Lecture 10 Revised 12.22.12 from 9.20.2012

Quadratic form geometry and development of mechanics of Lagrange and Hamilton

(Ch. 12 of Unit 1 and Ch. 4-5 of Unit 7)

Scaling transformation between Lagrangian and Hamiltonian views of KE (Review of Lecture 9) Introducing 1st Lagrange and Hamilton differential equations of mechanics (Review Of Lecture 9)

Introducing the Poincare' and Legendre contact transformations

Geometry of Legendre contact transformation

Example from thermodynamics

Legendre transform: special case of General Contact Transformation (lights, camera, ACTION!)

A general contact transformation from sophomore physics

Scaling transformation between Lagrangian and Hamiltonian views of KE (Review of Lecture 9)
Introducing the (partial) differential equations of mechanics (Review Of Lecture 9)

1st equations of Lagrange and Hamilton

Introducing the (partial ³?) differential equations of mechanics

Starts out with simple demands for explicit-dependence, "loyalty" or "fealty to the colors"

Lagrangian and Estrangian have <u>no</u> explicit dependence on momentum p

$$\frac{\partial L}{\partial p_k} \equiv 0 \equiv \frac{\partial E}{\partial p_k}$$

Hamiltonian and Estrangian have <u>no</u> explicit dependence on velocity v

$$\frac{\partial \mathbf{H}}{\partial \mathbf{v}_{k}} \equiv 0 \equiv \frac{\partial \mathbf{E}}{\partial \mathbf{v}_{k}}$$

Lagrangian and Hamiltonian have <u>no</u> explicit dependence on speedinum V

$$\frac{\partial L}{\partial V_k} \equiv 0 \equiv \frac{\partial H}{\partial V_k}$$

Such non-dependencies hold in spite of "under-the-table" matrix and partial-differential connections

$$\nabla_{v} L = \frac{\partial L}{\partial \mathbf{v}} = \frac{\partial}{\partial \mathbf{v}} \frac{\mathbf{v} \cdot \mathbf{M} \cdot \mathbf{v}}{2}$$
$$= \mathbf{M} \cdot \mathbf{v} = \mathbf{p}$$

$$\begin{pmatrix} \frac{\partial L}{\partial v_1} \\ \frac{\partial L}{\partial v_2} \end{pmatrix} = \begin{pmatrix} m_1 & 0 \\ 0 & m_2 \end{pmatrix} \begin{pmatrix} v_1 \\ v_2 \end{pmatrix} = \begin{pmatrix} p_1 \\ p_2 \end{pmatrix}$$

Lagrange's 1st equation(s)

$$\frac{\partial L}{\partial v_k} = p_k$$
 or: $\frac{\partial L}{\partial \mathbf{v}} = \mathbf{p}$

$$\nabla_{p} H = \mathbf{v} = \frac{\partial H}{\partial \mathbf{p}} = \frac{\partial}{\partial \mathbf{p}} \frac{\mathbf{p} \cdot \mathbf{M}^{-1} \cdot \mathbf{p}}{2}$$
$$= \mathbf{M}^{-1} \cdot \mathbf{p} = \mathbf{v}$$

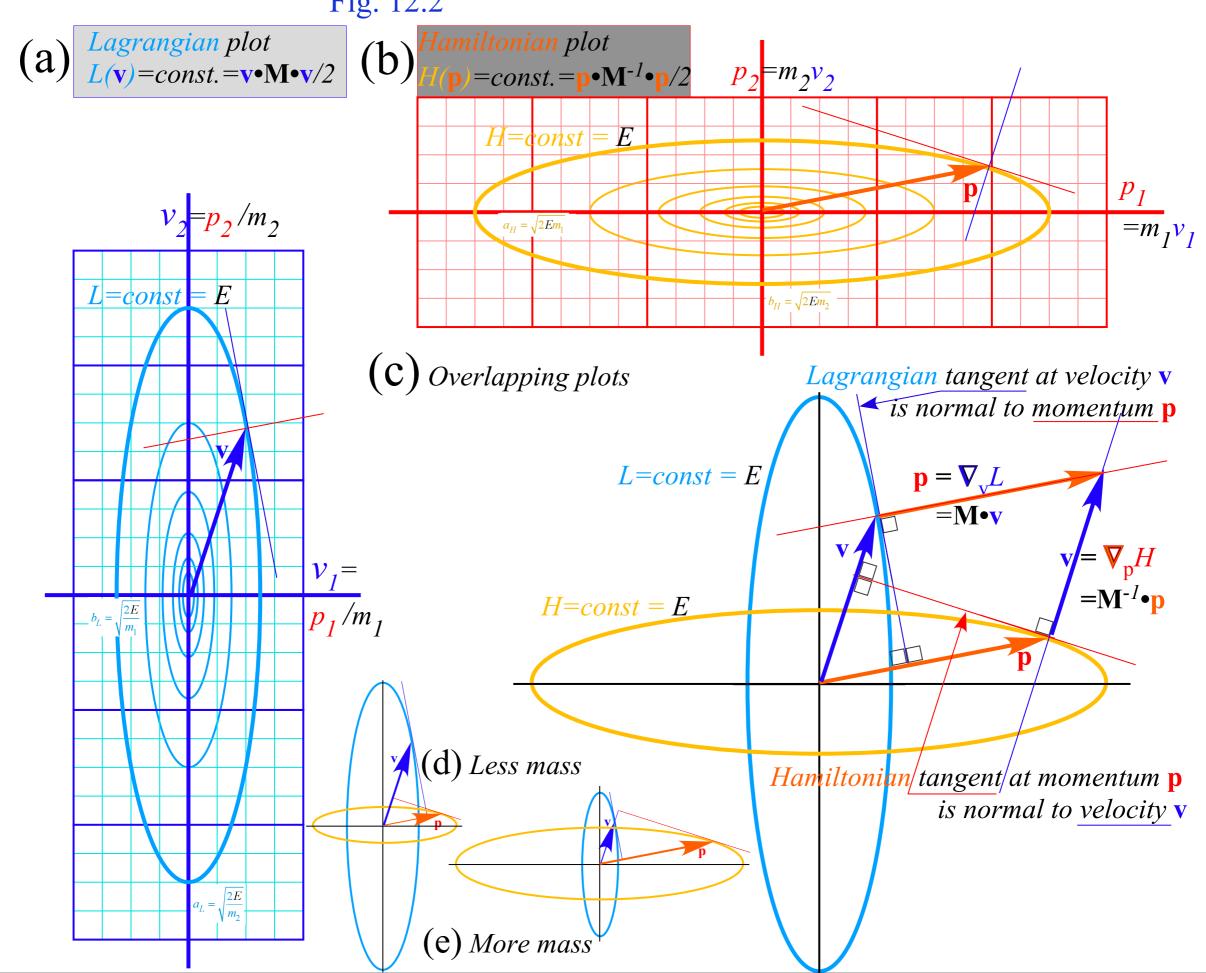
$$\begin{pmatrix} \frac{\partial \mathbf{H}}{\partial \mathbf{p}_1} \\ \frac{\partial \mathbf{H}}{\partial \mathbf{p}_2} \end{pmatrix} = \begin{pmatrix} m_1^{-1} & 0 \\ 0 & m_2^{-1} \end{pmatrix} \begin{pmatrix} \mathbf{p}_1 \\ \mathbf{p}_2 \end{pmatrix} = \begin{pmatrix} \mathbf{v}_1 \\ \mathbf{v}_2 \end{pmatrix}$$

Hamilton's 1st equation(s)

$$\frac{\partial H}{\partial p_k} = v_k \quad \text{or:} \quad \frac{\partial H}{\partial \mathbf{p}} = \mathbf{v}$$

(Forget Estrangian for now)

Unit 1 Fig. 12.2





Geometry of Legendre contact transformation

Example from thermodynamics

Legendre transform: special case of General Contact Transformation (lights,camera, ACTION!)

Saturday, December 22, 2012 5

Given matrix relation: $\mathbf{p} = \mathbf{M} \cdot \mathbf{v}$ or its inverse: $\mathbf{v} = \mathbf{M}^{-1} \cdot \mathbf{p}$ you might be tempted to rewrite

Q-forms $L(\mathbf{v}..)=(1/2)\mathbf{v}\cdot\mathbf{M}\cdot\mathbf{v}$ or $H(\mathbf{p}..)=(1/2)\mathbf{p}\cdot\mathbf{M}^{-1}\cdot\mathbf{p}$ to be $H=(1/2)\mathbf{p}\cdot\mathbf{v}$ or equivalently $L=(1/2)\mathbf{v}\cdot\mathbf{p}$.

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Numerically-CORRECT, but Differentially-WRONG!

Given matrix relation: $\mathbf{p} = \mathbf{M} \cdot \mathbf{v}$ or its inverse: $\mathbf{v} = \mathbf{M}^{-1} \cdot \mathbf{p}$ you might be tempted to rewrite

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Numerically-CORRECT, but Differentially-WRONG!

Instead try: $H(\mathbf{p}..) = \mathbf{p} \cdot \mathbf{v} - (1/2)\mathbf{v} \cdot \mathbf{p} = \mathbf{p} \cdot \mathbf{v} - L(\mathbf{v}..)$ or else: $L(\mathbf{v}..) = \mathbf{p} \cdot \mathbf{v} - H(\mathbf{p}..)$

Given matrix relation: $\mathbf{p} = \mathbf{M} \cdot \mathbf{v}$ or its inverse: $\mathbf{v} = \mathbf{M}^{-1} \cdot \mathbf{p}$ you might be tempted to rewrite

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Legendre contact transformation

$$L(\mathbf{v}) = \mathbf{p} \cdot \mathbf{v} - H(\mathbf{p})$$

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Now explicit dependency (non)-relations give the right derivatives

$$\frac{\partial L(\mathbf{v})}{\partial \mathbf{p}} = \frac{\partial}{\partial \mathbf{p}} \mathbf{p} \cdot \mathbf{v} - \frac{\partial H(\mathbf{p})}{\partial \mathbf{p}}$$
$$0 = \mathbf{v} - \frac{\partial H(\mathbf{p})}{\partial \mathbf{p}}$$

