

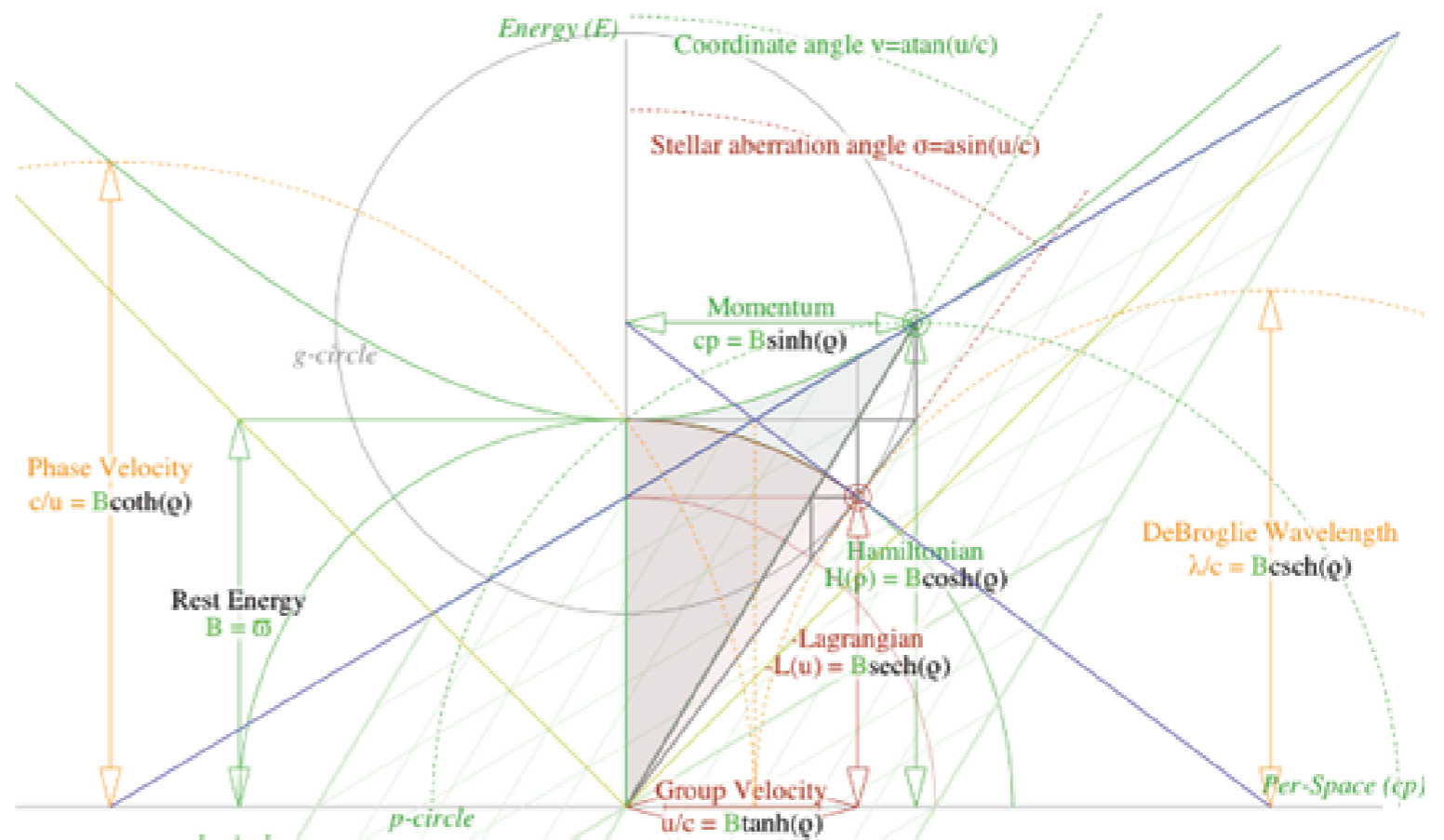
”SIMPLEST MOLECULE” CLARIFIES MODERN PHYSICS II. RELATIVISTIC QUANTUM MECHANICS

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A “simplest molecule” consisting of CW-laser beam pairs helps to clarify relativity in Talk I. In spite of a seemingly massless evanescence, an optical pair also clarifies classical and quantum mechanics of relativistic matter and anti-matter.

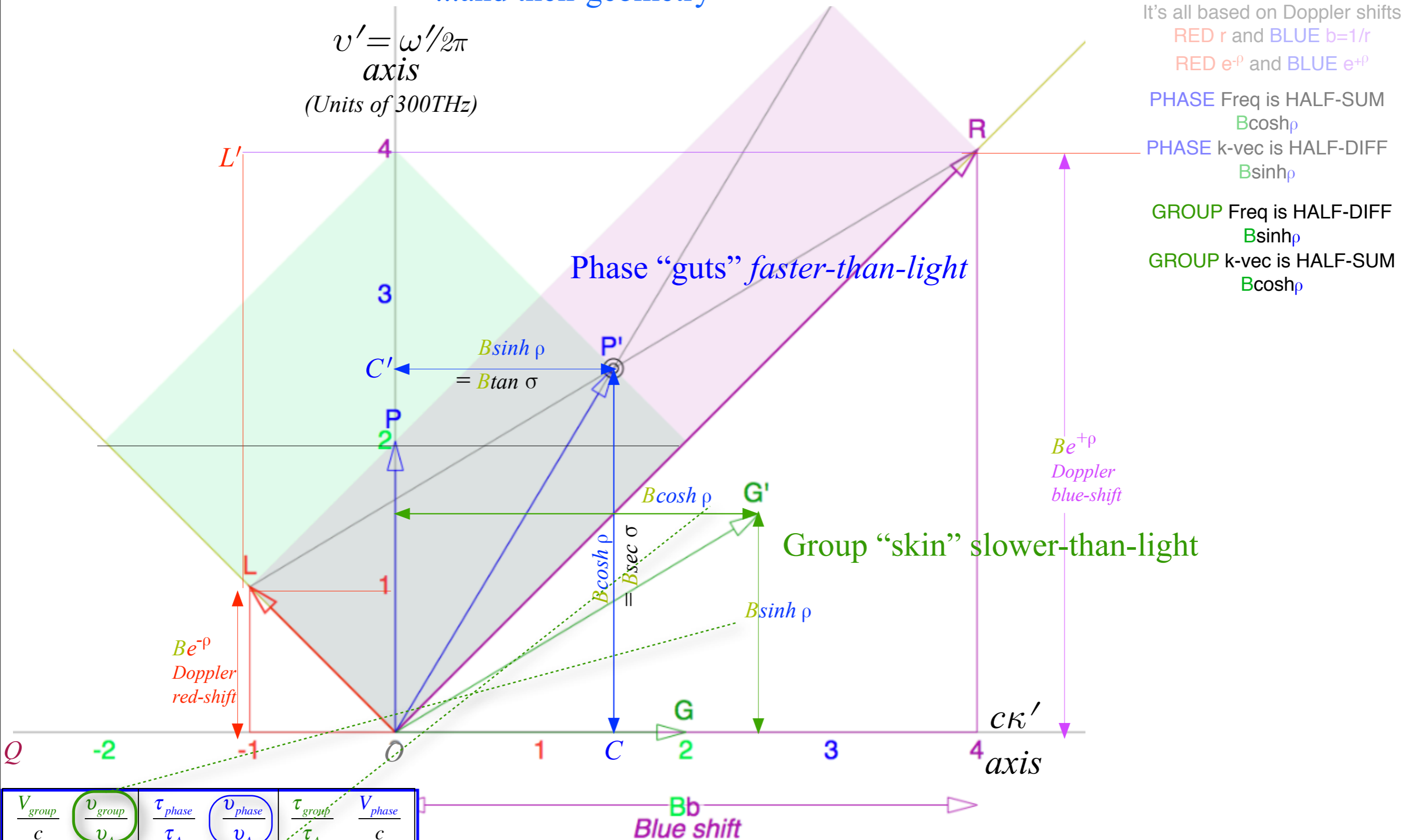
Logical extension of (x,ct) and (ω,ck) geometry gives relativistic action functions of Hamiltonian, Lagrangian, and Poincare that may be constructed in a few ruler-and-compass steps to relate relativistic parameters for group or phase velocity, momentum, energy, rapidity, stellar aberration, Doppler shifts, and DeBroglie wavelength. This exposes hyperbolic and circular trigonometry as two sides of one coin connected by Legendre contact transforms. One is Hamiltonian-like

with a longitudinal rapidity parameter ρ (log of Doppler shift). The other is Lagrange-like with a transverse angle parameter σ (stellar aberration). Optical geometry gives recoil in absorption, emission, and resonant Raman-Compton acceleration and distinguishes Einstein rest mass, Galilean momentum mass, and Newtonian effective mass. (Molecular photons appear less bullet-like and more rocket-like.) In conclusion, modern space-time physics appears as a simple result of the more self-evident Evenson’s axiom: “All colors go c.”



Review of optical wave parameters for relativity

...and their geometry



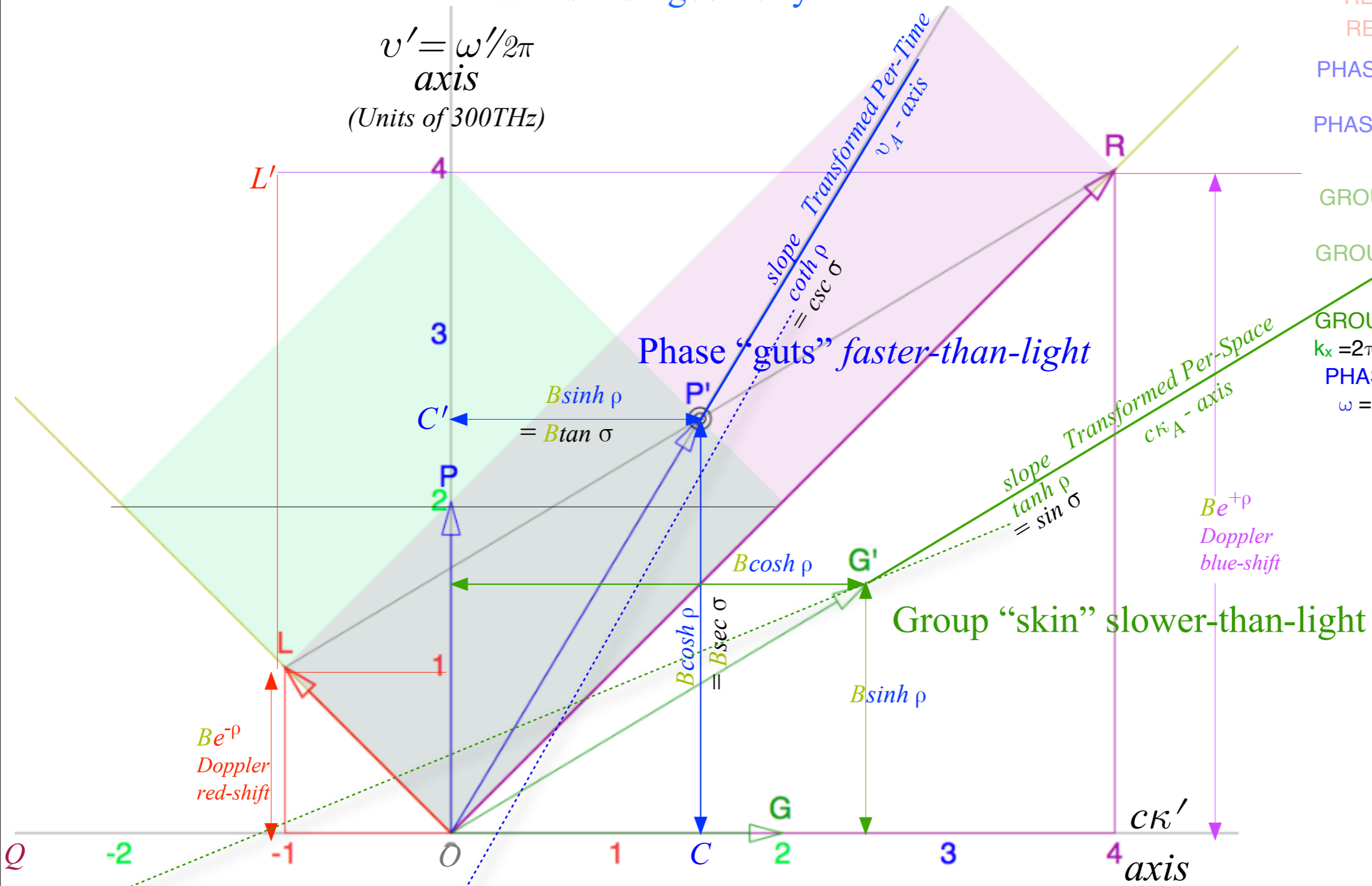
$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$
$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$
$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$
$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$

Phase "guts" of wave key to quantum dynamics (Difficult to observe)
 Group "skin" of wave key to classical mechanics (It's what we see.)

Review of optical wave parameters for relativity

...and their geometry

It's all based on Doppler shifts
 RED r and BLUE $b=1/r$
 RED $e^{-\rho}$ and BLUE $e^{+\rho}$
 PHASE Freq is HALF-SUM
 $B \cosh \rho$
 PHASE k-vec is HALF-DIFF
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 GROUP Freq is HALF-DIFF
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 GROUP is per-Space axis or
 $k_x = 2\pi \kappa_x = \text{Kappa}_x$ dimension
 PHASE is per-Time axis or
 $\omega = 2\pi \nu = \text{Nu}$ dimension



$\frac{V_{group}}{c}$	$\frac{\nu_{group}}{\nu_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{\nu_{phase}}{\nu_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$
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Phase “guts” of wave key to quantum dynamics (Difficult to observe)
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Bb
Blue shift

Review of optical wave parameters for relativity

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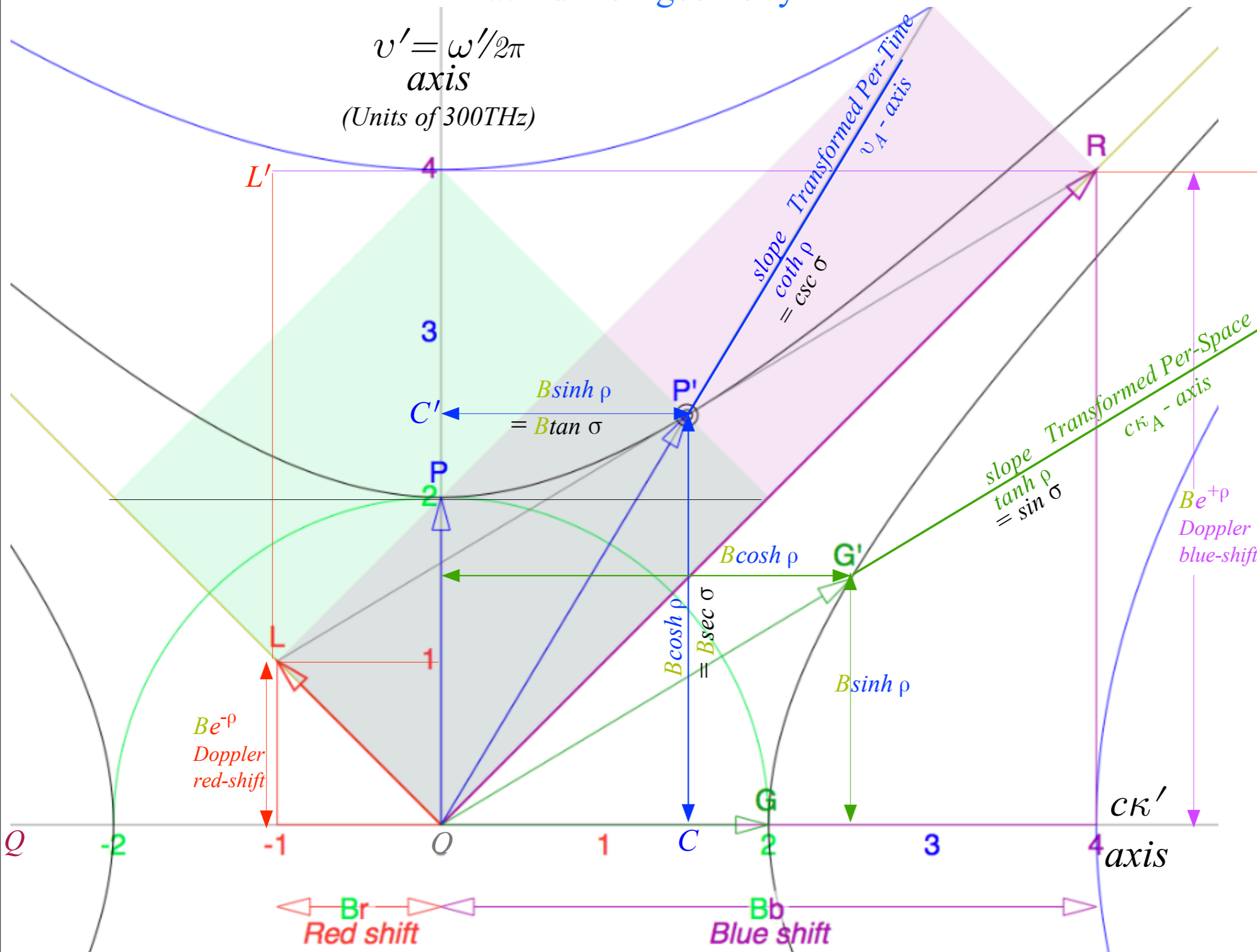
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PHASE and GROUP

hyperbolas are ρ -invariant

due to T-symmetry $b=1/r$



Review of optical wave parameters for relativity

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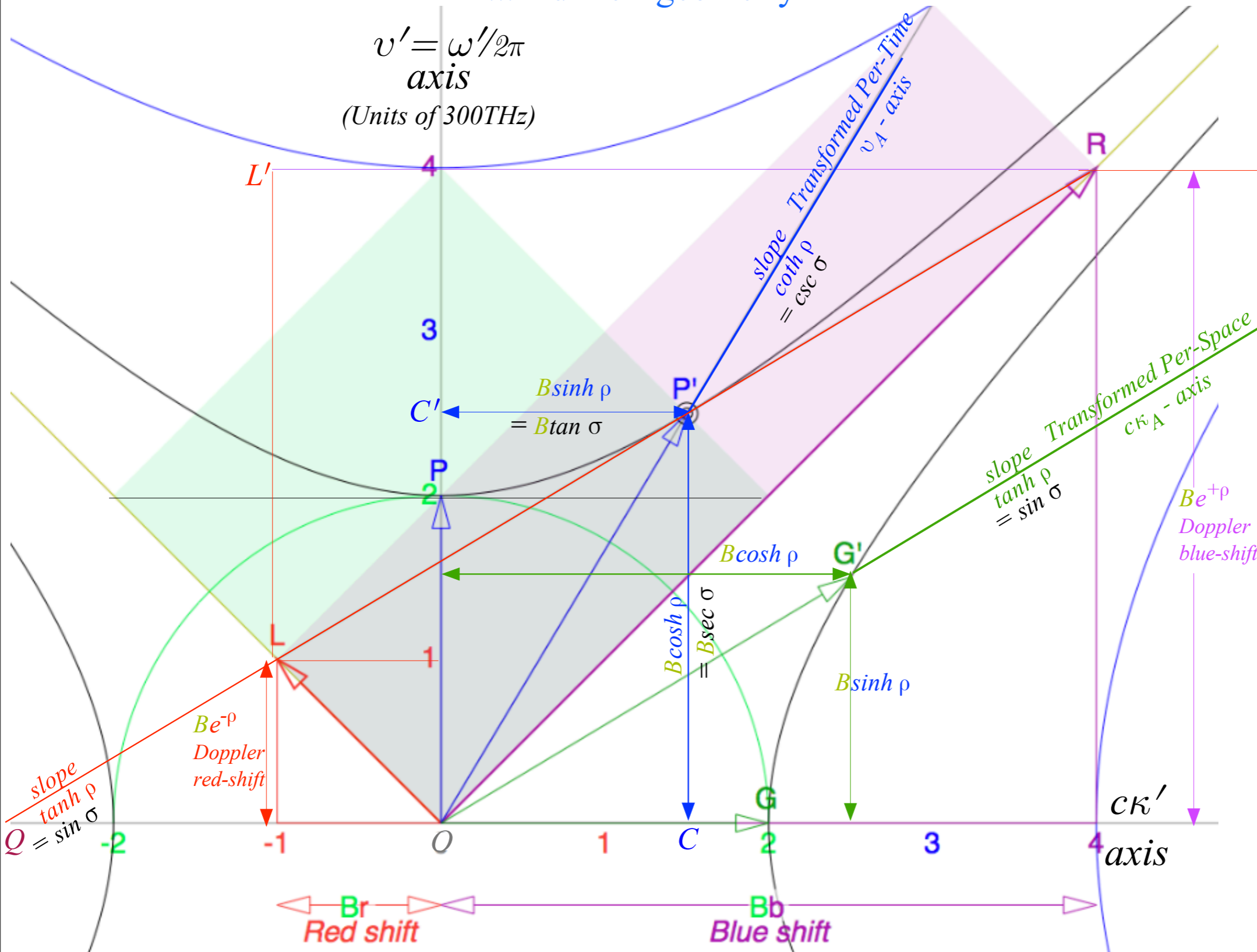
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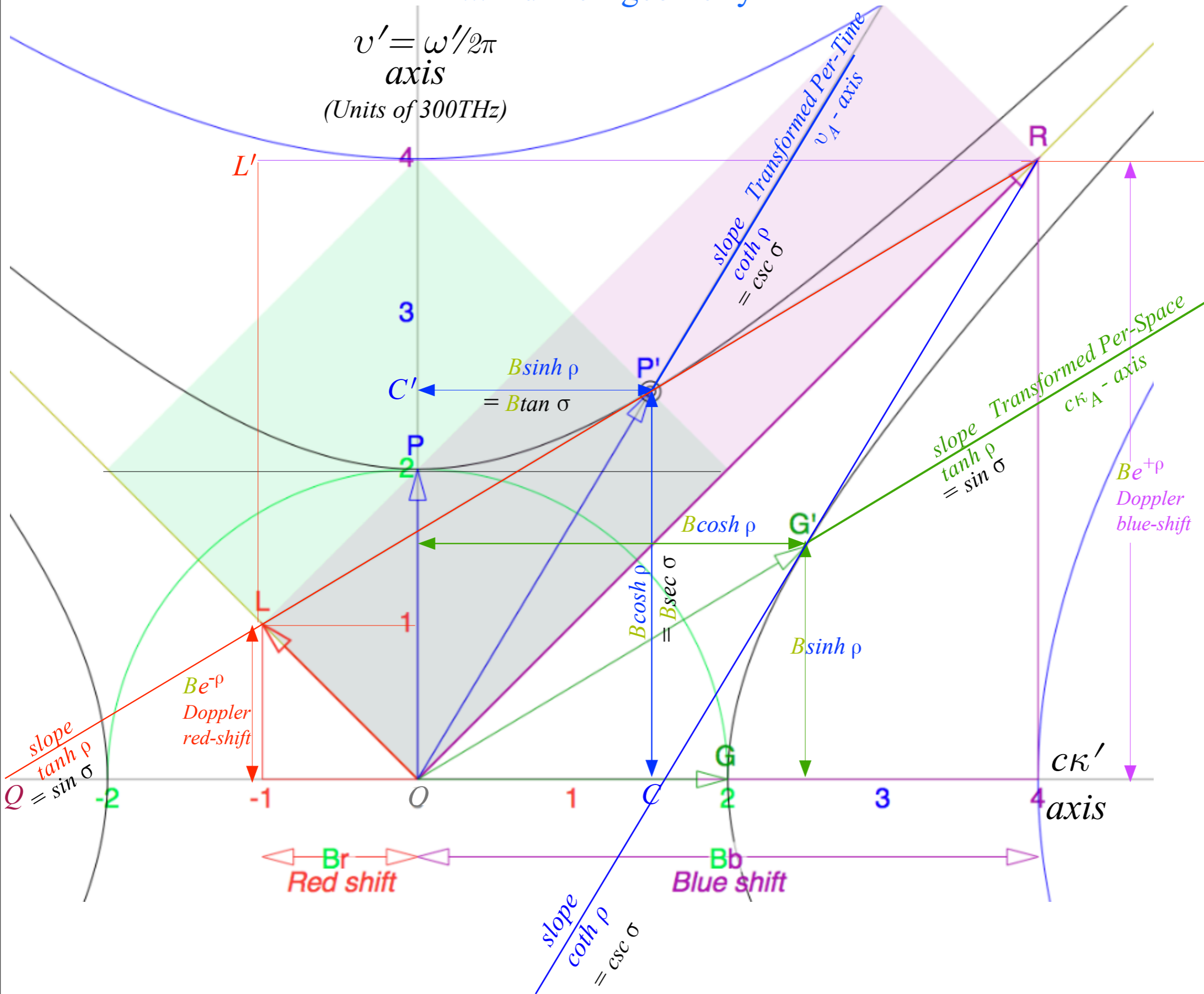
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Review of optical wave parameters for relativity

