

Relawavity : a novel introduction to relativistic mechanics III. (<u>CMwBang! Unit 8</u>, <u>AMOP Ch.0</u>,)

Review: Relawavity ρ functions and plots vs. ρ Derive relawavity parameters and Minkowski coordinates for v_R =2.5THz and v_L =0.5THz

Derivation of relativistic quantum mechanics What's the matter with mass? Shining some light on the Elephant in the room Relativistic action and Lagrangian-Hamiltonian relations Poincare' and Hamilton-Jacobi equations

Relativistic optical transitions and Compton recoil formulae

Feynman diagram geometry

Compton recoil related to rocket velocity formula

Comparing 2nd-quantization "photon" number N and 1st-quantization wavenumber κ

Relawavity in accelerated frames

Laser up-tuning by Alice and down-tuning by Carla makes *g*-acceleration grid Analysis of constant-*g* grid compared to zero-*g* Minkowsi grid Animation of mechanics and metrology of constant-*g* grid

This Lecture's Reference Link Listing

Web Resources - front page UAF Physics UTube channel <u>Quantum Theory for the Computer Age</u>

Principles of Symmetry, Dynamics, and Spectroscopy

Classical Mechanics with a Bang!

Modern Physics and its Classical Foundations

2017 Group Theory for QM 2018 Adv CM 2018 AMOP

2019 Advanced Mechanics

Lecture #28-30

In reverse order

AMOP Chapter 0: Space-Time Symmetry

AMOP Detailed Development of RelaWavity

2018 Rochester Talk (Auxilary Slides)

Special Relativity and Quantum Theory by Ruler and Compass - Earlier, expanded draft

Ruler & Compass - Relawavity Exercise

Relativity Visualized - Epstein-ip-1985 for sale here @www.allbookstores.com

GuideIt Web Simulations: $\underline{\sigma} = 30^{\circ}$, $\underline{\sigma} = 60^{\circ}$

Pirelli Site - A Colorful Road to Relativity Using Occam's Razors and Evenson's Lasers

World of Clocks - Animations - <u>12 hr. clock</u>, <u>24 hr. clock</u> <u>Phasor vs Thales(Pirelli Challenge)</u> - phasors_2_3_zoom_anim.html

RelativIt Web Simulations

Relativistic Events in Main Lighthouse's Space-Time Frame - scenario=22 Relativistic Events in Ship's Space-Time Frame - scenario=24 Epstein plot - scenario=600

BohrIt Web Simulations

 $\frac{2 \text{ CW ct vs x Plot (ck = \pm 2)} - scenario = -130022}{Multi-Panel 2 PW ct vs x Plot} - scenario = 30022}$ $\frac{1 \text{ CW ct vs x Plot (ck = -1)} - scenario = -30001}{1 \text{ CW ct vs x Plot (ck = +4)} - scenario = 30004}$ 2 CW Minkowski Plot (ck = -1, +4) - scenario = -30104

CMwBang Text 2012 Unit 6 page=5

BounceIt Web App/Scenarios: <u>5002</u>, <u>5003</u> Coullt Web App/Scenarios: <u>TwoParticleCollision_LToR</u>, <u>TwoParticleCollision_LToR_CM</u>, <u>TwoParticleOrbit_Coulomb</u>, <u>TwoParticleOrbit_Coulomb_CM</u>, <u>TwoParticleOrbit_Hooke, TwoParticleOrbit_Hooke_CM</u>

In development, but close to role out.

More Advanced QM and classical references will *soon* be available through our: <u>*References Page</u> Would be great to have our* <u>Apache SOLR</u> *Search & Index system up for a bigger* Bang!)</u>

RelaWavity Web Simulations

2019 RelaWavity Portal Page Relations between Hypergeometric and Hypergeometric functions - *plotType=0,9&...* RelaWavity Web Simulation {Physical Terms - All Terms} - *plotType=4,8*

Keyboard of the Gods

Per-Time vs Per-Space - plotType=7,1 Dual Plot #1 - plotType=7,2&bcStepInd=1 Dual Plot #2 - plotType=7,2&bcStepInd=2 Dual Plot #3 - plotType=7,2&bcStepInd=3 Dual Plot #7 - plotType=7,2&bcStepInd=7

<u>16 Relativity Dimensions</u> - *plotType=8,4* <u>Relativistic Terms (Expanded Table)</u> - *plotType=8,5* <u>Minkowski graph (Multi-plot)</u> - *plotType=8,8* <u>Detailed Thales Geometry</u> - *plotType=3,6* <u>PerSpace - PerTime {All}</u> - *plotType=3,6&minkGridPosCells=0* <u>Expanded Relativistic Relations</u> - *plotType=8,7* <u>Wavefronts in Space-Space</u> - *plotType=6,1* <u>Spectral Ellipse {PerSpace-PerSpace} { $\beta = u/c = 1/3$ } - *plotType=6,3&...* <u>Spectral Ellipse { $\beta = u/c = 3/4$ } - *plotType=6,3&...*</u></u>

Select, exciting, and/or related Research

Singular Motion of Asymetric Rotators AJP 44, 11 p1080 Harter-Kim-1976 Molecular Eigensolution Symmetry Analysis and Fine Structure - Harter-IJMS-2013 Lenz Vector and Orbital Analog Computers - AJP 44 p348 1976 Some Geometric Aspects of Classical Coulomb Scattering AJP 40 4 p1852 1972 How Molecules do Self-NMR - Harter-Mitchell-Columbus-2009 Classical Mechanics with a Bang! - Asymmetric Top Demo Allbookstores.com - Compare for Heller's SemiClassical Way - 0691163731 "My Bomerang Won't Come Back" (YouTube: Playlist) Rotating Solid Bodies in Microgravity (YouTube) Dancing T-handle in zero-g (YouTube)

Continued for 4 more pages ¬

This Lecture's Reference Link Listing

<u>Web Resources - front page</u> UAF Physics UTube channel Quantum Theory for the Computer Age

Principles of Symmetry, Dynamics, and Spectroscopy

Classical Mechanics with a Bang!

Modern Physics and its Classical Foundations

2017 Group Theory for QM 2018 Adv CM 2018 AMOP

2019 Advanced Mechanics

Lecture #22-27

In reverse order

Coullt Web App Simulations: <u>p19</u>, <u>p32</u>, <u>p72</u>, <u>p73</u>, <u>p92</u>, <u>R=-0.375</u>, <u>R=+0.5</u>, <u>Rutherford</u> OscillatorPE Web App: IHO Scenario 2, Coulomb Scenario 3 RelaWavity Web App/Simulator/Calculator: Elliptical - IHO orbits JerkIt Web App: 2-, 2+, Amp50Omega147-, Amp50Omega296, Amp50Omega602. Gap(1)MolVibes Web App: C3vN3 WaveIt Web App: Dim = 3 w/Wave Components;Static Char Table: <u>6</u>, <u>12</u>, <u>12(b)</u>, <u>16</u>, <u>36</u>, <u>256</u> Quantum Carpet with N=20: Gaussian, Boxcar Quantum Revivals of Morse Oscillators and Farey-Ford Geometry - Li-Harter-CPL-2015 QTCA Unit 5 Ch14 2013 Lester. R. Ford, Am. Math. Monthly 45,586(1938) John Farey, Phil. Mag.(1816) Wolfram Harter, J. Mol. Spec. 210, 166-182 (2001) Harter, Li IMSS (2013) Li, Harter, Chem. Phys. Letters (2015) Advanced Atomic and Molecular Optical Physics 2018 Class #9, pages: 5, 61 **BoxIt Web Simulations** Pure A-Type A=4.9, B=0, C=0, & D=4.0 Pure B-Type: A=4.0, B=-0.2, C=0, & D=4.0 Pure C-Type A,D=4.055, B=0, C=0.1 Mixed AB-Type w/Cosine Mixed AB Type A=4.0, BU2=0.866..., CU2=0, & D=1.0 w/Stokes & Freq rats Mixed AB Type A=5.086 B=-0.27 C=0 D=2.024 w/Stokes plot Mixed ABC Type A=4.833 B=0.2403 C=0.4162 D=4.277 w/Stokes plot Recent mixed ABC Type A=0.325 B=0.375 C=0.825 D=0.05 w/Stokes plot Select, exciting, and/or related Research

<u>This Indestructible NASA Camera Revealed Hidden Patterns on Jupiter</u> - seeker-yt-2019 <u>What did NASA's New Horizons discover around Pluto? - Astrum-yt-2018</u> <u>Synthetic_Chiral_Light_for_Efficient_Control_of_Chiral_Light-Matter_Interaction_-</u> <u>Ayuso-np-2019</u> Classical Mechanics with a Bang! 2018 Lectures 8, 9, 23 page 93 Text Unit 6, page=27 ColorU2 for the Web - in development Group Theory for Quantum Mechanics - 2017 Lectures: 6, 7, 8, and the combined 9-10 Quantum Theory for the Computer Age Unit 3 Ch.7-10, page=90 Spectral Decomposition with Repeated Eigenvalues - 2017 GTQM - Lecture 5 Web based 3D & XR ($x \in \{A, M, V\}$, R=Reality) <u>https://www.babylonis.com/</u> Web based 3D graphics WebGL API (Graphics Layer modeled after OpenGL) **Recent In-House draft Articles:** Springer handbook on Molecular Symmetry and Dynamics - Ch 32 -Molecular Symmetry AMOP Ch 0 Space-Time Symmetry - 2019 Seminar at Rochester Institute of Optics, Auxiliary slides, June 19, 2018 Quantum Computing - (Current) State of the Art - Reimer-www-2019 Geometric Algebra- A Guided Tour through Space and Time - Reimerwww-2019 Wildlife Monitoring Identification and Behavioral Study - Section 1 - Reimerwww-2019 Wildlife Monitoring Identification and Behavioral Study - Section 2 - Reimerwww-2019 Quantum Computing (*QC*) and Geometric Algebra (*GA*) references: **Ouantum Supremacy Using a Programmable Superconducting Processor - Arute-n-2019 Quantum Computing for Computer Scientists - Helwer-mr-yt-2018**, Slides Quantum Computing and Workforce, Curriculum, and App Devel - Roetteler-MS-2019 Quantum Computing - (Current) State of the Art - Reimer-www-2019 Excerpts (Page 44-47 in *Preliminary Draft*) for a GA take on the Complex Numbers Geometric Algebra- A Guided Tour through Space and Time - Reimer-www-2019 GA & OC references (Page 11-16 in Preliminary Draft)

This Lecture's Reference Link Listing

<u>Web Resources - front page</u> <u>UAF Physics UTube channel</u> Quantum Theory for the Computer Age

Principles of Symmetry, Dynamics, and Spectroscopy

Classical Mechanics with a Bang!

Modern Physics and its Classical Foundations

Lectures #12 through #21

In reverse order

2017 Group Theory for QM 2018 Adv CM 2018 AMOP 2019 Advanced Mechanics

Wiki on Pafnuty Chebyshev Nobelprize.org 2005 Physics Award

BoxIt Web Simulations:

A-Type w/Cosine, A-Type w/Freq ratios, AB-Type w/Cosine, AB-Type 2:1 Freq ratio

OscillIt Web Simulations:

Default/Generic, Weakly Damped #18, Forced : Way below resonance,On resonance Way above resonance,Underdamped Complex Response Plot

Coullt Web Simulations:

<u>Stark-Coulomb : Bound-state motion in parabolic coordinates</u> <u>Molecular Ion : Bound-state motion in hyperbolic coordinates</u> <u>Synchrotron Motion, Synchrotron Motion #2</u> <u>Mechanical Analog to EM Motion (YouTube video)</u> iBall demo - Quasi-periodicity (YouTube video)

Trebuchet Web Simulations:

Default/Generic URL, Montezuma's Revenge, Seige of Kenilworth, "Flinger", Position Space (Course), Position Space (Fine) Wacky Waving Solid Metal Arm Flailing Chaos Pendulum - Scooba_Steeve-yt-2015 Triple Double-Pendulum - Cohen-yt-2008 Punkin Chunkin - TheArmchairCritic-2011 Jersey Team Claims Title in Punkin Chunkin - sussexcountyonline-1999 Shooting range for medieval siege weapons. Anybody knows? - twcenter.net/forums The Trebuchet - Chevedden-SciAm-1995 NOVA Builds a Trebuchet

Recent Articles of Interest:

<u>A_Semi-Classical_Approach_to_the_Calculation_of_Highly_Excited_Rotational_Energies for</u> ...
 <u>Asymmetric-Top_Molecules_-_Schmiedt-pccp-2017</u>
 Tunable and broadband coherent perfect absorption by ultrathin blk phos metasurfaces - Guo-josab-2019
 Vortex Detection in Vector Fields Using Geometric Algebra - Pollock-aaca-2013.pdf

Pirelli Relativity Challenge (Introduction level) - Visualizing Waves:

Using Earth as a clock, Tesla's AC Phasors, Phasors using complex numbers. CM wBang Unit 1 - Chapter 10, pdf_page=135 Calculus of exponentials, logarithms, and complex fields, RelaWavity Web Simulation - Unit Circle and Hyperbola (Mixed labeling) Smith Chart, Invented by Phillip H. Smith (1905-1987)

