

Kepler Geometry of IHO (Isotropic Harmonic Oscillator) Elliptical Orbits

(Ch. 9 and Ch. 11 of Unit 1)

Constructing 2D IHO *orbital phasor “clock” dynamics* in uniform-body

Constructing 2D IHO orbits using *Kepler anomaly plots*

Mean-anomaly and eccentric-anomaly geometry

Calculus and vector geometry of IHO orbits

A confusing introduction to Coriolis-centrifugal force geometry (Derived better in Ch. 12)

Some Kepler’s “laws” for all central (isotropic) force $F(r)$ fields

Angular momentum invariance of IHO: $F(r)=-k\cdot r$ with $U(r)=k\cdot r^2/2$ (Derived here)

Angular momentum invariance of *Coulomb*: $F(r)=-GMm/r^2$ with $U(r)=-GMm\cdot/r$ (Derived in Unit 5)

Total energy $E=KE+PE$ invariance of IHO: $F(r)=-k\cdot r$ (Derived here)

Total energy $E=KE+PE$ invariance of *Coulomb*: $F(r)=-GMm/r^2$ (Derived in Unit 5)

Introduction to dual matrix operator contact geometry (based on IHO orbits)

Quadratic form ellipse $\mathbf{r}\cdot\mathbf{Q}\cdot\mathbf{r}=1$ vs. inverse form ellipse $\mathbf{p}\cdot\mathbf{Q}^{-1}\cdot\mathbf{p}=1$

Duality norm relations ($\mathbf{r}\cdot\mathbf{p}=1$)

Q -Ellipse tangents \mathbf{r}' normal to dual Q^{-1} -ellipse position \mathbf{p} ($\mathbf{r}'\cdot\mathbf{p}=0=\mathbf{r}\cdot\mathbf{p}'$)

Operator geometric sequences and eigenvectors

Alternative scaling of matrix operator geometry

Vector calculus of tensor operation

Web Links

Q: Where is this headed? A: Lagrangian-Hamiltonian duality

[BoxIt simulation: IHO orbits w/Stokes plot](#)

[RelaWavity Simulation: IHO orbital time rates of change](#)

[RelaWavity Simulation: Exegesis Plot](#)

This Lecture's Reference Link Listing

[Web Resources - front page](#)

[UAF Physics UTube channel](#)

[Quantum Theory for the Computer Age](#)

[Principles of Symmetry, Dynamics, and Spectroscopy](#)

[Classical Mechanics with a Bang!](#)

[Modern Physics and its Classical Foundations](#)

[2017 Group Theory for QM](#)

[2018 Adv CM](#)

[2018 AMOP](#)

[2019 Advanced Mechanics](#)

Lecture #7

[Pirelli Site: Phasors animimation](#)

[CMwithBang Lecture #6, page=70 \(9.10.18\)](#)

BoxIt Web Simulations:

[Generic/Default](#)

[Most Basic A-Type](#)

[Basic A-Type w/reference lines](#)

[Basic A-Type A-Type with Potential energy](#)

[A-Type with Potential energy and Stokes Plot](#)

[A-Type w/3 time rates of change](#)

[A-Type w/3 time rates of change with Stokes Plot](#)

[B-Type \(A=1.0, B=-0.05, C=0.0, D=1.0\)](#)

RelaWavity Web Elliptical Motion Simulations:

[Orbits with b/a=0.125](#)

[Orbits with b/a=0.5](#)

[Orbits with b/a=0.7](#)

[Exegesis with b/a=0.125](#)

[Exegesis with b/a=0.5](#)

[Exegesis with b/a=0.7](#)

[Contact Ellipsometry](#)

[Burning a hole in reality—design for a new laser may be powerful enough to pierce space-time](#)

[Trampoline mirror may push laser pulse through fabric of the Universe](#)

[Achieving Extreme Light Intensities using Optically Curved Relativistic Plasma Mirrors - Vincenti-prl-2019](#)

[A Soft Matter Computer for Soft Robots - Garrad-sr-2019](#)

[Correlated Insulator Behaviour at Half-Filling in Magic-Angle Graphene Superlattices - cao-n-2018](#)

Running Reference Link Listing

[RelaWavity Web Simulation: Contact Ellipsometry](#)

[BoxIt Web Simulation: Elliptical Motion \(A-Type\)](#)

[CMwBang Course: Site Title Page](#)

[Pirelli Relativity Challenge: Describing Wave Motion With Complex Phasors](#)

[UAF Physics UTube channel](#)

These **are** hot off the presses. Out in MISC for quick reference.

[Sorting ultracold atoms in a three-dimensional optical lattice in a realization of Maxwell's Demon - Kumar-n-2018](#)

[Synthetic three-dimensional atomic structures assembled atom by atom - Barredo-n-2018](#)

Older ones:

[Wave-particle duality of C60 molecules - Arndt-ltn-1999](#)

[Optical Vortex Knots - One Photon At A Time - Tempone-Wiltshire-Sr-2018](#)

[Baryon Deceleration by Strong Chromofields in Ultrarelativistic](#)

[Nuclear Collisions - Mishustin-PhysRevC-2007, APS Link & Abstract](#)

[Hadronic Molecules - Guo-x-2017](#)

[Hidden-charm pentaquark and tetraquark states - Chen-pr-2016](#)

BounceItIt Web Animation - Scenarios:

[49:1 y vs t, 49:1 V2 vs V1, 1:500:1 - 1D Gas Model w/ faux restorative force \(Cool\),](#)

[1:500:1 - 1D Gas \(Warm\), 1:500:1 - 1D Gas Model \(Cool, Zoomed in\),](#)

[Farey Sequence - Wolfram](#)

[Fractions - Ford-AMM-1938](#)

Monstermash BounceItIt Animations:

[1000:1 - V2 vs V1, 1000:1 with t vs x - Minkowski Plot](#)

[Quantum Revivals of Morse Oscillators and Farey-Ford Geometry - Li-Harter-2013](#)

[Quantum Revivals of Morse Oscillators and Farey-Ford Geometry - Li-Harter-cpl-2015](#)

[Quant. Revivals of Morse Oscillators and Farey-Ford Geom. - Harter-Li-CPL-2015 \(Publ.\)](#)

[Velocity Amplification in Collision Experiments Involving Superballs-Harter-1971](#)

WaveIt Web Animation - Scenarios:

[Quantum Carpet, Quantum Carpet wMBars,](#)

[Quantum Carpet BCar, Quantum Carpet BCar wMBars](#)

[Wave Node Dynamics and Revival Symmetry in Quantum Rotors - Harter-JMS-2001](#)

[Wave Node Dynamics and Revival Symmetry in Quantum Rotors - Harter-jms-2001 \(Publ.\)](#)

[AJP article on superball dynamics](#)

[AAPT Summer Reading List](#)

[Scitation.org - AIP publications](#)

[HarterSoft Youtube Channel](#)

Prior to Lecture #7

[Velocity Amplification in Collision Experiments Involving Superballs](#)

[MIT OpenCourseWare: High School/Physics/Impulse and Momentum](#)

[Hubble Site: Supernova - SN 1987A](#)

BounceIt Web Animation - Scenarios:

[Generic Scenario: 2-Balls dropped no Gravity \(7:1\) - V vs V Plot \(Power=4\)](#)

[1-Ball dropped w/Gravity=0.5 w/Potential Plot: Power=1, Power=4](#)

[7:1 - V vs V Plot: Power=1](#)

[3-Ball Stack \(10:3:1\) w/Newton plot \(y vs t\) - Power=4](#)

[3-Ball Stack \(10:3:1\) w/Newton plot \(y vs t\) - Power=1](#)

[3-Ball Stack \(10:3:1\) w/Newton plot \(y vs t\) - Power=1 w/Gaps](#)

[4-Ball Stack \(27:9:3:1\) w/Newton plot \(y vs t\) - Power=4](#)

[4-Newton's Balls \(1:1:1:1\) w/Newtonian plot \(y vs t\) - Power=4 w/Gaps](#)

[6-Ball Totally Inelastic \(1:1:1:1:1\) w/Gaps: Newtonian plot \(t vs x\), V6 vs V5 plot](#)

[5-Ball Totally Inelastic Pile-up w/ 5-Stationary-Balls - Minkowski plot \(t vs x1\) w/Gaps](#)

[1-Ball Totally Inelastic Pile-up w/ 5-Stationary-Balls - Vx2 vs Vx1 plot w/Gaps](#)

BounceIt Dual plots

$m_1:m_2 = 3:1$

[v2 vs v1 and V2 vs V1, \(v1, v2\)=\(1, 0.1\), \(v1, v2\)=\(1, 0\)](#)

[y2 vs y1 plots: \(v1, v2\)=\(1, 0.1\), \(v1, v2\)=\(1, 0\), \(v1, v2\)=\(1, -1\)](#)

[Estrangian plot V2 vs V1: \(v1, v2\)=\(0, 1\), \(v1, v2\)=\(1, -1\)](#)

$m_1:m_2 = 4:1$

[v2 vs v1, y2 vs y1](#)

$m_1:m_2 = 100:1$, (v1, v2)=(1, 0): [V2 vs V1 Estrangian plot, y2 vs y1 plot](#)

[With g=0 and 70:10 mass ratio](#)

[With non zero g, velocity dependent damping and mass ratio of 70:35](#)

[M1=49, M2=1 with Newtonian time plot](#)

[M1=49, M2=1 with V2 vs V1 plot](#)

[Example with friction](#)

[Low force constant with drag displaying a Pass-thru, Fall-Thru, Bounce-Off](#)

[m1:m2= 3:1 and \(v1, v2\) = \(1, 0\) Comparison with Estrangian](#)

X2 paper: [Velocity Amplification in Collision Experiments Involving Superballs - Harter, et. al. 1971 \(pdf\)](#)

Car Collision Web Simulator: <https://modphys.hosted.uark.edu/markup/CMMotionWeb.html>

Superball Collision Web Simulator: <https://modphys.hosted.uark.edu/markup/BounceItWeb.html>; with Scenarios: [1007](#)

[BounceIt web simulation with g=0 and 70:10 mass ratio](#)

[With non zero g, velocity dependent damping and mass ratio of 70:35](#)

[Elastic Collision Dual Panel Space vs Space: Space vs Time \(Newton\), Time vs. Space\(Minkowski\)](#)

[Inelastic Collision Dual Panel Space vs Space: Space vs Time \(Newton\), Time vs. Space\(Minkowski\)](#)

[Matrix Collision Simulator: M1=49, M2=1 V2 vs V1 plot <<Under Construction>>](#)

More Advanced QM and classical references at the end of this Lecture

→ *Introducing 2D IHO orbits and phasor geometry*
Phasor “clock” geometry

Isotropic Harmonic Oscillator *phase dynamics in uniform-body*

I.H.O. Force law

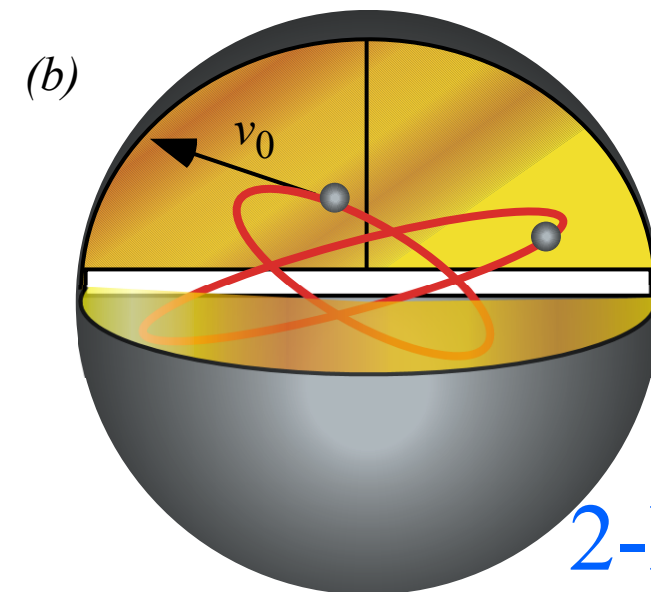
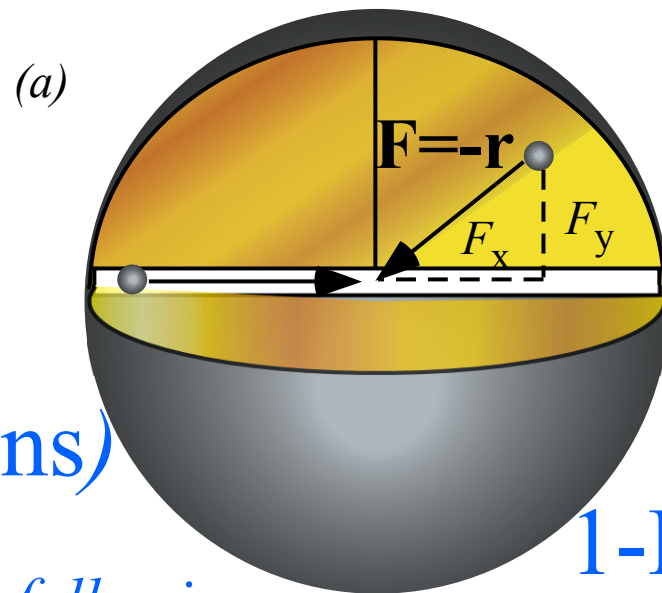
$$F = -x \quad (1\text{-Dimension})$$

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Each dimension x , y , or z obeys the following:

$$\text{Total } E = KE + PE = \frac{1}{2}mv^2 + U(x) = \frac{1}{2}mv^2 + \frac{1}{2}kx^2 = \text{const.}$$

Equations for x -motion
 $[x(t) \text{ and } v_x=v(t)]$ are
 given first. They apply
 as well to dimensions
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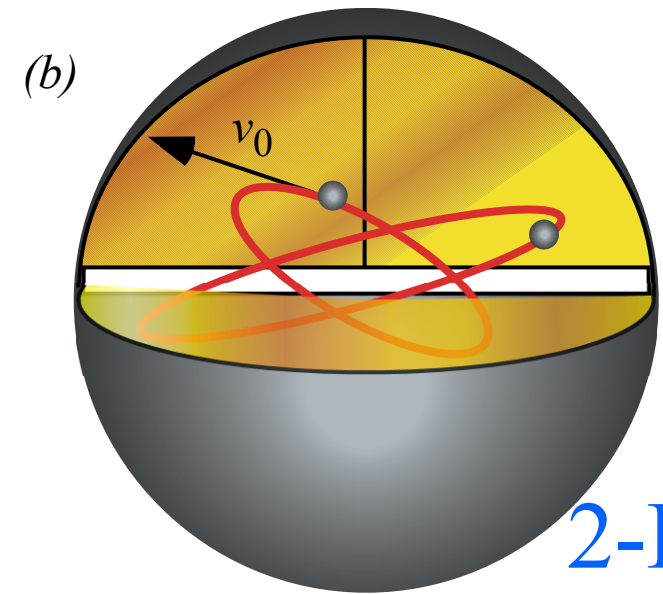
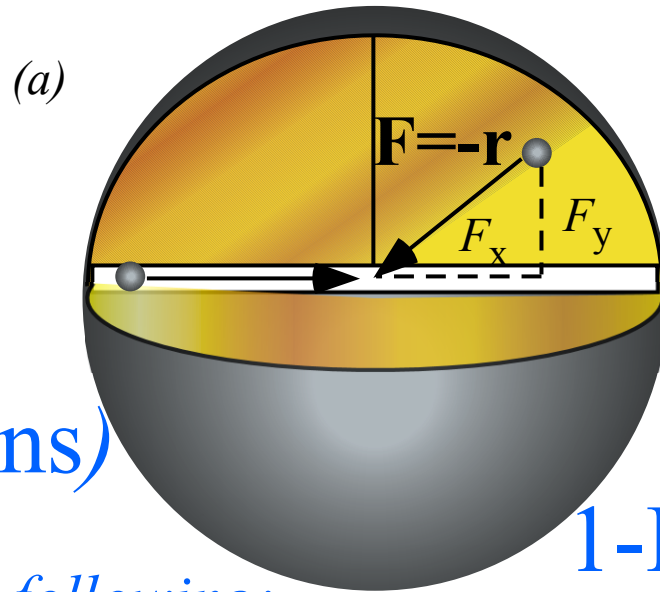


Unit 1
 Fig. 9.10

2-D or 3-D
 (Paths are *always* 2-D
 ellipses if viewed
 right!)

Isotropic Harmonic Oscillator *phase dynamics in uniform-body*

Unit 1
Fig. 9.10



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$$1 = \frac{mv^2}{2E} + \frac{kx^2}{2E} = \left(\frac{v}{\sqrt{2E/m}} \right)^2 + \left(\frac{x}{\sqrt{2E/k}} \right)^2$$

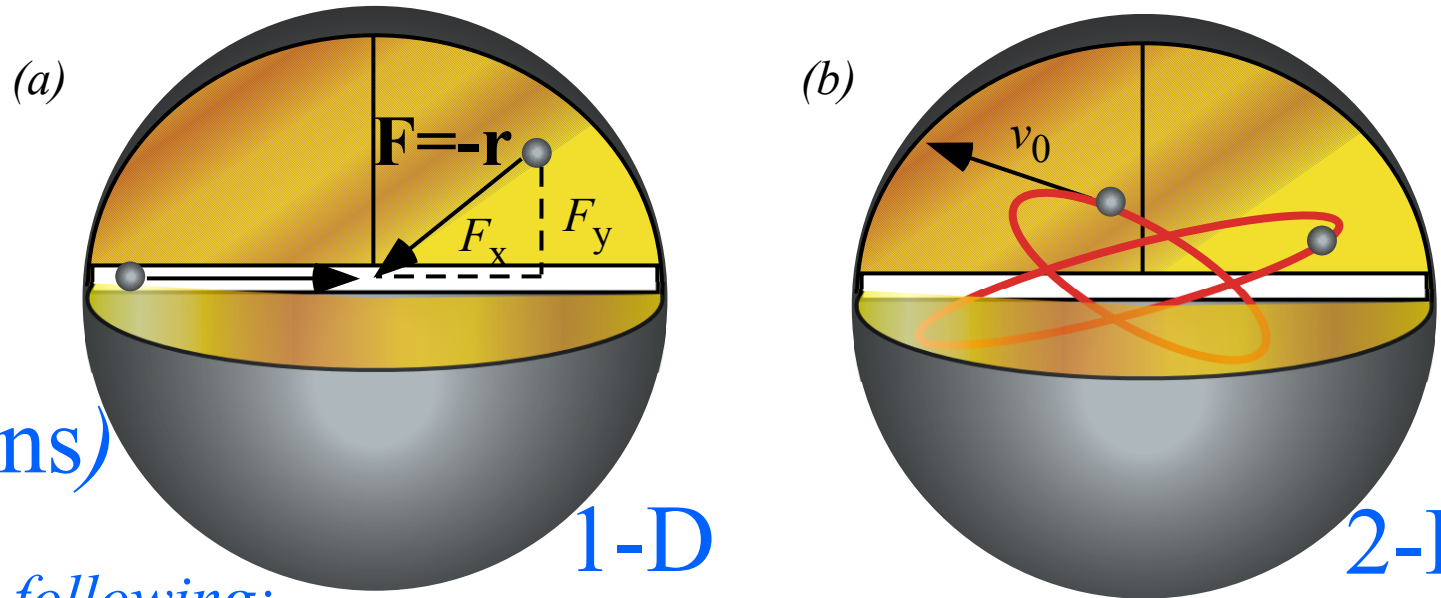
$$1 = \frac{mv^2}{2E} + \frac{kx^2}{2E} = (\cos\theta)^2 + (\sin\theta)^2$$

Another example of the old “scale-a-circle” trick...

Let : **(1)** *velocity:* $v = \sqrt{2E/m} \cos\theta$, and : **(2)** *position:* $x = \sqrt{2E/k} \sin\theta$

Isotropic Harmonic Oscillator *phase dynamics in uniform-body*

Unit 1
Fig. 9.10



1-D

2-D or 3-D

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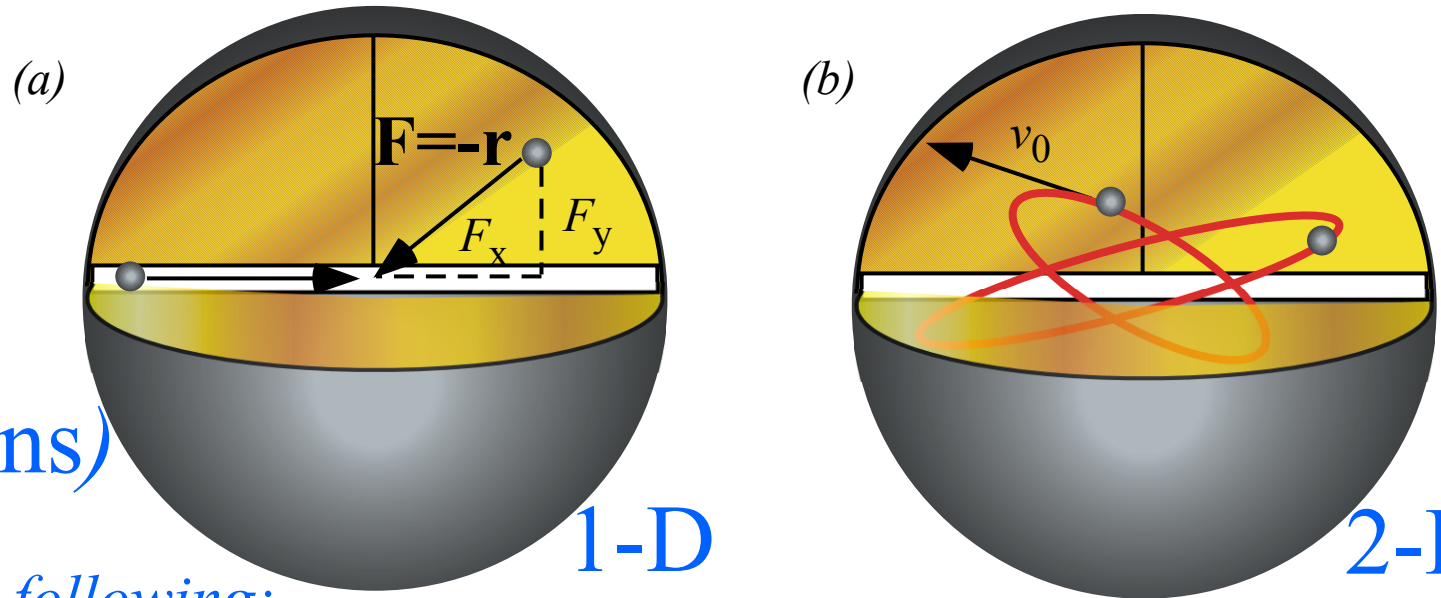
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Isotropic Harmonic Oscillator *phase dynamics in uniform-body*

Unit 1
Fig. 9.10



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Let : **(1)** $v = \sqrt{2E/m} \cos\theta$, and : **(2)** $x = \sqrt{2E/k} \sin\theta$ def. **(3)** $\omega = \frac{d\theta}{dt}$

velocity:

position:

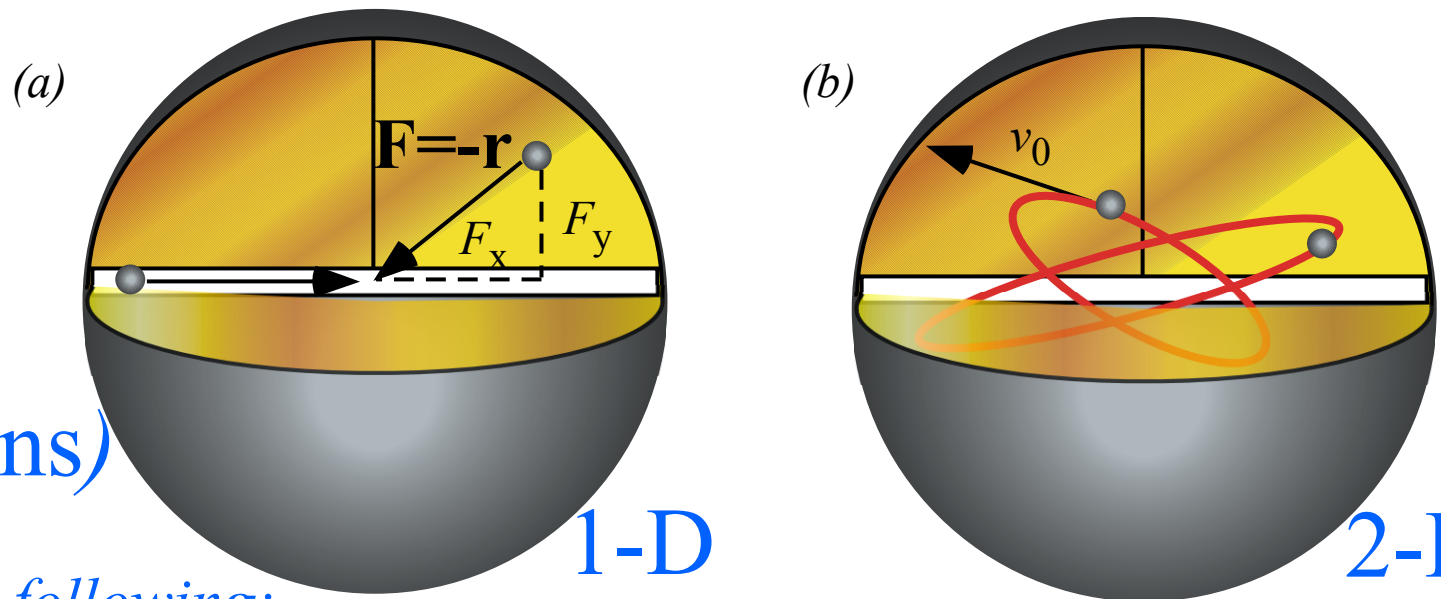
angular velocity:

$$\sqrt{\frac{2E}{m}} \cos\theta = v = \frac{dx}{dt}$$

by (1)

Isotropic Harmonic Oscillator *phase dynamics in uniform-body*

Unit 1
Fig. 9.10



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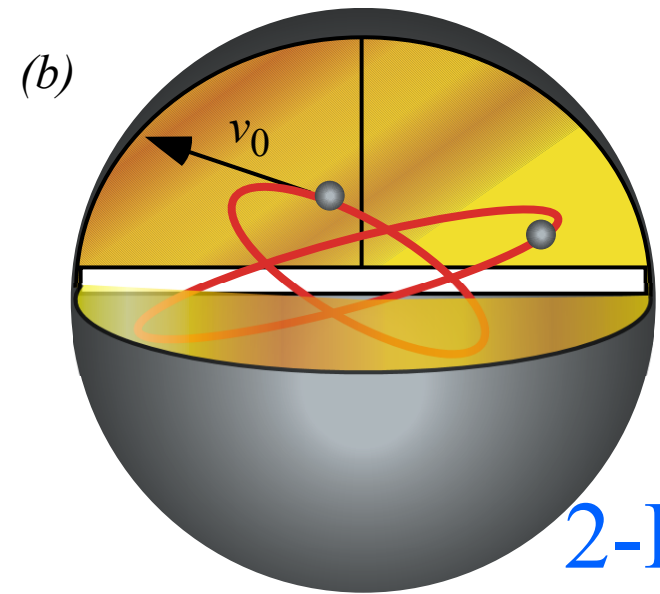
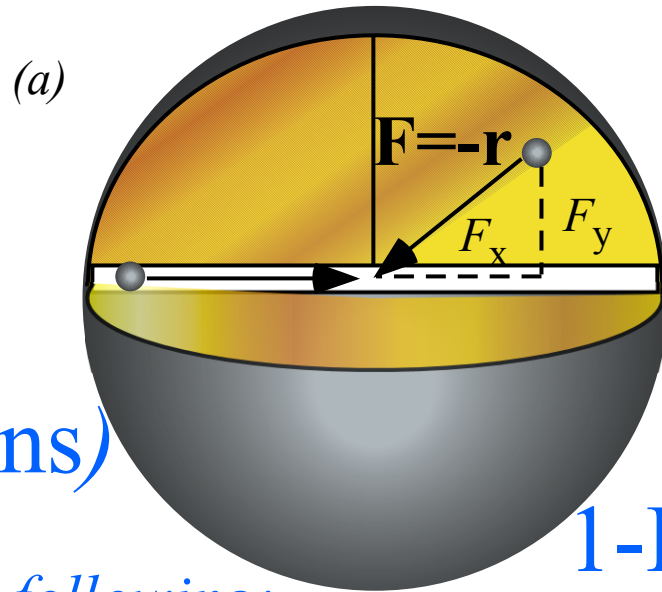
Let : **(1)** $v = \sqrt{2E/m} \cos\theta$, and : **(2)** $x = \sqrt{2E/k} \sin\theta$ angular velocity: $\omega = \frac{d\theta}{dt}$ def. **(3)**

$$\sqrt{\frac{2E}{m}} \cos\theta = v = \frac{dx}{dt} = \frac{d\theta}{dt} \frac{dx}{d\theta}$$

by (1)
simple calculus

Isotropic Harmonic Oscillator *phase dynamics* in uniform-body

Unit 1
Fig. 9.10



2-D or 3-D
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by (1) by def. (3) by (2)

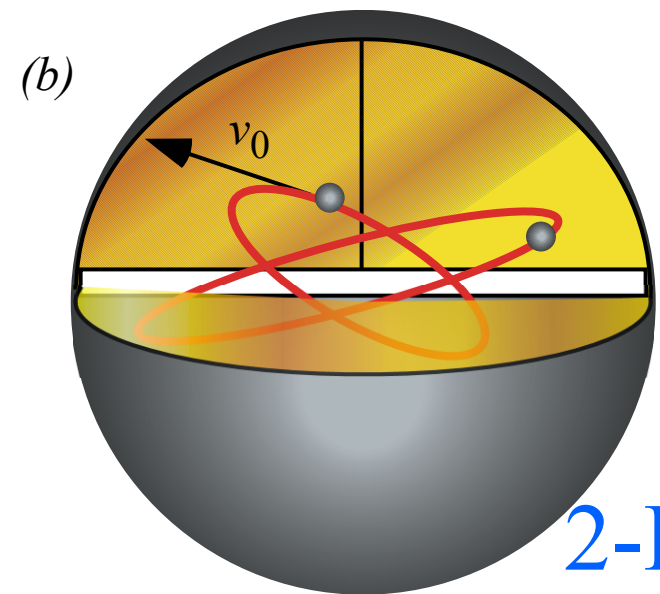
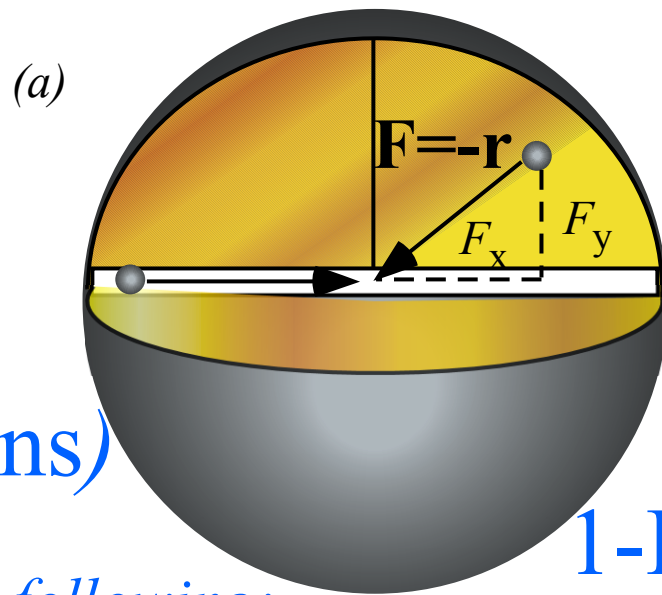
Isotropic Harmonic Oscillator *phase dynamics* in uniform-body

Unit 1
Fig. 9.10

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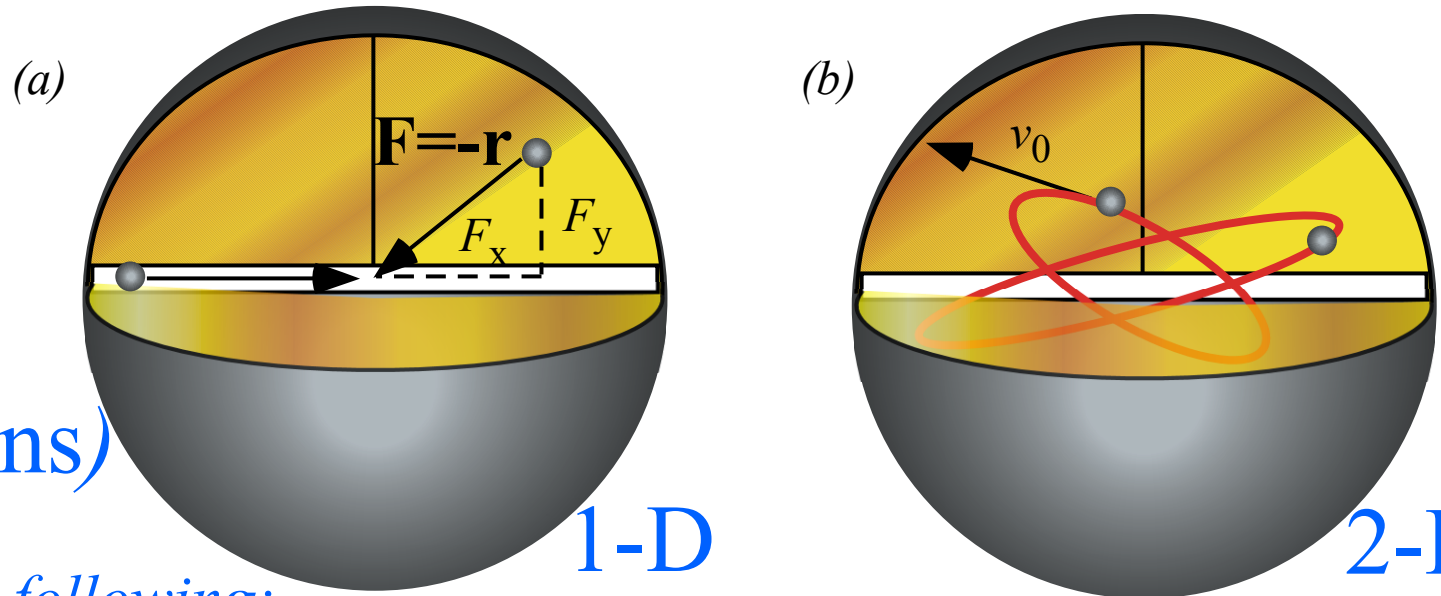
$$\sqrt{\frac{2E}{m}} \cos\theta \stackrel{\text{by (1)}}{=} v = \frac{dx}{dt} \stackrel{\text{by def. (3)}}{=} \frac{d\theta}{dt} \frac{dx}{d\theta} \stackrel{\text{by (2)}}{=} \omega \frac{dx}{d\theta} = \omega \sqrt{\frac{2E}{k}} \cos\theta$$

by def. (3)

$$\omega = \frac{d\theta}{dt}$$

Isotropic Harmonic Oscillator *phase dynamics* in uniform-body

Unit 1
Fig. 9.10



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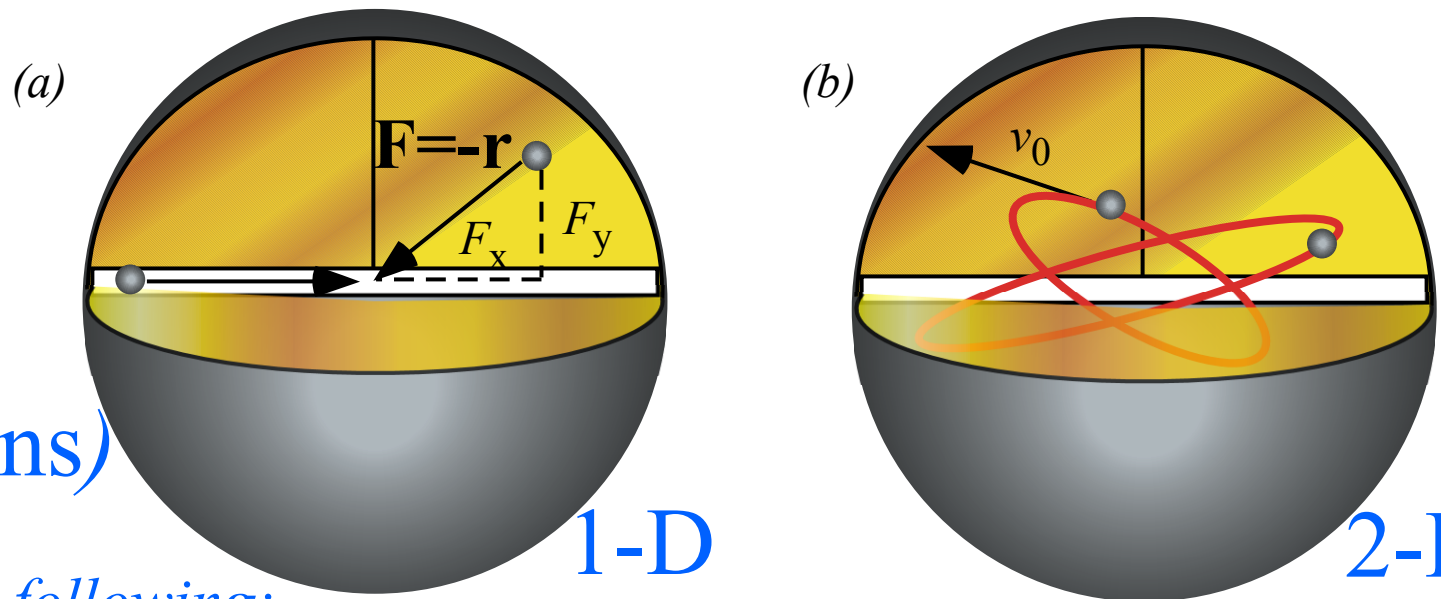
$$\sqrt{\frac{2E}{m}} \cos\theta \stackrel{\text{by (1)}}{=} v = \frac{dx}{dt} = \frac{d\theta}{dt} \frac{dx}{d\theta} = \omega \frac{dx}{d\theta} \stackrel{\text{by (2)}}{=} \omega \sqrt{\frac{2E}{k}} \cos\theta$$

$$\omega = \frac{d\theta}{dt} \stackrel{\text{by def. (3)}}{=} \frac{\sqrt{\frac{2E}{m}} \cos\theta}{\sqrt{\frac{2E}{k}} \cos\theta} \stackrel{\text{divide (1)}}{=} \sqrt{\frac{k}{m}}$$

by (2) derivative

Isotropic Harmonic Oscillator *phase dynamics in uniform-body*

Unit 1
Fig. 9.10



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$[y(t)$ and $v_y=v(t)]$ and $[z(t)$ and $v_z=v(t)]$ in the ideal isotropic case.

$$1 = \frac{mv^2}{2E} + \frac{kx^2}{2E} = \left(\frac{v}{\sqrt{2E/m}} \right)^2 + \left(\frac{x}{\sqrt{2E/k}} \right)^2$$

$$1 = \frac{mv^2}{2E} + \frac{kx^2}{2E} = (\cos\theta)^2 + (\sin\theta)^2$$

Another example of the old “scale-a-circle” trick...

Let : (1) $v = \sqrt{2E/m} \cos\theta$, and : (2) $x = \sqrt{2E/k} \sin\theta$ def. (3) $\omega = \frac{d\theta}{dt}$

$$\sqrt{\frac{2E}{m}} \cos\theta \stackrel{\text{by (1)}}{=} v = \frac{dx}{dt} = \frac{d\theta}{dt} \frac{dx}{d\theta} = \omega \frac{dx}{d\theta} \stackrel{\text{by (2)}}{=} \omega \sqrt{\frac{2E}{k}} \cos\theta$$

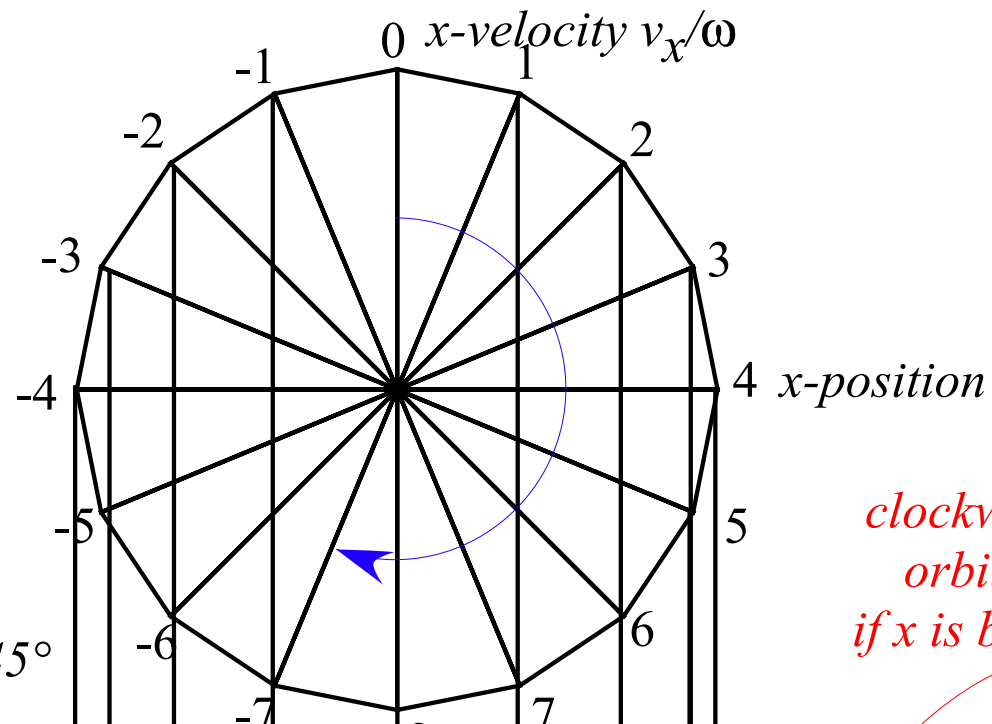
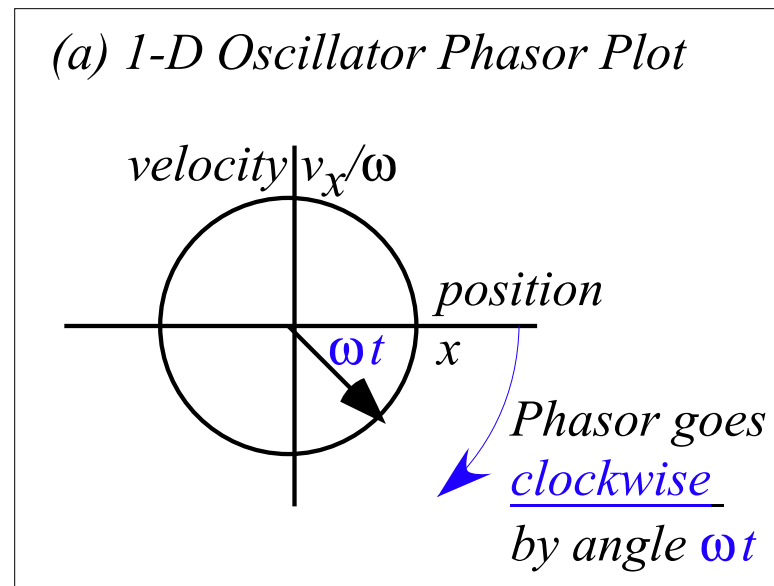
$$\omega = \frac{d\theta}{dt} = \frac{\sqrt{\frac{2E}{m}} \cos\theta}{\sqrt{\frac{2E}{k}} \cos\theta} \stackrel{\text{divide (1)}}{=} \sqrt{\frac{k}{m}}$$

by (2) derivative

→ *Review of IHO orbital phasor “clock” dynamics in uniform-body with two “movie” examples*

