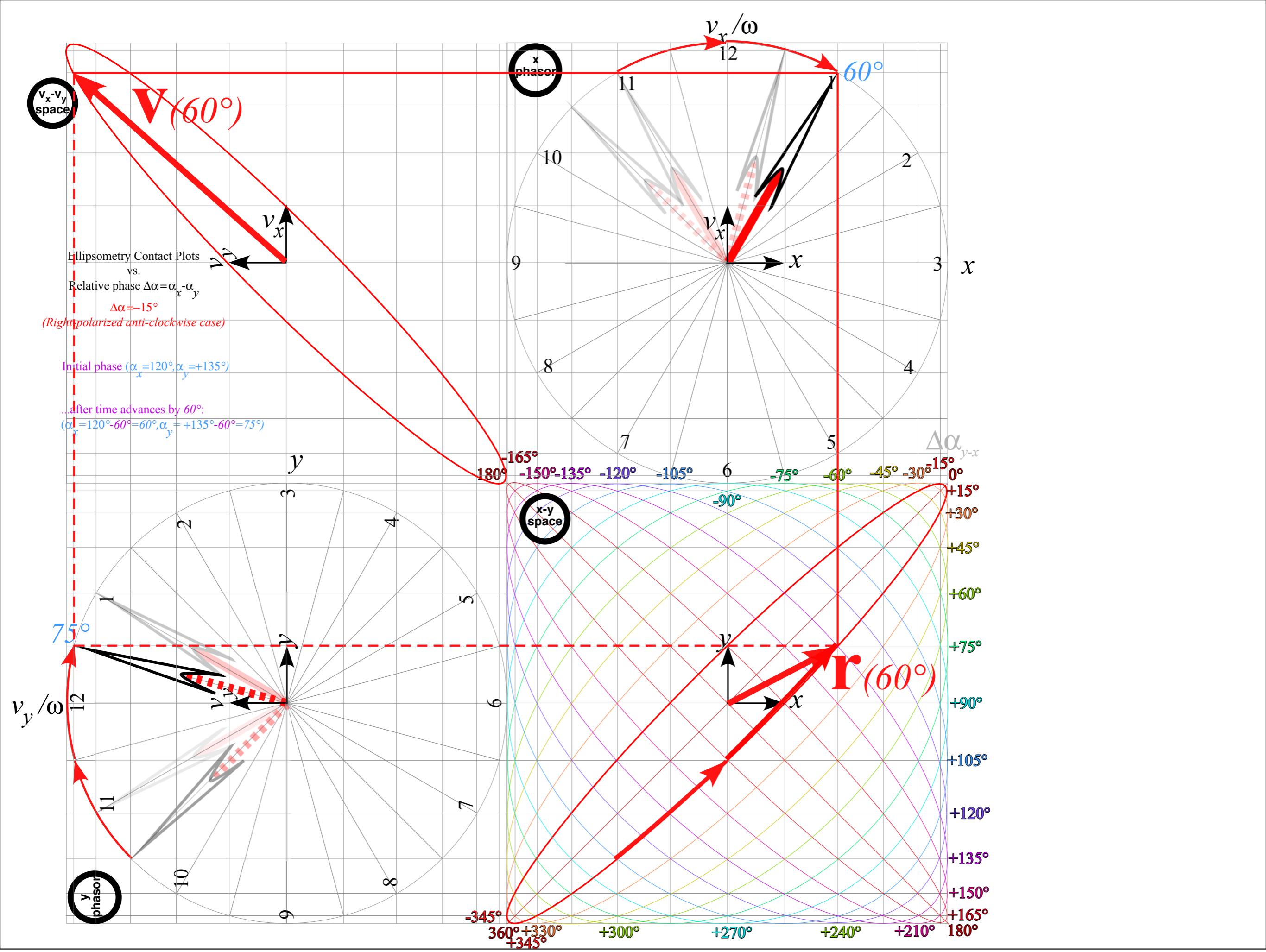
...after time advances by 135° :

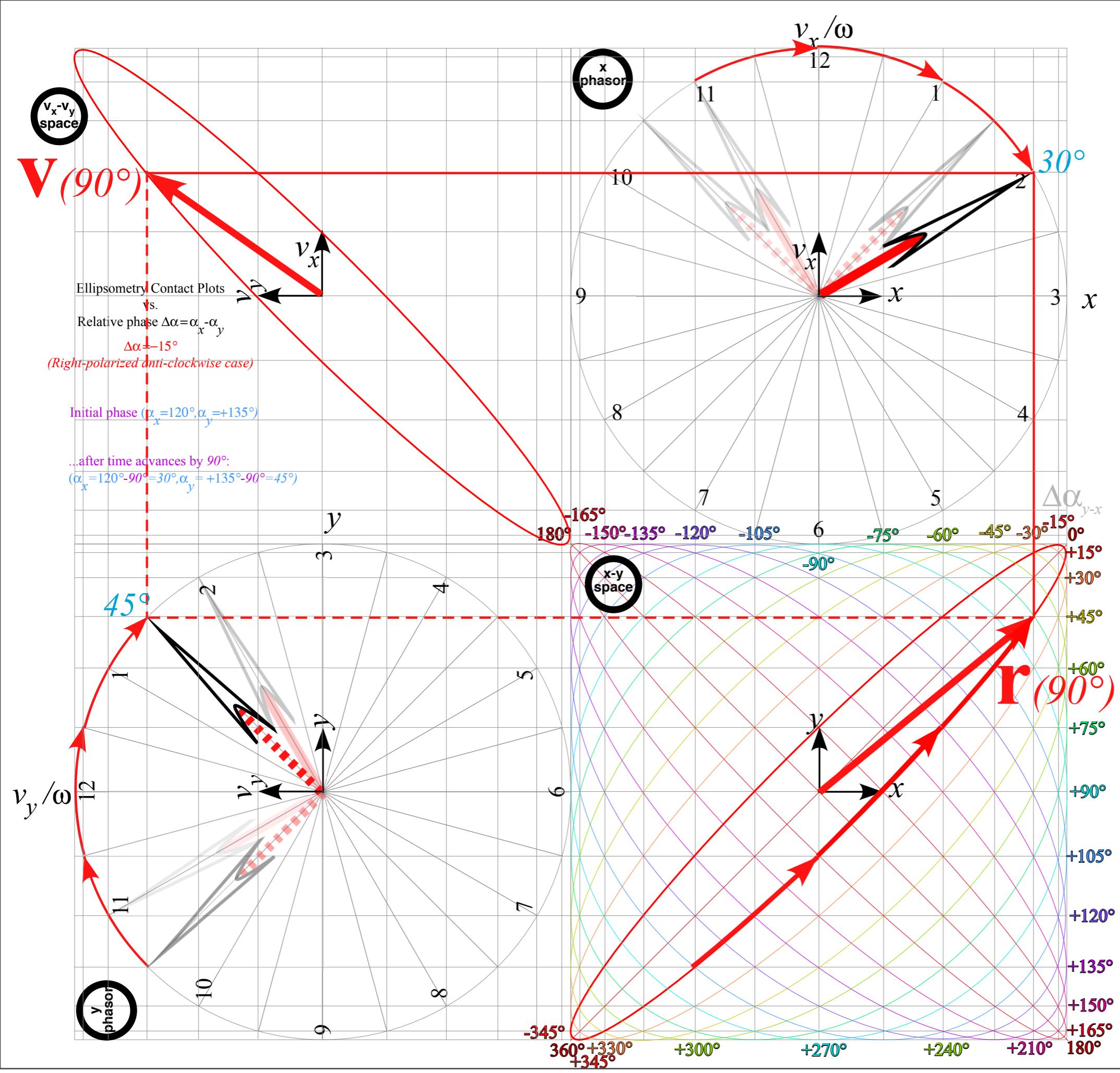
$$(\alpha_x = 90^\circ - 135^\circ = -45^\circ, \alpha_y = +165^\circ - 135^\circ = +30^\circ)$$