Select, exciting, and related Research

Clifford Algebra And The Projective Model Of Homogeneous Metric Spaces -Foundations - Sokolov-x-2013 Geometric Algebra 3 - Complex Numbers - MacDonald-yt-2015 Biquaternion - Complexified Quaternion - Roots of -1 - Sangwine-x-2015 An Introduction to Clifford Algebras and Spinors - Vaz-Rocha-op-2016 Unified View on Complex Numbers and Quaternions- Bongardt-wcmms-2015 Complex Functions and the Cauchy-Riemann Equations - complex2 - Friedman-columbia-2019 An sp-hybridized Molecular Carbon Allotrope- cyclo-18-carbon - Kaiser-s-2019 An Atomic-Scale View of Cyclocarbon Synthesis - Maier-s-2019 Discovery Of Topological Weyl Fermion Lines And Drumhead Surface States in a Room Temperature Magnet - Belopolski-s-2019 "Weyl"ing away Time-reversal Symmetry - Neto-s-2019 Non-Abelian Band Topology in Noninteracting Metals - Wu-s-2019 What Industry Can Teach Academia - Mao-s-2019 RoVib- quantum state resolution of the C60 fullerene - Changala-Ye-s-2019 (Alt) A Degenerate Fermi Gas of Polar molecules - DeMarco-s-2019

An assist from *Physics Girl* (YouTube Channel):

How to Make VORTEX RINGS in a Pool Crazy pool vortex - pg-yt-2014 Fun with Vortex Rings in the Pool - pg-yt-2014

Running Reference Link Listing

Lectures #11 through #7

In reverse order

Eric J Heller Gallery:

Main portal, Consonance and Dissonance II, Bessel 21, Chladni

The Semiclassical Way to Molecular Spectroscopy - Heller-acs-1981 Quantum_dynamical_tunneling_in_bound_states_-_Davis-Hellerjcp-1981

Pendulum Web Simulation Cycloidulum Web Simulation

Links to previous lecture: <u>Page=74</u>, <u>Page=75</u>, <u>Page=79</u>

Pendulum Web Sim

Cycloidulum Web Sim

JerkIt Web Simulations: Basic/Generic: Inverted, FVPlot

CMwithBang Lecture 8, page=20

WWW.sciencenewsforstudents.org: Cassini - Saturnian polar vortex

"RelaWavity" Web Simulations:
<u>2-CW laser wave, Lagrangian vs Hamiltonian,</u> <u>Physical Terms Lagrangian L(u) vs Hamiltonian H(p)</u>
<u>Coullt Web Simulation of the Volcanoes of Io</u>
BohrIt Multi-Panel Plot:
Relativistically shifted Time-Space plots of 2 CW light waves

BoxIt Web Simulations:

<u>Generic/Default</u> <u>Most Basic A-Type</u> <u>Basic A-Type w/reference lines</u> <u>Basic A-Type A-Type with Potential energy</u> <u>A-Type with Potential energy and Stokes Plot</u> <u>A-Type w/3 time rates of change</u> <u>A-Type w/3 time rates of change with Stokes Plot</u> <u>B-Type (A=1.0, B=-0.05, C=0.0, D=1.0)</u>

RelaWavity Web Elliptical Motion Simulations:

Orbits with b/a=0.125 Orbits with b/a=0.5 Orbits with b/a=0.7 Exegesis with b/a=0.125 Exegesis with b/a=0.5 Exegesis with b/a=0.7 Contact Ellipsometry

Coullt Web Simulations: Basic/Generic

Exploding Starlet Volcanoes of Io (Color Quantized)

JerkIt Web Simulations:

<u>Basic/Generic</u> Catcher in the Eye - IHO with Linear Hooke perturbation - Force-potential-Velocity Plot

OscillatorPE Web Simulation:

Coulomb-Newton-Inverse_Square, Hooke-Isotropic Harmonic, Pendulum-Circular Constraint

AMOP Ch 0 Space-Time Symmetry - 2019 Seminar at Rochester Institute of Optics, Aux. slides-2018

NASA Astronomy Picture of the Day -<u>Io: The Prometheus Plume (Just Image)</u> <u>NASA Galileo - Io's Alien Volcanoes</u> <u>New Horizons - Volcanic Eruption Plume on Jupiter's moon IO</u> <u>NASA Galileo - A Hawaiian-Style Volcano on Io</u>

<u>Pirelli Site: Phasors animimation</u> <u>CMwithBang Lecture #6, page=70 (9.10.18)</u>

Select, exciting, and related Research & Articles of Interest:

Burning a hole in reality—design for a new laser may be powerful enough to pierce space-time - Sumner-KOS-2019 Trampoline mirror may push laser pulse through fabric of the Universe - Lee-ArsTechnica-2019 Achieving Extreme Light Intensities using Optically Curved Relativistic Plasma Mirrors - Vincenti-prl-2019 <u>A Soft Matter Computer for Soft Robots - Garrad-sr-2019</u> <u>Correlated Insulator Behaviour at Half-Filling in Magic-Angle Graphene Superlattices - cao-n-2018</u> <u>Sorting ultracold atoms in a three-dimensional optical lattice in a</u> realization of Maxwell's Demon - Kumar-n-2018 Synthetic three-dimensional atomic structures assembled atom by atom - Barredo-n-2018 Older ones: Wave-particle duality of C60 molecules - Arndt-Itn-1999 Optical Vortex Knots - One Photon _ At A Time - Tempone-Wiltshire-Sr-2018 Baryon Deceleration by Strong Chromofields in Ultrarelativistic ,

<u>Baryon_Deceleration_by_Strong_Chromofields_in_Ottrarelativistic_</u>, <u>Nuclear_Collisions - Mishustin-PhysRevC-2007</u>, <u>APS Link & Abstract</u> Hadronic Molecules - Guo-x-2017

Hidden-charm pentaquark and tetraquark states - Chen-pr-2016

Running Reference Link Listing

Lectures #6 through #1

In reverse order

RelaWavity Web Simulation: Contact EllipsometryBoxIt Web Simulation: Elliptical Motion (A-Type)CMwBang Course: Site Title PagePirelli Relativity Challenge: Describing Wave Motion With Complex PhasorsUAF Physics UTube channel

Velocity Amplification in Collision Experiments Involving Superballs - Harter, 1971 <u>MIT OpenCourseWare: High School/Physics/Impulse and Momentum</u> <u>Hubble Site: Supernova - SN 1987A</u>

BounceItIt Web Animation - Scenarios:

49:1 y vs t, 49:1 V2 vs V1, 1:500:1 - 1D Gas Model w/ faux restorative force (Cool), 1:500:1 - 1D Gas (Warm), 1:500:1 - 1D Gas Model (Cool, Zoomed in),
Farey Sequence - Wolfram
Fractions - Ford-AMM-1938
Monstermash BounceItIt Animations: 1000:1 - V2 vs V1, 1000:1 with t vs x - Minkowski Plot
Quantum Revivals of Morse Oscillators and Farey-Ford Geometry - Li-Harter-2013
Quantum Revivals of Morse Oscillators and Farey-Ford Geometry - Li-Harter-cpl-2015
Quant. Revivals of Morse Oscillators and Farey-Ford Geom. - Harter-Li-CPL-2015 (Publ.)
Velocity Amplification in Collision Experiments Involving Superballs-Harter-1971
WaveIt Web Animation - Scenarios: Quantum Carpet, Quantum Carpet wMBars, Quantum Carpet BCar, Quantum Carpet BCar_wMBars
Wave Node Dynamics and Revival Symmetry in Quantum Rotors - Harter-JMS-2001 (Publ.)

AJP article on superball dynamics <u>AAPT Summer Reading List</u> <u>Scitation.org - AIP publications</u> <u>HarterSoft Youtube Channel</u>

BounceIt Web Animation - Scenarios:

Generic Scenario: <u>2-Balls dropped no Gravity (7:1) - V vs V Plot (Power=4)</u> 1-Ball dropped w/Gravity=0.5 w/Potential Plot: <u>Power=1, Power=4</u> <u>7:1 - V vs V Plot: Power=1</u> <u>3-Ball Stack (10:3:1) w/Newton plot (y vs t) - Power=4</u> <u>3-Ball Stack (10:3:1) w/Newton plot (y vs t) - Power=1 w/Gaps</u> <u>4-Ball Stack (27:9:3:1) w/Newton plot (y vs t) - Power=4</u> <u>4-Newton's Balls (1:1:1:1) w/Newtonian plot (y vs t) - Power=4</u> <u>5-Ball Totally Inelastic (1:1:1:1:1) w/Gaps: Newtonian plot (t vs x), V6 vs V5 plot</u> <u>5-Ball Totally Inelastic Pile-up w/ 5-Stationary-Balls - Minkowski plot (t vs x1) w/Gaps</u>

BounceIt Dual plots

 $m_{1}:m_{2} = 3:1$ $v_{2} vs v_{1} and V_{2} vs V_{1}, (v_{1}, v_{2}) = (1, 0.1), (v_{1}, v_{2}) = (1, 0)$ $y_{2} vs y_{1} plots: (v_{1}, v_{2}) = (1, 0.1), (v_{1}, v_{2}) = (1, 0), (v_{1}, v_{2}) = (1, -1)$ Estrangian plot $V_{2} vs V_{1}: (v_{1}, v_{2}) = (0, 1), (v_{1}, v_{2}) = (1, -1)$ $m_{1}:m_{2} = 4:1$ $v_{2} vs v_{1}, v_{2} vs y_{1}$ $m_{1}:m_{2} = 100:1, (v_{1}, v_{2}) = (1, 0): V_{2} vs V_{1} Estrangian plot, y_{2} vs y_{1} plot$ With g=0 and 70:10 mass ratio With non zero g, velocity dependent damping and mass ratio of 70:35 $M_{1}=49, M_{2}=1 with Newtonian time plot$ $M_{1}=49, M_{2}=1 with V_{2} vs V_{1} plot$ Example with friction Low force constant with drag displaying a Pass-thru, Fall-Thru, Bounce-Off $m_{1}:m_{2}= 3:1 and (v_{1}, v_{2}) = (1, 0) Comparison with Estrangian$

X2 paper: <u>Velocity Amplification in Collision Experiments Involving Superballs - Harter, et. al. 1971 (pdf)</u> Car Collision Web Simulator: <u>https://modphys.hosted.uark.edu/markup/CMMotionWeb.html</u> Superball Collision Web Simulator: <u>https://modphys.hosted.uark.edu/markup/BounceItWeb.html</u>; with Scenarios: <u>1007</u> <u>BounceIt web simulation with g=0 and 70:10 mass ratio</u> <u>With non zero g, velocity dependent damping and mass ratio of 70:35</u> Elastic Collision Dual Panel Space vs Space: <u>Space vs Time (Newton)</u>, <u>Time vs. Space(Minkowski)</u> Inelastic Collision Dual Panel Space vs Space: <u>Space vs Time (Newton)</u>, <u>Time vs. Space(Minkowski)</u> Matrix Collision Simulator: M_1 =49, M_2 =1 V_2 vs V_1 plot <<Under Construction>>



BohrIt Web Simulation: ±600THz































Fig. 11 in text <u>Relawavity...</u>





Space-time $(c\tau', x')$ geometry of 2-CW paths

Fig. 4 in Ch.0 text introducing <u>Relawavity...</u>





Review: Relawavity ρ functions and plots vs. ρ

Derive relawavity parameters and Minkowski coordinates for v_R =2.5THz and v_L =0.5THz *Rapidity* ρ related to *stellar aberration angle* σ and L. C. Epstein's approach to relativity Longitudinal hyperbolic ρ -geometry connects to transverse circular σ -geometry "Occams Sword" and summary of 16 parameter functions of ρ and σ Applications to optical waveguide, spherical waves, and accelerator radiation

Derivation of relativistic quantum mechanics What's the matter with mass? Shining some light on the Elephant in the room Relativistic action and Lagrangian-Hamiltonian relations Poincare' and Hamilton-Jacobi equations

Relativistic optical transitions and Compton recoil formulae

Feynman diagram geometry

Compton recoil related to rocket velocity formula

Comparing 2nd-quantization "photon" number N and 1st-quantization wavenumber κ

Relawavity in accelerated frames

Laser up-tuning by Alice and down-tuning by Carla makes *g*-acceleration grid Analysis of constant-*g* grid compared to zero-*g* Minkowsi grid Animation of mechanics and metrology of constant-g grid

Usi	ng (s	some	e) wa	ve pa	iram	eters	to de	evelo	p relativistic quantum theory
	υ _{phas} CK _{phas}	e = B c $s_e = B s$	$\cosh \rho$ $\sinh \rho$	$\approx B + \frac{1}{2}$ $\approx B\rho$	$B\rho^2$ (f (f At lo	for $u \ll c$ for $u \ll c$ for spee	c) c) ds:	coshp≈1 sinh p≈µ	$B = v_A \text{sec.}^{-1}$ $B = v_A = C \kappa_A \text{sec.}^{-1}$ $\frac{m \cdot 1}{\text{sec.} m \cdot 1}$
			1						
group	$b_{\scriptscriptstyle RED}^{\scriptscriptstyle Doppler}$	V _{group} C	$rac{arphi_{group}}{arphi_{A}}$	$rac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_g}{\kappa_A}$	$rac{ au_{group}}{ au_A}$	V _{phase} C	$b_{\scriptscriptstyle BLUE}^{\scriptscriptstyle Doppler}$	
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$rac{C}{V_{phase}}$	$\frac{\kappa_{phase}}{\kappa_A}$	$rac{{m au}_{phase}}{{m au}_{A}}$	$\left(\begin{array}{c} \upsilon_{phase} \\ \upsilon_A \end{array} ight)$	$rac{\lambda_{phase}}{\lambda_A}$	$rac{c}{V_{group}}$	$\frac{1}{b_{\tiny RED}^{\tiny Doppler}}$	
rapidity ρ	$e^{-\rho}$	$tanh \rho$	$\sinh ho$	$\operatorname{sech}\rho$	$\cosh \rho$	cschp	$\operatorname{coth} \rho$	$e^{+ ho}$	DeleWaritz Wah Simulation Deletivistic Terror
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos\sigma$	$\sec \sigma$	$\cot \sigma$	$\csc\sigma$	$1/e^{-\rho}$	<u>(Short version)</u>
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$	
value for $\beta = 3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4}$ =0.75	$\frac{4}{5} = 0.80$	$\frac{5}{4}$ =1.25	$\frac{4}{3}$ =1.33	$\frac{5}{3}$ =1.67	$\frac{2}{1} = 2.0$	