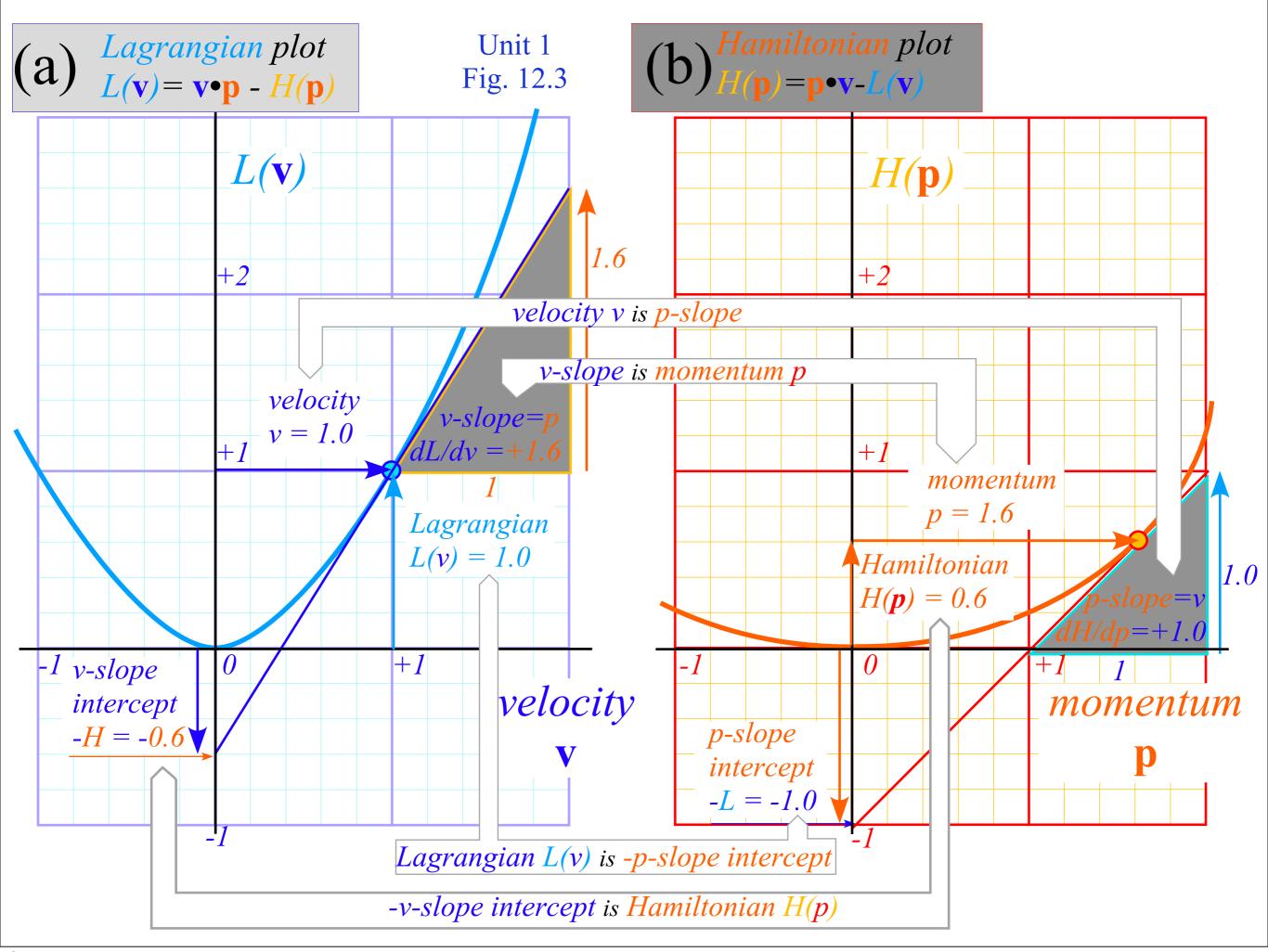
$$\frac{\partial H(\mathbf{p})}{\partial \mathbf{v}} = \frac{\partial}{\partial \mathbf{v}} \mathbf{p} \cdot \mathbf{v} - \frac{\partial L(\mathbf{v})}{\partial \mathbf{v}}$$
$$0 = \mathbf{p} - \frac{\partial L(\mathbf{v})}{\partial \mathbf{v}}$$

That is Hamilton's 1^{st} equation(s) and Lagrange's 1^{st} equation(s)

Geometry of Legendre contact transformation

Example from thermodynamics

Legendre transform: special case of General Contact Transformation (lights,camera, ACTION!)

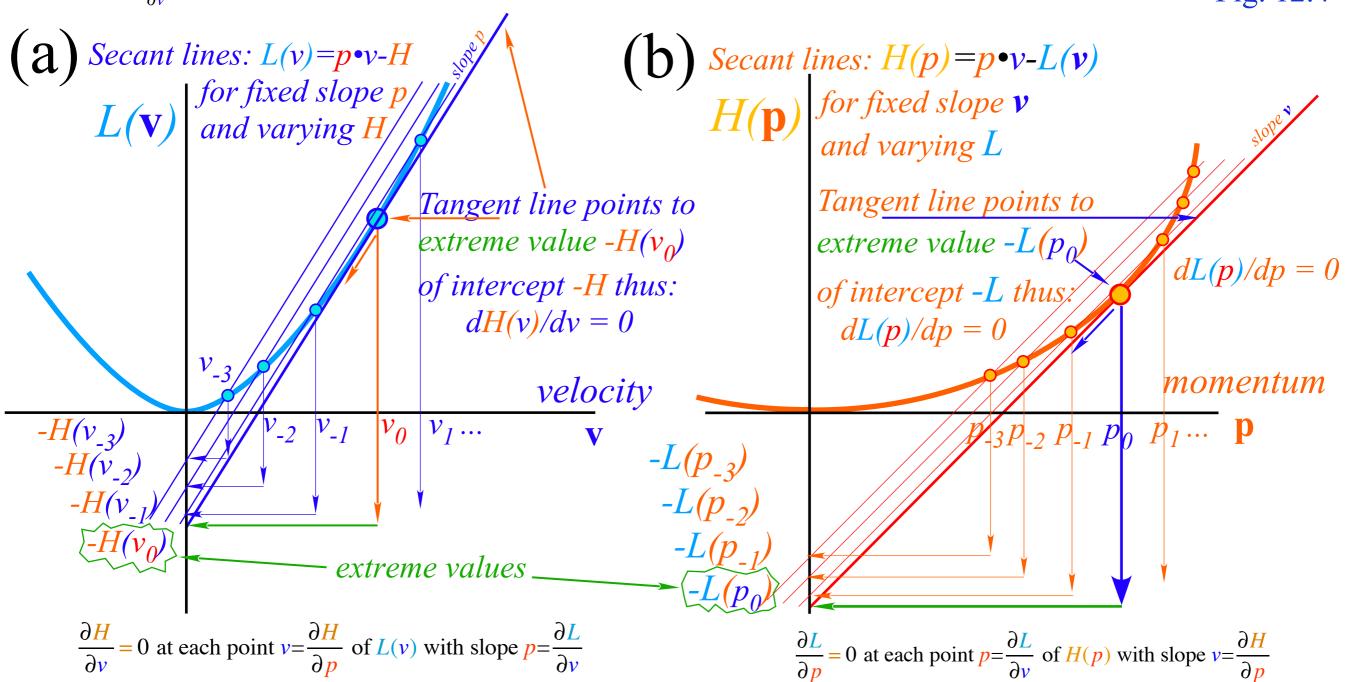


How Legendre contact transformations work...(to make $\frac{\partial H}{\partial v} = 0$ or $\frac{\partial L}{\partial p} = 0$)

Secant lines $L(\mathbf{v}) = p \cdot v - H$ of fixed slope $p = \frac{\partial L}{\partial v}$ and decreasing intercept $-H(v_{-2}) > -H(v_{-1}) > ...$ for increasing velocity $v_{-2} > v_{-1} > ... > v_0$ lead to unique tangent to $L(\mathbf{v})$ -curve at the tangent contact point $v = v_0$ that has $\max_{\mathbf{max}} H(p \ v_0)$ Thus $\frac{\partial H}{\partial v} = 0$

(Similarly...)

Unit 1 Fig. 12.4



Geometry of Legendre contact transformation

Example from thermodynamics

Legendre transform: special case of General Contact Transformation (lights,camera, ACTION!)

Example of Legendre contact transformation in thermodynamics

Internal energy U(S, V) is defined as a function of entropy S and volume V.

A new function *enthalpy* H(S,P) depends on entropy and *pressure* P.

It is a Legendre transform $H(S,P)=P\cdot V+U$ of energy U(S,V) to new variable $P=-(\frac{\partial U}{\partial V})_S$.

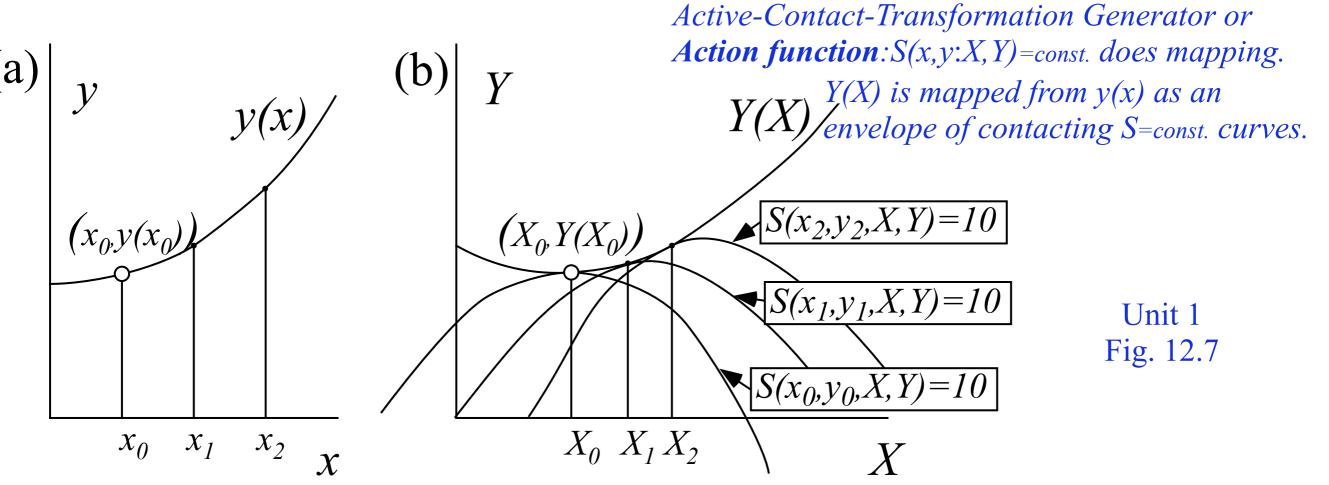
Except for \pm signs, it's our Hamiltonian $H(p) = p \cdot v \cdot L(v)$ going from Lagrangian L(v)

to use new variable momentum $p = (\frac{\partial L}{\partial v})_x$.

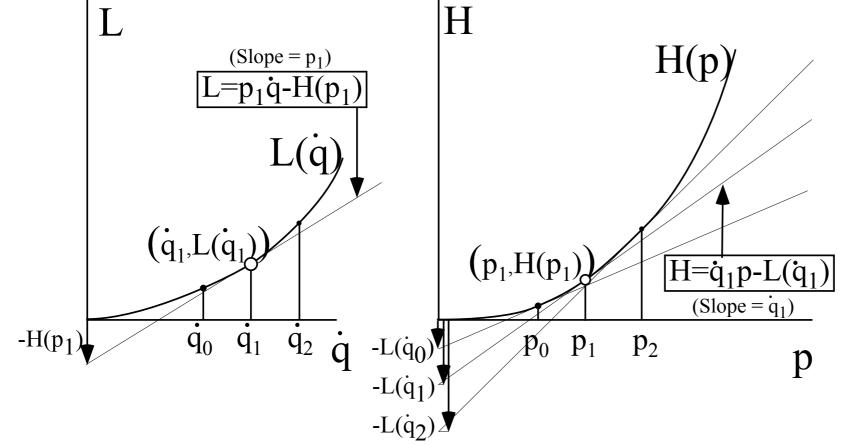
Geometry of Legendre contact transformation Example from thermodynamics

Legendre transform: special case of General Contact Transformation (lights, camera, ACTION!)

