

Lecture 7

Thur. 9.17.2013

Geometry and Motion of Isotropic Harmonic Oscillators (Ch. 9 and Ch. 11 of Unit 1)

Geometry of idealized “Sophomore-physics Earth”

Coulomb field outside Isotropic Harmonic Oscillator (IHO) field inside

Contact-geometry of potential curve(s)

“Crushed-Earth” models: 3 key energy “steps” and 4 key energy “levels”

Earth matter vs nuclear matter:

*Introducing the “neutron starlet” and “**Black-Hole-Earth**”*

Isotropic harmonic oscillator dynamics in 1D, 2D, and 3D

Sinusoidal space-time dynamics derived by geometry

Isotropic harmonic oscillator orbits in 1D and 2D (You get 3D for free!)

Constructing 2D Isotropic harmonic oscillator orbits using phasor plots

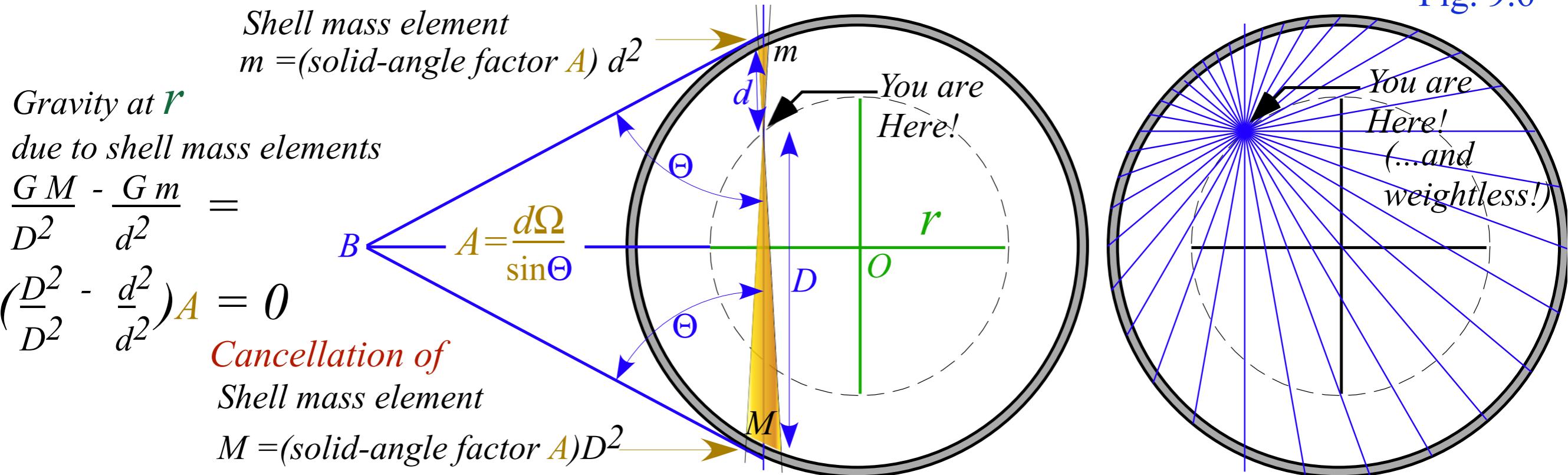
*Examples with x-y **phase lag** : $\alpha_{x-y} = \alpha_x - \alpha_y = 15^\circ, 30^\circ, \text{ and } \pm 75^\circ$*

Geometry of idealized “Sophomore-physics Earth”

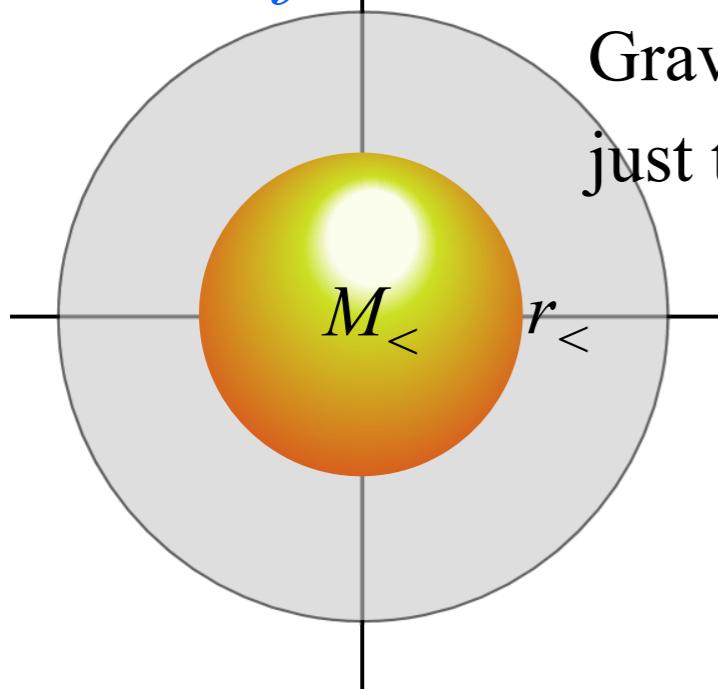
- *Coulomb field outside* *Isotropic Harmonic Oscillator (IHO) field inside*
- Contact-geometry of potential curve(s)*
- “Crushed-Earth” models: 3 key energy “steps” and 4 key energy “levels”*
- Earth matter vs nuclear matter:*
- Introducing the “neutron starlet” and “**Black-Hole-Earth**”*

Coulomb force vanishes inside-spherical shell (Gauss-law)

Unit 1
Fig. 9.6



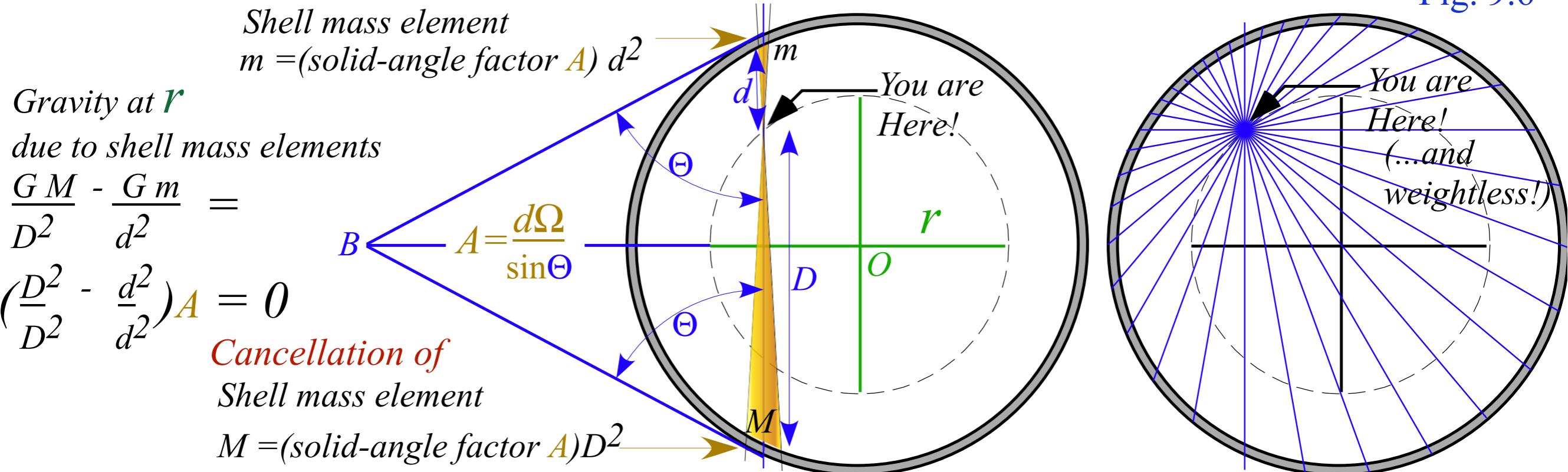
Coulomb force inside-spherical body due to stuff below you, only.



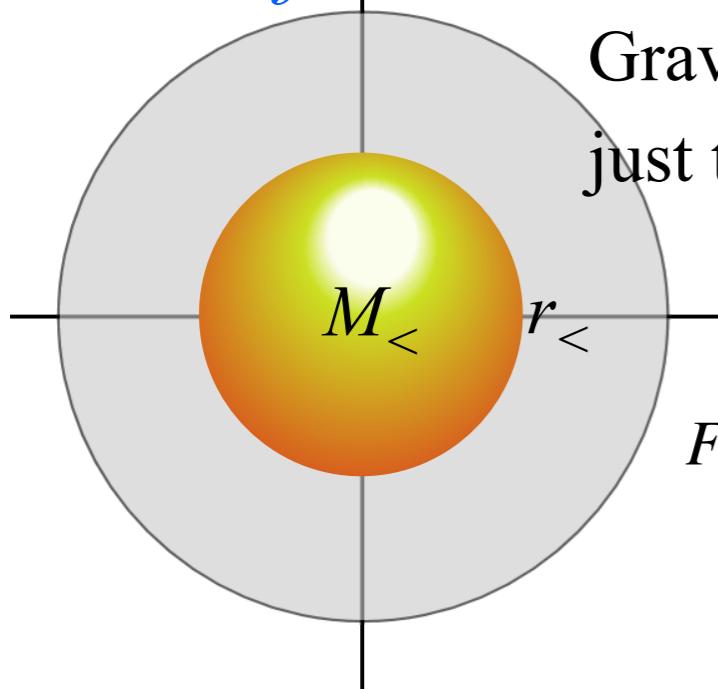
Gravitational force at $r_<$ is just that of planet $M_<$ below $r_<$

Coulomb force vanishes inside-spherical shell (Gauss-law)

Unit 1
Fig. 9.6



Coulomb force inside-spherical body due to stuff below you, only.

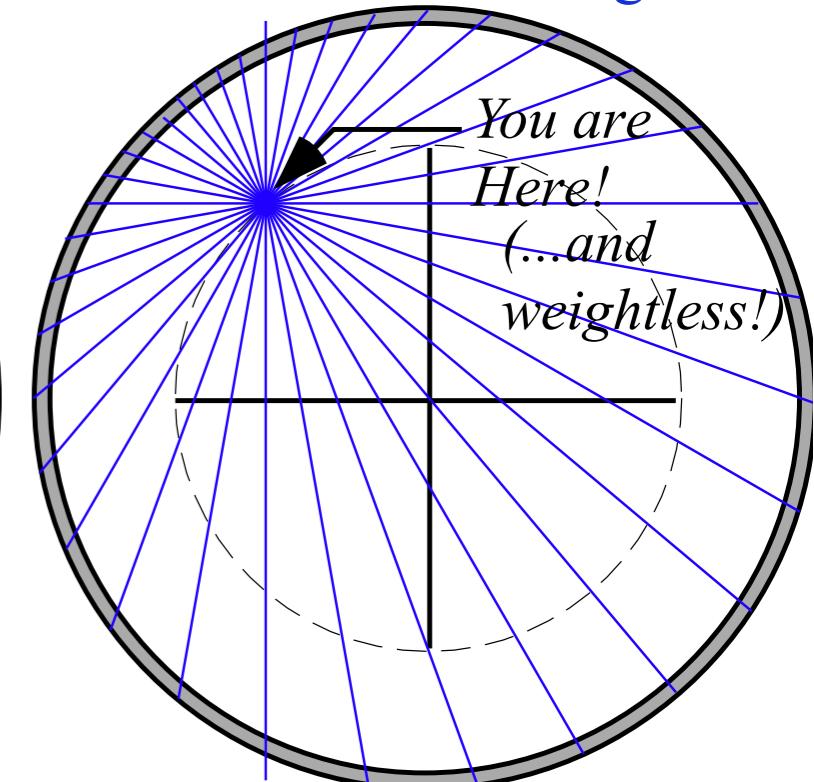
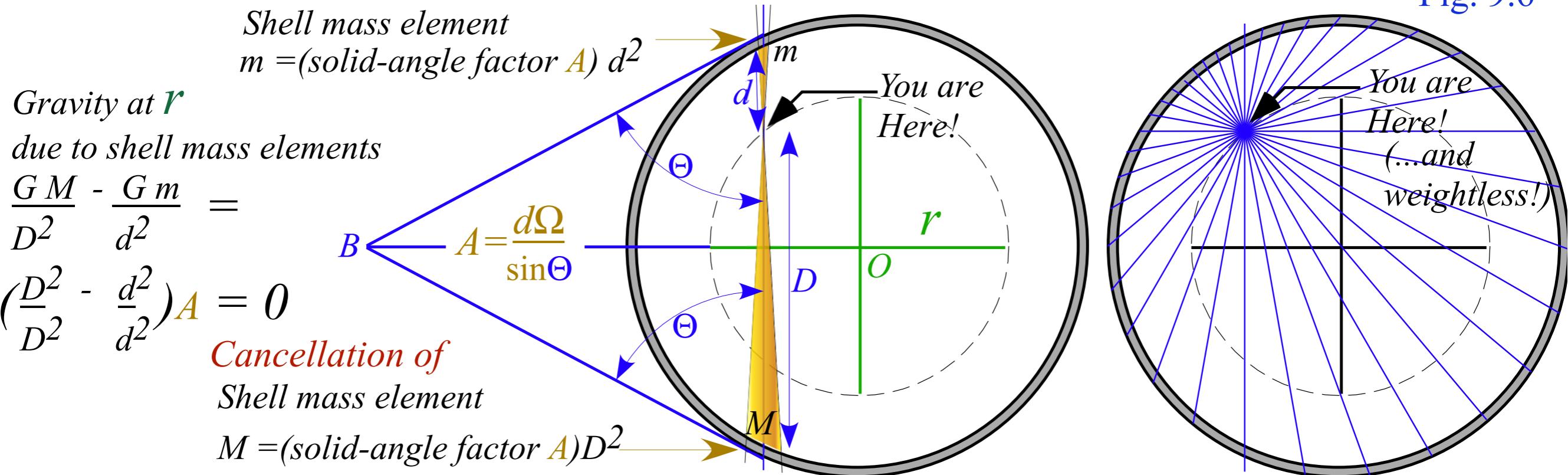


Gravitational force at $r_<$ is just that of planet $M_<$ below $r_<$

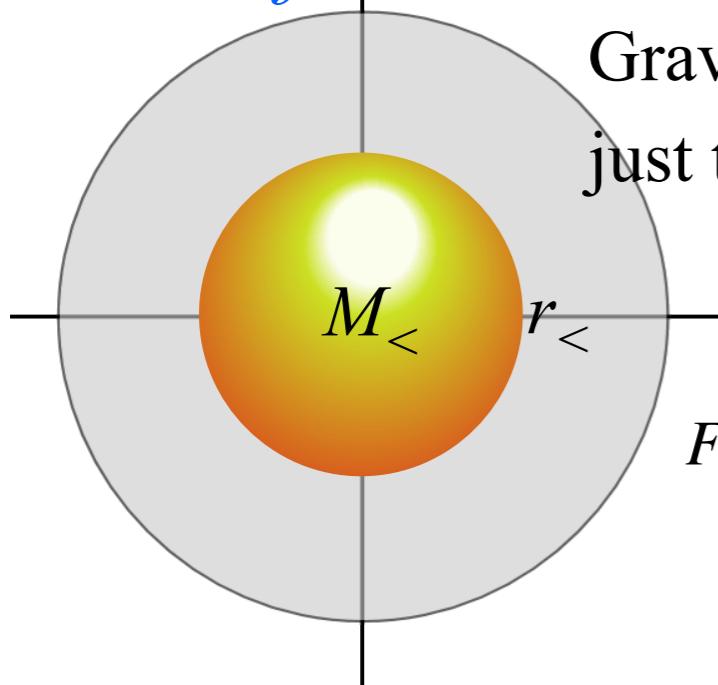
$$F^{inside}(r_<) = G \frac{mM_<}{r_<} = Gm \frac{4\pi}{3} \frac{M_<}{\frac{4\pi}{3} r_<} r_<$$

Coulomb force vanishes inside-spherical shell (Gauss-law)

Unit 1
Fig. 9.6



Coulomb force inside-spherical body due to stuff below you, only.



Gravitational force at $r_<$ is just that of planet $m_<$ below $r_<$

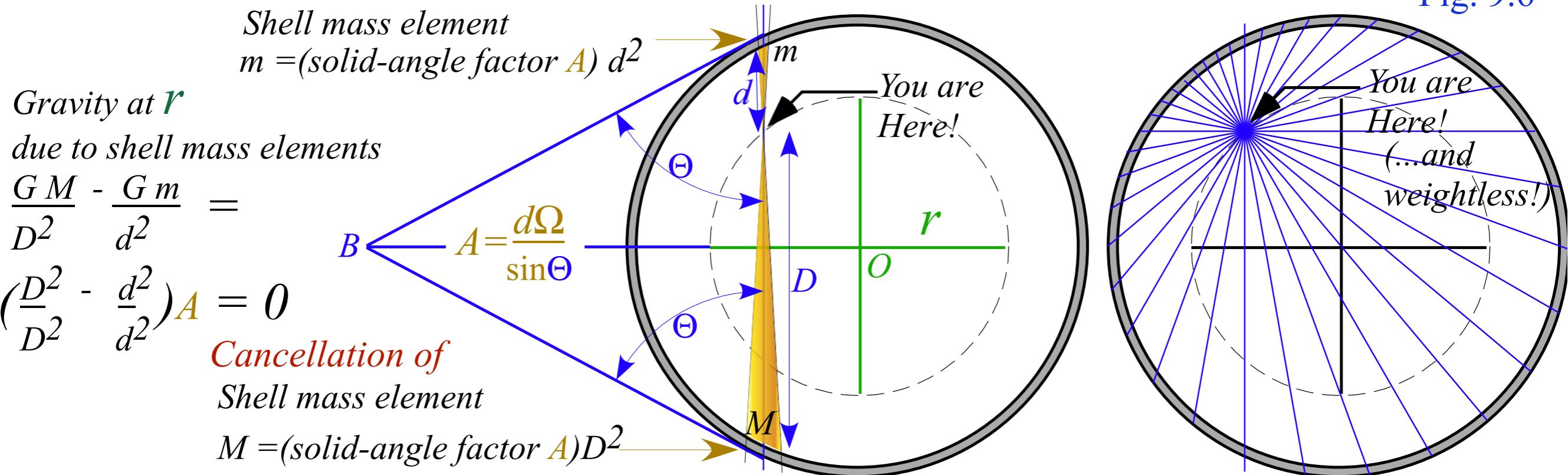
$$F^{inside}(r_<) = G \frac{mM_<}{r_<} = Gm \frac{4\pi}{3} \frac{M_<}{\frac{4\pi}{3}r_<} r_< = Gm \frac{4\pi}{3} \rho_{\oplus} r_<$$

Note:
Hooke's (linear) force law
for $r_<$ inside uniform body

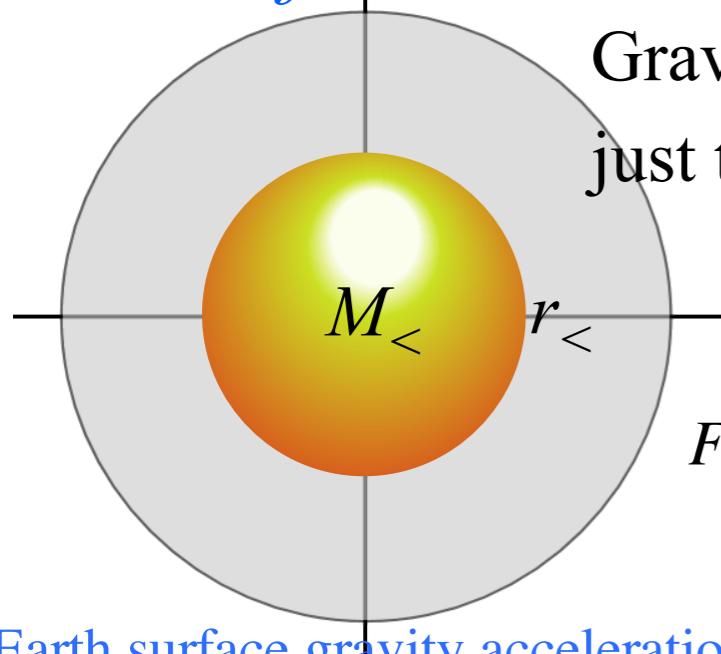
$$\downarrow \quad \downarrow \\ F^{inside}(r_<) = m\mathbf{g} \frac{r_<}{R_{\oplus}} \equiv m\mathbf{g} \cdot \mathbf{x}$$

Coulomb force vanishes inside-spherical shell (Gauss-law)

Unit 1
Fig. 9.6



Coulomb force inside-spherical body due to stuff below you, only.



Gravitational force at $r_<$ is just that of planet $m_<$ below $r_<$

Note:
Hooke's (linear) force law for $r_<$ inside uniform body

$$F_{\text{inside}}(r_<) = G \frac{m M_<}{r_<} = G m \frac{4\pi}{3} \frac{M_<}{4\pi r_<} r_< = G m \frac{4\pi}{3} \rho_+ r_<$$

Earth surface gravity acceleration: $g = G \frac{M_+}{R_+^2} = G \frac{M_+}{R_+^3} R_+ = G \frac{4\pi}{3} \frac{M_+}{4\pi R_+^3} R_+ = G \frac{4\pi}{3} \rho_+ R_+ = 9.8 \text{ m/s}^2$

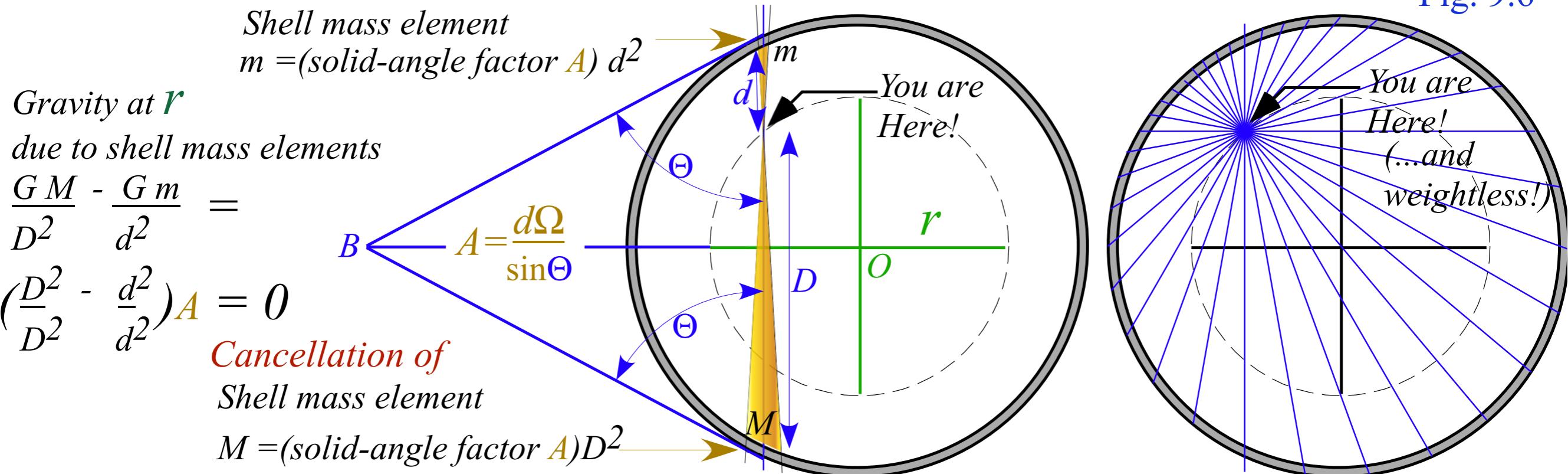
$$G = 6.67384(80) \cdot 10^{-11} \text{ Nm}^2/\text{C}^2 \sim (2/3) 10^{-10}$$

$$\downarrow \quad \downarrow$$

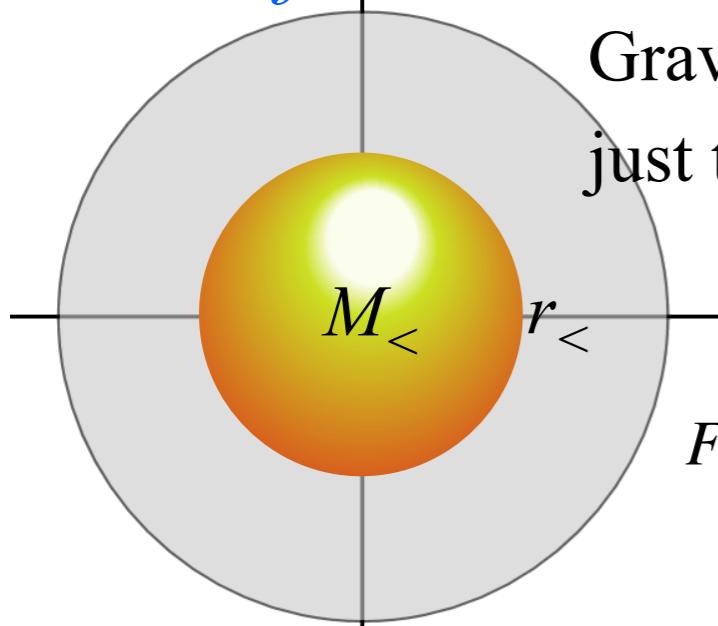
$$m g \frac{r_<}{R_+} \equiv m g \cdot x$$

Coulomb force vanishes inside-spherical shell (Gauss-law)

Unit 1
Fig. 9.6



Coulomb force inside-spherical body due to stuff below you, only.