Review of IHO orbital phase dynamics in uniform-body

Unit 1
Fig. 9.10

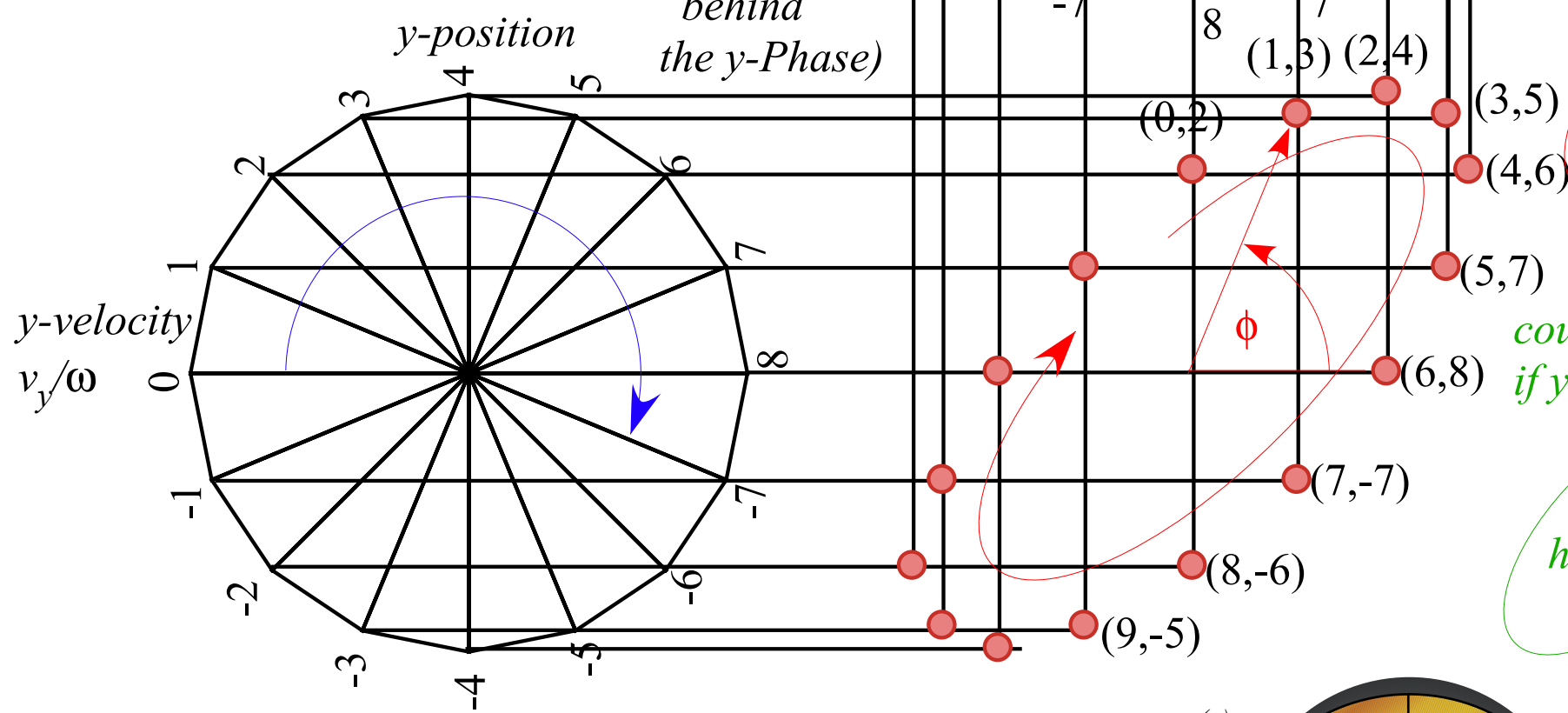


clockwise orbit if x is behind y

Left-handed

(b) 2-D Oscillator Phasor Plot

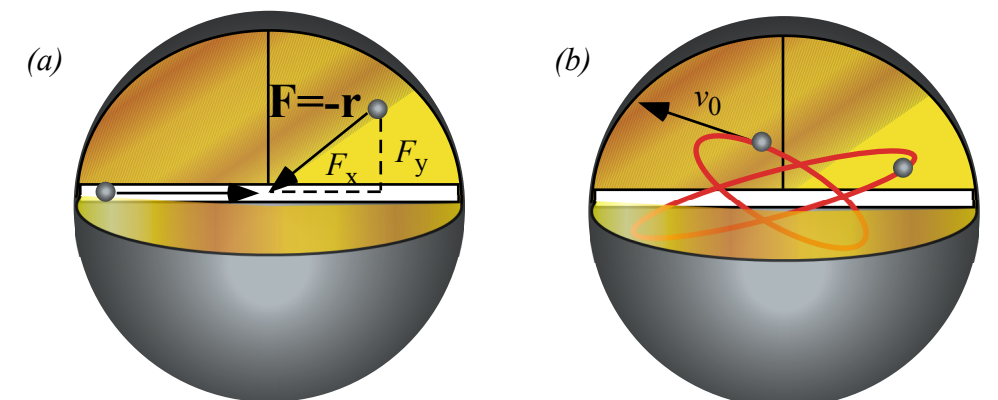
(x-Phase 45° behind the y-Phase)



counter-clockwise if y is behind x

Right-handed

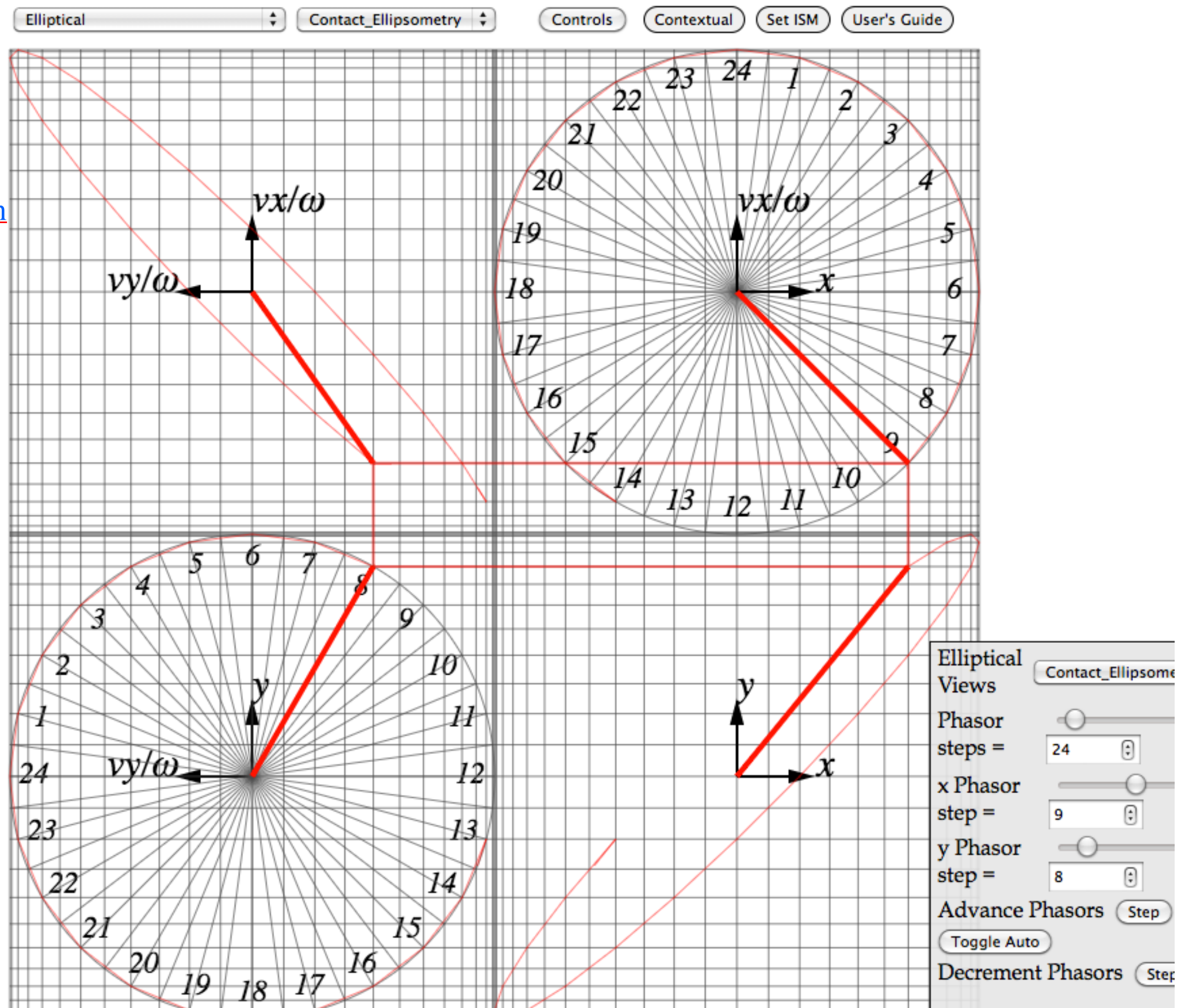
[RelaWavity web simulation - Contact ellipsometry](#)



[Introduction to Phasors at our Pirelli Relativity Site](#)

[BoxIt web simulation - With y-Phasor is on other side of xy plot](#)

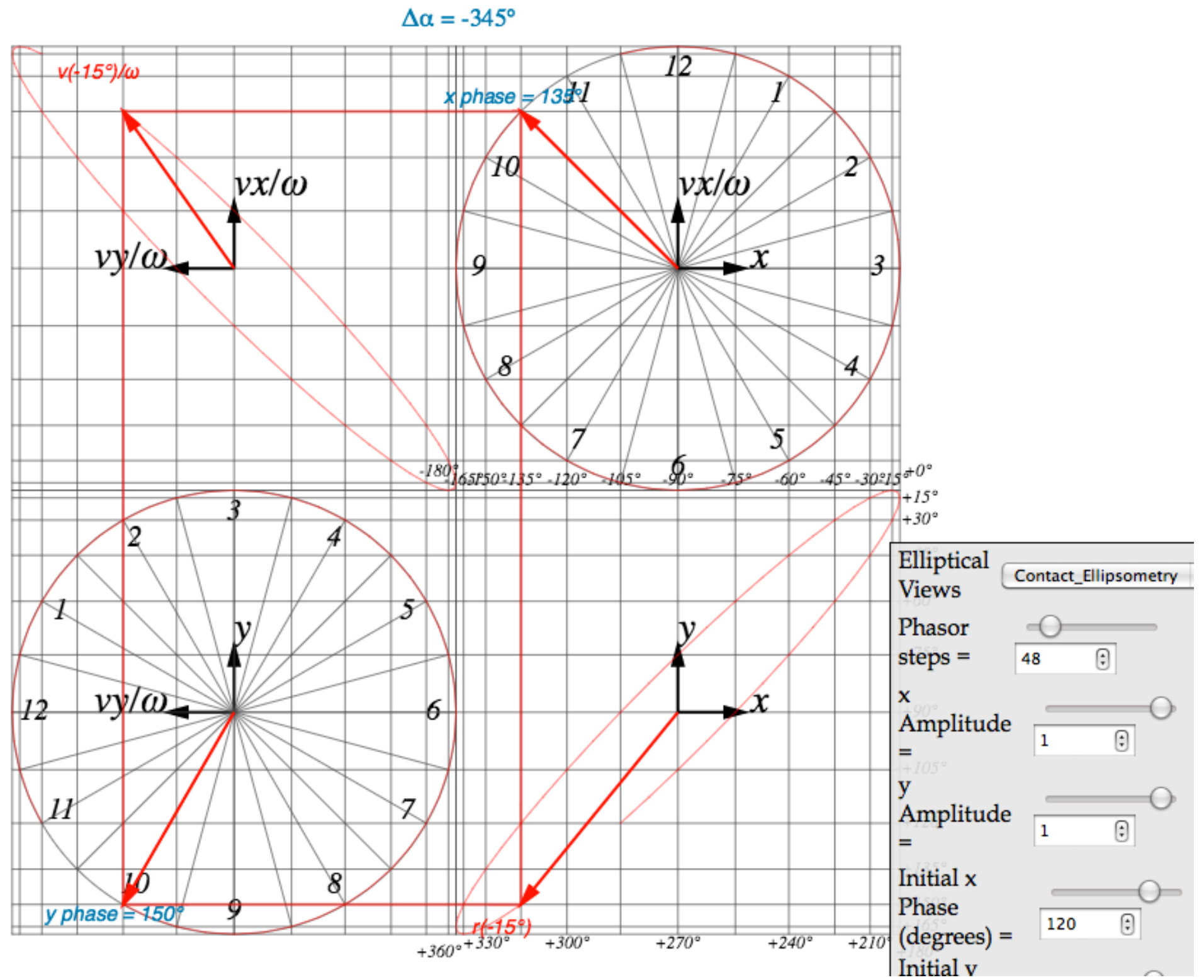
[RelaWavity Web Simulation](#)
[Ellipsometry](#)



Geometry of Kepler anomalies for vectors $[\mathbf{r}(\phi), \mathbf{v}(\phi)]$ in coordinate (x,y) space rendered by animation web-apps BoxIt and RelaWavity described below after p.70.

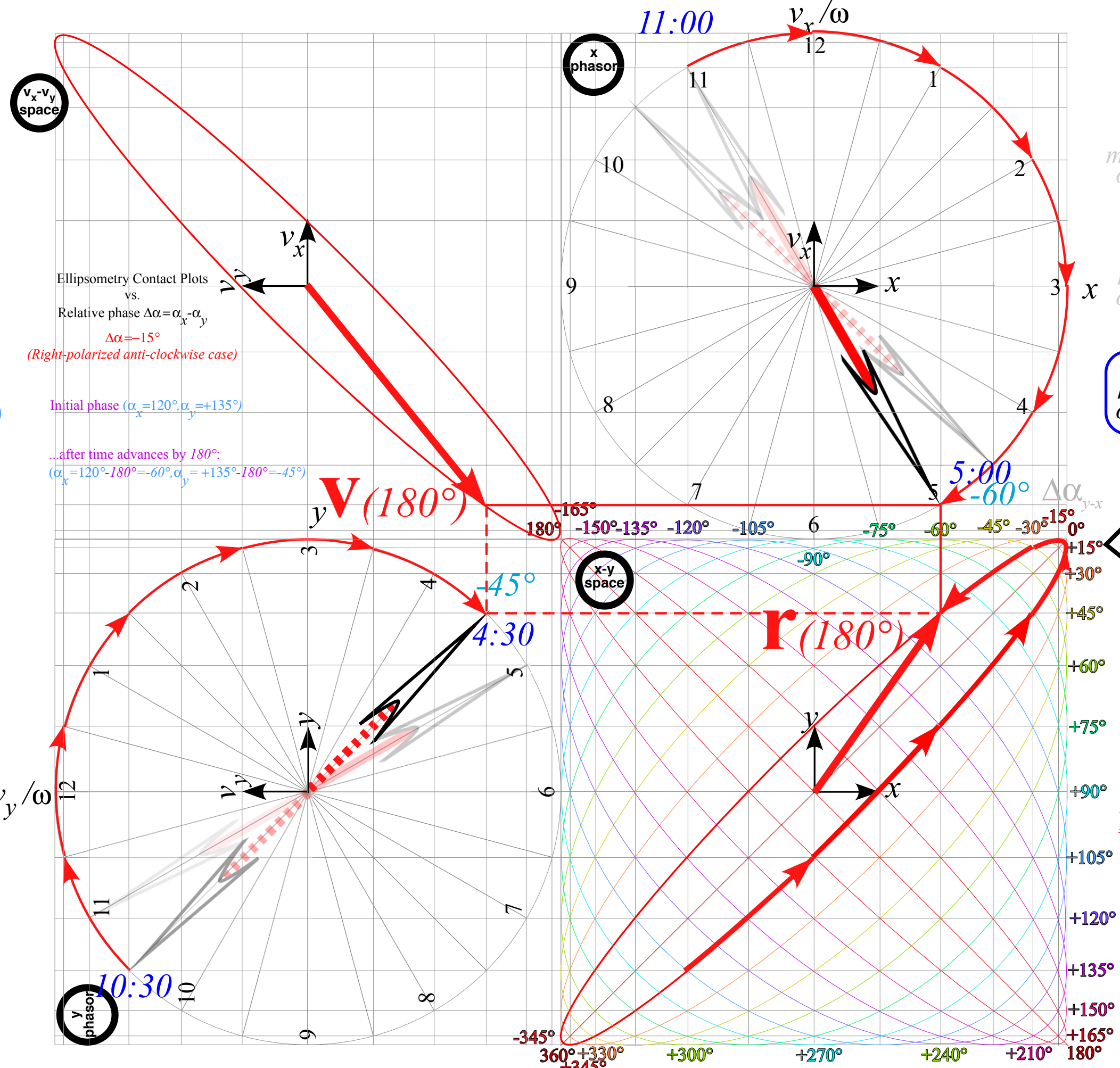
[RelaWavity web simulation - Contact ellipsometry](#) (User Mouse Input allowed for setting phasor values)

[RelaWavity Web Simulation](#)
[Ellipsometry](#)



Geometry of Kepler anomalies for vectors $[\mathbf{r}(\phi), \mathbf{v}(\phi)]$ in coordinate (x,y) space rendered by animation web-apps BoxIt and RelaWavity described below after p.7 and p.17.

[RelaWavity web simulation - Contact ellipsometry](#) (User Mouse Input allowed for setting phasor values)



v_x-v_y
space

x
phasor

11:00

v_x/ω

phase lag:
 $\Delta\alpha = \alpha_x - \alpha_y = 15^\circ$

1 minute orbit (2.5 seconds for second hand)

or
 1 hour orbit (2.5 minutes for minute hand)

12 hour orbit (1/2 hour for hour hand)

Ellipsometry Contact Plots vs. Relative phase $\Delta\alpha = \alpha_x - \alpha_y$
 $\Delta\alpha = -15^\circ$
 (Right-polarized anti-clockwise case)

Initial phase ($\alpha_x = 120^\circ, \alpha_y = +135^\circ$)

...after time advances by 180°:
 ($\alpha_x = 120^\circ - 180^\circ = -60^\circ, \alpha_y = +135^\circ - 180^\circ = -45^\circ$)

$V(180^\circ)$

5:00

$\Delta\alpha_{y-x}$

$\Delta\alpha = \alpha_x - \alpha_y = 15^\circ$

x-y
space

$r(180^\circ)$

v_y/ω

y
phasor

10:30

RelaWavity Web Simulation
 Ellipsometry

360° +330° +300° +270° +240° +210° 180°

*Constructing 2D IHO orbits using **Kepler anomaly plots***

 *Mean-anomaly and eccentric-anomaly geometry*

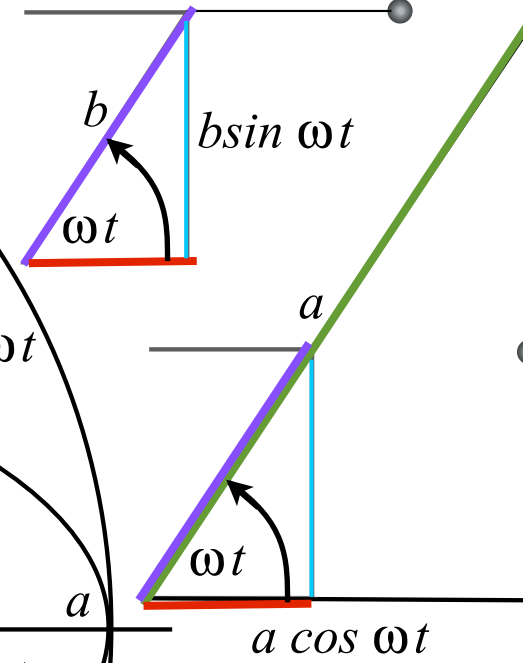
Calculus and vector geometry of IHO orbits

A confusing introduction to Coriolis-centrifugal force geometry (Derived better in Ch. 12)

Linear Harmonic
Force-Field
Orbits

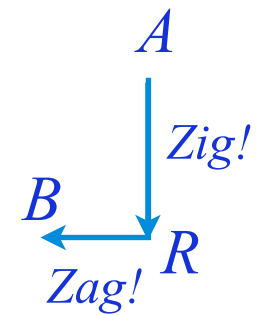
Kepler's
Mean Anomaly Line
(slope angle $\theta = \omega t$)

Kepler's
Eccentric Anomaly Line
(slope is polar angle $\phi = \text{atan}[y/x]$)



Unit 1
Fig. 11.1
(top 2/3's)

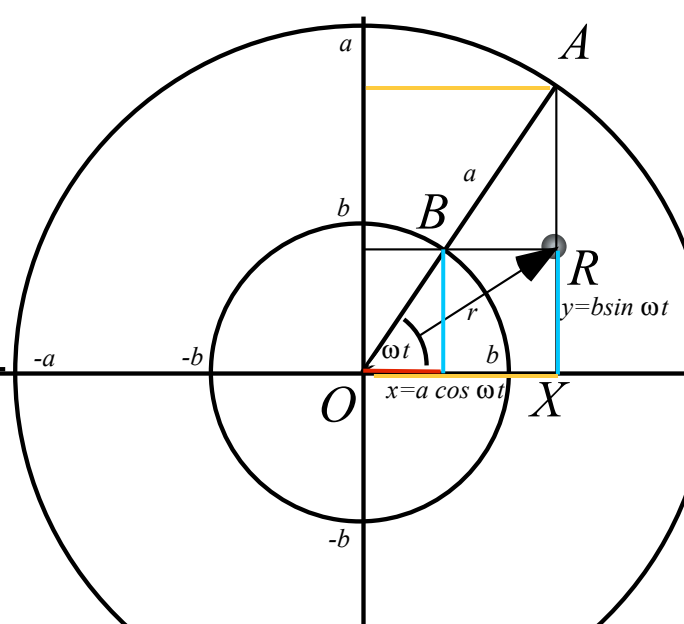
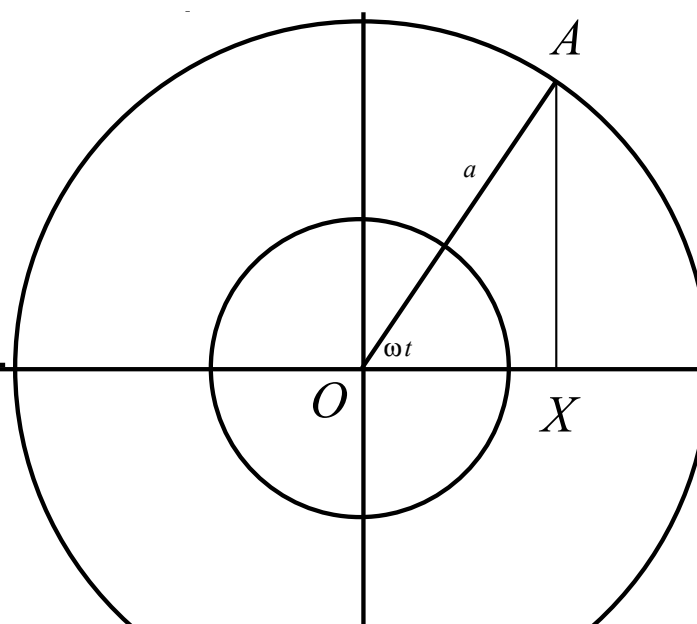
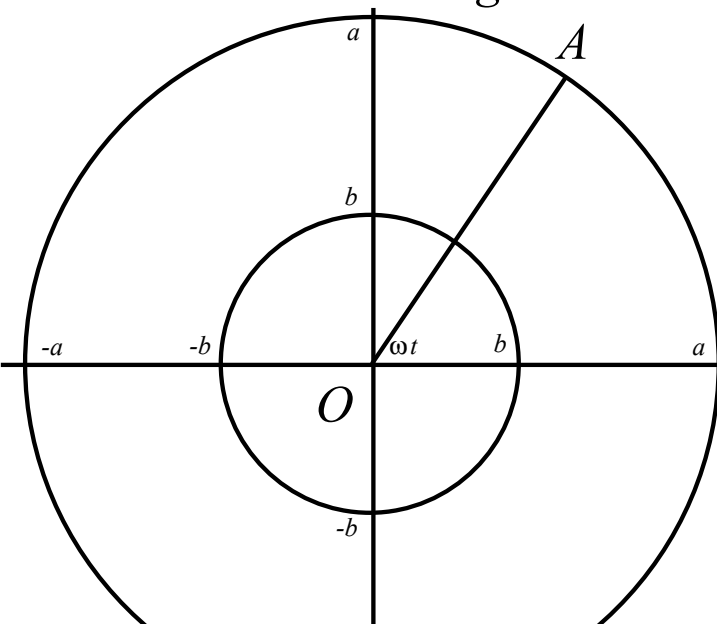
Another example
of a
Zig-Zag
construction:



Step 1. Draw concentric circles of radius a and b and a radius OA at angle ωt

Step 2. Draw vertical line AX from a -circle at ωt to x -axis

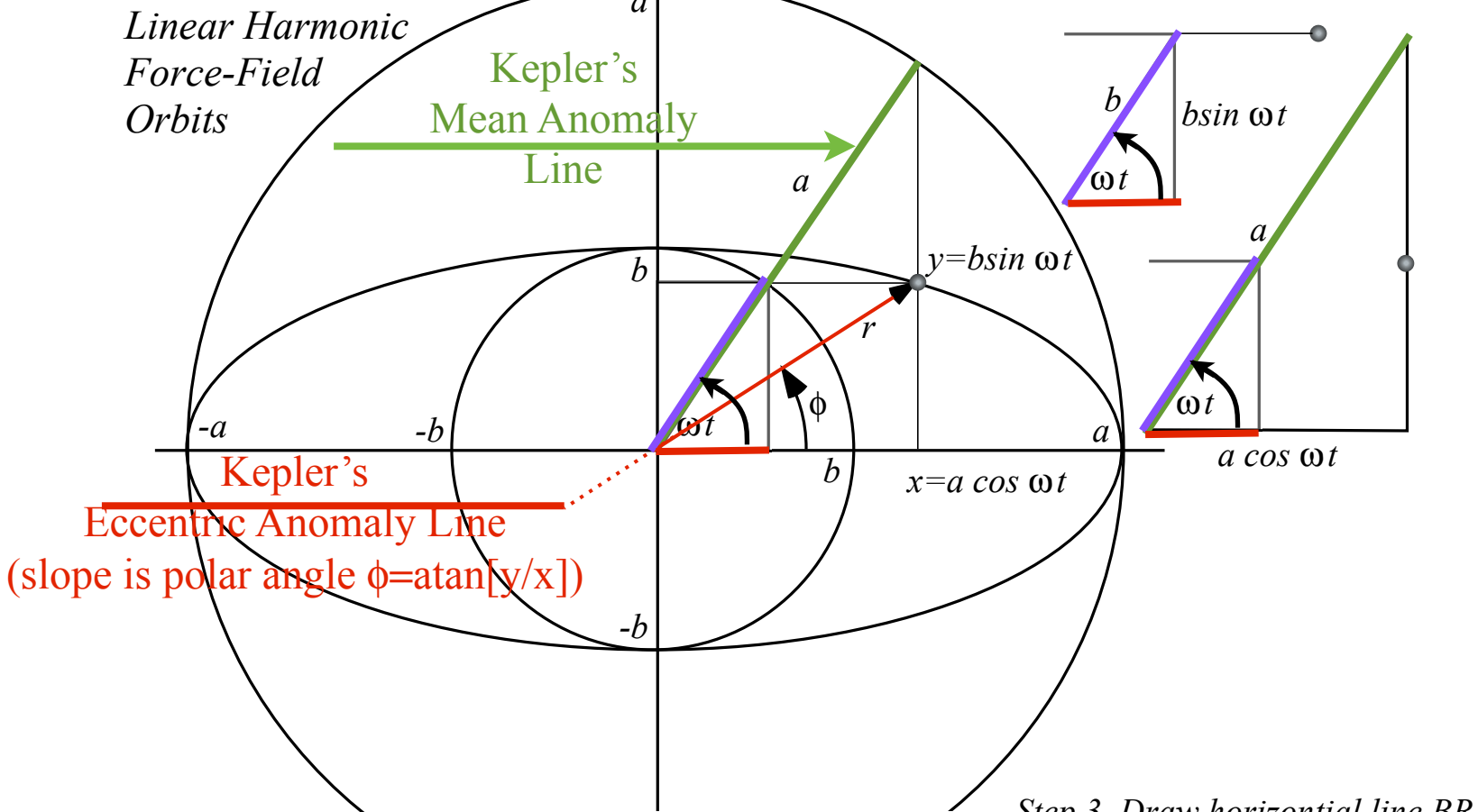
Step 3. Draw horizontal line BR from b -circle at ωt to line AX . Intersection is orbit point R .



Linear Harmonic
Force-Field
Orbits

Kepler's
Mean Anomaly
Line

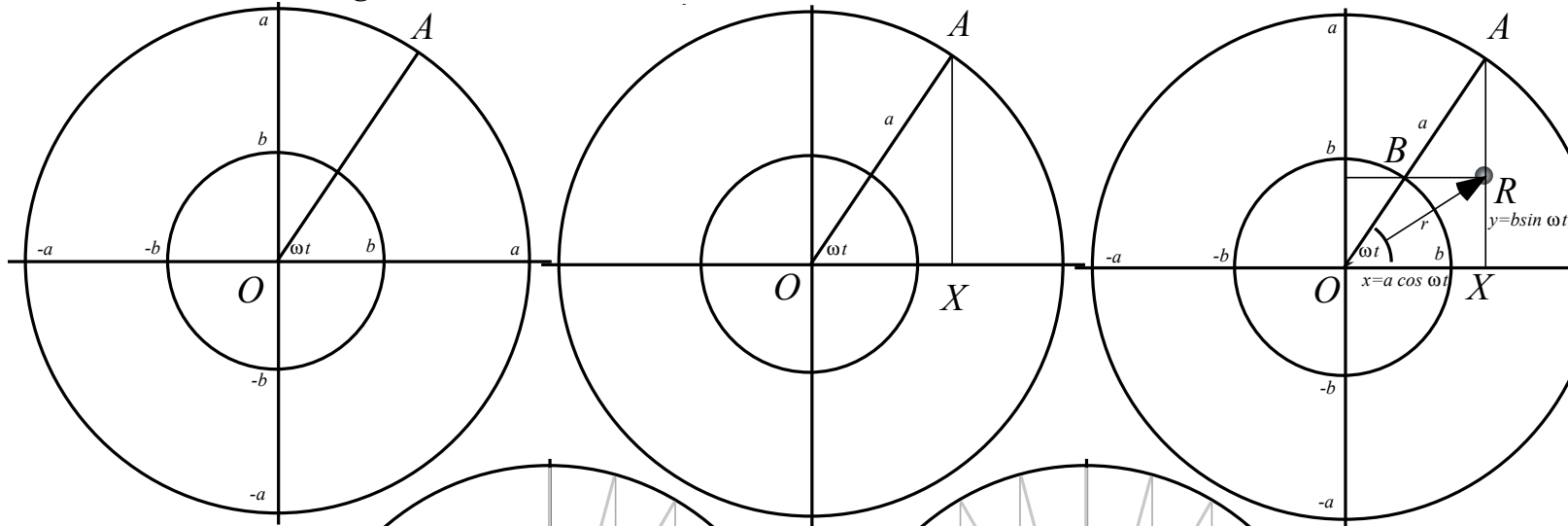
Kepler's
Eccentric Anomaly Line
(slope is polar angle $\phi = \text{atan}[y/x]$)



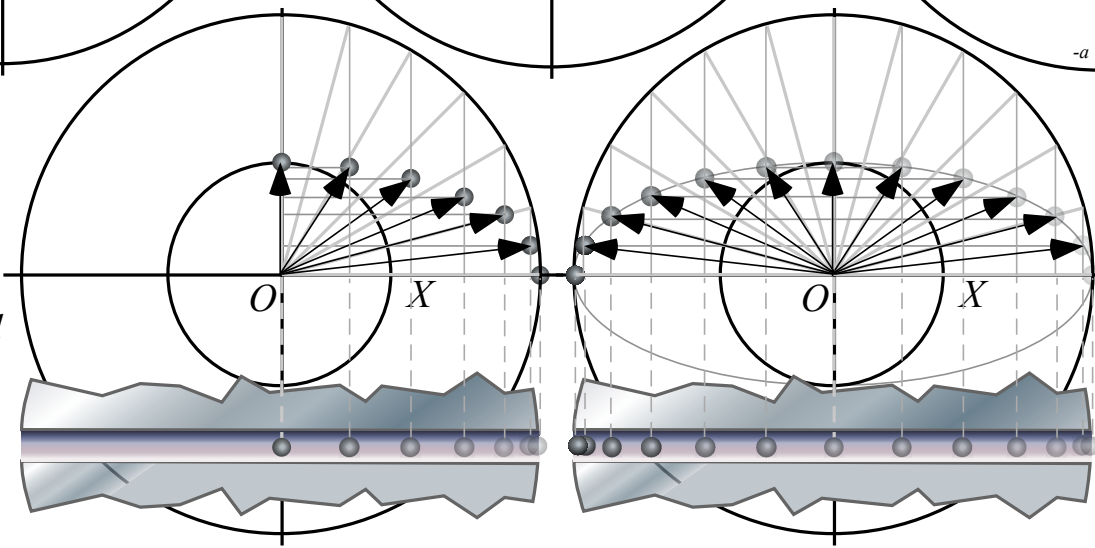
Step 1. Draw concentric circles of radius a and b and a radius OA at angle ωt

Step 2. Draw vertical line AX from a -circle at ωt to x -axis

Step 3. Draw horizontal line BR from b -circle at ωt to line AX . Intersection is orbit point R .



Step 4-N
Repeat
as often
as needed



Unit 1
Fig. 11.1

*Constructing 2D IHO orbits using **Kepler anomaly plots***

Mean-anomaly and eccentric-anomaly geometry

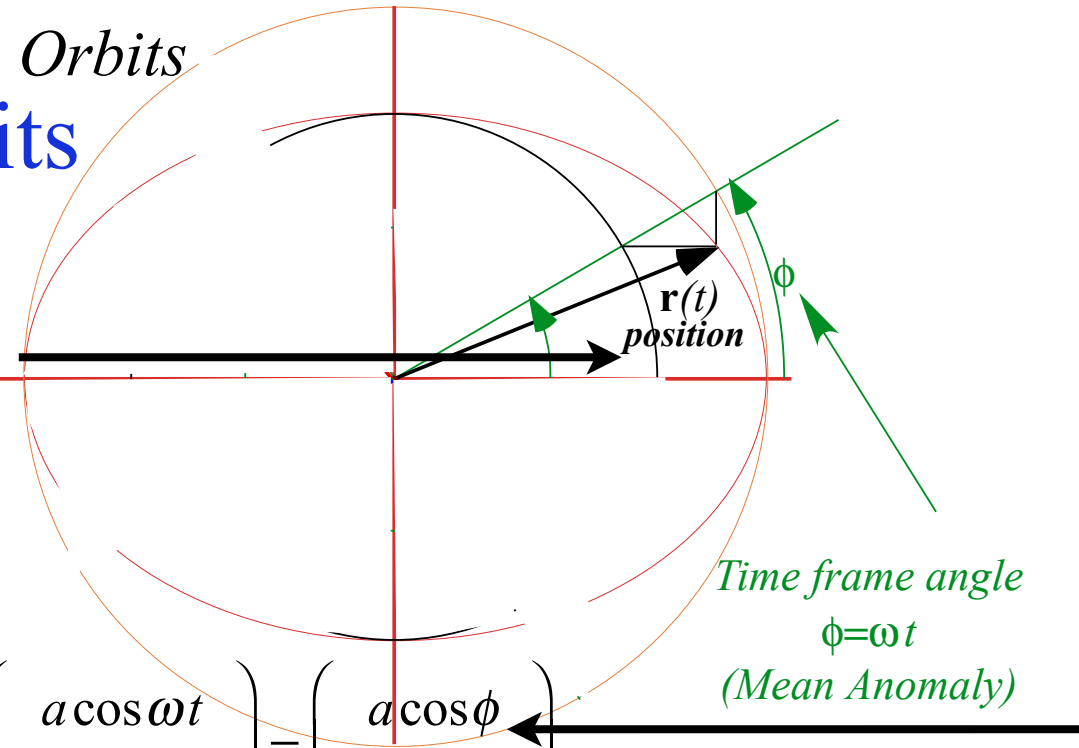
 *Calculus and vector geometry of IHO orbits*

A confusing introduction to Coriolis-centrifugal force geometry (Derived better in Ch. 12)

Calculus of IHO orbits

(a) Orbits

mean-anomaly ϕ of position vector \mathbf{r}



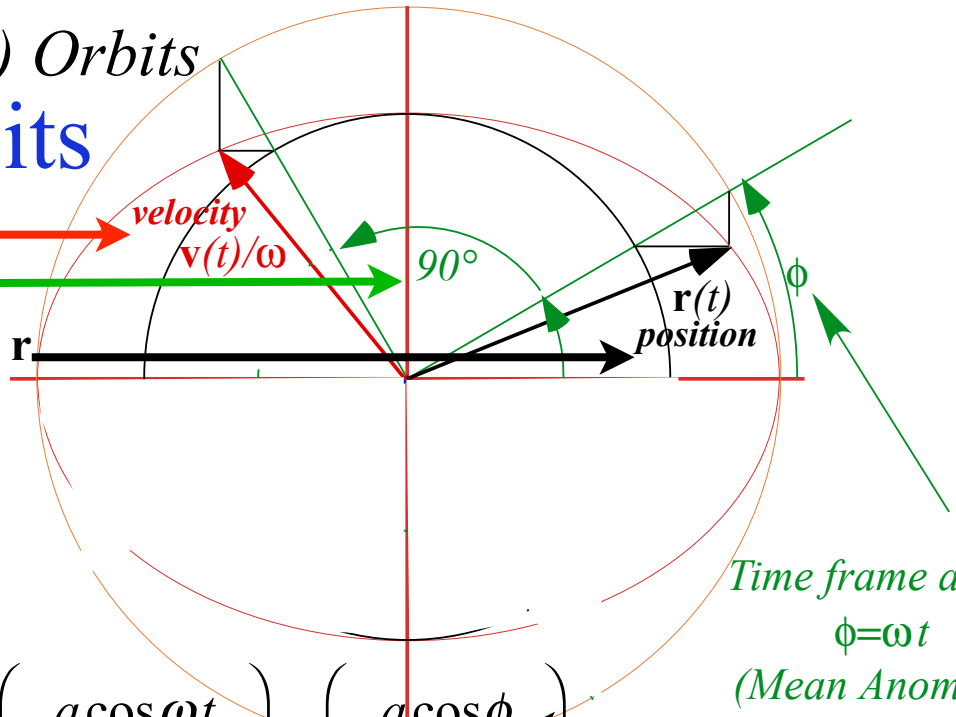
Time frame angle
 $\phi = \omega t$
 (Mean Anomaly)

$$\text{radius vector : } \mathbf{r} = \begin{pmatrix} r_x \\ r_y \end{pmatrix} = \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} a \cos \omega t \\ b \sin \omega t \end{pmatrix} = \begin{pmatrix} a \cos \phi \\ b \sin \phi \end{pmatrix}$$

Calculus of IHO orbits

(a) Orbits

To make velocity vector \mathbf{v} just rotate by $\pi/2$ or 90° the mean-anomaly ϕ of position vector \mathbf{r}



Time frame angle
 $\phi = \omega t$
 (Mean Anomaly)

$$\text{radius vector : } \mathbf{r} = \begin{pmatrix} r_x \\ r_y \end{pmatrix} = \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} a \cos \omega t \\ b \sin \omega t \end{pmatrix} = \begin{pmatrix} a \cos \phi \\ b \sin \phi \end{pmatrix}$$

mean-anomaly ϕ of position vector \mathbf{r} rotated by $\pi/2$ or 90° is *m.a.* of vector \mathbf{v}

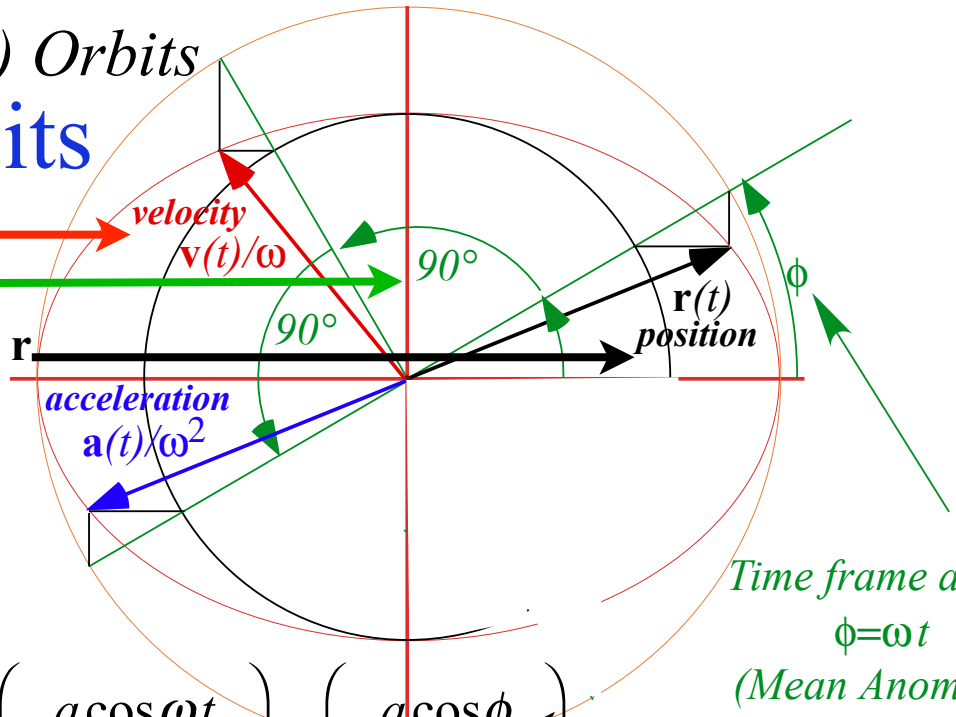
$$\text{velocity vector : } \mathbf{v} = \begin{pmatrix} v_x \\ v_y \end{pmatrix} = \begin{pmatrix} -a\omega \sin \omega t \\ b\omega \cos \omega t \end{pmatrix} = \frac{d\mathbf{r}}{dt} = \dot{\mathbf{r}} = \begin{pmatrix} a \cos \left(\phi + \frac{\pi}{2} \right) \\ b \sin \left(\phi + \frac{\pi}{2} \right) \end{pmatrix} \text{ (for } \omega = 1 \text{)}$$

Unit 1
 Fig. 11.5

Calculus of IHO orbits

(a) Orbits

To make velocity vector \mathbf{v} just rotate by $\pi/2$ or 90° the mean-anomaly ϕ of position vector \mathbf{r}



$$\text{radius vector : } \mathbf{r} = \begin{pmatrix} r_x \\ r_y \end{pmatrix} = \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} a \cos \omega t \\ b \sin \omega t \end{pmatrix} = \begin{pmatrix} a \cos \phi \\ b \sin \phi \end{pmatrix}$$

mean-anomaly ϕ of position vector \mathbf{r} rotated by $\pi/2$ or 90° is *m.a.* of vector \mathbf{v}

$$\text{velocity vector : } \mathbf{v} = \begin{pmatrix} v_x \\ v_y \end{pmatrix} = \begin{pmatrix} -a\omega \sin \omega t \\ b\omega \cos \omega t \end{pmatrix} = \frac{d\mathbf{r}}{dt} = \dot{\mathbf{r}} = \begin{pmatrix} a \cos\left(\phi + \frac{\pi}{2}\right) \\ b \sin\left(\phi + \frac{\pi}{2}\right) \end{pmatrix} \quad (\text{for } \omega = 1)$$

m.a. $\phi + \pi/2$ of vector \mathbf{v} rotated by another $\pi/2$ is *m.a.* of vector \mathbf{a}

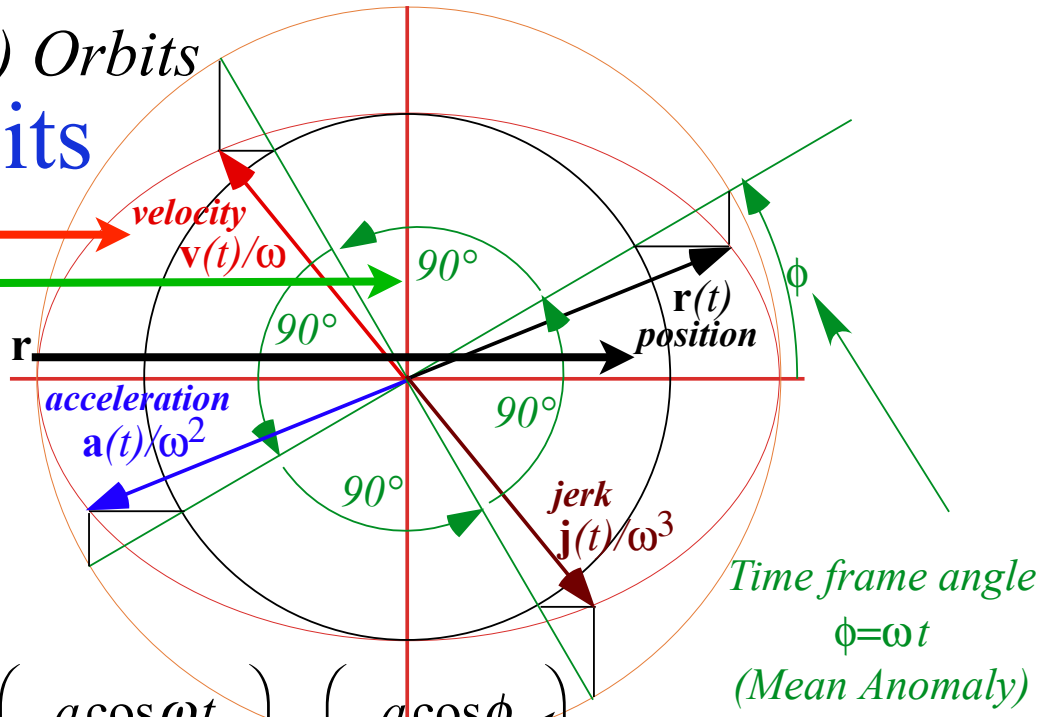
$$\text{acceleration or force vector : } \frac{\mathbf{F}}{m} = \mathbf{a} = \begin{pmatrix} a_x \\ a_y \end{pmatrix} = \begin{pmatrix} -a\omega^2 \cos \omega t \\ -b\omega^2 \sin \omega t \end{pmatrix} = \frac{d\mathbf{v}}{dt} = \dot{\mathbf{v}} = \ddot{\mathbf{r}} = \frac{d^2\mathbf{r}}{dt^2} = \begin{pmatrix} a \cos\left(\phi + \frac{2\pi}{2}\right) \\ b \sin\left(\phi + \frac{2\pi}{2}\right) \end{pmatrix}$$

Unit 1
Fig. 11.5

Calculus of IHO orbits

(a) Orbits

To make velocity vector \mathbf{v} just rotate by $\pi/2$ or 90° the mean-anomaly ϕ of position vector \mathbf{r}



$$\text{radius vector : } \mathbf{r} = \begin{pmatrix} r_x \\ r_y \end{pmatrix} = \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} a \cos \omega t \\ b \sin \omega t \end{pmatrix} = \begin{pmatrix} a \cos \phi \\ b \sin \phi \end{pmatrix}$$

mean-anomaly ϕ of position vector \mathbf{r} rotated by $\pi/2$ or 90° is *m.a.* of vector \mathbf{v}

Unit 1
Fig. 11.5

$$\text{velocity vector : } \mathbf{v} = \begin{pmatrix} v_x \\ v_y \end{pmatrix} = \begin{pmatrix} -a\omega \sin \omega t \\ b\omega \cos \omega t \end{pmatrix} = \frac{d\mathbf{r}}{dt} = \dot{\mathbf{r}} = \begin{pmatrix} a \cos\left(\phi + \frac{\pi}{2}\right) \\ b \sin\left(\phi + \frac{\pi}{2}\right) \end{pmatrix} \quad (\text{for } \omega = 1)$$

m.a. $\phi + \pi/2$ of vector \mathbf{v} rotated by another $\pi/2$ is *m.a.* of vector \mathbf{a}

$$\text{acceleration or force vector : } \frac{\mathbf{F}}{m} = \mathbf{a} = \begin{pmatrix} a_x \\ a_y \end{pmatrix} = \begin{pmatrix} -a\omega^2 \cos \omega t \\ -b\omega^2 \sin \omega t \end{pmatrix} = \frac{d\mathbf{v}}{dt} = \dot{\mathbf{v}} = \ddot{\mathbf{r}} = \frac{d^2\mathbf{r}}{dt^2} = \begin{pmatrix} a \cos\left(\phi + \frac{2\pi}{2}\right) \\ b \sin\left(\phi + \frac{2\pi}{2}\right) \end{pmatrix}$$

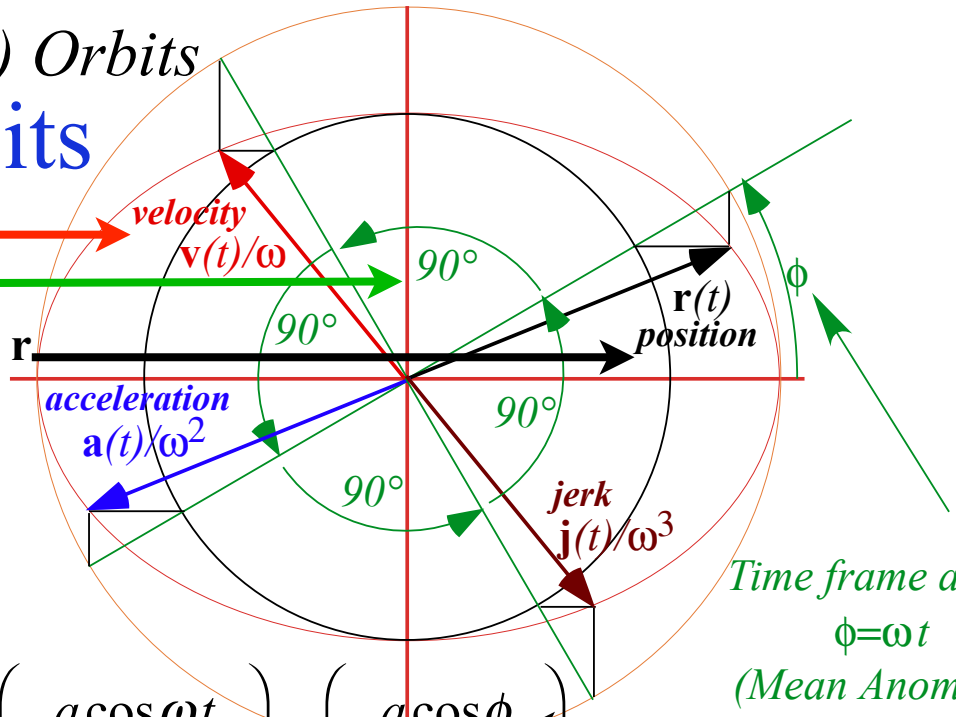
$$\text{jerk or change of acceleration : } \mathbf{j} = \begin{pmatrix} j_x \\ j_y \end{pmatrix} = \begin{pmatrix} +a\omega^3 \sin \omega t \\ -b\omega^3 \cos \omega t \end{pmatrix} = \frac{d\mathbf{a}}{dt} = \dot{\mathbf{a}} = \ddot{\mathbf{v}} = \ddot{\mathbf{r}} = \frac{d^3\mathbf{r}}{dt^3} = \begin{pmatrix} a \cos\left(\phi + \frac{3\pi}{2}\right) \\ b \sin\left(\phi + \frac{3\pi}{2}\right) \end{pmatrix}$$

...and so forth...

