

Relawavity: Simple trigonometry leads to understanding of relativity and quantum theory

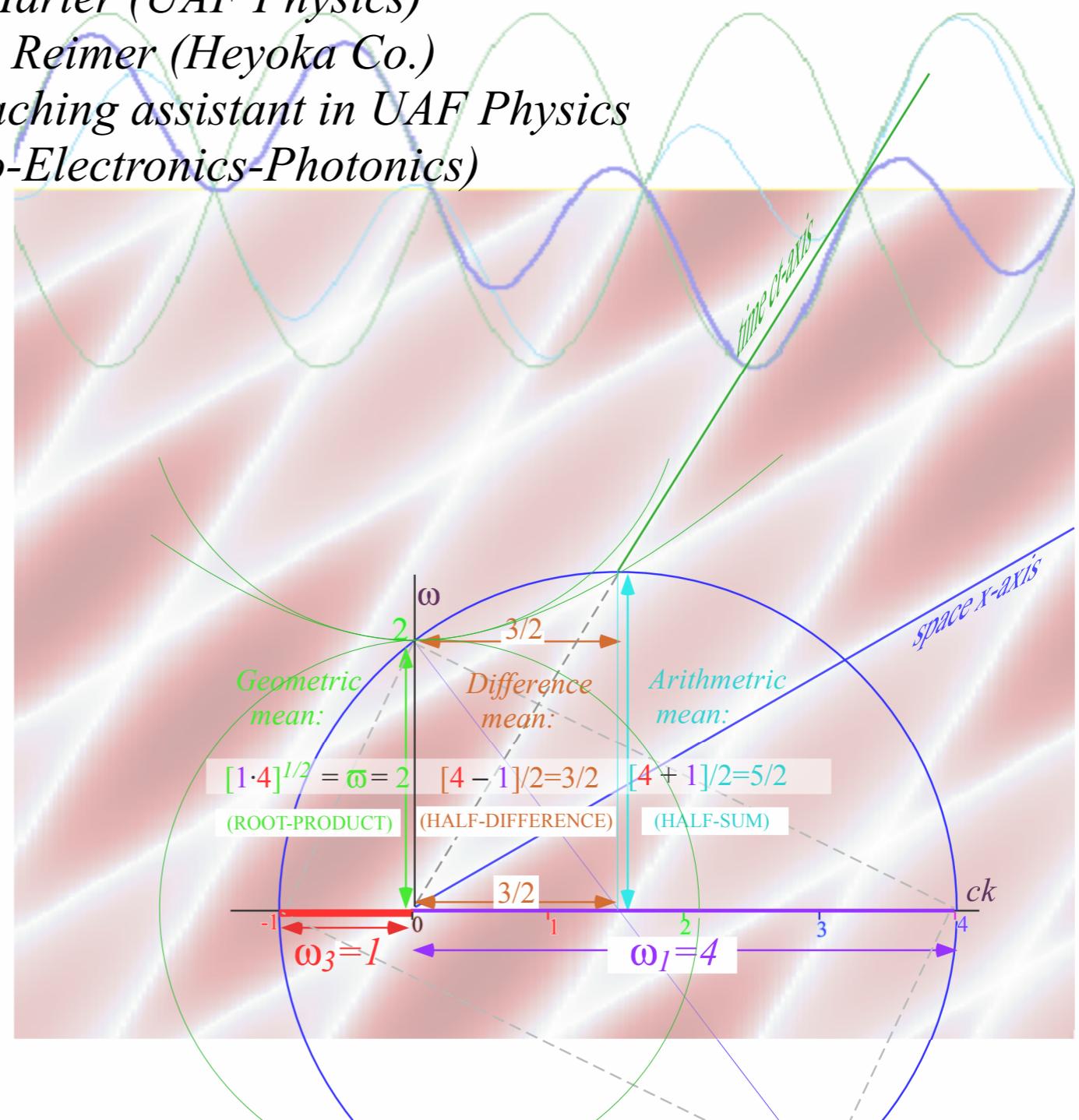
INBRE 2016

Presentation and production by

W. G. Harter (UAF Physics)

Dr. T.C. Reimer (Heyoka Co.)

*Al Calabrese (Teaching assistant in UAF Physics
and Micro-Electronics-Photonics)*



Relawavity: Simple trigonometry leads to understanding of relativity and quantum theory

Workshop by prof. W. G Harter (UAF Physics), Dr. T.C Reimer (Heyoka Co.), and Al Calabrese

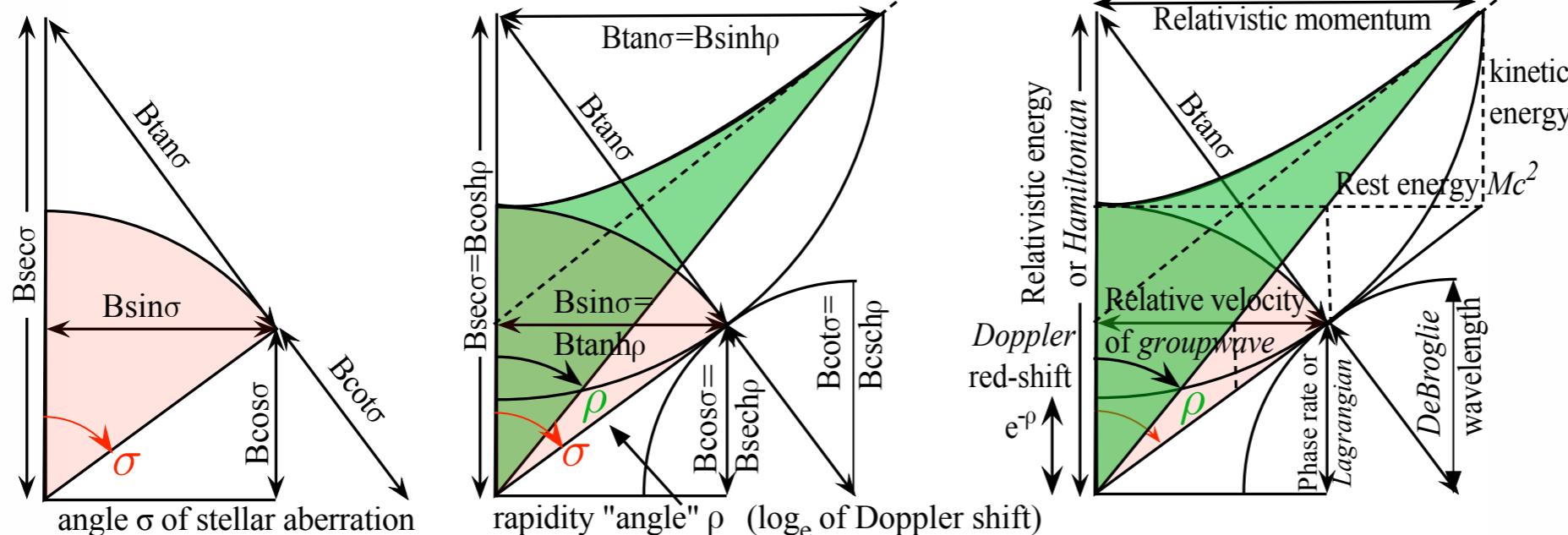
(Teaching assistant in UAF Physics and Micro-Electronics-Photonics lab).

Modern science from Astrophysics to Zoology depends increasingly on two pillars of modern physics, special relativity (SR) and quantum mechanics (QM), that are based on properties of light waves. The bioscience renaissance could not happen without the optics of maser, laser, UV, Xray, and synchrotron effects due to SR and QM theory that is still regarded as esoteric mystery. This workshop seeks to demystify SR and QM theory using high school trigonometry of plane light waves. Using diagrams below of circle trig functions ($\sin\sigma, \cos\sigma, \tan\sigma$) and inverses ($\csc\sigma, \sec\sigma, \cot\sigma$) we show each one is also a hyperbolic function that is key to SR and QM. A circle function like sine is a function $\sin\sigma$ of circular sector area σ that astronomers call a *stellar aberration*. $\sin\sigma$ equals a function $\tanh\rho$ of hyperbola sector area ρ that physicists call *rapidity*. (It happens that equality $\sin\sigma=\tanh\rho$ implies $\tan\sigma=\sinh\rho$, and similarly, $\csc\sigma=\coth\rho$ implies $\cot\sigma=\text{csch}\rho$. Finally, $\cos\sigma=\text{sech}\rho$ implies $\sec\sigma=\cosh\rho$, a key pair.)

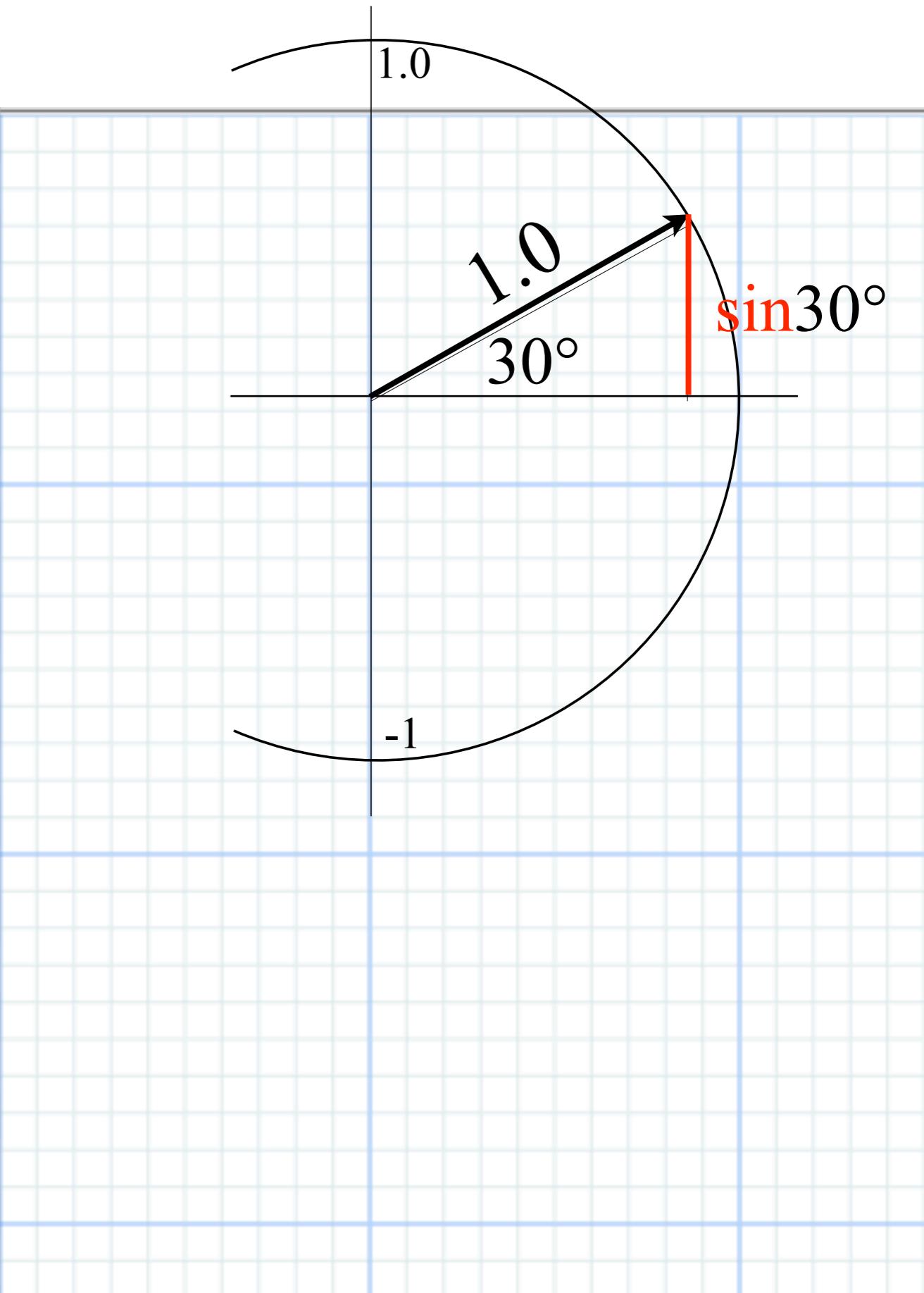
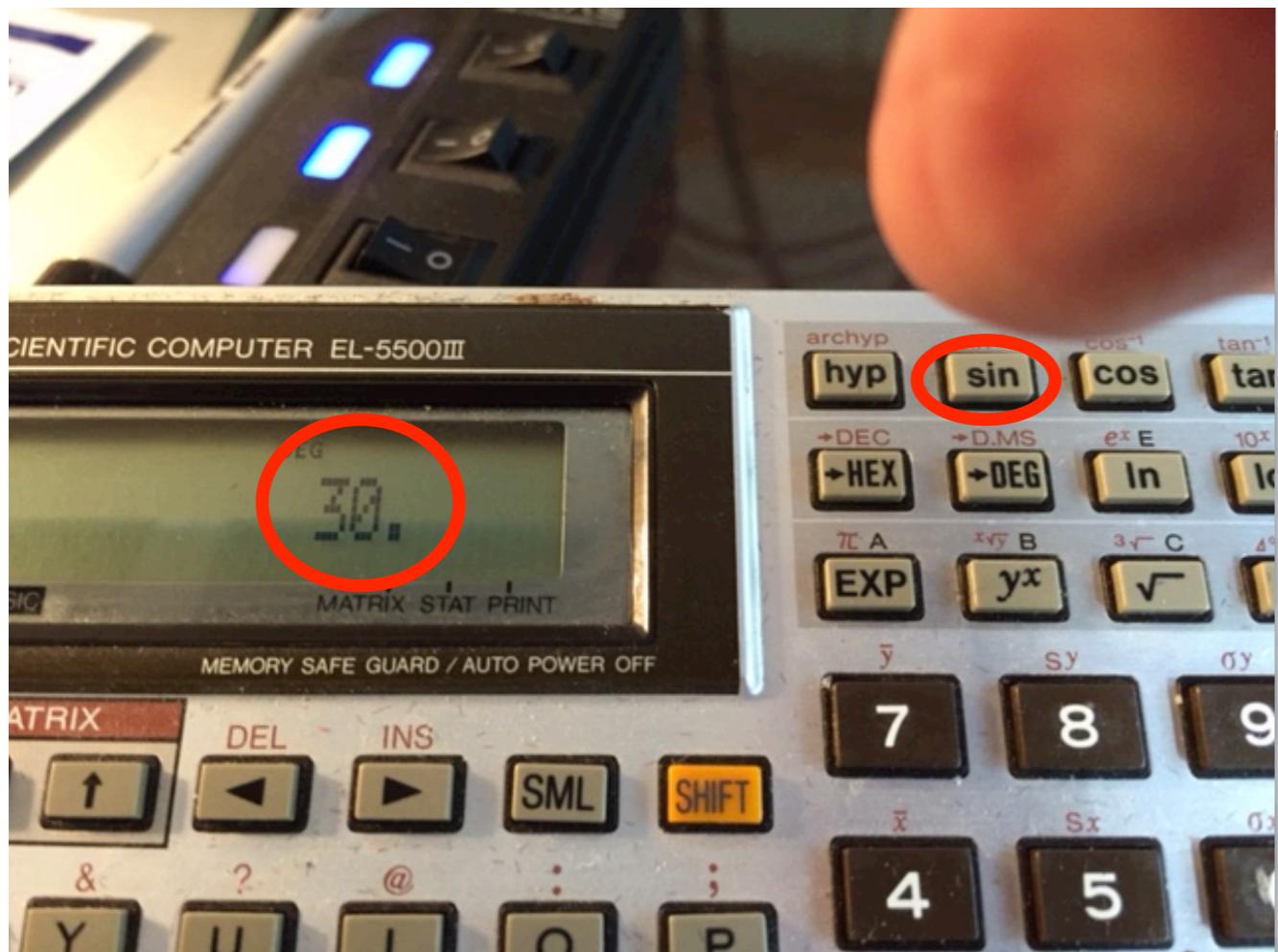
A diagram or “roadmap” of the functions follows from space-time plots of wave interference for a pair of laser plane waves colliding head-on. Similar diagrams arise for inverse or per-time-per-space plots (*i.e.*, frequency v vs wave-number κ). The (x, ct) -plots reveal relativistic space-time mechanics while $(c\kappa, v)$ -plots show quantum energy-momentum effects. Both plots vary with ρ or σ and Relawavity web apps provide animation of this.

Rapidity ρ is the natural logarithm of Doppler factor $b=e^\rho$ and key to understanding that a laser wave can be blue shifted two indistinguishable ways: (1) Tune up the laser, or (2) Accelerate the laser toward the observer. Red shifts act similarly. From this can be understood the super-constant nature of light speed $c=299,792,458$ meters per second. Both classical and quantum mechanics follow from this that we call **Evenson’s Axiom: All Colors go c !** The metrological precision revolution began with Ken Evenson’s CW laser speed of light measurements in 1972. What is needed now is similar improvement in our *precision of thinking* about quantum optics.

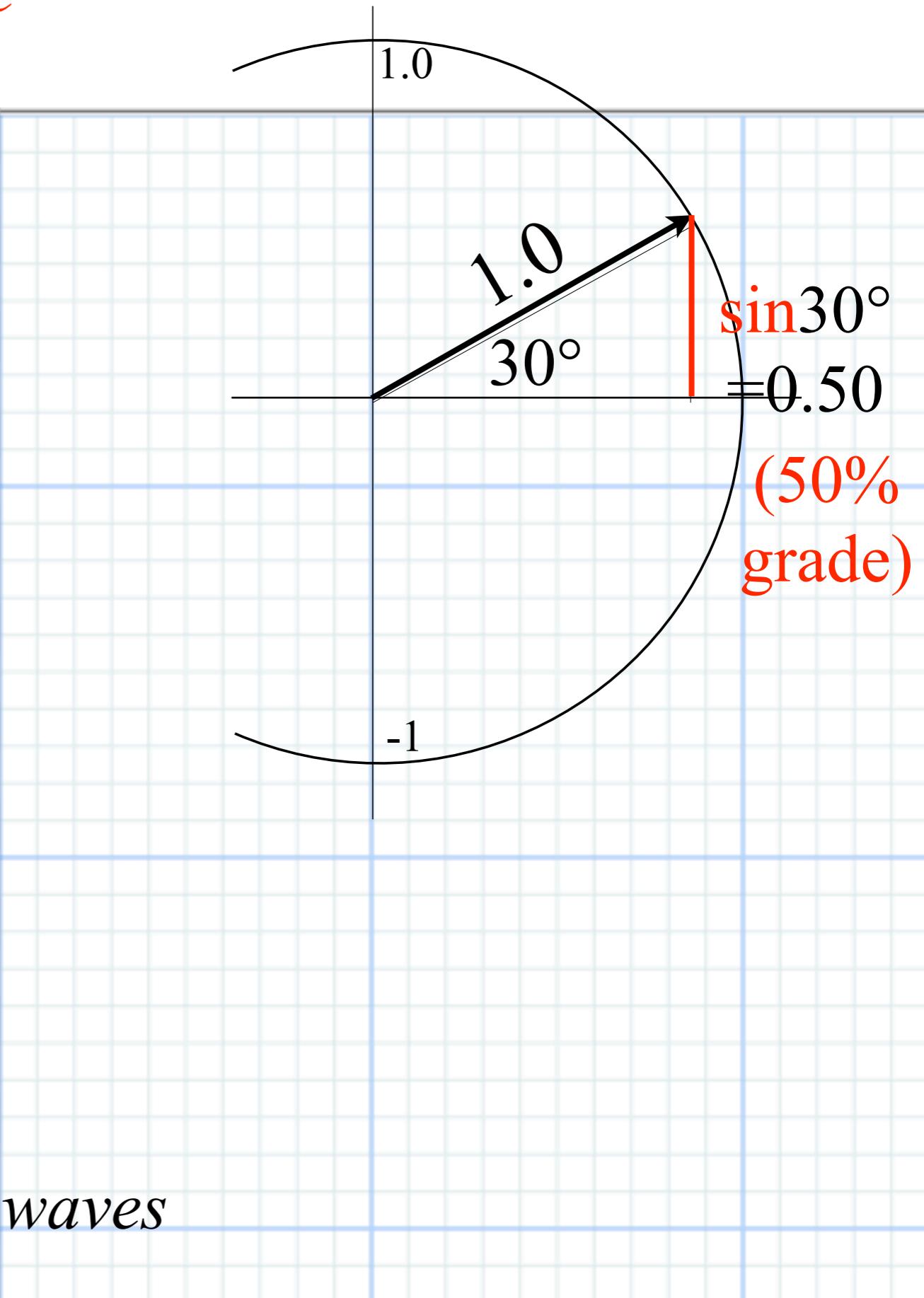
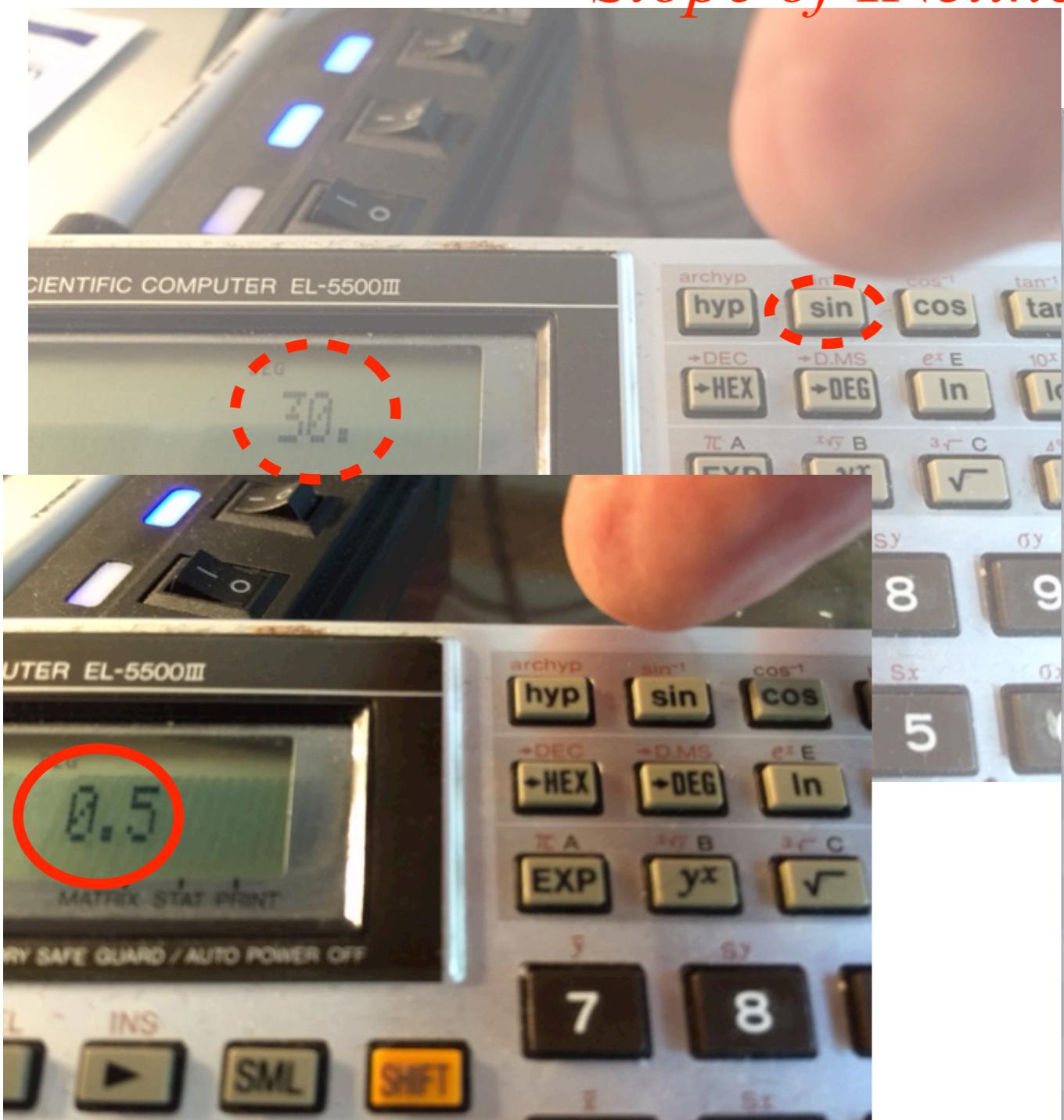
Circle trigonometry becomes Hyperbola trigonometry becomes Relativity and Quantum geometry.



Learning about SIN

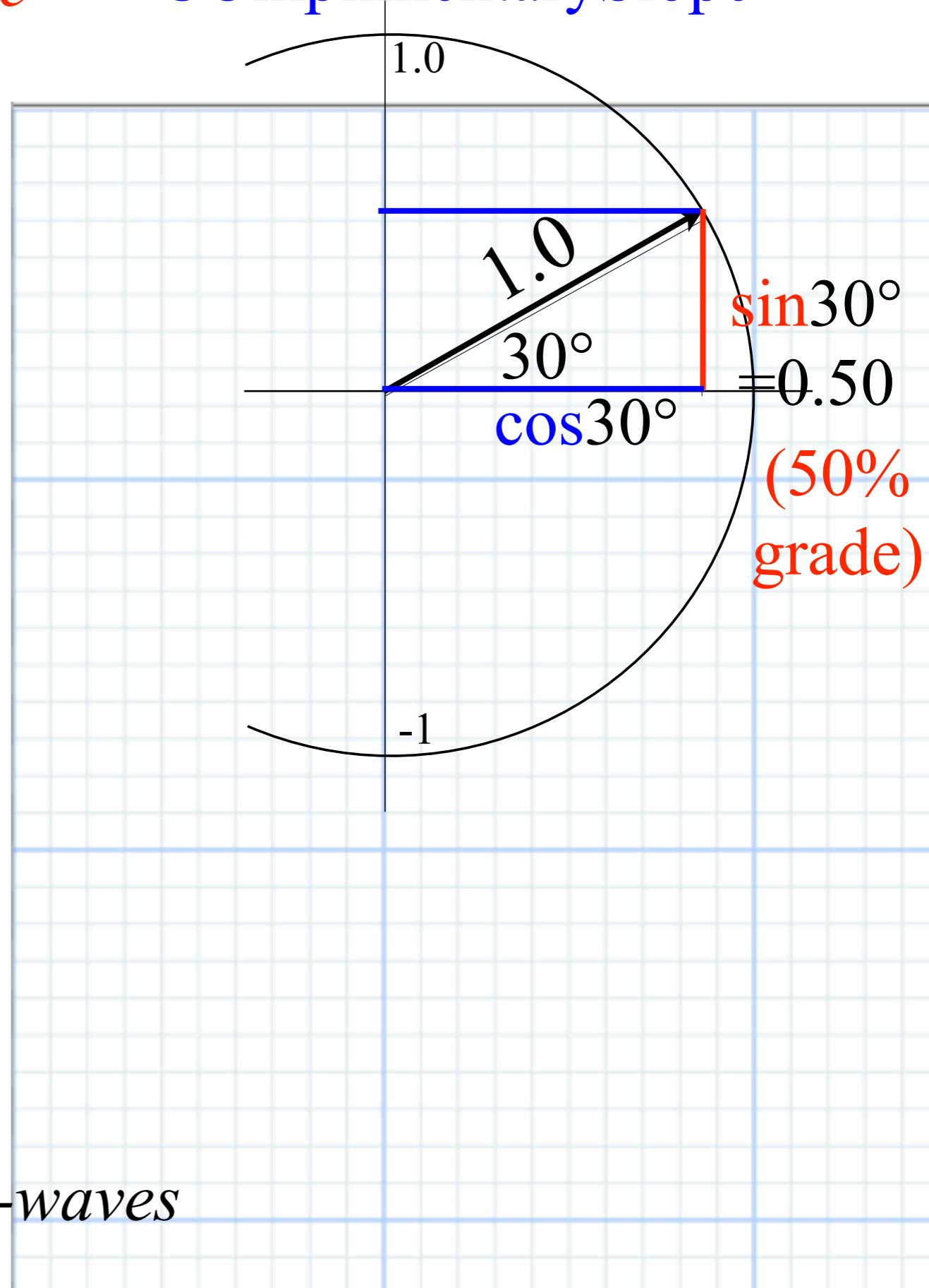
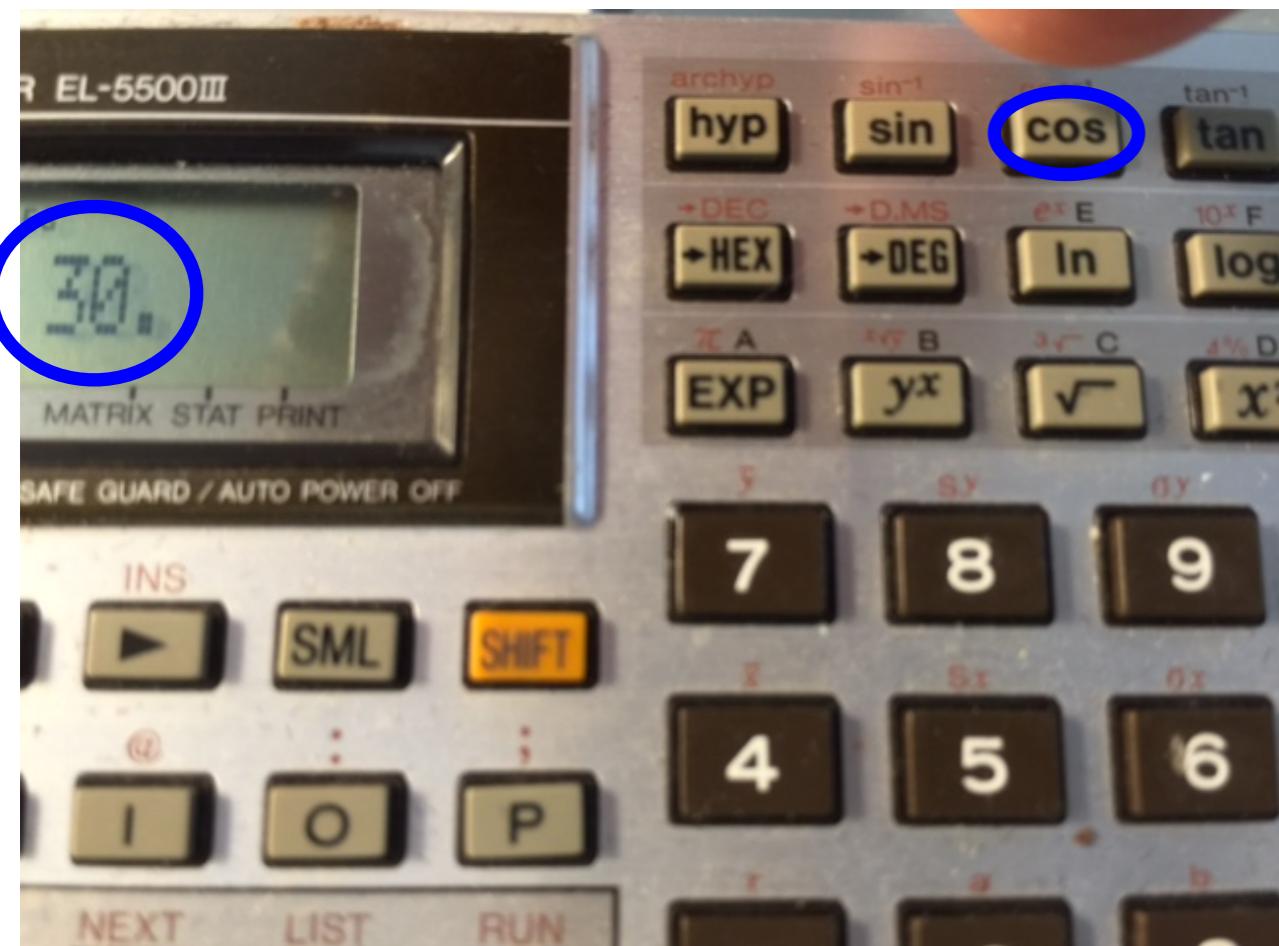


Learning about SIN “Slope of INcline”



It's mostly about triangles and *sine*-waves

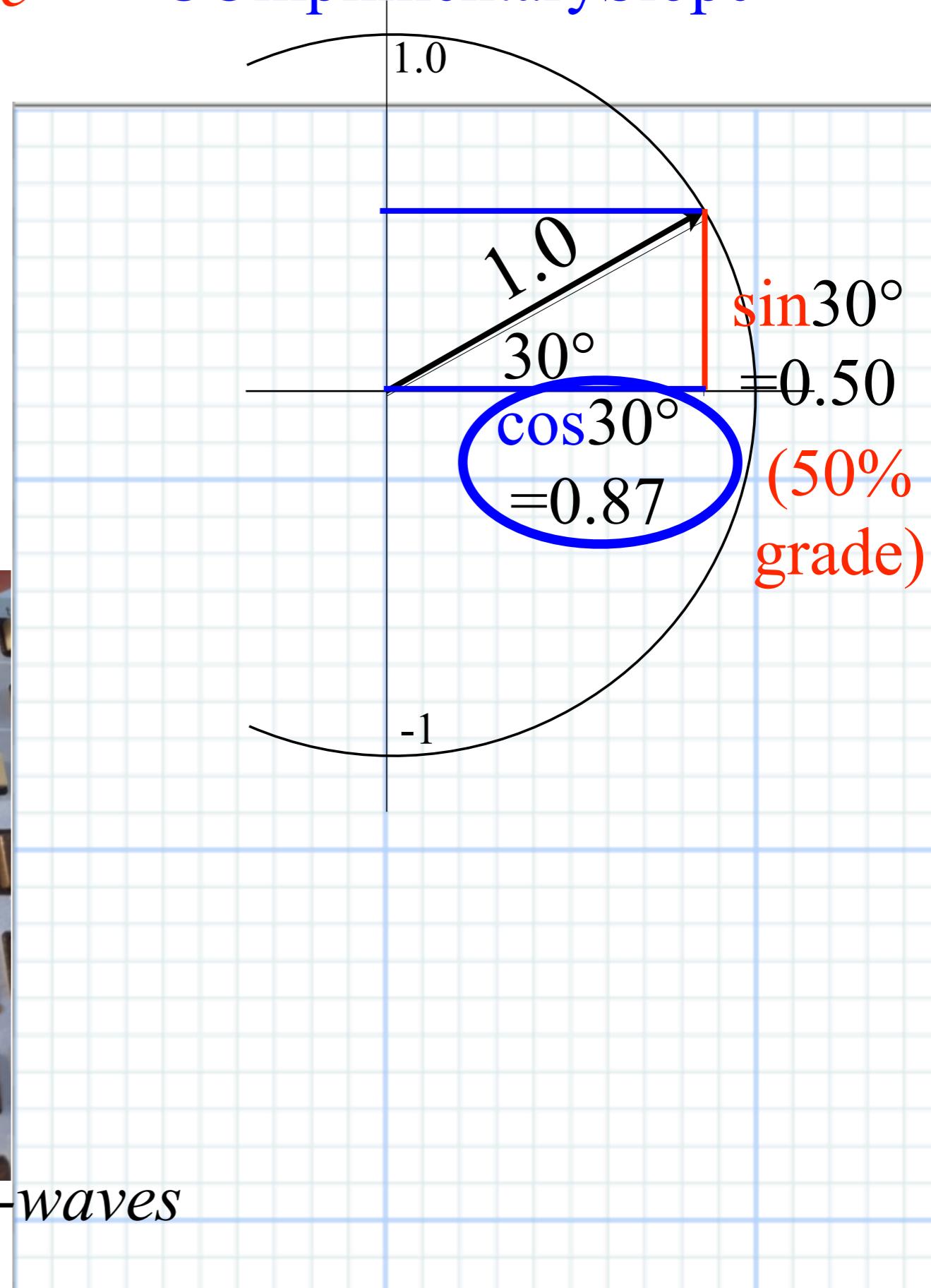
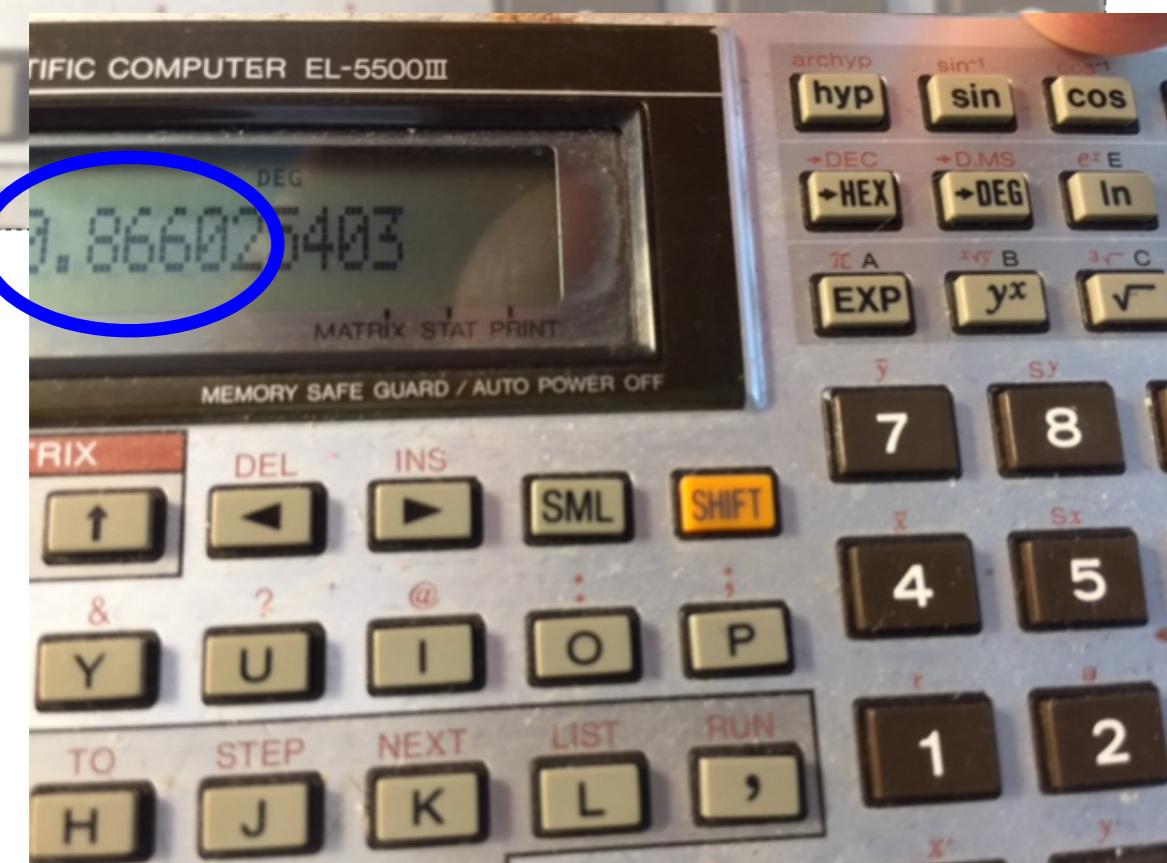
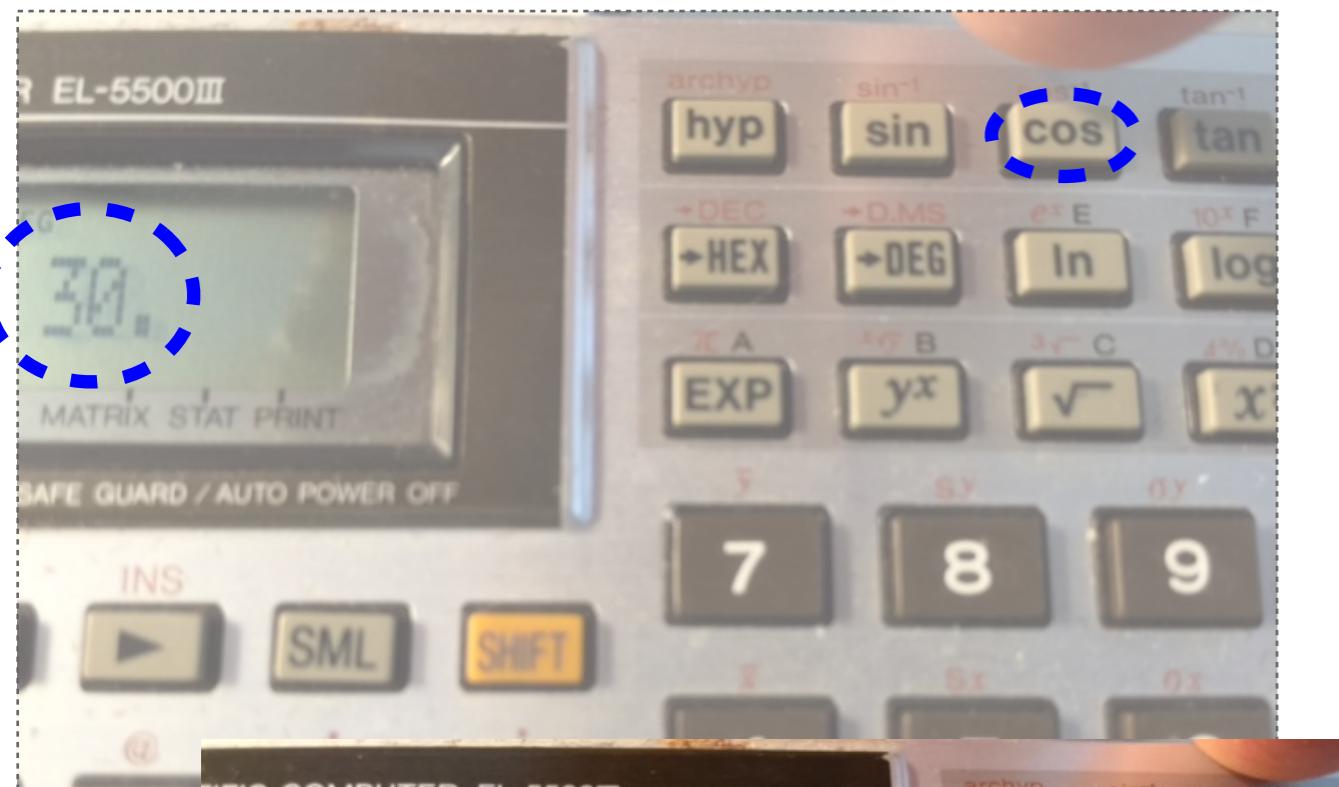
Learning about SIN and the COSin “Slope of INcline” “COmplimentarySlope”



It's mostly about triangles and sine-waves

Learning about SIN and the COSin

“Slope of INcline” “COmplimentary Slope”



It's mostly about triangles and sine-waves

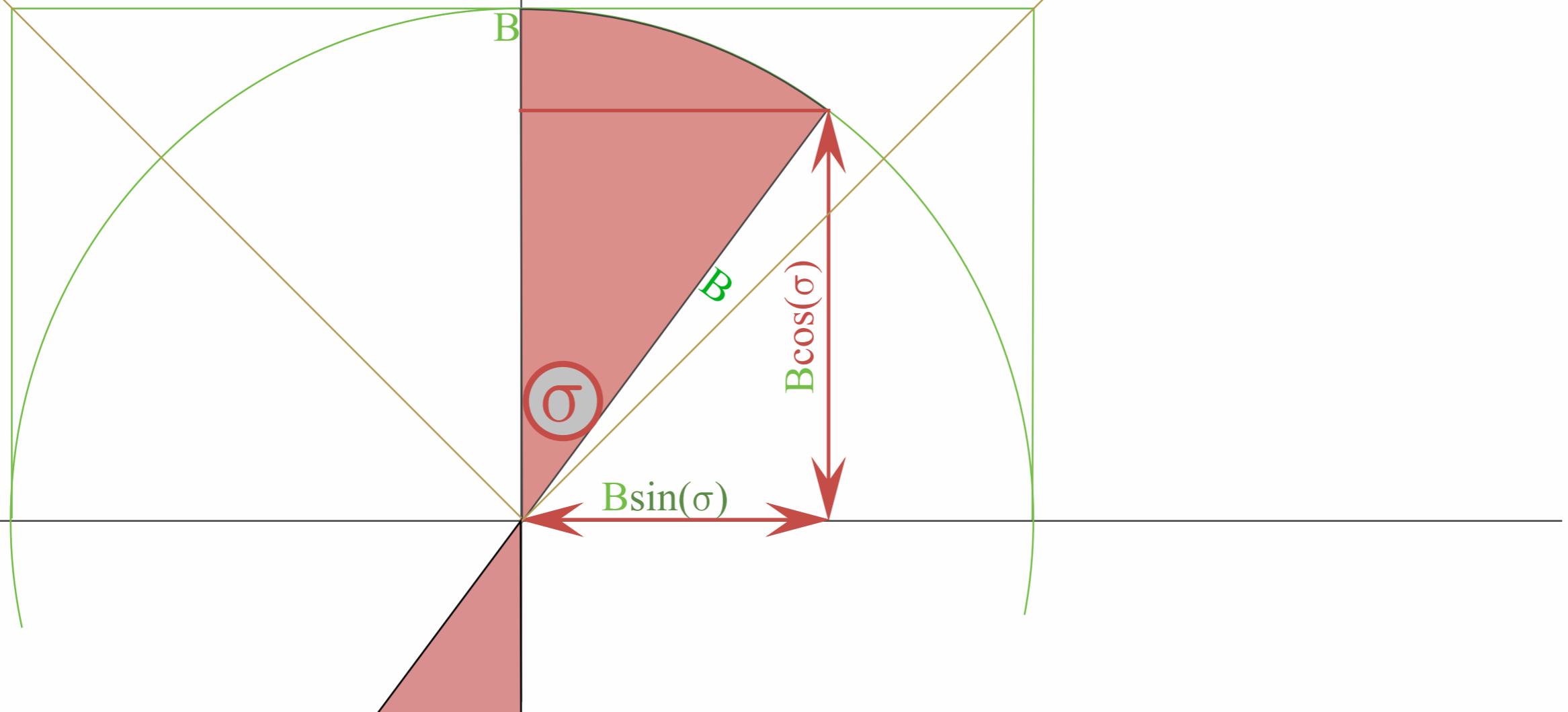
This map has circle sector arc-area $\sigma = 0.6435$

set to angle $\angle\sigma = 36.87^\circ = 0.6435 \text{radian}$

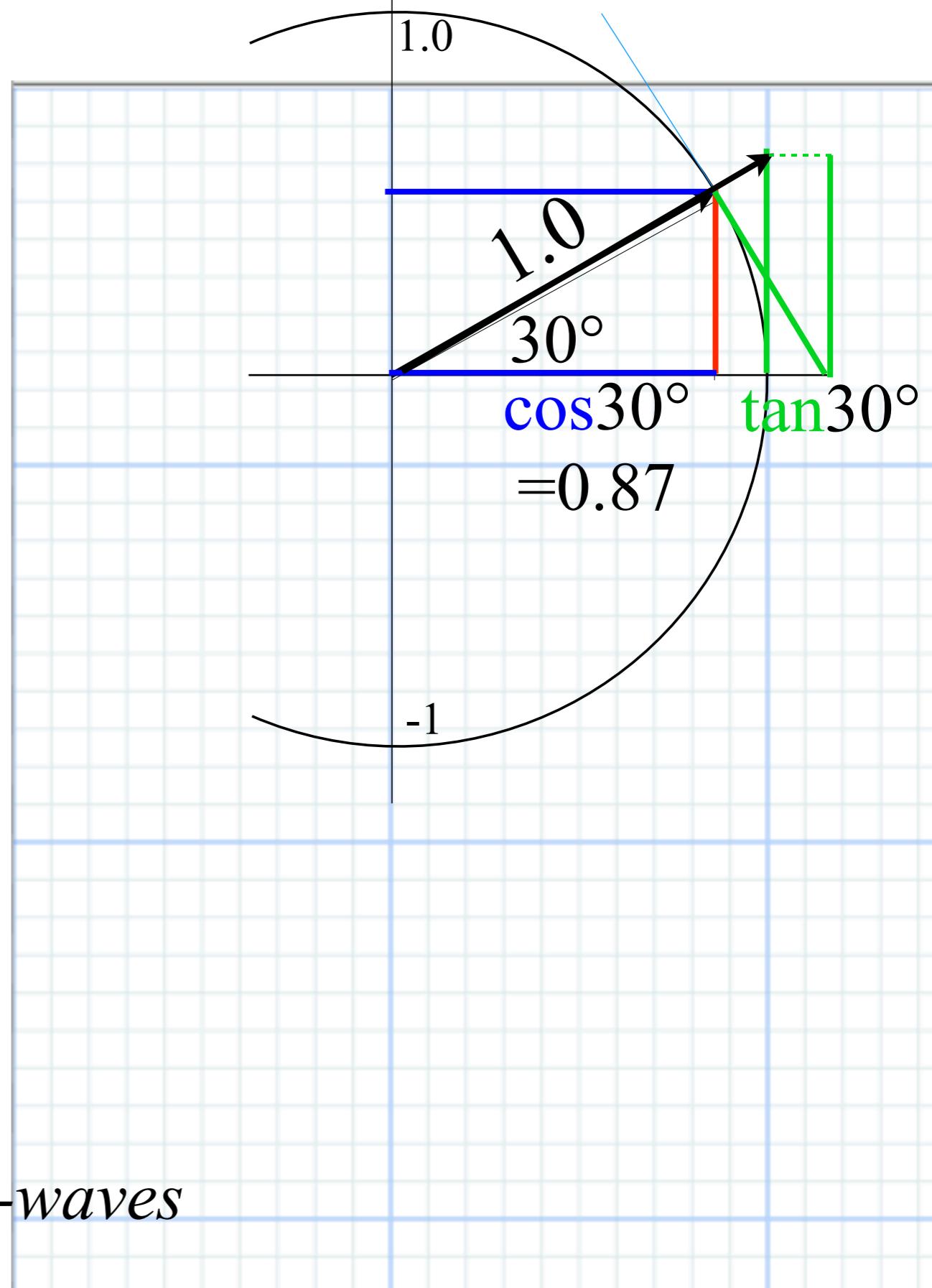
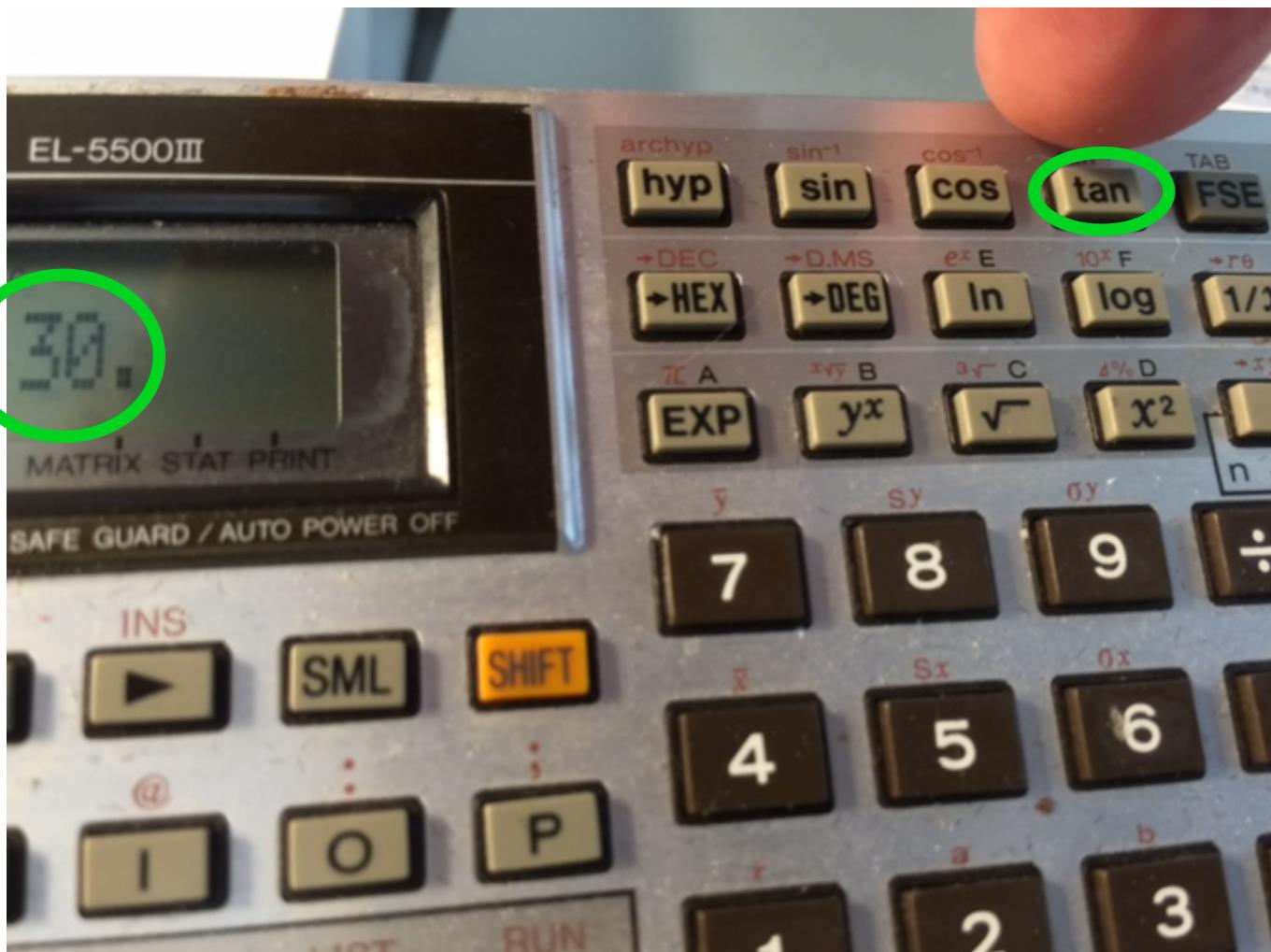
$$\sin(\sigma) = 0.6000 = 3/5$$

$$\cos(\sigma) = 0.8000 = 4/5$$

a small change : we measure angle by sector area

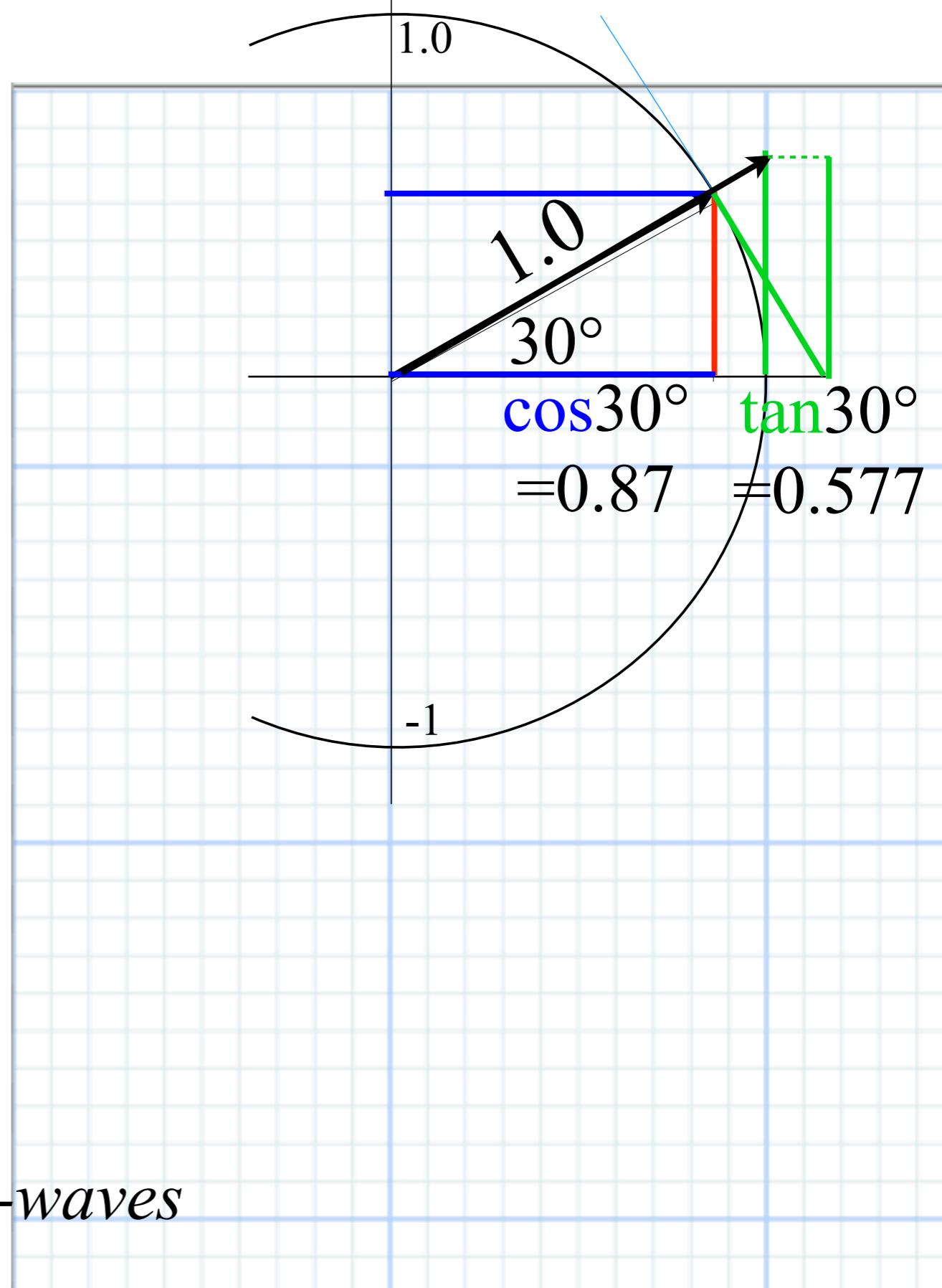
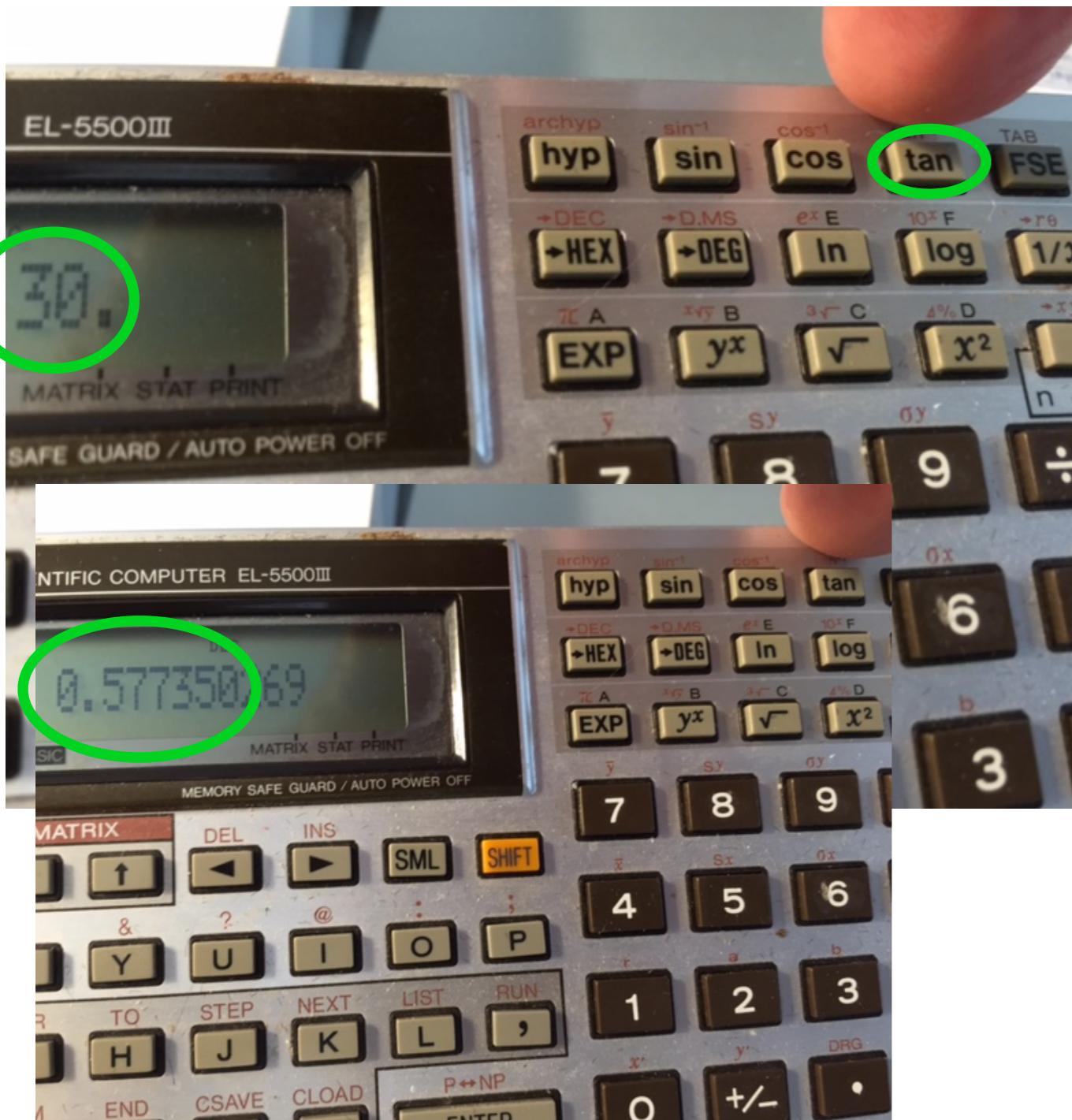


Learning about SIN and the COSin and TANgent “Slope of INcline” “COmplimentary Slope”