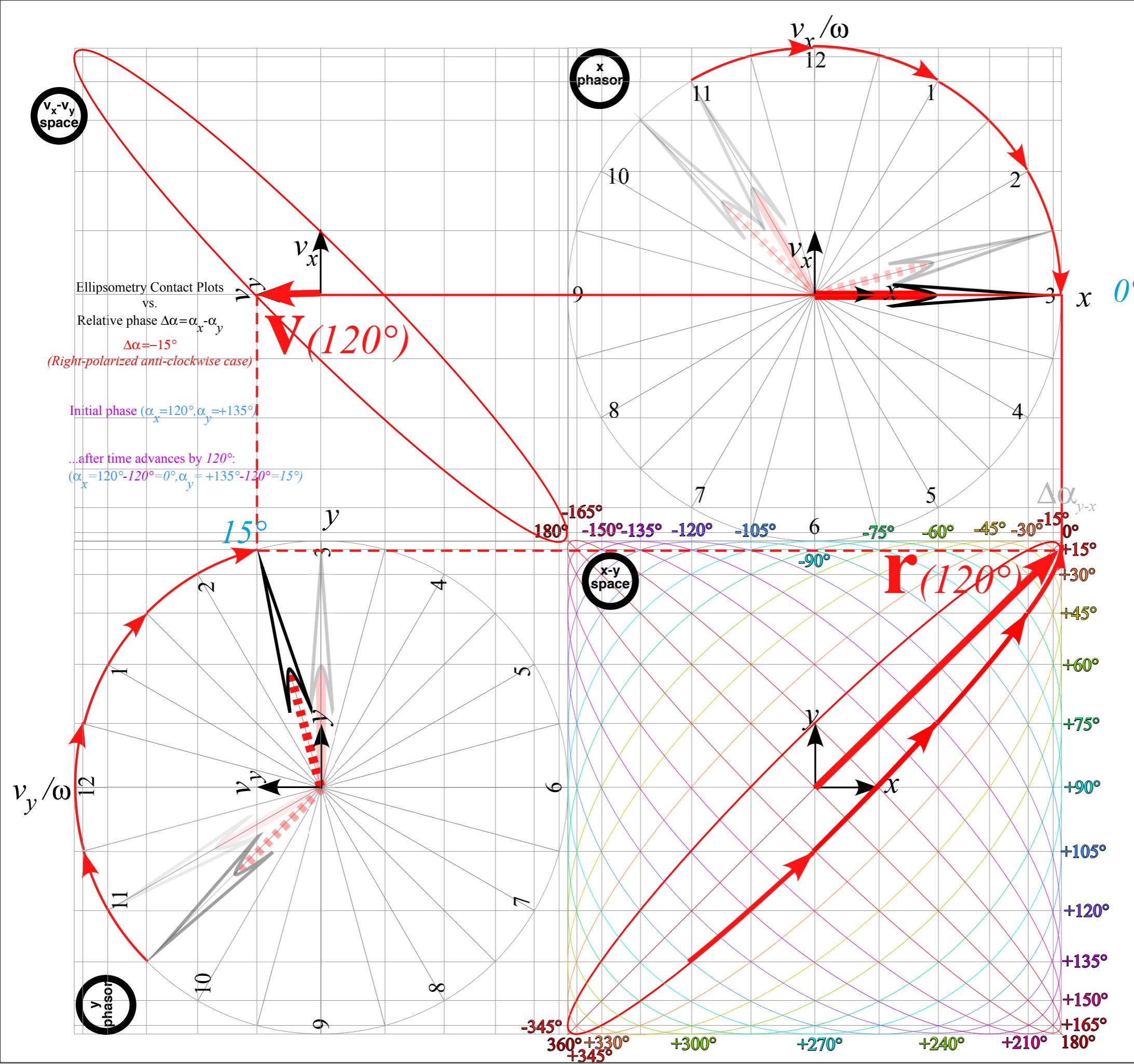


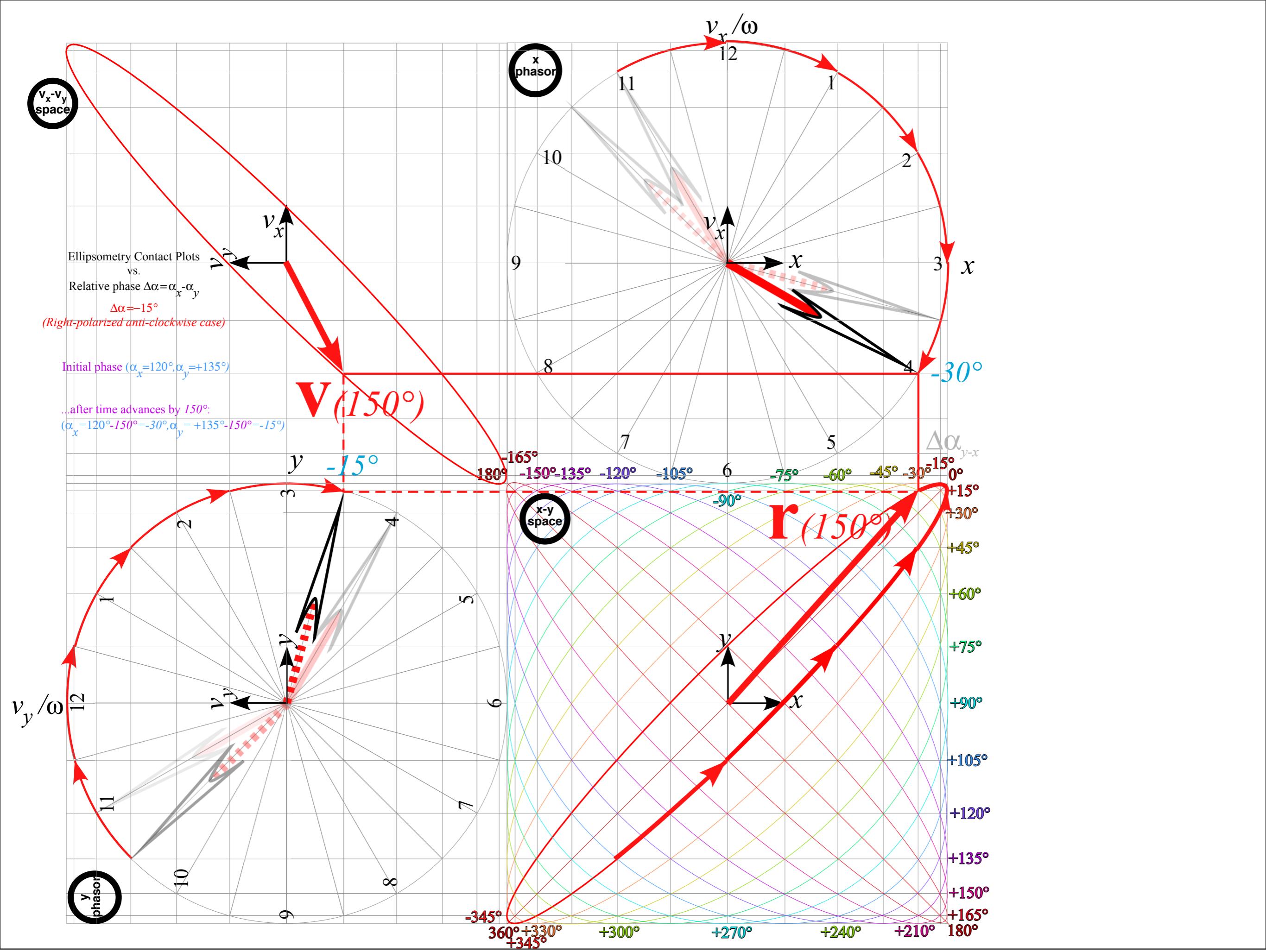


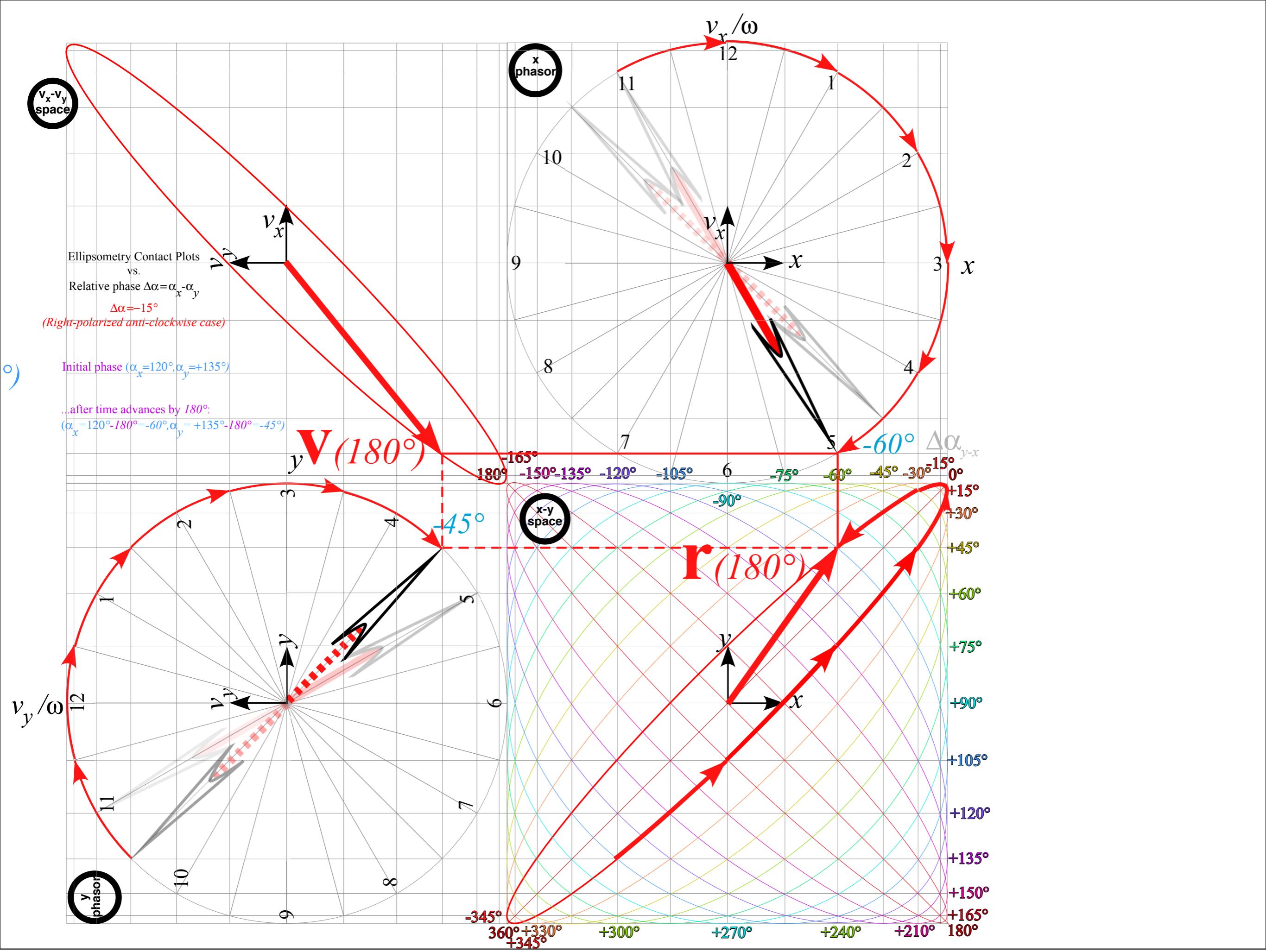


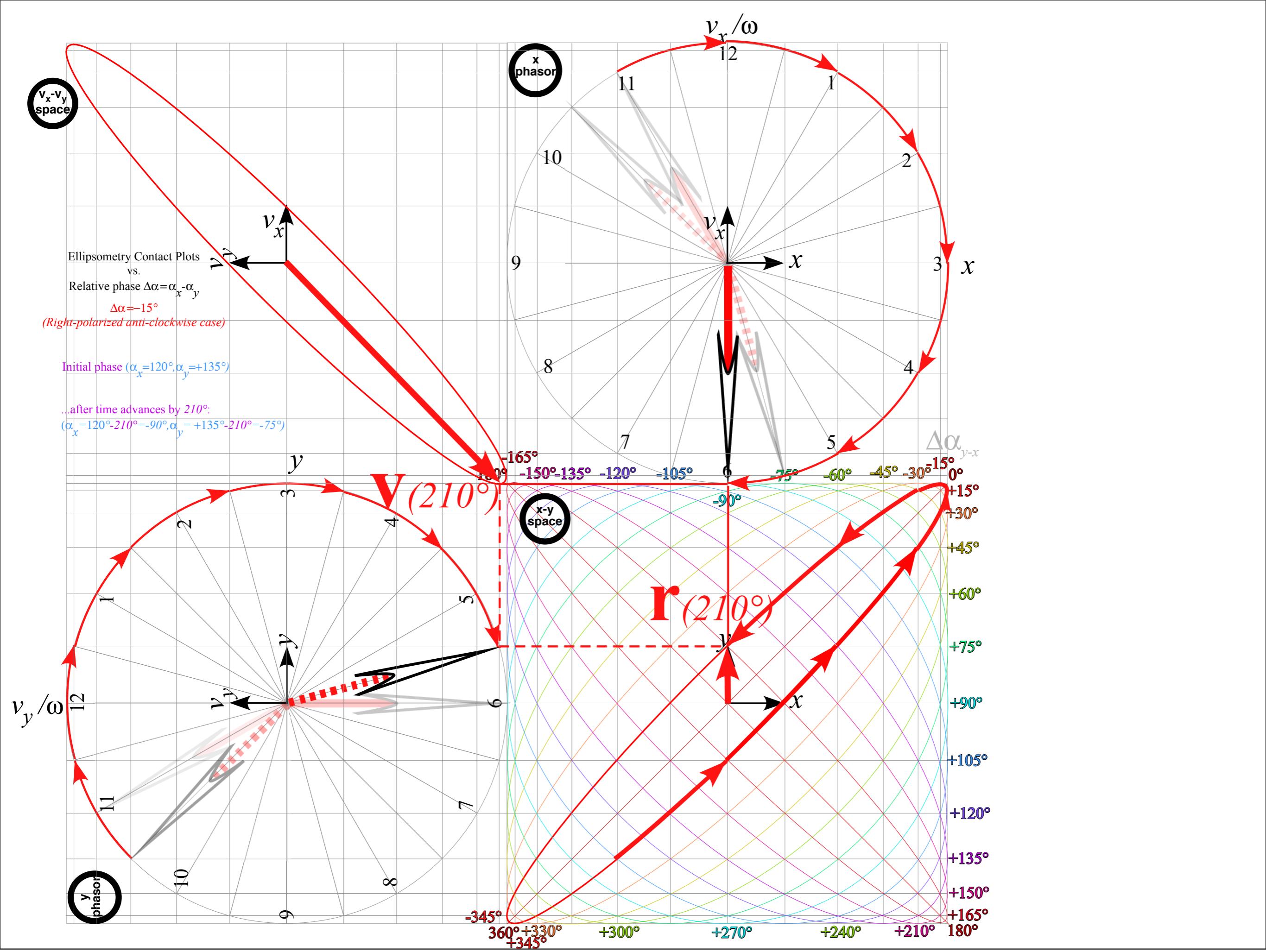


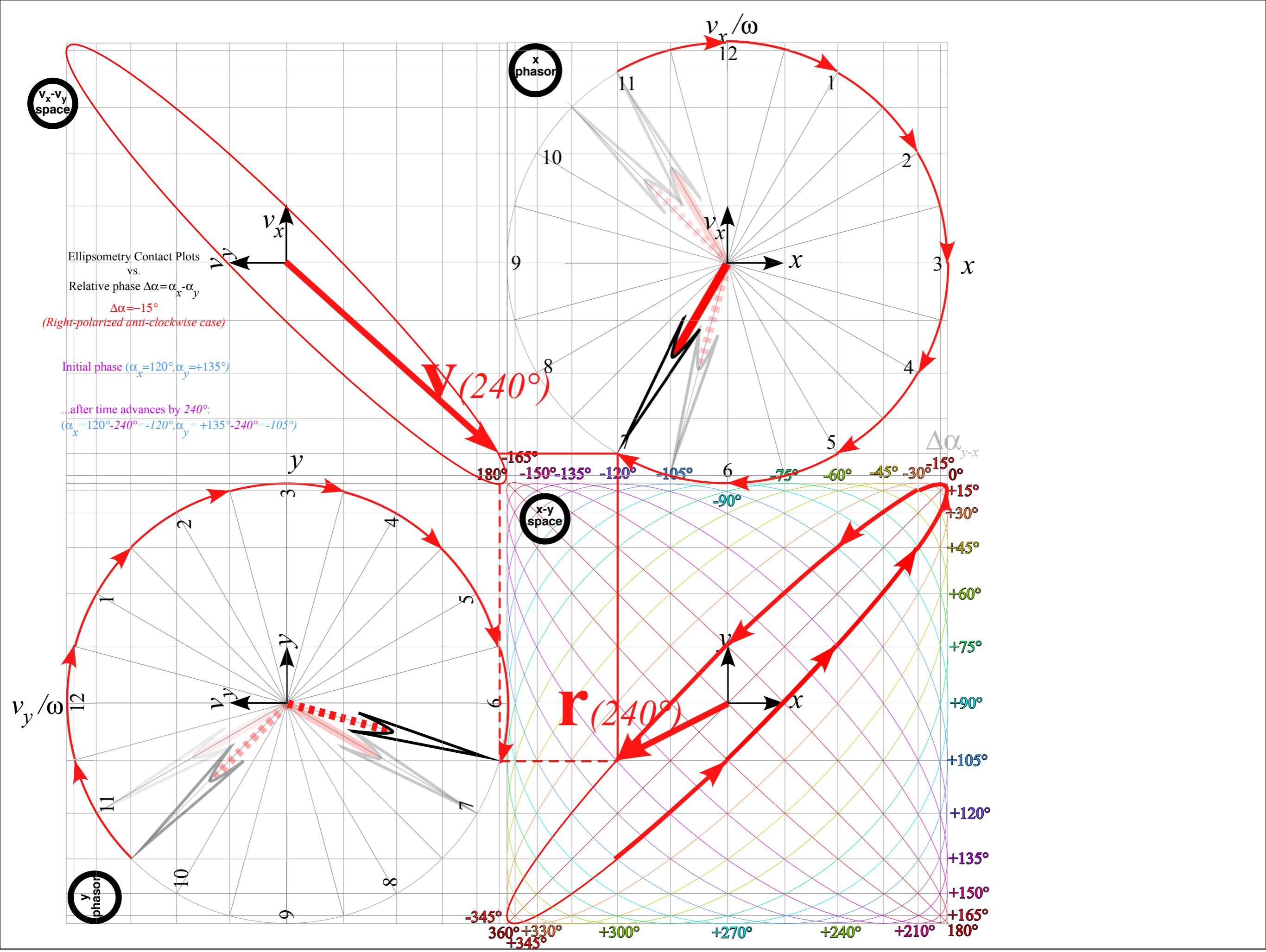