Usi	ng (s	some) wa	ve pa	ram	eters	to de	evelo	p relativistic	e quantum	n theory
	v_{phas}	e = B c	$\cosh \rho$	$\approx B + \frac{1}{2}$	$B\rho^2$ (f	or $u \ll c$	c)	cosh <i>p</i> ≈i	$1 + \frac{1}{2}\rho^2 \approx 1 + \frac{1}{2}\frac{u^2}{2}$	$B = v_A$	sec. ⁻¹
	СК _{phas}	$B_{e} = B S$	$\sinh \rho$	$\approx B\rho$	(f	or $u \ll c$	c)	' sinh ∩≈/	$n \approx \frac{u}{c^2}$	$B = v_A$	$= C \kappa_A \operatorname{sec.}^{-1}$
	$-\left(\frac{u}{u}\right)$	= ta	nh $ ho$	$\approx \rho$	(for $u \ll c$)		c) `		C		$\frac{m.}{\text{sec.}} \frac{1}{m.}$
	C				At low speeds:						
ΨE											
		\overline{V}	1)	λ	ĸ	τ	<i>V</i> .				
group	$b_{\scriptscriptstyle RED}^{\scriptscriptstyle Doppler}$	$\frac{c}{c}$	v_{group} v_A	$\frac{\lambda_{group}}{\lambda_A}$	K _a roup	$\frac{\sigma_{group}}{\tau_A}$	r phase C	$b_{\scriptscriptstyle BLUE}^{\scriptscriptstyle Doppler}$			
phase	1 1 Doppler	$\frac{c}{\mathbf{v}}$	K phase	$ au_{\it phase}$	v_{phase}	λ_{phase}	$\frac{c}{\mathbf{V}}$	1 1 Doppler			
rapidity	$b_{BLUE}^{boppler}$	V _{phase}	κ_A	τ_A	v_A	λ_A	V _{group}	$b_{RED}^{Boppler}$			
ρ stellar \forall	$1/e^{+\rho}$	$\sin \sigma$	$\tan p$	$\cos\sigma$	sec σ	$\cot \sigma$	$\cos\sigma$	$1/e^{-\rho}$			
angle σ	1-R	R	1	$\sqrt{1-\beta^2}$	1	$\sqrt{\beta^{-2}-1}$	1	1 + R			
$\beta \equiv \frac{\alpha}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{P}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\left \frac{\sqrt{1-p}}{1} \right $	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{r}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1-\beta}{1-\beta}}$			
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4}$ =0.75	$\frac{4}{5} = 0.80$	$\frac{5}{4}$ =1.25	$\frac{4}{3}$ =1.33	$\frac{5}{3}$ =1.67	$\frac{2}{1} = 2.0$			

$$\begin{array}{c}
 (v_{phase} = B \cosh \rho) \approx B + \frac{1}{2} B \rho^{2} (\text{for } u \ll c) \\
 (c \kappa_{phase} = B \sinh \rho) \approx B \rho & (\text{for } u \ll c) \\
 (\frac{u}{c} = \tanh \rho \approx \rho) & (\text{for } u \ll c) \\
 (\frac{u}{c} = \tanh \rho \approx \rho) & (\text{for } u \ll c) \\
 (\text{for } u \ll c) & \text{At low speeds:} \\
 v_{phase} \approx B + \frac{1}{2} \frac{B}{c^{2}} u^{2} & \leftarrow \text{for } (u \ll c)
\end{array}$$

group	$b_{\scriptscriptstyle RED}^{\scriptscriptstyle Doppler}$	$\left(\frac{V_{group}}{c} \right)$	$rac{oldsymbol{v}_{group}}{oldsymbol{v}_A}$	$rac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{g}}{\kappa_{A}}$	$rac{{m au}_{group}}{{m au}_A}$	V _{phase} C	$b_{\scriptscriptstyle BLUE}^{\scriptscriptstyle Doppler}$
phase	$rac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$rac{\kappa_{phase}}{\kappa_{A}}$	$rac{{m au}_{phase}}{{m au}_A}$	$\left(egin{array}{c} \upsilon_{phase} \ \upsilon_A \end{array} ight)$	$rac{\lambda_{phase}}{\lambda_A}$	$rac{c}{V_{group}}$	$rac{1}{b_{\scriptscriptstyle RED}^{\scriptscriptstyle Doppler}}$
rapidity ρ	$e^{-\rho}$	$tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\mathrm{csch} ho$	$\operatorname{coth} \rho$	$e^{+ ho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	cscσ	1/ <i>e</i> ^{-p}
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta = 3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4}$ =0.75	$\frac{4}{5} = 0.80$	$\frac{5}{4}$ =1.25	$\frac{4}{3}$ =1.33	$\frac{5}{3}$ =1.67	$\frac{2}{1} = 2.0$

$$\begin{array}{c}
\upsilon_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^{2} (\text{for } u \ll c) \\
c\kappa_{phase} = B \sinh \rho \approx B \rho \\
\left(\text{for } u \ll c) \\
\hline
u_{c} = \tanh \rho \approx \rho \\
\hline
v_{phase} \approx B + \frac{1}{2} \frac{B}{c^{2}} u^{2} \\
\hline
\psi_{phase} \approx B + \frac{1}{2} \frac{B}{c^{2}} u^{2} \\
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\psi_{phase} \approx B + \frac{1}{2} \frac{B}{c^{2}} u^{2} \\
\hline
\psi_{phase} \approx B + \frac{1}{2} \frac$$

tin	ne	$b_{\scriptscriptstyle RED}^{\scriptscriptstyle Doppler}$	$rac{V_{group}}{c}$	$rac{oldsymbol{arphi}_{group}}{oldsymbol{arphi}_A}$	$rac{{m au}_{phase}}{{m au}_A}$	$\left(egin{array}{c} \upsilon_{phase} \ \upsilon_A \end{array} ight)$	$rac{{{ au }_{group}}}{{{ au }_{A}}}$	$rac{V_{phase}}{c}$	$b_{\scriptscriptstyle BLUE}^{\scriptscriptstyle Doppler}$
spa	ace	$rac{1}{b_{\scriptscriptstyle BLUE}^{\scriptscriptstyle Doppler}}$	$\frac{c}{V_{phase}}$	$rac{\kappa_{phase}}{\kappa_{A}}$	$rac{\lambda_{group}}{\lambda_A}$	$rac{\kappa_{group}}{\kappa_A}$	$rac{\lambda_{phase}}{\lambda_A}$	$rac{c}{V_{group}}$	$rac{1}{b_{\scriptscriptstyle RED}^{\scriptscriptstyle Doppler}}$
rapi ρ	dity 9	$e^{-\rho}$	$tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\mathrm{csch} ho$	$\operatorname{coth} \rho$	$e^{+ ho}$
stella angle	ur∀ leσ	$1/e^{+\rho}$	$\sin \sigma$	$tan\sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	cscσ	1/ <i>e</i> ^{-p}
β≡	$=\frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value β=3/	e for /5	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4}$ =0.75	$\frac{4}{5} = 0.80$	$\frac{5}{4}$ =1.25	$\frac{4}{3}$ =1.33	$\frac{5}{3}$ =1.67	$\frac{2}{1}$ =2.0

$$\frac{v_{phase} = B \cosh \rho}{c} \approx B + \frac{1}{2} B \rho^{2} (\text{for } u \ll c) \qquad \cosh \rho \approx 1 + \frac{1}{2} \rho^{2} \approx 1 + \frac{1}{2} u^{2} \qquad B = v_{A} \approx c^{-1} \qquad B \approx v_{A} \approx c^{-1} \qquad C \approx$$

$$\begin{array}{c}
 \underbrace{v_{phase} = B \cosh \rho}_{CK \ phase} \approx B + \frac{1}{2} B \rho^{2} (\text{for } u \ll c) \\
 \underbrace{cK \ phase} = B \sinh \rho \approx B \rho \quad (\text{for } u \ll c) \\
 \underbrace{u} \ c = \tanh \rho \approx \rho \quad (\text{for } u \ll c) \\
 \underbrace{u} \ c = \tanh \rho \approx \rho \quad (\text{for } u \ll c) \\
 \underbrace{w_{phase} \approx B + \frac{1}{2} \frac{B}{c_{s}^{2}} u^{2}}_{Phase} \approx \text{for } (u \ll c) \Rightarrow \\
 \underbrace{w_{phase} \approx B + \frac{1}{2} \frac{B}{c_{s}^{2}} u^{2}}_{Rescale \ v_{phase} \text{ by } h} \quad \text{so: } M = \frac{hB}{c^{2}} \\
 Rescale \ v_{phase} \text{ by } h \quad \text{so: } M = \frac{hB}{c^{2}} \\
 Rescale \ v_{phase} \text{ by } h \quad \text{so: } M = \frac{hB}{c^{2}} \\
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 Rescale \ v_{phase} \text{ by } h \quad \text{so } h = \frac{hB}{c^{2}} \\
 Rescale \ v_{phase} \text{ by } h \quad \text{so } h \quad \text{so } h \quad$$

group	$b_{\scriptscriptstyle RED}^{\scriptscriptstyle Doppler}$	$rac{V_{group}}{c}$	$rac{oldsymbol{v}_{group}}{oldsymbol{v}_A}$	$rac{\lambda_{group}}{\lambda_A}$	κ _g . _{oup} κ _A	$rac{{m au}_{group}}{{m au}_A}$	$rac{V_{phase}}{c}$	$b_{\scriptscriptstyle BLUE}^{\scriptscriptstyle Doppler}$
phase	$rac{1}{b_{\scriptscriptstyle BLUE}^{\scriptscriptstyle Doppler}}$	$\frac{c}{V_{phase}}$	$rac{\kappa_{phase}}{\kappa_{A}}$	$rac{{m au}_{phase}}{{m au}_{A}}$	$\left(egin{array}{c} arpsilon_{phase} \ arpsilon_{A} \end{array} ight)$	$rac{\lambda_{phase}}{\lambda_A}$	$rac{c}{V_{group}}$	$rac{1}{b_{\scriptscriptstyle RED}^{\scriptscriptstyle Doppler}}$
rapidity ρ	$e^{-\rho}$	$tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	cschp	$\operatorname{coth} \rho$	$e^{+ ho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$tan\sigma$	$\cos \sigma$	sec σ	$\cot \sigma$	csco	1/ <i>e</i> ^{-p}
$\beta \equiv \frac{u}{c}$	$\sqrt{rac{1-eta}{1+eta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{eta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5}$ =0.80	$\frac{5}{4}$ =1.25	$\frac{4}{3}$ =1.33	$\frac{5}{3}$ =1.67	$\frac{2}{1}$ =2.0

group	$b_{\scriptscriptstyle RED}^{\scriptscriptstyle Doppler}$	$rac{V_{group}}{c}$	$rac{oldsymbol{v}_{group}}{oldsymbol{v}_A}$	$rac{\lambda_{group}}{\lambda_A}$	$\frac{\kappa_{g}}{\kappa_{A}}$	$rac{{m au}_{group}}{{m au}_A}$	$rac{V_{phase}}{c}$	$b_{\scriptscriptstyle BLUE}^{\scriptscriptstyle Doppler}$
phase	$rac{1}{b_{\scriptscriptstyle BLUE}^{\scriptscriptstyle Doppler}}$	$\frac{c}{V_{phase}}$	$rac{\kappa_{phase}}{\kappa_{A}}$	$rac{{ au _{phase}}}{{ au _A}}$	$\left(\begin{array}{c} \upsilon_{phase} \\ \upsilon_{A} \end{array} ight)$	$rac{\lambda_{phase}}{\lambda_A}$	$rac{c}{V_{group}}$	$rac{1}{b_{\scriptscriptstyle RED}^{\scriptscriptstyle Doppler}}$
rapidity ρ	$e^{-\rho}$	$tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	$\mathrm{csch} ho$	$\operatorname{coth} \rho$	$e^{+ ho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$tan\sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	csco	1/ <i>e^{-p}</i>
$\beta \equiv \frac{u}{c}$	$\sqrt{rac{1-eta}{1+eta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4}$ =0.75	$\frac{4}{5}$ =0.80	$\frac{5}{4}$ =1.25	$\frac{4}{3}$ =1.33	$\frac{5}{3}$ =1.67	$\frac{2}{1}$ =2.0