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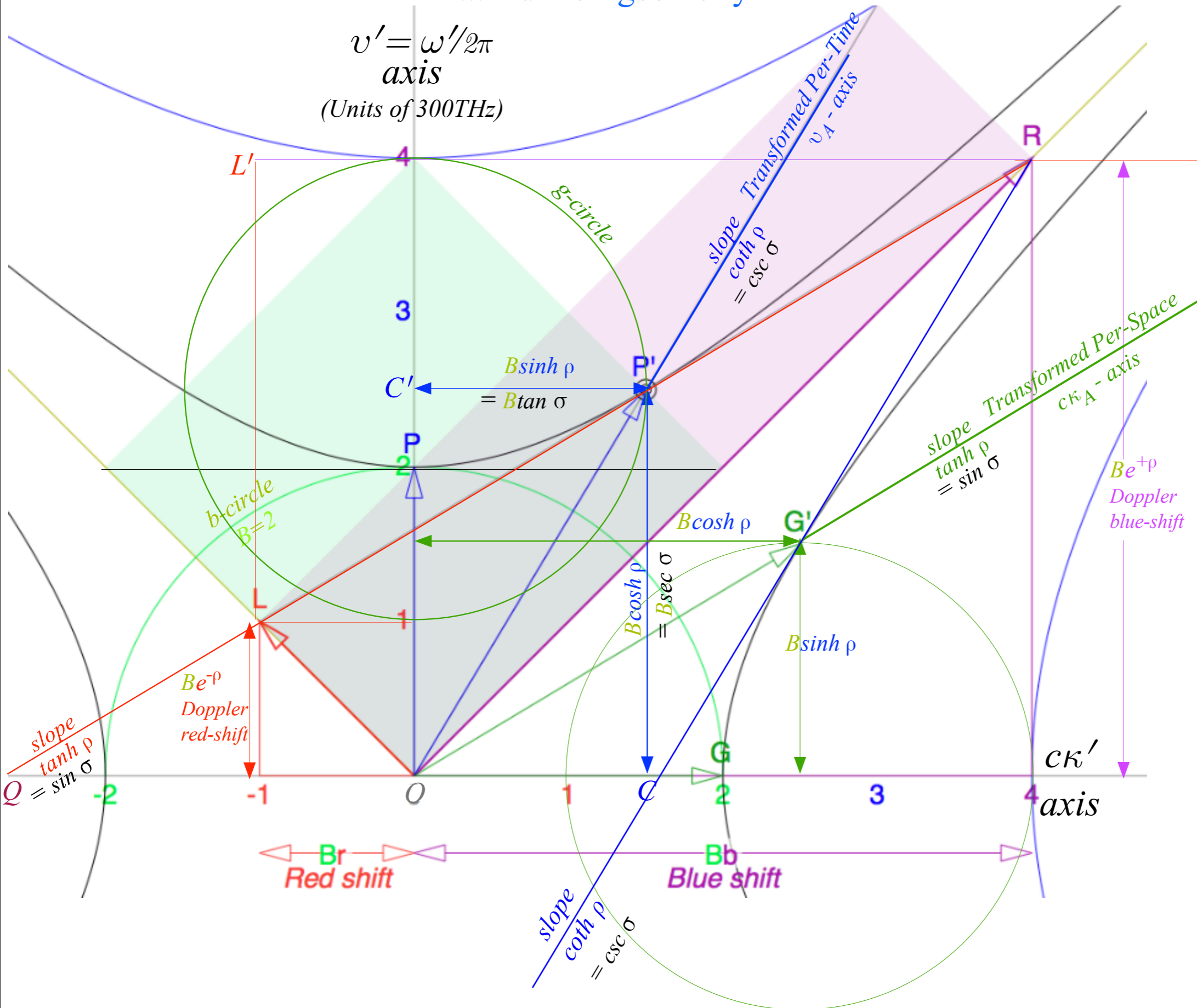
...and their geometry



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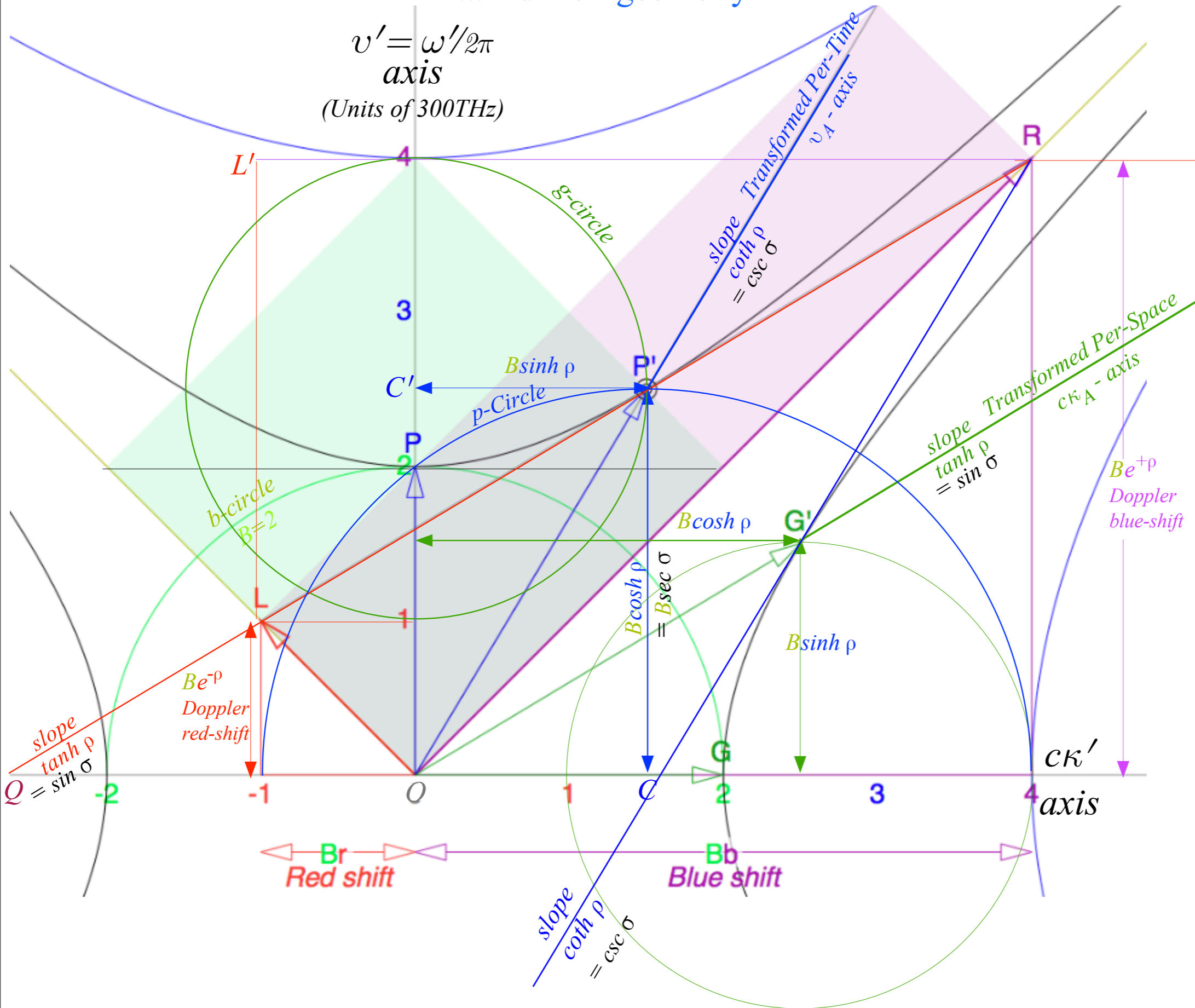
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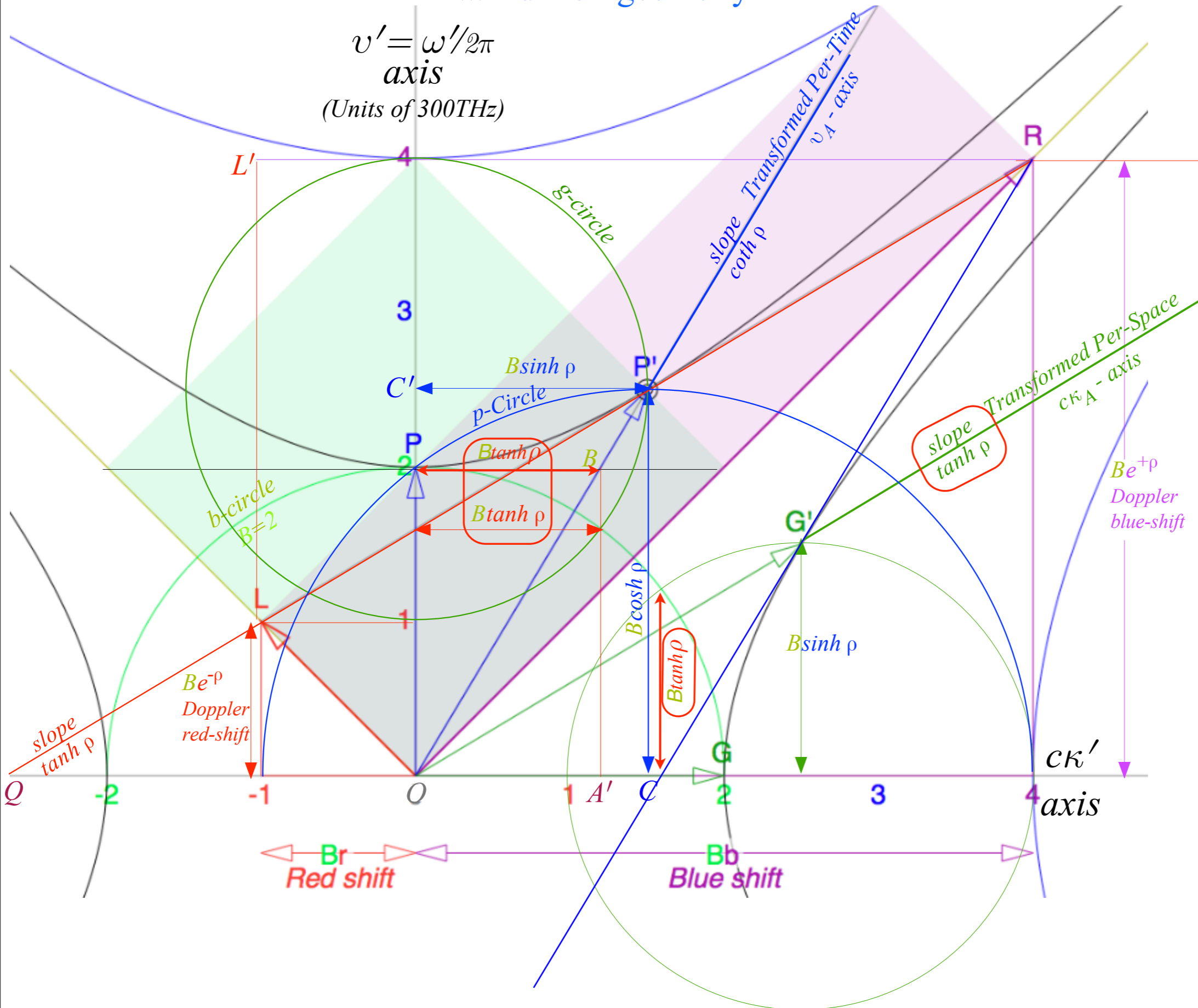
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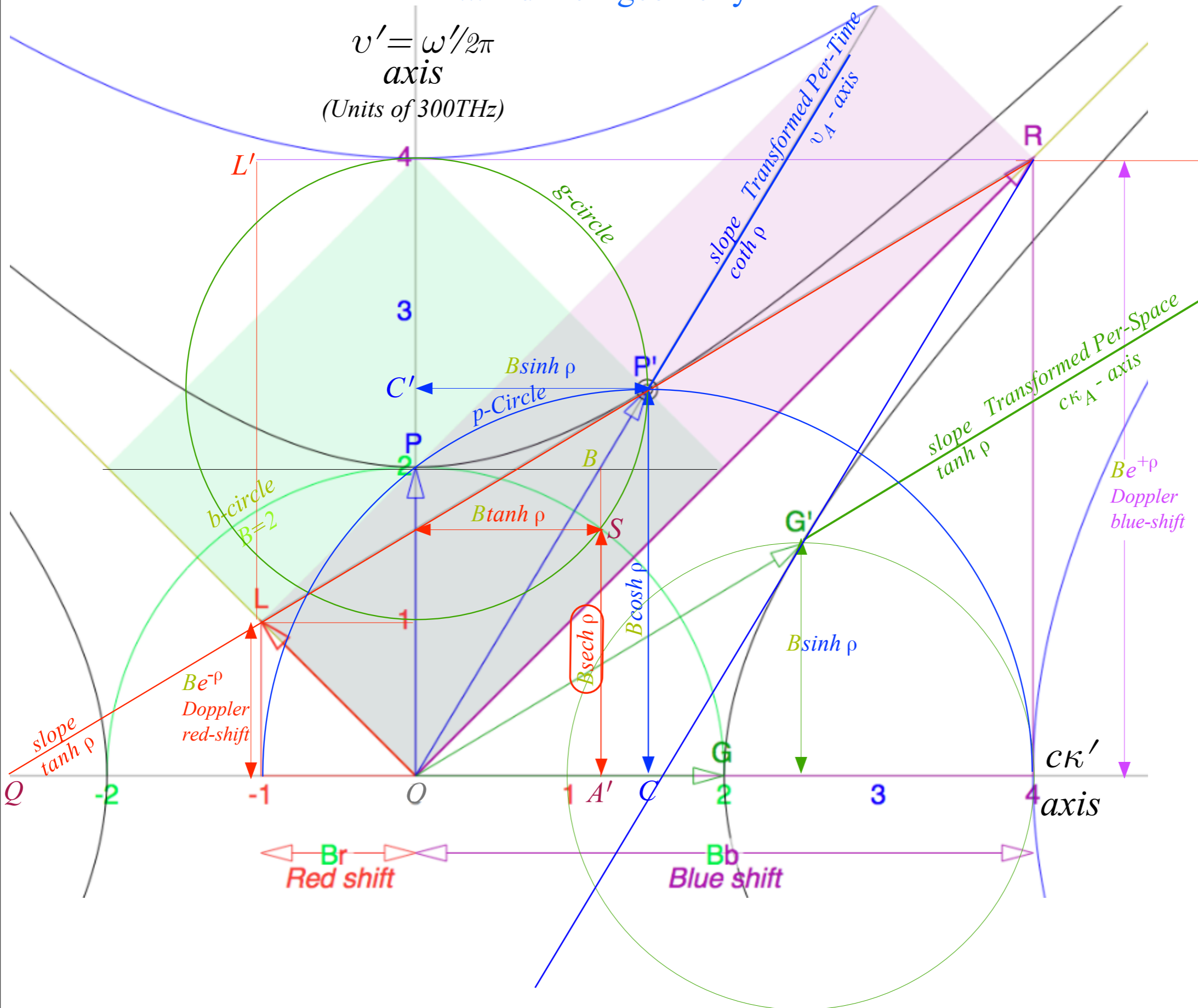
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 sets space-time axis slope
 $V_{\text{group}}/c = \tanh \rho$

Review of optical wave parameters for relativity

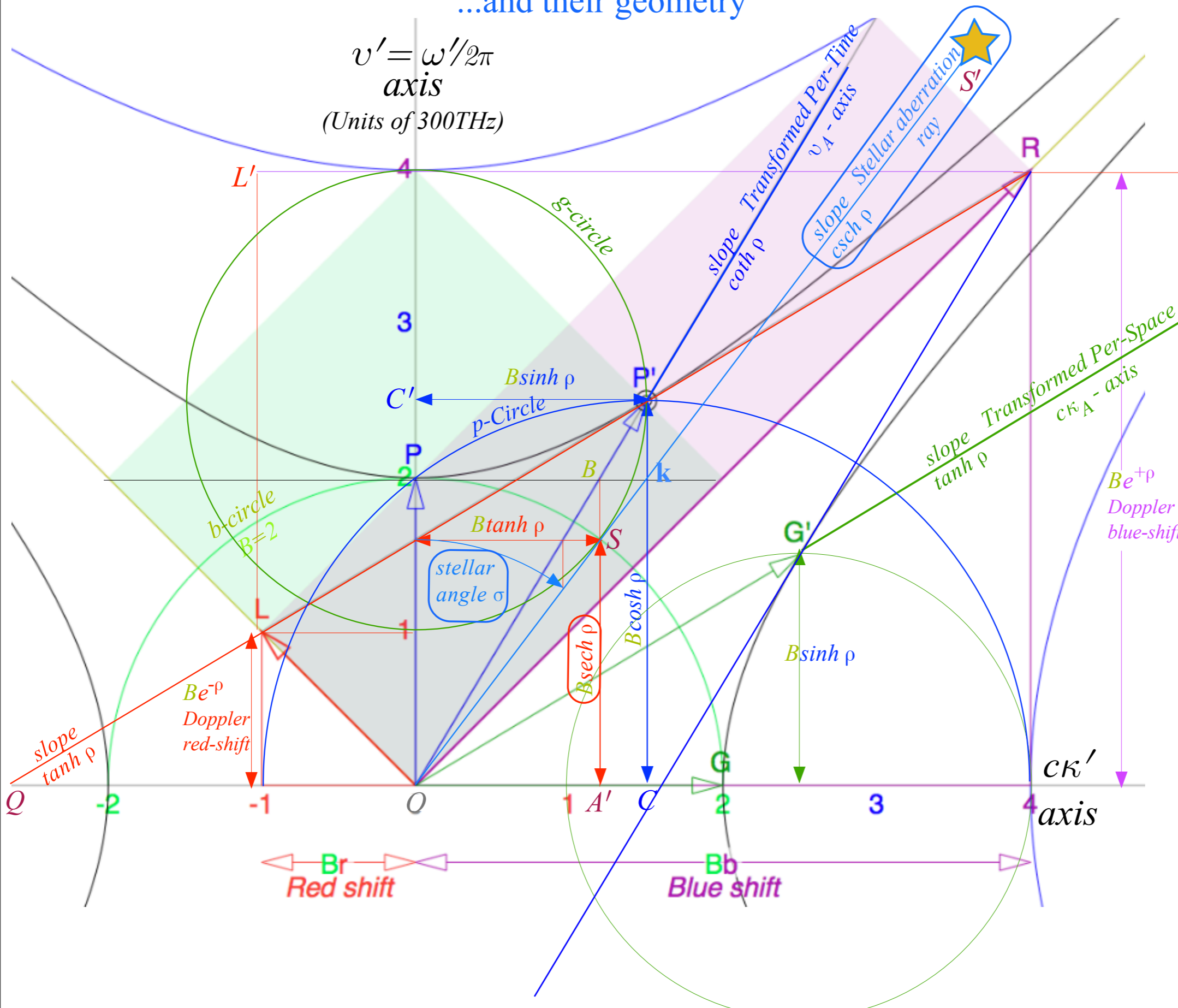
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- hyper-cotangent $B \coth \rho$ sets space-time axis slope $V_{\text{phase}}/c = \coth \rho$
- hyper-secant $B \operatorname{sech} \rho$ is compliment coord to hyper-tangent $B \tanh \rho$