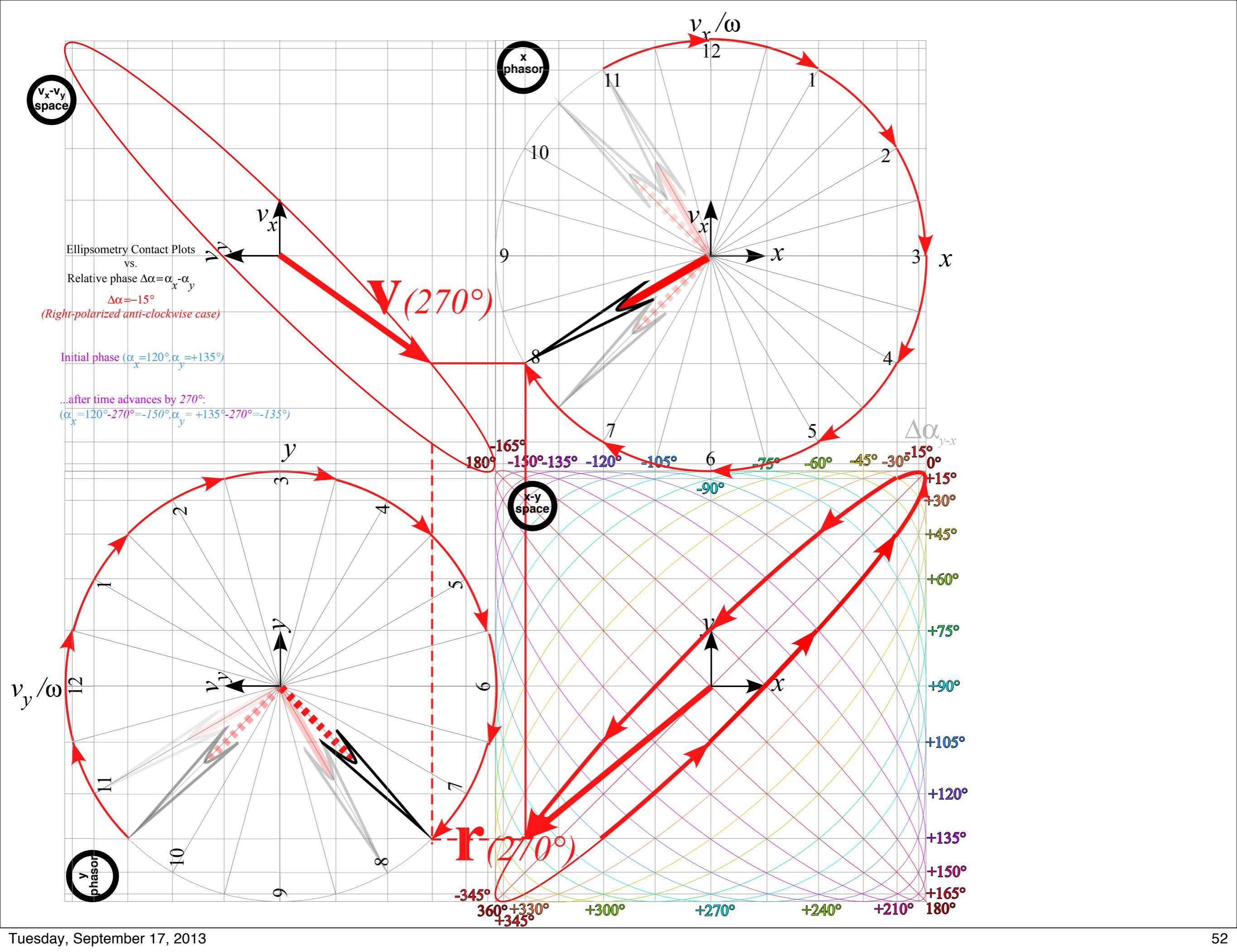


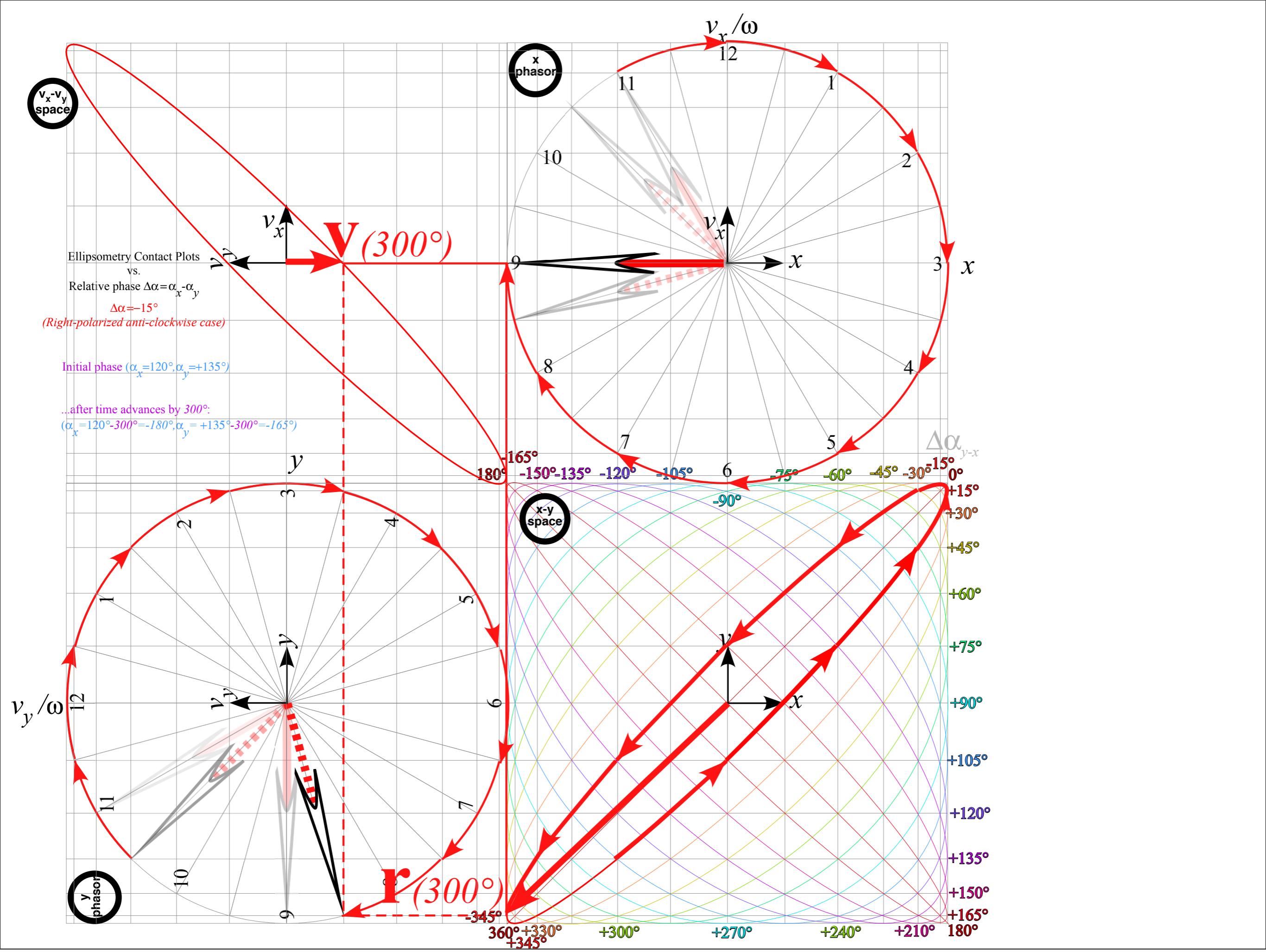


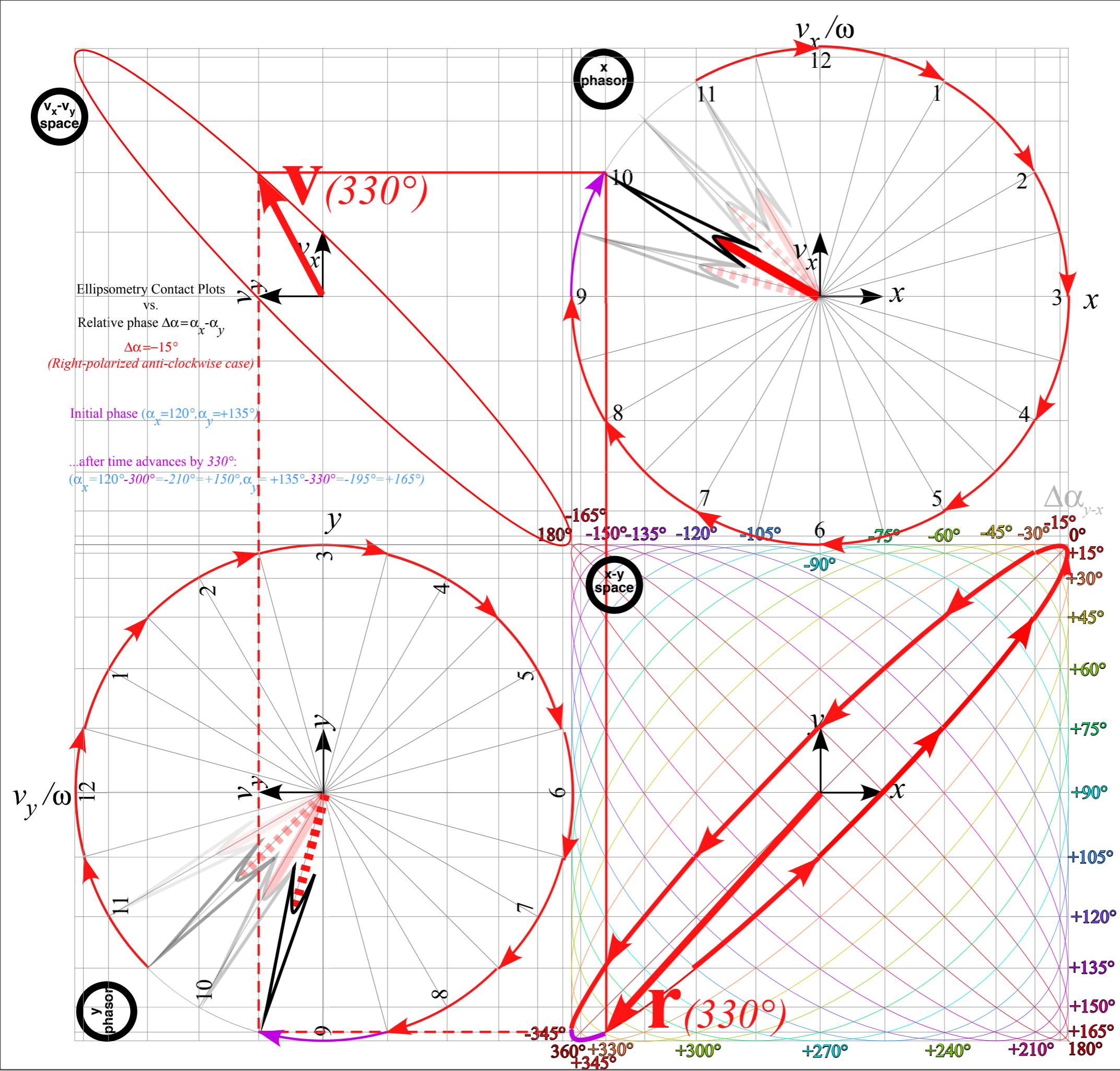


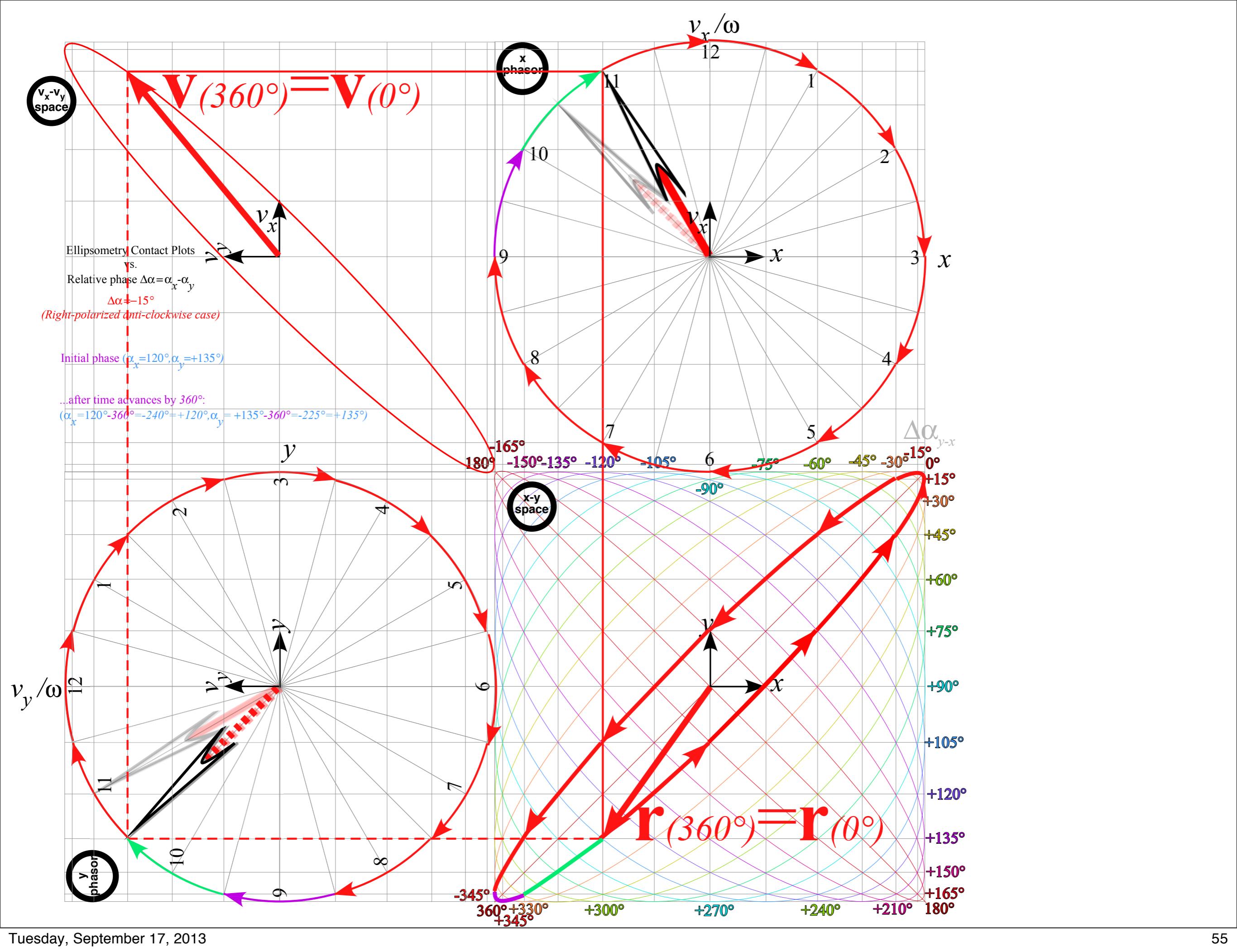


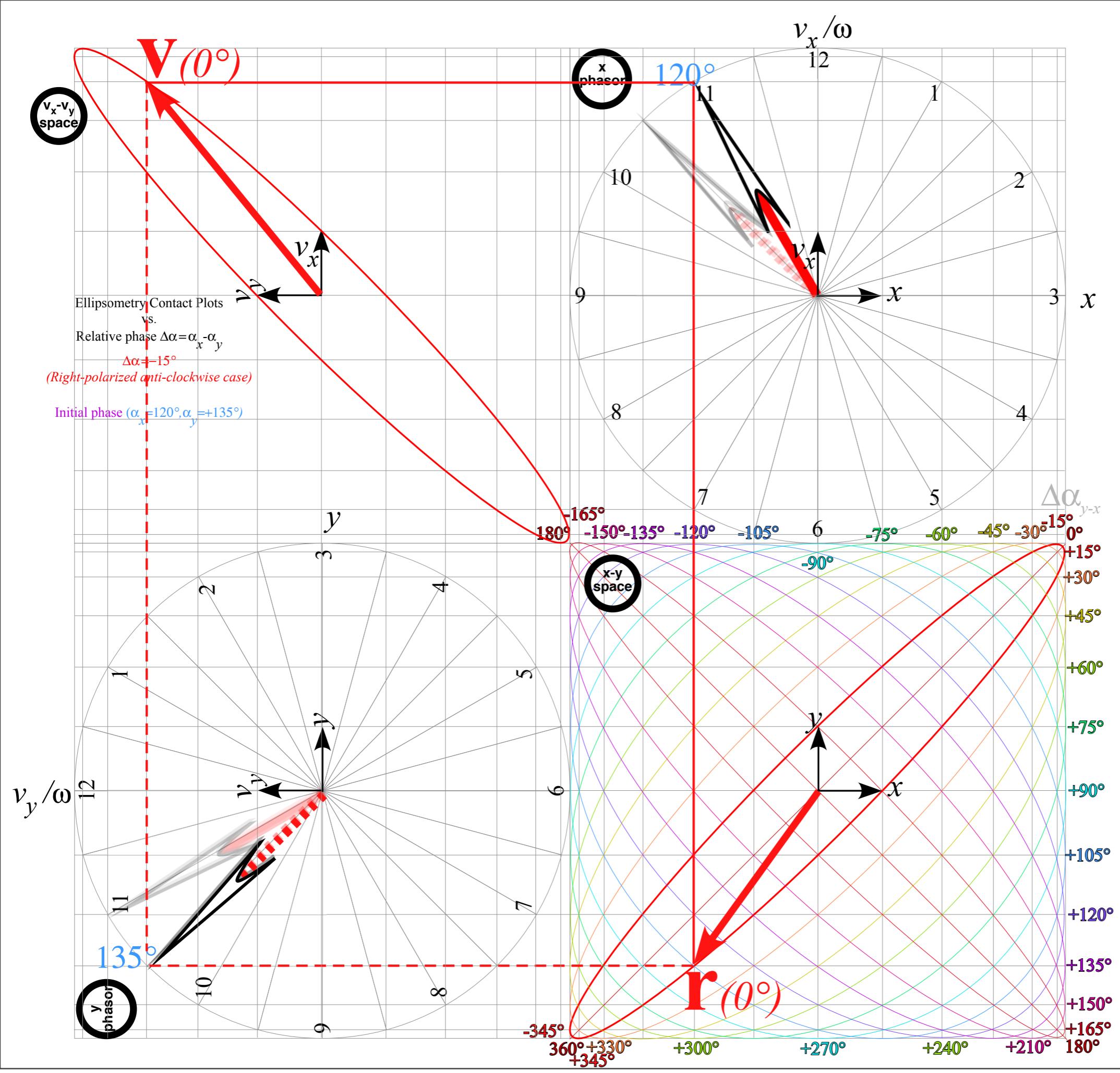


