

Current understanding of relativity and QM at UAF



NWAT photo by David Gottschalk

Current understanding of relativity and QM at UAF



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Is a clearer understanding possible...?

Level 1 Secrets (*which really shouldn't be secrets at all!*)

Special relativity and quantum mechanics
are very much a story of
the geometry of light-wave motion

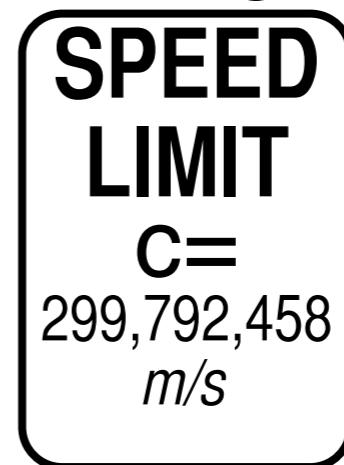
- How badly does Galilean relativity fail for light waves?

- How do you make sense of light-wave

The *Einstein Pulse Wave (PW)* axiom

versus

The *Evenson Continuous Wave (CW)* axiom



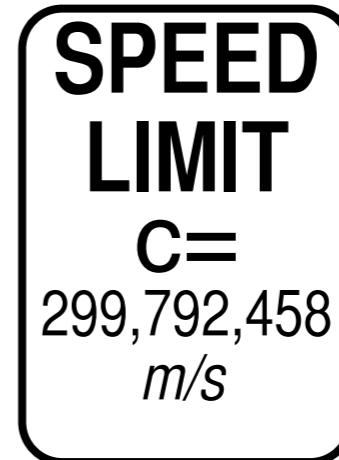
axiom(s)?

*Good approximation:
c=300 million m/s*

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- How badly does Galilean relativity fail for light waves?
- How do you make sense of light-wave axiom(s)?
 - The *Einstein Pulse Wave* (PW) axiom
 - versus*
 - The *Evenson Continuous Wave* (CW) axiom
- How does **space-time** and/or *per-space-per-time* carry light-waves?



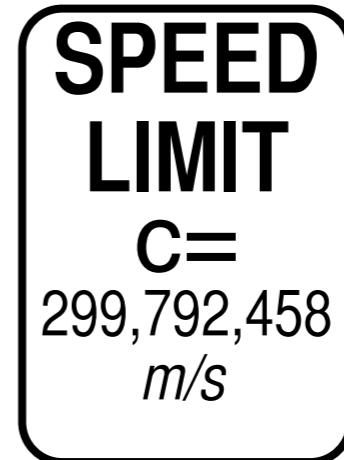
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- How does **space-time** and/or **per-space-per-time** carry light-waves?
(*wavelength λ - period τ*) and/or (*wavenumber κ - frequency v*)
($\lambda = 1/\kappa$ and $\tau = 1/v$) ($\kappa = 1/\lambda$ and $v = 1/\tau$)



axiom(s)?
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(*wavelength λ - period τ*)

and/or

(*wavenumber κ - frequency v*)

$$(\lambda = 1/\kappa \quad \text{and} \quad \tau = 1/v)$$

$$(\kappa = 1/\lambda \quad \text{and} \quad v = 1/\tau)$$

(λ = *meters per wave* and τ = *seconds per wave*) (κ = *waves per meter* and v = *waves per second*)

**SPEED
LIMIT**
C= 299,792,458 m/s

axiom(s)?

*Good approximation:
c=300 million m/s*

Heinrich
Hertz

1857-1894
1Hz=1sec⁻¹



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(*wavelength λ - period τ*)

($\lambda = 1/\kappa$ and $\tau = 1/v$)

($\lambda = \text{meters per wave}$ and $\tau = \text{seconds per wave}$)

Greek "L"
for Length

Greek "t"
for time

(*wavenumber κ - frequency v*)

($\kappa = 1/\lambda$ and $v = 1/\tau$)

($\kappa = \text{waves per meter}$ and $v = \text{waves per second}$)

Greek "k"
for Kayser
(or "kinks")

Heinreich
Hertz
1857-1894
1Hz=1sec⁻¹

Greek "n" for number
of waves per second
or Hertz (Hz)

**SPEED
LIMIT**
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299,792,458
m/s

*Good approximation:
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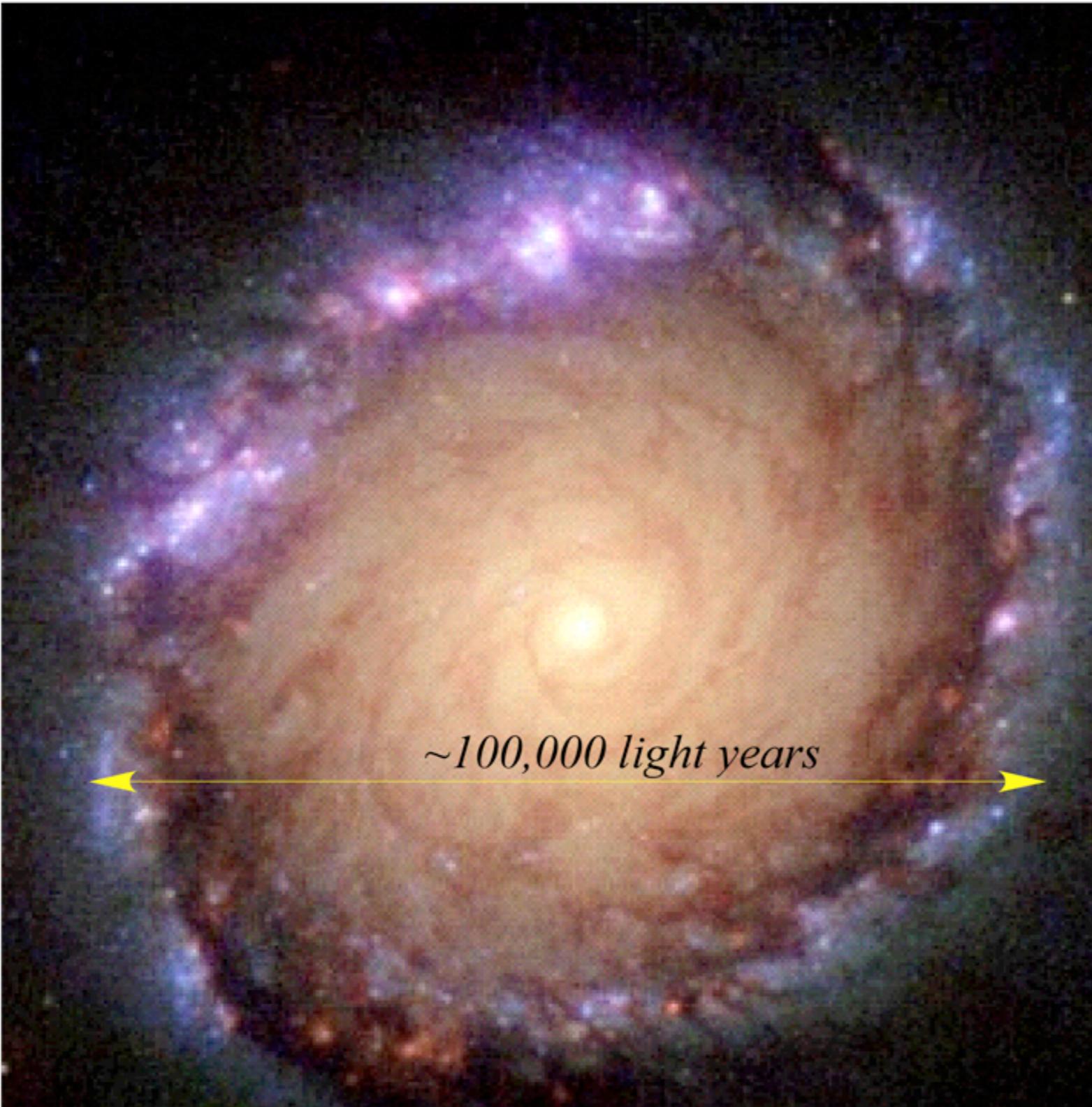


How fast is light? Light goes one foot in a nano-second .

This may seem quite fast to us.

But, on a cosmic scale lightspeed is positively sub-glacial.

In your lifetime light cannot move across one pixel (.) of a Hubble deep-sky photo.



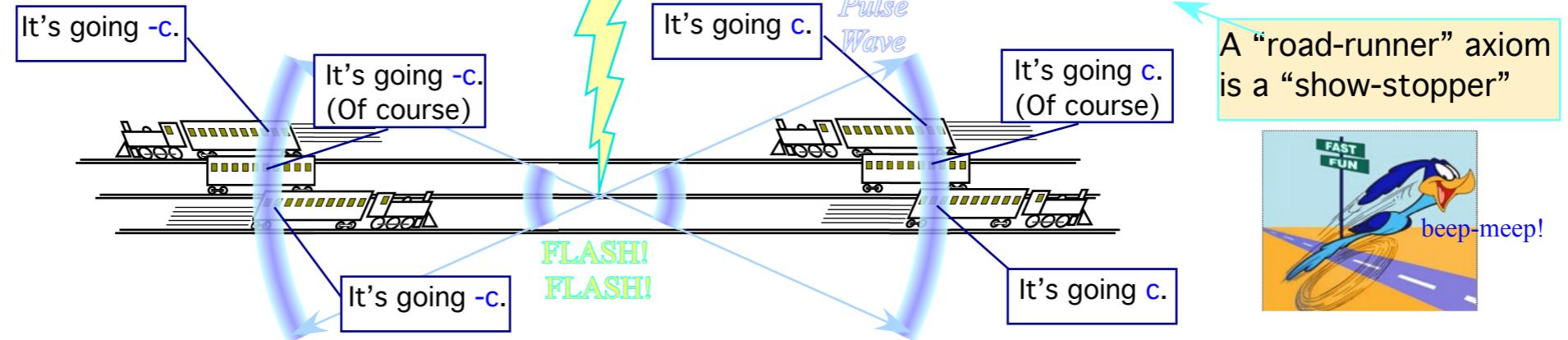
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Albert Einstein

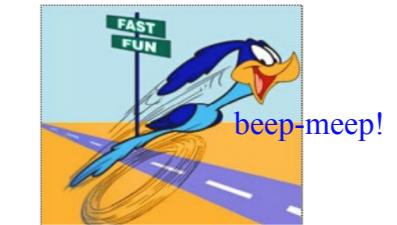


1879-1955

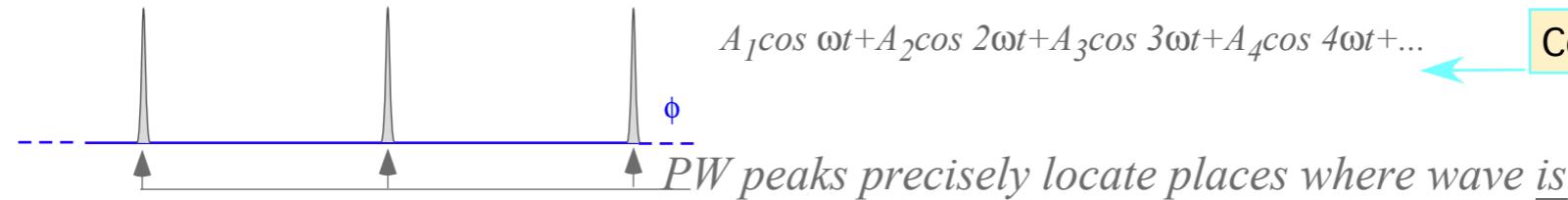
Einstein Pulse Wave (PW) Axiom: PW speed seen by all observers is c



A “road-runner” axiom is a “show-stopper”



Pulse wave (PW) train



Complicated

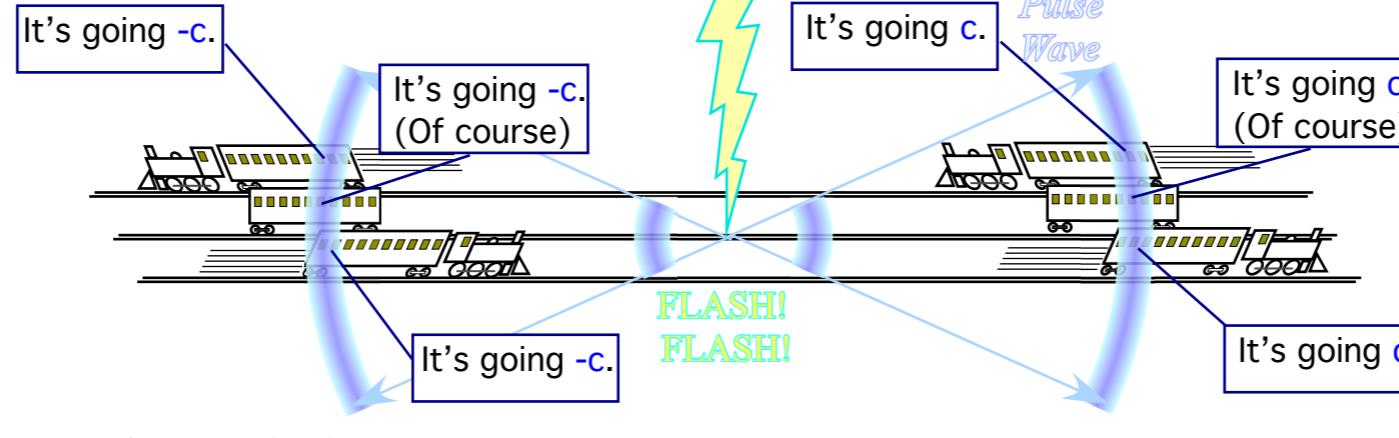
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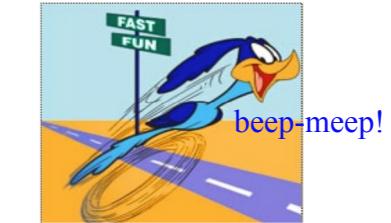


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Pulse wave (PW) train

$$A_1 \cos \omega t + A_2 \cos 2\omega t + A_3 \cos 3\omega t + A_4 \cos 4\omega t + \dots$$

Complicated

PW forms are also called Wave Packets (WP)

since

they are
interfering

sums of
many
CW terms

(10-Cosine Waves
make up this pulse)

CW terms are
also called

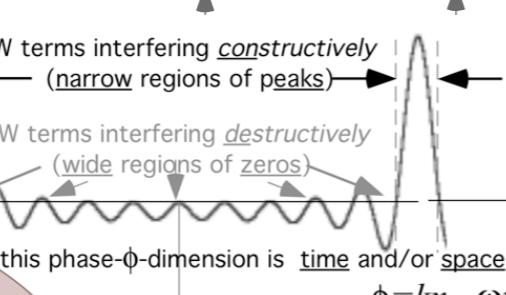
Color Waves

or

Fourier

Spectral

Components

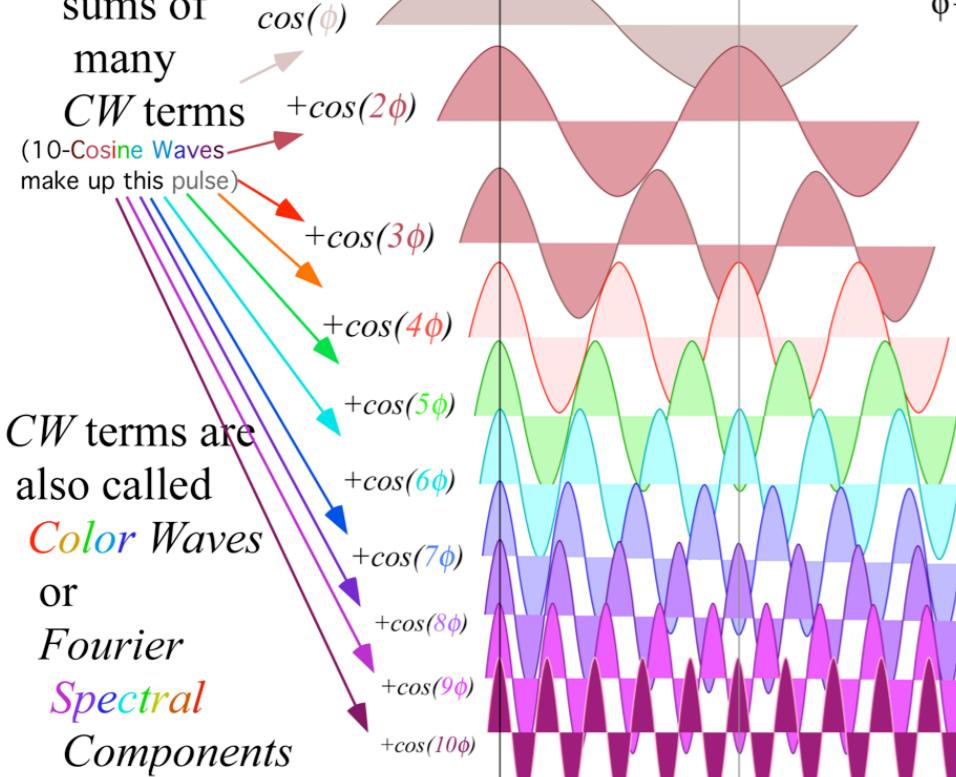


CW terms interfering constructively
(narrow regions of peaks)

CW terms interfering destructively
(wide regions of zeros)

(this phase- ϕ -dimension is time and/or space)

$$\phi = kx - \omega t$$



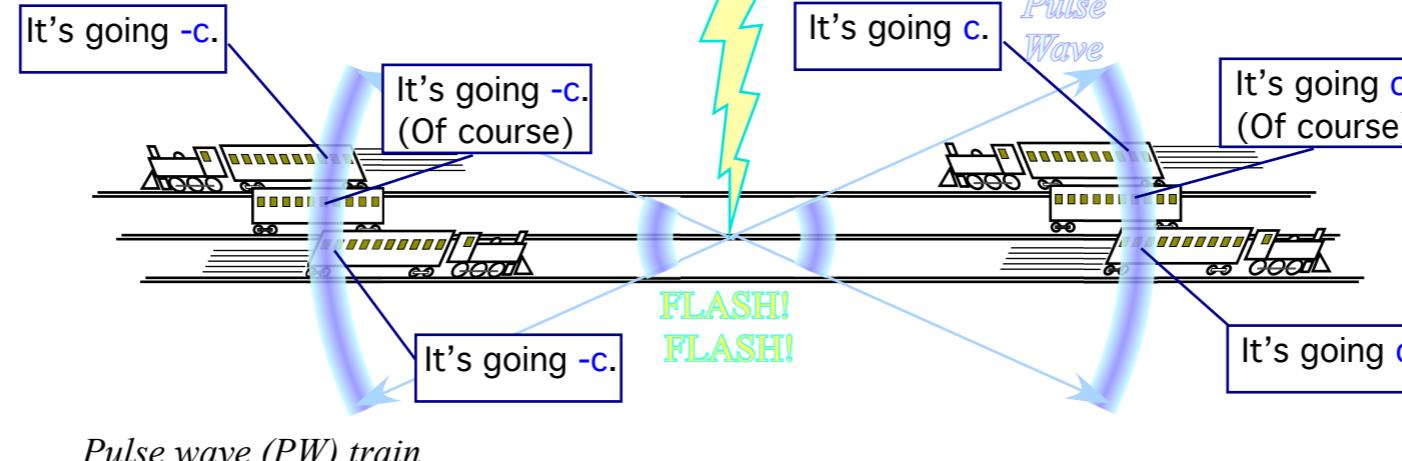
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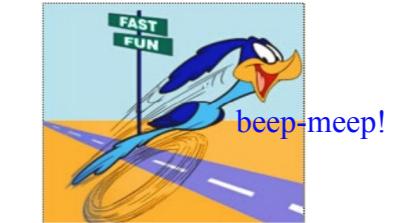


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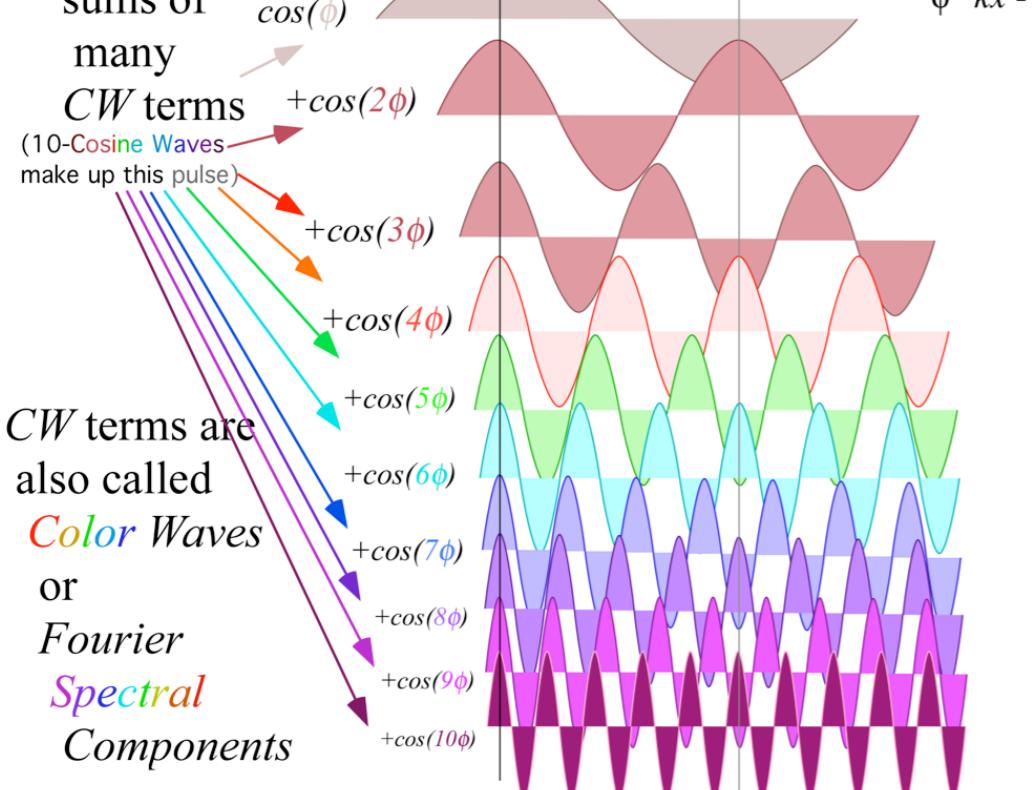


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PW forms are also called *Wave Packets (WP)* since

they are
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CW terms



PW peaks precisely locate places where wave is.
PW widths reduce proportionally with more CW terms (greater *Spectral width*)

Space-time width (pulse width)

$$\Delta t = \tau$$

$$\Delta t = \tau/2$$

$$\Delta t = \tau/5$$

$$\Delta t = \tau/10$$

$$\Delta t = \tau/50$$

this dimension is time

Spectral width (harmonic frequency range)

1 CW term

$$\Delta v = 1v = \text{fundamental frequency}$$

2 CW terms

$$\Delta v = 2v$$

5 CW terms

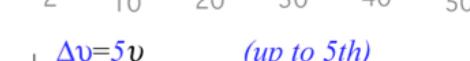
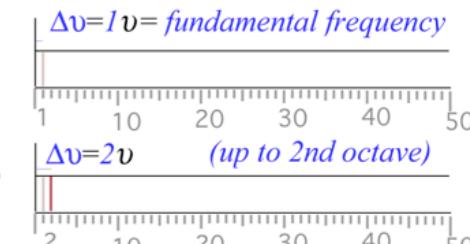
$$\Delta v = 5v$$

10 CW terms

$$\Delta v = 10v$$

50 CW terms

$$\Delta v = 50v$$



Fourier-Heisenberg product: $\Delta t \cdot \Delta v = 1$ (time-frequency uncertainty relation)

- How do you make sense of light-wave axiom(s)?

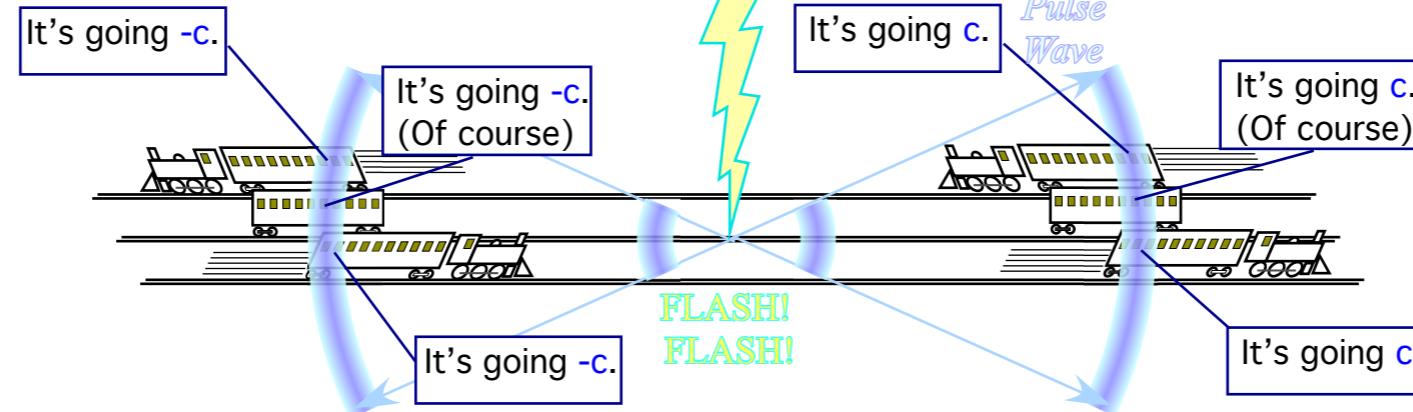
SPEED LIMIT
 $c =$
 299,792,458
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Albert Einstein



1879-1955

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Pulse wave (PW) train

$$A_1 \cos \omega t + A_2 \cos 2\omega t + A_3 \cos 3\omega t + A_4 \cos 4\omega t + \dots$$

Complicated

...many waves and Amplitude parameters

Continuous wave (CW) train

CW zeros precisely locate places where wave is not.

$$A \cos \omega t$$

...just one wave (a 1CW)

Simpler

William of Ockham

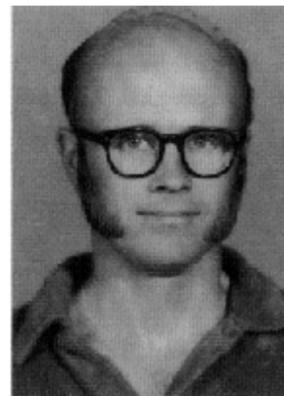


1285-1349

*Using
Occam's
Razor*

(and Evenson's lasers)

Kenneth Evenson



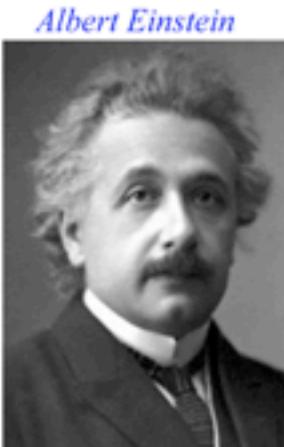
1929-2002

$c = 299,792,458 \text{ m/s}$

Cut a PW to just *one* Continuous Wave ($1CW$)

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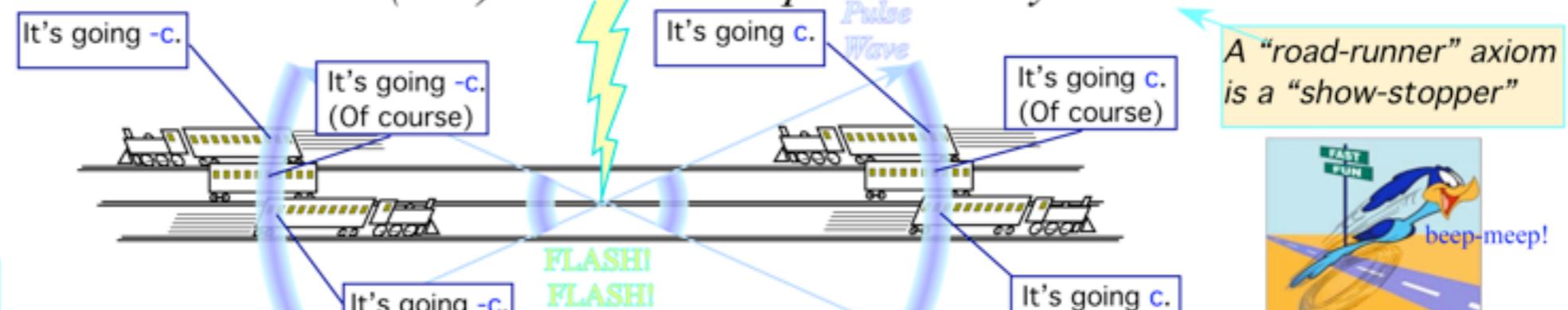
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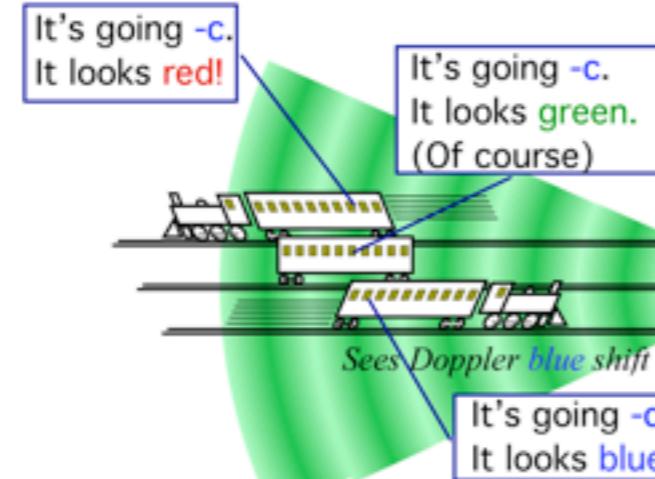
...just one wave (a 1CW)

Evenson Continuous Wave (CW) axiom: CW speed for all colors is c



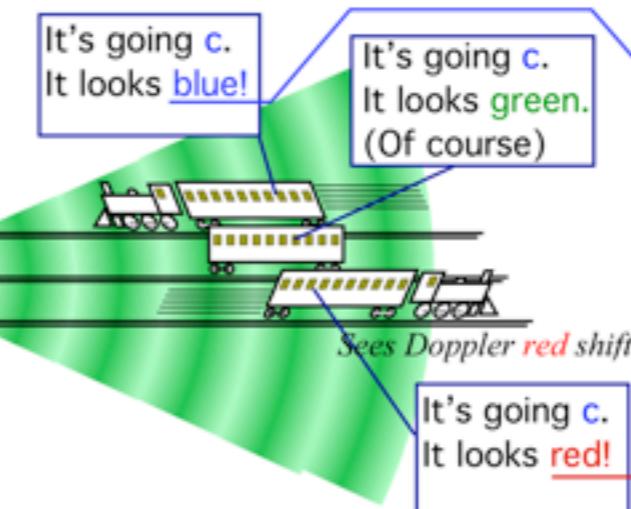
Kenneth Evenson

1929-2002
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600 THz
(green)

Laser
source



It's going c .
It looks blue!

It's going c .
It looks green.
(Of course)

CW affected by
1st order Doppler
Blue shifts $b = e^{+p}$
and
Red shifts $r = e^{-p}$
of frequency ν
and wavenumber κ



Sees Doppler blue shift

It's going c .
It looks blue!



Sees Doppler red shift

It's going c .
It looks red!

Cut a PW to one Continuous Wave (1CW) that changes Color if you accelerate!

- How do you make sense of light-wave axiom(s)?

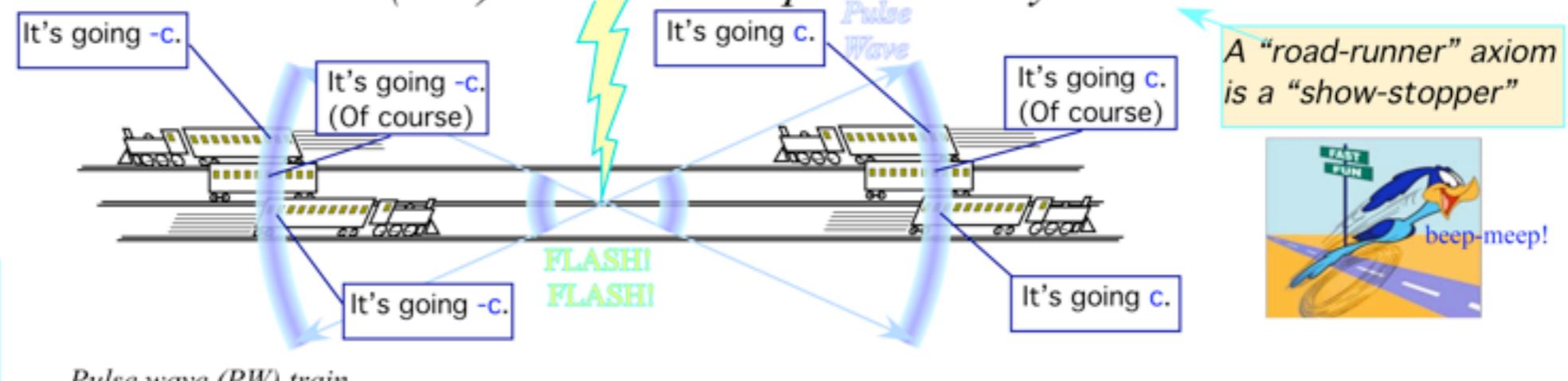
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Albert Einstein

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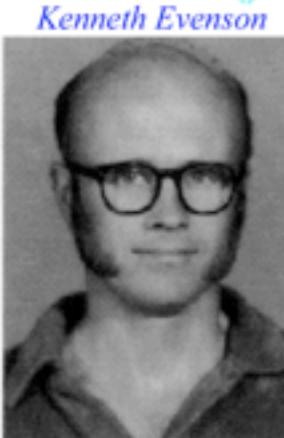
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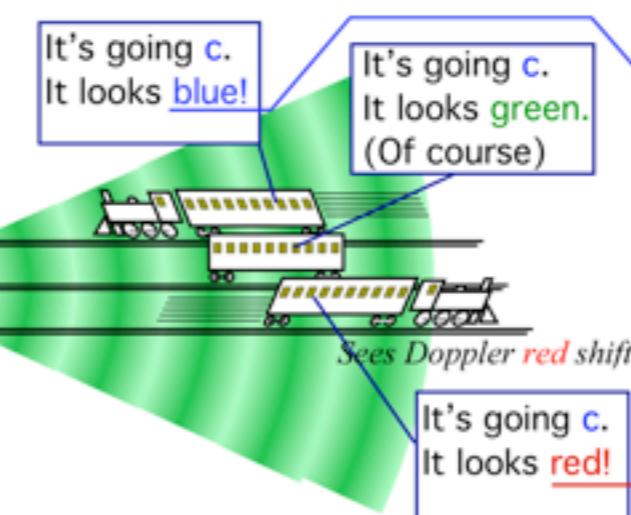
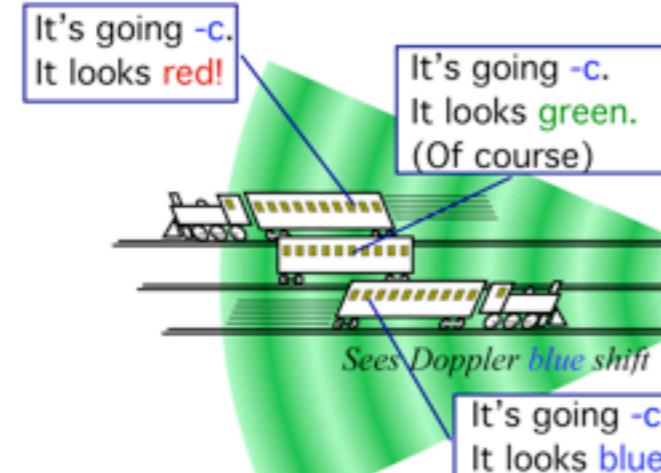
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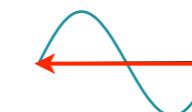
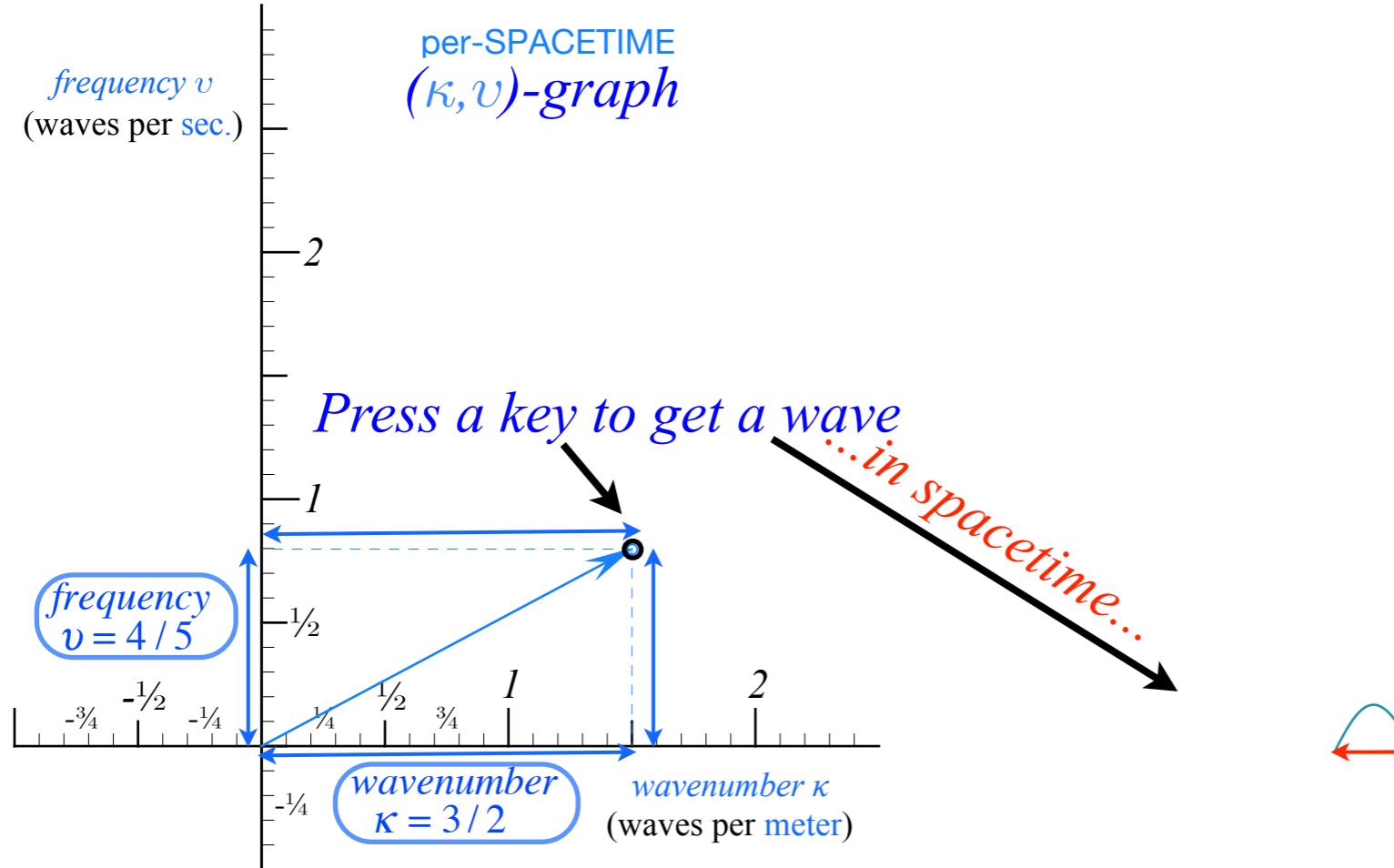
1929-2002
 $c = 299,792,458 \text{ m/s}$



CW affected by 1st order Doppler Blue shifts $b = e^{+p}$ and Red shifts $r = e^{-p}$ of frequency ν and wavenumber κ

Cut a PW to one Continuous Wave (1CW) that changes Color if you accelerate!
 CW also stands for “Cosine Wave” or “Coherent Wave” or “Colored Wave” (all helpful things!)

The “Keyboard of the gods” or per-space-per-time graphs versus space-time graphs



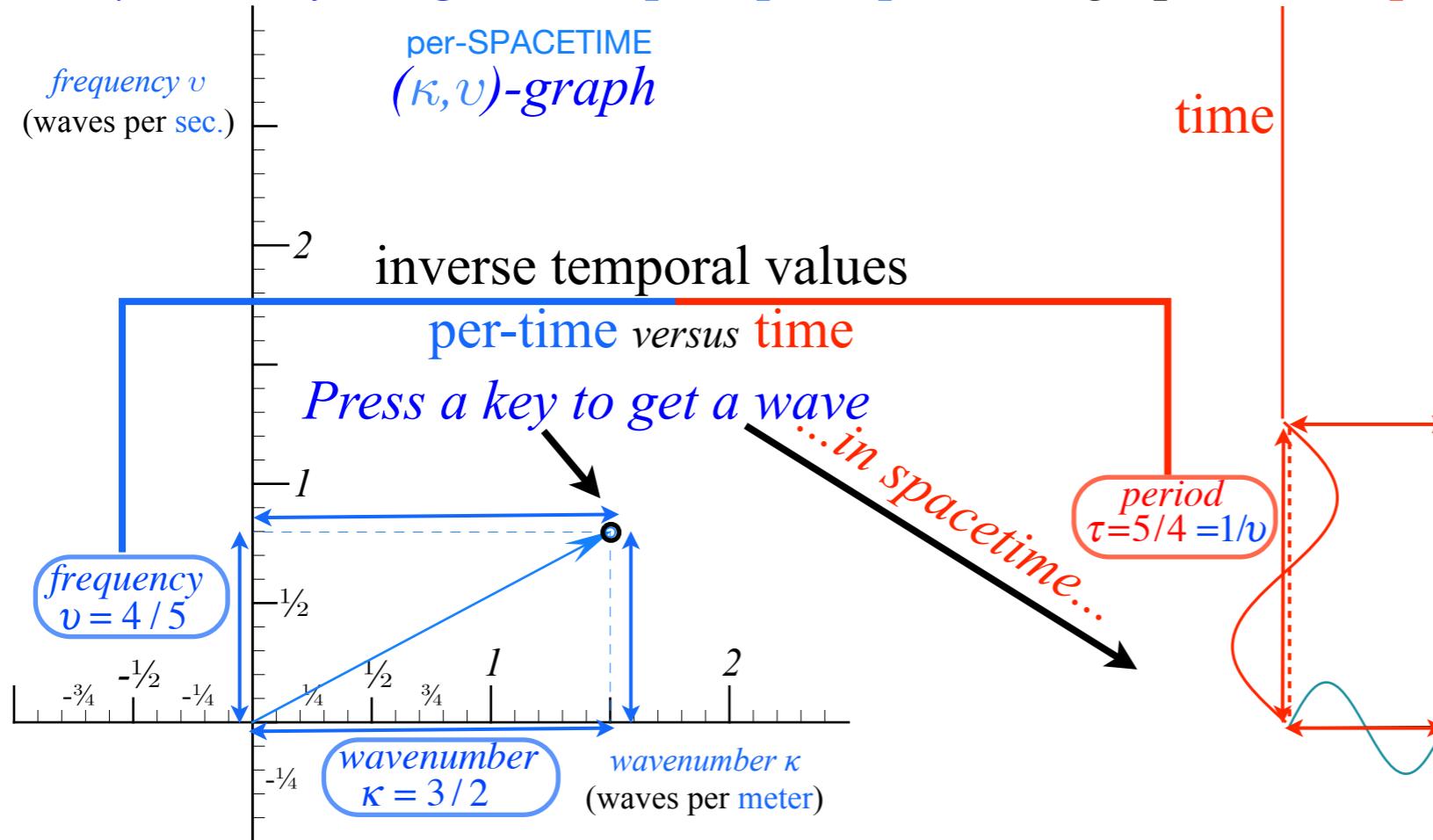
“Keyboard of the gods” is known as “Fourier-space”



Jean-Baptiste
Joseph Fourier
1768-1830

Ways to quantify general waves

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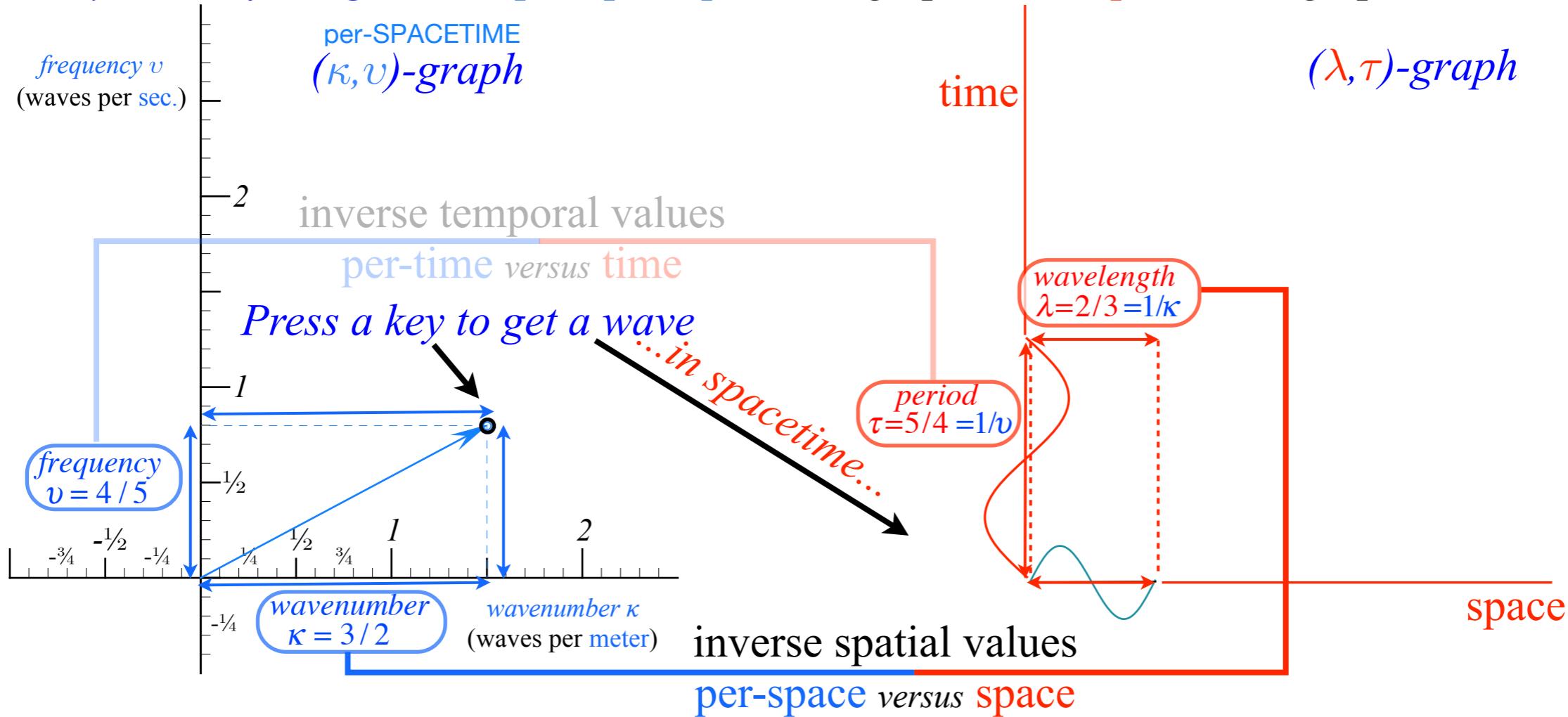
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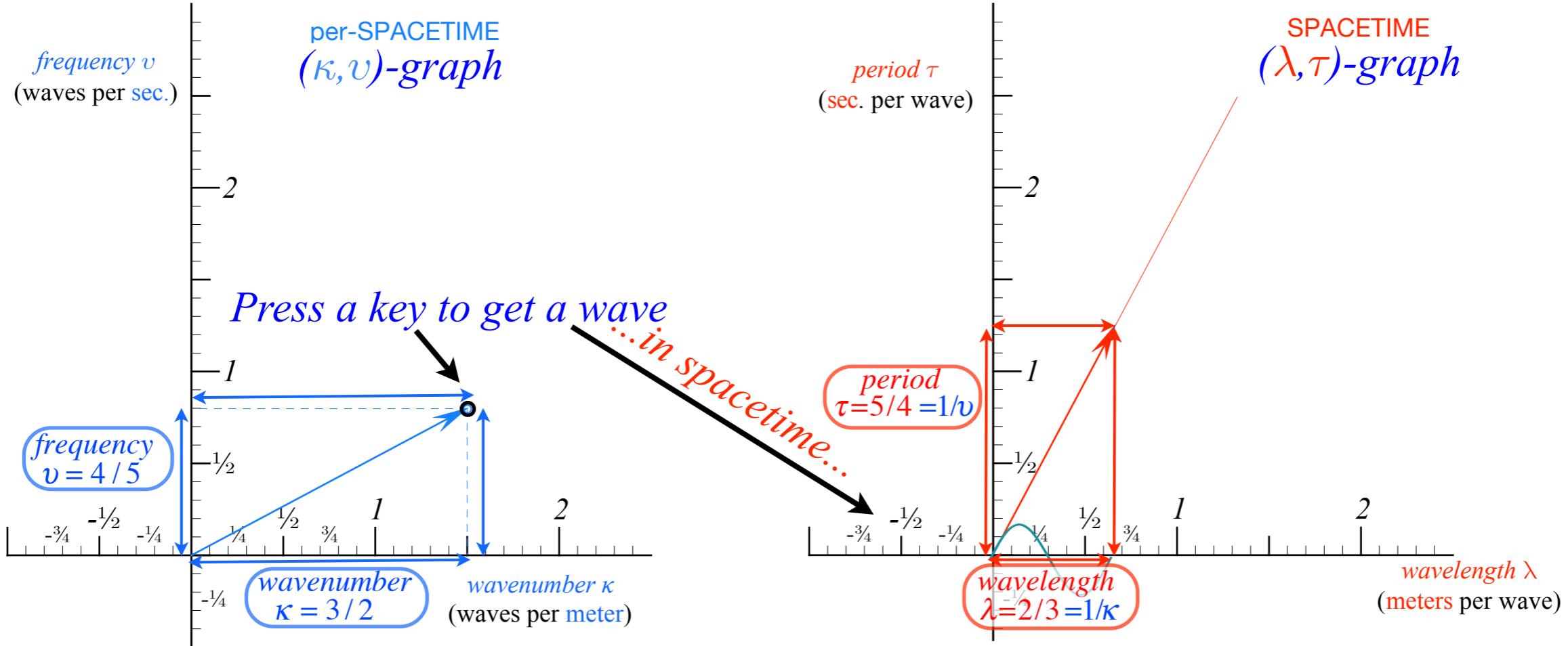
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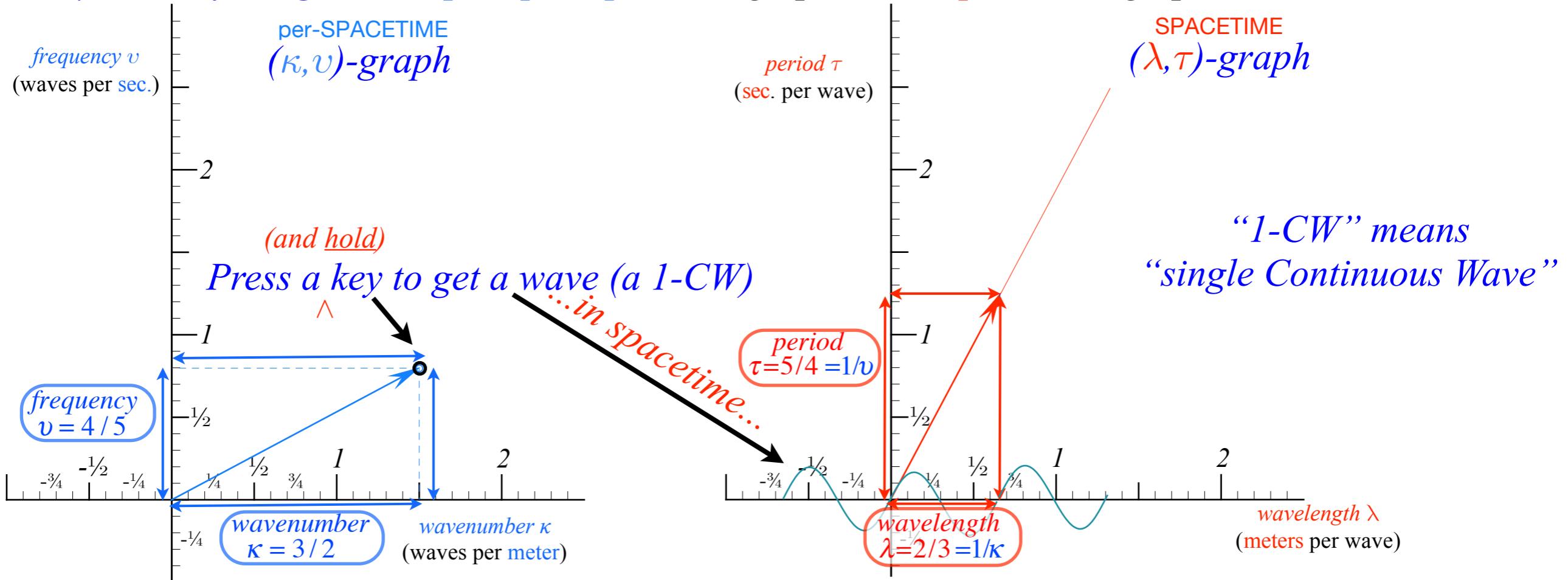
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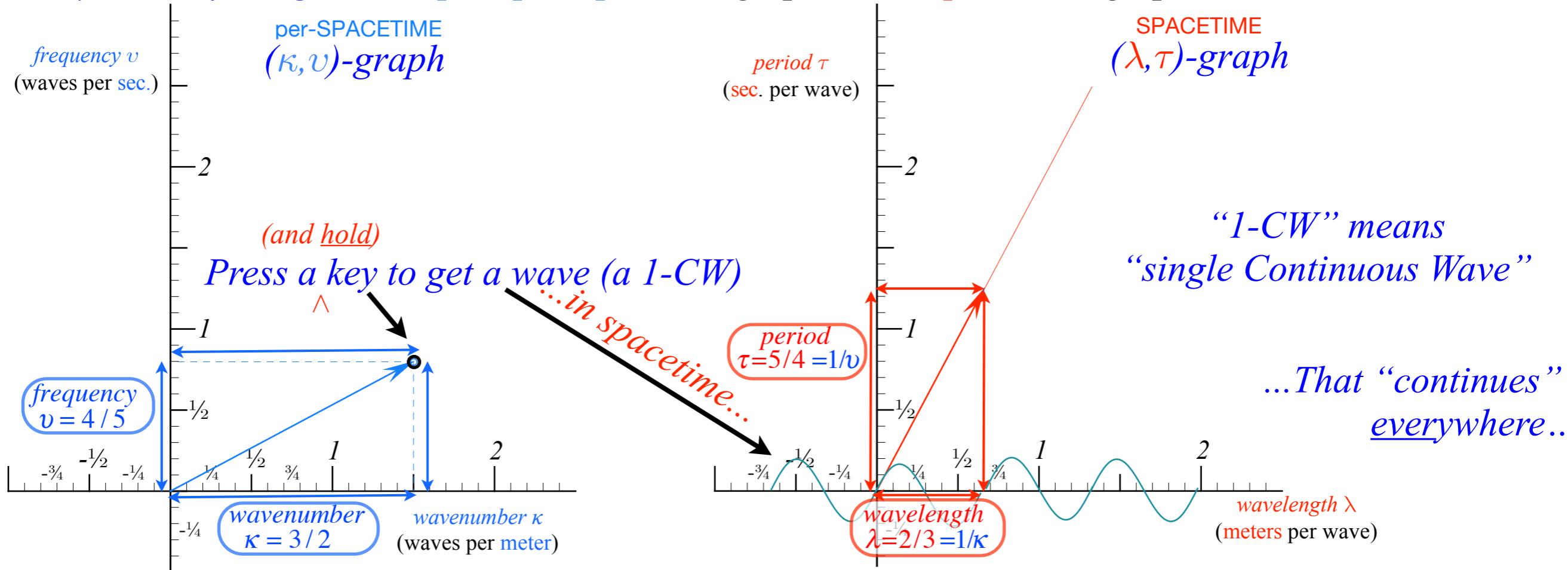
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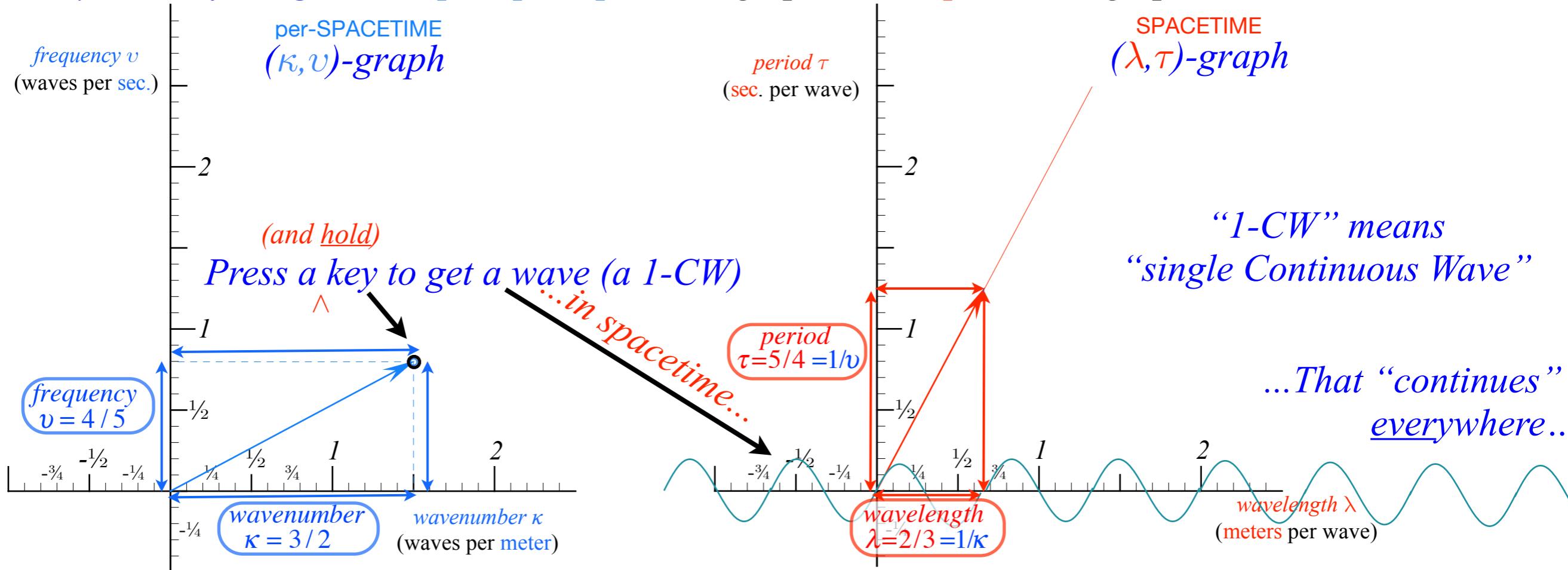
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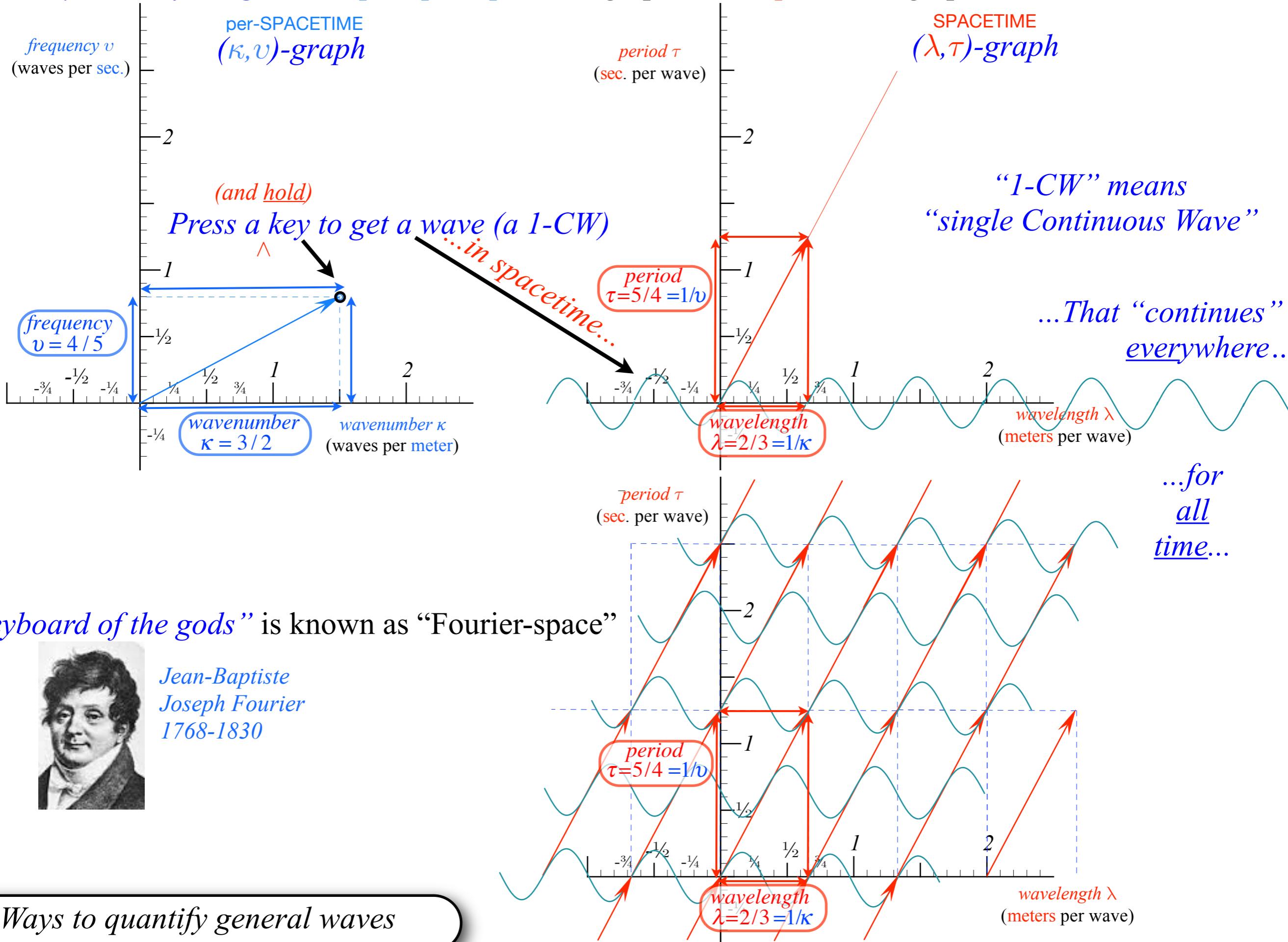
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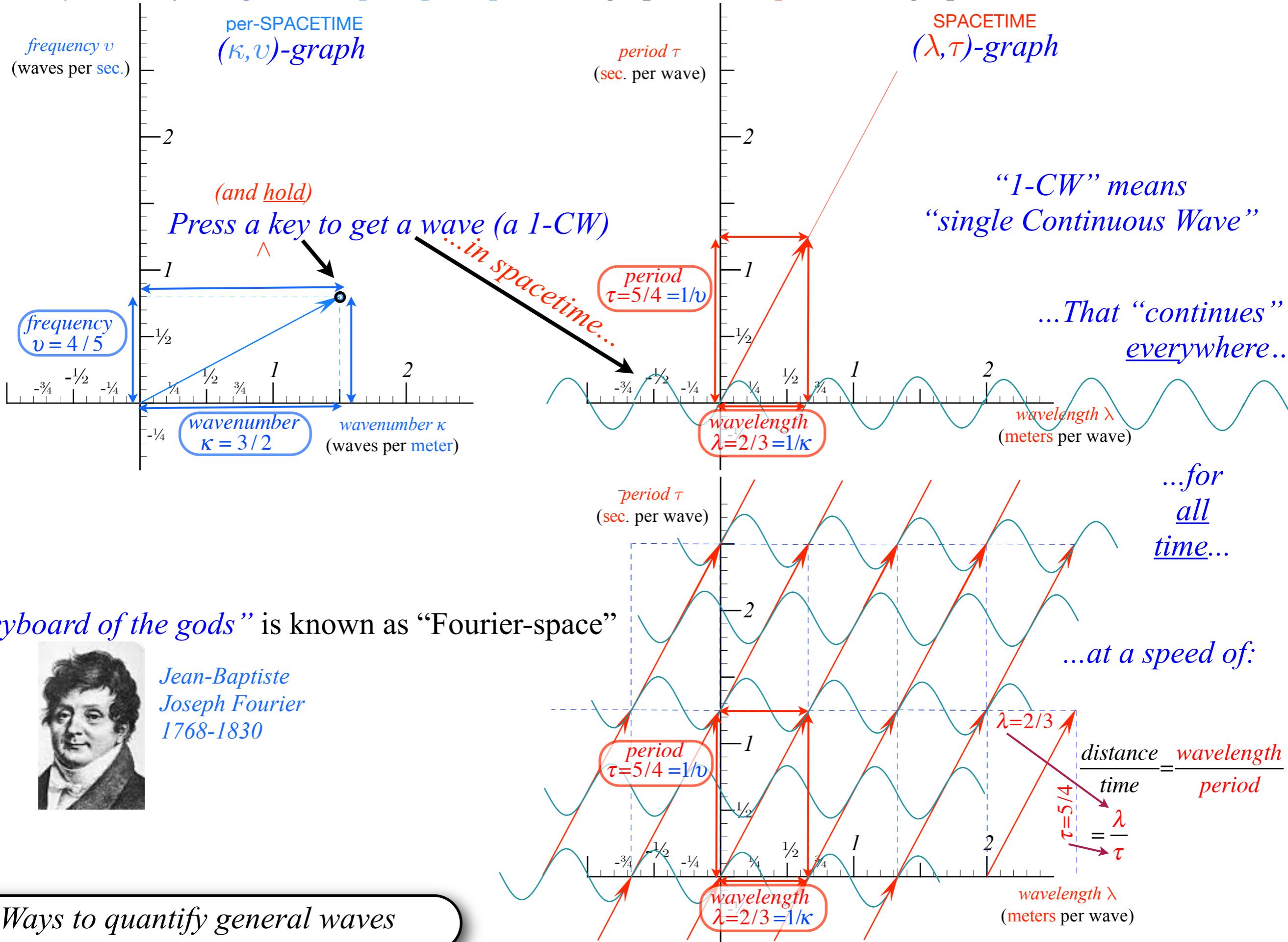
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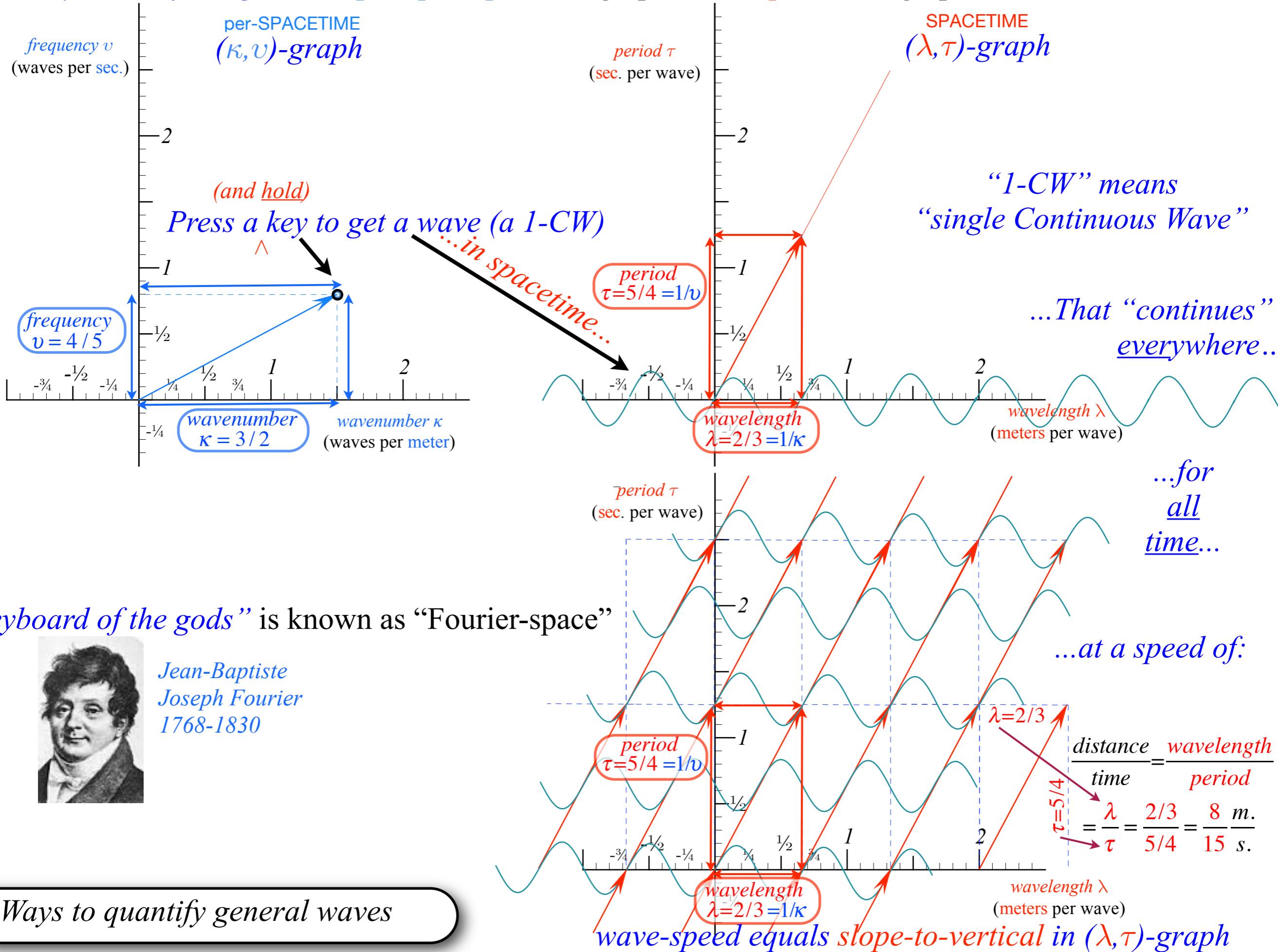
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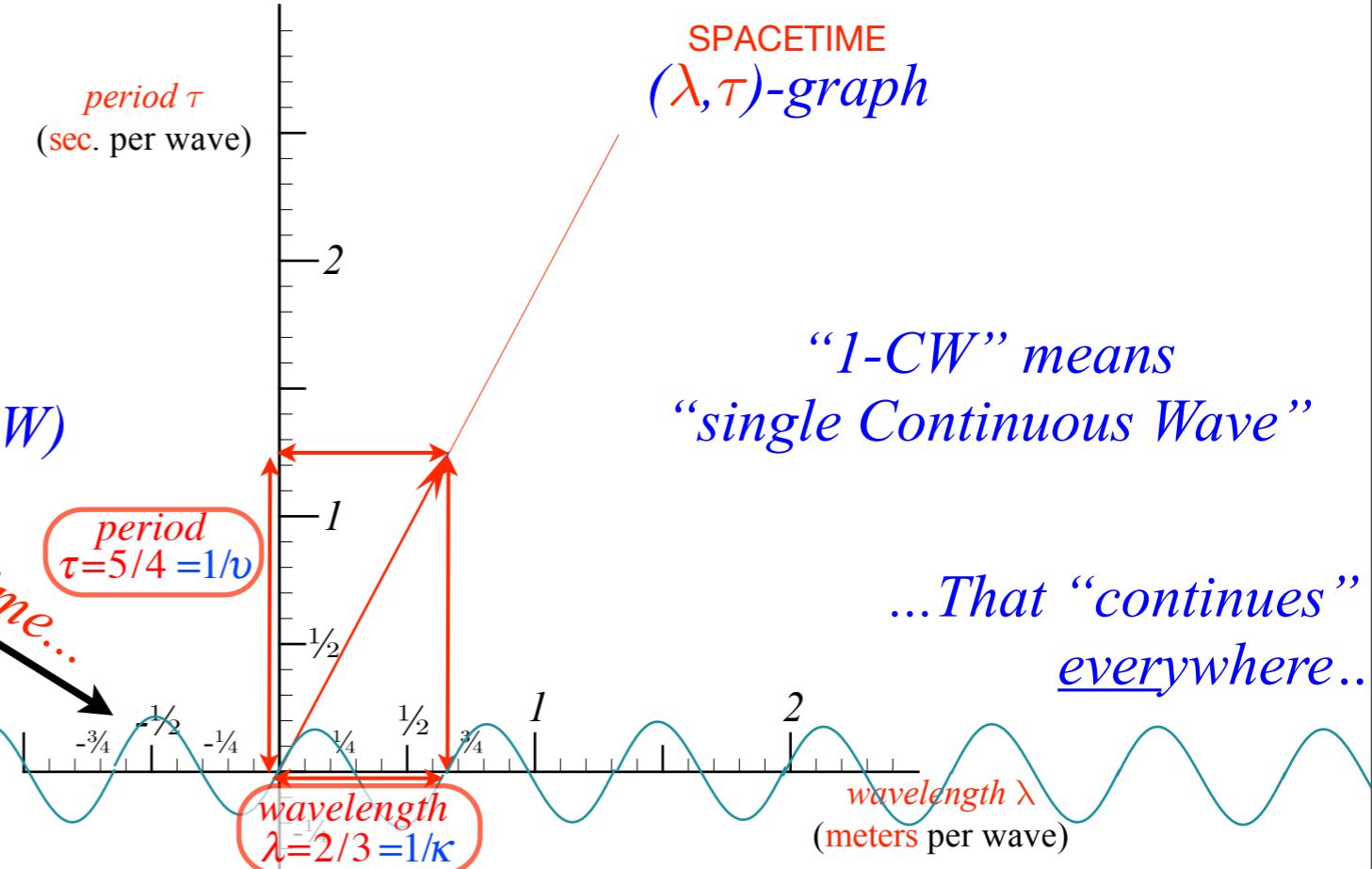
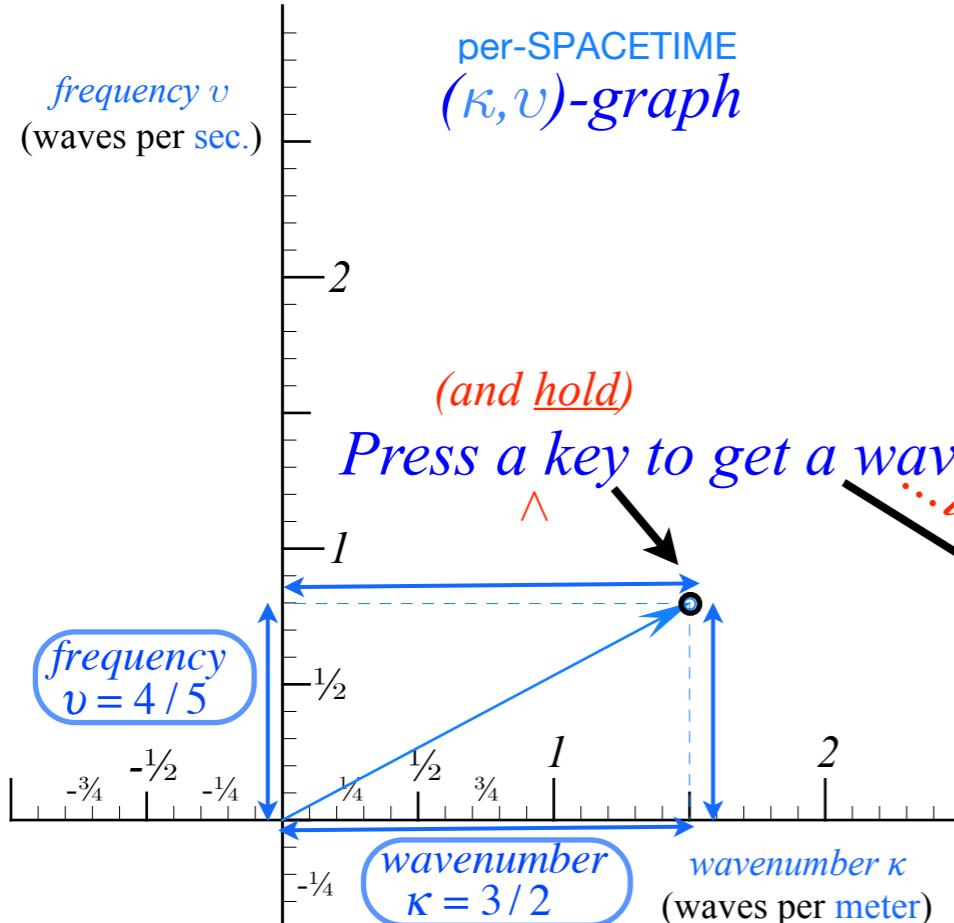
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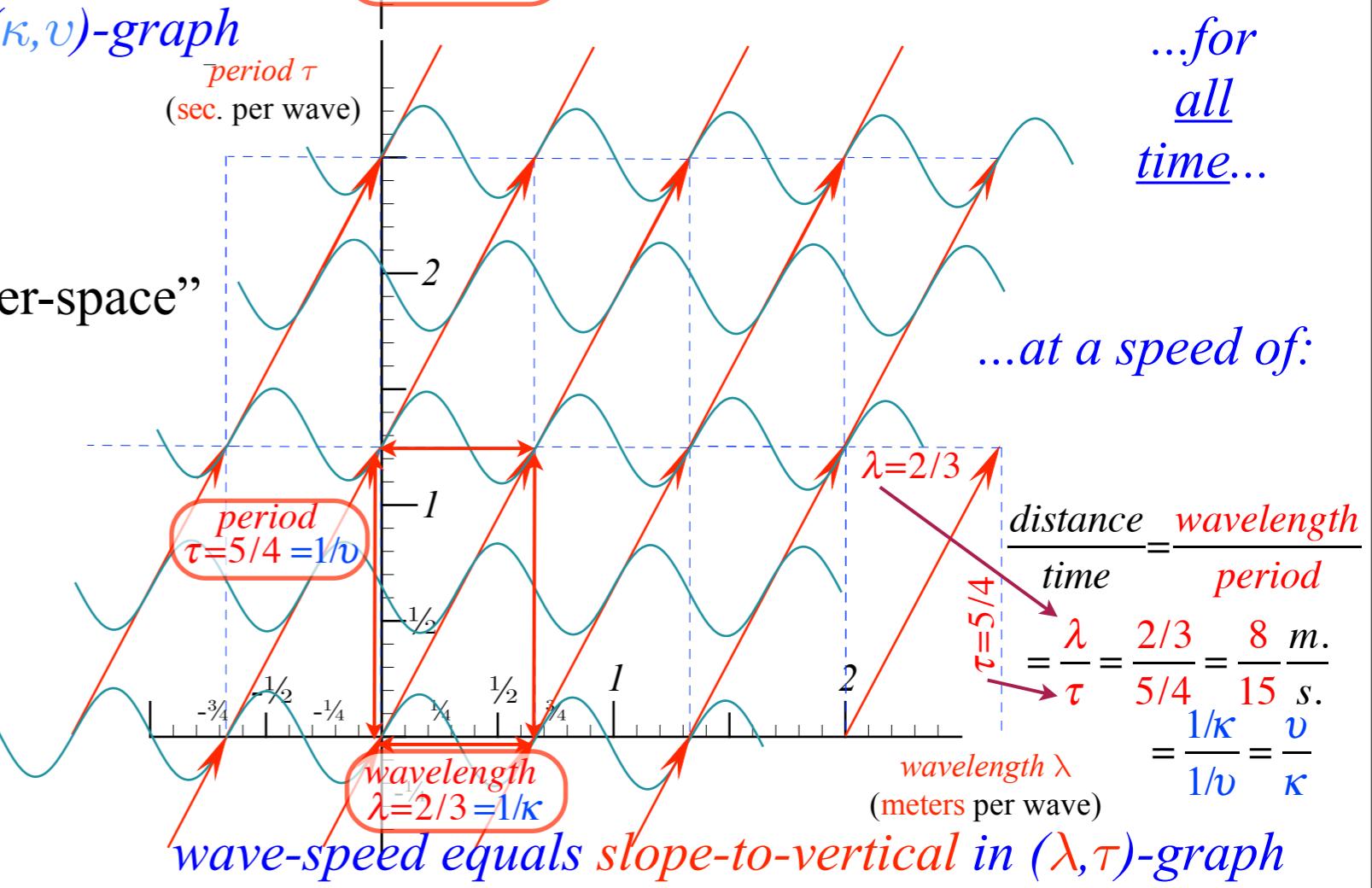
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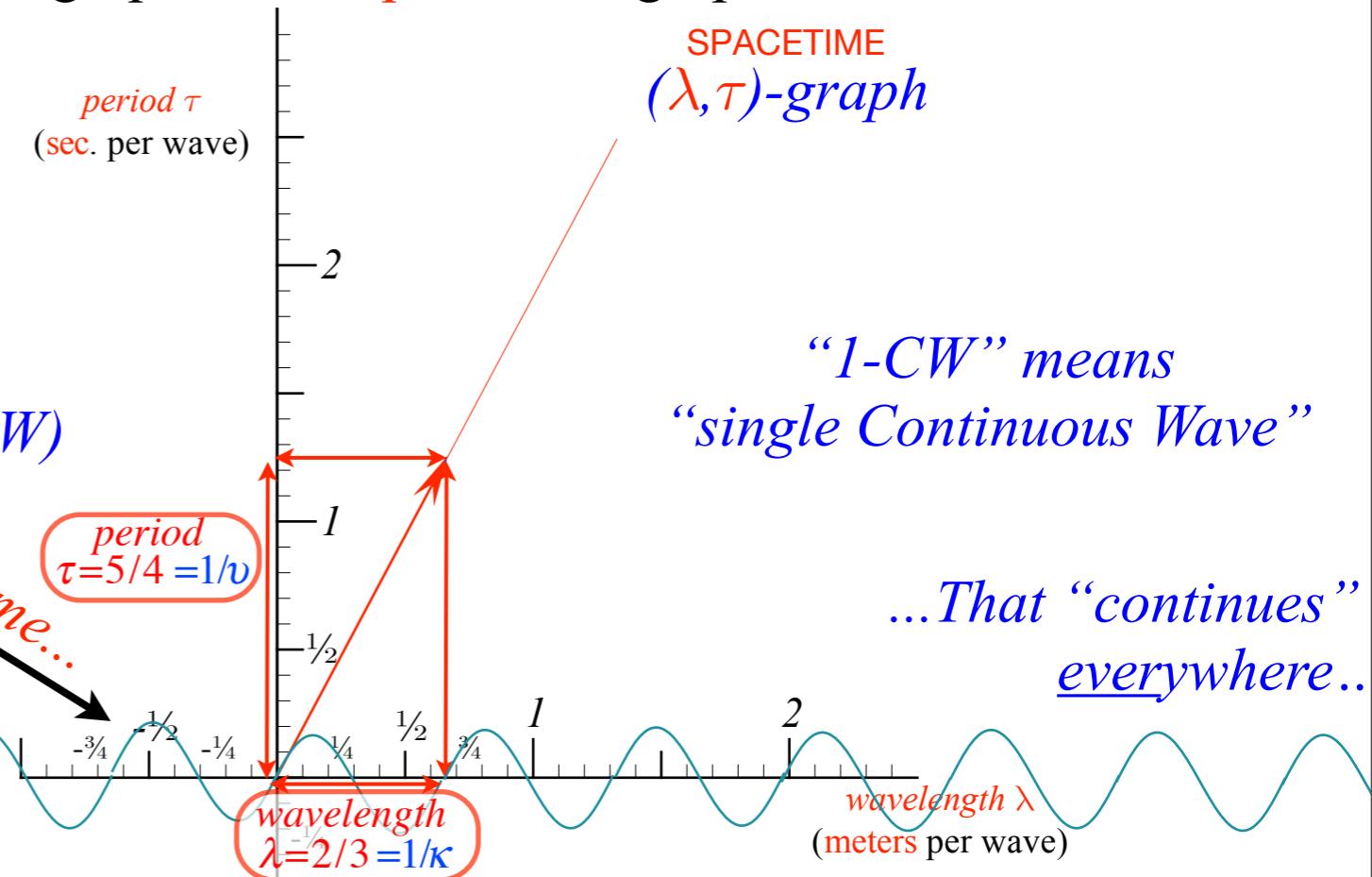
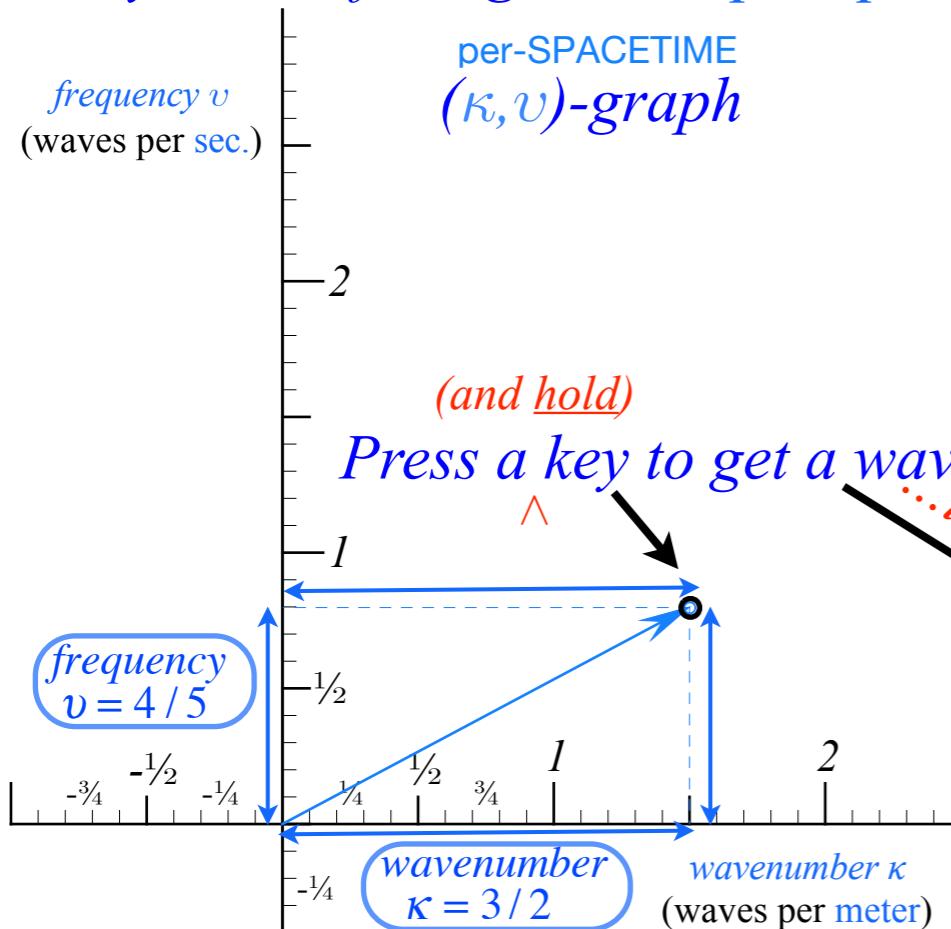
wave-speed equals slope-to-horizontal in (κ, v) -graph



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Ways to quantify general waves