Review of optical wave parameters for relativity

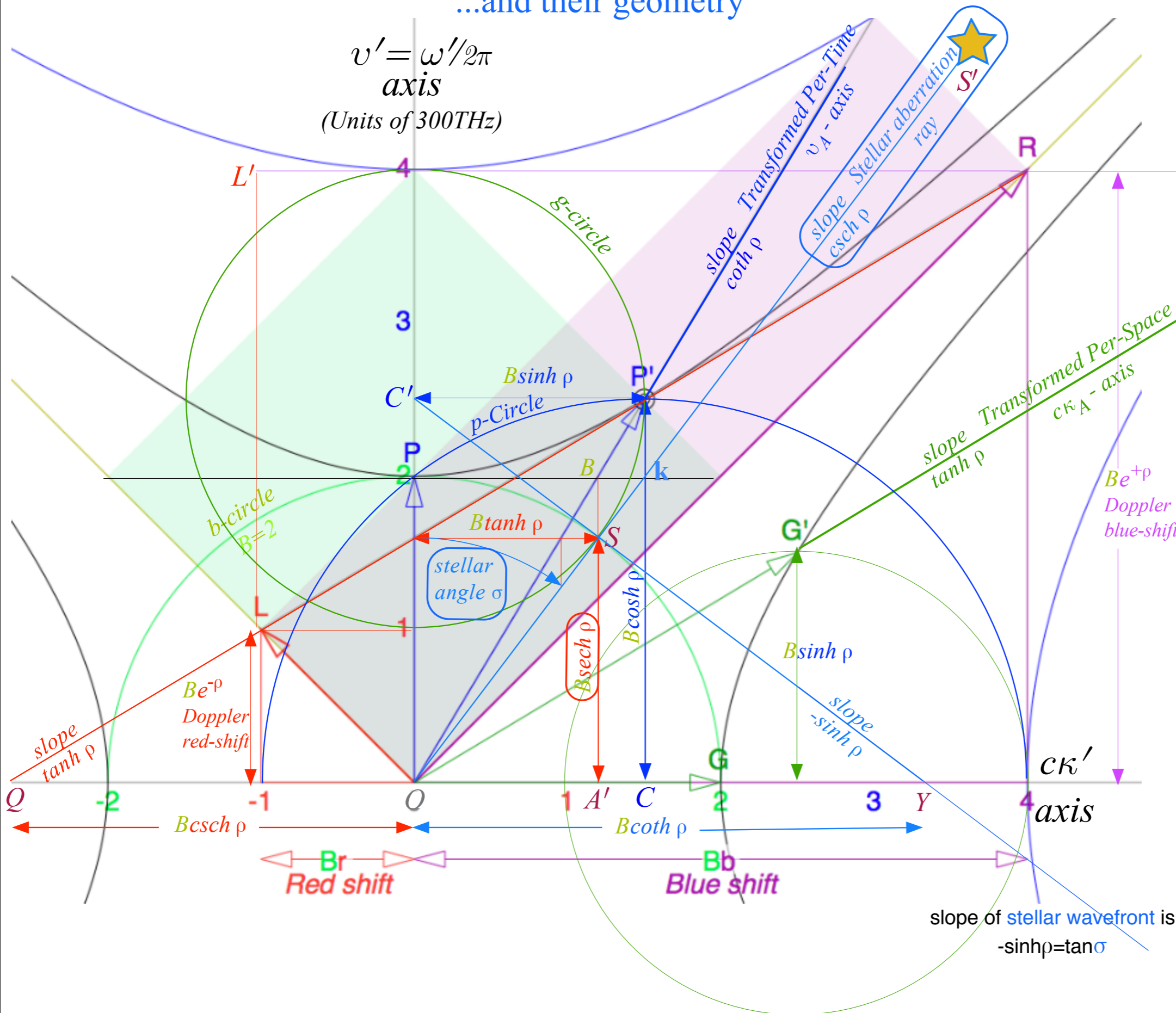
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- hyper-secant $B \operatorname{sech} \rho$ is compliment coord to hyper-tangent $B \tanh \rho$
- hyper-secant $B \operatorname{sech} \rho = B \cos \sigma$ is compliment coord to hyper-tangent $B \tanh \rho = B \sin \sigma$ for stellar aberration angle σ
- slope of stellar k-vector is $\operatorname{csch} \rho = \cot \sigma$

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hyper-secant $B \text{sech} \rho = B \cos \sigma$ is compliment coord to hyper-tangent $B \tanh \rho = B \sin \sigma$ for stellar aberration angle σ

slope of stellar wavefront is $-\sinh \rho = \tan \sigma$

slope of stellar k-vector is $\text{csch} \rho = \cot \sigma$

$B e^{+\rho}$ Doppler blue-shift

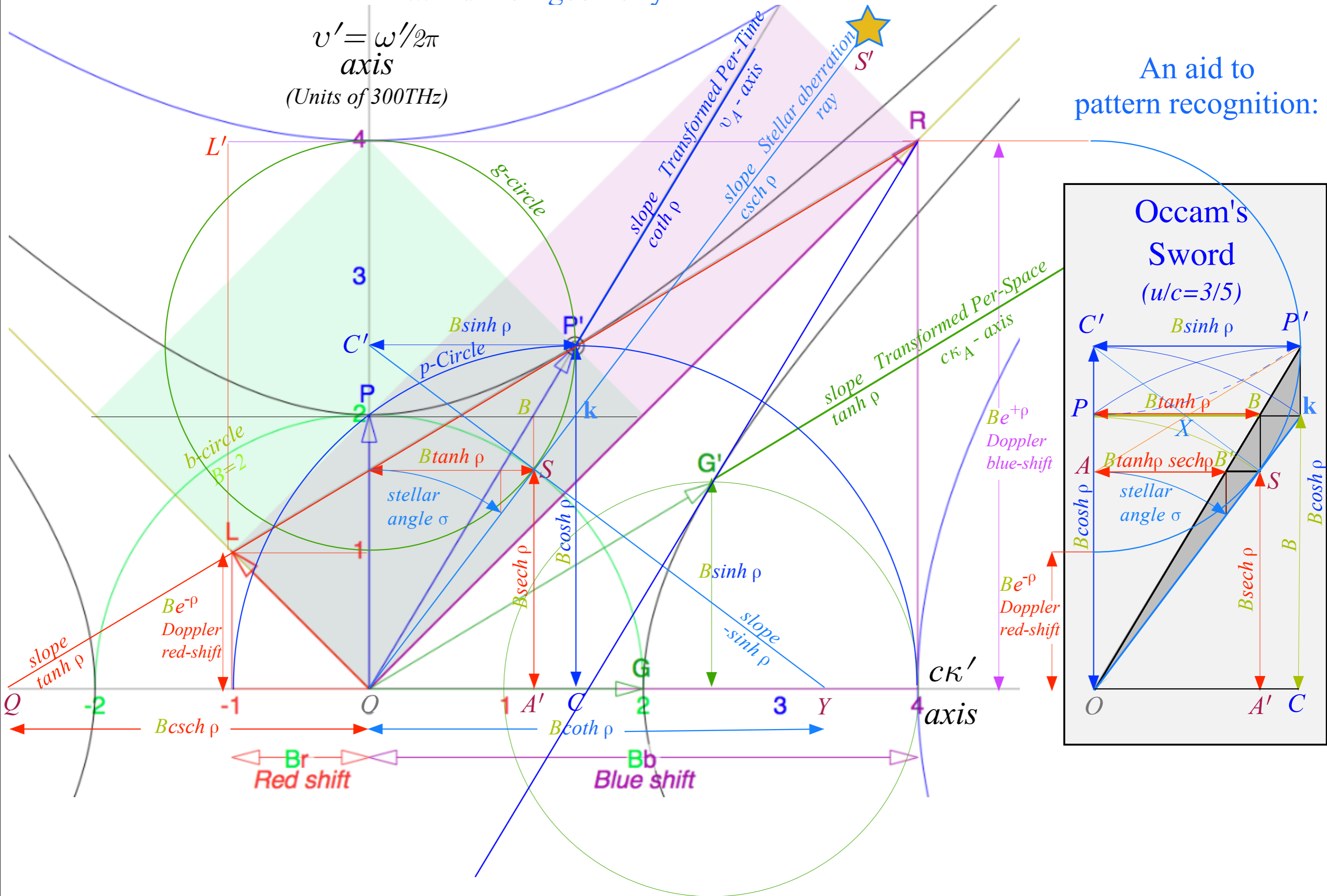
$B e^{-\rho}$ Doppler red-shift

$B r$ Red shift

$B b$ Blue shift

Review of optical wave parameters for relativity

...and their geometry



Using (some) wave parameters for relativistic quantum theory

$$v_{\text{phase}} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \text{ (for } u \ll c \text{)}$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2$$

$$B = v_A$$

At low speeds: ...

<i>time</i>	$b_{\text{RED}}^{\text{Doppler}}$	$\frac{V_{\text{group}}}{c}$	$\frac{v_{\text{group}}}{v_A}$	$\frac{\tau_{\text{phase}}}{\tau_A}$	$\frac{v_{\text{phase}}}{v_A}$	$\frac{\tau_{\text{group}}}{\tau_A}$	$\frac{V_{\text{phase}}}{c}$	$b_{\text{BLUE}}^{\text{Doppler}}$
<i>space</i>	$\frac{1}{b_{\text{BLUE}}^{\text{Doppler}}}$	$\frac{c}{V_{\text{phase}}}$	$\frac{\kappa_{\text{phase}}}{\kappa_A}$	$\frac{\lambda_{\text{group}}}{\lambda_A}$	$\frac{\kappa_{\text{group}}}{\kappa_A}$	$\frac{\lambda_{\text{phase}}}{\lambda_A}$	$\frac{c}{V_{\text{group}}}$	$\frac{1}{b_{\text{RED}}^{\text{Doppler}}}$
<i>rapidity</i> ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\text{coth } \rho$	$e^{+\rho}$
<i>stellar</i> ∇ <i>angle</i> σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
<i>value for</i> $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

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$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \text{ (for } u \ll c \text{)}$$

$$cK_{phase} = B \sinh \rho \approx B \rho \text{ (for } u \ll c \text{)}$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2$$

$$\sinh \rho \approx \rho$$

$$B = v_A$$

$$B = v_A = cK_A$$

At low speeds:

time	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar ∇ angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\operatorname{csc} \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
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At low speeds:

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

$$\sinh \rho \approx \rho \approx \frac{u}{c}$$

$$B = v_A$$

$$B = v_A = cK_A$$

time	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
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At low speeds: \Leftarrow for $(u \ll c)$

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2$$

time	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
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$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2$$

At low speeds:

← for ($u \ll c$) ⇒

$$K_{phase} \approx \frac{B}{c^2} u$$

time	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

Using (some) wave parameters for relativistic quantum theory

$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \quad (\text{for } u \ll c)$$

$$c\kappa_{phase} = B \sinh \rho \approx B \rho \quad (\text{for } u \ll c)$$

$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

$$\sinh \rho \approx \rho \approx \frac{u}{c}$$

$$B = v_A$$

$$B = v_A = c\kappa_A$$

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2$$

At low speeds:

← for ($u \ll c$) ⇒

$$\kappa_{phase} \approx \frac{B}{c^2} u$$

Looks like $\frac{1}{2} Mu^2$

Looks like Mu

time	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

Using (some) wave parameters for relativistic quantum theory

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$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

$$\sinh \rho \approx \rho \approx \frac{u}{c}$$

$$B = v_A$$

$$B = v_A = cK_A$$

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2$$

At low speeds:
 \Leftarrow for $(u \ll c) \Rightarrow$

$$K_{phase} \approx \frac{B}{c^2} u$$

Rescale v_{phase} by h so: $M = \frac{hB}{c^2}$

Looks like $\frac{1}{2} Mu^2$

Looks like Mu

time	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\text{coth } \rho$	$e^{+\rho}$
stellar ∇ angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

Using (some) wave parameters for relativistic quantum theory

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$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

$$cK_{phase} = B \sinh \rho \approx B \rho \text{ (for } u \ll c \text{)}$$

$$\sinh \rho \approx \rho \approx \frac{u}{c}$$

$$B = v_A$$

$$B = v_A = cK_A$$

$$\frac{u}{c} = \tanh \rho \approx \rho \text{ (for } u \ll c \text{)}$$

At low speeds:

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2$$

for $(u \ll c) \Rightarrow$

$$K_{phase} \approx \frac{B}{c^2} u$$

Rescale v_{phase} by h so: $M = \frac{hB}{c^2}$

or: $hB = Mc^2$

Looks like $\frac{1}{2} Mu^2$

Looks like Mu

time	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
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At low speeds:

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Rescale v_{phase} by h so: $M = \frac{hB}{c^2}$ or: $hB = Mc^2$

$$hv_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2 \quad \Leftarrow \text{for } (u \ll c) \Rightarrow \quad hK_{phase} \approx \frac{hB}{c^2} u$$

time	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

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Rescale v_{phase} by h so: $M = \frac{hB}{c^2}$ or: $hB = Mc^2$

$$hv_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2 \quad \Leftarrow \text{for } (u \ll c) \Rightarrow \quad hK_{phase} \approx \frac{hB}{c^2} u$$

$$hv_{phase} \approx Mc^2 + \frac{1}{2} Mu^2 \quad \Leftarrow \text{for } (u \ll c) \Rightarrow \quad hK_{phase} \approx Mu$$

time	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

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Rescale v_{phase} by h so: $M = \frac{hB}{c^2}$ or: $hB = Mc^2$

$$hv_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2 \quad \Leftarrow \text{for } (u \ll c) \Rightarrow \quad hK_{phase} \approx \frac{hB}{c^2} u$$

$$hv_{phase} \approx Mc^2 + \frac{1}{2} Mu^2 \quad \Leftarrow \text{for } (u \ll c) \Rightarrow \quad hK_{phase} \approx Mu$$

Lucky coincidences??

time	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

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Rescale v_{phase} by h so: $M = \frac{hB}{c^2}$ or: $hB = Mc^2$

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...Try exact v_{phase} ...

time	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar ∇ angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

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Rescale v_{phase} by h so: $M = \frac{hB}{c^2}$ or: $hB = Mc^2$

$$hv_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2 \quad \Leftarrow \text{for } (u \ll c) \Rightarrow \quad hK_{phase} \approx \frac{hB}{c^2} u$$

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$$hv_{phase} = hB \cosh \rho = Mc^2 \cosh \rho$$

time	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
stellar angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

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$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

$$\sinh \rho \approx \rho \approx \frac{u}{c}$$

$$B = v_A$$

$$B = v_A = cK_A$$

At low speeds:

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2$$

\Leftarrow for $(u \ll c) \Rightarrow$

$$K_{phase} \approx \frac{B}{c^2} u$$

Rescale v_{phase} by h so: $M = \frac{hB}{c^2}$ or: $hB = Mc^2$

$$hv_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2$$

\Leftarrow for $(u \ll c) \Rightarrow$

$$hK_{phase} \approx \frac{hB}{c^2} u$$

$$hv_{phase} \approx Mc^2 + \frac{1}{2} Mu^2$$

\Leftarrow for $(u \ll c) \Rightarrow$

$$hK_{phase} \approx Mu$$

Lucky coincidences??

...Try exact v_{phase} ...

$$hv_{phase} = hB \cosh \rho = Mc^2 \cosh \rho$$

$$= \text{Total Energy: } E = \frac{Mc^2}{\sqrt{1-u^2/c^2}}$$

Einstein (1905)

time	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\text{coth } \rho$	$e^{+\rho}$
stellar ∇ angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4} = 1.25$	$\frac{4}{3} = 1.33$	$\frac{5}{3} = 1.67$	$\frac{2}{1} = 2.0$

Using (some) wave parameters for relativistic quantum theory

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Einstein (1905)

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space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\text{coth } \rho$	$e^{+\rho}$
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$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
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$$hv_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2 \quad \Leftarrow \text{for } (u \ll c) \Rightarrow \quad hK_{phase} \approx \frac{hB}{c^2} u$$

$$hv_{phase} \approx Mc^2 + \frac{1}{2} Mu^2 \quad \Leftarrow \text{for } (u \ll c) \Rightarrow \quad hK_{phase} \approx Mu$$

Lucky coincidences??

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Einstein (1905)

$$hcK_{phase} = hB \sinh \rho = Mc^2 \sinh \rho$$

$$cp = \frac{Mc u}{\sqrt{1-u^2/c^2}}$$

time	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\text{coth } \rho$	$e^{+\rho}$
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$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
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$$\frac{1}{\sqrt{\beta^2-1}} = \frac{\frac{u}{c}}{\sqrt{1-\frac{u^2}{c^2}}}$$

Using (some) wave parameters for relativistic quantum theory

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Einstein (1905)

$$hcK_{phase} = hB \sinh \rho = Mc^2 \sinh \rho$$

$$\frac{1}{\sqrt{\beta^{-2}-1}} = \frac{\frac{u}{c}}{\sqrt{1-\frac{u^2}{c^2}}}$$

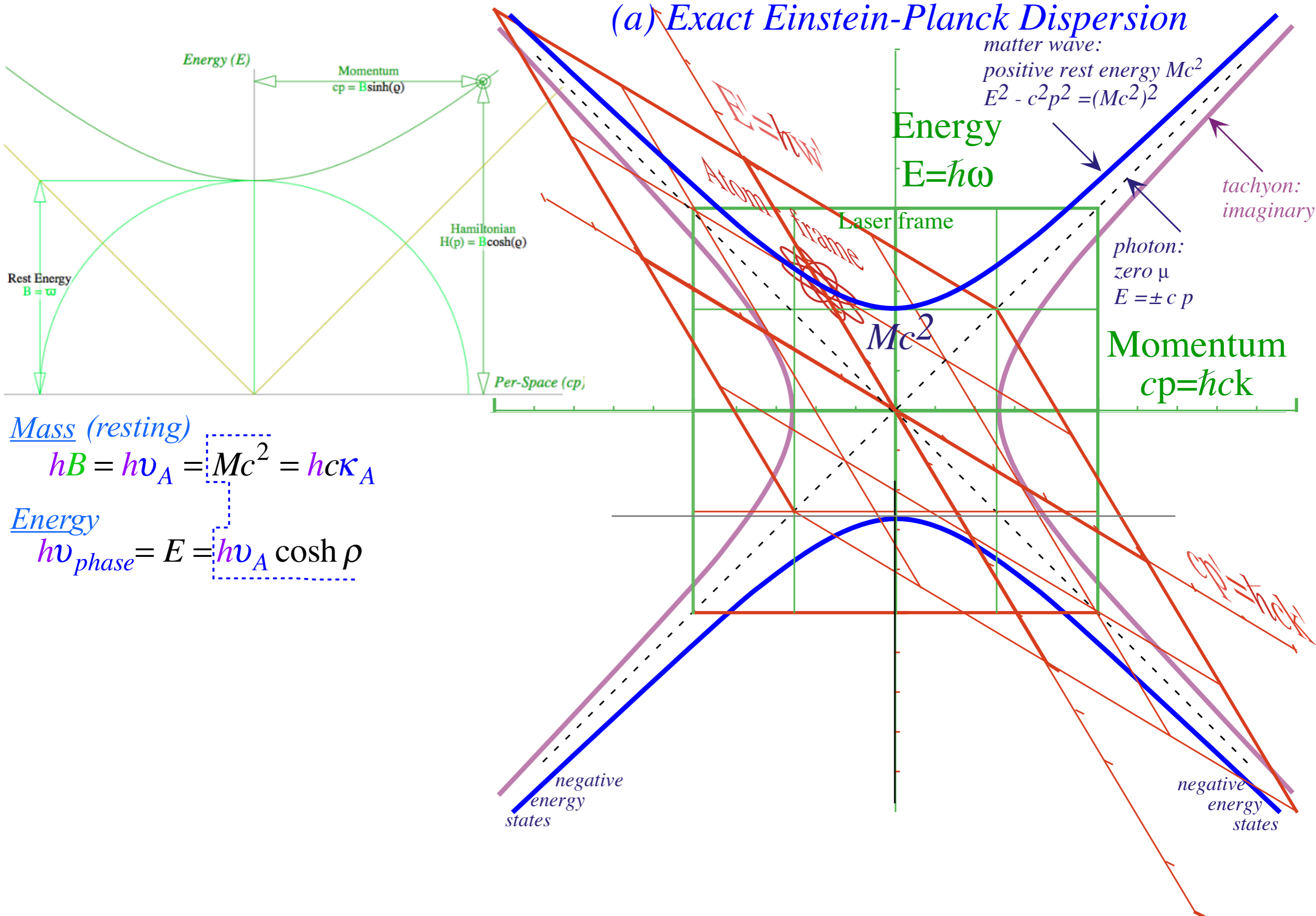
$$cp = \frac{Muc}{\sqrt{1-u^2/c^2}}$$

$$\text{Momentum: } hK_{phase} = p = \frac{Mu}{\sqrt{1-u^2/c^2}}$$

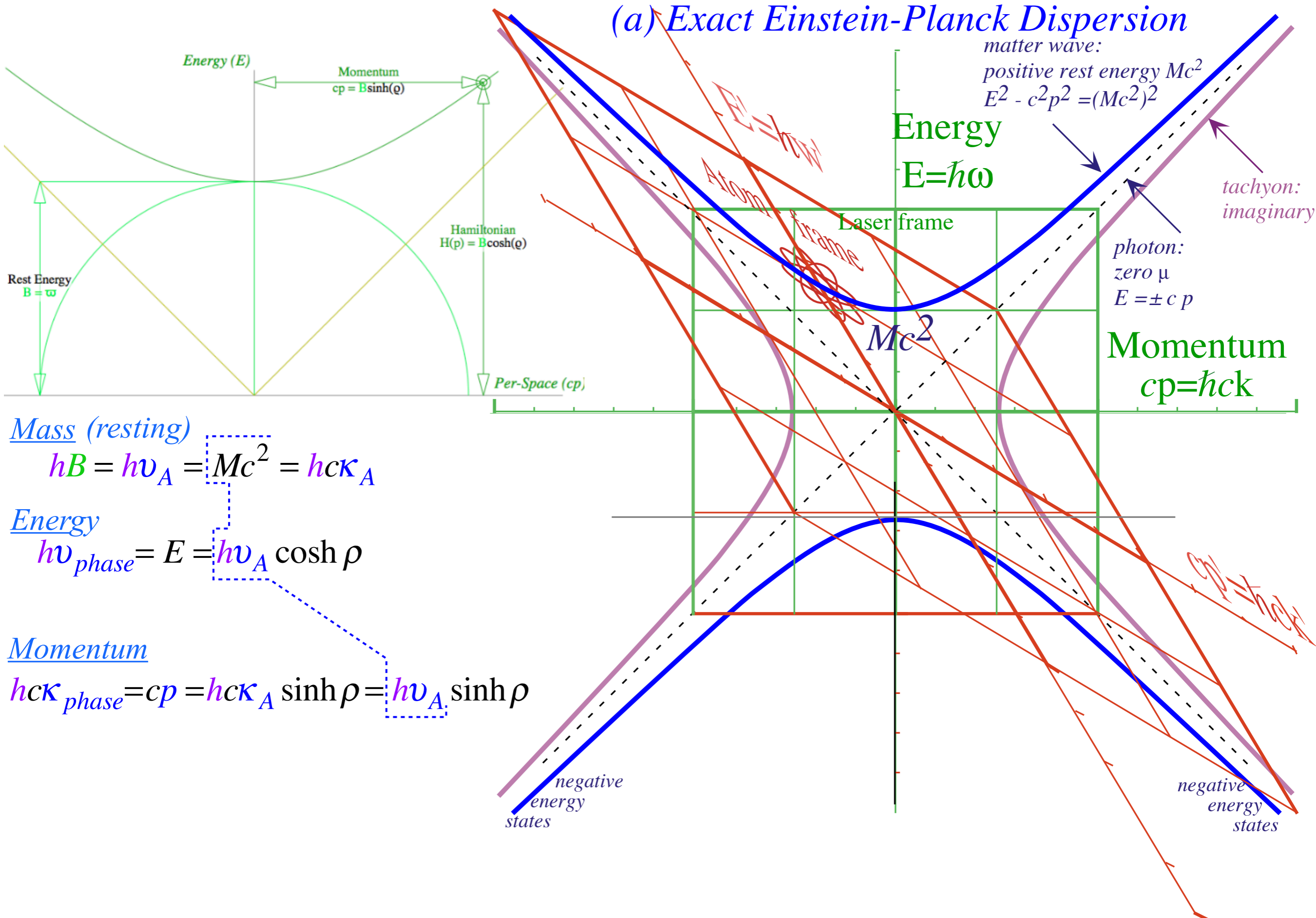
DeBroglie (1921)

time	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{K_{phase}}{K_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{K_{group}}{K_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
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$\beta \equiv \frac{u}{c}$	$\frac{\sqrt{1-\beta}}{\sqrt{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\frac{\sqrt{1+\beta}}{\sqrt{1-\beta}}$
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Using (some) wave coordinates for relativistic quantum theory