Calculus of IHO orbits

(a) Orbits

To make velocity vector \mathbf{v} just rotate by $\pi/2$ or 90° the mean-anomaly ϕ of position vector \mathbf{r}



$$\text{radius vector : } \mathbf{r} = \begin{pmatrix} r_x \\ r_y \end{pmatrix} = \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} a \cos \omega t \\ b \sin \omega t \end{pmatrix} = \begin{pmatrix} a \cos \phi \\ b \sin \phi \end{pmatrix}$$

mean-anomaly ϕ of position vector \mathbf{r} rotated by $\pi/2$ or 90° is *m.a.* of vector \mathbf{v}

Unit 1
Fig. 11.5

$$\text{velocity vector : } \mathbf{v} = \begin{pmatrix} v_x \\ v_y \end{pmatrix} = \begin{pmatrix} -a\omega \sin \omega t \\ b\omega \cos \omega t \end{pmatrix} = \frac{d\mathbf{r}}{dt} = \dot{\mathbf{r}} = \begin{pmatrix} a \cos\left(\phi + \frac{\pi}{2}\right) \\ b \sin\left(\phi + \frac{\pi}{2}\right) \end{pmatrix} \quad (\text{for } \omega = 1)$$

m.a. $\phi + \pi/2$ of vector \mathbf{v} rotated by another $\pi/2$ is *m.a.* of vector \mathbf{a}

$$\text{acceleration or force vector : } \frac{\mathbf{F}}{m} = \mathbf{a} = \begin{pmatrix} a_x \\ a_y \end{pmatrix} = \begin{pmatrix} -a\omega^2 \cos \omega t \\ -b\omega^2 \sin \omega t \end{pmatrix} = \frac{d\mathbf{v}}{dt} = \dot{\mathbf{v}} = \ddot{\mathbf{r}} = \frac{d^2\mathbf{r}}{dt^2} = \begin{pmatrix} a \cos\left(\phi + \frac{2\pi}{2}\right) \\ b \sin\left(\phi + \frac{2\pi}{2}\right) \end{pmatrix}$$

...and so forth...

$$\text{jerk or change of acceleration : } \mathbf{j} = \begin{pmatrix} j_x \\ j_y \end{pmatrix} = \begin{pmatrix} +a\omega^3 \sin \omega t \\ -b\omega^3 \cos \omega t \end{pmatrix} = \frac{d\mathbf{a}}{dt} = \dot{\mathbf{a}} = \ddot{\mathbf{v}} = \ddot{\mathbf{r}} = \frac{d^3\mathbf{r}}{dt^3} = \begin{pmatrix} a \cos\left(\phi + \frac{3\pi}{2}\right) \\ b \sin\left(\phi + \frac{3\pi}{2}\right) \end{pmatrix}$$

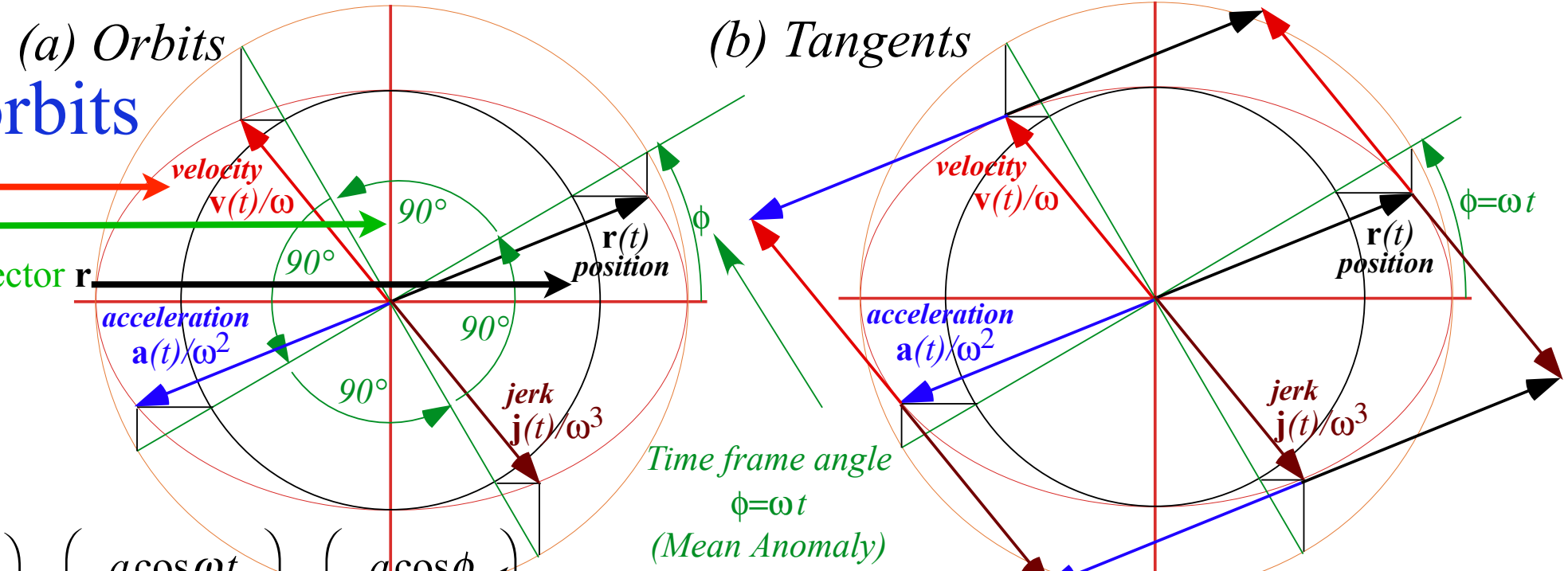
...and so on...
...But, now it repeats after 4 *t*-derivatives

$$\text{inauguration or change of jerk : } \mathbf{i} = \begin{pmatrix} i_x \\ i_y \end{pmatrix} = \begin{pmatrix} +a\omega^4 \cos \omega t \\ +b\omega^4 \sin \omega t \end{pmatrix} = \frac{d\mathbf{j}}{dt} = \dot{\mathbf{j}} = \ddot{\mathbf{a}} = \ddot{\mathbf{v}} = \ddot{\mathbf{r}} = \frac{d^4\mathbf{r}}{dt^4} = \begin{pmatrix} a \cos\left(\phi + \frac{4\pi}{2}\right) \\ b \sin\left(\phi + \frac{4\pi}{2}\right) \end{pmatrix}$$

Calculus of IHO orbits

To make velocity vector \mathbf{v} just rotate by $\pi/2$ or 90° the mean-anomaly ϕ of position vector \mathbf{r}

[Link \$\Rightarrow\$ BoxIt simulation of IHO orbits](#)



Unit 1
Fig. 11.5

$$\text{radius vector : } \mathbf{r} = \begin{pmatrix} r_x \\ r_y \end{pmatrix} = \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} a \cos \omega t \\ b \sin \omega t \end{pmatrix} = \begin{pmatrix} a \cos \phi \\ b \sin \phi \end{pmatrix}$$

mean-anomaly ϕ of position vector \mathbf{r} rotated by $\pi/2$ or 90° is *m.a.* of vector \mathbf{v}

$$\text{velocity vector : } \mathbf{v} = \begin{pmatrix} v_x \\ v_y \end{pmatrix} = \begin{pmatrix} -a\omega \sin \omega t \\ b\omega \cos \omega t \end{pmatrix} = \frac{d\mathbf{r}}{dt} = \dot{\mathbf{r}} = \begin{pmatrix} a \cos\left(\phi + \frac{\pi}{2}\right) \\ b \sin\left(\phi + \frac{\pi}{2}\right) \end{pmatrix} \quad (\text{for } \omega = 1)$$

[Link \$\rightarrow\$ IHO Exegesis Plot](#)
[Link \$\rightarrow\$ IHO orbital time rates of change](#)

m.a. $\phi + \pi/2$ of vector \mathbf{v} rotated by another $\pi/2$ is *m.a.* of vector \mathbf{a}

$$\text{acceleration or force vector : } \frac{\mathbf{F}}{m} = \mathbf{a} = \begin{pmatrix} a_x \\ a_y \end{pmatrix} = \begin{pmatrix} -a\omega^2 \cos \omega t \\ -b\omega^2 \sin \omega t \end{pmatrix} = \frac{d\mathbf{v}}{dt} = \dot{\mathbf{v}} = \ddot{\mathbf{r}} = \frac{d^2\mathbf{r}}{dt^2} = \begin{pmatrix} a \cos\left(\phi + \frac{2\pi}{2}\right) \\ b \sin\left(\phi + \frac{2\pi}{2}\right) \end{pmatrix}$$

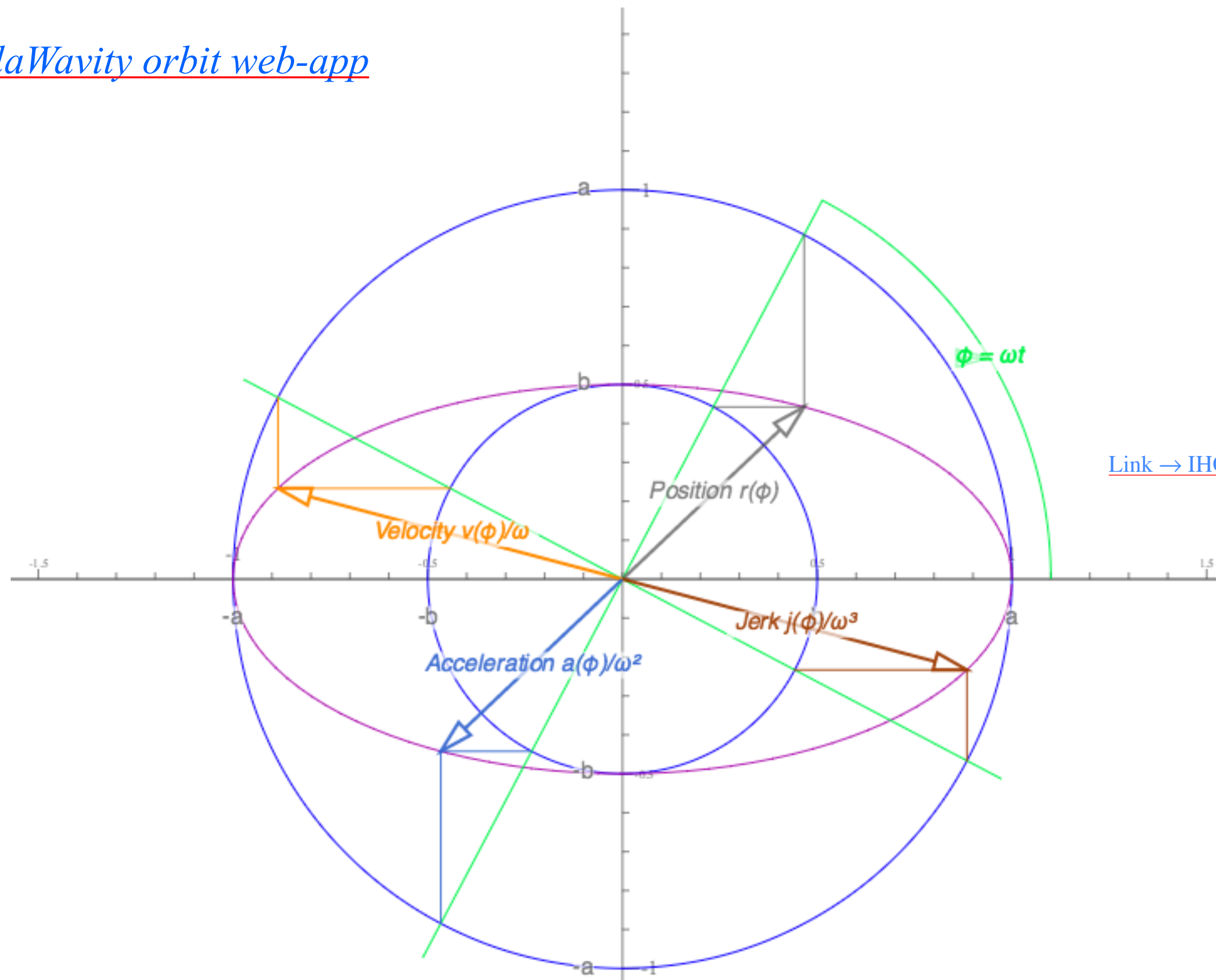
...and so forth...

$$\text{jerk or change of acceleration : } \mathbf{j} = \begin{pmatrix} j_x \\ j_y \end{pmatrix} = \begin{pmatrix} +a\omega^3 \sin \omega t \\ -b\omega^3 \cos \omega t \end{pmatrix} = \frac{d\mathbf{a}}{dt} = \dot{\mathbf{a}} = \ddot{\mathbf{v}} = \ddot{\mathbf{r}} = \frac{d^3\mathbf{r}}{dt^3} = \begin{pmatrix} a \cos\left(\phi + \frac{3\pi}{2}\right) \\ b \sin\left(\phi + \frac{3\pi}{2}\right) \end{pmatrix}$$

...and so on...
...But, now it repeats after 4 t -derivatives

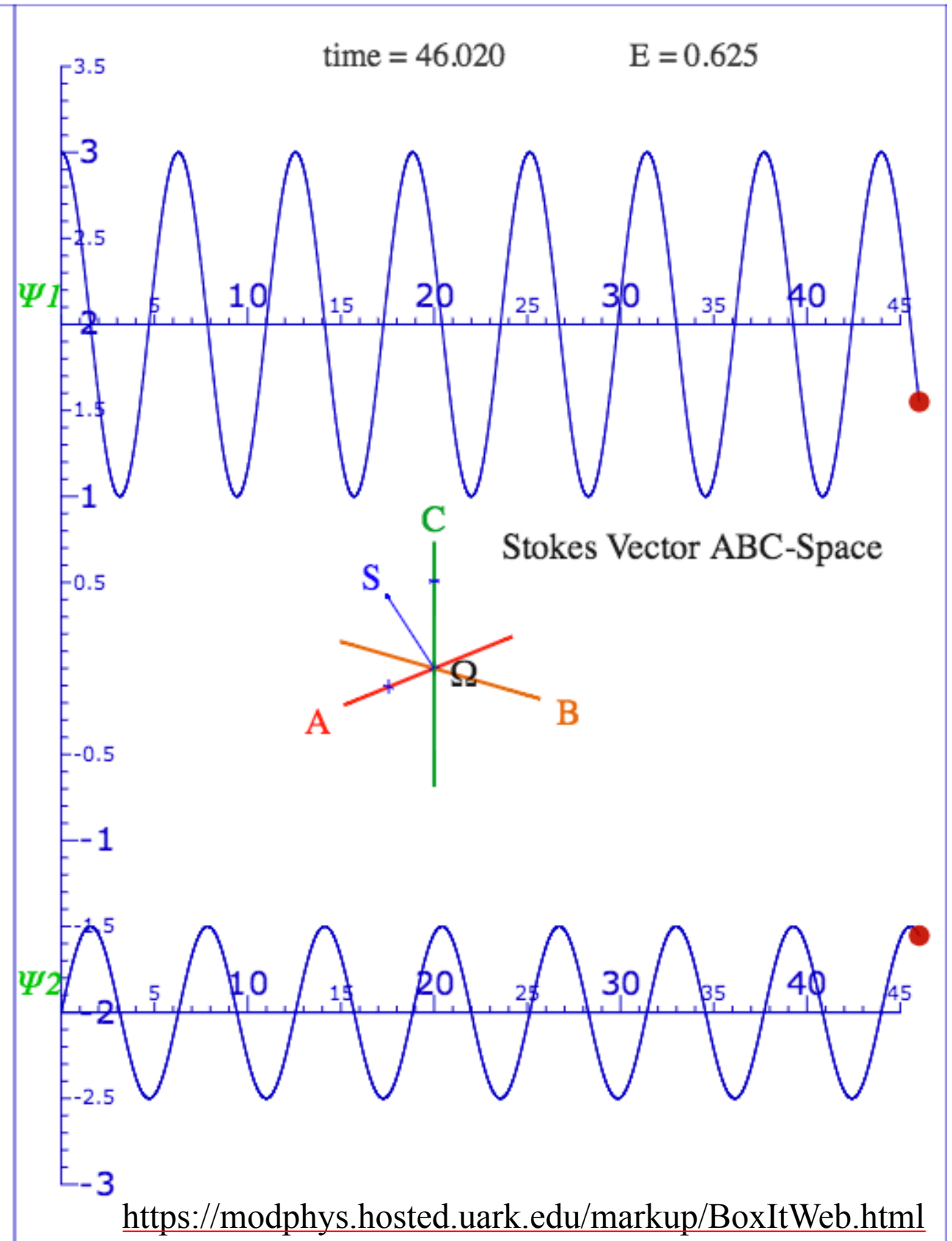
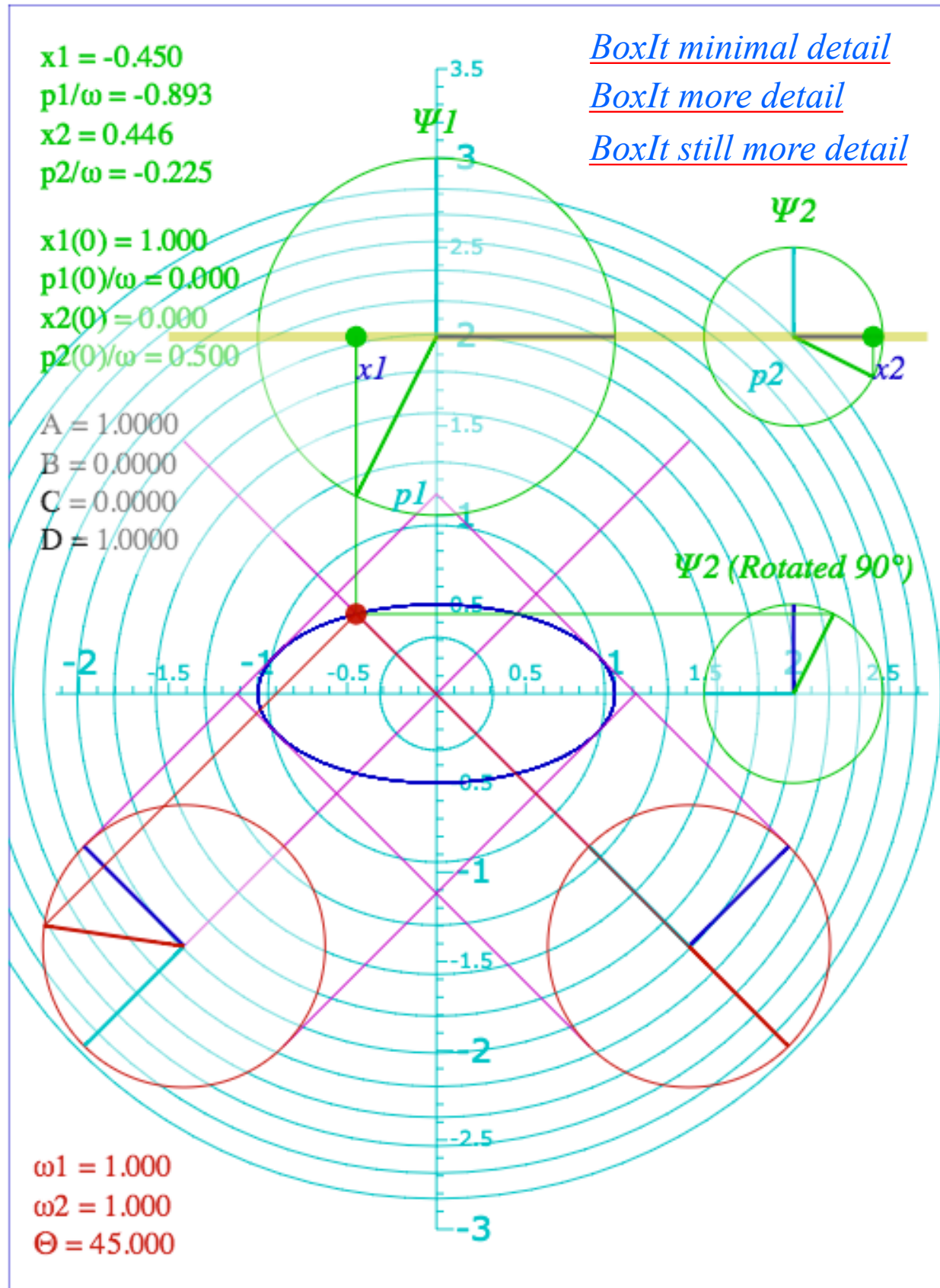
$$\text{inauguration or change of jerk : } \mathbf{i} = \begin{pmatrix} i_x \\ i_y \end{pmatrix} = \begin{pmatrix} +a\omega^4 \cos \omega t \\ +b\omega^4 \sin \omega t \end{pmatrix} = \frac{d\mathbf{j}}{dt} = \dot{\mathbf{j}} = \ddot{\mathbf{a}} = \ddot{\mathbf{v}} = \ddot{\mathbf{r}} = \frac{d^4\mathbf{r}}{dt^4} = \begin{pmatrix} a \cos\left(\phi + \frac{4\pi}{2}\right) \\ b \sin\left(\phi + \frac{4\pi}{2}\right) \end{pmatrix}$$

[RelaWavity orbit web-app](#)

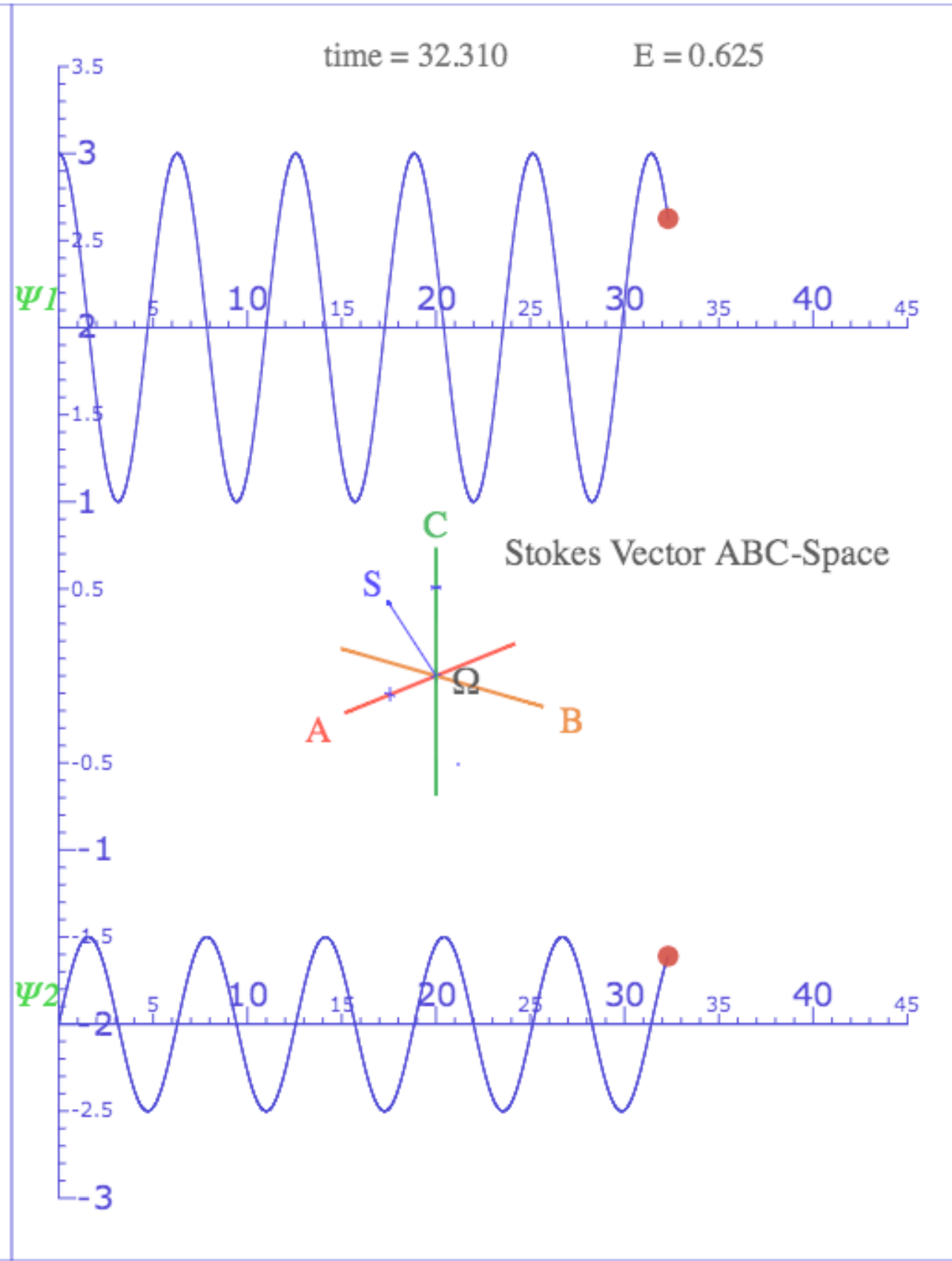
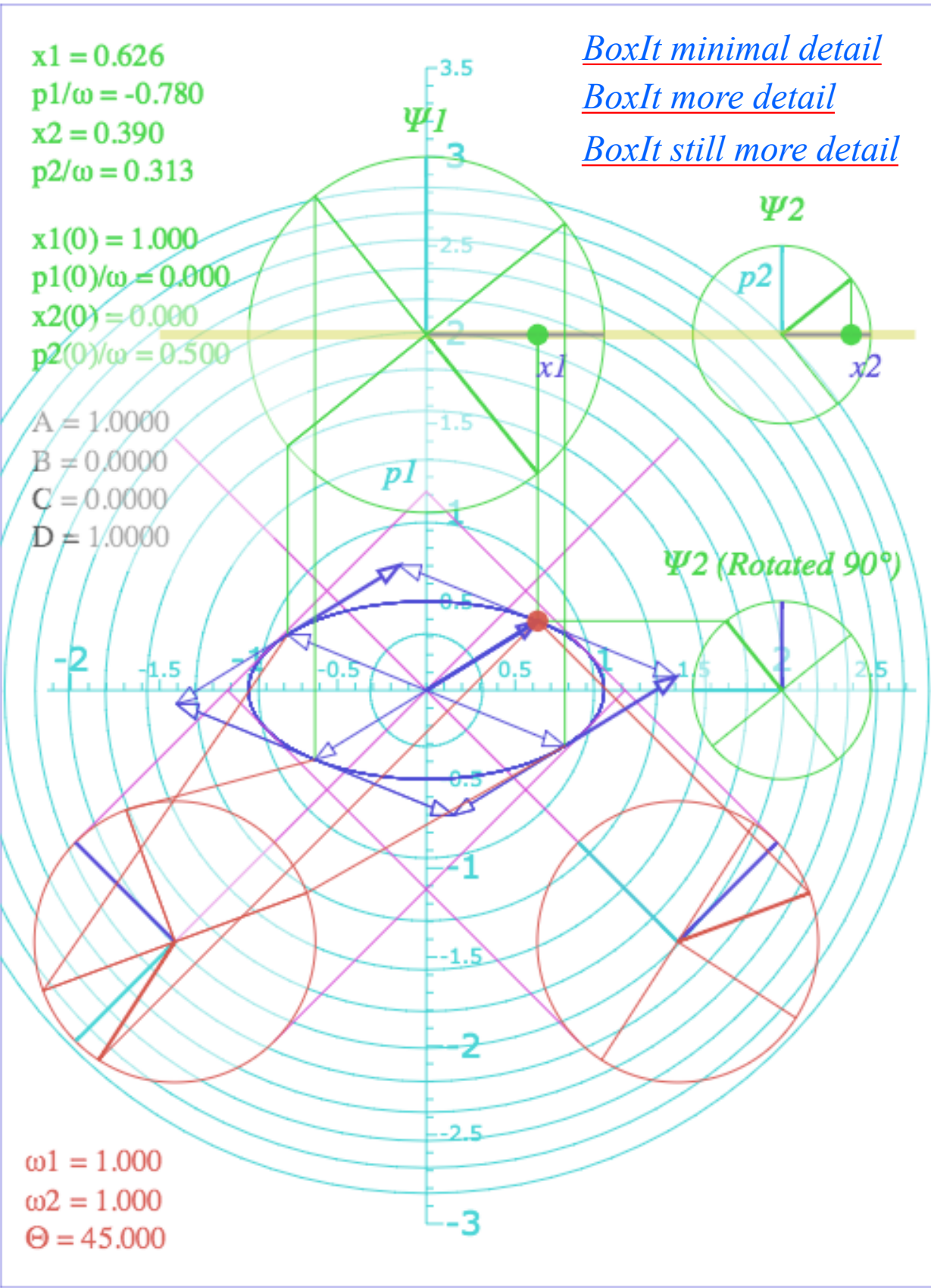


[Link \$\rightarrow\$ IHO orbital time rates of change](#)

Geometry of Kepler anomalies for vectors $[\mathbf{r}(\phi), \mathbf{v}(\phi), \mathbf{a}(\phi), \mathbf{j}(\phi),]$ in coordinate (x,y) space rendered by animation web-apps BoxIt and RelaWavity.



Geometry of Kepler anomalies for vectors $[\mathbf{r}(\phi)]$ in coordinate (x,y) space and 2-particle (x_1,x_2) space rendered by animation web-apps BoxIt.

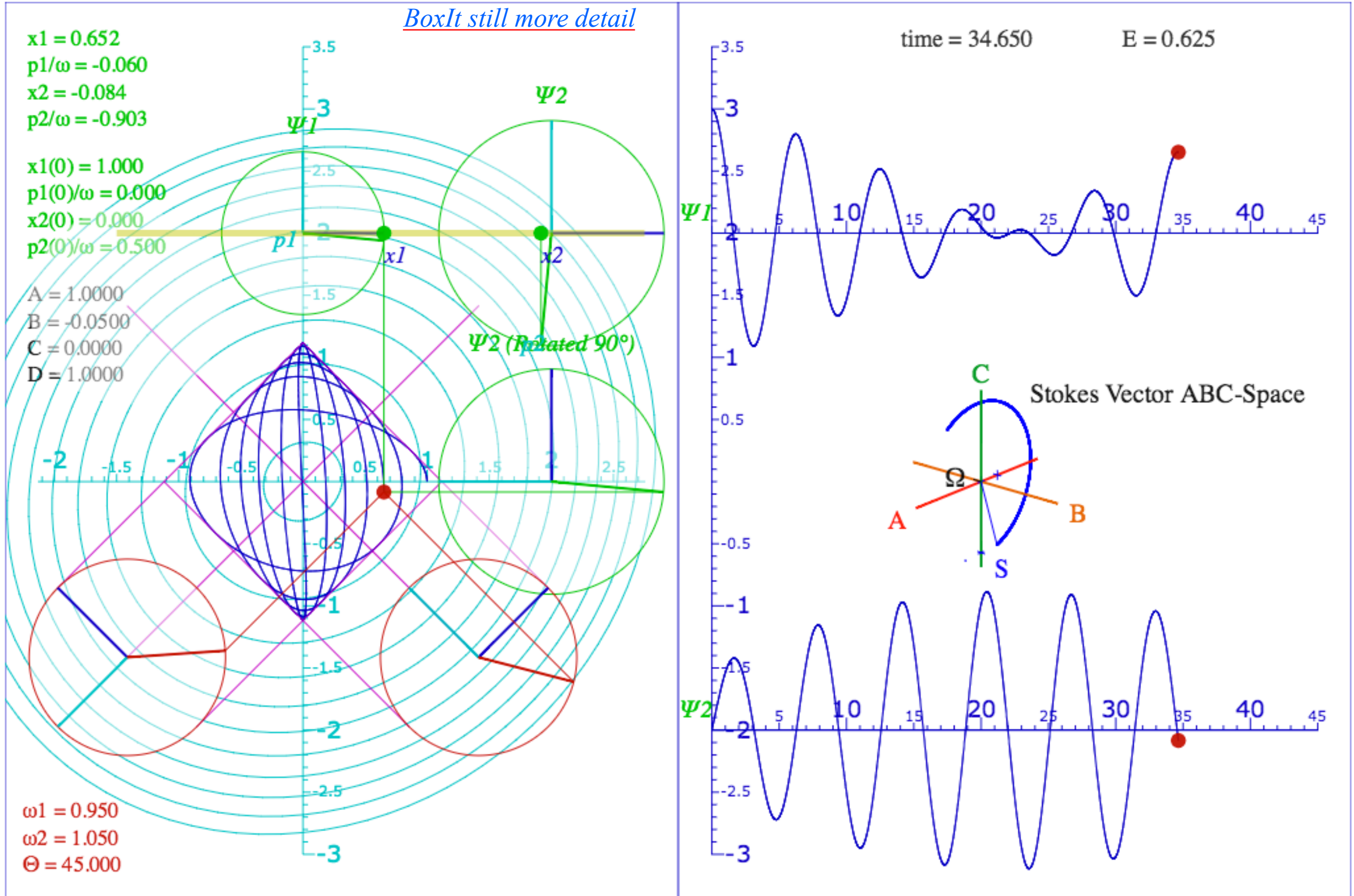


Geometry of Kepler anomalies for vectors $[\mathbf{r}(\phi), \mathbf{v}(\phi), \mathbf{a}(\phi), \mathbf{j}(\phi),]$ in coordinate (x,y) space and 2-particle (x_1,x_2) space rendered by animation web-apps BoxIt.

[BoxIt minimal detail](#)

[BoxIt more detail](#)

[BoxIt still more detail](#)



Geometry of vectors $[\mathbf{r}(\phi), \mathbf{p}(\phi)]$ and quantum spin S -space and 2-particle (x_1, x_2) space rendered by animation web-apps BoxIt.

[BoxIt Web Simulation - B-Type Motion](#)

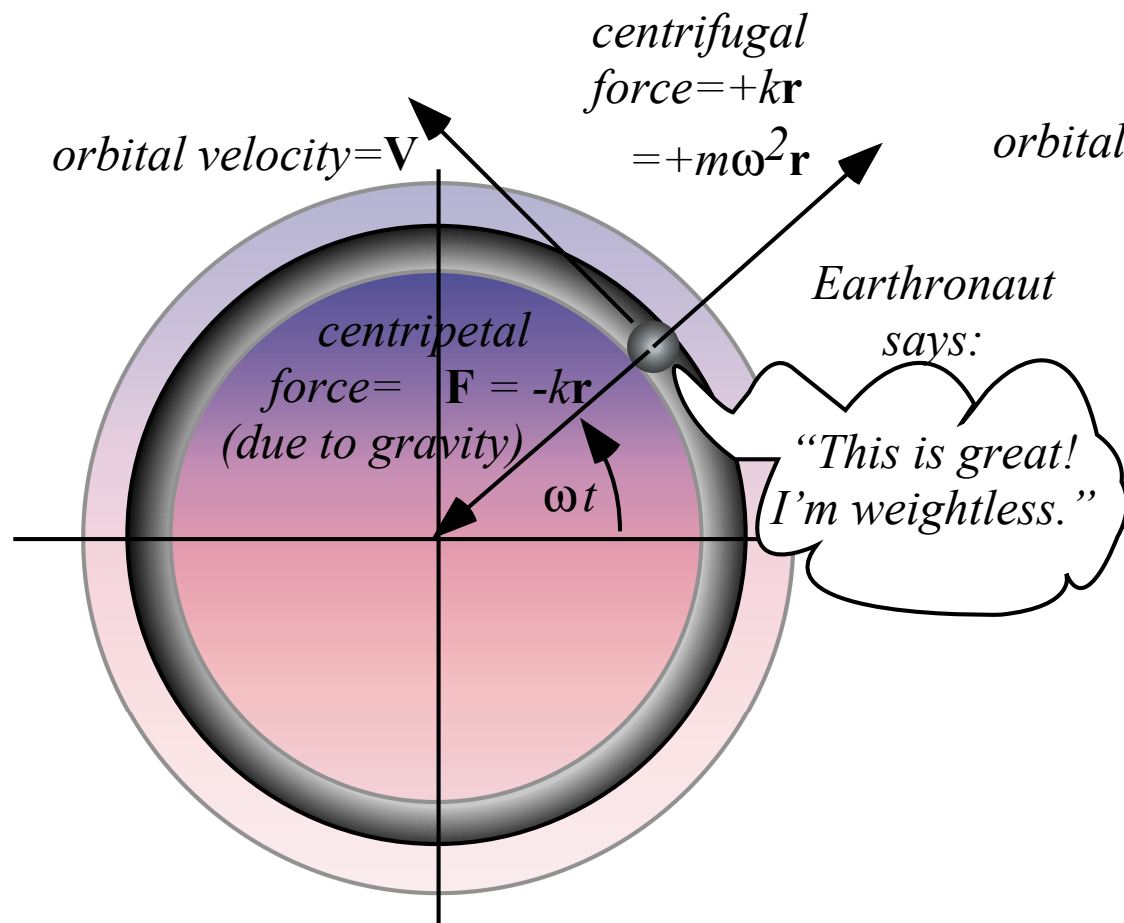
*Constructing 2D IHO orbits using **Kepler anomaly plots***

Mean-anomaly and eccentric-anomaly geometry

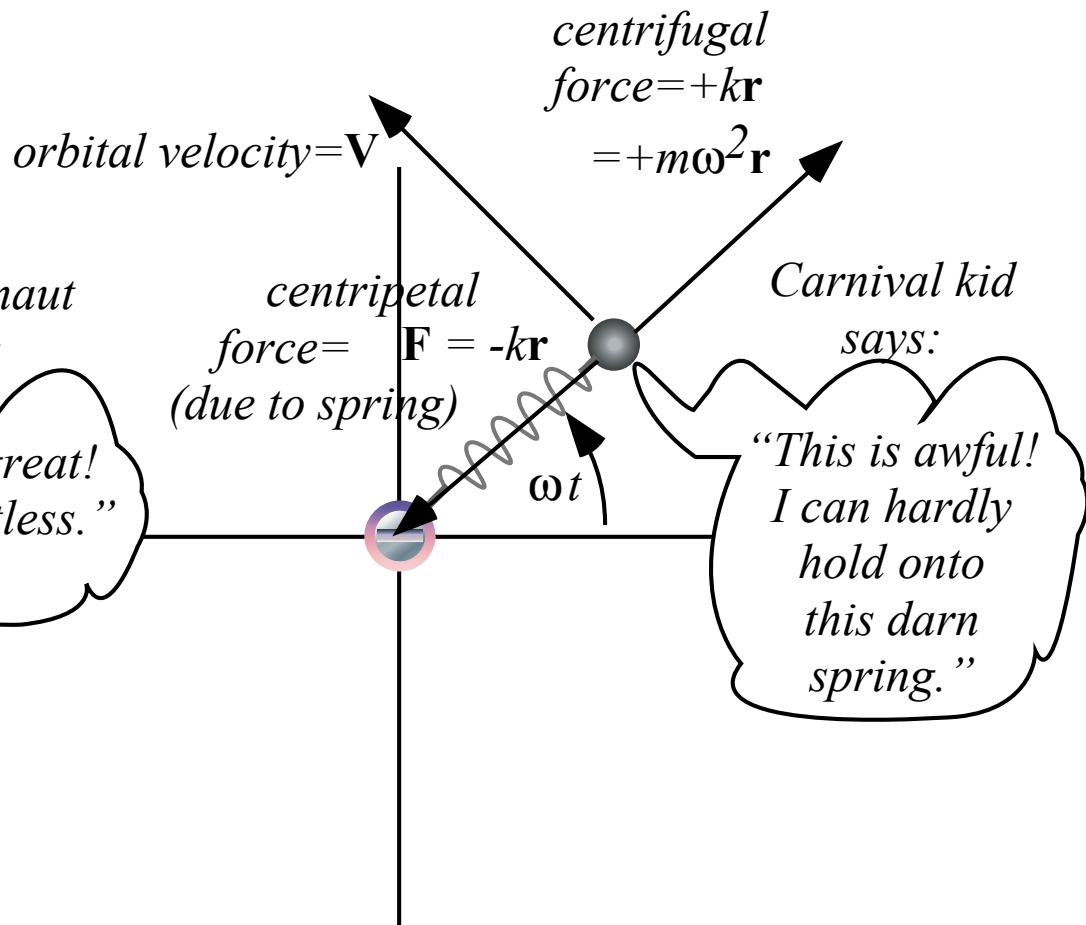
Calculus and vector geometry of IHO orbits

 *A confusing introduction to Coriolis-centrifugal force geometry* (Derived better in Ch. 12)

(a) "Earthronaut" orbiting tunnel inside Earth

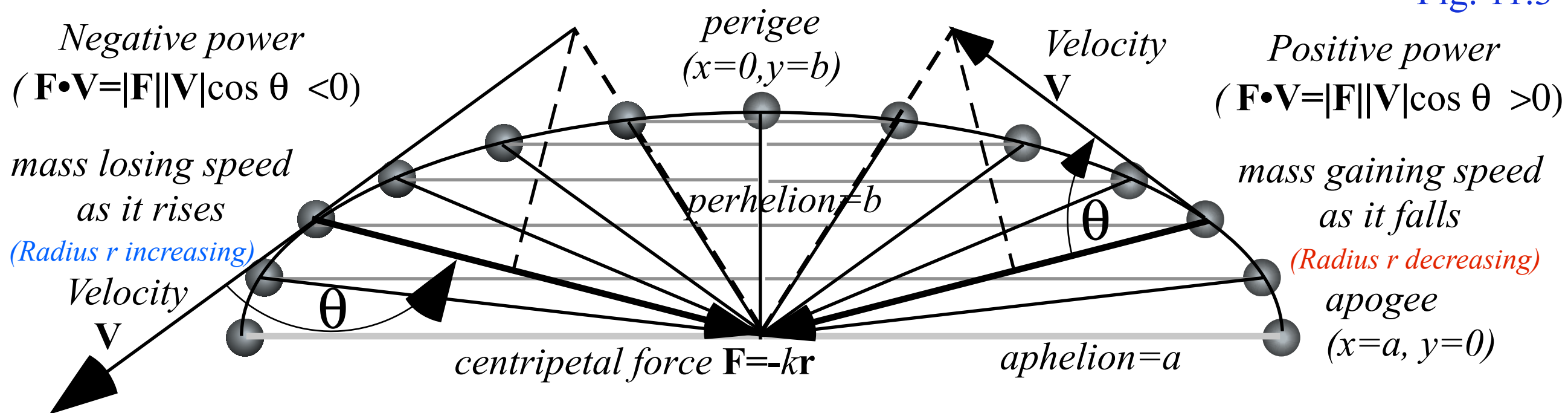


(b) "Carnival kid" orbiting in space attached to a spring



Unit 1
 Fig. 11.2

Unit 1
 Fig. 11.3

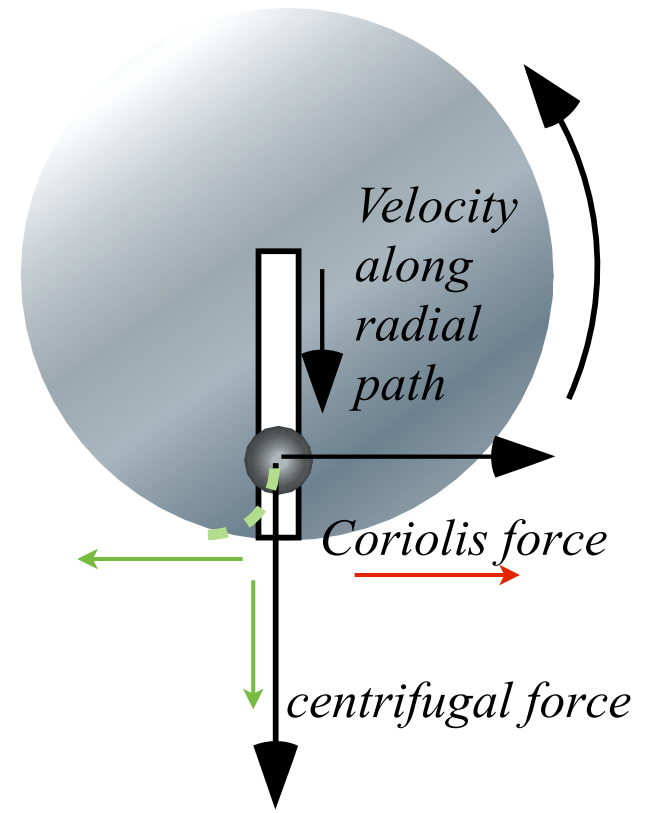
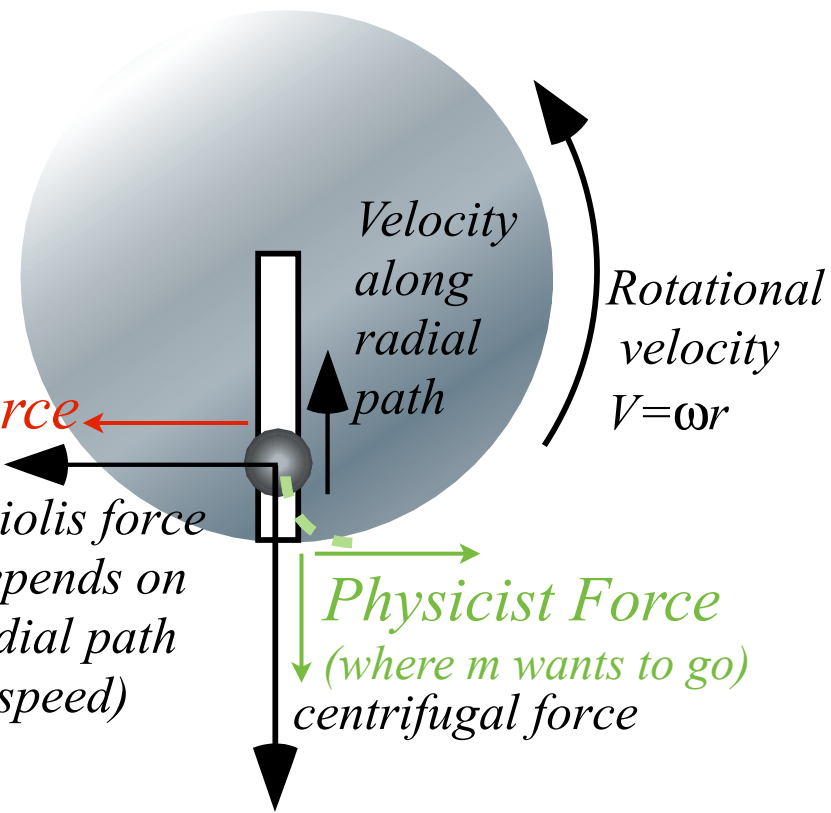