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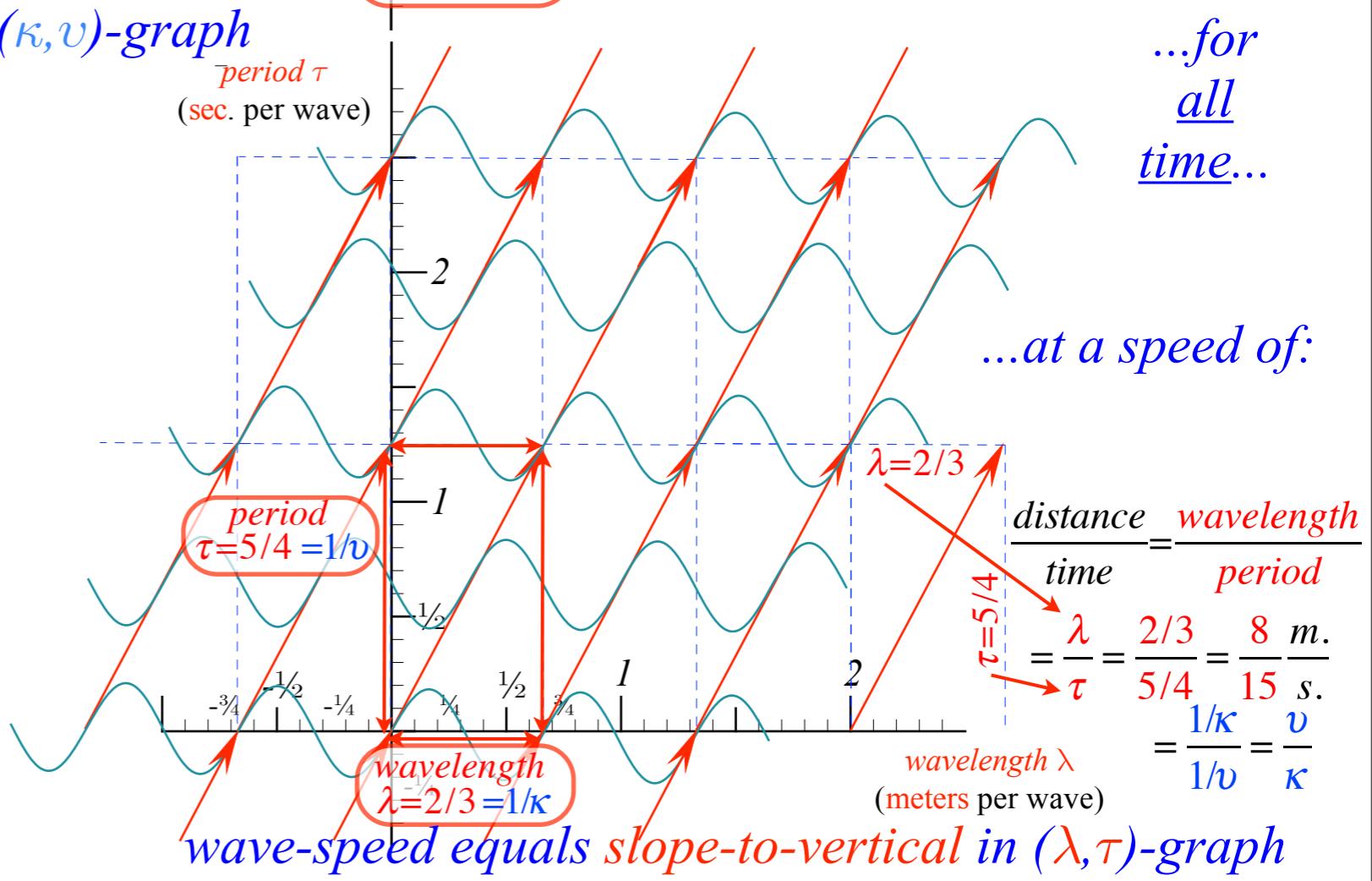
wave-velocity formula

$$\frac{\text{distance}}{\text{time}} = \frac{\text{wavelength}}{\text{period}} = \frac{\text{frequency}}{\text{wavenumber}}$$

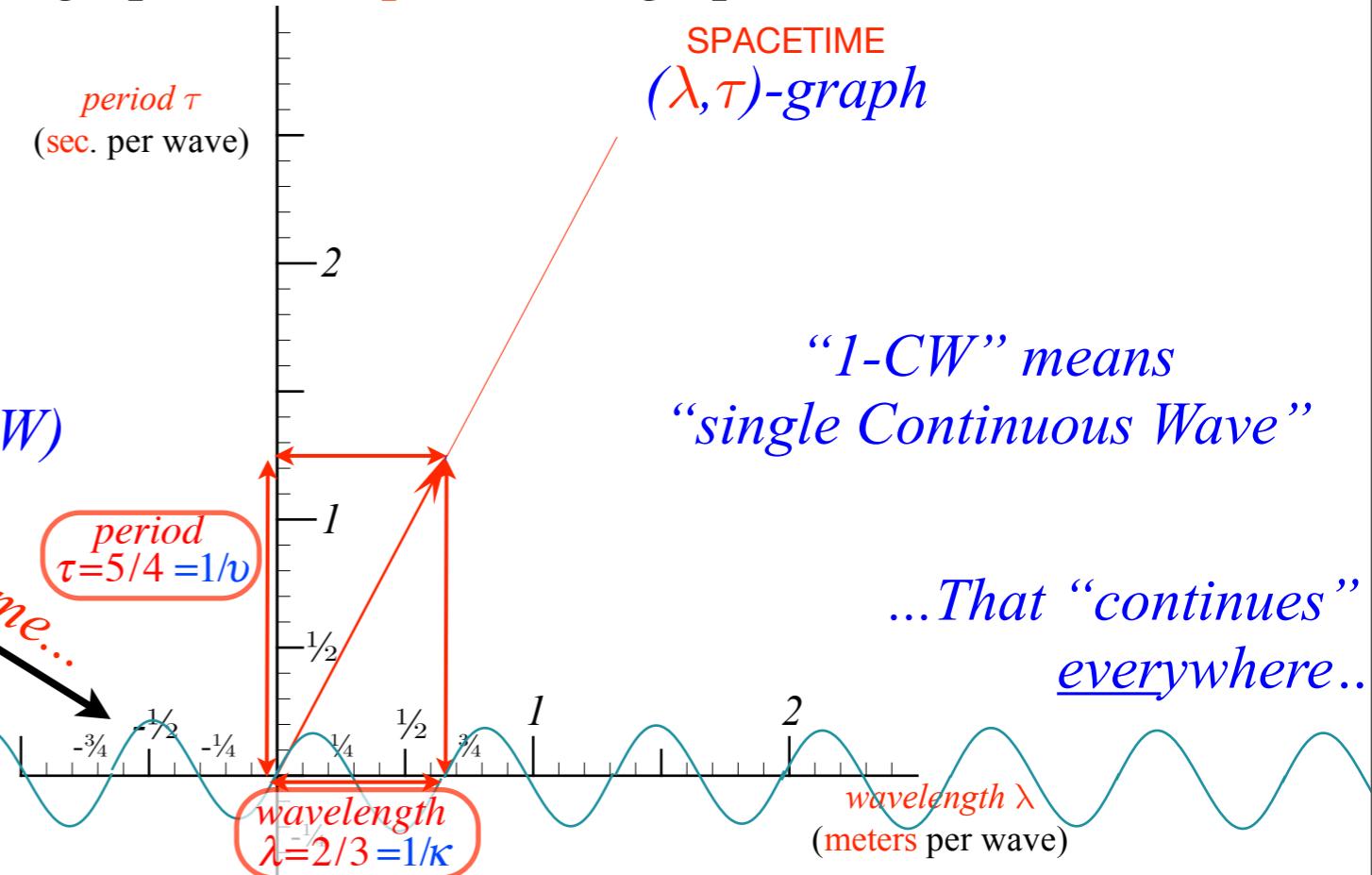
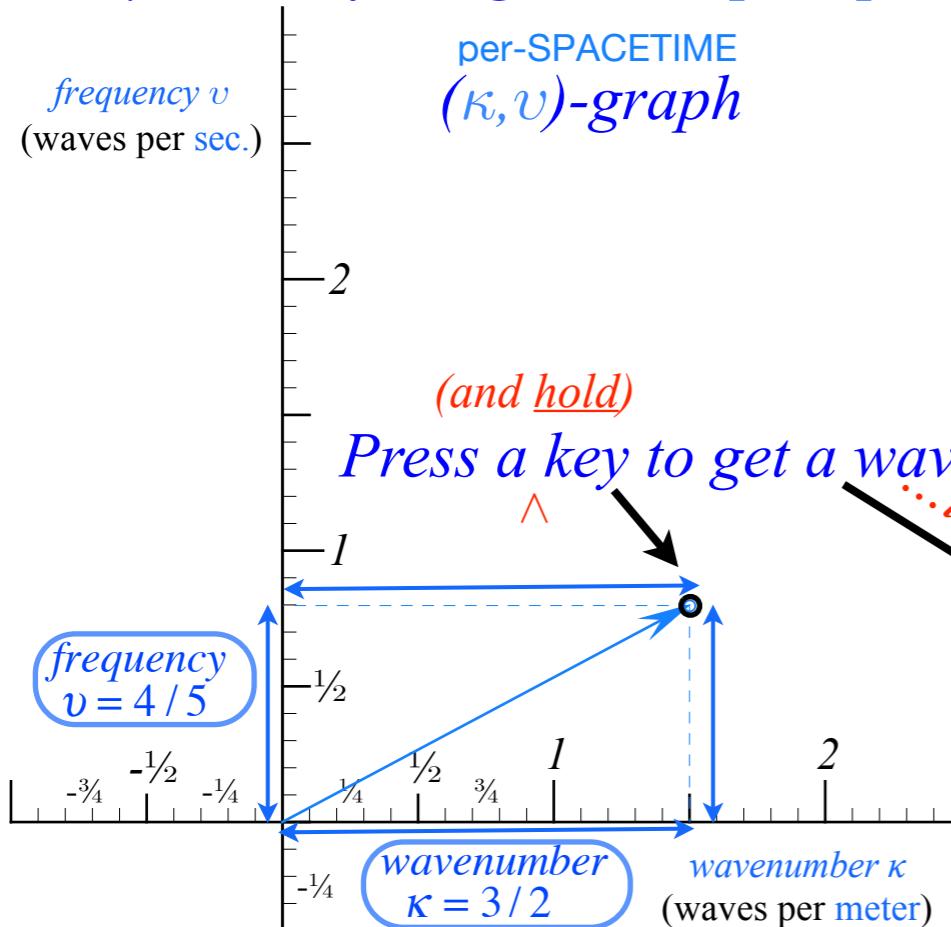
$$V_{\text{wave}} = \frac{\lambda}{\tau} = \frac{1/\kappa}{1/v} = \frac{v}{\kappa} = \frac{1/\tau}{1/\lambda}$$

$$= \frac{2/3}{5/4} = \frac{4/5}{3/2} = \frac{8}{15} \text{ m. s.}$$

Ways to quantify general waves



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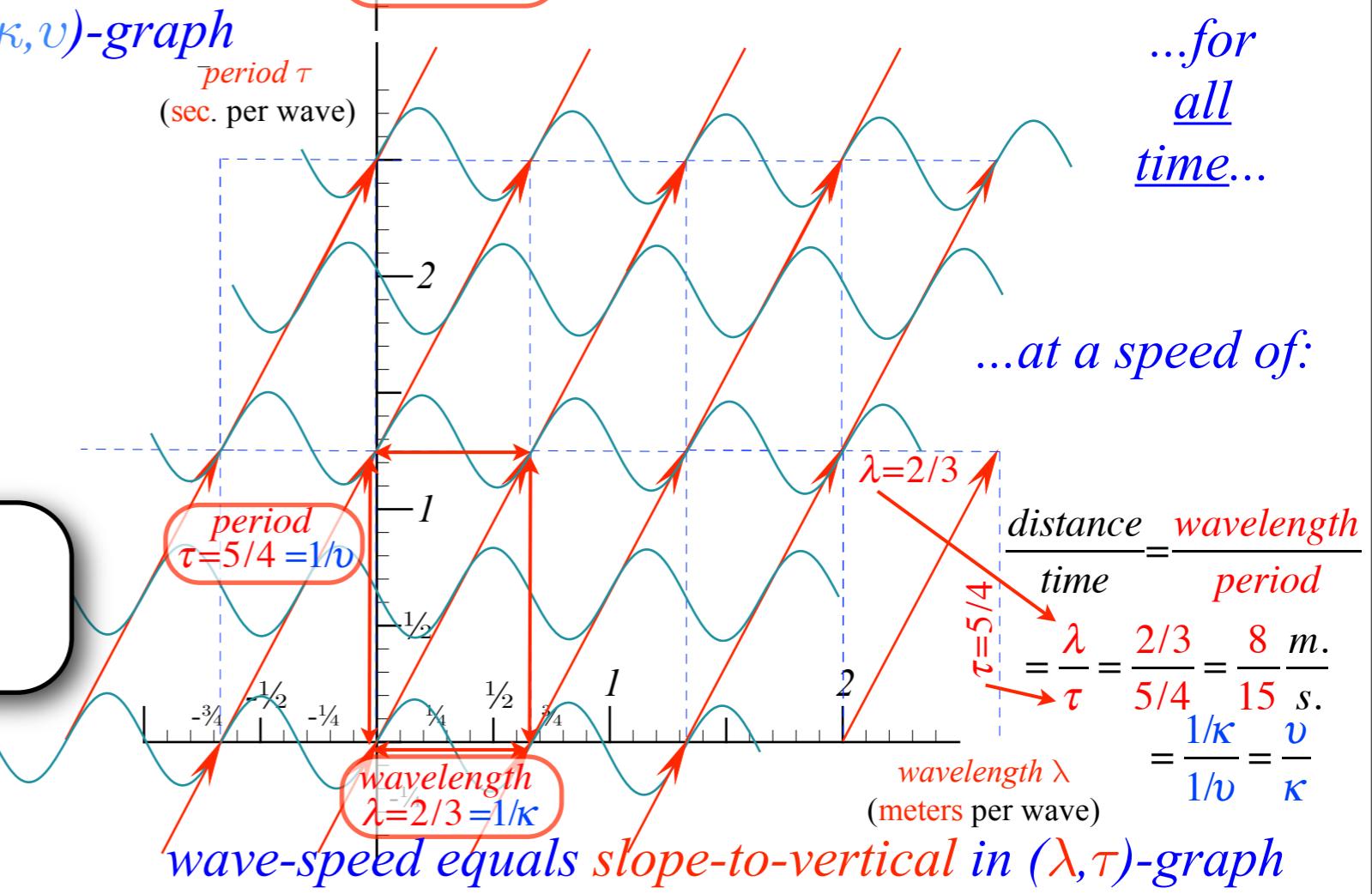
$$V_{\text{wave}} = \frac{\lambda}{\tau} = \frac{1/\kappa}{1/v} = \frac{v}{\kappa} = \frac{1/\tau}{1/\lambda}$$

$$= \frac{2/3}{5/4} = \frac{4/5}{3/2} = \frac{8}{15} \text{ m. s.}$$

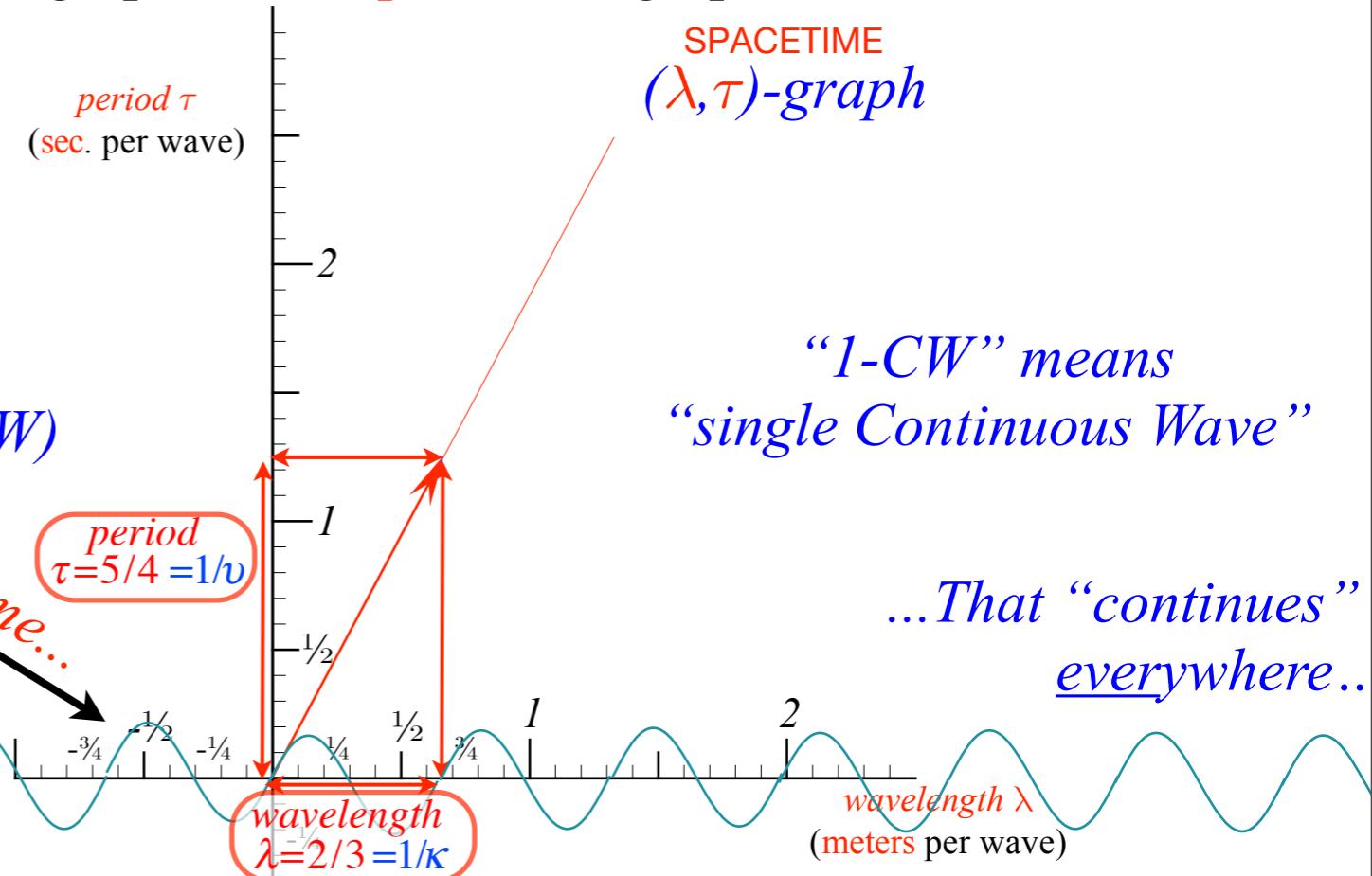
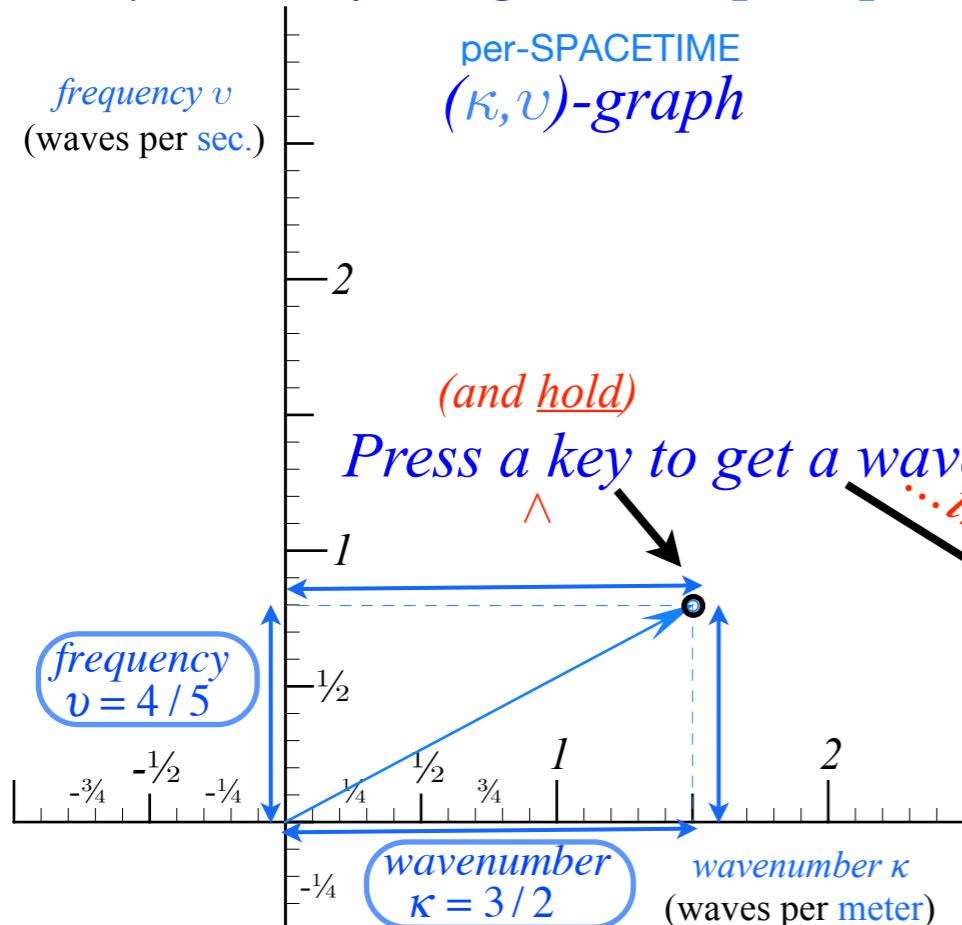
Light wave-velocity c (our main topic)

$$V_{\text{light}} = c = \frac{\lambda}{\tau} = \frac{1/\kappa}{1/v} = \frac{v}{\kappa} = \frac{1/\tau}{1/\lambda} = 299,792,458 \frac{\text{m.}}{\text{s.}}$$

(Next up:) Ways to quantify **light** waves



The “Keyboard of the gods” or per-space-per-time graphs versus space-time graphs



wave-speed equals slope-to-horizontal in (κ, v) -graph

wave-velocity formula

$$\frac{\text{distance}}{\text{time}} = \frac{\text{wavelength}}{\text{period}} = \frac{\text{frequency}}{\text{wavenumber}}$$

$$V_{\text{wave}} = \frac{\lambda}{\tau} = \frac{1/\kappa}{1/v} = \frac{v}{\kappa} = \frac{1/\tau}{1/\lambda}$$

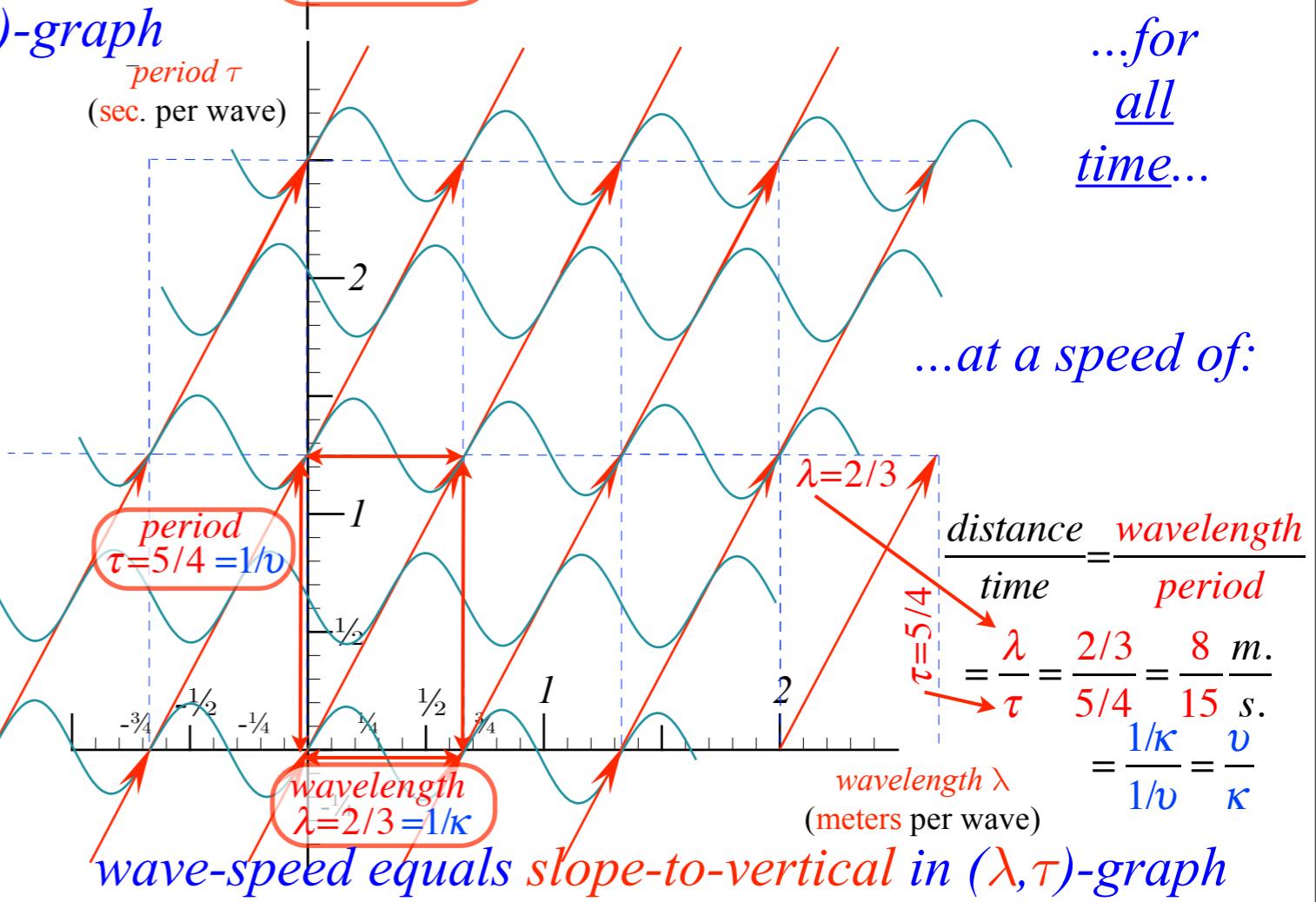
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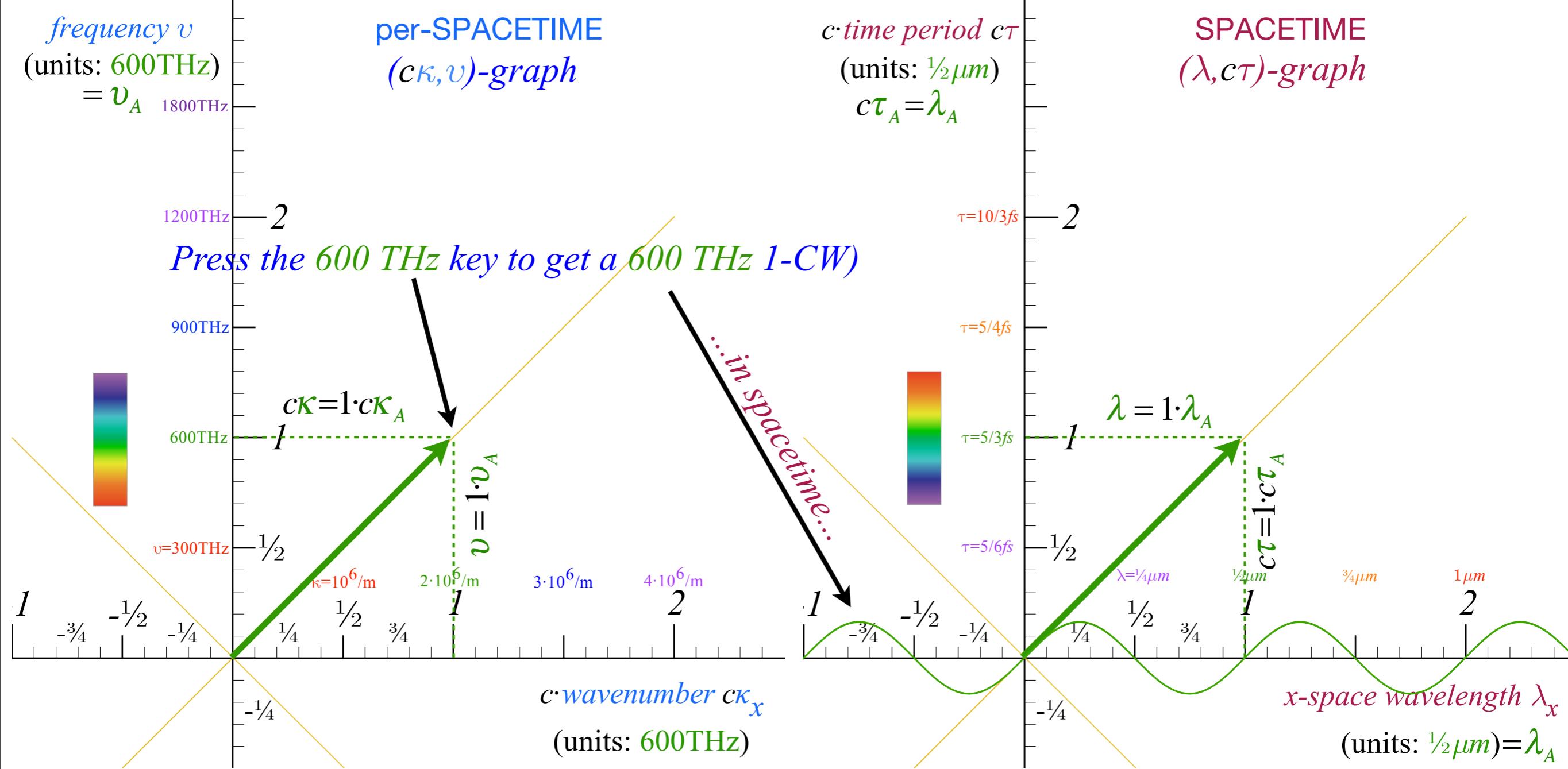
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Dimensionless **Light** wave-velocity $c/c = 1$

$$\frac{V_{\text{light}}}{c} = \frac{\lambda}{c\tau} = \frac{1/\kappa}{c/v} = \frac{v}{c\kappa} = \frac{1/\tau}{c/\lambda} = 1$$

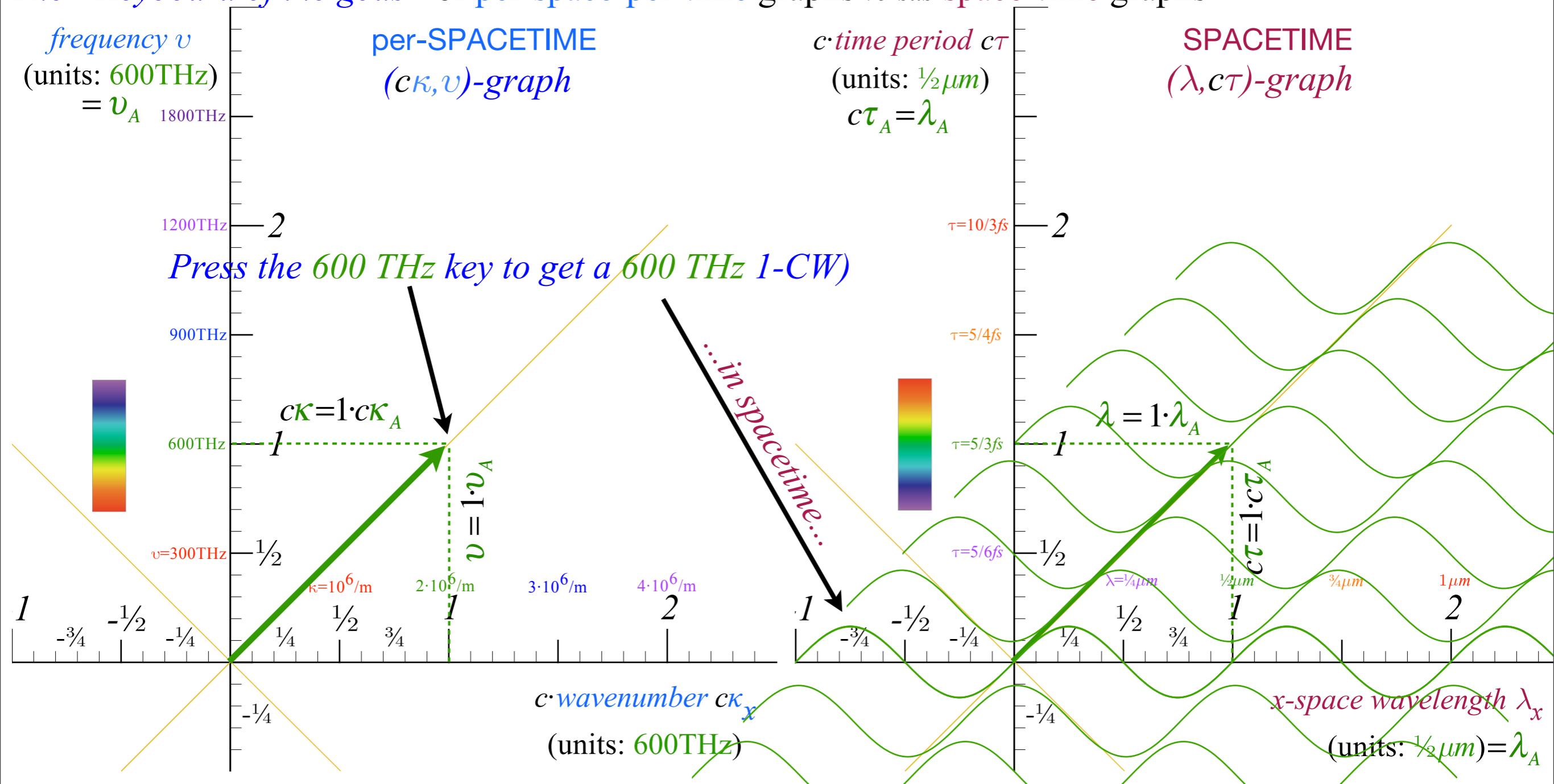


The “Keyboard of the gods” or per-space-per-time graphs versus space-time graphs



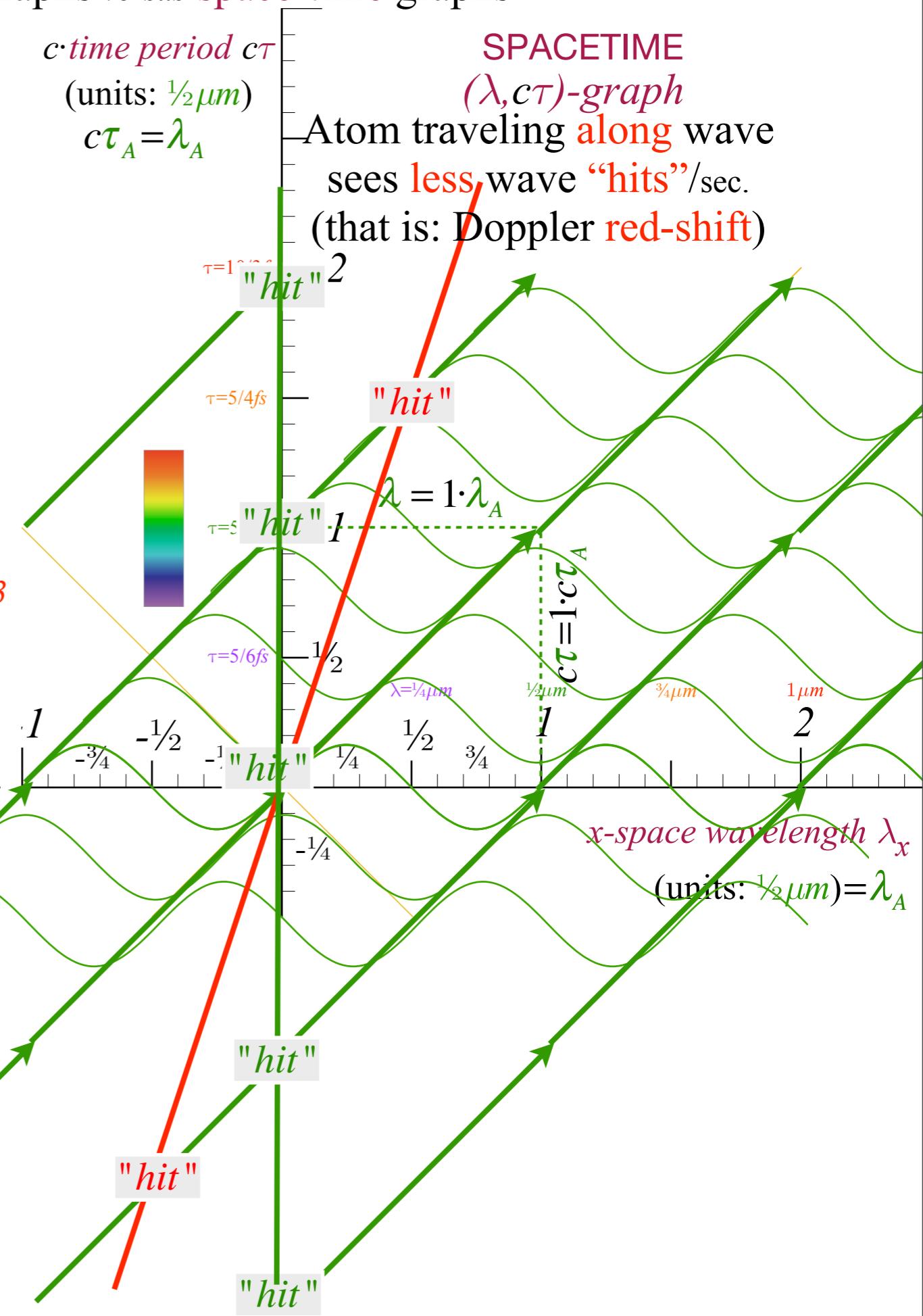
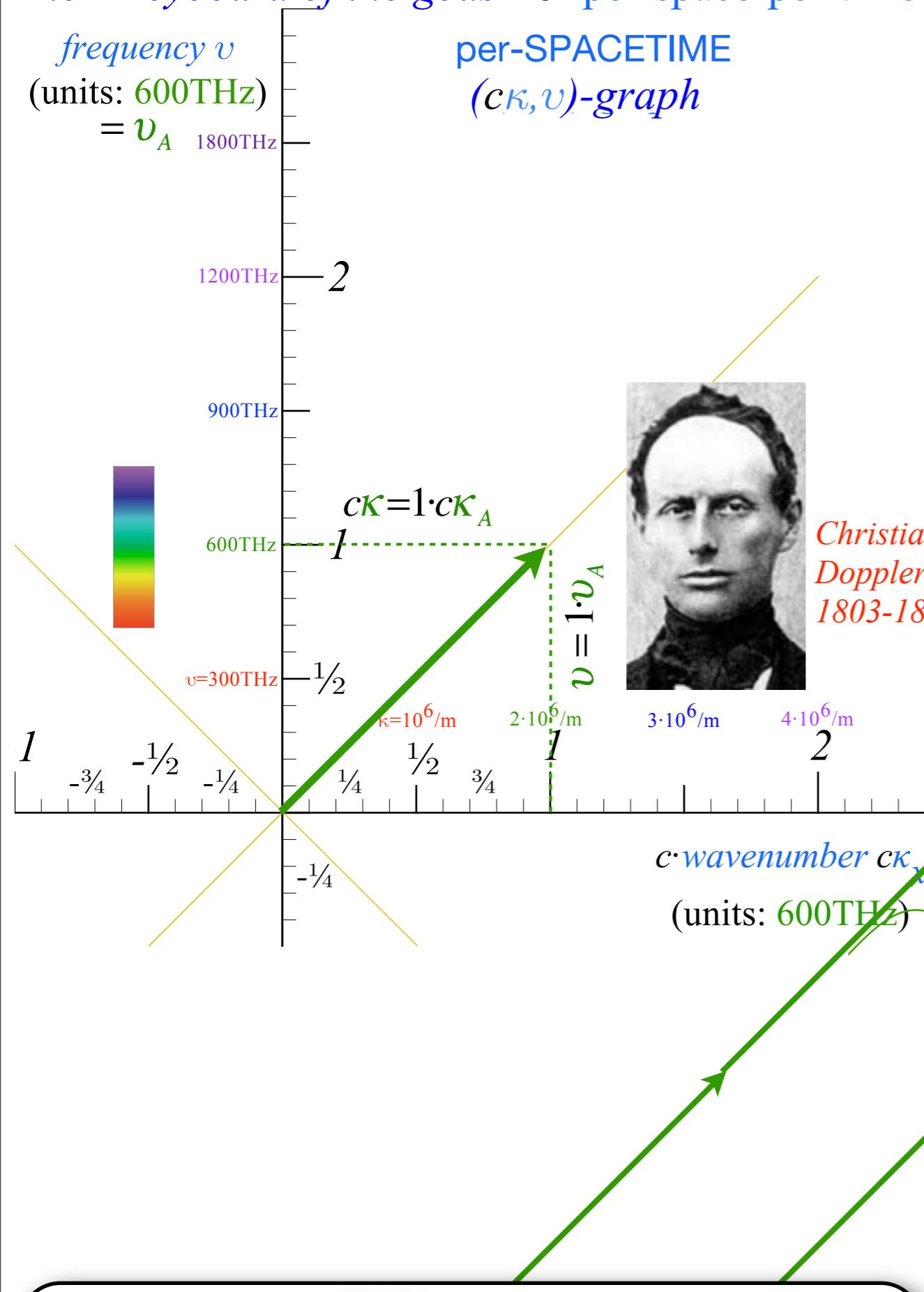
Ways to quantify **light** waves (600 THz example)

The “Keyboard of the gods” or per-space-per-time graphs versus space-time graphs



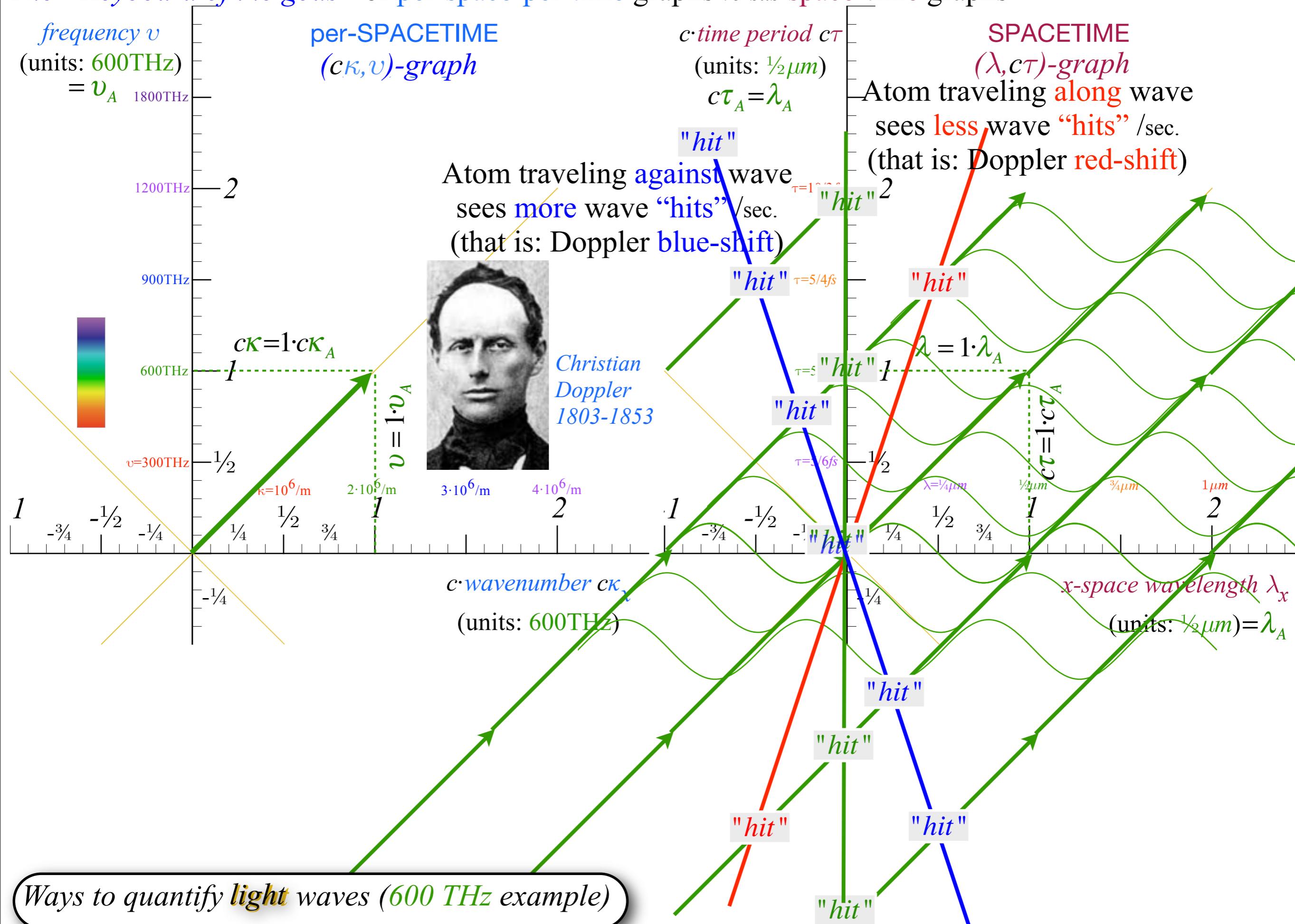
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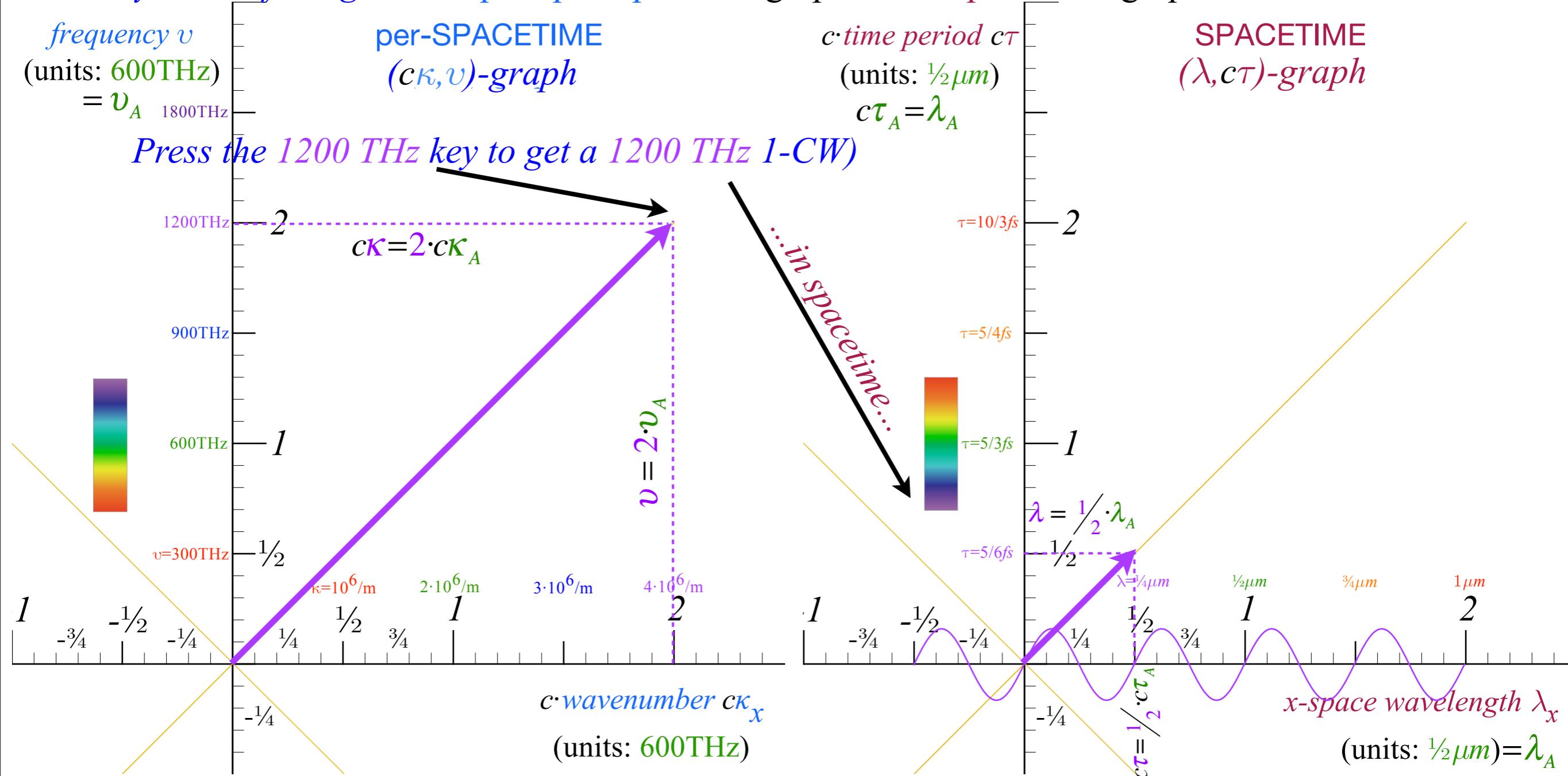


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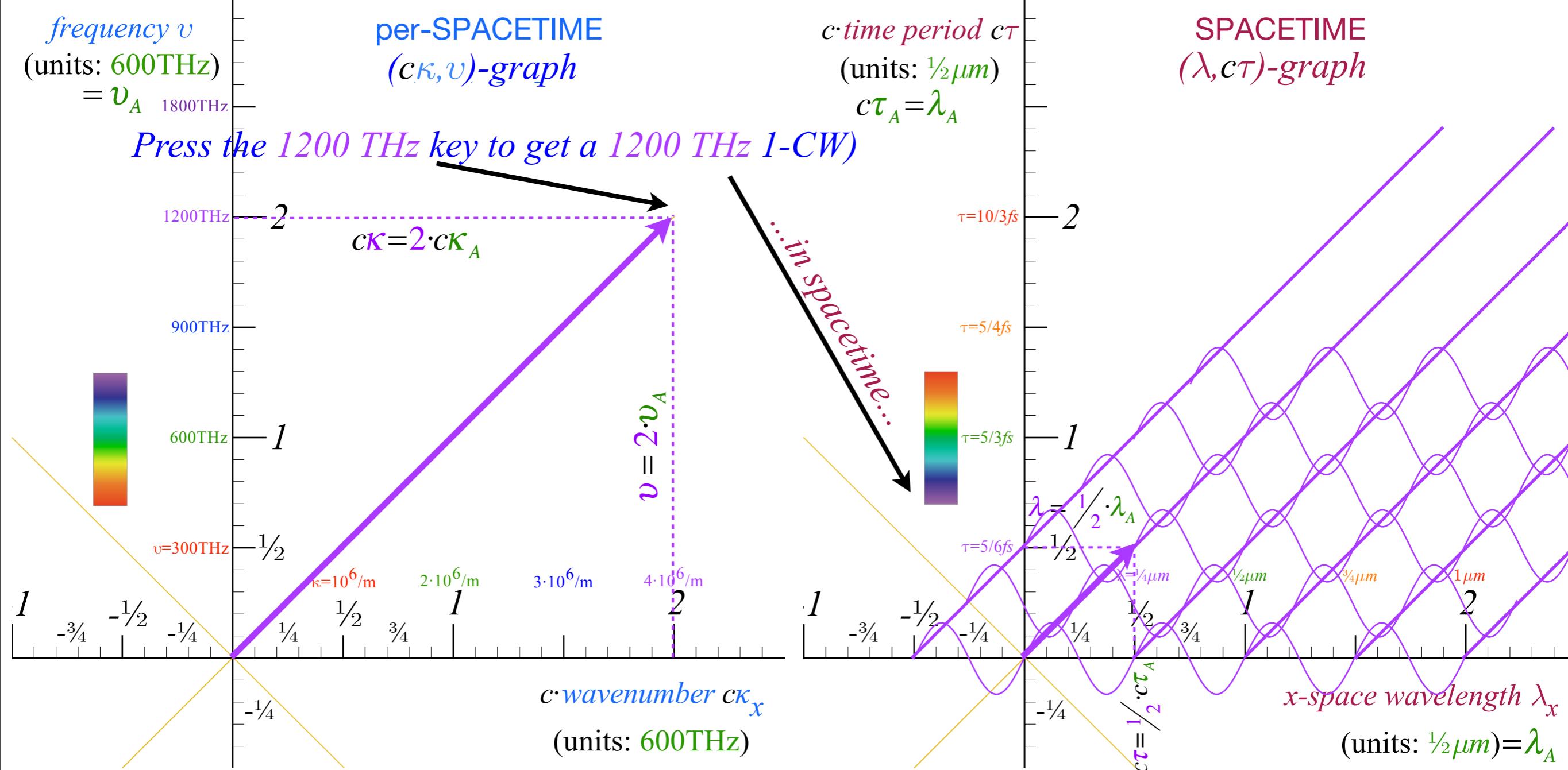


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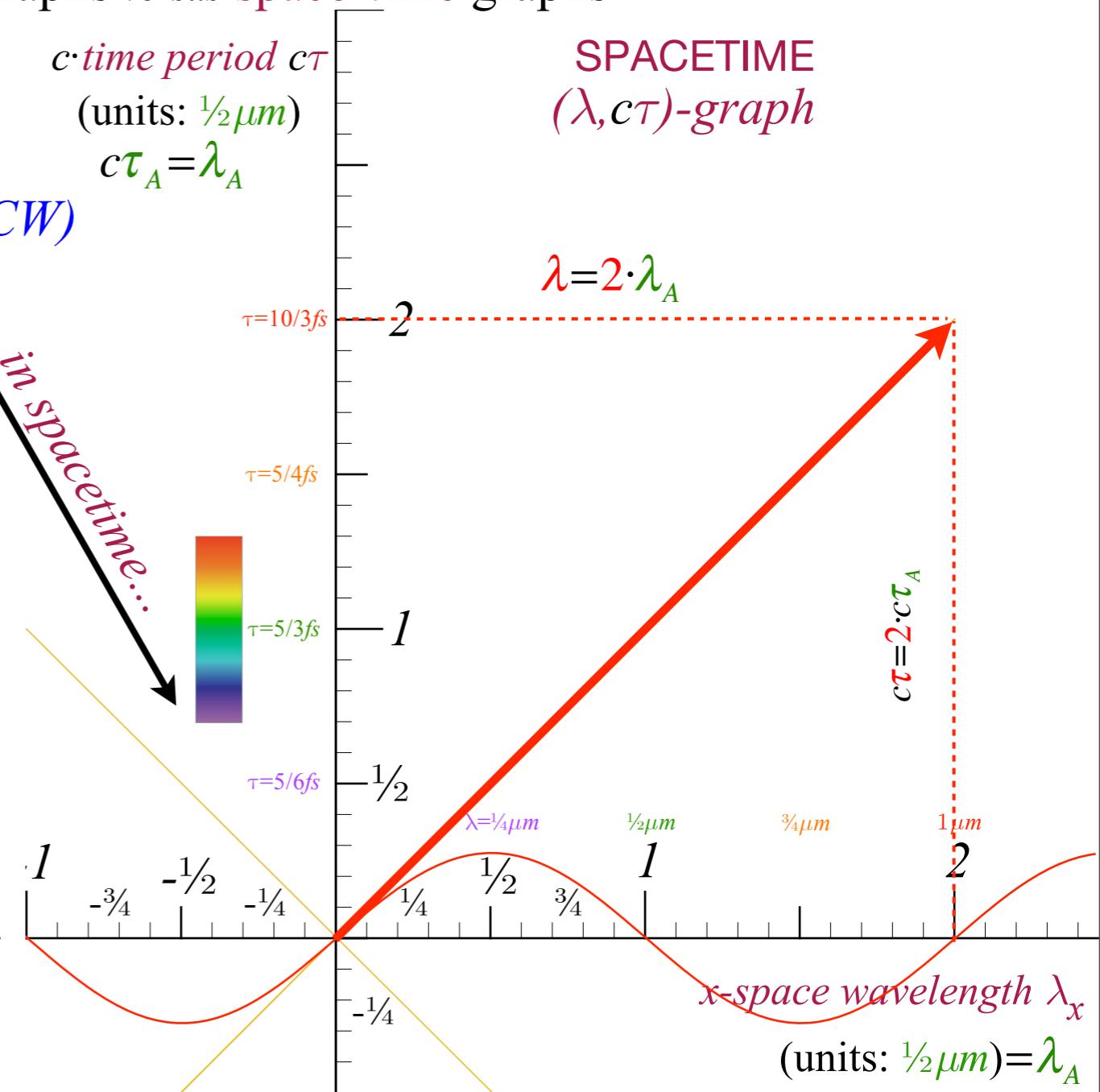
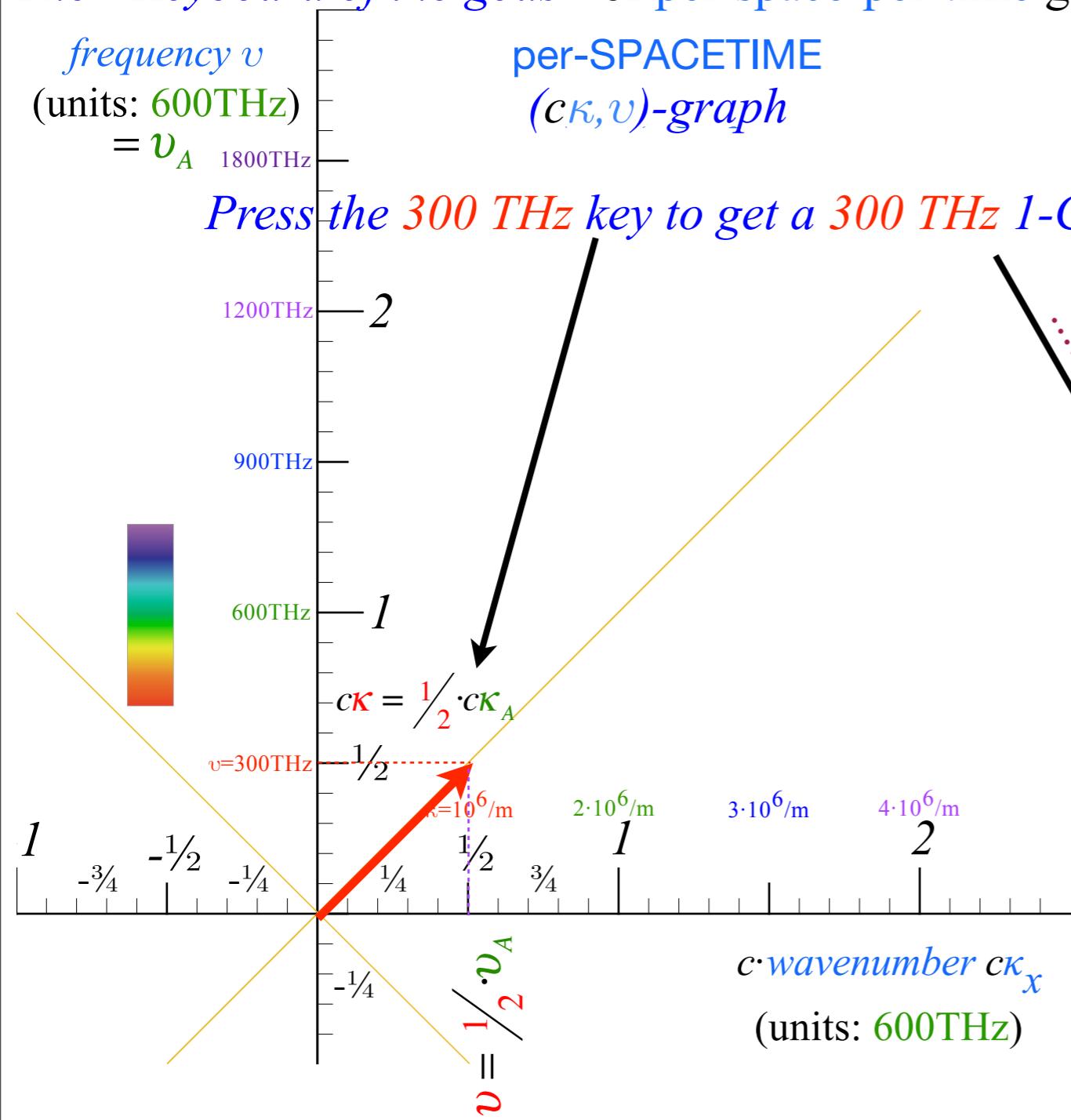
Ways to quantify **light** waves (1200 THz example)

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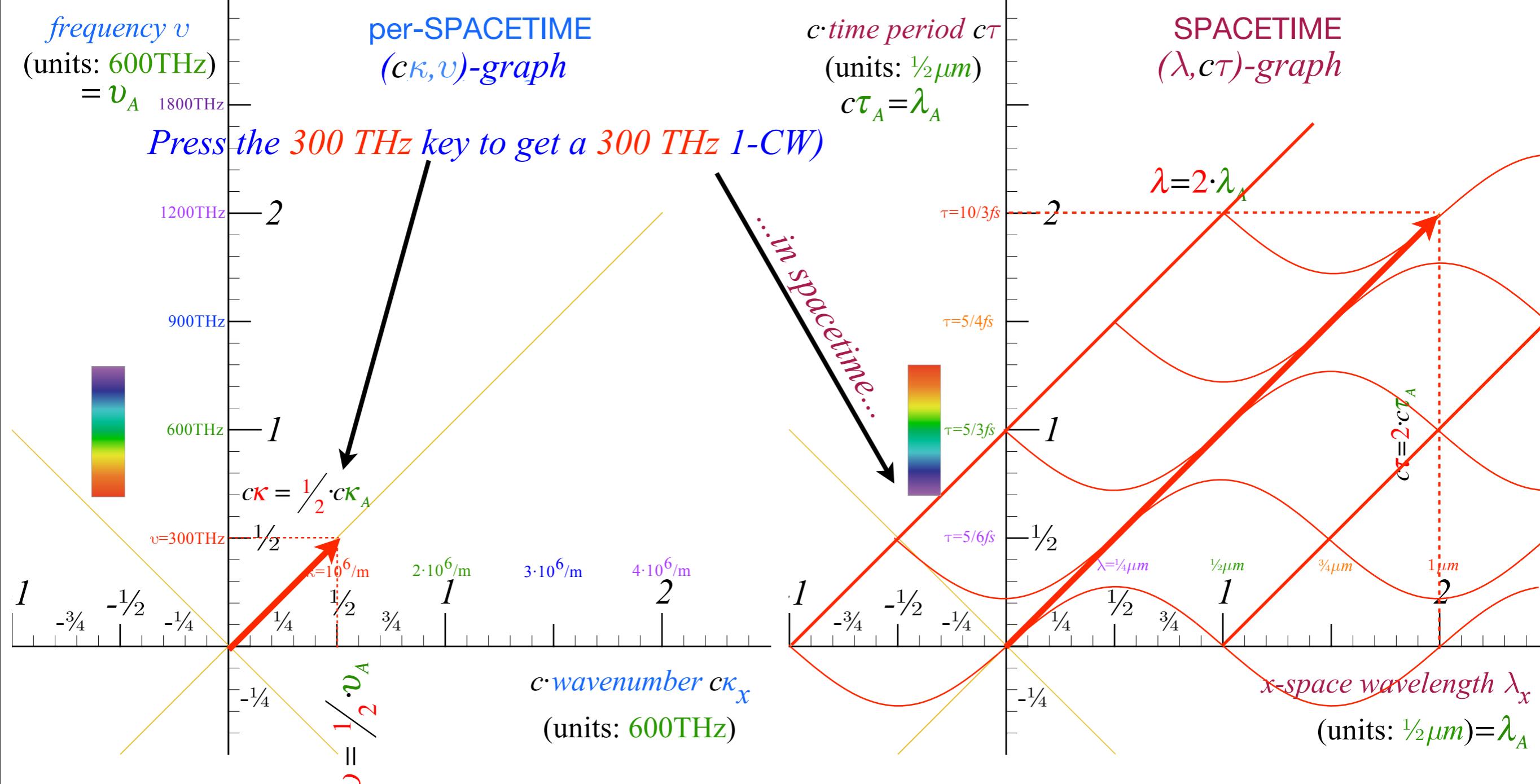
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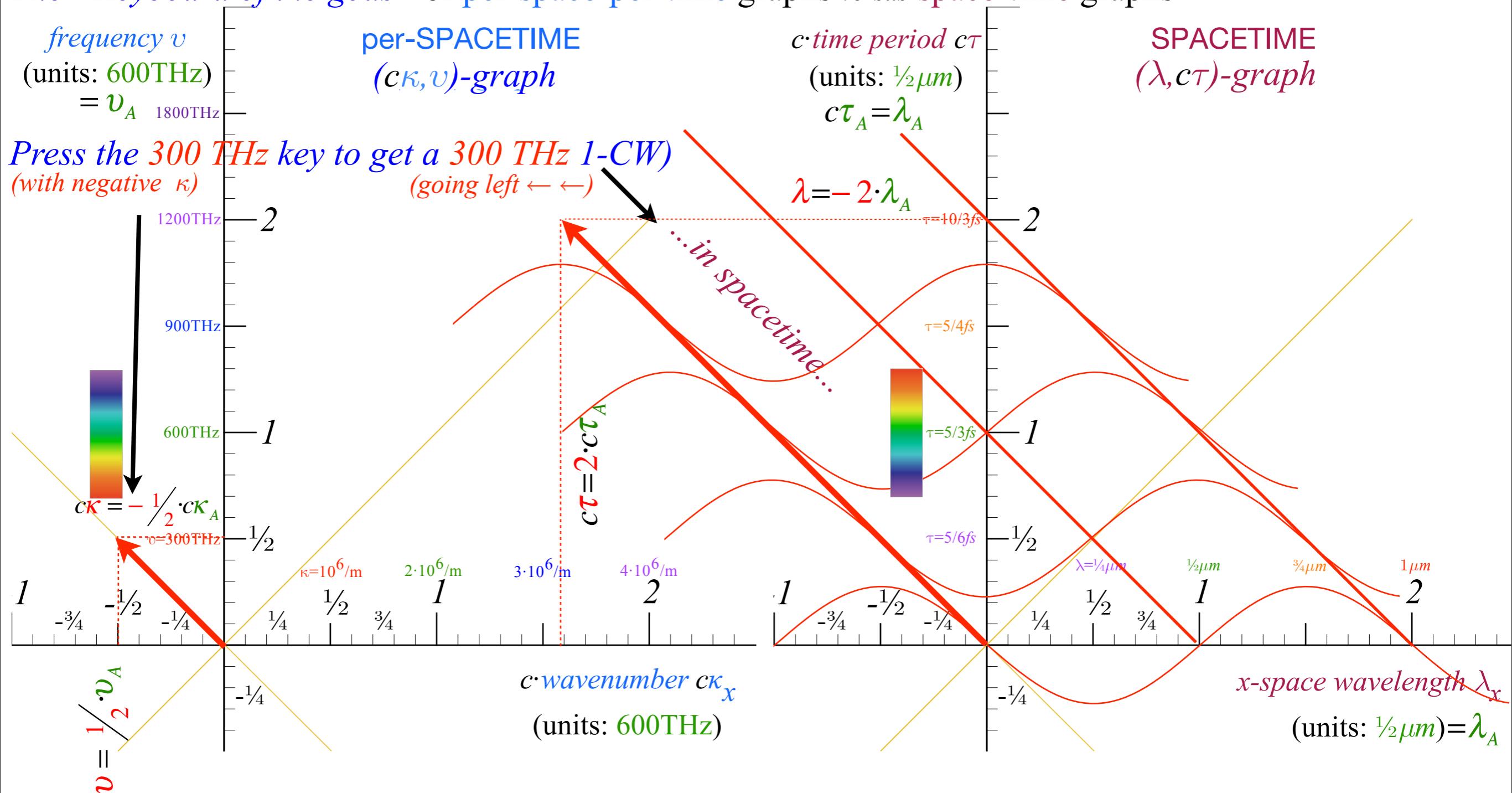
Ways to quantify light waves (*300 THz* example)

The “Keyboard of the gods” or per-space-per-time graphs versus space-time graphs



Ways to quantify **light** waves (300 THz example)

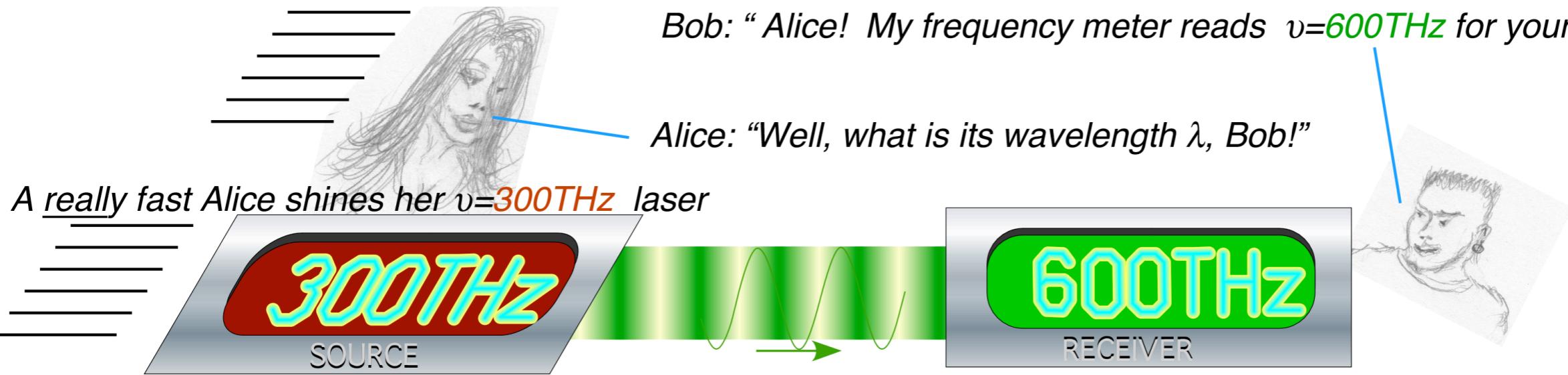
The “Keyboard of the gods” or per-space-per-time graphs versus space-time graphs



Ways to quantify **light** waves (**300 THz** example)

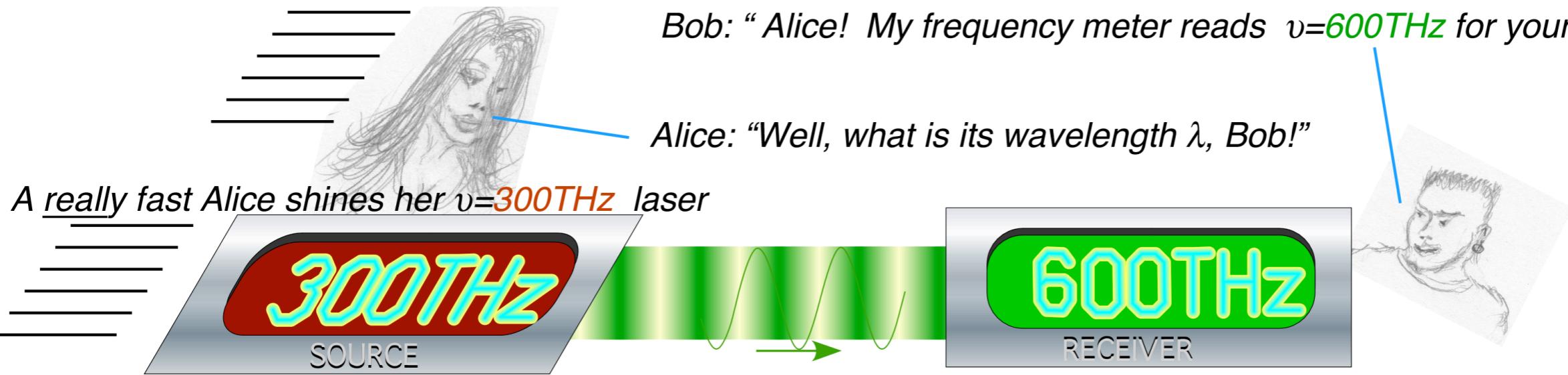
Clarify Evenson's CW Axiom (*All colors go c*) by Doppler effects

Alice tries to fool Bob that she's shining a 600THz laser. (Bob's unaware she's moving *really* fast...)



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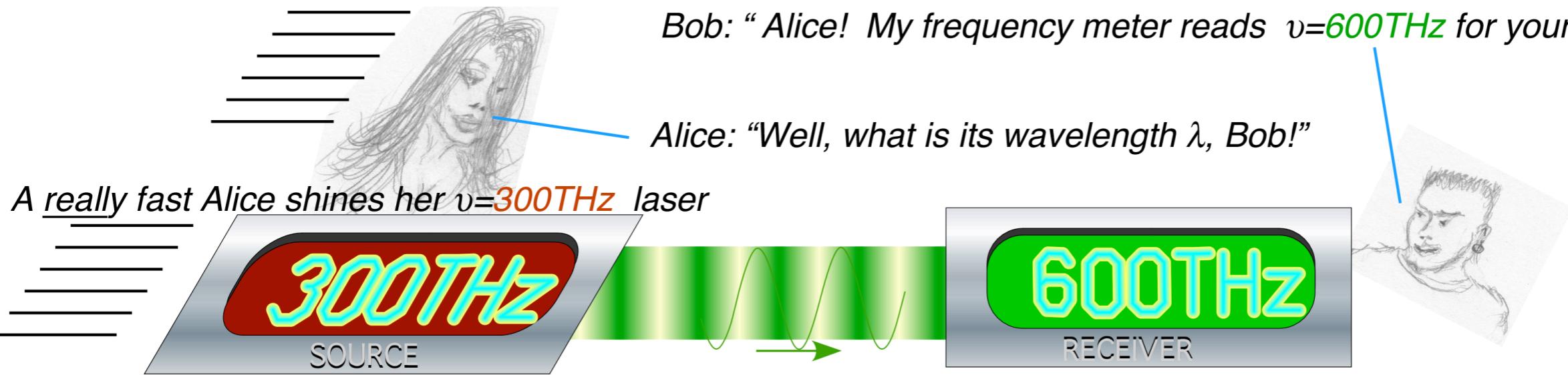
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Alice: "Well, what is its wavelength λ , Bob!"

600THz

RECEIVER

300THz

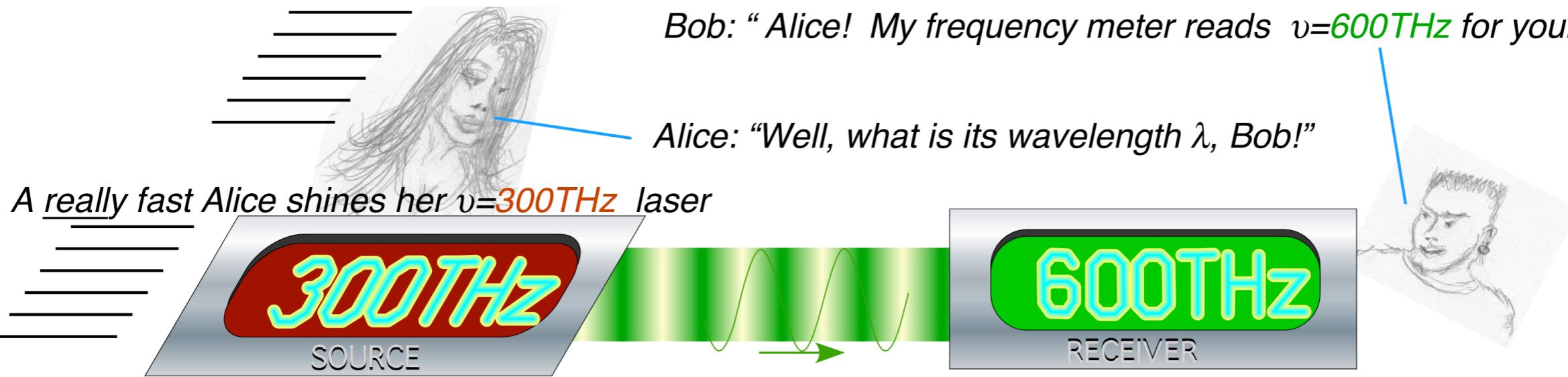
SOURCE

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Q2: If so, what "*phony*" λ does Bob see?

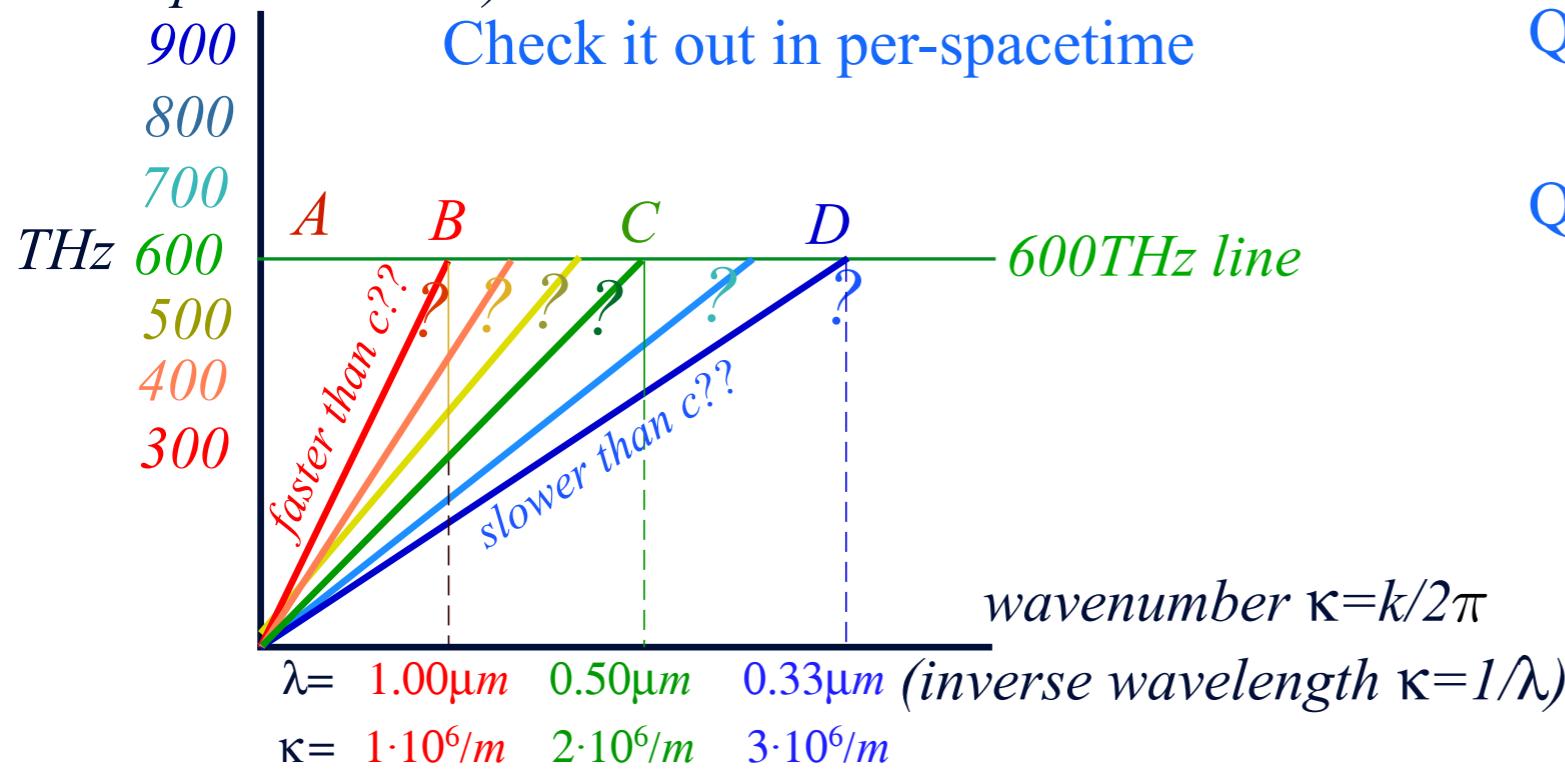
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frequency $\nu=\omega/2\pi$

(Inverse period $\nu=1/\tau$)

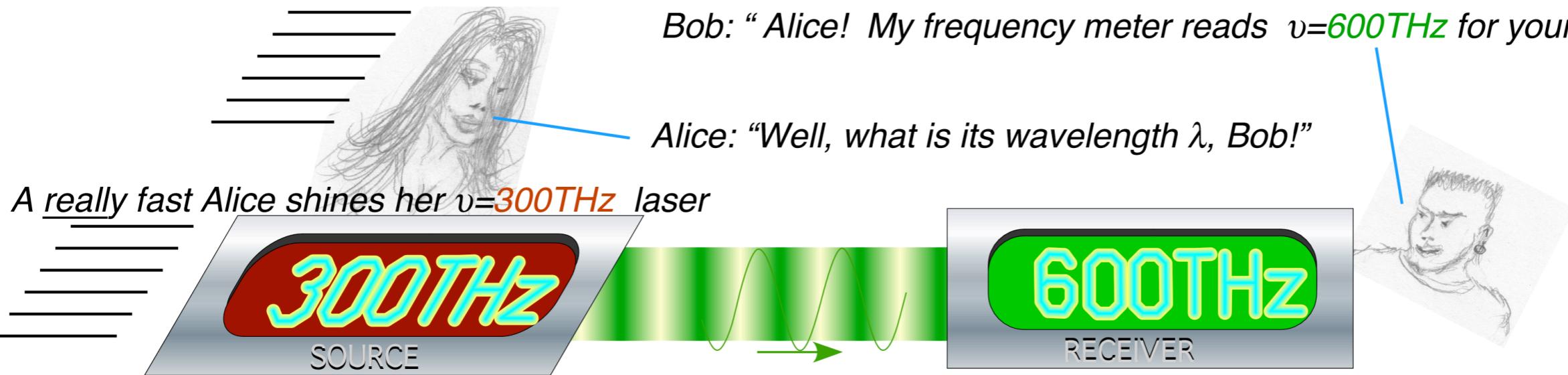


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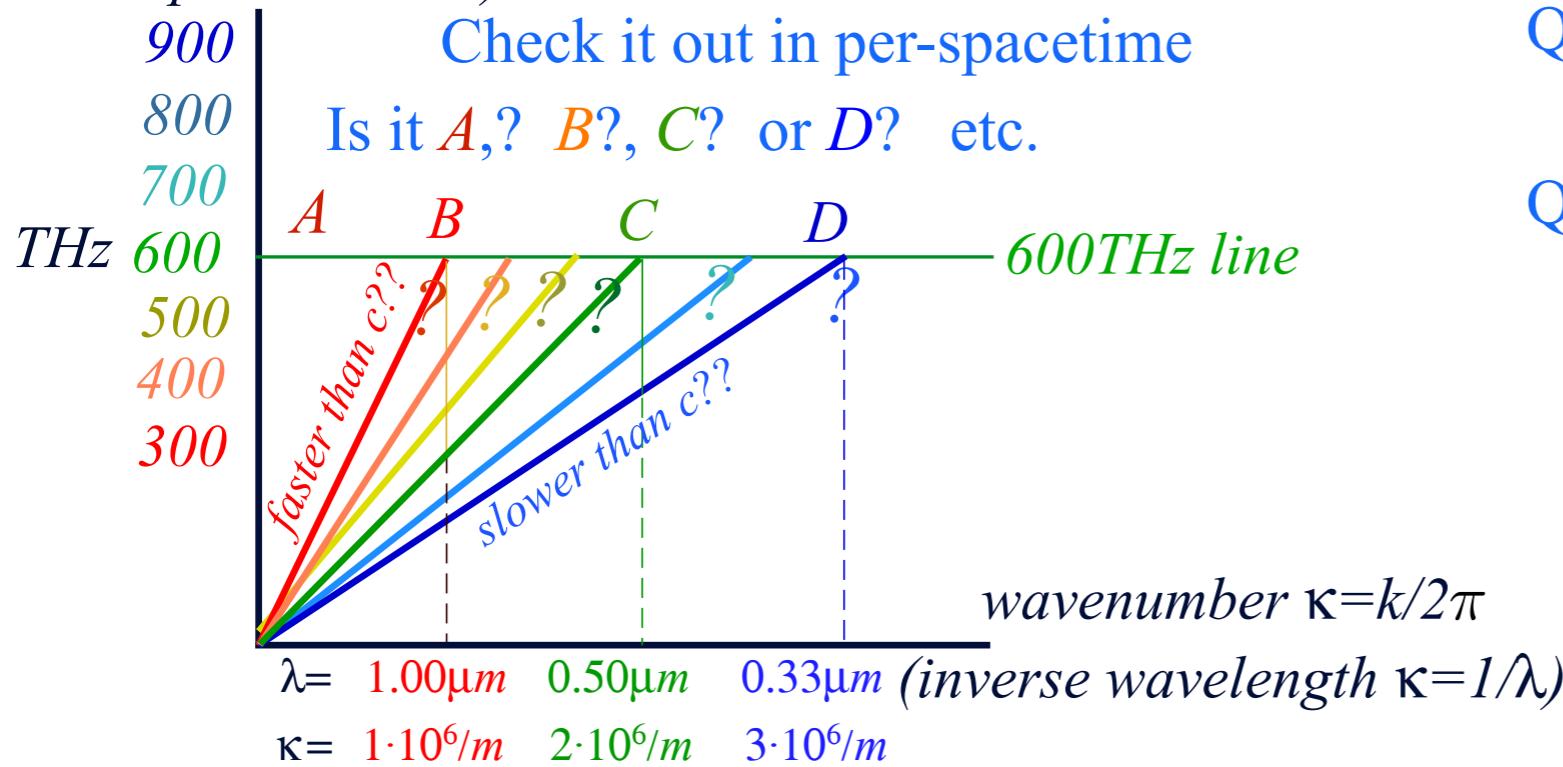
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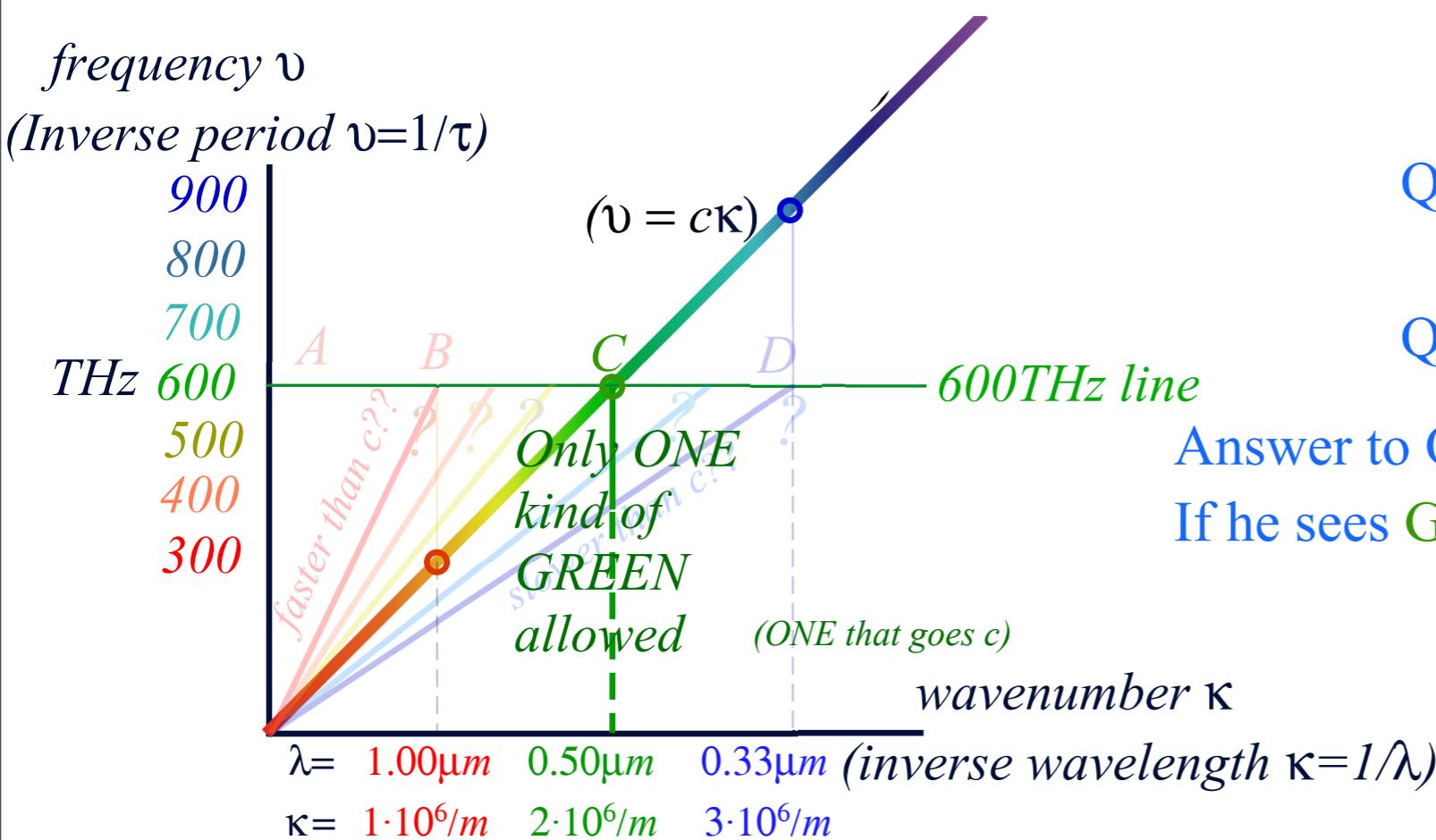
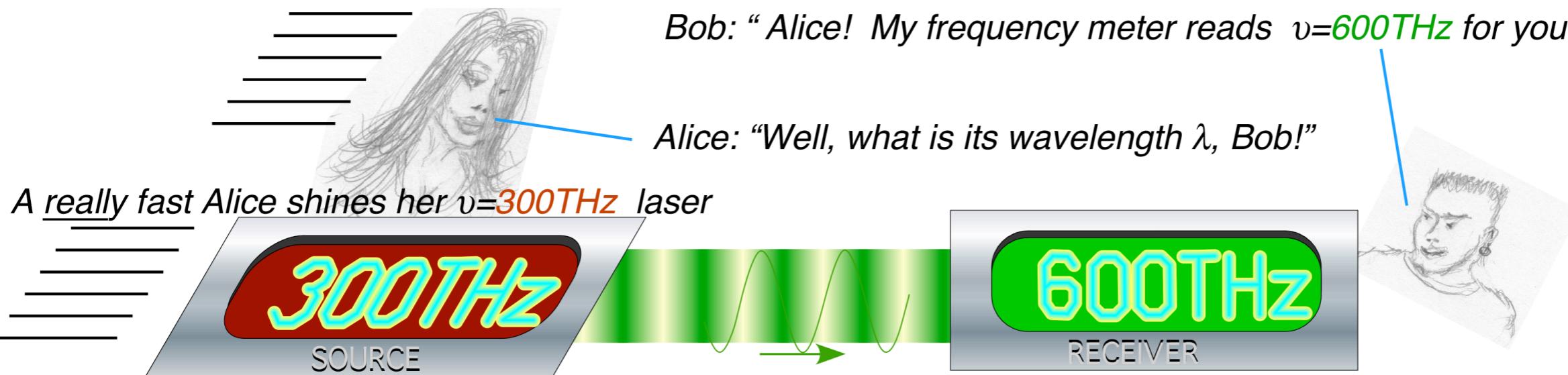


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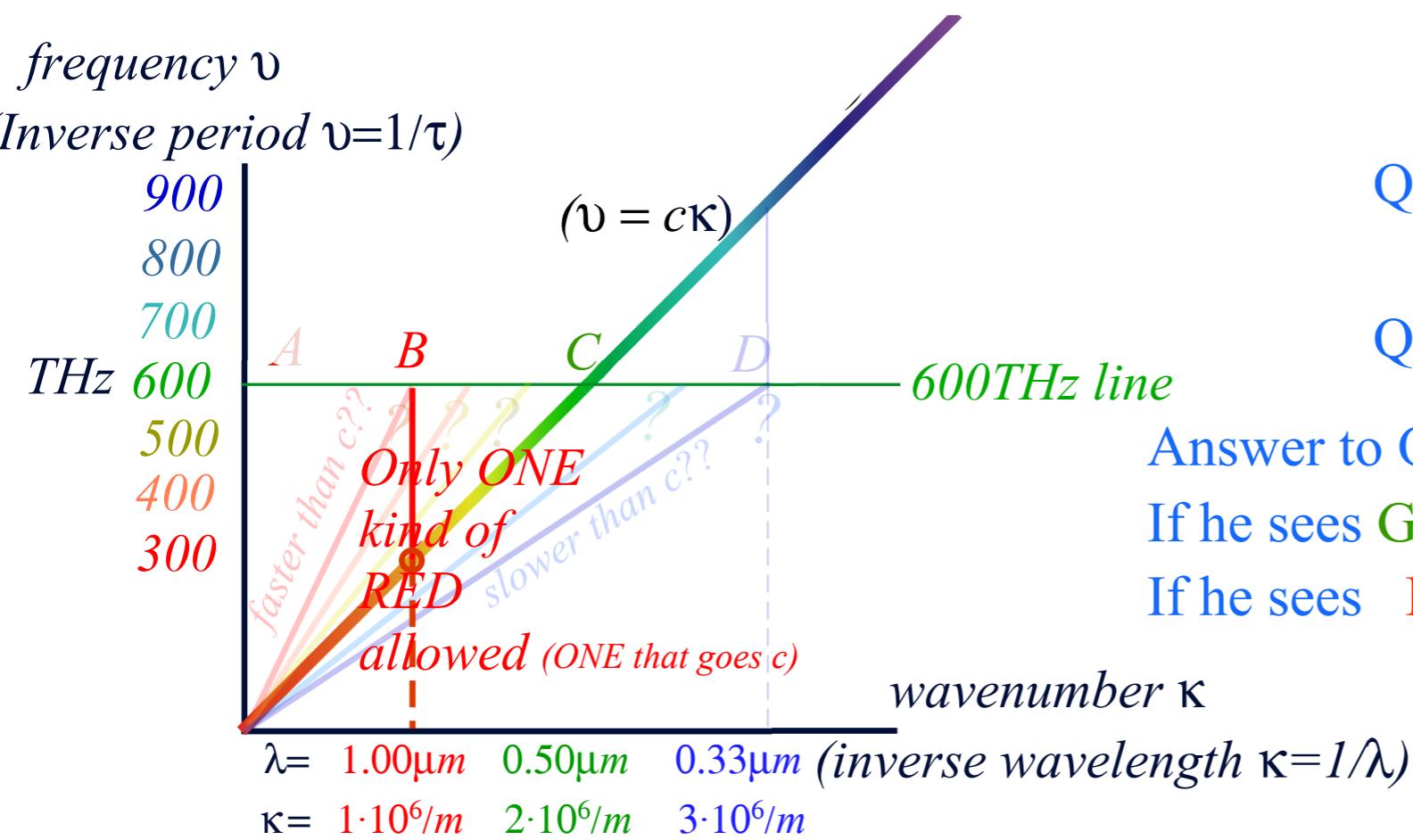
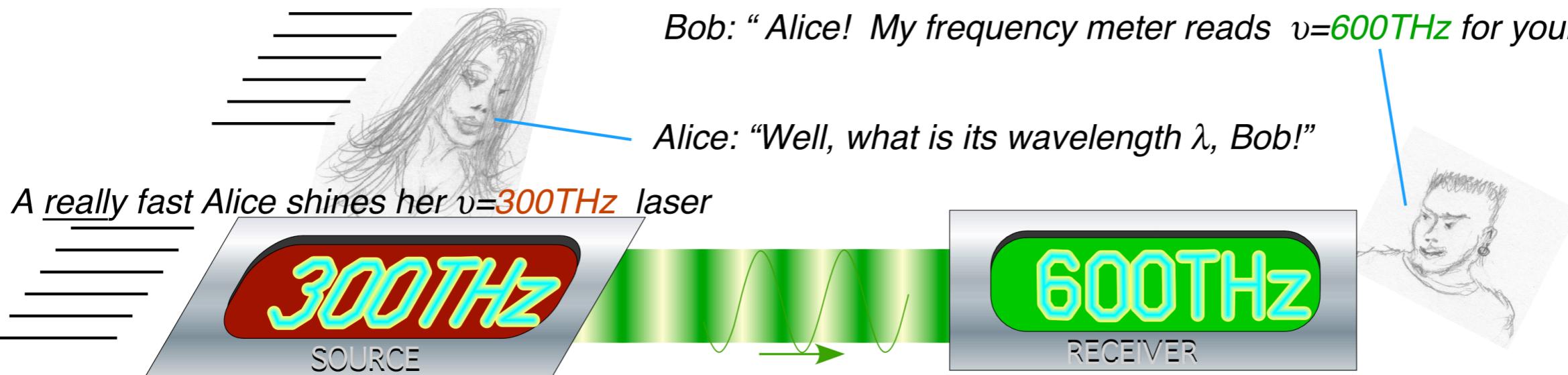
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If he sees Green 600THz then he measures $\lambda=0.5\mu\text{m}$.