Gravitational force at $r_<$ is
just that of planet $m_<$ below $r_<$

$$F^{inside}(r_<) = G \frac{m M_<}{r_<^2} = G m \frac{4\pi}{3} \frac{M_<}{4\pi r_<} r_< = G m \frac{4\pi}{3} \rho_< r_<^3$$

Earth surface gravity acceleration: $g = G \frac{M_\oplus}{R_\oplus^2} = G \frac{M_\oplus}{R_\oplus^3} R_\oplus = G \frac{4\pi}{3} \frac{M_\oplus}{4\pi R_\oplus^3} R_\oplus = G \frac{4\pi}{3} \rho_\oplus R_\oplus = 9.8 \text{ m/s}^2$

$G = 6.67384(80) \cdot 10^{-11} \text{ Nm}^2/\text{C}^2 \sim (2/3) 10^{-10}$

Earth radius: $R_\oplus = 6.371 \cdot 10^6 \text{ m} \approx 6.4 \cdot 10^6 \text{ m}$

Earth mass: $M_\oplus = 5.9722 \times 10^{24} \text{ kg} \approx 6.0 \cdot 10^{24} \text{ kg}$

Note:
Hooke's (linear) force law
for $r_<$ inside uniform body

$$\downarrow \quad \downarrow$$

$$F = m g \frac{r_<}{R_\oplus} = m g \cdot x$$

Solar radius: $R_\odot = 6.955 \times 10^8 \text{ m} \approx 7.0 \cdot 10^8 \text{ m}$

Solar mass: $M_\odot = 1.9889 \times 10^{30} \text{ kg} \approx 2.0 \cdot 10^{30} \text{ kg}$

Geometry of idealized “Sophomore-physics Earth”

Coulomb field outside

Isotropic Harmonic Oscillator (IHO) field inside

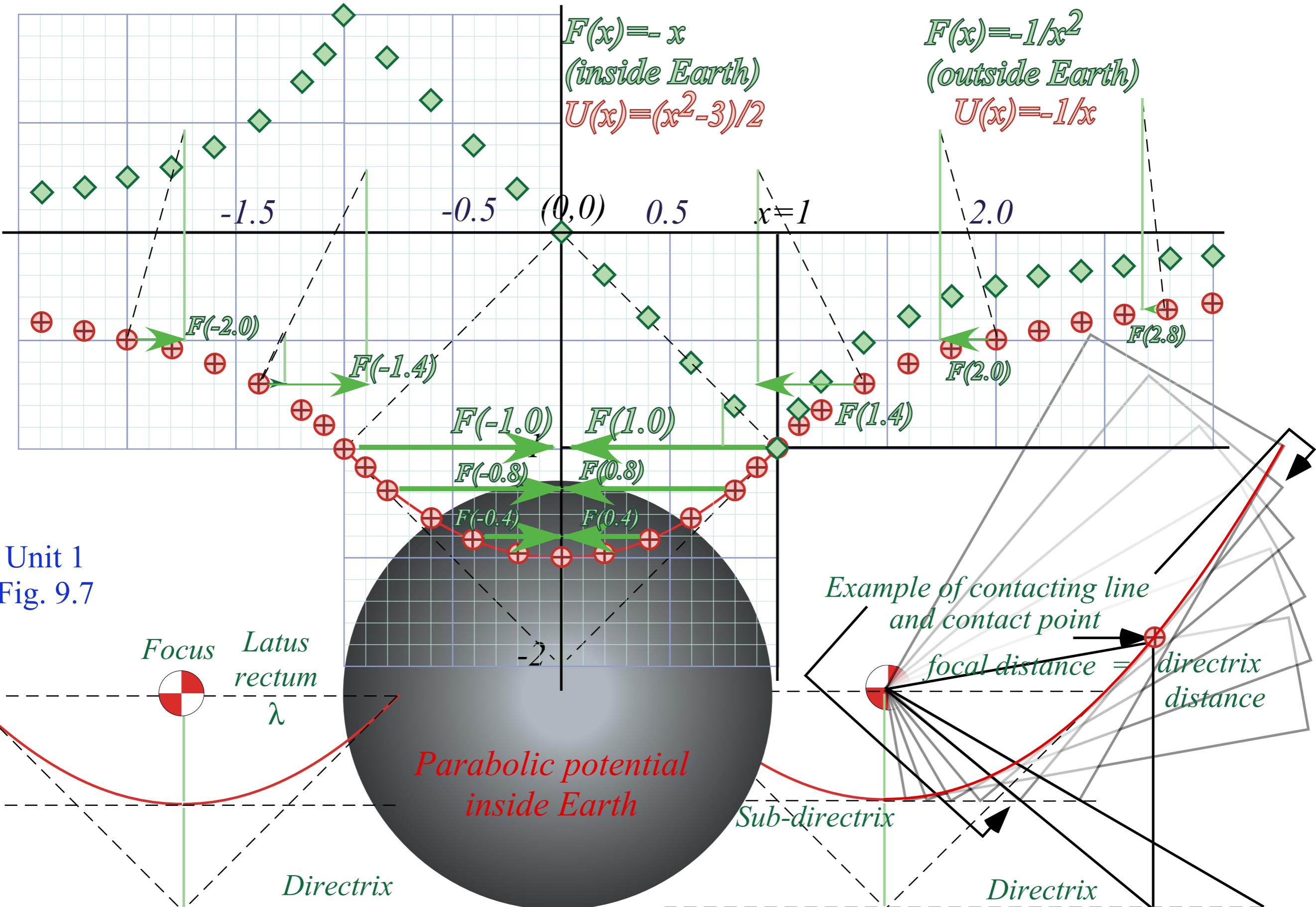
→ *Contact-geometry of potential curve(s)*

“Crushed-Earth” models: 3 key energy “steps” and 4 key energy “levels”

Earth matter vs nuclear matter:

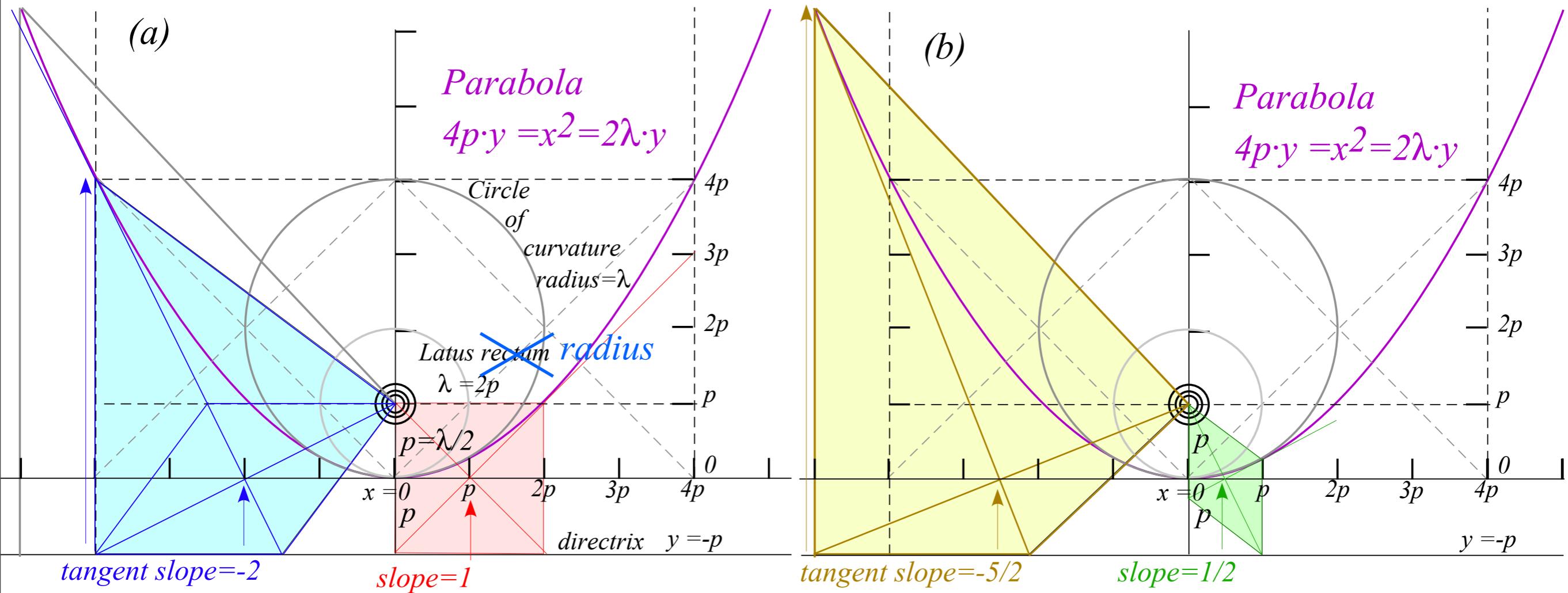
*Introducing the “neutron starlet” and “**Black-Hole-Earth**”*

The ideal “Sophomore-Physics-Earth” model of geo-gravity



...conventional parabolic geometry...carried to extremes...

(Review of Lect. 6 p.29)



Unit 1
Fig. 9.4

Geometry of idealized “Sophomore-physics Earth”

Coulomb field outside

Isotropic Harmonic Oscillator (IHO) field inside

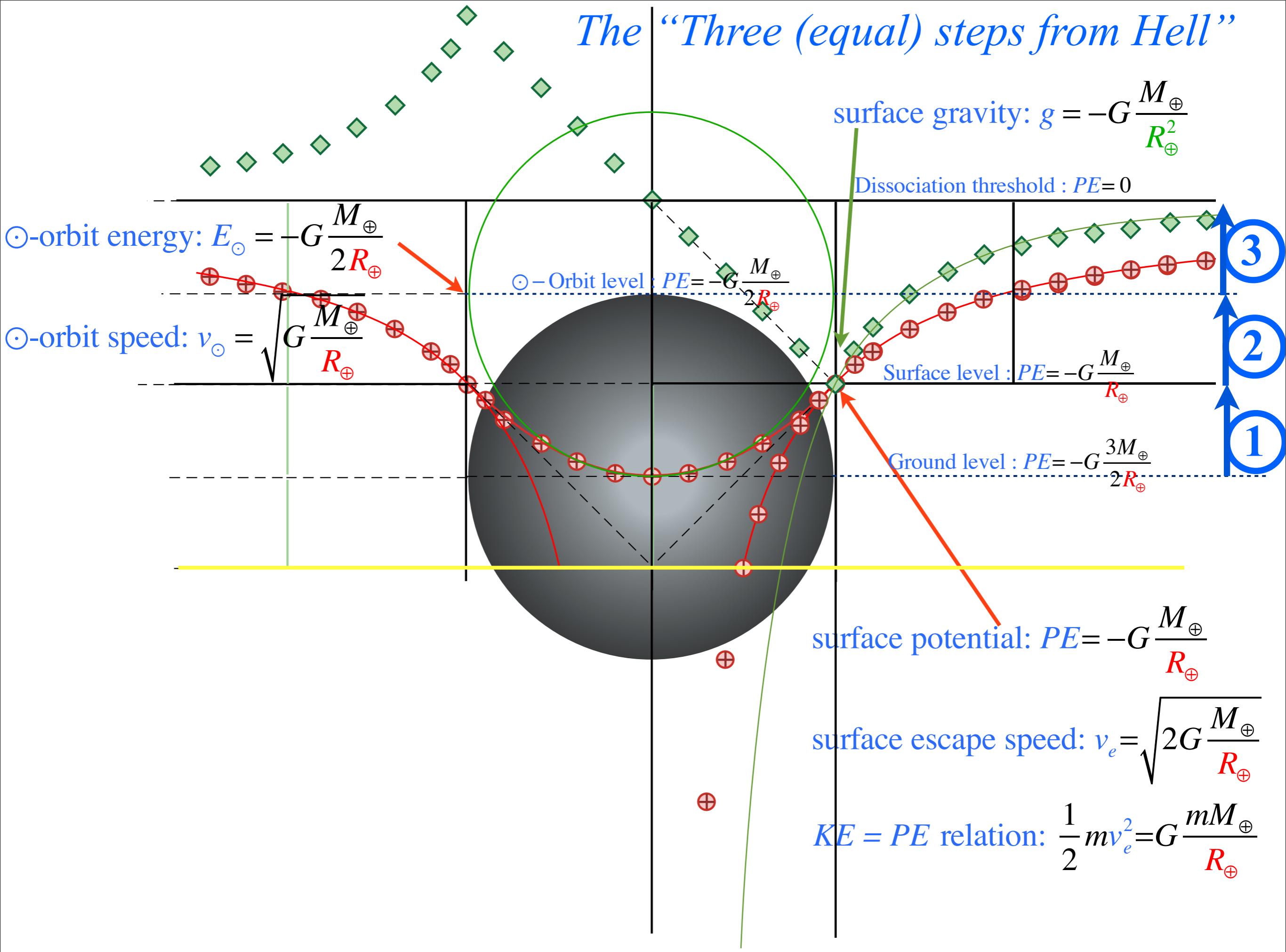
Contact-geometry of potential curve(s)

→ “Crushed-Earth” models: 3 key energy “steps” and 4 key energy “levels”

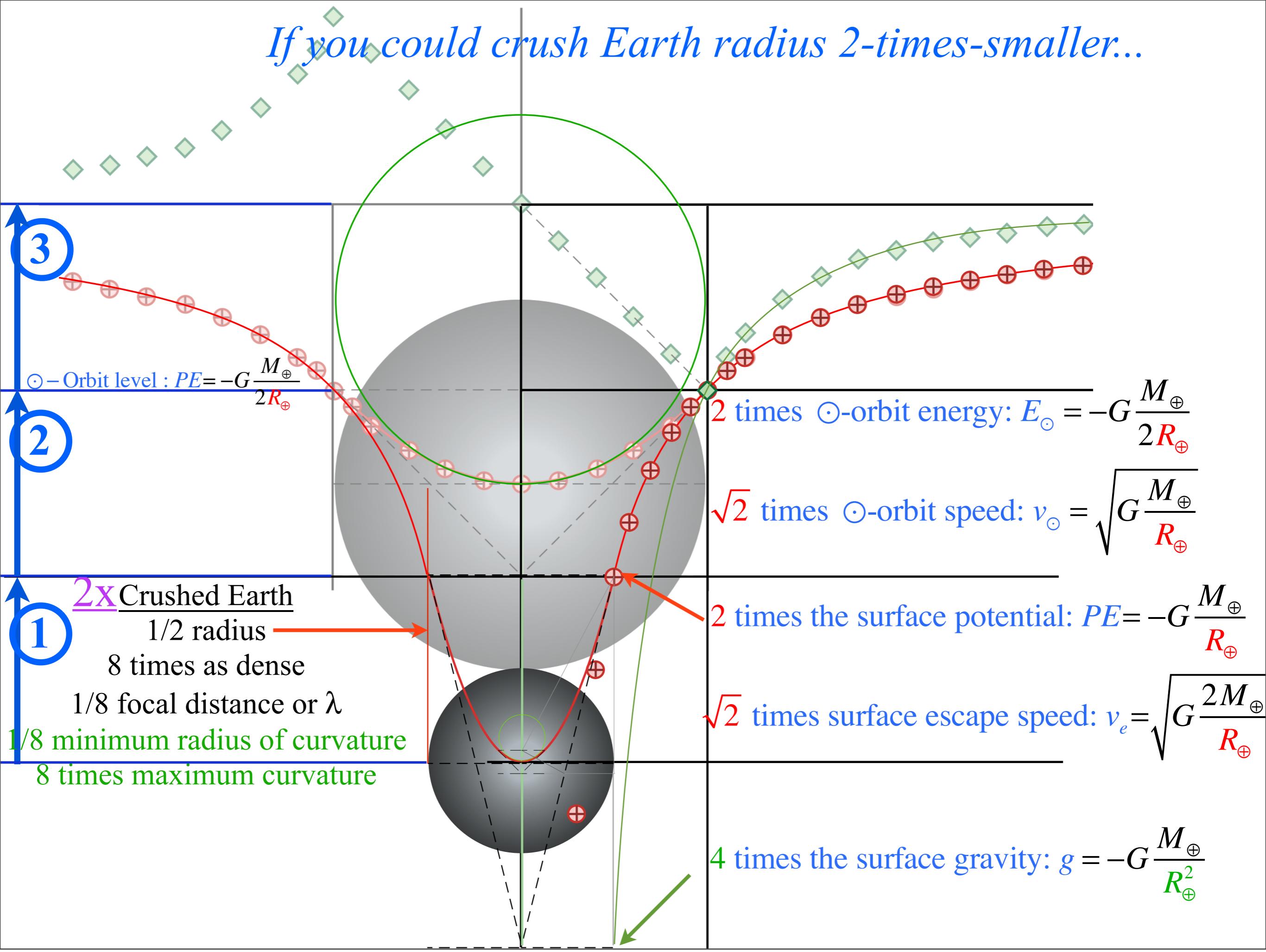
Earth matter vs nuclear matter:

*Introducing the “neutron starlet” and “**Black-Hole-Earth**”*

The “Three (equal) steps from Hell”



If you could crush Earth radius 2-times-smaller...



Geometry of idealized “Sophomore-physics Earth”

Coulomb field outside

Isotropic Harmonic Oscillator (IHO) field inside

Contact-geometry of potential curve(s)

“Crushed-Earth” models: 3 key energy “steps” and 4 key energy “levels”

Earth matter vs nuclear matter:



*Introducing the “neutron starlet” and “**Black-Hole-Earth**”*

Examples of “crushed” matter

Earth matter Earth mass : $M_{\oplus} = 5.9722 \times 10^{24} \text{ kg.} \approx 6.0 \cdot 10^{24} \text{ kg.}$ Density $\rho_{\oplus} ??$

Earth radius : $R_{\oplus} = 6.371 \cdot 10^6 \text{ m} \approx 6.4 \cdot 10^6 \text{ m}$ Earth volume : $(4\pi / 3)R_{\oplus}^3 \approx 4 \cdot 260 \cdot 10^{18} \approx 10^{21} \text{ m}^3$

$$(6.4)^3 \approx 262 \text{ and } (4\pi/3)260 = 1098 \approx 10^3$$

Examples of “crushed” matter

Earth matter Earth mass : $M_{\oplus} = 5.9722 \times 10^{24} \text{ kg.} \approx 6.0 \cdot 10^{24} \text{ kg.}$ Density $\rho_{\oplus} \sim 6.0 \cdot 10^{24-21} \sim 6 \cdot 10^3 \text{ kg/m}^3$

Earth radius : $R_{\oplus} = 6.371 \cdot 10^6 \text{ m} \approx 6.4 \cdot 10^6 \text{ m}$ Earth volume : $(4\pi / 3)R_{\oplus}^3 \approx 4 \cdot 260 \cdot 10^{18} \sim 10^{21} \text{ m}^3$

Examples of “crushed” matter

Earth matter Earth mass : $M_{\oplus} = 5.9722 \times 10^{24} \text{ kg.} \approx 6.0 \cdot 10^{24} \text{ kg.}$ Density $\rho_{\oplus} \sim 6.0 \cdot 10^{24-21} \sim 6 \cdot 10^3 \text{ kg/m}^3$

Earth radius : $R_{\oplus} = 6.371 \cdot 10^6 \text{ m} \approx 6.4 \cdot 10^6 \text{ m}$ Earth volume : $(4\pi / 3)R_{\oplus}^3 \approx 4 \cdot 260 \cdot 10^{18} \sim 10^{21} \text{ m}^3$

Density of solid Fe= $7.9 \cdot 10^3 \text{ kg/m}^3$
Density of liquid Fe= $6.9 \cdot 10^3 \text{ kg/m}^3$

Examples of “crushed” matter

Earth matter Earth mass : $M_{\oplus} = 5.9722 \times 10^{24} \text{ kg.} \approx 6.0 \cdot 10^{24} \text{ kg.}$ Density $\rho_{\oplus} \sim 6.0 \cdot 10^{24-21} \sim 6 \cdot 10^3 \text{ kg/m}^3$

Earth radius : $R_{\oplus} = 6.371 \cdot 10^6 \text{ m} \approx 6.4 \cdot 10^6 \text{ m}$ Earth volume : $(4\pi / 3)R_{\oplus}^3 \approx 4 \cdot 260 \cdot 10^{18} \sim 10^{21} \text{ m}^3$

Nuclear matter Nucleon mass = $1.67 \cdot 10^{-27} \text{ kg.} \sim 2 \cdot 10^{-27} \text{ kg.}$

Say a nucleus of atomic weight 50 has a radius of 3 fm , or 50 nucleons each with a mass $2 \cdot 10^{-27} \text{ kg.}$

Examples of “crushed” matter

Earth matter Earth mass : $M_{\oplus} = 5.9722 \times 10^{24} \text{ kg.} \approx 6.0 \cdot 10^{24} \text{ kg.}$ Density $\rho_{\oplus} \sim 6.0 \cdot 10^{24-21} \sim 6 \cdot 10^3 \text{ kg/m}^3$

Earth radius : $R_{\oplus} = 6.371 \cdot 10^6 \text{ m} \approx 6.4 \cdot 10^6 \text{ m}$ Earth volume : $(4\pi/3)R_{\oplus}^3 \approx 4 \cdot 260 \cdot 10^{18} \sim 10^{21} \text{ m}^3$

Nuclear matter Nucleon mass = $1.67 \cdot 10^{-27} \text{ kg.} \sim 2 \cdot 10^{-27} \text{ kg.}$

Say a nucleus of atomic weight 50 has a radius of 3 fm , or 50 nucleons each with a mass $2 \cdot 10^{-27} \text{ kg.}$

That's $100 \cdot 10^{-27} = 10^{-25} \text{ kg}$ packed into a volume of $4\pi/3r^3 = 4\pi/3 (3 \cdot 10^{-15})^3 \text{ m}^3$ or about $10^{-43} \text{ m}^3.$

$$36\pi=113 \sim 10^2$$

Examples of “crushed” matter

Earth matter Earth mass : $M_{\oplus} = 5.9722 \times 10^{24} \text{ kg.} \approx 6.0 \cdot 10^{24} \text{ kg.}$ Density $\rho_{\oplus} \sim 6.0 \cdot 10^{24-21} \sim 6 \cdot 10^3 \text{ kg/m}^3$

Earth radius : $R_{\oplus} = 6.371 \cdot 10^6 \text{ m} \approx 6.4 \cdot 10^6 \text{ m}$ Earth volume : $(4\pi/3)R_{\oplus}^3 \approx 4 \cdot 260 \cdot 10^{18} \sim 10^{21} \text{ m}^3$

Nuclear matter Nucleon mass = $1.67 \cdot 10^{-27} \text{ kg.} \sim 2 \cdot 10^{-27} \text{ kg.}$

Say a nucleus of atomic weight 50 has a radius of 3 fm , or 50 nucleons each with a mass $2 \cdot 10^{-27} \text{ kg.}$

That's $100 \cdot 10^{-27} = 10^{-25} \text{ kg}$ packed into a volume of $4\pi/3r^3 = 4\pi/3 (3 \cdot 10^{-15})^3 \text{ m}^3$ or about 10^{-43} m^3 .

Nuclear density is $10^{-25+43} = 10^{18} \text{ kg/m}^3$ or a *trillion* (10^{12}) kilograms in the size of a fingertip (1cc).

Earth radius crushed by a factor of $0.5 \cdot 10^{-5}$ to $R_{\text{crush}\oplus} \approx 300 \text{ m}$ would approach neutron-star density.

Examples of “crushed” matter

Earth matter Earth mass : $M_{\oplus} = 5.9722 \times 10^{24} \text{ kg.} \approx 6.0 \cdot 10^{24} \text{ kg.}$ Density $\rho_{\oplus} \sim 6.0 \cdot 10^{24-21} \sim 6 \cdot 10^3 \text{ kg/m}^3$

Earth radius : $R_{\oplus} = 6.371 \cdot 10^6 \text{ m} \approx 6.4 \cdot 10^6 \text{ m}$ Earth volume : $(4\pi/3)R_{\oplus}^3 \approx 4 \cdot 260 \cdot 10^{18} \sim 10^{21} \text{ m}^3$

Nuclear matter Nucleon mass = $1.67 \cdot 10^{-27} \text{ kg.} \sim 2 \cdot 10^{-27} \text{ kg.}$

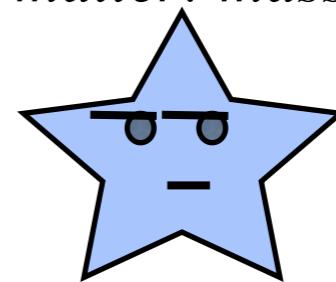
Say a nucleus of atomic weight 50 has a radius of 3 fm , or 50 nucleons each with a mass $2 \cdot 10^{-27} \text{ kg.}$

That's $100 \cdot 10^{-27} = 10^{-25} \text{ kg}$ packed into a volume of $4\pi/3r^3 = 4\pi/3 (3 \cdot 10^{-15})^3 \text{ m}^3$ or about 10^{-43} m^3 .

Nuclear density is $10^{-25+43} = 10^{18} \text{ kg/m}^3$ or a trillion (10^{12}) kilograms in the size of a fingertip.

Earth radius crushed by a factor of $0.5 \cdot 10^{-5}$ to $R_{\text{crush}\oplus} \approx 300 \text{ m}$ would approach neutron-star density.

Introducing the “Neutron starlet” 1 cm^3 of nuclear matter: mass = 10^{12} kg.



Examples of “crushed” matter

Earth matter Earth mass : $M_{\oplus} = 5.9722 \times 10^{24} \text{ kg.} \approx 6.0 \cdot 10^{24} \text{ kg.}$ Density $\rho_{\oplus} \sim 6.0 \cdot 10^{24-21} \sim 6 \cdot 10^3 \text{ kg/m}^3$

Earth radius : $R_{\oplus} = 6.371 \cdot 10^6 \text{ m} \approx 6.4 \cdot 10^6 \text{ m}$ Earth volume : $(4\pi/3)R_{\oplus}^3 \approx 4 \cdot 260 \cdot 10^{18} \sim 10^{21} \text{ m}^3$

Nuclear matter Nucleon mass = $1.67 \cdot 10^{-27} \text{ kg.} \sim 2 \cdot 10^{-27} \text{ kg.}$

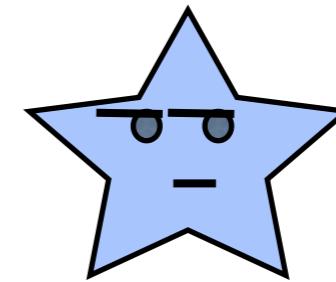
Say a nucleus of atomic weight 50 has a radius of 3 fm , or 50 nucleons each with a mass $2 \cdot 10^{-27} \text{ kg.}$

That's $100 \cdot 10^{-27} = 10^{-25} \text{ kg}$ packed into a volume of $4\pi/3r^3 = 4\pi/3 (3 \cdot 10^{-15})^3 \text{ m}^3$ or about 10^{-43} m^3 .

Nuclear density is $10^{-25+43} = 10^{18} \text{ kg/m}^3$ or a trillion (10^{12}) kilograms in the size of a fingertip.

Earth radius crushed by a factor of $0.5 \cdot 10^{-5}$ to $R_{\text{crush}\oplus} \approx 300 \text{ m}$ would approach neutron-star density.

Introducing the “Neutron starlet” 1 cm^3 of nuclear matter: mass = 10^{12} kg.



Introducing the “Black Hole Earth” Suppose Earth is crushed so that its

surface escape velocity is the speed of light $c \cong 3.0 \cdot 10^8 \text{ m/s.}$

$c \equiv 299,792,458 \text{ m/s (EXACTLY)}$

$$c = \sqrt{(2GM/R_{\odot})}$$

$$R_{\odot} = 2GM/c^2 = 8.9 \text{ mm} \sim 1 \text{ cm} \quad (\text{fingertip size!})$$

$$G = 6.67384(80) \cdot 10^{-11} \text{ Nm}^2/\text{C}^2 \sim (2/3)10^{-10}$$

Isotropic harmonic oscillator dynamics in 1D, 2D, and 3D

→ *Sinusoidal space-time dynamics derived by geometry*

Isotropic harmonic oscillator orbits in 1D and 2D (You get 3D for free!)

Constructing 2D Isotropic harmonic oscillator orbits using phasor plots

Examples with x-y phase lag : $\alpha_{x-y} = \alpha_x - \alpha_y = 15^\circ, 30^\circ, \text{ and } \pm 75^\circ$

[BoxIt simulation of U\(2\) orbits](#)

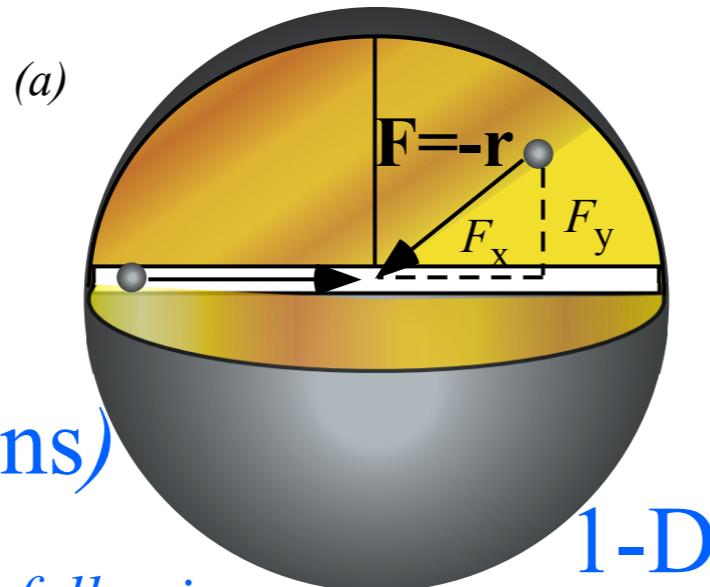
<http://www.uark.edu/ua/modphys/markup/BoxItWeb.html>

Isotropic Harmonic Oscillator phase dynamics in uniform-body

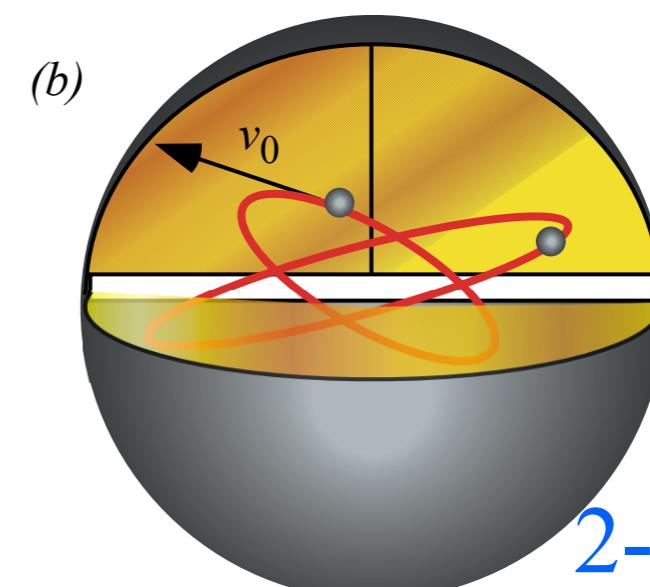
I.H.O. Force law

$$F = -x \quad (\text{1-Dimension})$$

$$\mathbf{F} = -\mathbf{r} \quad (\text{2 or 3-Dimensions})$$



1-D



2-D

Unit 1
Fig. 9.10

Each dimension x , y , or z obeys the following:

$$\text{Total } E = KE + PE = \frac{1}{2}mv^2 + U(x) = \frac{1}{2}mv^2 + \frac{1}{2}kx^2 = \text{const.}$$

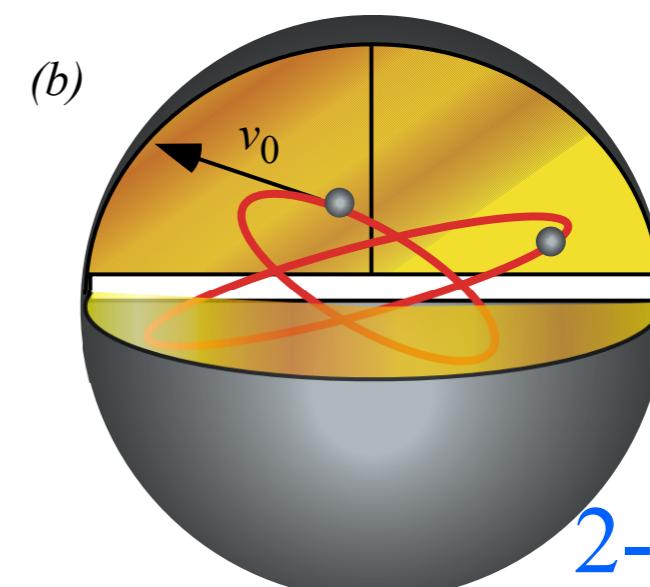
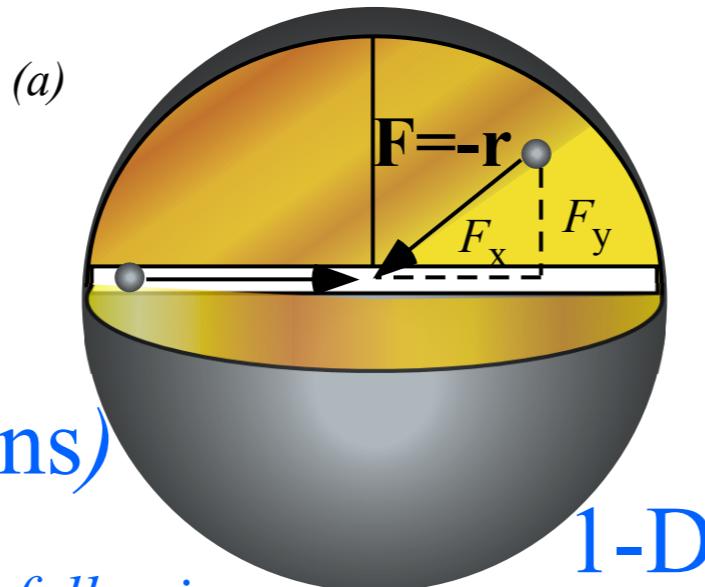
(Paths are always
2-D ellipses if
viewed right!)