Legendre transform: special case of General Contact Transformation

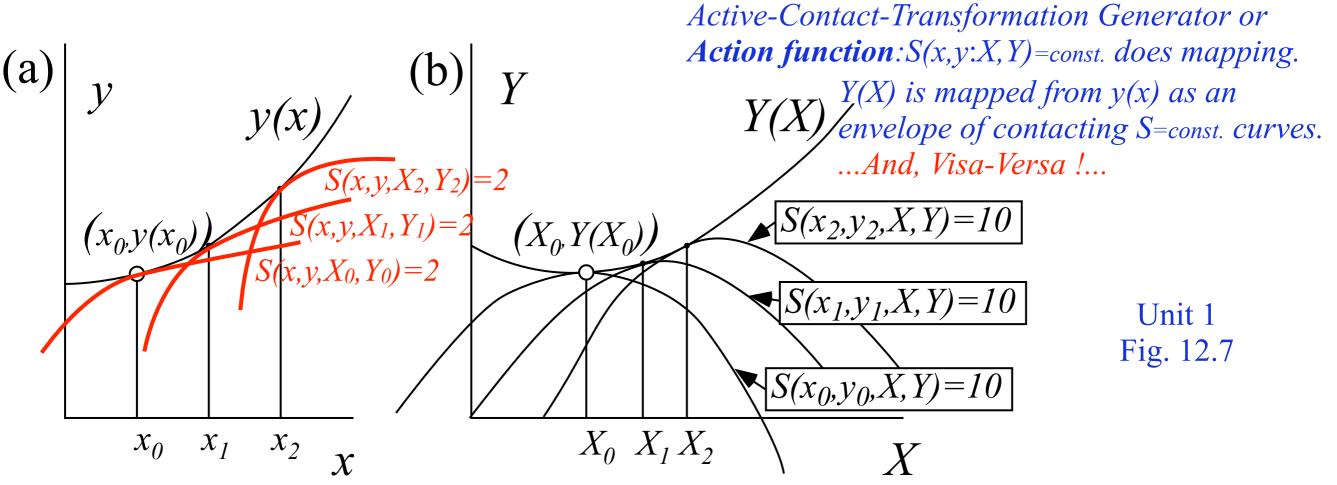


The Legendre transformation does it with contacting straight line tangents.

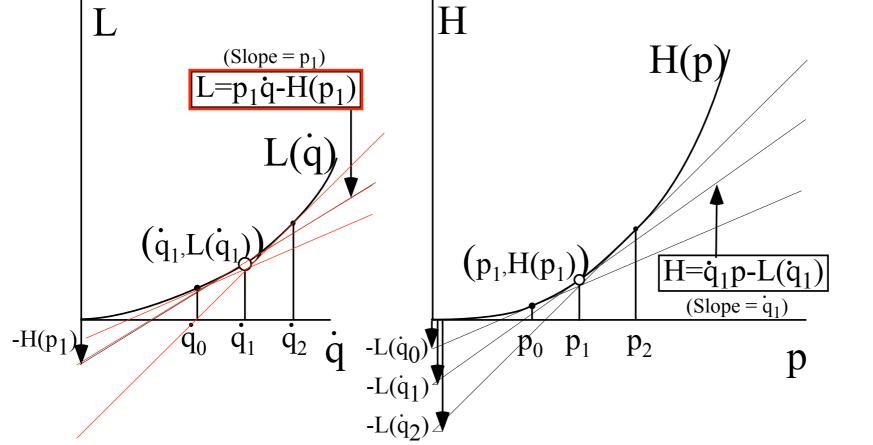


Unit 1 Fig. 12.9

Legendre transform: special case of General Contact Transformation

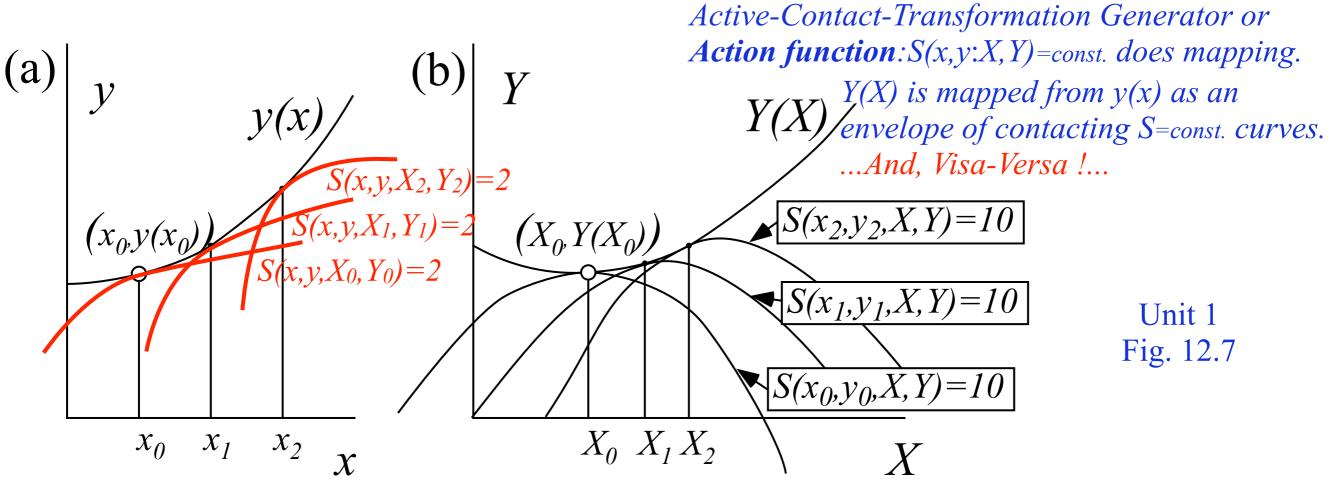


The Legendre transformation does it with contacting straight line tangents.