group	$b_{\scriptscriptstyle RED}^{\scriptscriptstyle Doppler}$	$rac{V_{group}}{c}$	$rac{oldsymbol{v}_{group}}{oldsymbol{v}_A}$	$rac{\lambda_{group}}{\lambda_A}$	K _g . _{oup} K _A	$rac{{m au}_{group}}{{m au}_A}$	V _{phase} C	$b_{\scriptscriptstyle BLUE}^{\scriptscriptstyle Doppler}$
phase	$rac{1}{b_{\scriptscriptstyle BLUE}^{\scriptscriptstyle Doppler}}$	$\frac{c}{V_{phase}}$	$rac{\kappa_{phase}}{\kappa_{A}}$	$rac{{m au}_{phase}}{{m au}_A}$	$\left(egin{array}{c} \upsilon_{phase} \ \upsilon_A \end{array} ight)$	$rac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$rac{1}{b_{\scriptscriptstyle RED}^{\scriptscriptstyle Doppler}}$
rapidity ρ	$e^{-\rho}$	$tanh \rho$	$\sinh \rho$	$\operatorname{sech} \rho$	$\cosh \rho$	cschp	$\operatorname{coth} \rho$	$e^{+ ho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$tan \sigma$	$\cos \sigma$	$\sec \sigma$	$\cot \sigma$	csco	1/ <i>e</i> ^{-p}
$\beta \equiv \frac{u}{c}$	$\sqrt{rac{1-eta}{1+eta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4}$ =0.75	$\frac{4}{5}$ =0.80	$\frac{5}{4}$ =1.25	$\frac{4}{3}$ =1.33	$\frac{5}{3}$ =1.67	$\frac{2}{1}$ =2.0

$$\frac{\nabla u_{phase} = B \cosh \rho}{\left(\frac{\kappa}{p_{phase}} = B \sinh \rho\right) \approx B\rho} \quad (\text{for } u \ll c) \qquad \cosh \rho \approx 1 + \frac{1}{2} \rho^2 \approx 1 + \frac{1}{2} u^2 \qquad B = v_A \sec^{-1} \\ B =$$

$$\frac{v_{phase} = B\cosh\rho}{v_{phase} = B\sinh\rho} \approx B + \frac{1}{2}B\rho^{2}(\text{for } u \ll c) \qquad \cosh\rho \approx 1 + \frac{1}{2}\rho^{2} \approx 1 + \frac{1}{2}u^{2} \qquad B = v_{A} \sec^{-1}$$

$$B = v_{A}$$

ρ	C			seenp		esemp	comp	, j
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	sec σ	$\cot \sigma$	csco	1/ <i>e</i> ⁻
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+eta}{1-eta}}$
value for β=3/5	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4} = 0.75$	$\frac{4}{5} = 0.80$	$\frac{5}{4}$ =1.25	$\frac{4}{3} = 1.33$	$\frac{5}{3}$ =1.67	$\frac{2}{1} = 2.0$

$$\frac{v_{phase} = B \cosh \rho}{v_{phase} = B \sinh \rho} \approx B\rho \quad (\text{for } u \ll c) \qquad \cosh \rho \approx 1 + \frac{1}{2} \rho^{2} (\text{for } u \ll c) \qquad \cosh \rho \approx 1 + \frac{1}{2} \rho^{2} \approx 1 + \frac{1}{2} u^{2} \qquad B = v_{A} \sec^{-1} \qquad B = v_{A}$$

 $\frac{1+\beta}{1-\beta}$

 $\frac{2}{1} = 2.0$

 $\frac{1}{\beta}$

 $\frac{5}{3}$ =1.67

 $\frac{\sqrt{1-\beta^2}}{1}$

 $\frac{1-\beta}{1+\beta}$

 $\frac{1}{2} = 0.5$

 $\beta \equiv \frac{u}{c}$

value for β=3/5

 $\frac{\beta}{1}$

 $\frac{1}{\sqrt{\beta^{-2}-1}}$

 $\frac{3}{5} = 0.6 \quad \frac{3}{4} = 0.75$

 $\frac{1}{\sqrt{1-\beta^2}}$

 $\frac{4}{5} = 0.80 \quad \frac{5}{4} = 1.25 \quad \frac{4}{3} = 1.33$
$$\frac{\psi_{phase} = B\cosh \rho}{\left(\kappa_{phase} = B\sinh \rho\right) \approx B\rho} (\text{for } u \ll c)} \qquad \cosh \rho \approx \left(h^{2} + \frac{1}{2}\rho^{2} \approx 1 + \frac{1}{$$

 $v_{phase} = B \cosh \rho \approx B + \frac{1}{2} B \rho^2 (\text{for } u \ll c)$ $\cosh \rho \approx 1 + \frac{1}{2}\rho^2 \approx 1 + \frac{1}{2}\frac{u^2}{c^2}$ $\sinh \rho \approx \rho \approx \frac{u}{c}$ $B = v_A$ sec.⁻¹ $C\kappa_{phase} = B \sinh \rho \approx B\rho$ (for $u \ll c$) $B = v_A = c \kappa_A \operatorname{sec.}^{-1}$ $\frac{u}{c} = \tanh \rho \approx \rho$ (for $u \ll c$) sec. m. At low speeds: $v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2$ At low speeds: $\Leftarrow \text{ for } (u \ll c) \Rightarrow$ $\kappa_{phase} \approx \frac{D}{c^2} u$ \mathcal{U}_{phase} and \mathcal{K}_{phase} resemble formulae for Newton's kinetic Rescale v_{phase} by h so: $M = \frac{hB}{c^2}$ or: $hB = Mc^2$ (The famous Mc^2 shows up here!) energy $\frac{1}{2}Mu^2$ and momentum Mu. $hv_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2 \quad \Leftarrow \text{ for } (u \ll c) \Rightarrow \quad h\kappa_{phase} \approx \frac{hB}{c^2} u \quad \text{(So attach scale to match units.)}$ So attach scale factor h-Lucky coincidences?? Cheap trick?? $hv_{phase} \approx Mc^2 + \frac{1}{2}Mu^2 \quad \Leftarrow \text{ for } (u \ll c) \Rightarrow \quad h\kappa_{phase} \approx Mu$... Try <u>exact</u> U_{phase} ... $hv_{phase} = hB \cosh \rho = Mc^2 \cosh \rho$ Planck (1900) Total Energy: $E = \frac{Mc^2}{\sqrt{1-u^2/c^2}}$ v_{group} λ_{group} V_{phase} au_{group} $b_{\scriptscriptstyle RED}^{\scriptscriptstyle Doppler}$ $b_{\scriptscriptstyle BLUE}^{\scriptscriptstyle Doppler}$ group group λ_A au_{A} С v_{A} v_{phase} λ_{phase} κ_{phase} $au_{\it phase}$ $\frac{c}{V_{group}}$ С phase $b_{\scriptscriptstyle RED}^{\scriptscriptstyle Doppler}$ $b_{\scriptscriptstyle BLUE}^{\scriptscriptstyle Doppler}$ V_{phase} κ_A λ_{A} au_{A} v_{A} (old-fashioned notation) Max Planck rapidity $\sinh \rho$ $e^{-\rho}$ $tanh \rho$ $\operatorname{sech}\rho$ $\cosh \rho$ $\operatorname{csch}\rho$ $\operatorname{coth} \rho$ ρ 1858-1947 stellar ∀ $1/e^{+\rho}$ $1/e^{-\rho}$ SCO $\sin \sigma$ sec of $\cot \sigma$ $\tan \sigma$ $\cos\sigma$ angle σ $\sqrt{\beta^{-2}-1}$ $\sqrt{1-\beta^2}$ $\frac{1}{\sqrt{\beta^{-2}-1}}$) $\sqrt{\frac{1+\beta}{1-\beta}}$ $\frac{\beta}{1}$ $\frac{1}{\sqrt{1-\beta^2}}$ $\frac{1}{\beta}$ $\beta \equiv \frac{u}{c}$ $\frac{3}{5} = 0.6 \quad \frac{3}{4} = 0.75$ $\frac{5}{4} = 1.25$ $\frac{4}{3}$ =1.33 $\frac{4}{5} = 0.80$ $\frac{5}{3} = 1.67 \left| \frac{2}{1} = 2.0 \right|$ $\frac{1}{2} = 0.5$ value for $\beta = 3/5$

$$\frac{v_{phase} = B \cosh \rho}{c\kappa_{phase} = B \sinh \rho} \approx B\rho \quad (for \ u \ll c)$$

$$\frac{u}{c} = \tanh \rho \approx \rho \quad (for \ u \ll c)$$

$$\frac{u}{c} = \tanh \rho \approx \rho \quad (for \ u \ll c)$$

$$\frac{u}{c} = \tanh \rho \approx \rho \quad (for \ u \ll c)$$

$$At \ low \ speeds:$$

$$v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2 \quad \Leftrightarrow \ for \ (u \ll c) \Rightarrow \qquad \kappa_{phase} \approx \frac{B}{c^2} u \quad V_{phase} \ and \ \kappa_{phase} \ resemble$$
formulae for Newton's kinetic
energy $\frac{1}{2} Mu^2$ and momentum Mu .

$$hv_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2 \quad \Leftrightarrow \ for \ (u \ll c) \Rightarrow \qquad h\kappa_{phase} \approx \frac{hB}{c^2} u \qquad V_{phase} \ and \ \kappa_{phase} \ resemble$$
formulae for Newton's kinetic
energy $\frac{1}{2} Mu^2$ and momentum Mu .

$$hv_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2 \quad \Leftrightarrow \ for \ (u \ll c) \Rightarrow \qquad h\kappa_{phase} \approx \frac{hB}{c^2} u \qquad So \ attach \ scale \ factor \ h \ (or \ hN)$$

$$hv_{phase} \approx Mc^2 + \frac{1}{2} Mu^2 \quad \iff \ for \ (u \ll c) \Rightarrow \qquad h\kappa_{phase} \approx Mu$$

$$Hv_{phase} \approx Mc^2 + \frac{1}{2} Mu^2 \quad \iff \ for \ (u \ll c) \Rightarrow \qquad h\kappa_{phase} \approx Mu$$

$$Hv_{phase} \approx hB + Och \ v_{phase} = hB \cosh \rho = Mc^2 \cosh \rho$$

$$\frac{Nccd \ or \ replace}{r} = Mc^2 \cosh \rho = Mc^2 \cosh \rho$$

$$\frac{Nccd \ or \ replace}{r} = hB \cosh \rho = Mc^2 \cosh \rho$$

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$$\frac{Nccd \ or \ replace}{r} = hB \cosh \rho = Mc^2 \cosh \rho$$

$$\frac{Nccd \ or \ replace}{r} = hB \cosh \rho = h$$

$$\frac{\nabla_{phase} = B\cosh \rho}{(\kappa_{phase} = B\sinh \rho) \approx B\rho} (\text{for } u \ll c)} \qquad \cosh \rho \approx 1 + \frac{1}{2}\rho^{2} \approx 1 + \frac{1}{2}r^{2} e^{2} \text{ inh} \rho \approx \rho^{2} e^{2} e^{2}$$

2

$$\frac{\nabla_{phase} = B\cosh p}{\binom{w}{k}} \approx B + \frac{1}{2}B\rho^{2}(\text{for } u \ll c)$$

$$\frac{\nabla_{phase} = B\sinh p}{\binom{w}{k}} \approx B\rho \quad (\text{for } u \ll c)$$

$$\frac{U}{c} = \tanh \rho \approx p \quad (\text{for } u \ll c)$$

$$\frac{U}{c} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

$$\sinh \rho \approx \rho \approx \frac{U}{c}$$

$$\sinh \rho \approx \rho \approx \frac{U}{c}$$

$$\sinh \rho \approx \rho \approx \frac{U}{c}$$

$$\lim_{M \gg 0} P \approx \frac{H}{c}$$

$$\lim_{M \propto 0} P \approx \frac{H}{c}$$

$$\lim_{M \gg 0} P \approx \frac{H}{c}$$

$$\lim_{M \gg 0} P \approx \frac{H}{c}$$

$$\lim_{M \propto 0} P \approx \frac{H}{c}$$

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$$\lim_{M \gg 0} P \approx \frac{H}{c}$$

$$\lim_{M \gg 0} P \approx \frac{H}{c}$$

$$\lim_{M \gg 0} P \approx \frac{H}{c}$$

$$\lim_{M \propto 0} P$$

$$\frac{|v_{phase} = B \cosh \rho| \approx B + \frac{1}{2} B \rho^{2} (\text{for } u \ll c)}{|c\kappa_{phase} = B \sinh \rho| \approx B \rho} \quad (\text{for } u \ll c)$$

$$\frac{|u_{c}|}{|c|} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

$$\frac{|u_{c}|}{|c|} = \tanh \rho \approx \rho \quad (\text{for } u \ll c)$$

$$\sinh \rho \approx \rho \approx \frac{u}{c}$$

$$\lim h \rho \approx \frac{u}{$$

$$\frac{v_{phase} = B \cosh p}{(\kappa_{phase} = B \sinh p)} \approx B\rho \quad (for \ u \ll c)}$$

$$\frac{u}{c} = \tanh p \approx \rho \quad (for \ u \ll c)}{(for \ u \ll c)}$$

$$\frac{u}{c} = \tanh p \approx \rho \quad (for \ u \ll c)}{(hv \ phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2 \quad (for \ u \ll c)}{(for \ u \ll c)}$$

$$\frac{u}{c} = \tanh p \approx \rho \quad (for \ u \ll c)}{(for \ u \ll c)}$$

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$$\frac{u}{c} = \tanh p \approx \rho \quad (for \ u \ll c)}{(for \ u \ll c)}$$

$$\frac{u}{c} = \tanh p \approx \rho \quad (for \ u \ll c)}{(for \ u \ll c)} \qquad K_{phase} \approx \frac{B}{c^2} u \quad U_{phase} \ and \ K_{phase} \ resemble}{(for \ u \ll c)}$$