Isotropic Harmonic Oscillator phase dynamics in uniform-body

I.H.O. Force law

$$F = -x \quad (\text{1-Dimension})$$

$$\mathbf{F} = -\mathbf{r} \quad (\text{2 or 3-Dimensions})$$



Unit 1
Fig. 9.10

Each dimension x , y , or z obeys the following:

$$\text{Total } E = KE + PE = \frac{1}{2}mv^2 + U(x) = \frac{1}{2}mv^2 + \frac{1}{2}kx^2 = \text{const.}$$

Equations for x -motion

[$x(t)$ and $v_x = v(t)$] are given first. They apply as well to dimensions [$y(t)$ and $v_y = v(t)$] and [$z(t)$ and $v_z = v(t)$] in the ideal isotropic case.

2-D or 3-D

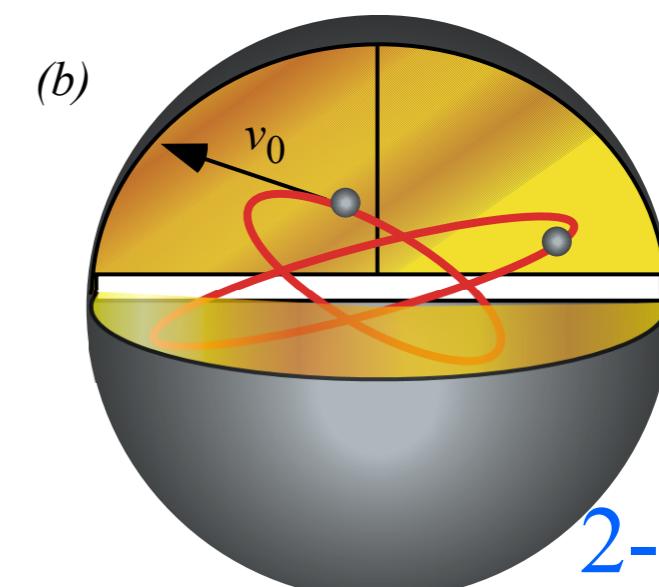
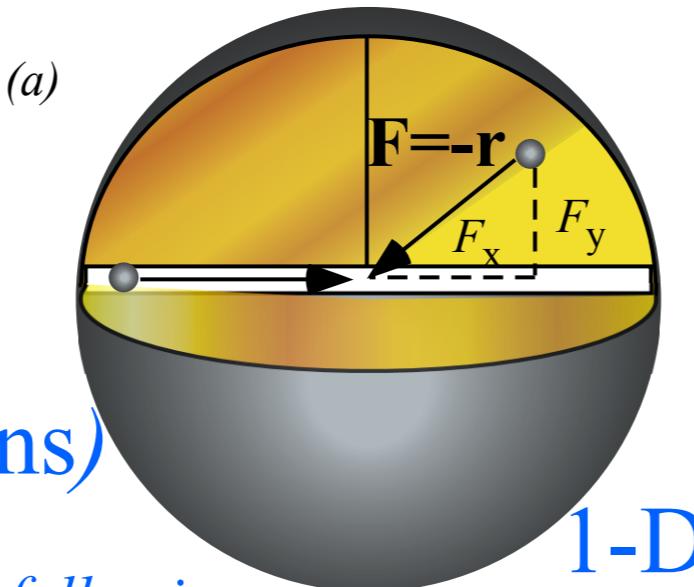
(Paths are always 2-D ellipses if viewed right!)

Isotropic Harmonic Oscillator phase dynamics in uniform-body

I.H.O. Force law

$$F = -x \quad (\text{1-Dimension})$$

$$\mathbf{F} = -\mathbf{r} \quad (\text{2 or 3-Dimensions})$$



Unit 1
Fig. 9.10

Each dimension x, y , or z obeys the following:

$$\text{Total } E = KE + PE = \frac{1}{2}mv^2 + U(x) = \frac{1}{2}mv^2 + \frac{1}{2}kx^2 = \text{const.}$$

Equations for x -motion

[$x(t)$ and $v_x = v(t)$] are given first. They apply as well to dimensions [$y(t)$ and $v_y = v(t)$] and [$z(t)$ and $v_z = v(t)$] in the ideal isotropic case.

$$1 = \frac{mv^2}{2E} + \frac{kx^2}{2E} = \left(\frac{v}{\sqrt{2E/m}} \right)^2 + \left(\frac{x}{\sqrt{2E/k}} \right)^2$$

$$1 = \frac{mv^2}{2E} + \frac{kx^2}{2E} = (\cos\theta)^2 + (\sin\theta)^2$$

Another example of the old “scale-a-circle” trick...

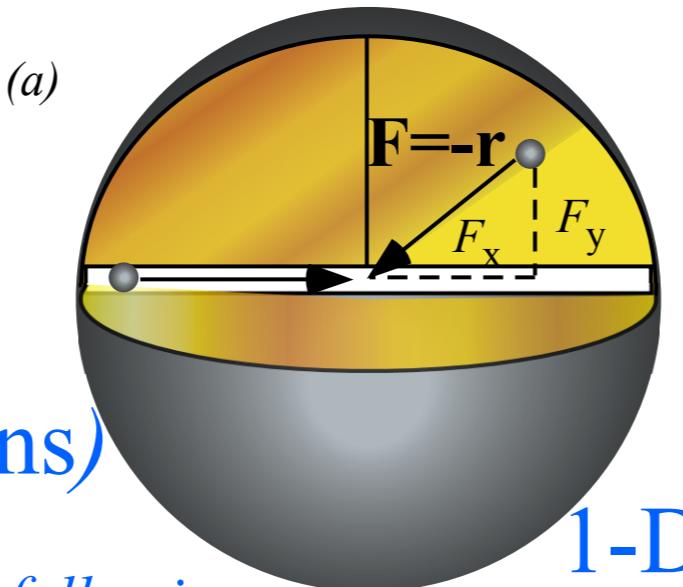
$$\text{Let : (1)} \quad v = \sqrt{2E/m} \cos\theta, \quad \text{and : (2)} \quad x = \sqrt{2E/k} \sin\theta$$

Isotropic Harmonic Oscillator phase dynamics in uniform-body

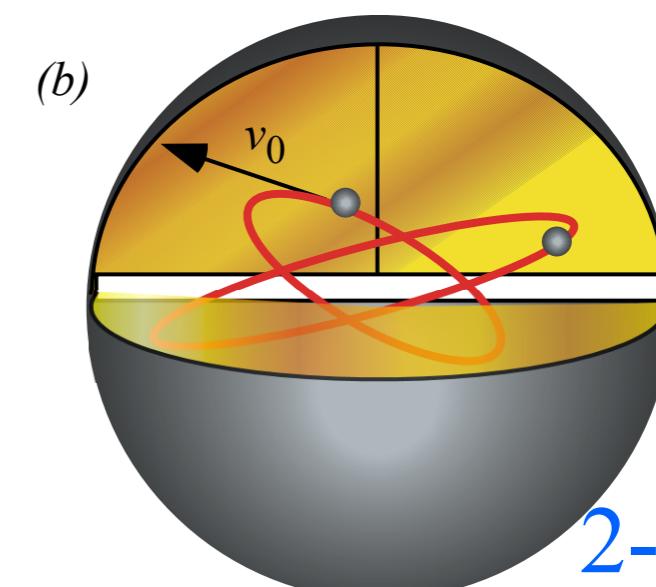
I.H.O. Force law

$$F = -x \quad (\text{1-Dimension})$$

$$\mathbf{F} = -\mathbf{r} \quad (\text{2 or 3-Dimensions})$$



1-D



Unit 1
Fig. 9.10

2-D or 3-D

(Paths are always 2-D ellipses if viewed right!)

Each dimension x, y , or z obeys the following:

$$\text{Total } E = KE + PE = \frac{1}{2}mv^2 + U(x) = \frac{1}{2}mv^2 + \frac{1}{2}kx^2 = \text{const.}$$

Equations for x -motion

[$x(t)$ and $v_x = v(t)$] are given first. They apply as well to dimensions [$y(t)$ and $v_y = v(t)$] and [$z(t)$ and $v_z = v(t)$] in the ideal isotropic case.

$$1 = \frac{mv^2}{2E} + \frac{kx^2}{2E} = \left(\frac{v}{\sqrt{2E/m}} \right)^2 + \left(\frac{x}{\sqrt{2E/k}} \right)^2$$

$$1 = \frac{mv^2}{2E} + \frac{kx^2}{2E} = (\cos\theta)^2 + (\sin\theta)^2$$

Another example of the old “scale-a-circle” trick...

$$\text{Let : (1)} \quad v = \sqrt{2E/m} \cos\theta, \quad \text{and : (2)} \quad x = \sqrt{2E/k} \sin\theta$$

$$\text{def. (3)} \quad \omega = \frac{d\theta}{dt}$$

$$\sqrt{\frac{2E}{m}} \cos\theta = v = \frac{dx}{dt} = \frac{d\theta}{dt} \frac{dx}{d\theta} = \omega \frac{dx}{d\theta}$$

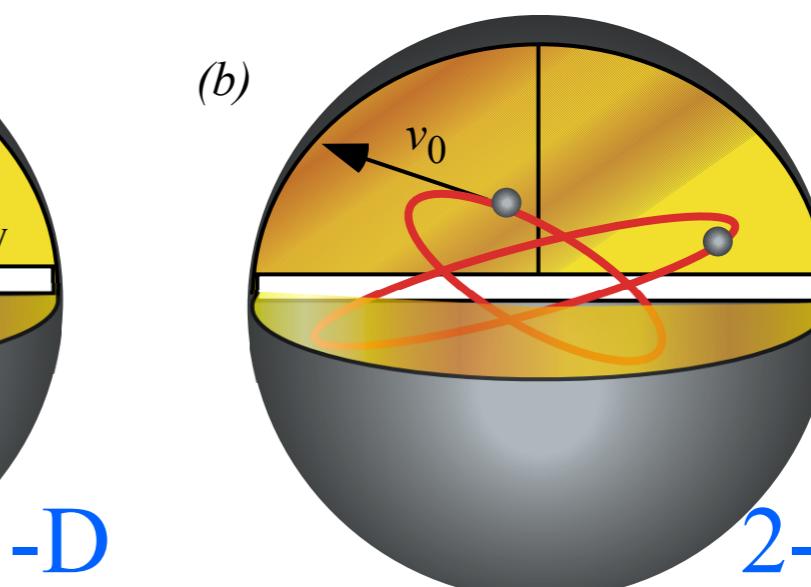
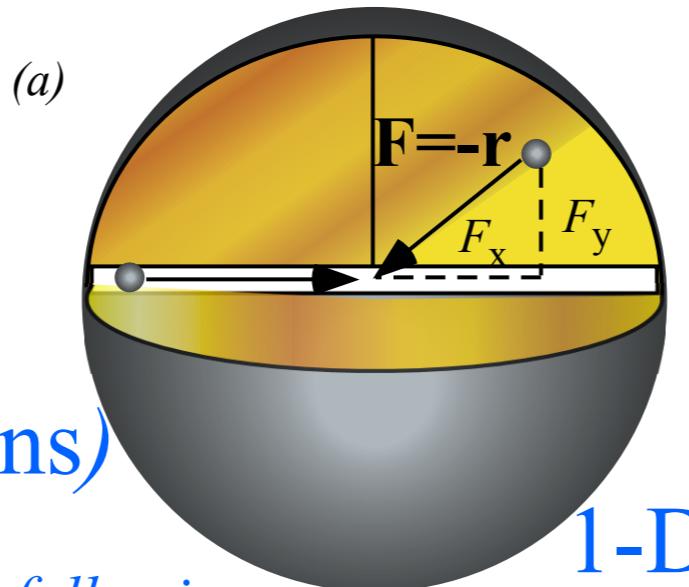
by (1) by def. (3)

Isotropic Harmonic Oscillator phase dynamics in uniform-body

I.H.O. Force law

$$F = -x \quad (\text{1-Dimension})$$

$$\mathbf{F} = -\mathbf{r} \quad (\text{2 or 3-Dimensions})$$



Unit 1
Fig. 9.10

1-D

2-D or 3-D

(Paths are always 2-D ellipses if viewed right!)

Each dimension x, y , or z obeys the following:

$$\text{Total } E = KE + PE = \frac{1}{2}mv^2 + U(x) = \frac{1}{2}mv^2 + \frac{1}{2}kx^2 = \text{const.}$$

Equations for x -motion

[$x(t)$ and $v_x = v(t)$] are given first. They apply as well to dimensions [$y(t)$ and $v_y = v(t)$] and [$z(t)$ and $v_z = v(t)$] in the ideal isotropic case.

$$1 = \frac{mv^2}{2E} + \frac{kx^2}{2E} = \left(\frac{v}{\sqrt{2E/m}} \right)^2 + \left(\frac{x}{\sqrt{2E/k}} \right)^2$$

$$1 = \frac{mv^2}{2E} + \frac{kx^2}{2E} = (\cos\theta)^2 + (\sin\theta)^2$$

Another example of the old “scale-a-circle” trick...

$$\text{Let : (1)} \quad v = \sqrt{2E/m} \cos\theta, \quad \text{and : (2)} \quad x = \sqrt{2E/k} \sin\theta$$

$$\text{def. (3)} \quad \omega = \frac{d\theta}{dt}$$

$$\sqrt{\frac{2E}{m}} \cos\theta = v = \frac{dx}{dt} = \frac{d\theta}{dt} \frac{dx}{d\theta} = \omega \frac{dx}{d\theta} = \omega \sqrt{\frac{2E}{k}} \cos\theta$$

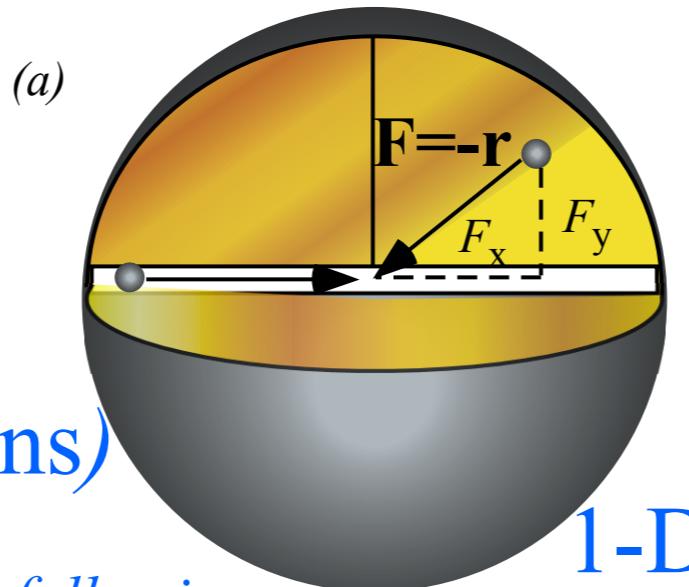
by (1) by def. (3) by (2)

Isotropic Harmonic Oscillator phase dynamics in uniform-body

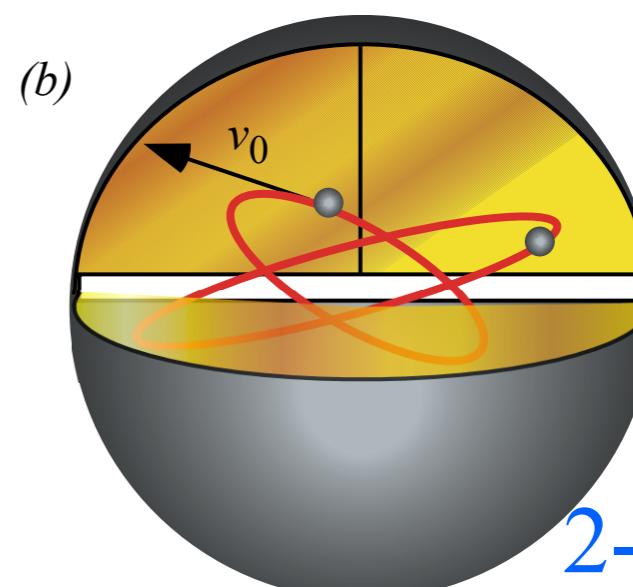
I.H.O. Force law

$F = -x$ (1-Dimension)

$\mathbf{F} = -\mathbf{r}$ (2 or 3-Dimensions)



1- Γ



Unit 1

Fig. 9.10

2-D or 3-D

(Paths are *always* 2-D ellipses if viewed right!)

Equations for x-motion

$[x(t) \text{ and } v_x=v(t)]$ are given first. They apply as well to dimensions $[y(t) \text{ and } v_y=v(t)]$ and $[z(t) \text{ and } v_z=v(t)]$ in the ideal isotropic case.

$$1 = \frac{mv^2}{2E} + \frac{kx^2}{2E} = \left(\frac{v}{\sqrt{2E/m}} \right)^2 + \left(\frac{x}{\sqrt{2E/k}} \right)^2$$

$$1 = \frac{mv^2}{2E} + \frac{kx^2}{2E} = (\cos\theta)^2 + (\sin\theta)^2$$

*Another example of
the old “scale-a-circle”
trick...*

Let : (1) $v = \sqrt{2E/m} \cos\theta$, and : (2) $x = \sqrt{2E/k} \sin\theta$

$$def.\,(3) \quad \omega = \frac{d\theta}{dt}$$

$$\sqrt{\frac{2E}{m}} \cos \theta = v = \frac{dx}{dt} = \frac{d\theta}{dt} \frac{dx}{d\theta} = \omega \frac{dx}{d\theta} = \omega \sqrt{\frac{2E}{k}} \cos \theta$$

by (1) *by def. (3)* *by (2)*

$$\omega = \frac{d\theta}{dt} = \sqrt{\frac{k}{m}}$$

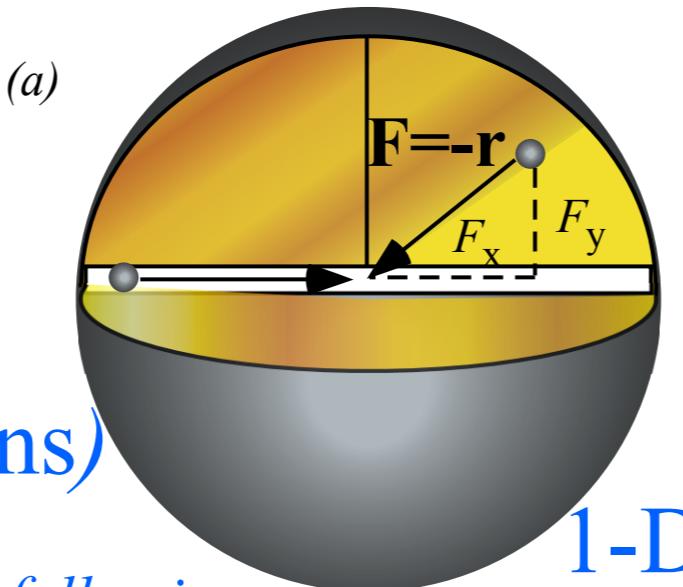
divide this by (1)

Isotropic Harmonic Oscillator phase dynamics in uniform-body

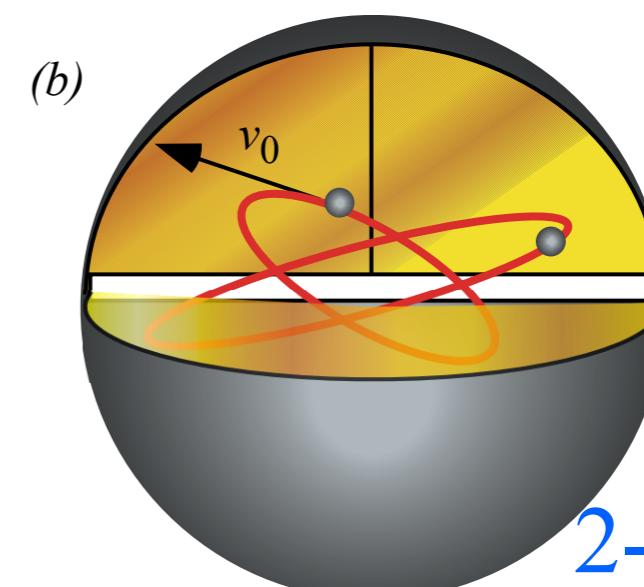
I.H.O. Force law

$$F = -x \quad (\text{1-Dimension})$$

$$\mathbf{F} = -\mathbf{r} \quad (\text{2 or 3-Dimensions})$$



1-D



Unit 1
Fig. 9.10

2-D or 3-D

(Paths are always 2-D ellipses if viewed right!)

Each dimension x, y , or z obeys the following:

$$\text{Total } E = KE + PE = \frac{1}{2}mv^2 + U(x) = \frac{1}{2}mv^2 + \frac{1}{2}kx^2 = \text{const.}$$

Equations for x -motion

[$x(t)$ and $v_x = v(t)$] are given first. They apply as well to dimensions [$y(t)$ and $v_y = v(t)$] and [$z(t)$ and $v_z = v(t)$] in the ideal isotropic case.

$$1 = \frac{mv^2}{2E} + \frac{kx^2}{2E} = \left(\frac{v}{\sqrt{2E/m}} \right)^2 + \left(\frac{x}{\sqrt{2E/k}} \right)^2$$

$$1 = \frac{mv^2}{2E} + \frac{kx^2}{2E} = (\cos\theta)^2 + (\sin\theta)^2$$

Another example of the old “scale-a-circle” trick...

$$\text{Let : (1)} \quad v = \sqrt{2E/m} \cos\theta, \quad \text{and : (2)} \quad x = \sqrt{2E/k} \sin\theta \quad \text{def. (3)} \quad \omega = \frac{d\theta}{dt}$$

$$\sqrt{\frac{2E}{m}} \cos\theta = v = \frac{dx}{dt} = \frac{d\theta}{dt} \frac{dx}{d\theta} = \omega \frac{dx}{d\theta} = \omega \sqrt{\frac{2E}{k}} \cos\theta$$

by (1) by def. (3) by (2)

$$\omega = \frac{d\theta}{dt} = \sqrt{\frac{k}{m}}$$

by def. (3)
divide this by (1)

$$\theta = \int \omega \cdot dt = \omega \cdot t + \alpha$$

by integration given constant ω :

Isotropic harmonic oscillator dynamics in 1D, 2D, and 3D

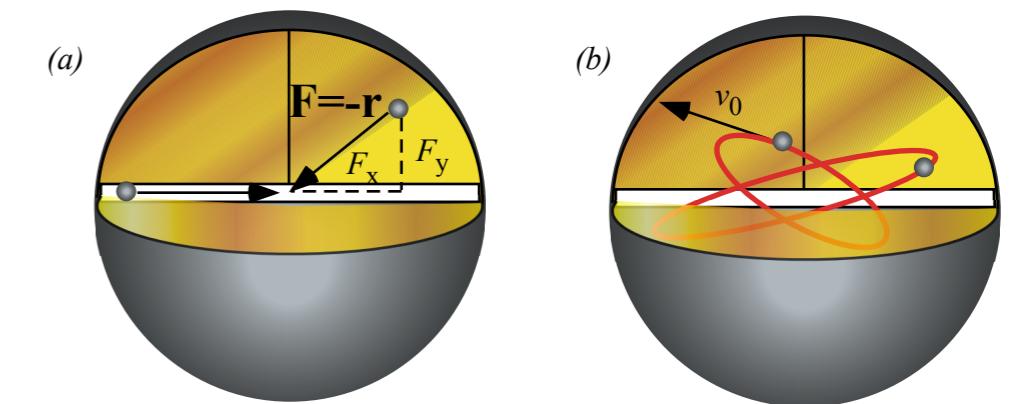
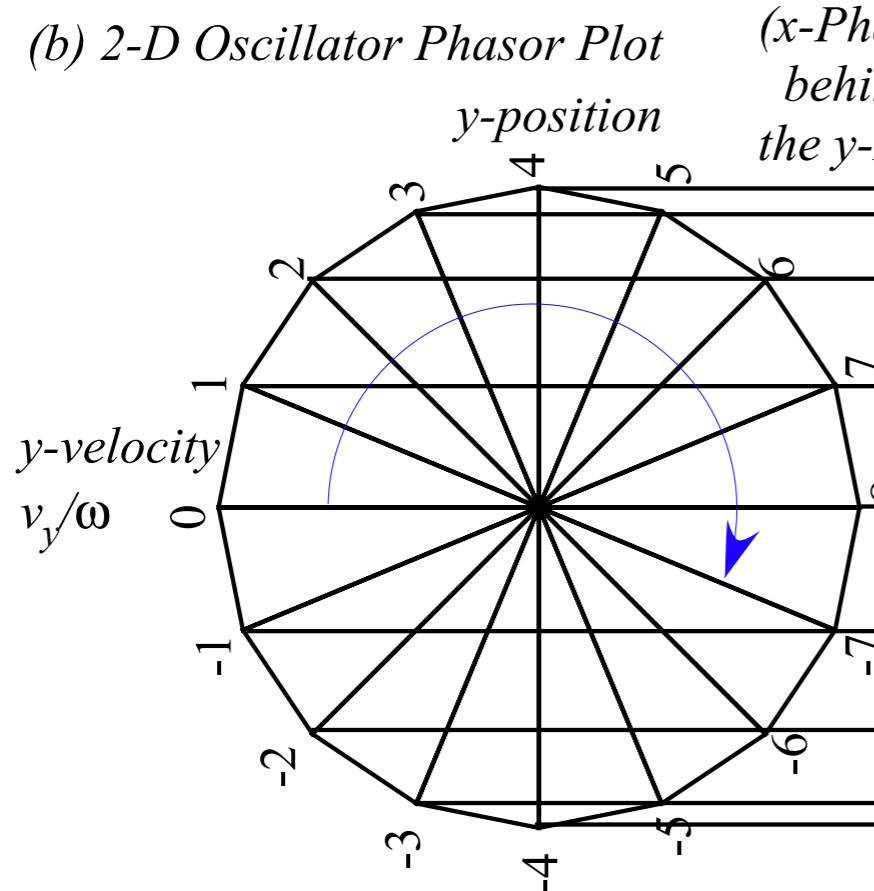
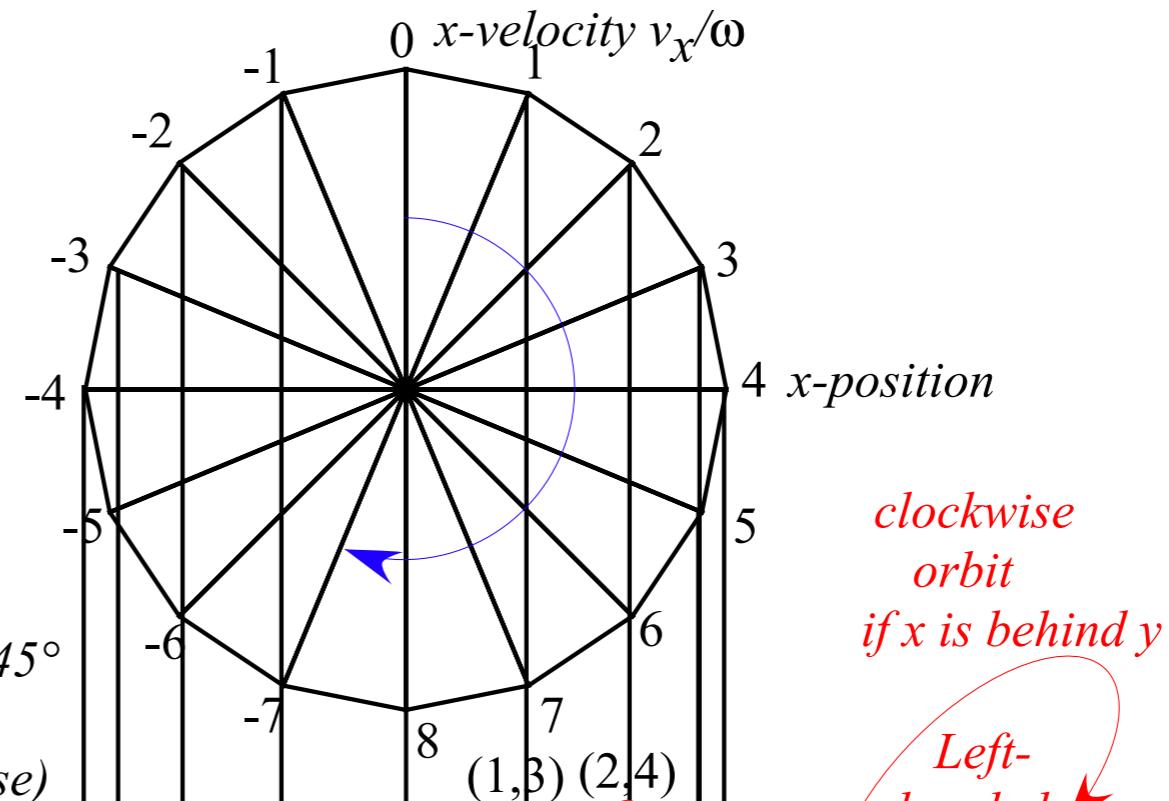
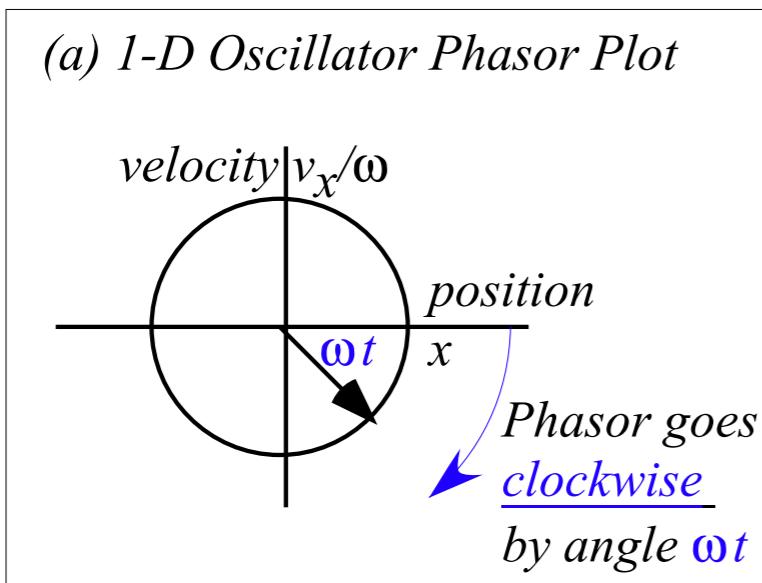
Sinusoidal space-time dynamics derived by geometry

Isotropic harmonic oscillator orbits in 1D and 2D (You get 3D for free!)