It's mostly about triangles and sine-waves

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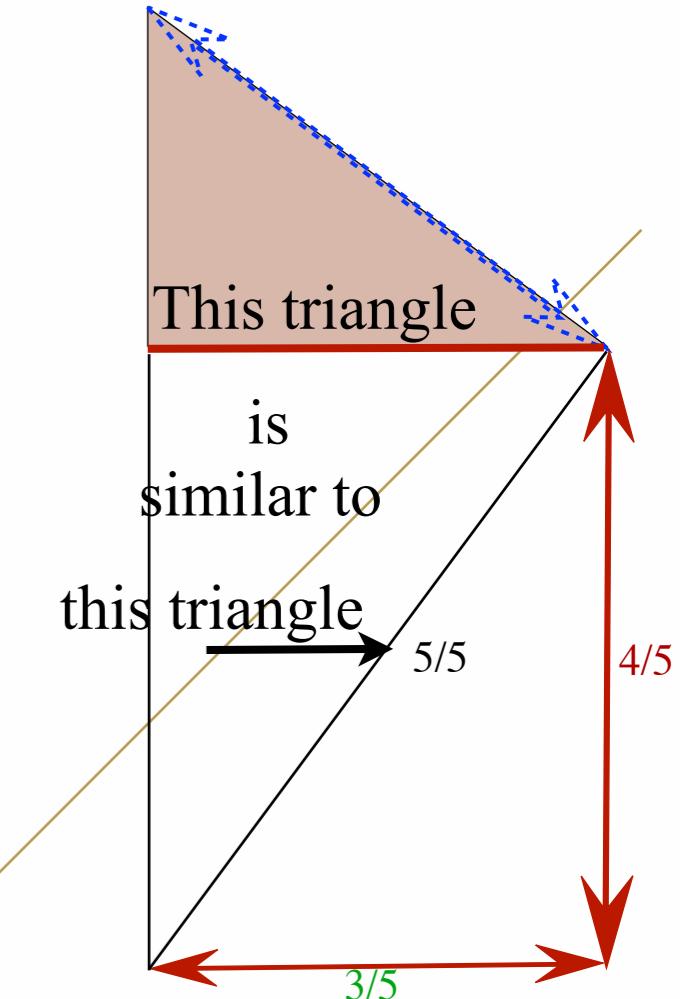
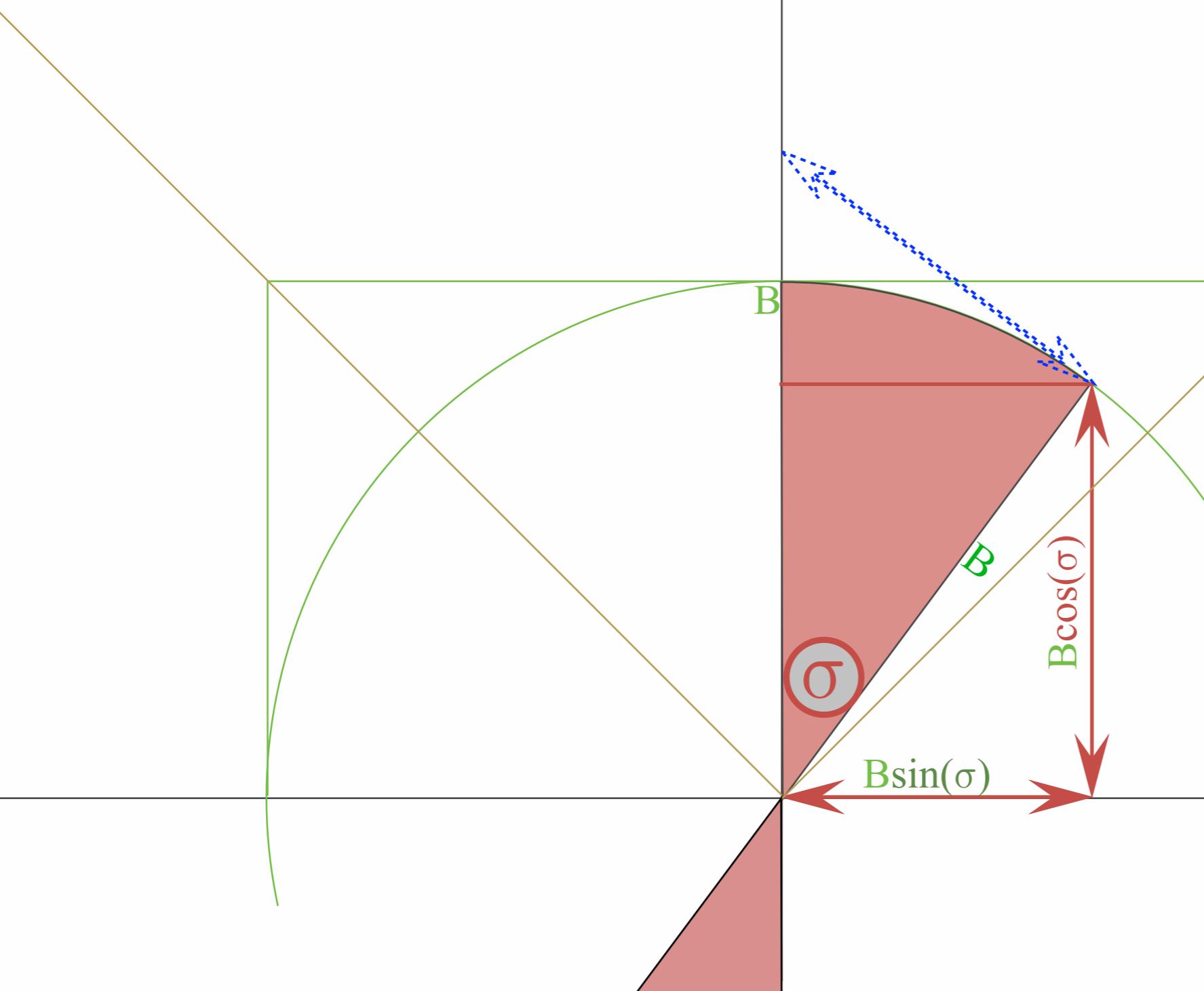
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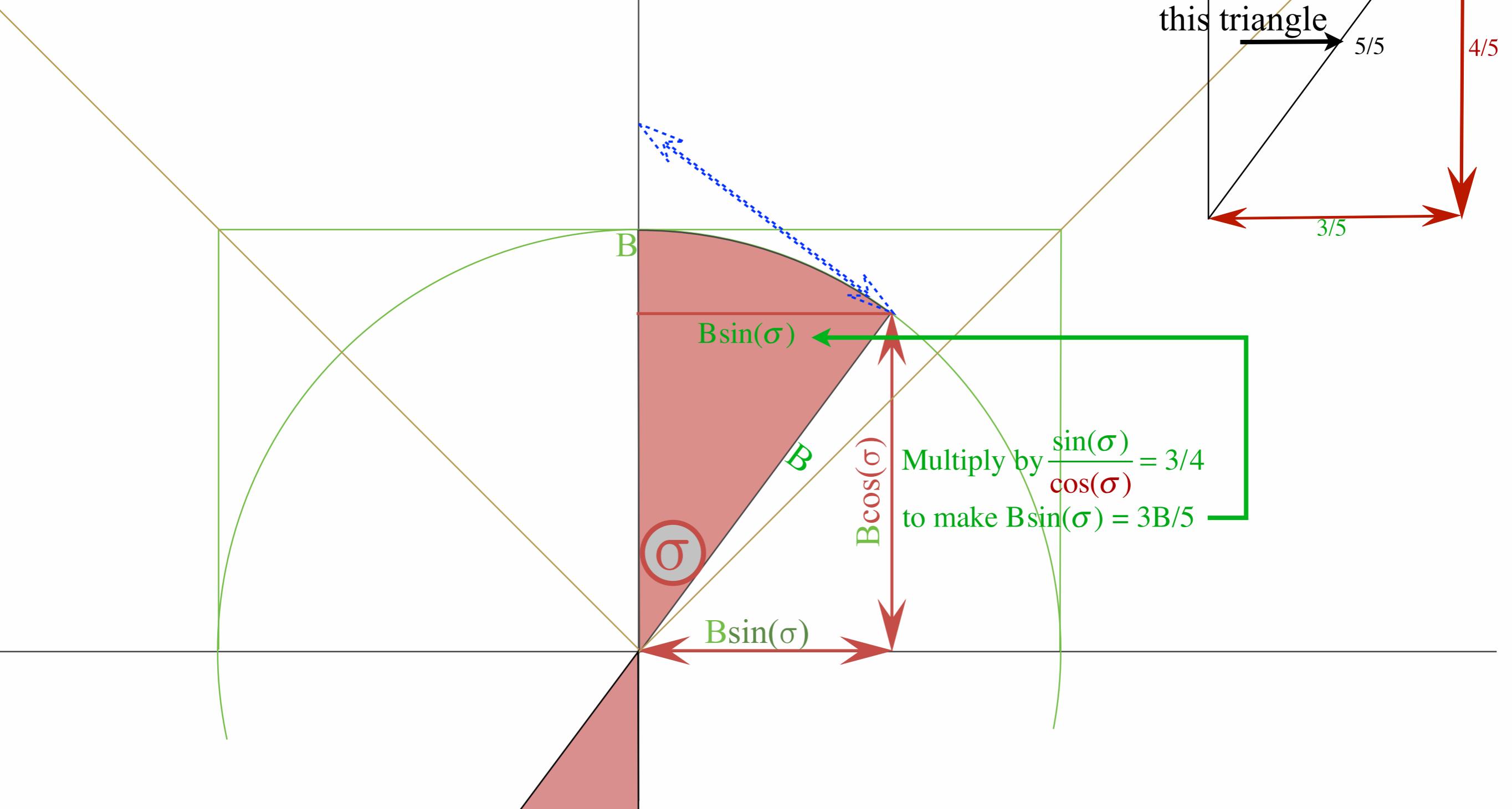


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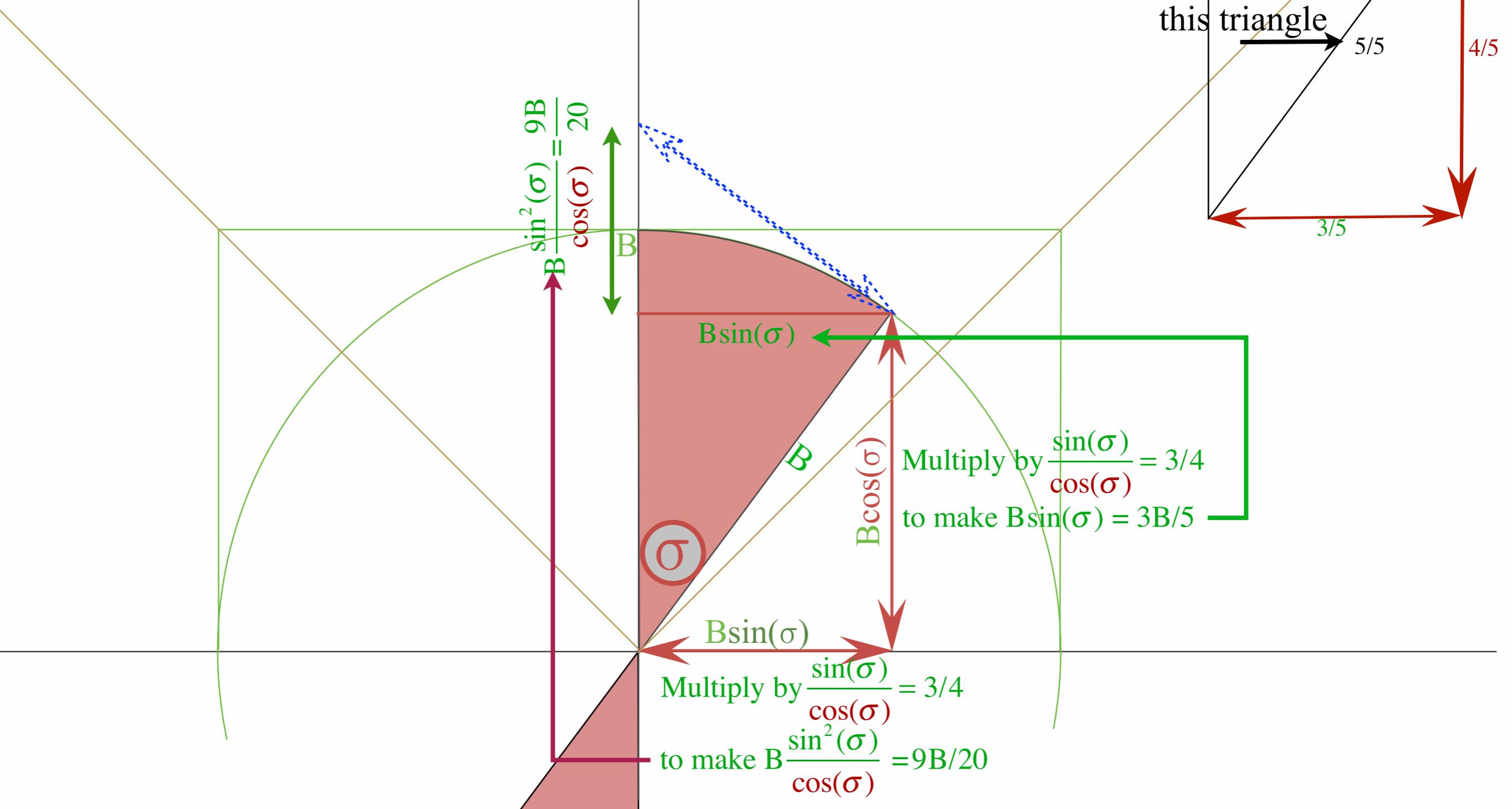


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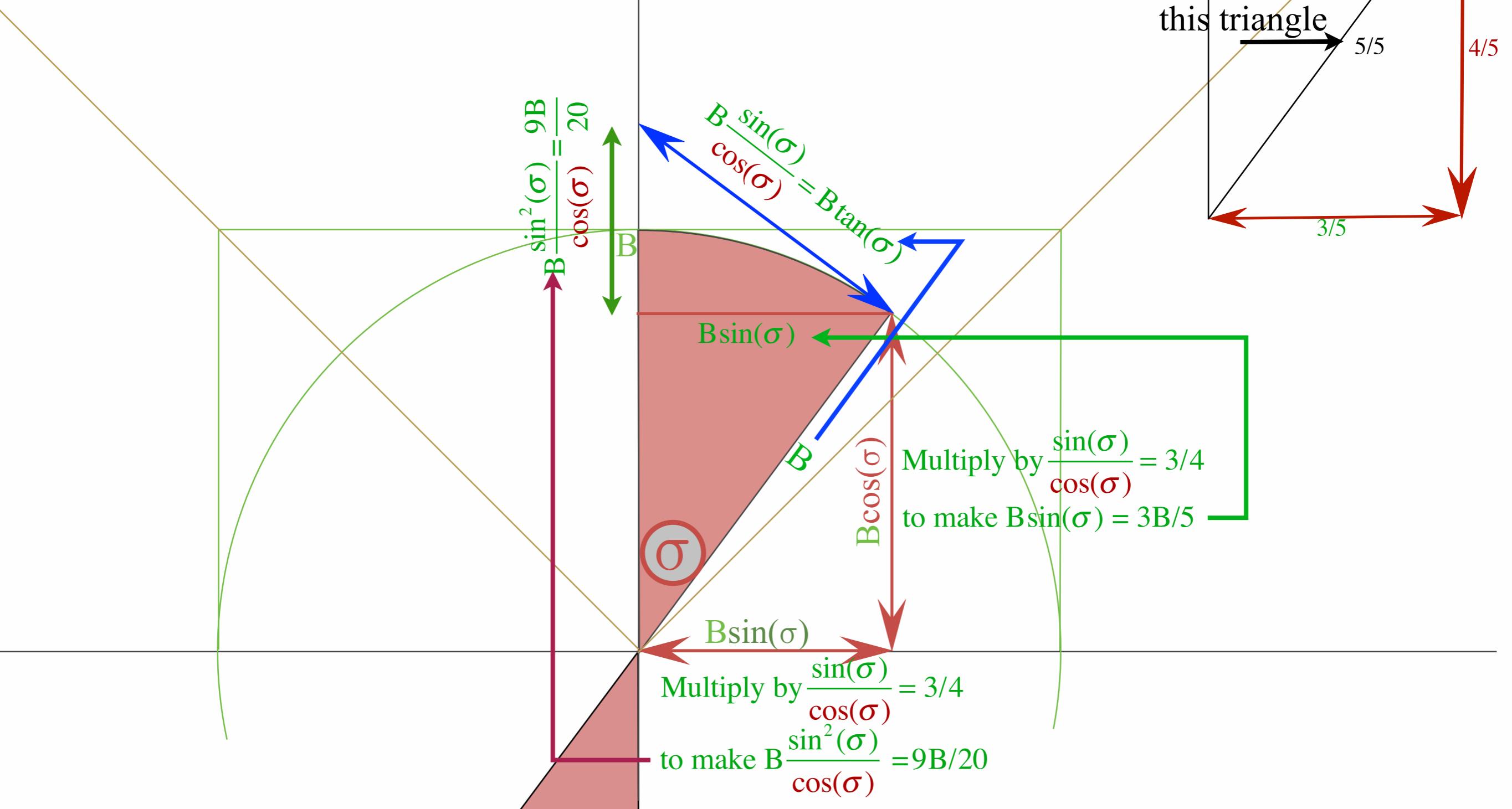


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$$\begin{aligned}\sin(\sigma) &= 0.6000 &= 3/5 \\ \tan(\sigma) &= 0.7500 &= 3/4\end{aligned}$$

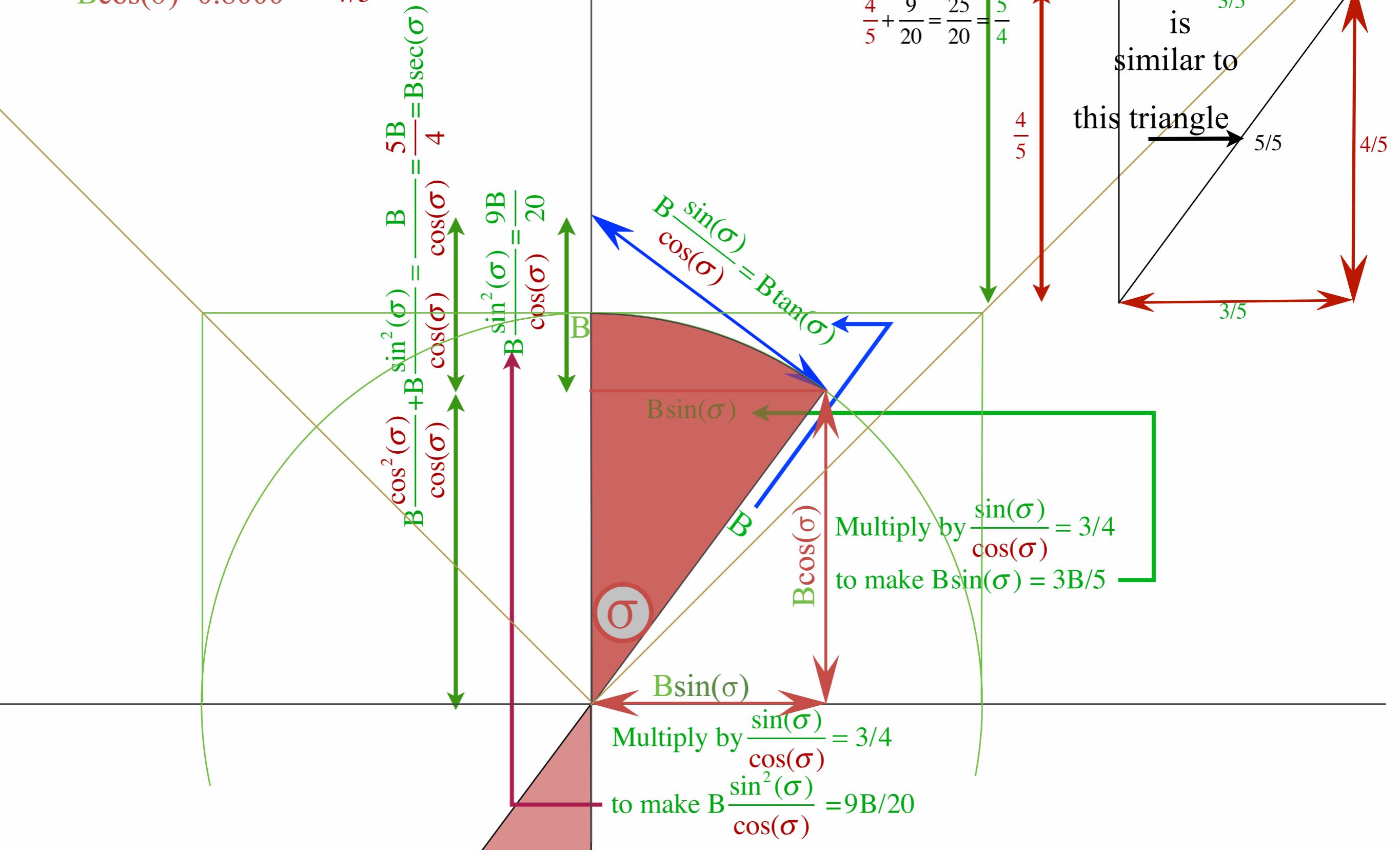
$$\cos(\sigma) = 0.8000 = 4/5$$



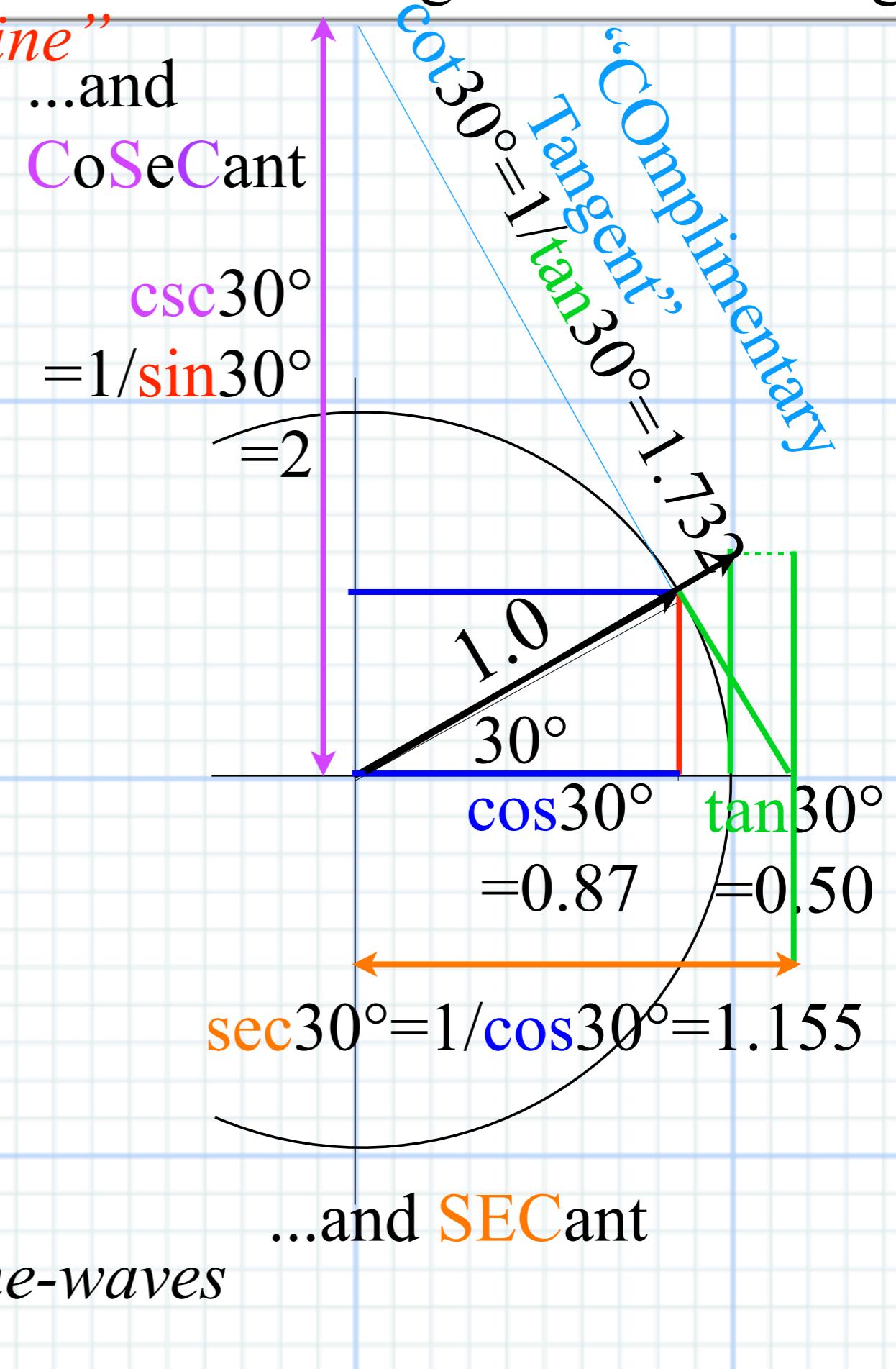
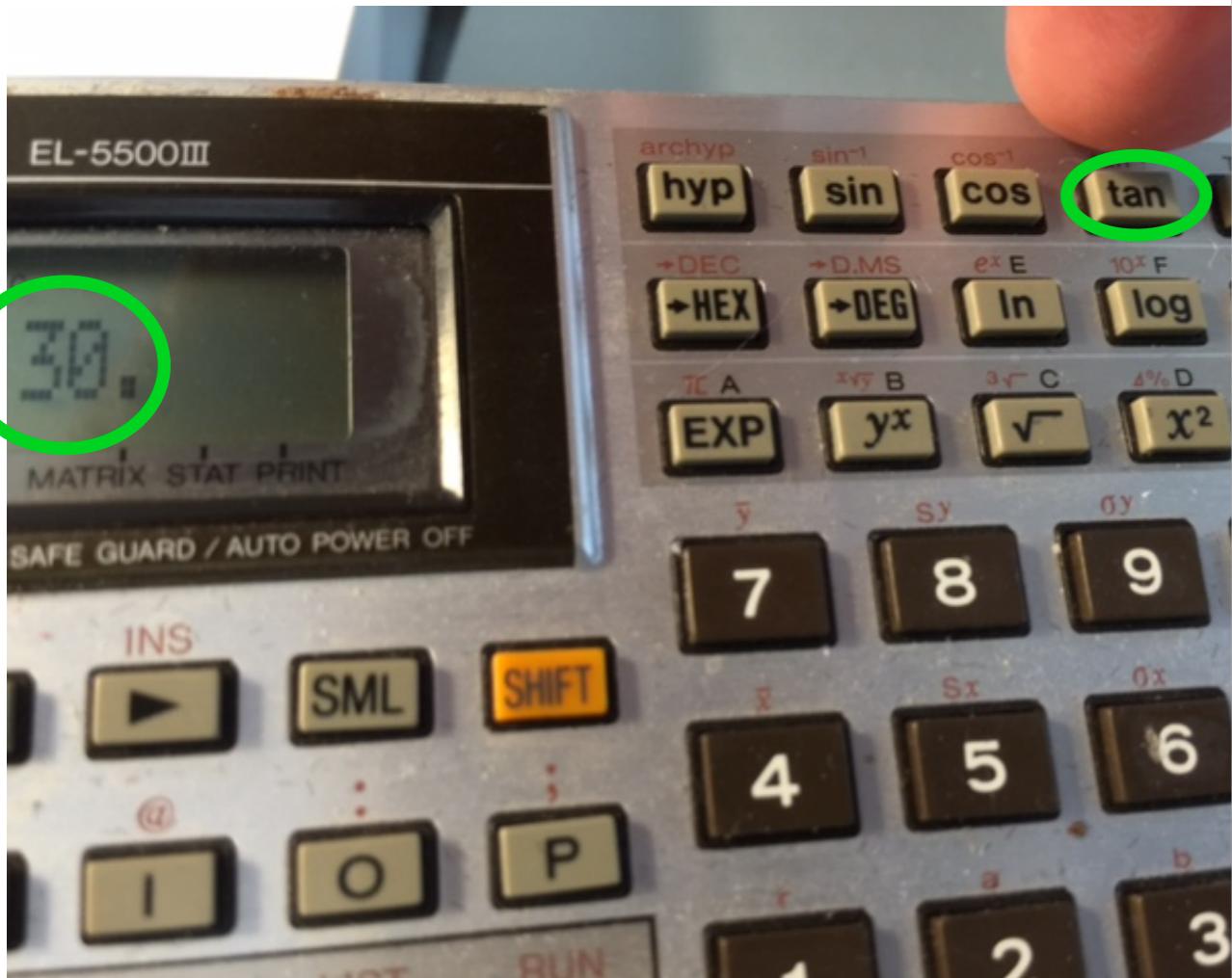
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$$\begin{aligned} B\sin(\sigma) &= 0.6000 & = 3/5 \\ \tan(\sigma) &= 0.7500 & = 3/4 \\ \sec(\sigma) &= 1.2500 & = 5/4 \\ B\cos(\sigma) &= 0.8000 & = 4/5 \end{aligned}$$



Learning about SIN and the COSin and TANgent and COTangent “Slope of INcline”

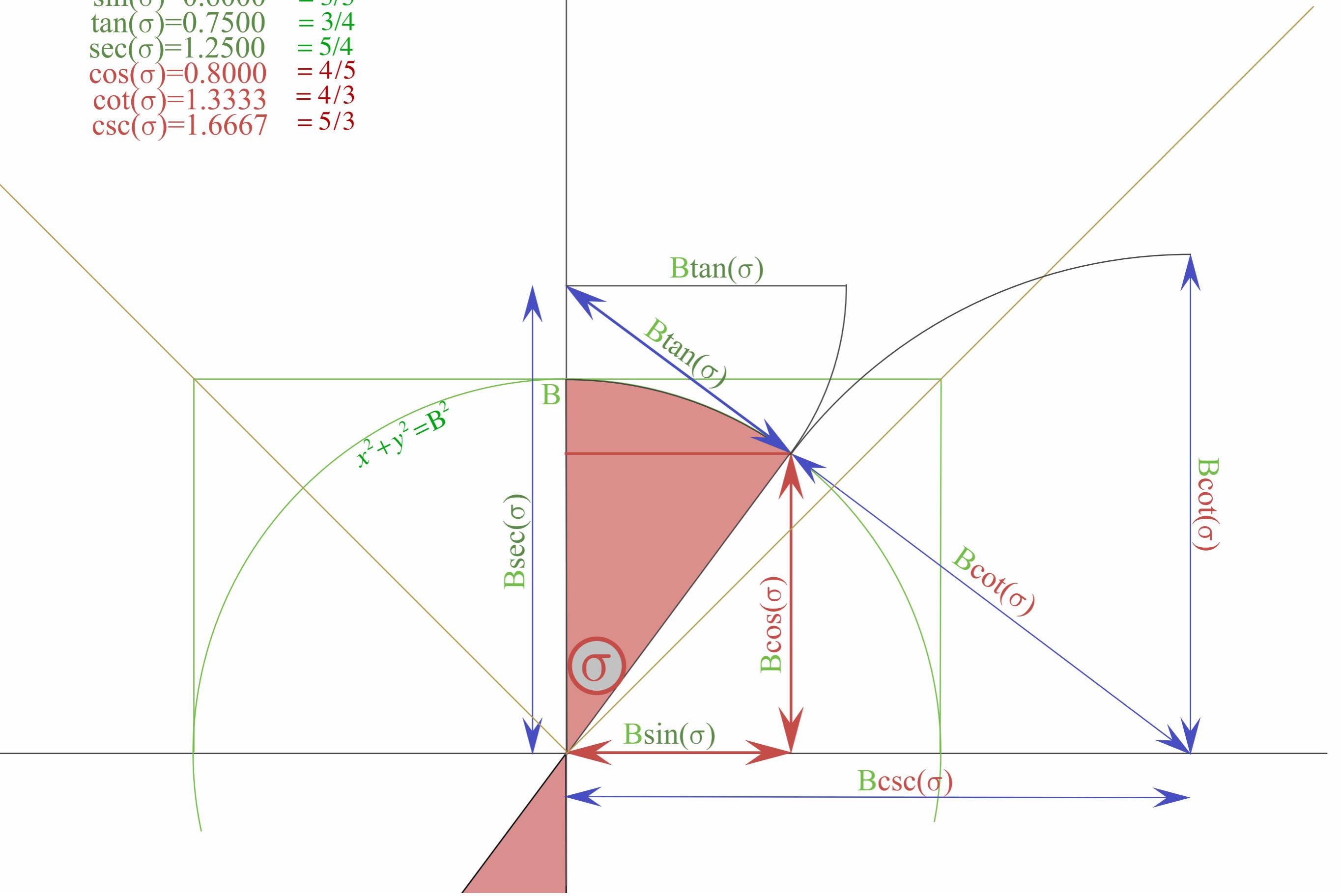


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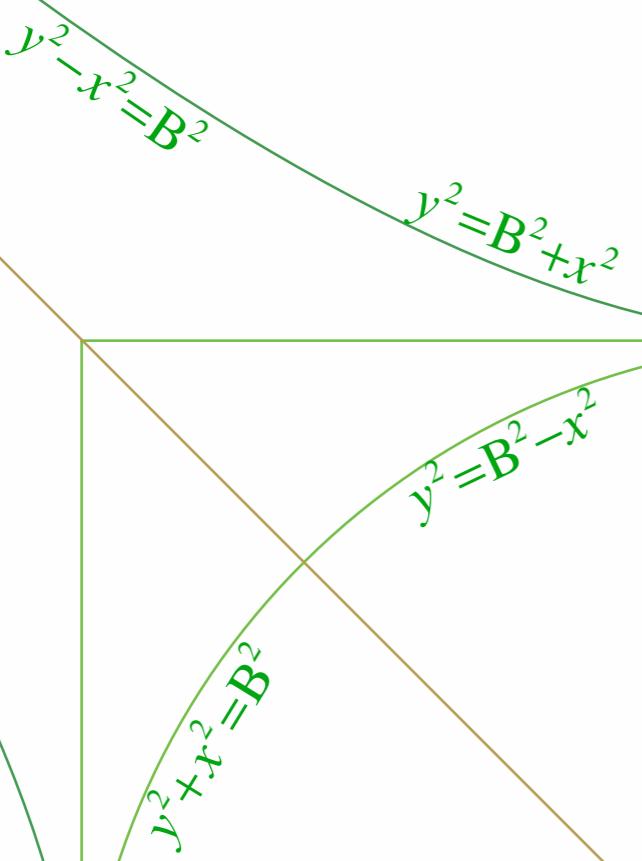
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$\cos(\sigma) = 0.8000$	$= 4/5$
$\cot(\sigma) = 1.3333$	$= 4/3$
$\csc(\sigma) = 1.6667$	$= 5/3$

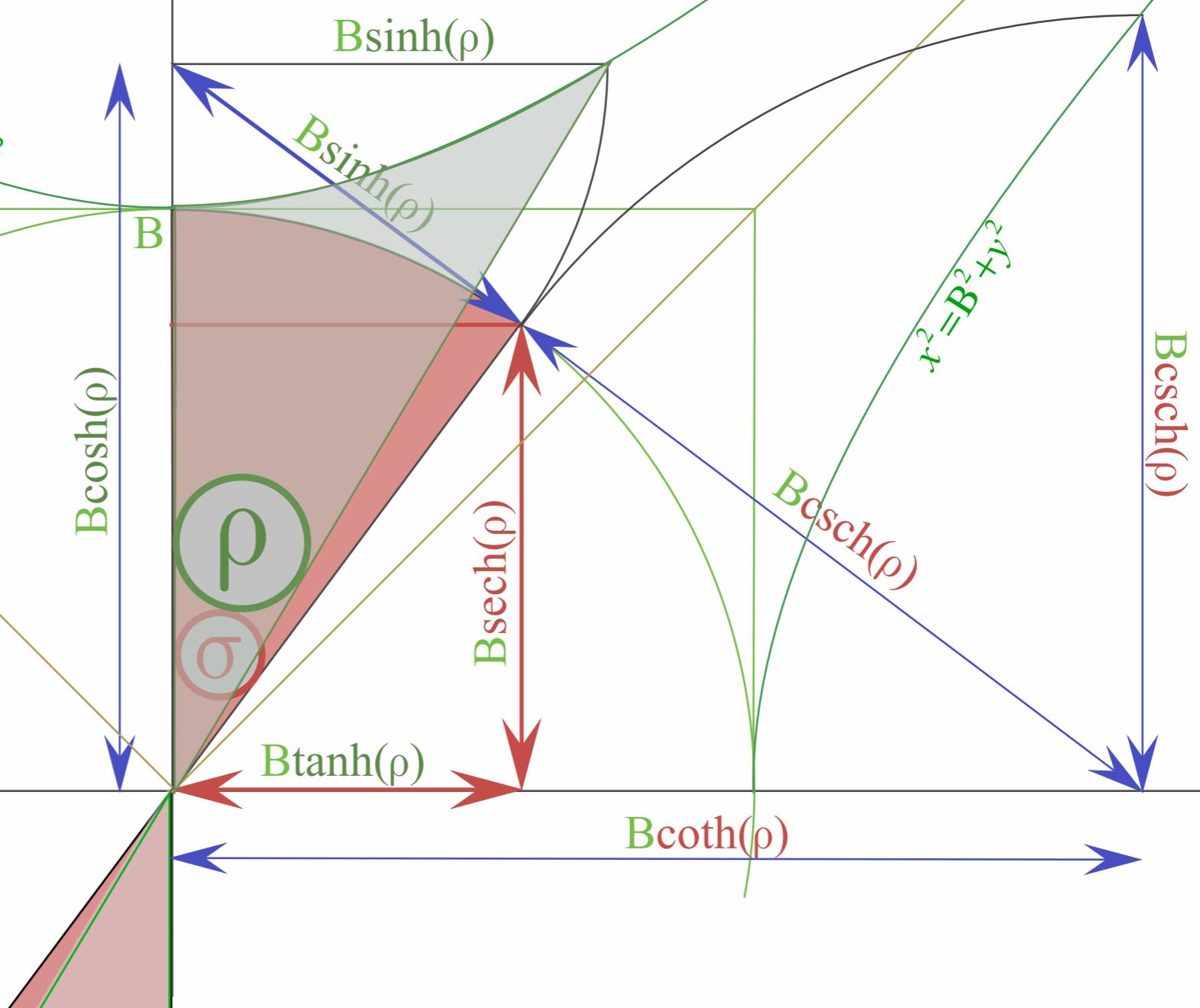


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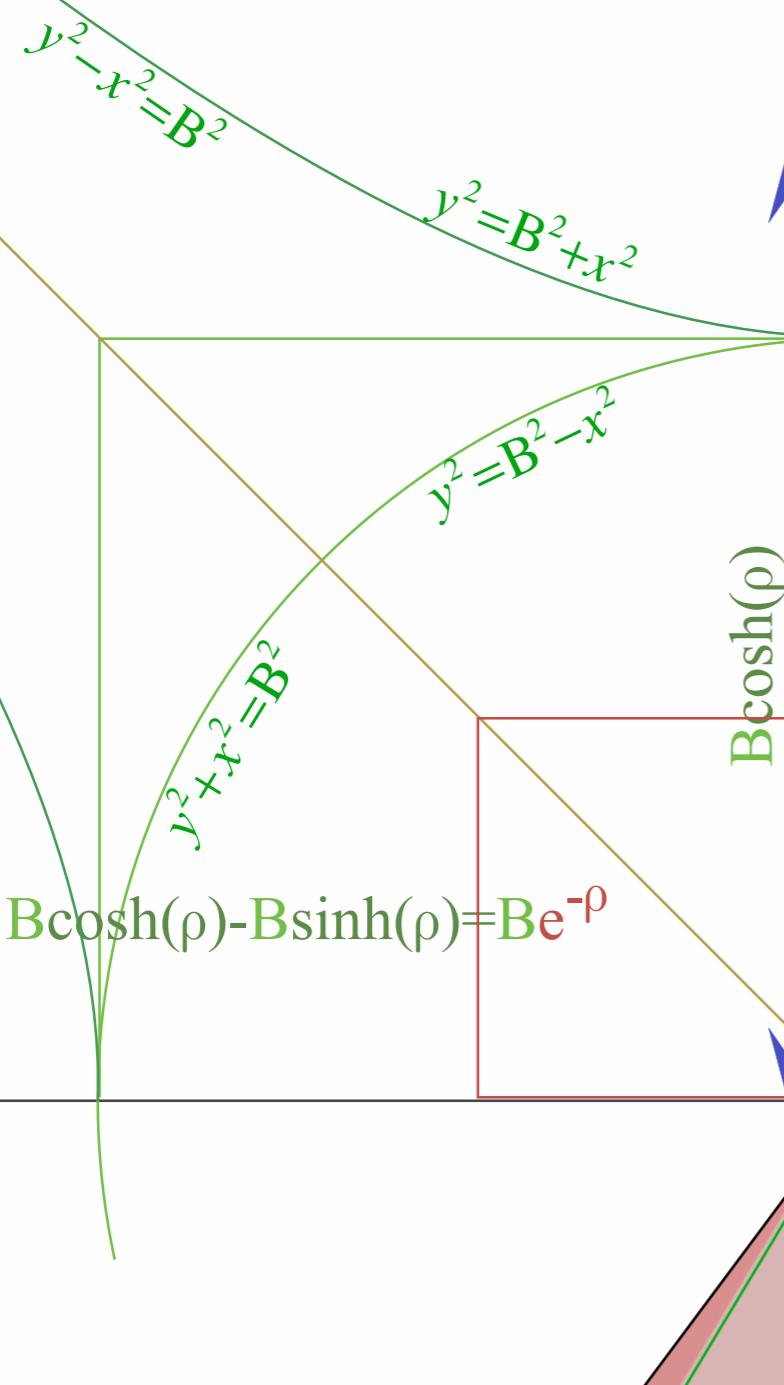


Also it is set to hyperbola sector arc-area $\rho = 0.6931$
angle $\angle\rho = v = 30.96^\circ$



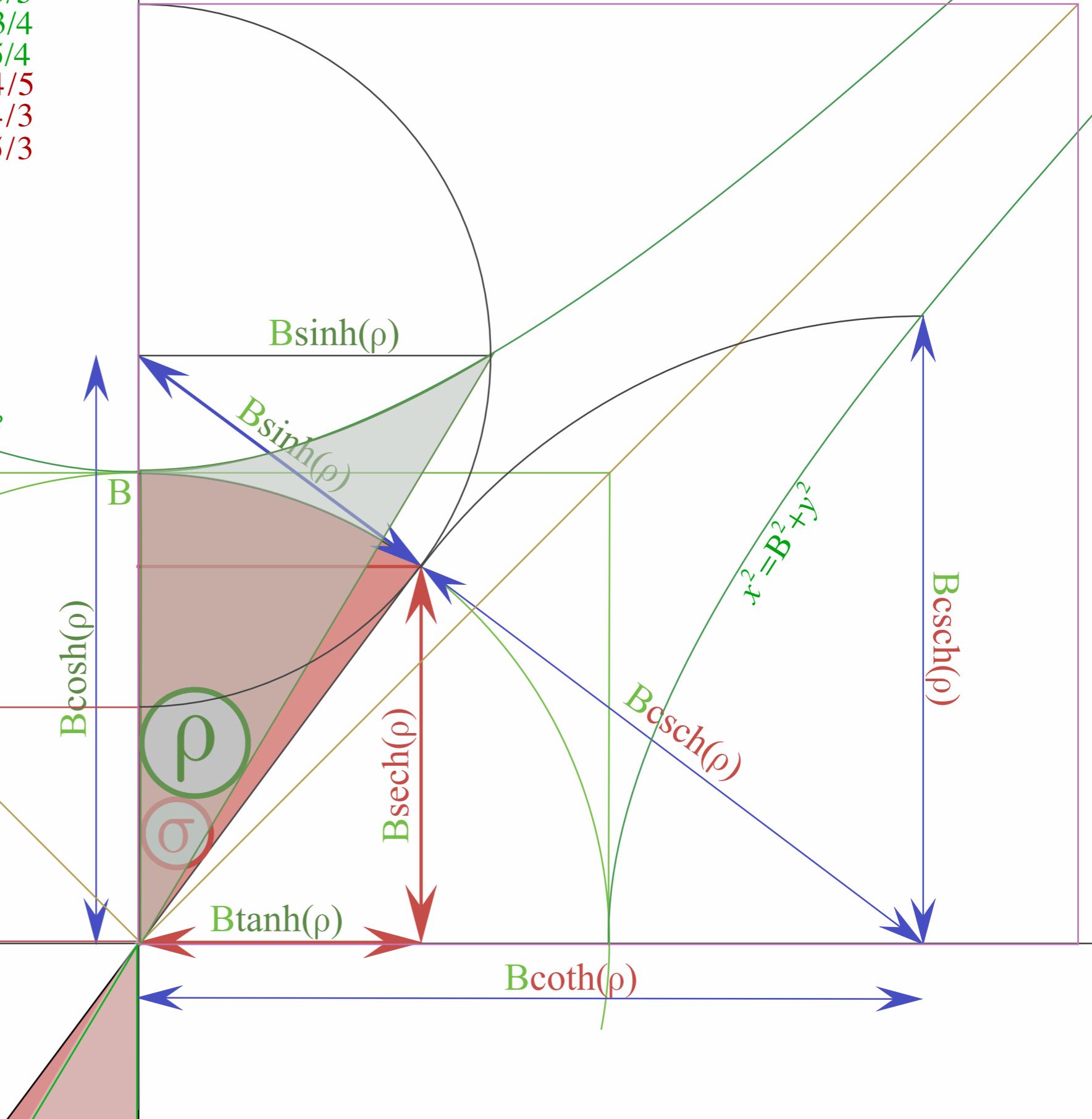
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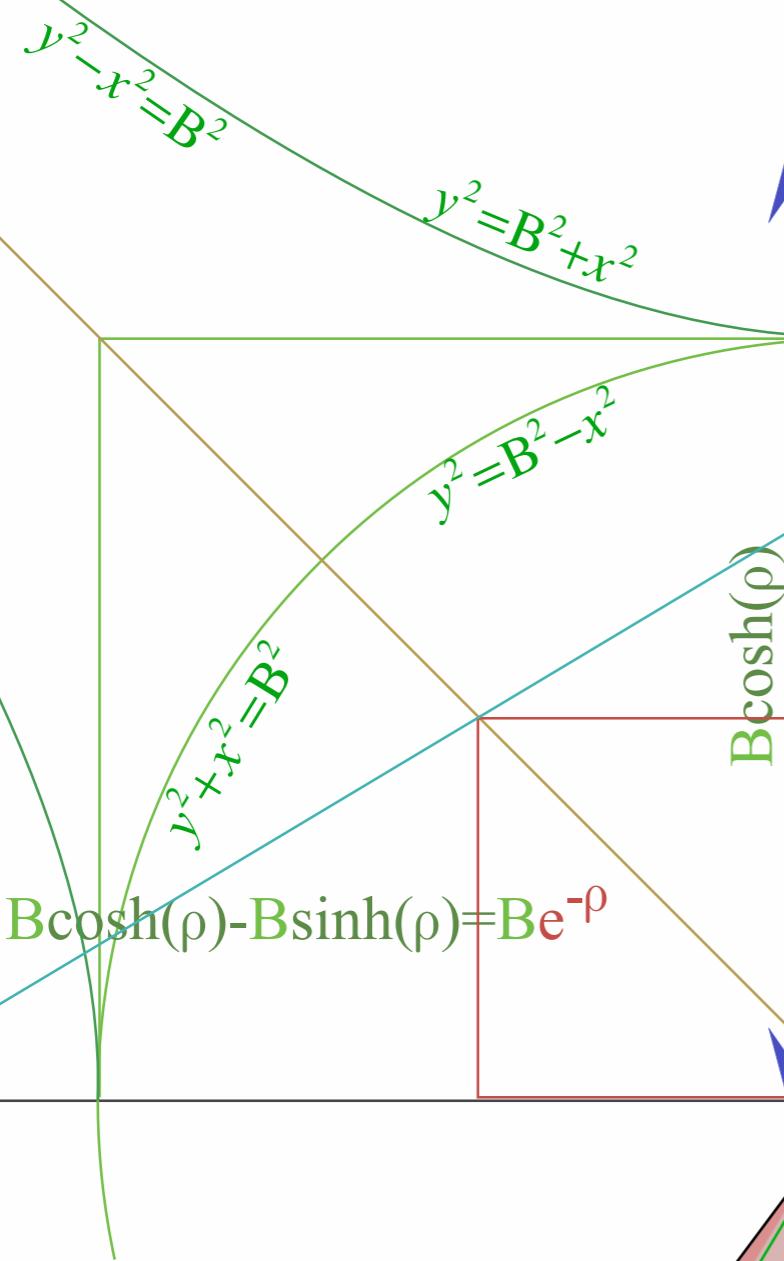
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$$B\cosh(\rho) + B\sinh(\rho) = Be^{+\rho}$$



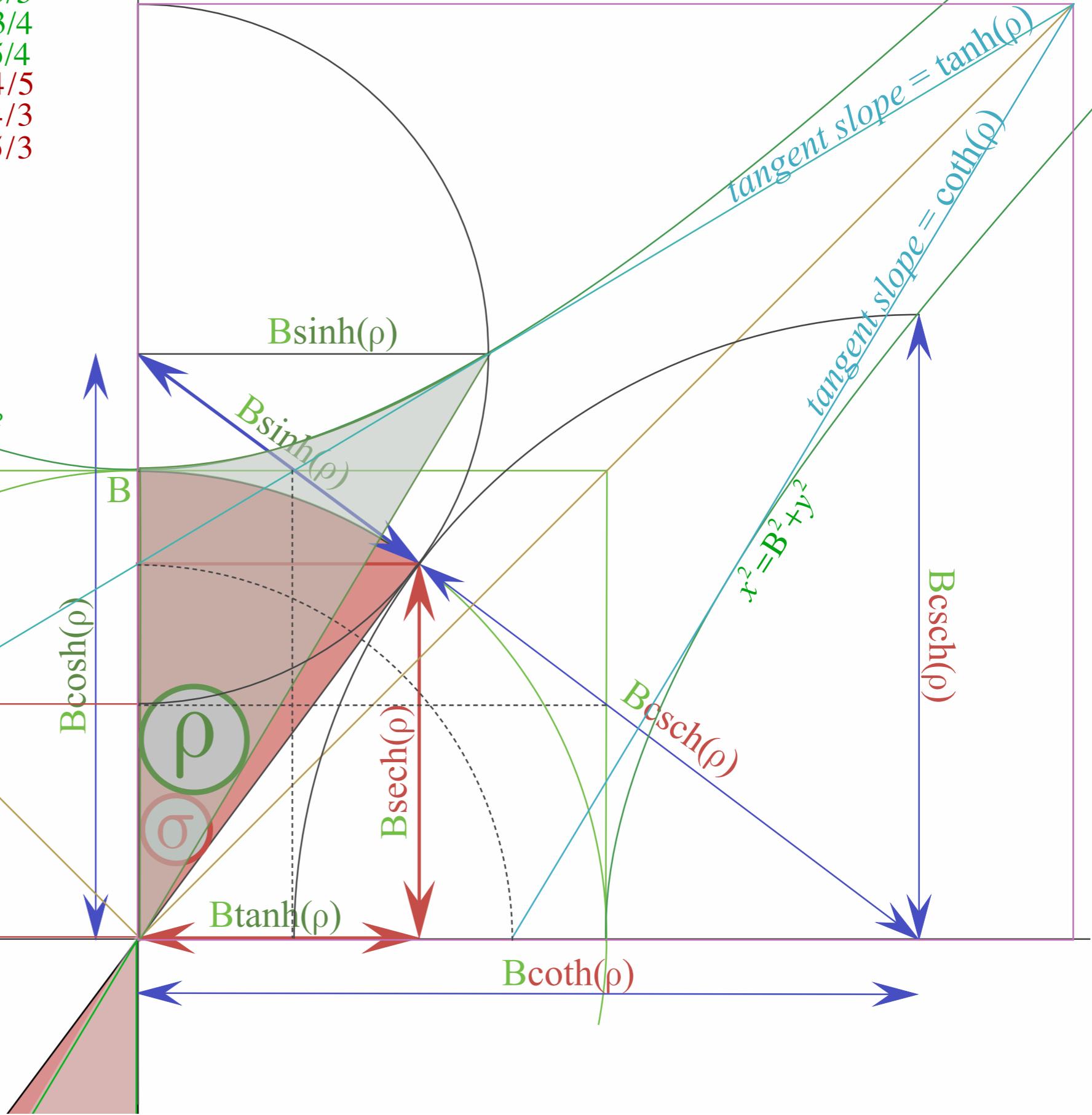
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$$\begin{aligned} \cosh(\rho) + \sinh(\rho) &= \frac{5}{4} + \frac{3}{4} = 2.0 = e^{+\rho} \\ \cosh(\rho) - \sinh(\rho) &= \frac{5}{4} - \frac{3}{4} = 1/2 = e^{-\rho} \end{aligned}$$

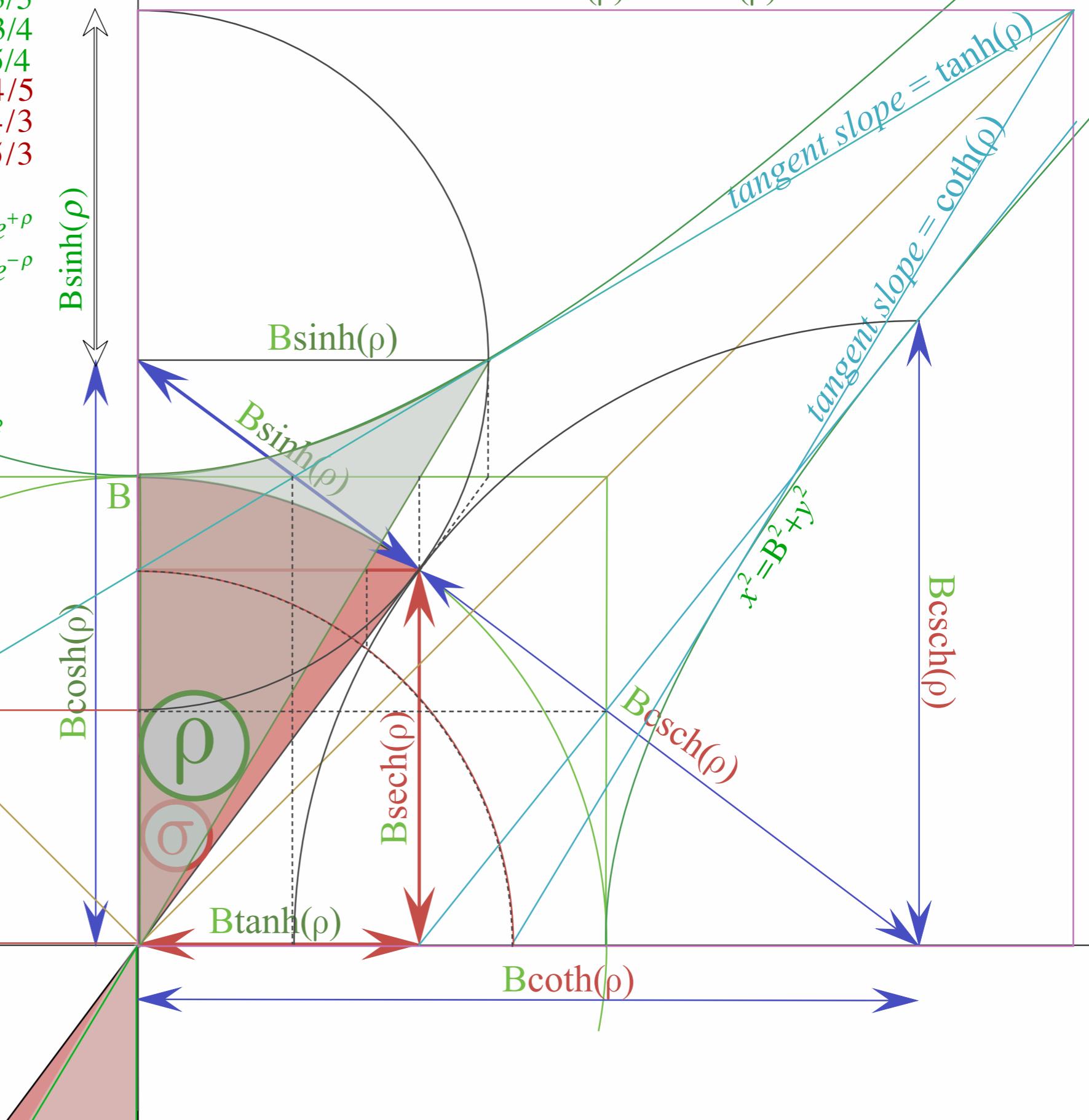
$$x^2 + y^2 = B^2$$

$$B\cosh(\rho) - B\sinh(\rho) = Be^{-\rho}$$

$$y^2 = B^2 + x^2$$

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$$\begin{aligned} \cosh(\rho) &= \frac{e^{+\rho} + e^{-\rho}}{2} & \text{Half-Sum-} \\ \sinh(\rho) &= \frac{e^{+\rho} - e^{-\rho}}{2} & \text{Half-Difference} \\ && \text{Trig-Formulae for} \\ && \text{exponentials } e^{\pm\rho} \end{aligned}$$

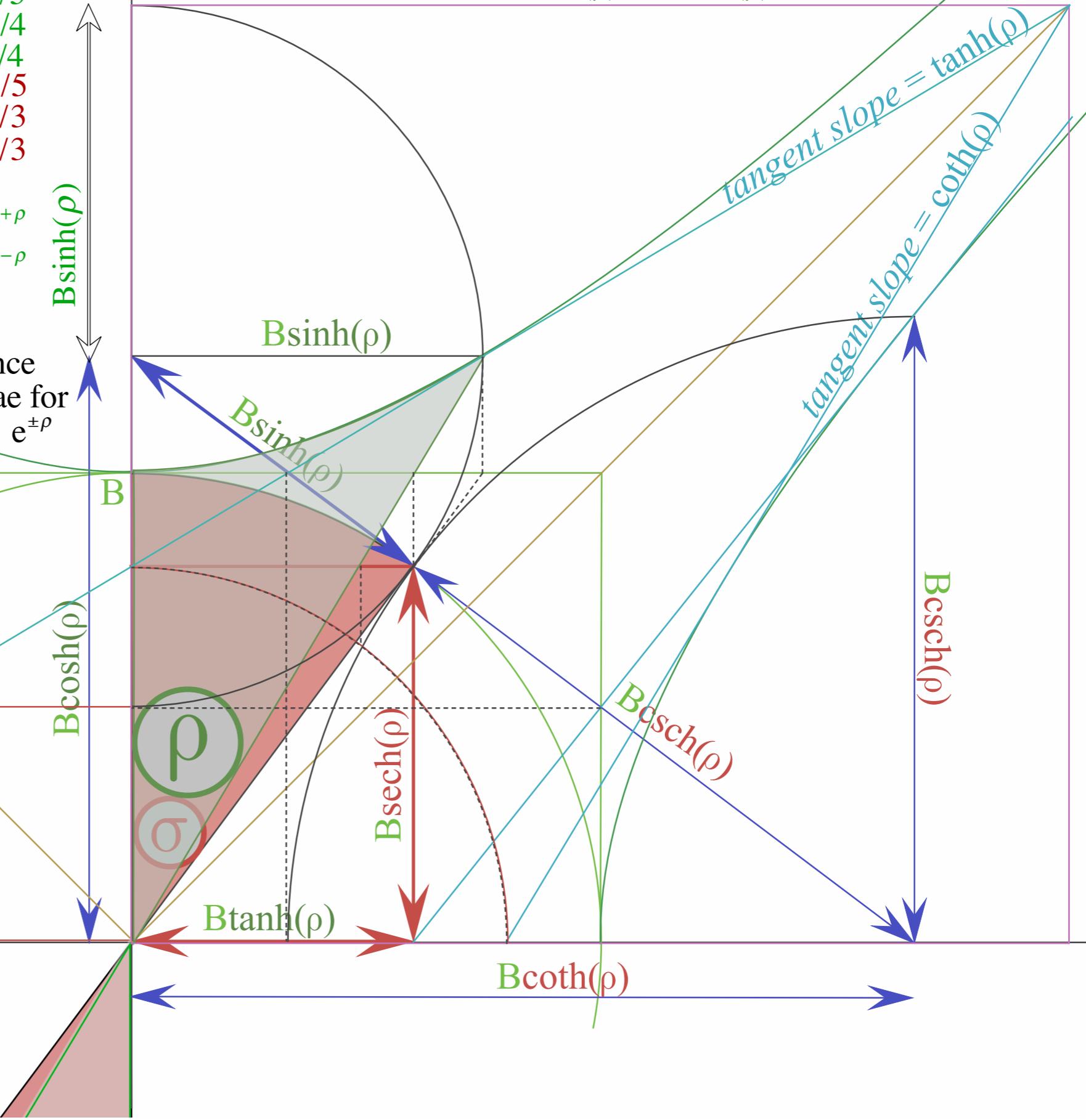
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$$B\cosh(\rho) + B\sinh(\rho) = Be^{+\rho}$$

$$\begin{aligned} \text{tangent slope} &= \tanh(\rho) \\ \text{tangent slope} &= \coth(\rho) \end{aligned}$$



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$$x^2 - y^2 = B^2$$

$$\operatorname{Bcosh}(\rho) - \operatorname{Bsinh}(\rho) = Be^{-\rho}$$

Thales Circle
(links $e^{-\rho}$ to $e^{+\rho}$)