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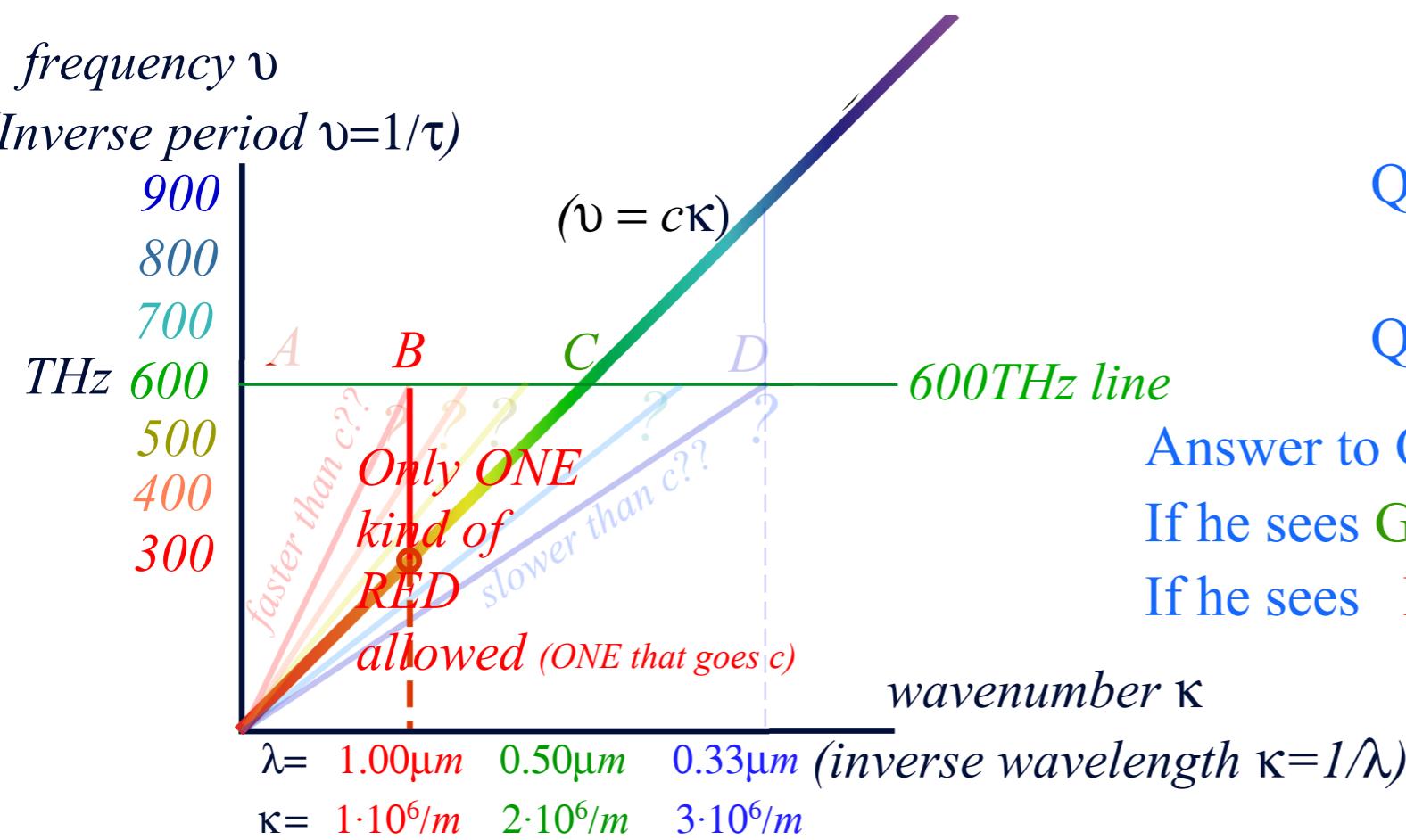
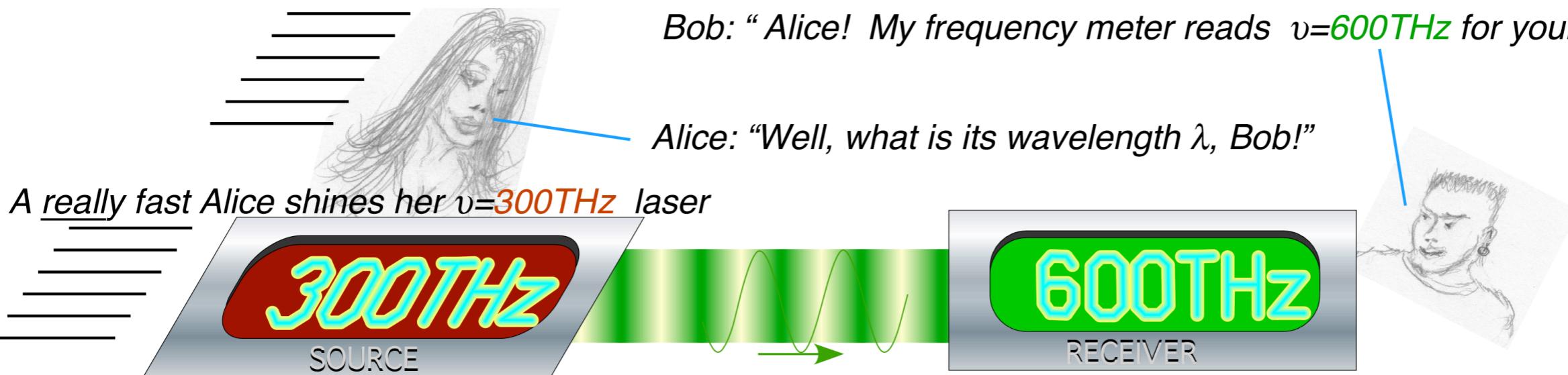
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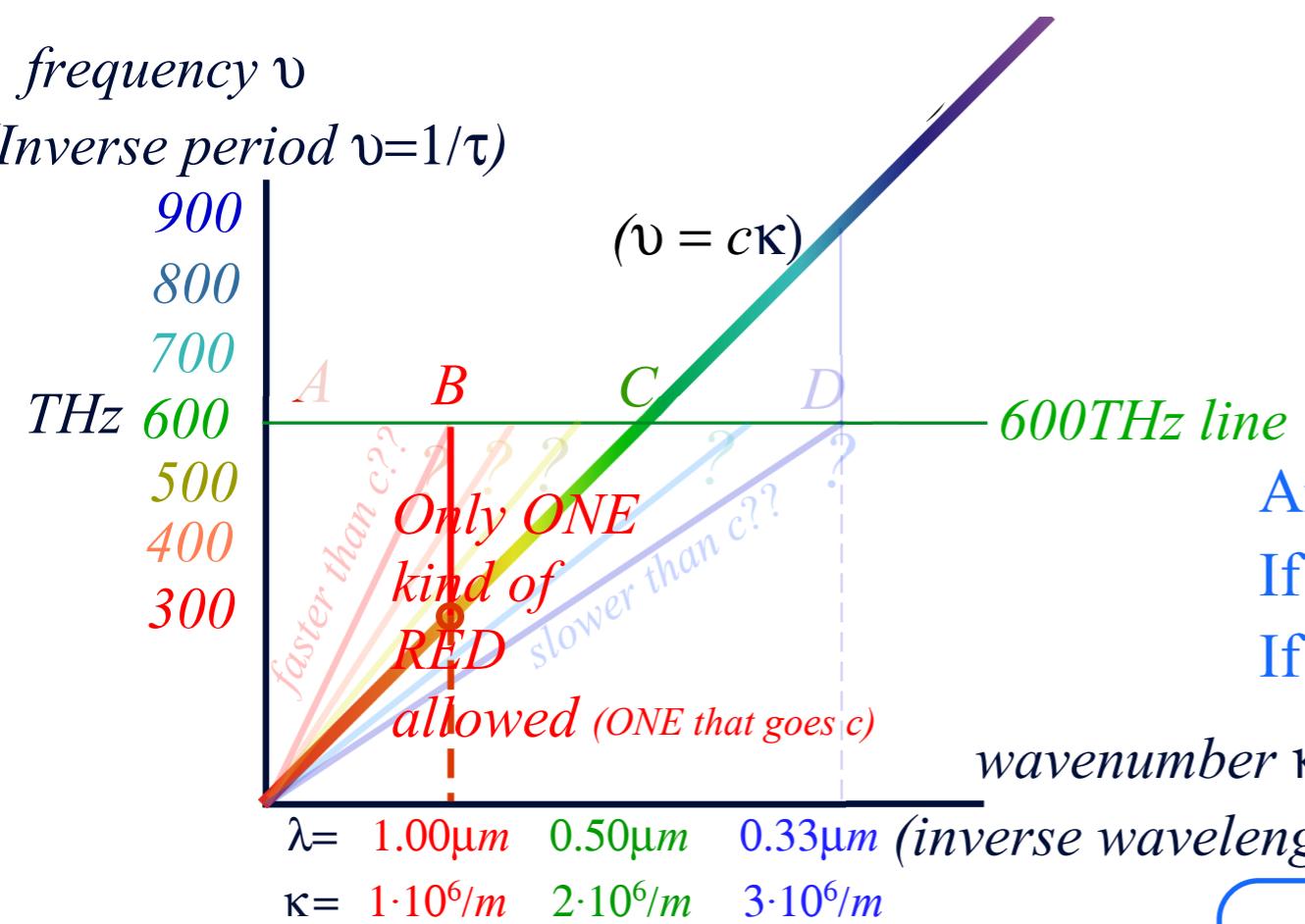
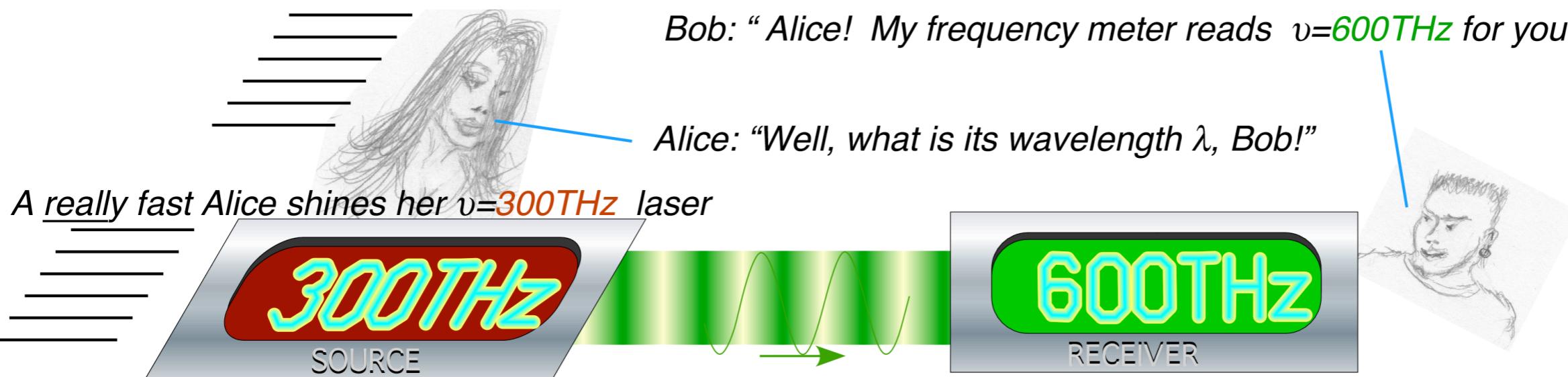
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Light carries no birth-certificate!

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Vacuum only makes one λ for each ν .*

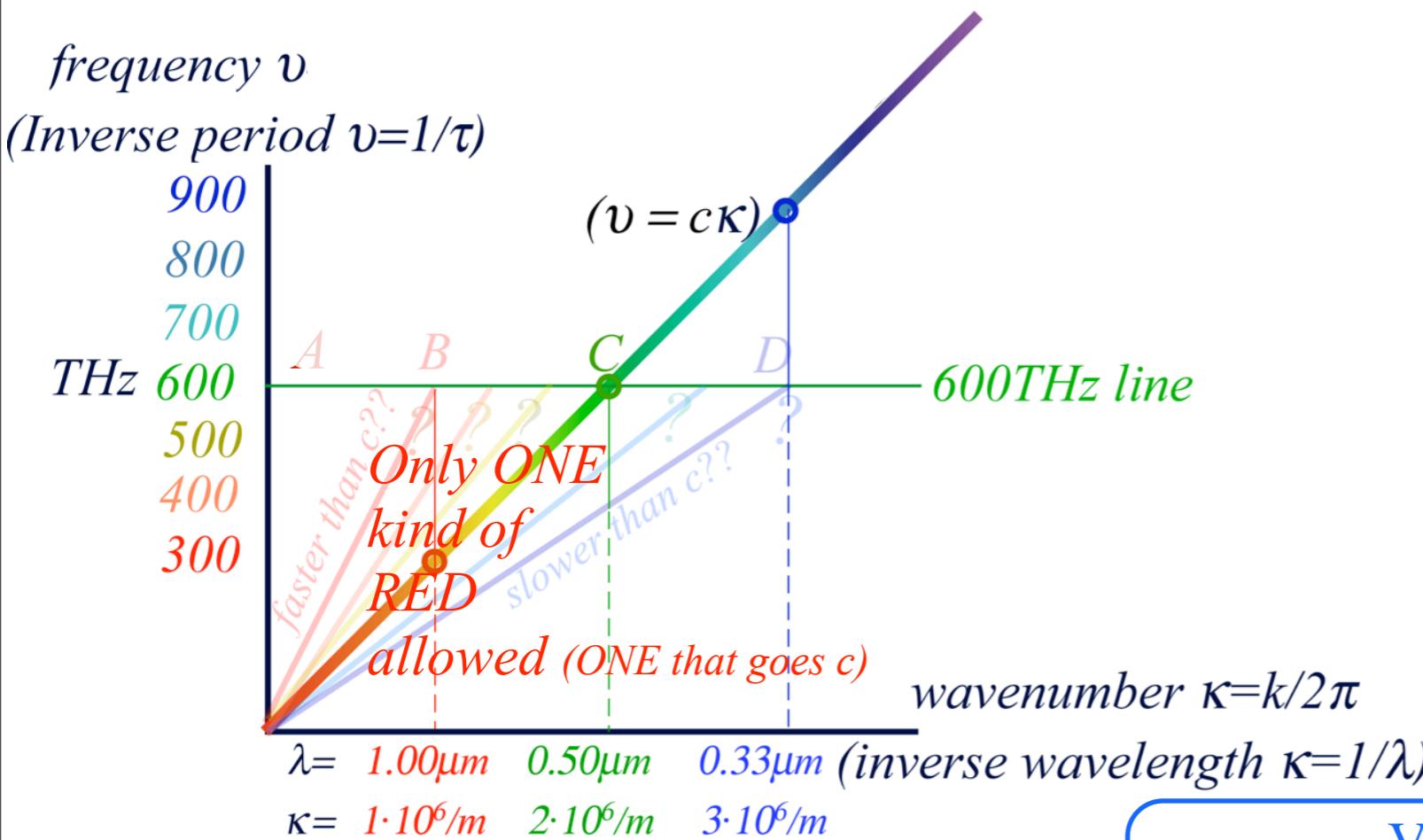
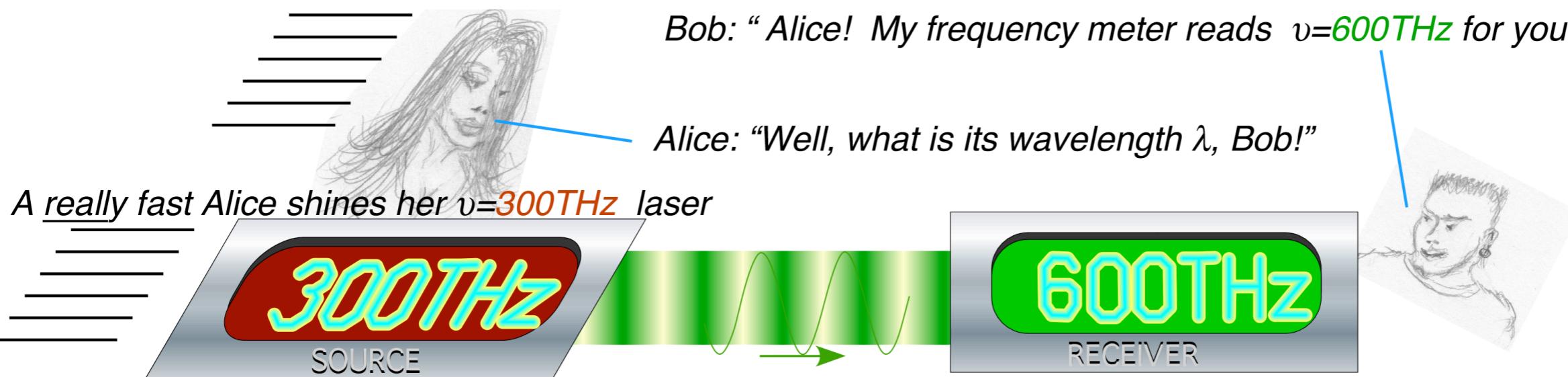
"All colors go $c = \lambda\nu = \nu/\kappa$ "

Then *Evenson's axiom* holds:

*for each beam and polarization orientation

Clarify Evenson's CW Axiom (*All colors go c*) by Doppler effects

Alice tries to fool Bob that she's shining a 600THz laser. (Bob's unaware she's moving *really fast...*)



Also could be labeled:
Linear-(non)-dispersion axiom: $\nu = ck$

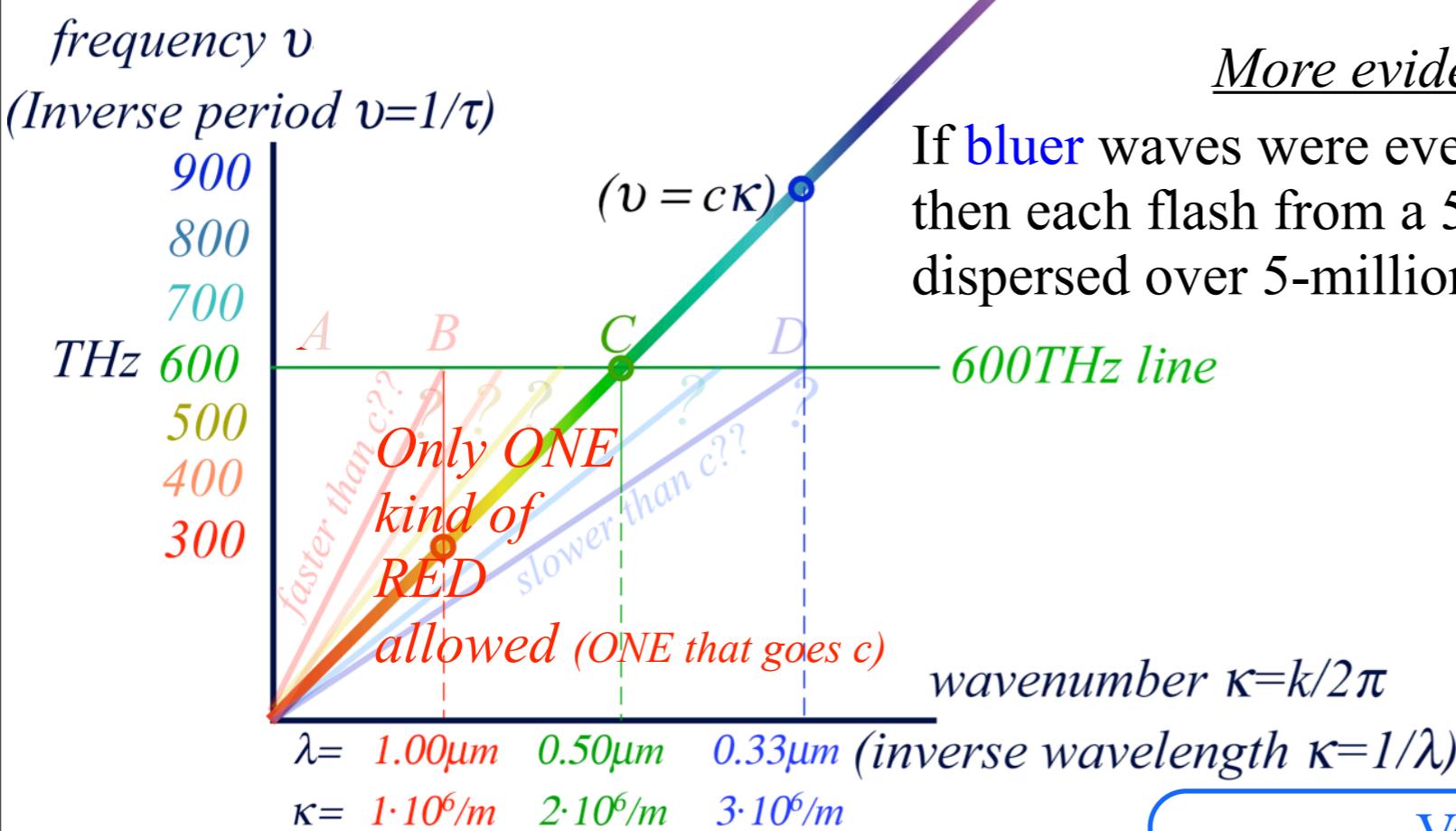
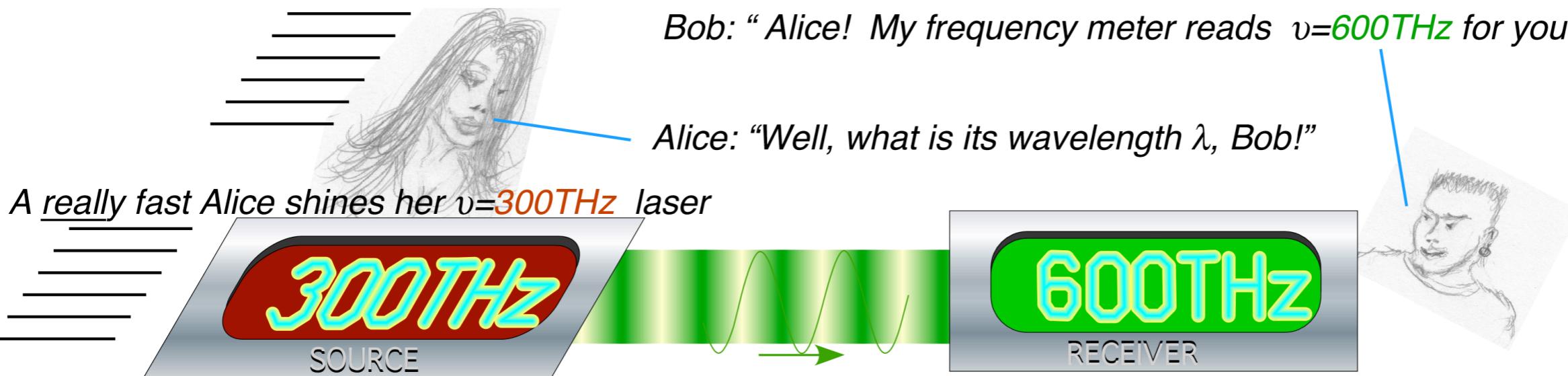
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Clarify Evenson's CW Axiom (*All colors go c*) by Doppler effects

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More evidence supporting Evenson's axiom
If bluer waves were even 0.1% faster (or slower) than redder ones
then each flash from a 5-billion light-year distant galaxy shows up
dispersed over 5-million years. (*Goodbye galactic astronomy!*)

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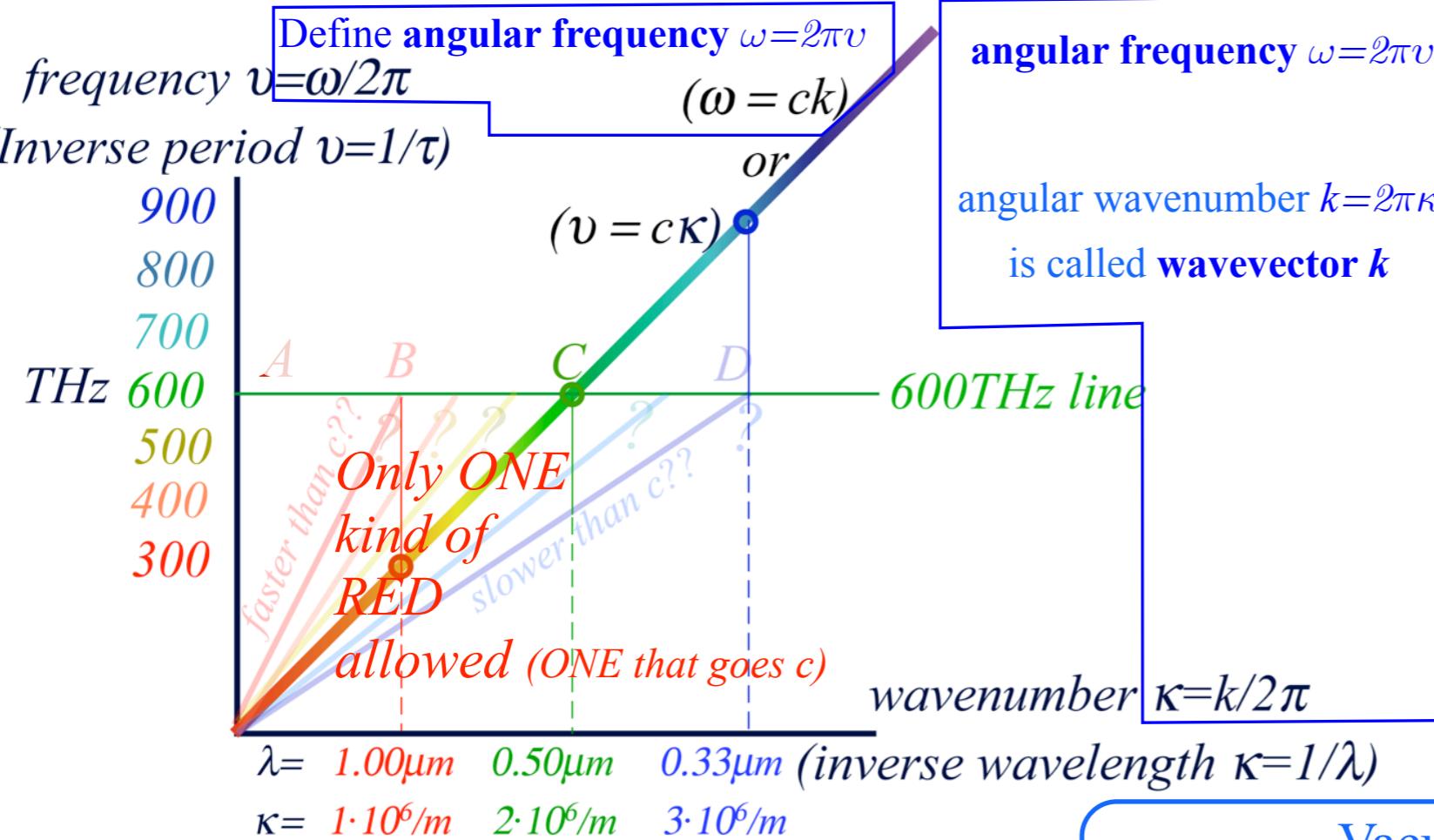
A really fast Alice shines her $\nu=300\text{THz}$ laser

Alice: "Well, what is its wavelength λ , Bob!"



300THz
SOURCE

600THz
RECEIVER



Coming Soon:
*Introduction
of
Laser-Phasor
clock
Parameters
 ω and k*

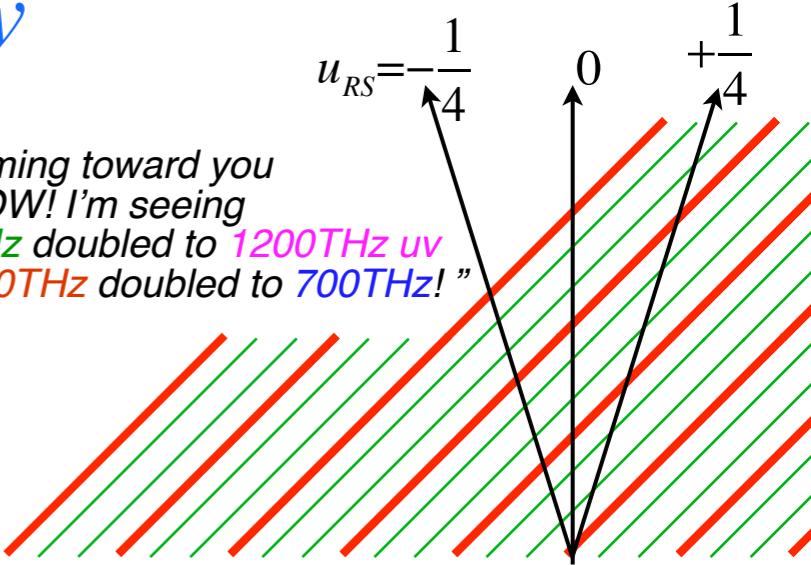
Also could be labeled:
Linear-(non)-dispersion axiom: $\nu = c\kappa$ or: $\omega = ck$

Vacuum only makes one λ for each ν .

"All colors go $c = \lambda\nu = \nu/\kappa = \omega/k$ "
Then *Evenson's axiom* holds:

Doppler shift-ratios and rapidity

VIEW FROM
ALICE'S LAB



Evenson's axiom is: "All **frequencies** march in lock-step." Hence, **Doppler shift ratio** $\langle R|S \rangle = \frac{v_R}{v_S}$ depends on relative velocity u_{RS} of RECEIVER **R** vs. SOURCE **S** but not on source frequency v_S :

$$v_{RECEIVER} = \langle R|S \rangle v_{SOURCE}$$

Light is GEOMETRIC

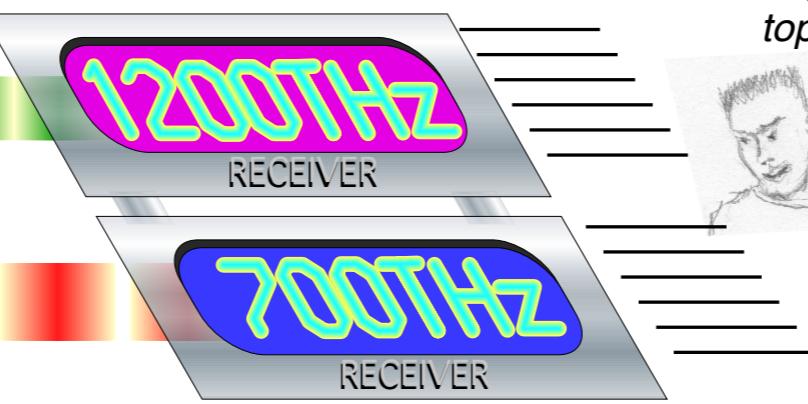
*(If light were ARITHMETIC
then $v_{RECEIVER} = v_{SOURCE} \pm \Delta_{RS}$
might be convenient.)*

Doppler shift-ratios and rapidity

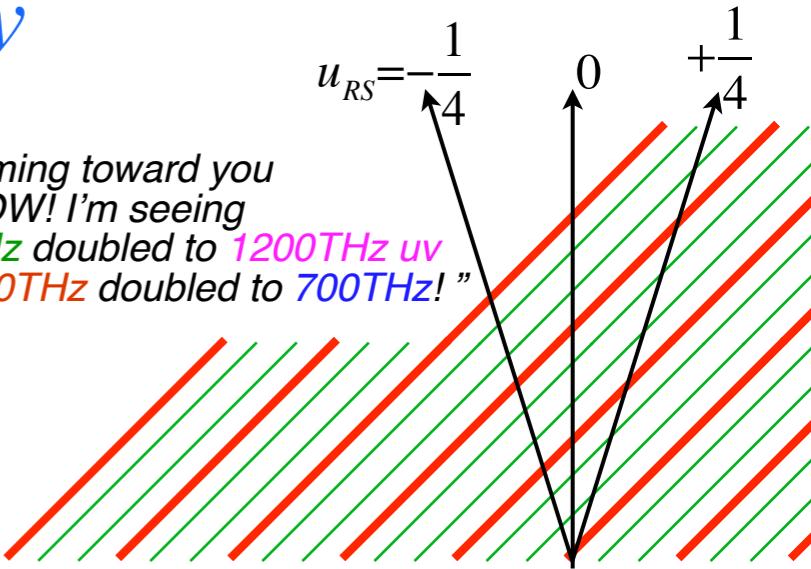
VIEW FROM
ALICE'S LAB



Alice: "Checkout my 600THz and 350THz beams!"



Bob: "Coming toward you
and WOW! I'm seeing
top 600THz doubled to 1200THz uv
and 350THz doubled to 700THz!"



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If Source-Receiver distance is *contracting*:

$$\frac{v_{RECEIVER}}{v_{SOURCE}} = \text{Blue shift} = \langle R|S \rangle > 1$$

If Source-Receiver distance is *expanding*:

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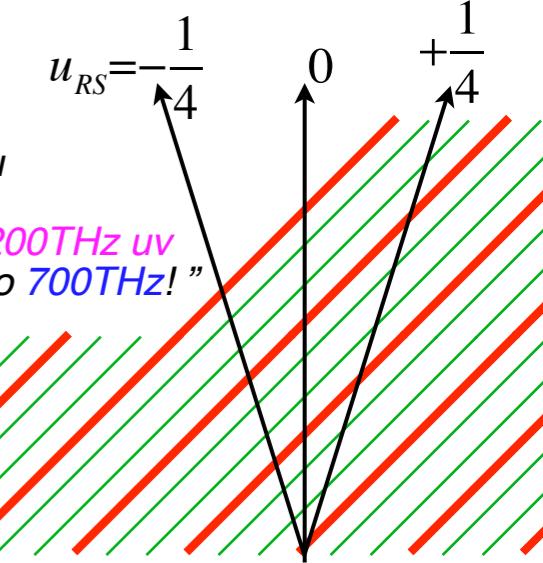
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Logarithm of $\langle R|S \rangle$ known as *Rapidity*: $\rho_{RS} = \log_e \langle R|S \rangle$ or: $\langle R|S \rangle = e^{\rho_{RS}}$

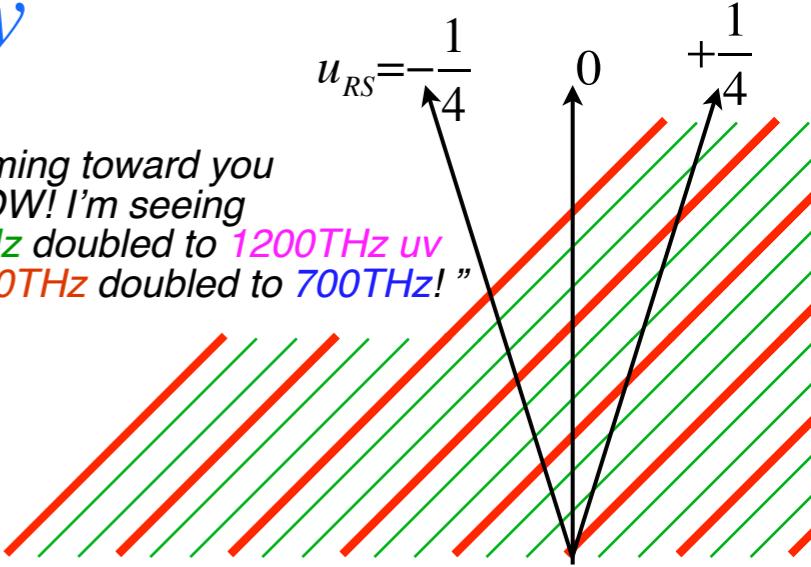
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$\langle R|S \rangle = e^{\rho_{RS}}$ with: $\rho_{RS} > 0$ for *contraction*,

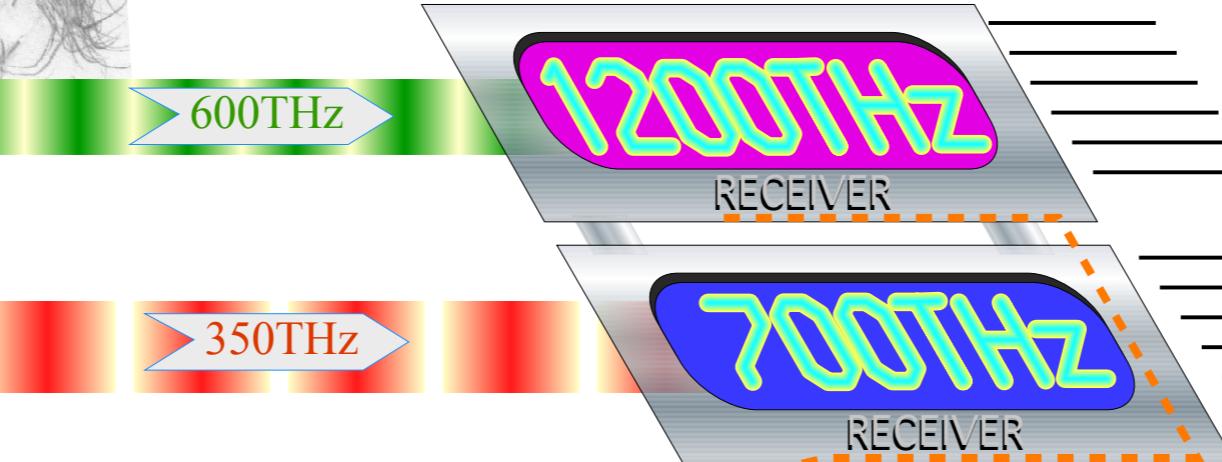
$\langle R|S \rangle = e^{\rho_{RS}}$ with: $\rho_{RS} < 0$ for *expansion*.

Doppler time-reversal symmetry

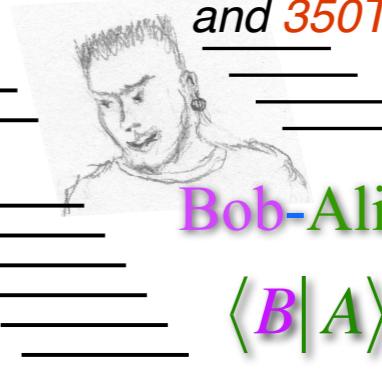
*VIEW FROM
ALICE'S LAB*



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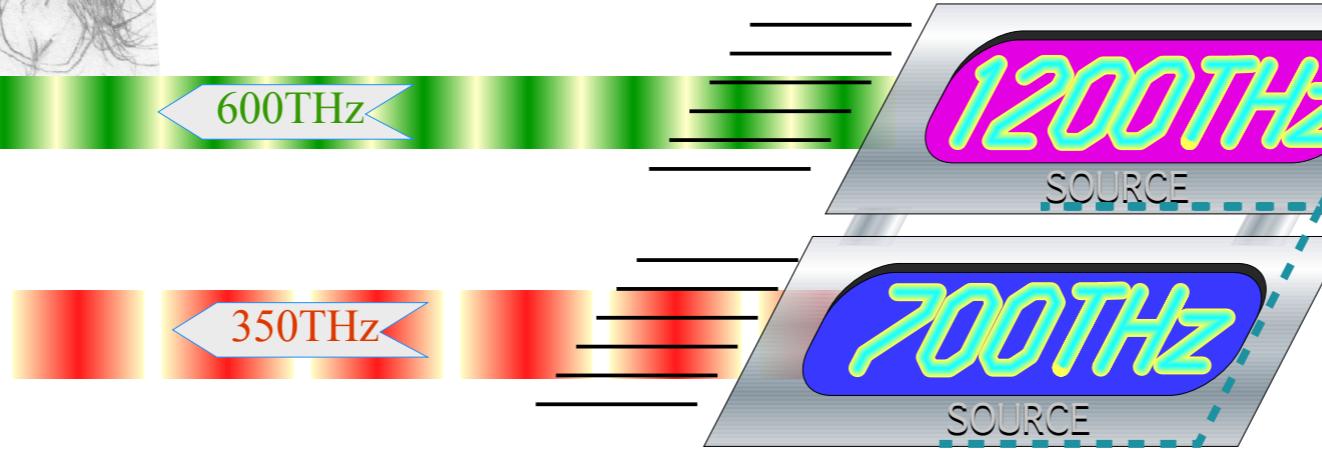
Bob-Alice Doppler ratio:
 $\langle B|A \rangle = \frac{v_B}{v_A} = \frac{1200}{600} = \frac{2}{1}$

Note: time-reversal switches SOURCE with RECEIVER, reverses motion of Bob's lasers and laser beams.
 (But, digital frequency readouts remain unchanged.)

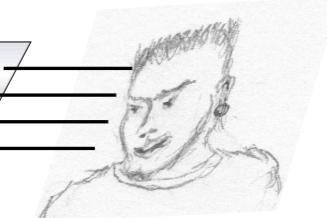
*VIEW FROM
ALICE'S LAB
(Time Reversed)*



Alice: "Well, I'm disappointed, Bob. Your so called 1200THz is a lousy 600THz, and I don't need any more Blue! (Fortunately, 700THz turned up as a warm 350THz .) "



Bob: "I'm leaving now! But, I'll send you a nice 1200THz uv beam and a Blue 700THz beam. "

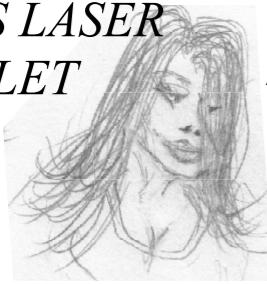


Alice-Bob Doppler ratio:
 $\langle A|B \rangle = \frac{v_A}{v_B} = \frac{600}{1200} = \frac{1}{2}$

Easy Doppler-shift and Rapidity calculation

ALICE'S LASER

GAUNTLET



Alice: "Hey Bob and Carla! Read your Doppler shifts of my **600THz** beam. What **rapidity** ρ_{BA} or ρ_{BC} do you'all have relative to me and each other?"



Doppler ratio:

$$\langle R|S \rangle = \frac{v_{RECEIVER}}{v_{SOURCE}}$$

Bob-Alice Doppler ratio:

$$\langle B|A \rangle = \frac{v_B}{v_A} = \frac{1200}{600} = \frac{2}{1}$$

Carla-Alice Doppler ratio:

$$\langle C|A \rangle = \frac{v_C}{v_A} = \frac{400}{600} = \frac{2}{3}$$

Easy Doppler-shift and Rapidity calculation

ALICE'S LASER

GAUNTLET



Alice: "Hey Bob and Carla! Read your Doppler shifts of my **600THz** beam. What **rapidity** ρ_{BA} or ρ_{BC} do you'all have relative to me and each other?"



Doppler ratio:

$$\langle R|S \rangle = \frac{v_{RECEIVER}}{v_{SOURCE}}$$

rapidity:

$$\rho_{RS} = \log_e \langle R|S \rangle$$

Definition of Rapidity

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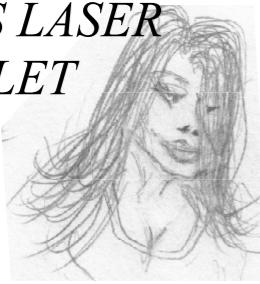
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Bob: I see Doppler Blue shift to 1200THz



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Carla-Alice Doppler ratio:

$$\langle C|A \rangle = \frac{v_C}{v_A} = \frac{400}{600} = \frac{2}{3}$$

Carla-Alice rapidity:

$$\rho_{CA} = \log_e \langle C|A \rangle = \log_e \frac{2}{3}$$

Easy Doppler-shift and Rapidity calculation

ALICE'S LASER

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Bob-Alice rapidity:

$$\rho_{BA} = \log_e \langle B|A \rangle = \log_e \frac{2}{1}$$

$$\rho_{BA} = 0.69 \quad (\text{so: } \rho_{AB} = -0.69)$$

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Carla: I see Doppler Red shift to 400THz



Carla-Alice Doppler ratio:

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Carla-Alice rapidity:

$$\rho_{CA} = \log_e \langle C|A \rangle = \log_e \frac{2}{3}$$

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Easy Doppler-shift and Rapidity calculation

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$$\rho_{RS} = \log_e \langle R|S \rangle$$

or:

$$\langle R|S \rangle = e^{\rho_{RS}}$$

Definition of Rapidity

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$$\rho_{CA} = -0.41$$

Carla-Bob Doppler ratio:

$$\langle C|B \rangle = \frac{v_C}{v_B} = \frac{v_C}{v_A} \frac{v_A}{v_B} = \langle C|A \rangle \langle A|B \rangle$$

Carla-Bob rapidity:

$$e^{\rho_{CB}} = e^{\rho_{CA}} e^{\rho_{AB}}$$

Easy Doppler-shift and Rapidity calculation

ALICE'S LASER

GAUNTLET



Alice: "Hey Bob and Carla! Read your Doppler shifts of my 600THz beam. What rapidity ρ_{BA} or ρ_{BC} do you'all have relative to me and each other?"



Doppler ratio:

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Carla-Alice Doppler ratio:

$$\langle C|A \rangle = \frac{v_C}{v_A} = \frac{400}{600} = \frac{2}{3}$$

Carla-Bob rapidity:

$$e^{\rho_{CB}} = e^{\rho_{CA}} e^{\rho_{AB}} \quad \text{implies: } \rho_{CB} = \rho_{CA} + \rho_{AB} = -0.41 - 0.69 = -1.10$$

Bob: I see Doppler Blue shift to 1200THz



Carla: I see Doppler Red shift to 400THz



Easy Doppler-shift and Rapidity calculation

ALICE'S LASER

GAUNTLET



Alice: "Hey Bob and Carla! Read your Doppler shifts of my **600THz** beam. What **rapidity** ρ_{BA} or ρ_{BC} do you'all have relative to me and each other?"



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Carla-Alice rapidity:

$$e^{\rho_{CB}} = e^{\rho_{CA}} e^{\rho_{AB}} \text{ implies:}$$

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Carla-Alice Doppler ratio:

$$\langle C|A \rangle = \frac{v_C}{v_A} = \frac{400}{600} = \frac{2}{3}$$

Carla-Alice rapidity:

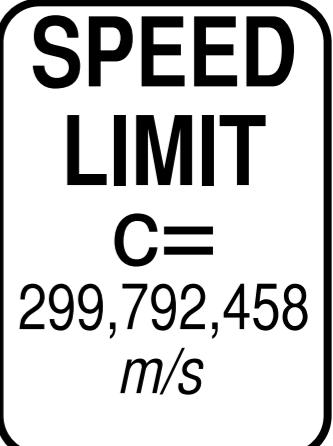
$$\rho_{CA} = \log_e \langle C|A \rangle = \log_e \frac{2}{3}$$

$$\rho_{CA} = -0.41$$

Galileo's Revenge (part 1)

Rapidity adds just like Galilean velocity

$$\rho_{CB} = \rho_{CA} + \rho_{AB} = -0.41 - 0.69 = -1.10$$



Level 2 Secrets (*which also shouldn't be secrets!*)

Special relativity and quantum mechanics
are very much a story of
the geometry of light-wave motion

- How do we measure space and time with light waves?
 Use *1CW laser-phasors* for a *phase-based* theory
- How do we make spacetime coordinate graph with light waves?
 Use *2CW laser-phasors* and wave interference geometry
 Get Einstein-Lorentz-Minkowski graphs for free!

1CW Laser-phasor wave function

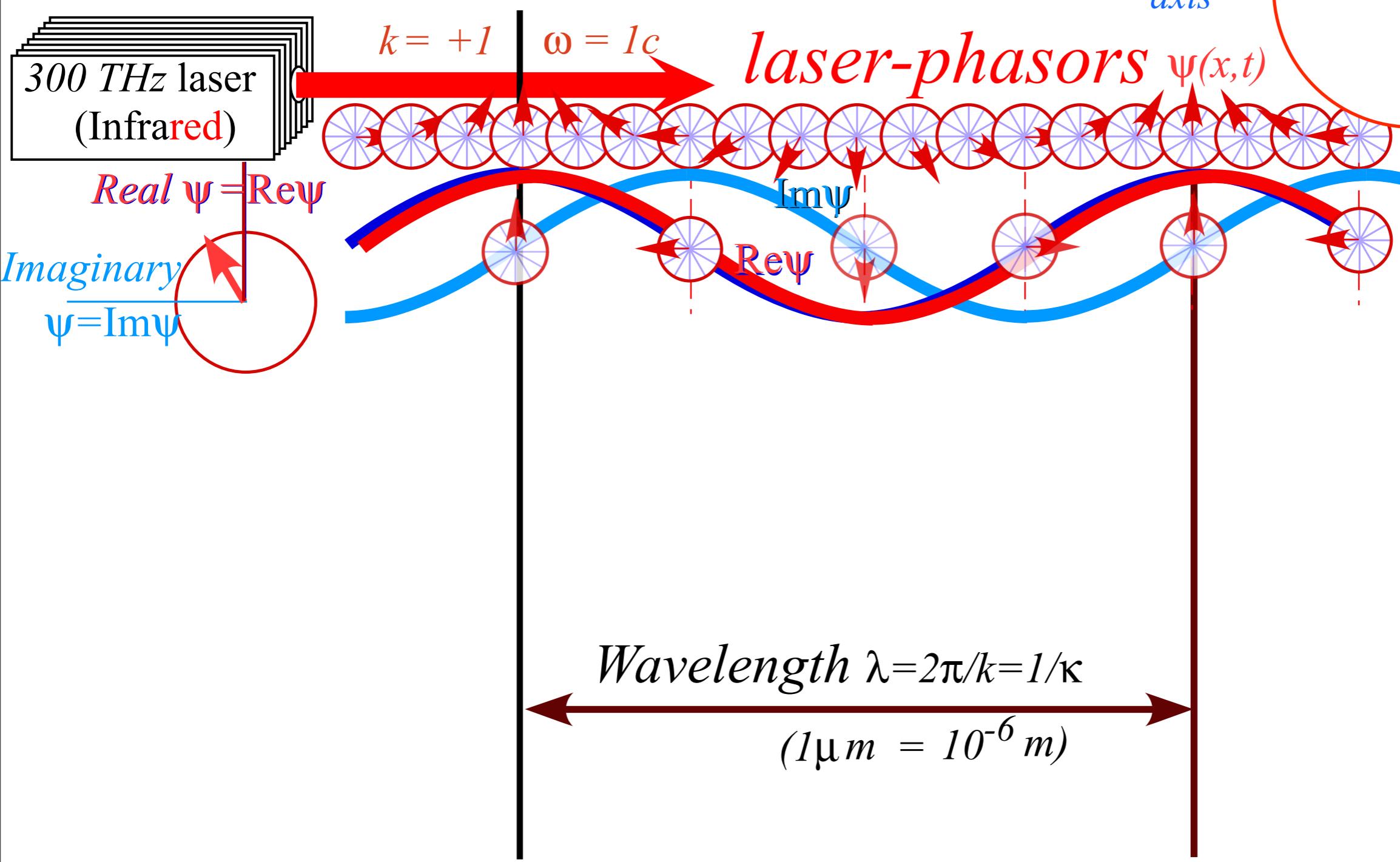
Dimensionless Light wave-velocity $c/c=1$

$$\frac{V_{light}}{c} = \frac{\lambda}{c\tau} = \frac{1/\kappa}{c/v} = \frac{v}{c\kappa} = \frac{1/\tau}{c/\lambda} = 1 = \frac{\omega}{ck} \text{ angular units}$$

$$\psi = A \cdot e^{i(kx - \omega t)} = A \cdot \cos(kx - \omega t) + iA \cdot \sin(kx - \omega t)$$

Amplitude A

phase-angle $(kx - \omega t)$



1CW Laser-phasor wave function

Dimensionless Light wave-velocity $c/c=1$

$$\frac{V_{light}}{c} = \frac{\lambda}{c\tau} = \frac{1/\kappa}{c/v} = \frac{v}{c\kappa} = \frac{1/\tau}{c/\lambda} = 1 = \frac{\omega \text{ angular}}{ck \text{ units}}$$

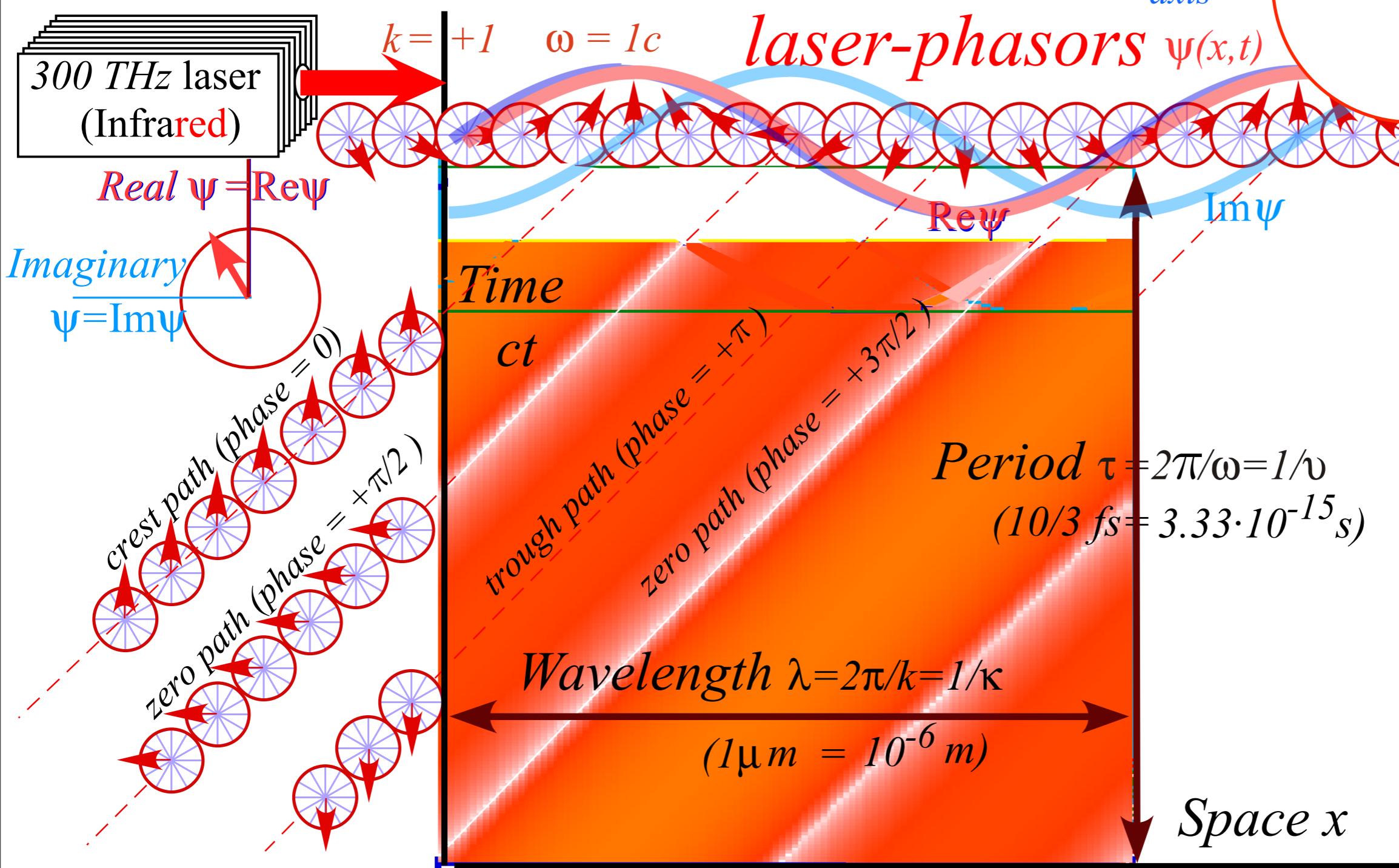
Q: Where is phase $= (kx - \omega t) = 0$?

A: It is wherever this is: $\frac{x}{t} = \frac{\omega}{k}$

$$\psi = A \cdot e^{i(kx - \omega t)} = A \cdot \cos(kx - \omega t) + iA \cdot \sin(kx - \omega t)$$

Amplitude A

phase-angle $(kx - \omega t)$



1CW Laser-phasor wave function

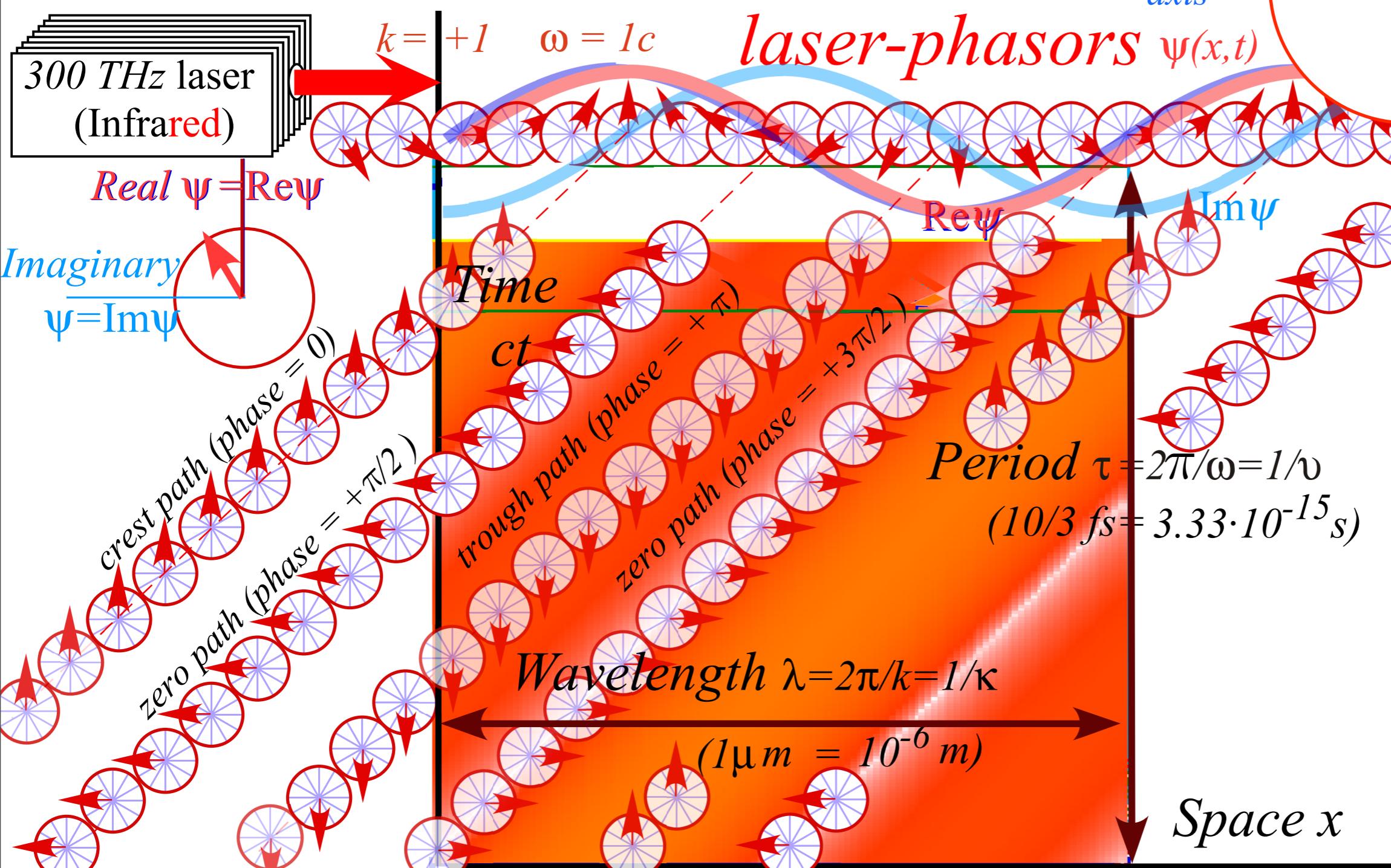
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1CW Laser-phasor wave function

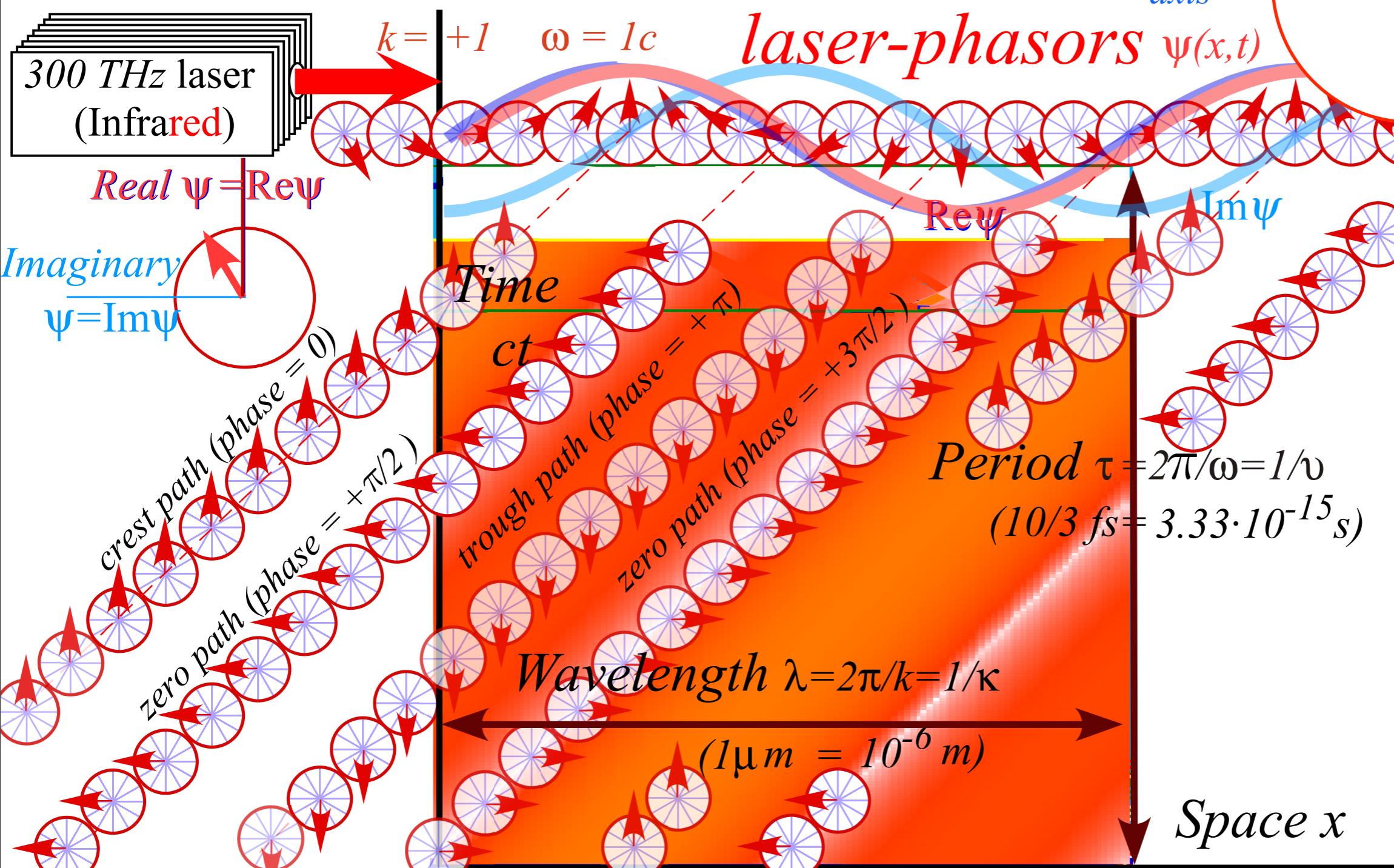
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Amplitude A

phase-angle $(kx - \omega t)$



*Alice: OK, Bob.
We're gonna' hit
you from both
sides, now!*

Colliding 2CW laser beams

*Carla:
Look out, Bob!*

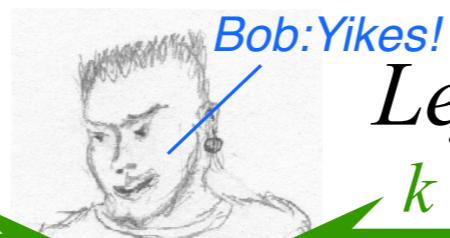
Right-moving wave $e^{i(kx-\omega t)}$

CW Dye-laser
600 THz

Alice's laser

$$k = +2$$

$$\omega = 2c$$



Bob: Yikes!

Left-moving wave $e^{i(-kx-\omega t)}$

CW Dye-laser
600 THz

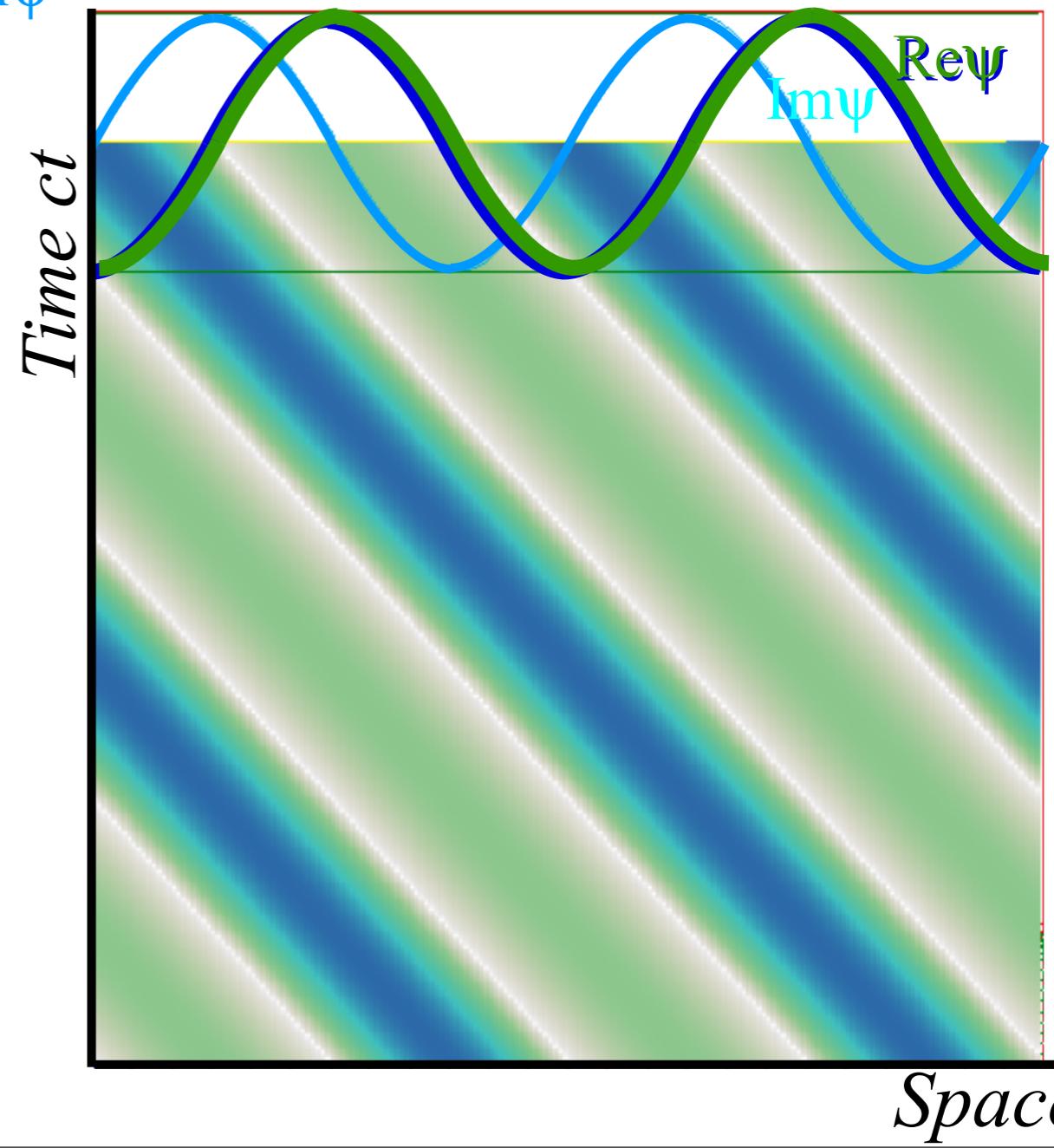
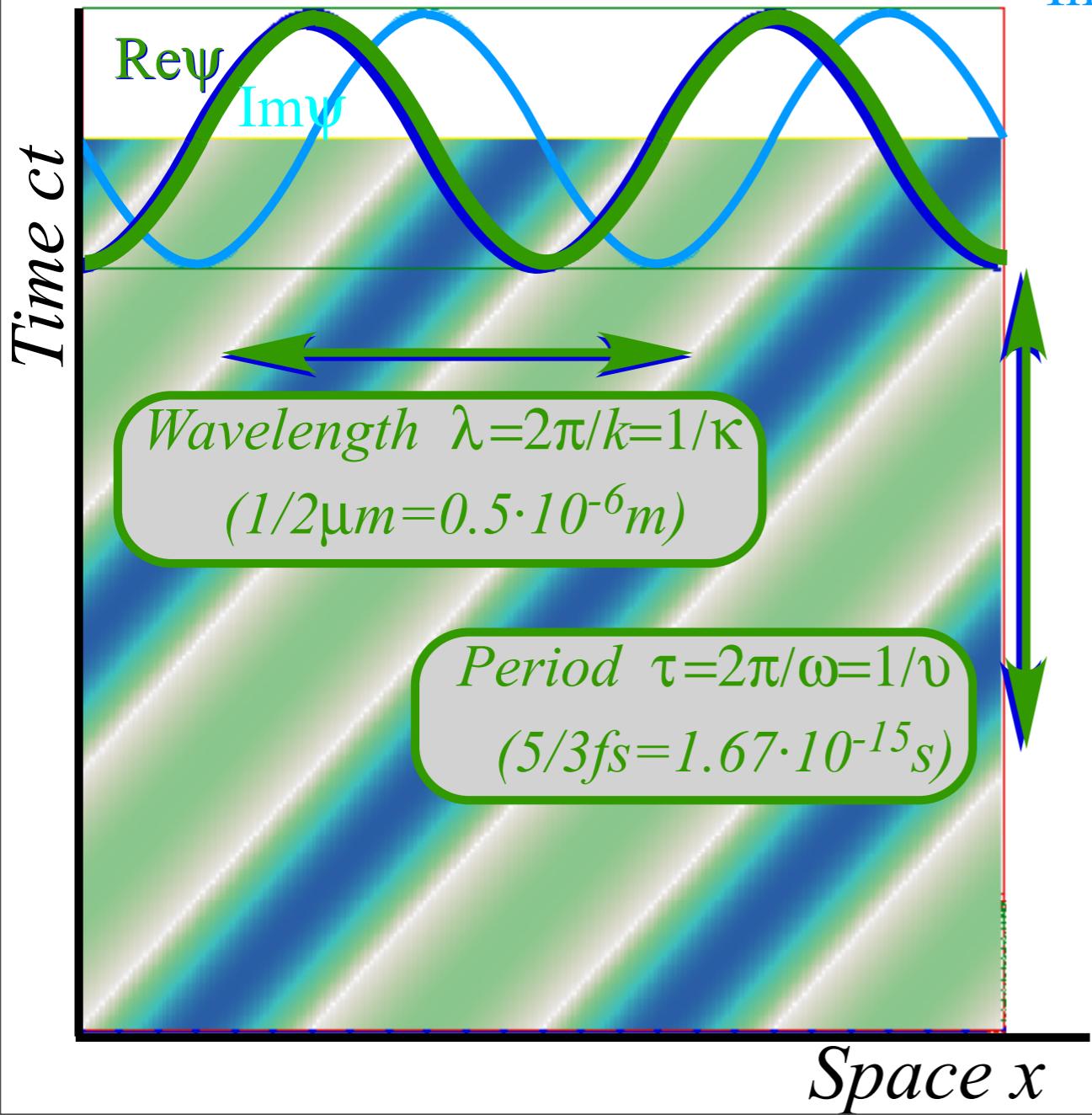
Carla's laser

$$k = -2$$

$$\omega = 2c$$



Reψ
Imψ



*Alice: OK, Bob.
We're gonna' hit
you from both
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Colliding 2CW laser beams

makes space-time coordinate frame

*Carla:
Look out, Bob!*

Right-moving wave $e^{i(kx-\omega t)}$

CW Dye-laser

600 THz

$k = +2$

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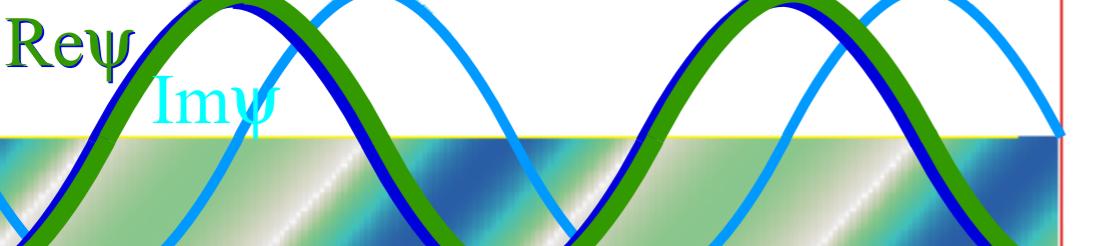
600 THz

Bob: Yikes!

$\text{Re}\psi$

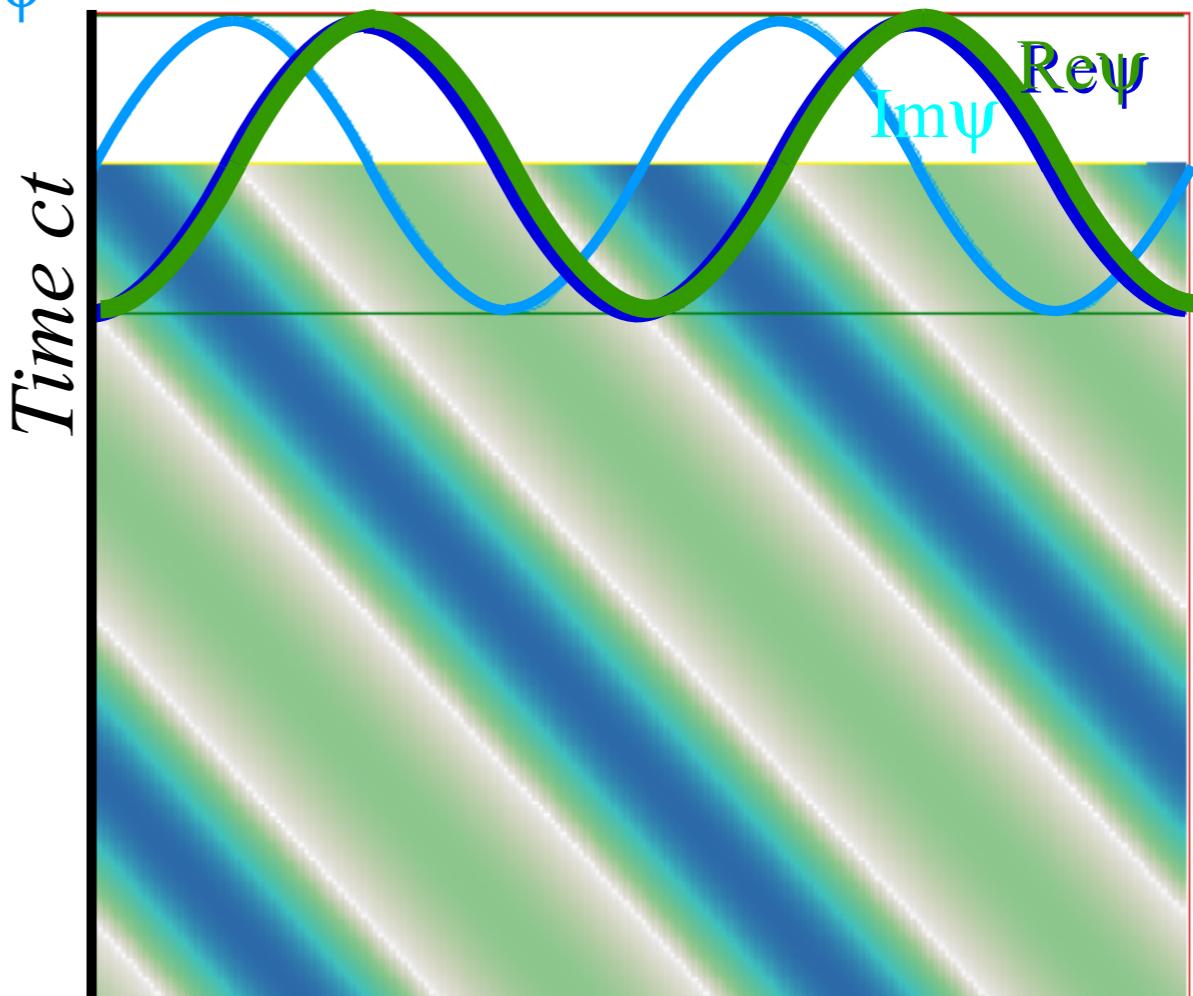
$\text{Im}\psi$

Time ct



*Wavelength $\lambda = 2\pi/k = 1/\kappa$
($1/2\mu\text{m} = 0.5 \cdot 10^{-6}\text{m}$)*

*Period $\tau = 2\pi/\omega = 1/\nu$
($5/3\text{fs} = 1.67 \cdot 10^{-15}\text{s}$)*

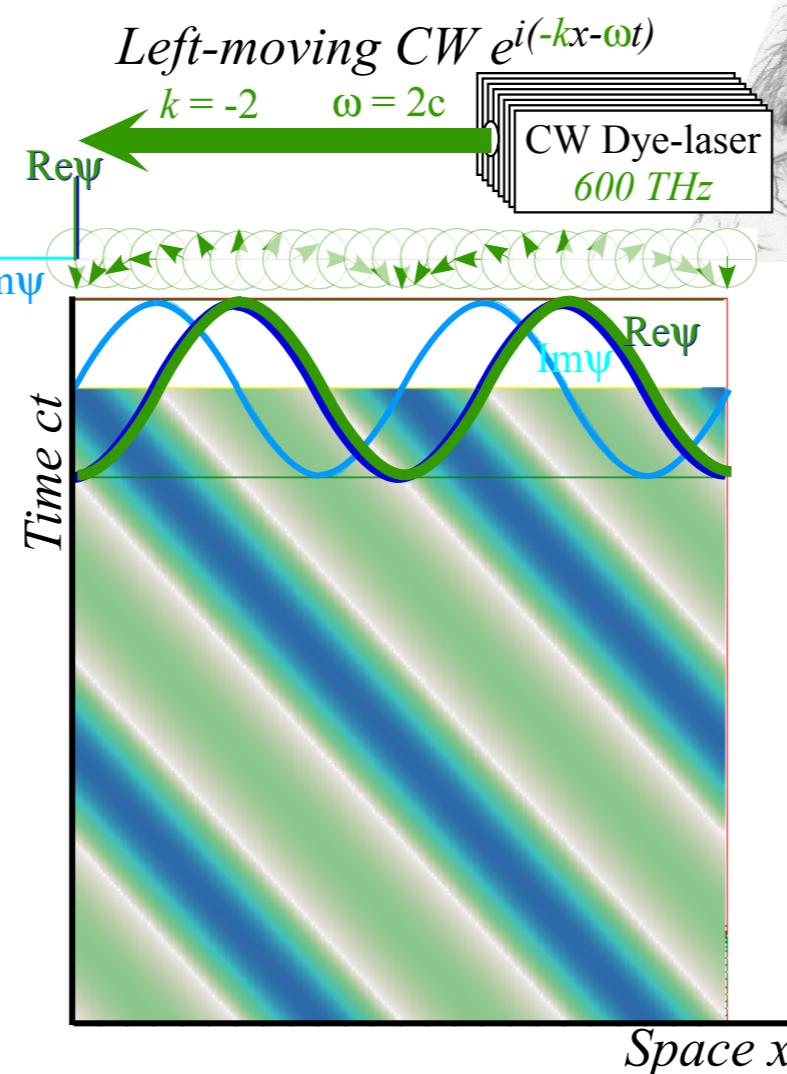
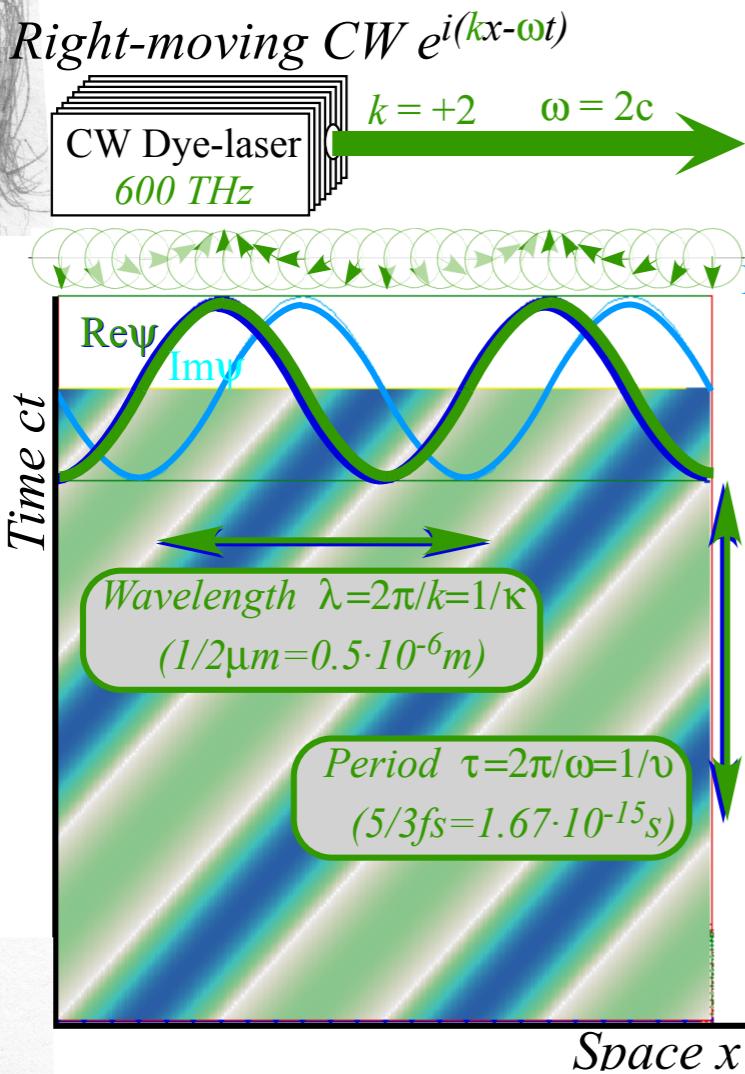


The result is the “simplest molecule” (a 2-γ “thing”)...

..with a *space-time frame* that eventually reveals relativistic/quantum matter-wave effects!