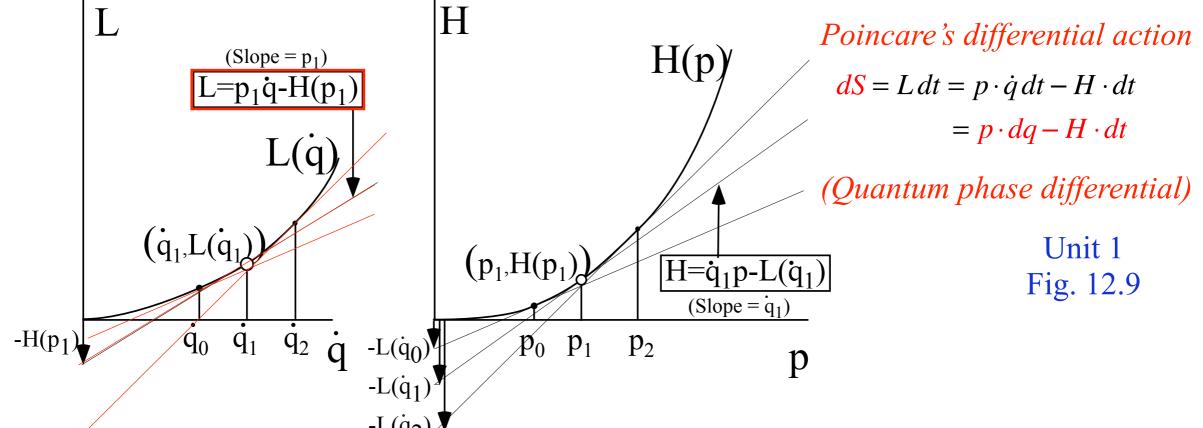


Unit 1 Fig. 12.9

Legendre transform: special case of General Contact Transformation

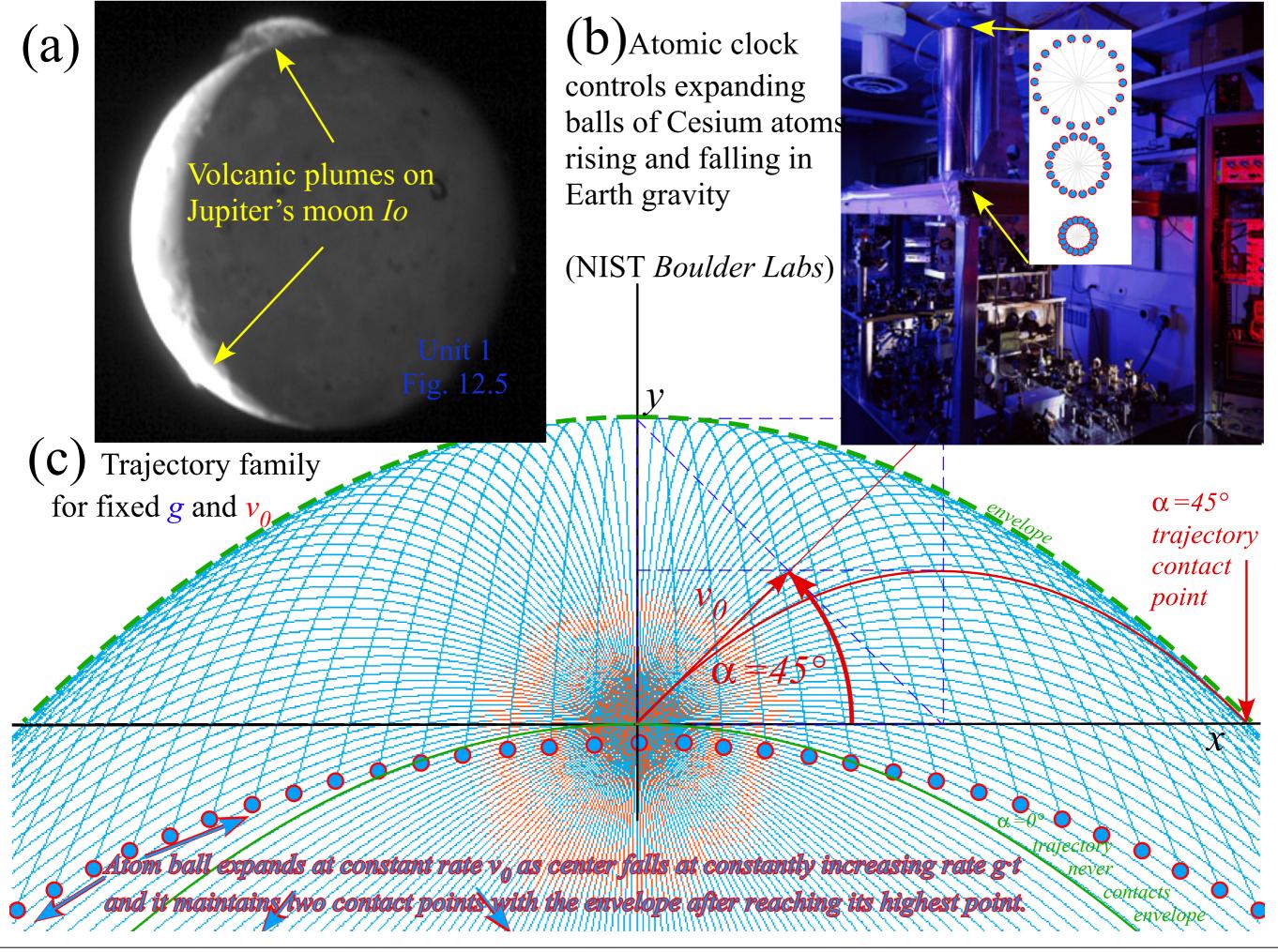


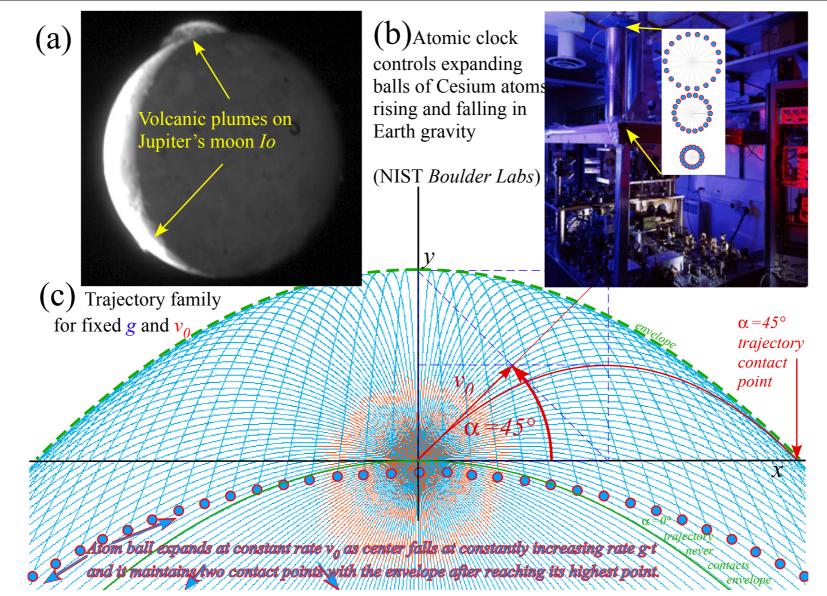
The Legendre transformation does it with contacting straight line tangents.



A general contact transformation from sophomore physics

Algebra-calculus development of "The Volcanoes of Io" and "The Atoms of NIST" Intuitive-geometric development of """ and """ """





Unit 1 Fig. 12.5

UP-1 formulas for trajectories in constant gravity g

$$x(t) = (v_0 \cos \alpha)t$$

$$y(t) = (v_0 \sin \alpha)t - \frac{1}{2}gt^2$$

$$\dot{x}(0) = v_x(0) = v_0 \cos \alpha$$

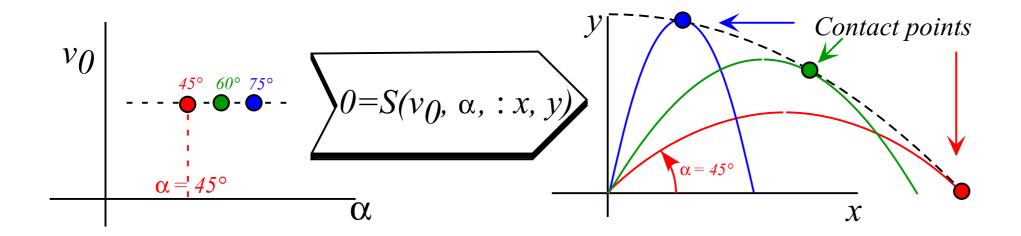
$$\dot{y}(0) = v_y(0) = v_0 \sin \alpha$$

Substitute time $t=x/(v_0 \cos \alpha)$ into y(t)

$$y(x) = \frac{v_0 \sin \alpha}{v_0 \cos \alpha} x - \frac{gx^2}{2v_0^2 \cos^2 \alpha}$$

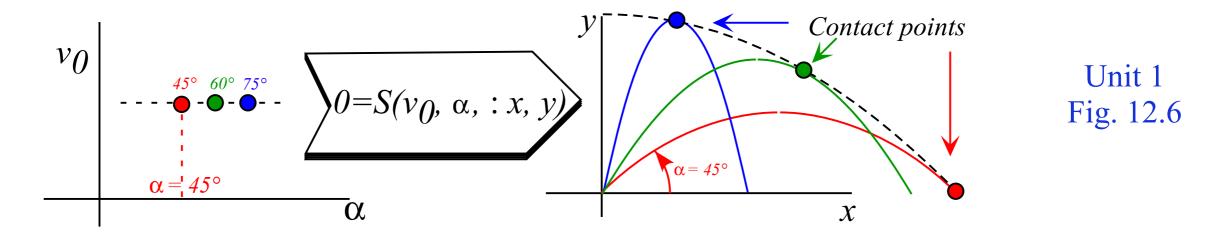
$$y(x) = x \tan \alpha - \frac{gx^2}{2v_0^2 \cos^2 \alpha}$$

$$y(x) = x \tan \alpha - \frac{gx^2}{2v_0^2 \cos^2 \alpha}$$
 becomes:
$$S(v_0, \alpha : x, y) = -y + x \tan \alpha - \frac{gx^2}{2v_0^2 \cos^2 \alpha} = 0$$



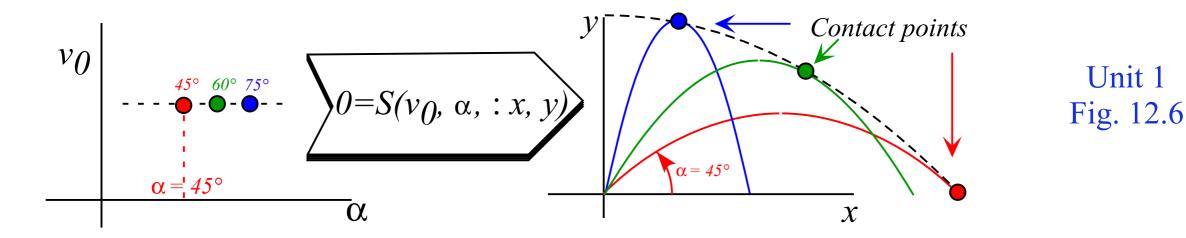
Unit 1 Fig. 12.6

$$y(x) = x \tan \alpha - \frac{gx^2}{2v_0^2 \cos^2 \alpha}$$
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Envelopes of the v_0 -trajectory region contain extremal contact points with each trajectory where: $\frac{\partial S(v_0,\alpha:x,y)}{\partial \alpha} = 0$

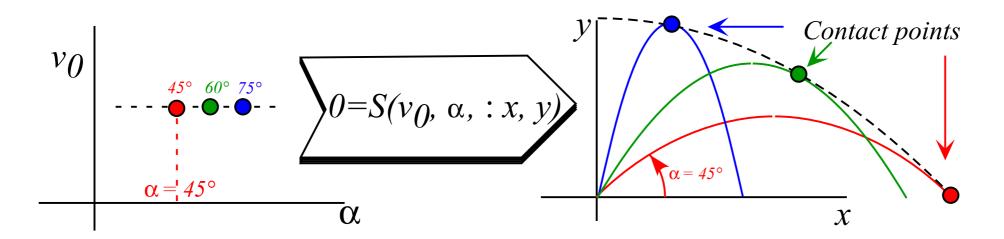
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Envelopes of the v_0 -trajectory region contain extremal *contact points* with each trajectory

where:
$$\frac{\partial S(v_0, \alpha; x, y)}{\partial \alpha} = 0$$
$$x \frac{\partial \tan \alpha}{\partial \alpha} - \frac{gx^2}{2v_0^2} \frac{\partial \cos^{-2} \alpha}{\partial \alpha} = 0 = \frac{x}{\cos^2 \alpha} - \frac{gx^2}{2v_0^2} \frac{2\sin \alpha}{\cos^3 \alpha}$$

$$y(x) = x \tan \alpha - \frac{gx^2}{2v_0^2 \cos^2 \alpha}$$
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Unit 1 Fig. 12.6

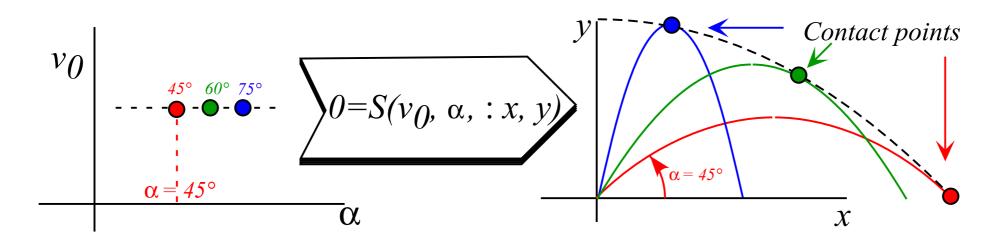
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$$x \frac{\partial \tan \alpha}{\partial \alpha} - \frac{gx^2}{2v_0^2} \frac{\partial \cos^{-2} \alpha}{\partial \alpha} = 0 = \frac{x}{\cos^2 \alpha} - \frac{gx^2}{2v_0^2} \frac{2\sin \alpha}{\cos^3 \alpha} \quad gives: \quad \tan \alpha = \frac{v_0^2}{gx} \text{ or: } x = \frac{v_0^2}{g \tan \alpha}.$$

$$y(x) = x \tan \alpha - \frac{gx^2}{2v_0^2 \cos^2 \alpha}$$
 becomes:
$$S(v_0, \alpha : x, y) = -y + x \tan \alpha - \frac{gx^2}{2v_0^2 \cos^2 \alpha} = 0$$