Also it is set to hyperbola sector arc-area $\rho = 0.6931$
angle $\angle\rho = v = 30.96^\circ$

$$\operatorname{Bcosh}(\rho) + \operatorname{Bsinh}(\rho) = Be^{+\rho}$$

$$\operatorname{tangent slope} = \tanh(\rho)$$

$$\operatorname{tangent slope} = \coth(\rho)$$

$$\operatorname{tangent slope} = \operatorname{coth}(\rho)$$

$$\operatorname{Bcsc}(\rho)$$

$$\operatorname{Bcsch}(\rho)$$

$$\operatorname{Bcoth}(\rho)$$

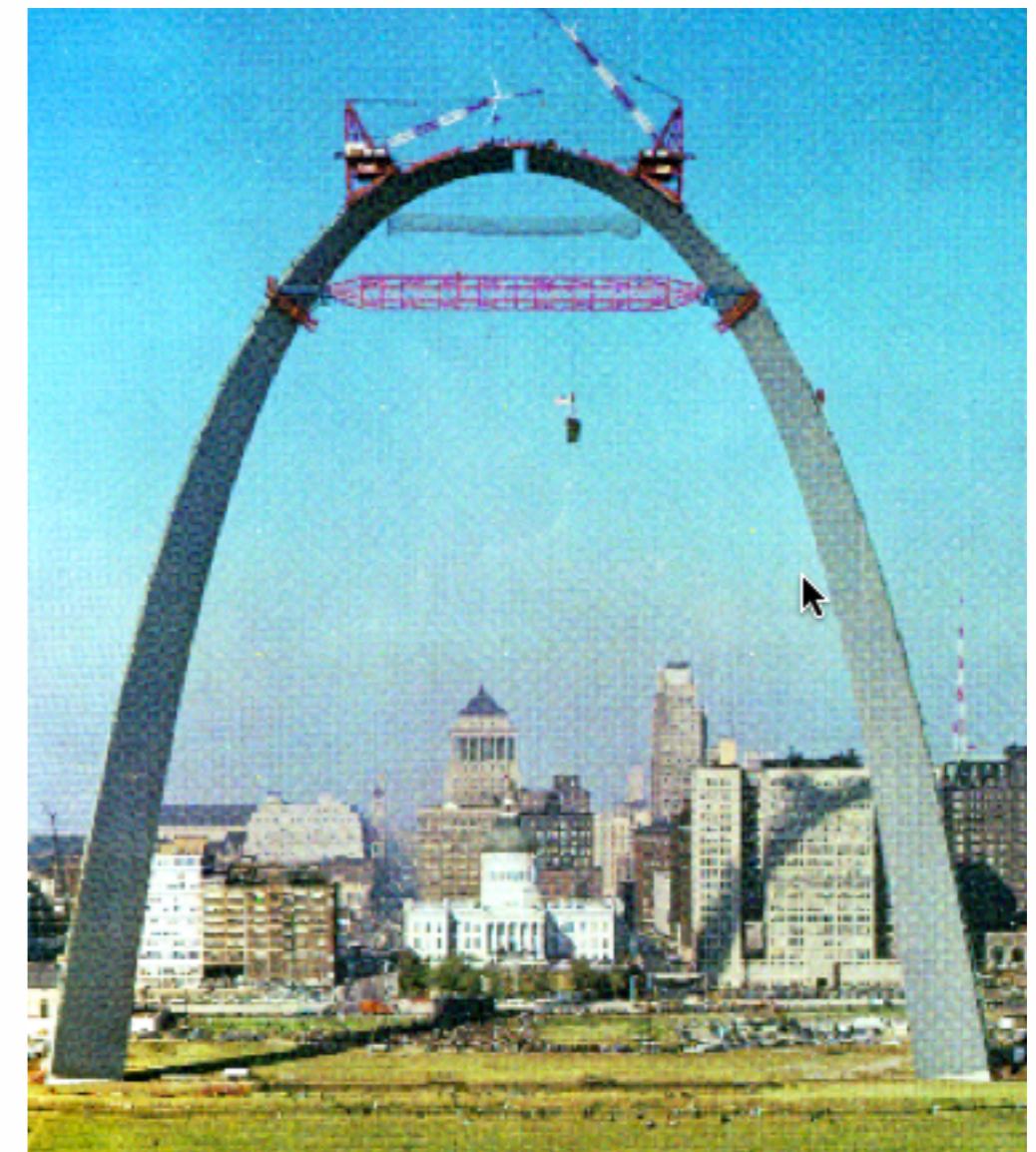
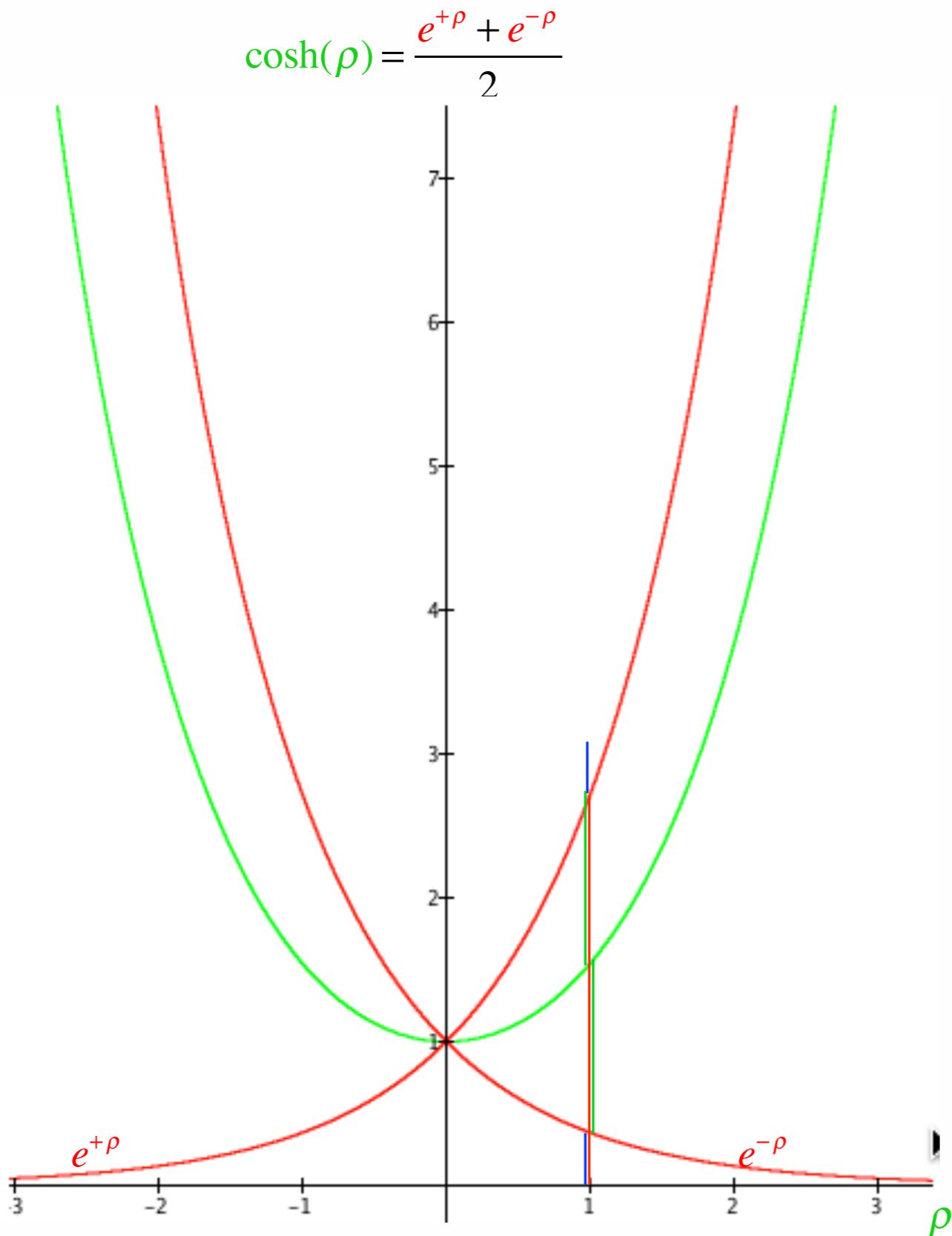
$$\operatorname{Bsech}(\rho)$$

$$\operatorname{Btanh}(\rho)$$

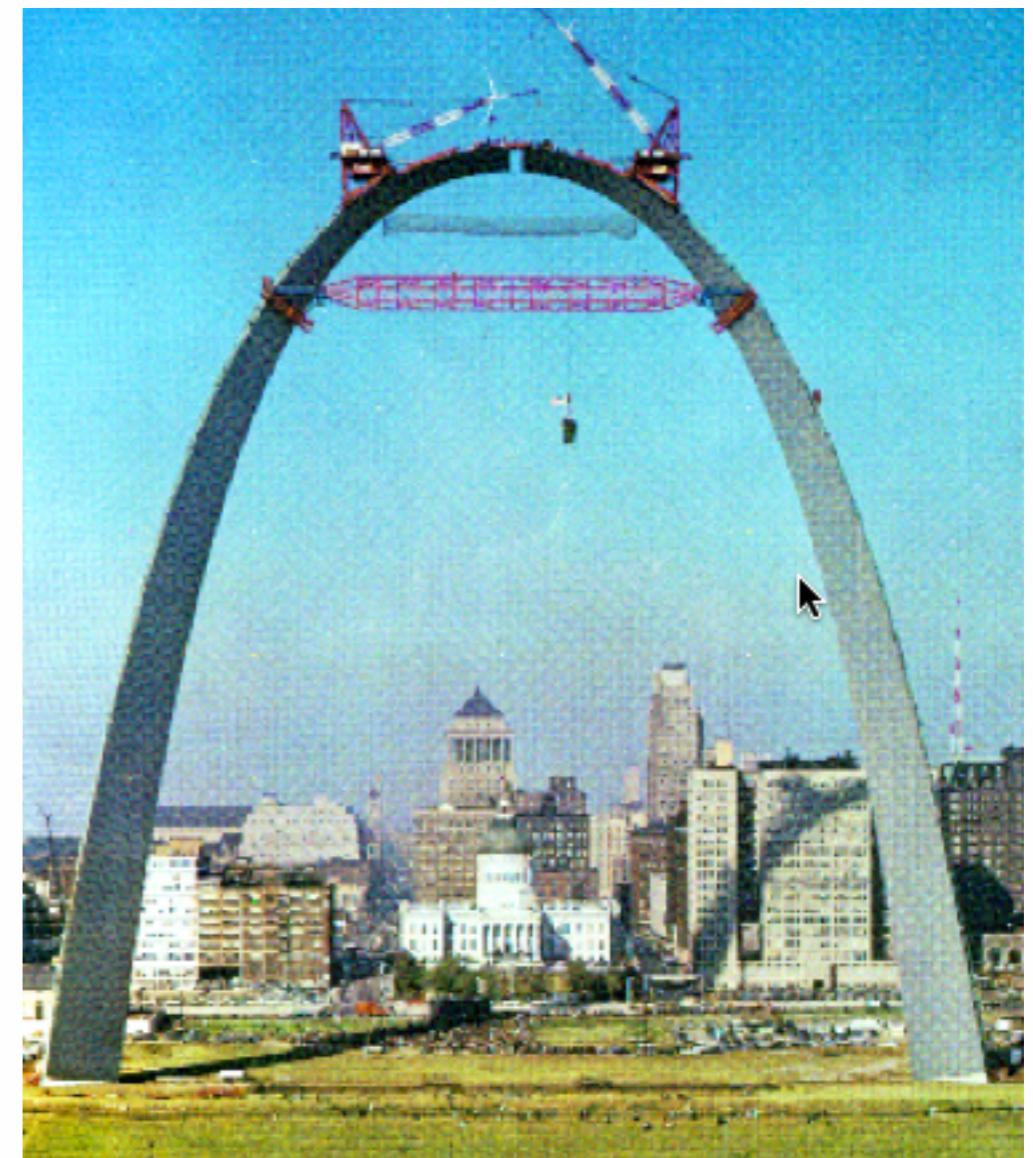
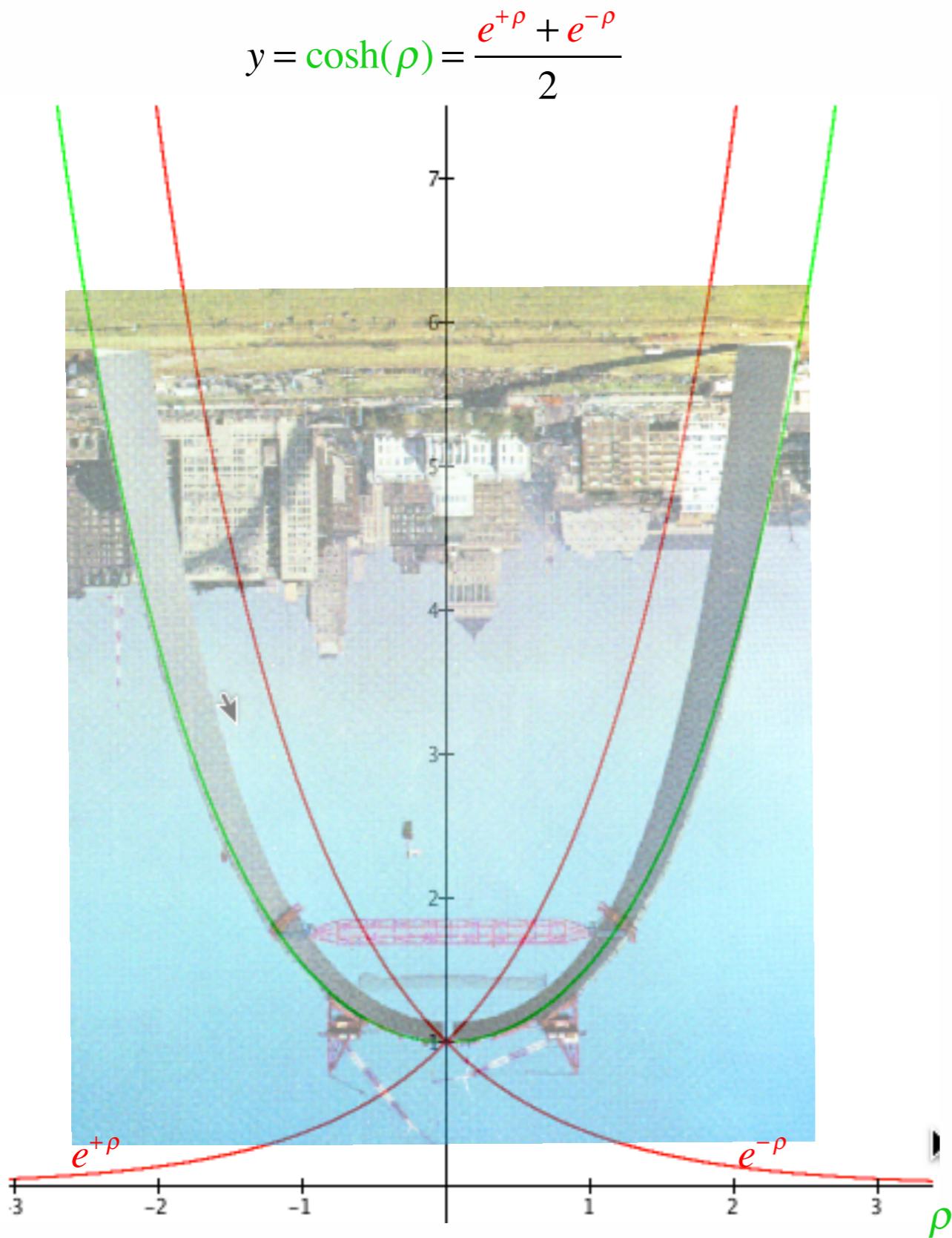
$$\operatorname{Bsinh}(\rho)$$

$$\operatorname{Bcosh}(\rho)$$

Hyperbolic cosine and the St. Louis Arch (being topped out in 1963)



Hyperbolic cosine and the St. Louis Arch (being topped out in 1963)



Formulae for Exponentials $e^{\pm\rho}$

$$\cosh(\rho) = \frac{e^{+\rho} + e^{-\rho}}{2}$$

$$\sinh(\rho) = \frac{e^{+\rho} - e^{-\rho}}{2}$$

$$\cosh(\rho) + \sinh(\rho) = e^{+\rho}$$

$$\cosh(\rho) - \sinh(\rho) = e^{-\rho}$$

Half-Sum-
Half-Difference

Trig-Formulae for
exponentials $e^{\pm\rho}$

Formulae for Exponentials $e^{\pm\rho}$ begin with interest-rate formula $e^{rt} = \lim_{n \rightarrow \infty} \left(1 + \frac{rt}{n}\right)^n$

$$\cosh(\rho) = \frac{e^{+\rho} + e^{-\rho}}{2}$$

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Half-Sum-
Half-Difference

Trig-Formulae for
exponentials $e^{\pm\rho}$

...and its binomial expansion series for exponentials...

$$e^{rt} = 1 + rt + \frac{(rt)^2}{2} + \frac{(rt)^3}{2 \cdot 3} + \frac{(rt)^4}{2 \cdot 3 \cdot 4} + \frac{(rt)^5}{2 \cdot 3 \cdot 4 \cdot 5} + \frac{(rt)^6}{2 \cdot 3 \cdot 4 \cdot 5 \cdot 6} + \dots$$

Formulae for Exponentials $e^{\pm\rho}$ begin with interest-rate formula $e^{rt} = \lim_{n \rightarrow \infty} \left(1 + \frac{rt}{n}\right)^n$

$$\cosh(\rho) = \frac{e^{+\rho} + e^{-\rho}}{2}$$

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Half-Sum-
Half-Difference
Trig-Formulae for
exponentials $e^{\pm\rho}$

...and its binomial expansion series for exponentials...

$$e^{rt} = 1 + rt + \frac{(rt)^2}{2} + \frac{(rt)^3}{2 \cdot 3} + \frac{(rt)^4}{2 \cdot 3 \cdot 4} + \frac{(rt)^5}{2 \cdot 3 \cdot 4 \cdot 5} + \frac{(rt)^6}{2 \cdot 3 \cdot 4 \cdot 5 \cdot 6} + \dots$$

...with rate r sign-flipped on odd powers (*decay-rate*)...

$$e^{-rt} = 1 - rt + \frac{(rt)^2}{2} - \frac{(rt)^3}{2 \cdot 3} + \frac{(rt)^4}{2 \cdot 3 \cdot 4} - \frac{(rt)^5}{2 \cdot 3 \cdot 4 \cdot 5} + \frac{(rt)^6}{2 \cdot 3 \cdot 4 \cdot 5 \cdot 6} - \dots$$

Formulae for Exponentials $e^{\pm\rho}$ begin with interest-rate formula $e^{rt} = \lim_{n \rightarrow \infty} \left(1 + \frac{rt}{n}\right)^n$

$$\cosh(\rho) = \frac{e^{+\rho} + e^{-\rho}}{2}$$

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$$\cosh(\rho) + \sinh(\rho) = e^{+\rho}$$

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Half-Sum-Half-Difference
Trig-Formulae for exponentials $e^{\pm\rho}$

...and its binomial expansion series for exponentials...

$$e^{rt} = 1 + rt + \frac{(rt)^2}{2} + \frac{(rt)^3}{2 \cdot 3} + \frac{(rt)^4}{2 \cdot 3 \cdot 4} + \frac{(rt)^5}{2 \cdot 3 \cdot 4 \cdot 5} + \frac{(rt)^6}{2 \cdot 3 \cdot 4 \cdot 5 \cdot 6} + \dots$$

...with rate r sign-flipped on odd powers (*decay-rate*)...

$$e^{-rt} = 1 - rt + \frac{(rt)^2}{2} - \frac{(rt)^3}{2 \cdot 3} + \frac{(rt)^4}{2 \cdot 3 \cdot 4} - \frac{(rt)^5}{2 \cdot 3 \cdot 4 \cdot 5} + \frac{(rt)^6}{2 \cdot 3 \cdot 4 \cdot 5 \cdot 6} - \dots$$

Half-sum has *even* powers of hyper-cosine ...

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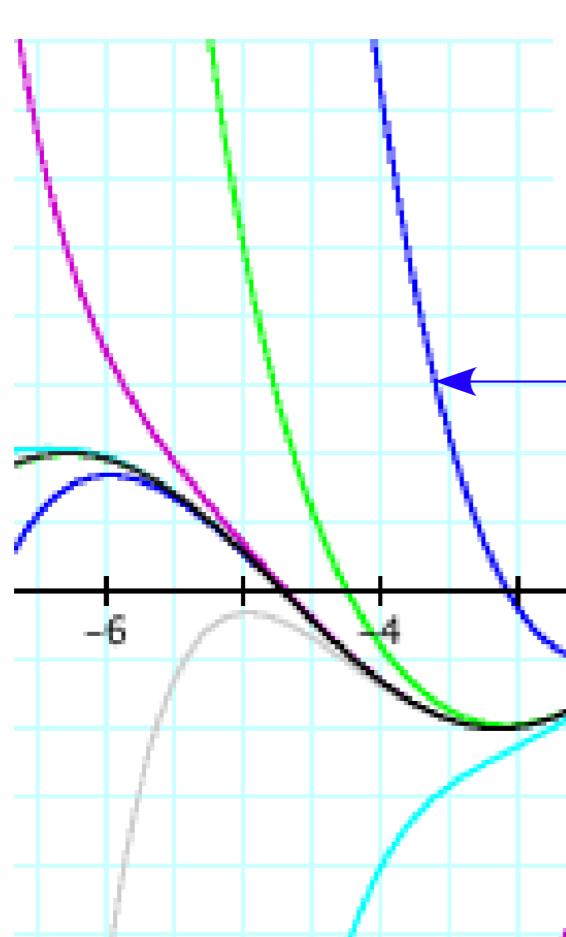
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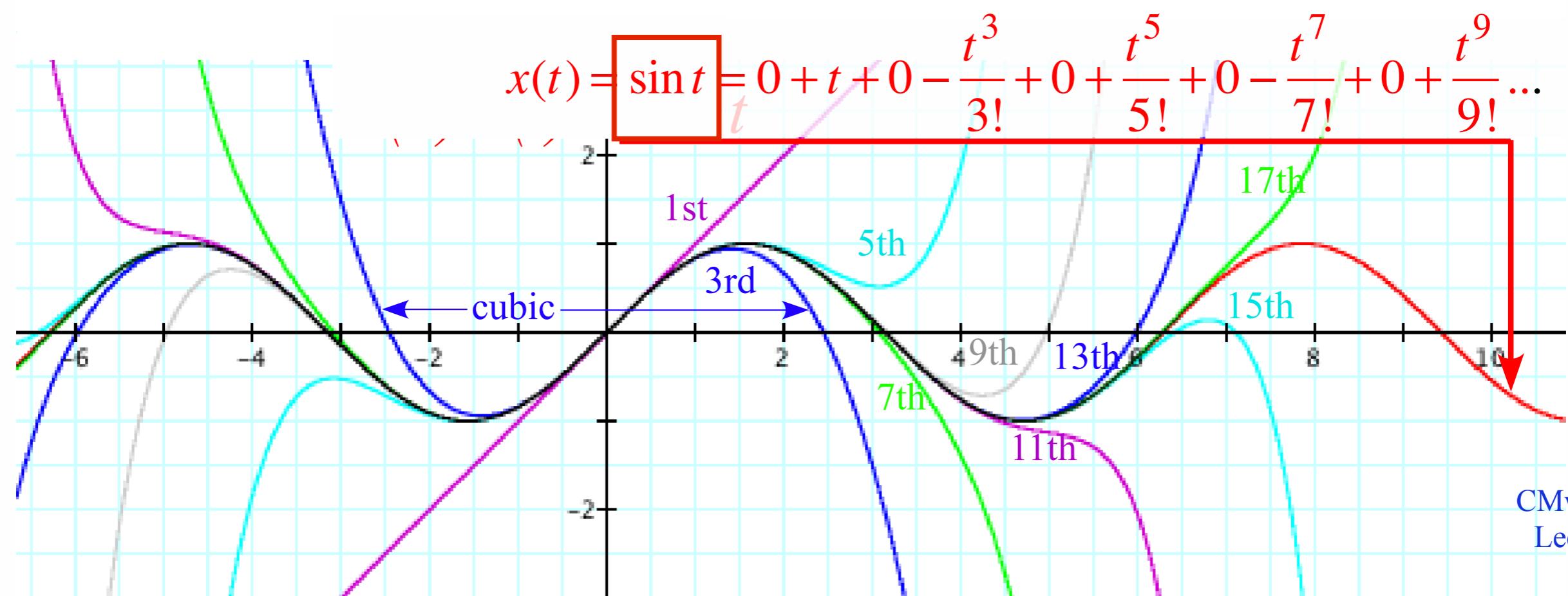
$$\begin{aligned}\frac{e^{+i rt} - e^{-i rt}}{2} &= i rt & - i \frac{(rt)^3}{2 \cdot 3} &+ i \frac{(rt)^5}{2 \cdot 3 \cdot 4 \cdot 5} - \dots &= i \sin rt \\ \frac{e^{+i rt} - e^{-i rt}}{2i} &= rt & - \frac{(rt)^3}{2 \cdot 3} &+ \frac{(rt)^5}{2 \cdot 3 \cdot 4 \cdot 5} - \dots &= \sin rt\end{aligned}$$



$$x(t) = \boxed{\cos t} = 1 + 0 - \frac{t^2}{2!} + 0 + \frac{t^4}{4!} + 0 - \frac{t^6}{6!} + 0 + \frac{t^8}{8!} \dots$$

quartic

quadratic
(parabola)



or
CMwBANG!
Lecture 12
p.22

Formulae for Exponentials $e^{\pm\rho}$ begin with interest-rate formula $e^{rt} = \lim_{n \rightarrow \infty} \left(1 + \frac{rt}{n}\right)^n$

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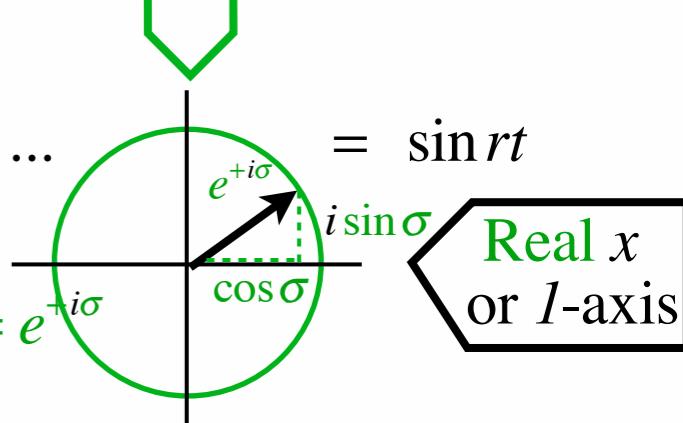
$$\frac{e^{+i rt} - e^{-i rt}}{2} = i rt - i \frac{(rt)^3}{2 \cdot 3}$$

$$\frac{e^{+i rt} - e^{-i rt}}{2i} = rt - \frac{(rt)^3}{2 \cdot 3}$$

$$+ i \frac{(rt)^5}{2 \cdot 3 \cdot 4 \cdot 5} - \dots = i \sin rt$$

$$+ \frac{(rt)^5}{2 \cdot 3 \cdot 4 \cdot 5} - \dots = \sin rt$$

Imaginary y
or i -axis



Phasor circle plot of $\cos \sigma + i \sin \sigma = e^{+i\sigma}$

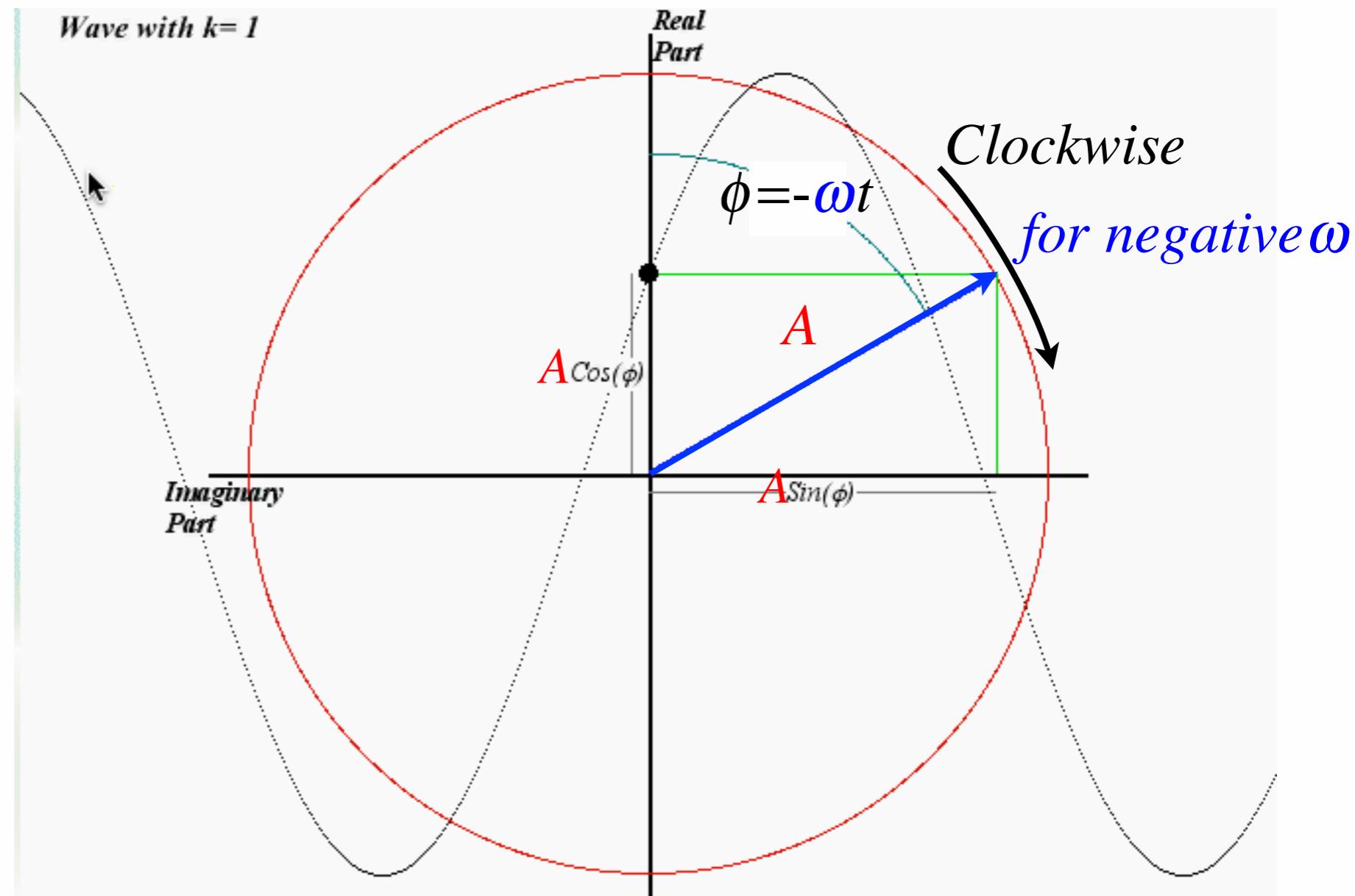
Complex Oscillator Phasors $Ae^{\pm i\phi} = A(\cos\phi \pm i\sin\phi)$

Clockwise angular velocity ω ($\phi=-\omega t$) and Amplitude A

$$Ae^{-i\omega t} = A\cos\omega t - iA\sin\omega t$$

[Animated phasor wave k=1](#)

[Animated waves k=-4, -1, +2, +5](#)



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Set: $A = e^{ikx} = \cos kx + i \sin kx$

to make an x -*moving wave*

$$e^{ikx}e^{-i\omega t} = e^{ikx}(\cos\omega t - i\sin\omega t)$$

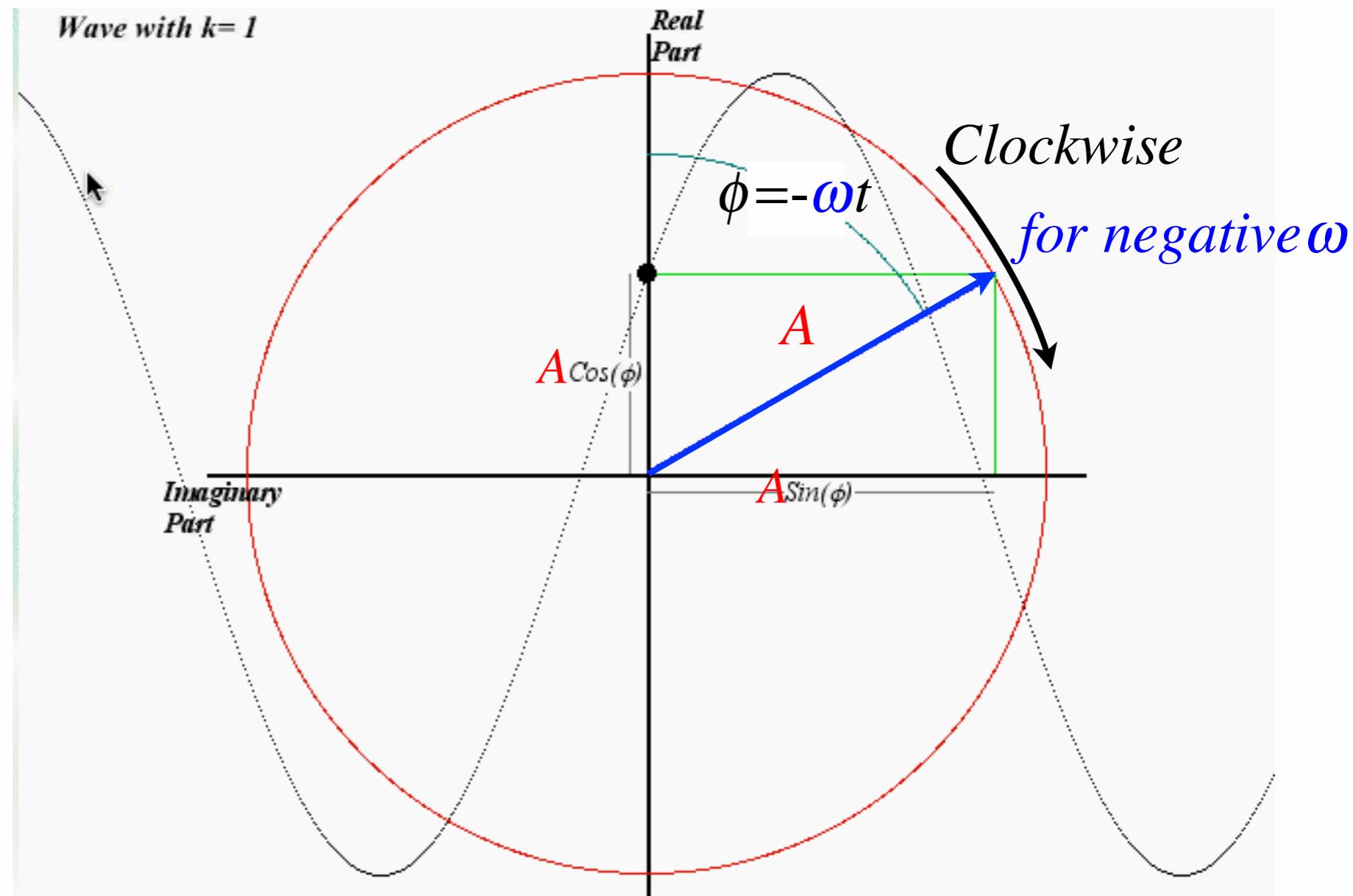
$$= e^{i(kx-\omega t)}$$

$$\cos(\sigma) = \frac{e^{+i\sigma} + e^{-i\sigma}}{2}$$

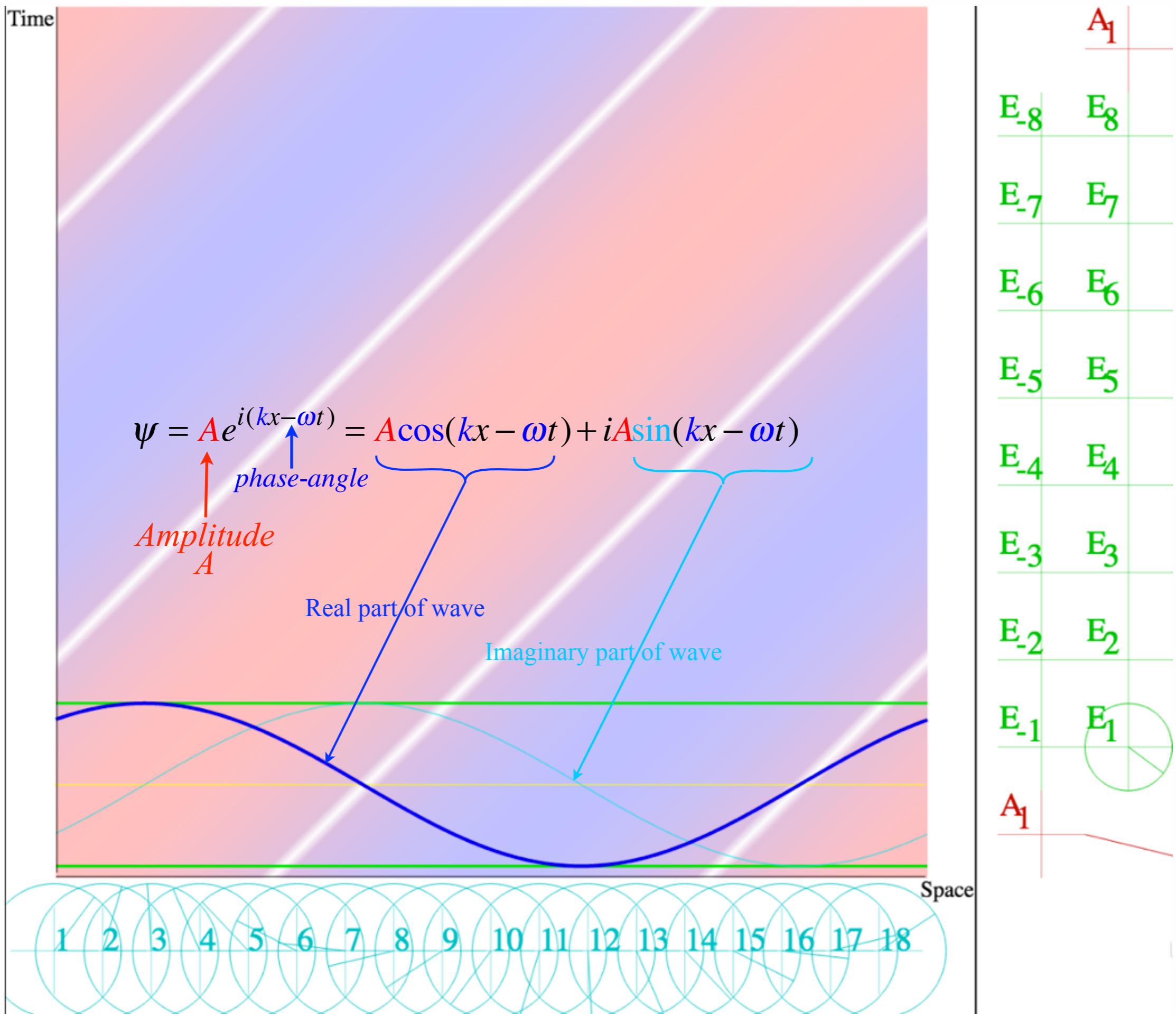
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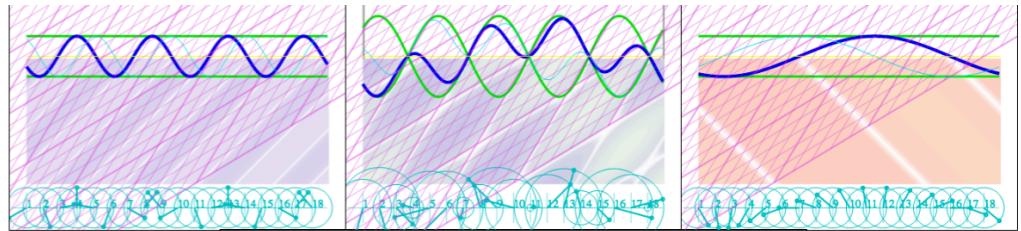
$$\cos(\sigma) - i \sin(\sigma) = e^{-i\sigma}$$



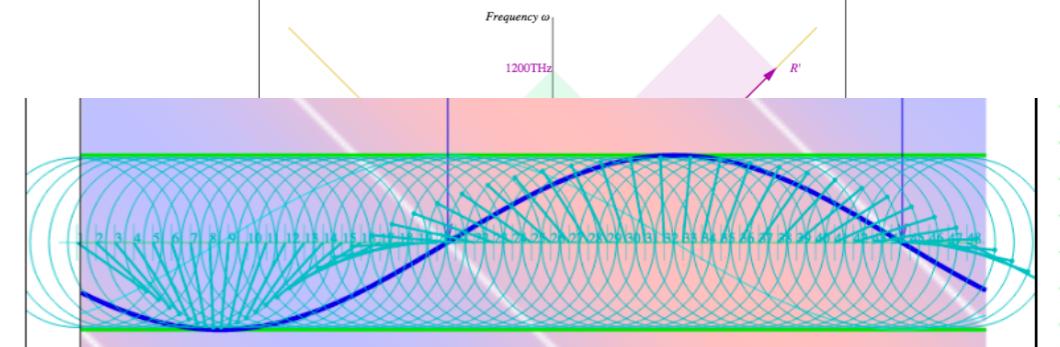
Animated k=+1 1-CW (continuous wave)



Multi-panel beta=0.6c, with extra coordinate grid. Not quite your black axes, but from other app
<http://www.uark.edu/ua/modphys/markup/BohrItWeb.html?scenario=-30104>

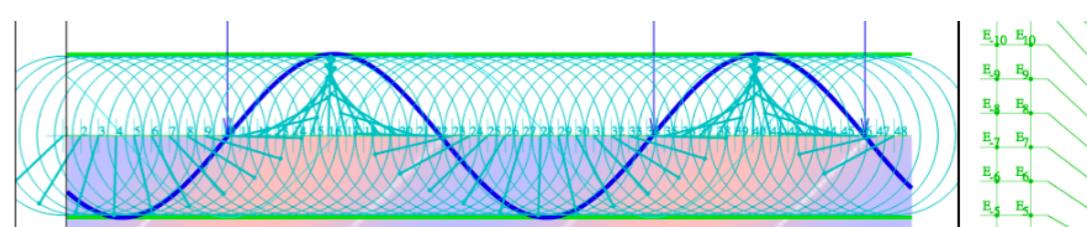


<http://www.uark.edu/ua/modphys/markup/BohrItWeb.html?scenario=-30022>



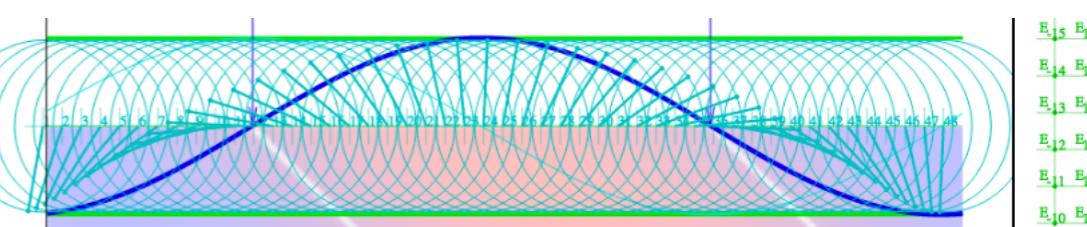
k=1; Hi Rez, points per well = 48 phasors; w/tracers; x phasors move with now line. URL qualifiers: ?

xPhasorLocationsInd=1& doPhaseZeroTracers=1& doGroupZeroTracers=1
<http://www.uark.edu/ua/modphys/markup/BohrItWeb.html?scenario=330002>



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<http://www.uark.edu/ua/modphys/markup/BohrItWeb.html?scenario=-330002>



Link to the talk:

http://www.uark.edu/ua/modphys/pdfs/Talk_Pdfs/INBRE_2016.pdf

Link to the RelaWavy Portal:

<http://www.uark.edu/ua/modphys/markup/RelaWavyPortal.html>

Link to the Harter-Soft Educational Resource Portal:

<http://www.uark.edu/ua/modphys/markup/Harter-SoftWebApps.html>

1CW Laser-phasor wave function

Dimensionless Light wave-velocity $c/c=1$

$$\frac{V_{light}}{c} = \frac{\lambda}{c\tau} = \frac{\nu}{c\kappa} = 1 = \frac{\omega}{ck} \text{ angular units}$$

"winks"
"n
"kinks"

$$\text{angular frequency: } \omega = 2\pi\nu$$

$$\text{angular wave number: } k = 2\pi\kappa$$

k = wavevector

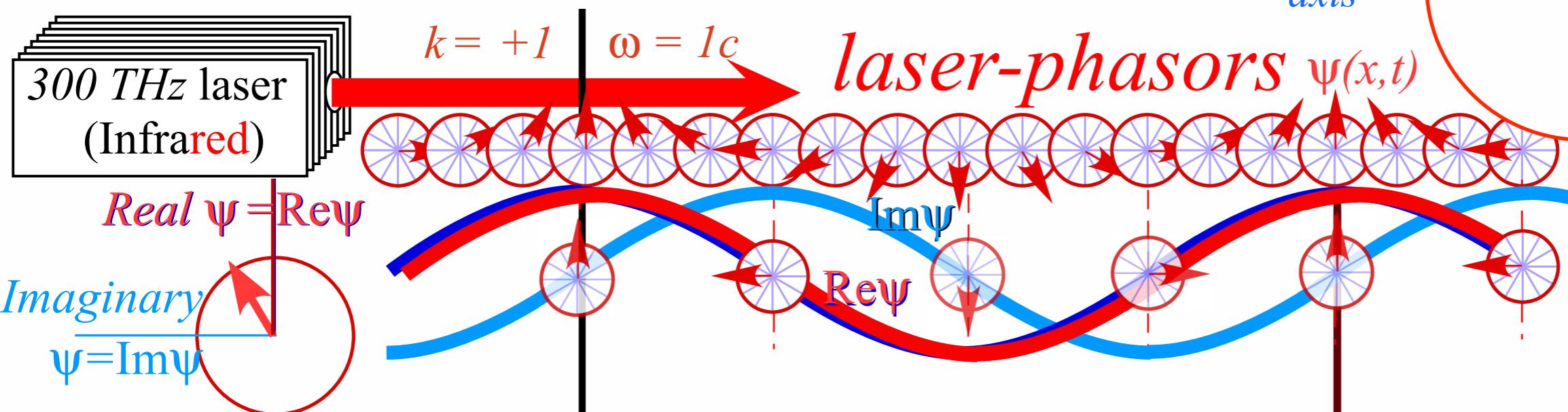
$$\psi = A \cdot e^{i(kx - \omega t)} = A \cdot \cos(kx - \omega t) + iA \cdot \sin(kx - \omega t)$$

Amplitude A

phase-angle $(kx - \omega t)$

Imaginary axis

Real axis



laser-phasors $\psi(x, t)$

Imaginary axis

Real axis

Amplitude A

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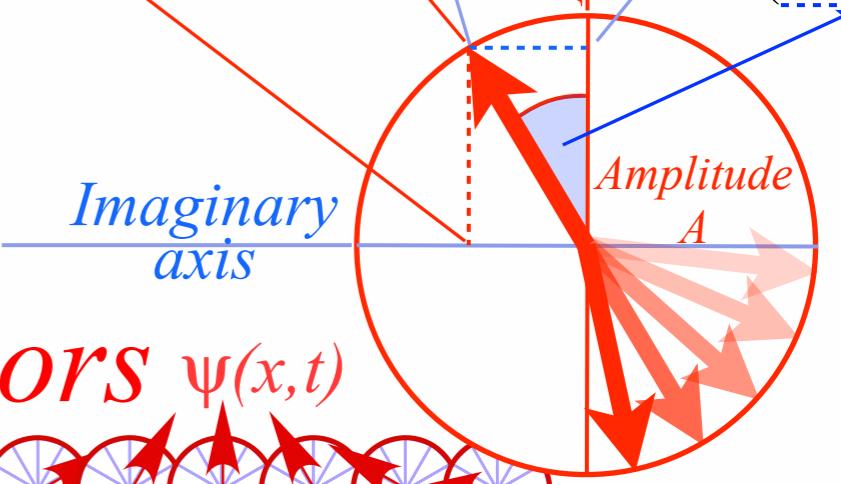
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Amplitude A

phase-angle $(kx - \omega t)$



300 THz laser
(Infrared)

Real $\psi = \operatorname{Re}\psi$

Imaginary $\psi = \operatorname{Im}\psi$

$k = +1$

$\omega = 1c$

laser-phasors $\psi(x, t)$

Imagination precedes Reality by exactly One Quarter!

Mantra for US publicly traded corporations

Wavelength $\lambda = 2\pi/k = 1/\kappa$

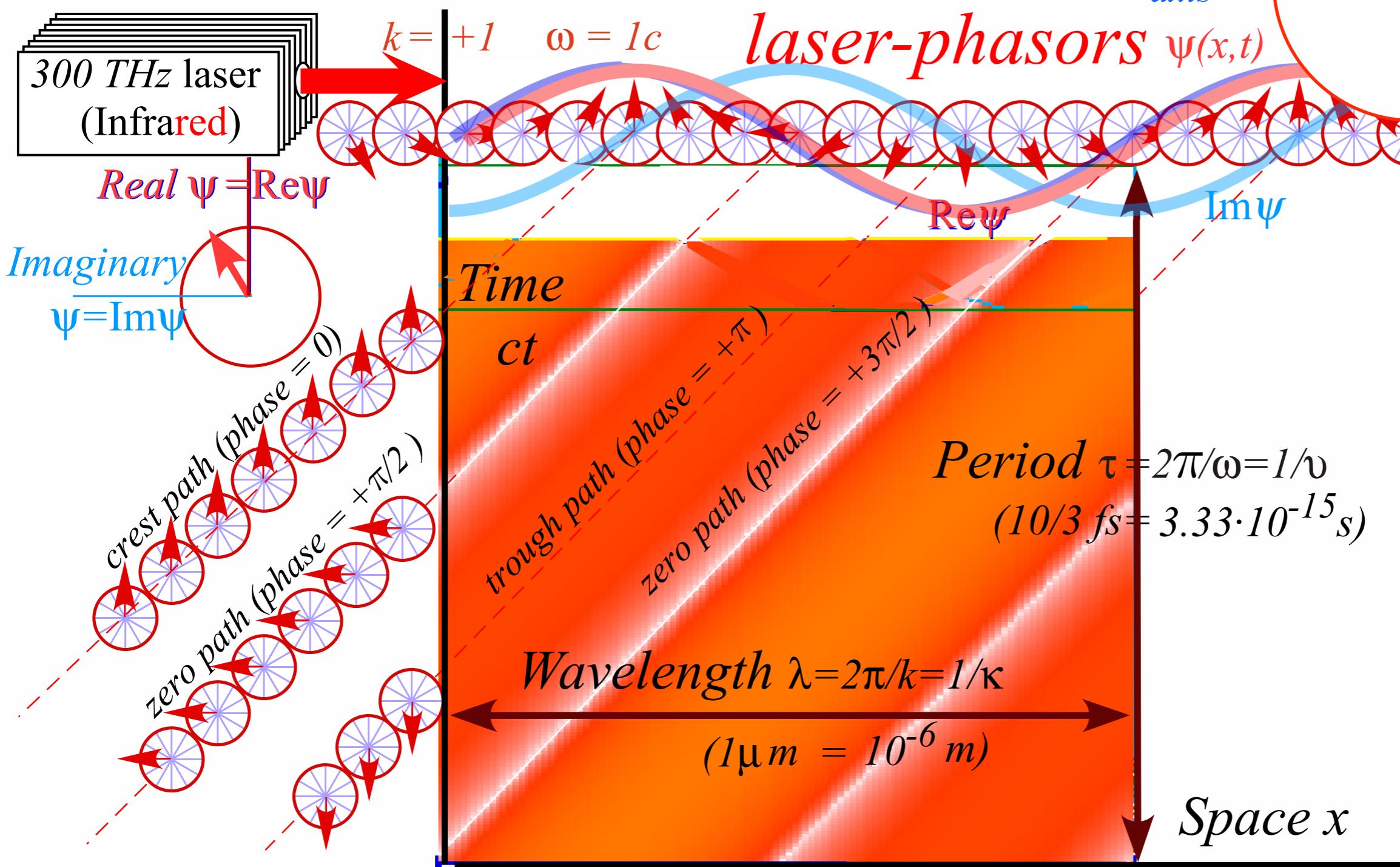
$(1\mu m = 10^{-6} m)$

1CW Laser-phasor wave function

$$\psi = A \cdot e^{i(kx - \omega t)} = A \cdot \cos(kx - \omega t) + iA \cdot \sin(kx - \omega t)$$

Amplitude A

phase-angle $(kx - \omega t)$



1CW Laser-phasor wave function

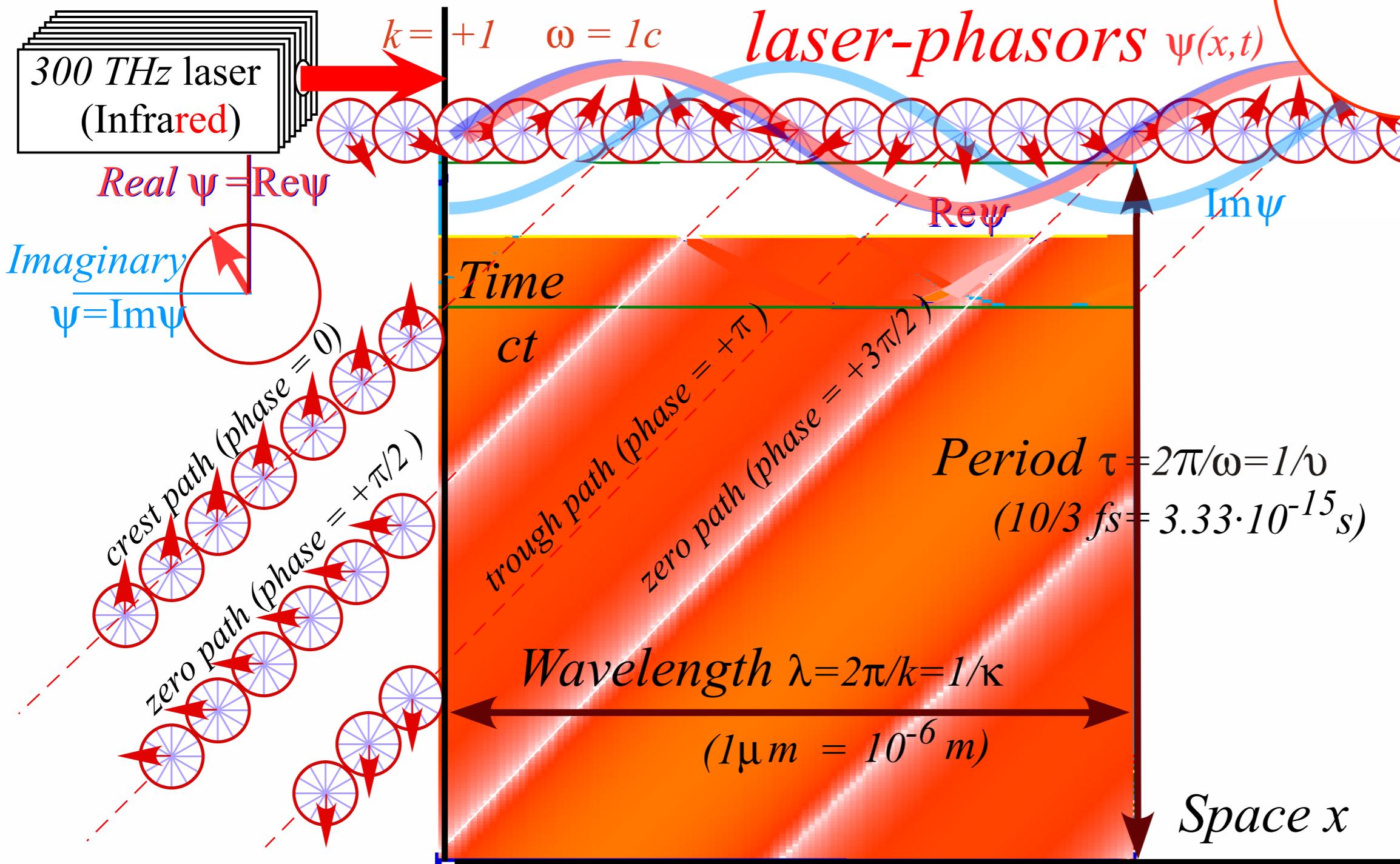
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Q: Where is phase $= (kx - \omega t) = 0$?

A: It is wherever this is: $\frac{x}{t} = \frac{\omega}{k}$



1CW Laser-phasor wave function

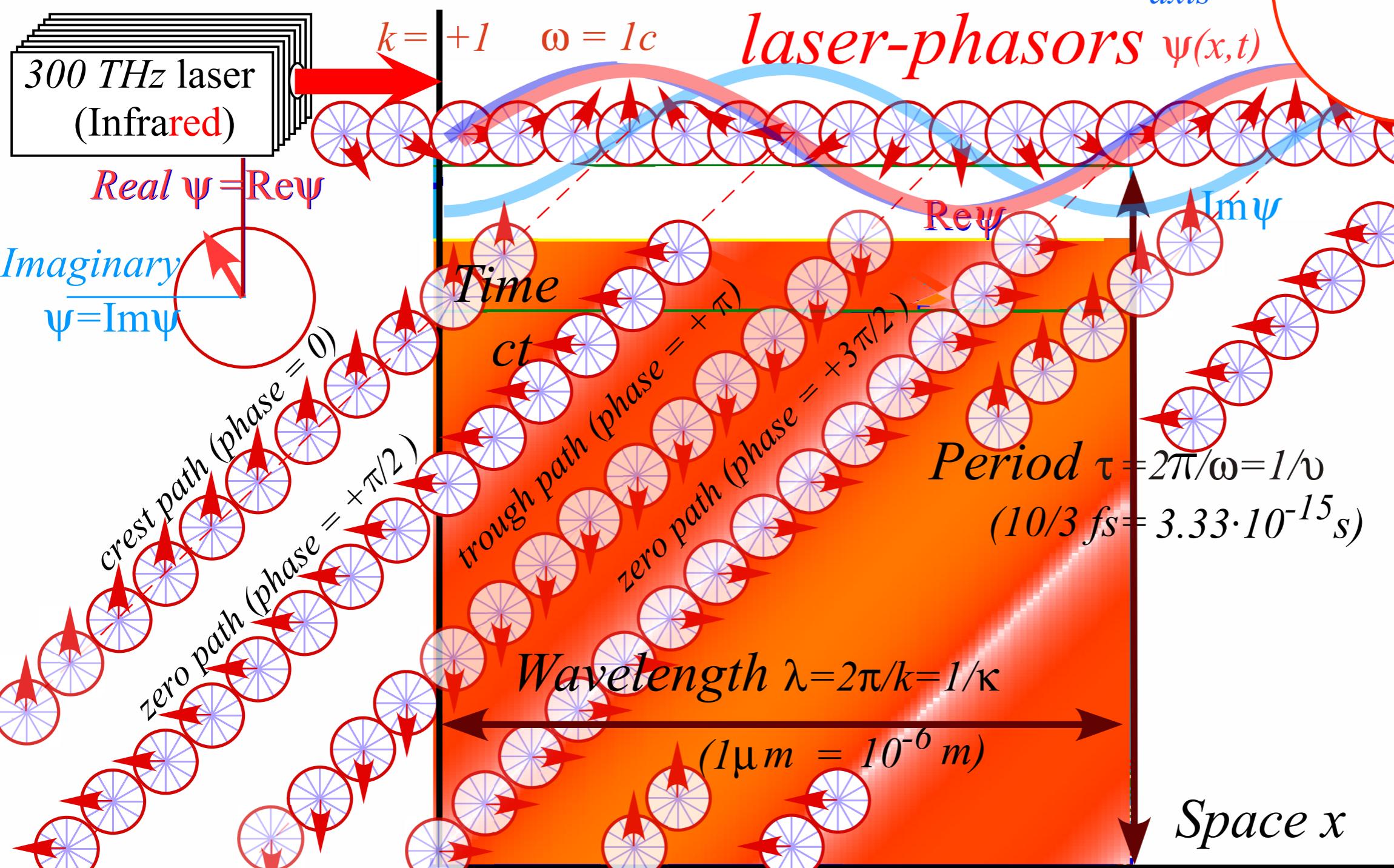
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Amplitude A

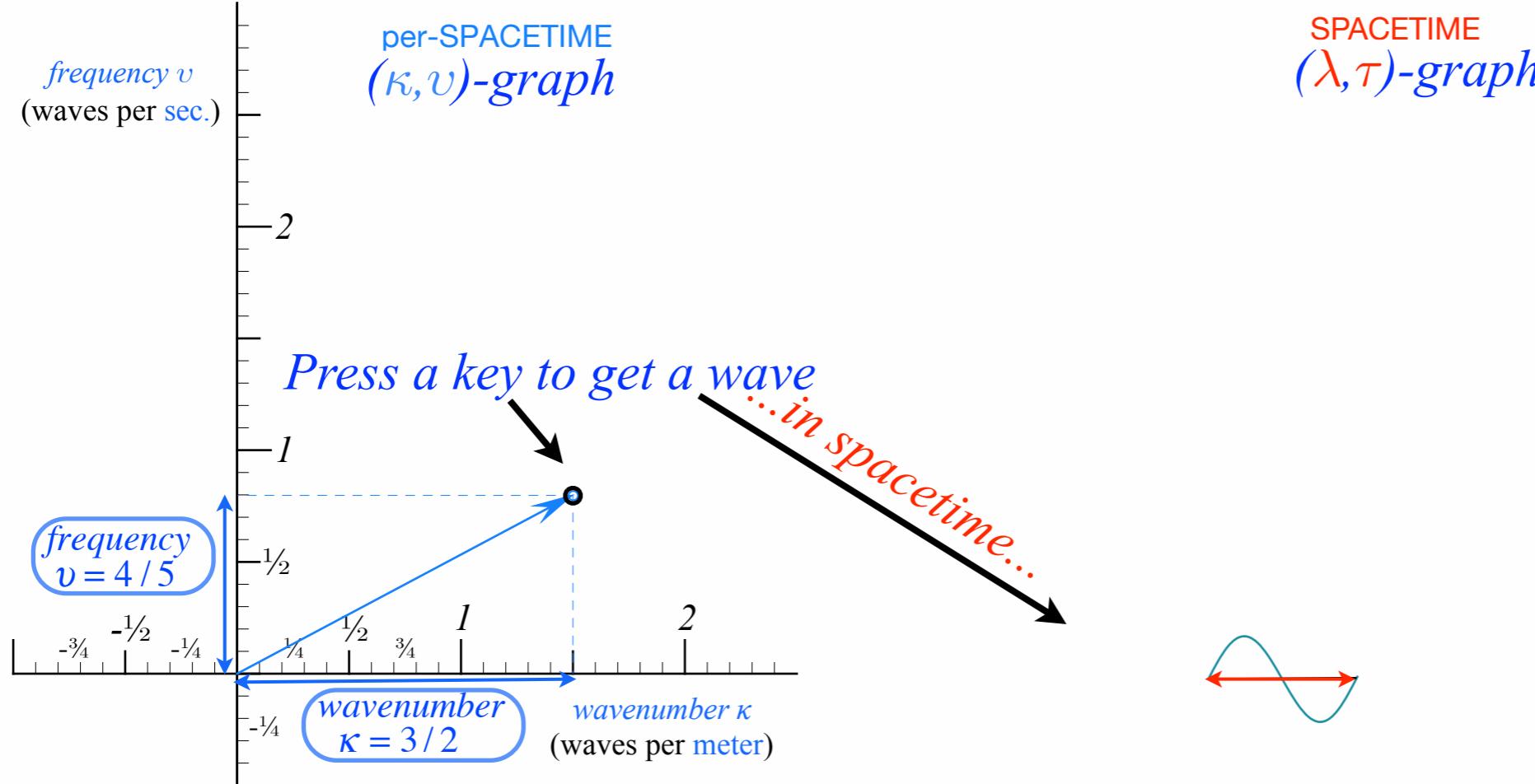
phase-angle $(kx - \omega t)$

Q: Where is phase $= (kx - \omega t) = 0$?