Space x

Space x



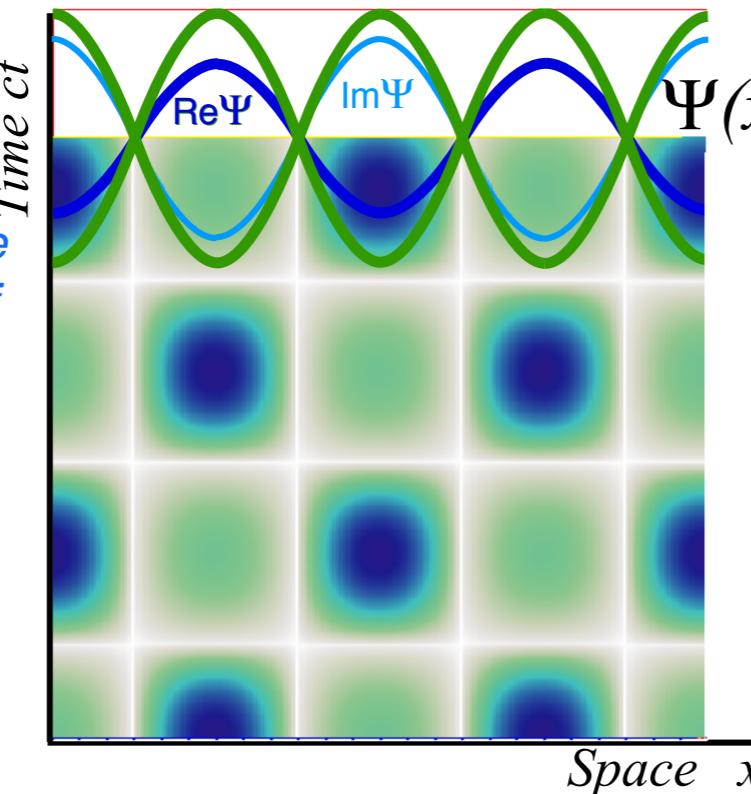
Carla:

Easy!

You get zeros of any wave-sum $e^{ia} + e^{ib}$ by factoring it into phase and group parts.

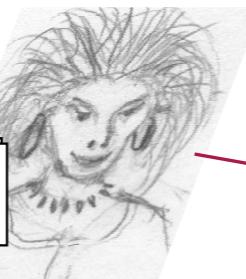
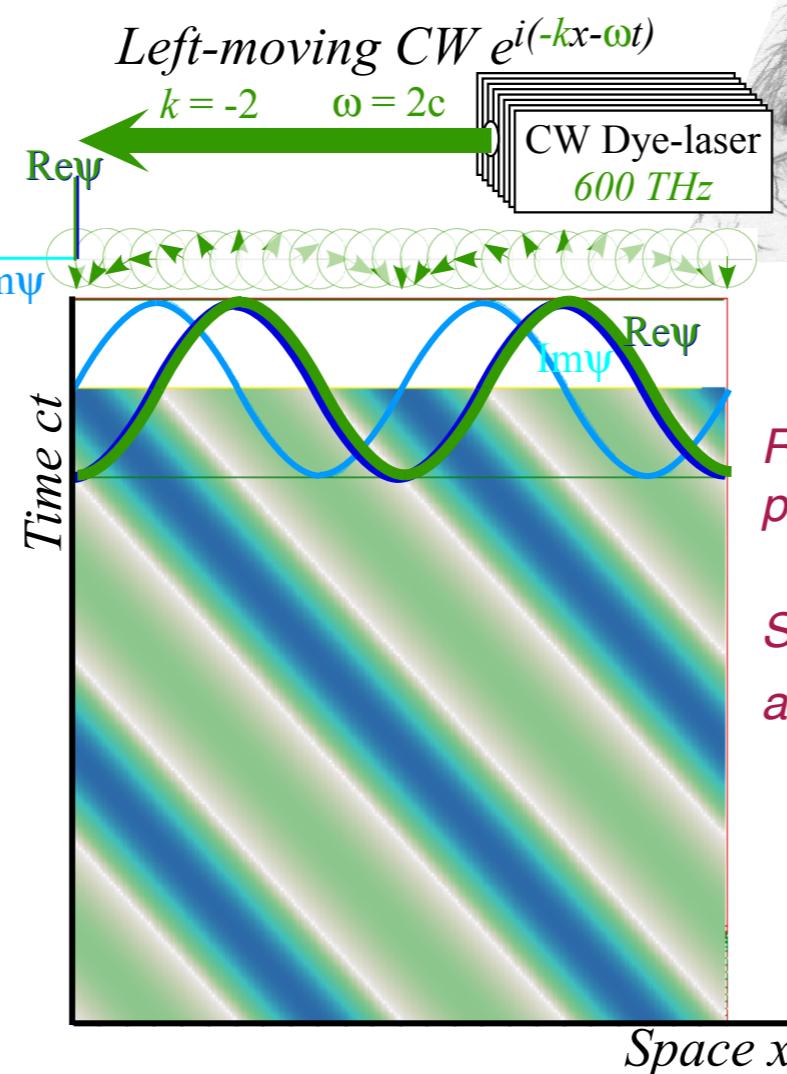
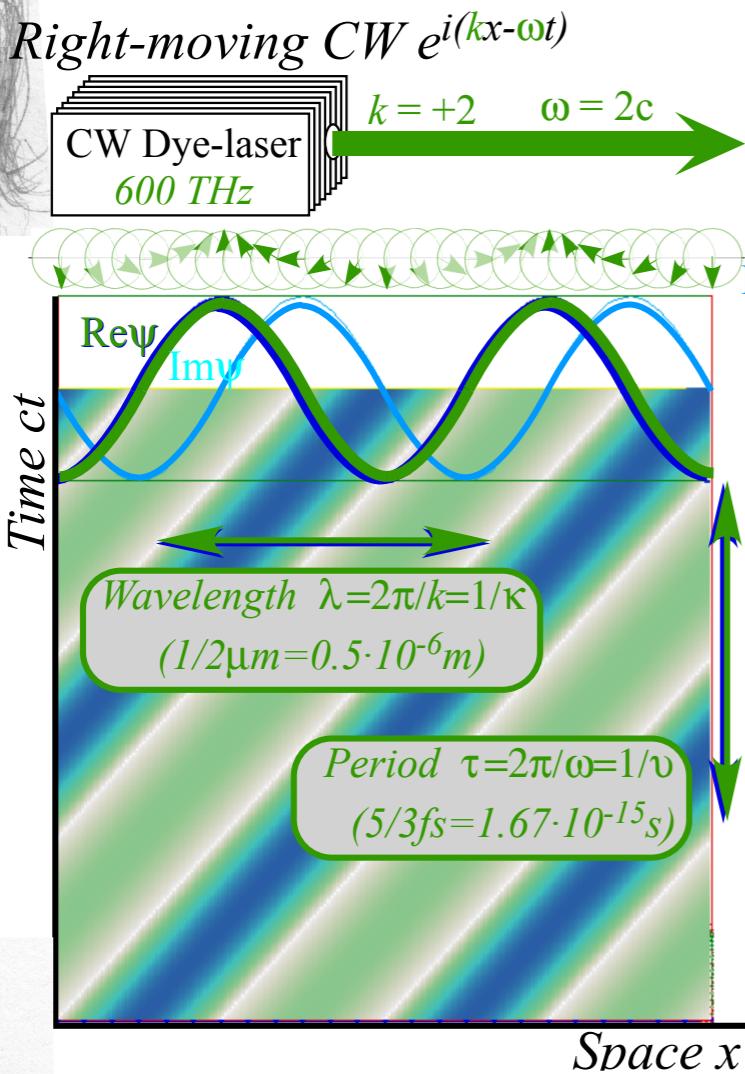
Bob:
Cool!
You guys
made me
a space-time
graph out of
real zeros.

How'd it
do that?



$$\Psi(x, t) = e^{i a} + e^{i b}$$

$kx - \omega t$ $-kx - \omega t$



Carla:

Easy!

You get zeros of any wave-sum $e^{ia} + e^{ib}$ by factoring it into phase and group parts.

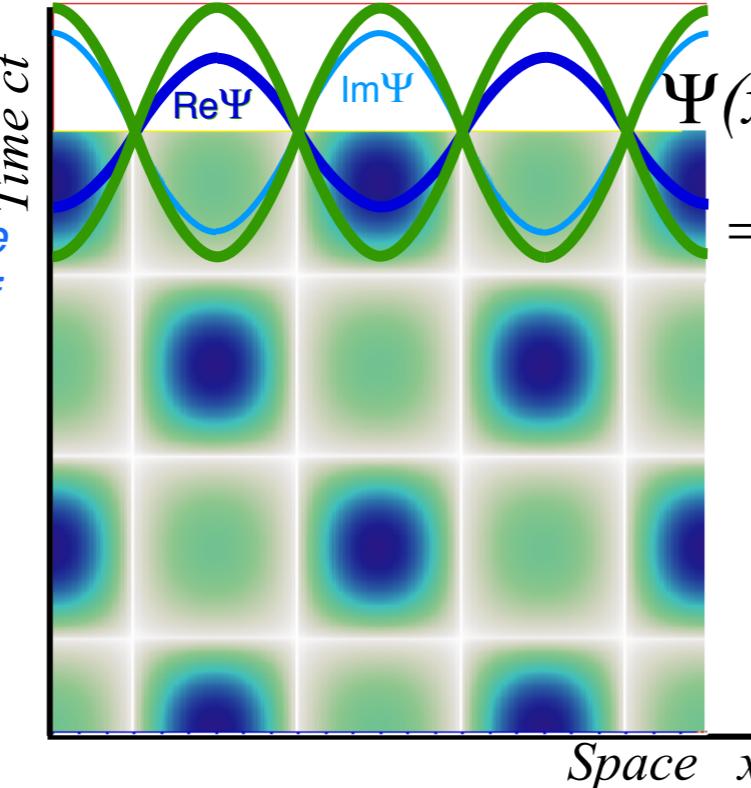
Remember your algebra? Exponents of products add.

So, half-sum $\frac{a+b}{2}$ plus half-diff $\frac{a-b}{2}$ gives a , and half-sum $\frac{a+b}{2}$ minus half-diff $\frac{a-b}{2}$ gives b .

Bob:

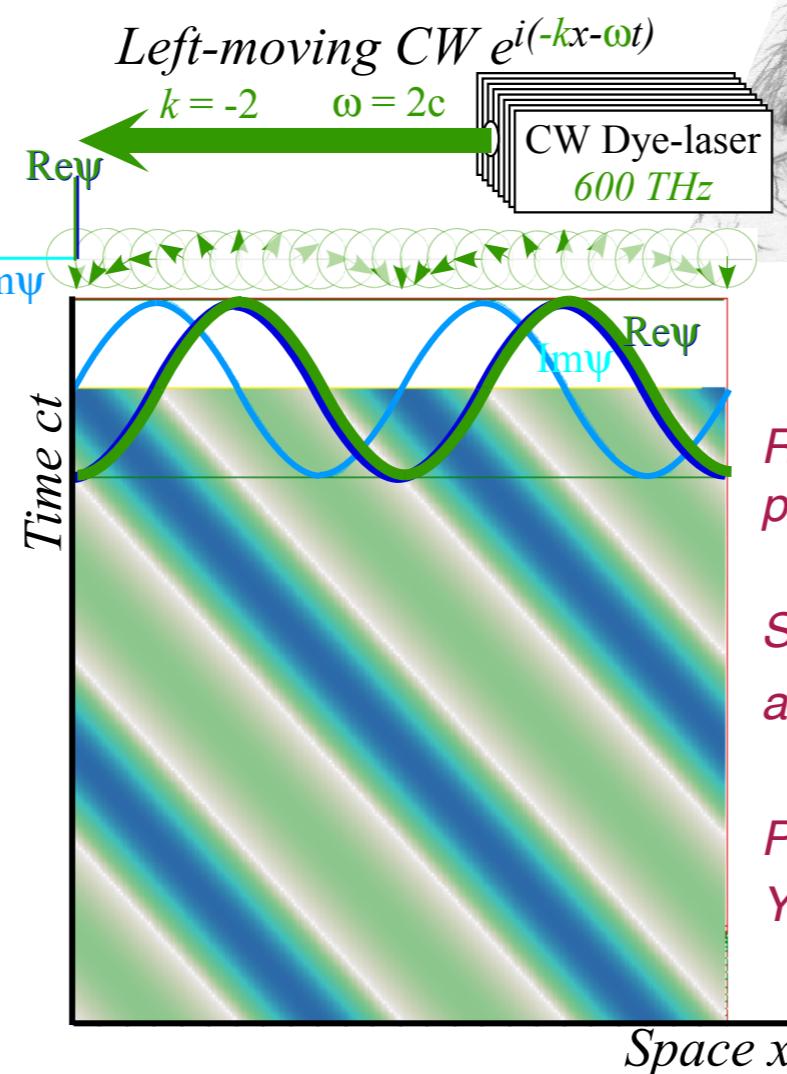
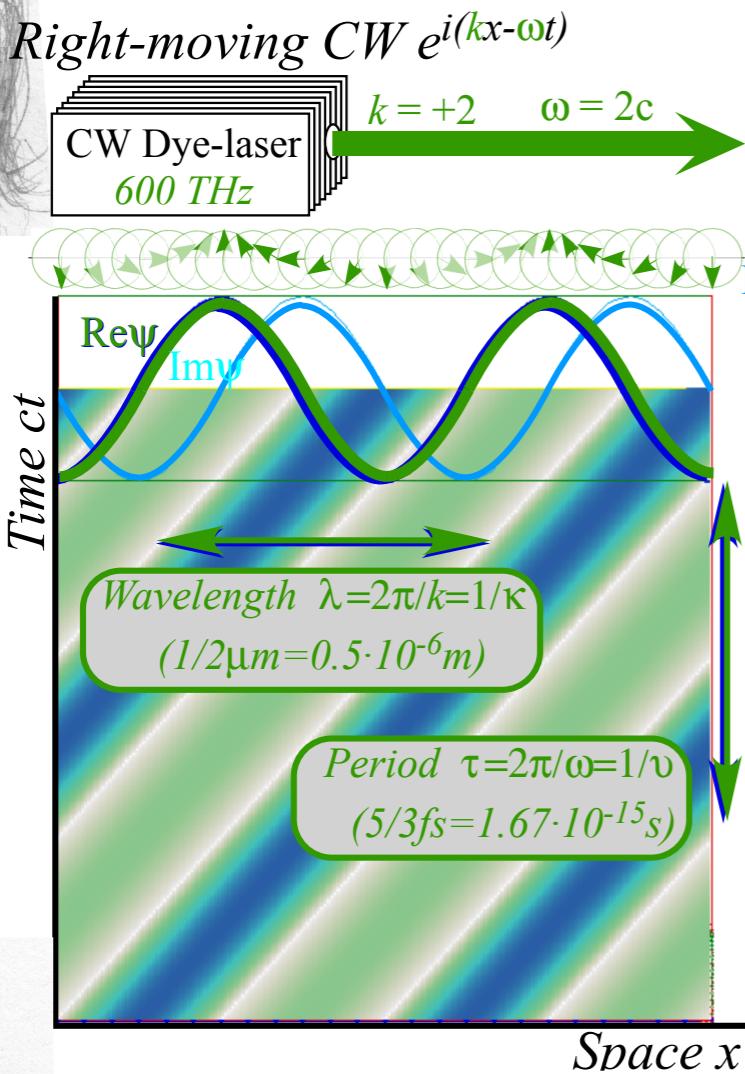
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$$\Psi(x, t) = e^{ia} + e^{ib}$$

$$= e^{i\frac{a+b}{2}} (e^{i\frac{a-b}{2}} + e^{-i\frac{a-b}{2}})$$



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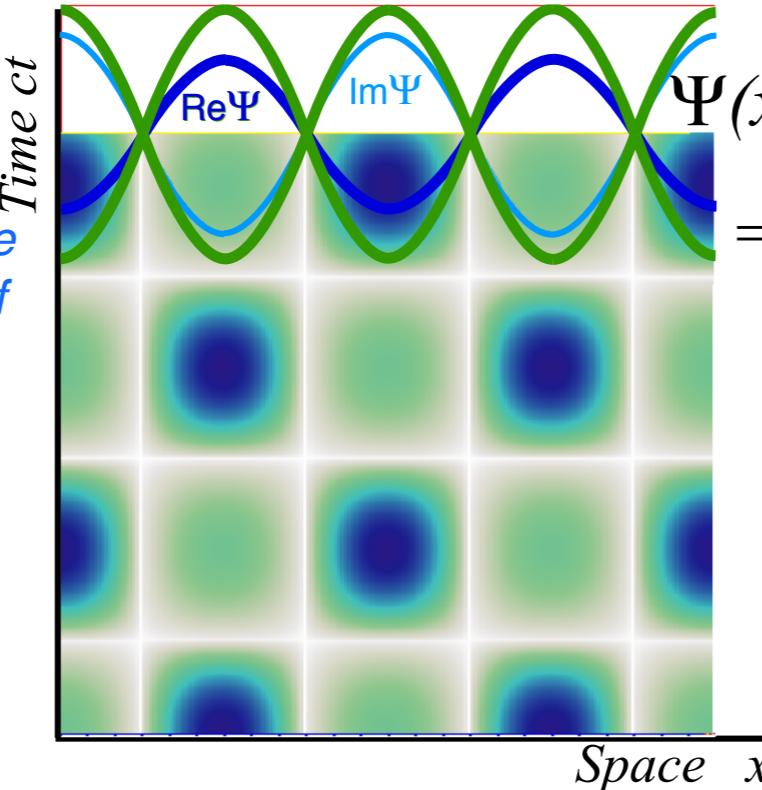
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Presto!
You factor $e^{ia} + e^{ib}$ into $e^{i\frac{a+b}{2}} \left(e^{i\frac{a-b}{2}} + e^{-i\frac{a-b}{2}} \right)$

Bob:

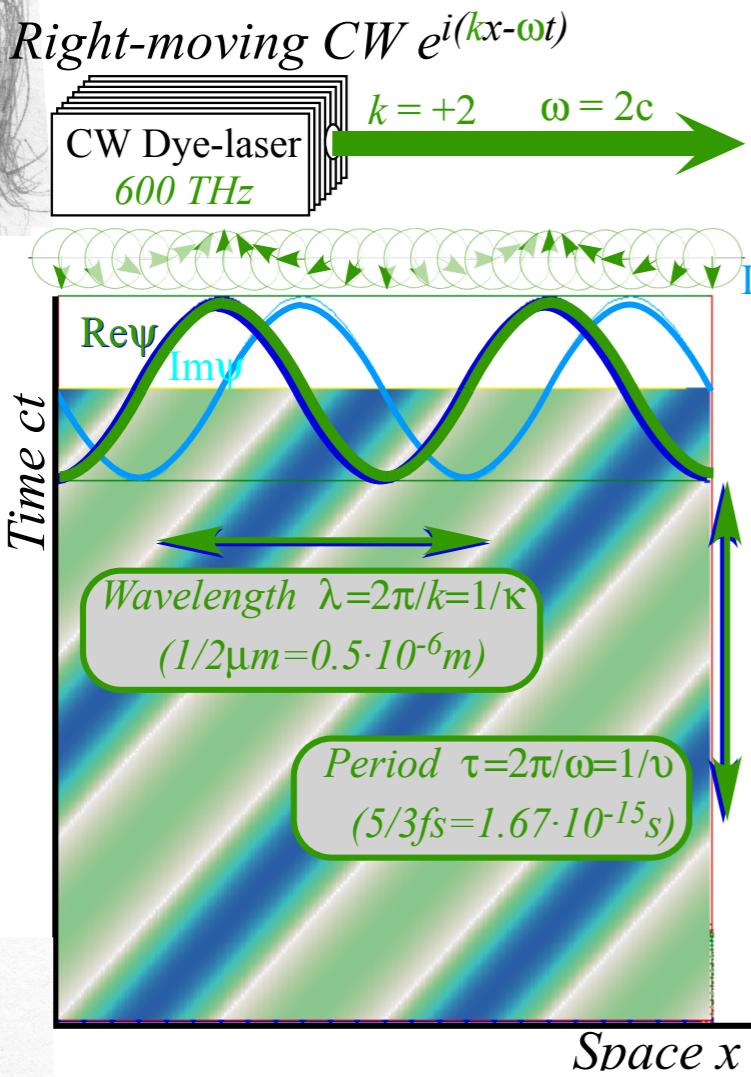
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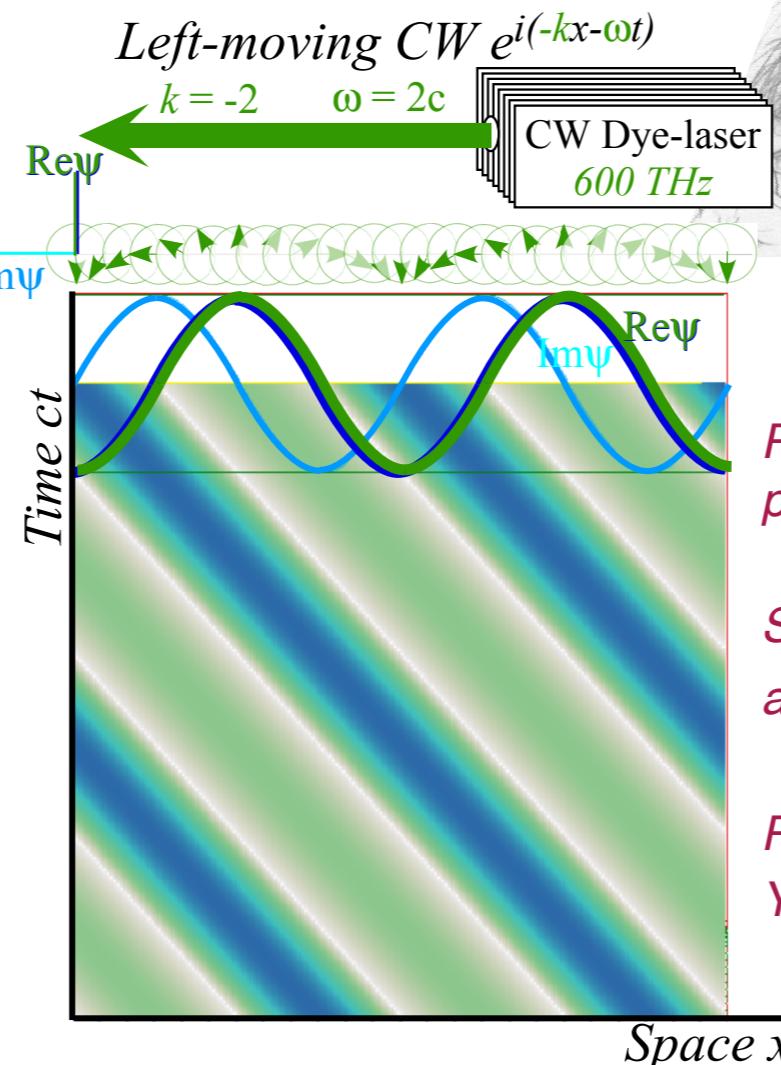
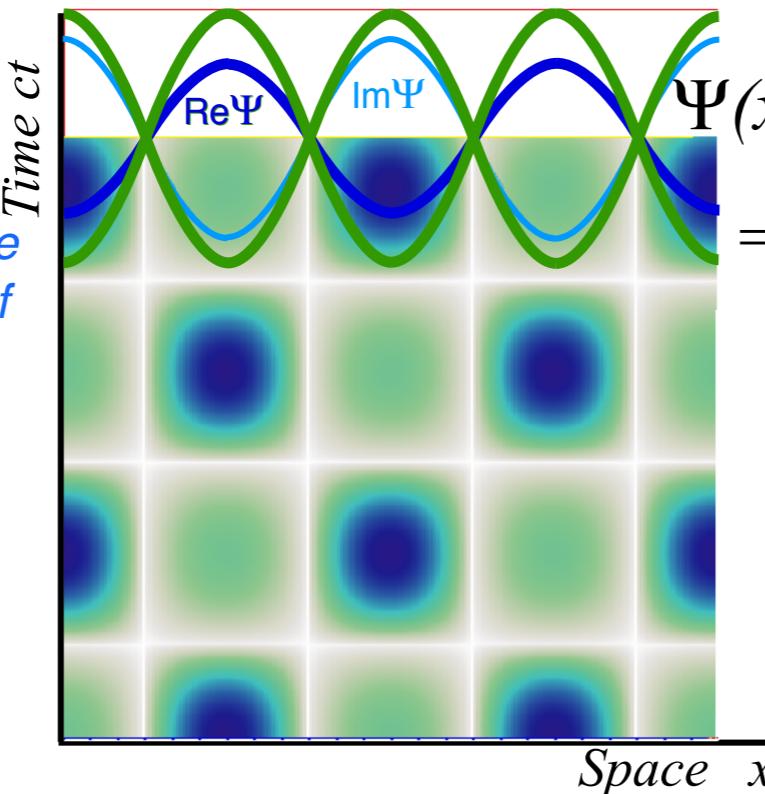
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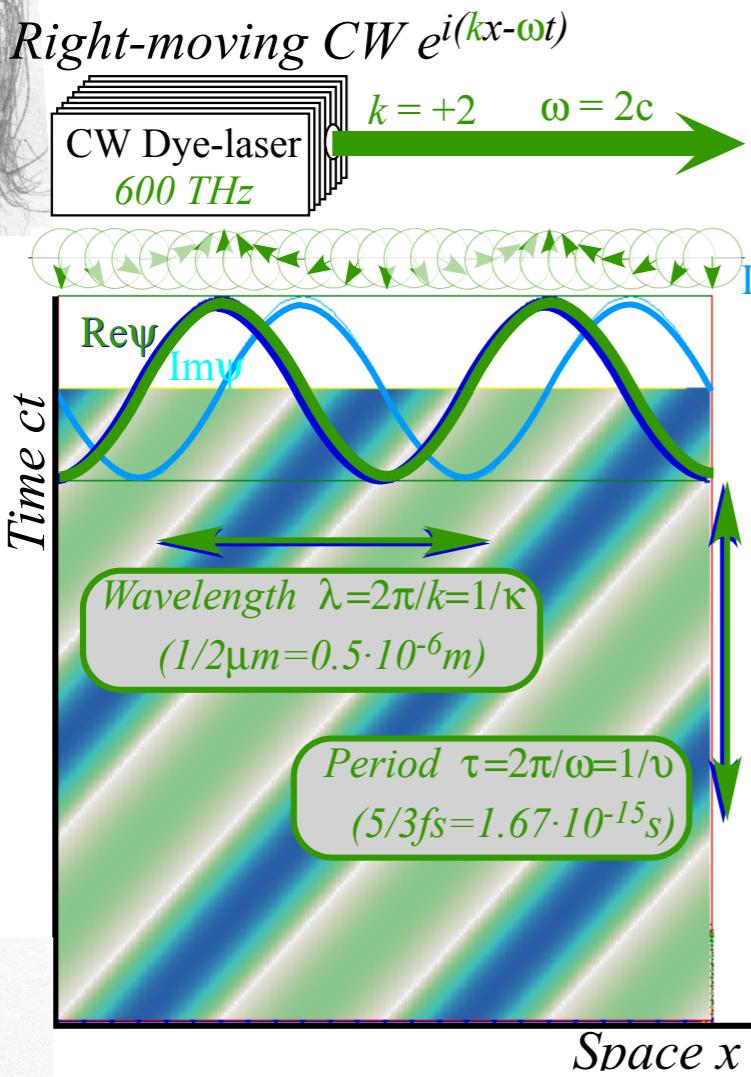
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Alice 1CW phase: $a = kx - \omega t$

Carla 1CW phase: $b = -kx - \omega t$

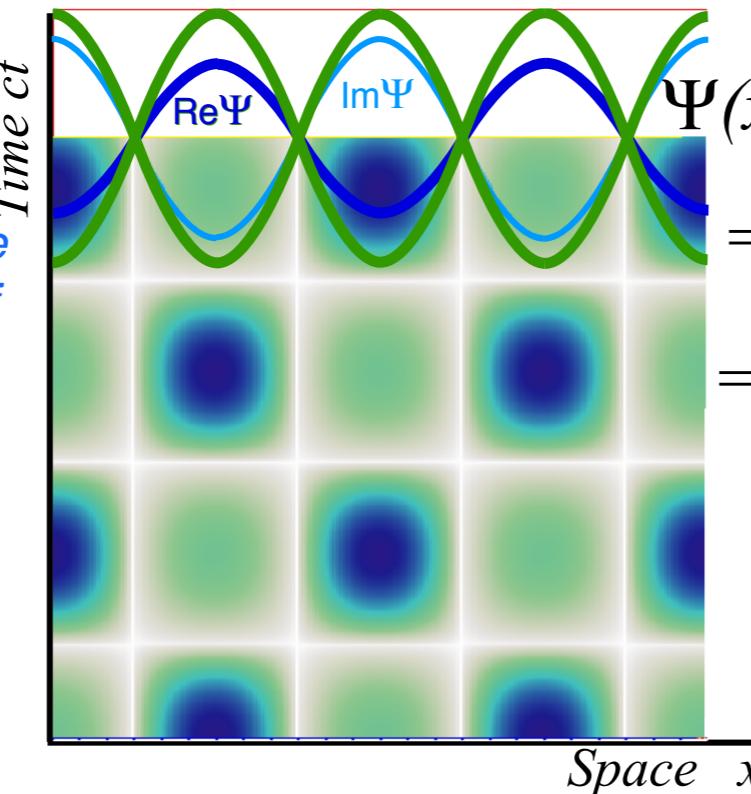
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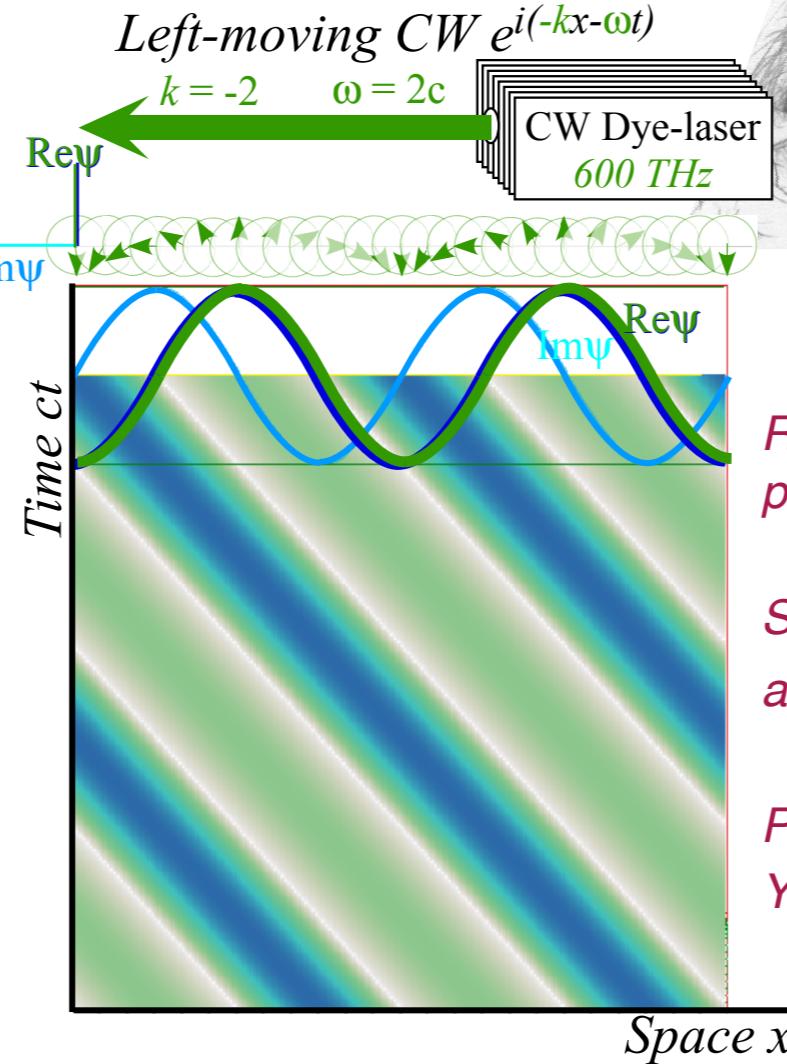
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$$\Psi(x, t) = e^{ia} + e^{ib}$$

$$= e^{i\frac{a+b}{2}} (e^{i\frac{a-b}{2}} + e^{-i\frac{a-b}{2}})$$

$$= e^{-i\omega t} (e^{ikx} + e^{-ikx})$$



Carla:

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Presto!
 You factor $e^{ia} + e^{ib}$ into $e^{i\frac{a+b}{2}} \left(e^{i\frac{a-b}{2}} + e^{-i\frac{a-b}{2}} \right)$

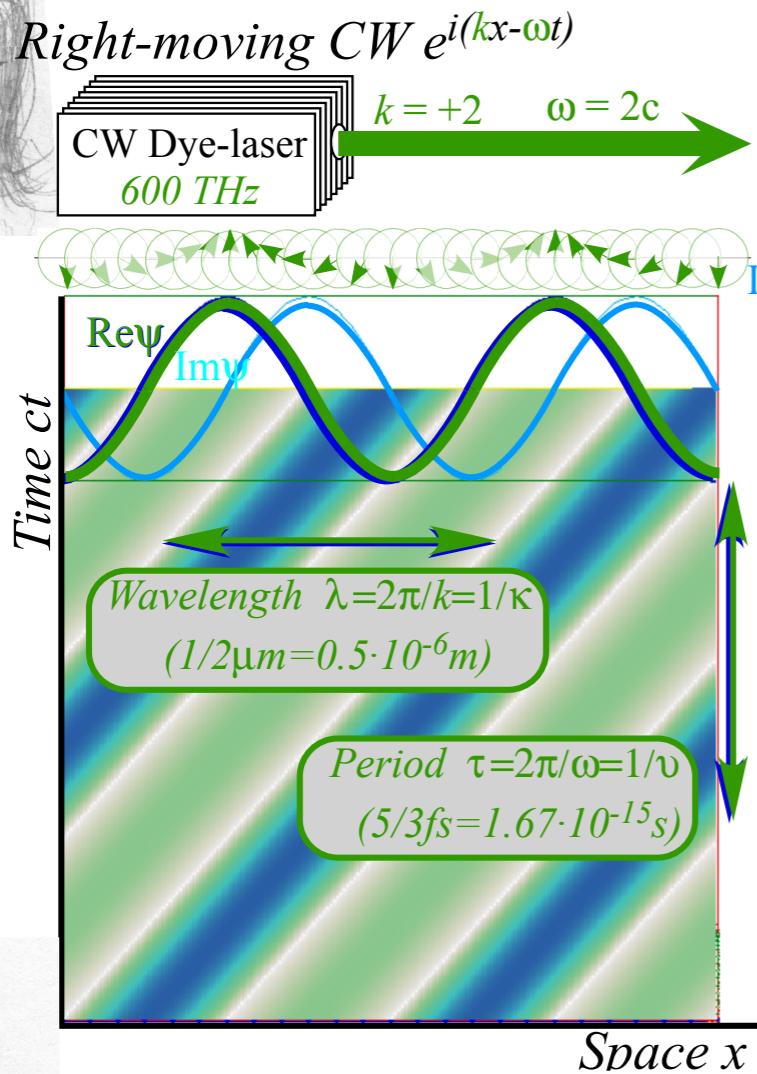
Alice 1CW phase: $a = kx - \omega t$

Carla 1CW phase: $b = -kx - \omega t$

Bob's 2CW Group-phase: $+k = \frac{a-b}{2}$

Group wave: $e^{-ikx} + e^{-ikx} = 2\cos{kx}$

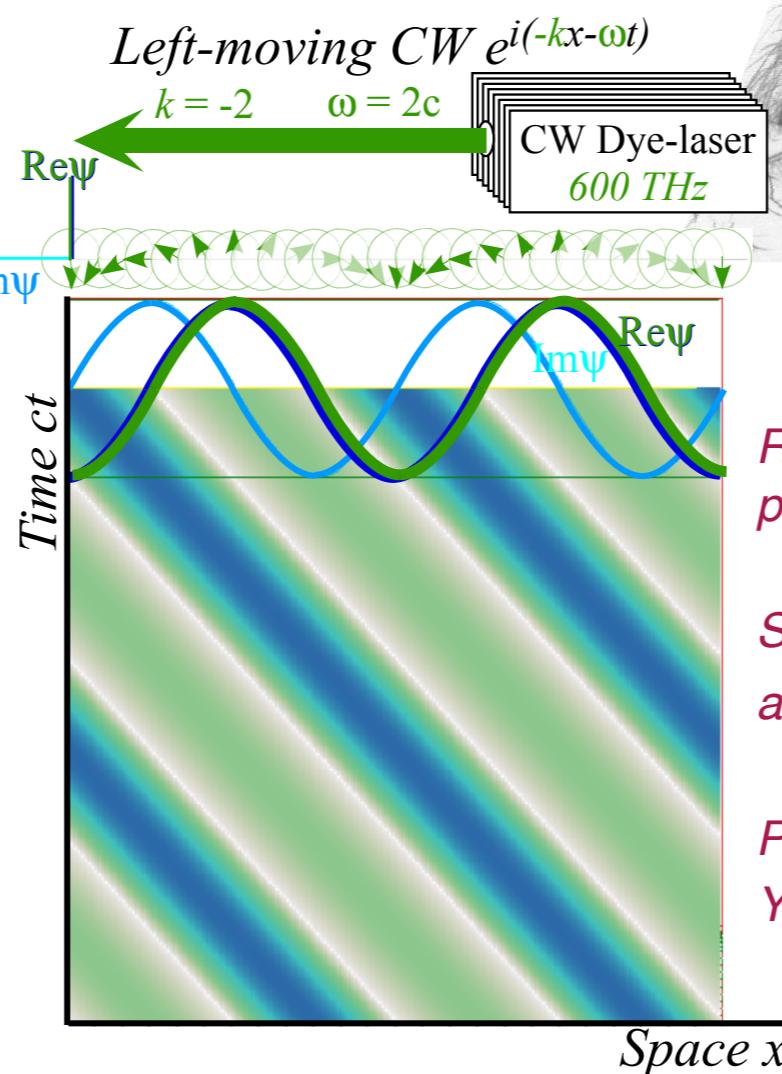
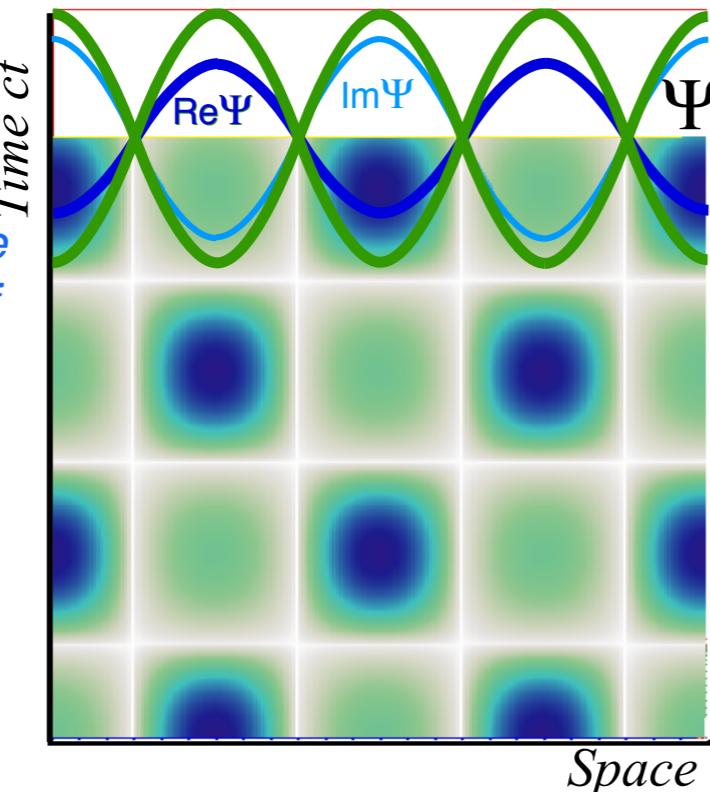
is standing wave (does not vary with time t)



Bob: Let's plot this in per-spacetime?!

Cool!
You guys
made me
a space-time
graph out of
real zeros.

How'd it
do that?



Carla:

Easy!

You get zeros of any wave-sum $e^{ia} + e^{ib}$ by factoring it into phase and group parts.

Remember your algebra? Exponents of products add.

So, half-sum $\frac{a+b}{2}$ plus half-diff $\frac{a-b}{2}$ gives a , and half-sum $\frac{a+b}{2}$ minus half-diff $\frac{a-b}{2}$ gives b .

Presto!
You factor $e^{ia} + e^{ib}$ into $e^{\frac{a+b}{2}} \left(e^{\frac{i(a-b)}{2}} + e^{-i\frac{a-b}{2}} \right)$

Alice 1CW phase: $a = kx - \omega t$

Carla 1CW phase: $b = -kx - \omega t$

Bob's 2CW Group-phase: $+k = \frac{a-b}{2}$
Wave

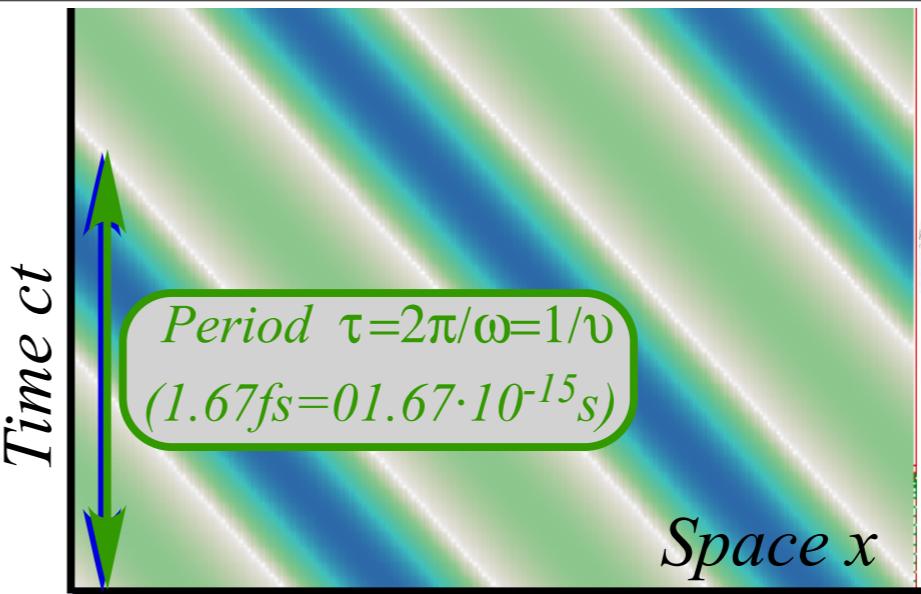
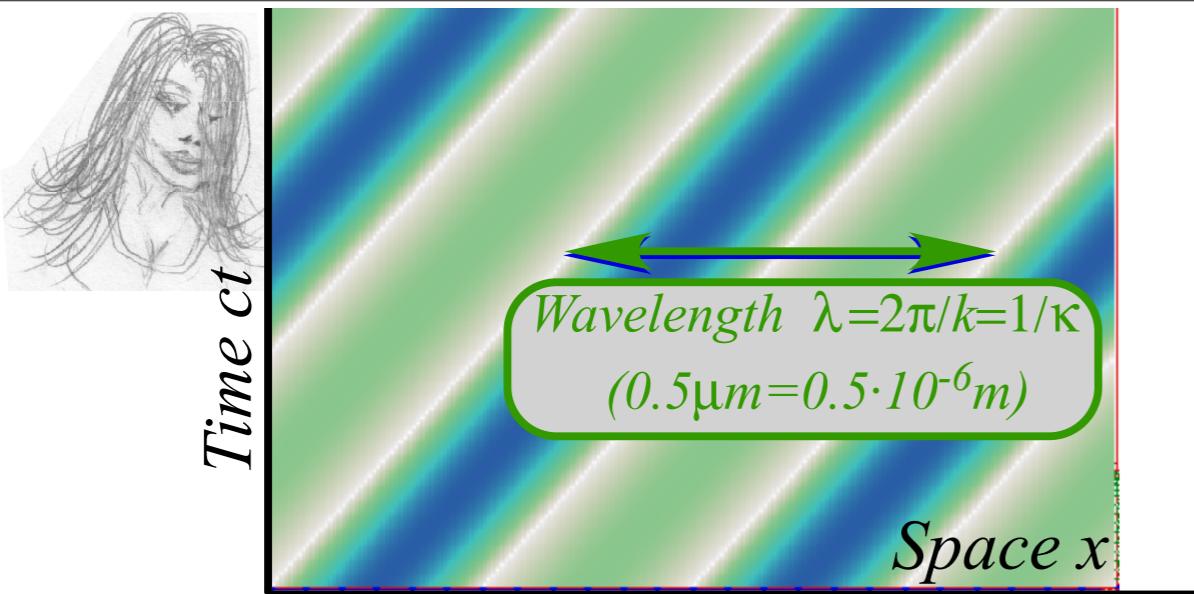
Group wave: $e^{-ikx} + e^{-ikx} = 2\cos kx$

is standing wave (does not vary with time t)

Bob's 2CW Phase-phase: $-\omega = \frac{a+b}{2}$
Wave

Phase wave real part: $\text{Re}(e^{-i\omega t}) = \cos(\omega t)$

is "instanton" wave (does not vary in space x)

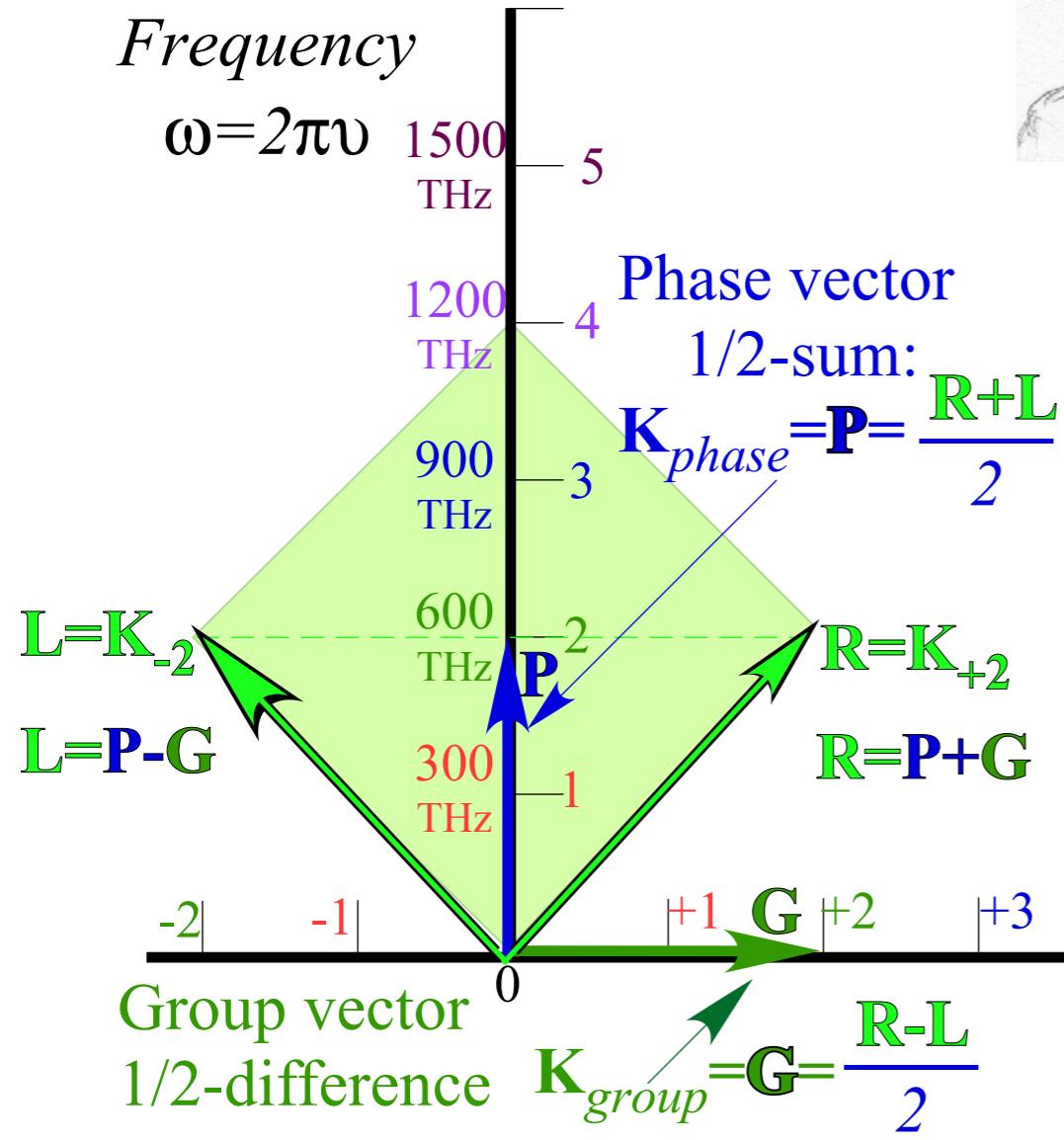


Carla:
OK, Bob!
It looks like a
baseball diamond
with
P at Pitcher's mound
and
G at the Grandstand*.
I'm on 1st base! (**R**)

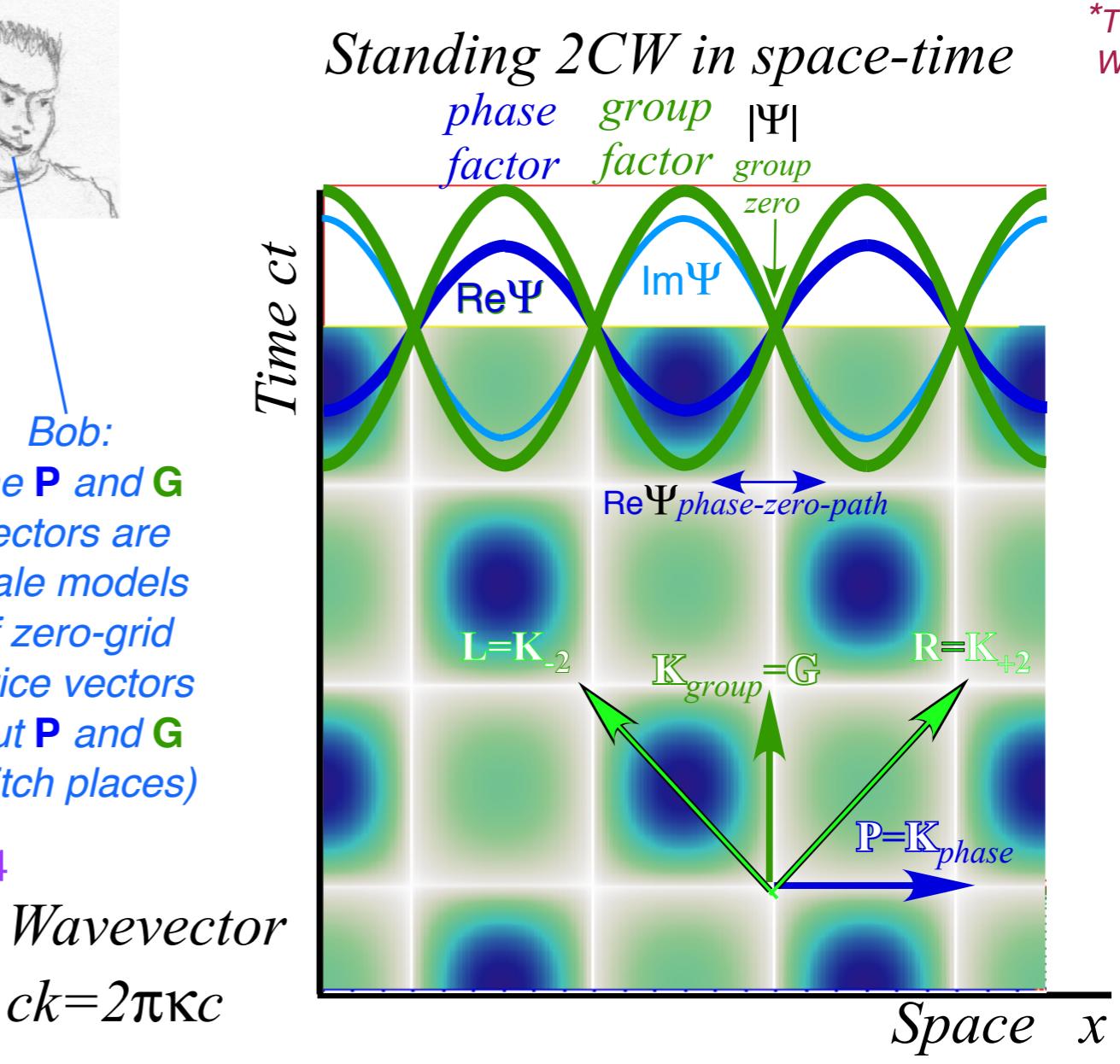
*Thanks,
Woody!

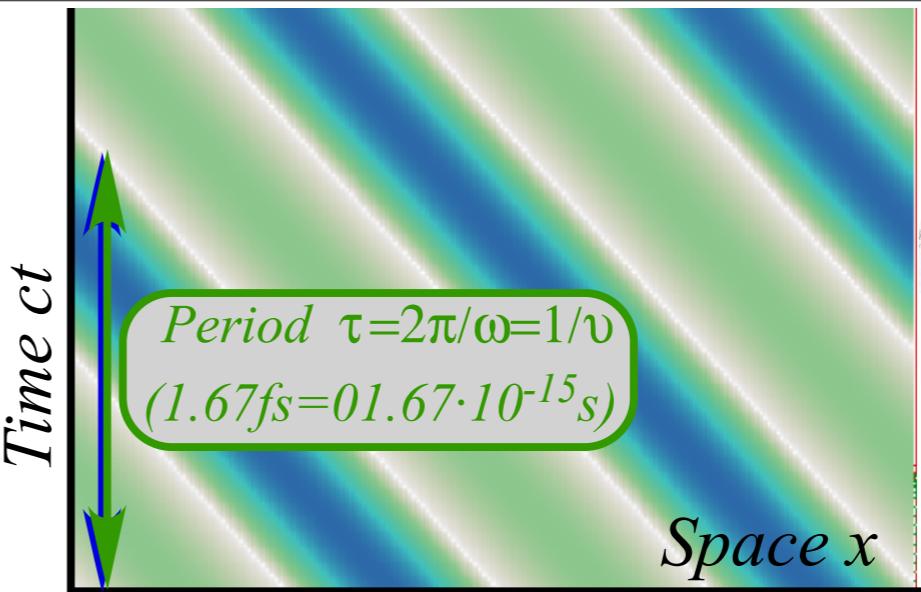
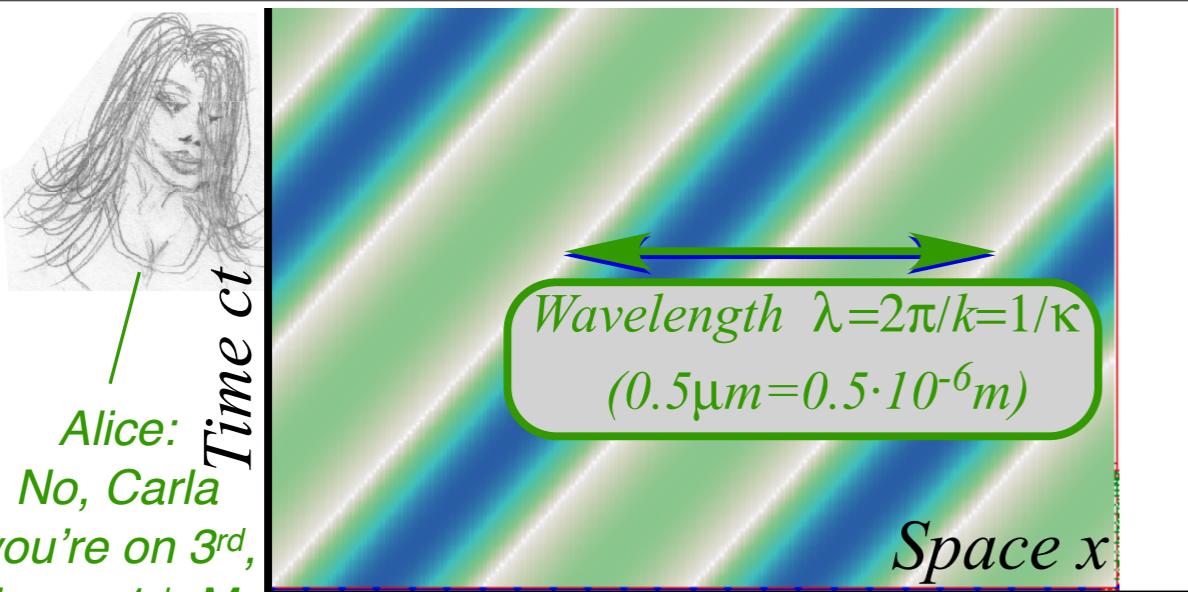
$$\Psi(x,t) = (e^{-i\omega t})(2\cos kx) = e^{i(kx-\omega t)} + e^{i(-kx-\omega t)}$$

Standing 2CW in per-space-time



Bob:
The **P** and **G**
vectors are
scale models
of zero-grid
lattice vectors
(but **P** and **G**
switch places)





*Thanks, Woody!

$$\Psi(x, t) = (e^{-i\omega t})(2 \cos kx) = e^{i(kx - \omega t)} + e^{i(-kx - \omega t)}$$

Standing 2CW in per-space-time

Frequency

$$\omega = 2\pi\nu$$

The (ν, κ) "Baseball Diamond"

3rd base
(Carla)

$$L = K_{-2}$$

$$L = P - G$$

Group vector

1/2-difference

1500 THz
2nd base
1200 THz
900 THz
600 THz
300 THz

Phase vector

1/2-sum:

$$K_{\text{phase}} = P = \frac{R+L}{2}$$

Pitcher's mound

1st base

(Alice)

$$R = K_{+2}$$

$$R = P + G$$

Grandstand

$$G = \frac{R-L}{2}$$

Group vector

1/2-difference

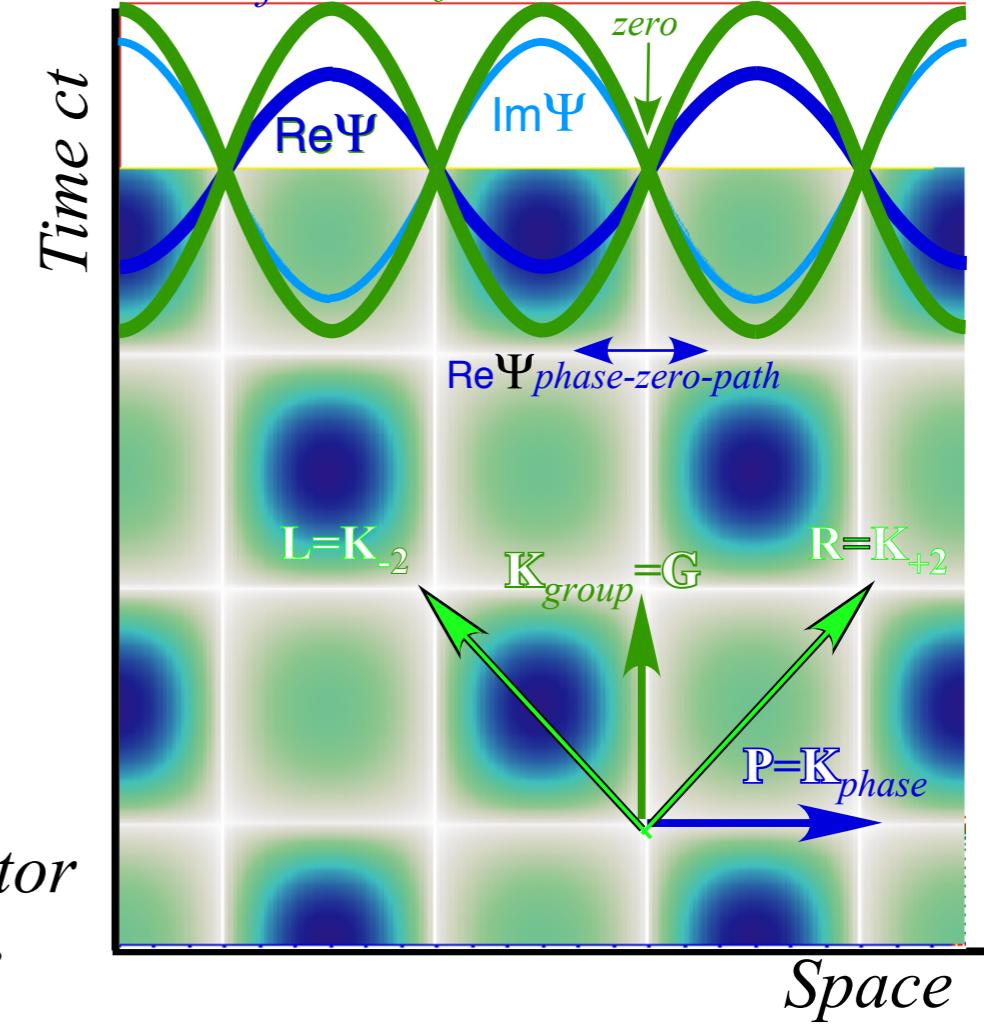


Bob: The P and G vectors are scale models of zero-grid lattice vectors (but P and G switch places)

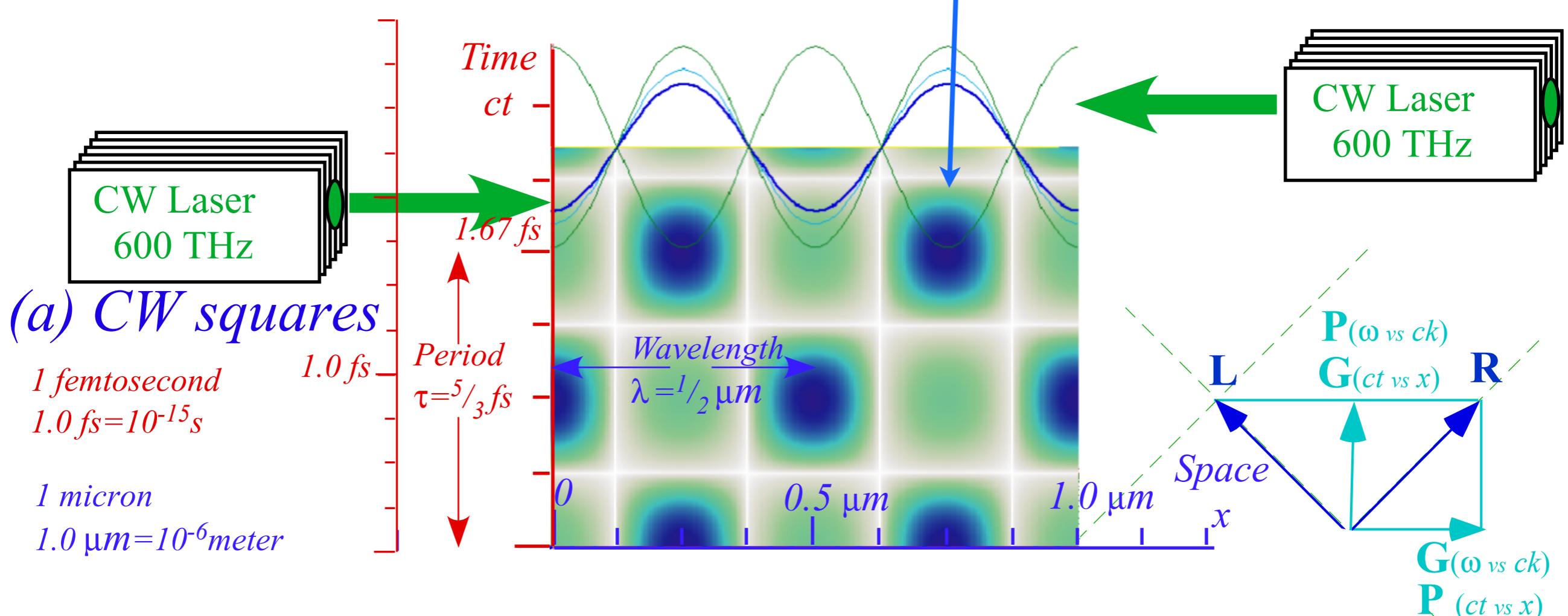
Wavevector
 $ck = 2\pi\kappa c$

Standing 2CW in space-time

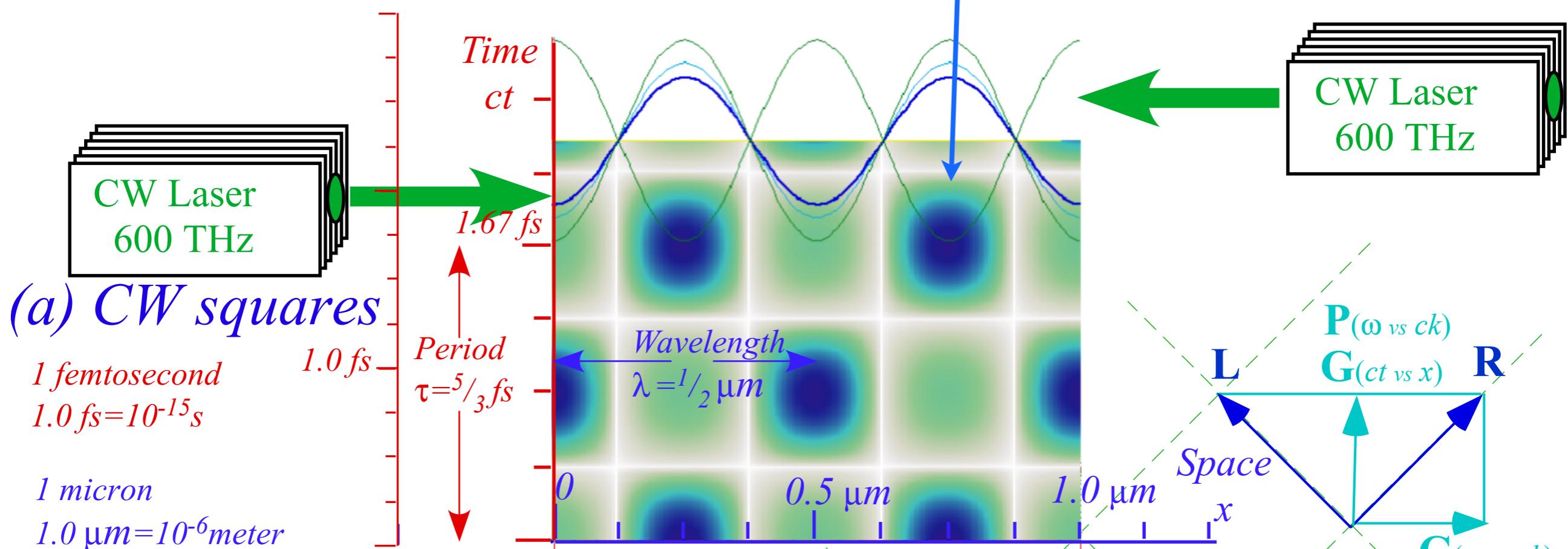
phase factor group factor |Ψ|
group zero



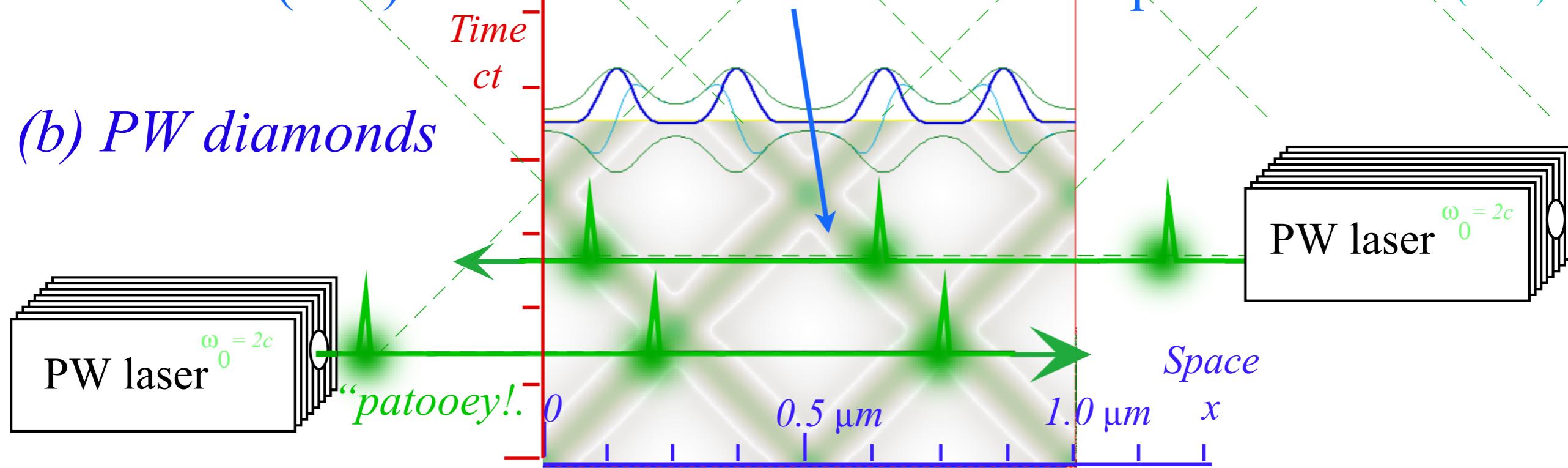
Continuous Waves (CW) trace “Cartesian squares” in space-time



Continuous Waves (CW) trace “Cartesian squares” in space-time



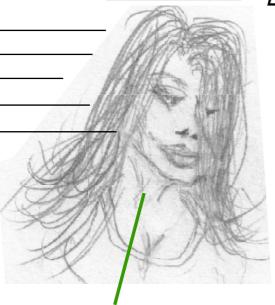
Pulse Waves (PW) trace “baseball diamonds” in space-time



Right-directed 1CW $e^{i(\underline{k}_4 x - \omega_4 t)}$

$$k_4 = +4 \quad \omega_4 = 4c$$

CW green-laser
600 THz Doppler blue shifted
to 1200THz



Left-directed 1CW $e^{i(\underline{k}_{-1} x - \omega_{-1} t)}$

$$k_{-1} = -1$$

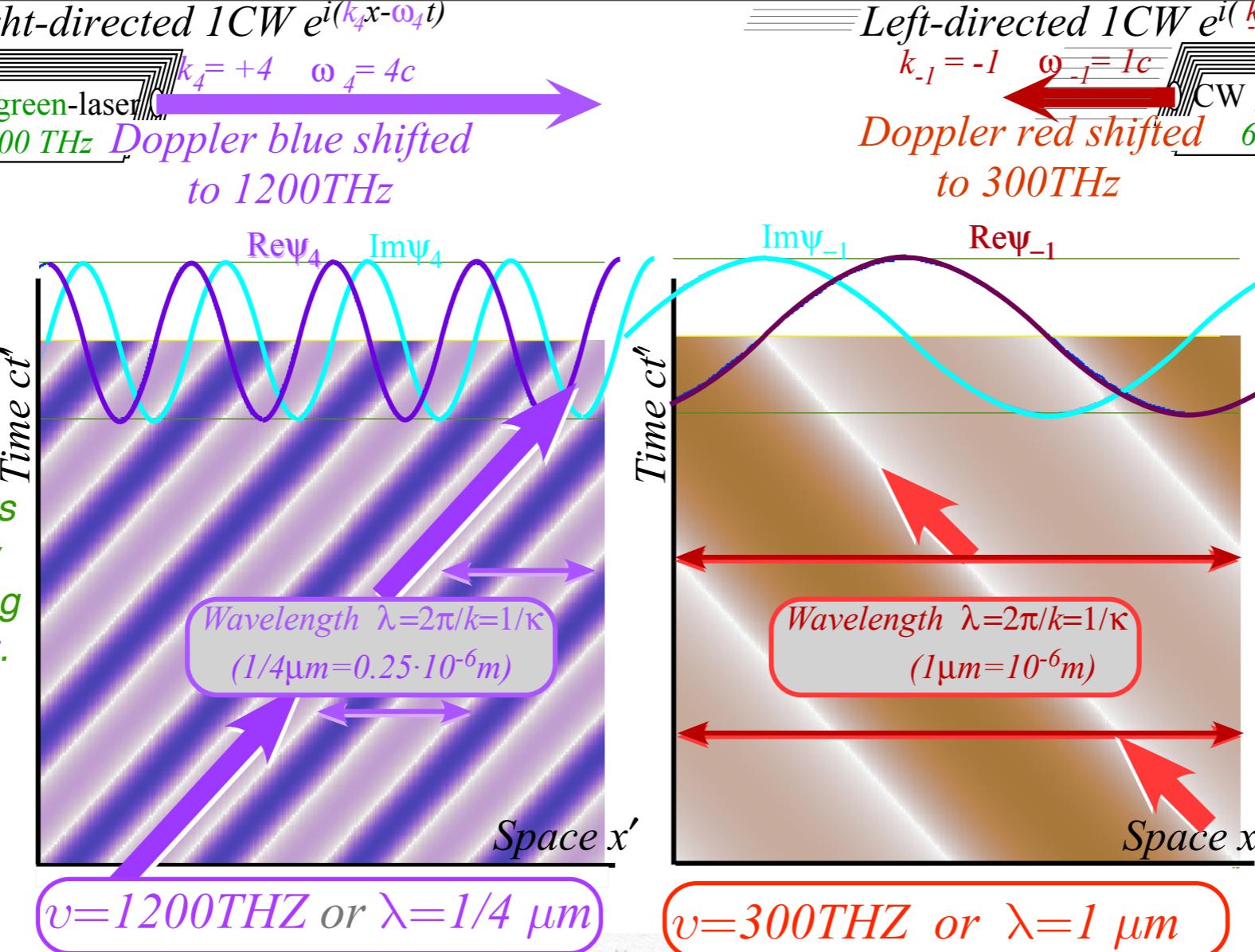
$$\omega_{-1} = 1c$$

CW green-laser
600 THz Doppler red shifted
to 300THz

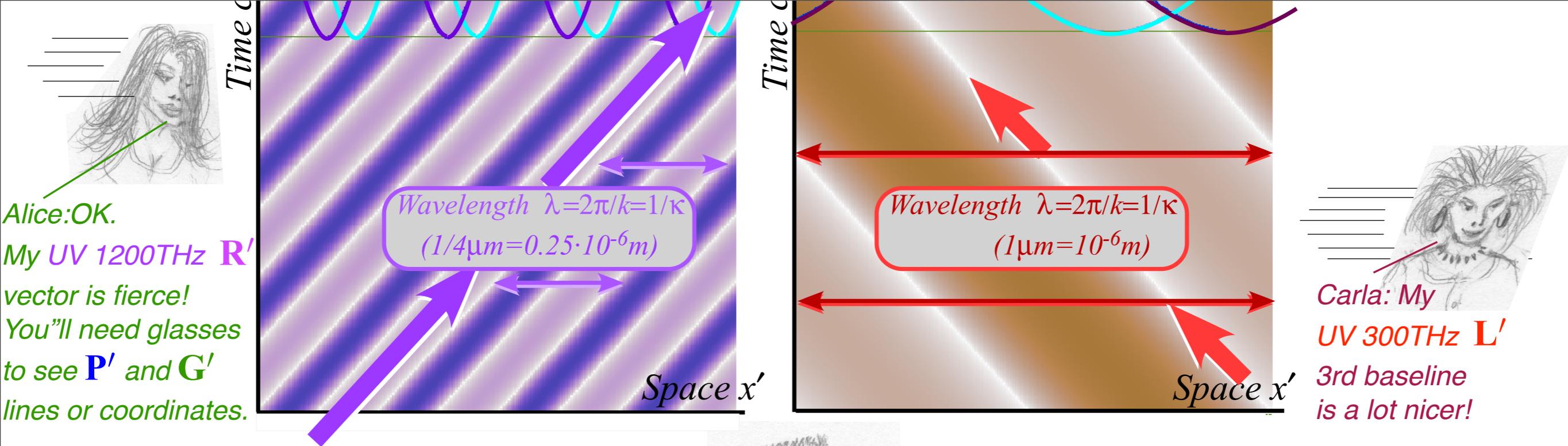


Alice:

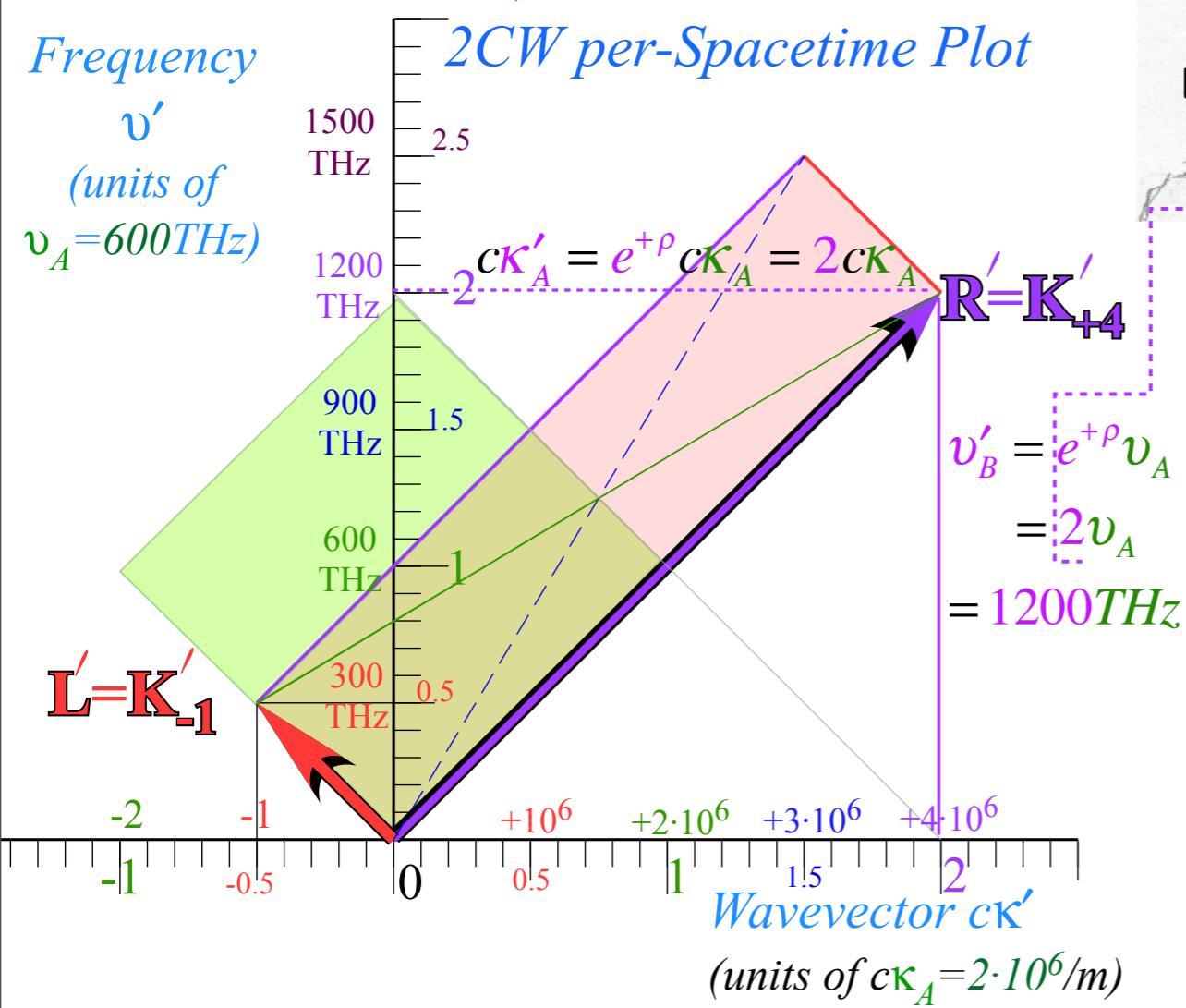
Now our 600THz lasers move left-to-right. My 600THz laser is blasting you with UV 1200THz. Carla's 600THz gives you a nice infrared 300THz.

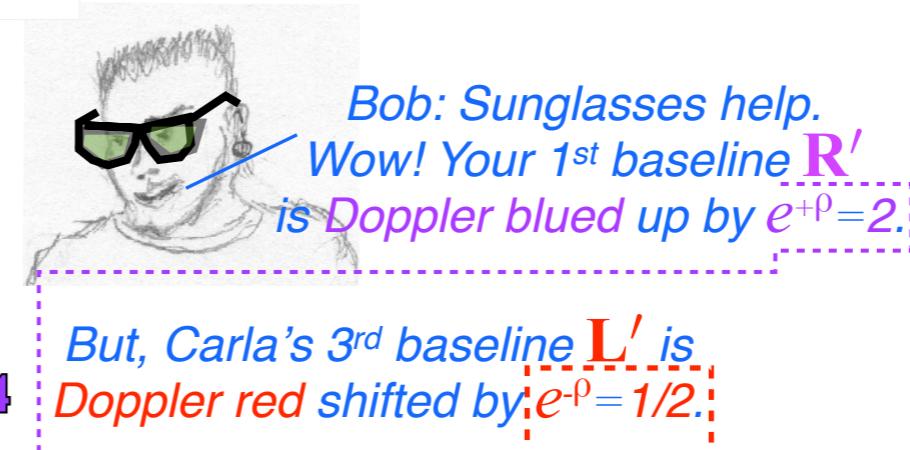
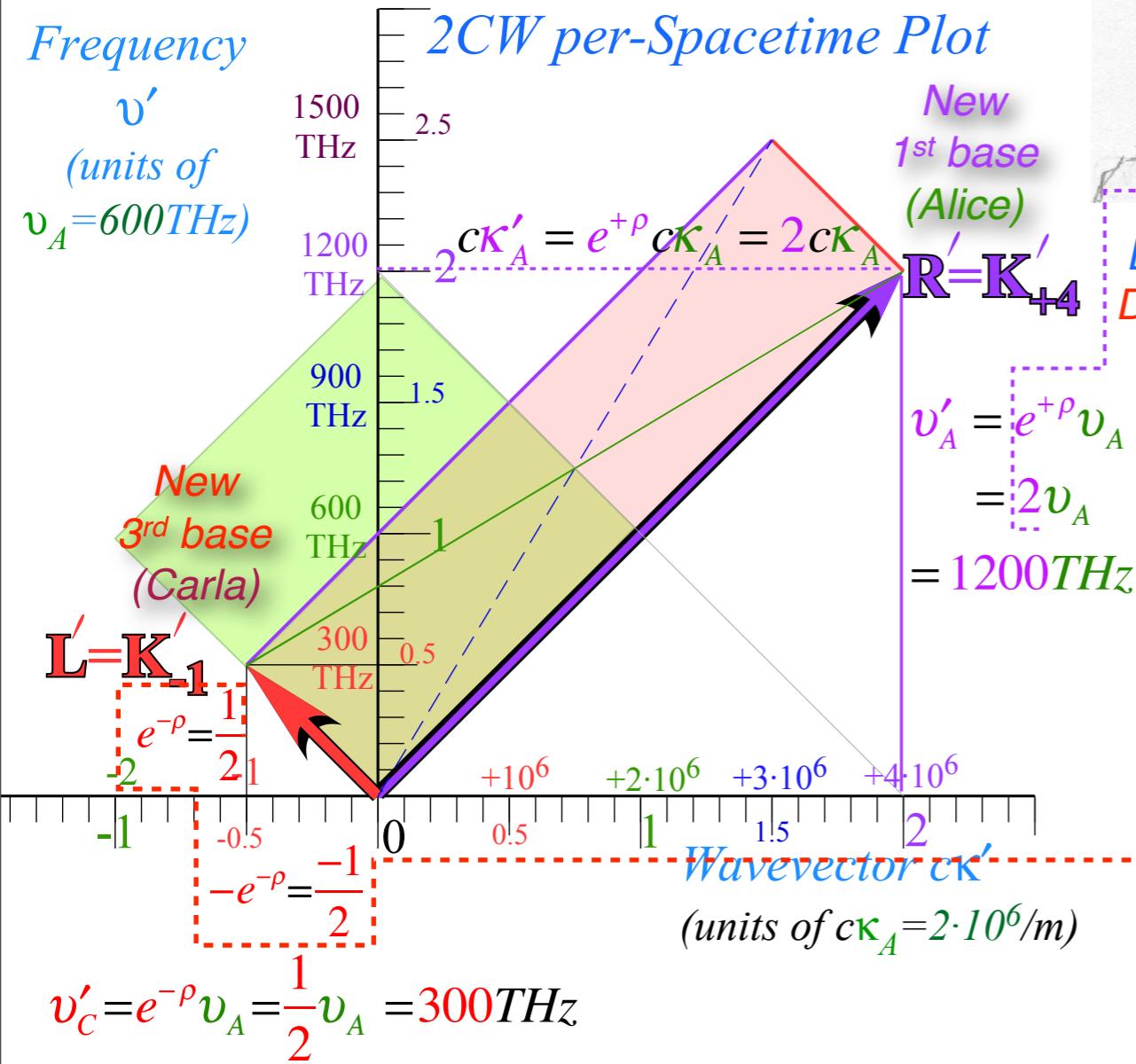
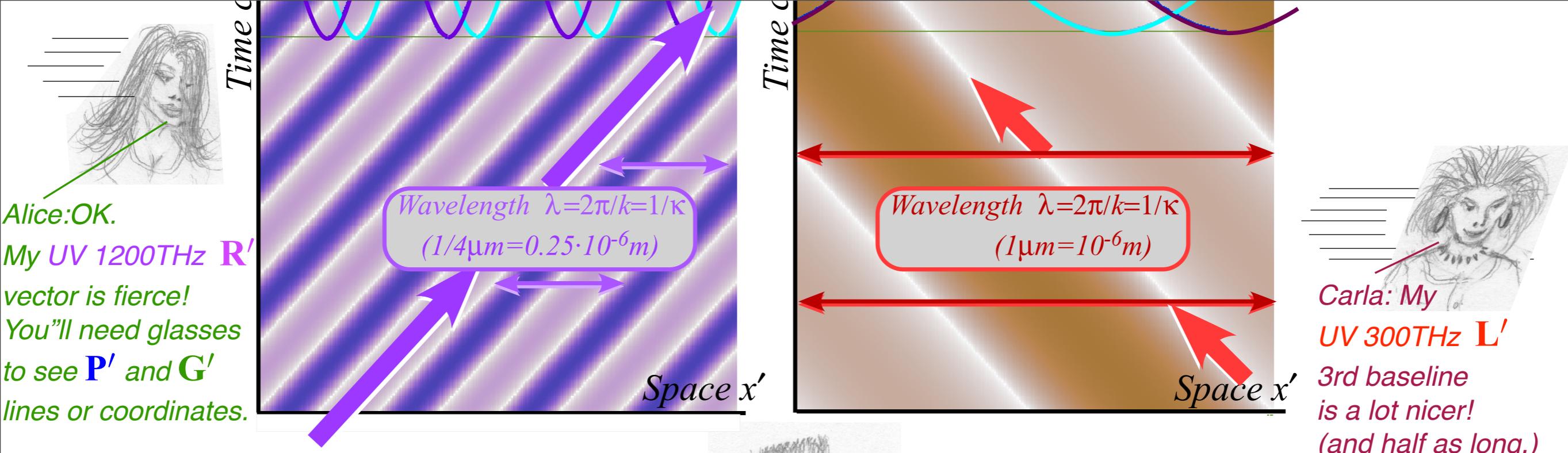


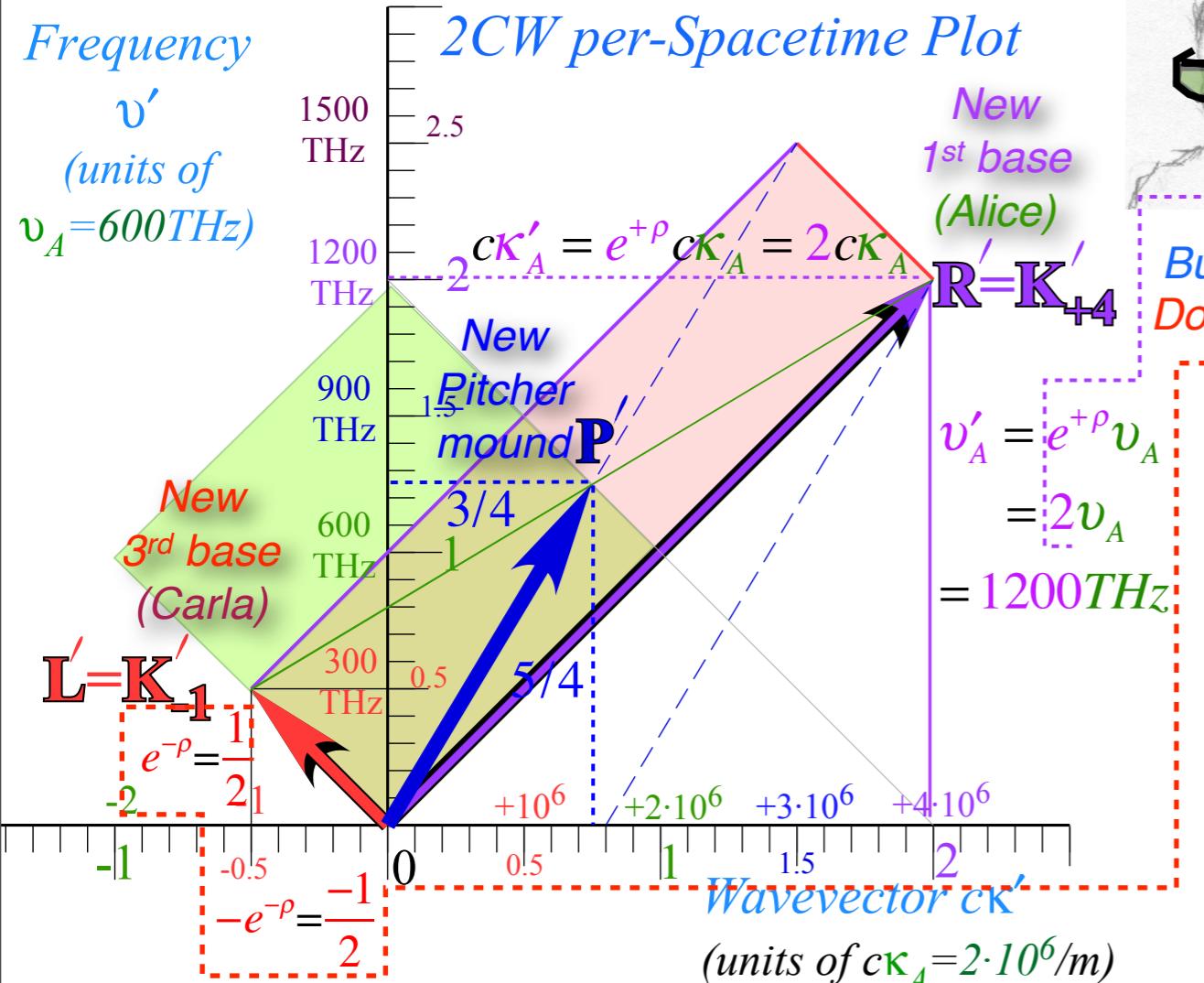
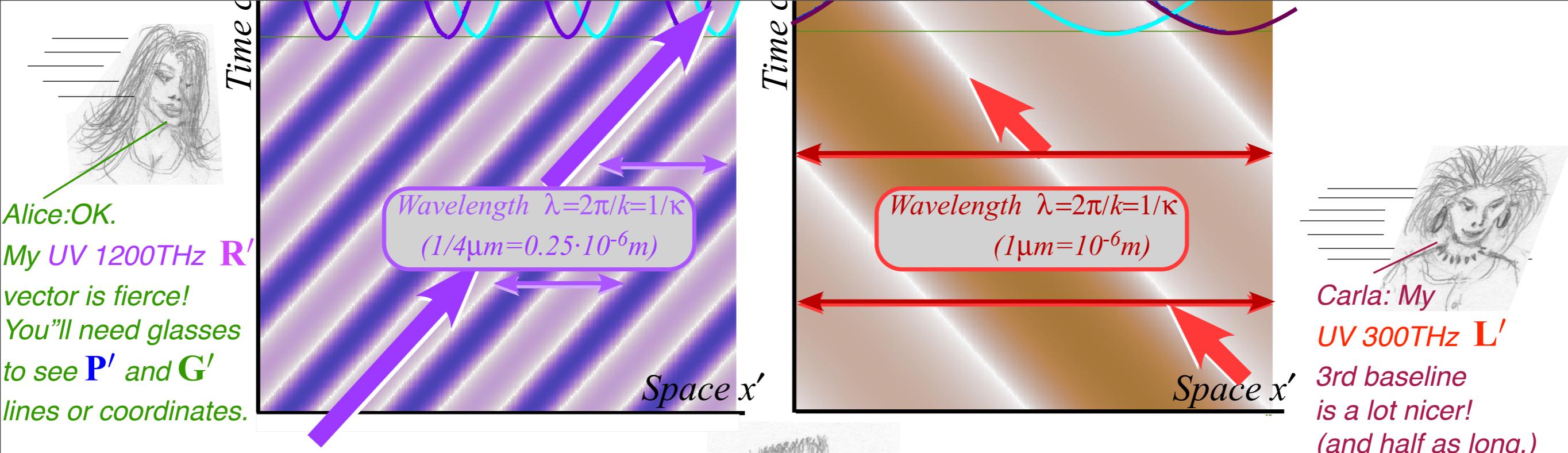
Bob: That UV burns!
I need to put on my sunglasses.