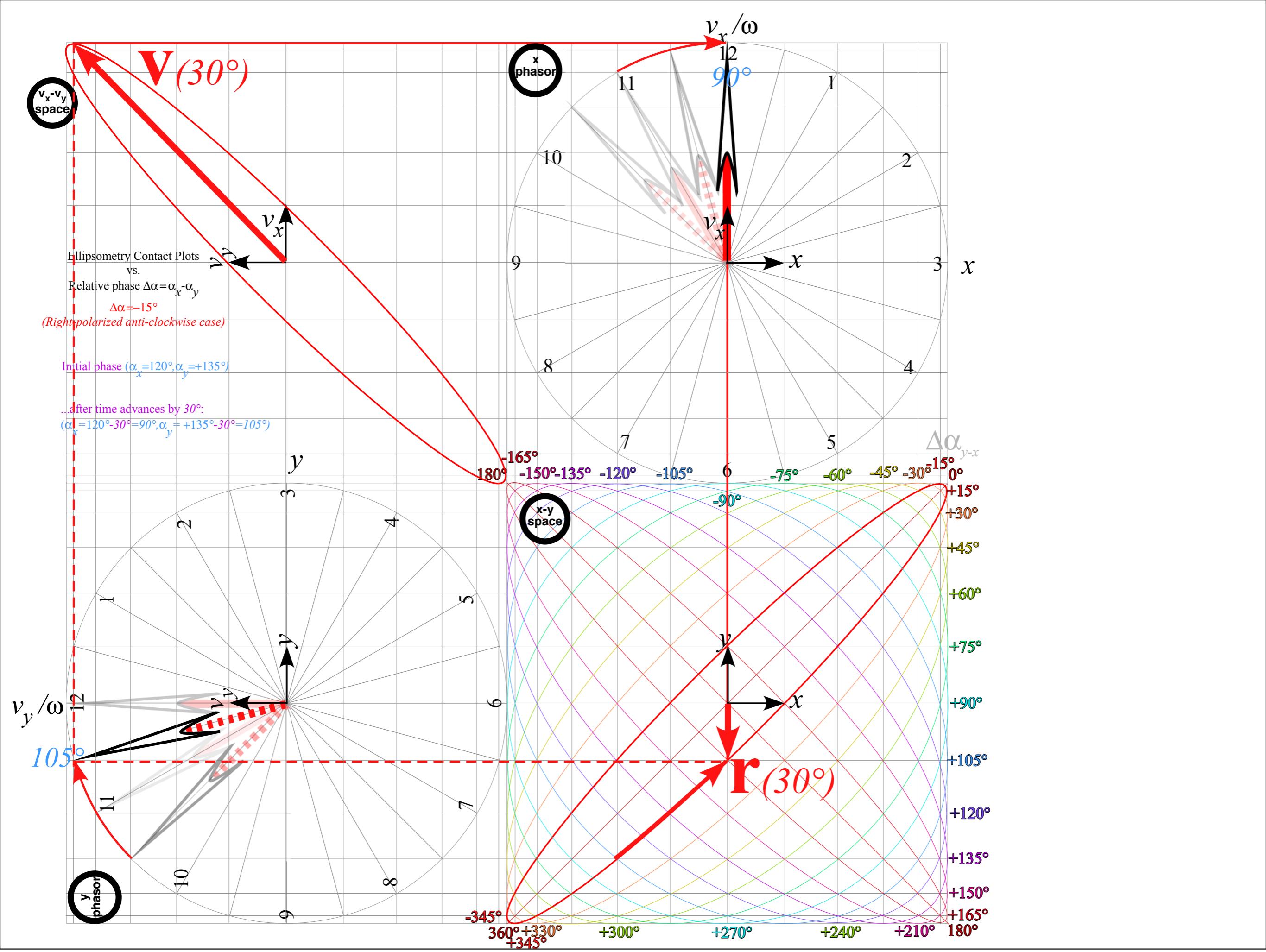


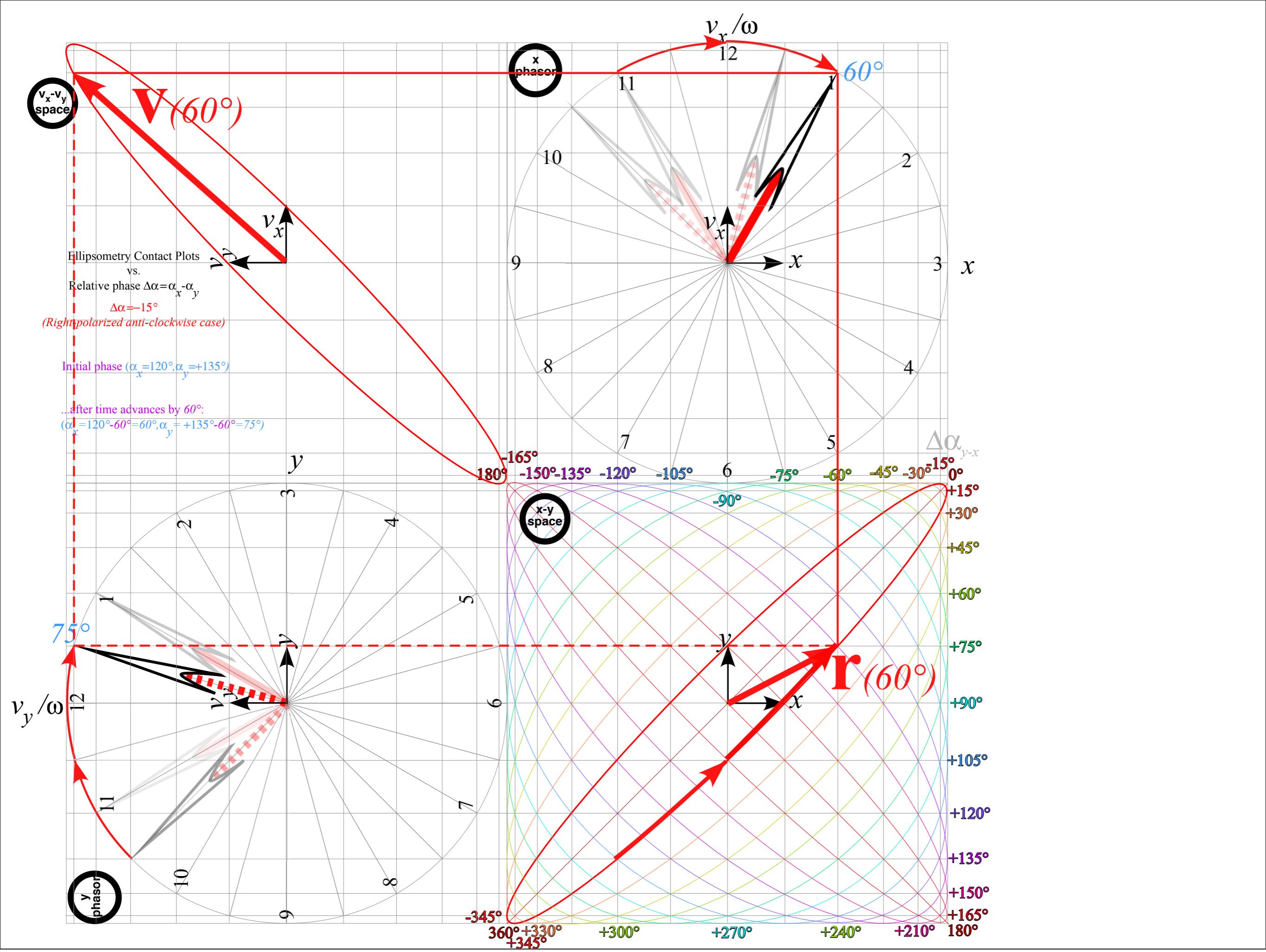


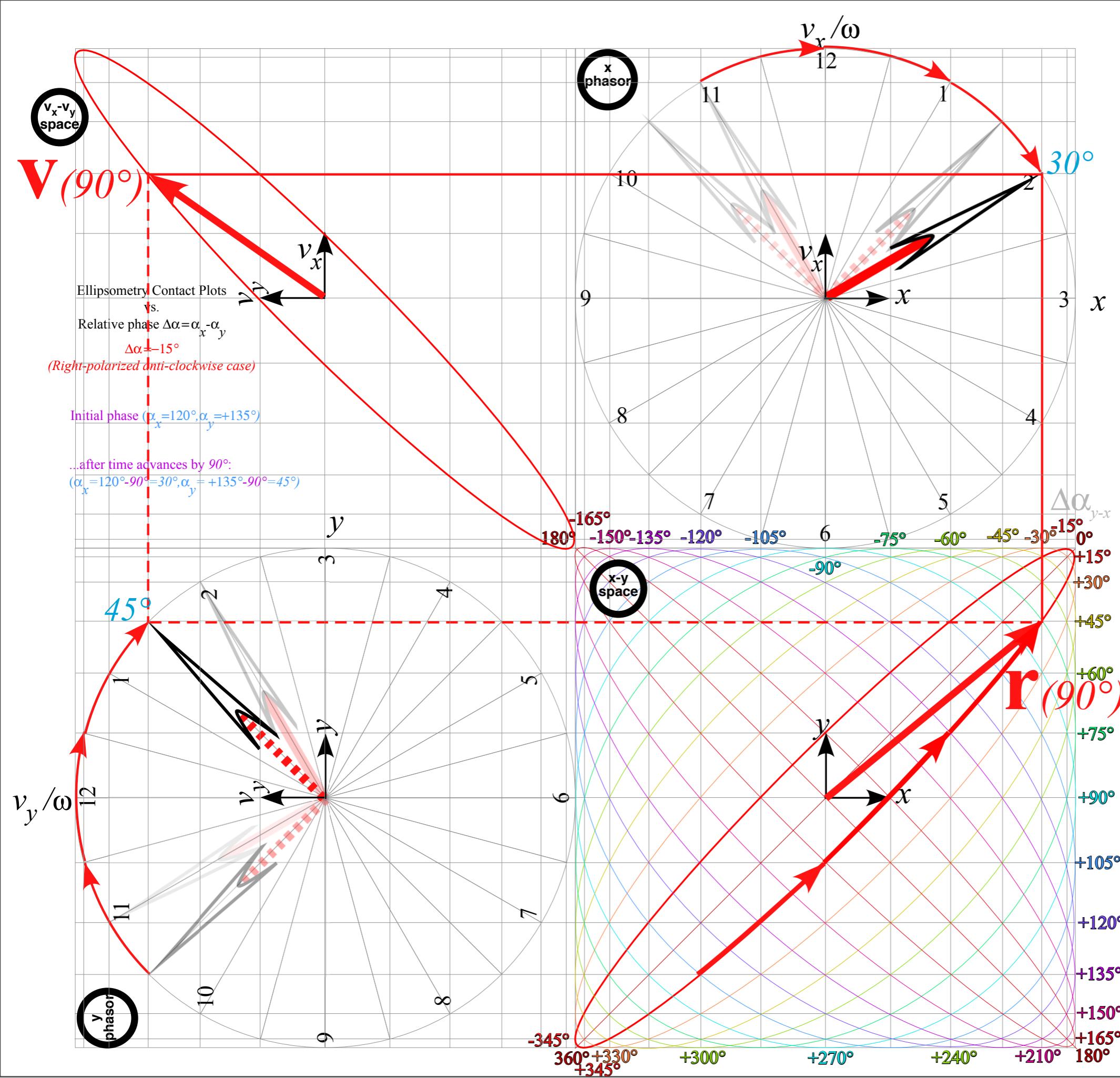








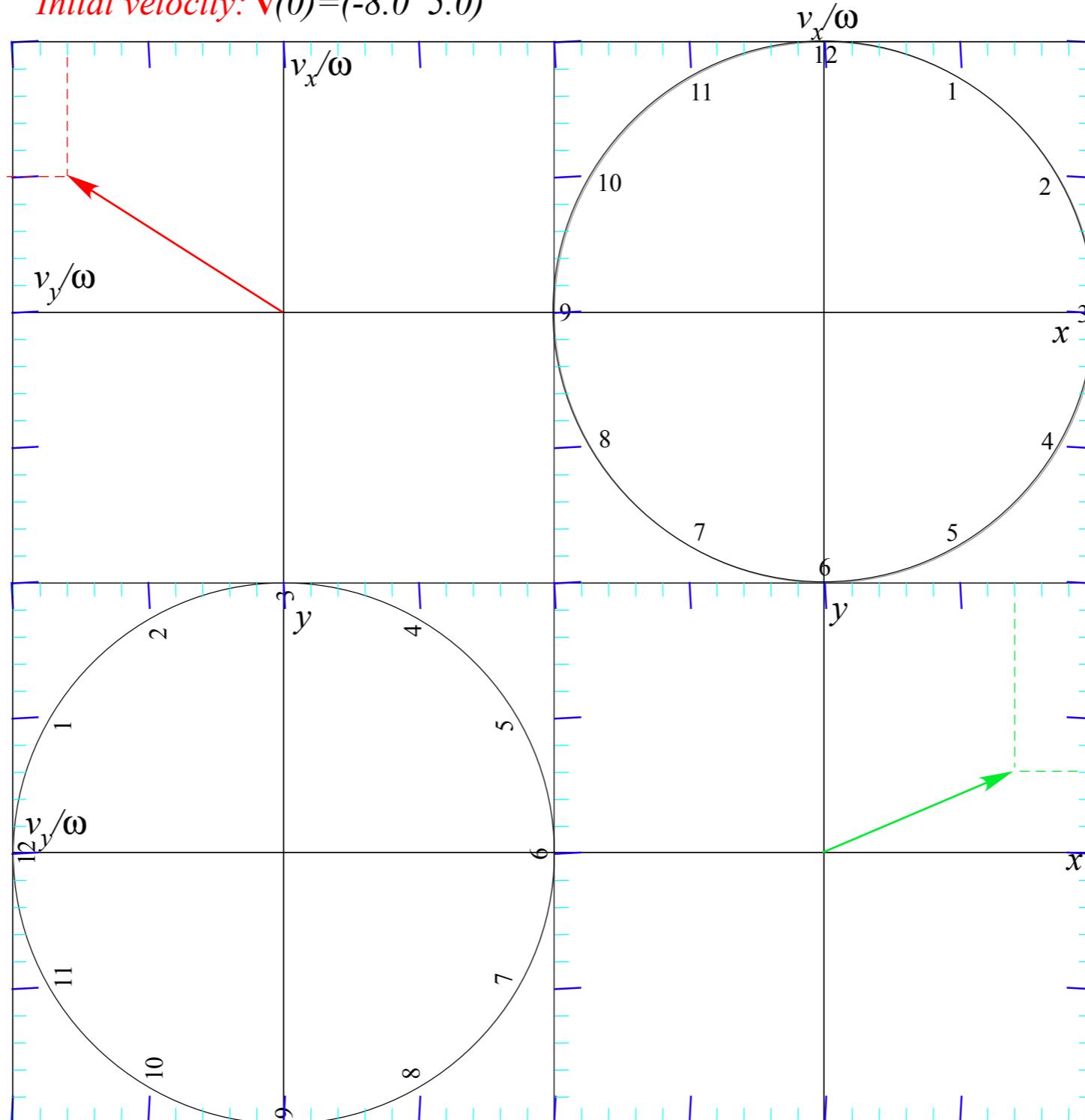




Constructing 2D IHO orbits by phasor plots