(a) Centrifugal and Coriolis Forces on Merry-Go-Round

Mathematician Force
(to hold m back)
Constraint force
keeps m in radial slot

Coriolis force
(depends on
radial path
speed)

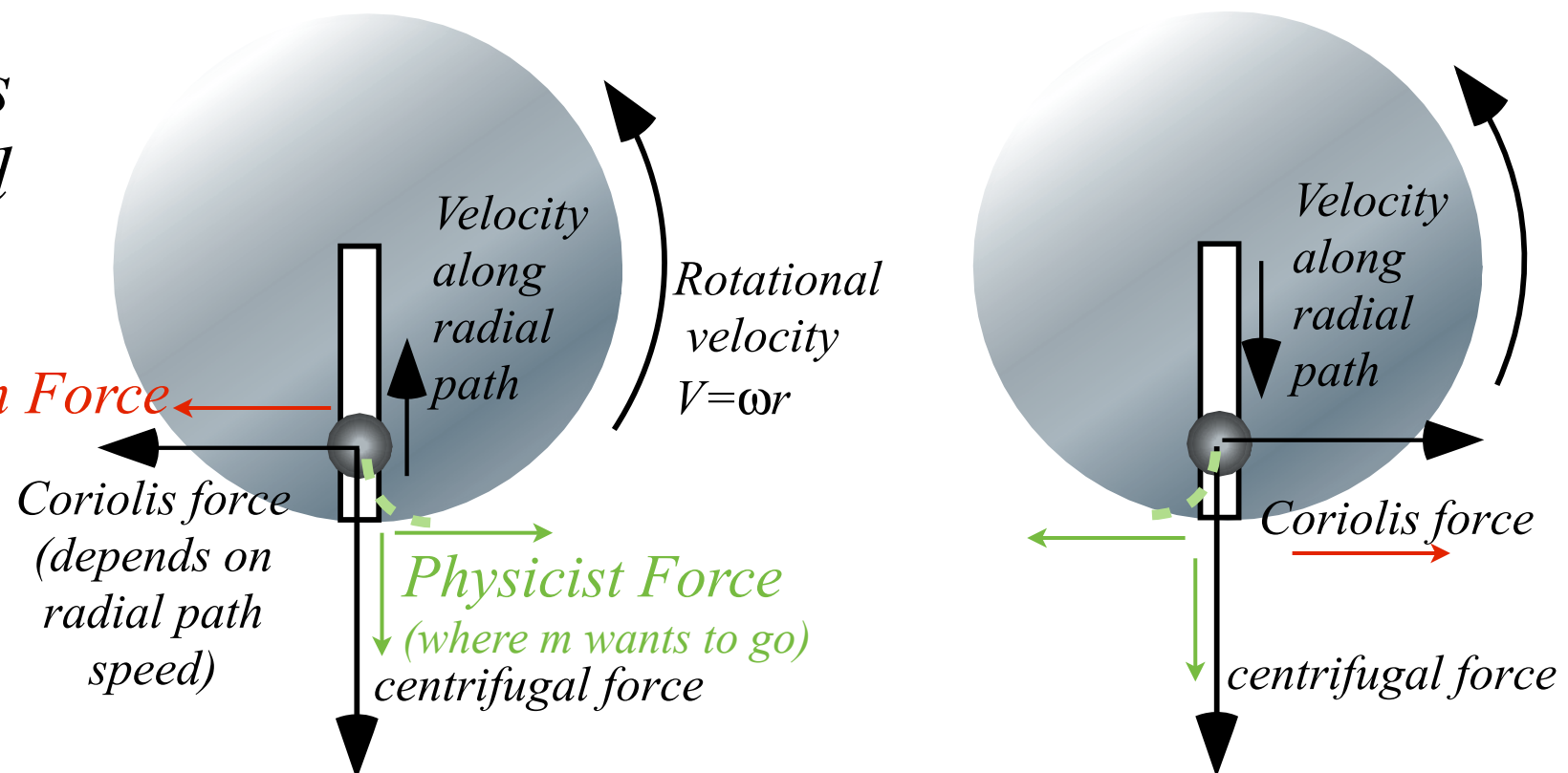
Physicist Force
(where m wants to go)
centrifugal force



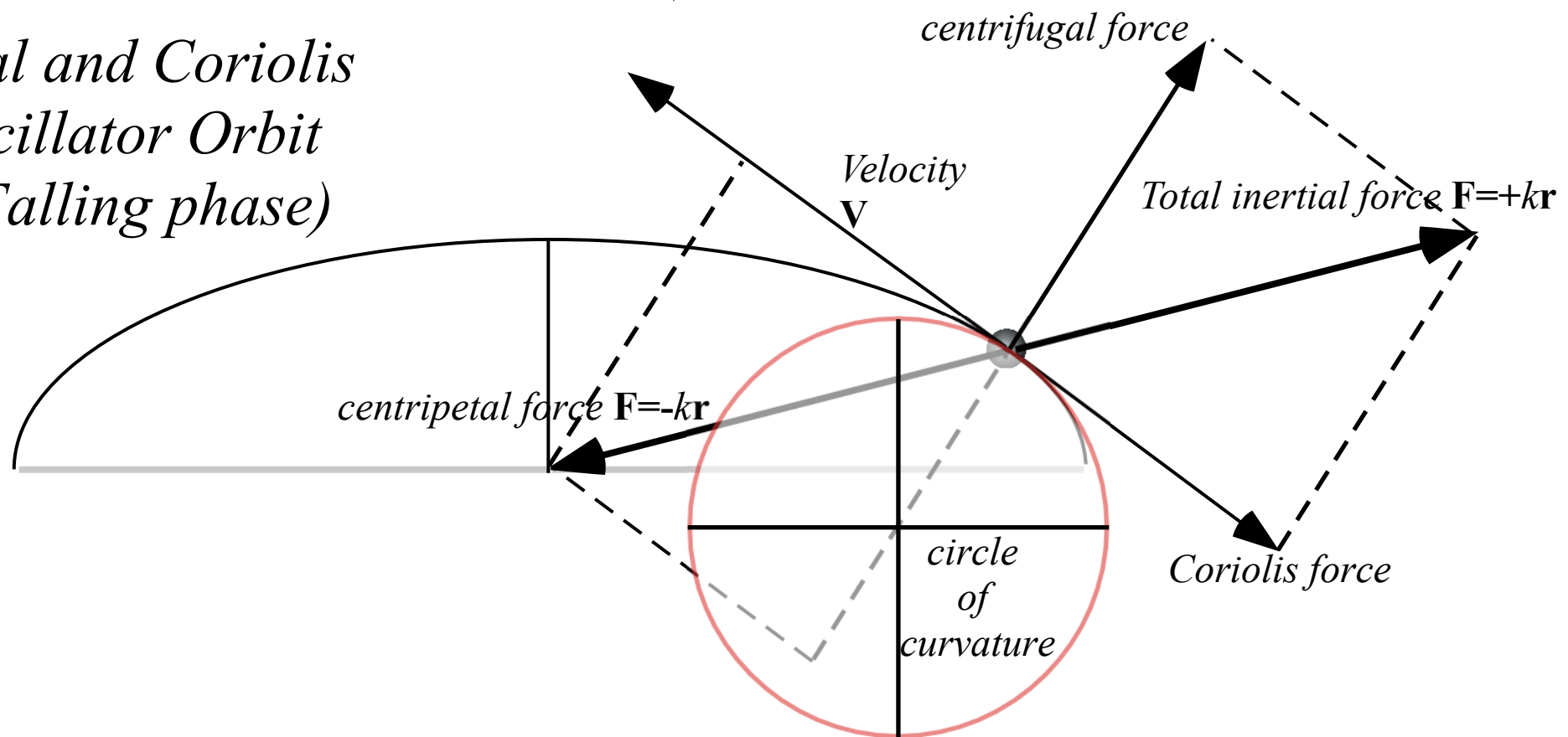
(a) Centrifugal and Coriolis Forces on Merry-Go-Round

Mathematician Force
(to hold m back)

Constraint force
keeps m in radial slot



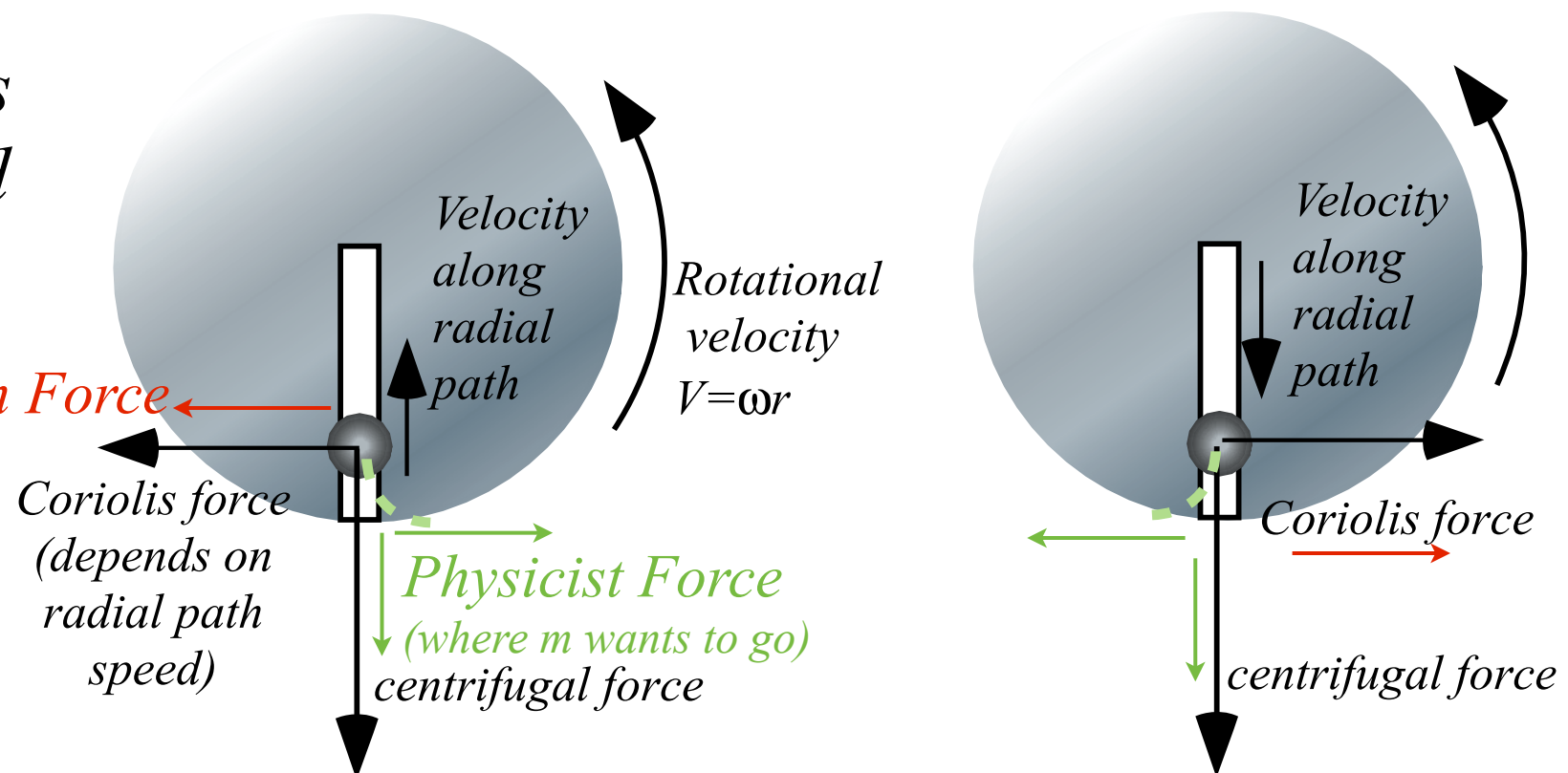
(b) Centrifugal and Coriolis Forces on Oscillator Orbit (Falling phase)



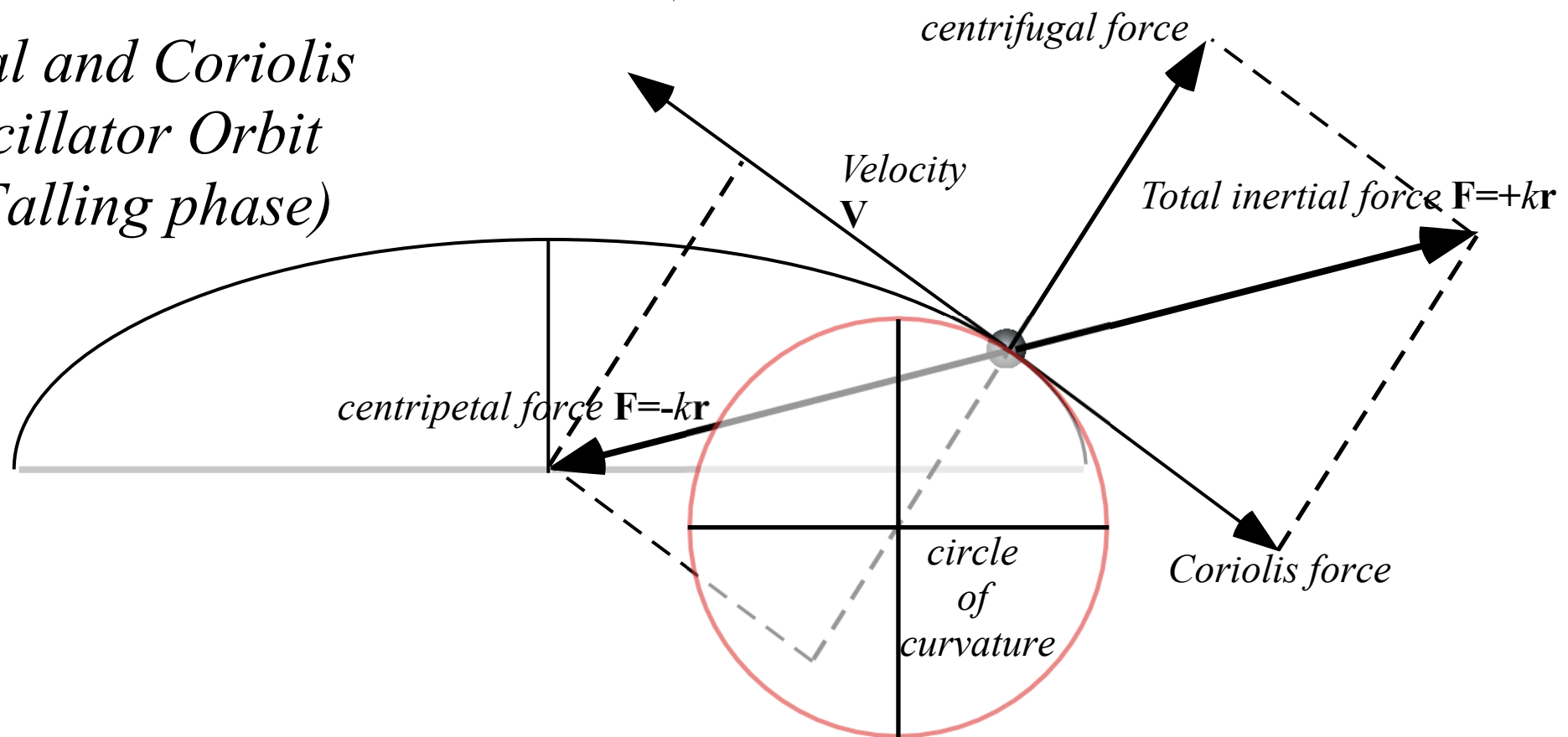
(a) Centrifugal and Coriolis Forces on Merry-Go-Round

Mathematician Force
(to hold m back)

Constraint force
keeps m in radial slot



(b) Centrifugal and Coriolis Forces on Oscillator Orbit (Falling phase)



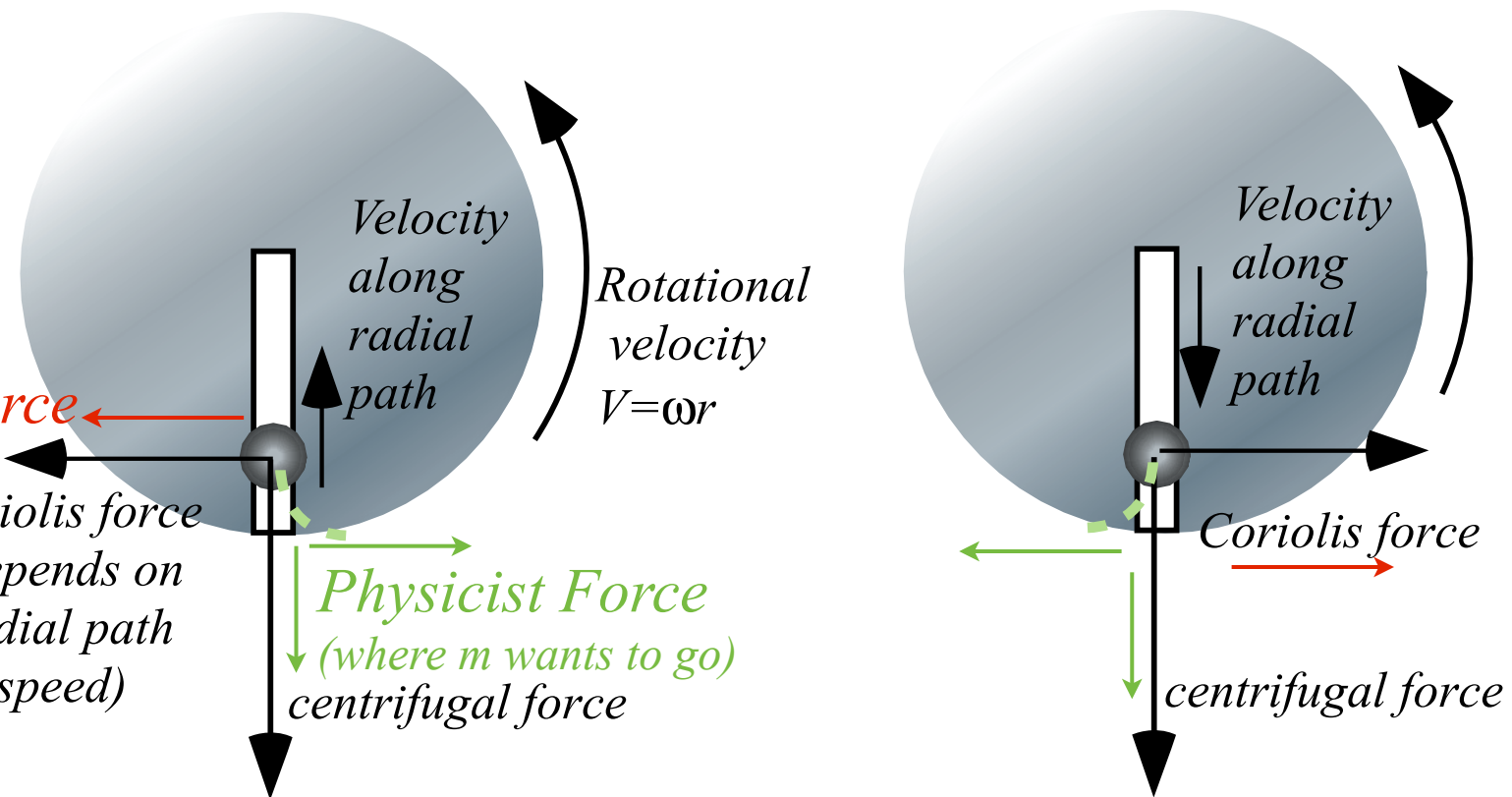
(a) Centrifugal and Coriolis Forces on Merry-Go-Round

Mathematician Force
(to hold m back)

Constraint force
keeps m in radial slot

Coriolis force
(depends on
radial path
speed)

Physicist Force
(where m wants to go)
centrifugal force



(c) Centrifugal and Coriolis Forces on Oscillator Orbit

centrifugal force (Rising phase)

Total inertial force $\mathbf{F} = +kr$

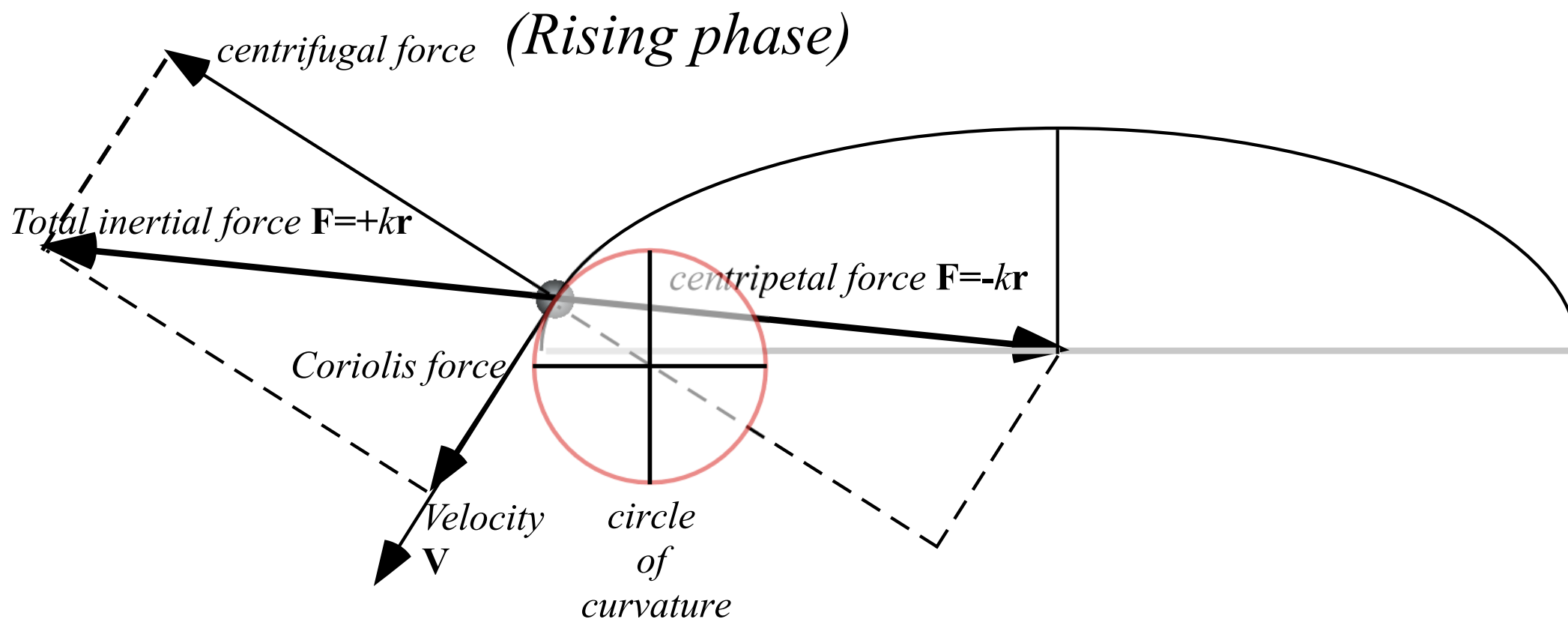
centripetal force $\mathbf{F} = -kr$

Coriolis force

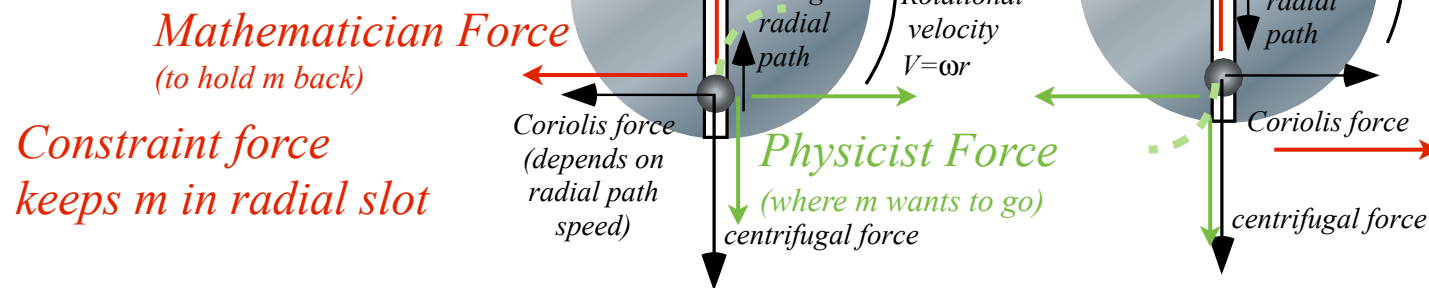
Velocity

\mathbf{v}

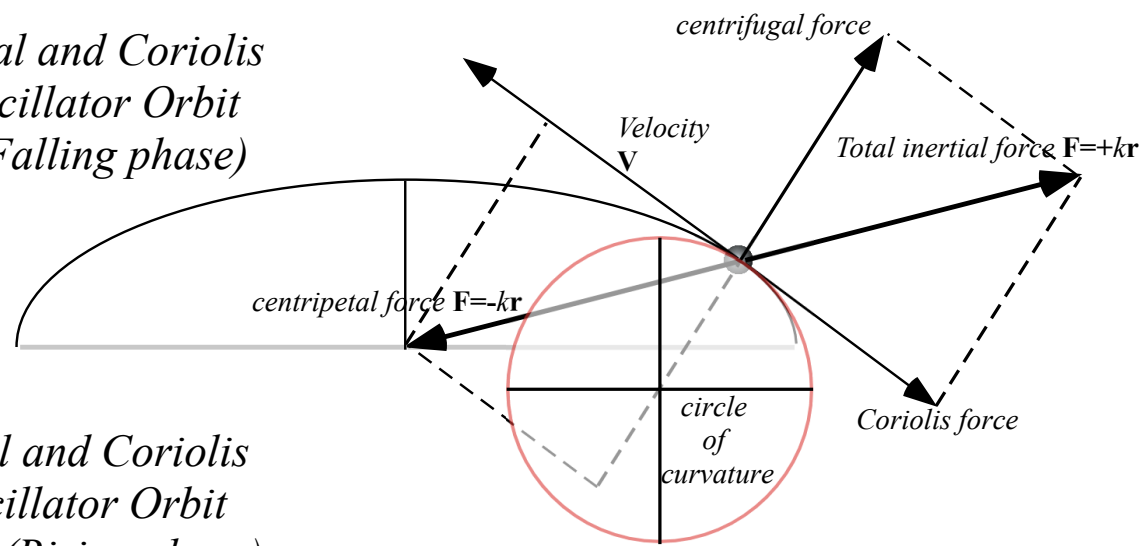
circle
of
curvature



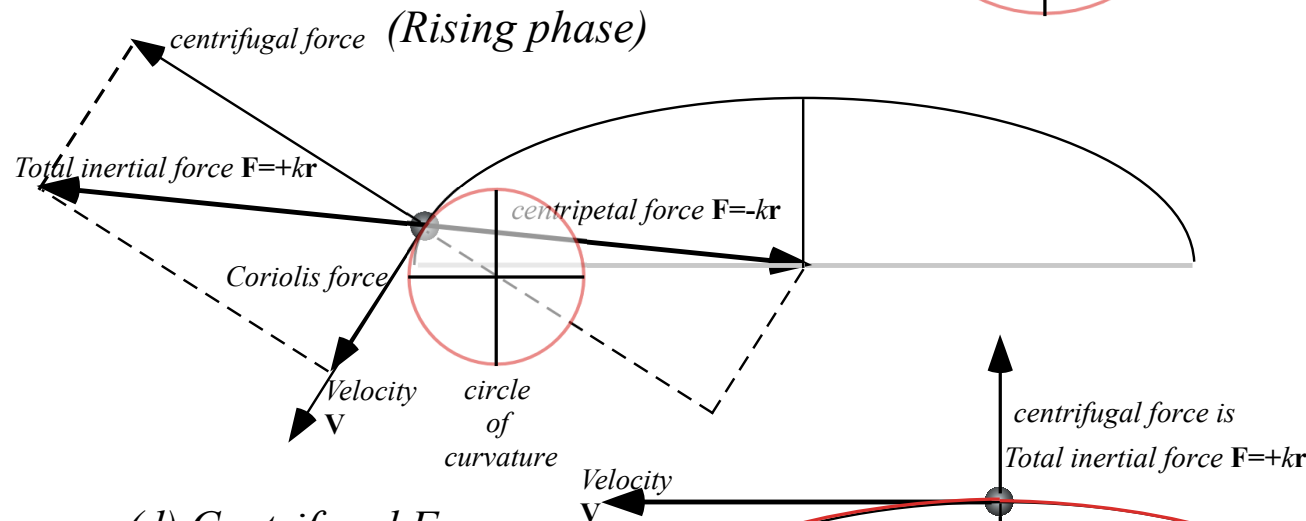
(a) Centrifugal and Coriolis Forces on Merry-Go-Round



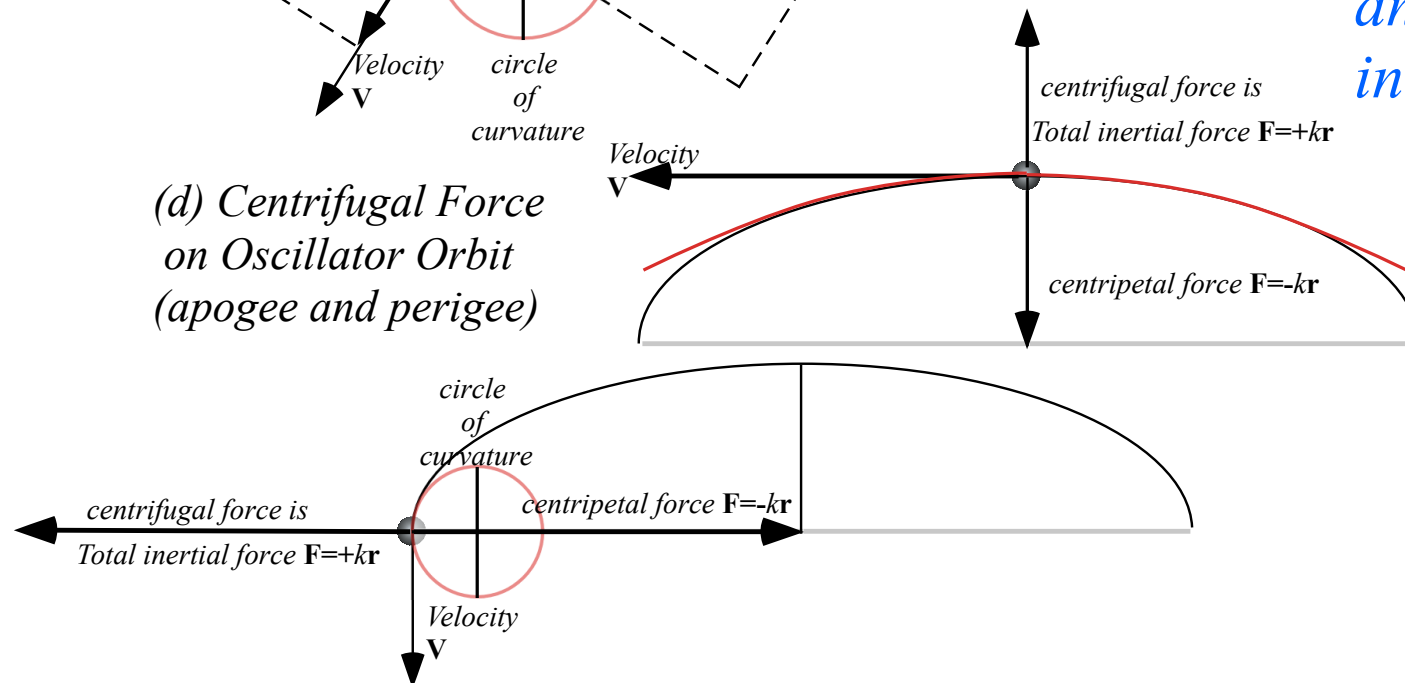
(b) Centrifugal and Coriolis Forces on Oscillator Orbit (Falling phase)



(c) Centrifugal and Coriolis Forces on Oscillator Orbit (Rising phase)




(d) Centrifugal Force on Oscillator Orbit (apogee and perigee)



Unit 1
Fig. 11.4
a-d

*A little confusing?
Discussion of Coriolis
forces will be done more elegantly
and made more physically intuitive
in Ch. 12 of Unit 1 and in Unit 6.*

Some Kepler's "laws" for all central (isotropic) force $F(r)$ fields

-  *Angular momentum invariance of IHO: $F(r) = -k \cdot r$ with $U(r) = k \cdot r^2 / 2$ (Derived here)*
- Angular momentum invariance of Coulomb: $F(r) = -GMm/r^2$ with $U(r) = -GMm \cdot /r$ (Derived in Unit 5)*
- Total energy $E = KE + PE$ invariance of IHO: $F(r) = -k \cdot r$ (Derived here)*
- Total energy $E = KE + PE$ invariance of Coulomb: $F(r) = -GMm/r^2$ (Derived in Unit 5)*

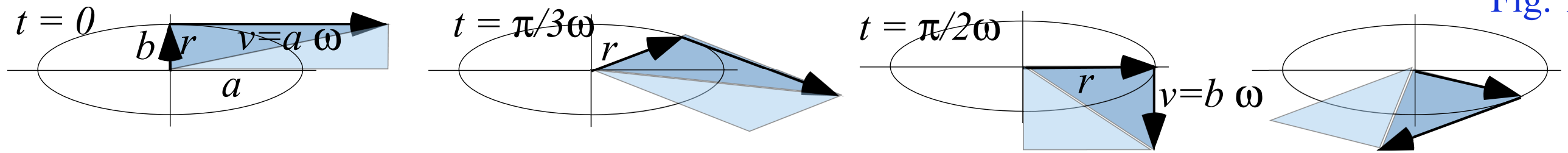
Some Kepler's "laws" for central (isotropic) force $F(r)$

...and certainly apply to the IHO: $F(r) = -k \cdot r$ with $U(r) = k \cdot r^2/2$

(Recall from Lecture 6: $k = Gm \frac{4\pi}{3} \rho_{\oplus}$)

Unit 1

Fig. 11.8



1. Area of triangle $\Delta_r^v = \mathbf{r} \times \mathbf{v}/2$ is constant

$$\mathbf{r} \times \mathbf{v} = r_x v_y - r_y v_x = a \cos \omega t \cdot (b \omega \cos \omega t) - b \sin \omega t \cdot (-a \omega \sin \omega t) = ab \cdot \omega (\cos^2 \omega t + \sin^2 \omega t) \quad \checkmark \text{ for IHO}$$

$$\begin{pmatrix} r_x \\ r_y \end{pmatrix} = \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} a \cos \omega t \\ b \sin \omega t \end{pmatrix} \quad \begin{pmatrix} v_x \\ v_y \end{pmatrix} = \begin{pmatrix} -a \omega \sin \omega t \\ b \omega \cos \omega t \end{pmatrix}$$

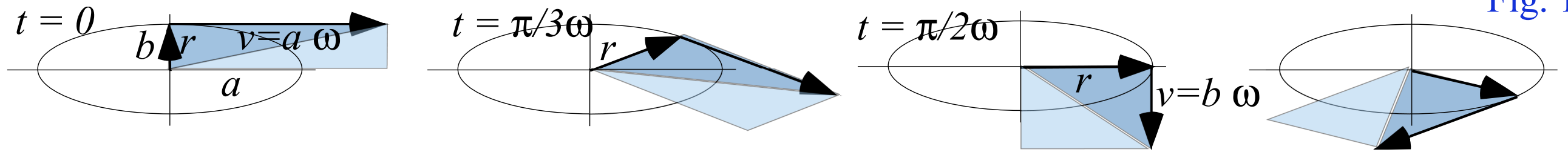
Some Kepler's "laws" that apply to any central (isotropic) force $F(r)$

...and certainly apply to the IHO: $F(r) = -k \cdot r$ with $U(r) = k \cdot r^2 / 2$

(Recall from Lecture 6: $k = Gm \frac{4\pi}{3} \rho_{\oplus}$)

Unit 1

Fig. 11.8



1. Area of triangle $\Delta_r^v = \mathbf{r} \times \mathbf{v} / 2$ is constant

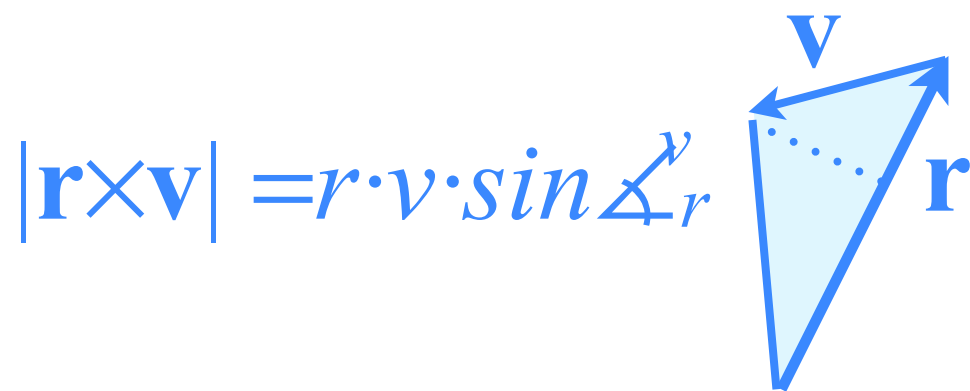
$$\mathbf{r} \times \mathbf{v} = r_x v_y - r_y v_x = a \cos \omega t \cdot (b \omega \cos \omega t) - a \sin \omega t \cdot (-b \omega \sin \omega t) = ab \cdot \omega$$

✓ for IHO

2. Angular momentum $\mathbf{L} = m \mathbf{r} \times \mathbf{v}$ is conserved

$$L = m |\mathbf{r} \times \mathbf{v}| = m (r_x v_y - r_y v_x) = m \cdot ab \cdot \omega$$

✓ for IHO



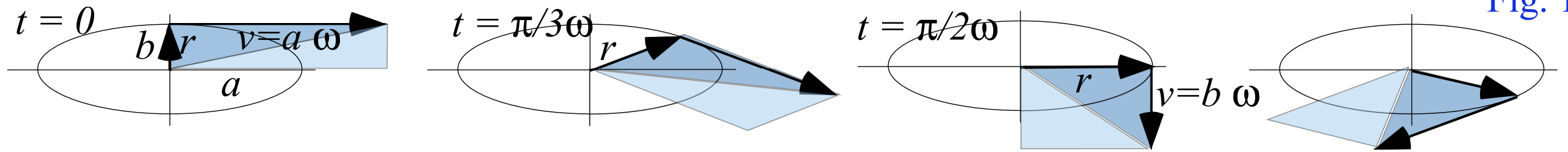
Some Kepler's "laws" that apply to any central (isotropic) force $F(r)$

...and certainly apply to the IHO: $F(r) = -k \cdot r$ with $U(r) = k \cdot r^2 / 2$

(Recall from Lecture 6: $k = Gm \frac{4\pi}{3} \rho_{\oplus}$)

Unit 1

Fig. 11.8



1. Area of triangle $\Delta_r^v = \mathbf{r} \times \mathbf{v} / 2$ is constant

$$\mathbf{r} \times \mathbf{v} = r_x v_y - r_y v_x = a \cos \omega t \cdot (b \omega \cos \omega t) - a \sin \omega t \cdot (-b \omega \sin \omega t) = ab \cdot \omega$$

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✓ for IHO

3. Equal area is swept by radius vector in each equal time interval T

$$A_T = \int_0^T \frac{\mathbf{r} \times d\mathbf{r}}{2} = \int_0^T \frac{\mathbf{r} \times \frac{d\mathbf{r}}{dt}}{2} dt = \int_0^T \frac{\mathbf{r} \times \mathbf{v}}{2} dt = \frac{L}{2m} \int_0^T dt = \frac{L}{2m} T$$

✓ for IHO

$$|\mathbf{r} \times d\mathbf{r}| = r \cdot dr \cdot \sin \Delta_r^{dr}$$

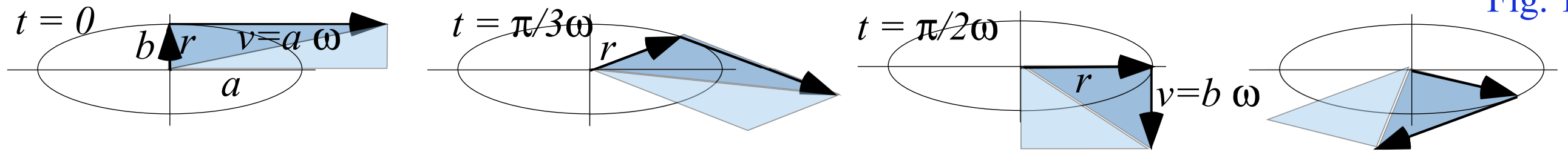
Some Kepler's "laws" that apply to any central (isotropic) force $F(r)$

...and certainly apply to the IHO: $F(r) = -k \cdot r$ with $U(r) = k \cdot r^2 / 2$

(Recall from Lecture 6: $k = Gm \frac{4\pi}{3} \rho_{\oplus}$)

Unit 1

Fig. 11.8



1. Area of triangle $\Delta_r^v = \mathbf{r} \times \mathbf{v} / 2$ is constant

$$\mathbf{r} \times \mathbf{v} = r_x v_y - r_y v_x = a \cos \omega t \cdot (b \omega \cos \omega t) - a \sin \omega t \cdot (-b \omega \sin \omega t) = ab \cdot \omega$$

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2. Angular momentum $L = m \mathbf{r} \times \mathbf{v}$ is conserved

$$L = m \mathbf{r} \times \mathbf{v} = m (r_x v_y - r_y v_x) = m \cdot ab \cdot \omega = m \cdot ab \cdot \frac{2\pi}{\tau}$$

✓ for IHO

3. Equal area is swept by radius vector in each equal time interval T

$$A_T = \int_0^T \frac{\mathbf{r} \times d\mathbf{r}}{2} = \int_0^T \frac{\mathbf{r} \times \frac{d\mathbf{r}}{dt}}{2} dt = \int_0^T \frac{\mathbf{r} \times \mathbf{v}}{2} dt = \frac{L}{2m} \int_0^T dt = \frac{L}{2m} T$$

✓ for IHO

In one period: $\tau = \frac{1}{\nu} = \frac{2\pi}{\omega} = \frac{2mA_\tau}{L}$ the area is: $A_\tau = \frac{L\tau}{2m}$ ($= ab \cdot \pi$ for ellipse orbit)

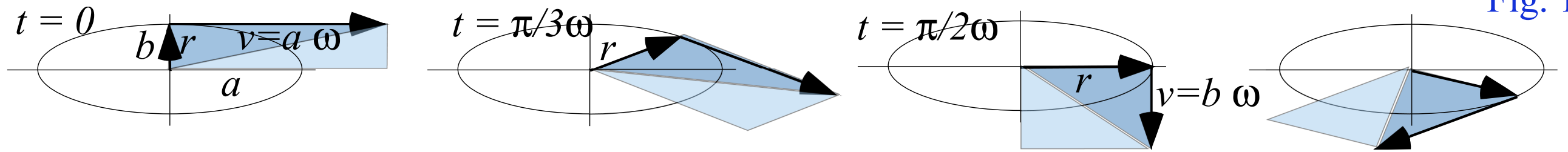
Some Kepler's "laws" that apply to any central (isotropic) force $F(r)$

...and certainly apply to the IHO: $F(r) = -k \cdot r$ with $U(r) = k \cdot r^2 / 2$

(Recall from Lecture 6: $k = Gm \frac{4\pi}{3} \rho_{\oplus}$)

Unit 1

Fig. 11.8



1. Area of triangle $\Delta_r^v = \mathbf{r} \times \mathbf{v} / 2$ is constant

$$\mathbf{r} \times \mathbf{v} = r_x v_y - r_y v_x = a \cos \omega t \cdot (b \omega \cos \omega t) - a \sin \omega t \cdot (-b \omega \sin \omega t) = ab \cdot \omega$$

✓ for IHO

2. Angular momentum $L = m \mathbf{r} \times \mathbf{v}$ is conserved

$$L = m \mathbf{r} \times \mathbf{v} = m (r_x v_y - r_y v_x) = m \cdot ab \cdot \omega = m \cdot ab \cdot \frac{2\pi}{\tau}$$

✓ for IHO

3. Equal area is swept by radius vector in each equal time interval T

$$A_T = \int_0^T \frac{\mathbf{r} \times d\mathbf{r}}{2} = \int_0^T \frac{\mathbf{r} \times \frac{d\mathbf{r}}{dt}}{2} dt = \int_0^T \frac{\mathbf{r} \times \mathbf{v}}{2} dt = \frac{L}{2m} \int_0^T dt = \frac{L}{2m} T$$

✓ for IHO

In one period: $\tau = \frac{1}{\nu} = \frac{2\pi}{\omega} = \frac{2mA_\tau}{L}$ the area is: $A_\tau = \frac{L\tau}{2m}$ ($= ab \cdot \pi$ for ellipse orbit)

(Recall from Lecture 6: $\omega = \sqrt{k/m} = \sqrt{G\rho_{\oplus} 4\pi/3}$)

(IHO formulas from Lect. 6 p.70-79)

Some Kepler's "laws" for all central (isotropic) force $F(r)$ fields

Angular momentum invariance of IHO: $F(r) = -k \cdot r$ with $U(r) = k \cdot r^2/2$ (Derived here)

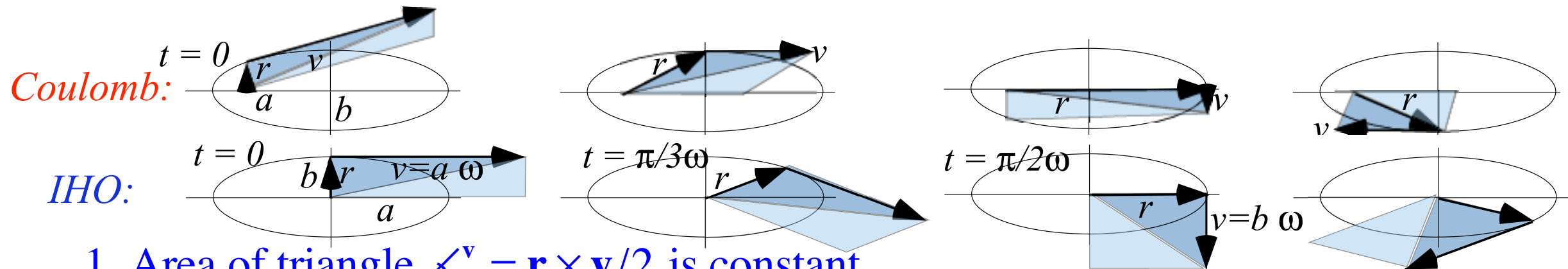
 *Angular momentum invariance of Coulomb: $F(r) = -GMm/r^2$ with $U(r) = -GMm \cdot /r$ (Derived in Unit 5)*

Total energy $E = KE + PE$ invariance of IHO: $F(r) = -k \cdot r$ (Derived here)

Total energy $E = KE + PE$ invariance of Coulomb: $F(r) = -GMm/r^2$ (Derived in Unit 5)

Some Kepler's "laws" that apply to any central (isotropic) force $F(r)$

Apply to IHO: $F(r) = -k \cdot r$ with $U(r) = k \cdot r^2/2$ and Coulomb: $F(r) = -GMm/r^2$ with $U(r) = -GMm \cdot /r$



1. Area of triangle $\triangle_r^v = \mathbf{r} \times \mathbf{v} / 2$ is constant

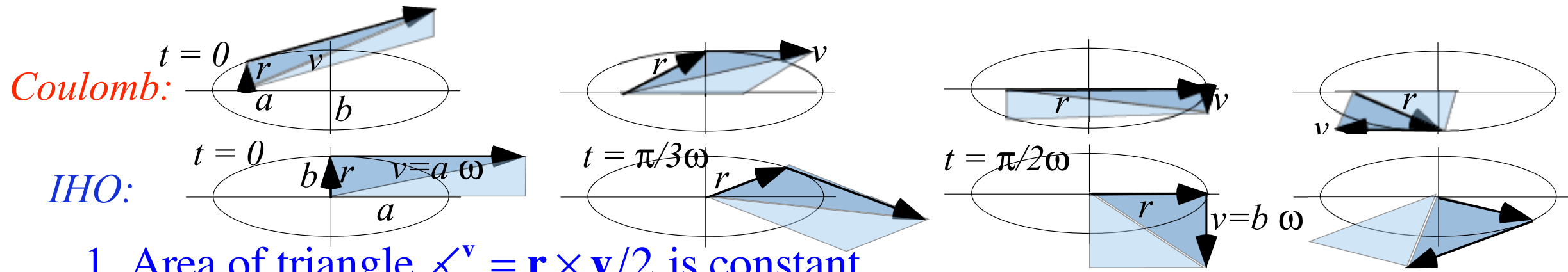
$$\mathbf{r} \times \mathbf{v} = r_x v_y - r_y v_x = \begin{cases} ab \cdot \sqrt{G\rho_{\oplus} 4\pi / 3} & \text{for IHO} \\ a^{-1/2} b \sqrt{GM_{\oplus}} & \text{for Coul.} \end{cases}$$

✓ for IHO

(Derived in Unit 5) ✓ for Coul.