*Carla: My
UV 300THz \mathbf{L}'
3rd baseline
is a lot nicer!*







$$v'_C = e^{-\rho} v_A = \frac{1}{2} v_A = 300 \text{ THz}$$

Bob: Sunglasses help.

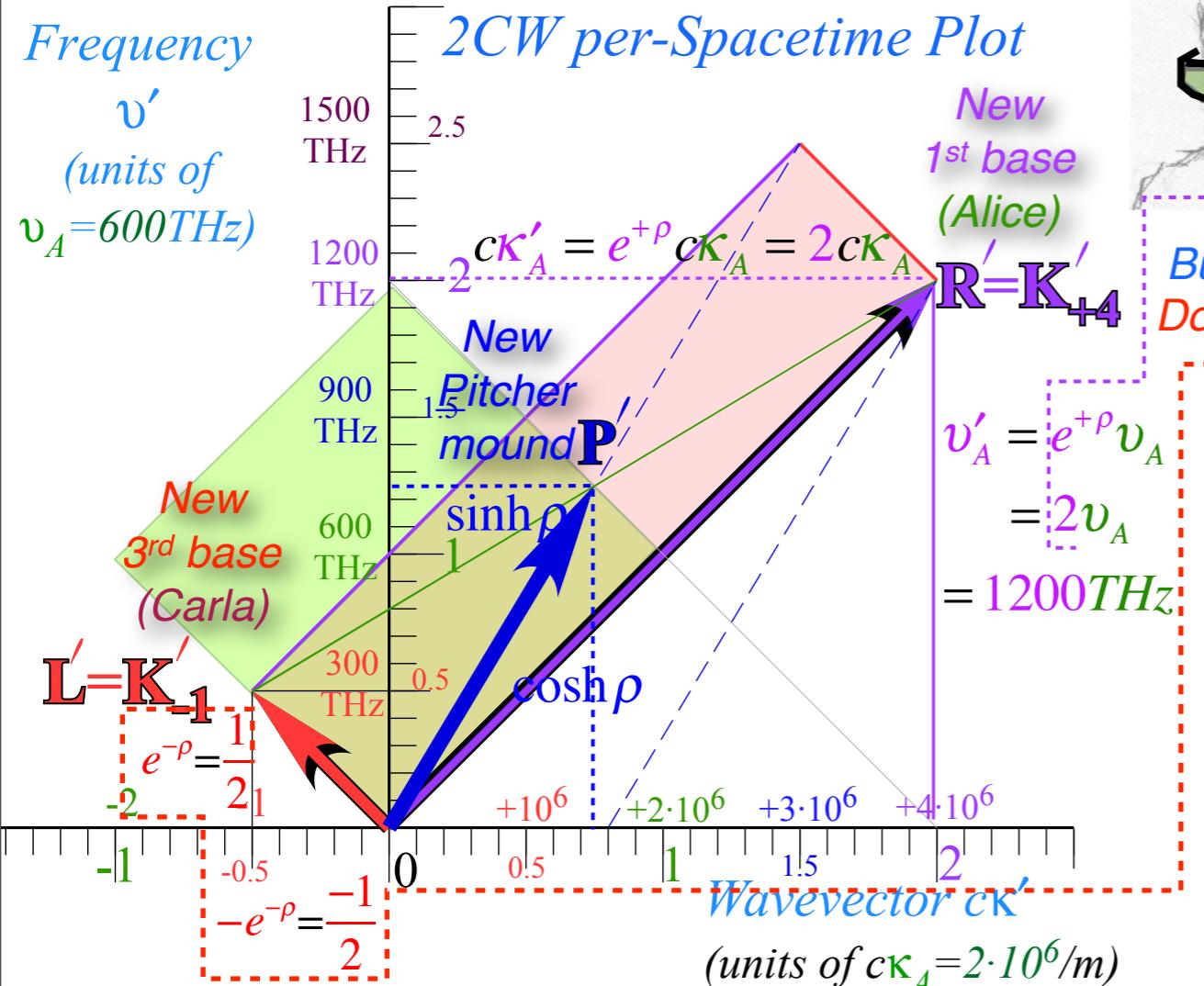
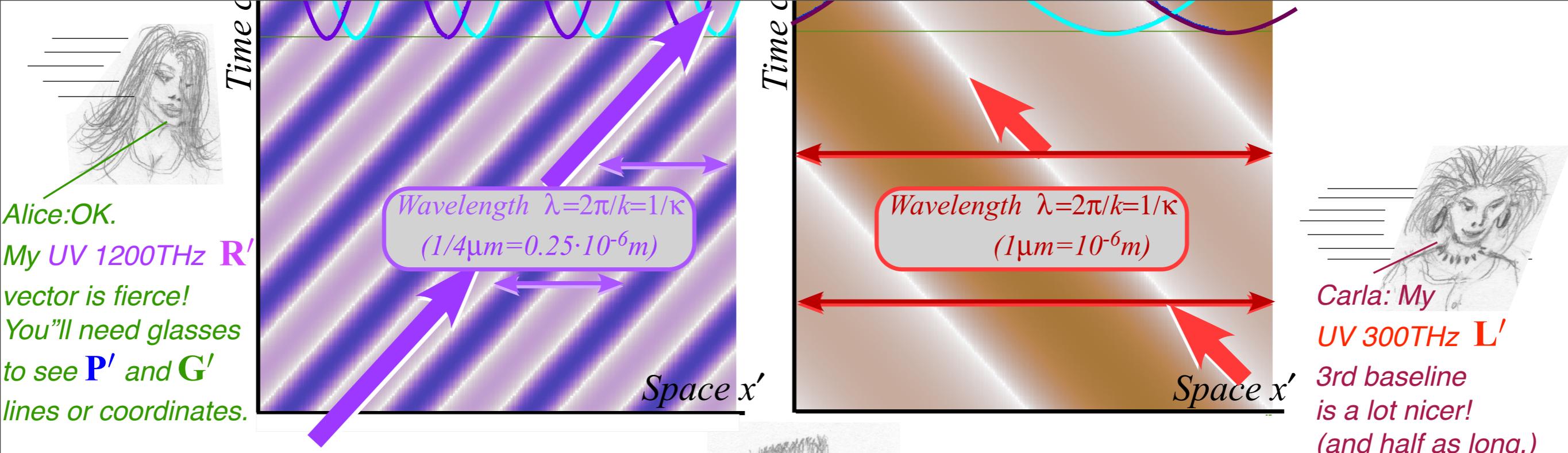
Wow! Your 1st baseline \mathbf{R}' is Doppler blued up by $e^{+p} = 2$.

But, Carla's 3rd baseline \mathbf{L}' is Doppler red shifted by $e^{-p} = 1/2$.

New “Pitcher-mound” \mathbf{P}' (Phase pt.) is 1/2-sum $(\mathbf{R}' + \mathbf{L}')/2$:

$$\begin{pmatrix} c\mathbf{K}'_{phase} \\ \mathbf{v}'_{phase} \end{pmatrix} = \frac{\mathbf{v}_A}{2} \begin{pmatrix} 2 \\ 2 \end{pmatrix} + \frac{\mathbf{v}_A}{2} \begin{pmatrix} -1/2 \\ +1/2 \end{pmatrix} = \mathbf{v}_A \begin{pmatrix} 2-1/2 \\ 2 \end{pmatrix} = \mathbf{v}_A \begin{pmatrix} 3/4 \\ 5/4 \end{pmatrix}$$

$$\begin{pmatrix} c\kappa'_{phase} \\ v'_{phase} \end{pmatrix} = \frac{v_A}{2} \begin{pmatrix} 2 \\ 2 \end{pmatrix} + \frac{v_A}{2} \begin{pmatrix} -1/2 \\ +1/2 \end{pmatrix} = v_A \begin{pmatrix} 3/4 \\ 5/4 \end{pmatrix}$$

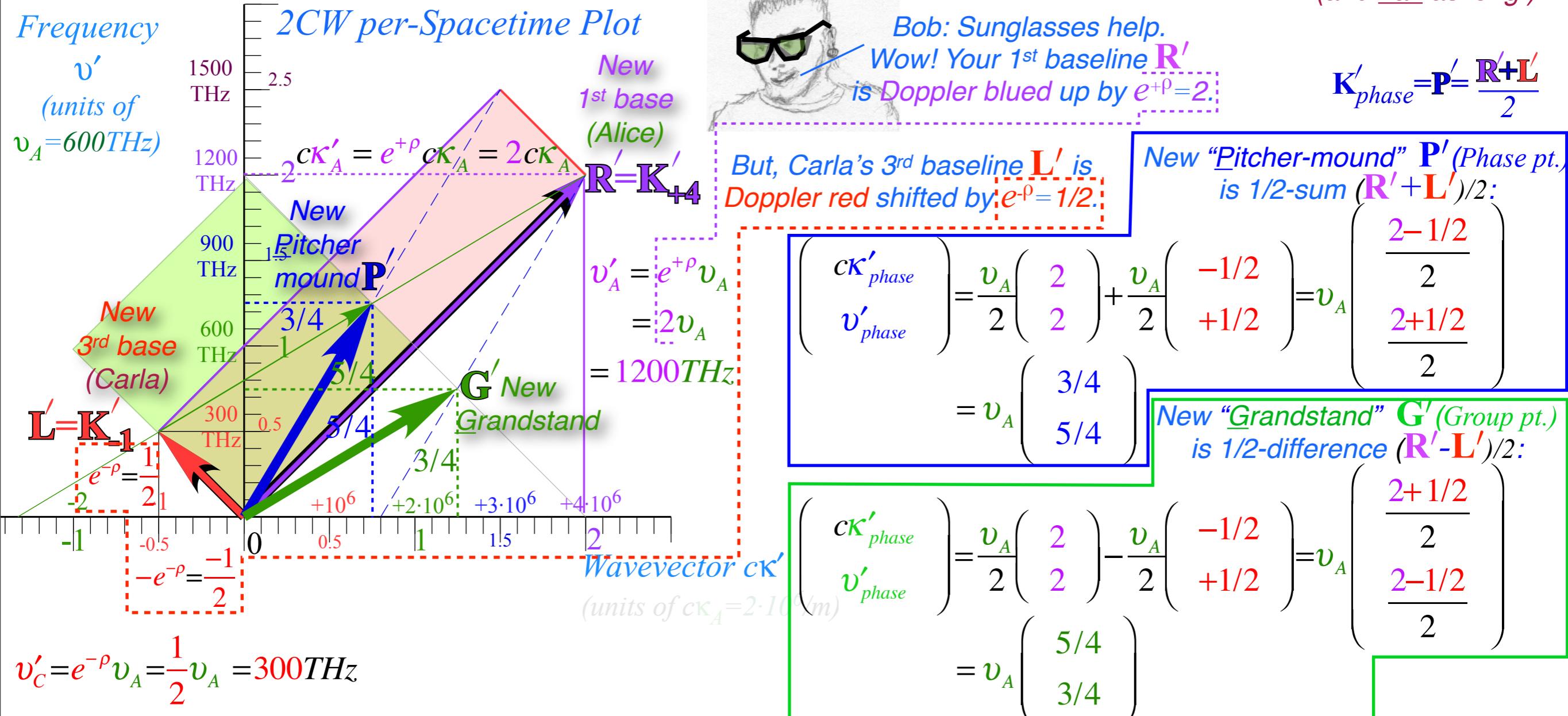
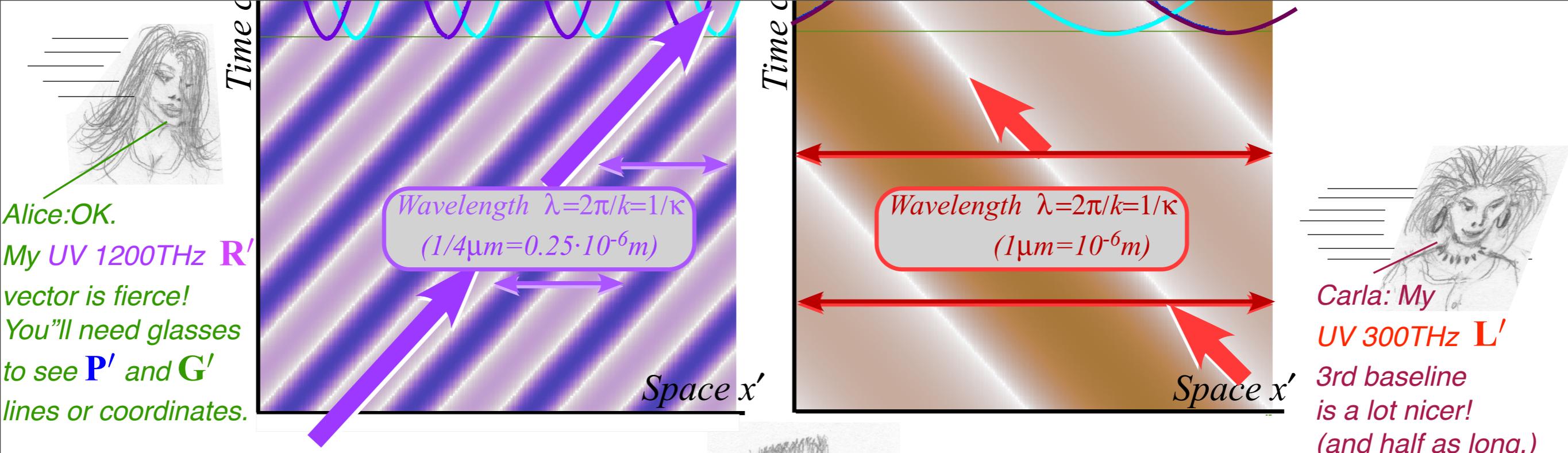


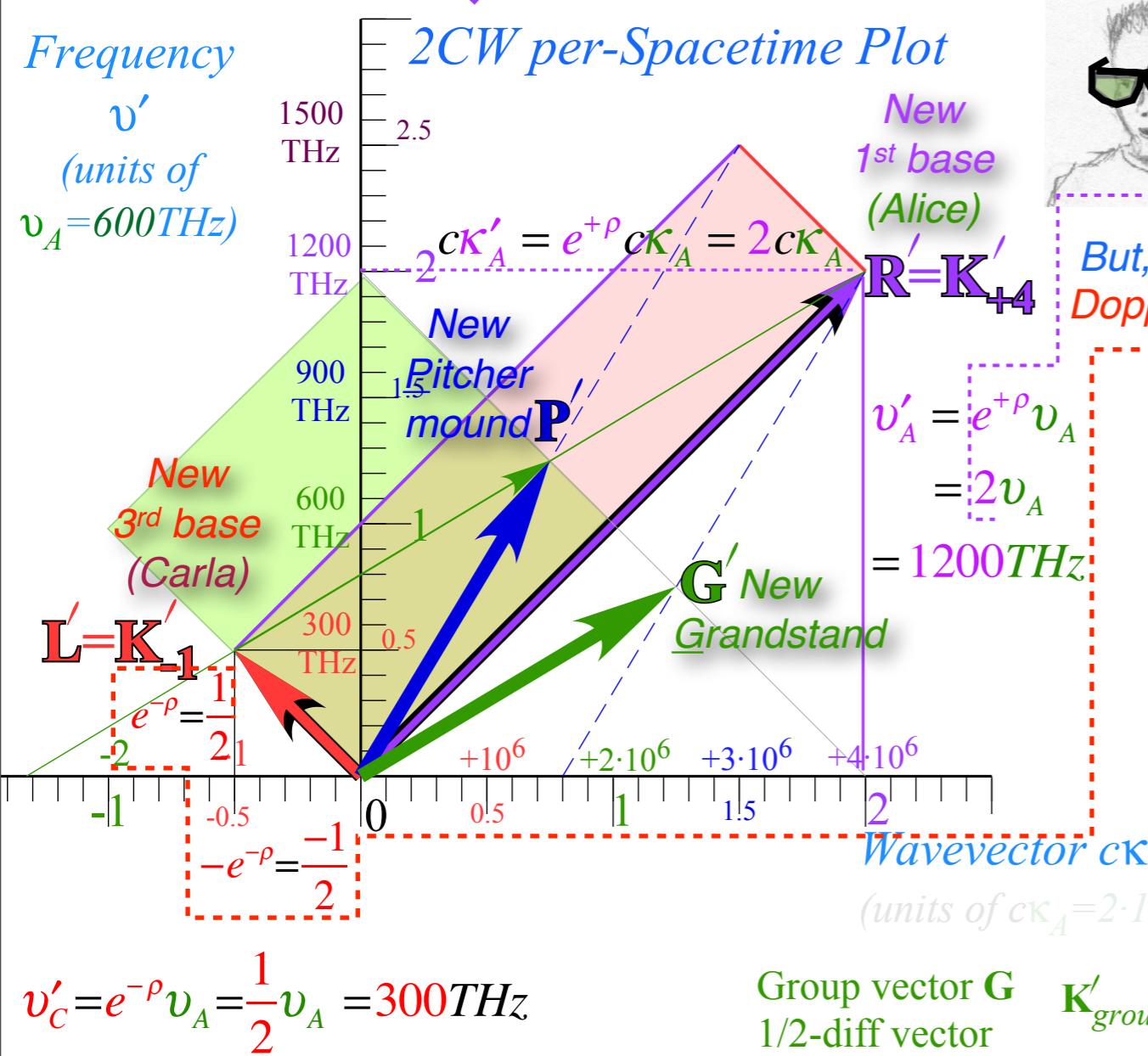
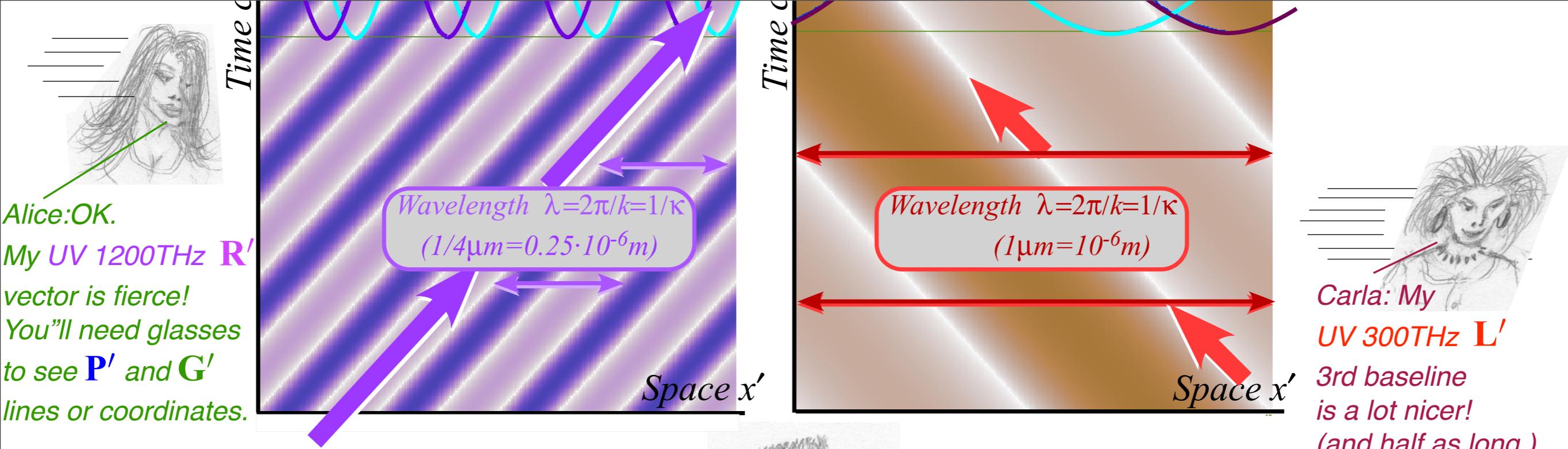
But, Carla's 3rd baseline \mathbf{L}' is Doppler redshifted by $e^{-\rho} = 1/2$.

New "Pitcher-mound" \mathbf{P}' (Phase pt.) is 1/2-sum $(\mathbf{R}' + \mathbf{L}')/2$:

$$\begin{pmatrix} cK'_{phase} \\ v'_{phase} \end{pmatrix} = \frac{v_A}{2} \begin{pmatrix} e^{+\rho} \\ e^{+\rho} \end{pmatrix} + \frac{v_A}{2} \begin{pmatrix} -e^{-\rho} \\ +e^{-\rho} \end{pmatrix} = v_A \begin{pmatrix} \sinh \rho \\ \cosh \rho \end{pmatrix} = v_A \begin{pmatrix} 3/4 \\ 5/4 \end{pmatrix}$$

$$K'_{phase} = \mathbf{P}' = \frac{\mathbf{R}' + \mathbf{L}'}{2}$$





Bob: Sunglasses help.

Wow! Your 1st baseline \mathbf{R}' is Doppler blueshifted by $e^{+\rho}=2$.

$\mathbf{K}'_{phase} = \mathbf{P}' = \frac{\mathbf{R} + \mathbf{L}'}{2}$

But, Carla's 3rd baseline \mathbf{L}' is Doppler redshifted by $e^{-\rho}=1/2$.

New “Pitcher-mound” \mathbf{P}' (Phase pt.) is 1/2-sum $(\mathbf{R}' + \mathbf{L}')/2$:

$$\begin{pmatrix} c\mathbf{K}'_{phase} \\ v'_{phase} \end{pmatrix} = \frac{v_A}{2} \begin{pmatrix} 2 \\ 2 \end{pmatrix} + \frac{v_A}{2} \begin{pmatrix} -1/2 \\ +1/2 \end{pmatrix} = v_A \begin{pmatrix} 2-1/2 \\ 2 \end{pmatrix} = v_A \begin{pmatrix} 3/4 \\ 5/4 \end{pmatrix}$$

New “Grandstand” \mathbf{G}' (Group pt.) is 1/2-difference $(\mathbf{R}' - \mathbf{L}')/2$:

$$\begin{pmatrix} c\mathbf{K}'_{group} \\ v'_{group} \end{pmatrix} = \frac{v_A}{2} \begin{pmatrix} e^{+\rho} \\ e^{+\rho} \end{pmatrix} - \frac{v_A}{2} \begin{pmatrix} -e^{-\rho} \\ +e^{-\rho} \end{pmatrix} = v_A \begin{pmatrix} e^{+\rho} + e^{-\rho} \\ 2 \end{pmatrix} = v_A \begin{pmatrix} \cosh \rho \\ \sinh \rho \end{pmatrix} = v_A \begin{pmatrix} 5/4 \\ 3/4 \end{pmatrix}$$

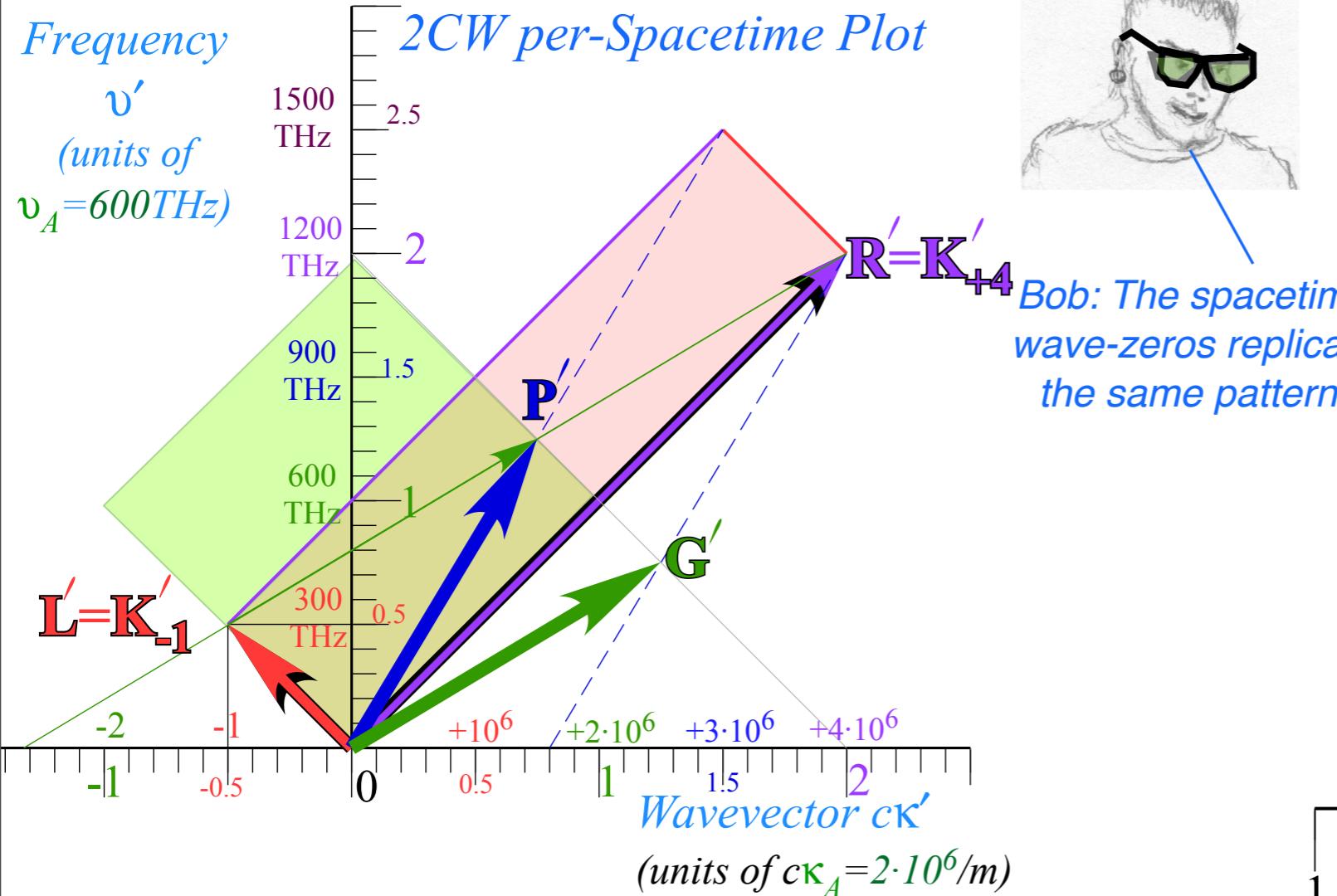
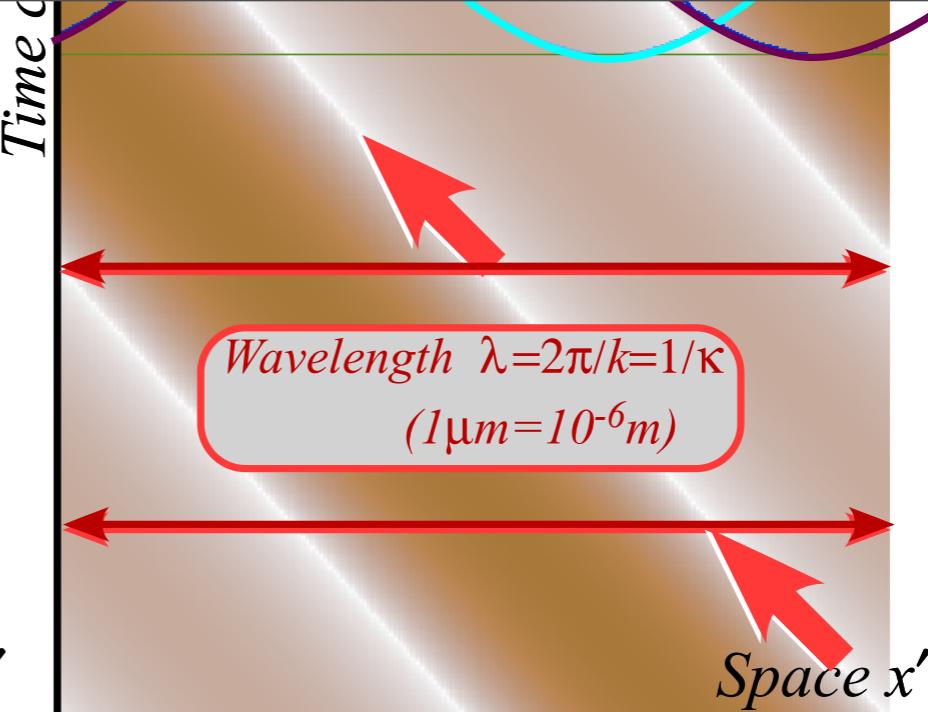
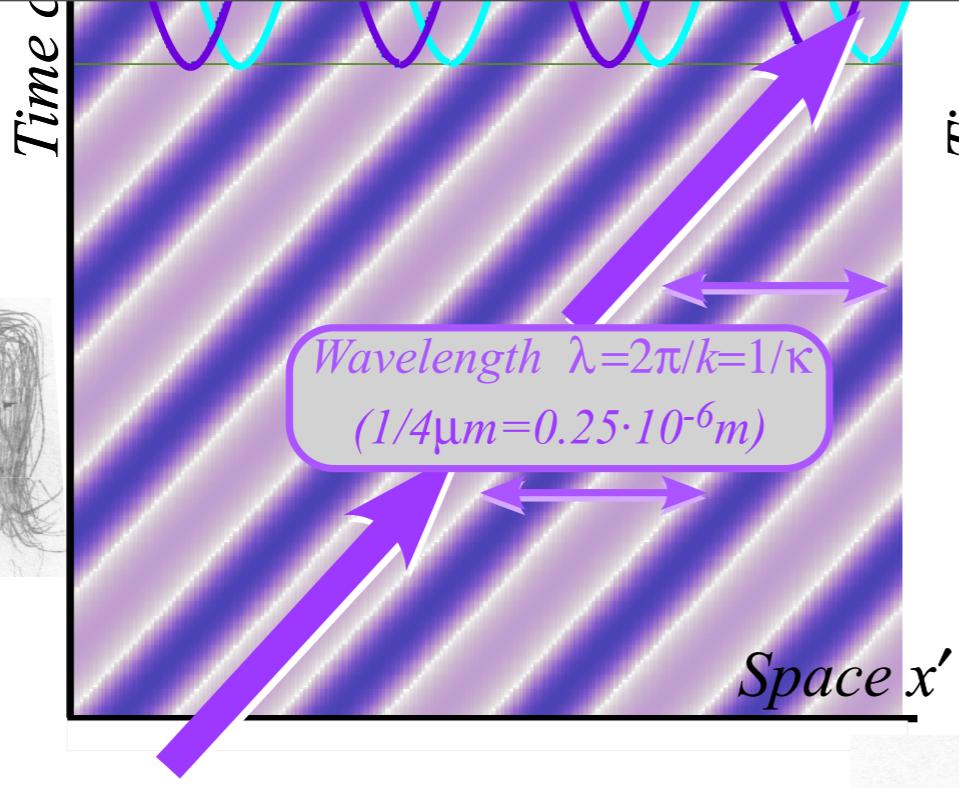
$$K'_{phase} = P' = \frac{R'+L'}{2}$$

$$\begin{pmatrix} cK'_{phase} \\ v'_{phase} \end{pmatrix} = \frac{v_A}{2} \begin{pmatrix} 2 \\ 2 \end{pmatrix} + \frac{v_A}{2} \begin{pmatrix} -1/2 \\ +1/2 \end{pmatrix} = v_A \begin{pmatrix} \frac{2-1/2}{2} \\ \frac{2+1/2}{2} \end{pmatrix}$$

$$= v_A \begin{pmatrix} 3/4 \\ 5/4 \end{pmatrix}$$

New “Grandstand” G' (Group pt.)
is $1/2$ -difference $(R' - L')/2$:

$$\kappa' \begin{pmatrix} cK'_{group} \\ v'_{group} \end{pmatrix} = \frac{v_A}{2} \begin{pmatrix} e^{+\rho} \\ e^{+\rho} \end{pmatrix} - \frac{v_A}{2} \begin{pmatrix} -e^{-\rho} \\ +e^{-\rho} \end{pmatrix} = v_A \begin{pmatrix} \frac{e^{+\rho} + e^{-\rho}}{2} \\ \frac{e^{+\rho} - e^{-\rho}}{2} \end{pmatrix}$$

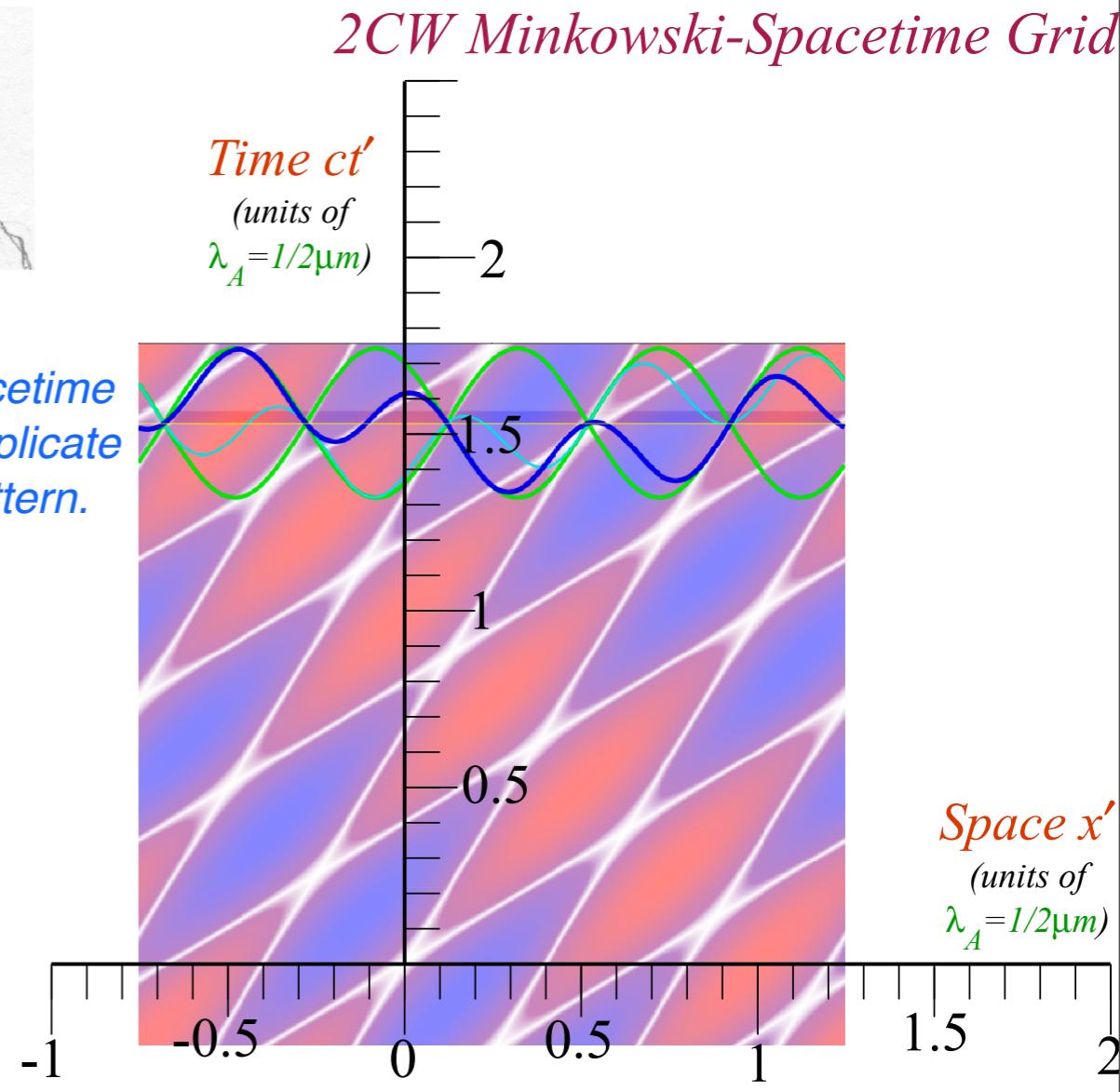


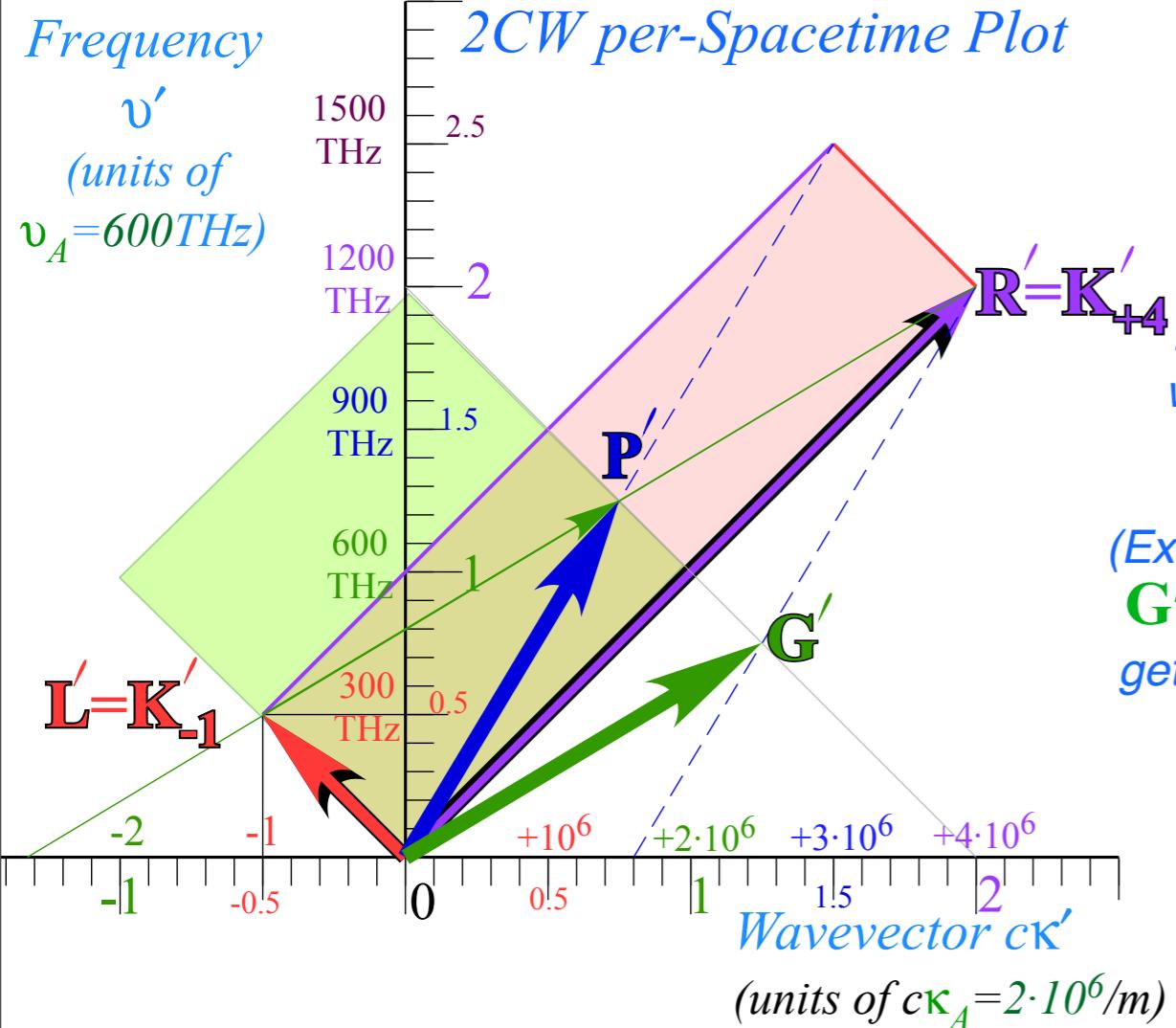
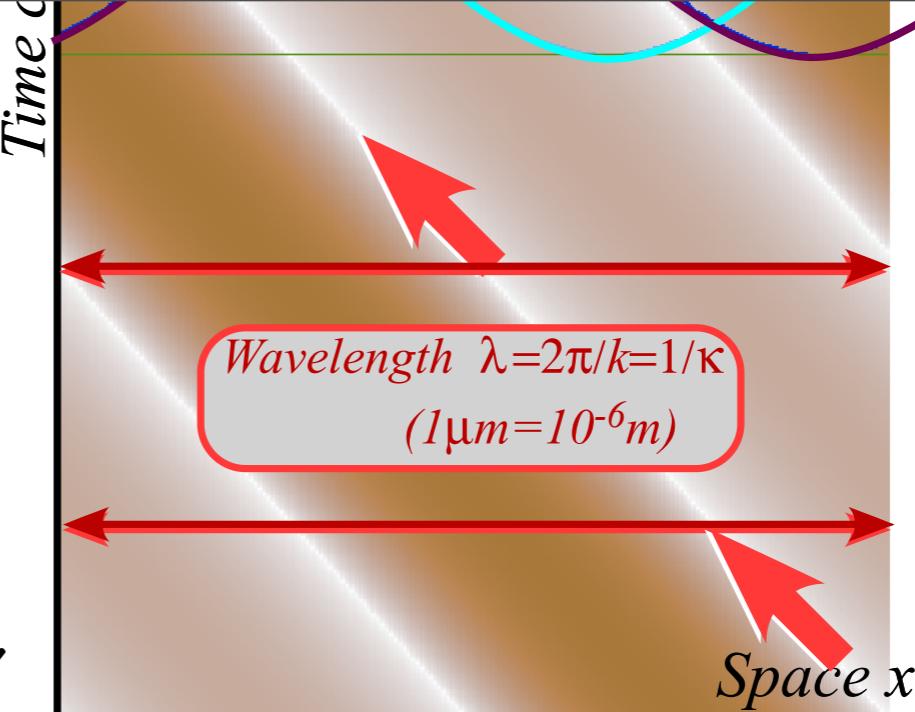
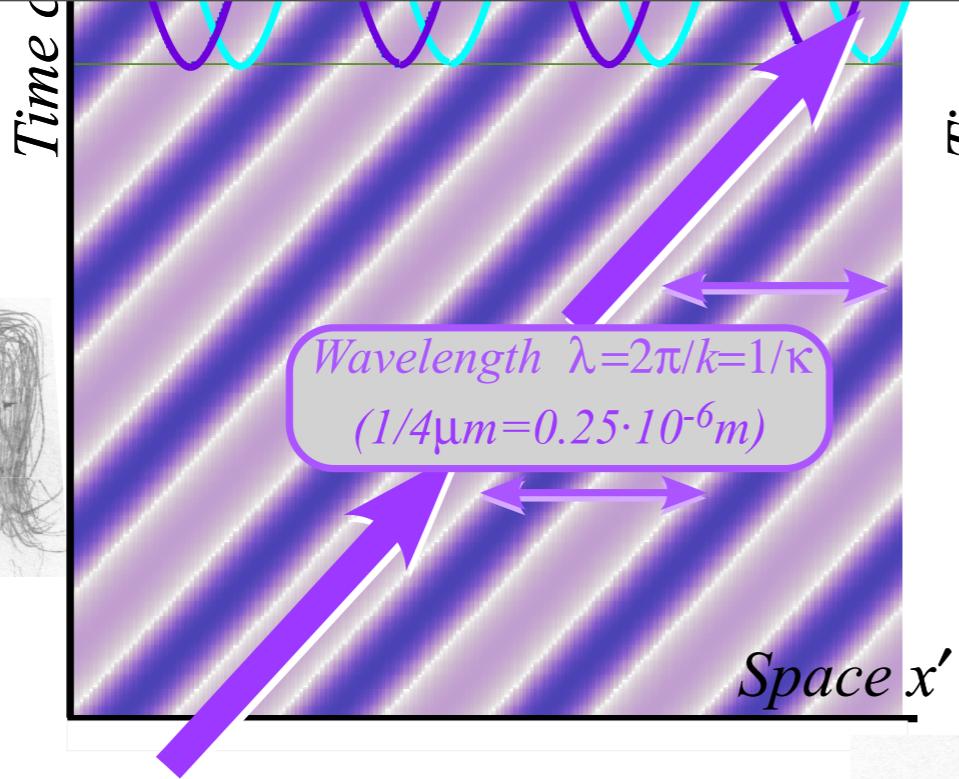
Phase vector P
1/2-sum vector

$$K'_{phase} = \frac{P + L'}{2}$$

Group vector G
1/2-diff vector

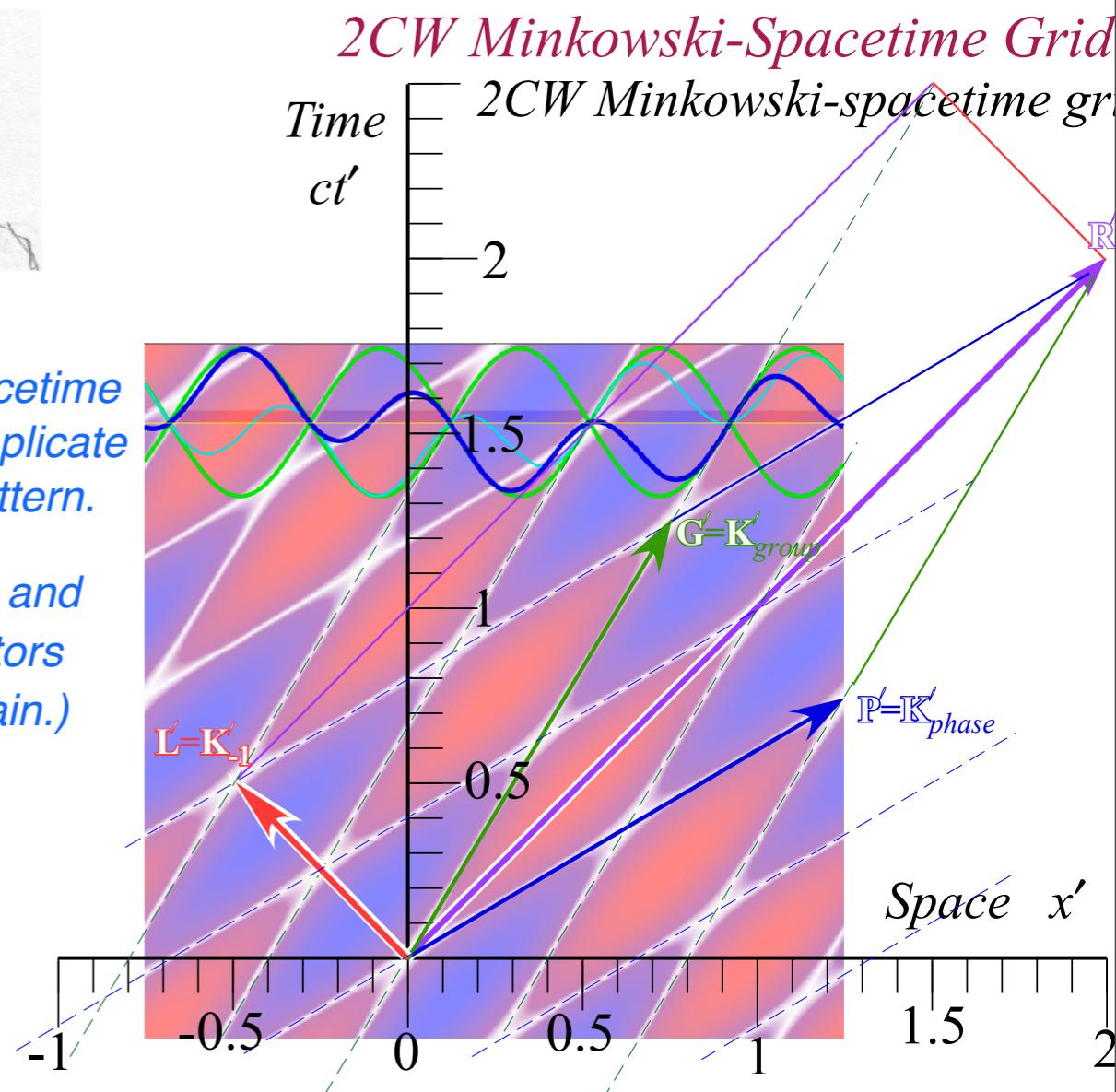
$$K'_{group} = \frac{G' - L'}{2}$$

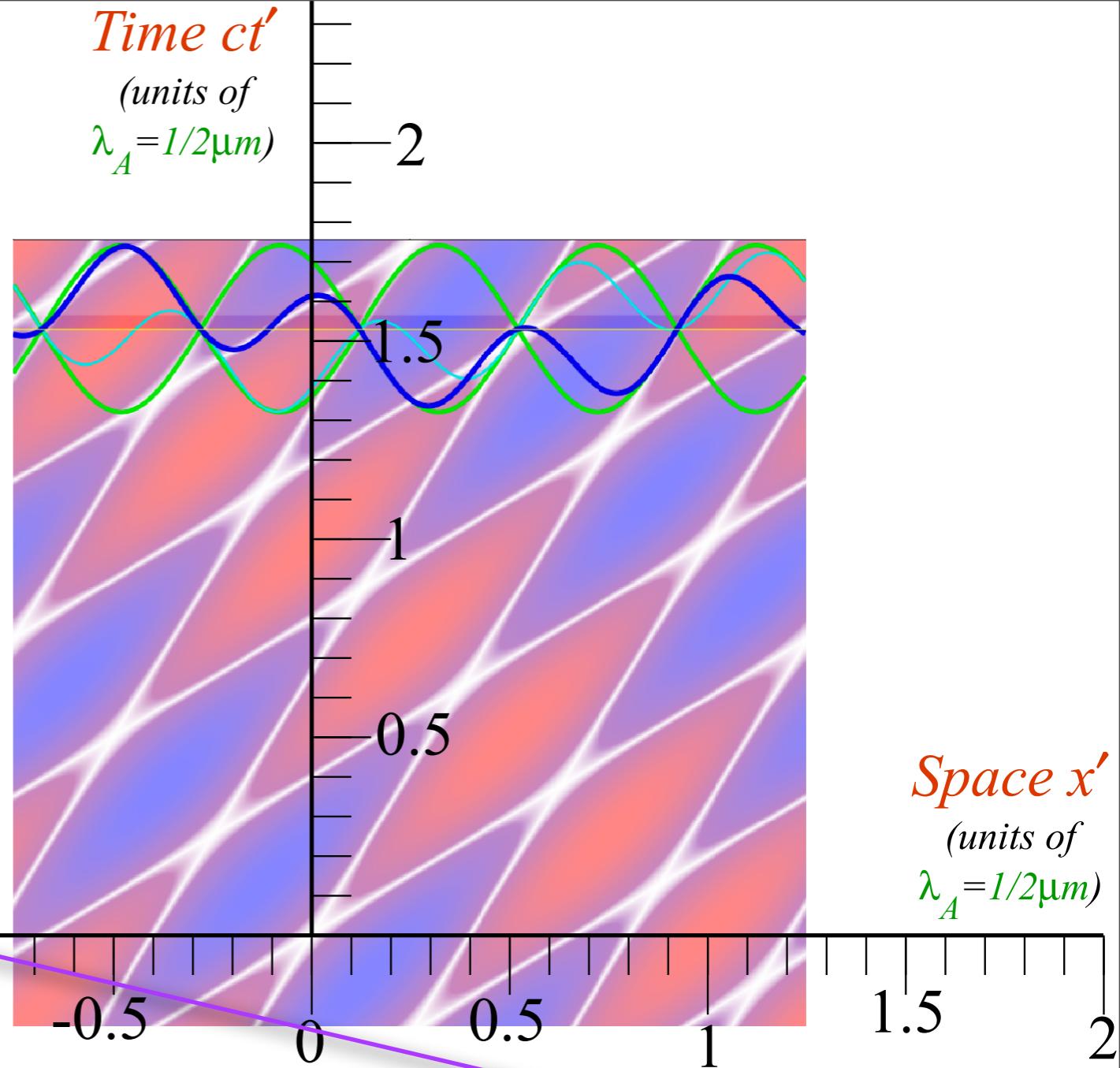
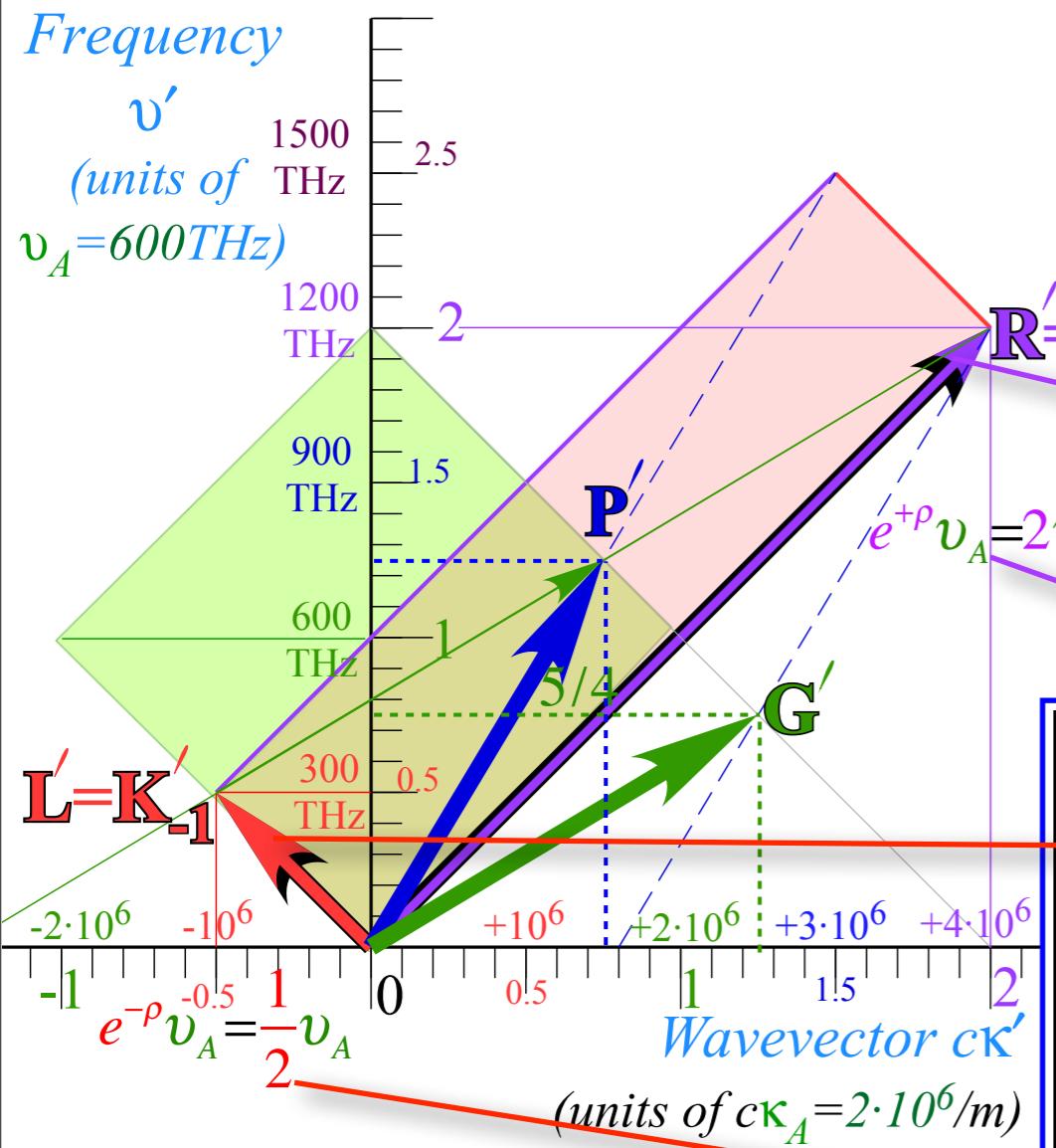




Phase vector \mathbf{P}'
1/2-sum vector $\mathbf{K}'_{\text{phase}} = \frac{\mathbf{P}' + \mathbf{L}'}{2}$

Group vector \mathbf{G}'
1/2-diff vector $\mathbf{K}'_{\text{group}} = \frac{\mathbf{G}' - \mathbf{L}'}{2}$





phase	$b_{RED}^{Doppler}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
group	$b_{BLUE}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{group}}{c}$	$b_{RED}^{Doppler}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

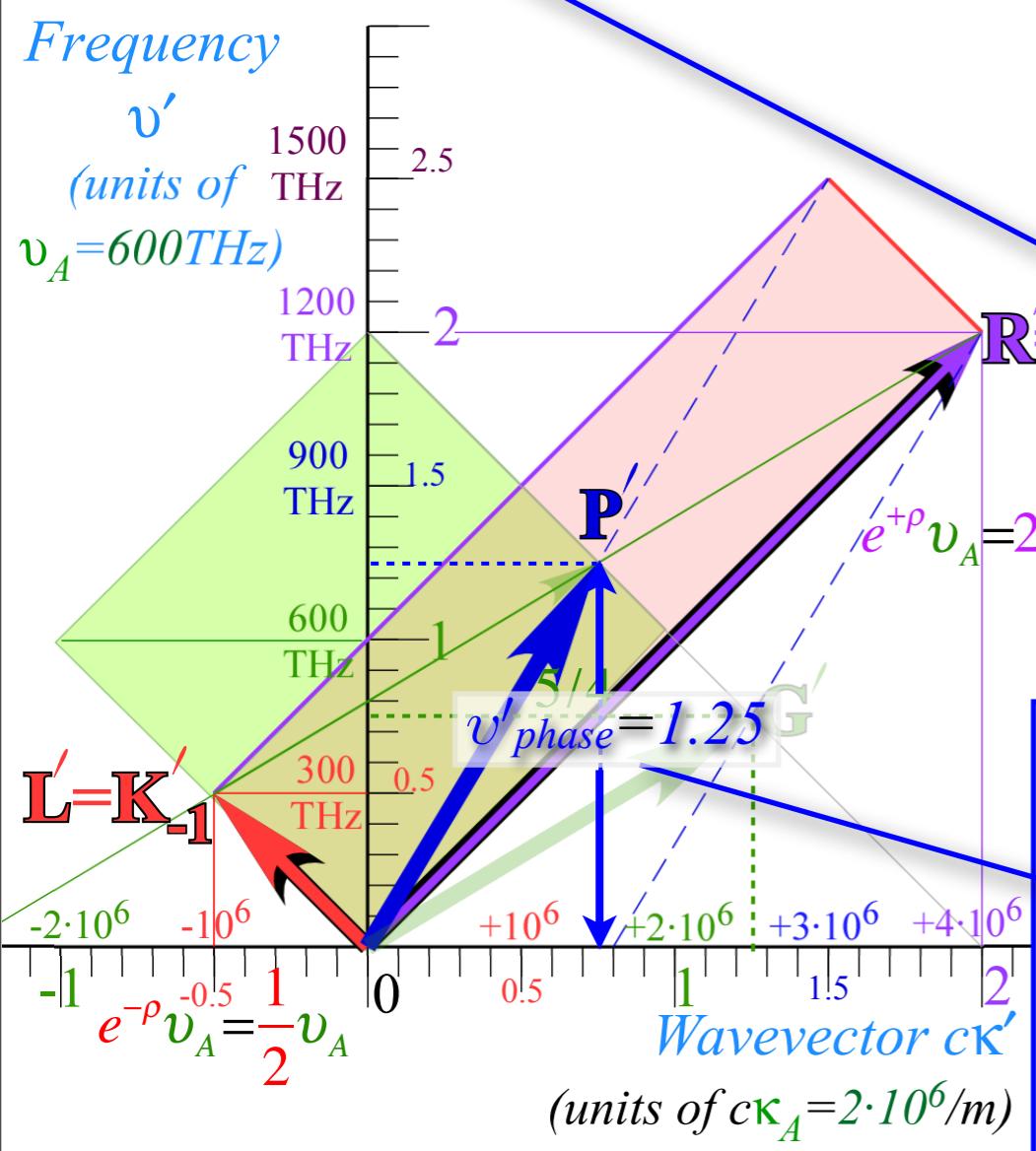
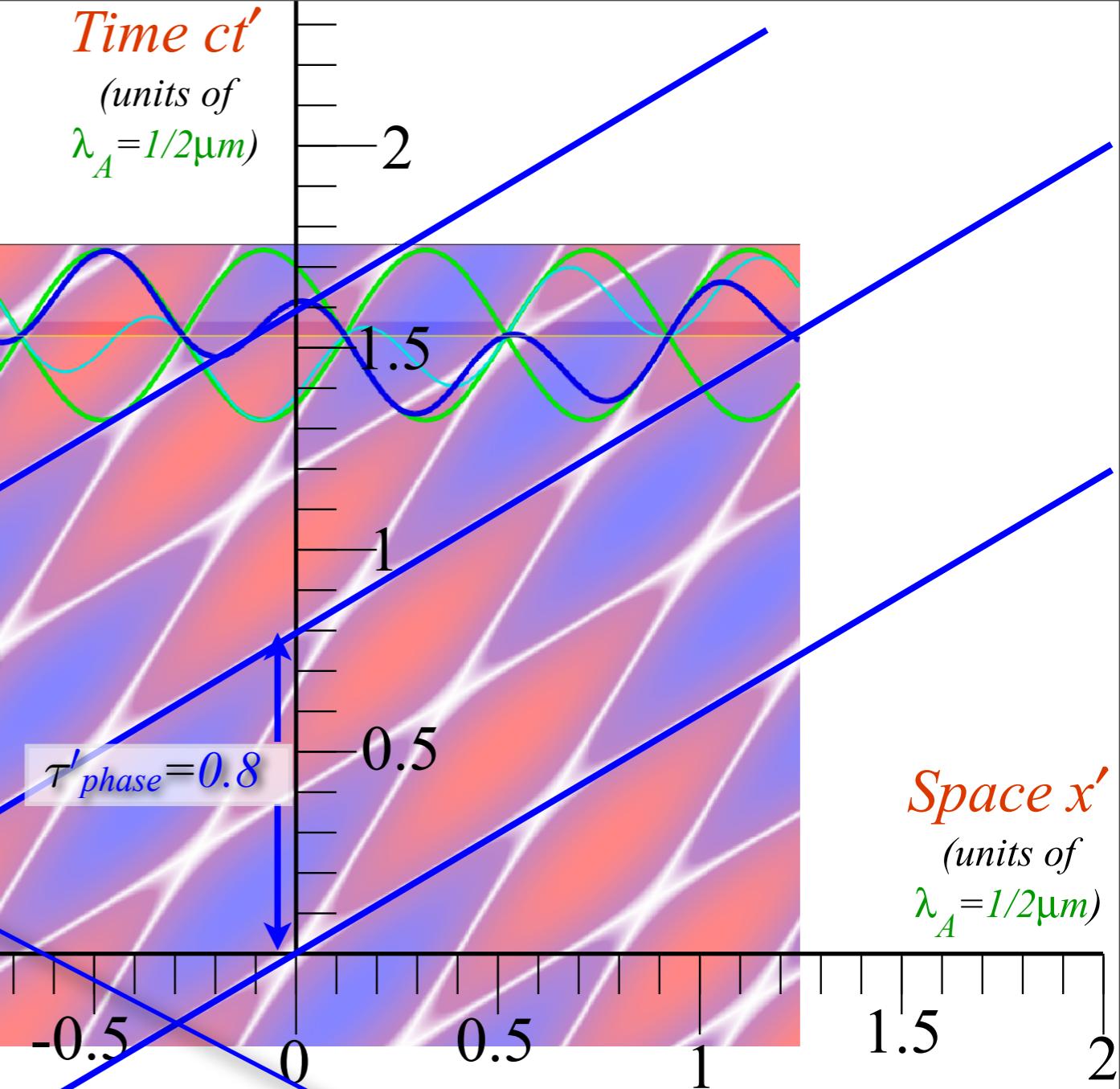
$$\mathbf{P}' = \begin{pmatrix} c\kappa'_{phase} \\ v'_{phase} \end{pmatrix} = v_A \begin{pmatrix} \sinh \rho \\ \cosh \rho \end{pmatrix} = v_A \begin{pmatrix} 3/4 \\ 5/4 \end{pmatrix}$$

Phase frequency
 $v'_{phase} = v_A \cosh \rho = 5/4 = 1.25$

flips to

Phase period $\tau = 1/v$
 $\tau'_{phase} = \tau_A \operatorname{sech} \rho = 4/5 = 0.8$

$$\tau'_{phase} = \tau_A \operatorname{sech} \rho = 4/5$$



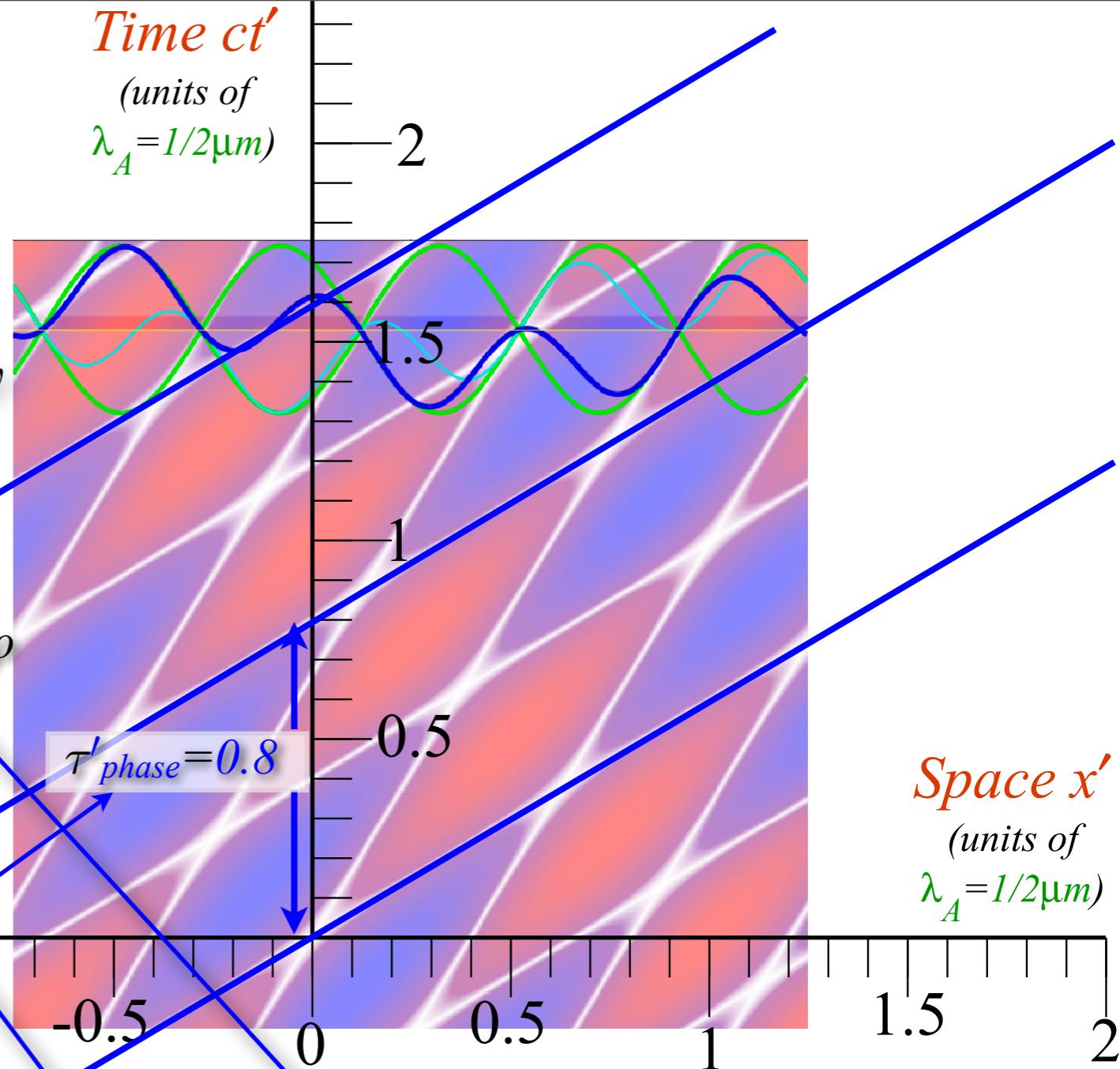
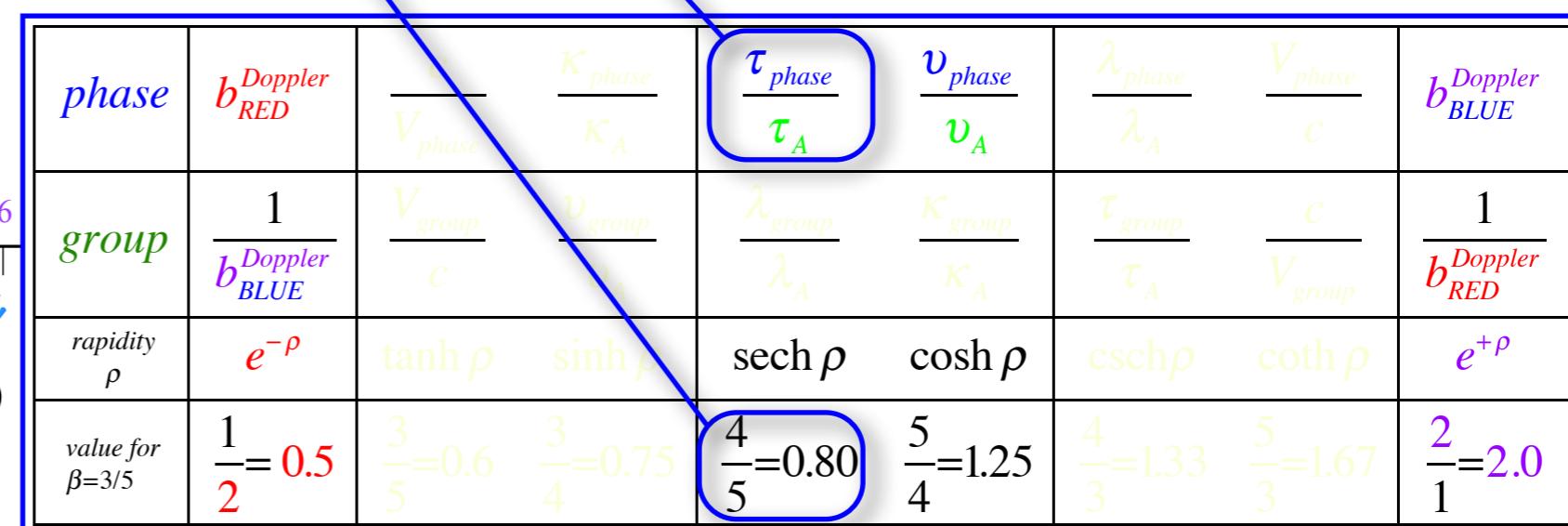
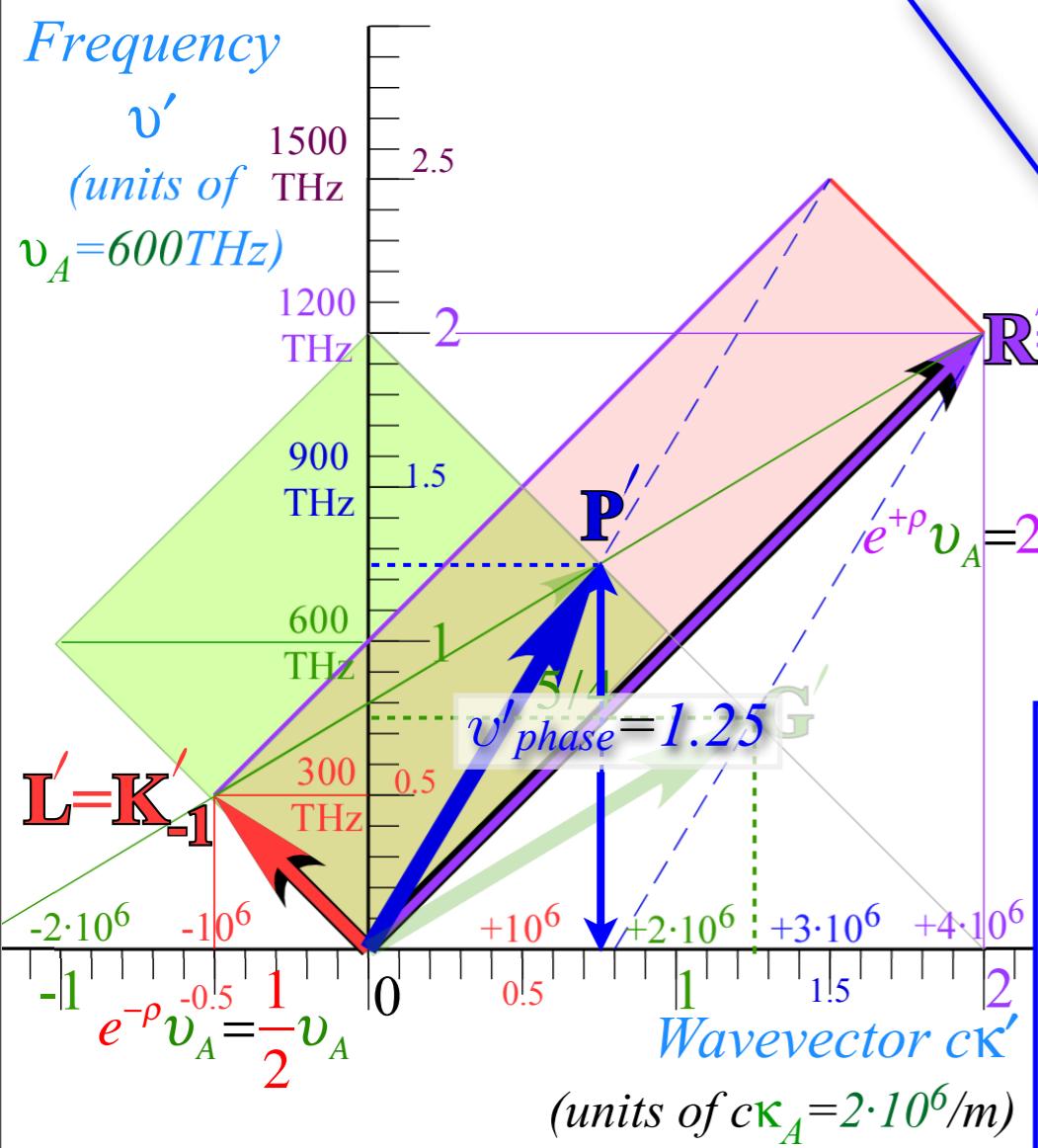
phase	$b_{Doppler}^{RED}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler}^{BLUE}$
group	1	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler}^{RED}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

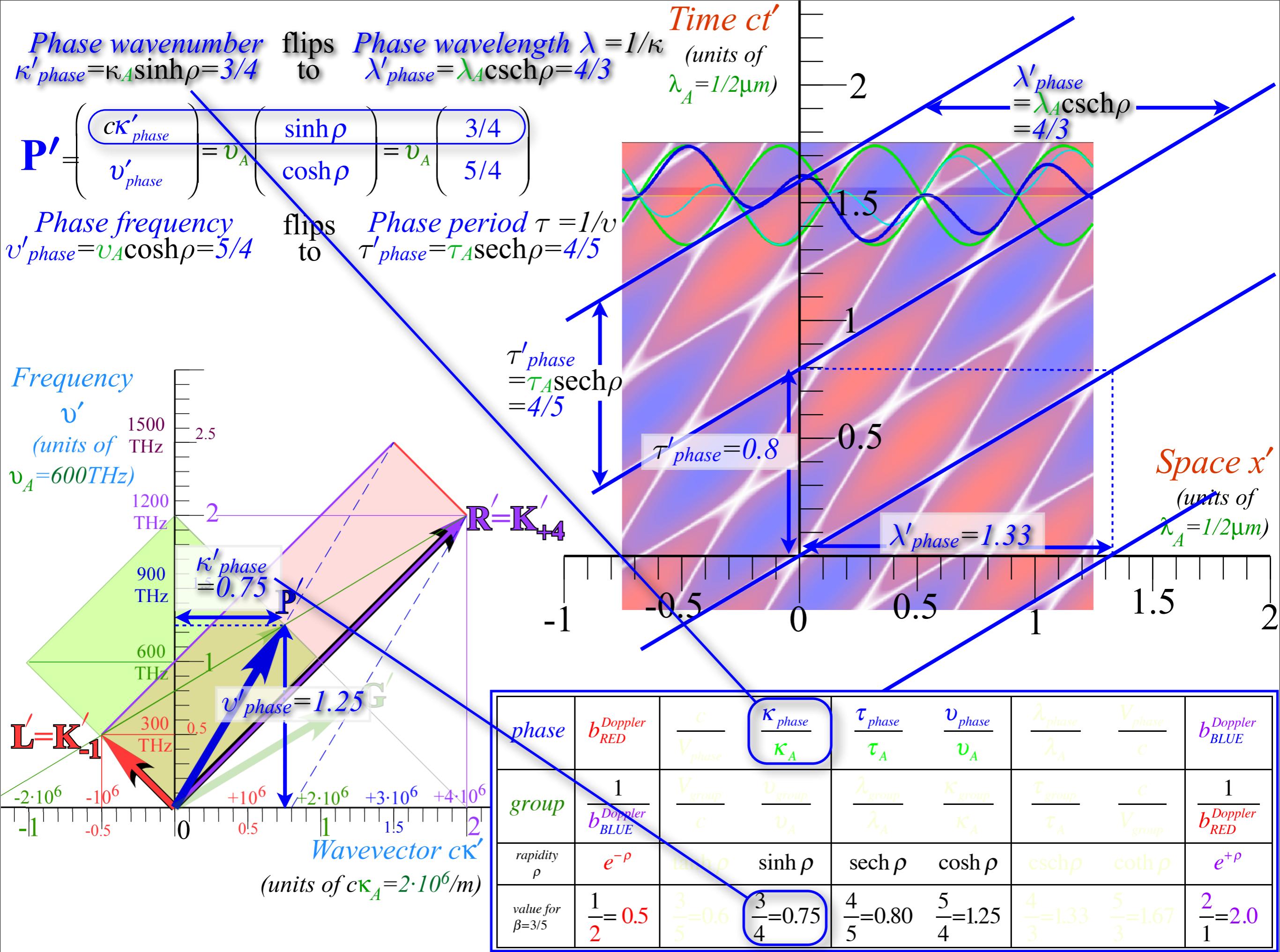
$$\mathbf{P}' = \begin{pmatrix} c\mathbf{k}'_{phase} \\ v'_{phase} \end{pmatrix} = \mathcal{V}_A \begin{pmatrix} \sinh \rho \\ \cosh \rho \end{pmatrix} = \mathcal{V}_A \begin{pmatrix} 3/4 \\ 5/4 \end{pmatrix}$$

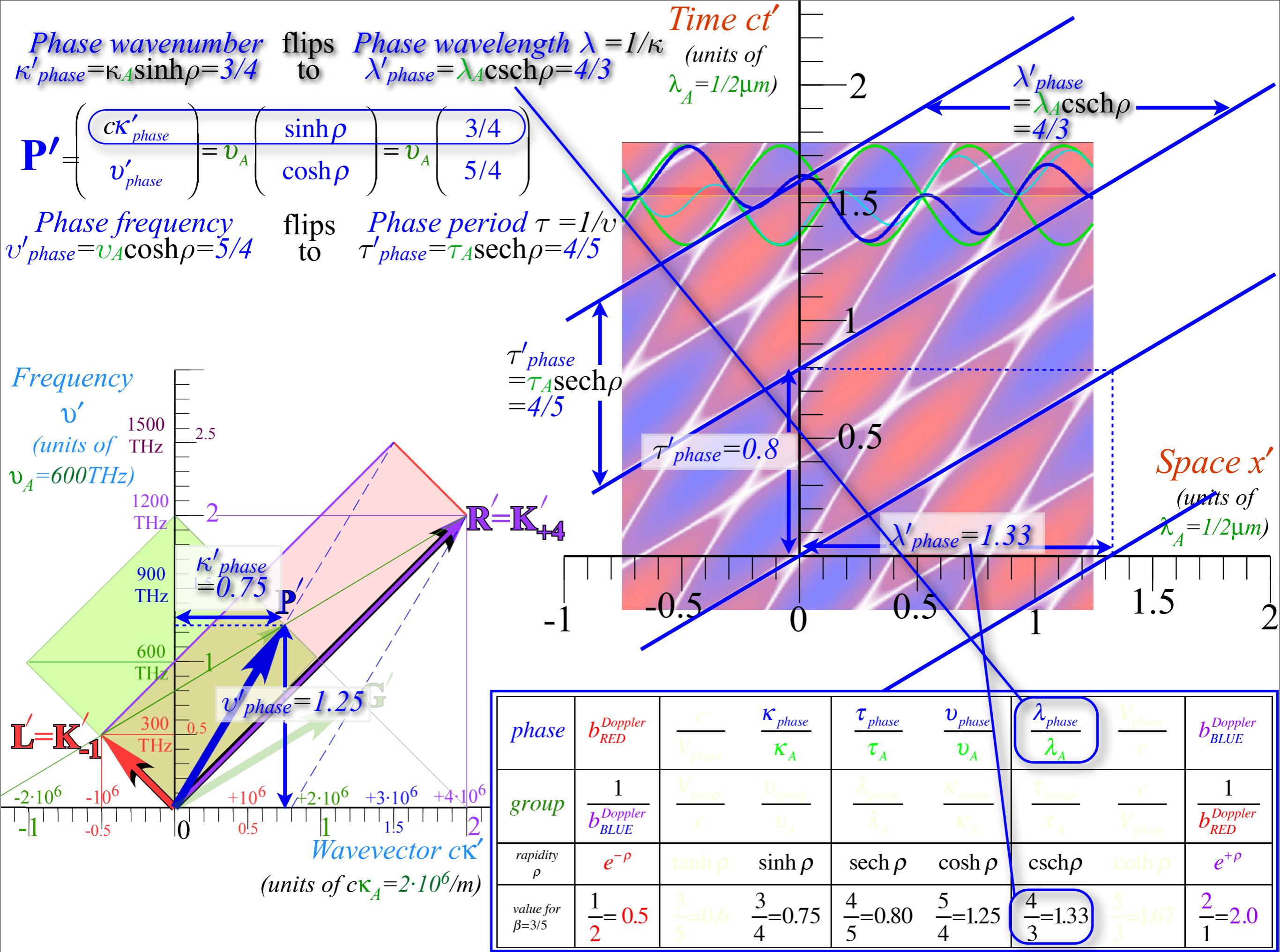
Phase frequency
 $v'_{phase} = v_A \cosh \rho = 5/4$
 $= 1.25$

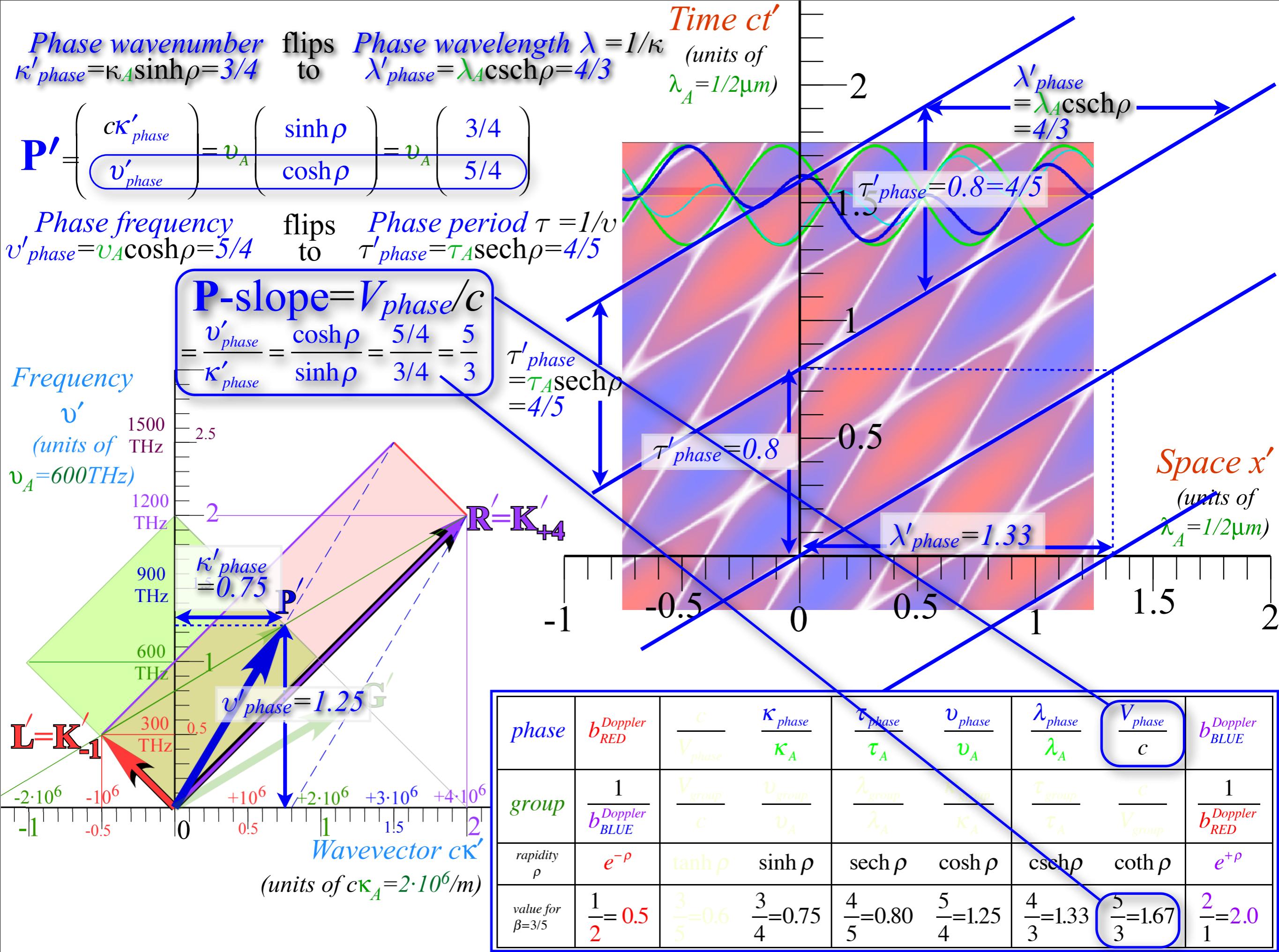
flips
to

Phase period $\tau = 1/v$
 $\tau'_{phase} = \tau_A \operatorname{sech} \rho = 4/5$









$$\text{Phase wavenumber } \kappa'_{phase} = \kappa_A \sinh \rho = 3/4 \quad \text{flips to} \quad \lambda'_{phase} = \lambda_A \cosh \rho$$

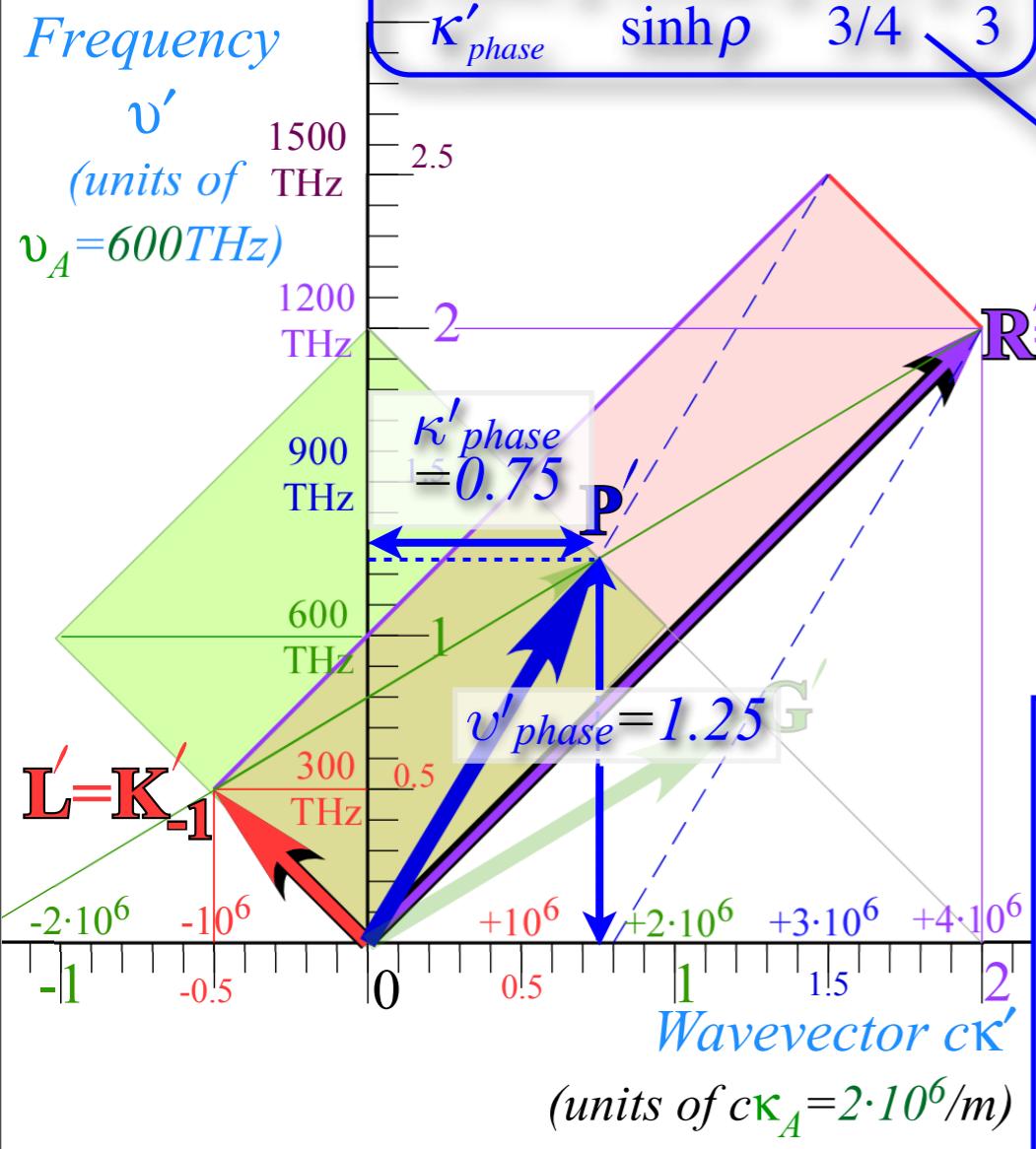
$\mathbf{P}' = \begin{pmatrix} c\kappa'_{phase} \\ v'_{phase} \end{pmatrix} = v_A \begin{pmatrix} \sinh \rho \\ \cosh \rho \end{pmatrix} = v_A \begin{pmatrix} 3/4 \\ 5/4 \end{pmatrix}$

Phase frequency $v'_{phase} = v_A \cosh \rho = 5/4$ flips *Phase period* $\tau = 1/v$
 $\tau'_{phase} = \tau_A \operatorname{sech} \rho = 4/5$

$$\text{P-slope} = V_{phase}/c$$

$$= \frac{\nu'_{phase}}{\kappa'_{phase}} = \frac{\cosh \rho}{\sinh \rho} = \frac{5/4}{3/4} = \frac{5}{3}$$

$$\begin{aligned}\tau'_{phase} &= \tau_A \operatorname{sech} \rho \\ &= 4/5\end{aligned}$$



Time ct'

κ (units of $\lambda_A = 1/2\mu\text{m}$)

$\tau'_{\text{phase}} = 0.8$

$\lambda'_{\text{phase}} = 1.33$

$\lambda'_{\text{phase}} = \lambda_A \text{csch} \rho = 4/3$

slope $v_{\text{phase}}/c = \coth \rho$

$v_{\text{phase}}/c = 4/3, 5/4, 5/3$

faster than light!