A: It is wherever this is: $\frac{x}{t} = \frac{\omega}{k}$ = wave phase velocity



The “Keyboard of the gods” : Introducing per-space-per-time graphs *versus* space-time graphs



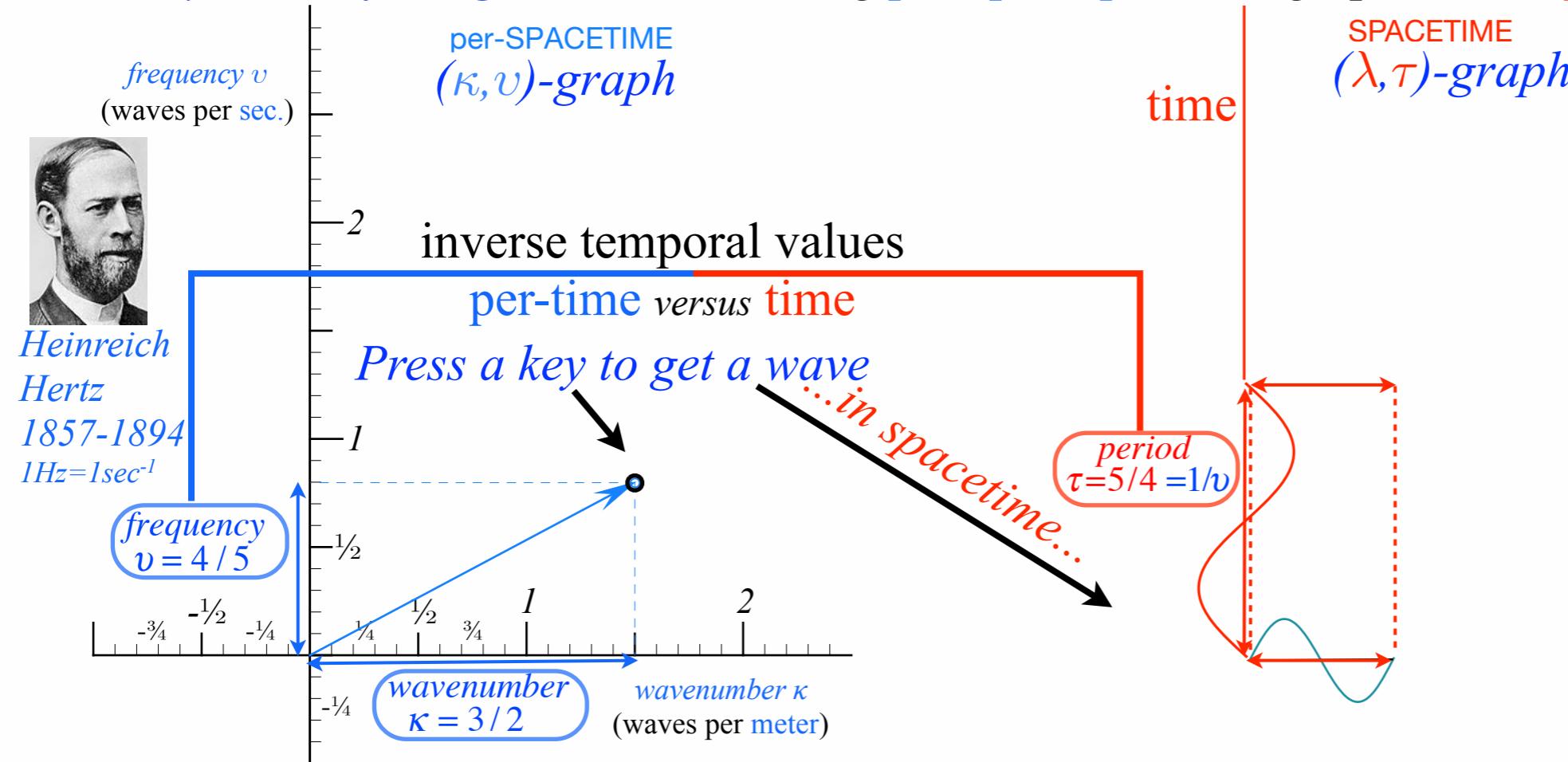
“Keyboard of the gods” is known as “Fourier-space”



Jean-Baptiste
Joseph Fourier
1768-1830

- How to understand waves
and
wave velocity V_{wave}

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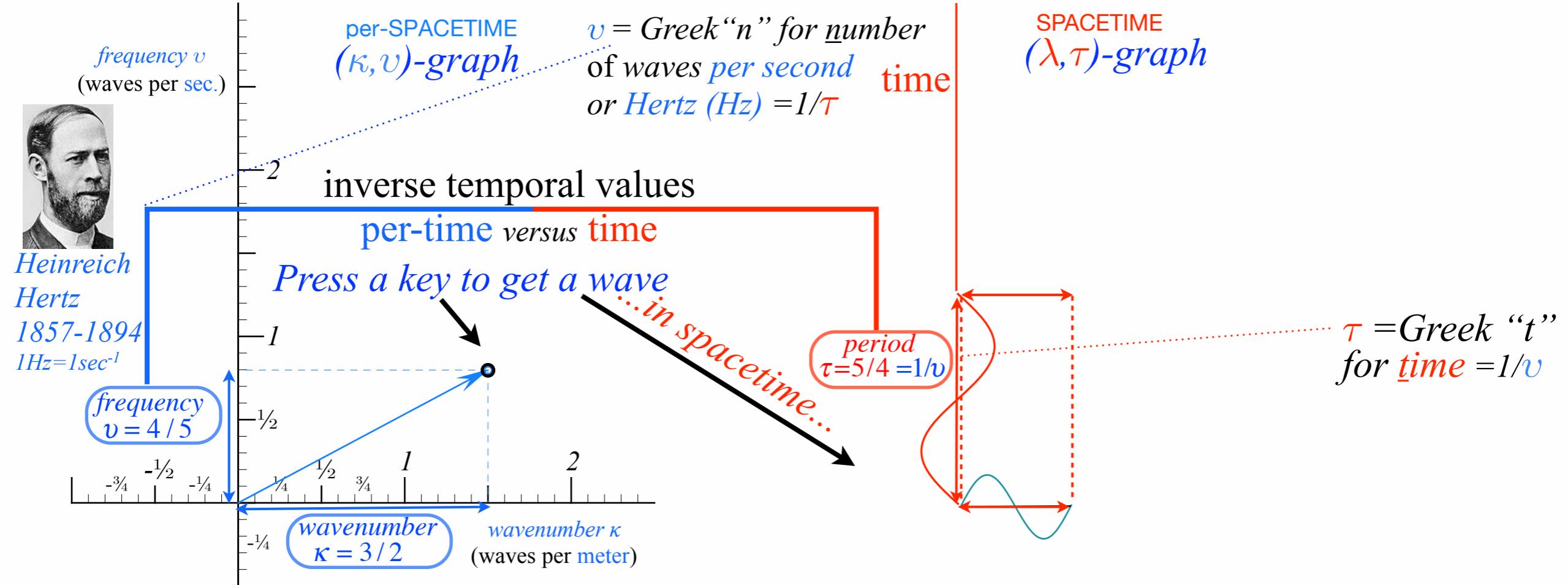
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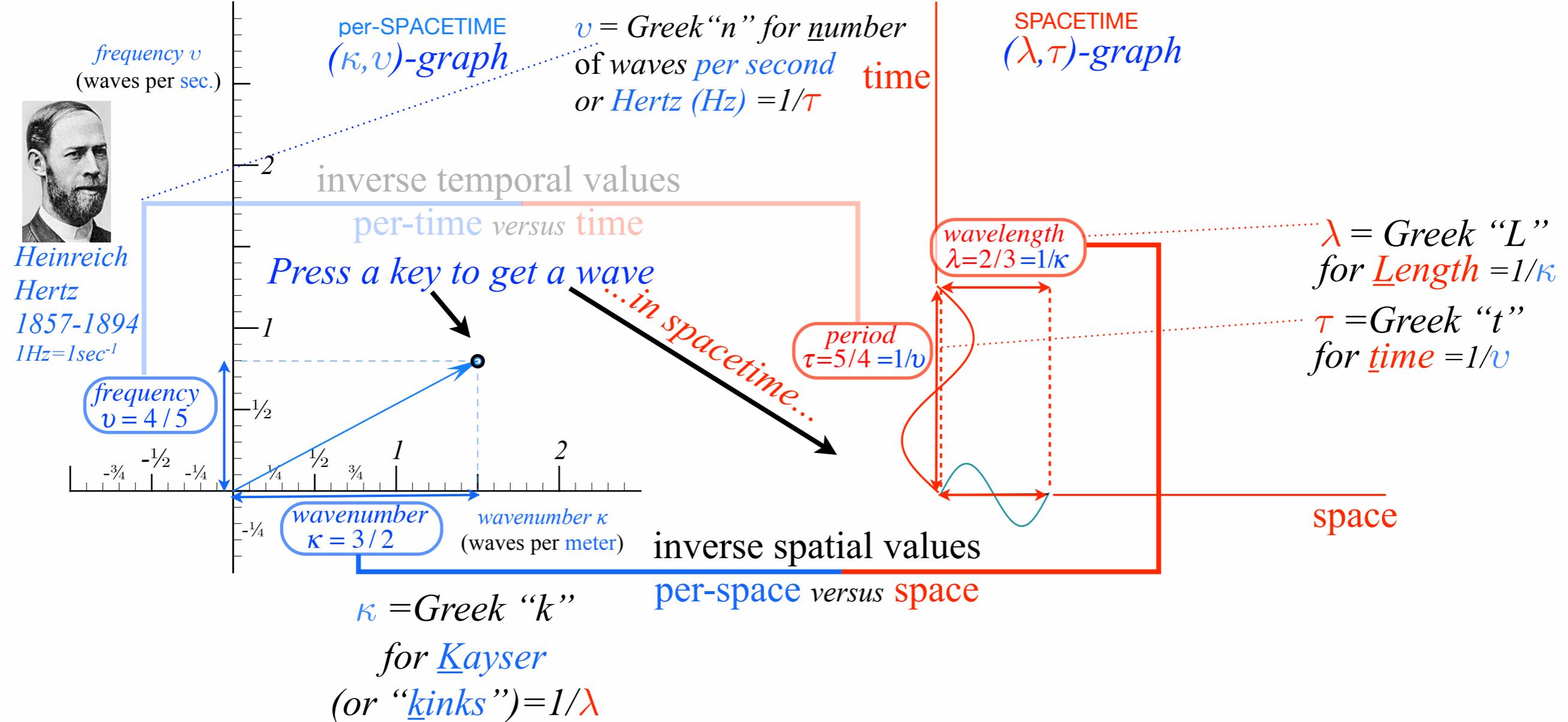
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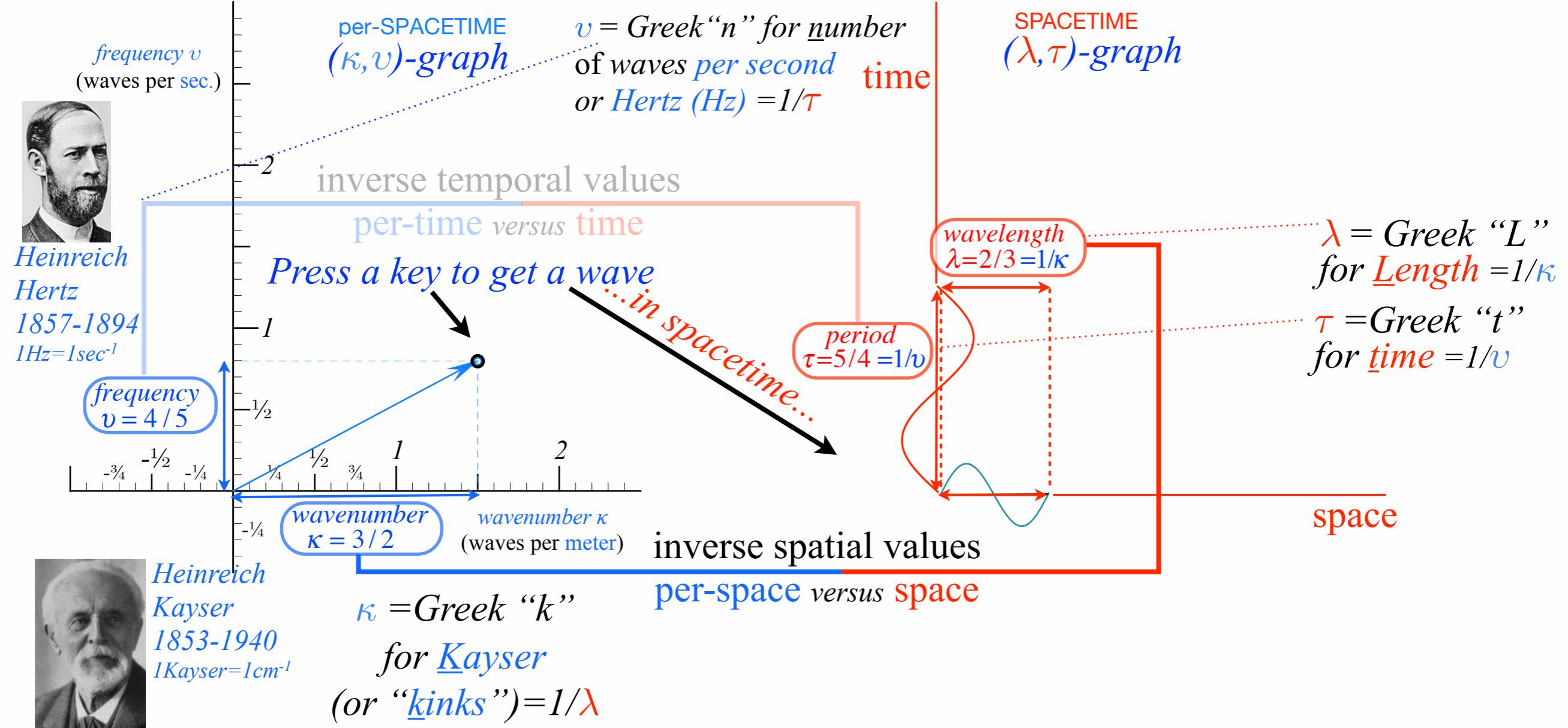
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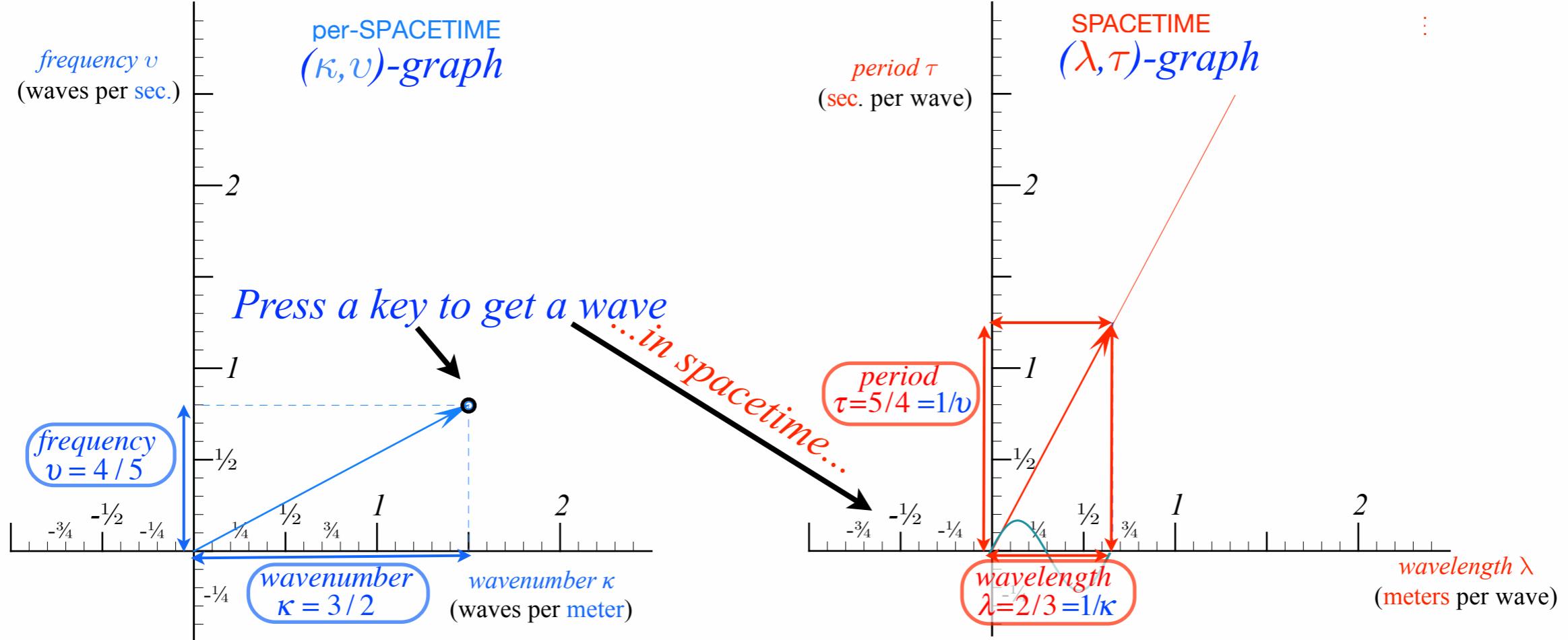
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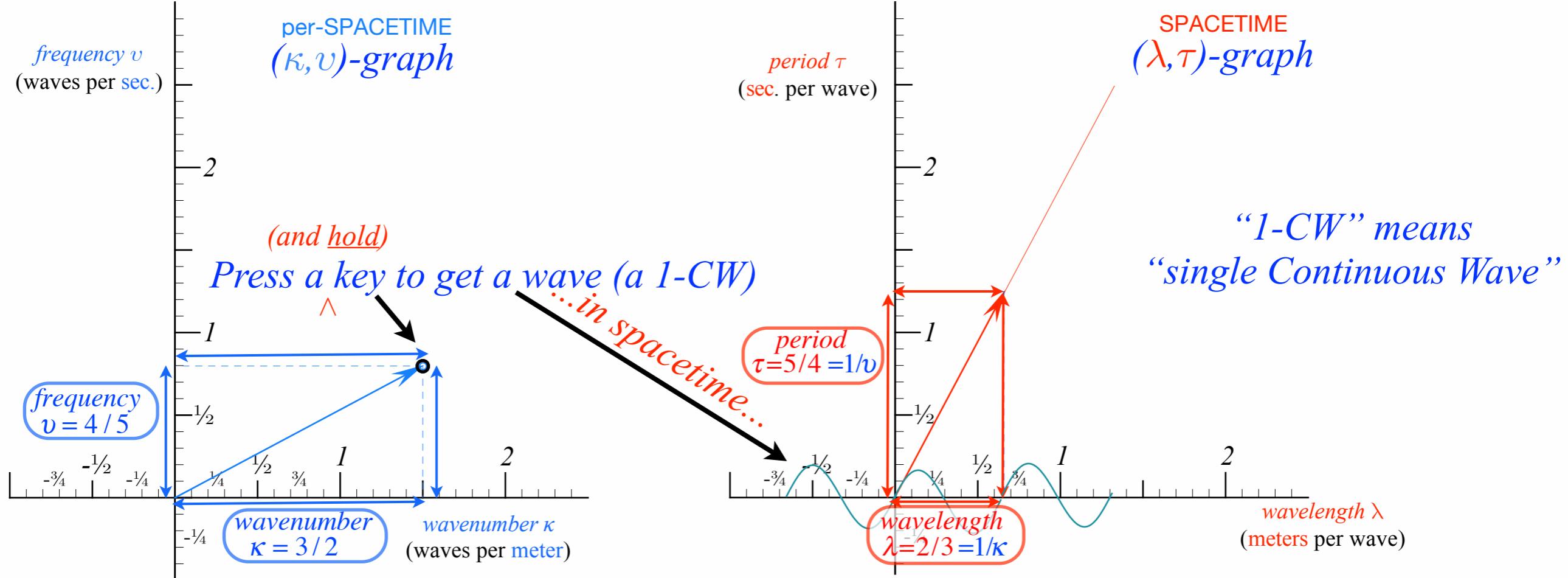
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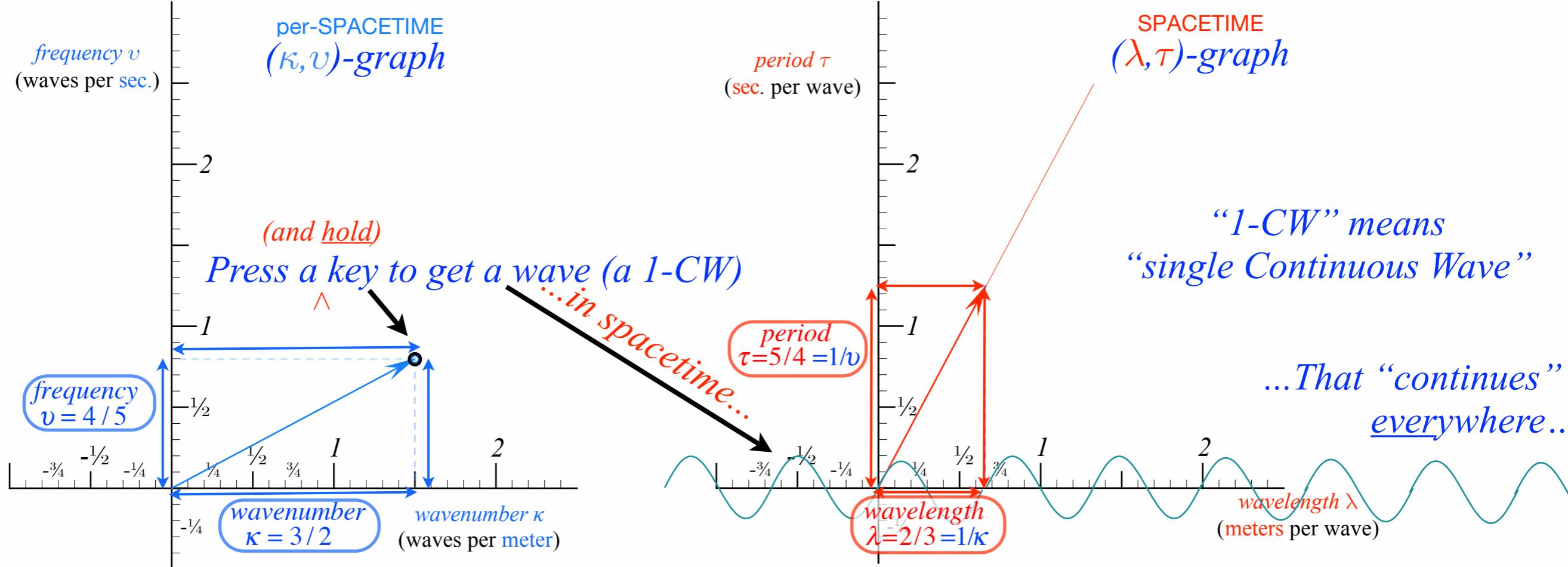
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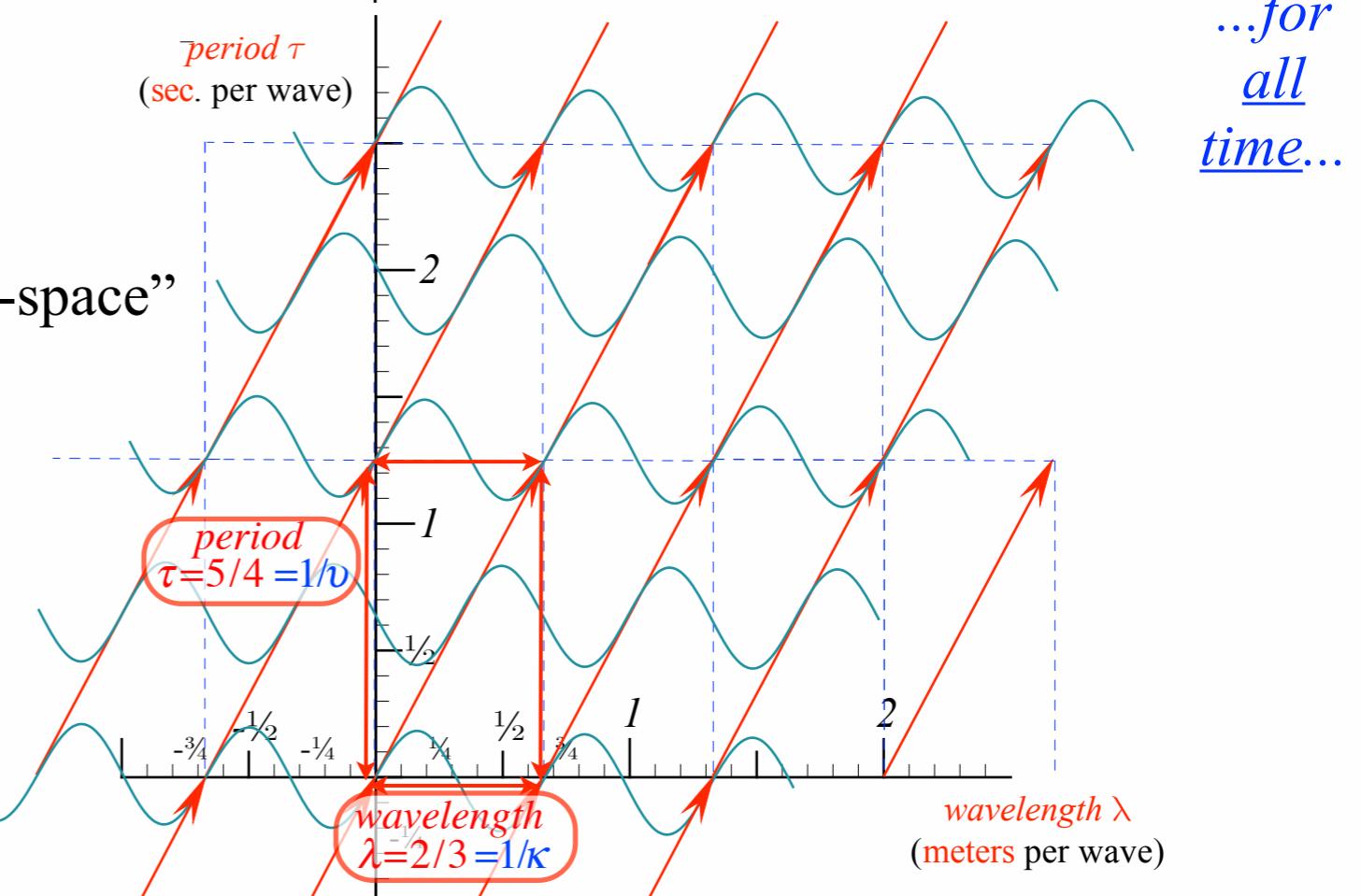
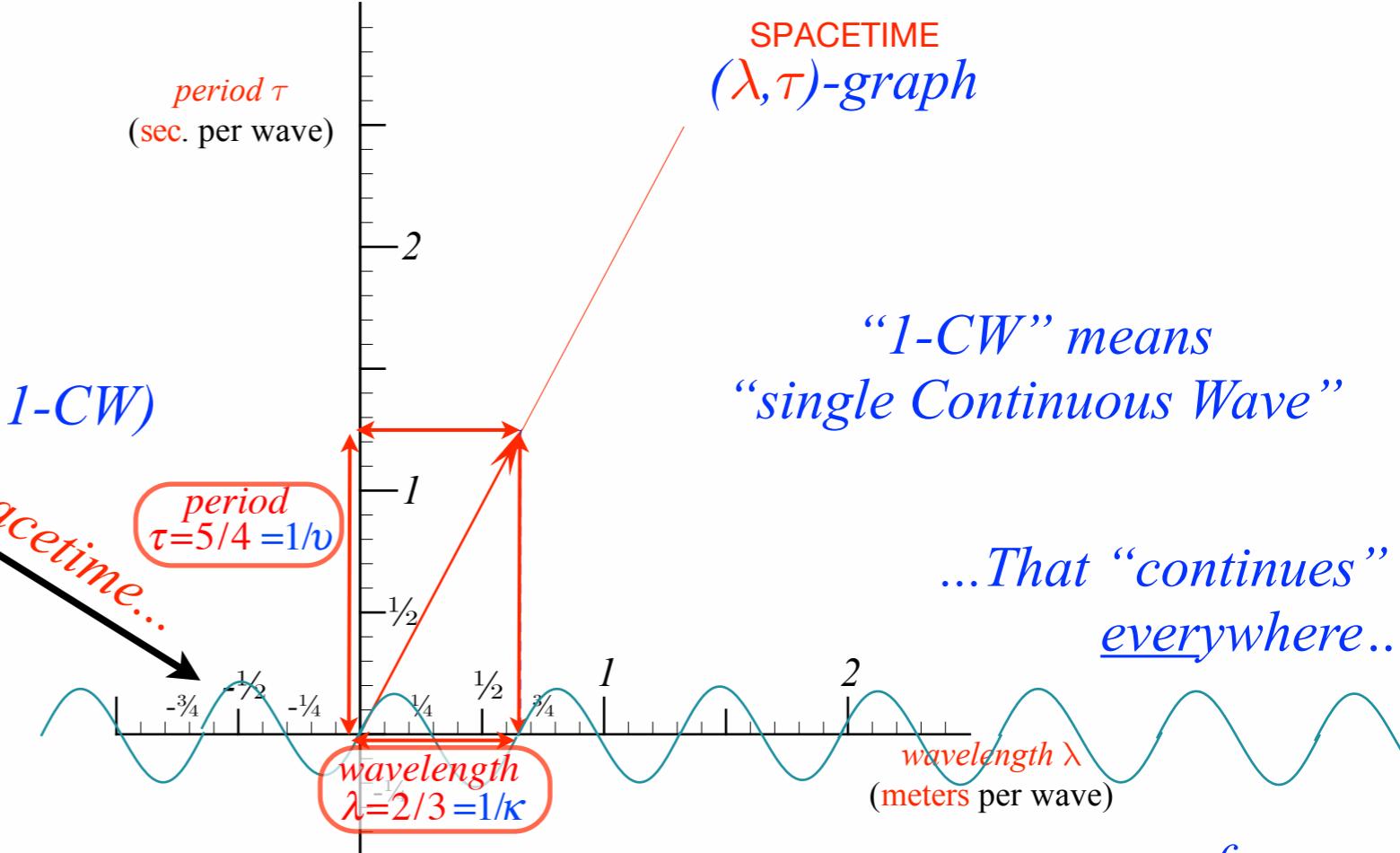
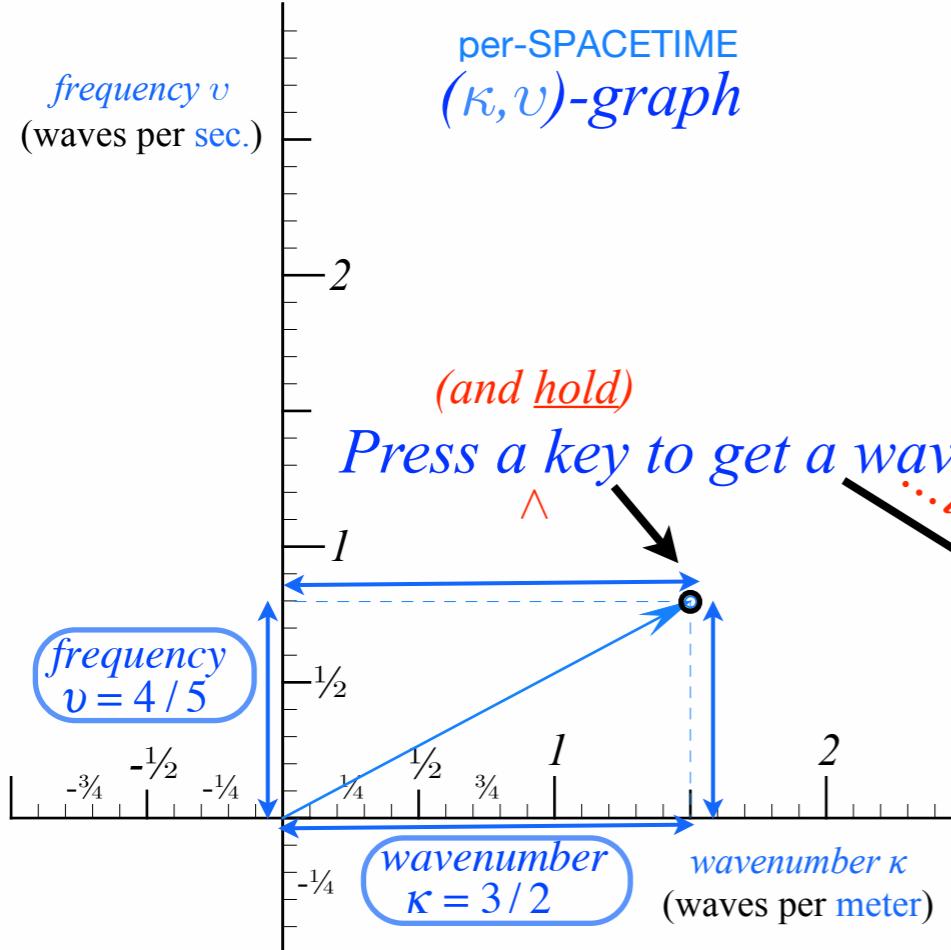
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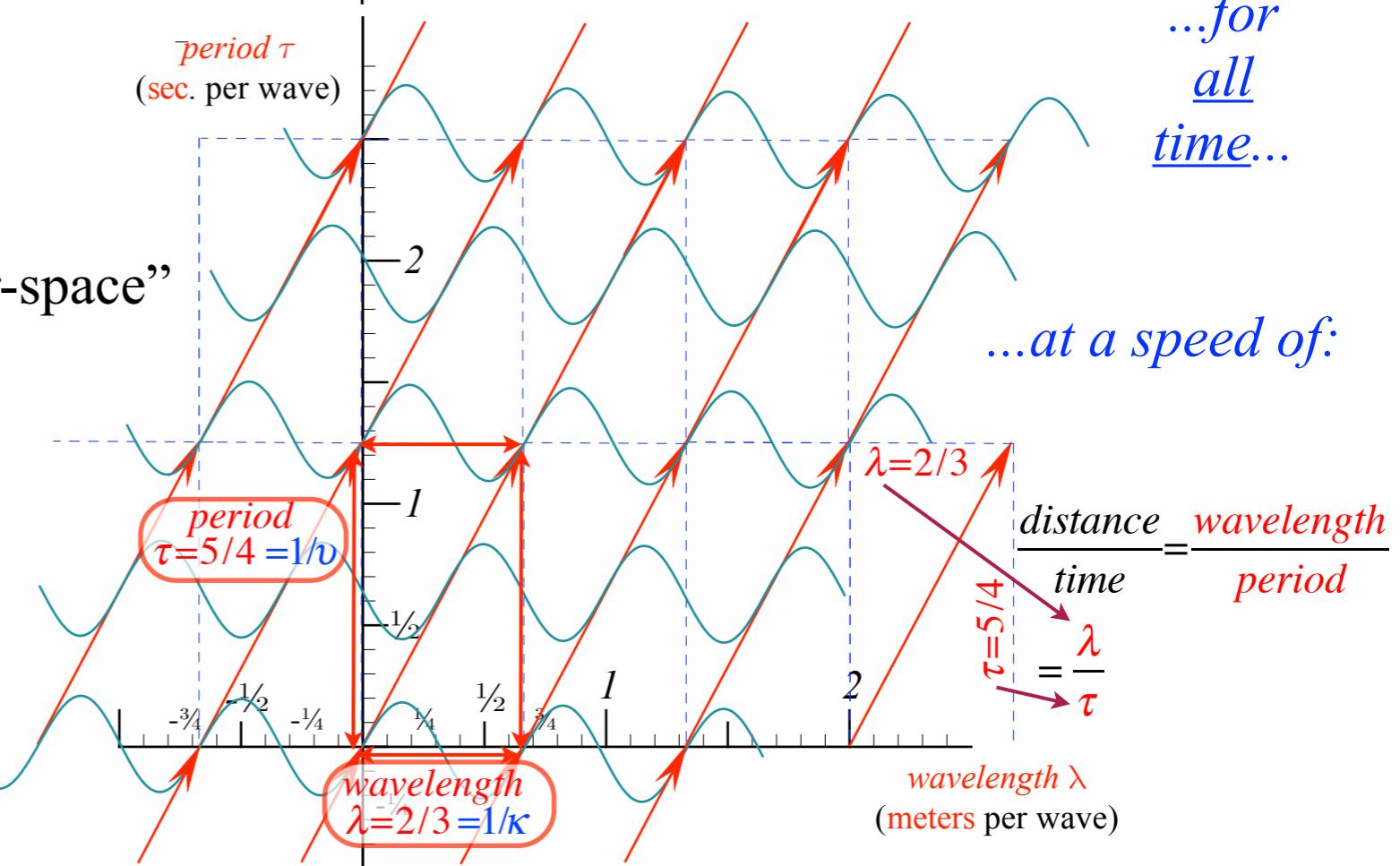
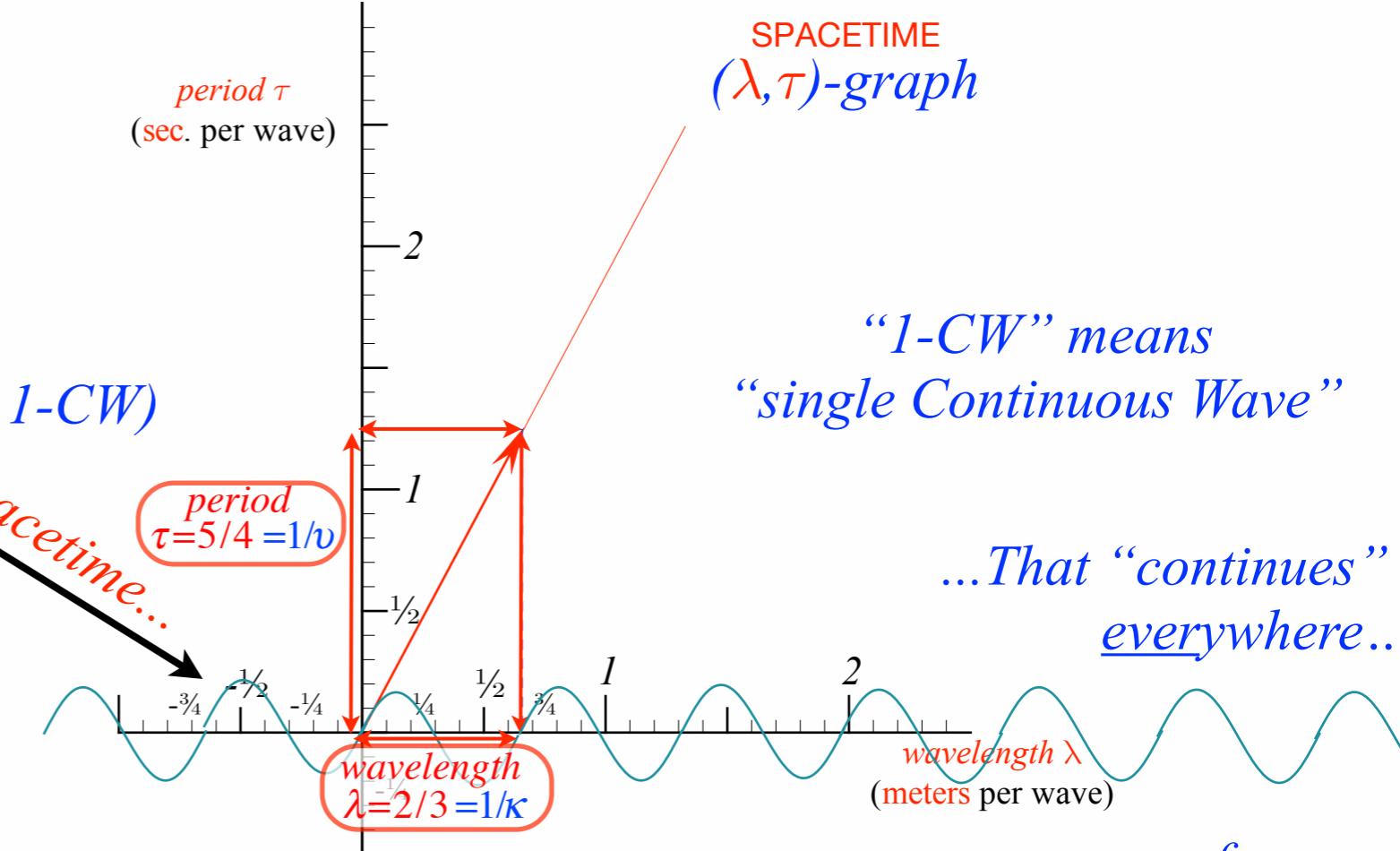
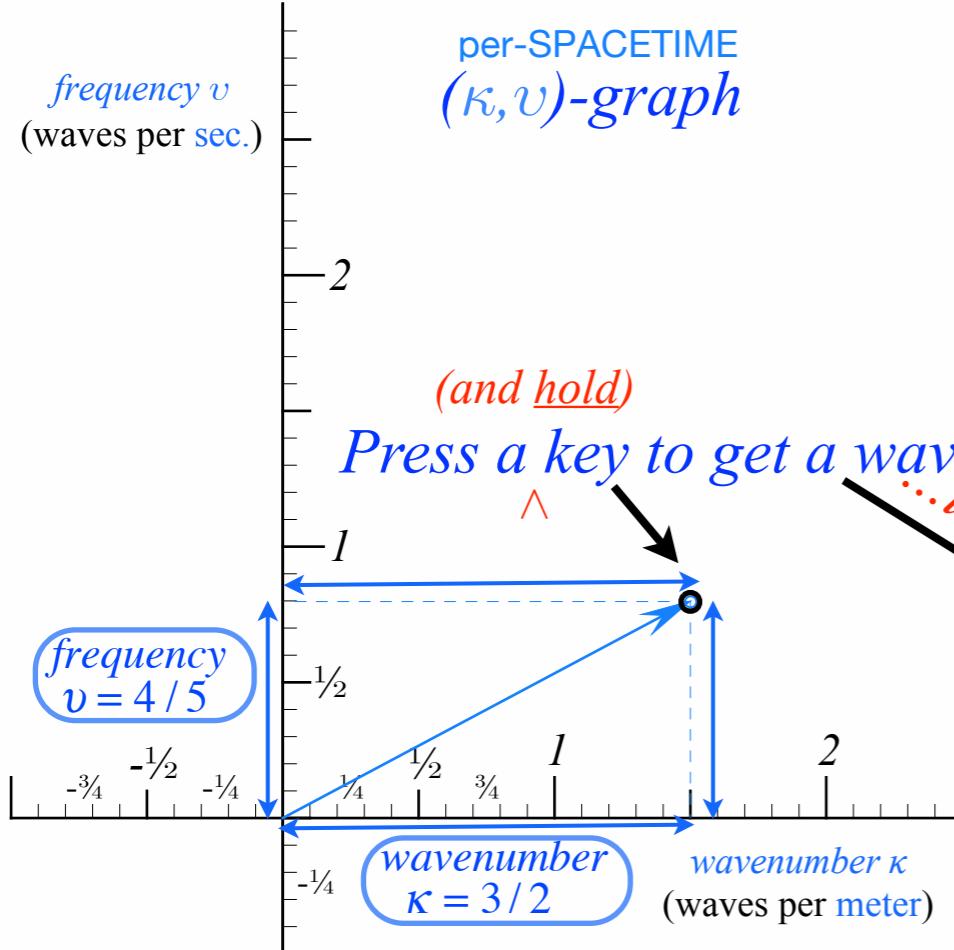
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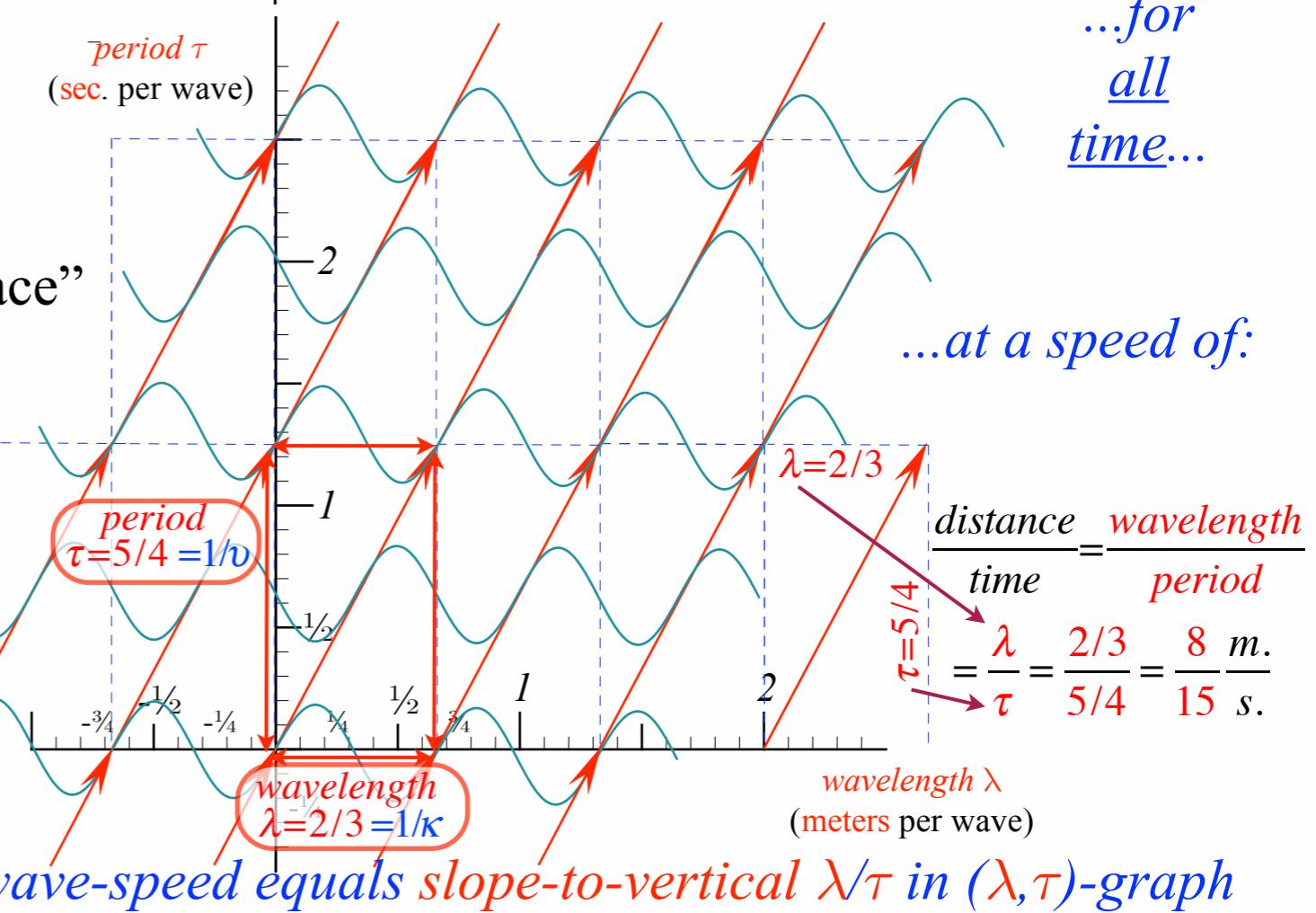
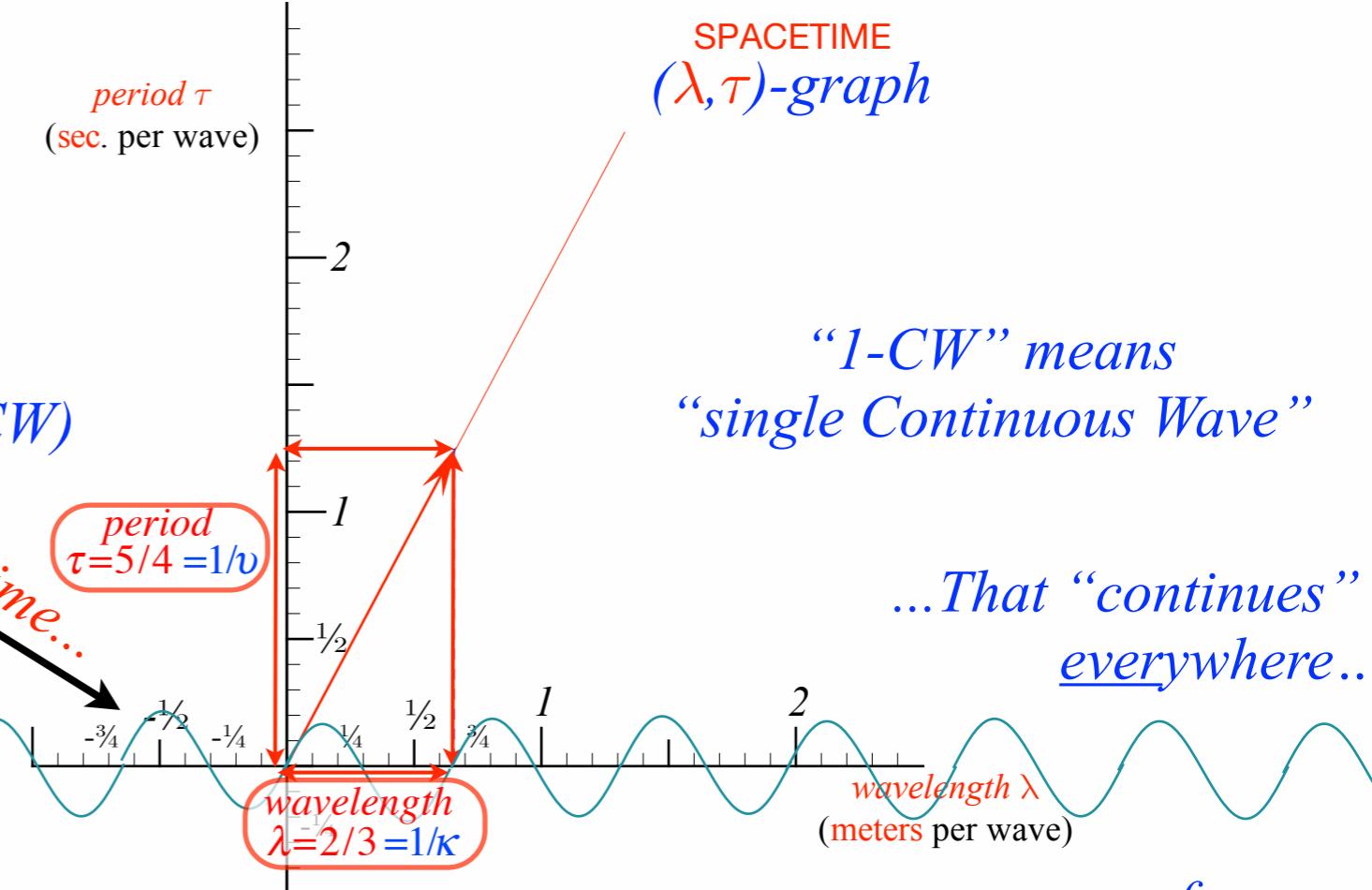
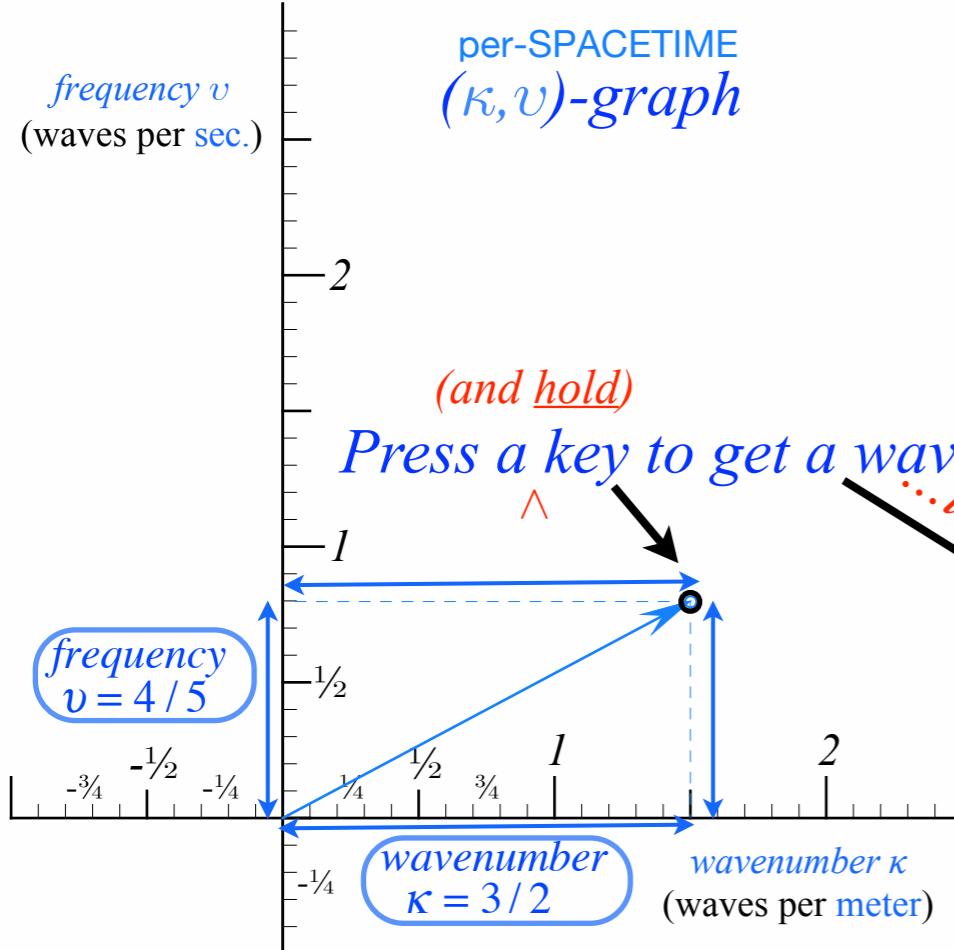
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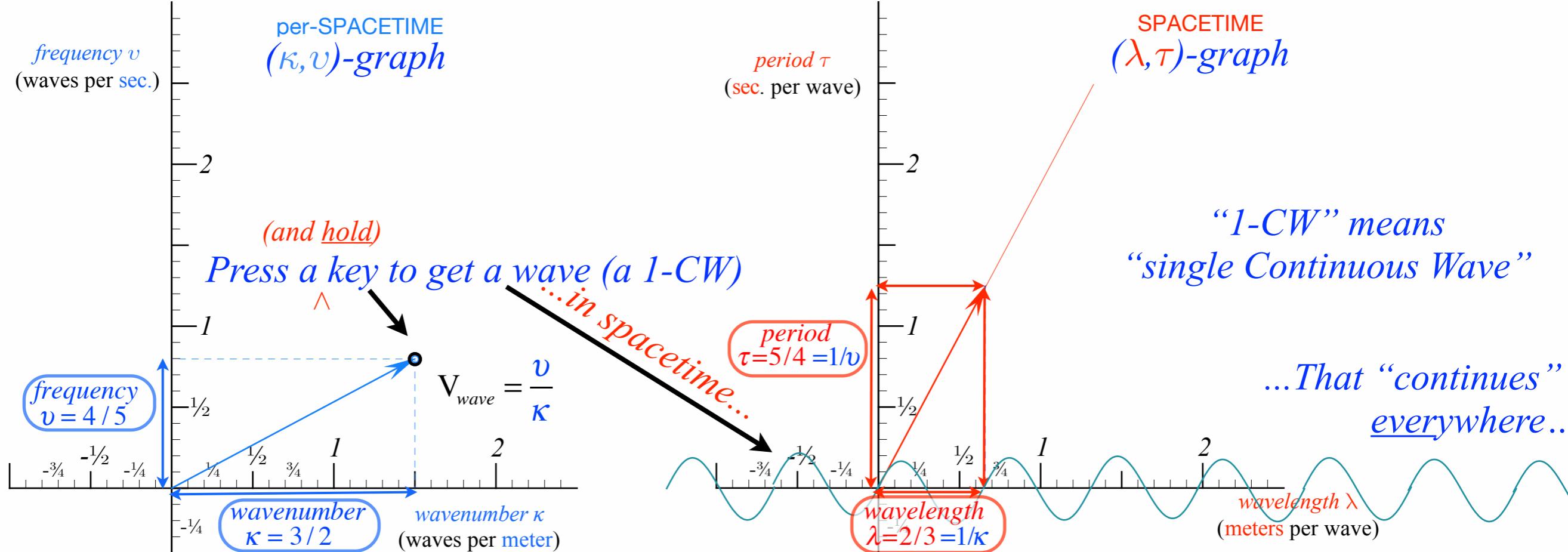
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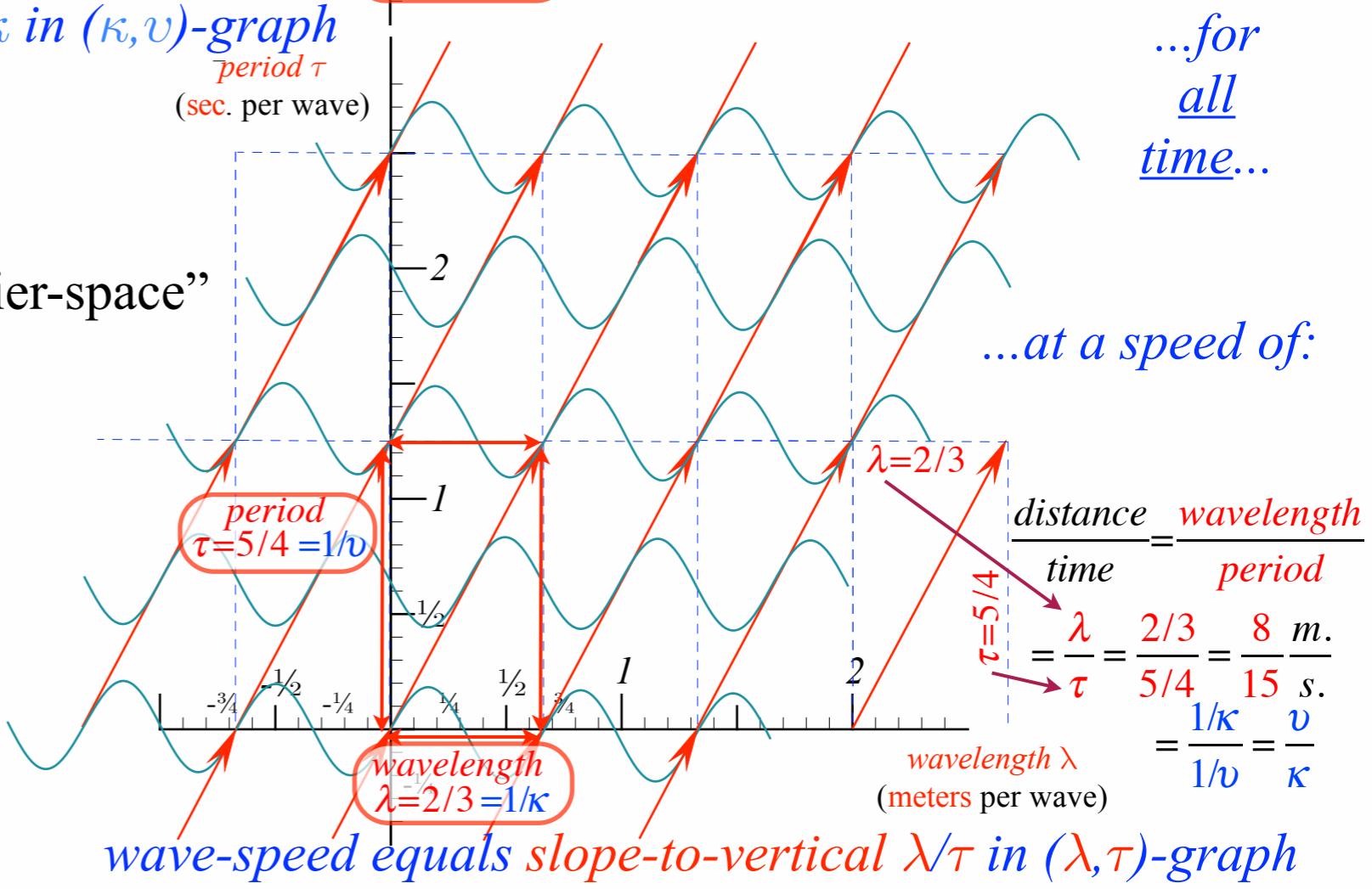


wave-speed equals slope-to-horizontal v/κ in (κ, v) -graph

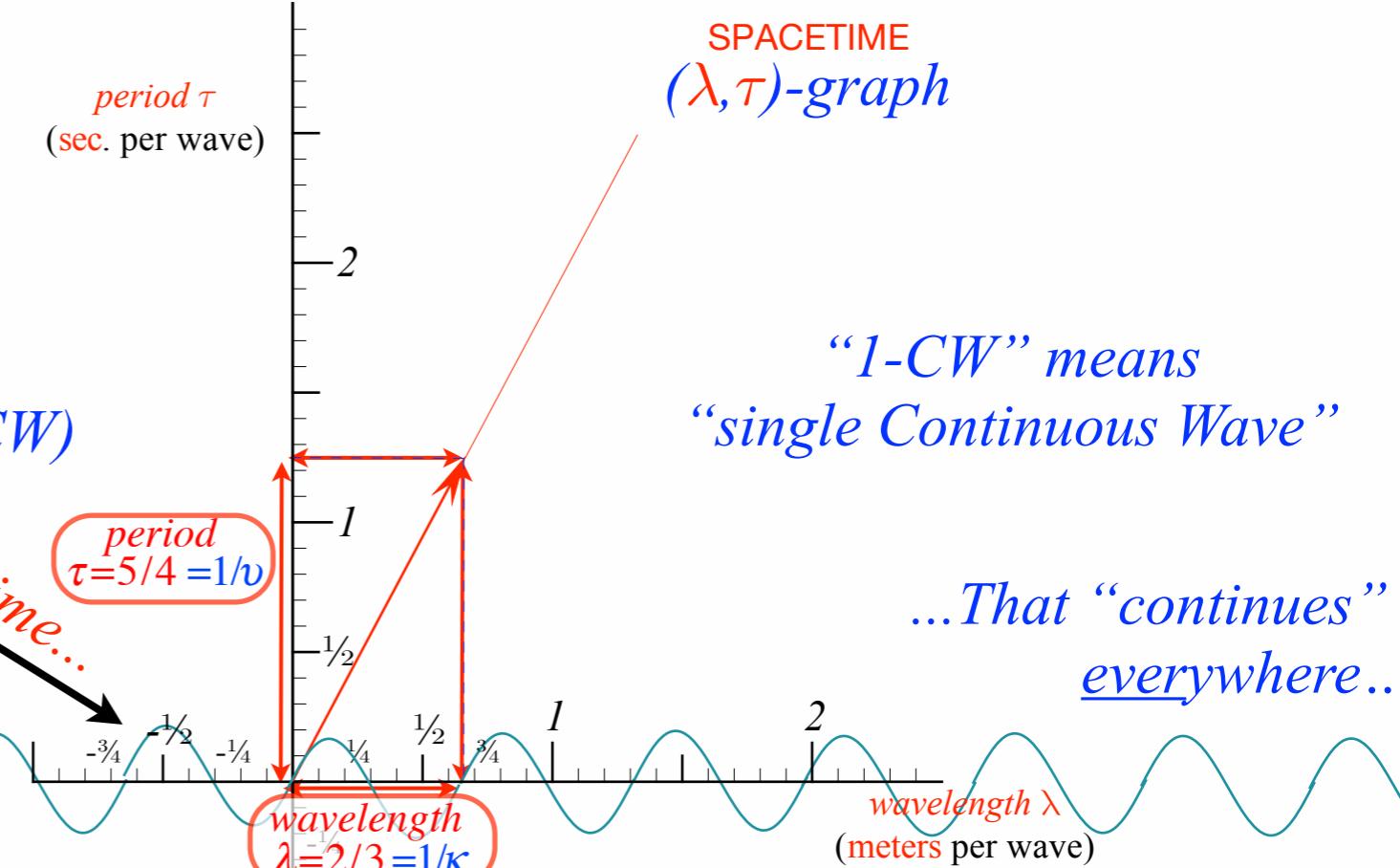
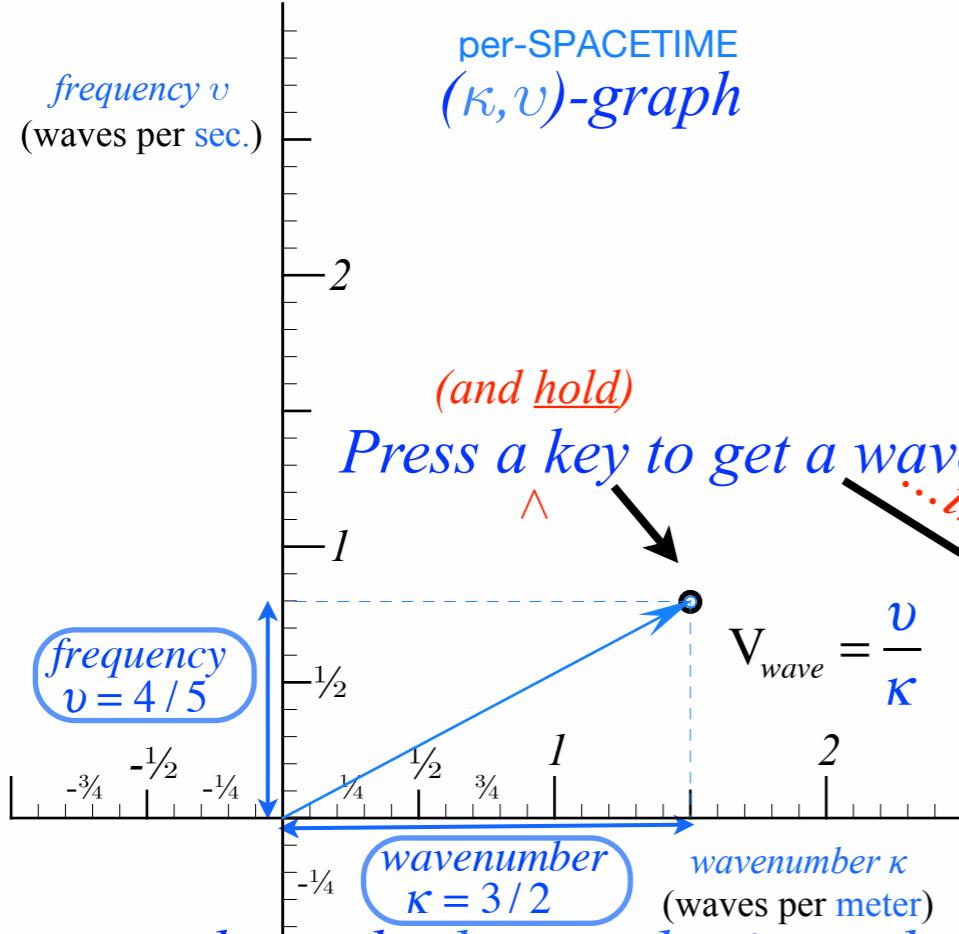