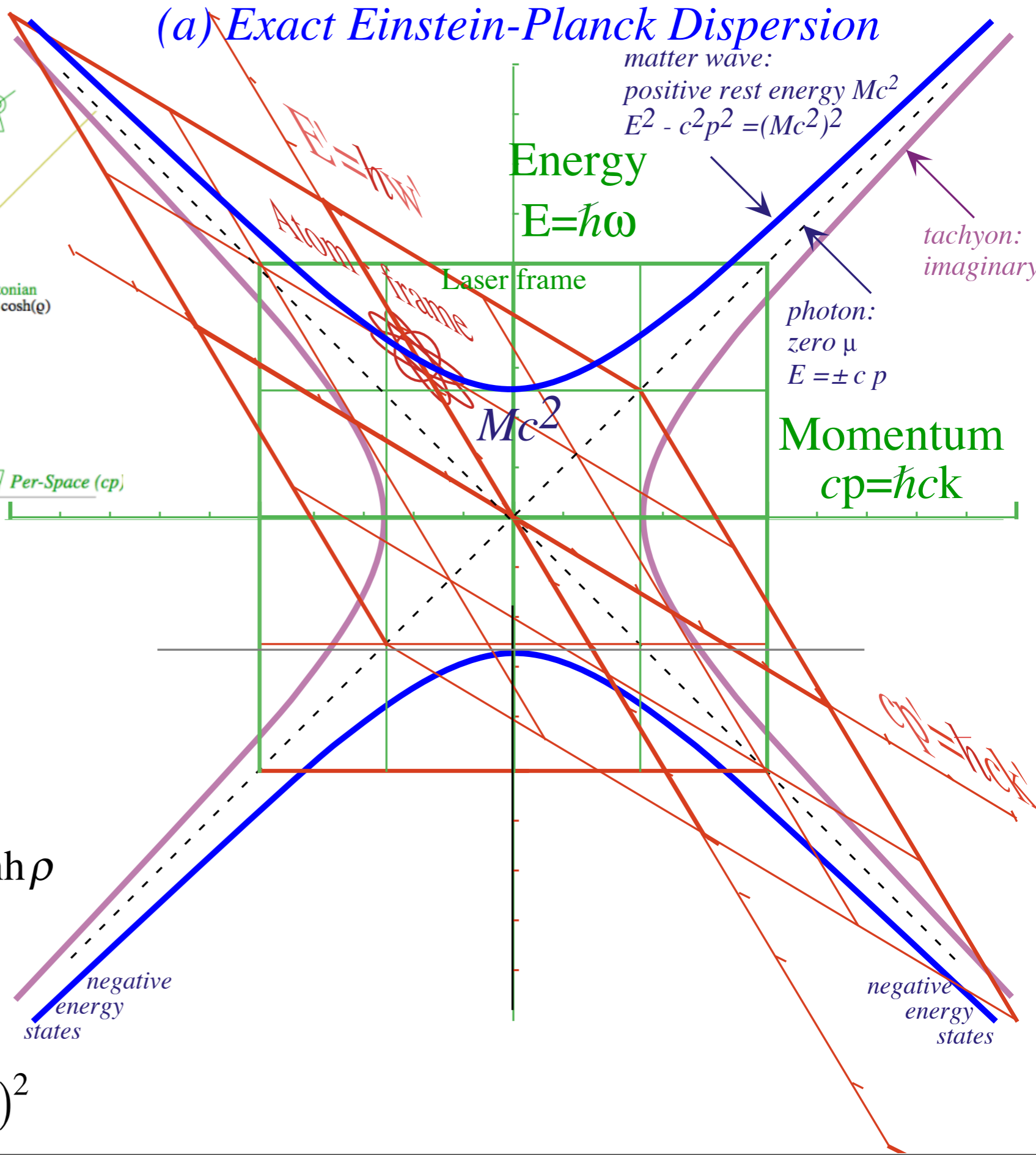
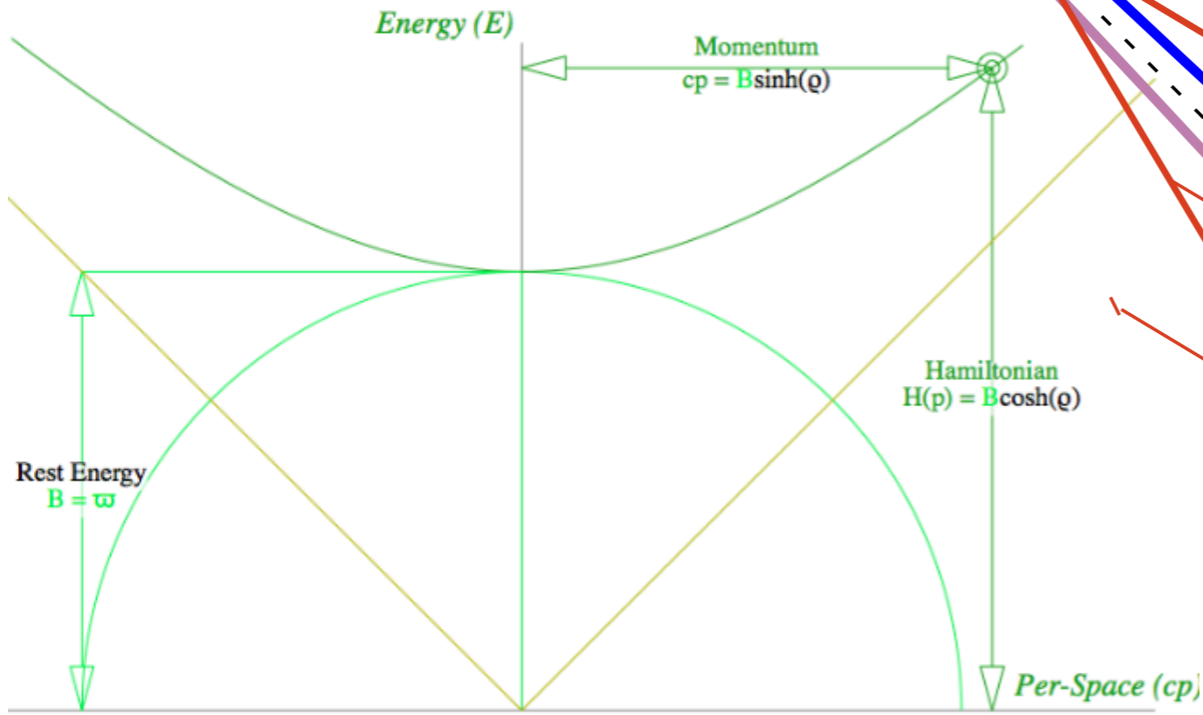


Using (some) wave coordinates for relativistic quantum theory



Using (some) wave coordinates for relativistic quantum theory

(a) Exact Einstein-Planck Dispersion



Mass (resting)

$$hB = h\omega_A = Mc^2 = hcK_A$$

Energy

$$h\omega_{phase} = E = h\omega_A \cosh \rho$$

Momentum

$$hcK_{phase} = cp = hcK_A \sinh \rho = h\omega_A \sinh \rho$$

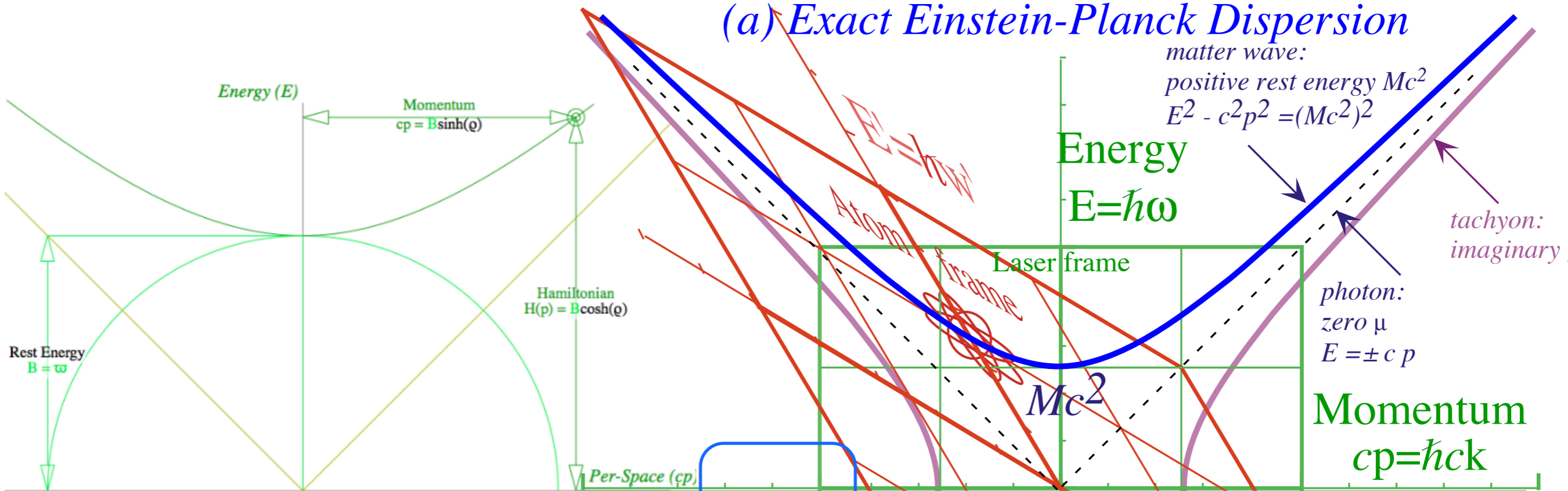
Energy versus Momentum

$$E^2 = (Mc^2)^2 \cosh^2 \rho$$

$$= (Mc^2)^2 (1 + \sinh^2 \rho) = (Mc^2)^2 + (cp)^2$$

Using (some) wave coordinates for relativistic quantum theory

(a) Exact Einstein-Planck Dispersion



Mass (resting)

$$hB = h\nu_A = Mc^2 = hc\kappa_A$$

Energy

$$h\nu_{phase} = E = h\nu_A \cosh \rho$$

Momentum

$$hc\kappa_{phase} = cp = hc\kappa_A \sinh \rho = h\nu_A \sinh \rho$$

Energy versus Momentum

$$E^2 = (Mc^2)^2 \cosh^2 \rho$$

$$= (Mc^2)^2 (1 + \sinh^2 \rho) = (Mc^2)^2 + (cp)^2 \Rightarrow E = \pm \sqrt{(Mc^2)^2 + (cp)^2}$$

The need for Negative Frequency arises!

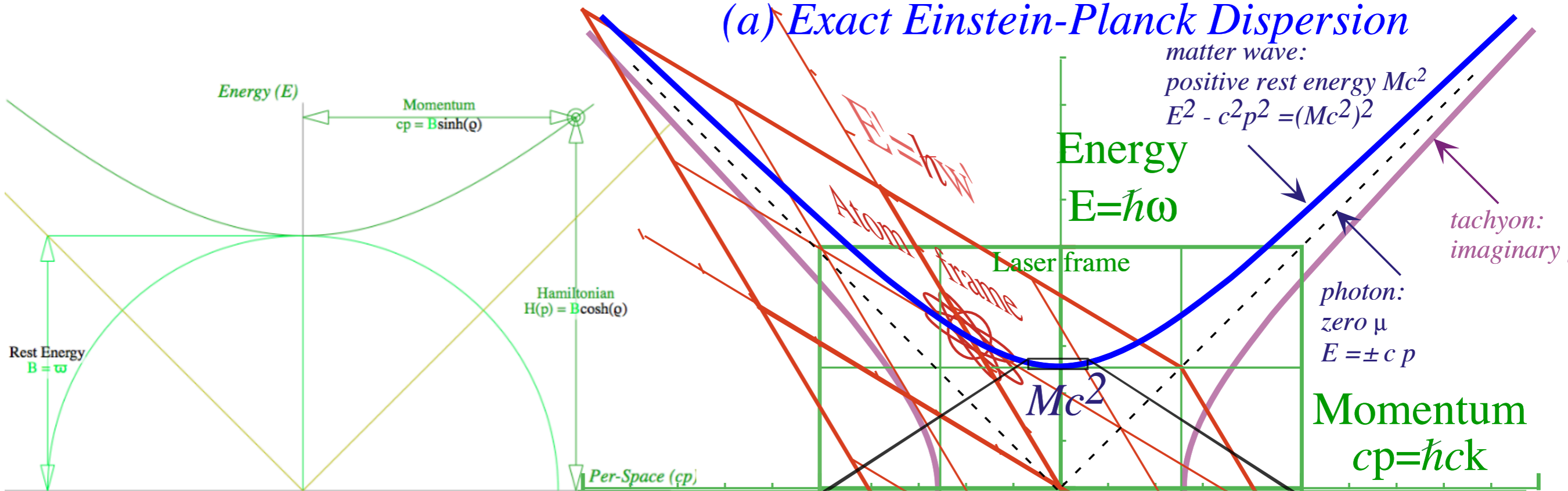
Phase conjugate light!

Counter-clockwise phasors!

negative energy states

negative energy states

Using (some) wave coordinates for relativistic quantum theory



Mass (resting)

$$hB = \hbar\omega_A = Mc^2 = \hbar ck_A$$

Energy

$$\hbar\omega_{phase} = E = \hbar\omega_A \cosh \rho$$

Momentum

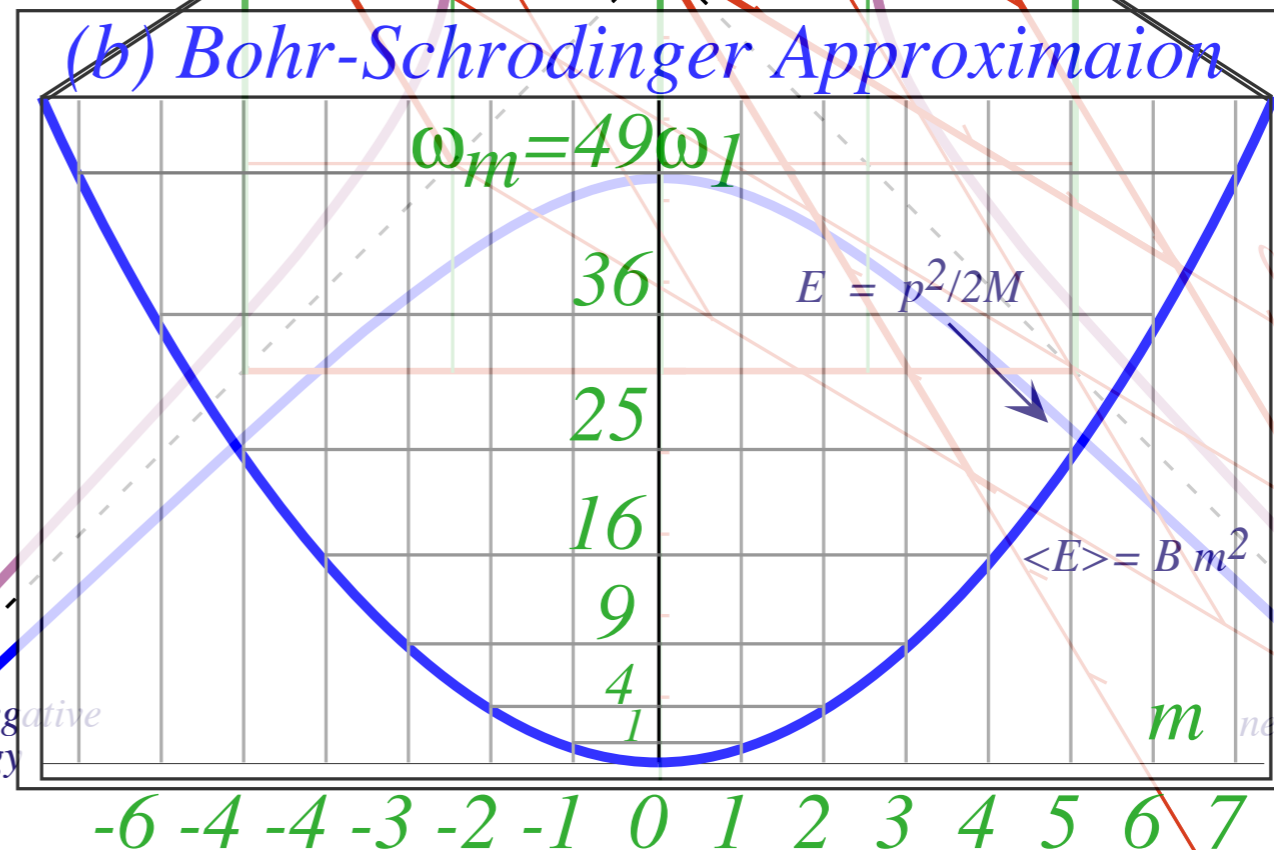
$$\hbar ck_{phase} = cp = \hbar ck_A \sinh \rho = \hbar\omega_A \sinh \rho$$

Energy versus Momentum

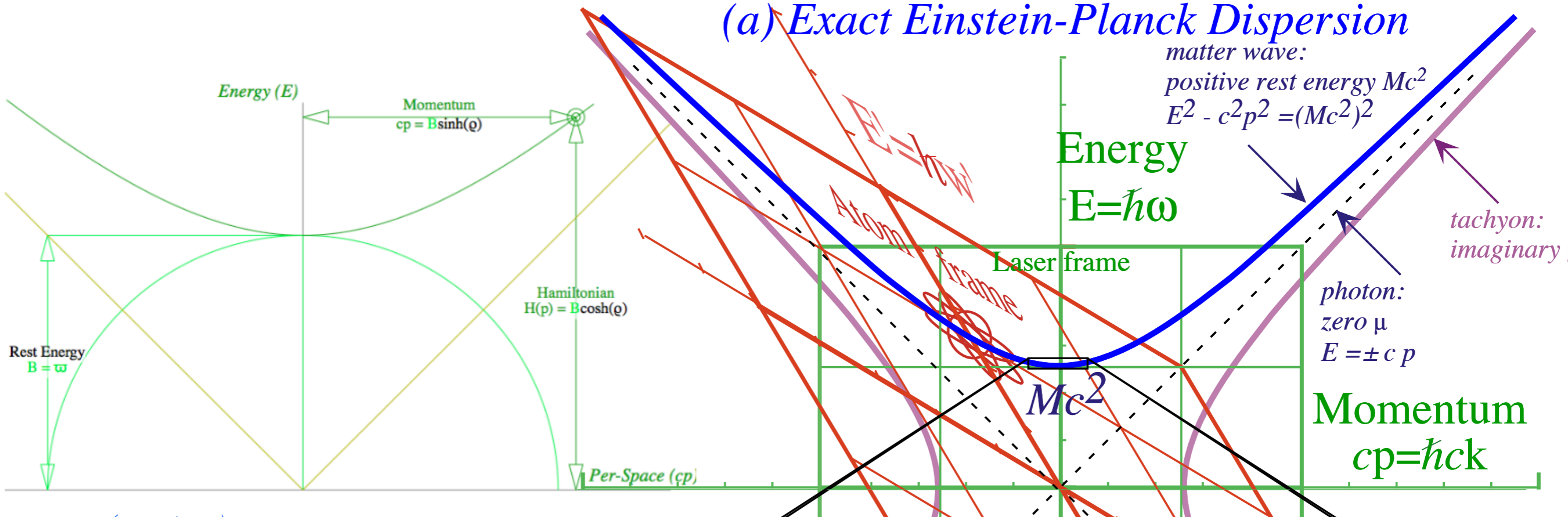
$$E^2 = (Mc^2)^2 \cosh^2 \rho$$

$$= (Mc^2)^2 (1 + \sinh^2 \rho) = (Mc^2)^2 + (cp)^2 \Rightarrow E = \pm \sqrt{(Mc^2)^2 + (cp)^2} \approx Mc^2 + \frac{p^2}{2M}$$

(b) *Bohr-Schrodinger Approximaion*



Using (some) wave coordinates for relativistic quantum theory



Mass (resting)

$$hB = h\nu_A = Mc^2 = hck_A$$

Energy

$$h\nu_{phase} = E = h\nu_A \cosh \rho$$

Momentum

$$hck_{phase} = cp = hck_A \sinh \rho = h\nu_A \sinh \rho$$

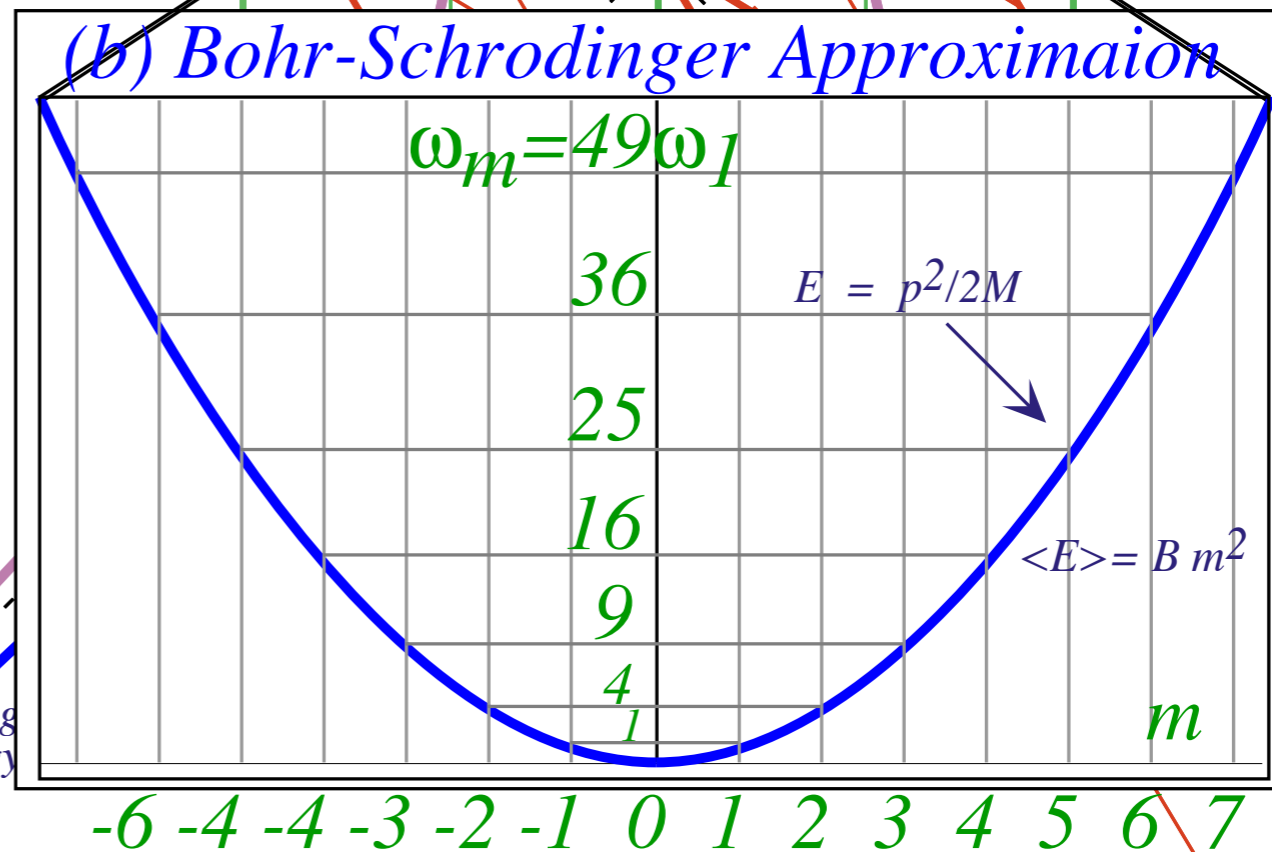
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low speed approximation