Unit 1 Fig. 12.6

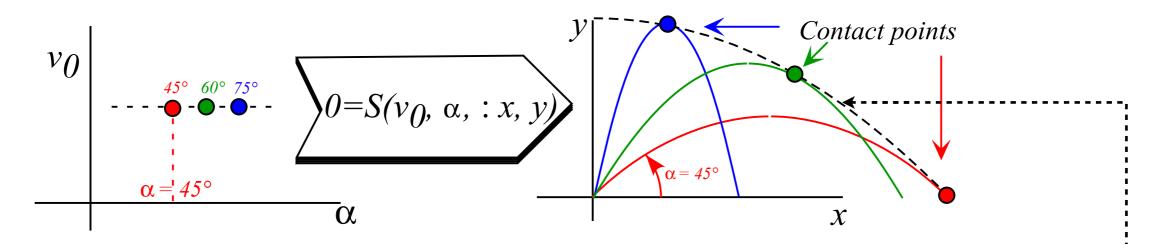
Envelopes of the v_0 -trajectory region contain extremal *contact points* with each trajectory where: $\partial S(v, \alpha; v)$

where:
$$\frac{\partial S\left(v_0,\alpha:x,y\right)}{\partial \alpha} = 0$$

$$x\frac{\partial \tan \alpha}{\partial \alpha} - \frac{gx^2}{2v_0^2}\frac{\partial \cos^{-2}\alpha}{\partial \alpha} = 0 = \frac{x}{\cos^2\alpha} - \frac{gx^2}{2v_0^2}\frac{2\sin\alpha}{\cos^3\alpha} \qquad \tan\alpha = \frac{v_0^2}{gx} \text{ or: } x = \frac{v_0^2}{g\tan\alpha}.$$

$$y_{env}(x) = x\tan\alpha - \frac{gx^2}{2v_0^2}\left(1 + \tan^2\alpha\right) \Rightarrow y_{env}(x) = x\frac{v_0^2}{gx} - \frac{gx^2}{2v_0^2}\left(1 + \frac{v_0^4}{g^2x^2}\right)$$

$$y(x) = x \tan \alpha - \frac{gx^2}{2v_0^2 \cos^2 \alpha}$$
 becomes:
$$S(v_0, \alpha : x, y) = -y + x \tan \alpha - \frac{gx^2}{2v_0^2 \cos^2 \alpha} = 0$$



Envelopes of the v_0 -trajectory region contain extremal *contact points* with each trajectory where: $\partial S(v_0, \alpha: x, y)$

where:
$$\frac{\partial S(v_0, \alpha : x, y)}{\partial \alpha} = 0$$

 $\frac{\partial \tan \alpha}{\partial x} = \frac{gx^2}{\partial \cos^{-2} \alpha} = 0 = \frac{x}{2\sin \alpha} = 0$

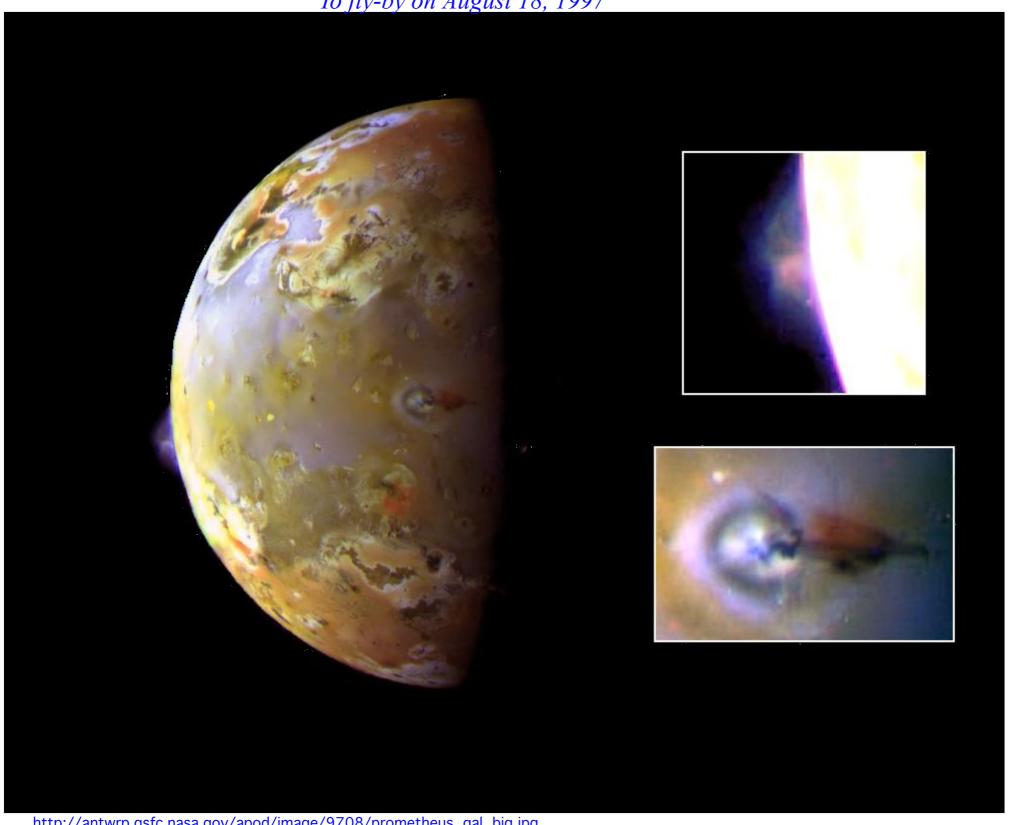
$$x\frac{\partial \tan \alpha}{\partial \alpha} - \frac{gx^2}{2v_0^2} \frac{\partial \cos^{-2} \alpha}{\partial \alpha} = 0 = \frac{x}{\cos^2 \alpha} - \frac{gx^2}{2v_0^2} \frac{2\sin \alpha}{\cos^3 \alpha} \qquad \tan \alpha = \frac{v_0^2}{gx} \text{ or: } x = \frac{v_0^2}{g \tan \alpha}.$$

$$y_{env}(x) = x \tan \alpha - \frac{gx^2}{2v_0^2} \left(1 + \tan^2 \alpha\right) \Rightarrow y_{env}(x) = x \frac{v_0^2}{gx} - \frac{gx^2}{2v_0^2} \left(1 + \frac{v_0^4}{g^2x^2}\right)$$

$$y_{env}(x) = \frac{v_0^2}{g} - \frac{gx^2}{2v_0^2} - \frac{gx^2}{2v_0^2} \frac{v_0^4}{g^2x^2} = \frac{v_0^2}{2g} - \frac{gx^2}{2v_0^2} - \frac{Envelope}{function}$$

The Plumes of Prometheus

NASA-Galileo Project Io fly-by on August 18, 1997



http://antwrp.gsfc.nasa.gov/apod/image/9708/prometheus_gal_big.jpg

http://antwrp.gsfc.nasa.gov/apod/ap970818.html

http://science.nasa.gov/science-news/science-at-nasa/1999/ast04oct99_1/

Io's ALIEN VOLCANOES



Pretty bad sketch of plumes (Las Vegas model of planetary ejecta?)

Do these guys need a geometry lesson?

Io's ALIEN VOLCANOES

Go fly a kite?

SCIENTISTS ARE EAGER FOR A CLOSER LOOK AT THE SOLAR SYSTEM'S STRANGEST AND MOST ACTIVE VOLCANOES WHEN GALILEO FLIES BY TO ON OCTOBER 11.

October 4, 1999: Thirty years ago, before the Voyager probes visited Jupiter, if you had described to a literary critic it would have been declared overwrought science fiction. Jupiter's strange moon is literally bursting with volcanoes. Dozens of active vents pepper the landscape which also includes gigantic frosty plains, towering mountains and volcanic rings the size of California. The volcanoes themselves are the hottest spots in the solar system with temperatures exceeding 1800 K (1527 C). The plumes which rise 300 km into space are so large they can be seen from Earth by the Hubble Space Telescope. Confounding common sense, these high-rising ejecta seem to be made up of, not blisteringly hot lava, but frozen sulfur

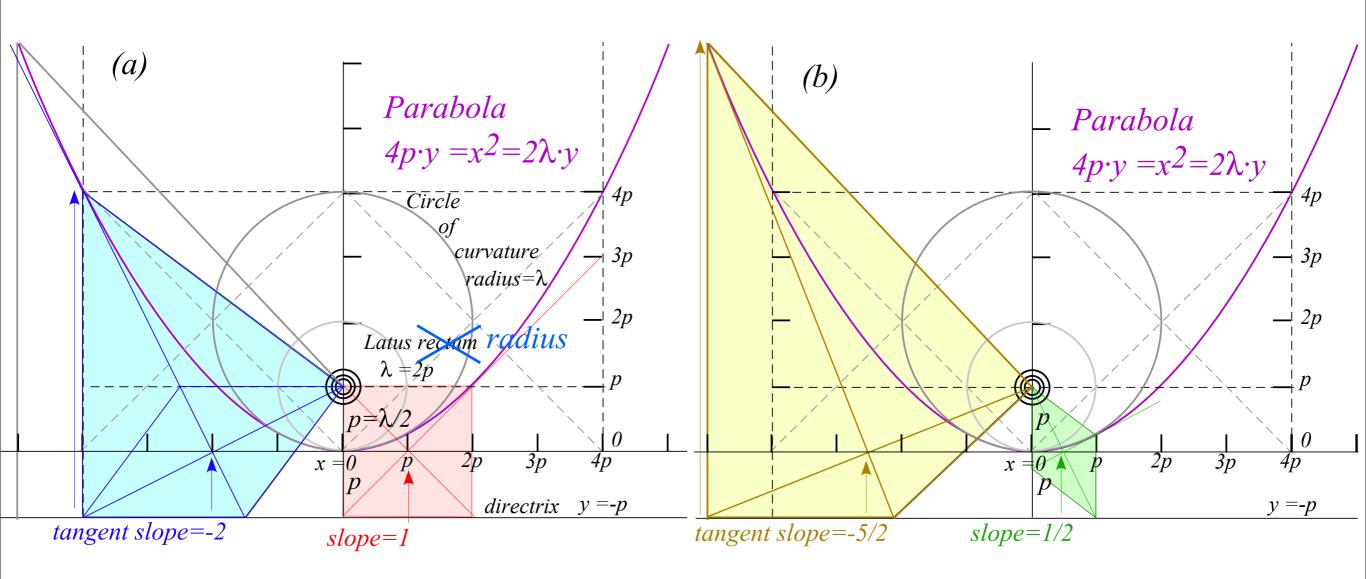


dioxide. And to top it all off, lo bears a striking resemblance to a pepperoni pizza. Simply unbelievable.

Right: Digital Radiance simulation of Pillan Patera just before the Galileo flyby. click for animation → .