Space x'

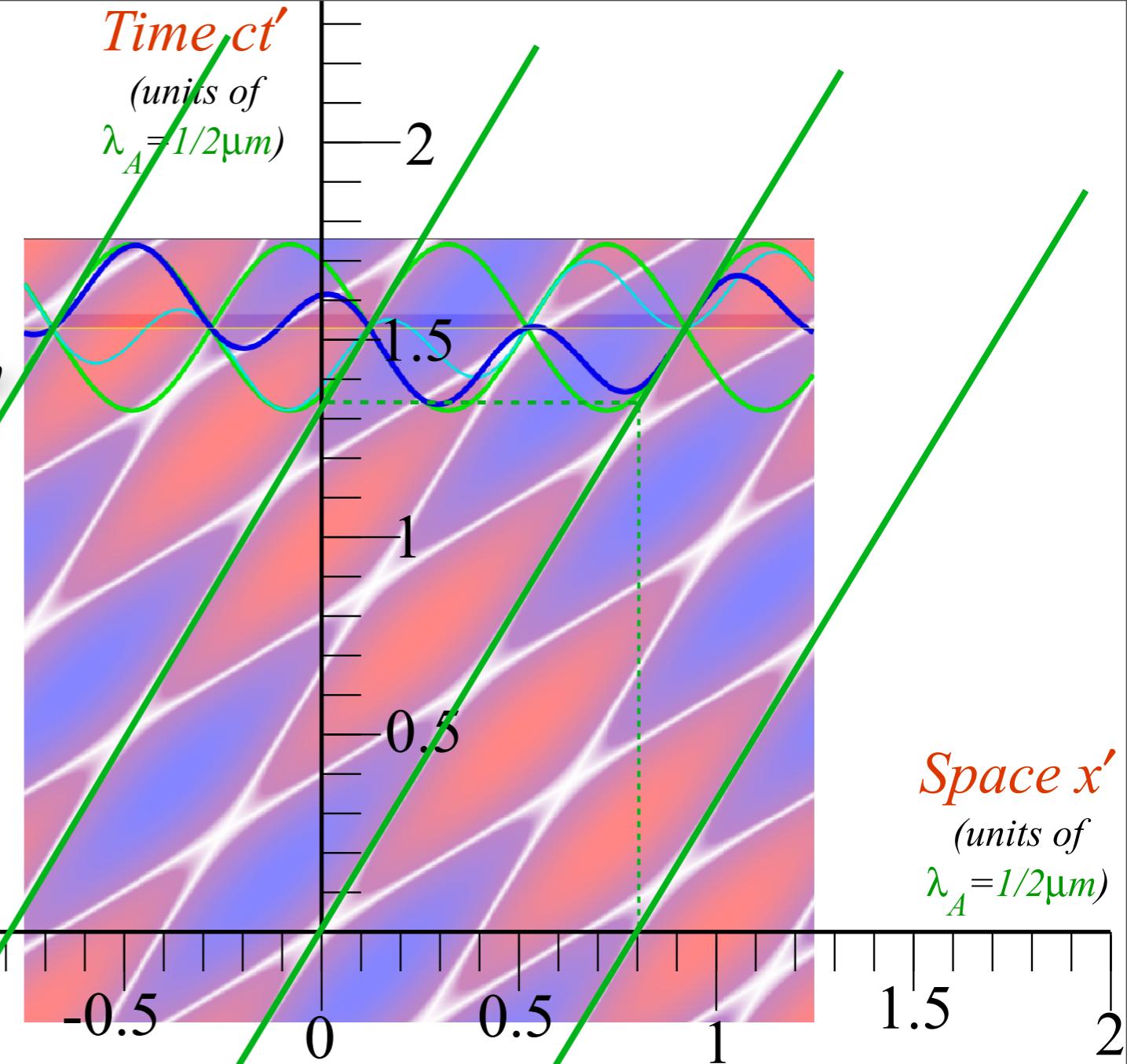
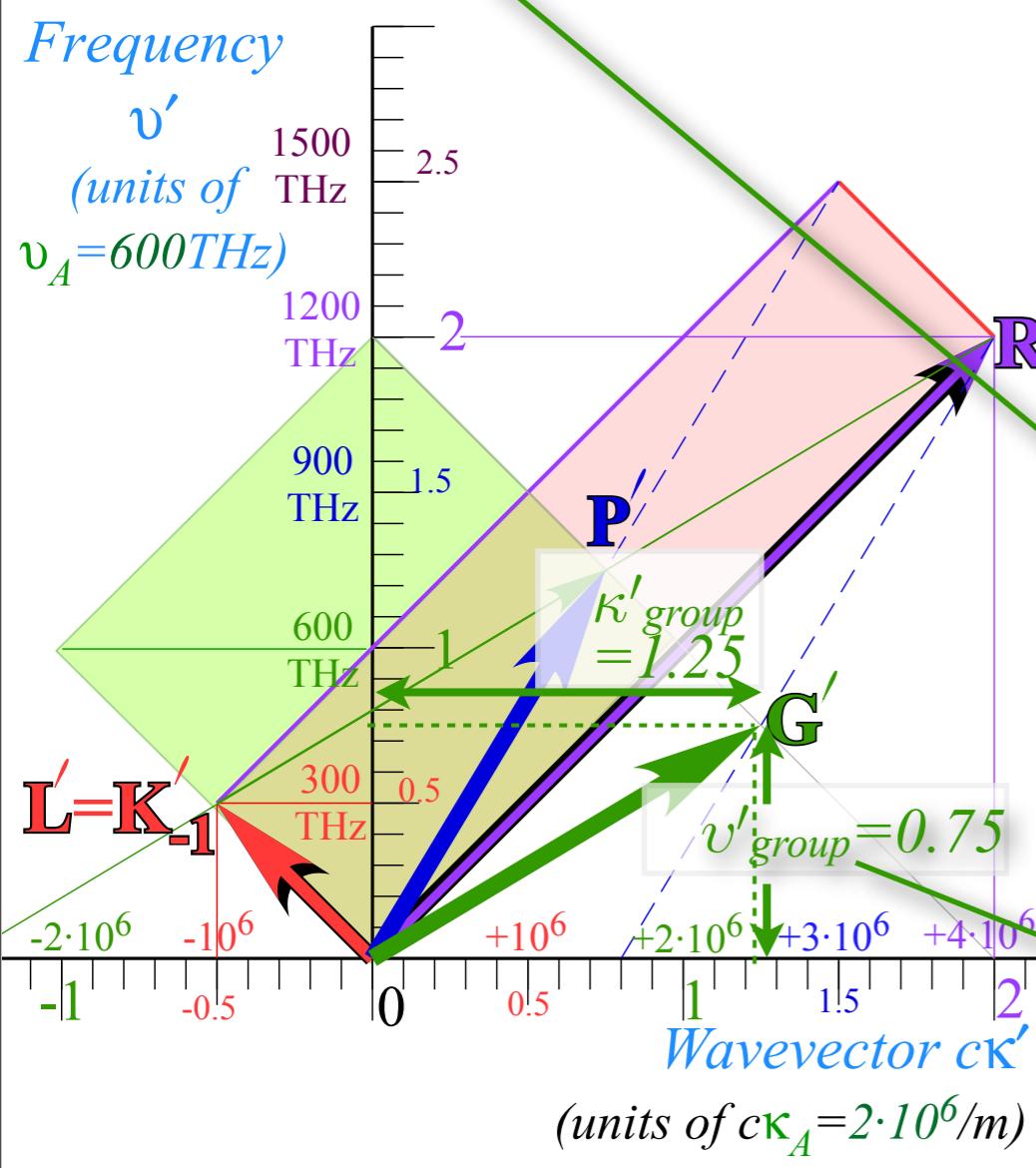
(units of $\lambda_A = 1/2\mu\text{m}$)

<i>phase</i>	$b_{RED}^{Doppler}$	$\frac{c}{V_{phase}}$	κ_{phase}	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
<i>group</i>	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
<i>rapidity</i> ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
<i>value for</i> $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

$$\mathbf{G}' \begin{pmatrix} c\kappa'_{group} \\ v'_{group} \end{pmatrix} = v_A \begin{pmatrix} \cosh \rho \\ \sinh \rho \end{pmatrix} = v_A \begin{pmatrix} 5/4 \\ 3/4 \end{pmatrix}$$

Group frequency
 $v'_{group} = v_A \sinh \rho = 3/4 = 0.75$

flips to
Group period $\tau = 1/v_A$
 $\tau'_{group} = \tau_A \text{csch} \rho = 4/3 = 1.33$

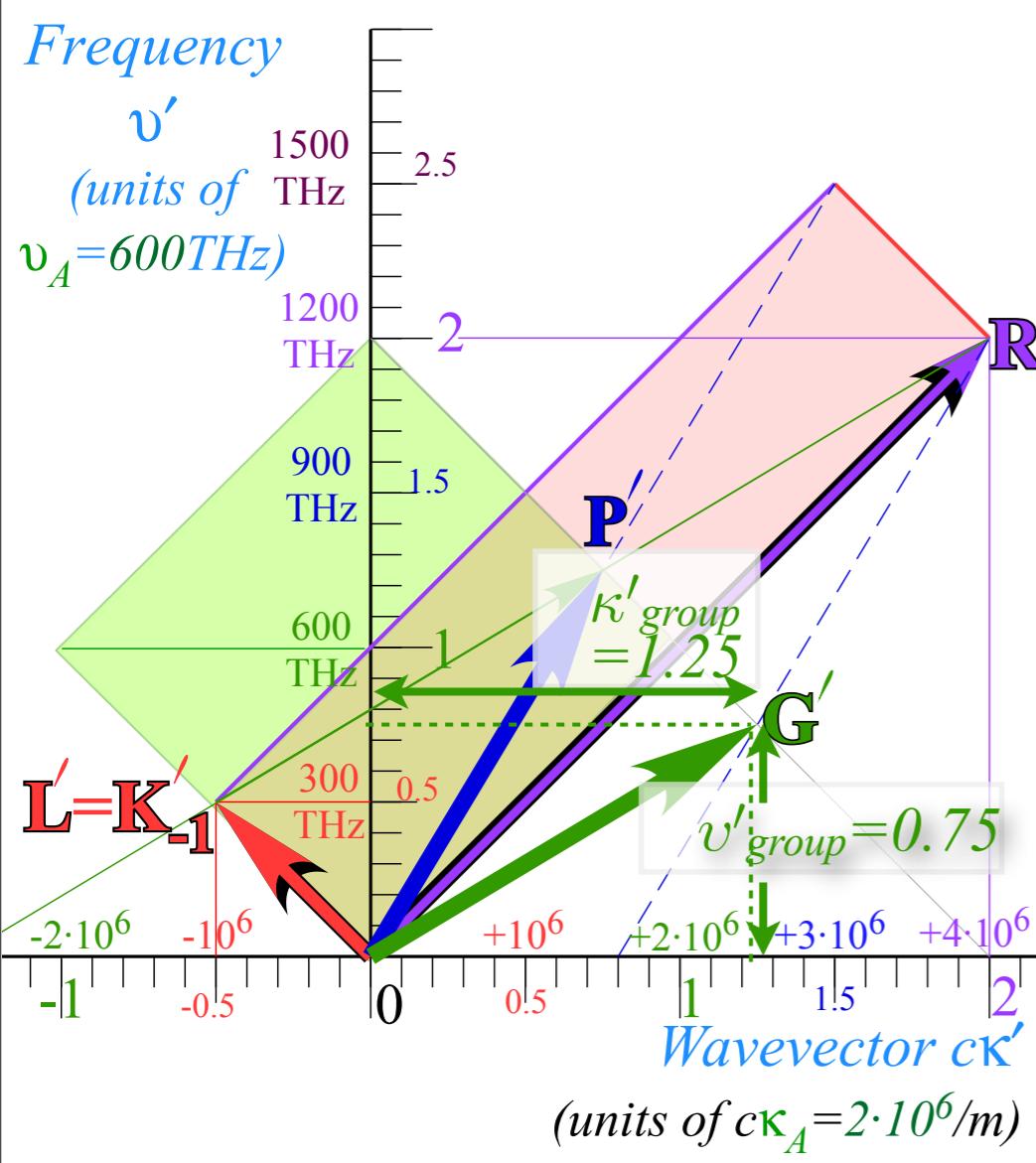


phase	$b_{RED}^{Doppler}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
group	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

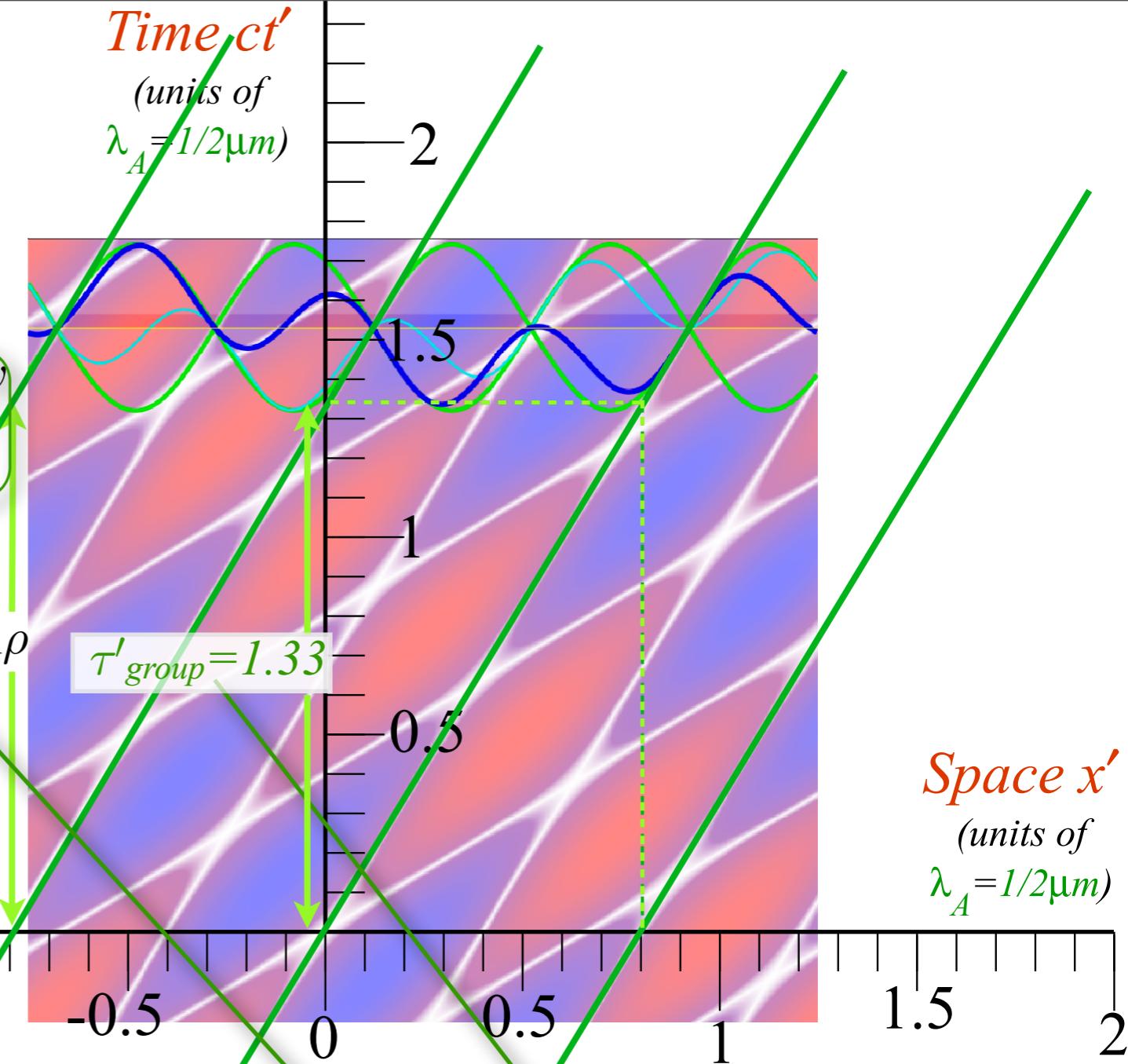
$$\mathbf{G}' \begin{pmatrix} c\kappa'_{group} \\ v'_{group} \end{pmatrix} = v_A \begin{pmatrix} \cosh \rho \\ \sinh \rho \end{pmatrix} = v_A \begin{pmatrix} 5/4 \\ 3/4 \end{pmatrix}$$

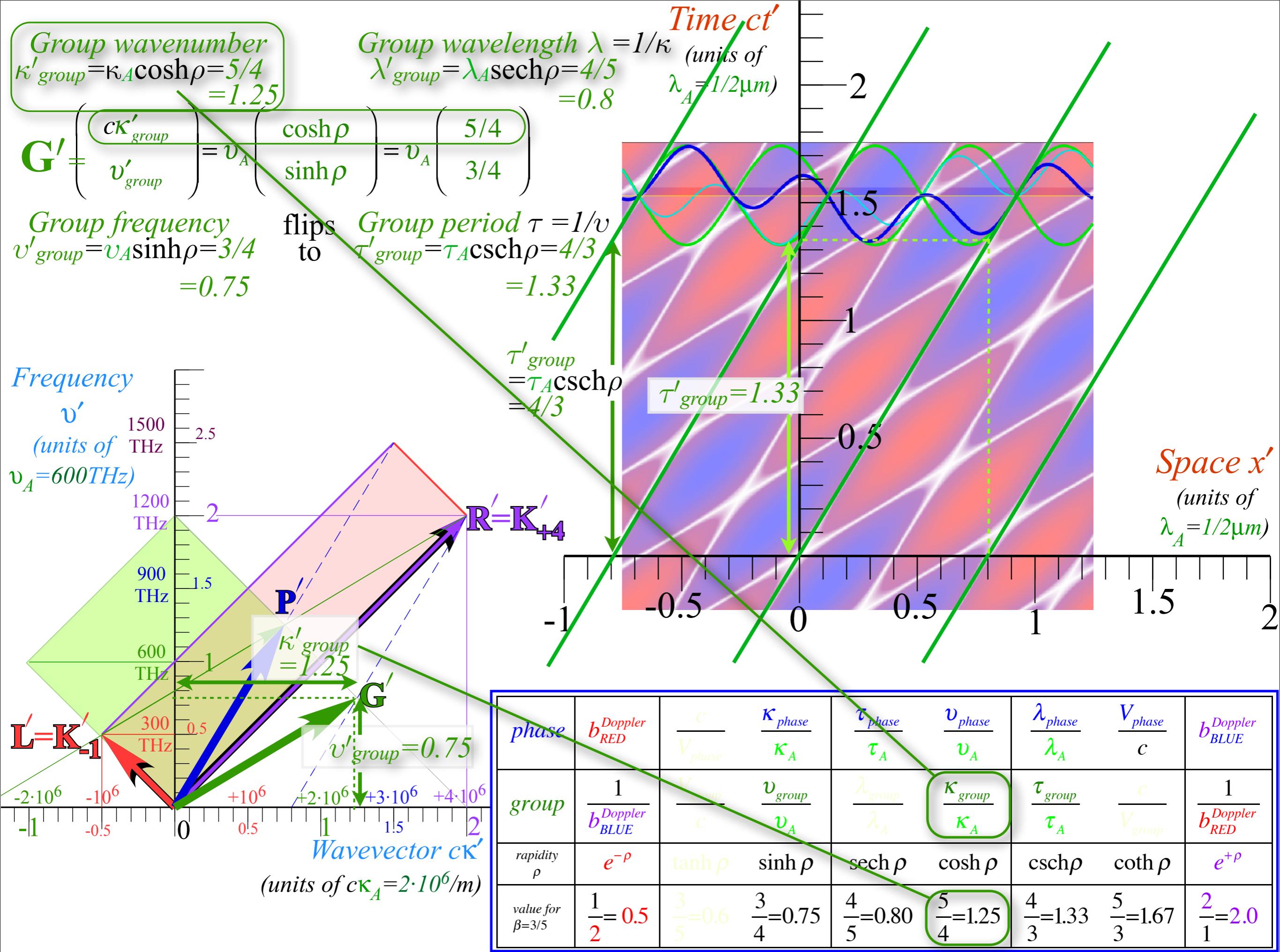
Group frequency
 $v'_{group} = v_A \sinh \rho = 3/4 = 0.75$

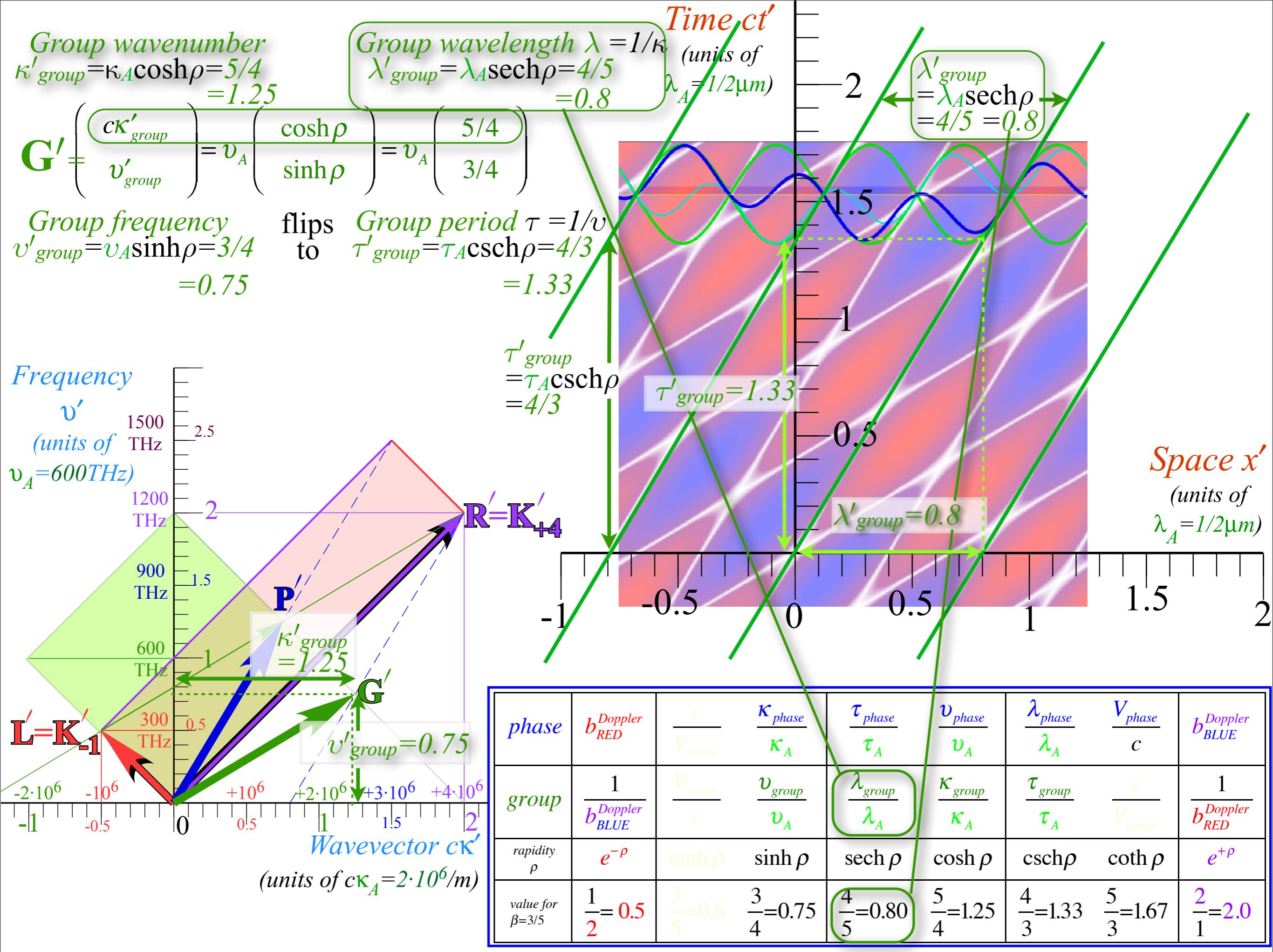
flips to
Group period $\tau = 1/v_A$
 $\tau'_{group} = \tau_A \operatorname{csch} \rho = 4/3 = 1.33$



phase	$b_{RED}^{Doppler}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
group	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$







Group wavenumber
 $\kappa'_{group} = \kappa_A \cosh \rho = 5/4 = 1.25$

$$\mathbf{G}' = \begin{pmatrix} c\mathbf{k}'_{group} \\ v'_{group} \end{pmatrix} = v_A \begin{pmatrix} \dots \end{pmatrix}$$

Group frequency
 $v'_{group} = v_A \sinh \rho = 3/4$
 $= 0.75$

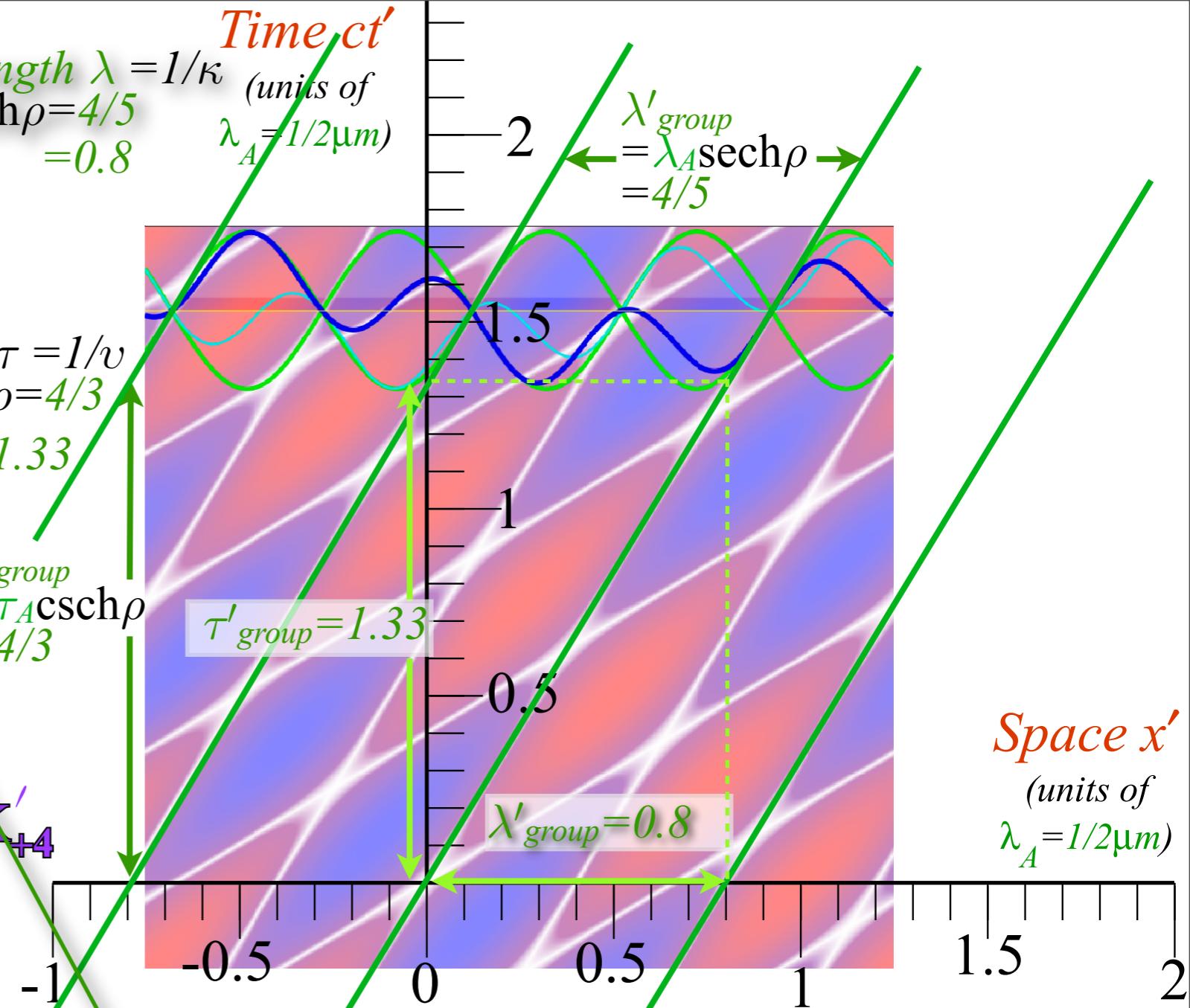
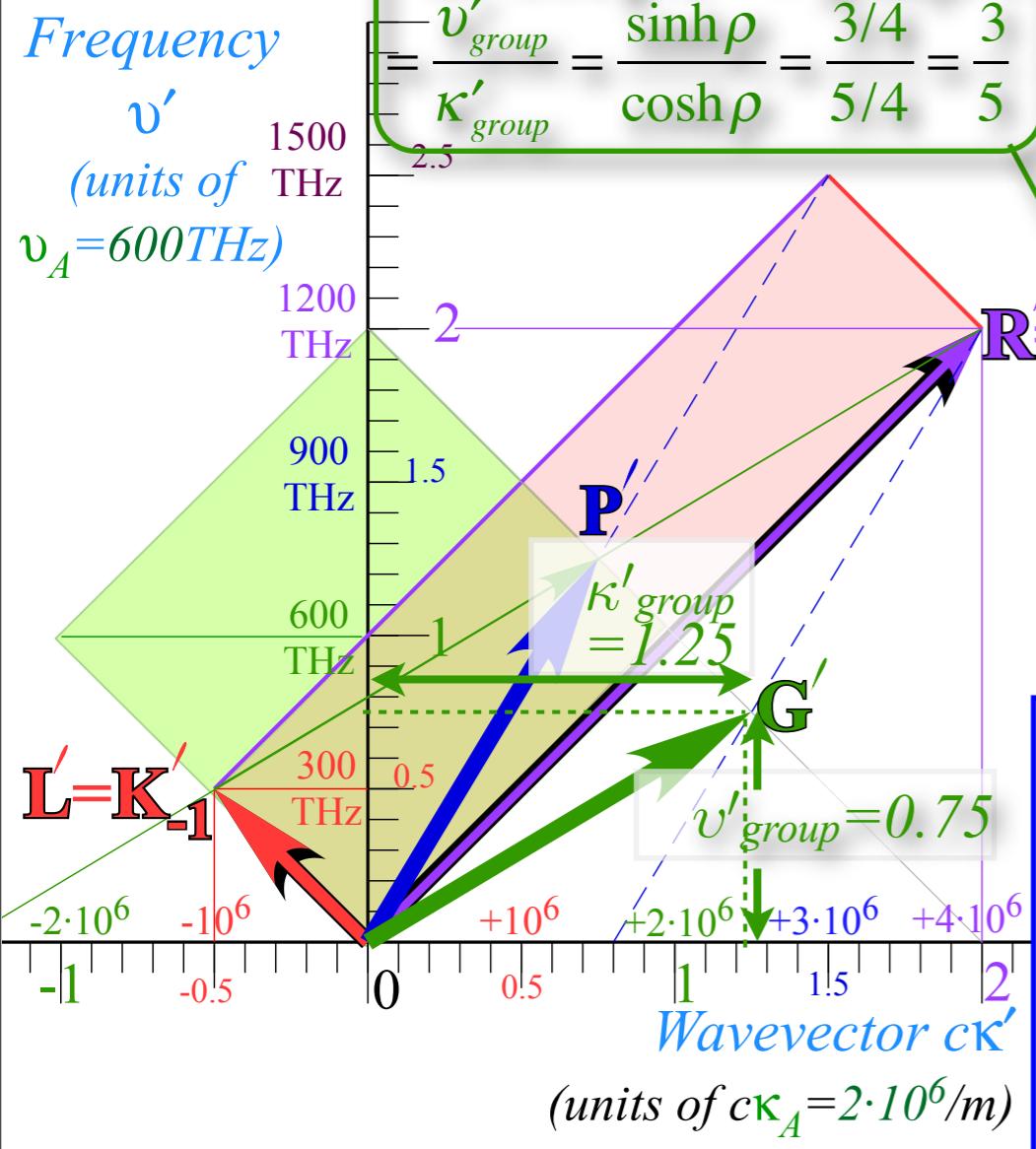
$$Group\ wavelength\ \lambda = 1/\kappa \quad (units\ of\ Time/ct')$$

$$\lambda'_{group} = \lambda_A \operatorname{sech} \rho = 4/5 = 0.8 \quad \lambda_A = 1/2 \mu m$$

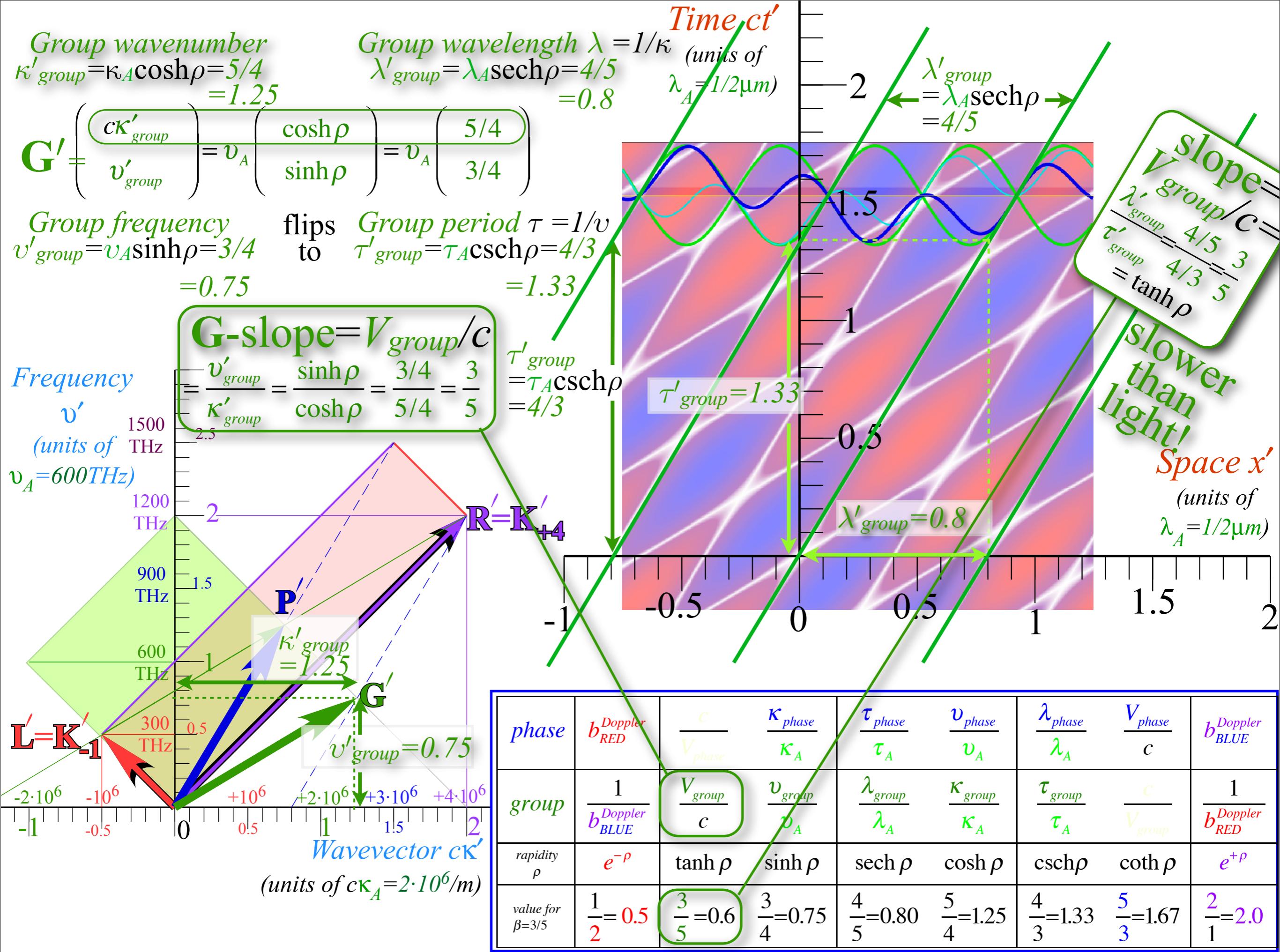
$$\begin{pmatrix} 0 \\ 0 \end{pmatrix} = v_A \begin{pmatrix} 5/4 \\ 3/4 \end{pmatrix}$$

Group period $\tau = 1/v$
 $\tau'_{group} = \tau_A \operatorname{csch} \rho = 4/3$

$$= V_{group}/c$$



<i>phase</i>	$b_{RED}^{Doppler}$	$\frac{c}{V_{phase}}$	κ_{phase}	τ_{phase}	v_{phase}	λ_{phase}	V_{phase}	$b_{BLUE}^{Doppler}$
<i>group</i>	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{V_{group}}{c}$	v_{group}	λ_{group}	κ_{group}	τ_{group}	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
<i>rapidity</i> ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
<i>value for</i> $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

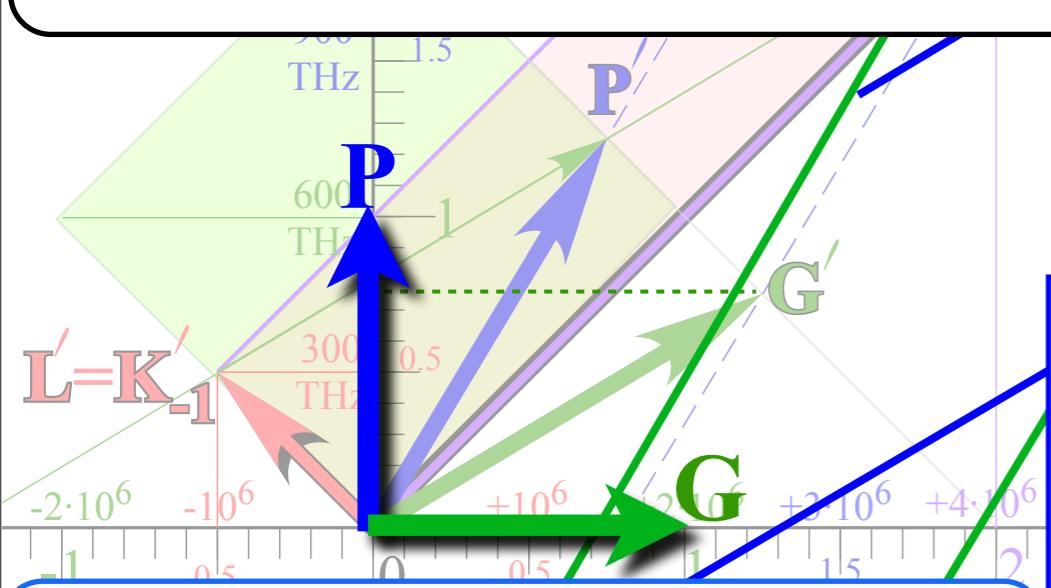


Lorentz transformations...

write \mathbf{G}' and \mathbf{P}' in terms of \mathbf{G} and \mathbf{P} using $\cosh\rho$ and $\sinh\rho$

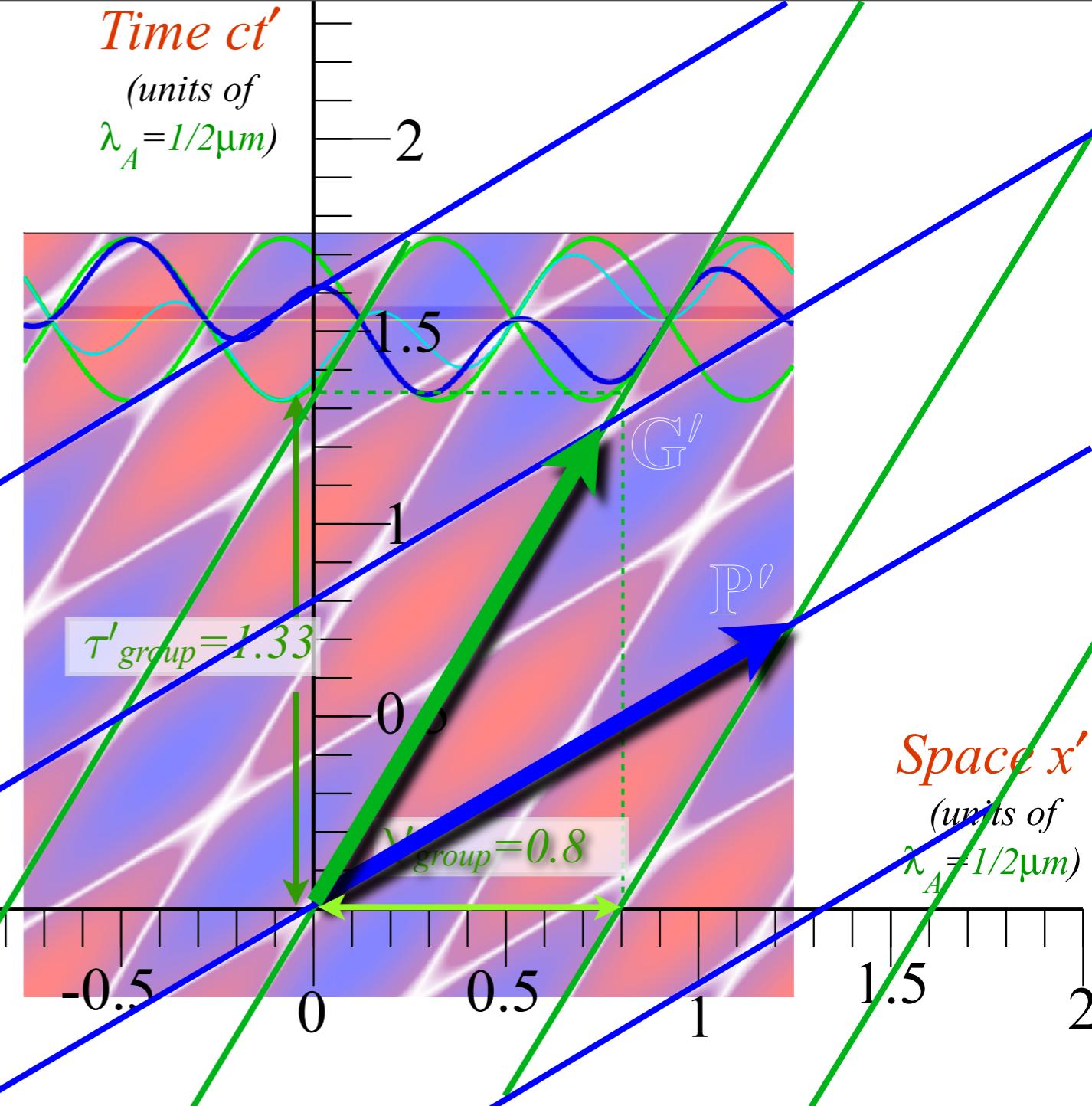
$$\begin{aligned}\mathbf{G}' &= \begin{pmatrix} cK'_{group} \\ v'_{group} \end{pmatrix} = v_A \begin{pmatrix} \cosh\rho \\ \sinh\rho \end{pmatrix} = v_A \begin{pmatrix} 5/4 \\ 3/4 \end{pmatrix} \\ &= v_A \begin{pmatrix} 1 \\ 0 \end{pmatrix} \cosh\rho + v_A \begin{pmatrix} 0 \\ 1 \end{pmatrix} \sinh\rho \\ \mathbf{G}' &= \mathbf{G} \cosh\rho + \mathbf{P} \sinh\rho\end{aligned}$$

$$\begin{aligned}\mathbf{P}' &= \begin{pmatrix} cK'_{phase} \\ v'_{phase} \end{pmatrix} = v_A \begin{pmatrix} \sinh\rho \\ \cosh\rho \end{pmatrix} = v_A \begin{pmatrix} 3/4 \\ 5/4 \end{pmatrix} \\ &= v_A \begin{pmatrix} 1 \\ 0 \end{pmatrix} \sinh\rho + v_A \begin{pmatrix} 0 \\ 1 \end{pmatrix} \cosh\rho \\ \mathbf{P}' &= \mathbf{G} \sinh\rho + \mathbf{P} \cosh\rho\end{aligned}$$



$$\begin{pmatrix} \cosh\rho & \sinh\rho \\ \sinh\rho & \cosh\rho \end{pmatrix} \text{ Lorentz transform matrix}$$

phase	$b_{RED}^{Doppler}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
group	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh\rho$	$\sinh\rho$	$\operatorname{sech}\rho$	$\cosh\rho$	$\operatorname{csch}\rho$	$\coth\rho$	$e^{+\rho}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

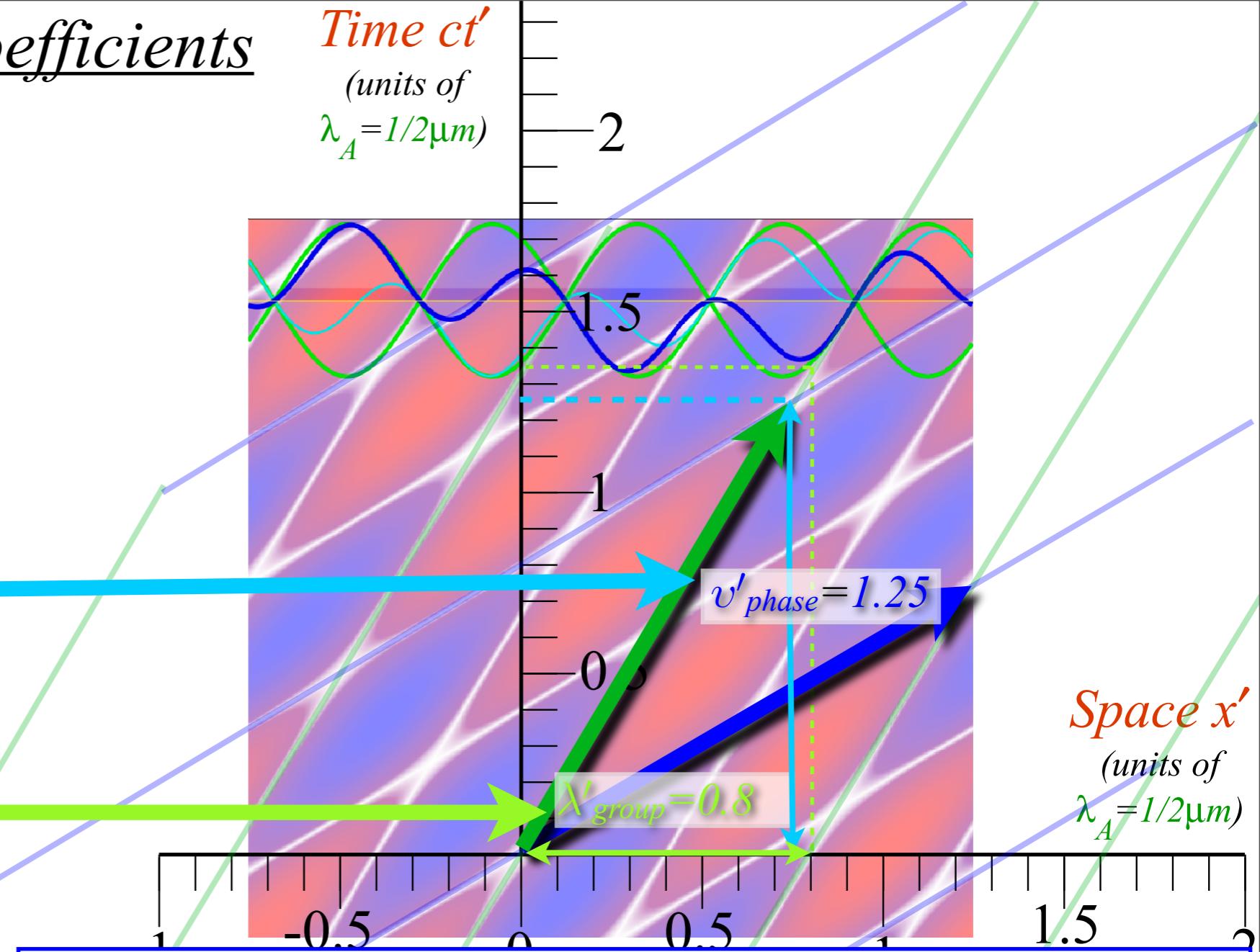


Two Famous-Name Coefficients

This number
is called an: **Einstein time-dilation**
(dilated by 25% here)

This number
is called a: **Lorentz length-contraction**
(contracted by 20% here)

Old-Fashioned Notation

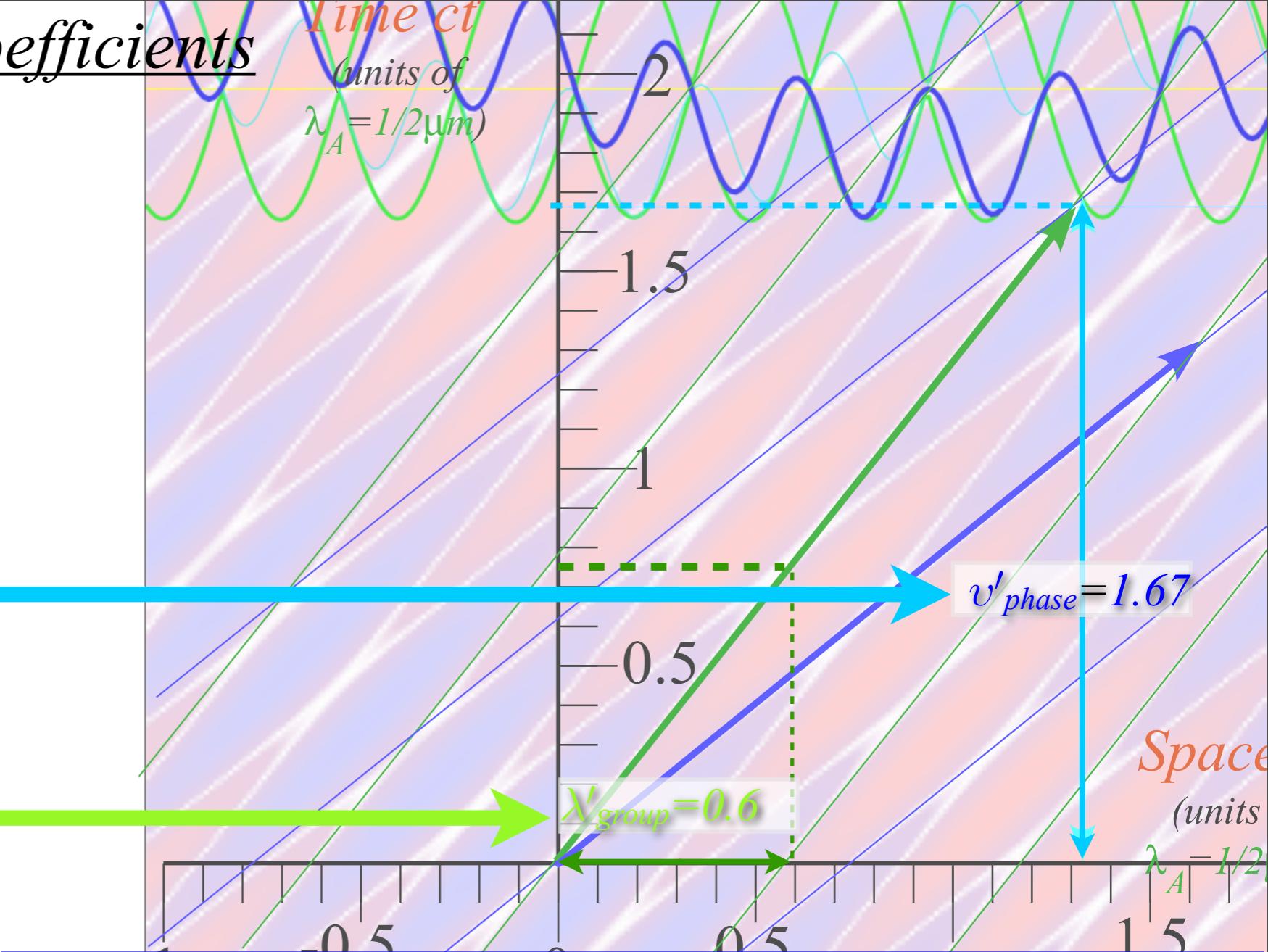


phase	$b_{\text{Doppler RED}}$	$\frac{c}{V_{\text{phase}}}$	$\frac{\kappa_{\text{phase}}}{\kappa_A}$	$\frac{\tau_{\text{phase}}}{\tau_A}$	$\frac{v_{\text{phase}}}{v_A}$	$\frac{\lambda_{\text{phase}}}{\lambda_A}$	$\frac{V_{\text{phase}}}{c}$	$b_{\text{Doppler BLUE}}$
group	$\frac{1}{b_{\text{Doppler BLUE}}}$	$\frac{V_{\text{group}}}{c}$	$\frac{v_{\text{group}}}{v_A}$	$\frac{\lambda_{\text{group}}}{\lambda_A}$	$\frac{\kappa_{\text{group}}}{\kappa_A}$	$\frac{\tau_{\text{group}}}{\tau_A}$	$\frac{c}{V_{\text{group}}}$	$\frac{1}{b_{\text{Doppler RED}}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

Two Famous-Name Coefficients

This number  is called an: Einstein time-dilation (dilated by 67% here)

This number  is called a: Lorentz length-contraction (contracted by 40% here)



phase	$b_{Doppler RED}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler BLUE}$
group	$\frac{1}{b_{Doppler BLUE}}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{Doppler RED}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=4/5$	$\frac{1}{3}=0.33$	$\frac{4}{5}=0.8$	$\frac{4}{3}=1.33$	$\frac{3}{5}=0.60$	$\frac{5}{3}=1.67$	$\frac{3}{4}=0.75$	$\frac{5}{4}=1.25$	$\frac{3}{1}=3.0$

Old-Fashioned Notation



$$\beta \equiv \frac{u}{c}$$

$$\sqrt{\frac{1-\beta}{1+\beta}}$$

$$\frac{\beta}{1}$$

$$\frac{1}{\sqrt{\beta^2-1}}$$

$$\frac{\sqrt{1-\beta^2}}{1}$$

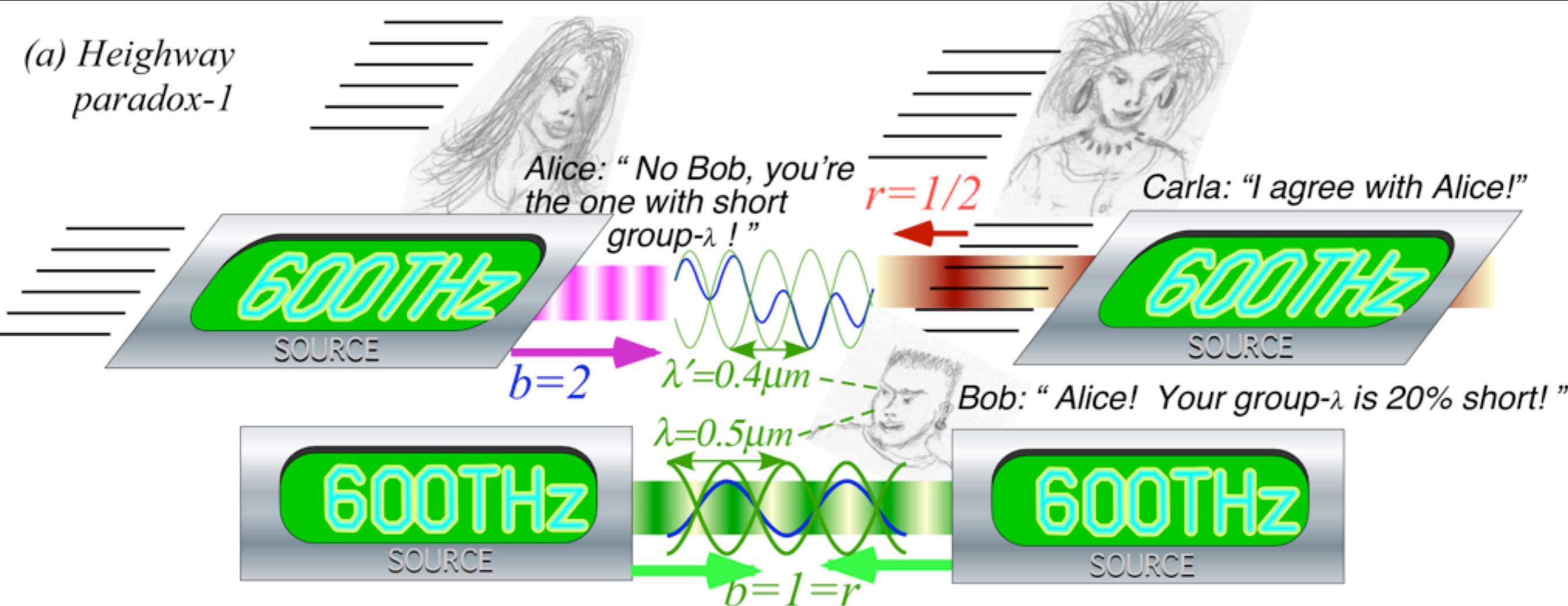
$$\frac{1}{\sqrt{1-\beta^2}}$$

$$\frac{\sqrt{\beta^2-1}}{1}$$

$$\frac{1}{\beta}$$

$$\sqrt{\frac{1+\beta}{1-\beta}}$$

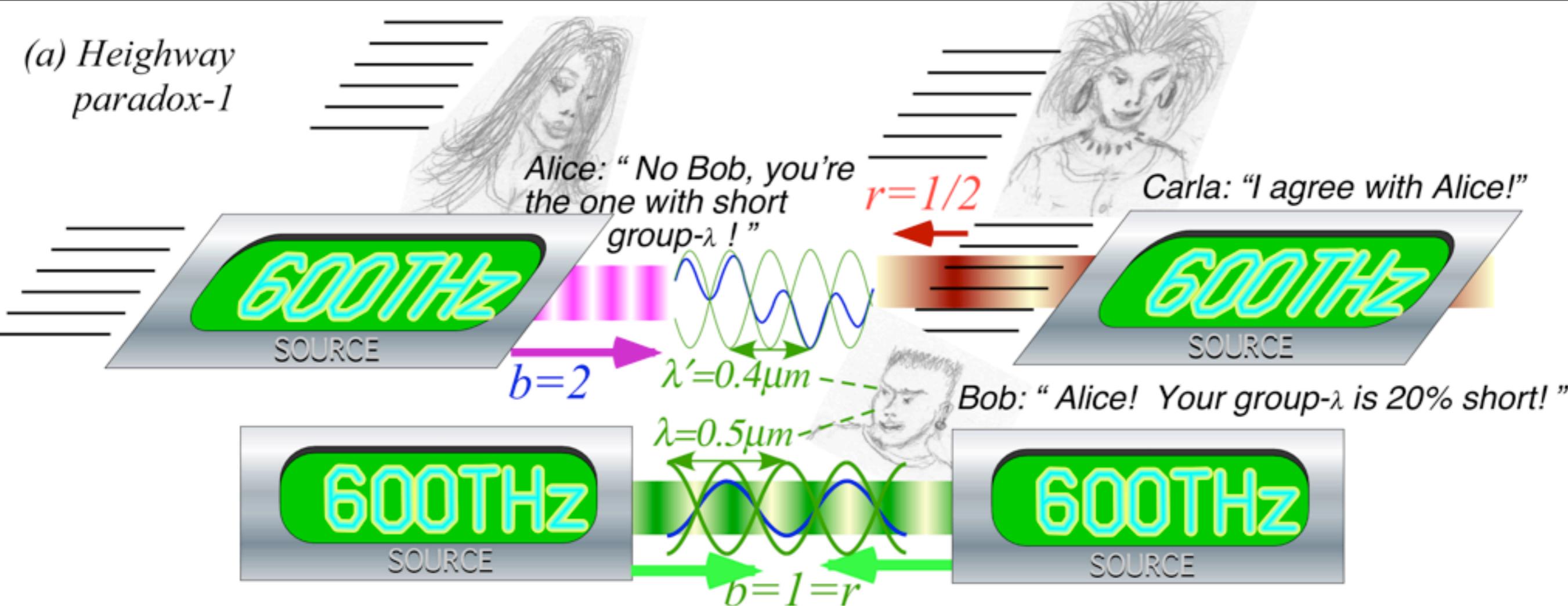
(a) Heighway paradox-1



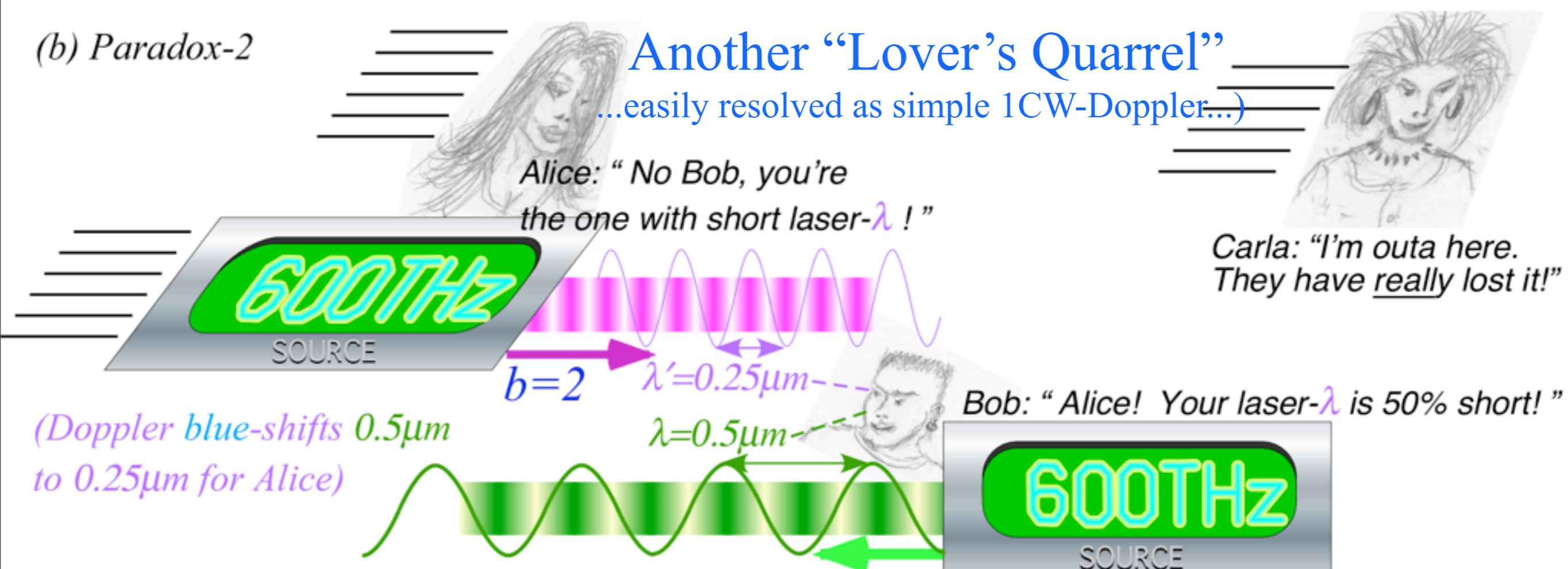
A “Lover’s Quarrel” about a 20% Lorentz contraction $\lambda'_{group} = 0.8 \lambda_A$
(You’re short! No, YOU'RE short!!, etc.)
...(The worst kind of quarrel is when both are right *and* wrong)

$$\frac{3}{5}c$$

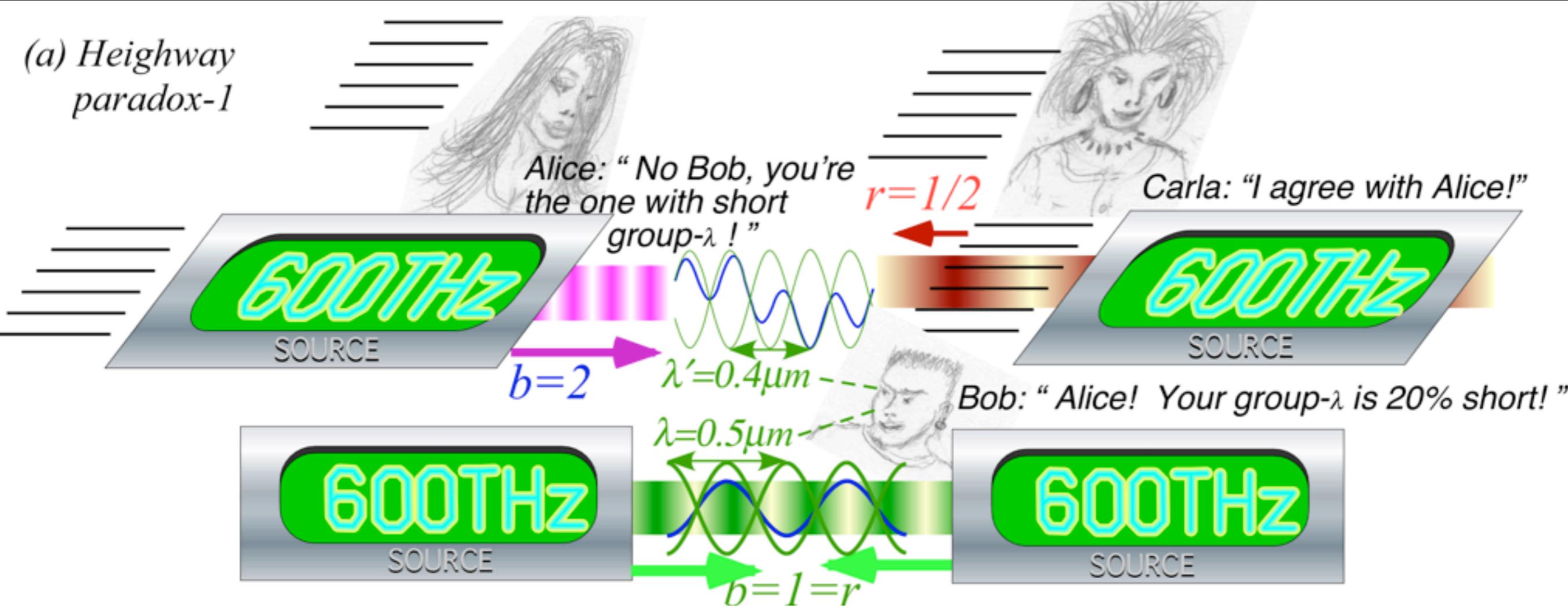
(a) Heighway paradox-1



(b) Paradox-2



(a) Heighway paradox-1

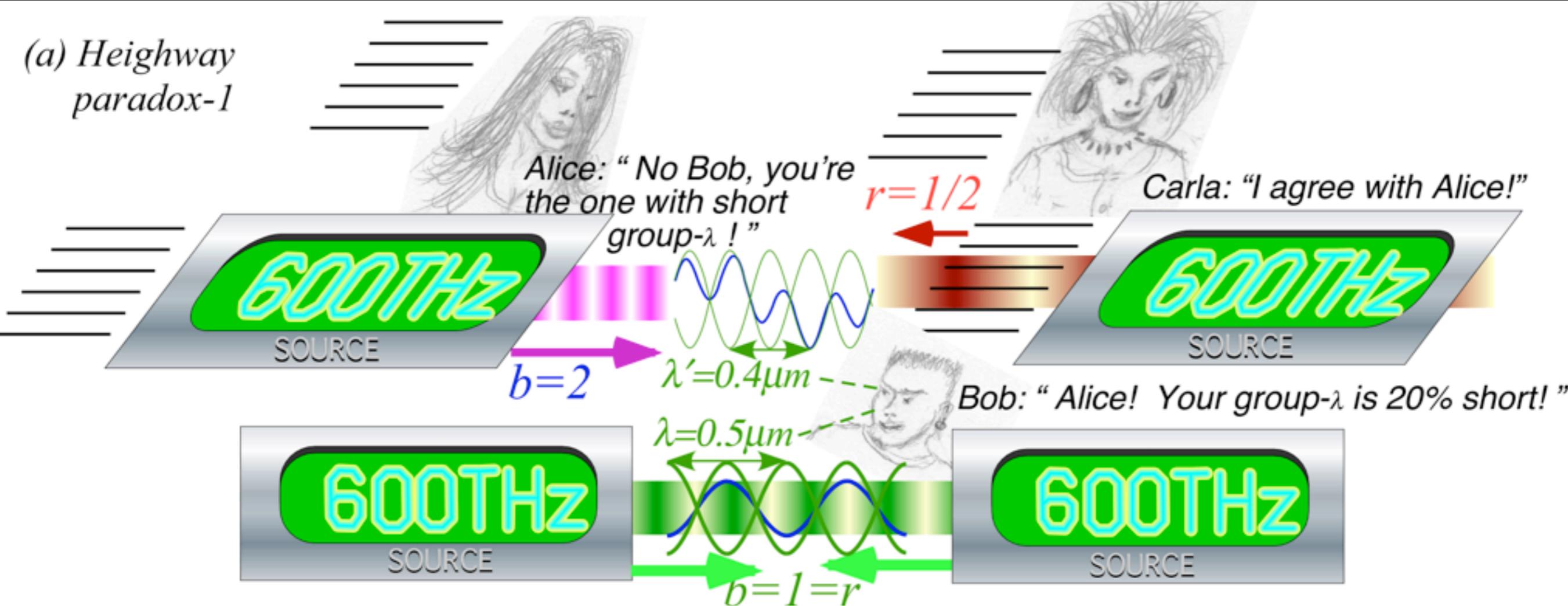


A “Lover’s Quarrel” about a 20% Lorentz contraction $\lambda'_{group} = 0.8 \lambda_A$
(You’re short! No, YOU'RE short!!, etc.)
...(The worst kind of quarrel is when both are right *and* wrong)

So we learn to accept that a group-wave shortens by 20% at this enormous speed of $\frac{3}{5}c$.
Q: But, does the *steel laser cavity* holding the wave *also shorten* by 20%??

A: ...

(a) Heighway paradox-1



A “Lover’s Quarrel” about a 20% Lorentz contraction $\lambda'_{group} = 0.8 \lambda_A$
(You’re short! No, YOU’RE short!!, etc.)
...(The worst kind of quarrel is when both are right *and* wrong)

So we learn to accept that a group-wave shortens by 20% at this enormous speed of $\frac{3}{5}c$.
Q: But, does the *steel laser cavity* holding the wave *also shorten* by 20%??

A: Yes, or else laser does not resonate! *Steel is made of waves, too.*
Contraction is what waves do.

Let's do the A_{lice}B_{ob}C_{arla} problem backwards...

Suppose Bob sees beam of frequency \mathcal{V}_L coming from the *LEFT*
and
opposing beam of frequency \mathcal{V}_R coming from the *RIGHT*.

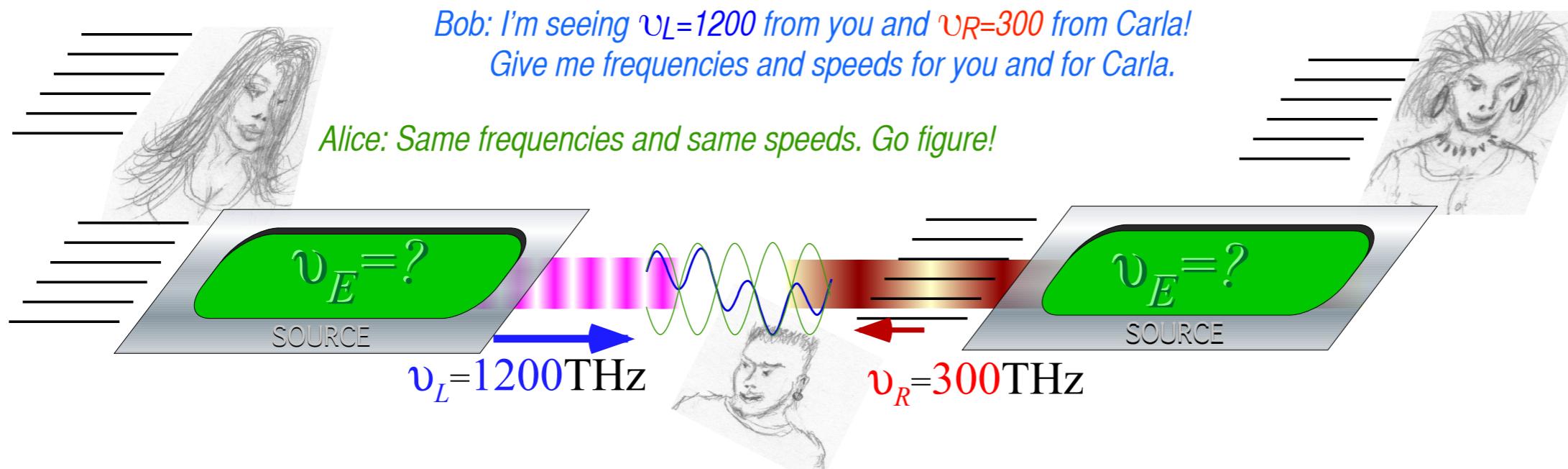
Question 1: To what velocity u_E must Bob accelerate to see beams of *EQUAL* frequency \mathcal{V}_E ?

Question 2: What is frequency \mathcal{V}_E ?

Alice: Hey Bob, speed up and join us!

Bob: I'm seeing $\mathcal{V}_L=1200$ from you and $\mathcal{V}_R=300$ from Carla!
Give me frequencies and speeds for you and for Carla.

Alice: Same frequencies and same speeds. Go figure!



Let's do the A_{lice}B_{ob}C_{arla} problem backwards...

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and
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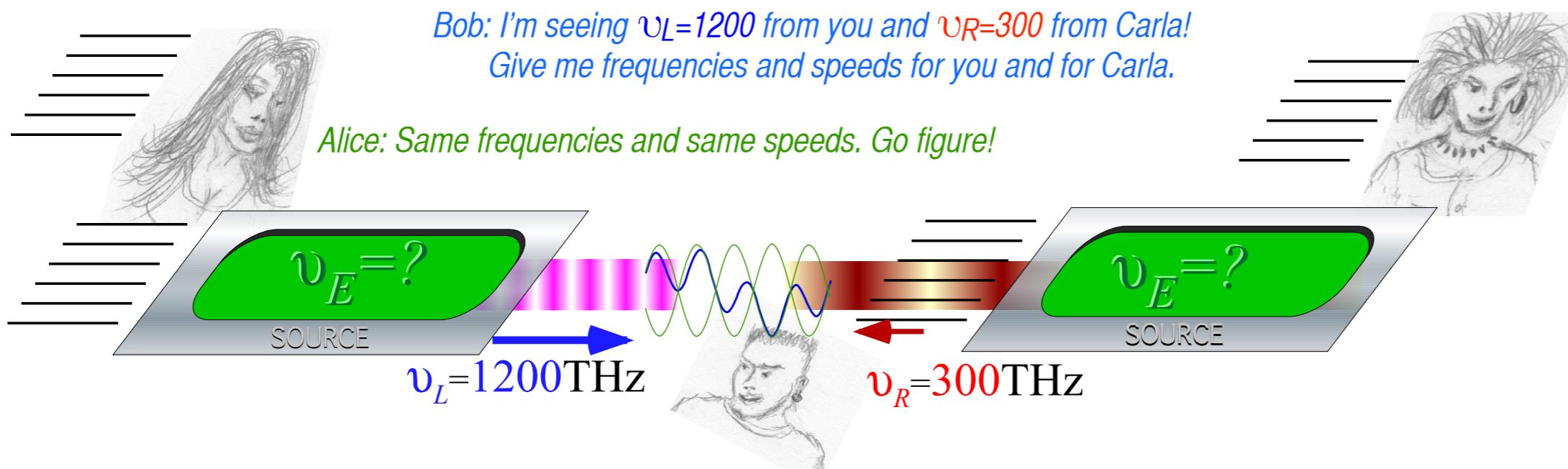
Question 1. has a Jeopardy-style answer-by-question: *What is beam group velocity?*

$$u_E = V_{group} = \frac{v_{group}}{k_{group}} = \frac{(\mathcal{V}_L - \mathcal{V}_R)/2}{(\kappa_L - \kappa_R)/2}$$

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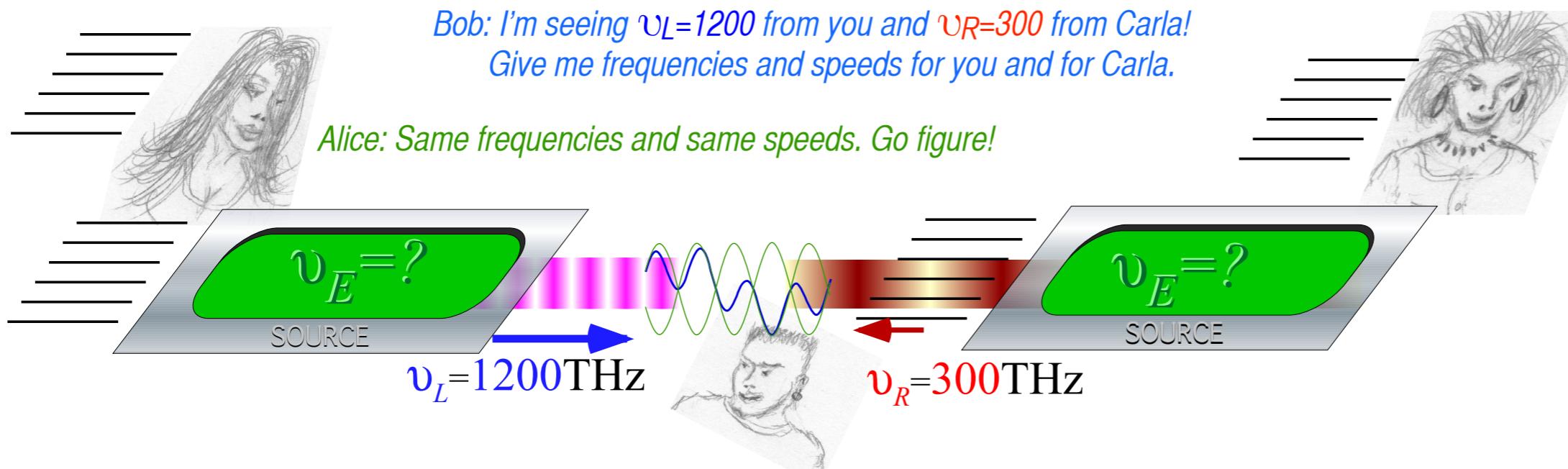
Question 1. has a Jeopardy-style answer-by-question: *What is beam group velocity?*

$$u_E = V_{group} = \frac{v_{group}}{k_{group}} = \frac{(\mathcal{V}_L - \mathcal{V}_R)/2}{(\kappa_L - \kappa_R)/2} = c \frac{(\mathcal{V}_L - \mathcal{V}_R)/2}{(\mathcal{V}_L + \mathcal{V}_R)/2} \text{ where: } \begin{array}{l} \mathcal{V}_L = +c\kappa_L \\ \text{and} \\ \mathcal{V}_R = -c\kappa_R \end{array}$$

Alice: Hey Bob, speed up and join us!

Bob: I'm seeing $\mathcal{V}_L=1200$ from you and $\mathcal{V}_R=300$ from Carla!
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and

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Question 1: To what velocity u_E must Bob accelerate to see beams of *EQUAL* frequency \mathcal{V}_E ?

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Question 1. has a Jeopardy-style answer-by-question: *What is beam group velocity?*

$$\frac{\text{Difference Mean}}{\text{Arithmetic Mean}} =$$

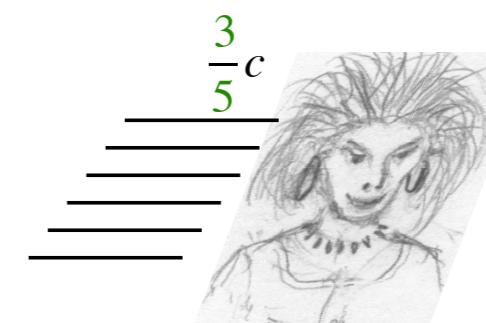
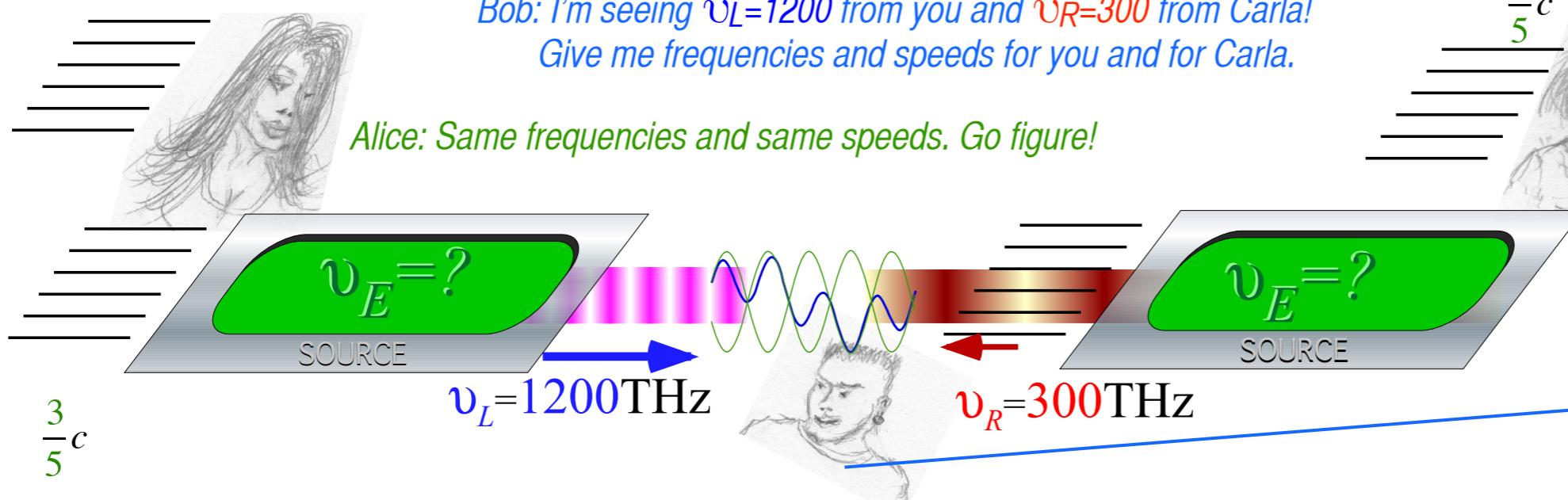
$$u_E = V_{group} = \frac{v_{group}}{\kappa_{group}} = \frac{(\mathcal{V}_L - \mathcal{V}_R)/2}{(\kappa_L - \kappa_R)/2} = c \frac{(\mathcal{V}_L - \mathcal{V}_R)/2}{(\mathcal{V}_L + \mathcal{V}_R)/2} \quad \text{where: } \begin{array}{l} \mathcal{V}_L = +c\kappa_L \\ \text{and} \\ \mathcal{V}_R = -c\kappa_R \end{array}$$

$$\frac{1200 - 300}{1200 + 300} c = \frac{900}{1500} c = \frac{3}{5} c$$

Alice: Hey Bob, speed up and join us!