Some Kepler's "laws" that apply to any central (isotropic) force $F(r)$

Apply to IHO: $F(r) = -k \cdot r$ with $U(r) = k \cdot r^2/2$ and Coulomb: $F(r) = -GMm/r^2$ with $U(r) = -GMm \cdot /r$



1. Area of triangle $\triangle_r^v = \mathbf{r} \times \mathbf{v} / 2$ is constant

$$\mathbf{r} \times \mathbf{v} = r_x v_y - r_y v_x = \begin{cases} ab \cdot \sqrt{G\rho_{\oplus} 4\pi / 3} & \text{for IHO} \\ a^{-1/2} b \sqrt{GM_{\oplus}} & \text{for Coul.} \end{cases}$$

✓ for IHO
(Derived in Unit 5) ✓ for Coul.

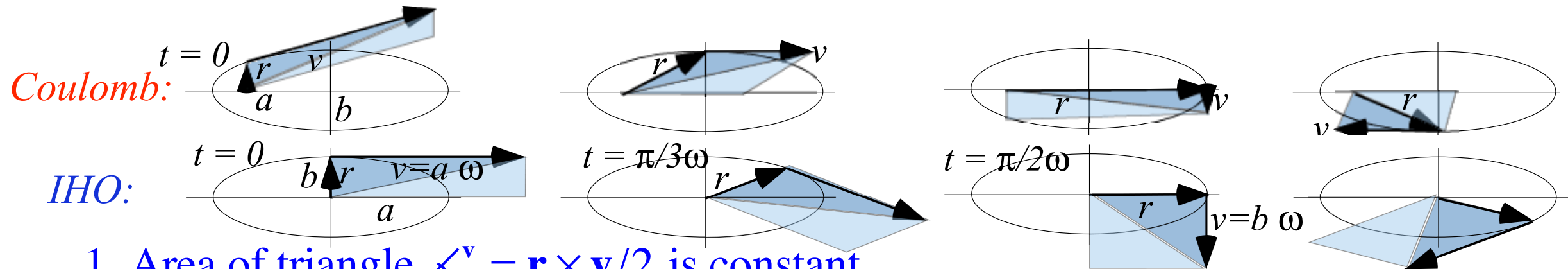
2. Angular momentum $L = m \mathbf{r} \times \mathbf{v}$ is conserved

$$L = m \mathbf{r} \times \mathbf{v} = m (r_x v_y - r_y v_x) = \begin{cases} m \cdot ab \cdot \sqrt{G\rho_{\oplus} 4\pi / 3} & \text{for IHO} \\ m \cdot a^{-1/2} b \sqrt{GM_{\oplus}} & \text{for Coul. (... in Unit 5)} \end{cases}$$

✓ for IHO
✓ for Coul.

Some Kepler's "laws" that apply to any central (isotropic) force $F(r)$

Apply to IHO: $F(r) = -k \cdot r$ with $U(r) = k \cdot r^2/2$ and Coulomb: $F(r) = -GMm/r^2$ with $U(r) = -GMm \cdot /r$



1. Area of triangle $\triangle_r^v = \mathbf{r} \times \mathbf{v} / 2$ is constant

$$\mathbf{r} \times \mathbf{v} = r_x v_y - r_y v_x = \begin{cases} ab \cdot \sqrt{G\rho_{\oplus} 4\pi / 3} & \text{for IHO} \\ a^{-1/2} b \sqrt{GM_{\oplus}} & \text{for Coul.} \end{cases}$$

✓ for IHO
(Derived in Unit 5) ✓ for Coul.

2. Angular momentum $L = m \mathbf{r} \times \mathbf{v}$ is conserved

$$L = m \mathbf{r} \times \mathbf{v} = m (r_x v_y - r_y v_x) = \begin{cases} m \cdot ab \cdot \sqrt{G\rho_{\oplus} 4\pi / 3} & \text{for IHO} \\ m \cdot a^{-1/2} b \sqrt{GM_{\oplus}} & \text{for Coul. (... in Unit 5)} \end{cases}$$

✓ for IHO
✓ for Coul.

3. Equal area is swept by radius vector in each equal time interval T

In one period:

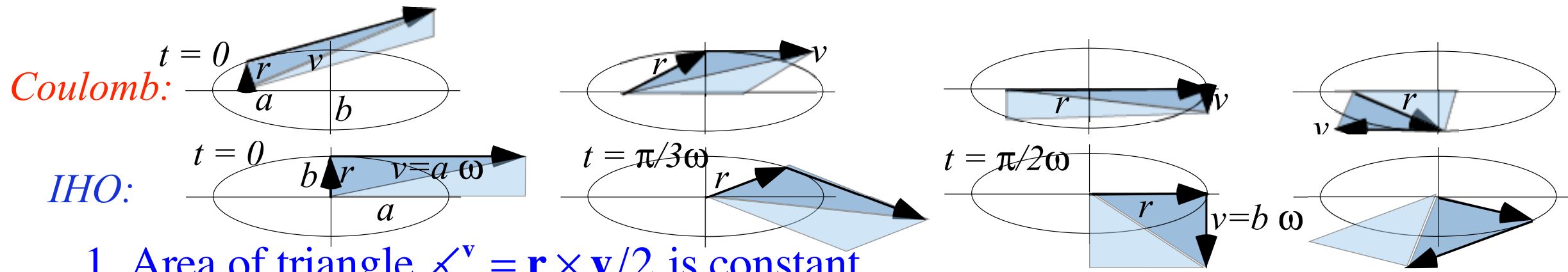
$$\tau = \frac{1}{\nu} = \frac{2\pi}{\omega} = \frac{2mA_{\tau}}{L} = \frac{2m \cdot ab \cdot \pi}{L} = \begin{cases} \frac{2m \cdot ab \cdot \pi}{m \cdot ab \cdot \sqrt{G\rho_{\oplus} 4\pi / 3}} \\ \frac{2m \cdot ab \cdot \pi}{m \cdot a^{-1/2} b \sqrt{GM_{\oplus}}} \end{cases}$$

Applies to any central $F(r)$ Applies to IHO and Coulomb

(IHO formulas from Lect. 6 p.70-79)

Some Kepler's "laws" that apply to any central (isotropic) force $F(r)$

Apply to IHO: $F(r) = -k \cdot r$ with $U(r) = k \cdot r^2/2$ and Coulomb: $F(r) = -GMm/r^2$ with $U(r) = -GMm \cdot /r$



1. Area of triangle $\triangle_r^v = \mathbf{r} \times \mathbf{v} / 2$ is constant

$$\mathbf{r} \times \mathbf{v} = r_x v_y - r_y v_x = \begin{cases} ab \cdot \sqrt{G\rho_{\oplus} 4\pi / 3} & \text{for IHO} \\ a^{-1/2} b \sqrt{GM_{\oplus}} & \text{for Coul.} \end{cases}$$

✓ for IHO
(Derived in Unit 5) ✓ for Coul.

2. Angular momentum $L = m \mathbf{r} \times \mathbf{v}$ is conserved

$$L = m \mathbf{r} \times \mathbf{v} = m (r_x v_y - r_y v_x) = \begin{cases} m \cdot ab \cdot \sqrt{G\rho_{\oplus} 4\pi / 3} & \text{for IHO} \\ m \cdot a^{-1/2} b \sqrt{GM_{\oplus}} & \text{for Coul. (... in Unit 5)} \end{cases}$$

✓ for IHO
✓ for Coul.

3. Equal area is swept by radius vector in each equal time interval T

In one period:

$$\tau = \frac{1}{\nu} = \frac{2\pi}{\omega} = \frac{2mA_{\tau}}{L} = \frac{2m \cdot ab \cdot \pi}{L}$$

Applies to any central $F(r)$

$$= \begin{cases} \frac{2m \cdot ab \cdot \pi}{m \cdot ab \cdot \sqrt{G\rho_{\oplus} 4\pi / 3}} = \frac{2\pi}{\sqrt{G\rho_{\oplus} 4\pi / 3}} & \text{for IHO} \\ \frac{2m \cdot ab \cdot \pi}{m \cdot a^{-1/2} b \sqrt{GM_{\oplus}}} = \frac{2\pi}{a^{-3/2} \sqrt{GM_{\oplus}}} & \text{for Coul.} \end{cases}$$

(not a function of a or b)
that is ω_{IHO}
that is ω_{Coul}
(not a function of b)

Some Kepler's "laws" for all central (isotropic) force $F(r)$ fields

Angular momentum invariance of IHO: $F(r) = -k \cdot r$ with $U(r) = k \cdot r^2/2$

(Derived here)

Angular momentum invariance of Coulomb: $F(r) = -GMm/r^2$ with $U(r) = -GMm \cdot /r$

(Derived in Unit 5)



Total energy $E = KE + PE$ invariance of IHO: $F(r) = -k \cdot r$

(Derived here)

Total energy $E = KE + PE$ invariance of Coulomb: $F(r) = -GMm/r^2$

(Derived in Unit 5)

Kepler laws involve ∇ -momentum conservation in isotropic force $F(r)$

Now consider orbital energy conservation of the IHO: $F(r)=-k \cdot r$ with $U(r)=k \cdot r^2/2$

Total energy= $KE + PE$ is constant

$$\begin{aligned} KE + PE &= \frac{1}{2} \mathbf{v} \cdot \mathbf{M} \cdot \mathbf{v} + \frac{1}{2} \mathbf{r} \cdot \mathbf{K} \cdot \mathbf{r} \\ &= \frac{1}{2} \begin{pmatrix} v_x & v_y \end{pmatrix} \cdot \begin{pmatrix} m & 0 \\ 0 & m \end{pmatrix} \cdot \begin{pmatrix} v_x \\ v_y \end{pmatrix} + \begin{pmatrix} r_x & r_y \end{pmatrix} \cdot \begin{pmatrix} k & 0 \\ 0 & k \end{pmatrix} \cdot \begin{pmatrix} r_x \\ r_y \end{pmatrix} \\ &= \frac{1}{2} m v_x^2 + \frac{1}{2} m v_y^2 + \frac{1}{2} k r_x^2 + \frac{1}{2} k r_y^2 \\ &= \frac{1}{2} m (-a\omega \sin \omega t)^2 + \frac{1}{2} m (b\omega \cos \omega t)^2 + \frac{1}{2} k (a \cos \omega t)^2 + \frac{1}{2} k (b \sin \omega t)^2 \end{aligned}$$

$$\begin{pmatrix} v_x \\ v_y \end{pmatrix} = \begin{pmatrix} -a\omega \sin \omega t \\ b\omega \cos \omega t \end{pmatrix}$$

$$\begin{pmatrix} r_x \\ r_y \end{pmatrix} = \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} a \cos \omega t \\ b \sin \omega t \end{pmatrix}$$

Kepler laws involve Δ -momentum conservation in isotropic force $F(r)$

Now consider orbital energy conservation of the IHO: $F(r) = -k \cdot r$ with $U(r) = k \cdot r^2 / 2$

Total IHO energy = KE + PE is constant

$$\begin{aligned} KE + PE &= \frac{1}{2} \mathbf{v} \cdot \mathbf{M} \cdot \mathbf{v} + \frac{1}{2} \mathbf{r} \cdot \mathbf{K} \cdot \mathbf{r} \\ &= \frac{1}{2} \begin{pmatrix} v_x & v_y \end{pmatrix} \cdot \begin{pmatrix} m & 0 \\ 0 & m \end{pmatrix} \cdot \begin{pmatrix} v_x \\ v_y \end{pmatrix} + \begin{pmatrix} r_x & r_y \end{pmatrix} \cdot \begin{pmatrix} k & 0 \\ 0 & k \end{pmatrix} \cdot \begin{pmatrix} r_x \\ r_y \end{pmatrix} \\ &= \frac{1}{2} m v_x^2 + \frac{1}{2} m v_y^2 + \frac{1}{2} k r_x^2 + \frac{1}{2} k r_y^2 \\ &= \frac{1}{2} m (-a\omega \sin \omega t)^2 + \frac{1}{2} m (b\omega \cos \omega t)^2 + \frac{1}{2} k (a \cos \omega t)^2 + \frac{1}{2} k (b \sin \omega t)^2 \\ &= \frac{1}{2} m a^2 \omega^2 (\sin^2 \omega t) + \frac{1}{2} m b^2 \omega^2 (\cos^2 \omega t) + \frac{1}{2} k a^2 (\cos^2 \omega t) + \frac{1}{2} k b^2 (\sin^2 \omega t) \\ &= \frac{1}{2} m \omega^2 (a^2 + b^2) \quad \text{Given : } k = m \omega^2 \end{aligned}$$

Kepler laws involve Δ -momentum conservation in isotropic force $F(r)$

Now consider orbital energy conservation of the IHO: $F(r)=-k \cdot r$ with $U(r)=k \cdot r^2/2$

Total IHO energy= $KE + PE$ is constant

$$\begin{aligned}
 KE + PE &= \frac{1}{2} \mathbf{v} \cdot \mathbf{M} \cdot \mathbf{v} + \frac{1}{2} \mathbf{r} \cdot \mathbf{K} \cdot \mathbf{r} \\
 &= \frac{1}{2} \begin{pmatrix} v_x & v_y \end{pmatrix} \cdot \begin{pmatrix} m & 0 \\ 0 & m \end{pmatrix} \cdot \begin{pmatrix} v_x \\ v_y \end{pmatrix} + \begin{pmatrix} r_x & r_y \end{pmatrix} \cdot \begin{pmatrix} k & 0 \\ 0 & k \end{pmatrix} \cdot \begin{pmatrix} r_x \\ r_y \end{pmatrix} \\
 &= \frac{1}{2} m v_x^2 + \frac{1}{2} m v_y^2 + \frac{1}{2} k r_x^2 + \frac{1}{2} k r_y^2 \\
 &= \frac{1}{2} m (-a\omega \sin \omega t)^2 + \frac{1}{2} m (b\omega \cos \omega t)^2 + \frac{1}{2} k (a \cos \omega t)^2 + \frac{1}{2} k (b \sin \omega t)^2 \\
 &= \frac{1}{2} m a^2 \omega^2 (\sin^2 \omega t) + \frac{1}{2} m b^2 \omega^2 (\cos^2 \omega t) + \frac{1}{2} k a^2 (\cos^2 \omega t) + \frac{1}{2} k b^2 (\sin^2 \omega t) \\
 &= \frac{1}{2} m \omega^2 (a^2 + b^2) \quad \text{Given : } k = m\omega^2 \\
 E = KE + PE &= \frac{1}{2} m \omega^2 (a^2 + b^2) = \frac{1}{2} k (a^2 + b^2) \quad \text{since: } \omega = \sqrt{\frac{k}{m}} = \sqrt{G \rho_{\oplus} 4\pi / 3} \quad \text{or: } m\omega^2 = k
 \end{aligned}$$

Some Kepler's "laws" for all central (isotropic) force $F(r)$ fields

Angular momentum invariance of IHO: $F(r) = -k \cdot r$ with $U(r) = k \cdot r^2/2$ (Derived here)

Angular momentum invariance of Coulomb: $F(r) = -GMm/r^2$ with $U(r) = -GMm \cdot /r$ (Derived in Unit 5)

Total energy $E = KE + PE$ invariance of IHO: $F(r) = -k \cdot r$ (Derived here)

 *Total energy $E = KE + PE$ invariance of Coulomb: $F(r) = -GMm/r^2$ (Derived in Unit 5)*

Kepler laws involve ∇ -momentum conservation in isotropic force $F(r)$

Now consider orbital energy conservation of the IHO: $F(r)=-k \cdot r$ with $U(r)=k \cdot r^2/2$

Total IHO energy= $KE + PE$ is constant

$$\begin{aligned}
 KE + PE &= \frac{1}{2} \mathbf{v} \cdot \mathbf{M} \cdot \mathbf{v} + \frac{1}{2} \mathbf{r} \cdot \mathbf{K} \cdot \mathbf{r} \\
 &= \frac{1}{2} \begin{pmatrix} v_x & v_y \end{pmatrix} \cdot \begin{pmatrix} m & 0 \\ 0 & m \end{pmatrix} \cdot \begin{pmatrix} v_x \\ v_y \end{pmatrix} + \begin{pmatrix} r_x & r_y \end{pmatrix} \cdot \begin{pmatrix} k & 0 \\ 0 & k \end{pmatrix} \cdot \begin{pmatrix} r_x \\ r_y \end{pmatrix} \\
 &= \frac{1}{2} m v_x^2 + \frac{1}{2} m v_y^2 + \frac{1}{2} k r_x^2 + \frac{1}{2} k r_y^2 \\
 &= \frac{1}{2} m (-a\omega \sin \omega t)^2 + \frac{1}{2} m (b\omega \cos \omega t)^2 + \frac{1}{2} k (a \cos \omega t)^2 + \frac{1}{2} k (b \sin \omega t)^2 \\
 &= \frac{1}{2} m a^2 \omega^2 (\sin^2 \omega t) + \frac{1}{2} m b^2 \omega^2 (\cos^2 \omega t) + \frac{1}{2} k a^2 (\cos^2 \omega t) + \frac{1}{2} k b^2 (\sin^2 \omega t) \\
 &= \frac{1}{2} m \omega^2 (a^2 + b^2) \quad \text{Given : } k = m\omega^2
 \end{aligned}$$

$$E = KE + PE = \frac{1}{2} m \omega^2 (a^2 + b^2) = \frac{1}{2} k (a^2 + b^2) \quad \text{since: } \omega = \sqrt{\frac{k}{m}} = \sqrt{G\rho_{\oplus} 4\pi / 3} \quad \text{or: } m\omega^2 = k$$

We'll see that the Coul. orbits are simpler:

(like the period...not a function of b)

Kepler laws involve ∇ -momentum conservation in isotropic force $F(r)$

Now consider orbital energy conservation of the IHO: $F(r) = -k \cdot r$ with $U(r) = k \cdot r^2 / 2$

Total IHO energy = KE + PE is constant

$$\begin{aligned}
 KE + PE &= \frac{1}{2} \mathbf{v} \cdot \mathbf{M} \cdot \mathbf{v} + \frac{1}{2} \mathbf{r} \cdot \mathbf{K} \cdot \mathbf{r} \\
 &= \frac{1}{2} \begin{pmatrix} v_x & v_y \end{pmatrix} \cdot \begin{pmatrix} m & 0 \\ 0 & m \end{pmatrix} \cdot \begin{pmatrix} v_x \\ v_y \end{pmatrix} + \begin{pmatrix} r_x & r_y \end{pmatrix} \cdot \begin{pmatrix} k & 0 \\ 0 & k \end{pmatrix} \cdot \begin{pmatrix} r_x \\ r_y \end{pmatrix} \\
 &= \frac{1}{2} m v_x^2 + \frac{1}{2} m v_y^2 + \frac{1}{2} k r_x^2 + \frac{1}{2} k r_y^2 \\
 &= \frac{1}{2} m (-a\omega \sin \omega t)^2 + \frac{1}{2} m (b\omega \cos \omega t)^2 + \frac{1}{2} k (a \cos \omega t)^2 + \frac{1}{2} k (b \sin \omega t)^2 \\
 &= \frac{1}{2} m a^2 \omega^2 (\sin^2 \omega t) + \frac{1}{2} m b^2 \omega^2 (\cos^2 \omega t) + \frac{1}{2} k a^2 (\cos^2 \omega t) + \frac{1}{2} k b^2 (\sin^2 \omega t) \\
 &= \frac{1}{2} m \omega^2 (a^2 + b^2) \quad \text{Given : } k = m\omega^2
 \end{aligned}$$

$$E = KE + PE = \frac{1}{2} m \omega^2 (a^2 + b^2) = \frac{1}{2} k (a^2 + b^2) \quad \text{since: } \omega = \sqrt{\frac{k}{m}} = \sqrt{G \rho_{\oplus} 4\pi / 3} \quad \text{or: } m\omega^2 = k$$

We'll see that the Coul. orbits are simpler:

(like the period...not a function of b)

$$E = KE + PE = \frac{1}{2} m v_x^2 + \frac{1}{2} m v_y^2 - \frac{k}{r} = \frac{1}{2} m v_x^2 + \frac{1}{2} m v_y^2 - \frac{GM_{\oplus} m}{r} = -\frac{GM_{\oplus} m}{a}$$

- Introduction to dual matrix operator contact geometry (based on IHO orbits)
- Quadratic form ellipse $\mathbf{r} \cdot \mathbf{Q} \cdot \mathbf{r} = 1$ vs. inverse form ellipse $\mathbf{p} \cdot \mathbf{Q}^{-1} \cdot \mathbf{p} = 1$
 - Duality norm relations ($\mathbf{r} \cdot \mathbf{p} = 1$)
 - \mathbf{Q} -Ellipse tangents \mathbf{r}' normal to dual \mathbf{Q}^{-1} -ellipse position \mathbf{p} ($\mathbf{r}' \cdot \mathbf{p} = 0 = \mathbf{r} \cdot \mathbf{p}'$)
 - Operator geometric sequences and eigenvectors
 - Alternative scaling of matrix operator geometry
 - Vector calculus of tensor operation

Quadratic forms and tangent contact geometry of their ellipses

A matrix Q that generates an ellipse by $\mathbf{r} \bullet Q \bullet \mathbf{r} = 1$ is called positive-definite (if $\mathbf{r} \bullet Q \bullet \mathbf{r}$ always > 0)

$$\begin{pmatrix} x & y \end{pmatrix} \bullet \overbrace{\begin{pmatrix} \frac{1}{a^2} & 0 \\ 0 & \frac{1}{b^2} \end{pmatrix}}^{\mathbf{r} \bullet Q \bullet \mathbf{r}} \bullet \begin{pmatrix} x \\ y \end{pmatrix} = 1 = \overbrace{\begin{pmatrix} x & y \end{pmatrix}}^{\mathbf{r}} \bullet \overbrace{\begin{pmatrix} \frac{x}{a^2} \\ \frac{y}{b^2} \end{pmatrix}}^{Q \bullet \mathbf{r}} = \frac{x^2}{a^2} + \frac{y^2}{b^2}$$

A inverse matrix Q^{-1} generates an ellipse by $\mathbf{p} \bullet Q^{-1} \bullet \mathbf{p} = 1$ called inverse or dual ellipse:

$$\begin{pmatrix} p_x & p_y \end{pmatrix} \bullet \overbrace{\begin{pmatrix} a^2 & 0 \\ 0 & b^2 \end{pmatrix}}^{\mathbf{p} \bullet Q^{-1} \bullet \mathbf{p}} \bullet \begin{pmatrix} p_x \\ p_y \end{pmatrix} = 1 = \overbrace{\begin{pmatrix} p_x & p_y \end{pmatrix}}^{\mathbf{p}} \bullet \overbrace{\begin{pmatrix} a^2 p_x \\ b^2 p_y \end{pmatrix}}^{Q^{-1} \bullet \mathbf{p}} = a^2 p_x^2 + b^2 p_y^2$$

Quadratic forms and tangent contact geometry of their ellipses

A matrix Q that generates an ellipse by $\mathbf{r} \bullet Q \bullet \mathbf{r} = 1$ is called positive-definite (if $\mathbf{r} \bullet Q \bullet \mathbf{r}$ always > 0)

$$\left(x \quad y \right) \bullet \overbrace{\begin{pmatrix} \frac{1}{a^2} & 0 \\ 0 & \frac{1}{b^2} \end{pmatrix}}^{\mathbf{r} \bullet Q \bullet \mathbf{r}} \bullet \begin{pmatrix} x \\ y \end{pmatrix} = 1 = \overbrace{\begin{pmatrix} x & y \end{pmatrix}}^{\mathbf{r}} \bullet \overbrace{\begin{pmatrix} \frac{x}{a^2} \\ \frac{y}{b^2} \end{pmatrix}}^{Q \bullet \mathbf{r} = \mathbf{p}} = \frac{x^2}{a^2} + \frac{y^2}{b^2}$$

Defined mapping between ellipses

A inverse matrix Q^{-1} generates an ellipse by $\mathbf{p} \bullet Q^{-1} \bullet \mathbf{p} = 1$ called inverse or dual ellipse:

$$\left(p_x \quad p_y \right) \bullet \overbrace{\begin{pmatrix} a^2 & 0 \\ 0 & b^2 \end{pmatrix}}^{\mathbf{p} \bullet Q^{-1} \bullet \mathbf{p}} \bullet \begin{pmatrix} p_x \\ p_y \end{pmatrix} = 1 = \overbrace{\begin{pmatrix} p_x & p_y \end{pmatrix}}^{\mathbf{p}} \bullet \overbrace{\begin{pmatrix} a^2 p_x \\ b^2 p_y \end{pmatrix}}^{Q^{-1} \bullet \mathbf{p} = \mathbf{r}} = a^2 p_x^2 + b^2 p_y^2$$

Introduction to dual matrix operator contact geometry (based on IHO orbits)

→ *Quadratic form ellipse $\mathbf{r} \cdot \mathbf{Q} \cdot \mathbf{r} = 1$ vs. inverse form ellipse $\mathbf{p} \cdot \mathbf{Q}^{-1} \cdot \mathbf{p} = 1$*

Duality norm relations ($\mathbf{r} \cdot \mathbf{p} = 1$)

\mathbf{Q} -Ellipse tangents \mathbf{r}' normal to dual \mathbf{Q}^{-1} -ellipse position \mathbf{p} ($\mathbf{r}' \cdot \mathbf{p} = 0 = \mathbf{r} \cdot \mathbf{p}'$)

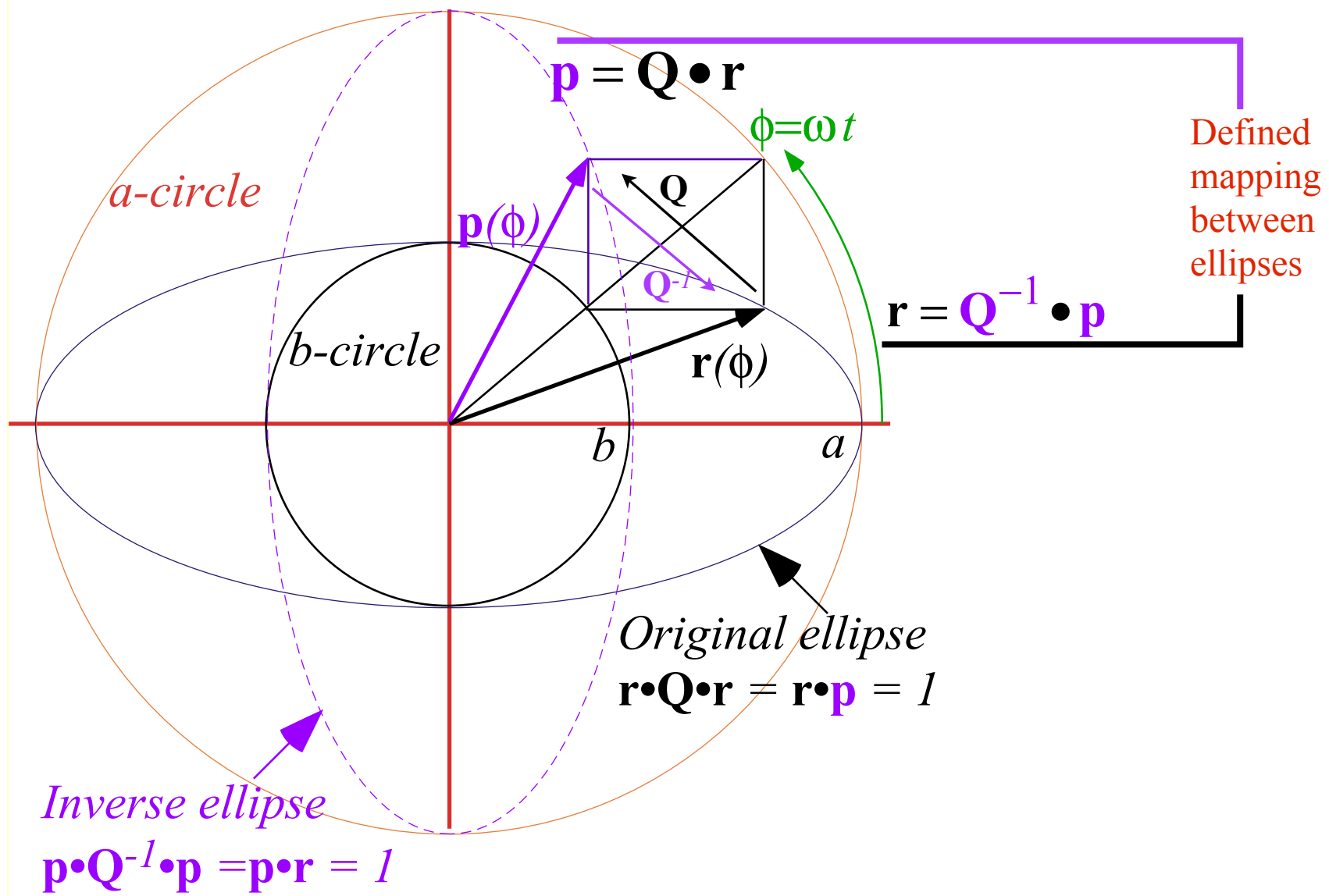
Operator geometric sequences and eigenvectors

Alternative scaling of matrix operator geometry

Vector calculus of tensor operation

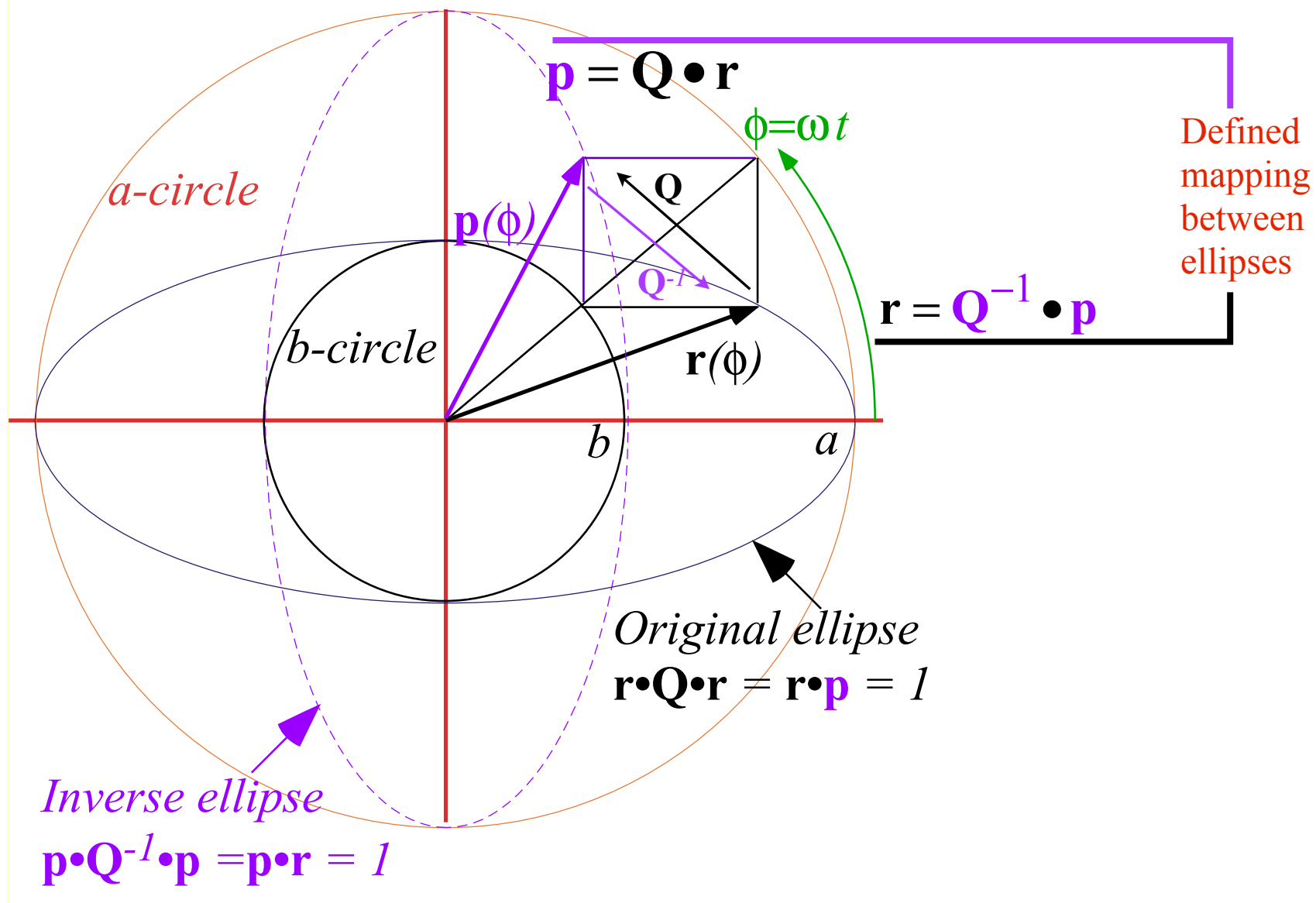
(a) Quadratic form ellipse and
Inverse quadratic form ellipse

based on
Unit 1
Fig. 11.6



(a) Quadratic form ellipse and
Inverse quadratic form ellipse

based on
Unit 1
Fig. 11.6



Here plot of \mathbf{p} -ellipse is re-scaled by scalefactor $S = a \cdot b$

\mathbf{p} -ellipse x -radius = $1/a$ plotted at: $S(1/a) = b$ ($=1$ for $a=2, b=1$)

\mathbf{p} -ellipse y -radius = $1/b$ plotted at: $S(1/b) = a$ ($=2$ for $a=2, b=1$)

Introduction to dual matrix operator geometry (based on IHO orbits)

Quadratic form ellipse $\mathbf{r} \cdot \mathbf{Q} \cdot \mathbf{r} = 1$ vs. inverse form ellipse $\mathbf{p} \cdot \mathbf{Q}^{-1} \cdot \mathbf{p} = 1$



Duality norm relations ($\mathbf{r} \cdot \mathbf{p} = 1$)

\mathbf{Q} -Ellipse tangents \mathbf{r}' normal to dual \mathbf{Q}^{-1} -ellipse position \mathbf{p} ($\mathbf{r}' \cdot \mathbf{p} = 0 = \mathbf{r} \cdot \mathbf{p}'$)

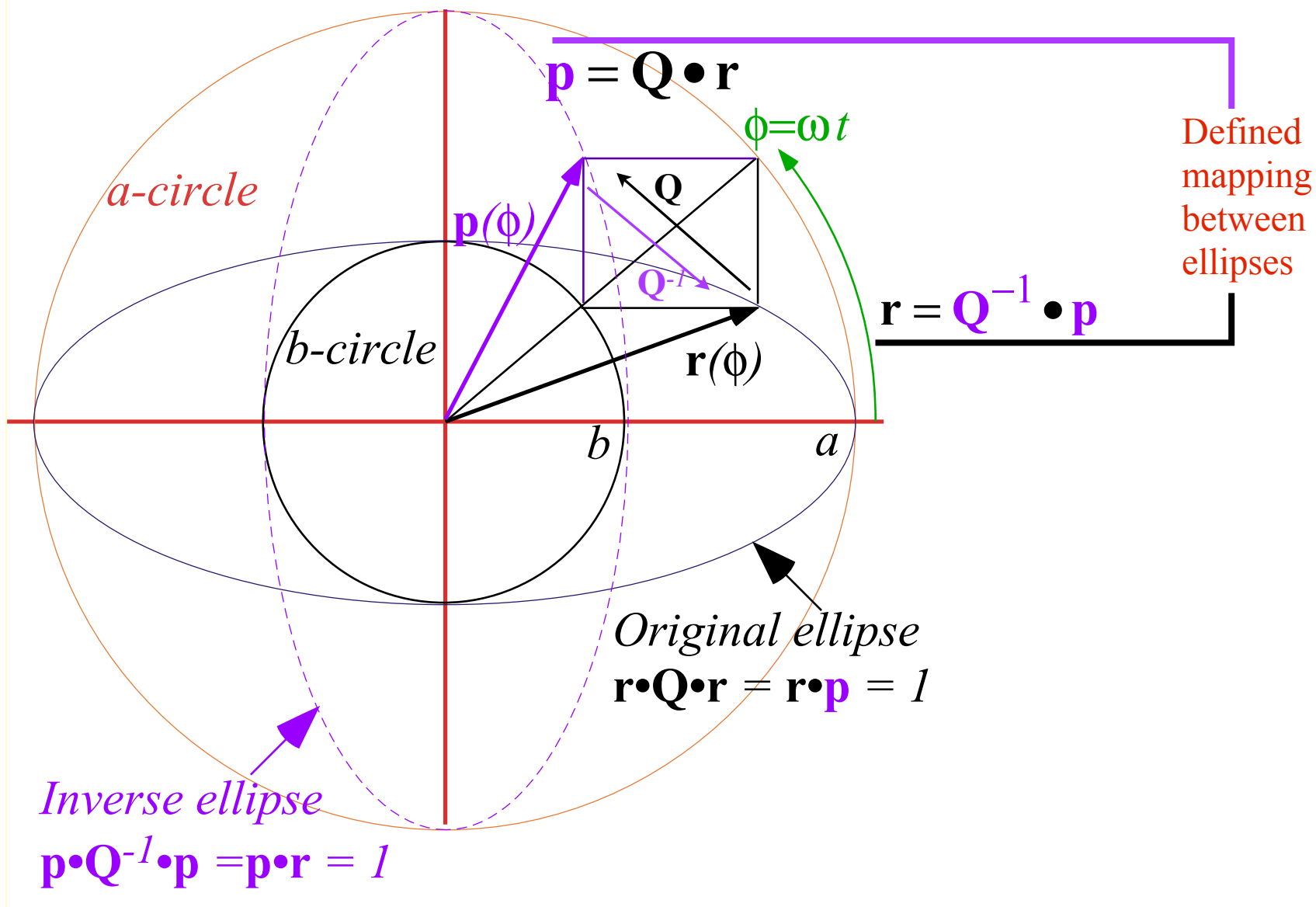
Operator geometric sequences and eigenvectors

Alternative scaling of matrix operator geometry

Vector calculus of tensor operation

(a) Quadratic form ellipse and
Inverse quadratic form ellipse

based on
Unit 1
Fig. 11.6



Quadratic form $\mathbf{r} \cdot \mathbf{Q} \cdot \mathbf{r} = 1$ has mutual duality relations with inverse form $\mathbf{p} \cdot \mathbf{Q}^{-1} \cdot \mathbf{p} = 1 = \mathbf{p} \cdot \mathbf{r}$

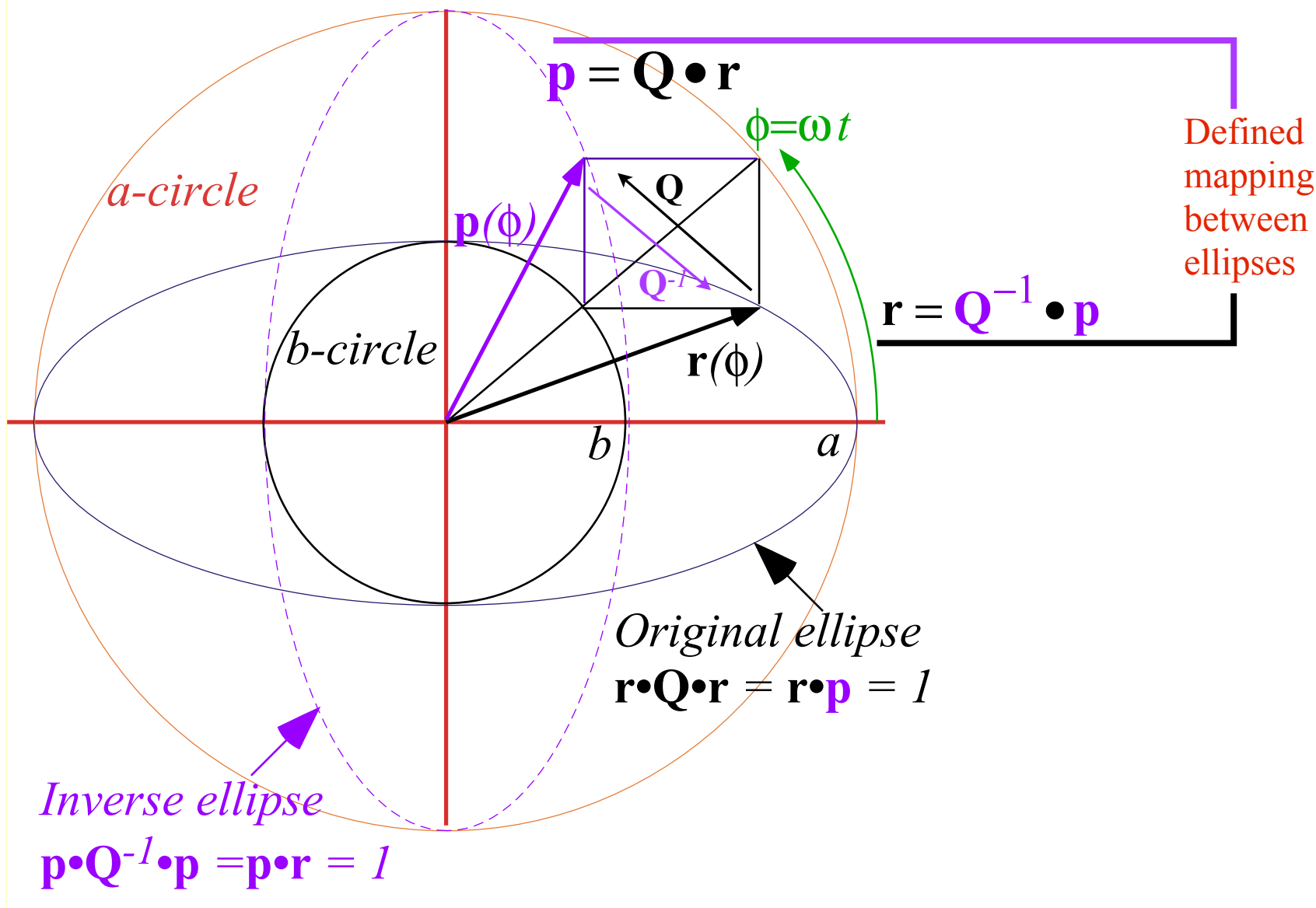
Here plot of \mathbf{p} -ellipse is re-scaled by scalefactor $S = a \cdot b$

\mathbf{p} -ellipse x -radius = $1/a$ plotted at: $S(1/a) = b$ ($=1$ for $a=2, b=1$)

\mathbf{p} -ellipse y -radius = $1/b$ plotted at: $S(1/b) = a$ ($=2$ for $a=2, b=1$)

(a) Quadratic form ellipse and Inverse quadratic form ellipse

based on
Unit 1
Fig. 11.6



Quadratic form $\mathbf{r} \cdot \mathbf{Q} \cdot \mathbf{r} = 1$ has mutual duality relations with inverse form $\mathbf{p} \cdot \mathbf{Q}^{-1} \cdot \mathbf{p} = 1 = \mathbf{p} \cdot \mathbf{r}$

$$\mathbf{p} = \mathbf{Q} \cdot \mathbf{r} = \begin{pmatrix} \overbrace{1/a^2}^{\mathbf{Q}} & 0 \\ 0 & \overbrace{1/b^2}^{\mathbf{Q}} \end{pmatrix} \cdot \begin{pmatrix} \overbrace{x}^{\mathbf{r}} \\ \overbrace{y}^{\mathbf{r}} \end{pmatrix} = \begin{pmatrix} \overbrace{x/a^2}^{\mathbf{p}} \\ \overbrace{y/b^2}^{\mathbf{p}} \end{pmatrix} = \begin{pmatrix} (1/a)\cos\phi \\ (1/b)\sin\phi \end{pmatrix} \text{ where: } \begin{matrix} x = r_x = a \cos\phi = a \cos\omega t \\ y = r_y = b \sin\phi = b \sin\omega t \end{matrix} \text{ so: } \boxed{\mathbf{p} \cdot \mathbf{r} = 1}$$