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wave-velocity formulas

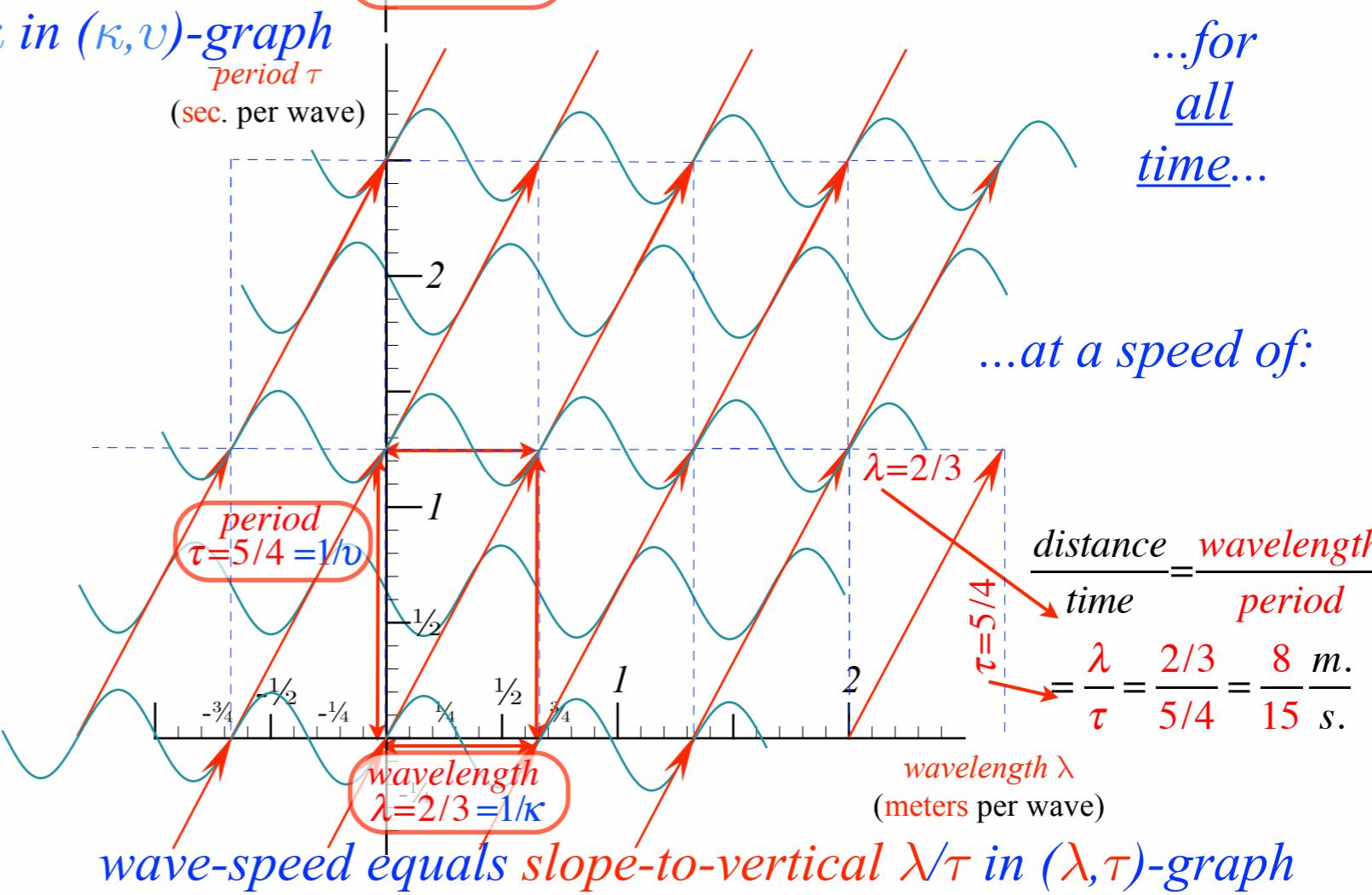
$$\frac{\text{distance}}{\text{time}} = \frac{\text{wavelength}}{\text{period}} = \frac{\text{frequency}}{\text{wavenumber}}$$

$$V_{wave} = \frac{\lambda}{\tau} = \frac{1/\kappa}{1/v} = \frac{v}{\kappa} = \frac{1/\tau}{1/\lambda}$$

$$= \frac{2/3}{5/4} = \frac{4/5}{3/2} = \frac{8}{15} \text{ m. s.}$$

wave arithmetic is simpler to explain using fractions

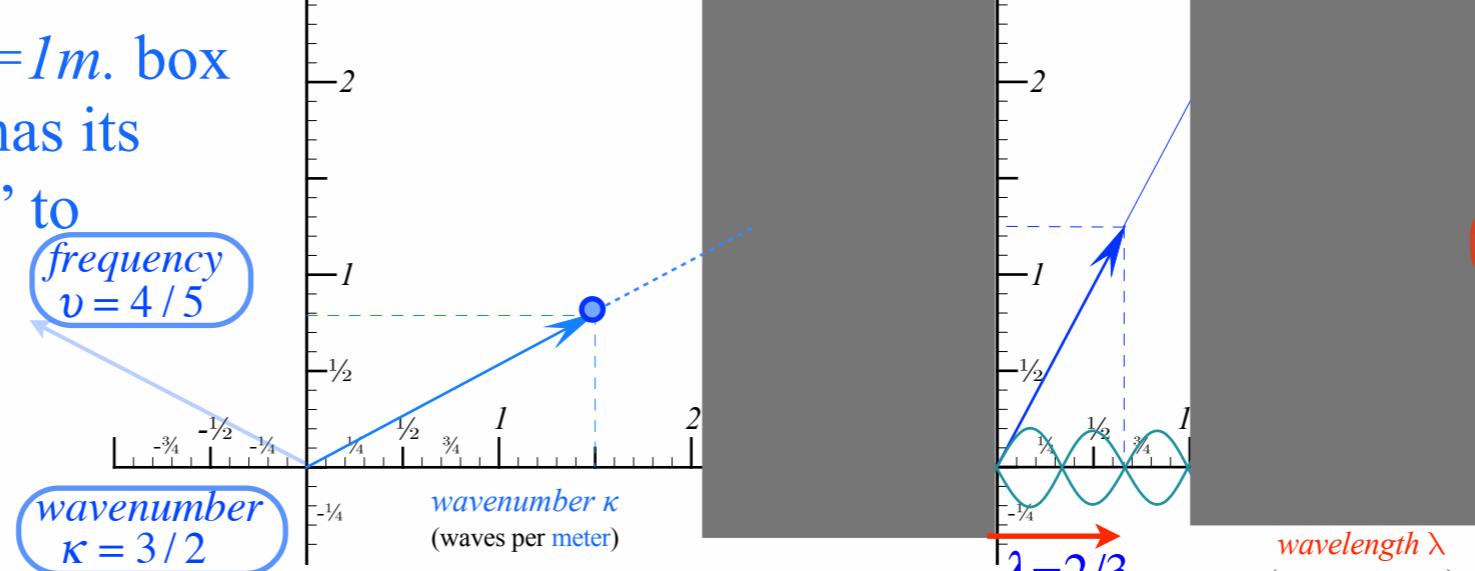
• How to understand waves
and
“1st quantization”



wave-speed equals slope-to-vertical λ/τ in (λ, τ) -graph

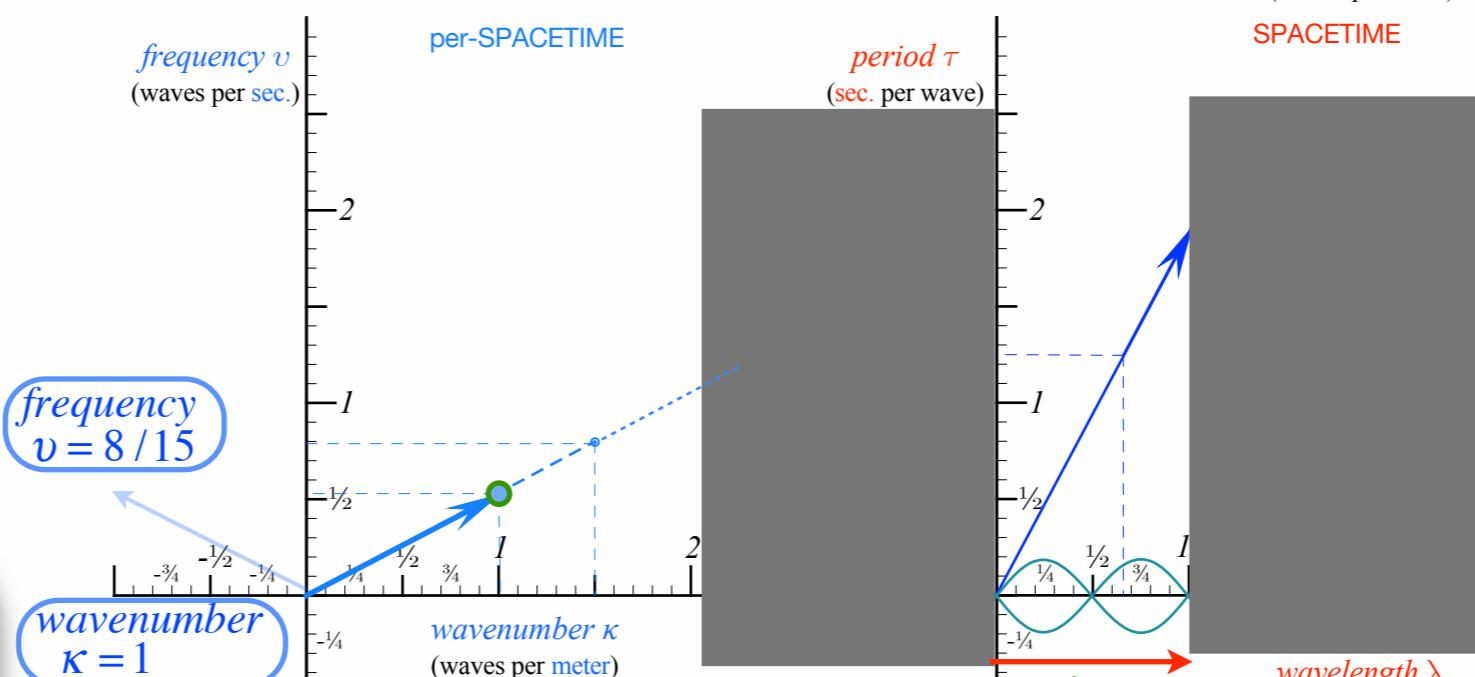
If a wave is confined to an $L=1\text{m.}$ box
the “Keyboard of the gods” has its
wavenumber κ is “quantized” to
multiples of $1/2L=1/2.$

$$\kappa = \frac{1}{2}, \frac{2}{2}, \frac{3}{2}, \frac{4}{2}, \dots$$



period
 $\tau = 5/4 = 1/v$

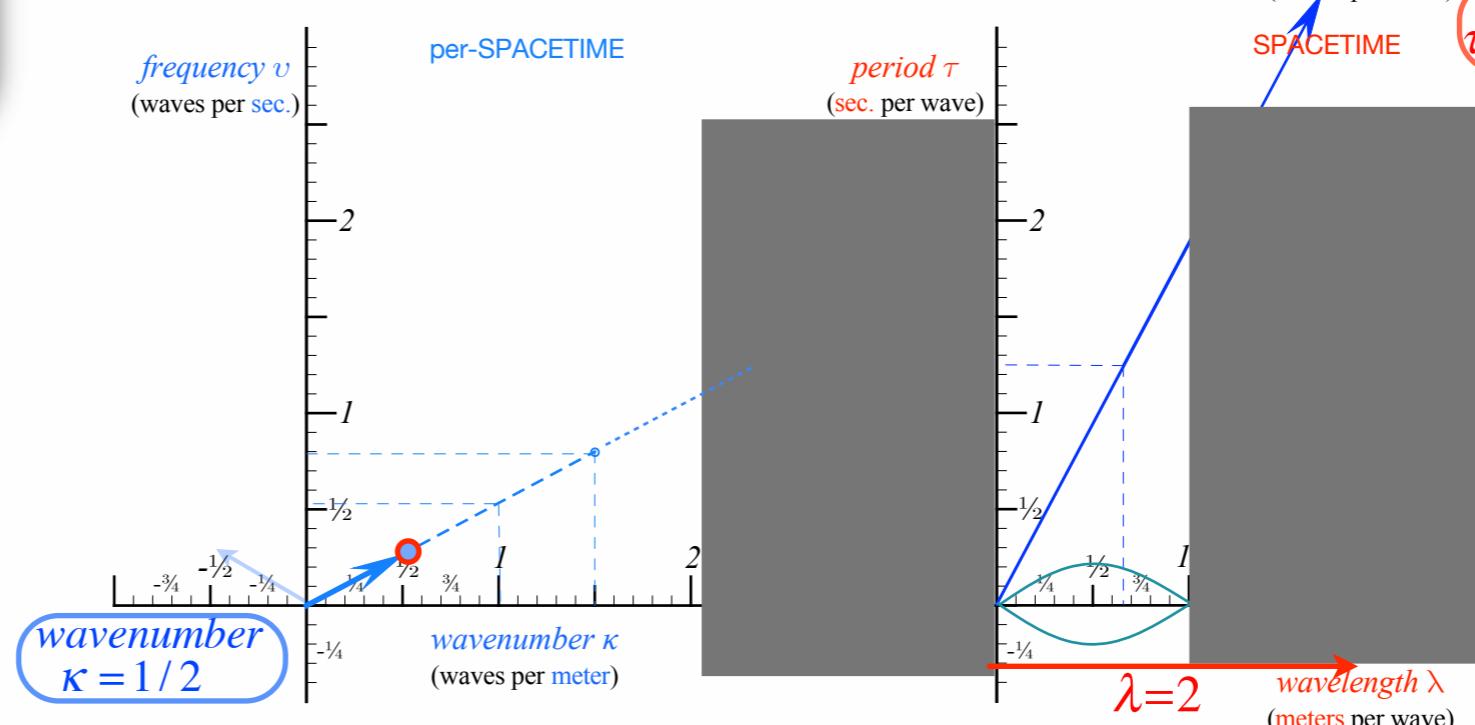
wavelength
 $\lambda = 2/3 = 1/\kappa$



period
 $\tau = 15/8 = 1/v$

wavelength
 $\lambda = 1 = 1/\kappa$

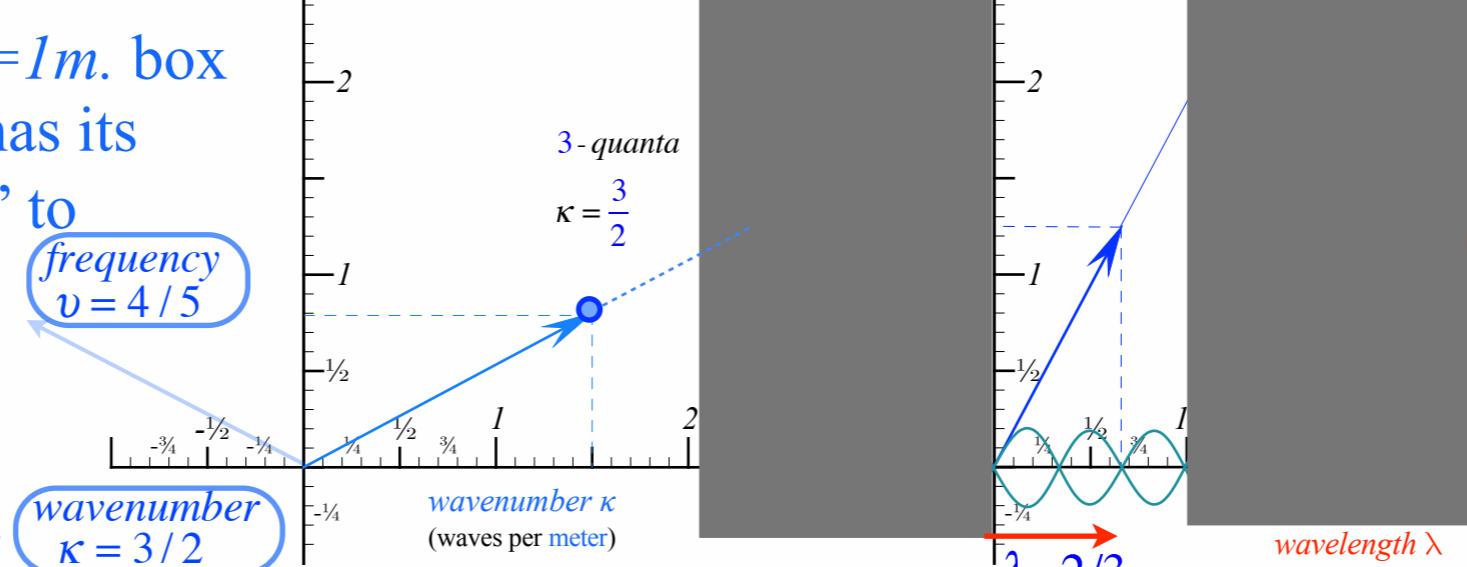
- How to understand waves
and
“1st quantization”
or κ -quantization



wavelength
 $\lambda = 2 = 1/\kappa$

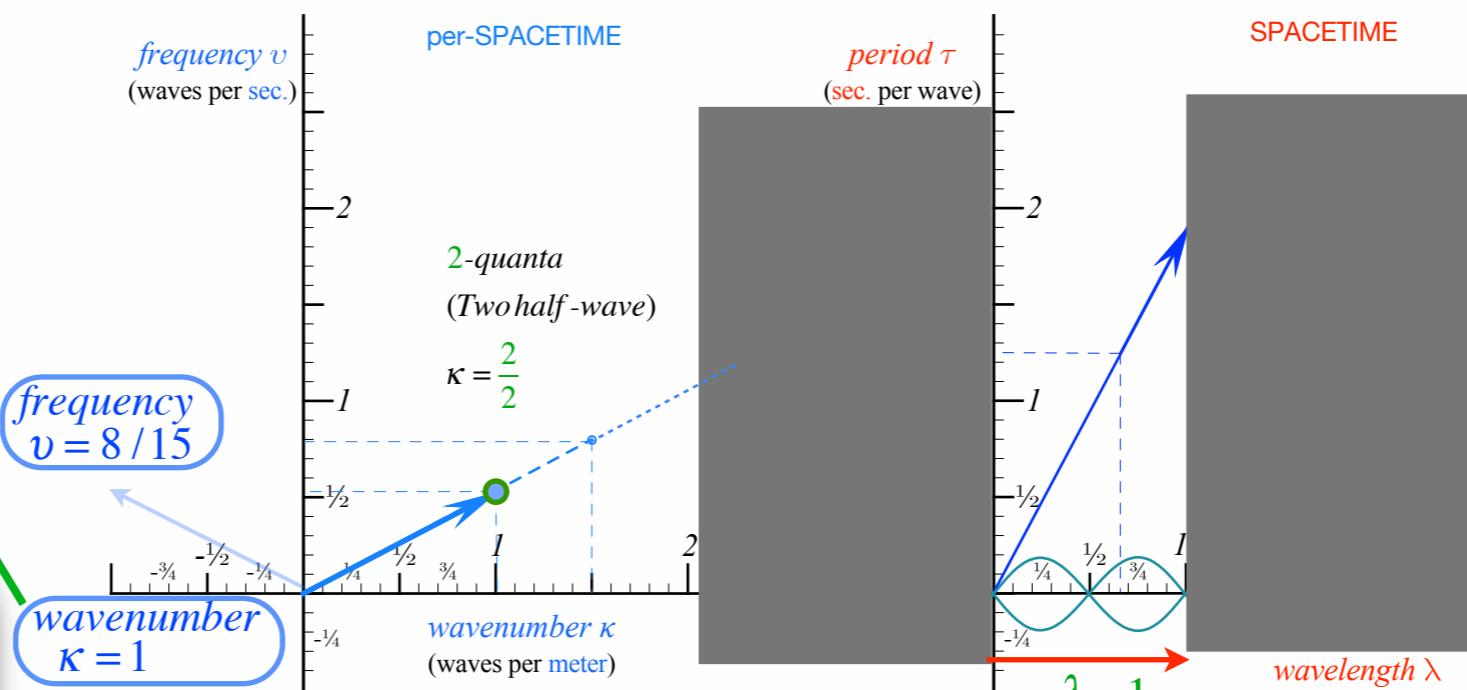
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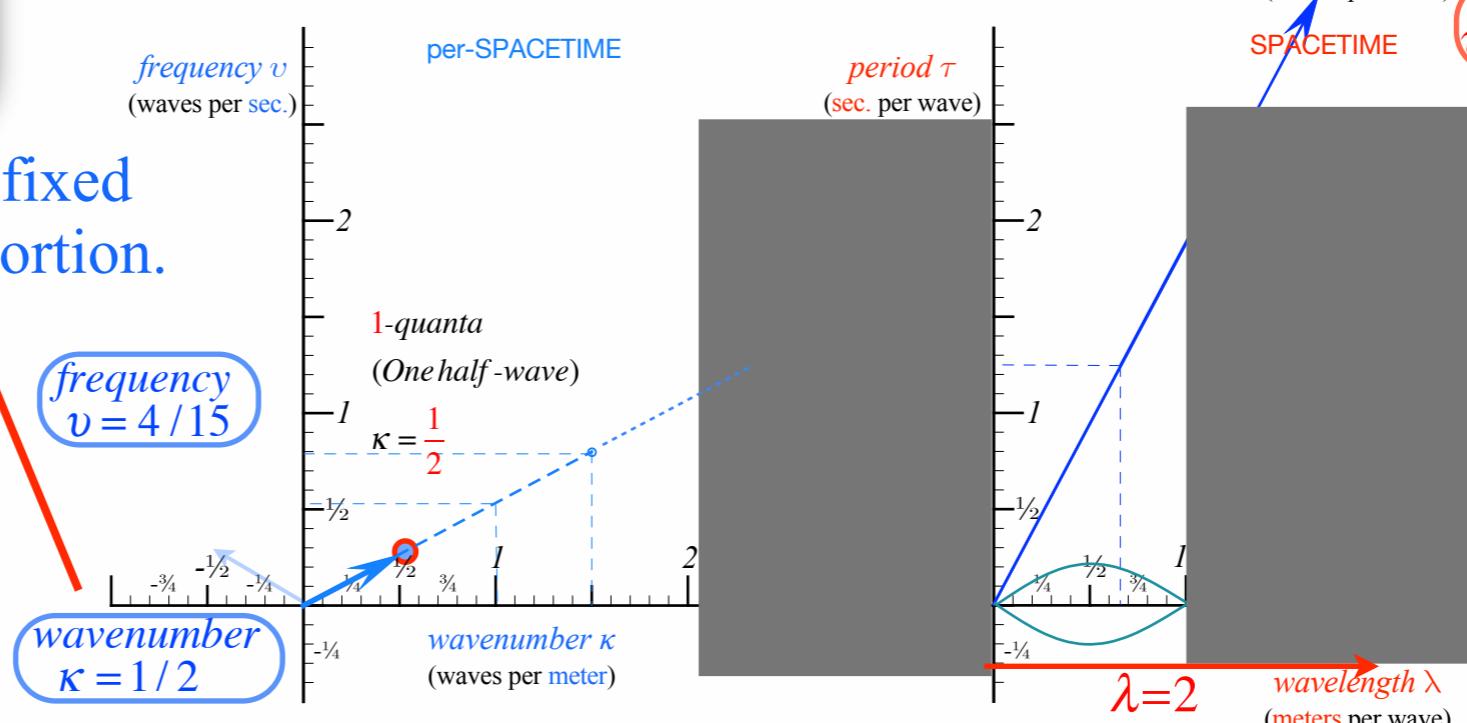
$$\text{period } \tau = 5/4 = 1/v$$

$$\text{wavelength } \lambda = 2/3 = 1/\kappa$$



$$\text{period } \tau = 15/8 = 1/v$$

$$\text{wavelength } \lambda = 1 = 1/\kappa$$



$$\text{period } \tau = 15/4 = 1/v$$

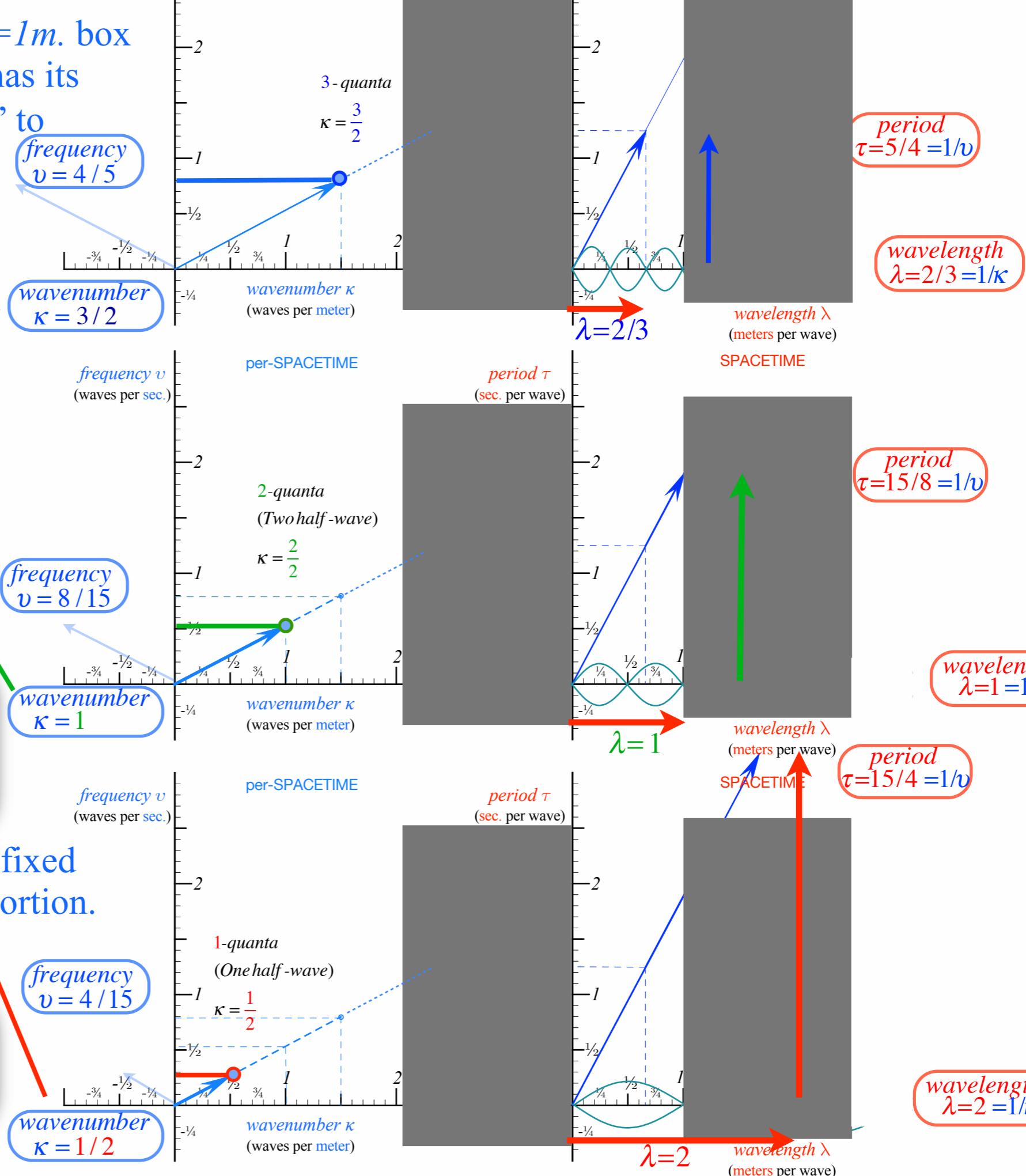
$$\text{wavelength } \lambda = 2 = 1/\kappa$$

**•How to understand waves
and
“1st quantization”
or κ-quantization**

If wave velocity $V_{\text{wave}}=8/15$ is fixed
frequency is quantized in proportion.

If a wave is confined to an $L=1\text{m}$. box
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$$\kappa = 1/2, 1, 3/2, 2, \dots$$

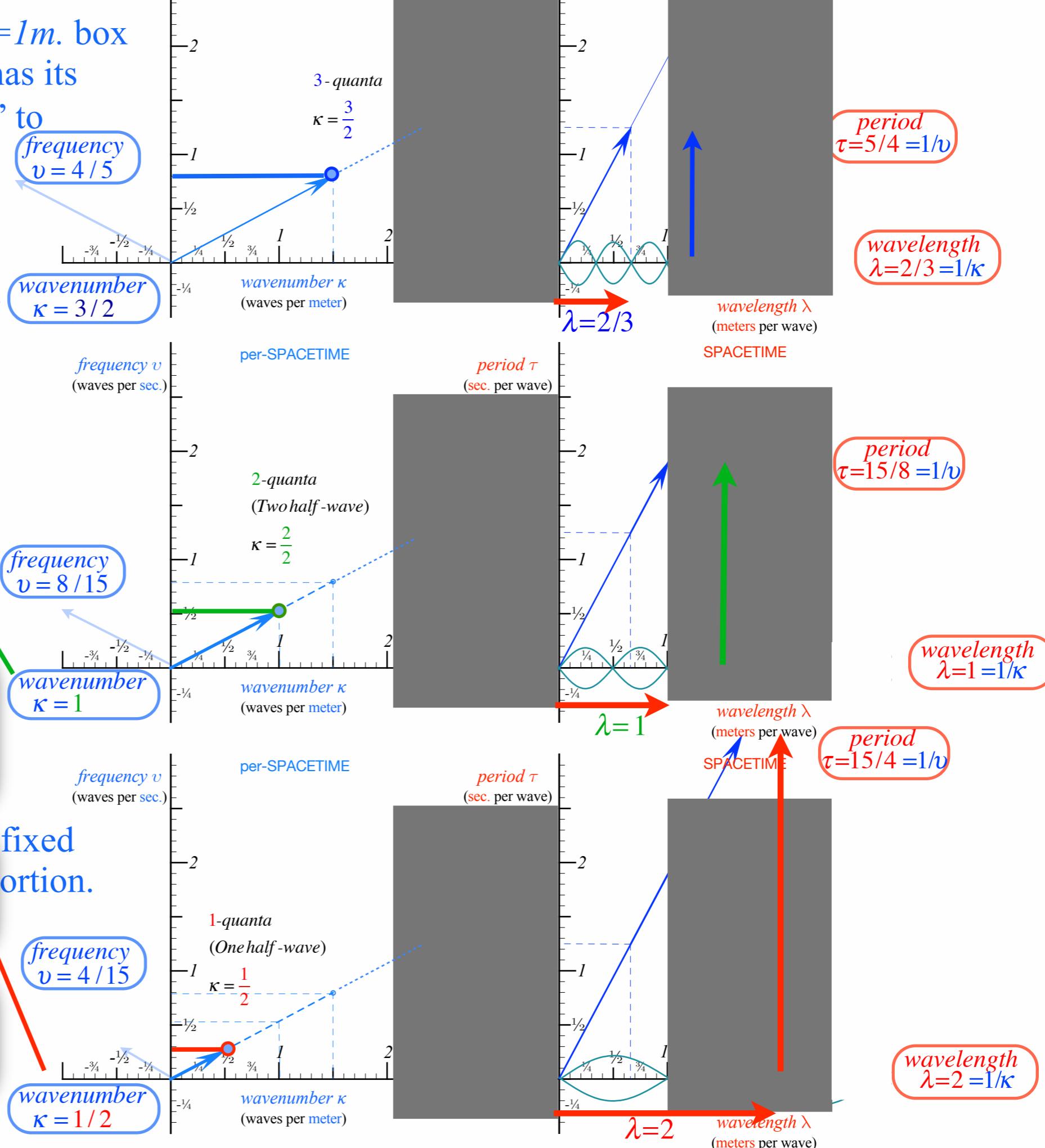
- How to understand waves and “1st quantization” or κ -quantization

If wave velocity $V_{\text{wave}}=8/15$ is fixed
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- Amplitude A -quantization is called “2nd quantization”

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$$\kappa = \frac{1}{2}, \frac{2}{2}, \frac{3}{2}, \frac{4}{2}, \dots$$



•How to understand waves and
“1st quantization”
or κ -quantization

...as *QUALITY* (color)
versus
QUANTITY (^{Number of photons})

•Amplitude A -quantization
is called
“2nd quantization”

As will be shown:

Light wave-velocity c is *VERY* fixed

$$V_{light} = c = \frac{\lambda}{\tau} = \frac{1/\kappa}{1/\nu} = \frac{\nu}{\kappa} = \frac{1/\tau}{1/\lambda} = 299,792,458 \frac{m}{s}.$$

As will be shown:

Light wave-velocity c is *VERY* fixed

$$V_{light} = c = \frac{v}{\kappa} = \frac{1/\kappa}{1/v} = \frac{\lambda}{\tau} = \frac{1/\tau}{1/\lambda} = 299,792,458 \frac{m}{s}.$$

Then it's convenient to use:

...or angular variables: $\omega = 2\pi v$

and: $k = 2\pi\kappa$

Dimensionless **Light** wave-velocity $c/c=1$

$$\frac{V_{light}}{c} = \frac{v}{c\kappa} = \frac{\lambda}{c\tau} = 1 \quad \text{instead of: } \frac{v}{\kappa} = \frac{\lambda}{\tau} = c$$

$$\frac{v}{\kappa} = \frac{\lambda}{\tau} = c = \frac{\omega}{k}$$

Such graphs use c -units of per-time $v=c\kappa$ and length $\lambda=c\tau$.

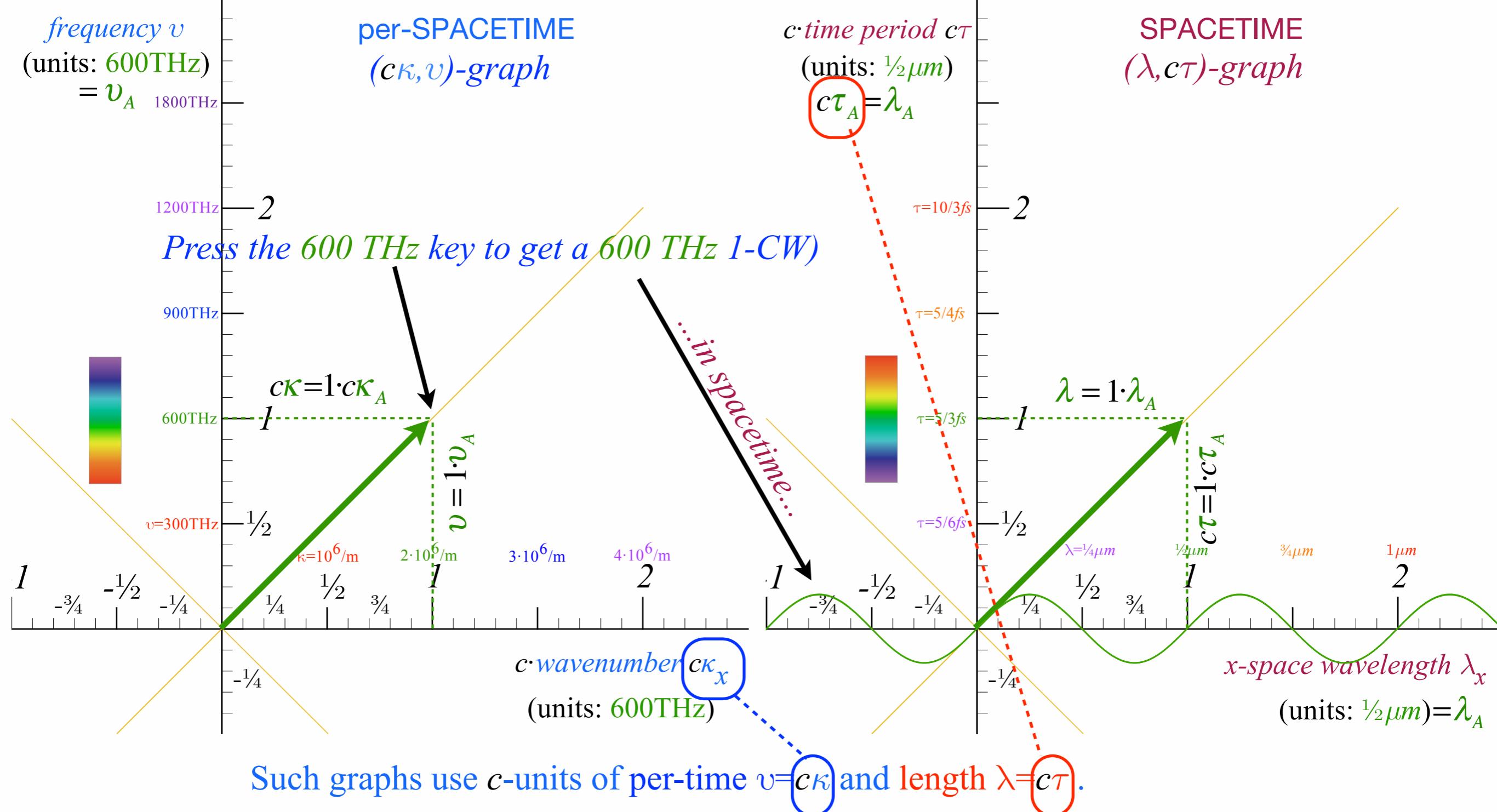
$$\frac{V_{light}}{c} = \frac{v}{c\kappa} = \frac{1/\kappa}{c/v} = \frac{\lambda}{c\tau} = \frac{1/\tau}{c/\lambda} = 1$$

$$\frac{v}{c\kappa} = 1 = \frac{\omega}{ck}$$

$$ck = \omega$$

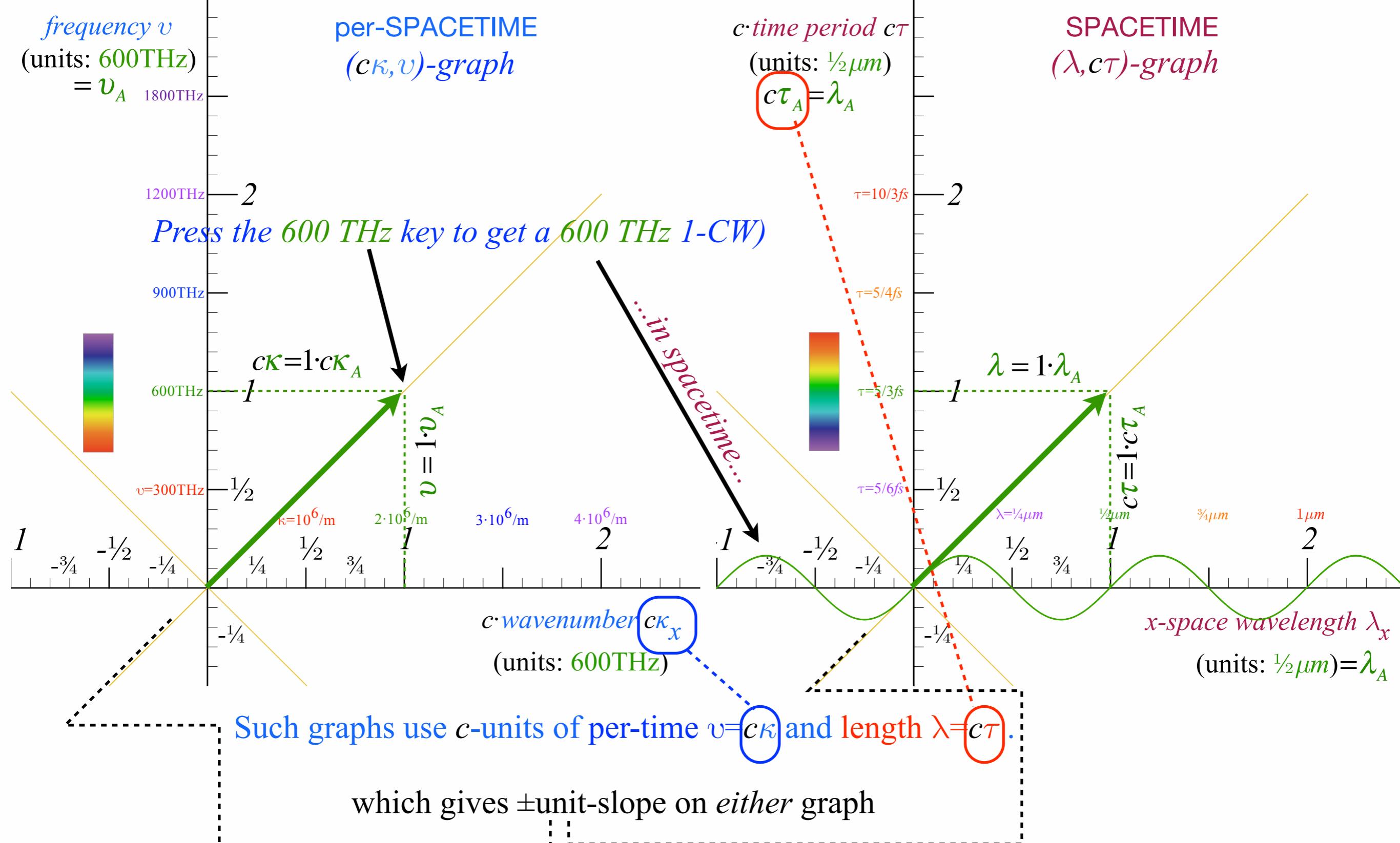
$$c\kappa = v$$

The “Keyboard of the gods” : Introducing per-space-per-time graphs versus space-time graphs



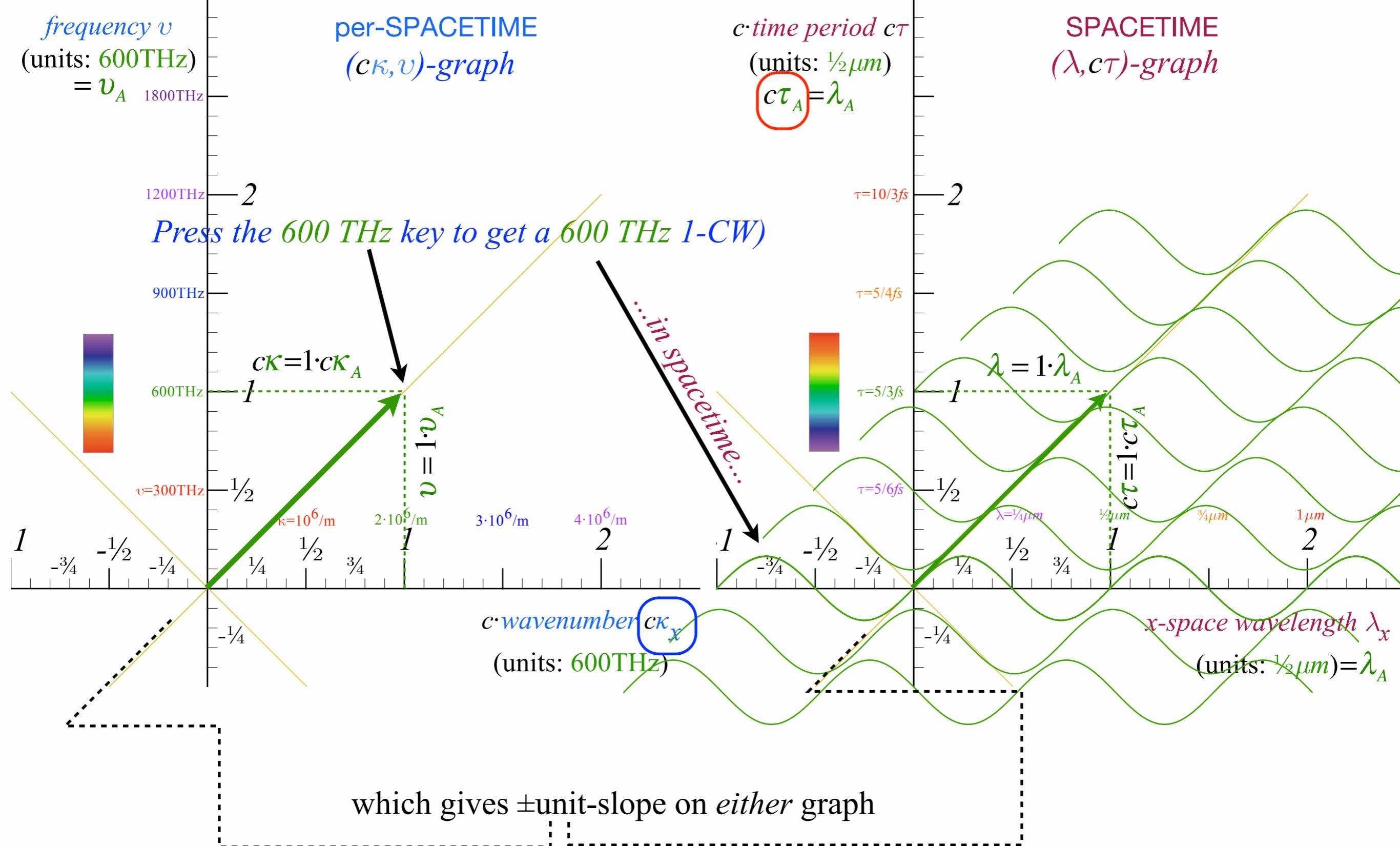
Ways to quantify **light** waves (600 THz example)

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Ways to quantify **light** waves (600 THz example)

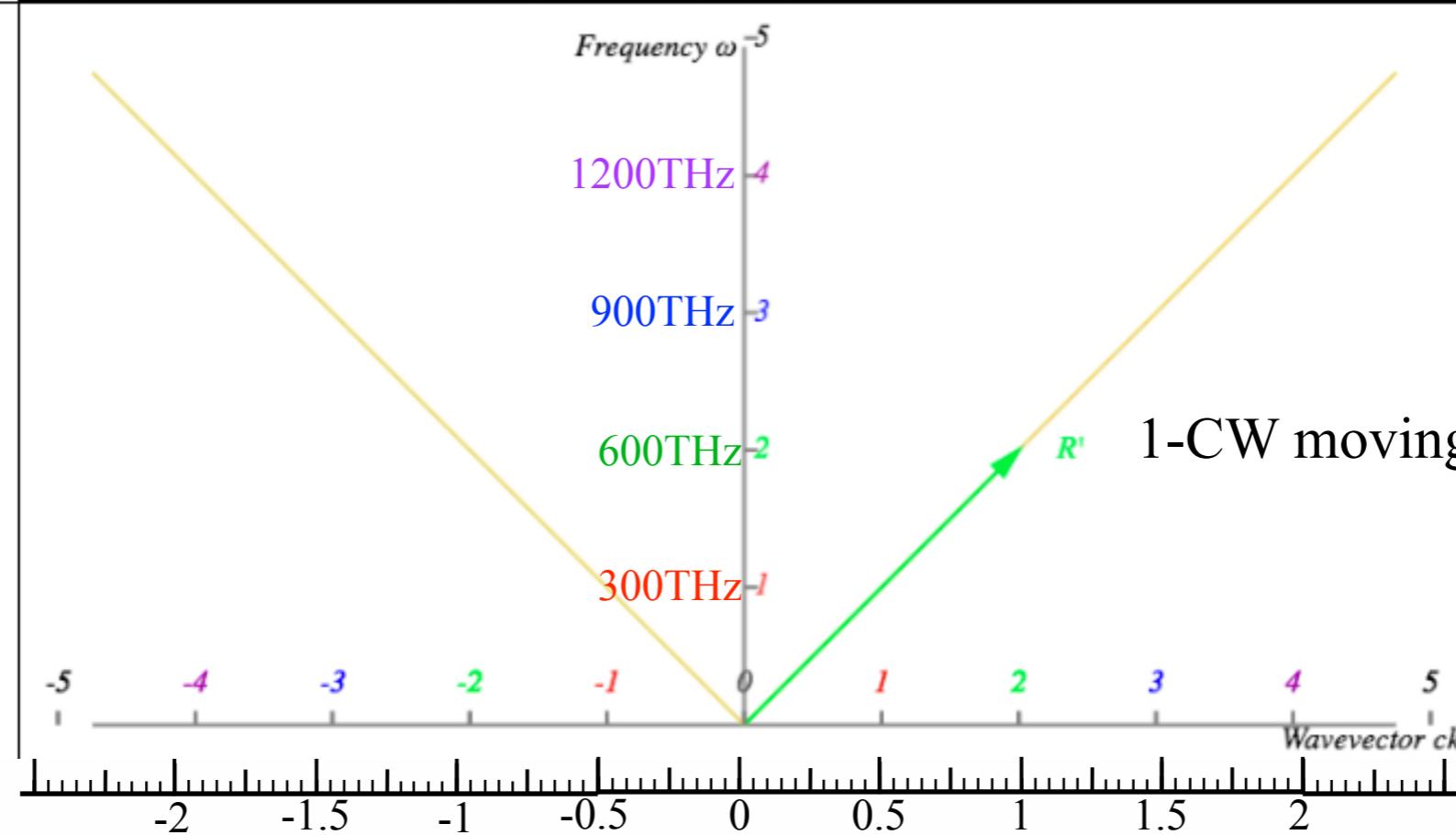
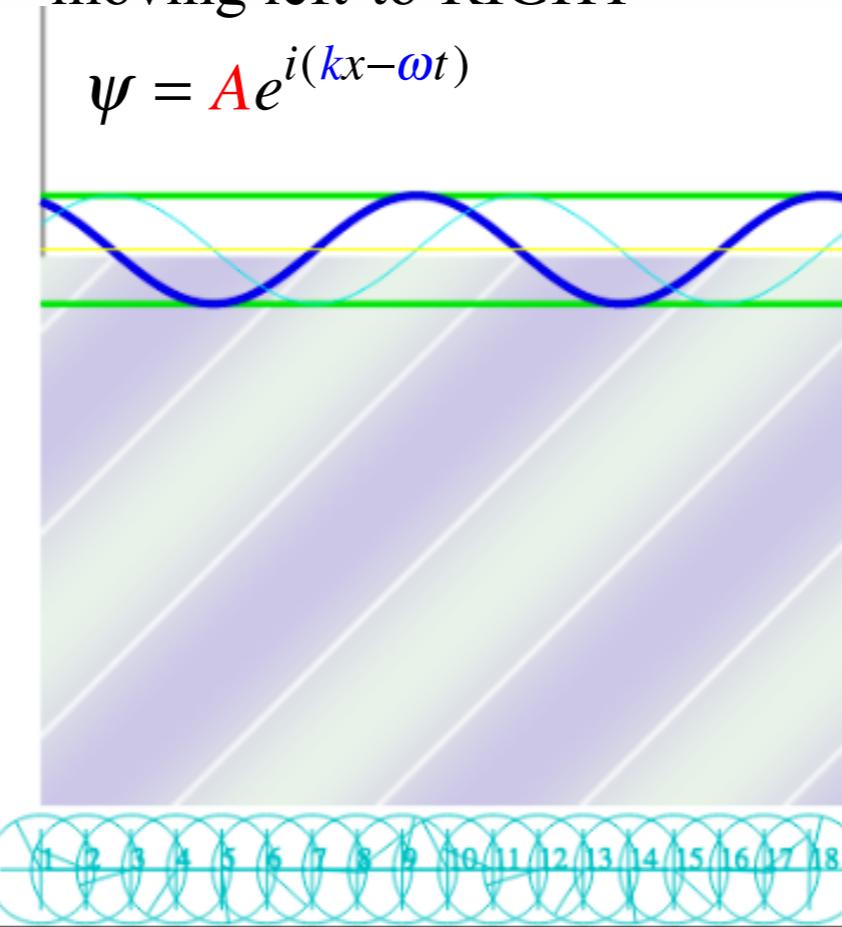
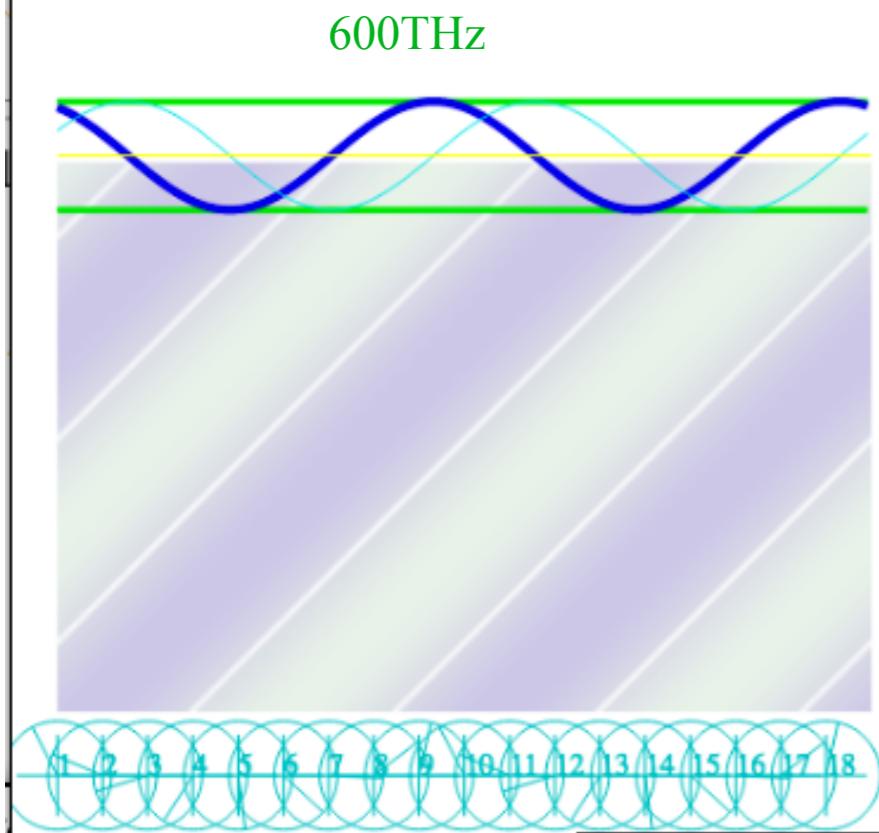
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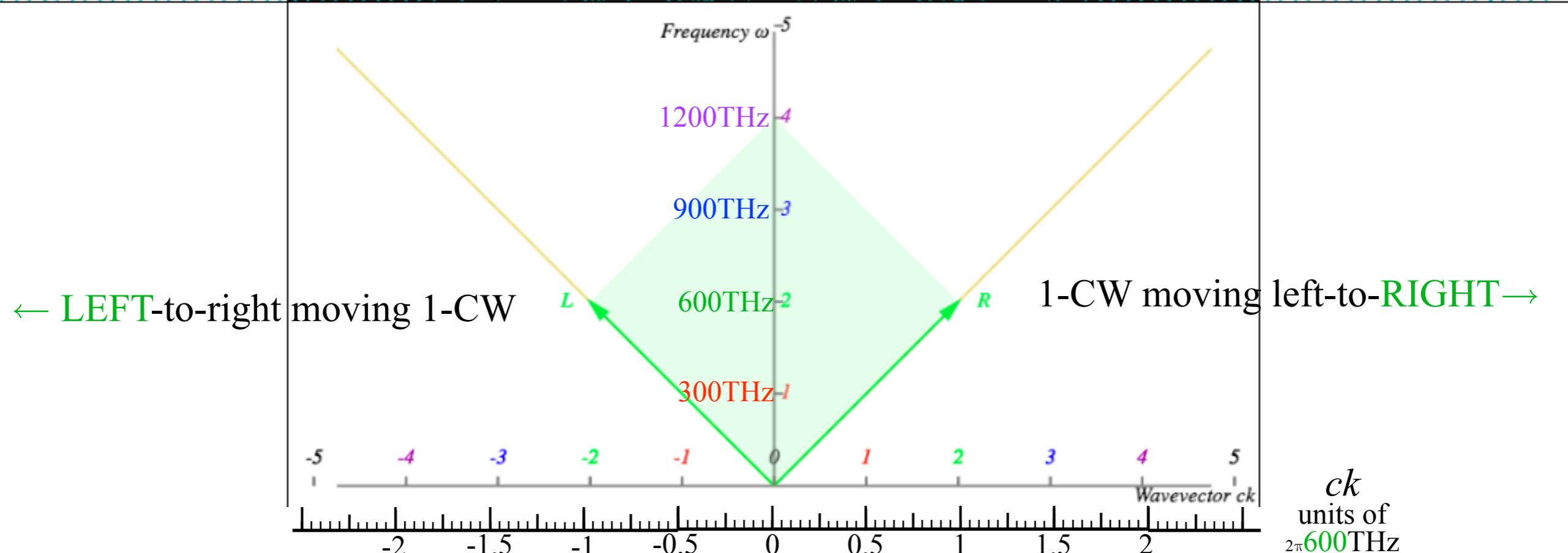
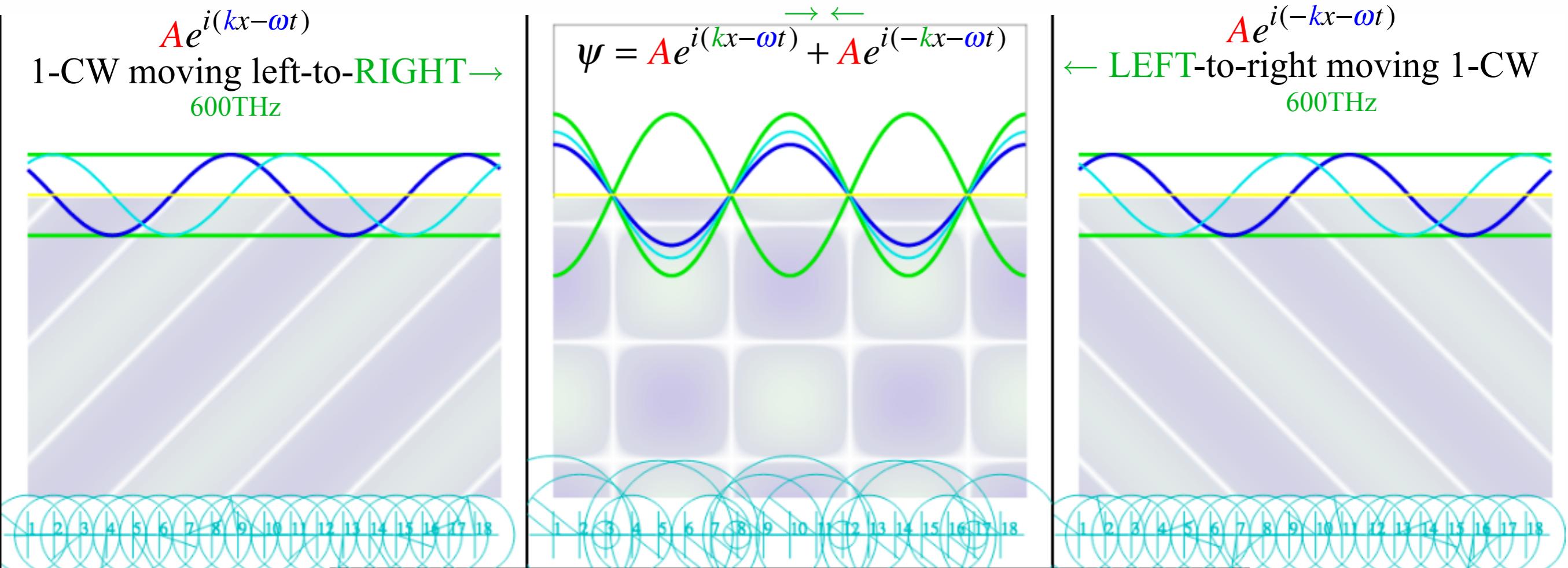
Ways to quantify **light** waves (600 THz example)

Single continuous wave (1-CW)
moving left-to-RIGHT →

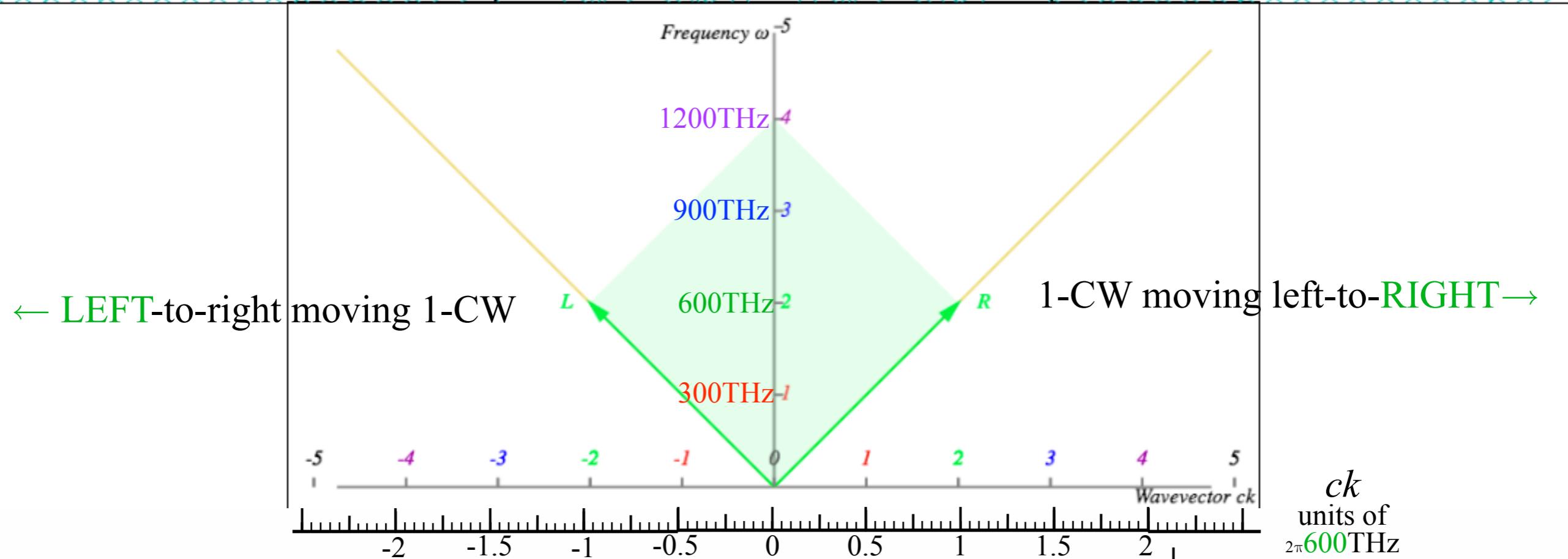
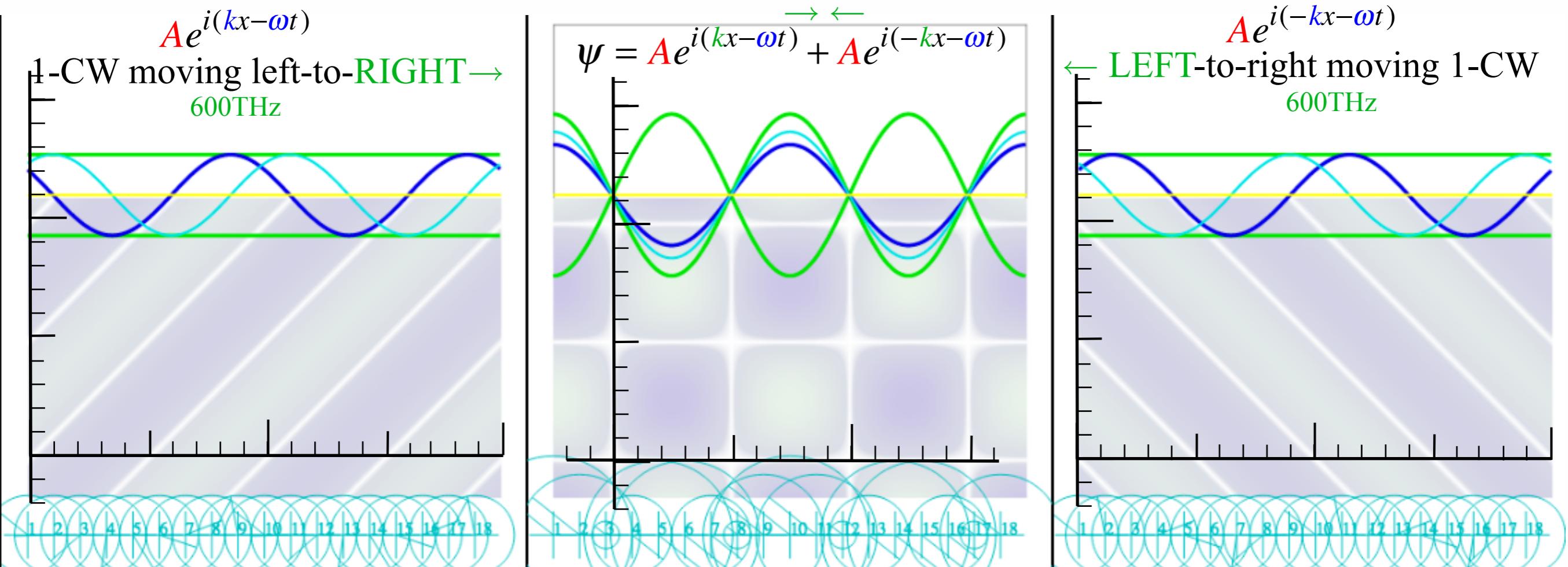
$$\psi = A e^{i(\mathbf{k}x - \omega t)}$$



2 colliding waves (2-CW)