Bob: I'm seeing $\mathcal{V}_L=1200$ from you and $\mathcal{V}_R=300$ from Carla!
Give me frequencies and speeds for you and for Carla.

Alice: Same frequencies and same speeds. Go figure!



Bob: OK. Now I know
they're both going $\frac{3}{5} c$
relative to me.

Let's do the A_{lice}B_{ob}C_{arla} problem backwards...

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Question 1: To what velocity u_E must Bob accelerate to see beams of *EQUAL* frequency \mathcal{V}_E ?

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Question 1. has a Jeopardy-style answer-by-question: *What is beam group velocity?*

$$\frac{\text{Difference Mean}}{\text{Arithmetic Mean}} =$$

$$u_E = V_{group} = \frac{v_{group}}{\kappa_{group}} = \frac{(\mathcal{V}_L - \mathcal{V}_R)/2}{(\kappa_L - \kappa_R)/2} = c \frac{(\mathcal{V}_L - \mathcal{V}_R)/2}{(\mathcal{V}_L + \mathcal{V}_R)/2} \quad \text{where: } \begin{array}{l} \mathcal{V}_L = +c\kappa_L \\ \text{and} \\ \mathcal{V}_R = -c\kappa_R \end{array}$$

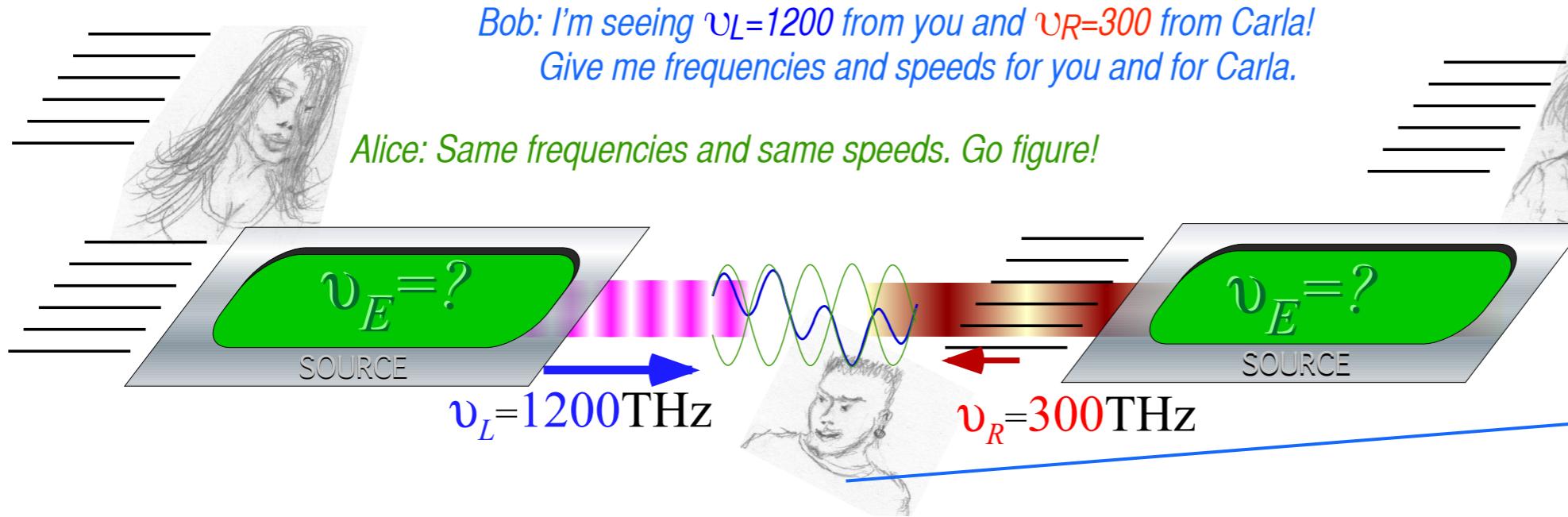
$$\frac{1200 - 300}{1200 + 300} c = \frac{900}{1500} c = \frac{3}{5} c$$

Question 2. ...similarly: *What \mathcal{V}_E is blue-shift $b\mathcal{V}_R$ of \mathcal{V}_R AND red-shift $r\mathcal{V}_L = \mathcal{V}_L/b$ of \mathcal{V}_L ?*

Alice: Hey Bob, speed up and join us!

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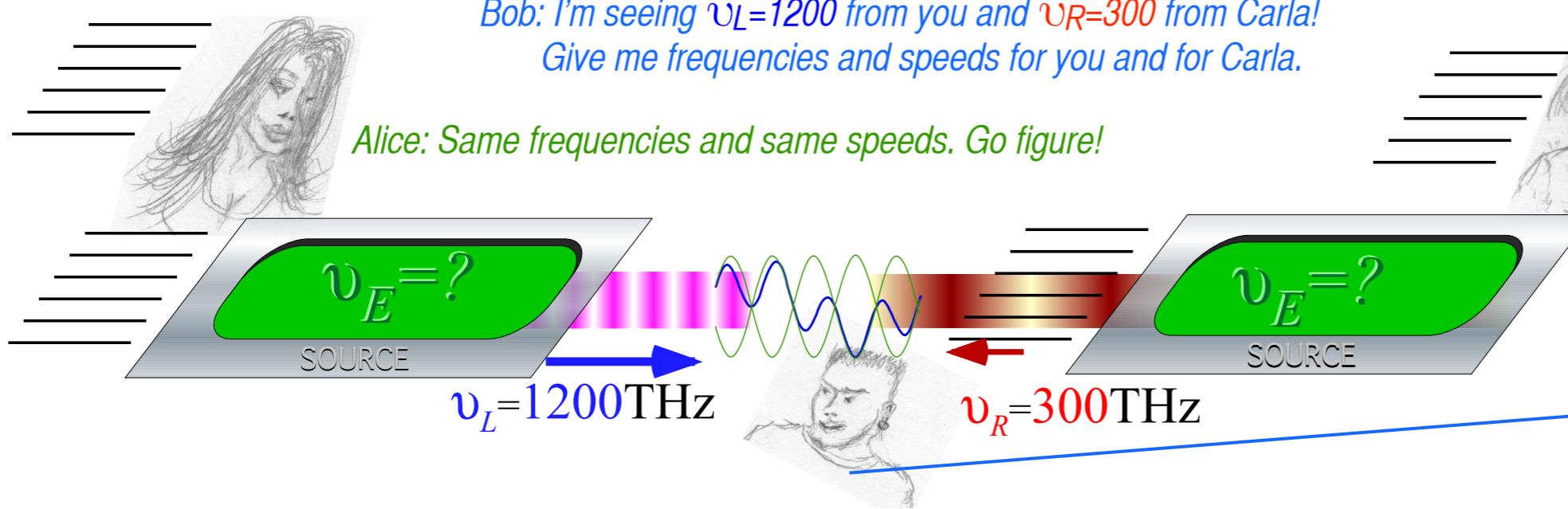
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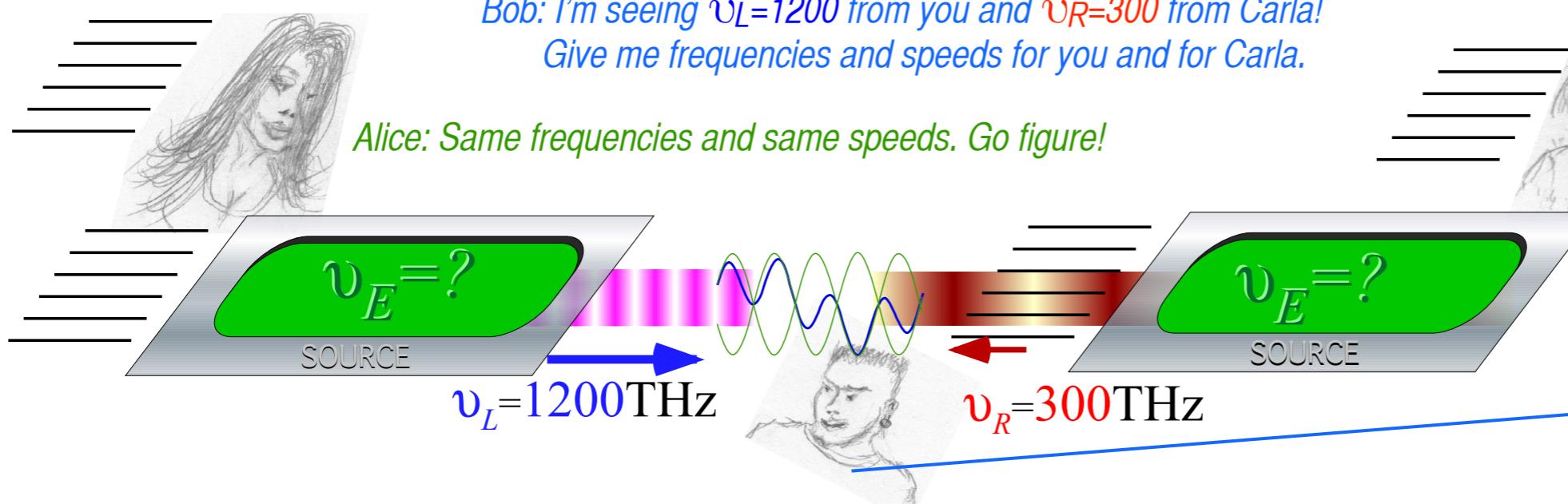
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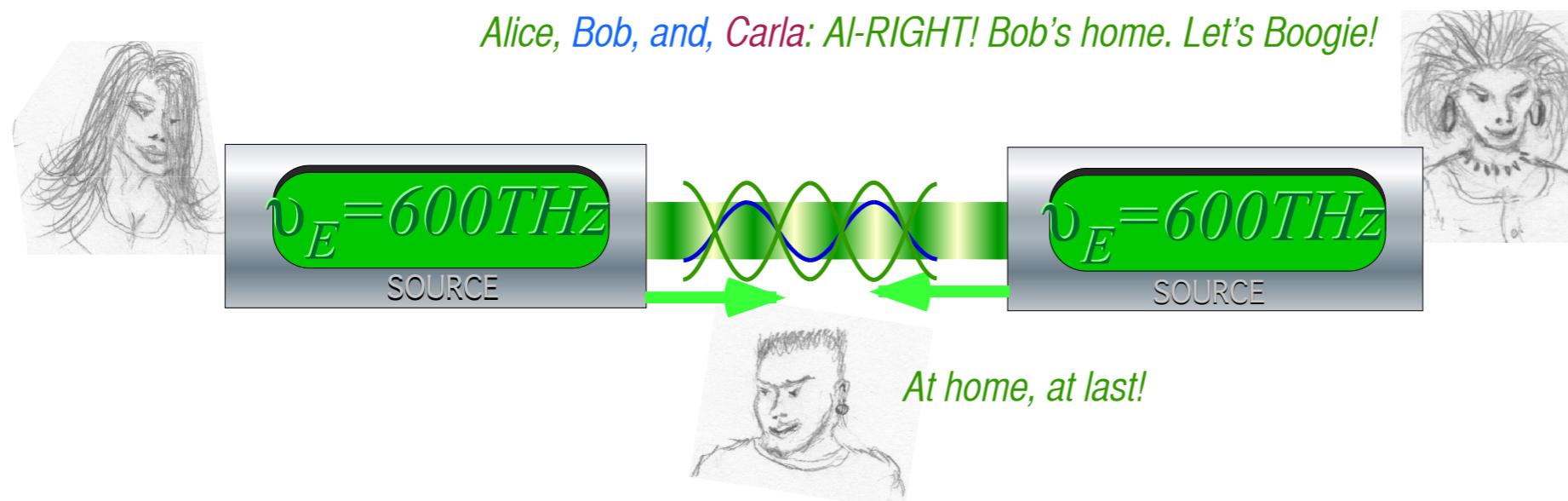
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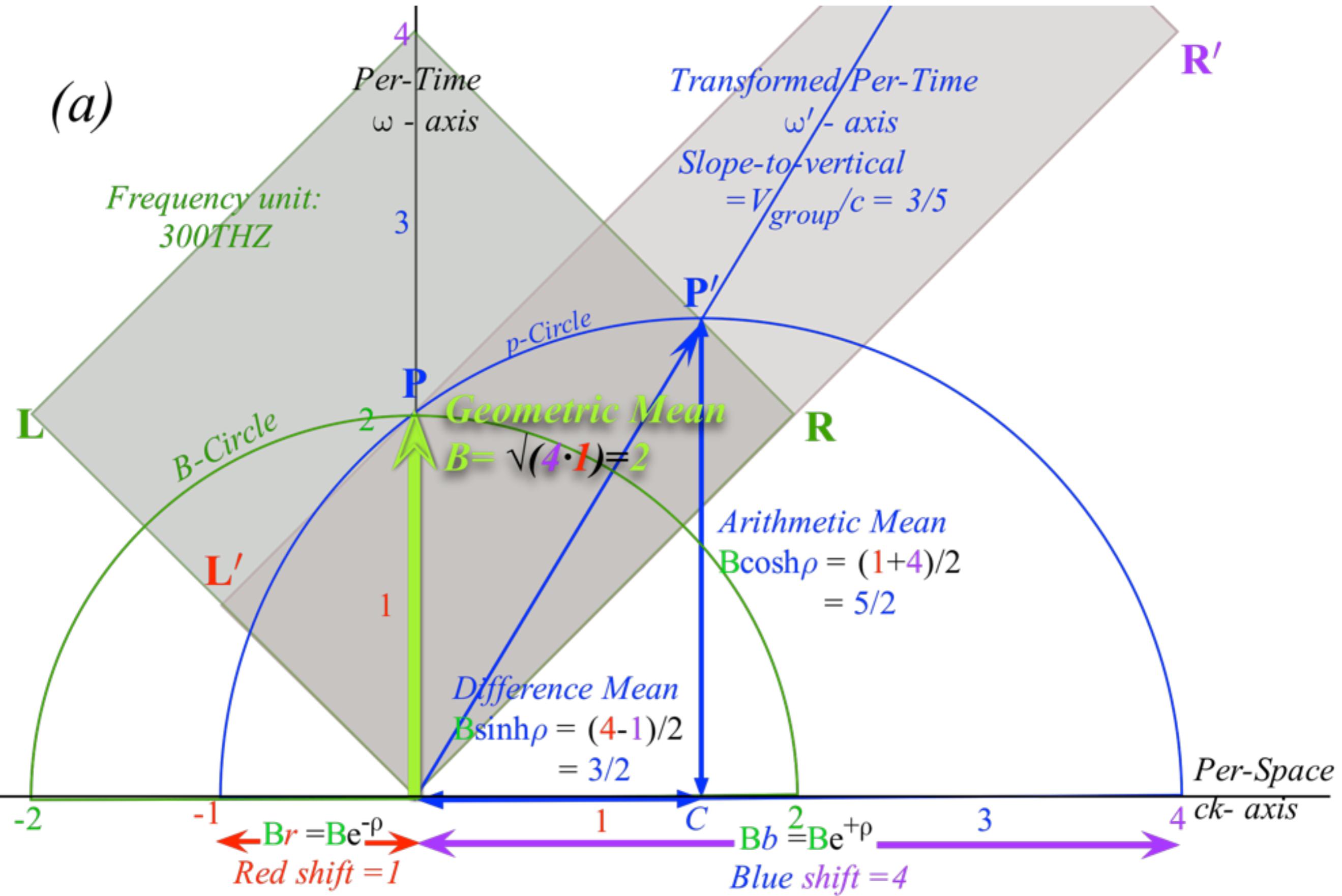
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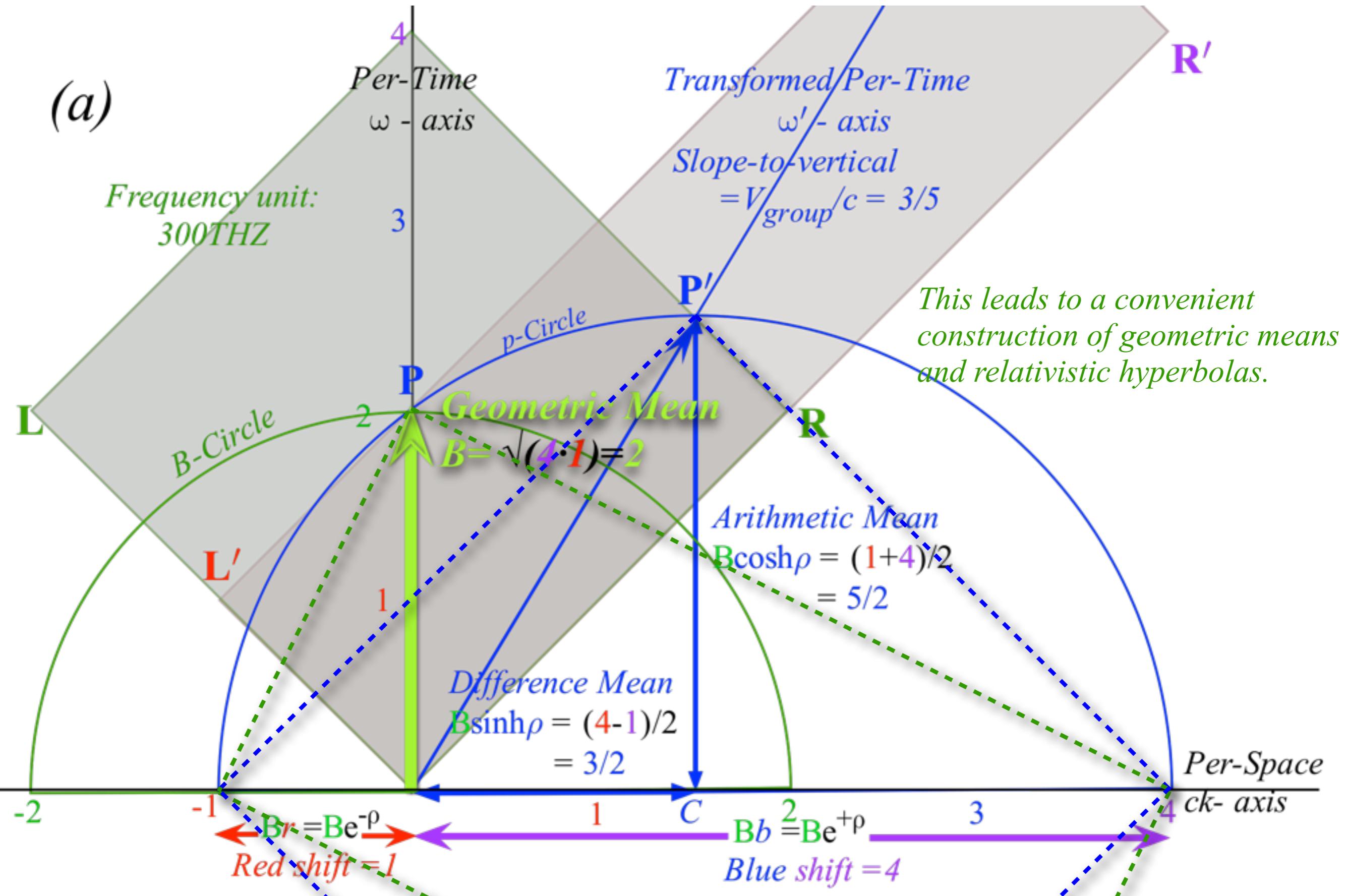
Thales Mean Geometry (600BCE)

helps “Relativity”



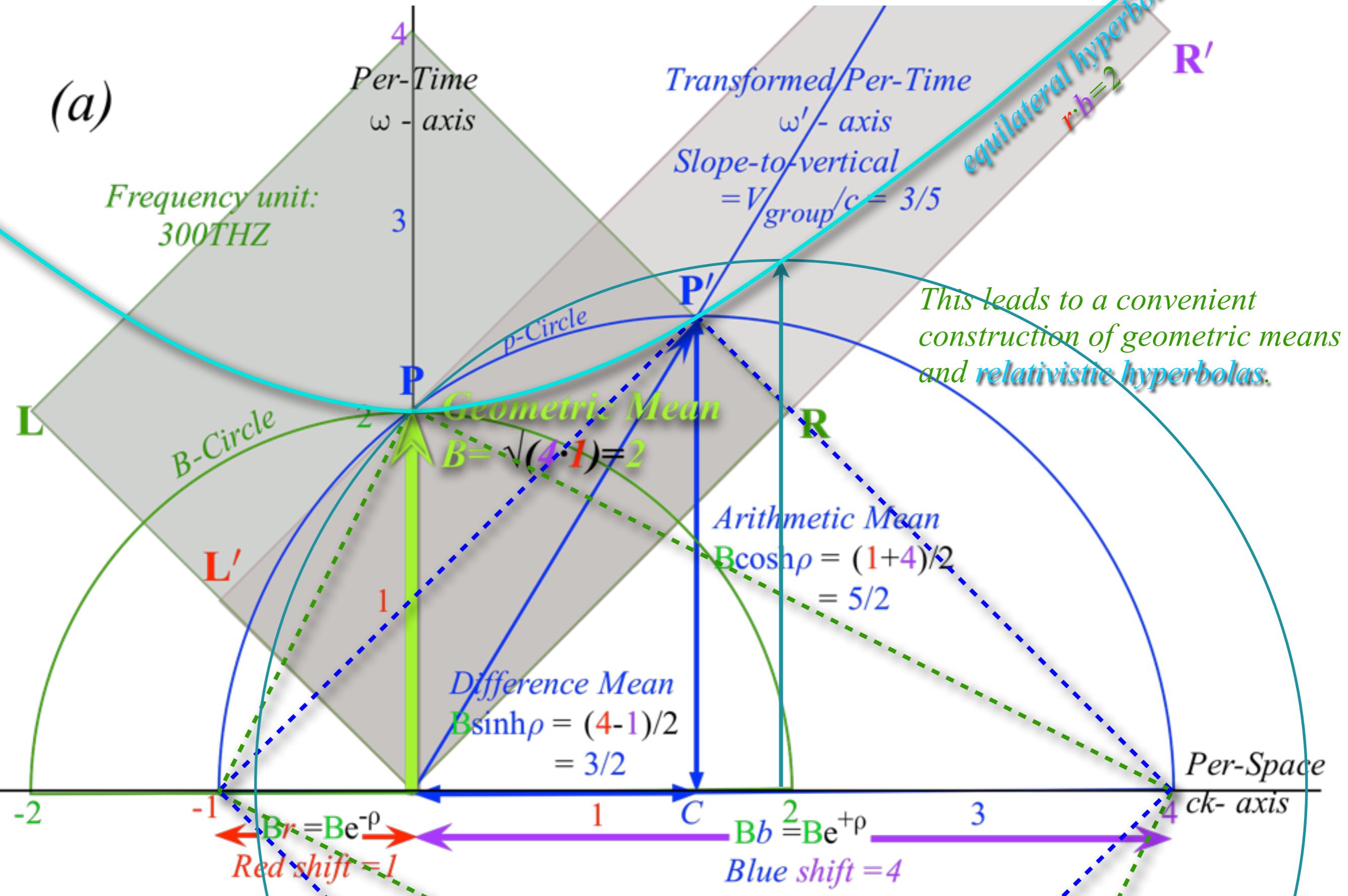
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helps “Relativity” Thales showed a circle diameter subtends a right angle with any circle point P



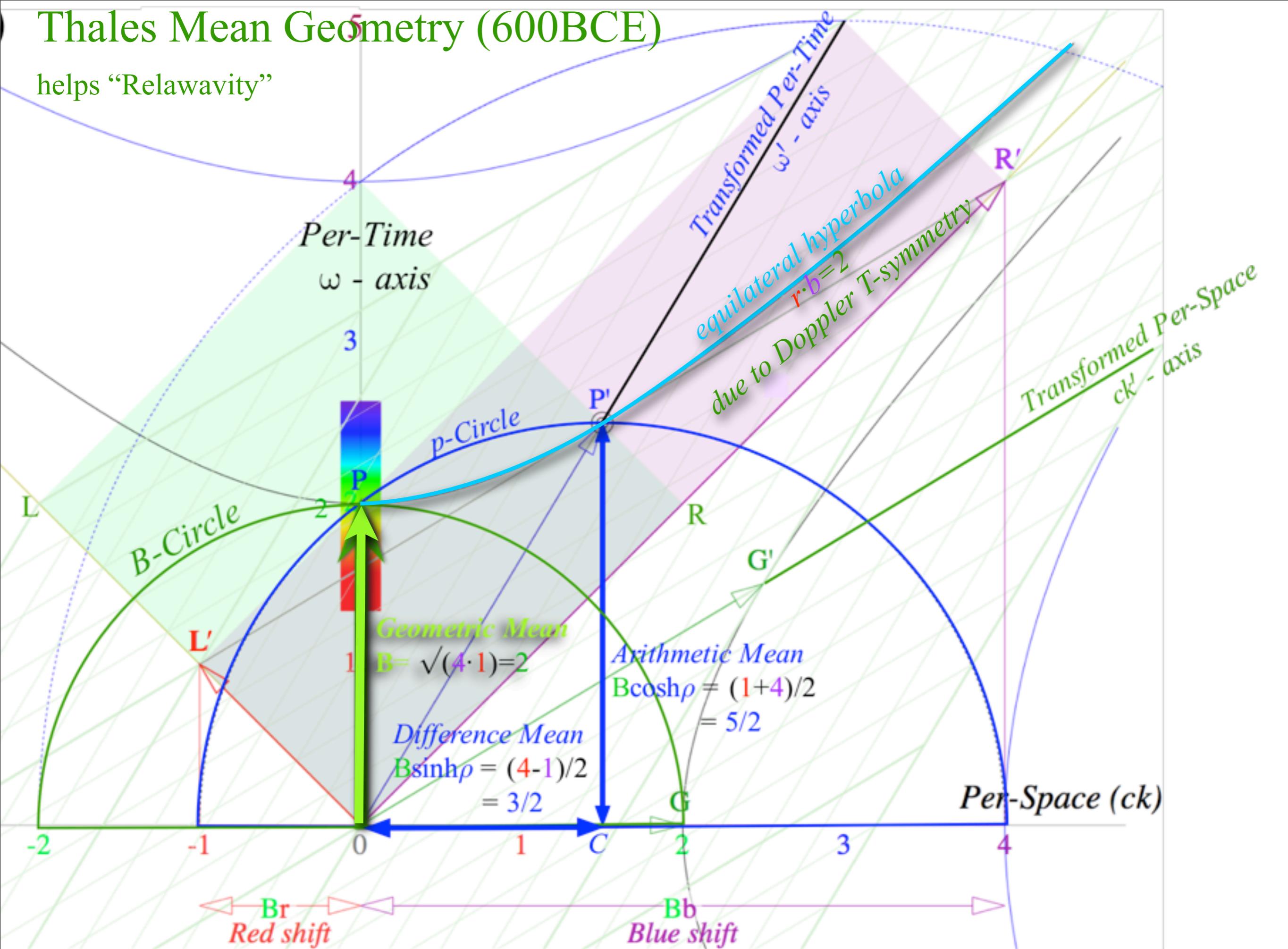
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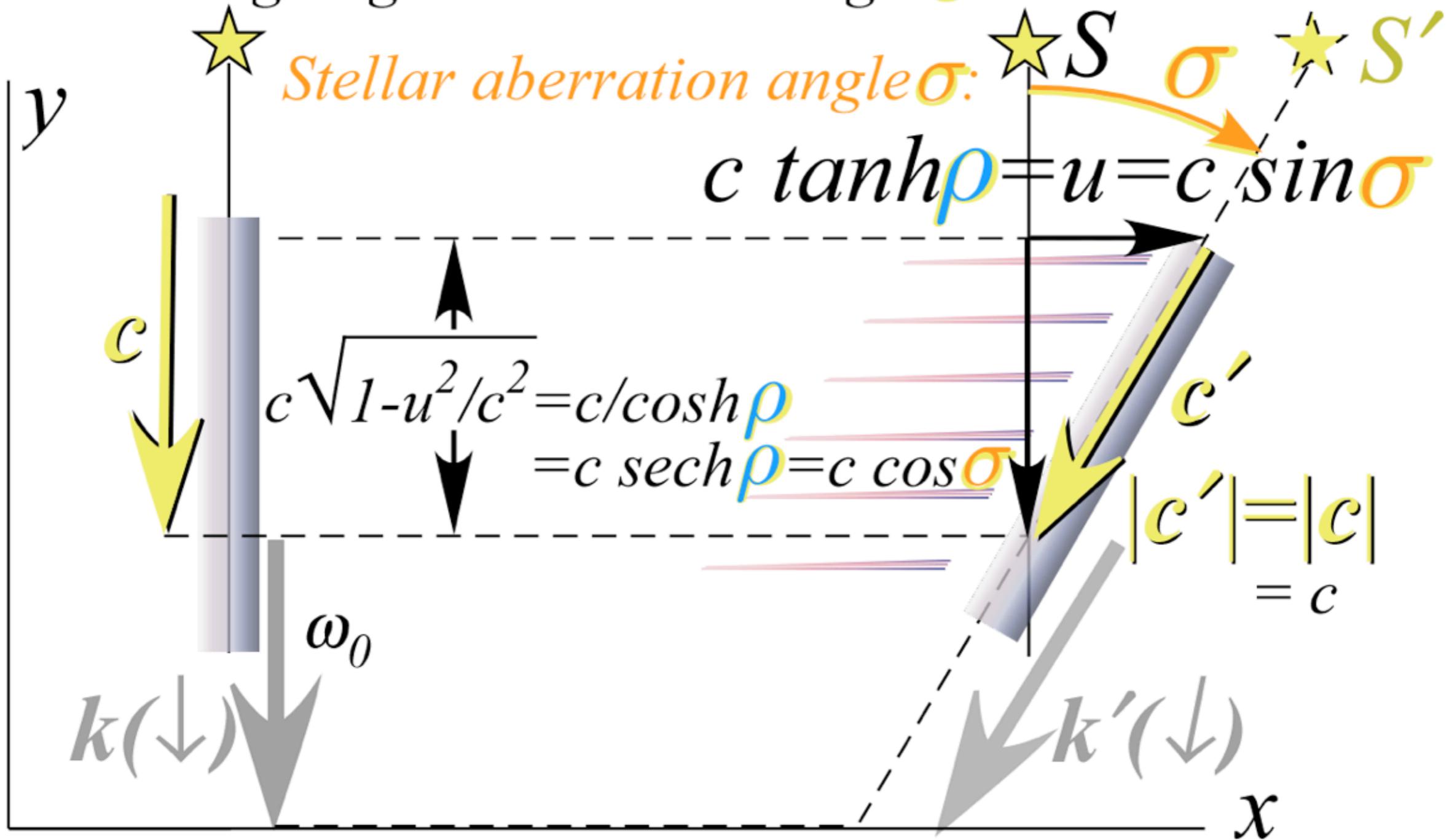


Comparing Longitudinal relativity parameter: Rapidity $\rho = \log_e(\text{Doppler Shift})$ to a Transverse* relativity parameter: Stellar aberration angle σ

*Lewis Carroll Epstein, *Relativitätstheorie*, Birkhäuser, (2004) Earlier English version (1985)-

Observer fixed below star sees it directly overhead.

Observer going u sees star at angle σ in u direction.



Comparing Longitudinal relativity parameter: Rapidity $\rho = \log_e(\text{Doppler Shift})$ to a Transverse* relativity parameter: Stellar aberration angle σ

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Proper time $c\tau$ vs. coordinate space x - (L. C. Epstein's "Cosmic Speedometer")

Particles P and P' have speed u in (x', ct') and speed c in $(x, c\tau)$

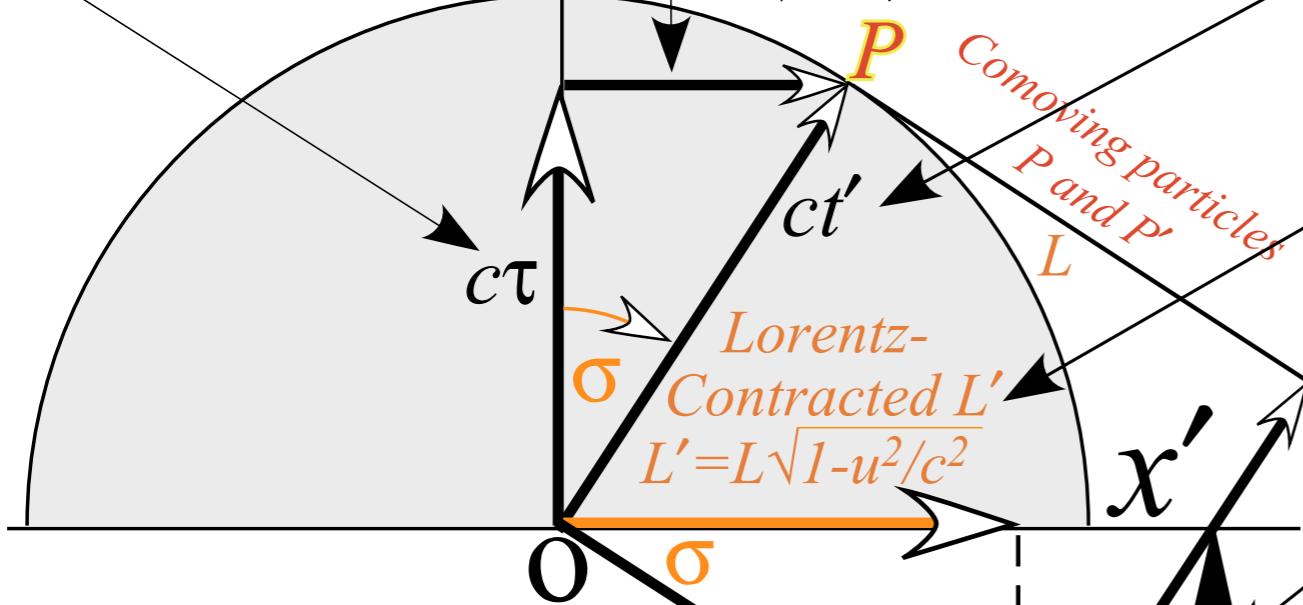
Proper time $C\tau$

$$c\tau = \sqrt{(ct')^2 - (x')^2}$$

Coordinate
 $x' = (u/c)ct' = ut'$

Einstein time dilation:

$$ct' = c\tau \sec \sigma = c\tau \cosh \rho = c\tau / \sqrt{1-u^2/c^2}$$



Lorentz length contraction:

$$L' = L \operatorname{sech} \rho = L \cos \sigma = L \cdot \sqrt{1-u^2/c^2}$$

Proper Time asimultaneity:

$$\begin{aligned} c \Delta \tau &= L' \sinh \rho = L \cos \sigma \sinh \rho \\ &= L \cos \sigma \tan \sigma \\ &= L \sin \sigma \quad = L / \sqrt{c^2/u^2 - 1} \sim L u/c \end{aligned}$$

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Epstein's trick is to
turn a hyperbolic form
into a circular form:

$$\sqrt{(c\tau)^2 + (x')^2} = (ct')$$

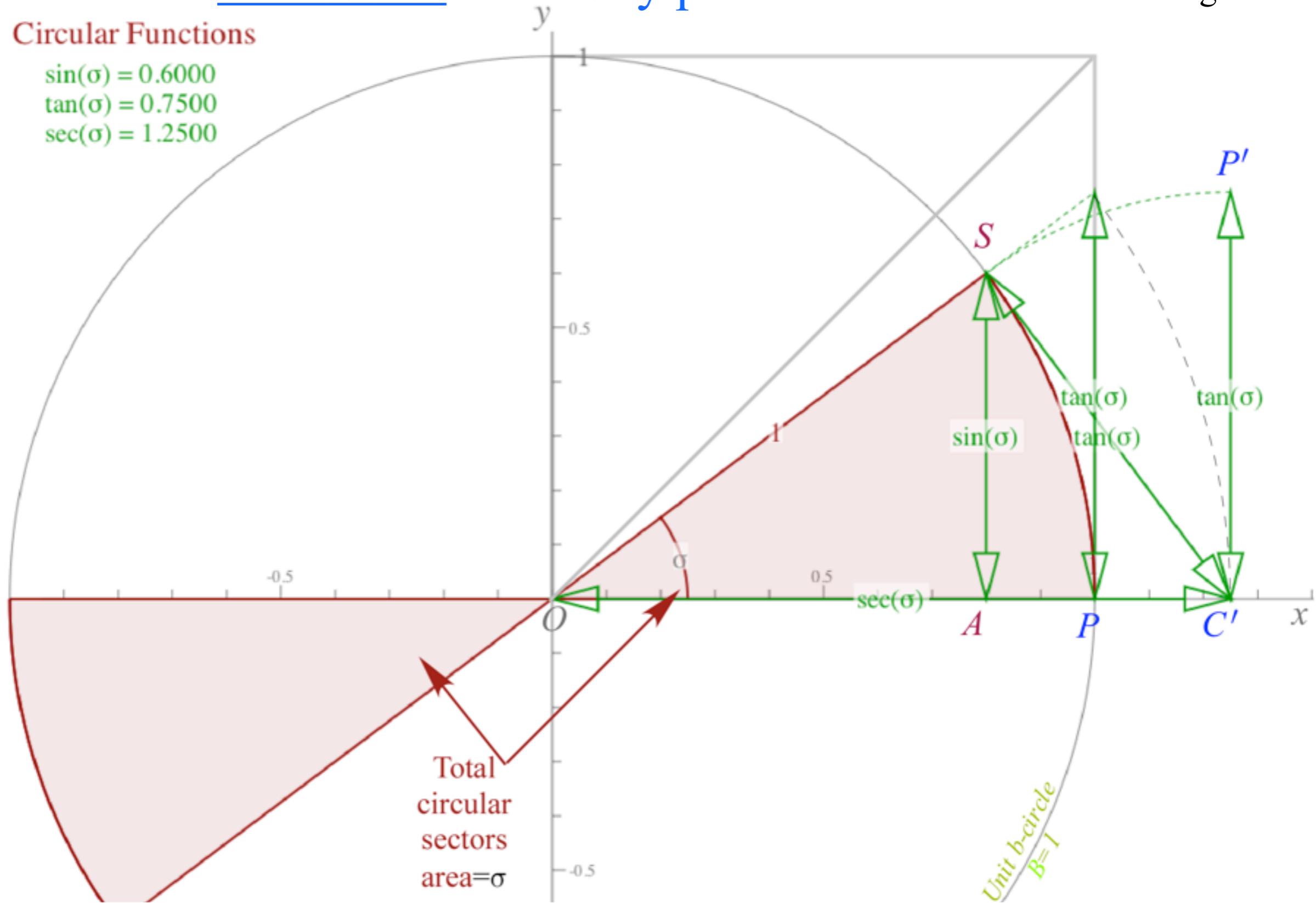
Comparing Longitudinal relativity parameter: Rapidity $\rho = \log_e(\text{Doppler Shift})$ to a Transverse relativity parameter: Stellar aberration angle σ

(a) Circular Functions

$$\sin(\sigma) = 0.6000$$

$$\tan(\sigma) = 0.7500$$

$$\sec(\sigma) = 1.2500$$



Comparing Longitudinal relativity parameter: Rapidity $\rho = \log_e(\text{Doppler Shift})$ to a Transverse relativity parameter: Stellar aberration angle σ

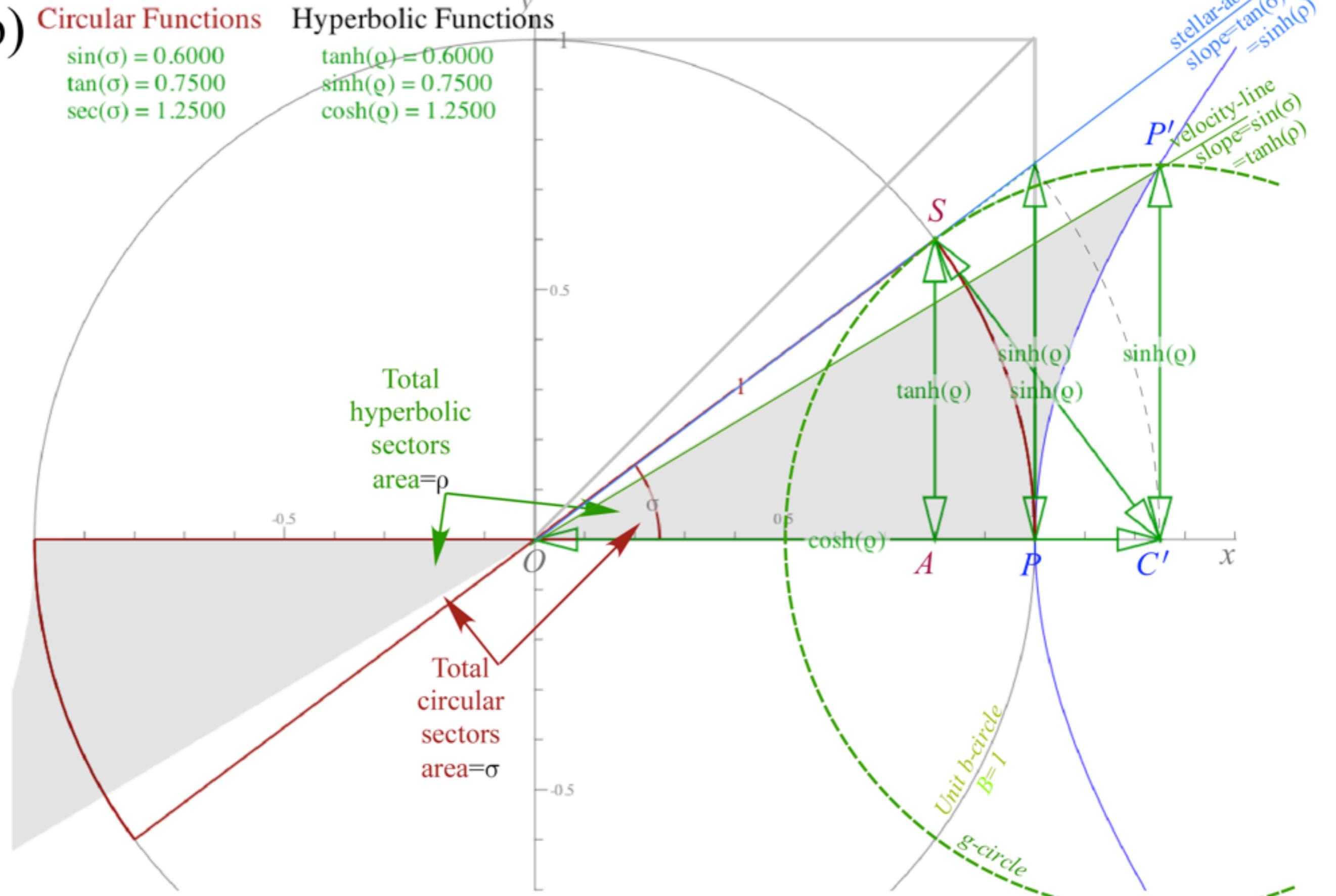
(b)

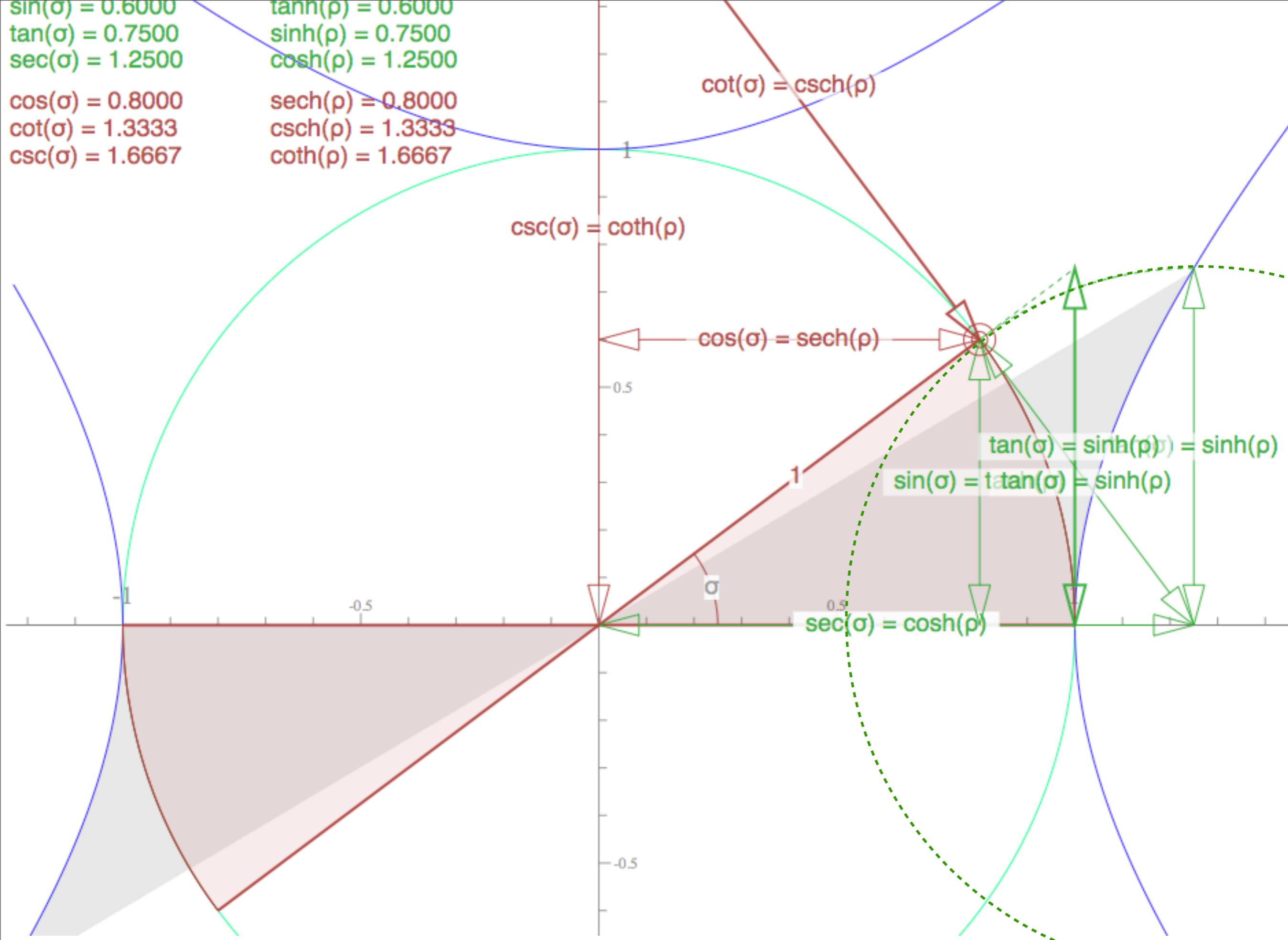
Circular Functions

$$\begin{aligned}\sin(\sigma) &= 0.6000 \\ \tan(\sigma) &= 0.7500 \\ \sec(\sigma) &= 1.2500\end{aligned}$$

Hyperbolic Functions

$$\begin{aligned}\tanh(\rho) &= 0.6000 \\ \sinh(\rho) &= 0.7500 \\ \cosh(\rho) &= 1.2500\end{aligned}$$





Circular Functions

$m_<(\sigma) = 0.6435$
 $\text{Length}(\sigma) = 0.6435$
 $\text{Area}(\sigma) = 0.6435$

 $\sin(\sigma) = 0.6000$
 $\tan(\sigma) = 0.7500$
 $\sec(\sigma) = 1.2500$

 $\cos(\sigma) = 0.8000$
 $\cot(\sigma) = 1.3333$
 $\csc(\sigma) = 1.6667$

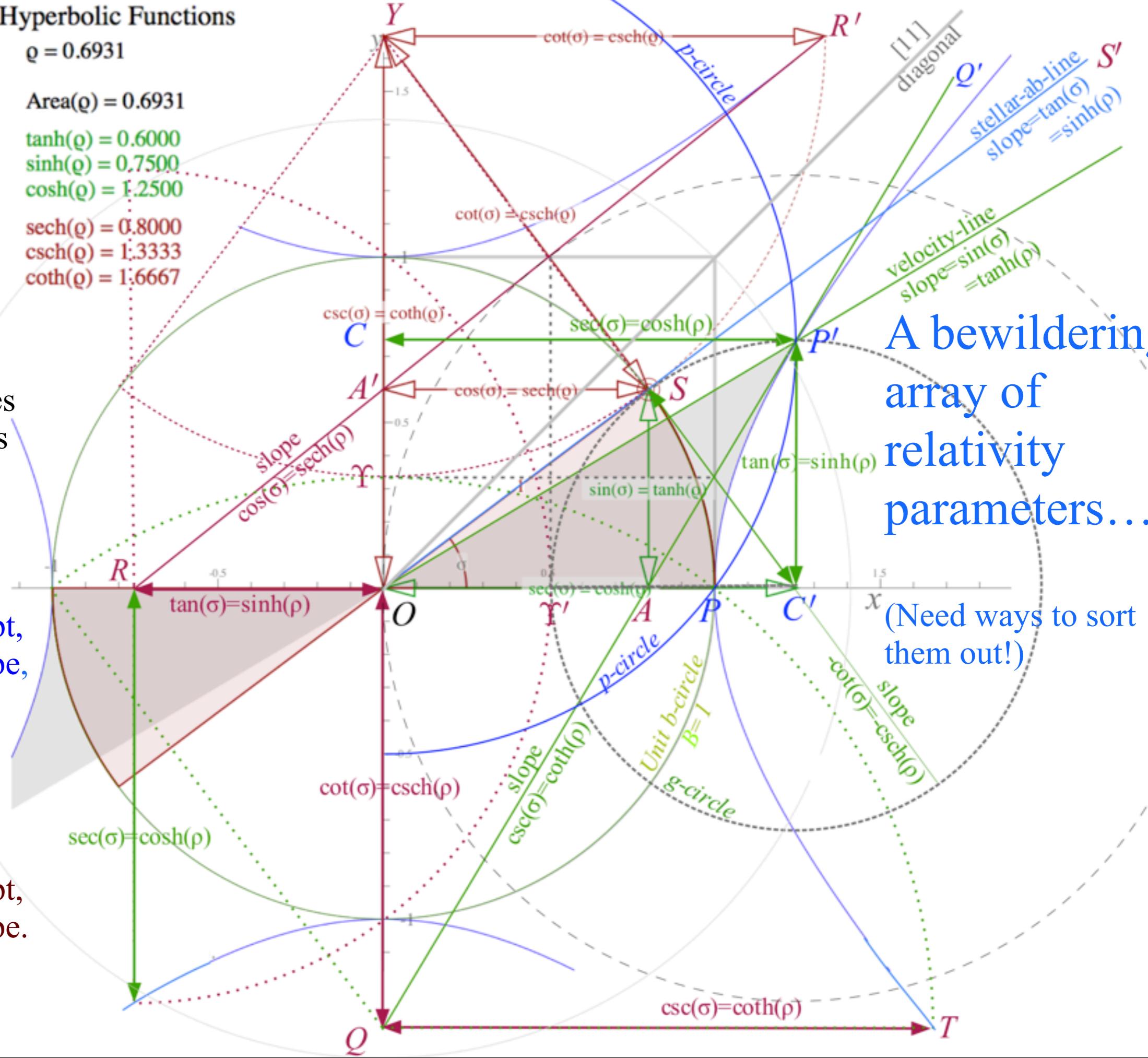
Each of 6 trig (or trigh) functions serves at least once as a hyperbolic x, y , and z coordinate, x, y , and z tangent intercept, and tangent slope, and a circular x, y , and z coordinate, x, y , and z tangent intercept, and tangent slope.

Hyperbolic Functions

$Q = 0.6931$

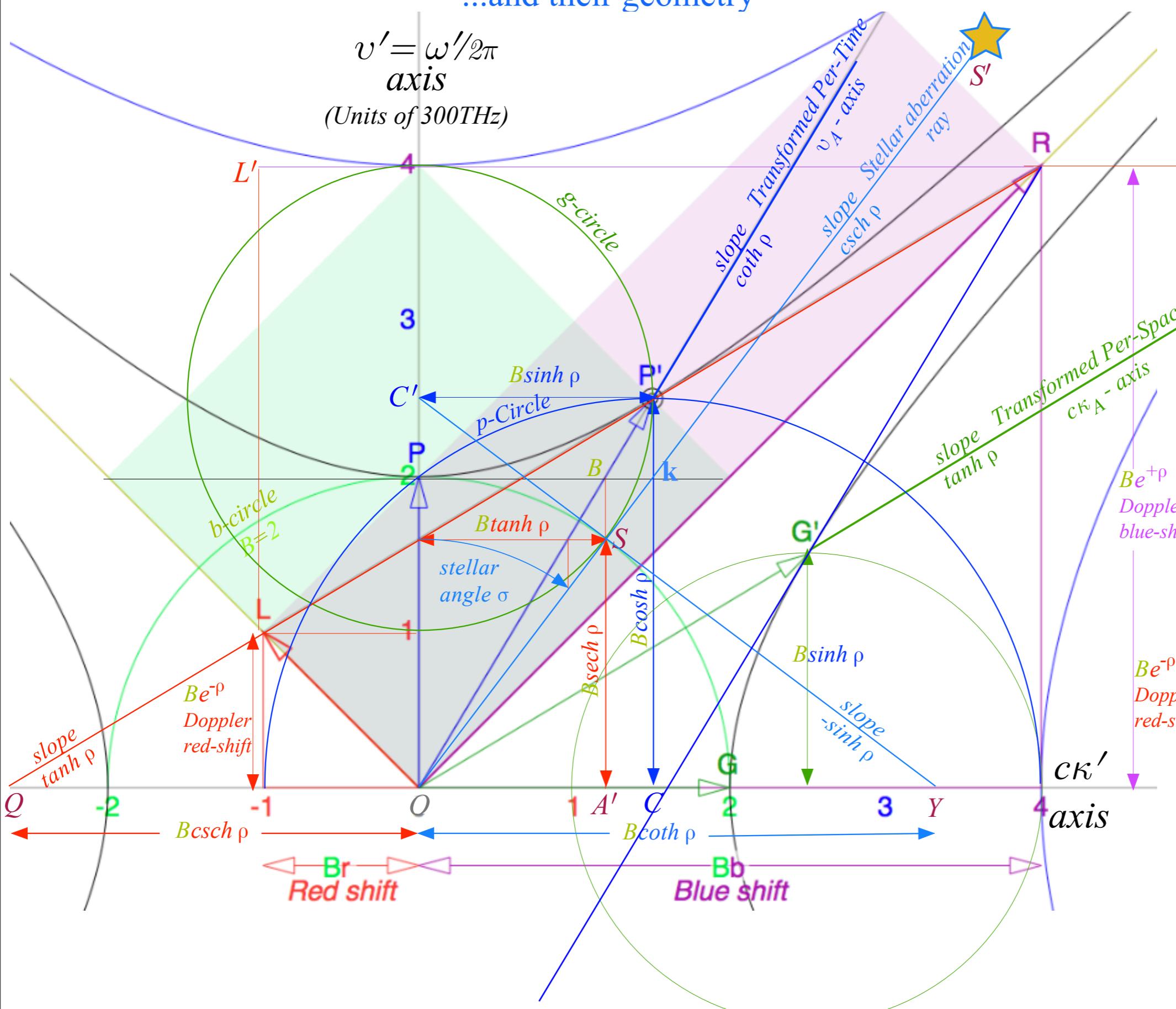
 $\text{Area}(Q) = 0.6931$
 $\tanh(Q) = 0.6000$
 $\sinh(Q) = 0.7500$
 $\cosh(Q) = 1.2500$

 $\text{sech}(Q) = 0.8000$
 $\text{csch}(Q) = 1.3333$
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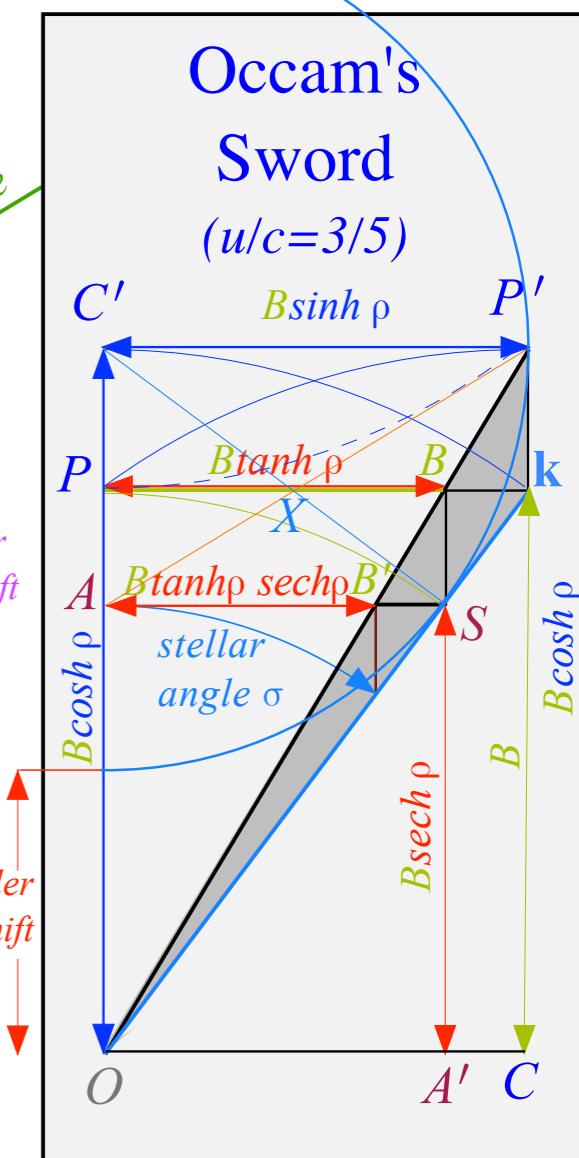


Summary of optical wave parameters for relativity and QM

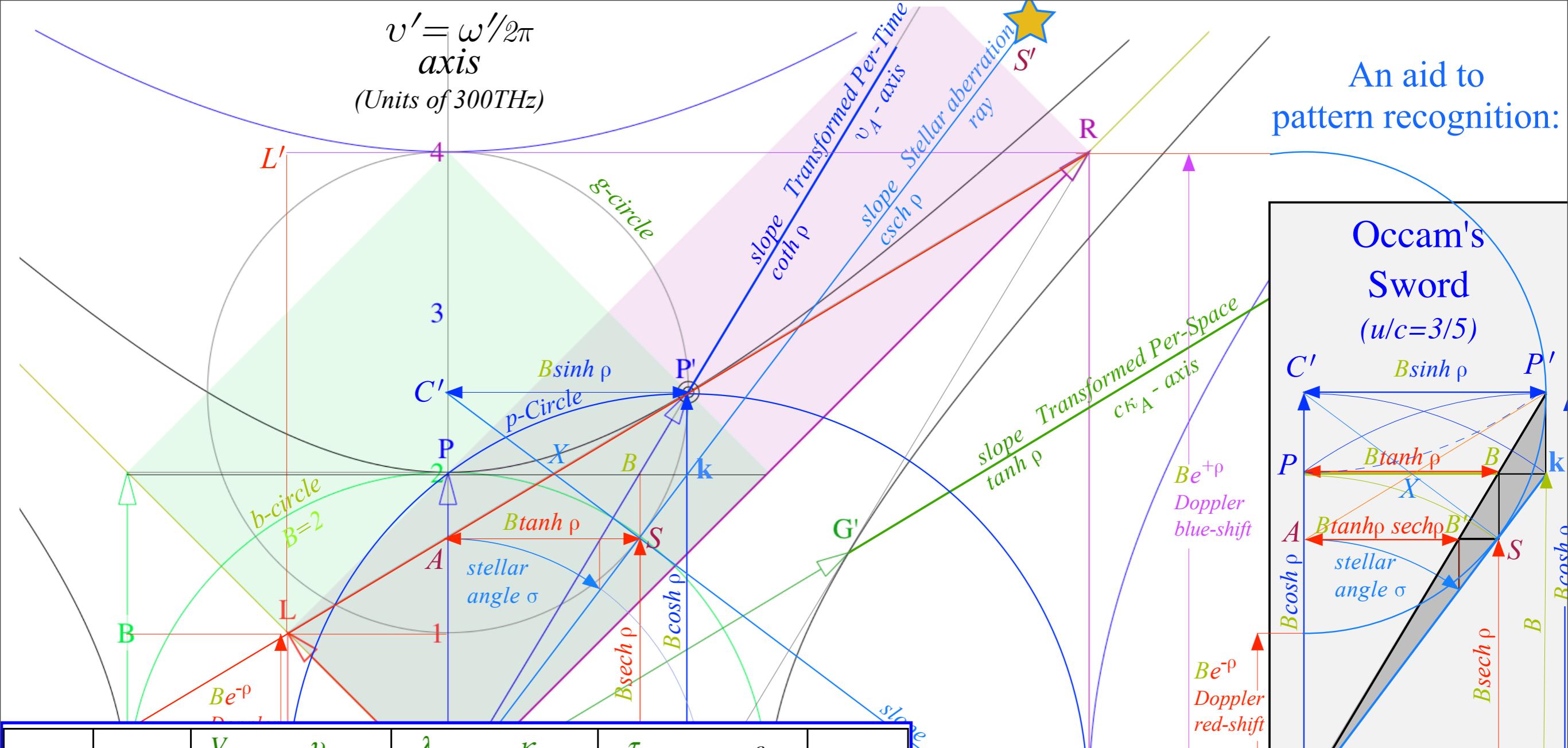
...and their geometry



An aid to pattern recognition:



An aid to
pattern recognition:



group	$b_{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{c}{V_{group}}$	$b_{Doppler}$
phase	$\frac{1}{b_{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{V_{phase}}{c}$	$b_{Doppler}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

Table of 12 wave parameters
(includes inverses) for relativity
...and values for $u/c=3/5$

Using (some) wave parameters for relativistic quantum theory

$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \text{ (for } u \ll c\text{)}$$

$$c\kappa_{phase} = B \sinh \rho \approx B \rho \quad \text{(for } u \ll c\text{)}$$

$$B = v_A$$

$$B = v_A = c\kappa_A$$

$$\cosh \rho \approx 1 + \frac{1}{2} \rho^2$$

$$\sinh \rho \approx \rho$$

At low speeds: ...

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	$\frac{v_{group}}{v_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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At low speeds:

$$\begin{aligned} \cosh \rho &\approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2} \\ \sinh \rho &\approx \rho \approx \frac{u}{c} \end{aligned}$$

$$B = v_A$$

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group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	v_{group}	λ_{group}	κ_{group}	τ_{group}	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
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At low speeds:

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2 \quad \Leftarrow \text{for } (u \ll c)$$

$$\begin{aligned} \cosh \rho &\approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2} \\ \sinh \rho &\approx \rho \approx \frac{u}{c} \end{aligned}$$

$$\begin{aligned} B &= v_A \\ B &= v_A = c\kappa_A \end{aligned}$$

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phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
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$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

Using (some) wave parameters for relativistic quantum theory

$$\begin{aligned} v_{phase} &= B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \text{ (for } u \ll c) \\ c\kappa_{phase} &= B \sinh \rho \approx B \rho \quad \text{(for } u \ll c) \\ \frac{u}{c} &= \tanh \rho \approx \rho \quad \text{(for } u \ll c) \end{aligned}$$

$$\begin{aligned} \cosh \rho &\approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} \frac{u^2}{c^2} \\ \sinh \rho &\approx \rho \approx \frac{u}{c} \end{aligned}$$

$$\begin{aligned} B &= v_A \\ B &= v_A = c\kappa_A \end{aligned}$$

At low speeds:

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2 \quad \Leftarrow \text{for } (u \ll c) \Rightarrow \quad \kappa_{phase} \approx \frac{B}{c^2} u$$

time	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	v_{group}	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$b_{RED}^{Doppler}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
stellar angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

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v_{phase} and κ_{phase} resemble
formulae for Newton's
kinetic energy and momentum

Resembles: $const. + \frac{1}{2} M u^2$

Resembles: $M u$

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	v_{group}	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

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$$B = v_A = c\kappa_A$$

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Rescale v_{phase} by \hbar so: $M = \frac{\hbar B}{c^2}$

Resembles: $const. + \frac{1}{2} Mu^2$

$$\kappa_{phase} \approx \frac{B}{c^2} u$$

Resembles: Mu

v_{phase} and κ_{phase} resemble formulae for Newton's kinetic energy $\frac{1}{2} Mu^2$ and momentum Mu .

time	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	v_{group}	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
space	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$b_{RED}^{Doppler}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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Rescale v_{phase} by \hbar so: $M = \frac{\hbar B}{c^2}$

Resembles: $const. + \frac{1}{2} Mu^2$

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At low speeds:

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v_{phase} and κ_{phase} resemble formulae for Newton's kinetic energy $\frac{1}{2} Mu^2$ and momentum Mu .

So attach scale factor \hbar (or $\hbar N$) to match units.

Resembles: Mu

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	v_{group}	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

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Rescale v_{phase} by \hbar so: $M = \frac{\hbar B}{c^2}$

$$\hbar v_{phase} \approx \hbar B + \frac{1}{2} \frac{\hbar B}{c^2} u^2 \Leftarrow \text{for } (u \ll c) \Rightarrow$$

Resembles: $const. + \frac{1}{2} Mu^2$

$$\kappa_{phase} \approx \frac{B}{c^2} u$$

$$\hbar \kappa_{phase} \approx \frac{\hbar B}{c^2} u$$

Resembles: Mu

v_{phase} and κ_{phase} resemble formulae for Newton's kinetic energy $\frac{1}{2} Mu^2$ and momentum Mu . So attach scale factor \hbar (or $\hbar N$) to match units.

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	v_{group}	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

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Resembles: $const. + \frac{1}{2} Mu^2$

$$\kappa_{phase} \approx \frac{B}{c^2} u$$

(The famous Mc^2 shows up!)

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v_{phase} and κ_{phase} resemble formulae for Newton's kinetic energy $\frac{1}{2} Mu^2$ and momentum Mu . So attach scale factor \hbar (or $\hbar N$) to match units.

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	v_{group}	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

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$$h\nu_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2 \Leftarrow \text{for } (u \ll c) \Rightarrow h\kappa_{phase} \approx \frac{hB}{c^2} u$$

$$h\nu_{phase} \approx Mc^2 + \frac{1}{2} Mu^2 \Leftarrow \text{for } (u \ll c) \Rightarrow h\kappa_{phase} \approx Mu$$

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group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	v_{group}	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

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Lucky coincidences??

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	v_{group}	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

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$$\kappa_{phase} \approx \frac{B}{c^2} u$$

$$\hbar v_{phase} \approx \hbar B + \frac{1}{2} \frac{\hbar B}{c^2} u^2 \Leftarrow \text{for } (u \ll c) \Rightarrow \hbar \kappa_{phase} \approx \frac{\hbar B}{c^2} u$$

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Lucky coincidences??
... Try exact v_{phase} ...

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	v_{group}	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2}=0.5$	$\frac{3}{5}=0.6$	$\frac{3}{4}=0.75$	$\frac{4}{5}=0.80$	$\frac{5}{4}=1.25$	$\frac{4}{3}=1.33$	$\frac{5}{3}=1.67$	$\frac{2}{1}=2.0$

Using (some) wave parameters for relativistic quantum theory

$$v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 \text{ (for } u \ll c)$$

$$c\kappa_{phase} = B \sinh \rho \approx B \rho \quad (\text{for } u \ll c)$$

$$\frac{u}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2 \Leftarrow \text{for } (u \ll c) \Rightarrow$$

Rescale v_{phase} by \hbar so: $M = \frac{\hbar B}{c^2}$ or: $\hbar B = Mc^2$

$$\hbar v_{phase} \approx \hbar B + \frac{1}{2} \frac{\hbar B}{c^2} u^2 \Leftarrow \text{for } (u \ll c) \Rightarrow \hbar \kappa_{phase} \approx \frac{\hbar B}{c^2} u$$

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phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	$\csc \sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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Rescale v_{phase} by h so: $M = \frac{hB}{c^2}$ or: $hB = Mc^2$

$$hv_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2 \Leftarrow \text{for } (u \ll c) \Rightarrow hv_{phase} \approx \frac{hB}{c^2} u$$

$$hv_{phase} \approx Mc^2 + \frac{1}{2} Mu^2 \Leftarrow \text{for } (u \ll c) \Rightarrow hv_{phase} \approx Mu$$

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Einstein (1905)

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phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
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$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^2-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^2-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
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$$\kappa_{phase} \approx \frac{B}{c^2} u$$

$$h\nu_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2 \Leftarrow \text{for } (u \ll c) \Rightarrow \quad h\kappa_{phase} \approx \frac{hB}{c^2} u$$

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$$cp = \frac{Mcu}{\sqrt{1 - u^2/c^2}}$$

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	v_{group}	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\operatorname{csch} \rho$	$\coth \rho$	$e^{+\rho}$
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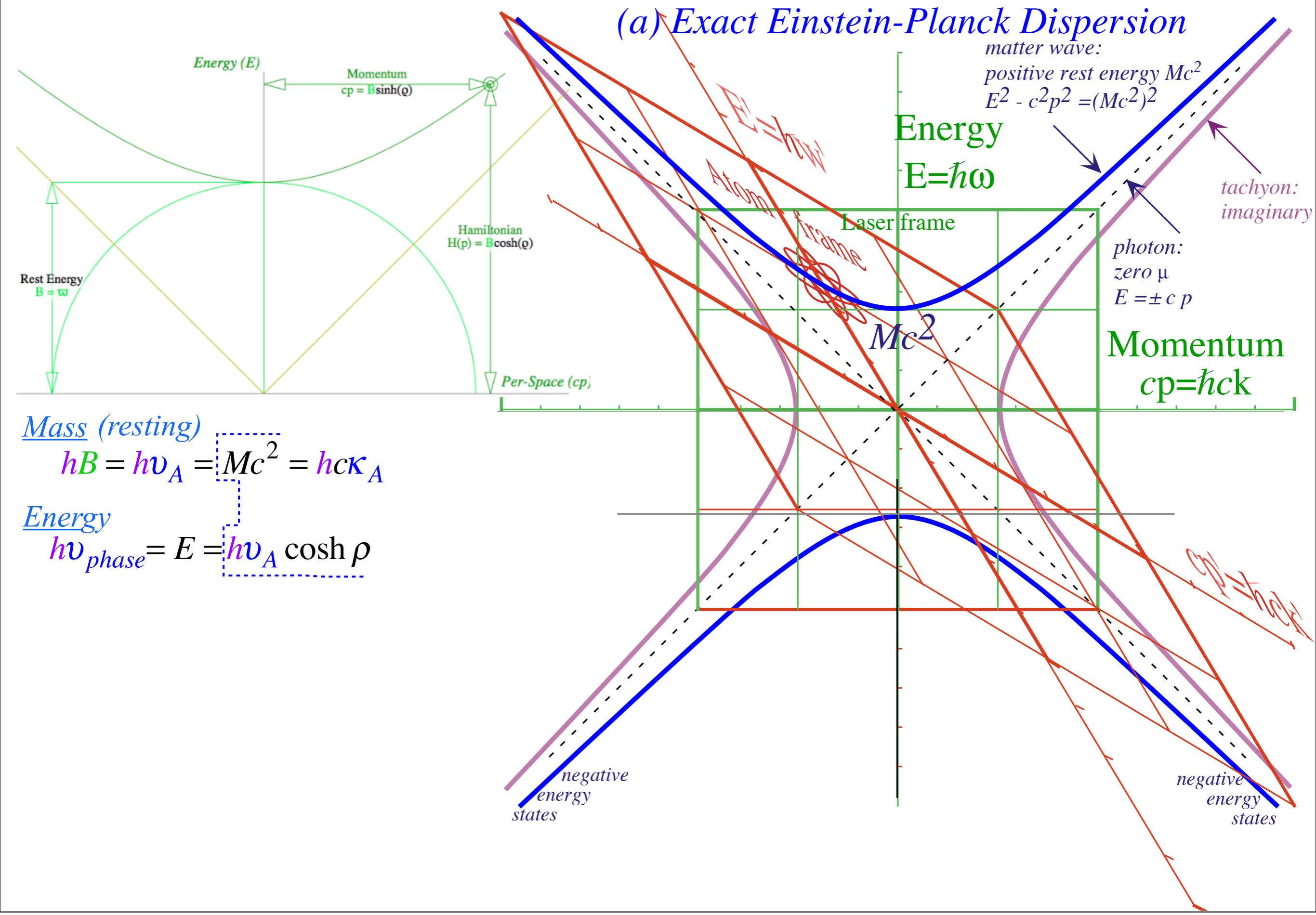
$$cp = \frac{Mu}{\sqrt{1-u^2/c^2}}$$

$$\text{Momentum: } h\kappa_{phase} = p = \frac{Mu}{\sqrt{1-u^2/c^2}}$$

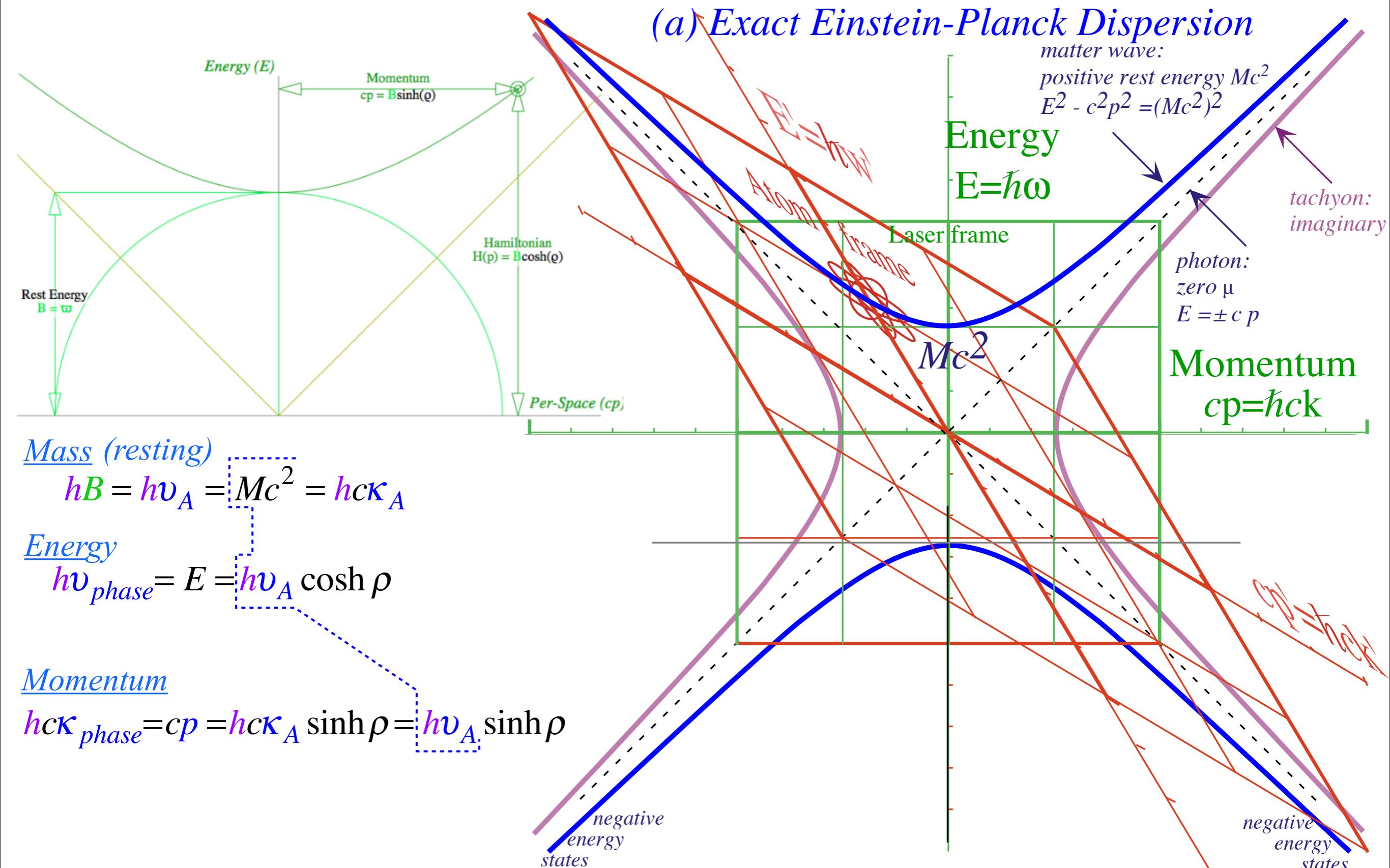
DeBroglie (1921)

group	$b_{RED}^{Doppler}$	$\frac{V_{group}}{c}$	v_{group}	$\frac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{group}}{\kappa_A}$	$\frac{\tau_{group}}{\tau_A}$	$\frac{V_{phase}}{c}$	$b_{BLUE}^{Doppler}$
phase	$\frac{1}{c}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$\frac{\tau_{phase}}{\tau_A}$	$\frac{v_{phase}}{v_A}$	$\frac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{RED}^{Doppler}}$
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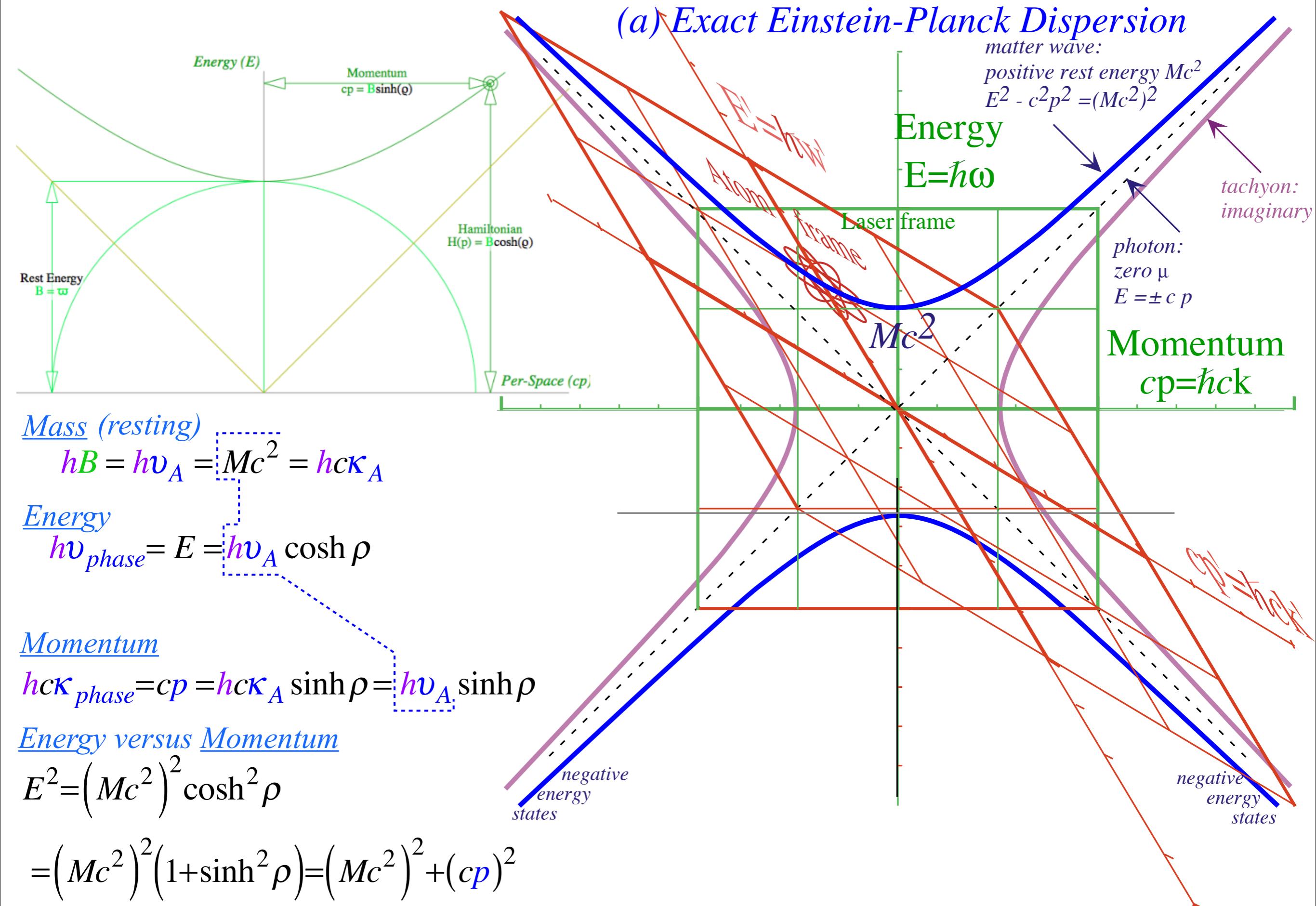
Using (some) wave coordinates for relativistic quantum theory



Using (some) wave coordinates for relativistic quantum theory

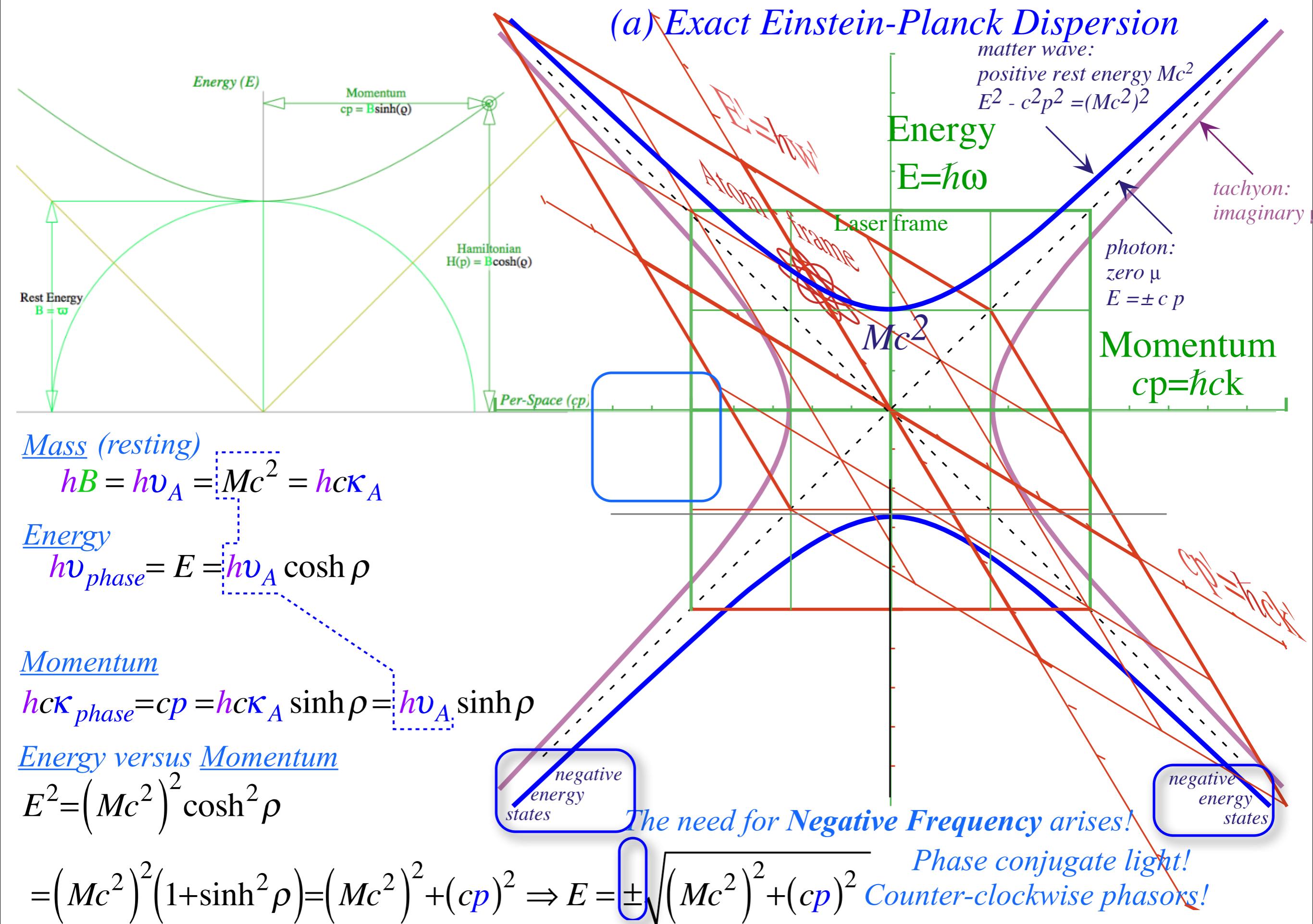


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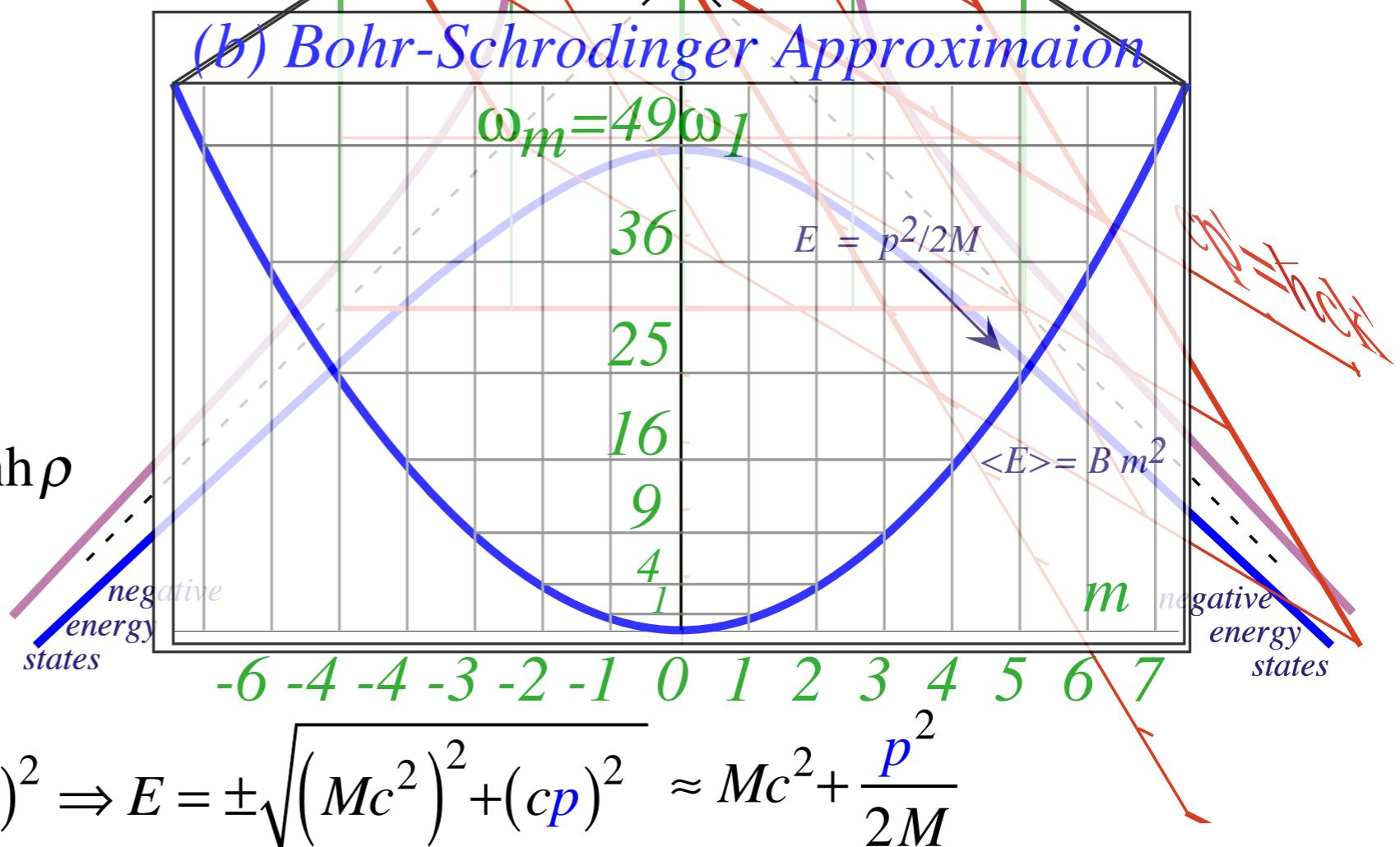
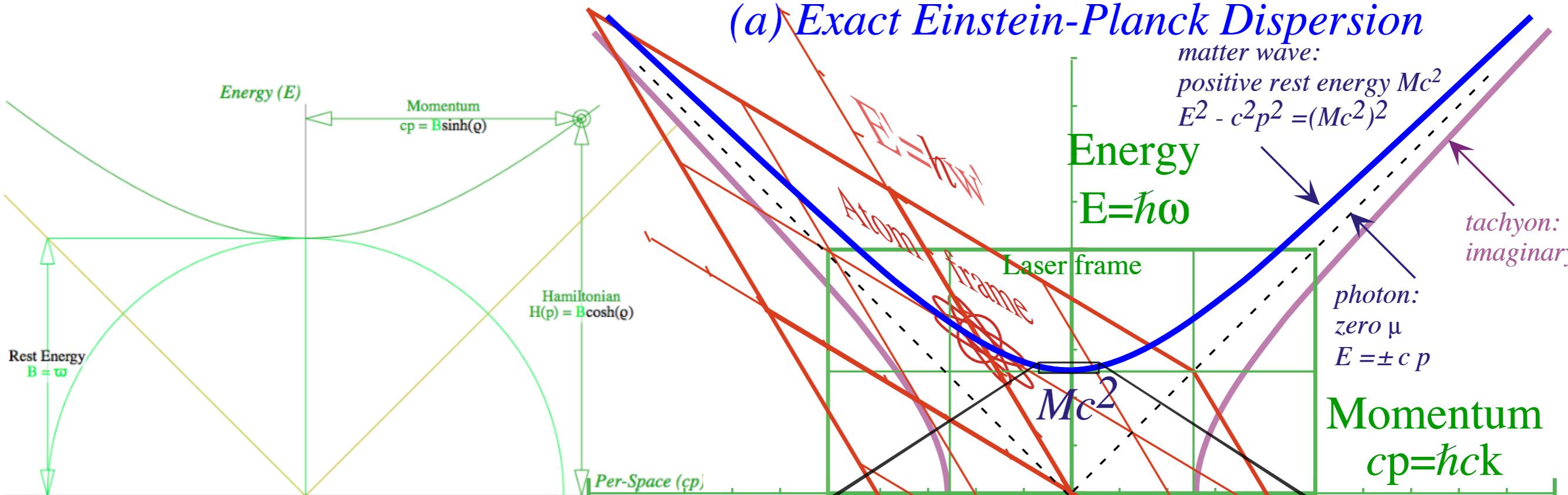


Using (some) wave coordinates for relativistic quantum theory

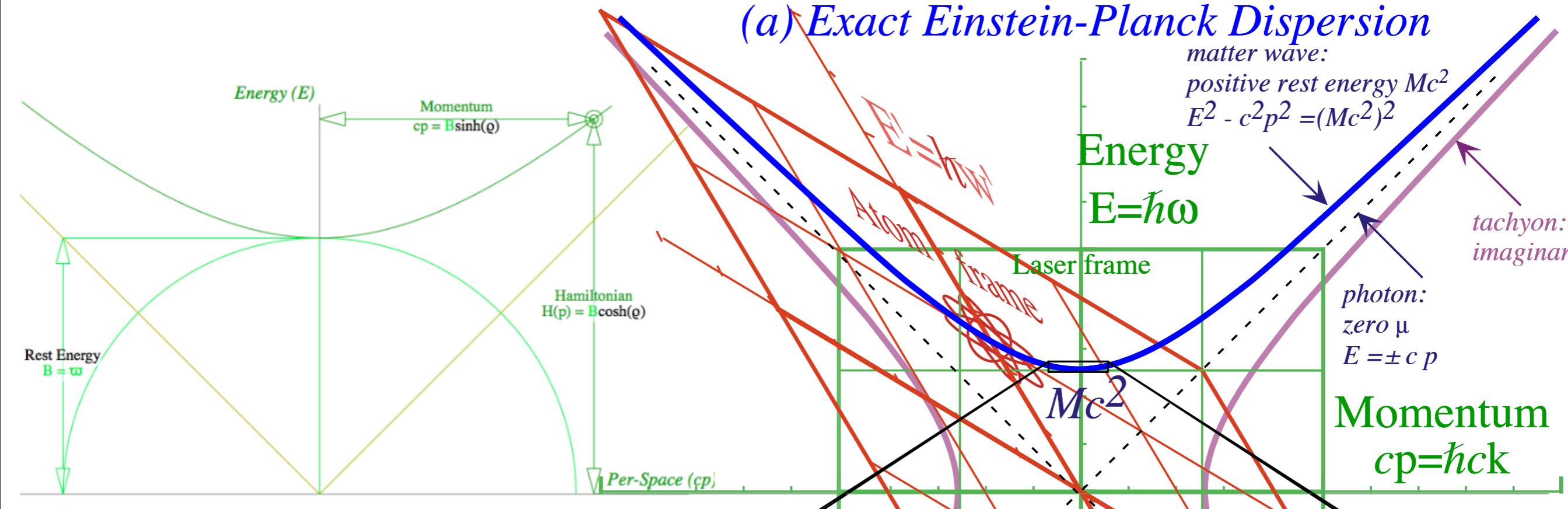
(a) Exact Einstein-Planck Dispersion



Using (some) wave coordinates for relativistic quantum theory



Using (some) wave coordinates for relativistic quantum theory



Mass (*resting*)

$$hB = h\nu_A = Mc^2 = hc\kappa_A$$

Energy

$$\hbar\nu_{phase} = E = \hbar\nu_A \cosh \rho$$

Momentum

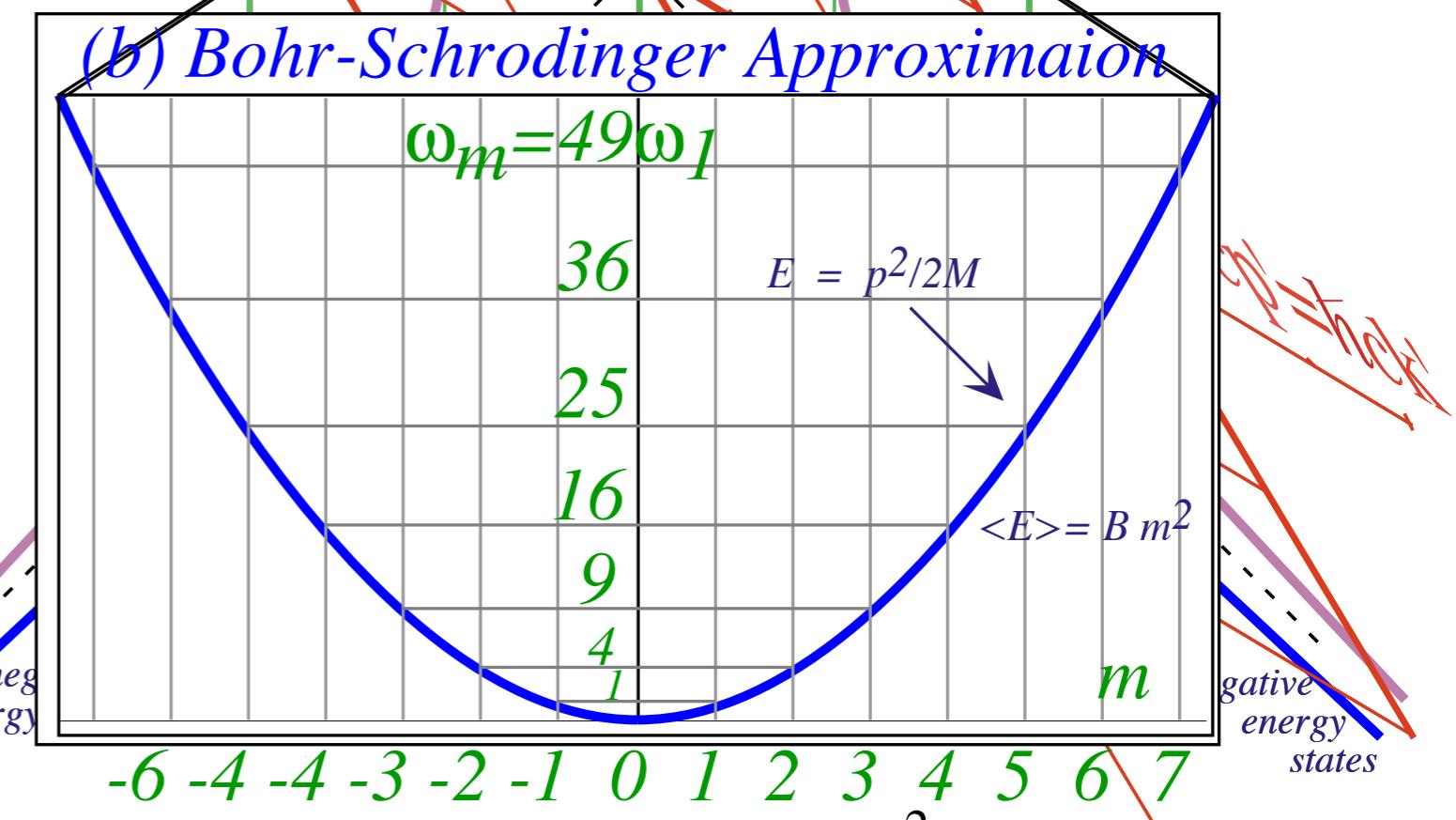
$$hck_{phase} = cp = hck_A \sinh \rho = hv_A \sinh \rho$$

Energy versus Momentum

$$E^2 = \left(Mc^2 \right)^2 \cosh^2 \rho$$

$$= \left(Mc^2 \right)^2 (1 + \sinh^2 \rho) = \left(Mc^2 \right)^2 + (cp)^2 \Rightarrow E = \pm \sqrt{\left(Mc^2 \right)^2 + (cp)^2} \approx Mc^2 + \frac{p^2}{2M}$$

low speed approximation



Definition(s) of mass for relativity/quantum

Given: Energy: $E = Mc^2 \cosh \rho$
= $\hbar v_{phase}$

Rest Mass M_{rest} (*Einstein's mass*)

$$\hbar B = \hbar v_A = Mc^2 = \hbar c \kappa_A$$

Defines invariant hyperbola(s)

$$E = \pm \sqrt{(Mc^2)^2 + (cp)^2}$$

momentum: $cp = Mc^2 \sinh \rho$

$$= \hbar c \kappa_{phase}$$

velocity: $u = c \tanh \rho = \frac{dv}{d\kappa}$

Definition(s) of mass for relativity/quantum

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$$= h\nu_{phase}$$

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$$hB = h\nu_A = Mc^2 = hc\kappa_A$$

$$\frac{h\nu_{phase}}{c^2} = M_{rest} = \frac{hc\kappa_{phase}}{c^2}$$

Rest
Mass

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Momentum Mass M_{mom} (*Galileo's mass*) Defined by ratio p/u of relativistic momentum to group velocity.