→ *Constructing 2D Isotropic harmonic oscillator orbits using phasor plots*

Examples with x-y phase lag : $\alpha_{x-y} = \alpha_x - \alpha_y = 15^\circ, 30^\circ, \text{ and } \pm 75^\circ$

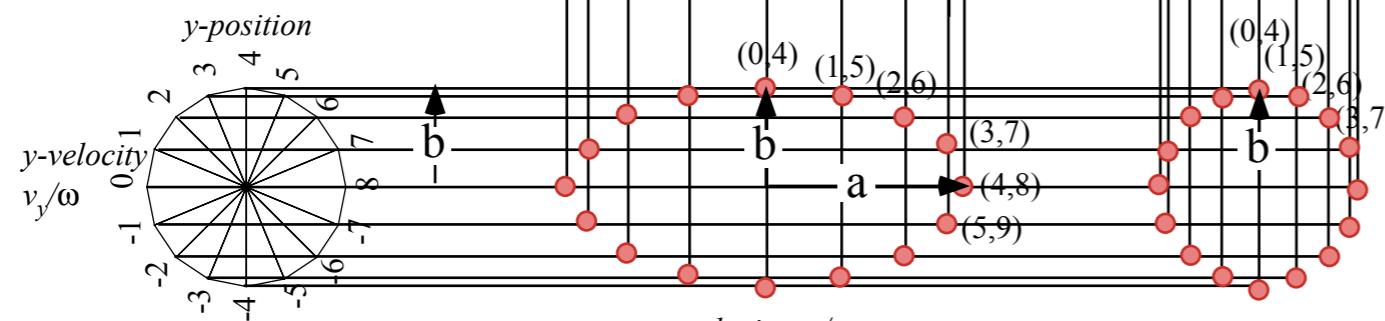
Isotropic Harmonic Oscillator phase dynamics in uniform-body



Unit 1
Fig. 9.10

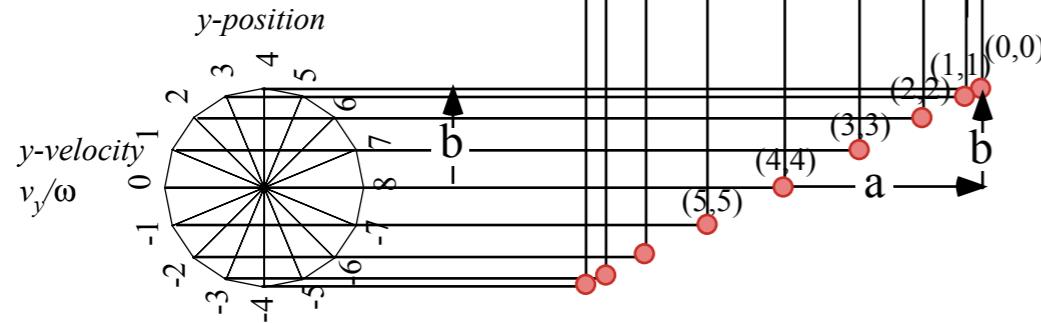
Unit 1
Fig. 9.12

(a) Phasor Plots
for
2-D Oscillator
or
Two 1D Oscillators
(x-Phase 90° behind
the y-Phase)



(b)
x-Phase 0° behind
the y-Phase

(In-phase case)



*These are more generic examples
with radius of x-phasor differing
from that of the y-phasor.*

Isotropic harmonic oscillator dynamics in 1D, 2D, and 3D

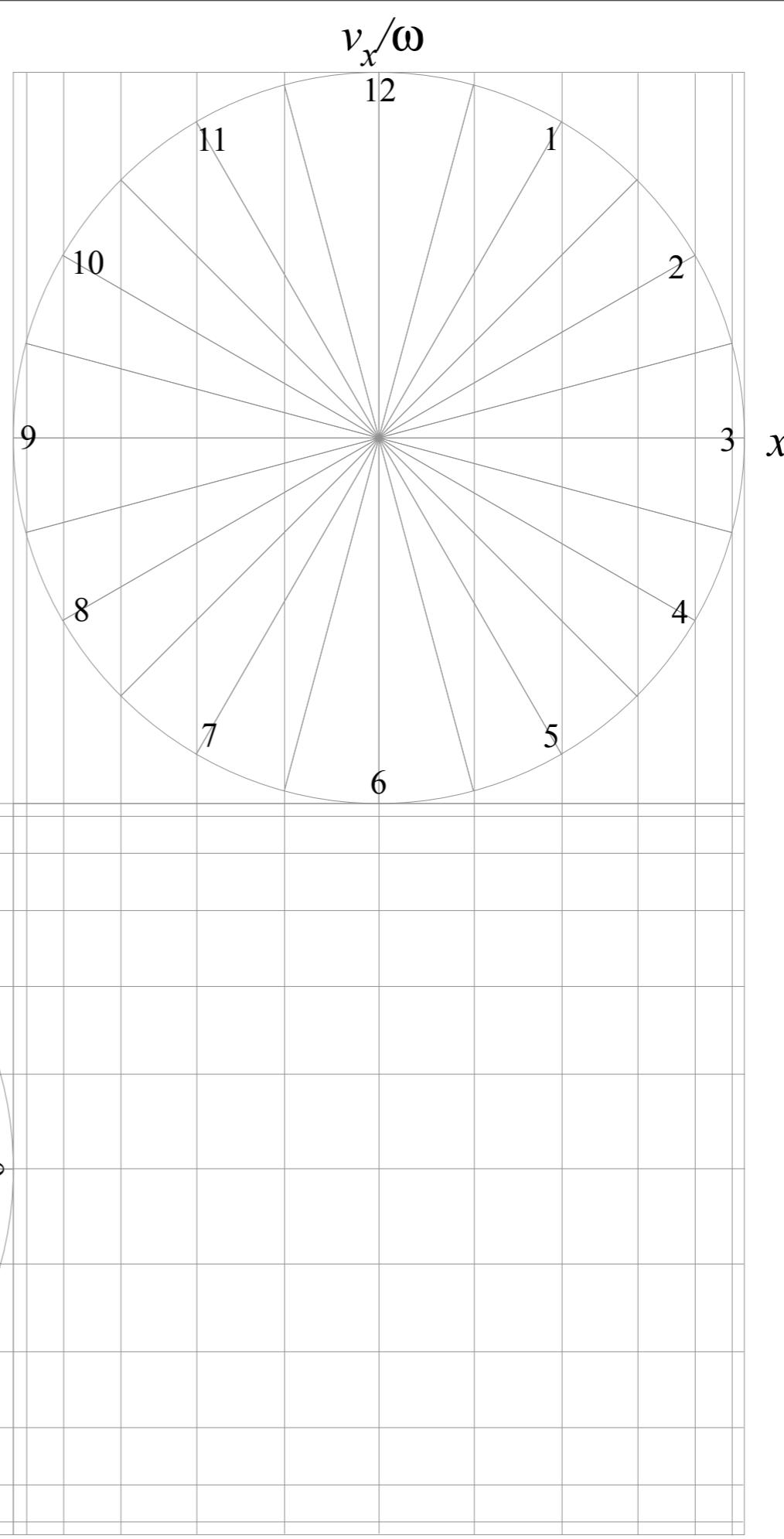
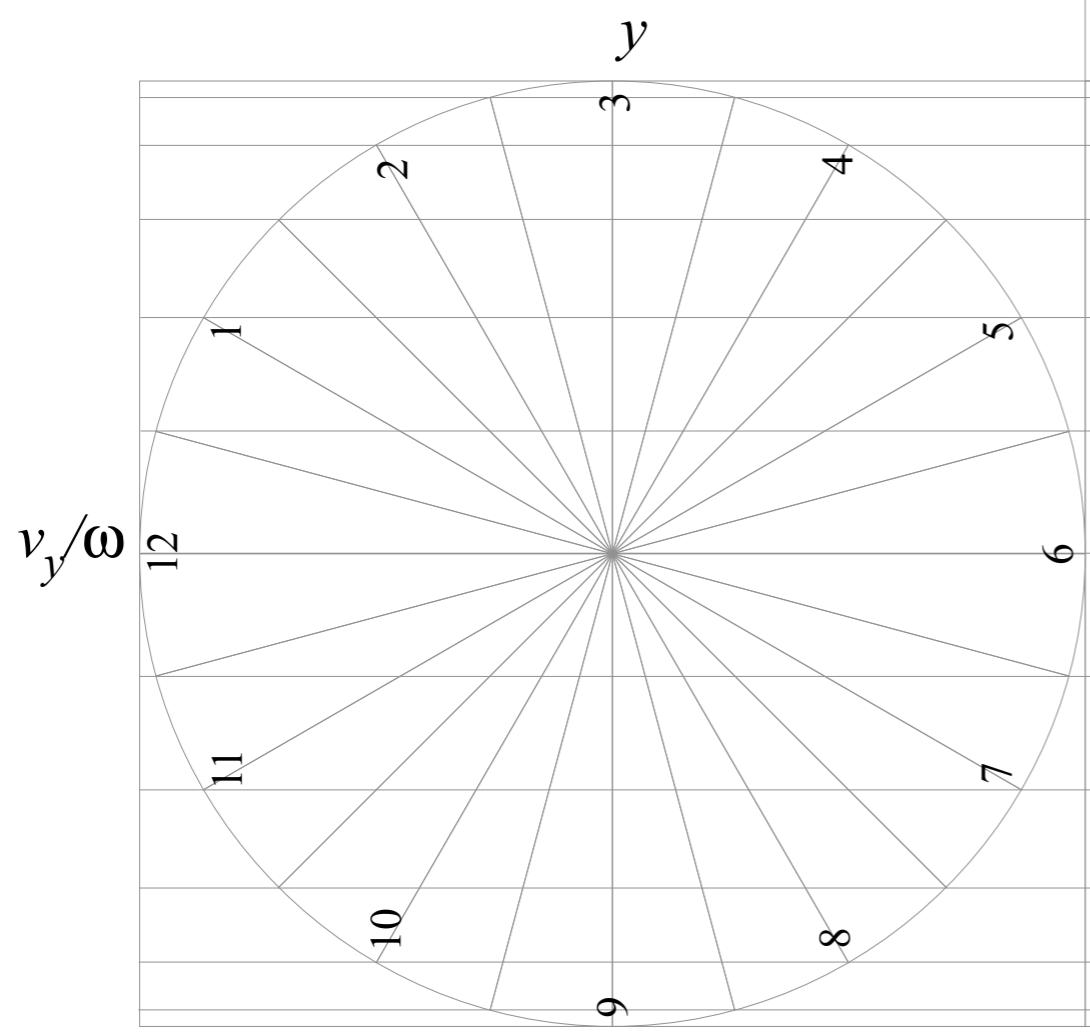
Sinusoidal space-time dynamics derived by geometry

Isotropic harmonic oscillator orbits in 1D and 2D (You get 3D for free!)

Constructing 2D Isotropic harmonic oscillator orbits using phasor plots



Examples with x-y phase lag : $\alpha_{x-y} = \alpha_x - \alpha_y = 15^\circ, 30^\circ, \text{ and } \pm 75^\circ$

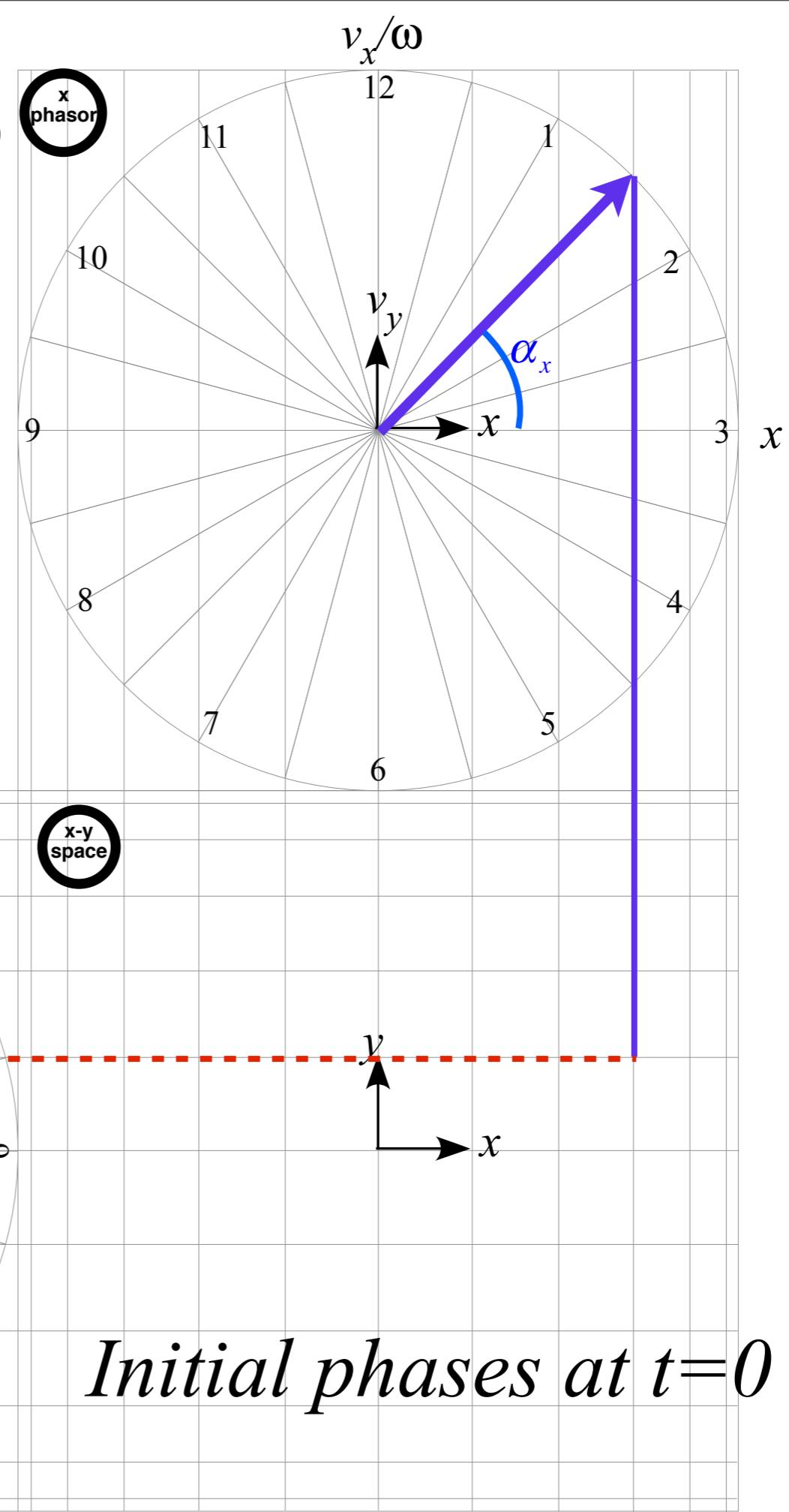


$$x(t) = A_x \cos(\omega \cdot t - \alpha_x) = A_x \cos(\alpha_x - \omega \cdot t)$$

$$\frac{v_x(t)}{\omega} = -A_x \sin(\omega \cdot t - \alpha_x) = A_x \sin(\alpha_x - \omega \cdot t)$$

$$y(t) = A_y \cos(\omega \cdot t - \alpha_y) = A_y \cos(\alpha_y - \omega \cdot t)$$

$$\frac{v_y(t)}{\omega} = -A_y \sin(\omega \cdot t - \alpha_y) = A_y \sin(\alpha_y - \omega \cdot t)$$

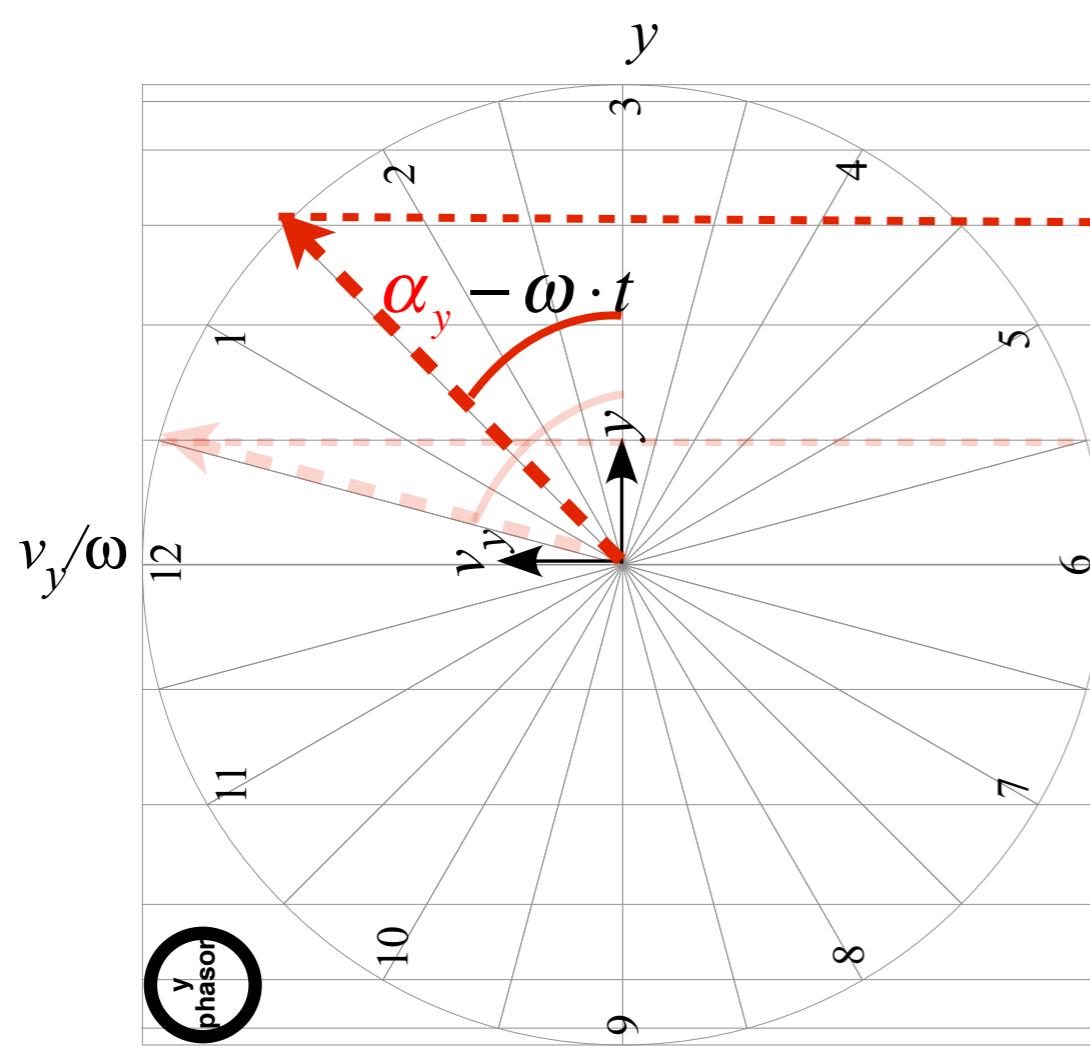
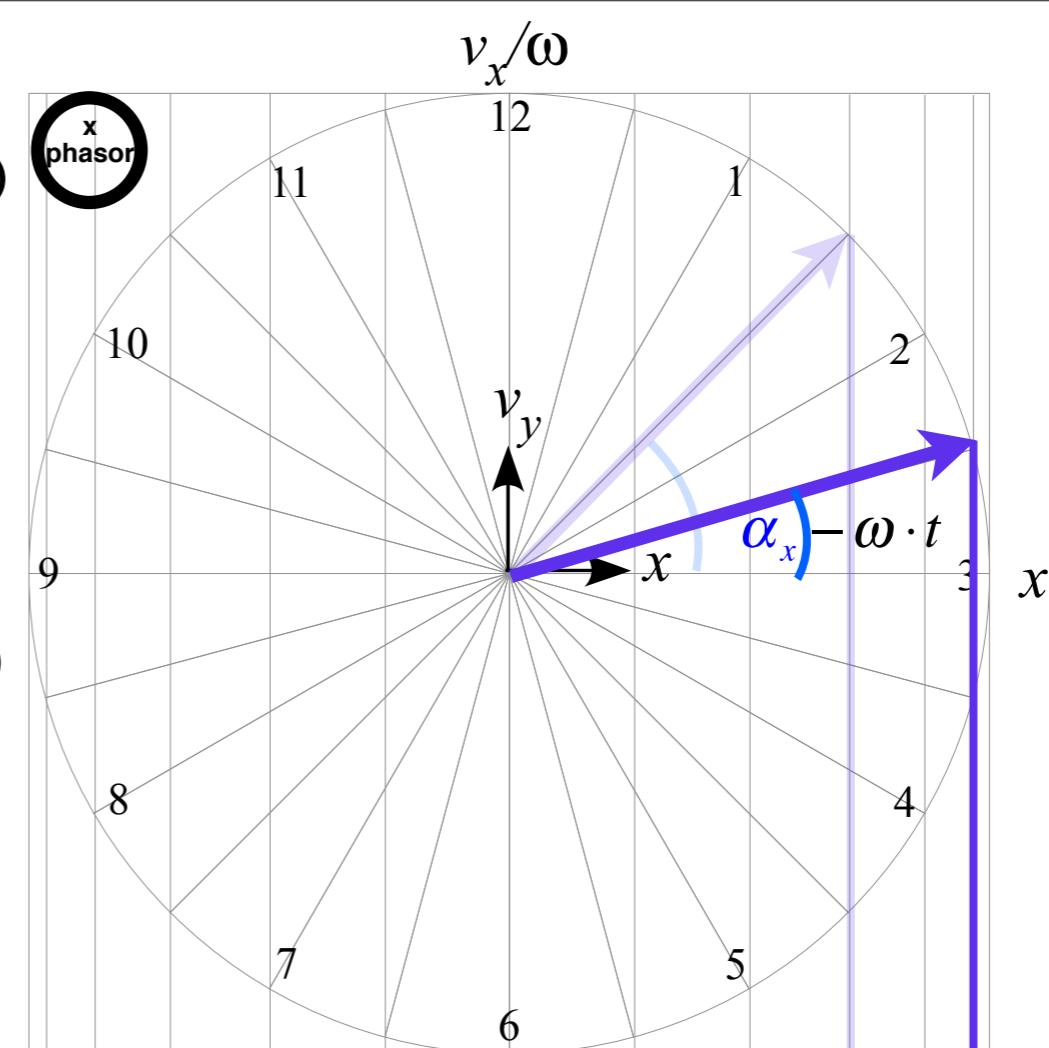


$$x(t) = A_x \cos(\omega \cdot t - \alpha_x) = A_x \cos(\alpha_x - \omega \cdot t)$$

$$\frac{v_x(t)}{\omega} = -A_x \sin(\omega \cdot t - \alpha_x) = A_x \sin(\alpha_x - \omega \cdot t)$$

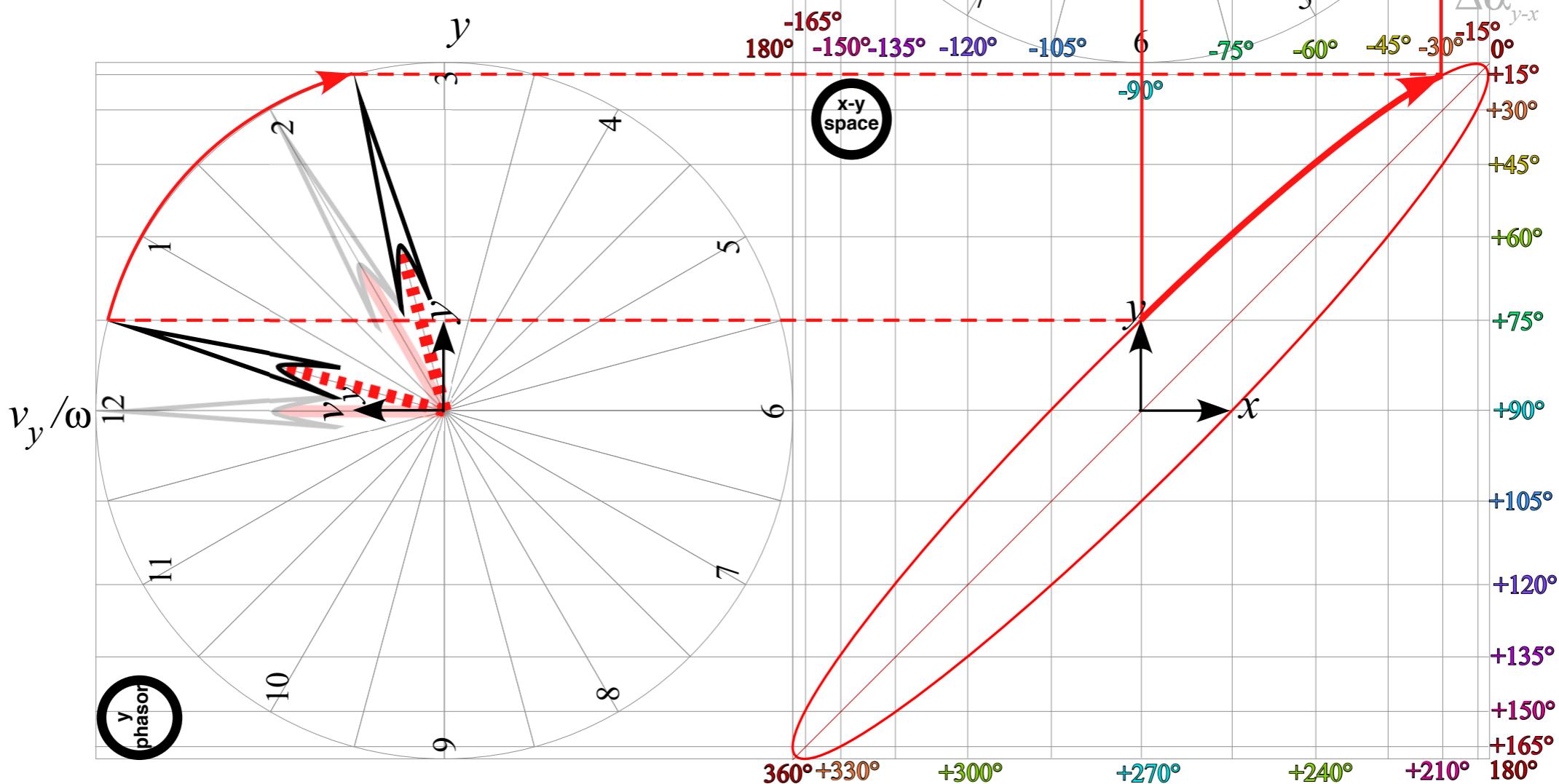
$$y(t) = A_y \cos(\omega \cdot t - \alpha_y) = A_y \cos(\alpha_y - \omega \cdot t)$$

$$\frac{v_y(t)}{\omega} = -A_y \sin(\omega \cdot t - \alpha_y) = A_y \sin(\alpha_y - \omega \cdot t)$$



Later phases at time t

Ellipsometry Contact Plots
vs.
Relative phase $\Delta\alpha = \alpha_y - \alpha_x$
 $\Delta\alpha = +15^\circ$
(Left-polarized clockwise case)



+30° case

Ellipsometry Contact Plots

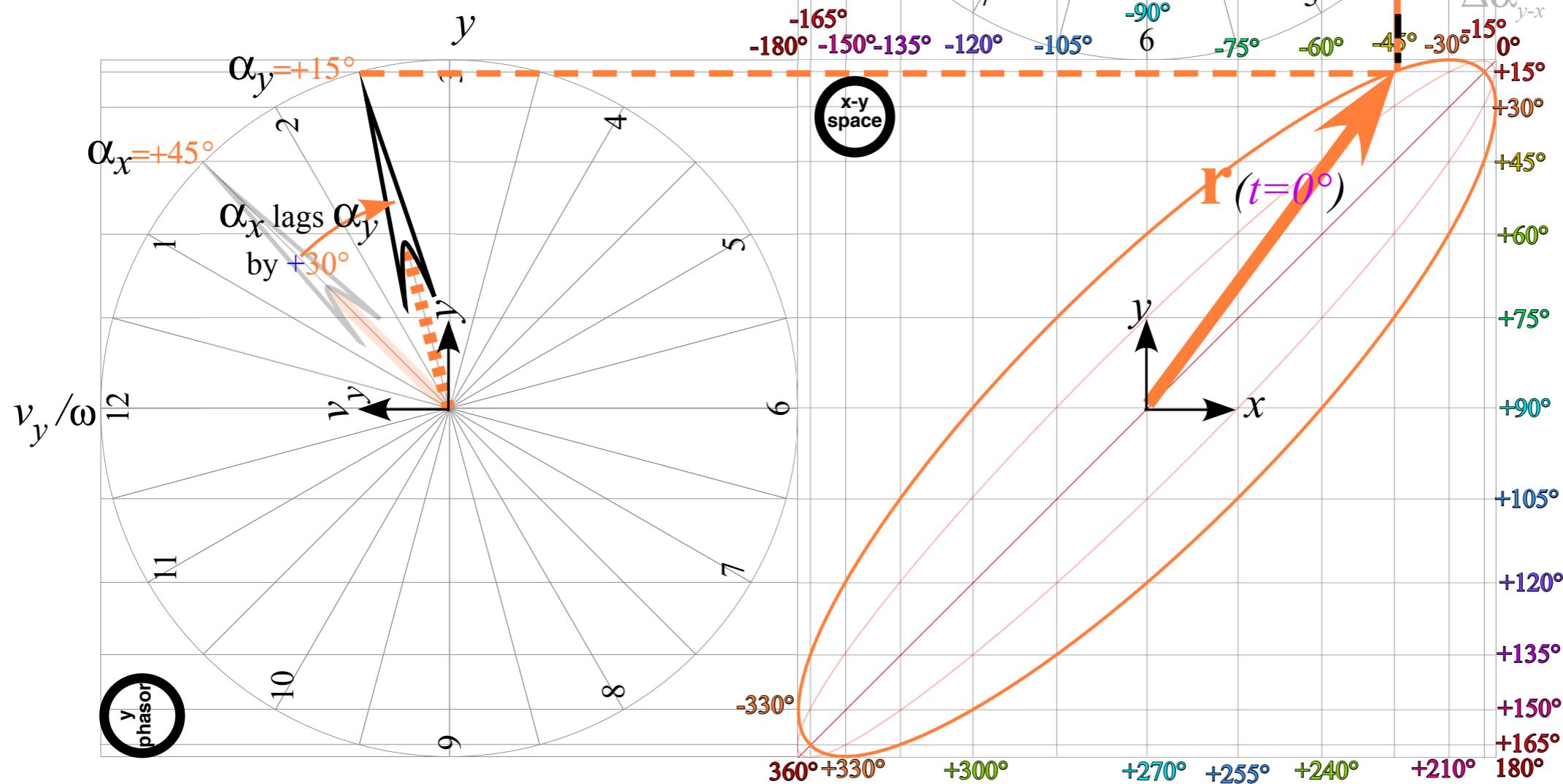
vs.