$$\frac{hv}{c} = \frac{hB}{c^2} u^2 \quad (for \ (u \ll c) \Rightarrow K_{phase} \approx \frac{B}{c^2} u \quad U_{phase} \ and \ K_{phase} \ resemble}{(for \ u \ll c)} \qquad K_{phase} \approx \frac{B}{c^2} u \quad U_{phase} \ and \ K_{phase} \ resemble}{(for \ u \ll c)}$$

$$\frac{hv}{c} = \frac{hB}{c^2} u^2 \quad (for \ (u \ll c) \Rightarrow hK_{phase} \approx \frac{hB}{c^2} u \quad (for \ u \ll c) \ (for \ u \ll c) \Rightarrow hK_{phase} \approx \frac{hB}{c^2} u \quad (for \ u \ll c) \ (for \ u$$

$$\frac{v_{phase} = B\cosh p}{c\kappa_{phase} = B\sinh p} \approx B\rho \quad (for \ u \ll c)$$

$$\frac{u}{c} = \tanh p \approx \rho \quad (for \ u \ll c)$$

$$\frac{u}{c} = \tanh p \approx \rho \quad (for \ u \ll c)$$

$$\frac{u}{c} = \tanh p \approx \rho \quad (for \ u \ll c)$$

$$\operatorname{At low speeds:} \quad v_{phase} \approx B + \frac{1}{2} \frac{B}{c^2} u^2 \quad \Leftrightarrow \text{ for } (u \ll c) \Rightarrow \quad \kappa_{phase} \approx \frac{B}{c^2} u \quad U_{phase} \text{ and } \kappa_{phase} \text{ resemble} \quad formulae for Newton's kinetic energy $\frac{1}{2}Mu^2$ and momentum Mu .
$$hv_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2 \quad \Leftrightarrow \text{ for } (u \ll c) \Rightarrow \quad h\kappa_{phase} \approx \frac{hB}{c^2} u \quad U_{phase} \text{ and } \kappa_{phase} \text{ resemble} \quad formulae for Newton's kinetic energy $\frac{1}{2}Mu^2$ and momentum Mu .
$$hv_{phase} \approx hB + \frac{1}{2} \frac{hB}{c^2} u^2 \quad \Leftrightarrow \text{ for } (u \ll c) \Rightarrow \quad h\kappa_{phase} \approx \frac{hB}{c^2} u \quad So \text{ attach scale factor } h \ (or \ hN) \text{ to match units.} \quad Try \ exact U_{phase} and \ \kappa_{phase} \dots \quad Try \ exact U_{phase} \dots \quad Try \ exact U_{phas$$$$$$





Relawavity variable tables

group	$b_{\scriptscriptstyle RED}^{\scriptscriptstyle Doppler}$	$\frac{V_{group}}{c}$	$rac{oldsymbol{v}_{group}}{oldsymbol{v}_A}$	$rac{\lambda_{group}}{\lambda_A}$	$rac{\kappa_{group}}{\kappa_A}$	$rac{ au_{group}}{ au_A}$	$rac{V_{phase}}{c}$	$b_{\scriptscriptstyle BLUE}^{\scriptscriptstyle Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$rac{\kappa_{phase}}{\kappa_{A}}$	$rac{oldsymbol{ au}_{phase}}{oldsymbol{ au}_{A}}$	$rac{oldsymbol{arphi}_{phase}}{oldsymbol{arphi}_A}$	$rac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{\tiny RED}^{\tiny Doppler}}$
rapidity ρ	$e^{- ho}$	$tanh \rho$	$\sinh ho$	$\operatorname{sech} \rho$	$\cosh ho$	csch <i>p</i>	$\operatorname{coth} \rho$	$e^{+ ho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec\sigma$	$\cot \sigma$	$\csc\sigma$	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$\frac{1}{\sqrt{eta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for β=3/5	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4}$ =0.75	$\frac{4}{5} = 0.80$	$\frac{5}{4}$ =1.25	$\frac{4}{3}$ =1.33	$\frac{5}{3}$ =1.67	$\frac{2}{1} = 2.0$
effects	$b_{\scriptscriptstyle RED}^{\scriptscriptstyle Doppler}$	V_{group}	past-future asymmetry (off-diagonal Lorentz-transform)	x-contraction ^(Lorentz) τ_{phase} -contraction	t-dilation ^(Einstein) v _{phase} -dilation (on-diagonal Lorentz-transform)	inverse asymmetry	V_{phase}	$b_{\scriptscriptstyle BLUE}^{\scriptscriptstyle Doppler}$

Relativistic quantum mechanics variable tables

group	$b_{\scriptscriptstyle RED}^{\scriptscriptstyle Doppler}$	V _{group}	$rac{oldsymbol{v}_{group}}{oldsymbol{v}_A}$	$rac{\lambda_{group}}{\lambda_A}$	$rac{\kappa_{group}}{\kappa_A}$	$rac{{m au}_{group}}{{m au}_A}$	V _{phase} C	$b_{\scriptscriptstyle BLUE}^{\scriptscriptstyle Doppler}$
phase	$\frac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$rac{\kappa_{phase}}{\kappa_{A}}$	$rac{{m au}_{phase}}{{m au}_{A}}$	$rac{oldsymbol{v}_{phase}}{oldsymbol{v}_A}$	$rac{\lambda_{phase}}{\lambda_A}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{\tiny RED}^{Doppler}}$
rapidity ρ	$e^{-\rho}$	$tanh \rho$	$\sinh ho$	$\operatorname{sech} \rho$	$\cosh ho$	$\mathrm{csch} ho$	$\mathrm{coth}\rho$	$e^{+\rho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$tan \sigma$	$\cos \sigma$	$\sec\sigma$	$\cot \sigma$	csco	$1/e^{-\rho}$
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	$\frac{\beta}{1}$	$rac{eta}{\sqrt{1-eta^2}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{1-eta^2}}{eta}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4}$ =0.75	$\frac{4}{5} = 0.80$	$\frac{5}{4}$ =1.25	$\frac{4}{3}$ =1.33	$\frac{5}{3}$ =1.67	$\frac{2}{1} = 2.0$
functions		$V_{group} = ctanh ho$	$momentum cp = Mc^2 \sinh \rho$	-Lagrangian $L=-Mc^2 \mathrm{sech}\rho$	Hamiltonian $H=Mc^2\cosh\rho$	$\begin{array}{l} DeBroglie\\ \lambda = \alpha \operatorname{csch} \rho \end{array}$	$\frac{V_{phase}}{c \cot h\rho} =$	

RelaWavity Web Simulation - Relativistic Terms (Expanded)



Review: Relawavity ρ functions and plots vs. ρ

Rapidity ρ related to stellar aberration angle σ and L. C. Epstein's approach to relativity Longitudinal hyperbolic ρ -geometry connects to transverse circular σ -geometry "Occams Sword" and summary of 16 parameter functions of ρ and σ Applications to optical waveguide, spherical waves, and accelerator radiation

Derivation of relativistic quantum mechanics

What's the matter with mass? Shining some light on the Elephant in the room Relativistic action and Lagrangian-Hamiltonian relations Poincare' and Hamilton-Jacobi equations

Relativistic optical transitions and Compton recoil formulae

Feynman diagram geometry

Compton recoil related to rocket velocity formula

Comparing 2nd-quantization "photon" number N and 1st-quantization wavenumber κ

Relawavity in accelerated frames

Laser up-tuning by Alice and down-tuning by Carla makes *g*-acceleration grid Analysis of constant-*g* grid compared to zero-*g* Minkowsi grid Animation of mechanics and metrology of constant-g grid Definition(s) of mass for relativity/quantum

Given: <u>Energy</u>: $E = Mc^2 \cosh \rho$

 $\frac{Rest Mass}{hB} M_{rest} (Einstein's mass)$ $hB = hv_A = Mc^2 = hc\kappa_A$

Defines invariant hyperbola(s) $E = \pm \sqrt{\left(Mc^2\right)^2 + (cp)^2}$

$$= hv_{phase}$$
momentum: $cp = Mc^{2} \sinh \rho$

$$= hc\kappa_{phase}$$
velocity: $u = c \tanh \rho = \frac{dv}{d\kappa}$

• What's the matter with Mass?



Shining some light on the elephant in the spacetime room

ct

Definition(s) of mass for relativity/quantum Given: <u>*Energy:*</u> $E = Mc^2 \cosh \rho$



• What's the matter with Mass?



Shining some light on the elephant in the spacetime room

ct





• What's the matter with Mass?



Shining some light on the elephant in the spacetime room

Ct



Rest Mass Mrest (Einstein's mass)
 Defines invariant hyperbola(s)
 momentum:

$$cp = Mc^2 \sinh \rho$$
 $hB = hv_A = Mc^2 = hc\kappa_A$
 $E = \pm \sqrt{(Mc^2)^2 + (cp)^2}$
 $momentum:$
 $cp = Mc^2 \sinh \rho$
 $\frac{hv_{phase}}{c^2} = M_{rest}$
 $\frac{Rest}{Mass}$
 $E = \pm \sqrt{(Mc^2)^2 + (cp)^2}$
 $momentum:$
 $cp = Mc^2 \sinh \rho$

 Momentum Mass M_{mom} (Galileo's mass) Defined by ratio p/u of relativistic momentum to group velocity.
 $u = c \tanh \rho = \frac{dv}{d\kappa}$

$$M_{mom} \equiv \frac{p}{u} = \frac{M_{rest}c \sinh \rho}{c \tanh \rho} \qquad \qquad \text{Limiting cases:} \qquad M_{mom} \xrightarrow{u \to c} M_{rest}e^{\rho}/2$$
$$= M_{rest}\cosh \rho = \frac{M_{rest}}{\sqrt{1 - u^2/c^2}} \qquad \underbrace{M_{omentum}}_{Mass}$$

Given: <u>Energy</u>: $E = Mc^2 \cosh \rho$ Definition(s) of mass for relativity/quantum $=hv_{phase}$ <u>Rest Mass</u> M_{rest} (Einstein's mass) $hB = hv_A = Mc^2 = hc\kappa_A$ Defines invariant hyperbola(s) $c_p = Mc^2 \sinh \rho$ momentum: $E = \pm \sqrt{\left(Mc^2\right)^2 + \left(cp\right)^2}$ $= hc\kappa_{phase}$ Rest $\frac{hase}{2} = M_{rest}$ <u>velocity:</u> $u = c \tanh \rho = \frac{dv}{d\kappa}$ Mass <u>Momentum Mass</u> M_{mom} (Galileo's mass) Defined by ratio p/u of relativistic momentum to group velocity. $M_{mom} \equiv \frac{p}{u} = \frac{M_{rest}c\sinh\rho}{c\tanh\rho}$ Limiting cases: $M_{mom} \xrightarrow{u \to c} M_{rest} e^{\rho}/2$ $M_{mom} \xrightarrow{u \ll c} M_{rest}$

<u>Effective Mass</u> M_{eff} (Newton's mass) Defined by ratio F/a=dp/du of relativistic force to acceleration.

 $= M_{rest} \cosh \rho = \frac{M_{rest}}{\sqrt{1 - u^2 / c^2}} \frac{Momentum}{Mass}$

Definition(s) of mass for relativity/quantum *Rest Mass Mrest (Einstein's mass)* $hB = hv_A = Mc^2 = hc\kappa_A$ $\underbrace{\frac{hv_{phase}}{c^2} = M_{rest}}_{c^2} = M_{rest}$ *Momentum Mass Mmom (Galileo's mass)* Defined by ratio p/u of relativistic momentum to group velocity. *E = Mc^2 cosh p a hv_phase for relativity/quantum B iven: E = Mc^2 cosh p a hv_phase for relativity/quantum B iven: E = Mc^2 cosh p a hv_phase for relativity/quantum B iven: E = Mc^2 cosh p a hv_phase for relativity for the phase for the p*

$$M_{mom} \equiv \frac{p}{u} = \frac{M_{rest}c\sinh\rho}{c\tanh\rho} \qquad \text{Limiting cases:} \qquad M_{mom} \xrightarrow{u \to c} M_{rest}e^{\rho/2}$$
$$= M_{rest}\cosh\rho = \frac{M_{rest}}{\sqrt{1 - u^2/c^2}} \qquad \underbrace{M_{mom}tum}_{Mass}$$

<u>Effective Mass</u> M_{eff} (Newton's mass) Defined by ratio F/a=dp/du of relativistic force to acceleration. That is ratio of change $dp=Mc \cosh\rho d\rho$ in momentum to change $du=c \operatorname{sech}^2\rho d\rho$ in group velocity.

$$= hv_{phase}$$

$$= hv_{phase}$$

$$= hv_{A} = Mc^{2} = hc\kappa_{A}$$

$$E = \pm \sqrt{\left(Mc^{2}\right)^{2} + (cp)^{2}}$$

$$= hc\kappa_{phase}$$

$$= hc\kappa_{phase}$$

$$= hc\kappa_{phase}$$

$$= hc\kappa_{phase}$$

$$= bc\kappa_{phase}$$

$$= bc\kappa_{phas$$

$$M_{mom} \equiv \frac{p}{u} = \frac{M_{rest}c\sinh\rho}{c\tanh\rho} \qquad \text{Limiting cases:} \qquad M_{mom} \xrightarrow{u \to c} M_{rest}e^{\rho}/2$$
$$= M_{rest}\cosh\rho = \frac{M_{rest}}{\sqrt{1 - u^2/c^2}} \qquad \frac{M_{omentum}}{Mass} \qquad M_{mom} \xrightarrow{u \to c} M_{rest}e^{\rho}/2$$

<u>Effective Mass</u> M_{eff} (Newton's mass) Defined by ratio F/a=dp/du of relativistic force to acceleration.