(b) Bohr-Schrodinger Approximaion



Definition(s) of mass for relativity/quantum

Given: Energy: $E = Mc^2 \cosh \rho$

$$= h\nu_{\text{phase}}$$

Rest Mass M_{rest} (Einstein's mass)

$$h\mathbf{B} = h\nu_A = Mc^2 = hc\mathbf{K}_A$$

Defines invariant hyperbola(s)

$$E = \pm \sqrt{(Mc^2)^2 + (cp)^2}$$

momentum: $cp = Mc^2 \sinh \rho$

$$= hc\mathbf{K}_{\text{phase}}$$

velocity: $u = c \tanh \rho = \frac{d\nu}{d\mathbf{K}}$

Definition(s) of mass for relativity/quantum

Given: Energy: $E = Mc^2 \cosh \rho$

$$= h\nu_{phase}$$

momentum: $cp = Mc^2 \sinh \rho$

$$= hc\kappa_{phase}$$

velocity: $u = c \tanh \rho = \frac{d\nu}{d\kappa}$

Defines invariant hyperbola(s)

$$E = \pm \sqrt{(Mc^2)^2 + (cp)^2}$$

Rest Mass M_{rest} (Einstein's mass)

$$hB = h\nu_A = Mc^2 = hc\kappa_A$$

$$\frac{h\nu_{phase}}{c^2} = M_{rest} = \frac{hc\kappa_{phase}}{c^2} \quad \frac{\text{Rest}}{\text{Mass}}$$

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$$= hcK_{phase}$$

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velocity: $u = c \tanh \rho = \frac{d\nu}{dK}$

Momentum Mass M_{mom} (Galileo's mass) Defined by ratio p/u of relativistic momentum to group velocity.

$$M_{mom} \equiv \frac{p}{u} = \frac{M_{rest} c \sinh \rho}{c \tanh \rho}$$

Definition(s) of mass for relativity/quantum

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Definition(s) of mass for relativity/quantum

Given: Energy: $E = Mc^2 \cosh \rho$

$= h\nu_{phase}$

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Defines invariant hyperbola(s)

momentum: $cp = Mc^2 \sinh \rho$

$hB = h\nu_A = Mc^2 = h\kappa_A$

$E = \pm \sqrt{(Mc^2)^2 + (cp)^2}$

$= h\kappa_{phase}$

$$\frac{h\nu_{phase}}{c^2} = M_{rest} = \frac{h\kappa_{phase}}{c^2} \quad \frac{\text{Rest}}{\text{Mass}}$$

velocity: $u = c \tanh \rho = \frac{d\nu}{d\kappa}$

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$$M_{mom} \equiv \frac{p}{u} = \frac{M_{rest} c \sinh \rho}{c \tanh \rho}$$

Limiting cases: $M_{mom} \xrightarrow{u \rightarrow c} M_{rest} e^{\rho} / 2$

$M_{mom} \xrightarrow{u \ll c} M_{rest}$

$$= M_{rest} \cosh \rho = \frac{M_{rest}}{\sqrt{1 - u^2 / c^2}} \quad \frac{\text{Momentum}}{\text{Mass}}$$

Definition(s) of mass for relativity/quantum

Given: Energy: $E = Mc^2 \cosh \rho$

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Defines invariant hyperbola(s)

momentum: $cp = Mc^2 \sinh \rho$

$hB = h\nu_A = Mc^2 = hck_A$

$E = \pm \sqrt{(Mc^2)^2 + (cp)^2}$

$= hck_{phase}$

$\frac{h\nu_{phase}}{c^2} = M_{rest} = \frac{hck_{phase}}{c^2}$ Rest Mass

velocity: $u = c \tanh \rho = \frac{d\nu}{dk}$

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$M_{mom} \xrightarrow{u \ll c} M_{rest}$

$= M_{rest} \cosh \rho = \frac{M_{rest}}{\sqrt{1 - u^2 / c^2}}$ Momentum Mass

Effective Mass M_{eff} (Newton's mass) Defined by ratio $F/a = dp/du$ of relativistic force to acceleration.

Definition(s) of mass for relativity/quantum

Given: Energy: $E = Mc^2 \cosh \rho$

$= h\nu_{phase}$

momentum: $cp = Mc^2 \sinh \rho$

$= hc\kappa_{phase}$

velocity: $u = c \tanh \rho = \frac{d\nu}{d\kappa}$

Defines invariant hyperbola(s)

$$E = \pm \sqrt{(Mc^2)^2 + (cp)^2}$$

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Effective Mass M_{eff} (Newton's mass) Defined by ratio $F/a = dp/du$ of relativistic force to acceleration.

That is ratio of change $dp = Mc \cosh \rho d\rho$ in momentum to change $du = c \operatorname{sech}^2 \rho d\rho$ in velocity

Definition(s) of mass for relativity/quantum

Given: Energy: $E = Mc^2 \cosh \rho$

$= h\nu_{phase}$

momentum: $cp = Mc^2 \sinh \rho$

$= hc\kappa_{phase}$

velocity: $u = c \tanh \rho = \frac{d\nu}{d\kappa}$

Defines invariant hyperbola(s)

$$E = \pm \sqrt{(Mc^2)^2 + (cp)^2}$$

Rest Mass M_{rest} (Einstein's mass)

$$hB = h\nu_A = Mc^2 = hc\kappa_A$$

$$\frac{h\nu_{phase}}{c^2} = M_{rest} = \frac{hc\kappa_{phase}}{c^2} \quad \frac{\text{Rest}}{\text{Mass}}$$

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Limiting cases: $M_{mom} \xrightarrow{u \rightarrow c} M_{rest} e^{\rho} / 2$

$M_{mom} \xrightarrow{u \ll c} M_{rest}$

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More common derivation using group velocity: $u \equiv V_{group} = \frac{d\omega}{dk} = \frac{d\nu}{dk}$

$M_{eff} \equiv \frac{dp}{du} = \frac{\hbar dk}{dV_{group}} = \frac{\hbar}{\frac{d}{dk} \frac{d\omega}{dk}} = \frac{\hbar}{\frac{d^2 \omega}{dk^2}} = \frac{M_{rest}}{(1 - u^2/c^2)^{3/2}}$

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general wave formula to accompany $V_{group} = \frac{d\omega}{dk}$

Definition(s) of mass for relativity/quantum

How much does a γ -photon weigh?

(a) γ -rest mass: $M_{rest}^{\gamma} = 0,$

(b) γ -momentum mass: $M_{mom}^{\gamma} = \frac{p}{c} = \frac{h\kappa}{c} = \frac{h\nu}{c^2},$

(c) γ -effective mass: $M_{eff}^{\gamma} = \infty.$

Newton complained about
his “corpuscles” of light having
“fits” (going crazy).

$$M_{mom}^{\gamma} = \frac{h\nu}{c^2} = \nu(1.2 \cdot 10^{-51}) \text{kg} \cdot \text{s} = 4.5 \cdot 10^{-36} \text{kg} \quad (\text{for: } \nu=600\text{THz})$$

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Newton complained about
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For him this would be evidence
of optical-triple-schizophrenia!

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Relativistic action S and Lagrangian-Hamiltonian relations

Define *Lagrangian* L using invariant wave phase $\Phi = kx - \omega t = k'x' - \omega't'$ for wave of $k = k_{phase}$ and $\omega = \omega_{phase}$.

$$L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega$$

$$\begin{aligned} h\nu_A &= Mc^2 = hc\kappa_A \\ h\nu_{phase} &= E = h\nu_A \cosh \rho \\ hc\kappa_{phase} &= cp = h\nu_A \sinh \rho \end{aligned}$$

Prior wave relations

← linear Hz format angular phasor →
format format

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Legendre transformation

$$p = \hbar k = Mc \sinh \rho$$

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Use *Group velocity* : $u = \frac{dx}{dt} = c \tanh \rho$

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$$L \text{ is : } Mc^2 \frac{-1}{\cosh \rho} = -Mc^2 \operatorname{sech} \rho$$

$$\hbar \nu_A = Mc^2 = \hbar c \kappa_A$$

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Prior wave relations

← linear Hz
format

angular phasor
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Note: $Mc u = Mc^2 \tanh \rho$

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Compare *Lagrangian* L

$$L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho$$

$$\hbar \nu_A = Mc^2 = \hbar c \kappa_A$$

Prior wave relations

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← linear Hz format

angular phasor →

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format

format

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Relativistic action S and Lagrangian-Hamiltonian relations

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Compare *Lagrangian* L

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with *Hamiltonian* $H=E$

$$H = \hbar\omega = -Mc^2 / \sqrt{1 - \frac{u^2}{c^2}} = Mc^2 \cosh \rho$$

$$\hbar\omega_A = Mc^2 = \hbar c k_A$$

Prior wave relations

$$\hbar\omega_{phase} = E = \hbar\omega_A \cosh \rho$$

← linear Hz

angular phasor →

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Also: $cp = Mc^2 \sinh \rho$

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Prior wave relations

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← linear Hz

angular phasor →

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Legendre transformation

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$$= Mc^2 \sin \sigma$$

Also: $cp = Mc^2 \sinh \rho$

$$= \hbar ck = Mc^2 \tan \sigma$$

Compare Lagrangian L

$$L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho = -Mc^2 \cos \sigma$$

with Hamiltonian $H = E$

$$H = \hbar \omega = -Mc^2 / \sqrt{1 - \frac{u^2}{c^2}} = Mc^2 \cosh \rho = Mc^2 \sec \sigma$$

$$= Mc^2 \sqrt{1 + \sinh^2 \rho} = Mc^2 \sqrt{1 + (cp)^2}$$

Including stellar angle σ

$$\hbar \nu_A = Mc^2 = \hbar c \kappa_A$$

Prior wave relations

$$\hbar \nu_{phase} = E = \hbar \nu_A \cosh \rho$$

← linear Hz format

angular phasor →

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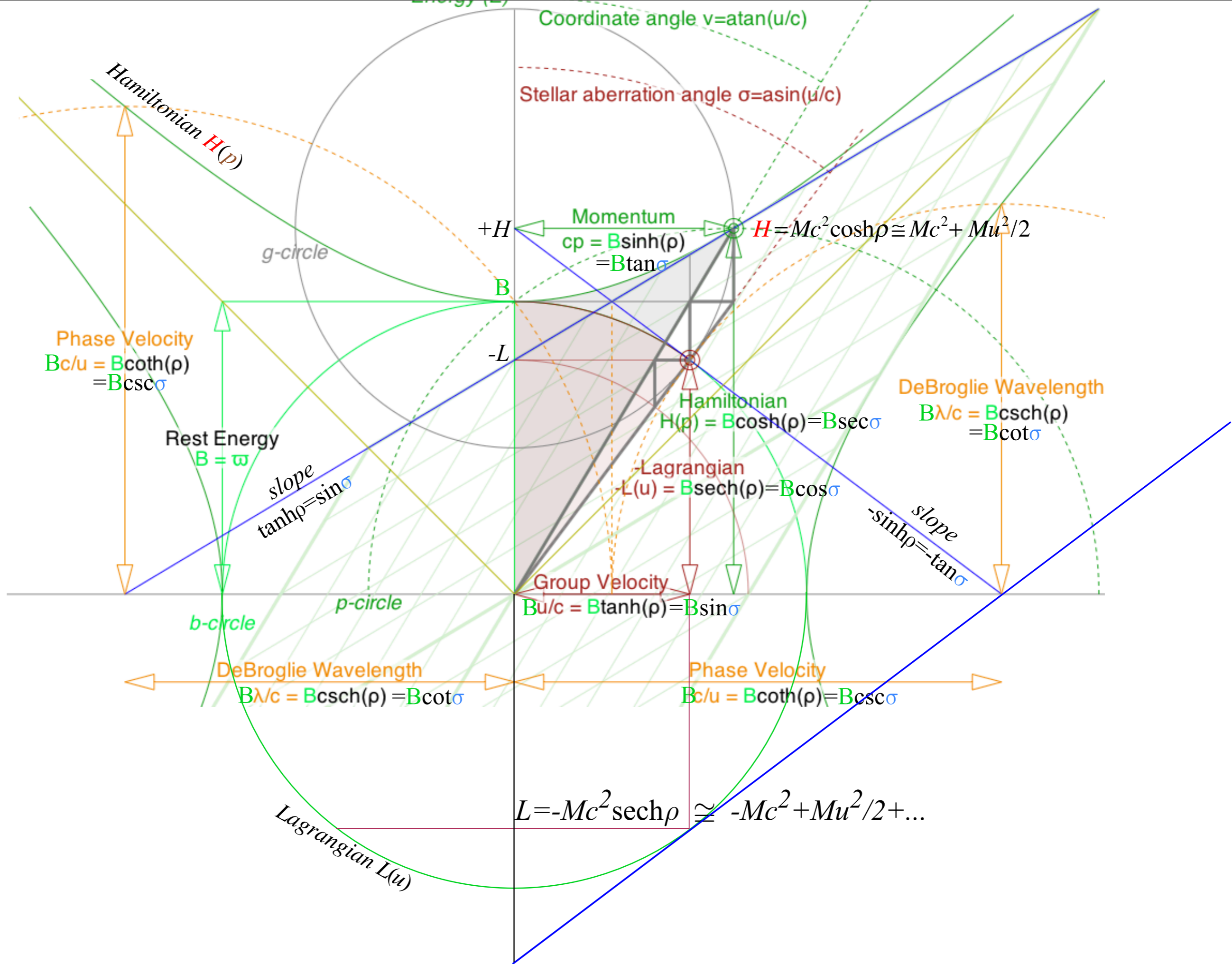
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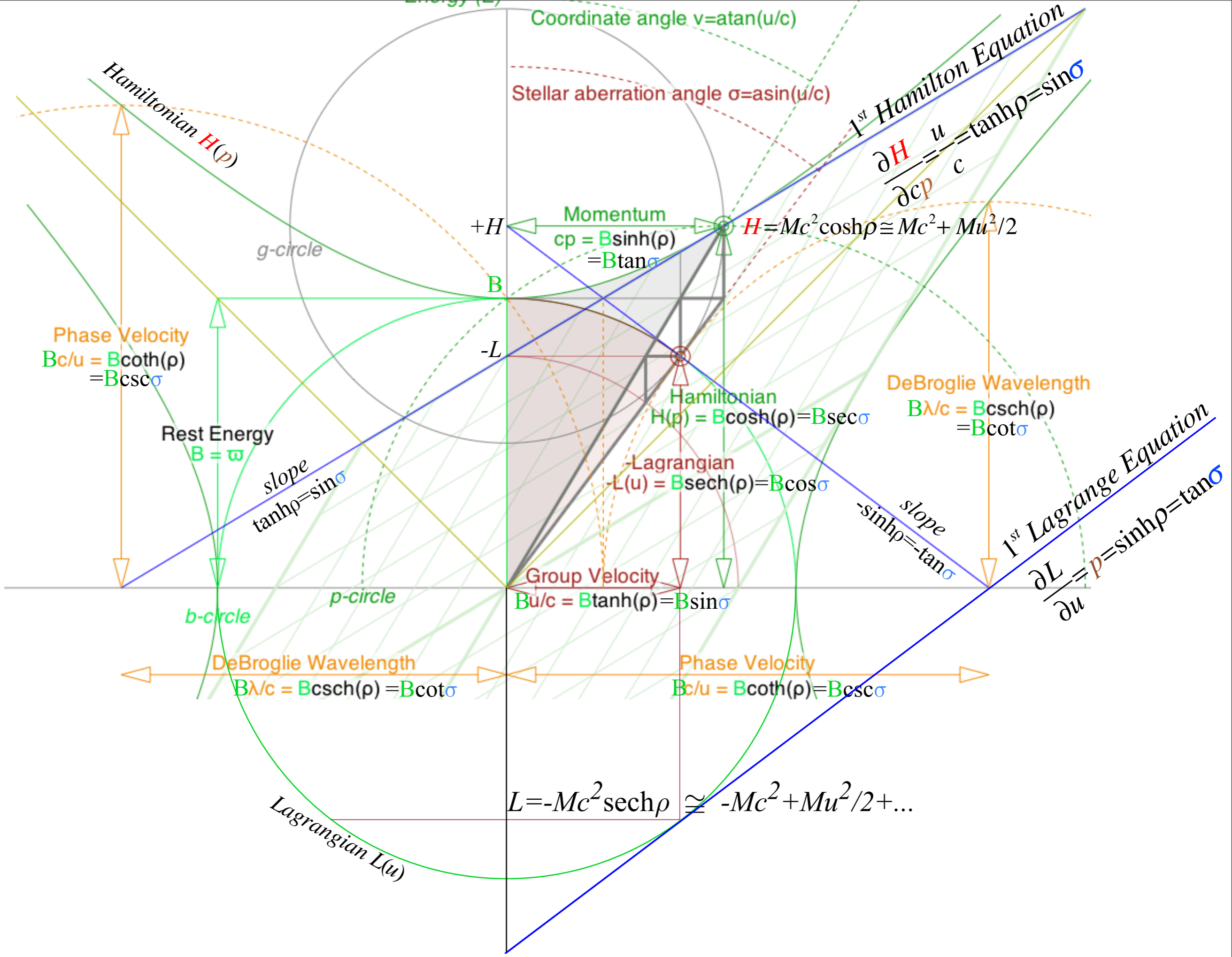
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Relativistic **action S** and Lagrangian-Hamiltonian relations

Define *Lagrangian L* using invariant wave phase $\Phi = kx - \omega t = k'x' - \omega't'$ for wave of $k = k_{phase}$ and $\omega = \omega_{phase}$.

Use DeBroglie-momentum $p = \hbar k$ relation and Planck-energy $E = \hbar \omega$ relation to define *Hamiltonian H = E*

$$\frac{dS}{dt} = L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E \equiv p\dot{x} - E \equiv pu - H = L \quad \text{Legendre transformation}$$

Compare *Lagrangian L*

$$\dot{S} = L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho = -Mc^2 \cos \sigma$$

with *Hamiltonian H = E*

$$H = \hbar \omega = -Mc^2 / \sqrt{1 - \frac{u^2}{c^2}} = Mc^2 \cosh \rho = Mc^2 \sec \sigma$$

$$= Mc^2 \sqrt{1 + \sinh^2 \rho} = Mc^2 \sqrt{1 + (cp)^2}$$

Define *Action S = \hbar \Phi*

$$\begin{aligned} \hbar \nu_A &= Mc^2 = \hbar c \kappa_A \\ \hbar \nu_{phase} &= E = \hbar \nu_A \cosh \rho \\ \hbar c \kappa_{phase} &= cp = \hbar \nu_A \sinh \rho \end{aligned}$$

Prior wave relations

← linear Hz format angular phasor format →

$$\begin{aligned} \hbar \omega_A &= Mc^2 = \hbar c k_A \\ \hbar \omega_{phase} &= E = \hbar \omega_A \cosh \rho \\ \hbar c k_{phase} &= cp = \hbar \omega_A \sinh \rho \end{aligned}$$

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$$dS \equiv Ldt \equiv \hbar d\Phi = \hbar k dx - \hbar \omega dt = p dx - H dt$$

Poincare Invariant action differential

Compare *Lagrangian* L

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Define *Action* $S = \hbar \Phi$

Prior wave relations

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← linear Hz
format

angular phasor
format

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Legendre transformation

Use *Group velocity* : $u = \frac{dx}{dt} = c \tanh \rho$

$$dS \equiv L dt \equiv \hbar d\Phi = \hbar k dx - \hbar \omega dt = p dx - H dt$$

Poincare Invariant action differential

$$\frac{\partial S}{\partial x} = p \quad \frac{\partial S}{\partial t} = -H$$

Hamilton-Jacobi equations

Compare *Lagrangian* L

$$\dot{S} = L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho = -Mc^2 \cos \sigma$$