Relative phase $\Delta\alpha = \alpha_y - \alpha_x$

$$\Delta\alpha = \alpha_x - \alpha_y = +30^\circ$$

Left-polarized (clockwise) orbit

Initial phase ($\alpha_x=45^\circ, \alpha_y=+15^\circ$)



+30° case

Ellipsometry Contact Plots

vs.

Relative phase $\Delta\alpha = \alpha_y - \alpha_x$

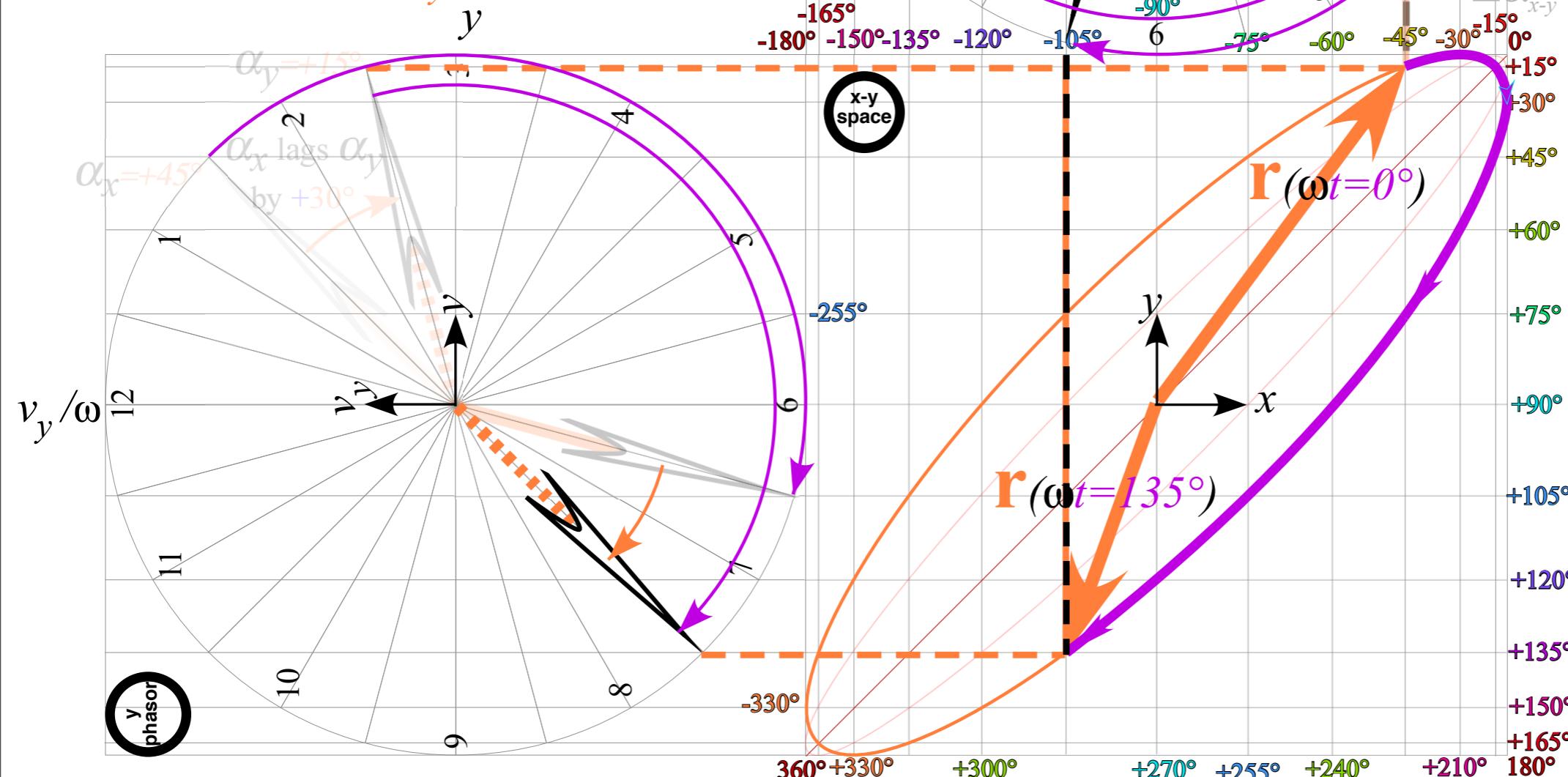
$$\Delta\alpha = \alpha_x - \alpha_y = +30^\circ$$

Left-polarized (clockwise) orbit

Initial phase ($\alpha_x = 45^\circ, \alpha_y = +15^\circ$)

...after time advances by 150°:

$$(\alpha_x = 45^\circ - 150^\circ = -105^\circ, \alpha_y = +15^\circ - 150^\circ = -135^\circ)$$



-75° case

Ellipsometry Contact Plots

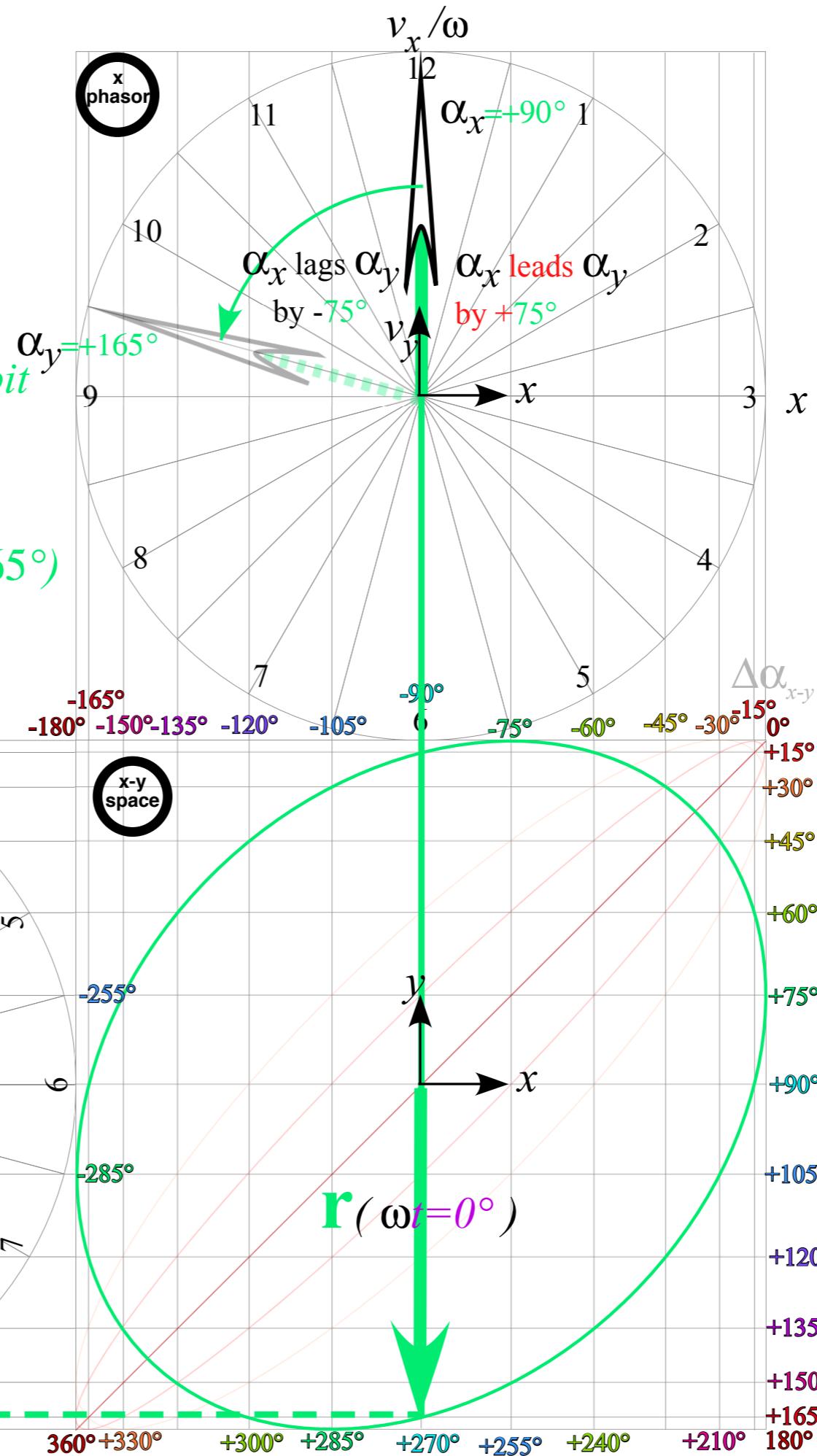
vs.

Relative phase $\Delta\alpha = \alpha_y - \alpha_x$

$$\Delta\alpha = \alpha_x - \alpha_y = -75^\circ$$

Right-polarized (anti-clockwise) orbit

Initial phase ($\alpha_x = 90^\circ, \alpha_y = +165^\circ$)



-75° case

Ellipsometry Contact Plots

vs.

Relative phase $\Delta\alpha = \alpha_y - \alpha_x$

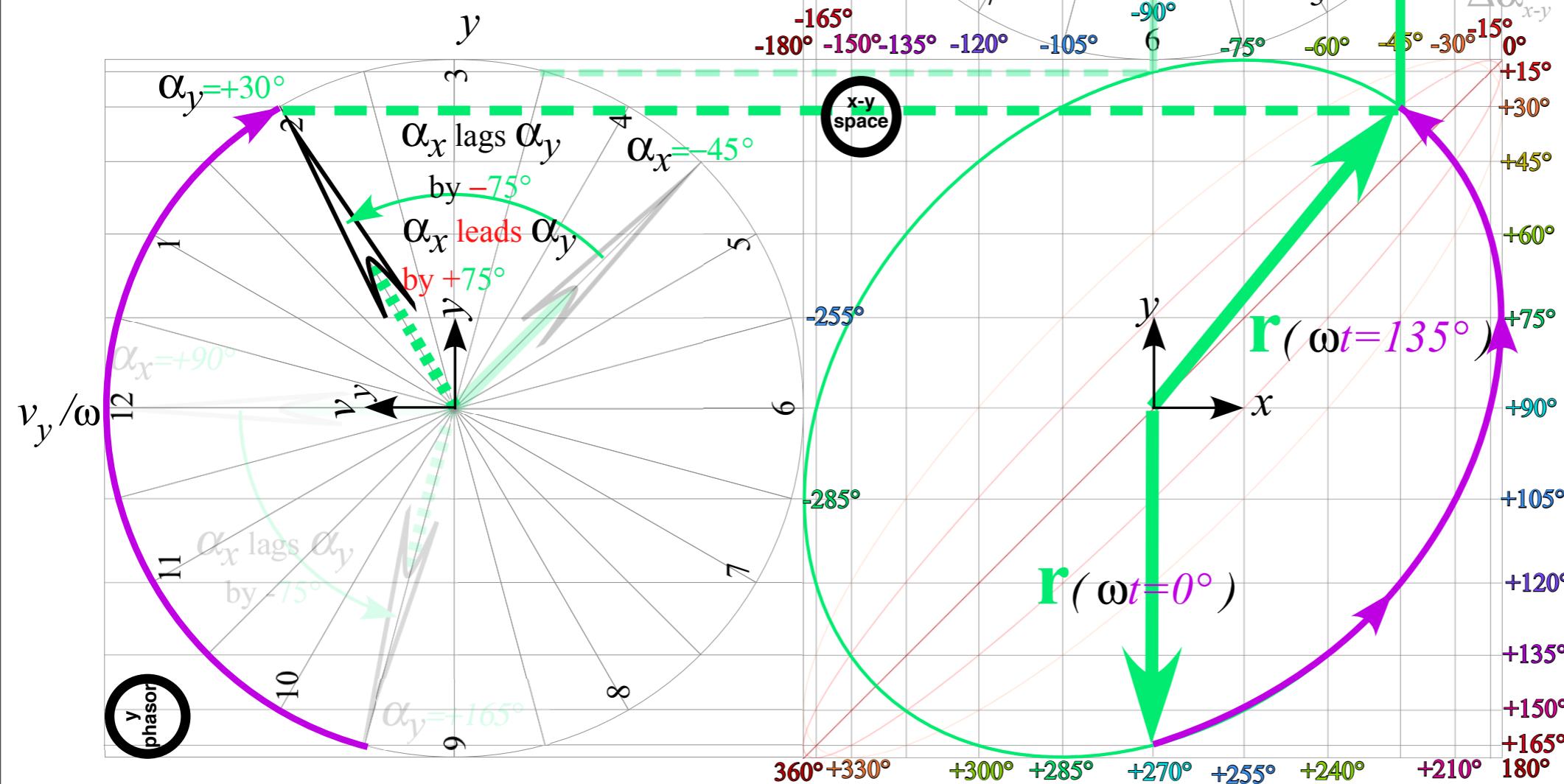
$$\Delta\alpha = \alpha_x - \alpha_y = -75^\circ$$

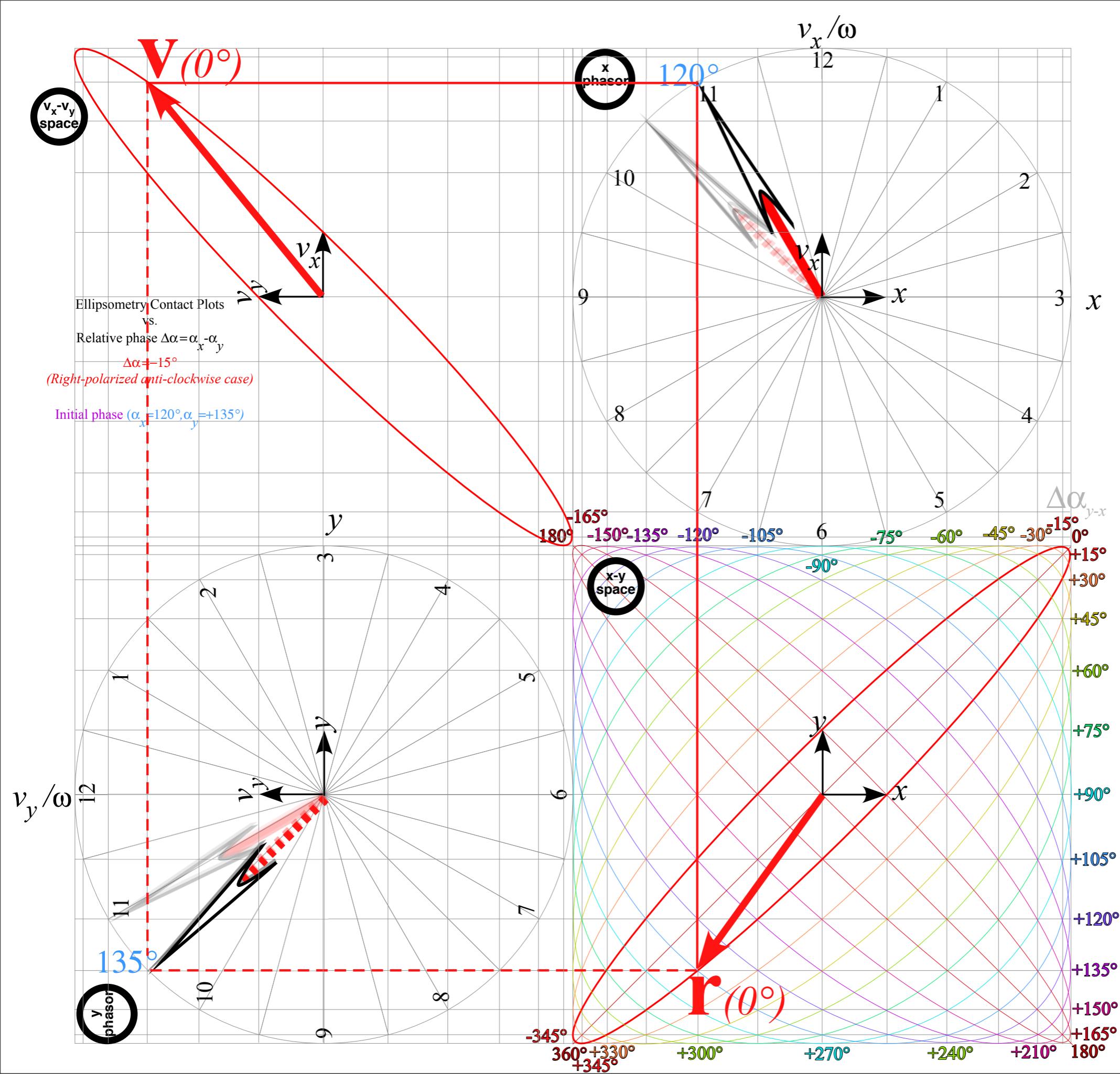
Right-polarized (anti-clockwise) orbit

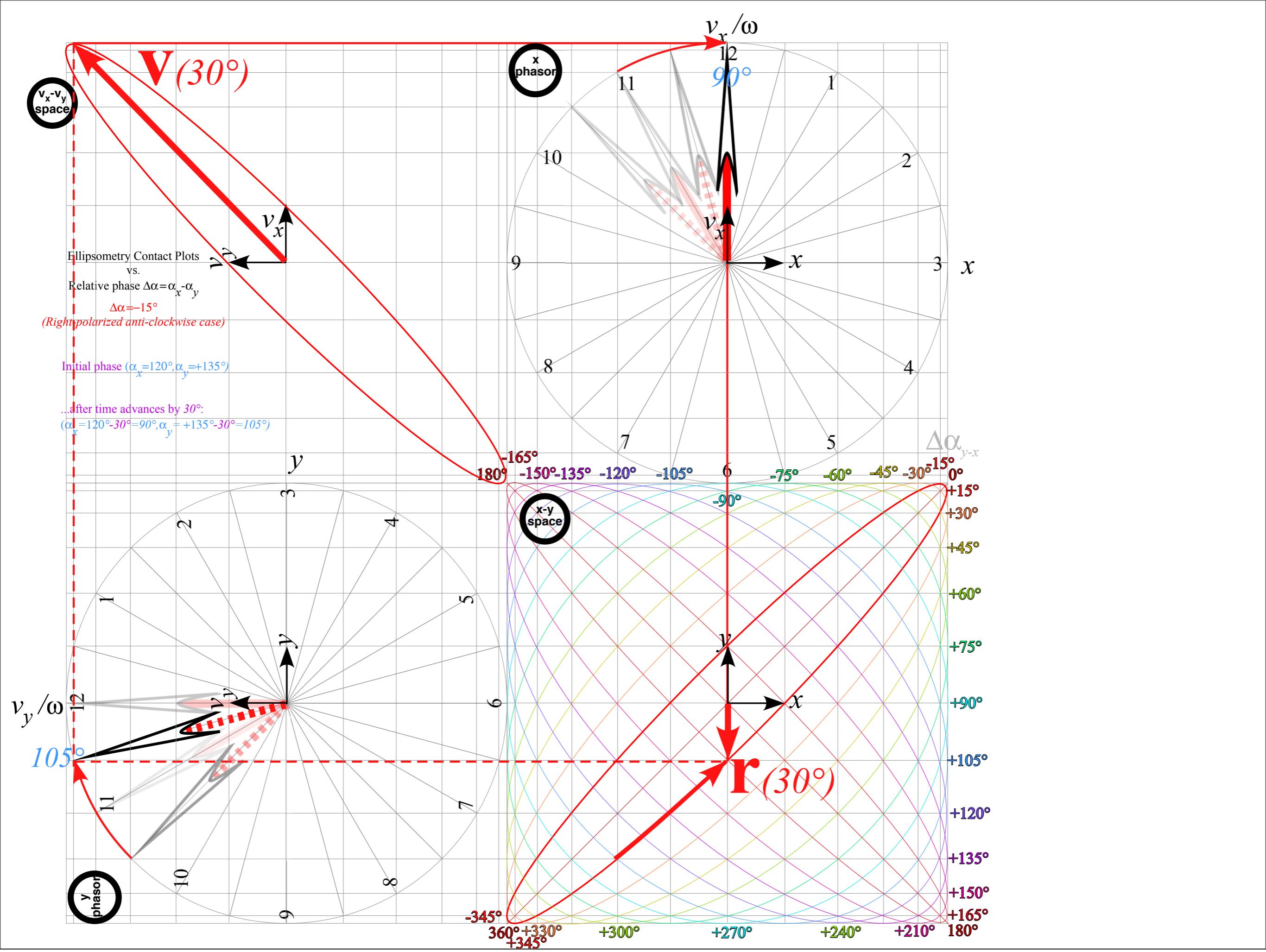
Initial phase ($\alpha_x = 90^\circ, \alpha_y = +165^\circ$)

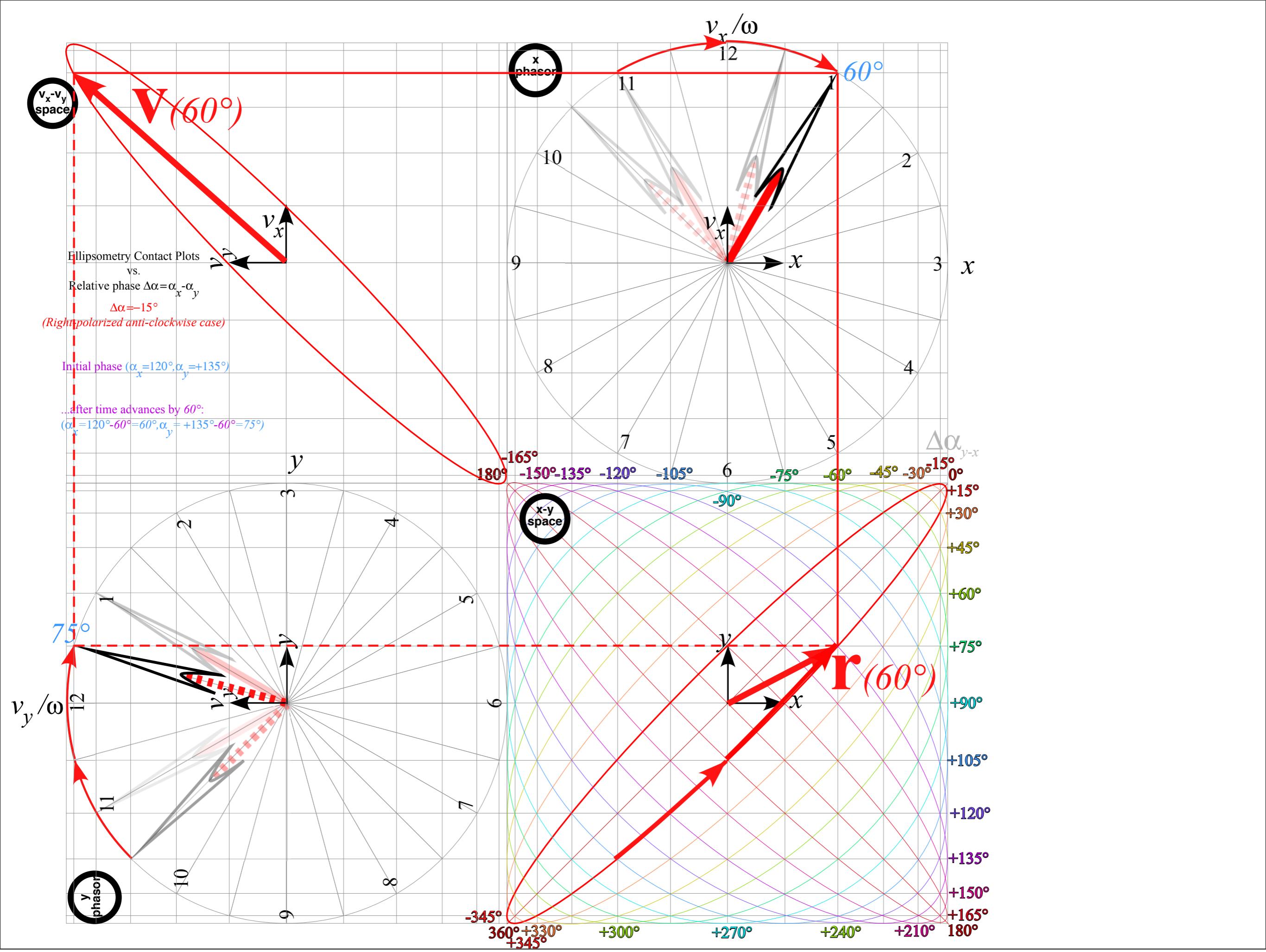
...after time advances by 135° :

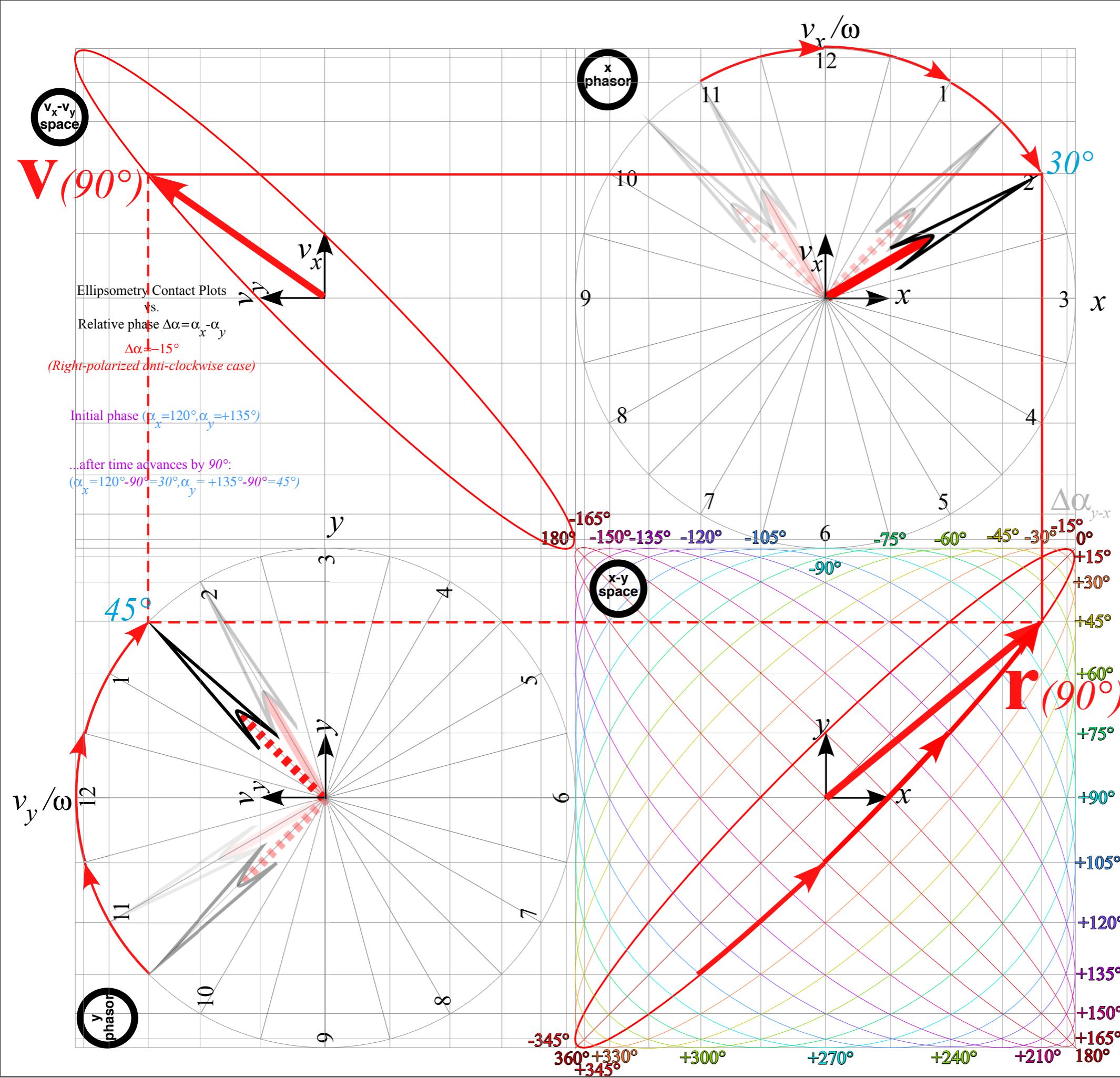
$$(\alpha_x = 90^\circ - 135^\circ = -45^\circ, \alpha_y = +165^\circ - 135^\circ = +30^\circ)$$