That is ratio of change $dp = Mc \cosh \rho d\rho$ in momentum to change $du = c \operatorname{sech}^2 \rho d\rho$ in group velocity.

$$M_{eff} \equiv \frac{dp}{du} = M_{rest} \frac{c \cosh \rho}{c \operatorname{sech}^2 \rho} = M_{rest} \cosh^3 \rho$$

$$\frac{Rest Mass}{hB} = hv_{A} = Mc^{2} = hc\kappa_{A}$$

$$\frac{hv_{phase}}{c^{2}} = M_{rest}$$

$$\frac{hv_{phase}}{c^{2}} = M_{rest}$$

$$\frac{Rest}{Mass}$$
Defines invariant hyperbola(s)
$$E = \pm \sqrt{\left(Mc^{2}\right)^{2} + (cp)^{2}}$$

$$\frac{momentum}{c}$$

$$\frac{momentu$$

$$M_{mom} \equiv \frac{p}{u} = \frac{M_{rest}c\sinh\rho}{c\tanh\rho} \qquad \text{Limiting cases:} \qquad M_{mom} \xrightarrow{u \to c} M_{rest}e^{\rho}/2$$
$$= M_{rest}\cosh\rho = \frac{M_{rest}}{\sqrt{1 - u^2/c^2}} \qquad \frac{M_{omentum}}{M_{ass}} \qquad M_{mom} \xrightarrow{u \to c} M_{rest}e^{\rho}/2$$

<u>Effective Mass</u> M_{eff} (Newton's mass) Defined by ratio F/a=dp/du of relativistic force to acceleration. That is ratio of change $dp=Mc \cosh\rho d\rho$ in momentum to change $du=c \operatorname{sech}^2\rho d\rho$ in group velocity.

$$M_{eff} \equiv \frac{dp}{du} = M_{rest} \frac{c \cosh \rho}{c \operatorname{sech}^2 \rho} \begin{bmatrix} -M_{rest} \cosh^3 \rho \\ \underline{Effective Mass} \end{bmatrix} \text{ Limiting cases: } M_{eff} \xrightarrow{u \to c} M_{rest} e^{3\rho}/2 \\ M_{eff} \xrightarrow{u \ll c} M_{rest} \end{bmatrix}$$

$$\frac{Rest Mass}{hB} = hv_{A} = Mc^{2} = hc\kappa_{A}$$

$$\frac{hv_{phase}}{c^{2}} = M_{rest}$$

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$$\frac{momentu$$

<u>Momentum Mass</u> M_{mom} (Galileo's mass) Defined by ratio p/u of relativistic momentum to group velocity.

$$M_{mom} \equiv \frac{p}{u} = \frac{M_{rest}c\sinh\rho}{c\tanh\rho} \qquad \text{Limiting cases:} \qquad M_{mom} \xrightarrow{u \to c} M_{rest}e^{\rho/2}$$
$$= M_{rest}\cosh\rho = \frac{M_{rest}}{\sqrt{1 - u^2/c^2}} \qquad \frac{M_{omentum}}{M_{ass}} \qquad M_{mom} \xrightarrow{u \to c} M_{rest}e^{\rho/2}$$

<u>Effective Mass</u> M_{eff} (Newton's mass) Defined by ratio F/a=dp/du of relativistic force to acceleration. That is ratio of change $dp=Mc \cosh\rho d\rho$ in momentum to change $du=c \operatorname{sech}^2\rho d\rho$ in group velocity.

$$M_{eff} \equiv \frac{dp}{du} = M_{rest} \frac{c \cosh \rho}{c \operatorname{sech}^2 \rho} = M_{rest} \cosh^3 \rho$$

$$\underbrace{\text{Limiting cases:}}_{Effective Mass} M_{eff} \xrightarrow{u \to c} M_{rest} e^{3\rho}/2$$

$$M_{eff} \xrightarrow{u \ll c} M_{rest}$$

More common derivation using group velocity: $u \equiv V_{group} = \frac{d\omega}{dk} = \frac{d\nu}{d\kappa}$

$$M_{eff} \equiv \frac{dp}{du} = \frac{\hbar dk}{dV_{group}} = \frac{\hbar}{\frac{d}{dk}\frac{d\omega}{dk}} = \frac{\hbar}{\frac{d^2\omega}{dk^2}} = \frac{M_{rest}}{\left(1 - u^2/c^2\right)^{3/2}}$$

$$\frac{Rest Mass}{hB} = hv_{A} = Mc^{2} = hc\kappa_{A}$$

$$\frac{hv_{phase}}{c^{2}} = M_{rest}$$

$$\frac{hv_{phase}}{c^{2}} = M_{rest}$$

$$\frac{momentum Mass}{M_{mom}} (Galileo's mass)$$
Defines invariant hyperbola(s)
$$E = \pm \sqrt{(Mc^{2})^{2} + (cp)^{2}}$$

$$\frac{momentum cp}{c^{2}} = M_{rest}$$

$$M_{mom} \equiv \frac{p}{u} = \frac{M_{rest} c \sinh \rho}{c \tanh \rho}$$
Limiting cases: $M_{mom} \xrightarrow{u \to c} M_{rest} e^{\rho/2}$

$$= M_{rest} \cosh \rho = \frac{M_{rest}}{\sqrt{1 - u^2/c^2}} \frac{Momentum}{Mass}$$

<u>Effective Mass</u> M_{eff} (Newton's mass) Defined by ratio F/a=dp/du of relativistic force to acceleration. That is ratio of change $dp=Mc \cosh\rho d\rho$ in momentum to change $du=c \operatorname{sech}^2\rho d\rho$ in group velocity.

$$M_{eff} = \frac{dp}{du} = M_{rest} \frac{c \cosh \rho}{c \operatorname{sech}^2 \rho} \left[= M_{rest} \cosh^3 \rho \right] \text{ Limiting cases: } M_{eff} \xrightarrow[u \to c]{} M_{rest} e^{3\rho/2} \\ M_{eff} \xrightarrow[u \to c]{} M_{rest} e^{3\rho/2} \\ M_{eff} \xrightarrow[u \to c]{} M_{rest} e^{3\rho/2} \\ M_{eff} = \frac{dp}{du} = \frac{\hbar dk}{dV_{group}} = \frac{\hbar}{\frac{d}{dk} \frac{d\omega}{dk}} = \frac{\hbar}{\frac{d^2 \omega}{dk^2}} = \frac{M_{rest}}{\left(1 - u^2/c^2\right)^{3/2}} = M_{rest} \cosh^3 \rho \\ \underbrace{Effective Mass}_{general wave formula} \text{ to accompany } V_{group} = \frac{d\omega}{dk} \\ \end{array} \right]$$

Definition(s) of mass for relativity/quantum

 $E = \hbar \omega$ <u>Rest Mass</u> M_{rest} (Einstein's mass) radius of curvature $hB = hv_A = Mc^2 = hc\kappa_A$ Finite-mass M $\frac{hv_{phase}}{c^2} = M_{rest} = \frac{hc\kappa_{phase}}{c^2} \frac{Rest}{Mass}$ dispersion function $\hbar\omega(ck)$ *Iomentum Mass M_{mom} (Galileo's mass)* Defined by *p/u E(p)* $M_{mom} \equiv \frac{p}{u} = \frac{M_{rest}c\sinh\rho}{c\tanh\rho}$ $Mc^2 = E$ minimum adius of curvature $= M_{rest} \cosh \rho = \frac{M_{rest}}{\sqrt{1 - \mu^2 / c^2}} \frac{Momentum}{Mass}$ momentum <u>Effective Mass</u> M_{eff} (Newton's mass) Defined by F/a=dp/duThat is ratio of $dp = Mc \cosh \rho \, d\rho$ to change $du = c \operatorname{sech}^2 \rho \, d\rho$ in velocity $cp = \hbar ck$ $M_{eff} \equiv \frac{dp}{du} = M_{rest} \frac{c \cosh \rho}{c \operatorname{sech}^2 \rho} \begin{bmatrix} =M_{rest} \cosh^3 \rho \\ \underline{Effective Mass} \end{bmatrix}$ More common derivation using group velocity: $u \equiv V_{group} = \frac{d\omega}{dk} = \frac{d\upsilon}{d\kappa}$ $M_{eff} = \frac{dp}{du} = \frac{\hbar dk}{dV_{group}} = \frac{\hbar}{\frac{d}{dk}\frac{d\omega}{dk}} = \frac{\hbar}{\frac{d^2\omega}{dk^2}} = \frac{\frac{\hbar}{d^2\omega}}{\left(1 - u^2/c^2\right)^{3/2}} = M_{rest}\cosh^3\rho$ <u>Effective Mass</u> Effective mass is proportional to the radius of curvature of $\omega(k)$ dispersion. general wave formula to accompany $V_{group} = \frac{d\omega}{dk}$

Definition(s) of mass for relativity/quantum How much mass does a γ -photon have?

Rest Mass $(a)\gamma$ -rest mass: $M_{rest}^{\gamma} = 0$,Newton complained about
his "corpuscles" of light having
"fits" (going crazy).Momentum Mass $(b)\gamma$ -momentum mass: $M_{mom}^{\gamma} = \frac{p}{c} = \frac{h\kappa}{c} = \frac{hv}{c^2}$,Newton complained about
his "corpuscles" of light having
"fits" (going crazy).Effective Mass $(c)\gamma$ -effective mass: $M_{eff}^{\gamma} = \infty$.Newton complained about
his "corpuscles" of light having
"fits" (going crazy).

$$M_{mom}^{\gamma} = \frac{h\nu}{c^2} = \nu(1.2 \cdot 10^{-51})kg \cdot s = 4.5 \cdot 10^{-36}kg \quad \text{(for: } \nu = 600\text{THz)}$$

Pirelli site discussion of optical mass-energy

https://pirelli.hosted.uark.edu/html/light_energy_flux_1.html

Pirelli site discussion of Wave amplitude effects

https://pirelli.hosted.uark.edu/html/amplitude_probability_1.html



Review: Relawavity ρ functions and plots vs. ρ

Rapidity ρ related to stellar aberration angle σ and L. C. Epstein's approach to relativity Longitudinal hyperbolic ρ -geometry connects to transverse circular σ -geometry "Occams Sword" and summary of 16 parameter functions of ρ and σ Applications to optical waveguide, spherical waves, and accelerator radiation

Derivation of relativistic quantum mechanics What's the matter with mass? Shining some light on the Elephant in the room Relativistic action and Lagrangian-Hamiltonian relations Poincare' and Hamilton-Jacobi equations

Relativistic optical transitions and Compton recoil formulae

Feynman diagram geometry

Compton recoil related to rocket velocity formula

Comparing 2nd-quantization "photon" number N and 1st-quantization wavenumber κ

Relawavity in accelerated frames

Laser up-tuning by Alice and down-tuning by Carla makes *g*-acceleration grid Analysis of constant-*g* grid compared to zero-*g* Minkowsi grid Animation of mechanics and metrology of constant-*g* grid

Define Lagrangian L using invariant wave phase $\Phi = kx - \omega t = k'x' - \omega't'$ for wave of $k = k_{phase}$ and $\omega = \omega_{phase}$.

$$L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega$$

$$h \equiv \frac{h}{2\pi}$$

 $\begin{pmatrix} hv_A = Mc^2 = hc\kappa_A \\ hv_{phase} = E = hv_A \cosh \rho \\ hc\kappa_{phase} = cp = hv_A \sinh \rho \end{pmatrix}$ Prior wave relations $\begin{pmatrix} \hbar\omega_A = Mc^2 = \hbar c\kappa_A \\ \hbar\omega_{phase} = E = \hbar\omega_A \cosh \rho \\ format & format \end{pmatrix}$ $\begin{pmatrix} \hbar\omega_A = Mc^2 = \hbar c\kappa_A \\ \hbar\omega_{phase} = E = \hbar\omega_A \cosh \rho \\ \hbar c\kappa_{phase} = cp = \hbar\omega_A \sinh \rho \end{pmatrix}$ $\hbar \equiv \frac{h}{2\pi}$

Define Lagrangian L using invariant wave phase $\Phi = kx - \omega t = k'x' - \omega't'$ for wave of $k = k_{phase}$ and $\omega = \omega_{phase}$.

Use DeBroglie-momentum $p=\hbar k$ relation and Planck-energy $E=\hbar \omega$ relation

 $L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega$ $\hbar \equiv \frac{n}{2\pi}$ $E = \hbar \omega = Mc^2 \cosh \rho$ $p = \hbar k = Mc \sinh \rho$ Prior wave relations $\hbar \omega_A = Mc^2 = \hbar c k_A$ $hv_A = Mc^2 = hc\kappa_A$ $\left| \begin{array}{c} \hbar \omega_{phase} = E = \hbar \omega_A \cosh \rho \\ \hbar c k_{phase} = cp = \hbar \omega_A \sinh \rho \end{array} \right| \hbar \equiv \frac{h}{2\pi}$ $hv_{phase} = E = hv_A \cosh \rho$ $hc\kappa_{phase} = cp = hv_A \sinh \rho$ -linear Hz angular phasor \rightarrow format format

Define Lagrangian L using invariant wave phase $\Phi = kx - \omega t = k'x' - \omega't'$ for wave of $k = k_{phase}$ and $\omega = \omega_{phase}$.