Here plot of \mathbf{p} -ellipse is re-scaled by scalefactor $S = a \cdot b$

\mathbf{p} -ellipse x -radius $= 1/a$ plotted at: $S(1/a) = b$ ($= 1$ for $a = 2, b = 1$)

\mathbf{p} -ellipse y -radius $= 1/b$ plotted at: $S(1/b) = a$ ($= 2$ for $a = 2, b = 1$)

[Link \$\Rightarrow\$ BoxIt simulation of IHO orbits](#)

[Link \$\rightarrow\$ IHO orbital time rates of change](#)

[Link \$\rightarrow\$ IHO Exegesis Plot](#)

Introduction to dual matrix operator geometry (based on IHO orbits)

Quadratic form ellipse $\mathbf{r} \cdot \mathbf{Q} \cdot \mathbf{r} = 1$ vs. inverse form ellipse $\mathbf{p} \cdot \mathbf{Q}^{-1} \cdot \mathbf{p} = 1$

Duality norm relations ($\mathbf{r} \cdot \mathbf{p} = 1$)

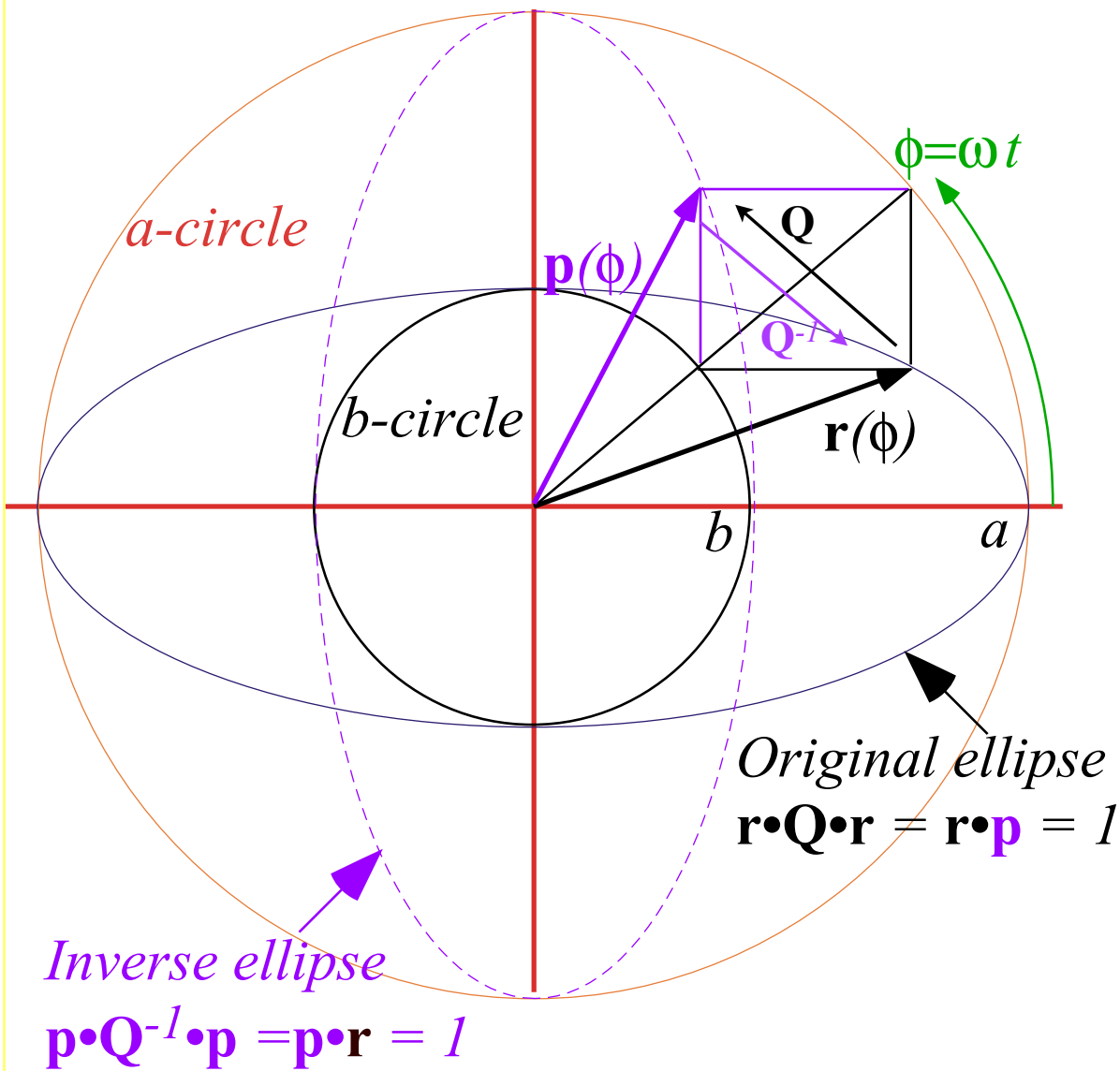
 *Q -Ellipse tangents \mathbf{r}' normal to dual Q^{-1} -ellipse position \mathbf{p} ($\mathbf{r}' \cdot \mathbf{p} = 0 = \mathbf{r} \cdot \mathbf{p}'$)*

Operator geometric sequences and eigenvectors

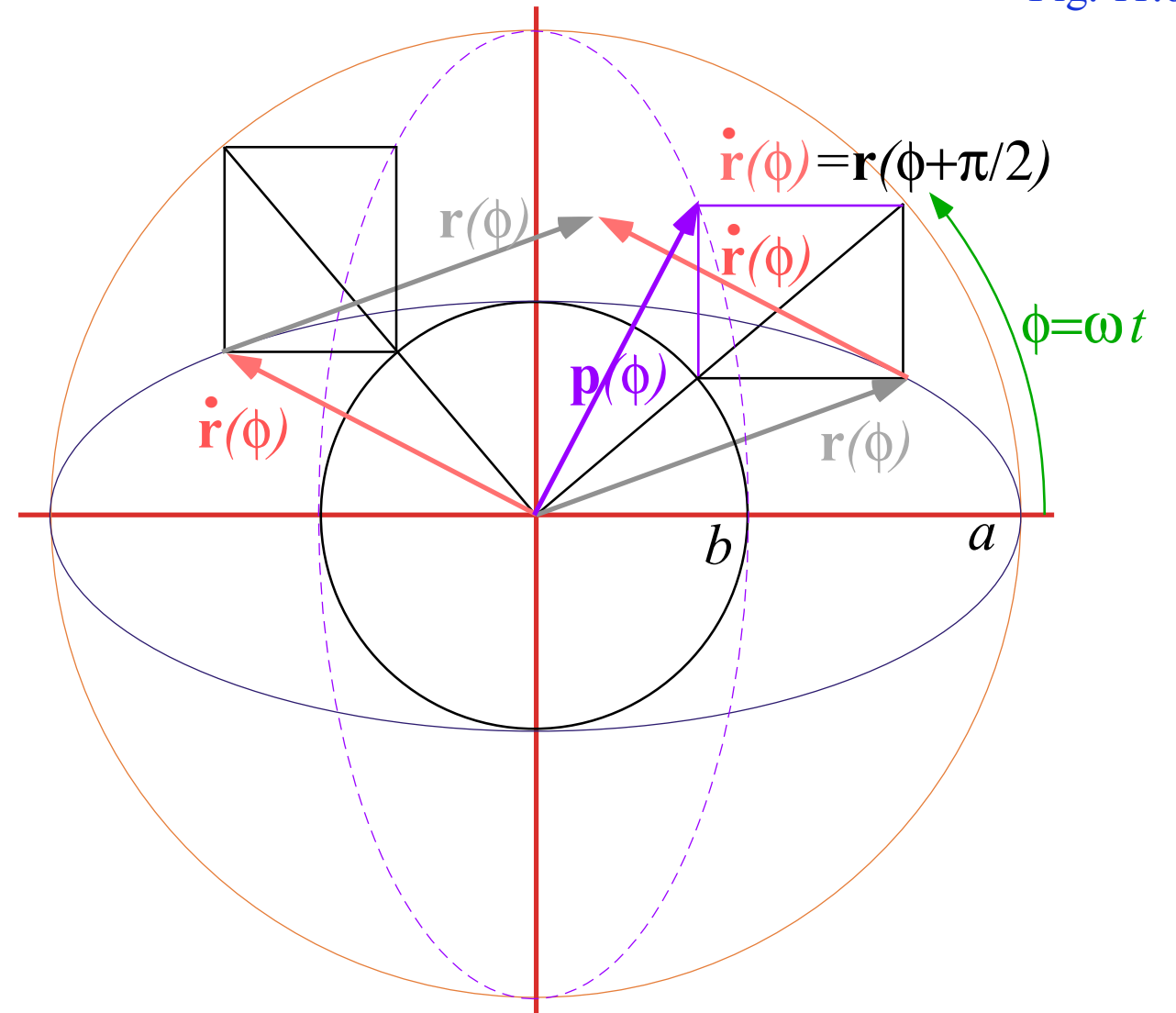
Alternative scaling of matrix operator geometry

Vector calculus of tensor operation

(a) Quadratic form ellipse and Inverse quadratic form ellipse



(b) Ellipse tangents



Quadratic form $\mathbf{r} \cdot \mathbf{Q} \cdot \mathbf{r} = 1$ has mutual duality relations with inverse form $\mathbf{p} \cdot \mathbf{Q}^{-1} \cdot \mathbf{p} = 1 = \mathbf{p} \cdot \mathbf{r}$

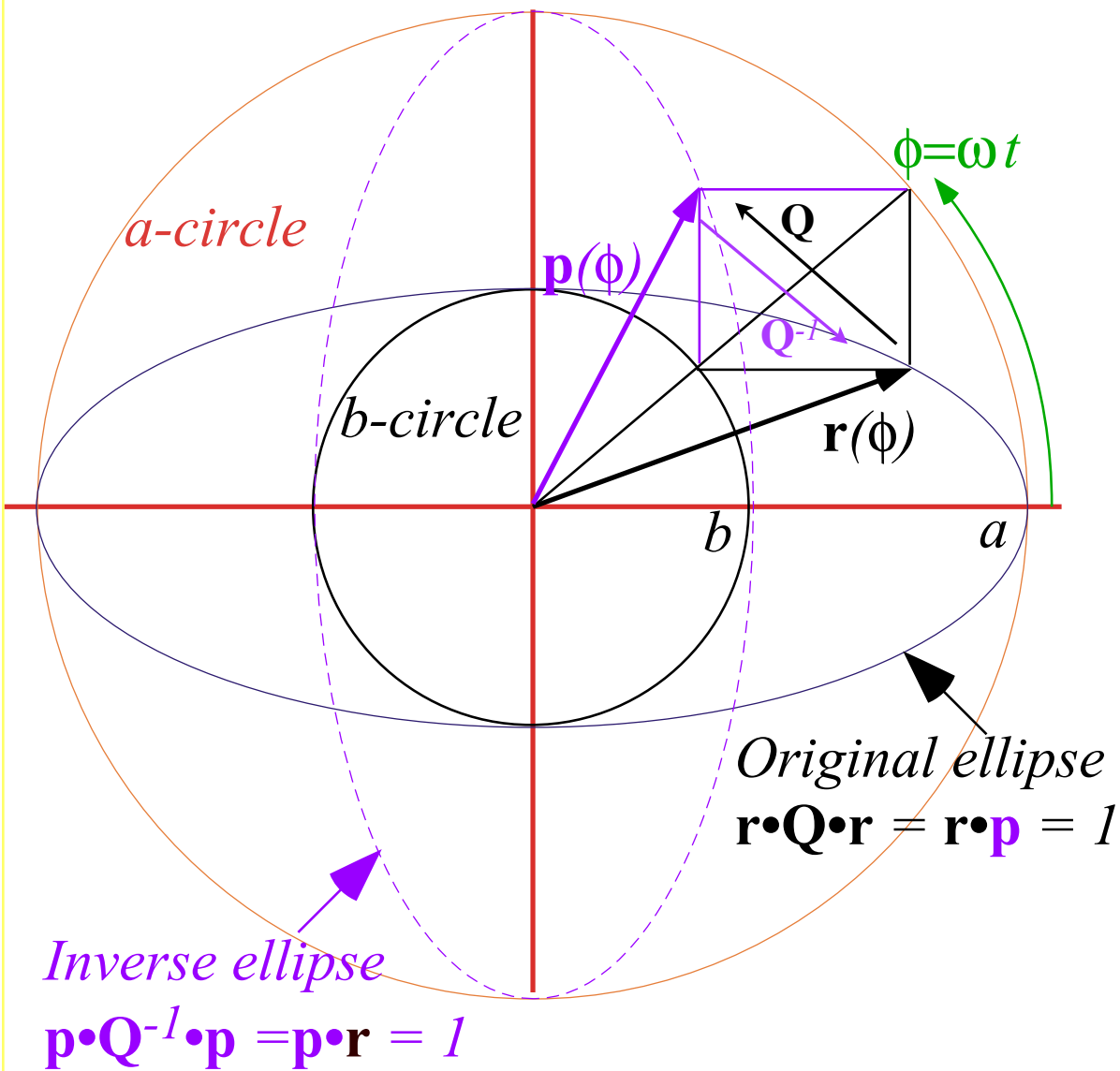
$$\mathbf{p} = \mathbf{Q} \cdot \mathbf{r} = \begin{pmatrix} 1/a^2 & 0 \\ 0 & 1/b^2 \end{pmatrix} \cdot \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} x/a^2 \\ y/b^2 \end{pmatrix} = \begin{pmatrix} (1/a)\cos\phi \\ (1/b)\sin\phi \end{pmatrix} \quad \text{where:} \quad \begin{matrix} x = r_x = a \cos\phi = a \cos\omega t \\ y = r_y = b \sin\phi = b \sin\omega t \end{matrix} \quad \text{so: } \boxed{\mathbf{p} \cdot \mathbf{r} = 1}$$

Here plot of \mathbf{p} -ellipse is re-scaled by scalefactor $S = a \cdot b$

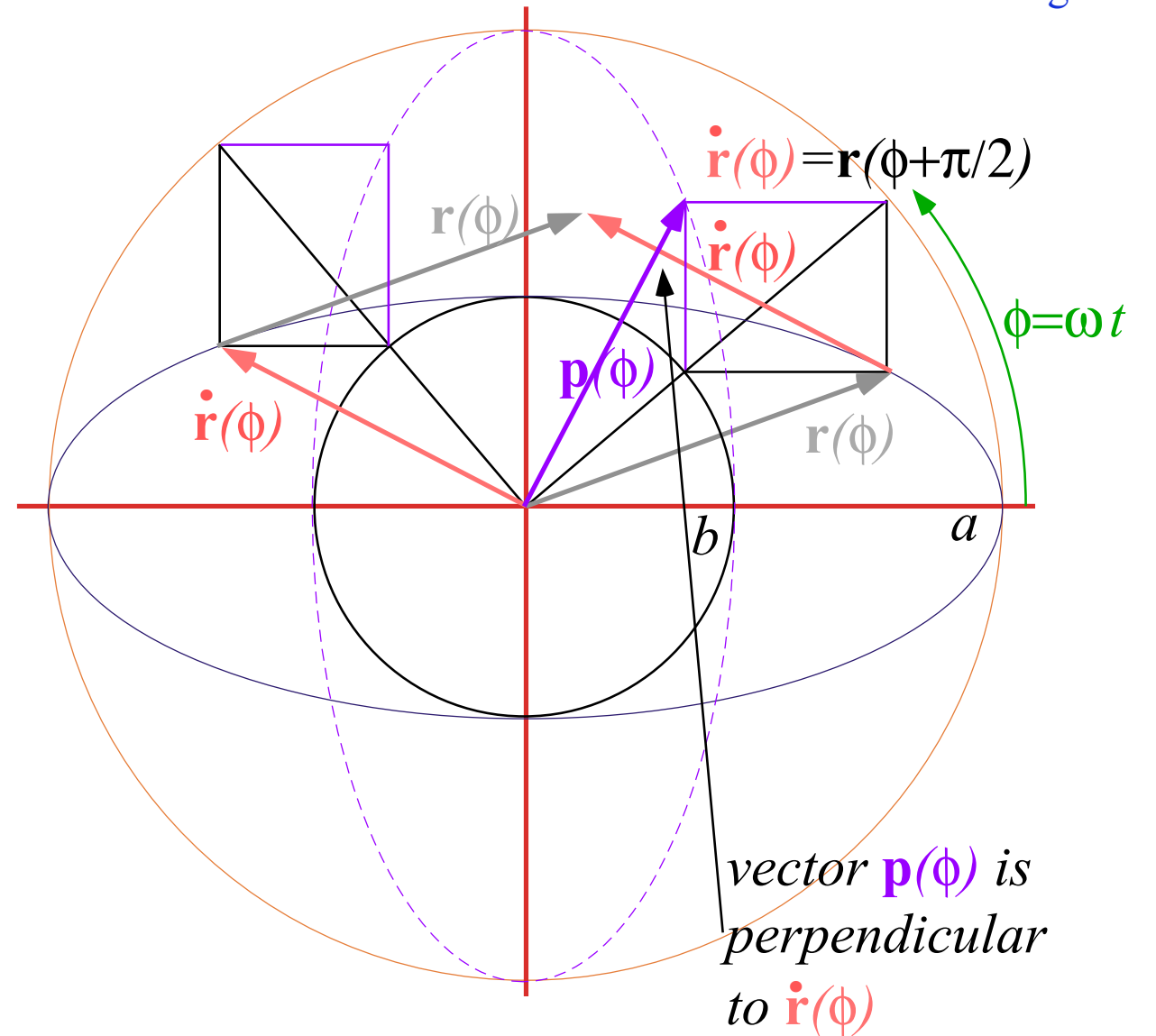
\mathbf{p} -ellipse x -radius $= 1/a$ plotted at: $S(1/a) = b$ ($= 1$ for $a=2, b=1$)

\mathbf{p} -ellipse y -radius $= 1/b$ plotted at: $S(1/b) = a$ ($= 2$ for $a=2, b=1$)

(a) Quadratic form ellipse and Inverse quadratic form ellipse



(b) Ellipse tangents



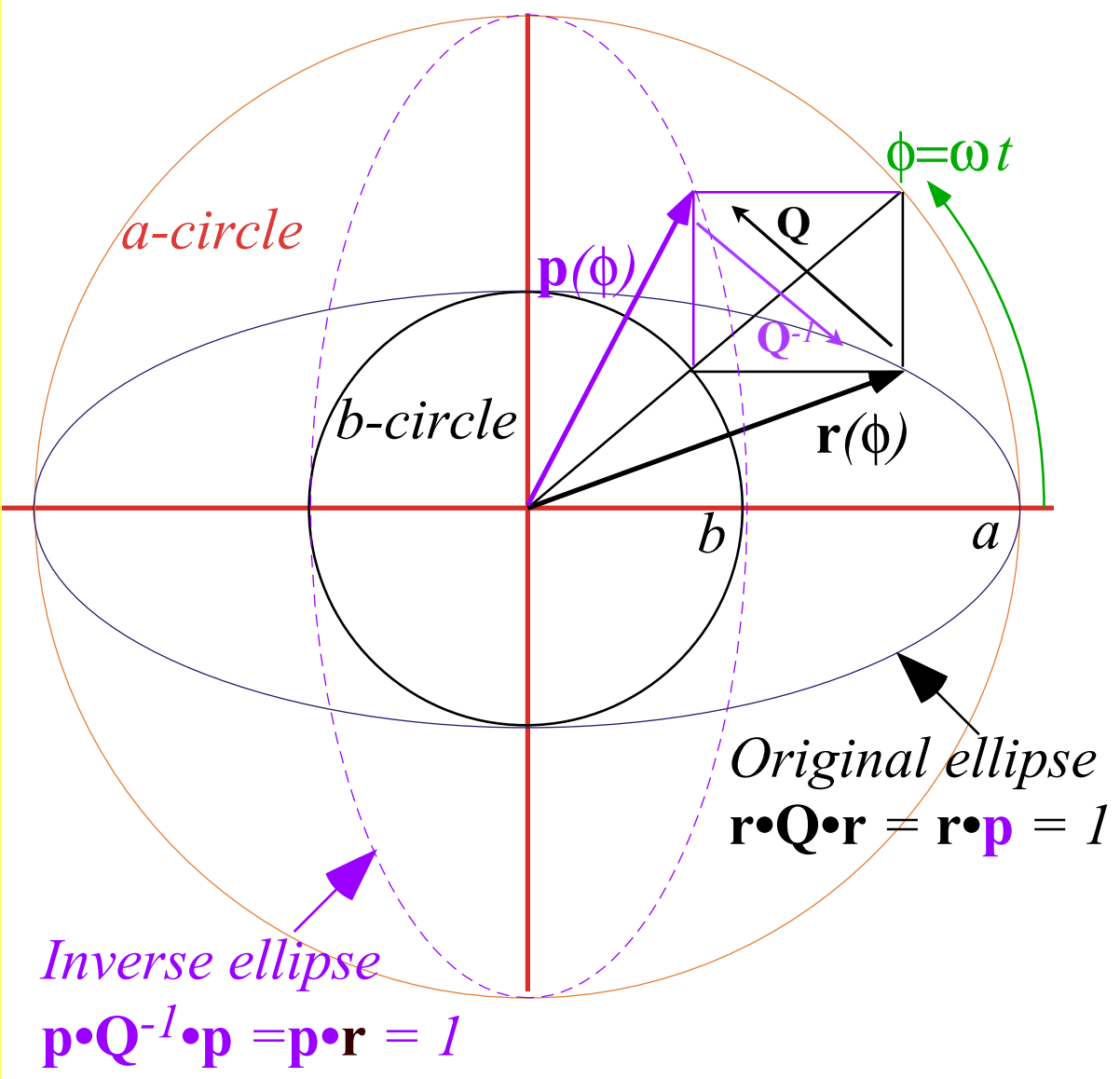
Quadratic form $\mathbf{r} \cdot \mathbf{Q} \cdot \mathbf{r} = 1$ has mutual duality relations with inverse form $\mathbf{p} \cdot \mathbf{Q}^{-1} \cdot \mathbf{p} = 1 = \mathbf{p} \cdot \mathbf{r}$

$$\mathbf{p} = \mathbf{Q} \cdot \mathbf{r} = \begin{pmatrix} 1/a^2 & 0 \\ 0 & 1/b^2 \end{pmatrix} \cdot \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} x/a^2 \\ y/b^2 \end{pmatrix} = \begin{pmatrix} (1/a)\cos\phi \\ (1/b)\sin\phi \end{pmatrix} \quad \text{where:} \quad \begin{matrix} x = r_x = a \cos\phi = a \cos\omega t \\ y = r_y = b \sin\phi = b \sin\omega t \end{matrix} \quad \text{so: } \boxed{\mathbf{p} \cdot \mathbf{r} = 1}$$

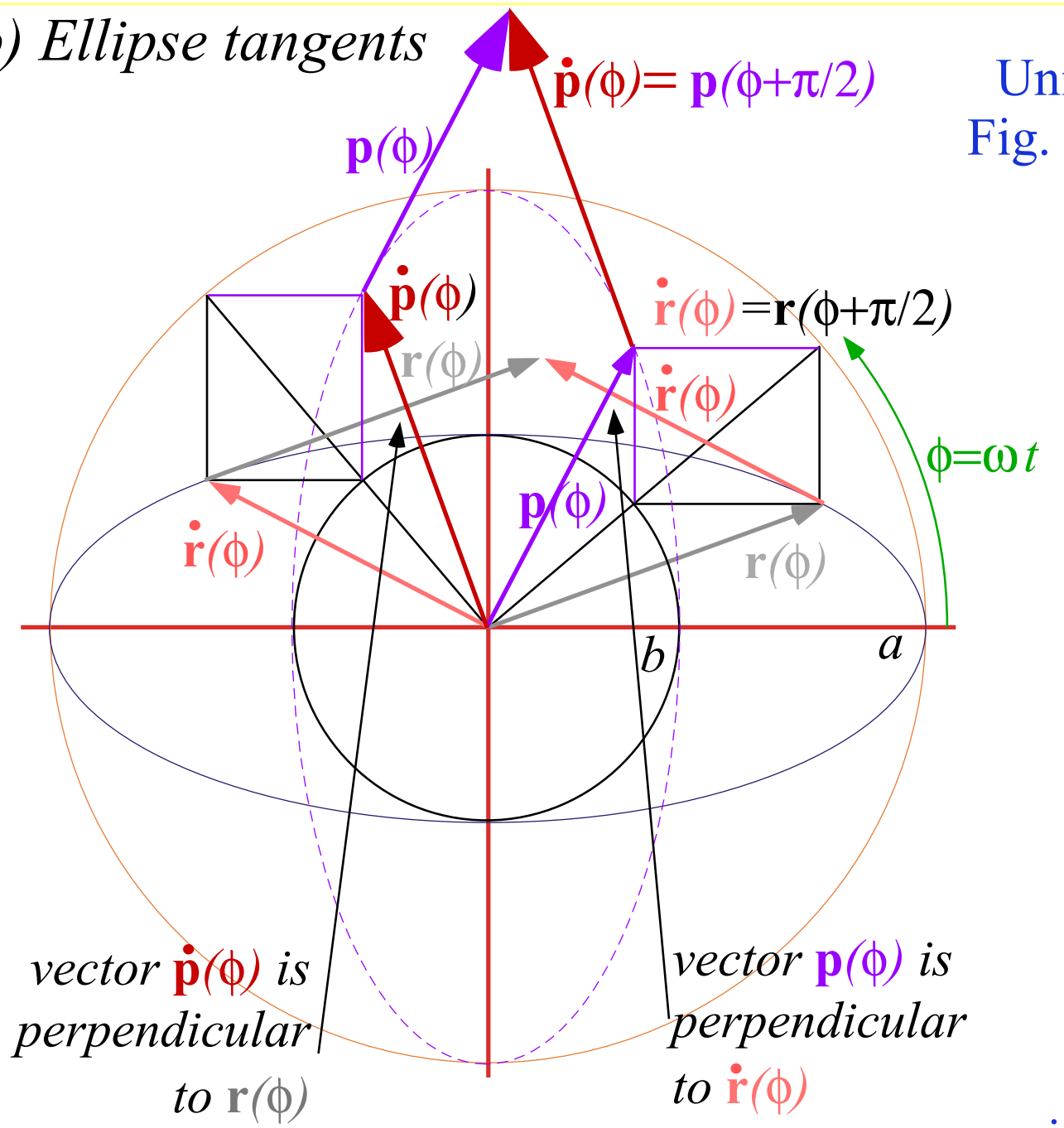
\mathbf{p} is perpendicular to velocity $\mathbf{v} = \dot{\mathbf{r}}$, a mutual orthogonality

$$\boxed{\dot{\mathbf{r}} \cdot \mathbf{p} = 0} = \begin{pmatrix} \dot{r}_x & \dot{r}_y \end{pmatrix} \cdot \begin{pmatrix} p_x \\ p_y \end{pmatrix} = \begin{pmatrix} -a \sin\phi & b \cos\phi \end{pmatrix} \cdot \begin{pmatrix} (1/a)\cos\phi \\ (1/b)\sin\phi \end{pmatrix} \quad \text{where:} \quad \begin{matrix} \dot{r}_x = -a \sin\phi \\ \dot{r}_y = b \cos\phi \end{matrix} \quad \text{and:} \quad \begin{matrix} p_x = (1/a)\cos\phi \\ p_y = (1/b)\sin\phi \end{matrix}$$

(a) Quadratic form ellipse and Inverse quadratic form ellipse



(b) Ellipse tangents



Quadratic form $\mathbf{r} \cdot \mathbf{Q} \cdot \mathbf{r} = 1$ has mutual duality relations with inverse form $\mathbf{p} \cdot \mathbf{Q}^{-1} \cdot \mathbf{p} = 1$

unit mutual projection

$$\mathbf{p} = \mathbf{Q} \cdot \mathbf{r} = \begin{pmatrix} 1/a^2 & 0 \\ 0 & 1/b^2 \end{pmatrix} \cdot \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} x/a^2 \\ y/b^2 \end{pmatrix} = \begin{pmatrix} (1/a)\cos\phi \\ (1/b)\sin\phi \end{pmatrix} \text{ where: } \begin{matrix} x = r_x = a \cos\phi = a \cos\omega t \\ y = r_y = b \sin\phi = b \sin\omega t \end{matrix} \text{ so: } \boxed{\mathbf{p} \cdot \mathbf{r} = 1}$$

\mathbf{p} is perpendicular to velocity $\mathbf{v} = \dot{\mathbf{r}}$, a mutual orthogonality. So is \mathbf{r} perpendicular to $\dot{\mathbf{p}}$: $\boxed{\dot{\mathbf{p}} \cdot \mathbf{r} = 0}$

$$\boxed{\dot{\mathbf{r}} \cdot \mathbf{p} = 0} = \begin{pmatrix} \dot{r}_x & \dot{r}_y \end{pmatrix} \cdot \begin{pmatrix} p_x \\ p_y \end{pmatrix} = \begin{pmatrix} -a \sin\phi & b \cos\phi \end{pmatrix} \cdot \begin{pmatrix} (1/a)\cos\phi \\ (1/b)\sin\phi \end{pmatrix} \text{ where: } \begin{matrix} \dot{r}_x = -a \sin\phi \\ \dot{r}_y = b \cos\phi \end{matrix} \text{ and: } \begin{matrix} p_x = (1/a)\cos\phi \\ p_y = (1/b)\sin\phi \end{matrix}$$

Elliptical

Exegesis

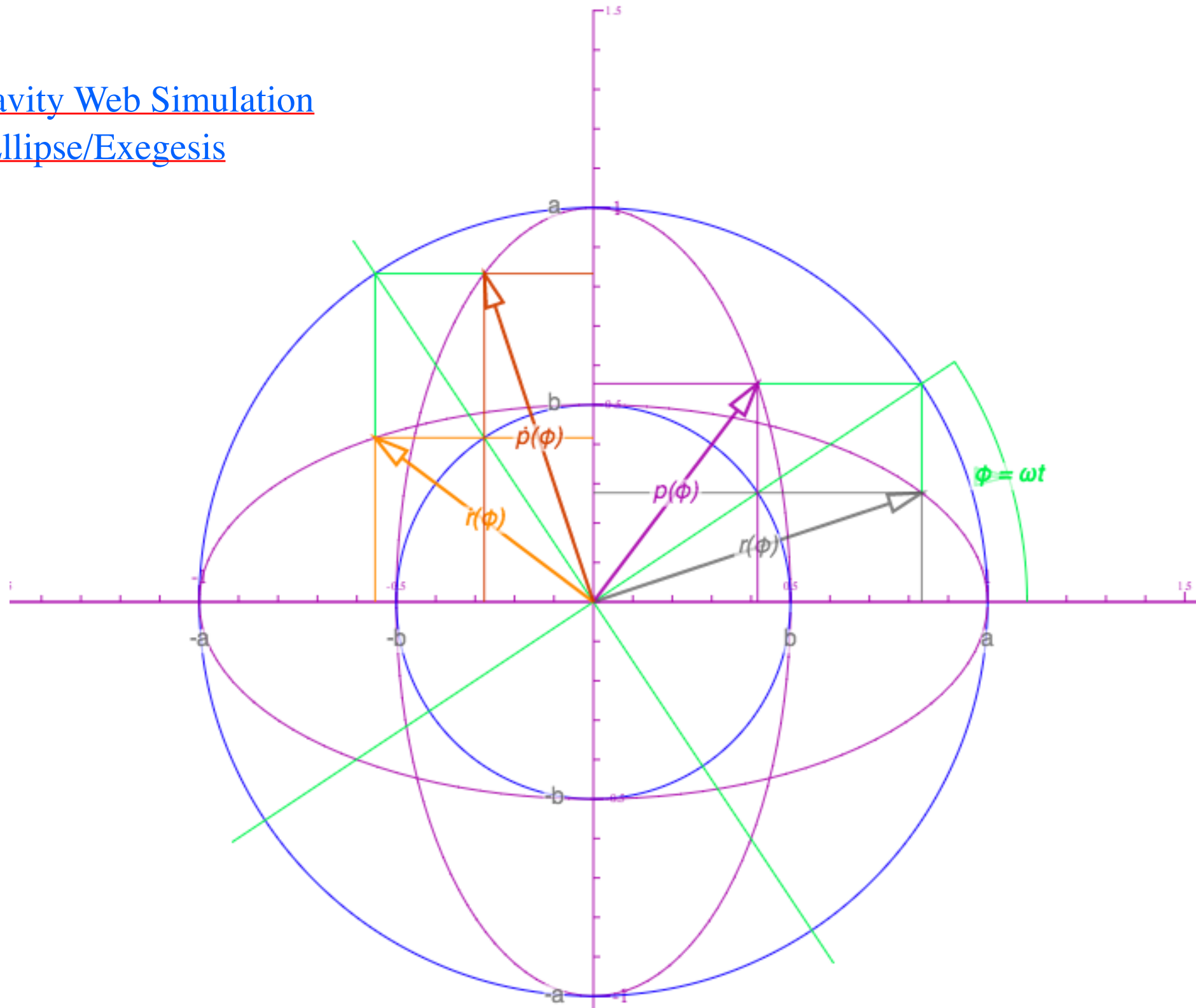
Controls

Contextual

Set ISM

User's Guide

[RelaWavity Web Simulation](#)
[Ellipse/Exegesis](#)



Geometry of dual ellipse Kepler anomalies for vectors $[\mathbf{r}(\phi), \mathbf{p}(\phi)]$ and $d/dt[\mathbf{r}(\phi), \mathbf{p}(\phi),]$ in coordinate (x,y) space rendered by animation web-app in RelaWavity and described in Lect. 12-advanced.

Introduction to dual matrix operator geometry (based on IHO orbits)

Quadratic form ellipse $\mathbf{r} \cdot \mathbf{Q} \cdot \mathbf{r} = 1$ vs. inverse form ellipse $\mathbf{p} \cdot \mathbf{Q}^{-1} \cdot \mathbf{p} = 1$

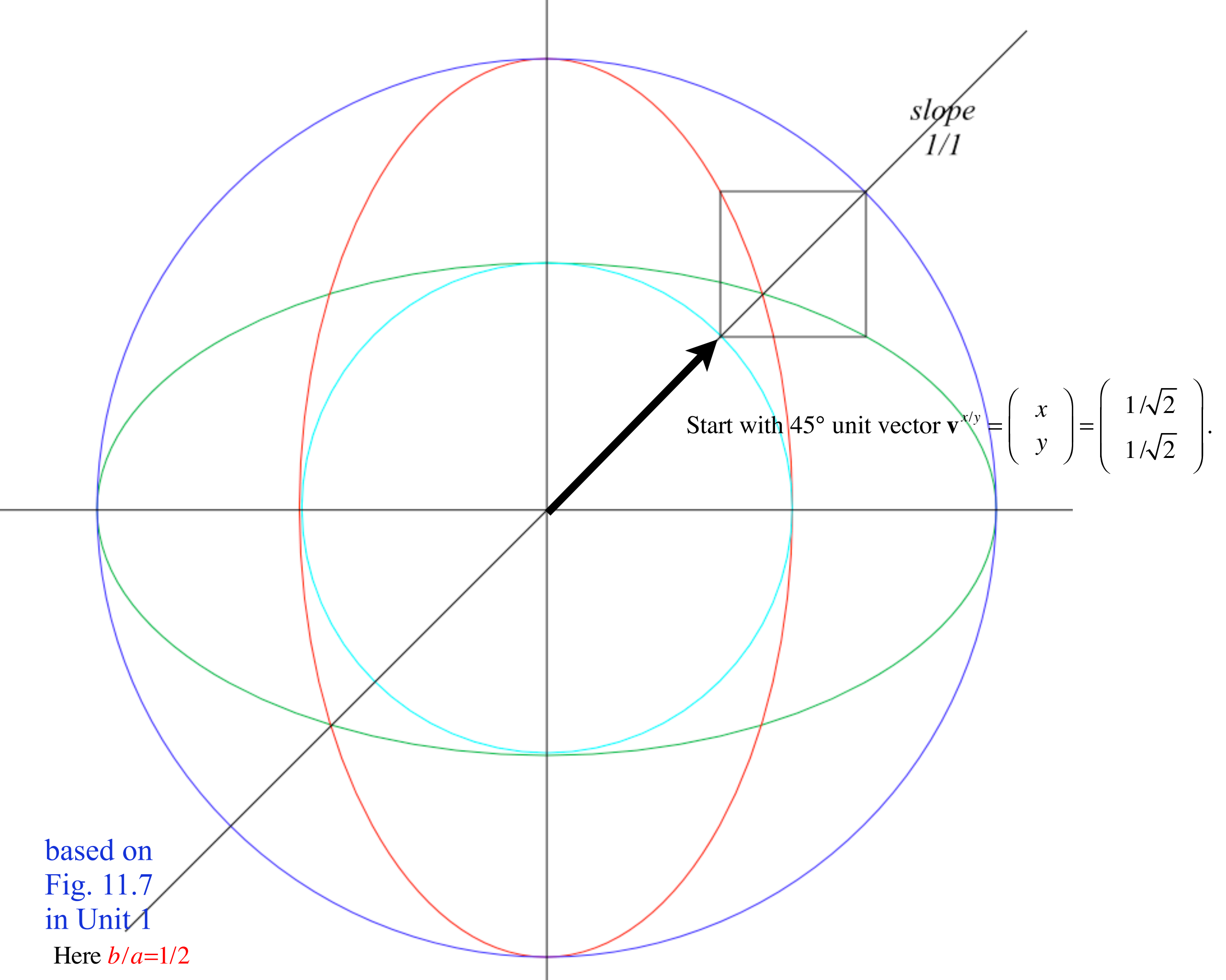
Duality norm relations ($\mathbf{r} \cdot \mathbf{p} = 1$)

Q -Ellipse tangents \mathbf{r}' normal to dual Q^{-1} -ellipse position \mathbf{p} ($\mathbf{r}' \cdot \mathbf{p} = 0 = \mathbf{r} \cdot \mathbf{p}'$)

 *Operator geometric sequences and eigenvectors*

Alternative scaling of matrix operator geometry

Vector calculus of tensor operation



Diagonal \mathbf{R} -matrix acts on vector $\mathbf{v}^{x/y}$.

Resulting vector has slope changed by factor $a/b=2$.

$$\mathbf{R} \cdot \mathbf{v}^{x/y} = \begin{pmatrix} 1/a & 0 \\ 0 & 1/b \end{pmatrix} \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} x/a \\ y/b \end{pmatrix}$$

(Slope increases if $a > b$.)

Action of "sqrt-" matrix $R = \sqrt{Q}$

slope a/b

slope $1/1$

slope b/a

Action of "sqrt⁻¹-" matrix $R^{-1} = \sqrt{Q^{-1}}$

Diagonal \mathbf{R}^{-1} -matrix acts on vector $\mathbf{v}^{x/y}$.

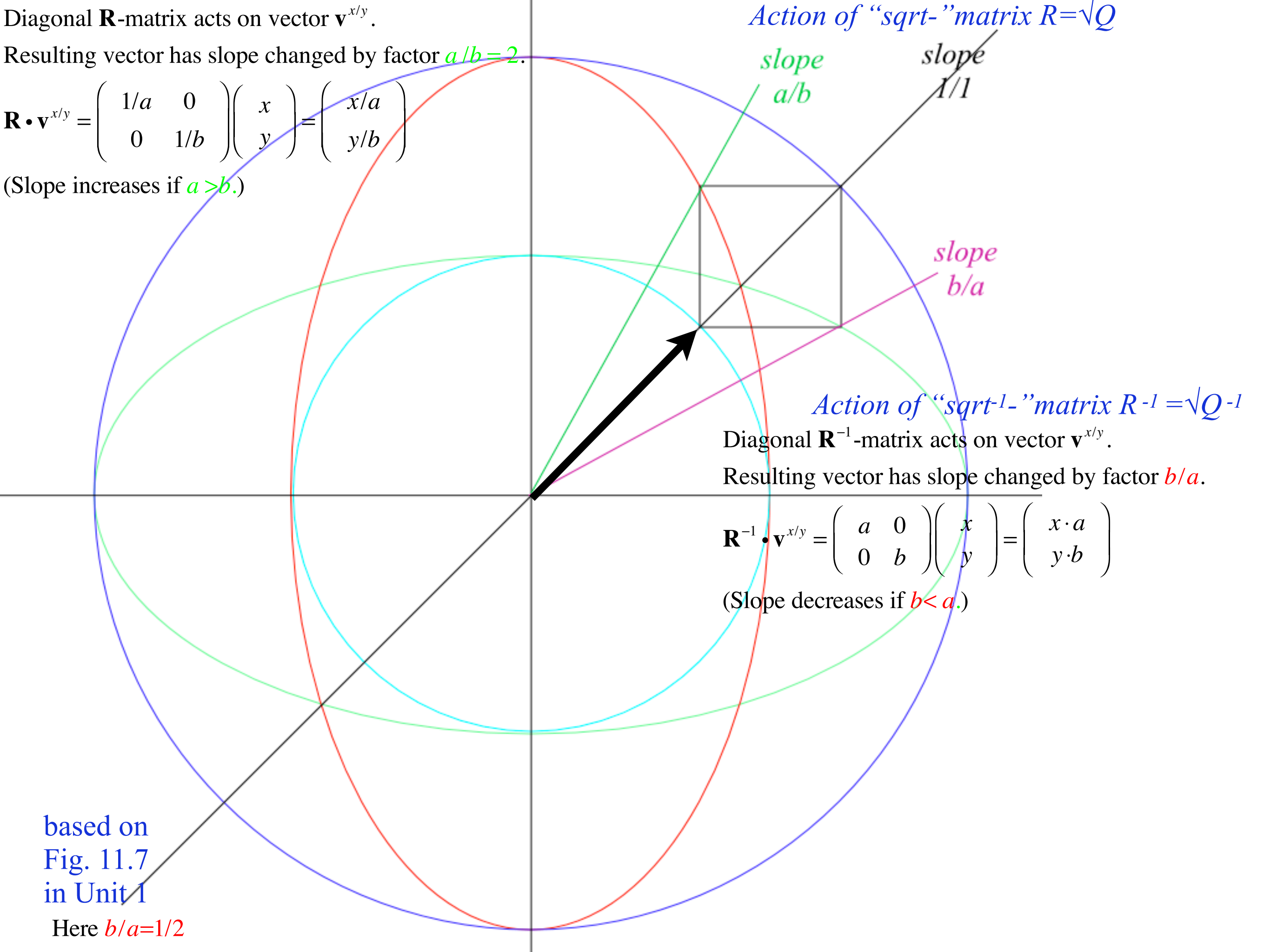
Resulting vector has slope changed by factor b/a .

$$\mathbf{R}^{-1} \cdot \mathbf{v}^{x/y} = \begin{pmatrix} a & 0 \\ 0 & b \end{pmatrix} \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} x \cdot a \\ y \cdot b \end{pmatrix}$$

(Slope decreases if $b < a$.)

based on
Fig. 11.7
in Unit 1

Here $b/a=1/2$



Diagonal \mathbf{R} -matrix acts on vector $\mathbf{v}^{x/y}$.

Resulting vector has slope changed by factor $a/b=2$.

$$\mathbf{R} \cdot \mathbf{v}^{x/y} = \begin{pmatrix} 1/a & 0 \\ 0 & 1/b \end{pmatrix} \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} x/a \\ y/b \end{pmatrix}$$

(It increases if $a > b$.)

Diagonal ($\mathbf{R}^2 = \mathbf{Q}$)-matrix acts on vector $\mathbf{v}^{x/y}$.

Resulting vector has slope changed by factor $a^2/b^2 = 4$.

$$\mathbf{Q} \cdot \mathbf{v}^{x/y} = \begin{pmatrix} 1/a^2 & 0 \\ 0 & 1/b^2 \end{pmatrix} \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} x/a^2 \\ y/b^2 \end{pmatrix}$$

(It increases if $a > b$.)

Action of "sqrt-" matrix $R = \sqrt{Q}$

slope a^2/b^2

slope a/b

slope $1/1$

slope b/a

slope b^2/a^2

Action of "sqrt-1-" matrix $R^{-1} = \sqrt{Q^{-1}}$

Diagonal \mathbf{R}^{-1} -matrix acts on vector $\mathbf{v}^{x/y}$.

Resulting vector has slope changed by factor $b/a=1/2$.

$$\mathbf{R}^{-1} \cdot \mathbf{v}^{x/y} = \begin{pmatrix} a & 0 \\ 0 & b \end{pmatrix} \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} x \cdot a \\ y \cdot b \end{pmatrix}$$

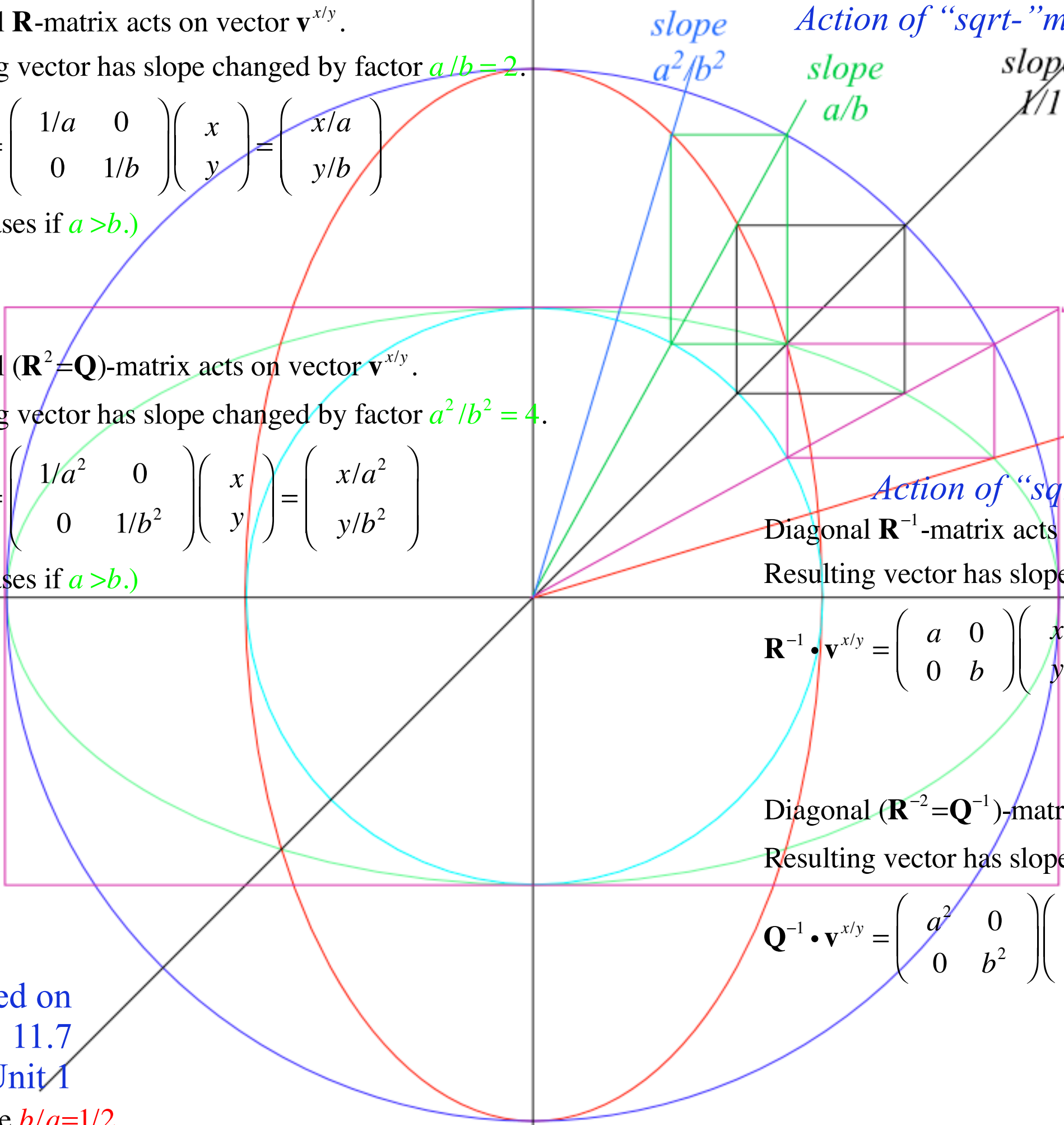
Diagonal ($\mathbf{R}^{-2} = \mathbf{Q}^{-1}$)-matrix acts on vector $\mathbf{v}^{x/y}$.