→ *Integrating IHO equations by phasor geometry (case of unequal x and y phasor area)*

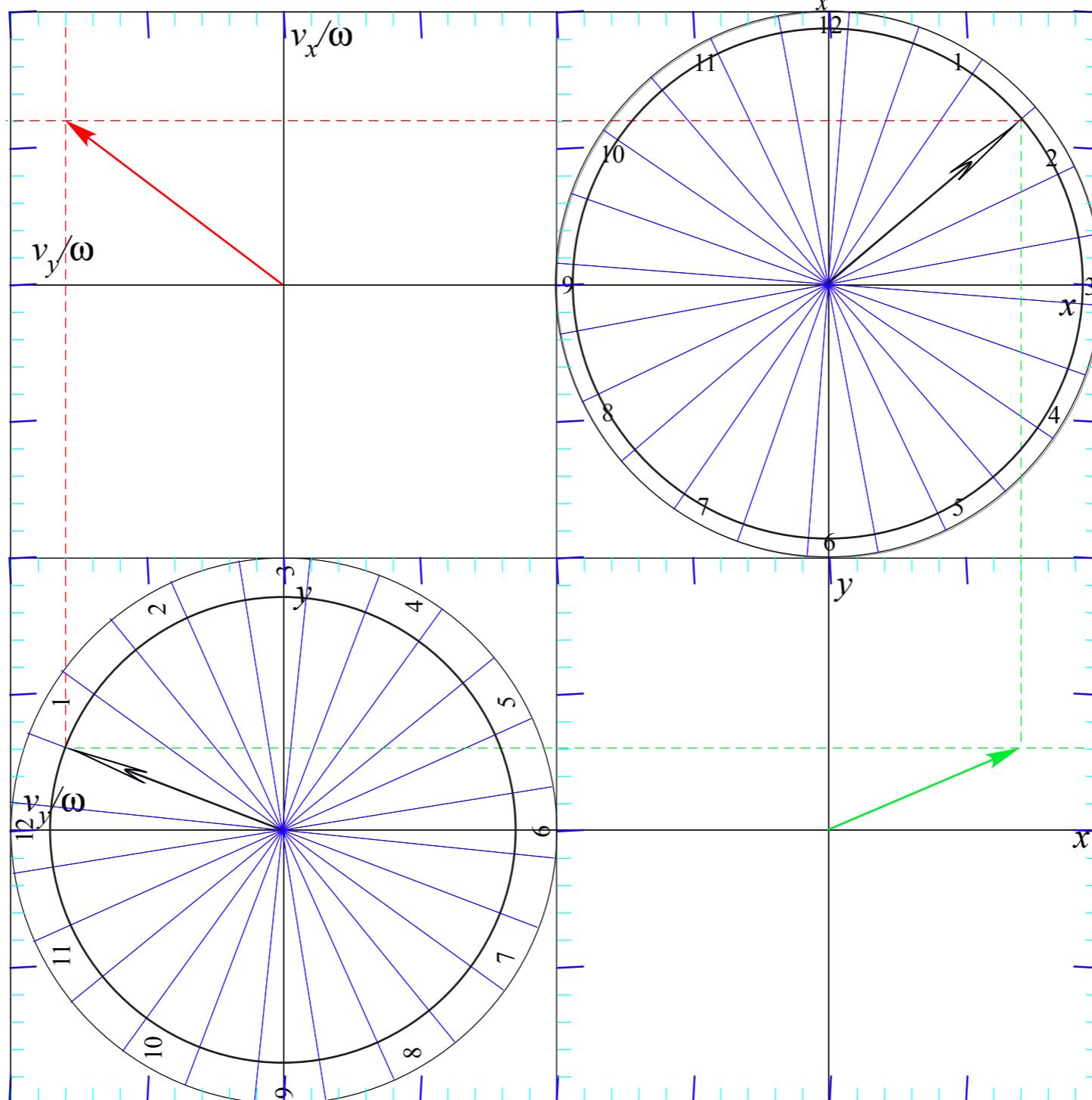
Initial velocity: $\mathbf{v}(0)=(-8.0 \ 5.0)$



Initial position: $\mathbf{r}(0)=(7.0 \ 3.0)$