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Prior wave relations

← linear Hz format angular phasor →
format format

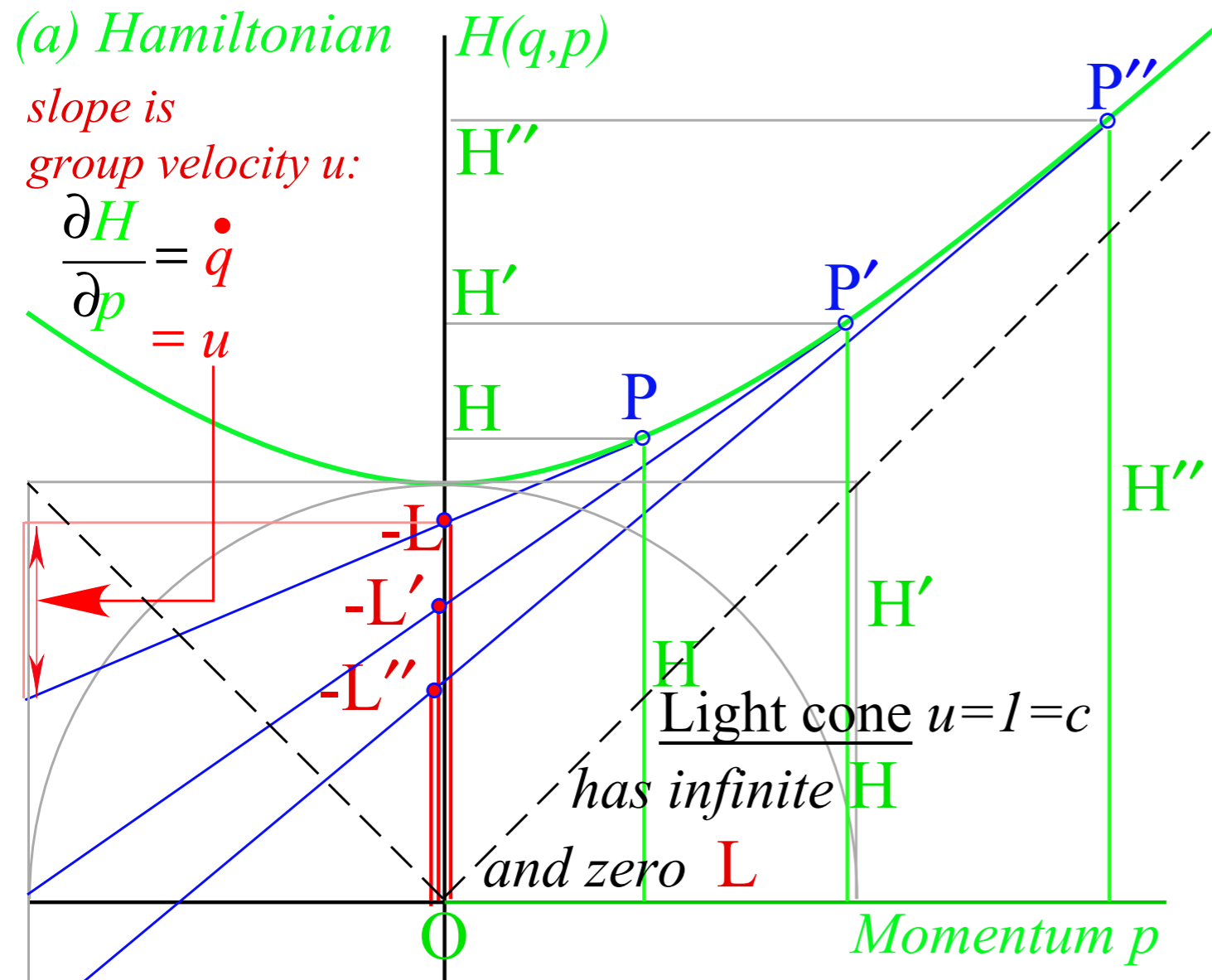
$$\begin{aligned} \hbar \omega_A &= Mc^2 = \hbar c k_A \\ \hbar \omega_{phase} &= E = \hbar \omega_A \cosh \rho \\ \hbar c k_{phase} &= cp = \hbar \omega_A \sinh \rho \end{aligned}$$

Poincare Invariant Action $dS=Ldt=p dq-H dt=\hbar d\Phi$ (phase)

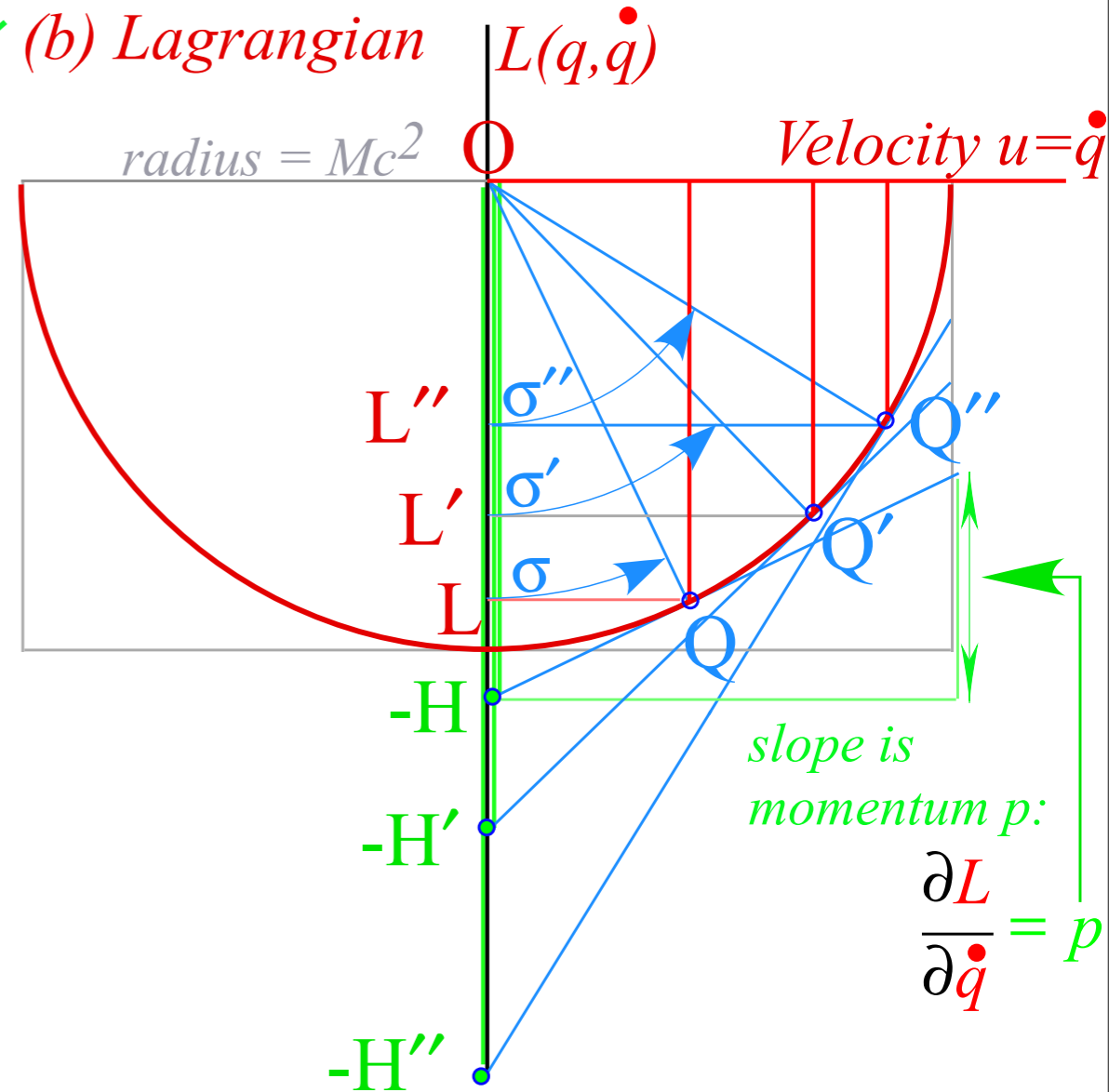
Hamiltonian $H(p,q)=p\dot{q}-L$ vs. Lagrangian $L(\dot{q},q)=p\dot{q}-H$

Contact transformation: (slope, -intercept) of H (or L) tangent determines the (X, Y coordinates) of L (or H).

(Also, called a Legendre contact transformation which is a special case of a Huygens transformation that uses contacting tangent curves instead of lines.)



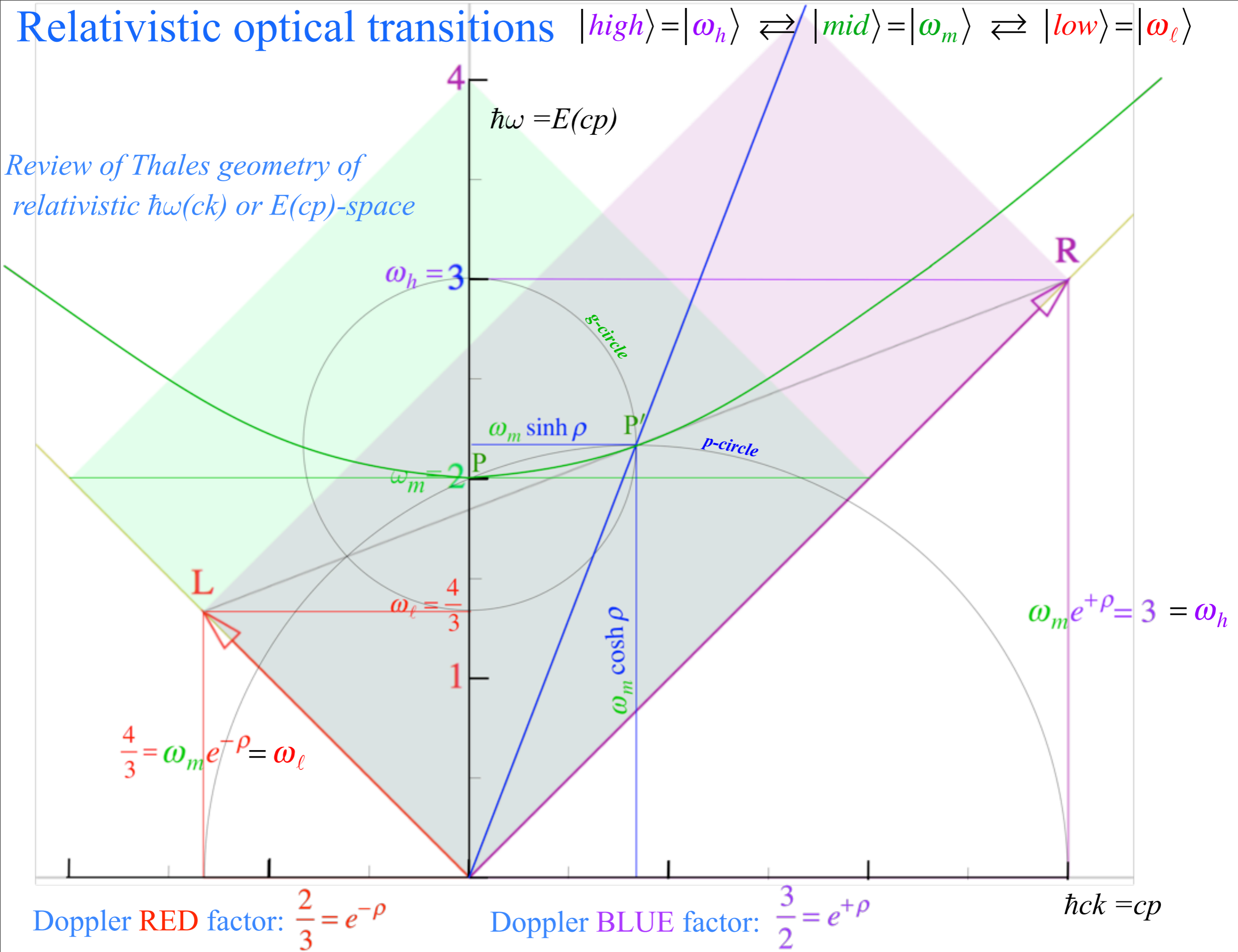
Here *slope* is group velocity $u=\dot{q}$
Y-coordinate is *energy* $H=\hbar\omega$



Here *slope* is momentum p
Y-coordinate is *phase rate* $L=\hbar\Phi$

Relativistic optical transitions $|high\rangle = |\omega_h\rangle \rightleftharpoons |mid\rangle = |\omega_m\rangle \rightleftharpoons |low\rangle = |\omega_l\rangle$

Review of Thales geometry of relativistic $\hbar\omega(ck)$ or $E(cp)$ -space



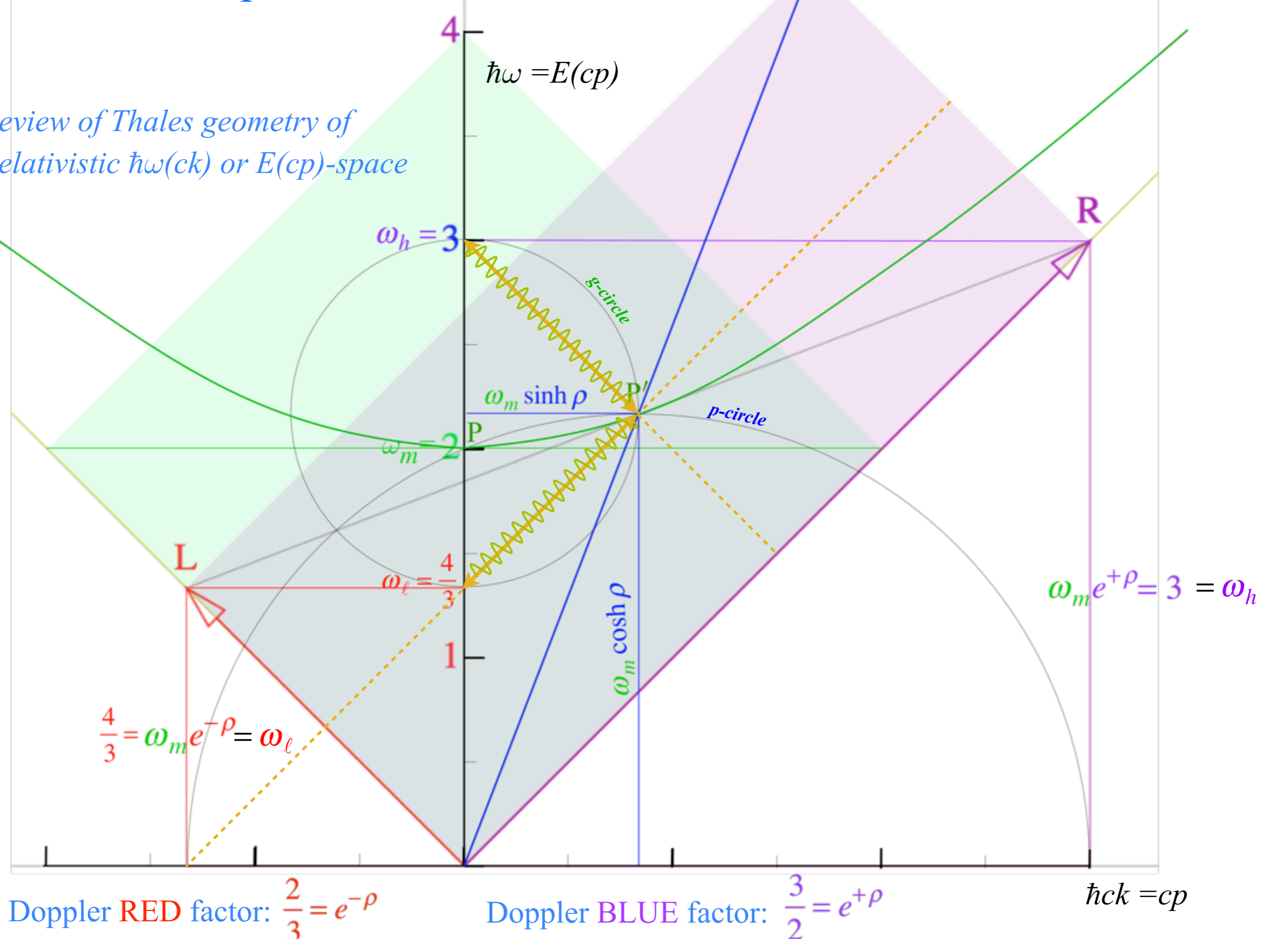
Doppler RED factor: $\frac{2}{3} = e^{-\rho}$

Doppler BLUE factor: $\frac{3}{2} = e^{+\rho}$

$\hbar ck = cp$

Relativistic optical transitions $|high\rangle = |\omega_h\rangle \rightleftharpoons |mid\rangle = |\omega_m\rangle \rightleftharpoons |low\rangle = |\omega_l\rangle$

Review of Thales geometry of relativistic $\hbar\omega(ck)$ or $E(cp)$ -space



Doppler RED factor: $\frac{2}{3} = e^{-\rho}$

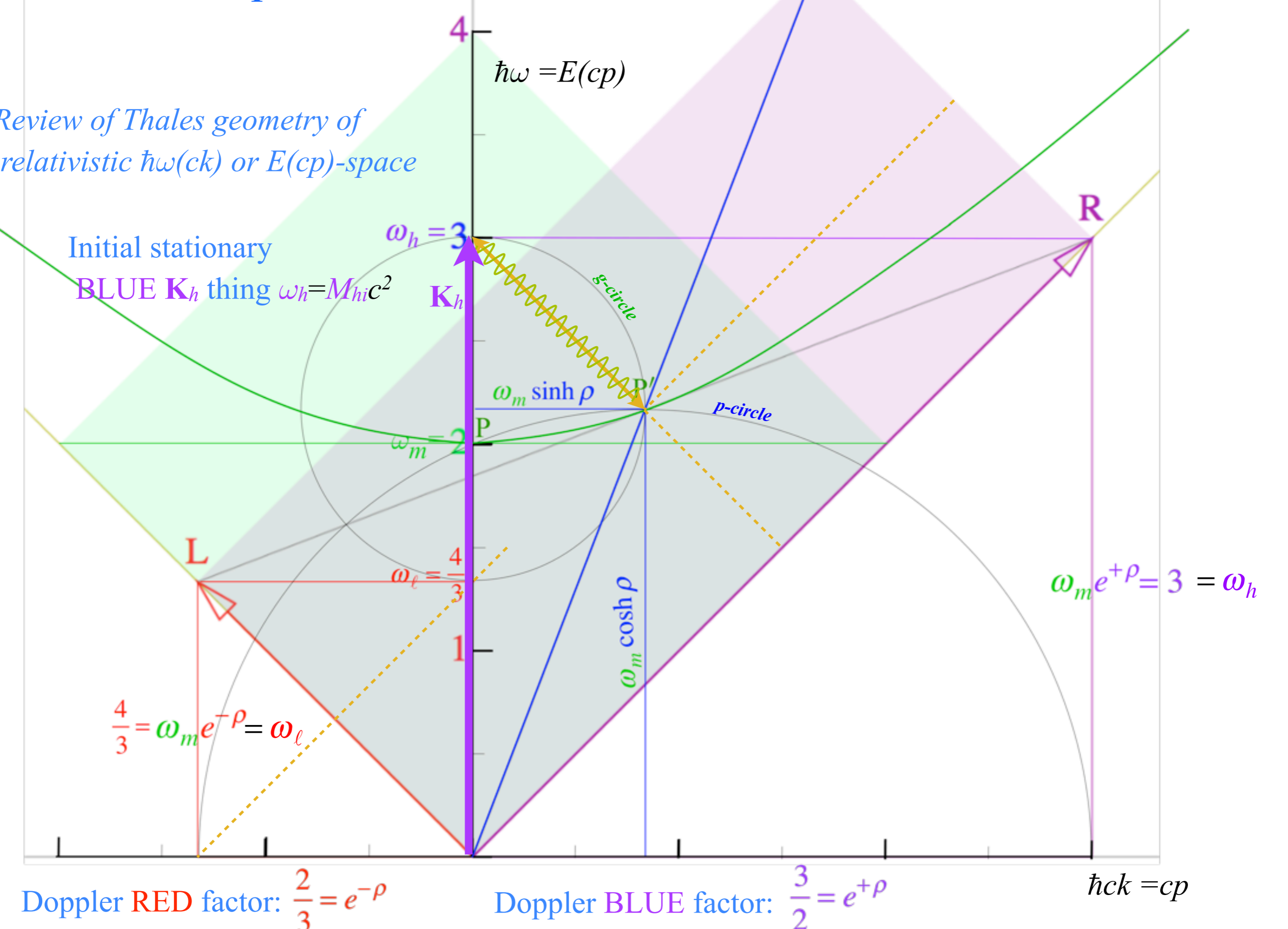
Doppler BLUE factor: $\frac{3}{2} = e^{+\rho}$

$\hbar ck = cp$

Relativistic optical transitions $|high\rangle = |\omega_h\rangle \rightleftharpoons |mid\rangle = |\omega_m\rangle \rightleftharpoons |low\rangle = |\omega_l\rangle$