Use DeBroglie-momentum $p=\hbar k$ relation and Planck-energy $E=\hbar \omega$ relation

$$L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E \equiv p\dot{x} - E \qquad \hbar \equiv \frac{\hbar}{2\pi}$$

 $p = \hbar k = Mc \sinh \rho \qquad \qquad E = \hbar \omega = Mc^2 \cosh \rho$

 $hv_{A} = Mc^{2} = hc\kappa_{A}$ $hv_{phase} = E = hv_{A}\cosh\rho$ $hc\kappa_{phase} = cp = hv_{A}\sinh\rho$

Prior wave relations-linear Hzangular phasor \rightarrow formatformat

 $\frac{\hbar\omega_A = Mc^2 = \hbar ck_A}{\hbar\omega_{phase}} = E = \hbar\omega_A \cosh\rho$ $\frac{\hbar ck_{phase}}{\hbar ck_{phase}} = cp = \hbar\omega_A \sinh\rho$

Define Lagrangian L using invariant wave phase $\Phi = kx - \omega t = k'x' - \omega't'$ for wave of $k = k_{phase}$ and $\omega = \omega_{phase}$.

Use DeBroglie-momentum $p=\hbar k$ relation and Planck-energy $E=\hbar\omega$ relation to define Hamiltonian H=E

$$L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E \equiv p\dot{x} - E \equiv pu - H = L \qquad \text{Legendre} \\ \text{transformation}$$

 $p = \hbar k = Mc \sinh \rho$ $E = \hbar \omega = Mc^2 \cosh \rho = H$

 $hv_{A} = Mc^{2} = hc\kappa_{A}$ $hv_{phase} = E = hv_{A}\cosh\rho$ $hc\kappa_{phase} = cp = hv_{A}\sinh\rho$

Prior wave relations-linear Hzangular phasor \rightarrow formatformat

 $\frac{\hbar\omega_A = Mc^2 = \hbar ck_A}{\hbar\omega_{phase} = E = \hbar\omega_A \cosh\rho} \hbar \equiv \frac{h}{2\pi} \frac{h}{2\pi}$

Define Lagrangian L using invariant wave phase $\Phi = kx - \omega t = k'x' - \omega't'$ for wave of $k = k_{phase}$ and $\omega = \omega_{phase}$.

Use DeBroglie-momentum $p=\hbar k$ relation and Planck-energy $E=\hbar\omega$ relation to define Hamiltonian H=E

$$L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E \equiv p\dot{x} - E \equiv pu - H = L \qquad \text{Legendre} \\ \text{Use Group velocity} : u = \frac{dx}{dt} = c \tanh \rho$$

 $p = \hbar k = Mc \sinh \rho \qquad \qquad E = \hbar \omega = Mc^2 \cosh \rho = H$



Define Lagrangian L using invariant wave phase $\Phi = kx - \omega t = k'x' - \omega't'$ for wave of $k = k_{phase}$ and $\omega = \omega_{phase}$.

Use DeBroglie-momentum $p=\hbar k$ relation and Planck-energy $E=\hbar\omega$ relation to define Hamiltonian H=E

$$L = \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E = p\dot{x} - E = pu - H = L \qquad \frac{Legendre}{transformation}$$

Use Group velocity : $u = \frac{dx}{dt} = c \tanh \rho$
 $L = pu - H = (Mc \sinh \rho)(c \tanh \rho) - Mc^2 \cosh \rho$

 $\begin{array}{c} h\upsilon_{A} = Mc^{2} = hc\kappa_{A} \\ h\upsilon_{phase} = E = h\upsilon_{A}\cosh\rho \\ hc\kappa_{phase} = cp = h\upsilon_{A}\sinh\rho \end{array} \begin{array}{c} Prior wave relations \\ \leftarrow \text{linear Hz} & \text{angular phasor} \rightarrow \end{array} \begin{array}{c} \hbar\omega_{A} = Mc^{2} = \hbar c\kappa_{A} \\ \hbar\omega_{phase} = E = \hbar\omega_{A}\cosh\rho \\ \hbar\sigma_{phase} = E = \hbar\omega_{A}\cosh\rho \\ \hbarc\kappa_{phase} = cp = \hbar\omega_{A}\sinh\rho \end{array} \begin{array}{c} \hbar = \frac{h}{2\pi} \end{array}$

p

Define Lagrangian L using invariant wave phase $\Phi = kx - \omega t = k'x' - \omega't'$ for wave of $k = k_{phase}$ and $\omega = \omega_{phase}$.

Use DeBroglie-momentum $p=\hbar k$ relation and Planck-energy $E=\hbar\omega$ relation to define Hamiltonian H=E

$$L = \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E \equiv p\dot{x} - E \equiv pu - H = L \qquad \text{transformation} \\ \text{Use Group velocity : } u = \frac{dx}{dt} = c \tanh \rho \\ \text{Use Group velocity : } u = \frac{dx}{dt} = c \tanh \rho \\ \frac{dt}{dt} = h \omega = Mc^2 \cosh \rho = H \\ L = pu - H = (Mc \sinh \rho)(c \tanh \rho) - Mc^2 \cosh \rho \\ = Mc^2 \frac{\sinh^2 \rho - \cosh^2 \rho}{\cosh \rho} = -Mc^2 \operatorname{sech} \rho \\ L \text{ is : } Mc^2 \frac{-1}{\cosh \rho} = -Mc^2 \operatorname{sech} \rho$$

 $\begin{array}{c} hv_A = Mc^2 = hc\kappa_A \\ hv_{phase} = E = hv_A \cosh\rho \\ hc\kappa_{phase} = cp = hv_A \sinh\rho \end{array} \begin{array}{c} Prior wave relations \\ \leftarrow \text{linear Hz} & \text{angular phasor} \rightarrow \end{array} \begin{array}{c} \hbar\omega_A = Mc^2 = \hbar ck_A \\ \hbar\omega_{phase} = E = \hbar\omega_A \cosh\rho \\ \hbar\sigma_{phase} = E = \hbar\omega_A \cosh\rho \\ \hbar ck_{phase} = cp = \hbar\omega_A \sinh\rho \end{array} \begin{array}{c} \hbar = \frac{h}{2\pi} \end{array}$

Relativistic action S and Lagrangian-Hamiltonian relations Define Lagrangian L using invariant wave phase $\Phi = kx - \omega t = k'x' - \omega't'$ for wave of $k = k_{phase}$ and $\omega = \omega_{phase}$. Use DeBroglie-momentum $p=\hbar k$ relation and Planck-energy $E=\hbar\omega$ relation to define Hamiltonian H=E $L \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E \equiv p\dot{x} - E \equiv pu - H = L$ Legenare
transformation
Use Group velocity: $u = \frac{dx}{dt} = c \tanh \rho$ dt $E = \hbar \omega = Mc^2 \cosh \rho = H$ $p = \hbar k = Mc \sinh \rho$ Note: $Mcu = Mc^2 \tanh \rho$ $L = \rho u - H = (Mc \sinh \rho)(c \tanh \rho) - Mc^2 \cosh \rho$ $= Mc^{2} \frac{\sinh^{2} \rho - \cosh^{2} \rho}{\cosh \rho} = -\frac{Mc^{2} \operatorname{sech} \rho}{\cosh \rho}$ Compare Lagrangian L $L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech}\rho$


Relativistic action S and Lagrangian-Hamiltonian relations Define Lagrangian L using invariant wave phase $\Phi = kx - \omega t = k'x' - \omega't'$ for wave of $k = k_{phase}$ and $\omega = \omega_{phase}$. For the problem of t Use DeBroglie-momentum $p=\hbar k$ relation and Planck-energy $E=\hbar\omega$ relation to define Hamiltonian H=E $E = \hbar \omega = Mc^2 \cosh \rho = H$ $p = \hbar k = Mc \sinh \rho$ Note: $Mcu = Mc^2 \tanh \rho$ $L = \rho u - H = (Mc \sinh \rho)(c \tanh \rho) - Mc^2 \cosh \rho$ $= Mc^{2} \frac{\sinh^{2} \rho - \cosh^{2} \rho}{\cosh \rho} = -Mc^{2} \operatorname{sech} \rho$ Compare Lagrangian L re Lagrangian L $L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech}\rho$ with Hamiltonian H=E $H = \hbar\omega = Mc^2 / \sqrt{1 - \frac{u^2}{c^2}} = Mc^2 \cosh \rho$

 $\begin{array}{c} h\upsilon_{A} = Mc^{2} = hc\kappa_{A} \\ h\upsilon_{phase} = E = h\upsilon_{A}\cosh\rho \\ hc\kappa_{phase} = cp = h\upsilon_{A}\sinh\rho \end{array} \begin{array}{c} Prior wave relations \\ \leftarrow \text{linear Hz} & \text{angular phasor} \rightarrow \end{array} \begin{array}{c} \hbar\omega_{A} = Mc^{2} = \hbar c\kappa_{A} \\ \hbar\omega_{phase} = E = \hbar\omega_{A}\cosh\rho \\ \hbar c\kappa_{phase} = cp = \hbar\omega_{A}\sinh\rho \end{array} \begin{array}{c} \hbar = \frac{h}{2\pi} \end{array}$

Define Lagrangian L using invariant wave phase $\Phi = kx \cdot \omega t = k'x' \cdot \omega't'$ for wave of $k = k_{phase}$ and $\omega = \omega_{phase}$.

For the problem of t Use DeBroglie-momentum $p=\hbar k$ relation and Planck-energy $E=\hbar\omega$ relation to define Hamiltonian H=E $E = \hbar \omega = Mc^2 \cosh \rho = H$ $p = \hbar k = Mc \sinh \rho$ $L = pu - H = (Mc \sinh \rho)(c \tanh \rho) - Mc^2 \cosh \rho$ Note: $Mcu = Mc^2 \tanh \rho$ $= Mc^{2} \frac{\sinh^{2} \rho - \cosh^{2} \rho}{\cosh \rho} = -Mc^{2} \operatorname{sech} \rho$ Also: $cp = Mc^2 \sinh \rho$ Compare Lagrangian L $L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech}\rho$ with Hamiltonian H=E $H = \hbar\omega = Mc^2 / \sqrt{1 - \frac{u^2}{c^2}} = Mc^2 \cosh\rho$ $=Mc^2\sqrt{1+\sinh^2\rho}=Mc^2\sqrt{1+(cp)^2}$

 $\begin{pmatrix} h\upsilon_A = Mc^2 = hc\kappa_A \\ h\upsilon_{phase} = E = h\upsilon_A \cosh\rho \\ hc\kappa_{phase} = cp = h\upsilon_A \sinh\rho \end{pmatrix}$ Prior wave relations $\begin{pmatrix} \hbar\omega_A = Mc^2 = \hbar c\kappa_A \\ \hbar\omega_{phase} = E = \hbar\omega_A \cosh\rho \\ \hbar\sigma_{phase} = E = \hbar\omega_A \cosh\rho \\ \hbar c\kappa_{phase} = cp = \hbar\omega_A \sinh\rho \end{pmatrix}$ $\hbar = \frac{h}{2\pi}$

Relativistic action S and Lagrangian-Hamiltonian relations Define Lagrangian L using invariant wave phase $\Phi = kx - \omega t = k'x' - \omega't'$ for wave of $k = k_{phase}$ and $\omega = \omega_{phase}$. Use DeBroglie-momentum $p=\hbar k$ relation and Planck-energy $E=\hbar\omega$ relation to define Hamiltonian H=E $\left(\frac{dS}{dt} = L\right) \equiv \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E \equiv p\dot{x} - E \equiv pu - H = L \qquad \frac{Legendre}{transformation}$ Legendre Use Group velocity: $u = \frac{dx}{dt} = c \tanh \rho$ $\boldsymbol{E} = \hbar\boldsymbol{\omega} = Mc^2 \cosh \rho = \boldsymbol{H}$ $p = \hbar k = Mc \sinh \rho$ $=c\sin\sigma$ Note: $Mcu = Mc^2 \tanh \rho$ $L = pu - H = (Mc \sinh \rho)(c \tanh \rho) - Mc^2 \cosh \rho$ $= Mc^{2} \frac{\sinh^{2} \rho - \cosh^{2} \rho}{\cosh \rho} = -Mc^{2} \operatorname{sech} \rho$ $= Mc^2 \sin \sigma$ Also: $cp = Mc^2 \sinh \rho$ Compare Lagrangian L $=\hbar ck = Mc^2 \tan \sigma$ $\dot{S} = L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}}$ $= -Mc^2 \operatorname{sech}\rho = -Mc^2 \cos\sigma$ Including with Hamiltonian H=E $H=\hbar\omega=Mc^2/\sqrt{1-\frac{u^2}{c^2}}$ stellar $= Mc^2 \cosh \rho = Mc^2 \sec \sigma$ angle σ $=Mc^2\sqrt{1+\sinh^2\rho}=Mc^2\sqrt{1+(cp)^2}$ **Define** Action $S = \hbar \Phi$)

 $\begin{pmatrix} hv_A = Mc^2 = hc\kappa_A \\ hv_{phase} = E = hv_A \cosh \rho \\ hc\kappa_{phase} = cp = hv_A \sinh \rho \end{pmatrix}$ Prior wave relations $\begin{pmatrix} \hbar\omega_A = Mc^2 = \hbar c\kappa_A \\ \hbar\omega_{phase} = E = \hbar\omega_A \cosh \rho \\ format & format \end{pmatrix}$ $\begin{pmatrix} \hbar\omega_A = Mc^2 = \hbar c\kappa_A \\ \hbar\omega_{phase} = E = \hbar\omega_A \cosh \rho \\ \hbar c\kappa_{phase} = cp = \hbar\omega_A \sinh \rho \end{pmatrix}$ $\hbar \equiv \frac{h}{2\pi}$