[BoxIt simulation of U\(2\) orbits](#)
<http://www.uark.edu/ua/modphys/markup/BoxItWeb.html>

Initial velocity: $\mathbf{v}(0) = (-8.0 \ 6.0)$



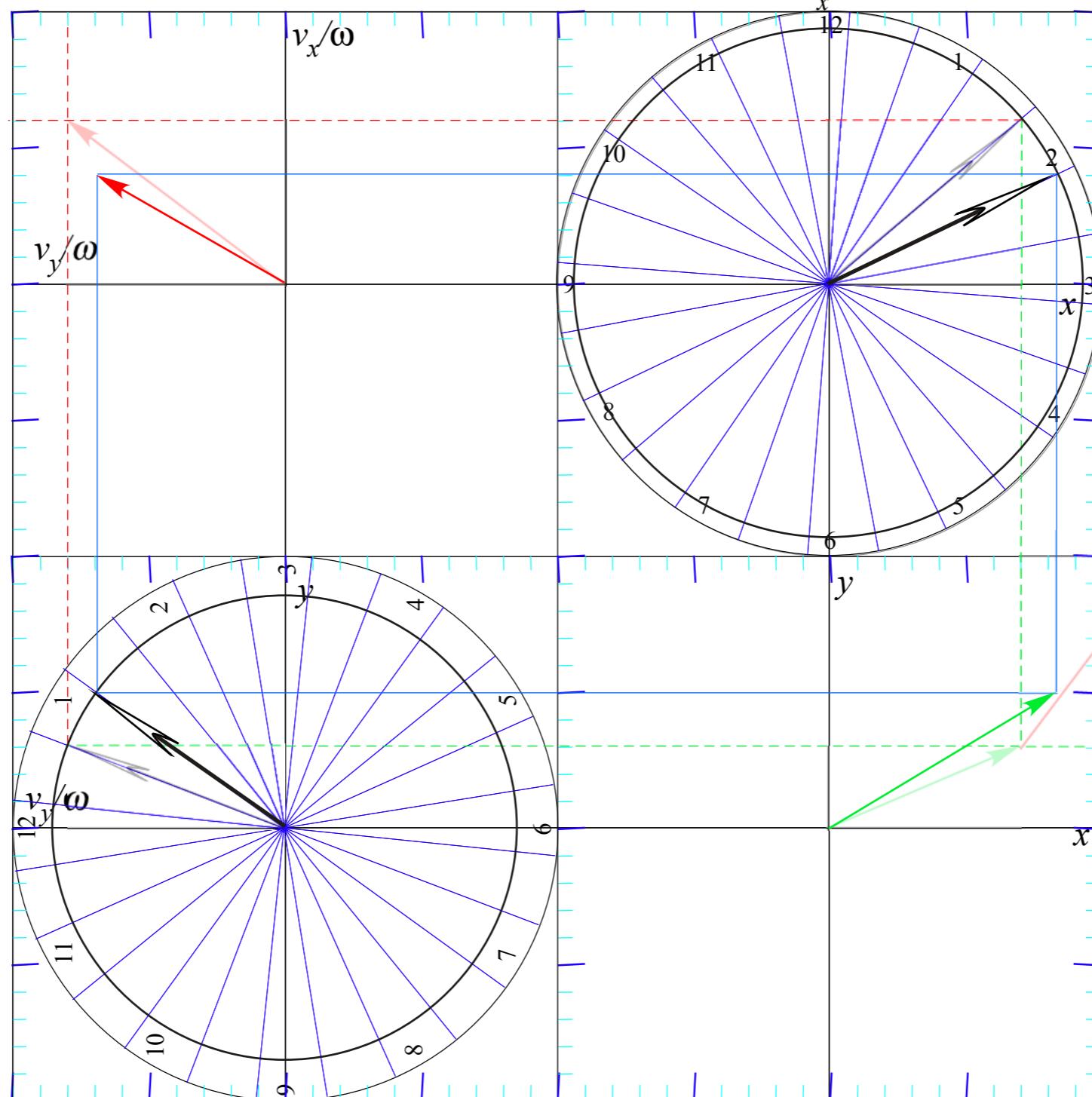
Initial position: $\mathbf{r}(0) = (7.0 \ 3.0)$