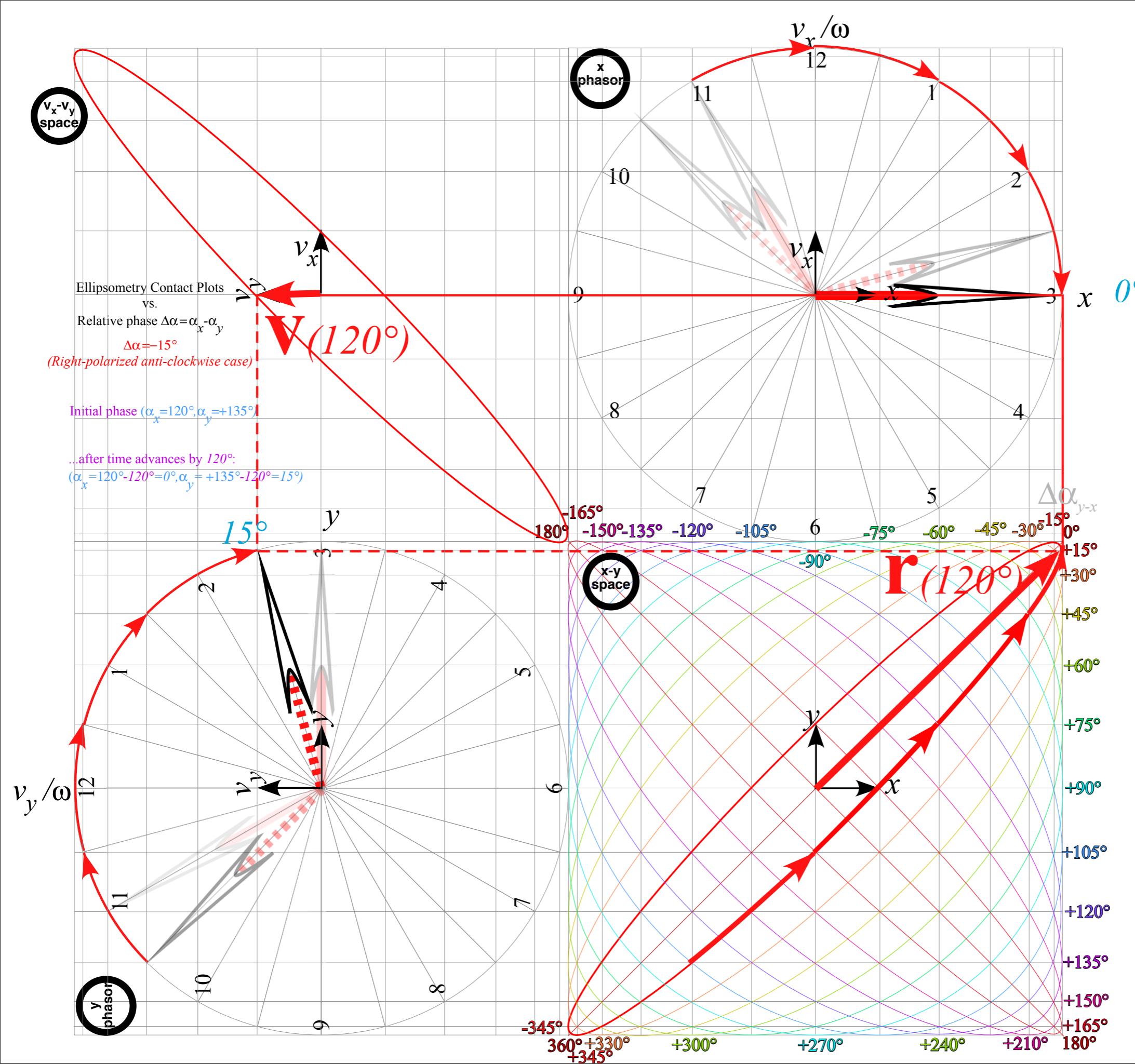


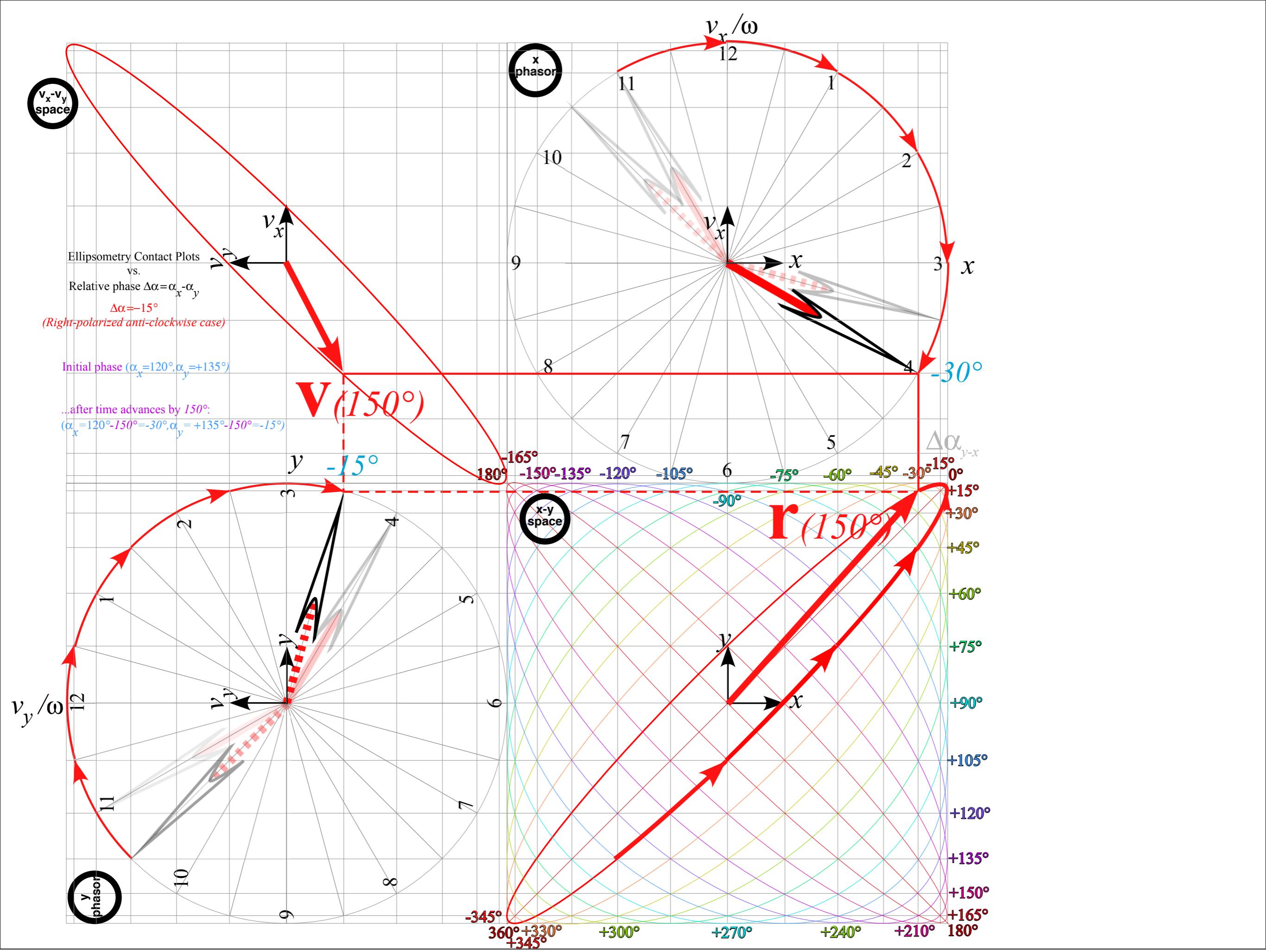


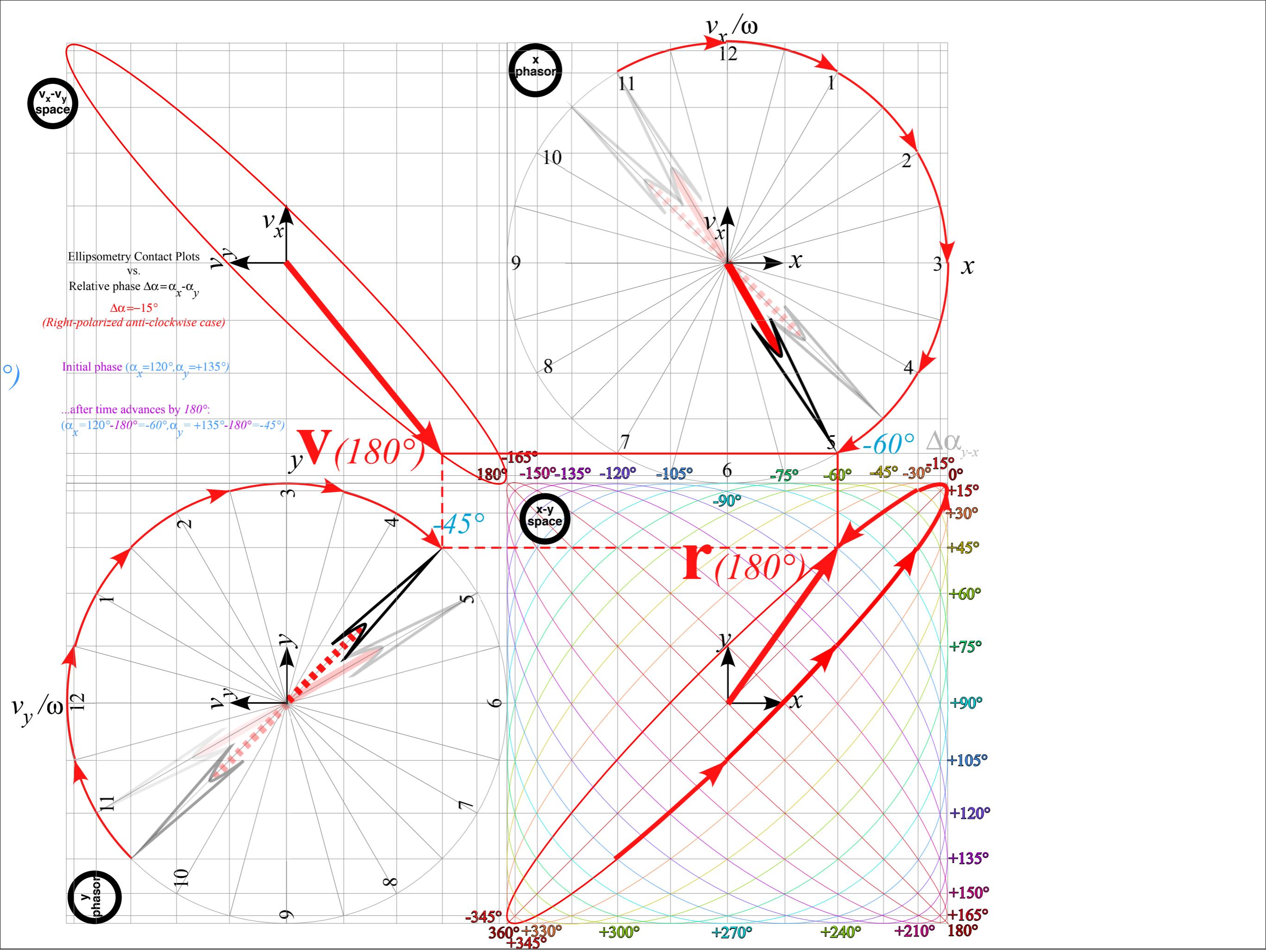


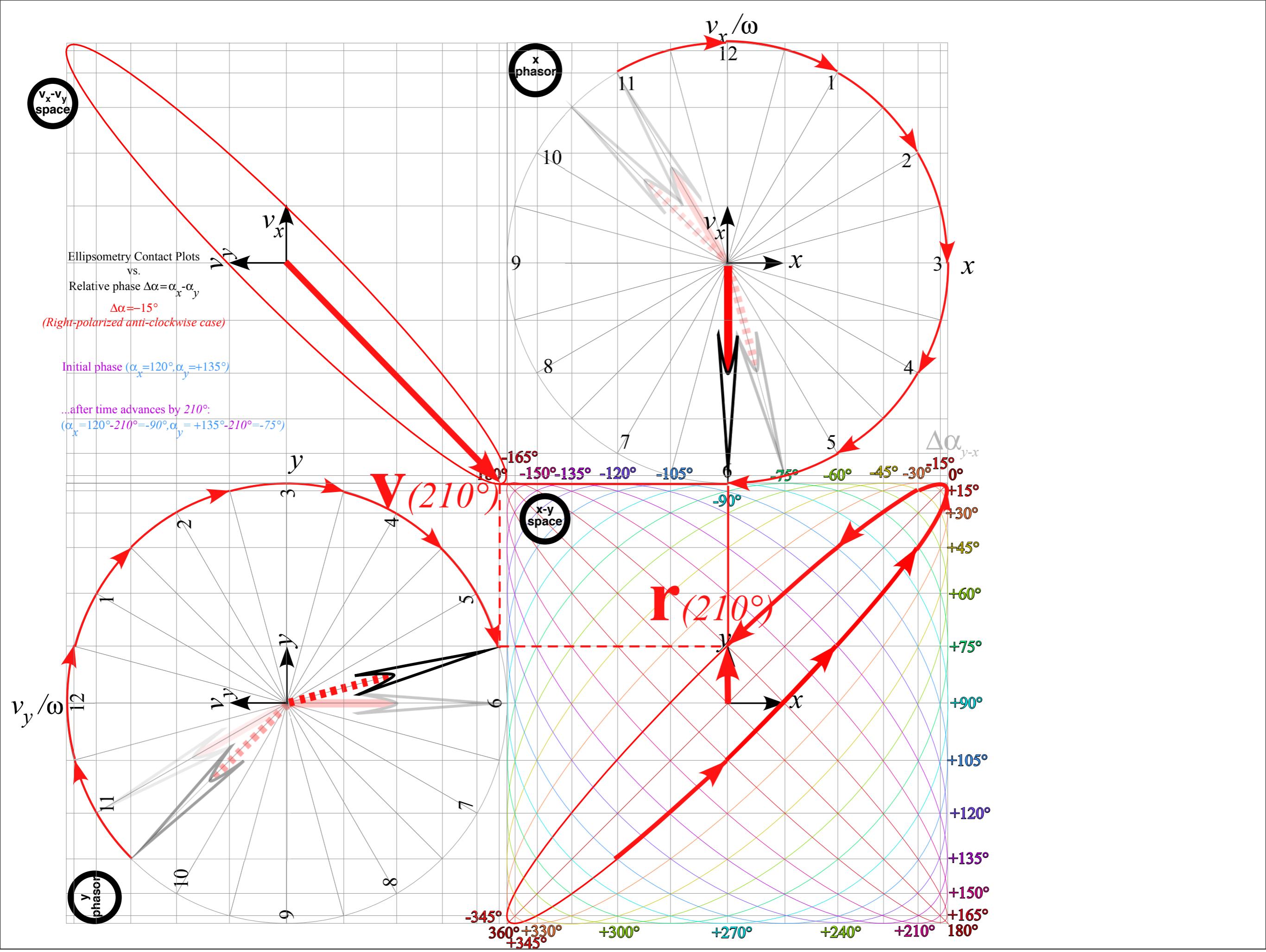


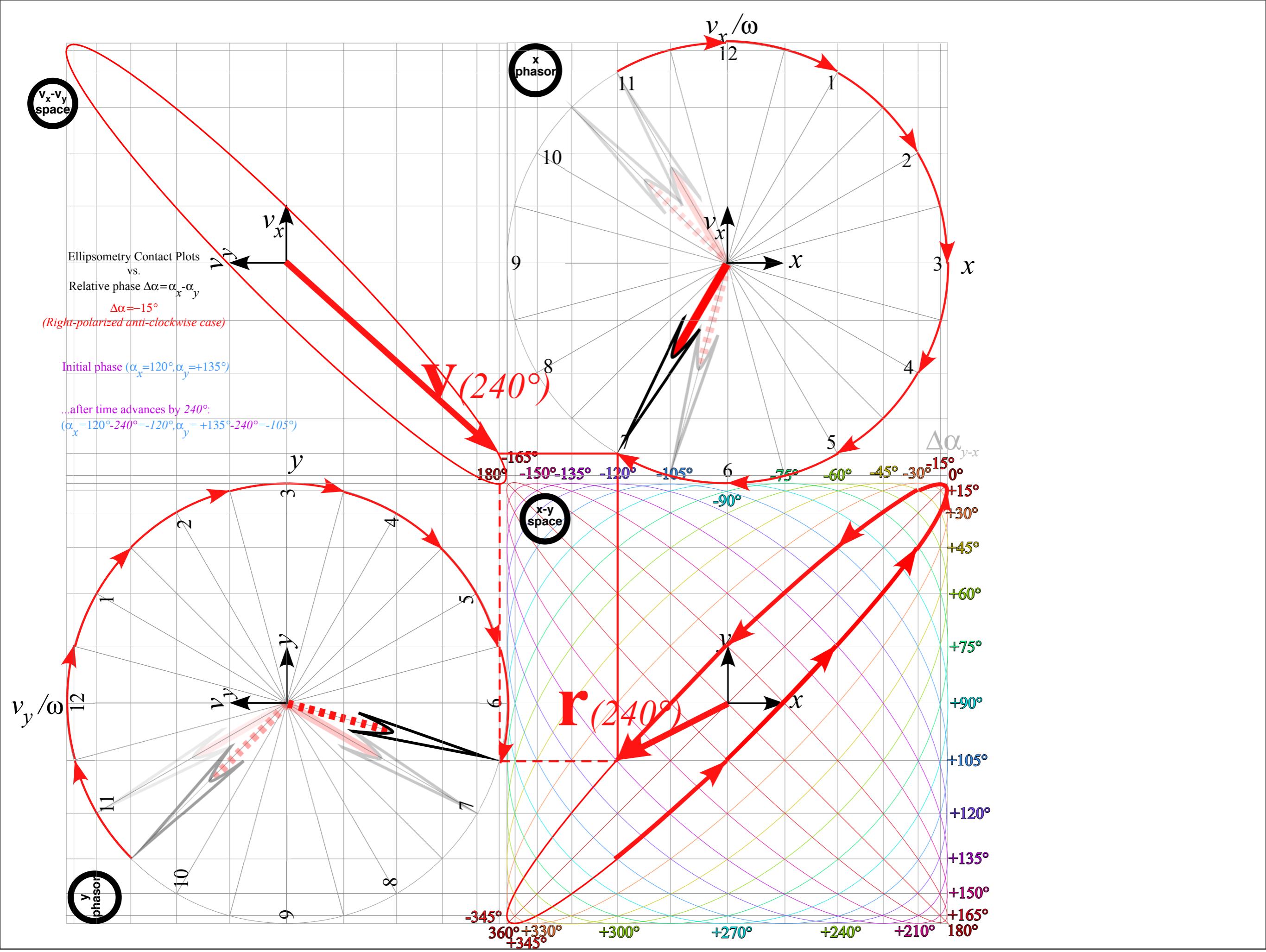


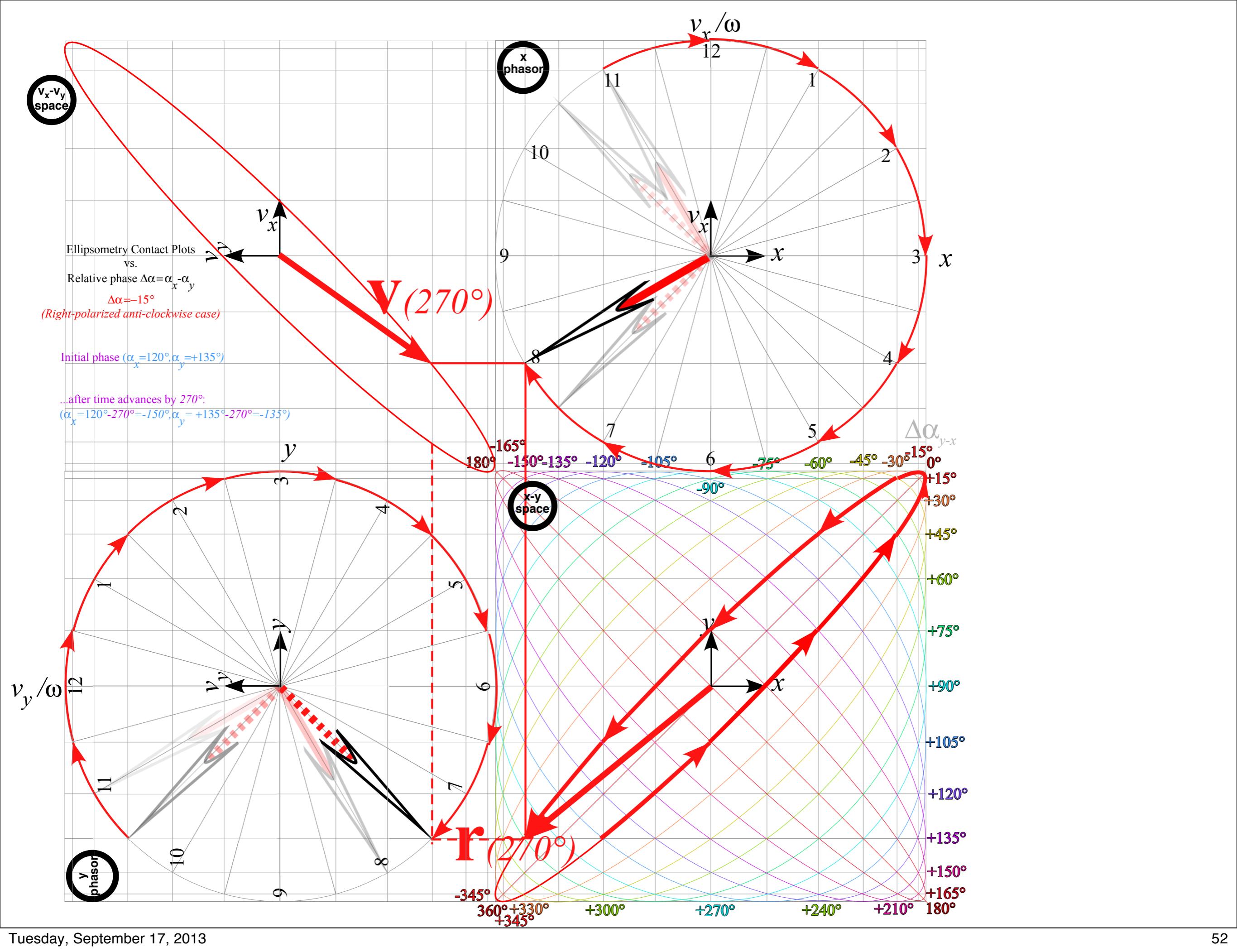


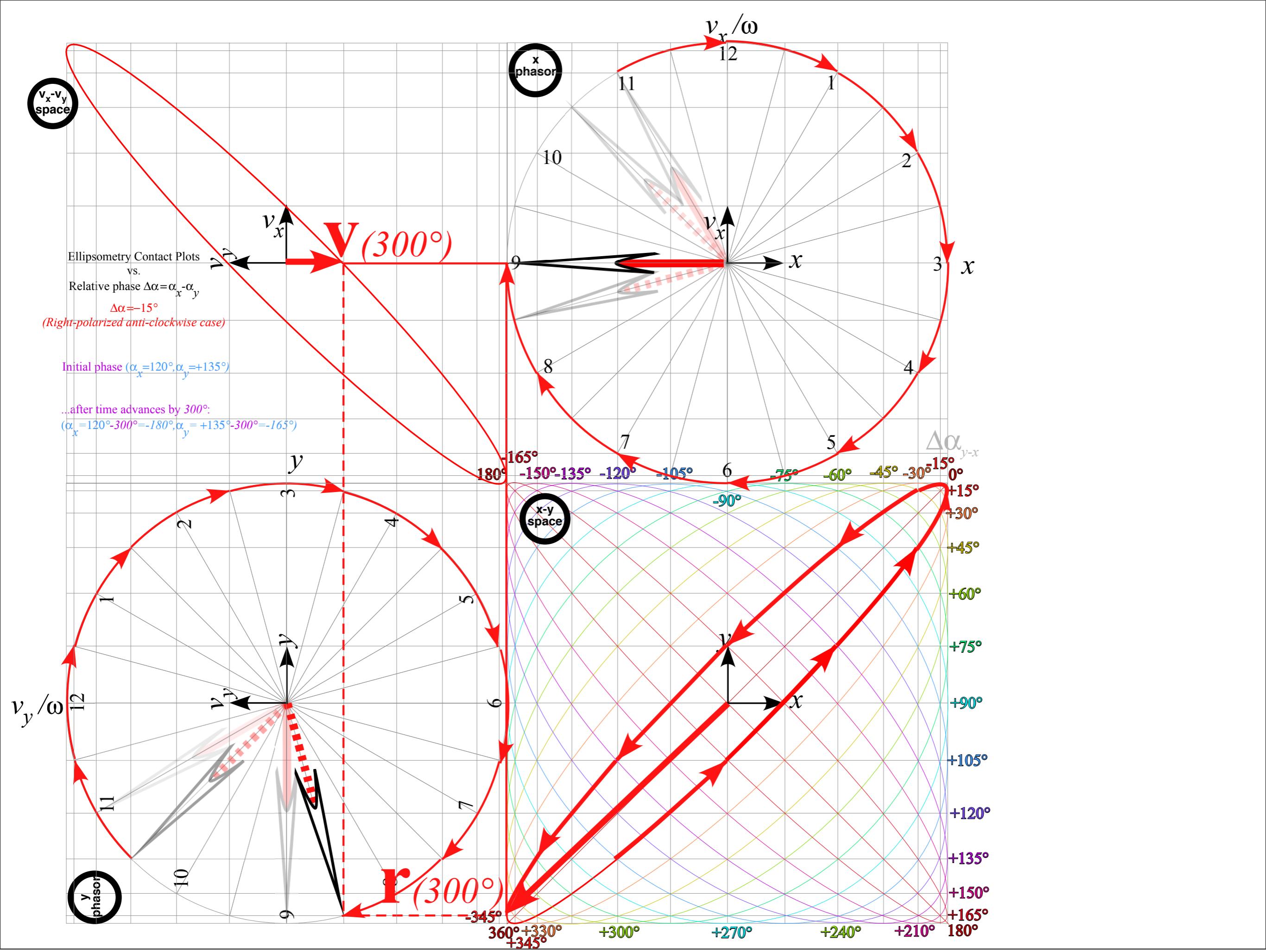


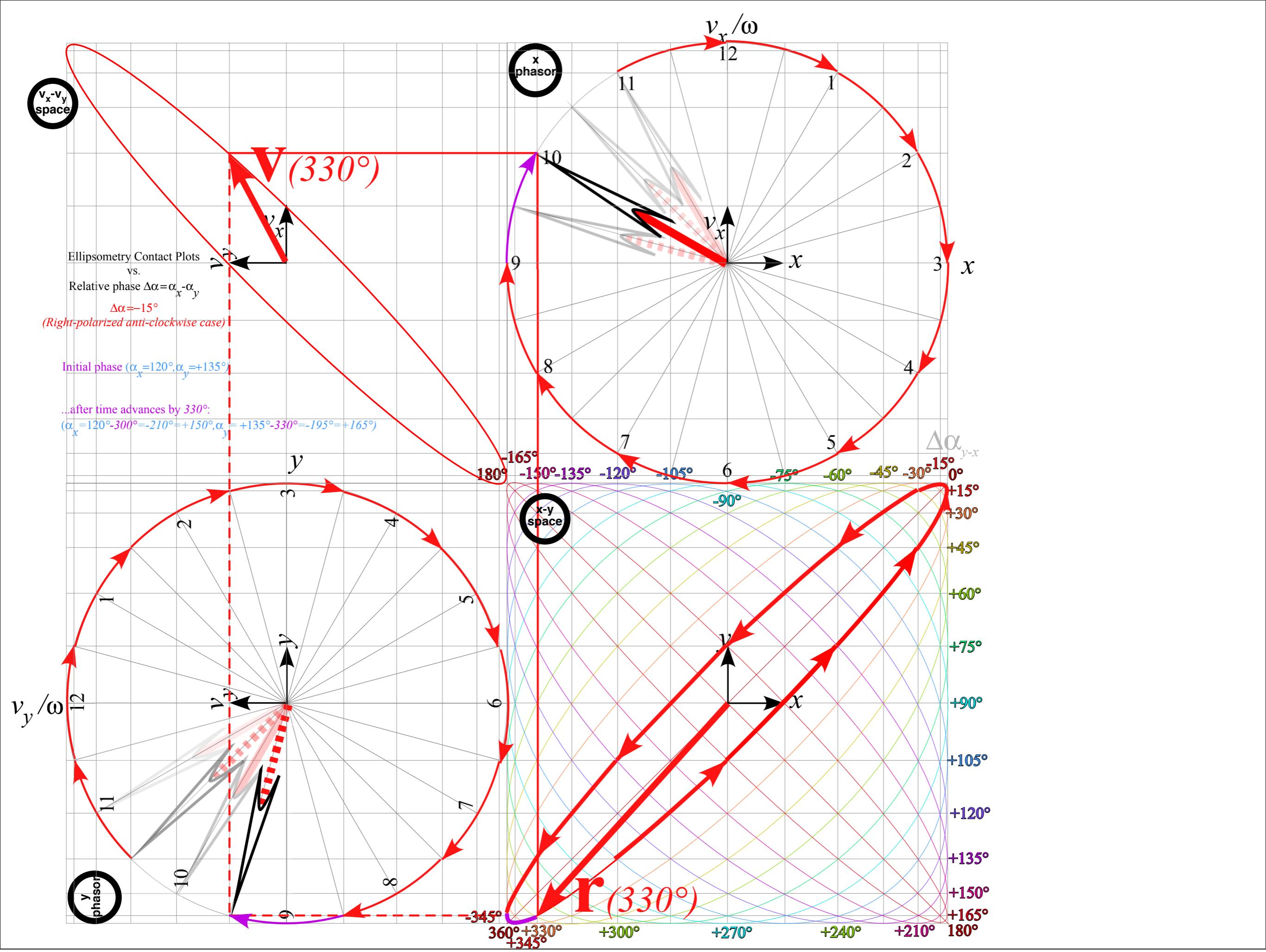


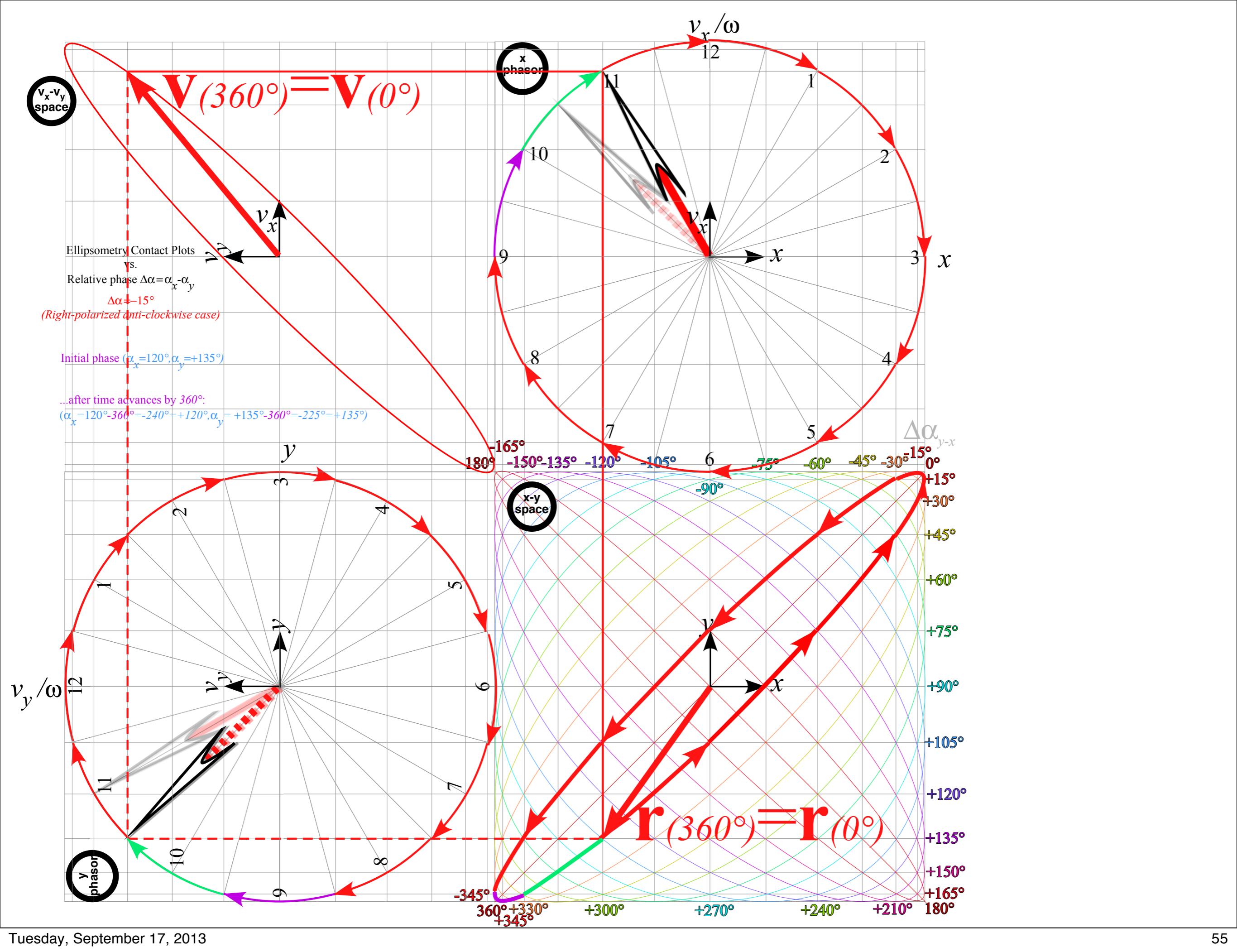


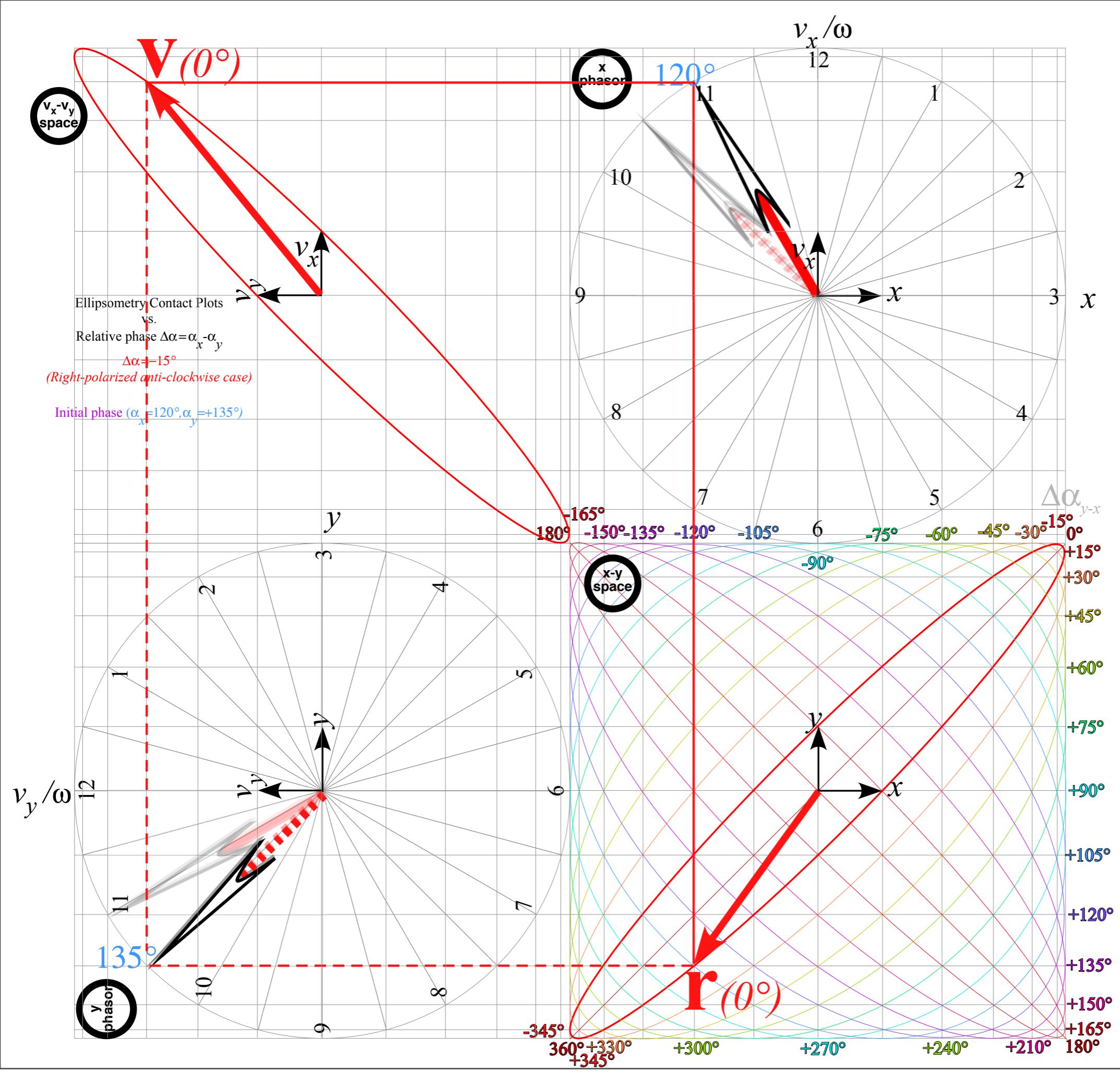


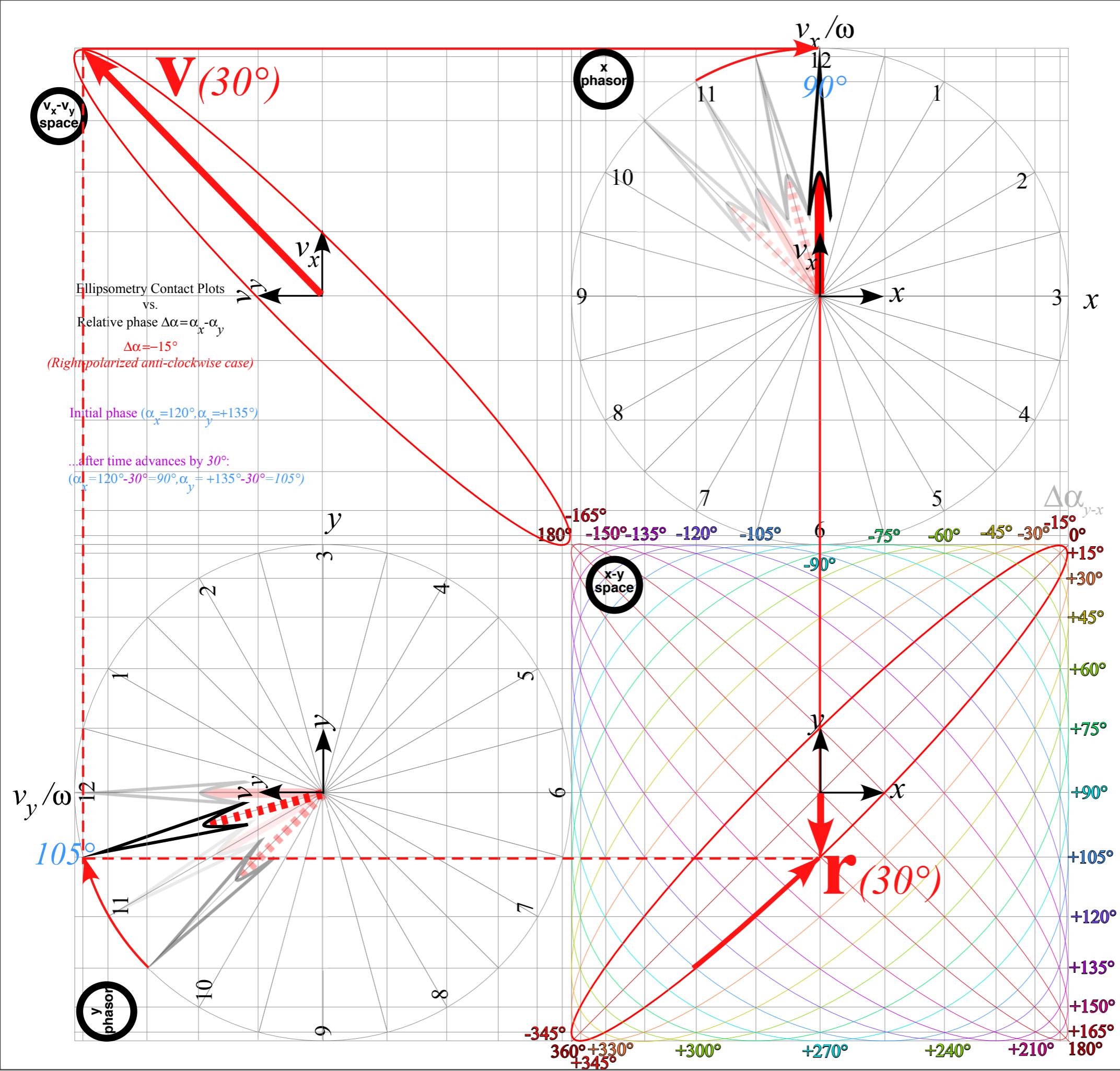