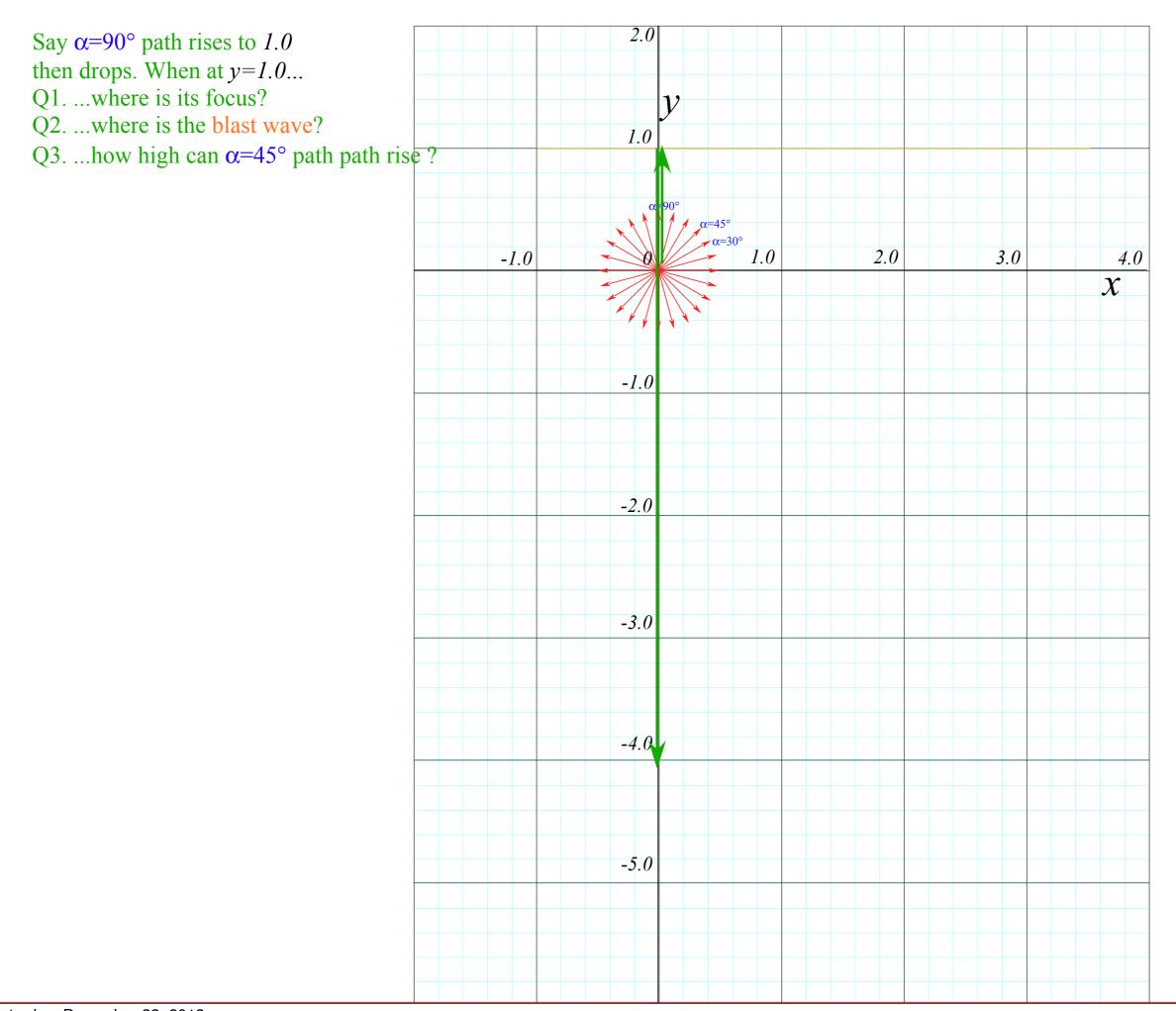
...conventional parabolic geometry...carried to extremes...