Resulting vector has slope changed by factor $b^2/a^2=1/4$.

$$\mathbf{Q}^{-1} \cdot \mathbf{v}^{x/y} = \begin{pmatrix} a^2 & 0 \\ 0 & b^2 \end{pmatrix} \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} x \cdot a^2 \\ y \cdot b^2 \end{pmatrix}$$

based on
Fig. 11.7
in Unit 1

Here $b/a=1/2$



Diagonal \mathbf{R} -matrix acts on vector $\mathbf{v}^{x/y}$.

Resulting vector has slope changed by factor $a/b = 2$.

$$\mathbf{R} \cdot \mathbf{v}^{x/y} = \begin{pmatrix} 1/a & 0 \\ 0 & 1/b \end{pmatrix} \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} x/a \\ y/b \end{pmatrix}$$

(It increases if $a > b$.)

Diagonal ($\mathbf{R}^2 = \mathbf{Q}$)-matrix acts on vector $\mathbf{v}^{x/y}$.

Resulting vector has slope changed by factor $a^2/b^2 = 4$.

$$\mathbf{Q} \cdot \mathbf{v}^{x/y} = \begin{pmatrix} 1/a^2 & 0 \\ 0 & 1/b^2 \end{pmatrix} \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} x/a^2 \\ y/b^2 \end{pmatrix}$$

(It increases if $a > b$.)

Either process can go on forever...

Diagonal ($\mathbf{R}^{2n} = \mathbf{Q}^n$)-matrix acts on vector $\mathbf{v}^{x/y}$.

Resulting vector has slope changed by factor $a^{2n}/b^{2n} = 4^n$.

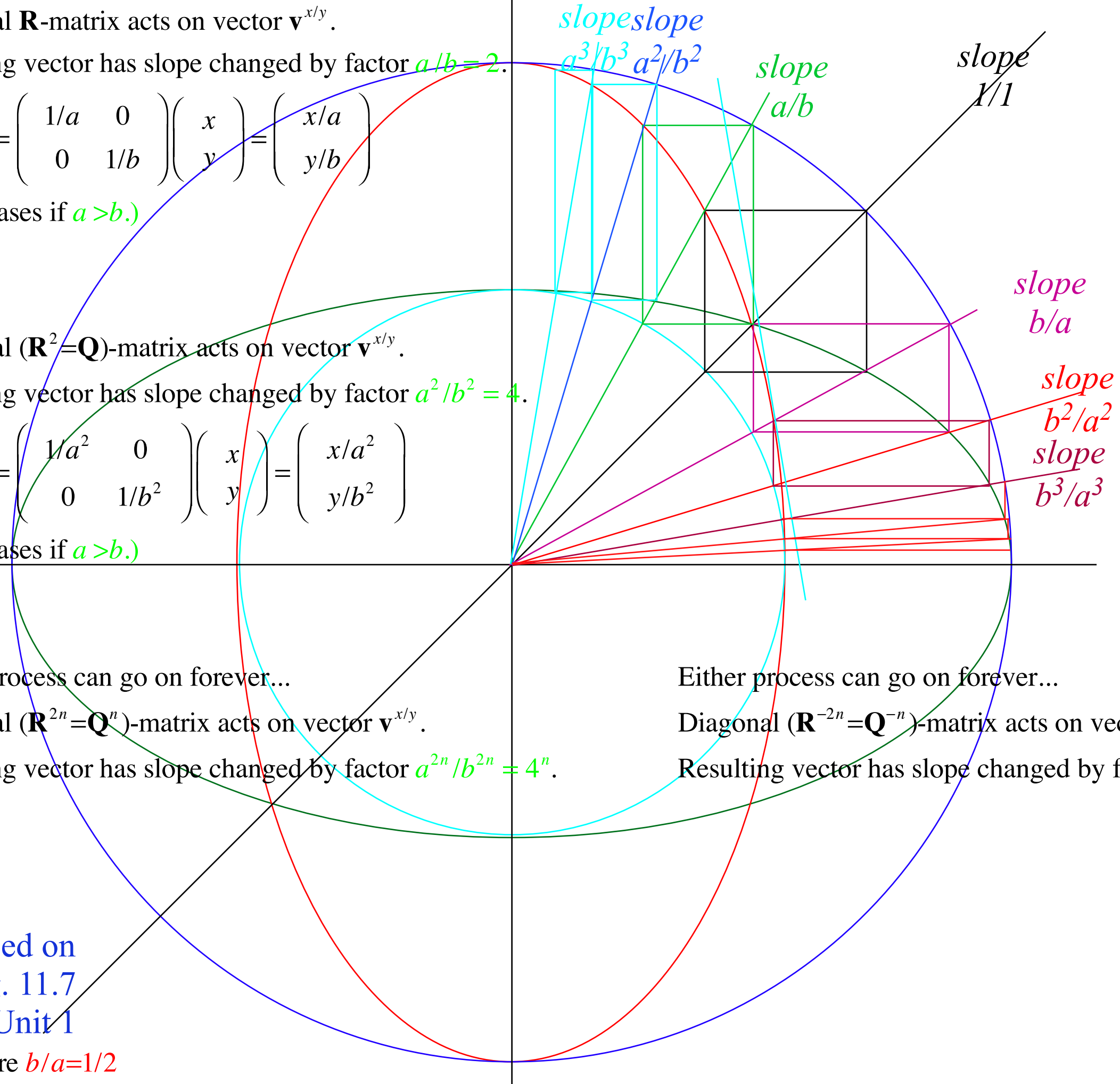
Either process can go on forever...

Diagonal ($\mathbf{R}^{-2n} = \mathbf{Q}^{-n}$)-matrix acts on vector $\mathbf{v}^{x/y}$.

Resulting vector has slope changed by factor $b^{2n}/a^{2n} = 4^{-n}$.

based on
Fig. 11.7
in Unit 1

Here $b/a = 1/2$



Diagonal \mathbf{R} -matrix acts on vector $\mathbf{v}^{x/y}$.

Resulting vector has slope changed by factor $a/b = 2$.

$$\mathbf{R} \cdot \mathbf{v}^{x/y} = \begin{pmatrix} 1/a & 0 \\ 0 & 1/b \end{pmatrix} \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} x/a \\ y/b \end{pmatrix}$$

(It increases if $a > b$.)

EIGENVECTOR

Diagonal ($\mathbf{R}^2 = \mathbf{Q}$)-matrix acts on vector $\mathbf{v}^{x/y}$.

Resulting vector has slope changed by factor $a^2/b^2 = 4$.

$$\mathbf{Q} \cdot \mathbf{v}^{x/y} = \begin{pmatrix} 1/a^2 & 0 \\ 0 & 1/b^2 \end{pmatrix} \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} x/a^2 \\ y/b^2 \end{pmatrix}$$

(It increases if $a > b$.)

EIGENVECTOR

Either process can go on forever...

Diagonal ($\mathbf{R}^{2n} = \mathbf{Q}^n$)-matrix acts on vector $\mathbf{v}^{x/y}$.

Resulting vector has slope changed by factor $a^{2n}/b^{2n} = 4^n$.

...Finally, the result approaches **EIGENVECTOR** $|y\rangle = \begin{pmatrix} 0 \\ 1 \end{pmatrix}$

of ∞ -slope which is "immune" to \mathbf{R} , \mathbf{Q} or \mathbf{Q}^n :

$$\mathbf{R}|y\rangle = (1/b)|y\rangle \quad \mathbf{Q}^n|y\rangle = (1/b^2)^n|y\rangle$$

Here $b/a = 1/2$

slopeslope

a^3/b^3 a^2/b^2

slope
 $/a/b$

slope
 $1/1$

slope
 b/a

slope
 b^2/a^2

slope
 b^3/a^3

Either process can go on forever...

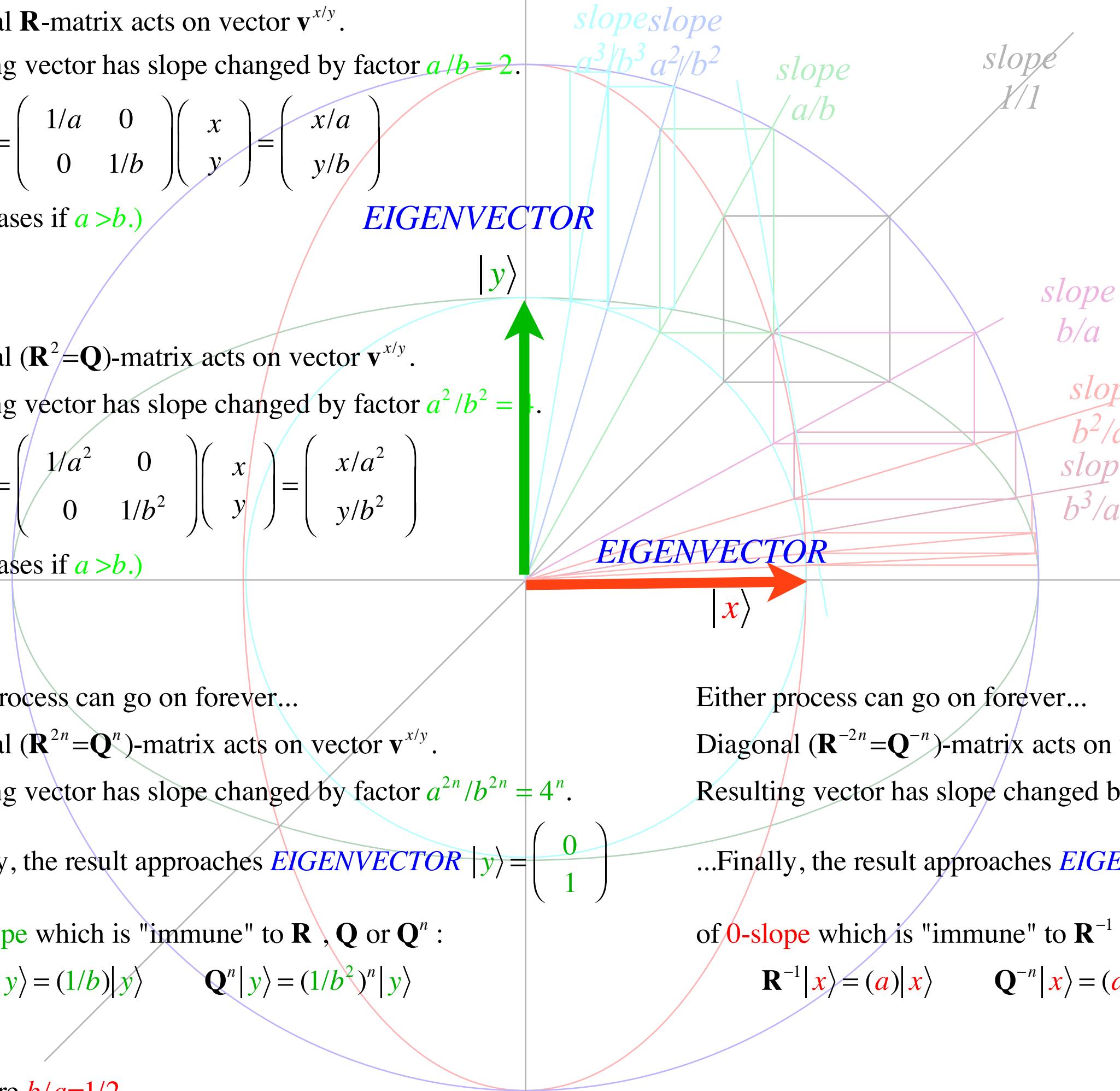
Diagonal ($\mathbf{R}^{-2n} = \mathbf{Q}^{-n}$)-matrix acts on vector $\mathbf{v}^{x/y}$.

Resulting vector has slope changed by factor $b^{2n}/a^{2n} = 4^{-n}$.

...Finally, the result approaches **EIGENVECTOR** $|x\rangle = \begin{pmatrix} 1 \\ 0 \end{pmatrix}$

of 0-slope which is "immune" to \mathbf{R}^{-1} , \mathbf{Q}^{-1} or \mathbf{Q}^{-n} :

$$\mathbf{R}^{-1}|x\rangle = (a)|x\rangle \quad \mathbf{Q}^{-n}|x\rangle = (a^2)^n|x\rangle$$



Diagonal \mathbf{R} -matrix acts on vector $\mathbf{v}^{x/y}$.

Resulting vector has slope changed by factor $a/b=2$.

$$\mathbf{R} \cdot \mathbf{v}^{x/y} = \begin{pmatrix} 1/a & 0 \\ 0 & 1/b \end{pmatrix} \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} x/a \\ y/b \end{pmatrix}$$

(It increases if $a > b$.)

EIGENVECTOR

Diagonal ($\mathbf{R}^2=\mathbf{Q}$)-matrix acts on vector $\mathbf{v}^{x/y}$.

Resulting vector has slope changed by factor $a^2/b^2=4$.

$$\mathbf{Q} \cdot \mathbf{v}^{x/y} = \begin{pmatrix} 1/a^2 & 0 \\ 0 & 1/b^2 \end{pmatrix} \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} x/a^2 \\ y/b^2 \end{pmatrix}$$

(It increases if $a > b$.)

EIGENVECTOR

Either process can go on forever...

Diagonal ($\mathbf{R}^{2n}=\mathbf{Q}^n$)-matrix acts on vector $\mathbf{v}^{x/y}$.

Resulting vector has slope changed by factor $a^{2n}/b^{2n}=4^n$.

...Finally, the result approaches **EIGENVECTOR** $|y\rangle = \begin{pmatrix} 0 \\ 1 \end{pmatrix}$

of ∞ -slope which is "immune" to \mathbf{R} , \mathbf{Q} or \mathbf{Q}^n :

$$\mathbf{R}|y\rangle = (1/b)|y\rangle \quad \mathbf{Q}^n|y\rangle = (1/b^2)^n|y\rangle$$

Eigenvalues

Eigensolution Relations

Either process can go on forever...

Diagonal ($\mathbf{R}^{-2n}=\mathbf{Q}^{-n}$)-matrix acts on vector $\mathbf{v}^{x/y}$.

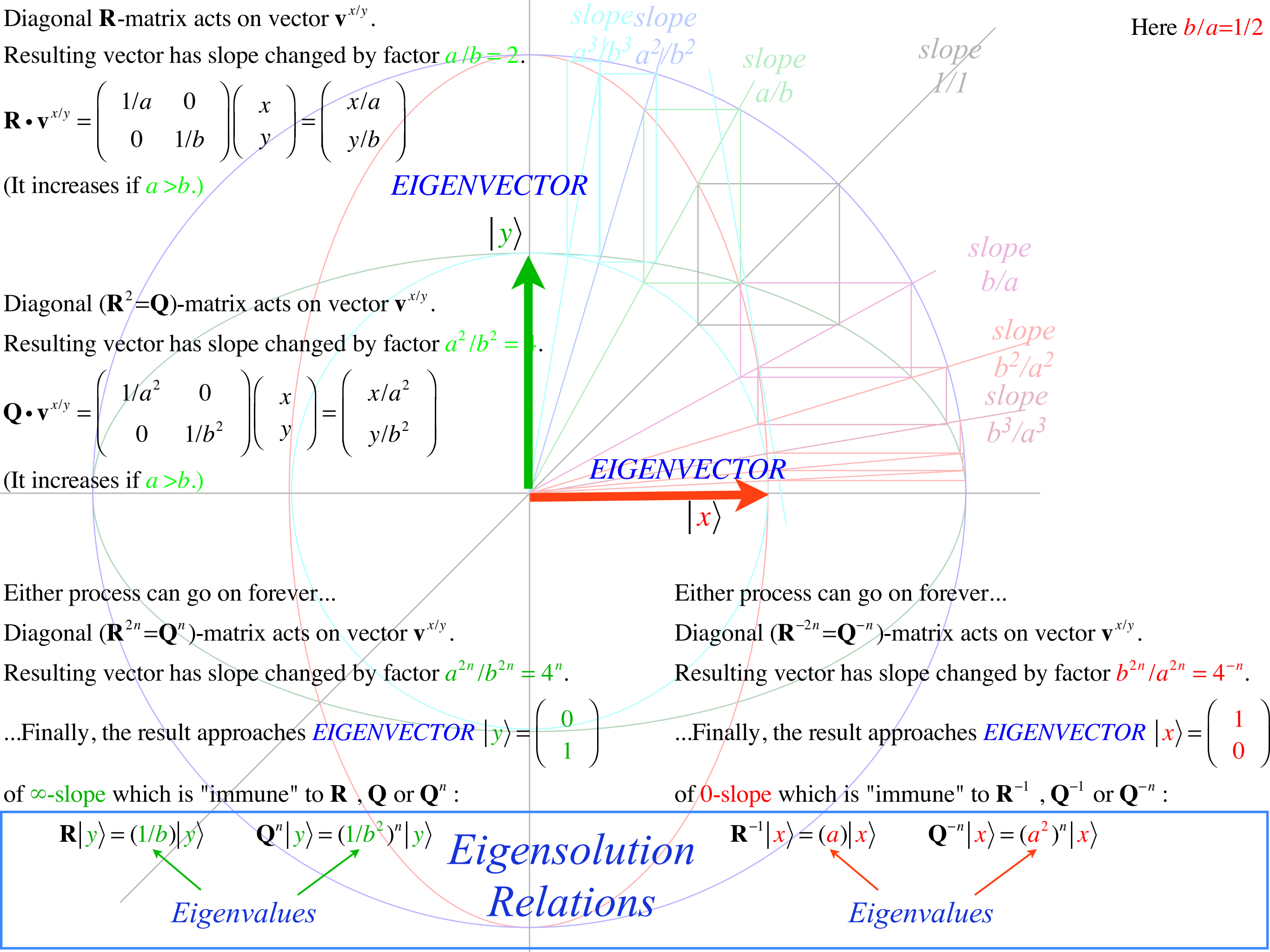
Resulting vector has slope changed by factor $b^{2n}/a^{2n}=4^{-n}$.

...Finally, the result approaches **EIGENVECTOR** $|x\rangle = \begin{pmatrix} 1 \\ 0 \end{pmatrix}$

of 0-slope which is "immune" to \mathbf{R}^{-1} , \mathbf{Q}^{-1} or \mathbf{Q}^{-n} :

$$\mathbf{R}^{-1}|x\rangle = (a)|x\rangle \quad \mathbf{Q}^{-n}|x\rangle = (a^2)^n|x\rangle$$

Eigenvalues



Introduction to dual matrix operator geometry (based on IHO orbits)

Quadratic form ellipse $\mathbf{r} \cdot \mathbf{Q} \cdot \mathbf{r} = 1$ vs. inverse form ellipse $\mathbf{p} \cdot \mathbf{Q}^{-1} \cdot \mathbf{p} = 1$

Duality norm relations ($\mathbf{r} \cdot \mathbf{p} = 1$)

Q -Ellipse tangents \mathbf{r}' normal to dual Q^{-1} -ellipse position \mathbf{p} ($\mathbf{r}' \cdot \mathbf{p} = 0 = \mathbf{r} \cdot \mathbf{p}'$)

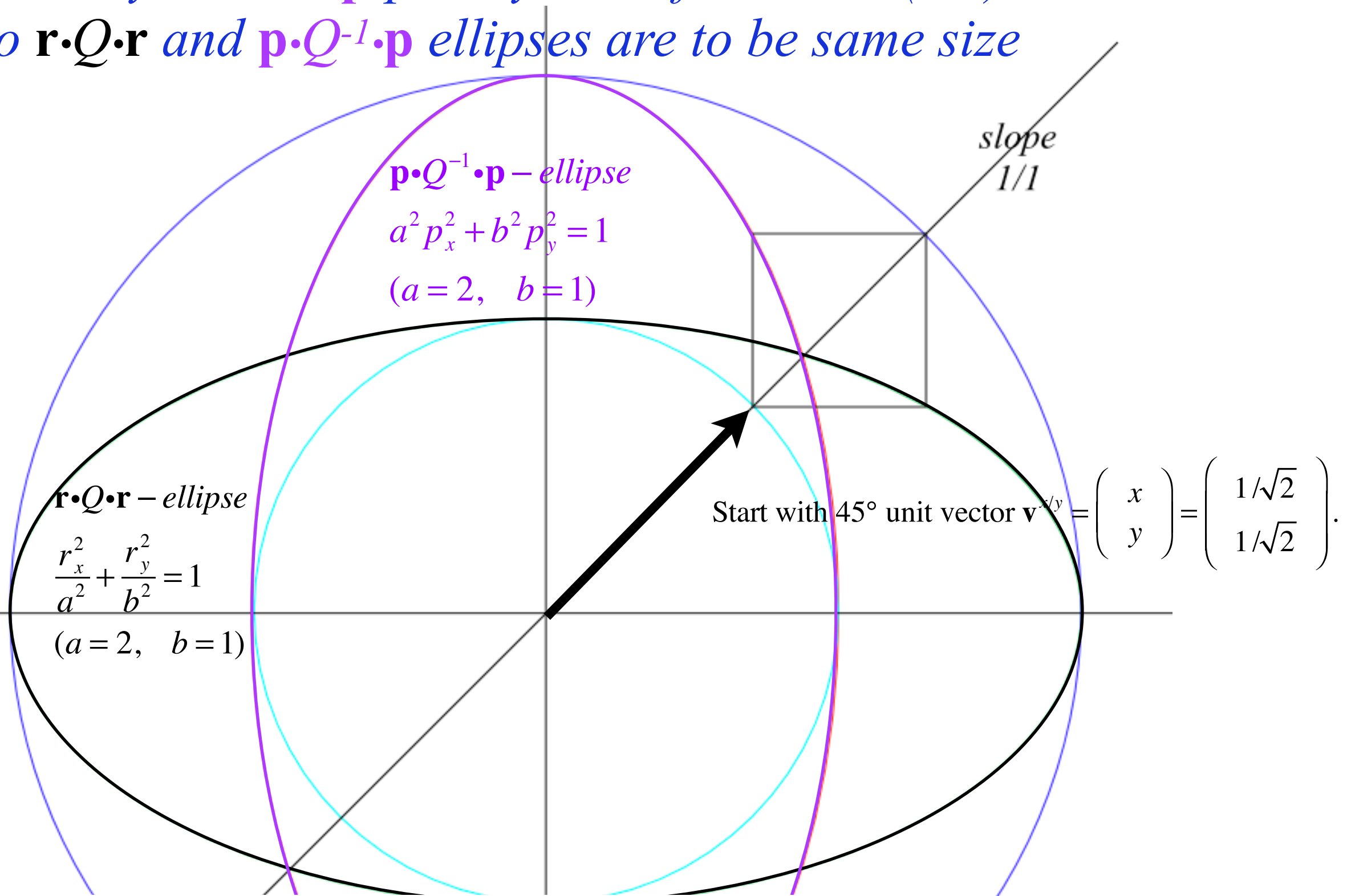
Operator geometric sequences and eigenvectors



Alternative scaling of matrix operator geometry

Vector calculus of tensor operation

You may rescale **p**-plot by scale factor $S=(a \cdot b)$
 so $\mathbf{r} \cdot \mathbf{Q} \cdot \mathbf{r}$ and $\mathbf{p} \cdot \mathbf{Q}^{-1} \cdot \mathbf{p}$ ellipses are to be same size



Here plot of **p**-ellipse is re-scaled by scalefactor $S=a \cdot b$

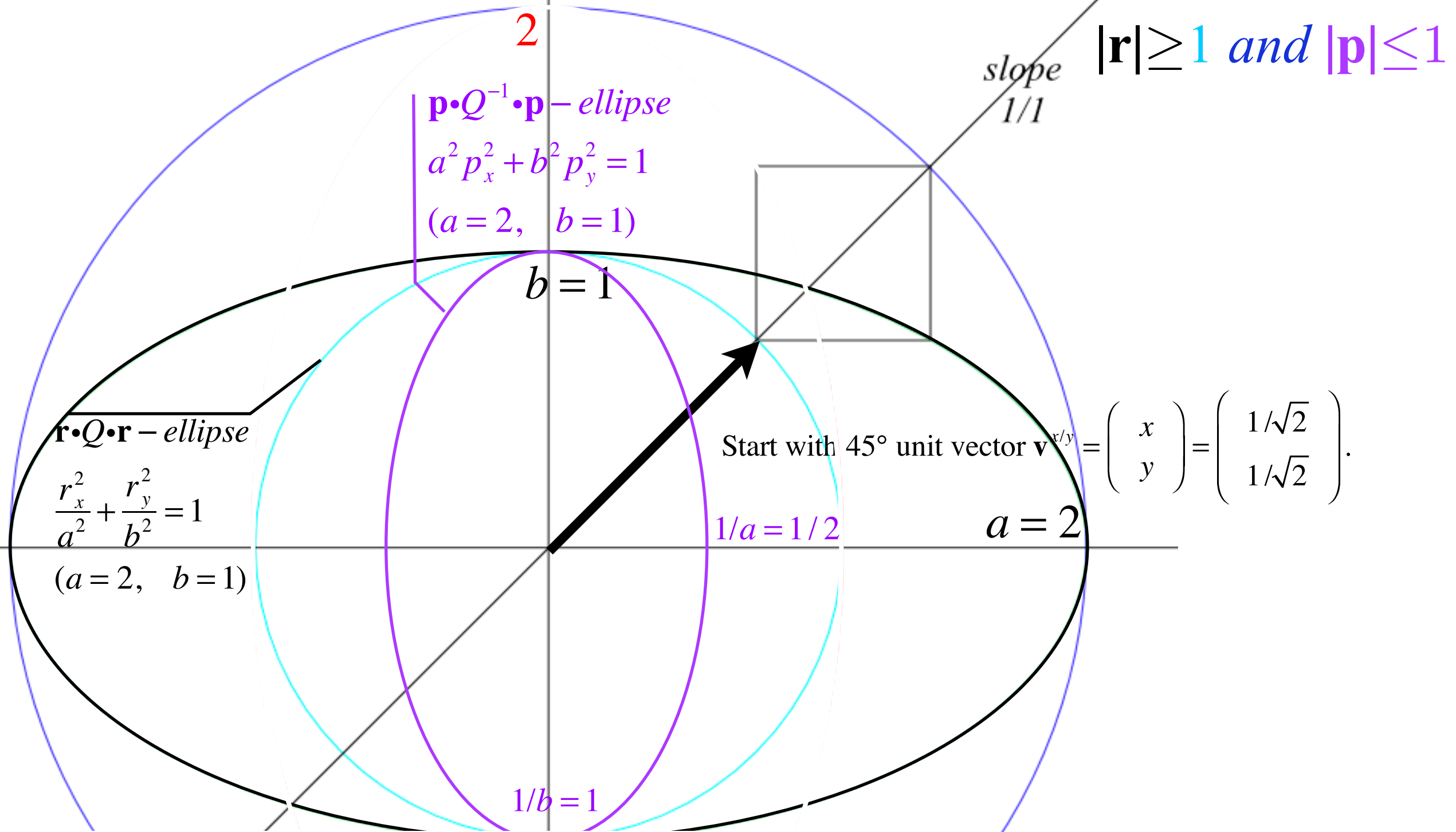
p-ellipse x -radius= $1/a$ plotted at: $S(1/a)=b$ ($=1$ for $a=2, b=1$)

p-ellipse y -radius= $1/b$ plotted at: $S(1/b)=a$ ($=2$ for $a=2, b=1$)

..or else rescale **p**-plot by scale factor $S=b$

Here $b/a=1/2$

to separate $\mathbf{r} \cdot \mathbf{Q} \cdot \mathbf{r}$ and $\mathbf{p} \cdot \mathbf{Q}^{-1} \cdot \mathbf{p}$ ellipses into different regions



Here plot of **p**-ellipse is re-scaled by scalefactor $S=b$

p-ellipse x -radius= $1/a$ plotted at: $S(1/a)=b/a$ ($=1/2$ for $a=2, b=1$)

p-ellipse y -radius= $1/b$ plotted at: $S(1/b)=1$

Action of matrix Q that generates an \mathbf{r} -ellipse ($\mathbf{r} \cdot Q \cdot \mathbf{r} = 1$)

on a single \mathbf{r} -vector $\mathbf{r}(\phi_{-1}) \dots$ is to rotate it to a new vector \mathbf{p} on the \mathbf{p} -ellipse ($\mathbf{p} \cdot Q^{-1} \cdot \mathbf{p} = 1$), that is, $Q \cdot \mathbf{r}(\phi_{-1}) = \mathbf{p}(\phi_{+1})$

$$\mathbf{p}(\phi_1) = Q \cdot \mathbf{r}(\phi_{-1})$$

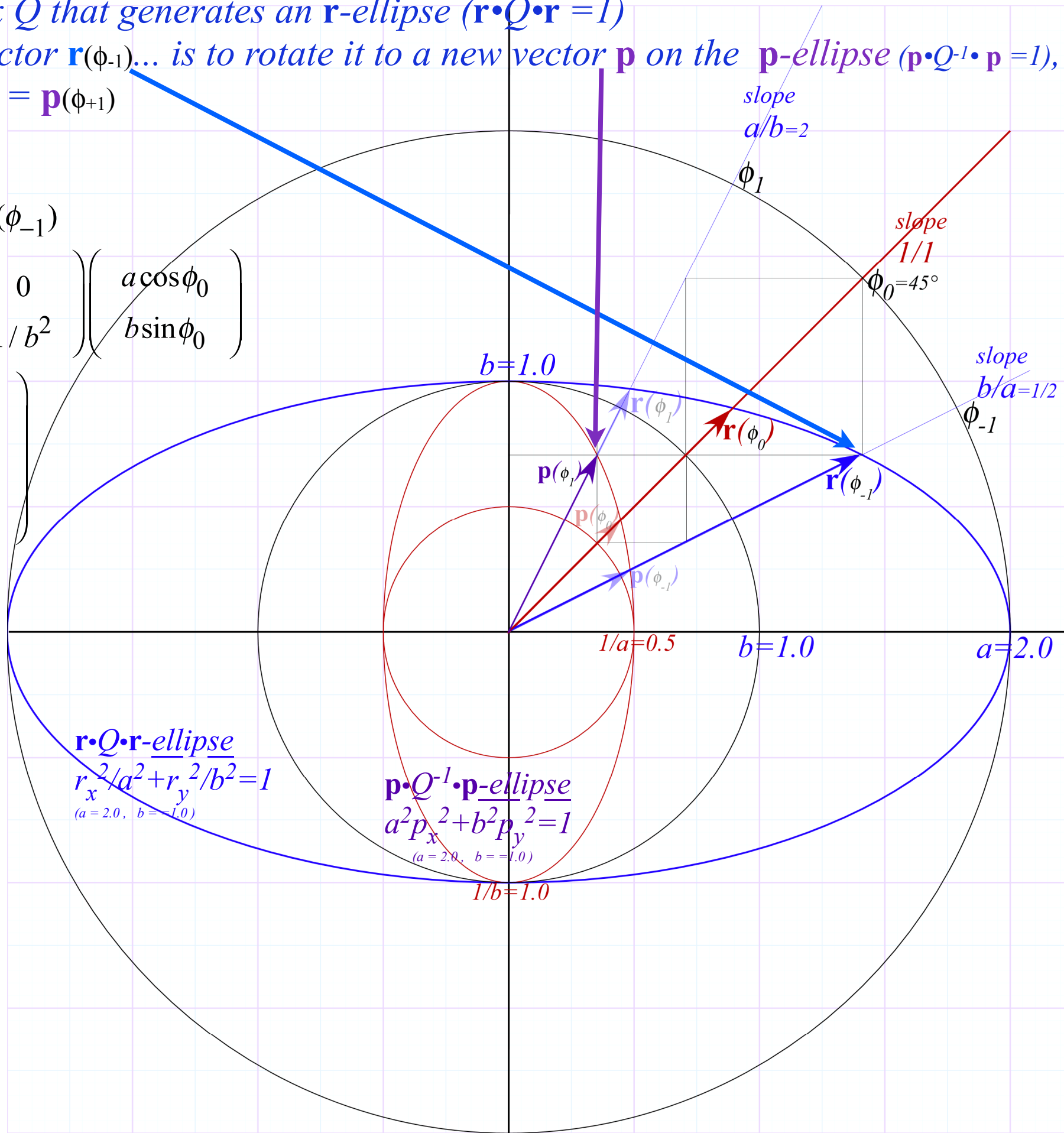
$$= \begin{pmatrix} 1/a^2 & 0 \\ 0 & 1/b^2 \end{pmatrix} \begin{pmatrix} a \cos \phi_0 \\ b \sin \phi_0 \end{pmatrix}$$

$$= \begin{pmatrix} \frac{1}{a} \cos \phi_0 \\ \frac{1}{b} \sin \phi_0 \end{pmatrix}$$

$$= \begin{pmatrix} \frac{1}{2} & \frac{1}{\sqrt{2}} \\ \frac{1}{1} & \frac{1}{\sqrt{2}} \end{pmatrix}$$

$\mathbf{r} \cdot Q \cdot \mathbf{r}$ -ellipse
 $r_x^2/a^2 + r_y^2/b^2 = 1$
 ($a = 2.0, b = 1.0$)

$\mathbf{p} \cdot Q^{-1} \cdot \mathbf{p}$ -ellipse
 $a^2 p_x^2 + b^2 p_y^2 = 1$
 ($a = 2.0, b = 1.0$)



Variation of Fig. 11.7 in Unit 1

Action of matrix Q that generates an \mathbf{r} -ellipse ($\mathbf{r} \cdot Q \cdot \mathbf{r} = 1$)

on a single \mathbf{r} -vector $\mathbf{r}(\phi_{-1})$... is to rotate it to a new vector \mathbf{p} on the \mathbf{p} -ellipse ($\mathbf{p} \cdot Q^{-1} \cdot \mathbf{p} = 1$), that is, $Q \cdot \mathbf{r}(\phi_{-1}) = \mathbf{p}(\phi_{+1})$

$$\mathbf{p}(\phi_1) = Q \cdot \mathbf{r}(\phi_{-1})$$

$$= \begin{pmatrix} 1/a^2 & 0 \\ 0 & 1/b^2 \end{pmatrix} \begin{pmatrix} a \cos \phi_0 \\ b \sin \phi_0 \end{pmatrix}$$

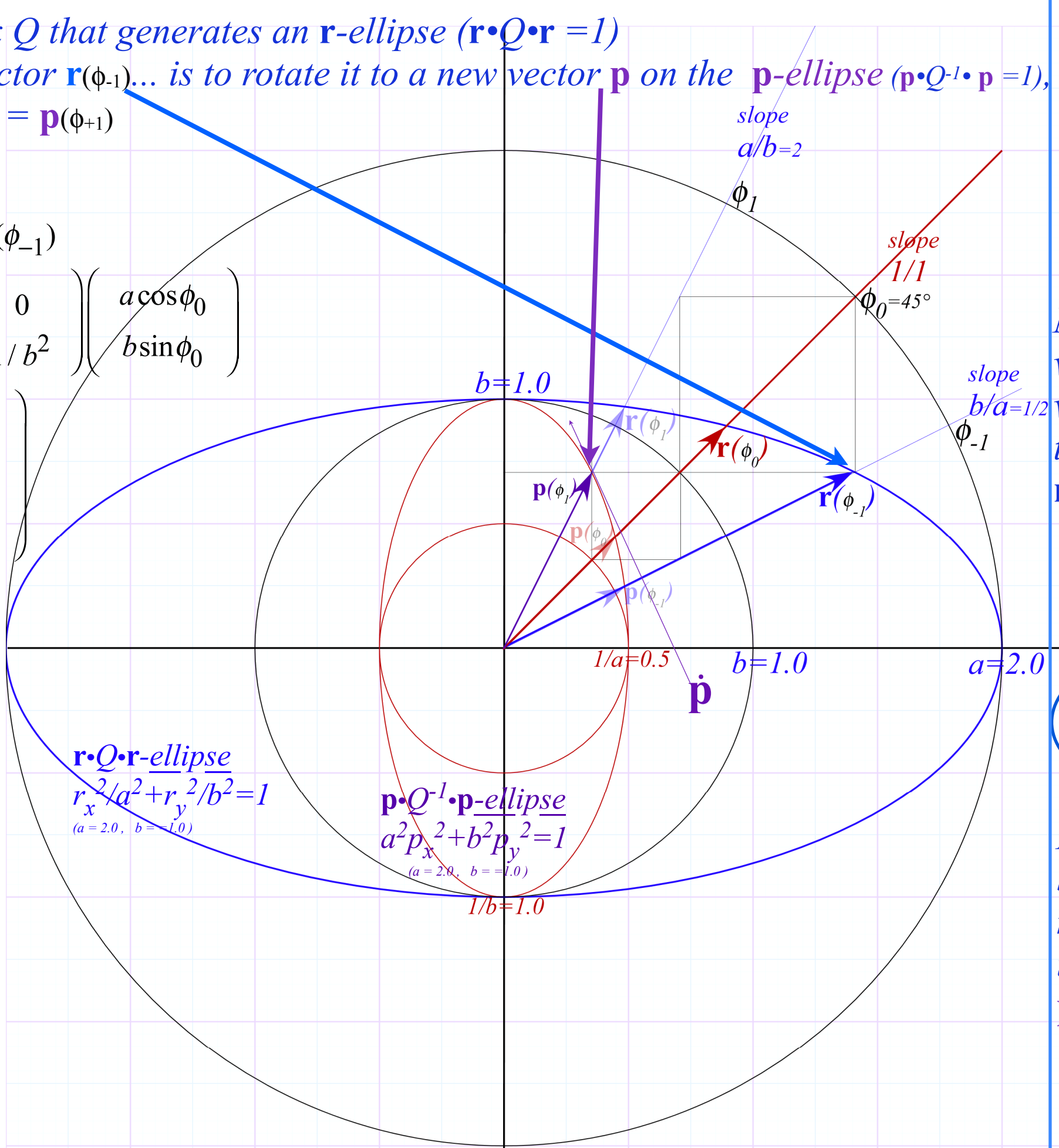
$$= \begin{pmatrix} \frac{1}{a} \cos \phi_0 \\ \frac{1}{b} \sin \phi_0 \end{pmatrix}$$

$$= \begin{pmatrix} \frac{1}{2} & \frac{1}{\sqrt{2}} \\ \frac{1}{1} & \frac{1}{\sqrt{2}} \end{pmatrix}$$

$\mathbf{r} \cdot Q \cdot \mathbf{r}$ -ellipse
 $r_x^2/a^2 + r_y^2/b^2 = 1$
 ($a = 2.0, b = 1.0$)

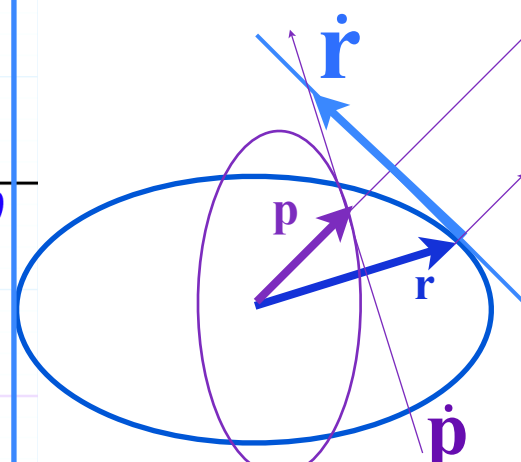
$\mathbf{p} \cdot Q^{-1} \cdot \mathbf{p}$ -ellipse
 $a^2 p_x^2 + b^2 p_y^2 = 1$
 ($a = 2.0, b = 1.0$)

Variation of Fig. 11.7 in Unit 1



Key points of matrix geometry:

Matrix Q maps any vector \mathbf{r} to a new vector \mathbf{p} normal to the tangent $\dot{\mathbf{r}}$ to its $\mathbf{r} \cdot Q \cdot \mathbf{r}$ -ellipse.



Matrix Q^{-1} maps \mathbf{p} back to \mathbf{r} that is normal to the tangent $\dot{\mathbf{p}}$ to its $\mathbf{p} \cdot Q^{-1} \cdot \mathbf{p}$ -ellipse.