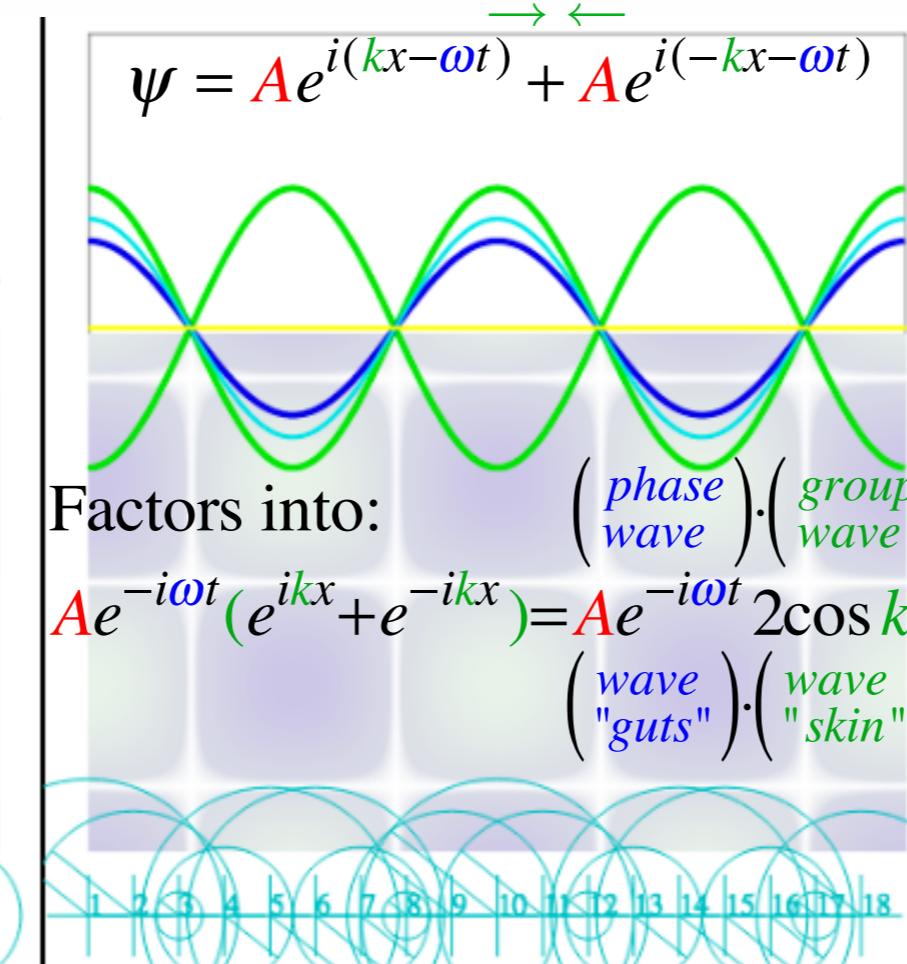
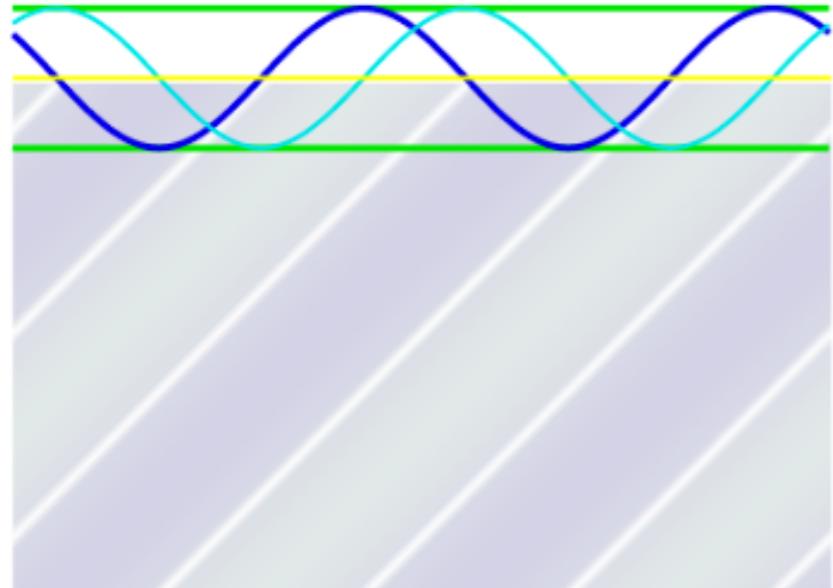
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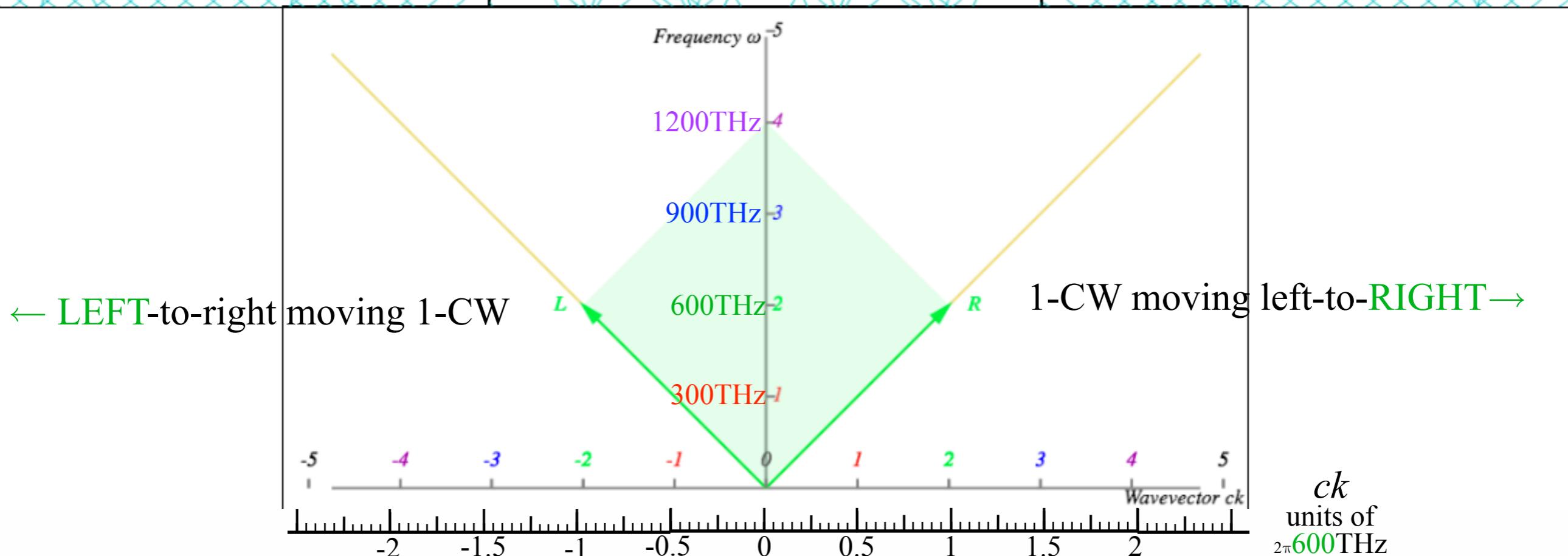
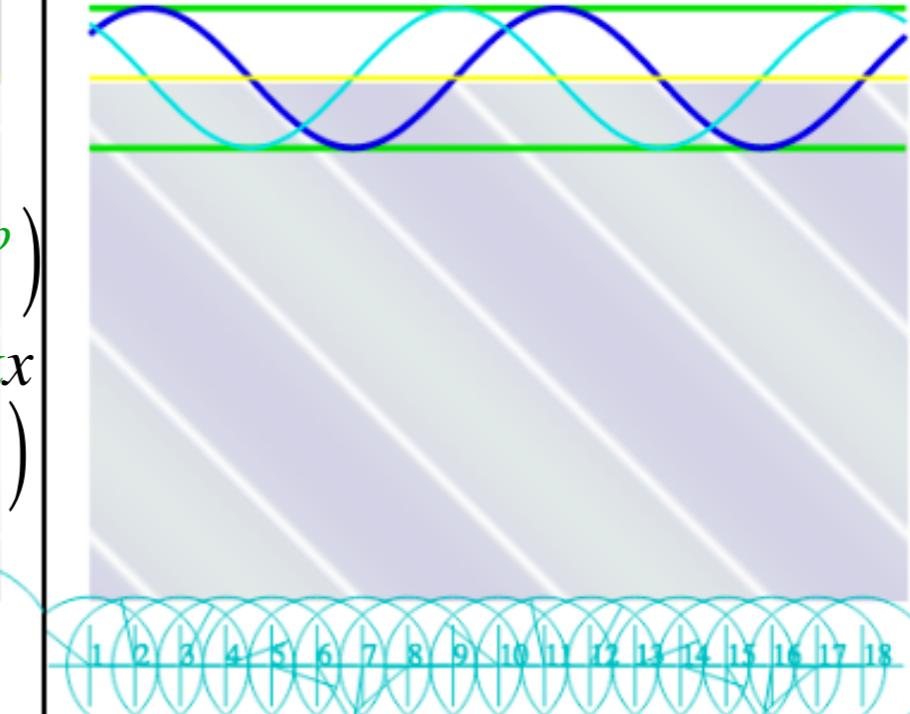
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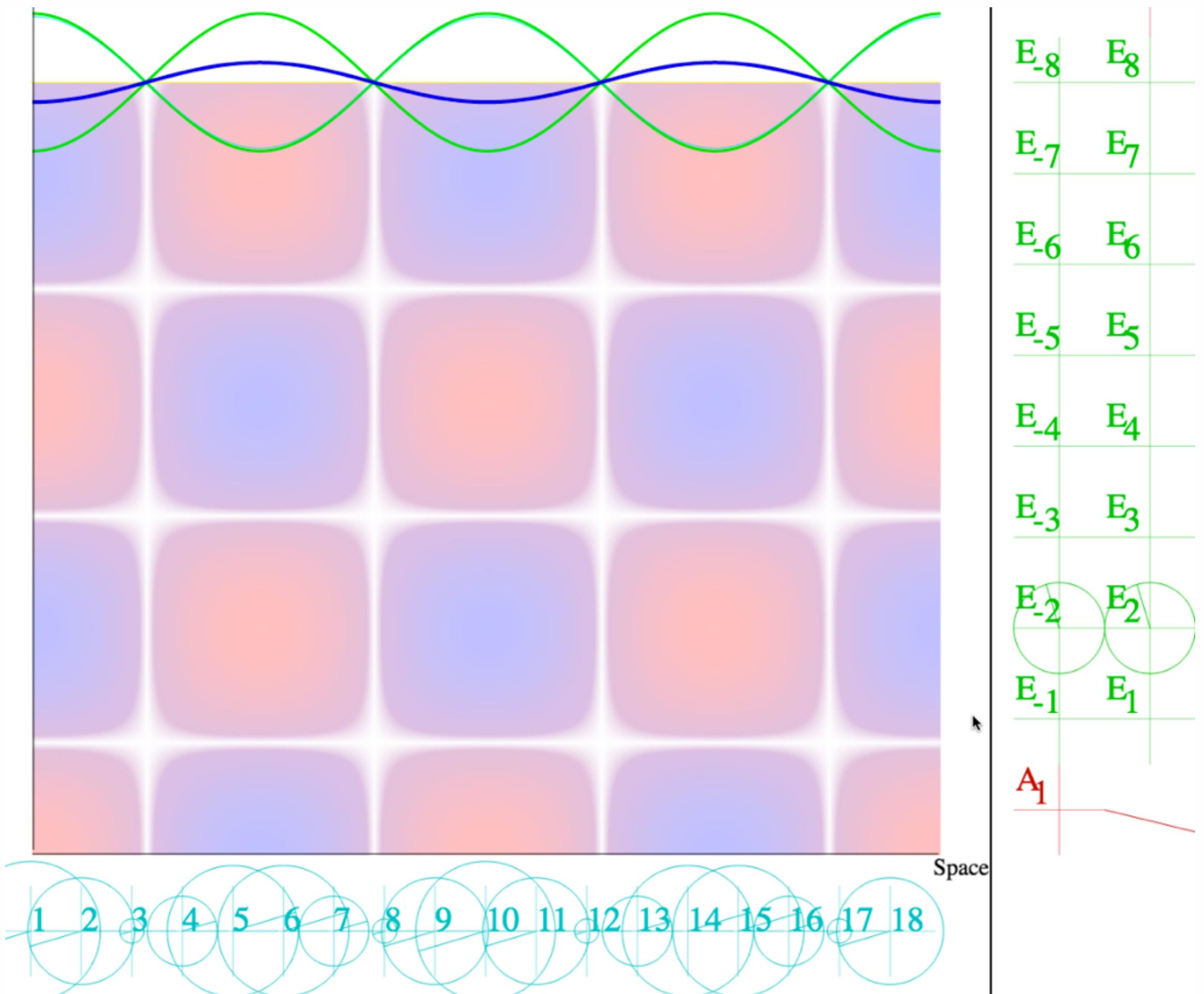
$$Ae^{i(kx-\omega t)}$$

1-CW moving left-to-**RIGHT**→
600THz

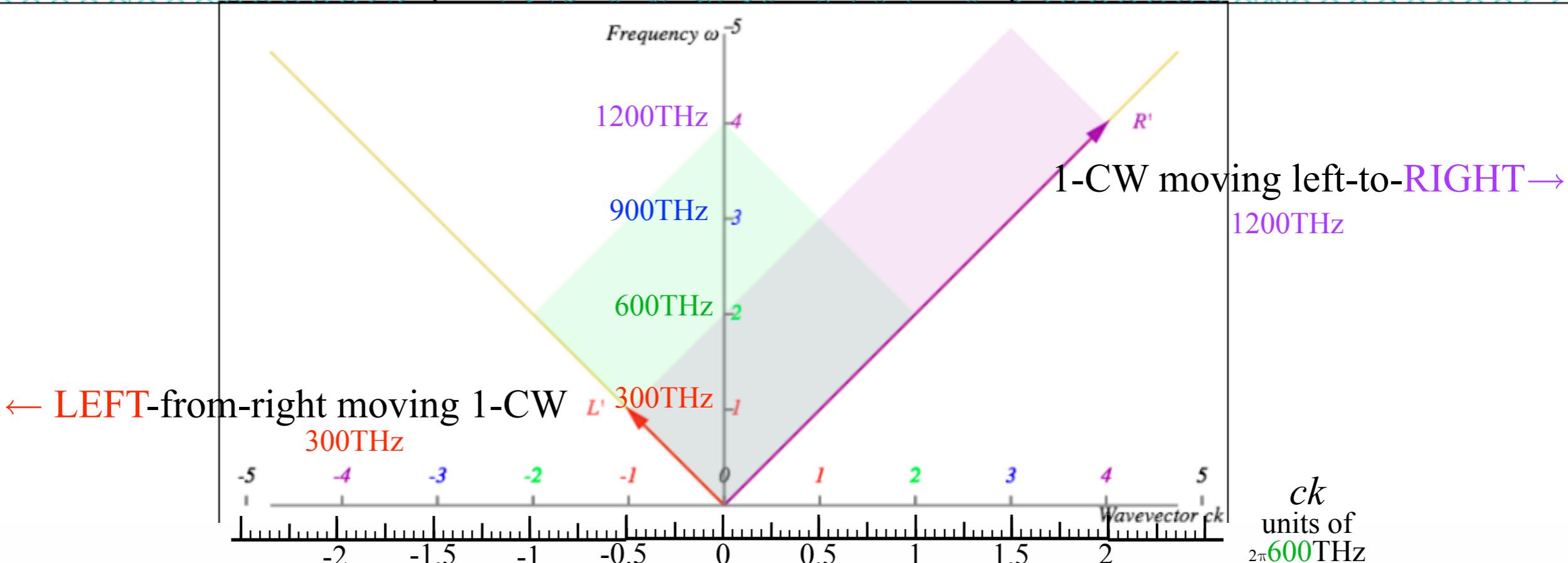
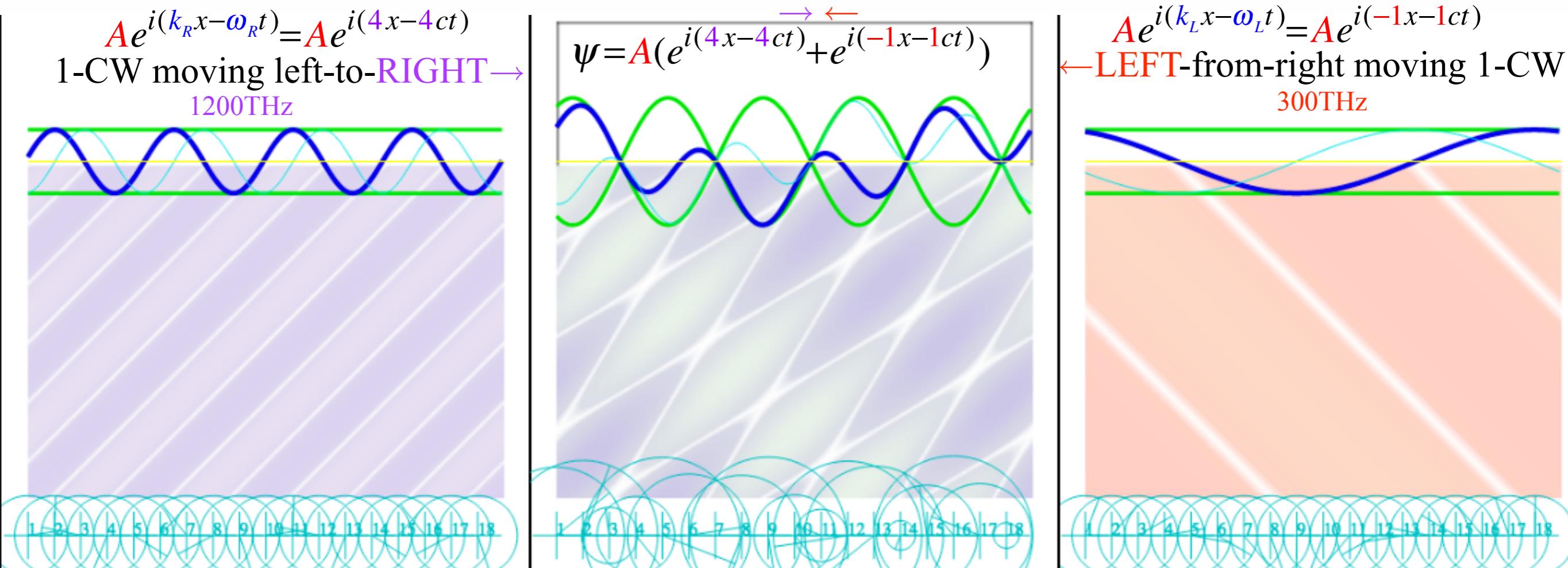


← LEFT-to-right moving 1-CW
600THz





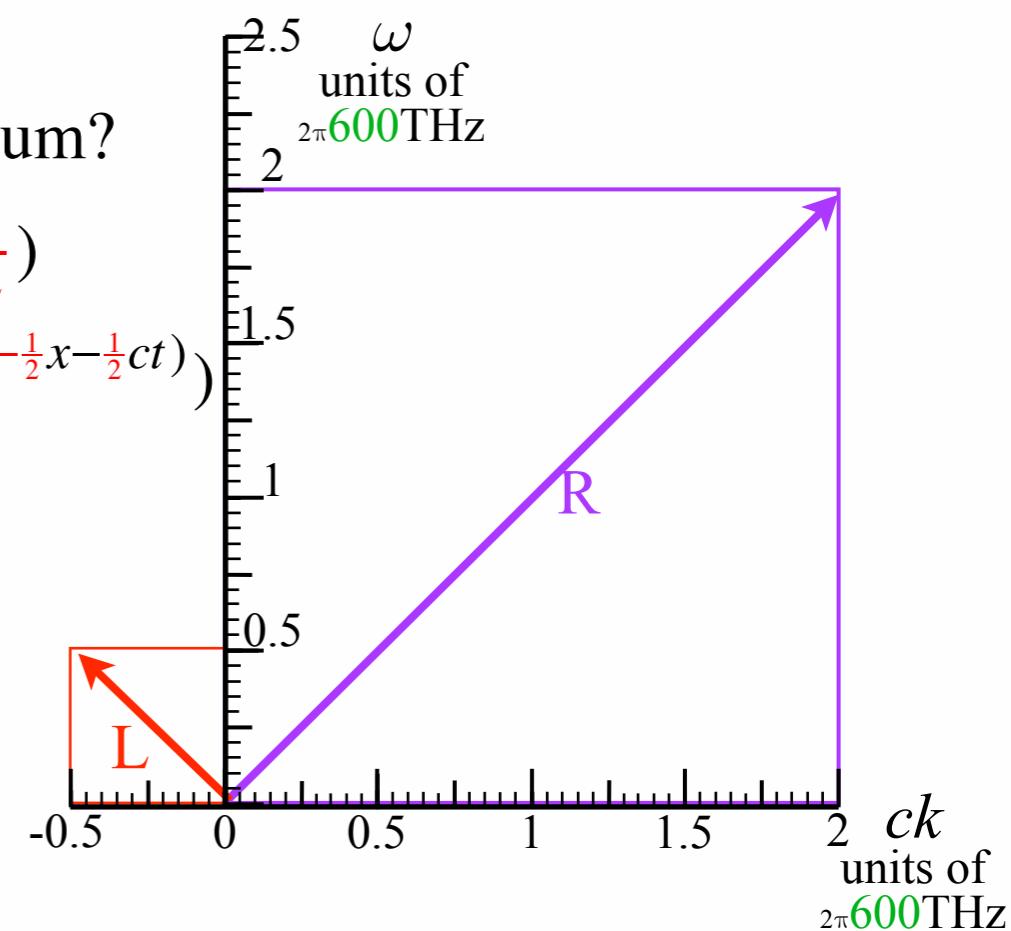
2 colliding waves (2-CW)



Where are the real-zeros of the colliding-light-wave (2-CW) sum?

RIGHT: ($\omega_R = c k_R = 2c$, $k_R = 2$) LEFT: ($\omega_L = -c k_L$, $k_L = -\frac{1}{2}$)

$$\Psi_{\omega_R k_R \omega_L k_L}(x, t) = A(e^{i(k_R x - \omega_R t)} + e^{i(k_L x - \omega_L t)}) = A(e^{i(2x - 2ct)} + e^{i(-\frac{1}{2}x - \frac{1}{2}ct)})$$

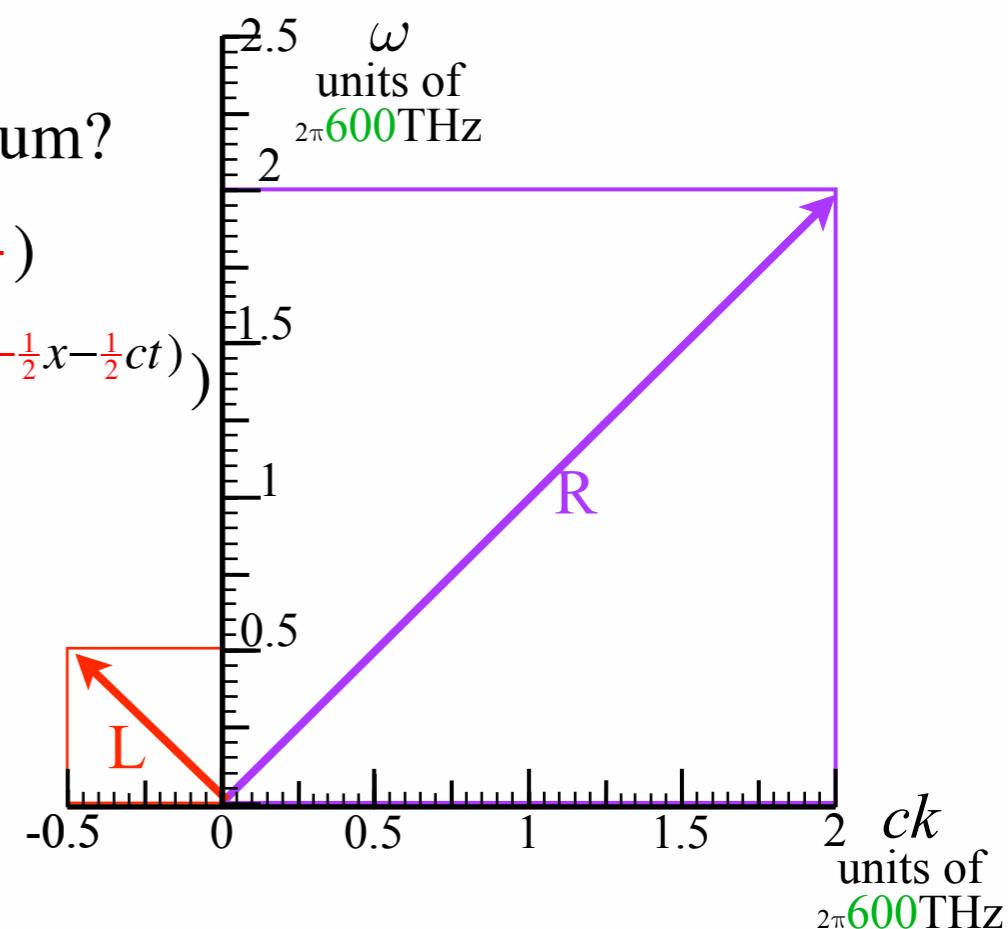


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To find zeros of a wave sum $e^{iR} + e^{iL}$ we need to factor it



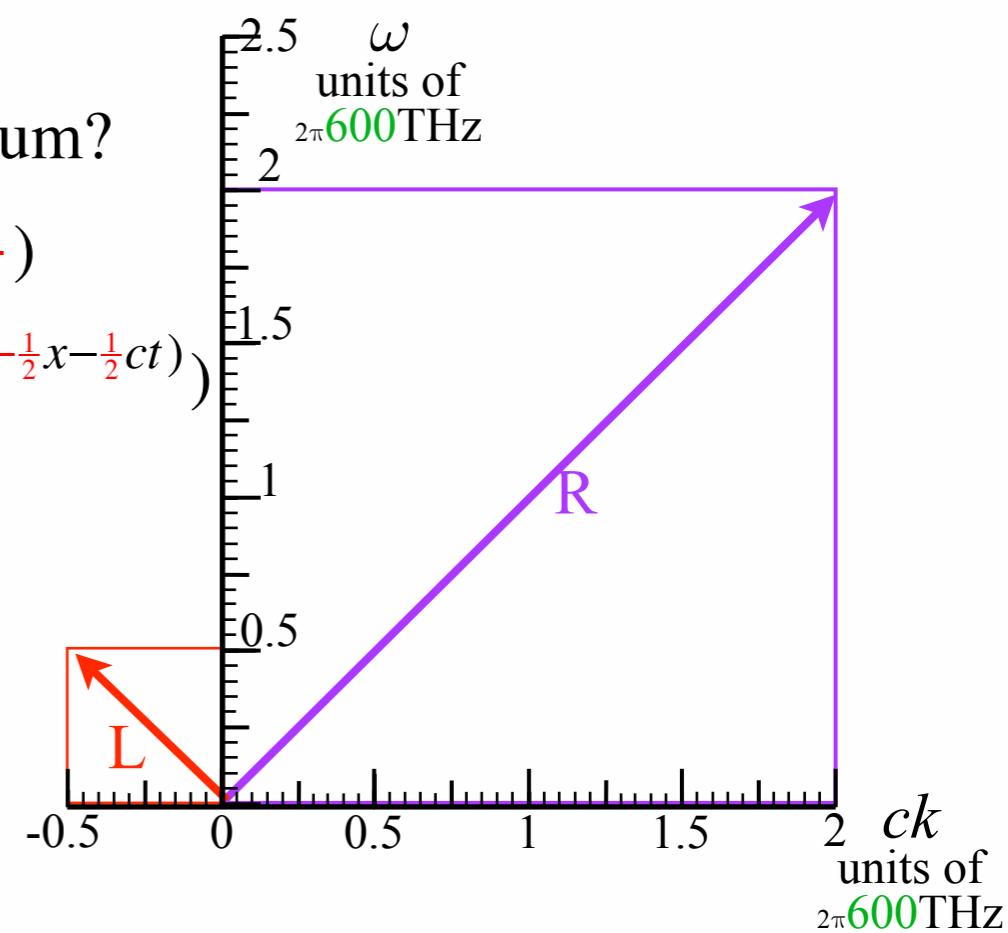
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To find zeros of a wave sum $e^{iR} + e^{iL}$ we need to factor it

$$e^{iR} + e^{iL} = e^{i\frac{R+L}{2}} \cdot \left(e^{i\frac{R-L}{2}} + e^{-i\frac{R-L}{2}} \right)$$



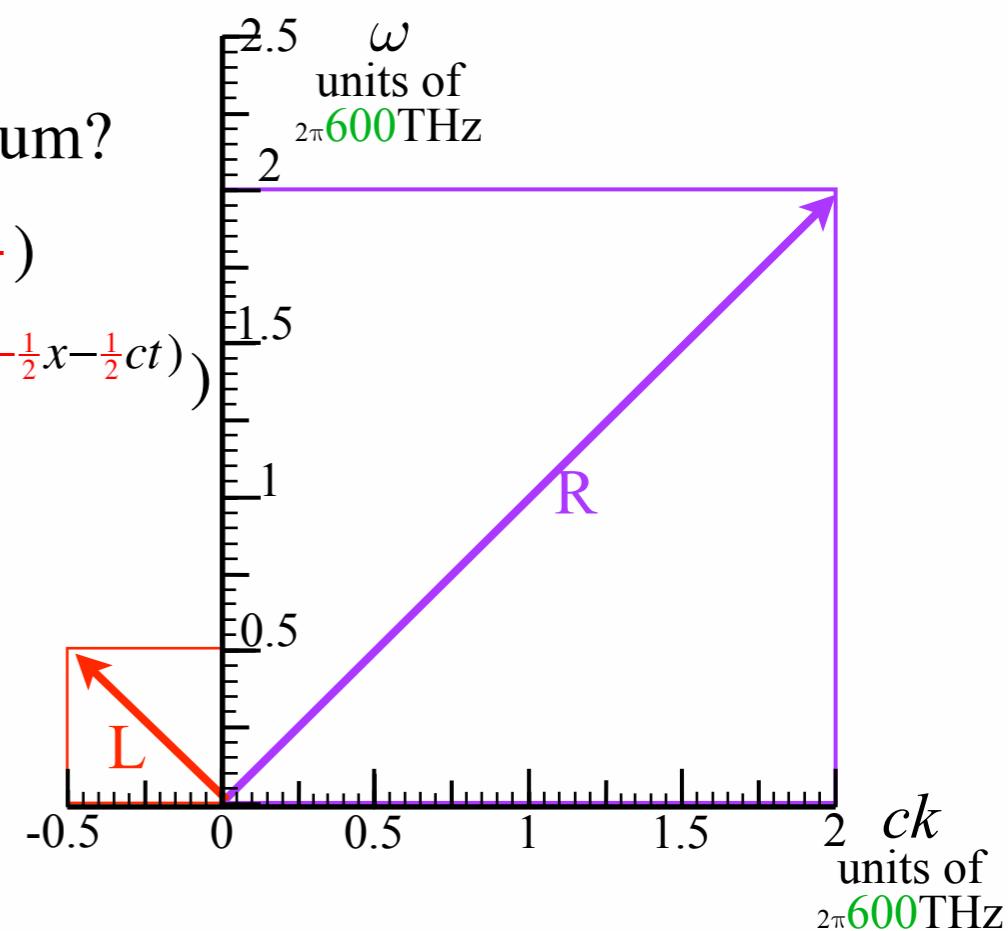
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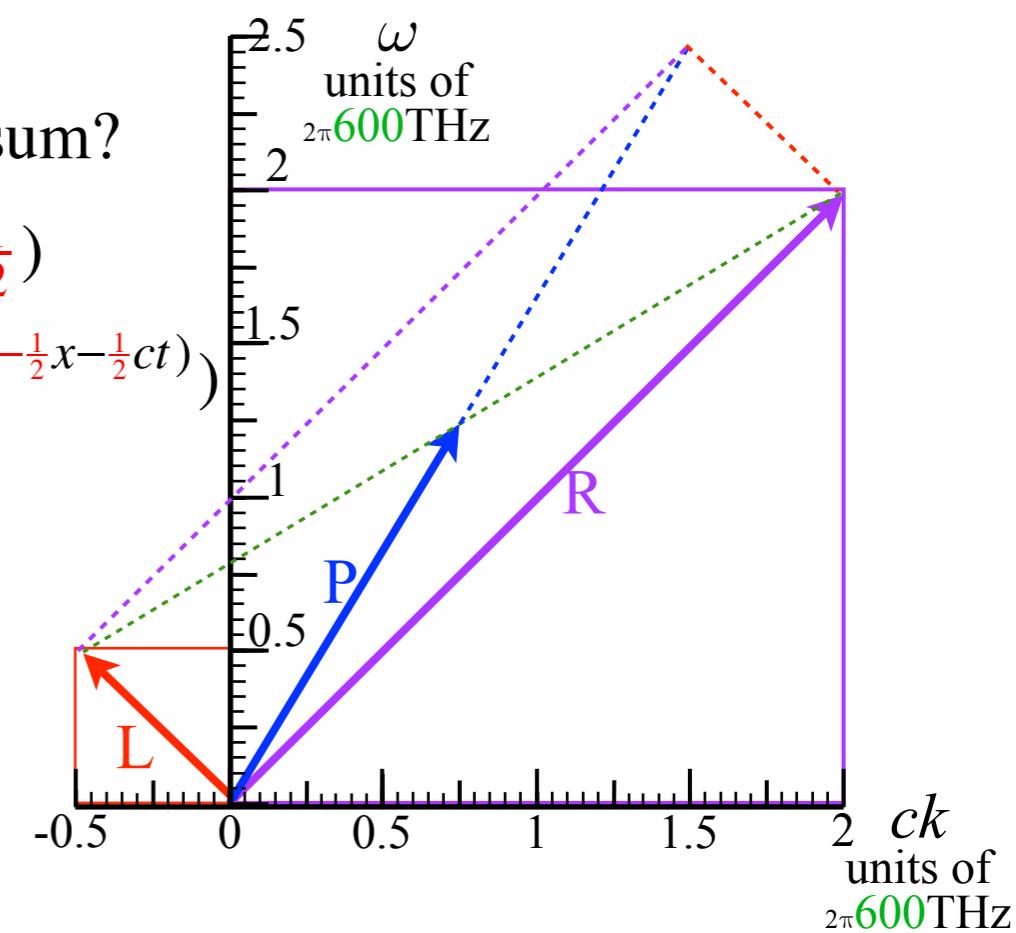
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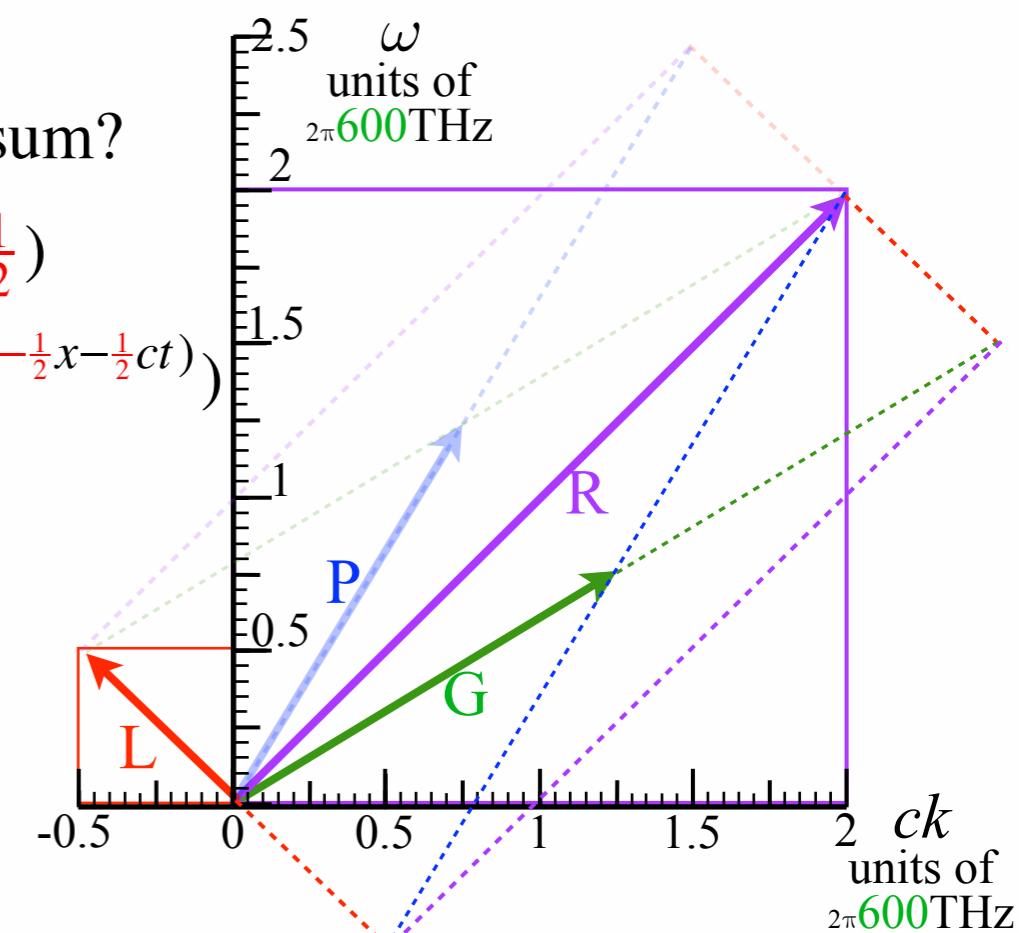
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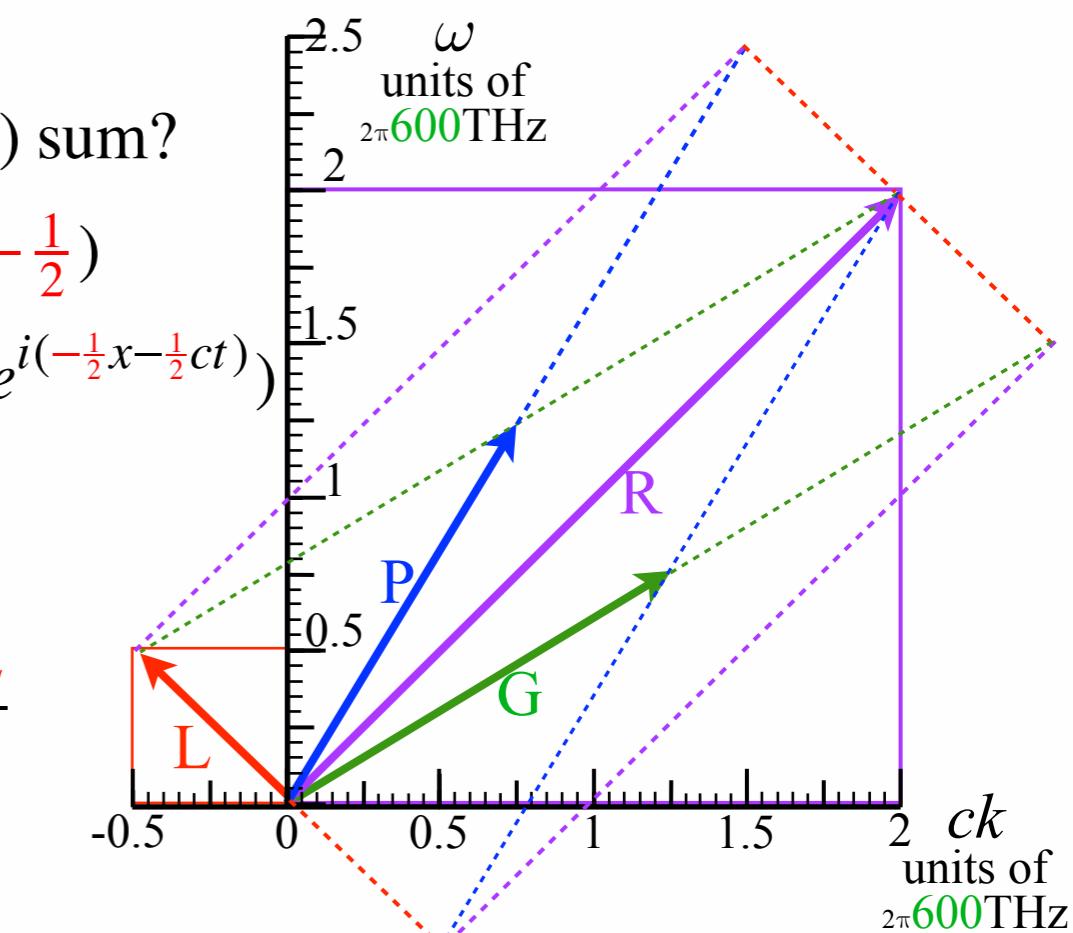
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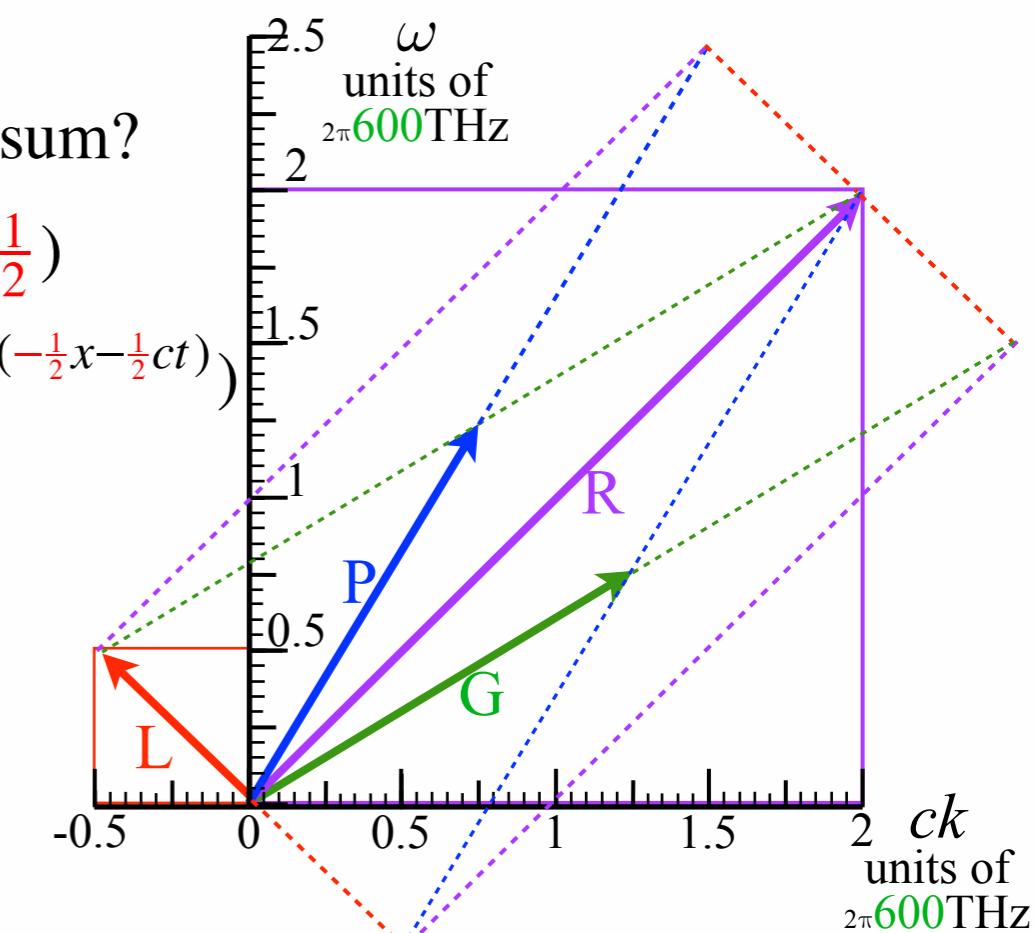
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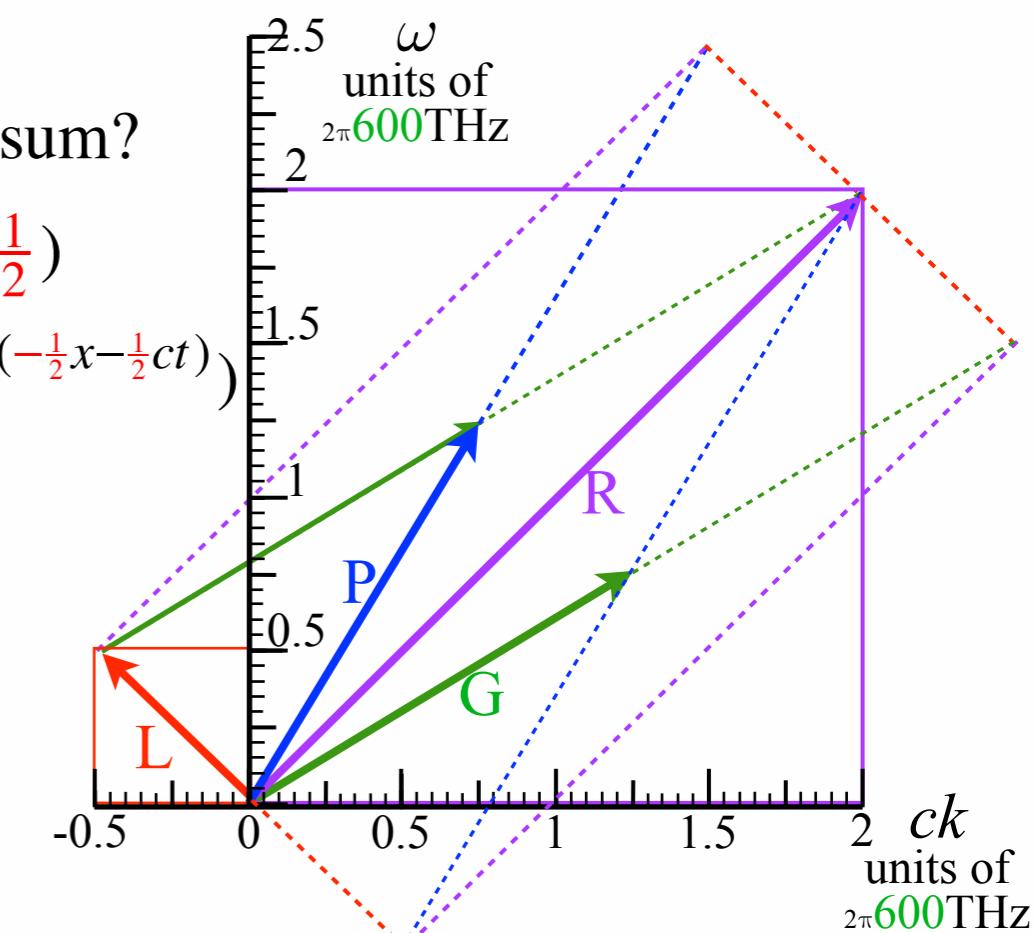
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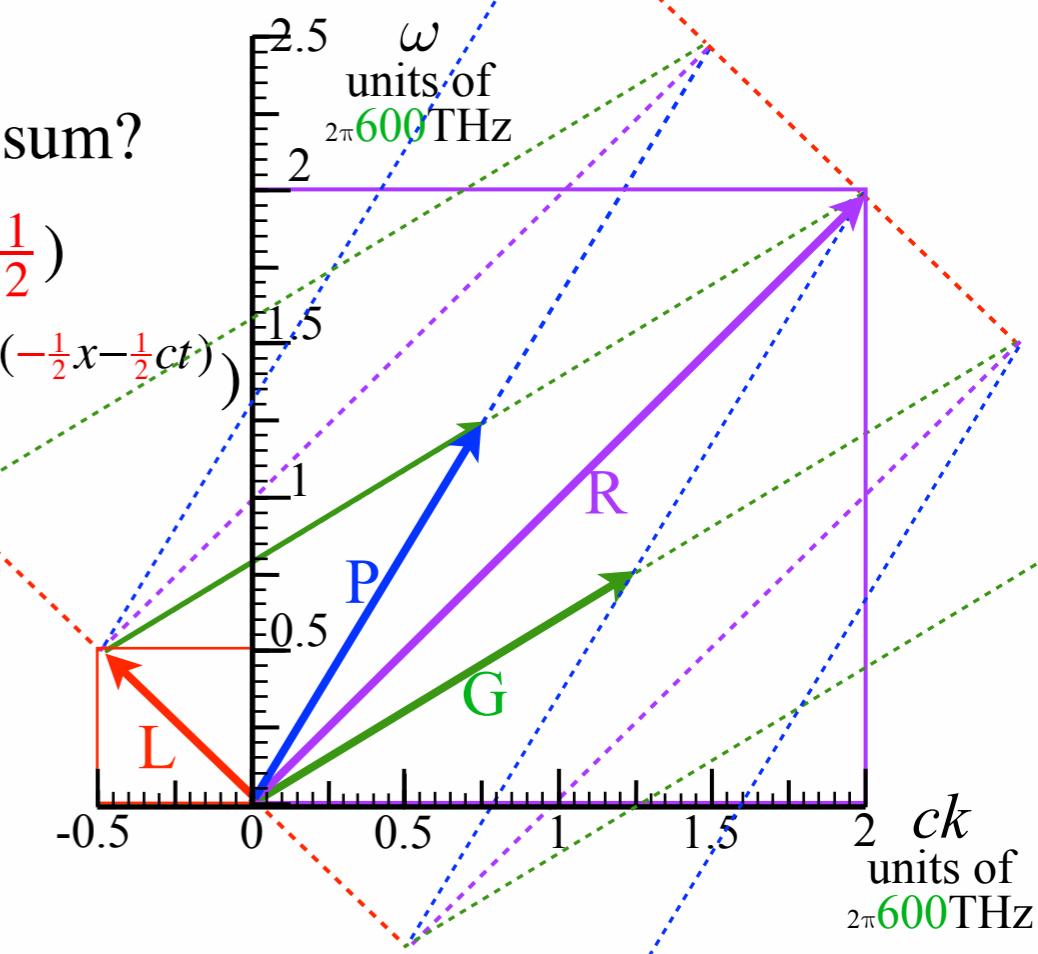
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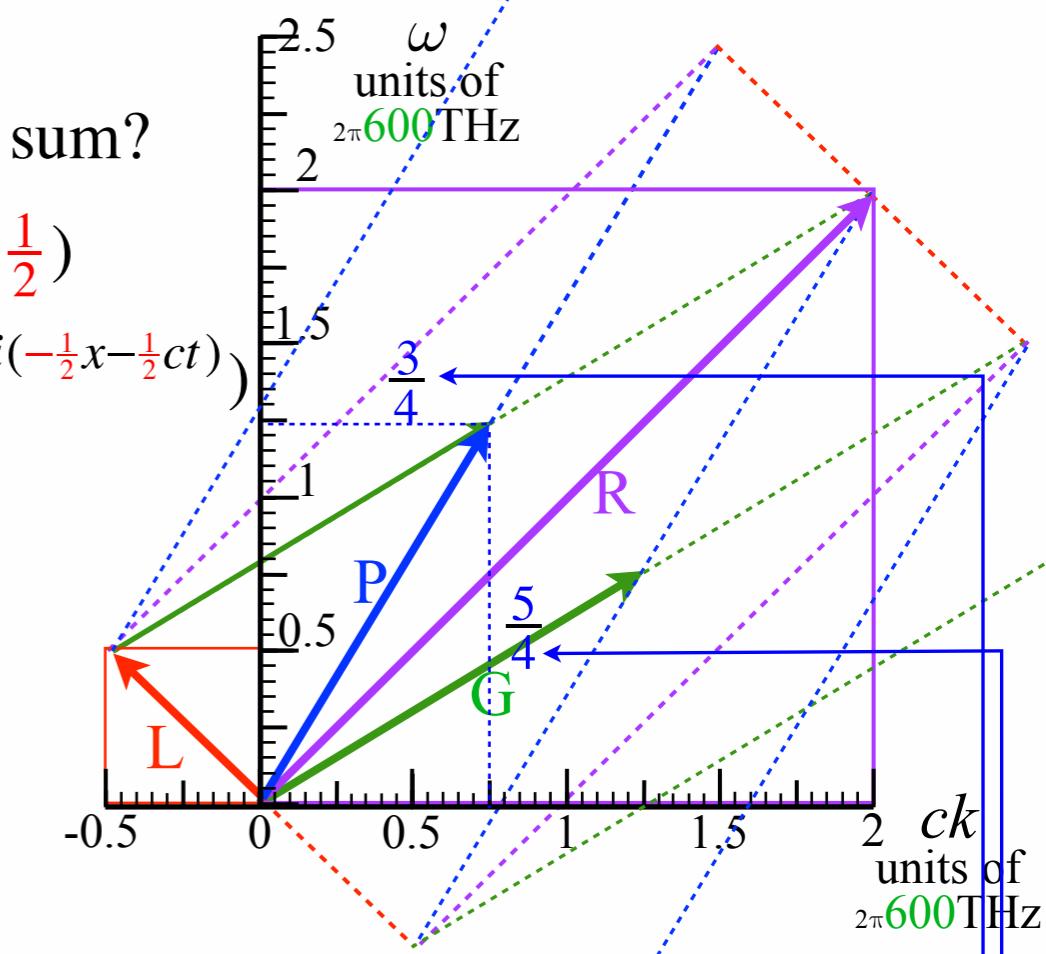
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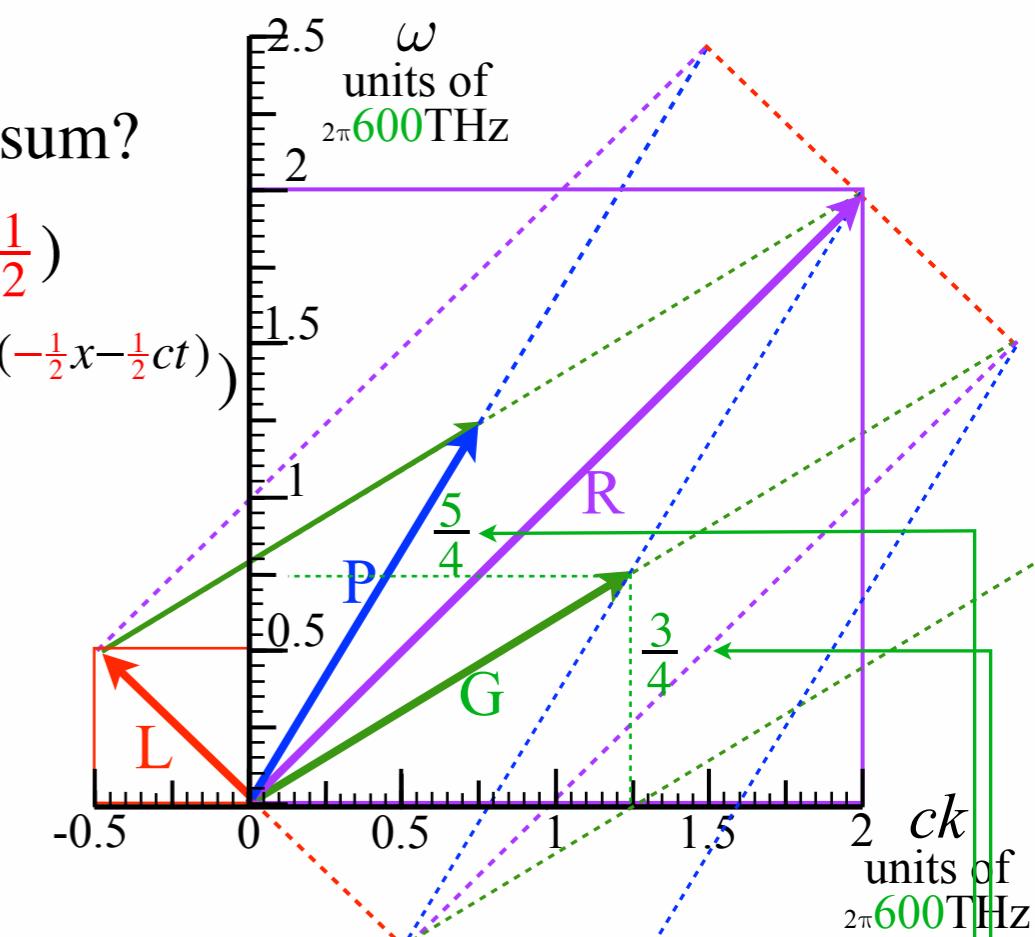
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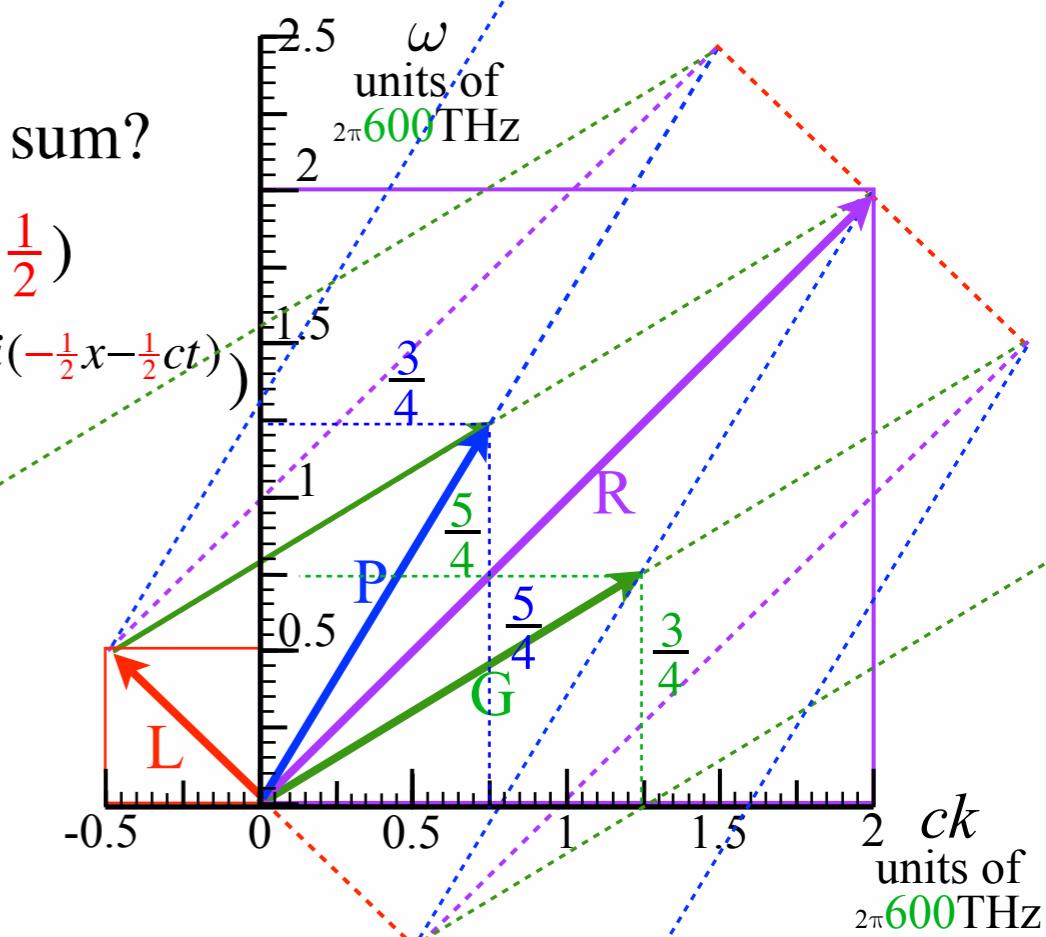
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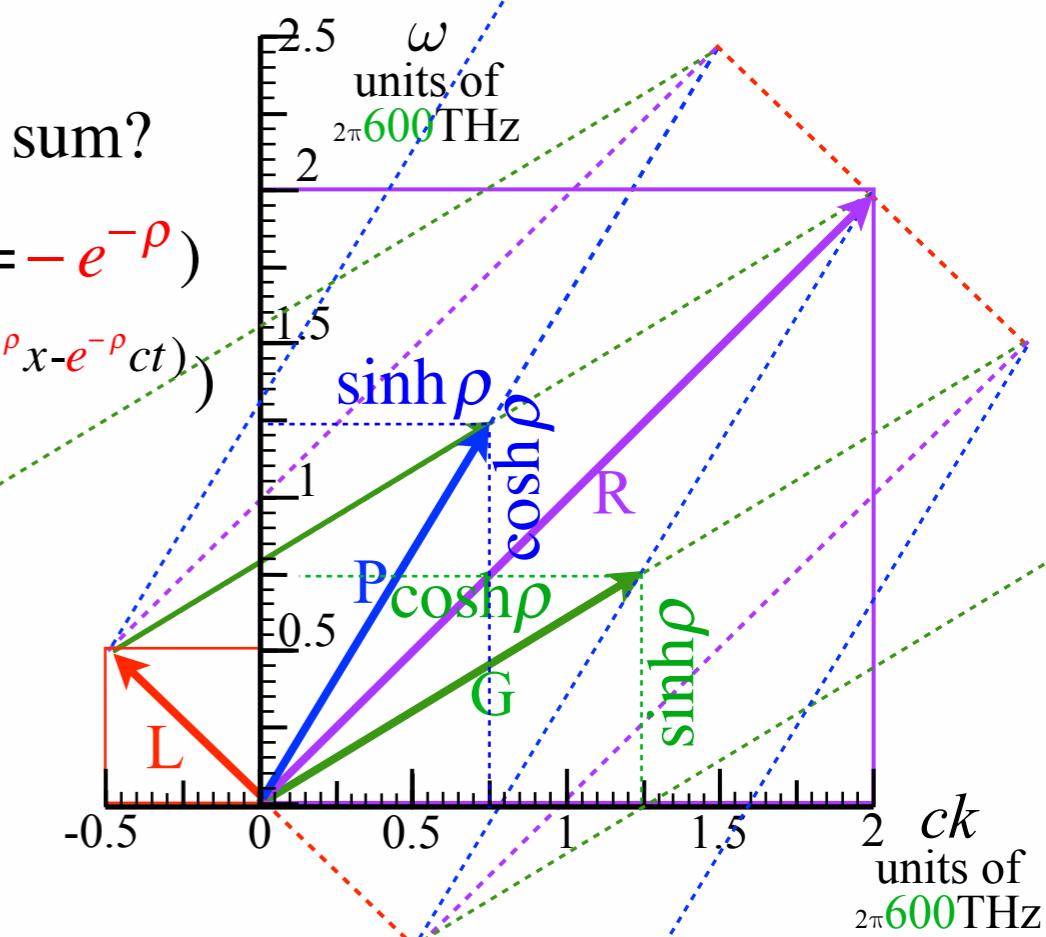
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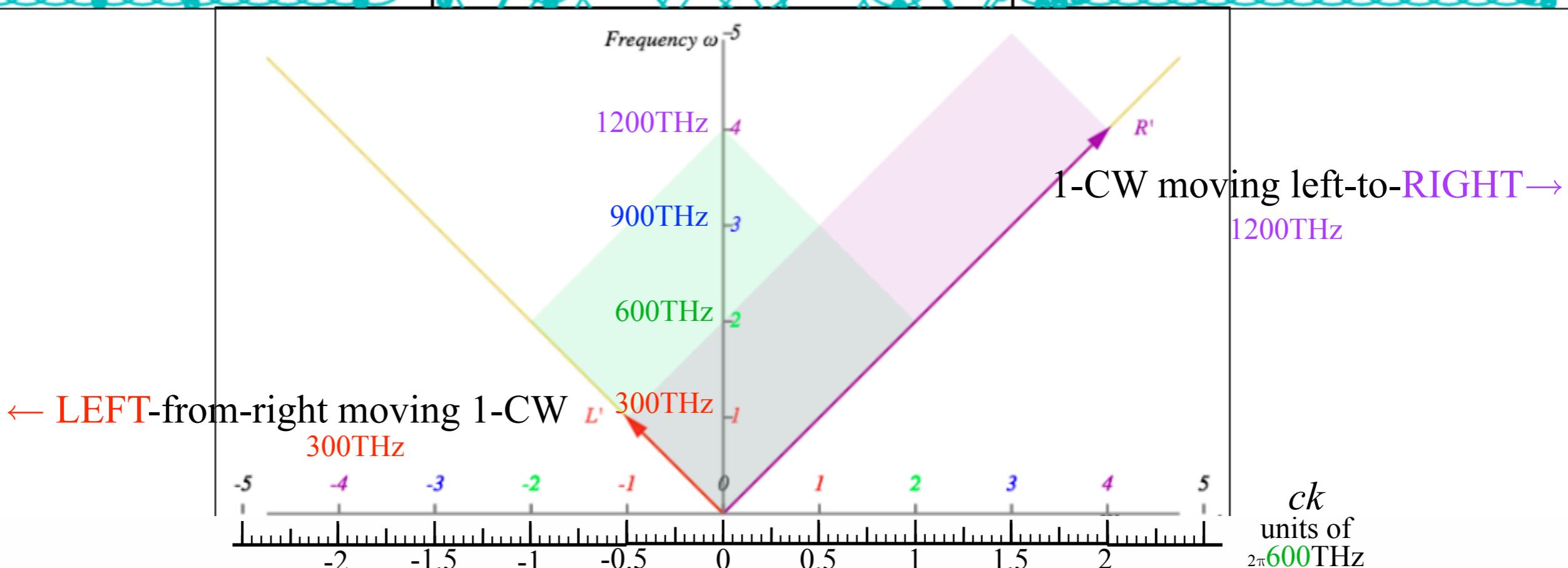
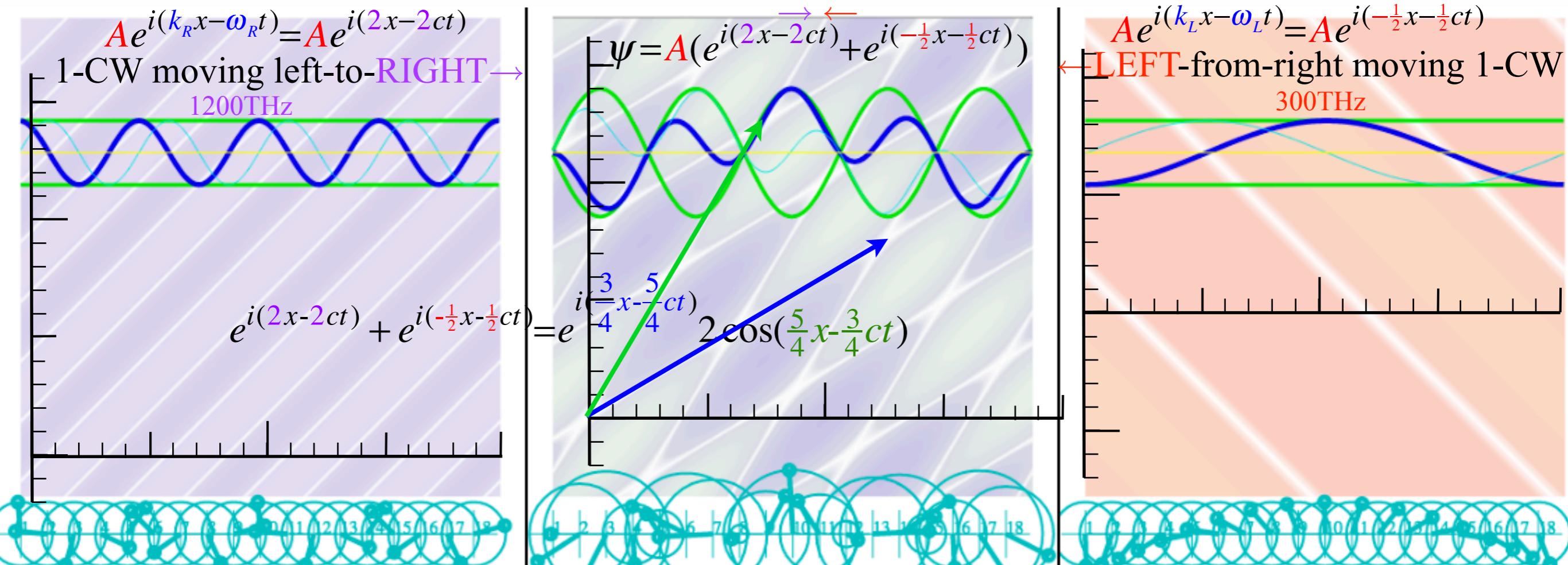
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2 colliding waves (2-CW)