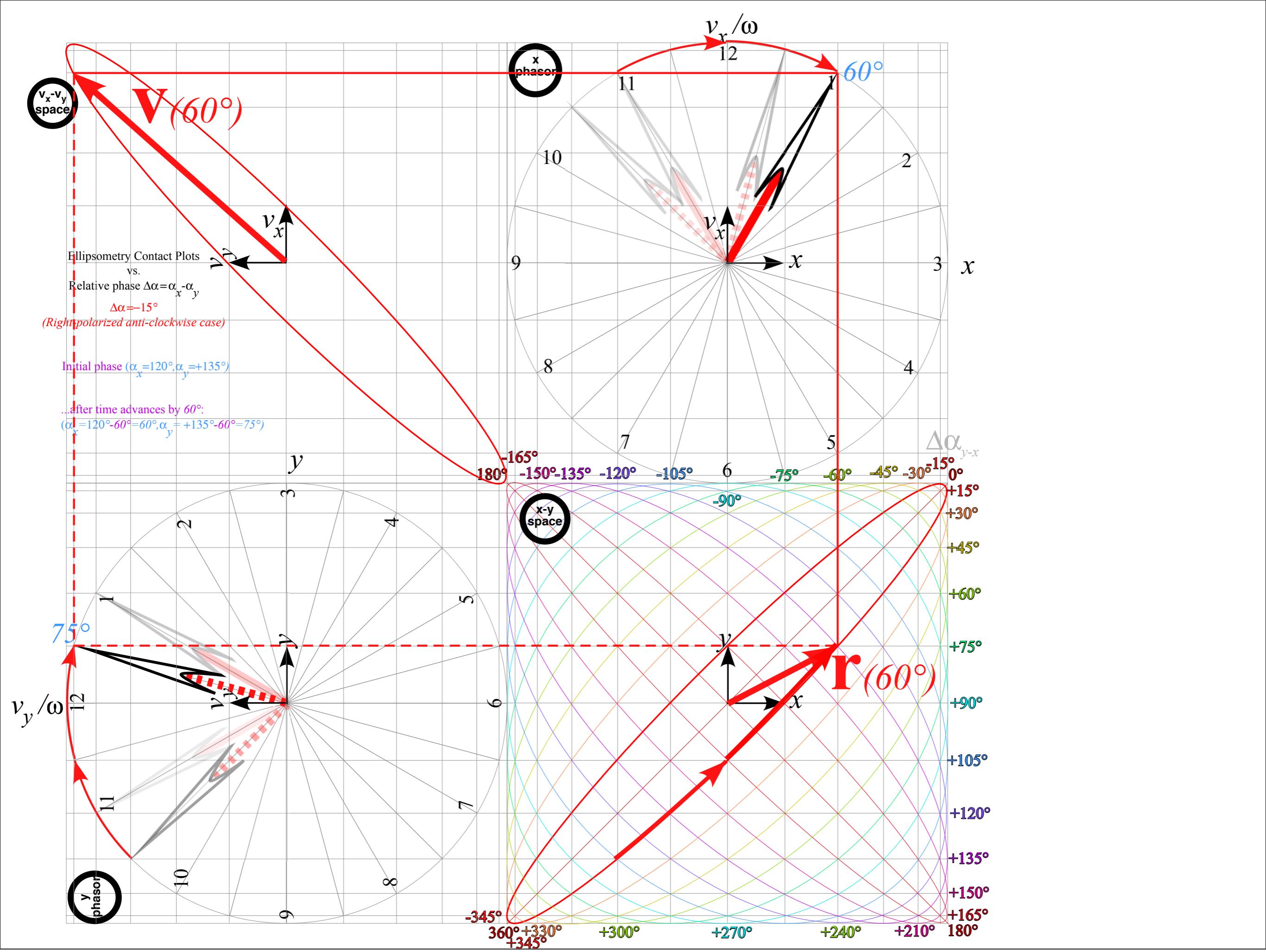


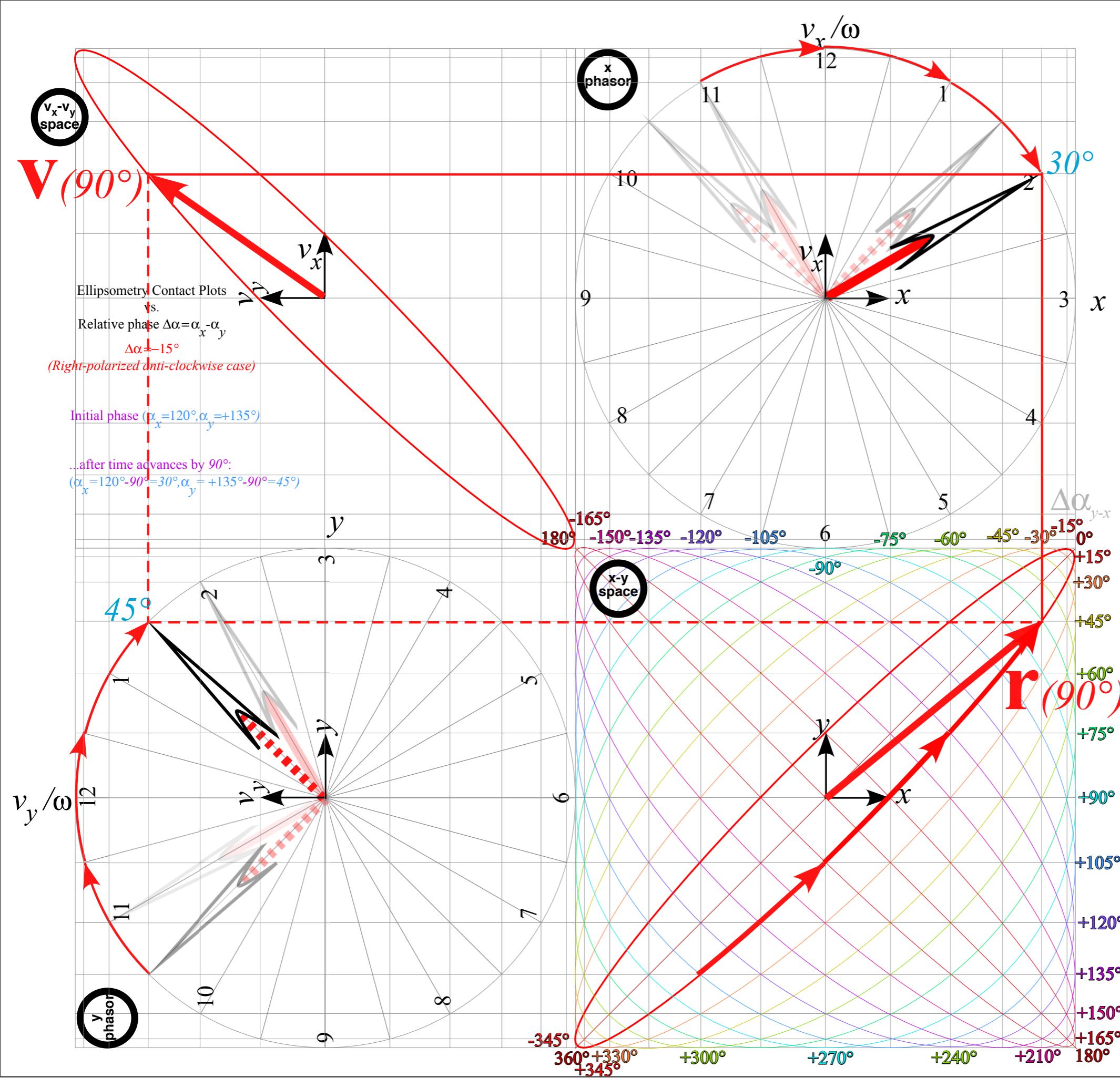








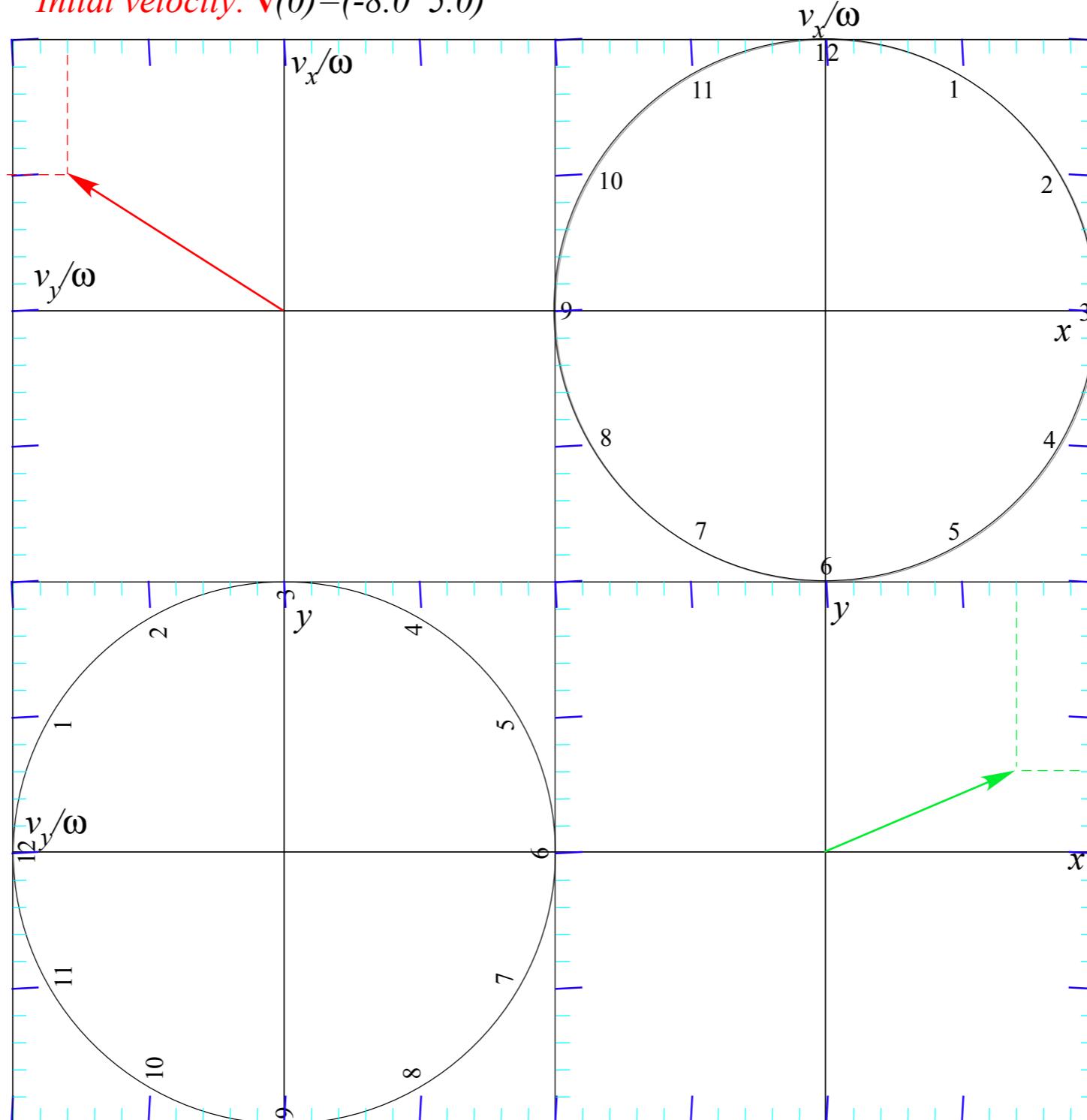




Constructing 2D IHO orbits by phasor plots

→ *Integrating IHO equations by phasor geometry (case of unequal x and y phasor area)*

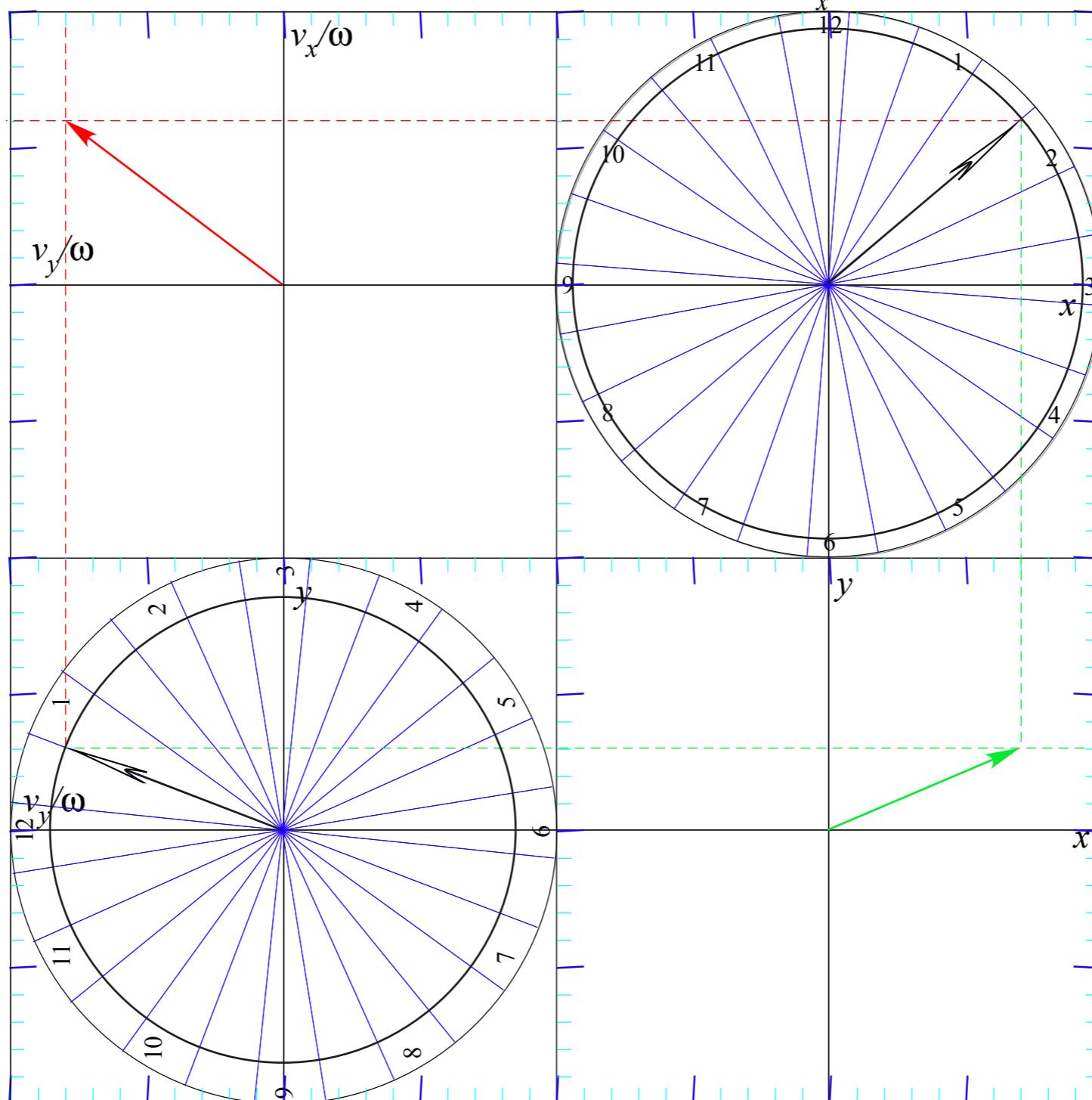
Initial velocity: $\mathbf{v}(0)=(-8.0 \ 5.0)$



Initial position: $\mathbf{r}(0)=(7.0 \ 3.0)$

[BoxIt simulation of U\(2\) orbits](#)
<http://www.uark.edu/ua/modphys/markup/BoxItWeb.html>