Introduction to dual matrix operator geometry (based on IHO orbits)

Quadratic form ellipse $\mathbf{r} \cdot \mathbf{Q} \cdot \mathbf{r} = 1$ vs. inverse form ellipse $\mathbf{p} \cdot \mathbf{Q}^{-1} \cdot \mathbf{p} = 1$

Duality norm relations ($\mathbf{r} \cdot \mathbf{p} = 1$)

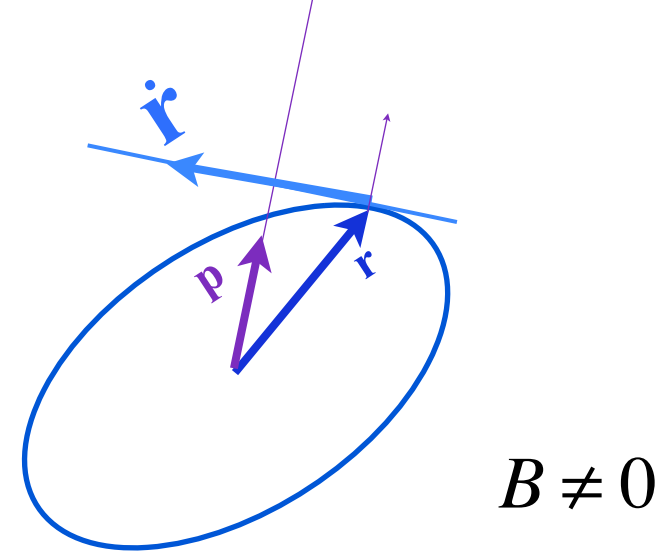
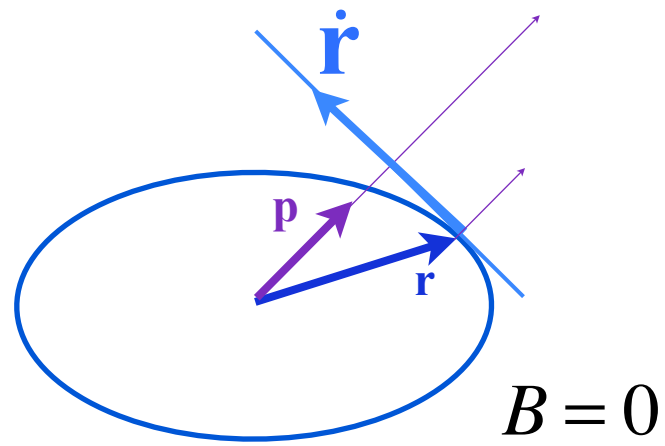
\mathbf{Q} -Ellipse tangents \mathbf{r}' normal to dual \mathbf{Q}^{-1} -ellipse position \mathbf{p} ($\mathbf{r}' \cdot \mathbf{p} = 0 = \mathbf{r} \cdot \mathbf{p}'$)

Operator geometric sequences and eigenvectors

Alternative scaling of matrix operator geometry



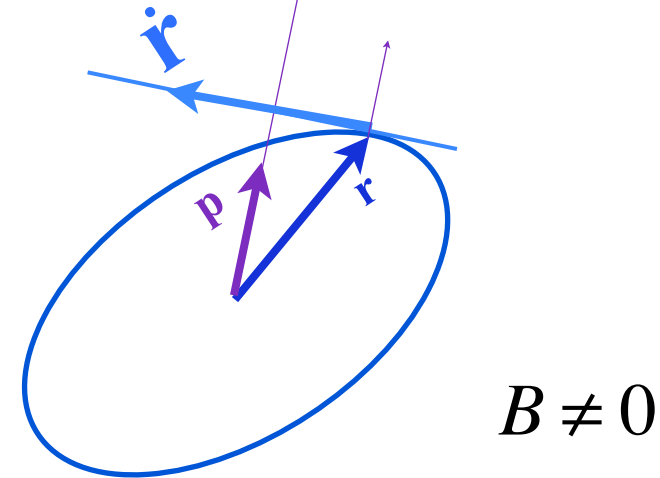
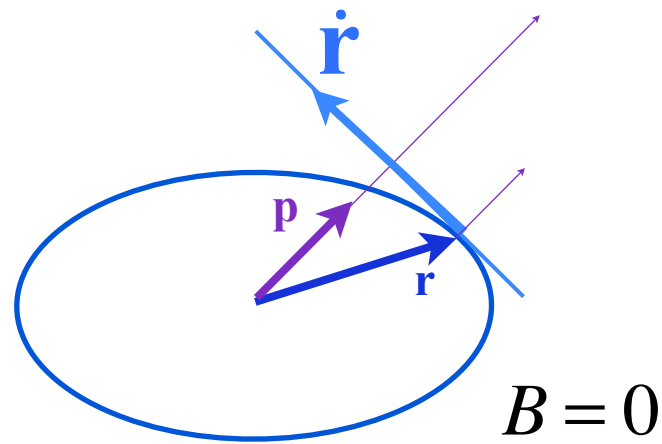
Vector calculus of tensor operation



Derive matrix “normal-to-ellipse” geometry by vector calculus:

Let matrix $Q = \begin{pmatrix} A & B \\ B & D \end{pmatrix}$

define the ellipse $1 = \mathbf{r} \cdot \mathbf{Q} \cdot \mathbf{r} = \begin{pmatrix} x & y \end{pmatrix} \cdot \begin{pmatrix} A & B \\ B & D \end{pmatrix} \cdot \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} x & y \end{pmatrix} \cdot \begin{pmatrix} A \cdot x + B \cdot y \\ B \cdot x + D \cdot y \end{pmatrix} = A \cdot x^2 + 2B \cdot xy + D \cdot y^2 = 1$



Derive matrix “normal-to-ellipse” geometry by vector calculus:

Let matrix $Q = \begin{pmatrix} A & B \\ B & D \end{pmatrix}$

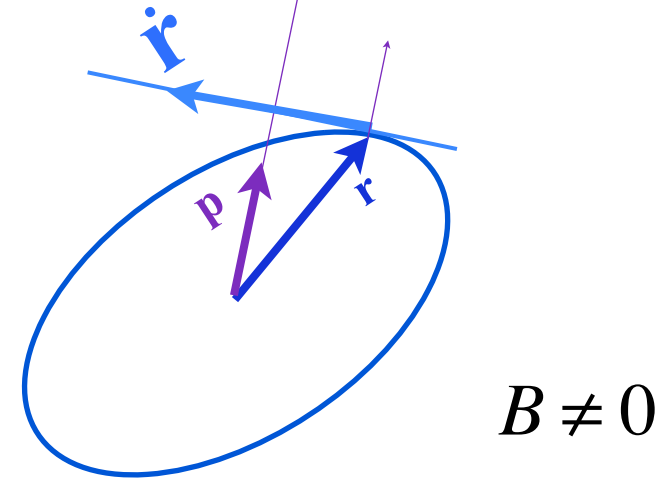
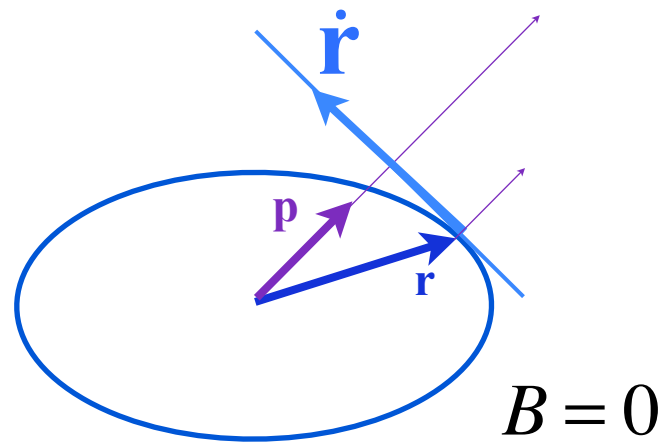
define the ellipse $1 = \mathbf{r} \cdot \mathbf{Q} \cdot \mathbf{r} = \begin{pmatrix} x & y \end{pmatrix} \cdot \begin{pmatrix} A & B \\ B & D \end{pmatrix} \cdot \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} x & y \end{pmatrix} \cdot \begin{pmatrix} A \cdot x + B \cdot y \\ B \cdot x + D \cdot y \end{pmatrix} = A \cdot x^2 + 2B \cdot xy + D \cdot y^2 = 1$

Compare operation by Q on vector \mathbf{r} with vector derivative or gradient of $\mathbf{r} \cdot \mathbf{Q} \cdot \mathbf{r}$

$$\begin{pmatrix} A & B \\ B & D \end{pmatrix} \cdot \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} A \cdot x + B \cdot y \\ B \cdot x + D \cdot y \end{pmatrix}$$

$$\frac{\partial}{\partial \mathbf{r}} (\mathbf{r} \cdot \mathbf{Q} \cdot \mathbf{r}) = \nabla (\mathbf{r} \cdot \mathbf{Q} \cdot \mathbf{r})$$

$$\begin{pmatrix} \frac{\partial}{\partial x} \\ \frac{\partial}{\partial y} \end{pmatrix} (A \cdot x^2 + 2B \cdot xy + D \cdot y^2) = \begin{pmatrix} 2A \cdot x + 2B \cdot y \\ 2B \cdot x + 2D \cdot y \end{pmatrix}$$



Derive matrix “normal-to-ellipse” geometry by vector calculus:

Let matrix $Q = \begin{pmatrix} A & B \\ B & D \end{pmatrix}$

define the ellipse $1 = \mathbf{r} \cdot Q \cdot \mathbf{r} = \begin{pmatrix} x & y \end{pmatrix} \cdot \begin{pmatrix} A & B \\ B & D \end{pmatrix} \cdot \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} x & y \end{pmatrix} \cdot \begin{pmatrix} A \cdot x + B \cdot y \\ B \cdot x + D \cdot y \end{pmatrix} = A \cdot x^2 + 2B \cdot xy + D \cdot y^2 = 1$

Compare operation by Q on vector \mathbf{r} with vector derivative or gradient of $\mathbf{r} \cdot Q \cdot \mathbf{r}$

$$\begin{pmatrix} A & B \\ B & D \end{pmatrix} \cdot \begin{pmatrix} x \\ y \end{pmatrix} = \begin{pmatrix} A \cdot x + B \cdot y \\ B \cdot x + D \cdot y \end{pmatrix} \quad \frac{\partial}{\partial \mathbf{r}} (\mathbf{r} \cdot Q \cdot \mathbf{r}) = \nabla (\mathbf{r} \cdot Q \cdot \mathbf{r})$$

$$\begin{pmatrix} \frac{\partial}{\partial x} \\ \frac{\partial}{\partial y} \end{pmatrix} (A \cdot x^2 + 2B \cdot xy + D \cdot y^2) = \begin{pmatrix} 2A \cdot x + 2B \cdot y \\ 2B \cdot x + 2D \cdot y \end{pmatrix}$$

Very simple result:

$$\frac{\partial}{\partial \mathbf{r}} \left(\frac{\mathbf{r} \cdot Q \cdot \mathbf{r}}{2} \right) = \nabla \left(\frac{\mathbf{r} \cdot Q \cdot \mathbf{r}}{2} \right) = Q \cdot \mathbf{r}$$

Introduction to dual matrix operator geometry (based on IHO orbits)

Quadratic form ellipse $\mathbf{r} \cdot \mathbf{Q} \cdot \mathbf{r} = 1$ vs. inverse form ellipse $\mathbf{p} \cdot \mathbf{Q}^{-1} \cdot \mathbf{p} = 1$

Duality norm relations ($\mathbf{r} \cdot \mathbf{p} = 1$)

Q-Ellipse tangents \mathbf{r}' normal to dual \mathbf{Q}^{-1} -ellipse position \mathbf{p} ($\mathbf{r}' \cdot \mathbf{p} = 0 = \mathbf{r} \cdot \mathbf{p}'$)

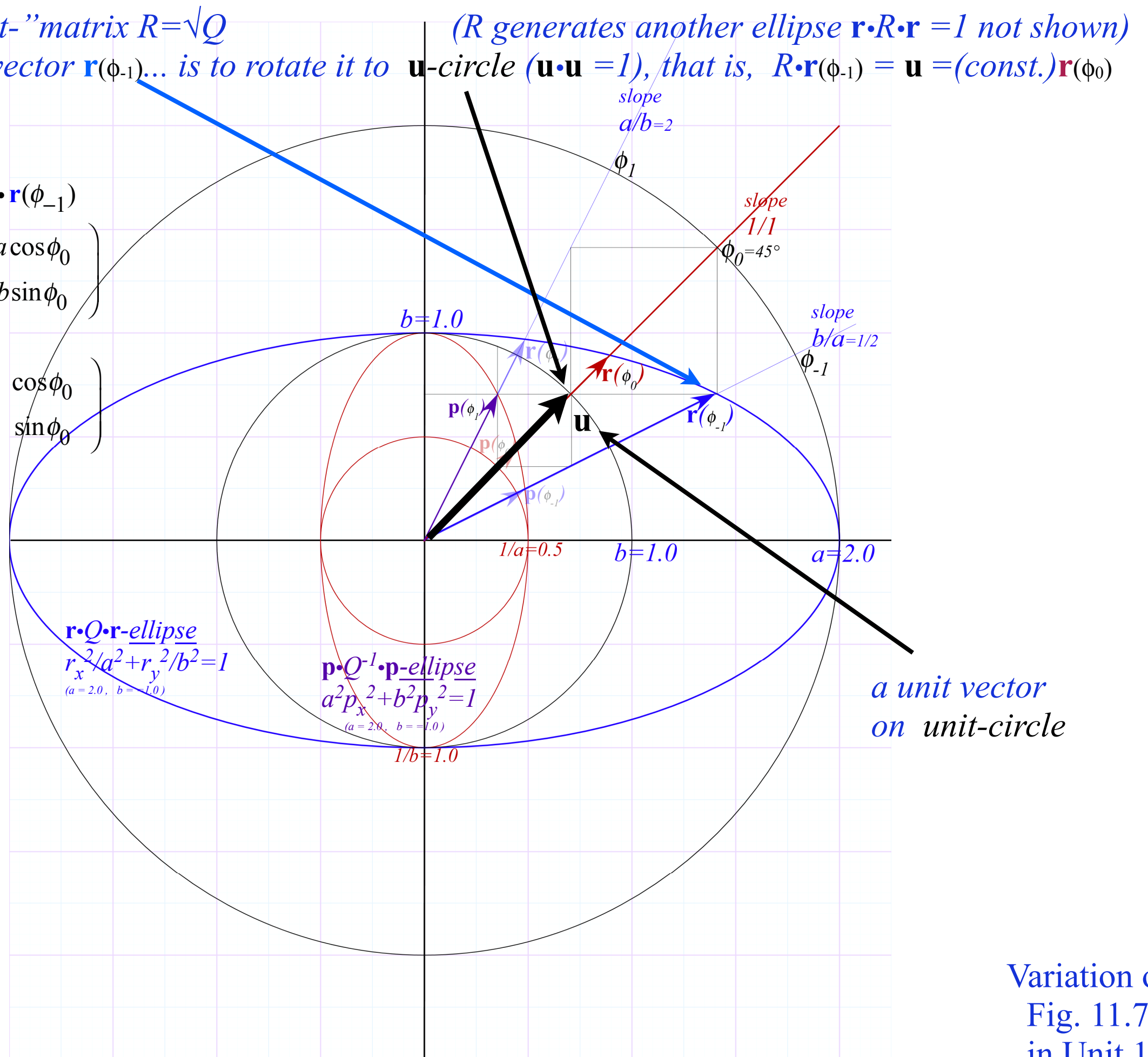
(Still more) Operator geometric sequences and eigenvectors

Alternative scaling of matrix operator geometry

Vector calculus of tensor operation

Action of "sqrt-" matrix $R=\sqrt{Q}$ (R generates another ellipse $\mathbf{r}\cdot R\cdot\mathbf{r} = 1$ not shown) on a single \mathbf{r} -vector $\mathbf{r}(\phi_{-1})$... is to rotate it to \mathbf{u} -circle ($\mathbf{u}\cdot\mathbf{u} = 1$), that is, $R\cdot\mathbf{r}(\phi_{-1}) = \mathbf{u} = (\text{const.})\mathbf{r}(\phi_0)$

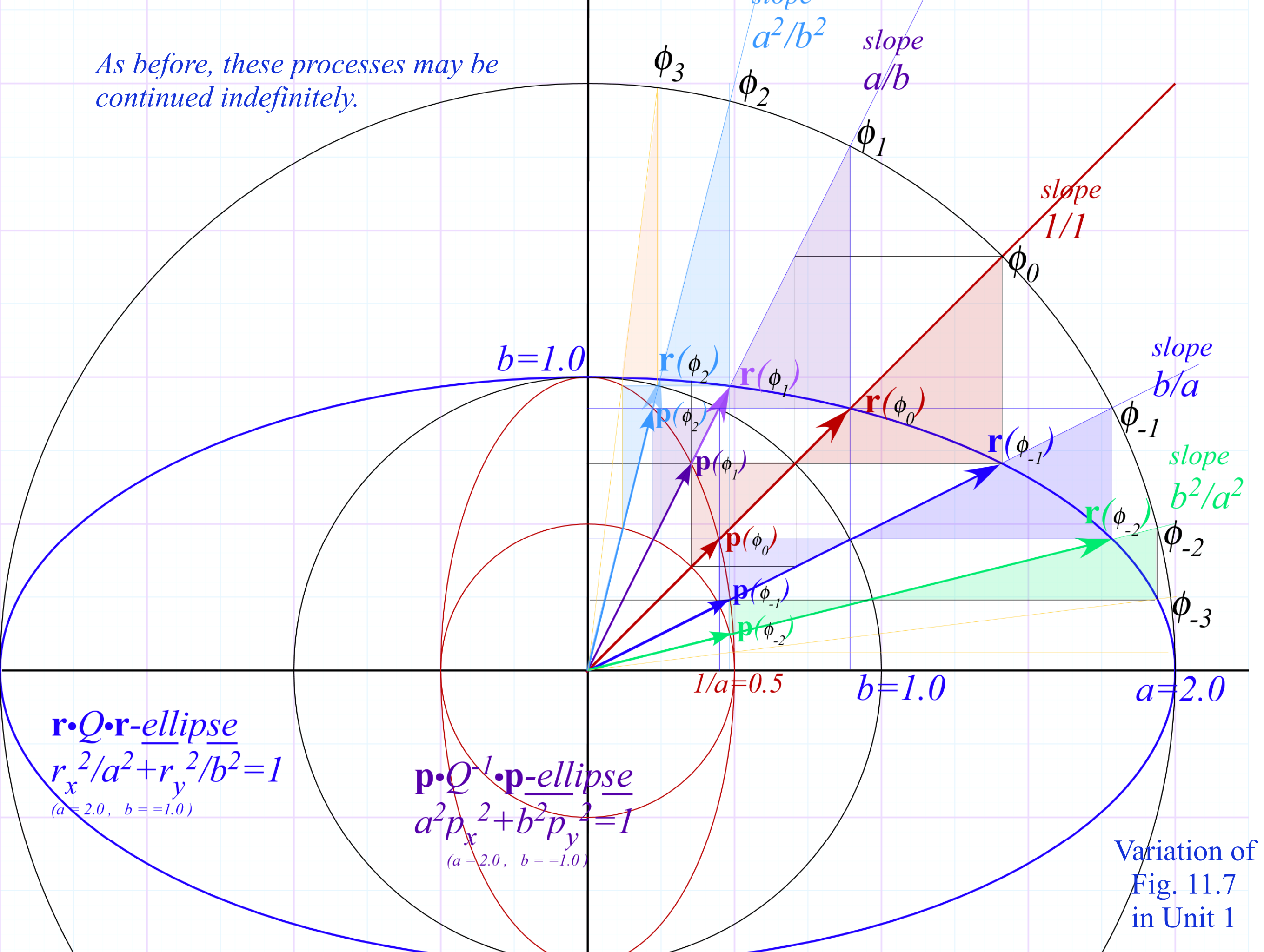
$$\begin{aligned} \mathbf{u} &= \sqrt{\mathbf{Q}} \cdot \mathbf{r}(\phi_{-1}) = \mathbf{R} \cdot \mathbf{r}(\phi_{-1}) \\ &= \begin{pmatrix} 1/a & 0 \\ 0 & 1/b \end{pmatrix} \begin{pmatrix} a \cos \phi_0 \\ b \sin \phi_0 \end{pmatrix} \\ &= \begin{pmatrix} \frac{1}{a} a \cos \phi_0 \\ \frac{1}{b} b \sin \phi_0 \end{pmatrix} = \begin{pmatrix} \cos \phi_0 \\ \sin \phi_0 \end{pmatrix} \\ &= \begin{pmatrix} \frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{2}} \end{pmatrix} \end{aligned}$$



a unit vector on unit-circle

Variation of Fig. 11.7 in Unit 1

As before, these processes may be continued indefinitely.

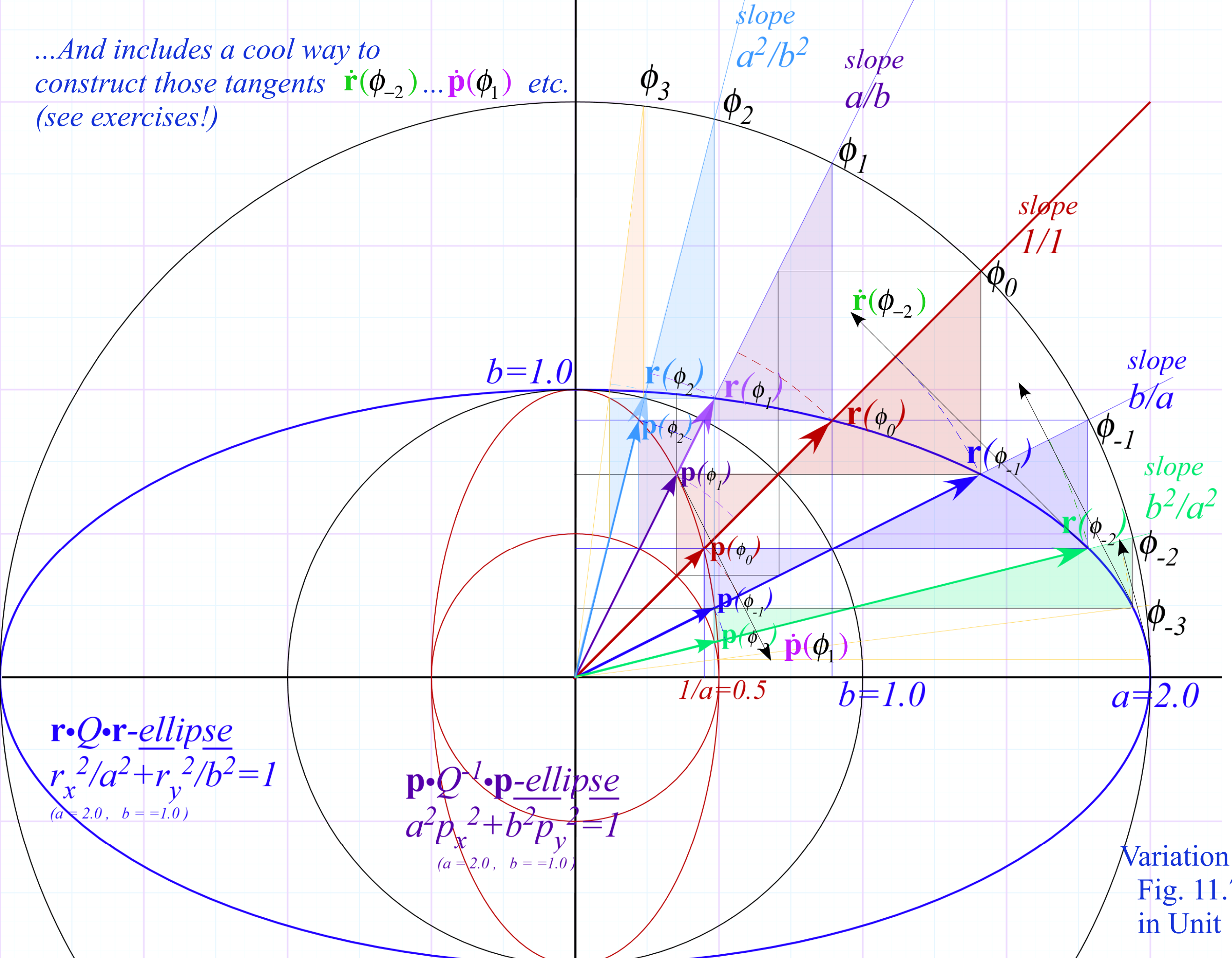


r·Q·r-ellipse
 $r_x^2/a^2 + r_y^2/b^2 = 1$
 (a=2.0, b=1.0)

p·Q⁻¹·p-ellipse
 $a^2 p_x^2 + b^2 p_y^2 = 1$
 (a=2.0, b=1.0)

Variation of Fig. 11.7 in Unit 1

...And includes a cool way to construct those tangents $\dot{\mathbf{r}}(\phi_{-2}) \dots \dot{\mathbf{p}}(\phi_1)$ etc. (see exercises!)



$\mathbf{r} \cdot \mathbf{Q} \cdot \mathbf{r}$ -ellipse
 $r_x^2/a^2 + r_y^2/b^2 = 1$
 ($a=2.0, b=1.0$)

$\mathbf{p} \cdot \mathbf{Q}^{-1} \cdot \mathbf{p}$ -ellipse
 $a^2 p_x^2 + b^2 p_y^2 = 1$
 ($a=2.0, b=1.0$)

Variation of Fig. 11.7 in Unit 1

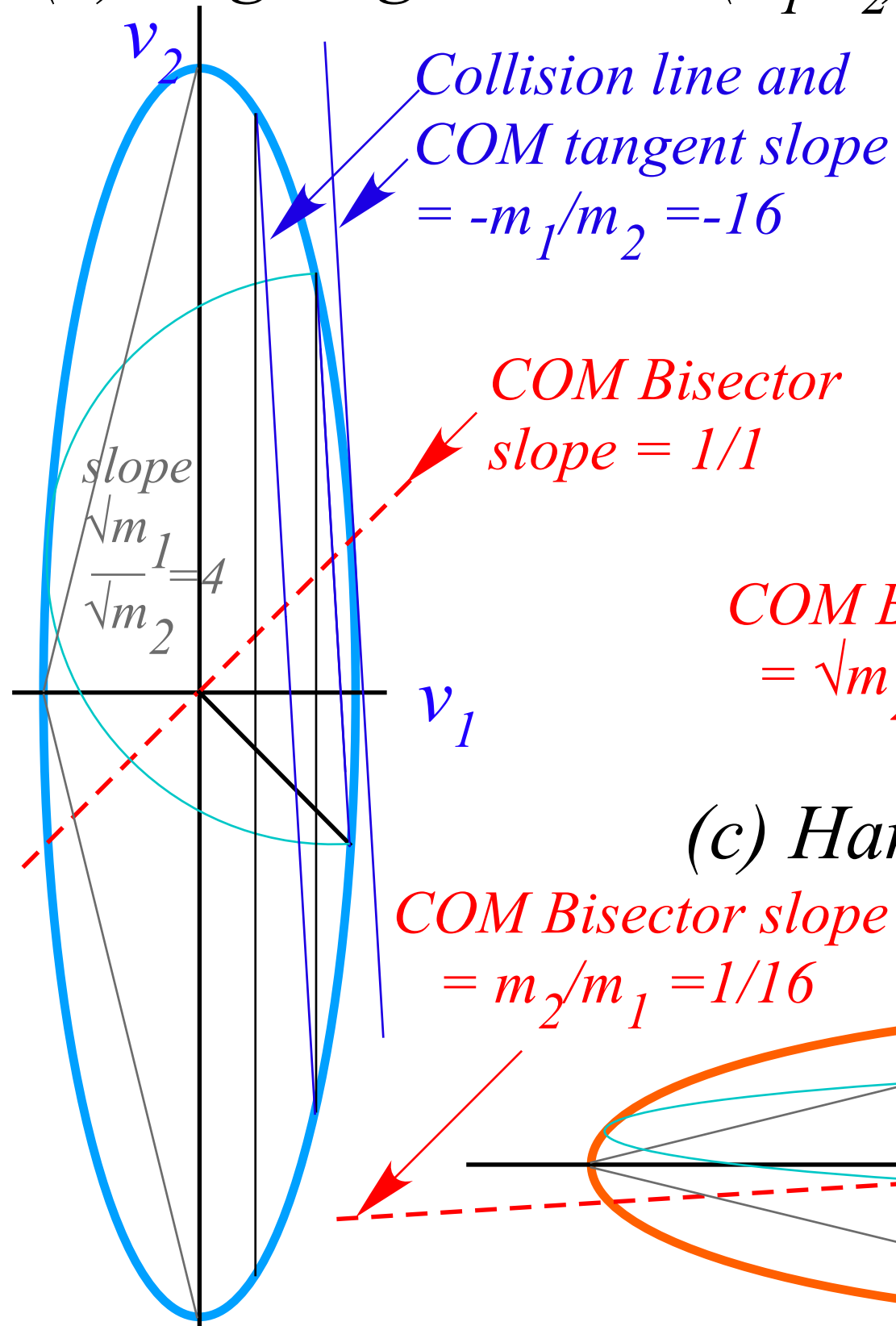
*Q: Where is this headed?
Preview of Lecture 8*

A: Lagrangian-Hamiltonian duality

The R and Q matrix transformations are like the mechanics rescaling matrices $\sqrt{\mathbf{M}}$ and \mathbf{M} :

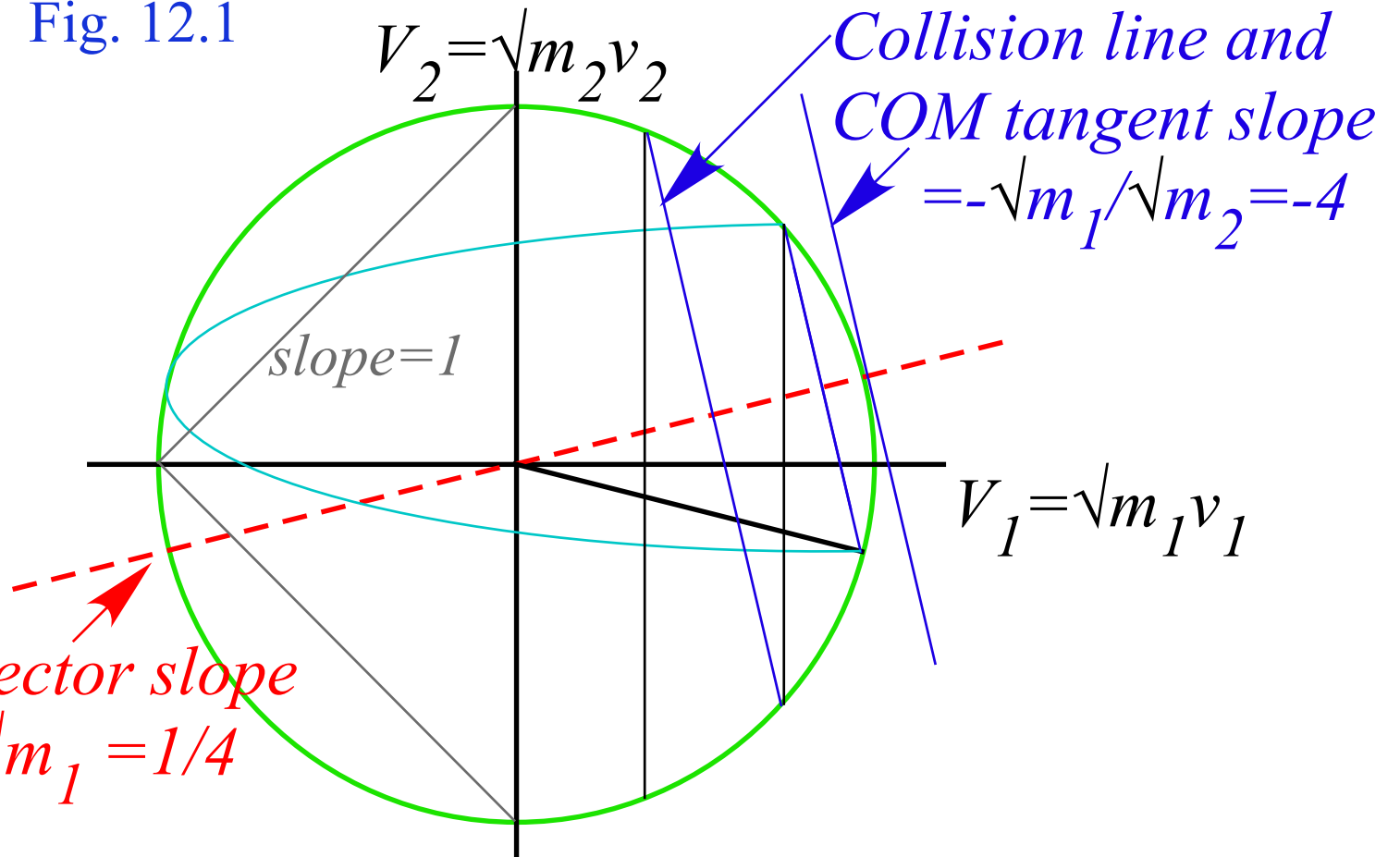
Like $Q=R^2$: $\mathbf{M} = \begin{pmatrix} m_1 & 0 \\ 0 & m_2 \end{pmatrix} = \mathbf{R}^2$ Like $\sqrt{Q}=R$: $\sqrt{\mathbf{M}} = \begin{pmatrix} \sqrt{m_1} & 0 \\ 0 & \sqrt{m_2} \end{pmatrix} = \mathbf{R}$ Like $Q^{-1}=R^{-2}$: $\mathbf{M}^{-1} = \begin{pmatrix} 1/m_1 & 0 \\ 0 & 1/m_2 \end{pmatrix} = \mathbf{R}^{-2}$

(a) Lagrangian $L = L(v_1, v_2)$

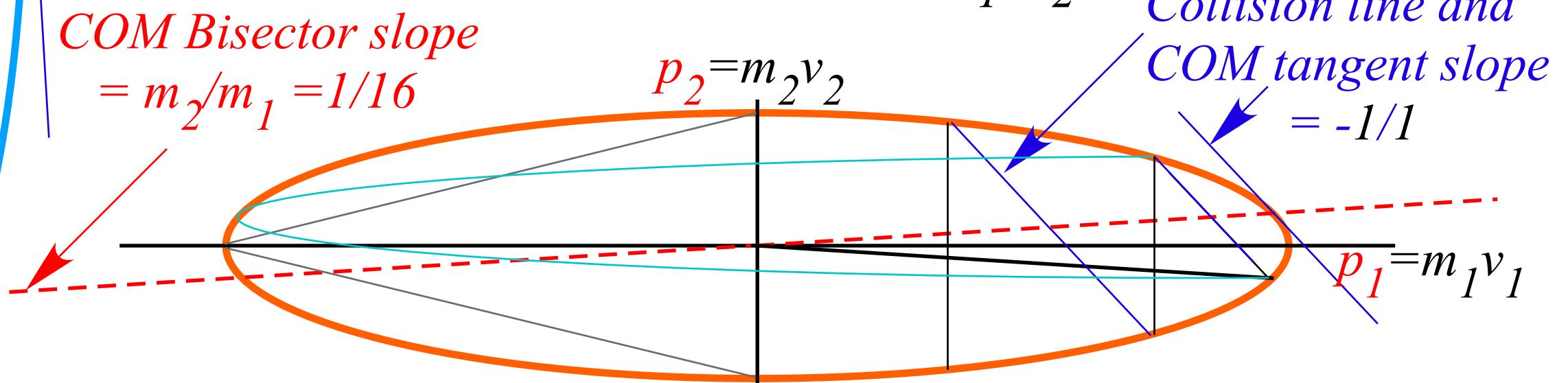


Unit 1
Fig. 12.1

(b) Estrangian $E = E(V_1, V_2)$

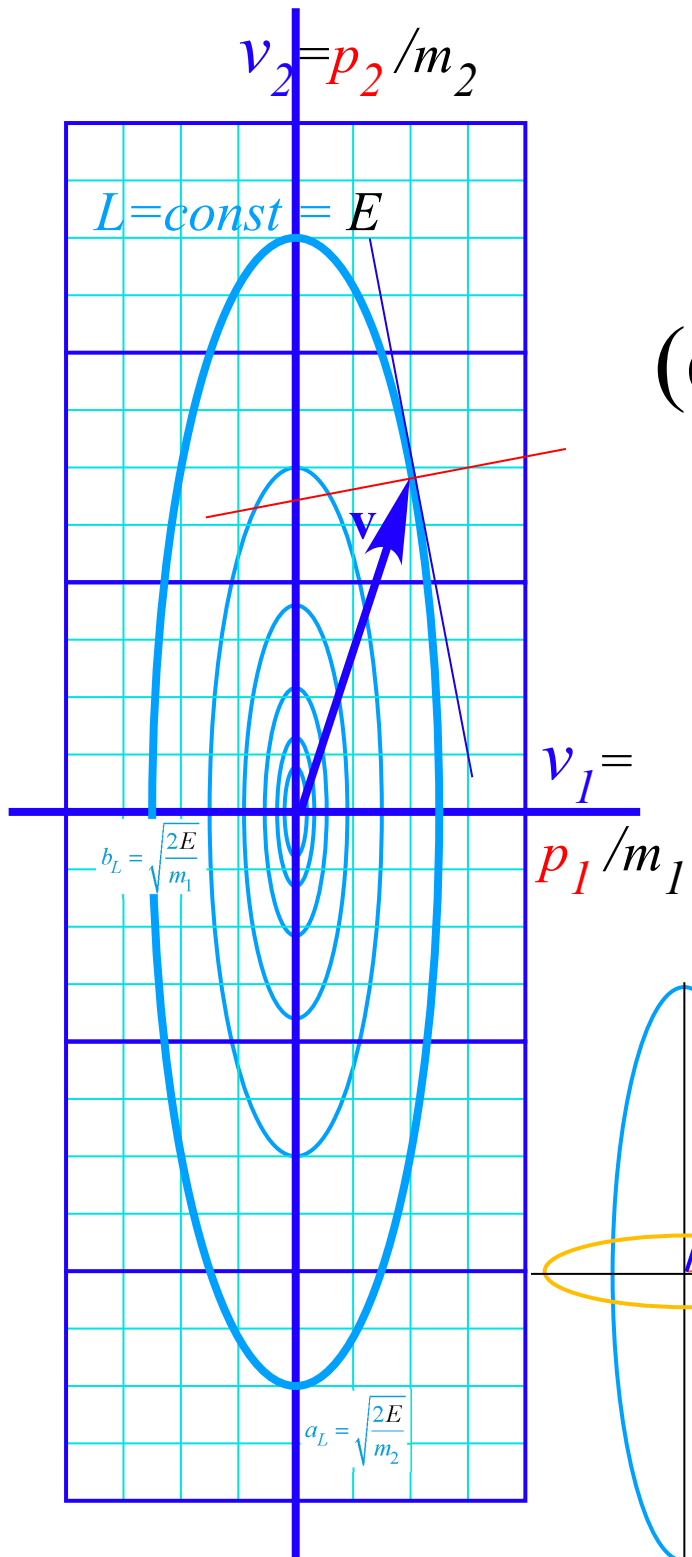


(c) Hamiltonian $H = H(p_1, p_2)$

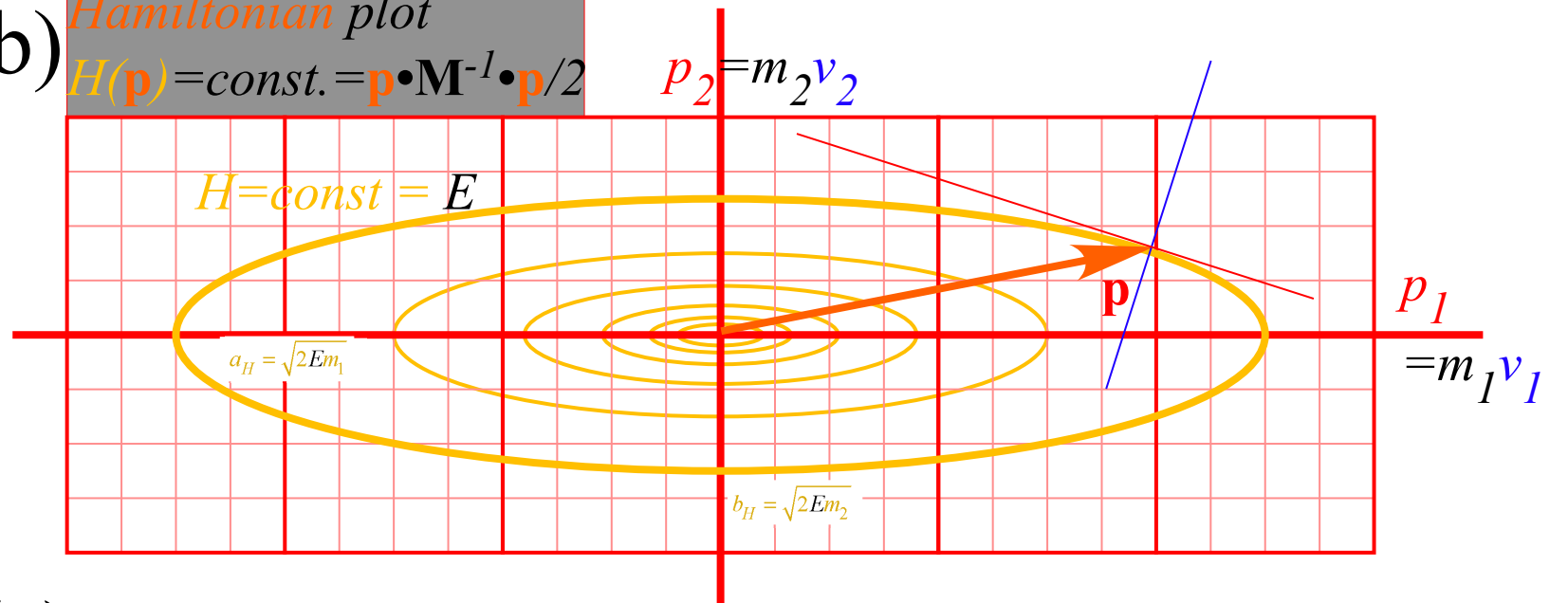


Unit 1
Fig. 12.2

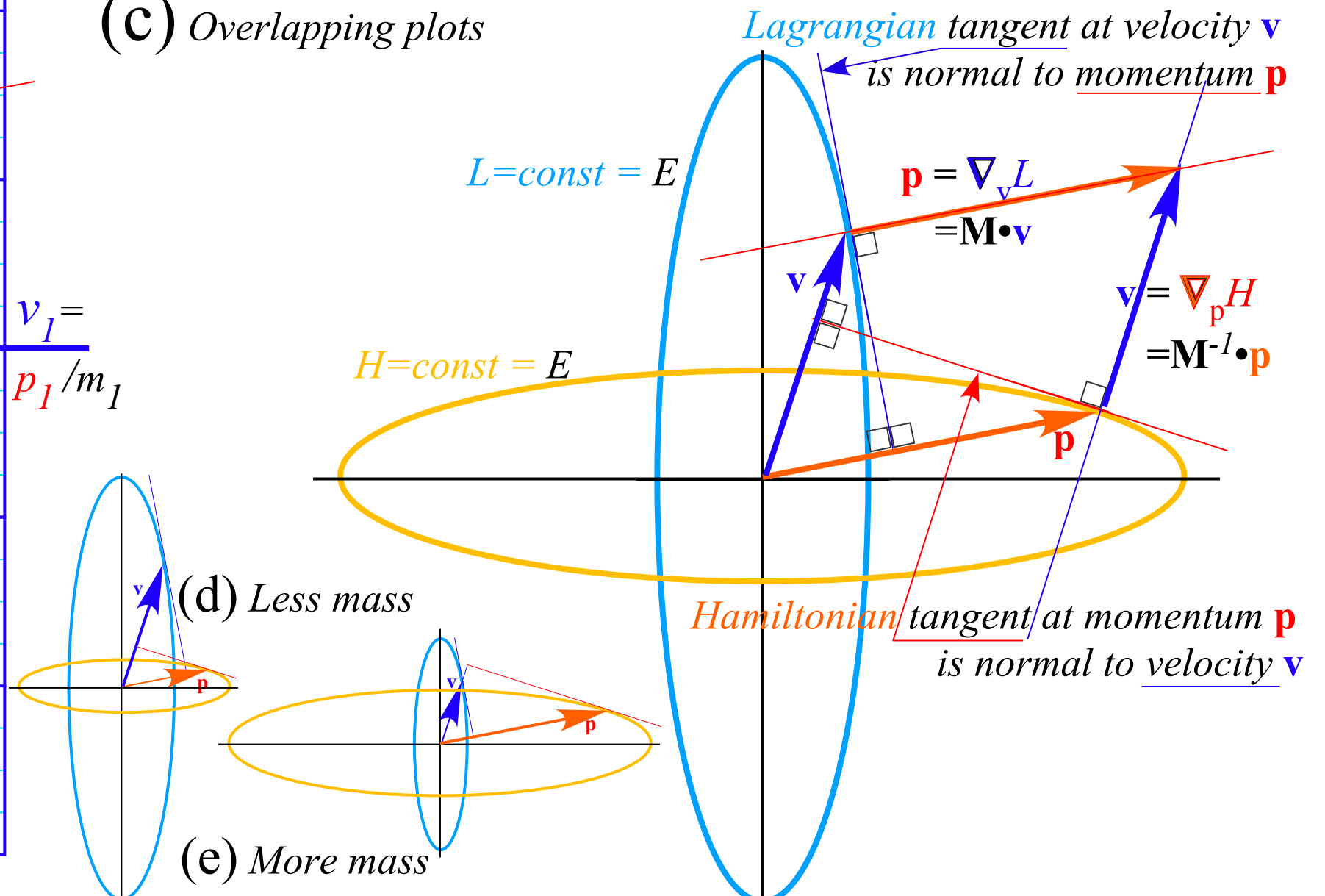
(a) *Lagrangian plot*
 $L(\mathbf{v}) = \text{const.} = \mathbf{v} \cdot \mathbf{M} \cdot \mathbf{v} / 2$



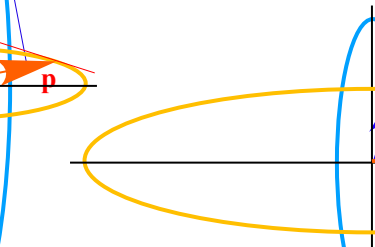
(b) *Hamiltonian plot*
 $H(\mathbf{p}) = \text{const.} = \mathbf{p} \cdot \mathbf{M}^{-1} \cdot \mathbf{p} / 2$



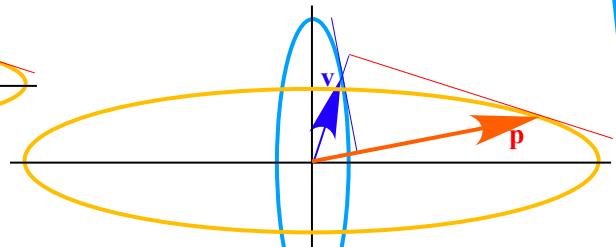
(c) *Overlapping plots*



(d) *Less mass*



(e) *More mass*



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Alternative Basis for the Theory of Complex Spectra

[Alternative Basis for the Theory of Complex Spectra I - harter-pra-1973](#)

[Alternative Basis for the Theory of Complex Spectra II - harter-patterson-pra-1976](#)

[Alternative Basis for the Theory of Complex Spectra III - patterson-harter-pra-1977](#)

[Frame Transformation Relations And Multipole Transitions In Symmetric Polyatomic Molecules - RMP-1978](#)

[Asymptotic eigensolutions of fourth and sixth rank octahedral tensor operators - Harter-Patterson-JMP-1979](#)

[Rotational energy surfaces and high- J eigenvalue structure of polyatomic molecules - Harter - Patterson - 1984](#)

[Galloping waves and their relativistic properties - ajp-1985-Harter](#)

[Rovibrational Spectral Fine Structure Of Icosahedral Molecules - Cpl 1986 \(Alt Scan\)](#)

Theory of hyperfine and superfine levels in symmetric polyatomic molecules.

I) [Trigonal and tetrahedral molecules: Elementary spin-1/2 cases in vibronic ground states - PRA-1979-Harter-Patterson \(Alt scan\)](#)

II) [Elementary cases in octahedral hexafluoride molecules - Harter-PRA-1981 \(Alt scan\)](#)

Rotation-vibration spectra of icosahedral molecules.

I) [Icosahedral symmetry analysis and fine structure - harter-weeks-jcp-1989 \(Alt scan\)](#)

II) [Icosahedral symmetry, vibrational eigenfrequencies, and normal modes of buckminsterfullerene - weeks-harter-jcp-1989 \(Alt scan\)](#)

III) [Half-integral angular momentum - harter-reimer-jcp-1991](#)

[Rotation-vibration scalar coupling zeta coefficients and spectroscopic band shapes of buckminsterfullerene - Weeks-Harter-CPL-1991 \(Alt scan\)](#)

[Nuclear spin weights and gas phase spectral structure of ¹²C₆₀ and ¹³C₆₀ buckminsterfullerene -Harter-Reimer-Cpl-1992 - \(Alt1, Alt2 Erratum\)](#)

[Gas Phase Level Structure of C₆₀ Buckyball and Derivatives Exhibiting Broken Icosahedral Symmetry - reimer-diss-1996](#)

[Fullerene symmetry reduction and rotational level fine structure/ the Buckyball isotopomer ¹²C ¹³C₅₉ - jcp-Reimer-Harter-1997 \(HiRez\)](#)

[Wave Node Dynamics and Revival Symmetry in Quantum Rotors - harter - jms - 2001](#)

[Molecular Symmetry and Dynamics - Ch32-Springer Handbooks of Atomic, Molecular, and Optical Physics - Harter-2006](#)

Resonance and Revivals

I) [QUANTUM ROTOR AND INFINITE-WELL DYNAMICS - ISMSLi2012 \(Talk\) OSU knowledge Bank](#)

II) [Comparing Half-integer Spin and Integer Spin - Alva-ISMS-Ohio2013-R777 \(Talks\)](#)

III) [Quantum Resonant Beats and Revivals in the Morse Oscillators and Rotors - \(2013-Li-Diss\)](#)

[Resonance and Revivals in Quantum Rotors - Comparing Half-integer Spin and Integer Spin - Alva-ISMS-Ohio2013-R777 \(Talk\)](#)

[Molecular Eigensolution Symmetry Analysis and Fine Structure - IJMS-harter-mitchell-2013](#)

[Quantum Revivals of Morse Oscillators and Farey-Ford Geometry - Li-Harter-cpl-2013](#)

[QTCA Unit 10 Ch 30 - 2013](#)

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(Int.J.Mol.Sci, 14, 714(2013) p.755-774 ,

QTCA Unit 7 Ch. 23-26),

(PSDS - Ch. 5, 7)

[Int.J.Mol.Sci, 14, 714\(2013\),](#)

[QTCA Unit 8 Ch. 23-25,](#)

[QTCA Unit 9 Ch. 26,](#)

[PSDS Ch. 5,](#)

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Intro spin ½ coupling

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[Deeper theory ends p53](#)

Wigner-Eckart tensor Theorem.

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Young Tableau Magic Formulae

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CG coupling derived (start)

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[Group Theory - PUP Lucy Day - Diagrammatic notation - Ch4](#)
[Simplification Rules for Birdtrack Operators - Alcock-Zeilinger-Weigert-zeilinger-jmp-2017](#)
[Group Theory - Birdtracks Lies and Exceptional Groups - Cvitanovic-2011](#)
[Simplification rules for birdtrack operators- jmp-alcock-zeilinger-2017](#)
[Birdtracks for SU\(N\) - 2017-Keppeler](#)

Frank Rioux's: UMA method of vibrational induction

[Quantum Mechanics Group Theory and C60 - Frank Rioux - Department of Chemistry Saint Johns U](#)
[Symmetry Analysis for H2O- H2OGrpTheory- Rioux](#)
[Quantum Mechanics-Group Theory and C60 - JChemEd-Rioux-1994](#)
[Group Theory Problems- Rioux- SymmetryProblemsX](#)
[Comment on the Vibrational Analysis for C60 and Other Fullerenes Rioux-RSP](#)

Supplemental AMOP Techniques & Experiment

[Many Correlation Tables are Molien Sequences - Klee \(Draft 2016\)](#)
[High-resolution spectroscopy and global analysis of CF4 rovibrational bands to model its atmospheric absorption- carlos-Boudon-iqsrt-2017](#)
[Symmetry and Chirality - Continuous Measures - Avnir](#)

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Special Topics & Colloquial References

[r-process nucleosynthesis from matter ejected in binary neutron star mergers-PhysRevD-Bovard-2017](#)

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