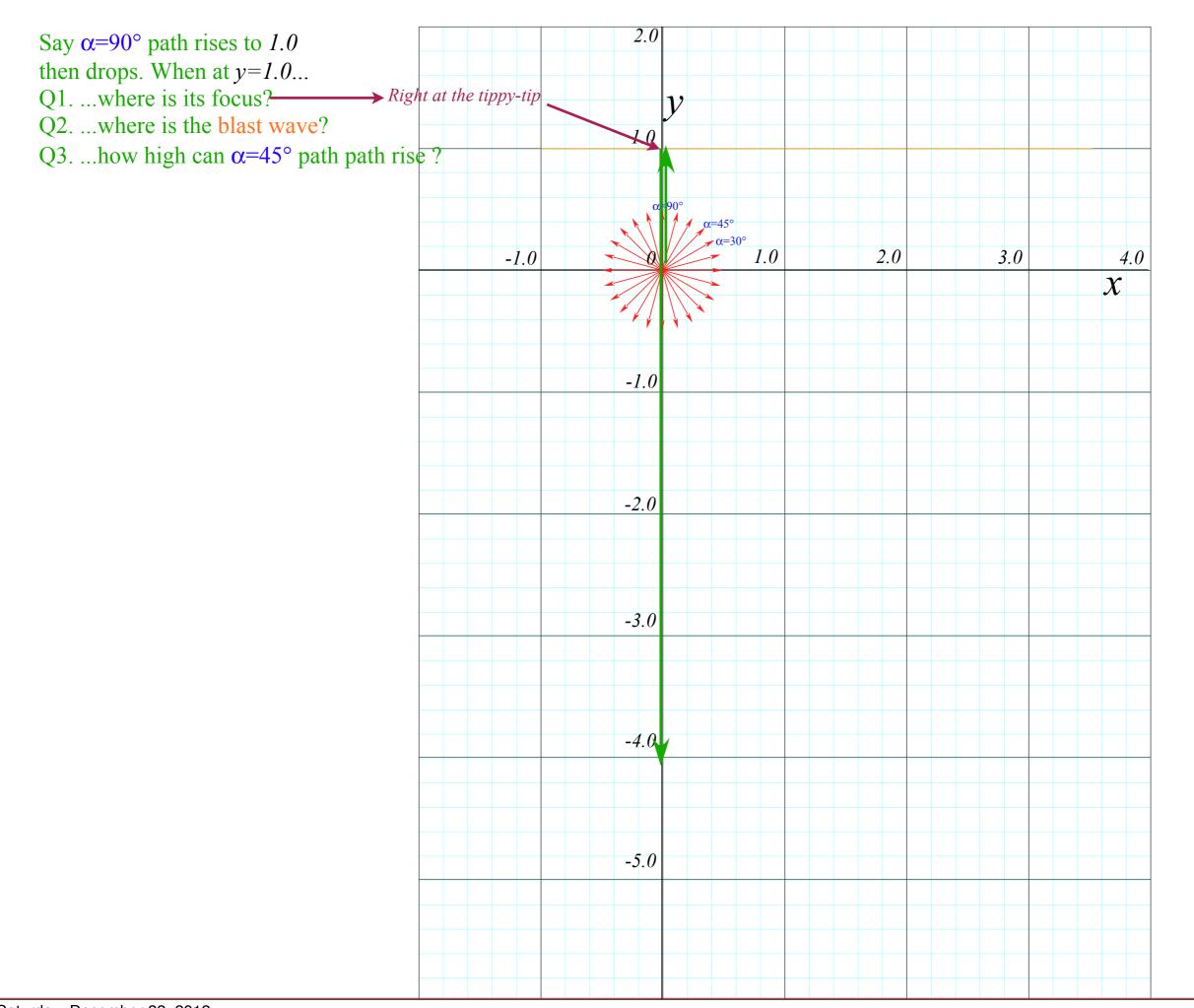
Recall Lecture 6 p.29

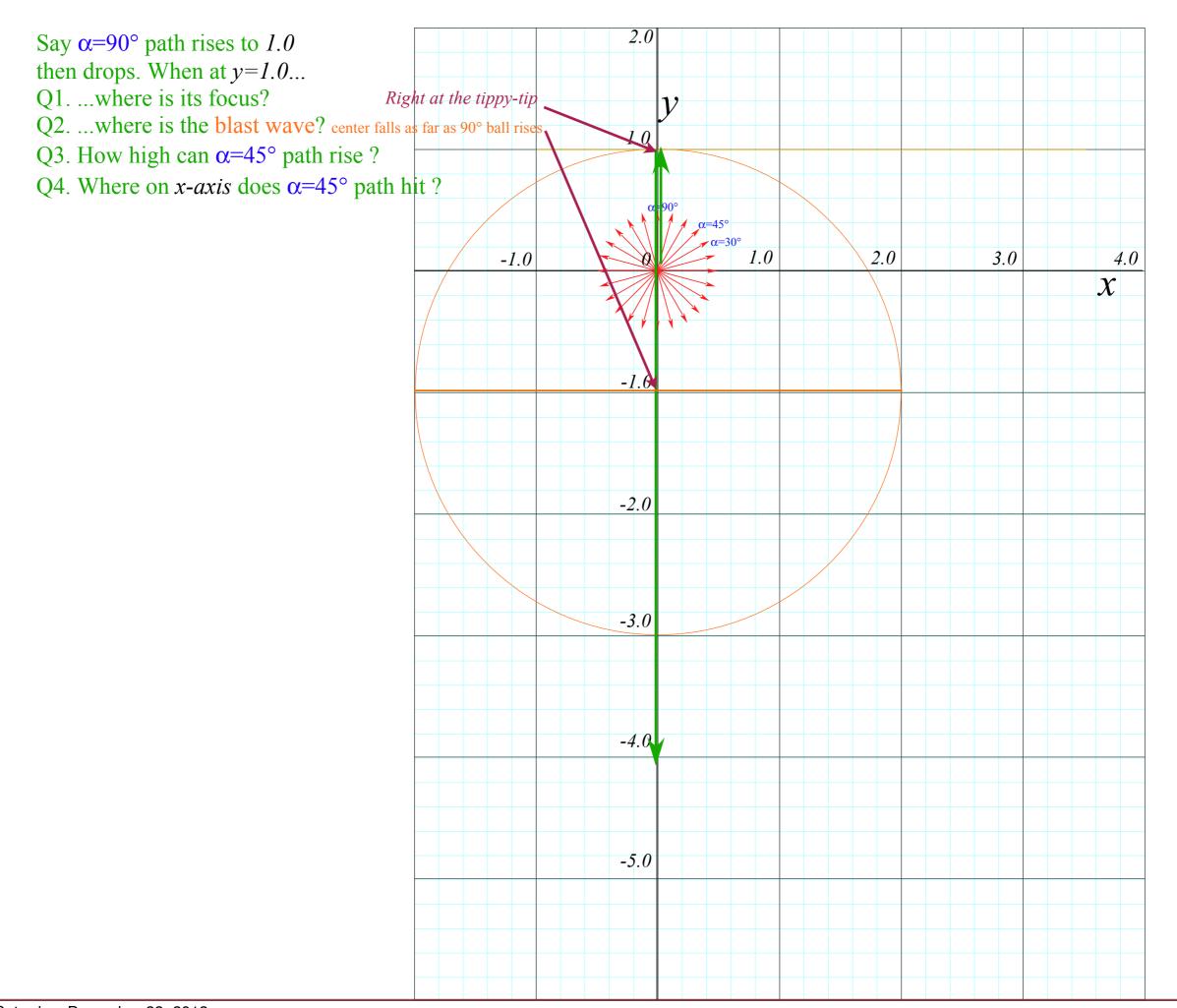


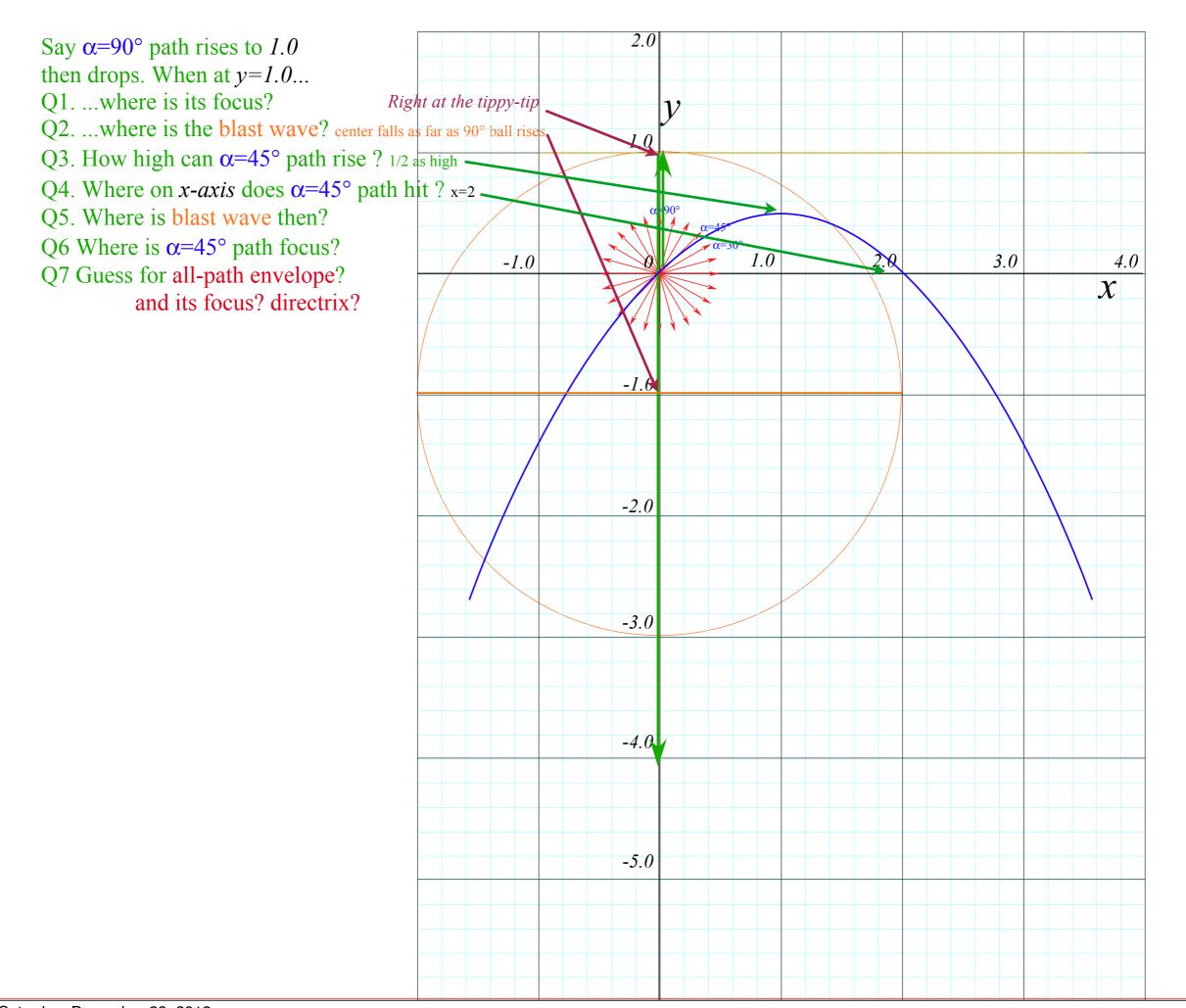
Unit 1 Fig. 9.4

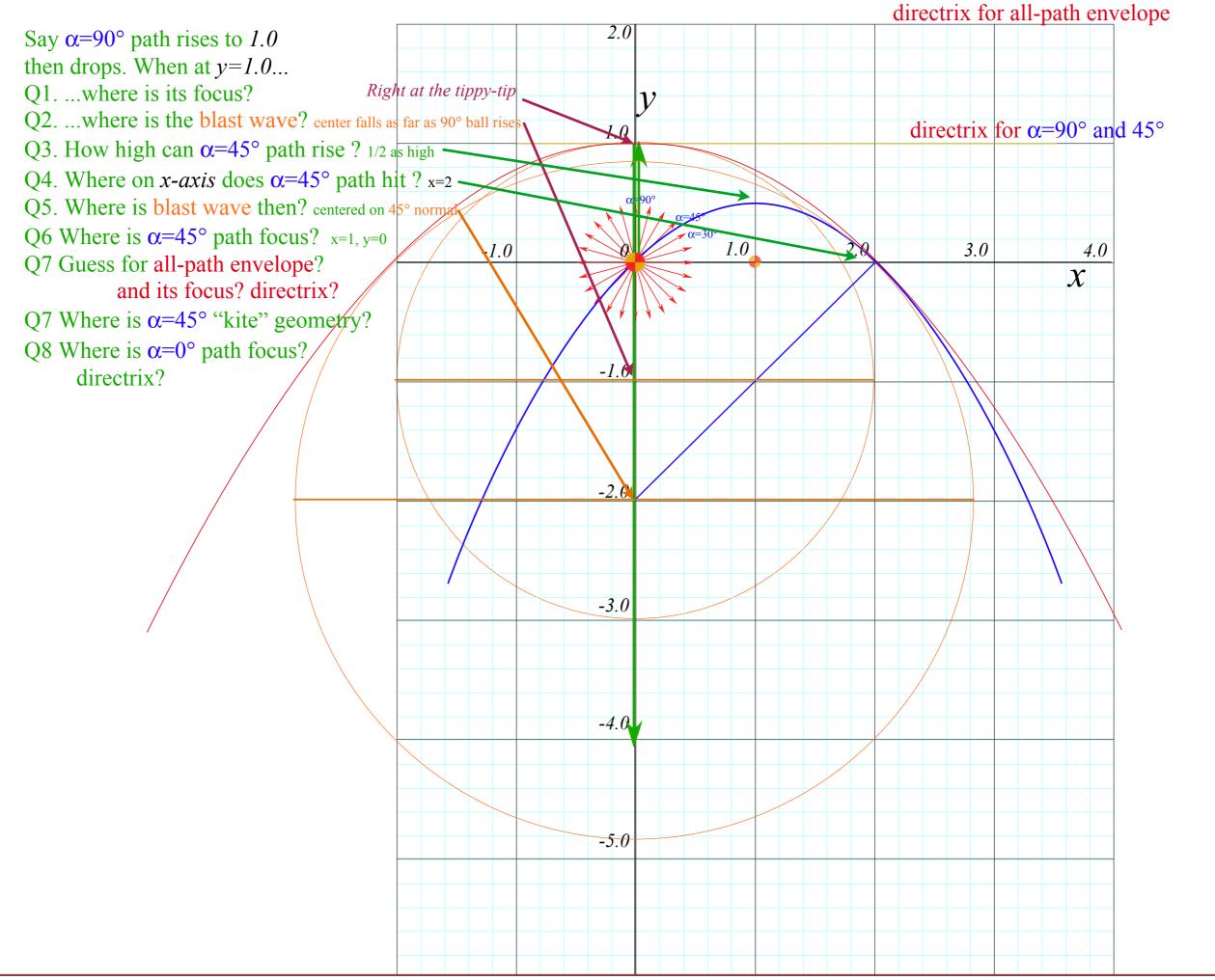
A general contact transformation from sophomore physics