Rapidity ρ related to *stellar aberration angle* σ and L. C. Epstein's approach to relativity Longitudinal hyperbolic ρ -geometry connects to transverse circular σ -geometry "Occams Sword" and summary of 16 parameter functions of ρ and σ Applications to optical waveguide, spherical waves, and accelerator radiation

Derivation of relativistic quantum mechanics

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Use DeBroglie-momentum $p=\hbar k$ relation and Planck-energy $E=\hbar\omega$ relation to define Hamiltonian H=E

$$\frac{dS}{dt} = L = \hbar \frac{d\Phi}{dt} = \hbar k \frac{dx}{dt} - \hbar \omega = p \frac{dx}{dt} - E \equiv p\dot{x} - E \equiv pu - H = L \quad \text{transformation}$$

Compare Lagrangian L

$$\dot{S} = L = \hbar \dot{\Phi} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} = -Mc^2 \operatorname{sech}\rho = -Mc^2 \cos\sigma$$

with Hamiltonian $H=E$
 $H = \hbar\omega = Mc^2 / \sqrt{1 - \frac{u^2}{c^2}} = Mc^2 \cosh\rho = Mc^2 \sec\sigma$
 $= Mc^2 \sqrt{1 + \sinh^2\rho} = Mc^2 \sqrt{1 + (cp)^2}$

 $\begin{array}{c} h\upsilon_{A} = Mc^{2} = hc\kappa_{A} \\ h\upsilon_{phase} = E = h\upsilon_{A}\cosh\rho \\ hc\kappa_{phase} = cp = h\upsilon_{A}\sinh\rho \end{array} \begin{array}{c} Prior \ wave \ relations \\ \leftarrow \text{linear Hz} & \text{angular phasor} \rightarrow \end{array} \begin{array}{c} \hbar\omega_{A} = Mc^{2} = \hbar ck_{A} \\ \hbar\omega_{phase} = E = \hbar\omega_{A}\cosh\rho \\ \hbar c\kappa_{phase} = cp = \hbar\omega_{A}\sinh\rho \end{array} \begin{array}{c} \hbar e = \frac{h}{2\pi} \end{array}$

Define Lagrangian L using invariant wave phase $\Phi = kx - \omega t = k'x' - \omega't'$ for wave of $k = k_{phase}$ and $\omega = \omega_{phase}$.

Use DeBroglie-momentum $p=\hbar k$ relation and Planck-energy $E=\hbar\omega$ relation to define Hamiltonian H=E $\begin{pmatrix}
\frac{dS}{dt} = L \\
\frac{dV}{dt} = \hbar k \frac{dx}{dt} - \hbar\omega = p \frac{dx}{dt} - E \equiv p\dot{x} - E \equiv pu - H = L$ Legendre
transformation

 $(dS \equiv Ldt \equiv \hbar d\Phi) = \hbar k dx - \hbar \omega dt = p dx - H dt$

Poincare Invariant action differential

$$\begin{array}{l} \begin{array}{l} \hline Compare \ Lagrangian \ L} \\ \hline \dot{S} = L = \hbar \dot{\Phi} \end{array} = -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} \\ \hline with \ Hamiltonian \ H = E \\ H = \hbar \omega = Mc^2 / \sqrt{1 - \frac{u^2}{c^2}} \\ \hline \end{array} = -Mc^2 \cosh \rho \\ \hline = Mc^2 \cosh \rho \\ = Mc^2 \operatorname{sec} \sigma \\ = Mc^2 \sqrt{1 + \sinh^2 \rho} = Mc^2 \sqrt{1 + (cp)^2} \end{array}$$

 $\begin{array}{c} h\upsilon_{A} = Mc^{2} = hc\kappa_{A} \\ h\upsilon_{phase} = E = h\upsilon_{A}\cosh\rho \\ hc\kappa_{phase} = cp = h\upsilon_{A}\sinh\rho \end{array} \begin{array}{c} Prior wave relations \\ \leftarrow \text{linear Hz} & \text{angular phasor} \rightarrow \end{array} \begin{array}{c} \hbar\omega_{A} = Mc^{2} = \hbar c\kappa_{A} \\ \hbar\omega_{phase} = E = \hbar\omega_{A}\cosh\rho \\ \hbar \sigma_{phase} = E = \hbar\omega_{A}\cosh\rho \\ \hbar \sigma_{phase} = E = \hbar\omega_{A}\cosh\rho \\ \hbar \sigma_{phase} = cp = \hbar\omega_{A}\sinh\rho \end{array} \begin{array}{c} \hbar \sigma_{A} = Mc^{2} = \hbar c\kappa_{A} \\ \hbar \sigma_{phase} = E = \hbar\omega_{A}\cosh\rho \\ \hbar \sigma_{phase} = cp = \hbar\omega_{A}\sinh\rho \end{array}$

Define Lagrangian L using invariant wave phase $\Phi = kx - \omega t = k'x' - \omega't'$ for wave of $k = k_{phase}$ and $\omega = \omega_{phase}$.

Use DeBroglie-momentum $p=\hbar k$ relation and Planck-energy $E=\hbar\omega$ relation to define Hamiltonian H=E $dS = Ldt = \hbar d\Phi = \hbar k \frac{dx}{dt} - \hbar\omega = p \frac{dx}{dt} - E = p\dot{x} - E = pu - H = L$ Legendre transformationUse Group velocity : $u = \frac{dx}{dt} = c \tanh \rho$ $dS = Ldt = \hbar d\Phi = \hbar k dx - \hbar\omega dt = p dx - H dt$ Poincare Invariant action differential $\frac{\partial S}{\partial x} = p$ $\frac{\partial S}{\partial t} = -H$ Hamilton-Jacobi equations

$$\begin{aligned} \overbrace{S = L = \hbar \dot{\Phi}}^{\text{Compare Lagrangian } L} &= -Mc^2 \sqrt{1 - \frac{u^2}{c^2}} &= -Mc^2 \operatorname{sech}\rho = -Mc^2 \cos \sigma \\ \text{with Hamiltonian } H = E \\ H = \hbar \omega = Mc^2 / \sqrt{1 - \frac{u^2}{c^2}} &= Mc^2 \cosh \rho = Mc^2 \operatorname{sec} \sigma \\ = Mc^2 \sqrt{1 + \sinh^2 \rho} = Mc^2 \sqrt{1 + (cp)^2} \end{aligned}$$

 $\begin{array}{c} h\upsilon_{A} = Mc^{2} = hc\kappa_{A} \\ h\upsilon_{phase} = E = h\upsilon_{A}\cosh\rho \\ hc\kappa_{phase} = cp = h\upsilon_{A}\sinh\rho \end{array} \begin{array}{c} Prior wave relations \\ \leftarrow \text{linear Hz} & \text{angular phasor} \rightarrow \end{array} \begin{array}{c} \hbar\omega_{A} = Mc^{2} = \hbar c\kappa_{A} \\ \hbar\omega_{phase} = E = \hbar\omega_{A}\cosh\rho \\ \hbar\omega_{phase} = E = \hbar\omega_{A}\cosh\rho \\ \hbar c\kappa_{phase} = cp = \hbar\omega_{A}\sinh\rho \end{array} \begin{array}{c} \hbar = \frac{h}{2\pi} \end{array}$



Rapidity ρ related to stellar aberration angle σ and L. C. Epstein's approach to relativity Longitudinal hyperbolic ρ -geometry connects to transverse circular σ -geometry "Occams Sword" and summary of 16 parameter functions of ρ and σ Applications to optical waveguide, spherical waves, and accelerator radiation

Derivation of relativistic quantum mechanics What's the matter with mass? Shining some light on the Elephant in the room Relativistic action and Lagrangian-Hamiltonian relations Poincare' and Hamilton-Jacobi equations

 Relativistic optical transitions and Compton recoil formulae Feynman diagram geometry Compton recoil related to rocket velocity formula Comparing 2nd-quantization "photon" number N and 1st-quantization wavenumber κ

Relawavity in accelerated frames

Relativistic optical transitions CMwBANG! Unit 8 Fig. 7.1 and 7.2



Relativistic optical transitions CMwBANG! Unit 8 Fig. 7.1 and 7.2



3-Level example













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2nd Quantization:

hv is actually hNv

 $(hv_{phase} = E = hv_A \cosh \rho)$ is actually $(hNv_{phase} = E_N = hNv_A \cosh \rho$ with quantum numbers)



2nd Quantization:

hv is actually hNv

 $(hv_{phase} = E = hv_A \cosh \rho)$ is actually $(hNv_{phase} = E_N = hNv_A \cosh \rho \quad (N=1,2,..))$





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Relawavity in accelerated frames

Laser up-tuning by Alice and down-tuning by Carla makes *g*-acceleration grid

Analysis of constant-g grid compared to zero-g Minkowsi grid Animation of mechanics and metrology of constant-g grid



Fig. 8.1 Optical wave frames by red-and-blue-chirped lasers (a) Varying acceleration (b) Constant g





Ecchure 31 Thm 12 10 2015

Review:Relawavity ρ functionsTwo famous onesExtremes and plot vs. ρ Doppler jeopardyGeometric mean and Relativistic hyperbolasAnimation of $e^{\rho}=2$ spacetime and per-spacetime plots

Rapidity ρ related to *stellar aberration angle* σ and L. C. Epstein's approach to relativity

Longitudinal hyperbolic ρ -geometry connects to transverse circular σ -geometry

"Occams Sword" and summary of 16 parameter functions of ρ and σ

Applications to optical waveguide, spherical waves, and accelerator radiation *Learning about* sin! and cos and...

Derivation of relativistic quantum mechanics

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Relawavity in accelerated frames









Fig. 8.2 Accelerated reference frames and their trajectories painted by chirped coherent light




time parameters	Per-space-t
$_{phase} = \lambda_A \operatorname{csch} \rho$	\mathcal{CK}_{phas}
$_{group} = \lambda_A \mathrm{sech}\rho$	<i>CK</i> _{grou}
$_{phase} = c\tau_A \mathrm{sech}\rho$	$oldsymbol{arphi}_{phase}$
$_{group} = c \tau_A \operatorname{csch} \rho$	$v_{_{group}}$

er-space-time parameters			
$c\kappa_{phase} = c\kappa_A \sinh\rho$			
$c\kappa_{group} = c\kappa_A \cosh\rho$			
$v_{phase} = v_A \cosh \rho$			
$v_{group} = v_A \sinh \rho$			

group	$b_{\scriptscriptstyle RED}^{\scriptscriptstyle Doppler}$	V _{group}	$rac{arphi_{group}}{arphi_{A}}$	$rac{\lambda_{group}}{\lambda_A}$	$rac{\kappa_{group}}{\kappa_A}$	$rac{{{ au }_{group}}}{{{ au }_{A}}}$	$rac{V_{phase}}{c}$	$b_{\scriptscriptstyle BLUE}^{\scriptscriptstyle Doppler}$
phase	$rac{1}{b_{BLUE}^{Doppler}}$	$\frac{c}{V_{phase}}$	$\frac{\kappa_{_{phase}}}{\kappa_{_A}}$	$rac{{m au}_{_{phase}}}{{m au}_{_A}}$	$rac{m{v}_{_{phase}}}{m{v}_{_A}}$	$rac{\lambda_{_{phase}}}{\lambda_{_A}}$	$\frac{c}{V_{group}}$	$\frac{1}{b_{\tiny RED}^{\tiny Doppler}}$
rapidity ρ	$e^{-\rho}$	$\tanh ho$	$\sinh ho$	$\operatorname{sech} \rho$	$\cosh ho$	$\mathrm{csch} ho$	$\mathrm{coth}\rho$	$e^{+ ho}$
stellar \forall angle σ	$1/e^{+\rho}$	$\sin \sigma$	$\tan \sigma$	$\cos \sigma$	$\sec\sigma$	$\cot \sigma$	$\csc\sigma$	1/ <i>e^{-p}</i>
$\beta \equiv \frac{u}{c}$	$\sqrt{\frac{1-\beta}{1+\beta}}$	<u>β</u> 1	$\frac{1}{\sqrt{\beta^{-2}-1}}$	$\frac{\sqrt{1-\beta^2}}{1}$	$\frac{1}{\sqrt{1-\beta^2}}$	$\frac{\sqrt{\beta^{-2}-1}}{1}$	$\frac{1}{\beta}$	$\sqrt{\frac{1+\beta}{1-\beta}}$
value for $\beta=3/5$	$\frac{1}{2} = 0.5$	$\frac{3}{5} = 0.6$	$\frac{3}{4}$ =0.75	$\frac{4}{5} = 0.80$	$\frac{5}{4}$ =1.25	$\frac{4}{3}$ =1.33	$\frac{5}{3}$ =1.67	$\frac{2}{1} = 2.0$
effects	$b_{\scriptscriptstyle RED}^{\scriptscriptstyle Doppler}$	V_{group}	past-future asymmetry (off-diagonal Lorentz-transform)	x-contraction ^(Lorentz) τ_{phase} -contraction	t-dilation ^(Einstein) v _{phase} -dilation (on-diagonal Lorentz-transform)	inverse asymmetry	V _{phase}	$b_{\scriptscriptstyle BLUE}^{\scriptscriptstyle Doppler}$