*Arbitrary initial position
 $\mathbf{r}(0) = (x(0), y(0))$*

*and initial velocity
 $\mathbf{v}(0) = (v_x(0), v_y(0))$*

*Usually have x and y
phasor circles of unequal size*

Initial velocity: $\mathbf{v}(0) = (-8.0 \ 6.0)$



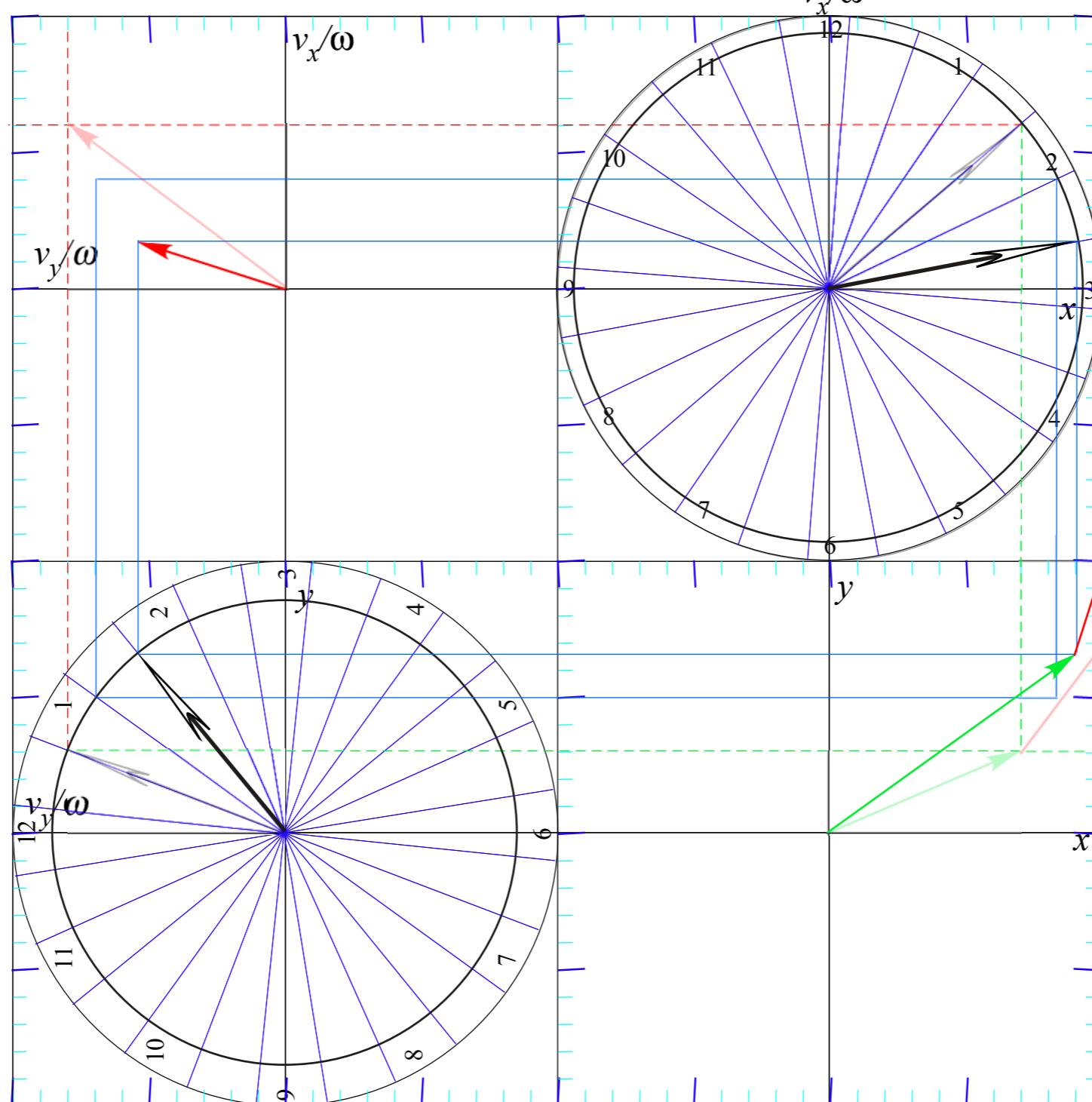
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Initial velocity: $\mathbf{v}(0) = (-8.0 \ 6.0)$



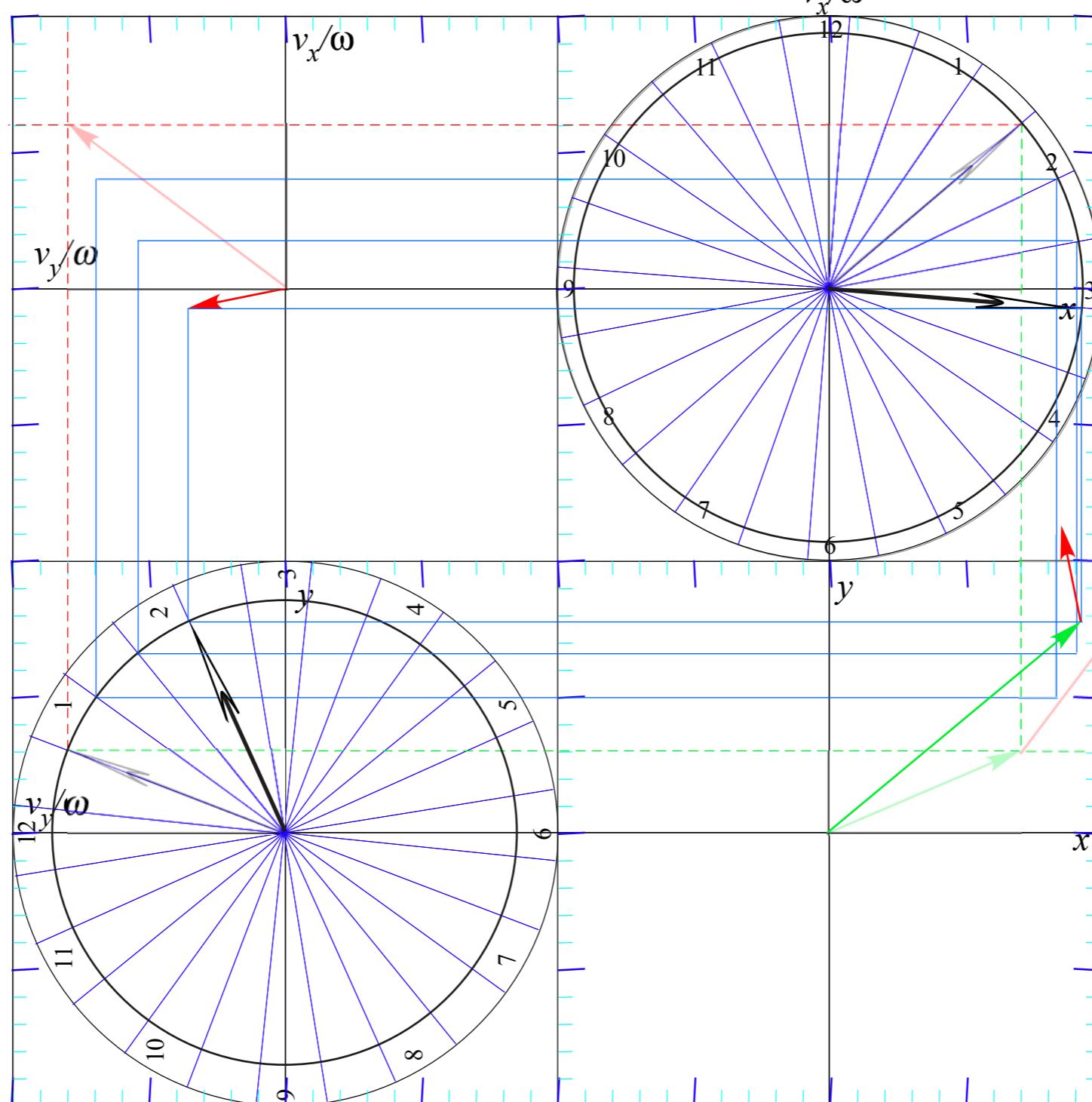
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*Usually have *x* and *y*
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Initial velocity: $\mathbf{v}(0) = (-8.0 \ 6.0)$



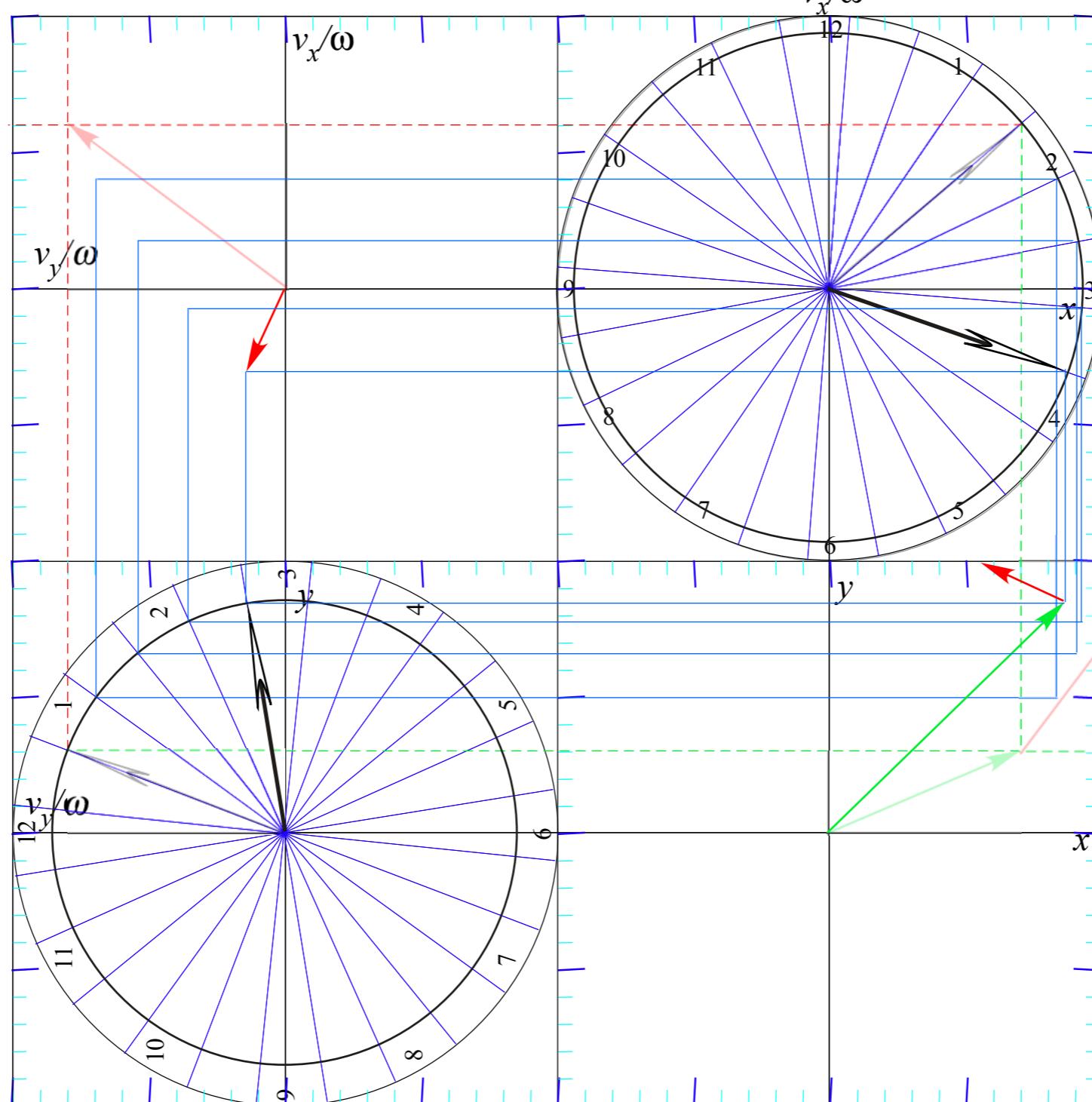
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phasor circles of unequal size*