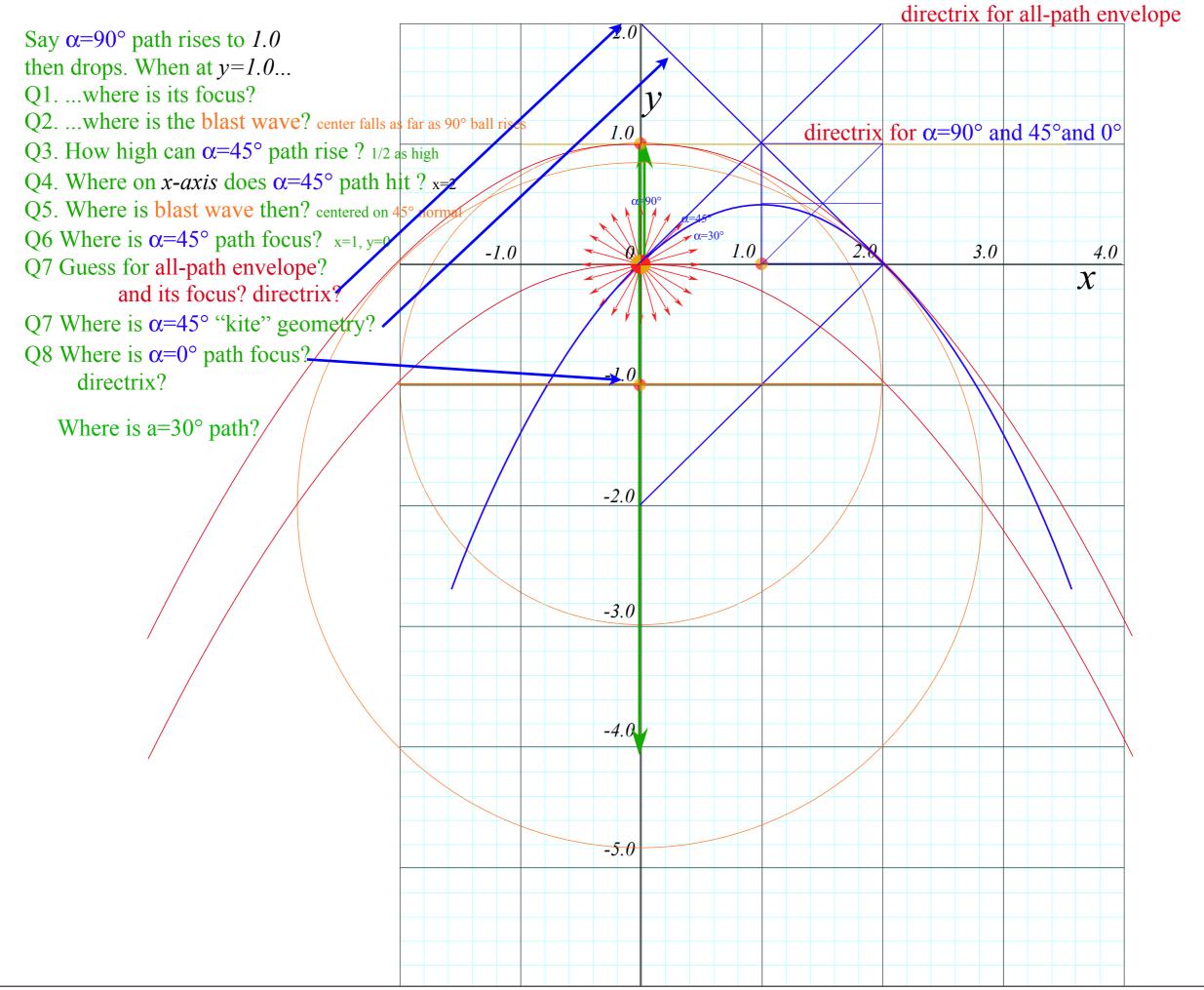


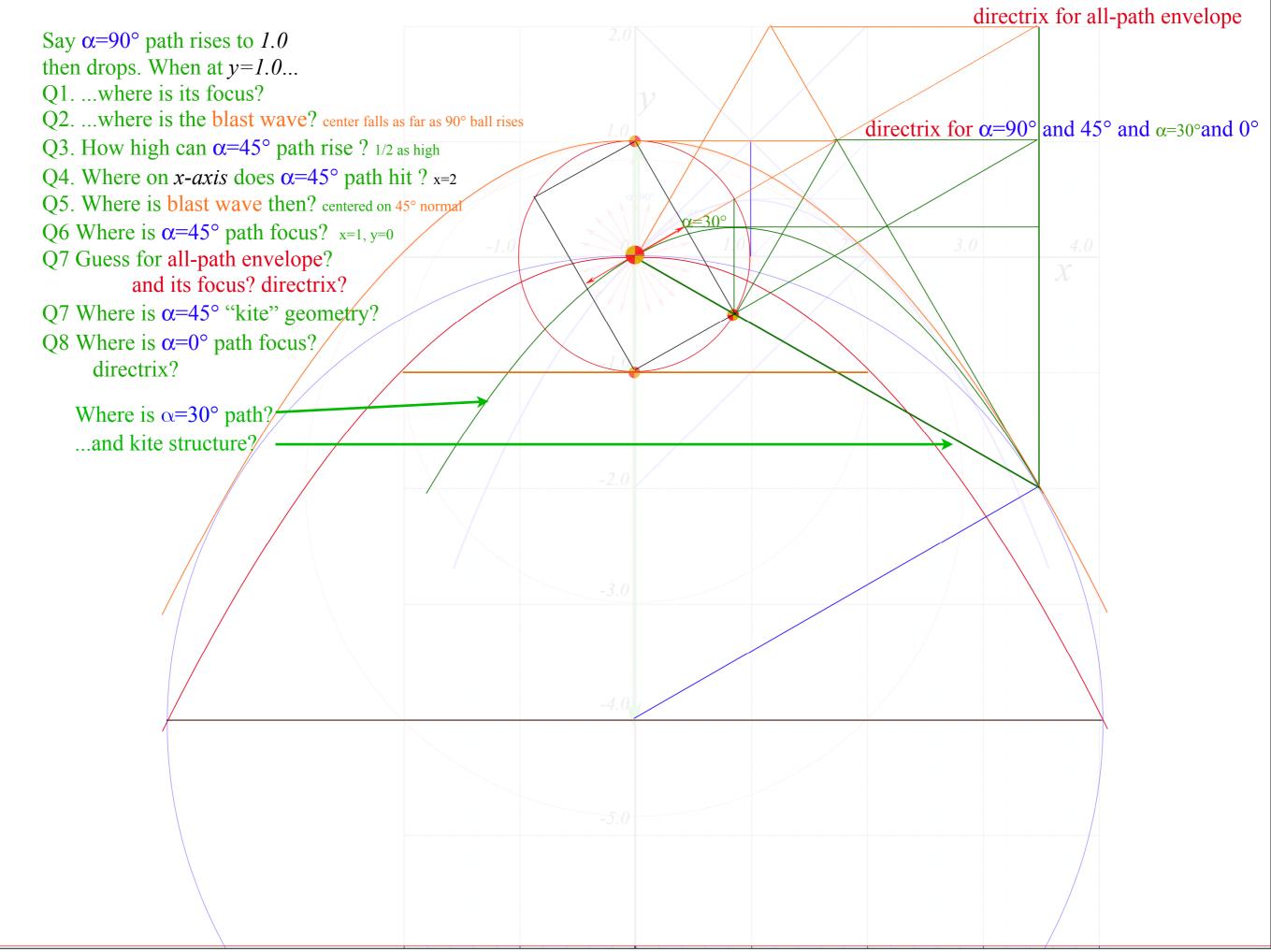


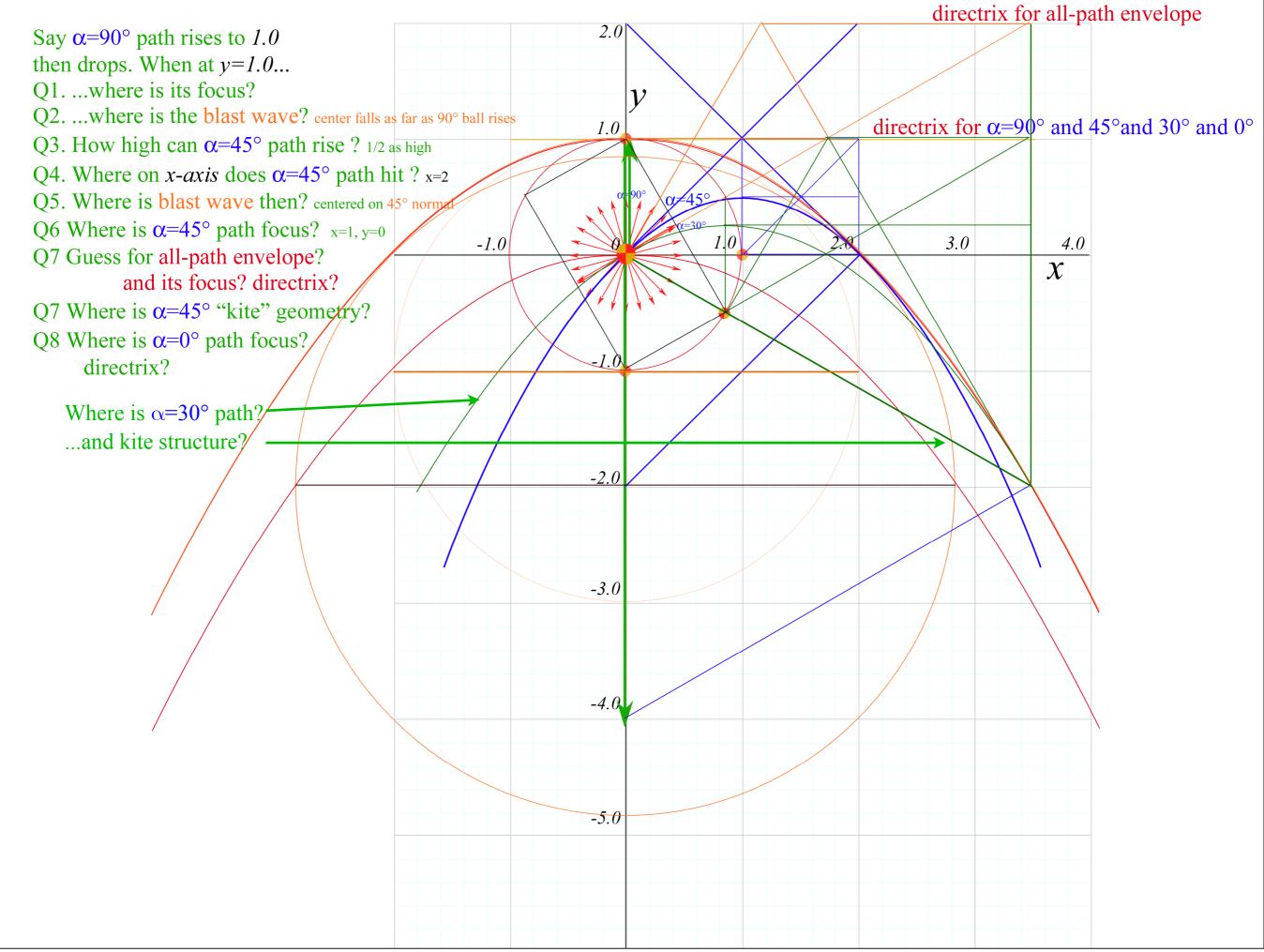


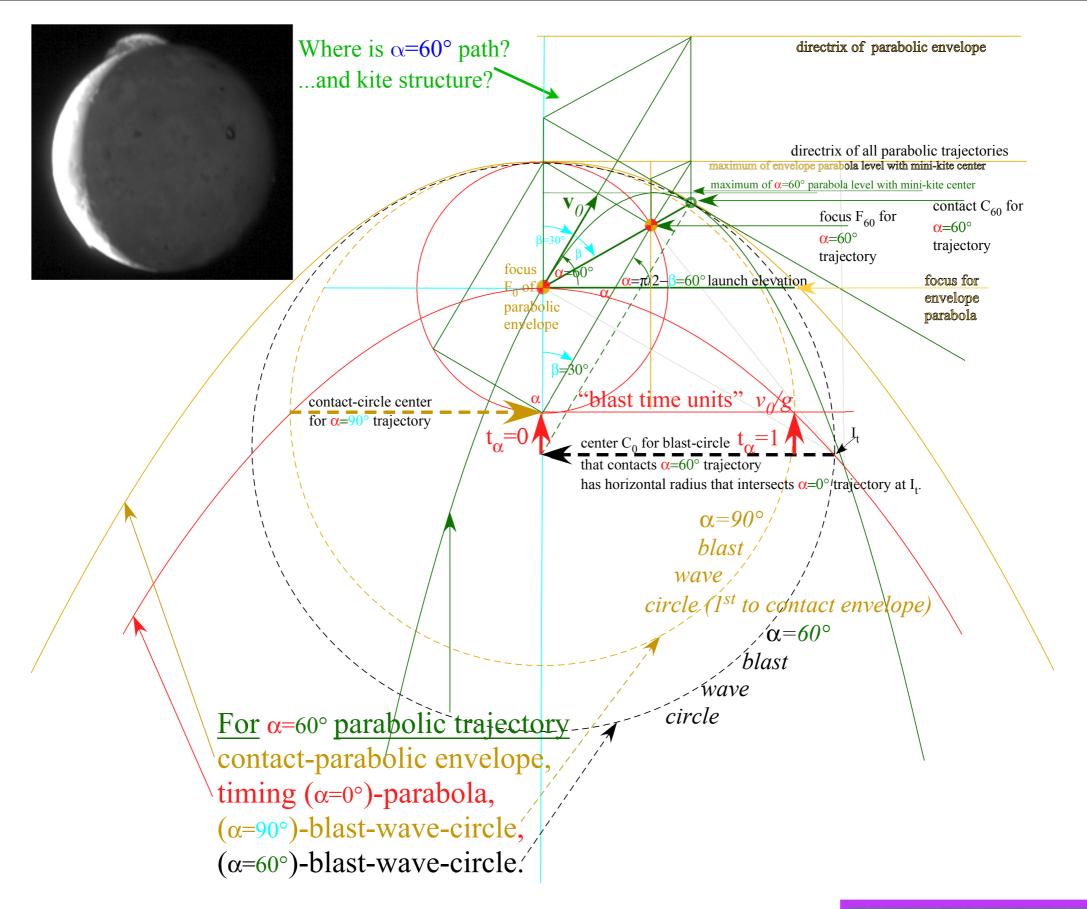












Lecture 10 ends here Thur. 9.20.2012