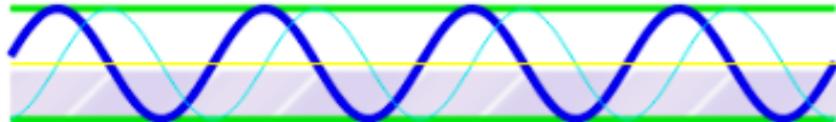


2 colliding waves (2-CW)

$$Ae^{i(k_R x - \omega_R t)} = Ae^{i(2x - 2ct)}$$

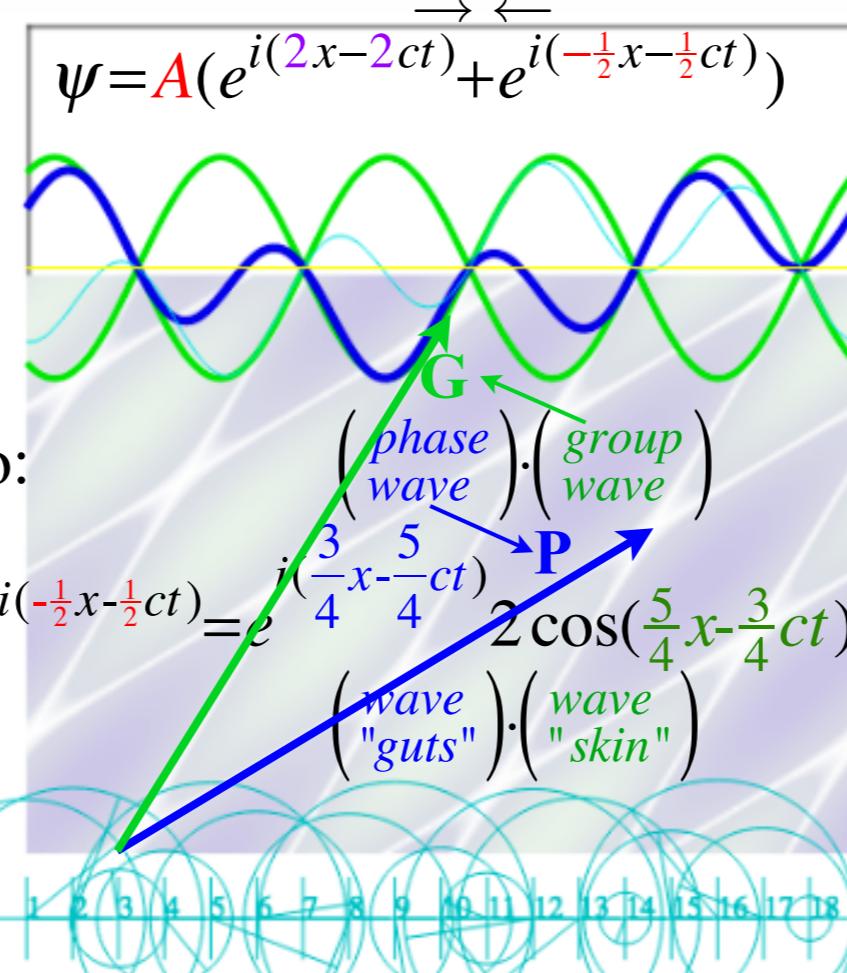
1-CW moving left-to-**RIGHT**→

1200THz



Factors into:

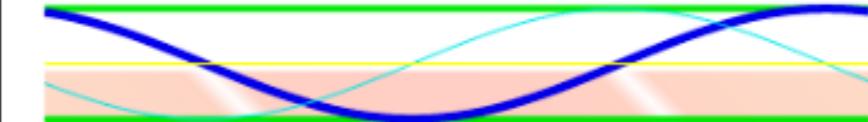
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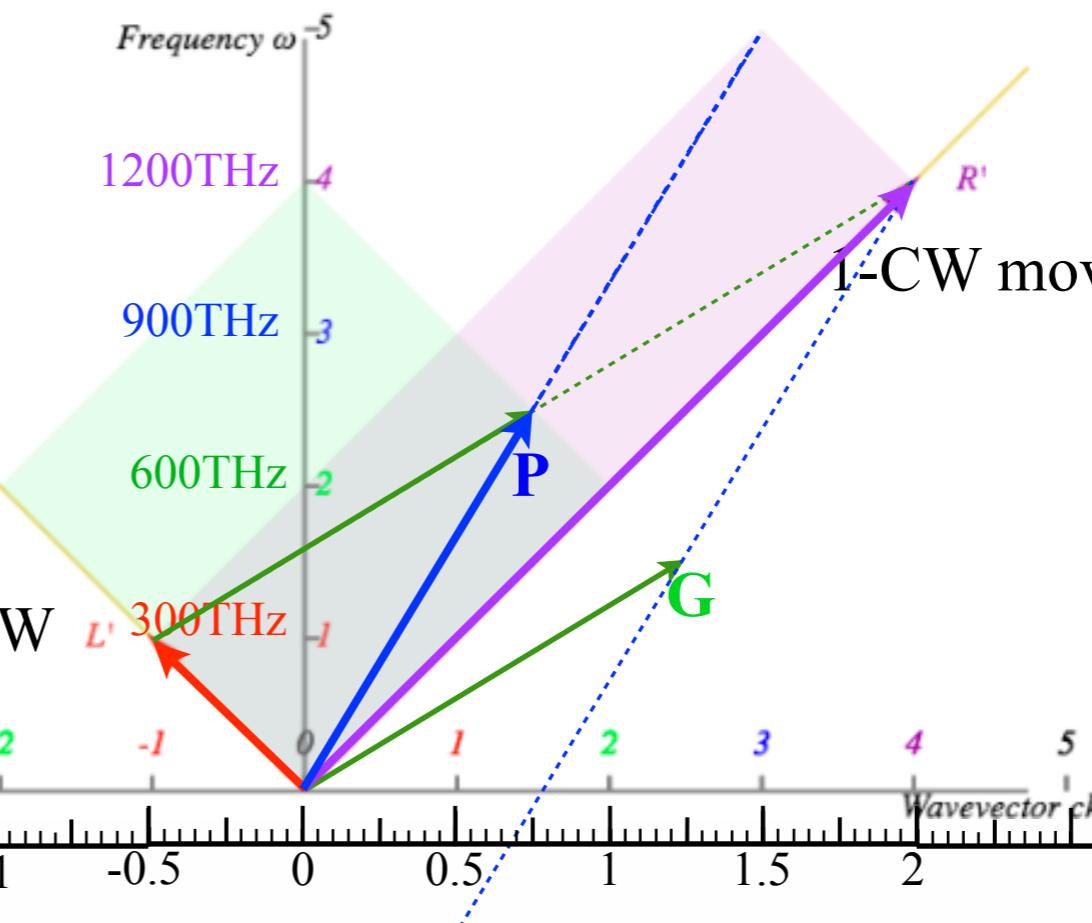
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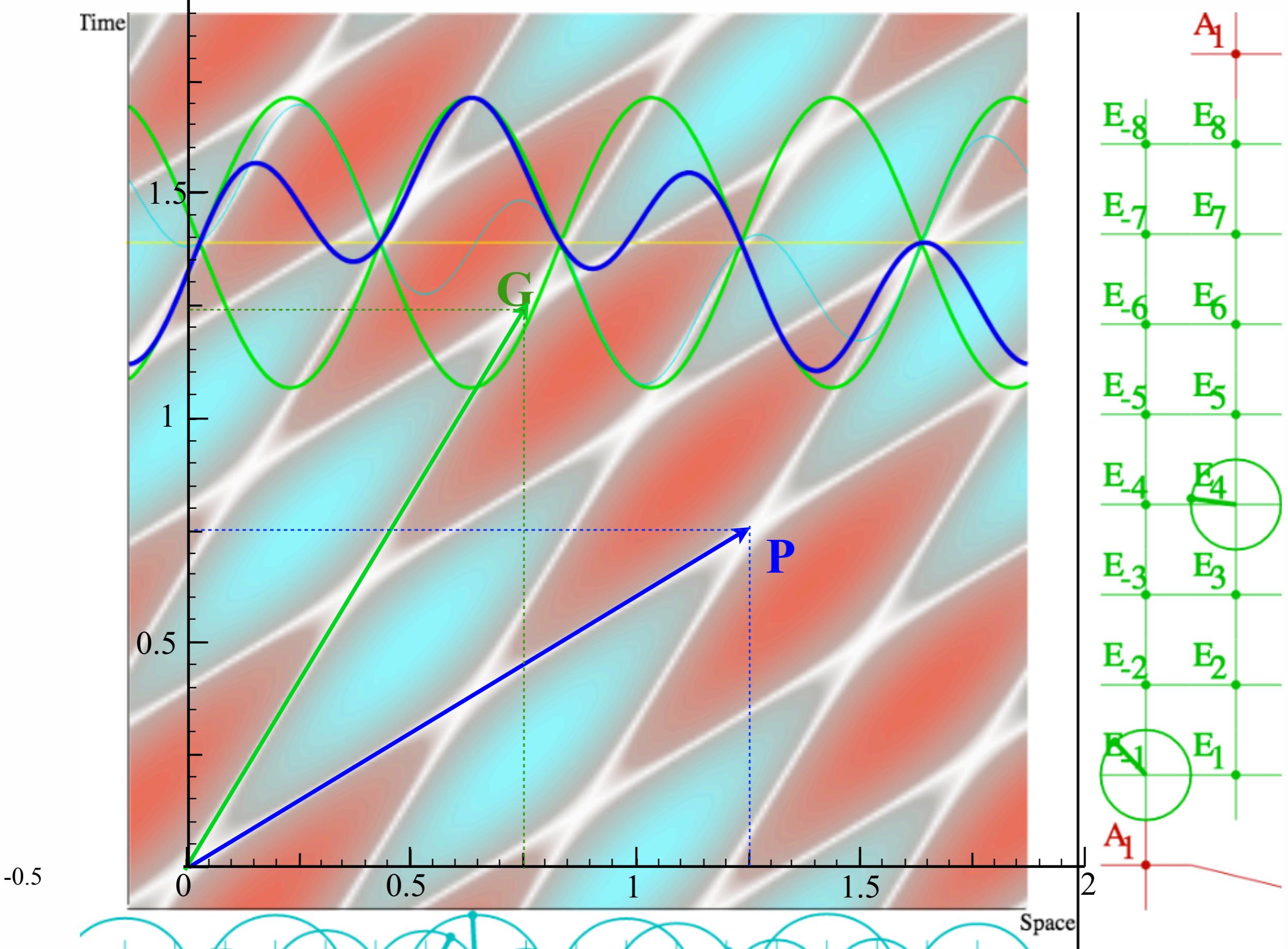
← LEFT-from-right moving 1-CW

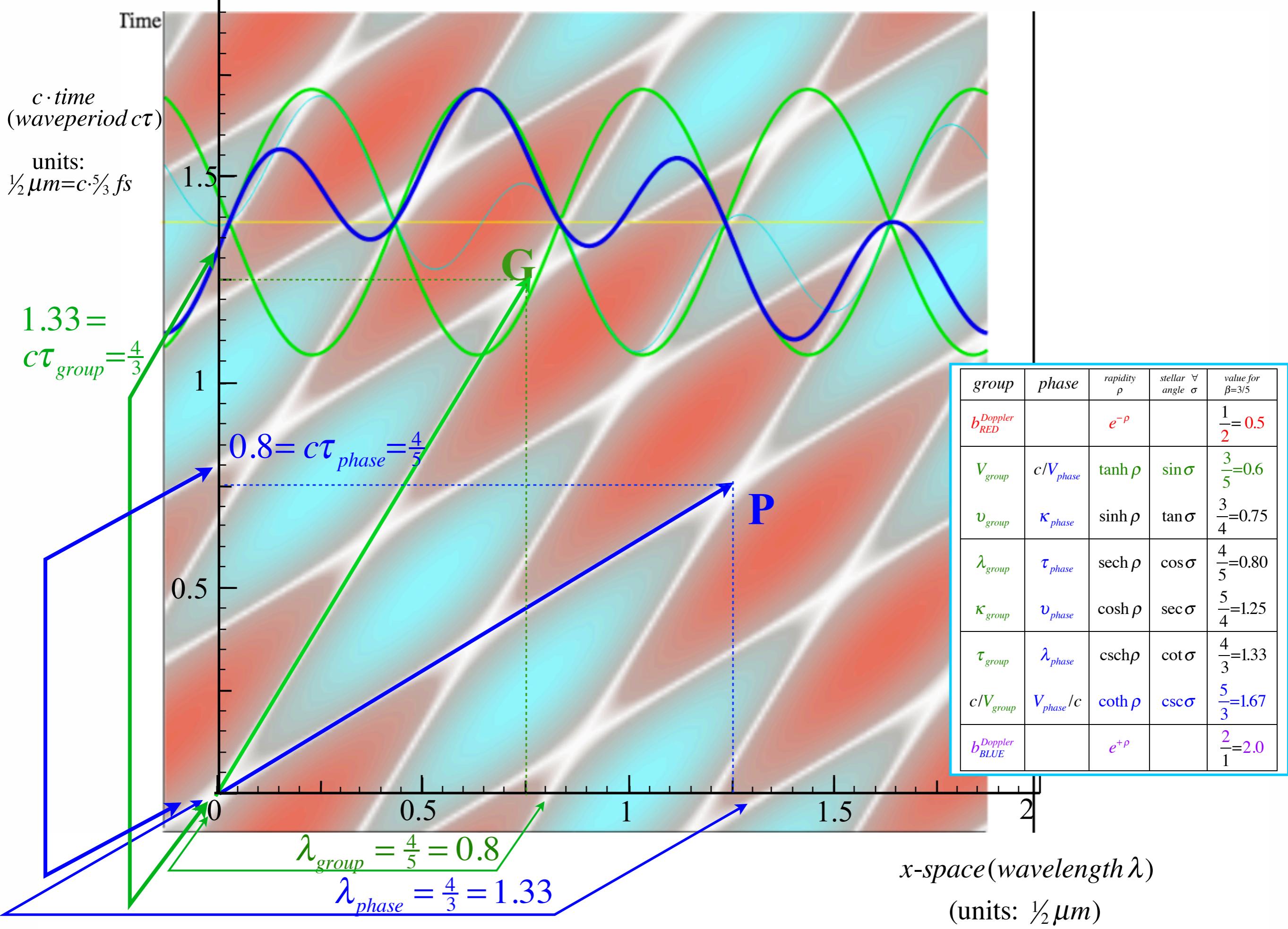
300THz

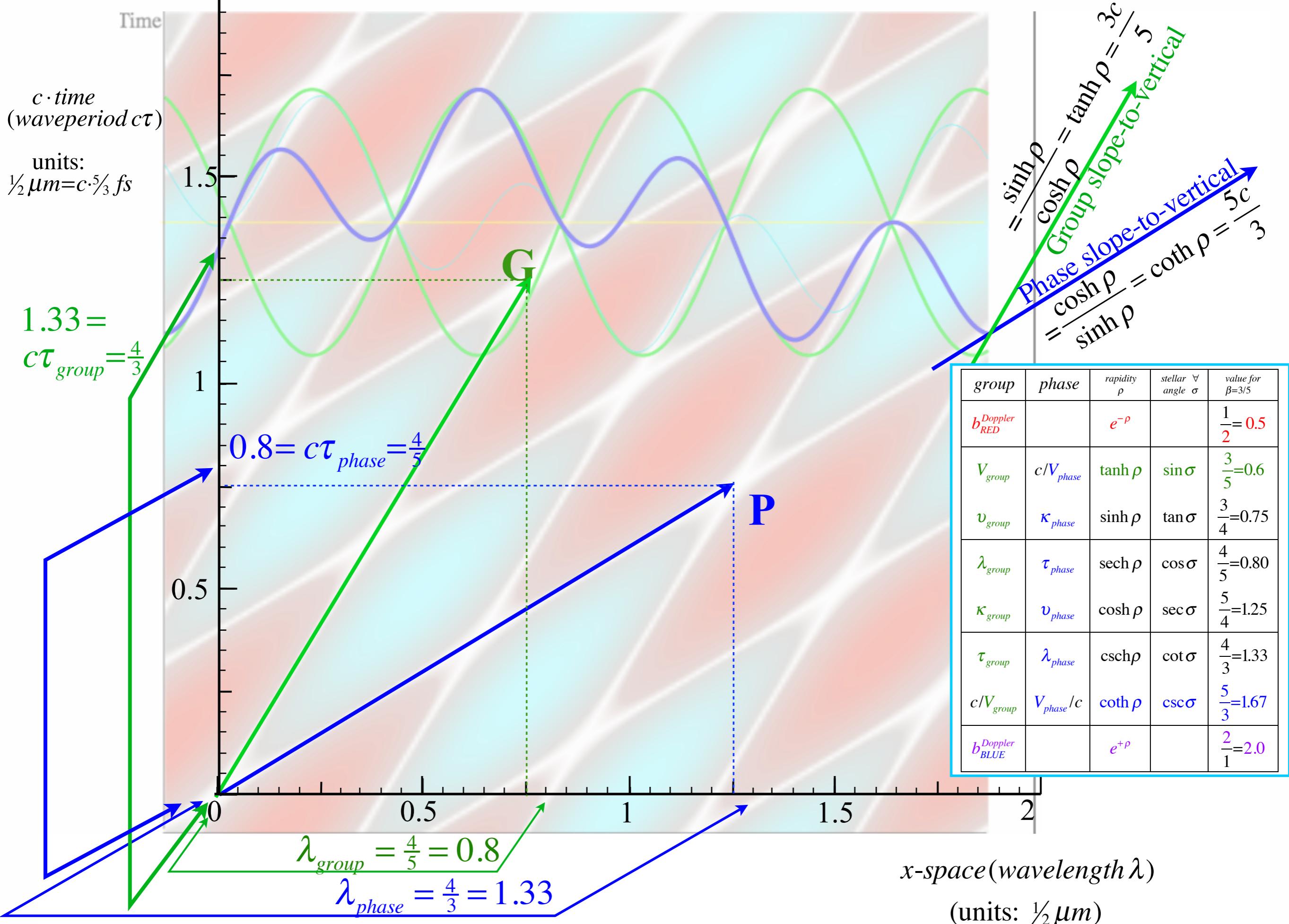


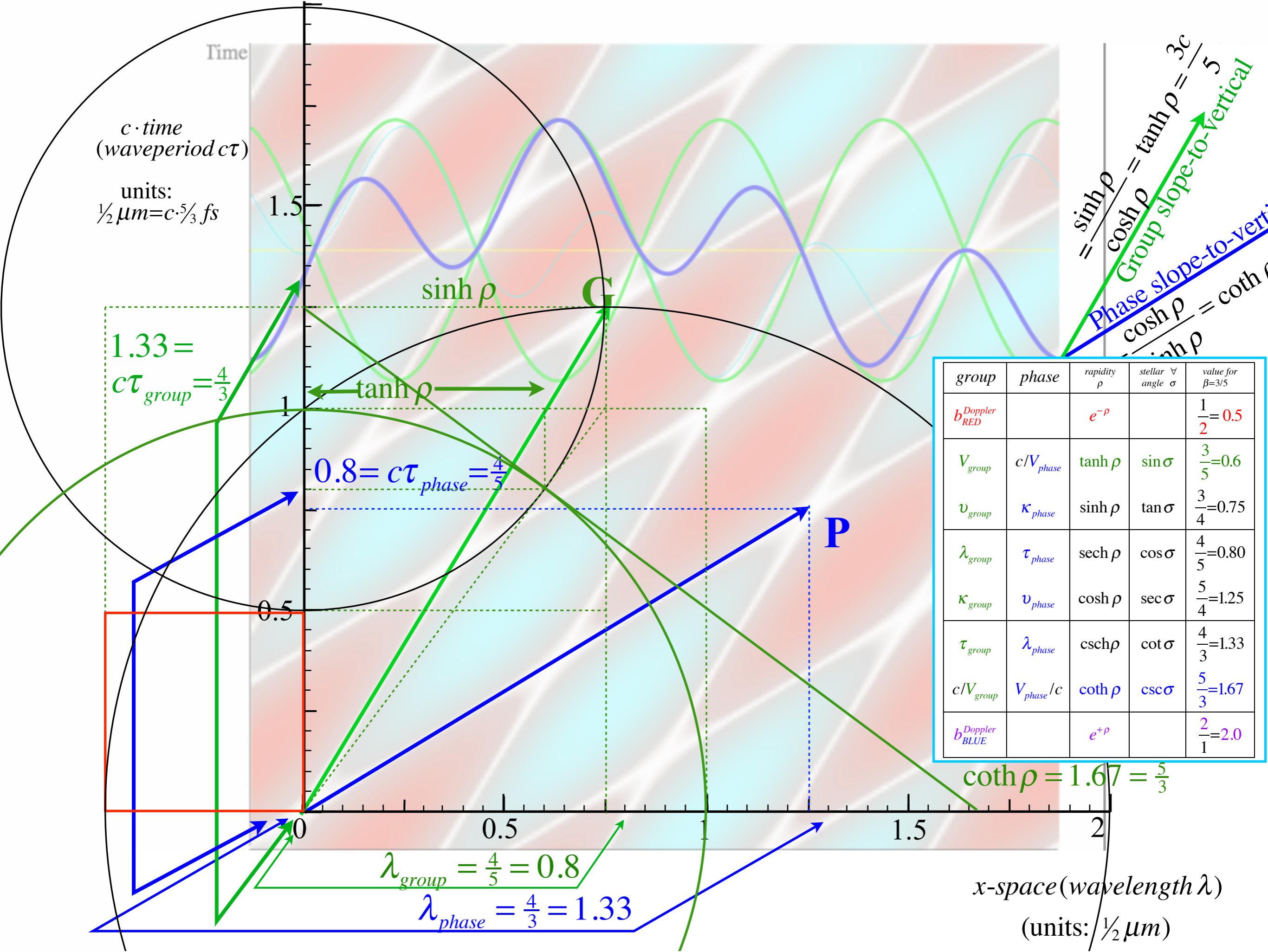
← LEFT-from-right moving 1-CW
300THz











This map has circle sector arc-area $\sigma = 0.6435$
set to angle $\angle\sigma = 36.87^\circ = 0.6435 \text{radian}$

$$\begin{array}{lll} \sin(\sigma) = 0.6000 & = \tanh(\rho) & = 3/5 \\ \tan(\sigma) = 0.7500 & = \sinh(\rho) & = 3/4 \\ \sec(\sigma) = 1.2500 & = \cosh(\rho) & = 5/4 \\ \cos(\sigma) = 0.8000 & = \operatorname{sech}(\rho) & = 4/5 \\ \cot(\sigma) = 1.3333 & = \operatorname{csch}(\rho) & = 4/3 \\ \csc(\sigma) = 1.6667 & = \operatorname{coth}(\rho) & = 5/3 \end{array}$$

$$\begin{aligned} \cosh(\rho) + \sinh(\rho) &= \frac{5}{4} + \frac{3}{4} = 2.0 = e^{+\rho} \\ \cosh(\rho) - \sinh(\rho) &= \frac{5}{4} - \frac{3}{4} = 1/2 = e^{-\rho} \end{aligned}$$

$$\begin{aligned} \cosh(\rho) &= \frac{e^{+\rho} + e^{-\rho}}{2} & \text{Half-Sum-} \\ \sinh(\rho) &= \frac{e^{+\rho} - e^{-\rho}}{2} & \text{Half-Difference} \\ && \text{Trig-Formulae for} \\ && \text{exponentials } e^{\pm\rho} \end{aligned}$$

$$x^2 - y^2 = B^2$$

$$\operatorname{Bcosh}(\rho) - \operatorname{Bsinh}(\rho) = Be^{-\rho}$$

Thales Circle
(links $e^{-\rho}$ to $e^{+\rho}$)

Also it is set to hyperbola sector arc-area $\rho = 0.6931$
angle $\angle\rho = v = 30.96^\circ$

$$\operatorname{Bcosh}(\rho) + \operatorname{Bsinh}(\rho) = Be^{+\rho}$$

$$\operatorname{tangent slope} = \tanh(\rho)$$

$$\operatorname{tangent slope} = \coth(\rho)$$

$$\operatorname{tangent slope} = \operatorname{coth}(\rho)$$

$$\operatorname{Bcsc}(\rho)$$

$$\operatorname{Bcsch}(\rho)$$

$$\operatorname{Bcoth}(\rho)$$

$$\operatorname{Bsech}(\rho)$$

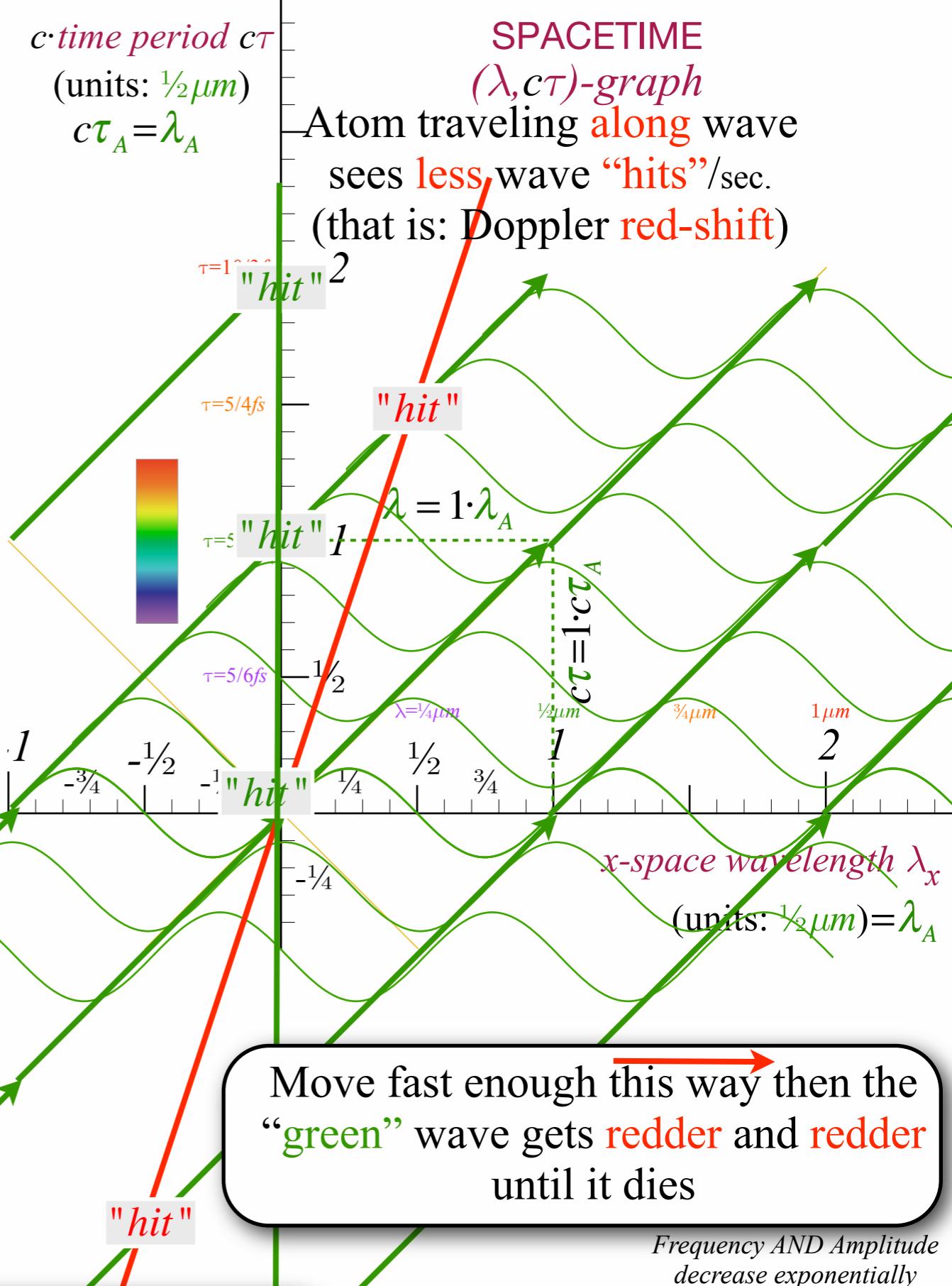
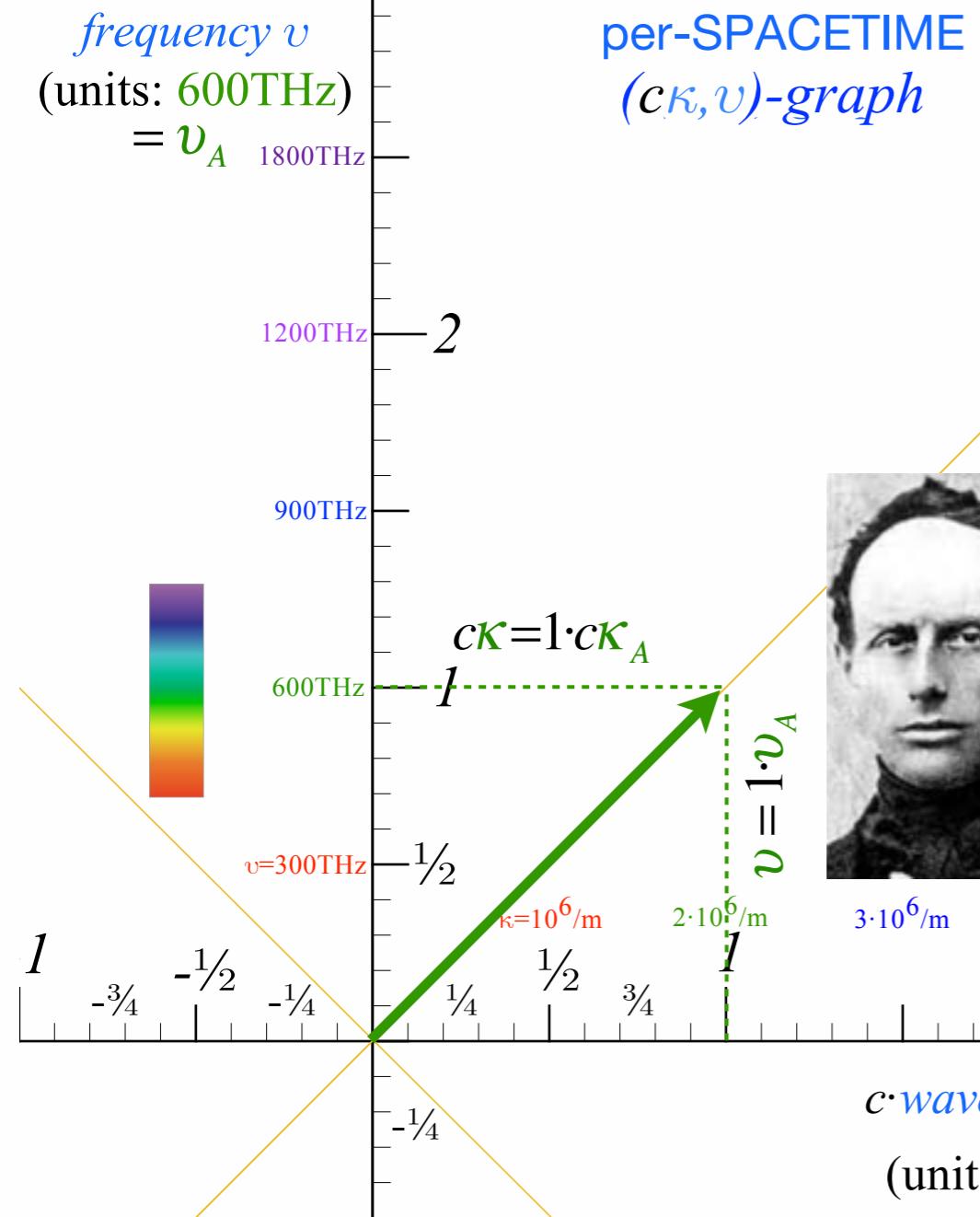
$$\operatorname{Btanh}(\rho)$$

$$\operatorname{Bsinh}(\rho)$$

$$\operatorname{Bcosh}(\rho)$$

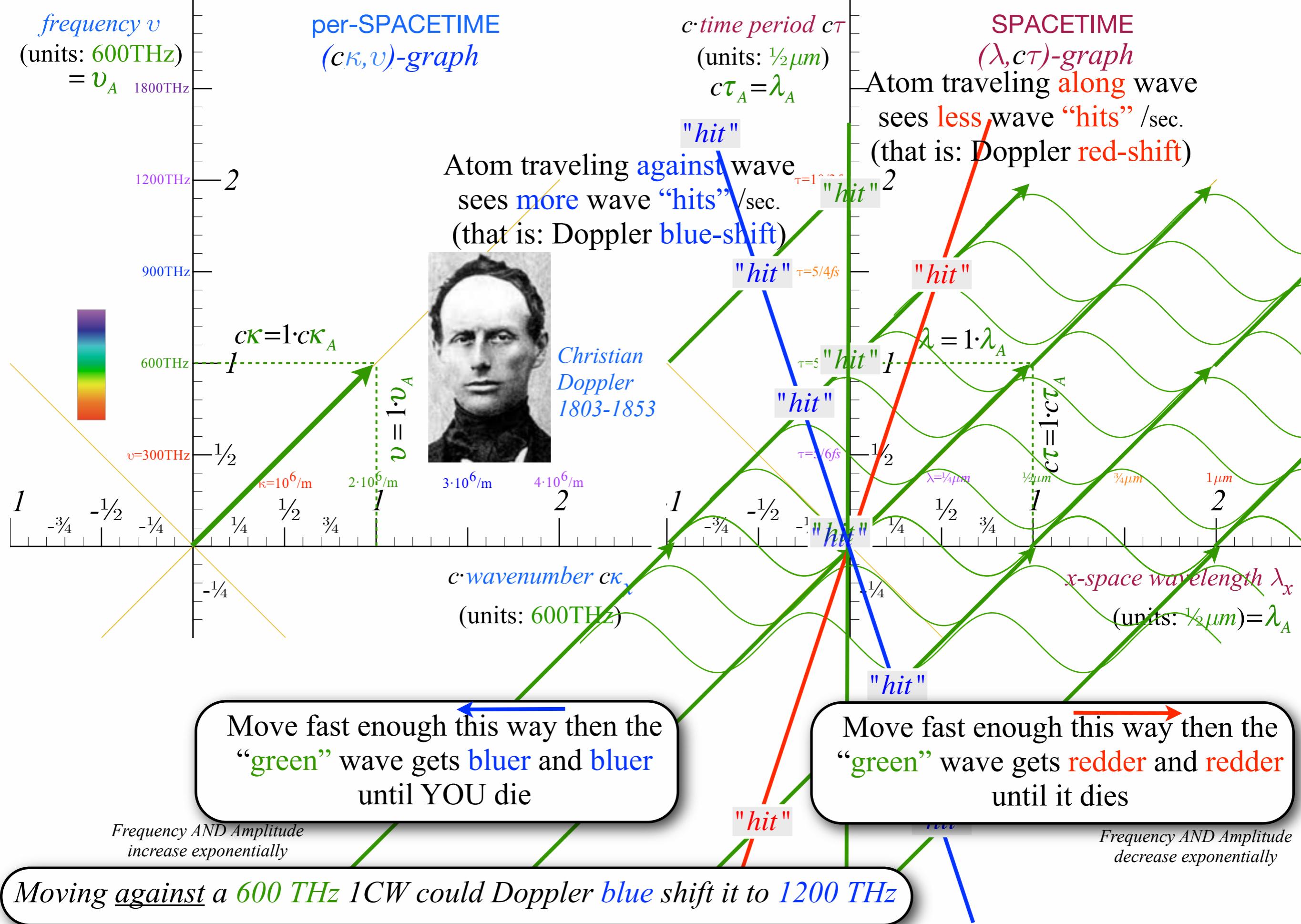
<i>group</i>	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$b_{BLUE}^{Doppler}$
<i>phase</i>	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$\frac{1}{b_{RED}^{Doppler}}$
<i>rapidity</i> ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
<i>stellar</i> \forall <i>angle</i> σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
<i>value for</i> $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

The “Keyboard of the gods” or per-space-per-time graphs versus space-time graphs



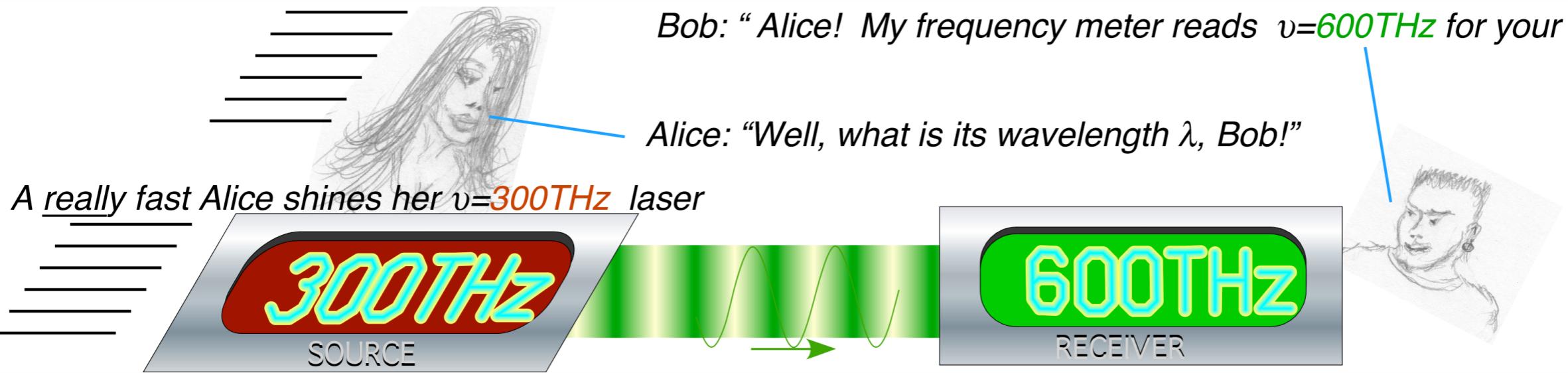
Moving along a 600 THz 1CW could Doppler **red** shift it to 300 THz

The “Keyboard of the gods” or per-space-per-time graphs versus space-time graphs



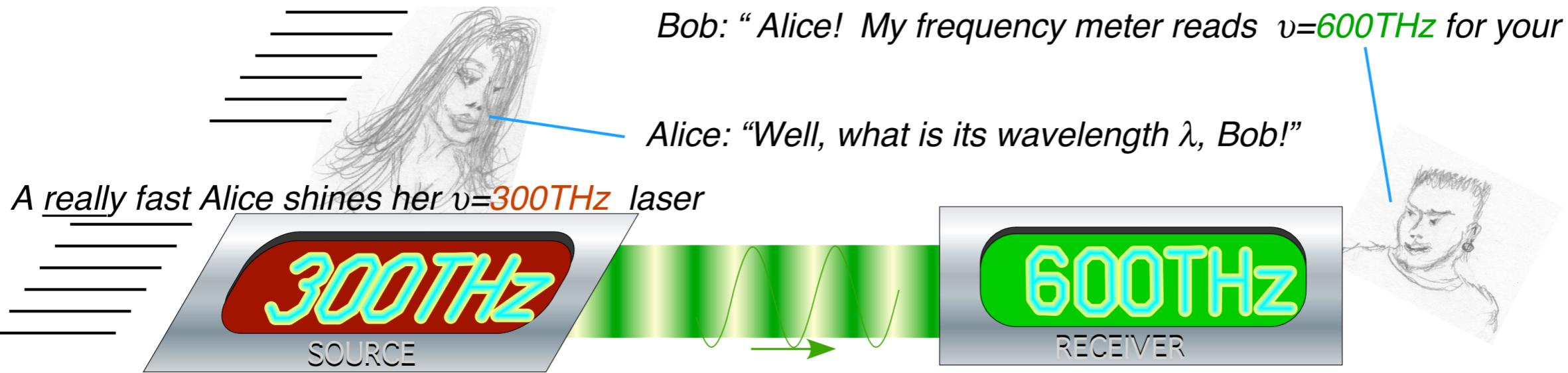
Clarify Evenson's CW Axiom (*All colors go c*) by Doppler effects

Alice tries to fool Bob that she's shining a 600THz laser. (Bob's unaware she's moving *really* fast...)



Clarify Evenson's CW Axiom (*All colors go c*) by Doppler effects

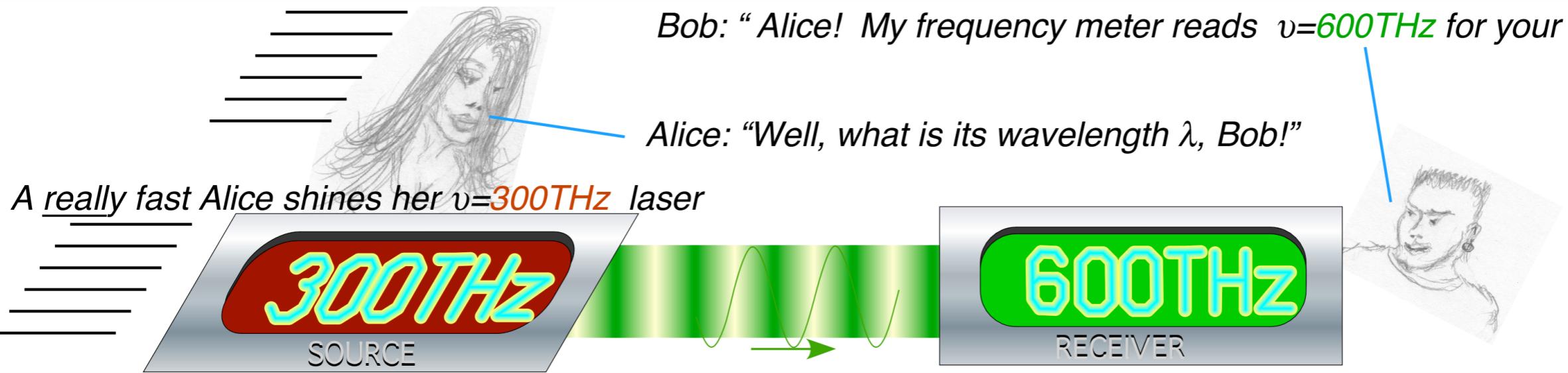
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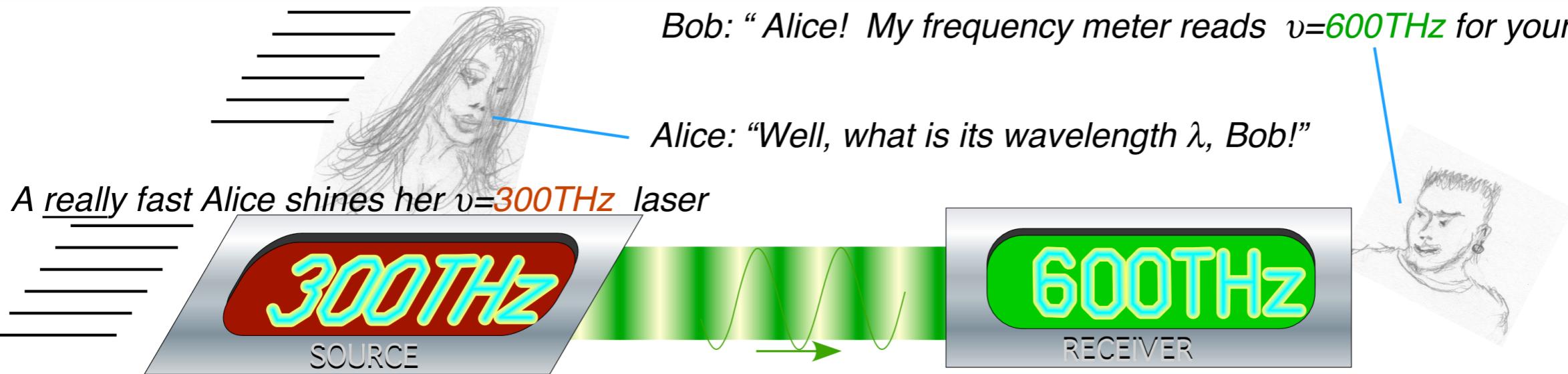


Q1: Can Bob tell it's a “*phony*” 600THz by measuring his received wavelength?

Q2: If so, what “*phony*” λ does Bob see?

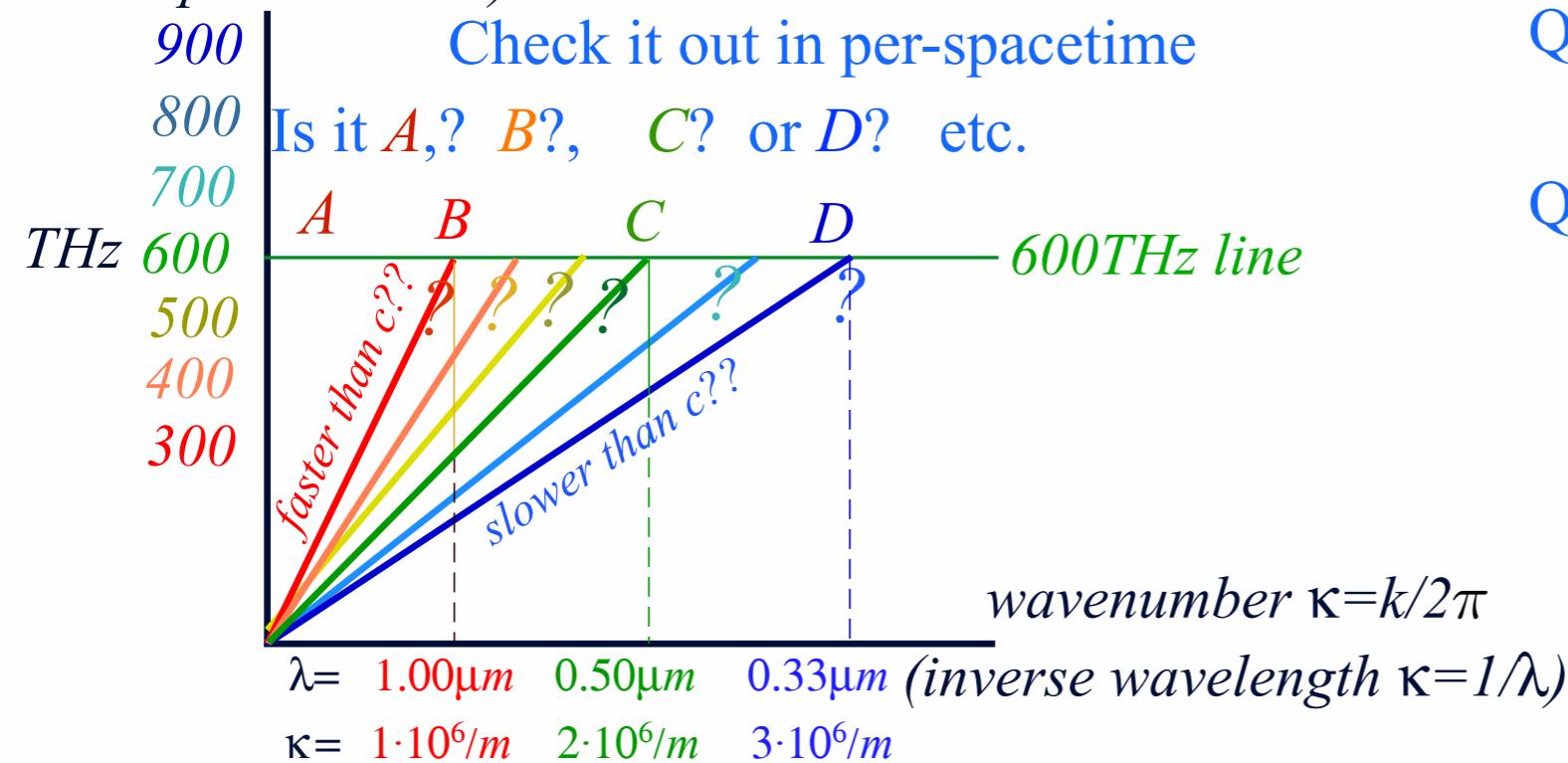
Clarify Evenson's CW Axiom (All colors go c) by Doppler effects

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frequency $\nu=\omega/2\pi$

(Inverse period $\nu=1/\tau$)

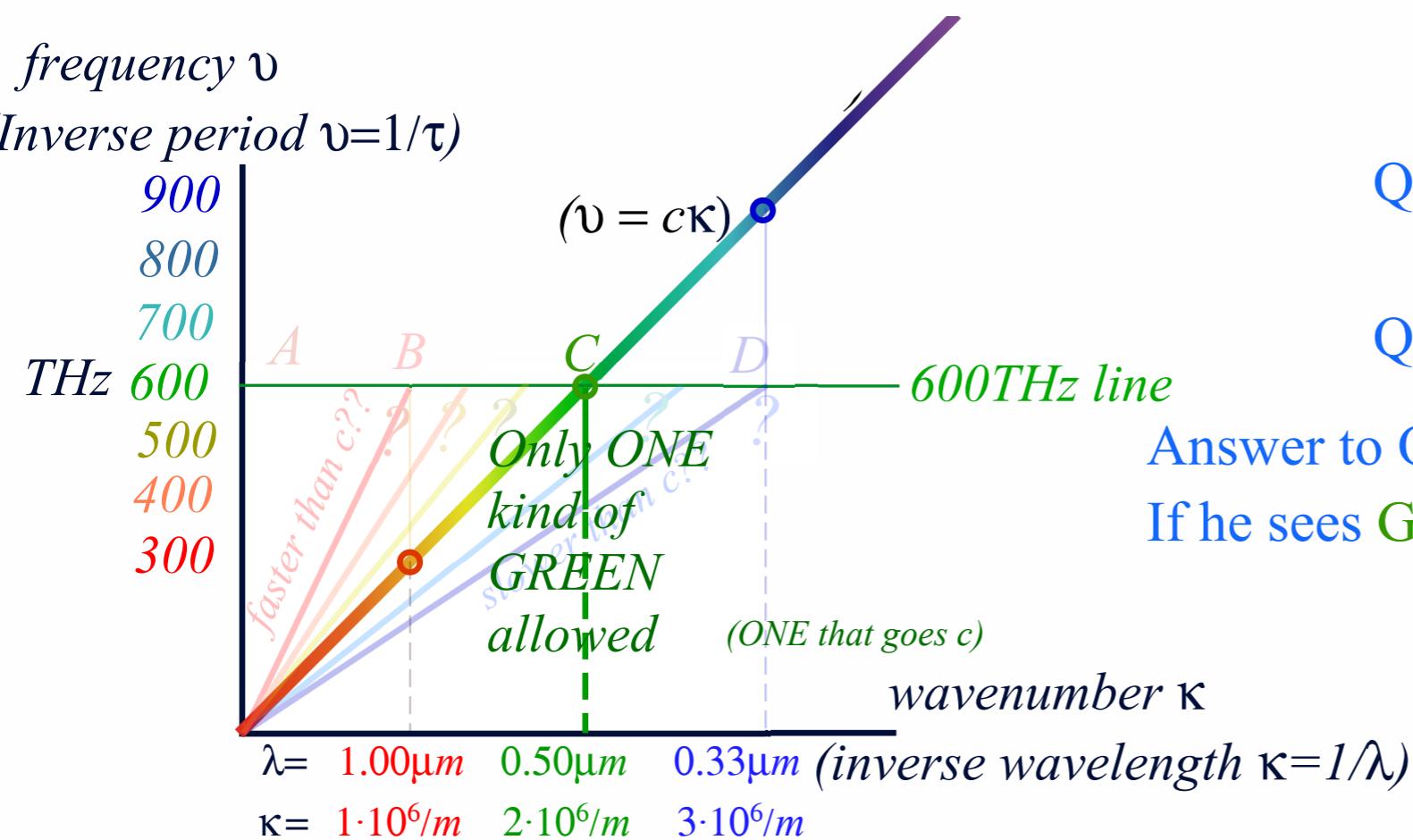
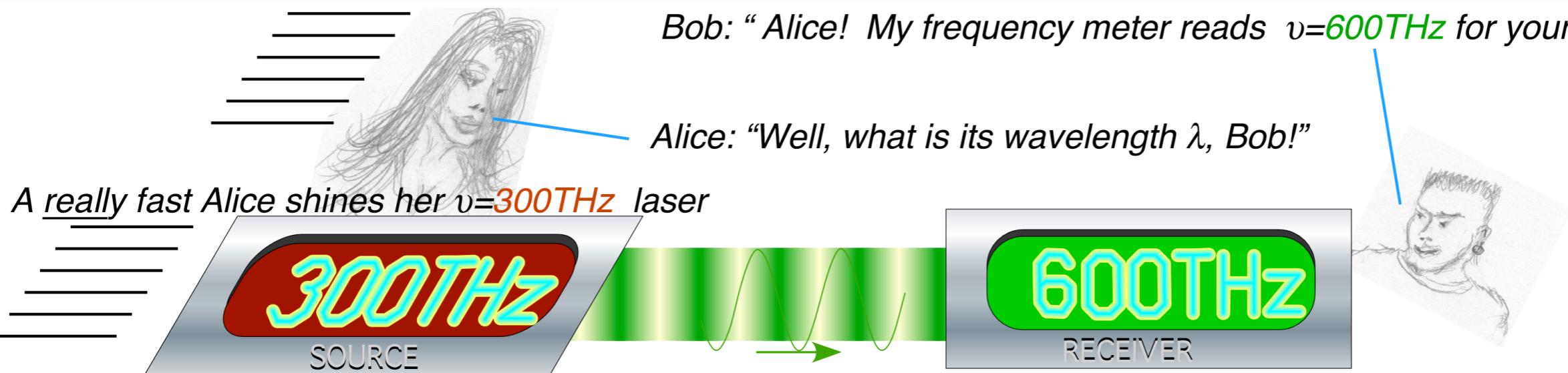


Q1: Can Bob tell it's a "phony" 600THz by measuring his received wavelength?

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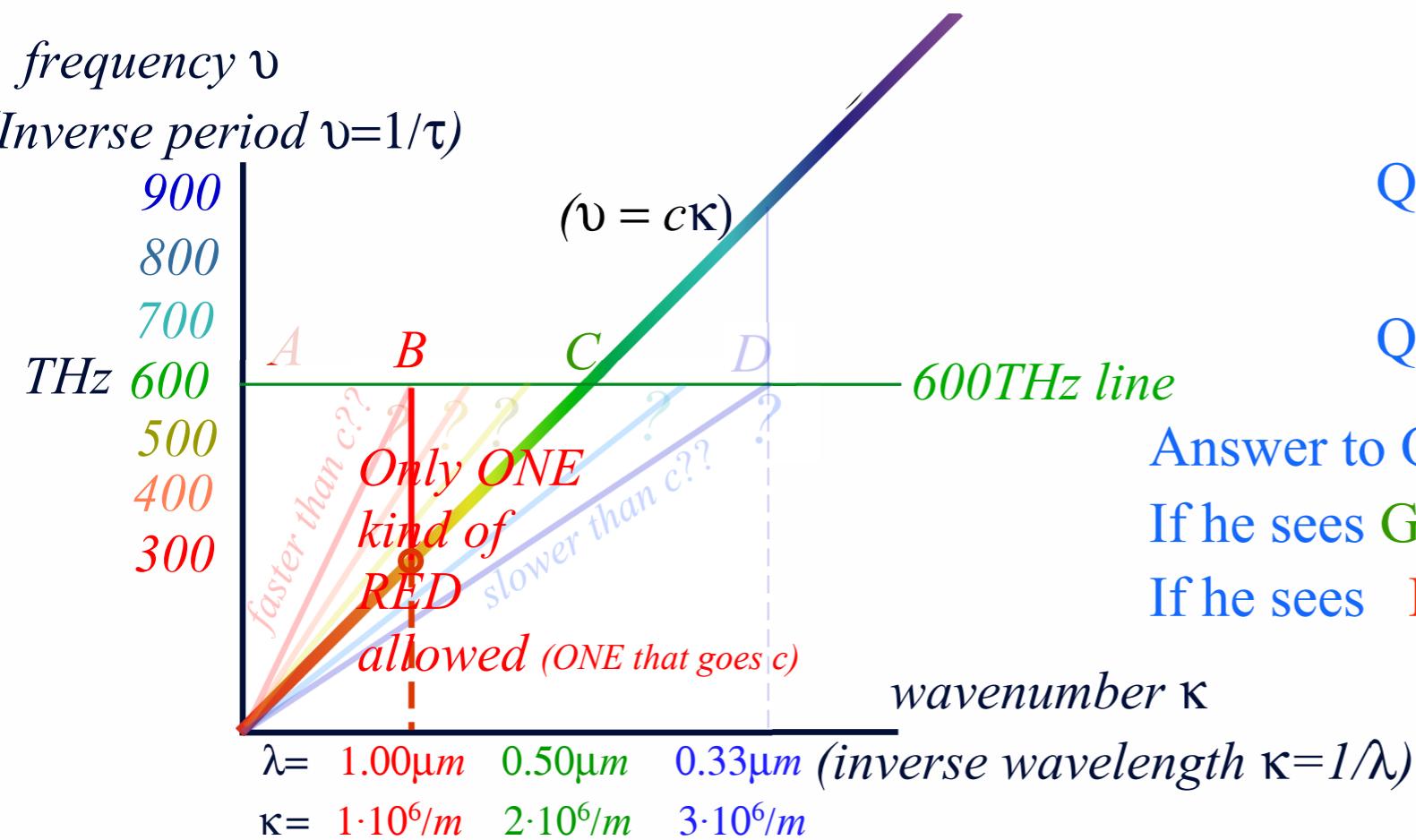
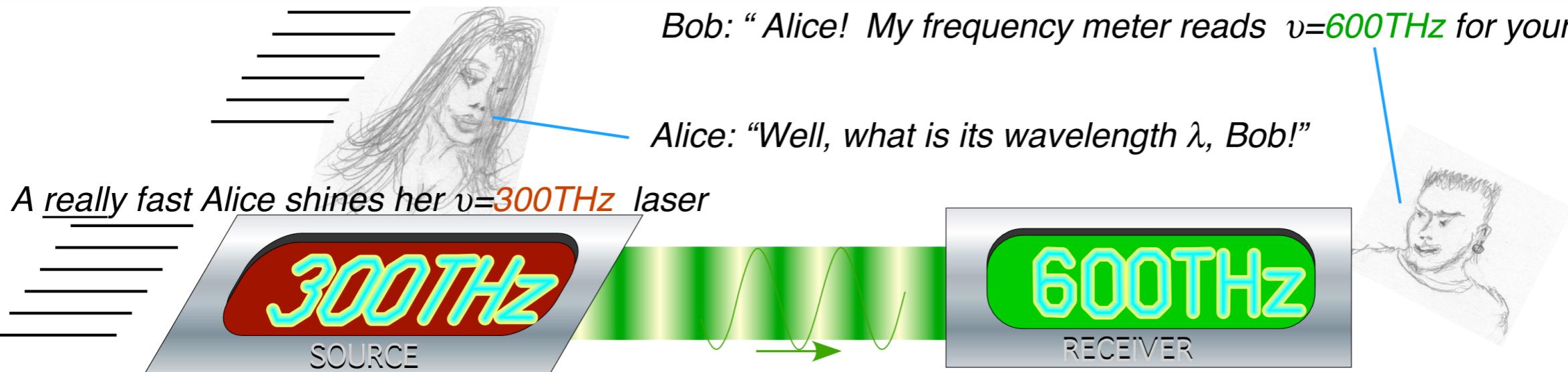
Q1: Can Bob tell it's a "phony" 600THz by measuring his received wavelength?

Q2: If so, what "phony" λ does Bob see?

Answer to Q2 is C, the one with slope $\nu/\kappa=\nu\cdot\lambda=c$.
If he sees Green 600THz then he measures $\lambda=0.5\mu\text{m}$.