Initial velocity: $\mathbf{v}(0) = (-8.0 \ 6.0)$



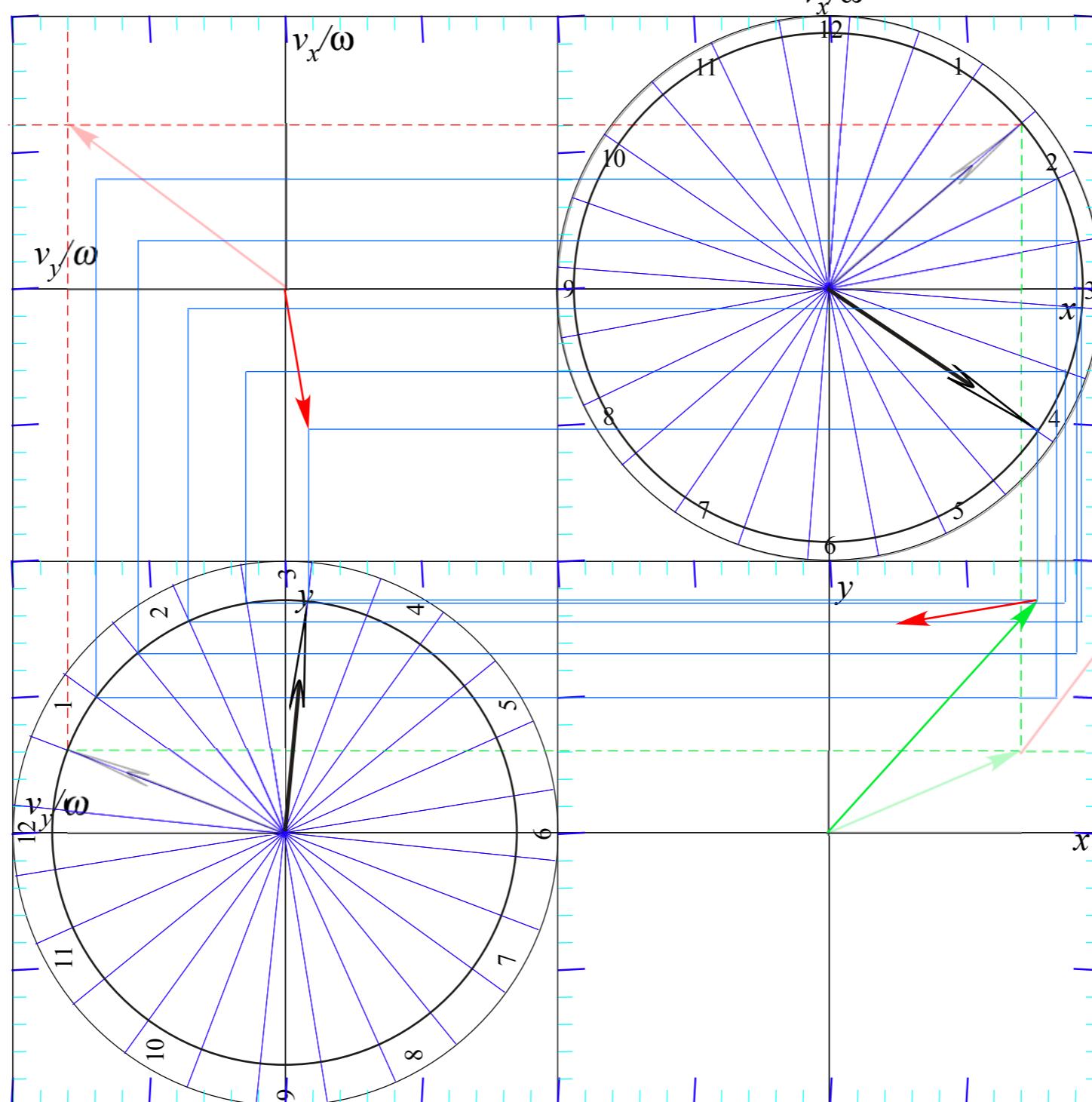
Initial position: $\mathbf{r}(0) = (7.0 \ 3.0)$

*Arbitrary initial position
 $\mathbf{r}(0) = (x(0), y(0))$*

*and initial velocity
 $\mathbf{v}(0) = (v_x(0), v_y(0))$*

*Usually have x and y
phasor circles of unequal size*

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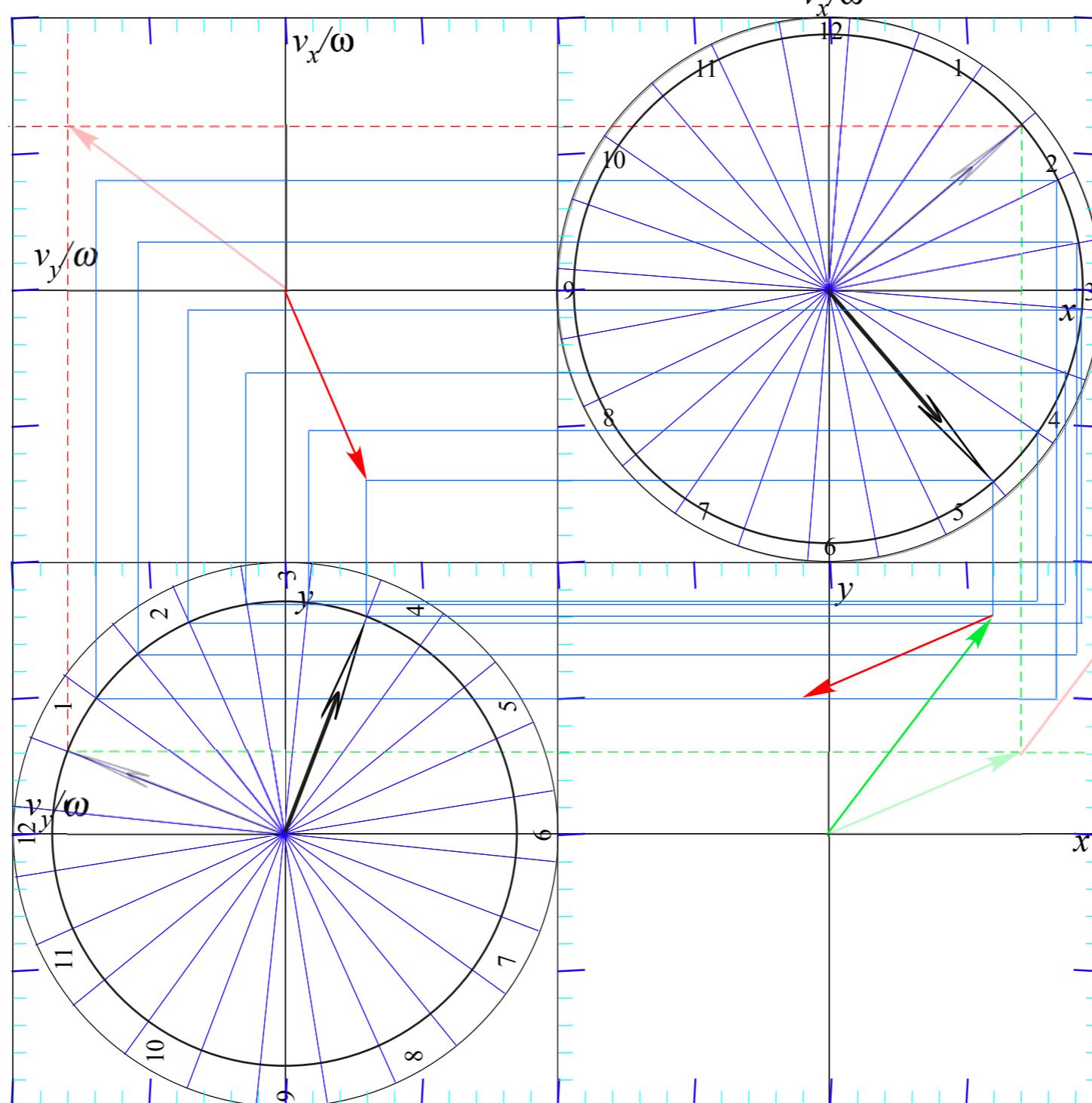
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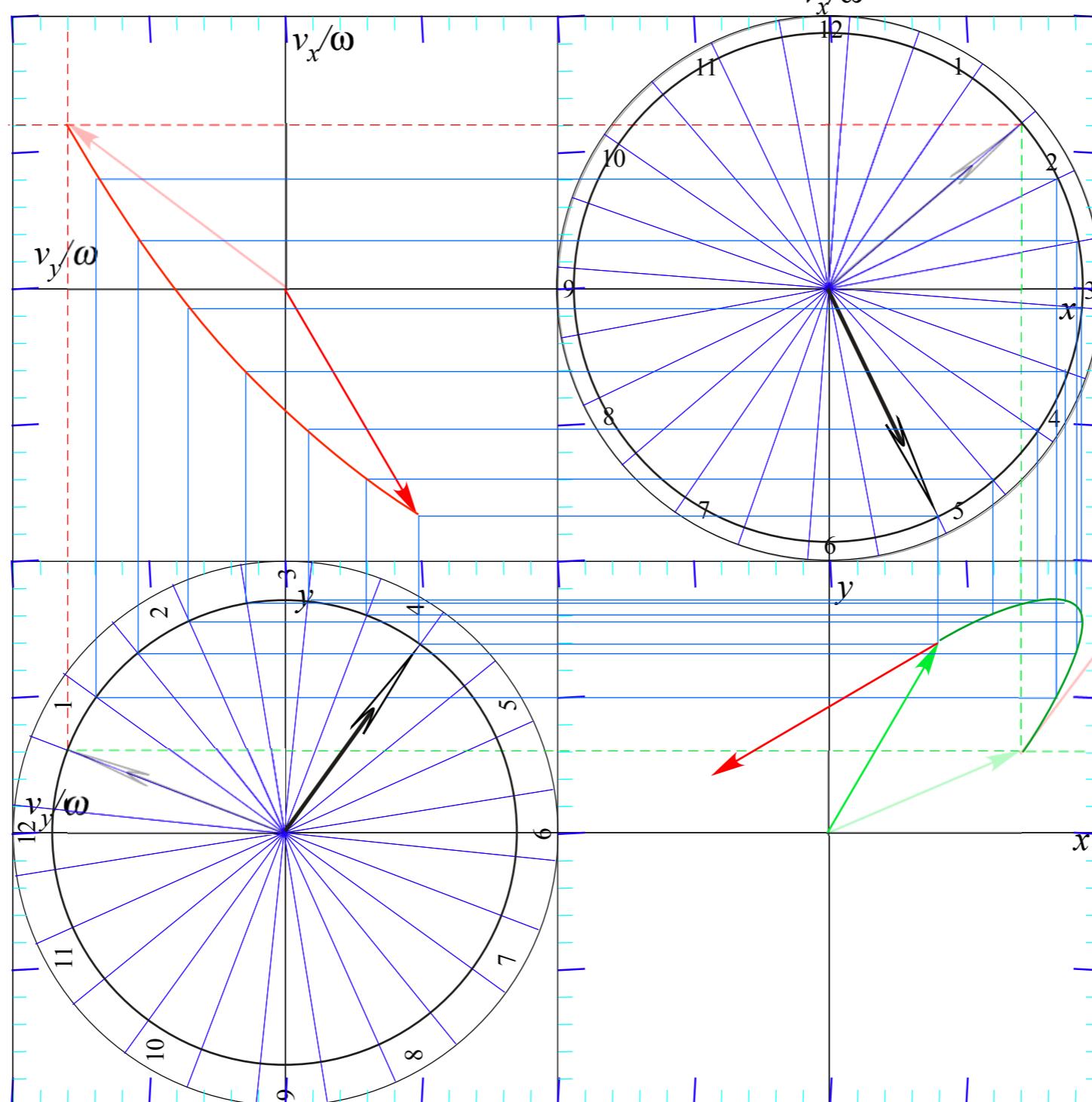
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