Review of Thales geometry of relativistic $\hbar\omega(ck)$ or $E(cp)$ -space

Initial stationary BLUE K_h thing $\omega_h = M_{hi}c^2$

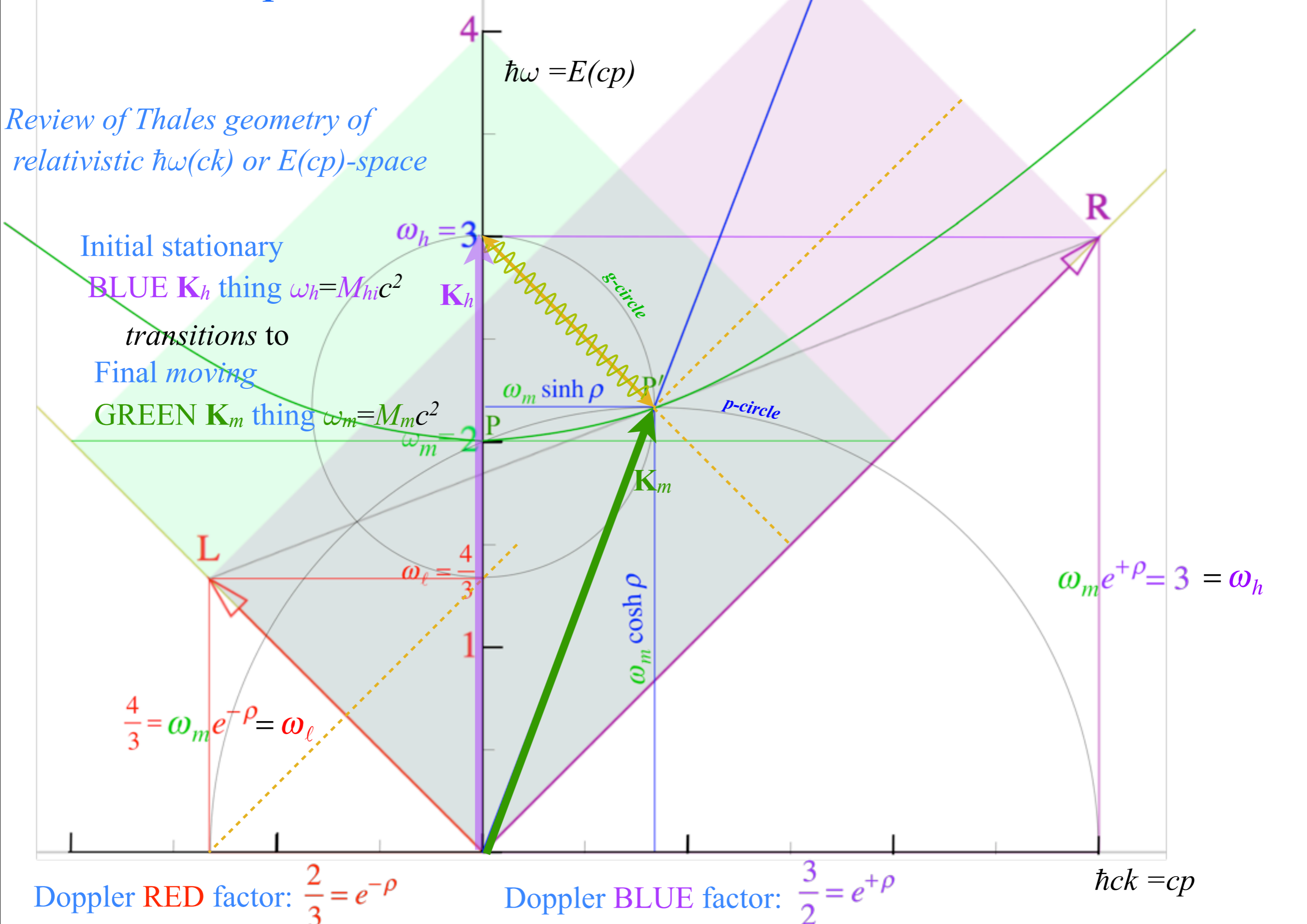


Doppler RED factor: $\frac{2}{3} = e^{-\rho}$

Doppler BLUE factor: $\frac{3}{2} = e^{+\rho}$

$\hbar ck = cp$

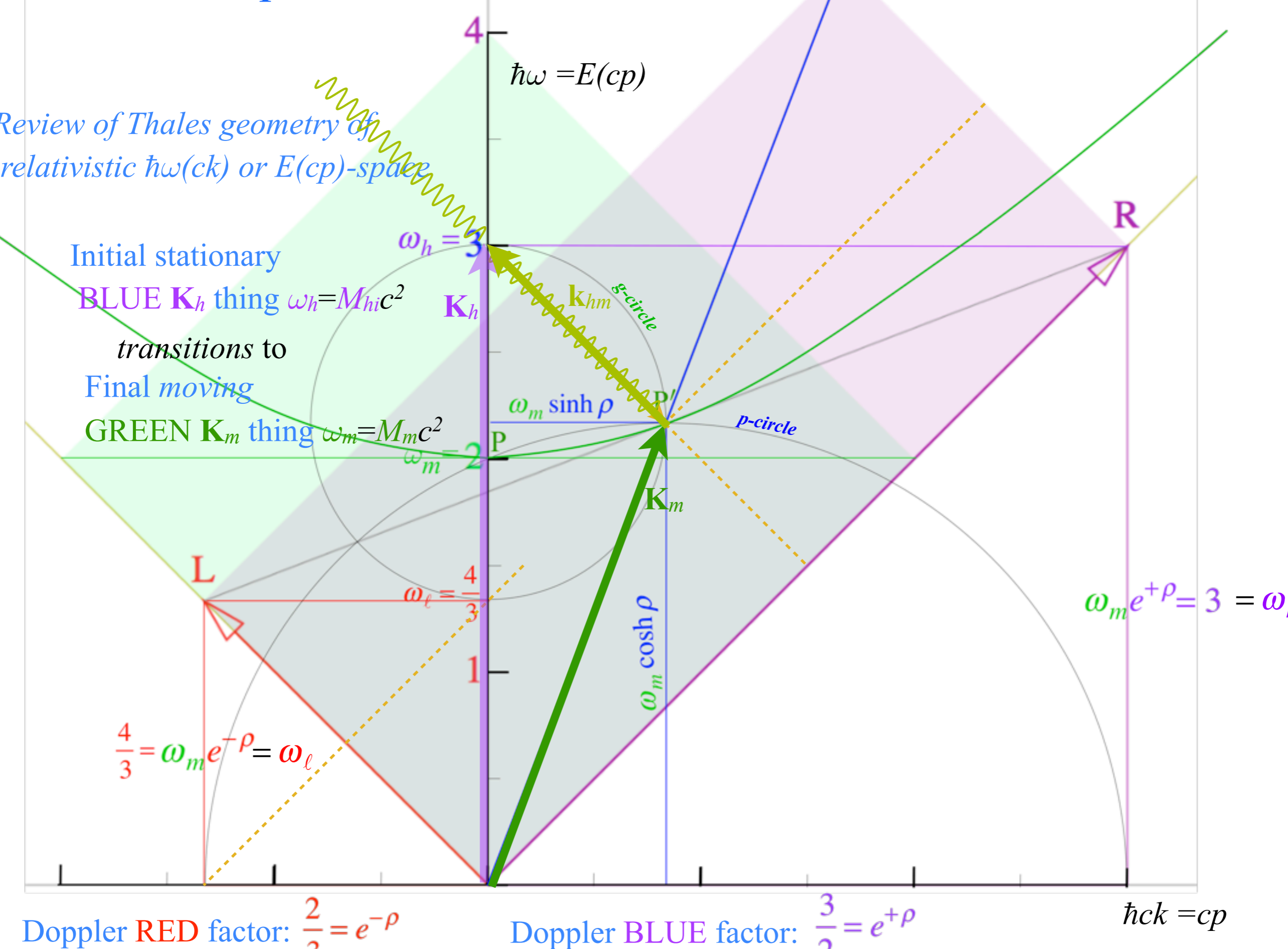
Relativistic optical transitions $|high\rangle = |\omega_h\rangle \rightleftharpoons |mid\rangle = |\omega_m\rangle \rightleftharpoons |low\rangle = |\omega_l\rangle$



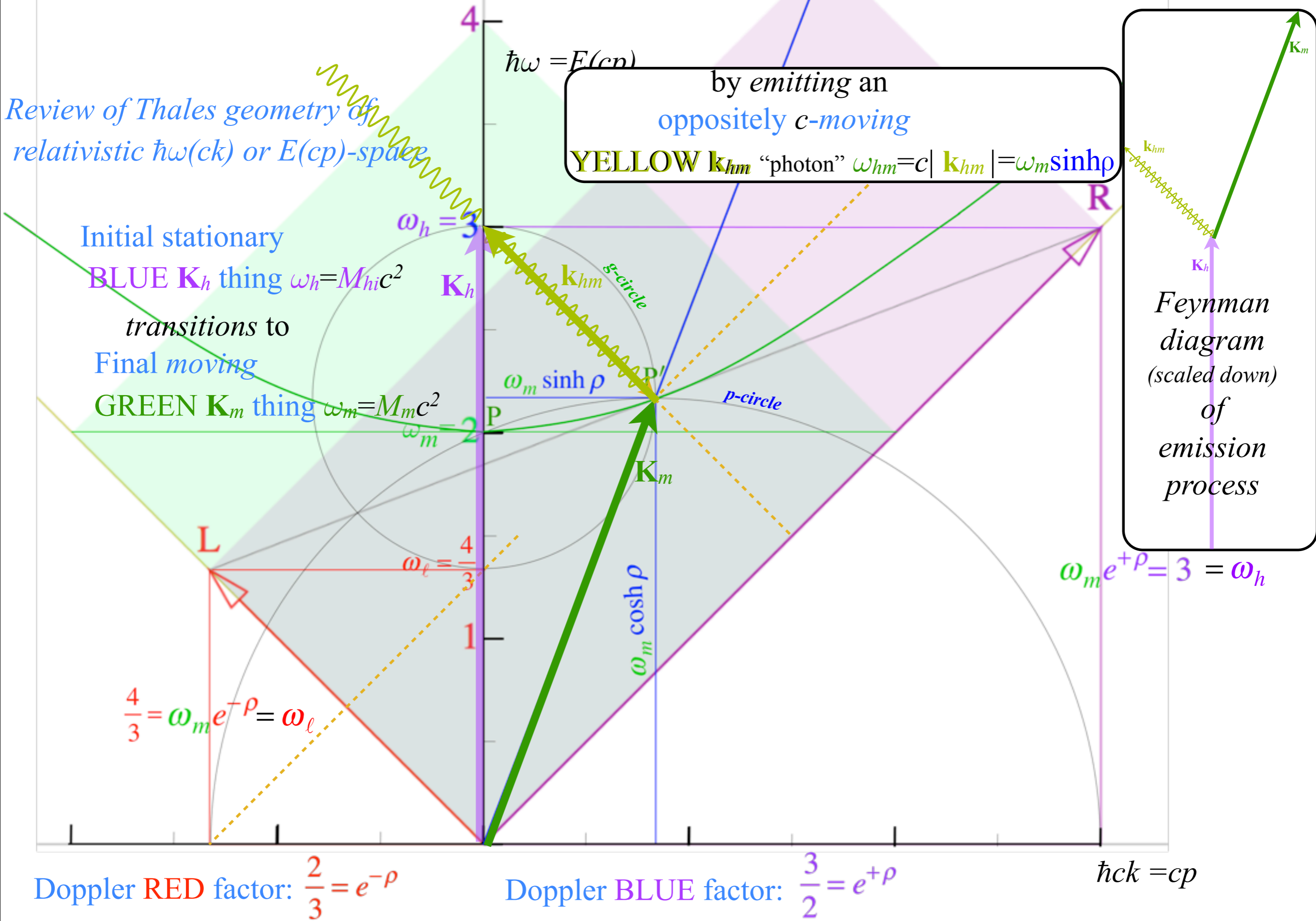
Relativistic optical transitions $|high\rangle = |\omega_h\rangle \rightleftharpoons |mid\rangle = |\omega_m\rangle \rightleftharpoons |low\rangle = |\omega_l\rangle$

Review of Thales geometry of relativistic $\hbar\omega(cp)$ or $E(cp)$ -space

Initial stationary BLUE K_h thing $\omega_h = M_h c^2$
 transitions to Final moving GREEN K_m thing $\omega_m = M_m c^2$



Relativistic optical transitions $|high\rangle = |\omega_h\rangle \rightleftharpoons |mid\rangle = |\omega_m\rangle \rightleftharpoons |low\rangle = |\omega_l\rangle$

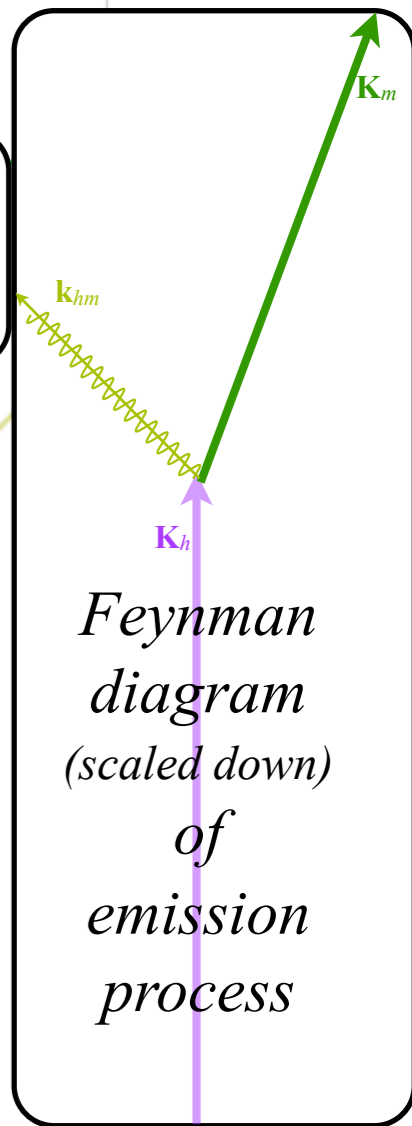


Relativistic optical transitions $|high\rangle = |\omega_h\rangle \rightleftharpoons |mid\rangle = |\omega_m\rangle \rightleftharpoons |low\rangle = |\omega_l\rangle$

Review of Thales geometry of relativistic $\hbar\omega(ck)$ or $E(cp)$ -space

Initial stationary BLUE K_h thing $\omega_h = M_h c^2$
 transitions to Final moving GREEN K_m thing $\omega_m = M_m c^2$

by emitting an oppositely c -moving YELLOW K_{hm} "photon" $\omega_{hm} = c |k_{hm}| = \omega_m \sinh \rho$



Take-away point 1
 Classical (and spectroscopic) Energy-momentum conservation is due to conservation quantum-phase space-time "wiggle-count"

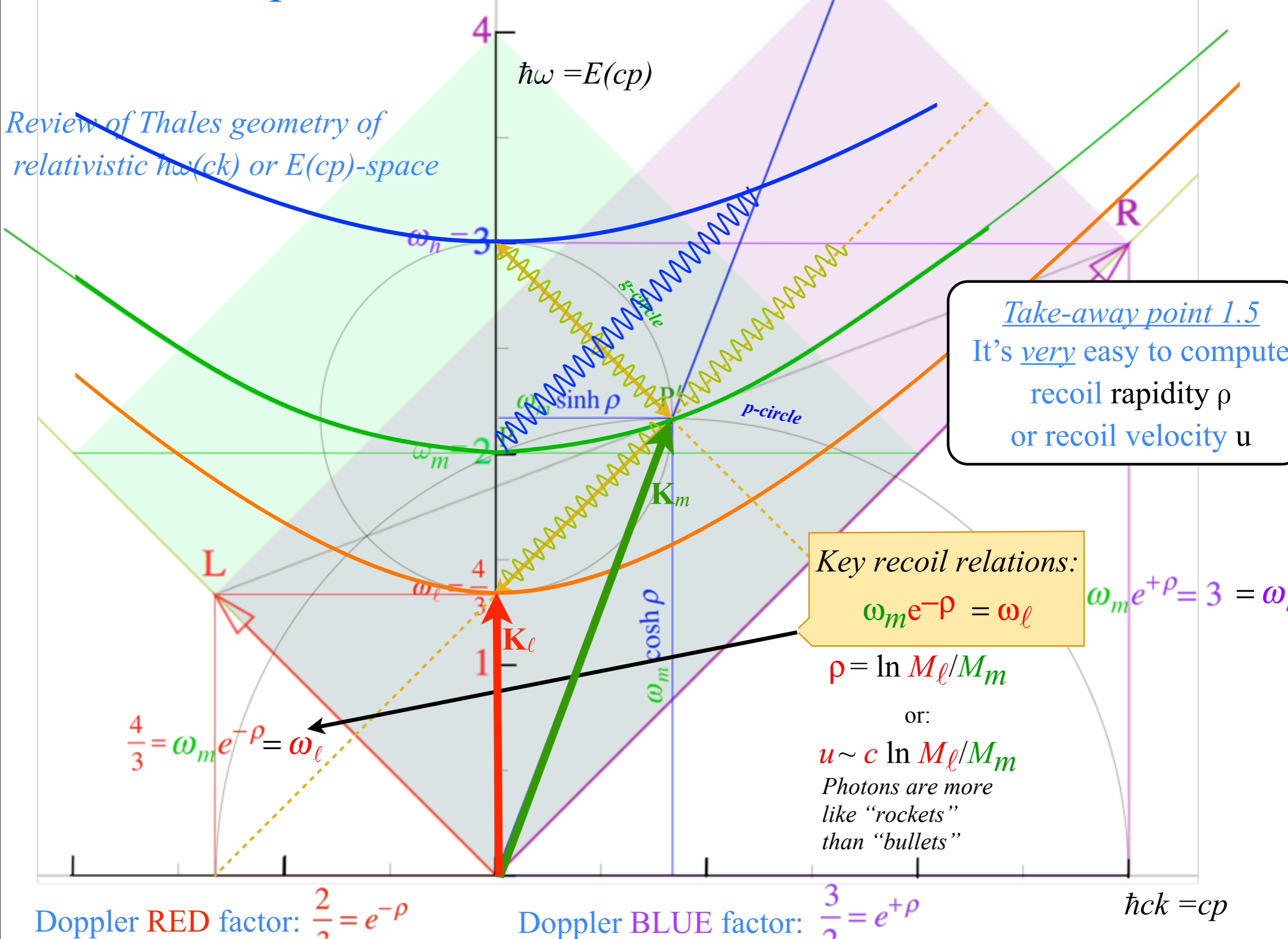
Doppler RED factor: $\frac{2}{3} = e^{-\rho}$

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Relativistic optical transitions $|high\rangle = |\omega_h\rangle \rightleftharpoons |mid\rangle = |\omega_m\rangle \rightleftharpoons |low\rangle = |\omega_l\rangle$

Review of Thales geometry of relativistic $\hbar\omega(c\mathbf{k})$ or $E(cp)$ -space



Take-away point 1.5
It's very easy to compute
recoil rapidity ρ
or recoil velocity u

Key recoil relations:
 $\omega_m e^{-\rho} = \omega_l$

$\omega_m e^{+\rho} = 3 = \omega_h$

$\rho = \ln M_l / M_m$

or:

$u \sim c \ln M_l / M_m$

Photons are more like "rockets" than "bullets"

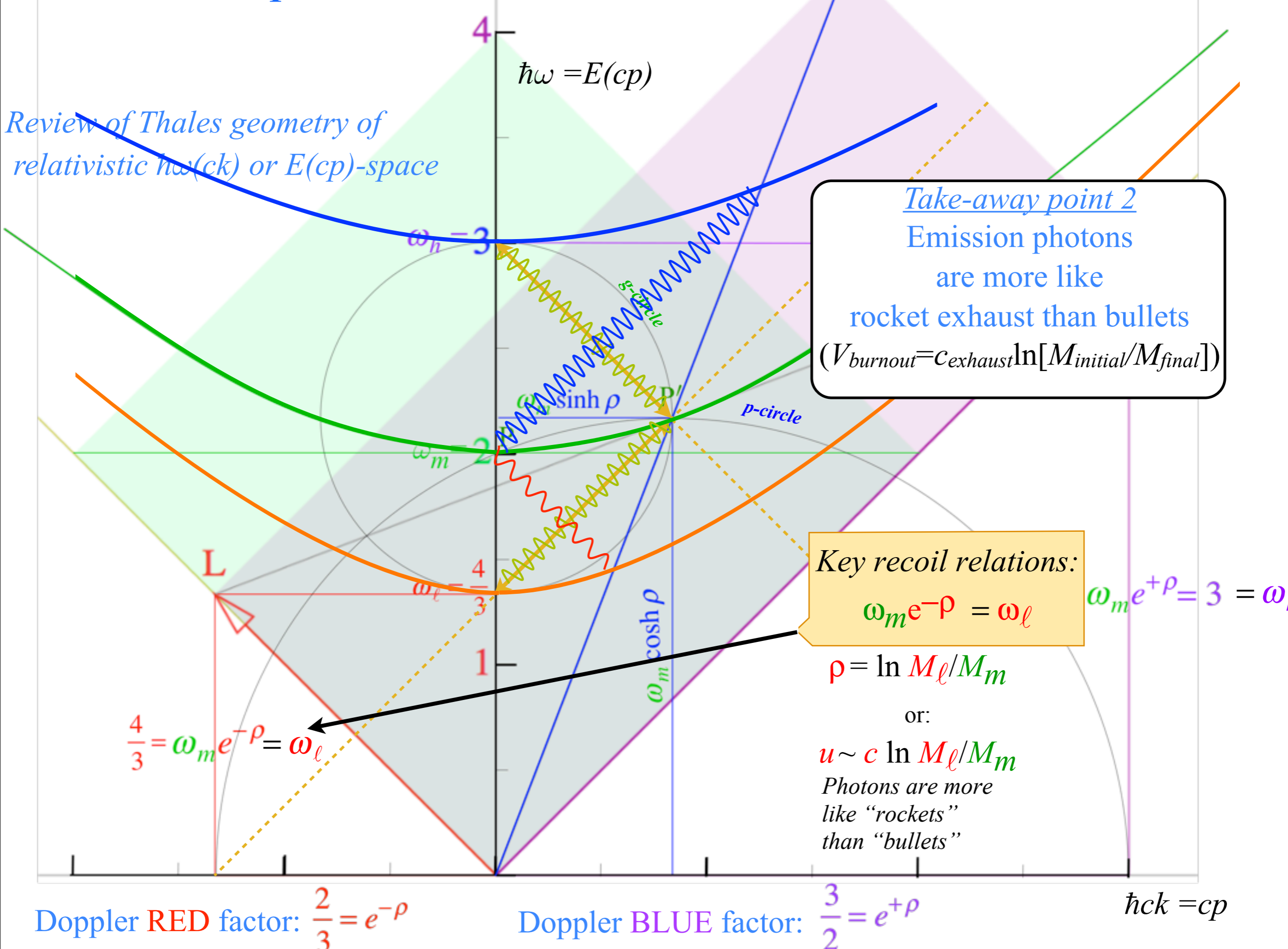
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$\hbar ck = cp$

Relativistic optical transitions $|high\rangle = |\omega_h\rangle \rightleftharpoons |mid\rangle = |\omega_m\rangle \rightleftharpoons |low\rangle = |\omega_l\rangle$

Review of Thales geometry of relativistic $\hbar\omega(c k)$ or $E(cp)$ -space



Take-away point 2
Emission photons are more like rocket exhaust than bullets
($V_{burnout} = c_{exhaust} \ln[M_{initial}/M_{final}]$)

Key recoil relations:
 $\omega_m e^{-\rho} = \omega_l$

$\omega_m e^{+\rho} = 3 = \omega_h$

$\rho = \ln M_\ell / M_m$

or:

$u \sim c \ln M_\ell / M_m$

Photons are more like "rockets" than "bullets"

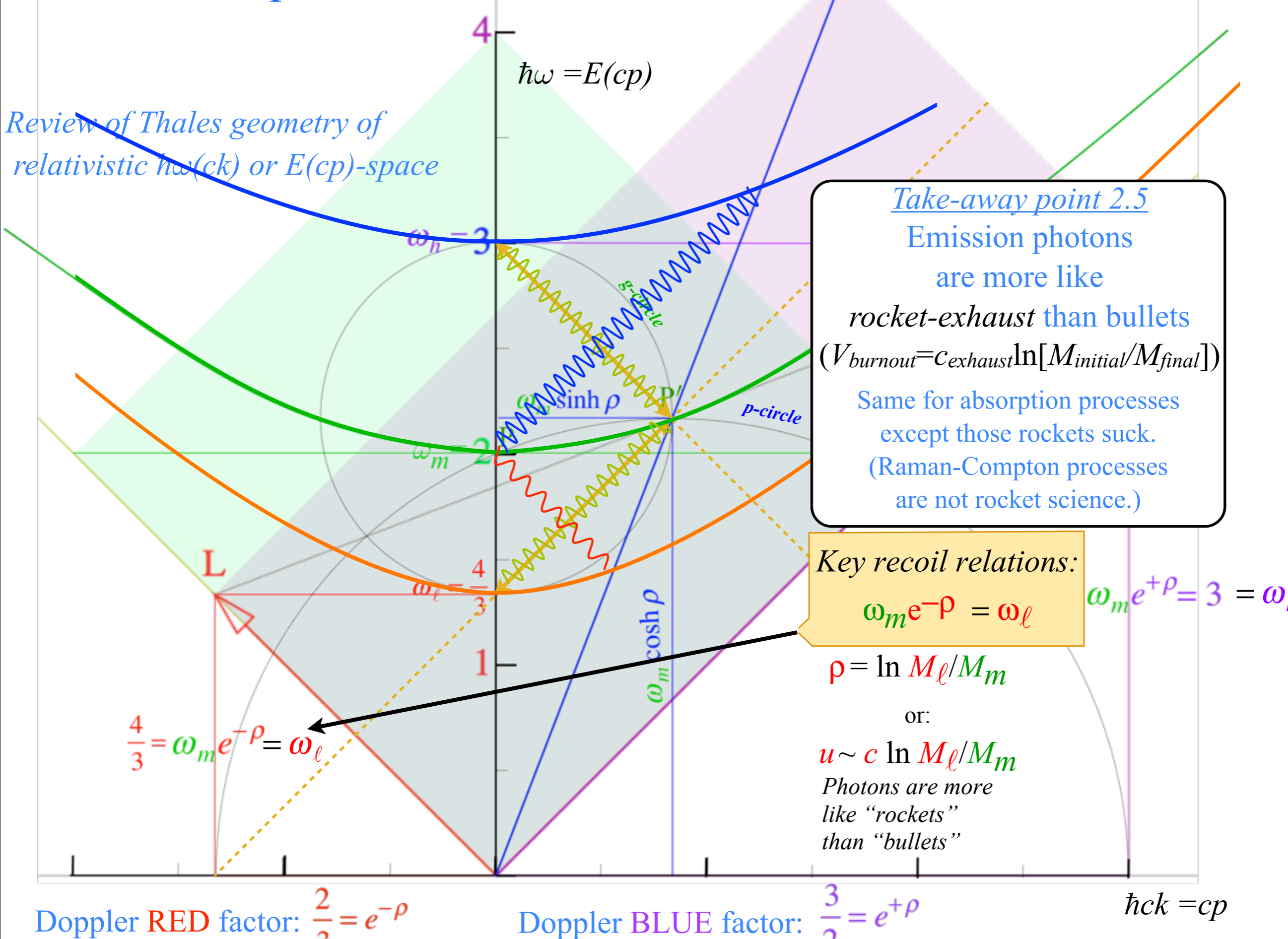
Doppler RED factor: $\frac{2}{3} = e^{-\rho}$

Doppler BLUE factor: $\frac{3}{2} = e^{+\rho}$

$\hbar ck = cp$

Relativistic optical transitions $|high\rangle = |\omega_h\rangle \rightleftharpoons |mid\rangle = |\omega_m\rangle \rightleftharpoons |low\rangle = |\omega_l\rangle$

Review of Thales geometry of relativistic $\hbar\omega(ck)$ or $E(cp)$ -space



Take-away point 2.5
 Emission photons are more like rocket-exhaust than bullets
 ($V_{burnout} = c_{exhaust} \ln[M_{initial}/M_{final}]$)
 Same for absorption processes except those rockets suck.
 (Raman-Compton processes are not rocket science.)