Initial velocity: $\mathbf{v}(0) = (-8.0 \ 6.0)$



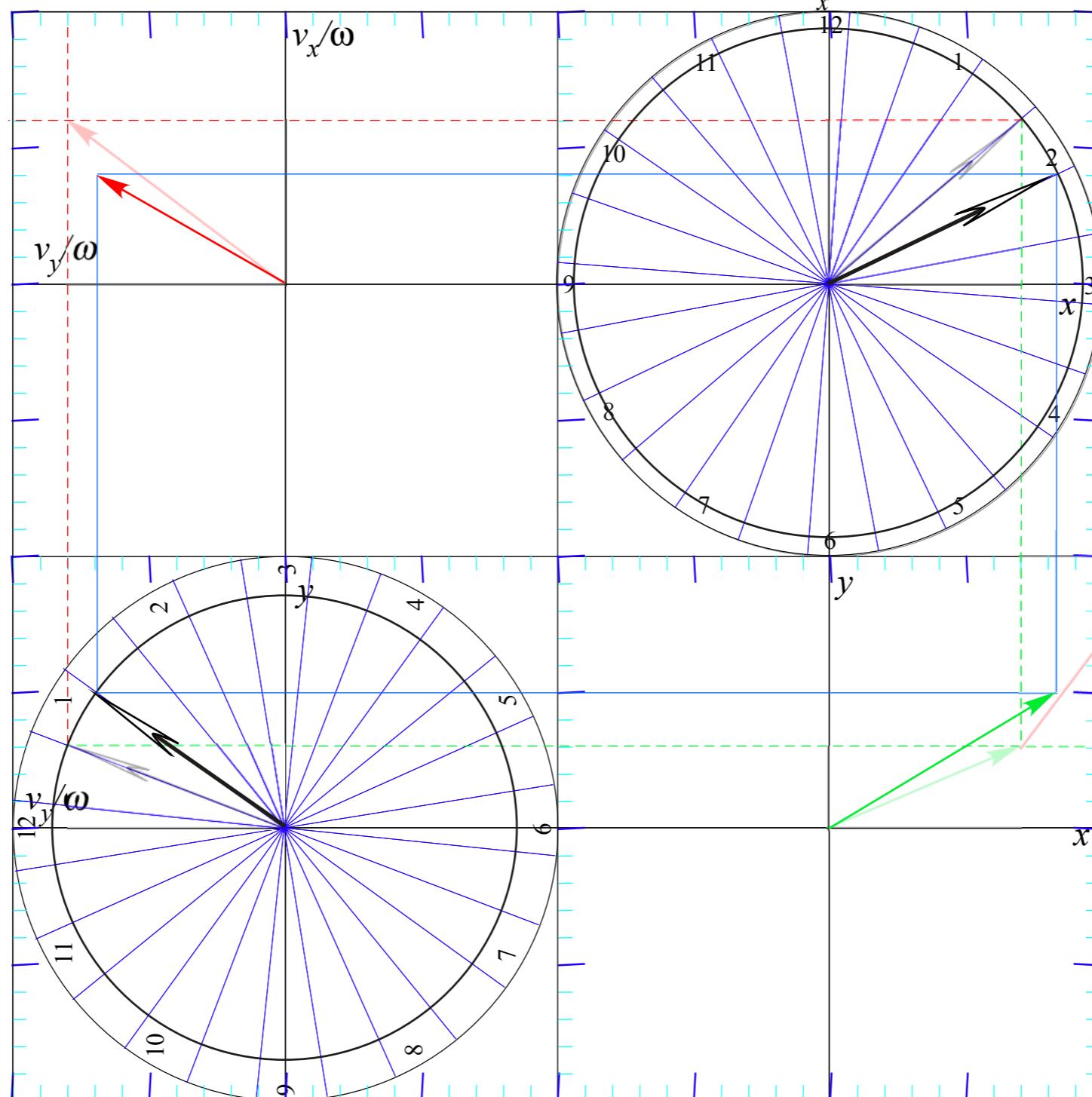
Initial position: $\mathbf{r}(0) = (7.0 \ 3.0)$

*Arbitrary initial position
 $\mathbf{r}(0) = (x(0), y(0))$*

*and initial velocity
 $\mathbf{v}(0) = (v_x(0), v_y(0))$*

*Usually have x and y
phasor circles of unequal size*

Initial velocity: $\mathbf{v}(0) = (-8.0 \ 6.0)$



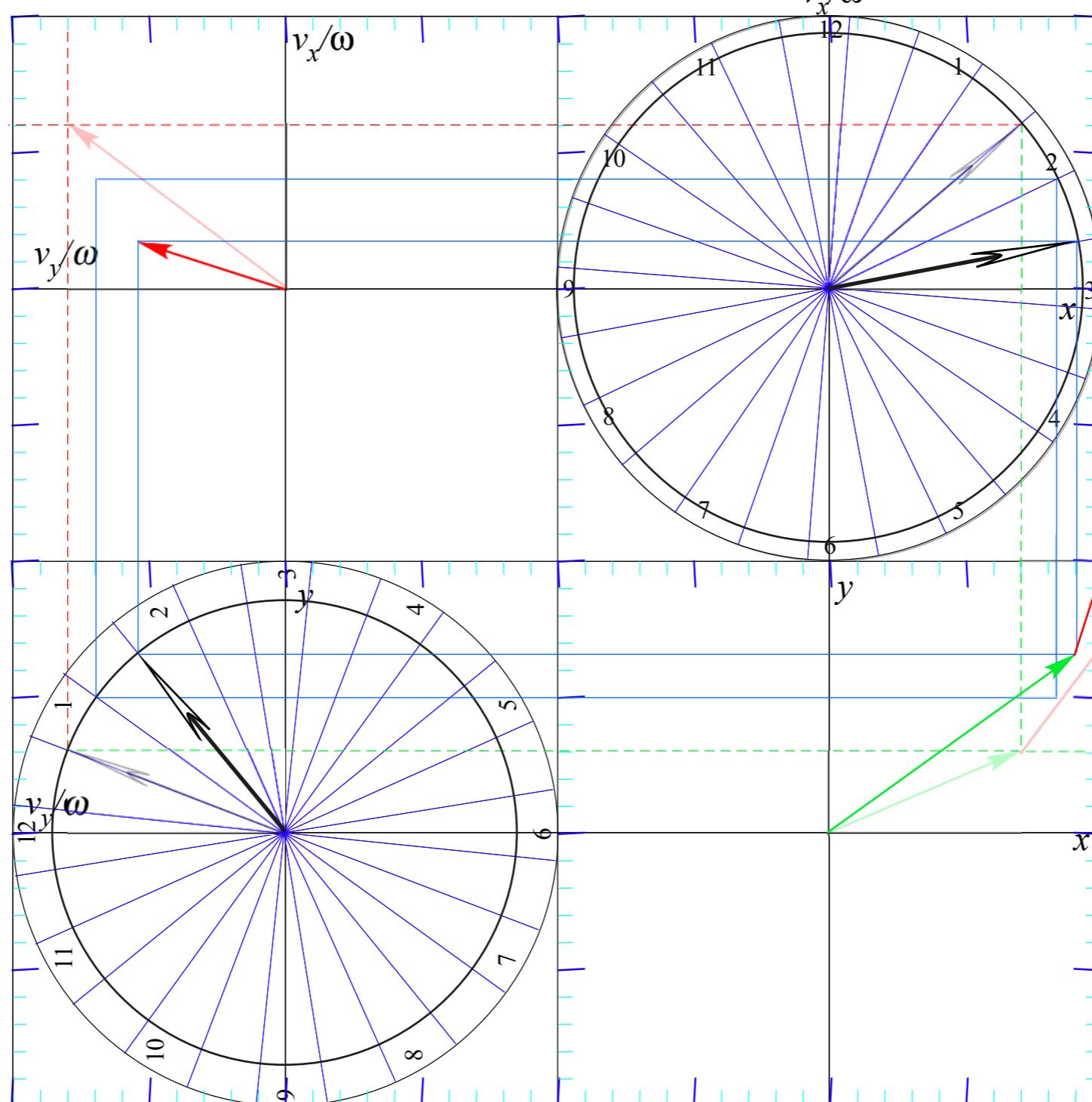
*Arbitrary initial position
 $\mathbf{r}(0) = (x(0), y(0))$*

*and initial velocity
 $\mathbf{v}(0) = (v_x(0), v_y(0))$*

*Usually have x and y
phasor circles of unequal size*

Initial position: $\mathbf{r}(0) = (7.0 \ 3.0)$

Initial velocity: $\mathbf{v}(0) = (-8.0 \ 6.0)$



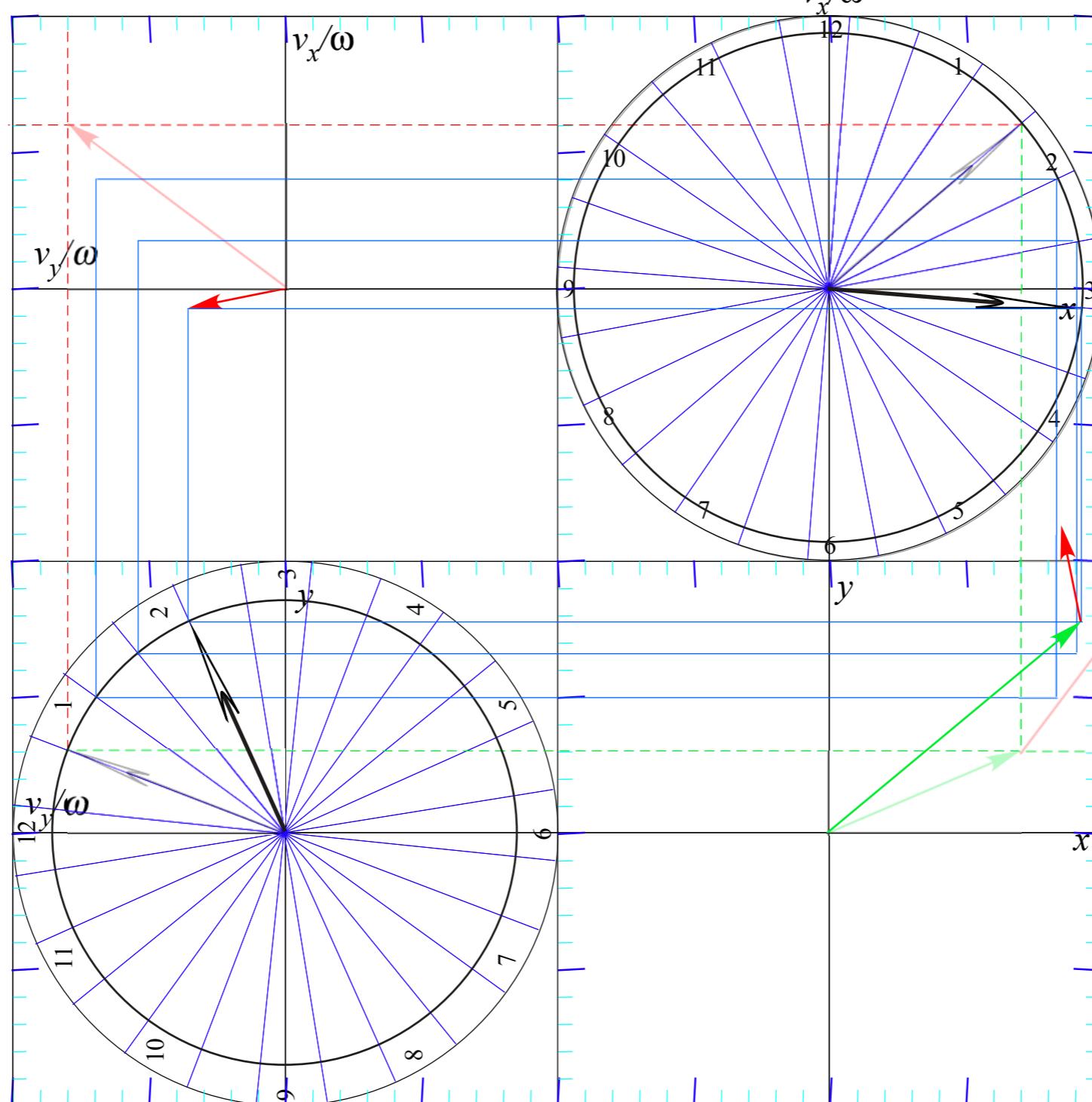
Initial position: $\mathbf{r}(0) = (7.0 \ 3.0)$

*Arbitrary initial position
 $\mathbf{r}(0) = (x(0), y(0))$*

*and initial velocity
 $\mathbf{v}(0) = (v_x(0), v_y(0))$*

*Usually have *x* and *y*
phasor circles of unequal size*

Initial velocity: $\mathbf{v}(0) = (-8.0 \ 6.0)$



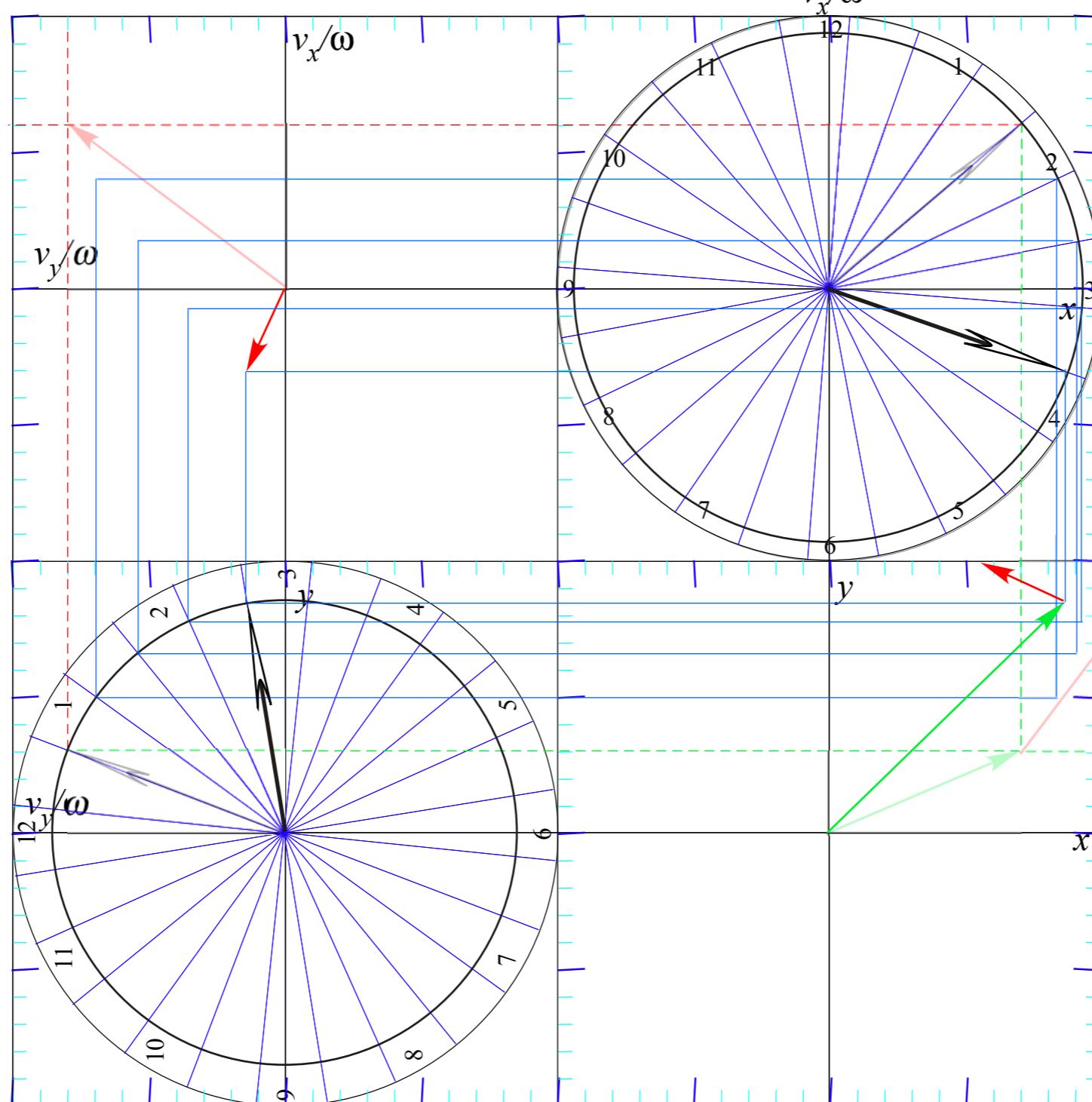
Initial position: $\mathbf{r}(0) = (7.0 \ 3.0)$

*Arbitrary initial position
 $\mathbf{r}(0) = (x(0), y(0))$*

*and initial velocity
 $\mathbf{v}(0) = (v_x(0), v_y(0))$*

*Usually have x and y
phasor circles of unequal size*

Initial velocity: $\mathbf{v}(0) = (-8.0 \ 6.0)$



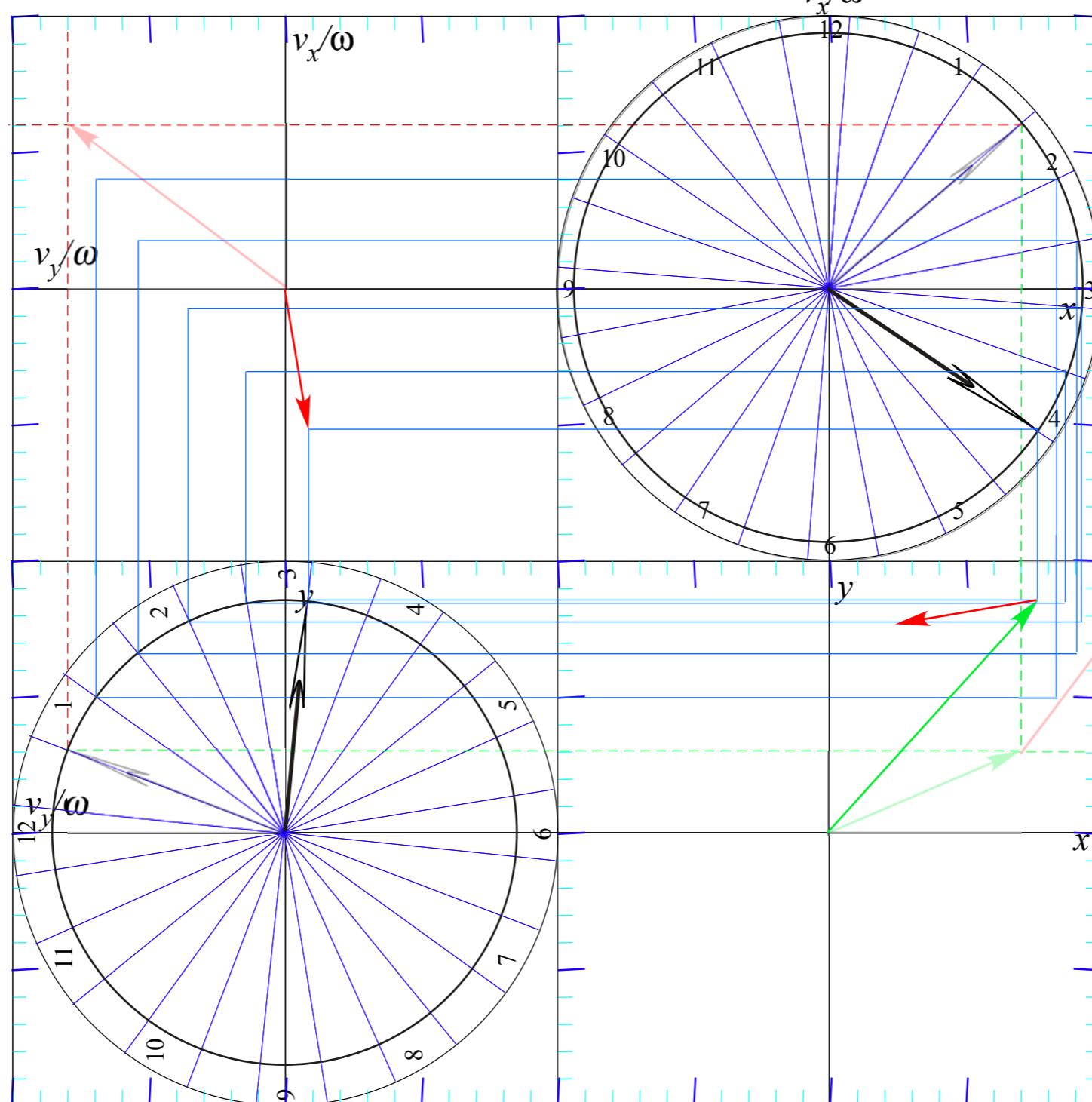
Initial position: $\mathbf{r}(0) = (7.0 \ 3.0)$

*Arbitrary initial position
 $\mathbf{r}(0) = (x(0), y(0))$*

*and initial velocity
 $\mathbf{v}(0) = (v_x(0), v_y(0))$*

*Usually have x and y
phasor circles of unequal size*

Initial velocity: $\mathbf{v}(0) = (-8.0 \ 6.0)$



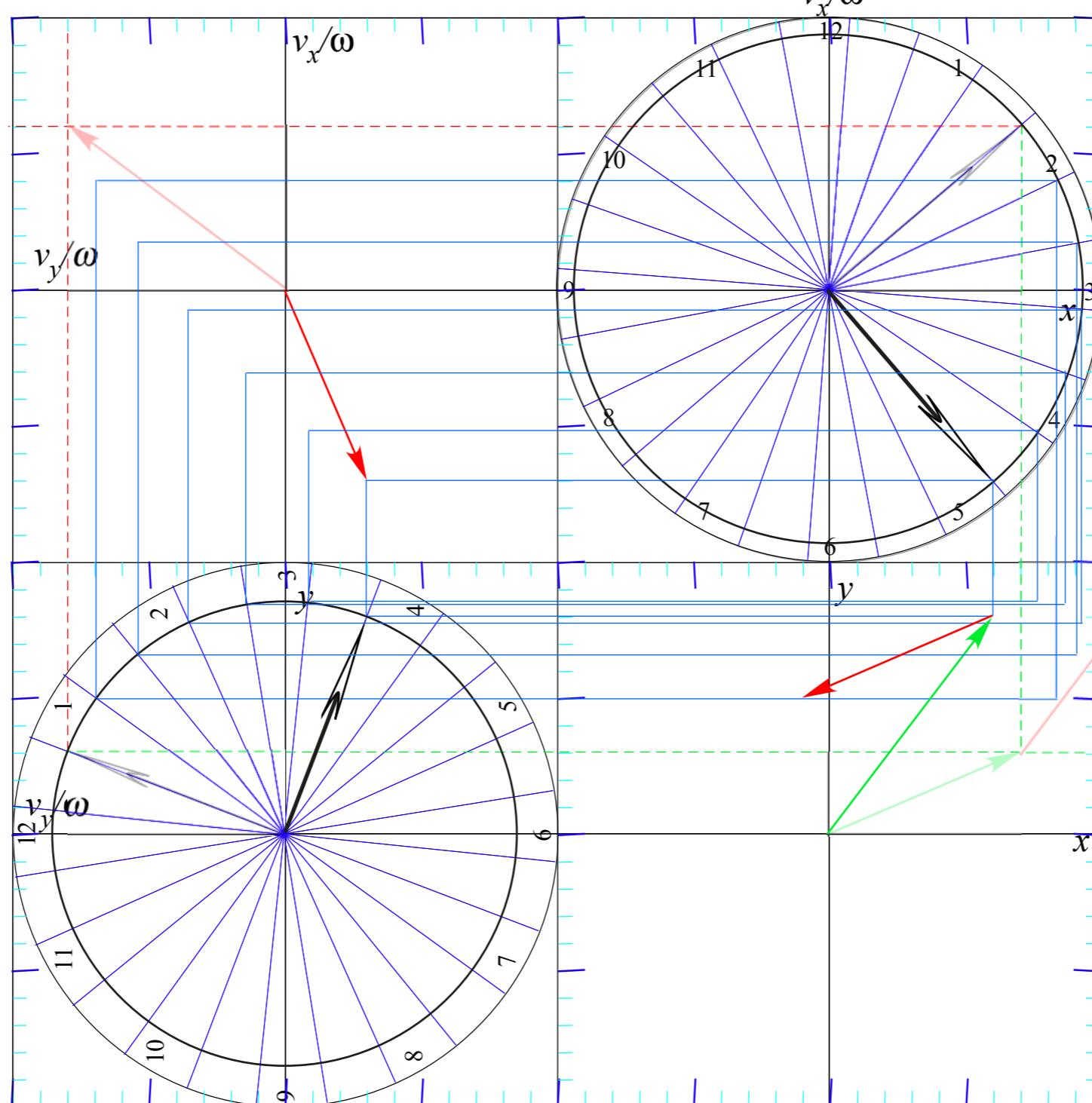
Initial position: $\mathbf{r}(0) = (7.0 \ 3.0)$

*Arbitrary initial position
 $\mathbf{r}(0) = (x(0), y(0))$*

*and initial velocity
 $\mathbf{v}(0) = (v_x(0), v_y(0))$*

*Usually have x and y
phasor circles of unequal size*

Initial velocity: $\mathbf{v}(0) = (-8.0 \ 6.0)$



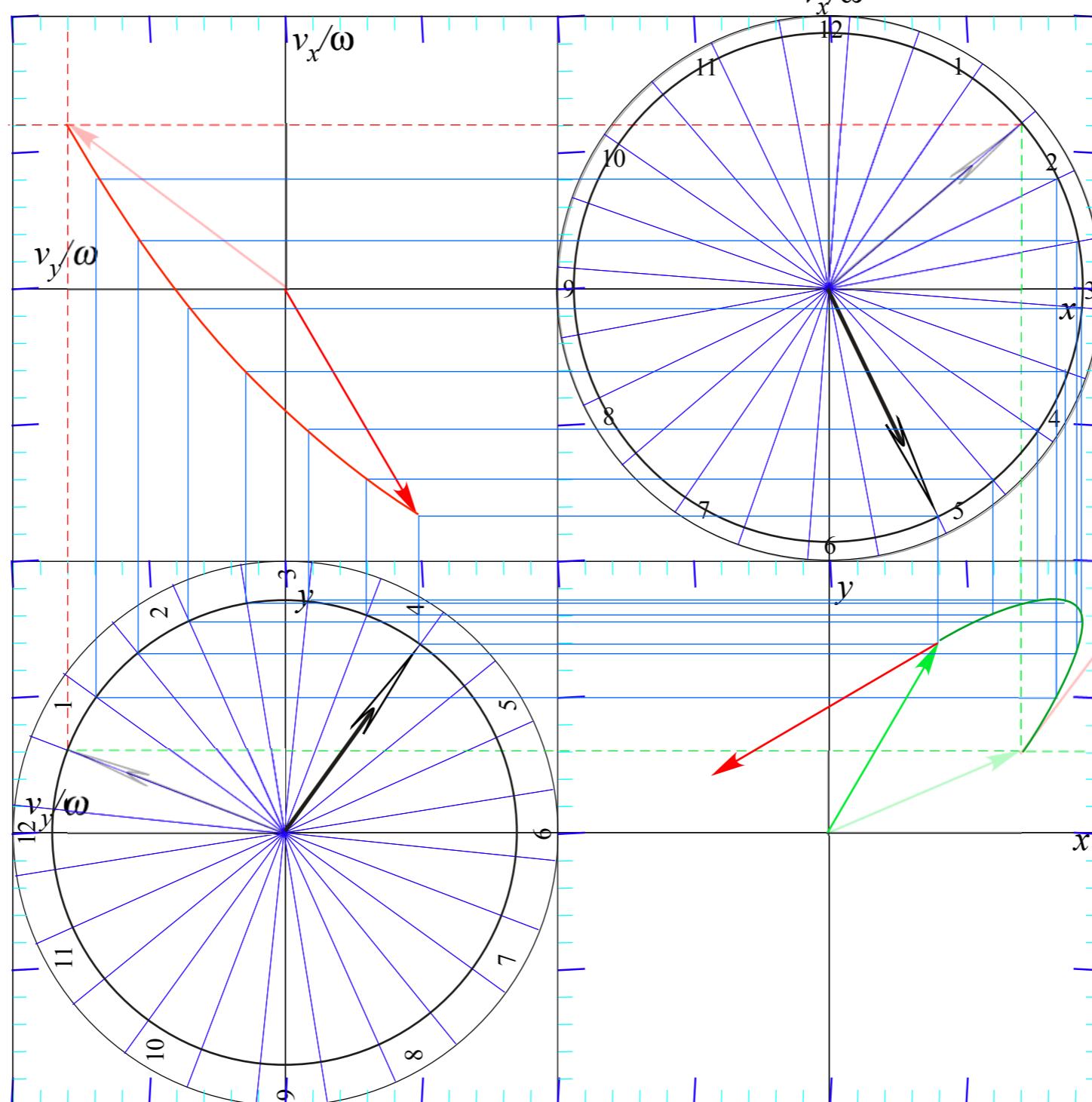
*Arbitrary initial position
 $\mathbf{r}(0) = (x(0), y(0))$*

*and initial velocity
 $\mathbf{v}(0) = (v_x(0), v_y(0))$*

*Usually have x and y
phasor circles of unequal size*

Initial position: $\mathbf{r}(0) = (7.0 \ 3.0)$

Initial velocity: $\mathbf{v}(0)=(-8.0 \ 6.0)$



Initial position: $\mathbf{r}(0)=(7.0 \ 3.0)$

*Arbitrary initial position
 $\mathbf{r}(0)=(x(0), y(0))$*

*and initial velocity
 $\mathbf{v}(0)=(v_x(0), v_y(0))$*

*Usually have x and y
phasor circles of unequal size*