Clarify Evenson's CW Axiom (All colors go c) by Doppler effects

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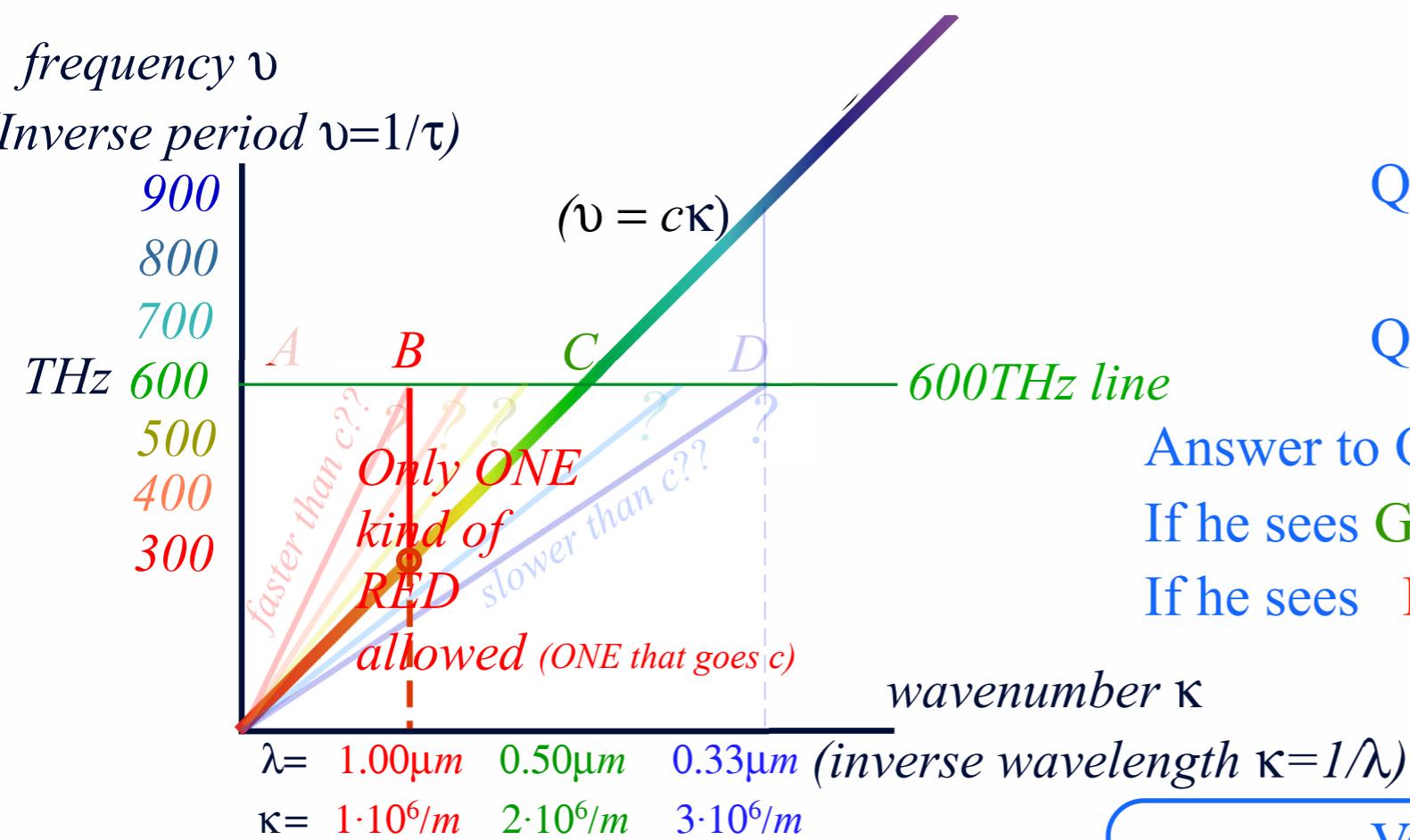
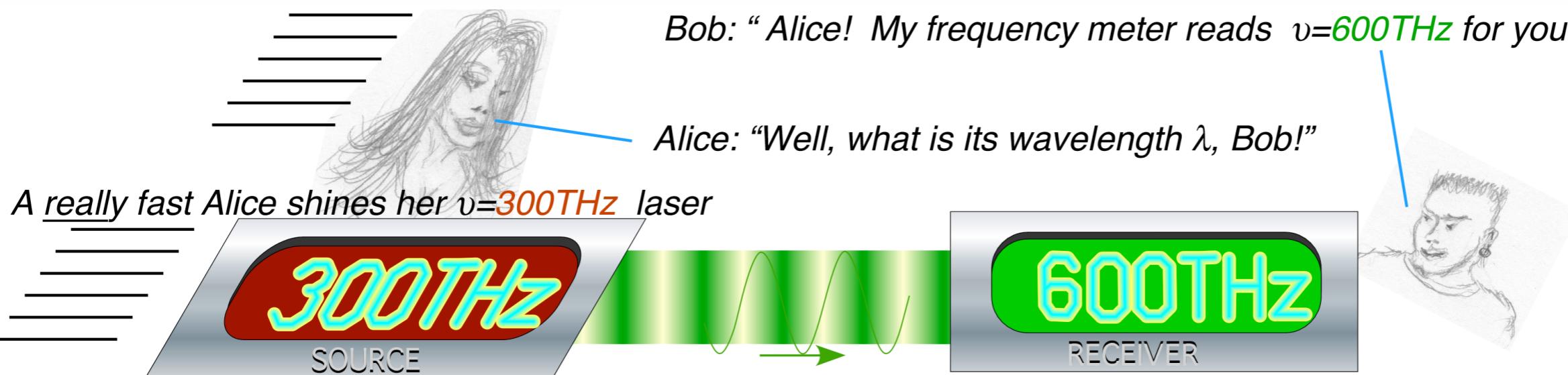
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If he sees Green 600THz then he measures $\lambda=0.5\mu\text{m}$.
If he sees Red 300THz then he measures $\lambda=1.0\mu\text{m}$.

Clarify Evenson's CW Axiom (All colors go c) by Doppler effects

Alice tries to fool Bob that she's shining a 600THz laser. (Bob's unaware she's moving *really* fast...)



Q1: Can Bob tell it's a "phony" 600THz by measuring his received wavelength?

Q2: If so, what "phony" λ does Bob see?

Answer to Q2 is C, the one with slope $v/\kappa=v\cdot\lambda=c$.

If he sees Green 600THz then he measures $\lambda=0.5\mu\text{m}$.

If he sees Red 300THz then he measures $\lambda=1.0\mu\text{m}$.

Answer to Q1 is NO!

CW Light carries **no** birth-certificate!

Vacuum only makes one λ for each v .*

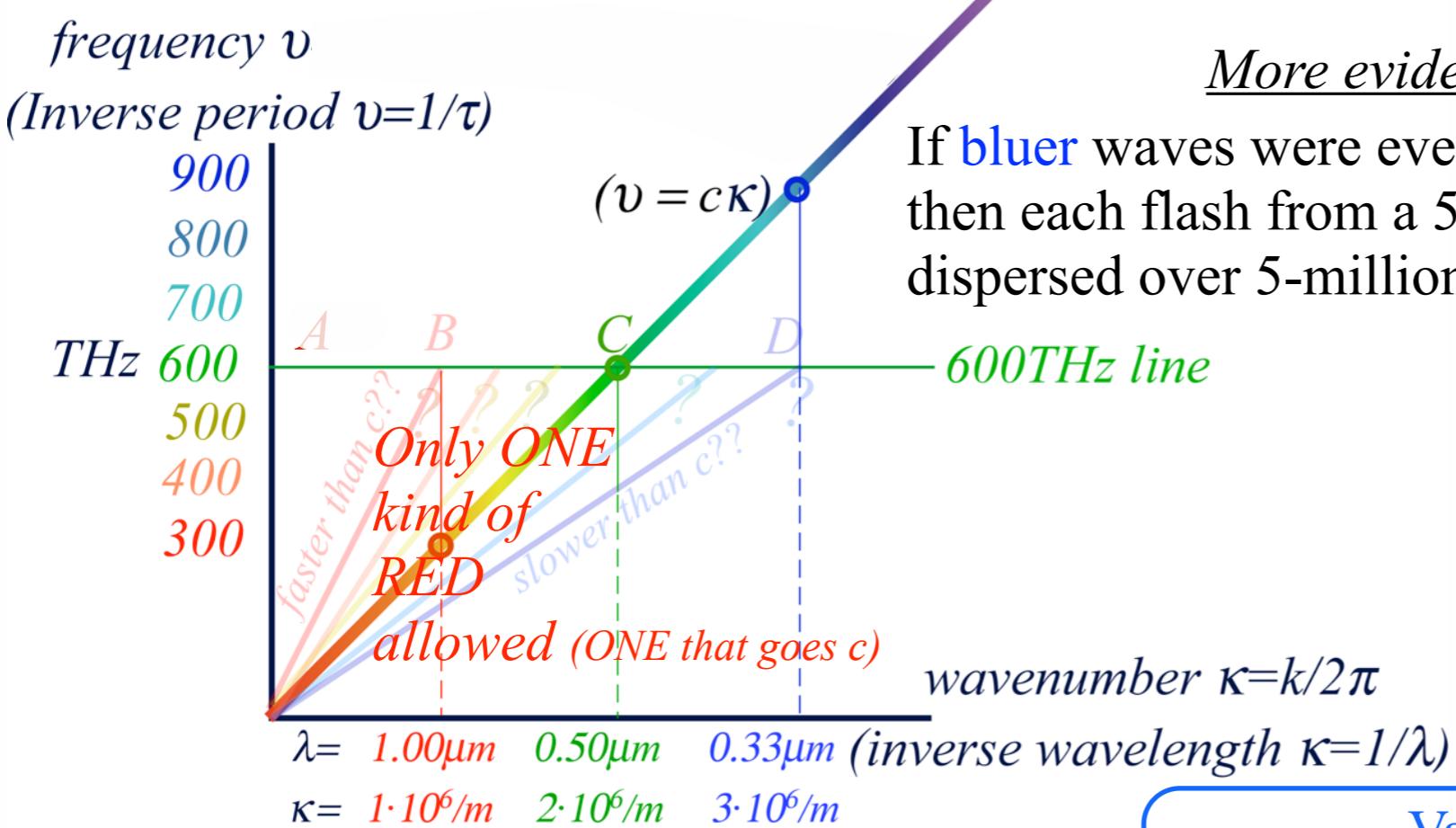
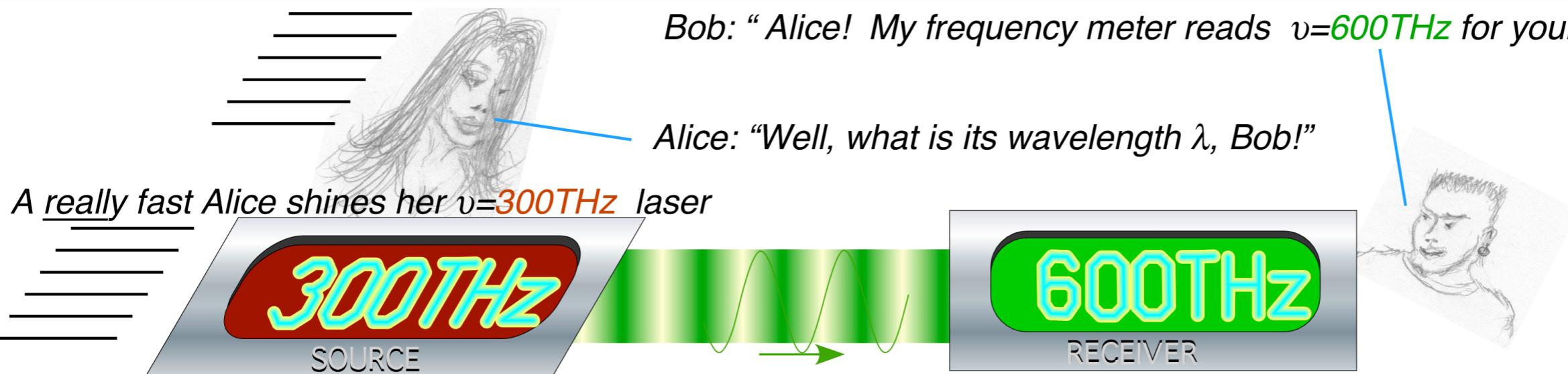
"All colors go $c = \lambda v = v/\kappa$ "

Then *Evenson's axiom* holds:

*for each beam and polarization orientation

Clarify Evenson's CW Axiom (*All colors go c*) by Doppler effects

Alice tries to fool Bob that she's shining a 600THz laser. (Bob's unaware she's moving *really fast...*)



More evidence supporting Evenson's axiom

If bluer waves were even 0.1% faster (or slower) than redder ones then each flash from a 5-billion light-year distant galaxy shows up dispersed over 5-million years. (*Goodbye galactic astronomy!*)

Also could be labeled :

Linear-(non)-dispersion

axiom: $\nu = ck$

Vacuum only makes one λ for each ν .*

"All **colors** go $c = \lambda\nu = \nu/\kappa$ "

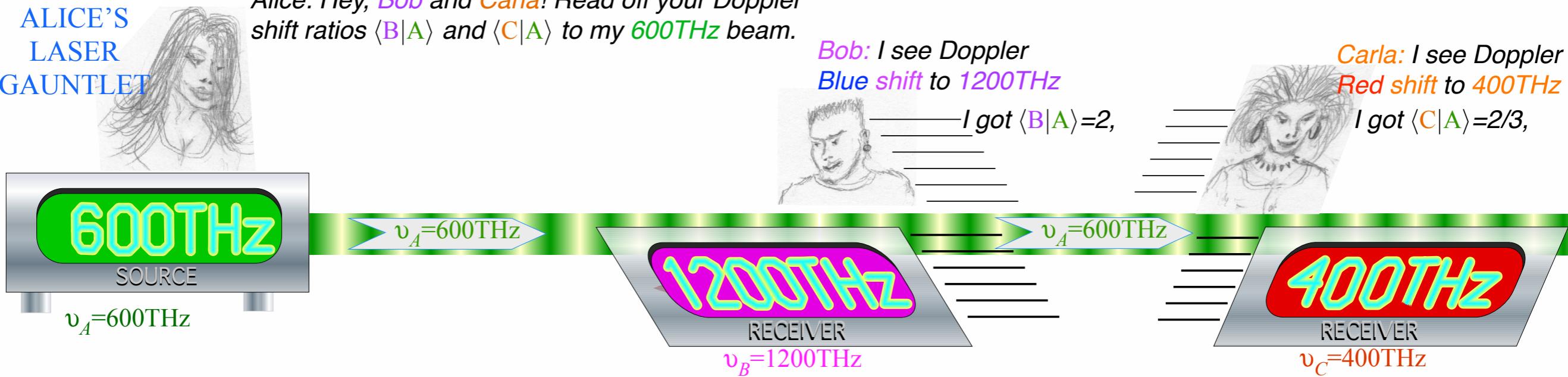
Then *Evenson's axiom* holds:

*for each beam and polarization orientation

Easy Doppler-shift and Rapidity calculation

ALICE'S
LASER
GAUNTLET

Alice: Hey, Bob and Carla! Read off your Doppler shift ratios $\langle B|A \rangle$ and $\langle C|A \rangle$ to my 600THz beam.



Doppler ratio:

$$\langle R|S \rangle = \frac{v_{RECEIVER}}{v_{SOURCE}}$$

Bob-Alice Doppler ratio:

$$\langle B|A \rangle = \frac{v_B}{v_A} = \frac{1200}{600} = \frac{2}{1}$$

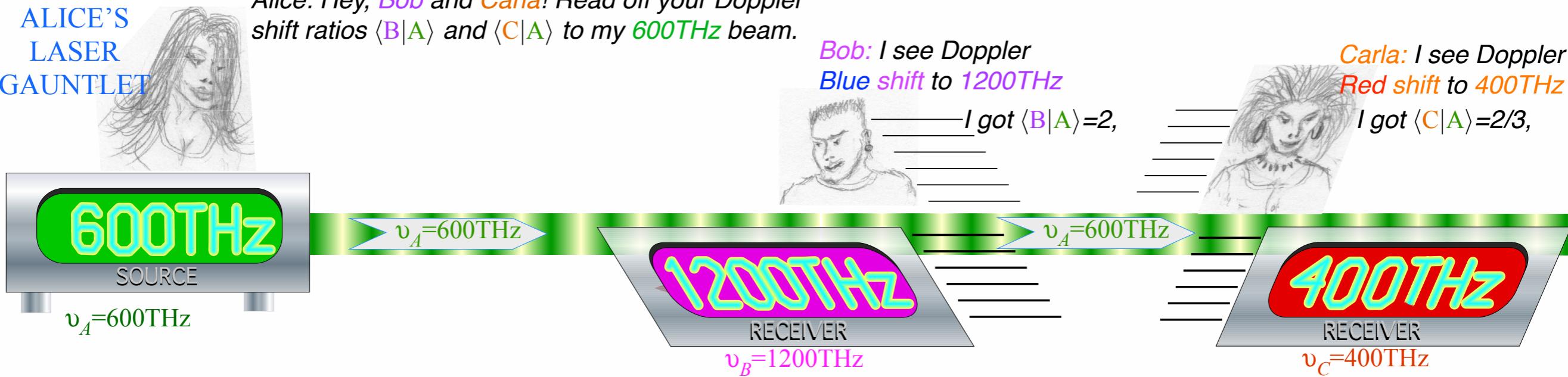
Carla-Alice Doppler ratio:

$$\langle C|A \rangle = \frac{v_C}{v_A} = \frac{400}{600} = \frac{2}{3}$$

Easy Doppler-shift and Rapidity calculation

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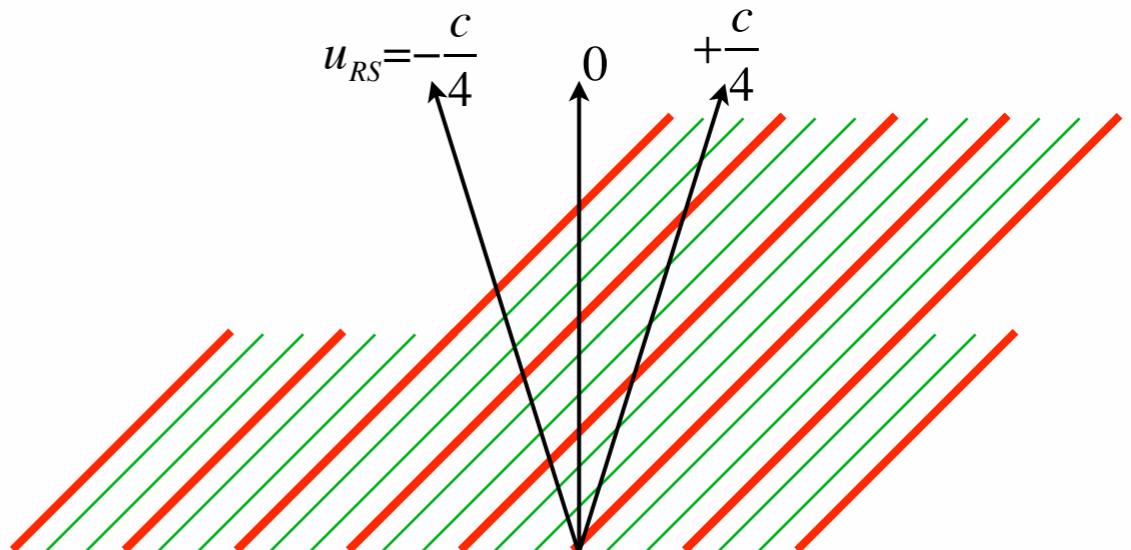
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IMPORTANT POINT:

Evenson axiom says Blue, Green, Red, etc. all march in lockstep and so *all* frequencies Doppler shift in same *geometric* proportion $\langle R|S \rangle$.



Easy Doppler-shift and Rapidity calculation

ALICE'S
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GAUNTLET

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IMPORTANT POINT:

Evenson axiom says Blue, Green, Red, etc. all march in lockstep and so *all* frequencies Doppler shift in same *geometric* proportion $\langle R|S \rangle$.

If Alice sends $v_A = 600\text{THz}$

Bob sees: $v_B = \langle B|A \rangle v_A = 1200\text{THz}$

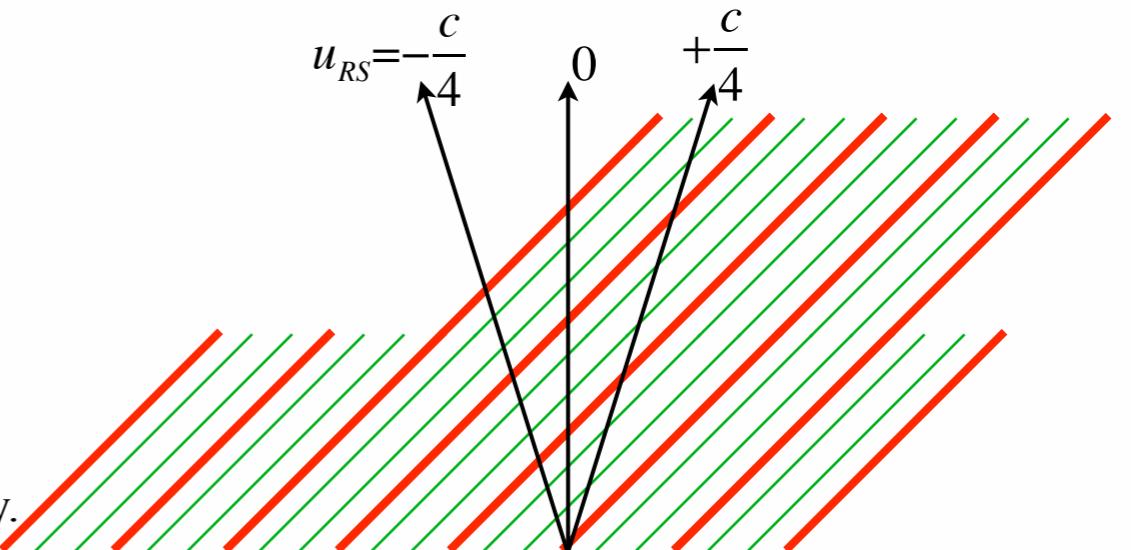
If Alice sends $v_A = 60\text{ THz}$

Bob sees: $v_B = \langle B|A \rangle v_A = 12\text{ THz}$

If Alice sends $v_A = 6\text{ Hz}$

Bob sees: $v_B = \langle B|A \rangle v_A = 12\text{ Hz}$

$\langle B|A \rangle = 2$ for any frequency Alice and Bob use while they maintain their relative velocity.



Easy Doppler-shift and Rapidity calculation

ALICE'S
LASER
GAUNTLET



Alice: Hey, Bob and Carla! Read off your Doppler shift ratios $\langle B|A \rangle$ and $\langle C|A \rangle$ to my 600THz beam.

Also, rapidity ρ_{BA} and ρ_{CA} relative to me.



Doppler ratio:

$$\langle R|S \rangle = \frac{v_{RECEIVER}}{v_{SOURCE}}$$

rapidity:

$$\rho_{RS} = \log_e \langle R|S \rangle$$

Definition of Rapidity

Rapidity is most convenient!

1 TeV proton has

$$u = 0.999995598 \cdot c \quad (\text{Pain in the A})$$

$$\text{or: } \langle R|S \rangle = 2131.6 \quad (\text{Better})$$

$$\text{or: } \rho_{RS} = 7.6646 \quad (\text{Best})$$

For low velocity $u \ll c$ rapidity ρ_{RS} approaches u/c

Bob-Alice Doppler ratio:

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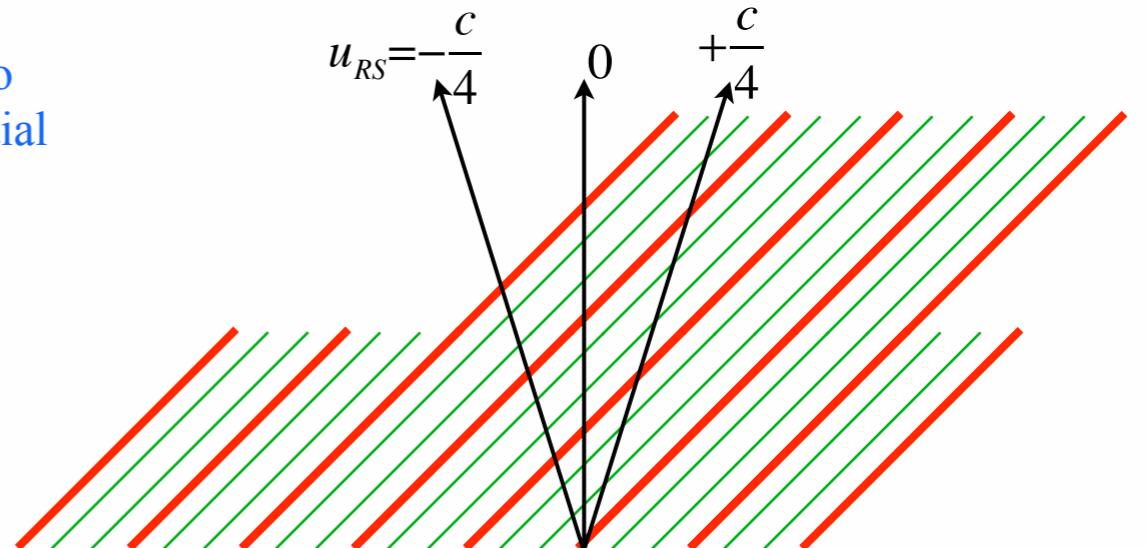
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IMPORTANT POINTS:

Evenson axiom says Blue, Green, Red, etc. all march in lockstep and so *all* frequencies Doppler shift in same *geometric* proportion $\langle R|S \rangle$.

Geometric phenomena tend to involve logarithmic/exponential functionality!



Easy Doppler-shift and Rapidity calculation

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GAUNTLET

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Definition of Rapidity

Bob: I see Doppler Blue shift to 1200THz

I got $\langle B|A \rangle = 2$,
and $\rho_{BA} = \ln 2$

Carla: I see Doppler Red shift to 400THz
I got $\langle C|A \rangle = 2/3$,

Bob-Alice Doppler ratio:

$$\langle B|A \rangle = \frac{v_B}{v_A} = \frac{1200}{600} = \frac{2}{1}$$

Bob-Alice rapidity:

$$\rho_{BA} = \log_e \langle B|A \rangle = \log_e \frac{2}{1}$$

Carla-Alice Doppler ratio:

$$\langle C|A \rangle = \frac{v_C}{v_A} = \frac{400}{600} = \frac{2}{3}$$

Easy Doppler-shift and Rapidity calculation

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Bob-Alice Doppler ratio:

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Definition of Rapidity

$$\langle B|A \rangle = \frac{v_B}{v_A} = \frac{2}{1}$$

is time-reversal of:

$$\langle A|B \rangle = \frac{v_A}{v_B} = \frac{1}{2}$$

Bob-Alice Doppler ratio:

$$\langle B|A \rangle = \frac{v_B}{v_A} = \frac{1200}{600} = \frac{2}{1}$$

Bob-Alice rapidity:

$$\rho_{BA} = \log_e \langle B|A \rangle = \log_e \frac{2}{1} \quad (\text{time-reversed})$$

$$\rho_{BA} = 0.69 \quad (\text{so: } \rho_{AB} = -0.69)$$

Carla-Alice Doppler ratio:

$$\langle C|A \rangle = \frac{v_C}{v_A} = \frac{400}{600} = \frac{2}{3}$$

Carla-Alice rapidity:

$$\rho_{CA} = \log_e \langle C|A \rangle = \log_e \frac{2}{3}$$

$$\rho_{CA} = -0.41$$

Mnemonic: You can think of rapidity ρ_{BA} as “R” for “Romance”... (+) positive on approach, (-) negative on reproach

Do the stars hate us?

Easy Doppler-shift and Rapidity calculation

ALICE'S
LASER
GAUNTLET



Alice: Hey, Bob and Carla! Read off your Doppler shift ratios $\langle B|A \rangle$ and $\langle C|A \rangle$ to my 600THz beam.

Also, rapidity ρ_{BA} and ρ_{CA} relative to me.

Now, Carla, what's your rapidity ρ_{CB} relative to Bob?



Doppler ratio:

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rapidity:

$$\rho_{RS} = \log_e \langle R|S \rangle$$

Definition of Rapidity

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Easy Doppler-shift and Rapidity calculation

ALICE'S
LASER
GAUNTLET



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Doppler ratio:

$$\langle R|S \rangle = \frac{v_{RECEIVER}}{v_{SOURCE}}$$

rapidity:

$$\rho_{RS} = \log_e \langle R|S \rangle$$

so:

$$\langle R|S \rangle = e^{\rho_{RS}}$$

Definition of Rapidity

$$\langle B|A \rangle = \frac{v_B}{v_A}$$

is time-reversed

$$\langle A|B \rangle = \frac{v_A}{v_B}$$

Alice: Hey, Bob and Carla! Read off your Doppler shift ratios $\langle B|A \rangle$ and $\langle C|A \rangle$ to my 600THz beam.

Also, rapidity ρ_{BA} and ρ_{CA} relative to me.

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$$I got \langle B|A \rangle = 2, \text{ and } \rho_{BA} = \ln(2) = +0.69$$

Carla: I see Doppler Red shift to 400THz

$$I got \langle C|A \rangle = 2/3, \text{ and } \rho_{CA} = \ln(2/3) = -0.41$$

Bob-Alice Doppler ratio:

$$\langle B|A \rangle = \frac{v_B}{v_A} = \frac{1200}{600} = \frac{2}{1}$$

Bob-Alice rapidity:

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Carla-Alice Doppler ratio:

$$\langle C|A \rangle = \frac{v_C}{v_A} = \frac{400}{600} = \frac{2}{3}$$

Carla-Alice rapidity:

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$$\rho_{CA} = -0.41$$

Carla-Bob Doppler ratio:

$$\langle C|B \rangle = \frac{v_C}{v_B} = \frac{v_C}{v_A} \frac{v_A}{v_B} = \langle C|A \rangle \langle A|B \rangle$$

Carla-Bob rapidity:

$$e^{\rho_{CB}} = e^{\rho_{CA}} e^{\rho_{AB}}$$

Easy Doppler-shift and Rapidity calculation

ALICE'S
LASER
GAUNTLET



Alice: Hey, Bob and Carla! Read off your Doppler shift ratios $\langle B|A \rangle$ and $\langle C|A \rangle$ to my 600THz beam.

Also, rapidity ρ_{BA} and ρ_{CA} relative to me.

Now, Carla, what's your rapidity ρ_{CB} relative to Bob?



Doppler ratio:

$$\langle R|S \rangle = \frac{v_{RECEIVER}}{v_{SOURCE}}$$

rapidity:

$$\rho_{RS} = \log_e \langle R|S \rangle$$

or:

$$\langle R|S \rangle = e^{\rho_{RS}} = e^{-\rho_{SR}}$$

Definition of Rapidity

$$\langle B|A \rangle = \frac{v_B}{v_A}$$

is time-reversed

$$\langle A|B \rangle = \frac{v_A}{v_B}$$

Alice: Hey, Bob and Carla! Read off your Doppler shift ratios $\langle B|A \rangle$ and $\langle C|A \rangle$ to my 600THz beam.

Also, rapidity ρ_{BA} and ρ_{CA} relative to me.

Bob: I see Doppler Blue shift to 1200THz



I got $\langle B|A \rangle = 2$,
and $\rho_{BA} = \ln(2) = +0.69$



Carla: I see Doppler Red shift to 400THz

I got $\langle C|A \rangle = 2/3$,
and $\rho_{CA} = \ln(2/3) = -0.41$

Bob-Alice Doppler ratio:

$$\langle B|A \rangle = \frac{v_B}{v_A} = \frac{1200}{600} = \frac{2}{1}$$

Bob-Alice rapidity:

$$\rho_{BA} = \log_e \langle B|A \rangle = \log_e \frac{2}{1}$$

$$\rho_{BA} = 0.69 \quad (\text{so: } \rho_{AB} = -0.69)$$

Carla-Alice Doppler ratio:

$$\langle C|A \rangle = \frac{v_C}{v_A} = \frac{400}{600} = \frac{2}{3}$$

Carla-Alice rapidity:

$$\rho_{CA} = \log_e \langle C|A \rangle = \log_e \frac{2}{3}$$

$$\rho_{CA} = -0.41$$

Carla-Bob Doppler ratio:

$$\langle C|B \rangle = \frac{v_C}{v_B} = \frac{v_C}{v_A} \frac{v_A}{v_B} = \langle C|A \rangle \langle A|B \rangle$$

Carla-Bob rapidity:

$$e^{\rho_{CB}} = e^{\rho_{CA}} e^{\rho_{AB}} \quad \text{implies: } \rho_{CB} = \rho_{CA} + \rho_{AB} \\ = -0.41 - 0.69 = -1.10$$

Easy Doppler-shift and Rapidity calculation

ALICE'S
LASER
GAUNTLET



Alice: Hey, Bob and Carla! Read off your Doppler shift ratios $\langle B|A \rangle$ and $\langle C|A \rangle$ to my 600THz beam.

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Carla: I see Doppler Red shift to 400THz

$$I got \langle C|A \rangle = 2/3, \text{ and } \rho_{CA} = \ln(2/3) = -0.41$$

I got $\langle C|B \rangle = \langle C|A \rangle \langle A|B \rangle = (2/3)(1/2) = 1/3$,
and $\rho_{CB} = \rho_{CA} + \rho_{AB} = -1.10$
We're in Splitsville!

Bob-Alice Doppler ratio:

$$\langle B|A \rangle = \frac{v_B}{v_A} = \frac{1200}{600} = \frac{2}{1}$$

Bob-Alice rapidity:

$$\rho_{BA} = \log_e \langle B|A \rangle = \log_e \frac{2}{1}$$

$$\rho_{BA} = 0.69 \quad (\text{so: } \rho_{AB} = -0.69)$$

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$$\langle C|A \rangle = \frac{v_C}{v_A} = \frac{400}{600} = \frac{2}{3}$$

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Carla-Bob rapidity:

$$e^{\rho_{CB}} = e^{\rho_{CA}} e^{\rho_{AB}} \text{ implies:}$$

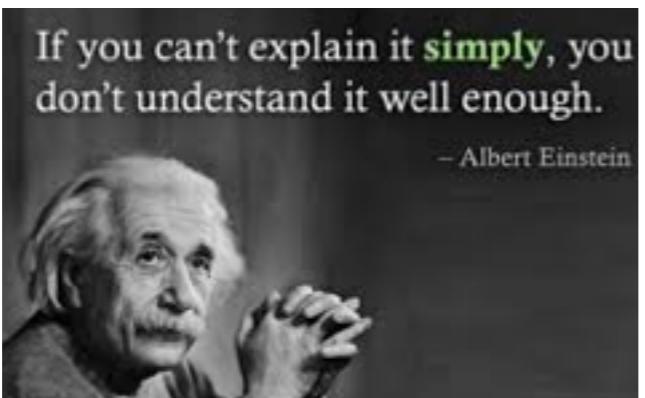
Galileo's Revenge (part 1)

Rapidity adds just like Galilean velocity

$$\rho_{CB} = \rho_{CA} + \rho_{AB} = -0.41 - 0.69 = -1.10$$

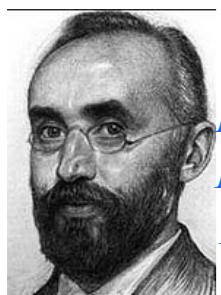
Two Famous-Name Coefficients

Albert Einstein
1859-1955



This number
is called an: Einstein time-dilation
(dilated by 25% here)

This number
is called a: Lorentz length-contraction
(contracted by 20% here)



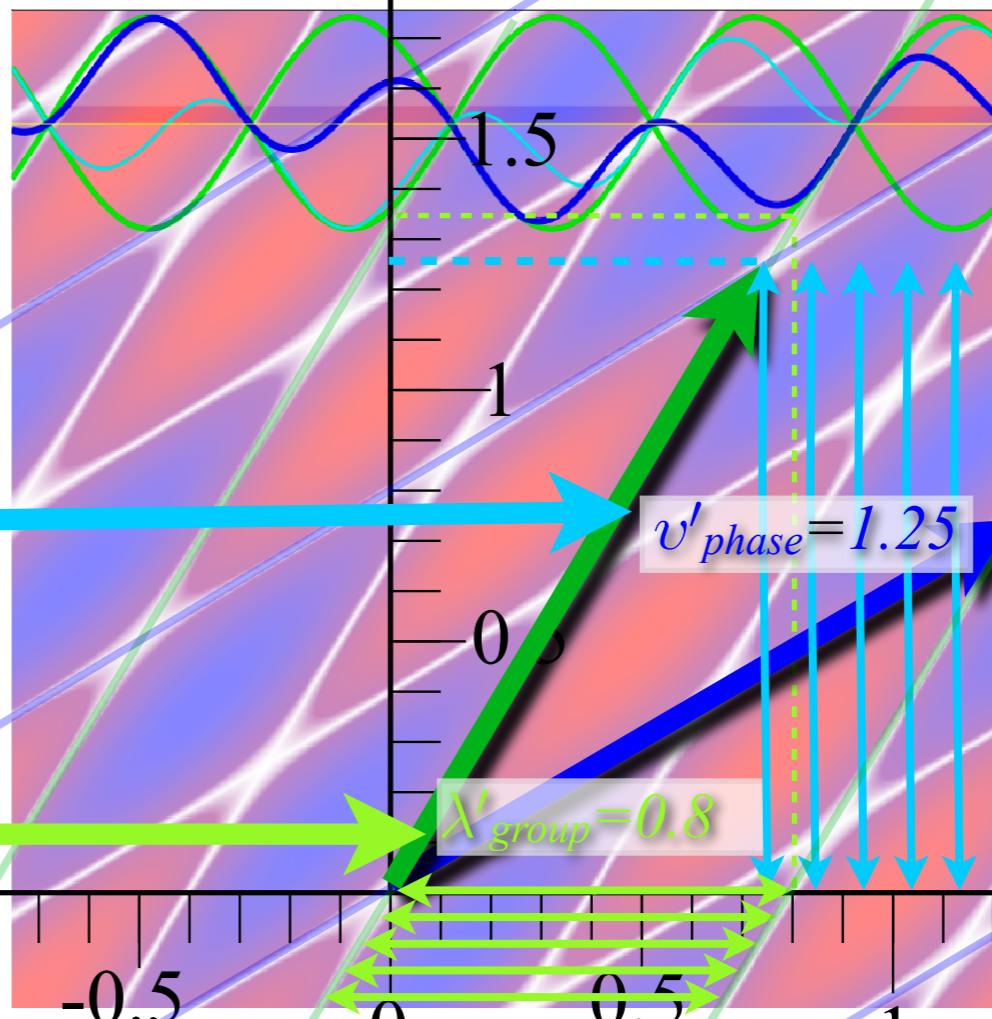
Hendrik A.
Lorentz
1853-1928

Old-Fashioned Notation

Time ct'
(units of
 $\lambda_A = 1/2\mu m$)



Herman
Minkowski
1864-1909



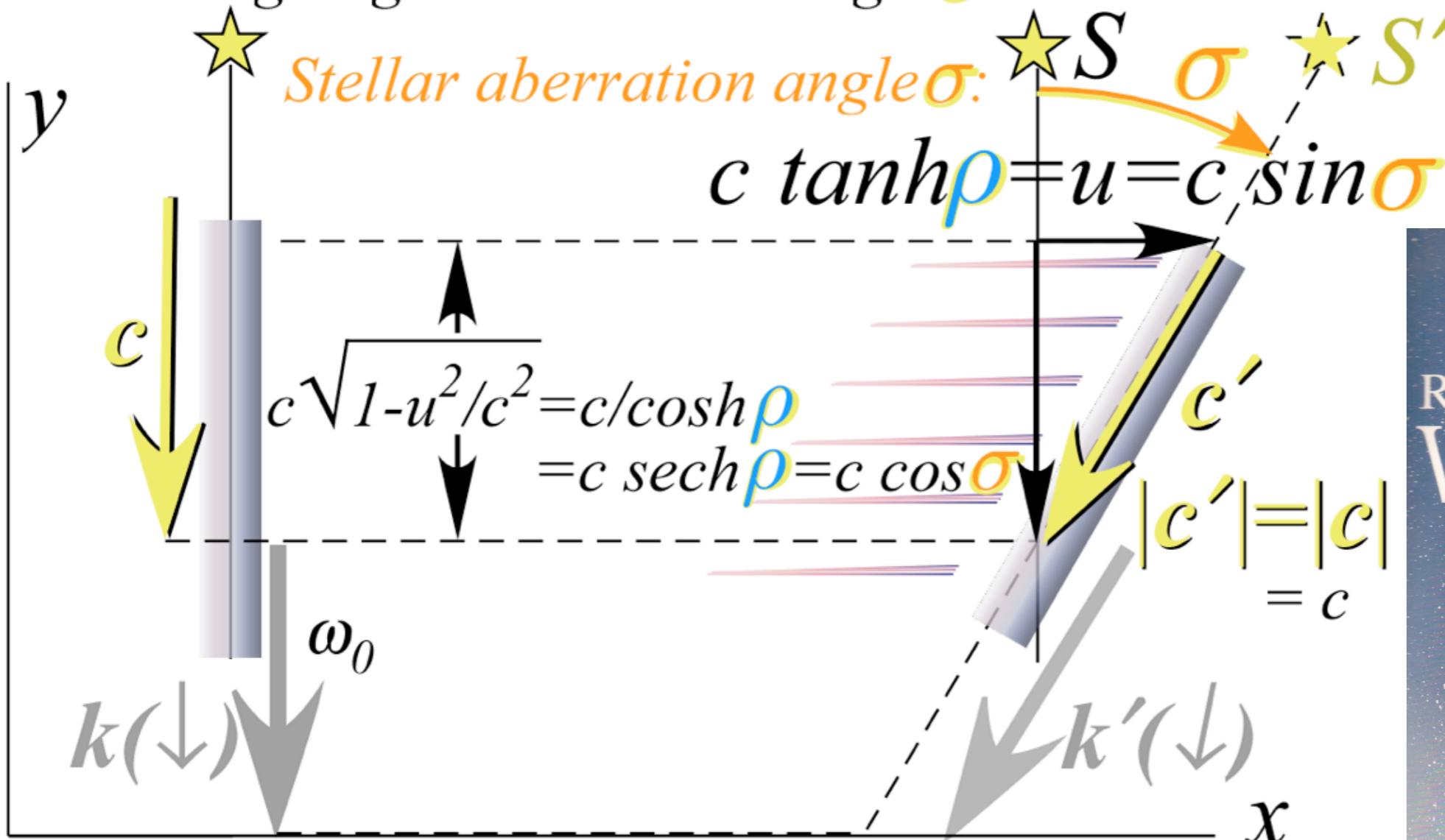
phase	$b_{RED}^{Doppler}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
group	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\text{sech } \rho$	$\cosh \rho$	$\text{csch } \rho$	$\coth \rho$	$e^{+\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

Comparing Longitudinal relativity parameter: Rapidity $\rho = \log_e(\text{Doppler Shift})$ to a Transverse* relativity parameter: Stellar aberration angle σ

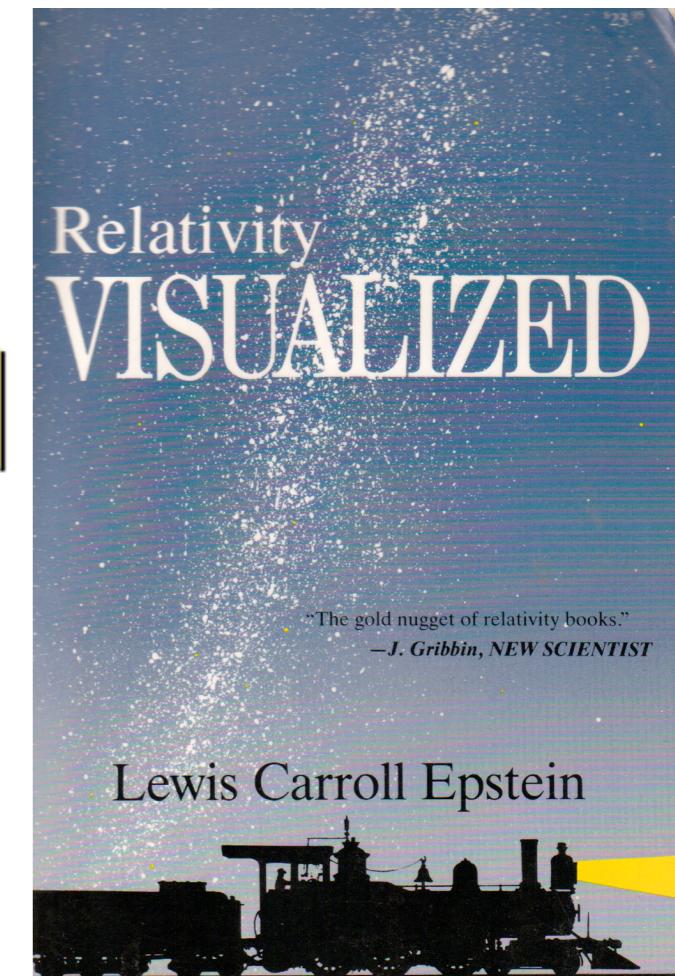
*Lewis Carroll Epstein, *Relativitätstheorie*, Birkhäuser, (2004) Earlier English version (1985)-

Observer fixed below star sees it directly overhead.

Observer going u sees star at angle σ in u direction.

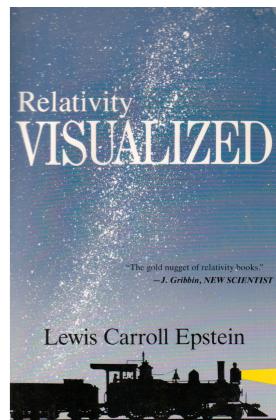


We used notion σ for stellar-ab-angle, (a “flipped-out” ρ). Epstein not interested in ρ analysis or in relation of σ and ρ .



Comparing Longitudinal relativity parameter: Rapidity $\rho = \log_e(\text{Doppler Shift})$ to a Transverse* relativity parameter: Stellar aberration angle σ

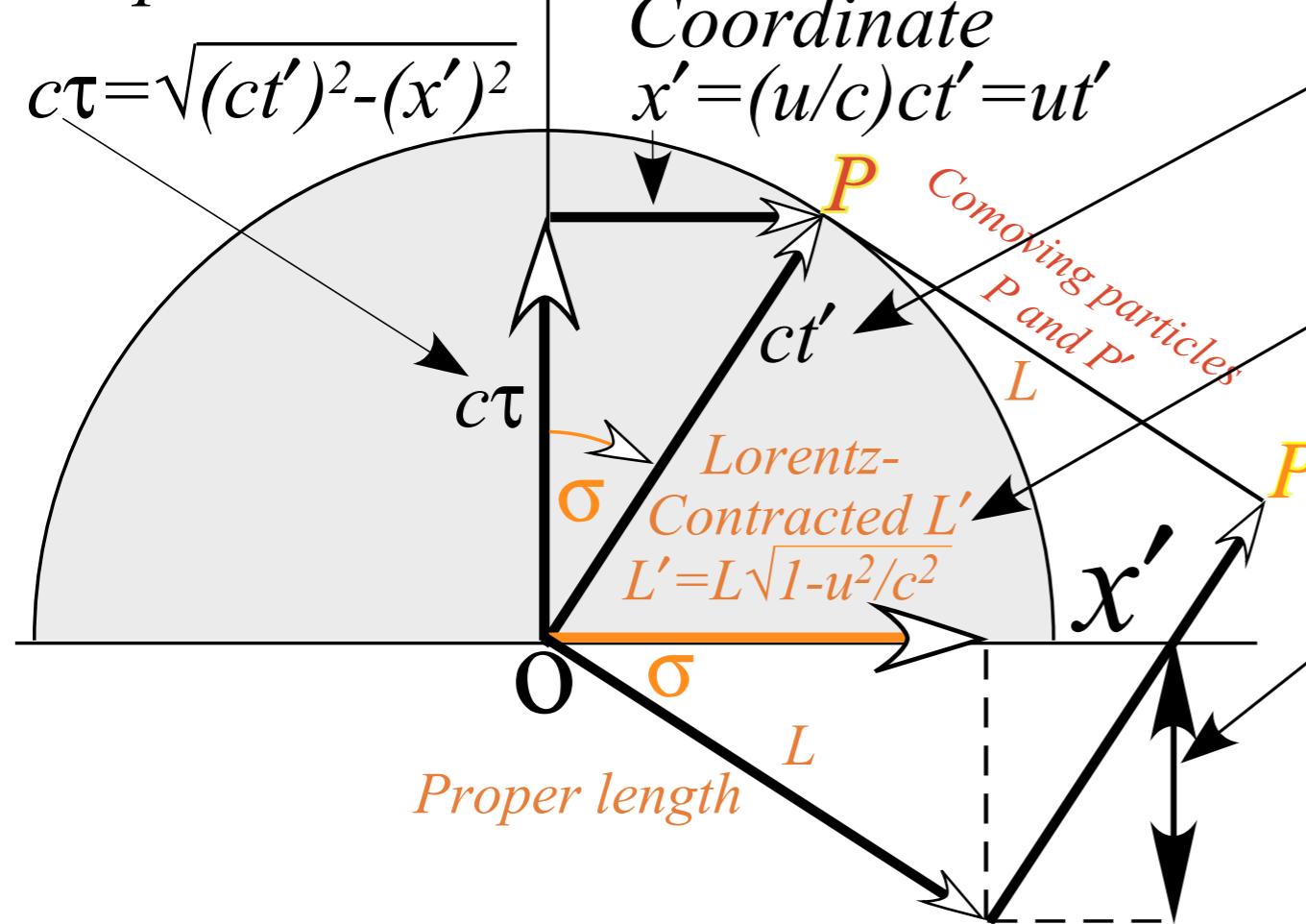
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Proper time $c\tau$ vs. coordinate space x - (L. C. Epstein's "Cosmic Speedometer")

Particles P and P' have speed u in (x', ct') and speed c in $(x, c\tau)$

Proper time $C\tau$



Einstein time dilation:

$$ct' = c\tau \sec \sigma = c\tau \cosh \rho = c\tau / \sqrt{1-u^2/c^2}$$

Lorentz length contraction:

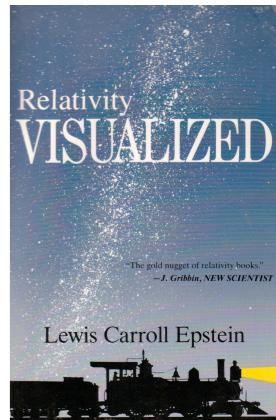
$$L' = L \operatorname{sech} \rho = L \cos \sigma = L \cdot \sqrt{1-u^2/c^2}$$

Proper Time asimultaneity:

$$\begin{aligned} c \Delta \tau &= L' \sinh \rho = L \cos \sigma \sinh \rho \\ &= L \cos \sigma \tan \sigma \\ &= L \sin \sigma = L / \sqrt{c^2/u^2 - 1} \sim L u/c \end{aligned}$$

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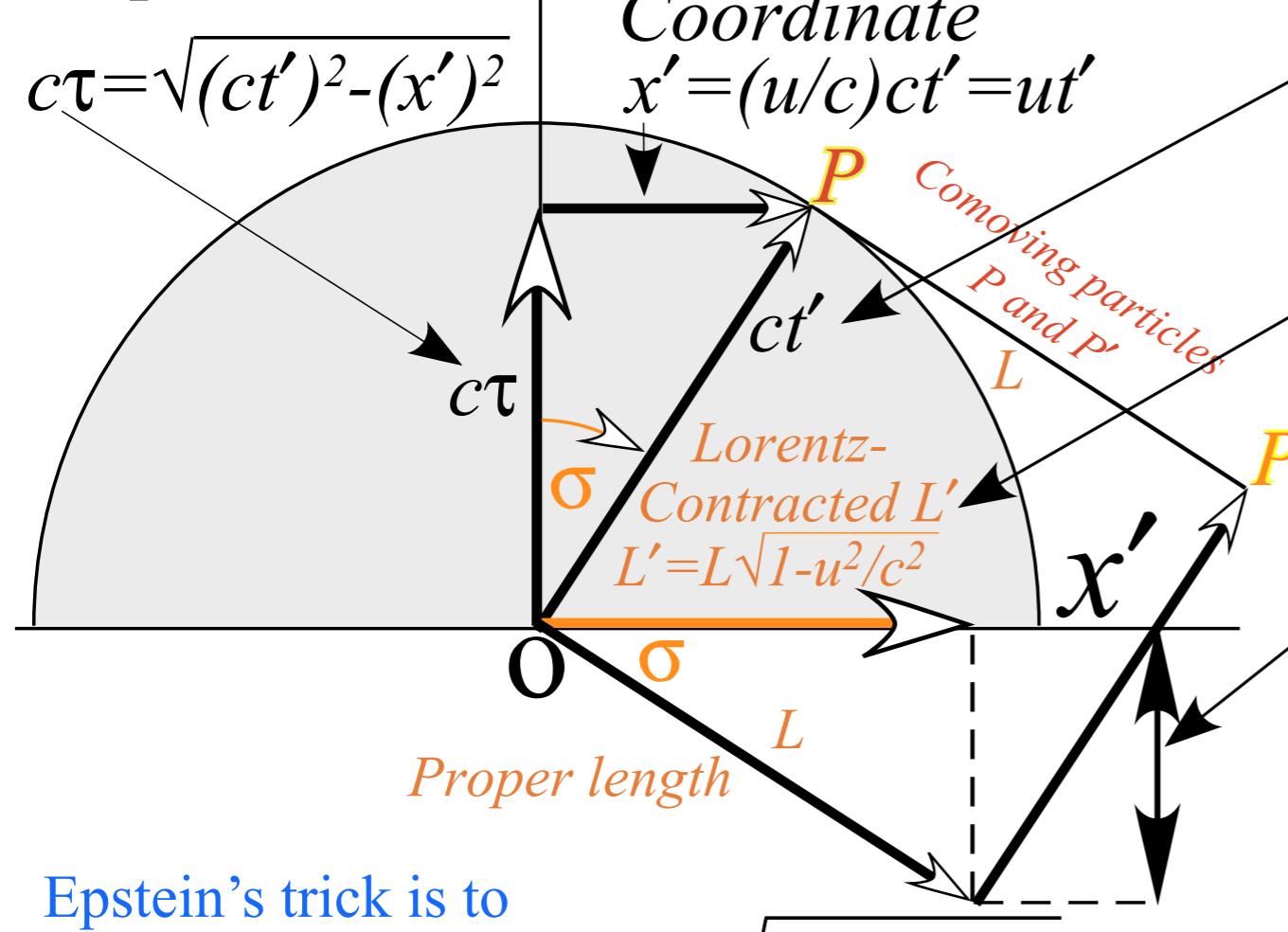
*Lewis Carroll Epstein, *Relativitätstheorie*, Birkhäuser, (2004) Earlier English version (1985)-



Proper time $c\tau$ vs. coordinate space x - (L. C. Epstein's "Cosmic Speedometer")

Particles P and P' have speed u in (x', ct') and speed c in $(x, c\tau)$

Proper time $C\Tau$



Epstein's trick is to

turn a hyperbolic form $c\tau = \sqrt{(ct')^2 - (x')^2}$

into a circular form:

$$\sqrt{(c\tau)^2 + (x')^2} = (ct')$$

Then everything (and everybody) always goes speed c through $(x', c\tau)$ space!

Einstein time dilation:

$$ct' = c\tau \sec \sigma = c\tau \cosh \rho = c\tau / \sqrt{1-u^2/c^2}$$

Lorentz length contraction:

$$L' = L \operatorname{sech} \rho = L \cos \sigma = L \cdot \sqrt{1-u^2/c^2}$$

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Using (some) wave parameters to develop relativistic quantum theory

$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \text{ (for } u \ll c)$$

$$c\kappa_{phase} = B \sinh \rho \approx B \rho \quad \text{(for } u \ll c)$$

$$B = v_A$$

$$B = v_A = c\kappa_A$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2$$

$$\sinh \rho \approx \rho$$

At low speeds: ..

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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At low speeds:

$$\begin{aligned} \cosh \rho &\approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2} \\ \sinh \rho &\approx \rho \approx \frac{u}{c} \end{aligned}$$

$$B = v_A$$

$$B = v_A = c\kappa_A$$

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	v_{group}	λ_{group}	κ_{group}	τ_{group}	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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 \end{aligned}$$

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 \end{aligned}$$

$$\begin{aligned}
 B &= v_A \\
 B &= v_A = c\kappa_A
 \end{aligned}$$

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phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
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$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2 \iff \text{for } (u \ll c) \Rightarrow$

$$\begin{aligned} \cosh \rho &\approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2} \\ \sinh \rho &\approx \rho \approx \frac{u}{c} \end{aligned}$$

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At low speeds:

$$\kappa_{phase} \approx \frac{B}{c^2} u$$

time	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$b_{RED}^{Doppler}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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At low speeds:

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$$\kappa_{phase} \approx \frac{B}{c^2} u$$

v_{phase} and κ_{phase} resemble
formulae for Newton's
kinetic energy and momentum

Resembles: $const. + \frac{1}{2} M u^2$

Resembles: $M u$

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	v_{group}	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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$$\sinh \rho \approx \rho \approx \frac{u}{c}$$

$$B = v_A$$

$$B = v_A = c\kappa_A$$

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2 \Leftarrow \text{for } (u \ll c) \Rightarrow$$

Rescale v_{phase} by \hbar so: $M = \frac{\hbar B}{c^2}$

Resembles: $const. + \frac{1}{2} Mu^2$

$$\kappa_{phase} \approx \frac{B}{c^2} u$$

Resembles: Mu

v_{phase} and κ_{phase} resemble formulae for Newton's kinetic energy $\frac{1}{2} Mu^2$ and momentum Mu .

time	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	v_{group}	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$b_{RED}^{Doppler}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

Using (some) wave parameters to develop relativistic quantum theory

$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \text{ (for } u \ll c)$$

$$c\kappa_{phase} = B \sinh \rho \approx B \rho \quad (\text{for } u \ll c)$$

$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2}$$

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$$B = v_A$$

$$B = v_A = c\kappa_A$$

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Rescale v_{phase} by \hbar so: $M = \frac{\hbar B}{c^2}$

Resembles: $const. + \frac{1}{2} Mu^2$

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phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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Resembles: $\text{const.} + \frac{1}{2} Mu^2$

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phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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Resembles: $\text{const.} + \frac{1}{2} Mu^2$

At low speeds: $\kappa_{phase} \approx \frac{B}{c^2} u$
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phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
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$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

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phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
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$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
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$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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(old-fashioned notation)

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(old-fashioned notation)

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phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

Using (some) wave parameters to develop relativistic quantum theory

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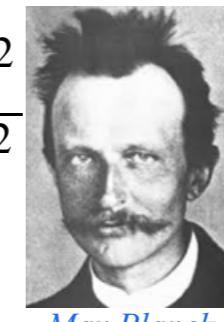
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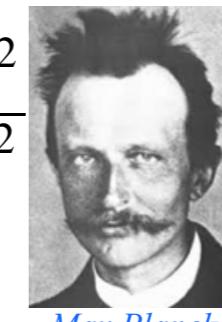
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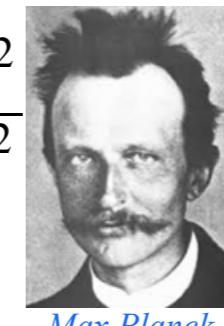
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Resolution and dirty secret: E , N , and v_{phase} are all frequencies!

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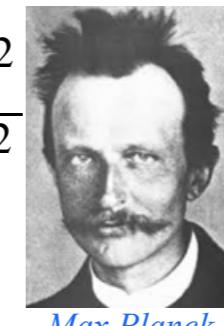
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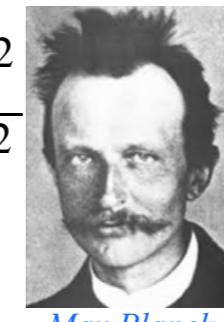
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$$cp = \frac{Mcu}{\sqrt{1-u^2/c^2}}$$

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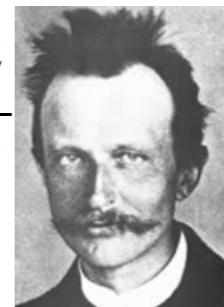
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$$h\nu_{phase} \approx Mc^2 + \frac{1}{2} Mu^2 \quad \Leftarrow \text{for } (u \ll c) \Rightarrow$$

$$\begin{aligned} \cosh \rho &\approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2} \\ \sinh \rho &\approx \rho \approx \frac{u}{c} \end{aligned}$$



Max Planck
1858-1947



Louis DeBroglie
1892-1987

$$\begin{aligned} B &= v_A \\ B &= v_A = c\kappa_A \end{aligned}$$

At low speeds:

$$\kappa_{phase} \approx \frac{B}{c^2} u$$

(The famous Mc^2 shows up here!)

v_{phase} and κ_{phase} resemble formulae for Newton's kinetic energy $\frac{1}{2}Mu^2$ and momentum Mu .

So attach scale factor h (or hN) to match units.

~~Natural wave conspiracy~~
~~Lucky coincidences??~~ ~~Expensive~~
~~Cheap trick??~~
... Try exact v_{phase} and κ_{phase} ...

$$h\nu_{phase} = hB \cosh \rho = Mc^2 \cosh \rho$$

↑ Planck (1900)

$$= \text{Total Energy: } E = \frac{Mc^2}{\sqrt{1-u^2/c^2}}$$

↓ Einstein (1905)

$$h\kappa_{phase} = hB \sinh \rho = Mc^2 \sinh \rho$$

$$\frac{1}{\sqrt{\beta^{-2}-1}} = \frac{u}{c} \quad (\text{old-fashioned notation})$$

$$cp = \frac{Mu}{\sqrt{1-u^2/c^2}}$$

$$\text{Momentum: } h\kappa_{phase} = p = \frac{Mu}{\sqrt{1-u^2/c^2}}$$

DeBroglie (1921)

group	$b_{Doppler RED}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$
phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$
stellar √ angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$
						$\frac{2}{1}=2.0$	

Using (some) wave parameters to develop relativistic quantum theory

$$\begin{aligned} v_{phase} &= B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \text{ (for } u \ll c) \\ c\kappa_{phase} &= B \sinh \rho \approx B \rho \quad \text{(for } u \ll c) \\ \frac{u}{c} &= \tanh \rho \approx \rho \quad \text{(for } u \ll c) \end{aligned}$$

$$\begin{aligned} \cosh \rho &\approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2} \\ \sinh \rho &\approx \rho \approx \frac{u}{c} \end{aligned}$$



Max Planck
1858-1947



Louis DeBroglie
1892-1987

$$\begin{aligned} B &= v_A \\ B &= v_A = c\kappa_A \end{aligned}$$

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2 \quad \Leftarrow \text{for } (u \ll c) \Rightarrow$$

$$\text{Rescale } v_{phase} \text{ by } h \text{ so: } M = \frac{hB}{c^2} \quad \text{or: } hB = Mc^2$$

$$\kappa_{phase} \approx \frac{B}{c^2} u \quad (\text{The famous } Mc^2 \text{ shows up here!})$$

$$h\nu_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2 \quad \Leftarrow \text{for } (u \ll c) \Rightarrow$$

$$h\nu_{phase} \approx Mc^2 + \frac{1}{2} Mu^2 \quad \Leftarrow \text{for } (u \ll c) \Rightarrow$$

$$h\kappa_{phase} \approx \frac{hB}{c^2} u$$

$$h\kappa_{phase} \approx Mu$$

v_{phase} and κ_{phase} resemble formulae for Newton's kinetic energy $\frac{1}{2}Mu^2$ and momentum Mu .

So attach scale factor h (or hN) to match units.

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$$= \text{Total Energy: } E = \frac{Mc^2}{\sqrt{1-u^2/c^2}}$$

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$$\text{Momentum: } h\kappa_{phase} = p = \frac{Mu}{\sqrt{1-u^2/c^2}}$$

DeBroglie (1921)

group	$b_{Doppler RED}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$
phase	$\frac{1}{b_{Doppler BLUE}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$		
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$		
stellar √ angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$
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