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$$= M_{rest} \cosh \rho = \frac{M_{rest}}{\sqrt{1 - u^2 / c^2}}$$

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 $= h\nu_{phase}$

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$$hB = h\nu_A = Mc^2 = hc\kappa_A$$

$$\frac{h\nu_{phase}}{c^2} = M_{rest} = \frac{hc\kappa_{phase}}{c^2}$$

Defines invariant hyperbola(s)

$$E = \pm \sqrt{(Mc^2)^2 + (cp)^2}$$

momentum: $cp = Mc^2 \sinh \rho$

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Limiting cases:

$$M_{mom} \xrightarrow{u \rightarrow c} M_{rest} e^\rho / 2$$

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Definition(s) of mass for relativity/quantum

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More common derivation using group velocity: $u \equiv V_{group} = \frac{d\omega}{dk} = \frac{dv}{d\kappa}$

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$$\frac{M_{rest}}{(1 - u^2 / c^2)^{3/2}} = M_{rest} \cosh^3 \rho$$

Effective Mass

general wave formula

to accompany $V_{group} = \frac{d\omega}{dk}$

Definition(s) of mass for relativity/quantum

How much mass does a γ -photon have?

(a) γ -rest mass: $M_{rest}^\gamma = 0$,

(b) γ -momentum mass: $M_{mom}^\gamma = \frac{p}{c} = \frac{h\kappa}{c} = \frac{h\nu}{c^2}$,

(c) γ -effective mass: $M_{eff}^\gamma = \infty$.

Newton complained about his “corpuscles” of light having “fits” (going crazy).

$$M_{mom}^\gamma = \frac{h\nu}{c^2} = \nu(1.2 \cdot 10^{-51}) \text{ kg} \cdot \text{s} = 4.5 \cdot 10^{-36} \text{ kg} \quad (\text{for: } \nu=600\text{THz})$$

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Relativistic action S and Lagrangian-Hamiltonian relations

Define *Lagrangian* L using invariant wave phase $\Phi = \textcolor{brown}{k}x - \omega t = \textcolor{brown}{k}'x' - \omega't'$ for wave of $k = k_{\text{phase}}$ and $\omega = \omega_{\text{phase}}$.

$$L \equiv \hbar \frac{d\Phi}{dt} = \hbar \textcolor{brown}{k} \frac{dx}{dt} - \hbar \omega$$

$$\begin{aligned}\hbar v_A &= Mc^2 = \hbar c \kappa_A \\ \hbar v_{\text{phase}} &= E = \hbar v_A \cosh \rho \\ \hbar c \kappa_{\text{phase}} &= cp = \hbar v_A \sinh \rho\end{aligned}$$

Prior wave relations
← linear Hz angular phasor →
format format

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Note: $Mcu = Mc^2 \tanh \rho$

Compare *Lagrangian* L

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$$\hbar v_A = Mc^2 = \hbar c k_A$$

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Also: $cp = Mc^2 \sinh \rho$

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$$L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho$$

with *Hamiltonian* $H = E$

$$\begin{aligned} H &= \hbar \omega = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = Mc^2 \cosh \rho \\ &= Mc^2 \sqrt{1 + \sinh^2 \rho} = Mc^2 \sqrt{1 + (cp)^2} \end{aligned}$$

$$\hbar v_A = Mc^2 = \hbar c K_A$$

$$\hbar v_{phase} = E = \hbar v_A \cosh \rho$$

$$\hbar c K_{phase} = cp = \hbar v_A \sinh \rho$$

Prior wave relations

← linear Hz
format

angular phasor
format

$$\hbar \omega_A = Mc^2 = \hbar c k_A$$

$$\hbar \omega_{phase} = E = \hbar \omega_A \cosh \rho$$

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Relativistic action S and Lagrangian-Hamiltonian relations

Define *Lagrangian* L using invariant wave phase $\Phi = \mathbf{k}x - \omega t = \mathbf{k}'x' - \omega't'$ for wave of $\mathbf{k} = k_{phase}$ and $\omega = \omega_{phase}$.

Use DeBroglie-momentum $p = \hbar k$ relation and Planck-energy $E = \hbar\omega$ relation to define *Hamiltonian* $H = E$

$$L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E \equiv p \dot{x} - E \equiv pu - H = L \quad \text{Legendre transformation}$$

Use Group velocity: $u = \frac{dx}{dt} = c \tanh \rho$

$$p = \hbar k = Mc \sinh \rho$$

$$E = \hbar \omega = Mc^2 \cosh \rho = H$$

$$= c \sin \sigma$$

$$\begin{aligned} L &= pu - H = (Mc \sinh \rho)(c \tanh \rho) - Mc^2 \cosh \rho \\ &= Mc^2 \frac{\sinh^2 \rho - \cosh^2 \rho}{\cosh \rho} = -Mc^2 \operatorname{sech} \rho \end{aligned}$$

$$\text{Note: } Mcu = Mc^2 \tanh \rho$$

$$= Mc^2 \sin \sigma$$

$$\text{Also: } cp = Mc^2 \sinh \rho$$

$$= \hbar ck = Mc^2 \tan \sigma$$

Compare *Lagrangian* L

$$L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho = -Mc^2 \cos \sigma$$

with *Hamiltonian* $H = E$

$$\begin{aligned} H &= \hbar \omega = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = Mc^2 \cosh \rho = Mc^2 \sec \sigma \\ &= Mc^2 \sqrt{1 + \sinh^2 \rho} = Mc^2 \sqrt{1 + (cp)^2} \end{aligned}$$

Including stellar angle σ

$$hv_A = Mc^2 = \hbar ck_A$$

$$hv_{phase} = E = hv_A \cosh \rho$$

$$\hbar ck_{phase} = cp = hv_A \sinh \rho$$

Prior wave relations

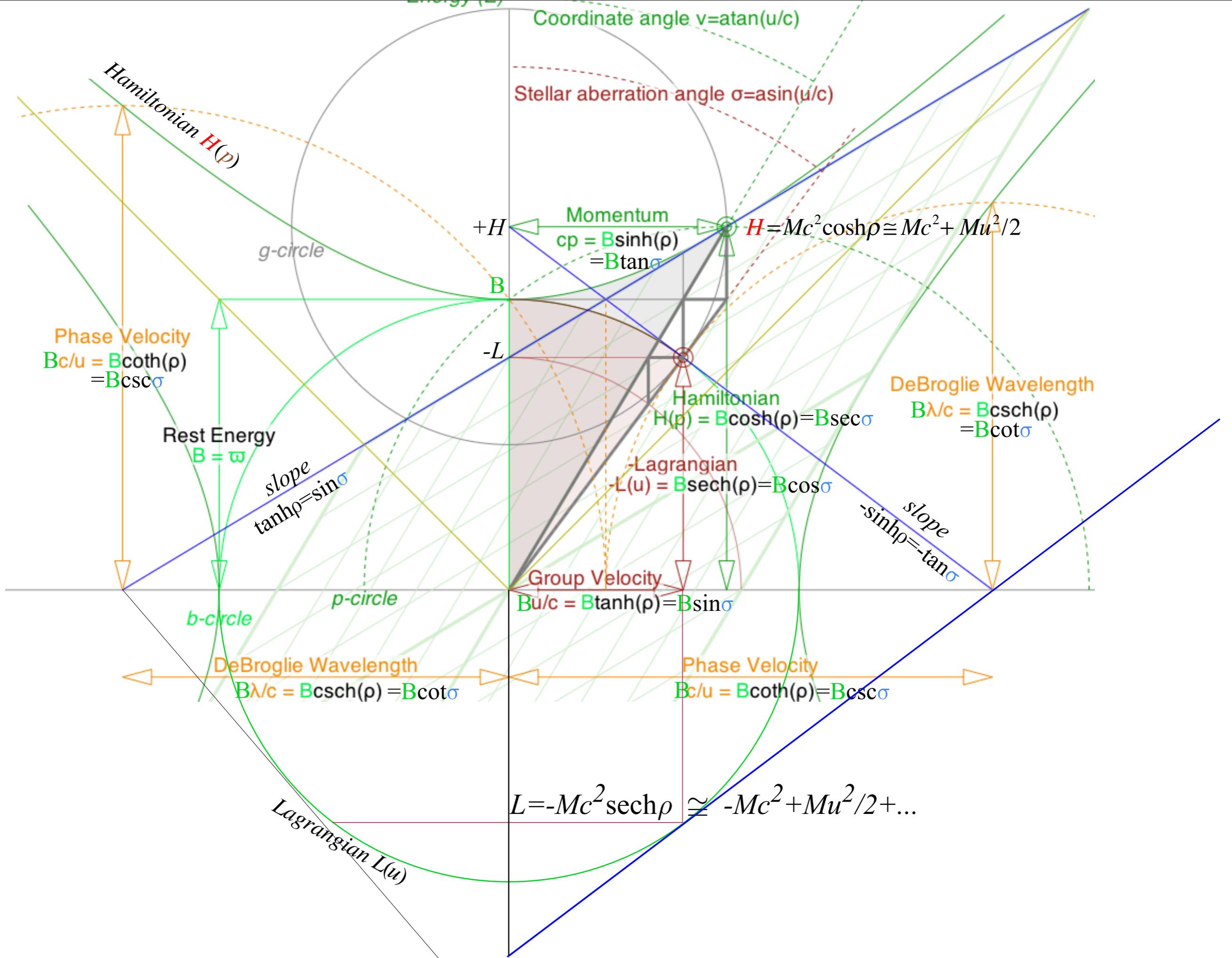
← linear Hz
format

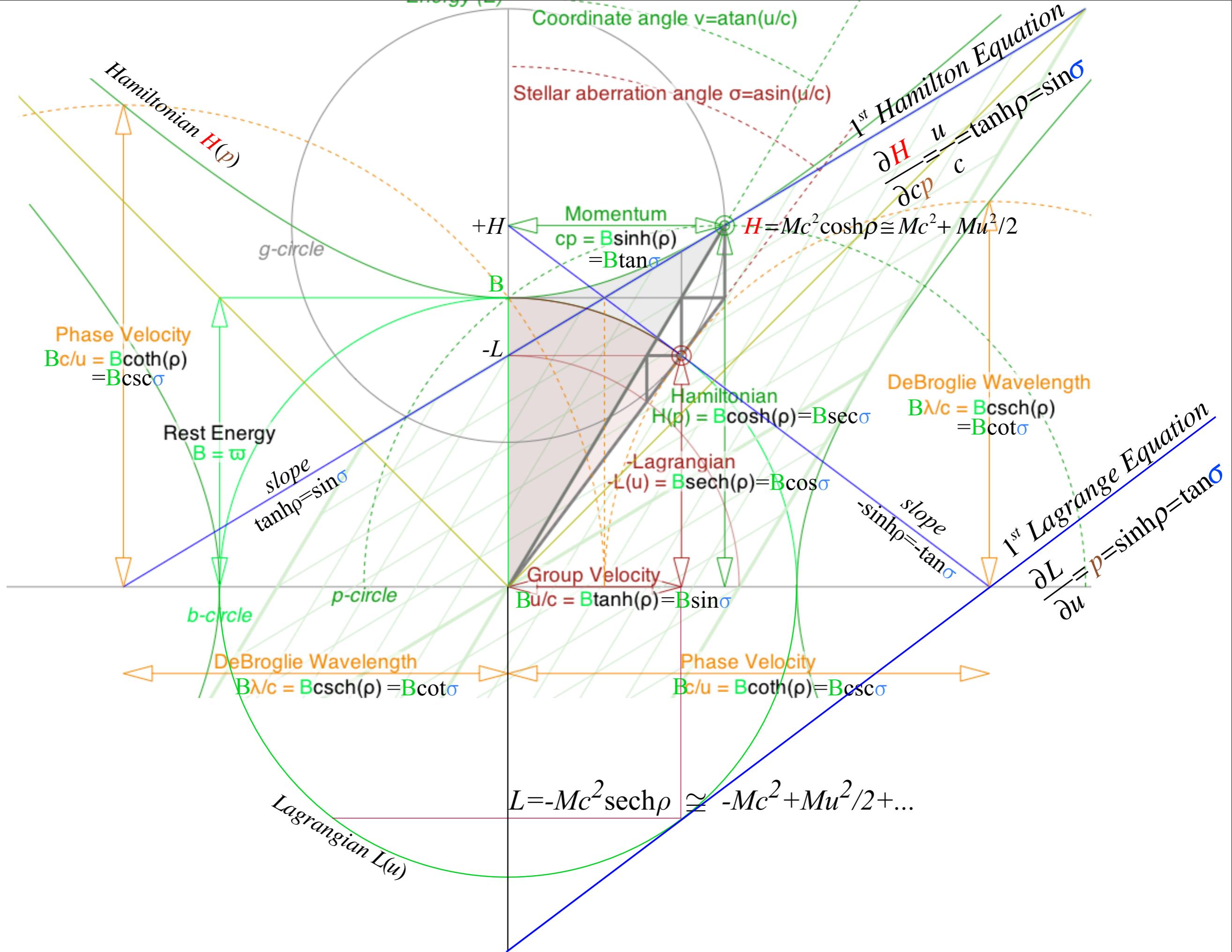
angular phasor
format

$$\hbar \omega_A = Mc^2 = \hbar ck_A$$

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Relativistic action S and Lagrangian-Hamiltonian relations

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Legendre transformation

Compare *Lagrangian* L

$$\dot{S} = L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho = -Mc^2 \cos \sigma$$

with *Hamiltonian* $H = E$

$$H = \hbar \omega = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = Mc^2 \cosh \rho = Mc^2 \sec \sigma \\ = Mc^2 \sqrt{1 + \sinh^2 \rho} = Mc^2 \sqrt{1 + (cp)^2}$$

Define *Action* $S = \hbar \Phi$

$$\hbar v_A = Mc^2 = \hbar c K_A$$

$$\hbar v_{phase} = E = \hbar v_A \cosh \rho$$

$$\hbar c K_{phase} = cp = \hbar v_A \sinh \rho$$

Prior wave relations

← linear Hz
format

angular phasor
format

$$\hbar \omega_A = Mc^2 = \hbar c K_A$$

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Relativistic action S and Lagrangian-Hamiltonian relations

Define *Lagrangian* L using invariant wave phase $\Phi = \mathbf{k}x - \omega t = \mathbf{k}'x' - \omega't'$ for wave of $\mathbf{k} = k_{phase}$ and $\omega = \omega_{phase}$.

Use DeBroglie-momentum $p = \hbar k$ relation and Planck-energy $E = \hbar\omega$ relation to define *Hamiltonian* $H = E$

$$\frac{dS}{dt} = L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E \equiv p \dot{x} - E \equiv pu - H = L$$

Legendre transformation

$$dS \equiv L dt \equiv \hbar d\Phi = \hbar k dx - \hbar \omega dt = p dx - H dt$$

Poincare Invariant action differential

Compare *Lagrangian* L

$$\dot{S} = L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho = -Mc^2 \cos \sigma$$

with *Hamiltonian* $H = E$

$$H = \hbar \omega = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = Mc^2 \cosh \rho = Mc^2 \sec \sigma \\ = Mc^2 \sqrt{1 + \sinh^2 \rho} = Mc^2 \sqrt{1 + (cp)^2}$$

Define *Action* $S = \hbar \Phi$

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$$\hbar c K_{phase} = cp = \hbar v_A \sinh \rho$$

Prior wave relations

← linear Hz angular phasor →
format format

$$\hbar \omega_A = Mc^2 = \hbar c k_A$$

$$\hbar \omega_{phase} = E = \hbar \omega_A \cosh \rho$$

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Relativistic action S and Lagrangian-Hamiltonian relations

Define *Lagrangian* L using invariant wave phase $\Phi = \mathbf{k}x - \omega t = \mathbf{k}'x' - \omega't'$ for wave of $\mathbf{k} = k_{phase}$ and $\omega = \omega_{phase}$.

Use DeBroglie-momentum $p = \hbar k$ relation and Planck-energy $E = \hbar\omega$ relation to define *Hamiltonian* $H = E$

$$\frac{dS}{dt} = L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E \equiv p \dot{x} - E \equiv pu - H = L \quad \text{Legendre transformation}$$

Use Group velocity: $u = \frac{dx}{dt} = c \tanh \rho$

$$dS \equiv L dt \equiv \hbar d\Phi = \hbar k dx - \hbar \omega dt = [p dx] - H dt$$

Poincare Invariant action differential

$$\frac{\partial S}{\partial x} = p$$

$$\frac{\partial S}{\partial t} = -H$$

Hamilton-Jacobi equations

Compare Lagrangian L

$$\dot{S} = L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech} \rho = -Mc^2 \cos \sigma$$

with Hamiltonian $H = E$

$$H = \hbar \omega = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = Mc^2 \cosh \rho = Mc^2 \sec \sigma \\ = Mc^2 \sqrt{1 + \sinh^2 \rho} = Mc^2 \sqrt{1 + (cp)^2}$$

Define Action $S = \hbar \Phi$

$$\hbar v_A = Mc^2 = \hbar c K_A$$

$$\hbar v_{phase} = E = \hbar v_A \cosh \rho$$

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Prior wave relations

← linear Hz angular phasor →
format format

$$\hbar \omega_A = Mc^2 = \hbar c k_A$$

$$\hbar \omega_{phase} = E = \hbar \omega_A \cosh \rho$$

$$\hbar c k_{phase} = cp = \hbar \omega_A \sinh \rho$$

Poincare Invariant Action $dS = L dt = p dq - H dt = \hbar d\Phi$ (phase)

Hamiltonian $H(p, q) = p \dot{q} - L$ vs. Lagrangian $L(\dot{q}, q) = p \dot{q} - H$

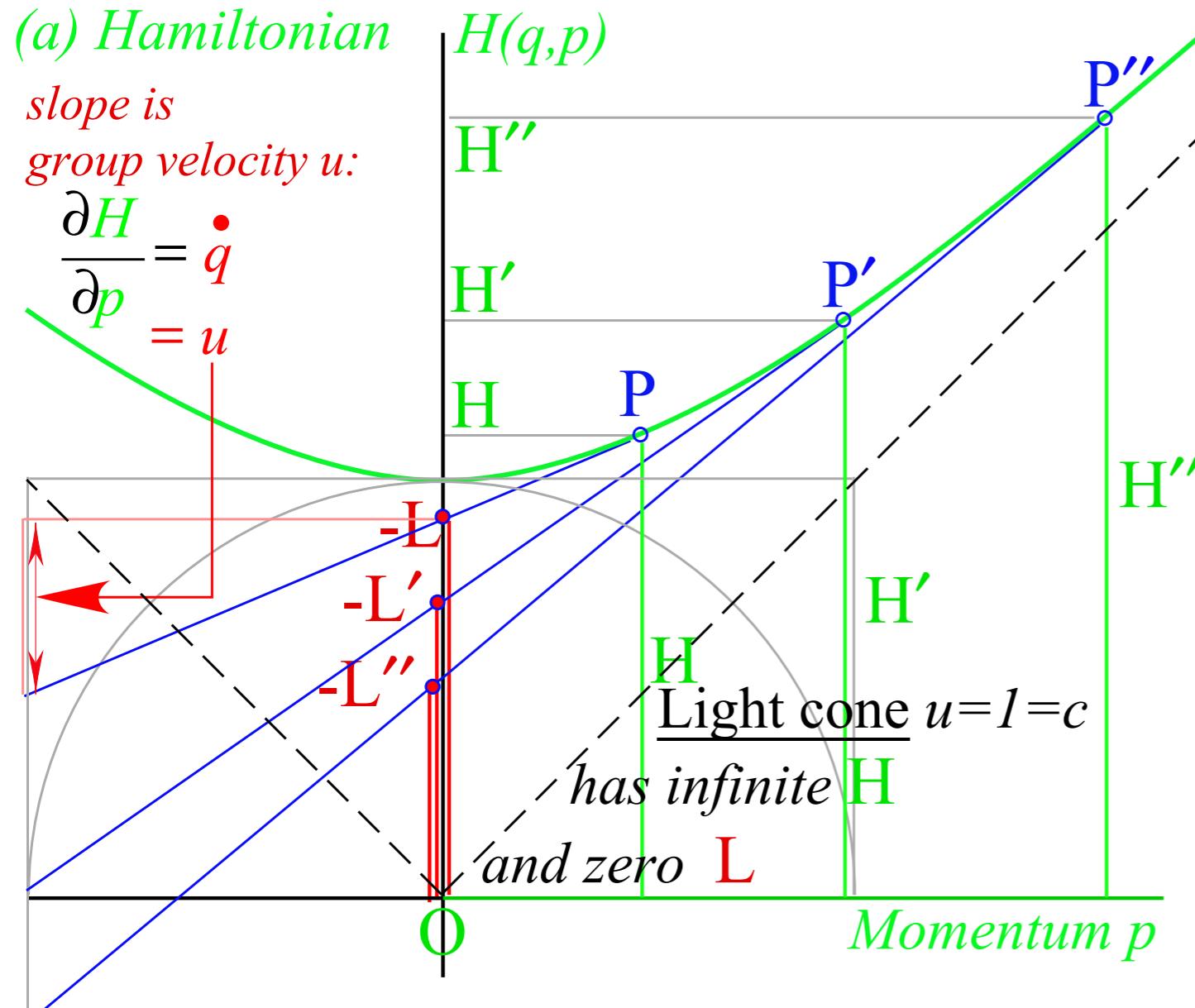
Contact transformation: (slope, -intercept) of H (or L) tangent determines the (X, Y coordinates) of L (or H).

(Also, called a *Legendre contact transformation* which is a special case of a *Huygens transformation* that uses contacting tangent *curves* instead of *lines*.)

(a) Hamiltonian

slope is group velocity u :

$$\frac{\partial H}{\partial p} = \dot{q} = u$$



Here *slope* is group velocity $u = \dot{q}$

Y-coordinate is *energy* $H = \hbar\omega$

(b) Lagrangian

$$radius = Mc^2$$

$$Velocity u = \dot{q}$$

$$L'' \quad L' \quad L \quad -H \quad -H' \quad -H''$$

$$\sigma'' \quad \sigma' \quad \sigma \quad -L \quad -L' \quad -L''$$

$$Q'' \quad Q' \quad Q \quad -Q \quad -Q' \quad -Q''$$

$$O \quad O' \quad O''$$

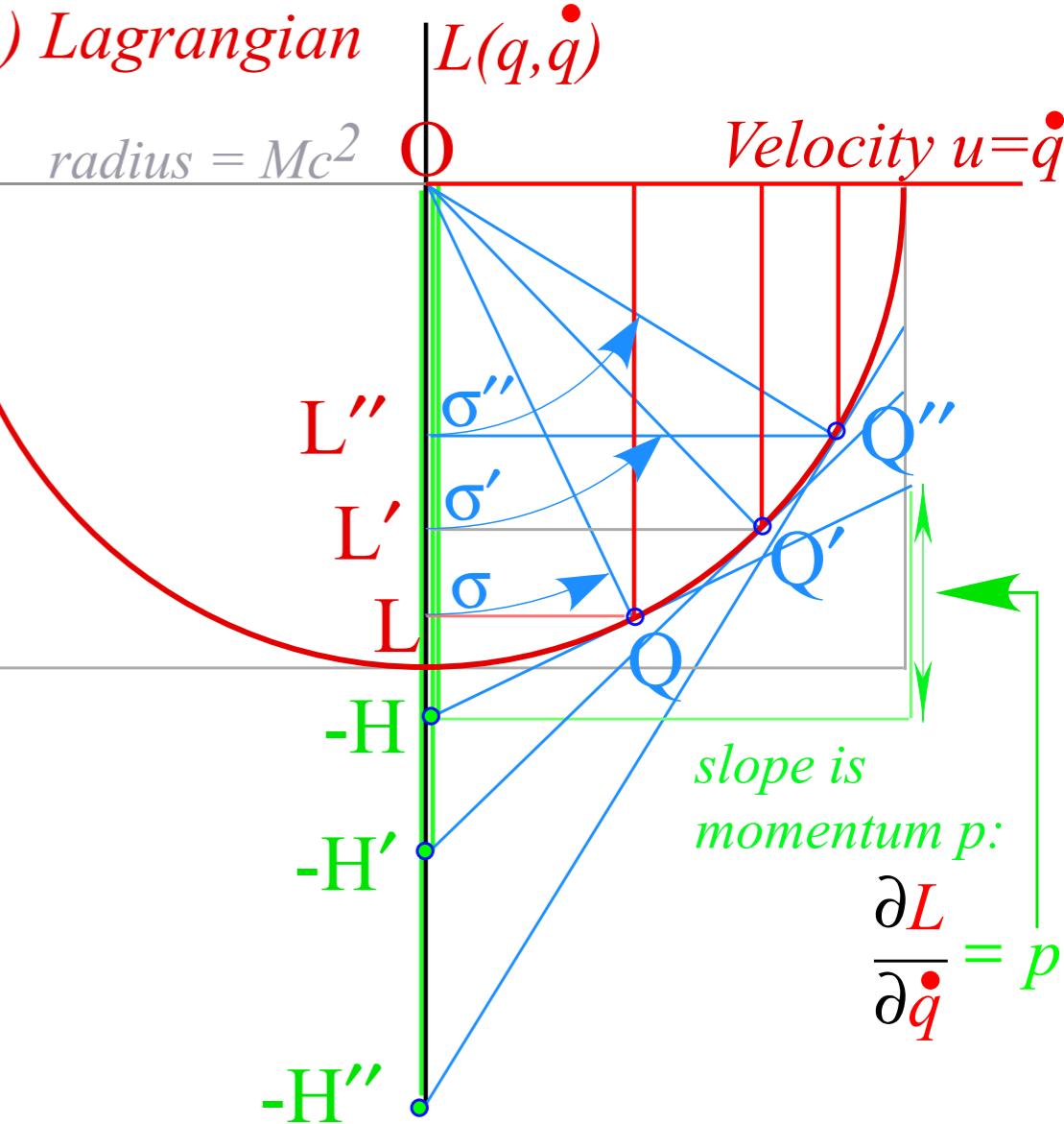
$$H'' \quad H' \quad H \quad -H \quad -H' \quad -H''$$

$$P'' \quad P' \quad P \quad -P \quad -P' \quad -P''$$

$$Momentum p$$

$$slope is momentum p:$$

$$\frac{\partial L}{\partial \dot{q}} = p$$

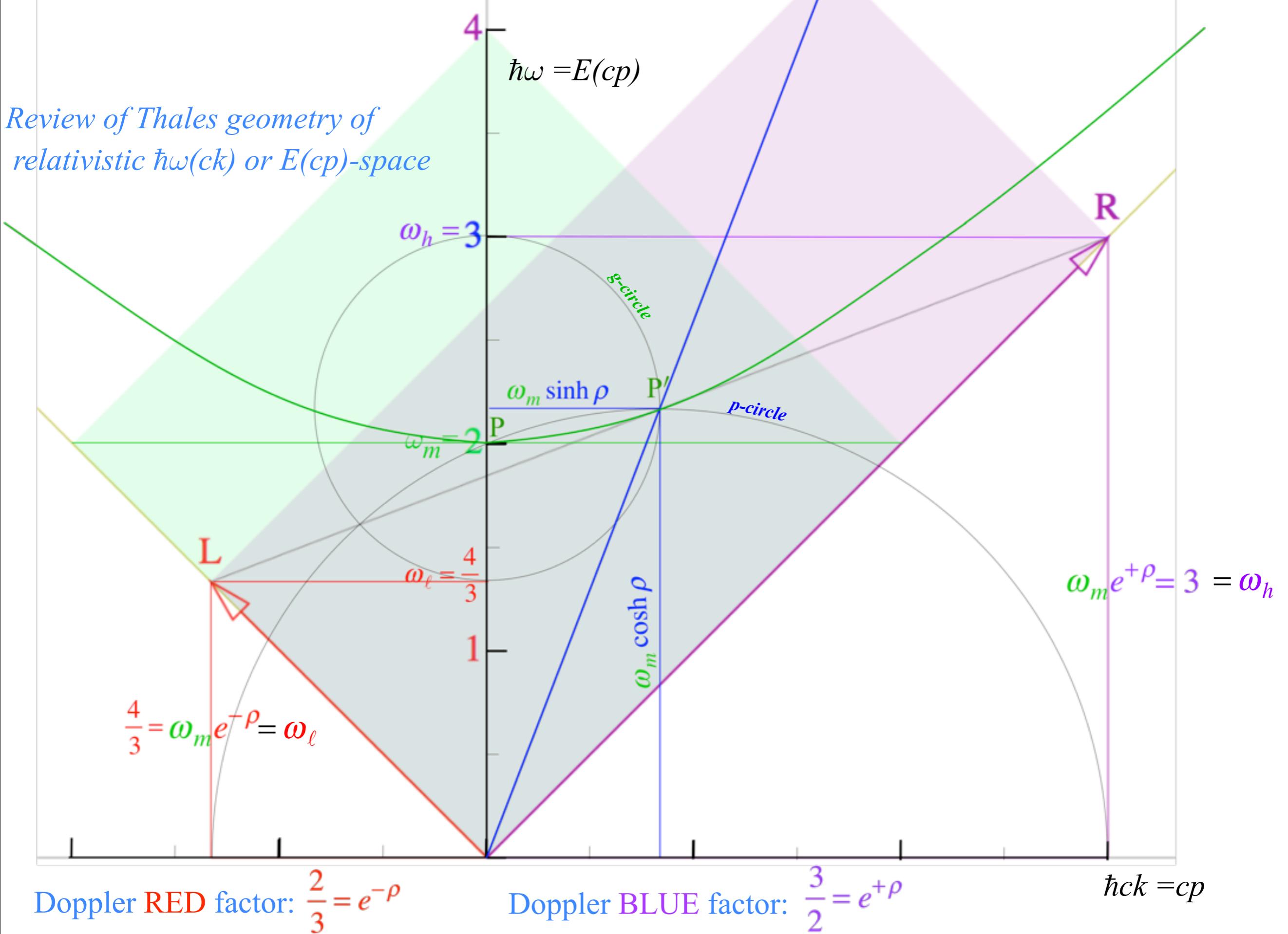


Here *slope* is momentum p

Y-coordinate is *phase rate* $L = \hbar\Phi$

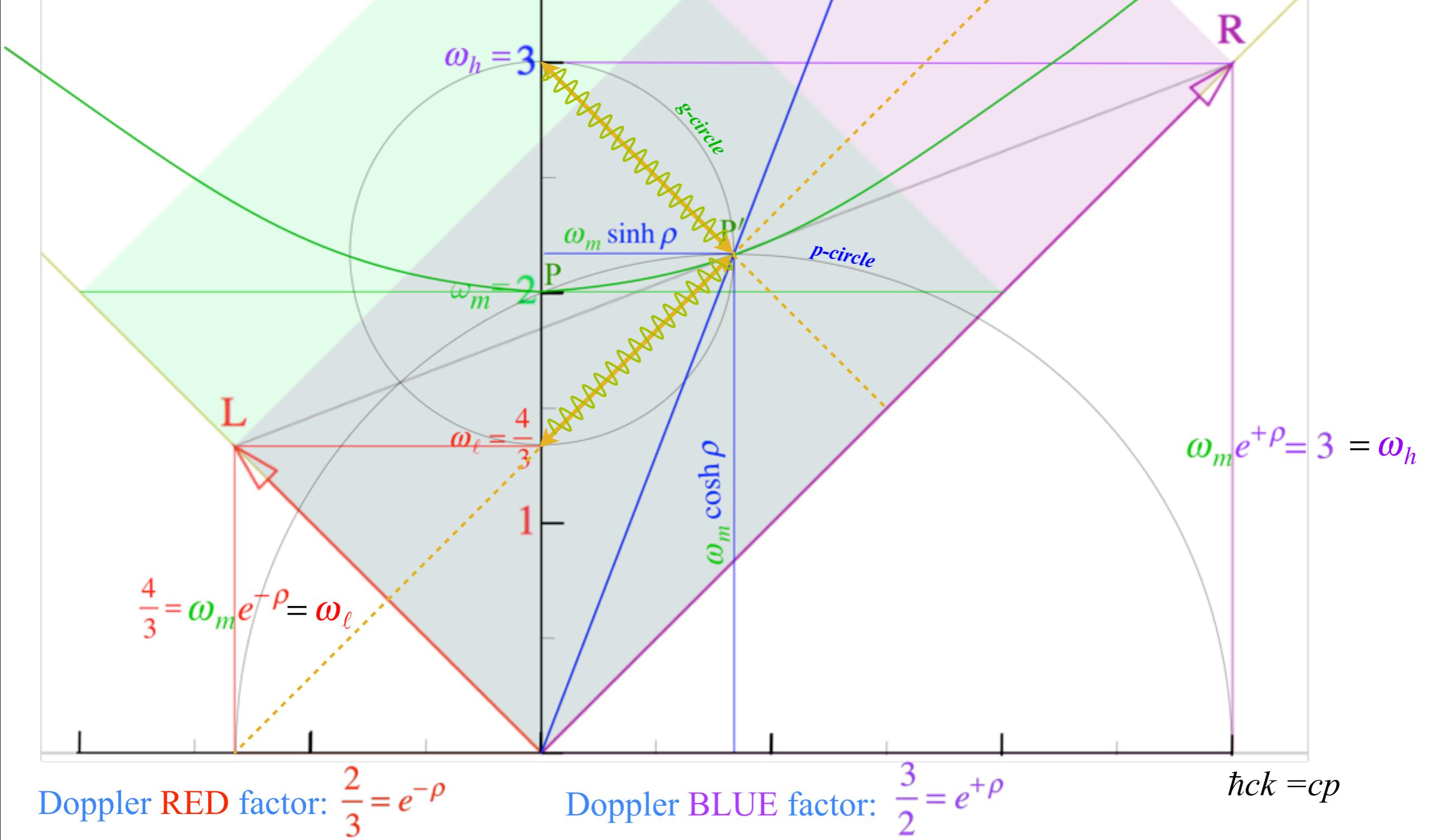
Relativistic optical transitions $|high\rangle = |\omega_h\rangle \Leftrightarrow |mid\rangle = |\omega_m\rangle \Leftrightarrow |low\rangle = |\omega_\ell\rangle$

Review of Thales geometry of relativistic $\hbar\omega(ck)$ or $E(cp)$ -space



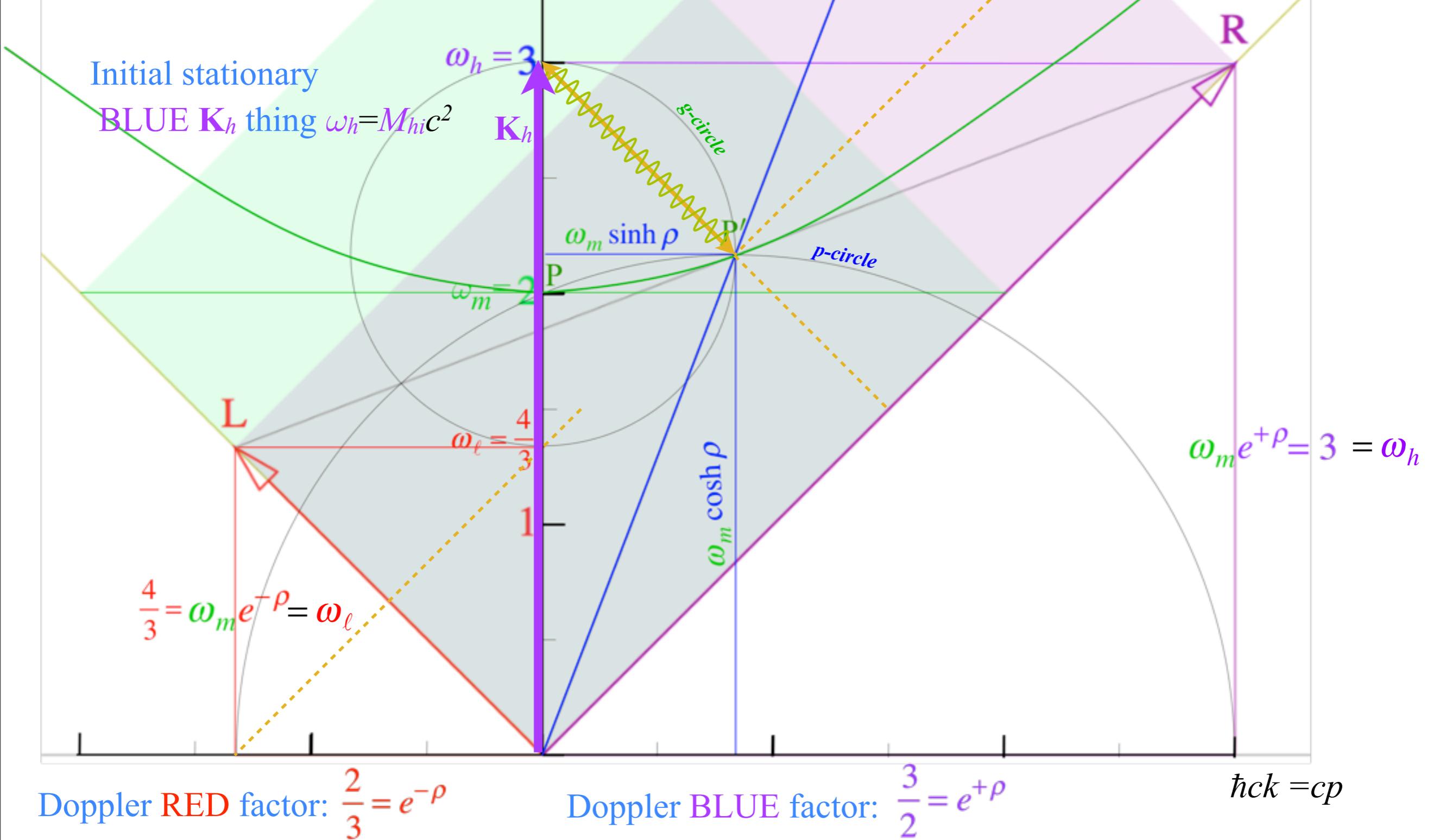
Relativistic optical transitions $|high\rangle=|\omega_h\rangle \Leftrightarrow |mid\rangle=|\omega_m\rangle \Leftrightarrow |low\rangle=|\omega_l\rangle$

Review of Thales geometry of relativistic $\hbar\omega(ck)$ or $E(cp)$ -space



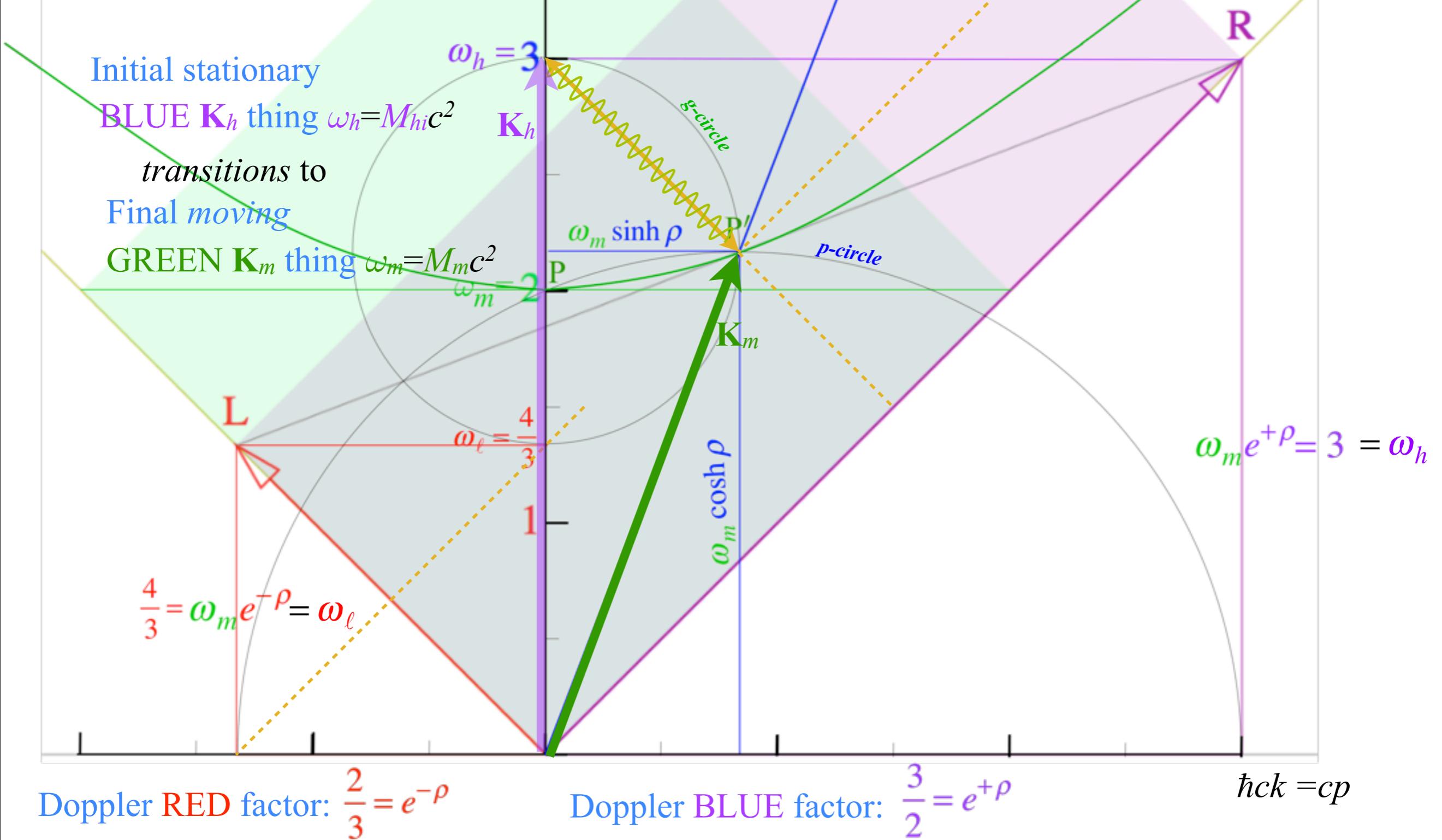
Relativistic optical transitions $|high\rangle=|\omega_h\rangle \Leftrightarrow |mid\rangle=|\omega_m\rangle \Leftrightarrow |low\rangle=|\omega_l\rangle$

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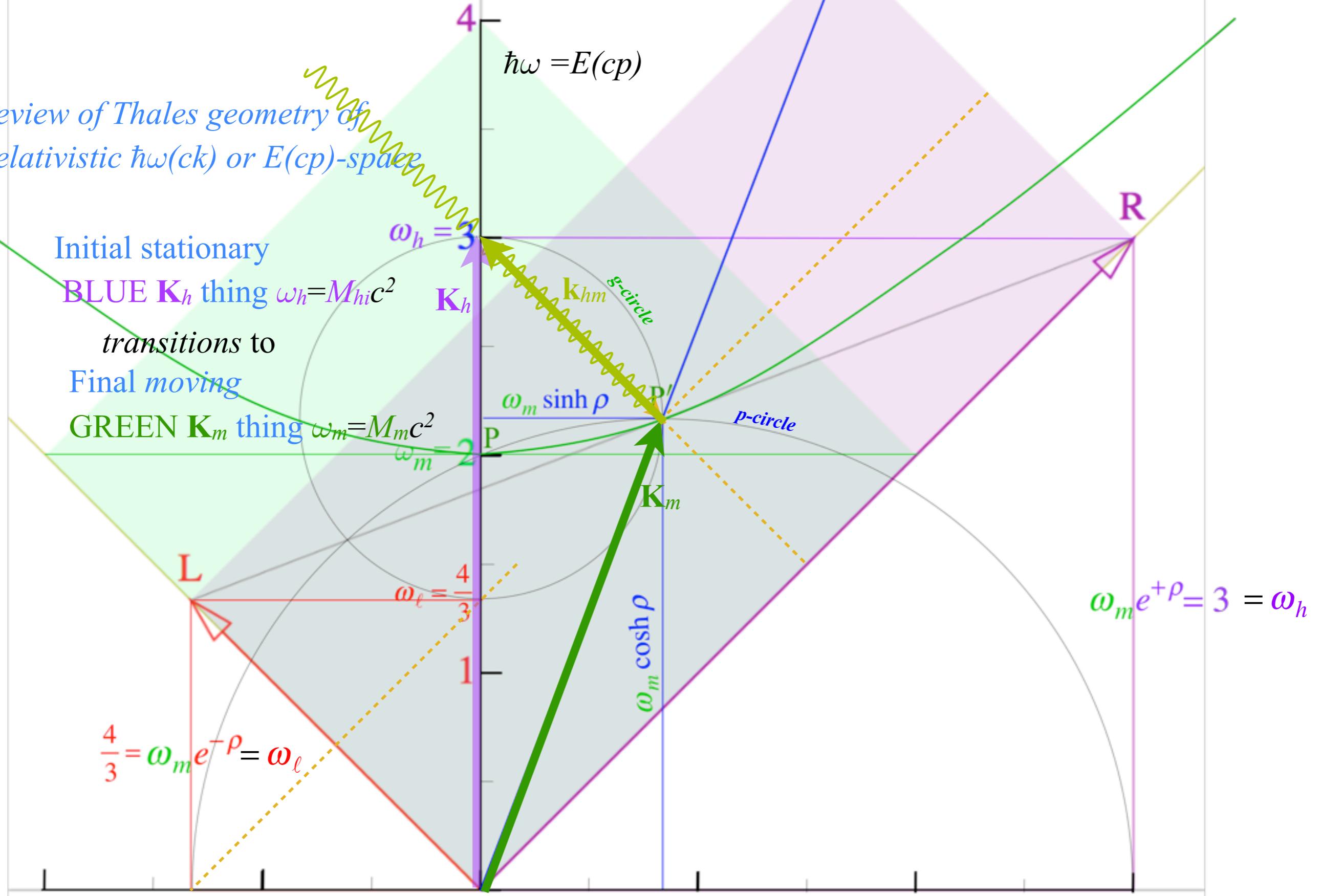
Review of Thales geometry of relativistic $\hbar\omega(ck)$ or $E(cp)$ -space



Relativistic optical transitions $|high\rangle = |\omega_h\rangle \leftrightarrow |mid\rangle = |\omega_m\rangle \leftrightarrow |low\rangle = |\omega_\ell\rangle$

Review of Thales geometry of relativistic $\hbar\omega(ck)$ or $E(cp)$ -space

Initial stationary
BLUE \mathbf{K}_h thing $\omega_h = M_{hi}c^2$
transitions to
Final moving
GREEN \mathbf{K}_m thing $\omega_m = M_{mi}c^2$



Doppler RED factor: $\frac{2}{3} = e^{-\rho}$

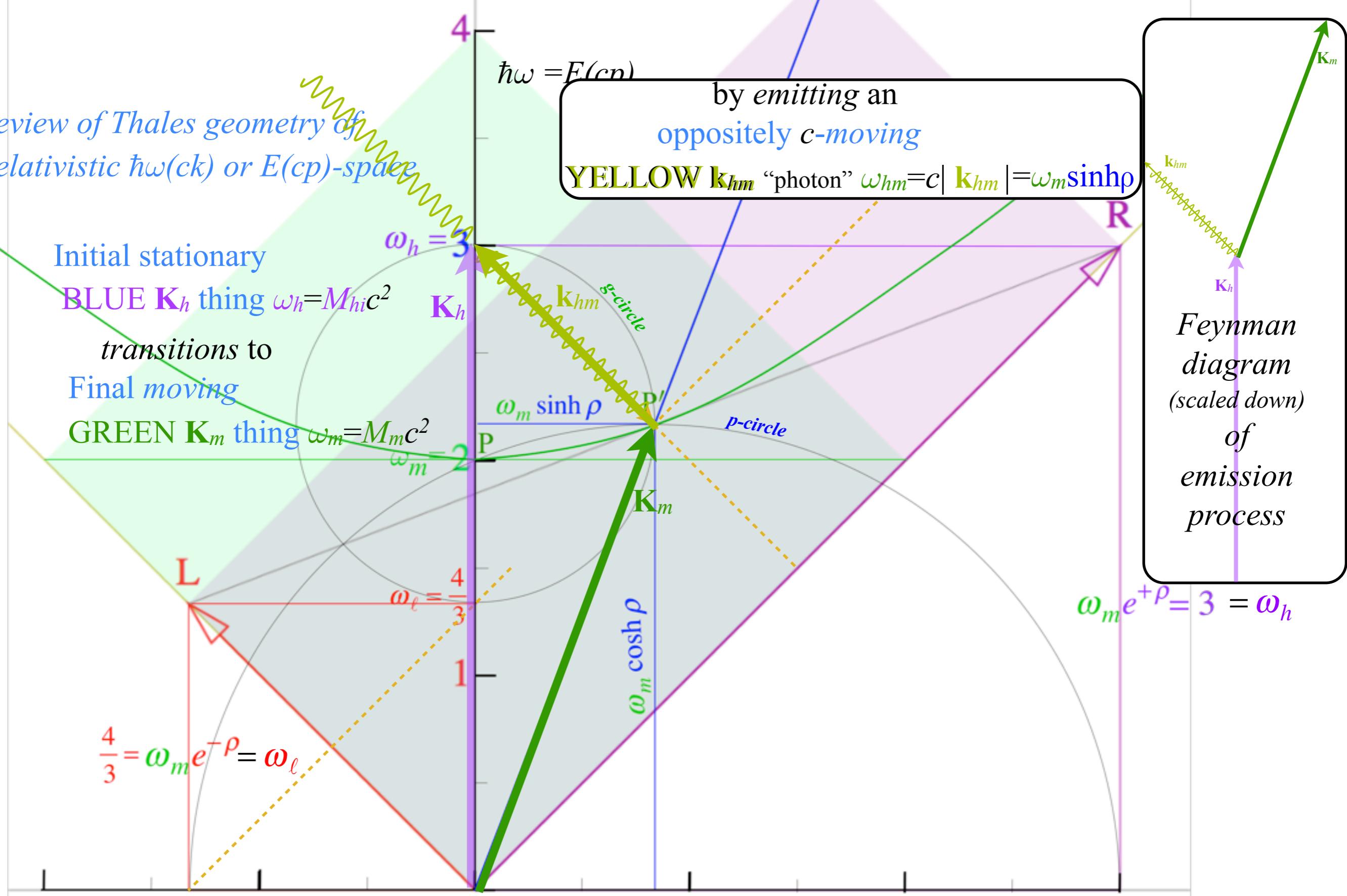
Doppler BLUE factor: $\frac{3}{2} = e^{+\rho}$

$\hbar ck = cp$

Relativistic optical transitions $|high\rangle = |\omega_h\rangle \leftrightarrow |mid\rangle = |\omega_m\rangle \leftrightarrow |low\rangle = |\omega_\ell\rangle$

Review of Thales geometry of relativistic $\hbar\omega(ck)$ or $E(cp)$ -space

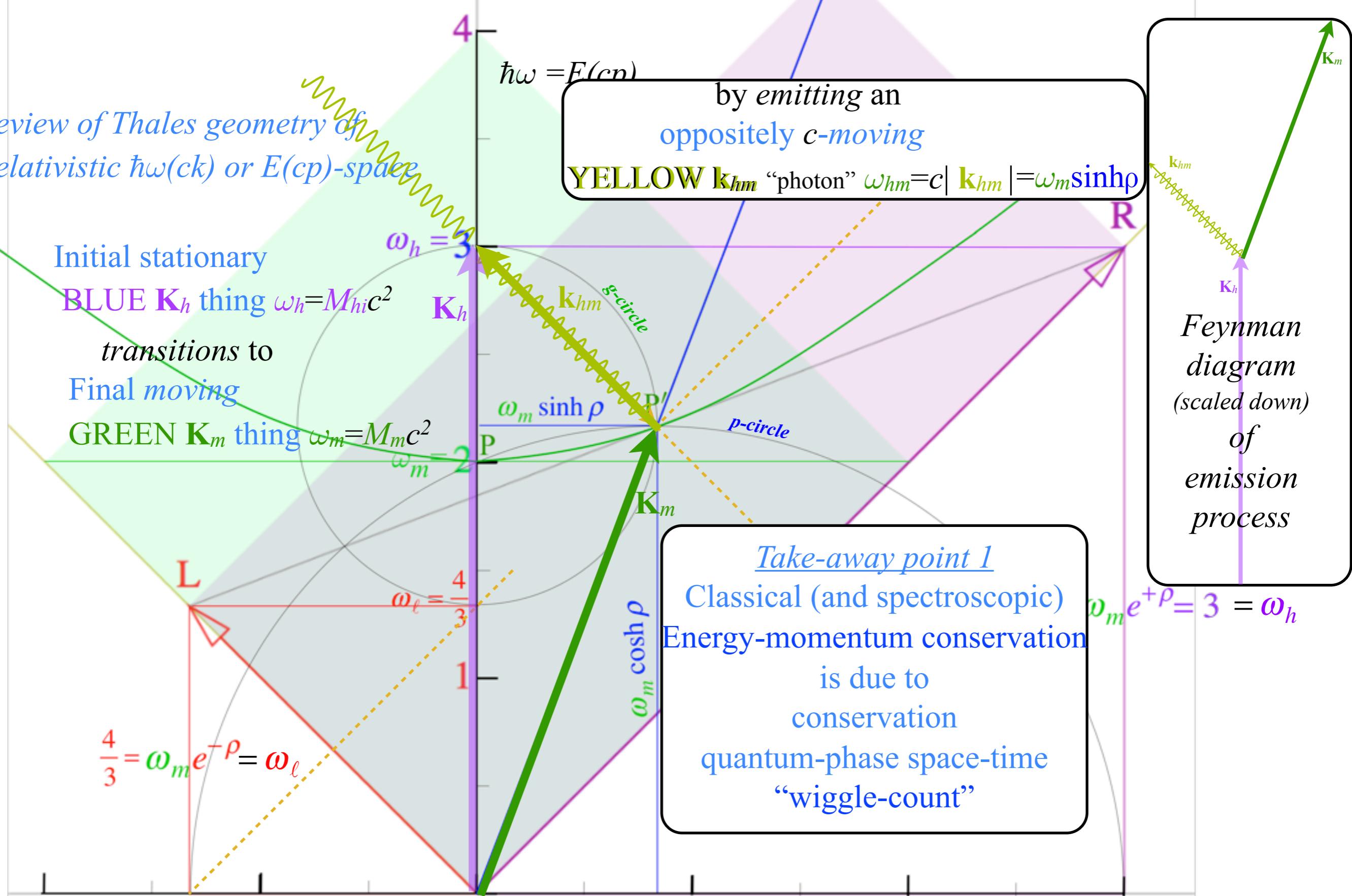
Initial stationary
BLUE \mathbf{K}_h thing $\omega_h = M_{hi}c^2$
transitions to
Final moving
GREEN \mathbf{K}_m thing $\omega_m = M_{mi}c^2$



Relativistic optical transitions $|high\rangle = |\omega_h\rangle \leftrightarrow |mid\rangle = |\omega_m\rangle \leftrightarrow |low\rangle = |\omega_\ell\rangle$

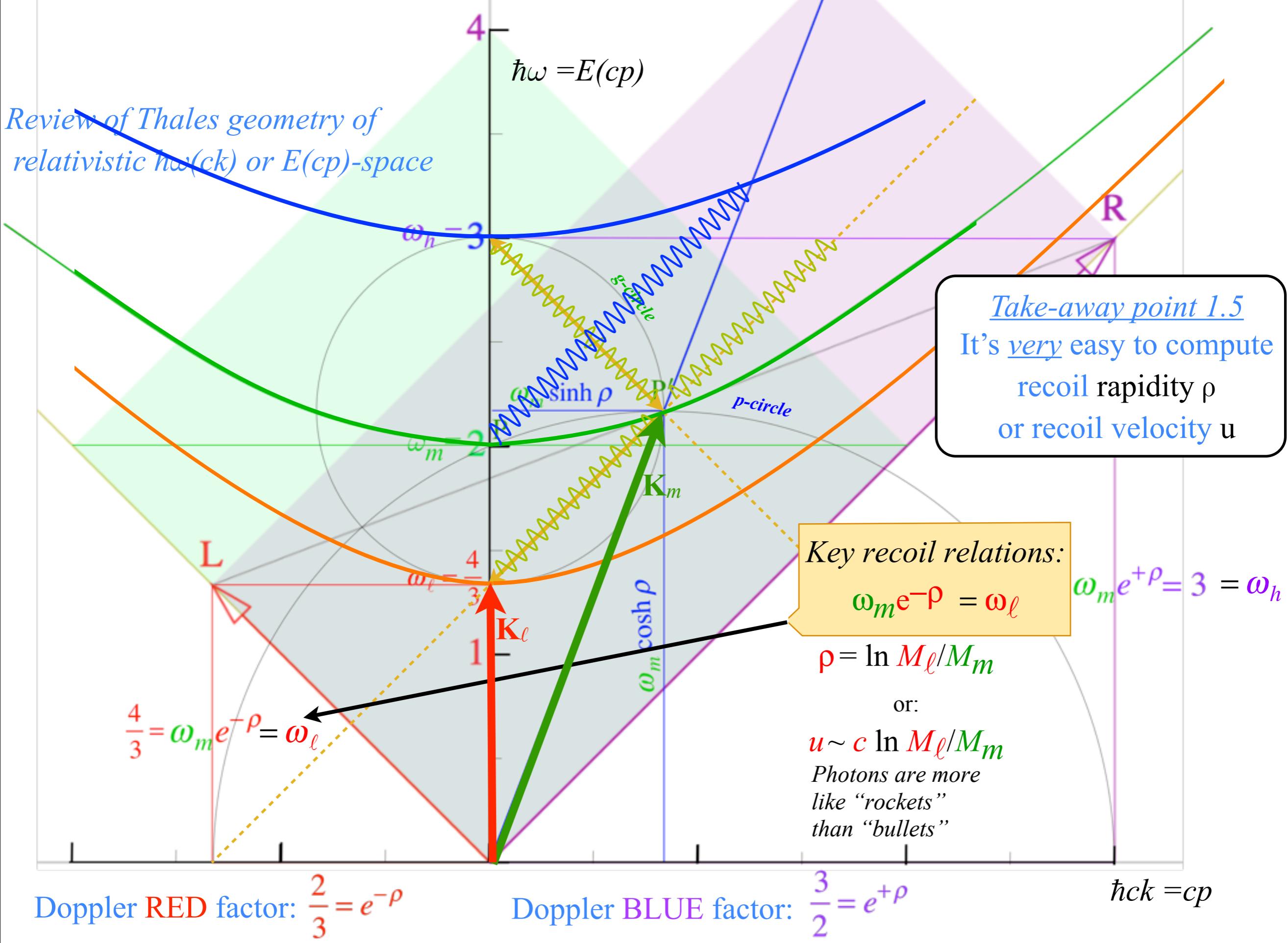
Review of Thales geometry of relativistic $\hbar\omega(ck)$ or $E(cp)$ -space

Initial stationary
BLUE \mathbf{K}_h thing $\omega_h = M_{hi}c^2$
transitions to
Final moving
GREEN \mathbf{K}_m thing $\omega_m = M_{mi}c^2$



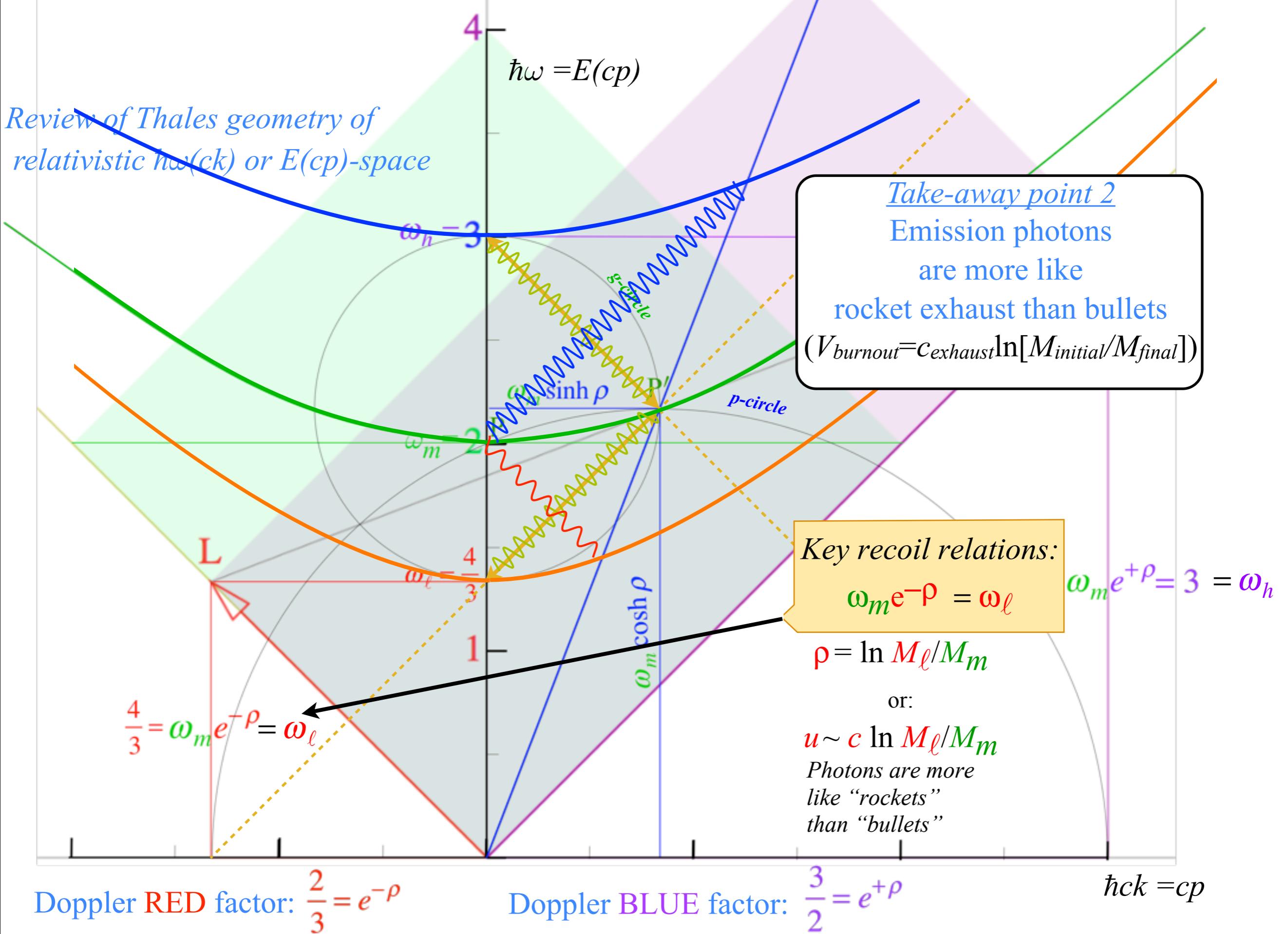
Relativistic optical transitions $|high\rangle=|\omega_h\rangle \Leftrightarrow |mid\rangle=|\omega_m\rangle \Leftrightarrow |low\rangle=|\omega_\ell\rangle$

~~Review of Thales geometry of relativistic $hw(ck)$ or $E(cp)$ -space~~



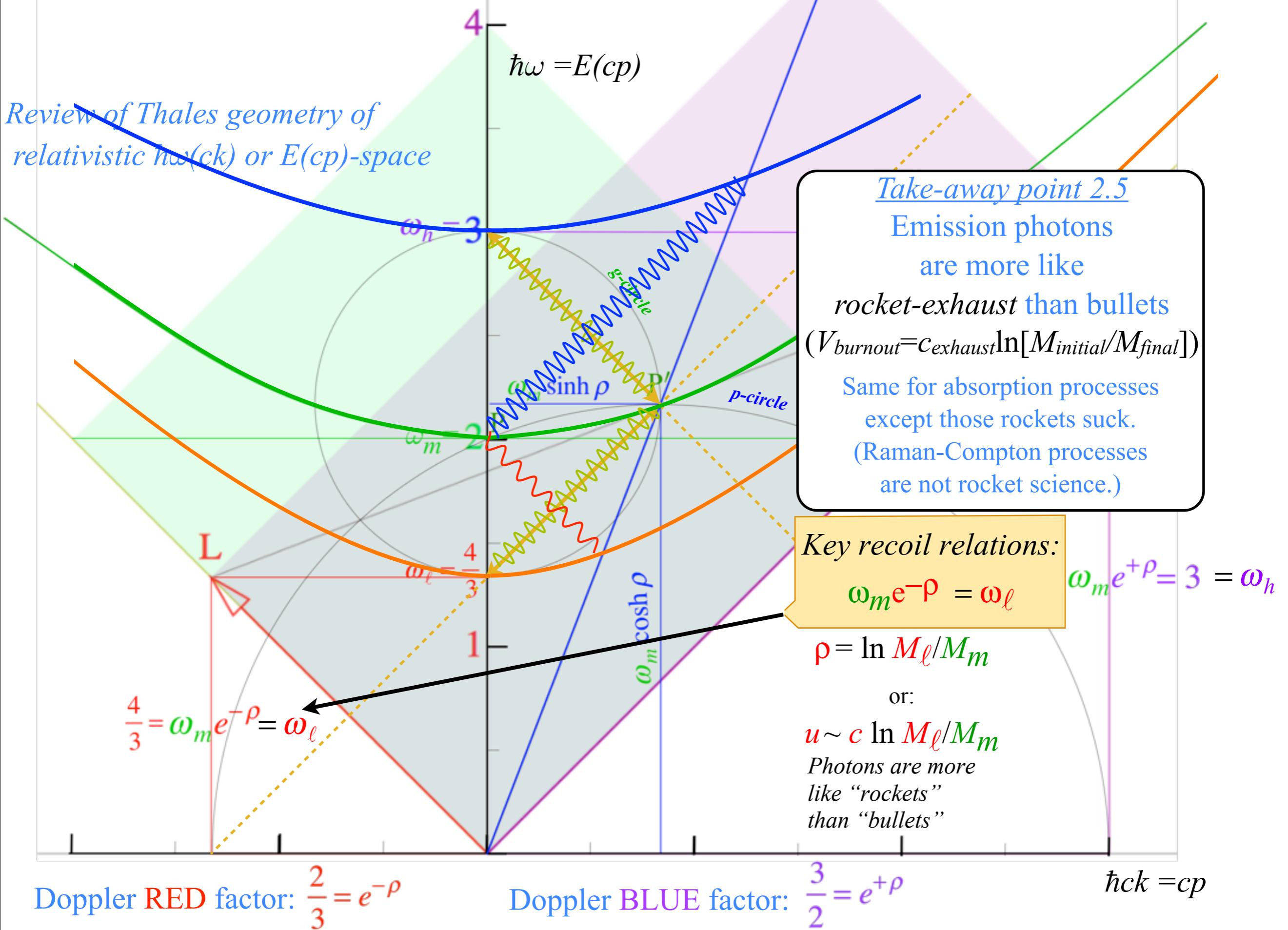
Relativistic optical transitions $|high\rangle=|\omega_h\rangle \Leftrightarrow |mid\rangle=|\omega_m\rangle \Leftrightarrow |low\rangle=|\omega_l\rangle$

~~Review of Thales geometry of relativistic $\hbar\omega(ck)$ or $E(cp)$ -space~~



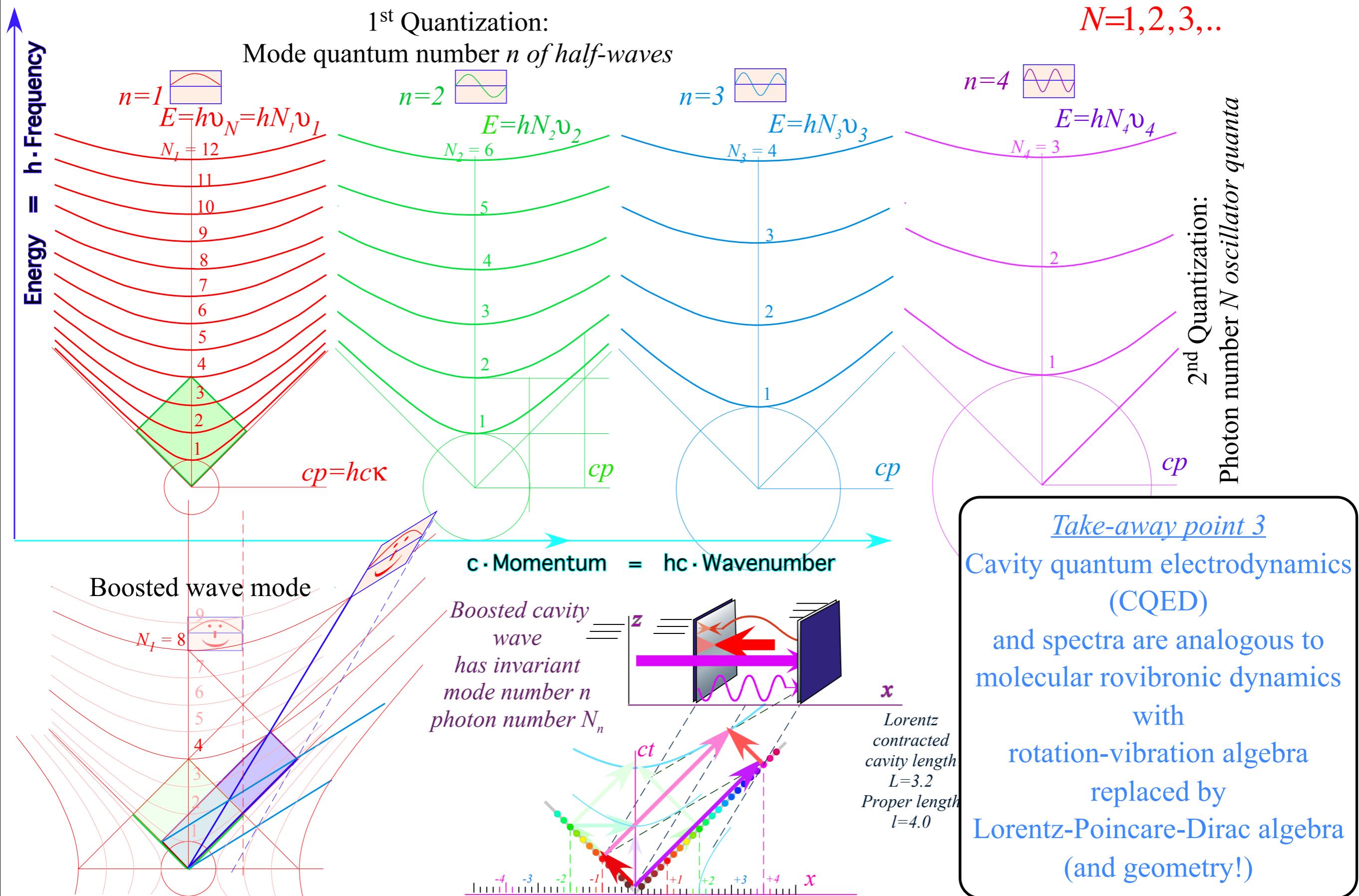
Relativistic optical transitions $|high\rangle = |\omega_h\rangle \leftrightarrow |mid\rangle = |\omega_m\rangle \leftrightarrow |low\rangle = |\omega_\ell\rangle$

Review of Thales geometry of relativistic $\hbar\omega(ck)$ or $E(cp)$ -space



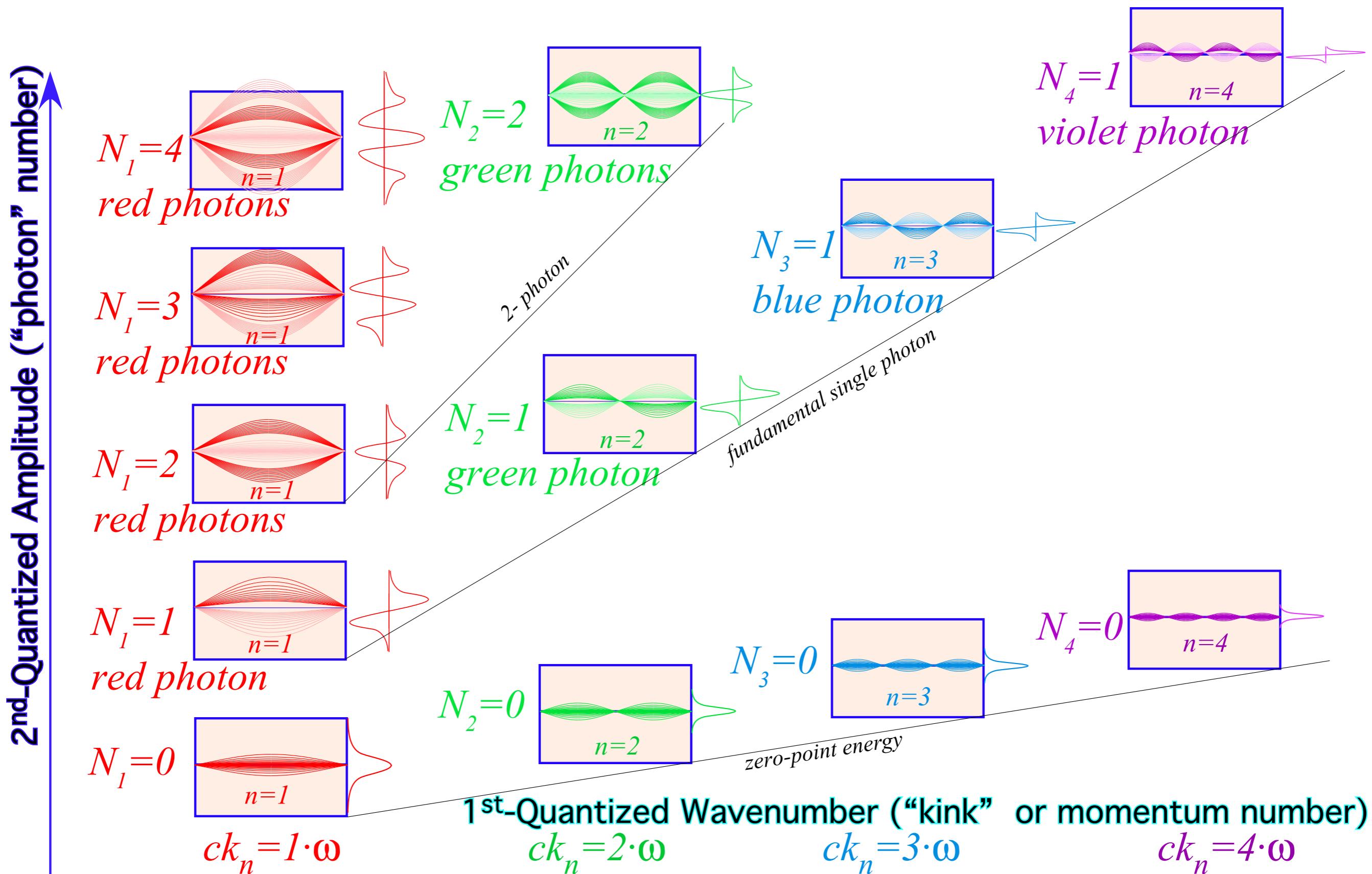
2nd Quantization: NEWS FLASH!!! $h\nu$ is actually $hN\nu$

($h\nu_{phase} = E = h\nu_A \cosh \rho$) is actually $(hN\nu_{phase} = E_N = hN\nu_A \cosh \rho$ with quantum numbers)



2nd Quantization: NEWS FLASH!!! $h\nu$ is actually $hN\nu$

$(h\nu_{phase}=E=h\nu_A \cosh \rho)$ is actually $(hN\nu_{phase}=E_N=hN\nu_A \cosh \rho \quad (N=1,2,\dots))$



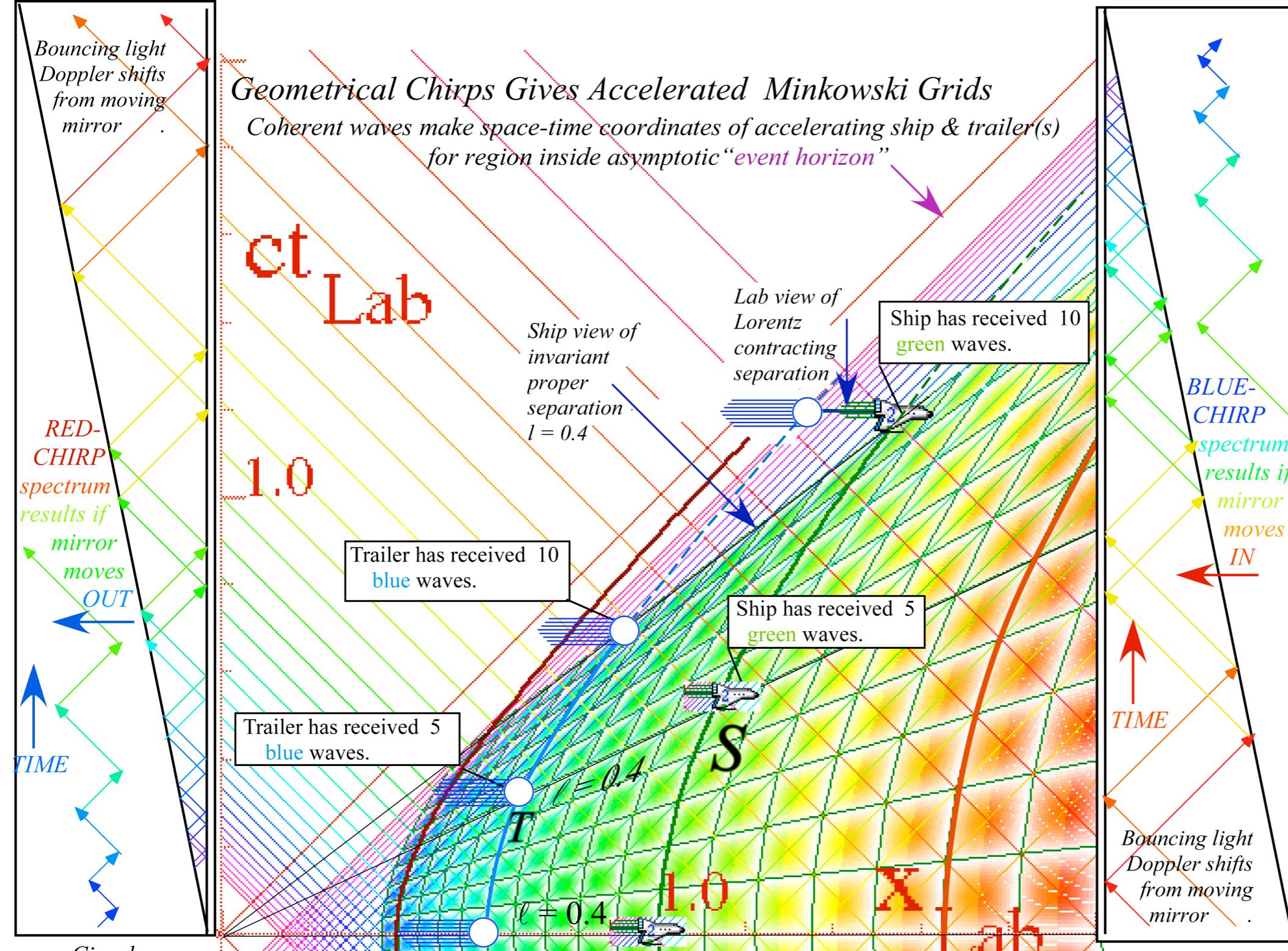


Fig. 8.2 Accelerated reference frames and their trajectories painted by chirped coherent light