Key recoil relations:
 $\omega_m e^{-\rho} = \omega_l$

$\omega_m e^{+\rho} = 3 = \omega_h$

$\rho = \ln M_l/M_m$

or:
 $u \sim c \ln M_l/M_m$

Photons are more like "rockets" than "bullets"

Doppler RED factor: $\frac{2}{3} = e^{-\rho}$

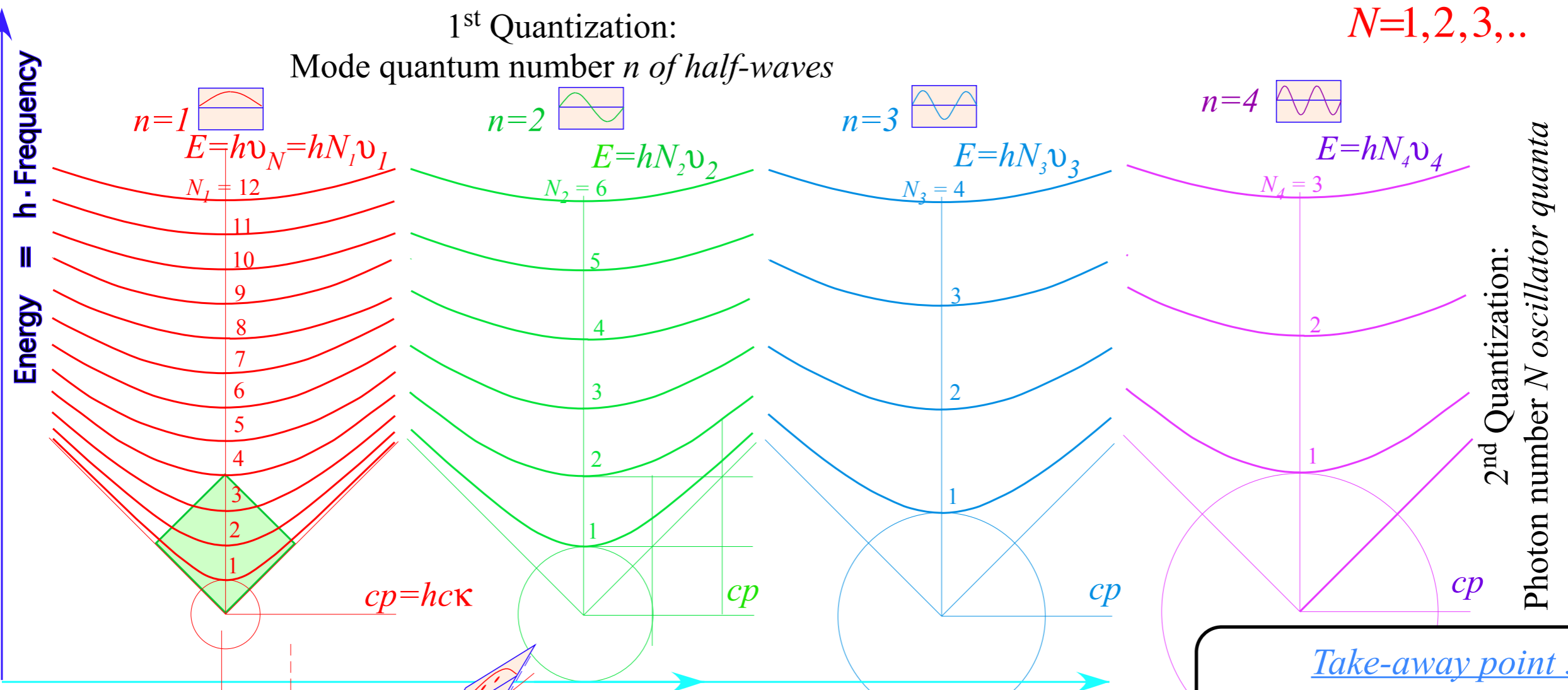
Doppler BLUE factor: $\frac{3}{2} = e^{+\rho}$

$\hbar ck = cp$

2nd Quantization: NEWS FLASH!!! $h\nu$ is actually $hN\nu$

($h\nu_{phase} = E = h\nu_A \cosh \rho$) is actually ($hN\nu_{phase} = E_N = hN\nu_A \cosh \rho$ with quantum numbers)

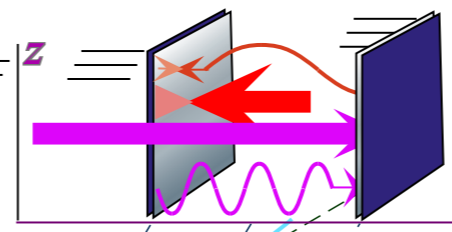
$N=1,2,3,..$



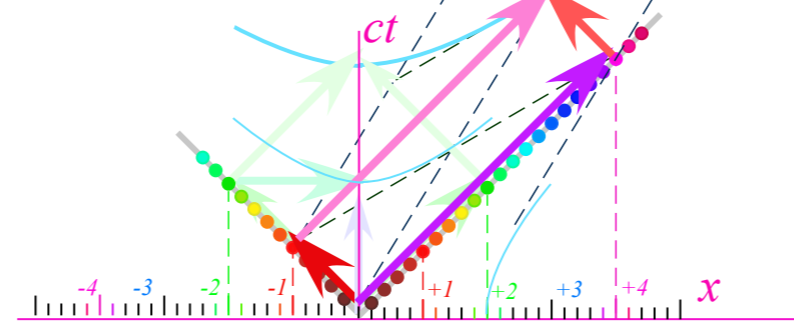
$c \cdot \text{Momentum} = hc \cdot \text{Wavenumber}$

Boosted wave mode

Boosted cavity wave has invariant mode number n photon number N_n



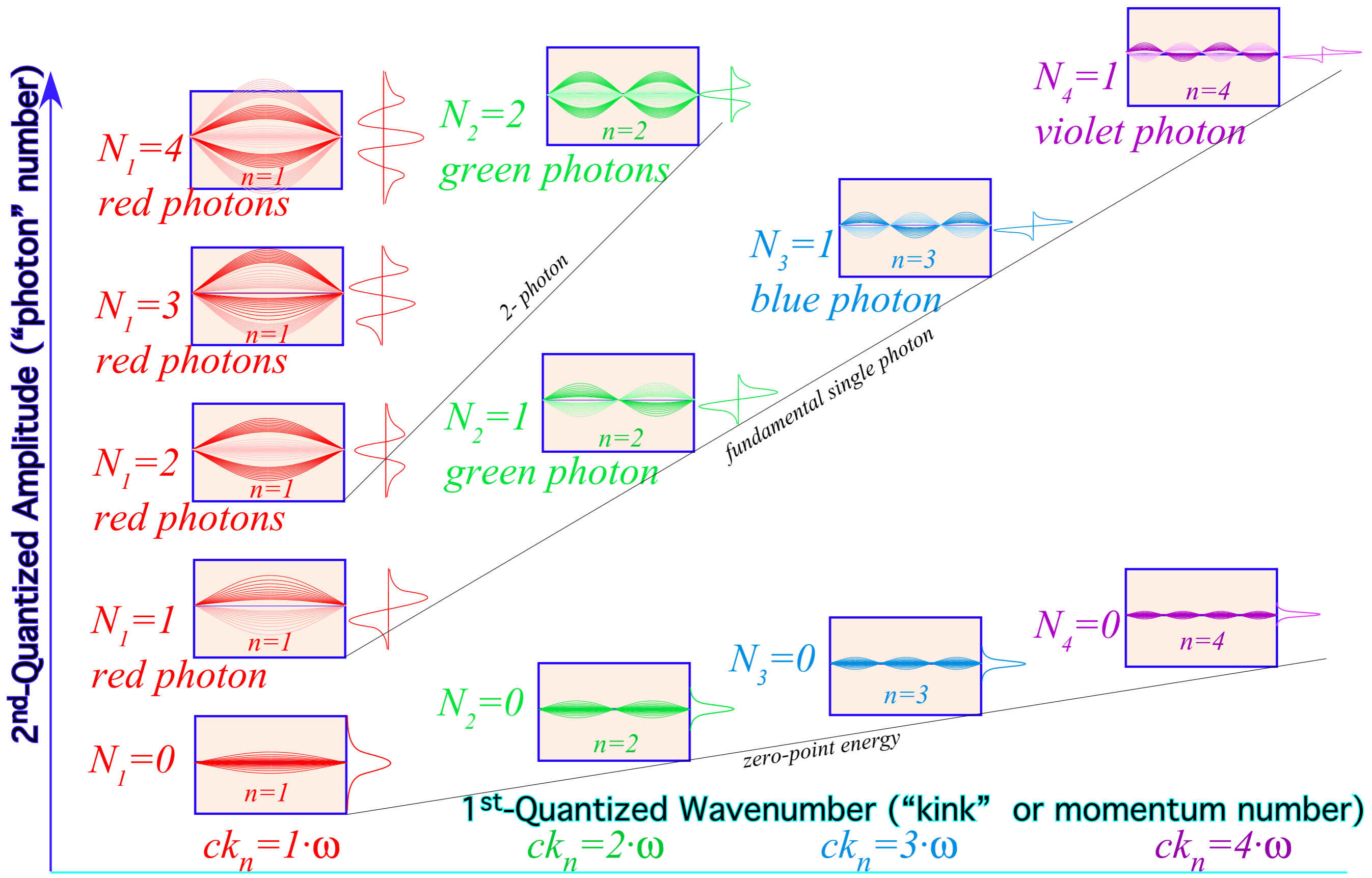
Lorentz contracted cavity length $L=3.2$
Proper length $l=4.0$



Take-away point 3
Cavity quantum electrodynamics (CQED) and spectra are analogous to molecular rovibronic dynamics with rotation-vibration algebra replaced by Lorentz-Poincare-Dirac algebra (and geometry!)

2nd Quantization: NEWS FLASH!!! $h\nu$ is actually $hN\nu$

$(h\nu_{phase}=E=h\nu_A \cosh \rho)$ is actually $(hN\nu_{phase}=E_N=hN\nu_A \cosh \rho \quad (N=1,2,..))$



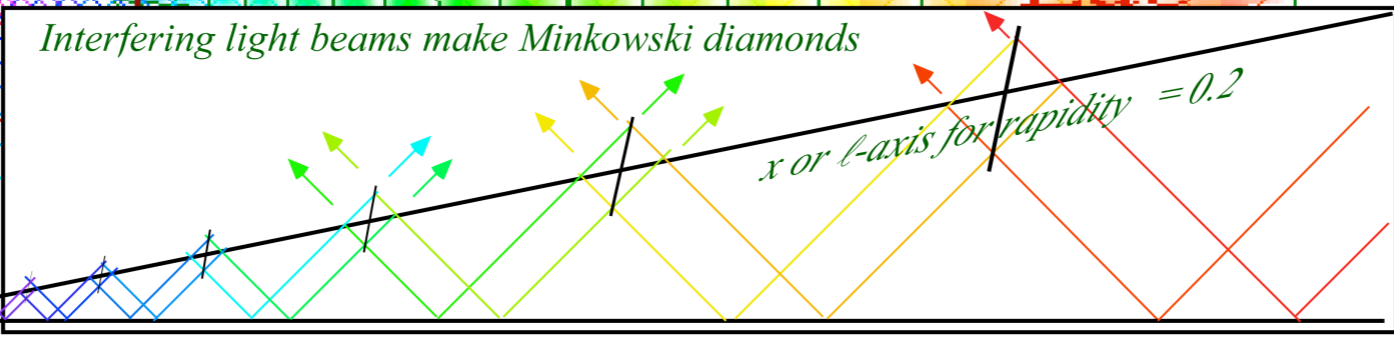
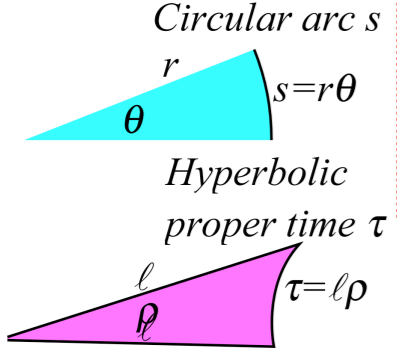
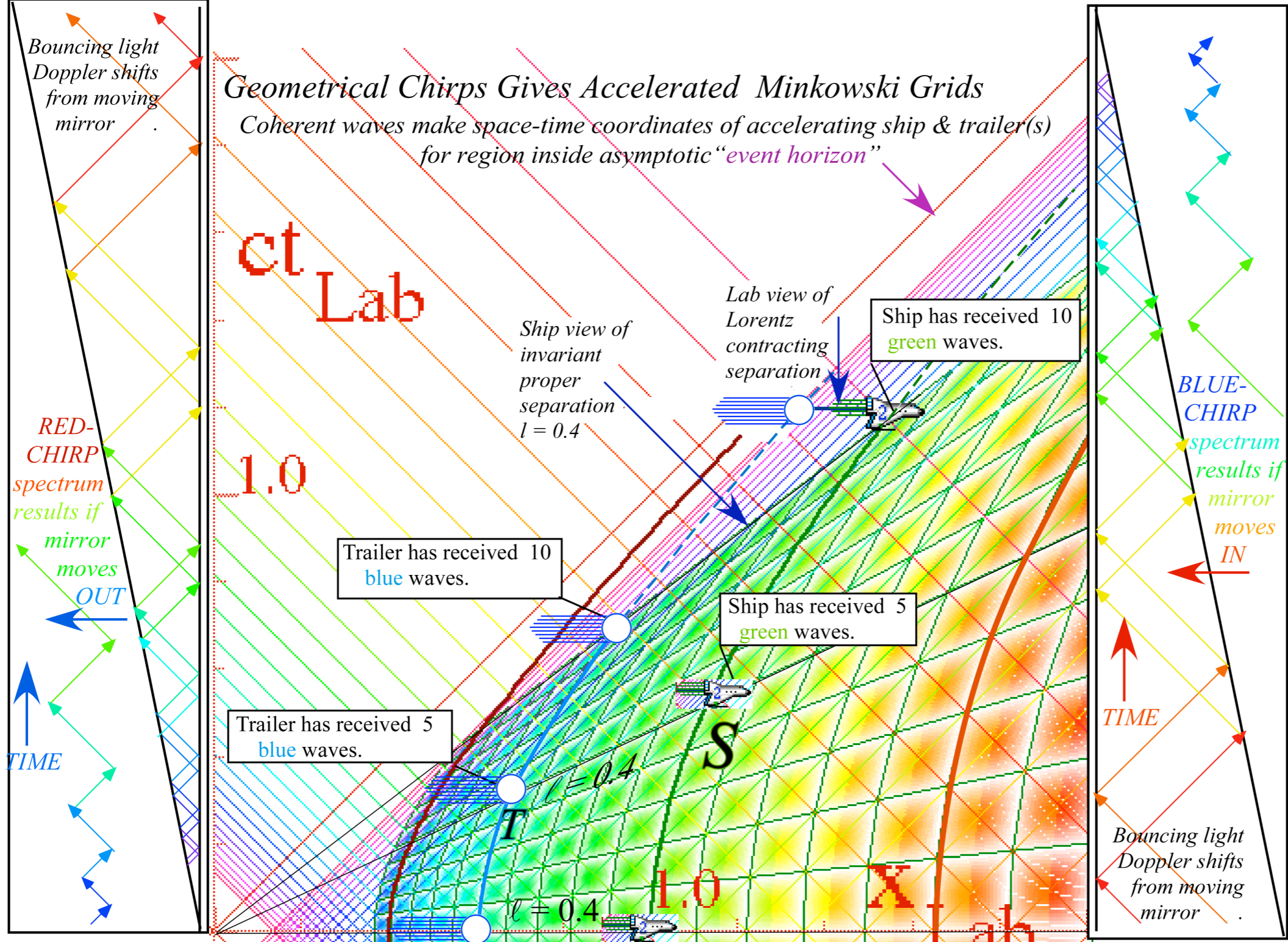


Fig. 8.2 Accelerated reference frames and their trajectories painted by chirped coherent light

Links to the current Harter-Soft Web Apps for

(**Bold** links have default redirect pages. *Italics* are not yet meant for production. **Red** are in the final stages of testing or

Production Links - *For the students & general public*

Textbooks & Lectures

[Classical Mechanics with a Bang!](http://www.uark.edu/ua/modphys/markup/CMwBangWeb.html) - URL is "<http://www.uark.edu/ua/modphys/markup/CMwBangWeb.html>"
[Group Theory in Quantum Mechanics {Lectures}](http://www.uark.edu/ua/modphys/markup/GTQMWeb.html) - URL is "<http://www.uark.edu/ua/modphys/markup/GTQMWeb.html>"
[Modern Physics and its Classical Foundations](http://www.uark.edu/ua/modphys/markup/MPCFWeb.html) - URL is "<http://www.uark.edu/ua/modphys/markup/MPCFWeb.html>"
[Principles of Symmetry, Dynamics, and Spectroscopy {Text}](http://www.uark.edu/ua/modphys/markup/PSDSWeb.html) - URL is "<http://www.uark.edu/ua/modphys/markup/PSDSWeb.html>"
[Quantum Theory for the Computer Age](http://www.uark.edu/ua/modphys/markup/QTCAWeb.html) - URL is "<http://www.uark.edu/ua/modphys/markup/QTCAWeb.html>"

LearnIt Web Applications

[Production Portal Page](http://www.uark.edu/ua/modphys/markup/LearnItWeb.html); URL is "<http://www.uark.edu/ua/modphys/markup/LearnItWeb.html>"

[BohrIt](http://www.uark.edu/ua/modphys/markup/BohrItWeb.html) - Production; URL is "<http://www.uark.edu/ua/modphys/markup/BohrItWeb.html>"
[BounceIt](http://www.uark.edu/ua/modphys/markup/BounceItWeb.html) - Production; URL is "<http://www.uark.edu/ua/modphys/markup/BounceItWeb.html>"
[BoxIt](http://www.uark.edu/ua/modphys/markup/BoxItWeb.html) - Production; URL is "<http://www.uark.edu/ua/modphys/markup/BoxItWeb.html>"
[CoullIt](http://www.uark.edu/ua/modphys/markup/CoullItWeb.html) - Production; URL is "<http://www.uark.edu/ua/modphys/markup/CoullItWeb.html>"
[Cycloidulum](http://www.uark.edu/ua/modphys/markup/CycloidulumWeb.html) - Production; URL is "<http://www.uark.edu/ua/modphys/markup/CycloidulumWeb.html>"
[JerkIt](http://www.uark.edu/ua/modphys/markup/JerkItWeb.html) - Production; URL is "<http://www.uark.edu/ua/modphys/markup/JerkItWeb.html>"
[MoIVibes](http://www.uark.edu/ua/modphys/markup/MoIVibesWeb.html) - Production; URL is "<http://www.uark.edu/ua/modphys/markup/MoIVibesWeb.html>"
[Pendulum](http://www.uark.edu/ua/modphys/markup/PendulumWeb.html) - Production; URL is "<http://www.uark.edu/ua/modphys/markup/PendulumWeb.html>"
[QuantIt](http://www.uark.edu/ua/modphys/markup/QuantItWeb.html) - Production; URL is "<http://www.uark.edu/ua/modphys/markup/QuantItWeb.html>"
[Relativity - 2005 Pirelli Challenge Entrant](http://www.uark.edu/ua/pirelli/html/default.html) - Production; URL is "<http://www.uark.edu/ua/pirelli/html/default.html>" or "<http://www.uark.edu/ua/pirelli/html/default.html>"
[RelativIt](http://www.uark.edu/ua/modphys/markup/RelativItWeb.html) Production; URL is "<http://www.uark.edu/ua/modphys/markup/RelativItWeb.html>"
[RelaWavity](http://www.uark.edu/ua/modphys/markup/RelaWavityWeb.html) Production; URL is "<http://www.uark.edu/ua/modphys/markup/RelaWavityWeb.html>"
[Trebuchet](http://www.uark.edu/ua/modphys/markup/TrebuchetWeb.html) Production; URL is "<http://www.uark.edu/ua/modphys/markup/TrebuchetWeb.html>"
[WaveIt](http://www.uark.edu/ua/modphys/markup/WaveItWeb.html) Production; URL is "<http://www.uark.edu/ua/modphys/markup/WaveItWeb.html>"

$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
$b_{BLUE}^{-1Doppler}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$b_{RED}^{-1Doppler}$
$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\operatorname{coth} \rho$	$e^{+\rho}$
	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	
0.5	0.6	0.75	0.80	1.25	1.33	1.67	2.0